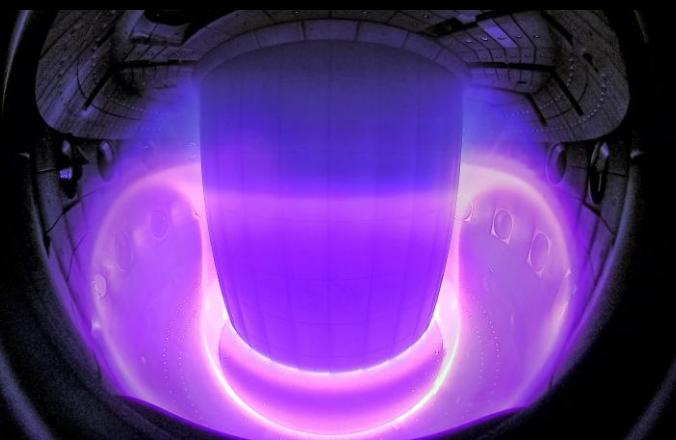
The background of the image features a plasma ball, likely a Van de Graaff generator, with numerous blue and purple glowing filaments radiating outwards from a central red sphere. The filaments are more concentrated at the top and bottom of the sphere, creating a starburst effect. The overall aesthetic is futuristic and scientific.

A Very Basic Introduction To Plasma Physics

What is a plasma

- A fully or partially ionised quasi-neutral state of matter
- Looking at a plasma on a macro scale it is neutral (quasi-neutrality)
- Look at it on a micro scale then you would see a soup of ions, electrons and maybe neutrals



Ionisation and Plasma Formation

- Ionisation of a gas occurs when a threshold energy is met for a given element (Xe: 12.1eV, Kr: 14.0eV, Ar: 15.8eV, Zn: 9.4eV)
- Several methods to cause ionisation for a plasma
 - Direct impact ionisation:
$$e^- + I \rightarrow 2e^- + I^+$$
 - Photo ionisation:
$$h\nu + I \rightarrow e^- + I^+$$
 - Thermal ionisation:
$$I \rightleftharpoons e^- + I^+, kT > \epsilon$$

Ionisation and Plasma Formation

Radiative Recombination:



Three Body Recombination:



- Ionisation fraction [χ] and equilibrium condition:

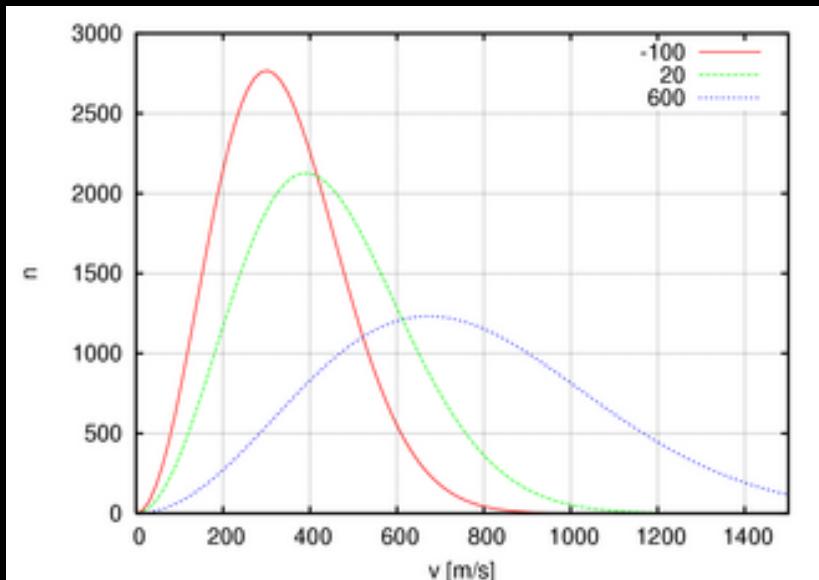
$$\chi = \frac{n_i}{n_i + n_n}$$

$$n_i \langle \sigma_r v \rangle = n_n \langle \sigma_i v \rangle, n_i \approx n_e$$

Debye Shielding and Quasi-Neutrality

- Boltzmann distributions: describes number of particles in a system at a given speed at thermal equilibrium

$$n_e = n_0 \exp\left(\frac{e\phi}{k_B T_e}\right) \approx n_0 \left(1 + \frac{e\phi}{k_B T_e}\right)$$



Debye Shielding and Quasi-Neutrality

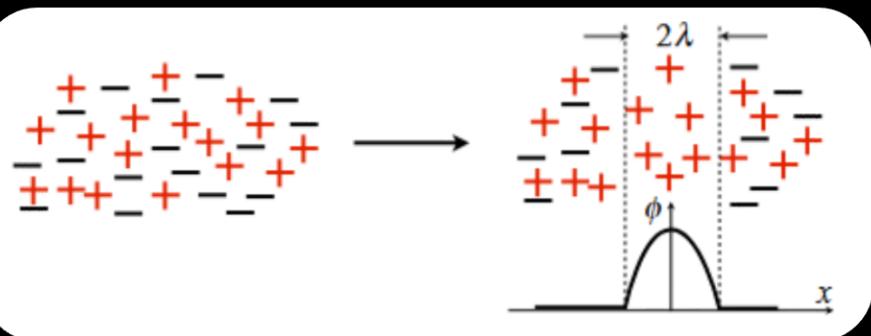
- Consider Gauss' law:

$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0}, \vec{E} = \nabla V \Rightarrow \nabla^2 V = \frac{\rho}{\epsilon_0}$$

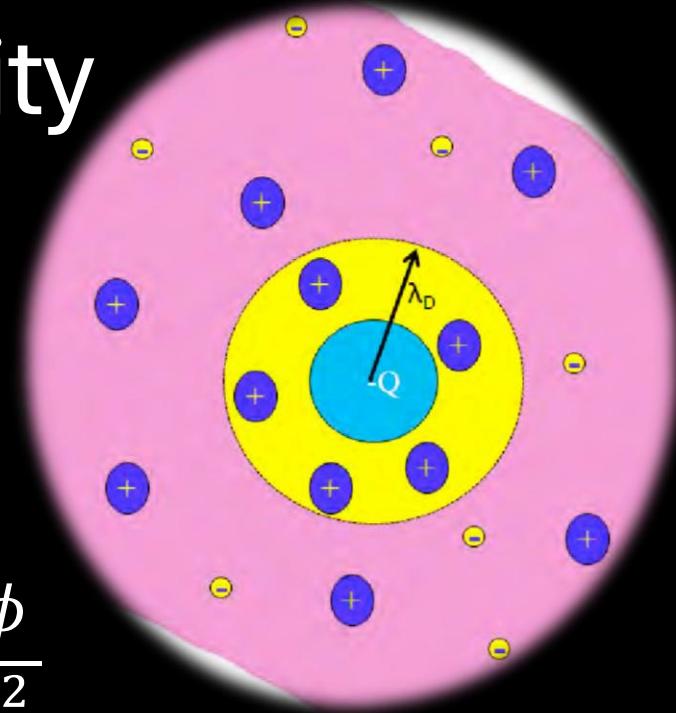
- Assume that ions are fixed with density $n_i \approx n_0$

$$\rho = -\frac{n_0 e^2}{k_B T_e} \phi \Rightarrow \nabla^2 V = \frac{1}{\epsilon_0} \cdot -\frac{n_0 e^2}{k_B T_e} \phi = \frac{\phi}{\lambda_D^2}$$

- Debye Length: The characteristic length over which electric potentials are screened in a plasma



$$\lambda_D = \sqrt{\frac{\epsilon_0 k_B T_e}{n_e e^2}}$$



Plasma Frequency and Timescales

Consider a displaced cloud of electrons:

Restoring force (per unit volume):

$$f = -en_e E$$

Equation of Motion for the cloud:

$$m_e n_e \frac{d^2 x}{dt^2} = -en_e E \Rightarrow m_e \frac{d^2 x}{dt^2} + eE = 0$$

$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0} = E = \frac{\rho}{\epsilon_0} x \text{ in 1D}$$

$$m_e \frac{d^2 x}{dt^2} + e \frac{\rho}{\epsilon_0} x = \frac{d^2 x}{dt^2} + e \frac{en_e}{m_e \epsilon_0} x \Rightarrow \omega^2 = \frac{e^2 n_e}{m_e \epsilon_0}$$

Plasma Frequency and Timescales

- Assume a quasi-neutral plasma $n_e = n_i = 10^{10}$
- Mass of ion and electron: $10^{-25}, 10^{-31}\text{kg}$
- Assume ion is singularly charged

$$\omega_i = \sqrt{\frac{e^2 n_i}{m_i \epsilon_0}} = 17\text{kHz}, \omega_e = \sqrt{\frac{e^2 n_e}{m_e \epsilon_0}} = 17\text{MHz}$$

$$\lambda_D \cdot \omega = v_{th} = \sqrt{\frac{k_B T_e}{m_e}}$$

Collisionality and Mean Free Path

- Mean free path: average distance a particle travels before colliding with another particle

$$\lambda = \frac{1}{n\sigma}$$

- Neutron-electron collisions:

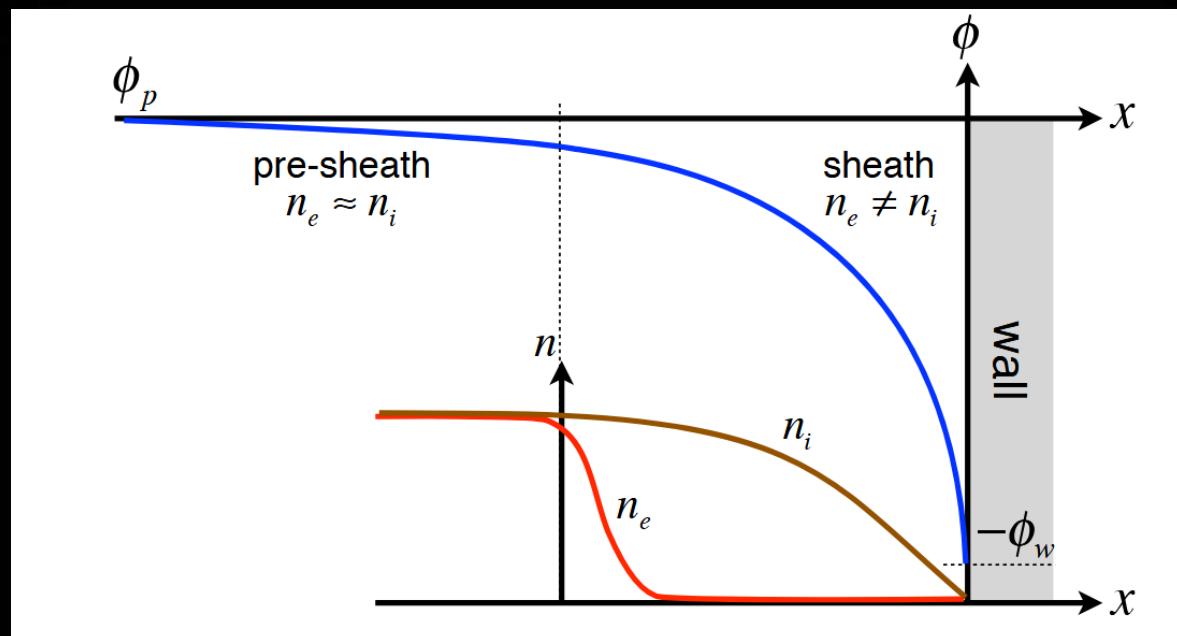
$$\lambda = \frac{v_n}{n_e \langle \sigma v_e \rangle}$$

- Describes the distance a neutral moves before it becomes ionised forming a plasma
- Collision frequencies: rate of collisions per unit time

$$\nu = \frac{v_{th}}{\lambda}$$

Plasma Sheaths

- Materials in contact with the plasma will become negatively charged with respect to the bulk plasma
- Ions directed at the wall will eventually hit it
- Electrons will lower energy than ϕ_w will be reflected away from it



Plasma Sheaths

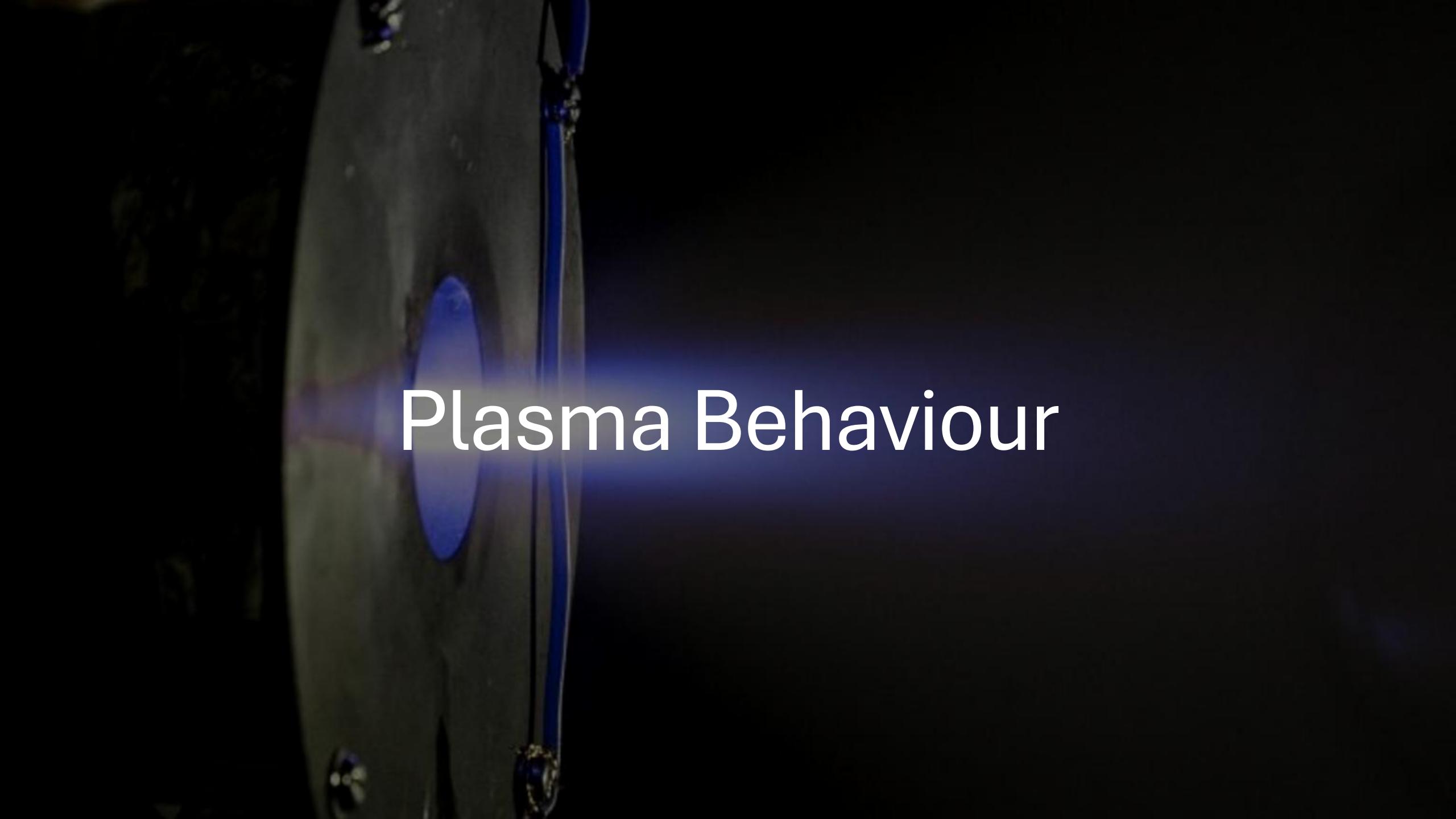
Bohm Criterion

- For a sheath to form and be stable ions must enter with sufficient speed

$$v_i \geq c_s = \sqrt{\frac{k_B T_e}{m_i}}$$

To describe the potential at which the plasma self biases (no current flow between plasma and walls)

$$\phi = \frac{k_B T_e}{e} \ln \sqrt{\frac{2m_i}{\pi m_e}}$$



Plasma Behaviour

Kinetic Description

- For modelling individual particle behaviour this method is often employed
- It is frequently used in collision less situations
- The distribution function:
 $f_s(r, v, t)$ r = position (3D) , v = velocity (3D), t = time (1D)
Where s can be e, i, or n

Kinetic Description

- The following macroscopic quantities can be extracted
 - Number density: $n_s(r, t)$
 - Mean velocity: $\mathbf{u}_s(r, t)$
 - Temperature: $n_s(r, t)$
 - Pressure: P_s

Kinetic Description

- The time dependent Boltzmann equation is as written:

$$\frac{\partial f_s}{\partial t} + \vec{v} \cdot \nabla_{\vec{r}} f_s + \frac{q_s}{m_s} (\vec{E} + \vec{v} \times \vec{B}) \cdot \nabla_{\vec{v}} f_s = \text{collisions}$$

- $\frac{\partial f_s}{\partial t}$
 - Is the local time rate of change of the function
- $\vec{v} \cdot \nabla_{\vec{r}} f_s$
 - How the particles move due to their velocity
- $\frac{q_s}{m_s} (\vec{E} + \vec{v} \times \vec{B}) \cdot \nabla_{\vec{v}} f_s$
 - Force term which describes how particles velocities change due to the Lorentz force
- E and B fields can be internal and external so must be calculated through maxwells equations

Fluid Models

- First consider the classical Navier-Stokes Equations
 - Continuity:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{u}) = 0$$

- Momentum:

$$\rho \left(\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} \right) = -\nabla p + \mu \nabla^2 \vec{u} + \rho g$$

Fluid Models

- Similar equations can be used to model plasmas on a macro scale
 - Continuity:

$$\frac{\partial n_s}{\partial t} + \nabla \cdot (n_s \vec{u}_s) = S_s$$

- Momentum:

$$m_s n_s \left(\frac{\partial \vec{u}_s}{\partial t} + \vec{u}_s \cdot \nabla \vec{u}_s \right) = -\nabla p_s - m_s n_s \nu_s (\vec{u}_s - u_{neutral}) + q_s n_s (\vec{E} + \vec{u}_s \times \vec{B})$$

Fluid Models

- Further equations are required to fully describe plasmas nature
 - Energy (simplified):

$$\frac{3}{2} n_s k_B \frac{dT_s}{dt} = Q$$

- Field Coupling:

$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0} = \frac{e(n_i - n_e)}{\epsilon_0} \Rightarrow \nabla^2 \phi = \frac{e(n_i - n_e)}{\epsilon_0}$$

Fluid Models

Term	Navier Stokes	Plasma Fluid
Inertia	$\rho \left(\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} \right)$	$m_s n_s \left(\frac{\partial \vec{u}_s}{\partial t} + \vec{u}_s \cdot \nabla \vec{u}_s \right)$
Pressure Gradient	$-\nabla p$	$-\nabla p_s$
Body Forces	ρg	$q_s n_s (\vec{E} + \vec{u}_s \times \vec{B})$
Internal viscous forces	$\mu \nabla^2 \vec{u}$	$-m_s n_s v_s (\vec{u}_s - \vec{u}_{neutral})$
Coupling	No coupling	$\nabla^2 \phi$
Multiple Species	Could have multiple species if multi-phase or a mixture but usually just one	Multiple species: ions, electrons and neutrons

Model Comparison

Factor	Kinetic Model	Fluid Model
Description	Describes individual particles and collisions between particles	Describes the plasma at a more macroscopic level capturing less detail
Assumptions	Very few assumptions, tries to capture all details	Assumes an averaged temperature and a near Maxwellian distribution
Computational cost	Very high	Less than kinetic
Physics	Captures all physics	Only captures bulk motion