

Table 2.1 Continued

ROTATION SEQUENCE	$\beta$	$\beta^{-1}$	ANGLES AS FUNCTIONS OF DIRECTION COSINES	SINGULAR AT $\alpha_2 =$
2-1-3	$\begin{bmatrix} c_2 s_3 & c_3 & 0 \\ c_2 c_3 & -s_3 & 0 \\ -s_2 & 0 & 1 \end{bmatrix}$	$\frac{1}{c_2} \begin{bmatrix} s_3 & c_3 & 0 \\ c_2 c_3 & -c_2 s_3 & 0 \\ s_2 s_3 & s_2 c_3 & c_2 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(C_{31}/C_{33})$ $\alpha_2 = s^{-1}(-C_{32})$ $\alpha_3 = \epsilon^{-1}(C_{12}/C_{22})$	$\pi/2$
2-3-1	$\begin{bmatrix} s_2 & 0 & 1 \\ c_2 c_3 & s_3 & 0 \\ -c_2 s_3 & c_3 & 0 \end{bmatrix}$	$\frac{1}{c_2} \begin{bmatrix} 0 & c_3 & -s_3 \\ 0 & -c_2 s_3 & c_2 c_3 \\ c_2 & -s_2 c_3 & s_2 s_3 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(-C_{13}/C_{11})$ $\alpha_2 = s^{-1}(C_{12})$ $\alpha_3 = \epsilon^{-1}(-C_{32}/C_{22})$	$\pi/2$
2-3-2	$\begin{bmatrix} s_2 c_3 & -s_3 & 0 \\ c_2 & 0 & 1 \\ s_2 s_3 & c_3 & 0 \end{bmatrix}$	$\frac{1}{s_2} \begin{bmatrix} c_3 & 0 & s_3 \\ -s_2 s_3 & 0 & s_2 c_3 \\ -c_2 c_3 & s_2 & -c_2 s_3 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(C_{23}/-C_{21})$ $\alpha_2 = \epsilon^{-1}(C_{22})$ $\alpha_3 = \epsilon^{-1}(C_{12}/C_{12})$	$0, \pi$
3-1-2	$\begin{bmatrix} -c_2 s_3 & c_3 & 0 \\ s_2 & 0 & 1 \\ c_2 c_3 & s_3 & 0 \end{bmatrix}$	$\frac{1}{c_2} \begin{bmatrix} -s_3 & 0 & c_3 \\ c_2 c_3 & 0 & c_2 s_3 \\ s_2 s_3 & c_2 & -s_2 c_3 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(-C_{21}/C_{22})$ $\alpha_2 = s^{-1}(C_{23})$ $\alpha_3 = \epsilon^{-1}(-C_{13}/C_{33})$	$\pi/2$
3-1-3	$\begin{bmatrix} s_2 s_3 & c_3 & 0 \\ s_2 c_3 & -s_3 & 0 \\ c_2 & 0 & 1 \end{bmatrix}$	$\frac{1}{s_2} \begin{bmatrix} s_3 & c_3 & 0 \\ c_3 s_2 & -s_3 s_2 & 0 \\ -s_3 c_2 & -c_3 c_2 & s_2 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(C_{31}/-C_{32})$ $\alpha_2 = \epsilon^{-1}(C_{33})$ $\alpha_3 = \epsilon^{-1}(C_{13}/C_{23})$	$0, \pi$
3-2-1	$\begin{bmatrix} -s_2 & 0 & 1 \\ c_2 s_3 & c_3 & 0 \\ -c_2 c_3 & -s_3 & 0 \end{bmatrix}$	$\frac{1}{c_2} \begin{bmatrix} 0 & s_3 & c_3 \\ 0 & c_2 c_3 & -c_2 s_3 \\ c_2 & s_2 s_3 & s_2 c_3 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(C_{12}/C_{11})$ $\alpha_2 = s^{-1}(-C_{13})$ $\alpha_3 = \epsilon^{-1}(C_{23}/C_{33})$	$\pi/2$
3-2-3	$\begin{bmatrix} -s_2 c_3 & s_3 & 0 \\ s_2 s_3 & c_3 & 0 \\ c_2 & 0 & 1 \end{bmatrix}$	$\frac{1}{s_2} \begin{bmatrix} -c_3 & s_3 & 0 \\ s_2 s_3 & s_2 c_3 & 0 \\ c_2 c_3 & -c_2 s_3 & s_2 \end{bmatrix}$	$\alpha_1 = \epsilon^{-1}(C_{32}/C_{31})$ $\alpha_2 = \epsilon^{-1}(C_{33})$ $\alpha_3 = \epsilon^{-1}(C_{23}/-C_{13})$	$0, \pi$

singularity. A moving frame with small Euler angles 2.1 reveals any indices (i.e., describing small or Euler angles taken as typical reference axes. In numerical situations it is a problem for the four Euler representations. For the four Euler representations, the singularities are:

singularity. A very familiar device in space vehicle dynamics is to define a moving frame which is the "nominal" or desired motion, then introduce three small Euler angles to describe departure motion. A casual inspection of Table 2.1 reveals any of the six  $\alpha$ - $\beta$ - $\gamma$  Euler angle sets which have *non-repeated indices* (i.e., 1-2-3, 3-2-1, 2-3-1, 3-1-2, 2-1-3, 1-3-2) are well-suited for describing small departure motions. The most common ("yaw, pitch, roll") set of Euler angles for aircraft and spacecraft applications is the 3-2-1 sequence, taken as typically small displacements from a moving "local vertical" set of reference axes. So long as these angles remain small, of course, no analytical or numerical difficulty associated with the geometric singularities will occur.

It is a remarkable truth that the corresponding kinematic relationships for the four Euler (quaternion) parameters of Section 2.6 are *rigorously*, and *universally* linearly related to the angular velocity components, for all orientations. These orientation parameters are closely related to Euler's principal rotation theorem developed in the following section.

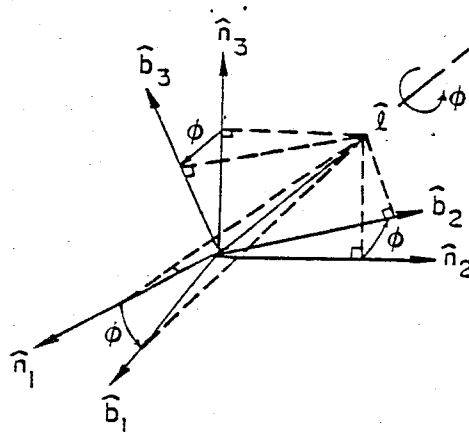


Figure 2.6 Euler's Principal Rotation

## 2.5 EULER'S PRINCIPAL ROTATION THEOREM

Euler (refs. 5, 6, 7) is generally credited with being responsible for the Principal Rotation Theorem:

A rigid body can be brought from an arbitrary initial orientation to an arbitrary final orientation by a single rotation of the body through a *principal angle* ( $\phi$ ) about a *principal line* ( $\hat{\mathbf{l}}$ ); the principal line being a judicious axis fixed in the body and fixed in space.

Letting (see Figure 2.6) the body fixed axes  $\{\hat{\mathbf{b}}\}$  be "initially" coincident with fixed axes  $\{\hat{\mathbf{n}}\}$ , we can use Euler's principal rotation theorem to develop several elegant parameterizations of the direction cosine matrix  $[C]$  defining  $\{\hat{\mathbf{b}}\}$ 's instantaneous angular position in the sense

$$\{\hat{\mathbf{b}}\} = [C]\{\hat{\mathbf{n}}\}.$$

Let us denote the  $\{\hat{\mathbf{b}}\}$  and  $\{\hat{\mathbf{n}}\}$  components of  $\hat{\mathbf{l}}$  as

$$\hat{\mathbf{l}} = l_{b1}\hat{\mathbf{b}}_1 + l_{b2}\hat{\mathbf{b}}_2 + l_{b3}\hat{\mathbf{b}}_3 \quad (2.46a)$$

and

$$\hat{\mathbf{l}} = l_{n1}\hat{\mathbf{n}}_1 + l_{n2}\hat{\mathbf{n}}_2 + l_{n3}\hat{\mathbf{n}}_3 \quad (2.46b)$$

As a direct consequence of the fact that  $l_{bi}$  and  $l_{ni}$  are constants (during a rotation  $\phi$  about a fixed  $\hat{\mathbf{l}}$ ) and  $\hat{\mathbf{b}}_i$  are "initially" coincident with  $\hat{\mathbf{n}}_i$ , we see that  $l_{ni} = l_{bi} = l_i$ , for  $i = 1, 2, 3$ . According to Eq. 2.15, we have

$$\begin{Bmatrix} l_{b1} \\ l_{b2} \\ l_{b3} \end{Bmatrix} = [C] \begin{Bmatrix} l_{n1} \\ l_{n2} \\ l_{n3} \end{Bmatrix} \quad (2.47)$$

but since  $l_{bi} = l_{ni} = l_i$ , we have

$$\begin{Bmatrix} l_1 \\ l_2 \\ l_3 \end{Bmatrix} = [C] \begin{Bmatrix} l_1 \\ l_2 \\ l_3 \end{Bmatrix} \quad (2.48)$$

Careful inspection of Eq. 2.48 reveals the truth that  $\hat{\mathbf{l}}$  exists if and only if

$[C]$  has an eigenvalue of  $+1$ ; in which case  $\hat{z}$  is the corresponding unit eigenvector of  $[C]$ . Goldstein (ref. 3) proves that all "proper" direction cosine matrices (those corresponding to displacement of right-handed axes imbedded in a rigid body) do in fact have an eigenvalue of  $+1$ ; the eigenvalue and corresponding eigenvector are unique (to within a sign on  $\hat{z}$  and  $\phi$ ) except for the case of zero angular displacement. The lack of sign uniqueness, as will be evident, does not cause a practical difficulty.

Since Euler's theorem reduces the general angular displacement to a single rotation about a fixed line, we can make immediate use of the developments in Section 2.3 to parameterize the direction cosine matrix in terms of  $\hat{z}$  and  $\phi$ . Specifically, if we take  $r = \hat{n}_i$  and  $r' = \hat{n}_i \equiv \hat{b}_i$  in the general Eq. 2.29, we obtain

$$\hat{b}_i = (1 - \cos\phi)(\hat{z} \cdot \hat{n}_i)\hat{z} + \cos\phi\hat{n}_i + \sin\phi(\hat{z} \times \hat{n}_i), \quad i = 1, 2, 3 \quad (2.49)$$

If we substitute  $\hat{z} = e_1\hat{n}_1 + e_2\hat{n}_2 + e_3\hat{n}_3$  and carry out the implied algebra in Eq. 2.49, we immediately obtain Eq. 2.11 with

$$[C] = \begin{bmatrix} e_1^2(1-\cos\phi)+\cos\phi & e_1e_2(1-\cos\phi)+e_3\sin\phi & e_1e_3(1-\cos\phi)-e_2\sin\phi \\ e_2e_1(1-\cos\phi)-e_3\sin\phi & e_2^2(1-\cos\phi)+\cos\phi & e_2e_3(1-\cos\phi)+e_1\sin\phi \\ e_3e_1(1-\cos\phi)+e_2\sin\phi & e_3e_2(1-\cos\phi)-e_1\sin\phi & e_3^2(1-\cos\phi)+\cos\phi \end{bmatrix} \quad (2.50)$$

Since  $e_1^2 + e_2^2 + e_3^2 = 1$ , we have  $[C(e_1, e_2, e_3, \phi)]$ , but only three degrees of freedom, as expected. Notice we can verify immediately (by summing the trace of Eq. 2.50) that

$$\cos\phi = \frac{1}{2} (C_{11} + C_{22} + C_{33} - 1) \quad (2.51)$$

and, by differencing the symmetric elements, we see that

$$2e_3\sin\phi = C_{12} - C_{21}$$

$$2e_2\sin\phi = C_{31} - C_{13}$$

$$2e_1\sin\phi = C_{23} - C_{32}$$

Multiplying the above three equations by  $e_3, e_2, e_1$ , respectively (and making use of  $\sum_{i=1}^3 e_i^2 = 1$ ), we have

$$\sin \phi = \frac{1}{2} [a_3(C_{12} - C_{21}) + a_2(C_{31} - C_{13}) + a_1(C_{23} - C_{32})] \quad (2.52)$$

Thus, given the direction cosines  $[C]$ , we can solve for  $\hat{i}$  from Eq. 2.48, normalized so that  $\hat{i}$  is a unit vector, and  $\phi$  from Eqs. 2.51 and 2.52. Notice that the sign choice on  $a_i$  correctly affects the quadrant of  $\phi$  in Eq. 2.52. Clearly a positive rotation about  $\pm \hat{i}$  is equivalent to a negative rotation about  $\mp \hat{i}$ .

## 2.6 EULER PARAMETERS

Following Euler, we define the four *Euler Parameters* in terms of the *Principal Line* ( $\hat{i}$ ) and *Principal Angle* ( $\phi$ ) as

$$\begin{aligned} a_0 &= \cos \frac{\phi}{2} \\ a_1 &= i_1 \sin \frac{\phi}{2} \\ a_2 &= i_2 \sin \frac{\phi}{2} \\ a_3 &= i_3 \sin \frac{\phi}{2} \end{aligned} \quad (2.53)$$

It is obvious that the four  $a_i$ 's satisfy the constraint equation

$$a_0^2 + a_1^2 + a_2^2 + a_3^2 = 1. \quad (2.54)$$

If we make use of the half-angle identities

$$\begin{aligned} \sin \phi &= 2 \sin \frac{\phi}{2} \cos \frac{\phi}{2} \\ \cos \phi &= 2 \cos^2 \frac{\phi}{2} - 1 \end{aligned}$$

and Eq. 2.54, the direction cosine matrix of Eq. 2.50 can be parameterized as a function of the Euler Parameters:

$$[C(a)] = \begin{bmatrix} a_0^2 + a_1^2 - a_2^2 - a_3^2 & 2(a_1 a_2 + a_0 a_3) & 2(a_1 a_3 - a_0 a_2) \\ 2(a_1 a_2 - a_0 a_3) & a_0^2 - a_1^2 + a_2^2 - a_3^2 & 2(a_2 a_3 + a_0 a_1) \\ 2(a_1 a_3 + a_0 a_2) & 2(a_2 a_3 - a_0 a_1) & a_0^2 - a_1^2 - a_2^2 + a_3^2 \end{bmatrix} \quad (2.55)$$

The inverse relationships ( $a$ 's in terms of the elements of  $[C]$ ) can be deduced from Eq. 2.55:

$$\begin{aligned} a_0 &= \pm \frac{1}{2} (C_{11} + C_{22} + C_{33} + 1)^{1/2} \\ a_1 &= (C_{23} - C_{32}) / 4a_0 \end{aligned}$$

## Section 2.6

$$a_2 = (C_{13} - C_{31}) / 4a_0$$

$$a_3 = (C_{12} - C_{21}) / 4a_0$$

The first of these

no loss of generality

changing all the

cosine matrix.

positive angle

2.56 contain

computational

$$a_0^2 = \frac{1}{4}$$

$$a_1^2 = \frac{1}{4}$$

$$a_2^2 = \frac{1}{4}$$

$$a_3^2 = \frac{1}{4}$$

where Stanley

$$T = \text{Trace}$$

$$C_{00} = T$$

and selects the

largest absolute

obtained by

equations (costs

$$a_0 a_1 =$$

$$a_0 a_2 =$$

$$a_0 a_3 =$$

$$a_2 a_3 =$$

$$a_3 a_1 =$$

$$a_1 a_2 =$$

(2.52)

from Eq. 2.48.

1 2.52. Notice

• in Eq. 2.52.

ative rotation

terms of the

(2.53)

(2.54)

terized as a

2.55)

be deduced

$$a_2 = (C_{31} - C_{13})/4a_0$$

$$a_3 = (C_{12} - C_{21})/4a_0$$

(2.56)

The first of these equations has an apparent sign ambiguity. However, there is no loss of generality in adopting the positive sign, since it is evident that changing all four signs of the  $a_i$ 's in Eq. 2.55 does not change the direction cosine matrix. This is another reflection of the equivalence of rotating by a positive angle  $\phi$  about  $\pm \hat{e}$  to a  $-\phi$  rotation about  $\mp \hat{e}$ . It is evident that Eqs. 2.56 contain a 0/0 indeterminacy whenever  $a_0$  goes through zero; a computationally superior algorithm has been developed by Stanley (ref. 8):

$$a_0^2 = \frac{1}{4} (1 + 2 C_{00} - T)$$

$$a_1^2 = \frac{1}{4} (1 + 2 C_{11} - T)$$

$$a_2^2 = \frac{1}{4} (1 + 2 C_{22} - T)$$

$$a_3^2 = \frac{1}{4} (1 + 2 C_{33} - T)$$

(2.57)

where Stanley defines

$$T = \text{Trace } [C] = C_{11} + C_{22} + C_{33}$$

(2.58a)

$$C_{00} = T$$

(2.58b)

and selects (for division by) the  $a_i$  computed from Eq. 2.57 which has the largest absolute value assuming it to be positive; the other three  $a_j$ 's can be obtained by dividing  $a_i$  into the appropriate three of the following six equations (obtained by differencing and summing the symmetric elements of  $[C]$ ):

$$a_0 a_1 = (C_{23} - C_{32})/4$$

$$a_0 a_2 = (C_{31} - C_{13})/4$$

$$a_0 a_3 = (C_{12} - C_{21})/4$$

$$a_2 a_3 = (C_{23} + C_{32})/4$$

$$a_3 a_1 = (C_{31} + C_{13})/4$$

$$a_1 a_2 = (C_{12} + C_{21})/4$$

(2.59)

The Euler parameters have many elegant properties; we shall develop the more important relationships in the following discussion.

One important property is the simple fashion in which Euler parameters of sequential rotations parameterize an equivalent single rotation. To develop these results, let  $\{\hat{b}\}$ ,  $\{\hat{b}'\}$ , and  $\{\hat{b}''\}$  be three arbitrary positions of a triad of unit vectors; the direction cosine matrix  $[C]$  defining the relative orientations of  $\{\hat{b}\}$ ,  $\{\hat{b}'\}$ , and  $\{\hat{b}''\}$  can be parameterized in terms of three sets of Euler parameters as

$$\{\hat{b}'\} = [C(a')]\{\hat{b}\} \quad (2.60a)$$

$$\{\hat{b}''\} = [C(a'')]\{\hat{b}'\} \quad (2.60b)$$

$$\{\hat{b}''\} = [C(a)]\{\hat{b}\} \quad (2.60c)$$

We seek a relationship of the form

$$a_i = \text{function}(a'_0, a'_1, a'_2, a'_3; a''_0, a''_1, a''_2, a''_3) \quad , \quad i = 0, 1, 2, 3 \quad (2.61)$$

relating the three sets of Euler parameters. From Eq. 2.60, it is obvious that

$$[C(a)] = [C(a'')][C(a')] \quad (2.62)$$

Direct substitution of Eq. 2.55 into Eq. 2.62 and equating the corresponding elements leads to the most elegant result

$$\begin{pmatrix} a_0 \\ a_1 \\ a_2 \\ a_3 \end{pmatrix} = \begin{bmatrix} a''_0 & -a''_1 & -a''_2 & -a''_3 \\ a''_1 & a''_0 & a''_3 & -a''_2 \\ a''_2 & -a''_3 & a''_0 & a''_1 \\ a''_3 & a''_2 & -a''_1 & a''_0 \end{bmatrix} \begin{pmatrix} a'_0 \\ a'_1 \\ a'_2 \\ a'_3 \end{pmatrix} \quad (2.63)$$

or, by transmutation of Eq. 2.63

$$\begin{pmatrix} a'_0 \\ a'_1 \\ a'_2 \\ a'_3 \end{pmatrix} = \begin{bmatrix} a'_0 & -a'_1 & -a'_2 & -a'_3 \\ a'_1 & a'_0 & -a'_3 & a'_2 \\ a'_2 & a'_3 & a'_0 & -a'_1 \\ a'_3 & -a'_2 & a'_1 & a'_0 \end{bmatrix} \begin{pmatrix} a''_0 \\ a''_1 \\ a''_2 \\ a''_3 \end{pmatrix} \quad (2.64)$$

It is obvious by inspection that the coefficient matrices in Eqs. 2.63 and 2.64 are orthogonal; thus, any set of  $a$ 's can be solved universally as a simple,

## Section 2.6

nonsingular, by properties of alternative transformations  $\alpha = \cos^{-1}$   $\beta = \tan^{-1}$   $\gamma = \tan^{-1}$

where the  $C_{ij}$  are the matrix multi-

$C(a, \dots)$

using Eq. 2.55

2.66. These are

the  $C_{ij}$  are the

directional cosines

of Eqs. 2.55 and 2.56

and as to the

and as to the

The transform

and transform

then apply, the

to the partial

and as to the

expression in

to derive the

expressions that

through the line

To derive

parameters, as

nonsingular, bilinear combination of the other two. The successive rotation properties of the Euler parameters are thus quite elegant compared to any alternative parameterizations of  $[C]$ . For example, the analogous transformations for 3-1-3 ( $\phi, \theta, \psi$ ) Euler angles can be verified to be

$$\begin{aligned} \phi &= \cos^{-1}(C_{33}) \\ \theta &= \tan^{-1}(C_{31}/-C_{32}) \\ \psi &= \tan^{-1}(C_{13}/C_{23}) \end{aligned} \quad (2.65)$$

where the  $C_{ij}$  are the lengthy transcendental functions obtained by carrying out the matrix multiplication

$$[C(\phi, \theta, \psi)] = [C(\phi'', \theta'', \psi'')][C(\phi', \theta', \psi')] \quad (2.66)$$

using Eq. 2.34 to parameterize the two matrices on the right side of Eq. 2.66. These nine equations can be inverted for  $(\phi, \theta, \psi)$  as in Eq. 2.65, where the  $C_{ij}$  are the functions of  $(\phi', \theta', \psi', \phi'', \theta'', \psi'')$  obtained from Eq. 2.66. It is a trivial observation that Eqs. 2.63 and 2.64 are vastly more attractive than Eqs. 2.65 and 2.66!

In applications, we often require a transformation from a set of Euler angles to the corresponding Euler parameters, and the inverse transformation. The transformation from Euler parameters into Euler angles is entirely straightforward; we simply calculate the direction cosine matrix from Eq. 2.55, then employ the inverse trigonometric relationships in Table 2.1, corresponding to the particular Euler angle rotation sequence. The transformation from Euler angles to Euler parameters can proceed by calculating numerical values for the direction cosines from Eq. 2.36, then using Eqs. 2.56, 2.57, or 2.59 to calculate the  $a$ 's. However it is possible to derive very compact analytical expressions that are more efficient and avoid the branching logic of the path through the direction cosines.

To develop the direct transformation from Euler angles to Euler parameters, we first note from Eq. 2.53 that the elementary Euler angle



rotations have the corresponding Euler parameter values

Rotation about a "1" axis:  $a_0 = \cos \frac{1}{2} \alpha$

$$a_1 = \sin \frac{1}{2} \alpha$$

$$a_2 = a_3 = 0$$

Rotation about a "2" axis:  $a_0 = \cos \frac{1}{2} \beta$

$$a_2 = \sin \frac{1}{2} \beta$$

$$a_1 = a_3 = 0$$

Rotation about a "3" axis:  $a_0 = \cos \frac{1}{2} \gamma$

$$a_3 = \sin \frac{1}{2} \gamma$$

$$a_1 = a_2 = 0$$

If we let  $c_i = \cos \frac{1}{2} \theta_i$  and  $s_i = \sin \frac{1}{2} \theta_i$ , for the 3-1-3 Euler angles  $(\theta_1, \theta_2, \theta_3)$ , we can write from Eq. 2.63 the equivalent Euler parameters (replacing the above three elementary rotations) as

$$\begin{pmatrix} a_0 \\ a_1 \\ a_2 \\ a_3 \end{pmatrix} = \begin{bmatrix} c_3 & 0 & 0 & -s_3 \\ 0 & c_3 & s_3 & 0 \\ 0 & -s_3 & c_3 & 0 \\ s_3 & 0 & 0 & c_3 \end{bmatrix} \begin{bmatrix} c_2 & -s_2 & 0 & 0 \\ s_2 & c_2 & 0 & 0 \\ 0 & 0 & c_2 & s_2 \\ 0 & 0 & -s_2 & c_2 \end{bmatrix} \begin{pmatrix} c_1 \\ 0 \\ 0 \\ s_1 \end{pmatrix}$$

Upon carrying out the matrix multiplications we obtain

$$a_0 = \cos \frac{1}{2} \theta_3 \cos \frac{1}{2} \theta_2 \cos \frac{1}{2} \theta_1 - \sin \frac{1}{2} \theta_3 \cos \frac{1}{2} \theta_2 \sin \frac{1}{2} \theta_1$$

$$a_1 = \cos \frac{1}{2} \theta_3 \sin \frac{1}{2} \theta_2 \cos \frac{1}{2} \theta_1 + \sin \frac{1}{2} \theta_3 \sin \frac{1}{2} \theta_2 \sin \frac{1}{2} \theta_1$$

$$a_2 = -\sin \frac{1}{2} \theta_3 \sin \frac{1}{2} \theta_2 \cos \frac{1}{2} \theta_1 + \cos \frac{1}{2} \theta_3 \sin \frac{1}{2} \theta_2 \sin \frac{1}{2} \theta_1$$

$$a_3 = \sin \frac{1}{2} \theta_3 \cos \frac{1}{2} \theta_2 \cos \frac{1}{2} \theta_1 + \cos \frac{1}{2} \theta_3 \cos \frac{1}{2} \theta_2 \sin \frac{1}{2} \theta_1$$

## Section 2.6

Finally, using

$$a_0 = \cos$$

$$a_1 = \sin$$

$$a_2 = \sin$$

$$a_3 = \cos$$

The above can be

transformations

for ease in

transformation

$$\begin{pmatrix} a_0 \\ a_1 \\ a_2 \\ a_3 \end{pmatrix} =$$

where

$$[R_{123}] =$$

$$R_1 = \begin{bmatrix} \cdot \\ \cdot \\ \cdot \\ \cdot \end{bmatrix}$$

$$R_2 = \begin{bmatrix} \cdot \\ \cdot \\ \cdot \\ \cdot \end{bmatrix}$$

Eq. 2.68 was derived  
Dec. 1973 (ref.)

Finally, using trigonometric identities to simplify to the final transformation

$$\begin{aligned} \theta_0 &= \cos \frac{1}{2} \theta_2 \cos \frac{1}{2} (\theta_1 + \theta_3) \\ \theta_1 &= \sin \frac{1}{2} \theta_2 \cos \frac{1}{2} (\theta_1 - \theta_3) \\ \theta_2 &= \sin \frac{1}{2} \theta_2 \sin \frac{1}{2} (\theta_1 - \theta_3) \\ \theta_3 &= \cos \frac{1}{2} \theta_2 \sin \frac{1}{2} (\theta_1 + \theta_3) \end{aligned} \tag{2.67}$$

The above can be paralleled for the other 11 sets of Euler angle sequences; the transformations for all 12 sets of Euler angles are summarized in Table 2.2. For ease in computer programming, we can employ the most useful universal transformation

$$\begin{Bmatrix} \theta_0 \\ \theta_1 \\ \theta_2 \\ \theta_3 \end{Bmatrix} = [R_{\alpha\beta\gamma}] \begin{Bmatrix} 1 \\ 0 \\ 0 \\ 0 \end{Bmatrix} \tag{2.68}^*$$

where

$$\begin{aligned} [R_{\alpha\beta\gamma}] &= [\cos \frac{1}{2} \theta_3 R_0 + \sin \frac{1}{2} \theta_3 R_Y] \cdot \\ &\quad [\cos \frac{1}{2} \theta_2 R_0 + \sin \frac{1}{2} \theta_2 R_\beta] \cdot \\ &\quad [\cos \frac{1}{2} \theta_1 R_0 + \sin \frac{1}{2} \theta_1 R_\alpha] \end{aligned} \tag{2.69}$$

$$\begin{aligned} R_0 &= \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}, \quad R_1 = \begin{bmatrix} 0 & -1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 \\ 0 & 0 & -1 & 0 \end{bmatrix} \\ R_2 &= \begin{bmatrix} 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \\ 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{bmatrix}, \quad R_3 = \begin{bmatrix} 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ 1 & 0 & 0 & 0 \end{bmatrix} \end{aligned}$$

\*Eq. 2.68 was developed by H. S. Morton, Jr. of the University of Virginia, Dec. 1973 (ref. 4).

TABLE 2.2

TRANSFORMATION FROM THE TWELVE SETS OF EULER ANGLES TO EULER PARAMETERS\*

$\alpha-\beta-\gamma$	$\theta_0$	$\theta_1$	$\theta_2$	$\theta_3$
1-2-1	$c \frac{\theta_2}{2} c(\frac{\theta_1 + \theta_3}{2})$	$c \frac{\theta_2}{2} s(\frac{\theta_1 + \theta_3}{2})$	$s \frac{\theta_2}{2} c(\frac{\theta_1 - \theta_3}{2})$	$s \frac{\theta_2}{2} s(\frac{\theta_1 - \theta_3}{2})$
2-3-2	$c \frac{\theta_2}{2} c(\frac{\theta_1 + \theta_3}{2})$	$s \frac{\theta_2}{2} s(\frac{\theta_1 - \theta_3}{2})$	$c \frac{\theta_2}{2} s(\frac{\theta_1 + \theta_3}{2})$	$s \frac{\theta_2}{2} c(\frac{\theta_1 - \theta_3}{2})$
3-1-3	$c \frac{\theta_2}{2} c(\frac{\theta_1 + \theta_3}{2})$	$s \frac{\theta_2}{2} c(\frac{\theta_1 - \theta_3}{2})$	$s \frac{\theta_2}{2} s(\frac{\theta_1 - \theta_3}{2})$	$c \frac{\theta_2}{2} s(\frac{\theta_1 + \theta_3}{2})$
1-3-1	$c \frac{\theta_2}{2} c(\frac{\theta_3 + \theta_1}{2})$	$c \frac{\theta_2}{2} s(\frac{\theta_3 + \theta_1}{2})$	$s \frac{\theta_2}{2} s(\frac{\theta_3 - \theta_1}{2})$	$s \frac{\theta_2}{2} c(\frac{\theta_3 - \theta_1}{2})$
2-1-2	$c \frac{\theta_2}{2} c(\frac{\theta_3 + \theta_1}{2})$	$s \frac{\theta_2}{2} c(\frac{\theta_3 - \theta_1}{2})$	$c \frac{\theta_2}{2} s(\frac{\theta_3 + \theta_1}{2})$	$s \frac{\theta_2}{2} s(\frac{\theta_3 - \theta_1}{2})$
3-2-3	$c \frac{\theta_2}{2} c(\frac{\theta_3 + \theta_1}{2})$	$s \frac{\theta_2}{2} s(\frac{\theta_3 - \theta_1}{2})$	$s \frac{\theta_2}{2} c(\frac{\theta_3 - \theta_1}{2})$	$c \frac{\theta_2}{2} s(\frac{\theta_3 + \theta_1}{2})$
1-2-3	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$s \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$
	$-s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$+c \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-s \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$+s \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$
2-3-1	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$s \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$
	$-s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$+s \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$+c \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-s \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$
3-1-2	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$s \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$
	$-s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-s \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$+s \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$+c \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$
1-3-2	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$s \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$
	$+s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-c \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$+s \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$
2-1-3	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$s \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$
	$+s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$+s \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-c \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-s \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$
3-2-1	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$c \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$s \frac{\theta_1}{2} c \frac{\theta_2}{2} c \frac{\theta_3}{2}$
	$+s \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-s \frac{\theta_1}{2} s \frac{\theta_2}{2} c \frac{\theta_3}{2}$	$+s \frac{\theta_1}{2} c \frac{\theta_2}{2} s \frac{\theta_3}{2}$	$-c \frac{\theta_1}{2} s \frac{\theta_2}{2} s \frac{\theta_3}{2}$

\*C = cos, s = sin

PARAMETERS\*

Equation 2.68 captures all 12 transformations of Table 2.2 in a general form suitable for a universal computer algorithm.

The above properties are significant, attractive features of the Euler parameters; however, the most impressive property is the kinematical differential equation which we now develop. We seek equations of the form

$\dot{a}_i = \text{function}_i(a_0, a_1, a_2, a_3, \omega_1, \omega_2, \omega_3)$ . These can be obtained by differentiating Eq. 2.56. For example, consider the derivation of the equation

for  $\dot{a}_0$ , by differentiating the first term of Eq. 2.56, we have

$$\dot{a}_0 = \frac{(C_{11} + C_{22} + C_{33})}{8a_0} \quad (2.70)$$

From Eq. 2.22, substitute for the  $C_{ij}$ 's

$$\begin{aligned} C_{11} &= \omega_3 C_{21} - \omega_2 C_{31} \\ C_{22} &= -\omega_3 C_{12} + \omega_1 C_{32} \\ C_{33} &= \omega_2 C_{13} - \omega_1 C_{23} \end{aligned} \quad (2.71)$$

so that Eq. 2.70 becomes

$$\dot{a}_0 = \frac{(C_{32} - C_{23})\omega_1 + (C_{13} - C_{31})\omega_2 + (C_{21} - C_{12})\omega_3}{8a_0} \quad (2.72)$$

and eliminating the  $C_{ij}$ 's in favor of the  $a_i$ 's using the first three terms of Eq. 2.59, we have the desired equation for  $\dot{a}_0$ .

$$\dot{a}_0 = -\frac{1}{2} (a_1 \omega_1 + a_2 \omega_2 + a_3 \omega_3) \quad (2.73)$$

Similarly, we can derive equations for  $\dot{a}_1$ ,  $\dot{a}_2$ ,  $\dot{a}_3$ ; the resulting four equations can be written in matrix form

$$\begin{pmatrix} \dot{a}_0 \\ \dot{a}_1 \\ \dot{a}_2 \\ \dot{a}_3 \end{pmatrix} = \frac{1}{2} \begin{bmatrix} 0 & -\omega_1 & -\omega_2 & -\omega_3 \\ \omega_1 & 0 & \omega_3 & -\omega_2 \\ \omega_2 & -\omega_3 & 0 & \omega_1 \\ \omega_3 & \omega_2 & -\omega_1 & 0 \end{bmatrix} \begin{pmatrix} a_0 \\ a_1 \\ a_2 \\ a_3 \end{pmatrix} \quad (2.74)$$

or, by transmutation of Eq. 2.74

$$\begin{pmatrix} \dot{\beta}_0 \\ \dot{\beta}_1 \\ \dot{\beta}_2 \\ \dot{\beta}_3 \end{pmatrix} = \frac{1}{2} \begin{bmatrix} \beta_0 & -\beta_1 & -\beta_2 & -\beta_3 \\ \beta_1 & \beta_0 & -\beta_3 & \beta_2 \\ \beta_2 & \beta_3 & \beta_0 & -\beta_1 \\ \beta_3 & -\beta_2 & \beta_1 & \beta_0 \end{bmatrix} \begin{pmatrix} 0 \\ \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix} \quad (2.75)$$

Eq. 2.74 or, alternatively, Eq. 2.75 are indeed useful results. Observe that the transformation matrix in Eq. 2.75 relating  $\dot{\beta}_i$ 's and  $\omega_j$ 's is orthogonal; therefore this relationship is universally nonsingular (whereas the corresponding kinematical equations for any three angle set is transcendental, nonlinear, and contains a 0/0 singularity). Further, note Eq. 2.74. For  $\omega_i(t)$  measured or integrated a priori, is a *rigorously linear* differential equation

$$\{\dot{\beta}\} = [\omega(t)]\{\beta\} \quad (2.76)$$

Since  $[\omega(t)]$  is skew symmetric (compare Eqs. 2.74 and 2.76), we can show that  $\|\beta\|^2 = 1$  is a rigorous integral of the solution. To see this, let

$$\|\beta\|^2 = \sum_{i=0}^3 \beta_i^2 = \{\beta\}^T \{\beta\} \quad (2.77)$$

then

$$\frac{d}{dt} \|\beta\|^2 = \{\dot{\beta}\}^T \{\beta\} + \{\beta\}^T \{\dot{\beta}\} \quad (2.78)$$

substitution of Eq. 2.76 into Eq. 2.78 yields

$$\frac{d}{dt} \|\beta\|^2 = \{\beta\}^T [\{\omega\}^T + \{\omega\}] \{\beta\} \quad (2.79)$$

and since  $\{\omega\}^T = -\{\omega\}$ , we see from Eq. 2.79 that  $\frac{d}{dt} \|\beta\|^2 = 0$ . Since  $\|\beta(t_0)\| = 1$  will be established by any valid choice of initial Euler parameters, we can see that any (accurate) solution of Eqs. 2.74, 2.75, or 2.76 will guarantee  $\|\beta(t)\| = 1$ . In fact, this is a standard (necessary) condition used to test numerical solutions of Eq. 2.75, to, for example, control step size.

The Euler parameters represent a fascinating example of *regularization* in mechanics. Through a judicious choice of coordinates, we are able to

eliminate the  
all possible  
dimensional  
parameters of

## 2.7 OTHER CP

Since  
the  
rotation  
matrix  
is  
orthogonal  
the  
Euler  
parameters

are

the

the

the

the

the

the

the

the

the

eliminate the singularities usually present. It is also a beautiful truth that all possible rotational motions correspond to a path on the surface of a four dimensional unit sphere. For certain cases, analytical solution for the Euler parameters are possible, see refs. 4 and 11.

### 2.7 OTHER ORIENTATION PARAMETERS

Aside from Euler angles, Euler parameters, and the direction cosines, there are an infinity of less commonly adopted possible descriptions of orientation. We summarize here the most prominent members of this large family of possibilities.

#### Rodriguez Parameters ( $\lambda_1, \lambda_2, \lambda_3$ )

These parameters (ref. 6) are intimately related to the principal rotation and, therefore, to the Euler parameters. In fact, the  $\lambda_i$ 's are simply

$$\begin{aligned}\lambda_1 &= a_1/a_0 = \epsilon_1 \tan \frac{1}{2} \phi \\ \lambda_2 &= a_2/a_0 = \epsilon_2 \tan \frac{1}{2} \phi\end{aligned}\quad (2.80)$$

$$\lambda_3 = a_3/a_0 = \epsilon_3 \tan \frac{1}{2} \phi$$

Clearly, these parameters have an unbounded behavior near  $\phi = (2n+1)\pi$  where  $n$  is integer and therefore are less attractive than the  $a$ 's themselves for most purposes. The geometric and kinematic relationships governing the  $\lambda_i$ 's can be immediately derived from the corresponding Euler parameter equations.

#### Cayley-Klein Parameters

These four complex parameters (ref. 7) are the combinations of the Euler parameters

$$\begin{aligned}\alpha &= a_0 + i a_3, \quad i^2 = -1 \\ \beta &= -a_2 + i a_1 \\ \gamma &= a_2 + i a_1 \\ \delta &= a_0 - i a_3\end{aligned}\quad (2.81)$$

with the inverse relations being

$$\begin{aligned}
 B_0 &= (\alpha + \delta)/2 \\
 B_1 &= -i(\beta + \gamma)/2 \\
 B_2 &= -(\beta - \gamma)/2 \\
 B_3 &= -i(\alpha - \delta)/2
 \end{aligned}
 \tag{2.82}$$

and the parameterization of the direction cosine matrix being

$$[C(\alpha, \beta, \gamma, \delta)] = \begin{bmatrix} (\alpha^2 - \beta^2 - \gamma^2 + \delta^2)/2 & i(-\alpha^2 + \beta^2 - \gamma^2 + \delta^2)/2 & (\beta\delta - \alpha\gamma) \\ i(\alpha^2 + \beta^2 - \gamma^2 - \delta^2)/2 & (\alpha^2 + \beta^2 + \gamma^2 + \delta^2)/2 & -i(\alpha\gamma + \beta\delta) \\ (\gamma\delta - \alpha\beta) & i(\alpha\beta + \gamma\delta) & (\alpha\delta + \beta\gamma) \end{bmatrix}
 \tag{2.83}$$

An alternate set of four complex Cayley-Klein-like parameters (having advantages for certain applications) is discussed in Ref. [11].

### Quaternions

Hamilton (ref. 7) developed quaternion algebra due to the intimate connection with Euler Parameters and rotation of vectors. Quaternions are four dimensional entities having one real and three complex parts.

$$q = q_0 + iq_1 + jq_2 + kq_3 \tag{2.84}$$

where

$$i \cdot i = -1, j \cdot j = -1, k \cdot k = -1$$

$$i \cdot j = k, j \cdot k = i, -k \cdot j = i, \text{ etc.}$$

For a unit quaternion  $\sum_{i=0}^3 q_i^2 = 1$ . From Ames and Murnahan's (ref. 10) development of quaternion algebra, it can be shown that the quaternion multiplication  $q = q''q'$  yields a new quaternion ( $q$ ) whose components relate to those of  $q''$  and  $q'$  as

$$\begin{pmatrix} q_0 \\ q_1 \\ q_2 \\ q_3 \end{pmatrix} = \begin{bmatrix} q_0'' & -q_1'' & -q_2'' & -q_3'' \\ q_1'' & q_0'' & -q_3'' & q_2'' \\ q_2'' & q_3'' & q_0'' & -q_1'' \\ q_3'' & -q_2'' & q_1'' & q_0'' \end{bmatrix} \begin{pmatrix} q_0' \\ q_1' \\ q_2' \\ q_3' \end{pmatrix}
 \tag{2.84}$$

Comparison of Eq. 2.84 with Eq. 2.63 allows us to identify

$$\begin{aligned} q_0 &= \beta_0 \\ q_i &= -\beta_i, \quad i = 1, 2, 3 \end{aligned} \quad (2.85)$$

Or, alternatively, we can consider the Euler parameters to be the elements of the quaternion

$$q = \beta_0 - i\beta_1 - j\beta_2 - k\beta_3. \quad (2.86)$$

Historically, quaternions played a "stepping stone" role en route to modern vector algebra and quaternion algebra is not presently widely used. Unfortunately, Hamilton's association of Euler parameters with quaternions served to make them more obscure when quaternion algebra fell largely into a "museum-piece" status. As we have seen in the foregoing, the basic kinematic and geometric relationships governing the Euler parameters can be easily developed without relying upon interpreting them as elements of a quaternion.

## REFERENCES

1. P. W. Likins, *Elements of Engineering Mechanics*, Ch. 3, McGraw-Hill, New York, 1973.
2. T. R. Kane, P. W. Likins, and D. A. Levinson, *SPACECRAFT DYNAMICS*, McGraw-Hill, New York, 1983.
3. H. Goldstein, *Classical Mechanics*, Addison Wesley, 1950.
4. H. S. Morton, Jr. and J. L. Junkins, *The Differential Equations of Rotational Motion*, Ch. 2, In Prep., Aug. 1986.
5. L. Euler, "Problema Algebraicum of Affectiones Prorsus Singulares Memorabile," *Nov. Comm. Petrop. T. SV1770*, p. 75; *Comm. Arith. Coll. T. I.*, pp. 427-443.
6. O. Rodrigues, "Des Lois Geometriques Qui Regissent Les Deplacements D'Un Systeme Solide Dans l'Espace, et de la Variation des Coordonnees Provenants de ces Deplacements Consideres Independamment des Causes Qui Peuvent les Produire," *LIUV, T. III*, pp. 380-440 (1840).
7. E. T. Whittaker, *Analytical Dynamics of Particles and Rigid Bodies*, Cambridge University Press, pp. 2-16, 1965 reprint.
8. W. S. Stanley, "Quaternion from Rotation Matrix," *AIAA J. Guidance and Control*, Vol. 1, No. 3, pp. 223-224, May 1978.



9. J. N. Blanton, Some New Results on the Free Motion of Triaxial Rigid Bodies. Ph.D. Dissertation, Univ. of Virginia, Aerospace Engineering, August 1976.
10. J. S. Ames and F. D. Murnahan, Theoretical Mechanics. Ginn & Co., 1929.
11. H. S. Morton, Jr., J. L. Junkins, and J. N. Blanton, "Analytical Solutions for Euler Parameters," Celestial Mechanics, Vol. 10, pp. 287-301, Sept. 1974.

CHAPTER 1

BASIC PRIN.