Induction term in spherical coordinates

Loren Matilsky

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1 The problem with the traditional induction terms

We consider the ideal (resistance-free) magnetohydrodynamic (MHD) induction equation:

$$\frac{\partial \boldsymbol{B}}{\partial t} = \nabla \times (\boldsymbol{u} \times \boldsymbol{B}) := \boldsymbol{I},\tag{1}$$

$$= \mathbf{B} \cdot \nabla \mathbf{u} - \mathbf{u} \cdot \nabla \mathbf{B} - (\nabla \cdot \mathbf{u}) \mathbf{B}, \tag{2}$$

where \boldsymbol{B} and \boldsymbol{u} are the vector magnetic and velocity fields (respectively) and \boldsymbol{I} is the vector induction. The three terms on the right-hand-side of Equation (2) are often interpreted as "shear," "advection," and "compression," respectively. However, this interpretation is problematic in general for two reasons:

- 1. The so-called shear and compression terms contain sub-terms that cancel; in particular, only velocity motions transverse to magnetic-field lines can shear or compress.
- 2. Solid-body rotation (which is a non-shearing motion that simply rotates the whole field configuration) shows up in the so-called shear and advection terms in a strange way.

Finally, even after these issues have been addressed, resolving the final terms into a particular curvilinear system (e.g., spherical coordinates) seems to present a major headache. Our goal here is to explain fully how these problems emerge and propose a tentative solution for the case of spherical coordinates.

2 Transverse shear and compression

To see how Problem 1 arises, we decompose the velocity field into components parallel and perpendicular to the local direction of B:

$$\boldsymbol{u} \coloneqq u_{\parallel} \hat{\boldsymbol{e}}_{\parallel} + \boldsymbol{u}_{\perp}. \tag{3}$$

Obviously $\mathbf{B} = Bx_{\parallel}$, where $B = |\mathbf{B}|$. We denote the Cartesian distance along \mathbf{B} by x_{\parallel} . We also decompose \mathbf{u} into its parallel and perpendicular components:

$$\nabla \cdot \boldsymbol{u} = \frac{\partial u_{\parallel}}{\partial x_{\parallel}} + \nabla_{\perp} \cdot \boldsymbol{u}_{\perp}. \tag{4}$$

We then calculate

$$\boldsymbol{B} \cdot \nabla \boldsymbol{u} - (\nabla \cdot \boldsymbol{u}) \boldsymbol{B} = B \frac{\partial}{\partial x_{\parallel}} (u_{\parallel} \hat{\boldsymbol{e}}_{\parallel} + \boldsymbol{u}_{\perp}) - \left(\frac{\partial u_{\parallel}}{\partial x_{\parallel}} + \nabla_{\perp} \cdot \boldsymbol{u}_{\perp} \right) B \hat{\boldsymbol{e}}_{\parallel}$$

$$= B \frac{\partial u_{\parallel}}{\partial x_{\parallel}} \hat{\boldsymbol{e}}_{\parallel} + B \frac{\partial \boldsymbol{u}_{\perp}}{\partial x_{\parallel}} - \frac{\partial u_{\parallel}}{\partial x_{\parallel}} B \hat{\boldsymbol{e}}_{\parallel} - (\nabla_{\perp} \cdot \boldsymbol{u}_{\perp}) B \hat{\boldsymbol{e}}_{\parallel}$$

$$= B \cdot \nabla \boldsymbol{u}_{\perp} - (\nabla_{\perp} \cdot \boldsymbol{u}_{\perp}) \boldsymbol{B}. \tag{5}$$

Thus, only motions transverse to the local field line (i.e., u_{\perp}) can shear or compress B.

3 Rigid rotation

To see how Problem 2 arises, we consider a velocity field due to rigid rotation at constant angular velocity Ω about the z-axis in a cylindrical coordinate system:

$$\Omega = \Omega \hat{e}_z = \text{constant} \tag{6a}$$

and
$$\mathbf{u} = \mathbf{\Omega} \times \mathbf{r} = \Omega \lambda \hat{\mathbf{e}}_{\phi}$$
. (6b)

Here, λ is the cylindrical radius, ϕ the longitude, and z the axial coordinate (also see Figure 1). In general, \hat{e}_{α} denotes a unit vector in the direction of the coordinate α . We calculate:

$$\nabla \cdot \boldsymbol{u} = \nabla \cdot (\boldsymbol{\Omega} \times \boldsymbol{r})$$

$$= \boldsymbol{\Omega} \cdot \nabla \times \boldsymbol{r} - \boldsymbol{r} \cdot \nabla \times \boldsymbol{\Omega}$$

$$= 0 \quad \text{no compression for rigid rotation (obviously)}. \tag{7}$$

Then:

$$B \cdot \nabla u = (B \cdot \nabla)(\Omega \times r)$$

$$= \Omega \times [(B \cdot \nabla)(r)]$$

$$= \Omega \times B \quad \text{"shear" for rigid rotation.}$$
(8)

Finally:

$$-\mathbf{u} \cdot \nabla \mathbf{B} = -\Omega \lambda \hat{\mathbf{e}}_{\phi} \cdot \nabla \mathbf{B}$$

$$= -\Omega \frac{\partial}{\partial \phi} (B_{\lambda} \hat{\mathbf{e}}_{\lambda} + B_{\phi} \hat{\mathbf{e}}_{\phi} + B_{z} \hat{\mathbf{e}}_{z})$$

$$= -\Omega \sum_{\alpha} \left(\frac{\partial B_{\alpha}}{\partial \phi} \right) \hat{\mathbf{e}}_{\alpha} - \Omega \sum_{\alpha} B_{\alpha} \frac{\partial \hat{\mathbf{e}}_{\alpha}}{\partial \phi},$$

where the index α runs over the three cylindrical coordinates. Note that in the cylindrical coordinate system (or indeed any coordinate system with an axis of rotational symmetry), $\partial \hat{e}_{\alpha}/\partial \phi = \hat{e}_z \times \hat{e}_{\alpha}$ for each α . Thus,

$$-\Omega \sum_{\alpha} B_{\alpha} \frac{\partial \hat{e}_{\alpha}}{\partial \phi} = -\Omega \sum_{\alpha} B_{\alpha} \hat{e}_{z} \times \hat{e}_{\alpha}$$
$$= -\Omega \hat{e}_{z} \times \sum_{\alpha} B_{\alpha} \hat{e}_{\alpha}$$
$$= -\Omega \times B$$

and so

$$-\boldsymbol{u} \cdot \nabla \boldsymbol{B} = -\Omega \sum_{\alpha} \left(\frac{\partial B_{\alpha}}{\partial \phi} \right) \hat{\boldsymbol{e}}_{\alpha} - \boldsymbol{\Omega} \times \boldsymbol{B} \quad \text{"advection" for rigid rotation.}$$
 (10)

Mathematically, in any coordinate system with a z-axis of rotational symmetry, the action of rigid rotation is as follows:

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times [(\mathbf{\Omega} \times \mathbf{r}) \times \mathbf{B}]$$

$$= \sum_{\alpha} \left(-\Omega \frac{\partial B_{\alpha}}{\partial \phi} \right) \hat{\mathbf{e}}_{\alpha} \tag{11a}$$

or
$$\left(\frac{\partial}{\partial t} + \Omega \frac{\partial}{\partial \phi}\right) B_{\alpha} = 0$$
 for each α . (11b)

If you think about it, this makes sense: All the rigid rotation does is rotate the whole field configuration around the z-axis at the rate Ω . If you decide to also rotate at Ω (so your personal Eulerian time derivative is $\partial/\partial t + \Omega \partial/\partial \phi$), then each component of the magnetic-field configuration should remain the same in your frame. Note that rotation does not advect the vector magnetic field (like the term $-\mathbf{u} \cdot \nabla \mathbf{B}$ viewed on its own would suggest), but rather advects the field components (as if they were scalars) in any coordinate system with an axis of rotational symmetry.

4 Solution for spherical coordinates

Resolving these issues fully for the spherical coordinate system seems complicated and I am not fully sure how to do it! In particular (for "full" resolution) we should, separately at each point (r, θ, ϕ) :

- 1. Form a local Cartesian coordinate system, say (x_1, x_2, x_3) , whose origin lies at the point (r, θ, ϕ) . At the origin, $(\hat{e}_1, \hat{e}_2, \hat{e}_3)$ will coincide with $(\hat{e}_r, \hat{e}_\theta, \hat{e}_\phi)$. But slightly away from the origin, $(\hat{e}_1, \hat{e}_2, \hat{e}_3)$ will stay fixed while $(\hat{e}_r, \hat{e}_\theta, \hat{e}_\phi)$ curve away.
- 2. Calculate the velocity-gradient tensor: $\partial u_1/\partial x_1$, $\partial u_2/\partial x_1$, $\partial u_3/\partial x_1$, etc. While calculating derivatives, be careful to differentiate along the Cartesian coordinates, *not* the curvilinear ones or along the actual \boldsymbol{B} -line). Express the final tensor components in spherical coordinates.
- 3. Rotate "into \boldsymbol{B} " to form a new primed coordinate system (such that $\hat{\boldsymbol{e}}'_1$ points along \boldsymbol{B}), calculate the $\partial u'_j/\partial x'_i$, and thus form $\partial u_{\parallel}/\partial x_{\parallel}$ and $\partial \boldsymbol{u}_{\perp}/\partial x_{\parallel}$ (note that $\nabla_{\perp} \cdot \boldsymbol{u}_{\perp} = \nabla \cdot \boldsymbol{u} \partial u_{\parallel}/\partial x_{\parallel}$). Note that the direction cosines $\hat{\boldsymbol{e}}'_1$ makes with $(\hat{\boldsymbol{e}}_1, \hat{\boldsymbol{e}}_2, \hat{\boldsymbol{e}}_3)$ are simply $B_r/|\boldsymbol{B}|$, $B_\theta/|\boldsymbol{B}|$, respectively. Note that both unit vectors and vector components transform like $x'_1 = \sum_{j=1}^3 R_{1j}x_j$, where R_{1j} is the j^{th} direction cosine.
- 4. Subtract the part of u_{\perp} corresponding to solid-body rotation and put it in the form of Equation (11a).

5. Exult, because I haven't been able to do this!

Instead of addressing the full problem as just described, I have brute-forced my way into a quasi-solution, canceling obvious terms and expressing what seem to be "transverse shear with no solid-body rotation," "transverse compression," and advection.

We write:

$$I_{r} = B_{r} \frac{\partial u_{r}}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial u_{r}}{\partial \theta} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial u_{r}}{\partial \phi} - \frac{B_{\theta} u_{\theta} + B_{\phi} u_{\phi}}{r}$$

$$- \mathbf{u} \cdot \nabla B_{r} + \frac{u_{\theta} B_{\theta} + u_{\phi} B_{\phi}}{r}$$

$$- \left[\frac{\partial u_{r}}{\partial r} + \frac{2u_{r}}{r} + \frac{1}{r} \left(\frac{\partial u_{\theta}}{\partial \theta} + \cot \theta u_{\theta} \right) + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} \right] B_{r}$$

or

$$I_r = \left(\frac{B_\theta}{r} \frac{\partial}{\partial \theta} + \frac{B_\phi}{r \sin \theta} \frac{\partial}{\partial \phi}\right) u_r \qquad r: \text{ transverse shear}$$
 (12a)

$$- \boldsymbol{u} \cdot \nabla B_r$$
 r: advection (12b)

$$-\left(\frac{1}{r}\frac{\partial u_{\theta}}{\partial \theta} + \frac{1}{r\sin\theta}\frac{\partial u_{\phi}}{\partial \phi} + \frac{2u_{r}}{r} + \frac{\cot\theta u_{\theta}}{r}\right)B_{r} \qquad r: \text{ transverse compression} \qquad (12c)$$

Then, for I_{θ} :

$$I_{\theta} = B_{r} \frac{\partial u_{\theta}}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial u_{\theta}}{\partial \theta} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial u_{\theta}}{\partial \phi} + \frac{B_{\theta} u_{r}}{r} - \frac{\cot \theta B_{\phi} u_{\phi}}{r}$$
$$- \mathbf{u} \cdot \nabla B_{\theta} - \frac{u_{\theta} B_{r}}{r} + \frac{\cot \theta u_{\phi} B_{\phi}}{r}$$
$$- \left[\frac{\partial u_{r}}{\partial r} + \frac{2u_{r}}{r} + \frac{1}{r} \left(\frac{\partial u_{\theta}}{\partial \theta} + \cot \theta u_{\theta} \right) + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} \right] B_{\theta}$$

or

$$I_{\theta} = r \left(B_r \frac{\partial}{\partial r} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial}{\partial \phi} \right) \left(\frac{u_{\theta}}{r} \right) \qquad \theta: \text{ transverse shear}$$
 (13a)

$$- \boldsymbol{u} \cdot \nabla B_{\theta}$$
 θ : advection (13b)

$$-\left(\frac{\partial u_r}{\partial r} + \frac{1}{r\sin\theta}\frac{\partial u_\phi}{\partial \phi} + \frac{u_r}{r} + \frac{\cot\theta u_\theta}{r}\right)B_\theta \qquad \theta: \text{ transverse compression}$$
 (13c)

Finally, for I_{ϕ} :

$$\begin{split} I_{\phi} = & B_{r} \frac{\partial u_{\phi}}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial u_{\phi}}{\partial \theta} + \frac{B_{\phi}}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} + \frac{B_{\phi} u_{r}}{r} + \frac{\cot \theta B_{\phi} u_{\theta}}{r} \\ & - \mathbf{u} \cdot \nabla B_{\phi} - \frac{u_{\phi} B_{r}}{r} - \frac{\cot \theta u_{\phi} B_{\theta}}{r} \\ & - \left[\frac{\partial u_{r}}{\partial r} + \frac{2u_{r}}{r} + \frac{1}{r} \left(\frac{\partial u_{\theta}}{\partial \theta} + \cot \theta u_{\theta} \right) + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} \right] B_{\phi} \end{split}$$

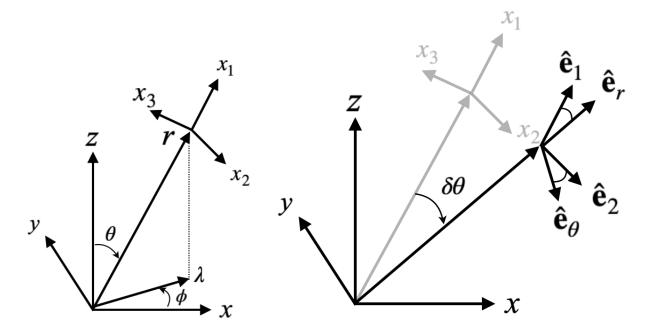


Figure 1: Left: Schematic of coordinate systems used here, their relation to one another, and each coordinate's meaning. Note that $(x, y, z) = (\lambda \cos \phi, \lambda \sin \phi, z) = (r \sin \theta \cos \phi, r \sin \theta \sin \phi, r \cos \theta)$. The local Cartesian system (x_1, x_2, x_3) has its origin at the point (r, θ, ϕ) , where $(\hat{e}_1, \hat{e}_2, \hat{e}_3) = (\hat{e}_r, \hat{e}_\theta, \hat{e}_\phi)$. Right: Away from the origin, the unit vectors no longer coincide. We show the relationship between the spherical-coordinate unit vectors and the Cartesian unit vectors when $\theta \to \theta + \delta\theta$ (or equivalently, $x_2 = 0 \to \delta x_2$).

or

$$I_{\phi} = r \sin \theta \left(B_r \frac{\partial}{\partial r} + \frac{B_{\theta}}{r} \frac{\partial}{\partial \theta} \right) \left(\frac{u_{\phi}}{r \sin \theta} \right) \qquad \phi: \text{ transverse shear}$$
 (14a)

$$- \boldsymbol{u} \cdot \nabla B_{\phi}$$
 ϕ : advection (14b)

$$-\left(\frac{\partial u_r}{\partial r} + \frac{1}{r}\frac{\partial u_\theta}{\partial \theta} + \frac{u_r}{r}\right)B_\phi \qquad \qquad \phi: \text{ transverse compression}$$
 (14c)

5 Transverse compression

Here we show that the "transverse compression" terms defined in Equations (12)–(14) make sense. The key point is that we must be careful when calculating the "transverse divergence." When we differentiate along θ or ϕ , the local spherical unit vectors will all change. For example, the divergence "transverse to $\hat{\boldsymbol{e}}_r$ " is not simply $\nabla \cdot (u_\theta \hat{\boldsymbol{e}}_\theta + u_\phi \hat{\boldsymbol{e}}_\phi)$; there are also curvature terms.

Figure 1 shows schematically how this works. We set up a local Cartesian coordinate system (with origin at the point (r, θ, ϕ)) denoted by (x_1, x_2, x_3) . As we differentiate \boldsymbol{u} , the spherical-coordinate unit vectors change, while the Cartesian ones stay fixed. In particular, if we move from $(r, \theta, \phi) \to (r + \delta r, \theta + \delta \theta, \phi + \delta \phi)$ (or equivalently in the Cartesian system,

 $(0,0,0) \to (\delta x_1, \delta x_2, \delta x_3)$, we find the following (to first order in the δ 's):

$$u_1 = u_r - (\delta\theta)u_\theta - (\sin\theta\delta\phi)u_\phi, \tag{15a}$$

$$u_2 = u_\theta + (\delta\theta)u_r - (\cos\theta\delta\phi)u_\phi, \tag{15b}$$

and
$$u_3 = u_{\phi} + (\sin \theta \delta \phi) u_r + (\cos \theta \delta \phi) u_{\theta}$$
. (15c)

And trivially:

$$\delta x_1 = \delta r \tag{16a}$$

$$\delta x_2 = r\delta\theta \tag{16b}$$

$$\delta x_3 = r \sin \theta \delta \phi. \tag{16c}$$

We thus compute:

$$\frac{\partial u_1}{\partial x_1} = \lim_{\delta x_1 \to 0} \frac{u_1(\delta x_1, 0, 0) - u_1(0, 0, 0)}{\delta x_1}$$

$$= \lim_{\delta r \to 0} \frac{u_r(r + \delta r, \theta, \phi) - u_r(r, \theta, \phi)}{\delta r}$$

$$= \frac{\partial u_r}{\partial r}.$$
(17)

That was the easy one! Now for the curvy ones:

$$\frac{\partial u_2}{\partial x_2} = \lim_{\delta x_2 \to 0} \frac{u_2(0, \delta x_2, 0) - u_2(0, 0, 0)}{\delta x_2}$$

$$= \lim_{\delta \theta \to 0} \frac{u_{\theta}(r, \theta + \delta \theta, \phi) + u_r(r, \theta + \delta \theta, \phi) \delta \theta - u_{\theta}(r, \theta, \phi)}{r \delta \theta}$$

$$= \lim_{\delta \theta \to 0} \frac{u_{\theta}(r, \theta + \delta \theta, \phi) - u_{\theta}(r, \theta, \phi)}{r \delta \theta} + \lim_{\delta \theta \to 0} \frac{u_r(r, \theta + \delta \theta, \phi)}{r}$$

$$= \frac{1}{r} \frac{\partial u_{\theta}}{\partial \theta} + \frac{u_r}{r}.$$
(18)

And finally:

$$\frac{\partial u_3}{\partial x_3} = \lim_{\delta x_3 \to 0} \frac{u_3(0, 0, \delta x_3) - u_3(0, 0, 0)}{\delta x_3}$$

$$= \lim_{\delta \phi \to 0} \frac{u_{\phi}(r, \theta, \phi + \delta \phi) + (\sin \theta u_r + \cos \theta u_{\theta}) \delta \phi - u_{\phi}(r, \theta, \phi)}{r \sin \theta \delta \phi}$$

$$= \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} + \frac{u_r}{r} + \frac{\cot \theta u_{\theta}}{r}.$$
(19)

We thus find the various components of transverse divergence:

$$(\nabla \cdot \boldsymbol{u})_{\perp \text{ to } r} = \frac{\partial u_{2}}{\partial x_{2}} + \frac{\partial u_{3}}{\partial x_{3}}$$

$$= \frac{1}{r} \frac{\partial u_{\theta}}{\partial \theta} + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} + \frac{2u_{r}}{r} + \frac{\cot \theta u_{\theta}}{r}, \qquad (20)$$

$$(\nabla \cdot \boldsymbol{u})_{\perp \text{ to } \theta} = \frac{\partial u_{1}}{\partial x_{1}} + \frac{\partial u_{3}}{\partial x_{3}}$$

$$= \frac{\partial u_{r}}{\partial r} + \frac{1}{r \sin \theta} \frac{\partial u_{\phi}}{\partial \phi} + \frac{u_{r}}{r} + \frac{\cot \theta u_{\theta}}{r}, \qquad (21)$$
and
$$(\nabla \cdot \boldsymbol{u})_{\perp \text{ to } \phi} = \frac{\partial u_{1}}{\partial x_{1}} + \frac{\partial u_{2}}{\partial x_{2}}$$

$$= \frac{\partial u_{r}}{\partial r} + \frac{1}{r} \frac{\partial u_{\theta}}{\partial \theta} + \frac{u_{r}}{r}. \qquad (22)$$

Comparing with Equations (12)–(14), we see that these are the exact same divergence terms showing up in the "transverse compression."