Induction term in spherical coordinates

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1 The problem with the traditional induction terms

We consider the ideal (resistance-free) magnetohydrodynamic (MHD) induction equation:

$$\frac{\partial \boldsymbol{B}}{\partial t} = \nabla \times (\boldsymbol{u} \times \boldsymbol{B}) := \boldsymbol{I},\tag{1}$$

$$= \mathbf{B} \cdot \nabla \mathbf{u} - \mathbf{u} \cdot \nabla \mathbf{B} - (\nabla \cdot \mathbf{u}) \mathbf{B}, \tag{2}$$

where \boldsymbol{B} and \boldsymbol{u} are the vector magnetic and velocity fields (respectively) and \boldsymbol{I} is the vector induction. The three terms on the right-hand-side of Equation (2) are often interpreted as "shear," "advection," and "compression," respectively. However, this interpretation is problematic in general for two reasons:

- 1. The so-called shear and compression terms contain sub-terms that cancel; in particular, only velocity motions *transverse* to magnetic-field lines can shear or compress.
- 2. Solid-body rotation (which is a non-shearing motion that simply rotates the whole field configuration) shows up in the so-called shear and advection terms in a strange way.

Finally, even after these issues have been addressed, resolving the final terms into a particular curvilinear system (e.g., spherical coordinates) seems to present a major headache. Our goal here is to explain fully how these problems emerge and propose a tentative solution for the case of spherical coordinates.

2 Transverse shear and compression

To see how Problem 1 arises, we decompose the velocity field into components parallel and perpendicular to the local direction of B:

$$\boldsymbol{u} \coloneqq u_{\parallel} \hat{\boldsymbol{e}}_{\parallel} + \boldsymbol{u}_{\perp} \tag{3}$$

Obviously $\mathbf{B} = Bx_{\parallel}$, where $B = |\mathbf{B}|$. We denote the Cartesian distance along \mathbf{B} by x_{\parallel} . We also decompose \mathbf{u} into its parallel and perpendicular components:

$$\nabla \cdot \boldsymbol{u} = \frac{\partial u_{\parallel}}{\partial x_{\parallel}} + \nabla_{\perp} \cdot \boldsymbol{u}_{\perp} \tag{4}$$

We then calculate

$$\boldsymbol{B} \cdot \nabla \boldsymbol{u} - (\nabla \cdot \boldsymbol{u}) \boldsymbol{B} = B \frac{\partial}{\partial x_{\parallel}} (u_{\parallel} \hat{\boldsymbol{e}}_{\parallel} + \boldsymbol{u}_{\perp}) - \left(\frac{\partial u_{\parallel}}{\partial x_{\parallel}} + \nabla_{\perp} \cdot \boldsymbol{u}_{\perp} \right) B \hat{\boldsymbol{e}}_{\parallel}$$

$$= B \frac{\partial u_{\parallel}}{\partial x_{\parallel}} \hat{\boldsymbol{e}}_{\parallel} + B \frac{\partial \boldsymbol{u}_{\perp}}{\partial x_{\parallel}} - \frac{\partial u_{\parallel}}{\partial x_{\parallel}} B \hat{\boldsymbol{e}}_{\parallel} - (\nabla_{\perp} \cdot \boldsymbol{u}_{\perp}) B \hat{\boldsymbol{e}}_{\parallel}$$

$$= B \cdot \nabla \boldsymbol{u}_{\perp} - (\nabla_{\perp} \cdot \boldsymbol{u}_{\perp}) \boldsymbol{B}. \tag{5}$$

Thus, only motions transverse to the local field line (i.e., u_{\perp}) can shear or compress B.

3 Rigid rotation

To see how Problem 2 arises, we consider a velocity field due to rigid rotation at constant angular velocity Ω about the z-axis in a cylindrical coordinate system:

$$\Omega = \Omega \hat{e}_z = \text{constant} \tag{6a}$$

$$\boldsymbol{u} = \boldsymbol{\Omega} \times \boldsymbol{r} = \Omega \lambda \hat{\boldsymbol{e}}_{\phi} \tag{6b}$$

Here, λ is the cylindrical radius, ϕ the longitude, and z the axial coordinate. In general, $\hat{e}_{(\cdots)}$ denotes a unit vector in the direction of its subscript. We calculate:

$$\nabla \cdot \boldsymbol{u} = \nabla \cdot (\boldsymbol{\Omega} \times \boldsymbol{r})$$

$$= \boldsymbol{\Omega} \cdot \nabla \times \boldsymbol{r} - \boldsymbol{r} \cdot \nabla \times \boldsymbol{\Omega}$$

$$= 0 \quad \text{no compression for rigid rotation (obviously)}. \tag{7}$$

Then:

$$\mathbf{B} \cdot \nabla \mathbf{u} = (\mathbf{B} \cdot \nabla)(\mathbf{\Omega} \times \mathbf{r})$$

$$= \mathbf{\Omega} \times [(\mathbf{B} \cdot \nabla)(\mathbf{r})]$$

$$= \mathbf{\Omega} \times \mathbf{B}$$
 "shear" for rigid rotation. (9)

Finally:

$$-\mathbf{u} \cdot \nabla \mathbf{B} = -\Omega \lambda \hat{\mathbf{e}}_{\phi} \cdot \nabla \mathbf{B}$$

$$= -\Omega \frac{\partial}{\partial \phi} (B_{\lambda} \hat{\mathbf{e}}_{\lambda} + B_{\phi} \hat{\mathbf{e}}_{\phi} + B_{z} \hat{\mathbf{e}}_{z})$$

$$= -\Omega \sum_{\alpha} \left(\frac{\partial B_{\alpha}}{\partial \phi} \right) \hat{\mathbf{e}}_{\alpha} - \Omega \sum_{\alpha} B_{\alpha} \frac{\partial \hat{\mathbf{e}}_{\alpha}}{\partial \phi},$$

where the index α runs over the three cylindrical coordinates. Note that in the cylindrical coordinate system (or indeed any coordinate system with an axis of rotational symmetry), $\partial \hat{e}_{\alpha}/\partial \phi = \hat{e}_z \times \hat{e}_{\alpha}$ for each α . Thus,

$$-\Omega \sum_{\alpha} B_{\alpha} \frac{\partial \hat{e}_{\alpha}}{\partial \phi} = -\Omega \sum_{\alpha} B_{\alpha} \hat{e}_{z} \times \hat{e}_{\alpha}$$
$$= -\Omega \hat{e}_{z} \times \sum_{\alpha} B_{\alpha} \hat{e}_{\alpha}$$
$$= -\Omega \times \mathbf{B}$$

and so

$$-\boldsymbol{u} \cdot \nabla \boldsymbol{B} = -\Omega \sum_{\alpha} \left(\frac{\partial B_{\alpha}}{\partial \phi} \right) \hat{\boldsymbol{e}}_{\alpha} - \boldsymbol{\Omega} \times \boldsymbol{B} \quad \text{"advection" for rigid rotation.}$$
 (10)

Mathematically, in any coordinate system with a z-axis of rotational symmetry, the action of rigid rotation is as follows:

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times [(\mathbf{\Omega} \times \mathbf{r}) \times \mathbf{B}]$$

$$= \sum_{\alpha} \left(-\Omega \frac{\partial B_{\alpha}}{\partial \phi} \right) \hat{\mathbf{e}}_{\alpha} \tag{11a}$$

or
$$\left(\frac{\partial}{\partial t} + \Omega \frac{\partial}{\partial \phi}\right) B_{\alpha} = 0$$
 for each α . (11b)

If you think about it, this makes sense: All the rigid rotation does is rotate the whole field configuration around the z-axis at the rate Ω . If you decide to also rotate at Ω (so your personal Eulerian time derivative is $\partial/\partial t + \Omega \partial/\partial \phi$), then each component of the magnetic-field configuration should remain the same in your frame. Note that rotation does not advect the vector magnetic field (like the term $-\mathbf{u} \cdot \nabla \mathbf{B}$ viewed on its own would suggest), but rather advects the field components (as if they were scalars) in any coordinate system with an axis of rotational symmetry.

4 Solution for spherical coordinates

Resolving these issues fully for the spherical coordinate system seems complicated and I am not fully sure how to do it! In particular (for "full" resolution) we should, separately at each point (r, θ, ϕ) :

- 1. Form a local Cartesian coordinate system, say (x_1, x_2, x_3) , whose origin lies at the point (r, θ, ϕ) . At the origin, $(\hat{e}_1, \hat{e}_2, \hat{e}_3)$ will coincide with $(\hat{e}_r, \hat{e}_\theta, \hat{e}_\phi)$. But slightly away from the origin, $(\hat{e}_1, \hat{e}_2, \hat{e}_3)$ will stay fixed while $(\hat{e}_r, \hat{e}_\theta, \hat{e}_\phi)$ curve away.
- 2. Calculate the velocity-gradient tensor: $\partial u_1/\partial x_1$, $\partial u_2/\partial x_1$, $\partial u_3/\partial x_1$, etc. While calculating derivatives, be careful to differentiate along the Cartesian coordinates, *not* the curvilinear ones or along the actual \boldsymbol{B} -line). Express the final tensor components in spherical coordinates.
- 3. Rotate "into \boldsymbol{B} " to form a new primed coordinate system (such that $\hat{\boldsymbol{e}}'_1$ points along \boldsymbol{B}), calculate the $\partial u'_j/\partial x'_i$, and thus form $\partial u_{\parallel}/\partial x_{\parallel}$ and $\partial \boldsymbol{u}_{\perp}/\partial x_{\parallel}$ (note that $\nabla_{\perp} \cdot \boldsymbol{u}_{\perp} = \nabla \cdot \boldsymbol{u} \partial u_{\parallel}/\partial x_{\parallel}$). Note that the direction cosines $\hat{\boldsymbol{e}}'_1$ makes with $(\hat{\boldsymbol{e}}_1, \hat{\boldsymbol{e}}_2, \hat{\boldsymbol{e}}_3)$ are simply $B_r/|\boldsymbol{B}|$, $B_\theta/|\boldsymbol{B}|$, respectively. Note that both unit vectors and vector components transform like $x'_1 = \sum_{j=1}^3 R_{1j}x_j$, where R_{1j} is the j^{th} direction cosine.
- 4. Subtract the part of u_{\perp} corresponding to solid-body rotation and put it in the form of Equation (11a).

5. Exult, because I haven't been able to do this!

Instead of addressing the full problem as just described, I have brute-forced my way into a quasi-solution, canceling obvious terms and expressing what seem to be "transverse shear with no solid-body rotation," "transverse compression," and advection.

We write:

$$I_{r} = B_{r} \frac{\partial u_{r}}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial u_{r}}{\partial \theta} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial u_{r}}{\partial \phi} - \frac{B_{\theta} u_{\theta} + B_{\phi} u_{\phi}}{r}$$

$$- \mathbf{u} \cdot \nabla B_{r} + \frac{u_{\theta} B_{\theta} + u_{\phi} B_{\phi}}{r}$$

$$- \left[\frac{\partial u_{r}}{\partial r} + \frac{2u_{r}}{r} + \frac{1}{r} \left(\frac{\partial u_{\theta}}{\partial \theta} + \cot \theta u_{\theta} \right) + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} \right] B_{r}$$

or

$$I_r = \left(\frac{B_\theta}{r} \frac{\partial}{\partial \theta} + \frac{B_\phi}{r \sin \theta} \frac{\partial}{\partial \phi}\right) u_r \qquad r: \text{ transverse shear}$$
 (12a)

$$- \boldsymbol{u} \cdot \nabla B_r$$
 r: advection (12b)

$$-\left(\frac{1}{r}\frac{\partial u_{\theta}}{\partial \theta} + \frac{1}{r\sin\theta}\frac{\partial u_{\phi}}{\partial \phi} + \frac{2u_{r}}{r} + \frac{\cot\theta u_{\theta}}{r}\right)B_{r} \qquad r: \text{ transverse compression} \qquad (12c)$$

Then, for I_{θ} :

$$I_{\theta} = B_{r} \frac{\partial u_{\theta}}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial u_{\theta}}{\partial \theta} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial u_{\theta}}{\partial \phi} + \frac{B_{\theta} u_{r}}{r} - \frac{\cot \theta B_{\phi} u_{\phi}}{r}$$
$$- \mathbf{u} \cdot \nabla B_{\theta} - \frac{u_{\theta} B_{r}}{r} + \frac{\cot \theta u_{\phi} B_{\phi}}{r}$$
$$- \left[\frac{\partial u_{r}}{\partial r} + \frac{2u_{r}}{r} + \frac{1}{r} \left(\frac{\partial u_{\theta}}{\partial \theta} + \cot \theta u_{\theta} \right) + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} \right] B_{\theta}$$

or

$$I_{\theta} = r \left(B_r \frac{\partial}{\partial r} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial}{\partial \phi} \right) \left(\frac{u_{\theta}}{r} \right) \qquad \theta: \text{ transverse shear}$$
 (13a)

$$- \boldsymbol{u} \cdot \nabla B_{\theta}$$
 θ : advection (13b)

$$-\left(\frac{\partial u_r}{\partial r} + \frac{1}{r\sin\theta}\frac{\partial u_\phi}{\partial \phi} + \frac{u_r}{r} + \frac{\cot\theta u_\theta}{r}\right)B_\theta \qquad \theta: \text{ transverse compression}$$
 (13c)

Finally, for I_{ϕ} :

$$\begin{split} I_{\phi} = & B_{r} \frac{\partial u_{\phi}}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial u_{\phi}}{\partial \theta} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} + \frac{B_{\phi} u_{r}}{r} + \frac{\cot \theta B_{\phi} u_{\theta}}{r} \\ & - \mathbf{u} \cdot \nabla B_{\phi} - \frac{u_{\phi} B_{r}}{r} - \frac{\cot \theta u_{\phi} B_{\theta}}{r} \\ & - \left[\frac{\partial u_{r}}{\partial r} + \frac{2u_{r}}{r} + \frac{1}{r} \left(\frac{\partial u_{\theta}}{\partial \theta} + \cot \theta u_{\theta} \right) + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} \right] B_{\phi} \end{split}$$

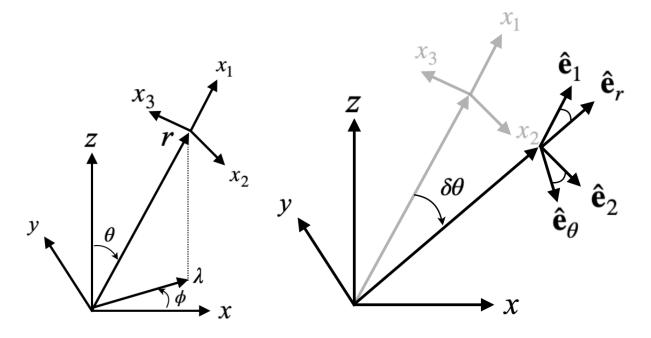


Figure 1: Left: Schematic of coordinate systems used here, their relation to one another, and each coordinate's meaning. Note that $(x, y, z) = (\lambda \cos \phi, \lambda \sin \phi, z) = (r \sin \theta \cos \phi, r \sin \theta \sin \phi, r \cos \theta)$. The local Cartesian system (x_1, x_2, x_3) has its origin at the point (r, θ, ϕ) , where $(\hat{e}_1, \hat{e}_2, \hat{e}_3) = (\hat{e}_r, \hat{e}_\theta, \hat{e}_\phi)$. Right: Away from the origin, the unit vectors no longer coincide. We show the relationship between the spherical-coordinate unit vectors and the Cartesian unit vectors when $\theta \to \theta + \delta\theta$ (or equivalently, $x_2 = 0 \to \delta x_2$).

or

$$I_{\phi} = r \sin \theta \left(B_r \frac{\partial}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial}{\partial \theta} \right) \left(\frac{u_{\phi}}{r \sin \theta} \right) \qquad \phi: \text{ transverse shear}$$
 (14a)

$$- \boldsymbol{u} \cdot \nabla B_{\phi}$$
 ϕ : advection (14b)

$$-\left(\frac{\partial u_r}{\partial r} + \frac{1}{r}\frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r}\right)B_\phi \qquad \qquad \phi: \text{ transverse compression}$$
 (14c)

5 Transverse compression

Here we show that the "transverse compression" terms defined in Equations (12)–(14) make sense. The key point is that we must be careful (in calculating $\nabla_{\perp} \cdot \boldsymbol{u}_{\perp}$) when differentiating along θ or ϕ , because the local spherical unit vectors will all be changing. For example the divergence "transverse to $\hat{\boldsymbol{e}}_r$ " is not simply $\nabla \cdot (u_{\theta}\hat{\boldsymbol{e}}_{\theta} + u_{\phi}\hat{\boldsymbol{e}}_{\phi})$; there are also curvature terms.

Figure 1 shows schematically how this works. We set up a local Cartesian coordinate system (with origin at the point (r, θ, ϕ)) denoted by (x_1, x_2, x_3) . As we differentiate \boldsymbol{u} , the spherical-coordinate unit vectors change, while the Cartesian ones stay fixed. In particular, if we move from $(r, \theta, \phi) \to (r + \delta r, \theta + \delta \theta, \phi + \delta \phi)$ (or in the Cartesian system, $(0, 0, 0) \to (\delta x_1, \delta x_2, \delta x_3)$)