

Angular Momentum in Terms of Toroidal and Poloidal Stream Functions

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1 Prior Work Considered

In this set of notes, we consider 3D nonlinear dynamo simulations that explored magnetic-field amplitudes at a variety of rotation rates (i.e., enough to make scatter plots of various quantities versus Rossby number and thus theoretically address the activity-rotation relation). We aim to determine what region of parameter space has been explored by global models, and in particular, how the rotation-activity relation may or may not have been addressed. The works considered are:

[Christensen & Aubert \(2006\)](#),

[Christensen et al. \(2009\)](#),

[Strugarek et al. \(2017\)](#),

[Guerrero et al. \(2019\)](#),

[Brun et al. \(2022\)](#),

References

- Brun, A. S., Strugarek, A., Noraz, Q., Perri, B., Varela, J., Augustson, K., Charbonneau, P., & Toomre, J. 2022, “Powering stellar magnetism: Energy transfers in cyclic dynamos of Sun-like stars”, *Astrophys. J.*, 926, 21, doi: [10.3847/1538-4357/ac469b](#)
- Christensen, U. R., & Aubert, J. 2006, “Scaling properties of convection-driven dynamos in rotating spherical shells and application to planetary magnetic fields”, *Geophysical Journal International*, 166, 97, doi: [10.1111/j.1365-246x.2006.03009.x](#)
- Christensen, U. R., Holzwarth, V., & Reiners, A. 2009, “Energy flux determines magnetic field strength of planets and stars”, *Nat.*, 457, 167, doi: [10.1038/nature07626](#)

- Guerrero, G., Zaire, B., Smolarkiewicz, P. K., de Gouveia Dal Pino, E. M., Kosovichev, A. G., & Mansour, N. N. 2019, “What sets the magnetic field strength and cycle period in solar-type stars?”, *Astrophys. J.*, 880, 6, doi: [10.3847/1538-4357/ab224a](https://doi.org/10.3847/1538-4357/ab224a)
- Strugarek, A., Beaudoin, P., Charbonneau, P., Brun, A. S., & do Nascimento, J.-D. 2017, “Reconciling solar and stellar magnetic cycles with nonlinear dynamo simulations”, *Sci.*, 357, 185, doi: [10.1126/science.aal3999](https://doi.org/10.1126/science.aal3999)