

Ray Tracing One Weekend

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1 Ray Tracing in One Weekend

- this is my take on Marc Andreessen's anti-todo¹ list concept that i have been doing for years no matter the nature of the project or type of work i am doing (i also maintain a bite-sized version in a series of tiny moleskin)
- test bib [1]
- bib working lets go
- now that bib is working, thanks to [Peter Shirley](#), [Trevor David Black](#), and [Steve Hollasch](#) for this incredible writeup! [1] ([course link](#))
- shoutout to [@ludwigABAP](#) for poasting this course (shoutout ml btw (for you page))
- this is my very first c or cpp project (op says it's c flavored cpp) beyond hello worlds and basic basic robotics stuff
- i love rust but i am not cracked at all so i would probably not be able to follow along in rust
- i will, however, follow op's advice to not copy pasta (besides most of makefile compiler flags hehehe) and build it up slowly by typing along
- going to try my best to thug this out by Sunday
- important setup for fresh arch install (not in order, and just off the dome, i likely am forgetting tons of things)
 - install unzip (will need for nvim clangd Mason lsp stuff)
 - install cmake, clangd, gcc stuff
 - setup debugger for nvim using dap, dap-ui, etc.
 - **build, compile, run:** (i think lol)
 1. `cmake -B build/Debug -DCMAKE_BUILD_TYPE=Debug`
 2. `cmake --build build/Debug`
 3. `build/rayTracing > image.ppm`
- op claims that if we can build project correctly in the beginning, then we are golden for rest of tutorial
- the provided cmake file is cash money and really easy to get working with my setup (Figure 1)

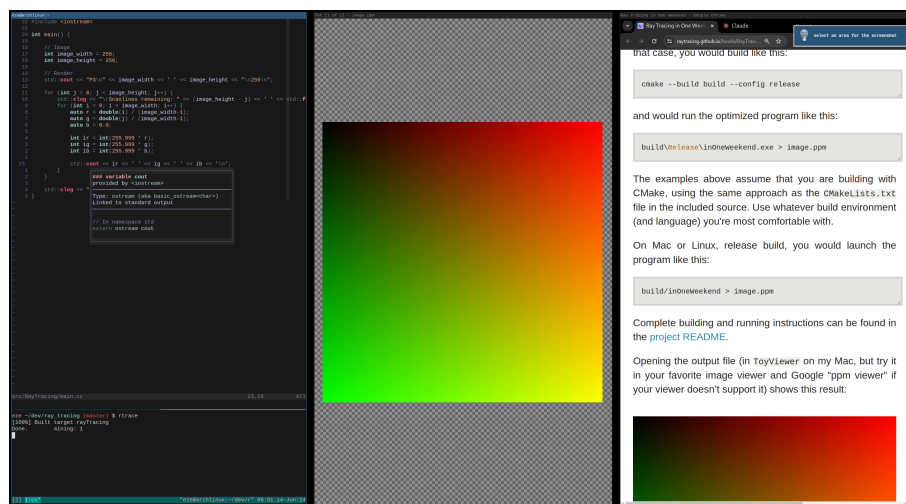


Figure 1: *build test*

- make `rtrace` alias for build, compile, run, then open image in feh, probably terrible idea but whatever
- got color header file with a `write_color` util function
- now working on a `ray` class which will use our `vec3` class.
- refresher on rays: think of them as functions (Equation 1):

$$\mathbf{P}(t) = \mathbf{A} + t\mathbf{b} \quad (1)$$

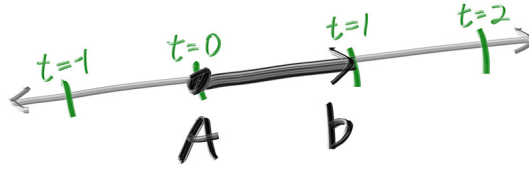


Figure 2: *linear interpolation*

- here \mathbf{P} is a 3D position along a line in 3D. \mathbf{A} is the ray origin and \mathbf{b} is the ray direction. The ray parameter t is a real number (`double` in the code). Plug in a different t and $\mathbf{P}(t)$ moves the point along the ray. Add in negative t values and you can go anywhere on the 3D line. For positive t , you get only the parts in front of \mathbf{A} , and this is what is often called a half-line or a ray. (Figure 2)
- to make the actual ray tracer we will make simple camera with 16:9 aspect ratio since it will be easier to debug x and y transpositions.
- we need the height to be at least 1 since we divide width by height since it's easier to set the aspect ratio to width then divide by height. e.g. $width/height = 16/9 = 1.7778$
- apparently this is just an *optimistic* (my words) ratio since these values are not integers. we approximate the aspect ratio as best we can by rounding height to the nearest integer (and don't allow it to be less than one)

¹thanks pmarcal [original blog post](#) (archived by someone)

- now we have a camera center in 3d space from which all the rays will originate (commonly referred to as the *eye point*). we initially set the distance between the viewport of the camera center point to be one unit. This is often referred to as the *focal length*.
- we will use *right-handed coordinates* (right hand rule gang) Figure 3

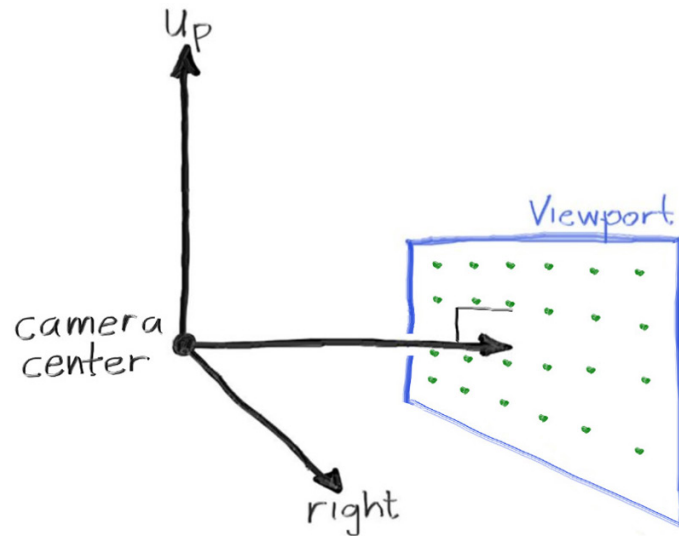


Figure 3: camera geometry

- unfortunately, the camera pose conflicts with the way we would like to render our image starting from the upper left pixel row by row scanning across from left to right. (Figure 4)

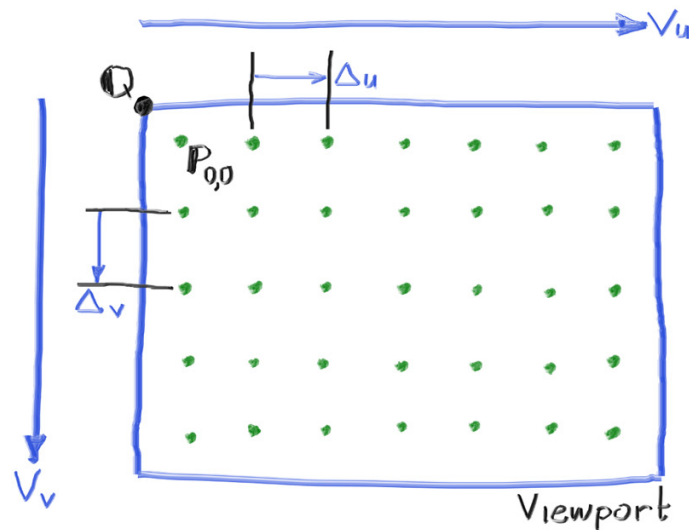


Figure 4: viewport and pixel grid

- we have example 7x5 resolution image, the viewport upper left corner Q , the pixel $P_{0,0}$ location, the viewport vector V_u (viewport_u), the viewport vector V_v (viewport_v), and the pixel delta vectors Δu and Δv .
- now we add a simple gradient to the `ray_color` function which will linearly blend white and blue depending on the height of y coordinate *after* scaling the ray direction to unit length. op uses standard graphics trick to linearly scale $0.0 \leq a \leq 1.0$. when $a = 1.0$, i want blue, when $a = 0.0$, i want white. in between, i want a blend. here's a *linear interpolation* or *linear interpolation*; commonly referred to as a *lerp* between two values. a lerp is always of the form

$$\text{blendedValue} = (1 - a) \cdot \text{startValue} + a \cdot \text{endValue} \quad (2)$$

with a going from zero to one (ayy lmao).

- now it's time to add a sphere (folks use spheres in ray tracers because calculating whether a ray hits a sphere is relatively simple.

- the equation for a radius r that is centered at the origin is an important mathematical equation

$$x^2 + y^2 + z^2 \quad (3)$$

- this of this as saying that if a given point (x, y, z) is on the surface of the sphere, then $x^2 + y^2 + z^2 = r^2$. if a given point (x, y, z) is *inside* the sphere, then $x^2 + y^2 + z^2 < r^2$, and if a given point (x, y, z) is *outside* the sphere, then $x^2 + y^2 + z^2 > r^2$.
- if we want to allow the sphere center to be at an arbitrary point (C_x, C_y, C_z) , then the equation is not so nice:

$$(C_x - x)^2 + (C_y - y)^2 + (C_z - z)^2 = r^2 \quad (4)$$

- in graphics, you almost always want formulas to be in terms of vectors so we don't need to write out so many terms.
- note that the vector from point $\mathbf{P} = (x, y, z)$ to center $\mathbf{C} = (C_x, C_y, C_z)$ is $(\mathbf{C} - \mathbf{P})$
- we can use the definition of the dot product:

$$(\mathbf{C} - \mathbf{P}) \cdot (\mathbf{C} - \mathbf{P}) = (C_x - x)^2 + (C_y - y)^2 + (C_z - z)^2 = r^2 \quad (5)$$

- then we can rewrite the equation of the sphere in vector form as:

$$(\mathbf{C} - \mathbf{P}) \cdot (\mathbf{C} - \mathbf{P}) = r^2 \quad (6)$$

- we can read this as “any point \mathbf{P} that satisfies this equation is on the sphere”. we want to know if our ray $\mathbf{P}(t) = \mathbf{Q} + t\mathbf{d}$ ever hits the sphere anywhere. if it does hit the sphere, there is some t for which $\mathbf{P}(t)$ satisfies the sphere equation. So we are looking for any t where this is true:

$$(\mathbf{C} - \mathbf{P}(t)) \cdot (\mathbf{C} - \mathbf{P}(t)) = r^2 \quad (7)$$

- which can be found by replacing $\mathbf{P}(t)$ with its expanded form:

$$(\mathbf{C} - (\mathbf{Q} + t\mathbf{d})) \cdot (\mathbf{C} - (\mathbf{Q} + t\mathbf{d})) = r^2 \quad (8)$$

- we have three vecs on the left dotted by three vecs on right, if we solved for the full dot product we would get nine vectors (slop), we need to solve for t with quadratic equation, so first shape equation above into quadratic by first isolating t terms, then distribute dot product, then move r^2 to left hand side:

- boom

$$(-t\mathbf{d} + (\mathbf{C} - \mathbf{Q})) \cdot (-t\mathbf{d} + (\mathbf{C} - \mathbf{Q})) = r^2 \quad (9)$$

- bop

$$t^2\mathbf{d} \cdot \mathbf{d} - 2t\mathbf{d} \cdot (\mathbf{C} - \mathbf{Q}) + (\mathbf{C} - \mathbf{Q}) \cdot (\mathbf{C} - \mathbf{Q}) = r^2 \quad (10)$$

- pow

$$t^2\mathbf{d} \cdot \mathbf{d} - 2t\mathbf{d} \cdot (\mathbf{C} - \mathbf{Q}) + (\mathbf{C} - \mathbf{Q}) \cdot (\mathbf{C} - \mathbf{Q}) - r^2 = 0 \quad (11)$$

- we can now run quadratic formula on this bad boy and all the vecs are reduced to scalars by dot prod.

$$\frac{-b \pm \sqrt{b^2 - 4ac}}{2a} \quad (12)$$

- can now for t using the terms a, b , and c :

$$a = \mathbf{d} \cdot \mathbf{d} \quad (13)$$

$$b = -2\mathbf{d} \cdot (\mathbf{C} - \mathbf{Q}) \quad (14)$$

$$c = (\mathbf{C} - \mathbf{Q}) \cdot (\mathbf{C} - \mathbf{Q}) - r^2 \quad (15)$$

- this will yield either positive (2 real solutions), negative (no real solutions) or zero (1 real solution)

- the algebra almost always relates very directly to the geometry (Figure 5)

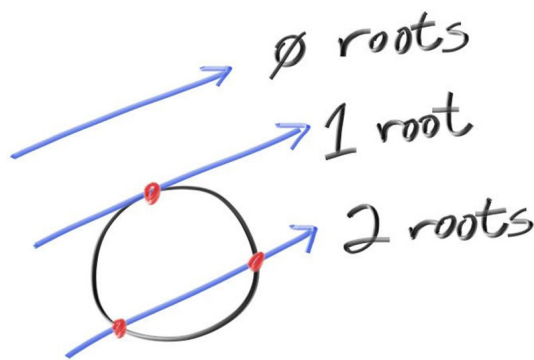


Figure 5: ray sphere intersection roots

- create first ray traced image by placing a small sphere at -1 on the z-axis and then coloring red any pixel that intersects it. (Figure 6)

```

czecharchlinux:~
13 #include "color.h"
12 #include "ray.h"
11 #include "vec3.h"
10
9 #include <iostream>
8
7 bool hit_sphere(const point3& center, double radius, const ray& r) {
6     vec3 oc = center - r.origin();
5     auto a = dot(r.direction(), r.direction());
4     auto b = -2.0 * dot(r.direction(), oc);
3     auto c = dot(oc, oc) - radius*radius;
2     auto discriminant = b*b - 4*a*c;
1     return (discriminant >= 0);
14 }

1
2 color ray_color(const ray& r) {
3     if (hit_sphere(point3(0,0,-1), 0.5, r))
4         return color(1, 0, 0);
5
6     vec3 unit_direction = unit_vector(r.direction());
7     auto a = 0.5*(unit_direction.y() + 1.0);
8     return (1.0-a)*color(1.0, 1.0, 1.0) + a*color(0.5, 0.7, 1.0);
9 }
10
11 int main() {
main.cc
-- VISUAL LINE --
[0] 1:notes- 2:nwtr 3:rtrace*Z

feh [1 of 1] - Image.ppm

```

Figure 6: simple red sphere

- this lacks all sorts like shading, reflection rays, and more than a single object, also this solution doesn't account for the camera really since it will also work with the sphere center at +1 (behind camera)

- i should start writing section titles... maybe later
- shading with surface normals, a normal is just perpendicular to the surface at the point of intersection.
- we have key design decision to make for normal vectors in code: whether normal vecs will have arbitrary length, or should we normalize? *much to think about...*
- it is tempting to skip the expensive sqrt op involved in normalizing the vector, in case it's not needed. in practice, however, there are three important observations. first, if a unit-length normal vector is *ever* required, then you might as well do it up from once, instead of over and over again "just in case" for every location where unit-length is required. Second, we *do* require unit-length normal vecors in several places. Third, if you require normal vectors to be unit length, then you can often efficiently generate that vec with an understanding of the specific geometry class, in its constructor, or in the `hit()` function. e.g., sphere normals can be made unit length simply by dividing by the sphere radius, avoiding the sqrt entirely.
- unit length sphere normals will be used up front for these reasons
- the outward normal for a sphere is in the direction of the hit point minus the center (Figure 7)

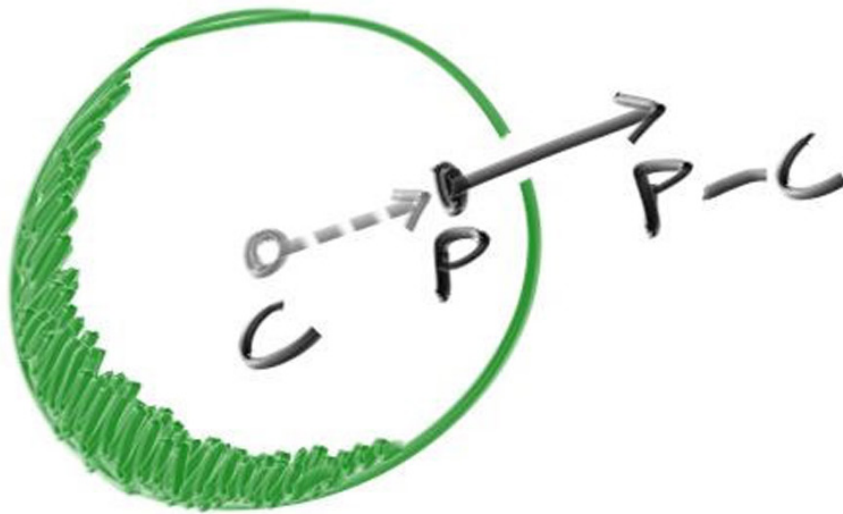


Figure 7: *sphere surface-normal geometry*

- we don't have light yet so let's use a common trick for visualizing normals: a color map. we can assume \mathbf{n} is a unit length vec, so each component is between -1 and 1; we just map each component to the interval from 0 to 1, and then map (x, y, z) to $(red, green, blue)$. for the normal, we need the hit point, not just whether we hit or not (which is all we are doing currently: Figure 6). we only have 1 sphere in the scene and it's right in front of the camera, so we won't worry about negative values of t yet. We'll just assume the closest hit point (smallest t) is the one that we want. These changes in the code let us compute and visualize \mathbf{n} .

References

- [1] S. H. Peter Shirley, Trevor David Black, “Ray tracing in one weekend,” April 2024.