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Homework 7: Linearized Gravity

1. Let you be the inverse of n. Then,

Up to first order in h,

Moreover

Thus $g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu}$. Therefore, using (3.31), we

have the Christoffel symbols

$$\Gamma^{\alpha}_{\beta Y} = \frac{1}{2} \left(\gamma^{\alpha \delta} - h^{\alpha \delta} \right) \left(h_{\beta \delta, Y} + h_{\gamma \delta, \beta} - h_{\beta Y, \delta} \right)$$

$$= \frac{1}{2} \gamma^{\alpha \delta} \left(h_{\beta \delta, Y} + h_{\gamma \delta, \beta} - h_{\beta Y, \delta} \right),$$

in our appropriate coordinates. Using (3.66) we

can now calculate the components of the Riemann tensor

$$R^{g}_{\sigma\mu\nu} = \frac{1}{2} \eta^{g\delta} \left(h_{\nu\delta,\sigma\mu} + h_{\sigma\delta,\nu\mu} - h_{\nu\sigma,\delta\mu} \right)$$

$$- \frac{1}{2} \eta^{g\delta} \left(h_{\mu\delta,\sigma\nu} + h_{\sigma\delta,\mu\nu} - h_{\mu\sigma,\delta\nu} \right)$$

$$= \frac{1}{2} \eta^{g\delta} \left(h_{\nu\delta,\sigma\mu} - h_{\nu\sigma,\delta\mu} - h_{\mu\delta,\sigma\nu} + h_{\mu\sigma,\delta\nu} \right),$$

We can further calculate the components of the Ricci tensor

$$R_{\mu\nu} = R^{9}_{\mu\rho\nu} = \frac{1}{2} \eta^{9\delta} (h_{\nu}\delta_{\mu} - h_{\nu}, \delta_{\rho} - h_{\rho}\delta_{\mu} + h_{\rho}\mu_{\nu}\delta_{\nu}),$$

$$= \frac{1}{2} (29 J_{\mu} h_{\nu} - 29 J_{\rho} h_{\mu\nu} - 2\mu J_{\nu} h + 29 J_{\nu} h_{\mu}s),$$

and the Ricci scalar

$$R = R^{\mu} = \frac{1}{2} \left(\partial^{g} \partial^{\mu} h_{\mu g} - \partial^{g} \partial_{g} h - \partial^{\mu} \partial_{\mu} h + \partial^{g} \partial^{\mu} h_{\mu g} \right)$$

$$= \partial^{g} \partial^{\mu} h_{\mu g} - \partial^{g} \partial_{g} h.$$

It is important to notice that, up to order he it doesn't matter whether we raise indices with y or g on the previous three tensors.

The Einstein tensor is finally

$$=\frac{1}{2}\left(3^{g}\partial_{\mu}h_{\nu\rho}-3^{\rho}\partial_{\rho}h_{\mu\nu}-3_{\mu}\partial_{\nu}h+3^{\rho}\partial_{\nu}h_{\mu\rho}-\gamma_{\mu\nu}\beta^{\rho}\partial_{\nu}h_{\rho\sigma}+\gamma_{\mu\nu}\beta^{\rho}\partial_{\rho}h\right)$$

$$=-\frac{1}{2}\left(3^{g}\partial_{\mu}h_{\nu\rho}-\frac{1}{2}\gamma_{\mu\nu}\beta^{\rho}\partial_{\nu}h\right)$$

$$+\frac{1}{2}\left(3^{g}\partial_{\mu}h_{\nu\rho}+2^{g}\partial_{\nu}h_{\mu\rho}-\frac{1}{2}\beta^{\rho}\partial_{\mu}h\right)$$

$$-\frac{1}{2}\gamma_{\mu\nu}\left(3^{g}\partial_{\nu}h_{\nu\rho}+\frac{1}{2}(3^{g}\partial_{\mu}h_{\nu\rho}-\frac{1}{2}3^{\rho}\partial_{\mu}g_{\nu\rho}h+3^{\rho}\partial_{\nu}h_{\mu\rho}-\frac{1}{2}3^{\rho}\partial_{\nu}g_{\rho}h\right)$$

$$-\frac{1}{2}\gamma_{\mu\nu}\left(3^{\rho}\partial_{\nu}h_{\rho\sigma}-\frac{1}{2}3^{\rho}\partial_{\mu}h_{\nu\rho}-\frac{1}{2}\gamma_{\mu\nu}\partial_{\nu}g_{\nu\rho}h+3^{\rho}\partial_{\nu}h_{\mu\rho}-\frac{1}{2}3^{\rho}\partial_{\nu}g_{\rho}h\right)$$

$$-\frac{1}{2}\gamma_{\mu\nu}\left(3^{\rho}\partial_{\nu}h_{\rho\sigma}-\frac{1}{2}3^{\rho}\partial_{\mu}h_{\nu\rho}-\frac{1}{2}\gamma_{\mu\nu}\partial_{\nu}g_{\nu}h+3^{\rho}\partial_{\nu}h_{\mu\rho}-\frac{1}{2}3^{\rho}\partial_{\nu}g_{\rho}h\right)$$

$$=-\frac{1}{2}3^{\rho}\partial_{\rho}h_{\mu\nu}+3^{\rho}\partial_{\nu}h_{\rho\sigma}-\frac{1}{2}\gamma_{\mu\nu}\partial_{\nu}g_{\nu}h$$
Einstein equations are
$$-\frac{1}{2}3^{\rho}\partial_{\rho}h_{\mu\nu}+3^{\rho}\partial_{\nu}h_{\rho\sigma}-\frac{1}{2}\gamma_{\mu\nu}\partial_{\nu}g_{\nu}h_{\rho\sigma}-g_{\mu\nu}=8\pi GT_{\mu\nu}-\frac{1}{2}\partial_{\nu}g_{\nu}h$$
2. a) Under such a coordinate transformation and up to trist order on g and g and g and g are have
$$\gamma_{\mu\nu}+h_{\mu\nu}=\gamma_{\mu\nu}=\gamma_{\mu\nu}=\gamma_{\mu\nu}g_{\nu}^{2}g_{\nu}^{2}g_{\nu}^{2}g_{\nu}^{2}$$

$$=(\gamma_{\mu\nu}+h_{\mu\nu})(\delta^{g}_{\mu}-\partial_{\mu}g^{g})(\delta^{g}_{\mu}-\partial_{\mu}g^{g})(\delta^{\sigma}_{\nu}-\partial_{\nu}g^{\sigma}+h_{\mu\nu})$$

$$=\gamma_{\mu\nu}-\gamma_{\mu\rho}g_{\nu}p_{\nu}g_{\nu}^{2}g_{\nu}^{2}g_{\nu}^{2}$$

$$=\gamma_{\mu\nu}-\gamma_{\mu\rho}g_{\nu}p_{\nu}g_{\nu}^{2}g$$

b) Under such a transformation,

Therefore, by choosing & s.t.

we obtain 2" h my = 0. This is under suitable conditions, always possible, as is clear from the theory of electromagnetism. Indeed, this corresponds to the inhomogeneous Maxwell's equations Lorenz gauge for a four potential sy created by a four current -2" hav. Under this transformation 2° 2(ph ,) p = = = (2 p 2 p + 2 , 28 p) = 0,

Thus, can (2) becomes

i.e.

3.a) We have $\frac{\partial h_{\mu\nu}}{\partial h_{\mu\nu}} = 0$ due to the static field condition. Thus, if we take the $\mu=\nu=0$ component of the linearized field eqns, we have

-16 rGg = -16 rG Too = 1 Thea = - 20 hoo + Ahoo = - 4 Ad.

Thus, we recover Poisson's eqn

$$\Delta \phi = 4\pi G_{g}$$

In this limit, the Christoffel symbols satisfy

$$\Gamma^{i}_{\mu\nu} = \frac{1}{2} \left(h_{\mu i,\nu} + h_{\nu i,\mu} - h_{\mu\nu,i} \right).$$

Moreover, For slow motion 11711<<1. Thus

$$-1 = u^{\mu}u_{\mu} = -\left(\frac{dx^{\circ}}{dt}\right)^{2} + \sum_{i=1}^{3} \left(\frac{dx^{i}}{dt}\right)^{2}$$

$$= -\left(\frac{dx^{\circ}}{dt}\right)^{2} \left(1 - ||\vec{v}||^{2}\right) \approx -\left(\frac{dx^{\circ}}{dt}\right)^{2}$$

and $\left|\frac{dx^{\circ}}{dz}\right|^{2} > \left|\left|\frac{d\bar{z}}{dz}\right|$. Therefore, the autoparallel eqn is

$$\frac{d^2x^i}{dt^2} = \frac{d}{dt} \left(\frac{dx^i}{dt} \frac{dy}{dx^0} \right) \frac{dx}{dx^0} = \frac{d^2x^i}{dt^2} = - \prod_{x \neq y} \frac{dx^x}{dt} \frac{dx^y}{dt}$$

$$= - \prod_{x \neq y} \frac{dx^y}{dt} \frac{dx^y}{dt} = \frac{1}{2} h_{00}(i)$$

To continue we need to find how. The other

Einstein equations become

$$\Delta \bar{h}_{iv} = \Box \bar{h}_{iv} = 0$$

Thus hormonico Since it is assumed to be bounded (as it is a small perturbation), it has to be constant. Thus, if our spacetimes is to approach Minkowski at infinity, hiv=0. Thus

$$-h_{00} = \bar{h} = h - 2h = -h$$

and

$$h_{00} = \bar{h}_{00} + \frac{1}{2} \eta_{00} h = \bar{h}_{00} = \frac{1}{2} \bar{h}_{00} = \frac{1}{2} h_{00} = -2 \phi.$$

Thus, indeed

$$\frac{d^2x^i}{dt^2} = -\phi_{ii} = -(\vec{\nabla}\phi)_{i}$$

recovering Newton's second law. Finally, the

components of the remaining components of the perturbation are
$$h_{\mu\nu} = \overline{h}_{\mu\nu} + \frac{1}{2}g_{\mu\nu}h = \overline{h}_{\mu\nu} + \frac{1}{2}g_{\mu\nu}\overline{h}_{00},$$

$$h_{i\nu} = \overline{h}_{i\nu} + \frac{1}{2}g_{i\nu}\overline{h}_{00}$$

$$= \frac{1}{2}\delta_{i\nu}\overline{h}_{00} = -2\delta_{i\nu}\phi.$$

We conclude

b) Taking the v=0 components of our linearized

Einstein aquations are hove

$$\Box A_{\mu} = -\frac{1}{4} \Box \overline{h}_{\mu 0} = -\frac{1}{4} (-16\pi T_{\mu 0}) = -4\pi J_{\mu}.$$

Taking this component for the De Donder gauge condition, $\partial^{\mu}A_{\mu}=-\frac{1}{4}\partial^{\mu}\bar{h}_{\mu 0}=0.$

We still have Thij = 0. Thus, as before, hij = 0.

$$h_{ij} = \overline{h}_{ij} - \frac{1}{2}g_{iy}\overline{h}_{00} = -2\delta_{iy}\phi.$$

$$h_{io} = \overline{h}_{io} + \frac{1}{z} g_{io} \overline{h}_{oo}$$

$$= -4A_{io} = \frac{2}{z} A_{io} + \frac{1}{z} g_{io} \overline{h}_{oo}$$

$$\int \mu v = \begin{bmatrix}
-(1+2\phi) & -4A_1 & -4A_2 & -4A_3 \\
-4A_1 & 1-2\phi & 0 \\
-4A_2 & 0 & 1-2\phi & 9 \\
-4A_3 & 0 & 4-2\phi - \mu v
\end{bmatrix}$$

Up to linear order in the velocity the geodesic equ

becomes

$$\frac{d^2x^i}{dt^2} = \frac{d}{dt} \left(\frac{dx^i}{dt} \frac{dt^3}{dx^0} \right) \frac{dy^1}{dx^0} = \frac{d^2x^i}{dt^2} = - \int_{-\infty}^{\infty} \frac{dx^{\infty}}{dt} \frac{dx^{\infty}}{dt} \frac{dx^{\beta}}{dt}$$

$$=-\Gamma^{i}_{00}-2\Gamma^{i}_{0j}v^{j}.$$

Much like before, $\prod_{i=0}^{i} = \frac{1}{2} h_{00} = \frac{1}{2} h_{00} = \frac{1}{2} \left(\overline{\forall} \phi \right)_{i}$. However,

naw

$$P^{i}_{oj} = \frac{1}{z} \left(h_{oi,j} + h_{ji,o} - h_{oj,i} \right) = 2 \left(2i A_{j} - 2j A_{i} \right)$$

$$= 2 \varepsilon_{ijk} B_{k}.$$

Thus,

$$\frac{d^2x^i}{dt^2} = -(\vec{\nabla}\phi)_i - 4 \epsilon_{ijk} B_k v^i$$

i.e.

$$\frac{d^2\vec{x}}{dt^2} = -\vec{E} - 4\vec{v} \times \vec{B}.$$

4. Hotice that

Thus

Then

Thus, the equations obtained through the variation of M are precisely the metricity conditions Mapy = 0. The variation with q is more complicated. Recalling (4.28) with many the menty recommendation

8 gea gorgan Mars Mean +

geagobdry Wall Hear +

gergabgun Haby Mear +

2 1-9 Mapy OMapy)

81-9 Mapy = - = 1-9 gmy dg my Mapy Mapy

ggagopgay Maby Mgox = -ggygay dg my gopgay Maby Mgox

=-Mypy Mapy dg ry

Jpa Sgopgar Maps Mgor =-gpagorgpv Sgragar Maps Mgor
=-Mars Mars Sgra

gragos Sgar Mas Mson =-gragos gargra Sgra Mas Mson

= - Mas, Mash Sgra

Moreover

dy Masn = - Da (g pr gro dyn) + Ms Baggr gro dyn

+ Mg ragpyggy dg my

We thus obtain the EOMs