

CONT08 - First Results and High Frequency Earth Rotation

T. Artz, S. Böckmann, A. Nothnagel

Abstract First results from the most recent continuous VLBI campaign (CONT08) are shown. CONT08 took place in August 2008. One of the scientific goals was the generation of high precision Earth Orientation Parameters (EOP) with sub-daily resolution. A general quality assessment of CONT08 is performed investigating station position variations and daily EOP. In addition, high-resolution Earth rotation time series are generated in a way that ensures consistency over the whole time span. Here, we demonstrate the effect of a modified scheduling procedure concerning (in-)consistency of sub-daily EOP at session borders. Results of prior continuous campaigns have differed significantly from each other in amplitude and phase of the spectral components. Thus, we also compare the CONT08 amplitude spectra with results from CONT02 and CONT05.

Keywords CONT08, High Frequency Earth Rotation

1 Introduction

The International VLBI Service for Geodesy and Astrometry (IVS) conducts continuous VLBI observations of up to 14 days duration (CONT) in irregular intervals. In contrast to the routine 24 h VLBI observations, which are performed two or three times per week, those CONT-campaigns are around-the-clock observations over several days. The last three CONT campaigns carried out in 2002, 2005 and 2008 took

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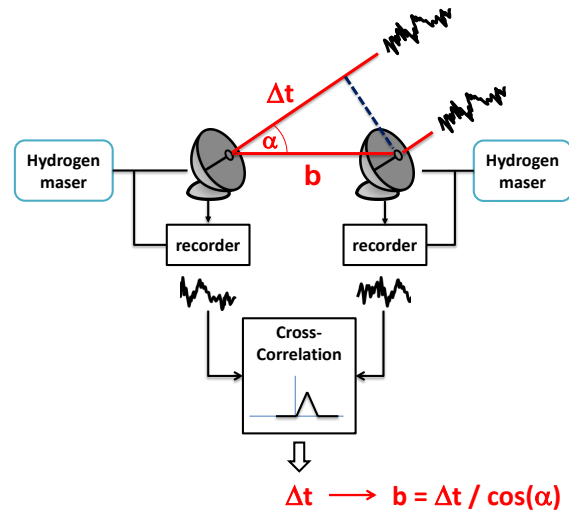


Fig. 1 CONT08 observing network

place over fortnightly time-spans. CONT-sessions should always demonstrate the highest available accuracy that can be achieved by VLBI observations. Additionally, one of the main scientific goals is the generation and interpretation of highly resolved Earth orientation parameters (EOP) due to the continuity of the campaign.

Most recently, CONT08 took place in 2008 from August 12-26 with individual sessions being set up for 0 h to 24 h UTC each. The observing network was arranged by 11 globally distributed VLBI telescopes (s. Fig. 1). The network has a strong European part since the stations in North-America at Fairbanks and Algonquin Park (that were part of prior continuous campaigns) were powered off and replaced by Medicina and Zelenchukskaya.

Station	Weekends [UT]	Weekdays [UT]
Hr	13-15	12-14
Kk	00-02	18-20
Mc	15-17	06-08
Ny	17-19	10-12
On	11-13	08-10
Sv	05-07	04-06
Tc	21-23	16-18
Ts	07-09	00-02
Wf	19-21	14-16
Wz	07-09	18-20
Zc	09-11	02-04

Table 1 Time slots for system checks during CONT08 (<http://ivscc.gsfc.nasa.gov/program/cont08/cont08-notes.txt>)

There are several changes in the scheduling of CONT08 with respect to prior continuous campaigns. The keypoint is the necessity of system checks that have to be performed on a daily basis. For CONT08 these have been decoupled in view on the observing sites. Prior to CONT08, the system checks were performed in the last ≈ 30 min of each session for all stations. Now, subsequent 2 h slots for each site have been planned for each session with exception of the first one. In those slots, the 30 min system checks should be performed (see Tab. 1). In reality, this idea has been fulfilled. Moreover, some stations performed system checks so quickly that nearly no gaps are visible in the observations.

Focussing on the observations itself, one can state that each session has an overall number of 9000 to 11000 observations. This is a great increase w.r.t. CONT02 (mean: 3000) and CONT05 (mean: 6000) due to a raise of the recording rate and, thus, shorter scan lengths. The variation between the individual sessions mainly depends on station problems. TIGO/Conception missed 3 sessions and had several interruptions of up to 6 h. Zelenchukskaya is absent in one session because of lost disks. Furthermore, post fit residuals of observations with Zelenchukskaya show bi-modal patterns suggesting the existence of sub-ambiguities. Since these could not be fixed, the observations affected have been eliminated from the solution.

In this paper, first results of the analysis of CONT08 observations are given with a special focus on the estimation of sub-daily EOP. In particular, the impact of the modified scheduling on the estimates is evaluated. Furthermore, CONT-campaigns of 2002, 2005 and 2008 are compared on the basis of station

position estimates and the frequency representation of polar motion.

2 Solution Description

Three different solutions have been performed for the analysis of CONT08 as well as of CONT02 and of CONT05.

1. Solution for session-wise station position estimates. The datum defect has been solved by NNR/NNT conditions w.r.t. ITRF2005 over the whole set of stations.
2. EOP solution with daily estimates of polar motion, ΔUT1 and their rates as well as nutation offsets.
3. EOP solution with hourly PM and ΔUT1 parameterized as continuous-piece-wise-linear-functions (CPWLF). Nutation parameters are fixed to a priori values estimated in a separate solution.

In solutions 2. and 3., the station positions are estimated only once for the middle epoch of the fortnightly time-span to de-correlate the estimates as shown by Artz et al. (2007).

All of these solutions have in common the same modelling and parameterization of nuisance parameters (Clocks and atmospheres). Source positions are fixed to ICRF and its extensions. Station clocks are estimated w.r.t. Kokee Park clock by a 2nd order polynomial with additional clock parameters modelled by 60 min CPWLF. Troposphere parameters are estimated as CPWLF, too. The zenith wet delay is parameterized with a temporal resolution of 20 min and gradients in 12 h intervals.

For all of these solutions a priori EOP are taken from USNO finals¹. Nutation is modelled by IAU2000A (McCarthy and Petit, 2004) plus additional corrections from a global VLBI solution. Ocean loading is modelled according to FES2004 (Letellier et al., 2004), furthermore, thermal expansion of the radio telescopes (Nothnagel, 2008) and atmospheric pressure loading (Petrov and Boy, 2004) have been applied.

The solutions presented here are all run in a two step procedure. First, the data is processed by the VLBI analysis software Calc/Solve (Petrov, 2008) and the

¹ <ftp://maia.usno.navy.mil/ser7/finals.daily>

3 Quality Assessment

The differences of the individual station position estimates and the campaign mean are quite homogeneous (Fig. 2). Most of them are below 10 mm for the horizontal and 20 mm for the up component. Only the variability of the east component of Kokee Park and Tsukuba is large compared to the other sites. Nevertheless, the impact on the estimation of other parameters is not big e.g. excluding these sites from the datum definition has nearly no impact on the other station positions. The bigger deviations of TIGO/Conception in the session of August 18th can be explained by the minor contribution to this session. Here, TIGO had only a few hours of observations due to some station problems and, thus, the error bars of these estimates are huge. As a consequence, there is no big impact on the estimation of other parameters as well.

Mk4/DiFX fourtil 3.9 rev 6188 **142715.ydvqcs, No. 0090, LV**
ONSALDES - WETZEL, Group L, pol RR

Amplitude

delay rate (ns/s)

delay (us)

singleband delay (us)

Amp. and Phase vs. time for each freq., 82 segs, 1 APs / seg (0.16 sec / seg.), time ticks 1 sec

Amplitude

delay (us)

Avgd. Xpower Spectrum (MHz)

Exp. # 168383
Yr/day 2014.018
Start 142716.92
Stop 142720.04
FRT 142722.00
Corr/FreqHz
S 2015:341:115400
2015:134:125811
2014:204:120420
RA & Dec (J2000)
000513m329.18s
+24 53'11.345"

Group delay (sec)(mode) **-1.41413838746E+03** **Apriori delay (sec)** **-1.41413797819E+03** **Resid mbdelay (usec)** **-3.91251E-04** **1. 6.0E-04**
amp/dly (sec) **-1.414025217E+03** **Apriori clock (usec)** **1.24183E-02** **Resid phdelay (usec)** **2.46847E-04** **1. 1.7E-08**
delay rate (sec) **-1.41413772698E+03** **Apriori clockrate (usec)** **0.000000000E+00** **Resid phdelay (usec)** **2.46847E-04** **1. 1.7E-08**
Delay rate (sec) **4.728770984E-01** **Apriori phase (usec)** **4.725769690E-01** **Resid rate (usec)** **2.98847E-04** **2. 2.3E-07**
Total phase (deg) **83.1** **Apriori ecorr (usec)** **4.61625864432E-01** **Resid phase (deg)** **141.7** **1.0 1.0**

freq (MHz)
Group
Apriori
Snd box
APs used
PC track
PC track
PC track
Mean PC
PC rms
Chain lts
Chain lts

phasing (deg) **6.3** **4.5** **Amplitude** **93.97** **at 0.815**
amp/dly (sec) **6.3** **4.5** **Search** **(25008)** **93.468**
amp/dly (sec) **6.5** **5.9** **Inc. seg. amp.** **93.958**
amp/dly (sec) **6.5** **5.9** **Inc. low amp.** **93.678**
L az 142.8 el 53.5 pa 21.1 **V az 135.8 el 60.1 pa 3.1** **u/v (Hz/sec)** **5.900 20.492**

File name: **McM4ML, NORMML (rds)** **Pcal period (APs)** **10** **-8.000 8.000**
Full rate: **3.702E-04, 6.740E-04 (rds)** **Full rate:** **3.702E-04, 6.740E-04 (rds)** **0 window (rds)** **0.000 0.000**
Sample rate(MHz): **16** **Sample rate(MHz):** **16** **0 window (rds)** **-1.959 1.959**
Rate rate(MHz): **32** **Rate rate(MHz):** **32** **0 window (rds)** **-0.900 0.900**

simultaneous integration

Fig. 3 Station position repeatabilities for each campaign. The datum definition has been applied by NNR/NT conditions over all stations in CONT02 (red) and over all stations but Kokee Park and Tsukuba in CONT05 (black) and CONT08 (purple). The two stations that were active in CONT08, only, (Medicina and Zelenchukskaya) are excluded from this figure.

Calculating the WRMS of the daily EOP differences between the VLBI estimates and external EOP series leads to an insight in the quality of the EOP from a continuous campaign. These WRMS differences are shown in Fig 4, where the official combined EOP series of the International GNSS Service² has been used

² <ftp://igscb.jpl.nasa.gov/pub/product/igs00p03.erp.Z>

Fig. 4 EOP repeatabilities for each campaign. The WRMS values are calculated w.r.t. IGS EOP series for polar motion and USNO finals for $\Delta UT1$.

Fig. 5 X pole differences to IGS with a hourly resolution from CONT05 (upper plot) and CONT08 (lower plot) from an independent solution.

for polar motion. For $\Delta UT1$ the USNO finals have been used to calculate the EOP repeatabilities. The differences have been detrended to eliminate the impact of differences in the underlying terrestrial reference frame. The UT1 variations from the three continuous VLBI campaigns agree with the reference series at the same level, whereas the CONT08 polar motion agrees much better with the IGS time series than those from CONT02 or CONT05.

4 Sub-daily Earth Rotation

The estimation of EOP with a sub-daily resolution from CONT campaigns is of great interest. Hourly EOP estimates over a time-span of two weeks provide the opportunity to analyze the characteristics of the EOP in the frequency domain.

Furthermore, the time series of hourly EOP shows the effects of the modified scheduling used in CONT08. Figure 5 displays the X pole differences w.r.t. the IGS time series for CONT05 and CONT08 from a solution where each session is analyzed independently. There are huge outliers in the CONT05 time series at the session borders. Those are due to the lack of observations in the last ≈ 30 min of each 24 h block where the system checks were performed. Such outliers are not visible in the CONT08 time series. For CONT08 the scheduling was changed in a way that the observations are really continuous as described in Sec. 1. Thus, the estimates in the last interval of each individual session are already stabilized through this type of scheduling.

In addition to the scheduling, the modified solution strategy of stacking the normal equations helps to further improve the sub-daily EOP estimates. Here, parameters at the session boundaries are estimated using observations from the last interval of the first session and the first interval of the second session and so on.

Fig. 6 X pole differences to IGS with a hourly resolution from CONT05 (upper plot) and CONT08 (lower plot) where the modified solution strategy has been applied.

Fig. 7 Polar Motion spectra from CONT-campaigns.

The resulting time series are shown in Fig. 6. Nearly all of the outliers in CONT05 are eliminated or at least minimized by the modified stacking approach. A minor improvement can be seen for CONT08 as well.

In the spectral domain several authors have reported about a retrograde ter-diurnal signal in the CONT02 Polar Motion data (e.g. Haas and Wunsch (2006)). This signal is visible in our analysis, too, as it is shown in Fig. 7. However, there is no such signal present in the other campaigns. To derive the spectra, total EOP series are reduced by the IERS sub-daily tidal model (McCarthy and Petit, 2004). The reason for this phenomena is still an open issue and under investigation. Subsequently, those values have been detrended by CPWLF with a resolution of one day. Afterwards, a FFT of the residuals has been calculated. Some peaks can be seen at the diurnal and semi-diurnal bands. The prograde diurnal and semi-diurnal as well as the semi-diurnal retrograde signal can be interpreted as inconsistency between the IERS sub-daily tidal model and the VLBI observations. The remaining retrograde diurnal signal is due to mismodelling of nutation. Nutation is fixed to a priori values for the estimation of sub-daily EOP, thus, some of the remaining signal appears as a near diurnal retrograde term. In the ΔUT spectra no significant variations besides the diurnal and semi-diurnal bands can be seen. There is a small irregular variation at a period of 6h in the CONT05 data only.

5 Conclusions

CONT08 serves as a high quality continuous data-set from VLBI observations. The variations of the station position estimates is quite homogeneous besides some bigger variations in the east component of Kokee Park and Tsukuba.

The modified scheduling improves the estimation of parameters with a sub-daily resolution as shown for the hourly EOP estimates. Not only EOP but all param-

eters with a sub-daily resolution as zenith wet delay or clock parameters benefit from this procedure.

The sub-daily EOP estimates match the GPS series quite well for all campaigns. The derived spectra for polar motion are inhomogeneous. The peaks at the well known tidal bands with periods around 12h and 24h vary. Furthermore, the retrograde ter-diurnal term is present in CONT02 polar motion but not in CONT05 or CONT08. For variations in ΔUT , no significant irregular variations could be found.

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