Covariances Between ISM and Timing Parameters in Millisecond Pulsars

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Refractive effects in the ISM

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From Shannon 2011 (PhD Dissertation)

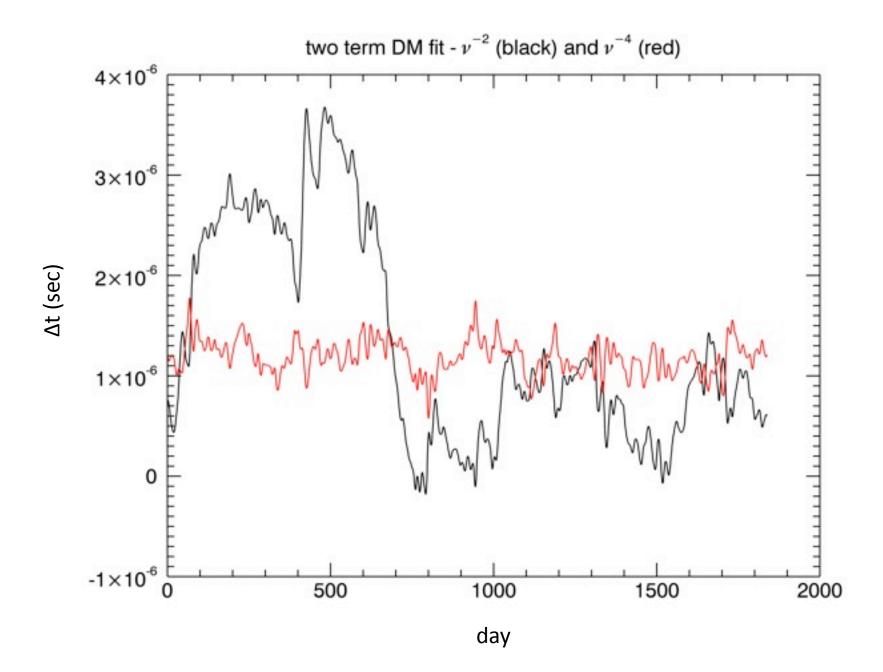
- Time delay due to angle-of-arrival θ is α θ^2
- Frequency dependence: θ «ν -²
- $\Delta t_{AOA} \propto v^{-4}$
- Here we assume Kolmogorov power spectrum for (isotropic) spatial fluctuations in refractive screen
- Focus here on refractive effects time delays due to diffraction as well
- Result varying DM as well as refractive component

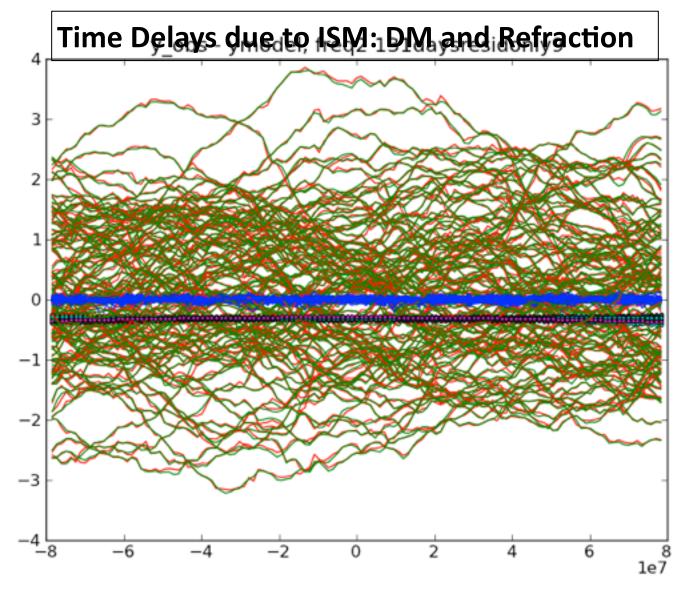
Model for timing residuals

- N_{days} = 128 (about every 2 weeks over 5 yr)
- Linear least squares no iterations

$$\Delta_t(\nu) = A + Bt_i + Ct_i^2 + \frac{DM(t_i)}{2}$$

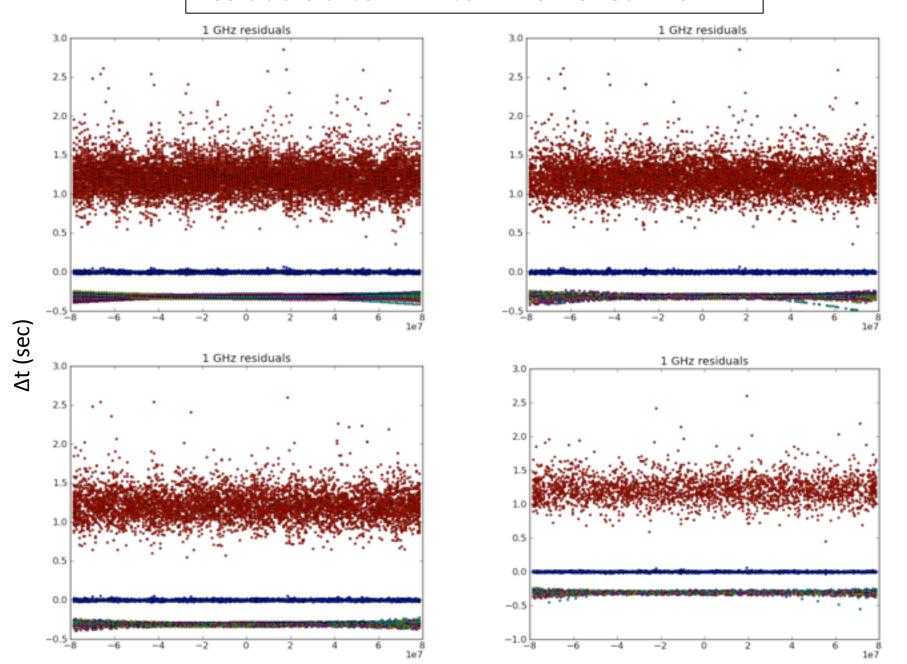
- DM at each time is a separate parameter for the model – 131 parameters total
- Time cadence is varied to be more and more "Poisson" -like



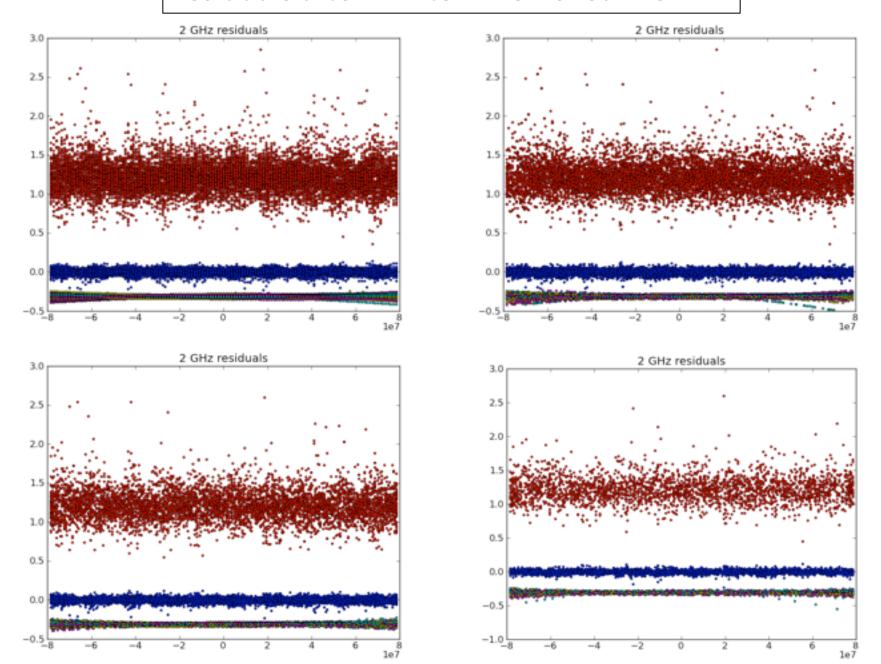


day

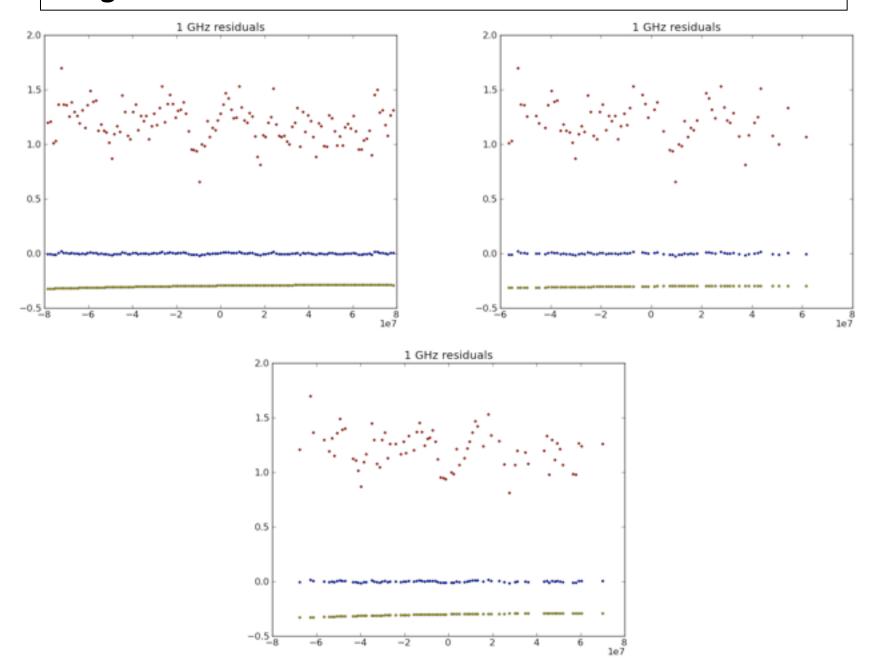
Residuals after DM term removed: 1GHz



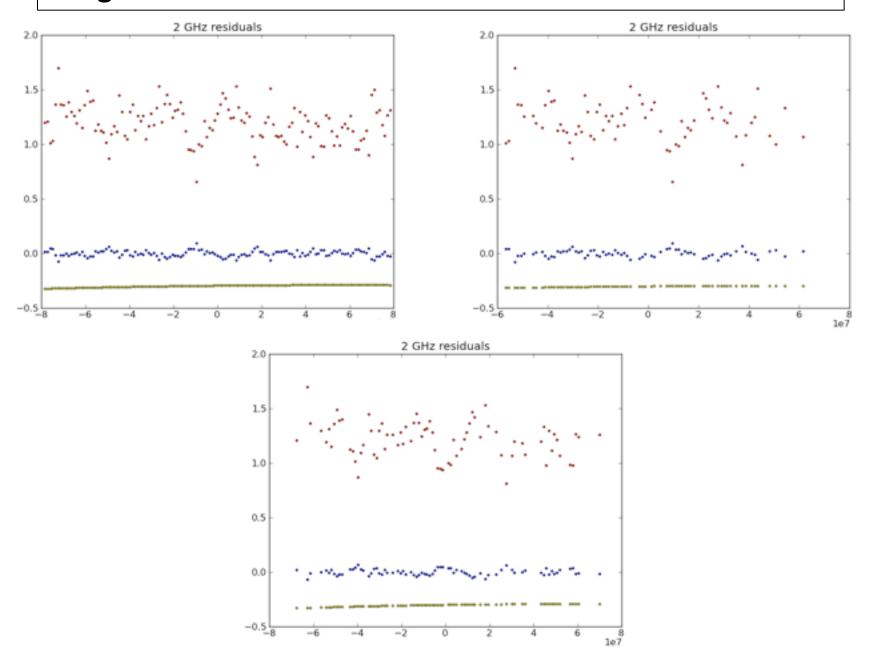
Residuals after DM term removed: 2GHz



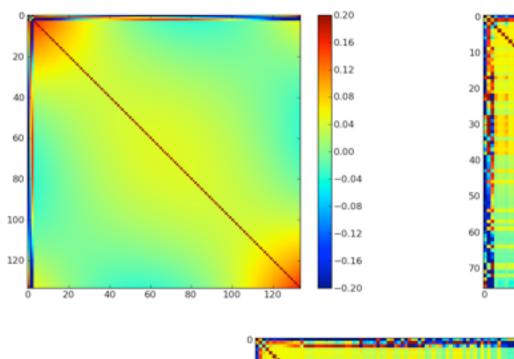
Single iteration residuals after DM term removed: 1 GHz

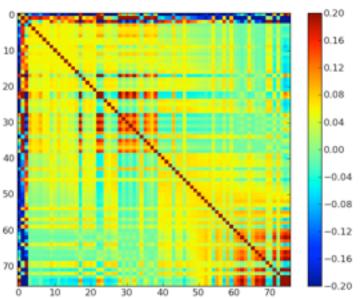


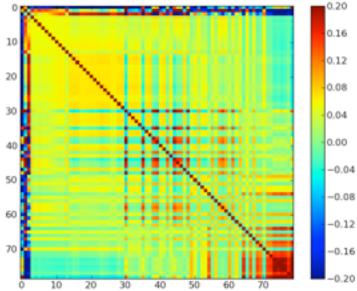
Single iteration residuals after DM term removed: 2 GHz



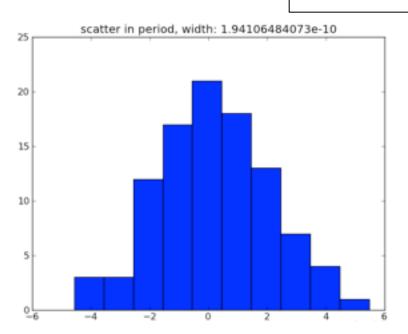
Covariance Matrix of Model Parameters

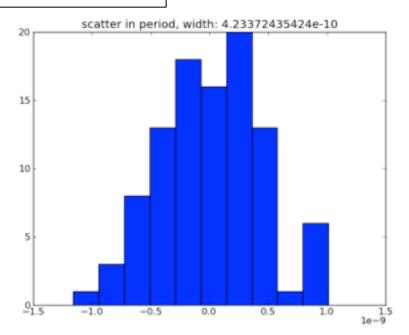


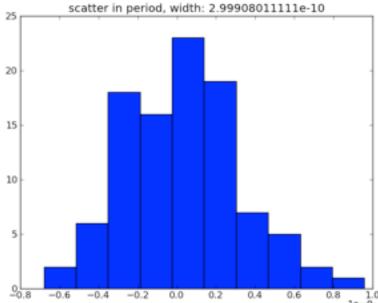




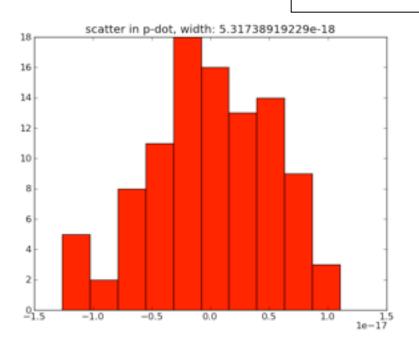
Scatter in B Parameter

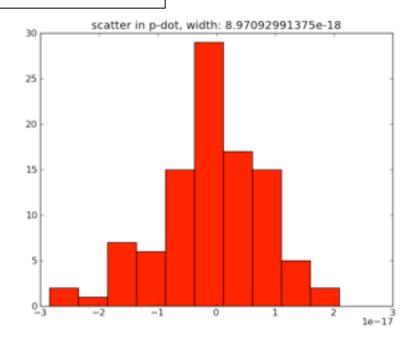


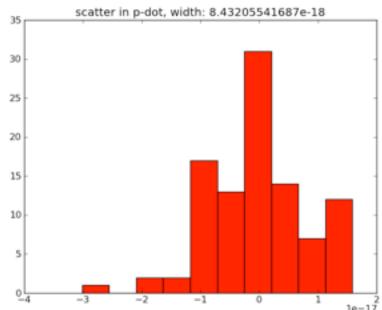




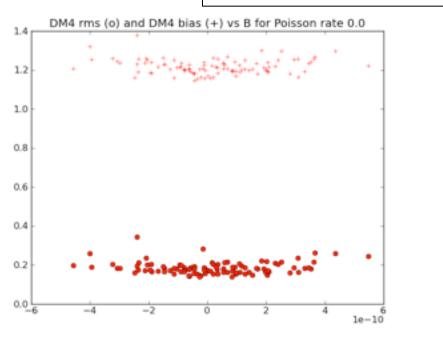
Scatter in C Parameter

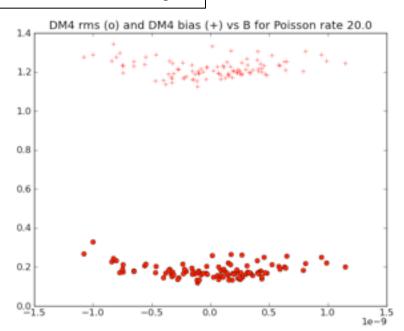


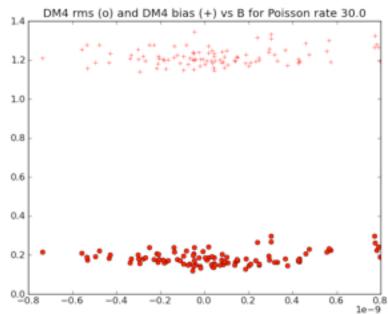




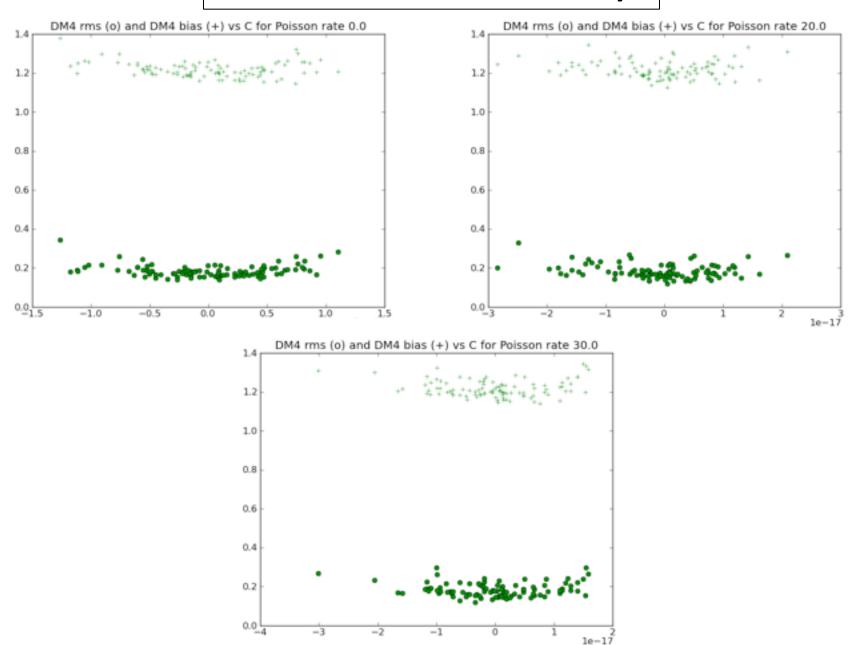
Scatter in Refractive Time Delay







Scatter in Refractive Time Delay



Towards modeling the ISM beyond DM

- Beyond Kolmogorov fluctuation spectrum
- Introduction of anisotropies
- as timing residuals approach 10ns, explicit modeling of refractive and diffractive effects should be employed in addition to DM variations

