

Towards the production of an anti-hydrogen beam

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(ASACUSA-CUSP collaboration)

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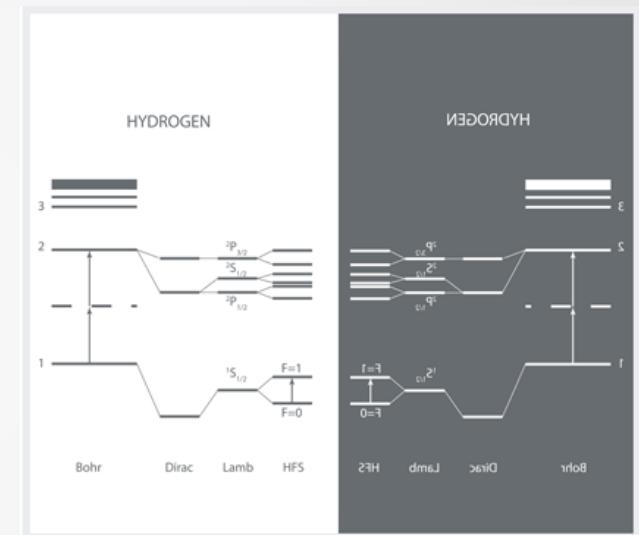


Why Antihydrogen ?

*Antihydrogen is the **simplest atom** consisting entirely of antimatter.

*Hydrogen counterpart is one of the **best understood** and most precisely measured atoms in physics.

*A comparison of antihydrogen and hydrogen offers one of the most **sensitive tests** of CPT symmetry.



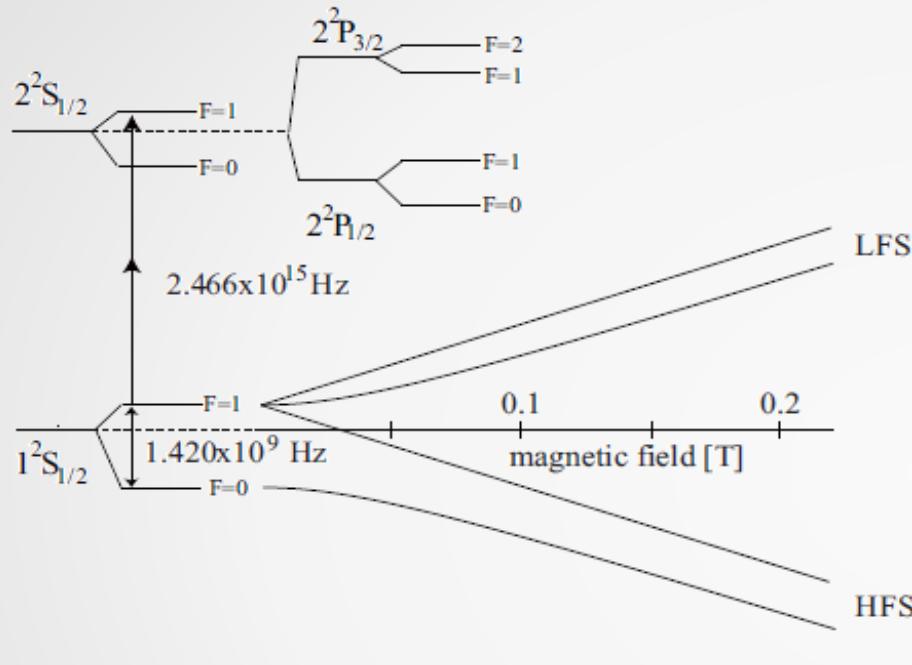
What ?

High precision spectroscopy of the ground state hyperfine splitting of antihydrogen.

How ?

Creation of a spin-polarized anti-hydrogen beam

Low energy anti-hydrogen atoms (Level diagram)



$e^+ \bar{p}$

↓

↓

Low Field Seekers (LFS)

$$\nu_{HFS} = \frac{16}{3} \left(\frac{m_p}{m_p + m_e} \right)^3 \frac{m_e \mu_p}{m_p \mu_N} \alpha^2 c R y + \Delta$$

↑

↓

↑

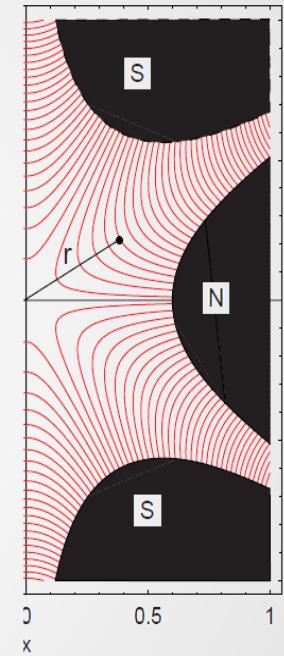
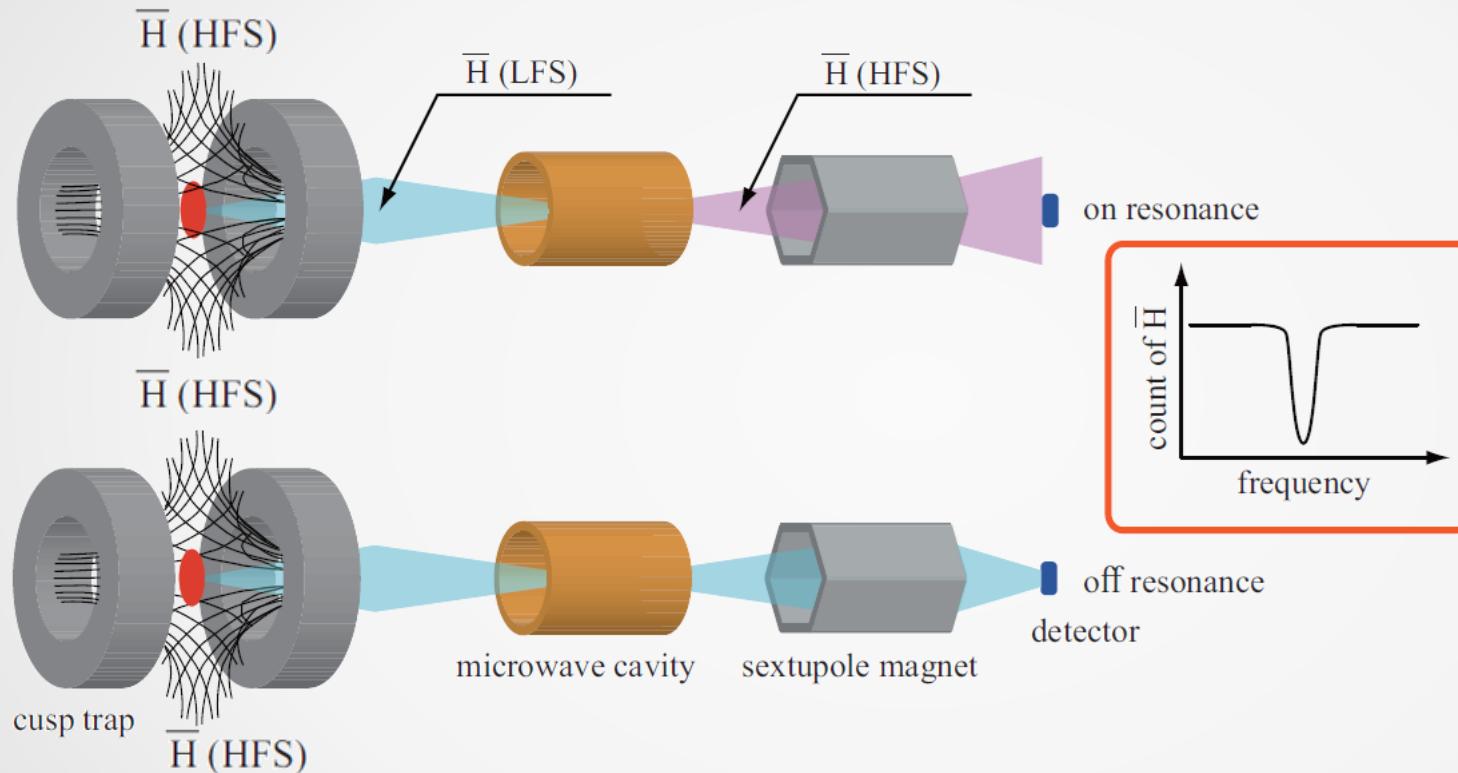
↑

High Field Seekers (HFS)

	experiments (Hz)	$\Delta\nu_{\text{exp}} / \nu$	$ \nu_{\text{theory}} - \nu_{\text{exp}} / \nu$
ν_{1S-2S}	2,466,061,413,187,103(46)	1.7×10^{-14}	1×10^{-11}
ν_{HFS}	1,420,405,751.768(1)	6.3×10^{-13}	$(3.5+0.9) \times 10^{-6}$

Aim to measure ν_{HFS} with a precision of 1.8×10^{-6}
 Measure antihydrogen under field-free conditions!

Measuring the hyperfine structure



Anti Helmholtz configuration:

HFS-states: de-focused

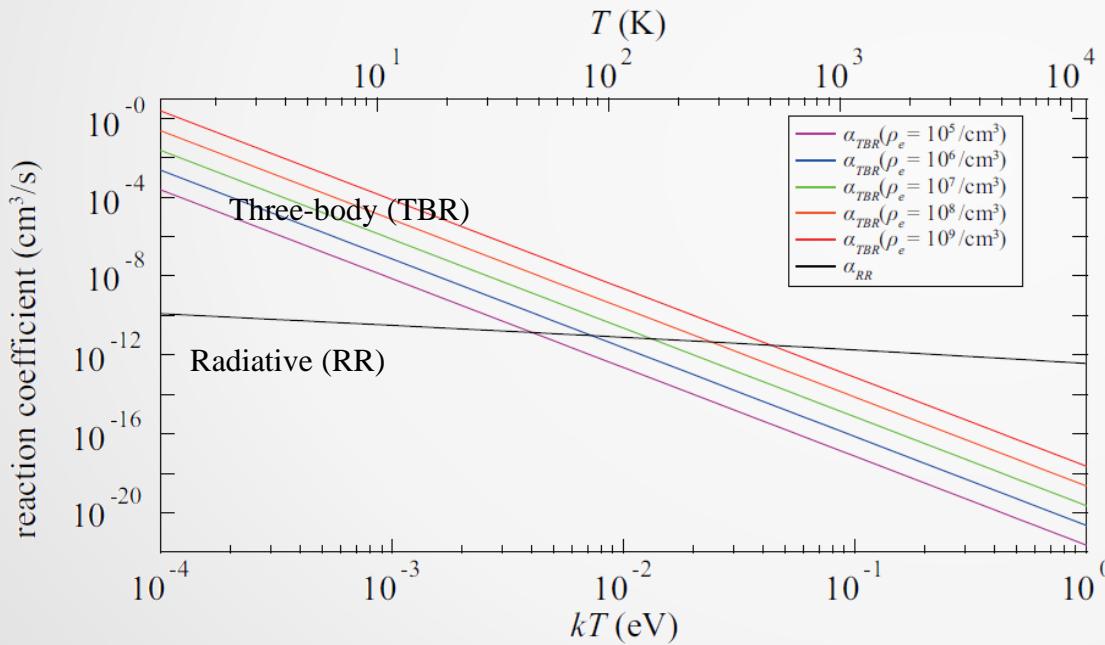
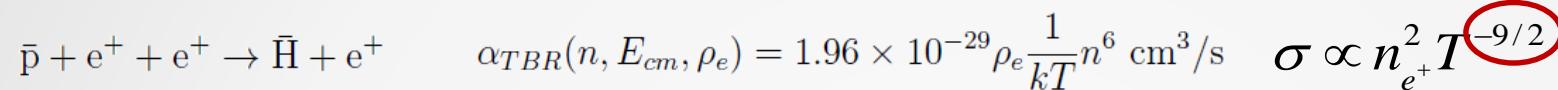
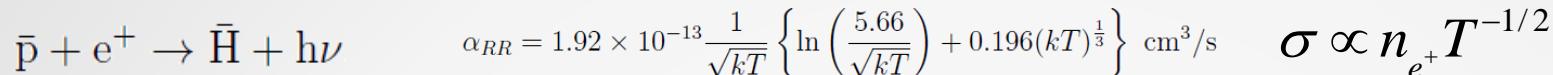
LFS-states: focused

We require an antihydrogen beam!

[1] A. Mohri and Y. Yamazaki, Europhys. Lett. **63** (2003) 207. [2] Y. Enomoto et al., Phys. Rev. Lett. **105** (2010) 243401

Cold = ☺ (1)

(a) \bar{H} formation rate

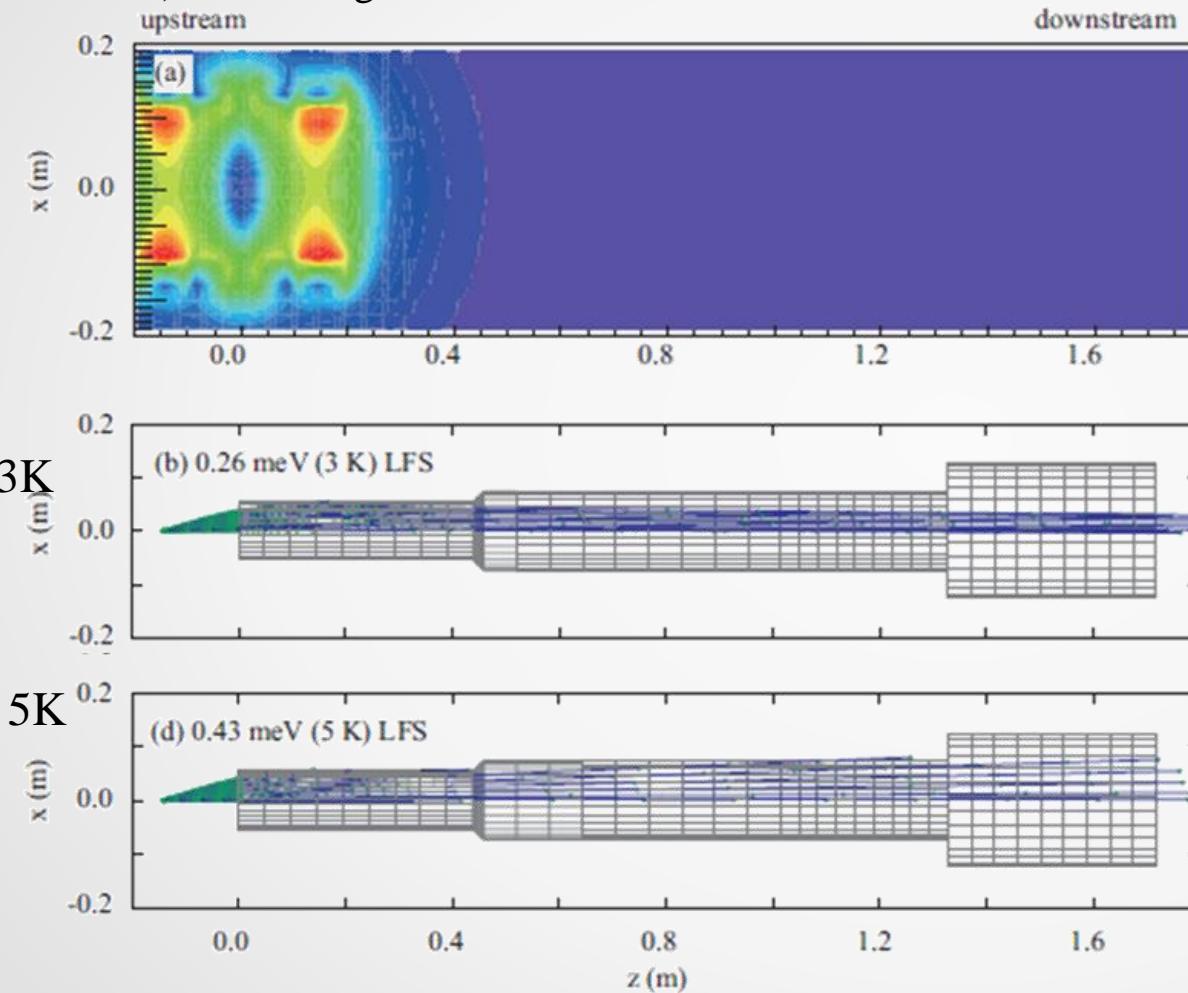


Due to the strong dependence on the density and temperature of e^+ , three body recombination is the dominant process at $\sim 10^8 \text{ cm}^{-3}$, $\sim 10 \text{ K}$

At a low temperature, a high density positron plasma is favorable for the effective production of anti-hydrogen atoms

Cold = ☺ (2)

2) \bar{H} focusing

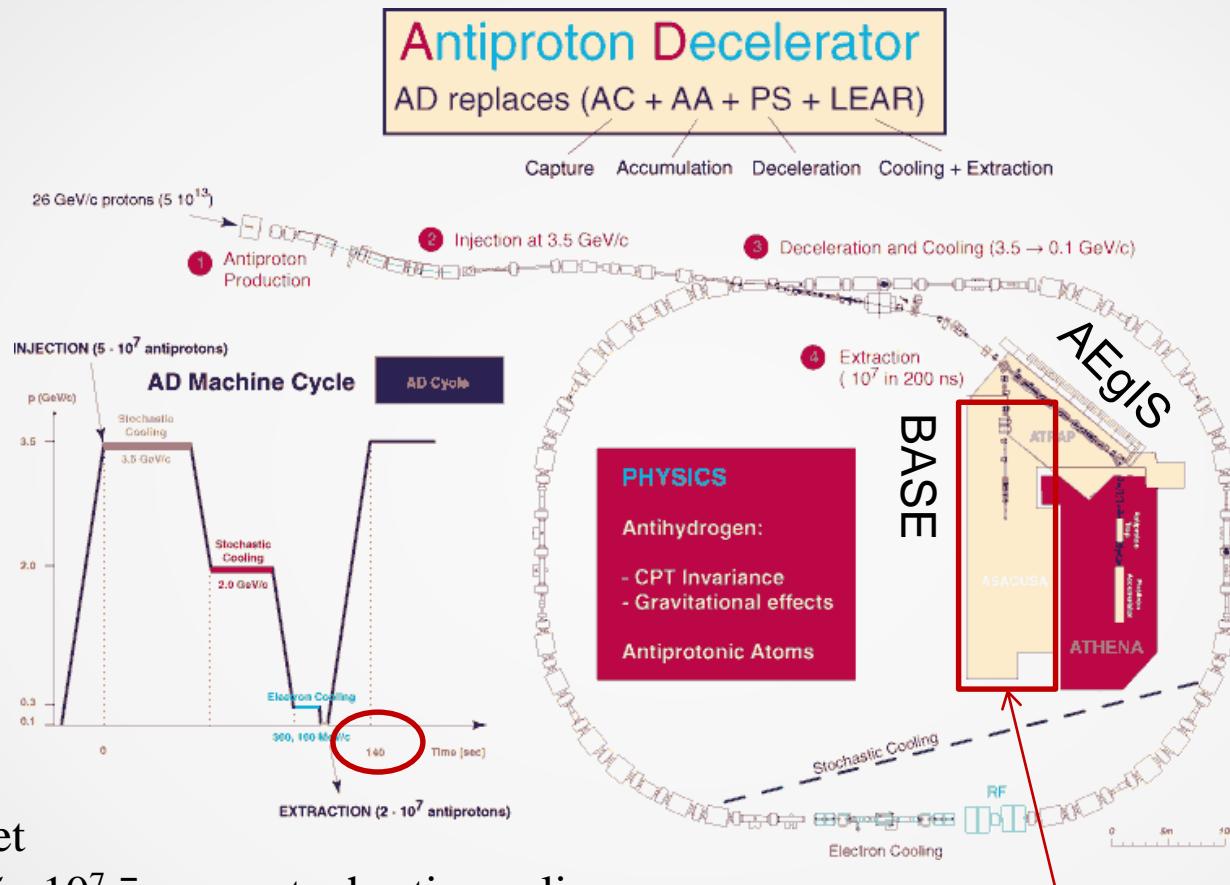


The intensity of LFS \bar{H} from CUSP to detector is much enhanced when the temperature is reduced

For 3K compared to the solid angle, the intensity changes:
→LFS x 50 !!
→HFS suppressed to 1/100

- 3) Interaction time with u-wave radiation in the cavity
- 4) Longer TOF \rightarrow lower n-state

Antiproton creation (AD)

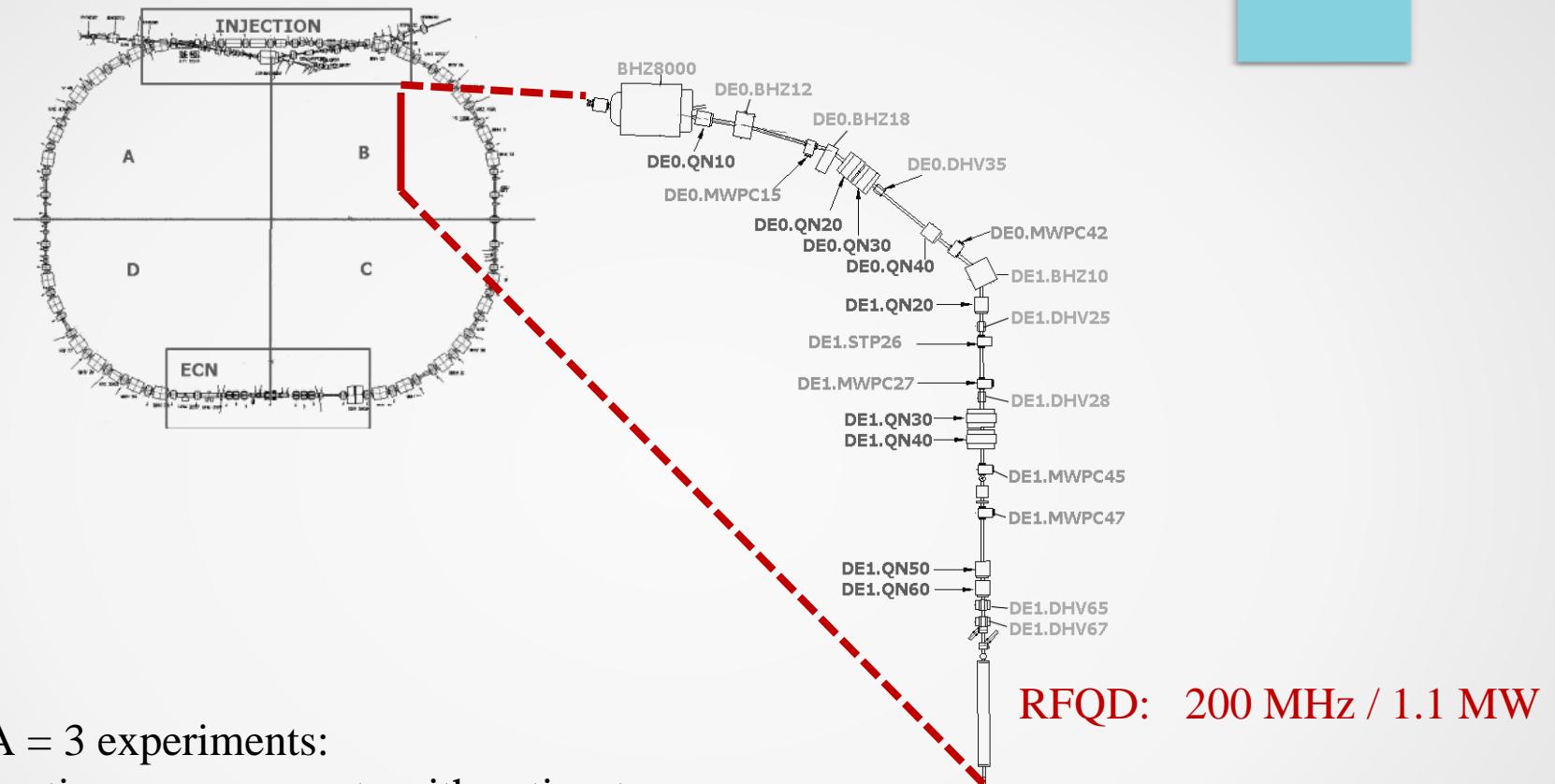


Production target

- 1) 3.5GeV/c ~ 5×10^7 \bar{p} stochastic cooling
 - 2) 2 GeV/c stochastic cooling
 - 3) 300MeV/c e^- cooling
 - 4) 5.3 MeV (100 MeV/c) e^- cooling
- ~ 10^7 \bar{p} pulse from AD (~100s cycle)
- a) 110keV <math><5 \times 10^6</math> pbar with RFQD

ASACUSA area

ASACUSA collaboration:

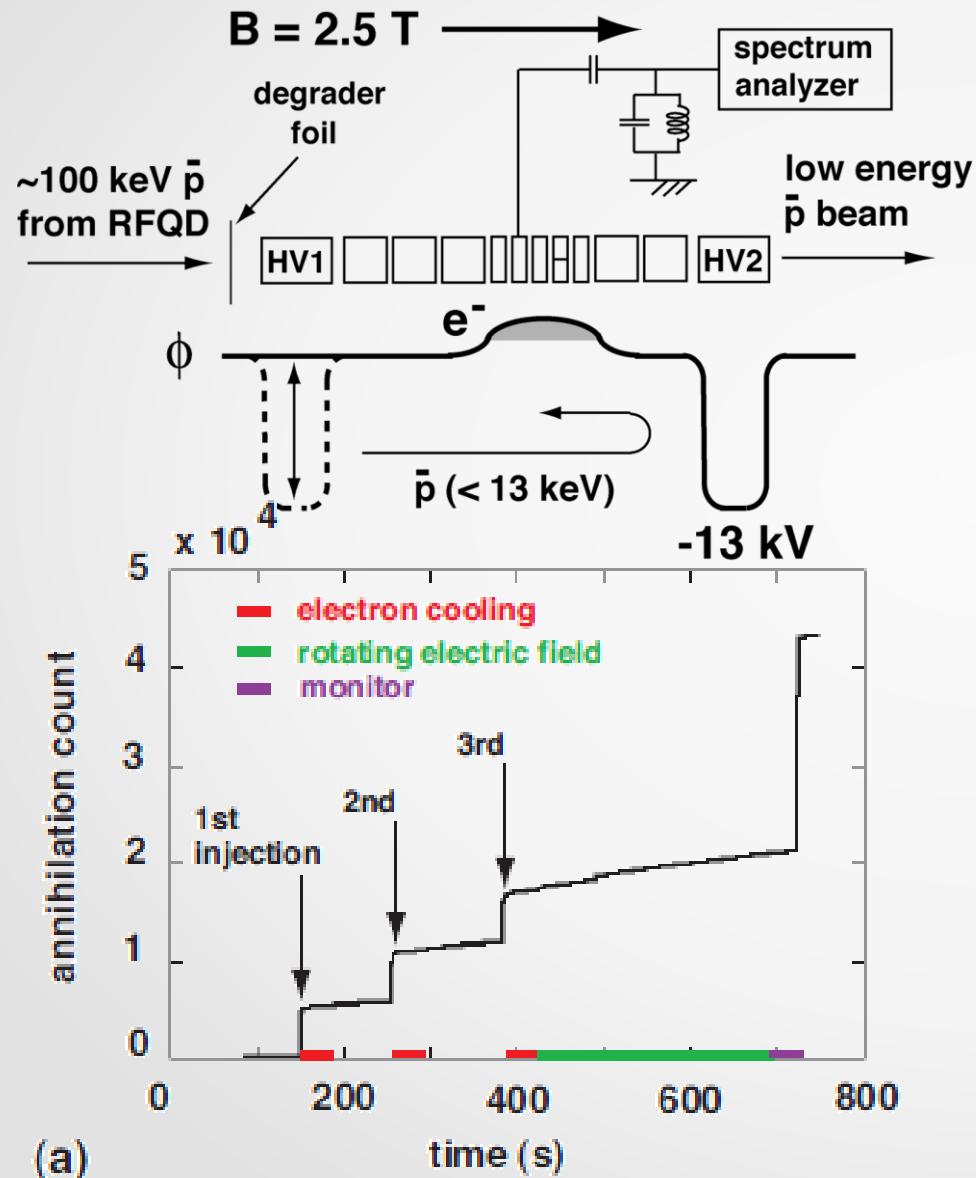


ASACUSA = 3 experiments:

- 1) Cross section measurements with antiprotons
 - 2) \bar{p} -Helium
 - 3) Antihydrogen: ASACUSA-CUSP experiment
- > 3 months/experiment

5-50 times more \bar{p} /AD cycle compared to other AD experiments

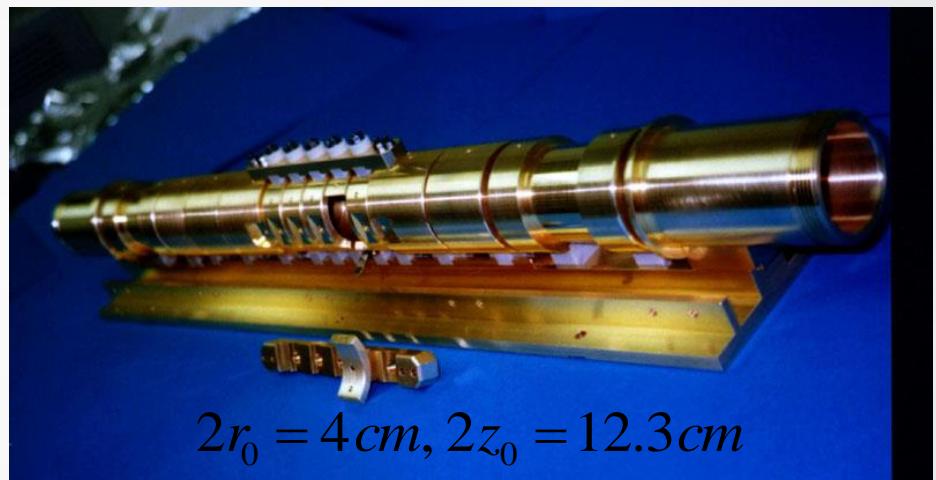
Antiproton accumulator + buncher (MUSASHI trap)



[4] N. Kuroda, H. A. Torii, K. Y. Franzen, et al., Phys. Rev. Lett. 94, 023401 (2005).

- [5] X.-P. Huang, et al., Phys. Rev. Lett. **78**(1997)875
 [6] H. Saitoh, et al., Phys. Rev. A **77**(2008)051403(R)

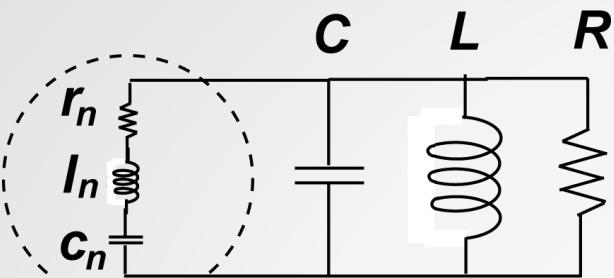
< 13 keV pbar confined in the trap
 $10^6 \bar{p} / \text{AD shot}$



Sympathetically cooled with $3 \times 10^8 e^- s$

A rotating wall electric field is superimposed on the trap ring electrodes to control the p-bar density

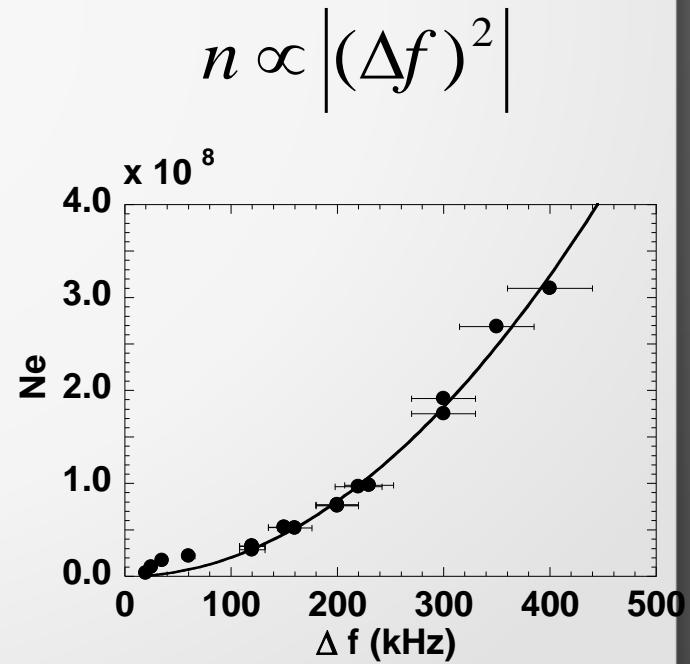
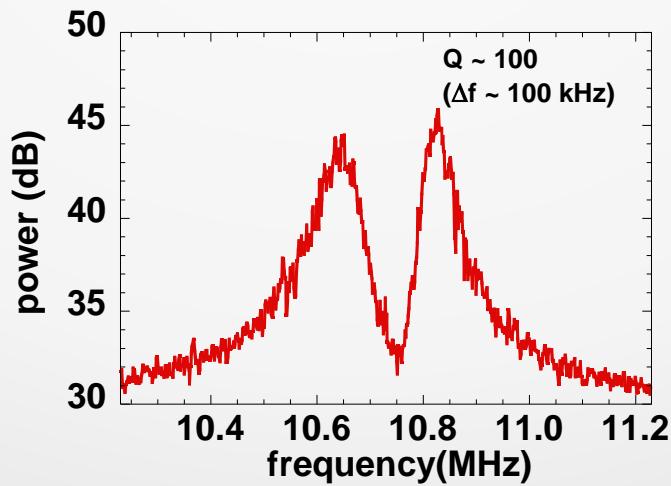
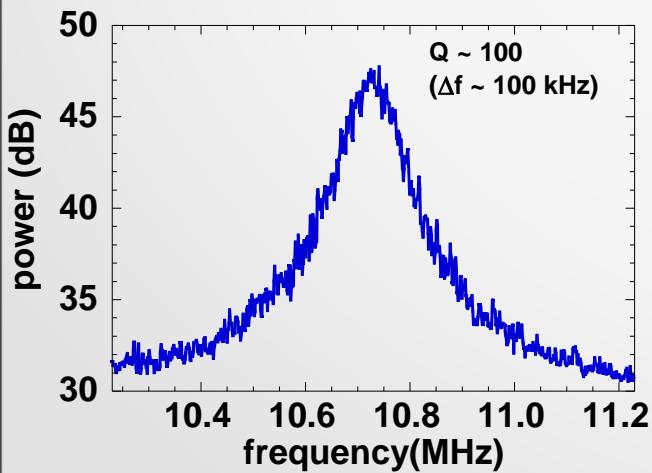
Resonant detection of antiproton number ~ Schottky pickup



$$V = l_s \frac{di}{dt} + \frac{1}{c_s} \int i dt + r_s i$$

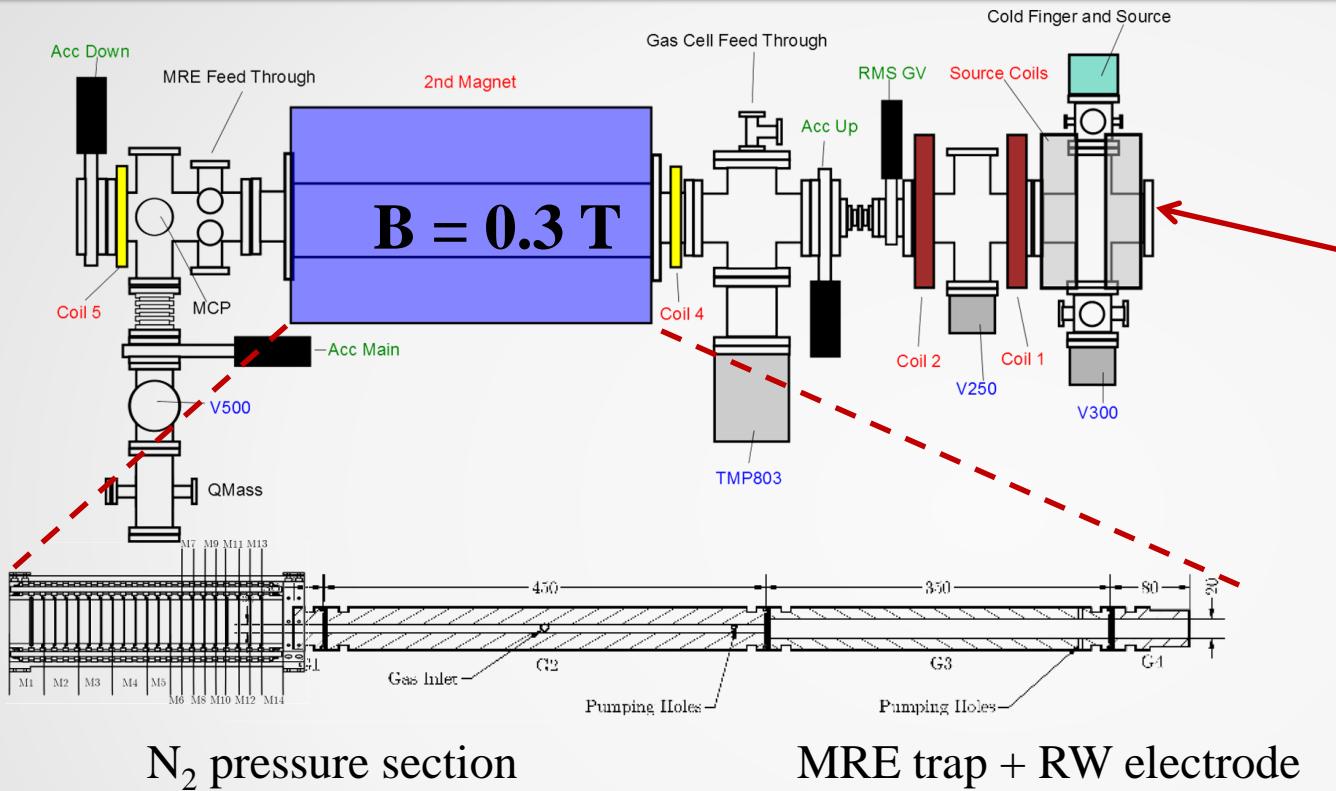
$$I = ni, l_n = l_s / n, c_n = c_s n, r_n = r_s / n$$

$$l_s = \frac{4mz_0^2}{e^2 d_1^2}, c_s = \frac{e^2 d_1^2}{4m\omega_z^2 z_0^2}, r_s = \frac{4\gamma mz_0^2}{e^2 d_1^2}$$



[7] D. J. Wineland and H. G. Dehmelt, *J. Appl. Phys.* **46**, 919 (1975) [8]: X.Feng, M.Charlton, M.Holtzscheiter, *et al.*, *J. Appl. Phys.* **79**, 8 (1996)

Positron trap



Gas cell: **Positrons** interact with N₂ and CF₄ and excite discrete, rotational and vibrational states of CF₄ and N₂ losing energy in this process.

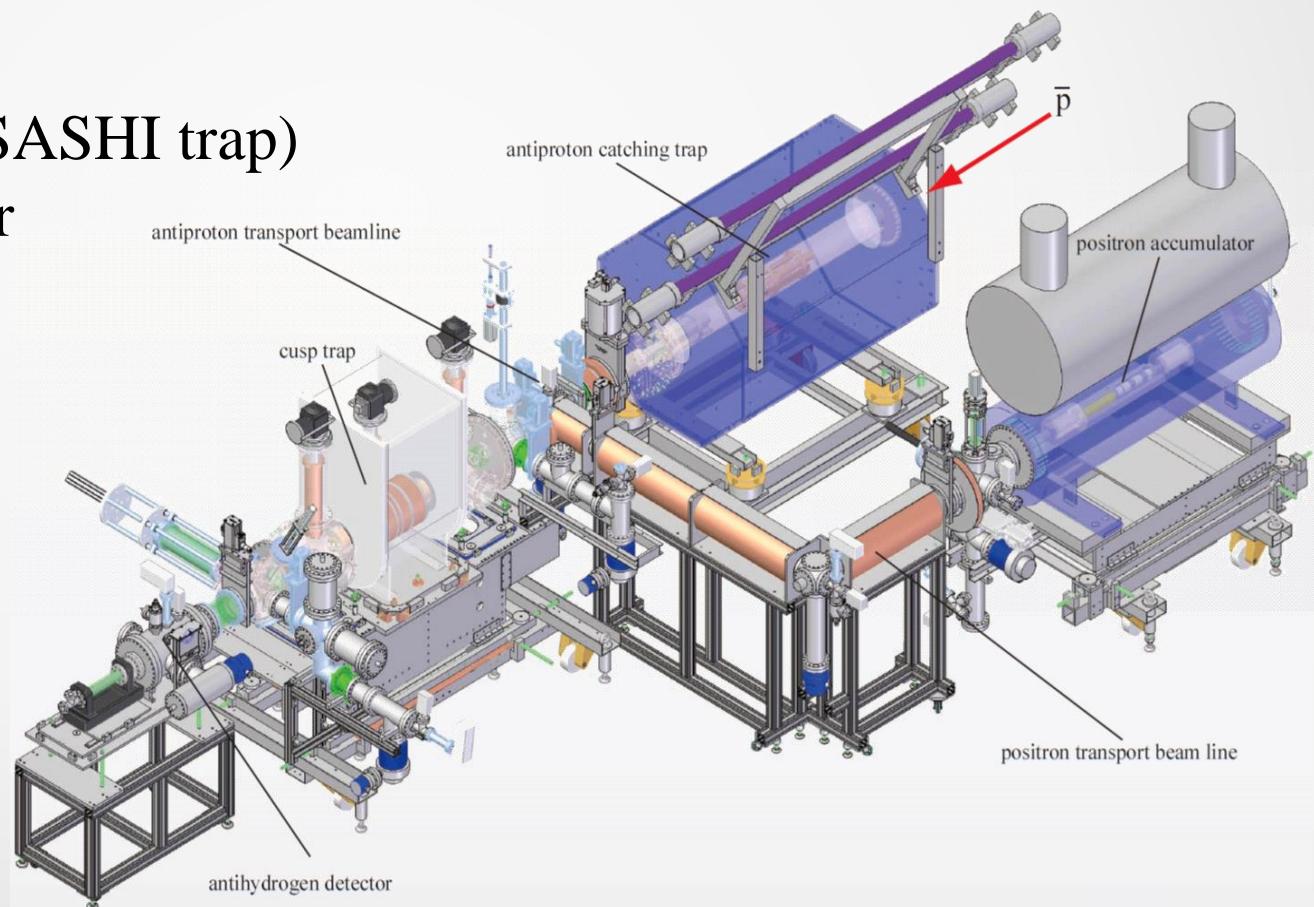
Tuning of gas cell (1V range) potentials according to competition with the positronium formation cross-section (=loss mechanism)

$$10^6 \text{ e}^+/\text{15s}$$

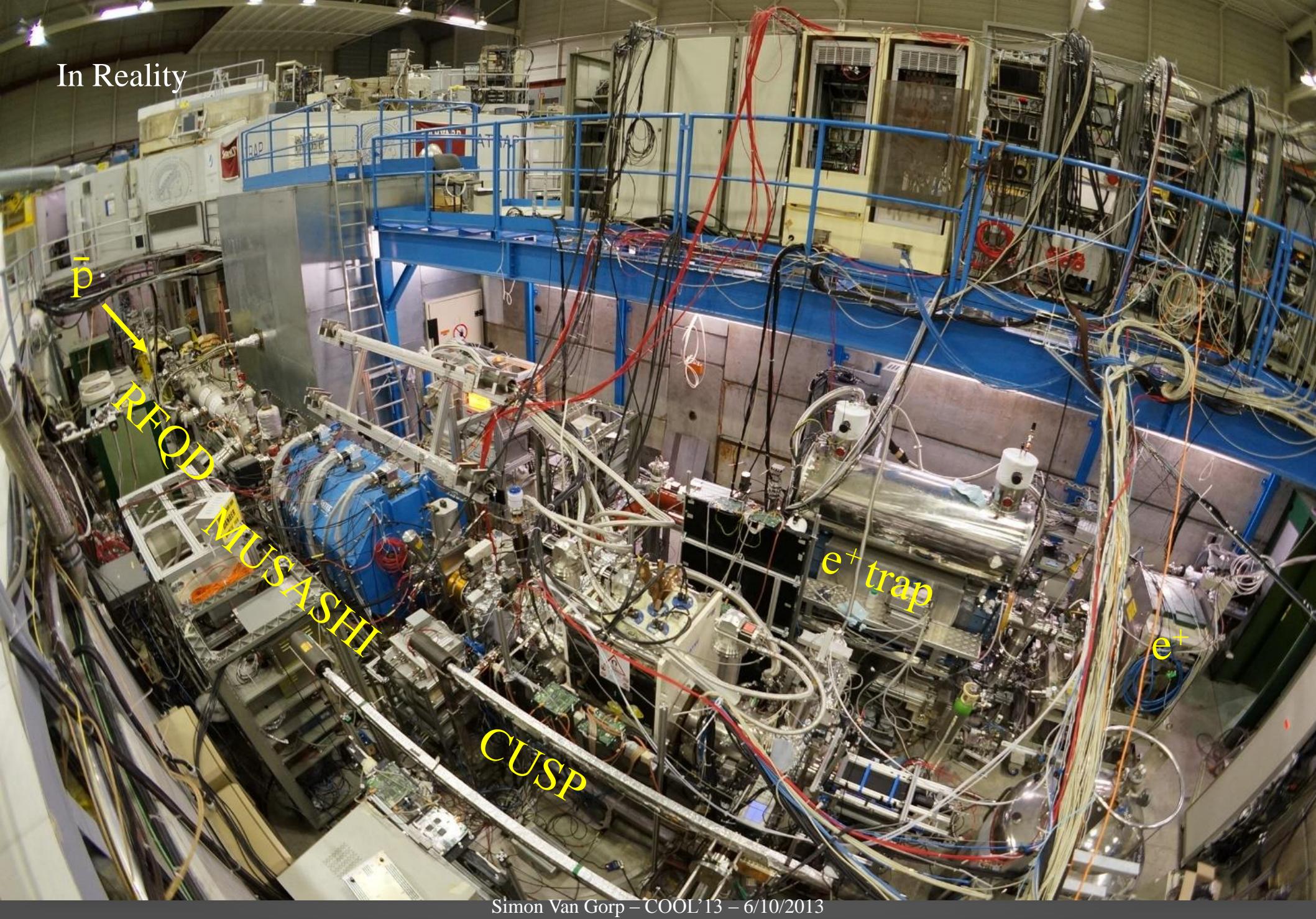
From 5.3 MeV \bar{p} to sub eV \bar{H} : the ingredients!

We need antihydrogen → has to be synthesized first

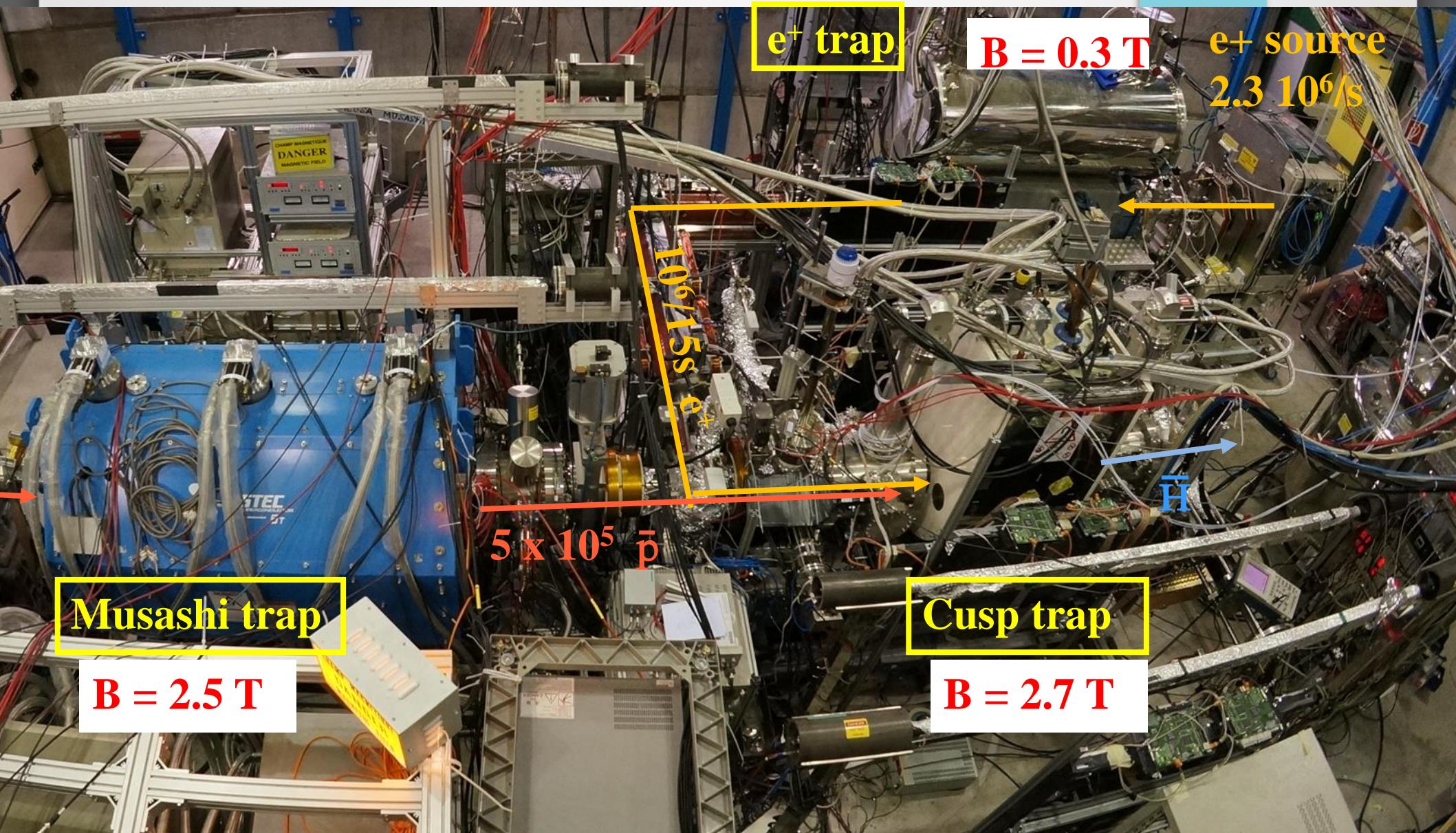
- 1) \bar{P} accumulator (MUSASHI trap)
- 2) Positron accumulator
- 3) CUSP trap
- 4) Spin flip cavity
- 5) Sextupole Magnet
- 6) Detectors



In Reality



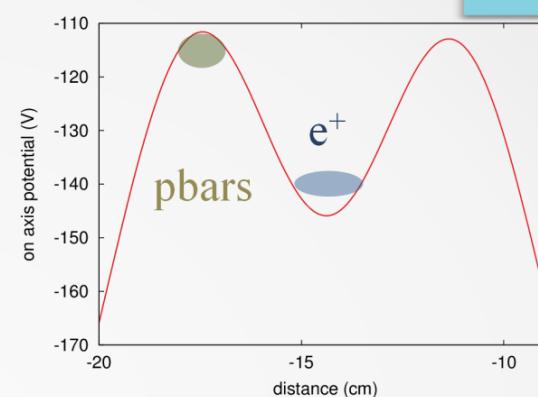
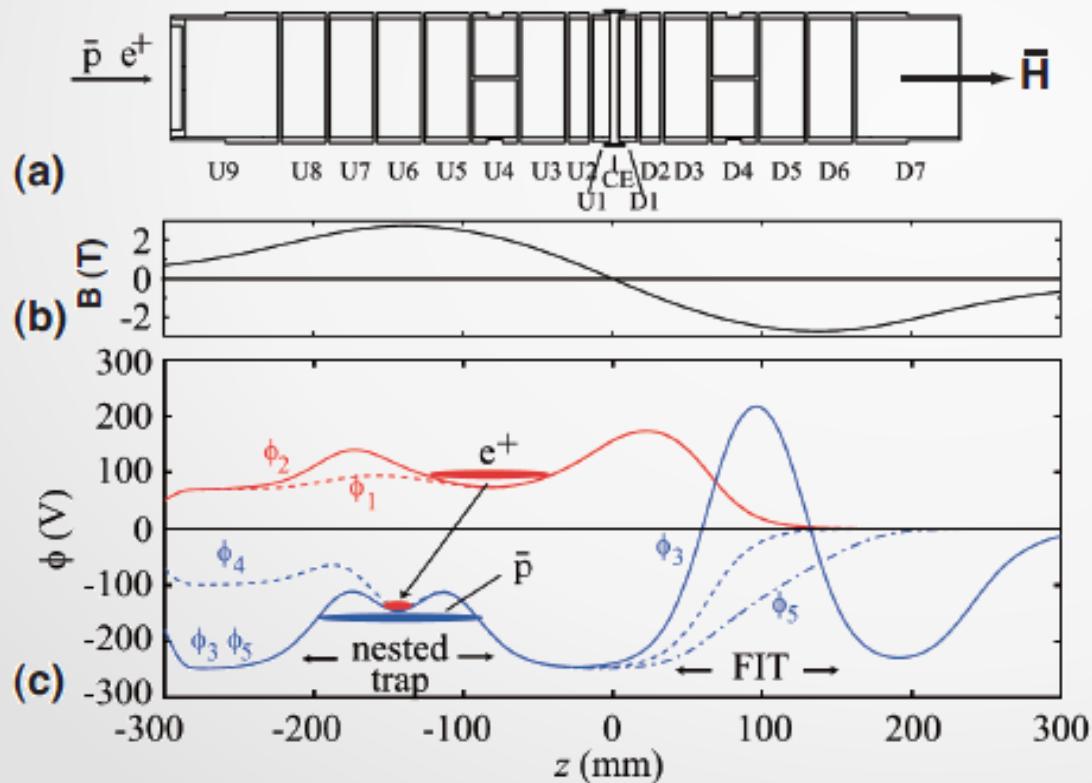
Experiment layout



How to create antihydrogen ?

Recipe

- 1) Load positrons ($6 \times 10^6 e^+$ /30 shots)
- 2) Bring positrons in a nested potential
- 3) Inject antiprotons (5×10^5)



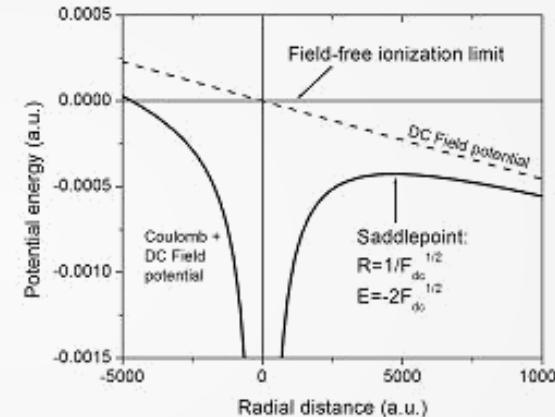
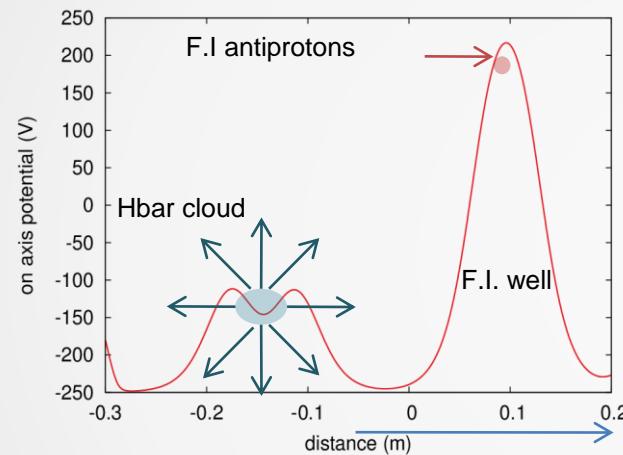
Antihydrogen production rate

Stops after ~ 10 s since:

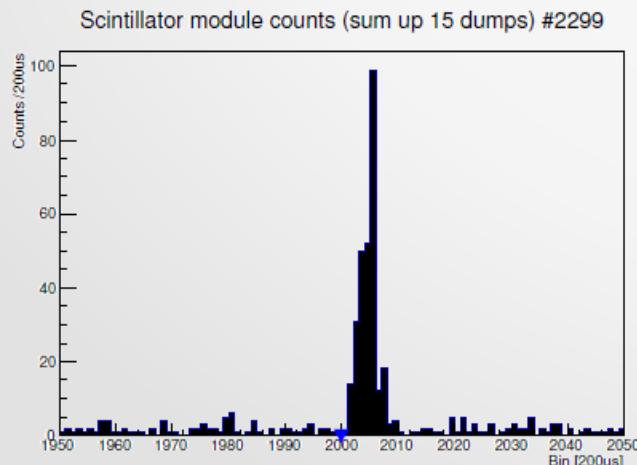
- 1) \bar{P} cools down with secondary e^- s from \bar{p} annihilations
 - 2) \bar{P} loses energy from scattering in e^+ cloud
- $\rightarrow \bar{p}$ cloud axially separated from e^+ cloud \rightarrow no \bar{H} formation

Detection of Antihydrogen

- 1) Neutral \bar{H} escapes from trap
- 2) Apply **field ionization** well \rightarrow strip positron, catch \bar{p}
- 3) Release field ionization well

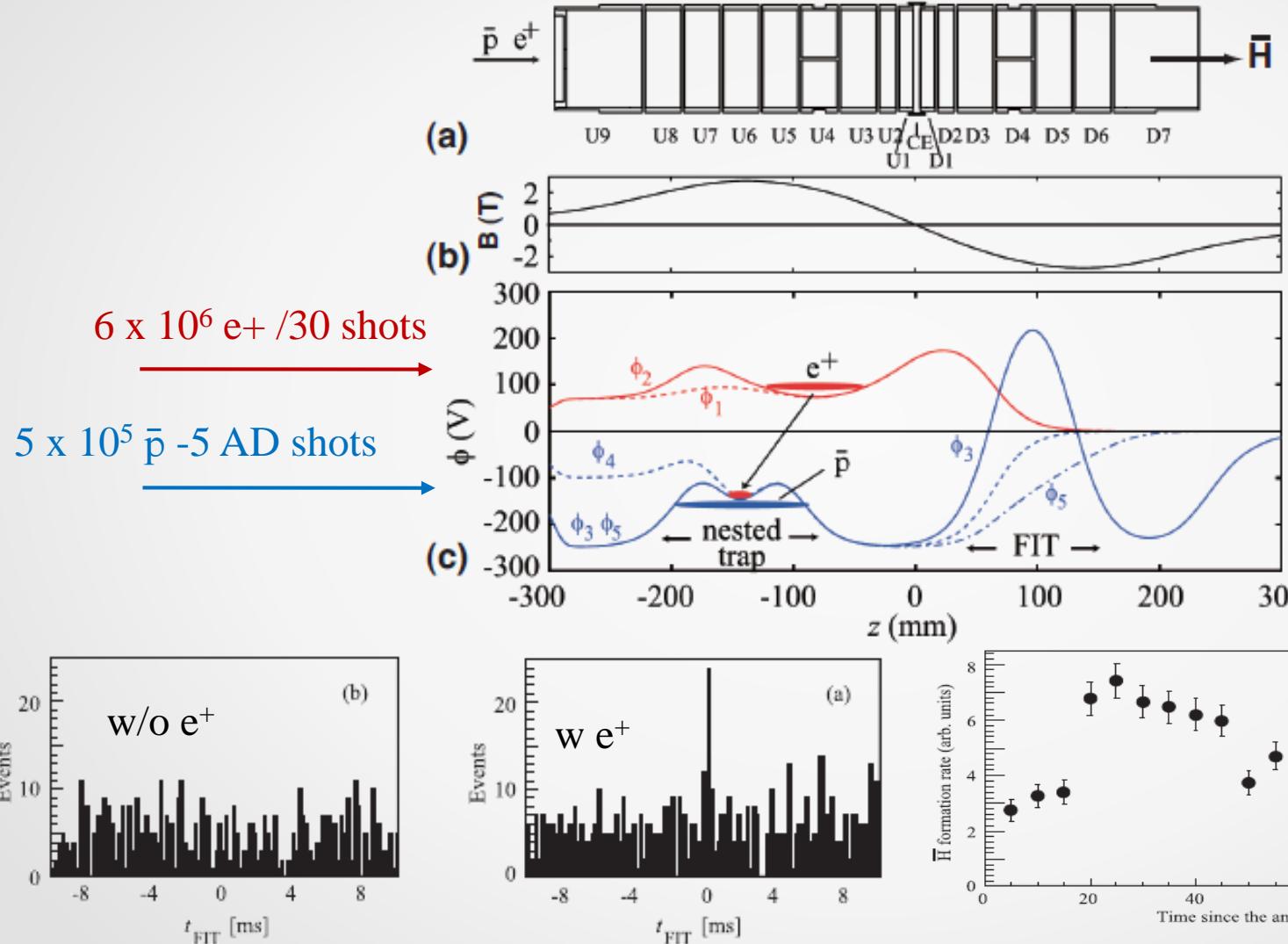


$$F = \frac{3.2 \cdot 10^8}{n^4} V/cm$$



Result after release
No peak without positrons
Clear indication for production of \bar{H}

Antihydrogen production in 2010

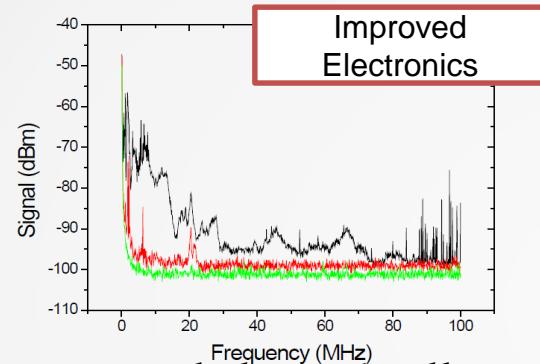


Direct injection of
Antiprotons to mix
With positrons

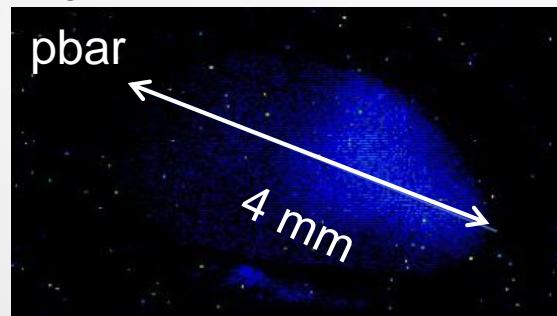
First \bar{H} results! Proof of principle of the CUSP field

Improvements since 2010

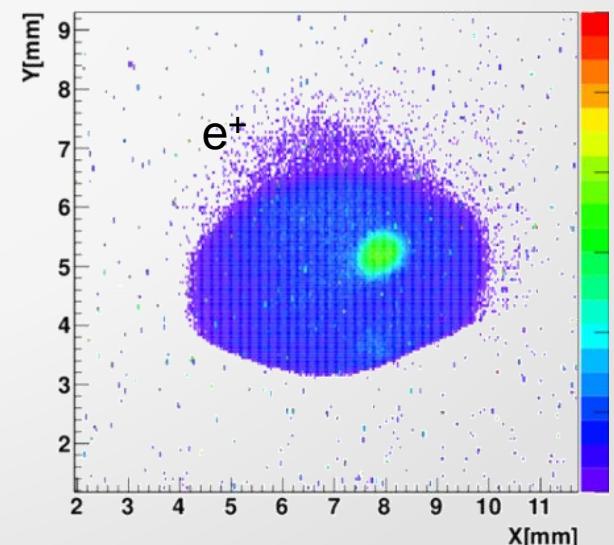
- 1) Better thermal contact (MRE: 18.5 K \rightarrow 12 K)
- 2) Extra aperture ($\varnothing = 50$ mm) to reduce thermal radiation
- 3) Electronic noise filters for the CUSP



- 4) Second RW segmented electrode allowed direct \bar{p} compression (100 kHz \rightarrow 200 kHz in 200 s)



- 5) Rare Gas moderator for positrons 60 million e^+ in 45 minutes (factor 15-20 improvement)
- 6) New \bar{H} detector
- 7) ...

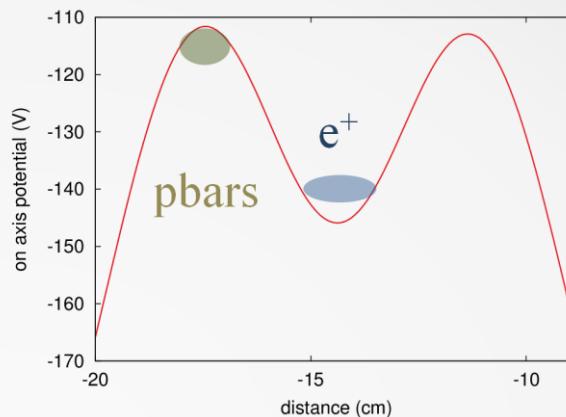
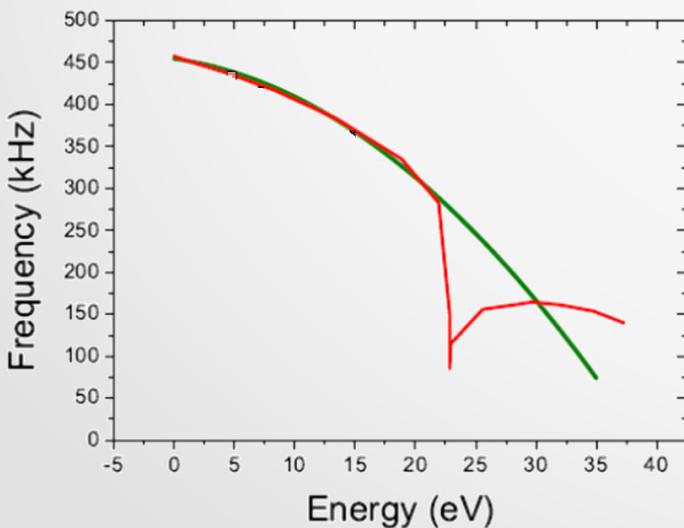


How to improve on the \bar{H} yield? → Autoresonant (AR) Excitation

Sweep the axial frequency to catch pbars and “park” them into e^+ cloud
 Note: E_{kin} is unaffected

Potential $\phi(z) = V_0 \sum_{k=0} C_k z^k$

Axial Frequency Scaling $\frac{3}{4} \left(\frac{C_4}{C_2^2} + \frac{5}{4} \frac{C_6}{C_2^3} \frac{E_z}{qV_0} \right) \frac{\omega_p}{qV_0}$

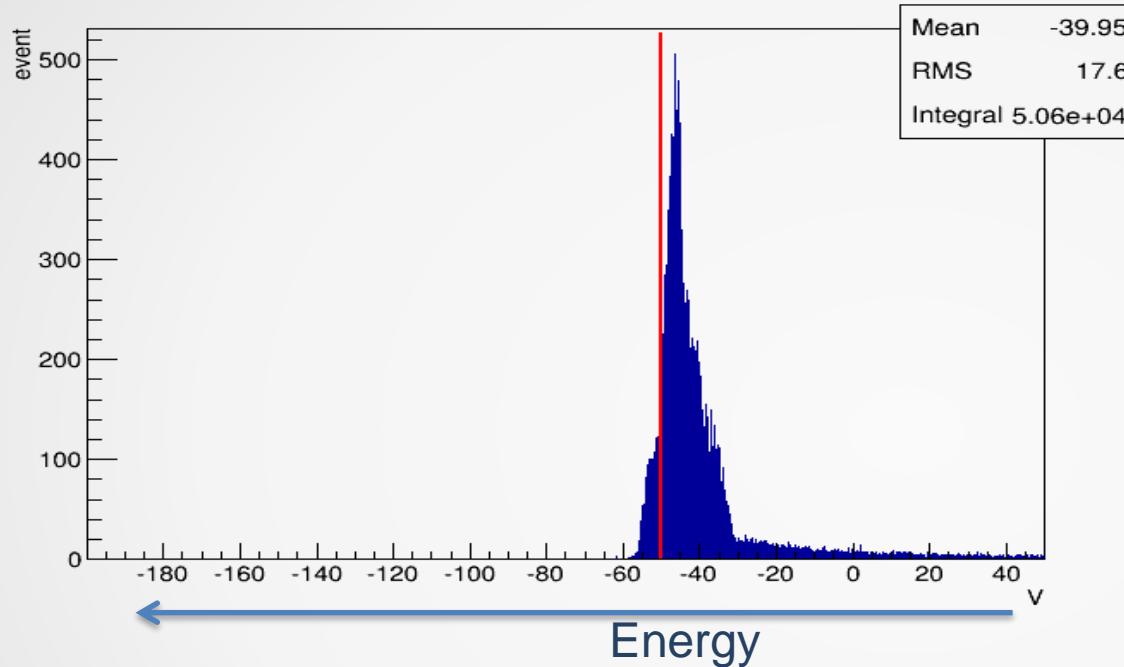


Exciting the particles as by a swept rf with a certain stop frequency:
Definition of interaction energy

- [7] J. Fajans, E. Gilson and L. Friedland. Autoresonant excitation of a collective nonlinear mode, *Phys. Plasmas*, **6** 4497, 1999.
 [8] Andresen et al., *Phys. Rev. Lett.* 106, 025002 (2011)

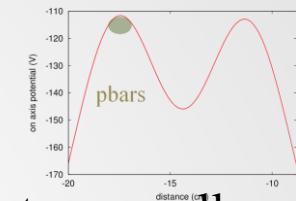
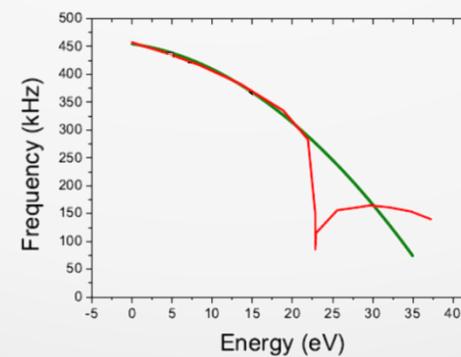
Autoresonance excitation (II)

Differential counts #1568

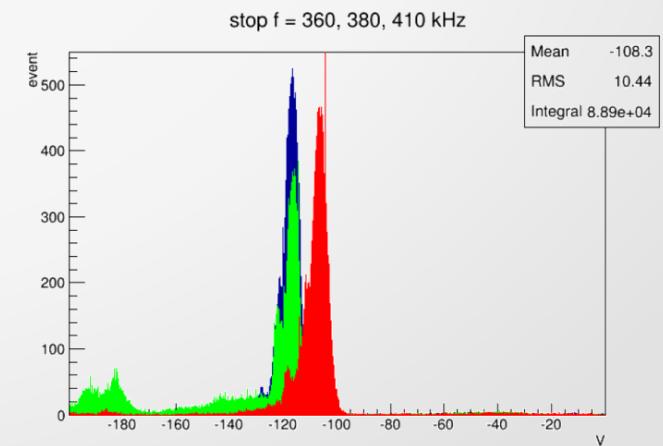


AR recipe

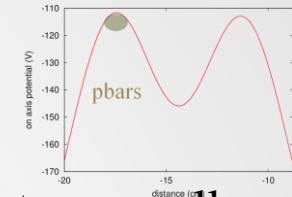
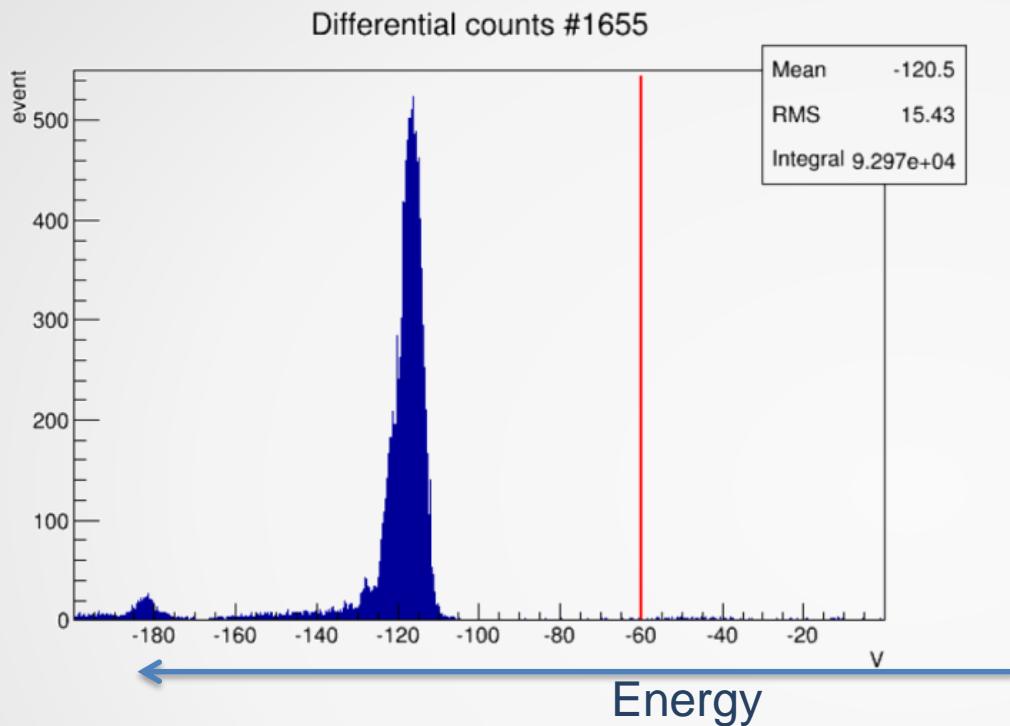
- 1) e^- cooling (30s)
- 2) e^- kickout
- 3) \bar{p} compression
(100 \rightarrow 250 kHz, 200s, 10 V
+ 220 kHz, 200 s, 10 V)
- 4) AR: 600 \rightarrow 380 kHz, 22 ms, 4 V
targeted slightly lower than the level
of e^+ well



- 1) \bar{p} stored in the upstream well
- 2) Energy distribution is measured by lowering the upstream electrode and counting #annihilations
- 3) The red line coincides with an empty trap (for $r=0$)

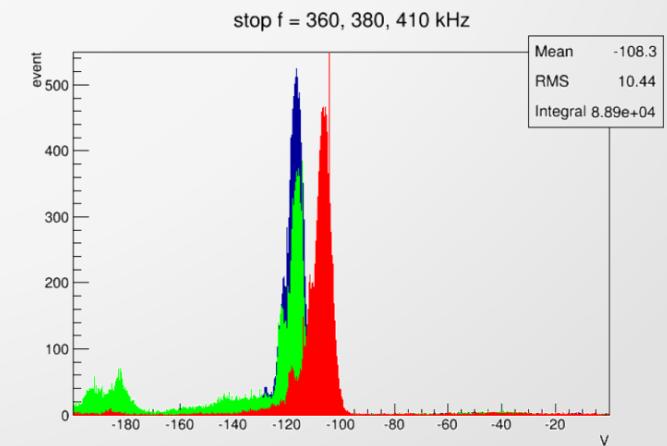
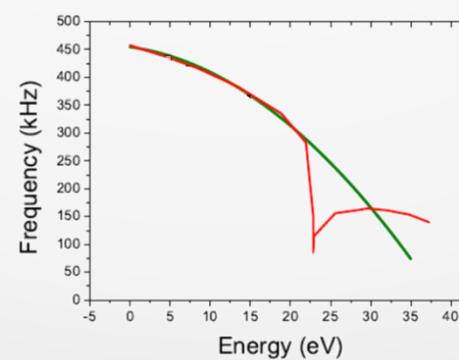


Autoresonance excitation (II)



- 1) \bar{p} stored in the upstream well
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- AR recipe**
- 1) e^- cooling (30s)
 - 2) e^- kickout
 - 3) \bar{p} compression
(100 \rightarrow 250 kHz, 200s , 10 V
+ 220 kHz, 200 s, 10 V)
 - 4) AR: 600 \rightarrow 380 kHz, 22 ms, 4 V
targeted slightly lower than the level
of e^+ well



Summary

Production of antihydrogen demonstrated.
Several major improvements resulted in a higher antihydrogen yield ...

Analysis of the 2012 data is being finalized and results will be presented soon !

Stay tuned!

Thank you + Special thanks to the AD operators/... for their support and \bar{p}



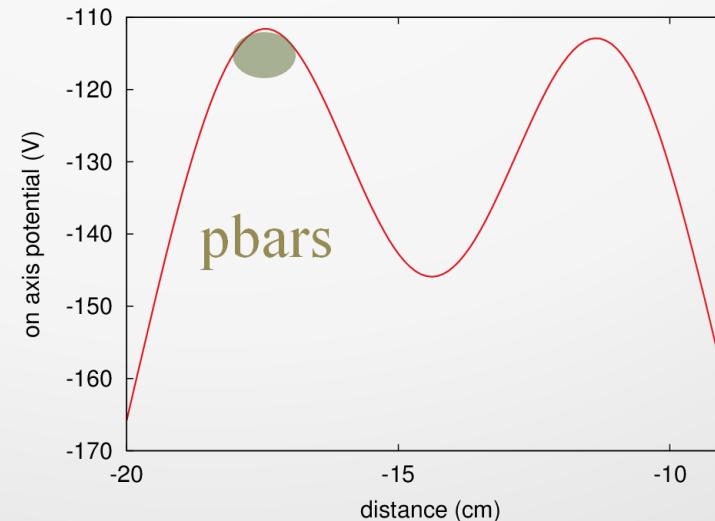
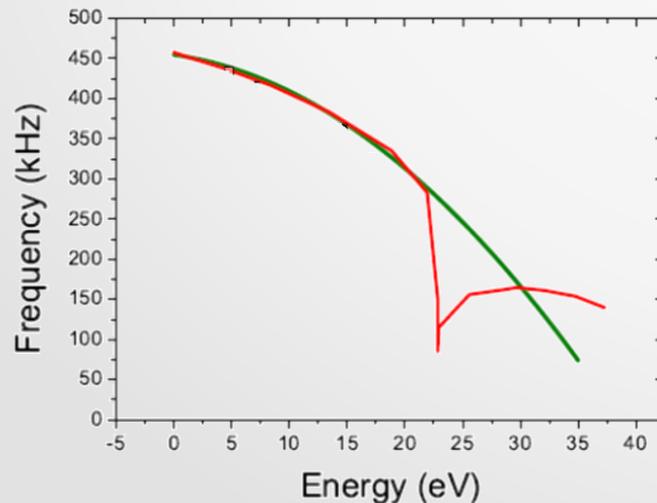
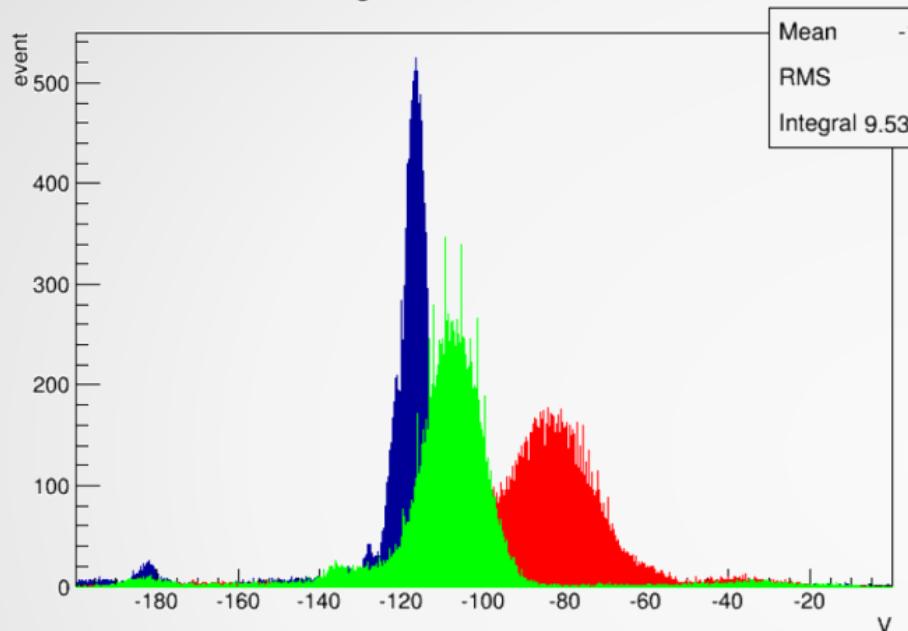
Antiproton trap
Positron trap
CUSP trap
 \bar{H} detector
CUSP detectors
SMI cavity
+sextupole
+ hodoscope

The boss

Backup slides

A.R. : wait time after A.R.

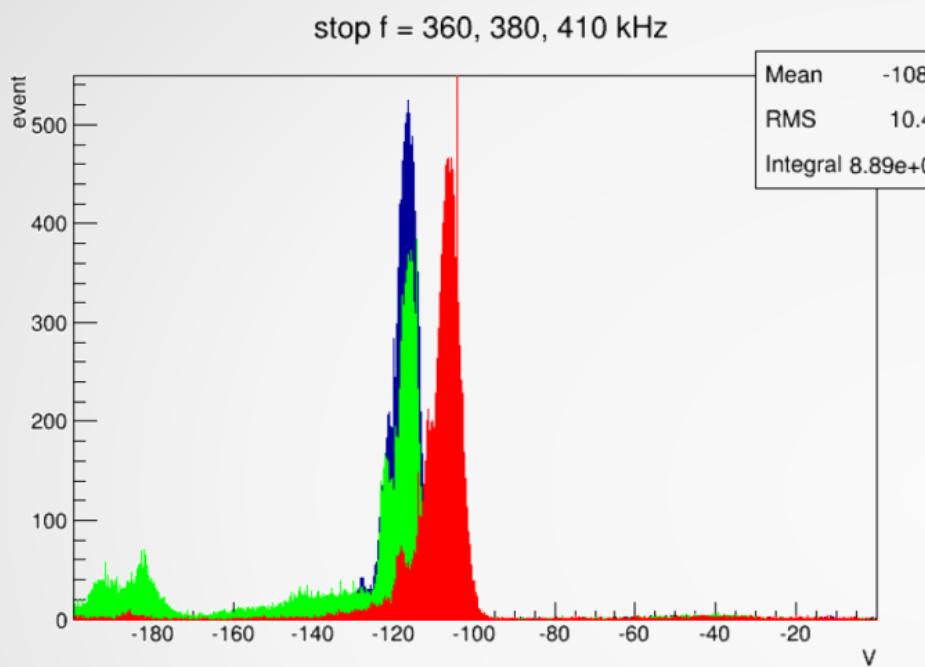
waiting after A.R. t = 0s, 5s, 20s



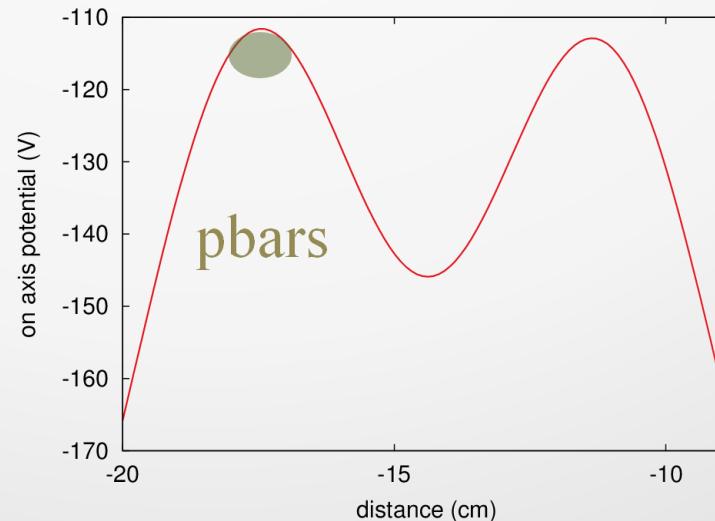
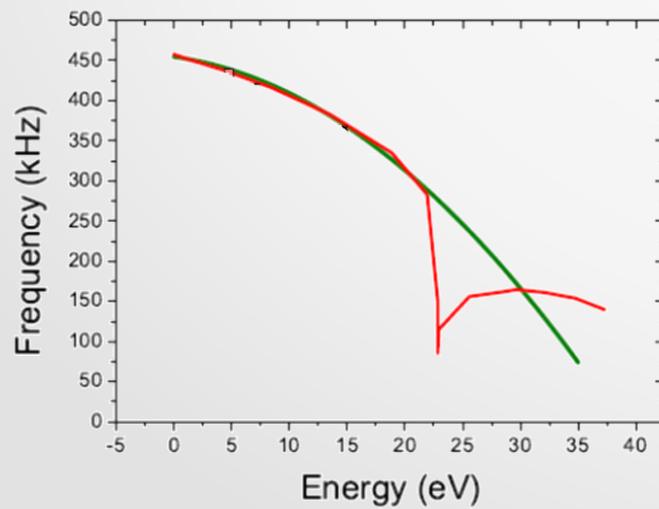
AR recipe

- 1) e^- cooling (30s)
 - 2) e^- kickout
 - 3) Pbar compression
($100 \rightarrow 250$ kHz, 200s , 10 V
+ 220 kHz, 200 s, 10 V)
 - 4) AR: $600 \rightarrow 380$ kHz, 22 ms, 4 V
- targeted slightly lower than the level of e^+
Well. wait: 0 s, 5 s, 20 s after AR
→Pbars cool again and fall again in the well

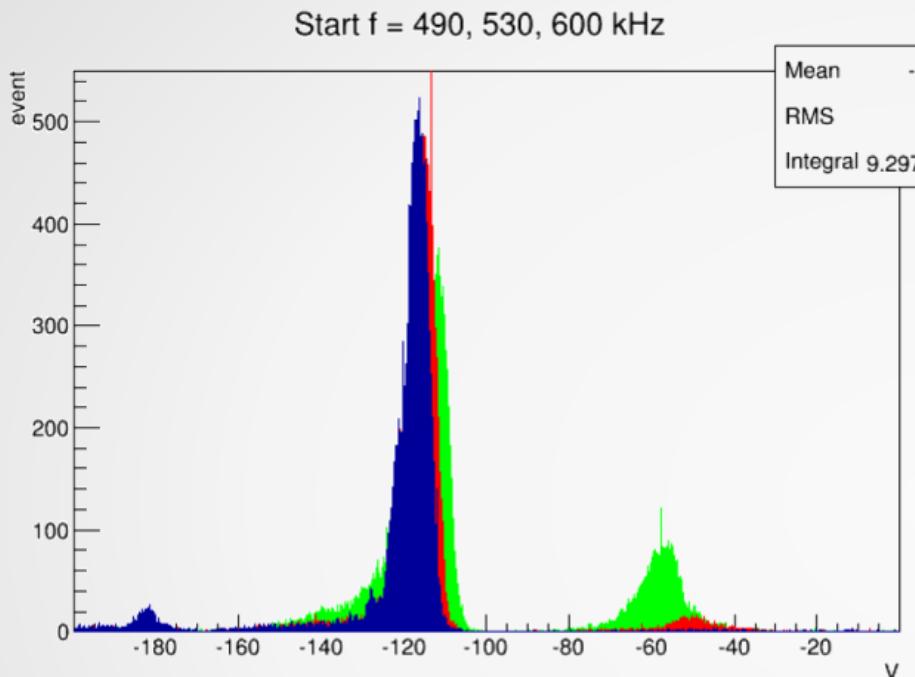
A.R.



- AR recipe
- 1) e^- cooling (30s)
 - 2) e^- kickout
 - 3) Pbar compression
(100 \rightarrow 250 kHz, 200s , 10 V
+ 220 kHz, 200 s, 10 V)
 - 4) AR: 600 \rightarrow f kHz, 22 ms, 4 V
scanned stop frequency f = 360, 380,
410 kHz

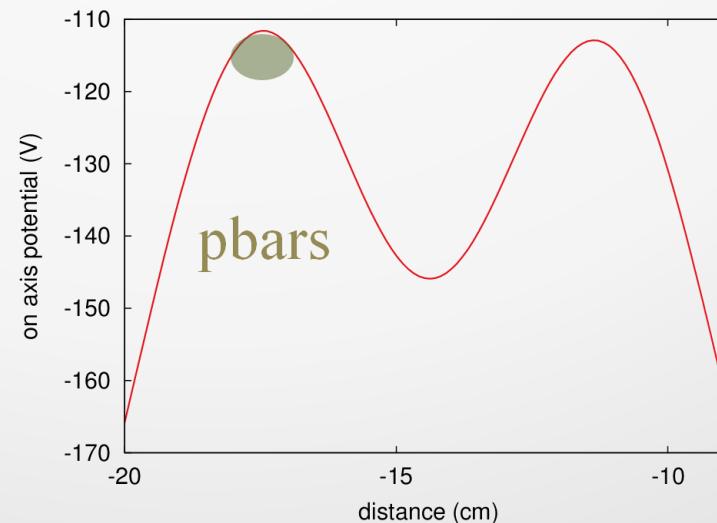
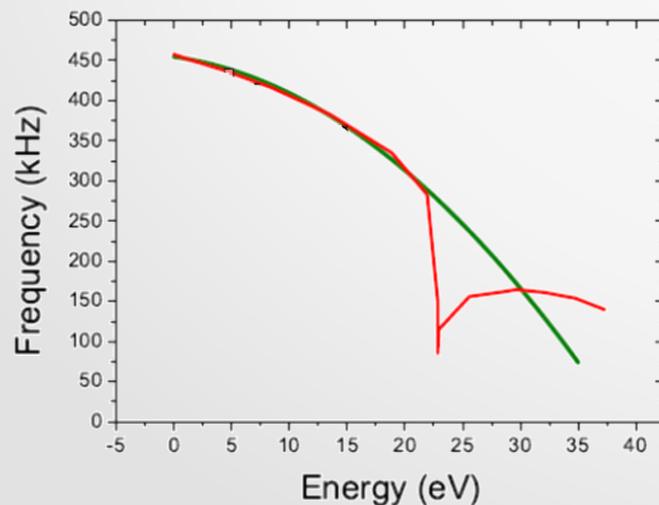


A.R. : change start frequency

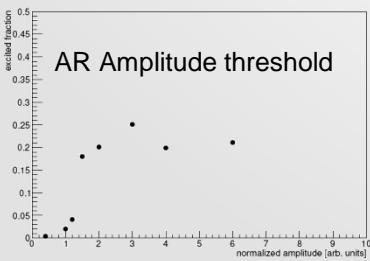
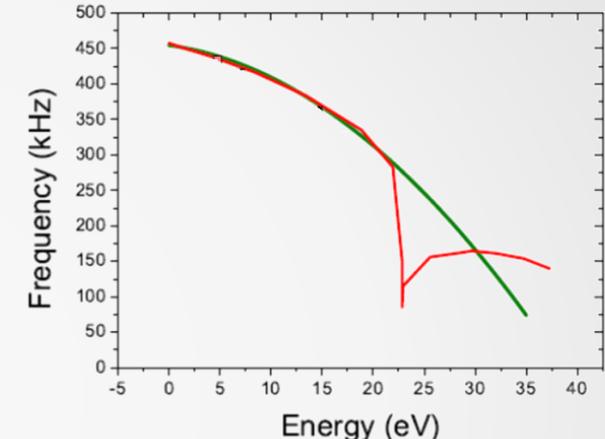
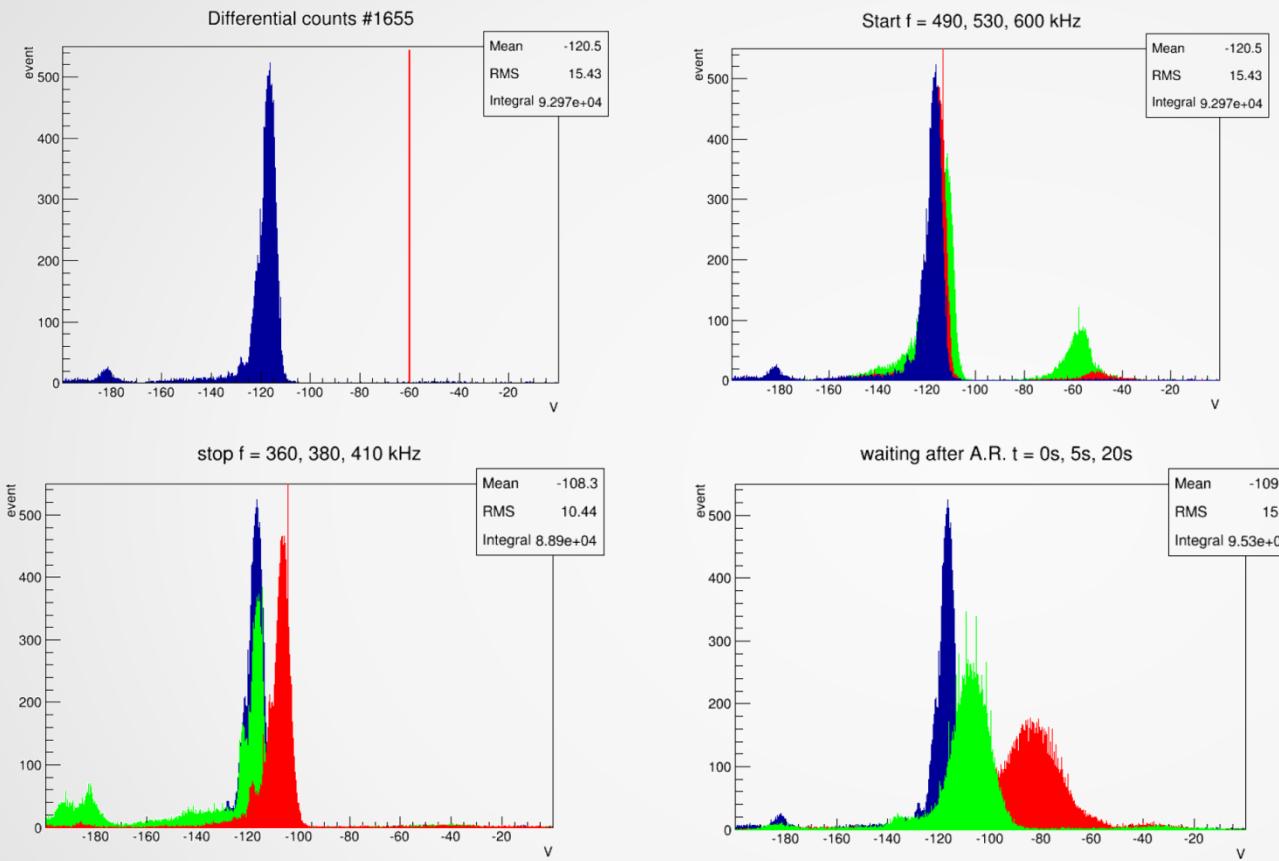


AR recipe

- 1) e^- cooling (30s)
- 2) e^- kickout
- 3) Pbar compression
(100 \rightarrow 250 kHz, 200s , 10 V
+ 220 kHz, 200 s, 10 V)
- 4) AR: $f \rightarrow 380$ kHz, 22 ms, 4 V
scanned start frequency $f = 490, 530,$
 600 kHz

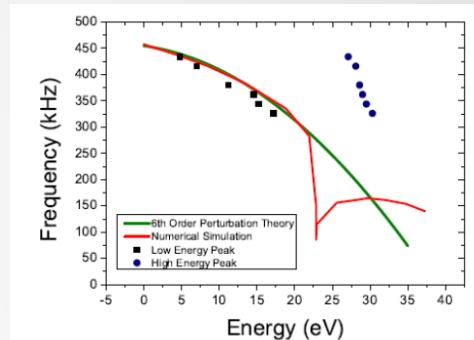
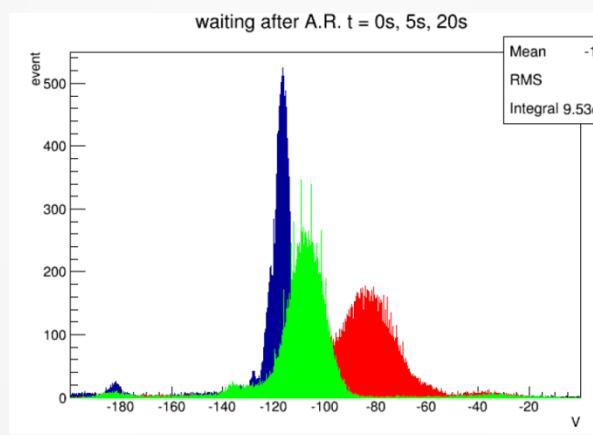
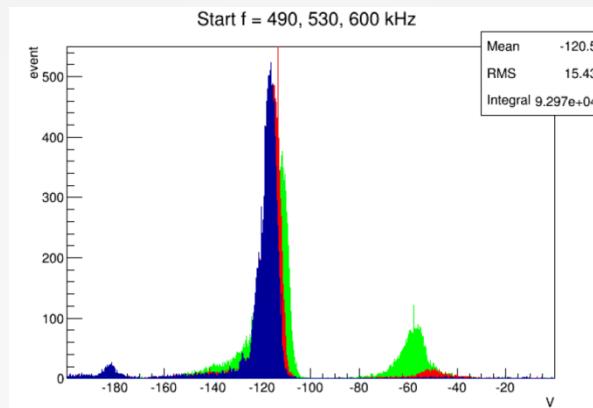
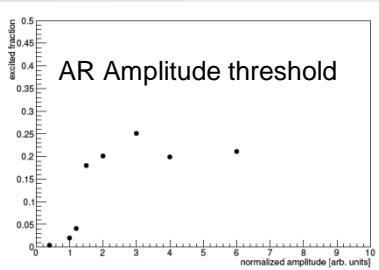
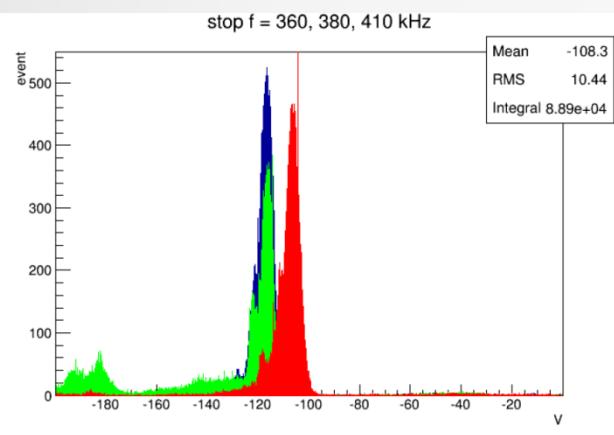
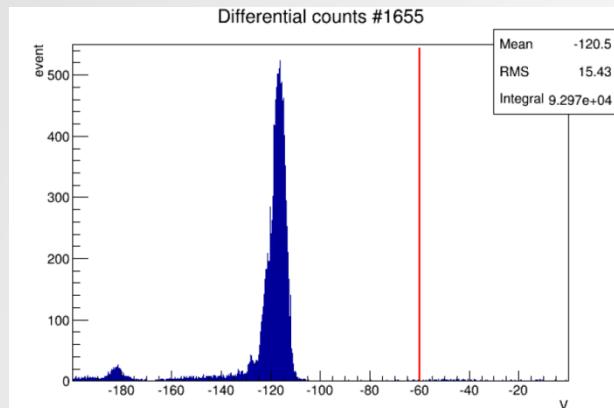


A.R is understood.



AR works as expected theoretically and is a tool that can be used to produce \bar{H}

A.R.



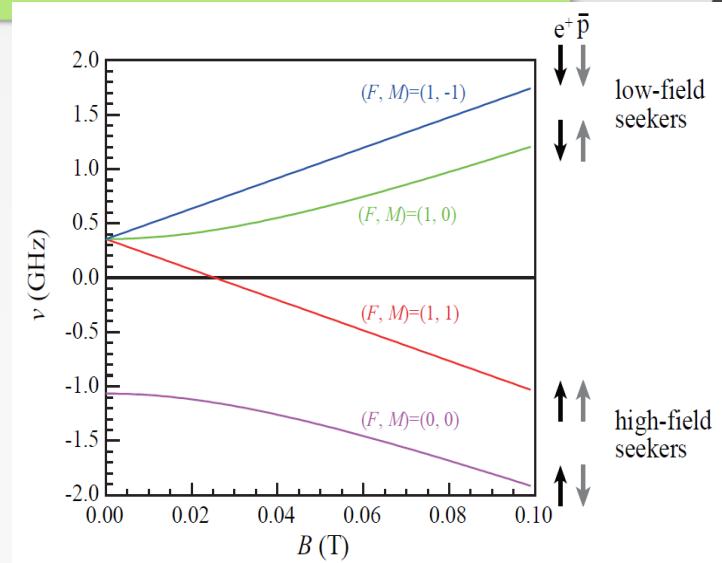
AR recipe

- 1) e^- cooling (30s)
- 2) e^- kickout
- 3) Pbar compression
(100 \rightarrow 250 kHz, 200s , 10 V
+ 220 kHz, 200 s, 10 V)
- 4) AR

Observations: as expected!

Spectroscopy

□ 1.420 405 751 766 7(9) GHz



- Correction: QED and proton/antiproton structure – level 10-6.
 - high precision proton g-factor measurement: constraints on antiproton mass
 - We aim at 10-6 or better.

Limitations

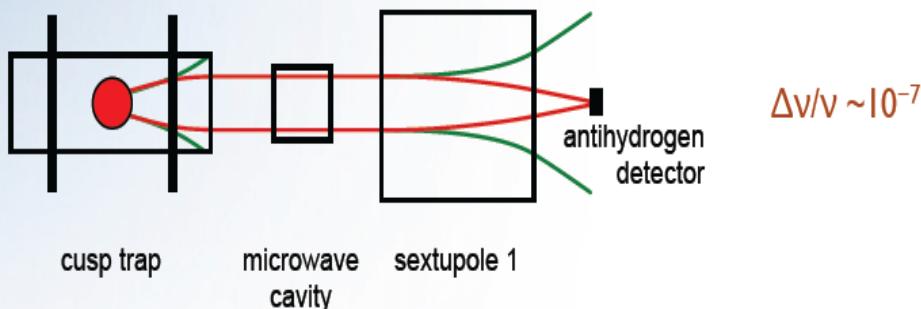
What limits the resolution of a Rabi type experiment?

- 1.) TOF
- 2.) Magnetic field homogeneity
- 3.) ...thus beam profile

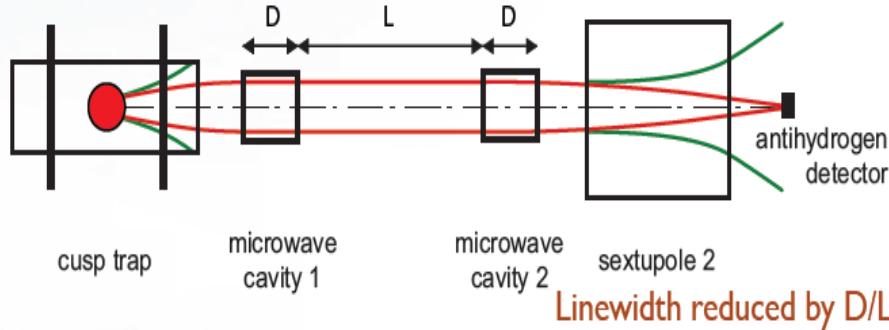
Experiment: Just limited by TOF (averaging over inhomogeneity) which can be arbitrarily long (

Proposals to measure more precisely

- Phase I (ongoing): Rabi method

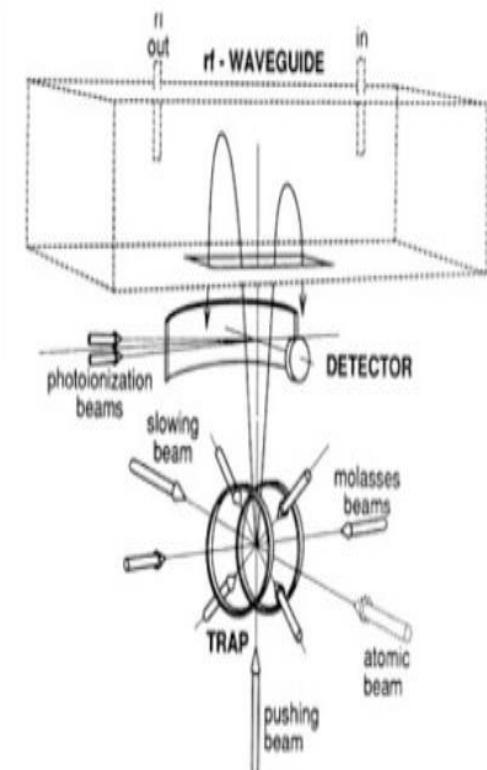


- Phase 2: Ramsey separated oscillatory fields



- Phase 3: trapped $H^{\bar{}}_{} \bar{}$

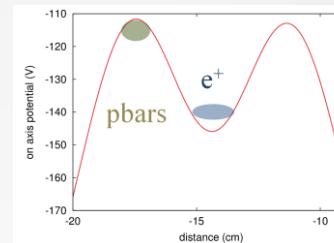
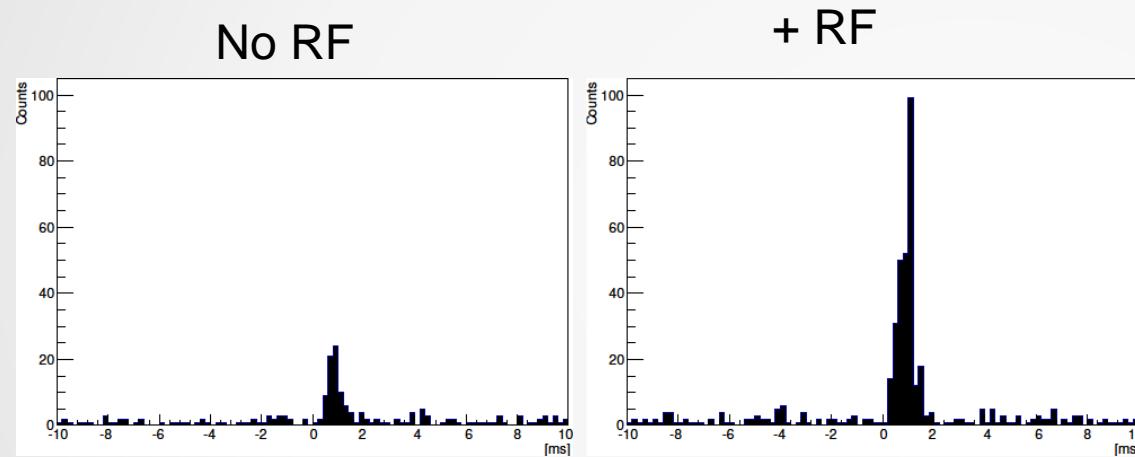
- Hyperfine spectroscopy in an atomic fountain of antihydrogen
- needs trapping and laser cooling outside of formation magnet
- slow beam & capture in measurement trap
- Ramsey method with $d=1\text{ m}$
- $\Delta f \sim 3\text{ Hz}$, $\Delta f/f \sim 2 \times 10^{-9}$



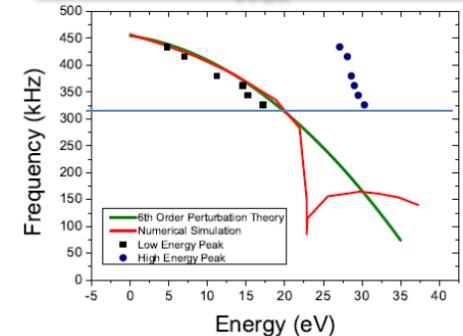
Far Future !

RF Assisted Direct Injection

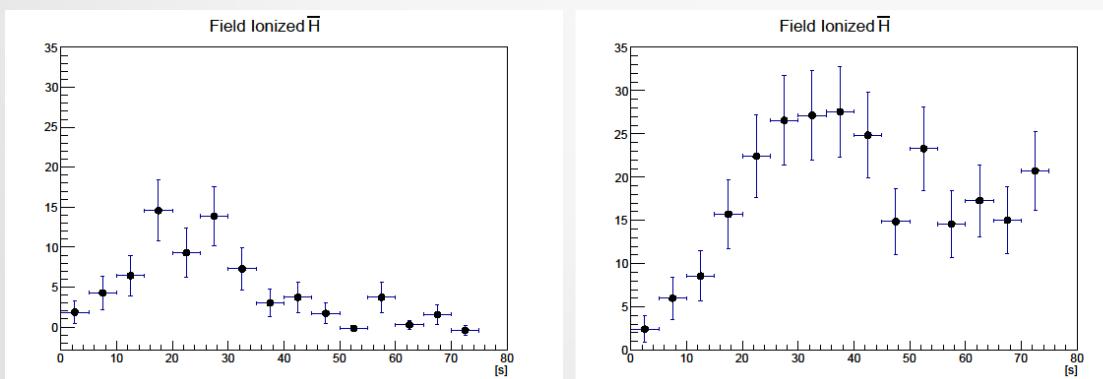
- 1) $40 \times 10^6 e^+ / 5 \times 10^5 pbars$
- 2) Inject directly from MUSASHI trap
- 3) Apply RF during interaction



Frequency Scaling in Nested Well



4 fold improvement in Hbar counts



Hbar production continues after 30 seconds

Produced encouraging results – further evaluation in progress

How to produce Specifications

positron

27mCi ^{22}Na

($<100\text{keV}$, $\sim 10^9/\text{s}$)

W moderator, N2buffer gas

($<3\text{ eV}$, $\sim 10^6/\text{s}$)

Pre-accumulator

(130eV , $2 \times 10^5/\text{30s}$)

antiproton (pbar)

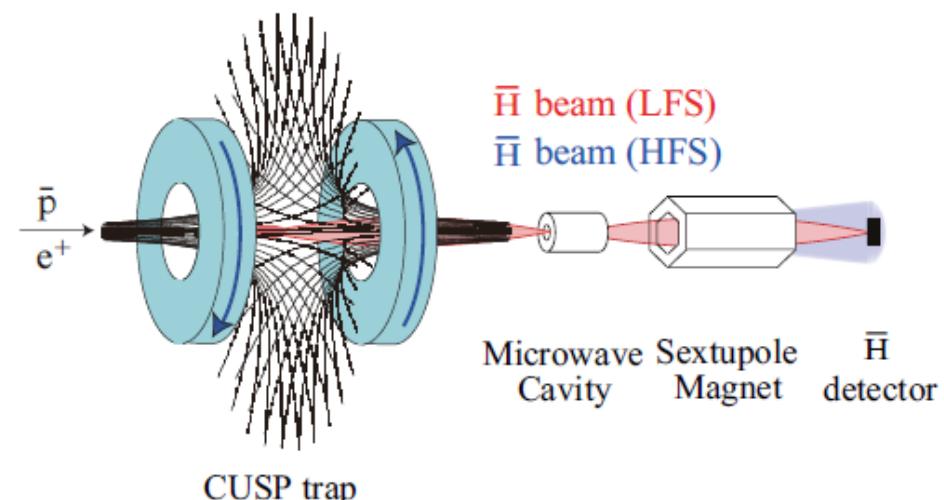
CERN PS ($3.5\text{GeV}/c$, 5×10^7)

AD (5MeV , 10^7)

RFQD (100keV , 5×10^6)

MUSASHI

(150eV , 5×10^5)

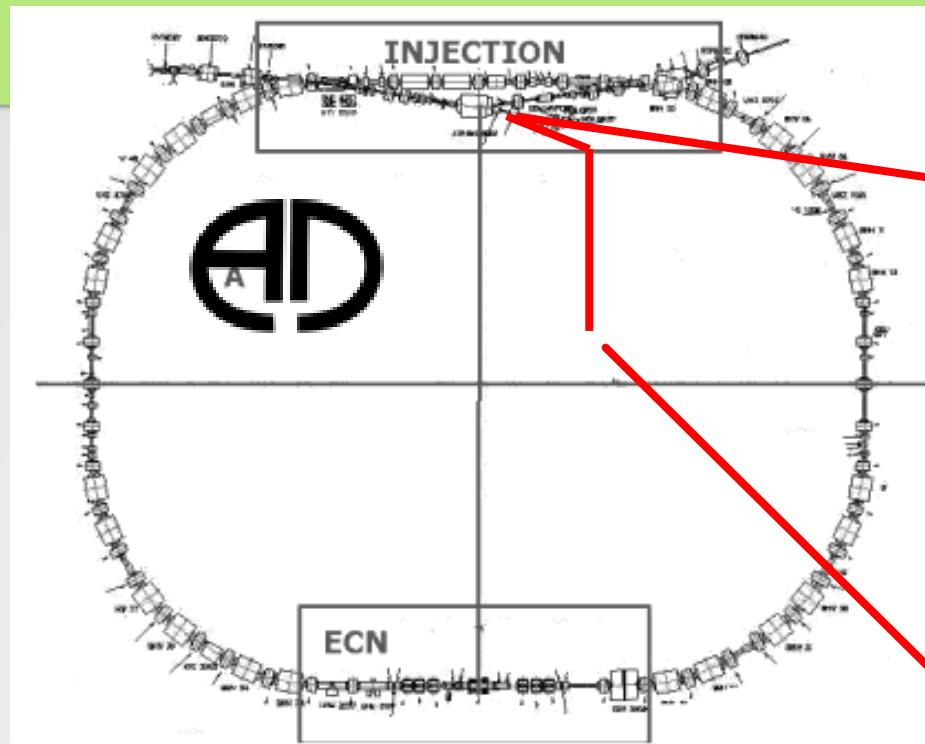


CUSP trap

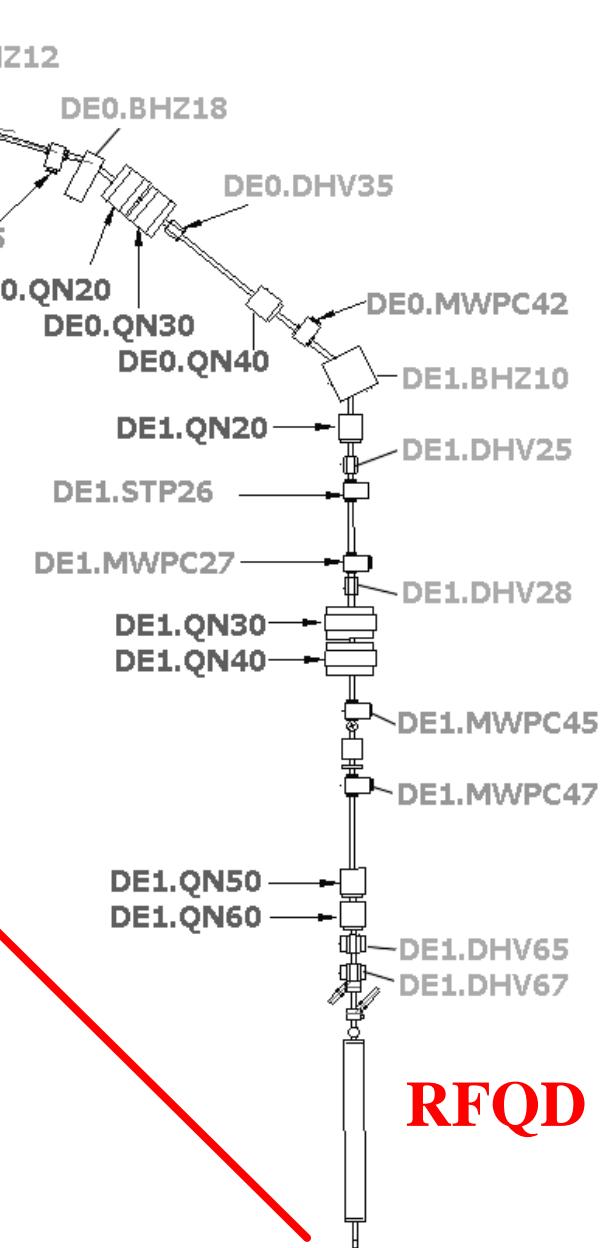
6×10^6 positron
($<0.1\text{eV}$, $n_{e+} \sim 10^8\text{cm}^{-3}$)

3×10^5 pbar
($< 20\text{ eV}$)

Low energy anti-proton beams (AD to RFQD)

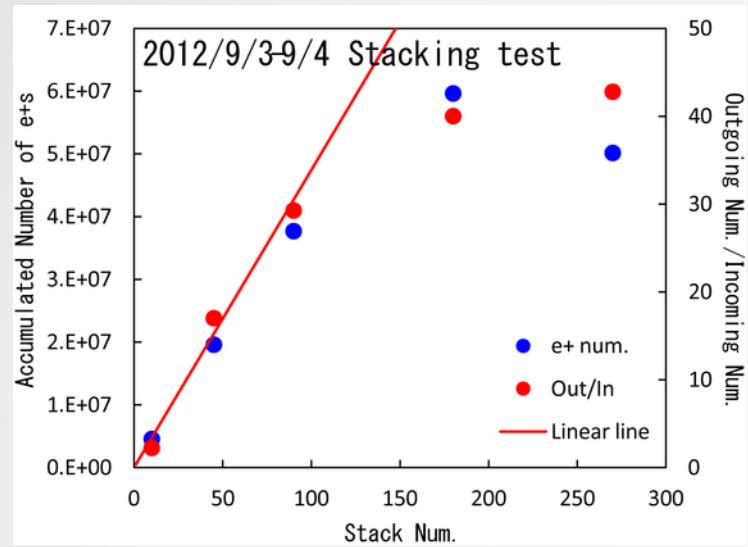


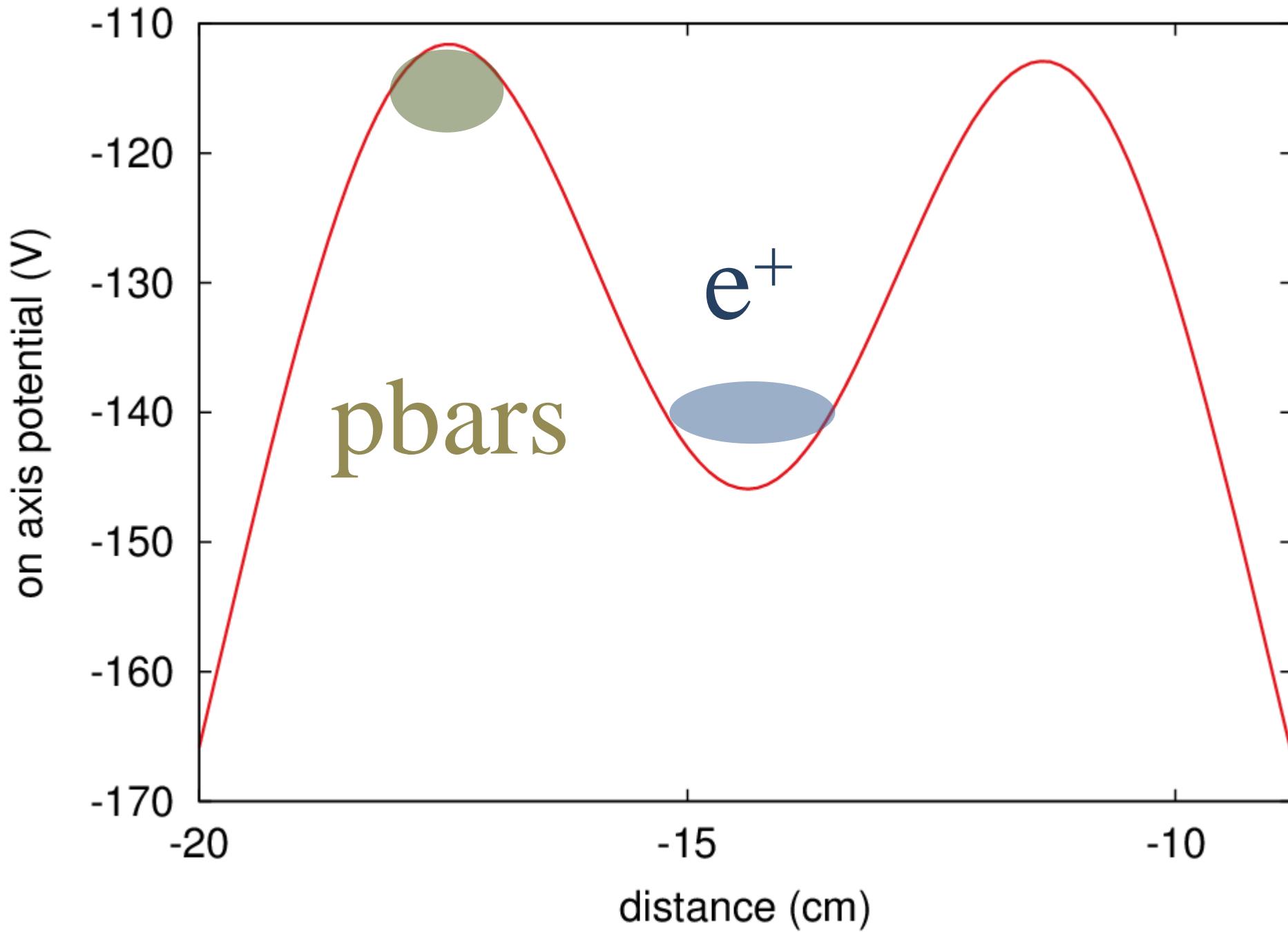
ASACUSA extraction line



Production target

- 3.5 GeV/c ~5x 10⁷ pbar stochastic cooling
- 2 GeV/c stochastic cooling
- 300 MeV/c stochastic cooling & e-coolling
- 100 MeV/c (~5 MeV) stochastic cooling & e-coolling
- ~ 10⁷ pbar pulse from AD (~100s cycle)
- 100 keV <5 x 10⁶ pbar with RFQD





How to produce \bar{H} (Low energy antiproton beam) Tank circuit signal with 3AD shot accumulated in MUSASHI

