

# TiCkS: A Flexible White-Rabbit Based Time-Stamping Board



C. Champion<sup>1</sup>, M. Punch<sup>1,2</sup>, R. Oger<sup>1</sup>, S. Colonges<sup>1</sup>, for the CTA Consortium<sup>4</sup>

& Y. Moudden<sup>3</sup>

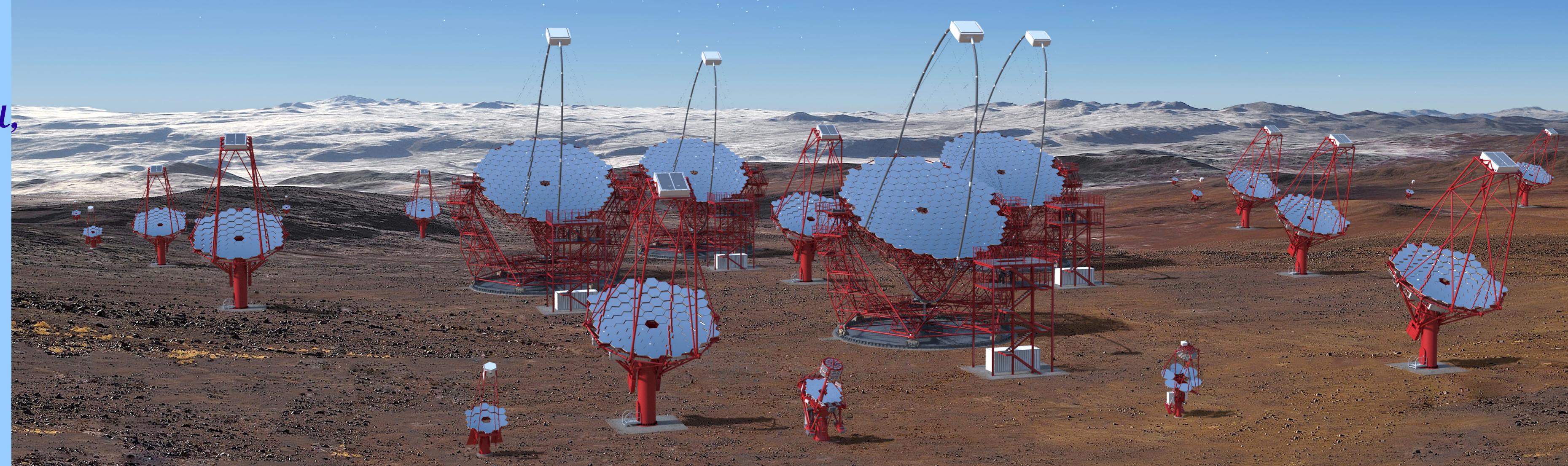
<sup>1</sup>APC, Univ Paris Diderot, CNRS/IN2P3, CEA/Irfu,

Obs de Paris, Sorbonne Paris Cité, France,

<sup>2</sup>Linnaeus University, Växjö, Sweden

<sup>3</sup>CEA – Cadarache / DRF /IRFM/STEP/GEAC

<sup>4</sup>See [www.cta-observatory.org](http://www.cta-observatory.org)



## ABSTRACT

The TiCkS board (Time and Clock Stamping) is based on the White Rabbit (WR) SPEC node (Simple PCIe FMC carrier), to provide ns-precision time-stamps (TSs) of input signals and transmission of these TSs to a central collection point, developed as one of the candidate CTA TS nodes, with a small form-factor for use in any CTA camera.

The main firmware changes in the Spartan-6 FPGA are: the addition of: 1) a ns-precision TDC for the TSs; and 2) a UDP stack to send TSs and auxiliary information over the WR fibre, and to receive configuration & slow control commands over the same fibre. It also provides a PPS (Pulse Per Second) and other clock signals to the connected device, from which it can receive auxiliary event-type information over an SPI link (Serial Peripheral Interface). A version of TiCkS will be made available in the WR Open Hardware repository, providing a cheap, flexible, and reliable solution for ns-precision time-stamping of trigger signals up to 400 kHz, for use in other experiments.

## CTA, Cherenkov Telescope Array

CTA [1] will be a gamma-ray observatory in the very-high-energy range (VHE, > 30 GeV), with >120 imaging atmospheric Cherenkov telescopes distributed over two sites (La Palma, Spain & Atacama desert, Chile).

CTA's Software Array Trigger (SWAT) will detect coincident Cherenkov light-flashes (<100 ns window) from showers of particles induced in the atmosphere by Gamma-rays & Cosmic rays.

The SWAT needs accurate relative Time-Stamps (TSs) from each telescope's Camera Trigger Management electronics (CTM), which it can correct in software for the telescope pointing

→ Coincidence identification with a flexible logic

## White Rabbit Technology for CTA

White Rabbit (WR) [2], adopted for time-stamping since it can achieve CTA's 2 ns rms relative accuracy requirement *WR for CTA*:

- Distribute time from a central clock system to WR "nodes" in each telescope camera,
- Sent over hierarchical network of WR-compatible switches at array control centre
- WR-nodes time-stamp trigger signals from CTMs
  - both for "read-out" events and "busy" triggers
- Event & PPS counters in both camera's CTM and its WR-node for combining event data with TSs
- WR network itself may be used to collect TSs from all telescopes at a central point for SWAT
- Coincidence information sent to each camera's data-processing pipeline
  - → allows to drop non-coincident event data

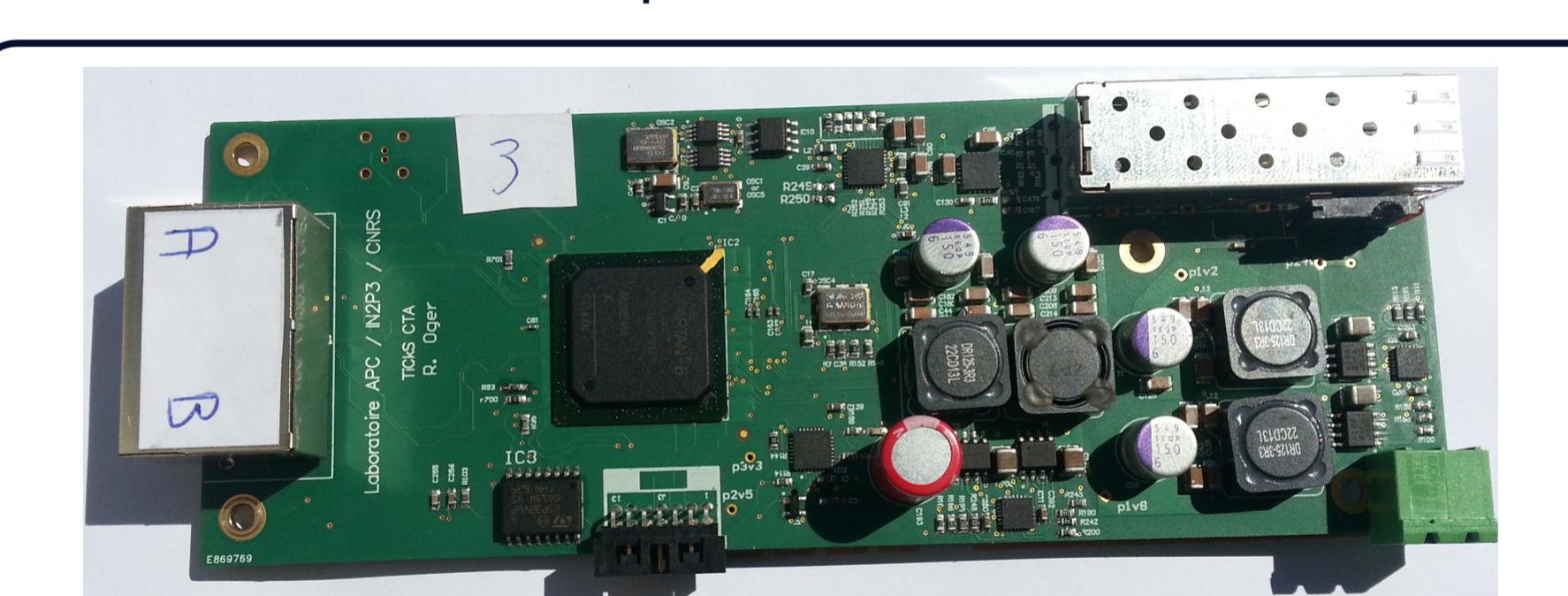
## TiCkS, CTA Time-Stamp node candidate

### Hardware

TiCkS board (Time and Clock Stamping), based on WR-SPEC node [3] & WR-core [4]

Operation based on:

- "fine-TDC", 1ns:
  - Based on 8-bit I/O-SerDes @ 1 GHz (from PLL on WR-125 MHz clock)
  - As shift register for input trigger signal
  - Trigger signal in register detected on WR-clock tick,
    - → flag + #ns stored
- "Coarse-TDC", 8 ns:
  - Counter of the WR-clock, Zeroed by WR-PPS
  - Read-out if trigger signal detected
  - PPS and Event counters
  - SPI link (50 MHz) from OpenCores [5]
    - Allows some bits (16) from CTM → TiCkS
    - FIFO from WR-core
      - To store/send event information
      - But "fix" of 2 FIFOs in ping-pong (by MUX), to avoid read-out dead-time
  - UDP Stack from OpenCores [6]



TiCkS board with 2xRJ45 CTA-Connector

### Data & Software

- Event data sent in "bunches"
  - ≤20 events (with bunch time-out)
- Close-packed into 12-byte words (LSB parts)
- "Tailer", 20 bytes (full LSB+MSB information)
- Bunch size ≤ 260 byte
  - (≤ 302 bytes with overhead)
- → Close to optimal for UDP network packets
- Average event size of 13 bytes
  - 75% gain over unpacked events, given 42-byte overhead
- Decoding/unpacking library in "c"

Firmware main characteristics:

- Spartan 6 xc6slx45t-3fgg484 version ISE 14.7
- Using 6735 slice registers (12%)
- 1/8 high-speed I/O buffers (BUFPLL)
- 4/4 Phase Locked Loops (PLL\_Advs)

Direction	Signal Name	Comment
Camera → TiCkS		
	Read-out Trigger	Required
	Busy Trigger	Optional (for cameras with dead.time)
	SPI clock, data, chip-select	Optional
TiCkS → Camera		
	PPS	Synchronized via WR
	x MHz clock	Aligned with PPS, x = 10 MHz here
	External Trigger	At defined TAI time

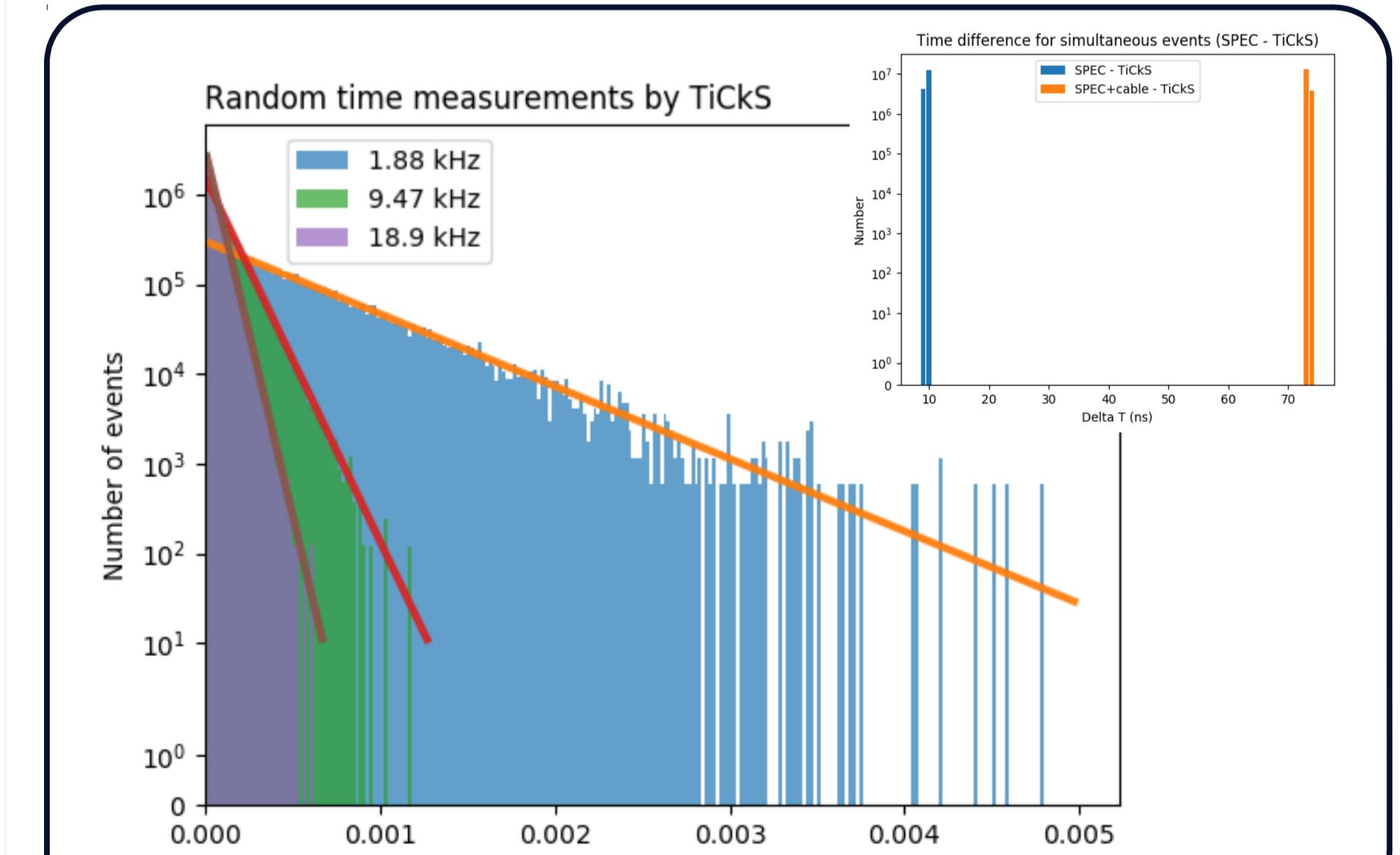
### Signals exchanged between Camera Trigger Management (CTM) and TiCkS

## Inter-board Time-Stamps

Tested to 400kHz (fixed freq.), 56kHz (random)

Distribution always within 2 adjacent ns bins

Addition of cable delay (63.2ns) → shift as expected

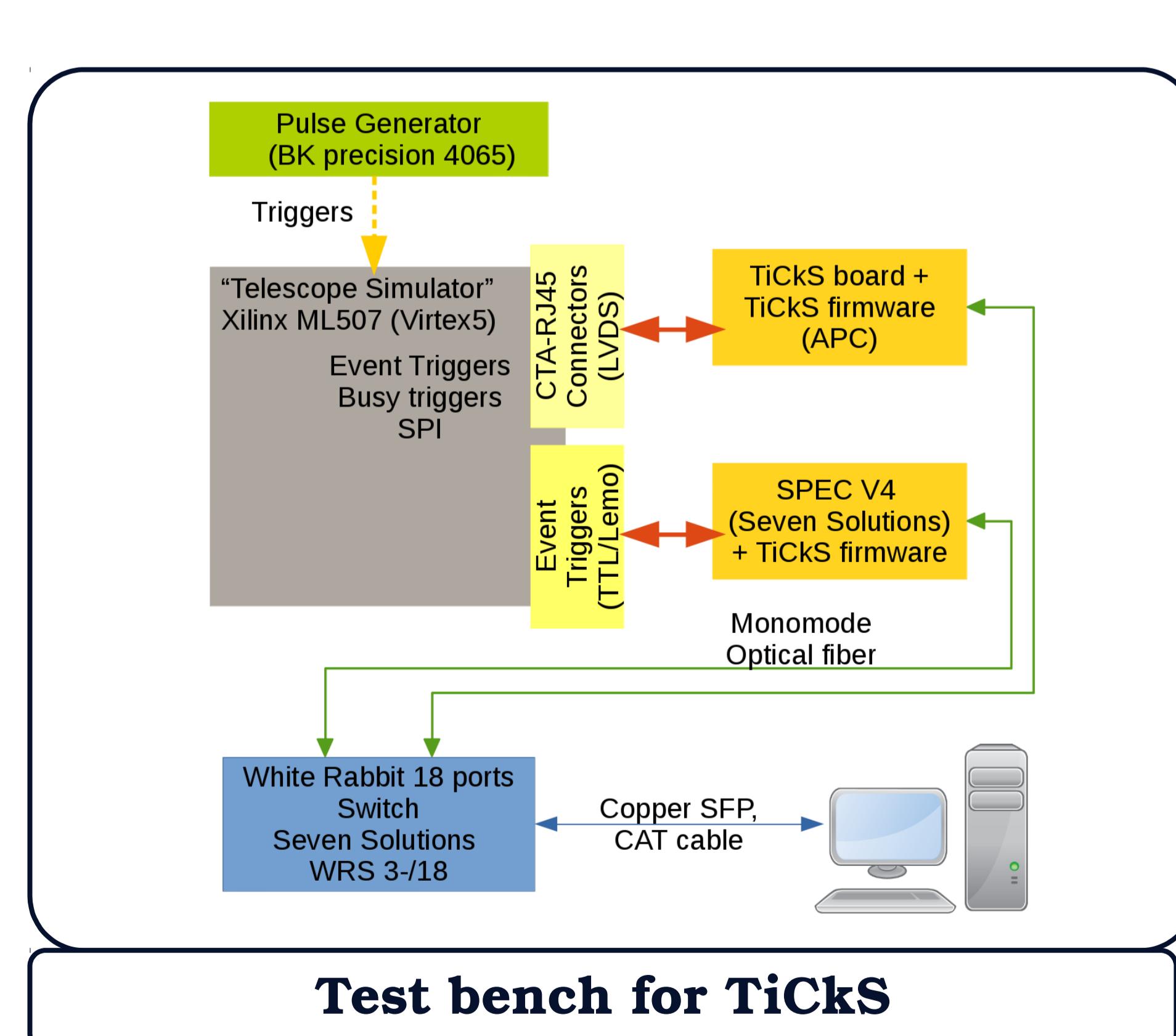


Inter-board trigger distributions for random Poisson inputs (inset: TiCkS intra-board distributions, with/without cable delay)

## Conclusions

TiCkS hardware and firmware built and tested:

- well-capable of time-stamping trigger signals
  - Tested for fixed & random time distribution
  - Little or no loss at CTA rates
- Soon to be made available in OpenHardware



Test bench for TiCkS

## References

- [1] CTA, [www.cta-observatory.org](http://www.cta-observatory.org)
- [2] White Rabbit Open Hardware, [www.ohwr.org/projects/white-rabbit](http://www.ohwr.org/projects/white-rabbit)
- [3] SPEC WR Open Hardware project, [www.ohwr.org/projects/spec](http://www.ohwr.org/projects/spec)
- [4] WR Core v4.0, [www.ohwr.org/projects/wr-cores/Wiki/Wrpc\\_release-v40](http://www.ohwr.org/projects/wr-cores/Wiki/Wrpc_release-v40)
- [5] SPI Master/Slave Interface, [www.opencores.org/project/spi\\_master\\_slave](http://www.opencores.org/project/spi_master_slave)
- [6] 1G eth UDP / IP Stack, [https://www.opencores.org/project/udp\\_ip\\_stack](http://www.opencores.org/project/udp_ip_stack)