#### Technischen Universität München Winter Semester 2018/2019

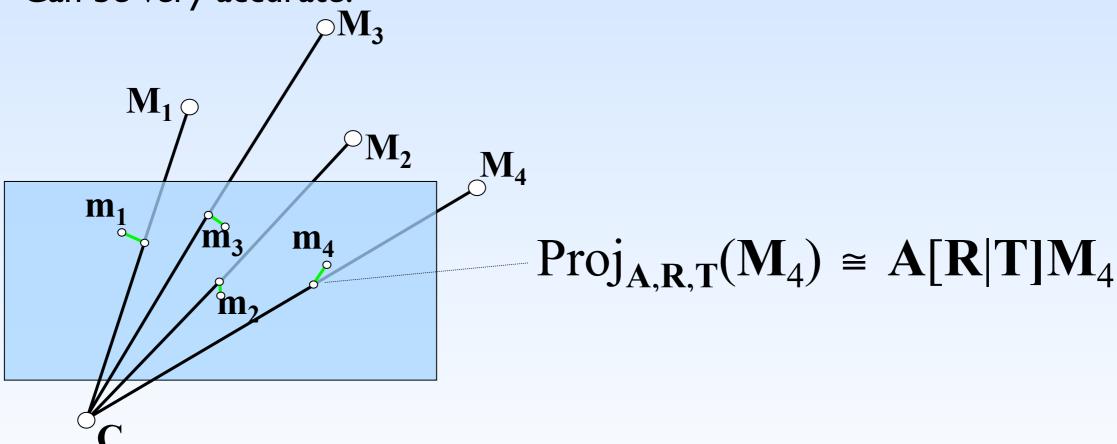
# TRACKING and DETECTION in COMPUTER VISION Non-linear optimization and robust estimation for tracking

Slobodan Ilić

## Minimization of the Reprojection Error

$$\min_{\mathbf{R},\mathbf{T}} \sum_{i} \left\| \operatorname{Proj}_{\mathbf{A},\mathbf{R},\mathbf{T}}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right\|^{2}$$

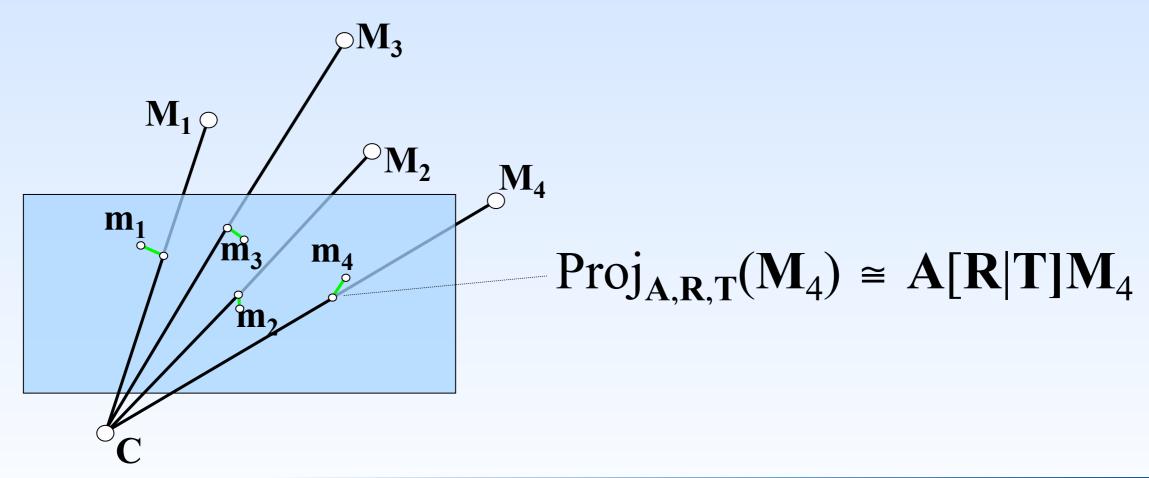
- Minimization of a physical, meaningful error (reprojection error, in pixels);
- No restriction on the number of correspondences;
- Can be very accurate.



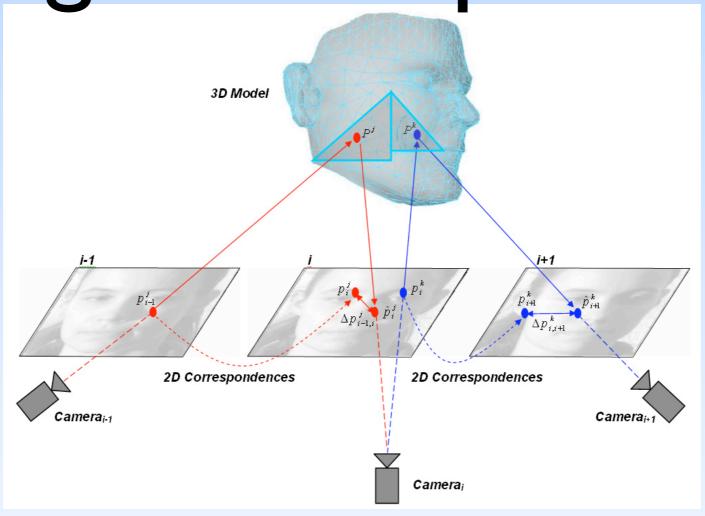
## Minimization of the Reprojection Error

$$\min_{\mathbf{R},\mathbf{T}} \sum_{i} \left\| \operatorname{Proj}_{\mathbf{A},\mathbf{R},\mathbf{T}}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right\|^{2}$$

- Non-linear least-squares minimization;
- Requires an iterative numerical optimization 
   Requires an initialization.

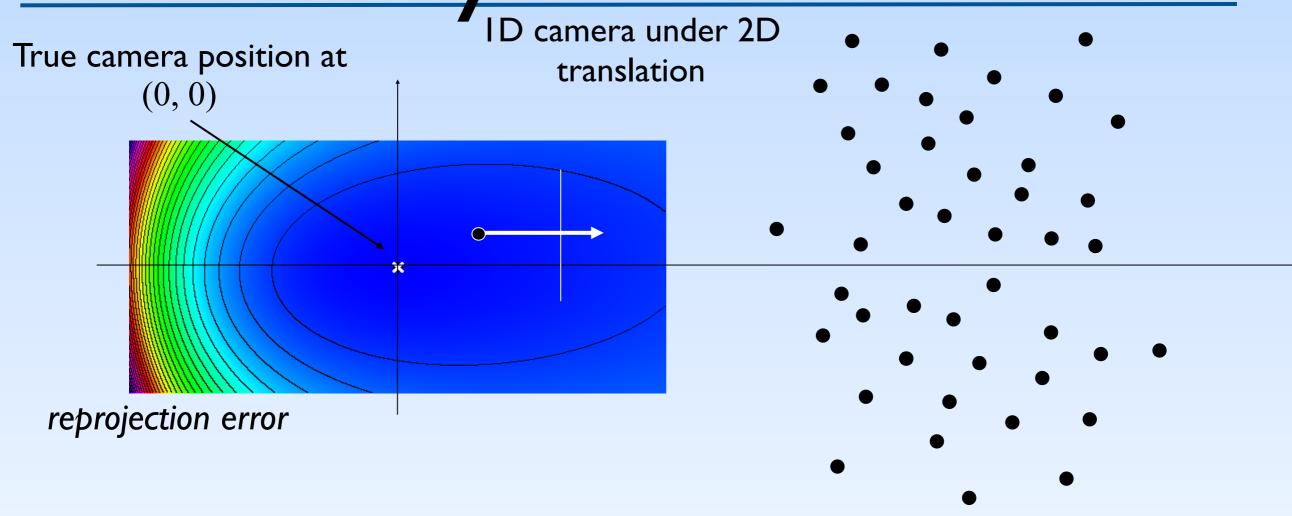


## Objective function handling image correspondences



$$\begin{split} \min_{R,T} &= \sum_{k=1}^N w_i^k \|\hat{\mathbf{p}}_{i+1}^k - \mathbf{p}_{i+1}^k\| = \sum_{k=1}^N w_i^k \|\psi(\mathbf{p}_i^k, \mathbf{\Theta}) - \mathbf{p}_{i+1}^k\| \\ & \psi(\mathbf{p}_i^k, \mathbf{\Theta}) \text{ - transfer function of a back projection} \end{split}$$

Toy Problem

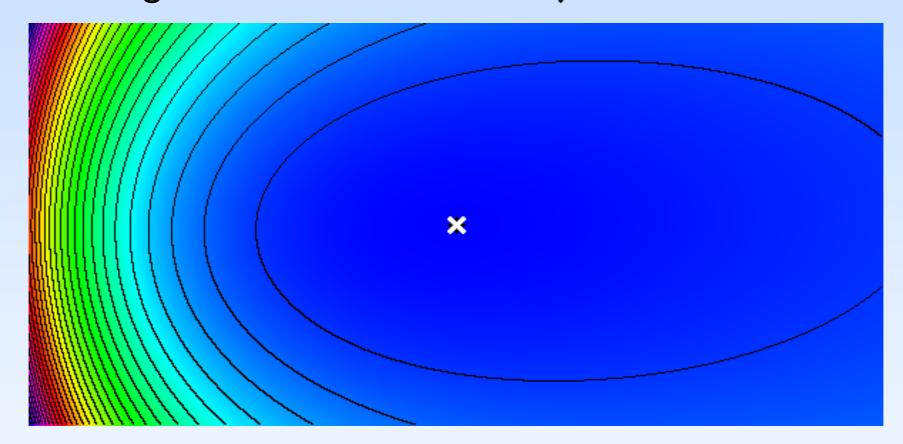


100 "3D points" taken at randomly in [400;1000]x[-500;+500]

## Gaussian Noise on the Projections

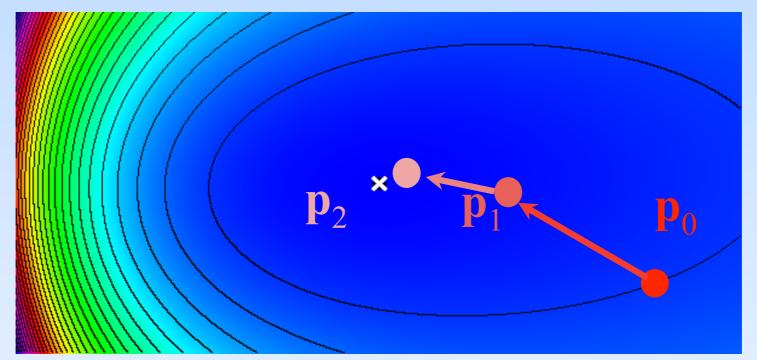
White cross: true camera position;

Black cross: global minimum of the objective function.



In case of Gaussian noise on projections, the global minimum of the objective function is very close(almost identical) to the true camera pose.

### Numerical Optimization



Start from an initial guess  $\mathbf{p}_0$ :

 $\mathbf{p}_0$  can be taken randomly but should be as close as possible to the global minimum:

- pose computed at time t-1;
- pose predicted from pose computed at time t-I and a motion model;

- ...

### Numerical Optimization

#### General methods:

- Gradient descent / Steepest Descent;
- Conjugate Gradient;
- •

#### Non-linear Least-squares optimization:

- Gauss-Newton;
- Levenberg-Marquardt;
- •

### Numerical Optimization

We want to find p that minimizes:

$$E(\mathbf{p}) = \sum_{i} \left\| \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})} (\mathbf{M}_{i}) - \mathbf{m}_{i} \right\|^{2}$$
$$= \left\| f(\mathbf{p}) - \mathbf{b} \right\|^{2}$$

where

$$f(\mathbf{p}) = \begin{bmatrix} u(\operatorname{Proj}_{\mathbf{A},\mathbf{R}(\mathbf{p}),\mathbf{T}(\mathbf{p})}(\mathbf{M}_{1})) \\ v(\operatorname{Proj}_{\mathbf{A},\mathbf{R}(\mathbf{p}),\mathbf{T}(\mathbf{p})}(\mathbf{M}_{1})) \\ \vdots \end{bmatrix} \quad b = \begin{bmatrix} u(\mathbf{m}_{1}) \\ v(\mathbf{m}_{1}) \\ \vdots \end{bmatrix}$$

- p is a vector of parameters that define the camera pose (translation vector + parameters of the rotation matrix);
- b is a vector made of the measurements (here the  $\mathbf{m}_i$ );
- f is the function that relates the camera pose to these measurements.

### Gradient descent / Steepest Descent

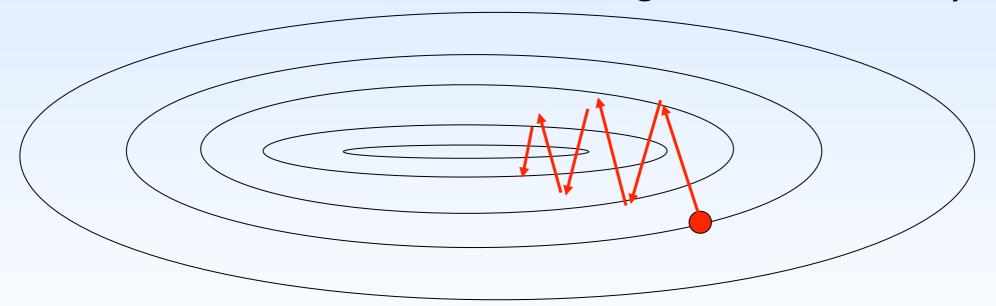
$$\mathbf{p}_{i+1} = \mathbf{p}_i - \lambda \nabla E(\mathbf{p}_i)$$

$$E(\mathbf{p}_i) = \|f(\mathbf{p}_i) - \mathbf{b}\|^2 = (f(\mathbf{p}_i) - \mathbf{b})^{\mathsf{T}} (f(\mathbf{p}_i) - \mathbf{b})$$

 $\rightarrow \nabla E(\mathbf{p}_i) = 2\mathbf{J}(f(\mathbf{p}_i) - \mathbf{b})$  with **J** the Jacobian matrix of f, computed at  $\mathbf{p}_i$ 

#### Weaknesses:

- How to choose  $\lambda$ ?
- Needs a lot of iterations in long and narrow valleys:



## The Gauss-Newton and the Levenberg-Marquardt alg.

$$E(\mathbf{p}) = \left\| f(\mathbf{p}) - \mathbf{b} \right\|^2$$

If the function f is linear ie  $f(\mathbf{p}) = \mathbf{A}\mathbf{p}$ ,  $\mathbf{p}$  can be estimated as:

$$p=A+b$$

where  $A^+$  is the pseudo-inverse of  $A: A^+=(A^TA)^{-1}A^T$ 

### Non-Linear Least-Squares: The Gauss-Newton

Iteration steps:

$$\mathbf{p}_{i+1} = \mathbf{p}_i + \Delta_i$$

 $\Delta_i$  is chosen to minimize the residual  $||f(\mathbf{p}_{i+1}) - \mathbf{b}||^2$ . It is computed by approximating f to the first order:

$$\begin{split} & \Delta_i &= \underset{\Delta}{\operatorname{argmin}} \left\| f(\mathbf{p}_i + \Delta) - \mathbf{b} \right\|^2 \\ &= \underset{\Delta}{\operatorname{argmin}} \left\| f(\mathbf{p}_i) + \mathbf{J}\Delta - \mathbf{b} \right\|^2 \quad \text{First order approximation: } f(\mathbf{p}_i + \Delta) \approx f(\mathbf{p}_i) + \mathbf{J}\Delta \\ &= \underset{\Delta}{\operatorname{argmin}} \left\| \varepsilon_i + \mathbf{J}\Delta \right\|^2 \quad \varepsilon_i = f(\mathbf{p}_i) - \mathbf{b} \text{ denotes the residual at iteration } i \end{split}$$

 $\Delta_i$  is the solution of the system  $J\Delta = -\varepsilon_i$  in the least – squares sense:  $\Delta_i = -J^+\varepsilon_i$  where  $J^+$  is the pseudo-inverse of J

## Non-Linear Least-Squares: The Levenberg-Marquardt Alg.

In the Gauss-Newton algorithm:

$$\mathbf{\Delta}_{i} = -(\mathbf{J}^{\mathsf{T}}\mathbf{J})^{-1}\mathbf{J}^{\mathsf{T}}\boldsymbol{\varepsilon}_{i}$$

In the Levenberg-Marquardt algorithm:

$$\mathbf{\Delta}_{i} = -\left(\mathbf{J}^{\mathrm{T}}\mathbf{J} + \lambda\mathbf{I}\right)^{-1}\mathbf{J}^{\mathrm{T}}\boldsymbol{\varepsilon}_{i}$$

Levenberg-Marquardt Algorithm:

- 0. Initialize  $\lambda$  with a small value:  $\lambda = 0.001$
- I. Compute  $\Delta_i$  and  $E(\mathbf{p}_i + \Delta_i)$
- 2. If  $E(\mathbf{p}_i + \Delta_i) > E(\mathbf{p}_i)$ :  $\lambda \leftarrow 10 \lambda$  and go back to 1 [happens when the linear approximation of f is too rough]
  - 3. If  $E(\mathbf{p}_i + \Delta_i) < E(\mathbf{p}_i)$ :  $\lambda \leftarrow \lambda / 10$ ,  $\mathbf{p}_{i+1} \leftarrow \mathbf{p}_i + \Delta_i$  and go back to 1.

## Non-Linear Least-Squares: The Levenberg-Marquardt Alg.

$$\boldsymbol{\Delta}_{i} = -\left(\mathbf{J}^{\mathrm{T}}\mathbf{J} + \lambda\mathbf{I}\right)^{-1}\mathbf{J}^{\mathrm{T}}\boldsymbol{\varepsilon}_{i}$$

- When  $\lambda$  is small, LM behaves similarly to the Gauss-Newton algorithm.
- When  $\lambda$  becomes large, LM behaves similarly to a steepest descent to guarantee convergence.

### Possible Parameterizations of the Rotation Matrix

Rotation in 3D space has only 3 degrees of freedom.

It would be awkward to use the nine elements as its parameters.

#### Possible parameterizations:

- Euler Angles;
- Quaternions;
- Exponential Map.

All have singularities, can be avoided by locally reparameterizing the rotation.

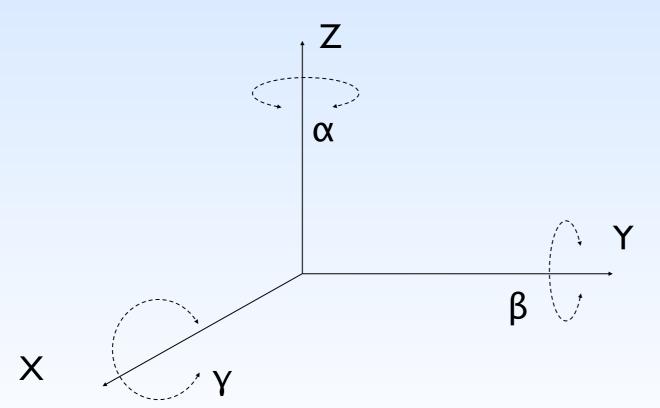
Exponential map has the best properties.

#### [From Grassia |GT98]

#### Euler Angles

Rotation defined by angles of rotation around the X-, Y-, and Z- axes. Different conventions. For example:

$$\mathbf{R} = \begin{bmatrix} \cos \alpha & -\sin \alpha & 0 \\ \sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \cos \beta & 0 & \sin \beta \\ 0 & 1 & 0 \\ -\sin \beta & 0 & \cos \beta \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \gamma & -\sin \gamma \\ 0 & \sin \gamma & \cos \gamma \end{bmatrix}$$



#### Gimbal Lock

When  $\beta = \pi/2$ ,

$$\mathbf{R} = \begin{bmatrix} \cos \alpha & -\sin \alpha & 0 \\ \sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \cos \beta & 0 & \sin \beta \\ 0 & 1 & 0 \\ -\sin \beta & 0 & \cos \beta \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \gamma & -\sin \gamma \\ 0 & \sin \gamma & \cos \gamma \end{bmatrix}$$

$$= \begin{bmatrix} \cos \alpha & -\sin \alpha & 0 \\ \sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 0 & 0 & 1 \\ 0 & 1 & 0 \\ -1 & 0 & 0 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \gamma & -\sin \gamma \\ 0 & \sin \gamma & \cos \gamma \end{bmatrix}$$

$$= \begin{bmatrix} \cos \alpha & -\sin \alpha & 0 \\ \sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 0 & \sin \gamma & \cos \gamma \\ 0 & \cos \gamma & -\sin \gamma \\ 1 & 0 & 0 \end{bmatrix}$$

$$= \begin{bmatrix} 0 & \cos\alpha\sin\gamma - \sin\alpha\cos\gamma & \cos\alpha\cos\gamma + \sin\alpha\sin\gamma \\ 0 & \sin\alpha\sin\gamma + \cos\alpha\cos\gamma & \sin\alpha\cos\gamma - \cos\alpha\sin\gamma \\ -1 & 0 & 0 \end{bmatrix} = \begin{bmatrix} 0 & \sin(\gamma - \alpha) & \cos(\gamma - \alpha) \\ 0 & \cos(\gamma - \alpha) & -\sin(\gamma - \alpha) \\ 0 & 0 & 0 \end{bmatrix}$$

## Gimbal Lock and Optimization

When  $\beta = \pi/2$ ,

$$\mathbf{R} = \begin{bmatrix} 0 & \sin(\gamma - \alpha) & \cos(\gamma - \alpha) \\ 0 & \cos(\gamma - \alpha) & -\sin(\gamma - \alpha) \\ -1 & 0 & 0 \end{bmatrix}$$

the rotation by  $\gamma$  can be cancelled by taking  $\alpha = \gamma$ .

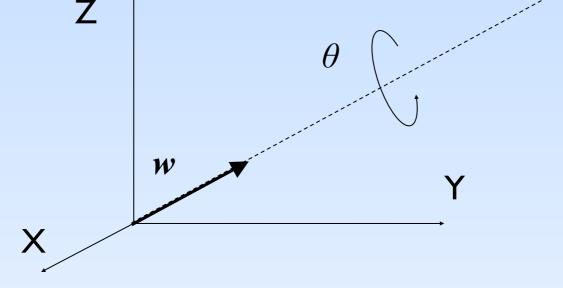
That means that

- for each possible angle  $\theta$ , all  $(\alpha, \beta = \pi/2, \gamma = \alpha + \theta)$  correspond to the same rotation matrix

=> for each possible angle  $\theta$ , there is a flat valley of axis  $(\alpha, \beta = \pi/2, \gamma = \alpha + \theta)$  in the energy to be minimized.

### Axis angle representation and quaternions

A rotation about the unit vector w by an angle  $\theta$  can be represented by the unit quaternion:



Quaternions are hyper-complex numbers that can be written as the linear combination a+bi+cj+dk, with  $i^2=j^2=k^2=ijk=-1$ .

Can also be interpreted as a scalar plus a 3- vector: (a, v).

$$q = \left(\cos\frac{\theta}{2}, w\sin\frac{\theta}{2}\right)$$

#### A Unit Quaternion

To rotate a 3D point M: write it as a quaternion p = (0, M), and take the rotated point p' to be

$$q = \left(\cos\frac{\theta}{2}, w\sin\frac{\theta}{2}\right)$$

No gimbal lock.

$$p' = q \cdot p \overline{q}$$
 with  $\overline{q} = \left(\cos \frac{\theta}{2}, -w \sin \frac{\theta}{2}\right)$ 

The norm of q must be equal to 1.  $||\mathbf{q}|| = 1$ 

In order to enforce this during optimization we have to add regularization term:  $k(1-||\mathbf{q}||^2)$ 

### Exponential Maps

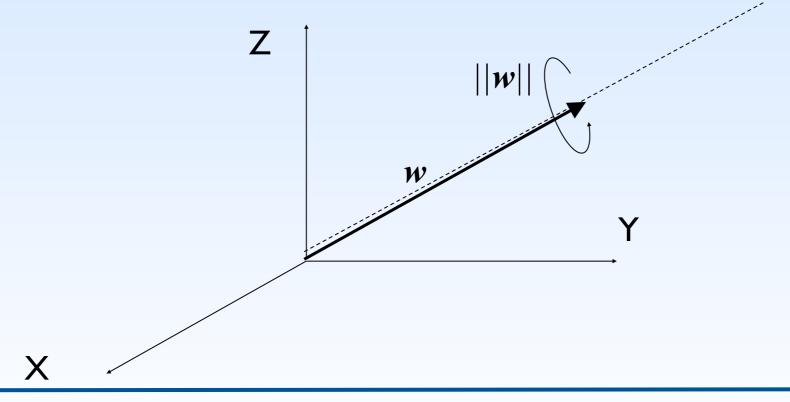
No gimbal lock;



No additional constraints;

Singularities occur in a region that can easily be avoided.

Parameterization by a 3D vector  $w = [w_1, w_2, w_3]^T$ : Rotation around the axis of direction w of an amount of ||w||



### Rodrigues' Formula

$$\mathbf{w} \times \mathbf{c} = \mathbf{\Omega} \mathbf{c} \ \hat{\Omega} = \begin{bmatrix} 0 & -w_z & w_y \\ w_z & 0 & -w_x \\ -w_y & w_x & 0 \end{bmatrix}, \|\mathbf{w}\|_2 = 1$$
 skew symmetric matrix  $\mathbf{w}$  normalised  $\frac{\omega}{||\omega||}$ ,  $\|\mathbf{w}\|_2 = 1$ 

The rotation matrix can be defined as an exponential map  $exp : so(3) \rightarrow SO(3)$  given by:

$$R = exp(\theta \hat{\Omega}) = \sum_{k=0}^{\infty} \frac{(\theta \hat{\Omega})^k}{k!} = I + \theta \hat{\Omega} + \frac{1}{2!} (\theta \hat{\Omega})^2 + \frac{1}{3!} (\theta \hat{\Omega})^3 + \dots$$

knowing 
$$\hat{\Omega}^3=-\hat{\Omega}, \hat{\Omega}^4=-\hat{\Omega}^2, \hat{\Omega}^5=\hat{\Omega}, \hat{\Omega}^6=\hat{\Omega}^2, \hat{\Omega}^7=-\hat{\Omega}^2$$

$$exp(\theta\hat{\Omega}) = I + (\theta - \frac{\theta^3}{3!} + \frac{\theta^5}{5!} - \dots)\hat{\Omega} + (\frac{\theta^2}{2!} - \frac{\theta^4}{4!} + \frac{\theta^6}{6!} - \dots)\hat{\Omega}^2$$

we obtain Rodrigues formula:

$$\mathbf{R}(\mathbf{\Omega}) = \mathbf{I} + \sin\theta \hat{\mathbf{\Omega}} + (1 - \cos\theta) \hat{\mathbf{\Omega}}^2$$

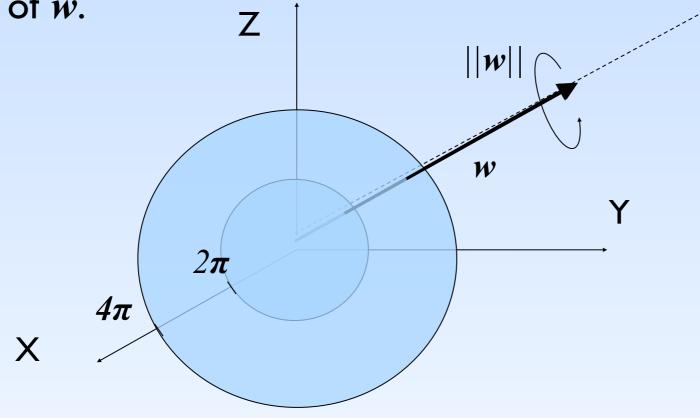
### Rodrigues' Formula

Given that  $\hat{\Omega} = \frac{\Omega}{\theta}$  we can rewrite Rodrigues formula as, where  $\Omega$  is screw symmetric matrix (not normalized) and  $\theta = ||\omega||$ 

$$R(\theta\Omega) = \exp(\theta\Omega) = I + \frac{\sin(\theta)}{\theta}\Omega + \frac{(1-\cos(\theta))}{\theta^2}\Omega^2$$

## The Singularities of Exponential Maps Rotation around the axis of direction w of an amount of ||w||

 $\rightarrow$  Singularities for w such that  $||w|| = 2n\pi$ : No rotation, whatever the direction of w.



Avoided during optimization as follows: when ||w|| becomes close to  $2n\pi$ , say higher than  $\pi$ , w can be replaced by  $\left(1-\frac{2\pi}{\|w\|}\right)w$ [From Grassia JGT98]

#### Linearization of small rotations

#### In 3D tracking:

- the camera motion between consecutive frames can often be assumed to remain small along with the corresponding rotation angles
- use a first order approximation of the rotation

$$\mathbf{M}^{'} = \mathbf{R}\mathbf{M}$$
  $pprox (\mathbf{I} + \mathbf{\Omega})\mathbf{M}$   $= \mathbf{M} + \mathbf{\Omega}\mathbf{M}$ 

 $\Omega$  is skew symmetric matrix

### Quaternions with axisangle parameterisation

To achieve the minimum number of DOF (i.e., 3), it is necessary to revert back to the axis-angle representation, this time using the following quaternion parameterization:

$$q = \left(\cos\frac{\theta}{2}, \sin\frac{\theta}{2}\left(\frac{w_1}{\theta}, \frac{w_2}{\theta}, \frac{w_3}{\theta}\right)\right)$$

where  $w = [w_1 \ w_2 \ w_3]^T$  is the axis of rotation and  $\theta = \sqrt{(w_1^2 + w_2^2 + w_3^2)}$  is norm of the axis of rotation equivalent to the angle of rotation around the axis.

This allows parameterisation of quaternions with 3 DOF instead of 4, so to avoid necessary constraint on the unit norm of the quaternion.

This also facilitate computation of the derivatives using chair ules:

$$\frac{\partial R(q(w))}{\partial w_i} = \sum_{j=0}^{3} \frac{\partial R(q)}{\partial q_j} \frac{\partial q_j(w)}{\partial w_i}$$

### Parameterization of the Rotation Matrix

Conclusion: Use exponential maps.

#### More details:

[Grassia JGT98] <a href="http://citeseerx.ist.psu.edu/viewdoc/download?doi=10.1.1.132.20&rep=rep1&type=pdf">http://citeseerx.ist.psu.edu/viewdoc/download?doi=10.1.1.132.20&rep=rep1&type=pdf</a>

George Terzakis, Phil Culverhouse, Guido Bugmann, Sanjay Sharma, and Robert Sutton

A Recipe on the Parameterization of Rotation Matrices for Non-Linear Optimization using Quaternions

### Computing J

We need to compute J, the Jacobian of f:

$$f(\mathbf{p}) = \begin{bmatrix} u(\operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_1)) \\ v(\operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_1)) \\ \vdots \end{bmatrix}$$

Solution I: Use Maple or Matlab to produce the analytical form, AND the code.

#### Solution 2

$$f(\mathbf{p}) = \begin{bmatrix} u(\operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_1)) \\ v(\operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_1)) \\ \vdots \end{bmatrix} = \begin{bmatrix} f_{\mathbf{M}_1}(\mathbf{p}) \\ \vdots \end{bmatrix}$$

#### First, decompose *f*:

$$f_{\mathbf{M}_1}(\mathbf{p}) = \mathbf{m} \Big( \tilde{\mathbf{m}} \Big( \mathbf{M}_{\mathbf{M}_1}^{cam}(\mathbf{p}) \Big) \Big)$$

#### where

- $\mathbf{M}_{\mathbf{M}_1}^{cam}(\mathbf{p})$  returns  $\mathbf{M}_1$  in the camera coordinates system defined by  $\mathbf{p}$ ;
- $\tilde{\mathbf{m}}(\mathbf{M}_{\mathbf{M}_1}^{\mathit{cam}})$  returns the projection of  $\mathbf{M}_{\mathbf{M}_1}^{\mathit{cam}}$  in homogeneous coordinates;
- $\mathbf{m}(\tilde{\mathbf{m}})$  returns the 2D vector corresponding to  $\tilde{\mathbf{m}}$ .

$$f(\mathbf{p}) = \begin{bmatrix} u(\operatorname{Proj}_{\mathbf{A},\mathbf{R}(\mathbf{p}),\mathbf{T}(\mathbf{p})}(\mathbf{M}_{1})) \\ v(\operatorname{Proj}_{\mathbf{A},\mathbf{R}(\mathbf{p}),\mathbf{T}(\mathbf{p})}(\mathbf{M}_{1})) \\ \vdots \end{bmatrix} = \begin{bmatrix} f_{\mathbf{M}_{1}}(\mathbf{p}) \\ \vdots \end{bmatrix} \text{ with } f_{\mathbf{M}_{1}}(\mathbf{p}) = \mathbf{m}(\tilde{\mathbf{m}}(\mathbf{M}_{\mathbf{M}_{1}}^{cam}(\mathbf{p})))$$

$$\mathbf{J} = \begin{bmatrix} \frac{\partial f}{\partial \mathbf{p}} \end{bmatrix} = \begin{bmatrix} \frac{\partial f_{\mathbf{M}_1}}{\partial \mathbf{p}} \\ \vdots \end{bmatrix} \text{ with } \begin{bmatrix} \frac{\partial f_{\mathbf{M}_1}}{\partial \mathbf{p}} \end{bmatrix}_{2 \times 6} = \begin{bmatrix} \frac{\partial \mathbf{m}}{\partial \tilde{\mathbf{m}}} \end{bmatrix}_{2 \times 3} \begin{bmatrix} \frac{\partial \tilde{\mathbf{m}}}{\partial \mathbf{M}_{\mathbf{M}_1}^{cam}} \end{bmatrix}_{3 \times 3} \begin{bmatrix} \frac{\partial \mathbf{M}_{\mathbf{M}_1}^{cam}}{\partial \mathbf{p}} \end{bmatrix}_{3 \times 6}$$

$$\mathbf{m}(\tilde{\mathbf{m}}) = \begin{bmatrix} u \\ v \end{bmatrix} = \begin{bmatrix} \frac{U}{W} \\ \frac{V}{W} \end{bmatrix} \text{ with } \tilde{\mathbf{m}} = \begin{bmatrix} U \\ V \\ W \end{bmatrix}$$

$$\frac{\partial \mathbf{m}}{\partial \tilde{\mathbf{m}}} = \begin{bmatrix} \frac{\partial u}{\partial U} & \frac{\partial u}{\partial V} & \frac{\partial u}{\partial W} \\ \frac{\partial v}{\partial U} & \frac{\partial v}{\partial V} & \frac{\partial v}{\partial W} \end{bmatrix} = \begin{bmatrix} 1/W & 0 & -\frac{U}{W^2} \\ 0 & 1/W & -\frac{V}{W^2} \end{bmatrix}$$

$$\widetilde{\mathbf{m}}\!\!\left(\mathbf{M}_{\mathbf{M}_1}^{cam}\right) = \mathbf{A}\mathbf{M}_{\mathbf{M}_1}^{cam}$$

$$\frac{\partial \tilde{\mathbf{m}}}{\partial \mathbf{M}_{\mathbf{M}_1}^{cam}} = \mathbf{A}$$

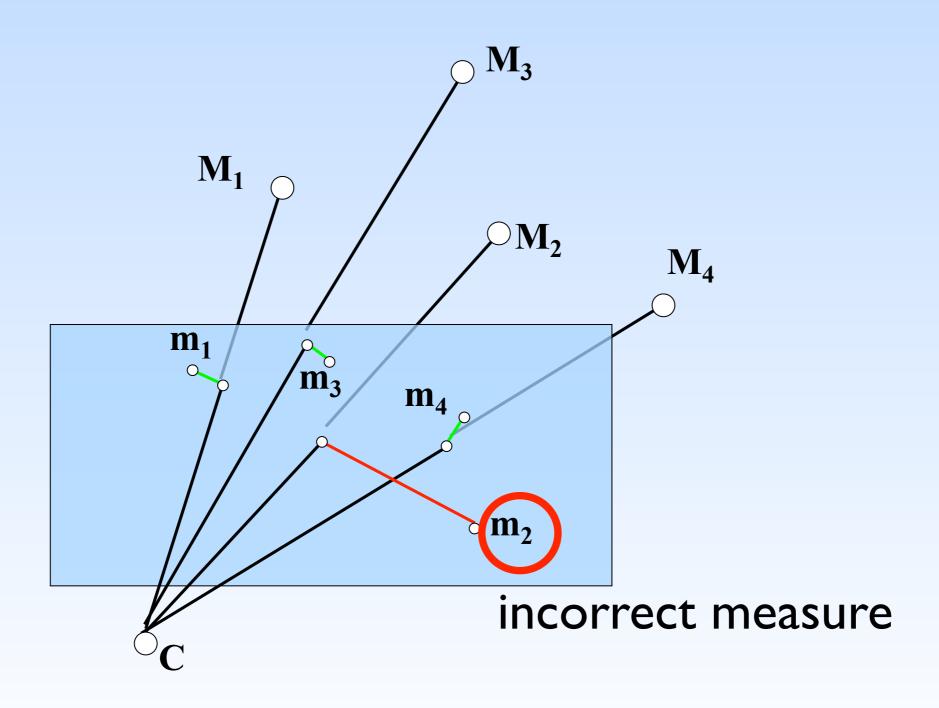
$$\mathbf{M}_{\mathbf{M}_{1}}^{cam}(\mathbf{p}) = \mathbf{R}(\mathbf{p})\mathbf{M}_{1} + \mathbf{T}$$
With  $\mathbf{p} = \begin{bmatrix} r_{1}, r_{2}, r_{3}, t_{1}, t_{2}, t_{3} \end{bmatrix}^{\mathsf{T}} \left( \mathbf{T} = \begin{bmatrix} t_{1}, t_{2}, t_{3} \end{bmatrix}^{\mathsf{T}} \right)$ :
$$\frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial \mathbf{p}} = \begin{bmatrix} \frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial r_{1}} & \frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial r_{2}} & \frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial r_{3}} & \frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial t_{1}} & \frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial t_{2}} & \frac{\partial \mathbf{M}_{\mathbf{M}_{1}}^{cam}}{\partial t_{3}} \end{bmatrix}$$

$$= \begin{bmatrix} \frac{\partial \mathbf{R}}{\partial r_{1}} \mathbf{M}_{1} & \begin{bmatrix} \frac{\partial \mathbf{R}}{\partial r_{2}} \mathbf{M}_{1} & \begin{bmatrix} \frac{\partial \mathbf{R}}{\partial r_{3}} \mathbf{M}_{1} & 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

The matrices 
$$\left[\frac{\partial \mathbf{R}}{\partial r_i}\right]_{3\times 3}$$
 can be computed from the expansion of  $\mathbf{R}$ .

In C, one can use the cvRodrigues function from the OpenCV library.

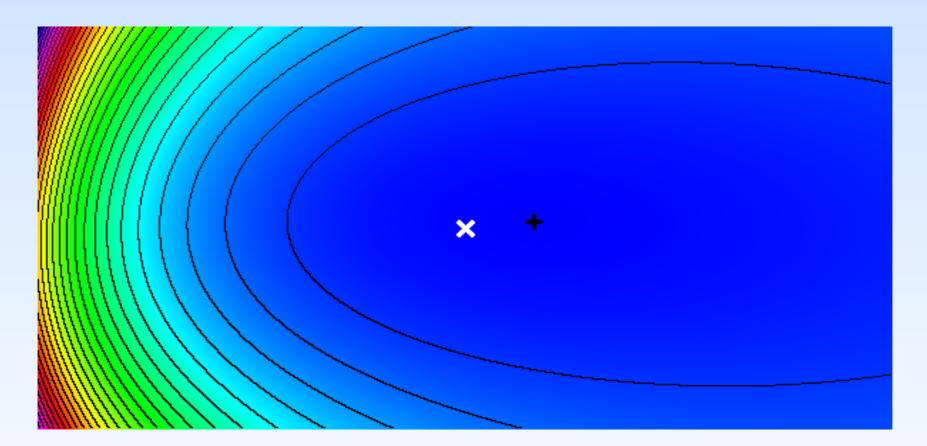
#### What if there are Outliers?



## Gaussian Noise on the Projections + 20% outliers

White cross: true camera position;

Black cross: global minimum of the objective function.



### What Happened?

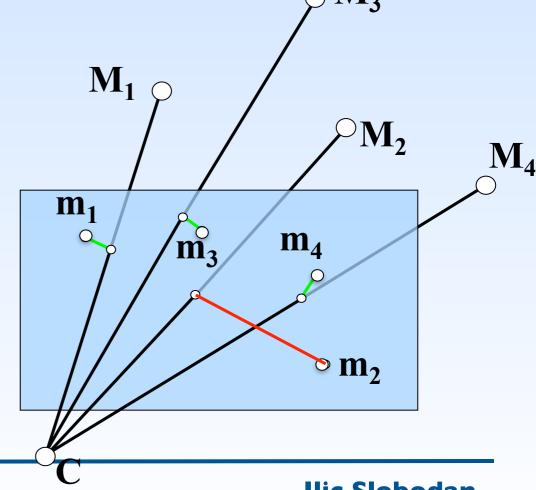
Bayesian interpretation:

$$\underset{\mathbf{p}}{\operatorname{arg min}} \sum_{i} \left\| \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})} (\mathbf{M}_{i}) - \mathbf{m}_{i} \right\|^{2}$$

$$= \underset{\mathbf{p}}{\operatorname{arg max}} \prod_{i} N \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})} (\mathbf{M}_{i}); \mathbf{m}_{i}, \sigma \mathbf{I} \right)$$

The error on the 2D point locations  $\mathbf{m}_i$  is assumed to have a Gaussian (Normal) distribution with identical covariance matrices σI, and independent;

This assumption is violated when  $\mathbf{m}_i$  is an outlier.



#### (the 2 equivalent formulations)

$$\begin{split} & \min_{\mathbf{p}} \sum_{i} \left\| \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right\|^{2} \\ &= \min_{\mathbf{p}} \sum_{i} \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right)^{\mathsf{T}} \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right) \\ &= \min_{\mathbf{p}} \sum_{i} \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right)^{\mathsf{T}} \left( \frac{1/\sigma}{0} \frac{0}{0} \frac{0}{1/\sigma} \right) \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right) \\ &= \max_{\mathbf{p}} \sum_{i} - \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right)^{\mathsf{T}} \sum_{\mathbf{m}} \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right) \\ &= \max_{\mathbf{p}} \prod_{i} \frac{1}{\sqrt{(2\pi)^{2} |\Sigma_{\mathbf{m}}|}} \exp \left( - \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right)^{\mathsf{T}} \sum_{\mathbf{m}} \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right) \right) \\ &= \max_{\mathbf{p}} \prod_{i} N \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}); \mathbf{m}_{i}, \Sigma_{\mathbf{m}} \right) \end{split}$$

### Robust estimation

#### Idea:

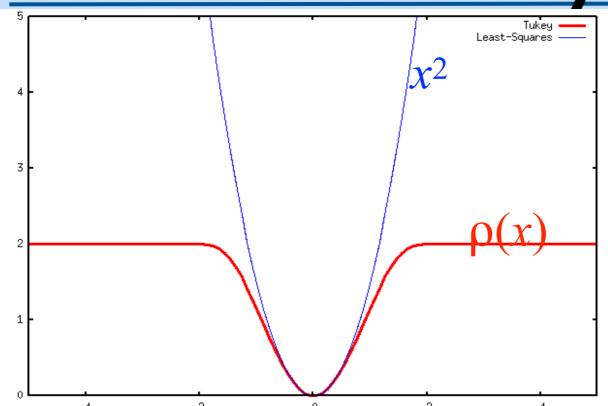
Replace the Normal distribution by a more suitable distribution,

or equivalently replace the least-squares estimator by a "robust estimator":

$$\min_{\mathbf{p}} \sum_{i} \left\| \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right\|^{2} = \sum_{i} r_{i}^{2}$$

$$\Rightarrow \min_{\mathbf{p}} \sum_{i} \rho \left( \operatorname{Proj}_{\mathbf{A}, \mathbf{R}(\mathbf{p}), \mathbf{T}(\mathbf{p})}(\mathbf{M}_{i}) - \mathbf{m}_{i} \right) = \sum_{i} \rho(r_{i})$$

### Example of an M-estimator: The Tukey Estimator

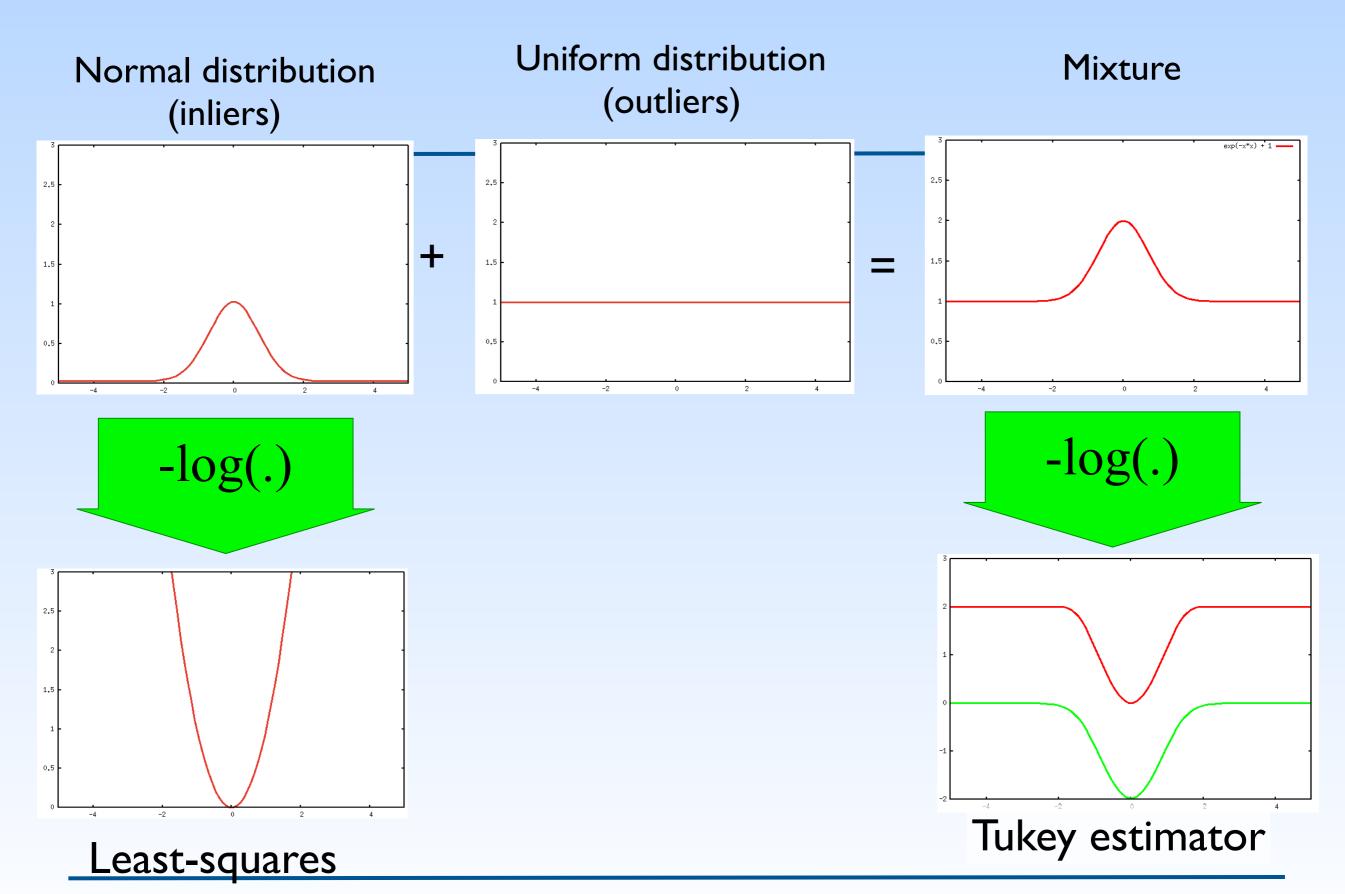


$$\begin{cases} |f|x| \le c \quad \rho(x) = \frac{c^2}{6} \left( 1 - \left[ 1 - \left( \frac{x}{c} \right)^2 \right]^3 \right) \\ |f|x| > c \qquad \rho(x) = \frac{c^2}{6} \end{cases}$$

The Tukey estimator assumes the measures follow a distribution that is a mixture of:

- a Normal distribution, for the inliers,
- a uniform distribution, for the outliers.

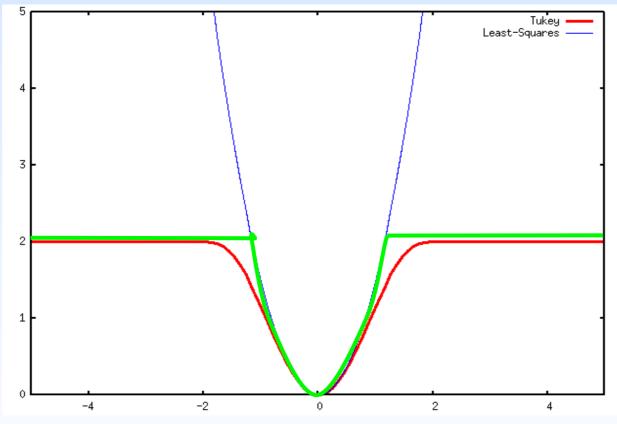
The threshold c is usually taken to be proportional to the measured standard deviation of the residual errors for inlier data.



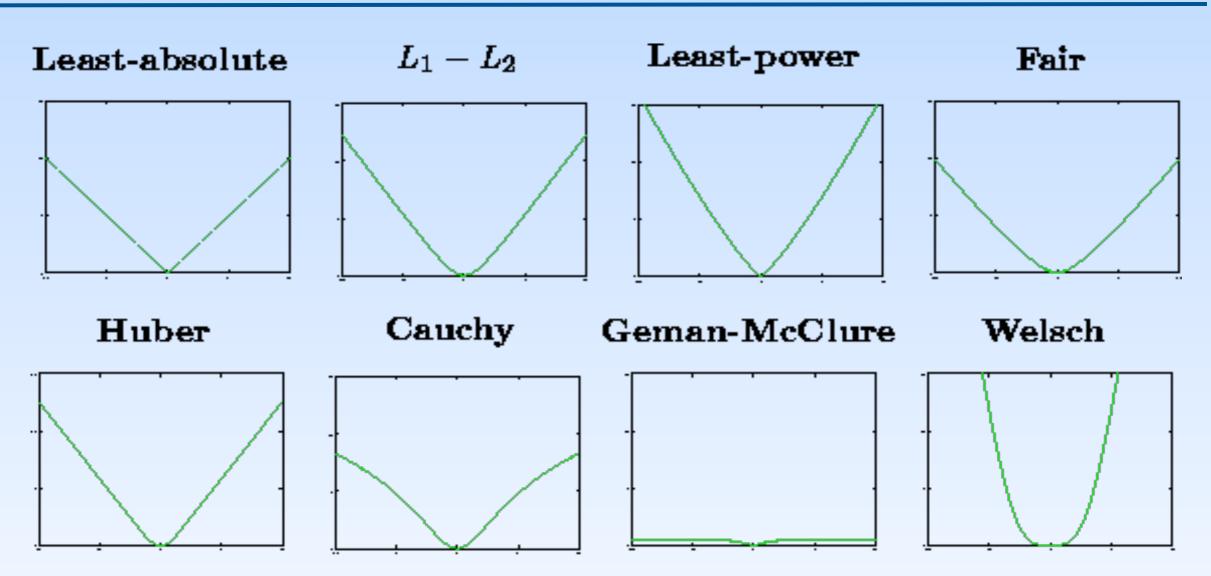
#### The Tukey Estimator in Levenberg-Marquardt Optimization

Use the following approximation:

$$\begin{cases} \text{if } |x| \le \tilde{c} & \rho(x) = \frac{1}{2}x^2 \text{ (least - squares)} \\ \text{if } |x| > \tilde{c} & \rho(x) = \frac{1}{2}\tilde{c}^2 \text{ (constante)} \end{cases}$$



### Other M-Estimators



# Use of robust estimator with GN or LM minimisation

The Gauss-Newton and Levenberg-Marquardt algorithms can still be applied to minimize the sum of residual errors  $E(\theta) = \sum r_i^2$  after the introduction of M-estimators  $E(\theta) = \sum \rho(r_i)$ , even if the M-estimators can be complex functions. We solve this by finding derivative of the objective function in aspect to parameters:

$$\frac{\partial E}{\partial \theta} = \sum \rho'(r_i)r_i\theta_i$$

This is simply done by weighting the residuals i at each iteration step: Each i is replaced by therefore the weight should be chosen as:

$$w_i = \frac{\rho(r_i)}{r_i}$$

In the case of the LevenbergMarquardt algorithm,  $\Delta$ i can be computed as be changed as:

$$\mathbf{\Delta_i} = -(\mathbf{J}^T \mathbf{W} \mathbf{J} + \lambda \mathbf{I})^{-1} \mathbf{J}^T \mathbf{W} \epsilon_i$$

where weight matrix is  $\mathbf{W} = diag(\dots w_i \dots)$ 

#### Scale of the residuals

- M-estimator (Tukey and Huber) constant **c** has been chosen assuming that measurements have normal distribution (standard deviation of I and zero mean) and therefore they provide asymptotic efficiency of 95% of linear regression. **c=4.685** Tukey and **c=1.345** (Huber)
- However, measurements with outliers are not normally (Gaussian) distributed, so the residuals must be scaled, i.e. every  $\rho(r_i)$  should be replaced with  $\rho(r_i/s)$  where **s** is estimated scale parameter.
- The simplest estimation of **s** is done using median absolut deviation of the residuals:

$$MAD = median\{|r_i|\}$$

where  $\hat{s} = MAD/0.6745$ , which is based on the idea expectation of MAD being E(MAD) = 0.6745 for normal distribution.

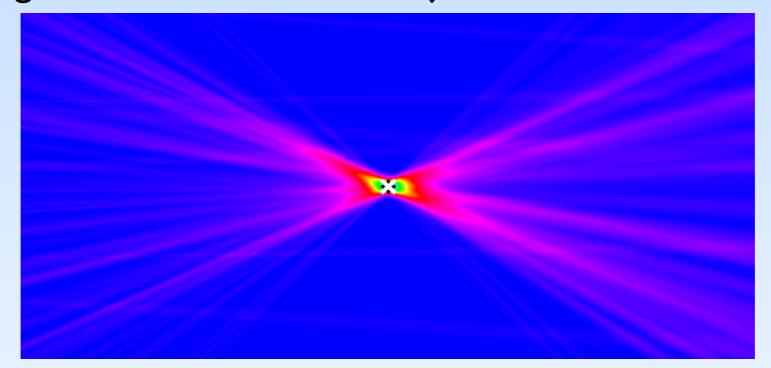
### Drawbacks of the Tukey Estimator

- Non-convex -> creates local minimas;
- Function becomes flat when too far from the global minimum.

### Gaussian Noise on the Projections + 20% outliers + Tukey estimator

White cross: true camera position;

Black cross: global minimum of the object function.



The global minimum is very close to the true camera pose.

#### **BUT**:

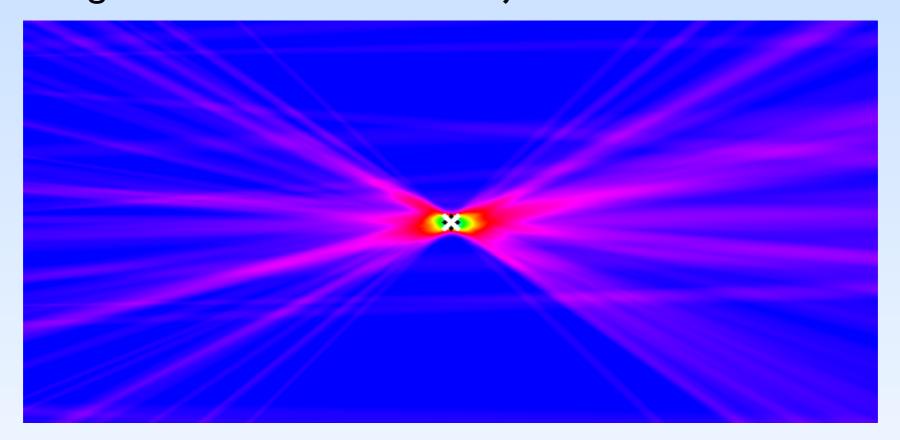
- local minima;
- the objective function is flat where all the correspondences are considered outliers.

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### Gaussian Noise on the Projections + 50% outliers + Tukey estimator

White cross: true camera position;

Black cross: global minimum of the object function.



Even more local minimums.

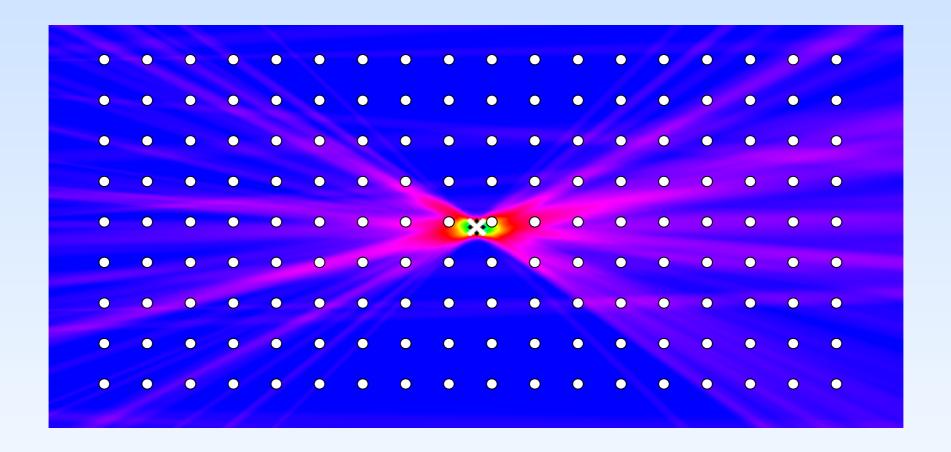
Numerical optimization can get trapped into a local minimum.

Non-liner optimization and robust estimation for tracking

Ilic Slobodan

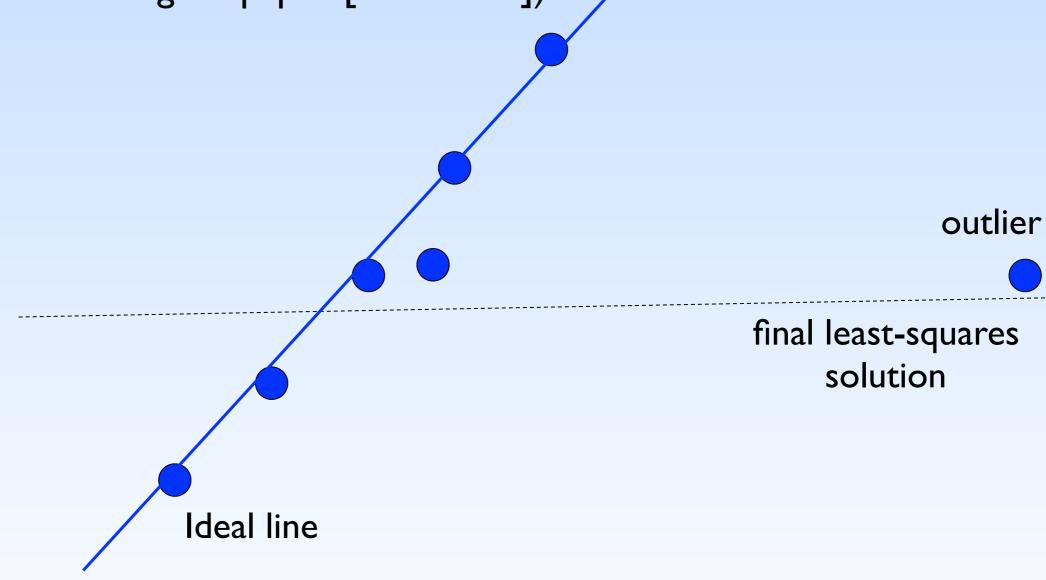
### RANSAC

Idea: sampling the space of solutions (the camera pose space here):



### RANSAC RANdom SAmple Consensus

Line fitting: the "Throwing Out the worst residual" heuristics can fail (Example for the original paper [Fischler81]):



### RANSAC

As before, we could do a regular sampling, but would not be optimal: Ideal line

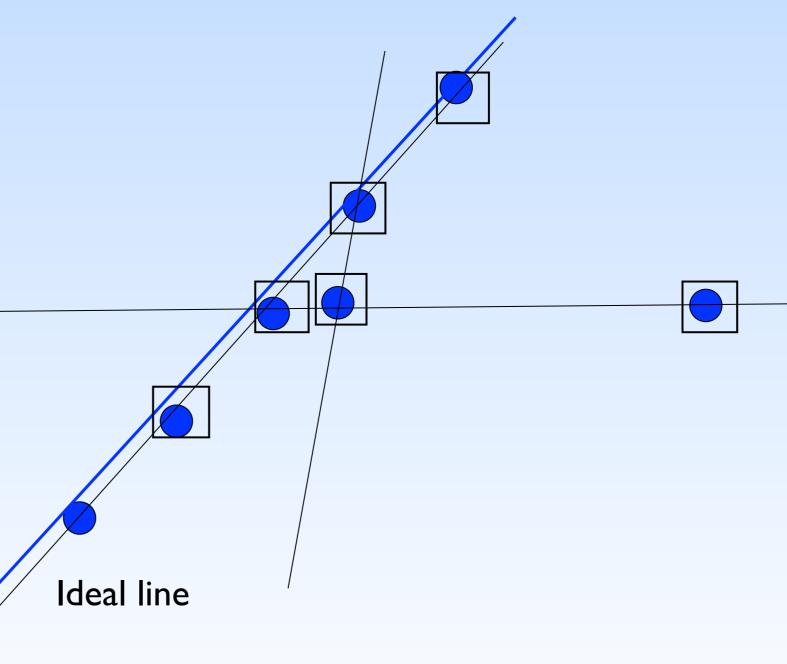
#### RANSAC

#### Idea:

•Generate hypotheses from subsets of the measurements.

• If a subset contains no gross errors, the estimated parameters (the hypothesis) are closed to the true ones.

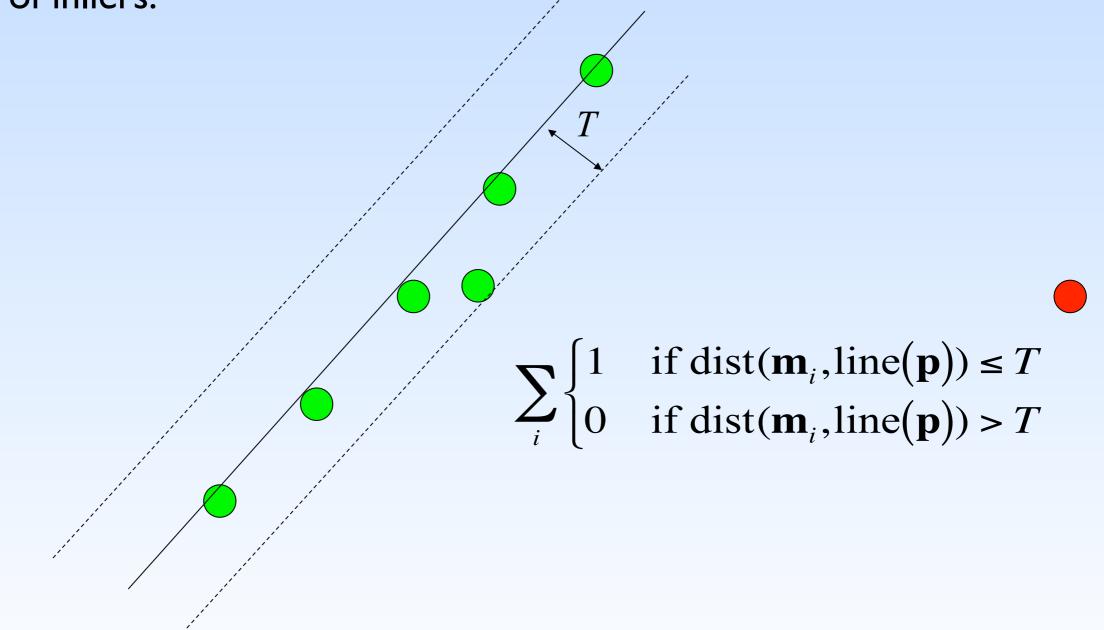
•Take several subsets at random, retain the best one.



The quality of a hypothesis is evaluated by the number of measures that lie "close enough" to the predicted line.

We need to choose a threshold (T) to decide if the measure is "close enough".

RANSAC returns the best hypothesis, i.e the hypothesis with the largest number of inliers.



e = probability that a point is an outlier

s = number of points in a sample

N = number of samples (we want to compute this)

p = desired probability that we get a good sample

#### Solve the following for N:

$$1 - (1 - (1 - e)^{s})^{N} = p$$

#### Where in the world did that come from? ....

From Robert Colins, Penn State University

e = probability that a point is an outlier

s = number of points in a sample

N = number of samples (we want to compute this)

p = desired probability that we get a good sample

$$1 - (1 - (1 - e)^{s})^{N} = p$$

Probability that choosing one point yields an inlier

e = probability that a point is an outlier

s = number of points in a sample

N = number of samples (we want to compute this)

p = desired probability that we get a good sample

$$1 - (1 - (1 - e)^{s})^{N} = p$$

Probability of choosing s inliers in a row (sample only contains inliers)

e = probability that a point is an outlier

s = number of points in a sample

N = number of samples (we want to compute this)

p = desired probability that we get a good sample

$$1 - (1 - (1 - e)^{s})^{N} = p$$

Probability that one or more points in the sample were outliers (sample is contaminated).

e = probability that a point is an outlier

s = number of points in a sample

N = number of samples (we want to compute this)

p = desired probability that we get a good sample

$$1 - (1 - (1 - e)^{s})^{N} = p$$

Probability that N samples were contaminated.

e = probability that a point is an outlier

s = number of points in a sample

N = number of samples (we want to compute this)

p = desired probability that we get a good sample

$$1 - (1 - (1 - e)^{s})^{N} = p$$

Probability that at least one sample was not contaminated (at least one sample of s points is composed of only inliers).

### How many samples?

Choose N so that, with probability p, at least one random sample is free from outliers. e.g. p=0.99

$$(1 - (1 - e)^s)^N = 1 - p$$

$$N = \frac{\log(1 - p)}{\log(1 - (1 - e)^s)}$$

	proportion of outliers $e$						
S	5%	10%	20%	25%	30%	40%	50%
2	2	3	5	6	7	11	17
3	3	4	7	9	11	19	35
4	3	5	9	13	17	34	72
5	4	6	12	17	26	57	146
6	4	7	16	24	37	97	293
7	4	8	20	33	54	163	588
8	5	9	26	44	78	272	1177

### Pose Estimation

To apply RANSAC to pose estimation, we need a way to compute a camera pose from a subset of measurements, for example a P3P algorithm.

Since RANSAC only provides a solution estimated with a limited number of data, it must be followed by a robust minimization to refine the solution.