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**Solvation Free Energy Calculations of
Molecules Mimicking Asphaltenes Using The
SAFT- γ Mie Force Field**

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Master's thesis presented to Engenharia de Processos Químicos e Bioquímicos graduate program, Escola de Química, Universidade Federal do Rio de Janeiro, as required for obtaining a Master's degree in Chemical Engineering.

Universidade Federal do Rio de Janeiro

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Engenharia de Processos Químicos e Bioquímicos Graduate Program

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Abstract

This dissertation studied the solvation free energy differences of molecules mimicking asphaltenes in different solvents with the SAFT- γ Mie force field. We obtained solvation free energy differences by carrying out molecular dynamics simulations at the expanded ensemble. The output from these simulations was then used to estimate the differences with the MBAR method. The results with solvents other than water had low absolute deviations to the experimental data. Meanwhile, the hydration free energy calculations required a binary interaction parameter estimated with output data from molecular dynamics to obtain accurate free energy differences. These results indicated some problems on the SAFT- γ Mie model for water, but, generally, proved that a coarse-grained model could represent the free energy differences of the studied sets of solute-solvent.

Keywords: solvation free energy. asphaltenes. SAFT- γ Mie force field.

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List of symbols

T	Temperature
P	Pressure
V	Volume
t	Time
p	Momentum
r	Coordinates
U, u	Potential Energy
m	mass
v	velocity
P	Pressure
\mathcal{H}	Hamiltonian
q	Generalized Coordinates
K	Kinetic Energy
N	Number of atoms/molecules
h	Planck Constant
S	Entropy
k_b	Boltzmann Constant
β	$1/k_bT$
μ	Chemical Potential
A	Helmholtz Free Energy
G	Gibbs Free Energy
f	Free Energy
F	Forces

ϵ	Depth of the potential well
σ	Distance correspondent to a zero intermolecular potential
λ_r	Repulsive exponent
λ_a	Attractive exponent
x_i	Molar fraction
w_i	Weight fraction
ρ	Density
λ	Coupling Parameter
η	Arbitrary Weight
k_{ij}	Binary Interaction Parameter

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1 Introduction

Solvation free energy calculations with molecular dynamics (MD) have a variety of applications ranging from drug design in the pharmaceutical industry to the development of separation technologies in the chemical industry. Solvation free energy is, more specifically, the difference in free energy related to the process of transferring the solute from the ideal gas phase condition to the liquid solvent phase condition (SHIRTS *et al.*, 2003). Through the study of the solvation phenomenon, it is possible to obtain information about the behavior of the solvent in different chemical environments and the influence of the solute's molecular geometry. It is also possible to calculate other important properties with the solvation free energy, namely the activity coefficient at infinite dilution, Henry constant and partition coefficients. Additionally, solvation free energy calculations can be part of the process of calculating solubility with molecular dynamics.

The solvation phenomenon described above is intrinsically complex. There are many competing forces interfering in the behavior of the solute-solvent interaction, and free energy simulations are susceptible to sampling problems in low energy regions. Various simulation methodologies were developed to enable estimations of energy differences such as the expanded ensemble (LYUBARTSEV *et al.*, 1992), thermodynamic integration (KIRKWOOD, 1935), free energy perturbation (ZWANZIG, 1954; BENNETT, 1976; SHIRTS; CHODERA, 2008) and umbrella sampling (TORRIE; VALLEAU, 1977).

Another influencing factor in the output of these calculations is the choice of force field used to model the solvent and solute molecules. Force field is the name given to the group of parameters and equations used to represent the potential energy function of a system in molecular simulations. They have different levels of description, such as quantum mechanics, atomistic and coarse-grained. The quantum mechanics approach describes the motion of electrons and requires the solution of the Schrodinger equation during the simulation. While, in the atomistic description, only the atomic motions

are represented, and this is done by solving the newton's motion equations. Finally, in the coarse-grained description, atoms are grouped into pseudo atoms or beads, and the motion's equations are solved for them. Coarse-grained models generally reproduce free energy differences since the effects of reducing degrees of freedom in the entropy are counterbalanced by the reduction of enthalpic terms (KMIECIK *et al.*, 2016). Additionally, the success of a coarse-grained force field is essential to decrease the computational time of solvation free energy calculations and to reveal deficiencies in the description of small molecules by these models (MOBLEY *et al.*, 2007; SHIRTS *et al.*, 2003). Hence, we, in this study, assess the efficiencies and shortcomings of the SAFT- γ Mie coarse-grained force field (AVENDAÑO *et al.*, 2011) with free energy calculations for a variety of pairs solute-solvent. This force field was chosen because it uses, unlike the majority of the force fields, the Mie potential (MIE, 1903) and because its method of obtaining parameters is more straightforward than other models. It was initially parameterized with pure component equilibrium and interfacial tension data (AVENDAÑO *et al.*, 2011), and this strategy has provided satisfactory results. Examples include the prediction of phase equilibrium of aromatic compounds, alkanes, light gases, and water (HERDES; TOTTON; MÜLLER, 2015; MÜLLER; MEJÍA, 2017; LOBANOVA *et al.*, 2015), thermodynamic properties of carbon dioxide and methane (AIMOLI; MAGINN; ABREU, 2014a), multiphase equilibrium of mixtures of water, carbon dioxide, and n-alkanes (LOBANOVA *et al.*, 2016), and water/oil interfacial tension (HERDES *et al.*, 2017).

The solvents and solutes in our free energy calculations were picked to test the force field with standard sets used as a benchmark in solvation free energy calculations and with aromatic substances used as a model to asphaltenes. Asphaltenes are complicated to characterize by determining their composition on a molecular basis, but the literature broadly accepts that they can be described as a fraction of crude oil soluble in toluene and insoluble in n-alkenes (pentane, hexane, heptane) (SJÖBLOM *et al.*, 2003). They have motivated many studies with interest in developing models for their structure and behavior due to all the problems they can cause during their transportation

and refining such as precipitation during the oil processing (SJÖBLOM; SIMON; XU, 2015). This precipitation issue is a recurrent problem due to the growing market of the production of crude oil in deep waters, which have conditions favorable to precipitation, such as pressure depletion and acid stimulation (BUENROSTRO-GONZALEZ *et al.*, 2004). As an example, asphaltene precipitation due to pressure drop can clog oil production equipment and cause an almost exponential increase in the cost of production (BUENROSTRO-GONZALEZ *et al.*, 2004; JOSHI *et al.*, 2001). All these factors make the understanding of the behavior of asphaltenes in different chemical and physical environments relevant to the oil industry.

As said in the previous paragraph, asphaltene characterization still faces some issues. Hence, we choose to use polycyclic aromatic hydrocarbons (PAH'S), which have well-defined characteristics, to initially test the efficiency of the SAFT- γ Mie force field in describing the solvation phenomenon. PAH's are a group of organic compounds that have fused rings, carbon and hydrogen in their structure (RAVINDRA; SOKHI; GRIEKEN, 2008). The ones utilized in this work were phenanthrene, anthracene, and pyrene since they have similarities with asphaltenes regarding their solubility. Meanwhile, we selected compounds that are used to characterize asphaltenes (toluene, hexane) as solvents in our free energy calculations. We also tested the anti-solvent/solvent effect of carbon dioxide due to its influence in asphaltene precipitation during the oil processing (SOROUSH *et al.*, 2014). With this study of solvation free energies with the SAFT- γ Mie model, we intend to improve this force field and provide accurate free energy calculations of PAH's with a coarse-grained model. The correct description of these smaller asphaltene-like compounds by this force field opens up the possibility of obtaining satisfactory results for more complex asphaltene models with a less computational expensive force field.

2 Literature Review

2.1 Molecular Simulations of Molecules Mimicking Asphaltenes

Asphaltenes, unlike the majority of the compounds, are not defined on a molecular basis. Their most accepted definition is that they are a fraction of crude oil insoluble in n-alkanes (pentane, hexane, and heptane) and soluble in toluene (SJÖBLOM *et al.*, 2003). Due to uncertainties related to its structures, much work has been done to develop model compounds that have a well-defined structure and can represent asphaltenes. The two categories of models presented in the literature are the archipelago and continental models. In the archipelago model, asphaltenes consist of polyaromatic parts linked together by aliphatic or naphthenic moieties and, in the continental model, they consist of a single polyaromatic ring with linked aliphatic or naphthenic chains (MULLINS, 2010; MURGICH, 2003). Studies with molecular simulation use these models or compounds, such as toluene and pyrene, that have similar solubility properties. The choice of the model's structure, such as chemical bonding, is highly essential since some structures can cause the occurrence of high energy regions during the simulation (LI; GREENFIELD, 2011) .

Ervik *et al.* (2016) obtained the correct interfacial orientation of asphaltenes using coarse-grained molecular dynamics simulations of the interface, with a representative model for the asphaltene molecules. Also using a coarse-grained force field, Jover *et al.* (2015) carried molecular simulations with a continental asphaltene model. The results reproduced experimental data of the strong aggregation of asphaltene molecules in n-heptane and high solubility in toluene. Gao *et al.* (2014) performed a molecular dynamics study with the force field GROMOS 45a3 force field (SCHULER; DAURA; GUNSTEREN, 2001) to identify the structural features of different asphaltene molecules. Mikami *et al.* (2013) employed the archipelago model in their work to investigate the interfacial behavior of asphaltene molecules at the oil-water interface using molecular

dynamics simulations with the OPLS-AA force field (JORGENSEN; TIRADO-RIVES, 1988). They found that asphaltenes are preferably distributed in the oil phase in the case of pure toluene and at the oil-water interface in the case of pure heptane and the oscillatory behavior of asphaltene molecules at the oil-water interface was also discovered.

Teklebrhan *et al.* (2014) used a perylene based model to study molecular association and interaction as well as the adsorption properties of the perylene molecule at the water/toluene or water/heptane interface. Molecular dynamics simulations were also utilized in the study of the nanoaggregation of four types of model asphaltene molecules in binary mixtures of toluene and water (KUZNICKI; MASLIYAH; BHATTACHARJEE, 2009). The authors observed that in thin films of toluene trapped between two aqueous phases, both interface-bound and core-bound asphaltenes have similar diffusion behavior. Headen *et al.* (2017) reported molecular dynamics simulations of four model asphaltenes. They alleged that there is no formation of nanoaggregates, and the distribution of asphaltene clusters is continuous for mixtures of asphaltene in heptane. Generally, the work on molecular simulation with asphaltenes models showed a difference in packing tendencies according to the model (GREENFIELD, 2011). For a thorough review of studies of different models of asphaltenes, the reader is referred elsewhere (GREENFIELD, 2011).

2.2 Coarse-Grained (CG) Force Fields

Molecular simulations can be carried out at different levels of description. The detailed atomistic level or *ab initio* level is described by the laws of quantum mechanics. The system consists of a set of subatomic particles in which Schrodinger's equation is solved for all of them. The next level is the atomistic description. It considers that the system is made up of atoms following the laws of classical mechanics. Force fields at this level are based on van der Waals interactions, which may include neutral or Coulombic charged sites. Contributions due to intramolecular interactions such as bond-stretching, angle-bending, and torsion are also usually accounted in these kinds of force fields.

When the simulation scale needs to be increased, and the atomistic simulations become too computationally expensive such as in the study of biological systems, the coarse-grained (CG) description is more suited. It considers that the system is made up of pseudo-atoms or beads that contain multiple atoms or even an entire molecule.

There is an evident loss of information in grouping atoms; hence it is necessary to assure that the process of eliminating unnecessary or unimportant information ('coarse-graining') does not affect the system's physical behavior. Ideally, coarse-grained models need to have representability, robustness, transferability, and computational efficiency. Representability means that you can use a model at a state point other than the one in which it was parameterized. The other characteristic, robustness, is related to the model's ability to enable reliable predictions for various structural, thermodynamic, or transport properties. Finally, a transferable model is one in which the representation of atomic or chemical moieties have the same behavior in different molecules- e. g., a pseudo-atom representing CH_2 should have the same properties both in an alkene molecule and in a polyethylene molecule (MÜLLER; JACKSON, 2014). To achieve the cited goals, coarse-grained force fields are usually developed by mapping the atomistic model to define the pseudo-atoms, which are generally formed by similar functional groups. The level of coarse-graining also needs to be established. Hadley and McCabe (2012) affirm that six heavy atoms (non-hydrogen atoms) per bead are traditionally used in order to not lose valuable details and to maintain isotropic representations of the beads. After the mapping, CG force fields need to be parametrized. There are two different approaches, bottom up and top down, to link the simulations on the coarse-grained scale to another scale (schematically represented in Figure 2.2.1). The bottom-up approach uses information of a more detailed scale such as the quantum mechanics description or the atomistic description to obtain information necessary to the parametrization. This method highly depends on the more detailed model quality to succeed. Meanwhile, the top down methodology obtains parameters from larger scales, namely experimental thermodynamic properties or structure based properties.

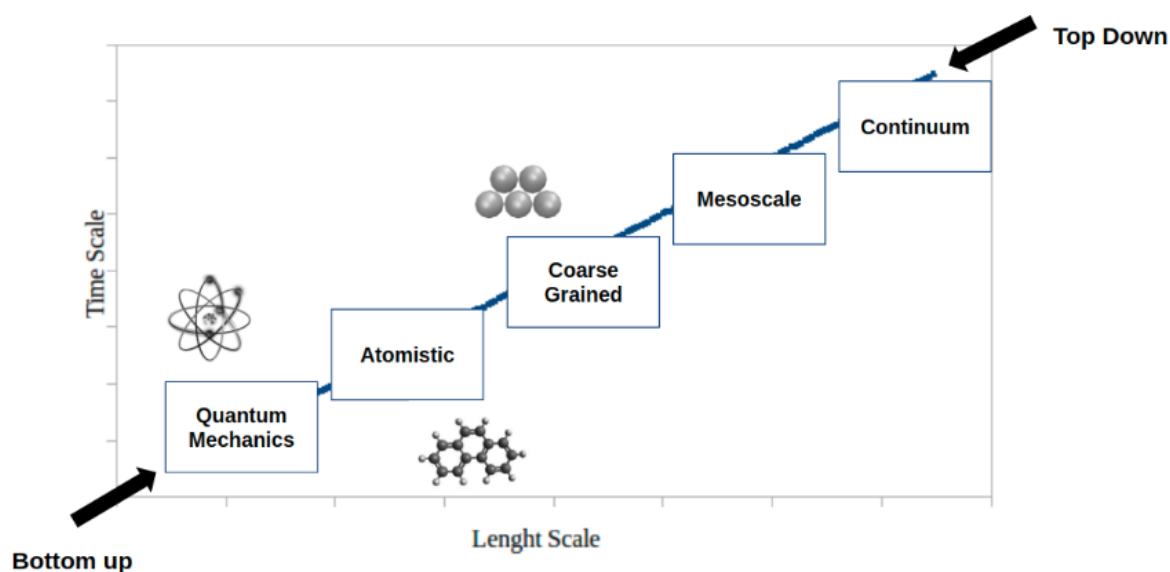


Figure 2.2.1 – Schematic representation for the two different approaches of coarse-graining.

One of the first applications of coarse-grained models is in the study of protein folding ([LEVITT; WARSHE, 1975](#); [LEVITT, 1976](#)). These earlier protein CG models were based on known molecular structure, and they contributed to the knowledge of physicochemical forces associated with protein folding and protein interactions ([KOGE; TAKADA, 2001](#)). More recently, models focused on retaining the protein's chemical specificity. The Bereau and Deresmo model ([BEREAU; DESERNO, 2009](#)) represents a single amino acid with a maximum of four beads, and it was used in studies of protein folding and aggregation. However, this model still needs tuning to improve protein stability ([BEREAU; BACHMANN; DESERNO, 2010](#)). The OPEP (Optimized Potential for Efficient Protein Structure Prediction) model ([STERPONE *et al.*, 2014](#); [STERPONE; DERREUMAUX; MELCHIONNA, 2015](#)) represents a single amino acid with a maximum of six beads. It was used to investigate a variety of phenomena, ranging from protein folding to the modeling of DNA-RNA complexes ([BARDUCCI; BONOMI; DERREUMAUX, 2011](#); [CHEBARO *et al.*, 2009](#); [STERPONE *et al.*, 2014](#)). Other CG protein models used in the literature are the Scorpion (solvated coarse-grained protein interaction) ([BASDEVANT; BORGIS; HA-DUONG, 2013](#)), the UNRES (United Residue) ([ADAM *et al.*, 2014](#)) and the MARTINI model ([LARS *et al.*, 2013](#)). The latter

one is the most popular model for CG modeling of membrane proteins ([MARRINK; TIELEMAN, 2013](#)). The MARTINI force field is also extensively used as a CG model for water. This force field represents four water molecules as one bead using a shifted Lennard-Jones potential for non-bonded interactions. Despite its extensive use, the MARTINI water model does not properly reproduce properties such as interfacial tension and compressibility ([HE *et al.*, 2010](#)). Besides, it can freeze at room temperature ([WINGER *et al.*, 2009](#); [MARRINK *et al.*, 2007](#)), which requires the use of anti-freeze agents during the simulations. This behavior can be explained by the high level of coarse-graining (4:1), the lack of explicit charges, and the use of a 12-6 potential. [Chiu, Scott and Jakobsson \(2010\)](#) used the Morse Potential, which is softer than the LJ potential, to improve the MARTINI model. Meanwhile, [Shinoda, Devane and Klein \(2007\)](#) used different forms of the Mie potential to build a versatile and transferable coarse-grained model for surfactants/water systems using density, interfacial, and hydration free energies data. They selected a 12-4 Mie potential for water cross interactions and a 9-6 Mie potential for the surfactant (alkanes, oxyethylenes, ethylene glycols, ethers, and alcohols) interactions.

Outside of the Martini framework, [Orsi and Essex \(2011\)](#) proposed the ELBA coarse-grained model for molecular dynamics simulations of lipids membranes. In this model, electrostatics are modeled explicitly by charges, and water molecules are represented by a single Lennard-Jones bead embedded with a point dipole. [Genheden \(2016\)](#) expanded the Elba force field to model 1-hexanol, 1-nonanol, n-hexane and n-nonane by representing three carbons with a single bead. Using different Mie and Morse potentials, [He *et al.* \(2010\)](#) studied different levels of coarse-graining for water ranging from one to four molecules per bead. Other investigations also assessed the use of Soft-core potentials to study aqueous solutions of surfactants ([SHINODA; DEVANE; KLEIN, 2007](#)), ionic liquids ([BHARGAVA; KLEIN, 2009](#)), lipids ([SHINODA; DEVANE; KLEIN, 2010](#)), and membranes ([PANTANO; KLEIN, 2009](#)). Another CG force field for water based on the Mie Potential is the SAFT- γ Mie ([LOBANOVA *et al.*, 2015](#)). In this strategy, there are two different models: CGW1-vle and CGW1-ift. Both of them

represent the water molecule as one bead, and the Mie Potential has a repulsive and attractive exponent equal to eight and six, respectively. The CGW1-vle model was parameterized using saturated-liquid density and vapor pressure data and should be used for simulations of aqueous systems in fluid-phase equilibrium at high temperatures and pressures. This model still suffers from premature freezing with a triple point at 343 K. The other model, CGW1-ift, was parameterized using saturated-liquid density and vapor-liquid interfacial tension. Hence it is best suited for interfacial property calculations. Both models had temperature-dependent size and energy parameters and performed well for these properties over the entire liquid temperature range. The SAFT- γ Mie force field has also been applied to other compounds with satisfactory results. [Müller and Mejía \(2017\)](#) parameterized the force field for aromatic compounds and tested it with simulations of fluid phase equilibrium. [Herdes, Totton and Müller \(2015\)](#) carried out simulations of alkanes and light gases. Binary and ternary mixtures of water, carbon dioxide, and water ([LOBANOVA *et al.*, 2016](#)), there are also thermodynamic and transport properties of carbon dioxide and methane ([AIMOLI; MAGINN; ABREU, 2014a](#); [AIMOLI; MAGINN; ABREU, 2014b](#)) and water/oil interfacial tension ([HERDES *et al.*, 2017](#)) studies with this force field.

2.3 Solvation Free Energies

Solvation free energy calculations with molecular dynamics can be used to evaluate the quality of a coarse-grained force field, such as the ones described in the section above, since these estimations can reveal deficiencies in a force field. Besides this application, solvation free energies are used to obtain information about the behavior of the solvent in different chemical environments and to assess the influence of the solute's molecular geometry on the solvation phenomenon. Due to their range of application and inherent complexity, free energy calculations were the subject of a variety of studies in the last decade interested in improving free energy simulations and post-processing methods ([SHIRTS; CHODERA, 2008](#); [PALIWAL; SHIRTS, 2011](#); [SHIRTS; PANDE, 2005](#); [YTREBERG; SWENDSEN; ZUCKERMAN, 2006](#))s.

Recent papers ([MOBLEY; GUTHRIE, 2014](#); [MATOS *et al.*, 2017](#)) made available a big database of hydration free energy of small molecules using the GAFF force field. [Beckstein, Fourrier and Iorga \(2014\)](#) also calculated the hydration free energies for fifty two compounds with the OPLS-AA force field. They obtained an overall root mean square deviation to the experimental data of 1.75 kcal/mol and concluded that the reproducibility of the Lennard-Jones parameters main limiter of the precision of the results. A comparison of polar and nonpolar contributions to these hydration free energies indicated the significance of each the terms ([IZAIRI; KAMBERAJ, 2017](#)). [GARRIDO *et al.* \(2009, 2011\)](#) calculated the free energy of solvation of large alkanes in 1-octanol and water with three different force fields (TraPPE, GROMOS, OPLS-AA/TraPPE) and the solvation free energy of propane and benzene in non aqueous solvents like n-hexadecane, n-hexane, ethylbenzene, and acetone with the force fields TraPPE-UA and TraPPE-AA. [Roy, Blinov and Kovalenko \(2017\)](#) addressed the choice of the Lennard Jones parameters for predicting solvation free energy in 1-octanol. They calculated the solvation free energy of a set of 205 small organic molecules in 1-octanol and found that the force field parametrization of n-octanol proposed by [Kobryn and Kovalenko \(2008\)](#) provided the best agreement. [Goncalves and Stassen \(2005\)](#) calculated the free energy of solvation using the polarizable continuum model coupled to molecular dynamics computer simulation with the GROMOS force field. These calculations were done with a representative set of solutes and with the solvents tetrachloride, chloroform, and benzene. Using the GAFF and the polarizable AMOEBA force fields, [Mohamed, Bradshaw and Essex \(2016\)](#) evaluated the solvation free energy of small molecules in toluene, chloroform, and acetonitrile, and obtained a mean unsigned error of 1.22 kcal/mol for AMOEBA and 0.66 kcal/mol for GAFF. To define the role of solvent water in the docking structure determination of proteins, [Matubayasi \(2017\)](#) developed a method to compute the solvation free energy of proteins while using OPLS-AA force field for the solutes and TIP3P for water. [Genheden \(2016\)](#) expanded the Elba force field to calculate solvation free energies of more than 150 solutes taken from the Minnesota solvation database in polar (water, hexanol, octanol and nonanol) and apolar (hexane, octane, and nonane) solvents. They obtained mean absolute deviations

of 1 kcal/mol for water and 1.5 kcal/mol for hexane. In this model, three carbons are represented by a single bead and water is represented by a single bead.

Though this variety of data using the intramolecular Lennard-Jones potential, we are not aware of work using the Mie Potential in free energy calculations. We, in the present study, try to provide information about these predictions with the SAFT- γ Mie coarse-grained force field. As said before, the output of these calculations are highly dependent on the force field, and some coarse-grained models obtained satisfactory results for these kinds of simulations. Hence, knowing if other coarse-grained approaches have similar performances to the all-atoms force fields can help increase the scale of solvation free energy calculations.

2.4 Solvation Free Energy Calculation Methods

The data from solvation free energy simulations provide potential energies that need to be post-processed and analyzed in order to calculate the solvation free energies effectively. Since these calculations can have slow convergences, a lot of papers in the last decades focused on developing methods to improve sampling and analysis techniques. Almost all methods rely on o three different approaches: free energy of perturbation (FEP), histogram and thermodynamic integration methods.

2.4.1 Thermodynamic integration

The thermodynamic integration method ([KIRKWOOD, 1935](#)) uses equilibrium averages to evaluate the energy derivative with respect to the coupling parameter (λ):

$$\frac{\partial(1/k_bT)G}{\partial\lambda} = \left\langle \frac{\partial\mathcal{H}}{\partial\lambda} \right\rangle \quad (2.1)$$

In Eq. (2.1), k_b is the Boltzmann constant, G is the Gibbs free energy and \mathcal{H} is the Halmitonian of the system. The derivative is obtained by interpolating the output data between the states from simulations. Some examples of methods for the interpolations

are the trapezoidal rule or natural cubic spline (PALIWAL; SHIRTS, 2011). There are also more complex schemes that are usually system specific as the works of Jorge *et al.* (2010) and Shyu and Ytreberg (2010). MD simulations for each coupling parameter k are carried out and the average over the derivative at each state is compute in order to obtain the final solvation free energy:

$$\Delta G \approx \sum_k \left\langle \frac{\partial \mathcal{H}}{\partial \lambda} \right\rangle \quad (2.2)$$

2.4.2 Histograms

Histograms are used to compute probability distributions. Usually every histogram bin is treated as number of visits to a specific state. The standard practice when using histograms is to use the weighted histogram analysis method (WHAM) developed by Ferrenberg and Swendsen (1989) and generalized by Kumar *et al.* (1992) (CHIPOT; POHORILLE, 2007) to put together different histograms by minimizing the statistical error in the computed density of states and entropy function. This method describes the total probability distribution as a weighted sum of probability distributions without bias obtained from biased simulations. This method was developed to avoid problems related to data loss, high uncertainties and the calculation of the constant of the free energy added to the system by the biased potential (ROUX, 1995). The probability distribution dependent on the potential energy (U) and temperature (T) ($\tilde{\varrho}_r^*(U, T)$) for the WHAM is:

$$\tilde{\varrho}_r^*(U, T) = \frac{\sum_i f_i(U) \exp(-\beta U)}{\sum_i f_{tot,i} \exp(\beta_i \tilde{A}_i - \beta_i U)} \quad (2.3)$$

$$\exp(-\beta_i \tilde{A}_i) = \sum_U \tilde{\varrho}_r^*(U, T) \quad (2.4)$$

$$\tilde{\varrho}_r^*(U, T) = \frac{\tilde{\varrho}_r^*(U, T)}{\sum_U \tilde{\varrho}_r^*(U, T)} \quad (2.5)$$

where $\beta = 1/k_bT$, \tilde{A}_i gives the free energy for run i , $f_i(U)$ is the number of counts of energy U for run i and $f_{tot,i}$ is the total number of counts in run i . The Eq. (2.3) and (2.4) are solved self consistently with the initial value for \tilde{A}_i equal to zero. The final probability distribution is then given by Eq. (2.5), and is used to calculate various moments of the potential energy (CHIPOT; POHORILLE, 2007).

2.4.3 Free Energy of Perturbation (FEP)

The free energy of perturbation method (ZWANZIG, 1954) is the oldest and one of the most general purpose strategy to calculate free energy differences. In this method, the thermodynamics of two different systems(A and B) are related with the intention of evaluating differences in intermolecular potentials. This energy change from state A to state B is calculated by:

$$\Delta G_{AB} = -k_bT \ln \langle e^{-\beta(U_B - U_A)} \rangle_A \quad (2.6)$$

According to the equation above, the free energy difference is calculated by doing an average over the potential energies of state A and B obtained during the simulation of state A. This method requires a great overlap between states(the state B needs to represent a small perturbation in state A) in order to obtain a rapid convergence of the free energy difference. To assure overlap, it is possible to carry out simulations in N intermediate states between A and B, so Eq. (2.6) becomes:

$$\Delta G_{AB} = -k_bT \ln \left(\frac{1}{N} \sum_{i=0}^{N+1} e^{-\beta(U_{i+1} - U_i)} \right) \quad (2.7)$$

The way of calculating ΔG of Eq. (2.7) is called Exponential Averaging (EXP) (ZWANZIG, 1955; PALIWAL; SHIRTS, 2011). The direction of the transformation is also important in this method. If the direction is of decreasing entropy, the step is of insertion (ΔG_{AB}), and the method is called insertion exponential averaging (IEXP). The direction of increasing entropy is called a deletion step (ΔG_{BA}), and the method is labeled as

deletion exponential averaging (DEXP). These directions can yield different values of free energy differences due to undersampling in the tail regions of the ΔG_{AB} distribution (KLIMOVICH; SHIRTS; MOBLEY, 2015; POHORILLE; JARZYNSKI; CHIPOT, 2010). These problems make the EXP methods not suited to calculate free energy differences when the system hasn't a sufficient overlap. For these cases, the Bennet Acceptance Ratio or the Multi-State Bennet Acceptance Ratio is more indicated.

2.4.3.1 Bennet Acceptance Ratio (BAR)

The BAR method (BENNETT, 1976) was developed with the intent of eliminating the bias in the free energy estimation. It uses the uncorrelated samples of the potential energy in both directions ($A \rightarrow B$ and $B \rightarrow A$) to obtain the free energy differences using the information in a statically optimal way. The free energy difference between two intermediate states (i and j) is calculated by the self-consistent solution of the following equations:

$$\Delta G_{ij} = \frac{1}{\beta} \ln \left(\frac{\sum_{k=1}^{N_j} \frac{1}{1 + \exp[-\beta(\Delta U_k^j - C)]}}{\sum_{l=1}^{N_i} \frac{1}{1 + \exp[-\beta(\Delta U_l^i - C)]}} \right) + C - \frac{1}{\beta} \ln \left(\frac{N_j}{N_i} \right) \quad (2.8)$$

$$C = \Delta G_{ij} + \frac{1}{\beta} \ln \left(\frac{N_j}{N_i} \right) \quad (2.9)$$

The total free energy difference between end states is then given by the sum over differences of consecutive intermediate states. This method also provides a function to obtain the minimum variance for free energy differences. The variance equation for any value of C is given by:

$$s_{ij}^2 = \frac{1}{\beta^2 N_i} \left[\frac{\langle f^2(x) \rangle_i}{\langle f(x) \rangle_i^2} - 1 \right] + \frac{1}{\beta^2 N_j} \left[\frac{\langle f^2(x) \rangle_j}{\langle f(x) \rangle_j^2} - 1 \right] \quad (2.10)$$

where $f(x) = 1/(1+x)$ is the Fermi equation and $x = \exp[\beta(\Delta U - C)]$. The variance of the

free energy difference between end states can be calculated by assuming independent errors and summing over the variance of consecutive intermediate states. However, this assumption is not correct and there is no general formula to obtain a statistically unbiased estimate of an entire transformation with the BAR method (PALIWAL; SHIRTS, 2011).

There are two other methods related to the BAR method that do not solve Eqs. (2.8) and (2.9) self consistently. By doing that, free energy differences will not have minimum variance, but significant space and disk memory can be saved since the averages of Eqs. (2.8) - (2.10) are accumulated (PALIWAL; SHIRTS, 2011). The two methods are the Unoptimized Bennett Acceptance Ratio (UBAR) and the Range-Based Bennett Acceptance Ratio (RBAR). The first one avoids the self consistently resolution of the BAR equations by defining $C = \beta^{-1} \ln(N_j/N_i)$. The UBAR method also requires that the intermediate free energy differences are approximately equal to zero to obtain optimal estimations. Meanwhile, the RBAR method selects a range of initial guesses of the constant C to calculate a range of ΔG_{ij} . The value of free energy difference correspondent to the minimum variance is then used as input in Eq. (2.9) to calculate the value of C . Hence, this method requires a good estimation of the initial range of the values of C . Regarding accuracy, the UBAR method can be as accurate as the BAR method, but it may end up being as computationally costly (PALIWAL; SHIRTS, 2011).

2.4.3.2 Multistate Bennet Acceptance Ratio (MBAR)

The MBAR method (SHIRTS; CHODERA, 2008) is a further development of the BAR method. It proposes an estimator that computes free energies and their uncertainties of all K states by minimizing the $K \times K$ matrix of variances simultaneously. The estimator solves the following equation for each G_i self consistently:

$$G_i = \frac{1}{\beta} \ln \sum_{k=1}^K \sum_{n=1}^{N_k} \frac{\exp[-\beta U_i(x_{kn})]}{\sum_{l=1}^K N_l \exp[\beta(G_l - U_l(x_{kn}))]} \quad (2.11)$$

The equation above requires the evaluation of the potential energy of uncor-

related configuration n for all K states ($U_i(x_{kn})$) and for all uncorrelated configuration snapshots (N_k) from state k . Free energy changes between states are given then by $\Delta G_{ij} = G_j - G_i$. The statistical variance of $S_{ij}^2 \Delta G_{ij}$ is given by the matrix covariances:

$$s_{ij}^2 \Delta G_{ij} \equiv \text{cov}(-\ln \hat{Z}_j / \hat{Z}_i, -\ln \hat{Z}_j / \hat{Z}_i) \quad (2.12)$$

where \hat{Z}_j and \hat{Z}_i are the partition functions of states i and j .

3 Fundamentals of the Computational Methods

3.1 Molecular Dynamics

3.1.1 Background and Formalution

Molecular Dynamics (MD) is a kind of molecular simulation that uses molecular configurations (Cartesian coordinates and momentum) to extract structural , thermodynamical and dynamical information of a system. This information is extracted from the time evolution of a system, which are obtained through the Newton motion' equations numerical integration ([TUCKERMAN, 2010](#)):

$$\frac{d\vec{p}_i}{dt} = -\frac{\partial U(\vec{r}_N)}{\partial \vec{r}_i} \quad (3.1)$$

where p_i is the momentum and r^N are the coordinates of all the atoms $(x_1, y_1, z_1, \dots, x_N, y_N, z_N)$. Alternatively, we can write the equation relative to the velocity (v_i):

$$m_i \frac{d\vec{v}_i}{dt} = -\frac{\partial U(\vec{r}_N)}{\partial \vec{r}_i} \quad (3.2)$$

In order to develop equations for any coordinate system ($q^N = (r_1, \theta_1, \phi_1)$), the Halmitonian formulation, a more generalized formulation of classical mechanics, is used to develop the equations of motion:

$$\mathcal{H}(q^N, p^N) = K(p^N) + U(q^N) \quad (3.3)$$

In the equation above, K is the kinetic energy and U is the potential energy. The

equations of motion can now be represented by:

$$\frac{d\vec{p}_i}{dt} = -\frac{\partial \mathcal{H}}{\partial \vec{q}_i} \quad (3.4)$$

$$\frac{d\vec{q}_i}{dt} = \frac{\partial \mathcal{H}}{\partial \vec{p}_i} \quad (3.5)$$

The coordinates and momentum axes for each atom in a 6N dimensional space is defined as the phase space. The trajectory through the phase space is then the time evolution of a system in a molecular dynamics simulation. This evolution of the simulation may be used to calculate the thermodynamic properties if the system is ergodic. That is, a trajectory in this system will explore with the same probability all regions of the phase space of microstates with the same energy (SHELL, 2015).

3.1.2 Statistical Ensembles

A statistical ensemble corresponds to the set of configurations of a system. For an isolated system at equilibrium, the ensemble corresponds to the microcanonical ensemble. This ensemble has as control variables the number of particles (N), volume (V) and total energy (E). Following the ergodic hypothesis, the system at these conditions will spend the same amount of time in each of the microstates (points in phase space) with the fixed E. The number of accessible microstates is defined by the partition function or the density of states, and, is given by the following equation for the microcanonical ensemble:

$$\Omega(N, V, E) = \frac{1}{h^{3N} N!} \int dp^N dr^N \delta[\mathcal{H}(p^N, r^N) - E] \quad (3.6)$$

where h is the Planck constant and the delta function is a Dirac delta function. As mentioned above, the system will spend the same amount of time at each of the microstates,

i. e. each of these microstates have equal probabilities (ϱ) of being visited:

$$\varrho(p^N, r^N) = \frac{[\mathcal{H}(p^N, r^N) - E]}{\Omega(N, V, E)} \quad (3.7)$$

In order to obtain the macroscopic properties from molecular dynamics, we need to have the relation of the microcanonical partition function to the entropy:

$$S = k_b \ln \Omega(N, V, E) \quad (3.8)$$

where k_b is the Boltzmann constant. With this equations, we can derive other relations to macroscopic properties with the fundamental equation:

$$dS = \frac{1}{T}dE + \frac{P}{T}dV + \frac{\mu}{T}dN \quad (3.9)$$

$$dE = TdS + PdV + \mu dN \quad (3.10)$$

As said above, the microcanonical ensemble have N,V,E as its control variable. Other ensembles can also be defined according to the macroscopic properties held constant. In the canonical ensemble, N, V and the temperature (T) are held constant and N, pressure (P) and T are held constant in the isothermal-isobaric ensemble. Other ensembles include the isenthalpic-isobaric (constant number of particles, pressure and enthalpy) and the grand canonical (constant chemical potential, volume and temperature). A variety of physical properties are measure at the conditions of the isenthalpic-isobaric ensemble such as enthalpies, entropies, redox potential, equilibrium constants and free energies, what makes this ensemble one of the most important (TUCKERMAN, 2010). This is also the ensemble in which solvation free energy simulations are carried at this work, hence we are going to briefly talk about it. This ensemble is obtained from a Legendre transformation on the canonical ensemble. The Helmholtz free energy $A(N, V, T)$ becomes the Gibbs free energy $G(N, P, T)$ by transforming the volume into

the external pressure:

$$G(N, P, T) = A(N, V(P), T) + PV(P) \quad (3.11)$$

The Gibbs free energy is related to its partition function $\Delta(N, P, T)$ by:

$$G(N, P, T) = -\kappa_b T \ln \Delta(N, P, T) \quad (3.12)$$

where $\Delta(N, P, T)$ is given by:

$$\Delta(N, P, T) = \frac{1}{V_0} \int_0^\infty dV \int d^N p d^N r \exp \left[-\beta \left(\sum_{i=1}^N \frac{p_i^2}{2m_i} + U(r^N) + PV(r^N) \right) \right] \quad (3.13)$$

In the equation above, $Q(N, V, T)$ is the partition function of the canonical ensemble:

$$Q(N, V, T) = \int d^N p d^N r \exp \left[-\beta \left(\sum_{i=1}^N \frac{p_i^2}{2m_i} + U(r^N) \right) \right] \quad (3.14)$$

From these relations and a differential change in G , we can obtain the chemical potential (μ), volume and entropy (S) relations for isenthalpic-isobaric ensemble:

$$\mu = \left(\frac{\partial G}{\partial N} \right)_{P,T} = -\kappa_b T \left(\frac{\partial \ln \Delta(N, P, T)}{\partial N} \right)_{N,T} \quad (3.15)$$

$$\langle V \rangle = \left(\frac{\partial G}{\partial P} \right)_{N,T} = \kappa_b T \left(\frac{\partial \ln \Delta(N, P, T)}{\partial P} \right)_{N,P} \quad (3.16)$$

$$S = \left(\frac{\partial G}{\partial T} \right)_{N,P} = \kappa_b (\ln Q(N, V, T) + T \left(\frac{\partial \ln Q(N, V, T)}{\partial T} \right)_{V,N}) \quad (3.17)$$

3.1.3 Thermostats and Barostats

The isothermal-isobaric and canonical ensembles have external conditions being applied to it (temperature and pressure). For temperature control, the method employed

mimics the effect of a thermal reservoir through a thermostat. The thermostats need to be capable of capturing the correct energy fluctuations in the system since the kinetic energy will fluctuate when using a heat bath to control temperature in a canonical ensemble of a finite system.

Among the available thermostat are the Berendsen (BERENDSEN *et al.*, 1984), the Andersen (ANDERSEN, 1980) and the Nose (NOSÉ, 1984) thermostats, but here we are going to discuss the most widely used thermostat: the Nosé-Hoover (HOOVER, 1985). This thermostat is based on the formulation of Nosé (NOSÉ, 1984), whom used a Lagrangian that contains additional, artificial coordinates and velocities (FRENKEL; SMIT, 2001). In this method the Halmitonian in a canonical ensemble is constructed as:

$$\mathcal{H} = K(p^N) + U(q^N) + \frac{\xi^2 Q}{2} + 3Nk_b T \ln s \quad (3.18)$$

where ξ is a friction coefficient related to the conjugate momentum of the thermal reservoir to which the system is coupled, s is the position of the thermal reservoir and Q is the effective mass associated with s . The velocity update is then done with the Newton's equation with the addition of the friction term (SHELL, 2015):

$$\frac{dr_i}{dt} = v_i \quad (3.19)$$

$$\frac{dv_i}{dt} = -\frac{1}{m_i} \frac{\partial U(r^N)}{\partial r_i} - \xi v_i \quad (3.20)$$

$$\frac{\xi}{dt} = \frac{\sum m_i v_i - 3Nk_b T}{Q} \quad (3.21)$$

$$\frac{\xi}{dt} = -\frac{1}{m_i} \frac{\partial U(r^N)}{\partial r_i} - \xi v_i \quad (3.22)$$

To increase the robustness of the Nosé-Hoover thermostat, the Nosé-Hoover chains of thermostats method was developed. It proposes the use of multiple thermal

reservoirs that are linked to enhance temperature equilibration (SHELL, 2015).

Meanwhile, the pressure is controlled with a barostat. It maintains the pressure constant during the simulation by adjusting the simulation volume. Among the available barostats methodologies are the Berendsen (BERENDSEN *et al.*, 1984), in which the pressure is coupled to a pressure bath and the volume is periodically rescaled and the Anderson barostat (ANDERSEN, 1980), which serves as basis for other barostating methods such as the ones developed by Hoover (HOOVER, 1985), Martina-Tobias-Klein (MARTYNA; TOBIAS; KLEIN, 1994) and Parrinello-Rahman (PARRINELLO; RAHMAN, 1981). The Andersen's idea was couple the system to a fictitious pressure bath and incorporate the volume into the phase space as an additional degree of freedom (TUCKERMAN, 2010). As in the Nosé-Hoover thermostat, Andersen added terms into the Hamiltonian in order to control the variable of interest:

$$\mathcal{H} = \sum_i \frac{V^{-2/3} \pi_i^2}{2m_i} + U(V^{1/3} \mathbf{s}^N) + \frac{p_V^2}{2\mathcal{W}} + PV \quad (3.23)$$

with \mathbf{s} and π being scale transformations:

$$\mathbf{s}_i = V^{-1/3} \mathbf{r}_i \quad (3.24)$$

$$\pi_i = V^{1/3} p_i \quad (3.25)$$

Here, p_V is the momentum conjugate to the volume and \mathcal{W} is a mass parameter that determines the time scale of the volume motion. With the equations above, we can derive the equations of motion in Cartesian coordinates for a constant pressure ensemble:

$$\frac{dr_i}{dt} = \frac{p_i}{m_i} + \frac{1}{3} \frac{dV}{dt} \frac{r_i}{V} \quad (3.26)$$

$$\frac{dp_i}{dt} = -\frac{\partial U}{\partial r_i} - \frac{1}{3} \frac{dV}{dt} \frac{p_i}{V} \quad (3.27)$$

$$\frac{dV_i}{dt} = \frac{p_V}{\mathcal{W}} \quad (3.28)$$

$$\frac{dp_V}{dt} = \frac{1}{3V} \sum_i \left[\frac{p_i^2}{m_i} - \frac{\partial U}{\partial r_i} \cdot r_i \right] - P \quad (3.29)$$

The equation of motions above are integrated numerically using the methodologies described in the next section.

3.1.4 Integration of the motion's equations

With the formalism defined for the equations of motions and for the statistical ensemble, we can then derive discrete-time numerical approximations for them. The basic idea is to solve the trajectory of atoms as function of time ($r^N(t)$) by updating the positions in discrete time intervals or time steps. One of the possible algorithms is The Verlet ([VERLET, 1967](#)). Its proposed equation for the coordinates update is:

$$r(t + \Delta t) = 2r(t) - r(t - \Delta t) + \frac{F(t)}{m} \Delta t^2 + \mathcal{O}(\Delta t^4) \quad (3.30)$$

In this equation, the accuracy is of order $\mathcal{O}(\Delta t^4)$ and the positions are updated by a time step of Δt by using the positions at the previous two time steps and the forces (F) at Δt . The forces acting on every particle are calculated by doing directional derivatives of the intermolecular potential $U(r)$ in relation to the atoms' positions:

$$\vec{F}(r)_i = -\nabla U(r) \quad (3.31)$$

The intermolecular potential is generally described by a force field. The veloci-

ties of this method can be approximated by:

$$v(t + \Delta t) = \frac{r(t + \Delta t) - r(t - \Delta t)}{2\Delta t} + \mathcal{O}(\Delta t^3) \quad (3.32)$$

The Verlet method doesn't use the velocity directly to update the positions. Hence, the Velocity Verlet algorithm was proposed to serve as a reformulation of the Verlet algorithm that uses the velocity:

$$r(t + \Delta t) = r(t) + v(t)\Delta t + \frac{F(t)}{2m}\Delta t^2 \quad (3.33)$$

The velocity in the equation above is updated following:

$$v(t + \Delta t) = v(t) + \frac{F(t + \Delta t) + F(t)}{2m}\Delta t \quad (3.34)$$

Instead of using a time step of Δt , the velocities can be updated at $\Delta t/2$. This is the strategy proposed by the leap frog algorithm:

$$v(t + \Delta t/2) = v(t - \Delta t/2) + \frac{F(t + \Delta t) + F(t)}{m}\Delta t \quad (3.35)$$

$$r(t + \Delta t) = r(t) + v(t + \Delta t/2)\Delta t \quad (3.36)$$

3.1.5 Initial Configuration

The equations above require a non overlapping initial positions and velocities for every atom in the system in MD simulations. The initial velocities follow a temperature dependent Maxwell-Boltzmann distribution:

$$\varrho(v_{x,i}) = \frac{m_i}{2\pi k_b T}^{\frac{1}{2}} \exp\left(-\frac{m_i v_{x,i}^2}{2\pi k_b T}\right) \quad (3.37)$$

Random velocities are then found with the equation above for each of the $3N$

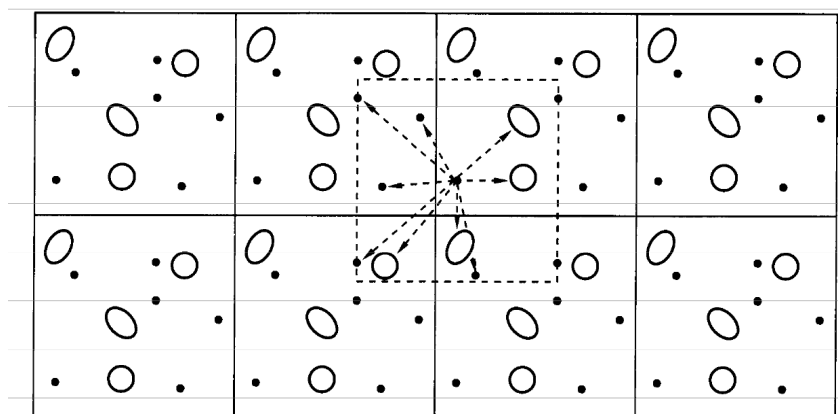


Figure 3.1.1 – Periodic boundary condition representation. Taken from [Frenkel and Smit \(2001\)](#)

components of the velocity. Meanwhile, the initial positions can be obtained by several approaches. The initial configuration can be taken from x ray or NMR spectroscopy, the atoms can be placed randomly in the simulation volume or the atoms can be placed in idealized or approximate geometries. The generally used method to acquire the configurations places the molecules on a cubic lattice. Independent of the method used, certain restrictions should be applied in the initial configurations in order to carry out molecular simulations. As an example, the cubic lattice has a finite size, but a finite size box would result in simulations dominated by surface effects. To avoid that, we can create a box periodically repeated in all directions by applying the so called Periodic Boundary Conditions (PBC). The periodic box has a primitive cell, which contains the system N particles, replicated in periodic lattice of infinite cells as represented in Figure 3.1.1 (FRENKEL; SMIT, 2001).

3.1.6 Force Fields

Force fields are models used to describe structural characteristics such as van der Waals interactions, bond lengths, bond angles and torsion. The description is done by approximating the potential energy function ($U(r^N)$). $U(r^N)$ has contributions due to intermolecular and intramolecular interactions. The intramolecular interactions include bond stretching, bond angle bending and bond torsion. The contribution to the bond stretching is usually given by the following Taylor expansion around the energy

minimum:

$$u_{bs}(d) = k_{bs}(d - d_0)^2 \quad (3.38)$$

Here, d is the bond length, d_0 is the equilibrium bond length and k_{bs} is bond stretching constant. The bond angle bending contribution corresponds to deviations from the preferred geometry, and is given by:

$$u_{ab}(\theta) = k_{ab}(\theta - \theta_0)^2 \quad (3.39)$$

where k_{ab} and θ_0 are constants defined by the force field and θ is the bond angle between three atoms. The bond torsion interactions corresponds to the energies of rotations along bonds and it happens among four atoms:

$$u_{bt}(\omega) = \sum_{n=0}^N c_n \cos(\omega)^n \quad (3.40)$$

where N is the number of terms, c_n is the summation index defined by the force field and ω is the torsional angle defined by the force field.

The intermolecular interactions include electrostatics interactions, van der Waal interactions and volume repulsions. The electrostatics interactions of two atoms i and j with partial charges follow the Coulomb's Law:

$$u_q(r_{ij}) = \frac{q_i q_j}{4\pi\epsilon_0 r_{ij}} \quad (3.41)$$

In the equation above, q_i and q_j are the partial charges and ϵ_0 is the free space permittivity constant. In many force fields, the van der Waal interactions and volume repulsions between particles i and j are modeled in the same equation by the Lennard Jones Potential:

$$u_{LJ}(r_{ij}) = 4\epsilon \left[\left(\frac{\sigma}{r} \right)^{12} - \left(\frac{\sigma}{r} \right)^6 \right] \quad (3.42)$$

where r is the distance the molecules, ϵ is the depth of the potential well, σ is the distance

correspondent to a zero intermolecular potential. The graphical representation of the Lennard Jones potential is:

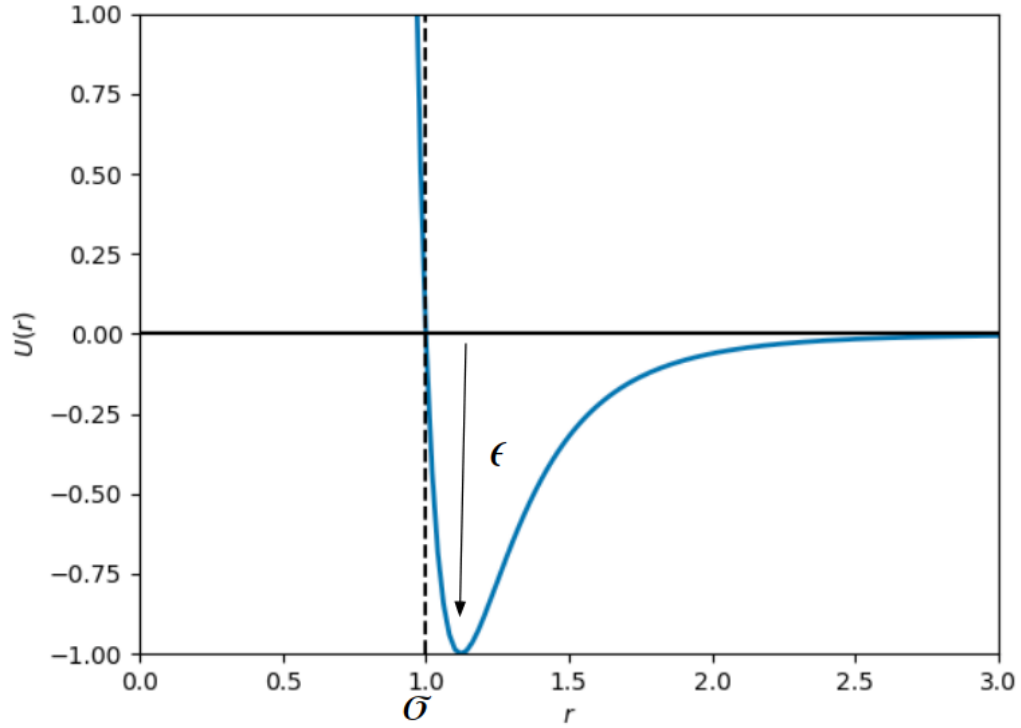


Figure 3.1.2 – Lennard Jones potential representation.

The interactions in Figure 3.1.2 tend to zero after a certain value of r , so this value can be set as the cutoff radius in which the potential energy is considered to be zero after it. Also, only interactions with the nearest periodic image of the cell are considered for short range interactions (Minimum image convention condition). With this conditions, the calculations of forces and velocities are viable.

The final potential energy function defined by the force field can then be expressed by summing all the interactions above:

$$U(r^N) = u_{bs}(d) + u_{ab}(\theta) + u_{bt}(\omega) + u_q(r_{ij}) + u_{LJ}(r_{ij}) \quad (3.43)$$

3.2 SAFT- γ Mie Force Field

3.2.1 SAFT-VR Mie EoS

The SAFT-VR Mie equation of state (LAFITTE *et al.*, 2013) is the basis for the SAFT- γ Mie coarse-grained force field (AVENDAÑO *et al.*, 2011). This EoS was initially developed to describe chain molecule formed from fused Mie segments using the Mie attractive and repulsive potential. The Mie potential is a type of generalized Lennard-Jones potential that can be used to explicitly describe repulsive interactions of different hardness/softness and attractive interactions of different ranges, and is given by:

$$U_{Mie}(r) = \epsilon \frac{\lambda_r}{\lambda_r - \lambda_a} \left(\frac{\lambda_r}{\lambda_a} \right)^{\left(\frac{\lambda_a}{\lambda_r - \lambda_a} \right)} \left[\left(\frac{\sigma}{r} \right)^{\lambda_r} - \left(\frac{\sigma}{r} \right)^{\lambda_a} \right] \quad (3.44)$$

where λ_r is the repulsive exponent and λ_a is the attractive exponent. This equation uses the Barker and Henderson (1976) high perturbation expansion of the Helmholtz free energy up to third order and an improved expression for the radial distribution function (RDF) of Mie monomers at contact to obtain an equation able to give an accurate theoretical description of the vapor-liquid equilibrium and second derivative properties (LAFITTE *et al.*, 2013). For a non-associating fluid, the Helmholtz free energy is:

$$\frac{A}{N\kappa_b T} = a = a^{IDEAL} + a^{MONO} + a^{CHAIN} \quad (3.45)$$

3.2.1.1 Ideal Contribution

The ideal contribution for a mixture is given by:

$$a^{IDEAL} = \sum_{i=1}^{N_c} x_i \ln (\rho_i \Lambda_i^3) - 1 \quad (3.46)$$

where $x_i = N_i/N$ is the molar fraction of component i , $\rho_i = N_i/V$ is the number density, N_i is the number of molecules of each component and Λ_i^3 is the de Broglie wavelength.

3.2.1.2 Monomer Contribution

The monomer contribution describes interactions between Mie segments and can be expressed, for a mixture, as:

$$a^{MONO} = \left(\sum_{i=1}^{N_c} x_i m_{s,i} \right) a^M \quad (3.47)$$

In the equation above, $m_{s,i}$ is the number of spherical segments making up the molecule i and a^M is the monomer dimensionless Helmholtz free energy and it is expressed as a third order perturbation expansion in the inverse temperature (BARKER; HENDERSON, 1976):

$$a^M = a^{HS} + \beta a_1 + \beta a_2^2 + \beta a_3^3 \quad (3.48)$$

where $\beta = \kappa_b T$ and a^{HS} is the hard-sphere dimensionless Helmholtz free energy for a mixture :

$$a^{HS} = \frac{6}{\pi \rho_s} \left[\left(\frac{\zeta_2^3}{\zeta_3^2} - \zeta_0 \right) \ln(1 - \zeta_3) + \frac{3\zeta_1 \zeta_2}{1 - \zeta_3} + \frac{\zeta_2^3}{\zeta_3(1 - \zeta_3)^2} \right] \quad (3.49)$$

The variable $\rho_s = \rho \sum_i^{N_c} x_i m_{s,i}$ is the total number density of spherical segments and ζ_l are the moments of the number density:

$$\zeta_l = \frac{\pi \rho_s}{6} \left(\sum_{i=1}^{N_c} x_{s,i} d_{ii}^l \right), l = 0, 1, 2, 3 \quad (3.50)$$

where $x_{s,i}$ is the mole fraction of segments and is related through the mole fraction of component i (x_i) by:

$$x_{s,i} = \frac{m_{s,i} x_i}{\sum_{k=1}^{N_c} m_{s,k} x_k} \quad (3.51)$$

The effective hard-sphere diameter d_{ii} for the segments is:

$$d_{ii} = \int_0^{\sigma_{ii}} (1 - \exp(-\beta U_{ii}^{Mie}(r))) dr \quad (3.52)$$

The integral in Eq. (3.52) is normally obtained by means of Gauss-Legendre with a 5-point quadrature (PAPAIOANNOU *et al.*, 2014). The detailing of terms of Eq. (3.47)

can be found in [Lafitte *et al.* \(2013\)](#).

3.2.1.3 Chain Contribution

The chain formation of m_s tangentially bonded Mie segments contribution is based on the first-order perturbation theory (TPT1) ([PAPAIOANNOU *et al.*, 2014](#)) and can be expressed as:

$$a^{CHAIN} = - \sum_{i=1}^{N_c} x_i (m_{s,i} - 1) \ln(g_{ii}^{Mie}(\sigma_{ii})) \quad (3.53)$$

The $g_{ij}^{Mie}(\sigma_{ij})$ term correspond to the radial distribution function (RDF) of the hypothetical Mie system evaluated at the effective diameter and can be obtained with the perturbation expansion:

$$g_{ij}^{Mie}(\sigma_{ij}) = g_{d,ij}^{HS}(\sigma_{ij}) \exp[\beta \epsilon g_{1,ij}(\sigma_{ij}) / g_{d,ij}^{HS}(\sigma_{ij}) + (\beta \epsilon)^2 g_{2,ij}(\sigma_{ij}) / g_{d,ij}^{HS}(\sigma_{ij})] \quad (3.54)$$

The other terms in the equations above are explicitly exposed in the original article ([LAFITTE *et al.*, 2013](#)).

3.2.1.4 Ring Contribution

There are two forms for the Helmholtz free energy for rings formed from m_s tangentially bonded segments in the literature. The first one ([LAFITTE *et al.*, 2012](#)) considered that the difference between a chain and a ring molecule is that the latter has one more bond that is connecting the first segment to the last. With this assumption, Eq. (3.53) can be adapted to rings by:

$$a^{RING} = - \sum_{i=1}^{N_c} x_i m_{s,i} \ln(g_{ii}^{Mie}(\sigma_{ii})) \quad (3.55)$$

According to [Lafitte *et al.* \(2012\)](#), Eq. (3.55) needs an additional parameterization with molecular simulation data so the EoS can be used in molecular simulations, but this procedure is not necessary for chain molecules. Recently, [Müller and Mejía \(2017\)](#)

tried to correct this inconsistency by means of developing the ring free energy based on the work of [Müller and Gubbins \(1993\)](#), who obtained rigorous expressions for hard fluids with molecular geometries of rings of $m_s = 3$. The final expression developed for the ring dimensionless Helmholtz free energy is:

$$a^{RING} = - \sum_{i=1}^{N_c} x_i (m_{s,i} - 1 + \chi_i \eta_i) \ln(g_{ii}^{Mie}(\sigma_{ii})) \quad (3.56)$$

$\eta_i = m_{s,i} \rho_i \sigma_{ii}^3 / 6$ is the packing fraction and χ_i is a parameter which depends on $m_{s,i}$ and on the geometry of the ring of each component i . For a value of $\chi = 0$, Eq. (3.56) is equal to Eq. (3.53). Meanwhile, the equation corresponds to a hard sphere system of triangles when $\chi = 1.3827$. [Müller and Mejía \(2017\)](#) also calculated values of ζ for $m_s = 3, m_s = 4, m_s = 5, m_s = 7$ with pseudo-experimental data from molecular dynamics (MD) for a defined pure fluid. Values of χ for some of the geometry estimated can be seen in [Figure 3.2.1](#).

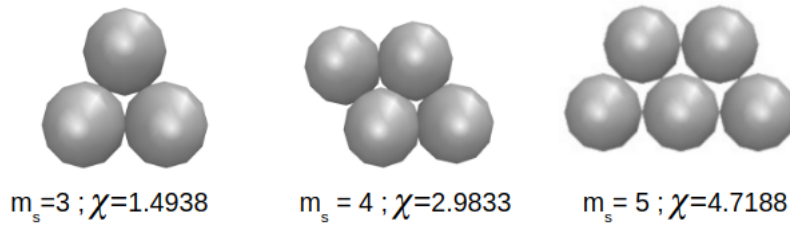


Figure 3.2.1 – Values for parameter χ according to the ring geometry. Adapted from [Müller and Mejía \(2017\)](#).

3.2.1.5 Combining rules for the intermolecular potential parameters

[Lafitte et al. \(2013\)](#) also suggested mixing rules for the potential parameters based on Lorentz-Berthelot combining rules ([ROWLINSON; SWINTON, 1982](#)):

$$\sigma_{ij} = \frac{\sigma_{ii} + \sigma_{jj}}{2} \quad (3.57)$$

$$\lambda_{k,ij} - 3 = \sqrt{(\lambda_{k,ii} - 3)(\lambda_{k,jj} - 3)}, k = r, a \quad (3.58)$$

$$\epsilon_{ij} = (1 - k_{ij}) \frac{\sqrt{\sigma_{ii}^3 \sigma_{jj}^3}}{\sigma_{ij}^3} \sqrt{\epsilon_{ii} \epsilon_{jj}} \quad (3.59)$$

The k_{ij} is a binary interaction parameter to correct the deviations of the Lorentz-Berthelot rule for chemically distinct compounds. This parameter can be fitted to experimental data or pseudo experimental data.

3.2.2 Parameter Estimation for the SAFT- γ Mie Force Field

The SAFT- γ Mie Force Field uses a coarse-graining top down methodology in its parameterization. This methodology aims to obtain the intermolecular parameters from macroscopic experimental data such as fluid-phase equilibrium or interfacial tension data. The idea is that the force field parameters estimated with the SAFT-VR Mie EoS can be used on molecular simulations since both the equation of state and the force field use the same explicit intermolecular potential model (Mie potential). This correspondence between models has already been seen for a variety of fluids in which this force field was parameterized. This success in the representation of the properties of real fluids can be imputed to the degrees of freedom of Mie Potential (HERDES; TOTTON; MÜLLER, 2015). Furthermore, this flexibility provides the exploration of a very large parameter space without using an iterative simulation scheme (AVENDAÑO *et al.*, 2011).

Each substance has initially five parameters to be estimated ($m_s, \sigma, \epsilon, \lambda_r$ and λ_a) according to Eq. (3.44). The number of segments are usually fixed in an integer value since each segment represents one pseudo atom. The attractive parameter is generally fixed due to its high correlation with the repulsive parameter. Usually, the parameter is fixed in the London value of 6, which is a good representation of the dispersion scale of most simple fluids that don't have strong polar interactions (RAMRATTAN *et al.*, 2015; HERDES; TOTTON; MÜLLER, 2015). There are two strategies to obtain the parameters: one is by fitting the Saft-Vr Mie EoS to experimental data as vapor pressure and liquid density and the other one is using correspondent state parametrization. The

first, generally, minimizes the following unweighted least-squares objective function:

$$\min_{\sigma, \epsilon, \lambda_r} F_{obj}(\sigma, \epsilon, \lambda_r) = \sum_{i=1}^{N_p} \left(\frac{P_v^{SAFT}(T_i, \sigma, \epsilon, \lambda_r) - P_v^{exp}(T_i)}{P_v^{exp}(T_i)} \right)^2 + \sum_{i=1}^{N_p} \left(\frac{\rho_l^{SAFT}(T_i, \sigma, \epsilon, \lambda_r) - \rho_l^{exp}(T_i)}{\rho_l^{exp}(T_i)} \right)^2 \quad (3.60)$$

where N_p is the number of experimental points, P_v is the vapor pressure and ρ_l is the saturated liquid density. Another properties that can be used in the estimation are superficial tension and speed of sound. The multiple parameters of the model make it necessary the use of a wide range of experimental data since multiple solutions may be found for the fit. Therefore, one needs to be careful in deciding the level of coarse-graining (i.e. the parameter m_s) and the subsequent parameter space that will not result in some physical inconsistencies such as a premature freezing fluid.

[Lafitte *et al.* \(2012\)](#) suggested that two corrections factors (c_σ and c_ϵ) should be estimated with simulation data when using Eq. (3.55) for the ring contribution. They are related to the EoS parameters by scaled parameters:

$$\sigma^{scaled} = c_\sigma \sigma^{SAFT} \quad (3.61)$$

$$\epsilon^{scaled} = c_\epsilon \epsilon^{SAFT} \quad (3.62)$$

According to [Lafitte *et al.* \(2012\)](#), these corrections are necessary because the approximations employed in the EoS theory generate discrepancies between molecular simulations and the EoS for ring molecules modeled with Eq. (3.55). The objective

function for this estimation is given by:

$$\min_{c_\sigma, c_\epsilon} F_{obj}(c_\sigma, c_\epsilon) = \sum_{i=1}^{N_p} \left(\frac{P_v^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT}) - P_v^{SAFT}(T_i, \sigma^{scaled}, \epsilon^{scaled})}{P_v^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT})} \right)^2 + \sum_{i=1}^{N_p} \left(\frac{\rho_{liq}^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT}) - \rho_{liq}^{SAFT}(T_i, \sigma^{scaled}, \epsilon^{scaled})}{\rho_{liq}^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT})} \right)^2 \quad (3.63)$$

The repulsive parameter is maintained in the value found on the minimization of Eq. (3.60). The refined values for σ and ϵ are:

$$\sigma^{sim} = \sigma^{SAFT} / c_\sigma \quad (3.64)$$

$$\epsilon^{sim} = \epsilon^{SAFT} / c_\epsilon \quad (3.65)$$

It is interesting to point out that this new parameterization is not necessary when using Eq. (3.56) as the ring contribution. The other method to obtain the force field parameters is the correspondent state parametrization (MEJÍA; HERDES; MÜLLER, 2014). This method considers that the unweighted volume average of the attractive contribution to the Mie intermolecular potential, a_1 , is a mean field approximation:

$$a_1 = 2\pi\rho\sigma^3\epsilon\alpha \quad (3.66)$$

The van der Waals constant, α , considering $\lambda_a = 6$ is related by the Mie exponents by:

$$\alpha = \frac{1}{\epsilon\sigma^3} \int_{\sigma}^{\infty} \phi(r)r^2 dr = \frac{\lambda_r}{3(\lambda_r - 3)} \left(\frac{\lambda_r}{6} \right)^{6/(\lambda_r - 6)} \quad (3.67)$$

The parameterization in this method starts by using the experimental acentric factor, ω , for each molecule with a fixed value of m_s to obtain the value of the repulsive

exponent with the following Padé series:

$$\lambda_r = \frac{\sum_{i=0} a_i \omega^i}{1 + \sum_{i=1} b_i \omega^i} \quad (3.68)$$

a_i and b_i are dependent parameters of the number of segments and a table with their values is presented in the original paper (MEJÍA; HERDES; MÜLLER, 2014). The van der Waals constant can be found substituting λ_r into Eq. (3.67). The reduced critical potential T_c^* is related to α by a Padé series:

$$T_c^* = \frac{\sum_{i=0} c_i \alpha^i}{1 + \sum_{i=1} d_i \alpha^i} \quad (3.69)$$

The values of c_i and d_i are also available in the original paper. The reduced temperature of the equation above is used in conjunction with the experimental critical temperature, T_c , to find the energy parameter with the relation below:

$$T_c^* = \frac{\kappa_b T_c}{\epsilon} \quad (3.70)$$

The diameter parameter, however, is not obtained with the critical properties, but with the reduced liquid density, $\rho_{T_r=0.7}$, at the reduced temperature, T_r , of 0.7. This density is also obtained with a Padé series using parameters by Mejía, Herdes and Müller (2014):

$$\rho_{T_r=0.7}^* = \frac{\sum_{i=0} j_i \alpha^i}{1 + \sum_{i=1} k_i \alpha^i} \quad (3.71)$$

The relation among the equation above, σ and the experimental density is given by:

$$\rho_{T_r=0.7}^* = \rho_{T_r=0.7} \sigma^3 N_{av} \quad (3.72)$$

where N_{av} is The Avogadro number. This correspondent state method has the advantage

of only requiring critical data, which is available for a great range of fluids, and liquid density data. The parameters found with this strategy are available at an online database (ERVIK; MEJÍA; MÜLLER, 2016).

The binary interaction parameter k_{ij} of Eq. (3.59) is necessary to adjust the mixture behavior of chemically distinct components. Normally, it is estimated minimizing the difference between experimental binary vapor liquid equilibrium or interfacial tension data and the SAFT-VR Mie EoS output data (MÜLLER; MEJÍA, 2017; LOBANOVA *et al.*, 2016). The objective function is similar to:

$$\min_{k_{ij}} F_{obj}(k_{ij}) = \sum_{k=1}^{N_p} \left(\frac{P_v^{SAFT}(T_k, x, k_{ij}) - P_v^{exp}(T_k, x)}{P_v^{exp}(T_k, x)} \right)^2 + \sum_{k=1}^{N_p} \left(\frac{\rho_l^{SAFT}(T_k, x, k_{ij}) - \rho_l^{exp}(T_i)}{\rho_l^{exp}(T_i)} \right)^2 \quad (3.73)$$

However, Ervik *et al.* (2016) used molecular simulation results to fit the parameter to the superficial tension data. The strategy followed by them was to do simulations in three values of k_{ij} first and, after, they refined the parameter until a value in good agreement with the experimental data was found.

3.3 Solvation Free Energy Simulations Based on Molecular Dynamics

Free energies can be expressed as averages over ensembles of atomic configurations generated using Monte Carlo or molecular dynamics techniques. In the canonical ensemble, the free energy is given by:

$$F(N, V, T) = -\kappa_b T \ln Q(N, V, T) \quad (3.74)$$

Remembering that $Q(N, V, T)$ is the partition function of the canonical ensemble:

$$Q(N, V, T) = \int d^n p d^n r \exp \left[-\beta \left(\sum_{i=1}^N \frac{p_i^2}{2m_i} + U(r_1, \dots, r_n) \right) \right] \quad (3.75)$$

Meanwhile, the average over the isothermal-isobaric ensemble gives the Gibbs free energy:

$$G(N, P, T) = -\kappa_b T \ln \Delta(N, P, T) \quad (3.76)$$

where $\Delta(N, P, T)$ is the partition function of the isothermal-isobaric ensemble:

$$\Delta(N, P, T) = \frac{1}{V_0} \int_0^\infty dV \int d^n p d^n r \exp \left[-\beta \left(\sum_{i=1}^N \frac{p_i^2}{2m_i} + U(r_1, \dots, r_n) + PV(r_1, \dots, r_n) \right) \right] \quad (3.77)$$

Evaluating the partition function is a difficult task, but we are interested in calculating only the Gibbs free energy difference between two states of a system, that is given by:

$$\Delta G_{AB} = G_B - G_A = -\kappa_b T \ln \left(\frac{\Delta_B}{\Delta_A} \right) \quad (3.78)$$

Since the masses of particles in systems 0 and 1 are the same, the moment integrals in the ratio Δ_B/Δ_A can be simplified into to the ratio of the configuration integrals:

$$\frac{Z_B}{Z_A} = \frac{\int_0^\infty dV \int d^n r \exp [-\beta (U(r_1, \dots, r_n) + PV(r_1, \dots, r_n))]_B}{\int_0^\infty dV \int d^n r \exp [-\beta (U(r_1, \dots, r_n) + PV(r_1, \dots, r_n))]_A} \quad (3.79)$$

What results in the following equation for the Gibbs free energy difference that

doesn't require the calculation of the partition function at each state:

$$\Delta G_{AB} = G_B - G_A = -\kappa_b T \ln \left(\frac{Z_B}{Z_A} \right) \quad (3.80)$$

The Gibbs free energy difference between end states A and B are, more specifically, the difference between the solute alone in the gas phase and the solute interacting with the solvent. In order to have accurate results for free energy differences, the states' phase integral must have sufficient overlap ([KLIMOVICH; SHIRTS; MOBLEY, 2015](#)). This can be achieved by calculating the free energy difference between a series of intermediates states. The result of these differences are independent of the path chosen since free energy is a state function. That's why alchemical states (no physical sense) are used to link physical states of interest. The solvation free energy calculations are done through a thermodynamic cycle to gradually insert the solute molecule into the solvent as illustrated in [Figure 3.3.1](#). According to this cycle, the free energy of solvation is expressed as:

$$\Delta G_{solv} = \Delta G_{1 \rightarrow 4} = \Delta G_{1 \rightarrow 2} + \Delta G_{2 \rightarrow 3} + \Delta G_{3 \rightarrow 4} \quad (3.81)$$

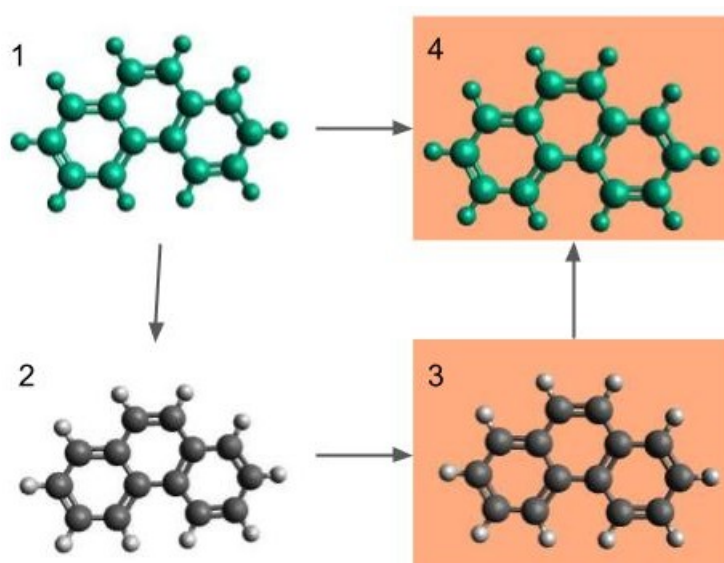


Figure 3.3.1 – Thermodynamic cycle for solvation free energy calculations with molecular dynamics. Adapted from [Klimovich, Shirts and Mobley \(2015\)](#).

The solvation free energy between states 1 and 2 in [Figure 3.3.1](#) is associated with turning off the molecule's non bonded interactions in the gas phase. The following transformation, $\Delta G_{2 \rightarrow 3}$, is the free energy of moving the non-interacting molecule in the gas phase to the solvent, and is equal to zero since the transformation of a non-interacting molecule doesn't depend on the environment. Lastly, $\Delta G_{3 \rightarrow 4}$ is the free energy required to the non-interaction molecule in the aqueous phase regain its non-bonded interactions. The solvation free energy calculation can be classified according to the types of non bonded interactions that are turned off in the $1 \rightarrow 2$ and $3 \rightarrow 4$ parts of the cycle. If both non-bonded interactions with the environment and internal interactions are turned off, this is an annihilation free energy calculation. Meanwhile, if only non-bonded interactions with the environment are turned off, this is a decoupling free energy calculation. In the later case, $\Delta G_{1 \rightarrow 2} = 0$ and the $\Delta G_{solv} = \Delta G_{3 \rightarrow 4}$. The methods used to carry out these transformations scale the solute charges to zero and then turn off the interactions corresponding to the Lennard Jones potential. In order to carry out the later process, a modified potential with a coupling parameter(λ) is used. Each λ represent an alchemical state and, when $\lambda = 0$, there is no interaction with the solvent and, when $\lambda = 1$, interactions are fully activated. The coupling of the λ parameter could be linear, but it could generate numerical problems related to the exponential part of the potential. That's why the non-linear soft-core scheme ([BEUTLER et al., 1994](#)) is used to make the potential behave more smoothly in relation to the change of λ as can be seem in [Figure 3.3.2](#). The generalized soft core potential is given by:

$$U^{sc}(r) = \lambda \epsilon \frac{\lambda_r}{\lambda_r - \lambda_a} \left(\frac{\lambda_r}{\lambda_a} \right)^{\left(\frac{\lambda_a}{\lambda_r - \lambda_a} \right)} \left\{ \frac{1}{[\alpha(1 - \lambda) + (r/\sigma)^{\lambda_a}]^{\lambda_r/\lambda_a}} - \frac{1}{\alpha(1 - \lambda) + (r/\sigma)^{\lambda_a}} \right\} \quad (3.82)$$

where α is a constant which the value of 0.5 is normally assumed to it, r is the distance the molecules, ϵ is the depth of the potential well, σ is the distance correspondent to a zero intermolecular potential, λ_r is the repulsive exponent and λ_a is the attractive

exponent. Using the Lennard Jones exponents ($\lambda_r = 12$ and $\lambda_a = 6$), Eq. (3.82) becomes:

$$U_{LJ}^{sc}(r) = 4\lambda\epsilon \left\{ \frac{1}{[\alpha(1-\lambda) + (r/\sigma)^6]^2} - \frac{1}{\alpha(1-\lambda) + (r/\sigma)^6} \right\} \quad (3.83)$$

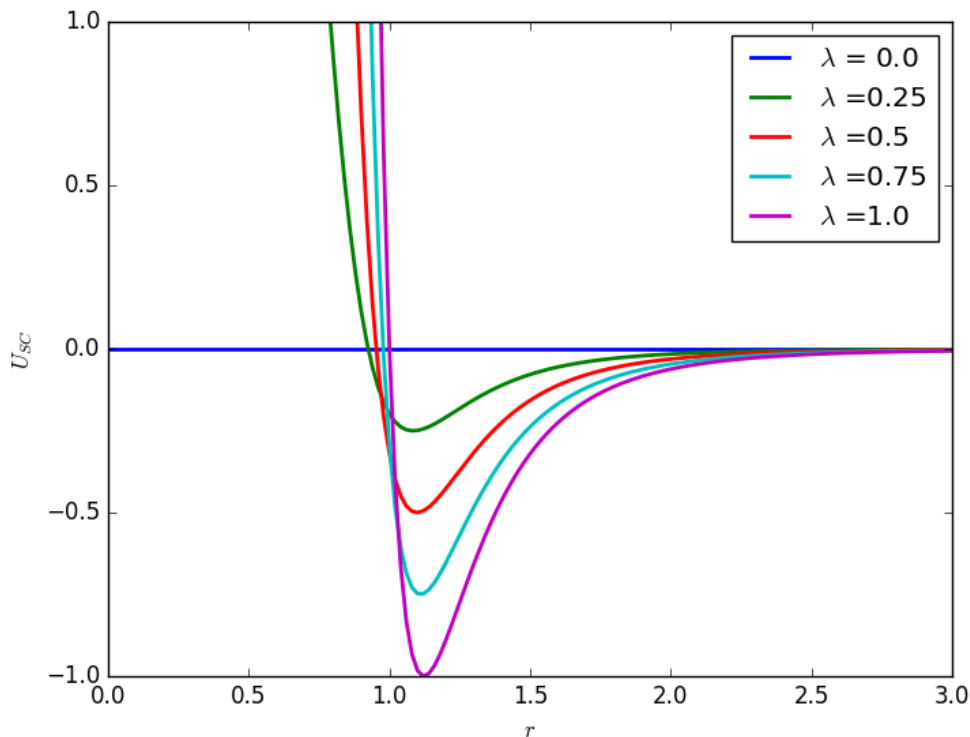


Figure 3.3.2 – Soft-core potential in reduced units.

The $\Delta G_{3 \rightarrow 4}$ can be then obtained by doing independent simulations in different values of λ or by doing expanded ensemble simulations (LYUBARTSEV *et al.*, 1992), which samples all states in a single simulation. This method allows a faster sampling across the alchemical states, provided that the kinetic barriers are not substantial. The free energies of solvation obtained can then be used to calculate other properties such as the partition coefficient, that is a measure of the partitioning of one solute in two solvents (a and b) with different physicochemical characteristics at a temperature T:

$$\log P^{a/b} = \frac{\Delta G_{solv}^a - \Delta G_{solv}^b}{2.303RT} \quad (3.84)$$

3.4 Multistate Bennet Acceptance Ratio (MBAR)

The MBAR method is a free energy of perturbation based method, the MBAR is equal to the BAR method when the calculations are made between only two states. It proposes an estimator that computes free energies and their uncertainties of all K states by minimizing the $K \times K$ matrix of variances simultaneously for a simulation with N_j uncorrelated samples in equilibrium. For each of the $x_{i,n=1}^{N_i}$ configurations of i , the following probability distributions is sampled:

$$p_i(x) = \frac{q_i(x)}{c_i} \quad (3.85)$$

$$c_i = \int dx q_i(x) \quad (3.86)$$

where $q_i(x) = \exp(-u_i(x))$ and u_i is the potential energy of each state, defined by $u_i = \beta[U_i(x) + P_i V(x) + \mu_i^T n(x)]$. Meanwhile, c_i is normalization constant. The free energies are estimated with the ratio of this constant in each state:

$$\Delta f_{ij} = f_i - f_j = -\ln \frac{c_j}{c_i} = -\ln \frac{\int dx q_j(x)}{\int dx q_i(x)} \quad (3.87)$$

[Shirts and Chodera \(2008\)](#) then proposed the following arbitrary function:

$$c_i \langle \alpha_{ij} q_j \rangle_i = c_j \langle \alpha_{ij} q_i \rangle_j \quad (3.88)$$

Using the above equation for every state K , the relation is obtained:

$$\sum_{j=1}^K \frac{\hat{c}_i}{N_i} \sum_{n=1}^{N_i} \alpha_{ij} q_j(x_{i,n}) = \sum_{j=1}^K \frac{\hat{c}_j}{N_j} \sum_{n=1}^{N_j} \alpha_{ij} q_i(x_{j,n}) \quad (3.89)$$

[Shirts and Chodera \(2008\)](#) suggested the following equation for the arbitrary

term α_{ij} in order to decrease the variance:

$$\alpha_{ij}(x) = \frac{N_j \hat{c}_i^{-1}}{\sum_{k=1}^K N_k c_i^{-1} q_k(x)} \quad (3.90)$$

Assuming the sampling is carried out following Boltzmann statistics, Eqs. (3.89) and (3.90) can be rearranged to obtain the free energy estimator, which is solved self consistently:

$$f_i = \frac{1}{\beta} \ln \sum_{k=1}^K \sum_{n=1}^{N_k} \frac{\exp[-\beta u_i(x_{kn})]}{\sum_{l=1}^K N_l \exp[\beta(f_l - u_l(x_{kn}))]} \quad (3.91)$$

The equation above requires the evaluation of the potential energy of uncorrelated configuration n for all K states ($u_i(x_{kn})$) and for all uncorrelated configuration snapshots (N_k) from state k . Free energy changes between states are then:

$$\Delta f_{ij} = f_j - f_i \quad (3.92)$$

The statistical variance resulting from free energy estimation $S_{ij}^2 \Delta f_{ij}$ is given by the matrix covariances:

$$s_{ij}^2 \Delta f_{ij} \equiv \text{cov}(-\ln \hat{Z}_j / \hat{Z}_i, -\ln \hat{Z}_j / \hat{Z}_i) \quad (3.93)$$

where \hat{Z}_j and \hat{Z}_i are the partition functions of states i and j . The MBAR method can be considered as limit case of the Weighted Histogram Analysis Method (WHAM) (KUMAR *et al.*, 1992) for computing free energies. Eq. (3.91) becomes equal to the WHAM equations if the histogram width tends to zero. Despite this, the MBAR is still more suited than the WHAM because it doesn't have the bias associated with the discretization, allows the calculation of an error estimate and allows free energy calculations of states that weren't sampled (SHIRTS; CHODERA, 2008).

3.5 Expanded Ensemble Method

Instead of doing various simulations in different values of λ , expanded ensemble simulations (LYUBARTSEV *et al.*, 1992) were developed to allow a non-Boltzmann sampling scheme of different states in only one simulation. In this scheme, the sampling is done by biasing the phase space exploration process with weights not related to the statistical ensemble. The statistical expanded ensemble, Z^{EE} , is obtained from the probability distributions correspondent to each λ , hence Z^{EE} is defined as a sum of sub ensembles Z_i in different values of λ :

$$Z^{EE} = \sum_{i=1}^N Z_i(\lambda_i) \exp(\eta_i) \quad (3.94)$$

where N is the number of alchemical states, η_i is the arbitrary weight of the sub ensemble at each state and Z_i is configuration partition function at state i . For the isothermal-isobaric ensemble, Z is:

$$Z = \int_0^\infty dV \int d^n r \exp [-\beta (U(\lambda, r_1, \dots, r_n) + PV(r_1, \dots, r_n))] \quad (3.95)$$

In solvation free energy calculations with molecular dynamics, λ corresponds to the coupling parameter of the soft-core potential (Eq. 3.82). For molecular dynamic simulations, the sampling of the expanded ensemble is done by performing an arbitrary number of MD steps followed by a λ transition. Chodera and Shirts (2011) proved that type of sampling of the expanded ensemble is similar to the Gibbs sampling method (GEMAN; GEMAN, 1984; LIU, 2002). Following the Gibbs method, the sampling of the configuration space x for one state λ_k during the MD steps is done by using the conditional distribution:

$$\pi(x|\lambda_k) = \frac{\exp[-\beta u(x, \lambda_k)]}{\int dx \exp[-\beta u(x, \lambda_k)]} \quad (3.96)$$

Meanwhile, the state transition in the MD simulation uses the following condi-

tional distribution:

$$\pi(\lambda_k|x) = \frac{\exp[-\beta u(x, \lambda_k) + \eta_k]}{\sum_{k=1}^K \exp[-\beta u(x, \lambda_k) + \eta_k]} \quad (3.97)$$

where $u(x, \lambda_k) = U(x, \lambda_k) + PV(x, \lambda_k)$ is the reduced potential function for the NPT ensemble. There are a variety of acceptance schemes to do the expanded sampling using Eq. (3.97), but [Chodera and Shirts \(2011\)](#) suggested that the independence sampling ([LIU, 2002](#)) is the best strategy to increase the number of uncorrelated configurations. The implementation suggested by them updates the state index from i to j by first generating a uniform random number R on the interval $[0, 1)$ and then selecting the smallest new value of j that satisfies the relation below:

$$R < \sum_{i=1}^j \pi(\lambda_i|x) \quad (3.98)$$

The sampling strategy above depends on the weights in order to assure an adequate sampling of the states. If there isn't a sufficient number of states sampled, the expanded ensemble becomes deficient in obtaining input data to estimate free energy differences with the methods exposed in Chapter 2. The weights can be calculated following the flat-histogram approach ([BERG; NEUHAUS, 1992](#); [LEE, 1993](#); [DAYAL *et al.*, 2004](#)). This strategy aims to obtain adequate sampling by assuring that all the states have an equal number of samples, i.e. the ratio of the probability of sampling state (π_i) to the probability of sampling state j (π_j) is equal to one. Given that π_i has the following equation:

$$\pi_i = \frac{Z_i(\lambda_i) \exp(\eta_i)}{Z^{EE}} \quad (3.99)$$

and using Eqs. 3.80 and 3.79, the following relation can be obtained for $\pi_i/\pi_j = 1$:

$$(\eta_i - \eta_j) = \beta(G_i - G_j) \quad (3.100)$$

Eq. (3.100) is solved iteratively with trial simulations. For the first simulation, the values of η are chosen or set to zero and the histogram of the states visited is obtained. With this histogram, it is possible to estimate the free energy differences and, since the weights are related to the free energies by Eq. (3.100), the next values of η can be calculated. This iteration goes on until a uniform distribution is secured. The weights found are then used in a longer simulation to obtain the final solvation free energy differences.

The choice of the λ set correspondent to overlapping alchemical states are crucial to acquire accurate energy differences. In this work, the method chosen to obtain the optimal stage of the λ domain is the one developed by Escobedo and Martinez-Veracoechea (2007) with basis in the study of Trebst, Huse and Troyer (2004). This method targets "bottlenecks" in the simulation. It does that by optimizing λ through the minimization of the number of round trips per CPU time between the lowest λ ($= 0$) and highest λ ($= 1$). This is specifically done by maximizing the steady-state stream ϕ of the simulation, which "walks" among the values of λ . This stream is estimated from Fick's diffusion type of law:

$$\phi = D(\Lambda)\Pi(\Lambda)\frac{dx(\Lambda)}{d\Lambda} \quad (3.101)$$

In the equation above, Λ is the actual continue value of the coupling parameter. This continue function of λ s is obtained by interpolating the λ set linearly. $D(\Lambda)$ is the diffusivity at state Λ and $x(\Lambda)$ is the fraction of times that the trial simulation at state Λ_i has most recently visited the state $\lambda = 1$ as opposed to state $\lambda = 0$. The derivative $dx(\Lambda)/d\Lambda$ is approximated with the central finite differences method. Finally, $\Pi(\Lambda)$ is the probability of visiting Λ :

$$\Pi(\Lambda) = \frac{C'\bar{\Pi}(\lambda)}{\Lambda_{i+1} - \Lambda_i} \quad (3.102)$$

The C' term in the equation above represents a constant and $\bar{\Pi}(\lambda)$ is the arithmetic

average of visiting the Λ states:

$$\bar{\Pi}(\lambda) = \frac{\pi_{i+1} - \pi_i}{2} \quad (3.103)$$

The ϕ is maximum when the probability $\Pi'(\Lambda_i)$ is proportional to $1/\sqrt{D(\Lambda)}$. With that information, it is possible to estimate the diffusivity using one trial simulation with the following equation:

$$D(\Lambda) = \frac{\Lambda_{i+1} - \Lambda_i}{\bar{\Pi}(\lambda) dx(\Lambda)/d\Lambda} \quad (3.104)$$

Hence, we can calculate $\bar{\Pi}$ and, consequently, the cumulative probability, which is used to obtain the new λ states:

$$\Phi = \int_{\lambda=0}^{\lambda=1} \Pi'(\Lambda_i) d\Lambda = \frac{i}{K} \quad (3.105)$$

where, K is the total number of λ states.

3.6 Gibbs Ensemble Monte Carlo (GEMC)

The Monte Carlo (MC) approach is another method for generating atomic trajectories in order to obtain macroscopic properties. The trajectories are obtained stochastically, unlike the molecular dynamics approach. The positions are evolved by random moves (MC steps) or perturbations obtained with the Metropolis method ([Metropolis et al., 1953](#)), hence the trajectories are not predictable from the set of initial positions. The Metropolis method is a Markov process, that is a stochastic process in which the configurations change randomly with time and have no memories of the previous configurations. The random move is constructed in such way that the probability of visiting a particular point r^N is proportional to the Boltzmann factor ($\exp[-\beta U(r^N)]$) ([FRENKEL; SMIT, 2001](#)). The construction of a particle displacement MC step according to [Metropolis et al. \(1953\)](#) can be briefly summed up in:

1. Pick a random particle, and calculate its energy $U(r^N)$.
2. Perturb the particle by randomly displacing its configuration, $r' = r + \Delta$. Where Δ is a perturbation chosen from a defined interval of maximum displacement $([-\delta_{max}, \delta_{max}])$. Calculate the energy with the new positions $U(r'^N)$.
3. Accept the move from r^N to r'^N with the probability:

$$acc_{A \rightarrow B} = \min(1, \exp[-\beta U(r'^N) + \beta U(r^N)]) \quad (3.106)$$

The values of maximum displacement are defined in order to obtain acceptance rates of 25-50 % in the step 3 (FRENKEL, 2013). Monte Carlo simulations are interesting when we need to calculate properties in different thermodynamic ensembles, such as the Gibbs Ensemble (PANAGIOTOPOULOS, 1987). This ensemble is used to study phase coexistence with simultaneous Monte Carlo (MC) simulations of two boxes, representing a two phase system, with periodic conditions. The boxes exchange molecules, energy and volume between them. Equilibrium is obtained through MC steps that consist of translation and rotation moves, volume exchange moves and randomly exchanges of molecules between the boxes. For the phase equilibrium of multi-component systems, the GEMC simulations should be carried out at the NPT (constant number of particles, pressure and temperature) ensemble to obey the requirement of an additional degree of freedom for mixtures. Meanwhile, the simulation of single component systems is carried out at constant number of particles, temperature and volume (NVT) since the two phase region would be a line for this system at constant pressure and temperature. The constant volume GEMC ensemble is rigorously equivalent to the canonical ensemble in the thermodynamic limit as demonstrated by Frenkel and Smit (2001). The partition function of the GEMC-NVT ensemble is obtained considering that the particles in both boxes are subjected to the same intermolecular interactions. Also, the boxes' volumes and number of particles (N_1, N_2, V_1 and V_2) can vary while the total volume (V)

and total number of particles (N) remain constant ($N = N_1 + N_2, V = V_1 + V_2$):

$$Q(NVT) \equiv \sum_{N_1}^N \frac{1}{V \Lambda^{3N} N_1! (N - N_1)!} \int_0^V dV_1 V_1^{N_1} V_2^{N_2} \int dx_1^{N_1} \exp[-\beta U(x_1^{N_1})] \int dx_2^{N_2} \exp[-\beta U(x_2^{N_2})] \quad (3.107)$$

In order to define the acceptance rules for the MC moves and compute any property of interest, it is necessary to know the probability of finding the configuration with N_1 particles in box 1 with volume V_1 and positions $x_1^{N_1}$ and $x_2^{N_2}$. This probability is given by:

$$\pi(x_1^{N_1}, x_2^{N_2}, N_1, N_2, V_1, V_2) \propto \frac{V_1^{N_1} V_2^{N_2}}{N_1! N_2!} \exp[-\beta U(x_1^{N_1}) - \beta U(x_2^{N_2})] \quad (3.108)$$

The acceptance criterion for the translation and rotation moves of configuration A to B is similar to the conventional NVT MC ensembles and is equal to:

$$acc_{A \rightarrow B} = \min(1, \exp[-\beta U(x_A^{N_1}) - \beta U(x_B^{N_1})]) \quad (3.109)$$

The exchange volume moves happen by exchanging an amount ΔV between the boxes to achieve pressure equilibrium. ΔV can be chosen from a uniform distribution based on the maximum variation of volume defined (δV_{max}) with probability $1/\delta V_{max}$. The acceptance rule for these moves is:

$$acc_{A \rightarrow B} = \min \left(1, \left(\frac{V_1^B}{V_1^A} \right)^{N_1-1} \left(\frac{V_2^B}{V_2^A} \right)^{N_2+1} \exp[-\beta U(x_A^N) - \beta U(x_B^N)] \right) \quad (3.110)$$

Particle exchange moves are carried out to obtain the equality of chemical potential between the boxes. One particle from one box is removed and then added to a

random location in the other box. The criteria to accept or reject this type of move is:

$$acc_{A \rightarrow B} = \min \left(1, \frac{N_1 V_2}{N_2 V_1} \exp[-\beta U(x_A^N) - \beta U(x_B^N)] \right) \quad (3.111)$$

This method has been widely used to calculate phase equilibrium, but it underperforms for the region near the critical point due to large density fluctuations. The GEMC also has a poor performance for dense systems since the particle exchange moves have a low acceptance ratio.

4 Methodology

4.1 Phenanthrene Parameterization

We implemented the two parameterization strategies for molecules with aromatic rings described in section 3.2.2 for phenanthrene. For both of them, only vapor pressure data (MORTIMER; MURPHY, 1923) were used due to the unavailability of saturated liquid density. We did not estimate the attractive exponent, λ_a . Instead, the value of six was given to it due to its high correlation with the repulsive exponent. The parameterization with the ring equation of Müller and Mejía (2017) was done with the number of segments equals to five and with the geometry in Figure 4.1.1, since this level of coarse-graining was also used for a similar molecule (anthracene) in the original paper.

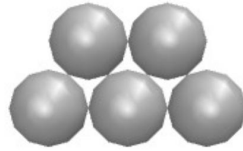


Figure 4.1.1 – Geometry for $m_s = 5$

The minimization was done using the Particle Swarm Optimization (PSO) method (SCHWAAB *et al.*, 2008) with the following objective function:

$$\min_{\sigma, \epsilon, \lambda_r} F_{obj}(\sigma, \epsilon, \lambda_r) = \sum_{i=1}^{N_p} \left(\frac{P_v^{SAFT}(T_i, \sigma, \epsilon, \lambda_r) - P_v^{exp}(T_i)}{P_v^{exp}(T_i)} \right)^2 \quad (4.1)$$

Here, P_v^{exp} is the experimental vapor pressure and P_v^{SAFT} is the vapor pressure obtained with SAFT-VR Mie EoS. We used the routine proposed by Smith, Ness and Abbot (2007) to calculate the bubble point with the EoS. The resulting parameters (σ , ϵ and λ_r) of Eq. (4.1) are the final force field parameters used in molecular simulations.

The parameterization with the Lafitte *et al.* (2012) ring equation was done with

$m_s = 3$ so every bead would represent one aromatic ring:

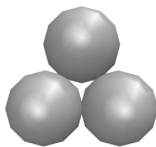


Figure 4.1.2 – Geometry for $m_s = 3$

The first part of the estimation followed the same procedure described above for the Müller and Mejía (2017) equation. As explained in section 3.2.2, the Lafitte *et al.* (2012) equation requires the estimation of correction factors c_σ and c_ϵ (Eqs. (3.61) and (3.62)). We then estimated these parameters using the PSO method with Eq. (3.63). In this equation, the vapor pressures and saturated liquid densities from molecular simulations were obtained using the Gibbs Ensemble Monte Carlo method on the NVT ensemble (section 3.6).

The boxes for the GEMC-NVT simulations were generated inserting 400 molecules of phenanthrene into the liquid box and 100 molecules of phenanthrene into the vapor one using the Playmol package (ABREU, 2017) integrated with the Packmol package (MARTÍNEZ *et al.*, 2009). Initial densities of each box were equal to the saturated densities found with the SAFT-VR Mie Eos to avoid the migration of all molecules to a single phase throughout the simulation. The GEMC-NVT simulations were carried using the Cassandra software (SHAH; MAGINN, 2011), which was developed to perform Monte Carlo simulations. The equilibration and production times lasted around 10^4 and $5 \cdot 10^4$ MC cycles respectively. Each MC cycle corresponded to 10^3 rotation trials, 10^3 translation trials, 10^2 molecule insertion trials, 10^2 molecule deletion trials, and 10 volume exchange trials. The cut-off distance was equal to 20 \AA and we did not use long-range interactions. The saturated vapor density (ρ_{vap}), the saturated liquid density (ρ_{liq}), and the vapor pressure (P_v) were sampled at each 100 MC cycles. Later, these data were divided into five blocks to calculate the averages and standard deviations. With the correction factors found after the estimation with the simulation data, we calculated with Eqs. (3.64) and (3.65) the σ and ϵ parameters, which are the ones to be used in molecular simulations.

4.2 Solvation Free Energy Calculations

Using the parameters for phenanthrene estimated with the Müller and Mejía (2017) approach and the available SAFT- γ Mie force field parameters, we carried out molecular dynamic simulations to estimate solvation free energy differences. The chosen software package to perform the simulations was the Lammmps package (PLIMPTON, 1995). In it, the motion equations were integrated with the velocity-Verlet algorithm (VERLET, 1967) with a time step of 2 fs. As required by the coarse-grained model, molecules were treated as rigid bodies. The thermostat and the barostat were the Nose/Hoover with chains with damping factors of 100 and 1000 time steps respectively. Electrostatics interactions are not explicitly accounted by the SAFT- γ Mie force field, hence there were no shifting of forces or long-range corrections. The potential cutoff was equal to 20 Å (MÜLLER; MEJÍA, 2017) with a neighbor skin of 2 Å. The initial configurations of the solvated systems were also generated using the Playmol package integrated with the Packmol package. For the binary mixtures, one molecule of solute and a varying number of solvent molecules- 700 molecules of toluene, 700 molecules of octanol, 1024 molecules of hexane, 3000 molecules of water - were randomly added to a cubic box. We did the simulations to study solvation free energy of phenanthrene in a mixture of toluene and carbon dioxide with different weight fractions of carbon dioxide. The system consisted of one molecule of phenanthrene for all the fractions and 123 molecules of CO_2 and 618 molecules of toluene for $w_{CO_2} = 0.087$; 166 molecules of CO_2 and 589 molecules of toluene for $w_{CO_2} = 0.119$; 232 molecules of CO_2 and 545 molecules of toluene for $w_{CO_2} = 0.169$; 380 molecules of CO_2 and 446 molecules of toluene for $w_{CO_2} = 0.289$.

All simulations were performed with the constant temperature and pressure values of 298 K and 1 bar, except the ones containing carbon dioxide. These had the temperature of 298 K and the pressure of the liquid phase equilibrium correspondent to the CO_2 fraction (CHANG, 1992). For all the simulations, the initial box was equilibrated at the NPT ensemble for two ns, and the resulting configurations were used as the initial configuration of the expanded ensemble simulations. These were carried out

with the Lammmps user package for expanded ensemble simulations with the Mie Potential developed by our group, available at <https://github.com/atoms-ufjr/USER-ALCHEMICAL>.

During these expanded ensemble simulations, the sampling of a new alchemical state was tried at every 10 MD steps. To define the optimal values of λ and η related to each state, short trials simulations, having around nine ns of production time, were carried out. In the first simulation, we chose the group of λ arbitrarily, and we set the η s to zero, or we assigned to them the values found for similar pairs of solute-solvent. The subsequent group of η were estimated with the flat histogram approach (Eq. (3.100)). We then did another trial simulation with the new weights. The results of this simulation were used to optimize the group of λ s by minimizing the number of round trips, as described in section 3.5. The η s correspondent to the newest group of λ s were interpolated linearly from the free energy differences. With the final values of η and λ defined for each mixture, larger simulations with a production time of 20 ns were carried out.

Since the force field used considers that the beads don't have charges, there are no Coulombic interactions and Eq. (3.81) becomes equal to $\Delta G_{3 \rightarrow 4}$. The post-processing method used to effectively calculate free energy differences with the potential energies obtained from the expanded ensemble simulations was the Multistate Bennet Acceptance Ratio (MBAR), described in section 3.4. The software alchemical-analysis (KLIMOVICH; SHIRTS; MOBLEY, 2015) was utilized to obtain the ΔG_{solv} with MBAR and to assess the quality of the results. After the first estimations, we realized the need of the binary interaction parameter of Eq. (3.59) for pairs with water as a solvent. Hence, we estimated k_{ij} for these pairs and, for all the other pairs, we set k_{ij} to zero. The estimation was done by performing trial expanded ensemble simulations in three values of k_{ij} , as suggested by Ervik *et al.* (2016). With the ΔG_{solv} obtained with these simulations, we did a linear fit to obtain the refined value of the parameter. We used this strategy because the estimation with SAFT VR Mie EoS gave poor results for the solvation free energies.

5 Results and Discussion

5.1 Solvation free energies

The first part of this work consisted of obtaining phenanthrene parameters for the SAFT- γ Mie Force Field as described in section 4.1. This part was necessary since these parameters were not available for the ring geometry on the force field database. The parameters obtained and the mean percentage error (MPE) to the experimental data (MORTIMER; MURPHY, 1923) were:

Table 1 – Estimated SAFT- γ Mie Force Field parameters for phenanthrene

m_s	ϵ/k_B (K)	σ (Å)	λ_r	MPE(%)
3 (LAFITTE <i>et al.</i> , 2012)	485.55	4.197	14.34	1.64 9.74
5 (MÜLLER; MEJÍA, 2017)	262.74	4.077	9.55	0.88

In the table above, the first value of MPE for the Lafitte *et al.* (2012) strategy corresponds to the estimation with experimental data and the second corresponds to the corrections factors estimation with MD data. This strategy has an inconsistency since parameters solely estimated with the EoS do not perform well in molecular simulations. Hence, we only studied the solvation free energy of phenanthrene with the set of parameters estimated with Müller and Mejía (2017) strategy. In fact, we only followed the (LAFITTE *et al.*, 2012) strategy because it was the sole one available when we first started this work. The sets of parameters for the other compounds were retrieved from the literature (LOBANOVA *et al.*, 2016; HERDES; TOTTON; MÜLLER, 2015; ERVIK; MEJÍA; MÜLLER, 2016; MÜLLER; MEJÍA, 2017), and all the utilized parameters are available in Table 2.

Table 2 – SAFT- γ Mie Force Field for each substance used in this work

	m_s	ϵ/k_B (K)	$\sigma(\text{\AA})$	λ_r
Water	1	305.21	2.902	8.0
Propane	1	426.08	4.871	34.29
Carbon dioxide	2	194.94	2.848	14.65
Hexane	2	376.35	4.508	19.57
Octanol	3	495.71	4.341	28.79
Toluene	3	268.24	3.685	11.80
Benzene	3	230.30	3.441	10.45
Pyrene	4	459.04	4.134	15.79
Anthracene	5	259.68	3.631	9.55

The solvation free energies of aromatic solutes in nonpolar (hexane), aromatic (toluene) and hydrogen bonding (1-octanol) solvents were examined with binary interaction parameters equal to zero. Since the force field does not account for charges, we only needed to calculate the Mie contribution (Eq. (3.82)) to the solvation free energy. A total of 15-18 λ_s , depending on the pairs solute-solvent, and their respective η_s were estimated as described in the methodology. The final λ set was found using the cumulative probability distribution (Eq. (3.105)) for all pairs. The distribution for the pair hexane+benzene can be seen in Figure 5.1.1 and the optimized values of λ and η in Table 3. The λ_s and η_s for the other pairs are available at Appendix A. Observing the coupling parameters found for all the pairs, we can see that they are concentrated on the region with a steeper slope as it is expected.

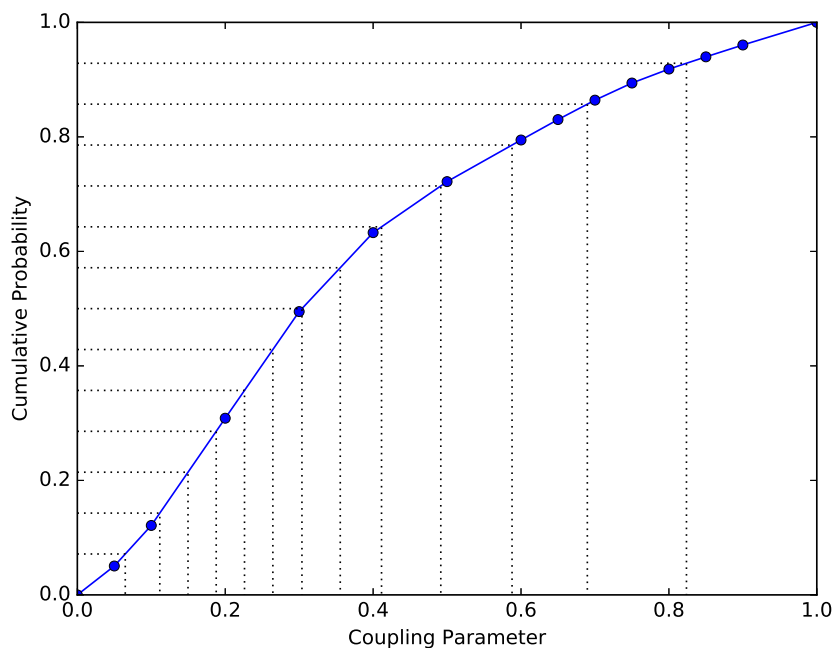


Figure 5.1.1 – Cumulative probability

Table 3 – Optimized values of λ and η for the pair benzene + hexane

λ	η
0	0
0.065	0.708
0.112	1.385
0.15	1.892
0.188	2.399
0.226	2.519
0.264	2.457
0.304	2.367
0.356	1.921
0.411	1.411
0.492	0.524
0.588	-0.663
0.69	-2.016
0.824	-3.922
1	-6.583

It is also important to analyze the reliability of solvation free energy estimations through the overlapping of the intermediate states. Insufficient overlap among states when using FEP based methods such as MBAR may result in the underestimation of variance, and, consequently, in substantially incorrect free energies (KLIMOVICH; SHIRTS; MOBLEY, 2015). The overlapping matrix for the solvation free energy of benzene in hexane is presented in Figure 5.1.2 and the matrices for the other pairs

are available at Appendix B. The elements of these matrices are the probabilities of observing a sample from state i in state j . As an example, the probability of observing state 6 in a simulation of state 8 is 0.16 in Figure 5.1.2. It can be observed in the figures that the matrices for all the pairs in this work have more than three diagonals. In addition to that, the probabilities in the three main diagonals are higher than 0.03, which are one of the characteristics of a reliable free energy estimation (KLIMOVICH; SHIRTS; MOBLEY, 2015). After this analysis, we present in Table 4 the results for solvation free energy calculations and the absolute deviations to experimental data (KATRITZKY *et al.*, 2003).

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.42	.29	.16	.08	.03	.01	.01								
1	.33	.27	.18	.10	.06	.03	.02	.01							
2	.23	.23	.18	.12	.09	.06	.04	.02	.01	.01					
3	.14	.16	.16	.13	.11	.10	.08	.06	.03	.02	.01				
4	.06	.09	.11	.11	.12	.12	.13	.11	.07	.04	.02	.01	.01		
5	.02	.03	.05	.07	.09	.12	.16	.16	.12	.08	.05	.02	.01	.01	
6		.01	.02	.03	.06	.09	.15	.18	.16	.12	.08	.04	.02	.01	
7			.01	.02	.03	.07	.13	.17	.18	.16	.11	.07	.04	.02	.01
8				.01	.02	.04	.09	.14	.17	.18	.15	.10	.06	.03	.01
9					.01	.02	.06	.11	.15	.18	.17	.13	.09	.05	.02
10						.01	.04	.07	.11	.16	.18	.17	.14	.08	.03
11						.01	.02	.04	.08	.12	.17	.19	.19	.14	.06
12							.01	.02	.05	.08	.13	.18	.22	.19	.11
13							.01	.01	.03	.05	.09	.15	.22	.25	.19
14								.01	.01	.03	.06	.11	.19	.29	.30

Figure 5.1.2 – Overlapping matrix for hexane+benzene.

Table 4 – Calculated and experimental values for the solvation free energy differences (kcal/mol) of solutes in non aqueous solvents

Solvent	Solute	ΔG_{solv}^{exp}	ΔG_{solv}^{Mie}	Absolute Deviation
hexane	benzene	-3.96	-3.76 ± 0.01	0.20
hexane	pyrene	-11.53	-10.82 ± 0.02	0.71
hexane	phenanthrene	-10.01	-9.16 ± 0.01	0.85
1-octanol	propane	-1.32	-1.36 ± 0.02	0.04
1-octanol	anthracene	-11.72	-8.16 ± 0.03	3.61
1-octanol	phenanthrene	-10.22	-8.34 ± 0.03	1.47
toluene	pyrene	-12.86	-11.74 ± 0.01	1.11
toluene	anthracene	-11.31	-9.90 ± 0.01	1.41

The numerical values for solvation free energies in hexane had smaller absolute deviations to experimental data, what shows that the SAFT- γ Mie force field performs better for a non-polar solvent. Additionally, this force field presented better results for the pair hexane+benzene than the Trappe force field (-4.35 ± 0.05 kcal/mol) (GARRIDO *et al.*, 2011) and the ELBA coarse-grained force field (-2.92 ± 0.01 kcal/mol) (GENHEDEN, 2016). We also observed the effect of molecule's size on the entropic region of the free energy curve in Figure 5.1.3. It was expected that a force field based on an EoS that does not explicitly account for hydrogen bond would not perform well for 1-octanol. Despite this, the solvation free energies of propane and phenanthrene in 1-octanol stayed in the desired deviation range of 1-2 kcal/mol (MOBLEY; GILSON, 2017). The solvation free energy absolute deviation for propane was much smaller when compared to the other solutes, what can be attributed to propane's non-polarity and smoother free energy curve (Figure 5.1.4). This solvation free energy of propane in 1-octanol also had a smaller deviation than the prediction of the ELBA force field (-0.92 ± 0.01) (GENHEDEN, 2016). The anthracene and phenanthrene molecules have the same geometry in the model and similar physical properties, but the absolute deviation of the solvation free energy of anthracene in 1-octanol is much higher than the one of phenanthrene 1-octanol. This high deviation may indicate a problem in the parameterization of anthracene.

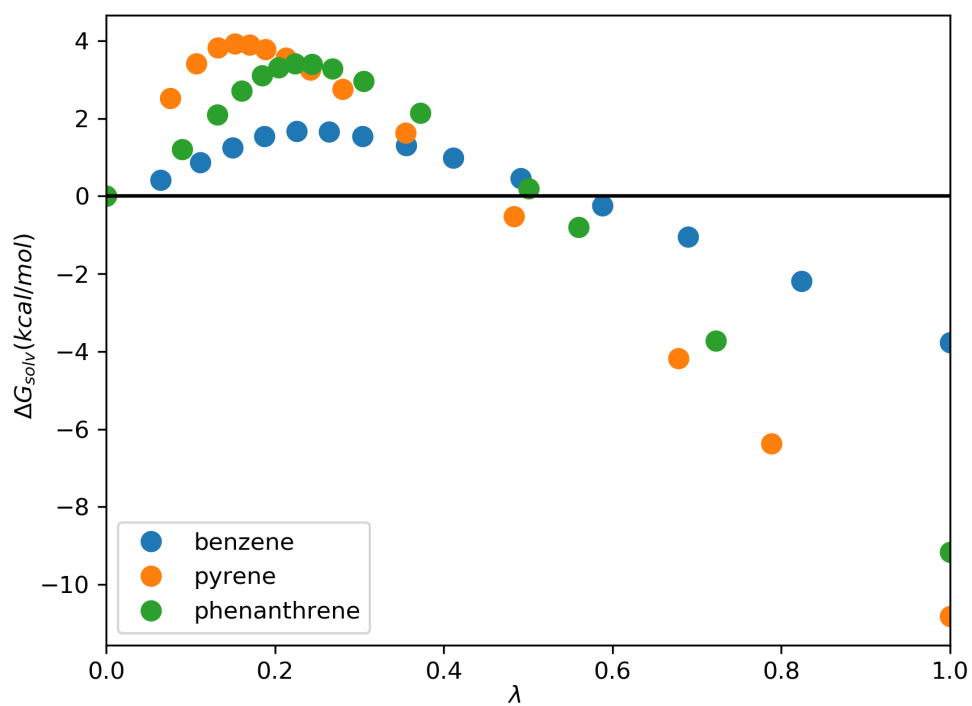


Figure 5.1.3 – Solvation free energy profiles of different solutes in hexane.

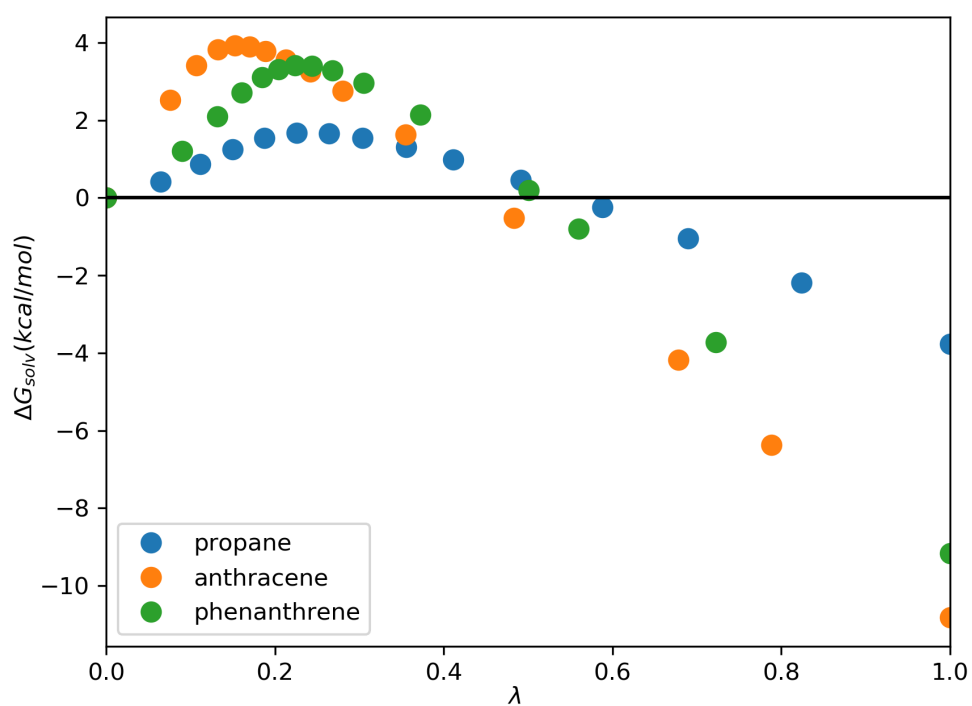


Figure 5.1.4 – Solvation free energy profiles of different solutes in 1-octanol.

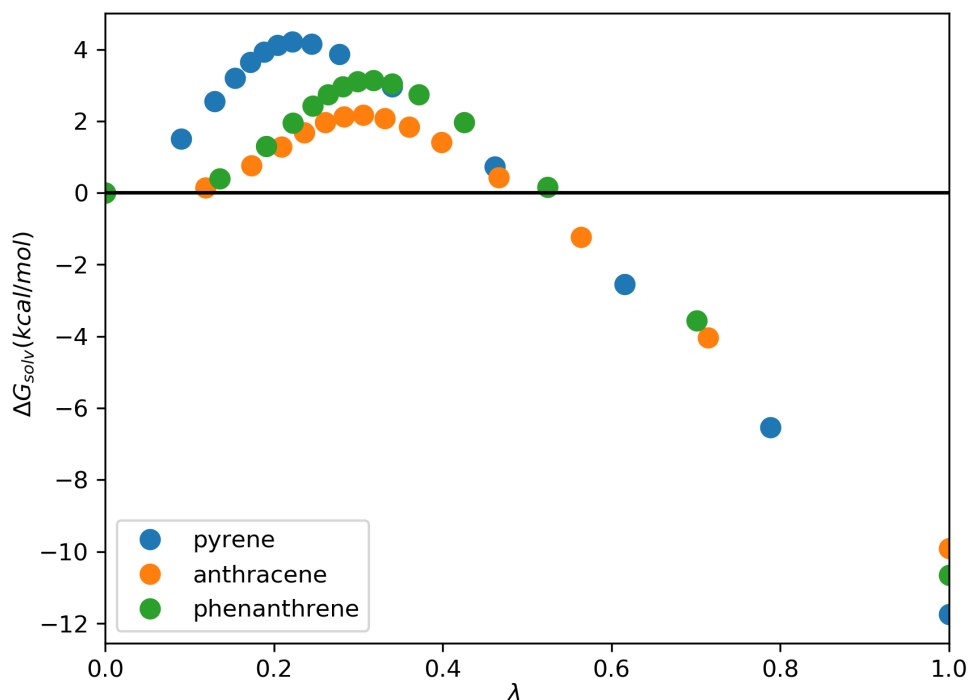


Figure 5.1.5 – Solvation free energy profiles of different solutes in toluene.

The results also indicated the prediction capability of the force field for pairs of aromatic solute and solvent. The influence of the molecule's geometry on the free energy curves was the same as the one observed for other solvents (Figure 5.1.5). ΔG_{solv} was also calculated for phenanthrene in toluene and in toluene+ CO_2 . To the best of our knowledge, there were no available experimental data for these solvation free energies, but the previous results for phenanthrene in other solvents and for the pair anthracene+toluene showed that the force field is adequate to describe the solvation phenomenon of phenanthrene in an aromatic solvent. The results for these sets are exposed below:

Table 5 – Calculated values for the solvation free energy differences (kcal/mol) of phenanthrene in toluene+ CO_2 .

w_{CO_2}	ΔG_{solv}^{Mie}
0.0	-10.65 ± 0.02
0.087	-10.73 ± 0.02
0.119	-10.78 ± 0.02
0.169	-10.71 ± 0.02
0.289	-10.69 ± 0.02

The increasing of CO_2 mass fraction in toluene caused a small effect on solvation free energies. First, the ΔG_{solv} decreased with the increase of w_{CO_2} . After the 0.119 fraction, the effect was reversed and carbon dioxide became an anti-solvent. Soroush *et al.* (2014) reported that asphaltene precipitation occurs when carbon dioxide mass fractions became higher than 0.10 in the system asphaltene+toluene+carbon dioxide, what is in agreement with the anti-solvent effect of carbon dioxide observed on the calculated values. It is also important to point out that the small differences observed in the free energy profiles (Figure 5.1.6) may indicate the insignificance of CO_2 in the solvation of phenanthrene in toluene when using the Saft- γ Mie force field. But, more studies need to be done to make a secure assertion about it since this is a qualitative study due to the lack of experimental data.

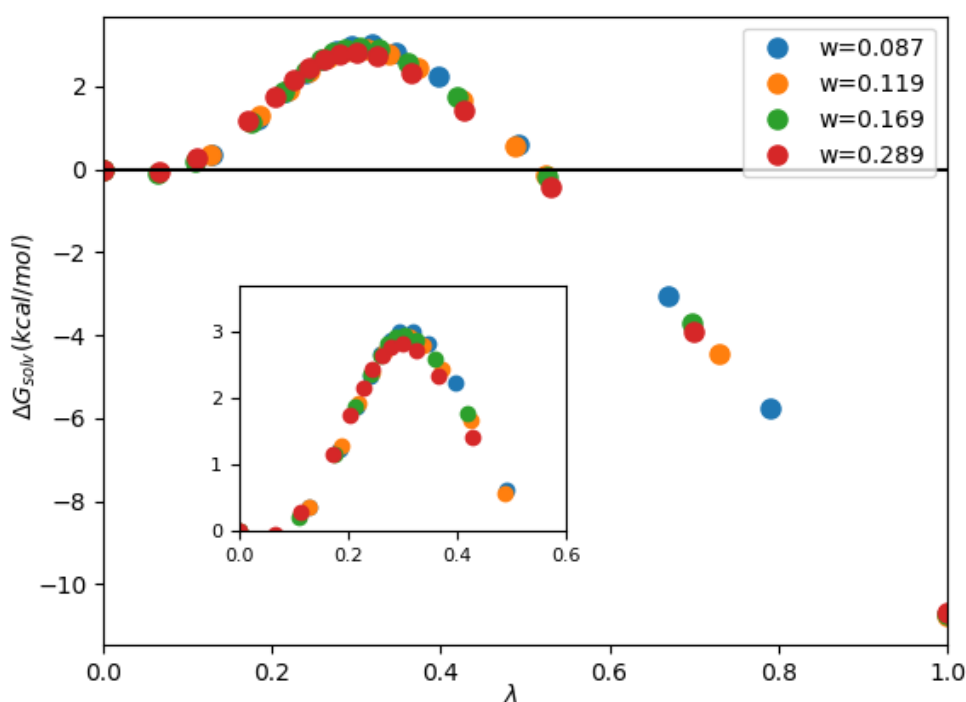


Figure 5.1.6 – Solvation free energy profiles of phenanthrene in toluene+ CO_2 .

5.2 Hydration free energies

We also calculated the hydration free energies of widely studied solutes (propane, benzene) and aromatic solutes (toluene, phenanthrene) with a group of fifteen intermediate states. First, the binary interaction parameter was set to zero, but the preliminary results for hydration free energies, exposed in Table 6, had a high deviation from the experimental data (ABRAHAM *et al.*, 1990; RIZZO *et al.*, 2006).

Table 6 – Calculated values using $k_{ij} = 0$ and experimental values for the hydration free energy differences (kcal/mol) of solutes in water.

Solute	ΔG_{solv}^{exp}	ΔG_{solv}^{Mie}	Absolute Deviation
propane	2.00 ± 0.20	1.10 ± 0.01	0.90
benzene	-0.86 ± 0.20	-4.45 ± 0.03	3.59
toluene	-0.83 ± 0.20	-10.98 ± 0.30	10.15
phenanthrene	-3.88 ± 0.60	-10.90 ± 0.04	7.02

After these results, the need for binary interaction parameters was clear. First, we estimated k_{ij} with the SAFT VR Mie EoS and experimental vapor pressure data, but this strategy also did not provide good results. Hence, we used the approach of estimating the k_{ij} with the output from solvation free energy calculations with molecular dynamics, as described in the last paragraph of section 4.2. We initially found individual values for the interaction parameter of each pair, but, since the parameters for aromatic solutes were very similar (0.148, 0.162, 0.152), we averaged these values. By doing that, we obtained a general parameter for the water+aromatic pairs:

Table 7 – Binary interaction parameters employed.

Pair	k_{ij}
water + propane	0.067
water + aromatic	0.154

The relatively large k_{ij} value of the aromatic solutes can be pinned on the lack of an explicit association term in the model and on the water model itself since the force field did not need a k_{ij} for mixtures with the other hydrogen bonding solvent (1-octanol). This SAFT- γ Mie model for water (LOBANOVA *et al.*, 2016) has two different temperature-dependent sets of parameters. The parameters utilized in this work was the one estimated with experimental interfacial tension data. Hence, we tested the binary

interaction parameter for water+toluene estimated with MD interfacial data by [Herdes et al. \(2017\)](#). Nevertheless, the result was not satisfactory and this parameter could not be transferable to the solvation free energy of toluene in water.

These issues faced by SAFT- γ Mie model are related to the problems of modeling water with a coarse-grained force field. One of the main difficulties is the choice of which water molecules are going to be represented by which specific beads since water molecules move independently and are only bound by non bonded interactions ([HADLEY; MCCABE, 2010](#); [HADLEY; MCCABE, 2012](#)). The SAFT- γ Mie water considers that one water molecule corresponds to one bead. This strategy only saves small simulation time, but it can predict properties at physiological temperatures unlike other more aggressive models, which consider that one bead represents various water molecules. In light of all this, the SAFT- γ Mie force field appears to be a good alternative when working close to room temperatures, but the necessity of additional parameters estimated with molecular simulation indicates problems on the model. Using these parameters, we then obtained the final hydration free energy differences presented in Table 8.

Table 8 – Calculated and experimental hydration free energy differences (kcal/mol) of solutes in water.

Solute	ΔG_{solv}^{GAFF}	ΔG_{solv}^{ELBA}	ΔG_{solv}^{exp}	ΔG_{solv}^{Mie}	Absolute Deviation
propane	2.50 ± 0.02	2.76 ± 0.02	2.00 ± 0.20	2.01 ± 0.01	0.01
benzene	-0.81 ± 0.02	-0.69 ± 0.01	-0.86 ± 0.20	-1.12 ± 0.01	0.26
toluene	-0.79 ± 0.03	-0.76 ± 0.01	-0.83 ± 0.20	-0.84 ± 0.01	0.01
phenanthrene	-5.26 ± 0.03	-	-3.88 ± 0.60	-3.47 ± 0.02	0.41

Hydration free energy differences calculated using the SAFT- γ Mie force field with $k_{ij} \neq 0$ had low absolute deviations to the experimental data, as expected since the parameters were adjusted to fit the experimental data. In the table above, we also show the results obtained by [Mobley and Guthrie \(2014\)](#) with the GAFF force field and by [Genheden \(2016\)](#) with the ELBA force field. The GAFF (General Amber Force Field) force field is an all-atom model that consists of bonded and non bonded parameters and is suitable for the study of a significant number of molecules. Meanwhile, the ELBA

force field is a coarse-grained model that comprises six independent parameters. This force field models three carbons as one Lennard-Jones site and water molecules as Lennard Jones sites with a point dipole.

Comparing the three force fields, the root mean square error (RMSE) for the pairs tested with the SAFT- γ Mie model was 0.24, the RMSE for hydration free energy differences with the GAFF force field was 0.73, and the RMSE for the ELBA coarse-grained force field was 0.44. The difference in absolute deviations between the GAFF and SAFT- γ Mie force fields is significantly high for phenanthrene, hence the coarse-grained force field with a binary parameter is preferred if the application requires a higher level of accuracy. The results also indicated that the SAFT- γ Mie Model with the binary interaction parameter performed better than the ELBA force field in modeling the solvation phenomenon of the pairs studied in this work and worst with the binary parameter equal to zero. This fact occurred despite the fact that both models have the same level of coarse-graining (one bead represents one water molecule). Hence, the choice between the two coarse-grained models is dependent on the availability and transferability of binary interaction parameters for the Mie Model. We also present, for the SAFT- γ Mie force field, the hydration free energy profiles in Figure 5.2.1. The geometry dependence on the free energy profiles is apparent as it was for the solvation free energy study in other solvents. We also observe that the hydration free energy for the first non zero λ is negative for benzene and toluene when a positive value is expected since energy is required to 'open space' in the solvent for the solute's insertion. This anomaly can be caused by numerical errors during the estimation or by another inconsistency in the force field.

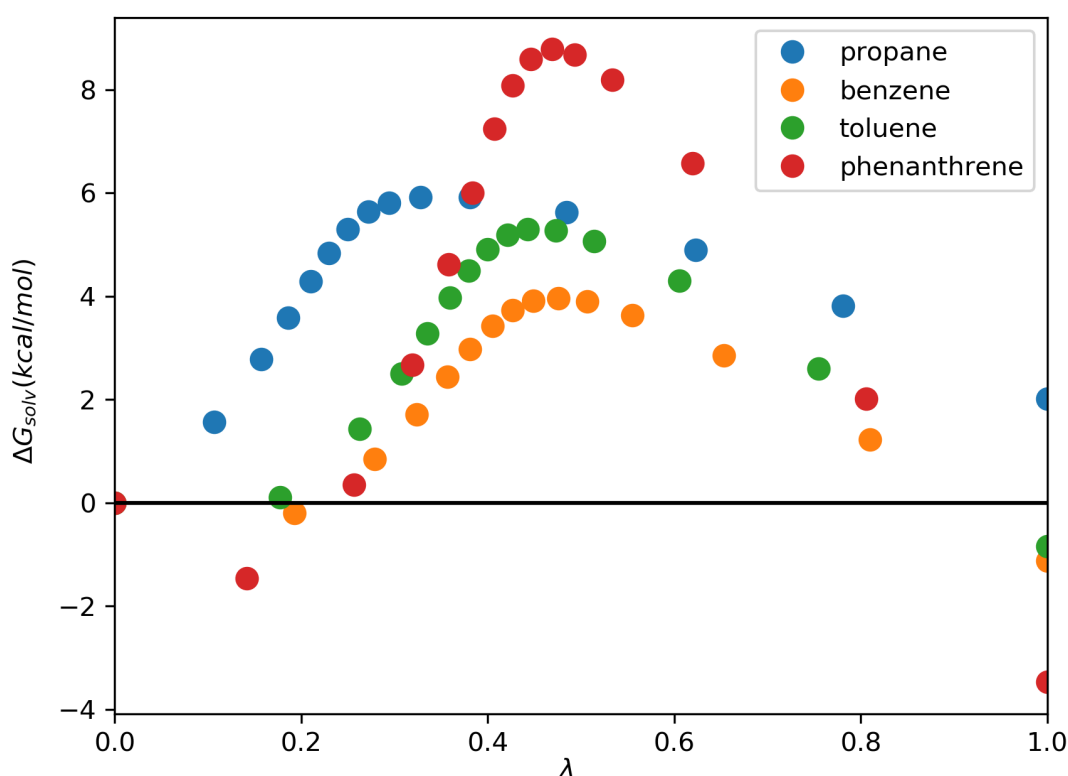


Figure 5.2.1 – Hydration free energy profiles for different solutes.

5.3 Partition Coefficients

Using the solvation free energies estimated in the sections above, we also calculated partition coefficients, Eq. (3.84), for toluene/hexane, water/hexane, and water/1-octanol. The later is the one with most experimental data in the literature because 1-octanol is used to quantify hydrophobicity and it can serve as a model for biological lipids and different soils (RUELLE, 2000). Our interest in calculating the other two coefficients is due to the use of hexane as a model for an apolar, hydrophobic phase and because of the characterization of asphaltenes by hexane and toluene. Calculated values and experimental data are in Table 9. The experimental data of partition coefficients were only available for water/1-octanol (POOLE; DURHAM; KIBBEY, 2000; SANGSTER, 1997) and water/hexane (SCHULTE *et al.*, 1998). Hence, for the partition coefficient of toluene/ hexane, we used the solvation free energy experimental data exposed in Table 4 as input in Eq. (3.84) to generate estimated experimental data.

Table 9 – Partition Coefficient Calculated from MD simulations and from experimental data.

	Molecular Dynamics	Experimental	Absolute Deviation
	$\log P^{\text{water}/1\text{-octanol}}$		
propane	2.47	2.40	0.07
phenanthrene	3.57	4.46	0.89
	$\log P^{\text{water}/\text{hexane}}$		
benzene	1.93	2.06	0.13
phenanthrene	4.17	4.49	0.32
	$\log P^{\text{toluene}/\text{hexane}}$		
pyrene	-0.67	-0.97	0.30
phenanthrene	-0.47	-	-

Overall absolute deviations were small for pairs with smaller solvation free energy deviations such as propane and benzene. The water/1-octanol partition coefficient of phenanthrene had higher deviation due to the higher deviation of the free energy of solvation of this compound in 1-octanol. The toluene/hexane partition coefficient of pyrene has its high deviation related to the prediction of the solvation free energy in toluene. Comparing with other force fields, [Garrido *et al.* \(2009\)](#) reported average absolute deviations for the water/1-octanol partition coefficient of 0.4 for Gromos, 0.3 for TraPPE and 0.9 for OPLS-AA/TraPPE force fields. However, they attribute the low deviations of TraPPE to the cancellation of errors between the two solvation free energies. Additionally, [Genheden \(2016\)](#) found average absolute deviations of 0.86 for the water/hexane partition coefficient and of 0.75 for the water/1-octanol partition coefficient with the ELBA coarse-grained force field. Although these papers didn't predict the partition coefficient of phenanthrene and pyrene in the pair of solvents, we can say that the predictions for our small sets of solute-solvent stayed in the same overall range of absolute deviations observed for the other predictions in this work. However, a larger set would be necessary to do a complete evaluation of this force field using partition coefficients.

6 Conclusions

This dissertation consisted of the study of solvation free energy calculations of aromatic solutes that can mimic asphaltenes in different solvents with the SAFT- γ Mie coarse-grained force field. Solvation free energy studies are mostly done using water as a solvent and with all-atom force fields based on the Lennard Jones Potential, therefore, with this study, we provided data that were lacking in the literature. Additionally, the free energy estimations done can help improve the SAFT- γ Mie force field since these calculations are helpful in identifying errors in the modeling process. The SAFT- γ Mie uses the SAFT-VR Mie EoS in its parameterization, which results in a more straightforward method of obtaining parameters. Following this strategy, the phenanthrene parameters, which were not available in this force field database, were obtained using phase equilibrium data and two different ring equations and geometries. However, only the parameters estimated with the ring equation proposed by [Müller and Mejía \(2017\)](#) were utilized in the solvation free energy simulations since this equation did not have inconsistencies in its parameterization.

To obtain accurate solvation free energies, we carefully selected and optimized the coupling parameter and their respective simulation weights used in our Expanded Ensemble simulations. The resulting potential energies from these simulations were then served as input to estimate solvation free energy differences with the MBAR method. The results for solvation free energy differences with non-aqueous solvents had absolute deviations to the experimental data of less than 2.0 kcal/mol, except for the pair 1-octanol+anthracene. We also observed the geometry effect on the free energy curves - larger molecules had steeper curves and more substantial absolute deviations. The influence of carbon dioxide on the solvation free energy of phenanthrene in toluene was found to be minimum. The ΔG_{solv} decreased slightly until the mass fraction of CO_2 was equal to 0.119 and, after this point, solvation free energies increased.

Hydration free energy differences calculations with the SAFT- γ Mie model

required the use of relatively larger values of k_{ij} to obtain satisfactory results. We chose to estimate the parameter with the output from molecular dynamics data since the strategy of using the SAFT-VR Mie EoS did not provide good results. This necessity of one additional parameter happens probably due to the lack of a term to account for the hydrogen bond on the EoS that the model is based and due to the problems associated with the coarse-graining of water molecules. The results with k_{ij} estimated with MD output were great, the absolute deviations to the experimental data found were smaller than the ones for the GAFF and ELBA force field. We also used the solvation free energies to calculate partition coefficients in water/1-octanol, water/hexane, and toluene/hexane. The absolute deviations to the experimental data obtained were similar to the ones found for all-atom force fields (Gromos, TraPPE and OPLS-AA/TraPPE) and other coarse-grained force field (ELBA).

Overall, the SAFT- γ Mie force field proved to be an excellent model to represent the solvation phenomenon. It correctly described solvation free energy differences of solutes mimicking asphaltenes in hexane, toluene, 1-octanol, and water. The requirement of binary interaction parameter estimated with MD output for hydration free energies increases the simulation time, which is already more significant for this water model due to its coarse-graining level. Nevertheless, the SAFT- γ Mie force field for water used does not predict freezing at room temperature as other force fields, which is essential for our hydration free energy calculations. Based on this work, we have some ideas for future development. We intend to test the binary interaction parameter representability to calculations with other aromatic solutes in water. Additionally, we want to use the SAFT- γ Mie force field to model more complex asphaltenes models and, consequently, increase the scale of the simulations we performed. The final step to expand this work would be to develop new methodologies to use solvation free energies to calculate solubility efficiently.

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APPENDIX A – Optimized values of λ and η

Table 10 – Optimized values of λ and η for the solutes in hexane

pyrene		phenanthrene	
λ	η	λ	η
0.000	0.000	0.000	0.000
0.076	4.234	0.090	1.981
0.107	5.620	0.132	3.461
0.132	6.499	0.161	4.494
0.152	6.690	0.185	5.185
0.170	6.643	0.205	5.552
0.189	6.461	0.224	5.725
0.213	6.091	0.244	5.722
0.242	5.566	0.268	5.523
0.280	4.729	0.305	4.975
0.355	2.853	0.372	3.576
0.483	-0.778	0.500	0.297
0.678	-6.947	0.560	-1.390
0.788	-10.631	0.722	-6.309
1.000	-18.141	1.000	-15.448

Table 11 – Optimized values of λ and η for the solutes in 1-octanol

propane		anthracene		phenanthrene	
λ	η	λ	η	λ	η
0.000	0.000	0.000	0.000	0.000	0.000
0.027	3.126	0.078	3.932	0.049	2.578
0.050	5.109	0.111	6.178	0.091	5.663
0.073	6.093	0.130	7.426	0.125	8.575
0.095	6.570	0.143	8.201	0.144	10.069
0.117	6.826	0.154	8.717	0.157	10.978
0.142	6.956	0.164	9.085	0.169	11.599
0.174	6.969	0.174	9.357	0.180	12.040
0.215	6.847	0.184	9.556	0.192	12.340
0.269	6.554	0.197	9.676	0.206	12.499
0.337	6.050	0.214	9.681	0.225	12.478
0.427	5.228	0.238	9.490	0.253	12.161
0.545	3.955	0.274	8.958	0.298	11.280
0.720	1.843	0.326	7.906	0.371	9.406
1.000	-1.903	0.399	6.088	0.484	5.891
		0.515	2.777	0.664	-0.516
		0.695	-2.960	0.802	-5.908
		1.000	-13.768	1.000	-14.073

Table 12 – Optimized values of λ and η for the solutes in toluene

pyrene		anthracene		phenanthrene	
λ	η	λ	η	λ	η
0.000	0.000	0.000	0.000	0.000	0.000
0.090	2.563	0.119	0.218	0.136	0.726
0.130	4.338	0.174	1.210	0.191	2.307
0.154	5.439	0.209	2.052	0.223	3.430
0.172	6.181	0.236	2.664	0.246	4.233
0.188	6.670	0.261	3.122	0.264	4.780
0.204	6.986	0.283	3.378	0.281	5.149
0.222	7.121	0.306	3.449	0.299	5.354
0.244	7.025	0.332	3.311	0.318	5.389
0.278	6.520	0.360	2.936	0.340	5.222
0.340	5.010	0.399	2.209	0.372	4.717
0.462	1.247	0.466	0.567	0.425	3.440
0.616	-4.283	0.564	-2.211	0.524	0.444
0.788	-11.032	0.715	-6.983	0.701	-5.814
1.000	-19.814	1.000	-16.923	1.000	-17.803

Table 13 – Optimized values of λ and η for the phenanthrene in different mass fractions of CO_2 in toluene

0.087		0.119		0.169		0.289	
λ	η	λ	η	λ	η	λ	η
0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
0.128	0.604	0.128	0.732	0.064	0.883	0.066	0.806
0.184	2.067	0.186	2.223	0.108	0.764	0.111	0.760
0.217	3.164	0.219	3.319	0.175	1.969	0.172	1.983
0.240	3.940	0.244	4.098	0.214	3.156	0.204	2.967
0.260	4.472	0.267	4.704	0.240	3.974	0.227	3.627
0.277	4.823	0.289	5.031	0.258	4.457	0.245	4.082
0.295	5.035	0.313	5.084	0.273	4.750	0.262	4.395
0.318	5.059	0.339	4.950	0.287	4.921	0.279	4.583
0.347	4.762	0.373	4.371	0.305	4.962	0.299	4.621
0.397	3.753	0.425	3.055	0.326	4.885	0.325	4.423
0.491	1.031	0.488	1.196	0.361	4.401	0.365	3.739
0.670	-5.148	0.525	-0.027	0.419	2.990	0.428	2.198
0.791	-9.713	0.730	-7.185	0.527	-0.299	0.530	-0.842
1.000	-18.098	1.000	-17.769	0.697	-6.180	0.701	-6.763
				1.000	-17.998	1.000	-18.163

Table 14 – Optimized values of λ and η for the solutes in water

propane		benzene		toluene		phenanthrene	
λ	η	λ	η	λ	η	λ	η
0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
0.107	2.673	0.193	-0.295	0.177	0.182	0.142	-2.462
0.157	4.703	0.279	1.468	0.262	2.432	0.256	0.597
0.186	6.047	0.324	2.931	0.307	4.244	0.319	4.504
0.210	7.148	0.357	4.168	0.336	5.552	0.358	7.762
0.230	8.017	0.381	5.091	0.360	6.696	0.384	10.104
0.250	8.883	0.405	5.891	0.380	7.558	0.407	12.185
0.272	9.291	0.427	6.443	0.400	8.233	0.427	13.607
0.294	9.700	0.449	6.770	0.422	8.678	0.446	14.490
0.328	9.900	0.476	6.900	0.443	8.859	0.469	14.834
0.381	9.930	0.506	6.805	0.473	8.810	0.494	14.667
0.484	9.463	0.555	6.392	0.514	8.452	0.533	13.832
0.623	8.195	0.653	5.109	0.606	7.148	0.620	11.069
0.781	6.378	0.810	2.421	0.755	4.273	0.806	3.279
1.000	3.333	1.000	-1.480	1.000	-1.547	1.000	-6.122

APPENDIX B – Overlapping Matrices

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.75	.19	.05	.01											
1	.28	.33	.19	.09	.05	.03	.02	.01							
2	.08	.21	.19	.15	.12	.10	.07	.04	.02	.01					
3	.01	.08	.11	.13	.14	.15	.15	.11	.07	.03	.01				
4		.03	.05	.08	.12	.16	.18	.16	.12	.07	.03	.01			
5		.01	.02	.05	.09	.14	.18	.18	.15	.10	.05	.01			
6			.01	.03	.07	.12	.17	.19	.18	.13	.07	.02			
7				.02	.04	.09	.14	.19	.20	.16	.11	.04	.01		
8				.01	.03	.06	.12	.17	.20	.19	.15	.06	.01		
9					.02	.04	.08	.14	.19	.20	.19	.09	.02	.01	
10					.01	.02	.05	.09	.14	.19	.25	.17	.06	.03	
11						.01	.02	.04	.07	.11	.21	.27	.17	.09	.01
12								.01	.02	.03	.08	.18	.32	.29	.07
13									.01	.01	.03	.10	.30	.38	.16
14											.01	.03	.16	.36	.44

Figure B.0.1 – Overlapping matrix for hexane+pyrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.54	.29	.11	.04	.01										
1	.29	.31	.19	.11	.05	.03	.01	.01							
2	.14	.25	.20	.15	.10	.07	.04	.02	.01						
3	.06	.15	.17	.16	.14	.11	.09	.06	.04	.02					
4	.02	.07	.11	.13	.14	.13	.13	.11	.09	.05	.02				
5	.01	.03	.06	.09	.12	.13	.14	.14	.13	.10	.04	.01			
6		.01	.03	.05	.08	.11	.13	.15	.17	.16	.07	.02	.01		
7			.01	.03	.05	.08	.11	.15	.19	.21	.12	.03	.02		
8				.01	.03	.05	.08	.12	.19	.25	.17	.06	.03		
9					.01	.02	.05	.09	.16	.26	.24	.10	.06	.01	
10						.01	.02	.04	.09	.20	.27	.19	.14	.03	
11								.01	.03	.08	.17	.28	.28	.13	.02
12									.01	.04	.12	.27	.31	.20	.03
13										.01	.04	.17	.27	.37	.14
14											.01	.05	.09	.31	.55

Figure B.0.2 – Overlapping matrix for hexane+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.66	.26	.06	.01											
1	.40	.30	.14	.08	.04	.02	.01								
2	.10	.14	.17	.22	.15	.09	.06	.03	.02	.01					
3	.01	.03	.09	.20	.20	.16	.12	.08	.05	.03	.01	.01			
4		.01	.05	.14	.18	.17	.15	.12	.08	.05	.03	.01	.01		
5			.02	.09	.15	.16	.16	.14	.11	.08	.04	.02	.01		
6			.01	.06	.11	.14	.15	.16	.13	.11	.06	.04	.02	.01	
7			.01	.04	.07	.11	.13	.16	.15	.14	.09	.06	.03	.01	
8				.02	.05	.07	.10	.14	.16	.17	.13	.09	.05	.02	
9				.01	.02	.04	.07	.11	.14	.18	.16	.14	.09	.03	.01
10					.01	.02	.04	.07	.10	.16	.18	.18	.15	.06	.01
11					.01	.01	.02	.04	.06	.11	.15	.21	.24	.13	.02
12							.01	.02	.03	.06	.10	.19	.30	.24	.05
13								.01	.01	.02	.05	.11	.26	.38	.16
14										.01	.02	.04	.13	.34	.47

Figure B.0.3 – Overlapping matrix for 1-octanol+propane.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17
0	.57	.28	.09	.03	.01	.01												
1	.28	.32	.17	.09	.06	.03	.02	.01	.01									
2	.13	.25	.19	.12	.09	.07	.05	.04	.03	.02	.01							
3	.06	.17	.17	.13	.11	.09	.07	.06	.05	.04	.03	.01						
4	.03	.11	.13	.12	.11	.10	.09	.08	.08	.06	.05	.03	.01					
5	.01	.07	.10	.10	.10	.10	.09	.09	.09	.09	.08	.05	.02	.01				
6	.01	.04	.07	.07	.08	.09	.09	.10	.11	.11	.10	.07	.04	.02				
7		.02	.04	.05	.07	.08	.08	.10	.11	.12	.13	.10	.06	.03	.01			
8		.01	.02	.03	.05	.06	.07	.09	.11	.13	.15	.13	.09	.04	.01			
9			.01	.02	.03	.04	.05	.07	.10	.13	.17	.16	.12	.06	.02	.01		
10				.01	.02	.02	.04	.05	.08	.11	.17	.19	.16	.10	.04	.01		
11					.01	.01	.02	.03	.05	.09	.14	.19	.20	.15	.07	.02		
12							.01	.01	.03	.05	.10	.17	.22	.21	.13	.05	.01	
13								.01	.01	.03	.06	.11	.19	.25	.22	.11	.02	
14										.01	.02	.05	.12	.21	.29	.22	.06	
15											.01	.02	.04	.11	.24	.36	.21	.01
16													.01	.03	.09	.28	.49	.10
17															.01	.06	.33	.59

Figure B.0.4 – Overlapping matrix for 1-octanol+anthracene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17
0	.51	.34	.11	.03	.01													
1	.37	.36	.17	.06	.02	.01												
2	.22	.32	.23	.12	.06	.03	.02	.01										
3	.10	.20	.22	.18	.11	.07	.05	.03	.02	.01	.01							
4	.04	.12	.16	.18	.13	.11	.08	.07	.05	.03	.02	.01						
5	.02	.06	.10	.14	.13	.12	.10	.10	.09	.07	.05	.02	.01					
6	.01	.03	.06	.10	.11	.11	.11	.11	.11	.10	.08	.05	.02					
7		.01	.03	.06	.07	.09	.09	.11	.13	.13	.13	.09	.05	.01				
8			.01	.03	.04	.06	.07	.10	.13	.14	.17	.14	.09	.02				
9				.01	.02	.03	.05	.07	.11	.14	.19	.19	.15	.04				
10					.01	.01	.02	.04	.07	.11	.19	.23	.22	.07	.01			
11							.01	.02	.04	.07	.15	.23	.30	.14	.02			
12								.01	.02	.04	.09	.18	.35	.24	.06	.01		
13										.01	.03	.10	.28	.34	.16	.06	.01	
14											.01	.03	.12	.27	.28	.23	.06	
15													.02	.06	.14	.44	.31	.03
16														.01	.04	.33	.49	.13
17															.01	.13	.45	.41

Figure B.0.5 – Overlapping matrix for 1-octanol+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.53	.30	.11	.04	.01	.01									
1	.21	.34	.22	.12	.06	.03	.01	.01							
2	.09	.25	.24	.16	.11	.07	.04	.02	.01						
3	.04	.17	.20	.17	.14	.11	.08	.05	.03	.01					
4	.02	.10	.15	.15	.14	.13	.11	.09	.07	.04	.01				
5	.01	.05	.10	.12	.13	.13	.13	.12	.11	.08	.02				
6		.02	.05	.08	.10	.12	.13	.14	.15	.14	.05	.01			
7		.01	.02	.04	.06	.09	.11	.14	.19	.21	.11	.02			
8			.01	.02	.03	.05	.08	.12	.19	.27	.18	.05	.01		
9					.01	.02	.04	.07	.15	.29	.28	.10	.02		
10							.01	.03	.08	.22	.35	.23	.06	.01	
11								.01	.02	.08	.22	.38	.22	.07	.01
12										.02	.07	.25	.37	.25	.05
13											.01	.07	.26	.45	.20
14												.01	.08	.36	.55

Figure B.0.6 – Overlapping matrix for toluene+pyrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.48	.29	.14	.05	.02	.01									
1	.24	.30	.23	.12	.06	.03	.01								
2	.12	.24	.24	.17	.11	.06	.03	.02	.01						
3	.06	.16	.21	.19	.14	.09	.07	.04	.02	.01					
4	.03	.10	.16	.17	.15	.12	.10	.07	.05	.03	.02				
5	.01	.05	.10	.13	.14	.13	.13	.11	.09	.06	.04	.01			
6		.02	.05	.08	.10	.11	.13	.13	.12	.11	.09	.03	.01		
7		.01	.02	.04	.06	.08	.11	.13	.15	.15	.15	.07	.02		
8			.01	.02	.03	.05	.08	.11	.14	.18	.21	.12	.05	.01	
9				.01	.01	.02	.05	.08	.12	.18	.26	.18	.08	.02	
10						.01	.02	.04	.08	.15	.27	.24	.14	.04	
11							.01	.02	.04	.09	.21	.28	.25	.10	.01
12								.01	.01	.04	.12	.23	.33	.23	.03
13										.01	.04	.11	.28	.41	.14
14												.02	.08	.29	.60

Figure B.0.7 – Overlapping matrix for toluene+anthracene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.52	.31	.11	.04	.01										
1	.20	.34	.23	.12	.06	.03	.01	.01							
2	.07	.24	.25	.17	.11	.07	.04	.02	.01						
3	.03	.15	.21	.18	.14	.11	.07	.05	.03	.01	.01				
4	.01	.09	.16	.16	.15	.13	.11	.08	.06	.03	.02				
5		.05	.11	.13	.14	.13	.13	.11	.09	.06	.04	.01			
6		.02	.06	.09	.11	.12	.13	.13	.12	.10	.08	.03	.01		
7		.01	.03	.05	.08	.10	.12	.13	.14	.14	.13	.07	.02		
8			.01	.03	.04	.07	.09	.11	.14	.16	.18	.13	.04		
9				.01	.02	.04	.06	.08	.12	.16	.22	.20	.07	.01	
10					.01	.01	.03	.05	.08	.13	.24	.29	.14	.02	
11							.01	.02	.04	.08	.19	.34	.26	.07	
12									.01	.03	.09	.25	.39	.22	.01
13											.02	.07	.25	.53	.13
14													.03	.27	.69

Figure B.0.8 – Overlapping matrix for toluene+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.54	.30	.11	.04	.01										
1	.23	.34	.23	.11	.05	.03	.01	.01							
2	.09	.25	.25	.17	.10	.06	.04	.02	.01						
3	.04	.16	.22	.18	.13	.10	.08	.05	.03	.01					
4	.02	.10	.16	.17	.14	.12	.11	.08	.05	.03	.01				
5	.01	.05	.11	.13	.13	.13	.13	.12	.09	.06	.03				
6		.02	.06	.10	.11	.12	.14	.14	.13	.11	.06	.01			
7		.01	.03	.06	.07	.10	.13	.14	.15	.16	.11	.03			
8			.01	.02	.04	.06	.09	.13	.16	.21	.20	.07	.01		
9				.01	.01	.03	.05	.08	.13	.22	.29	.14	.02		
10						.01	.02	.03	.07	.17	.35	.27	.06	.02	
11								.01	.02	.07	.24	.39	.19	.07	.01
12										.01	.05	.19	.37	.32	.07
13											.01	.07	.29	.43	.20
14												.01	.11	.34	.54

Figure B.0.9 – Overlapping matrix for toluene+CO₂(0.087)+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.55	.31	.10	.03	.01										
1	.24	.36	.22	.10	.05	.02	.01								
2	.09	.27	.26	.17	.10	.06	.03	.01	.01						
3	.04	.17	.22	.18	.14	.10	.07	.04	.02	.01					
4	.01	.09	.16	.17	.15	.14	.12	.08	.05	.03	.01				
5		.04	.09	.12	.13	.15	.15	.12	.10	.07	.03	.01			
6		.01	.04	.07	.10	.12	.15	.15	.15	.12	.06	.02	.01		
7			.01	.03	.05	.08	.12	.14	.17	.19	.12	.05	.03		
8				.01	.02	.04	.08	.11	.17	.22	.19	.09	.05	.01	
9					.01	.02	.03	.06	.12	.22	.26	.16	.10	.01	
10							.01	.02	.06	.15	.27	.24	.19	.04	
11								.01	.03	.09	.22	.28	.26	.11	.01
12								.01	.02	.06	.18	.28	.28	.16	.01
13										.01	.05	.13	.18	.48	.14
14											.01	.02	.03	.32	.62

Figure B.0.10 – Overlapping matrix for toluene+ CO_2 (0.119)+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
0	.32	.36	.24	.06	.02											
1	.24	.34	.27	.09	.03	.01	.01									
2	.18	.30	.28	.13	.06	.02	.01	.01								
3	.08	.19	.24	.19	.12	.06	.04	.03	.02	.01						
4	.03	.10	.16	.19	.16	.10	.08	.06	.05	.03	.02	.01				
5	.01	.05	.09	.14	.15	.12	.11	.09	.08	.07	.05	.02	.01			
6	.01	.03	.05	.10	.12	.11	.11	.11	.11	.10	.08	.05	.02			
7		.01	.03	.07	.09	.09	.11	.11	.12	.12	.11	.08	.04	.01		
8		.01	.01	.04	.06	.07	.09	.11	.12	.13	.14	.12	.07	.01		
9				.02	.03	.05	.07	.09	.11	.14	.16	.17	.13	.03		
10				.01	.01	.02	.04	.06	.08	.12	.17	.21	.21	.06	.01	
11						.01	.02	.03	.05	.08	.14	.23	.31	.12	.02	
12								.01	.02	.03	.07	.18	.37	.24	.07	
13										.01	.02	.07	.26	.37	.25	.01
14												.01	.07	.24	.55	.12
15													.01	.04	.30	.66

Figure B.0.11 – Overlapping matrix for toluene+ CO_2 (0.169)+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
0	.33	.33	.22	.07	.02	.01										
1	.25	.31	.25	.11	.05	.02	.01									
2	.17	.27	.26	.15	.08	.04	.02	.01								
3	.08	.17	.22	.20	.13	.09	.05	.03	.02	.01						
4	.04	.10	.15	.19	.16	.12	.09	.07	.04	.03	.01	.01				
5	.02	.06	.10	.15	.15	.13	.11	.10	.07	.05	.03	.02				
6	.01	.03	.06	.11	.13	.13	.12	.12	.09	.08	.06	.03	.01			
7		.01	.03	.07	.10	.11	.12	.12	.11	.11	.10	.07	.03	.01		
8		.01	.01	.04	.07	.09	.10	.12	.12	.14	.13	.11	.05	.01		
9				.02	.04	.05	.07	.09	.11	.14	.17	.17	.10	.03		
10				.01	.01	.02	.04	.06	.08	.12	.17	.23	.18	.06	.01	
11						.01	.01	.02	.04	.08	.14	.25	.28	.14	.03	
12								.01	.01	.03	.07	.19	.33	.27	.08	
13										.01	.02	.09	.25	.38	.23	.02
14												.02	.08	.27	.50	.12
15													.01	.05	.30	.65

Figure B.0.12 – Overlapping matrix for toluene+ CO_2 (0.289)+phenanthrene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.45	.28	.16	.08	.02	.01									
1	.24	.26	.24	.16	.06	.02	.01	.01							
2	.13	.21	.26	.21	.09	.05	.02	.02	.01						
3	.07	.16	.24	.23	.12	.06	.04	.03	.03	.01					
4	.04	.11	.21	.22	.13	.08	.05	.05	.06	.04	.01				
5	.02	.07	.16	.19	.12	.08	.06	.08	.10	.08	.03	.01			
6	.01	.04	.09	.13	.09	.07	.07	.10	.16	.16	.07	.02			
7		.01	.04	.06	.05	.05	.06	.10	.20	.24	.14	.04	.01		
8			.01	.02	.03	.03	.04	.09	.20	.29	.19	.07	.02		
9				.01	.01	.01	.02	.06	.17	.29	.25	.13	.03	.01	
10							.01	.03	.11	.24	.28	.21	.08	.02	
11								.01	.04	.12	.21	.31	.22	.09	.01
12									.01	.03	.08	.22	.33	.26	.07
13										.01	.02	.08	.25	.40	.25
14												.02	.08	.31	.59

Figure B.0.13 – Overlapping matrix for water+propane.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.44	.33	.16	.05	.02	.01									
1	.19	.32	.27	.13	.05	.03	.01								
2	.08	.23	.30	.19	.10	.06	.03	.01	.01						
3	.04	.16	.27	.20	.13	.09	.05	.03	.02	.01					
4	.02	.10	.21	.19	.14	.11	.08	.06	.04	.03	.02				
5	.01	.06	.15	.16	.13	.12	.11	.08	.07	.06	.04	.01			
6		.03	.08	.11	.11	.11	.11	.10	.10	.10	.09	.04	.01		
7		.01	.04	.06	.07	.08	.10	.10	.12	.14	.16	.09	.02		
8			.01	.03	.04	.05	.07	.09	.11	.17	.23	.15	.05	.01	
9				.01	.01	.02	.04	.06	.09	.16	.27	.22	.09	.02	
10						.01	.02	.03	.06	.14	.27	.27	.14	.04	.01
11							.01	.01	.03	.09	.23	.29	.23	.08	.01
12									.01	.04	.12	.22	.31	.22	.06
13										.01	.03	.08	.22	.39	.28
14												.02	.07	.29	.62

Figure B.0.14 – Overlapping matrix for water+benzene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.46	.32	.15	.04	.01										
1	.19	.31	.28	.13	.06	.02	.01								
2	.07	.21	.32	.20	.11	.05	.02	.01							
3	.03	.13	.28	.23	.15	.08	.05	.03	.01	.01					
4	.01	.08	.23	.22	.17	.11	.07	.05	.03	.02	.01				
5	.01	.05	.16	.19	.16	.12	.09	.07	.05	.04	.03	.01			
6		.02	.10	.14	.14	.12	.11	.09	.08	.08	.07	.03	.01		
7		.01	.05	.08	.09	.10	.10	.10	.11	.12	.14	.08	.02		
8			.01	.03	.05	.06	.07	.09	.11	.15	.22	.16	.04	.01	
9				.01	.02	.03	.04	.06	.09	.14	.27	.24	.08	.01	
10						.01	.02	.03	.06	.12	.27	.31	.15	.03	
11							.01	.01	.03	.08	.23	.33	.23	.07	.01
12									.01	.03	.12	.25	.34	.22	.03
13										.01	.03	.08	.24	.45	.19
14												.01	.04	.25	.69

Figure B.0.15 – Overlapping matrix for water+toluene.

λ	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14
0	.50	.37	.10	.02											
1	.29	.38	.23	.08	.02										
2	.08	.21	.33	.26	.10	.03	.01								
3	.02	.07	.24	.34	.22	.09	.02	.01							
4		.02	.13	.31	.29	.16	.07	.02	.01						
5		.01	.06	.22	.29	.20	.12	.06	.03	.01					
6			.02	.12	.21	.21	.17	.12	.08	.04	.02	.01			
7			.01	.04	.11	.14	.16	.16	.15	.11	.09	.04			
8				.01	.03	.06	.09	.13	.17	.17	.19	.14	.02		
9						.01	.03	.06	.11	.16	.26	.30	.06		
10							.01	.02	.05	.11	.25	.43	.13	.01	
11									.01	.05	.17	.48	.27	.02	
12										.01	.05	.29	.48	.15	.01
13												.03	.17	.58	.22
14													.01	.24	.75

Figure B.0.16 – Overlapping matrix for water+phenanthrene.

APPENDIX C – Work Published in Scientific Conference



PARAMETRIZAÇÃO E AVALIAÇÃO DO CAMPO DE FORÇA SAFT- γ CG PARA SIMULAÇÃO MOLECULAR DE FENANTRENO

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O fenantreno é um hidrocarboneto policíclico aromático que pode ser usado para estudar moléculas mais complexas como os asfaltenos. Portanto, é interessante desenvolver campos de forças para simulação molecular eficientes computacionalmente que consigam descrever esses tipos de molécula. O campo de força SAFT- γ CG, escolhido para o estudo, é do tipo *coarse-grained* e foi desenvolvido com base na equação de estado SAFT-VR Mie. Essa equação modela uma substância subdividindo sua estrutura em segmentos que interagem através de potencial de Mie. Os parâmetros desse campo de força para o fenantreno foram estimados e avaliados em relação à determinação de propriedades de equilíbrio líquido-vapor (ELV). A estratégia de estimação possui duas etapas. A primeira consiste em minimizar o erro quadrático entre a pressão de vapor calculada com a equação de estado (EdE) SAFT-VR Mie e os valores experimentais. Os parâmetros dessa minimização foram, então, usados no cálculo de ELV com o método de *Gibbs Ensemble Monte Carlo* com volume total constante (GEMC-NVT). A segunda etapa consistiu em estimar novamente os parâmetros através do método de mínimos quadrados envolvendo a pressão de vapor e a densidade de líquido saturado calculadas por Monte Carlo (MC) e as calculadas com a EdE SAFT-VR Mie. Essa segunda parte é necessária devido a aproximações teóricas que geram diferenças entre os resultados da simulação molecular e da EdE. Esse campo de força apresentou uma boa descrição da pressão de vapor do fenantreno com método de GEMC quando comparado com os dados experimentais. As densidades de líquido e vapor saturados e as propriedades críticas apresentaram resultados similares aos obtidos com o campo de força atômico TraPEE-EH, sendo que o SAFT- γ CG exige menor esforço computacional. Para continuação do trabalho, pretende-se usar esses parâmetros para calcular a energia de solvatação do fenantreno, em diluição infinita, em tolueno e em solução tolueno + CO₂ com dinâmica molecular.

Palavras-chave: campo de força, SAFT- γ CG, simulação molecular, fenantreno.

Introdução

Os campos de força do tipo *coarse-grained* parametrizados a partir de propriedades experimentais são uma alternativa aos campos de força desenvolvidos com cálculos *ab initio* quando a escala das simulações moleculares precisa ser aumentada. O método *coarse-grained* consiste, basicamente, em dividir a substância em pseudo-átomos representativos de grupos de átomos. Normalmente, a parametrização do campo de força é feita a partir de informações obtidas em uma escala mais detalhada. Quando as informações utilizadas provêm do comportamento em uma escala maior, a estratégia para o modelo *coarse-grained* é dita ser do tipo *top down*. Um dos campos de força que possuem essa estratégia é o SAFT- γ CG (Avendaño *et al.*, 2011), o qual foi desenvolvido com base na equação de estado (EdE) SAFT-VR Mie (Lafitte *et al.*, 2013). Esse modelo usa o potencial de Mie para descrever moléculas formadas por segmentos conectados. A vantagem de usar essa equação como base

para o campo de força é a sua capacidade de descrever bem as propriedades do fluido, incluindo derivadas de segunda ordem, de uma variedade de sistemas (Lafitte *et al.*, 2013). Com base nessas ideias, o presente trabalho teve como objetivo parametrizar o campo de força SAFT- γ CG para o fenantreno, cuja estrutura é apresentada na Figura 1, utilizando dados de equilíbrio líquido-vapor e considerando que ele é formado por três segmentos esféricos. O fenantreno foi escolhido como forma de testar a capacidade de representação desse campo de força simplificado, já que sua parametrização para moléculas menores tem sido bem sucedida (Lafitte *et al.*, 2012). Outra razão para o estudo do fenantreno é o fato de ele servir como modelo simplificado para moléculas mais complexas, tais como os asfaltenos, já que possui uma estrutura de anéis policondensados e é solúvel em tolueno.

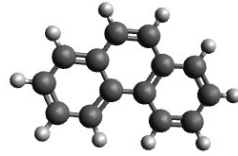


Figura1. Estrutura molecular do fenantreno.

Metodologia

Estimação com a Equação de Estado SAFT-VR Mie

A Ede SAFT-VR Mie descreve moléculas formadas por segmentos conectados e que interagem através do potencial de Mie, dado por:

$$U_{Mie}(r) = \varepsilon \frac{\lambda_r}{\lambda_r - \lambda_a} \left(\frac{\lambda_r}{\lambda_a} \right)^{\lambda_a/\lambda_r - \lambda_a} \left[\left(\frac{\sigma}{r} \right)^{\lambda_r} - \left(\frac{\sigma}{r} \right)^{\lambda_a} \right] \quad (1)$$

em que λ_r é o expoente repulsivo, λ_a é o expoente atrativo, σ corresponde à distância entre os centros dos segmentos e ε é o parâmetro de energia do segmento. A energia livre de Helmholtz específica para a Ede SAFT-VR Mie para um fluido não-associativo é definida como

$$a = a^{IDEAL} + a^{MONO} + a^{CHAIN}, \quad (2)$$

em que a^{IDEAL} é a contribuição de gás ideal, a^{MONO} é contribuição dos monômeros (segmentos desconectados) e a^{CHAIN} é a contribuição relativa à formação das cadeias de segmentos. Para cadeias que possuem m_s segmentos do mesmo tipo tangencialmente ligados, a contribuição da cadeia é

$$a^{CHAIN} = -(m_s - 1) \ln g^{Mie}(\sigma), \quad (3)$$

onde $\ln g^{Mie}(\sigma)$ é o valor da função de distribuição radial de pares para o fluido monomérico de referência (fluido de Mie). Devido à estrutura aromática do fenantreno, usou-se a seguinte expressão para anéis formados por m_s segmentos no lugar da contribuição de cadeia:

$$a^{RING} = -m_s \ln g^{Mie}(\sigma). \quad (4)$$

Essa substituição é feita porque a diferença entre uma molécula formada por cadeias e uma formada por anéis, tendo ambas o mesmo número de segmentos, é que a última possui uma ligação a mais (Lafitte *et al.*, 2012). Para o equacionamento completo da EdE, o leitor é referido ao artigo de Lafitte *et al.* (2013). Seguindo essa formulação, os parâmetros ε , λ_r e σ foram estimados e os parâmetros m_s e λ_a foram fixados em 3 e 6, respectivamente. A razão para fixar o número de segmentos em três deve-se à própria estrutura do fenantreno, que consiste em três anéis aromáticos condensados. Já o parâmetro atrativo foi fixado no valor London para facilitar a estimação, já que é comprovada a alta correlação entre os parâmetros repulsivo e atrativo do potencial (Ramrattan *et al.*, 2015). A minimização foi feita através do método PSO (*Particle Swarm Optimization*) e com apenas dados de pressão de vapor, por uma questão de indisponibilidade de dados experimentais de densidade do fenantreno como líquido saturado. A função objetivo possui a seguinte forma:

$$F_{\text{exp}}(\sigma^{SAFT}, \varepsilon^{SAFT}, \lambda_r^{SAFT}) = \sum_{i=1}^{N_p} \left(\frac{P_v^{\text{exp}}(T_i) - P_v^{SAFT}(T_i)}{P_v^{\text{exp}}(T_i)} \right)^2 \quad (5)$$

em que N_p corresponde ao número de pontos experimentais, P_v^{exp} aos pontos de pressão de vapor experimental (Mortimer e Murphy, 1923) e P_v^{SAFT} à pressão de vapor calculada com a EdE SAFT-VR Mie. Esse cálculo de equilíbrio foi feito usando como base a rotina do ponto de bolha proposta por Smith *et al.* (2007).

Cálculo com o Método “Gibbs Ensemble Monte Carlo” (GEMC)

Os parâmetros estimados com equação SAFT-VR Mie foram usados para realizar simulações no GEMC (Panagiotopoulos, 1987) com o simulador CASSANDRA (Shah e Maginn, 2011). O método de GEMC foi desenvolvido com o intuito de estudar a coexistência entre fases através da simulação simultânea de duas caixas com condições de contorno periódicas e que trocam moléculas, energia e volume entre si, mas de forma a manter o volume total constante. O equilíbrio entre elas é obtido através de passos de Monte Carlo (MC) que incluem o deslocamento aleatório das moléculas, mudança de volume e transferências aleatórias de moléculas entre as caixas. Para sistemas com apenas um componente, os cálculos são realizados mantendo-se o volume e o número de partículas total constantes (NVT), mas de maneira a permitir a variação de volume ($V = V^{liq} + V^{vap}$) e partículas ($N = N^{liq} + N^{vap}$) dentro de cada caixa. Essas simulações no GEMC-NVT foram feitas inserindo-se aleatoriamente 400 moléculas na caixa líquida e 100 moléculas na caixa vapor. As densidades iniciais das caixas foram escolhidas ajustando-as às densidades obtidas com a EdE SAFT-VR Mie, para evitar que todas as moléculas migrassem para uma única caixa ao longo da simulação. A simulação consistiu em, no mínimo, 10000 ciclos de equilibração e 100000 ciclos de produção, sendo que cada ciclo de MC corresponde a 1000 tentativas de rotação, 1000 de translação, 100 de inserção, 100 de exclusão e 10 de variação de volume. A distância de corte usada foi igual a quatro vezes o valor do diâmetro do segmento estimado e as interações de van der Waals foram calculadas através do potencial Mie com correção de longa distância (*tail correction*). As propriedades densidade de vapor (ρ_{vap}), densidade de líquido (ρ_{liq}) e pressão (P_v^{sim}) foram amostradas a cada 100 ciclos de MC e essas amostragens foram divididas em cinco blocos para os cálculos da média e do desvio padrão. Os resultados

obtidos nessas simulações foram usados para estimar coeficientes de correção para os parâmetros conformacionais do campo de força (c_σ e c_ϵ), que são relacionados aos parâmetros provenientes da EdE SAFT-VR Mie através de parâmetros oriundos de ajuste,

$$\sigma^{ajuste} = c_\sigma \sigma^{SAFT} \quad (6)$$

$$\epsilon^{ajuste} = c_\epsilon \epsilon^{SAFT} \quad (7)$$

Os parâmetros de ajuste são necessários porque as aproximações teóricas feitas na EdE geram discrepâncias entre os resultados da simulação molecular e os da EdE (Lafitte *et al.*, 2012). Por isso, a estimação desses parâmetros foi feita de maneira a diminuir as diferenças entre a pressão de vapor e a densidade de líquido saturado obtidas com a equação de estado e as obtidas via simulação molecular. A função objetivo, minimizada através do método PSO, possuiu a seguinte forma:

$$F_{sim}(c_\sigma, c_\epsilon) = \sum_{i=1}^{N_p} \left(\frac{P_v^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT}) - P_v^{SAFT}(T_i, \sigma^{ajuste}, \epsilon^{ajuste})}{P_v^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT})} \right)^2 + \sum_{i=1}^{N_p} \left(\frac{\rho_{liq}^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT}) - \rho_{liq}^{SAFT}(T_i, \sigma^{ajuste}, \epsilon^{ajuste})}{\rho_{liq}^{sim}(T_i, \sigma^{SAFT}, \epsilon^{SAFT})} \right)^2 \quad (8)$$

Com os parâmetros atrativo e repulsivo fixados nos valores encontrados com a EdE SAFT-VR Mie, o espaço paramétrico pode ser redefinido de maneira a se encontrar um conjunto de parâmetros refinado para o campo de força SAFT- γ CG (Lafitte *et al.*, 2012), que são:

$$\sigma^{sim} = \sigma^{SAFT} / c_\sigma \quad (9)$$

$$\epsilon^{sim} = \epsilon^{SAFT} / c_\epsilon \quad (10)$$

As simulações pelo método GEMC-NVT foram refeitas seguindo-se a mesma metodologia e os resultados foram comparados com os dados de referência. A determinação do ponto crítico não foi realizada por meio do método GEMC-NVT, pois ele apresenta grandes flutuações perto do ponto crítico, que fazem as caixas de simulação mudarem de identidade durante a simulação. A temperatura crítica (T_c) foi então ajustada através do método PSO com a seguinte equação:

$$(\rho_{liq} - \rho_{vap}) = A(T_c - T)^\eta \quad (11)$$

As densidades de líquido e vapor no equilíbrio na equação acima foram provenientes de simulações feitas em uma faixa de temperatura entre 476,75 e 825 K. O expoente crítico (η) foi fixado no valor correspondente ao seu valor universal (0,325) (Hansen e McDonald, 2006) e a constante A foi determinada pelo ajuste. A densidade crítica (ρ_c) foi obtida através da lei linear dos retângulos (Equação 12) na mesma faixa de temperatura usada para obter T_c .

$$\frac{(\rho_{liq} - \rho_{vap})}{2} = \rho_c + D(T_c - T) \quad (12)$$

A curva de coexistência líquido-vapor do campo de força SAFT- γ CG foi comparada com os resultados disponíveis na literatura para o campo de força TraPPE-EH (Rai e Siepmann, 2013), e as demais propriedades foram avaliadas por comparação com dados experimentais (Mortimer e Murphy, 1923; Nelson e Senseman, 1922; Linstrom e Mallard, 2017). A equação de desvio relativo absoluto médio usada possui a seguinte forma:

$$\Delta X = \frac{1}{N_p} \sum_{i=1}^{N_p} \left| \frac{X_{i^{ref}} - X_{i^{calc}}}{X_{i^{ref}}} \right| 100\%, \quad (13)$$

X^{ref} é o valor da propriedade de referência e X^{calc} é o valor da propriedade calculado.

Resultados e Discussão

Os parâmetros do fenantreno com três segmentos para o campo de força SAFT- γ CG estimados com a metodologia descrita estão expostos na Tabela 1.

Tabela 1. Parâmetros do campo de força SAFT- γ CG para o fenantreno.

$\sigma^{sim} / \text{\AA}$	$(\epsilon^{sim} / \kappa_b) / \text{K}$	λ_r^{sim}
4,008	529,646	14,339

A Figura 2 mostra a pressão de vapor do fenantreno em função de temperatura. O campo de força com parâmetros estimados conseguiu apresentar um comportamento similar ao observado nos dados experimentais, considerando a sua simplicidade e eficiência computacional. O valor da pressão de vapor para temperaturas mais próximas do ponto normal de ebulição foi subestimado, enquanto houve uma superestimação dos valores de pressão de vapor para temperaturas mais baixas. Esse mesmo comportamento e o desvio relativo da pressão de vapor, exposto na Tabela 2, foram similares aos resultados observados para os cálculos de propriedades de equilíbrio e transporte do dióxido de carbono e metano com o campo de força SAFT- γ CG em trabalhos anteriores (Aimoli *et al.*, 2014; Aimoli, 2015). Observa-se, também, que os resultados obtidos ficaram próximos dos resultados do campo de força TraPPE-EH disponíveis na literatura, o que indica que a simplificação do modelo acarreta somente em pequenos desvios no cálculo da pressão de vapor. O desvio relativo médio está disponível na Tabela 2.

Tabela 2. Desvio relativo médio entre os resultados obtidos com o campo de força SAFT- γ CG e os dados experimentais.

	$\Delta / \%$
P_v	9,01

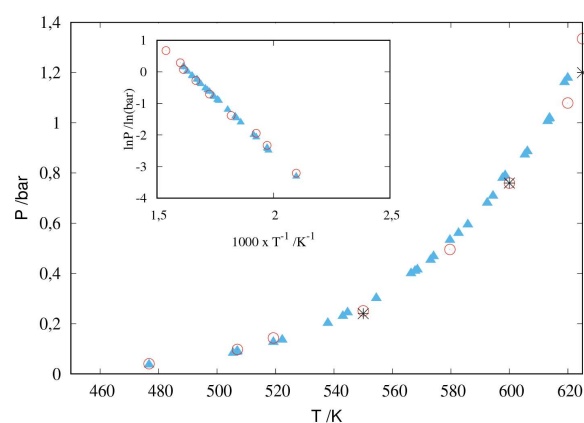


Figura 2. Pressão de vapor em função da temperatura calculada com o campo de força SAFT- γ CG (círculos vermelhos), dados experimentais (triângulos azuis) e resultados do campo de força TraPPE-EH (asteriscos pretos). A figura inserida é o mesmo gráfico, porém no formato de Clausius-Clapeyron.

Na Figura 3, as curvas de densidade de líquido e vapor saturados obtidas para o campo de força SAFT- γ CG foram comparadas com os dados disponíveis do TraPPE-EH. O campo de força estimado apresentou comportamento razoavelmente similar, mas com discrepâncias maiores do que aquelas observadas para os resultados de pressão de vapor. Adicionalmente, a densidade do fenantreno foi calculada fora do ELV no ensemble NPT-MC a uma temperatura de 298 K e a uma pressão de 1 bar como forma de comparar os resultados com o valor experimental disponível que é igual 1180 kg/m^3 (Linstrom e Mallard, 2017). O campo de força TraPPE-EH forneceu um valor de densidade igual 1110 kg/m^3 e o campo de força SAFT- γ CG forneceu um valor igual a 1455 kg/m^3 . Esses resultados mostram que o campo de força SAFT- γ CG tem mais dificuldade em prever propriedades em estados diferentes daquele que foi estimado do que o TraPPE-EH por ser um modelo coarse-grained top down. Então, essa diferença na densidade pode ter sido causada pelo uso de apenas a pressão de vapor na estimação dos parâmetros da equação de estado tida como base.

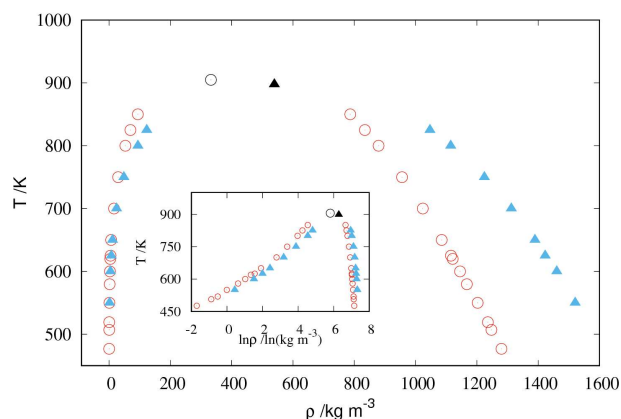


Figura 3. Curva de coexistência líquido vapor calculada com campo de força SAFT- γ CG (círculos vermelhos) e com o campo de força TraPPE-EH (triângulos azuis). O círculo e o triângulo pretos representam as propriedades críticas calculadas com o ajuste.

Os desvios relativos da densidade de líquido e de vapor saturados e das propriedades críticas encontradas com o ajuste das Equações 12-13 estão resumidos na Tabela 3. Os desvios deixam clara a existência de uma diferença significativa para densidade no equilíbrio ocasionados pela diferença de rigor teórico entre os modelos. Essas discrepâncias talvez possam ser reduzidas com uso dos dados da densidade do campo de força TraPPE-EH para estimação inicial, o que permitiria um melhor ajuste do expoente atrativo do potencial de Mie. Já com relação à previsão da temperatura crítica, os resultados entre os campos de força foram mais próximos.

Tabela 3. Propriedades críticas e desvios relativos entre o campo de força SAFT- γ CG e os dados de referência para o fenantreno.

$\Delta / \%$			
T_c	ρ_c	ρ_{liq}	ρ_{vap}
4,08	38,17	21,42	38,64

Conclusão

O campo de força SAFT- γ CG foi obtido para o fenantreno com uma metodologia em que se usa a equação de estado SAFT-VR Mie como base para parametrização desse campo de força. Os parâmetros encontrados foram avaliados através de cálculos das propriedades de equilíbrio com o método GEMC-NVT. Os resultados para pressão de vapor tiveram concordância com os dados experimentais e com o resultados da literatura obtidos para o campo de força atomístico TraPPE-EH. As densidades de líquido e vapor no equilíbrio apresentaram maiores desvios em relação ao TraPPE-EH do que a pressão de vapor. Isso mostra que a simplificação da molécula não representou uma grande perda na representação da pressão de vapor do fenantreno e que esse modelo de campo de força pode ser uma alternativa a modelos atomísticos. Com esse campo de força avaliado, pretende-se fazer cálculos de energia de solvatação do fenantreno em tolueno e em soluções tolueno+CO₂ utilizando-se dinâmica molecular.

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APPENDIX D – Paper for Publication in Scientific Journal