Novel, Fast, Open-Source Code for Synchrotron Radiation Computation on Arbitrary 3D Geometries

Dean Andrew Hidas





Outline

Synchrotron Radiation Basics

Introduction to OSCARS

Basic SR Calculations

Calculations on 3D Geometries

• Time Dependent E, B



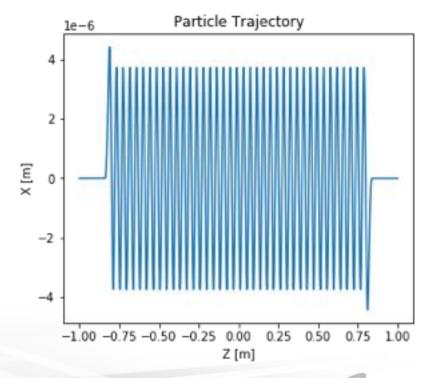


SR Basics

- Goal of looking at a beam and calculating radiative properties
 - Beam, Bending Magnets, Quads, Insertion Devices, Arbitrary Fields
- First is to solve for the trajectory
 - 2nd order differential equation

$$\frac{d\vec{p}}{dt} = q(\vec{E} + \vec{\beta}c \times \vec{B})$$

Typically solved via 4th order Runge-Kutta method







SR Basics

- Calculate a quantity of interest (from trajectory information)
- Flux, from E-field

$$\vec{E}(\vec{x},\omega) = C \int_{-\infty}^{+\infty} \left[\frac{1}{\gamma^2} \frac{\hat{n} - \vec{\beta}}{R^2 (1 - \hat{n} \cdot \vec{\beta})^2} + \frac{\hat{n} \times (\hat{n} - \vec{\beta}) \times \dot{\vec{\beta}}}{R (1 - \hat{n} \cdot \vec{\beta})^2} \right] e^{i\omega(\tau + R/c)} d\tau$$

- Numerically integrate this (and the like) to desired accuracy
 - Smaller trajectory steps = higher accuracy, cost is time and memory
 - GPU and multi-threading good candidates for this

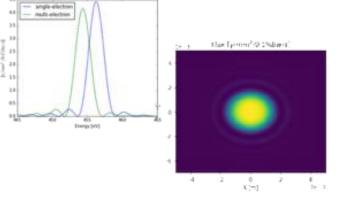


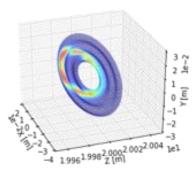


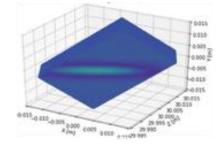
Introduction to OSCARS

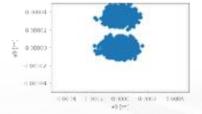
Open Source Code for Advanced Radiation Simulation

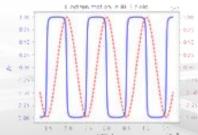
- Spectra
- Flux
- Power Density
- 3D Power Density (including CAD)
- Multiple beams
- Time dependent E(t) and B(t) fields















About OSCARS

- Multi-threaded internally
 - Simply add argument: nthreads=72 for faster calculation
- Capable of using Multiple-GPUs simultaneously



- Simply add: gpu=1 (for all available), or gpus=[0, 1, 4] to specify which GPUs to use
- Using CUDA, GPUs communicate directly (direct GPU-GPU data transfers)
- MPI compatible, Used on large-scale grid computing (i.e. Open Science Grid)
- Written in c++ (for speed) with python API
 - Technically: C-Python extension





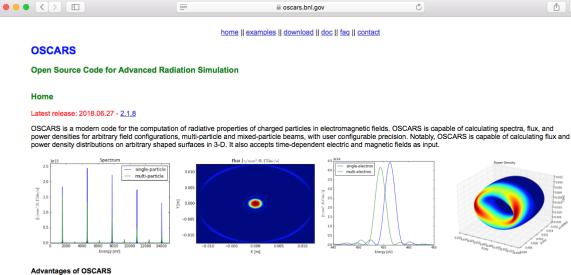
Where to find OSCARS

- Anywhere (pypi)
 - pip install oscars
- Anywhere (conda)
 - conda install oscars –c lightsource2-tag
- https://oscars.bnl.gov
- oscars@bnl.gov
- https://github.com/dhidas/OSCARS
- Full documentation and many examples









- · Accurate calculation with user defined precision.
- Multi-threaded and capable of using your GPU.
- . Input of time dependent fields.
- . Designed with very large scale computing in mind.
- Simple and very powerful python API (application program interface).
- 100% Open Source.

The core of OSCARS is written in modern c++ for speed with a simple python user interface. No additional packages are required to run the core of OSCARS. One can easily run OSCARS on their desktop or laptop computer. It also comes with utilities to use MPI for your local machine and cluster usage. Significant gains are achieved through the use of graphical processing units (GPUs) and OSCARS makes this very easy for compatible nvidia GPUs. OSCARS was also designed with very large scale computing in

OSCARS is also intended to be an open source and community based project with contributions from anyone willing and able regardless of institution or position. We also encourage user feedback in order to continually improve the ease of use and calculations within. If you are interested in contributing or giving us feedback, please send an

Last modified: 27 June 2018.

oscars@bnl.go





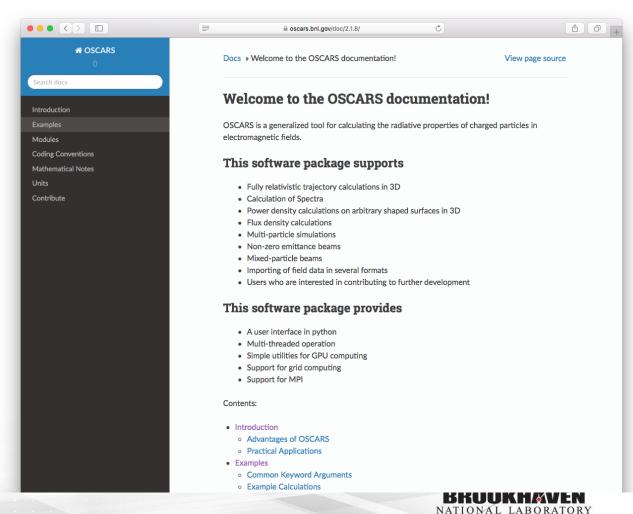
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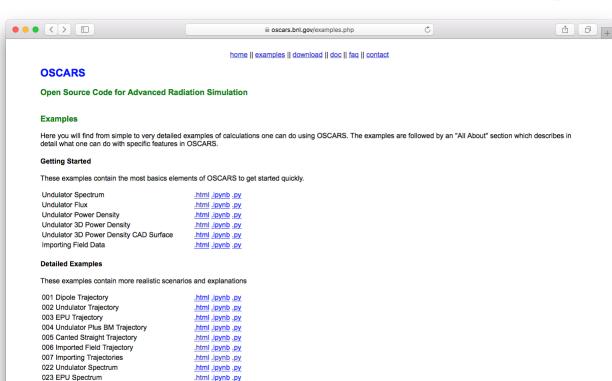
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.html .ipynb .py MPI





032 Undulator Flux

042 Undulator Power Density

122 Undulator Spectrum Multi-Particle

142 Undulator Power Density Multi-Particle

123 EPU Spectrum Multi-Particle

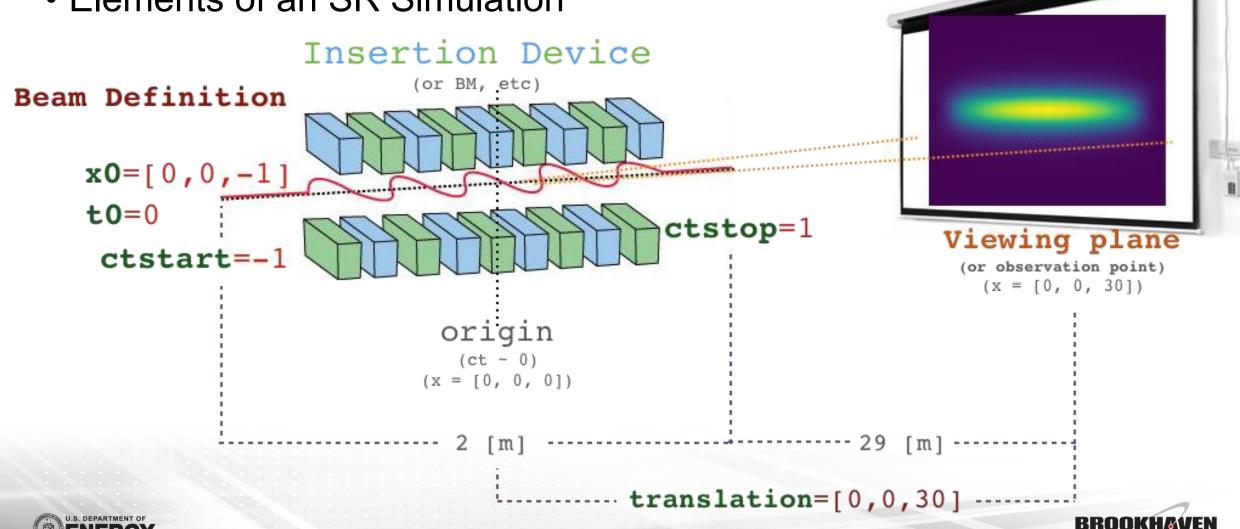
132 Undulator Flux Multi-Particle 133 EPU Flux Multi-Particle

043 EPU Power Density

033 EPU Flux

SR Simulation Construction

Elements of an SR Simulation



SR Full Simulation (from previous page)

Should be easy, here to help you

```
import oscars.sr
osr = oscars.sr.sr(gpu=1,
                   nthreads=16)
osr.set_particle_beam(type='electron',
                      energy GeV=3,
                      ctstartstop=[-1, 1]
osr.add bfield undulator(bfield=[0, 0.8375, 0],
                         period=[0, 0, 0.042],
                         nperiods=38)
osr.calculate_power_density_rectangle(plane='XY',
                                       width=[0.04, 0.04],
                                       npoints=[101, 101],
                                       translation=[0, 0, 30])
```

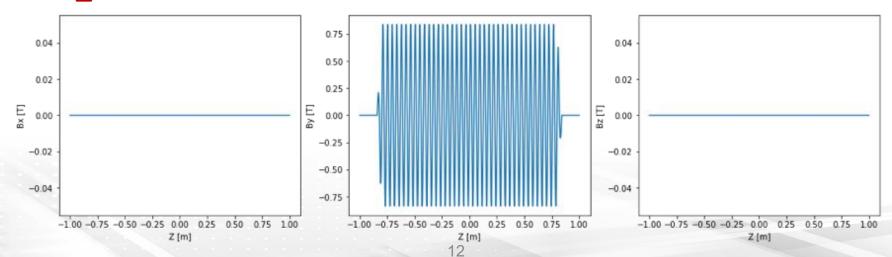




Magnetic Fields

- Many types of built-in fields
 - Undulator, bending magnet, gaussian, quadrupole, arbitrary Python function, data files, interpolated data
- Add a simulated undulator

osr.add_bfield_undulator(bfield=[0, 0.8375, 0], period=[0, 0, 0.042], nperiods=38)



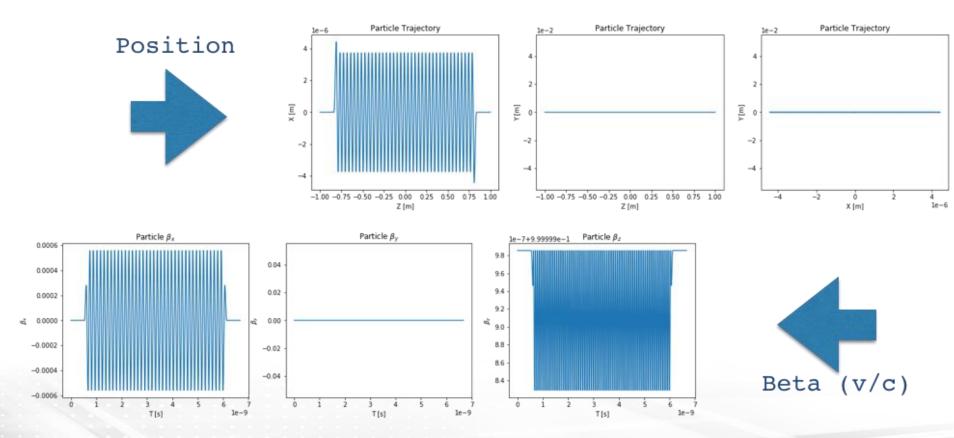




Calculate Trajectory

Calculate trajectory as a sanity check (not necessary)

osr.calculate_trajectory()



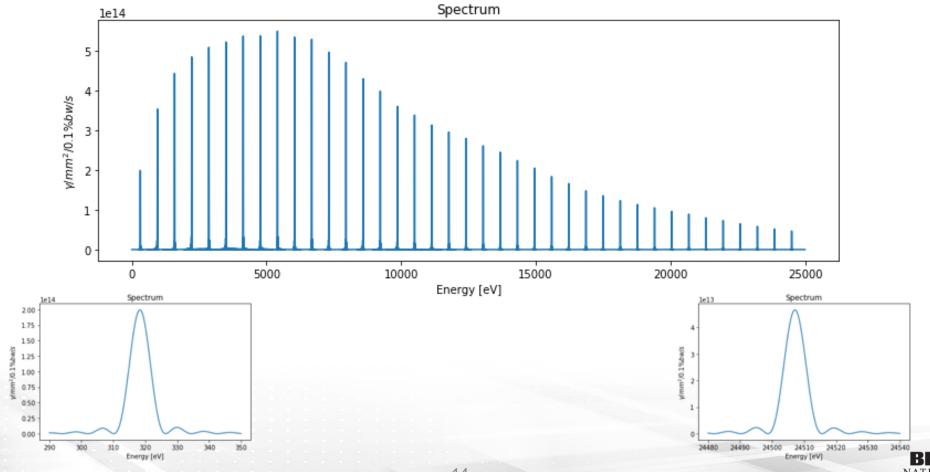




Calculating Spectra

Calculate spectrum at observation point

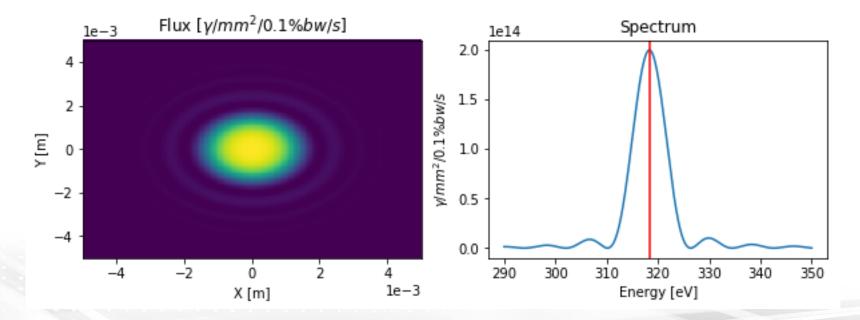
osr.calculate_spectrum(obs=[0, 0, 30], energy_range_eV=[10, 25000])





Calculating Flux

Calculating the flux at specific wavelength on an observation plane

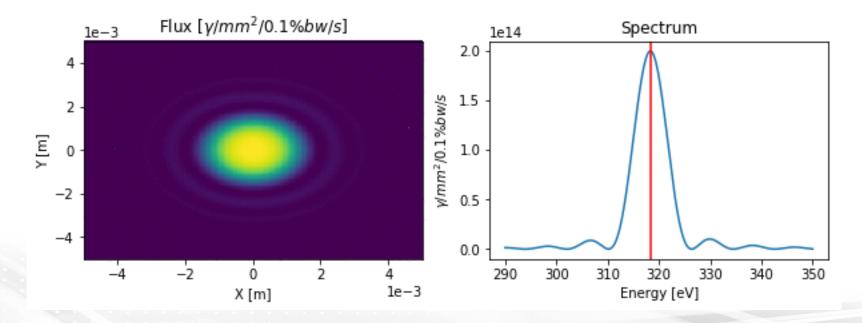






Calculating Flux

Calculating the flux at specific wavelength on an observation plane

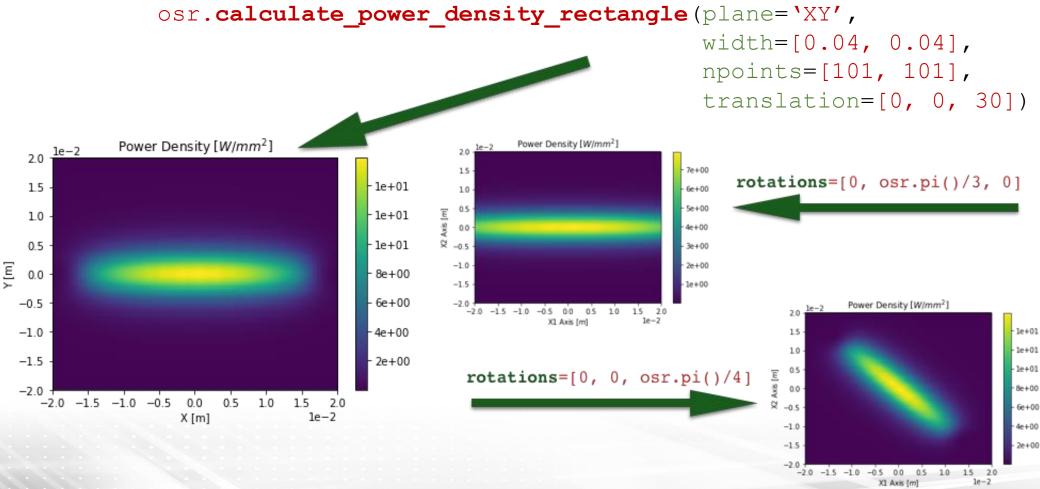






Calculating Power Density

Calculate power density on observation plane



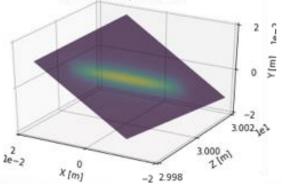


Calculating Power Density

Calculate power density on observation plane (3D)

osr.calculate power density rectangle (plane='XY', width=[0.04, 0.04], npoints=[101, 101], translation=[0, 0, 30])Power Density [W/mm2] Power Density [W/mm2] rotations=[0, osr.pi()/3, 0] 3.002,02 Power Density [W/mm2] 3.002,02 3.000 ml -2 2.998 X [m] -2 2.998



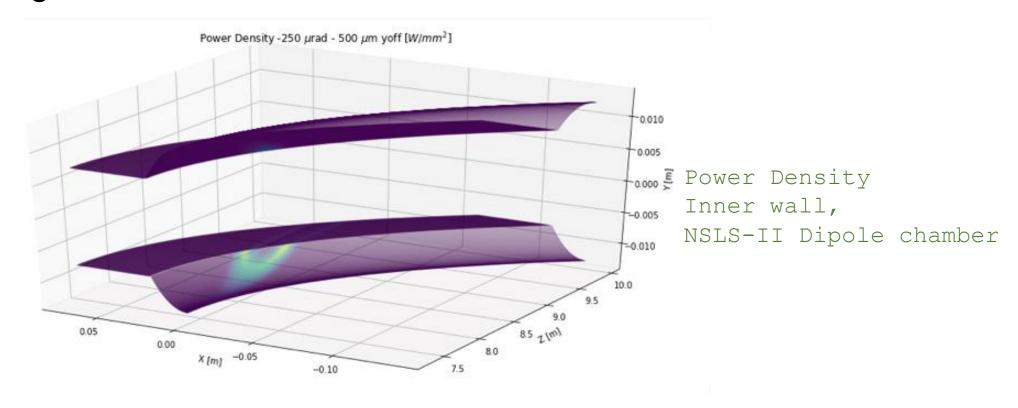






Power Density – Parametric 3D

Advantage: Much easier visualization



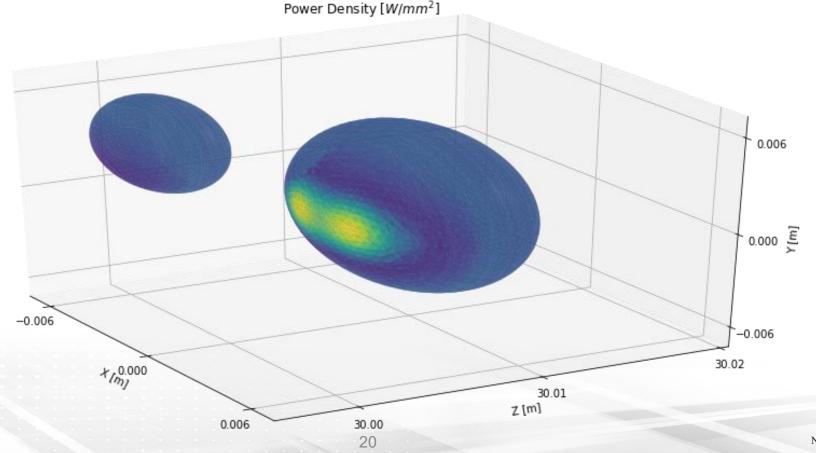
Major disadvantage: Very difficult for complex objects





Power Density CAD

- Can now import CAD model in STL format
- Allows for substantially more complex objects (than shown here)

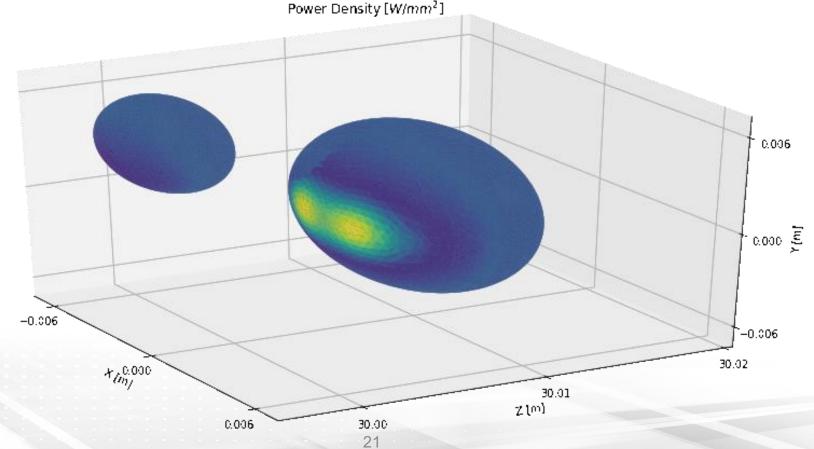




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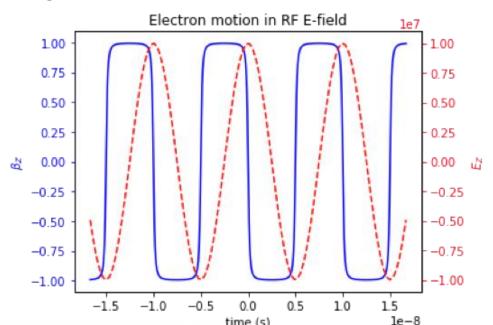


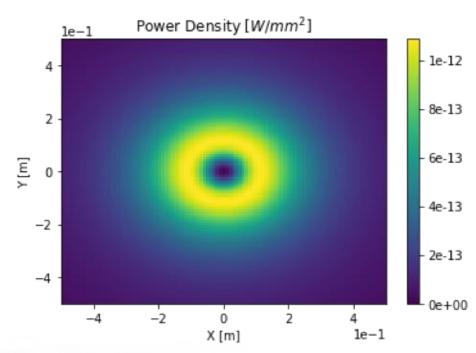


Time Dependence

• Any functional form B(x, y, z, t), E(x, y, z, t) or real field data in resonance

• E.g. 100 [MHz], 10 [MV/m]





Also valid for very high fields > 100 [GV/m]





Conclusion

- Synchrotron Radiation Basics
- Introduction to OSCARS and basic calculations
- Calculations on 3D Geometries
- Time Dependent E, B
- Visit, contribute: https://oscars/bnl.gov

Thank You



