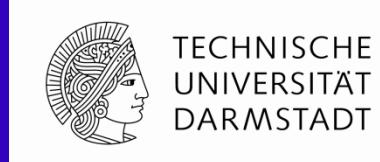
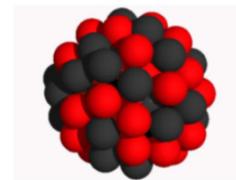
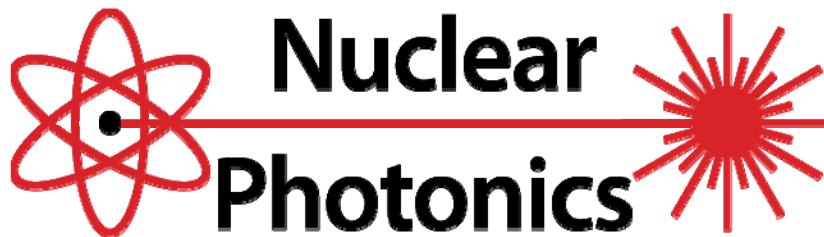


Research opportunities for Nuclear Photonics with an ERL-based hard X-ray source



Norbert Pietralla, TU Darmstadt

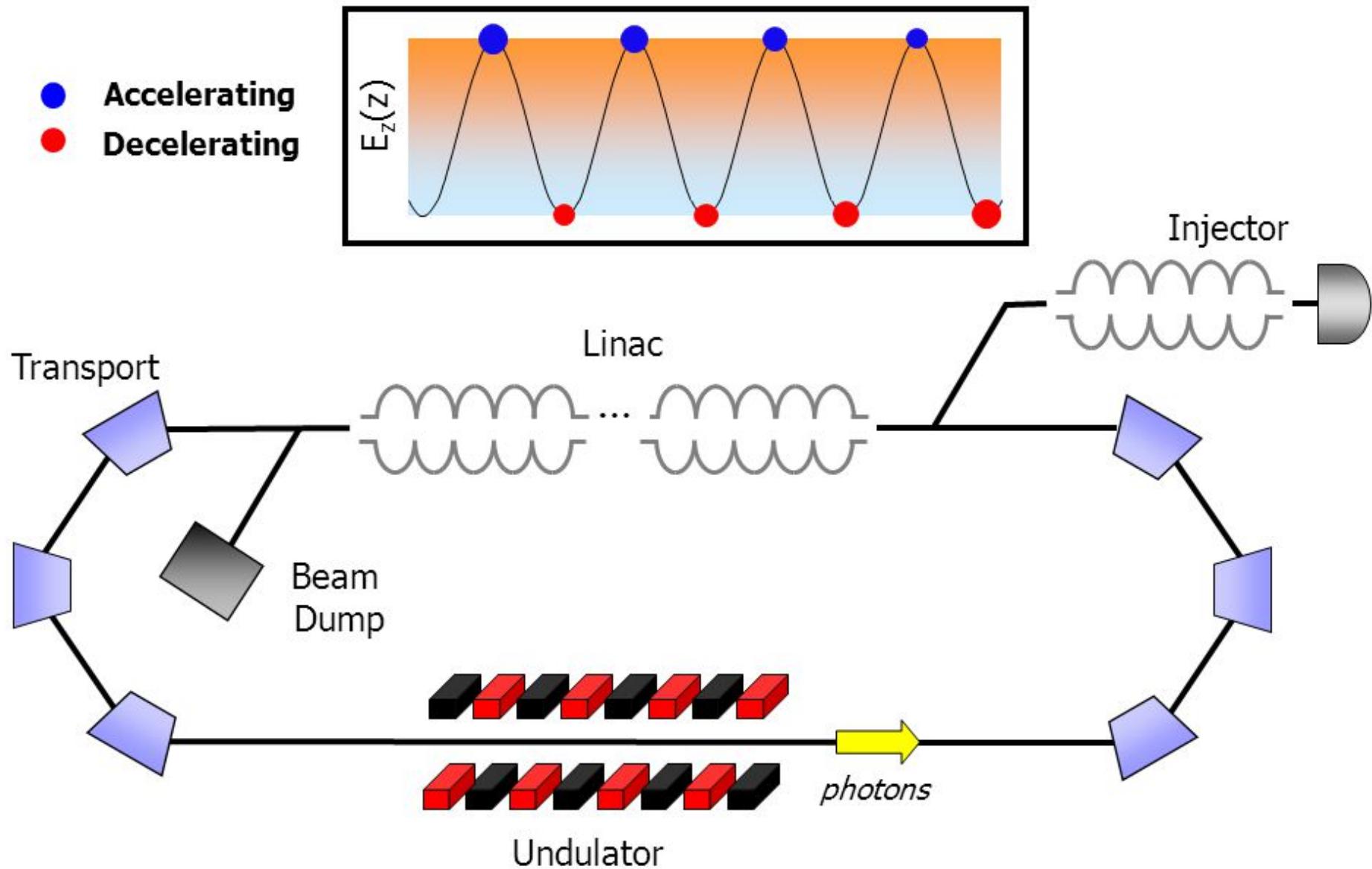


Photonuclear reactions

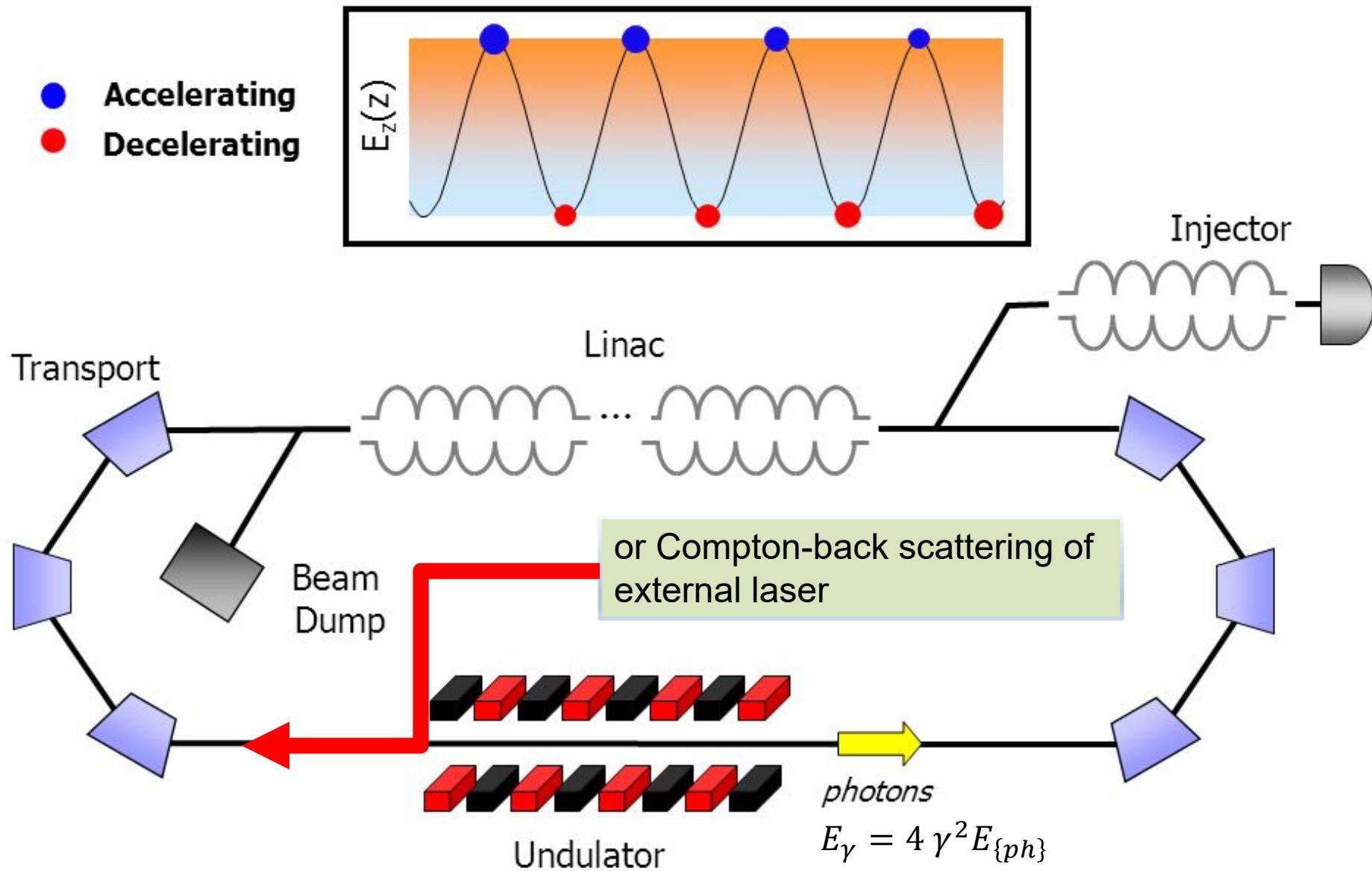


Conference photo
Nuclear Photonics 2018
Brasov, Romania

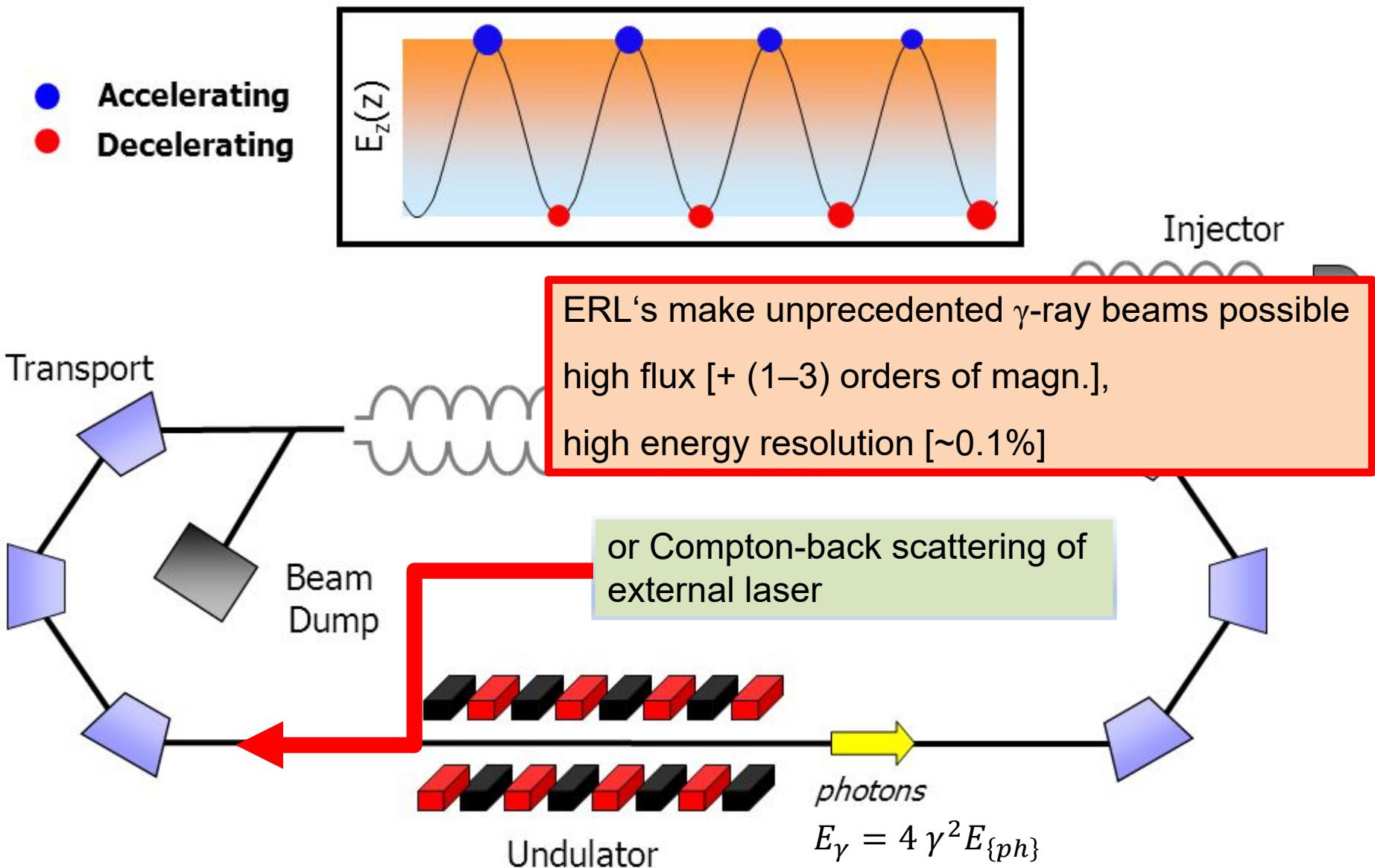
Generic ERL-based Light Source



Generic ERL-based Light Source



Generic ERL-based Light Source



Outline



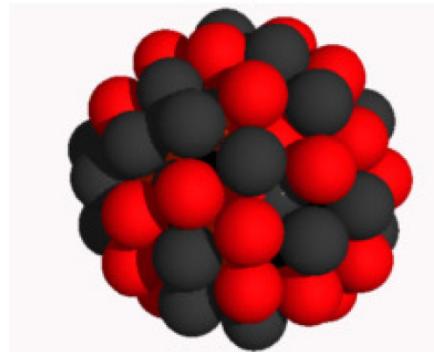
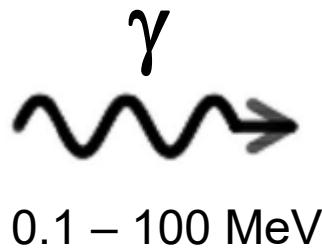
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- Photonuclear reactions with MeV-range γ rays
- Methodological developments for nuclear photonics
- Motivation from astrophysics and particle physics
- Examples for recent achievements
- Concluding remarks

Photonuclear Reactions



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What happens?

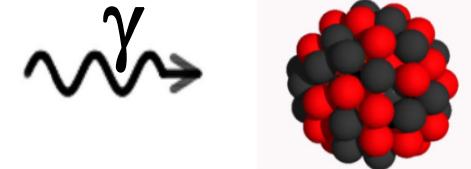
Nuclear Physics with photon beams



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Pure EM-interaction

- (nuclear-)model independent
- “small“ cross sections, intense beams



Minimum projectile mass

- min. angular momentum transfer,
spin-selective: low-spin modes [E1,M1,E2,(E3?)]

Polarisation

- “Parity Physics“, channel selectivity

Narrow Bandwidth (at 4th generation LCB source)

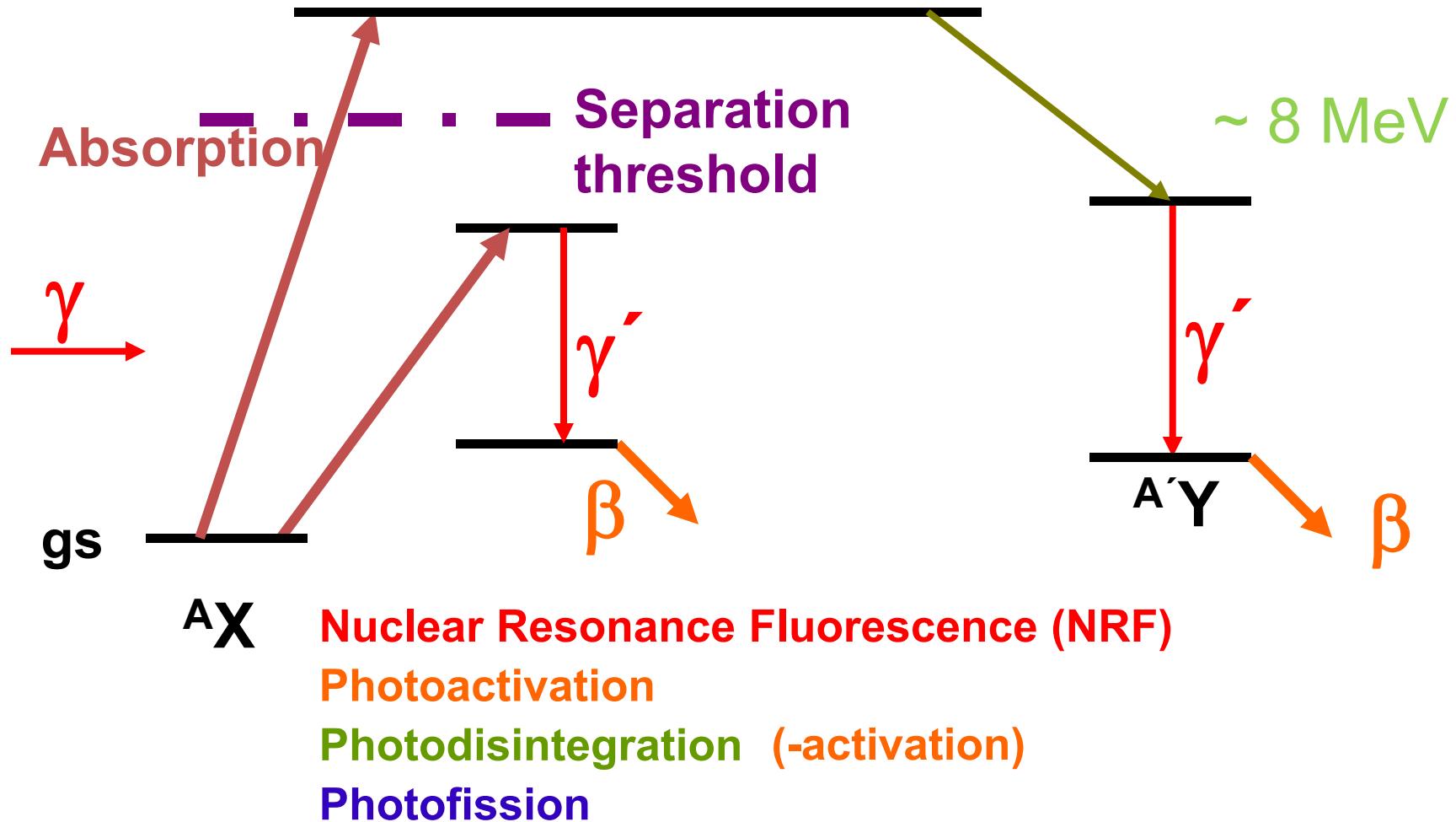
- Explore specific excitation energy
“Selective Manipulation of Nuclear States“:



Photonuclear Reactions



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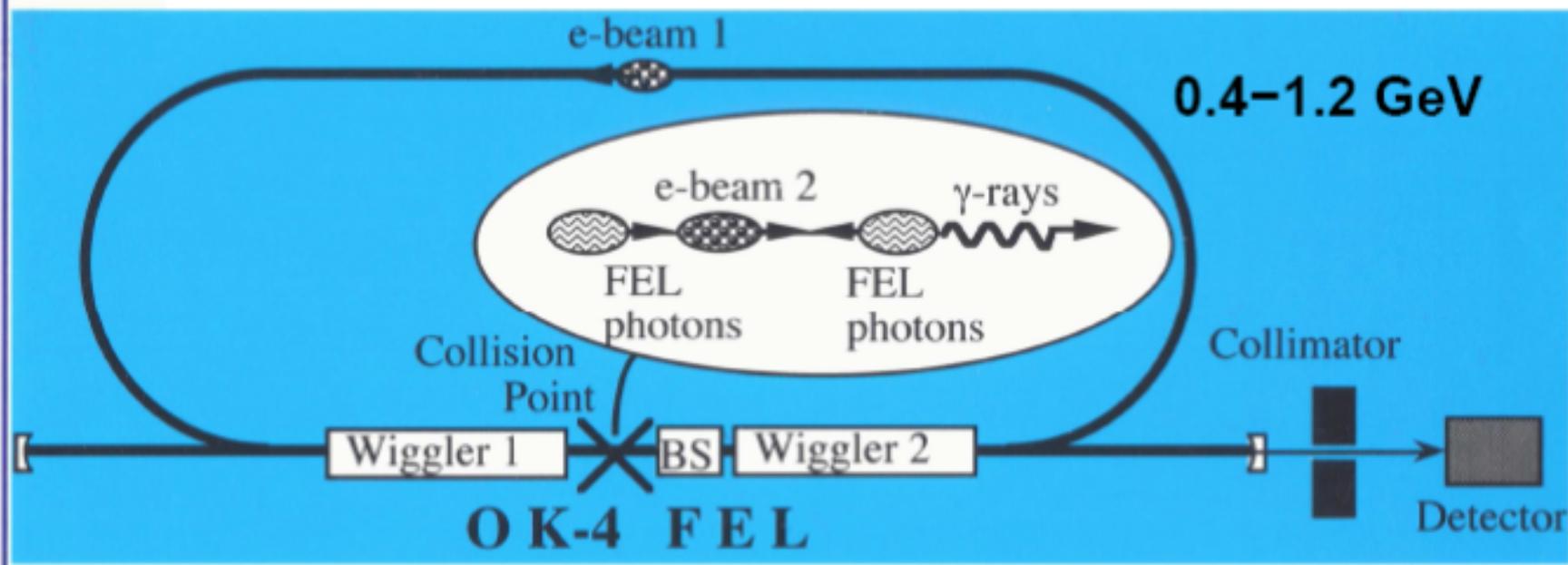


High Intensity γ -Ray Source (HIGS)



H.R.Weller, V.N.Litvinenko
Duke University, Durham, NC, U.S.A.

Compton Backscattering of Intra-cavity Laser Light



2 – 60 MeV

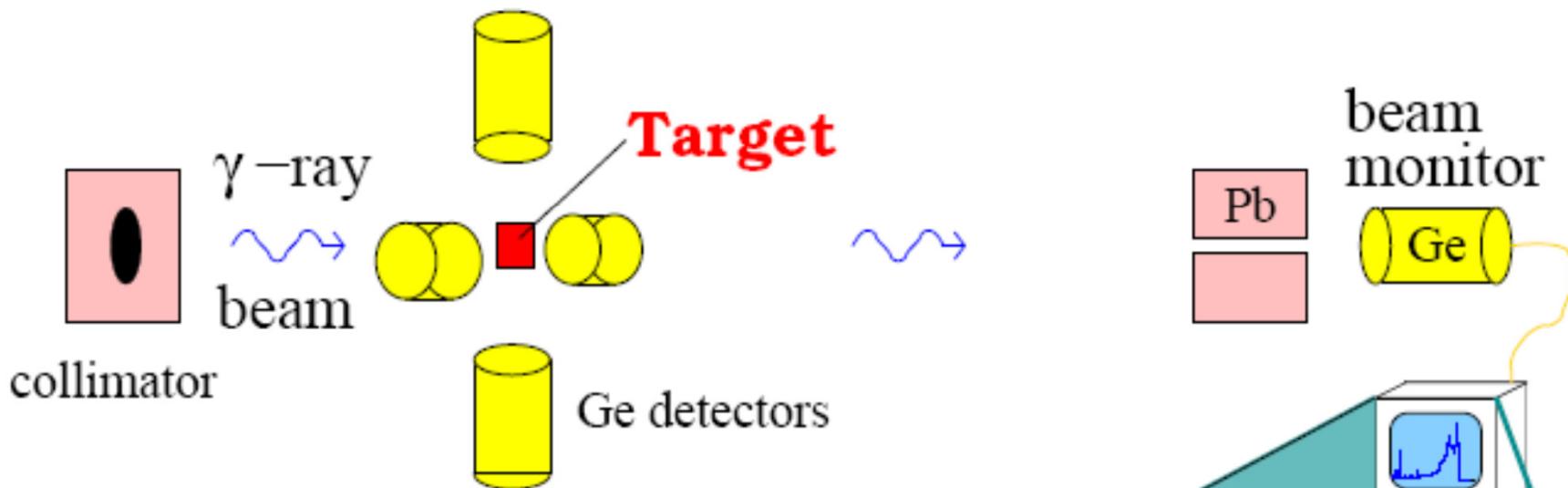
$$E_\gamma = \frac{4\gamma^2 E_{ph}}{(1 + r + \gamma^2 \theta^2)}; \quad r = \frac{4\gamma E_{ph}}{mc^2}; \quad E_{ph} = \frac{2\gamma^2 hc}{\lambda_w (1 + K_w^2 / 2)}; \quad \gamma = \frac{E_e}{mc^2};$$

1.7 – 6.4 eV

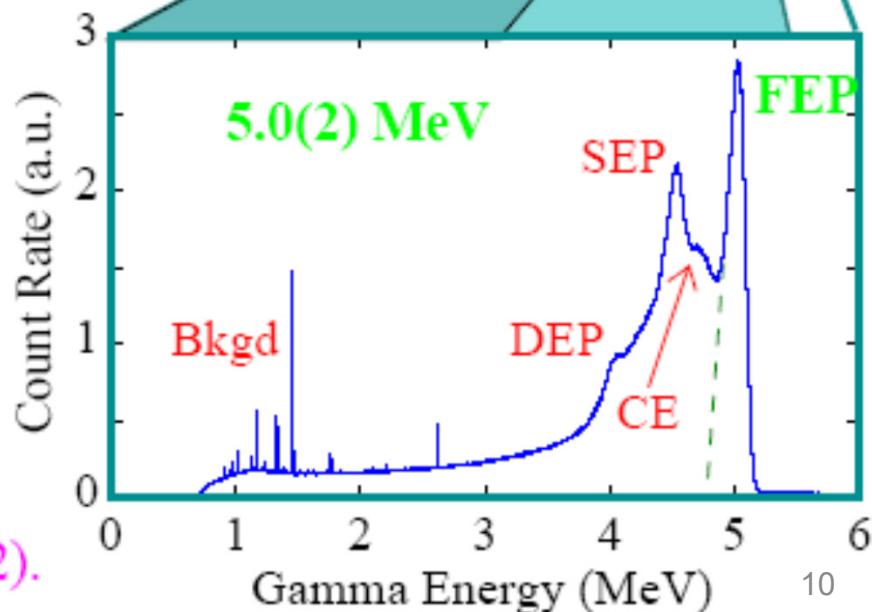
~ 1000

nearly monochromatic, tunable, completely polarized

Looking at the HIGS Gamma-Ray Beam



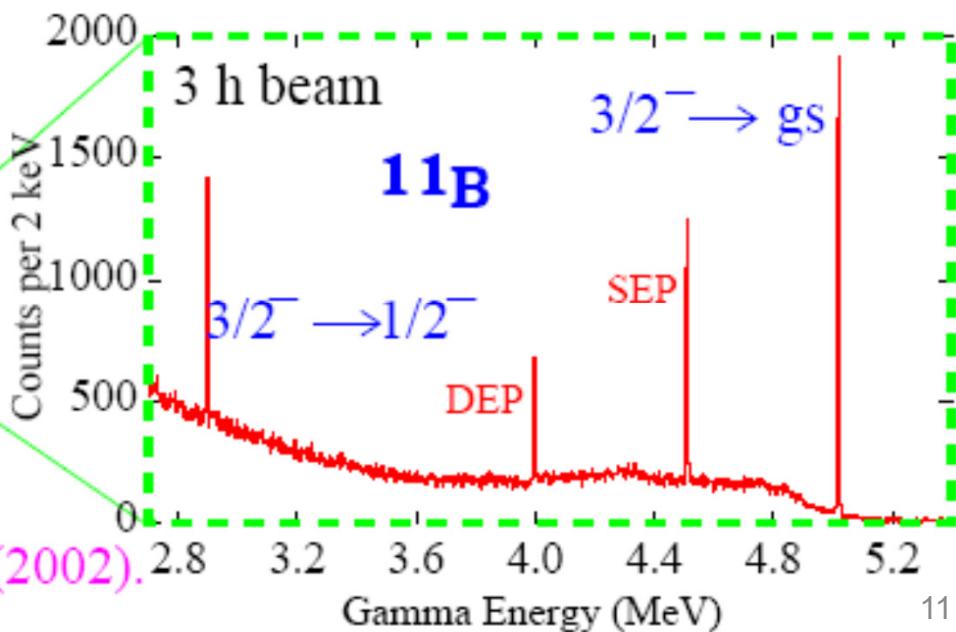
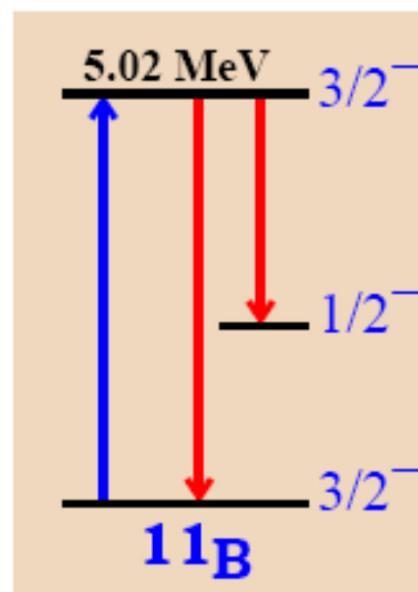
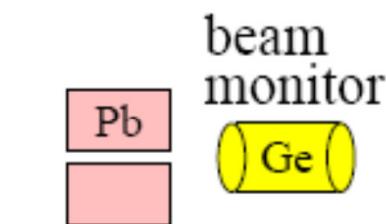
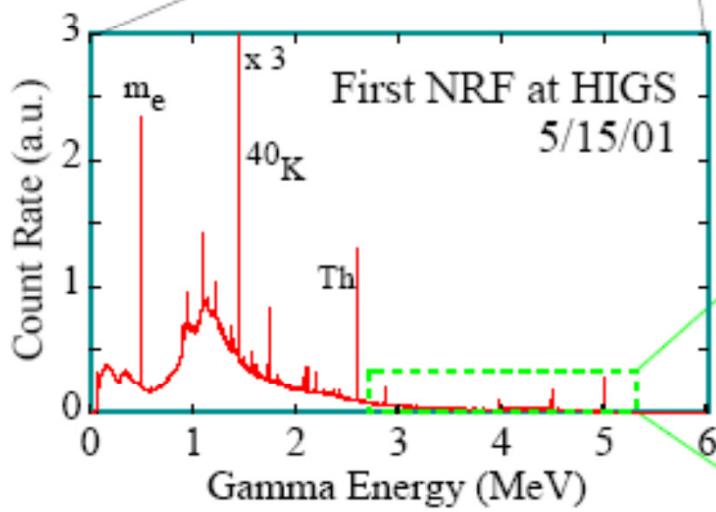
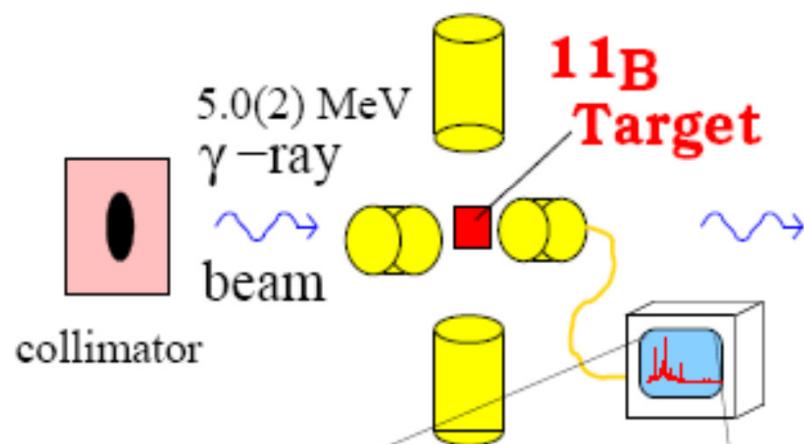
Flux at target: $10^7/\text{s}$
at maximum: $10^5/(\text{s keV})$
Resolution: 3%
(with 1" collimator)



N.Pietralla et al.

Nucl.Instrum.Methods A 483, 556 (2002).

Looking at the Target

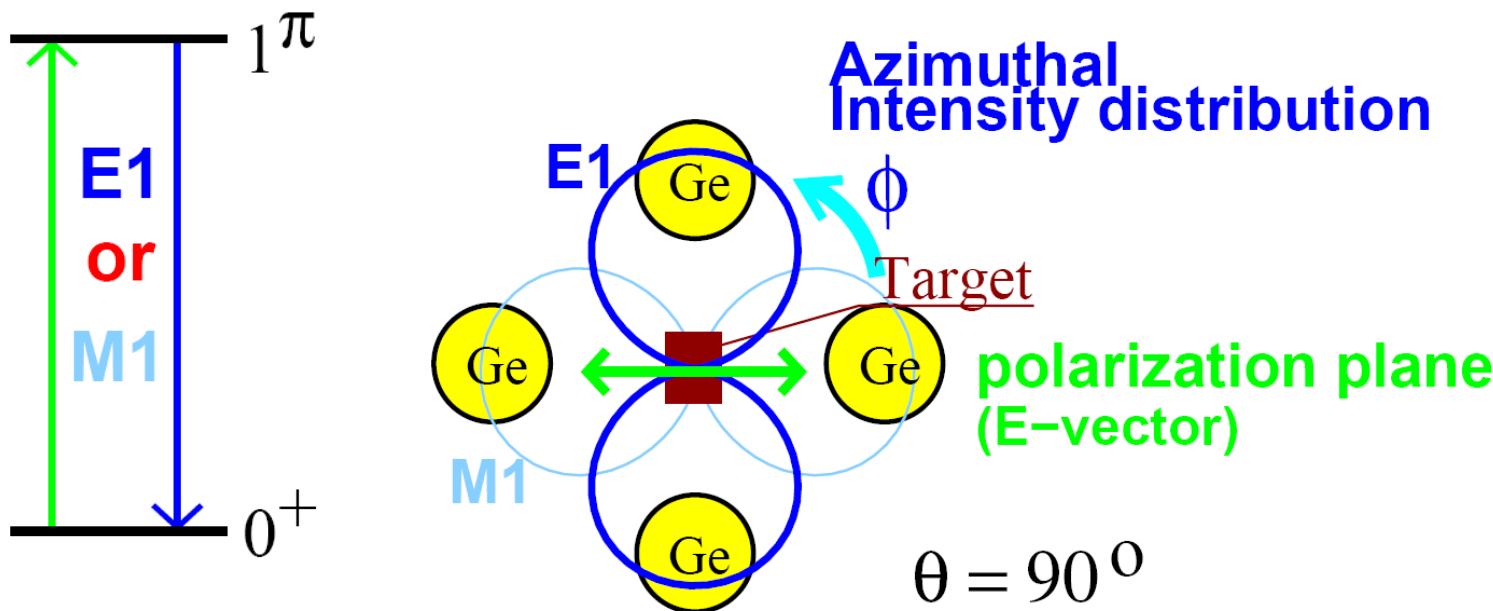
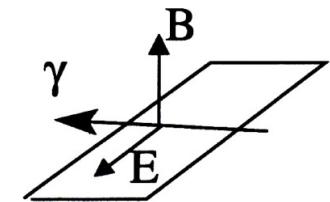
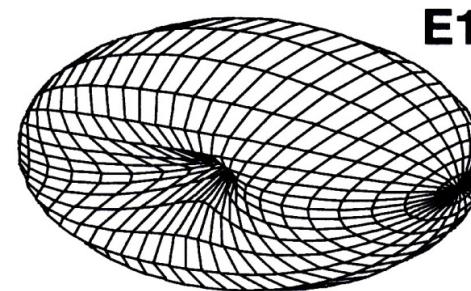


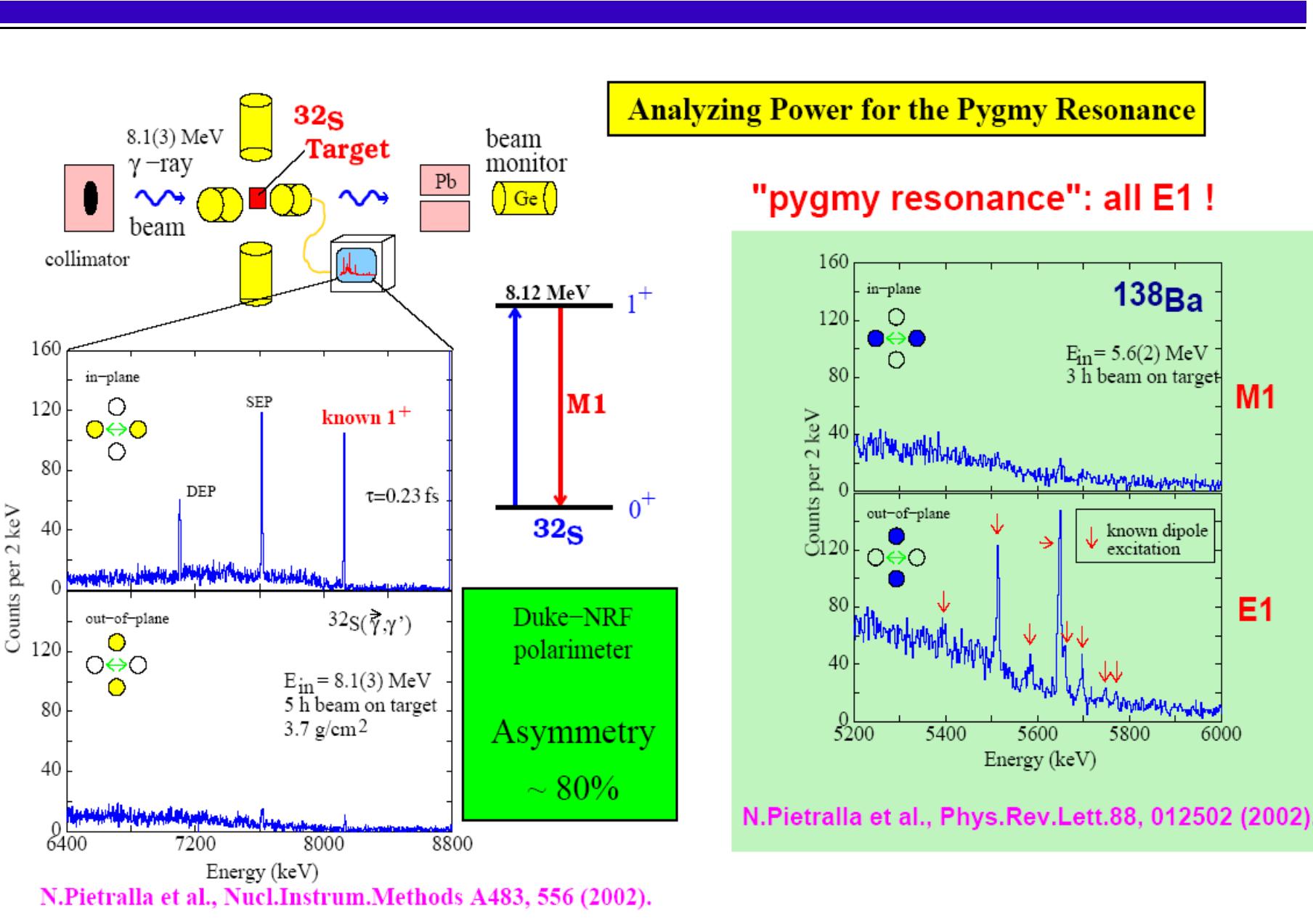
N.Pietralla et al.
Nucl.Instrum.Methods A483, 556 (2002).

Polarimetry with linearly polarized γ -ray beams



Scattering intensity distribution for linearly polarized photons



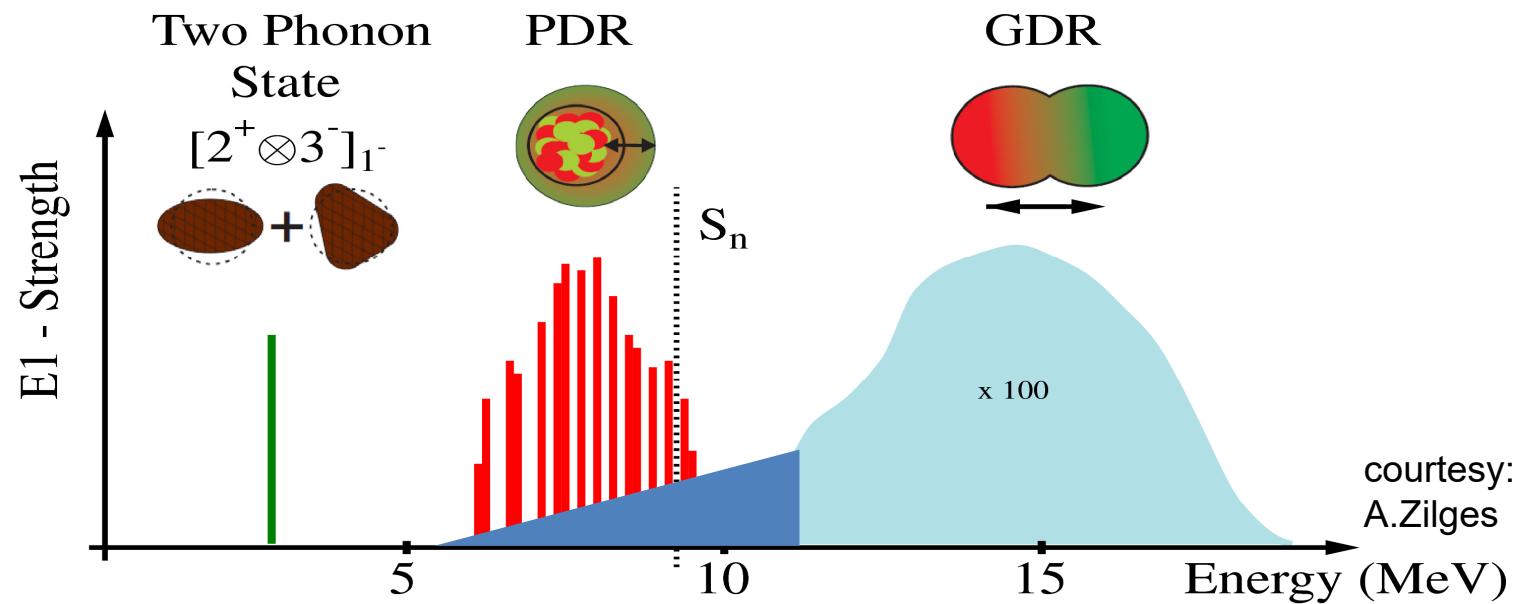
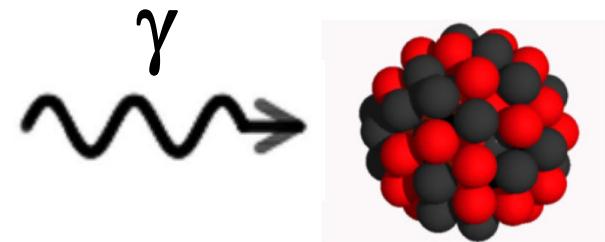


Electromagnetic dipole-response



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Typical E1 strength distribution in heavy atomic nuclei

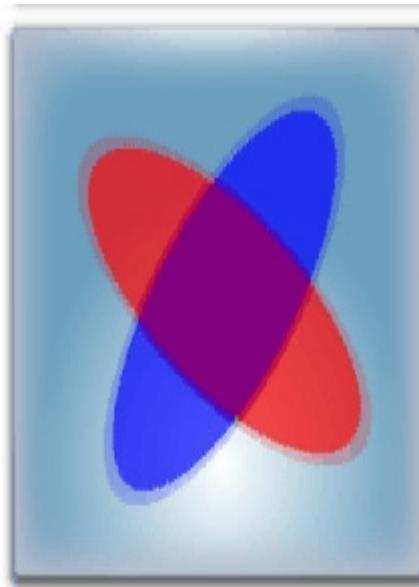


Discovery of Scissors mode: deformed nuclei

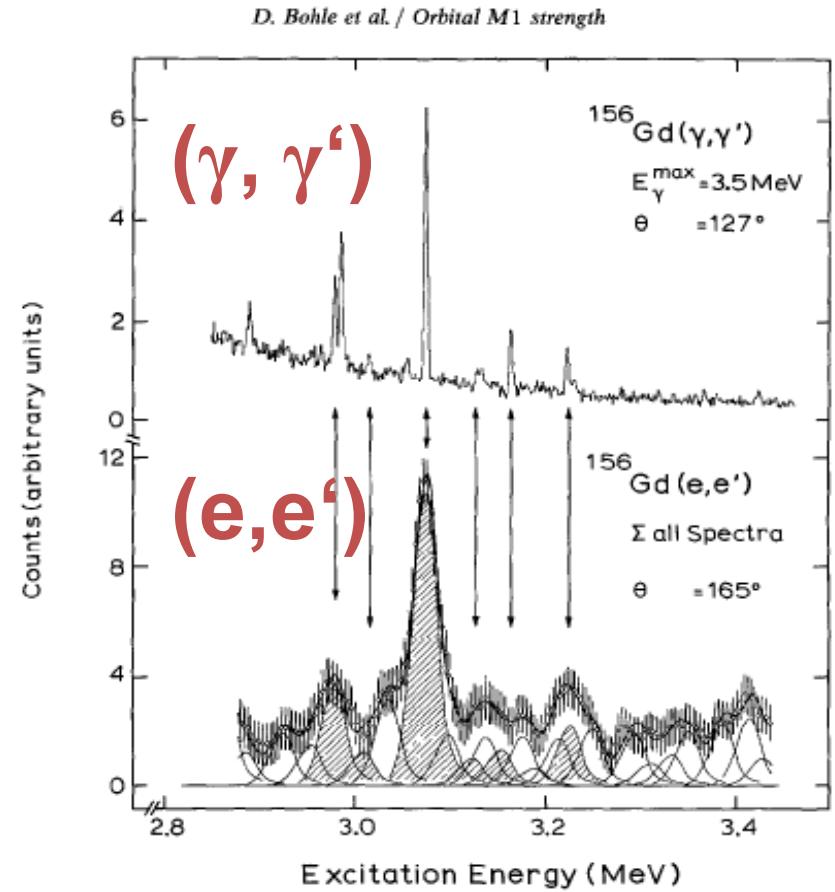


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Darmstadt, 1983
Achim Richter



Predicted by
Lo Iudice & Palumbo, 1978
@ 3 MeV: *Iachello, 1981 in IBM-2*



Bohle et al., NPA 458, 205 (1986).
DALINAC data

E2 decay strength of the Scissors mode



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PRL 118, 212502 (2017)

PHYSICAL REVIEW LETTERS

week ending
26 MAY 2017

E2 decay strength of the M1 scissors mode of ^{156}Gd and its first excited rotational state

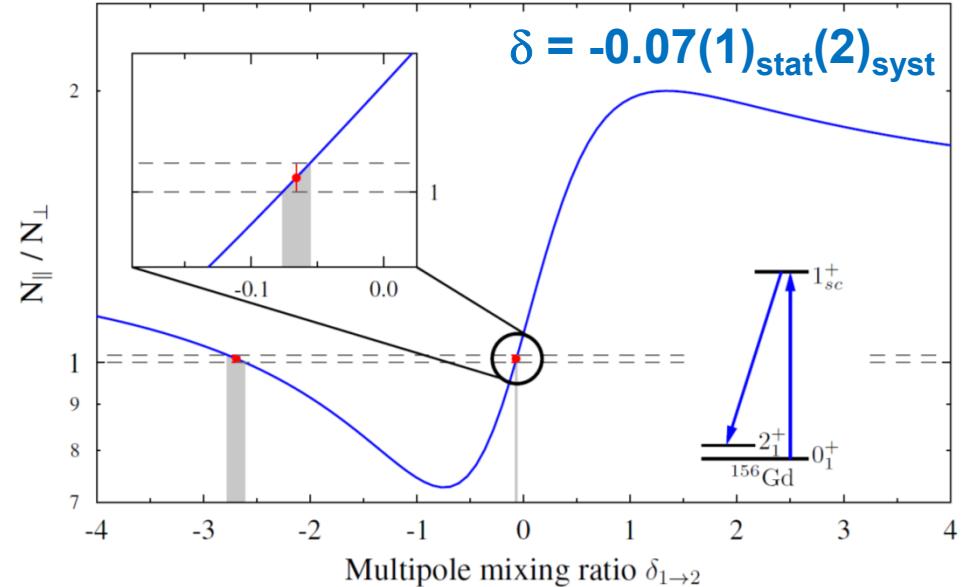
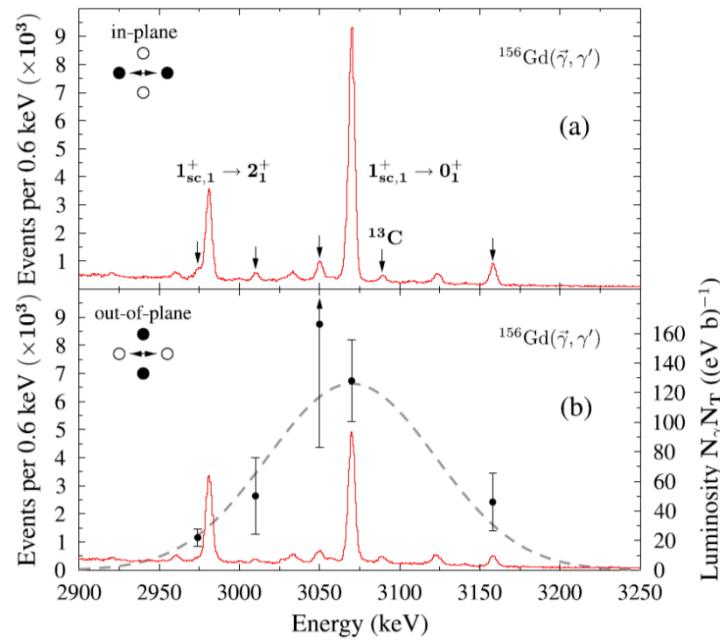
T. Beck,^{1,*} J. Beller,¹ N. Pietralla,¹ M. Bhike,² J. Birkhan,¹ V. Derya,³ U. Gayer,¹ A. Hennig,³ J. Isaak,^{4,5,†} B. Löher,^{4,5} V. Yu. Ponomarev,¹ A. Richter,¹ C. Romig,^{1,‡} D. Savran,^{4,5} M. Scheck,^{1,6,7} W. Tornow,² V. Werner,¹ A. Zilges,³ and M. Zweidinger¹

¹Institut für Kernphysik, TU Darmstadt, Schlossgartenstr. 9, D-64289 Darmstadt, Germany

²Department of Physics, Duke University and Triangle Universities Nuclear Laboratory, Durham, North Carolina 27708-0308, USA

³Institut für Kernphysik, Universität zu Köln, Zülpicher Str. 77, D-50937 Köln, Germany

B(E2; $1^+_1 \text{ScM} \rightarrow 2^+_1$) = 0.037(19) W.u.

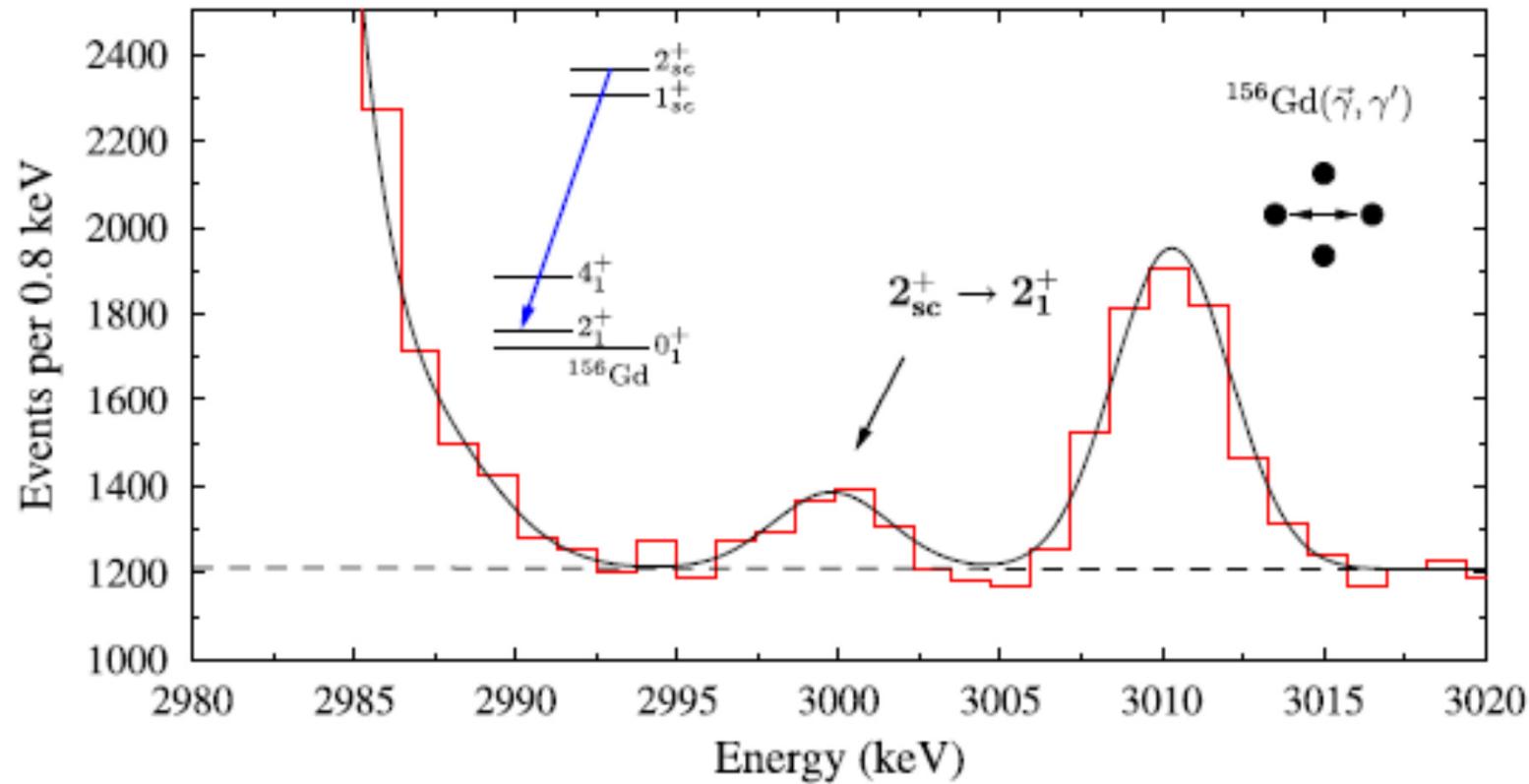


Evidence for rotational excitation of the Scissors Mode



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S-DALINAC & HI γ S data



T. Beck et al., Phys. Rev. Lett. 118 212502 (2017).

Supernova 1987A



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22.02.1987

Sanduleak-69 202



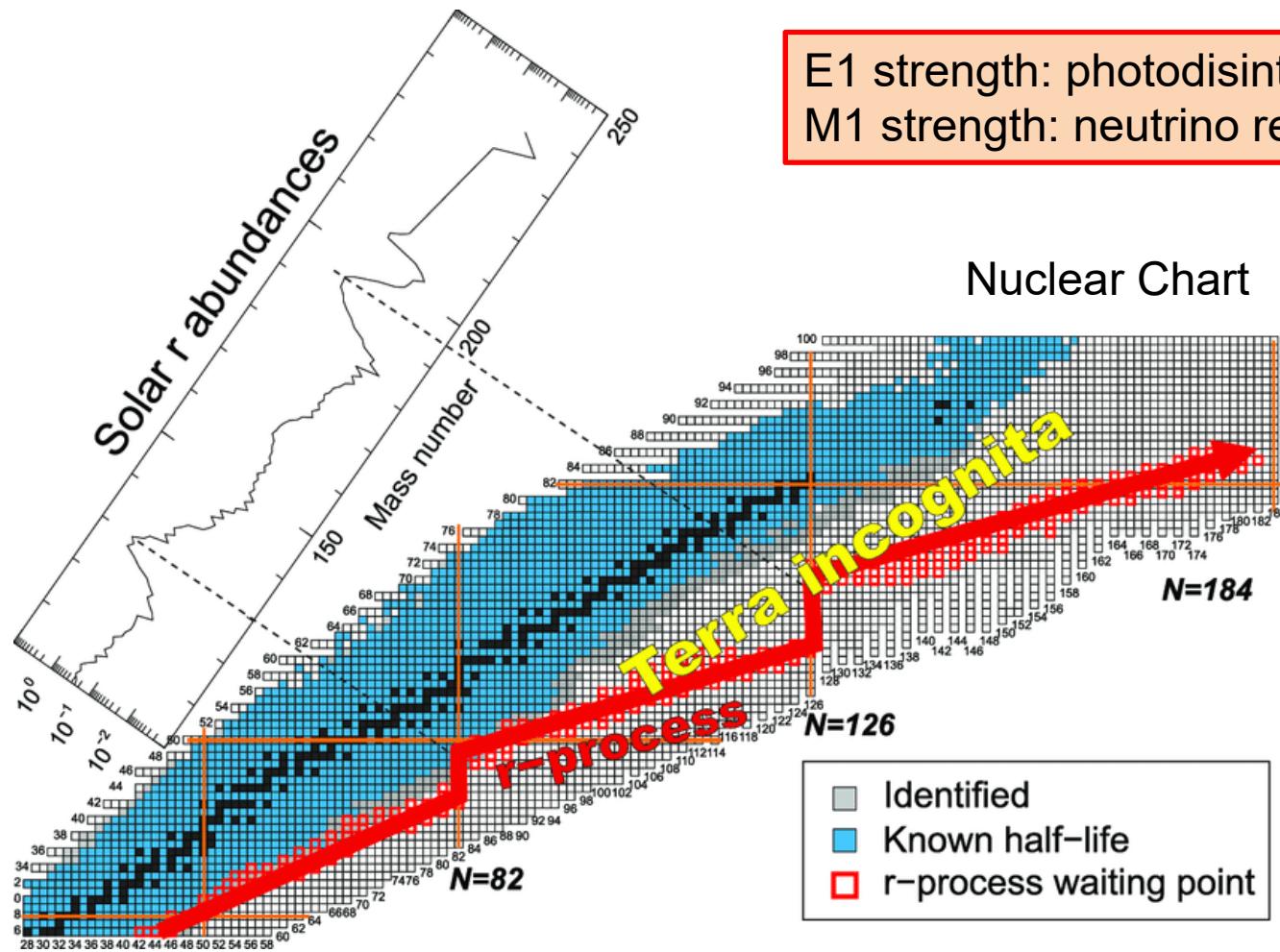
Great Magellanic Cloud
160.000 lightyears away

23.02.1987

Cosmic Nucleosynthesis



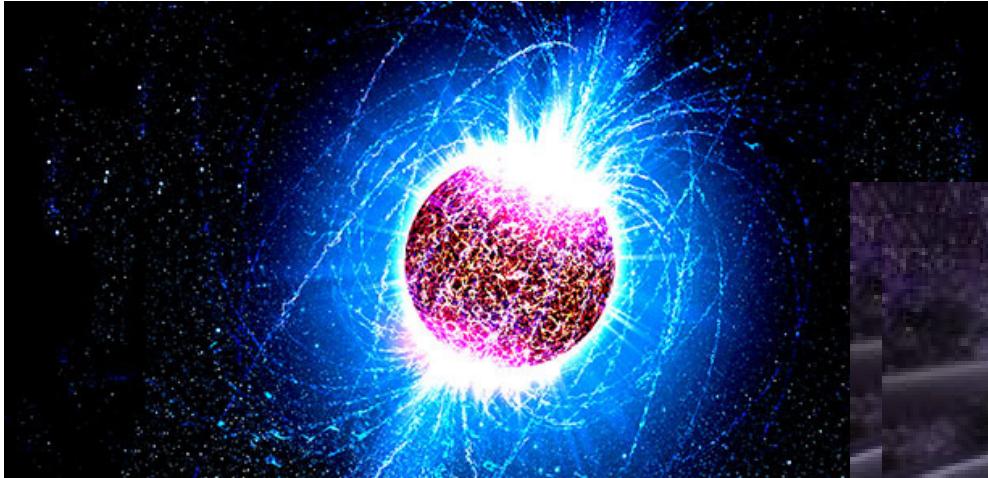
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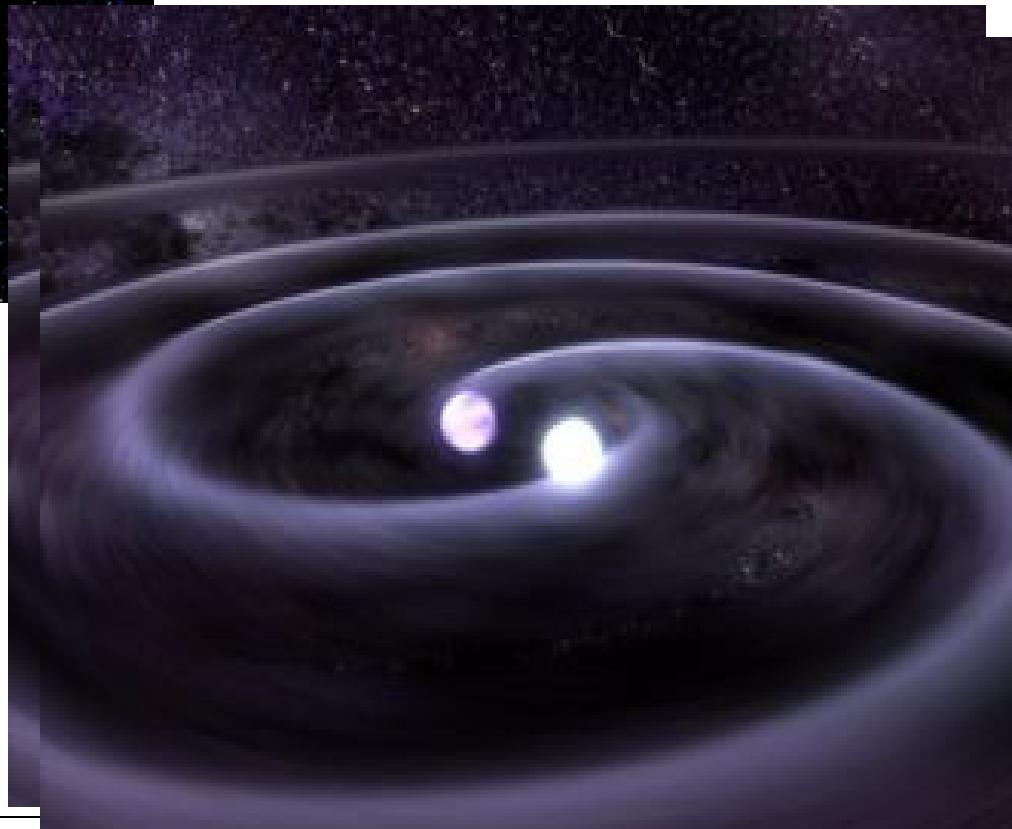
Neutron-Star Mergers



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Artist's views !



Abundant source of neutrons
for nucleosynthesis

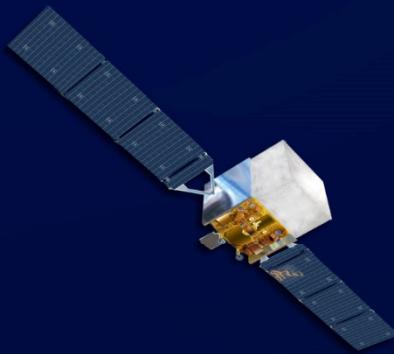
Neutron-star merger:
Explosive ejection of neutron-
rich matter into space

Multi-Messenger Observation of N-Star Merger

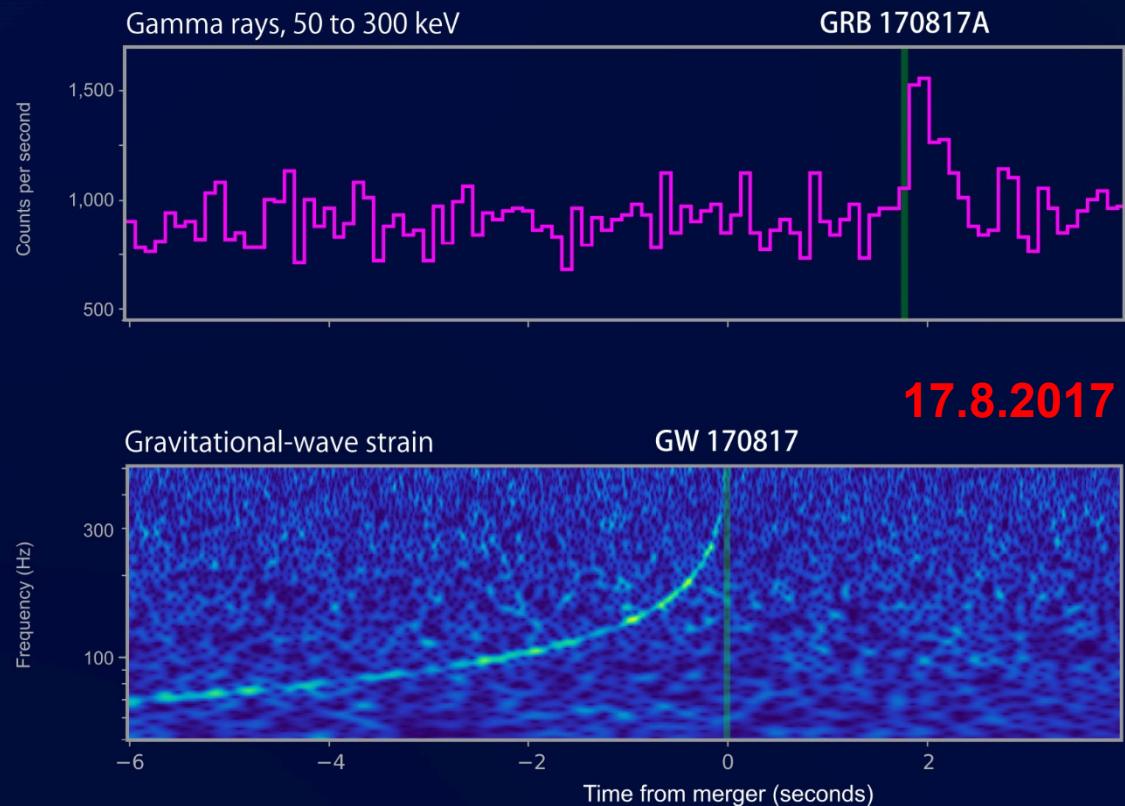


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Fermi



LIGO

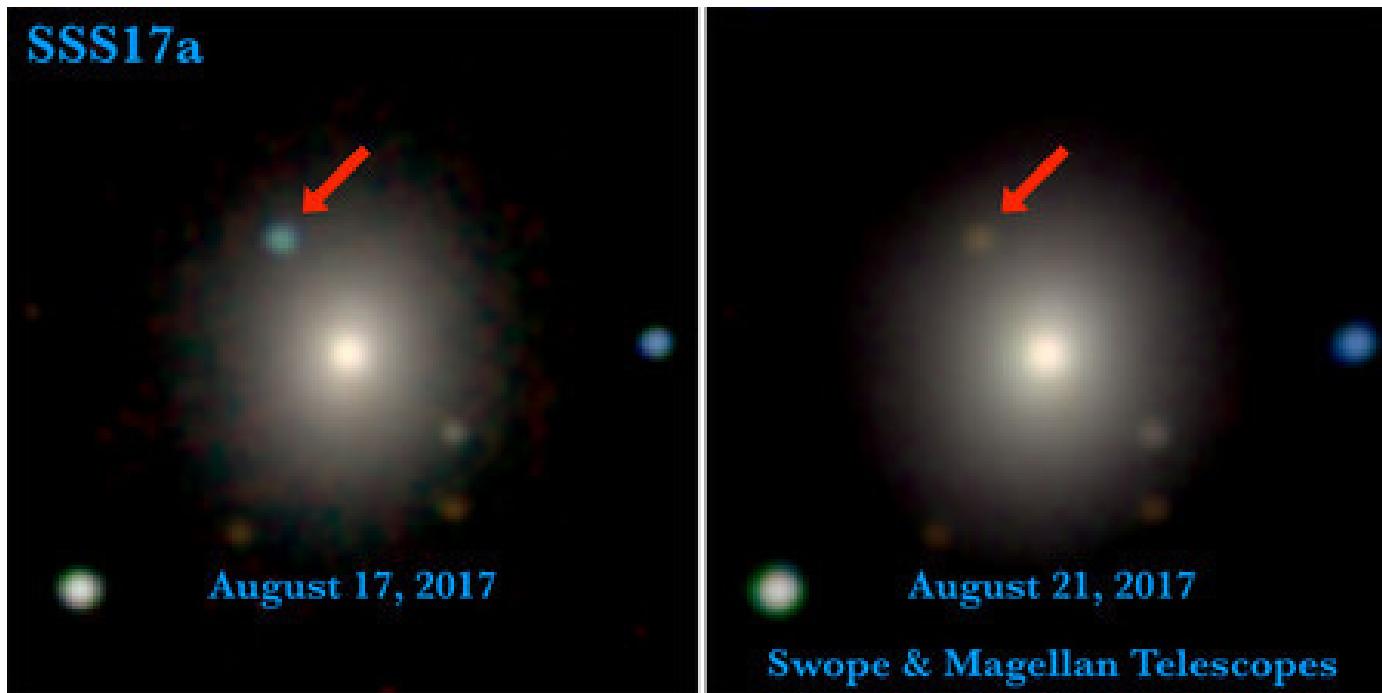


LIGO/VIRGO – collaboration, Fermi Satellite

Neutron-Star Merger in the Visible Spectrum



A star changes colour in 4 days ! „Kilonova“

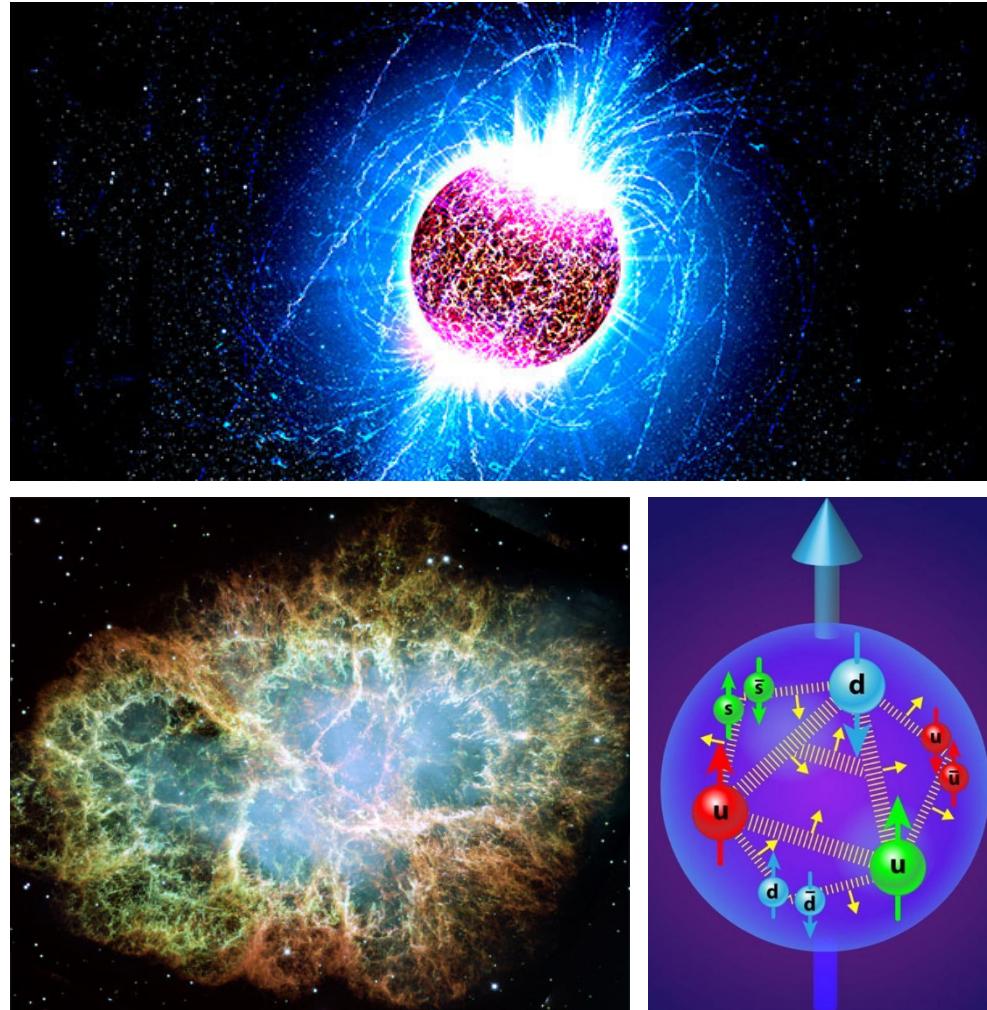


Light curve of neutron-star mergers due to synthesized rare-earth nuclei
B. D. Metzger et al., Mon. Not. Roy. Ast. Soc. 406, 2650 (2010).

Motivation for Nuclear Photonics

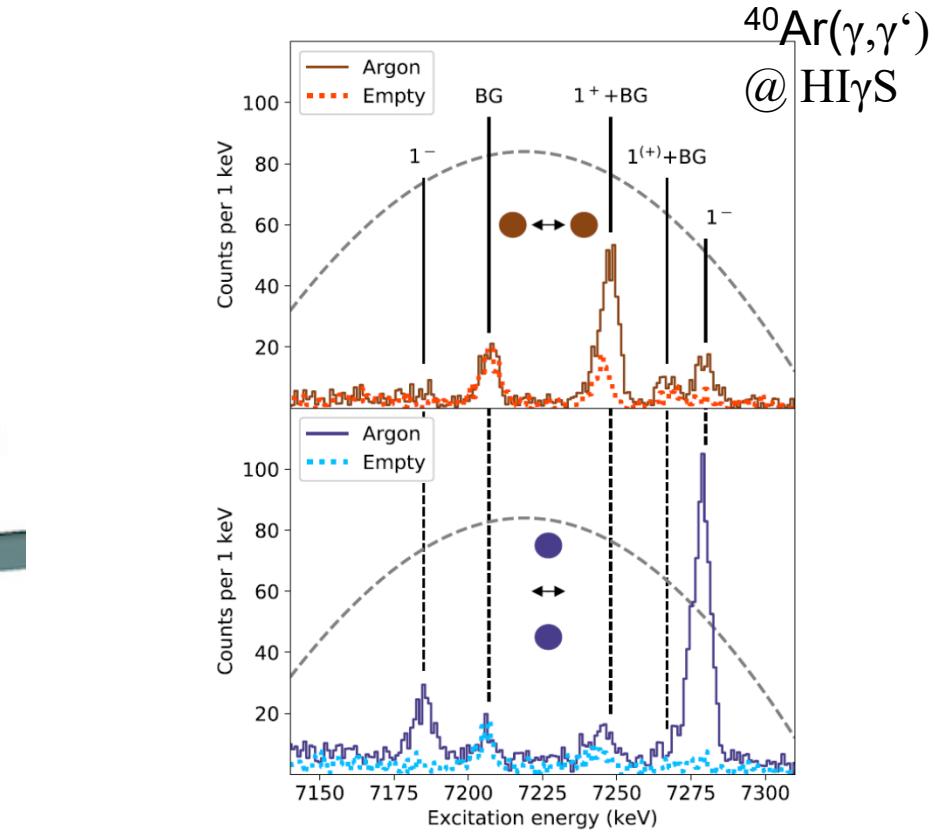
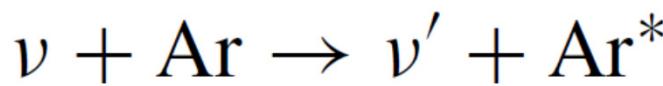
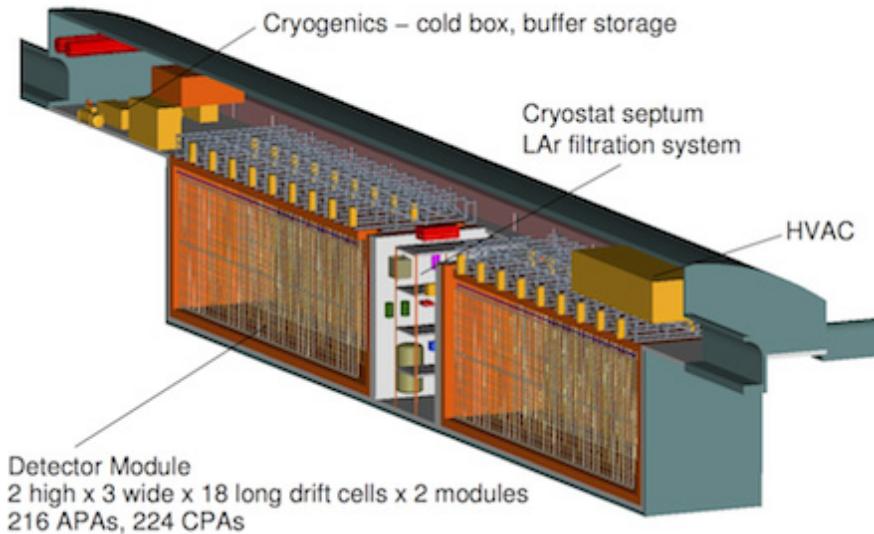


- Photons provide a sensitive probe for nuclear structure
- Properties of nuclear isotopes as a function of mass (neutron number)
- Low-energy frontier of Quantum Chromo-Dynamics
- Photonuclear reactions impact the structure and dynamics of stars
- Formation of chemical elements in the Universe
- Nuclei as detector material for neutrino experiments and searches beyond the Standard Model
- → **Nuclear photonics**



Deep Underground Neutrino Experiment

DUNE LAr – detector for cosmic ν 's

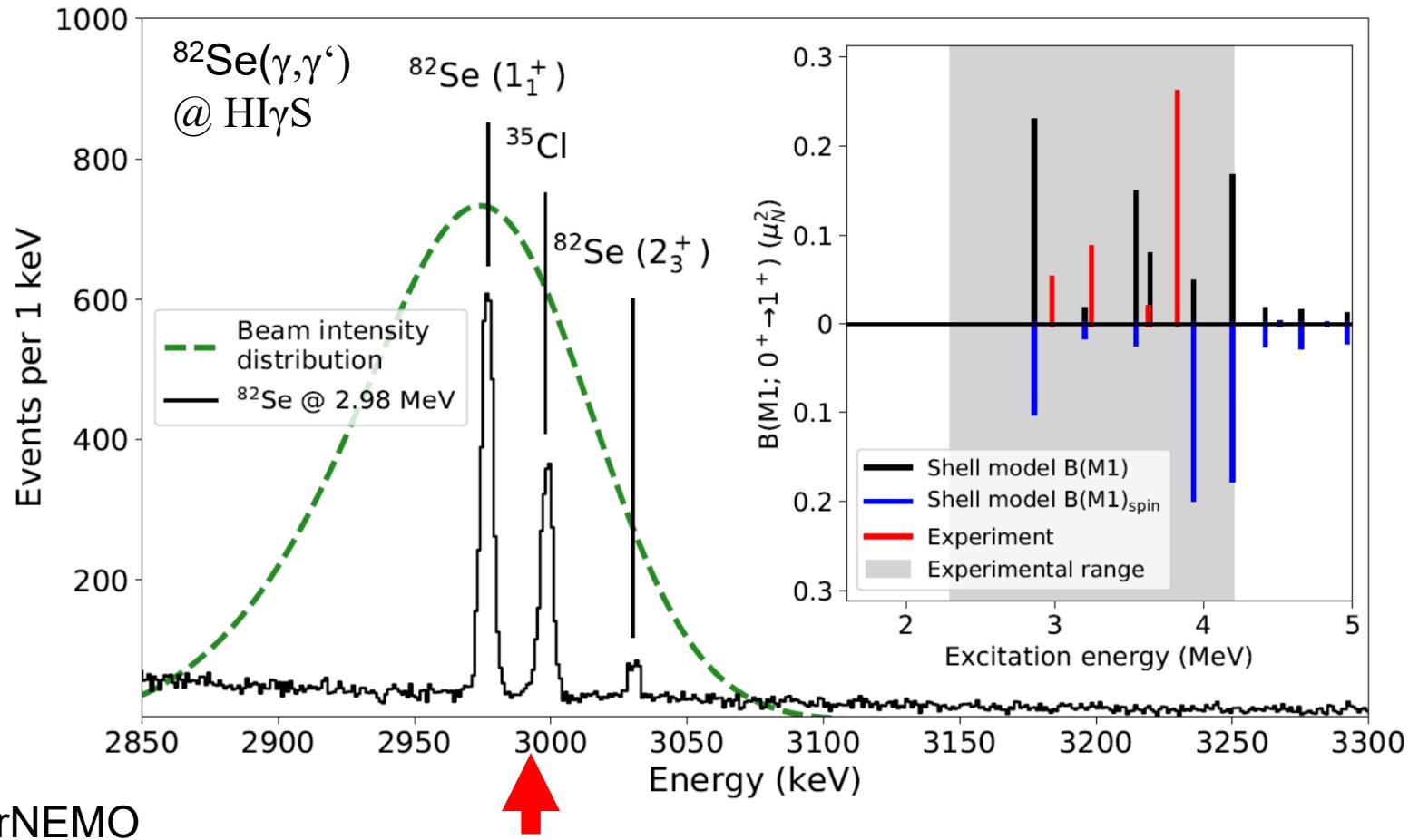


U.Gayer et al., Phys. Rev. C 100, 034305 (2019)
 → total low-energy neutrino cross section

Nuclear Structure for $0\nu\beta\beta$ searches



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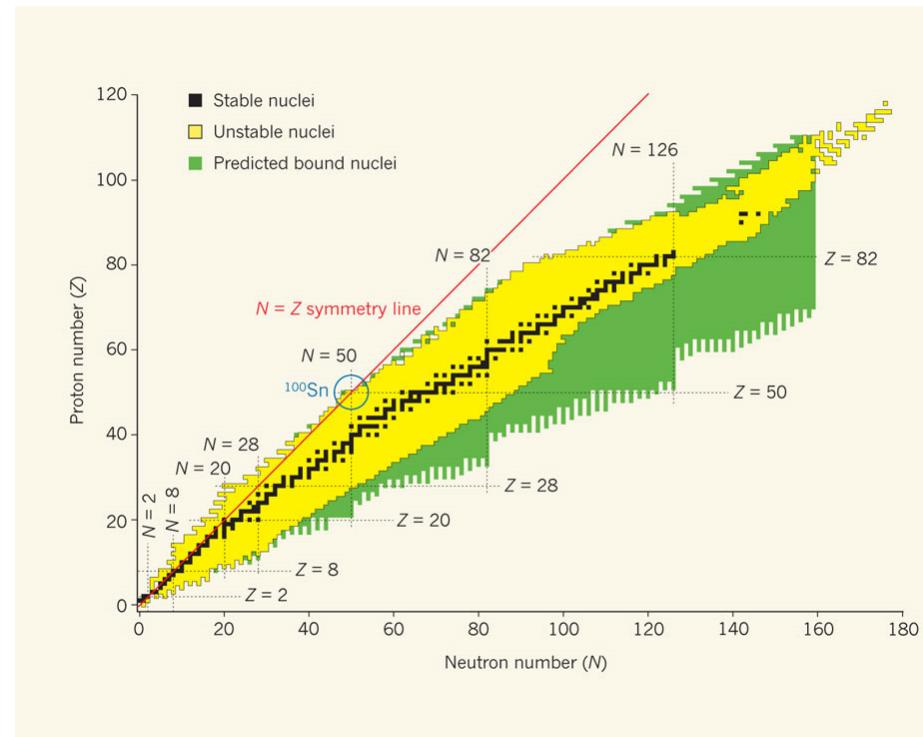


Physics with Photon Beams



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- Nuclear Structure Physics
 - Nuclear single particle structure
 - Collective nuclear structures
 - Photofission
- Particle-Physics Metrology
 - Neutrino detectors
 - Nuclear matrix elements for $\beta\beta$ -decay
- Nuclear Astrophysics
 - Capture / desintegration reactions
 - Nuclear synthesis
- Applications
 - Radiotomography of fuel rods
 - Cultural heritage, etc.



„Discovery Frontier“ for Nuclear Photonics at 4th Generation γ Source

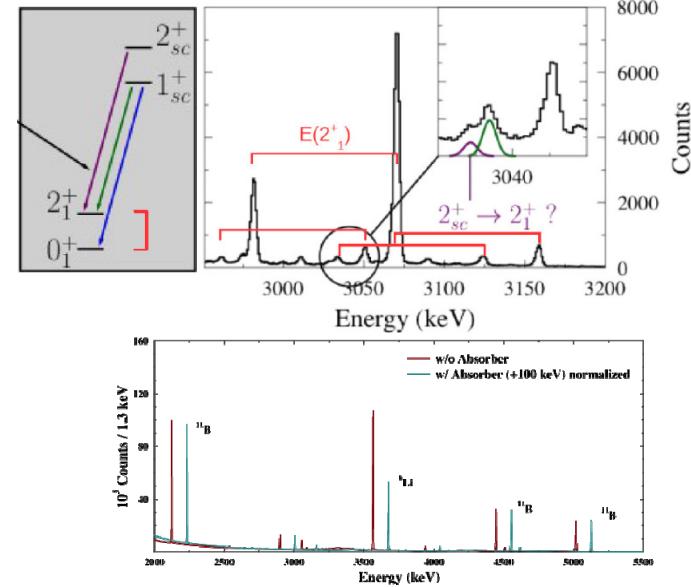


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**High-Intensity Frontier =
„Discovery Frontier“
(scientific opportunities)**



- „Availability Frontier“
Access to rare isotopes,
→ broader „nuclear gene pool“
- „Sensitivity Frontier“
weak channels: strong physics
- „Precision Frontier“
high count rates, new methods



Thank you very much!