

DAMWARE

(Data Mining & Exploration Web Application and Resources)

on behalf of the DAMWARE collaboration:

Co PI's:

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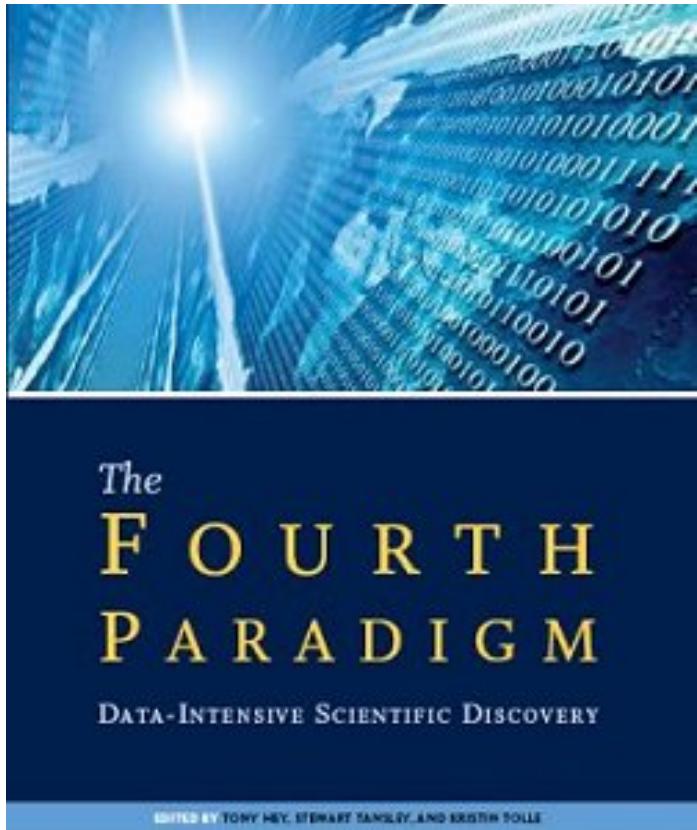
Giuseppe Longo – UNINA (Italy)

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Why Machine learning ?



The four legs of modern science

1. **Experiment** (ca. 3000 yrs)
2. **Theory** (few hundreds yrs) mathematical description, theoretical models, analytical laws (e.g. Newton, Maxwell, etc.)
3. **Simulations** (few tens of yrs) Complex phenomena
4. **Data-Intensive science** (now!!!)

<http://research.microsoft.com/fourthparadigm/>

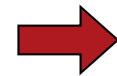
“One of the greatest challenges for 21st-century science is how we respond to this new era of data intensive science”

ASTRONOMY CAN BENEFIT FROM WHAT HAPPENS ELSEWHERE

Pucon, August 2013



As a result of large surveys, astronomy has entered an era where



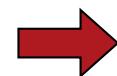
Most data will never be seen by humans!

The need for data storage, network, database-related technologies standards, etc.



Most knowledge hidden behind data complexity is potentially lost

Most (if not all) empirical relationships known so far depend on 3 parameters (e.g. fundamental plane of E galaxies and bulges). Simple universe or rather human bias?



Most data (and data constructs) cannot be comprehended by humans directly!

The need for data mining, KDD, data understanding technologies, hyperdimensional visualization, AI/Machine-assisted discovery

The various components of the data challenge

Data Gathering (e.g., from sensor networks, telescopes...)

→ **Data Farming:**

Storage/Archiving
Indexing, Searchability
Data Fusion, Interoperability

Specific projects
(pipelines, etc.)
&
Virtual observatory

→ **Data Mining** (or Knowledge Discovery in Databases):

Pattern or correlation search
Clustering analysis, automated classification
Outlier / anomaly searches
Hyperdimensional visualization.

Uncharted
land

→ **Data understanding**

Computer aided understanding
KDD
Etc.

→ **New Knowledge**



DAMEWARE

Is a web-based application (FREE AND OPEN TO THE PUBLIC) for massive data mining based on a suite of machine learning methods on top of a virtualized hybrid computing infrastructure

A joint effort between University Federico II, INAF-OACN & Caltech, aimed at implementing (as web 2.0 apps and services) a scientific gateway for data exploration on top of a virtualized distributed computing environment

Alfa released NOW

<http://dame.dsf.unina.it/>

Science and management

Technical documents

Template science cases

Newsletters

Tutorials



Effective DM requires complex work-flows

Use case

pre-processing

feature selection

choice of DM model

experiments

evaluation
of results

The logic behind DAMEWARE

Use case

Functionality

Classification

Regression

Clustering

Feature selection

DM models

GAME S, C,R

MLPBP S, C,R

MLPGA S, C,R

MLPQNA S, C,R

SVM S, C,R

K-Means U, CI

ESOM U, CI

SOFM U, CI

SOM U, CI

PPS U, CI, FS

Experiments

1-st

2-nd

3-rd

4-th

....

N-th



DAMEWARE the GUI

The screenshot displays the DAMEWARE application interface. At the top, there is a navigation bar with links for App Manuals, Model Manuals, Cloud Services, Science Cases, Documents, and Info. A user icon and 'LogOut' button are also present. The main area is titled 'RESOURCE MANAGER'. On the left, under 'Workspace', there is a 'New Workspace' button and a workspace named 'trial' with options to Rename, Workspace, Upload, Experiment, or Delete. On the right, there are two sections: 'File Manager' and 'My Experiments'. The 'File Manager' section shows a file named 'dataset2_2class_train.csv' last accessed on 2011-07-14. The 'My Experiments' section shows an experiment named 'mlpqnaClass1' which ended on 2011-07-15. This experiment has five associated files listed:

File	Type	Description
mlpqna_TRAIN_weights.txt	ASCII	final weights frozen at the end of the batch training
mlpqna_TRAIN.log	txt	log file
mlpqna_TRAIN_errorPlot.jpeg	JPEG	Plotting
dataset2_2class_train_mlpqna_TRAIN_output.txt	ASCII	confusion matrix calculated at the end of training
MLPQNA_Train_params.xml	xml	Experiment Configuration File



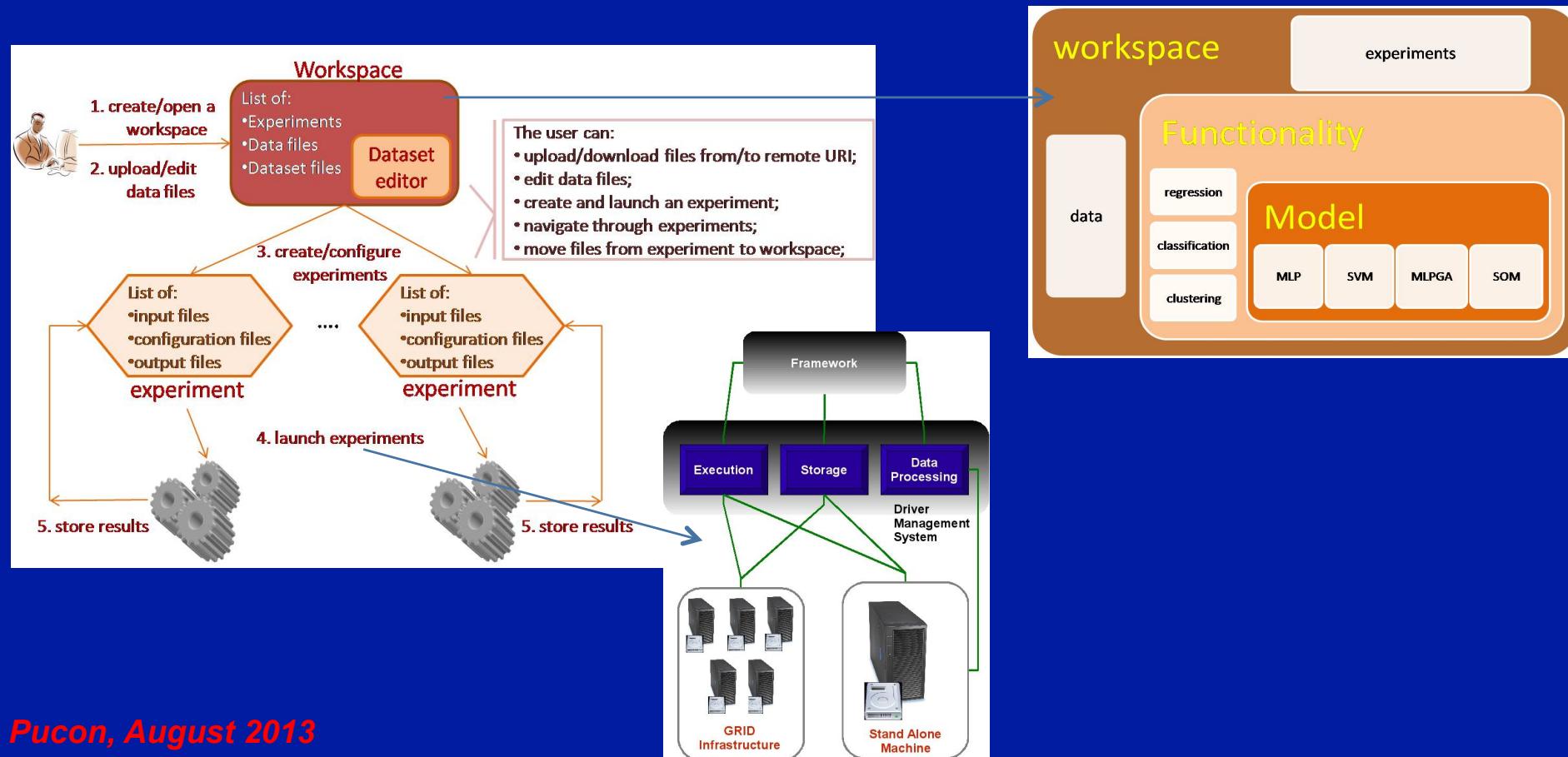
DAMEWARE



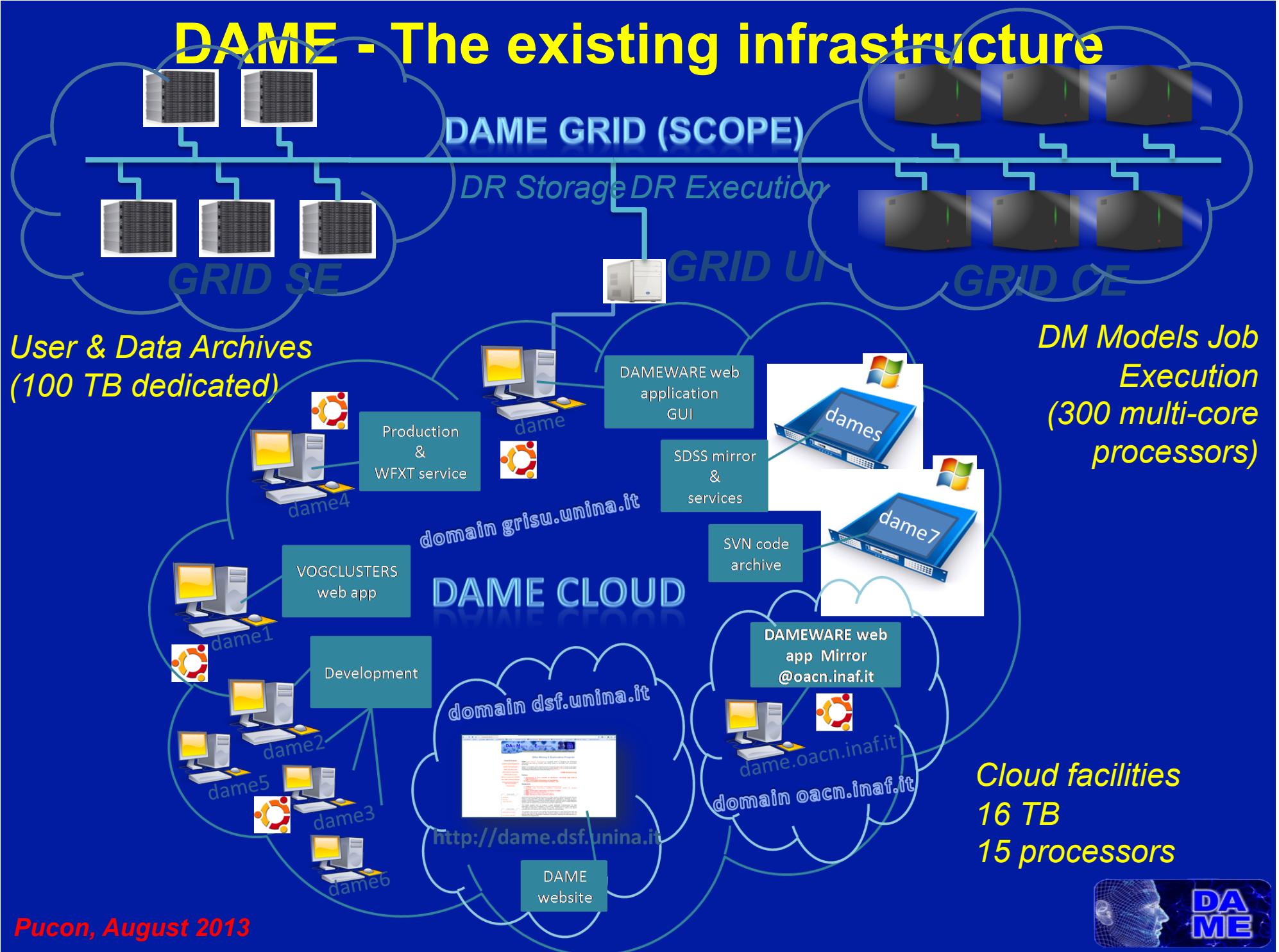
It is multi-disciplinary platform (astronomy, bioinformatics and medical diagnostics)

End users can remotely exploit high computing and storage power to process massive datasets (in principle they can do data mining on their smartphone...)

User can automatically plug-in his/her own algorithm and launch experiments through the Suite via a simple web browser



DAME - The existing infrastructure

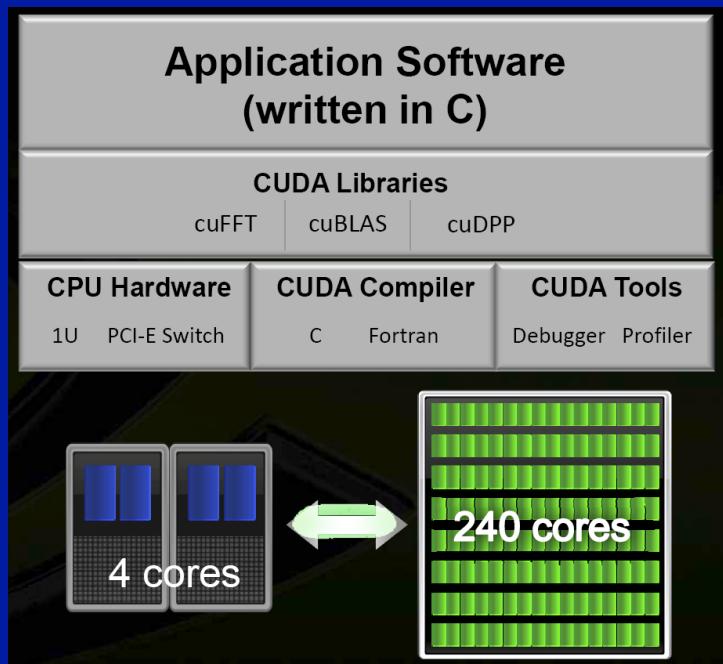
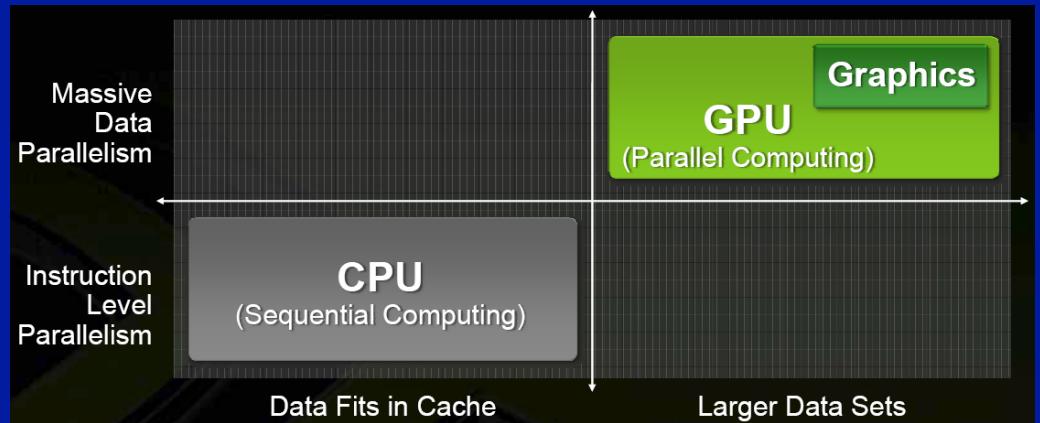


GPU technology ... sometime useful



The Graphical Processing Unit is specialized for compute-intensive, highly parallel computation (exactly what graphics rendering is about).

« GPU have evolved to the point where many real world apps are easily implemented on them and run significantly faster than on multi-core systems.»



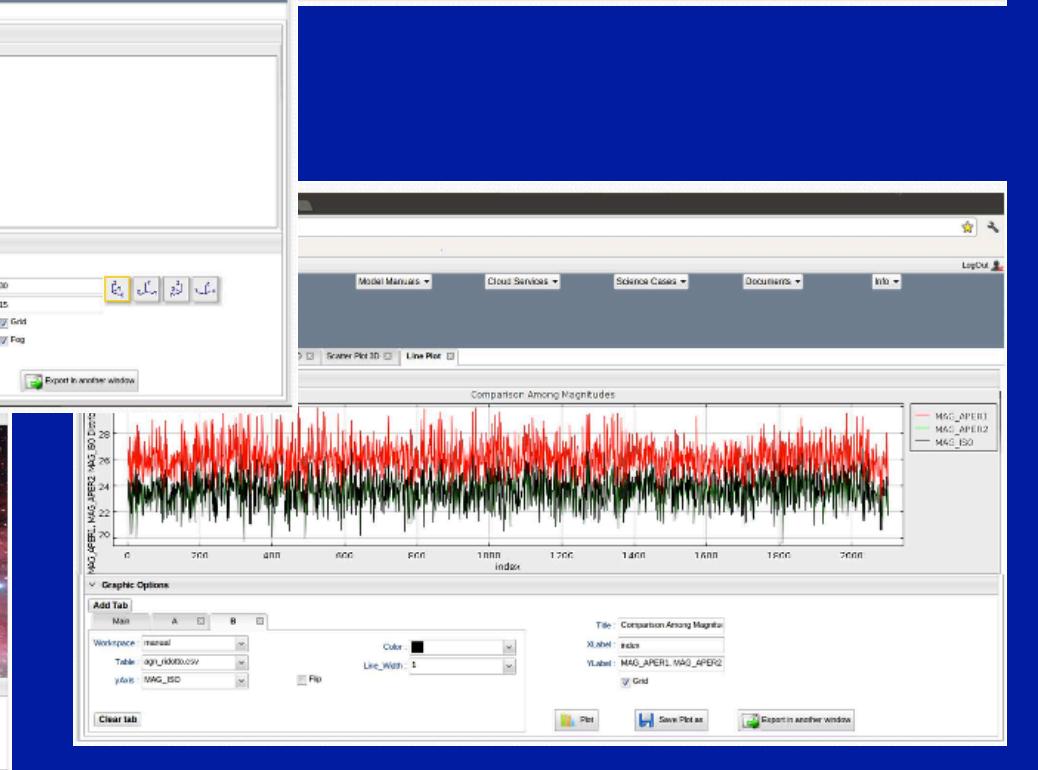
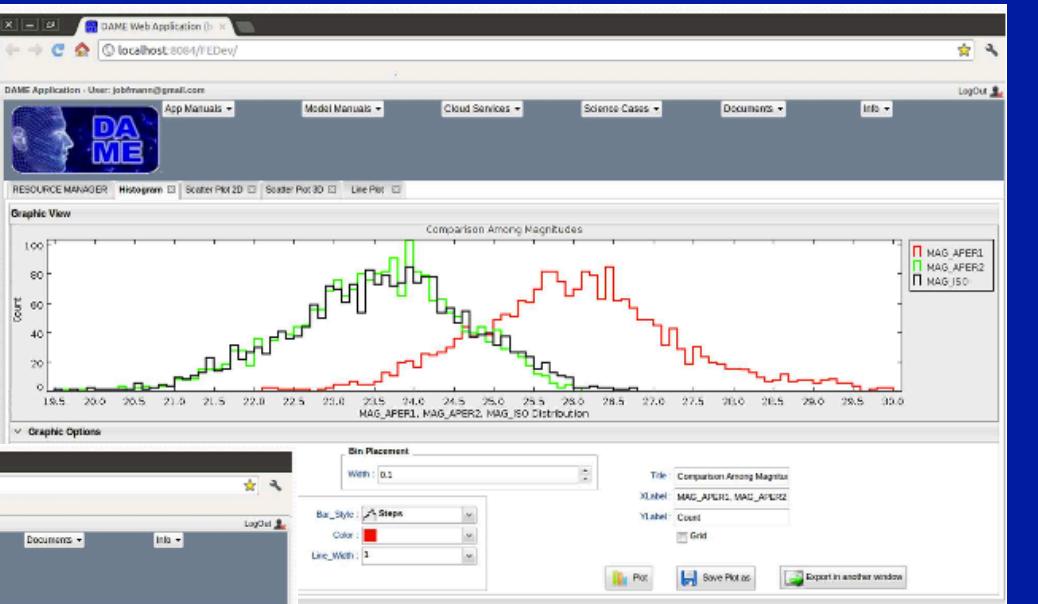
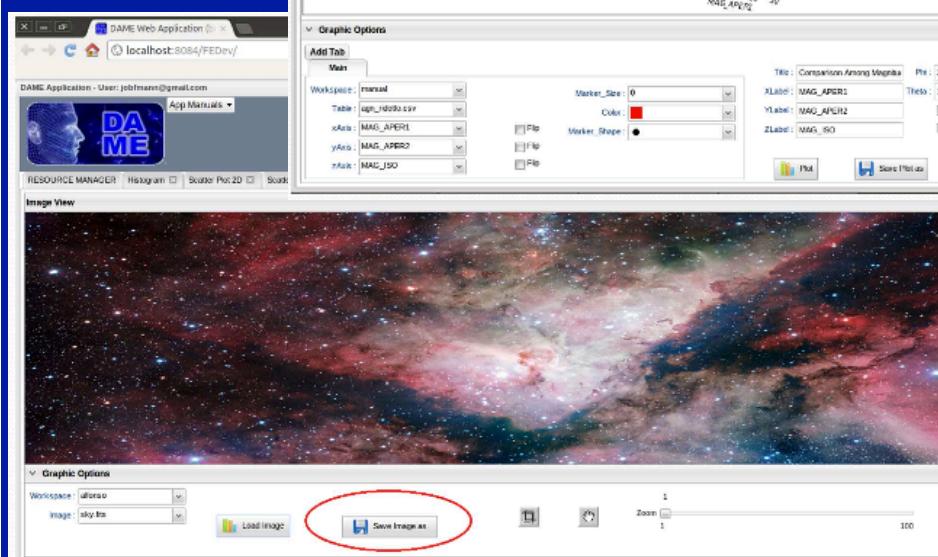
DAME - GAME Genetic Algorithm Mining Experiment

GAME is a pure genetic algorithm developed in order to solve supervised problems of regression or classification, able to work on Massive Data Sets (MDS).

Graphical capabilities in DAMEWARE

Histograms
2-D & 3-D plots
Line plots
Image visualization

Java client



AGN identification and classification

Photometric classification of emission line galaxies with Machine Learning methods, Cavuoti et al., 2013, MNRAS, submitted

Star/Galaxy separation

The detection of globular clusters as a data mining problem, Brescia et al., 2012, MNRAS, 421, 1155-1165 (arXiv:1110.2144)

GPUs for astrophysical data mining. A test on the search for candidate globular clusters in external galaxies. S. Cavuoti, et al., New Astronomy, april 20, 2013, <http://dx.doi.org/10.1016/j.newast.2013.04.004> (astro-ph: 1304.0597)

Photometric redshifts

Mining the SDSS archive. I. Photometric redshifts in the nearby universe, D'Abrusco, Logno G., Walton N., 2007, ApJ, 663, 752

Astroinformatics of galaxies and quasars: a new general method for photometric redshifts estimation , O. Laurino, R. D'Abrusco, G. Longo, and G. Riccio, MNRAS, 2011, 418, 2165 (arXiv/1107.3160);

Photometric redshifts with Quasi Newton Algorithm (MLPQNA) Results in the PHAT1 context, Cavuoti et al. 2012, , Astronomy and Astrophysics 546, 13, (ArXiv:1206.0876)

Photometric redshifts for quasars in multiband surveys, M. Brescia et al., 2013, ApJ, 772, 140 (astro-ph: 1305.5641)

Inside catalogs: a comparison of source extraction software, M. Annunziatella, et al., 2012, PASP, 125, 68 (astro-ph:1212.0564).

Other

Astroinformatics, data mining and the future of astronomical research, M. Brescia & G. Longo, 2012, invited to appear in proceed. of IFDT2 - 2nd International conference frontiers on diagnostic technologies (arXiv:1201.1867)

CLASPS: a new methodology for knowledge extraction from complex astronomical data sets, R. D'Abrusco, G. Fabbiano, S.G. Djorgovski, C. Donalek, O. Laurino & G. Longo, 2012, ApJ, 755, 92 (ArXiv: 1206.2919)

Current Applications in other fields

Medical diagnosis of alzheimer

(S. Cocozza et al.)



Brain tomography analysis

(Bellotti M. et al.)

Real time classification of ethernet data flows

(G. Ventre et al.)

Etc...

Some statistics (2013)

Ca. 100 users, > 12.000 experiments

An operative example

Use case:

Photometric redshifts evaluation for quasars in panchromatic surveys

Functionality: regression

Pre-processing: preparation of KB (10^5 objects)
removal of NaN,
splitting of train, validation, test sets

Feature selection (>50 experiments)

Selection of best DM model: SVM; MLPBP, MLP-GA, GAME, MLPQNA

Training

Validation (10 experiments)

Test

TOTAL of ca. 2000 experiments

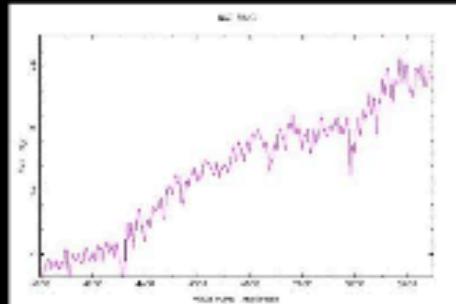
Visualization, comparison & Evaluation of results

Pucon, August 2013



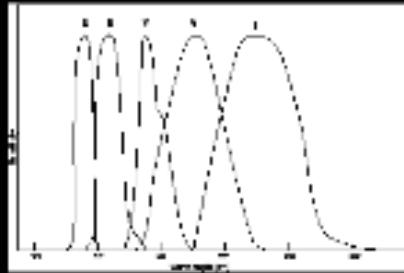
PHOTOMETRIC REDSHIFTS AS A INVERSE PROBLEM

Spectral Energy Distribution convolved with band filters



Galaxy spectrum - $F(\lambda)$

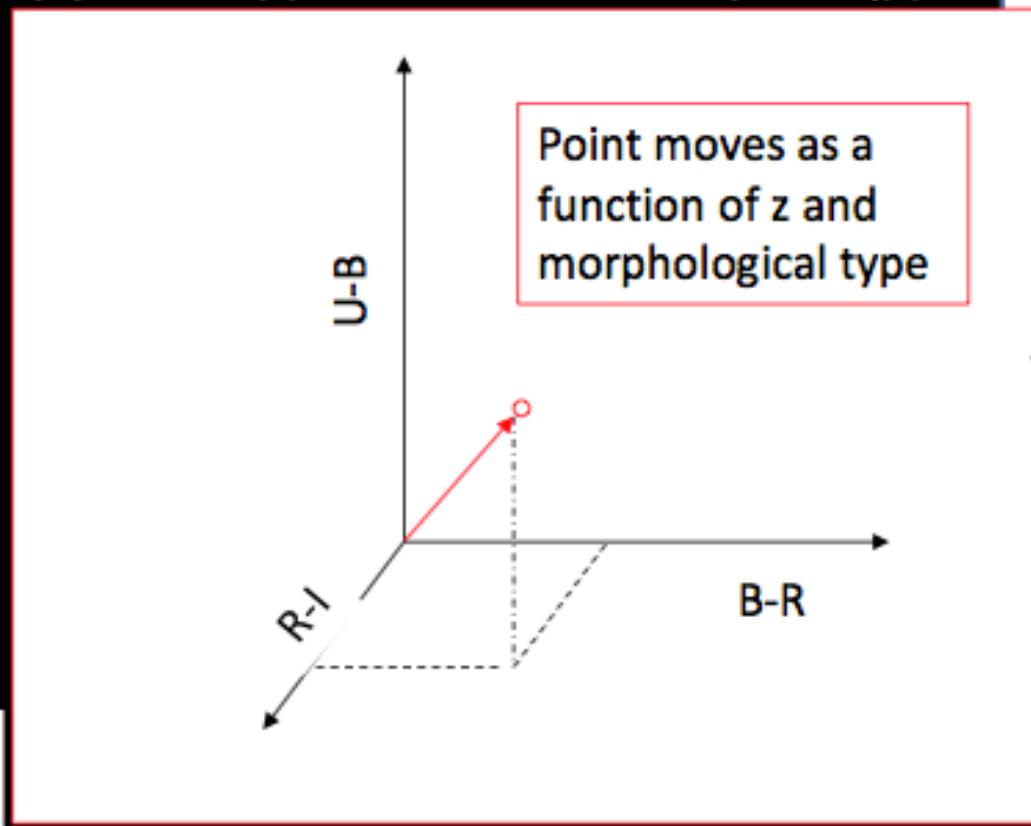
X



Photometric system - $S_i(\lambda)$

$$m_U = -2.5 \log_{10} \frac{\int F(\lambda) S_U(\lambda) d\lambda}{\int S_U(\lambda) d\lambda} + c_u$$
$$m_B = -2.5 \log_{10} \frac{\int F(\lambda) S_B(\lambda) d\lambda}{\int S_B(\lambda) d\lambda} + c_B$$

=



Color indexes

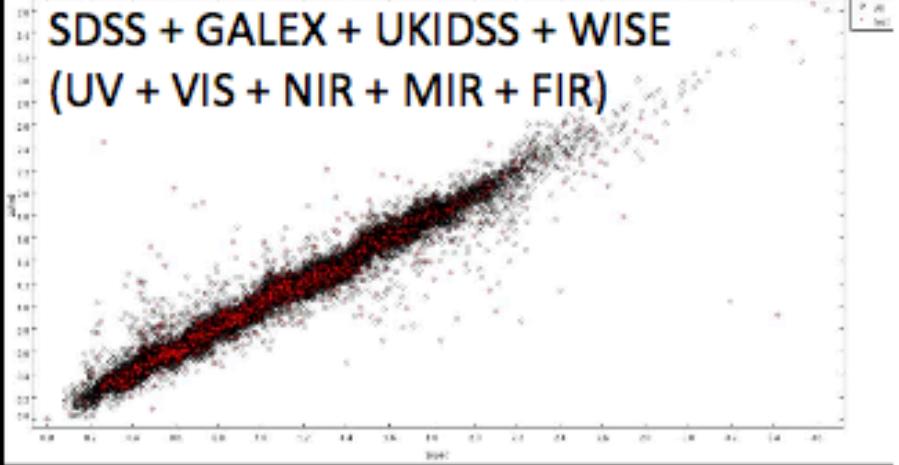
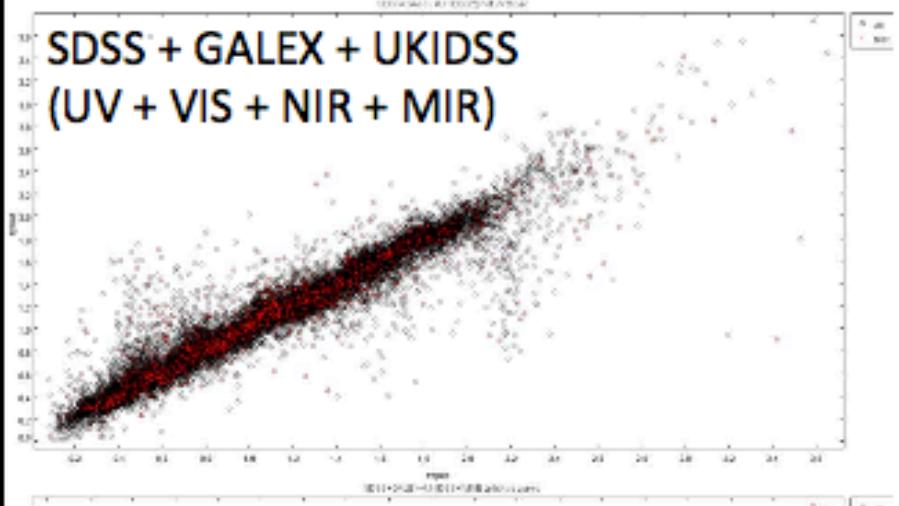
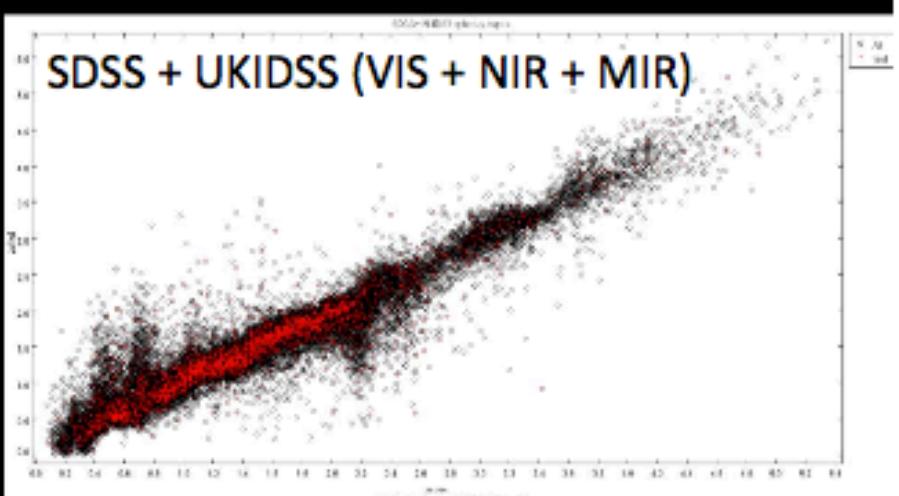
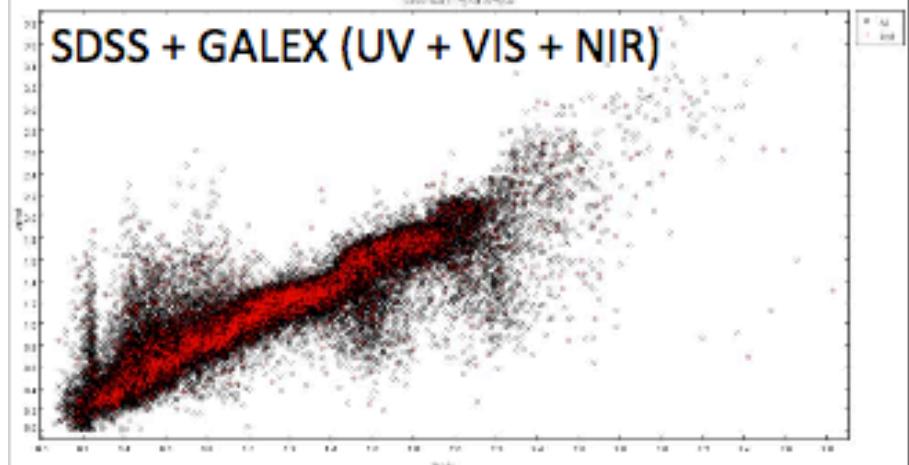
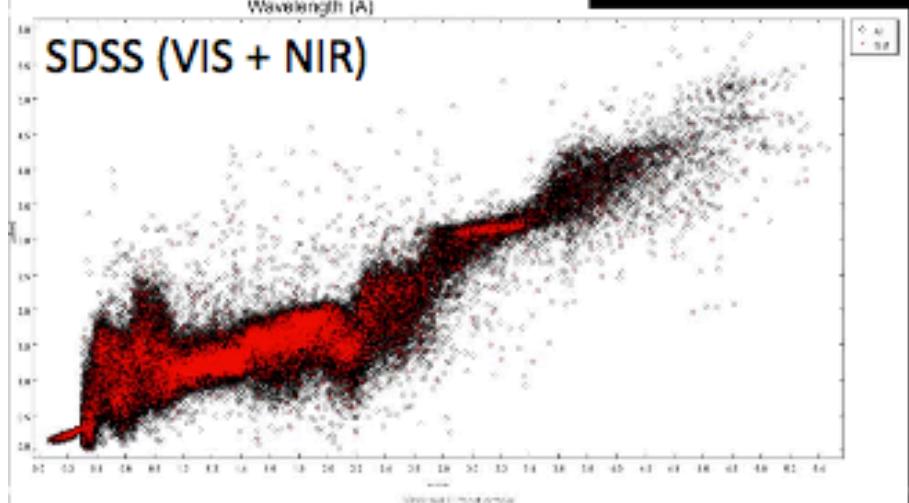
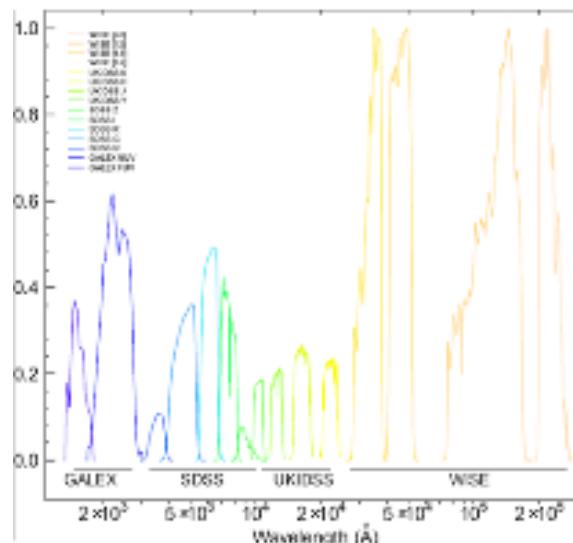
$$U - B \equiv m_U - m_B$$

$$B - R \equiv m_B - m_R$$

etc.

Phot-z are an inverse problem





TEST	MEAN	σ	out. 1 σ	out. 2 σ	out. 3 σ	out. 4 σ	TOTAL OBJECTS
E5	0,0005	0,118	18,67%	4,01%	1,51%	0,87%	3787
E16	-0,0004	0,154	18,11%	4,75%	1,98%	0,98%	3787

Table 3. Summary of the statistical indicators already used in Table xx (bias, σ and the percentage of outliers at, respectively, 1,2,3 and 4 σ computed as in citebovy2012 on all objects (test and training set).

ZSPEC BIN	EXP	BIAS	SIGMA	$ \Delta Z > 0.1$	$ \Delta Z > 0.2$	$ \Delta Z > 0.3$	$ \Delta Z > 0.4$	OBJECTS
TRAIN Only								
[0.2, 1.0]	E23	-0.0897	0.206	44.94%	22.78%	13.29%	8.86%	316
[0.2, 1.0]	E5	-0.0183	0.118	27.53%	7.28%	2.22%	1.27%	316
[0.2, 1.0]	E16	-0.029	0.127	29.43%	10.76%	3.48%	1.90%	316
[0.2, 1.0]	E10	-0.1807	0.281	64.87%	39.87%	26.58%	18.67%	316
[1.4, 3.0]	E23	0.1209	0.273	59.05%	32.33%	21.98%	14.22%	232
[1.4, 3.0]	E5	0.0364	0.18	38.36%	14.66%	8.19%	4.74%	232
[1.4, 3.0]	E16	0.0408	0.183	40.09%	15.52%	8.62%	4.74%	232
[1.4, 3.0]	E10	0.2188	0.367	62.50%	41.38%	28.88%	22.84%	232
TRAIN+TEST								
[0.2, 1.0]	E23	-0.0911	0.23	46.24%	23.18%	13.77%	9.03%	1583
[0.2, 1.0]	E5	-0.0174	0.101	21.04%	4.17%	1.58%	0.82%	1583
[0.2, 1.0]	E16	-0.0326	0.142	30.01%	9.85%	4.67%	2.65%	1583
[0.2, 1.0]	E10	-0.1877	0.287	63.93%	39.55%	27.48%	19.08%	1583
[1.4, 3.0]	E23	0.1238	0.269	56.24%	30.74%	18.69%	12.49%	1145
[1.4, 3.0]	E5	0.0271	0.139	31.44%	9.61%	3.93%	2.10%	1145
[1.4, 3.0]	E16	0.0492	0.183	39.83%	14.93%	7.95%	4.37%	1145
[1.4, 3.0]	E10	0.2488	0.37	64.28%	44.02%	32.23%	24.72%	1145

The new mantra

Discovery of rare and unknown...

Search for higher order correlations etc...

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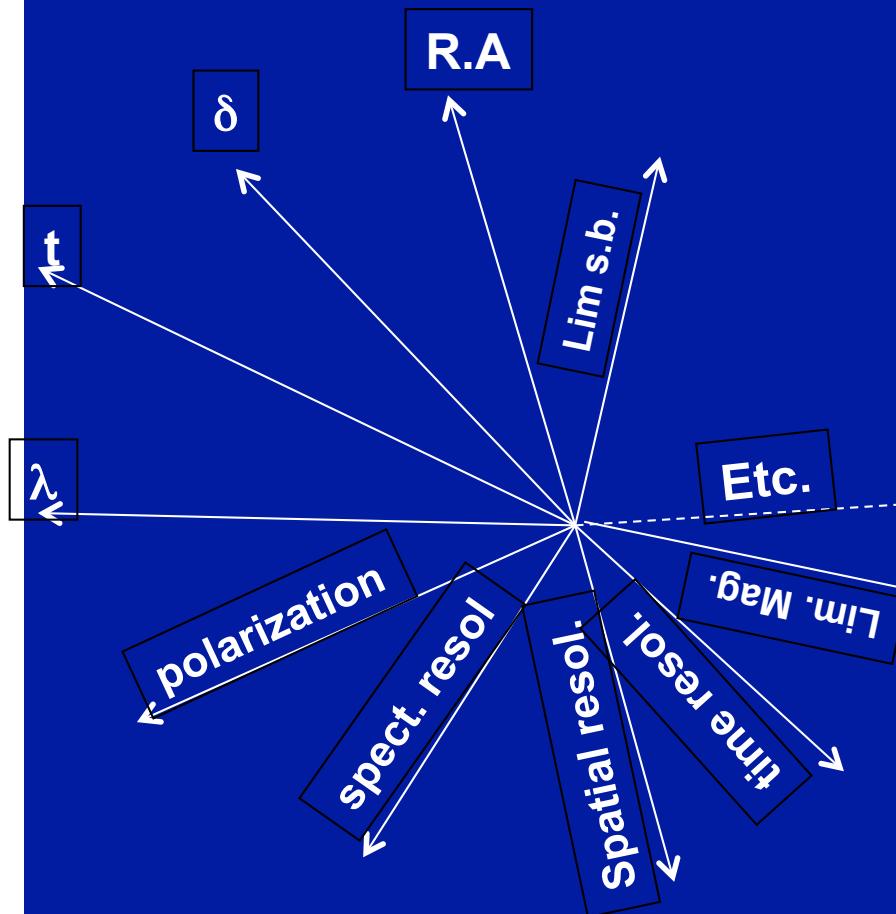
Small data sets: serendipity or guided luck

Large Data sets: clustering....

The new mantra

Discovery of rare and unknown...

Search for higher order correlations etc...



Each datum is defined by n measured parameters.

- X,y,t
- Flux
- Polarization
- wavelength
- Etc..

New sensor

$$p \in \Re^N \quad N \gg 100$$

Exploration of PS with $N > 10^9$, $D \gg 100$, $K > 10$ Is anything but simple

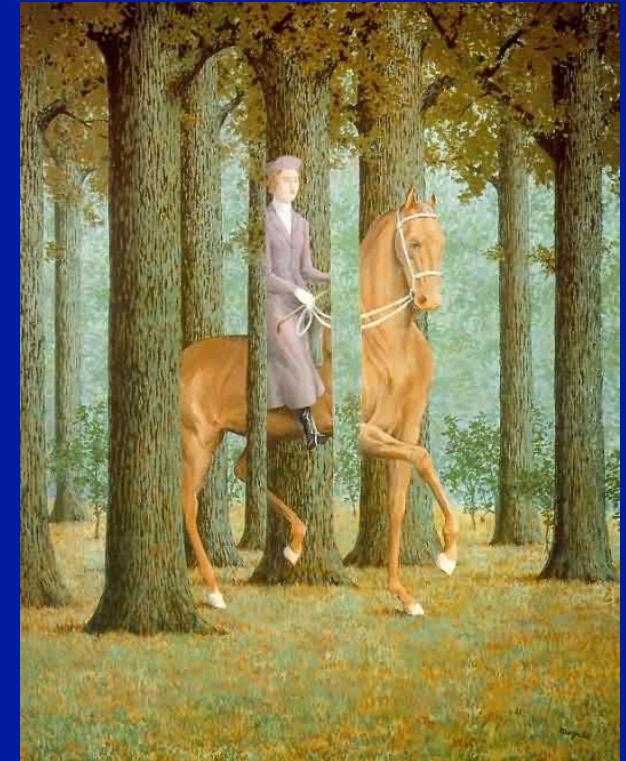
N = no. of data vectors,

D = no. of data dimensions

K = no. of clusters chosen,

K_{\max} = max no. of clusters tried

I = no. of iterations, M = no. of Monte Carlo trials/partitions



K-means: $K \times N \times I \times D$

Expectation Maximisation: $K \times N \times I \times D^2$

Monte Carlo Cross-Validation: $M \times K_{\max}^2 \times N \times I \times D^2$

Correlations $\sim N \log N$ or N^2 , $\sim D^k$ ($k \geq 1$)

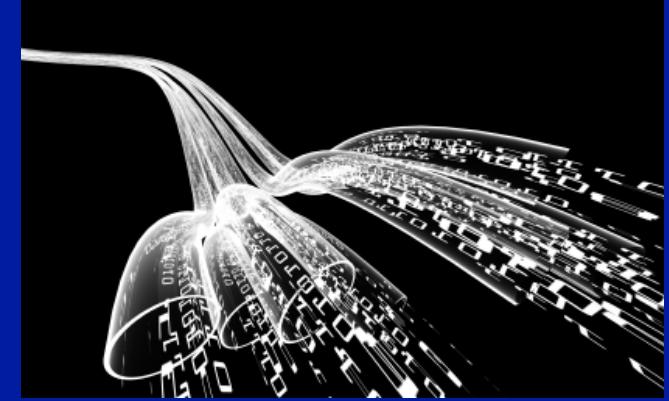
Likelihood, Bayesian $\sim N^m$ ($m \geq 3$), $\sim D^k$ ($k \geq 1$)

SVM $> \sim (N \times D)^3$

**Lots (...truly
lots and
lots...) of
computing
power**

Moving programs not data: the true bottle neck

Data Mining + Data
Warehouse =
Mining of Warehouse Data



- For organizational learning to take place, data from must be gathered together and organized in a consistent and useful way – hence, Data Warehousing (DW);
- DW allows an organization to remember what it has noticed about its data;
- Data Mining apps should be interoperable with data organized and shared between DW.

Interoperability scenarios



Full interoperability between DA (Desktop Applications)



Local user desktop fully involved (requires computing power)



Full WA → DA interoperability

Partial DA → WA interoperability (such as remote file storing)

MDS must be moved between local and remote apps

user desktop partially involved (requires minor computing and storage power)



Except from URI exchange, no interoperability and different accounting

MDS must be moved between remote apps (but larger

bandwidth) computing power required



The Lernaean Hydra KDD

After a certain number of such iterations...



The scenario will become:

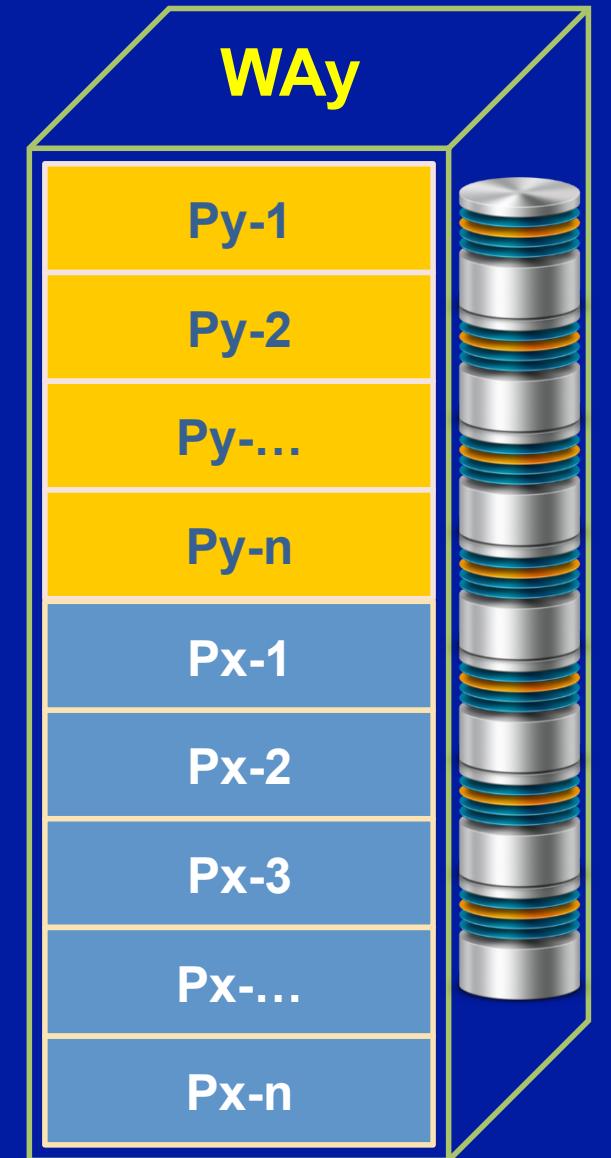
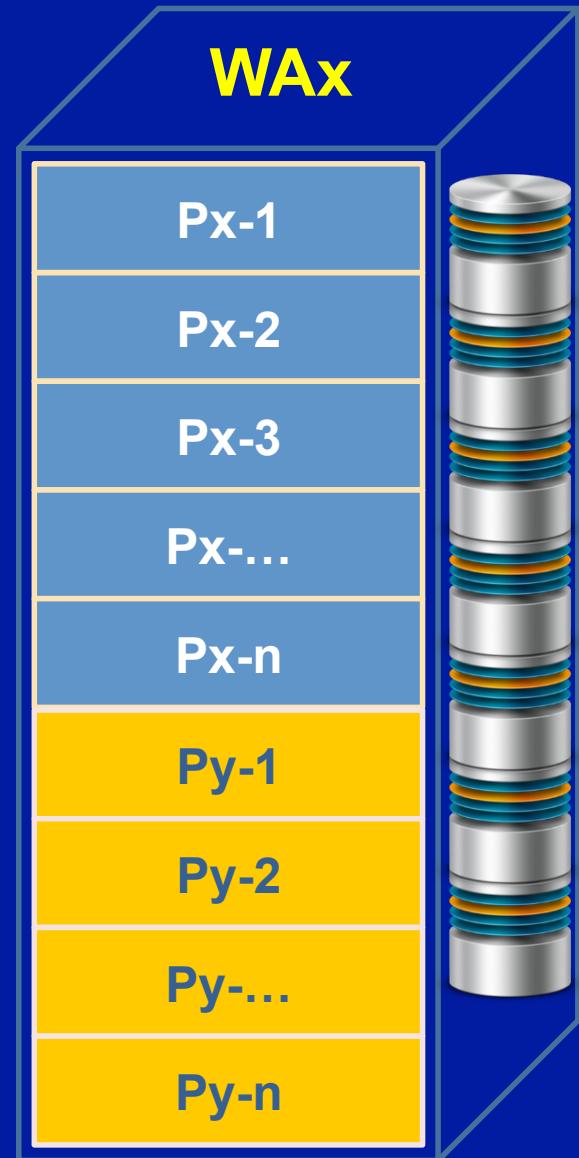
No different WSs, but simply one WS with several sites (eventually with different GUIs and computing environments)

All WS sites can become a mirror site of all the others

The synchronization of plugin releases between WSs is performed at request time

Minimization of data exchange flow (just few plugins in case of synchronization between mirrors)

YES
MDS!





The Data Growth Monster
How Much Digital Information Exists in the World Today?

astronomical problems are a piece of cake....

Growth of digital data worldwide (2012)
1 ZB/yr or = 10^9 Terabyte





NSA listening station in Bluffdale, South Dakota

1 YB of storage = 10^{12} TB
Indexed, searched, mined...

*With mainly unknown technology
which will slowly leak out to the
scientific community*

Pucon, August 2013

astronomical problems are a piece of cake....

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**Thanks for the attention ...
... and to Eduardo for organizing
the meeting ...**