

Quantum Optics¹

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Introduction

Quantum optics studies optical phenomena where the quantum nature of light is evident. It can be said that quantum optics examines optical phenomena in which light and the system interacting with it must be described by quantum equations.

In quantum optics, the frequency range is approximately from $f_1 \simeq 10^{13}\text{Hz}$ to $f_2 \simeq 10^{18}\text{Hz}$, that is, from the infrared range to X-rays. The lower limit is determined by the condition that the quantum energy exceeds the thermal motion energy: $\omega_1\hbar > kT$. The upper limit is set as in quantum optics, non-relativistic electron energies are generally considered, and therefore the quantum energy should be noticeably less than the electron rest energy: $\omega_2\hbar \ll mc^2$.

Currently, there are few educational materials dedicated to quantum optics. Among domestic works, one should mention [50]. The proposed material reflects a number of questions of the quantum theory of light, presented in the course "Quantum Optics." The manual consists of three parts.

The first part of the manual provides an introduction to quantum electrodynamics. It examines the simplest tasks related to the interaction of the light field and matter. A quantum theory of the laser is built on a simple model. The quantum description of light coherence and its connection with the classical description is considered. Chapter 1 gives a brief introduction to quantum electrodynamics within the scope necessary for presenting the course on quantum optics. Various quantum states of the light field used in quantum optics are considered. Special attention is paid to coherent states, which allow the quantum description of optical phenomena to be as close as possible to the classical one. In Chapter 2, the simplest tasks for the light field associated with both individual atoms and a large number of atoms in thermal equilibrium are considered. Special attention is given to the connection of the field with the thermostat, leading to relaxation processes that play a significant role in quantum optics and quantum electronics.

The second part of the manual deals with some applied issues of quantum optics. In Chapter 3, a simple model is used to examine the quantum theory of the laser. An expression for the statistics of laser photons and a formula estimating the "natural" width of the laser generation line are obtained. Chapter 4 reviews the quantum theory of the laser in the Heisenberg representation.

Chapter 5 is dedicated to photon optics. It outlines the quantum description of light coherence and its connection with classical description, coherence functions of different orders, the problem of photon statistics, and its connection with the statistics of photo-counts (Mandel's formula). Numerous examples are provided. Further, the connection of photon statistics with the spectral properties of light beams and its application for spectral measurements are considered.

In Chapter 6, a quantum description of interferometric experiments is given. The Mach-Zehnder interferometer and the error of interference measurements are examined. It will be shown later in Chapter 9 how to improve the accuracy of measurements using squeezed states.

Chapter 7 is devoted to the quantum description of the polarization properties of light. Later, this material is used in the description of entangled states in Chapter 10.

The third part of the manual discusses non-classical states of light. The definition of a non-classical state is provided (see Chapter 8). Squeezed (Chapter 9) and entangled (see chapter 10) non-classical states of light are described. For each type of non-classical state, the definition, ways of obtaining, application, and proof of non-classicality are given.

The fourth part of the manual considers quantum information theory. In particular, Chapter 11 introduces quantum information theory. The concept of the amount of information in the classical and quantum case is defined. Shannon's classical coding theorem and its quantum analogue are considered. Chapter 12 is dedicated to information security theory (cryptography): the classical theory of cryptography and its main drawbacks are described. A description of quantum cryptography is given. Chapter 13 describes the main quantum algorithms: Shor's algorithm for factoring integers and Grover's algorithm for searching in an unstructured data array.

The appendices contain material that complements the main course and mainly serves a reference function, allowing one not to resort to special literature.

The manual is intended for senior students specializing in the field of quantum electronics.

Part 1

Quantum Electrodynamics

Chapter 1

Quantization of the Electromagnetic Field

Quantum electrodynamics serves as the foundation of quantum optics. Below are outlined the principles of the quantum theory of the electromagnetic field necessary for studying and understanding quantum optics.

1.1 Decomposition of the Electromagnetic Field by Modes (Types of Oscillations)

The electromagnetic field satisfies the system of Maxwell's equations:

$$\text{rot} \vec{H} = \frac{\partial \vec{D}}{\partial t} + \vec{j}, \quad \text{rot} \vec{E} = -\frac{\partial \vec{B}}{\partial t}, \quad \text{div} \vec{D} = \rho, \quad \text{div} \vec{B} = 0. \quad (1.1)$$

In free space ($\vec{j} = 0, \rho = 0$) we have a simpler system:

$$\begin{aligned} \text{rot} \vec{H} &= \varepsilon_0 \frac{\partial \vec{E}}{\partial t}, \quad \text{rot} \vec{E} = -\mu_0 \frac{\partial \vec{H}}{\partial t}, \\ \text{div} \vec{H} &= 0, \quad \text{div} \vec{E} = 0, \\ \vec{B} &= \mu_0 \vec{H}, \quad \vec{D} = \varepsilon_0 \vec{E}, \quad \mu_0 \varepsilon_0 = \frac{1}{c^2}. \end{aligned} \quad (1.2)$$

Eliminating the term \vec{H} from (1.2), we obtain the equation:

$$\text{rot rot} \vec{E} = -\mu_0 \frac{\partial}{\partial t} \text{rot} \vec{H} = -\varepsilon_0 \mu_0 \frac{\partial^2 \vec{E}}{\partial t^2} = -\frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2}.$$

Considering $\text{rot rot} \vec{E} = \text{grad div} \vec{E} - \Delta \vec{E} = -\Delta \vec{E}$ ¹ we obtain the following system

¹under the condition $\text{div} \vec{E} = 0$

$$\begin{aligned}
\Delta \vec{E} - \frac{1}{c^2} \frac{\partial^2 \vec{E}}{\partial t^2} &= 0; \\
\operatorname{div} \vec{E} &= 0, \quad \frac{\partial H}{\partial t} = -\frac{1}{\mu_0} \operatorname{rot} \vec{E} \\
\vec{B} &= \mu_0 \vec{H}, \quad \vec{D} = \varepsilon_0 \vec{E}, \quad \mu_0 \varepsilon_0 = \frac{1}{c^2}.
\end{aligned} \tag{1.3}$$

which is fully equivalent to the original system of equations.

For quantizing the electromagnetic field, its equations are conveniently represented in the so-called Hamiltonian form. The essence of the method, well known to radiophysicists, is that the field is decomposed into modes, and solving the equations is reduced to solving a system of ordinary differential equations for the expansion coefficients, which depend on time. For this purpose, a system of orthogonal vector functions is introduced, which are the distribution of fields corresponding to the natural oscillations of the electromagnetic field in a certain volume V , bounded by an ideally conducting surface S . The fields must satisfy the system of equations (1.3). On the surface S certain boundary conditions must be satisfied, for example:

$$\left[\vec{n} \vec{E} \right] \Big|_S = 0 \text{ or } \left(\vec{n} \vec{H} \right) \Big|_S = 0, \tag{1.4}$$

corresponding to an ideally conducting surface.

An arbitrary electromagnetic field in volume V is represented by decompositions:

$$\vec{E}(r, t) = \sum_{(n)} Q_n(t) \vec{E}_n(r), \quad \vec{H}(r, t) = \sum_{(n)} P_n(t) \vec{H}_n(r), \tag{1.5}$$

where $Q_n(t) \vec{E}_n(r)$ and $P_n(t) \vec{H}_n(r)$ are partial solutions of the equations satisfying boundary conditions (1.4); $\vec{E}_n(r)$ and $\vec{H}_n(r)$ correspond to the n -th type of oscillation in volume V (mode). From the equations (1.3), we have:

$$Q_n(t) \Delta \vec{E}_n(r) = \frac{1}{c^2} \frac{d^2 Q_n}{dt^2} \vec{E}_n(r). \tag{1.6}$$

Separating variables, we get:

$$\frac{d^2 Q_n}{dt^2} + \omega_n^2 Q_n = 0, \quad \Delta \vec{E}_n(r) = -\frac{\omega_n^2}{c^2} \vec{E}_n(r) \tag{1.7}$$

where ω_n is the separation constant, which is the frequency of natural oscillations.

The second of the equations (1.7) takes the following form:

$$\Delta \vec{E}_n(r) + k_n^2 \vec{E}_n(r) = 0, \quad k_n^2 = \frac{\omega_n^2}{c^2}, \quad (1.8)$$

with boundary conditions $\left[\vec{n} \vec{E}_n \right] \Big|_S = 0$. This problem has a solution only at certain values of $k_n(\omega_n)$.

To determine $\vec{H}_n(r)$, we use the equations (1.3):

$$Q_n(t) \operatorname{rot} \vec{E}_n(r) = -\mu_0 \frac{dP_n}{dt} \vec{H}_n(r). \quad (1.9)$$

First, we make the following substitution:

$$Q_n(t) = \frac{\omega_n}{\sqrt{\varepsilon_0}} q_n(t), \quad P_n(t) = \frac{1}{\sqrt{\mu_0}} p_n(t). \quad (1.10)$$

For the new variable q_n , it is obvious from (1.7) that we have

$$\frac{d^2 q_n}{dt^2} + \omega_n^2 q_n = 0.$$

The decomposition (1.5) in the new variables has the form:

$$\vec{E}(r, t) = \sum_{(n)} \frac{q_n(t) \omega_n}{\sqrt{\varepsilon_0}} \vec{E}_n(r), \quad \vec{H}(r, t) = \sum_{(n)} \frac{p_n(t)}{\sqrt{\mu_0}} \vec{H}_n(r), \quad (1.11)$$

Considering the substitution (1.10) and the expression for the speed of light $c = \frac{1}{\sqrt{\varepsilon_0 \mu_0}}$ (1.2), we can rewrite (1.9) in the following form:

$$q_n(t) \operatorname{rot} \vec{E}_n(r) = -\frac{1}{c \omega_n} \frac{dp_n(t)}{dt} \vec{H}_n(r) \quad (1.12)$$

The equations (1.12) will be satisfied if

$$p_n = \frac{dq_n}{dt}, \quad \frac{dp_n}{dt} = \frac{d^2 q_n}{dt^2} = -\omega_n^2 q_n,$$

then

$$\begin{aligned} \operatorname{rot} \vec{E}_n(r) &= k_n \vec{H}_n(r), \\ \vec{H}_n(r) &= \frac{1}{k_n} \operatorname{rot} \vec{E}_n(r) \end{aligned} \quad (1.13)$$

From the course "Electromagnetic Oscillations," it is known that the system of functions $\vec{E}_n(r)$, $\vec{H}_n(r)$ is orthogonal and can be normalized:

$$\int_{\nu} \left(\vec{E}_n \vec{E}_m \right) d\nu = \delta_{nm}, \quad \int_{\nu} \left(\vec{H}_n \vec{H}_m \right) d\nu = \delta_{nm}, \quad \int_{\nu} \left(\vec{H}_n \vec{E}_m \right) d\nu = 0. \quad (1.14)$$

Solving the original electrodynamic problem reduces to finding the coefficients $q_n(t)$, $p_n(t)$ in the decompositions (1.11).

1.2 Hamiltonian Form of the Equations of Free Electromagnetic Field

Let's represent the field in volume V in the form of expansions (1.11):

$$\vec{E}(r, t) = \sum_{(n)} \frac{q_n(t) \omega_n}{\sqrt{\varepsilon_0}} \vec{E}_n(r), \quad \vec{H}(r, t) = \sum_{(n)} \frac{p_n(t)}{\sqrt{\mu_0}} \vec{H}_n(r), \quad (1.15)$$

where $p_n = \frac{dq_n}{dt}$ and q_n satisfies the equation

$$\frac{d^2 q_n}{dt^2} + \omega_n^2 q_n = 0.$$

This is the equation of a harmonic oscillator. Its solution:

$$q_n(t) = q_n(0) \cos \omega_n t + \frac{\dot{q}_n(0)}{\omega_n} \sin \omega_n t \quad (1.16)$$

$q_n(0)$ and $\dot{q}_n(0)$ are determined from the expansion of the initial field over \vec{E}_n , \vec{H}_n .

Now let's write down the equations of electrodynamics in Hamiltonian form. The Hamiltonian function for the electromagnetic field (field energy)

$$\mathcal{H} = \frac{1}{2} \int_{\nu} \left(\varepsilon_0 \left(\vec{E} \right)^2 + \mu_0 \left(\vec{H} \right)^2 \right) d\nu \quad (1.17)$$

Substituting the expansions (1.15) into (1.17) and using the orthogonality of eigen-

functions, we obtain

$$\begin{aligned}\mathcal{H} &= \frac{1}{2} \int_{\nu} \left(\sum_{(n)} \sum_{(m)} \varepsilon_0 \frac{q_n q_m \omega_n \omega_m}{\varepsilon_0} \left(\vec{E}_n \vec{E}_m \right) \right) d\nu + \\ &\quad + \frac{1}{2} \int_{\nu} \left(\sum_{(n)} \sum_{(m)} \mu_0 \frac{p_n p_m}{\mu_0} \left(\vec{H}_n \vec{H}_m \right) \right) d\nu = \\ &= \frac{1}{2} \sum_{(n)} (\omega_n^2 q_n^2 + p_n^2) .\end{aligned}\tag{1.18}$$

This expression corresponds to the Hamiltonian function for a set of independent harmonic oscillators.

For one mode, the Hamiltonian function is

$$\mathcal{H}_n = \frac{1}{2} (\omega_n^2 q_n^2 + p_n^2) .\tag{1.19}$$

then

$$\mathcal{H} = \sum_{(n)} \mathcal{H}_n\tag{1.20}$$

The equations of motion are obtained from the Hamiltonian function by the known method (14.49):

$$\dot{q}_n = \frac{\partial \mathcal{H}_n}{\partial p_n} = p_n, \quad \dot{p}_n = -\frac{\partial \mathcal{H}_n}{\partial q_n} = -\omega_n^2 q_n = \ddot{q}_n.\tag{1.21}$$

The obtained equations coincide with the equations (1.7). The Hamiltonian function of a mechanical oscillator

$$\mathcal{H} = \frac{1}{2} \left(M \Omega^2 x^2 + \frac{p^2}{M} \right)$$

Therefore, (1.19) formally corresponds to an oscillator with unit mass and natural frequency ω_n . The Hamiltonian representation of the electromagnetic field equations is convenient for the procedure of quantizing the field. This approach is suitable for solving various classical problems in many cases. More details can be found in the book [41].

1.3 Quantization of the Free Electromagnetic Field

Using the previously noted analogy, the electromagnetic field can be quantized similarly to a simple mechanical harmonic oscillator.

In quantization, q_n and p_n become operators \hat{q}_n , \hat{p}_n , satisfying the same commutation relations as position and momentum:

$$[\hat{q}_n, \hat{p}_n] = \hat{q}_n \hat{p}_n - \hat{p}_n \hat{q}_n = i\hbar \quad (1.22)$$

where \hat{q}_n and \hat{p}_n are Hermitian (self-adjoint) operators. In the Schrödinger picture, they are time-independent, while in the Heisenberg picture, they depend on time. It is convenient to introduce new operators via these operators, which are not self-adjoint but are conjugate to each other:

$$\hat{a}_n = \frac{1}{\sqrt{2\hbar\omega_n}} (\omega_n \hat{q}_n + i\hat{p}_n), \quad \hat{a}_n^\dagger = \frac{1}{\sqrt{2\hbar\omega_n}} (\omega_n \hat{q}_n - i\hat{p}_n), \quad (1.23)$$

The inverse relation gives

$$\hat{q}_n = \sqrt{\frac{\hbar}{2\omega_n}} (\hat{a}_n + \hat{a}_n^\dagger), \quad \hat{p}_n = i\sqrt{\frac{\hbar\omega_n}{2}} (\hat{a}_n^\dagger - \hat{a}_n), \quad (1.24)$$

Using expressions (1.22), (1.23), (1.24), one can obtain the commutation relations for the operators \hat{a}_n and \hat{a}_n^\dagger :

$$\begin{aligned} [\hat{a}_n, \hat{a}_n^\dagger] &= \frac{1}{2\hbar\omega_n} (\omega_n \hat{q}_n + i\hat{p}_n) (\omega_n \hat{q}_n - i\hat{p}_n) - \\ &\quad - \frac{1}{2\hbar\omega_n} (\omega_n \hat{q}_n - i\hat{p}_n) (\omega_n \hat{q}_n + i\hat{p}_n) = \\ &= \frac{1}{2\hbar\omega_n} (-2i\omega_n (\hat{q}_n \hat{p}_n - \hat{p}_n \hat{q}_n)) = 1. \end{aligned} \quad (1.25)$$

Thus,

$$[\hat{a}_n, \hat{a}_n^\dagger] = 1. \quad (1.26)$$

Additionally, it is obvious that

$$[\hat{a}_n, \hat{a}_n] = 0, \quad [\hat{a}_n^\dagger, \hat{a}_n^\dagger] = 0. \quad (1.27)$$

The Hamiltonian, expressed through the new operators, takes the form

$$\begin{aligned} \hat{\mathcal{H}}_n &= \frac{1}{2} (\omega_n^2 \hat{q}_n^2 + \hat{p}_n^2) = \\ &= \frac{1}{2} \frac{\omega_n \hbar}{2} ((\hat{a}_n + \hat{a}_n^\dagger) (\hat{a}_n + \hat{a}_n^\dagger) - (\hat{a}_n^\dagger - \hat{a}_n) (\hat{a}_n^\dagger - \hat{a}_n)) = \\ &= \frac{\omega_n \hbar}{2} (\hat{a}_n \hat{a}_n^\dagger + \hat{a}_n^\dagger \hat{a}_n) = \frac{\omega_n \hbar}{2} (1 + \hat{a}_n^\dagger \hat{a}_n + \hat{a}_n^\dagger \hat{a}_n) = \\ &= \omega_n \hbar \left(\hat{a}_n^\dagger \hat{a}_n + \frac{1}{2} \right) \end{aligned} \quad (1.28)$$

Here, the relation of \hat{q}_n , \hat{p}_n with \hat{a}_n^\dagger , \hat{a}_n (1.23), as well as the commutation relation (1.26), are used.

The energy $\frac{1}{2}\omega_n\hbar$ corresponds to zero oscillations. The total energy of zero oscillations

$$\frac{1}{2} \sum_{(n)} \omega_n \hbar \rightarrow \infty,$$

as the number of modes is infinite. This does not lead to significant difficulties since we are only interested in energy differences, and the constant part can be discarded. Then the Hamiltonian takes the form

$$\hat{\mathcal{H}} = \sum_{(n)} \hat{\mathcal{H}}_n = \sum_{(n)} \hbar \omega_n \hat{a}_n^\dagger \hat{a}_n \quad (1.29)$$

Operators of the electric and magnetic fields can be represented as

$$\begin{aligned} \hat{\vec{E}} &= \sum_{(n)} \frac{\hat{q}_n \omega_n}{\sqrt{\varepsilon_0}} \vec{E}_n(r) = \sum_{(n)} \sqrt{\frac{\hbar \omega_n}{2\varepsilon_0}} (\hat{a}_n^\dagger + \hat{a}_n) \vec{E}_n(r), \\ \hat{\vec{H}} &= \sum_{(n)} \frac{\hat{p}_n}{\sqrt{\mu_0}} \vec{H}_n(r) = i \sum_{(n)} \sqrt{\frac{\hbar \omega_n}{2\mu_0}} (\hat{a}_n^\dagger - \hat{a}_n) \vec{H}_n(r). \end{aligned} \quad (1.30)$$

Let's provide a simple example. In laser theory, a resonator is considered, formed by two parallel mirrors. In this case, an approximate field representation is often used, assuming that the field depends only on the longitudinal coordinate (Figure 1.1). In this approximation, the normalized eigenfunction (mode) can be presented as follows:

$$E_{nx} = \sqrt{\frac{2}{V}} \sin k_n z, \quad H_{ny} = \sqrt{\frac{2}{V}} \cos k_n z, \quad k_n = \frac{\pi n}{L}, \quad \omega_n = \frac{c\pi n}{L}, \quad (1.31)$$

where $V = SL$, L - length of the resonator, S - cross-section of the light beam.

The operator of the electric field mode will then have the form:

$$\begin{aligned} \hat{E}_{nx} &= \sqrt{\frac{\hbar \omega_n}{2\varepsilon_0}} (\hat{a}_n^\dagger + \hat{a}_n) \sqrt{\frac{2}{V}} \sin k_n z = E_1 (\hat{a}_n^\dagger + \hat{a}_n) \sin k_n z, \\ \hat{H}_{ny} &= i E_1 \sqrt{\frac{\varepsilon_0}{\mu_0}} (\hat{a}_n^\dagger - \hat{a}_n) \cos k_n z, \end{aligned} \quad (1.32)$$

where $E_1 = \sqrt{\frac{\hbar \omega}{\varepsilon_0 V}}$ - electric field corresponding to one photon (quantum) in the mode.

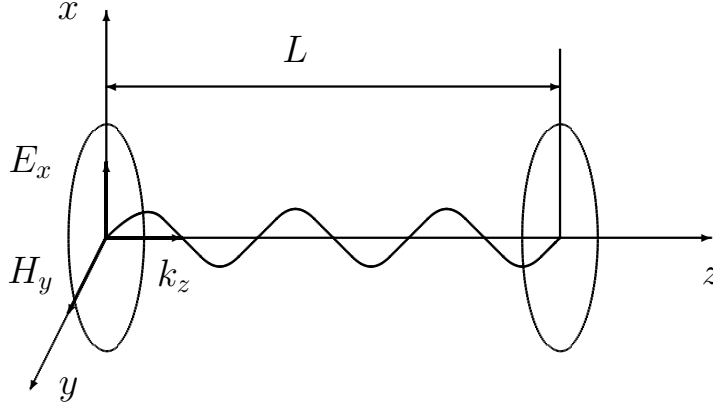


Figure 1.1: Electromagnetic field in a resonator formed by two mirrors. The length of the resonator is L . The wave vector of the electromagnetic field is directed along the z axis, the electric field along the x axis, and the magnetic field along the y axis.

1.4 Field Decomposition by Plane Waves in Free Space

So far, we have considered the electromagnetic field in a physically defined volume - in a resonator. If we are dealing with free space, we can artificially select a sufficiently large volume containing the region of interest, define modes (types of oscillations) with appropriate boundary conditions, and then proceed as already considered. If necessary, the volume in the final result can be taken to infinity. A sufficiently large cubic volume with a side L is usually used (Figure 1.2). In this case, it is customary to use periodic boundary conditions:

$$\begin{aligned}\vec{E}(0, y, z) &= \vec{E}(L, y, z), \\ \vec{E}(x, 0, z) &= \vec{E}(x, L, z), \\ \vec{E}(x, y, 0) &= \vec{E}(x, y, L).\end{aligned}\tag{1.33}$$

It is convenient to decompose by plane waves. As known from the course of electromagnetic oscillations, the solution corresponding to a plane wave has the form

$$\vec{E}_k(r, t) = A_k(t) \vec{e}_k e^{i(\vec{k}\vec{r})} + (\text{c.c.}),\tag{1.34}$$

where \vec{e}_k is the unit vector of wave polarization; k is the wave number; \vec{k} is the wave vector; $A_k(t) = A_k(0) e^{-i\omega_k t}$.

The magnetic field is related to the electric field by the relation

$$\vec{H}_k(r, t) = \sqrt{\frac{\varepsilon_0}{\mu_0}} \frac{1}{k} [\vec{k} \vec{e}_k] A_k(t) e^{i(\vec{k}\vec{r})} + (\text{c.c.}).\tag{1.35}$$

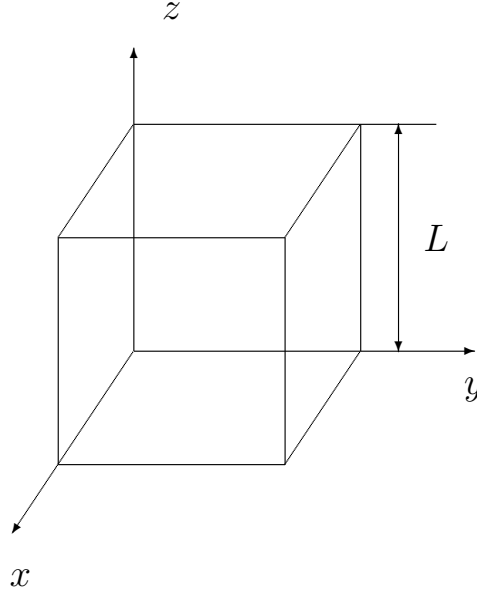


Figure 1.2: Decomposition of the electromagnetic field into plane waves in free space. The region of space in which the decomposition is performed is selected in the form of a cube with side L . Periodic boundary conditions are used on the boundary of the cube.

The following equalities hold:

$$\left(\vec{k}\vec{E}_k\right) = \left(\vec{k}\vec{H}_k\right) = \left(\vec{E}_k\vec{H}_k\right) = 0, \quad k^2 = \left(\vec{k}\vec{k}\right) = \frac{\omega_k^2}{c^2} \quad (1.36)$$

indicating that \vec{E}_k and \vec{H}_k are perpendicular to the direction of wave propagation and to each other (the wave is transverse).

The periodicity condition (1.33) will be satisfied if

$$\vec{k} = \frac{2\pi}{L} (n_x \vec{x}_0 + n_y \vec{y}_0 + n_z \vec{z}_0), \quad n_x, n_y, n_z = 0, \pm 1, \pm 2, \dots \quad (1.37)$$

Then we obtain

$$\begin{aligned} \left(\vec{k}\vec{r}\right) &= \frac{2\pi}{L} (n_x x + n_y y + n_z z), \\ \left(\vec{k}\vec{r}\right)\Big|_{x=L} &= 2\pi n_x + \frac{2\pi}{L} (n_y y + n_z z) = 2\pi n_x + \left(\vec{k}\vec{r}\right)\Big|_{x=0}. \end{aligned} \quad (1.38)$$

Thus, $e^{i(\vec{k}\vec{r})}$ is periodic in x . The periodicity is demonstrated in the same way for y and z .

In the future, instead of real functions (1.34), (1.35), we will use complex eigenfunctions

$$\vec{E}_k(r) = \vec{e}_k e^{i(\vec{k}\vec{r})}, \quad \vec{H}_k(r) = \sqrt{\frac{\varepsilon_0}{\mu_0}} \frac{1}{k} \left[\vec{k} \vec{E}_k(r) \right]. \quad (1.39)$$

The following relations hold

$$\left(\vec{k}\vec{E}_k\right) = \left(\vec{k}\vec{H}_k\right) = \left(\vec{E}_k\vec{H}_k\right) = 0, \quad k^2 = \left(\vec{k}\vec{k}\right) = \frac{\omega_k^2}{c^2}.$$

The eigenfrequencies ω_k are determined by equation (1.37). Indeed, from (1.37), considering $k^2 = \frac{\omega_k^2}{c^2}$, we have

$$\omega_k = c\sqrt{k_x^2 + k_y^2 + k_z^2} = \frac{2\pi c}{L}\sqrt{n_x^2 + n_y^2 + n_z^2}. \quad (1.40)$$

For linear polarization, the polarization vectors \vec{e}_k are:

$$\vec{e}_{k_1} = \vec{\xi}_0, \quad \vec{e}_{k_2} = \vec{\eta}_0,$$

where $\vec{\xi}_0, \vec{\eta}_0$, are unit vectors of directions forming a rectangular system with the vector \vec{k} , therefore,

$$\left(\vec{\xi}_0\vec{k}\right) = \left(\vec{\eta}_0\vec{k}\right) = (\vec{e}_{k_1}\vec{e}_{k_2}) = 0.$$

For circular polarization

$$\vec{e}_{k_1} = \frac{\vec{\xi}_0 + i\vec{\eta}_0}{\sqrt{2}}, \quad \vec{e}_{k_2} = \frac{\vec{\xi}_0 - i\vec{\eta}_0}{\sqrt{2}}.$$

In this case, the orthogonality condition holds: $(\vec{e}_{k_1}\vec{e}_{k_2}^*)$ and normalization conditions: $(\vec{e}_{k_1}\vec{e}_{k_1}^*) = (\vec{e}_{k_2}\vec{e}_{k_2}^*) = 1$.

An arbitrary electromagnetic field can be represented using complex functions (1.39) by decompositions

$$\begin{aligned} \vec{E}(r, t) &= \sum_{(k)} A_k(t) \vec{E}_k(r) + (\text{c.c.}), \\ \vec{H}(r, t) &= \sum_{(k)} A_k(t) \vec{H}_k(r) + (\text{c.c.}) \end{aligned} \quad (1.41)$$

1.5 Density of States

If the quantization volume L^3 is large, the density of states (the number of modes per unit frequency interval) will be quite large. The wave vector of the mode is defined by the relation (1.37)

$$\vec{k}_{n_x, n_y, n_z} = \frac{2\pi}{L} (n_x \vec{x}_0 + n_y \vec{y}_0 + n_z \vec{z}_0) \quad (1.42)$$



Figure 1.3: Allowed modes of the electromagnetic field (indicated by dots) for a confined space in the form of a cube with side L . Rectangular coordinate system.

Each set of integers corresponds to two waves differing in polarization. The modes can be visually represented by points in a Cartesian coordinate system with axes n_x, n_y, n_z (Figure 1.3). The number of oscillations in the volume $\Delta n_x \Delta n_y \Delta n_z$ is obviously $\Delta N = 2\Delta n_x \Delta n_y \Delta n_z$, or, considering the relationship between k and n (1.37),

$$\Delta N = 2 \left(\frac{L}{2\pi} \right)^3 \Delta k_x \Delta k_y \Delta k_z. \quad (1.43)$$

The two in formula (1.43) appears because one value of k corresponds to two modes with different polarizations. For large L ($L \rightarrow \infty$), the distribution becomes quasi-continuous, and summing over the modes, which is almost always required when solving quantum optics problems, can be replaced by integration over k :

$$2 \sum_{(n)} (\dots) = 2 \left(\frac{L}{2\pi} \right)^3 \iiint_{-\infty}^{+\infty} (\dots) dk_x dk_y dk_z. \quad (1.44)$$

The transition from rectangular coordinates k_x, k_y, k_z to spherical coordinates k, θ, φ (Figure 1.4), gives

$$\begin{aligned} k_x &= k \sin \theta \cos \varphi, & k_y &= k \sin \theta \sin \varphi, & k_z &= k \cos \theta, \\ dk_x dk_y dk_z &= k^2 \sin \theta dk d\theta d\varphi = k^2 dk d\Omega, \end{aligned} \quad (1.45)$$

where $d\Omega = \sin \theta d\theta d\varphi$ is the solid angle element in the direction of \vec{k} . Thus, we obtain

$$dN = 2 \left(\frac{L}{2\pi} \right)^3 k^2 dk d\Omega \quad (1.46)$$

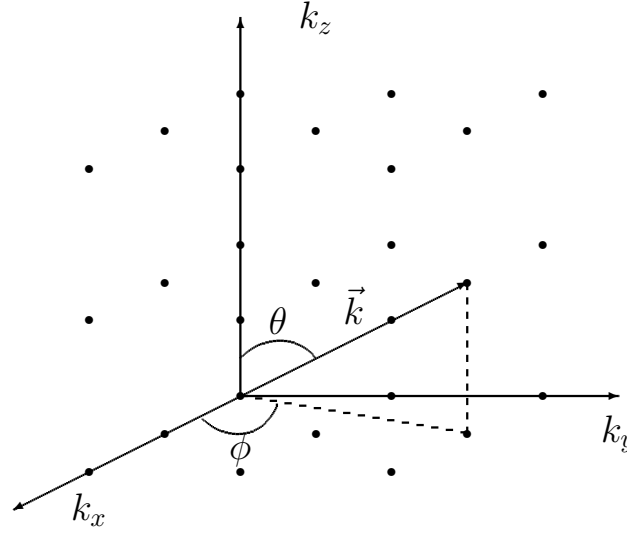


Figure 1.4: Allowed modes of the electromagnetic field (indicated by dots) for a confined space in the shape of a cube with side L . Spherical coordinate system.

or, considering that $k^2 c^2 = \omega^2$, i.e., $k^2 dk = \frac{\omega^2 d\omega}{c^3}$:

$$dN = 2 \left(\frac{L}{2\pi c} \right)^3 \omega^2 d\omega d\Omega \quad (1.47)$$

i.e., the number of states per unit volume, unit frequency, and unit solid angle

$$g(\omega) = \frac{2\omega^2}{(2\pi c)^3} \quad (1.48)$$

(spectral density of eigenstates).

1.6 Hamiltonian Form of Field Equations in Plane Wave Decomposition

Eigenfunctions (1.39) are orthogonal, i.e.,

$$\begin{aligned} \int_{\nu} \left(\vec{E}_k \vec{E}_{k'}^* \right) d\nu &= C \int_{\nu} e^{i(\vec{k}\vec{r})} e^{-i(\vec{k}'\vec{r})} d\nu = \\ &= C \int_0^L dx \int_0^L dy \int_0^L dz \left[e^{i((k_x - k'_x)x + (k_y - k'_y)y + (k_z - k'_z)z)} \right] = 0, \end{aligned} \quad (1.49)$$

if at least one of the inequalities $k_x \neq k'_x$, $k_y \neq k'_y$, $k_z \neq k'_z$, (or $\vec{k} - \vec{k}' \neq 0$) holds. This is due to the periodicity of the integrand functions.

Similarly, we obtain:

$$\int_{\nu} \left(\vec{E}_k \vec{E}_{k'} \right) d\nu = 0, \quad \vec{k} \neq -\vec{k}'. \quad (1.50)$$

The eigenfunctions \vec{E}_k can be normalized by setting

$$\int_{\nu} \left(\vec{E}_k \vec{E}_k^* \right) d\nu = 1. \quad (1.51)$$

Then the eigenfunctions will take the form

$$\vec{E}_k(r) = \frac{1}{\sqrt{\nu}} \vec{e}_k e^{i(\vec{k}\vec{r})}, \quad \vec{H}_k(r) = \frac{1}{k} \left[\vec{k} \vec{e}_k(r) \right] \sqrt{\frac{\varepsilon_0}{\nu \mu_0}} e^{i(\vec{k}\vec{r})}.$$

Note that under the index k we imply k_j , where $j = 1, 2$ correspond to two different polarizations. Functions \vec{E}_k and \vec{H}_k , presented in this form, satisfy the following orthogonality and normalization conditions:

$$\begin{aligned} \int_{\nu} \left(\vec{E}_k \vec{E}_{k'}^* \right) d\nu &= 0 \quad \text{when } \vec{k} \neq \vec{k}', \\ \int_{\nu} \left(\vec{E}_k \vec{E}_{k'} \right) d\nu &= 0 \quad \text{when } \vec{k} \neq -\vec{k}', \\ \int_{\nu} \left(\vec{E}_k \vec{E}_k^* \right) d\nu &= 1, \quad \int_{\nu} \left(\vec{E}_k \vec{E}_{-k} \right) d\nu = 1, \\ \int_{\nu} \left(\vec{H}_k \vec{H}_{k'}^* \right) d\nu &= 0 \quad \text{when } \vec{k} \neq \vec{k}', \\ \int_{\nu} \left(\vec{H}_k \vec{H}_{k'} \right) d\nu &= 0 \quad \text{when } \vec{k} \neq -\vec{k}', \\ \int_{\nu} \left(\vec{H}_k \vec{H}_k^* \right) d\nu &= \frac{\varepsilon_0}{\mu_0}, \quad \int_{\nu} \left(\vec{H}_k \vec{H}_{-k} \right) d\nu = -\frac{\varepsilon_0}{\mu_0}. \end{aligned} \quad (1.52)$$

Orthogonality of modes with the same k value but different polarizations follows from the identities:

$$(\vec{e}_{k_1} \vec{e}_{k_2}^*) = 0, \quad (\vec{e}_{k_1} \vec{e}_{-k_2}) = 0.$$

The electric and magnetic fields are expressed through these functions using decomposition:

$$\begin{aligned} \vec{E}(r, t) &= \sum_{(k)} A_k(t) \vec{E}_k(r) + \text{c.c.}, \\ \vec{H}(r, t) &= \sum_{(k)} A_k(t) \vec{H}_k(r) + \text{c.c.} \end{aligned} \quad (1.53)$$

Summation over $k = k_j$ implies summation over all k values and two polarizations ($j = 1, 2$): $\sum_{(k)} \equiv \sum_{j=1,2} \sum_{(k_j)}$.

The Hamiltonian function for the electromagnetic field is:

$$\mathcal{H} = \frac{1}{2} \int_{\nu} \left(\varepsilon_0 \left(\vec{E}(r, t) \right)^2 + \mu_0 \left(\vec{H}(r, t) \right)^2 \right) d\nu \quad (1.54)$$

Substituting in (1.54) the series (1.53), multiplying them and considering the orthogonality and normalization conditions, we obtain

$$\begin{aligned} \mathcal{H} = & \frac{1}{2} \varepsilon_0 \int_{(\nu)} \left\{ \sum_{(k)} \sum_{(k')} \left(A_k \vec{E}_k + A_k^* \vec{E}_k^* \right) \left(A_{k'} \vec{E}_{k'} + A_{k'}^* \vec{E}_{k'}^* \right) \right\} d\nu + \\ & + \frac{1}{2} \mu_0 \int_{(\nu)} \left\{ \sum_{(k)} \sum_{(k')} \left(A_k \vec{H}_k + A_k^* \vec{H}_k^* \right) \left(A_{k'} \vec{H}_{k'} + A_{k'}^* \vec{H}_{k'}^* \right) \right\} d\nu = \\ & \frac{\varepsilon_0}{2} \sum_{(k)} \left(A_k A_k^* + A_k A_{-k} + A_k^* A_{-k}^* + A_k^* A_k + \right. \\ & \left. + A_k A_k^* - A_k A_{-k} - A_k^* A_{-k}^* + A_k^* A_k \right) = \varepsilon_0 \sum_{(k)} \left(A_k A_k^* + A_k^* A_k \right). \quad (1.55) \end{aligned}$$

1.7 Quantization of the Electromagnetic Field by Decomposing it into Plane Waves

For the quantization of the electromagnetic field in this case, we note the analogy between (1.55) and (1.28):

$$\hat{\mathcal{H}}_n = \frac{\omega_n \hbar}{2} (\hat{a}_n \hat{a}_n^\dagger + \hat{a}_n^\dagger \hat{a}_n) = \omega_n \hbar \left(\hat{a}_n^\dagger \hat{a}_n + \frac{1}{2} \right)$$

and

$$\mathcal{H} = \varepsilon_0 \sum_{(k)} \left(A_k A_k^* + A_k^* A_k \right).$$

From this analogy, it follows that the following quantization procedure can be used:

$$\sqrt{\varepsilon_0} A_k \rightarrow \sqrt{\frac{\omega_k \hbar}{2}} \hat{a}_k, \quad \sqrt{\varepsilon_0} A_k^* \rightarrow \sqrt{\frac{\omega_k \hbar}{2}} \hat{a}_k^\dagger.$$

This substitution leads to the following expression for the Hamiltonian:

$$\hat{\mathcal{H}}_k = \frac{\omega_k \hbar}{2} \left(\hat{a}_k \hat{a}_k^\dagger + \hat{a}_k^\dagger \hat{a}_k \right).$$

Using the commutation relations

$$[\hat{a}, \hat{a}^\dagger] = 1, \quad \hat{a}\hat{a}^\dagger - \hat{a}^\dagger\hat{a} = 1, \quad \hat{a}\hat{a}^\dagger = \hat{a}^\dagger\hat{a} + 1,$$

we have $\hat{\mathcal{H}}_k = \omega_k \hbar \left(\hat{a}_k^\dagger \hat{a}_k + \frac{1}{2} \right)$ - an expression completely matching (1.28). The total Hamiltonian is obtained by summing over all modes:

$$\hat{\mathcal{H}} = \sum_{(k)} \hat{\mathcal{H}}_k = \sum_{(k)} \omega_k \hbar \left(\hat{a}_k^\dagger \hat{a}_k + \frac{1}{2} \right). \quad (1.56)$$

The electromagnetic field momentum can also be represented in Hamiltonian form. The classical momentum of an electromagnetic field in a volume V is defined by the formula

$$\vec{G} = \frac{1}{c^2} \int_{(\nu)} [\vec{E} \vec{H}] d\nu. \quad (1.57)$$

Using the decomposition of the field into plane waves and considering orthogonality relations, we obtain

$$\begin{aligned} \vec{G} &= \frac{1}{c^2} \int_{(\nu)} [\vec{E} \vec{H}] d\nu = \\ &= \frac{1}{c^2} \int_{(\nu)} \left[\left(\sum_{(k)} A_k \vec{E}_k + \sum_{(k)} A_k^* \vec{E}_k^* \right) \left(\sum_{(k')} A_{k'} \vec{H}_{k'} + \sum_{(k')} A_{k'}^* \vec{H}_{k'}^* \right) \right] d\nu = \\ &= \sum_{(k)} \frac{1}{c^2} \int_{(\nu)} \left(A_k A_k^* [\vec{E}_k \vec{H}_k^*] + A_k^* A_k [\vec{E}_k^* \vec{H}_k] \right) d\nu = \\ &= \frac{1}{c^2} \sum_{(k)} \frac{\vec{k}}{k} \sqrt{\frac{\varepsilon_0}{\mu_0}} (A_k A_k^* + A_k^* A_k). \end{aligned}$$

Transitioning to operators, we have

$$\begin{aligned} \hat{\vec{G}} &= \frac{1}{c^2} \sum_{(k)} \frac{\vec{k}}{k} \sqrt{\frac{\varepsilon_0}{\mu_0}} \frac{\hbar \omega_k}{2\varepsilon_0} \left(\hat{a}_k^\dagger \hat{a}_k + \hat{a}_k \hat{a}_k^\dagger \right) = \\ &= \sum_{(k)} \hbar \vec{k} \frac{\omega_k}{ck} \frac{1}{c\sqrt{\varepsilon_0\mu_0}} \frac{\hat{a}_k^\dagger \hat{a}_k + \hat{a}_k \hat{a}_k^\dagger}{2} = \sum_{(k)} \vec{k} \hbar \left(\hat{a}_k^\dagger \hat{a}_k + \frac{1}{2} \right). \end{aligned}$$

From symmetry, it follows that $\sum_{(k)} \vec{k} \hbar = 0$ and, therefore,

$$\hat{\vec{G}} = \sum_{(k)} \vec{k} \hbar \hat{a}_k^\dagger \hat{a}_k. \quad (1.58)$$

The operator of the electric field is now expressed as follows:

$$\hat{\vec{E}} = \sum_{(k)} \hat{a}_k \sqrt{\frac{\hbar\omega_k}{2\nu\varepsilon_0}} \vec{e}_k e^{i(\vec{k}\vec{r})} + \sum_{(k)} \hat{a}_k^\dagger \sqrt{\frac{\hbar\omega_k}{2\nu\varepsilon_0}} \vec{e}_k^* e^{-i(\vec{k}\vec{r})}. \quad (1.59)$$

In the future, we will use the Dirac formulation of quantum mechanics equations, so in [section 14.1](#) a brief exposition of the Dirac formalism is given: the concept of state vector is introduced and ways of operating with it are outlined.

1.8 Properties of Operators \hat{a} and \hat{a}^\dagger (Annihilation and Creation Operators)

First of all, let's consider the properties of the operators \hat{a} and \hat{a}^\dagger . We will show that

$$[\hat{\mathcal{H}}, \hat{a}] = -\hbar\omega\hat{a}, \quad [\hat{\mathcal{H}}, \hat{a}^\dagger] = \hbar\omega\hat{a}^\dagger. \quad (1.60)$$

Since

$$\hat{\mathcal{H}} = \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right)$$

then, using commutation relations

$$[\hat{a}, \hat{a}^\dagger] = 1,$$

we get

$$\begin{aligned} [\hat{\mathcal{H}}, \hat{a}] &= \hbar\omega \left(\hat{a}^\dagger \hat{a} \hat{a} + \frac{1}{2} \hat{a} - \hat{a} \hat{a}^\dagger \hat{a} - \frac{1}{2} \hat{a} \right) = \\ &= \hbar\omega (\hat{a}^\dagger \hat{a} \hat{a} - (1 + \hat{a}^\dagger \hat{a}) \hat{a}) = -\hbar\omega \hat{a}, \end{aligned}$$

The second equality (1.60) is proved similarly.

In the Schrödinger representation, the operators do not depend on time, while in the Heisenberg representation, they do depend and obey the equations of motion:

$$\dot{\hat{a}}(t) = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{a}] = -i\omega\hat{a}(t), \quad \dot{\hat{a}}^\dagger(t) = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{a}^\dagger] = i\omega\hat{a}^\dagger(t). \quad (1.61)$$

Here, the equalities (1.60) are used. The solution of these equations:

$$\hat{a}(t) = \hat{a}(0) e^{-i\omega t}, \quad \hat{a}^\dagger(t) = \hat{a}^\dagger(0) e^{i\omega t}, \quad (1.62)$$

where $\hat{a}(0)$, $\hat{a}^\dagger(0)$ are the operators in the Schrödinger representation.

1.9 Quantum state of the electromagnetic field with a definite energy (with a definite number of photons)

First, let's consider a single mode of the electromagnetic field (a simple harmonic oscillator). The state vector with a definite energy $|E_n\rangle$ satisfies the following equation

$$\hat{\mathcal{H}} |E_n\rangle = E_n |E_n\rangle \quad (1.63)$$

Here and henceforth, we will use Dirac's formalism (see [section 14.1](#)). Using relation (1.60), we obtain

$$\hat{\mathcal{H}}\hat{a} |E_n\rangle = \left(\hat{a}\hat{\mathcal{H}} - \hbar\omega\hat{a}\right) |E_n\rangle = (E_n - \hbar\omega) \hat{a} |E_n\rangle \quad (1.64)$$

i.e., $\hat{a} |E_n\rangle$ is also a state vector with energy $E_n - \hbar\omega$. From this, it follows that

$$\hat{a} |E_n\rangle = |E_n - \hbar\omega\rangle$$

Thus, the operator \hat{a} lowers the energy of the state by $\hbar\omega$, where ω is the frequency of the mode (oscillator). The operator \hat{a} is often called the lowering operator, or the annihilation operator. The lowest energy must be positive, and reducing the energy cannot continue indefinitely. For an arbitrary state vector, the expected energy value

$$\langle\Phi| \hbar\omega \left(a^\dagger a + \frac{1}{2}\right) |\Phi\rangle = \hbar\omega \langle\Phi'| \Phi'\rangle + \frac{1}{2}\hbar\omega, \quad (1.65)$$

where $\hat{a} |\Phi\rangle = |\Phi'\rangle$, $\langle\Phi| \hat{a}^\dagger = \langle\Phi'|$. Since the norm of the state vector must be positive, the minimum energy value will be at $\langle\Phi'| \Phi'\rangle = 0$. This means $\hat{a} |0\rangle = 0$, where $|\Phi\rangle = |0\rangle$ is the state vector with the lowest energy. The lowest energy

$$E_0 = \frac{\hbar\omega}{2} \quad (1.66)$$

is called the zero-point energy. To verify this claim, we can write

$$\begin{aligned} \hat{\mathcal{H}} |0\rangle &= \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2}\right) |0\rangle = \\ &= \hbar\omega \hat{a}^\dagger \hat{a} |0\rangle + \frac{\hbar\omega}{2} |0\rangle = \\ &= \frac{\hbar\omega}{2} |0\rangle = E_0 |0\rangle. \end{aligned} \quad (1.67)$$

Using (1.60), we can obtain

$$\begin{aligned}\hat{\mathcal{H}}\hat{a}^\dagger|0\rangle &= \hat{\mathcal{H}}|1\rangle = \left(\hat{a}^\dagger\hat{\mathcal{H}} + \hbar\omega\hat{a}^\dagger\right)|0\rangle = \\ &= \hbar\omega\left(1 + \frac{1}{2}\right)\hat{a}^\dagger|0\rangle = \hbar\omega\left(1 + \frac{1}{2}\right)|1\rangle\end{aligned}\quad (1.68)$$

where $|1\rangle = \hat{a}^\dagger|0\rangle$ denotes the state with energy $\hbar\omega\left(1 + \frac{1}{2}\right)$.

By induction, we have

$$\hat{\mathcal{H}}(\hat{a}^\dagger)^n|0\rangle = \hat{\mathcal{H}}|n\rangle = \hbar\omega\left(n + \frac{1}{2}\right)(\hat{a}^\dagger)^n|0\rangle = \hbar\omega\left(n + \frac{1}{2}\right)|n\rangle \quad (1.69)$$

where $|n\rangle = (\hat{a}^\dagger)^n|0\rangle$ without normalization, which we will perform later - the state with energy $\hbar\omega\left(n + \frac{1}{2}\right)$, n is a positive integer.

We see that the operator \hat{a}^\dagger raises the energy of the state by $\hbar\omega$. It can be considered as the creation operator of a photon with energy $\hbar\omega$. The photon is better understood as a particle with energy $\hbar\omega$ and momentum $\hbar\vec{k}$ when considering field decomposition into plane waves, as follows from (1.58).

The relations $\hat{a}|n\rangle = |n-1\rangle$ and $\hat{a}^\dagger|n\rangle = |n+1\rangle$ define unnormalized state vectors. Let's define the normalization factor. Assume that $\hat{a}|n\rangle = S_n|n-1\rangle$, where $|n\rangle$ and $|n-1\rangle$ are normalized to 1, and S_n is the normalization factor. From this, we find

$$S_n^2\langle n-1|n-1\rangle = \langle n|\hat{a}^\dagger\hat{a}|n\rangle = n\langle n|n\rangle$$

since the operator $\hat{a}^\dagger\hat{a} = \hat{n}$ is the photon number operator, whose eigenvalue is the photon number. This is evident from the formula (1.69). Indeed, from the equality

$$\hat{\mathcal{H}}|n\rangle = \hbar\omega\left(\hat{a}^\dagger\hat{a} + \frac{1}{2}\right)|n\rangle = \hbar\omega\left(n + \frac{1}{2}\right)|n\rangle,$$

we obtain:

$$\hat{n}|n\rangle = \hat{a}^\dagger\hat{a}|n\rangle = n|n\rangle.$$

From the normalization condition, it follows: $\langle n|n\rangle = 1$ and $S_n^2 = n$, thus $S_n = \sqrt{n}$ and hence:

$$\hat{a}|n\rangle = \sqrt{n}|n-1\rangle \quad (1.70)$$

Similarly, using the commutation relations (1.26), and 14.7 and 14.8 from section 14.1, we find

$$\begin{aligned}\hat{a}^\dagger|n\rangle &= S_{n+1}|n+1\rangle, \quad \langle n|\hat{a} = S_{n+1}\langle n+1|, \\ \langle n|\hat{a}\hat{a}^\dagger|n\rangle &= S_{n+1}^2\langle n+1|n+1\rangle = \langle n|\hat{a}^\dagger\hat{a} + 1|n\rangle = (n+1)\langle n|n\rangle, \\ S_{n+1}^2 &= n+1.\end{aligned}\quad (1.71)$$

Therefore, we have the equality

$$\hat{a}^\dagger |n\rangle = \sqrt{n+1} |n+1\rangle, \quad (1.72)$$

The eigenstates of the photon number operator \hat{n} are orthonormal. Indeed, from the fact that the operator \hat{n} is Hermitian:

$$\hat{n}^\dagger = (\hat{a}^\dagger \hat{a})^\dagger = \hat{a}^\dagger (\hat{a}^\dagger)^\dagger = \hat{a}^\dagger \hat{a} = \hat{n}$$

it follows (see [section 14.1](#)) that the eigenfunctions of this operator, corresponding to different eigenvalues, are orthogonal, i.e.

$$\langle n|n'\rangle = 0, \text{ if } n \neq n'. \quad (1.73)$$

Let's summarize the relations involving the operators \hat{a} and \hat{a}^\dagger :

$$\begin{aligned} \hat{\mathcal{H}} &= \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right), \quad \hat{a} |0\rangle = 0, \quad \hat{a}^\dagger \hat{a} |n\rangle = \hat{n} |n\rangle, \\ [\hat{a}, \hat{a}^\dagger] &= \hat{a} \hat{a}^\dagger - \hat{a}^\dagger \hat{a} = 1, \quad \hat{a} |n\rangle = \sqrt{n} |n-1\rangle \\ \hat{\mathcal{H}} |n\rangle &= \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right) |n\rangle, \\ \hat{a}^\dagger |n\rangle &= \sqrt{n+1} |n+1\rangle, \quad |n\rangle = \frac{1}{\sqrt{n!}} (\hat{a}^\dagger)^n |0\rangle \end{aligned} \quad (1.74)$$

and the conjugate equalities

$$\begin{aligned} \langle 0| \hat{a}^\dagger &= 0, \quad \langle n| \hat{a} = \sqrt{n+1} \langle n+1| \\ \langle n| \hat{a}^\dagger &= \sqrt{n} \langle n-1|, \quad \langle n| = \frac{1}{\sqrt{n!}} \langle 0| (\hat{a}^\dagger)^n. \end{aligned} \quad (1.75)$$

For the simplest resonator model, we have

$$\hat{E}(z, t) = E_1 (\hat{a} + \hat{a}^\dagger) \sin k_n z$$

where $E_1 = \sqrt{\frac{\hbar\omega}{\varepsilon_0 V}}$ is the field, corresponding to one photon in the mode.

Consider some properties of energy states, i.e., states with a definite number of photons. We will show that the average value of the electric field in this state is zero:

$$\begin{aligned} \langle n| \hat{E} |n\rangle &= E_1 \sin k_n z (\langle n| \hat{a} |n\rangle + \langle n| \hat{a}^\dagger |n\rangle) = \\ &= E_1 \sin k_n z (\langle n|n-1\rangle \sqrt{n} + \langle n|n+1\rangle \sqrt{n+1}) = 0 \end{aligned} \quad (1.76)$$

which follows from the orthogonality of states (1.73) $\langle n|n'\rangle = 0$.

The average value of the square of the electric field operator is non-zero:

$$\begin{aligned}\langle n|\hat{E}^2|n\rangle &= E_1^2 \sin^2 k_n z \langle n|(\hat{a}^\dagger \hat{a}^\dagger + \hat{a} \hat{a}^\dagger + \hat{a}^\dagger \hat{a} + \hat{a} \hat{a})|n\rangle = \\ &= 2E_1^2 \sin^2 k_n z \left(n + \frac{1}{2}\right).\end{aligned}\quad (1.77)$$

The fields usually considered in quantum optics are not in a stationary state with a definite energy (with a definite number of photons). However, an arbitrary state can be represented as a superposition of states $|n\rangle$:

$$|\psi\rangle = \sum_{(n)} C_n |n\rangle \quad (1.78)$$

where $|C_n|^2$ is the probability of finding n photons in the mode upon measurement; $\sum_{(n)} |C_n|^2 = 1$. In [section 14.1](#), it is shown that

$$C_n = \langle n|\psi\rangle, \quad |\psi\rangle = \sum_{(n)} \langle n|\psi\rangle |n\rangle = \sum_{(n)} |n\rangle \langle n|\psi\rangle, \quad \sum_{(n)} |n\rangle \langle n| = \hat{I},$$

where \hat{I} is the identity operator.

Remark 1.9.1 (On states with definite energy in quantum mechanics). *It is worth noting that states with definite energy do not violate Heisenberg's uncertainty relation for the pair energy - time*

$$\Delta E \Delta t \geq \frac{\hbar}{2},$$

(see [subsection 14.3.2](#) for more details).

Meanwhile, if we look at the energy operator of the harmonic oscillator, written as

$$\hat{\mathcal{H}} = \frac{1}{2} (\hat{p}^2 + \omega^2 \hat{q}^2)$$

and use expressions (1.24), we find that $\langle n|\hat{q}|n\rangle = \langle n|\hat{p}|n\rangle = 0$. Meanwhile,

$$\begin{aligned}\langle n|\hat{q}^2|n\rangle &= \frac{\hbar}{2\omega} \left[\langle n|\hat{a}^2|n\rangle + \langle n|(\hat{a}^\dagger)^2|n\rangle + \right. \\ &\quad \left. + \langle n|\hat{a}^\dagger \hat{a}|n\rangle + \langle n|\hat{a} \hat{a}^\dagger|n\rangle \right] = \\ &= \frac{\hbar}{2\omega} [n + n + 1],\end{aligned}$$

and also

$$\begin{aligned}\langle n | \hat{p}^2 | n \rangle &= -\frac{\hbar\omega}{2} \left[\langle n | \hat{a}^2 | n \rangle + \langle n | (\hat{a}^\dagger)^2 | n \rangle - \right. \\ &\quad \left. - \langle n | \hat{a}^\dagger \hat{a} | n \rangle - \langle n | \hat{a} \hat{a}^\dagger | n \rangle \right] = \\ &= \frac{\hbar\omega}{2} [n + n + 1].\end{aligned}$$

Thus,

$$\Delta p = \sqrt{\langle n | \hat{p}^2 | n \rangle - \langle n | \hat{p} | n \rangle^2} = \sqrt{\frac{\hbar\omega}{2} (2n + 1)}$$

and

$$\Delta q = \sqrt{\langle n | \hat{q}^2 | n \rangle - \langle n | \hat{q} | n \rangle^2} = \sqrt{\frac{\hbar}{2\omega} (2n + 1)}$$

or

$$\Delta p \Delta q = \frac{\hbar}{2} (2n + 1) \geq \frac{\hbar}{2},$$

which is consistent with Heisenberg's uncertainty relations (14.47).

Thus, states with definite energy are states in which, despite the impossibility of precisely determining p and q , it is possible to determine the value of $\frac{1}{2} (p^2 + \omega^2 q^2)$. This reflects the fact that for quantum systems, situations are possible where composite events exist in the absence of elementary ones (see [section 15.3](#)). In particular, such a situation occurs for the photon - states with definite energy $|1\rangle$.

This may lead us to consider that a photon is a virtual particle, which does not exist in the real physical world [17]. Meanwhile, the mathematical apparatus related to the concept of the photon such as creation operators \hat{a}^\dagger and annihilation \hat{a} , electromagnetic field states with definite energy $\{|n\rangle\}$ (which can be used as basis states), is convenient for theoretical descriptions.

It is also worth noting that from the perspective of the formal definition of a non-classical state (8.5), $|n\rangle$ is a non-classical state of light since from (5.29) it follows $G^{(2)} < 1$.

On the other hand, if we consider the minimum possible energy of the electromagnetic field mode, from the relation (1.19)

$$\mathcal{H} = \frac{1}{2} (\omega^2 q^2 + p^2)$$

it follows that in the classical case, the minimum possible zero energy is reached at $p = 0, q = 0$, but due to (14.47), zero values are impossible in the quantum

case, and considering measurement uncertainties we find

$$\begin{aligned} \frac{1}{2} (\omega^2 (\Delta q)^2 + (\Delta p)^2) &\geq \\ &\geq \omega \Delta q \Delta p \geq \frac{\hbar \omega}{2} \end{aligned}$$

that is, the minimum possible energy (vacuum energy) is determined by Heisenberg's inequalities for the coordinate - momentum pair.

1.10 Multimode States

So far we have considered single-mode fields. The results are easily generalized to a multimode state. For the electromagnetic field in some volume, we have the decomposition (1.11):

$$\vec{E}(r, t) = \sum_{(s)} \frac{q_s(t) \omega_s}{\sqrt{\varepsilon_0}} \vec{E}_s(r), \quad \vec{H}(r, t) = \sum_{(s)} \frac{p_s(t)}{\sqrt{\mu_0}} \vec{H}_s(r), \quad (1.79)$$

The Hamiltonian function is equal to the sum of the Hamiltonian functions of all modes, as the modes correspond to independent oscillators:

$$\mathcal{H} = \frac{1}{2} \sum_{(s)} (\omega_s^2 q_s^2 + p_s^2) \quad (1.80)$$

Quantization is reduced to replacing q_s and p_s with operators \hat{q}_s and \hat{p}_s with commutation relations:

$$[\hat{q}_s, \hat{p}_{s'}] = i\hbar \delta_{ss'}, \quad [\hat{q}_s, \hat{q}_{s'}] = [\hat{p}_s, \hat{p}_{s'}] = 0. \quad (1.81)$$

The Hamiltonian function thus becomes a Hamiltonian operator:

$$\hat{\mathcal{H}} = \frac{1}{2} \sum_{(s)} (\omega_s^2 \hat{q}_s^2 + \hat{p}_s^2) \quad (1.82)$$

As in section 1.3, we introduce creation and annihilation operators for each mode (1.23):

$$\hat{a}_s = \frac{1}{\sqrt{2\hbar\omega_s}} (\omega_s \hat{q}_s + i\hat{p}_s), \quad \hat{a}_s^\dagger = \frac{1}{\sqrt{2\hbar\omega_s}} (\omega_s \hat{q}_s - i\hat{p}_s), \quad (1.83)$$

With their help, the Hamiltonian can be represented as

$$\hat{\mathcal{H}}_s = \omega_s \hbar \left(\hat{a}_s^\dagger \hat{a}_s + \frac{1}{2} \right)$$

The general state, when there are n_1 photons in the first mode, n_2 in the second, and n_s photons in the s -th mode, etc., can be represented as the product of the state vectors of each mode (see [section 14.5](#)):

$$|\{n_s\}\rangle = |n_1, n_2, \dots, n_s, \dots\rangle = |n_1\rangle \otimes |n_2\rangle \otimes \dots \otimes |n_s\rangle \otimes \dots,$$

where $\{n_s\}$ denotes a set of mode occupation numbers.

The action of the operators \hat{a}_s^\dagger and \hat{a}_s , related to the s -th mode, on the state vector is described by the equalities

$$\begin{aligned} \hat{a}_s |n_1, n_2, \dots, n_s, \dots\rangle &= \sqrt{n_s} |n_1, n_2, \dots, n_s - 1, \dots\rangle, \\ \hat{a}_s^\dagger |n_1, n_2, \dots, n_s, \dots\rangle &= \sqrt{n_s + 1} |n_1, n_2, \dots, n_s + 1, \dots\rangle. \end{aligned} \quad (1.84)$$

The state vector in general can be represented as a linear superposition of states $|\{n_s\}\rangle$:

$$|\Psi\rangle = \sum_{n_1} \sum_{n_2} \dots \sum_{n_s} \dots C_{n_1, n_2, \dots, n_s, \dots} |n_1, n_2, \dots, n_s, \dots\rangle = \sum_{\{n_s\}} C_{\{n_s\}} |\{n_s\}\rangle \quad (1.85)$$

In the decomposition of the field into plane waves, the electric field operator has the form

$$\hat{\vec{E}} = \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \left(\hat{a}_k(t) \vec{e}_k e^{i(\vec{k}\vec{r})} + \hat{a}_k^\dagger(t) \vec{e}_k^* e^{-i(\vec{k}\vec{r})} \right) \quad (1.86)$$

This expression is written in the Heisenberg representation (the operator depends on time). In this representation

$$\hat{a}_k(t) = \hat{a}_k(0) e^{-i\omega_k t}, \quad \hat{a}_k^\dagger(t) = \hat{a}_k^\dagger(0) e^{i\omega_k t},$$

where $\hat{a}_k(0)$, $\hat{a}_k^\dagger(0)$ are operators in the Schrödinger representation. In the Schrödinger representation in expression (1.86), the operators do not depend on time. For each mode, we have the equalities (1.76), (1.77):

$$\langle n_k | \hat{a}_k | n_k \rangle = \langle n_k | \hat{a}_k^\dagger | n_k \rangle = 0$$

From here for the complete field we get

$$\langle \{n_k\} | \hat{\vec{E}} | \{n_k\} \rangle = 0, \quad \langle \{n_k\} | \hat{\vec{E}}^2 | \{n_k\} \rangle = \sum_{(k)} \frac{\hbar\omega_k}{\varepsilon_0 V} \left(n + \frac{1}{2} \right) \quad (1.87)$$

Expression (1.86) can be split into two terms: a positive frequency part, which includes annihilation operators, and a negative frequency part, which includes creation operators:

$$\begin{aligned}\hat{\vec{E}} &= \hat{\vec{E}}^{(+)} + \hat{\vec{E}}^{(-)}, \\ \hat{\vec{E}}^{(+)} &= \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \hat{a}_k(t) \vec{e}_k e^{i(\vec{k}\vec{r})}, \\ \hat{\vec{E}}^{(-)} &= \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \hat{a}_k^\dagger(t) \vec{e}_k^* e^{-i(\vec{k}\vec{r})}\end{aligned}\tag{1.88}$$

$\hat{\vec{E}}^{(+)}$ corresponds to the analytic signal in the classical case.

1.11 Coherent States

The states $|n\rangle$ do not approach the classical solution for the oscillator as n increases. This property is held by the so-called coherent states. They were introduced by E. Schrödinger and further developed in modern times by R. Glauber [56] and E. Sudarshan [47] in the context of quantum optics problems. A coherent state can be defined in various ways. We use the definition as a superposition of the states $|n\rangle$:

$$|\alpha\rangle = e^{-\frac{1}{2}|\alpha|^2} \sum_{(n)} \frac{\alpha^n}{\sqrt{n!}} |n\rangle, \quad \langle\alpha| = e^{-\frac{1}{2}|\alpha|^2} \sum_{(n)} \frac{\alpha^{*n}}{\sqrt{n!}} \langle n|, \tag{1.89}$$

where α is a certain complex parameter, whose meaning will be clarified later.

From the definition (1.89), several properties of coherent states follow. First of all,

$$\langle\alpha|\alpha\rangle = e^{-|\alpha|^2} \sum_{(n)} \frac{(\alpha\alpha^*)^n}{n!} = e^{|\alpha|^2} e^{-|\alpha|^2} = 1, \tag{1.90}$$

i.e., coherent states are normalized. However, these states are not orthogonal:

$$\langle\alpha|\beta\rangle = e^{-\frac{1}{2}(|\alpha|^2+|\beta|^2)} \sum_{(n)} \frac{(\alpha^*\beta)^n}{n!} = e^{-\frac{1}{2}|\alpha|^2-\frac{1}{2}|\beta|^2+\alpha^*\beta} \tag{1.91}$$

From this,

$$|\langle\alpha|\beta\rangle|^2 = e^{-|\alpha|^2-|\beta|^2+\alpha^*\beta+\alpha\beta^*} = e^{-|\alpha-\beta|^2}, \tag{1.92}$$

i.e., these states can be considered approximately orthogonal if $|\alpha - \beta|$ is sufficiently large.

The non-orthogonality of coherent states is a consequence of the over-completeness of the system (α is any complex number, while n is only integer numbers).

Coherent states are eigenstates of the annihilation operator with eigenvalue α :

$$\hat{a} |\alpha\rangle = e^{-\frac{1}{2}|\alpha|^2} \sum_{(n)} \frac{\alpha^n}{\sqrt{n!}} \sqrt{n} |n-1\rangle = e^{-\frac{1}{2}|\alpha|^2} \sum_{(m=n-1)} \frac{\alpha \alpha^m}{\sqrt{m!}} |m\rangle = \alpha |\alpha\rangle \quad (1.93)$$

The conjugate relation has the form

$$\langle \alpha | \hat{a}^\dagger = \alpha^* \langle \alpha | \quad (1.94)$$

The probability of detecting (measuring) n photons in a coherent state

$$|\langle n | \alpha \rangle|^2 = e^{-|\alpha|^2} \frac{(|\alpha|^2)^n}{n!} \quad (1.95)$$

corresponds to the Poisson distribution with an average number of photons $\langle n \rangle = |\alpha|^2$. Using the known relation

$$|n\rangle = \frac{1}{\sqrt{n!}} (\hat{a}^\dagger)^n |0\rangle \quad (1.96)$$

a coherent state can be represented as

$$|\alpha\rangle = e^{-\frac{1}{2}|\alpha|^2} \sum_{(n)} \frac{(\alpha \hat{a}^\dagger)^n}{n!} |0\rangle = e^{\alpha \hat{a}^\dagger - \frac{1}{2}|\alpha|^2} |0\rangle \quad (1.97)$$

We rewrite (1.97) in a different form, using the following relation

$$e^{-\alpha^* \hat{a}} |0\rangle = (1 - \alpha^* \hat{a} + \dots) |0\rangle = |0\rangle,$$

resulting in

$$|\alpha\rangle = e^{\alpha \hat{a}^\dagger} e^{-\frac{1}{2}|\alpha|^2} e^{-\alpha^* \hat{a}} |0\rangle. \quad (1.98)$$

Next, we use the operator identity (Baker-Campbell-Hausdorff formula)

$$e^{\hat{c} + \hat{d}} = e^{-\frac{1}{2}[\hat{c}, \hat{d}]} e^{\hat{c}} e^{\hat{d}} \quad (1.99)$$

valid under the condition $[\hat{c}, [\hat{c}, \hat{d}]] = [\hat{d}, [\hat{c}, \hat{d}]] = 0$. If we set $\hat{c} = \alpha \hat{a}^\dagger$, $\hat{d} = -\alpha^* \hat{a}$ then from (1.98) we obtain

$$\begin{aligned} |\alpha\rangle &= e^{\alpha \hat{a}^\dagger} e^{-\frac{1}{2}|\alpha|^2} e^{-\alpha^* \hat{a}} |0\rangle = e^{-\frac{1}{2}|\alpha|^2} e^{\alpha \hat{a}^\dagger} e^{-\alpha^* \hat{a}} |0\rangle = \\ &= e^{\alpha \hat{a}^\dagger - \alpha^* \hat{a}} |0\rangle. \end{aligned} \quad (1.100)$$

To visually represent coherent states, the coordinate representation of the coherent state is given in [section 14.6](#).

We provide some more useful relations derived from (1.97). By differentiating (1.97) with respect to α , we have:

$$\frac{\partial}{\partial \alpha} |\alpha\rangle = \left(\hat{a}^\dagger - \frac{1}{2} \alpha^* \right) |\alpha\rangle,$$

or otherwise

$$\hat{a}^\dagger |\alpha\rangle = \left(\frac{\partial}{\partial \alpha} + \frac{1}{2} \alpha^* \right) |\alpha\rangle.$$

the conjugate equality has the form

$$\langle \alpha | \hat{a} = \left(\frac{\partial}{\partial \alpha^*} + \frac{1}{2} \alpha \right) \langle \alpha |$$

In general, the operator of the positive frequency part of the electric field has the form

$$\hat{\vec{E}}^{(+)} = \sum_{(k)} \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} \hat{a}_k(t) \vec{e}_k e^{-i \omega_k t + i(\vec{k} \vec{r})},$$

Assume that the field (multimode) is in the state

$$|\{\alpha_s\}\rangle = |\alpha_1, \alpha_2, \dots, \alpha_s, \dots\rangle = |\alpha_1\rangle \otimes |\alpha_2\rangle \otimes \dots \otimes |\alpha_s\rangle \otimes \dots,$$

where $\{\alpha_s\} = \{\alpha_{k_s}\}$ indicates a certain set of parameters α . If the operator $\hat{\vec{E}}^{(+)}$ acts on the state $|\{\alpha_{k_s}\}\rangle$, we get

$$\begin{aligned} \hat{\vec{E}}^{(+)} |\{\alpha_{k_s}\}\rangle &= \left\{ \sum_{(k_s)} \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} \hat{a}_{k_s} \vec{e}_k e^{-i \omega_k t + i(\vec{k} \vec{r})} \right\} |\{\alpha_{k_s}\}\rangle = \\ &= \sum_{(k_s)} \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} \alpha_{k_s} \vec{e}_k e^{-i \omega_k t + i(\vec{k} \vec{r})} |\{\alpha_{k_s}\}\rangle, \end{aligned} \quad (1.101)$$

i.e., $|\{\alpha_{k_s}\}\rangle$ is an eigenvector of the operator of the positive-frequency part of the electric field, and the eigenvalue is the classical field (analytic signal), whose mode complex amplitudes are equal to $\sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} \alpha_k$. From this, it follows that each classical field corresponds to a certain coherent state.

The average value of the electric field operator $\hat{\vec{E}} = \hat{\vec{E}}^{(+)} + \hat{\vec{E}}^{(-)}$ in a coherent state is obtained by first considering one mode:

$$\langle \alpha_k | \hat{\vec{E}}_k | \alpha_k \rangle = 2 \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} |\alpha_k| \cos \left(\omega_k t - (\vec{k} \vec{r} + \theta_k) \right),$$

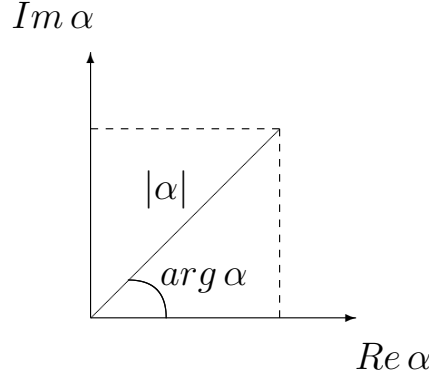


Figure 1.5: Complex plane on which coherent states are defined. Polar coordinates.

where $\alpha_k = |\alpha_k| e^{i\theta_k}$, as follows from the relations

$$\langle \alpha_k | \hat{a}_k | \alpha_k \rangle = \alpha_k, \quad \langle \alpha_k | \hat{a}_k^\dagger | \alpha_k \rangle = \alpha_k^*$$

Generalizing this to the multi-mode case, we obtain

$$\langle \{\alpha_k\} | \hat{\vec{E}} | \{\alpha_k\} \rangle = 2 \sum_{(k)} \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} |\alpha_k| \cos \left(\omega_k t - (\vec{k} \vec{r}) + \theta_k \right), \quad (1.102)$$

which corresponds to the classical multi-mode field, i.e., the average value of $\hat{\vec{E}}$ in the coherent state is a classical field with the corresponding amplitudes and phases of the modes.

Coherent states, as we have seen, are not orthogonal due to the over-completeness of the system of eigenvectors, but they satisfy the completeness condition:

$$\frac{1}{\pi} \int |\alpha\rangle \langle \alpha| d^2 \alpha = \hat{I}, \quad (1.103)$$

Here $d^2 \alpha = d \operatorname{Re} \alpha d \operatorname{Im} \alpha$, \hat{I} is the identity operator, and integration is carried out over the entire complex plane. We will prove this by using polar coordinates, depicted in [Figure 1.5](#).

Using the expansion of $|\alpha\rangle$ in terms of $|n\rangle$ in polar coordinates, we obtain

$$\begin{aligned}
& \int d^2\alpha |\alpha\rangle \langle\alpha| = \\
&= \int_0^\infty d|\alpha| \int_0^{2\pi} |\alpha| \sum_{(n)} \sum_{(m)} e^{-i(n-m)\varphi} e^{-|\alpha|^2} |m\rangle \langle n| \frac{|\alpha|^m |\alpha|^n}{\sqrt{n!m!}} d\varphi = \\
&= \pi \sum_{(n)} \frac{1}{n!} \int_0^\infty 2|\alpha| |\alpha|^{2n} e^{-|\alpha|^2} |n\rangle \langle n| d|\alpha| = \\
&= \pi \sum_{(n)} \frac{|n\rangle \langle n|}{n!} \int_0^\infty x^n e^{-x} dx = \pi \sum_{(n)} |n\rangle \langle n| = \pi \hat{I}, \quad (1.104)
\end{aligned}$$

where we have taken into account that

$$\int_0^{2\pi} e^{-i(n-m)\varphi} d\varphi = 2\pi \delta_{nm},$$

a change of variables was made:

$$|\alpha|^2 = x, \quad 2|\alpha| d|\alpha| = dx,$$

and we used the integral representation of the factorial:

$$n! = \int_0^\infty x^n e^{-x} dx$$

Thus, the formula (1.103) is proven. Using this relation, one coherent state can be expressed through all the others:

$$|\alpha\rangle = \frac{1}{\pi} \int |\beta\rangle \langle\beta| \alpha\rangle d^2\beta,$$

but in (1.91) we had $\langle\alpha| \beta\rangle = e^{-\frac{1}{2}|\alpha|^2 - \frac{1}{2}|\beta|^2 + \alpha^*\beta}$, therefore,

$$|\alpha\rangle = \frac{1}{\pi} \int |\beta\rangle e^{-\frac{1}{2}|\alpha|^2 - \frac{1}{2}|\beta|^2 + \alpha^*\beta} d^2\beta. \quad (1.105)$$

1.12 Mixed States of the Electromagnetic Field

So far, we have considered only pure states of the electromagnetic field described by state vectors. One of the most distinctive properties of pure states is the principle of superposition.

Only isolated systems can be pure, whereas mixed systems are not isolated from their environment. To describe mixed states, we use the density matrix apparatus. The density operator can be introduced as follows.

Consider a mixed state, for which the probabilities P_M of the field being in each of the states $|M\rangle$ are known. Let's calculate the average value of the operator \hat{O} corresponding to the observable O in this statistically mixed state. The average for the statistical mixture of states

$$\langle \hat{O} \rangle = \sum_{(M)} P_M \langle M | \hat{O} | M \rangle, \quad (1.106)$$

where it is clear that

$$\sum_{(M)} P_M = 1.$$

This expression can be represented differently. Using some complete set of states $|S\rangle$ and the condition of its completeness

$$\sum_{(S)} |S\rangle \langle S| = \hat{I},$$

we obtain

$$\langle \hat{O} \rangle = \sum_{(M)} P_M \sum_{(S)} \langle M | \hat{O} | S \rangle \langle S | M \rangle = \sum_{(M)} \sum_{(S)} P_M \langle S | M \rangle \langle M | \hat{O} | S \rangle. \quad (1.107)$$

Based on (1.107), the statistical operator $\hat{\rho}$ can be introduced using the expression

$$\hat{\rho} = \sum_{(M)} P_M |M\rangle \langle M|. \quad (1.108)$$

Then the equality (1.107) can be presented in the form

$$\langle \hat{O} \rangle = \sum_{(S)} \langle S | \hat{\rho} \hat{O} | S \rangle = Sp(\hat{\rho} \hat{O}), \quad (1.109)$$

This relation does not depend on the choice of $|S\rangle$, as the trace of the operator does not depend on the representation. Note that

$$Sp(\hat{\rho}) = \sum_{(M)} \langle M | \hat{\rho} | M \rangle = \sum_{(M)} P_M = 1 \quad (1.110)$$

As an example, consider the thermal excitation of photons in one mode:

$$\hat{\rho} = \sum_{(n)} P_n |n\rangle \langle n|, \quad (1.111)$$

where P_n is determined by the Boltzmann distribution

$$P_n = e^{-\beta n \hbar \omega} (1 - e^{-\beta \hbar \omega}),$$

here $\beta = \frac{1}{k_b T}$, and $(1 - e^{-\beta \hbar \omega})$ is the normalization factor.

The average number of photons in a mode for thermal excitation:

$$\begin{aligned} \bar{n} &= \langle \hat{n} \rangle = Sp(\hat{\rho} \hat{n}) = Sp(\hat{\rho} \hat{a}^\dagger \hat{a}) = \\ &= \sum_{(m)} \sum_{(n)} P_n \langle m | n \rangle \langle n | \hat{a}^\dagger \hat{a} | m \rangle = \\ &= \sum_{(n)} P_n \langle n | \hat{n} | n \rangle = \sum_{(n)} n e^{-\beta n \hbar \omega} (1 - e^{-\beta \hbar \omega}) = \frac{1}{e^{\beta \hbar \omega} - 1} \end{aligned} \quad (1.112)$$

This is the well-known expression obtained by Planck. Here, we used the following expression

$$\sum_{n=0}^{\infty} n r^{n-1} = \frac{d}{dr} \sum_{n=0}^{\infty} r^n = \frac{1}{(1-r)^2},$$

valid for $|r| < 1$.

The equalities (1.111), (1.112) can be written in another form. From (1.112) we get

$$e^{-\beta \hbar \omega} = \frac{\bar{n}}{1 + \bar{n}},$$

therefore,

$$P_n = e^{-\beta \hbar \omega n} (1 - e^{-\beta \hbar \omega}) = \frac{\bar{n}^n}{(1 + \bar{n})^{n+1}}. \quad (1.113)$$

In this case, the density matrix acquires the form:

$$\hat{\rho} = \sum_{(n)} \frac{\bar{n}^n}{(1 + \bar{n})^{n+1}} |n\rangle \langle n|.$$

This formula describes the density matrix of chaotic light through \bar{n} - the average number of photons in the mode. \bar{n} does not necessarily depend on the frequency according to Planck's formula (1.113). The dependence of \bar{n} on frequency determines the shape of the chaotic light line (different frequencies correspond to different modes). Generalizing the obtained result, we can write the expression for the statistical operator of chaotic light:

$$\hat{\rho} = \sum_{\{n_k\}} P_{\{n_k\}} |\{n_k\}\rangle \langle \{n_k\}| = \sum_{\{n_k\}} |\{n_k\}\rangle \langle \{n_k\}| \prod_{\{k\}} \frac{\bar{n}_k^{n_k}}{(1 + \bar{n}_k)^{n_k+1}}, \quad (1.114)$$

where \bar{n}_k depends on the frequency. For example, for chaotic light with a Lorentzian line,

$$\bar{n}_k = \frac{IS}{\hbar\omega_{k_0}} \frac{\gamma}{(\omega_{k_0} - \omega_k)^2 + \gamma^2},$$

where $\frac{IS}{\hbar\omega_{k_0}}$ is the average number of photons in the beam; I is the intensity of the beam (energy flow); S is the beam cross-section.

1.13 Representation of the Density Operator through Coherent States

An important aspect of quantum optics is the description of optical phenomena using coherent states. By utilizing the completeness condition twice for states with different numbers of photons, for an arbitrary statistical operator we get:

$$\hat{\rho} = \hat{I} \hat{\rho} \hat{I} = \sum_{(n)} \sum_{(m)} |n\rangle \langle n| \hat{\rho} |m\rangle \langle m| = \sum_{(n)} \sum_{(m)} \rho_{mn} |n\rangle \langle m|$$

Similarly for coherent states

$$\begin{aligned} \hat{\rho} &= \hat{I} \hat{\rho} \hat{I} = \frac{1}{\pi^2} \int d^2\alpha \int |\alpha\rangle \langle \alpha| \hat{\rho} |\beta\rangle \langle \beta| d^2\beta = \\ &= \frac{1}{\pi^2} \int d^2\alpha \int R(\alpha^*, \beta) e^{-\frac{1}{2}|\alpha|^2 - \frac{1}{2}|\beta|^2} |\alpha\rangle \langle \beta| d^2\beta, \end{aligned} \quad (1.115)$$

where

$$R(\alpha^*, \beta) = \langle \alpha | \hat{\rho} | \beta \rangle e^{\frac{1}{2}|\alpha|^2 + \frac{1}{2}|\beta|^2}$$

If we express $|\alpha\rangle$ and $|\beta\rangle$ through states $|n\rangle$, we obtain the formula

$$R(\alpha^*, \beta) = \sum_{(n)} \sum_{(m)} \frac{(\alpha^*)^n (\beta)^m}{\sqrt{n!m!}} \rho_{nm} \quad (1.116)$$

i.e., $R(\alpha^*, \beta)$ is easily found if the density matrix is known in the occupation number representation (photon numbers). As an example, consider the thermal excitation of a mode:

$$\begin{aligned} R(\alpha^*, \beta) &= \sum_{(n)} \sum_{(m)} \frac{(\alpha^*)^n (\beta)^m}{\sqrt{n!m!}} e^{-\beta n \hbar \omega} (1 - e^{-\beta \hbar \omega}) \delta_{nm} = \\ &= \sum_{(n)} \frac{(\alpha^* \beta)^n}{n!} e^{-\beta n \hbar \omega} (1 - e^{-\beta \hbar \omega}) = (1 - e^{-\beta \hbar \omega}) e^{\alpha^* \beta e^{-\beta \hbar \omega}}. \end{aligned} \quad (1.117)$$

This representation depends on two parameters α^* and β . A more convenient representation of the density matrix, introduced by R. Glauber [56], called the diagonal representation and depending on one parameter, is of the form

$$\hat{\rho} = \int P(\alpha) |\alpha\rangle \langle \alpha| d^2\alpha, \quad (1.118)$$

where $P(\alpha)$ is a real function that satisfies the condition

$$\int P(\alpha) d^2\alpha = 1.$$

Moreover, $P(\alpha)$ is a real function of a complex argument. This all follows from the conditions $Sp(\hat{\rho}) = 1$ and $\hat{\rho} = \hat{\rho}^\dagger$.

Such decomposition is possible due to the overcompleteness of the coherent state system. If $P(\alpha)$ is a positive function, it can be interpreted as a probability distribution. This applies to some of the most practically interesting field states, such as the completely chaotic state, but this is not true in the general case. Sometimes $P(\alpha)$ may be negative over a limited range of α , then it cannot be interpreted as a probability distribution. $P(\alpha)$ may also be a generalized function (an example being the δ -function).

A rigorous justification for the possibility of the representation (1.118) is contained in the literature [56], [47]. The possibility of introducing the diagonal representation (1.118) can be justified as follows. Suppose the statistical operator $\hat{\rho}$, like any other operator acting on the electromagnetic field, can be represented as a function of creation and annihilation operators:

$$\hat{\rho} = \bar{f}(\hat{a}^\dagger, \hat{a}).$$

Then the operator can be represented in ordered form: normal and antinormal. In the first case, creation operators are positioned to the right of annihilation operators, and in the second case, the opposite - on the left. For example,

$$(\hat{a}^\dagger)^m (\hat{a})^n$$

is a normally ordered operator. In the antinormally ordered operator, the order is reversed: creation operators are on the right of the annihilation operator. For example, the operator

$$(\hat{a})^n (\hat{a}^\dagger)^m$$

is antinormally ordered. Ordering can be achieved by multiple applications of the commutation relation. For example, the operator

$$(\hat{a}^\dagger \hat{a})^2 = \hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{a}$$

is neither normal nor antinormal. Using the condition $[\hat{a}, \hat{a}^\dagger] = 1$, bring it to normal form:

$$(\hat{a}^\dagger \hat{a})^2 = \hat{a}^\dagger (1 + \hat{a}^\dagger \hat{a}) \hat{a} = \hat{a}^\dagger \hat{a} + (\hat{a}^\dagger)^2 (\hat{a})^2$$

and this operator can be presented in antinormal form:

$$\begin{aligned} (\hat{a}^\dagger \hat{a})^2 &= (\hat{a} \hat{a}^\dagger - 1) (\hat{a} \hat{a}^\dagger - 1) = \hat{a} \hat{a}^\dagger \hat{a} \hat{a}^\dagger - 2 \hat{a} \hat{a}^\dagger + 1 = \\ &= \hat{a} (\hat{a} \hat{a}^\dagger - 1) \hat{a}^\dagger - 2 \hat{a} \hat{a}^\dagger + 1 = (\hat{a})^2 (\hat{a}^\dagger)^2 - 3 \hat{a} \hat{a}^\dagger + 1. \end{aligned}$$

There are also other, more efficient methods of operator ordering [61].

Let's represent the density operator in antinormal form (considering the single-mode case):

$$\hat{\rho}^{(a)}(\hat{a}, \hat{a}^\dagger) = \sum_{(n)} \sum_{(m)} C_{nm}^{(a)} (\hat{a})^n (\hat{a}^\dagger)^m \quad (1.119)$$

Now use the expansion (1.103) for the unit operator \hat{I}

$$\hat{I} = \frac{1}{\pi} \int |\alpha\rangle \langle \alpha| d^2 \alpha$$

as a result, we get

$$\begin{aligned} \hat{\rho}^{(a)}(\hat{a}, \hat{a}^\dagger) &= \sum_{(n)} \sum_{(m)} C_{nm}^{(a)} (\hat{a})^n \hat{I} (\hat{a}^\dagger)^m = \\ &= \frac{1}{\pi} \sum_{(n)} \sum_{(m)} C_{nm}^{(a)} \int d^2 \alpha (\hat{a})^n |\alpha\rangle \langle \alpha| (\hat{a}^\dagger)^m = \\ &= \frac{1}{\pi} \sum_{(n)} \sum_{(m)} \int d^2 \alpha C_{nm}^{(a)} \alpha^n \alpha^{*m} |\alpha\rangle \langle \alpha|. \end{aligned} \quad (1.120)$$

Now denote

$$P(\alpha, \alpha^*) = \frac{1}{\pi} \sum_{(n)} \sum_{(m)} C_{nm}^{(a)} \alpha^n \alpha^{*m}$$

and obtain that $\hat{\rho}$ can be represented in the form

$$\hat{\rho} = \int d^2 \alpha P(\alpha, \alpha^*) |\alpha\rangle \langle \alpha|. \quad (1.121)$$

The average value of the operator using the diagonal representation

$$\begin{aligned}
\langle \hat{O} \rangle &= Sp(\hat{\rho} \hat{O}) = \\
&= \sum_{(n)} \langle n | \int d^2 \alpha P(\alpha, \alpha^*) |\alpha\rangle \langle \alpha| \hat{O} |n\rangle = \\
&= \sum_{(n)} \int d^2 \alpha P(\alpha, \alpha^*) \langle \alpha | \hat{O} |n\rangle \langle n | \alpha \rangle = \\
&= \int d^2 \alpha P(\alpha, \alpha^*) \langle \alpha | \hat{O} | \alpha \rangle, \tag{1.122}
\end{aligned}$$

since $\sum_{(n)} |n\rangle \langle n| = \hat{I}$.

The expression $\langle \alpha | \hat{O} | \alpha \rangle$ is easily calculated if the operator \hat{O} is represented in normal form:

$$\hat{O} = \hat{O}^{(n)} = \sum_{(n)} \sum_{(m)} d_{nm} (\hat{a}^\dagger)^n (\hat{a})^m, \tag{1.123}$$

then

$$\langle \alpha | \hat{O} | \alpha \rangle = \sum_{(n)} \sum_{(m)} d_{nm} \langle \alpha | (\hat{a}^\dagger)^n (\hat{a})^m | \alpha \rangle = \sum_{(n)} \sum_{(m)} d_{nm} \alpha^{*n} \alpha^m. \tag{1.124}$$

From here, we derive a simple rule: to obtain the required matrix element, in the operator represented in normally ordered form, make the replacement $\hat{a} \rightarrow \alpha$, $\hat{a}^\dagger \rightarrow \alpha^*$.

As an example, we show that for the thermal excitation of a mode

$$P(\alpha, \alpha^*) = \frac{1}{\pi \bar{n}} e^{-\frac{|\alpha|^2}{\bar{n}}} \tag{1.125}$$

and, therefore

$$\hat{\rho} = \frac{1}{\pi \bar{n}} \int d^2 \alpha e^{-\frac{|\alpha|^2}{\bar{n}}} |\alpha\rangle \langle \alpha|.$$

Now find the matrix elements of the operator $\hat{\rho}$ in the occupation number representation (photon numbers)

$$\begin{aligned}
\langle n | \hat{\rho} | n \rangle &= \frac{1}{\pi \bar{n}} \int d^2 \alpha e^{-\frac{|\alpha|^2}{\bar{n}}} \langle n | \alpha \rangle \langle \alpha | n \rangle = \\
&= \frac{1}{\pi \bar{n}} \int d^2 \alpha e^{-\frac{|\alpha|^2}{\bar{n}}} \frac{|\alpha|^{2n}}{n!} e^{-|\alpha|^2}
\end{aligned}$$

The resulting integral can be computed by writing in polar coordinate system:

$$\alpha = r e^{i\varphi}, \quad d^2 \alpha = r dr d\varphi, \quad \text{i.e. } |\alpha| = r, \quad \arg \alpha = \varphi,$$

from where

$$\begin{aligned}
& \frac{1}{\pi \bar{n}} \int d^2 \alpha e^{-\frac{|\alpha|^2}{\bar{n}}} \frac{|\alpha|^{2n}}{n!} e^{-|\alpha|^2} = \\
&= \frac{1}{\pi \bar{n}} \int_0^{2\pi} d\varphi \int_0^\infty r dr \frac{\exp\left(-r^2 \frac{\bar{n}+1}{\bar{n}}\right)}{n!} r^{2n} = \\
&= \frac{1}{\bar{n}} \int_0^\infty 2r dr \frac{\exp\left(-r^2 \frac{\bar{n}+1}{\bar{n}}\right)}{n!} r^{2n} = \\
&= \frac{1}{\bar{n}} \int_0^\infty \frac{e^{-x}}{n!} \frac{\bar{n}^{n+1}}{(\bar{n}+1)^{n+1}} x^n dx = \\
&= \frac{\bar{n}^n}{(\bar{n}+1)^{n+1}} \frac{1}{n!} \int_0^\infty e^{-x} x^n dx = \frac{\bar{n}^n}{(\bar{n}+1)^{n+1}}, \tag{1.126}
\end{aligned}$$

where the substitution $x = r^2 \frac{\bar{n}+1}{\bar{n}}$ was made.

Thus, we obtained the well-known matrix element of the statistical operator of the chaotic field in the occupation number representation. From here, it follows that expression (1.125) is indeed the statistical operator of the chaotic field in the representation of coherent states. Therefore, for the thermal excitation of photons in the mode we have

$$P(\alpha) = \frac{1}{\pi \bar{n}} e^{-\frac{|\alpha|^2}{\bar{n}}}, \text{ where } \bar{n} = \frac{1}{e^{\hbar\omega\beta} - 1} \tag{1.127}$$

The generalization of the obtained expressions to the multimode case is obvious:

$$\begin{aligned}
\langle \hat{O} \rangle &= \int \cdots \int_{\{\alpha_k\}} P^{(a)}(\{\alpha_k\}) O^{(n)}(\{\alpha_k\}) \prod_k d^2 \alpha_k, \tag{1.128} \\
P^{(a)}(\{\alpha_k\}) &= \prod_k \frac{1}{\pi \bar{n}_k} e^{-\frac{|\alpha_k|^2}{\bar{n}_k}}.
\end{aligned}$$

1.14 Exercises

1. Prove the orthogonality of the fields of eigen-oscillations (1.14).
2. Prove for plane waves the equalities (1.52).
3. Based on the equality (1.57), derive the expression (1.58) for the momentum operator of the quantized electromagnetic field.
4. Prove the Baker-Hausdorff formula (1.99)

5. How many modes of the electromagnetic field with wavelength $\lambda \geq 500\text{nm}$ are in a quantization cube with side $L = 1\text{mm}$ [14]

Chapter 2

Interaction of Light with an Atom

Questions related to the interaction of quantized electromagnetic fields with an atom are considered. A simplified model of an atom is used - a two-level atom. Such simplification is justified in resonant interaction and is widely used in problems of quantum electronics and quantum optics. Great attention is paid to the consideration of the interaction of an atom, a mode of a resonator (a dynamic system) with a thermostat (a dissipative system), which is responsible for the relaxation of the dynamic system.

2.1 Radiation and Absorption of Light by an Atom

Consider the simplest problem of interaction between a single-mode radiation and a two-level atom. The simplifications are justified because laser radiation can often be considered as single-mode, and in resonant interaction, all levels of the atom other than two, whose transition frequency is close to the frequency of the mode interacting with the atom, can be neglected. This simplest situation is depicted in [Figure 2.1](#)

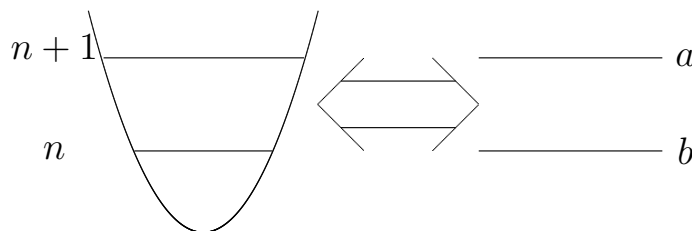


Figure 2.1: Interaction of the electromagnetic field mode with a two-level atom.

Assume that the atom and the field are initially in the states

$$|\psi_A\rangle = C_a |a\rangle + C_b |b\rangle, |\psi_F\rangle = \sum_{(n)} C_n |n\rangle \quad (2.1)$$

where: $|a\rangle, |b\rangle$ are the vectors of the upper and lower states of the atom, respectively, $|n\rangle$ is the state vector of the mode containing n photons.

The total state vector of the atom-field system is

$$|\psi_{AF}\rangle = \sum_{(n)} (C_{an} |a\rangle |n\rangle + C_{bn} |b\rangle |n\rangle). \quad (2.2)$$

Here C_{an}, C_{bn} are the corresponding probability amplitudes. Due to the interaction between the atom and the field, the probability amplitudes change over time. For example, if the atom was initially at the upper level and the mode field contained n photons, i.e., $|\psi_{AF}\rangle = |a\rangle |n\rangle$, then at subsequent moments in time the field will be in the state

$$|\psi_{AF}\rangle = C_{an} |a\rangle |n\rangle + C_{b,n+1} |b\rangle |n+1\rangle. \quad (2.3)$$

2.2 Hamiltonian of Atom-Field System

The Hamiltonian of the atom-field system can be represented as

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}^{(A)} + \hat{\mathcal{H}}^{(F)} + \hat{V}$$

where $\hat{\mathcal{H}}^{(A)}$ is the Hamiltonian of a two-level atom, $\hat{\mathcal{H}}^{(F)}$ is the Hamiltonian of the electromagnetic field, and \hat{V} is the interaction Hamiltonian.

For the electromagnetic field Hamiltonian, we have the expression (1.28):

$$\hat{\mathcal{H}}^{(F)} = \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right).$$

In the case of the two-level atom Hamiltonian, we use the following relations:

$$\begin{aligned} \hat{\mathcal{H}}^{(A)} |a\rangle &= \hbar\omega_a |a\rangle, \\ \hat{\mathcal{H}}^{(A)} |b\rangle &= \hbar\omega_b |b\rangle, \\ |a\rangle \langle a| + |b\rangle \langle b| &= \hat{I}, \end{aligned}$$

from which we obtain

$$\begin{aligned} \hat{\mathcal{H}}^{(A)} &= \hat{\mathcal{H}}^{(A)} \hat{I} = \hat{\mathcal{H}}^{(A)} |a\rangle \langle a| + \hat{\mathcal{H}}^{(A)} |b\rangle \langle b| = \\ &= \hbar\omega_a |a\rangle \langle a| + \hbar\omega_b |b\rangle \langle b| = \\ &= \hbar\omega_a \hat{\sigma}_a + \hbar\omega_b \hat{\sigma}_b, \end{aligned} \quad (2.4)$$

where the following notations are introduced

$$\hat{\sigma}_a = |a\rangle \langle a|, \quad \hat{\sigma}_b = |b\rangle \langle b|.$$

Expression (2.4) can be rewritten in matrix form:

$$\hat{\mathcal{H}}^{(A)} = \hbar \begin{pmatrix} \omega_a & 0 \\ 0 & \omega_b \end{pmatrix}. \quad (2.5)$$

For the interaction Hamiltonian \hat{V} in the semiclassical approach under the dipole approximation, we have:

$$\hat{V} = -e \left(\hat{\vec{r}} \vec{E}(t) \right). \quad (2.6)$$

Applying the completeness condition twice to (2.6) we obtain

$$\begin{aligned} \hat{V} &= \hat{I} \hat{V} \hat{I} = -e \left(\vec{E}(t) \hat{I} \hat{\vec{r}} \hat{I} \right) = \\ &= -e \left(\vec{E}(t) (|a\rangle \langle a| + |b\rangle \langle b|) \hat{\vec{r}} (|a\rangle \langle a| + |b\rangle \langle b|) \right) = \\ &= - \left(\vec{E}(t) \left(\vec{P}_{ab} |a\rangle \langle b| + \vec{P}_{ba} |b\rangle \langle a| \right) \right), \end{aligned} \quad (2.7)$$

where

$$\vec{P}_{ab} = \vec{P}_{ba}^* = e \langle a | \hat{\vec{r}} | b \rangle$$

is the matrix element of the electric dipole moment. For simplicity, we will assume $\vec{P}_{ab} = \vec{P}_{ba} = \vec{p}$ - a real quantity. As a result, the expression (2.7) can be written as

$$\begin{aligned} \hat{V} &= - \left(\vec{E}(t) \vec{p} \right) (|a\rangle \langle b| + |b\rangle \langle a|) = \\ &= - \left(\vec{E}(t) \vec{p} \right) (\hat{\sigma}^\dagger + \hat{\sigma}), \end{aligned} \quad (2.8)$$

where $\hat{\sigma}^\dagger = |a\rangle \langle b|$ and $\hat{\sigma} = |b\rangle \langle a|$ are transition operators. $\hat{\sigma}^\dagger$ is the transition operator from the lower state to the upper one (raising operator):

$$\hat{\sigma}^\dagger |b\rangle = |a\rangle \langle b|b\rangle = |a\rangle,$$

and $\hat{\sigma}$ is the operator for the transition from the upper state to the lower (lowering operator):

$$\hat{\sigma} |a\rangle = |b\rangle \langle a|a\rangle = |b\rangle.$$

In addition, the equalities hold

$$\hat{\sigma}^\dagger |a\rangle = |a\rangle \langle b|a\rangle = 0$$

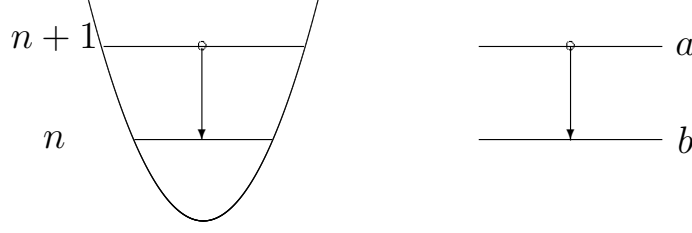


Figure 2.2: Hamiltonian of interaction. The process described by $\hat{a}\hat{\sigma}$ - absorption of a photon and transition of the atom from the upper state to the lower state.

and

$$\hat{\sigma} |b\rangle = |b\rangle \langle a|b\rangle = 0.$$

For operators $\hat{\sigma}$ and $\hat{\sigma}^\dagger$ the following relation can be written:

$$\begin{aligned} \hat{\sigma}\hat{\sigma}^\dagger + \hat{\sigma}^\dagger\hat{\sigma} &= \\ &= |b\rangle \langle a|a\rangle \langle b| + |a\rangle \langle b|b\rangle \langle a| = \\ &= |b\rangle \langle b| + |a\rangle \langle a| = \hat{I}, \end{aligned} \quad (2.9)$$

where \hat{I} is the identity operator.

Expression (2.8) can also be written in matrix form

$$\hat{V} = - \left(\vec{p} \vec{E} \right) \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}. \quad (2.10)$$

The transition from the semiclassical interaction Hamiltonian to the fully quantum expression is made by replacing the classical electric field with the electric field operator. Thus, we obtain

$$\hat{V} = -pE_1 \sin kz (\hat{a} + \hat{a}^\dagger) (\hat{\sigma} + \hat{\sigma}^\dagger) \quad (2.11)$$

Here, the simplest expression for the electric field operator of the mode (1.32) was used.

The complete Hamiltonian of the atom-field system is

$$\hat{\mathcal{H}}_{AF} = \hbar\omega_a \sigma_a + \hbar\omega_b \sigma_b + \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right) + \hbar g (\hat{a} + \hat{a}^\dagger) (\hat{\sigma} + \hat{\sigma}^\dagger)$$

where $g = -\frac{p}{\hbar} E_1 \sin kz$ is the interaction constant, dependent on the mode field distribution in space.

Not all terms in the interaction Hamiltonian (2.11) are equivalent. When multiplying the brackets, we obtain terms:

1. $\hat{a}\hat{\sigma}$ - This term corresponds to the absorption of a photon and the transition of the atom from the upper state to the lower state. This process is conditionally depicted in Figure 2.2;

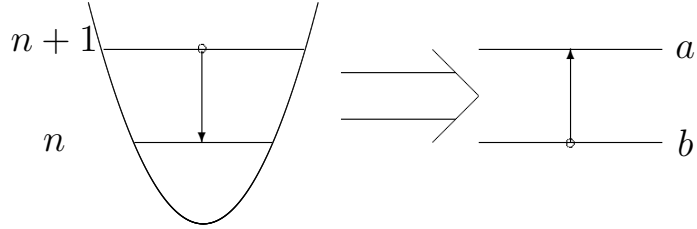


Figure 2.3: Hamiltonian of interaction. The process described by $\hat{a}\hat{\sigma}^\dagger$ - absorption of a photon and transition of the atom from the lower state to the upper state.

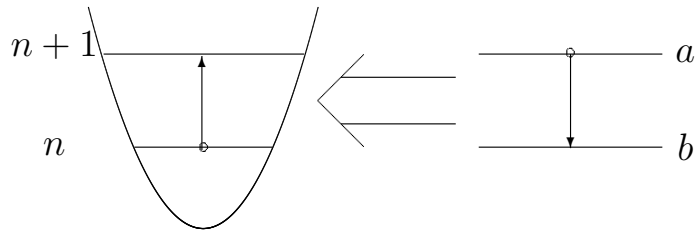


Figure 2.4: Hamiltonian of interaction. The process described by $\hat{a}^\dagger\hat{\sigma}$ - emission of a photon and transition of the atom from the upper state to the lower state.

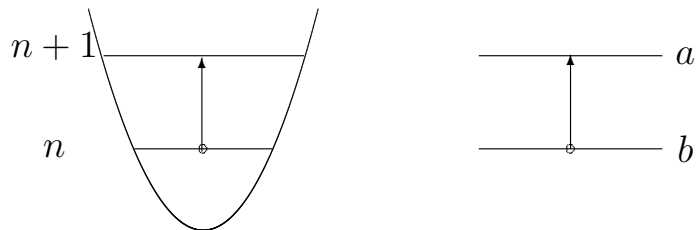


Figure 2.5: Hamiltonian of interaction. The process described by $\hat{a}^\dagger\hat{\sigma}^\dagger$ - emission of a photon and transition of the atom from the lower state to the upper state.

2. $\hat{a}\hat{\sigma}^\dagger$ - This term corresponds to the absorption of a photon and the transition of the atom from the lower state to the upper state (see [Figure 2.3](#));
3. $\hat{a}^\dagger\hat{\sigma}$ - This term corresponds to the emission of a photon and the transition of the atom from the upper state to the lower state (see [Figure 2.4](#));
4. $\hat{a}^\dagger\hat{\sigma}^\dagger$ - This term corresponds to the emission of a photon and the transition of the atom from the lower state to the upper state (see [Figure 2.5](#)).

Cases 1 and 4 refer to processes (see [Figure 2.2](#) and [2.5](#)) in which the law of energy conservation is violated. These processes can be attributed to the so-called virtual processes. Their probability approaches 0 over time, and their contribution to the interaction Hamiltonian can be neglected. The remaining terms 2 and 3 (see [Figure 2.3](#) and [2.4](#)) correspond to processes that occur without violating the law of energy conservation. Considering all that was said, the interaction Hamiltonian takes the form:

$$\hat{V} = g\hbar (\hat{a}\hat{\sigma}^\dagger + \hat{a}^\dagger\hat{\sigma}) \quad (2.12)$$

It is convenient to transition to the interaction representation (see [section 14.2.1](#)), using the known formula ([14.31](#))

$$\hat{V}_I = e^{i\frac{\hat{\mathcal{H}}_0 t}{\hbar}} \hat{V} e^{-i\frac{\hat{\mathcal{H}}_0 t}{\hbar}} \quad (2.13)$$

where $\hat{\mathcal{H}}_0 = \hat{\mathcal{H}}^{(A)} + \hat{\mathcal{H}}^{(F)}$ is the unperturbed Hamiltonian of the atom-field system. It can be shown that in the interaction representation the interaction Hamiltonian takes the form

$$\hat{V}_I = g\hbar \left(\hat{a}^\dagger\hat{\sigma} e^{i(\omega - \omega_{ab})t} + \hat{\sigma}^\dagger\hat{a} e^{-i(\omega - \omega_{ab})t} \right). \quad (2.14)$$

This is most easily done by calculating the matrix elements $\langle n+1 | \langle b | \hat{V}_I | a \rangle | n \rangle$ and $\langle n | \langle a | \hat{V}_I | b \rangle | n+1 \rangle$ for the interaction operator \hat{V}_I , represented in the form ([2.13](#)) and ([2.14](#)), and verifying that they are identical in both cases.

2.3 Interaction of an atom with the mode of the electromagnetic field

The Schrödinger equation in the interaction representation has the form

$$i\hbar \frac{\partial}{\partial t} |\psi(r, t)\rangle_I = \hat{V}_I |\psi(r, t)\rangle_I, \quad (2.15)$$

where $|\psi(r, t)\rangle_I$ is the wave function in the interaction representation:

$$|\psi(r, t)\rangle_I = e^{i\frac{\hat{H}_0 t}{\hbar}} |\psi(r, t)\rangle. \quad (2.16)$$

We solve this equation for the initial state $|\psi(r, 0)\rangle_I = |b, n+1\rangle_I$, i.e., we assume that initially, the atom is in the lower state, and the field contains $n+1$ photons. Another case corresponds to the initial state of the system $|a, n\rangle_I$, i.e., the atom is in the upper state, and the field contains n photons. After some time, as a result of the interaction between the atom and the field, the system will be in a superposition state

$$|\psi(r, t)\rangle_I = C_{a,n}(t)_I |a, n\rangle_I + C_{b,n+1}(t)_I |b, n+1\rangle_I. \quad (2.17)$$

Using (2.16) in the left part of the expression (2.17), we get

$$\begin{aligned} & C_{a,n}(t)_I |a, n\rangle_I + C_{b,n+1}(t)_I |b, n+1\rangle_I = \\ & = C_{a,n}(t)_I e^{i\frac{\hat{H}_0 t}{\hbar}} |a, n\rangle + C_{b,n+1}(t)_I e^{i\frac{\hat{H}_0 t}{\hbar}} |b, n+1\rangle = \\ & = C_{a,n}(t)_I \exp\left\{i\left(\omega_a + \omega\left(n + \frac{1}{2}\right)\right)t\right\} |a, n\rangle + \\ & + C_{b,n+1}(t)_I \exp\left\{i\left(\omega_b + \omega\left(n + 1 + \frac{1}{2}\right)\right)t\right\} |b, n+1\rangle \\ & = C_{a,n}(t) |a, n\rangle + C_{b,n+1}(t) |b, n+1\rangle, \end{aligned}$$

where $C_{a,n}(t)$ and $C_{b,n+1}(t)$ are denoted as

$$\begin{aligned} C_{a,n}(t) &= C_{a,n}(t)_I \exp\left\{i\left(\omega_a + \omega\left(n + \frac{1}{2}\right)\right)t\right\}, \\ C_{b,n+1}(t) &= C_{b,n+1}(t)_I \exp\left\{i\left(\omega_b + \omega\left(n + 1 + \frac{1}{2}\right)\right)t\right\}. \end{aligned}$$

Thus, equation (2.15) takes the following form

$$\begin{aligned} & i\left(\dot{C}_{a,n}(t) |a, n\rangle + \dot{C}_{b,n+1}(t) |b, n+1\rangle\right) = \\ & = g\left(\hat{a}\hat{\sigma}^\dagger e^{-i(\omega-\omega_{ab})t} + \hat{a}^\dagger\hat{\sigma} e^{i(\omega-\omega_{ab})t}\right)(C_{a,n} |a, n\rangle + C_{b,n+1} |b, n+1\rangle) = \\ & = g\sqrt{n+1}\left(C_{b,n+1} e^{-i(\omega-\omega_{ab})t} |a, n\rangle + C_{a,n} e^{i(\omega-\omega_{ab})t} |b, n+1\rangle\right). \quad (2.18) \end{aligned}$$

Multiply this equation on the left by $\langle a, n|$ and $\langle b, n+1|$ respectively. Considering the properties of the operators \hat{a} , \hat{a}^\dagger , $\hat{\sigma}$, $\hat{\sigma}^\dagger$, and the orthogonality of state vectors, we obtain the system of equations for the probability amplitudes

$$\begin{aligned} \dot{C}_{a,n}(t) &= -ig\sqrt{n+1}e^{-i(\omega-\omega_{ab})t}C_{b,n+1}(t), \\ \dot{C}_{b,n+1}(t) &= -ig\sqrt{n+1}e^{i(\omega-\omega_{ab})t}C_{a,n}(t). \end{aligned} \quad (2.19)$$

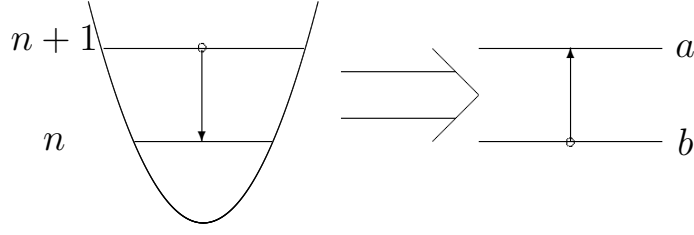


Figure 2.6: Absorption of a photon by a two-level atom. As a result of the interaction, the atom transitions from the lower level to the upper level, and there is one less photon in the mode of the electromagnetic field.

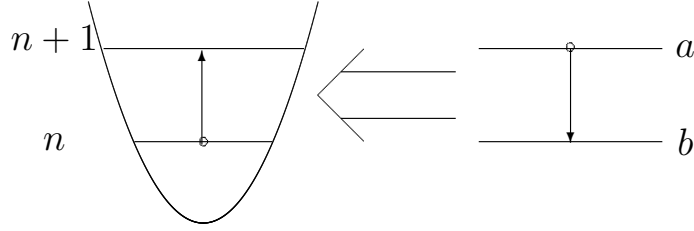


Figure 2.7: Emission of a photon by a two-level atom. As a result of the interaction, the atom transitions from the upper level to the lower one, and the mode of the electromagnetic field has one more photon.

For the case when the atom is in the lower state (photon absorption case), initially we have $C_{a,n}(0) = 0$, $C_{b,n+1}(0) = 1$ (Figure 2.6). Formal integration of (2.19) gives

$$C_{a,n}(t) = -ig\sqrt{n+1} \int_0^t e^{-i(\omega-\omega_{ab})t'} C_{b,n+1}(t') dt' \quad (2.20)$$

In the first approximation for small times, we can assume under the integral in (2.20) $C_{b,n+1}(t') = 1$. Integrating, we obtain

$$C_{a,n}(t) = -ig\sqrt{n+1} \frac{e^{-i(\omega-\omega_{ab})t} - 1}{-i(\omega - \omega_{ab})} \quad (2.21)$$

from which for the probability of exciting the atom and absorbing a photon, we have

$$|C_{a,n}(t)|^2 = g^2(n+1)t^2 \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2(t/2)^2}. \quad (2.22)$$

In this approximation, we can also solve the second problem (Figure 2.7) $C_{a,n}(0) = 1$, $C_{b,n+1}(0) = 0$. We obtain for the probability of the atom transitioning to the lower state and emitting a photon, an expression similar to (2.22)

$$|C_{b,n+1}(t)|^2 = g^2(n+1)t^2 \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2(t/2)^2}. \quad (2.23)$$

In the case when the field does not have any photons ($n = 0$), we obtain a non-zero probability $|C_{b,1}(t)|^2 \cong g^2 t^2 (\frac{\sin^2 x}{x^2} \cong 1, \text{ if } x \ll 1)$, from which it is seen that an excited atom can transition to the lower state even in the absence of photons. This process is called spontaneous emission or spontaneous transition. It is associated with zero point field fluctuations. Note that in a semiclassical consideration, spontaneous emission does not follow from semiclassical equations and is introduced from additional considerations. From the quantum equations, spontaneous emission follows naturally. The system of equations (2.19) can be solved exactly. For simplicity, we will consider the resonant case $\omega = \omega_{ab}$. Eliminating $C_{b,n+1}(t)$ from the equations:

$$\ddot{C}_{a,n}(t) = -ig\sqrt{n+1}\dot{C}_{b,n+1}(t) = -g^2(n+1)C_{a,n}(t).$$

Solution of the obtained equation:

$$C_{a,n}(t) = A \sin(g\sqrt{n+1}t) + B \cos(g\sqrt{n+1}t)$$

In absorption, the initial conditions ($C_{a,n}(0) = 0, C_{b,n+1}(0) = 1$) give $A = i, B = 0$, therefore,

$$C_{a,n}(t) = -i \sin\left(\frac{\omega_R t}{2}\right), \quad C_{b,n+1}(t) = \cos\left(\frac{\omega_R t}{2}\right), \quad (2.24)$$

where $\omega_R = 2g\sqrt{n+1}$ is the Rabi frequency in the quantum case. For the emission case, from the initial conditions ($C_{a,n}(0) = 1, C_{b,n+1}(0) = 0$) we have $A = 0, B = 1$, therefore,

$$C_{a,n}(t) = \cos\left(\frac{\omega_R t}{2}\right), \quad C_{b,n+1}(t) = -i \sin\left(\frac{\omega_R t}{2}\right), \quad (2.25)$$

In [Figure 2.8](#), the dependencies of the probabilities $|C_{a,n}(t)|^2$ and $|C_{b,n+1}(t)|^2$ on time for the case of photon emission are shown.

In the semiclassical expression, the Rabi frequency ω_R is equal to $\frac{|pE|}{\hbar}$, in the quantum one - $2g\sqrt{n+1}$, i.e. E corresponds to $E_1\sqrt{n+1}$, where E_1 is the field corresponding to one photon in the mode. For large n , these two expressions practically coincide. But for small n , the difference can be significant. From the quantum expression, it follows that even when there are no photons in the mode, the probabilities oscillate with frequency g . In fact, it is known that the oscillations at the Rabi frequency dampen. The discrepancy is due to the fact that the atom interacts with all the modes of space, and we only considered the interaction with one mode.

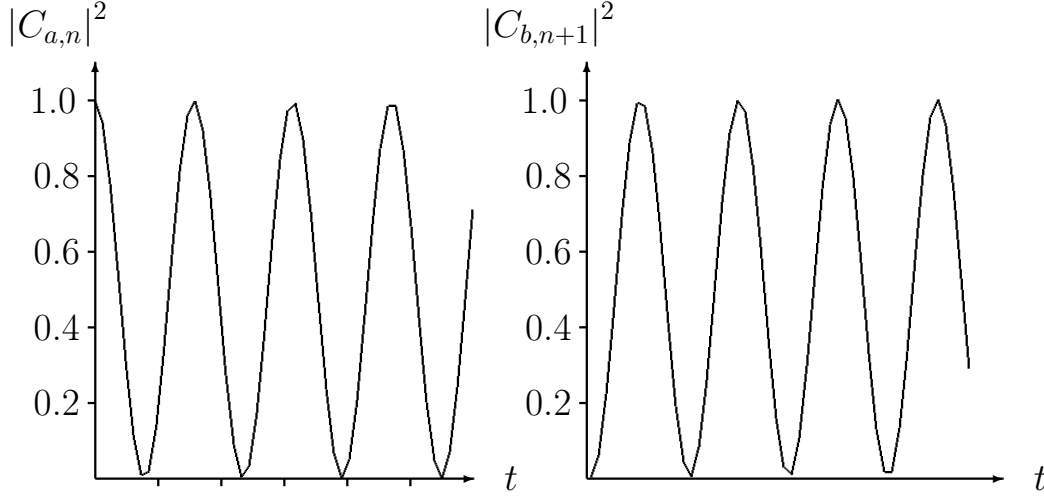


Figure 2.8: Dependencies of probabilities $|C_{a,n}|^2$ and $|C_{b,n+1}|^2$ on time for the photon emission case.

Now let's try to find the general solution for the system of equations (2.19). The relations considered can be written in the following form

$$\begin{aligned}\dot{C}_{a,n}(t) &= -i\frac{\omega_R}{2}e^{-i\delta t}C_{b,n+1}(t), \\ \dot{C}_{b,n+1}(t) &= -i\frac{\omega_R}{2}e^{i\delta t}C_{a,n}(t),\end{aligned}$$

where $\delta = \omega - \omega_{ab}$ is the detuning frequency. The problem will be solved only for the emission case, i.e., the initial conditions are given by

$$\begin{aligned}C_{a,n}(0) &= 1, \\ C_{b,n+1}(0) &= 0.\end{aligned}$$

In this case, we are primarily interested in the behavior of the probabilities $|C_{b,n+1}(t)|^2$. So, denoting $x(t) = C_{b,n+1}(t)$, we get

$$\begin{aligned}\ddot{x} &= -i\frac{\omega_R}{2}\frac{d(e^{i\delta t}C_{a,n}(t))}{dt} = \\ &= i\delta\left(-i\frac{\omega_R}{2}e^{i\delta t}C_{a,n}(t)\right) - i\frac{\omega_R}{2}e^{i\delta t}\frac{dC_{a,n}(t)}{dt} = \\ &= i\delta\dot{x} - \frac{\omega_R^2}{4}e^{-i\delta t}e^{i\delta t}x = \\ &= i\delta\dot{x} - \frac{\omega_R^2}{4}x,\end{aligned}$$

i.e.

$$\ddot{x} - i\delta\dot{x} + \frac{\omega_R^2}{4}x = 0. \quad (2.26)$$

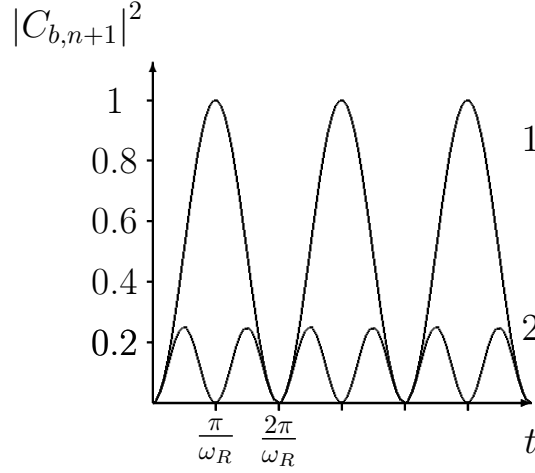


Figure 2.9: Graph 1: Rabi oscillations with zero detuning $\delta = 0$. Oscillation period $\frac{2\pi}{\omega_R}$, $\Omega_R = \omega_R$, maximum probability amplitude $|C_{b,n+1}|^2 = 1$. Graph 2: Rabi oscillations with non-zero detuning $\delta = \sqrt{3}\omega_R$. Oscillation period $\frac{\pi}{\omega_R}$, $\Omega_R = 2 \cdot \omega_R$, maximum probability amplitude $|C_{b,n+1}|^2 = \frac{\omega_R^2}{\Omega_R^2} = \frac{1}{4}$.

Equation (2.26) should be solved under the following initial conditions:

$$\begin{aligned} x|_{t=0} &= C_{b,n+1}(0) = 0, \\ \dot{x}|_{t=0} &= -i\frac{\omega_R}{2}C_{a,n}(0) = -i\frac{\omega_R}{2}. \end{aligned}$$

Result:

$$x(t) = -ie^{\frac{i\delta t}{2}} \frac{\omega_R}{\sqrt{\omega_R^2 + \delta^2}} \sin\left(\frac{\sqrt{\omega_R^2 + \delta^2}}{2}t\right),$$

denoting $\Omega_R = \sqrt{\omega_R^2 + \delta^2}$ - the generalized Rabi frequency - for the desired probability $|C_{b,n+1}|^2$, we get, see Figure 2.9

$$|C_{b,n+1}|^2 = \frac{\omega_R^2}{\Omega_R^2} \sin^2 \frac{\Omega_R}{2} t. \quad (2.27)$$

Remark 2.3.1 (Variable Stark Effect). *It should be noted that in the case of $\delta \neq 0$, the states we used in the expression 2.3 do not satisfy the conservation law of energy. Indeed, if the system was initially in the state $|a\rangle|n\rangle$ with energy $E_a + \hbar n\omega$, then after some time it can, with a non-zero probability, be registered in the state $|b\rangle|n+1\rangle$ with energy*

$$E_b + \hbar(n+1)\omega \neq E_a + \hbar n\omega.$$

That is, there appears to be a violation of the conservation law of energy. The indicated contradiction is resolved in the following way. Due to the variable Stark

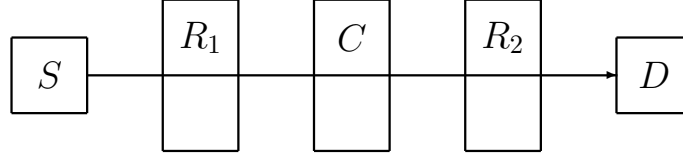


Figure 2.10: Ramsey Interferometer. The atom from source S sequentially passes through 3 zones R_1 , C , and R_2 , and at the output D , the state of the atom is determined after passing through the 3 zones.



Figure 2.11: Three-level atom used in the experiment.

effect (known in English literature also as the AC-Stark effect or Autler-Townes effect)[27], the energy levels of the atom shift, which ensures the conservation law of energy.

For example, at small detunings $\delta \approx 0$, this shift is [27]

$$\Delta E \approx \pm \frac{\hbar \omega_R}{2}.$$

2.4 Ramsey Interferometer. Quantum Non-Demolition Measurements

Observation of a photon typically involves its destruction, for example when using a photodetector, the energy of the photon is converted into an electrical signal, causing the photon itself to disappear.

An intriguing experiment [20] using Rabi oscillations demonstrated the possibility of observing a photon without destroying it. A Ramsey interferometer was used in this experiment (see [Figure 2.10](#)).

The operation scheme is as follows. Atoms with level structures depicted in [Figure 2.11](#) are used. The transition frequency ω_{ab} between levels a and b matches the frequency of the radiation in zones R_1 and R_2 of the interferometer (see [Figure 2.10](#)). Zone C contains the electromagnetic field under investigation, for which we want to determine whether it contains one photon or not, i.e., whether it is in state $|1\rangle$ or in vacuum state $|0\rangle$. The field under investigation in zone C has a frequency that matches the transition frequency ω_{ac} between levels a and c .

The size of zones $R_{1,2}$ is chosen such that the interaction time t_R of the atom with the electromagnetic field in them corresponds to the following relation

$$\omega_R^{(R)} t_R = \frac{\pi}{2},$$

where $\omega_R^{(R)}$ is the Rabi frequency corresponding to the transition frequency ω_{ab} . The interaction time t_C in zone C is determined by the expression

$$\omega_R^{(C)} t_C = 2\pi,$$

where $\omega_R^{(C)}$ is the Rabi frequency corresponding to the transition frequency ω_{ac} .

For interaction in zones $R_{1,2}$, the following relations hold (2.24), 2.25:

$$\begin{aligned} |a\rangle &\rightarrow \frac{1}{\sqrt{2}} (|a\rangle - i|b\rangle), \\ |b\rangle &\rightarrow \frac{1}{\sqrt{2}} (-i|a\rangle + |b\rangle). \end{aligned} \quad (2.28)$$

The expression (2.28) can be rewritten in matrix form $|\psi\rangle \rightarrow R|\psi\rangle$, where

$$R = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & -i \\ -i & 1 \end{pmatrix}.$$

Zone C does not affect state $|b\rangle$; for state $|a\rangle$ in the case of vacuum state $|0\rangle$ (absence of a photon),

$$|a\rangle \rightarrow |a\rangle,$$

i.e., in the absence of a photon, the state of the atom does not change. Indeed, if we consider a two-level system formed by states $|a\rangle$ and $|c\rangle$, then $|a\rangle|0\rangle$ will be the minimally possible energy state, from which the system (two-level atom and electromagnetic field) cannot transition to any other state. In this case, for C we have

$$C_0 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}.$$

In the presence of a photon, from (2.24)

$$|a\rangle \rightarrow -|a\rangle,$$

and in this case C is represented as

$$C_1 = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}.$$

Thus, if initially the atom emitted by source S is in state $|a\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$, then in the absence of a photon in zone C , we will detect the atom in state

$$\begin{aligned} R_2 C_0 R_1 |a\rangle &= \\ &= \frac{1}{2} \begin{pmatrix} 1 & -i \\ -i & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & -i \\ -i & 1 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \\ &= -i \begin{pmatrix} 0 \\ 1 \end{pmatrix} = -i |b\rangle, \end{aligned}$$

i.e., the atom will be observed in the unexcited state $|b\rangle$.

In the presence of one photon in zone C , we have

$$\begin{aligned} R_2 C_1 R_1 |a\rangle &= \\ &= \frac{1}{2} \begin{pmatrix} 1 & -i \\ -i & 1 \end{pmatrix} \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & -i \\ -i & 1 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \\ &= - \begin{pmatrix} 1 \\ 0 \end{pmatrix} = -|a\rangle, \end{aligned}$$

i.e., the atom will be observed in the excited state $|a\rangle$. Thus, it is possible to determine the presence of a photon in zone C without destructively affecting it.

2.5 Interaction of an Atom with a Multimode Field. Induced and Spontaneous Transitions

Previously, we established that the probability of an atom transitioning from a lower to an upper state when interacting with a single mode of the field in a state with n photons is given by (2.22)

$$|C_{a,n}(t)|^2 = 4g^2(n+1) \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2}$$

We can express this probability as follows

$$P_a = |C_{a,n}(t)|^2 = P_a^{\text{ind}} + P_a^{\text{spont}}, \quad (2.29)$$

where

$$\begin{aligned} P_a^{\text{ind}} &= 4g^2 n \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2}, \\ P_a^{\text{spont}} &= 4g^2 \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2} \end{aligned} \quad (2.30)$$

are the probabilities of induced (P_a^{ind}) and spontaneous (P_a^{spont}) transitions.

2.5.1 Induced Transitions

We will now focus on the probability of the induced transition P_a^{ind} . To obtain the transition probability considering interactions with all modes, the probabilities (2.30) corresponding to each mode need to be summed.

In the case of an atom interacting with a free space field (field expansion in plane waves), the coupling constant is given by [61]:

$$g = -\frac{(\vec{p}\vec{e})}{\hbar} \sqrt{\frac{\hbar\omega}{2\varepsilon_0 V}} \quad (2.31)$$

where \vec{p} is the matrix element of the dipole moment operator, and \vec{e} is the field's polarization vector.

If in expression (2.30) time t is not very small (but significantly less than the characteristic time of probability changes), the dependence of $|C_{a,n}(t)|^2$ on the frequency will have a sharp peak at $\omega = \omega_{ab}$, indicating energy conservation in the elementary act (the energy of the absorbed photon equals the change in the atom's energy). Assume that the incident field in the direction \vec{k} within the solid angle $d\Omega$ has a spectrum that changes little near frequency ω_{ab} .

The number of modes corresponding to the frequency interval $d\omega$ near ω_{ab} and the solid angle $d\Omega$ around the direction \vec{k} is (1.47),

$$dN = 2 \left(\frac{L}{2\pi c} \right)^3 \omega^2 d\omega d\Omega \quad (2.32)$$

The probability of photon absorption during atom interaction with a single mode is given by formula (2.30). The total probability can be obtained by summing over all modes. Assuming the mode spectrum is quasi-continuous (see (1.44)), summation can be replaced by integration

$$W_{b \rightarrow a} = \int_{\Omega} 8 |g|^2 n(\vec{k}) \left(\frac{L}{2\pi c} \right)^3 \omega_{ab}^2 d\Omega \int_{-\infty}^{+\infty} \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2} d\omega. \quad (2.33)$$

It's noted that $\frac{\sin^2 x}{x^2}$ has a narrow peak near $x = 0$, while in expression (2.33) we assume that t is small compared to the characteristic time change of C , but sufficiently large to exhibit the filter properties of the integral. All slowly varying terms are taken at $\omega = \omega_{ab}$ and factored out of the integral. It is known that $\int_{-\infty}^{+\infty} \frac{\sin^2 x}{x^2} dx = \pi$. Thus, we easily obtain

$$\int_{-\infty}^{+\infty} \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2} d\omega = \frac{\pi t}{2}.$$

Therefore, we have

$$W_{b \rightarrow a} = t \frac{\omega^2 2\pi\omega}{\hbar\epsilon_0 (2\pi c)^3} \int_{\Omega} n(\vec{k}) |(\vec{p}\vec{e})|^2 d\Omega. \quad (2.34)$$

Note that the quantization volume does not enter into the final expression, and the number of photons n depends on the direction, i.e., the solid angle Ω . Formula (2.34) shows that the transition probability is proportional to time. This allows the introduction of the transition rate, i.e., the probability of transition per unit time

$$w_{b \rightarrow a} = \frac{W_{b \rightarrow a}}{t} = \frac{2\pi\omega^3}{\hbar\epsilon_0 (2\pi c)^3} \int_{\Omega} n(\vec{k}) |(\vec{p}\vec{e})|^2 d\Omega. \quad (2.35)$$

The transition rate or photon absorption rate can be expressed [61] in terms of the energy flux (photon flux) traveling in the direction \vec{k} within the solid angle $d\Omega$. The energy flux is defined as the energy transferred per unit area dS per unit time dt , i.e.

$$dI = \frac{dH}{dSdt}.$$

From ??, we have that the number of photons passing through area dS per unit time dt is equal to the number contained in the cylinder shown in ??, i.e., the transferred energy can be written as

$$dH = n\hbar\omega dN,$$

where dN is the number of modes corresponding to the interval $d\omega d\Omega$ in the considered volume $c \cdot dSdt$ (1.47):

$$dN = 2 \left(\frac{1}{2\pi c} \right)^3 \omega^2 c \cdot dSdt d\omega d\Omega,$$

from which we obtain

$$dI = I(\omega, \vec{k}) d\omega d\Omega = \frac{2n\hbar\omega c}{(2\pi c)^3} \omega^2 d\omega d\Omega, \quad (2.36)$$

where $I(\omega, \vec{k})$ is the photon energy flux in the direction \vec{k} per unit frequency interval and unit solid angle. Thus, from (2.36), we obtain

$$I(\omega, \vec{k}) = \frac{2n\hbar\omega^3 c}{(2\pi c)^3}$$

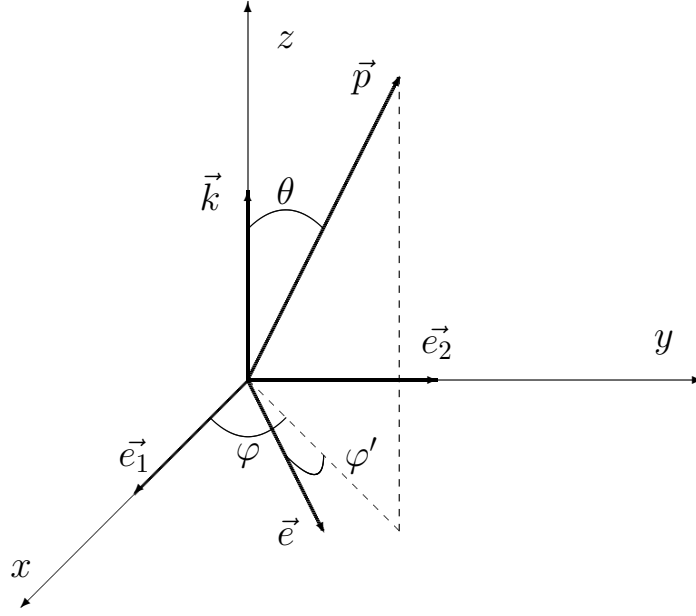


Figure 2.12: Coordinate system in which it is convenient to average the multimode field over all directions and polarizations. The wave vector of the electromagnetic wave \vec{k} is directed along the z -axis, while the position of the dipole moment matrix element \vec{p} is defined by the angles θ and ϕ . Polarization vectors \vec{e}_1 and \vec{e}_2 are directed along the x and y axes. The angle ϕ' determines the direction of polarization of the electromagnetic wave.

Substituting this into expression (2.35), we obtain:

$$w_{ab} = \frac{\pi}{\hbar^2} \sqrt{\frac{\mu_o}{\varepsilon_0}} \int_{\Omega} I(\omega, \vec{k}) |(\vec{p}\vec{e})|^2 d\Omega. \quad (2.37)$$

The relation $\mu_0\varepsilon_0 = \frac{1}{c^2}$ is used here.

To find the total absorption rate, integrate (2.37) over all wave propagation directions. Additionally, the incident radiation is usually unpolarized. To account for this, all polarization directions should be averaged. We use the coordinate system shown in Figure 2.12. Direction \vec{k} is chosen as the polar axis z . Vectors \vec{e}_1 and \vec{e}_2 are directed along x and y , respectively. Angles φ and θ define the direction of \vec{p} . The angle φ' defines the polarization direction of the incident wave. From the figure, it is seen that

$$|(\vec{p}\vec{e})|^2 = |p|^2 \sin^2 \theta \cos^2 \varphi'.$$

Averaging over all polarizations gives:

$$\frac{|p|^2}{2\pi} \int_0^{2\pi} \cos^2 \varphi' d\varphi' = \frac{|p|^2}{2}. \quad (2.38)$$

Summation over all wave arrival directions leads to expression

$$w_{ab} = \frac{\pi}{2\hbar^2} \sqrt{\frac{\mu_o}{\varepsilon_0}} |p|^2 \int_{\Omega} I(\omega, \vec{k}) \sin^3 \theta d\theta d\varphi. \quad (2.39)$$

Assuming radiation comes from all directions and is isotropic, obtain:

$$w_{ab} = \frac{\pi}{2\hbar^2} \sqrt{\frac{\mu_o}{\varepsilon_0}} |p|^2 I(\omega) \int_0^{2\pi} d\varphi \int_0^\pi \sin^3 \theta d\theta = \frac{\pi}{\hbar^2} \sqrt{\frac{\mu_o}{\varepsilon_0}} \frac{|p|^2}{3} I_0, \quad (2.40)$$

where $I_0 = 4\pi I(\omega)$ is the total energy flux incident on the atom. In deriving (2.40), the following relation was used:

$$\int_0^\pi \sin^3 \theta d\theta = \frac{4}{3}$$

If, however, the radiation comes from a small region $\Delta\Omega$ of directions near angles θ_0, φ_0 , then we can write:

$$\begin{aligned} w_{ab} &= \frac{\pi}{2\hbar^2} \sqrt{\frac{\mu_o}{\varepsilon_0}} |p|^2 \sin^2 \theta_0 \int_{\Delta\Omega} I(\theta, \varphi) d\Omega = \\ &= \frac{\pi}{2\hbar^2} \sqrt{\frac{\mu_o}{\varepsilon_0}} |p|^2 I_0 \sin^2 \theta_0, \end{aligned} \quad (2.41)$$

where $I_0 = \int_{\Delta\Omega} I(\theta, \varphi) d\Omega$ is the total flux illuminating the atom.

2.5.2 Spontaneous Transitions

Now, consider another problem. Determine the probability of photon emission or the transition of an excited atom to a lower state.

For the probability of atom emission of a photon into one mode, we have the expression (2.23):

$$|C_{b,n+1}(t)|^2 = 4g^2(n+1) \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2},$$

which, like (2.29), can be considered as a sum of two terms

$$P_b = |C_{b,n+1}(t)|^2 = P_b^{\text{ind}} + P_b^{\text{spont}}$$

where the first term

$$P_b^{\text{ind}} = 4g^2 n \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2}$$

corresponds to induced emission, and the second

$$P_b^{\text{spont}} = 4g^2 \frac{\sin^2((\omega - \omega_{ab})t/2)}{(\omega - \omega_{ab})^2}$$

to spontaneous emission. The term corresponding to induced emission is considered in the same way as for photon absorption. The result will obviously be the same. The corresponding formulas will match (2.40), (2.41). It follows that the probabilities of induced absorption and emission processes are equal. The probability of spontaneous emission per unit time will obviously be equal to

$$w_{\text{spont}} = \frac{2\pi\omega^3}{\hbar\varepsilon_0(2\pi c)^3} \int_{\Omega} |(\vec{p}\vec{e})|^2 d\Omega \quad (2.42)$$

To obtain the total probability, (2.42) must be integrated over all directions, since spontaneous emission can occur into any mode. For polarizations, use the average value (2.38). Thus, the entire procedure reduces to calculating the integral

$$w_{\text{spont}} = \frac{|p|^2 \omega^3}{\varepsilon_0 \hbar (2\pi)^2 c^3} \int_0^{2\pi} d\varphi \int_0^\pi \sin^3 \theta d\theta = \frac{4\pi}{3} \frac{|p|^2 \omega^3}{\hbar \varepsilon_0 (2\pi)^2 c^3} \quad (2.43)$$

Ultimately, we get

$$w_{\text{spont}} = \frac{|p|^2 \omega^3}{3\pi c^2 \hbar} \sqrt{\frac{\mu_0}{\varepsilon_0}} \quad (2.44)$$

The relation

$$\frac{1}{c} = \sqrt{\mu_0 \varepsilon_0}$$

is used here.

Formula (2.44) indicates a strong dependence of the spontaneous transition probability on frequency, as ω^3 . Thus, based on quantum electrodynamics equations, we directly, without invoking external considerations, obtained expressions for the probabilities of induced and spontaneous atomic transitions per unit time. A shortcoming of this consideration is the use of perturbation theory, necessitating limitations to small times. For instance, when solving the problem of the lifetime of an excited atom, one cannot restrict to short times alone. In this case, another approximation known as the Weisskopf-Wigner approximation is applied [61]. Interestingly, the results obtained in these two cases do not contradict each other.

2.6 Spontaneous Emission. Weisskopf-Wigner Approximation.

Above, we derived a formula for the rate of spontaneous emission, valid for sufficiently small times. At the same time, when solving problems such as the lifetime of an excited atom, one cannot limit oneself to small times. In this case, a different method is applied, called the Weisskopf-Wigner approximation.

Let's consider this method. We found that an atom transitions from an excited state to the ground state due to interaction with all modes of space, even if they are unexcited and in a vacuum state. The interaction Hamiltonian in this case has the form

$$\hat{V}_I = \hbar \sum_k g_k \hat{\sigma}^\dagger \hat{a}_k e^{i(\omega_{ab} - \omega_k)t} + \text{h. c.}, \quad (2.45)$$

where g_k is the coupling constant with the k -th mode. The Hamiltonian (2.45) differs from the one we previously used (2.14) by summing over all quantization space modes.

Suppose at the initial moment $t = 0$ the atom is excited, and there are no photons in the field. Then the initial state is

$$|\psi(0)\rangle = |a, \{0\}\rangle,$$

meaning the atom is excited, and all modes are in the vacuum state. At later times, due to the atom-field interaction, the state of the system will be

$$|\psi(t)\rangle = \sum_k C_{bk}(t) |b, 1_k\rangle + C_a(t) |a, \{0\}\rangle, \quad (2.46)$$

where $|b, 1_k\rangle$ is the state when the atom is in the ground state, and there is one photon excited in the k -th mode.

The Schrödinger equation in the interaction representation is

$$\frac{d}{dt} |\psi(t)\rangle = -\frac{i}{\hbar} \hat{V}_I |\psi(t)\rangle. \quad (2.47)$$

Substituting (2.46) and multiplying (2.47) sequentially by $\langle a, \{0\}|$ and $\langle b, 1_k|$, we obtain a system of equations for the probability amplitudes $C_a(t)$ and $C_{bk}(t)$.

$$\begin{aligned} \dot{C}_a(t) &= -i \sum_k g_k e^{i(\omega_{ab} - \omega_k)t} C_{bk}(t), \\ \dot{C}_{bk}(t) &= -i g_k e^{-i(\omega_{ab} - \omega_k)t} C_a(t). \end{aligned} \quad (2.48)$$

Integrating the second equation of (2.48) over time from 0 to t .

$$C_{bk}(t) = -i g_k \int_0^t e^{-i(\omega_{ab} - \omega_k)t'} C_a(t') dt'. \quad (2.49)$$

Substituting (2.49) into the first equation of the system (2.48), we get the following integro-differential equation

$$\dot{C}_a(t) = - \sum_k g_k^2 \int_0^t e^{i(\omega_{ab}-\omega_k)(t-t')} C_a(t') dt'. \quad (2.50)$$

Let's make some simplifications (Weisskopf-Wigner approximation). We will consider the distribution of modes as quasi-continuous and replace the summation over k with integration, using the relation (1.46):

$$dN = 2 \left(\frac{L}{2\pi} \right)^3 k^2 dk d\Omega = 2 \frac{V}{(2\pi)^3} k^2 dk d\Omega$$

Then we obtain (1.44)

$$\begin{aligned} & \sum_k g_k^2 \int_0^t e^{i(\omega_{ab}-\omega_k)(t-t')} C_a(t') dt' = \\ & = 2 \frac{V}{(2\pi)^3} \int_{4\pi} d\Omega \int_0^\infty g_k^2 k^2 dk \int_0^t dt' e^{i(\omega_{ab}-\omega_k)(t-t')} C_a(t'). \end{aligned} \quad (2.51)$$

In (2.51)

$$g_k^2 = \frac{\omega_k |p|^2 \sin^2 \theta}{4\hbar \varepsilon_0 V}, \quad (2.52)$$

where θ is the angle between the directions \vec{k} and \vec{p} . The expression (2.52) is obtained considering the averaging over polarizations (2.38).

Transitioning in (2.51) from k to ω using the relations

$$k = \frac{\omega_k}{c}, \quad k^2 dk = \frac{\omega_k^2 d\omega_k}{c^3}$$

and denoting for convenience $\omega_k = \omega$, we obtain

$$\dot{C}_a(t) = -2 \frac{V}{(2\pi c)^3} \int_{4\pi} d\Omega \int_0^\infty g^2(\omega) \omega^2 d\omega \int_0^t dt' e^{i(\omega_{ab}-\omega)(t-t')} C_a(t'). \quad (2.53)$$

For the approximate calculation of the integral in (2.53), we further apply a series of simplifying assumptions. First, integrate (2.53) over frequency (i.e., change the order of integration, assuming it is possible). From the structure (2.53), it is evident that the main contribution to the time integral is given by frequency regions $\omega \approx \omega_{ab}$. For this reason, we can assume

$$\omega^2 g^2(\omega) \approx \omega_{ab}^2 g^2(\omega_{ab}).$$

In this approximation, the frequency integral will look like (see the integral representation of the delta function (18.3)):

$$\begin{aligned} \int_0^\infty d\omega e^{i(\omega_{ab}-\omega)(t-t')} &= |\nu = \omega - \omega_{ab}| = \int_{-\omega_{ab}}^\infty d\nu e^{-i\nu(t-t')} \approx \\ &\approx \int_{-\infty}^\infty d\nu e^{-i\nu(t-t')} = \int_{-\infty}^\infty d\nu e^{i\nu(t'-t)} = 2\pi\delta(t'-t). \end{aligned} \quad (2.54)$$

Substituting (2.54) into equation (2.53) and integrating over time, we obtain

$$\dot{C}_a(t) = -2 \frac{V}{(2\pi c)^3} \int_{4\pi} d\Omega g^2(\omega_{ab}) \omega_{ab}^2 2\pi C_a(t) = -\frac{\Gamma}{2} C_a(t). \quad (2.55)$$

Substituting here the expression for g_k^2 (2.52) and performing integration over angles (over $d\Omega$), we find the expression for the damping coefficient Γ :

$$\Gamma = \frac{\omega_{ab}^3 |p|^2}{3\pi c^2 \hbar} \sqrt{\frac{\mu_0}{\varepsilon_0}}. \quad (2.56)$$

Note that the expression (2.56) coincides with the one obtained earlier by another method (2.44).

From (2.55), it follows that Γ characterizes the rate of change of the probability $|C_a|^2$. Indeed, from (2.55) we have:

$$\begin{aligned} \dot{C}_a C_a^* &= -\frac{\Gamma}{2} C_a C_a^*, \\ \dot{C}_a^* C_a &= -\frac{\Gamma}{2} C_a^* C_a, \end{aligned}$$

from which we have

$$\frac{dC_a C_a^*}{dt} = -\Gamma (C_a C_a^*),$$

or otherwise

$$\frac{d|C_a|^2}{dt} = -\Gamma |C_a|^2.$$

The solution has the form (see Figure 2.13)

$$|C_a|^2 = e^{-\Gamma t},$$

if the initial value $|C_a|^2 \Big|_{t=0} = 1$, i.e., the atom was initially excited.

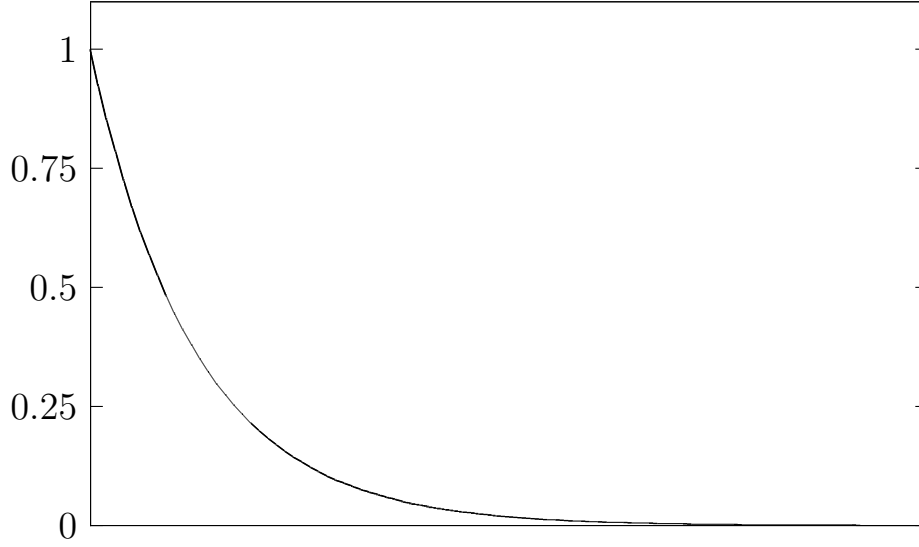


Figure 2.13: Dependence of the probability of finding an atom in an excited state $|C_a|^2$ in the Weisskopf-Wigner approximation

Remark 2.6.1 (Weisskopf-Wigner Approximation and Rabi Oscillations). *If we compare Figure 2.8 and Figure 2.13, we can see that these two graphs are very different, despite describing seemingly the same model system. This raises the question of when we can apply the single-mode approximation, and hence obtain Rabi oscillations (see Figure 2.8), and in which cases it is necessary to use multi-mode approximations like Weisskopf-Wigner (see Figure 2.13).*

Formally, we can consider that interaction with each mode of the electromagnetic field contributes to the resulting probabilities; however, according to (2.23), the change in probability (for small t)

$$\Delta |C_{a,n}(t)|^2 = -\Delta |C_{b,n+1}(t)|^2 \sim (n+1).$$

Thus, those modes in which the number of photons is large contribute more than those where it is small. Consequently, if there is a mode with a large number of photons $n \gg 1$, the influence of vacuum modes with a photon number $n = 0$ can be neglected in this case. Otherwise, especially if all modes are equally significant, it is necessary to consider all modes and apply approximations like Weisskopf-Wigner.

2.7 Relaxation of a Dynamic System. Method of Density Matrix

In quantum electronics and quantum optics, we deal with open systems. Such systems must be considered using the density matrix. A small dynamic system

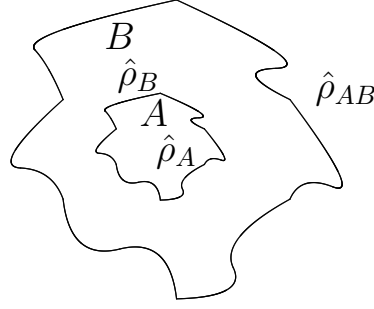


Figure 2.14: Interaction of dynamic system A with reservoir B

(atom, molecules, resonator mode, etc.) is considered, weakly connected with a large system called a reservoir (dissipative system, thermostat), consisting of a large number of atoms, molecules, or oscillators in equilibrium at a certain temperature. Thus, in general, we consider a dynamic system (A) interacting with a reservoir (B) (see [Figure 2.14](#)). The density operator of the combined system (AB) is denoted by $\hat{\rho}_{AB}$. The density operator collapsed over the reservoir (thermostat) variables describes the behavior of the dynamic system taking into account the averaged interaction with the thermostat. We have:

$$\hat{\rho}_A = Sp_B \hat{\rho}_{AB} = \sum_{(B)} \langle B | \hat{\rho}_{AB} | B \rangle. \quad (2.57)$$

At the initial moment of time t_0 , the dynamic system and the reservoir are considered independent (uncorrelated). Then

$$\hat{\rho}_{AB}(t_0) = \hat{\rho}_A(t_0) \otimes \hat{\rho}_B(t_0) \quad (2.58)$$

where $Sp_B \hat{\rho}_B = 1$, \otimes is the sign of the external or tensor product.

The equation of motion of the density operator in the interaction representation is

$$\frac{d}{dt} \hat{\rho}_{AB} = -\frac{i}{\hbar} [\hat{V}_{AB}, \hat{\rho}_{AB}],$$

where \hat{V}_{AB} is the interaction Hamiltonian between systems A and B . This equation can be solved using the following iterative process.

$$\hat{\rho}_{AB}(t) = \hat{\rho}_{AB}^{(0)}(t) + \hat{\rho}_{AB}^{(1)}(t) + \hat{\rho}_{AB}^{(2)}(t) + \dots, \quad (2.59)$$

where the term $\hat{\rho}_{AB}^{(0)}(t)$ is found assuming the absence of interaction, i.e.

$$\hat{\rho}_{AB}^{(0)}(t) = \hat{\rho}_{AB}(t_0) = \hat{\rho}_A(t_0) \otimes \hat{\rho}_B(t_0).$$

The subsequent terms of the iterative relation (2.59) are defined by the following relation

$$\begin{aligned}
\hat{\rho}_{AB}^{(s)}(t) &= -\frac{i}{\hbar} \int_{t_0}^t dt_s \left[\hat{V}_{AB}(t_s), \hat{\rho}_{AB}^{(s-1)}(t_s) \right] = \\
&= \left(-\frac{i}{\hbar} \right)^s \int_{t_0}^t dt_s \int_{t_0}^{t_s} dt_{s-1} \dots \\
&\dots \int_{t_0}^{t_2} dt_1 \left[\hat{V}_{AB}(t_s), \left[\hat{V}_{AB}(t_{s-1}), \dots \left[\hat{V}_{AB}(t_1), \hat{\rho}_{AB}^{(0)}(t_0) \right] \dots \right] \right] = \\
&= \left(-\frac{i}{\hbar} \right)^s \int_{t_0}^t dt_s \int_{t_0}^{t_s} dt_{s-1} \dots \\
&\dots \int_{t_0}^{t_2} dt_1 \left[\hat{V}_{AB}(t_s), \left[\hat{V}_{AB}(t_{s-1}), \dots \left[\hat{V}_{AB}(t_1), \hat{\rho}_A(t_0) \otimes \hat{\rho}_B(t_0) \right] \dots \right] \right]. \quad (2.60)
\end{aligned}$$

Substituting (2.60) into (2.59) and averaging over the thermostat variables, we obtain

$$\begin{aligned}
\hat{\rho}_A(t) &= \hat{\rho}_A(t_0) + \\
&+ Sp_B \sum_{(s)} \left(-\frac{i}{\hbar} \right)^s \int_{t_0}^t dt_s \int_{t_0}^{t_s} dt_{s-1} \dots \\
&\dots \int_{t_0}^{t_2} dt_1 \left[\hat{V}_{AB}(t_s), \left[\hat{V}_{AB}(t_{s-1}), \dots \left[\hat{V}_{AB}(t_1), \hat{\rho}_A(t_0) \otimes \hat{\rho}_B(t_0) \right] \dots \right] \right]. \quad (2.61)
\end{aligned}$$

Usually, only the first few terms of the expansion are considered. For example, the first two terms of the expansion are

$$\begin{aligned}
\hat{\rho}_A(t) &= \hat{\rho}_A(t_0) - \frac{i}{\hbar} Sp_B \left\{ \int_{t_0}^t \left[\hat{V}_{AB}(t_1), \hat{\rho}_A(t_0) \otimes \hat{\rho}_B(t_0) \right] dt_1 \right\} - \\
&- \frac{1}{\hbar^2} Sp_B \left\{ \int_{t_0}^t dt_2 \int_{t_0}^{t_2} dt_1 \left[\hat{V}_{AB}(t_2), \left[\hat{V}_{AB}(t_1), \hat{\rho}_A(t_0) \otimes \hat{\rho}_B(t_0) \right] \right] \right\}.
\end{aligned}$$



Figure 2.15: Interaction of the electromagnetic field of the resonator F with a reservoir of atoms R at temperature T

2.8 Interaction of Electromagnetic Field of the Resonator with a Reservoir of Atoms at Temperature T

We apply the general approach outlined above to a system consisting of a harmonic oscillator (electromagnetic field mode) interacting with a reservoir in the form of a beam of two-level atoms in equilibrium at temperature T . As we will see later, the specific model of the reservoir does not matter for the final result.¹ The considered model is depicted in [Figure 2.15](#).

The initial atomic density matrix corresponds to the Boltzmann distribution at temperature T :

$$\hat{\rho}_{at} = \begin{pmatrix} \rho_{aa} & 0 \\ 0 & \rho_{bb} \end{pmatrix} = z^{-1} \begin{pmatrix} e^{-\frac{\hbar\omega_a}{k_B T}} & 0 \\ 0 & e^{-\frac{\hbar\omega_b}{k_B T}} \end{pmatrix}. \quad (2.62)$$

Here $\hbar\omega_a = E_a$ is the energy of the upper level; $\hbar\omega_b = E_b$ is the energy of the lower level; $z = e^{-\frac{\hbar\omega_a}{k_B T}} + e^{-\frac{\hbar\omega_b}{k_B T}}$ is the partition function; T is the temperature of the reservoir. We take the interaction Hamiltonian in the form of (2.14). For simplicity, assume that $\omega_{ab} = \omega$, i.e., the transition frequency equals the mode frequency. Then

$$\hat{V} = \hbar g (\hat{\sigma} \hat{a}^\dagger + \hat{\sigma}^\dagger \hat{a}) = \hbar g \begin{pmatrix} 0 & \hat{a} \\ \hat{a}^\dagger & 0 \end{pmatrix} \quad (2.63)$$

since $\hat{\sigma} = \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$; $\hat{\sigma}^\dagger = \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix}$; g is the interaction constant.

¹If the interaction is weak, i.e., the binding energy is small compared to the energy of the dynamic system.

We retain terms up to the second order in the iterative series (2.61). In our case, after integration, we get a simple expression

$$\begin{aligned} \hat{\rho}_f(t) &= \hat{\rho}_f(t_0 + \tau) = \\ &= \hat{\rho}_f(t_0) + Sp_{at} \left\{ -\frac{i}{\hbar} \tau [\hat{V}, \hat{\rho}_{at,f}] - \frac{1}{2} \left(\frac{\tau}{\hbar} \right)^2 [\hat{V}, [\hat{V}, \hat{\rho}_{at,f}]] \right\}, \end{aligned} \quad (2.64)$$

where $\hat{\rho}_{at,f}$ is the statistical operator of the atom-field system; $\hat{\rho}_f$ is the statistical operator of the field, folded over atomic variables.

In deriving (2.64), it was assumed that the atom beam interacts with the field for a time τ (time of flight through the resonator). From the initial conditions (at time t_0), we have:

$$\hat{\rho}_{at,f} = \hat{\rho}_f \begin{pmatrix} \rho_{aa} & 0 \\ 0 & \rho_{bb} \end{pmatrix} = \begin{pmatrix} \rho_{aa} \hat{\rho}_f & 0 \\ 0 & \rho_{bb} \hat{\rho}_f \end{pmatrix} \quad (2.65)$$

The first-order term in (2.64) gives a matrix with zero diagonal elements in the reservoir variables. The trace of the matrix in these variables is obviously zero

$$\begin{aligned} Sp_{at} [\hat{V}, \hat{\rho}_{at,f}] &= \hbar g Sp_{at} \left\{ \begin{pmatrix} 0 & \hat{a} \rho_{bb} \hat{\rho}_f \\ \hat{a}^\dagger \rho_{aa} \hat{\rho}_f & 0 \end{pmatrix} - \right. \\ &\quad \left. - \begin{pmatrix} 0 & \rho_{aa} \hat{\rho}_f \hat{a} \\ \rho_{bb} \hat{\rho}_f \hat{a}^\dagger & 0 \end{pmatrix} \right\} = 0. \end{aligned} \quad (2.66)$$

The matrix included in the second-order term has non-zero diagonal elements, its trace in the reservoir variables is non-zero and can be calculated:

$$\begin{aligned} Sp_{at} [\hat{V}, [\hat{V}, \hat{\rho}_{at,f}]] &= \hbar^2 g^2 \cdot \\ \cdot Sp_{at} \begin{pmatrix} \hat{a} \hat{a}^\dagger \rho_{aa} \hat{\rho}_f - \hat{a} \rho_{bb} \hat{\rho}_f \hat{a}^\dagger & 0 \\ 0 & \hat{a}^\dagger \hat{a} \rho_{bb} \hat{\rho}_f - \hat{a}^\dagger \rho_{aa} \hat{\rho}_f \hat{a} \end{pmatrix} &+ \text{c.c.} \end{aligned} \quad (2.67)$$

From here, using equations (2.64), (2.66), (2.67), we obtain the final equation for the statistical operator of the field $\hat{\rho}$:

$$\begin{aligned} \hat{\rho}_f(t_0 + \tau) &= \hat{\rho}_f(t_0) - \frac{1}{2} g^2 \tau^2 \{ (\hat{a} \hat{a}^\dagger \hat{\rho}_f - \hat{a}^\dagger \hat{\rho}_f \hat{a}) \rho_{aa} + \\ &\quad + (\hat{a}^\dagger \hat{a} \hat{\rho}_f - \hat{a} \hat{\rho}_f \hat{a}^\dagger) \rho_{bb} \} + \text{c.c.} \end{aligned} \quad (2.68)$$

Thus, the statistical operator of the mode due to interaction with one atom changes by an amount $\delta \hat{\rho}_f = \hat{\rho}_f(t_0 + \tau) - \hat{\rho}_f(t_0)$, determined by (2.68). If the injection rate of atoms into the resonator is r atoms per second, during the time

Δt $r\Delta t$ atoms will interact. The change in the density operator over this time will be $\Delta\hat{\rho}_f = \delta\hat{\rho}_f r\Delta t$. We define a smoothed derivative as

$$\dot{\hat{\rho}}_f = \frac{\Delta\hat{\rho}_f}{\Delta t} = \delta\hat{\rho}_f r = r(\hat{\rho}_f(t_0 + \tau) - \hat{\rho}_f(t_0)).$$

Let us denote: $r_a = r\rho_{aa}$ as the injection rate of atoms in the upper state; $r_b = r\rho_{bb}$ as the injection rate of atoms in the lower state. This leads us to the final form of the equation for the statistical operator of the resonance mode field

$$\begin{aligned} \dot{\hat{\rho}}_f = & -\frac{1}{2}R_a(\hat{a}\hat{a}^\dagger\hat{\rho}_f - \hat{a}^\dagger\hat{\rho}_f\hat{a}) - \\ & -\frac{1}{2}R_b(\hat{a}^\dagger\hat{a}\hat{\rho}_f - \hat{a}\hat{\rho}_f\hat{a}^\dagger) + \text{c.c.} \end{aligned} \quad (2.69)$$

Here it is denoted

$$\begin{aligned} R_a &= r_a g^2 \tau^2, \\ R_b &= r_b g^2 \tau^2. \end{aligned} \quad (2.70)$$

Equation (2.69) is written in operator form. It can be written in different representations.

2.9 Equation for the Density Matrix of the Field in the Occupation Number Representation

In the occupation number representation (photon numbers), the matrix elements are defined by the equality

$$\rho_{nm} = \langle n | \hat{\rho} | m \rangle.$$

To obtain the required equation, multiply (2.69) on the left by $\langle n |$ and on the right by $|m\rangle$. Using the properties of the operators \hat{a} and \hat{a}^\dagger , we obtain a system of equations relating the matrix elements

$$\begin{aligned} \dot{\rho}_{nm} = & -\frac{1}{2}(R_a(n+1+m+1) + R_b(n+m))\rho_{nm} + \\ & + R_a\sqrt{nm}\rho_{n-1,m-1} + R_b\sqrt{(n+1)(m+1)}\rho_{n+1,m+1}. \end{aligned} \quad (2.71)$$

For diagonal elements $n = m$, therefore

$$\begin{aligned} \dot{\rho}_{nn} = & -(R_a(n+1) + R_b(n))\rho_{nn} + \\ & + R_a n \rho_{n-1,n-1} + R_b(n+1)\rho_{n+1,n+1}. \end{aligned} \quad (2.72)$$

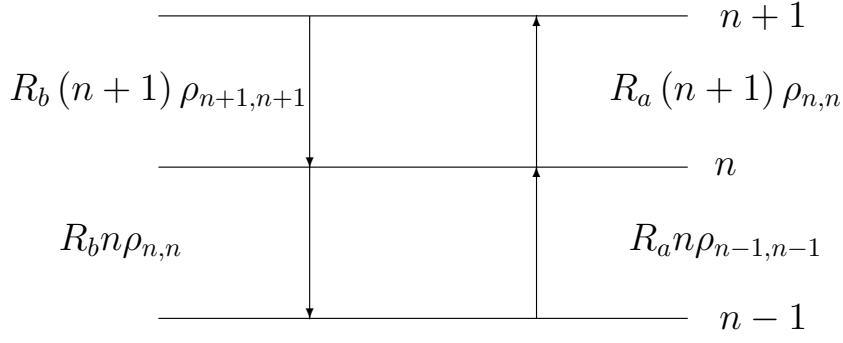


Figure 2.16: Balance of probability flows of photon numbers

This equation can be considered as the balance of probability flows of photon numbers. This is conditionally depicted in Figure 2.16. In thermal equilibrium, the flows should be equal. Hence, using the principle of detailed balance, we obtain the equalities:

$$\begin{aligned} R_a n \rho_{n-1,n-1} &= R_b n \rho_{nn}, \\ R_a (n+1) \rho_{nn} &= R_b (n+1) \rho_{n+1,n+1} \end{aligned} \quad (2.73)$$

From (2.73) we have

$$\rho_{n+1,n+1} = \frac{R_a}{R_b} \rho_{nn} = e^{-\frac{\hbar\omega}{k_B T}} \rho_{nn}, \quad (2.74)$$

because $\frac{R_a}{R_b} = \exp\left(-\frac{\hbar\omega}{k_B T}\right)$ (the reservoir is in equilibrium at temperature T). Using (2.74) successively, starting with $n = 0$, we obtain:

$$\rho_{nn} = \rho_{00} e^{-\frac{n\hbar\omega}{k_B T}} = \left(1 - e^{-\frac{\hbar\omega}{k_B T}}\right) e^{-\frac{n\hbar\omega}{k_B T}}, \quad (2.75)$$

where ρ_{00} is found from the condition $\sum_{(n)} \rho_{nn} = 1$.

The average number of photons in the mode, as expected, is determined by the Planck formula

$$\bar{n} = \langle n \rangle = \sum_{(n)} n \rho_{nn} = \frac{1}{e^{\frac{\hbar\omega}{k_B T}} - 1}. \quad (2.76)$$

From all said, it follows that over time the temperature of the radiation becomes equal to the temperature of the atomic beam (reservoir). The change in

the average number of photons over time is obtained from the equation (2.72)

$$\begin{aligned}
\frac{d}{dt} \langle n(t) \rangle &= \sum_{(n)} n \dot{\rho} = \\
&= \sum_{(n)} \left(-R_a (n^2 + n) \rho_{nn} - R_b n^2 \rho_{nn} + \right. \\
&\quad \left. + R_b (n^2 + n) \rho_{n+1, n+1} + R_a n^2 \rho_{n-1, n-1} \right). \tag{2.77}
\end{aligned}$$

Replace the summation variables $m = n + 1$ in the third sum and $m = n - 1$ in the fourth sum. We get:

$$\begin{aligned}
\frac{d}{dt} \langle n \rangle &= -R_a \sum_{(n)} (n^2 + n) \rho_{nn} - R_b \sum_{(n)} n^2 \rho_{nn} + \\
&+ R_b \sum_{(m)} (m^2 - m) \rho_{m,m} + R_a \sum_{(m)} (m^2 + 2m + 1) \rho_{m,m} = \\
&= R_a \sum_{(m)} (m + 1) \rho_{m,m} - R_b \sum_{(m)} m \rho_{m,m} = \\
&= (R_a - R_b) \langle n \rangle + R_a. \tag{2.78}
\end{aligned}$$

At equilibrium

$$\frac{d}{dt} \langle n \rangle = 0,$$

and we obtain the previous relation

$$\langle n_{(\infty)} \rangle = \bar{n} = \frac{R_a}{R_b - R_a} = \frac{1}{\frac{R_b}{R_a} - 1} = \frac{1}{e^{\frac{\hbar\omega}{k_B T}} - 1}. \tag{2.79}$$

Definition 2.9.1 (Q-factor). The Q-factor of an oscillatory system Q is called [48] the number of oscillations performed by the system during the characteristic time of their damping τ . That is, if the change in the energy of the system U (complex amplitude) obeys the equation

$$\frac{dU}{dt} = -\frac{1}{\tau} (U - U_0), \tag{2.80}$$

then

$$U = (U|_{t=0} - U_0) e^{-\frac{t}{\tau}} + U_0$$

Thus

$$Q = \omega\tau$$

therefore the equation (2.80) can be rewritten as

$$\frac{dU}{dt} = -\frac{\omega}{Q} (U - U_0)$$

Equation (2.78) describes the change in the number of photons (energy) over time as a result of relaxation (interaction with a dissipative system). The classical equation describing this process has the form (see def. 2.9.1)

$$\frac{d}{dt} \langle n \rangle = -\frac{\omega}{Q} \langle n \rangle + \frac{\omega}{Q} \langle n \rangle_{\text{eq.}},$$

where $\langle n \rangle_{\text{eq.}} = \bar{n}_T$ is the equilibrium value of $\langle n \rangle$ at temperature T , and Q is the Q-factor of the resonator (see def. 2.9.1). Therefore, it is possible to assume

$$\begin{aligned} R_b - R_a &= \frac{\omega}{Q}, \\ R_a &= \bar{n}_T \frac{\omega}{Q}, \\ R_b &= \frac{\omega}{Q} (1 + \bar{n}_T), \end{aligned} \tag{2.81}$$

i.e., the quantities R_a and R_b introduced by us are expressed through the classical quantity $\frac{\omega}{Q}$, characterizing the losses in the resonator.

Another quantity of interest is the average electromagnetic field

$$\begin{aligned} \bar{E}_{(f)} &= E_0 \sin kz Sp(\hat{\rho}(\hat{a}^\dagger + \hat{a})) = \\ &= E_0 \sin kz Sp(\hat{\rho}\hat{a}) + \text{c.c.} = \\ &= E_1 Sp(\hat{\rho}\hat{a}) + \text{c.c.} = \langle E \rangle + \text{c.c.}, \end{aligned}$$

where $\langle E \rangle$ is the analytical signal of the classical field. The equation satisfied by the field can be obtained using the equation of motion for the density matrix. Write the equation for the statistical operator of the field mode (2.69) and using for R_a , R_b their expressions through Q and \bar{n}_T (2.81), we have

$$\begin{aligned} \dot{\hat{\rho}} &= -\frac{\omega}{2Q} \bar{n}_T (\hat{a}\hat{a}^\dagger \hat{\rho} - 2\hat{a}^\dagger \hat{\rho} \hat{a} + \hat{\rho} \hat{a}\hat{a}^\dagger) - \\ &- \frac{\omega}{2Q} (\bar{n}_T + 1) (\hat{a}^\dagger \hat{a} \hat{\rho} - 2\hat{a} \hat{\rho} \hat{a}^\dagger + \hat{\rho} \hat{a}^\dagger \hat{a}) \end{aligned} \tag{2.82}$$

Since

$$\langle E \rangle = E_1 Sp(\hat{\rho}\hat{a}), \quad \frac{d\langle E \rangle}{dt} = E_1 Sp\left(\frac{d\hat{\rho}}{dt}\hat{a}\right), \tag{2.83}$$

Therefore, equation (2.82) must be multiplied by $E_1 \hat{a}$ and Sp taken. As a result, we obtain:

$$\begin{aligned} \langle \dot{E} \rangle &= -\frac{\omega E_1}{2Q} \bar{n}_T \{ Sp(\hat{a}\hat{a}^\dagger \hat{\rho}\hat{a} - 2\hat{a}^\dagger \hat{\rho}\hat{a}\hat{a} + \hat{\rho}\hat{a}\hat{a}^\dagger \hat{a}) + \\ &+ Sp(\hat{a}^\dagger \hat{a} \hat{\rho}\hat{a} - 2\hat{a} \hat{\rho}\hat{a}^\dagger \hat{a} + \hat{\rho}\hat{a}^\dagger \hat{a}\hat{a}) \} - \\ &- \frac{\omega E_1}{2Q} Sp(\hat{a}^\dagger \hat{a} \hat{\rho}\hat{a} - 2\hat{a} \hat{\rho}\hat{a}^\dagger \hat{a} + \hat{\rho}\hat{a}^\dagger \hat{a}\hat{a}) \end{aligned} \tag{2.84}$$

It is known that under the sign Sp , circular permutation of operators can be performed. Apply this to (2.84). For example

$$\begin{aligned}
& Sp(\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} - 2\hat{a}^\dagger\hat{\rho}\hat{a}\hat{a} + \hat{\rho}\hat{a}\hat{a}^\dagger\hat{a}) = \\
& = Sp(\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} - 2\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} + \hat{\rho}\hat{a}\hat{a}^\dagger\hat{a}) = \\
& = Sp(\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} - 2\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} + \hat{\rho}\hat{a}(\hat{a}\hat{a}^\dagger - 1)) = \\
& Sp(\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} - 2\hat{a}\hat{a}^\dagger\hat{\rho}\hat{a} + (\hat{a}\hat{a}^\dagger - 1)\hat{\rho}\hat{a}) = -Sp(\hat{\rho}\hat{a}).
\end{aligned} \tag{2.85}$$

Doing the same with the second bracket in (2.84), we get

$$\begin{aligned}
& Sp(\hat{a}^\dagger\hat{a}\hat{\rho}\hat{a} - 2\hat{a}\hat{\rho}\hat{a}^\dagger\hat{a} + \hat{\rho}\hat{a}^\dagger\hat{a}\hat{a}) = \\
& = Sp(\hat{\rho}\hat{a}\hat{a}^\dagger\hat{a} - 2\hat{\rho}\hat{a}^\dagger\hat{a}\hat{a} + \hat{\rho}\hat{a}^\dagger\hat{a}\hat{a}) = \\
& = Sp(\hat{\rho}(\hat{a}^\dagger\hat{a} + 1)\hat{a} - 2\hat{\rho}\hat{a}^\dagger\hat{a}\hat{a} + \hat{\rho}\hat{a}^\dagger\hat{a}\hat{a}) = Sp(\hat{\rho}\hat{a}).
\end{aligned} \tag{2.86}$$

Substituting (2.85) and (2.86) into (2.84), we have

$$\begin{aligned}
\langle \dot{E} \rangle &= -\frac{\omega E_1}{2Q} \bar{n}_T \{Sp(\hat{\rho}\hat{a}) - Sp(\hat{\rho}\hat{a})\} - \\
& -\frac{\omega E_1}{2Q} Sp(\hat{\rho}\hat{a}) = -\frac{\omega}{2Q} \langle E \rangle.
\end{aligned} \tag{2.87}$$

Thus, we have related the parameters of the reservoir (parameters of the beam) with the classical value Q and with the average number of photons in the mode at the reservoir temperature. This allows us to write the equations of motion for the density matrix of the field in general form, while the specific model of the reservoir does not matter

$$\begin{aligned}
\dot{\rho}_{nm} &= -\frac{\omega}{2Q} (2\bar{n}_T(n+m+1) + n+m) \rho_{nm} + \\
& + \frac{\omega \bar{n}_T}{Q} \sqrt{nm} \rho_{n-1, m-1} + \frac{\omega}{Q} (\bar{n}_T + 1) \sqrt{(n+1)(m+1)} \rho_{n+1, m+1}
\end{aligned} \tag{2.88}$$

In operator form, this equation is written as

$$\begin{aligned}
\dot{\rho} &= -\frac{\omega}{2Q} \{ \bar{n}_T (\hat{a}\hat{a}^\dagger\hat{\rho} - \hat{a}^\dagger\hat{\rho}\hat{a}) + \\
& + (\bar{n}_T + 1) (\hat{a}^\dagger\hat{a}\hat{\rho} - \hat{a}\hat{\rho}\hat{a}^\dagger) \} + \text{h.c.}
\end{aligned} \tag{2.89}$$

The recording of the equation (2.69), 2.89 in the occupancy number representation is one of many. It is often convenient to use other representations.

2.10 Equation of Motion for the Statistical Operator of the Field Mode in the Coherent State Representation

In quantum optics and quantum electronics, it is often convenient to use coherent states $|\alpha\rangle$. In this case, the statistical operator can be written in diagonal representation (1.118)

$$\hat{\rho}(t) = \int P(\alpha, t) |\alpha\rangle \langle\alpha| d^2\alpha, \quad (2.90)$$

Substituting this expression into the equation of motion for the statistical operator (2.69), 2.89, we get:

$$\begin{aligned} & \int \dot{P}(\alpha, t) |\alpha\rangle \langle\alpha| d^2\alpha = \\ & = -\frac{1}{2} \int P(\alpha, t) [R_a(\hat{a}\hat{a}^\dagger |\alpha\rangle \langle\alpha| - \hat{a}^\dagger |\alpha\rangle \langle\alpha| \hat{a}) + \\ & \quad + R_b(\hat{a}^\dagger \hat{a} |\alpha\rangle \langle\alpha| - \hat{a} |\alpha\rangle \langle\alpha| \hat{a}^\dagger)] d^2\alpha + \text{h.c.} \end{aligned} \quad (2.91)$$

The following equalities hold

$$\begin{aligned} \hat{a}^\dagger |\alpha\rangle \langle\alpha| &= \left(\frac{\partial}{\partial \alpha} + \alpha^* \right) |\alpha\rangle \langle\alpha|, \\ |\alpha\rangle \langle\alpha| \hat{a} &= \left(\frac{\partial}{\partial \alpha^*} + \alpha \right) |\alpha\rangle \langle\alpha|, \end{aligned} \quad (2.92)$$

These equalities are easily proven if we use one of the forms of the definition of the coherent state

$$|\alpha\rangle = e^{\alpha\hat{a}^\dagger - \frac{1}{2}\alpha\alpha^*} |0\rangle.$$

From here we have

$$|\alpha\rangle \langle\alpha| = e^{\alpha\hat{a}^\dagger - \frac{1}{2}\alpha\alpha^*} |0\rangle \langle 0| e^{\alpha^*\hat{a} - \frac{1}{2}\alpha^*\alpha}.$$

Differentiating this expression by α , we get:

$$\begin{aligned} \frac{\partial}{\partial \alpha} |\alpha\rangle \langle\alpha| &= (\hat{a}^\dagger - \alpha^*) e^{\alpha\hat{a}^\dagger - \frac{1}{2}\alpha\alpha^*} |0\rangle \langle 0| e^{\alpha^*\hat{a} - \frac{1}{2}\alpha^*\alpha} = \\ &= (\hat{a}^\dagger - \alpha^*) |\alpha\rangle \langle\alpha| \end{aligned}$$

thus,

$$\hat{a}^\dagger |\alpha\rangle \langle\alpha| = \left(\frac{\partial}{\partial \alpha} + \alpha^* \right) |\alpha\rangle \langle\alpha|. \quad (2.93)$$

In the same way, the validity of the second equality is shown

$$\frac{\partial}{\partial \alpha^*} |\alpha\rangle \langle \alpha| = e^{\alpha \hat{a}^\dagger - \frac{1}{2} \alpha \alpha^*} |0\rangle \langle 0| e^{\alpha^* \hat{a} - \frac{1}{2} \alpha^* \alpha} (\hat{a} - \alpha).$$

Thus,

$$|\alpha\rangle \langle \alpha| \hat{a} = \left(\frac{\partial}{\partial \alpha^*} + \alpha \right) |\alpha\rangle \langle \alpha|, \quad (2.94)$$

Using (2.92) and the known relations for the creation and annihilation operators $\hat{a} |\alpha\rangle = \alpha |\alpha\rangle$, $\langle \alpha| \hat{a}^\dagger = \alpha^* \langle \alpha|$, we transform the equation (2.91) to the form:

$$\begin{aligned} & \int \dot{P}(\alpha, t) |\alpha\rangle \langle \alpha| d^2 \alpha = \\ & = -\frac{1}{2} R_a \int P(\alpha, t) \left\{ - \left(\frac{\partial^2}{\partial \alpha \partial \alpha^*} + \alpha^* \frac{\partial}{\partial \alpha^*} \right) |\alpha\rangle \langle \alpha| \right\} d^2 \alpha - \\ & \quad - \frac{1}{2} R_b \int P(\alpha, t) \alpha \frac{\partial}{\partial \alpha} |\alpha\rangle \langle \alpha| d^2 \alpha + \text{h.c.} \end{aligned} \quad (2.95)$$

We need to consider the integrals involved in (2.95), transforming them using integration by parts. In the integrals involved in (2.95) $d^2 \alpha = dx dy$, if $\alpha = x + iy$, $x = \text{Re } \alpha$, $y = \text{Im } \alpha$. Thus, for integration by parts, it is necessary to express the underintegral function in terms of x , y , perform integration by parts, and then return to the previous variables. But we can proceed differently. Formally consider α and α^* independent variables and express $dx dy$ through $d\alpha d\alpha^*$.

$$d^2 \alpha = dx dy = d\alpha d\alpha^* |J|,$$

where $|J|$ is the Jacobian of the transformation, which in our case is a constant $|J| = \frac{1}{2}$. Now we can integrate by parts, and in the final expression, switch back to the previous notation

$$d\alpha d\alpha^* |J| \rightarrow d^2 \alpha.$$

We have

$$\begin{aligned} & \int P \alpha \frac{\partial}{\partial \alpha} |\alpha\rangle \langle \alpha| d^2 \alpha = \\ & = \alpha P |\alpha\rangle \langle \alpha| \Big|_{-\infty}^{+\infty} - \int \frac{\partial}{\partial \alpha} (\alpha P) |\alpha\rangle \langle \alpha| d^2 \alpha = \\ & = - \int \frac{\partial}{\partial \alpha} (\alpha P) |\alpha\rangle \langle \alpha| d^2 \alpha. \end{aligned} \quad (2.96)$$

Here, we assumed that $\alpha P(\alpha, t) \rightarrow 0$, as $|\alpha| \rightarrow \infty$. The integral with the second derivative, after double integration by parts, is transformed to the form:

$$\begin{aligned} & \int P \frac{\partial^2}{\partial \alpha \partial \alpha^*} |\alpha\rangle \langle \alpha| d^2 \alpha = \\ & = \int \left(\frac{\partial^2}{\partial \alpha \partial \alpha^*} P \right) |\alpha\rangle \langle \alpha| d^2 \alpha \end{aligned} \quad (2.97)$$

Substituting all this into equality (2.95), we obtain:

$$\begin{aligned} & \int \dot{P}(\alpha, t) |\alpha\rangle \langle \alpha| d^2 \alpha = \\ & = - \int \left\{ \frac{1}{2} (R_a - R_b) \left[\frac{\partial}{\partial \alpha} (P \alpha) + \text{c.c.} \right] - \right. \\ & \quad \left. - R_a \frac{\partial^2 P}{\partial \alpha \partial \alpha^*} \right\} |\alpha\rangle \langle \alpha| d^2 \alpha \end{aligned} \quad (2.98)$$

This equality will be satisfied if we set the multiplier for $|\alpha\rangle \langle \alpha|$ equal to zero. Thus, we obtain

$$\dot{P}(\alpha, t) = -\frac{1}{2} (R_a - R_b) \left(\frac{\partial}{\partial \alpha} (\alpha P) + \text{c.c.} \right) + R_a \frac{\partial^2 P}{\partial \alpha \partial \alpha^*} \quad (2.99)$$

This equation is of the Fokker-Planck type, which is used in solving classical problems of statistical physics, such as problems about Brownian motion and other similar tasks. By solving (2.99) with appropriate initial and boundary conditions, it is easy to use (2.90) to write an expression for the statistical operator

$$\hat{\rho}(t) = \int P(\alpha, t) |\alpha\rangle \langle \alpha| d^2 \alpha$$

and using (1.122) and (1.128) to calculate the average values of observable quantities related to the system under consideration

$$\langle O \rangle = \int P(\alpha, t) \langle \alpha | \hat{O} | \alpha \rangle d^2 \alpha = \int P(\alpha, t) O^{(n)}(\alpha, \alpha^*) d^2 \alpha,$$

where $O^{(n)}$ is the normal representation of the operator \hat{O} (1.123). The coefficients R_a and R_b , involved in (2.99), can, as before (2.81), be expressed through Q and \bar{n}_T , then the equation takes the form

$$\frac{\partial}{\partial t} P(\alpha, t) = \frac{1}{2} \frac{\omega}{Q} \left[\frac{\partial}{\partial \alpha} (\alpha P(\alpha, t)) + \text{c.c.} \right] + \frac{\omega \bar{n}_T}{Q} \frac{\partial^2 P(\alpha, t)}{\partial \alpha \partial \alpha^*}. \quad (2.100)$$

As an example, we will calculate the average value of the operator of the positive frequency part of the electric field $E_1 \hat{a}$.

$$\begin{aligned} \frac{d \langle E_1 \hat{a} \rangle}{dt} &= E_1 \int \alpha \dot{P} d^2 \alpha = \\ &= E_1 \frac{1}{2} \frac{\omega}{Q} \int \left[\frac{\partial}{\partial \alpha} (\alpha P) + \text{c.c.} \right] \alpha d^2 \alpha + \\ &\quad + E_1 \frac{\omega \bar{n}_T}{Q} \int \alpha \frac{\partial^2 P}{\partial \alpha \partial \alpha^*} d^2 \alpha \end{aligned} \quad (2.101)$$

Integrating the second term by α^* we get

$$\begin{aligned} \int \alpha \frac{\partial^2 P}{\partial \alpha \partial \alpha^*} d^2 \alpha &= \left. \frac{\partial P}{\partial \alpha} \alpha \right|_{-\infty}^{\infty} - \\ &\quad - \int \frac{\partial P}{\partial \alpha} \alpha \frac{\partial \alpha}{\partial \alpha^*} = 0, \end{aligned}$$

since $\frac{\partial P}{\partial \alpha}$ is assumed to rapidly approach 0 at infinity, and $\frac{\partial \alpha}{\partial \alpha^*} = 0$, since α and α^* are independent variables. The first term can be written as follows

$$\begin{aligned} &\int \alpha \left(\frac{\partial}{\partial \alpha} (\alpha P) + \frac{\partial}{\partial \alpha^*} (\alpha^* P) \right) d^2 \alpha = \\ &= - \int \alpha P \frac{\partial \alpha}{\partial \alpha} d^2 \alpha - \int \alpha^* P \frac{\partial \alpha}{\partial \alpha^*} d^2 \alpha = - \int \alpha P d^2 \alpha = - \langle \hat{a} \rangle. \end{aligned}$$

Thus, finally, we have:

$$\frac{d \langle E_1 \hat{a} \rangle}{dt} = - \frac{\omega}{2Q} \langle E_1 \hat{a} \rangle \quad (2.102)$$

- the result we obtained earlier by another method. Equation (2.100) can be written in various coordinate systems, for example, in polar $\alpha = r e^{i\theta}$, $\alpha^* = r e^{-i\theta}$. By performing the appropriate transformations, we obtain

$$\begin{aligned} \frac{\partial}{\partial t} P(r, \theta, t) &= \frac{1}{2} \frac{\omega}{Q} \frac{\partial}{\partial r} (r^2 P(r, \theta, t)) + \\ &+ \frac{1}{2} \frac{\omega}{Q} \frac{\bar{n}_T}{2r^2} \left(r \frac{\partial}{\partial r} r \frac{\partial}{\partial r} + \frac{\partial^2}{\partial \theta^2} \right) P(r, \theta, t). \end{aligned} \quad (2.103)$$

If we use the coordinates x and y , $\alpha = x + iy$, $\alpha^* = x - iy$, that is $x = \frac{1}{2}(\alpha + \alpha^*)$, $y = -\frac{i}{2}(\alpha - \alpha^*)$, the equation can be represented as

$$\begin{aligned} \frac{\partial}{\partial t} P(x, y, t) &= \frac{1}{2} \frac{\omega}{Q} \left(\frac{\partial}{\partial x} x + \frac{\partial}{\partial y} y \right) P(x, y, t) + \\ &+ \frac{1}{4} \frac{\omega}{Q} \bar{n}_T \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) P(x, y, t). \end{aligned} \quad (2.104)$$

In solving a specific problem, it is convenient to use the equation written either in rectangular or in polar coordinate systems.

2.11 General Theory of Interaction of a Dynamic System with a Thermostat

So far, we have dealt with a specific model of the interaction of a resonator mode (dynamic system) with an atomic beam in equilibrium at temperature T (with a reservoir). Here we consider a more general approach to the problem of interaction between a dynamic system and a reservoir [51]. We are dealing with two interacting systems: A - the dynamic system and B - the reservoir (dissipative system). The systems A and B are weakly connected, meaning the binding energy is small compared to the energy of the dynamic system. We denote the interaction Hamiltonian as \hat{V} . The dissipative system B is a vast system in thermodynamic equilibrium at temperature T . We will initially consider systems A and B as equal. Let the statistical operator of the full system be $\hat{\rho}_{AB}$. We introduce statistical operators $\hat{\rho}_A$ and $\hat{\rho}_B$ related respectively to systems A and B . These are related to the operator $\hat{\rho}_{AB}$ by the relations

$$\hat{\rho}_A = Sp_B \{ \hat{\rho}_{AB} \}, \quad \hat{\rho}_B = Sp_A \{ \hat{\rho}_{AB} \}, \quad (2.105)$$

The equation of motion of the statistical operator of the entire combined system in the interaction representation has the form

$$i\hbar\dot{\hat{\rho}}_{AB} = [\hat{V}, \hat{\rho}_{AB}]. \quad (2.106)$$

The equations of motion for reduced operators are evidently equal to

$$\begin{aligned} i\hbar\dot{\hat{\rho}}_A &= Sp_B [\hat{V}, \hat{\rho}_{AB}], \\ i\hbar\dot{\hat{\rho}}_B &= Sp_A [\hat{V}, \hat{\rho}_{AB}]. \end{aligned} \quad (2.107)$$

We represent the statistical operator $\hat{\rho}_{AB}$ in the following form

$$\hat{\rho}_{AB} = \hat{\rho}_A \otimes \hat{\rho}_B + \hat{\rho}_C, \quad (2.108)$$

where $\hat{\rho}_C$ is a consequence of the interaction between systems, \otimes denotes the outer (tensor) product. At the initial moment t_0 , the systems are assumed to be non-interacting, then the full statistical operator equals the outer product

$$\hat{\rho}_{AB}|_{t=t_0} = \hat{\rho}_A \otimes \hat{\rho}_B,$$

and

$$\hat{\rho}_C|_{t=t_0} = 0.$$

Let us now transform the term $[\hat{V}, \hat{\rho}_A \otimes \hat{\rho}_B]$, included in equation (2.107), and we obtain

$$\begin{aligned} [\hat{V}, \hat{\rho}_A \otimes \hat{\rho}_B] &= \hat{V} \hat{\rho}_A \otimes \hat{\rho}_B - \hat{\rho}_A \otimes \hat{\rho}_B \hat{V} = \\ &= \hat{V} \hat{\rho}_B \otimes \hat{\rho}_A - \hat{\rho}_A \otimes \hat{\rho}_B \hat{V}. \end{aligned} \quad (2.109)$$

since $\hat{\rho}_A \otimes \hat{\rho}_B = \hat{\rho}_B \otimes \hat{\rho}_A$ because $\hat{\rho}_A$ and $\hat{\rho}_B$ act on different variables. From here, performing the operation Sp_B , we have

$$\begin{aligned} Sp_B [\hat{V}, \hat{\rho}_A \otimes \hat{\rho}_B] &= Sp_B (\hat{V} \hat{\rho}_B) \hat{\rho}_A - \\ &- \hat{\rho}_A Sp_B (\hat{\rho}_B \hat{V}) = [\hat{V}_A, \hat{\rho}_A], \end{aligned} \quad (2.110)$$

where $\hat{V}_A = Sp_B (\hat{V} \hat{\rho}_B)$. Similarly, we obtain

$$Sp_A [\hat{V}, \hat{\rho}_A \otimes \hat{\rho}_B] = [\hat{V}_B, \hat{\rho}_B], \quad (2.111)$$

where $\hat{V}_B = Sp_A (\hat{V} \hat{\rho}_A)$. Considering the equalities (2.108), (2.109), (2.110), (2.111), the equations of motion for $\hat{\rho}_A$ and $\hat{\rho}_B$ can be represented as

$$\begin{aligned} i\hbar \frac{d}{dt} \hat{\rho}_A(t) &= [\hat{V}_A(t), \hat{\rho}_A] + Sp_B [\hat{V}, \hat{\rho}_C], \\ i\hbar \frac{d}{dt} \hat{\rho}_B(t) &= [\hat{V}_B(t), \hat{\rho}_B] + Sp_A [\hat{V}, \hat{\rho}_C]. \end{aligned} \quad (2.112)$$

From the form of (2.112) it is clear that \hat{V}_A and \hat{V}_B have the meaning of the energy change of one system connected with its interaction with another, such as the interaction of a dynamic system with a thermostat. These changes are small and in our case can be neglected. Thus (2.112) transforms into

$$\begin{aligned} i\hbar \frac{d}{dt} \hat{\rho}_A(t) &= Sp_B [\hat{V}, \hat{\rho}_C], \\ i\hbar \frac{d}{dt} \hat{\rho}_B(t) &= Sp_A [\hat{V}, \hat{\rho}_C]. \end{aligned} \quad (2.113)$$

For the operator $\hat{\rho}_C$, using the fact that

$$\hat{\rho}_C = \hat{\rho}_{AB} - \hat{\rho}_A \otimes \hat{\rho}_B$$

we can write the equation of motion as follows

$$\begin{aligned} i\hbar \frac{d\hat{\rho}_C}{dt} &= i\hbar \frac{d\hat{\rho}_{AB}}{dt} - i\hbar \frac{d}{dt} (\hat{\rho}_A \otimes \hat{\rho}_B) = \\ &= i\hbar \left(\frac{d\hat{\rho}_{AB}}{dt} - \frac{d\hat{\rho}_A}{dt} \otimes \hat{\rho}_B - \hat{\rho}_A \otimes \frac{d\hat{\rho}_B}{dt} \right). \end{aligned} \quad (2.114)$$

Using the equations of motion for $\hat{\rho}_{AB}$, $\hat{\rho}_A$, $\hat{\rho}_B$ (2.113), for (2.114) we obtain

$$\begin{aligned} i\hbar \frac{d\hat{\rho}_C}{dt} &= \left[\hat{V}, (\hat{\rho}_A \otimes \hat{\rho}_B + \hat{\rho}_C) \right] - \\ &- \left(Sp_B \left[\hat{V}, \hat{\rho}_C \right] \right) \hat{\rho}_B - \hat{\rho}_A \left(Sp_A \left[\hat{V}, \hat{\rho}_C \right] \right) \end{aligned} \quad (2.115)$$

Neglecting second-order small terms, we obtain a simple equation:

$$i\hbar \frac{d\hat{\rho}_C}{dt} = \left[\hat{V}, (\hat{\rho}_A \otimes \hat{\rho}_B) \right]. \quad (2.116)$$

First-order small terms are \hat{V} , $\hat{\rho}_C$. This follows from the assumption that interaction between the dynamic system and the thermostat is small. Terms involving products like $\hat{V}\hat{\rho}_C$ are considered small of the second order. A formal solution to equation (2.116) is

$$\hat{\rho}_C(t) = -\frac{i}{\hbar} \int_{t_0}^t \left[\hat{V}(t'), (\hat{\rho}_A \otimes \hat{\rho}_B) \right] dt' \quad (2.117)$$

with initial conditions $\hat{\rho}_C(t_0) = 0$.

So far, we have considered systems A and B as equal. Now we take into account that the dynamic system A is small, and the dissipative system B is so large that interaction with the dynamic system cannot significantly disrupt its equilibrium. We account for this by setting $\frac{d\hat{\rho}_B}{dt} = 0$:

$$\hat{\rho}_B(t') = \hat{\rho}_B(t_0). \quad (2.118)$$

Substituting in the equation for $\hat{\rho}_A$ (2.112) the equalities (2.115) and (2.117), we obtain

$$\begin{aligned} \frac{d\hat{\rho}_A}{dt} &= -\frac{1}{\hbar^2} \int_{t_0}^t Sp_B \left[\hat{V}(t), \left[\hat{V}(t'), \hat{\rho}_A(t') \otimes \hat{\rho}_B(t_0) \right] \right] dt' = \\ &= -\frac{1}{\hbar^2} \int_{t_0}^t dt' Sp_B \left\{ \left[\hat{V}(t) \hat{V}(t') \hat{\rho}_A(t') \otimes \hat{\rho}_B(t_0) - \right. \right. \\ &\quad \left. \left. - \hat{V}(t) \hat{\rho}_A(t') \otimes \hat{\rho}_B(t_0) \hat{V}(t') \right] - \right. \\ &\quad \left. - \left[\hat{V}(t') \hat{\rho}_A(t') \otimes \hat{\rho}_B(t_0) \hat{V}(t) - \hat{\rho}_A(t') \otimes \hat{\rho}_B(t_0) \hat{V}(t') \hat{V}(t) \right] \right\}. \end{aligned} \quad (2.119)$$

This is the equation for the density operator of the dynamic system A interacting with the dissipative system B . For the vast system B , the correlation time of $\hat{V}(t)$ and $\hat{V}(t')$ is very small, significantly less than the characteristic time change of $\hat{\rho}_A$. Thus, the significant integration region is in the vicinity of $t' = t$. Then, in many cases during integration, we can set

$$\hat{\rho}_A(t') = \hat{\rho}_A(t).$$

For further details on the general theory of interaction of a dynamic system with a thermostat, see [51]. We will apply the general theory to specific cases.

2.12 Relaxation of an Atom in the Simplest Reservoir Consisting of Harmonic Oscillators

Consider a multimode electromagnetic field described using the creation and annihilation operators $\hat{a}_k^\dagger, \hat{a}_k$. The frequency of the k -th mode is denoted as ω_k .

This electromagnetic field interacts with a certain two-level atom. The atom can be in two states: excited $|a\rangle$ and unexcited $|b\rangle$. The transition operators remain $\hat{\sigma}$ and $\hat{\sigma}^\dagger$. The transition frequency is denoted as Ω . The interaction Hamiltonian in this case is

$$\hat{V}(t) = \hbar \sum_{(k)} g_k \hat{a}_k^\dagger \hat{\sigma} e^{-i(\Omega - \omega_k)t} + \text{h.c.} \quad (2.120)$$

where g_k is the coupling constant.

Thus, system B (the reservoir) will be a multimode electromagnetic field, and the dynamic system under study A is a two-level atom. The density operator of the reservoir (electromagnetic field) is denoted as $\hat{\rho}_R$. The density operator of the dynamic system under study (atom) is denoted as $\hat{\rho}_{at}$.

The following relations hold:

$$Sp_R \left(\hat{a}_k^\dagger \hat{a}_k \hat{\rho}_R \right) = \bar{n}_k \quad (2.121)$$

where \bar{n}_k is the average number of photons in the mode of the reservoir;

$$\begin{aligned} Sp_R \left(\hat{a}_k \hat{a}_k^\dagger \hat{\rho}_R \right) &= \bar{n}_k + 1, \\ Sp_R \left(\hat{a}_k \hat{a}_k \hat{\rho}_R \right) &= Sp_R \left(\hat{a}_k^\dagger \hat{a}_k^\dagger \hat{\rho}_R \right) = 0. \end{aligned} \quad (2.122)$$

Using these equalities, the general equation (2.119) can be represented as:

$$\begin{aligned} \dot{\hat{\rho}}_{at}(t) = & - \int_{t_0}^t \sum_{(k)} g_k^2 \left\{ \hat{\sigma}^\dagger \hat{\sigma} \hat{\rho}_{at}(t') (\bar{n}_k + 1) e^{i(\Omega - \omega_k)(t-t')} + \right. \\ & + \hat{\sigma} \hat{\sigma}^\dagger \hat{\rho}_{at}(t') \bar{n}_k e^{-i(\Omega - \omega_k)(t-t')} - \hat{\sigma}^\dagger \hat{\rho}_{at}(t') \hat{\sigma} \bar{n}_k e^{i(\Omega - \omega_k)(t-t')} - \\ & \left. - \hat{\sigma} \hat{\rho}_{at}(t') \hat{\sigma}^\dagger (\bar{n}_k + 1) e^{-i(\Omega - \omega_k)(t-t')} \right\} dt' + \text{h.c.} \end{aligned} \quad (2.123)$$

Summation over k , considering the quasi-continuous spectrum of the dissipative system's states, can be replaced by integration (1.44)

$$\sum_{(k)} \rightarrow \int D(\omega) d\omega,$$

where $D(\omega)$ is the spectral density of states. Replace $D(\omega)$, $\bar{n}(\omega)$, $g^2(\omega)$ with D , \bar{n} , g^2 - their values at $\omega = \Omega$, assuming these functions slowly depend on frequency. Using the equality (see integral representation of the delta function (18.3))

$$\begin{aligned} \int_0^\infty e^{\pm(\Omega - \omega)(t-t')} d\omega &= |\nu = \omega - \Omega| = \\ &= \int_{-\Omega}^\infty e^{-\nu(t \mp t')} d\nu \approx \int_{-\infty}^\infty e^{-\nu(t \mp t')} d\nu = \\ &= 2\pi\delta(t \mp t') \end{aligned} \quad (2.124)$$

and integrate over time. This leads us to the equation

$$\begin{aligned} \dot{\hat{\rho}}_{at} = & -\frac{1}{2}\gamma \left\{ \bar{n} [\hat{\sigma} \hat{\sigma}^\dagger \hat{\rho}_{at} - \hat{\sigma}^\dagger \hat{\rho}_{at} \hat{\sigma}] + \right. \\ & \left. + (\bar{n} + 1) [\hat{\sigma}^\dagger \hat{\sigma} \hat{\rho}_{at} - \hat{\sigma} \hat{\rho}_{at} \hat{\sigma}^\dagger] \right\} + \text{h.c.} \end{aligned} \quad (2.125)$$

where $\gamma = 4\pi D g^2$, \bar{n} taken at the mode frequency Ω . Equation (2.125) has the same form as equation (2.89), obtained earlier. The equations fully coincide if we accept $\gamma = \frac{\omega}{Q}$ and make the substitution $\hat{\sigma} \rightarrow \hat{a}$, $\hat{\sigma}^\dagger \rightarrow \hat{a}^\dagger$. This again confirms that the final result does not depend on the nature of the reservoir.

At a reservoir temperature $T = 0$, photons are absent and $\bar{n} = 0$. Assume that initially the atom is excited

$$\hat{\rho}_{at}(t_0) = 1.$$

Using the properties (2.10) - (2.9) of the operators $\hat{\sigma}$ and $\hat{\sigma}^\dagger$, we obtain

$$\begin{aligned}\dot{\rho}_{aa} &= \langle a | \dot{\rho}_{at} | a \rangle = -\frac{1}{2}\gamma \langle a | \hat{\sigma}^\dagger \hat{\sigma} \hat{\rho}_{at} | a \rangle + \\ &\quad + \frac{1}{2}\gamma \langle a | \hat{\sigma} \hat{\rho}_{at} \hat{\sigma}^\dagger | a \rangle + \text{c.c.} = \\ &= -\frac{1}{2}\gamma \langle a | \hat{\rho}_{at} | a \rangle + \text{c.c.} = -\gamma \rho_{aa}.\end{aligned}\tag{2.126}$$

Here we used the relations $\hat{\sigma}^\dagger |a\rangle = 0$, $\hat{\sigma}^\dagger |b\rangle = |a\rangle$, $\hat{\sigma} |a\rangle = |b\rangle$, $\hat{\sigma} |b\rangle = 0$ and their conjugates. The final result

$$\dot{\rho}_{aa} = -\gamma \rho_{aa}.\tag{2.127}$$

This means that as a result of spontaneous emission, the probability of finding the atom at the upper level exponentially decreases.

2.13 Resonator Mode Damping. Heisenberg-Langevin Equation.

2.13.1 Problem Statement.

The problem of field relaxation (mode) of a resonator is considered using the density matrix method, i.e., the Schrödinger representation. There is another approach, attributed to Langevin, where the equations of a dynamic system for accounting for the influence of a dissipative system are supplemented by random forces with known statistical properties. In the quantum case, the dynamic equations are Heisenberg equations for operators of observable quantities, and random influences associated with the thermostat are accounted for by adding a certain noise operator. The resulting equations are known as Heisenberg-Langevin equations.

We consider the problem of relaxing the mode of a resonator (harmonic oscillator) interacting with a thermostat using this method. In the Heisenberg representation, operators depend on time and satisfy the Heisenberg equations:

$$\frac{d\hat{O}}{dt} = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{O}],$$

where $\hat{\mathcal{H}}$ is the Hamiltonian of the system considered. For example, for the free field, we previously obtained the equations (1.61):

$$\frac{d\hat{a}}{dt} = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{a}] = -i\omega\hat{a}, \quad \frac{d\hat{a}^\dagger}{dt} = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{a}^\dagger] = i\omega\hat{a}^\dagger,$$



Figure 2.17: Interaction of the electromagnetic field mode of the resonator F with a reservoir of atoms R at temperature T . \hat{a}^\dagger and \hat{a} - creation and annihilation operators of the studied electromagnetic field mode. \hat{b}_k^\dagger and \hat{b}_k - creation and annihilation operators of the k -th mode of the reservoir (phonons).

where ω is the mode frequency, and \hat{a} and \hat{a}^\dagger are the annihilation and creation operators.

The problem in question is schematically shown in Figure 2.17. The electromagnetic field mode, considered as a dynamic system, interacts with a vast reservoir consisting of a large number of harmonic oscillators (phonons) in equilibrium at temperature T . The Hamiltonian of such a system can be represented as

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_0 + \hat{V},$$

where

$$\hat{\mathcal{H}}_0 = \hbar\omega\hat{a}^\dagger\hat{a} + \sum_{(k)} \hbar\omega_k\hat{b}_k^\dagger\hat{b}_k$$

is the Hamiltonian of the system without taking into account the interaction between its parts,

$$\hat{V} = \hbar \sum_{(k)} g_k \left(\hat{b}_k^\dagger \hat{a} + \hat{a}^\dagger \hat{b}_k \right) \quad (2.128)$$

is the Hamiltonian of interaction between the dynamic system and the thermostat. In (2.128), g_k denotes the interaction constant for the k -th mode of the reservoir. \hat{a} and \hat{a}^\dagger are annihilation and creation operators for the field mode, \hat{b}_k and \hat{b}_k^\dagger are annihilation and creation operators for the k -th mode of the reservoir (phonons). Vacuum terms in the Hamiltonian are not accounted for, as they cancel out during the derivation of the equations.

2.13.2 Heisenberg Equations of Motion for Operators.

The Heisenberg equations of motion for \hat{a} and \hat{b}_k are

$$\begin{aligned}\frac{d\hat{a}}{dt} &= \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{a}], \\ \frac{d\hat{b}_k}{dt} &= \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{b}_k]\end{aligned}$$

using the commutation relations $[\hat{a}, \hat{a}^\dagger] = 1$ and $[\hat{b}_k, \hat{b}_k^\dagger] = 1$ are reduced to the form

$$\begin{aligned}\frac{d\hat{a}(t)}{dt} &= -i\omega\hat{a}(t) - i \sum_{(k)} g_k \hat{b}_k(t), \\ \frac{d\hat{b}_k(t)}{dt} &= -i\omega_k \hat{b}_k(t) - ig_k \hat{a}(t).\end{aligned}\tag{2.129}$$

From the system (2.129) \hat{b}_k can be eliminated by integrating the second equation and substituting the result into the first equation. Integration is easiest to perform using operational calculus. We obtain

$$\hat{b}_k(t) = \hat{b}_k(0) e^{-i\omega_k t} - ig_k \int_0^t dt' \hat{a}(t') e^{-i\omega_k(t-t')}$$

and further

$$\begin{aligned}\frac{d\hat{a}(t)}{dt} &= -i\omega\hat{a}(t) - \sum_{(k)} g_k^2 \int_0^t dt' \hat{a}(t') e^{-i\omega_k(t-t')} + \hat{f}(t), \\ \hat{f}(t) &= -i \sum_{(k)} g_k \hat{b}_k(0) e^{-i\omega_k t}.\end{aligned}\tag{2.130}$$

The operator $\hat{f}(t)$ depends on the thermostat variables and is a noise operator describing the impact of the reservoir on the dynamic system.

The equation (2.130) is integro-differential. Under certain approximations, it can be transformed into a purely differential one. First, we switch to slow variables by writing

$$\begin{aligned}\hat{a}(t) &= \hat{A}(t) e^{-i\omega t}, \\ \hat{f}(t) &= \hat{F}(t) e^{-i\omega t},\end{aligned}$$

where ω is the mode frequency. From this, we have

$$\frac{d\hat{A}(t)}{dt} = - \sum_{(k)} g_k^2 \int_0^t dt' \hat{A}(t') e^{i(\omega - \omega_k)(t-t')} + \hat{F}(t).\tag{2.131}$$

The first term of expression (2.131) resembles the form encountered when considering spontaneous emission using the Weisskopf-Wigner method. The approach used there can also be applied here. From expression (2.131) it is clear that the main contribution comes from terms with frequencies close to the mode frequency: $\omega_k \approx \omega$. For this reason, all slowly varying terms can be taken at frequency ω and taken out of the integral. Summation over k , assuming the phonon spectrum is quasi-continuous, can be replaced by integration over frequency:

$$\begin{aligned} & \sum_{(k)} g_k^2 \int_0^t dt' \hat{A}(t') e^{i(\omega - \omega_k)(t-t')} = \\ &= \int_0^t dt' \hat{A}(t') \int_0^\infty g^2(\omega) D(\omega) e^{i(\omega - \omega')(t-t')} d\omega' = \\ &= g^2(\omega) D(\omega) \int_0^t dt' \hat{A}(t') \int_0^\infty e^{i(\omega - \omega')(t-t')} d\omega', \end{aligned}$$

where $D(\omega)$ is the frequency density of states. $D(\omega)$ and $g^2(\omega)$ are taken at the field mode frequency and factored out of the integral.

Consider the frequency integral. Proceeding as in the consideration of spontaneous emission using the Weisskopf-Wigner method, we obtain

$$\begin{aligned} \int_0^\infty e^{-i(\omega' - \omega)(t-t')} d\omega' &= \int_{-\omega}^\infty e^{-i\nu(t-t')} d\nu \approx \int_{-\infty}^\infty e^{-i\nu(t-t')} d\nu = \\ &= \int_{-\infty}^\infty e^{i\nu(t'-t)} d\nu = 2\pi \delta(t' - t). \end{aligned}$$

Consequently, we can write

$$\begin{aligned} \frac{d\hat{A}(t)}{dt} &= -2\pi g^2(\omega) D(\omega) \int_0^t dt' \hat{A}(t') \delta(t - t') + \hat{F}(t) = \\ &= -2\pi g^2(\omega) D(\omega) \hat{A}(t) + \hat{F}(t). \end{aligned}$$

Let us denote $2\pi D g^2|_\omega = \frac{\gamma}{2}$ - the damping coefficient. We obtain

$$\frac{d\hat{A}(t)}{dt} = -\frac{\gamma}{2} \hat{A}(t) + \hat{F}(t), \quad (2.132)$$

where $\hat{F}(t)$ is a random noise operator whose properties need to be determined. As will be seen later, the noise operator ensures the preservation of commutation relations for the field mode operator.

Let us now examine the statistical properties characterized by the correlation functions of \hat{F} . Suppose the reservoir is in thermal equilibrium at temperature

T . Then we have:

$$\begin{aligned}
\langle \hat{b}_k(0) \rangle_R &= \langle \hat{b}_k^\dagger(0) \rangle_R = 0, \\
\langle \hat{b}_k^\dagger(0) \hat{b}_{k'}(0) \rangle_R &= \bar{n}_k \delta_{kk'}, \\
\langle \hat{b}_k(0) \hat{b}_{k'}^\dagger(0) \rangle_R &= (\bar{n}_k + 1) \delta_{kk'}, \\
\langle \hat{b}_k(0) \hat{b}_{k'}(0) \rangle_R &= \langle \hat{b}_k^\dagger(0) \hat{b}_{k'}^\dagger(0) \rangle_R = 0,
\end{aligned} \tag{2.133}$$

where $\langle \dots \rangle_R$ means averaging over the thermostat variables, and \bar{n}_k is the average number of phonons in mode k . The noise operator was defined by the formula

$$\hat{F}(t) = -i \sum_{(k)} g_k \hat{b}_k(0) e^{-i(\omega_k - \omega)t}. \tag{2.134}$$

Using (2.133), it is easy to see

$$\begin{aligned}
\langle \hat{F}(t) \rangle_R &= \langle \hat{F}^\dagger(t) \rangle_R = 0, \\
\langle \hat{F}^\dagger(t) \hat{F}(t') \rangle_R &= \\
&= \sum_k \sum_{k'} g_k g_{k'} \langle \hat{b}_k^\dagger(0) \hat{b}_{k'}(0) \rangle_R e^{i(\omega_k - \omega)t} e^{-i(\omega_{k'} - \omega)t'} = \\
&= \sum_k g_k^2 \bar{n}_k e^{i(\omega_k - \omega)(t - t')} = \\
&= \int_0^\infty g^2(\omega) D(\omega) \bar{n}_T e^{i(\omega' - \omega)(t - t')} d\omega' = \\
&= 2\pi g^2(\omega) D(\omega) \bar{n}_T \delta(t - t') = \frac{\gamma \bar{n}_T}{2} \delta(t - t'),
\end{aligned} \tag{2.135}$$

where \bar{n}_T is the average number of phonons in the reservoir mode at temperature T , determined by Planck's formula, $\gamma = 4\pi g^2(\omega) D(\omega)$. If, according to Langevin, we introduce the diffusion coefficient $\mathcal{D} = \frac{\gamma \bar{n}_T}{2}$, then the last equation of the system (2.135) can be written as

$$\langle \hat{F}^\dagger(t) \hat{F}(t') \rangle_R = \mathcal{D} \delta(t - t').$$

Note that $\langle \dots \rangle_R$ means averaging over the thermostat variables, so the result of averaging will be a correlation function that depends only on the variables of the dynamic system (field).

Similarly proved

$$\langle \hat{F}(t) \hat{F}^\dagger(t') \rangle_R = \frac{\gamma(\bar{n}_T + 1)}{2} \delta(t - t') \tag{2.136}$$

and

$$\left\langle \hat{F}(t) \hat{F}(t') \right\rangle_R = \left\langle \hat{F}^\dagger(t) \hat{F}^\dagger(t') \right\rangle_R = 0. \quad (2.137)$$

Thus, in our approximation, the noise is δ -correlated.

We will also need a correlation function of the form $\left\langle \hat{F}^\dagger(t) \hat{A}(t') \right\rangle_R$ - necessary to derive the equation that $\left\langle \hat{A}^\dagger(t) \hat{A}(t') \right\rangle_R$ satisfies. For this, we integrate the equation for $\hat{A}(t)$ (2.132) and, as a result, using operational calculus, we obtain

$$\hat{A}(t) = \hat{A}(0) e^{-\frac{\gamma}{2}t} + \int_0^t dt' e^{-\frac{\gamma}{2}(t-t')} \hat{F}(t').$$

Multiply this expression by $\hat{F}^\dagger(t)$ and average over the thermostat. We obtain

$$\begin{aligned} \left\langle \hat{F}^\dagger(t) \hat{A}(t) \right\rangle_R &= \int_0^t \left\langle \hat{F}^\dagger(t) \hat{F}(t') \right\rangle_R e^{-\frac{\gamma}{2}(t-t')} dt' + \\ &\quad + \left\langle \hat{F}^\dagger(t) \hat{A}(0) \right\rangle_R e^{-\frac{\gamma}{2}t}. \end{aligned} \quad (2.138)$$

The last term in (2.138) is 0, as $\hat{F}^\dagger(t)$ and $\hat{A}(0)$ are independent, and their averages are 0. Taking into account the second equation of the system (2.135) i.e., $\left\langle \hat{F}^\dagger(t) \hat{F}(t') \right\rangle_R = \frac{\gamma \bar{n}_T}{2} \delta(t - t')$, we obtain

$$\left\langle \hat{F}^\dagger(t) \hat{A}(t) \right\rangle_R = \frac{\gamma \bar{n}_T}{2}.$$

Now we can consider the correlation function $\left\langle \hat{A}^\dagger(t) \hat{A}(t) \right\rangle_R$. Differentiating this function over time, we obtain

$$\frac{d \left\langle \hat{A}^\dagger(t) \hat{A}(t) \right\rangle_R}{dt} = \left\langle \frac{d \hat{A}^\dagger(t)}{dt} \hat{A}(t) \right\rangle_R = \left\langle \frac{d \hat{A}^\dagger(t)}{dt} \hat{A}(t) \right\rangle_R + \left\langle \hat{A}^\dagger(t) \frac{d \hat{A}(t)}{dt} \right\rangle_R. \quad (2.139)$$

Then, using the equation (2.132) and its conjugate, we obtain

$$\frac{d \left\langle \hat{A}^\dagger(t) \hat{A}(t) \right\rangle_R}{dt} = -\gamma \left\langle \hat{A}^\dagger(t) \hat{A}(t) \right\rangle_R + \left\langle \hat{F}^\dagger(t) \hat{A}(t) \right\rangle_R + \left\langle \hat{A}^\dagger(t) \hat{F}(t) \right\rangle_R.$$

Considering the expressions

$$\left\langle \hat{F}^\dagger(t) \hat{A}(t) \right\rangle_R = \left\langle \hat{A}^\dagger(t) \hat{F}(t) \right\rangle_R = \frac{\gamma \bar{n}_T}{2},$$

finally, we have

$$\frac{d \langle \hat{A}^\dagger(t) \hat{A}(t) \rangle_R}{dt} = -\gamma \langle \hat{A}^\dagger(t) \hat{A}(t) \rangle_R + \gamma \bar{n}_T.$$

Similarly, we obtain

$$\frac{d \langle \hat{A}(t) \hat{A}^\dagger(t) \rangle_R}{dt} = -\gamma \langle \hat{A}(t) \hat{A}^\dagger(t) \rangle_R + \gamma (\bar{n}_T + 1). \quad (2.140)$$

Subtracting these expressions from each other, we arrive at the commutator

$$\frac{d \langle [\hat{A}(t), \hat{A}^\dagger(t)] \rangle_R}{dt} = \gamma \left(1 - \langle [\hat{A}(t), \hat{A}^\dagger(t)] \rangle_R \right). \quad (2.141)$$

From equation (2.141) it follows that if at the initial time

$$\langle [\hat{A}(0), \hat{A}^\dagger(0)] \rangle_R = 1,$$

i.e., commutation relations hold, then these relations will always hold. In other words, commutation relations averaged over the thermostat variables are preserved over time.

The picture would be completely different if quantum noises were not taken into account. In this case, the equation (2.141) would take the form

$$\frac{d \langle [\hat{A}(t), \hat{A}^\dagger(t)] \rangle_R}{dt} = -\gamma \langle [\hat{A}(t), \hat{A}^\dagger(t)] \rangle_R,$$

whose solution is

$$\langle [\hat{A}(t), \hat{A}^\dagger(t)] \rangle_R = \langle [\hat{A}(0), \hat{A}^\dagger(0)] \rangle_R e^{-\gamma t} = e^{-\gamma t}.$$

This means that, over a long enough time, the operators \hat{A} and \hat{A}^\dagger commute, which is not observed in reality. Therefore, taking into account the quantum noise term is mandatory.

The equation of motion for the field averaged over the thermostat variables is easily obtained from (2.132). To do this, we first write the averaged field

$$\begin{aligned} \langle \hat{E}(t) \rangle_R &= \langle E_0 \sin kz (\hat{a}(t) + \hat{a}^\dagger(t)) \rangle_R = \\ &= E_0 \sin kz \left(\langle \hat{A}(t) \rangle_R e^{-i\omega t} + \langle \hat{A}^\dagger(t) \rangle_R e^{i\omega t} \right), \end{aligned}$$

where it is considered $\hat{a} = \hat{A}e^{-i\omega t}$. From (2.132) we have

$$\frac{d\langle \hat{A}(t) \rangle_R}{dt} = -\frac{\gamma}{2} \langle \hat{A}(t) \rangle_R,$$

since $\langle \hat{F}(t) \rangle_R = 0$. Thus, over time, $\langle \hat{A}(t) \rangle_R$ and $\langle \hat{A}^\dagger(t) \rangle_R$ tend to zero along with $\langle \hat{E}(t) \rangle_R$:

$$\langle \hat{E}(t) \rangle_R \rightarrow 0.$$

2.13.3 Fluctuation-Dissipation Formula

We obtained expression (2.135):

$$\langle \hat{F}^\dagger(t) \hat{F}(t') \rangle_R = \frac{\gamma \bar{n}_T}{2} \delta(t - t').$$

Integrating both sides by dt' , we obtain

$$\gamma = \frac{2}{\bar{n}_T} \int_0^\infty \langle \hat{F}^\dagger(t) \hat{F}(t') \rangle_R dt', \quad (2.142)$$

i.e., the damping rate and noise fluctuations are related by the relation (2.142). Greater damping corresponds to greater noise.

2.14 Exercises

1. Based on (2.13), derive expression (2.14) for the operator \hat{V} in the interaction representation.
2. Derive the system of equations (2.19).
3. Obtain formula (2.64).
4. Derive from (2.69) the equation for the density matrix in the occupation number representation (2.71).
5. Transform formula (2.91) to the form (2.95).
6. By integrating by parts, obtain expressions (2.96), (2.97).
7. Write the general equation (2.99) in polar coordinates (2.103).

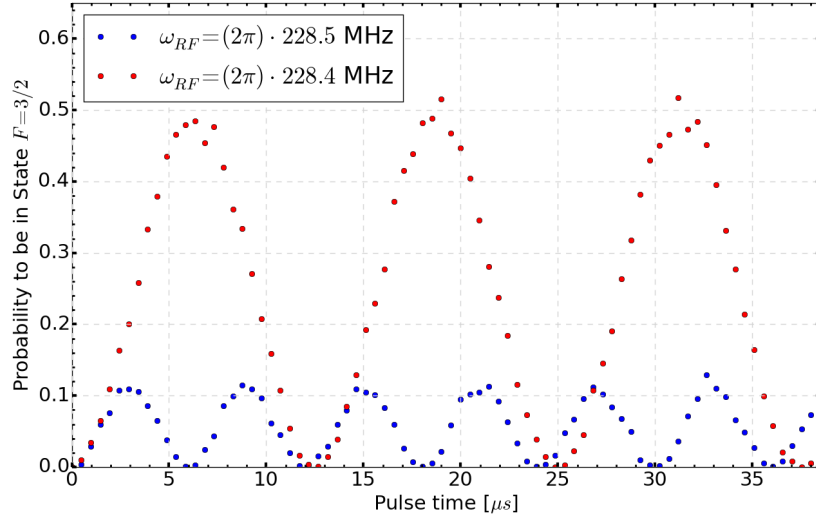


Figure 2.18: You want to measure the transition frequency $f_0 = \frac{\omega_o}{2\pi}$ between two states of a lithium atom by evaluating Rabi oscillations. You prepare a cold cloud of lithium atoms in state $|1\rangle$ (the so-called $|F = 1/2\rangle$ state) and you can apply an electromagnetic pulse of fixed intensity and variable duration τ . Ultimately, you measure the number of atoms in state $|2\rangle$ (the so-called $|F = 3/2\rangle$ state). You know that the frequency difference between these two levels is close to $2\pi \cdot 228.5\text{MHz}$, so you apply an electromagnetic pulse with a frequency close to this value: first with frequency $\omega_1 = 2\pi \cdot 228.5\text{MHz}$, then you repeat the measurements with frequency $\omega_2 = 2\pi \cdot 228.4\text{MHz}$. The result is shown in the presented figure. What is the transition frequency $f_0 = \frac{\omega_o}{2\pi}$ between the two states of the lithium atom?

8. Write the general equation (2.99) in rectangular coordinates (2.104).
9. Derive relations (2.121) and (2.122).
10. Derive equation (2.123).
11. Prove (2.136) and (2.137).
12. Obtain expression (2.140).
13. What is the transition frequency $f_0 = \frac{\omega_o}{2\pi}$ between two states of a lithium atom if experimental data have been obtained as shown in Figure 2.18 ²

²Taken from exercises [14].

Part 2

Quantum Optics

Chapter 3

Quantum Theory of the Laser

The semiclassical theory of the laser cannot answer all the questions that arise in connection with its operation. According to this theory, the laser does not generate at all before reaching the threshold, and when the threshold is exceeded, it begins to generate a classical electromagnetic field (light).

In reality, noticeably below the threshold, the laser generates chaotic light, and significantly above the threshold, its emission is close to classical. At and near the threshold, there is a transition region from chaotic light to ordered emission. Only a fully quantum theory can adequately describe this.

Another task, also requiring quantization of the electromagnetic field, is determining the ultimate (natural) linewidth of laser emission when the linewidth is determined by quantum fluctuations of the field, and various external influences, fundamentally eliminable, are not considered.

The laser is an open system in which active atoms and the resonator field are connected to large external systems that we will call reservoirs, providing pumping and losses. Therefore, the laser, as an open system, must be considered using a density matrix.

3.1 Laser Model

Earlier in (2.69) we considered the interaction of a resonator mode (harmonic oscillator) with a reservoir in thermal equilibrium at temperature T . The same model can be used to study the operation of a laser, but it needs to be supplemented by an additional reservoir that provides pumping. See Figure 3.1: the first reservoir, R_1 , contains atoms at temperature T . The second, R_2 , contains atoms (of a different type) in an inverted state.

In the linear approximation (to second order perturbation theory) we already solved the problem of the interaction of the resonator mode with a reservoir. For the density matrix describing the state of the resonator mode field, the equations

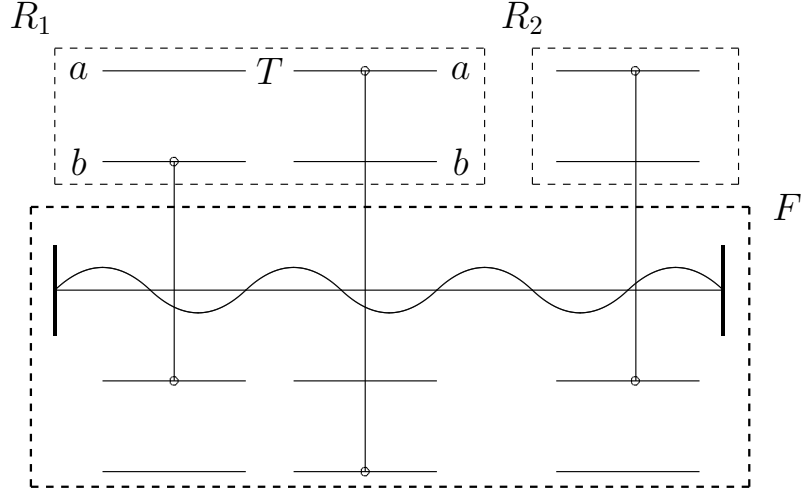


Figure 3.1: The considered laser model. The resonator mode (electromagnetic field) F interacts with reservoirs R_1 and R_2 . R_1 contains atoms in thermal equilibrium at temperature T (losses). R_2 contains atoms in an inverted state (pumping).

obtained were:

$$\begin{aligned} \dot{\rho} = & -\frac{1}{2} \frac{\omega}{Q} \{ \bar{n}_T (\hat{a} \hat{a}^\dagger \hat{\rho} - \hat{a}^\dagger \hat{\rho} \hat{a}) + \\ & + (\bar{n}_T + 1) (\hat{a}^\dagger \hat{a} \hat{\rho} - \hat{a} \hat{\rho} \hat{a}^\dagger) \} + \text{h.c.} \end{aligned} \quad (3.1)$$

or in another form (2.69)

$$\begin{aligned} \dot{\rho} = & -\frac{1}{2} R_a (\hat{a} \hat{a}^\dagger \hat{\rho} - \hat{a}^\dagger \hat{\rho} \hat{a}) - \\ & -\frac{1}{2} R_b (\hat{a}^\dagger \hat{a} \hat{\rho} - \hat{a} \hat{\rho} \hat{a}^\dagger) + \text{h.c.} \end{aligned} \quad (3.2)$$

If we limit ourselves to the linear approximation (in the field), the equation for the laser density matrix can be written using these equations. To describe the action of the reservoir introducing losses, it is convenient to use (3.1), taking into account that at not too high temperatures $\bar{n}_T \ll 1$ and \bar{n}_T can be neglected compared to unity.

To describe the reservoir providing pumping, it is convenient to use (3.2), discarding the second term, since for simplicity it is assumed that all atoms of the pumping reservoir are in the upper state.

All this leads to the following equation:

$$\begin{aligned} \dot{\rho} = & -\frac{1}{2} \frac{\omega}{Q} (\hat{a}^\dagger \hat{a} \hat{\rho} - \hat{a} \hat{\rho} \hat{a}^\dagger) - \\ & -\frac{1}{2} A (\hat{a} \hat{a}^\dagger \hat{\rho} - \hat{a}^\dagger \hat{\rho} \hat{a}) + \text{h.c.} \end{aligned} \quad (3.3)$$

where $A = R_a$ is determined by the pumping intensity.

Equation (3.3) is obtained in the second-order perturbation approximation. It can describe the behavior of the laser below the generation threshold, allows determining threshold conditions, but cannot describe the laser above the generation threshold.

Let us assume losses are linear, and the first term in (3.3) accurately describes the losses. The second term in (3.3) should be found in the next non-zero approximation (fourth order perturbation theory).

The procedure to find this approximation is analogous to the one used earlier in the derivation of equations (3.1) - (3.2). One just needs to continue it to higher order terms. Third-order terms will give zero, since the resulting matrix has zero diagonal elements and its trace is zero. The next non-zero term is fourth order and has the form:

$$Sp_{at} \left\{ \left(-\frac{i}{\hbar} \right)^4 \int_t^{t+\tau} dt_1 \int_t^{t_1} dt_2 \int_t^{t_2} dt_3 \int_t^{t_3} \left[\hat{V}, \left[\hat{V}, \left[\hat{V}, \left[\hat{V}, \hat{\rho}_{at}(t) \otimes \hat{\rho}_f(t) \right] \right] \right] \right] \right] dt_4 \right\} \quad (3.4)$$

Calculations similar to those done earlier lead to the expression (see [49]):

$$\frac{1}{8}B \left[(\hat{a}\hat{a}^\dagger)^2 \hat{\rho} + 3\hat{a}\hat{a}^\dagger \hat{\rho} \hat{a}\hat{a}^\dagger - 4\hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{\rho} \hat{a} \right] + \text{h.c.} \quad (3.5)$$

Taking (3.5) into account, the equation for the density matrix (statistical operator) of the laser mode takes the form:

$$\begin{aligned} \dot{\hat{\rho}}(t) = & -\frac{1}{2} \frac{\omega}{Q} (\hat{a}^\dagger \hat{a} \hat{\rho} - \hat{a} \hat{\rho} \hat{a}^\dagger) - \frac{1}{2} A (\hat{a} \hat{a}^\dagger \hat{\rho} - \hat{a}^\dagger \hat{\rho} \hat{a}) + \\ & + \frac{1}{8} B \left[(\hat{a}\hat{a}^\dagger)^2 \hat{\rho} + 3\hat{a}\hat{a}^\dagger \hat{\rho} \hat{a}\hat{a}^\dagger - 4\hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{\rho} \hat{a} \right] + \text{h.c.} \end{aligned} \quad (3.6)$$

where A is the linear (unsaturated) gain, $B = \frac{g^2 \tau^2 A}{3} = \frac{1}{3} r_a g^4 \tau^4$ is the saturation parameter.

Equation (3.6) is the motion equation for the density matrix of the laser field interacting with a nonlinear medium consisting of active atoms and linear losses.

Equation (3.6) can be written in various representations. Here we restrict ourselves to the photon number (photon number states) representation and the coherent state representation.

In the first case, from (3.6) it is easy to obtain a system of equations for the matrix elements $\langle m | \hat{\rho} | n \rangle = \rho_{mn}$.

For the diagonal elements we have:

$$\begin{aligned} \dot{\rho}_{nn}(t) = & -[A - (n+1)B](n+1)\rho_{nn} + \\ & + (A - nB)n\rho_{n-1,n-1} - \frac{\omega}{Q}n\rho_{nn} + \frac{\omega}{Q}(n+1)\rho_{n+1,n+1}. \end{aligned} \quad (3.7)$$

Note that only diagonal terms enter this equation.

Equations for off-diagonal terms are obtained in the same way. They are somewhat more complicated and are not presented here.

To write equation (3.6) in the coherent state representation, one can proceed as before when considering the decay of the resonator mode. Write

$$\hat{\rho} = \int P(\alpha) |\alpha\rangle \langle \alpha| d^2\alpha$$

and substitute it into equation (3.6). Then perform the familiar procedure used previously for resonator mode relaxation. For the linear terms characterizing losses and unsaturated gain, the previously obtained results can be used. Additionally, the terms characterizing saturation must be considered. The calculations are more cumbersome here, although not fundamentally different from those done before. The resulting equation is quite complex, but recalling that perturbation theory (small field) was used, one can neglect smaller terms, leaving only the main ones [49]. In conclusion, one obtains an equation of the Fokker-Planck type for the quasi-probability $P(\alpha, t)$:

$$\frac{\partial}{\partial t} P(\alpha, t) = -\frac{1}{2} \left\{ \frac{\partial}{\partial \alpha} \left[A - \frac{\omega}{Q} - B|\alpha|^2 \right] \alpha P + \text{c.c.} \right\} + A \frac{\partial^2 P}{\partial \alpha \partial \alpha^*} \quad (3.8)$$

That is, compared to formula (2.100), $-\frac{\omega}{Q}$ is replaced by $A - \frac{\omega}{Q} - B|\alpha|^2$, where $A - \frac{\omega}{Q} = G$ is the unsaturated gain minus losses, and B characterizes gain reduction due to saturation.

3.2 Theory of Laser Generation

Using equation (3.7), one can obtain the equation for the average number of photons in the mode

$$\langle n(t) \rangle = \sum_{(n)} n \rho_{nn}.$$

To do this, one needs to multiply equation (3.7) by n and sum over all n . We have

$$\frac{d}{dt} \langle n \rangle = \frac{d}{dt} \sum_n \rho_{nn} n = \sum_n \dot{\rho}_{nn} n.$$

Thus, from (3.7)

$$\begin{aligned} \frac{d}{dt} \langle n \rangle = & - \sum_n n [A - (n+1) B] (n+1) \rho_{nn} + \\ & + \sum_n n (A - nB) n \rho_{n-1, n-1} - \sum_n n \frac{\omega}{Q} n \rho_{nn} + \\ & + \sum_n n \frac{\omega}{Q} (n+1) \rho_{n+1, n+1}. \end{aligned} \quad (3.9)$$

We replace the summation indices in (3.9). In the second sum, $n-1 = m$, i.e., $n = m+1$. In the last sum, $n+1 = m$, i.e., $n = m-1$. Substituting this into (3.9) we get

$$\begin{aligned} \frac{d}{dt} \langle n \rangle = & - \sum_n n [A - (n+1) B] (n+1) \rho_{nn} + \\ & + \sum_m (m+1) (A - (m+1) B) (m+1) \rho_{m, m} - \sum_n n \frac{\omega}{Q} n \rho_{nn} + \\ & + \sum_m (m-1) \frac{\omega}{Q} m \rho_{m, m} = \sum_m (A - (m+1) B) (m+1) \rho_{m, m} - \sum_m \frac{\omega}{Q} m \rho_{m, m} = \\ & = \sum_n (A - (n+1) B) (n+1) \rho_{n, n} - \sum_n \frac{\omega}{Q} n \rho_{n, n}. \end{aligned}$$

Here we return the summation index m to n and combine the sums. From here we have

$$\frac{d}{dt} \langle n(t) \rangle = \left(A - \frac{\omega}{Q} \right) \langle n(t) \rangle + A - B \langle (n+1)^2 \rangle \quad (3.10)$$

Here the first term $\left(A - \frac{\omega}{Q} \right)$ corresponds to unsaturated gain minus losses. The term $B \langle (n+1)^2 \rangle = B [\langle n^2 \rangle + 2 \langle n \rangle + 1]$ characterizes the decrease in gain due to saturation. The term A takes into account spontaneous emission, which is absent in the semiclassical consideration. Otherwise, (3.10) resembles the classical laser equation [56]:

$$\dot{I}_n = 2I_n (\alpha_n - \beta_n I_n).$$

3.3 Statistics of Laser Photons

For equation (3.7), one can visually represent the probability flows shown in Figure 3.2.

The term $\frac{\omega}{Q} n \rho_{nn}$ represents the flow from the state $|n\rangle$ to the state $|n-1\rangle$ due to photon absorption by atoms of the first reservoir. The same can be said

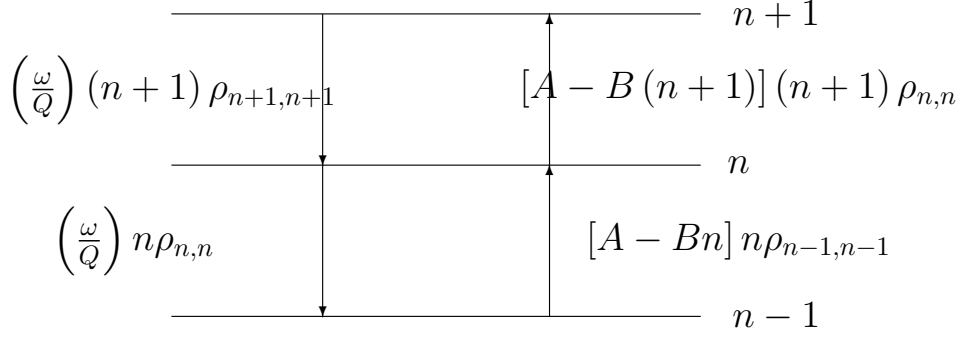


Figure 3.2: Balance of probability flows of photon numbers.

for the term $\frac{\omega}{Q}(n+1)\rho_{n+1,n+1}$ — the probability flow from the state $|n+1\rangle$ to the state $|n\rangle$ for the same reasons.

The term $[A - (n+1)B](n+1)\rho_{nn}$ represents the flow from the state $|n\rangle$ to the state $|n+1\rangle$ and characterizes the birth of photons due to interaction with active atoms, taking saturation effects into account. The same applies to the term $-(A - nB)n\rho_{n-1,n-1}$ — the probability flow from the state $|n-1\rangle$ to the state $|n\rangle$.

ρ_{nn} has the meaning of the probability of detecting n photons in the laser mode. In the transient regime, $\rho_{nn}(t)$ depends on time. In the steady-state regime, $\dot{\rho}_{nn} = 0$, and from equation Equation 3.7 one can establish the photon statistics in the steady generation regime [62].

Equating, for example, the probability flows between the states $|n+1\rangle$ and $|n\rangle$ (using the principle of detailed balance), we obtain:

$$\frac{\omega}{Q}(n+1)\rho_{n+1,n+1} = [A - (n+1)B](n+1)\rho_{nn} \quad (3.11)$$

Equality (3.11) gives an iterative relation

$$\rho_{n+1,n+1} = \frac{A - (n+1)B}{\omega/Q} \rho_{nn} \quad (3.12)$$

which allows expressing ρ_{nn} through ρ_{00} . We have

$$\rho_{nn} = \rho_{00} \prod_{k=1}^n \frac{A - kB}{\omega/Q} \quad (3.13)$$

ρ_{00} can be found from the normalization condition $\sum_{(n)} \rho_{nn} = 1$.

Let us qualitatively consider using (3.13) ρ_{nn} as a function of n , when the laser operates above threshold, i.e., when $A > \frac{\omega}{Q}$. In the steady-state regime, the

gain equals the losses (on average), that is

$$A - \bar{n}_{st}B = \frac{\omega}{Q},$$

whence the average number of photons in the steady-state regime equals

$$\bar{n}_{st} = \frac{A - \omega/Q}{B}.$$

For $n \ll \bar{n}_{st}$, $Bn \ll A$ the formula (3.13) can be represented as

$$\rho_{nn} \approx \left(\frac{AQ}{\omega} \right)^n.$$

This quantity grows exponentially with increasing n , since $\frac{AQ}{\omega} > 1$, meaning that for small n the probability increases with n , but the growth slows down as n increases. For large n the factor $\frac{A-nB}{\frac{\omega}{Q}}$ approaches unity when n approaches \bar{n}_{st} . At $n \approx \bar{n}_{st}$ the factor becomes equal to 1, and the distribution reaches a maximum:

$$\left. \frac{A - Bk}{\omega/Q} \right|_{k=\bar{n}_{st}} \rightarrow \frac{A - B\bar{n}_{st}}{\omega/Q} = \frac{A - A + \omega/Q}{\omega/Q} = 1.$$

For $n > \bar{n}_{st}$ and $k > \bar{n}_{st}$, the factors decrease, reaching zero at $k = \frac{A}{B}$. Note that for $k = \frac{A}{B}$ the factor becomes approximately zero, and terms with larger k should be discarded, since ρ_{nn} cannot be negative. The arising difficulty is related to the fact that we used perturbation theory and therefore assumed that the number of photons is not too large. Our theory is valid while $n < \frac{A}{B}$. These difficulties disappear if one uses the large signal theory developed in [62]. From the above, it follows that ρ_{nn} as a function of n (the photon distribution in the laser field) first increases, reaching a maximum near $n = \bar{n}_{st}$, and then decreases to zero.

In Figure 3.3 a qualitative curve of the photon distribution is shown for the case of threshold excess. At threshold, $A = \frac{\omega}{Q}$. Then the factor $\frac{A-Bk}{\omega/Q} = 1 - \left(\frac{B}{A}\right)k$ is less than unity for any k . The distribution curve will be monotonically decreasing. Below threshold, $A < \frac{\omega}{Q}$, then $\frac{A-Bk}{\omega/Q} \approx \frac{AQ}{\omega} \ll 1$ and ρ_{nn} decays exponentially, as $\left(\frac{AQ}{\omega}\right)^n$. The dependence of ρ_{nn} on n for all three cases can be numerically calculated using formula (3.13). The results of such calculation are shown in Figure 3.4.

From the large signal theory [62], which we do not consider here, it follows that for very large threshold excess, $\rho_{nn} \rightarrow e^{-\bar{n}} \frac{\bar{n}^n}{n!}$, i.e., tends to the distribution characteristic of a coherent state. For moderate threshold excess, the photon

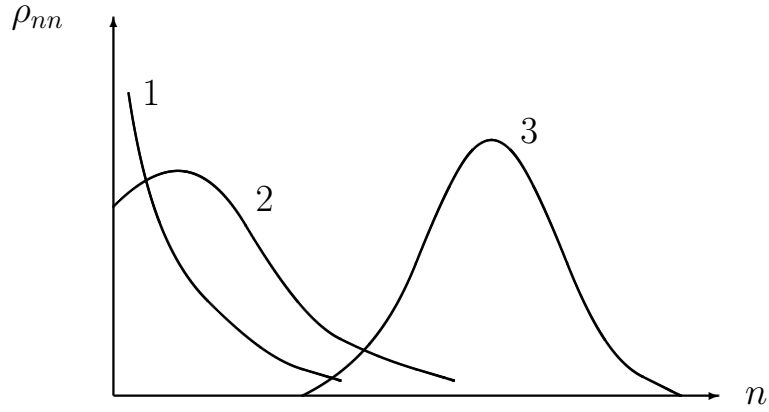


Figure 3.3: Qualitative distribution of photons for three cases. Graph 1: $A < \frac{\omega}{Q}$ - below the generation threshold. Graph 2: $A \approx \frac{\omega}{Q}$ - at the generation threshold. Graph 3: $A > \frac{\omega}{Q}$ - above the generation threshold.

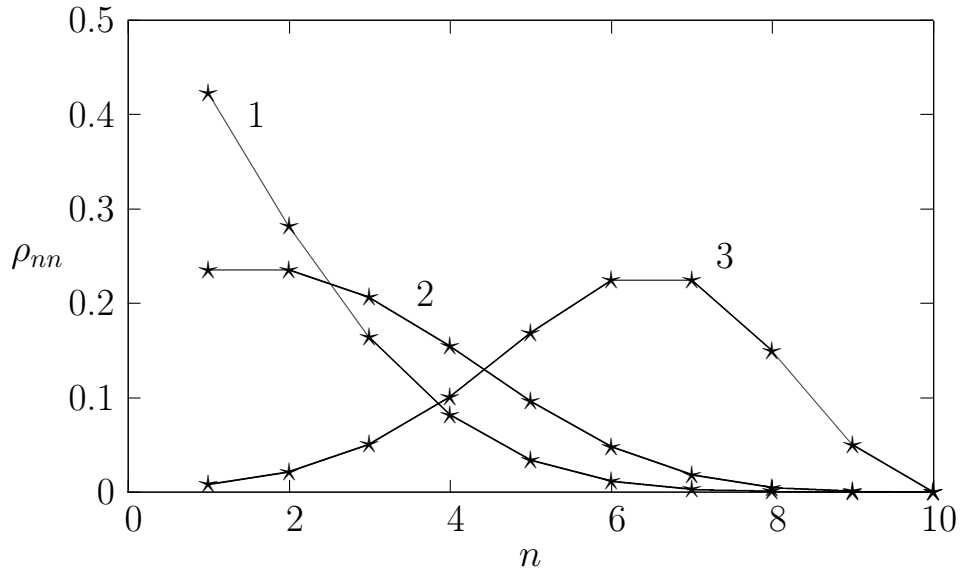


Figure 3.4: Results of the numerical calculation of the photon distribution. Graph 1, which describes the photon distribution before the generation threshold, was obtained with the following parameters: $A = 10$, $\frac{\omega}{Q} = 12$, $B = 1$. Graph 2, corresponding to the generation threshold, was obtained with $A = 10$, $\frac{\omega}{Q} = 8$, $B = 1$. For graph 3, $A = 10$, $\frac{\omega}{Q} = 3$, $B = 1$ was accepted, which corresponds to exceeding the threshold.

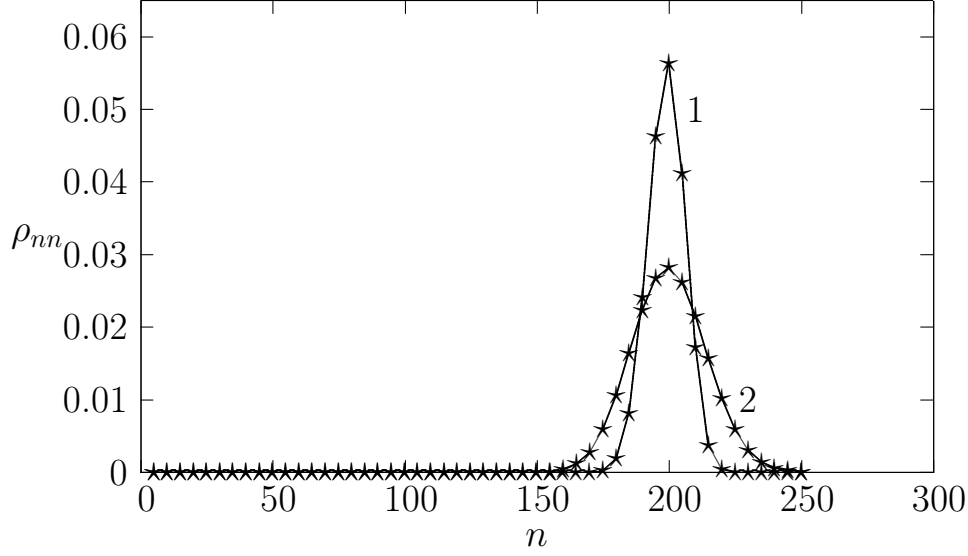


Figure 3.5: Difference in photon distribution in the laser field with moderate pumping (graph 1) from the Poisson distribution (graph 2).

distribution in laser radiation noticeably differs from the photon distribution in a coherent state. The difference between the photon distribution in the laser field and the Poisson distribution (corresponding to a coherent state) under moderate pumping is clearly shown in [Figure 3.5](#).

3.4 Laser theory. Representation of coherent states

It is necessary to use the equation for the quasiprobability function $P(\alpha, t)$ ([3.8](#)):

$$\frac{\partial}{\partial t} P(\alpha, t) = -\frac{1}{2} \left\{ \frac{\partial}{\partial \alpha} \left[A - \frac{\omega}{Q} - B|\alpha|^2 \right] \alpha P + \text{c.c.} \right\} + A \frac{\partial^2 P}{\partial \alpha \partial \alpha^*}.$$

For our purposes, it is convenient to represent this equation in polar coordinates $\alpha = r e^{i\theta}$. We have

$$\begin{aligned} \frac{\partial}{\partial t} P(r, \theta, t) = & -\frac{1}{2r} \cdot \frac{\partial}{\partial r} \left[r^2 \left(A - \frac{\omega}{Q} - B r^2 \right) P(r, \theta, t) \right] + \\ & + \frac{A}{4r^2} \left(r \frac{\partial}{\partial r} r \frac{\partial}{\partial r} + \frac{\partial^2}{\partial \theta^2} \right) P(r, \theta, t) \end{aligned} \quad (3.14)$$

Suppose we are interested in the steady state. Then $\frac{\partial P}{\partial t} = 0$, and the phase θ is uniformly distributed from 0 to 2π . It follows that $\frac{\partial^2 P}{\partial \theta^2} = 0$ and the quasiprob-

ability $P(r, \theta, t)$ does not depend on t or θ . Equation (3.14) then takes the form

$$0 = -\frac{\partial}{\partial r} \left[r^2 \left(A - \frac{\omega}{Q} - Br^2 \right) P(r) \right] + \frac{1}{2} A \frac{\partial}{\partial r} r \frac{\partial}{\partial r} P(r). \quad (3.15)$$

This equation is easily integrated. The first integration gives

$$r \left(A - \frac{\omega}{Q} - Br^2 \right) P(r) = \frac{1}{2} A \frac{\partial}{\partial r} P(r) + C. \quad (3.16)$$

The constant $C = 0$, since P and $\frac{\partial P}{\partial r}$ must tend sufficiently quickly to zero as $r \rightarrow \infty$.

From here follows the equation

$$\frac{dP}{P} = \frac{2}{A} r \left(A - \frac{\omega}{Q} - Br^2 \right) dr. \quad (3.17)$$

Its solution is

$$\begin{aligned} P &= N \exp \left(\frac{1}{A} \left[r^2 \left(A - \frac{\omega}{Q} \right) - \frac{Br^4}{2} \right] \right) = \\ &= N \exp \left(\frac{1}{A} \left[r^2 G - \frac{Br^4}{2} \right] \right), \end{aligned} \quad (3.18)$$

where N is the normalization factor, equal to

$$\frac{1}{N} = 2\pi \int_0^\infty \exp \left(\frac{1}{A} \left[r^2 G - \frac{Br^4}{2} \right] \right) r dr.$$

Distribution (3.18) is a function of $r^2 = |\alpha|^2 = \bar{n}$. When threshold is exceeded $A > \frac{\omega}{Q}$, i.e. when $G > 0$, $P(r)$ first increases with increasing r^2 , reaching a maximum at $r^2 = \frac{A - \omega/Q}{B} = \frac{G}{B}$, and then decreases.

This same result was obtained earlier. The obtained value $r = \sqrt{\frac{G}{B}}$ corresponds, at $G > 0$, to the most probable value of $|\alpha|$.

For $A < \frac{\omega}{Q}$, $G < 0$, the laser is below threshold. With increasing r^2 the curve monotonically decreases. At threshold $A = \frac{\omega}{Q}$, $G = 0$, the curve also decreases, but at $r^2 = 0$ we have an extremum.

On Figure 3.6 are shown dependencies of the quasiprobability on $r^2 = |\alpha|^2$, for different values of parameter G , corresponding to below-threshold, threshold, and above-threshold regimes.

On the complex plane $\alpha = r e^{i\theta}$, when the threshold is substantially exceeded, the region occupied by the generated field can be visually represented as shown

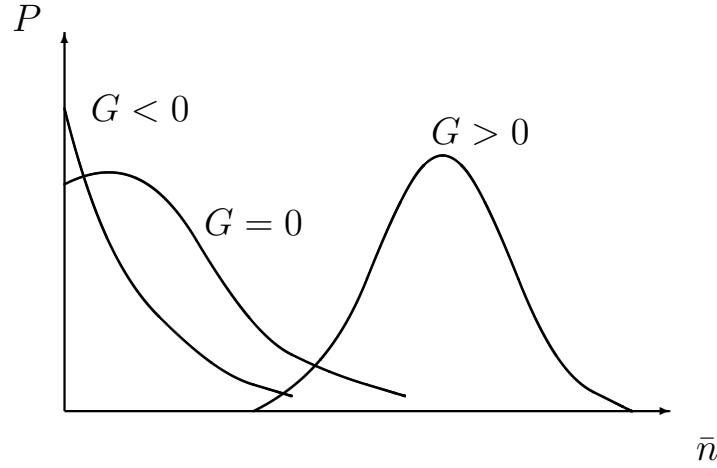


Figure 3.6: Qualitative view of the dependence of the quasiprobability P on the average number of photons $\bar{n} = r^2$ for three different values of the parameter G . $G < 0$ before the generation threshold; $G \approx 0$ at the generation threshold; $G > 0$ exceeding the generation threshold.

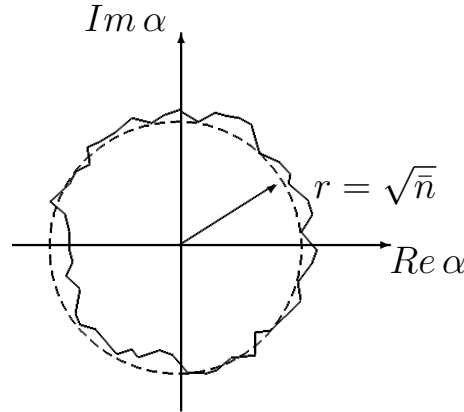


Figure 3.7: Fluctuation of the amplitude of the field generated by the laser in a narrow region near the circle of radius $r = \sqrt{\bar{n}_{st}}$.

on [Figure 3.7](#). The broken curve characterizes the distribution of the oscillation amplitude, i.e., corresponds to amplitude noise. The phase θ is uniformly distributed in the interval $0 - 2\pi$, therefore the broken curve lies close to the circle $r = \sqrt{\bar{n}_{st}}$ with the radius corresponding to the most probable value of $|\alpha|$.

From [Figure 3.7](#) it is seen that the amplitude fluctuates in a narrow region around the circle of radius $r = \sqrt{\bar{n}_{st}}$, while the phase freely diffuses (wanders) along this circle.

Phase diffusion in the laser occurs slowly, therefore, if we consider not too long time intervals, it can be assumed that the laser generates radiation with a well-defined phase. When averaged over a long time interval, during which the phase can assume any value in the range $0 - 2\pi$ with equal probability, the mean value of the laser field over this interval will be zero.

Let us consider the phase diffusion process more carefully. Suppose the amplitude distribution is established and $P(r)$ corresponds to the stationary amplitude distribution. Then

$$P(r, \theta, t) = P(r) P(\theta, t).$$

In this case, equation (3.14) takes the form

$$\frac{\partial P(\theta, t)}{\partial t} = \frac{A}{4r^2} \frac{\partial^2}{\partial \theta^2} P(\theta, t) \quad (3.19)$$

The function $P(r)$ cancels out because it satisfies the stationary equation (3.17).

Equation (3.19) can be used to obtain the equation satisfied by the average field. The average electric field of the laser mode, expressed through P , has the form

$$\langle \hat{E} \rangle = E_1 \int P(\alpha, t) \alpha d^2 \alpha.$$

In polar coordinates this looks like:

$$\langle \hat{E} \rangle = E_1 \int_0^\infty r dr \int_0^{2\pi} r e^{i\theta} P(r, \theta, t) d\theta.$$

Using now equation (3.19):

$$\begin{aligned} \frac{\partial \langle \hat{E} \rangle}{\partial t} &= E_1 \frac{\partial}{\partial t} \int_0^\infty r^2 dr \int_0^{2\pi} e^{i\theta} P(r, \theta, t) d\theta = \\ &= E_1 \int_0^\infty r^2 dr \int_0^{2\pi} e^{i\theta} \frac{\partial P}{\partial t} d\theta = \\ &= \frac{E_1 A}{\bar{n}_{st} 4} \int_0^\infty r^2 dr \int_0^{2\pi} e^{i\theta} \frac{\partial^2}{\partial \theta^2} P(r, \theta, t) d\theta. \end{aligned} \quad (3.20)$$

Here we approximately replaced $\frac{1}{r^2} \approx \frac{1}{\bar{n}_{st}}$, assuming the amplitude distribution is sufficiently narrow.

We integrate the inner integral twice by parts. We have

$$\begin{aligned} \int_0^{2\pi} e^{i\theta} \frac{\partial^2 P}{\partial \theta^2} d\theta &= e^{i\theta} \frac{\partial P}{\partial \theta} \Big|_0^{2\pi} - i \int_0^{2\pi} e^{i\theta} \frac{\partial P}{\partial \theta} d\theta = \\ &= -ie^{i\theta} P \Big|_0^{2\pi} - \int_0^{2\pi} e^{i\theta} P d\theta. \end{aligned} \quad (3.21)$$

The first terms in (3.21) equal zero due to periodicity of P and $\frac{\partial P}{\partial \theta}$. Using

(3.21), equation (3.20) can be written as

$$\begin{aligned} \frac{\partial \langle E \rangle}{\partial t} &= E_1 \int_0^\infty r^2 dr \int_0^{2\pi} e^{i\theta} \frac{\partial P}{\partial t} d\theta = \\ &= -\frac{AE_1}{4\bar{n}_{st}} \int_0^\infty r dr \int_0^{2\pi} r e^{i\theta} P d\theta = -\frac{A}{4\bar{n}_{st}} \langle E \rangle. \end{aligned} \quad (3.22)$$

The solution of this equation is

$$\langle E(t) \rangle = \langle E(0) \rangle e^{-\frac{D}{2}t}, \quad (3.23)$$

where

$$D = \frac{A}{2\bar{n}_{st}}. \quad (3.24)$$

Thus, we have obtained that the mean field indeed tends to zero with the characteristic time $\tau = \frac{1}{D}$. In lasers this interval can be quite long compared to the oscillation period.

To determine the natural linewidth of the laser emission it is necessary to find the correlation function of laser radiation; its spectrum by the Wiener-Khinchin theorem (Theorem 15.2.1) will be the energy spectrum of laser radiation. We have

$$\left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle = E_1^2 \int d^2\alpha P(\alpha) \alpha^*(0) \alpha(t) e^{-i\omega_0 t},$$

where $P(\alpha)$ is the quasiprobability of the laser state, defined by formula (3.14). When switching to polar coordinates (amplitude-phase) $\alpha = r e^{i\theta}$ we assume that the amplitude is already established and does not change, while the phase changes slowly, so that

$$\begin{aligned} \alpha^*(0) &= r e^{i\theta(0)} = r e^{i0} = r, \\ \alpha(t) &= r e^{i\theta(t)} = r e^{i\theta}, \end{aligned} \quad (3.25)$$

In (3.25) it is assumed that at the initial time $\theta(0) = 0$ and at time $t - \theta(t) = \theta$. Thus we have:

$$\left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle = E_1^2 e^{-i\omega_0 t} \int_0^\infty r^3 dr \int_0^{2\pi} d\theta P(r e^{i\theta}, t) e^{i\theta}.$$

Let us derive an equation satisfied by $\left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle$. From equation (3.19)

$\frac{\partial P(\theta, t)}{\partial t} = \frac{A}{4\bar{n}_{st}} \frac{\partial^2}{\partial \theta^2} P(\theta, t)$, therefore

$$\begin{aligned} \frac{d}{dt} \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle &= -i\omega_0 \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle + \\ &+ E_1^2 e^{-i\omega_0 t} \int_0^\infty r^3 dr \int_0^{2\pi} d\theta \frac{\partial}{\partial t} P(re^{i\theta}, t) e^{i\theta} = \\ &= -i\omega_0 \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle + \\ &+ \frac{E_1^2 A}{4\bar{n}_{st}} \int_0^\infty r^3 dr \int_0^{2\pi} d\theta \frac{\partial^2}{\partial \theta^2} P(re^{i\theta}, t) e^{i\theta}. \end{aligned} \quad (3.26)$$

Performing two integrations by parts by θ in (3.26), we obtain:

$$\begin{aligned} \int_0^{2\pi} d\theta \frac{\partial^2}{\partial \theta^2} P e^{i\theta} &= \frac{\partial}{\partial \theta} P e^{i\theta} \Big|_0^{2\pi} - i \int_0^{2\pi} d\theta \frac{\partial}{\partial \theta} P e^{i\theta} = \\ &= -i \int_0^{2\pi} d\theta \frac{\partial}{\partial \theta} P e^{i\theta} = -i P e^{i\theta} \Big|_0^{2\pi} - \int_0^{2\pi} d\theta P e^{i\theta} = - \int_0^{2\pi} d\theta P e^{i\theta}. \end{aligned}$$

Substituting the obtained result into the original expression (3.26), we get:

$$\begin{aligned} \frac{d}{dt} \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle &= -i\omega_0 \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle - \\ &- \frac{E_1^2 A}{4\bar{n}_{st}} \int_0^\infty r^3 dr \int_0^{2\pi} d\theta P(re^{i\theta}, t) e^{i\theta} = \\ &= -i\omega_0 \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle - \frac{A}{4\bar{n}_{st}} \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle. \end{aligned}$$

The solution of this equation is

$$r_+(t) = \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \right\rangle = \left\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(0) \right\rangle e^{-i\omega_0 t - \frac{D}{2}t}, \quad (3.27)$$

where

$$\frac{D}{2} = \frac{A}{4\bar{n}_{st}}.$$

Expression (3.27) defines the correlation function $r(t)$ in the domain $t \geq 0$. For $t \leq 0$, taking into account (15.5), the correlation function can be written as

$$r_-(t) = r_-(-|t|) = r_+^*(|t|). \quad (3.28)$$

Using expressions (3.27) and (3.28), one can write the following expression for the Fourier transform of the correlation function, which according to the Wiener-

Khinchin theorem (Theorem 15.2.1) defines the energy spectrum $S(\omega)$:

$$\begin{aligned}
S(\omega) &= \tilde{r}(\omega) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{i\omega t} r(t) dt = \\
&= \frac{1}{2\pi} \int_{-\infty}^0 e^{i\omega t} r_-(t) dt + \frac{1}{2\pi} \int_0^{+\infty} e^{i\omega t} r_+(t) dt = \\
&= \frac{1}{2\pi} \int_0^{+\infty} e^{-i\omega t} r_+^*(t) dt + \frac{1}{2\pi} \int_0^{+\infty} e^{i\omega t} r_+(t) dt = \\
&= \frac{1}{\pi} \operatorname{Re} \int_0^{+\infty} e^{i\omega t} r_+(t) dt = \\
&= \frac{1}{\pi} \operatorname{Re} \int_0^{+\infty} e^{i\omega t} \langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(t) \rangle dt = \\
&= \frac{\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(0) \rangle}{\pi} \operatorname{Re} \int_0^{+\infty} e^{(i(\omega - \omega_0) - \frac{D}{2})t} dt = \\
&= \frac{\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(0) \rangle}{\pi} \operatorname{Re} \frac{1}{\frac{D}{2} - i(\omega - \omega_0)} = \\
&= \frac{\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(0) \rangle}{\pi} \operatorname{Re} \frac{\frac{D}{2} + i(\omega - \omega_0)}{(\omega - \omega_0)^2 + (D/2)^2} = \\
&= \frac{\langle \hat{E}^{(-)}(0) \hat{E}^{(+)}(0) \rangle}{\pi} \frac{D/2}{(\omega - \omega_0)^2 + (D/2)^2}. \tag{3.29}
\end{aligned}$$

In the derivation of (3.29), the definition of the Fourier transform (18.1) from section 18.1 was used.

It follows that the generation line is Lorentzian, and the linewidth equals D (Figure 3.8). Since we considered only the "natural", fundamentally unavoidable cause of line broadening, associated with quantum fluctuations, and neglected fundamentally removable technical causes, the obtained linewidth is extremely narrow.

Here we considered the simplest laser model. A more realistic model of a three-level system is considered in [44]. Results obtained for various laser models are in good agreement with each other and correspond to the results obtained with the simple model used here.

3.5 Exercises

1. Derive (3.5) from (3.4).



Figure 3.8: To the calculation of the generation line width of the laser. Dependence of the spectral power density $S(\omega)$ on the frequency ω . The line width is determined by half the power, i.e., those values of ω are considered for which $\frac{S(\omega)}{S(\omega_0)} \geq \frac{1}{2}$. From (3.29), it follows that the interval boundaries correspond to frequencies $\omega = \omega_0 \pm \frac{D}{2}$. Thus, the line width is D .

2. Obtain the expression for the diagonal elements of the density matrix (3.7) from (3.6).
3. Obtain the expression for the density operator in the coherent state representation (3.8) from (3.6).
4. Represent equation (3.8) in polar coordinates.
5. Derive equation (3.19) from the general equation (3.14).
6. Using the formulas (3.24) and (3.29), estimate the "natural" linewidth of the laser generation for $A \approx \frac{\omega}{Q}$ (slightly above threshold), $\frac{\omega}{Q} = \Delta\omega = 10^6 \text{Hz}$, $n_{st} = 10^6 - 10^7$ photons in the mode.

Chapter 4

Quantum Theory of the Laser in the Heisenberg Representation

So far we have considered the quantum theory of the laser using the Schrödinger representation, in which the density matrix depends on time, while operators do not depend on time. As we have seen from the example of the problem of the decay of the resonator mode, another approach is possible, using the Heisenberg representation [52], in which operators depend on time, while the density matrix does not depend on time. In some cases, this approach can be more convenient for investigating subtle questions of the quantum theory of lasers, such as laser generation of the field in a squeezed state.

4.1 Laser Model

The scheme of laser levels is shown in [Figure 4.1](#). The level c is the main level, level b is the lower working level, a is the upper working level. In fact, the scheme is four-level. Pumping is carried out by incoherent light through a rather wide absorption line, depicted by a dashed line in the figure. But since a high rate of non-radiative transition to the upper working level is assumed, it can be considered that pumping occurs directly to level a . The presence of the fourth level allows a significant reduction of the reverse transition from level a to the main level c . [Figure 4.1](#) shows the transitions that will be considered in the theory. γ_a and γ_b characterize the relaxation of populations from levels a and b due to the connection with the dissipative system. r_a is the rate of pumping of the upper working level a due to incoherent optical pumping. The transition $a \rightarrow c$ is an induced transition caused by the laser field being generated.

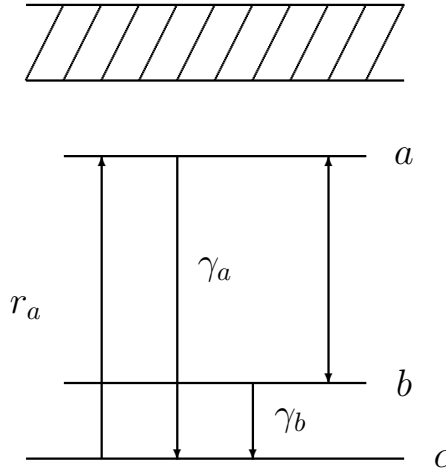


Figure 4.1: Diagram of laser levels. Level c - ground level, level b - lower working level, a - upper working level. γ_a and γ_b characterize relaxation of populations from levels a and b due to the connection with the dissipative system. r_a - pumping rate of the upper working level a due to incoherent optical pumping.

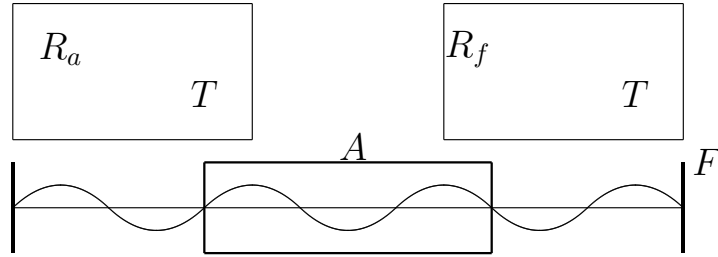


Figure 4.2: Laser model. The resonator, where the generated mode interacting with active atoms of the working medium A is excited, is denoted by F . Additionally, there are two reservoirs at temperature T : R_a , which is connected with the active atoms, and R_F , connected with the mode.

The laser scheme is presented in [Figure 4.2](#). The scheme contains a resonator F in which the generated mode is excited, interacting with the active atoms of the working medium, the scheme of which is presented in [Figure 4.1](#). In addition, there are two reservoirs at temperature T : R_a (associated with the active atoms) and R_F (associated with the mode), causing relaxation of the atoms and the mode field.

4.2 Heisenberg equation for the laser system

The considered system consists of 4 subsystems, which are connected to each other through 3 interaction Hamiltonians. The first subsystem represents the mode of the electromagnetic field with frequency ω (denoted by F in [Figure 4.2](#)),

which is described by creation and annihilation operators of the mode: \hat{a}^\dagger and \hat{a} .

The electromagnetic field mode is connected with the reservoir R_F (see Figure 4.2), which is a collection of harmonic oscillators with frequencies $\omega_k^{(F)}$, described by creation and annihilation operators $\hat{b}_k^\dagger, \hat{b}_k$. The following commutation relations hold:

$$\begin{aligned} [\hat{b}_k, \hat{b}_{k'}^\dagger] &= \delta_{kk'}, \\ [\hat{b}_k, \hat{b}_{k'}] &= 0, \\ [\hat{b}_k^\dagger, \hat{b}_{k'}^\dagger] &= 0. \end{aligned}$$

The atomic subsystem consists of a large number of atoms. The j -th atom can be in the following ground states: $|a\rangle^j, |b\rangle^j$, and $|c\rangle^j$. The behavior of this atom can be described using projection operators onto the states $|a\rangle^j$ and $|b\rangle^j$:

$$\hat{\sigma}_a^j = |a\rangle^j \langle a|^j, \quad \hat{\sigma}_b^j = |b\rangle^j \langle b|^j,$$

as well as transition operators

$$\begin{aligned} \hat{\sigma}_{a \rightarrow b}^j &= \hat{\sigma}_{ab}^j = |b\rangle^j \langle a|^j, \quad \hat{\sigma}_{b \rightarrow a}^j = \hat{\sigma}_{ba}^j = (\hat{\sigma}_{ab}^j)^\dagger, \\ \hat{\sigma}_{a \rightarrow c}^j &= \hat{\sigma}_{ac}^j = |c\rangle^j \langle a|^j, \quad \hat{\sigma}_{c \rightarrow a}^j = \hat{\sigma}_{ca}^j = (\hat{\sigma}_{ac}^j)^\dagger, \\ \hat{\sigma}_{b \rightarrow c}^j &= \hat{\sigma}_{bc}^j = |c\rangle^j \langle b|^j, \quad \hat{\sigma}_{c \rightarrow b}^j = \hat{\sigma}_{cb}^j = (\hat{\sigma}_{bc}^j)^\dagger. \end{aligned}$$

The atomic subsystem is connected to reservoir R_A , which consists of a large number of harmonic oscillators, each with frequency $\omega_k^{(A)}$, described by creation and annihilation operators $\hat{c}_k^\dagger, \hat{c}_k$ satisfying the following commutation relations:

$$\begin{aligned} [\hat{c}_k, \hat{c}_{k'}^\dagger] &= \delta_{kk'}, \\ [\hat{c}_k, \hat{c}_{k'}] &= 0, \\ [\hat{c}_k^\dagger, \hat{c}_{k'}^\dagger] &= 0. \end{aligned}$$

In the general case, the Heisenberg equation has the form:

$$\frac{d\hat{O}}{dt} = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{O}], \quad (4.1)$$

where $\hat{\mathcal{H}}$ is the Hamiltonian of the system, and \hat{O} is an operator of some observable related to this system. For simplicity, we will use the interaction picture, where

equation (4.1) is rewritten as:

$$\frac{d\hat{O}_I}{dt} = \frac{i}{\hbar} [\hat{V}_I, \hat{O}_I], \quad (4.2)$$

where the interaction operators \hat{V}_I and observable \hat{O}_I in the interaction picture are defined as:

$$\begin{aligned} \hat{O}_I &= e^{i\frac{\hat{\mathcal{H}}_0 t}{\hbar}} \hat{O} e^{-i\frac{\hat{\mathcal{H}}_0 t}{\hbar}}, \\ \hat{V}_I &= e^{i\frac{\hat{\mathcal{H}}_0 t}{\hbar}} \hat{V} e^{-i\frac{\hat{\mathcal{H}}_0 t}{\hbar}}. \end{aligned} \quad (4.3)$$

Here, $\hat{\mathcal{H}}_0$ denotes the full Hamiltonian of the considered system without including the interaction between its parts.

In what follows, all operators will be assumed to be written in the interaction picture (4.3), and thus the index I will be omitted.

The interaction Hamiltonian between the j -th atom and the mode of the electromagnetic field in the interaction picture is (see 2.14):

$$\hat{V}_{AF}^j = \Delta(t - t_j) g\hbar \left(\hat{a}^\dagger \hat{\sigma}_{ab}^j e^{i(\omega - \omega_{ab})t} + \left(\hat{\sigma}_{ab}^j \right)^\dagger \hat{a} e^{-i(\omega - \omega_{ab})t} \right), \quad (4.4)$$

where $\Delta(t - t_j)$ is a step function describing the process of atom injection, i.e. taking into account the fact that the interaction between the j -th atom and the field begins at time t_j :

$$\Delta(t - t_j) = \begin{cases} 1; & t \geq t_j \\ 0; & t < t_j \end{cases}$$

The full interaction Hamiltonian is the sum of Hamiltonians of individual atoms (4.4):

$$\hat{V}_{AF} = g\hbar \sum_j \Delta(t - t_j) \left(\hat{a}^\dagger \hat{\sigma}_{ab}^j + \left(\hat{\sigma}_{ab}^j \right)^\dagger \hat{a} \right). \quad (4.5)$$

In (4.5) we assume that the electromagnetic field mode frequency coincides with the transition frequency: $\omega = \omega_{ab}$.

The interaction Hamiltonian of the field with reservoir R_F is written as:

$$\hat{V}_{FR_F} = \hbar \sum_k g_k^{(F)} \left(\hat{b}_k^\dagger \hat{a} e^{i(\omega_k^{(F)} - \omega)t} + \hat{a}^\dagger \hat{b}_k e^{-i(\omega_k^{(F)} - \omega)t} \right). \quad (4.6)$$

The case is more complicated for the interaction Hamiltonian between the atomic subsystem and reservoir R_A . Based on the considered model (see Figure 4.2) we deal with two transitions: $a \rightarrow c$ and $b \rightarrow c$. Accordingly, the full Hamiltonian can be written as

$$\hat{V}_{AR_A} = \hat{V}_{AR_A}^{a \rightarrow c} + \hat{V}_{AR_A}^{b \rightarrow c}.$$

Assuming that the interaction constants for the considered transitions are the same and real, we get

$$\begin{aligned}\hat{V}_{AR_A}^{a \rightarrow c} &= \hbar \sum_{k,j} g_k^{(A)} \left(\hat{c}_k^\dagger \hat{\sigma}_{ac}^j e^{i(\omega_k^{(A)} - \omega_{ac})t} + (\hat{\sigma}_{ac}^j)^\dagger \hat{c}_k e^{-i(\omega_k^{(A)} - \omega_{ac})t} \right), \\ \hat{V}_{AR_A}^{b \rightarrow c} &= \hbar \sum_{k,j} g_k^{(A)} \left(\hat{c}_k^\dagger \hat{\sigma}_{bc}^j e^{i(\omega_k^{(A)} - \omega_{bc})t} + (\hat{\sigma}_{bc}^j)^\dagger \hat{c}_k e^{-i(\omega_k^{(A)} - \omega_{bc})t} \right).\end{aligned}\quad (4.7)$$

In (4.7) summation is assumed over all atoms (index j) and all reservoir modes (k), where ω_{ac} and ω_{bc} denote transition frequencies.

4.2.1 Equations of motion of the atomic subsystem

First, consider the equations satisfied by operators \hat{c}_k and \hat{c}_k^\dagger . Since these operators enter only in the Hamiltonian \hat{V}_{AR_A} , the equation of motion can be written as

$$\begin{aligned}\frac{d\hat{c}_k}{dt} &= \frac{i}{\hbar} \left[\hat{V}_{AR_A}, \hat{c}_k \right] = \\ &= \frac{i}{\hbar} \sum_{k',j} g_{k'}^{(A)} \left[\hat{c}_{k'}^\dagger, \hat{c}_k \right] \left(\hat{\sigma}_{ac}^j e^{i(\omega_{k'}^{(A)} - \omega_{ac})t} + \hat{\sigma}_{bc}^j e^{i(\omega_{k'}^{(A)} - \omega_{bc})t} \right) + \\ &+ \frac{i}{\hbar} \sum_{k',j} g_{k'}^{(A)} \left((\hat{\sigma}_{ac}^j)^\dagger e^{-i(\omega_{k'}^{(A)} - \omega_{ac})t} + (\hat{\sigma}_{bc}^j)^\dagger e^{-i(\omega_{k'}^{(A)} - \omega_{bc})t} \right) [\hat{c}_{k'}, \hat{c}_k] = \\ &= -i \sum_{k',j} g_{k'}^{(A)} \delta_{kk'} \left(\hat{\sigma}_{ac}^j e^{i(\omega_{k'}^{(A)} - \omega_{ac})t} + \hat{\sigma}_{bc}^j e^{i(\omega_{k'}^{(A)} - \omega_{bc})t} \right) = \\ &= -i g_k^{(A)} \sum_j \left(\hat{\sigma}_{ac}^j e^{i(\omega_k^{(A)} - \omega_{ac})t} + \hat{\sigma}_{bc}^j e^{i(\omega_k^{(A)} - \omega_{bc})t} \right).\end{aligned}\quad (4.8)$$

Formal integration of (4.8) gives

$$\begin{aligned}\hat{c}_k(t) &= \hat{c}_k(0) - \\ &- i g_k^{(A)} \sum_j \int_0^t dt' \left(\hat{\sigma}_{ac}^j(t') e^{i(\omega_k^{(A)} - \omega_{ac})t'} + \hat{\sigma}_{bc}^j(t') e^{i(\omega_k^{(A)} - \omega_{bc})t'} \right),\end{aligned}\quad (4.9)$$

from which for the Hermitian conjugate operator \hat{c}_k^\dagger we have

$$\begin{aligned}\hat{c}_k^\dagger(t) &= \hat{c}_k^\dagger(0) + \\ &+ i g_k^{(A)} \sum_j \int_0^t dt' \left(\hat{\sigma}_{ca}^j(t') e^{-i(\omega_k^{(A)} - \omega_{ac})t'} + \hat{\sigma}_{cb}^j(t') e^{-i(\omega_k^{(A)} - \omega_{bc})t'} \right),\end{aligned}\quad (4.10)$$

where we introduced the notation $\hat{\sigma}_{ca}^j = (\hat{\sigma}_{ac}^j)^\dagger$ and $\hat{\sigma}_{cb}^j = (\hat{\sigma}_{bc}^j)^\dagger$.

The next step is to derive the equations satisfied by operators $\hat{\sigma}_a^j$ and $\hat{\sigma}_b^j$. For $\hat{\sigma}_a^j$ we have

$$\frac{d\hat{\sigma}_a^j}{dt} = \frac{i}{\hbar} [\hat{V}_{AF}, \hat{\sigma}_a^j] + \frac{i}{\hbar} [\hat{V}_{AR_A}, \hat{\sigma}_a^j]. \quad (4.11)$$

For the first term of the sum (4.11) it holds

$$\frac{i}{\hbar} [\hat{V}_{AF}, \hat{\sigma}_a^j] = \frac{i}{\hbar} g \hbar \sum_i \Delta(t - t_i) \left(\hat{a}^\dagger [\hat{\sigma}_{ab}^i, \hat{\sigma}_a^j] + [(\hat{\sigma}_{ab}^i)^\dagger, \hat{\sigma}_a^j] \hat{a} \right),$$

from which, taking into account the commutation relations

$$\begin{aligned} [\hat{\sigma}_{ab}^i, \hat{\sigma}_a^j] &= |b^i\rangle \langle a^i| a^j \rangle \langle a^j| - |a^j\rangle \langle a^j| b^i \rangle \langle a^i| = \delta_{ij} \hat{\sigma}_{ab}^j, \\ [(\hat{\sigma}_{ab}^i)^\dagger, \hat{\sigma}_a^j] &= |a^i\rangle \langle b^i| a^j \rangle \langle a^j| - |a^j\rangle \langle a^j| a^i \rangle \langle b^i| = -\delta_{ij} (\hat{\sigma}_{ab}^j)^\dagger \end{aligned}$$

we get

$$\frac{i}{\hbar} [\hat{V}_{AF}, \hat{\sigma}_a^j] = ig \Delta(t - t_j) \left(\hat{a}^\dagger \hat{\sigma}_{ab}^j - (\hat{\sigma}_{ab}^j)^\dagger \hat{a} \right). \quad (4.12)$$

For the second term of the sum (4.11) we have

$$\begin{aligned} & \frac{i}{\hbar} [\hat{V}_{AR_A}, \hat{\sigma}_a^j] = \\ &= i \sum_{k,i} g_k^{(A)} \hat{c}_k^\dagger \left([\hat{\sigma}_{ac}^i, \hat{\sigma}_a^j] e^{i(\omega_k^{(A)} - \omega_{ac})t} + [\hat{\sigma}_{bc}^i, \hat{\sigma}_a^j] e^{i(\omega_k^{(A)} - \omega_{bc})t} \right) + \\ &+ i \sum_{k,i} g_k^{(A)} \left([(\hat{\sigma}_{ac}^i)^\dagger, \hat{\sigma}_a^j] e^{-i(\omega_k^{(A)} - \omega_{ac})t} + [(\hat{\sigma}_{bc}^i)^\dagger, \hat{\sigma}_a^j] e^{-i(\omega_k^{(A)} - \omega_{bc})t} \right) \hat{c}_k, \end{aligned}$$

which, using the commutation relations

$$\begin{aligned} [\hat{\sigma}_{ac}^i, \hat{\sigma}_a^j] &= \delta_{ij} \hat{\sigma}_{ac}^j, \\ [(\hat{\sigma}_{ac}^i)^\dagger, \hat{\sigma}_a^j] &= -\delta_{ij} (\hat{\sigma}_{ac}^j)^\dagger, \\ [\hat{\sigma}_{bc}^i, \hat{\sigma}_a^j] &= 0, \\ [(\hat{\sigma}_{bc}^i)^\dagger, \hat{\sigma}_a^j] &= 0 \end{aligned} \quad (4.13)$$

is rewritten as

$$\begin{aligned} \frac{i}{\hbar} [\hat{V}_{AR_A}, \hat{\sigma}_a^j] &= i \sum_k g_k^{(A)} \hat{c}_k^\dagger \hat{\sigma}_{ac}^j e^{i(\omega_k^{(A)} - \omega_{ac})t} - \\ &- i \sum_k g_k^{(A)} (\hat{\sigma}_{ac}^j)^\dagger \hat{c}_k e^{-i(\omega_k^{(A)} - \omega_{ac})t}. \end{aligned} \quad (4.14)$$

Combining (4.12) and (4.14) together we obtain

$$\frac{d\hat{\sigma}_a^j}{dt} = ig\Delta(t - t_j) \hat{a}^\dagger \hat{\sigma}_{ab}^j + i \sum_k g_k^{(A)} \hat{c}_k^\dagger \hat{\sigma}_{ac}^i e^{i(\omega_k^{(A)} - \omega_{ac})t} + \text{h.c.} \quad (4.15)$$

Using (4.10) we can compute the second term of (4.15):

$$\begin{aligned} i \sum_k g_k^{(A)} \hat{c}_k^\dagger \hat{\sigma}_{ac}^i e^{i(\omega_k^{(A)} - \omega_{ac})t} &= i \sum_k g_k^{(A)} \hat{c}_k^\dagger(0) \hat{\sigma}_{ac}^i e^{i(\omega_k^{(A)} - \omega_{ac})t} - \\ &\quad - \sum_{k,i} \left(g_k^{(A)}\right)^2 \int_0^t dt' \hat{\sigma}_{ca}^i(t') \hat{\sigma}_{ac}^j(t) e^{i(\omega_k^{(A)} - \omega_{ac})(t-t')} - \\ &\quad - \sum_{k,i} \left(g_k^{(A)}\right)^2 \int_0^t dt' \hat{\sigma}_{cb}^i(t') \hat{\sigma}_{ac}^j(t) e^{i(\omega_k^{(A)} - \omega_{bc})(t-t')} e^{i(\omega_{bc} - \omega_{ac})t}. \end{aligned}$$

Denoting

$$\hat{f}_a^j(t) = i \sum_k g_k^{(A)} \hat{c}_k^\dagger(0) \hat{\sigma}_{ac}^i e^{i(\omega_k^{(A)} - \omega_{ac})t} \quad (4.16)$$

we get

$$\begin{aligned} i \sum_k g_k^{(A)} \hat{c}_k^\dagger \hat{\sigma}_{ac}^i e^{i(\omega_k^{(A)} - \omega_{ac})t} &= \hat{f}_a^j - \\ &\quad - \sum_i \int_0^t dt' \sum_k \left(g_k^{(A)}\right)^2 \hat{\sigma}_{ca}^i(t') \hat{\sigma}_{ac}^j(t) e^{i(\omega_k^{(A)} - \omega_{ac})(t-t')} - \\ &\quad - D(\omega_{bc}) \left(g^{(A)}(\omega_{bc})\right)^2 e^{i(\omega_{bc} - \omega_{ac})t} \cdot \\ &\quad \cdot \sum_i \int_0^t dt' \int_0^\infty d\omega' \hat{\sigma}_{cb}^i(t') \hat{\sigma}_{ac}^j(t) e^{i(\omega' - \omega_{bc})(t-t')} \approx \\ &\quad \approx \hat{f}_a^j - \frac{\gamma_a}{2} \sum_i \int_0^t dt' \hat{\sigma}_{ca}^i(t') \hat{\sigma}_{ac}^j(t) \delta(t - t') - \\ &\quad - \frac{\gamma_b}{2} e^{i(\omega_{bc} - \omega_{ac})t} \sum_i \int_0^t dt' \hat{\sigma}_{cb}^i(t') \hat{\sigma}_{ac}^j(t) \delta(t - t') = \\ &\quad = \hat{f}_a^j - \frac{\gamma_a}{2} \hat{\sigma}_a^j(t) - \frac{\gamma_b}{2} e^{i(\omega_{bc} - \omega_{ac})t} \hat{\sigma}_{ba}^j(t) \approx \\ &\quad \approx \hat{f}_a^j - \frac{\gamma_a}{2} \hat{\sigma}_a^j(t), \end{aligned} \quad (4.17)$$

where

$$\gamma_a = 4\pi D(\omega_{ac}) \left(g^{(A)}(\omega_{ac})\right)^2$$

and

$$\gamma_b = 4\pi D (\omega_{bc}) \left(g^{(A)} (\omega_{bc}) \right)^2.$$

In deriving (4.17) we also neglected rapidly oscillating term $\frac{\gamma_b}{2} e^{i(\omega_{bc}-\omega_{ac})t} \hat{\sigma}_{ba}^j(t)$ compared to $\frac{\gamma_a}{2} \hat{\sigma}_a^j(t)$. Thus, as a result we get

$$\frac{d\hat{\sigma}_a^j}{dt} = -\gamma_a \hat{\sigma}_a^j + ig\Delta(t-t_j) \left(\hat{a}^\dagger \hat{\sigma}_{ab}^j - \left(\hat{\sigma}_{ab}^j \right)^\dagger \hat{a} \right) + \hat{f}_a^j. \quad (4.18)$$

By analogy, one can obtain the expression for $\hat{\sigma}_b^j$:

$$\frac{d\hat{\sigma}_b^j}{dt} = -\gamma_b \hat{\sigma}_b^j - ig\Delta(t-t_j) \left(\hat{a}^\dagger \hat{\sigma}_{ab}^j - \left(\hat{\sigma}_{ab}^j \right)^\dagger \hat{a} \right) + \hat{f}_b^j, \quad (4.19)$$

where

$$\hat{f}_b^j(t) = i \sum_k g_k^{(A)} \hat{c}_k^\dagger(0) \hat{\sigma}_{bc}^i e^{i(\omega_k^{(A)} - \omega_{bc})t} + \text{h.c.} \quad (4.20)$$

For $\hat{\sigma}_{ab}^j$ we have

$$\frac{d\hat{\sigma}_{ab}^j}{dt} = \frac{i}{\hbar} [\hat{V}_{AF}, \hat{\sigma}_{ab}^j] + \frac{i}{\hbar} [\hat{V}_{AR_A}, \hat{\sigma}_{ab}^j]. \quad (4.21)$$

For the first term of the sum (4.21) it holds

$$\frac{i}{\hbar} [\hat{V}_{AF}, \hat{\sigma}_{ab}^j] = \frac{i}{\hbar} g\hbar \sum_i \Delta(t-t_i) \left(\hat{a}^\dagger [\hat{\sigma}_{ab}^i, \hat{\sigma}_{ab}^j] + [(\hat{\sigma}_{ab}^i)^\dagger, \hat{\sigma}_{ab}^j] \hat{a} \right),$$

from which, taking into account

$$[(\hat{\sigma}_{ab}^i)^\dagger, \hat{\sigma}_{ab}^j] = |a^i\rangle \langle b^i| b^j \rangle \langle a^j| - |b^j\rangle \langle a^j| a^i \rangle \langle b^i| = -\delta_{ij} (\hat{\sigma}_a^j - \hat{\sigma}_b^j)$$

we obtain

$$\frac{i}{\hbar} [\hat{V}_{AF}, \hat{\sigma}_{ab}^j] = ig\Delta(t-t_j) (\hat{\sigma}_a^j - \hat{\sigma}_b^j) \hat{a}. \quad (4.22)$$

For the second term of the sum (4.21) we have

$$\begin{aligned} & \frac{i}{\hbar} [\hat{V}_{AR_A}, \hat{\sigma}_{ab}^j] = \\ &= i \sum_{k,i} g_k^{(A)} \hat{c}_k^\dagger \left([\hat{\sigma}_{ac}^i, \hat{\sigma}_{ab}^j] e^{i(\omega_k^{(A)} - \omega_{ac})t} + [\hat{\sigma}_{bc}^i, \hat{\sigma}_{ab}^j] e^{i(\omega_k^{(A)} - \omega_{bc})t} \right) + \\ &+ i \sum_{k,i} g_k^{(A)} \left([(\hat{\sigma}_{ac}^i)^\dagger, \hat{\sigma}_{ab}^j] e^{-i(\omega_k^{(A)} - \omega_{ac})t} + [(\hat{\sigma}_{bc}^i)^\dagger, \hat{\sigma}_{ab}^j] e^{-i(\omega_k^{(A)} - \omega_{bc})t} \right) \hat{c}_k, \end{aligned}$$

which, using commutation relations

$$\begin{aligned}
\left[\hat{\sigma}_{ac}^i, \hat{\sigma}_{ab}^j\right] &= 0, \\
\left[\left(\hat{\sigma}_{ac}^i\right)^\dagger, \hat{\sigma}_{ab}^j\right] &= -\delta_{ij} \left(\hat{\sigma}_{bc}^j\right)^\dagger, \\
\left[\hat{\sigma}_{bc}^i, \hat{\sigma}_{ab}^j\right] &= \delta_{ij} \hat{\sigma}_{ac}^j, \\
\left[\left(\hat{\sigma}_{bc}^i\right)^\dagger, \hat{\sigma}_{ab}^j\right] &= 0
\end{aligned} \tag{4.23}$$

is rewritten as

$$\begin{aligned}
\frac{i}{\hbar} \left[\hat{V}_{AR_A}, \hat{\sigma}_{ab}^j\right] &= i \sum_k g_k^{(A)} \hat{c}_k^\dagger \hat{\sigma}_{ac}^j e^{i(\omega_k^{(A)} - \omega_{bc})t} - \\
&\quad - i \sum_k g_k^{(A)} \left(\hat{\sigma}_{bc}^j\right)^\dagger \hat{c}_k e^{-i(\omega_k^{(A)} - \omega_{ac})t},
\end{aligned} \tag{4.24}$$

Combining (4.22) and (4.24) together we get

$$\begin{aligned}
\frac{d\hat{\sigma}_{ab}^j}{dt} &= ig\Delta(t - t_j) \left(\hat{\sigma}_a^j - \hat{\sigma}_b^j\right) \hat{a} + \\
&+ i \sum_k g_k^{(A)} \hat{c}_k^\dagger \hat{\sigma}_{ac}^j e^{i(\omega_k^{(A)} - \omega_{bc})t} - i \sum_k g_k^{(A)} \left(\hat{\sigma}_{bc}^j\right)^\dagger \hat{c}_k e^{-i(\omega_k^{(A)} - \omega_{ac})t}
\end{aligned} \tag{4.25}$$

Substituting into (4.25) the expressions for \hat{c}_k (4.9) and \hat{c}_k^\dagger (4.10) and neglecting rapidly oscillating terms, we have:

$$\begin{aligned}
\frac{d\hat{\sigma}_{ab}^j}{dt} &= ig\Delta(t - t_j) \left(\hat{\sigma}_a^j - \hat{\sigma}_b^j\right) \hat{a} - \\
&- \frac{\gamma_b}{2} \int_0^\infty dt' \hat{\sigma}_{cb}^j(t') \hat{\sigma}_{ac}^j(t) \delta(t - t') - \\
&- \frac{\gamma_a}{2} \int_0^\infty dt' \hat{\sigma}_{cb}^j(t) \hat{\sigma}_{ac}^j(t') \delta(t - t') + \hat{f}_{ab}^j,
\end{aligned} \tag{4.26}$$

where

$$\begin{aligned}
\hat{f}_{ab}^j(t) &= i \sum_k g_k^{(A)} \hat{c}_k^\dagger(0) \hat{\sigma}_{ac}^j e^{i(\omega_k^{(A)} - \omega_{bc})t} - \\
&- i \sum_k g_k^{(A)} \left(\hat{\sigma}_{bc}^j\right)^\dagger \hat{c}_k(0) e^{-i(\omega_k^{(A)} - \omega_{ac})t}.
\end{aligned} \tag{4.27}$$

In what follows, we will use the assumption that

$$\gamma_a = \gamma_b = \gamma. \tag{4.28}$$

Taking into account

$$\hat{\sigma}_{cb}^j \hat{\sigma}_{ac}^j = \hat{\sigma}_{ab}^j,$$

as a result from (4.26) we get

$$\begin{aligned} \frac{d\hat{\sigma}_{ab}^j}{dt} &= ig\Delta(t-t_j) \left(\hat{\sigma}_a^j - \hat{\sigma}_b^j \right) \hat{a} - \gamma \hat{\sigma}_{cb}^j \hat{\sigma}_{ac}^j + \hat{f}_{ab}^j = \\ &= ig\Delta(t-t_j) \left(\hat{\sigma}_a^j - \hat{\sigma}_b^j \right) \hat{a} - \gamma \hat{\sigma}_{ab}^j + \hat{f}_{ab}^j \end{aligned} \quad (4.29)$$

For the noise operator \hat{f}_{ab}^j one can write the following correlation functions:

$$\begin{aligned} \left\langle \left(\hat{f}_{ab}^i(t_1) \right)^\dagger \hat{f}_{ab}^j(t_2) \right\rangle &= \delta_{ij} \frac{\gamma}{2} \langle \hat{\sigma}_a^j \rangle \delta(t_1 - t_2), \\ \left\langle \hat{f}_{ab}^j(t_1) \left(\hat{f}_{ab}^i(t_2) \right)^\dagger \right\rangle &= \delta_{ij} \frac{\gamma}{2} \langle \hat{\sigma}_b^j \rangle \delta(t_1 - t_2). \end{aligned} \quad (4.30)$$

Indeed, for the first relation (4.30), assuming the average phonon number in the reservoir mode at the initial time is zero:

$$\left\langle \hat{c}_{k_1}^\dagger(0) \hat{c}_{k_2}(0) \right\rangle = 0,$$

so that

$$\left\langle \hat{c}_{k_1}(0) \hat{c}_{k_2}^\dagger(0) \right\rangle = \delta_{k_1, k_2},$$

we obtain

$$\begin{aligned} &\left\langle \left(\hat{f}_{ab}^i(t_1) \right)^\dagger \hat{f}_{ab}^j(t_2) \right\rangle = \\ &= \sum_{k_1, k_2} g_{k_1}^{(A)} g_{k_2}^{(A)} \left\langle \hat{c}_{k_1}(0) \hat{c}_{k_2}^\dagger(0) \right\rangle \left\langle \left(\hat{\sigma}_{ac}^i \right)^\dagger \hat{\sigma}_{ac}^j \right\rangle e^{-i(\omega_{k_1}^{(A)} - \omega_{bc})t_1} e^{i(\omega_{k_2}^{(A)} - \omega_{bc})t_2} + \\ &+ \sum_{k_1, k_2} g_{k_1}^{(A)} g_{k_2}^{(A)} \left\langle \hat{c}_{k_1}^\dagger(0) \hat{c}_{k_2}(0) \right\rangle \left\langle \hat{\sigma}_{bc}^i \left(\hat{\sigma}_{bc}^j \right)^\dagger \right\rangle e^{i(\omega_{k_1}^{(A)} - \omega_{ac})t_1} e^{-i(\omega_{k_2}^{(A)} - \omega_{ac})t_2} = \\ &= \sum_k \left(g_k^{(A)} \right)^2 \left\langle \left(\hat{\sigma}_{ac}^i \right)^\dagger \hat{\sigma}_{ac}^j \right\rangle e^{-i(\omega_k^{(A)} - \omega_{bc})(t_1 - t_2)} = \\ &= \delta_{ij} \sum_k \left(g_k^{(A)} \right)^2 \langle \hat{\sigma}_a^j \rangle e^{-i(\omega_k^{(A)} - \omega_{bc})(t_1 - t_2)} = \delta_{ij} \frac{\gamma}{2} \langle \hat{\sigma}_a^j \rangle \delta(t_1 - t_2). \end{aligned}$$

4.2.2 Equations of motion for macroscopic variables

From the equations of motion describing the behavior of individual atoms (4.18, 4.19, 4.29), we proceed to equations describing the behavior of the entire

atomic subsystem collectively. For this, from operators $\hat{\sigma}_a^j$, $\hat{\sigma}_b^j$, and $\hat{\sigma}_{ab}^j$ we go to new operators which represent the collective effect of all atoms:

$$\begin{aligned}\hat{N}_a &= \sum_j \Delta(t - t_j) \hat{\sigma}_a^j, \\ \hat{N}_b &= \sum_j \Delta(t - t_j) \hat{\sigma}_b^j, \\ \hat{N}_{ab} &= \sum_j \Delta(t - t_j) \hat{\sigma}_{ab}^j.\end{aligned}\tag{4.31}$$

Next, consider $\frac{d\hat{N}_a}{dt}$:

$$\begin{aligned}\frac{d\hat{N}_a}{dt} &= \sum_j \frac{d\Delta(t - t_j)}{dt} \hat{\sigma}_a^j + \sum_j \Delta(t - t_j) \frac{d\hat{\sigma}_a^j}{dt} = \\ &= \sum_j \delta(t - t_j) \hat{\sigma}_a^j + \sum_j \Delta(t - t_j) \frac{d\hat{\sigma}_a^j}{dt}.\end{aligned}\tag{4.32}$$

Substituting (4.18) into (4.32) and using (4.28) we obtain:

$$\begin{aligned}\frac{d\hat{N}_a}{dt} &= \sum_j \delta(t - t_j) \hat{\sigma}_a^j - \gamma \sum_j \Delta(t - t_j) \hat{\sigma}_a^j + \\ &+ ig \sum_j \Delta(t - t_j) \left(\hat{a}^\dagger \hat{\sigma}_{ab}^j - \left(\hat{\sigma}_{ab}^j \right)^\dagger \hat{a} \right) + \sum_j \Delta(t - t_j) \hat{f}_a^j = \\ &= \sum_j \delta(t - t_j) \hat{\sigma}_a^j - \gamma \hat{N}_a + \\ &+ ig \left(\hat{a}^\dagger \hat{N}_{ab} - \hat{N}_{ab}^\dagger \hat{a} \right) + \sum_j \Delta(t - t_j) \hat{f}_a^j\end{aligned}\tag{4.33}$$

Next, consider the average value of the first term in (4.33):

$$\begin{aligned}\left\langle \sum_j \delta(t - t_j) \hat{\sigma}_a^j(t) \right\rangle &= \left\langle \sum_j \delta(t - t_j) \hat{\sigma}_a^j(t_j) \right\rangle = \\ &= \left\langle \sum_j \delta(t - t_j) \langle \hat{\sigma}_a^j(t_j) \rangle \right\rangle_T,\end{aligned}\tag{4.34}$$

where $\langle \dots \rangle_T$ means averaging over injection times of atoms:

$$\langle \dots \rangle_T = r_a \int_{-\infty}^{\infty} (\dots) dt_j,\tag{4.35}$$

where r_a is the atom injection rate. For the average value of operator $\hat{\sigma}_a^j$ assuming that at time t_j , the j -th atom is injected in the excited state, we have

$$\langle \hat{\sigma}_a^j(t_j) \rangle = \langle a | \hat{\sigma}_a^j | a \rangle = 1. \quad (4.36)$$

Substituting (4.35) and (4.36) into (4.34) we get

$$\left\langle \sum_j \delta(t - t_j) \hat{\sigma}_a^j(t) \right\rangle = r_a. \quad (4.37)$$

Next, add and subtract the average (4.37) from (4.33), resulting in:

$$\frac{d\hat{N}_a}{dt} = r_a - \gamma\hat{N}_a + ig \left(\hat{a}^\dagger \hat{N}_{ab} - \hat{N}_{ab}^\dagger \hat{a} \right) + \hat{F}_a, \quad (4.38)$$

where a new noise operator \hat{F}_a is introduced, which accounts for not only quantum fluctuations but also fluctuations of injection times of active atoms:

$$\hat{F}_a = \sum_j \delta(t - t_j) \hat{\sigma}_a^j - r_a + \sum_j \Delta(t - t_j) \hat{f}_a^j, \quad (4.39)$$

and it is obvious that the average of this operator is zero:

$$\langle \hat{F}_a \rangle = \left\langle \sum_j \delta(t - t_j) \hat{\sigma}_a^j \right\rangle - r_a + \left\langle \sum_j \Delta(t - t_j) \hat{f}_a^j \right\rangle = 0.$$

From (4.19, 4.29) one can get the following relations for \hat{N}_b and \hat{N}_{ab} :

$$\begin{aligned} \frac{d\hat{N}_b}{dt} &= -\gamma\hat{N}_b - ig \left(\hat{a}^\dagger \hat{N}_{ab} - \hat{N}_{ab}^\dagger \hat{a} \right) + \hat{F}_b, \\ \frac{d\hat{N}_{ab}}{dt} &= -\gamma\hat{N}_{ab} + ig \left(\hat{N}_a - \hat{N}_b \right) \hat{a} + \hat{F}_{ab}, \end{aligned} \quad (4.40)$$

where the following noise operators are introduced:

$$\begin{aligned} \hat{F}_b &= \sum_j \Delta(t - t_j) \hat{f}_b^j + \sum_j \delta(t - t_j) \hat{\sigma}_b^j, \\ \hat{F}_{ab} &= \sum_j \Delta(t - t_j) \hat{f}_{ab}^j + \sum_j \delta(t - t_j) \hat{\sigma}_{ab}^j. \end{aligned} \quad (4.41)$$

Considering that

$$\langle \hat{\sigma}_b^j(t_j) \rangle = \langle \hat{\sigma}_{ab}^j(t_j) \rangle = 0,$$

these operators have zero average.

For the correlator $\langle \hat{F}_{ab}^\dagger(t_1) \hat{F}_{ab}(t_2) \rangle$ from (4.30) we obtain the following relation

$$\begin{aligned} \langle \hat{F}_{ab}^\dagger(t_1) \hat{F}_{ab}(t_2) \rangle &= \sum_{ij} \Delta(t_1 - t_i) \Delta(t_2 - t_j) \left\langle \left(\hat{f}_{ab}^i \right)^\dagger \hat{f}_{ab}^j \right\rangle + \\ &\quad + \sum_{ij} \delta(t_1 - t_i) \delta(t_2 - t_j) \left\langle \left(\hat{\sigma}_{ab}^i \right)^\dagger \hat{\sigma}_{ab}^j \right\rangle = \\ &= \frac{\gamma}{2} \delta(t_1 - t_2) \sum_j \Delta(t_1 - t_j) \langle \hat{\sigma}_a^j \rangle + \delta(t_1 - t_2) \sum_j \delta(t_1 - t_j) \langle \hat{\sigma}_a^j \rangle = \\ &= \left(\frac{\gamma}{2} \bar{N}_a + r_a \right) \delta(t_1 - t_2), \quad (4.42) \end{aligned}$$

where \bar{N}_a is the average number of atoms in the excited state at the considered time t_1

$$\bar{N}_a(t_1) = \sum_j \delta(t_1 - t_j) \langle \hat{\sigma}_a^j \rangle$$

Similarly, for the correlator $\langle \hat{F}_{ab}(t_1) \hat{F}_{ab}^\dagger(t_2) \rangle$ from (4.30) we obtain the following relation

$$\begin{aligned} \langle \hat{F}_{ab}(t_1) \hat{F}_{ab}^\dagger(t_2) \rangle &= \sum_{ij} \Delta(t_1 - t_i) \Delta(t_2 - t_j) \left\langle \hat{f}_{ab}^i \left(\hat{f}_{ab}^j \right)^\dagger \right\rangle + \\ &\quad + \sum_{ij} \delta(t_1 - t_i) \delta(t_2 - t_j) \left\langle \hat{\sigma}_{ab}^j \left(\hat{\sigma}_{ab}^i \right)^\dagger \right\rangle = \\ &= \frac{\gamma}{2} \delta(t_1 - t_2) \sum_j \Delta(t_1 - t_j) \langle \hat{\sigma}_b^j \rangle + \delta(t_1 - t_2) \sum_j \delta(t_1 - t_j) \langle \hat{\sigma}_b^j \rangle = \\ &= \frac{\gamma}{2} \bar{N}_b \delta(t_1 - t_2), \quad (4.43) \end{aligned}$$

where again the average number of atoms in the unexcited state is introduced:

$$\bar{N}_b(t_1) = \sum_j \delta(t_1 - t_j) \langle \hat{\sigma}_b^j \rangle$$

It will be convenient to use the operator of population difference $\hat{N}_z = \hat{N}_a - \hat{N}_b$, for which from (4.38) and (4.40) we get:

$$\frac{d\hat{N}_z}{dt} = r_a - \gamma \hat{N}_z + 2ig \left(\hat{a}^\dagger \hat{N}_{ab} - \hat{N}_{ab}^\dagger \hat{a} \right) + \hat{F}_z, \quad (4.44)$$

where

$$\hat{F}_z = \hat{F}_a - \hat{F}_b. \quad (4.45)$$

4.2.3 Equations of motion for the electromagnetic field

Similarly to the atomic subsystem, we start by considering the equations satisfied by reservoir R_A operators: \hat{b}_k . Since these operators enter only in the Hamiltonian \hat{V}_{FR_F} , the equation of motion takes the form

$$\frac{d\hat{b}_k}{dt} = \frac{i}{\hbar} [\hat{V}_{FR_F}, \hat{b}_k]. \quad (4.46)$$

Substituting into (4.46) the expression for the Hamiltonian (4.6) we get

$$\begin{aligned} \frac{d\hat{b}_k}{dt} &= \frac{i}{\hbar} [\hat{V}_{FR_F}, \hat{b}_k] = \\ &= \frac{i}{\hbar} \sum_{k',j} g_{k'}^{(F)} [\hat{b}_{k'}^\dagger, \hat{b}_{k'}] \hat{a} e^{i(\omega_{k'}^{(F)} - \omega)t} = -ig_k^{(F)} \hat{a} e^{i(\omega_k^{(F)} - \omega)t}. \end{aligned} \quad (4.47)$$

Formal integration of (4.47) gives

$$\hat{b}_k(t) = \hat{b}_k(0) - ig_k^{(F)} \int_0^t dt' \hat{a}(t') e^{i(\omega_k^{(F)} - \omega)t'}. \quad (4.48)$$

The operator of the electromagnetic field mode \hat{a} appears in the Hamiltonians \hat{V}_{AF} and \hat{V}_{FR_F} ; thus, taking into account the expressions for these Hamiltonians (4.5) and (4.6) and the definition (4.31), we have

$$\begin{aligned} \frac{d\hat{a}}{dt} &= \frac{i}{\hbar} [\hat{V}_{AF}, \hat{a}] + \frac{i}{\hbar} [\hat{V}_{FR_F}, \hat{a}] = \\ &= \frac{i}{\hbar} \hbar g \sum_j \Delta(t - t_j) [\hat{a}^\dagger, \hat{a}] \hat{\sigma}_{ab}^j + \\ &+ \frac{i}{\hbar} \hbar \sum_k g_k^{(F)} [\hat{a}^\dagger, \hat{a}] \hat{b}_k e^{-i(\omega_k^{(F)} - \omega)t} = \\ &= -ig \sum_j \Delta(t - t_j) \hat{\sigma}_{ab}^j - i \sum_k g_k^{(F)} \hat{b}_k e^{-i(\omega_k^{(F)} - \omega)t} = \\ &= -ig \hat{N}_{ab} - i \sum_k g_k^{(F)} \hat{b}_k e^{-i(\omega_k^{(F)} - \omega)t}. \end{aligned} \quad (4.49)$$

Substituting (4.48) into (4.49) we get

$$\begin{aligned} \frac{d\hat{a}}{dt} &= -ig \hat{N}_{ab} - 2\pi D(\omega) \left(g^{(F)}(\omega) \right)^2 \int_0^t dt' \hat{a}(t') \delta(t - t') - \\ &- i \sum_k g_k^{(F)} \hat{b}_k(0) e^{-i(\omega_k^{(F)} - \omega)t}. \end{aligned}$$

Denoting

$$\gamma^{(F)} = 4\pi D(\omega) \left(g^{(F)}(\omega) \right)^2 = \frac{\omega}{Q}$$

and introducing the noise operator

$$\hat{F}_F(t) = -i \sum_k g_k^{(F)} \hat{b}_k(0) e^{-i(\omega_k^{(F)} - \omega)t}, \quad (4.50)$$

we finally obtain

$$\frac{d\hat{a}}{dt} = -\frac{1}{2} \frac{\omega}{Q} \hat{a} - ig \hat{N}_{ab} + \hat{F}_F. \quad (4.51)$$

Since the definition of the noise operator (4.50) coincides with the definition (2.134), we can write the correlation relations as in (2.135) and (2.136):

$$\begin{aligned} \left\langle \hat{F}_F^\dagger(t_1) \hat{F}_F(t_2) \right\rangle &= \frac{\gamma \bar{n}_T}{2} \delta(t_1 - t_2) = \frac{1}{2} \frac{\omega}{Q} \bar{n}_T \delta(t_1 - t_2), \\ \left\langle \hat{F}_F(t_1) \hat{F}_F^\dagger(t_2) \right\rangle &= \frac{1}{2} \frac{\omega}{Q} (\bar{n}_T + 1) \delta(t_1 - t_2), \end{aligned} \quad (4.52)$$

where \bar{n}_T is the average number of phonons in the reservoir mode at temperature T .

4.3 Heisenberg-Langevin Equation System Describing the Laser

Combining the equations we obtained (4.51), (4.40), and (4.44) we can write the following system:

$$\begin{aligned} \frac{d\hat{a}}{dt} &= -\frac{1}{2} \frac{\omega}{Q} \hat{a} - ig \hat{N}_{ab} + \hat{F}_F, \\ \frac{d\hat{N}_{ab}}{dt} &= -\gamma \hat{N}_{ab} + ig \hat{N}_z \hat{a} + \hat{F}_{ab}, \\ \frac{d\hat{N}_z}{dt} &= r_a - \gamma \hat{N}_z + 2ig \left(\hat{a}^\dagger \hat{N}_{ab} - \hat{N}_{ab}^\dagger \hat{a} \right) + \hat{F}_z, \end{aligned} \quad (4.53)$$

which we will call the Heisenberg-Langevin equation system describing the laser operation.

By their form, these equations (4.53) are very similar to the laser equations obtained in the semiclassical approximation. The difference is that instead of classical quantities, the equations use operators, which may not commute with

each other, so the order of factors must be preserved during transformations. Moreover, the equations include Langevin terms describing quantum noise present in the laser system.

The same approximations applied in the semiclassical case can be applied to equations (4.53).

If, as is usually the case, the adiabatic approximation can be applied, neglecting in the second equation (4.53) the derivative $\frac{d\hat{N}_{ab}}{dt}$ compared to the term $\gamma\hat{N}_{ab}$, we obtain

$$\hat{N}_{ab} = \frac{1}{\gamma} \left(ig\hat{N}_z\hat{a} + \hat{F}_{ab} \right).$$

Substituting this expression for \hat{N}_{ab} into the remaining two equations of system (4.53) yields a new system of equations:

$$\begin{aligned} \frac{d}{dt}\hat{a} &= -\frac{1}{2} \left(\frac{\omega}{Q} \right) \hat{a} + \frac{g^2}{\gamma} \hat{N}_z\hat{a} + \hat{F}_F - i\frac{g}{\gamma} \hat{F}_{ab}, \\ \frac{d}{dt}\hat{N}_z &= r_a - \gamma\hat{N}_z - \frac{4g^2}{\gamma} \hat{N}_z\hat{a}^\dagger\hat{a} + \\ &\quad + \hat{F}_z + \frac{2ig}{\gamma} \left(\hat{a}^\dagger\hat{F}_{ab} - \hat{F}_{ab}^\dagger\hat{a} \right). \end{aligned} \quad (4.54)$$

Introducing a new notation

$$\hat{F}_\Sigma = - \left(\frac{g}{\gamma} \hat{F}_{ab} + i\hat{F}_F \right), \quad (4.55)$$

the system (4.54) can be rewritten as

$$\begin{aligned} \frac{d}{dt}\hat{a} &= -\frac{1}{2} \left(\frac{\omega}{Q} \right) \hat{a} + \frac{g^2}{\gamma} \hat{N}_z\hat{a} + i\hat{F}_\Sigma, \\ \frac{d}{dt}\hat{N}_z &= r_a - \gamma\hat{N}_z - \frac{4g^2}{\gamma} \hat{N}_z\hat{a}^\dagger\hat{a} + \\ &\quad + \hat{F}_z + \frac{2ig}{\gamma} \left(\hat{a}^\dagger\hat{F}_{ab} - \hat{F}_{ab}^\dagger\hat{a} \right). \end{aligned} \quad (4.56)$$

For the noise operator (4.55), using (4.42), (4.43), and (4.52), the following correlation relations can be written:

$$\begin{aligned} \left\langle \hat{F}_\Sigma^\dagger(t_1) \hat{F}_\Sigma(t_2) \right\rangle &= \frac{g^2}{\gamma^2} \left\langle \hat{F}_{ab}^\dagger(t_1) \hat{F}_{ab}(t_2) \right\rangle + \left\langle \hat{F}_F^\dagger(t_1) \hat{F}_F(t_2) \right\rangle = \\ &= \frac{1}{2} \left(\frac{g^2}{\gamma} \bar{N}_a + 2\frac{g^2}{\gamma^2} r_a + \frac{\omega}{Q} \bar{n}_T \right) \delta(t_1 - t_2), \\ \left\langle \hat{F}_\Sigma(t_1) \hat{F}_\Sigma^\dagger(t_2) \right\rangle &= \frac{1}{2} \left(\frac{g^2}{\gamma} \bar{N}_b + \frac{\omega}{Q} (\bar{n}_T + 1) \right) \delta(t_1 - t_2) \end{aligned} \quad (4.57)$$

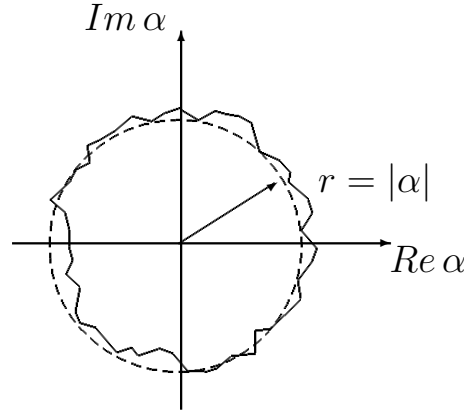


Figure 4.3: In the presence of noise sources (Langevin forces), the uncertainty of the amplitude occupies a narrow region near the circle $|\alpha| = \text{const}$, while the phase freely diffuses (wanders) along this circle.

Equations (4.56) resemble the semiclassical laser equation and can, to a large extent, be solved using the same approximations and techniques as those used in the semiclassical problem.

4.4 Natural linewidth of emission

Equations (4.56) are suitable for solving many problems related to the quantum nature of the laser: determining the statistics of laser photons in various operating modes, determining the linewidth of emission, and with some additions for solving more subtle issues, for example, the study of direct laser generation of squeezed states (for more on squeezed states see Ch. 9).

Here we will consider only the problem of the natural linewidth of laser emission. We will consider the steady-state laser operation mode that is sufficiently far from the generation threshold. In this case, the problem can be greatly simplified since the linewidth depends little on fluctuations of the population difference and amplitude fluctuations of the generated radiation, as we established earlier by considering this problem using the density matrix method. The main influence on the linewidth is due to phase fluctuations (drift). This can be visually represented on the complex amplitude plane, as shown in Figure 4.3.

If there were no noise sources (Langevin forces), the generated field would have a definite amplitude and phase. In reality, with their presence, the amplitude uncertainty occupies a narrow region near the circle $|\alpha| = \text{const}$, while the phase freely diffuses (wanders) along this circle.

In light of this, the simplification of the equations reduces to the following.

The equation for the population will be replaced by the averaged equation

$$r_a - \gamma \bar{N}_z - \frac{4g^2}{\gamma} \bar{N}_z \bar{n} = 0, \quad (4.58)$$

where \bar{N}_z is the averaged population value, $\bar{n} = \langle \hat{a}^\dagger \hat{a} \rangle$ is the average photon number in the generated mode. Zero on the right means that a steady-state regime is considered.

From (4.58) we find \bar{N}_z

$$\bar{N}_z = \frac{r_a}{\gamma \left(1 + \left(\frac{2g}{\gamma} \right)^2 \bar{n} \right)} \approx \frac{r_a}{\gamma} \left(1 - \left(\frac{2g}{\gamma} \right)^2 \bar{n} \right). \quad (4.59)$$

The final value in (4.59) is obtained under the approximation that the second term in parentheses is small compared to 1 (laser is not at threshold). Substituting this result into the field equation, we get

$$\begin{aligned} \frac{d}{dt} \hat{a} = & -\frac{1}{2} \left(\frac{\omega}{Q} \right) \hat{a} + r_a \frac{g^2}{\gamma^2} \hat{a} - \\ & - 4r_a \frac{g^4}{\gamma^4} \bar{n} \hat{a} + i \hat{F}_\Sigma, \end{aligned} \quad (4.60)$$

where the first term on the right corresponds to losses in the resonator, the second - gain by the active medium, the third characterizes the level saturation degree.

The field correlation function we need to determine is expressed by the equality

$$I(t, 0) = \langle \hat{a}^\dagger(t) \hat{a}(0) \rangle,$$

where $\hat{a}(t)$ satisfies equation (4.60).

In our approximation, when we neglect amplitude fluctuations, but do not neglect phase diffusion, the operator \hat{a} can be represented in the form

$$\begin{aligned} \hat{a}(t) &= A e^{i\hat{\varphi}(t)}, \\ \hat{a}^\dagger(t) &= A e^{-i\hat{\varphi}(t)}, \end{aligned} \quad (4.61)$$

where A is the average amplitude, and $\hat{\varphi}(t)$ is the phase operator. Then the correlation function can be presented as

$$\begin{aligned} \langle \hat{a}^\dagger(t) \hat{a}(0) \rangle &= |A|^2 \langle e^{-i(\hat{\varphi}(t) - \hat{\varphi}(0))} \rangle = \\ &= |A|^2 \langle e^{-\frac{1}{2}(\hat{\varphi}(t) - \hat{\varphi}(0))^2} \rangle. \end{aligned} \quad (4.62)$$

The last follows from the fact that

$$\left\langle e^{-i(\Delta\varphi)} \right\rangle = \left\langle e^{-\frac{1}{2}(\Delta\varphi)^2} \right\rangle,$$

where $\Delta\varphi = \hat{\varphi}(t) - \hat{\varphi}(0)$, and it can be proved by expanding the exponential in a series, averaging each series term, and then summing the result.

We have

$$\Delta\varphi = \int_0^t \frac{d\varphi}{dt} dt,$$

then

$$\left\langle (\Delta\varphi)^2 \right\rangle = \int_0^t dt_1 \left\langle \frac{d\varphi}{dt} \right\rangle \int_0^t dt_2 \left\langle \frac{d\varphi}{dt} \right\rangle,$$

where $\frac{d\varphi}{dt}$ can be obtained from equation (4.60) through the relations

$$\begin{aligned} i \frac{d\varphi}{dt} &= \frac{1}{\hat{a}} \frac{d\hat{a}}{dt} = -\frac{1}{2} \left(\frac{\omega}{Q} \right) + r_a \frac{g^2}{\gamma^2} - \\ &\quad - 4r_a \frac{g^4}{\gamma^4} \bar{n} + i \frac{\hat{F}_\Sigma}{\hat{a}}, \\ -i \frac{d\varphi}{dt} &= \frac{1}{\hat{a}^\dagger} \frac{d\hat{a}^\dagger}{dt} = -\frac{1}{2} \left(\frac{\omega}{Q} \right) + r_a \frac{g^2}{\gamma^2} - \\ &\quad - 4r_a \frac{g^4}{\gamma^4} \bar{n} - i \frac{\hat{F}_\Sigma^\dagger}{\hat{a}^\dagger}, \end{aligned} \quad (4.63)$$

thus subtracting the second equation (4.63) from the first we obtain

$$2 \frac{d\varphi}{dt} = \frac{\hat{F}_\Sigma}{\hat{a}} + \frac{\hat{F}_\Sigma^\dagger}{\hat{a}^\dagger}. \quad (4.64)$$

Substituting now (4.61) into (4.64) we have:

$$\frac{d\varphi}{dt} = \frac{1}{2A} \left\{ e^{i\hat{\varphi}(t)} \hat{F}_\Sigma^\dagger(t) + e^{-i\hat{\varphi}(t)} \hat{F}_\Sigma(t) \right\}.$$

It remains to calculate $\left\langle (\Delta\varphi)^2 \right\rangle$. This leads to a double integral

$$\begin{aligned} \left\langle (\Delta\varphi)^2 \right\rangle &= \frac{1}{4A^2} \int_0^t dt_1 \int_0^t dt_2 \left\langle \left(e^{i\hat{\varphi}(t_1)} \hat{F}_\Sigma^\dagger(t_1) + e^{-i\hat{\varphi}(t_1)} \hat{F}_\Sigma(t_1) \right) \right. \\ &\quad \left. \left(e^{i\hat{\varphi}(t_2)} \hat{F}_\Sigma^\dagger(t_2) + e^{-i\hat{\varphi}(t_2)} \hat{F}_\Sigma(t_2) \right) \right\rangle = \\ &= \frac{1}{4} \frac{1}{\bar{n}} \int_0^t dt_1 \int_0^t dt_2 \left\langle \hat{F}_\Sigma^\dagger(t_1) \hat{F}_\Sigma(t_2) e^{-i(\hat{\varphi}(t_1) - \hat{\varphi}(t_2))} + \text{h.c.} \right\rangle. \end{aligned}$$

From here, using (4.57) we obtain

$$\langle (\Delta\varphi)^2 \rangle = \frac{1}{4} \frac{1}{\bar{n}} \left(\frac{g^2}{2\gamma} (\bar{N}_a + \bar{N}_b) + \frac{g^2}{\gamma^2} r_a + \frac{\omega}{Q} \left(\bar{n}_T + \frac{1}{2} \right) \right) t. \quad (4.65)$$

From equations (4.38) and (4.40) one can write the equation for $\bar{N}_a + \bar{N}_b$:

$$\frac{d(\bar{N}_a + \bar{N}_b)}{dt} = r_a - \gamma (\bar{N}_a + \bar{N}_b),$$

hence in the steady-state regime:

$$r_a = \gamma (\bar{N}_a + \bar{N}_b).$$

Substituting this relation into (4.65) we get:

$$\langle (\Delta\varphi)^2 \rangle = \frac{1}{4} \frac{1}{\bar{n}} \left(\frac{3g^2}{2\gamma} (\bar{N}_a + \bar{N}_b) + \frac{\omega}{Q} \left(\bar{n}_T + \frac{1}{2} \right) \right) t. \quad (4.66)$$

Consider the laser at threshold:

$$A = R_a \approx \frac{\omega}{Q},$$

substituting here the expression for R_a (2.70) in which we take $\tau \approx \frac{1}{\gamma}$:

$$r_a \frac{g^2}{\gamma^2} \approx \frac{\omega}{Q}.$$

Using this relation from (4.59) we have:

$$\bar{N}_z \approx \frac{r_a}{\gamma} \approx \frac{\omega}{Q} \frac{\gamma}{g^2},$$

hence

$$\frac{g^2}{\gamma} \approx \frac{\omega}{Q} \frac{1}{\bar{N}_z}.$$

Thus (4.66) can be rewritten as

$$\langle (\Delta\varphi)^2 \rangle = \frac{1}{4} \frac{1}{\bar{n}} \frac{\omega}{Q} \left(\frac{3}{2} \frac{\bar{N}_a + \bar{N}_b}{\bar{N}_z} + \bar{n}_T + \frac{1}{2} \right) t. \quad (4.67)$$

If now in (4.67) we assume complete inversion: $\bar{N}_b = 0$, and also neglect \bar{n}_T compared to 1, then we obtain

$$\langle (\Delta\varphi)^2 \rangle \approx \frac{1}{2} \frac{1}{\bar{n}} \frac{\omega}{Q} t. \quad (4.68)$$

Substituting this into (4.62) we get

$$\langle \hat{a}^\dagger(t) \hat{a}(0) \rangle = |A|^2 e^{-\frac{1}{4} \frac{1}{\bar{n}} \frac{\omega}{Q} t}, \quad (4.69)$$

which coincides with expression (3.27) previously obtained, if we accept the threshold generation mode:

$$A \approx \frac{\omega}{Q}.$$

Thus the expression for the natural linewidth will be the same:

$$D \approx \frac{1}{2} \frac{1}{\bar{n}} \frac{\omega}{Q}.$$

4.5 Exercises

1. Prove the commutation relations (4.13) and (4.23).
2. Derive the equations of motion for the operator $\hat{\sigma}_b^j$: (4.19) and (4.20).
3. Derive the relations (4.40) for the operators \hat{N}_b and \hat{N}_{ab} .
4. Prove the second relation (4.30).
5. Prove (4.62).
6. Prove (4.65).

Chapter 5

Photon Optics (Quantum Phenomena in Optics)

We will consider those optical phenomena in which the quantum properties of light manifest to varying degrees.

Although many optical phenomena can be considered from classical positions, many phenomena can only be fully understood and described within the framework of a fully quantum description.

Quantum consideration allows for a more complete understanding of the essence of interference experiments and, on this basis, to understand the connection between classical and quantum descriptions. Moreover, the quantum approach allows for new types of experiments in which the statistics of photons in light beams and their connection with the spectral properties of light are studied.

5.1 Photoeffect

The photoeffect is used for photon detection, in which, as is known, upon absorption of a photon a bound electron transitions to a free state and is registered accordingly. One of the devices used for these purposes is the photomultiplier, schematically shown in [Figure 5.1](#). Its principle of operation is well known. Semiconductor avalanche photodiodes with an amplifier may also be used. Let us consider the relationship between the rate of photoelectron transitions and the light field. The electric field operator can be decomposed into frequency-positive and frequency-negative parts ([1.88](#)):

$$\hat{\vec{E}}(\vec{r}, t) = \hat{\vec{E}}^{(+)}(\vec{r}, t) + \hat{\vec{E}}^{(-)}(\vec{r}, t), \quad (5.1)$$

where the frequency-positive part includes annihilation operators, and the frequency-negative part includes creation operators.

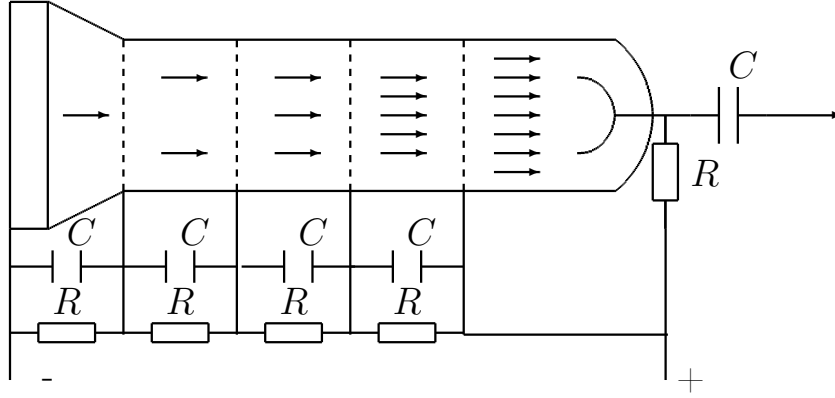


Figure 5.1: Photomultiplier

When decomposing the field into plane waves, we have:

$$\begin{aligned}\hat{\vec{E}}^{(+)}(\vec{r}, t) &= \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \hat{a}_k \vec{e}_k e^{-i\omega_k t + i(\vec{k}\vec{r})}, \\ \hat{\vec{E}}^{(-)}(\vec{r}, t) &= \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \hat{a}_k^\dagger \vec{e}_k^* e^{i\omega_k t - i(\vec{k}\vec{r})}.\end{aligned}\quad (5.2)$$

We consider the electric dipole interaction approximation, that is, we assume that the size of the atom (or other electronic system) interacting with the light is much smaller than the wavelength. The interaction Hamiltonian in this case has the form:

$$\hat{\mathcal{H}}_{ED} = -\left(\hat{\vec{p}}\hat{\vec{E}}\right), \quad (5.3)$$

where $\hat{\vec{p}}$ is the dipole moment operator of the system. Furthermore, we assume that only processes accompanied by photon annihilation are considered. Thus, in (5.3), one must retain only the annihilation operators \hat{a} , i.e., replace $\hat{\vec{E}}$ by $\hat{\vec{E}}^{(+)}$:

$$\hat{\mathcal{H}}_{ED} = -\left(\hat{\vec{p}}\hat{\vec{E}}^{(+)}\right). \quad (5.4)$$

Assume that initially the atom is in the ground state, and the multimode field contains $\{n_k\}$ photons, so that the initial state vector is

$$|i\rangle = |\{n_k\}\rangle |b\rangle \quad (5.5)$$

After the interaction, the atom-field system will be in state $|f\rangle$ (we will not detail it). The probability of transition per unit time in first-order perturbation theory according to Fermi's golden rule (Theorem 14.9.1) is given by the square

modulus of the matrix element of the transition:

$$\begin{aligned} \left| \langle f | \hat{\mathcal{H}}_{ED} | i \rangle \right|^2 &= \langle f | \left(\hat{\vec{p}} \hat{\vec{E}}^{(+)} \right) | i \rangle \langle f | \left(\hat{\vec{p}} \hat{\vec{E}}^{(+)} \right) | i \rangle^* = \\ &= \langle f | \left(\hat{\vec{p}} \hat{\vec{E}}^{(+)} \right) | i \rangle \langle i | \left(\hat{\vec{p}} \hat{\vec{E}}^{(-)} \right) | f \rangle = \langle i | \left(\hat{\vec{p}} \hat{\vec{E}}^{(-)} \right) | f \rangle \langle f | \left(\hat{\vec{p}} \hat{\vec{E}}^{(+)} \right) | i \rangle \end{aligned} \quad (5.6)$$

The final state is unknown. It can be any state, and it is necessary to sum (5.6) over all final states (sum the probabilities). We consider the system of final states complete, i.e., $\sum_{(f)} |f\rangle \langle f| = \hat{I}$. From this we have:

$$\begin{aligned} \sum_{(f)} \langle i | \left(\hat{\vec{p}} \hat{\vec{E}}^{(-)} \right) | f \rangle \langle f | \left(\hat{\vec{p}} \hat{\vec{E}}^{(+)} \right) | i \rangle &= \langle i | \hat{p}_E^2 \hat{E}^{(-)} \hat{E}^{(+)} | i \rangle = \\ &= \langle \{n_k\} | \hat{E}^{(-)} \hat{E}^{(+)} | \{n_k\} \rangle \langle b | \hat{p}_E^2 | b \rangle = \alpha \langle \{n_k\} | \hat{E}^{(-)} \hat{E}^{(+)} | \{n_k\} \rangle, \end{aligned} \quad (5.7)$$

where \hat{p}_E is the projection of the dipole moment onto the direction of the field, and α is some interaction constant that does not depend on the field magnitude. Now let us generalize (5.7) to the case of a mixed initial state. Suppose we know the probability $P_{\{n_k\}}$ of the state $|\{n_k\}\rangle$. The density matrix of this state can be represented as

$$\hat{\rho} = \sum_{\{n_k\}} P_{\{n_k\}} |\{n_k\}\rangle \langle \{n_k\}|. \quad (5.8)$$

To obtain the result in the mixed state case, it is necessary to average (5.7) using the probabilities $P_{\{n_k\}}$. For the photo-count rate, we have:

$$\begin{aligned} W &= \alpha \sum_{\{n_k\}} P_{\{n_k\}} \langle \{n_k\} | \hat{E}^{(-)} \hat{E}^{(+)} | \{n_k\} \rangle = \\ &= \alpha Sp \left(\hat{\rho} \hat{E}^{(-)} \hat{E}^{(+)} \right) \end{aligned} \quad (5.9)$$

- in an arbitrary representation, since Sp does not depend on the representation.

Thus, in the general case of a statistically mixed state, the average photon count rate is proportional to the expectation value of the operators $\hat{E}^{(-)} \hat{E}^{(+)}$ in the initial field state, where $\hat{\rho}$ is the statistical operator of the initial field state.

We considered an atom or a similar microsystem as a photodetector. A real photodetector contains many atoms. If we assume that the detector contains N non-interacting atoms, and its size is small enough so that all atoms can be regarded as being in identical conditions with good accuracy, then the probability (5.9) can simply be increased by a factor of N .

Let us determine the meaning of the operator $\hat{E}^{(-)} \hat{E}^{(+)}$. Introduce the operator

$$\hat{\vec{J}} = 2\varepsilon_0 \left(\hat{E}^{(-)} \hat{E}^{(+)} \right) c \vec{k}_0, \quad (5.10)$$

where \vec{k}_0 is the unit vector in the propagation direction of the wave; the vector operator $\hat{\vec{J}}$ can be considered as the operator of energy flux in the direction \vec{k}_0 . For example, for a single-mode field, this operator equals

$$\hat{\vec{J}} = \frac{c\hbar\omega_k}{V}\vec{k}_0\hat{n}_k, \quad (5.11)$$

where \vec{k}_0 is the unit vector in the direction \vec{k} . The mean value of the operator $\hat{\vec{J}}$ in a state with a definite photon number is

$$\langle \hat{\vec{J}} \rangle = \frac{c\hbar\omega_k}{V}n_k\vec{k}_0,$$

In a single-mode mixed state, we have

$$\langle \hat{\vec{J}} \rangle = \frac{c\hbar\omega_k}{V}\bar{n}\vec{k}_0, \text{ where } \bar{n} = \sum_{(n_k)} P_{n_k} n_k.$$

From the examples given, it is clear that $\hat{\vec{J}}$ indeed has the meaning of the energy flux operator. In the single-mode case, the intensity does not depend on time. In the multimode case, the intensity can vary in time and space. From all the above, it can be concluded that the photoemission rate is proportional to the average photon flux, i.e., the light intensity. In a semiclassical treatment, this corresponds to the average energy flux over a period.

5.2 Coherent Properties of Light

The coherent properties of light determine the interference phenomena in optics, which are well described by classical wave theory. In classical wave optics, the coherent properties of light are described by the corresponding coherence functions [59, 49, 60]. Let us consider the phenomenon of interference from a quantum point of view. To this end, consider Young's interference experiment (see Figure 5.2). On the screen (photographic plate), the arrival of a photon is recorded as a dark spot. To obtain an interference pattern similar to the classical one, a sufficient amount of time is needed. Then individual points merge into fringes and the pattern will not differ from the classical one (Figure 5.3).

Field operators on the screen can be represented as a superposition of fields passing through holes 1 and 2:

$$\hat{E}^{(+)}(\vec{r}, t) = \gamma \left[\hat{E}^{(+)}(\vec{r}_1, t_1) + \hat{E}^{(+)}(\vec{r}_2, t_2) \right], \quad (5.12)$$

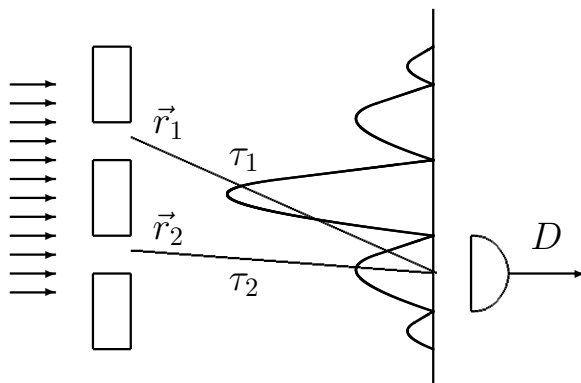


Figure 5.2: Young's interference experiment. $\vec{r}_{1,2}$ - coordinates of the two openings, $\tau_{1,2}$ - delays in the propagation of light from the openings to the screen, D - photodetector.

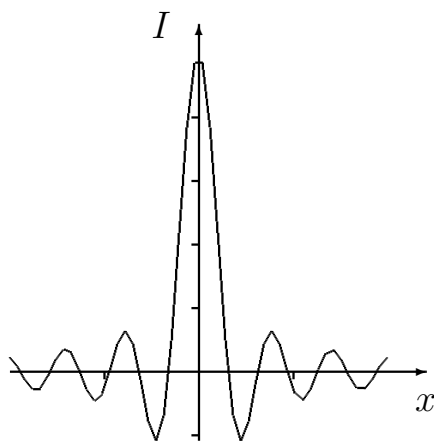


Figure 5.3: Classical interference pattern. Dependence of light intensity I on the coordinate on the screen x .

where \vec{r}_1, \vec{r}_2 are the coordinates of the holes; $t_1 = t - \tau_1, t_2 = t - \tau_2$; τ_1, τ_2 are the delays due to propagation of light from the holes to the screen; the coefficient γ characterizes the attenuation of the field during propagation from the holes to the screen. Using photodetector D , we examine the intensity of light at different points on the screen.

The photon counting rate will be, as we know, proportional to

$$W = \alpha Sp \left(\hat{\rho} \hat{E}^{(-)}(\vec{r}, t) \hat{E}^{(+)}(\vec{r}, t) \right) = \alpha |\gamma|^2 Sp \left[\hat{\rho} \left(\hat{E}^{(-)}(\vec{r}_1, t_1) + \hat{E}^{(-)}(\vec{r}_2, t_2) \right) \left(\hat{E}^{(+)}(\vec{r}_1, t_1) + \hat{E}^{(+)}(\vec{r}_2, t_2) \right) \right].$$

Multiplying out term by term, we obtain:

$$W = g Sp \left[\hat{\rho} \left(\hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_1, t_1) + \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_2, t_2) + \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) + \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_1, t_1) \right) \right], \quad (5.13)$$

where the following notation is introduced: $g = \alpha |\gamma|^2$.

The first two terms give the intensity of the fields passing through the first and second holes when the other hole is closed. The last two terms describe interference. Since

$$\left(\hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_1, t_1) \right)^\dagger = \left(\hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) \right)$$

then $Sp \left(\hat{\rho} \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_1, t_1) \right)$ and $Sp \left(\hat{\rho} \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) \right)$ will be complex conjugates. The term

$$Sp \left[\hat{\rho} \left(\hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) + \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_1, t_1) \right) \right]$$

gives the oscillating interference term (Figure 5.3). The envelope of the interference term is proportional to the expression

$$\left| Sp \left[\hat{\rho} \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) \right] \right| \quad (5.14)$$

From this follows the definition of the first-order coherence function

$$G(\vec{r}_1, t_1, \vec{r}_2, t_2) = G_{12}^1 = \frac{\left| Sp \left[\hat{\rho} \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) \right] \right|}{\sqrt{Sp \left[\hat{\rho} \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_1, t_1) \right] Sp \left[\hat{\rho} \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_2, t_2) \right]}} \quad (5.15)$$

This expression resembles the classical definition of the coherence function, but here instead of the analytic signal we have field operators and quantum averaging is performed using the density matrix.

Let us consider some particular cases as examples. We start with a single-mode field in the states $|n\rangle$ or $|\alpha\rangle$. In this case, we have:

$$\begin{aligned}\hat{E}^{(+)}(\vec{r}, t) &= \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \hat{a}_k \vec{e}_k e^{-i\omega_k t + i(\vec{k}\vec{r})}, \\ \hat{E}^{(-)}(\vec{r}, t) &= \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \hat{a}_k^\dagger \vec{e}_k^* e^{i\omega_k t - i(\vec{k}\vec{r})} = \left(\hat{E}^{(+)}(\vec{r}, t) \right)^\dagger.\end{aligned}\quad (5.16)$$

Using these expressions to calculate the coherence function (5.15), we get $G_{12}^1 = 1$ since $\langle n | \hat{a}^\dagger \hat{a} | n \rangle = n$ and $|e^{-ix}| = 1$. The same result is obtained for the coherent state case: $G_{12}^1 = 1$, as $\langle \alpha | \hat{a}^\dagger \hat{a} | \alpha \rangle = |\alpha|^2$.

In general, a single-mode field excited in an arbitrary pure state possesses full first-order coherence. Moreover, a single-mode field excited in an arbitrary statistical mixed state also exhibits full coherence. In this case, one needs to compute $Sp(\hat{\rho} \hat{a}^\dagger \hat{a})$, where $\hat{\rho} = \sum_{(m)} \sum_{(l)} \rho_{ml} |m\rangle \langle n|$. We have

$$\begin{aligned}Sp(\hat{\rho} \hat{a}^\dagger \hat{a}) &= \sum_{(n)} \sum_{(m)} \sum_{(l)} \rho_{ml} \langle n | m \rangle \langle l | \hat{a}^\dagger \hat{a} | n \rangle = \\ &= \sum_{(n)} \sum_{(l)} \rho_{nl} \langle l | n \rangle n = \sum_{(n)} \rho_{nn} n = \bar{n}.\end{aligned}\quad (5.17)$$

A similar expression appears in the denominator, therefore, $G_{12}^{(1)} = 1$.

More commonly, one deals with multimode states of the field, so let us consider the coherence of light in this case. Consider two limiting cases: light in a coherent state and fully chaotic light.

In the first case, the state vector is

$$|\{\alpha_k\}\rangle = |\{\alpha_{k_1}\}\rangle |\{\alpha_{k_2}\}\rangle \dots |\{\alpha_{k_s}\}\rangle \dots$$

Using the properties of coherent states,

$$\hat{E}^{(+)}(\vec{r}, t) |\{\alpha_k\}\rangle = E(\vec{r}, t) |\{\alpha_k\}\rangle$$

and

$$\langle \{\alpha_k\} | \hat{E}^{(-)}(\vec{r}, t) = E^*(\vec{r}, t) \langle \{\alpha_k\} |,$$

we get:

$$\langle \{\alpha_k\} | \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) | \{\alpha_k\} \rangle = E^*(\vec{r}_1, t_1) E(\vec{r}_2, t_2)$$

where E is the eigenvalue of the operator $\hat{E}^{(+)}$, being the analytic signal of the classical field. From this, we have

$$G_{12}^{(1)} = \frac{|E^*(\vec{r}_1, t_1) E(\vec{r}_2, t_2)|}{\sqrt{E^*(\vec{r}_1, t_1) E(\vec{r}_1, t_1) E^*(\vec{r}_2, t_2) E(\vec{r}_2, t_2)}} = 1 \quad (5.18)$$

that is, a multimode field in a coherent state is fully coherent.

In the case of an arbitrary multimode field state, we have only partial coherence. Consider the most common case of chaotic light (radiation from a heated body, gas discharge, etc.), where light is emitted by many independent sources (atoms, ions, molecules). The expression for the density matrix in this case has the form (1.114):

$$\begin{aligned} \hat{\rho} &= \sum_{\{n_k\}} |\{n_k\}\rangle \langle \{n_k\}| \prod_{\{n_k\}} \frac{\bar{n}_k^{n_k}}{(1 + \bar{n}_k)^{n_k+1}} = \\ &= \sum_{\{n_k\}} \rho_{\{n_k\}, \{n_k\}} |\{n_k\}\rangle \langle \{n_k\}| \end{aligned} \quad (5.19)$$

where

$$\begin{aligned} \rho_{\{n_k\}, \{n_k\}} &= P_{\{n_k\}} = \prod_k \frac{\bar{n}_k^{n_k}}{(1 + \bar{n}_k)^{n_k+1}} = \prod_k P_{n_k}, \\ \sum_{\{n_k\}} \dots &= \sum_{n_1} \sum_{n_2} \dots \sum_{n_k} \dots \end{aligned}$$

The electric field operators are

$$\begin{aligned} \hat{\vec{E}}^{(+)} &= \sum_{(k)} \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} \vec{e}_k \hat{a}_k e^{-i \omega_k t + i(\vec{k} \vec{r})}, \\ \hat{\vec{E}}^{(-)} &= \sum_{(k)} \sqrt{\frac{\hbar \omega_k}{2 \varepsilon_0 V}} \vec{e}_k^* \hat{a}_k^\dagger e^{i \omega_k t - i(\vec{k} \vec{r})}, \end{aligned} \quad (5.20)$$

from which it follows:

$$\begin{aligned} &Sp \left(\hat{\rho} \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) \right) = \\ &= \sum_{\{n_{k'}\}} \sum_{\{n_k\}} \langle \{n_{k'}\} | \{n_k\} \rangle P_{\{n_k\}} \langle \{n_k\} | \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) | \{n_{k'}\} \rangle = \\ &= \sum_{\{n_k\}} \langle \{n_k\} | \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) | \{n_k\} \rangle P_{\{n_k\}}. \end{aligned} \quad (5.21)$$

The product of operators equals

$$\hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) = \sum_{(k)} \sum_{(k')} \frac{\hbar \sqrt{\omega_k \omega_{k'}}}{2\varepsilon_0 V} (\vec{e}_k \vec{e}_{k'}) \hat{a}_k^\dagger \hat{a}_{k'} e^{-ix_2 + ix_1}, \quad (5.22)$$

where $x_1 = \omega_k t_1 - (\vec{k} \vec{r}_1)$, $x_2 = \omega_k t_2 - (\vec{k} \vec{r}_2)$. The operator $\hat{a}_k^\dagger \hat{a}_{k'}$ is averaged. Since

$$|\{n_k\}\rangle = |n_{k_1}\rangle |n_{k_2}\rangle \dots,$$

we have

$$\langle \{n_k\} | \hat{a}_k^\dagger \hat{a}_{k'} | \{n_k\} \rangle = n_k \delta_{kk'},$$

and, therefore,

$$\begin{aligned} Sp \left(\hat{\rho} \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) \right) &= \\ &= \sum_k \sum_{\{n_k\}} \frac{\hbar \omega_k}{2\varepsilon_0 V} n_k e^{-i(x_2 - x_1)} P_{\{n_k\}} = \\ &= \sum_k \frac{\hbar \omega_k}{2\varepsilon_0 V} \sum_{\{n_k\}} n_k \prod_k P_{n_k} = \sum_k \frac{\hbar \omega_k}{2\varepsilon_0 V} \bar{n}_k e^{-i(x_2 - x_1)}, \end{aligned}$$

since

$$\begin{aligned} \sum_{\{n_k\}} n_k \prod_k P_{n_k} &= \sum_{n_1} P_{n_1} \sum_{n_2} P_{n_2} \dots \sum_{n_k} n_k P_{n_k} \dots = \\ &= \sum_{n_k} n_k P_{n_k} = \bar{n}_k, \end{aligned}$$

because

$$\sum_{n_k} P_{n_k} = 1.$$

From this we get

$$G_{12}^{(1)} = \frac{\left| \sum_k \frac{\hbar \omega_k}{2\varepsilon_0 V} \bar{n}_k e^{-i\omega_k(t_2 - t_1) + i(\vec{k}, \vec{r}_2 - \vec{r}_1)} \right|}{\sum_k \frac{\hbar \omega_k}{2\varepsilon_0 V} \bar{n}_k} \quad (5.23)$$

The summation over k can be replaced by integration over frequency. Using the expression for the density of states, we get:

$$\begin{aligned} \sum_k (\dots) &= \frac{L^3}{(2\pi)^3} \int \int \int (\dots) dk_x dk_y dk_z = \\ &= \frac{V}{(2\pi)^3} \int_\Omega d\Omega \int \omega^2 (\dots) d\omega. \end{aligned}$$

This leads us to the expression

$$G_{12}^{(1)} = \frac{\left| \int d\Omega \int \omega^3 \bar{n}(\omega, \Omega) e^{-i\omega(t_2-t_1)+i(\vec{k}, \vec{r}_2-\vec{r}_1)} d\omega \right|}{\int d\Omega \int \omega^3 \bar{n}(\omega, \Omega) d\omega}. \quad (5.24)$$

If the line is narrow compared to the carrier frequency, that is $\bar{n}(\omega, \Omega)$ has a narrow peak near the frequency $\sim \omega_0$, the function ω^3 can be taken out of the integral and canceled; moreover, if the light beam is narrow, that is, the integration domain over Ω is small, the expression (5.24) can be simplified to

$$G_{12}^{(1)} = \frac{\left| \int \bar{N}(\omega) e^{-i\omega(t_2-t_1)+i(\vec{k}, \vec{r}_2-\vec{r}_1)} d\omega \right|}{\int \bar{N}(\omega) d\omega}, \quad (5.25)$$

where $\bar{N}(\omega) = \int_{\Delta\Omega} \bar{n}(\omega, \Omega) d\Omega$. For example, when the spectral line is Lorentzian, we have:

$$\bar{N}(\omega) = \bar{N}_0 \frac{\gamma}{(\omega_0 - \omega)^2 + \gamma^2}.$$

The problem considered can be solved using the representation of coherent states. Then

$$\hat{\rho} = \int \cdots \int P(\{\alpha_k\}) |\{\alpha_k\}\rangle \langle \{\alpha_k\}| d^2\{\alpha_k\},$$

where

$$P(\{\alpha_k\}) = \prod_k \frac{1}{\pi \bar{n}_k} e^{-\frac{|\alpha_k|^2}{\bar{n}_k}} = \prod_k P_k(\alpha_k).$$

We have

$$\begin{aligned} & Sp \left(\hat{\rho} \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) \right) = \\ & = \int \cdots \int P(\{\alpha_k\}) \langle \{\alpha_k\} | \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) | \{\alpha_k\} \rangle d^2\{\alpha_k\}. \end{aligned} \quad (5.26)$$

The product of operators equals

$$\hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) = \sum_k \sum_{k'} \frac{\hbar \sqrt{\omega_k \omega_{k'}}}{2\varepsilon_0 V} (\vec{e}_k \vec{e}_{k'}) \hat{a}_k^\dagger \hat{a}_{k'} e^{-ix_2 + ix_1}.$$

The product of operators is written in normal order (annihilation operators to the right of creation operators), so the matrix element appearing in (5.26) can be

easily written, replacing $\hat{a}_{k'} \rightarrow \alpha_{k'}$, $\hat{a}_k^\dagger \rightarrow \alpha_k^*$. From this, we have

$$\begin{aligned}
& Sp \left(\hat{\rho} \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) \right) = \\
& = \sum_{k \neq k'} \frac{\hbar \sqrt{\omega_k \omega_{k'}}}{2\varepsilon_0 V} (\vec{e}_k \vec{e}_{k'}) e^{-ix_2 + ix_1} \cdot \\
& \cdot \int P(\alpha_k) \alpha_k^* d^2 \alpha_k \int P(\alpha_{k'}) \alpha_{k'} d^2 \alpha_{k'} + \\
& + \sum_k \frac{\hbar \omega_k}{2\varepsilon_0 V} e^{-ix_2 + ix_1} \int |\alpha_k|^2 d^2 \alpha_k.
\end{aligned} \tag{5.27}$$

In the derivation of (5.27), the following relation was used:

$$\int P(\alpha_k) d^2 \alpha_k = 1.$$

Further, the following relations are used:

$$\begin{aligned}
& \int P(\alpha_k) \alpha_k d^2 \alpha_k = 0, \\
& \int |\alpha_k|^2 d^2 \alpha_k = \bar{n}_k,
\end{aligned}$$

which can be easily verified by switching to polar coordinates:

$$\alpha = |\alpha| e^{i\theta} = r e^{i\theta}.$$

As a result, for (5.27) we obtain

$$Sp \left(\hat{\rho} \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_2) \right) = \sum_k \frac{\hbar \omega_k}{2\varepsilon_0 V} e^{-ix_2 + ix_1} \bar{n}_k.$$

Using these results, which coincide with those obtained previously using the number state representation n , we get the same final expressions (5.24) and (5.25).

5.3 Second-Order Coherence

Second-order coherence can be introduced based on the analysis of various experiments where the rate of simultaneous photon registration by two detectors is measured. The analysis leads to the following definition of the second-order



Figure 5.4: Scheme for measuring second-order coherence. Two photodetectors $D^{(1)}$ and $D^{(2)}$ are located at points in space, defined by vectors \vec{r}_1 and \vec{r}_2 . The signal from $D^{(2)}$ passes through a delay τ , so the coincidence circuit E registers the simultaneous absorption of two photons, one at point \vec{r}_1 at time $t_1 = t$, and the second at point \vec{r}_2 at time $t_2 = t - \tau$.

degree of coherence:

$$G^{(2)}(\vec{r}_1, t_1, \vec{r}_2, t_2) = G_{12}^{(2)} = \frac{\langle \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) \rangle}{\langle \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_1, t_1) \rangle \langle \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(+)}(\vec{r}_2, t_2) \rangle}. \quad (5.28)$$

Angle brackets denote quantum mechanical averaging over the ensemble using the statistical operator (density matrix)

$$\langle (\dots) \rangle = Sp \{ \hat{\rho} (\dots) \}$$

Expression (5.28) resembles the classical expression that defines second-order coherence, but instead of classical fields, operators are used, and averaging is carried out using the density matrix.

Formula (5.28) can be justified as follows. Consider the operator $\hat{E}^{(+)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2)$. It corresponds to the absorption of one photon at point \vec{r}_1 and time $t_1 = t$, and a second photon at point \vec{r}_2 at time $t_2 = t - \tau$. This procedure can be realized using the setup shown in Figure 5.4. Points \vec{r}_1 and \vec{r}_2 are defined by the positions of the photodetectors. Times $t_1 = t$ and $t_2 = t - \tau$ are determined by coincident photo counts. τ is an adjustable delay. Signals from both detectors are fed into a coincidence circuit.

Applying the procedure used for the photoelectric effect to the operator $\hat{E}^{(+)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2)$ the probability of detecting the first photon at time $t_1 = t$ and the second at time $t_2 = t - \tau$ within the interval Δt is:

$$w(\vec{r}_1, \vec{r}_2, t_1, t_2) \Delta t = \alpha \left| \langle f | \hat{E}^{(+)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) | i \rangle \right|^2 \Delta t,$$

where $|i\rangle$ is the initial state of the system, $|f\rangle$ is the final state of the system, α is a quantity depending on the properties of the photodetectors. Then, the counting rate (number of counts per unit time) after summation over final states is given by:

$$w(\vec{r}_1, \vec{r}_2, t_1, t_2) = \alpha \langle i | \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) | i \rangle.$$

Here, only the field part is of interest; the atomic part is included in the coefficient α .

If the initial field is in a statistically mixed state, averaging must be done using the statistical operator of the initial field. Then we have:

$$w(\vec{r}_1, \vec{r}_2, t_1, t_2) = \alpha Sp \left\{ \hat{\rho} \hat{E}^{(-)}(\vec{r}_2, t_2) \hat{E}^{(-)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_1, t_1) \hat{E}^{(+)}(\vec{r}_2, t_2) \right\}.$$

Using this expression for the normalized degree of coherence, we obtain formula (5.28).

Let us provide some examples of calculating the second-order degree of coherence. We start with a simple one. Find the second-order coherence for a single-mode state with a definite number of photons. For this, we need to consider the matrix element

$$\langle n | \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} | n \rangle = n(n-1).$$

This expression will be in the numerator. In the denominator, we have

$$\left(\langle n | \hat{a}^\dagger \hat{a} | n \rangle \right)^2 = n^2.$$

The numerical coefficients in numerator and denominator cancel out. We get:

$$\begin{aligned} G_{12}^{(2)} &= \frac{n(n-1)}{n^2} = \frac{n-1}{n} \text{ for } n > 2, \\ G_{12}^{(2)} &= 0 \text{ for } n = 1, \\ G_{12}^{(2)} &\text{ is undefined for } n = 0. \end{aligned} \tag{5.29}$$

The obvious fact in (5.29) is worth noting: if there is one photon, the probability of detecting two photons equals zero, which leads to $G_{12}^{(2)} = 0$ at $n = 1$.

Consider another simple case: a single-mode field in a coherent state. We have

$$\begin{aligned} \langle \alpha | \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} | \alpha \rangle &= (\alpha^* \alpha)^2 \text{ - in numerator,} \\ \left(\langle \alpha | \hat{a}^\dagger \hat{a} | \alpha \rangle \right)^2 &= (\alpha^* \alpha)^2 \text{ - in denominator.} \end{aligned}$$

Altogether this gives

$$G_{12}^{(2)} = 1.$$

So far, we have considered the coherence of pure states. Let us proceed to the consideration of mixed states. Consider single-mode chaotic light, whose density matrix is

$$\hat{\rho} = \sum_n \frac{\bar{n}^n}{(\bar{n} + 1)^{n+1}} |n\rangle \langle n| = \sum_n \rho_{nn} |n\rangle \langle n|.$$

From this, we obtain

$$\begin{aligned} Sp \{ \hat{\rho} \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \} &= \sum_n \sum_m \langle n|m \rangle \langle m| \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} |m \rangle \rho_{mm} = \\ &= \sum_n \rho_{nn} (n-1)n = \bar{n}^2 - \bar{n}. \end{aligned} \quad (5.30)$$

For chaotic light, the relation holds

$$\bar{n}^2 = 2(\bar{n})^2 + \bar{n}. \quad (5.31)$$

Indeed,

$$\begin{aligned} \bar{n}^2 &= \sum_n n^2 \frac{\bar{n}^n}{(\bar{n} + 1)^{n+1}} = \frac{\bar{n}}{(\bar{n} + 1)} \sum_n n^2 \frac{\bar{n}^{n-1}}{(\bar{n} + 1)^n} = \\ &= \frac{\bar{n}}{(\bar{n} + 1)} \sum_{m=n-1} (m+1)^2 \frac{\bar{n}^m}{(\bar{n} + 1)^{m+1}} = \\ &= \frac{\bar{n}}{(\bar{n} + 1)} \sum_m (m^2 + 2m + 1) \frac{\bar{n}^m}{(\bar{n} + 1)^{m+1}} = \\ &= \frac{\bar{n}}{(\bar{n} + 1)} (\bar{n}^2 + 2\bar{n} + 1), \end{aligned}$$

thus we get

$$\bar{n}^2 (\bar{n} + 1) = \bar{n} (\bar{n}^2 + 2\bar{n} + 1)$$

from which follows the desired expression (5.31):

$$\bar{n}^2 = 2(\bar{n})^2 + \bar{n}.$$

Thus, from (5.30) and (5.31), we have

$$Sp \{ \dots \} = 2(\bar{n})^2$$

In the denominator, we have $(\bar{n})^2$. Hence, for chaotic light,

$$G_{12}^{(2)} = 2, \quad (5.32)$$

which means photon pairs are registered more frequently than observed for more ordered light. Physically, this is related to fluctuations of chaotic light. Because of this, it is sometimes said that photons tend to bunch. Generally, for an arbitrary single-mode field, the second-order coherence is

$$G^{(2)} = \frac{\bar{n}^2 - \bar{n}}{(\bar{n})^2}. \quad (5.33)$$

We will consider this issue in more detail in the third part of this book (see ch. 8 Non-classical light).

In practice, multimode fields are more common. Define the second-order degree of coherence in two limiting cases: multimode coherent state and multimode chaotic light. The first case is simple to consider. The state vector can be represented as

$$|\{\alpha_k\}\rangle = |\{\alpha_{k_1}\}\rangle |\{\alpha_{k_2}\}\rangle \dots |\{\alpha_{k_s}\}\rangle \dots \quad (5.34)$$

Using the equalities

$$\hat{E}^{(+)}(x) |\{\alpha_k\}\rangle = E(x) |\{\alpha_k\}\rangle$$

and

$$\langle\{\alpha_k\}|\hat{E}^{(-)}(x) = \langle\{\alpha_k\}|E^*(x),$$

where $x = (t, r)$, $E(x)$ is the analytic signal (positive-frequency part) of the classical field obtained from the operator $\hat{E}^{(+)}$ by replacing $\hat{a}_k \rightarrow \alpha_k$. From this, we get:

$$\begin{aligned} \langle\{\alpha_k\}|\hat{E}^{(-)}(x_2)\hat{E}^{(-)}(x_1)\hat{E}^{(+)}(x_1)\hat{E}^{(+)}(x_2)|\{\alpha_k\}\rangle = \\ = E^*(x_2)E^*(x_1)E(x_1)E(x_2). \end{aligned}$$

Similarly, we see that the denominator has the quantity $E^*(x_1)E(x_1)E^*(x_2)E(x_2)$. Therefore, in this case,

$$G_{12}^{(2)} = 1, \quad (5.35)$$

The case of a chaotic multimode light field is more difficult. The statistical operator in this case will have the form (1.114)

$$\begin{aligned} \hat{\rho} = \sum_{\{n_k\}} P_{\{n_k\}} |\{n_k\}\rangle \langle\{n_k\}| = \sum_{\{n_k\}} |\{n_k\}\rangle \langle\{n_k\}| \prod_{\{n_k\}} \frac{\bar{n}_k^{n_k}}{(1 + \bar{n}_k)^{n_k+1}} = \\ = \sum_{\{n_k\}} |\{n_k\}\rangle \langle\{n_k\}| \prod_{\{n_k\}} P_{\{n_k\}}. \end{aligned} \quad (5.36)$$

When multiplying the electric field operators, we obtain a fourfold sum, since $\hat{E}^{(+)}$ and $\hat{E}^{(-)}$ are expanded in plane waves:

$$\begin{aligned}\hat{E}^{(+)}(\vec{r}, t) &= \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \vec{e}_k \hat{a}_k e^{-i\omega_k t + i(\vec{k}\vec{r})}, \\ \hat{E}^{(-)}(\vec{r}, t) &= \sum_{(k)} \sqrt{\frac{\hbar\omega_k}{2\varepsilon_0 V}} \vec{e}_k^* \hat{a}_k^\dagger e^{i\omega_k t - i(\vec{k}\vec{r})}.\end{aligned}\quad (5.37)$$

General term in the product of sums (5.37) will contain operator products of the form

$$\hat{a}_{k^I}^\dagger \hat{a}_{k^{II}}^\dagger \hat{a}_{k^{III}} \hat{a}_{k^{IV}}. \quad (5.38)$$

Averaging this term using the statistical operator (5.36) leads to

$$Sp \left(\hat{\rho} \hat{a}_{k^I}^\dagger \hat{a}_{k^{II}}^\dagger \hat{a}_{k^{III}} \hat{a}_{k^{IV}} \right) = \sum_{\{n_k\}} P_{\{n_k\}} \langle \{n_k\} | \hat{a}_{k^I}^\dagger \hat{a}_{k^{II}}^\dagger \hat{a}_{k^{III}} \hat{a}_{k^{IV}} | \{n_k\} \rangle.$$

It is easy to see that terms in which all modes are different:

$$k^I \neq k^{II} \neq k^{III} \neq k^{IV},$$

will be zero. Non-zero terms satisfy the conditions:

$$k^I = k^{III} = k_1, \quad k^{II} = k^{IV} = k_2$$

or

$$k^I = k^{IV} = k_1, \quad k^{II} = k^{III} = k_2$$

or

$$k^I = k^{II} = k^{III} = k^{IV} = k.$$

We have: in the first case -

$$\sum_{n_{k_1}} \sum_{n_{k_2}} P_{n_{k_1}} P_{n_{k_2}} n_{k_1} n_{k_2} = \bar{n}_{k_1} \bar{n}_{k_2}, \quad (5.39)$$

in the second -

$$\sum_{n_{k_1}} \sum_{n_{k_2}} P_{n_{k_1}} P_{n_{k_2}} n_{k_1} n_{k_2} = \bar{n}_{k_1} \bar{n}_{k_2}. \quad (5.40)$$

The last case was considered earlier (5.33), which gives

$$\sum_{n_k} P_{n_k} \langle n_k | \hat{a}_k^\dagger \hat{a}_k^\dagger \hat{a}_k \hat{a}_k | n_k \rangle = 2 (\bar{n}_k)^2. \quad (5.41)$$

All the above allows writing the coherence function as

$$\begin{aligned}
G_{12}^{(2)} = & \frac{\sum_{k_1} \sum_{k_2 \neq k_1} \bar{n}_{k_1} \bar{n}_{k_2} \omega_{k_1} \omega_{k_2} e^{i\omega_{k_1}\tau} e^{-i\omega_{k_2}\tau}}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2} + \\
& + \frac{2 \sum_k \bar{n}_k^2 \omega_k^2}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2} + \\
& + \frac{\sum_{k_1} \sum_{k_2 \neq k_1} \bar{n}_{k_1} \bar{n}_{k_2} \omega_{k_1} \omega_{k_2}}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2}
\end{aligned} \tag{5.42}$$

where

$$\omega\tau = \omega(t_2 - t_1) - \left(\vec{k}, \vec{r}_2 - \vec{r}_1\right) = \omega \left[(t_2 - t_1) - \frac{1}{c} \left(\vec{k}_0, \vec{r}_2 - \vec{r}_1\right) \right],$$

if we have a narrow light beam in which all modes propagate approximately in the same direction.

Half of the average sum $2 \sum_k \bar{n}_k^2 \omega_k^2$ can be combined with the left sum, and the other half with the right one. We get:

$$\begin{aligned}
G_{12}^{(2)} = & \frac{\sum_{k_1} \sum_{k_2} \bar{n}_{k_1} \bar{n}_{k_2} \omega_{k_1} \omega_{k_2} e^{i\omega_{k_1}\tau} e^{-i\omega_{k_2}\tau}}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2} + \\
& + \frac{\sum_{k_1} \sum_{k_2} \bar{n}_{k_1} \bar{n}_{k_2} \omega_{k_1} \omega_{k_2}}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2} = \\
= & \frac{\sum_{k_1} \bar{n}_{k_1} \omega_{k_1} e^{i\omega_{k_1}\tau} \sum_{k_2} \bar{n}_{k_2} \omega_{k_2} e^{-i\omega_{k_2}\tau} + \left(\sum_{(k)} \bar{n}_k \omega_k\right)^2}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2} = \\
= & \frac{\left|\sum_{(k)} \bar{n}_k \omega_k e^{i\omega_k\tau}\right|^2 + \left(\sum_{(k)} \bar{n}_k \omega_k\right)^2}{\left(\sum_{(k)} \bar{n}_k \omega_k\right)^2} = \left(G_{12}^{(1)}\right)^2 + 1,
\end{aligned} \tag{5.43}$$

where $G_{12}^{(1)}$ is the first-order coherence function.

An important result has been obtained: the second-order coherence for chaotic light is expressed via the first-order coherence. This fact forms the basis of so-called intensity interferometry. It was first observed in the experiments by Hanbury Brown and Twiss, whose setup is shown in [Figure 5.5](#).

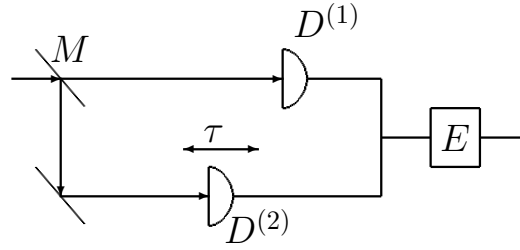


Figure 5.5: Experiment for observing second-order coherence. Using a 50% mirror M , the light beam is directed to two photodetectors $D^{(1)}$ and $D^{(2)}$. The adjustable delay τ is implemented by moving photodetector $D^{(2)}$.

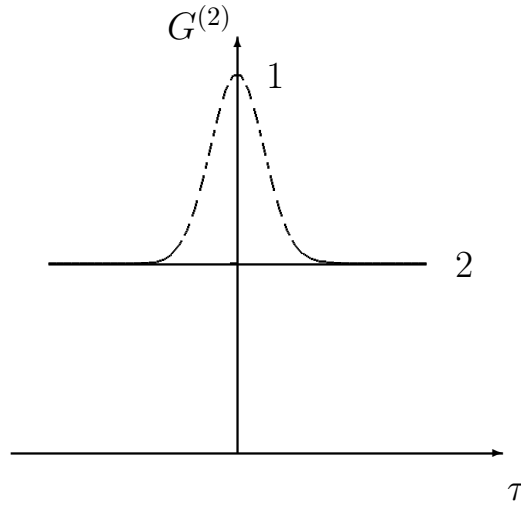


Figure 5.6: Dependence of second-order coherence $G^{(2)}$ on delay τ . Plot 1 describes the case of chaotic light. Plot 2 describes a coherent state (ideal laser).

As seen from [Figure 5.5](#), using a 50% beam splitter, the light beam is directed to two photodetectors. Delay is introduced by moving one of the photodetectors. The coincidence circuit records detection of two photons with a given time delay.

[Figure 5.6](#) shows the dependence of the number of coincidences on the delay. In the classical case, the experimental setup remains the same, only the coincidence circuit is replaced by a correlator. The experimental result in this case has the form shown in [Figure 5.6](#).

In these experiments, intensity correlation was measured, i.e., the second-order coherence function. Using the relation between correlation functions of different orders, one can calculate the first-order function, and then, using the relation between the correlation function and the power spectrum (Wiener-Khinchin theorem), determine the emission spectrum. Thus, by "counting photons," spectral measurements can be indirectly performed. This is widely used nowadays. As will be seen later, this procedure allows measurement of spectra of chaotic light with high resolution, complementing traditional spectral measurement methods (using spectral devices such as diffraction gratings, prisms, etc.).

Expression (5.43) can also be obtained using the coherent state representation. In this case,

$$\hat{\rho} = \int \cdots \int P(\{\alpha_k\}) |\{\alpha_k\}\rangle \langle \{\alpha_k\}| d^2 \{\alpha_k\},$$

where

$$P(\{\alpha_k\}) = \prod_k \frac{1}{\pi \bar{n}_k} e^{-\frac{|\alpha_k|^2}{\bar{n}_k}} = \prod_k P_k(\alpha_k).$$

The operator product to be averaged is

$$\hat{a}_{kI}^\dagger \hat{a}_{kII}^\dagger \hat{a}_{kIII} \hat{a}_{kIV}.$$

This leads to the integral

$$\begin{aligned} & Sp \left\{ \hat{\rho} \hat{a}_{kI}^\dagger \hat{a}_{kII}^\dagger \hat{a}_{kIII} \hat{a}_{kIV} \right\} = \\ & = \int \cdots \int d^2 \{\alpha_k\}, P(\{\alpha_k\}) (\alpha_{kI}^* \alpha_{kII}^* \alpha_{kIII} \alpha_{kIV}). \end{aligned}$$

Again, non-zero contributions come from paired products satisfying

$$k^I = k^{III} = k_1, \quad k^{II} = k^{IV} = k_2$$

or

$$k^I = k^{IV} = k_1, \quad k^{II} = k^{III} = k_2$$

and also

$$k^I = k^{II} = k^{III} = k^{IV} = k.$$

It is easy to show by integrating in polar coordinates $\alpha = r e^{i\theta}$ that the non-zero integrals are:

$$\int d^2 \alpha_{k_1} P(\alpha_{k_1}) \alpha_{k_1}^* \alpha_{k_1} \int d^2 \alpha_{k_2} P(\alpha_{k_2}) \alpha_{k_2}^* \alpha_{k_2} = \bar{n}_{k_1} \bar{n}_{k_2}.$$

In the third case, given

$$P(\alpha_k) = \frac{1}{\pi \bar{n}_k} e^{-\frac{r^2}{\bar{n}_k}}$$

we have

$$\begin{aligned} \int d^2 \alpha_k P(\alpha_k) (\alpha_k^* \alpha_k)^2 &= \int_0^\infty r dr \int_0^{2\pi} P(\alpha_k) r^4 d\theta = \\ &= \frac{2\pi}{\pi \bar{n}_k} \int_0^\infty r^4 dr e^{-\frac{r^2}{\bar{n}_k}} = \bar{n}_k^2 \int_0^\infty x^2 dx e^{-x} = 2\bar{n}_k^2. \end{aligned}$$

Here, the substitution $x = \frac{r^2}{\bar{n}_k}$, $dx = \frac{2r dr}{\bar{n}_k}$ is made. Thus, we arrive at the same results as previously obtained. Therefore, the final result (5.43) remains unchanged.

5.4 Higher-Order Coherence

First- and second-order coherence functions are special, although fundamental cases of coherence functions. The n th-order coherence function in the quantum case is defined as follows

$$G^{(n)}(x_1, x_2, \dots, x_n) = \frac{\langle \hat{E}^{(-)}(x_1) \hat{E}^{(-)}(x_2) \dots \hat{E}^{(-)}(x_n) \hat{E}^{(+)}(x_n) \hat{E}^{(+)}(x_{n-1}) \dots \hat{E}^{(+)}(x_1) \rangle}{\langle \hat{E}^{(-)}(x_1) \hat{E}^{(+)}(x_1) \rangle \dots \langle \hat{E}^{(-)}(x_n) \hat{E}^{(+)}(x_n) \rangle}. \quad (5.44)$$

Angle brackets denote quantum-mechanical averaging using the statistical operator (density matrix), x denotes the set of variables t, \vec{r} . The degree of n th-order coherence determines the counting rate in an experiment where n photons are registered in some way.

5.5 Photon Counting and Statistics

We have verified that the quantization of the optical field and the existence of light quanta (photons) preserve the general picture of interference phenomena if the observation is conducted long enough for the averaged picture to emerge. In our formulas, this corresponds to averaging using the density matrix. However, the existence of photons allows for new types of experiments based on photon counting and the investigation of their statistical patterns. These methods are known as photon counting methods. The essence is as follows: the studied light is fed to a photodetector connected to a counter that counts the number of photoelectrons recorded over a certain period. A shutter in front of the photodetector (or circuit lock) controls the counting duration. With the counter set to zero, the shutter opens for a time T , and the number of photoelectrons is recorded. After a time longer than the correlation time τ_c , everything repeats many times. Based on the measurements, one can determine $P_m(T)$ - the probability of registering m counts of photoelectrons over time T :

$$P_m(T) = \frac{N_m}{N}, \quad (5.45)$$

where N_m is the number of measurements in which m photoelectrons were recorded, N is the total number of measurements, which should be large. Obviously, it is assumed that the light flow is stationary. The obtained distribution contains information about the spectral properties of the light beams. The primary task here is to find the photon statistics from the photo-count statistics, which we can measure, required to obtain information about the properties of the light beams.

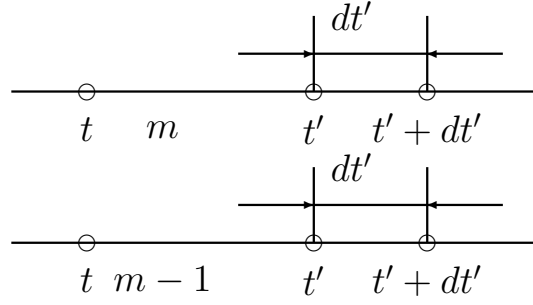


Figure 5.7: To the conclusion $P_m(t, t' + dt')$. There are two possibilities for obtaining m counts in the time interval $t, t' + dt'$: in the interval t, t' , m counts are made, and in the interval dt , none; in the interval t, t' , $m - 1$ counts are made, and in the interval dt' - one.

5.6 Connection of photon statistics with photoelectron count statistics

Let us start with a semiclassical consideration and then generalize the results to a fully quantum treatment. We have seen that the photoemission rate is proportional to the intensity of the light flux, which is determined by the mean value of the operator $\langle \hat{E}^{(-)} \hat{E}^{(+)} \rangle$. Since the quantum-mechanical intensity is analogous, in the classical case, to the intensity $I(t)$ averaged over a period, the electron emission rate can be considered proportional to $I(t)$ in the semiclassical approximation. Let $P(t) dt$ denote the probability of the appearance of a photoelectron in the time interval t to $t + dt$. Then, $P(t) dt = \xi I(t) dt$, where ξ characterizes the efficiency of the photocathode. Denote the probability of registering m photo counts in the time interval t to $t' + dt'$ by $P_m(t, t' + dt')$. There are two possibilities to obtain m counts in the given time interval, shown in Figure 5.7:

1. In the interval t to t' , m counts were made, and in the interval dt - none.
2. In the interval t to t' , $m - 1$ counts were made, and in the interval dt' - one count.

Note: the interval is so small that the probabilities of two or more counts are negligibly small. For these two cases, we can write

$$\begin{aligned} P_m^{(1)}(t, t' + dt') &= P_m(t, t') (1 - P(t') dt'), \\ P_m^{(2)}(t, t' + dt') &= P_{m-1}(t, t') P(t') dt', \end{aligned} \quad (5.46)$$

where $P(t) dt = \xi I(t) dt$. The total probability will be the sum of the probabilities of these two events:

$$\begin{aligned} P_m(t, t' + dt') &= P_m(t, t') (1 - P(t') dt') + \\ &\quad + P_{m-1}(t, t') P(t') dt'. \end{aligned} \quad (5.47)$$

From this, we get a chain of differential recurrence equations

$$\frac{dP_m}{dt'} = \xi I(t') \{P_{m-1}(t') - P_m(t')\}, \quad (5.48)$$

which can be solved by sequentially integrating, starting with $m = 0$. The initial equation has the form:

$$\frac{dP_0}{dt'} = -\xi I(t') P_0(t'), \quad (5.49)$$

with obvious initial conditions $P_0(t') = 1$ at $t' = t$. The solution of the equation under these initial conditions is

$$P_0(t, T) = e^{-\xi \int_t^{t+T} I(t') dt'}, \quad (5.50)$$

where T is the counting time. If we introduce the intensity averaged over the counting time

$$\bar{I}(t, T) = \frac{1}{T} \int_t^{t+T} I(t') dt',$$

then equation (5.50) can be rewritten as

$$P_0(t, T) = e^{-\xi \bar{I} T} \quad (5.51)$$

The other probabilities can be successively expressed through $P_0(t, T)$. By induction it is easy to show that

$$P_m(t, T) = \frac{(\xi \bar{I} T)^m}{m!} e^{-\xi \bar{I} T} \quad (5.52)$$

One can verify the validity of this solution by substituting it into the original equation (5.48). Generally, the intensity $\bar{I}(t, T)$ fluctuates from one counting interval to another. To account for this, an averaging over the ensemble of measurements needs to be done. The result is Mandel's formula:

$$P_m(T) = \langle P_m(t, T) \rangle = \left\langle \frac{(\xi \bar{I}(t, T) T)^m}{m!} e^{-\xi \bar{I}(t, T) T} \right\rangle. \quad (5.53)$$

Alternatively, this formula can be written as

$$P_m(T) = \langle P_m(t, T) \rangle = \int_0^\infty P(\bar{I}) \frac{(\xi \bar{I} T)^m}{m!} e^{-\xi \bar{I} T} d\bar{I}, \quad (5.54)$$

where $P(\bar{I})$ is the probability density for \bar{I} .

5.7 Distribution of Photo-Counts for Coherent and Chaotic Light

To understand how the statistics of photo-counts relate to the statistics of light incident on the photodetector, we apply formula (5.54) to two extreme cases: constant amplitude (intensity) light and chaotic light with fluctuating intensity. Let's start with the simplest case, when the intensity of the incident light is constant. In this case, $\bar{I}(t, T)$ does not depend on t and T , $\bar{I}(t, T) = I_0 = \text{const}$, so the second averaging is not needed. We get the same result if in (5.54) we assume $P(\bar{I}) = \delta(\bar{I} - I_0)$. The final expression for the case of light with constant intensity is:

$$P_m = \frac{\bar{m}^m}{m!} e^{-\bar{m}}. \quad (5.55)$$

where $\bar{m} = \xi I_0 T$ is the average number of photo-counts recorded by the photodetector over time T . Thus, for constant intensity, the distribution of photo-counts is a Poisson distribution. Note that for sufficiently large T (significantly larger than the correlation time τ_c), $\bar{I}(t, T)$ in any case¹ will tend to a constant value, and the distribution of photo-counts will become a Poisson distribution.

In another extreme case, when $T \ll \tau_c$, we can consider $\bar{I}(t, T) = I(t)$ - the instantaneous intensity. In the case of chaotic light, $P(I) = \frac{1}{I} e^{-\frac{1}{I}}$ [57]. By Mandel's formula, we have:

$$\begin{aligned} P_m(T) &= \frac{1}{\bar{I}} \int_0^\infty e^{-\frac{I}{\bar{I}}} \frac{(\xi I T)^m}{m!} e^{-\xi I T} dI = \\ &= \int_0^\infty \frac{y^m (\bar{I} \xi T)^m}{m! (1 + \bar{m})^{m+1}} e^{-y} dy. \end{aligned} \quad (5.56)$$

Here a change of variables is made:

$$\begin{aligned} y &= I \left(\frac{1}{\bar{I}} + \xi T \right), \\ I &= \frac{y \bar{I}}{1 + \xi \bar{I} T}, \\ \bar{m} &= \xi \bar{I} T. \end{aligned}$$

Finally, we obtain

$$P_m(T) = \frac{\bar{m}^m}{(1 + \bar{m})^{m+1} m!} \int_0^\infty y^m e^{-y} dy = \frac{\bar{m}^m}{(1 + \bar{m})^{m+1}} \quad (5.57)$$

¹including chaotic light

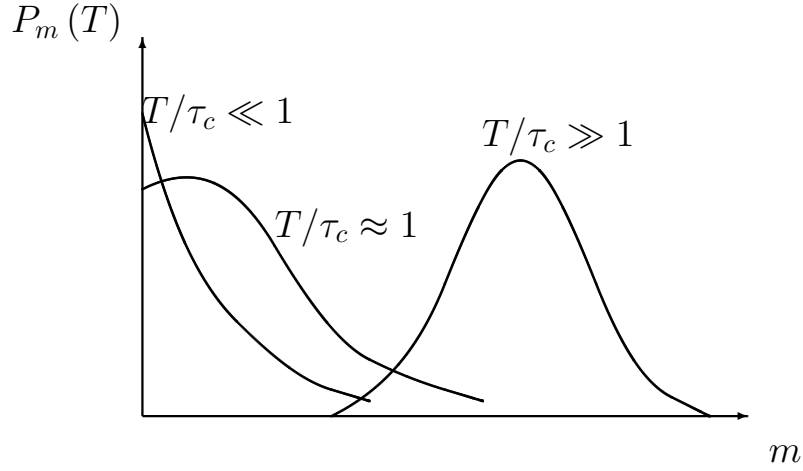


Figure 5.8: Qualitative view of the distribution of photodetections $P_m(T)$ for three values of T/τ_c .

since

$$\int_0^\infty y^m e^{-y} dy = m!.$$

We have found that the distribution of photo-counts follows the distribution of photons for a chaotic field, but with a modified scale. It is much more difficult to determine the distribution of photo-counts for the counting time T , corresponding to the intermediate case $T \approx \tau_c$. Here, only numerical calculations are possible. The results of such calculations are provided in [Figure 5.8](#) [57]. They allow us to understand the nature of the change in the distribution of photo-counts with an increasing ratio of T/τ_c . From the graphs, it is evident that the nature of the distribution changes near $T = \tau_c$. This allows, by conducting measurements at different T , to estimate τ_c and the width of the incident light spectrum $\sim 1/\tau_c$.

5.8 Determination of photon statistics through the distribution of photocounts

Usually, the main interest is the photon statistics characterizing the light beam. To determine it from the measured $P_m(T)$, one should invert (5.54), expressing $P(\bar{I})$ through $P_m(T)$.

There are several methods to find $P(\bar{I})$. (See [43] for inversion methods.) We will focus on one of them. Consider the expression (5.54)

$$P_m(T) = \int_0^\infty P(\bar{I}) \frac{(\xi \bar{I} T)^m}{m!} e^{-\xi \bar{I} T} d\bar{I}.$$

From this equality, given the measured $P_m(T)$, one needs to determine the probability density $P(\bar{I})$ that characterizes the statistics of the incident light.

Let us change the variable to $u = \xi \bar{I} T$ and define a new probability density

$$P_T(u) = \frac{1}{\xi T} P(\bar{I}).$$

We have:

$$P_T(u) du = \frac{1}{\xi T} P(\bar{I}) \xi T d\bar{I} = P(\bar{I}) d\bar{I}.$$

Hence:

$$\int_0^\infty P_T(u) du = \int_0^\infty P(\bar{I}) d\bar{I} = 1,$$

as it should be. Rewrite equation (5.54) using these variables:

$$P_m(T) = \int_0^\infty P_T(u) \frac{u^m}{m!} e^{-u} du. \quad (5.58)$$

Next, we use the orthogonal Laguerre polynomials

$$\begin{aligned} L_n(y) &= \sum_{k=0}^n \binom{n}{k} \frac{(-y)^k}{k!}, \\ \binom{n}{k} &= C_k^n = \frac{n!}{k!(n-k)!} \end{aligned} \quad (5.59)$$

which satisfy the orthogonality conditions

$$\int_0^\infty L_p(y) L_q(y) e^{-y} dy = \delta_{pq}, \quad (5.60)$$

which can be put into the form by substituting $y \rightarrow 2y$:

$$2 \int_0^\infty L_p(2y) L_q(2y) e^{-2y} dy = \delta_{pq}. \quad (5.61)$$

Now, expand $P_T(u)$ in a series of Laguerre polynomials:

$$P_T(u) = \sum_{n=0}^{\infty} A_n L_n(2u) e^{-u} \quad (5.62)$$

Using the orthogonality conditions (5.61), we get for the expansion coefficients A_n the expression

$$A_n = 2 \int_0^\infty L_n(2u) P_T(u) e^{-u} du. \quad (5.63)$$

Substituting $L_n(2u)$ in the form of the series (5.59) into (5.63), we obtain:

$$\begin{aligned} A_n &= 2 \sum_{k=0}^{k=n} \binom{n}{k} (-2)^k \int_0^\infty P_T(y) \frac{y^k}{k!} e^{-y} dy = \\ &= 2 \sum_{k=0}^{k=n} \binom{n}{k} (-2)^k P_k(T), \end{aligned} \quad (5.64)$$

where the expression

$$\int_0^\infty P(y) \frac{y^k}{k!} e^{-y} dy = P_k(T)$$

was used.

Assume that from measurements we know a sufficient number of values $P_k(T)$. This will allow us to compute a sufficient number of expansion coefficients A_n using formula (5.64). Substituting them into the series (5.62) for $P_T(u)$ and reverting to the original variables, we get

$$P(\bar{I}) = \xi T \sum_{k=0}^{\infty} A_n L_n(2\xi T \bar{I}) e^{-\xi T \bar{I}}. \quad (5.65)$$

Using this expression, it is in principle possible to invert Mandel's formula. The difficulty here lies in the fact that only a finite number of $P_n(T)$ are known with limited accuracy.

5.9 Quantum expression for the photocount distribution

The expression for the photocount distribution obtained above is based on the semiclassical approach. A fully quantum expression is given by [57]:

$$P_m(T) = Sp \left\{ \hat{\rho} \hat{N} \left[\frac{(\beta \hat{I}(t) T)^m}{m!} e^{-\beta \hat{I}(t) T} \right] \right\} \quad (5.66)$$

where

$$\begin{aligned} \hat{I}(t) &= \frac{1}{T} \int_t^{t+T} 2\varepsilon_0 c \hat{E}^{(-)}(t') \hat{E}^{(+)}(t') dt' = \\ &= \frac{2}{T} \int_t^{t+T} \sqrt{\frac{\varepsilon_0}{\mu_0}} \hat{E}^{(-)}(t') \hat{E}^{(+)}(t') dt' \end{aligned}$$

corresponds to the operator of the energy flux averaged over the counting period, β is the quantum efficiency of the photodetector. Formula (5.66) looks very similar to Mandel's formula. The difference is that here the classical fields are replaced by their operators, averaging is performed quantum mechanically using the statistical operator, and there is the normal ordering operator \hat{N} . The action of this operator reduces to rearranging the operators so that annihilation operators stand to the right of creation operators. When expanding the operator inside (5.66) in a power series, terms of the form $\left(\hat{E}^{(-)}\hat{E}^{(+)}\right)^n$ appear. The action of \hat{N} on them is given by the expression

$$\hat{N} \left(\hat{E}^{(-)}\hat{E}^{(+)}\right)^n = \left(\hat{E}^{(-)}\right)^n \left(\hat{E}^{(+)}\right)^n.$$

Such terms correspond to higher-order coherences, meaning the photocount distribution depends on coherences of all orders. Formula (5.66) can be justified as follows: Assume that the field is in a coherent state $|\{\alpha_k\}\rangle$. Then the probability of electron emission in the time interval dt is

$$Pdt = \sigma \langle \{\alpha_k\} | \hat{E}^{(-)}\hat{E}^{(+)} | \{\alpha_k\} \rangle dt = \sigma (E^* E) dt,$$

where σ characterizes the efficiency of the photocathode, and E is the classical field corresponding to the state $|\{\alpha_k\}\rangle$

$$\begin{aligned} \hat{E}^{(+)} |\{\alpha_k\}\rangle &= E |\{\alpha_k\}\rangle, \\ \langle \{\alpha_k\} | \hat{E}^{(-)} &= E^* \langle \{\alpha_k\} |. \end{aligned}$$

Further consideration proceeds in the same way as in the semiclassical approach. One obtains a similar formula:

$$P_m(T | \{\alpha_k\}) = \frac{(\sigma \overline{E^* E} T)^m}{m!} e^{-\sigma T \overline{E^* E}} \quad (5.67)$$

where

$$\overline{E^* E} = \frac{1}{T} \int_t^{t+T} E^*(t') E(t') dt'.$$

This expression can be written in operator form:

$$P_m(T | \{\alpha_k\}) = \langle \{\alpha_k\} | \hat{N} \left\{ \frac{(\sigma \overline{\hat{E}^{(-)} \hat{E}^{(+)}} T)^m}{m!} e^{-\sigma T \overline{\hat{E}^{(-)} \hat{E}^{(+)}}} \right\} | \{\alpha_k\} \rangle, \quad (5.68)$$

where $P_m(T | \{\alpha_k\})$ is the probability of counting m photoelectrons in time T if the field is in the state $|\{\alpha_k\}\rangle$. The operator inside the curly braces can be

considered as the operator counting m photoelectrons in time T . If the field is in a statistically mixed state with statistical operator, defined in the diagonal representation by the function $P_m(T|\{\alpha_k\})$,

$$\hat{\rho} = \int_{\{\alpha_k\}} P(\{\alpha_k\}) |\{\alpha_k\}\rangle \langle\{\alpha_k\}| \prod_k d^2\alpha_k,$$

the average value of the operator in expression (5.68), according to (1.122), will be

$$P_m(T) = \int_{\{\alpha_k\}} P(\{\alpha_k\}) \langle\{\alpha_k\}| \cdot \hat{N} \left\{ \frac{(\sigma \overline{\hat{E}^{(-)}} \hat{E}^{(+)} T)^m}{m!} e^{-\sigma T \overline{\hat{E}^{(-)}} \hat{E}^{(+)}} |\{\alpha_k\}\rangle \right\} \prod_k d^2\alpha_k \quad (5.69)$$

The obtained formula can be written in any representation, since the operation Sp does not depend on the representation. We have

$$P_m = Sp \left(\hat{\rho} \hat{N} \left\{ \frac{(\sigma \overline{\hat{E}^{(-)}} \hat{E}^{(+)} T)^m}{m!} e^{-\sigma T \overline{\hat{E}^{(-)}} \hat{E}^{(+)}} \right\} \right), \quad (5.70)$$

where

$$\overline{\hat{E}^{(-)}} \hat{E}^{(+)} = \frac{1}{T} \int_t^{t+T} \hat{E}^{(-)}(t') \hat{E}^{(+)}(t') dt'.$$

Formula (5.70) fully corresponds to formula (5.66) taken from [57]. Calculations using formula (5.70) are somewhat more complicated than using Mandel's formula. As an example, let's consider a few simple cases. First, consider single-mode states. Then (5.70) can be written as

$$P_m = Sp \left(\hat{\rho} \hat{N} \left\{ \frac{(\gamma \overline{\hat{a}^\dagger} \hat{a} T)^m}{m!} e^{-\gamma T \overline{\hat{a}^\dagger} \hat{a}} \right\} \right)$$

where γ is a coefficient characterizing the photocathode efficiency in this notation. The operator

$$\hat{N} \left\{ \frac{(\gamma \overline{\hat{a}^\dagger} \hat{a} T)^m}{m!} e^{-\gamma T \overline{\hat{a}^\dagger} \hat{a}} \right\}$$

has only diagonal elements in the number representation (photon number). Then $Sp(\dots)$ will be determined only by the diagonal elements of the density matrix, which we denote as $P_n = \langle n | \hat{\rho} | n \rangle$.

Then (5.70) can be represented as:

$$\begin{aligned}
P_m &= \sum_n P_n \langle n | \hat{N} \frac{(\gamma \hat{a}^\dagger \hat{a} T)^m}{m!} e^{-\gamma T \hat{a}^\dagger \hat{a}} | n \rangle = \\
&= \sum_n P_n \frac{(\gamma T)^m}{m!} \langle n | \sum_l (-1)^l \frac{(\gamma T)^l}{l!} (\hat{a}^\dagger)^{l+m} (\hat{a})^{l+m} | n \rangle = \\
&= \sum_{n=m} P_n \frac{(\gamma T)^m}{m!} \sum_{l=0}^{n-m} (-1)^l \frac{(\gamma T)^l}{l!} \frac{n!}{(n-m-l)!}, \tag{5.71}
\end{aligned}$$

which follows from the relation:

$$\begin{aligned}
&\langle n | (\hat{a}^\dagger)^{l+m} (\hat{a})^{l+m} | n \rangle = \\
&= \{n(n-1)(n-2)\dots(n-m-l+1)\} = \frac{n!}{(n-m-l)!}.
\end{aligned}$$

Expression (5.71) can be further transformed. The known expansion (Newton's binomial) is:

$$(1 - \gamma T)^{n-m} = \sum_{l=0}^{n-m} (-1)^l (\gamma T)^l \frac{(n-m)!}{l! (n-m-l)!}.$$

From here we have:

$$P_m(T) = \sum_{n=m}^{\infty} P_n \frac{n!}{m! (n-m)!} (\gamma T)^m (1 - \gamma T)^{n-m}. \tag{5.72}$$

Note that, as known from probability theory (see Bernoulli distribution), the general term of the sum gives the probability that out of n photons m are registered, and $n-m$ remain unregistered. Let's apply the general formula (5.71) to the case of chaotic light. Then

$$P_n = \frac{\bar{n}^n}{(\bar{n} + 1)^{n+1}},$$

and from (5.71) we get

$$\begin{aligned}
P_m(T) &= \sum_{n=m}^{\infty} \frac{\bar{n}^n}{(\bar{n} + 1)^{n+1}} (\gamma T)^m (1 - \gamma T)^{n-m} \frac{n!}{m! (n-m)!} = \\
&= (\gamma T)^m \sum_{l=0}^{\infty} \frac{\bar{n}^{l+m}}{(\bar{n} + 1)^{l+m+1}} (1 - \gamma T)^l \frac{(l+m)!}{m! l!}. \tag{5.73}
\end{aligned}$$

In the transformation, the summation index substitution is made: $l = n - m$, $n = l + m$.

The known expansion [43] is:

$$(1 - x)^{-(m+1)} = \sum_{l=0}^{\infty} \frac{x^l (m + l)!}{m! l!}.$$

Using it, we get:

$$\begin{aligned} P_m(T) &= \frac{(\gamma T \bar{n})^m}{(1 + \bar{n})^{m+1}} \left(1 - \frac{\bar{n}}{1 + \bar{n}} (1 - \gamma T) \right)^{-(m+1)} = \\ &= \frac{(\gamma T \bar{n})^m}{(1 + \bar{n})^{m+1}} \left\{ \frac{1 - \gamma T \bar{n}}{1 + \bar{n}} \right\}^{-(m+1)} = \frac{(\gamma T \bar{n})^m}{(1 + \gamma T \bar{n})^{m+1}} = \frac{\bar{m}^m}{(\bar{m} + 1)^{m+1}}, \end{aligned} \quad (5.74)$$

where $\bar{m} = \gamma T \bar{n}$. We obtained the same result as in the semiclassical consideration. The distribution $P_m(T)$ mirrors P_n , but at a different scale. Instead of \bar{n} , there is $\bar{m} = \gamma T \bar{n}$. Now consider another limiting case: the field is in a coherent state. It is convenient to consider it in the representation of coherent states. We have

$$\begin{aligned} P_m(T) &= \langle \alpha | \hat{N} \frac{(\gamma \hat{a}^\dagger \hat{a} T)^m}{m!} e^{-\gamma T \hat{a}^\dagger \hat{a}} | \alpha \rangle = \\ &= \frac{(\gamma |\alpha|^2 T)^m}{m!} e^{-\gamma |\alpha|^2 T} = \frac{\bar{m}^m}{m!} e^{-\bar{m}}, \end{aligned} \quad (5.75)$$

where $\bar{m} = \gamma |\alpha|^2 T = \gamma \bar{n} T$.

The photon number distribution for a coherent state is known (1.95):

$$P_n = \frac{\bar{n}^n e^{-\bar{n}}}{n!}.$$

Thus, in this case as well, the photocount distribution mirrors the photon distribution but at a different scale. Instead of \bar{n} , there is \bar{m} . In the general case, such a simple relation does not hold. From the derivation of the quantum photocount formula it follows that as long as $P(\{\alpha_k\})$ in the expression for the statistical operator can be interpreted as a probability distribution, there is no discrepancy between semiclassical results (Mandel's formula) and the fully quantum treatment. However, if $P(\{\alpha_k\})$ is negative in some regions, the results will differ.

5.10 Photon Counting Experiments. Application of Photon Counting Technique for Spectral Measurements

We have seen that the distribution of photocounts depends on the counting time T . When the counting time is compared with the correlation time, the nature of the distribution changes, and a peak appears on the curve. With further increase in counting time, the nature of the curve is preserved, and for large counting times, the distribution will tend towards the Poisson distribution.

Therefore, by measuring $P_m(T)$ at different T , one can estimate τ_c and the width of the light beam spectrum $\sim 1/\tau_c$. Such methods, where spectral parameters of light are determined from photon counting experiments, are called intensity fluctuation spectroscopy. The minimum counting time is determined by the resolution of the photodetector, which is about $10^{-8} - 10^{-9}$ seconds, corresponding to a frequency of $10^8 - 10^9$ Hz. This is the upper limit of the frequency changes of the method. The lower limit is determined by the maximum counting time, which is usually 1 second, corresponding to a resolution of 1 Hz.

Thus, the photon counting method can investigate the frequency interval from 1 Hz to 10^8 Hz. Consequently, this method complements conventional spectroscopy, which operates in the frequency range from 10^7 Hz to 10^{15} Hz. It is more convenient to use another type of photon counting experiment. In these, the correlation between the numbers of photons m_1 and m_2 , that is $\langle m_1 m_2 \rangle$, recorded in two short time intervals $\Delta t_1 = \Delta t_2 = \Delta t$, delayed relative to each other by time τ , is measured. Both intervals have the same duration Δt , less than τ , and the correlation time τ_c . In this case, the second-order coherence is measured, defined by the formula

$$G_{12}^{(2)} = \frac{\langle m_1 m_2 \rangle}{(\bar{m})^2} = 1 + \left(G_{12}^{(1)} \right)^2$$

where \bar{m} is the average number of counts over time Δt .

Knowing $G_{12}^{(2)}$ for chaotic light, one can calculate the first-order coherence function

$$G_{12}^{(1)} = \sqrt{G_{12}^{(2)} - 1}$$

and the form and width of the spectral line of the light beam related to it by Fourier transformation. For chaotic light with a Lorentzian spectral line, we have:

$$\langle m_1 m_2 \rangle = \bar{m}^2 \left(e^{-2\gamma|\tau|} + 1 \right), \quad (5.76)$$

and for a Doppler line -

$$\langle m_1 m_2 \rangle = \bar{m}^2 \left(e^{-\delta^2 \tau^2} + 1 \right). \quad (5.77)$$

Here δ and γ define the line widths in these cases. By measuring these parameters in the experiment, the spectral width of the lines can be determined.

5.11 Exercises

1. Show that the formula (5.52) indeed provides a solution to the problem.
2. Derive expression (5.57) from formula (5.56).
3. Derive formula (5.74) using the representation of coherent states.
4. Prove the orthogonality conditions for Laguerre polynomials (5.60) and (5.61).
5. Expand $P_T(u)$ in a series using Laguerre polynomials: (5.62)-(5.63)

Chapter 6

Quantum Description of Optical Interference Experiments

6.1 General Considerations

In quantum optics, as in classical optics, one of the main devices used for studying optical fields (light) is the interferometer.

In quantum physics, the device is described based on classical physics concepts, while the quantum object interacting with the device is described by quantum equations.

The boundary between the device and the quantum object is somewhat conditional and can be chosen depending on the problem to be solved.

In quantum optics, the interferometer itself is described by classical equations, just as in classical optics. The light field is described by quantum equations for field operators, usually using the Heisenberg picture.

The most important element of any interferometer is a mirror. According to the above, we consider the mirror as part of the device and describe it classically, introducing reflection coefficients r , r' and transmission coefficients t , t' (see [Figure 6.1](#)). The figure shows coefficients r , t and r' , t' corresponding to different sides of the mirror. Also shown are the annihilation operators of two input field modes \hat{a}_0 and \hat{a}_1 and two output mode operators \hat{a}_2 and \hat{a}_3 .

If the mirror is lossless, then the “scattering matrix” is unitary, as in the classical case:

$$\hat{S} = \begin{bmatrix} t' & r \\ r' & t \end{bmatrix},$$

where r , t and r' , t' are the transmission and reflection coefficients from different

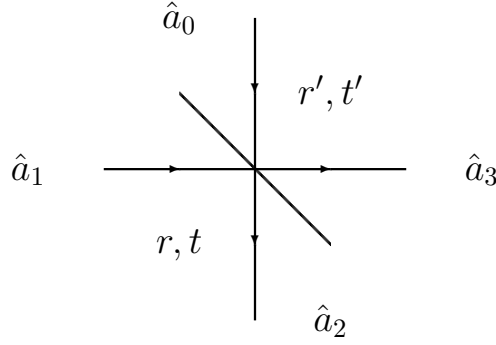


Figure 6.1: Mirror model, \hat{a}_0, \hat{a}_1 - annihilation operators of input field modes, \hat{a}_2, \hat{a}_3 - annihilation operators of output field modes, r, r', t , and t' - reflection and transmission coefficients from different sides of the mirror.

sides of the mirror.

The unitarity condition is written as follows:

$$\begin{aligned}
 \hat{S}^{-1} &= \hat{S}^\dagger = \begin{bmatrix} t'^* & r'^* \\ r^* & t^* \end{bmatrix}, \\
 \hat{S}\hat{S}^\dagger &= \begin{bmatrix} t' & r \\ r' & t \end{bmatrix} \begin{bmatrix} t'^* & r'^* \\ r^* & t^* \end{bmatrix} = \\
 &= \begin{bmatrix} |t'|^2 + |r|^2 & t'r'^* + rt^* \\ r't'^* + tr^* & |r'|^2 + |t|^2 \end{bmatrix} = \hat{I}, \\
 \hat{S}^\dagger\hat{S} &= \begin{bmatrix} t'^* & r'^* \\ r^* & t^* \end{bmatrix} \begin{bmatrix} t' & r \\ r' & t \end{bmatrix} = \\
 &= \begin{bmatrix} |t'|^2 + |r'|^2 & t'^*r + r'^*t \\ r^*t' + t^*r' & |r|^2 + |t|^2 \end{bmatrix} = \hat{I}
 \end{aligned} \tag{6.1}$$

For the unitarity condition (6.1) to hold, the energy conservation law and the reciprocity theorem must be satisfied. This leads to certain conditions that the coefficients r, r', t , and t' must satisfy:

$$\begin{aligned}
 |r| &= |r'|, \quad |t| = |t'|, \\
 |r|^2 + |t|^2 &= 1, \\
 r't'^* + tr^* &= 0, \\
 t'r^* + r't^* &= 0.
 \end{aligned} \tag{6.2}$$

It should be noted that the last expression is not independent. Indeed, assuming $t \neq 0$ (otherwise one can consider the relation for the reflection coefficients) and from the relation $|t| = |t'|$ we have $tt^* = t't'^*$ from which

$$t^* = \frac{t't'^*}{t}.$$

Therefore, using $r't'^* + tr^* = 0$, we have

$$\begin{aligned} t'r^* + r't^* &= t'r^* + r'\frac{t't'^*}{t} = \\ &= \frac{t'}{t}(tr^* + r't'^*) = 0. \end{aligned}$$

The reference phase plane can be shifted slightly with respect to the mirror plane, especially since a real mirror is not infinitely thin. In some cases, the reference plane can be moved by a distance of the order of the wavelength. By choosing the reference plane, the scattering matrix can be brought to a simpler form, with conditions (6.2) automatically satisfied. Assuming $t = t' = Te^{i\theta}$ and $r = r' = Re^{i\phi}$, from (6.2) we get

$$tr^* + r't^* = Tre^{i\theta}e^{-i\phi} + Tre^{i\phi}e^{-i\theta} = e^{i(\phi-\theta)}TR(e^{-2i(\phi-\theta)} + 1) = 0,$$

i.e., $\phi - \theta = \frac{\pi}{2}$. Taking $\theta = 0$ and redefining $t = T, r = R$ we have

$$\hat{S} = \begin{pmatrix} t & ir \\ ir & t \end{pmatrix}. \quad (6.3)$$

If we take $t, t', r, r' \in \mathbb{R}$, then obviously the necessary conditions will also be satisfied for $t' = t$ and $r' = -r$:

$$\hat{S} = \begin{pmatrix} t & r \\ -r & t \end{pmatrix}. \quad (6.4)$$

In equations (6.3) and (6.4), the coefficients r and t are real and positive.

In expression (6.3) the phase of the reflection coefficient is shifted by $\frac{\pi}{2}$ relative to the phase of the transmission coefficient. In the second variant (6.4), all coefficients are real, but the reflection coefficients from different sides have different signs. Since by choosing the reference plane the general case can be reduced either to (6.3) or to (6.4), we will subsequently use one of these expressions.

Example 6.1.1 (Symmetric Splitter. Classical Case). *Suppose our splitter is symmetric, i.e., $r = r'$ and $t = t'$ (see Figure 6.2). Obviously, the law of energy conservation must hold:*

$$|E_0|^2 + |E_1|^2 = |E_2|^2 + |E_3|^2,$$

or

$$\begin{aligned} |E_0|^2 + |E_1|^2 &= |tE_0 + rE_1|^2 + |rE_0 + tE_1|^2 = \\ &= \left(|t|^2 + |r|^2\right) \left(|E_0|^2 + |E_1|^2\right) + \\ &+ E_0E_1^*(tr^* + rt^*) + E_1E_0^*(tr^* + rt^*) \end{aligned}$$

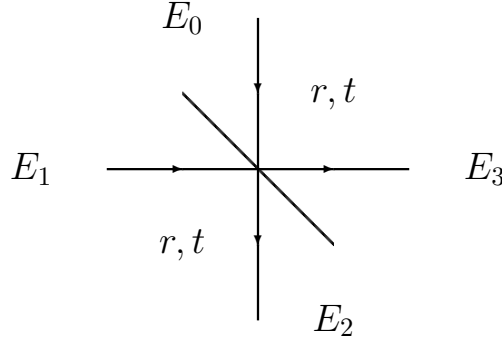


Figure 6.2: Symmetric splitter. Classical case. E_0, E_1 - two input light beams, E_2, E_3 - output fields, destruction operators of the output field modes, r and t - reflection and transmission coefficients.

That is,

$$\begin{aligned} |t|^2 + |r|^2 &= 1, \\ tr^* + rt^* &= 0 \end{aligned} \quad (6.5)$$

The last equality will hold if the phases of t and r differ by $\frac{\pi}{2}$, in particular if we take $t = |t|$, $r = i|r|$, which corresponds to expression (6.3).

Moreover, expressions (6.5) correspond to (6.1), indeed

$$\begin{aligned} \hat{S}\hat{S}^\dagger &= \begin{pmatrix} t & r \\ r & t \end{pmatrix} \begin{pmatrix} t^* & r^* \\ r^* & t^* \end{pmatrix} = \\ &= \begin{pmatrix} |t|^2 + |r|^2 & tr^* + rt^* \\ t^*r + r^*t & |t|^2 + |r|^2 \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} = \hat{I}. \end{aligned}$$

Thus the unitarity condition (6.1) represents the law of energy conservation.

6.2 Operator relations

We will consider the incident and transmitted fields as two-dimensional, see Figure 6.3. The operators \hat{a}_0 and \hat{a}_1 are operators of the incident field, corresponding to two different modes (zero and first). The operators of the transmitted field are \hat{a}_2 and \hat{a}_3 (second and third modes). Naturally, the operators must satisfy the commutation relations:

$$\begin{aligned} [\hat{a}_0, \hat{a}_0^\dagger] &= [\hat{a}_1, \hat{a}_1^\dagger] = [\hat{a}_2, \hat{a}_2^\dagger] = [\hat{a}_3, \hat{a}_3^\dagger] = 1, \\ [\hat{a}_0, \hat{a}_1^\dagger] &= [\hat{a}_2, \hat{a}_3^\dagger] = 0. \end{aligned} \quad (6.6)$$

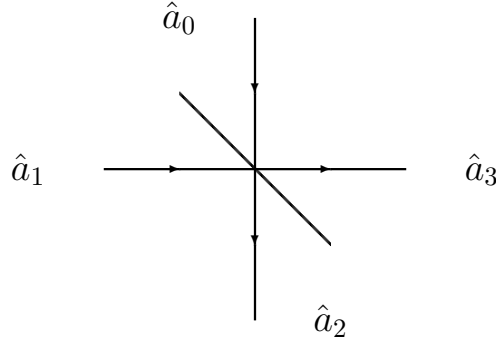


Figure 6.3: Mirror Model. Operators \hat{a}_0 and \hat{a}_1 are the operators of the incoming field related to two different modes (zero and first). Operators of the transmitted field \hat{a}_2 and \hat{a}_3 (second and third modes).

The relation between the modes arises upon reflection from the mirror and passing through it. It is described by the following equations:

$$\begin{aligned}\hat{a}_2 &= t'\hat{a}_0 + r\hat{a}_1, \\ \hat{a}_3 &= r'\hat{a}_0 + t\hat{a}_1.\end{aligned}\tag{6.7}$$

The incident and transmitted fields must satisfy the commutation relations (6.6), from which, taking into account (6.2), we have

$$\begin{aligned}[\hat{a}_2, \hat{a}_2^\dagger] &= [t'\hat{a}_0 + r\hat{a}_1, t'^*\hat{a}_0^\dagger + r^*\hat{a}_1^\dagger] = \\ &= |r|^2 + |t|^2 = 1.\end{aligned}\tag{6.8}$$

Similarly, we get

$$[\hat{a}_3, \hat{a}_3^\dagger] = 1\tag{6.9}$$

and further

$$[\hat{a}_2, \hat{a}_3^\dagger] = [t'\hat{a}_0 + r\hat{a}_1, r'\hat{a}_0^\dagger + t^*\hat{a}_1^\dagger] = t'r'^* + rt^* = 0.\tag{6.10}$$

Thus, we have obtained the fulfillment of all the required commutation relations. Note that if only mode 1 (input 1) is excited, then the field of mode zero (input 0) cannot be canceled as is done in the classical case. In the quantum case, the mode is in the vacuum state, and its field is not zero. If this is not taken into account, then the commutation relations (6.6) will be violated.

Let us assume that the mirror (beam splitter) shown in Figure 6.4 receives at input 1 (first mode) - a coherent state $|\alpha\rangle$, and at the zero input (zero mode) a vacuum state $|0\rangle$, i.e.

$$\begin{aligned}\hat{a}_0 |0\rangle &= 0, \\ \hat{a}_1 |\alpha\rangle &= \alpha |\alpha\rangle.\end{aligned}$$

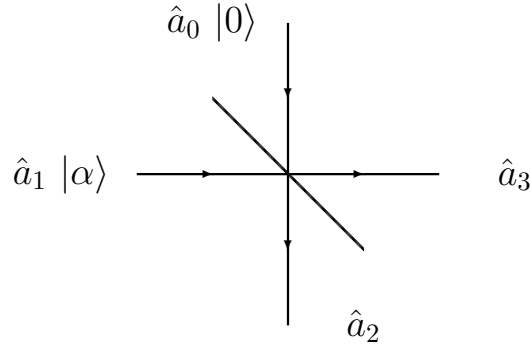


Figure 6.4: Mirror model. Input 1 (first mode) receives a coherent state $|\alpha\rangle$, and the zero input (zero mode) receives a vacuum state $|0\rangle$.

Thus, at the input, we have a two-mode state

$$|\psi\rangle = |0\rangle_0 |\alpha\rangle_1.$$

From (6.7) it follows

$$\hat{a}_3 |\psi\rangle = (r'\hat{a}_0 + t\hat{a}_1) |0\rangle_0 |\alpha\rangle_1 = t\alpha |\psi\rangle,$$

from which it is seen that the state at the output 3 is a coherent state with a reduced amplitude $\alpha_3 = t\alpha$. Similarly, we get

$$\hat{a}_2 |\psi\rangle = r\alpha |\psi\rangle, \quad \alpha_2 = r\alpha. \quad (6.11)$$

Thus, the change in amplitude α is the same as in the classical case, and the state remains coherent.

6.3 Mach-Zehnder Interferometer. Error of Phase Measurements.

6.3.1 Interferometer Equation

Let us now consider in more detail the operation of the Mach-Zehnder interferometer, the schematic of which is shown in Figure 6.5.

Assume that the input M_1 and output M_2 mirrors are partially transparent (50:50), i.e.

$$t = r = \frac{1}{\sqrt{2}},$$

and the other two mirrors are fully reflective

$$t = 0, \quad r = 1.$$

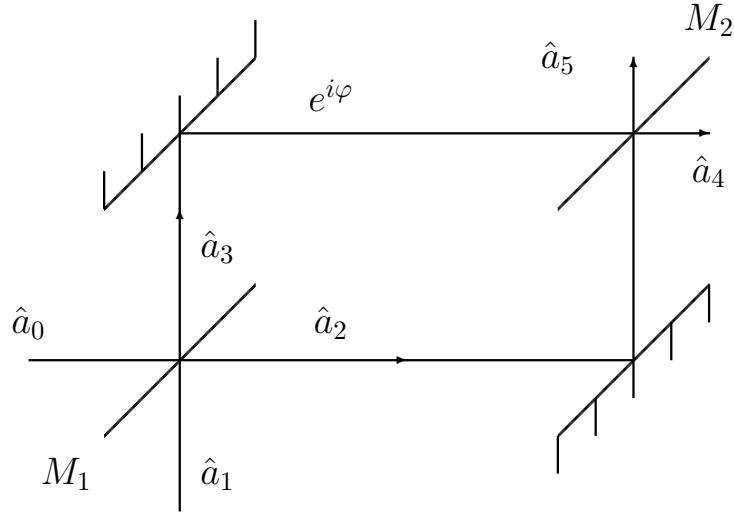


Figure 6.5: Mach-Zehnder interferometer. We assume that the input M_1 and output M_2 mirrors are semi-transparent (50:50), and the other two mirrors are opaque.

The “scattering matrix” of the beam splitters has the form

$$\hat{S} = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & i \\ i & 1 \end{pmatrix}.$$

Here we used the notation (6.3). The interferometer equation is:

$$\begin{aligned} \hat{a}_2 &= \frac{1}{\sqrt{2}} (\hat{a}_0 + i\hat{a}_1), \quad \hat{a}_3 = \frac{1}{\sqrt{2}} (i\hat{a}_0 + \hat{a}_1), \\ \hat{a}_4 &= \frac{1}{\sqrt{2}} (i\hat{a}_2 + e^{i\varphi}\hat{a}_3) = \frac{1}{2} [i(1 + e^{i\varphi})\hat{a}_0 - (1 - e^{i\varphi})\hat{a}_1], \\ \hat{a}_5 &= \frac{1}{\sqrt{2}} (\hat{a}_2 + ie^{i\varphi}\hat{a}_3) = \frac{1}{2} [(1 - e^{i\varphi})\hat{a}_0 + i(1 + e^{i\varphi})\hat{a}_1], \end{aligned} \quad (6.12)$$

where φ is the difference in optical path lengths of the interferometer arms (the phase difference accrued in the arms).

Formulas (6.12) connect the input fields with the output fields. Let us assume that the input field at port zero is in the vacuum state $|0\rangle$, and the input field at port one is in a coherent state $|\alpha\rangle$. Thus the input field is in a two-mode state $|\psi\rangle = |0\rangle_0 |\alpha\rangle_1$. Next, one can find the mean number of photons at outputs 4 and 5 depending on φ . The result will be indistinguishable from the classical one. Indeed,

$$\begin{aligned} \hat{a}_5^\dagger \hat{a}_5 &= \frac{1}{4} \left[(1 - e^{-i\varphi}) \hat{a}_0^\dagger - i(1 + e^{-i\varphi}) \hat{a}_1^\dagger \right] \\ &\quad \left[(1 - e^{i\varphi}) \hat{a}_0 + i(1 + e^{i\varphi}) \hat{a}_1 \right] = \\ &= \frac{1}{2} \left[(1 - \cos \varphi) \hat{a}_0^\dagger \hat{a}_0 + (1 + \cos \varphi) \hat{a}_1^\dagger \hat{a}_1 \right] - \frac{\sin \varphi}{2} \left[\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right] \end{aligned} \quad (6.13)$$

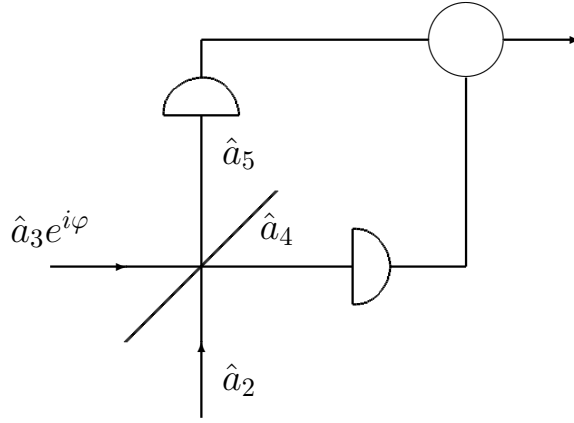


Figure 6.6: Balanced Detector

Thus, for the state $|\psi\rangle = |0\rangle_0 |\alpha\rangle_1$ the detector at port 5 produces the following signal

$$\langle\psi| \hat{a}_5^\dagger \hat{a}_5 |\psi\rangle = \frac{1}{2} (1 + \cos \varphi) |\alpha|^2.$$

For the photodetector at port 4 similarly we obtain

$$\begin{aligned} \hat{a}_4^\dagger \hat{a}_4 &= \frac{1}{4} \left[-i (1 + e^{-i\varphi}) \hat{a}_0^\dagger - (1 - e^{-i\varphi}) \hat{a}_1^\dagger \right] \\ &\quad \left[i (1 + e^{i\varphi}) \hat{a}_0 - (1 - e^{i\varphi}) \hat{a}_1 \right] = \\ &= \frac{1}{2} \left[(1 + \cos \varphi) \hat{a}_0^\dagger \hat{a}_0 + (1 - \cos \varphi) \hat{a}_1^\dagger \hat{a}_1 \right] + \frac{\sin \varphi}{2} \left[\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right] \end{aligned} \quad (6.14)$$

At photodetector 4 we have the following signal

$$\langle\psi| \hat{a}_4^\dagger \hat{a}_4 |\psi\rangle = \frac{1}{2} (1 - \cos \varphi) |\alpha|^2.$$

We consider a more complex balanced detection scheme, shown in [Figure 6.6](#). Each channel is detected by its own photodetector. The signals coming from the photodetectors are subtracted and recorded. This detection scheme essentially is a synchronous detector (homodyne detector), where the local oscillator is the field in the coherent state, and the signal consists of vacuum fluctuations.

Using [\(6.14, 6.13\)](#) one can calculate the signal at the input of the balanced detector, which is defined as the expectation value of the operator

$$\hat{R} = \hat{a}_5^\dagger \hat{a}_5 - \hat{a}_4^\dagger \hat{a}_4 = \left(\hat{a}_1^\dagger \hat{a}_1 - \hat{a}_0^\dagger \hat{a}_0 \right) \cos \varphi - \left(\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right) \sin \varphi. \quad (6.15)$$

In deriving [\(6.15\)](#) the relations [\(6.12\)](#) were used.

If the mode of the local oscillator (input 1, mode \hat{a}_1) is in a coherent state with a large amplitude α , while the signal mode (input 0, mode \hat{a}_0) is in the vacuum

state, then the output signal is

$$\begin{aligned}\langle \hat{R} \rangle &= \langle \psi | \hat{R} | \psi \rangle = \langle 0 |_0 \langle \alpha |_1 \left(\hat{a}_1^\dagger \hat{a}_1 - \hat{a}_0^\dagger \hat{a}_0 \right) | 0 \rangle_0 | \alpha \rangle_1 \cos \varphi - \\ &\quad - \langle 0 |_0 \langle \alpha |_1 \left(\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right) | 0 \rangle_0 | \alpha \rangle_1 = \\ &= |\alpha|^2 \cos \varphi.\end{aligned}\quad (6.16)$$

In deriving (6.16) it was taken into account that the operators of the first and zero modes act only on the state of their own mode. Therefore,

$$\hat{a}_0 | 0 \rangle = 0, \quad \langle 0 | \hat{a}_0^\dagger = 0,$$

and hence

$$\langle 0 |_0 \langle \alpha |_1 \left(\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right) | 0 \rangle_0 | \alpha \rangle_1 = 0,$$

while

$$\begin{aligned}\langle \psi | \hat{R} | \psi \rangle &= \langle 0 |_0 \langle \alpha |_1 \left(\hat{a}_1^\dagger \hat{a}_1 - \hat{a}_0^\dagger \hat{a}_0 \right) | 0 \rangle_0 | \alpha \rangle_1 = \\ &= \langle 0 |_0 \langle \alpha |_1 \left(\hat{a}_1^\dagger \hat{a}_1 \right) | 0 \rangle_0 | \alpha \rangle_1 = |\alpha|^2 = \langle \hat{n} \rangle\end{aligned}$$

is equal to the mean photon number in the local oscillator mode.

If $\varphi = \frac{\pi}{2}$, then $\langle \hat{R} \rangle = 0$ and the output signal is absent. At the same time, the operator $\hat{R} \neq 0$:

$$\hat{R} = - \left(\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right).$$

Although the mean of this operator is zero, the mean of its square will be nonzero and will describe the noise which limits the accuracy of measuring φ , which itself describes the difference in optical path lengths of the interferometer arms.

6.3.2 Measurement Accuracy of the Interferometer

First, let us find the mean square of the noise term:

$$\begin{aligned}\langle \psi | \hat{R} \hat{R}^\dagger | \psi \rangle &= \langle \psi | \left(\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 \right) \left(\hat{a}_0 \hat{a}_1^\dagger + \hat{a}_1 \hat{a}_0^\dagger \right) | \psi \rangle = \\ &= \langle \psi | \left(\hat{a}_0^\dagger \hat{a}_1 \hat{a}_0 \hat{a}_1^\dagger + \hat{a}_0^\dagger \hat{a}_1 \hat{a}_1 \hat{a}_0^\dagger + \hat{a}_1^\dagger \hat{a}_0 \hat{a}_0 \hat{a}_1^\dagger + \hat{a}_1^\dagger \hat{a}_0 \hat{a}_1 \hat{a}_0^\dagger \right) | \psi \rangle.\end{aligned}\quad (6.17)$$

Of the four terms in (6.17), the first three vanish upon averaging, while the last one is nonzero. For example, for the first term we have

$$\begin{aligned}\langle \psi | \hat{a}_0^\dagger \hat{a}_1 \hat{a}_0 \hat{a}_1^\dagger | \psi \rangle &= \langle 0 |_0 \langle \alpha |_1 \hat{a}_0^\dagger \hat{a}_1 \hat{a}_0 \hat{a}_1^\dagger | 0 \rangle_0 | \alpha \rangle_1 = \\ &= \langle 0 |_0 \langle \alpha |_1 \hat{a}_0^\dagger \hat{a}_0 \hat{a}_1 \hat{a}_1^\dagger | 0 \rangle_0 | \alpha \rangle_1 = \langle 0 |_0 \hat{a}_0^\dagger \hat{a}_0 | 0 \rangle_0 \langle \alpha |_1 \hat{a}_1 \hat{a}_1^\dagger | \alpha \rangle_1 = 0.\end{aligned}$$

The last term equals

$$\begin{aligned}
 \langle \psi | \hat{a}_1^\dagger \hat{a}_0 \hat{a}_1 \hat{a}_0^\dagger | \psi \rangle &= \langle 0 |_0 \langle \alpha |_1 \hat{a}_1^\dagger \hat{a}_0 \hat{a}_1 \hat{a}_0^\dagger | 0 \rangle_0 | \alpha \rangle_1 = \\
 &= \langle 0 |_0 \langle \alpha |_1 \hat{a}_1^\dagger \hat{a}_1 \hat{a}_0 \hat{a}_0^\dagger | 0 \rangle_0 | \alpha \rangle_1 = \langle \alpha |_1 \hat{a}_1^\dagger \hat{a}_1 | \alpha \rangle_1 \langle 0 |_0 \hat{a}_0 \hat{a}_0^\dagger | 0 \rangle_0 = \\
 &= \langle \alpha |_1 \hat{a}_1^\dagger \hat{a}_1 | \alpha \rangle_1 \langle 1 | 1 \rangle_0 = \langle \alpha |_1 \hat{a}_1^\dagger \hat{a}_1 | \alpha \rangle_1 = |\alpha|^2. \quad (6.18)
 \end{aligned}$$

Thus, we have found that the root mean square value of the noise is equal to the mean photon number in the mode that is in the coherent state:

$$|\bar{R}|^2 = \langle \hat{R} \hat{R}^\dagger \rangle = |\alpha|^2 = \bar{n}.$$

The standard deviation is

$$\sqrt{|\bar{R}|^2} = \sqrt{\bar{n}}.$$

The signal arises due to the variation

$$\varphi = \frac{\pi}{2} + \Delta\varphi.$$

Its magnitude is

$$\bar{R} = \bar{n} \cos\left(\frac{\pi}{2} + \Delta\varphi\right) \approx -\bar{n}\Delta\varphi.$$

The minimal distinguishable signal, i.e., the threshold signal, will be considered the signal equal to the noise standard deviation:

$$|\bar{n}\Delta\varphi| \geq \sqrt{|\bar{R}|^2} = \sqrt{\bar{n}},$$

from which it follows that

$$\Delta\varphi \geq \frac{1}{\sqrt{\bar{n}}}.$$

Equality corresponds to the threshold value of $\Delta\varphi$.

As we will see later (see ch. 9), the measurement accuracy can be significantly improved if the “squeezed vacuum” is supplied to the signal arm of the interferometer.

6.4 Exercises

1. Verify the validity of the commutation relations (6.8), (6.9), and (6.10).
2. Prove the validity of (6.11).
3. Derive the equation of the interferometer (6.12).

Chapter 7

Quantum Description of Polarization Properties of Light

7.1 Classical description of polarization properties of light

The field of an arbitrary monochromatic plane wave can be represented as follows:

$$\vec{E} = \vec{E}_0 e^{-i(\omega t - \vec{k}\vec{r})},$$

where the vector

$$\vec{E}_0 = E_x \vec{e}_x + E_y \vec{e}_y$$

defines the intensity and polarization properties of the electromagnetic radiation. For describing polarization properties, the complex polarization vector—the Jones vector—is used:

$$\vec{e} = \alpha \vec{e}_x + \beta \vec{e}_y. \quad (7.1)$$

The components of this vector, α and β , can be represented as points on a certain sphere called the Poincaré sphere. The coordinates of these points are determined by two angles θ and φ

$$\alpha = \frac{E_x}{|E_x|^2 + |E_y|^2} = \cos \frac{\theta}{2},$$

$$\beta = \frac{E_y}{|E_x|^2 + |E_y|^2} = e^{i\varphi} \sin \frac{\theta}{2}.$$

To measure the components of the Jones vector α and β , the scheme shown in [Figure 7.1](#) can be used. Here, a beam of polarized light is directed onto a Nicol

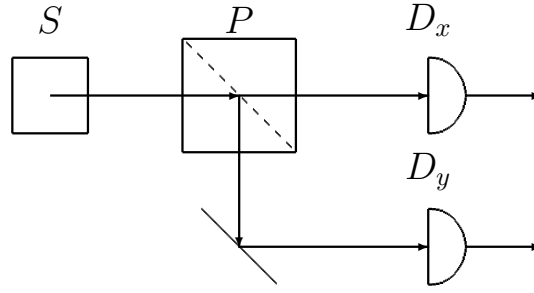


Figure 7.1: Diagram for measuring the parameters of the Jones vector. S - light source, P - Nicol prism, D_x and D_y - photodetectors.

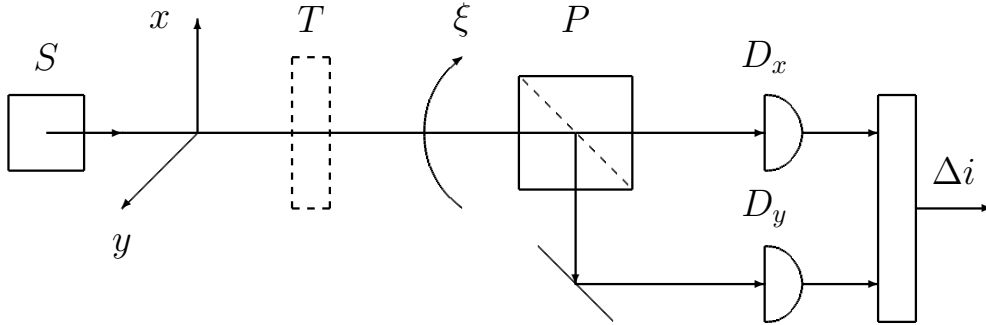


Figure 7.2: Scheme for measuring Stokes parameters. S - light source, P - Nicol prism, ξ - angle of its orientation relative to the x axis, equal to 0 when measuring S_0 and S_1 , $\frac{\pi}{4}$ - when measuring S_2 , T - additional quarter-wave plate $\frac{\lambda}{4}$ for measuring S_3 , D_x and D_y - photodetectors, Δi - difference in detector currents.

prism P , which separates the x and y components of this beam, which are then sent to two photodetectors D_x and D_y .

In addition to the Jones vector, the Stokes vector is often used to describe polarization properties. Its four components have the dimension of intensity and can be easily measured experimentally. The Stokes vector can be defined as follows:

$$\begin{aligned}
 S_0 &= |E_x|^2 + |E_y|^2, \\
 S_1 &= |E_x|^2 - |E_y|^2, \\
 S_2 &= E_x^* E_y + E_x E_y^* = 2\text{Re}(E_x^* E_y), \\
 S_3 &= \frac{E_x^* E_y - E_x E_y^*}{i} = 2\text{Im}(E_x^* E_y),
 \end{aligned} \tag{7.2}$$

where the amplitude values $E_{x,y}$ are taken at some moment in time t . The parameter S_0 in (7.2) defines the intensity of the wave at time t , and the other three parameters S_1 , S_2 , and S_3 describe the polarization properties.

To measure the Stokes parameters, the scheme shown in Figure 7.2 can be used [45]. To measure S_1 , S_2 , and S_3 , it is necessary to split the original beam into three parts, each of which is sent to a detector depicted in Figure 7.2.

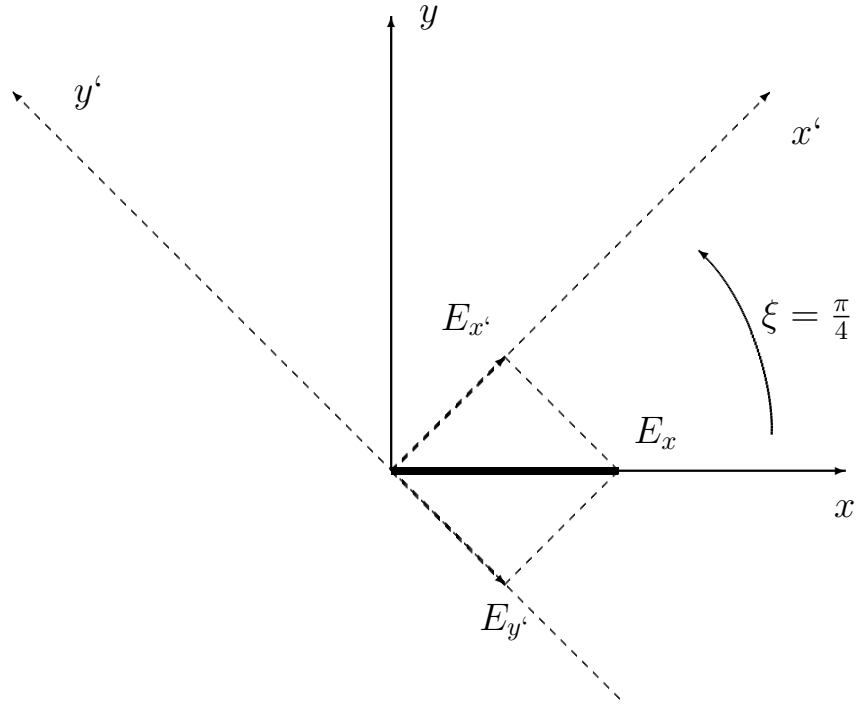


Figure 7.3: Polarization in a rotated coordinate system. $E_{x'} = \frac{E_y + E_x}{\sqrt{2}}$, $E_{y'} = \frac{E_y - E_x}{\sqrt{2}}$, thus $|E_{x'}|^2 - |E_{y'}|^2 = E_{x'}E_{x'}^* - E_{y'}E_{y'}^*$ or $|E_{x'}|^2 - |E_{y'}|^2 = \frac{|E_x|^2 + |E_y|^2 + E_x E_y^* + E_x^* E_y}{2} - \frac{|E_x|^2 + |E_y|^2 - E_x E_y^* - E_x^* E_y}{2}$, i.e. $|E_{x'}|^2 - |E_{y'}|^2 = E_x^* E_y + E_x E_y^* = 2\text{Re}(E_x^* E_y) = S_2$

To measure S_1 , a Nicol prism P is used, which separates the x and y polarized components of the beam, so that the difference of the currents of the two photodetectors is proportional to S_1 , while the sum of the currents is proportional to S_0 . To measure S_2 , the prism is rotated by an angle $\xi = \frac{\pi}{4}$, so that the difference of the currents is proportional to S_2 (see Figure 7.3). To measure S_3 , a quarter-wave plate $\frac{\lambda}{4}$ is placed before the prism with an orientation at $\frac{\pi}{4}$, resulting in the difference of the photodetector currents being proportional to S_3 .

The detector readings are averaged over time. Thus, one can speak of measuring time-averaged Stokes parameters - $\langle S_k \rangle$, which can also be used to describe partially polarized light. For quantitative characterization of such radiation, a quantity called the degree of polarization is used:

$$P = \frac{\sqrt{\langle S_1 \rangle^2 + \langle S_2 \rangle^2 + \langle S_3 \rangle^2}}{\langle S_0 \rangle}.$$

The degree of polarization of fully polarized light is $P = 1$. For completely unpolarized light, $P = 0$.

7.2 Quantum description of the polarization properties of light

When describing the polarization properties of a single-photon state, it is convenient to use the Jones vector, in which case the wave function will have the form:

$$|\psi\rangle = \alpha |x\rangle + \beta |y\rangle, \quad (7.3)$$

where $|\alpha|^2 + |\beta|^2 = 1$, and $|x\rangle = |1\rangle_x \otimes |0\rangle_y$ denotes a single-photon state polarized along x . Similarly, $|y\rangle = |0\rangle_x \otimes |1\rangle_y$ is a single-photon state polarized along y .

The measurable quantities for us will be the Stokes parameters, so the variables (7.2) must be replaced by operators. For this, we replace the electric field operators $E_{x,y}$ with annihilation operators $\hat{a}_{x,y}$, resulting in

$$\begin{aligned} \hat{S}_0 &= \hat{a}_x^\dagger \hat{a}_x + \hat{a}_y^\dagger \hat{a}_y, \\ \hat{S}_1 &= \hat{a}_x^\dagger \hat{a}_x - \hat{a}_y^\dagger \hat{a}_y, \\ \hat{S}_2 &= \hat{a}_x^\dagger \hat{a}_y + \hat{a}_x \hat{a}_y^\dagger, \\ \hat{S}_3 &= \frac{\hat{a}_x^\dagger \hat{a}_y - \hat{a}_x \hat{a}_y^\dagger}{i}. \end{aligned} \quad (7.4)$$

Further, we will be primarily interested in the two operators \hat{S}_1 and \hat{S}_2 , for which we will find the commutator, eigenvalues, and eigenvectors.

For the commutator $[\hat{S}_1, \hat{S}_2]$, we have:

$$\begin{aligned}
[\hat{S}_1, \hat{S}_2] &= \hat{S}_1 \hat{S}_2 - \hat{S}_2 \hat{S}_1 = \\
&= \hat{a}_x^\dagger \hat{a}_x \hat{a}_x^\dagger \hat{a}_y - \hat{a}_y^\dagger \hat{a}_y \hat{a}_y^\dagger \hat{a}_x + \\
&\quad + \hat{a}_x^\dagger \hat{a}_x \hat{a}_x \hat{a}_y^\dagger - \hat{a}_y^\dagger \hat{a}_y \hat{a}_y^\dagger \hat{a}_x - \\
&\quad - \hat{a}_x^\dagger \hat{a}_x^\dagger \hat{a}_x \hat{a}_y + \hat{a}_y \hat{a}_y^\dagger \hat{a}_y \hat{a}_x^\dagger - \\
&\quad - \hat{a}_x \hat{a}_x^\dagger \hat{a}_x \hat{a}_y^\dagger + \hat{a}_y^\dagger \hat{a}_y^\dagger \hat{a}_y \hat{a}_x = \\
&= \hat{a}_x^\dagger [\hat{a}_x \hat{a}_x^\dagger] \hat{a}_y - [\hat{a}_x \hat{a}_x^\dagger] \hat{a}_x \hat{a}_y^\dagger + \\
&\quad + [\hat{a}_y \hat{a}_y^\dagger] \hat{a}_y \hat{a}_x^\dagger - \hat{a}_y^\dagger [\hat{a}_y \hat{a}_y^\dagger] \hat{a}_x = \\
&= \hat{a}_x^\dagger \hat{a}_y + \hat{a}_y \hat{a}_x^\dagger - \hat{a}_x \hat{a}_y^\dagger - \hat{a}_y^\dagger \hat{a}_x = \\
&= 2 (\hat{a}_x^\dagger \hat{a}_y - \hat{a}_x \hat{a}_y^\dagger) = 2i\hat{S}_3 \neq 0.
\end{aligned} \tag{7.5}$$

In the derivation of (7.5), we used the commutation relations for creation and annihilation operators:

$$[\hat{a}_x, \hat{a}_x^\dagger] = [\hat{a}_y, \hat{a}_y^\dagger] = 1,$$

as well as the fact that operators acting on different polarization components x and y commute with each other.

To find the eigenvalues and eigenvectors of the operators \hat{S}_1 and \hat{S}_2 , it is convenient to represent them in matrix form. To do this, we will use the basis formed by the vectors $|x\rangle$ and $|y\rangle$:

$$\begin{aligned}
|x\rangle &= \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \\
|y\rangle &= \begin{pmatrix} 0 \\ 1 \end{pmatrix},
\end{aligned}$$

For the operator \hat{S}_1 , we have:

$$\begin{aligned}
\hat{S}_1 |x\rangle &= \hat{a}_x^\dagger \hat{a}_x |1\rangle_x \otimes |0\rangle_y - \hat{a}_y^\dagger \hat{a}_y |1\rangle_x \otimes |0\rangle_y = \\
&= \hat{a}_x^\dagger \hat{a}_x |1\rangle_x \otimes |0\rangle_y = |1\rangle_x \otimes |0\rangle_y = |x\rangle, \\
\hat{S}_1 |y\rangle &= \hat{a}_x^\dagger \hat{a}_x |0\rangle_x \otimes |1\rangle_y - \hat{a}_y^\dagger \hat{a}_y |0\rangle_x \otimes |1\rangle_y = \\
&= -\hat{a}_y^\dagger \hat{a}_y |0\rangle_x \otimes |1\rangle_y = -|0\rangle_x \otimes |1\rangle_y = -|y\rangle,
\end{aligned} \tag{7.6}$$

from which we obtain the following matrix representation:

$$\hat{S}_1 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \tag{7.7}$$

From (7.7) one can write the eigenvalue equation:

$$(1 - s)(1 + s) = 0,$$

from which two eigenvalues can be found, $s_1 = 1$ and $s_2 = -1$. As can be easily verified, the eigenvector for $s_1 = 1$ will be $|s_1\rangle = |x\rangle$. Indeed, from (7.6):

$$\hat{S}_1 |s_1\rangle = \hat{S}_1 |x\rangle = 1 \cdot |s_1\rangle. \quad (7.8)$$

For the second eigenvalue, the eigenvector will be $|s_2\rangle = |y\rangle$:

$$\hat{S}_1 |s_2\rangle = -|y\rangle = -1 \cdot |s_2\rangle. \quad (7.9)$$

For the operator \hat{S}_2 , we have:

$$\begin{aligned} \hat{S}_2 |x\rangle &= \hat{a}_x^\dagger \hat{a}_y |1\rangle_x \otimes |0\rangle_y + \hat{a}_y^\dagger \hat{a}_x |1\rangle_x \otimes |0\rangle_y = \\ &= \hat{a}_y^\dagger \hat{a}_x |1\rangle_x \otimes |0\rangle_y = |0\rangle_x \otimes |1\rangle_y = |y\rangle, \\ \hat{S}_2 |y\rangle &= \hat{a}_x^\dagger \hat{a}_y |0\rangle_x \otimes |1\rangle_y + \hat{a}_y^\dagger \hat{a}_x |0\rangle_x \otimes |1\rangle_y = \\ &= \hat{a}_x^\dagger \hat{a}_y |0\rangle_x \otimes |1\rangle_y = |1\rangle_x \otimes |0\rangle_y = |x\rangle. \end{aligned} \quad (7.10)$$

From (7.10) we obtain the following matrix representation of the operator \hat{S}_2 :

$$\hat{S}_2 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}. \quad (7.11)$$

Analogous to the operator \hat{S}_1 , from (7.10) and (7.11) one can find two eigenvalues $s_1 = 1$ and $s_2 = -1$. For the first eigenvalue, the eigenvector is

$$|s_1\rangle = \frac{1}{\sqrt{2}} (|x\rangle + |y\rangle), \quad (7.12)$$

and for the second

$$|s_2\rangle = \frac{1}{\sqrt{2}} (|x\rangle - |y\rangle). \quad (7.13)$$

For the operator \hat{S}_3 , we have:

$$\begin{aligned} \hat{S}_3 |x\rangle &= \frac{1}{i} \left(\hat{a}_x^\dagger \hat{a}_y |1\rangle_x \otimes |0\rangle_y - \hat{a}_y^\dagger \hat{a}_x |1\rangle_x \otimes |0\rangle_y \right) = \\ &= -\frac{1}{i} \hat{a}_y^\dagger \hat{a}_x |1\rangle_x \otimes |0\rangle_y = -\frac{1}{i} |0\rangle_x \otimes |1\rangle_y = -\frac{|y\rangle}{i} = i |y\rangle, \\ \hat{S}_3 |y\rangle &= \frac{1}{i} \left(\hat{a}_x^\dagger \hat{a}_y |0\rangle_x \otimes |1\rangle_y + \hat{a}_y^\dagger \hat{a}_x |0\rangle_x \otimes |1\rangle_y \right) = \\ &= \frac{1}{i} \hat{a}_x^\dagger \hat{a}_y |0\rangle_x \otimes |1\rangle_y = \frac{1}{i} |1\rangle_x \otimes |0\rangle_y = -i |x\rangle. \end{aligned} \quad (7.14)$$

Thus, from (7.14) we obtain the following matrix representation of the operator \hat{S}_3 :

$$\hat{S}_3 = \begin{pmatrix} 0 & i \\ -i & 0 \end{pmatrix}. \quad (7.15)$$

From (7.15), one can find that the eigenvalues are again $s_1 = 1$ and $s_2 = -1$. At the same time, the eigenstates of the operator \hat{S}_3 are the states with left and right circular polarizations:

$$\begin{aligned} |s_1\rangle &= |-\rangle = \frac{1}{\sqrt{2}} (|x\rangle - i|y\rangle), \\ |s_2\rangle &= |+\rangle = \frac{1}{\sqrt{2}} (|x\rangle + i|y\rangle). \end{aligned} \quad (7.16)$$

TBD: Add the connection with Pauli matrices

7.3 Exercises

1. TBD

Part 3

Non-Classical Light

Chapter 8

Nonclassical Light

Quantum states of light can be divided into two groups: states that asymptotically approach a classical state (classical light) as the intensity of the light increases, and states that do not possess this property. For example, among the states we considered earlier, the coherent state has a classical limit, while the number (Fock) state does not.

Earlier we defined the concept of the second-order coherence function $G^{(2)}$. For single-mode light (5.33)

$$\begin{aligned} G^{(2)} &= \frac{\langle \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \rangle}{\langle \hat{a}^\dagger \hat{a} \rangle^2} = \\ &= \frac{\langle \hat{a}^\dagger \hat{a} (\hat{a}^\dagger \hat{a} - 1) \rangle}{\langle \hat{a}^\dagger \hat{a} \rangle^2} = \frac{\langle n^2 \rangle - \langle n \rangle}{\langle n \rangle^2}, \end{aligned} \quad (8.1)$$

where the commutation relations $[\hat{a}, \hat{a}^\dagger] = 1$, $\hat{n} = \hat{a}^\dagger \hat{a}$ is the particle number operator, have been used.

In the classical case, $\langle \hat{a}^\dagger \hat{a} \rangle = |\alpha|^2$, where $|\alpha|^2 = I$ is the intensity of the field (the squared modulus of the field amplitude). Since classical quantities commute, formula (8.1) in the classical case can be replaced by

$$G_{\text{cl}}^{(2)} = \frac{\langle |\alpha|^4 \rangle}{\langle |\alpha|^2 \rangle^2} = \frac{\langle I^2 \rangle}{\langle I \rangle^2}, \quad (8.2)$$

where averaging is performed with the help of a positive definite classical distribution function $P(\alpha)$

$$\langle |\alpha|^{2n} \rangle = \int_0^\infty P(|\alpha|^2) |\alpha|^{2n} d^2\alpha, \quad n = 1, 2.$$

Consider the difference $G_{\text{cl}}^{(2)} - G_{\text{cl}}^{(1)} = G_{\text{cl}}^{(2)} - 1$, since for a single-mode field, as we know, $G_{\text{cl}}^{(1)} = 1$ always. From (8.2) it follows

$$\begin{aligned}
G_{\text{cl}}^{(2)} - G_{\text{cl}}^{(1)} &= G_{\text{cl}}^{(2)} - 1 = \\
&= \frac{\langle |\alpha|^4 \rangle}{\langle |\alpha|^2 \rangle^2} - 1 = \frac{\langle |\alpha|^4 \rangle - \langle |\alpha|^2 \rangle^2}{\langle |\alpha|^2 \rangle^2} = \\
&= \frac{\int_0^\infty P |\alpha|^4 d^2\alpha - 2 \langle |\alpha|^2 \rangle^2 + \langle |\alpha|^2 \rangle^2}{\langle |\alpha|^2 \rangle^2} = \\
&= \frac{\int_0^\infty P \left(|\alpha|^4 - 2 |\alpha|^2 \langle |\alpha|^2 \rangle + \langle |\alpha|^2 \rangle^2 \right) d^2\alpha}{\langle |\alpha|^2 \rangle^2} = \\
&= \frac{\int_0^\infty P \left(|\alpha|^2 - \langle |\alpha|^2 \rangle \right)^2 d^2\alpha}{\langle |\alpha|^2 \rangle^2} \geq 0.
\end{aligned} \tag{8.3}$$

From inequality (8.3) it follows that

$$G_{\text{cl}}^{(2)} \geq 1.$$

In the quantum case, due to operator noncommutativity, the situation will be different. Formula (8.1) can be rewritten in the following form:

$$\begin{aligned}
G^{(2)} &= \frac{\langle \hat{n}^2 \rangle - \langle \hat{n} \rangle}{\langle \hat{n} \rangle^2} = \\
&= 1 + \frac{\left(\langle \hat{n}^2 \rangle - \langle \hat{n} \rangle^2 \right) - \langle \hat{n} \rangle}{\langle \hat{n} \rangle^2} = 1 + \frac{\sigma^2 - \langle \hat{n} \rangle}{\langle \hat{n} \rangle^2},
\end{aligned} \tag{8.4}$$

where σ^2 is the variance, i.e., σ is the standard deviation.

From (8.4) it follows that $G^{(2)}$ can be either greater or less than 1, depending on whether σ^2 or $\langle \hat{n} \rangle$ is larger. In the classical case $G^{(2)} \geq 1$, so a criterion of nonclassicality can be taken as

$$G^{(2)} < 1. \tag{8.5}$$

From (8.4, 8.5) it follows that the condition for nonclassicality is $\sigma^2 - \langle \hat{n} \rangle < 0$, i.e., the variance must be less than the mean photon number. It is known that a



Figure 8.1: Examples of photodetection realizations corresponding to different statistical properties of light: a) bunching and super-Poissonian statistics, b) Poissonian statistics, c) anti-bunching and sub-Poissonian statistics (according to [46]).

photon flux for which $\sigma^2 - \langle \hat{n} \rangle = 0$, has Poissonian statistics (an example is the coherent state). The case $\sigma^2 - \langle \hat{n} \rangle < 0$ corresponds to a more regular flux, called sub-Poissonian. This is the case of photon antibunching. An example of a state for which this holds is, for instance, a state resulting from parametric scattering, which we will consider later. The case $\sigma^2 - \langle \hat{n} \rangle > 0$ corresponds to a less regular process where photons bunch (taken as the thermal excitation of light). Such a state is called super-Poissonian. All three cases are shown in Figure 8.1, where the timing of photodetection events is illustrated.

We conclude that nonclassical states correspond to the sub-Poissonian threshold. The number $G^{(2)}$ is simply related to other parameters characterizing photon number fluctuations: the variance and the Fano factor, which frequently appear in the literature.

$$\Phi = \frac{\sigma^2}{\langle \hat{n} \rangle},$$

$$G^{(2)} - 1 = \frac{\Phi - 1}{\langle \hat{n} \rangle}$$

The statistical nature of the experiment can be established from an experiment in which the second-order correlation function is measured (the Brown-Twiss experiment). The experimental setup is shown in Figure 8.2.

The adjustable delay τ is implemented by changing the length of one of the arms.

Figure 8.3 shows three curves obtained from the experiment. Curve (b) corresponds to the coherent state. Curve (a) corresponds to photon bunching and super-Poissonian statistics. Curve (c) corresponds to antibunching and sub-Poissonian statistics. Such a curve is characteristic of nonclassical light. The

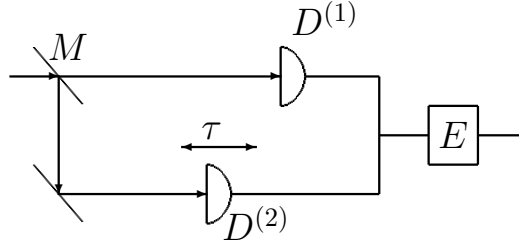


Figure 8.2: Diagram of the Brown-Twiss experiment. Using a 50% mirror M , the light beam is directed to two photodetectors $D^{(1)}$ and $D^{(2)}$. The adjustable delay τ is achieved by moving photodetector $D^{(2)}$. The result from the photodetectors is fed into the coincidence circuit E .

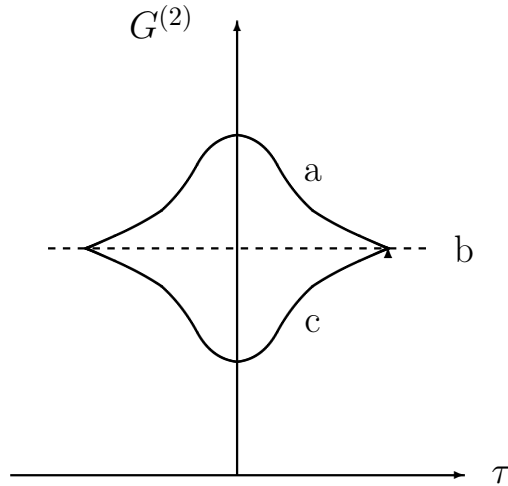


Figure 8.3: Qualitative view of the dependence of the correlation function $G^{(2)}$ on the delay τ for 3 different cases: a) bunching and super-Poissonian statistics, b) Poissonian statistics, c) antibunching and sub-Poissonian statistics.

explanation of the behavior of the curves is simple. If τ is large, photodetection events are random and independent of each other. Therefore, the behavior of all three curves is determined by random coincidences and will be the same for all three. For small τ the behavior of the curves differs. In the case of photon bunching, there is a maximum at $\tau = 0$, and for antibunching – a minimum.

Another criterion for nonclassicality is the absence of positive definiteness of the quasiprobability distribution $P(\alpha)$ when using the coherent state representation.

In the classical case, the function $P(\alpha)$ must be everywhere positive definite, i.e., $P(\alpha) \geq 0$ for all values of α . Let us consider this issue in more detail. The condition of nonclassicality $G^{(2)} < 1$, using (8.1), can be represented in the form

$$\langle \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \rangle - \langle \hat{a}^\dagger \hat{a} \rangle^2 < 0. \quad (8.6)$$

Consider the expression

$$\int P(\alpha) \left(|\alpha|^2 - \langle \hat{n} \rangle \right)^2 d^2\alpha < 0 \quad (8.7)$$

$$\begin{aligned} \int P(\alpha) \left(|\alpha|^2 - \langle \hat{n} \rangle \right)^2 d^2\alpha &= \int P(\alpha) \left(|\alpha|^4 - 2|\alpha|^2 \langle \hat{n} \rangle + \langle \hat{n} \rangle^2 \right) d^2\alpha = \\ &= \langle \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \rangle - \langle \hat{a}^\dagger \hat{a} \rangle^2 < 0, \end{aligned}$$

i.e., expressions (8.6) and (8.7) are equivalent. This follows from the following relations

$$\begin{aligned} \langle \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \rangle &= \text{Tr} (\hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \rho) = \\ &= \sum_n \langle n | \left(\hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} \int P(\alpha) |\alpha\rangle \langle \alpha| d^2\alpha \right) | n \rangle = \\ &= \int P(\alpha) \langle \alpha | \hat{a}^\dagger \hat{a}^\dagger \hat{a} \hat{a} | \alpha \rangle d^2\alpha = \int P(\alpha) |\alpha|^4 d^2\alpha, \\ &\int P(\alpha) |\alpha|^2 d^2\alpha = \langle \hat{a}^\dagger \hat{a} \rangle = \langle \hat{n} \rangle. \end{aligned}$$

We have found that expression (8.7) must be negative. Since $\left(|\alpha|^2 - \langle \hat{n} \rangle \right)^2$ is always positive, a negative result will be obtained if $P(\alpha)$ is negative in at least part of the domain of α . Thus, for nonclassical light, the quasiprobability $P(\alpha)$ is not a positive definite function.

Chapter 9

Squeezed States

In squeezed states, the possibility allowed by the Heisenberg uncertainty relation is realized, to reduce the uncertainty of one of the conjugate observables at the cost of increasing the uncertainty of the other. Such states may prove useful in quantum optics for practical needs. For example, for improving the accuracy of interference experiments.

The theory of squeezed states in parametric processes and the applications of squeezed states in interference experiments are further considered.

9.1 Heisenberg Uncertainty Principle

In quantum mechanics, two conjugate observables A and B correspond to non-commuting operators \hat{A} and \hat{B} , satisfying the commutation relation

$$[\hat{A}, \hat{B}] = \hat{A}\hat{B} - \hat{B}\hat{A} = i\hat{C},$$

where \hat{C} is some Hermitian operator. In this case, observables A and B can only be measured with a certain uncertainty, expressed by the Heisenberg uncertainty relation (for more details, see [subsection 14.3.1](#))

$$(\Delta A \Delta B) \geq \frac{1}{2} |\langle \hat{C} \rangle|, \quad (9.1)$$

where

$$(\Delta A)^2 = \langle \hat{A}^2 \rangle - \langle \hat{A} \rangle^2$$

and

$$(\Delta B)^2 = \langle \hat{B}^2 \rangle - \langle \hat{B} \rangle^2$$

are the mean square deviations from the average value. The average $\langle \dots \rangle$ corresponds to the quantum state in which the quantum object is found.

The relation (9.1) only restricts the product of uncertainties, allowing for states where the uncertainty of one of the observables is significantly less than that of the other. If

$$(\Delta A)^2 < \frac{1}{2} \left| \langle \hat{C} \rangle \right|,$$

such a state is called a squeezed state for the observable A . If at the same time the condition of minimal uncertainty is observed

$$(\Delta A \Delta B) = \frac{1}{2} \left| \langle \hat{C} \rangle \right|,$$

such a state is called perfectly squeezed.

9.2 Operators of Quadrature Components of the Electromagnetic Field

Consider the single-mode electromagnetic field operator

$$\hat{\vec{E}} = E_1 \vec{e} [\hat{a} e^{-i\omega t} + \hat{a}^\dagger e^{i\omega t}]$$

and introduce two new operators related to it

$$\begin{aligned} \hat{X}_1 &= \frac{1}{2} (\hat{a} + \hat{a}^\dagger), \\ \hat{X}_2 &= \frac{1}{2i} (\hat{a} - \hat{a}^\dagger), \end{aligned}$$

which are proportional to operators \hat{q} and \hat{p} (coordinate, momentum) that we encountered earlier.

The operators \hat{X}_1 and \hat{X}_2 are Hermitian operators. Their commutation relations are as follows.

$$\begin{aligned} [\hat{X}_1, \hat{X}_2] &= (\hat{X}_1 \hat{X}_2 - \hat{X}_2 \hat{X}_1) = \\ &= \frac{1}{4i} \{ (\hat{a} + \hat{a}^\dagger) (\hat{a} - \hat{a}^\dagger) - (\hat{a} - \hat{a}^\dagger) (\hat{a} + \hat{a}^\dagger) \} = \\ &= \frac{1}{4i} \{ \hat{a}^\dagger \hat{a} - \hat{a} \hat{a}^\dagger - \hat{a} \hat{a}^\dagger + \hat{a}^\dagger \hat{a} \} = \frac{1}{2i} \{ \hat{a}^\dagger \hat{a} - \hat{a} \hat{a}^\dagger \} = \\ &= \frac{1}{2i} [\hat{a}^\dagger, \hat{a}] = \frac{i}{2} \end{aligned}$$

Using the operators \hat{X}_1 and \hat{X}_2 , the electric field operator can be represented as

$$\vec{E} = 2E_1 \vec{e} (\hat{X}_1 \cos \omega t + \hat{X}_2 \sin \omega t),$$

i.e., these operators can be considered as operators of the quadrature components of the electromagnetic field. From the uncertainty relation (9.1) we have

$$(\Delta X_1 \Delta X_2) \geq \frac{1}{4}.$$

The condition for a squeezed state is

$$(\Delta X_i) < \frac{1}{4}, \quad i = 1, 2.$$

If at the same time the minimal uncertainty relation holds

$$(\Delta X_1 \Delta X_2) = \frac{1}{4},$$

then the squeezing is ideal.

As an example, consider the uncertainties of \hat{X}_1 , \hat{X}_2 in the case when the field is in a coherent state:

$$\begin{aligned} (\Delta X_1)^2 &= \langle \alpha | \hat{X}_1^2 | \alpha \rangle - \langle \alpha | \hat{X}_1 | \alpha \rangle^2 = \\ &= \frac{1}{4} \left\{ \langle \alpha | (\hat{a} + \hat{a}^\dagger)^2 | \alpha \rangle - \langle \alpha | (\hat{a} + \hat{a}^\dagger) | \alpha \rangle^2 \right\} = \\ &= \frac{1}{4} \left\{ \alpha^{*2} + \alpha^2 + 2|\alpha|^2 + 1 - (\alpha^{*2} + \alpha^2 + 2|\alpha|^2) \right\} = \frac{1}{4}. \end{aligned}$$

Similarly, one can obtain

$$(\Delta X_2)^2 = \frac{1}{4} \tag{9.2}$$

and

$$(\Delta X_1 \Delta X_2) = \frac{1}{4}. \tag{9.3}$$

Thus, we have found that the considered case is a state with minimal uncertainty, but there is no squeezing.

Now consider what the energy eigenstate $|n\rangle$ gives. We have

$$\begin{aligned} (\Delta X_1)^2 &= \frac{1}{4} \left\{ \langle n | (\hat{a} + \hat{a}^\dagger)^2 | n \rangle - \langle n | (\hat{a} + \hat{a}^\dagger) | n \rangle^2 \right\} = \\ &= \frac{1}{4} \langle n | (\hat{a} + \hat{a}^\dagger)^2 | n \rangle = \frac{1}{4} \langle n | (\hat{a}\hat{a}^\dagger + \hat{a}^\dagger\hat{a}) | n \rangle = \\ &= \frac{1}{4} \langle n | (1 + 2\hat{a}^\dagger\hat{a}) | n \rangle = \frac{1}{4} (2n + 1). \end{aligned}$$

Similarly,

$$(\Delta X_2)^2 = \frac{1}{4} (2n + 1) \tag{9.4}$$

and

$$(\Delta X_1 \Delta X_2) = \frac{1}{4} (2n + 1). \quad (9.5)$$

Therefore, the energy eigenstate is neither a state with minimal uncertainty nor a squeezed state for the observables X_1 or X_2 .

9.3 Squeezed Quadrature State

First, let's consider the theoretical possibility of creating squeezed quadrature states. Consider constructing a squeezed state from a coherent state, in which one of the quadrature components will be squeezed. The state transformation is performed using a certain unitary transformation.

Consider the unitary operator known as the squeezing operator (the reason for this name will become clear later)

$$\hat{S}(z) = e^{\frac{1}{2}z^* \hat{a}^2 - \frac{1}{2}z \hat{a}^{\dagger 2}},$$

where $z = re^{i\theta}$ is an arbitrary complex number. The conjugate operator has the form

$$\hat{S}^\dagger(z) = e^{\frac{1}{2}z \hat{a}^{\dagger 2} - \frac{1}{2}z^* \hat{a}^2},$$

from which it follows that

$$\hat{S}\hat{S}^\dagger = \hat{S}^\dagger\hat{S} = \hat{I},$$

where \hat{I} is the identity operator. Thus, $\hat{S}^\dagger = \hat{S}^{-1}$ is indeed a unitary operator.

Consider the action of the operator $\hat{S}(z)$ on the operators \hat{a} and \hat{a}^\dagger and on the coherent state. For this purpose, we apply the operator decomposition theorem (see detailed derivation in the theorem on operator identity (Theorem 17.1.1)):

$$e^{\hat{A}}\hat{B}e^{-\hat{A}} = \hat{B} + [\hat{A}, \hat{B}] + \frac{1}{2!} [\hat{A}, [\hat{A}, \hat{B}]] + \dots \quad (9.6)$$

Using (9.6), it is easy to show, by taking

$$e^{\hat{A}} = \hat{S}(z), \quad \hat{B} = \hat{a},$$

that

$$\begin{aligned} \hat{A} &= \hat{S}(z) \hat{a} \hat{S}^\dagger(z) = \hat{a} + z \hat{a}^\dagger + \frac{|z|^2 \hat{a}}{2!} + \frac{z |z|^2 \hat{a}^\dagger}{3!} + \dots = \\ &= \hat{a} \cosh r + \hat{a}^\dagger e^{i\theta} \sinh r = \mu \hat{a} + \nu \hat{a}^\dagger, \end{aligned} \quad (9.7)$$

where $\mu = \cosh r$, $\nu = e^{i\theta} \sinh r$, $|\mu|^2 - |\nu|^2 = 1$. Similarly, it can be shown that for the reverse order of operators $\hat{S}^\dagger \dots \hat{S}$ the formula differs from (9.7) only by the sign between the terms:

$$\begin{aligned}\hat{S}^\dagger(z) \hat{a} \hat{S}(z) &= \hat{a} \cosh r - \hat{a}^\dagger e^{i\theta} \sinh r, \\ \hat{S}^\dagger(z) \hat{a}^\dagger \hat{S}(z) &= \hat{a}^\dagger \cosh r - \hat{a} e^{-i\theta} \sinh r.\end{aligned}\quad (9.8)$$

Let's act with the operator \hat{S} on the coherent state vector. Then we obtain a new state:

$$|\alpha, z\rangle = \hat{S}(z) |\alpha\rangle. \quad (9.9)$$

Now show that the state $|\alpha, z\rangle$ is an eigenstate of the operator \hat{A} :

$$\begin{aligned}\hat{A} |\alpha, z\rangle &= \hat{S}(z) \hat{a} \hat{S}^\dagger(z) \hat{S}(z) |\alpha\rangle = \hat{S}(z) \hat{a} |\alpha\rangle = \\ &= \alpha \hat{S}(z) |\alpha\rangle = \alpha |\alpha, z\rangle.\end{aligned}\quad (9.10)$$

Here the definition of \hat{A} and $\hat{S} \hat{S}^\dagger = \hat{I}$ were used. From (9.10) it follows that the state $|\alpha, z\rangle$ is an eigenstate of the operator \hat{A} , and the eigenvalue coincides with the eigenvalue of the coherent state from which the state $|\alpha, z\rangle$ is obtained. The conjugate equality is

$$\langle\alpha, z| \hat{A}^\dagger = \alpha^* \langle\alpha, z| \quad (9.11)$$

Based on relations (9.10) and (9.11), by analogy with the operators \hat{a} and \hat{a}^\dagger , \hat{A} and \hat{A}^\dagger are called quasiparticle creation and quasiparticle annihilation operators.

It remains to determine whether the state $|\alpha, z\rangle$ is a squeezed state. To this end, the uncertainty relation for ΔX_1 and ΔX_2 should be written. We have

$$(\Delta X_{1,2})^2 = \langle\alpha, z| \hat{X}_{1,2}^2 |\alpha, z\rangle - \langle\alpha, z| \hat{X}_{1,2} |\alpha, z\rangle^2.$$

This expression, using (9.9), can be rewritten as

$$(\Delta X_{1,2})^2 = \langle\alpha| \hat{S}^\dagger(z) \hat{X}_{1,2}^2 \hat{S}(z) |\alpha\rangle - \langle\alpha| \hat{S}^\dagger(z) \hat{X}_{1,2} \hat{S}(z) |\alpha\rangle^2. \quad (9.12)$$

The operators $\hat{X}_{1,2}$ can be expressed through \hat{a} and \hat{a}^\dagger :

$$\begin{aligned}\hat{X}_1 &= \frac{1}{2} (\hat{a} + \hat{a}^\dagger), \\ \hat{X}_2 &= \frac{1}{2i} (\hat{a} - \hat{a}^\dagger).\end{aligned}$$

Substituting this into (9.12) and transforming the operators \hat{a} and \hat{a}^\dagger using (9.8), we obtain the final expressions for ΔX_1 and ΔX_2 : (under the condition $z = r$,

i.e. $\theta = 0$)

$$\begin{aligned}(\Delta X_1)^2 &= \frac{1}{4}e^{-2r}, \\(\Delta X_2)^2 &= \frac{1}{4}e^{2r}, \\(\Delta X_1 \Delta X_2) &= \frac{1}{4},\end{aligned}\tag{9.13}$$

where $r = |z|$ is the squeezing parameter. From (9.13) it follows that the state $|\alpha, z\rangle$ is indeed a squeezed state for one of the quadrature components. Since the uncertainty product has the minimum value, the state is an ideally squeezed state. We will call this state the ideally squeezed quadrature state.

If the operator $\hat{S}(z)$ acts on the vacuum state ($\alpha = 0$), we obtain the squeezed vacuum state (squeezed vacuum)

$$\hat{S}(z)|0\rangle = |z, 0\rangle.$$

The average photon number in the quadrature squeezed coherent state is given by the expression:

$$\begin{aligned}\langle\alpha, z|\hat{a}^\dagger\hat{a}|\alpha, z\rangle &= \langle\alpha|\hat{S}^\dagger\hat{a}^\dagger\hat{a}\hat{S}|\alpha\rangle = \\&= \langle\alpha|\hat{S}^\dagger\hat{a}^\dagger\hat{S}\hat{S}^\dagger\hat{a}\hat{S}|\alpha\rangle = \\&= \langle\alpha|(\hat{a}^\dagger \cosh r - \hat{a}e^{-i\theta} \sinh r)(\hat{a} \cosh r - \hat{a}^\dagger e^{i\theta} \sinh r)|\alpha\rangle = \\&= (|\alpha|^2 (\cosh^2 r + \sinh^2 r) - (\alpha^*)^2 e^{i\theta} \sinh r \cosh r - \alpha^2 \sinh r \cosh r e^{-i\theta} + \sinh^2 r)\end{aligned}\tag{9.14}$$

In deriving (9.14) we used relation (9.8).

For the squeezed vacuum state, $\alpha = 0$. Then the average photon number in this state is

$$\langle 0, z|\hat{a}^\dagger\hat{a}|0, z\rangle = \sinh^2 r \neq 0,\tag{9.15}$$

where r is the squeezing parameter.

Obviously, the squeezed vacuum state is not a true vacuum state, since it can be detected by a photodetector using the photoelectric effect, which is impossible for a true vacuum state.

Visually, quadrature squeezed states are shown in Figures 9.1-9.3, where initial states and time evolution of oscillations in various squeezed states are presented. The graphs on the right depict the uncertainty regions, which have the shape of ellipses. Such an image could be seen if we had something like an optical stroboscopic oscilloscope, which in reality does not exist.

Let's examine the obtained states in more detail. The formula (9.9) can be written somewhat differently by using the representation of the coherent state via

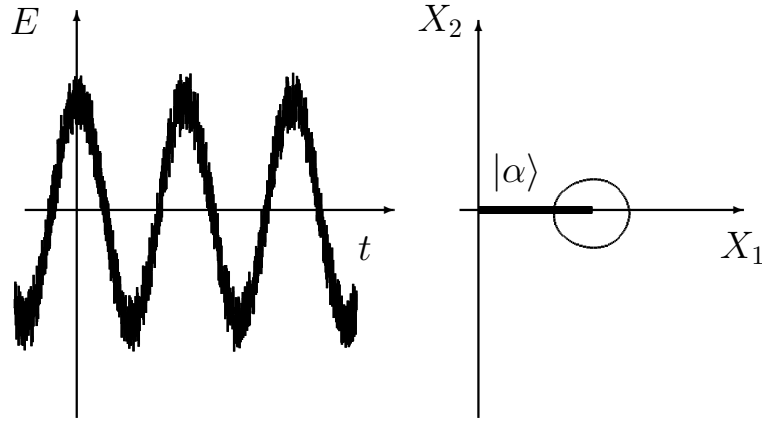


Figure 9.1: Uncertainty region and the corresponding dependence of the electric field on time for a coherent state. The complex number α , corresponding to the considered state $|\alpha\rangle$ on the complex plane defined by X_1 and X_2 lies along the X_1 axis. In this case, the number α is real, the average value $\langle\alpha|\hat{X}_1|\alpha\rangle = \frac{\alpha+\alpha^*}{2} = \text{Re } \alpha = \alpha$. Thus, the uncertainty ΔX_1 can be interpreted as amplitude uncertainty, and ΔX_2 as phase uncertainty. In the case of the coherent (not squeezed) state depicted in the figure, the phase and amplitude uncertainty are the same.

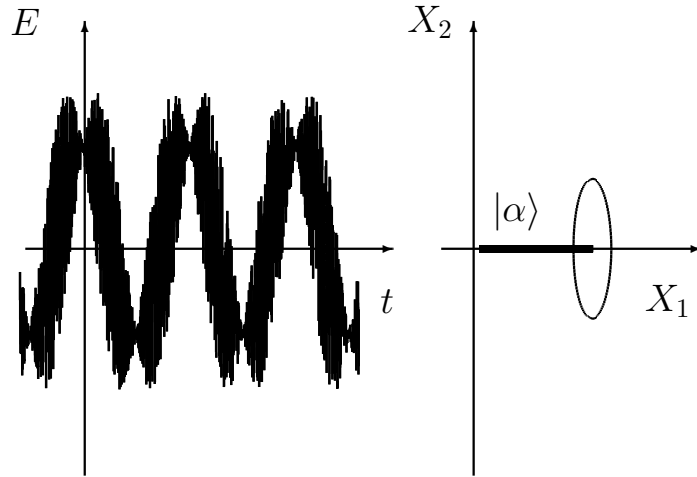


Figure 9.2: Region of uncertainty and the corresponding dependence of the electric field on time for a coherent squeezed state in X_1 . The uncertainty in amplitude is minimal, allowing for highly accurate measurements of electric field intensity peaks.

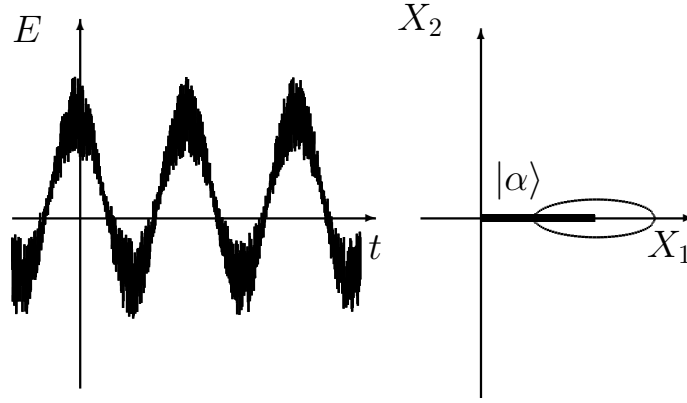


Figure 9.3: Uncertainty area and corresponding dependence of the electric field on time for a coherent squeezed state by X_2 . The phase uncertainty is minimal, allowing for high precision in measuring the oscillation period: time values at which the electric field intensity is zero.

the vacuum state in the form (1.100):

$$|\alpha\rangle = e^{\alpha\hat{a}^\dagger - \alpha^*\hat{a}} |0\rangle = \hat{D} |0\rangle,$$

where $\hat{D} = e^{\alpha\hat{a}^\dagger - \alpha^*\hat{a}}$ is called the displacement operator. \hat{D} is a unitary operator, since

$$\hat{D}^\dagger = e^{-(\alpha\hat{a}^\dagger - \alpha^*\hat{a})},$$

from which

$$\hat{D}^\dagger \hat{D} = \hat{D} \hat{D}^\dagger = \hat{I},$$

i.e. $\hat{D}^\dagger = \hat{D}^{-1}$.

The name of the operator is related to the fact that the action of $\hat{D}(\alpha)$ on the operators \hat{a} and \hat{a}^\dagger leads to their displacement by the amount α (α^*):

$$\begin{aligned} \hat{D}^\dagger(\alpha) \hat{a} \hat{D}(\alpha) &= \hat{a} + \alpha, \\ \hat{D}^\dagger(\alpha) \hat{a}^\dagger \hat{D}(\alpha) &= \hat{a}^\dagger + \alpha^*, \end{aligned} \quad (9.16)$$

This result follows from the operator decomposition theorem (9.6) when taking $\hat{A} = -\alpha\hat{a}^\dagger + \alpha^*\hat{a}$, $B = \hat{a}, \hat{a}^\dagger$.

Thus, the coherent state $|\alpha\rangle$ is a vacuum state displaced by α (see Figure 9.4). Based on the above, a squeezed state can be considered as a result of two actions: displacement of the vacuum state and its subsequent squeezing, as shown in Figure 9.5.

The squeezing procedure considered here is not unique. One can apply another sequence of operations – first squeeze the vacuum state, then displace it to obtain a new squeezed state (see Figure 9.6). In operator form, this operation can be written as:

$$|\alpha, z\rangle = \hat{D}(\alpha) \hat{S}(z) |0\rangle. \quad (9.17)$$

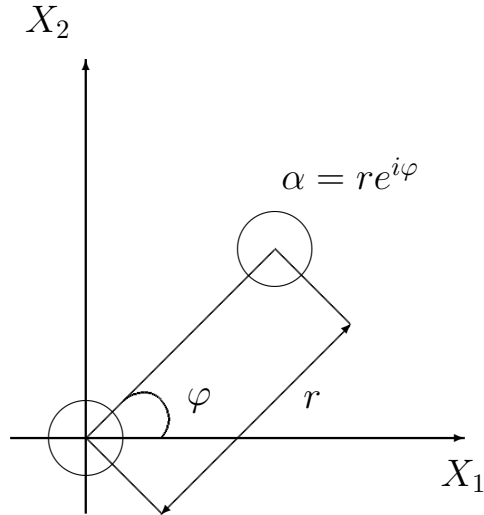


Figure 9.4: Coherent state as a displaced vacuum state. Uncertainties of observables \hat{X}_1 and \hat{X}_2 are: $\Delta X_1 = \Delta X_2 = \frac{1}{2}$.

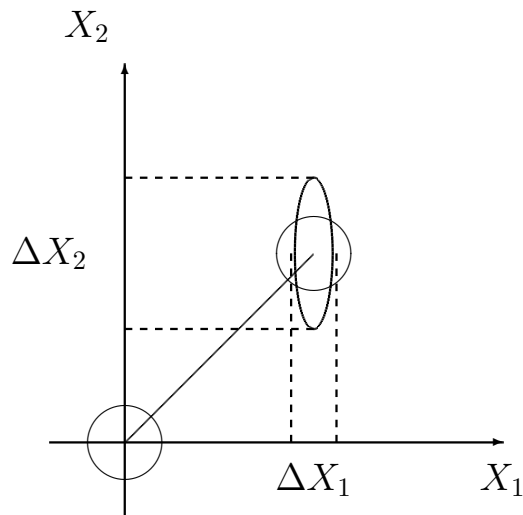


Figure 9.5: State squeezed by X_1 as a result of two actions: displacement of the vacuum state followed by its squeezing.

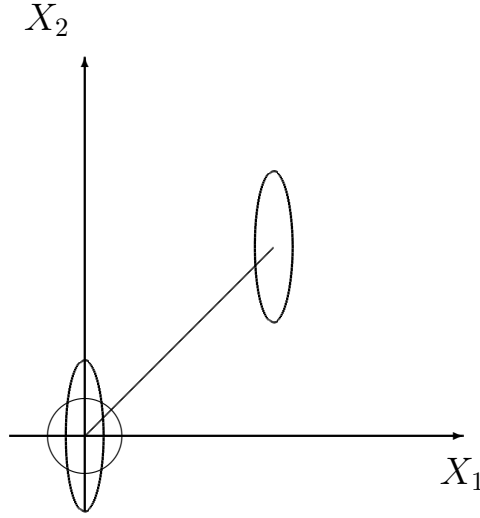


Figure 9.6: Squeezed state as a result of squeezing the vacuum state followed by its displacement.

To distinguish the states, the parameter order here is reversed compared to

$$|z, \alpha\rangle = \hat{S}(z) \hat{D}(\alpha) |0\rangle.$$

The state (9.17) also corresponds to a quadrature squeezed state, but they are not entirely identical. We will not consider this case in detail here.

The above results and illustrations shown in Figure 9.1-9.3 were obtained for the case when parameters $z = re^{i\theta}$ and $\alpha = |\alpha|e^{i\varphi}$ are real ($\varphi = \theta = 0$). If the parameters are complex, the picture changes somewhat (see Figure 9.7): the squeezing axis is rotated by an angle $\frac{\theta}{2}$ relative to the X_1, X_2 axes, and the position of the uncertainty region center is rotated by an angle φ relative to X_1 . By changing the phase of the light beams used in the generation and detection process, we can move to new quadratures Y_1, Y_2 as shown in Figure 9.7. If we further compensate the phase shift φ , we return to the situation depicted in Figures 9.1-9.3. From Figure 9.7, it is clear that by measuring the angles θ and φ it is possible, for example, to transform a state squeezed in X_1 into a state squeezed in X_2 , and vice versa.

9.4 Generation of Quadrature-Squeezed States

So far, we have considered the theoretical possibility of the existence and generation of squeezed states. The question arises — in what real processes can one obtain light in a squeezed state? Here we consider the possibility of obtaining a quadrature squeezed state through parametric interaction occurring when a strong pump field passes through a medium with quadratic nonlinearity. In such

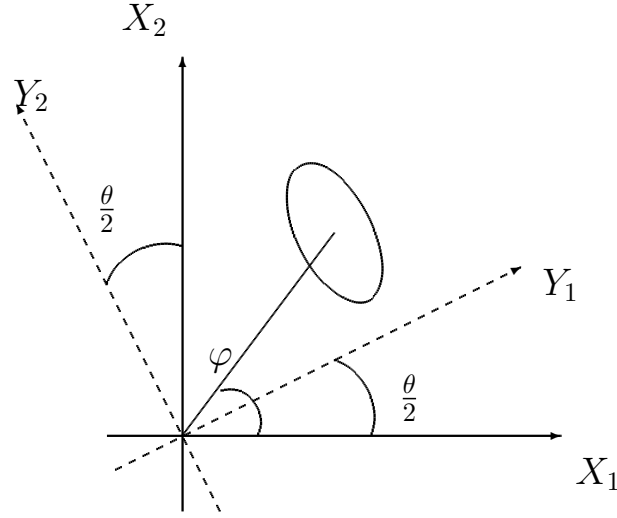


Figure 9.7: Squeezed coherent state. The squeezing axis is rotated by an angle $\frac{\theta}{2}$ relative to the axes X_1, X_2 , and the position of the uncertainty region's center is rotated by an angle φ relative to X_1 .

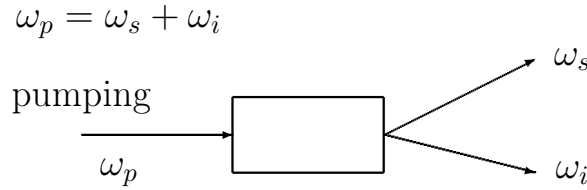


Figure 9.8: Generation of a squeezed state in parametric interaction

an interaction, two waves arise, related to the pump frequency by the relations:

$$\omega_p = \omega_s + \omega_i,$$

where ω_p is the pump frequency, ω_s is the signal frequency, ω_i is the idler frequency. This is schematically shown in [Figure 9.8](#).

The interaction Hamiltonian between the pump, signal, and idler waves in the interaction picture is given by:

$$\hat{V} = \hbar\kappa \left(\hat{a}_s^\dagger \hat{a}_i^\dagger \hat{b} + \hat{a}_s \hat{a}_i \hat{b}^\dagger \right), \quad (9.18)$$

where \hat{a}_s^\dagger , \hat{a}_i^\dagger , \hat{a}_s , and \hat{a}_i are creation and annihilation operators for the signal and idler waves, and κ is the interaction constant.

In the degenerate regime, the signal and idler waves have the same frequency equal to half the pump frequency (see [Figure 9.9](#)):

$$\omega_s = \omega_i = \frac{\omega_p}{2} = \omega.$$

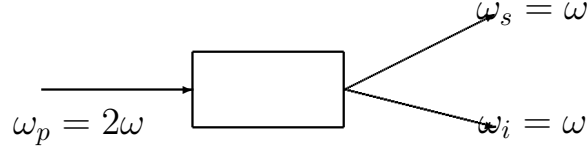


Figure 9.9: Generation of squeezed state in degenerate parametric interaction

For the degenerate parametric process, the interaction Hamiltonian (9.18) simplifies to:

$$\hat{V} = \hbar\kappa \left((\hat{a}^\dagger)^2 \hat{b} + \hat{a}^2 \hat{b}^\dagger \right).$$

If the pump field is in a coherent state with large amplitude $\alpha_p = A_p e^{i\varphi}$, $|\alpha_p| \gg 1$, the pump can be replaced by a classical field. Then the interaction Hamiltonian further simplifies to:

$$\hat{V} = \hbar\kappa A_p \left((\hat{a}^\dagger)^2 e^{-i\varphi} + \hat{a}^2 e^{i\varphi} \right),$$

where A_p and φ are the real amplitude and phase of the pump. Obviously, in this approximation, we neglect pump depletion. This is valid as long as the amplitude of the signal (idler) wave remains small compared to the pump field amplitude. The Heisenberg equation for the operator \hat{a} takes the form:

$$\begin{aligned} \frac{d\hat{a}}{dt} &= \frac{i}{\hbar} [\hat{V}, \hat{a}] = i\kappa A_p \left[(\hat{a}^\dagger)^2 e^{-i\varphi} + \hat{a}^2 e^{i\varphi}, \hat{a} \right] = \\ &= i\kappa A_p e^{-i\varphi} \left((\hat{a}^\dagger)^2 \hat{a} - \hat{a} (\hat{a}^\dagger)^2 \right) = \\ &= i\kappa A_p e^{-i\varphi} \left((\hat{a}^\dagger)^2 \hat{a} - (\hat{a}^\dagger \hat{a} + 1) \hat{a}^\dagger \right) = \\ &= i\kappa A_p e^{-i\varphi} (\hat{a}^\dagger \hat{a}^\dagger \hat{a} - \hat{a}^\dagger \hat{a} \hat{a}^\dagger - \hat{a}^\dagger) = \\ &= i\kappa A_p e^{-i\varphi} \hat{a}^\dagger (\hat{a}^\dagger \hat{a} - \hat{a} \hat{a}^\dagger - 1) = i\kappa A_p e^{-i\varphi} \hat{a}^\dagger ((-1) - 1) = \\ &= -2i\kappa A_p e^{-i\varphi} \hat{a}^\dagger. \end{aligned}$$

Thus we have

$$\frac{d\hat{a}}{dt} = -i\Omega_p e^{-i\varphi} \hat{a}^\dagger, \quad (9.19)$$

where

$$\Omega_p = 2\kappa A_p$$

is the effective Rabi frequency (energy expressed through frequency). Similarly,

$$\frac{d\hat{a}^\dagger}{dt} = i\Omega_p e^{i\varphi} \hat{a}. \quad (9.20)$$

Eliminating \hat{a}^\dagger from the system (9.19) and (9.20), we get

$$\frac{d^2 \hat{a}(t)}{dt^2} = -i\Omega_p e^{-i\varphi} \frac{d\hat{a}^\dagger(t)}{dt} = \Omega_p^2 \hat{a}(t). \quad (9.21)$$

The initial conditions for this equation are

$$\begin{aligned} \hat{a}(0) &= \hat{a}_0, \\ \hat{a}^\dagger(0) &= \hat{a}_0^\dagger, \end{aligned}$$

from which

$$\begin{aligned} \hat{a}|_{t=0} &= \hat{a}_0, \\ \left. \frac{d\hat{a}}{dt} \right|_{t=0} &= -i\Omega_p e^{-i\varphi} \hat{a}_0^\dagger, \end{aligned} \quad (9.22)$$

where \hat{a}_0 and \hat{a}_0^\dagger are the initial values of the operators at $t = 0$. The solution of (9.21) with the initial conditions (9.22) is:

$$\hat{a}(t) = \hat{a}_0 \operatorname{ch}(\Omega_p t) - i\hat{a}_0^\dagger \operatorname{sh}(\Omega_p t) e^{-i\varphi}. \quad (9.23)$$

Similarly,

$$\hat{a}^\dagger(t) = \hat{a}_0^\dagger \operatorname{ch}(\Omega_p t) + i\hat{a}_0 \operatorname{sh}(\Omega_p t) e^{i\varphi}. \quad (9.24)$$

Note that at $\varphi = \frac{\pi}{2}$ expressions (9.23) and (9.24) reduce to the relations (9.8):

$$\begin{aligned} \hat{a}(t) &= \hat{a}_0 \operatorname{ch}(\Omega_p t) - \hat{a}_0^\dagger \operatorname{sh}(\Omega_p t), \\ \hat{a}^\dagger(t) &= \hat{a}_0^\dagger \operatorname{ch}(\Omega_p t) - \hat{a}_0 \operatorname{sh}(\Omega_p t). \end{aligned}$$

As we see, the operator $\hat{a}(t)$ results from a unitary transformation (9.8) which in our case has the form:

$$\hat{a}(t) = \hat{S}^\dagger(t) \hat{a}(0) \hat{S}(t) = \hat{a}_0 \operatorname{ch}(\Omega_p t) - \hat{a}_0^\dagger \operatorname{sh}(\Omega_p t). \quad (9.25)$$

Thus, moving from the Heisenberg picture (14.36) to the Schrödinger picture (14.34), one can conclude that the initial coherent state undergoes the same unitary transformation as the annihilation operator (9.25):

$$\hat{S}(t) |\alpha\rangle = |t, \alpha\rangle,$$

i.e., the parametric interaction is equivalent to the action of the squeezing operator. In this case, for the quadrature components

$$\hat{X}_1 = \frac{\hat{a} + \hat{a}^\dagger}{2}$$

and

$$\hat{X}_2 = \frac{\hat{a} - \hat{a}^\dagger}{2i}$$

a squeezed state is realized with uncertainties

$$\begin{aligned} (\Delta X_1)^2 &= \frac{1}{4}e^{-2\Omega_p t}, \\ (\Delta X_2)^2 &= \frac{1}{4}e^{2\Omega_p t}, \\ (\Delta X_1 \Delta X_2) &= \frac{1}{4}. \end{aligned}$$

The squeezing parameter $r = \Omega_p t$ grows unbounded in time and, therefore, the degree of squeezing increases indefinitely. In reality, this does not happen. Unlimited squeezing is a consequence of overly idealized assumptions: pump depletion, deviation of the real pump from a monochromatic classical light, and several other factors were neglected. All these lead to limitations on the degree of squeezing. The closer the real conditions to ideal, the higher the degree of squeezing that can be obtained.

The case considered above of obtaining squeezed states via parametric processes pertains to the generation of squeezed vacuum, since the generation of the squeezed state occurs from the always-present vacuum seed field. This is sometimes called parametric scattering of the pump field. As we have found, the "squeezed vacuum" is not a vacuum field in the strict sense. The mean photon number in this field is not zero but depends on the degree of squeezing according to formula (9.15) and can be quite large at high degrees of squeezing.

9.5 Observation of the squeezed state. Measurement of the squeezing degree

In quadrature squeezing of light, the fluctuations of one of the quadrature components can be significantly smaller than the other. If this component is selected and used for measurement purposes, one can expect a strong reduction of noise that limits measurement accuracy.

Selection can be carried out using a homodyne (synchronous) detector. The simplest scheme of a homodyne detector is shown in Figure 9.10. The beam splitter has the following parameters: the transmission coefficient is close to 1 ($t \approx 1$) and, consequently, $r \ll 1$, since $|t|^2 + |r|^2 = 1$. This is necessary so that the signal \hat{a}_0 is not significantly attenuated. At the same time, the amplitude of the local oscillator, whose field is in a coherent state $|\alpha_H\rangle$, is so large that the local oscillator field intensity strongly exceeds the signal field. For the homodyne detector to

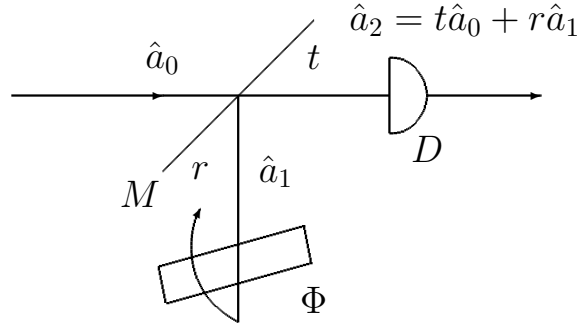


Figure 9.10: Scheme of the homodyne detector. M - mirror with reflection coefficient r and transmission coefficient t . D - detector. Φ - phase shifter.

operate, the local oscillator field and signal must have the same frequency. This is achieved by forming the signal and local oscillator fields from the same specified laser. For example, if \hat{a}_0 corresponds to the squeezed vacuum, the pumping of the parametric generator is done by the second harmonic of the master laser $2\omega_p$. During degenerate parametric interaction, a signal with frequency ω_p is generated, which coincides with the frequency of the local oscillator ω_p . Thus, the condition of frequency equality is satisfied.

The field operator incident on the photodetector \hat{a}_2 can be expressed through the operators \hat{a}_0 and \hat{a}_1 — the signal field and local oscillator field operators:

$$\hat{a}_2 = t\hat{a}_0 + r\hat{a}_1.$$

Then the photon number operator is

$$\begin{aligned} \hat{a}_2^\dagger \hat{a}_2 &= (t\hat{a}_0^\dagger + r\hat{a}_1^\dagger)(t\hat{a}_0 + r\hat{a}_1) = \\ &= t^2 \hat{a}_0^\dagger \hat{a}_0 + rt (\hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0) + r^2 \hat{a}_1^\dagger \hat{a}_1. \end{aligned}$$

Assume that the local oscillator is in a coherent state, and the signal is in a squeezed vacuum state. Then the input state is two-mode

$$|\psi\rangle = |\alpha_p\rangle |r, 0\rangle.$$

The expectation value of the photon number operator $\hat{n}_2 = \hat{a}_2^\dagger \hat{a}_2$ in this case is:

$$\begin{aligned} \langle \hat{n}_2 \rangle &= \langle \psi | \hat{a}_2^\dagger \hat{a}_2 | \psi \rangle = \\ &= t^2 \langle r, 0 | \hat{a}_0^\dagger \hat{a}_0 | r, 0 \rangle + r^2 |\alpha_p|^2 + 2tr |\alpha_p| \langle r, 0 | \hat{X}(\varphi) | r, 0 \rangle, \end{aligned} \quad (9.26)$$

where

$$\begin{aligned} \hat{X}(\varphi) &= \frac{1}{2} (\hat{a}_0 e^{-i\varphi} + \hat{a}_0^\dagger e^{i\varphi}), \\ \alpha_p &= |\alpha_p| e^{i\varphi}. \end{aligned}$$

Here φ is the local oscillator phase. Then at $\varphi = 0$ and at $\varphi = \frac{\pi}{2}$ we have

$$\begin{aligned}\hat{X}(0) &= \frac{1}{2} (\hat{a}_0 + \hat{a}_0^\dagger) = \hat{X}_1, \\ \hat{X}\left(\frac{\pi}{2}\right) &= \frac{1}{2i} (\hat{a}_0 - \hat{a}_0^\dagger) = \hat{X}_2,\end{aligned}$$

i.e., by changing the local oscillator phase with a phase shifter, one can select the corresponding quadrature component.

Equation (9.26) contains three terms: those containing $\langle \hat{a}_0^\dagger \hat{a}_0 \rangle$ — the average photon number in the signal mode, $r^2 |\alpha_p|^2$ — the average photon number in the local oscillator mode, and the interference term $\langle r, 0 | \hat{X}(\varphi) | r, 0 \rangle$, which selects the respective quadrature component. Assume that the local oscillator intensity is so high that

$$r^2 |\alpha_p|^2 \gg \langle \hat{a}_0^\dagger \hat{a}_0 \rangle,$$

then the first term in (9.26) can be neglected. In this case, the average photon number incident on the photodetector is

$$\langle \hat{n}_2 \rangle = r^2 |\alpha_p|^2 + 2tr |\alpha_p| \langle r, 0 | \hat{X}(\varphi) | r, 0 \rangle,$$

where the first term is a known number and can be subtracted. Only the term containing the signal quadrature remains. Next, the photon number fluctuations (uncertainty) can be determined using the formula

$$(\Delta n_2)^2 = \langle \hat{n}_2^2 \rangle - \langle \hat{n}_2 \rangle^2.$$

Calculations lead to the formula

$$(\Delta n_2)^2 = r^2 |\alpha_p|^2 \left\{ r^2 + 4t^2 (\Delta X(\varphi))^2 \right\}. \quad (9.27)$$

From (9.27) it follows that noise has two components, one of which $r^2 |\alpha_p|^2 r^2$ — expresses the local oscillator noise, the other $r^2 |\alpha_p|^2 4t^2 (\Delta X(\varphi))^2$ — is related to the signal noise. If the signal is vacuum fluctuations,

$$(\Delta X(\varphi))^2 = \frac{1}{4},$$

then the squeezing condition is

$$(\Delta X_1)^2 < \frac{1}{4}$$

for the squeezed component.

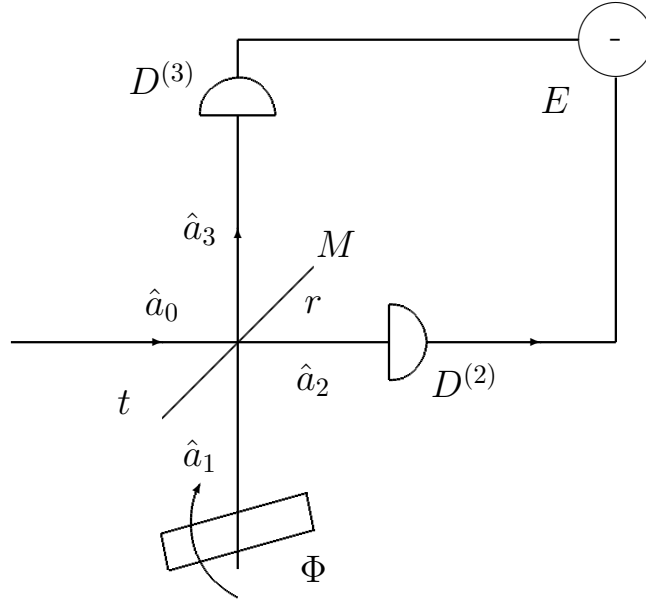


Figure 9.11: Balanced scheme of the homodyne detector. The investigated signal (\hat{a}_0) and the homodyne mode signal (\hat{a}_1) are applied to the mirror M with a reflection coefficient $r = \frac{1}{\sqrt{2}}$ and a transmission coefficient $t = \frac{1}{\sqrt{2}}$. The signals after passing through the mirror are registered by two photodetectors $D^{(2)}$ and $D^{(3)}$, whose output is fed to a subtractor E .

The squeezing effect can be detected using a homodyne detector in the following experiment. First, only the vacuum state is present at the input, and the noise level is measured. Then, the studied oscillation is fed to the input, and the noise level is measured depending on φ . If a squeezed state is applied at the input, at some value $\varphi = \varphi^{(1)}$, the noise will be minimal (less than in the case of vacuum fluctuations at the input). At another value $\varphi = \varphi^{(2)}$, differing by π , the noise will be maximal (higher than in the case of vacuum fluctuations). Such experiments have been carried out repeatedly. At the same time, the squeezing effect was observed and its degree measured (see for example [30]).

9.5.1 Balanced homodyne detector scheme

A more complex homodyne receiver scheme is the balanced scheme, shown in Figure 9.11. It allows substantial reduction of the local oscillator noise. It differs from the previously discussed scheme by having two photodetectors, one for each output beam. The signals from each photodetector are subtracted, and the difference is recorded. The beam splitter is semi-transparent with $t = r = \frac{1}{\sqrt{2}}$ (this time we assume t and r are real, but the reflection coefficient on different sides of the beam splitter has opposite sign).

In the scheme, \hat{a}_0 is the signal operator, and \hat{a}_1 is the local oscillator mode

operator. It is assumed that the local oscillator is in a coherent state with a large amplitude. The fields of modes 2 and 3 (operators \hat{a}_2 and \hat{a}_3) are detected by separate photodetectors $D^{(2)}$ and $D^{(3)}$, and the signals from each photodetector are subtracted. Thus, the output signal is defined by the expectation value of the operator

$$\hat{n}_{23} = \hat{a}_2^\dagger \hat{a}_2 - \hat{a}_3^\dagger \hat{a}_3.$$

The operators \hat{a}_2 and \hat{a}_3 are related to the input field operators \hat{a}_0 and \hat{a}_1 by the beam splitter through the relations:

$$\begin{aligned}\hat{a}_2 &= \frac{1}{\sqrt{2}} (\hat{a}_0 + \hat{a}_1), \\ \hat{a}_3 &= \frac{1}{\sqrt{2}} (-\hat{a}_0 + \hat{a}_1),\end{aligned}$$

so

$$\begin{aligned}\hat{n}_{23} &= \frac{1}{2} \left((\hat{a}_0^\dagger + \hat{a}_1^\dagger) (\hat{a}_0 + \hat{a}_1) - (-\hat{a}_0^\dagger + \hat{a}_1^\dagger) (-\hat{a}_0 + \hat{a}_1) \right) = \\ &= \frac{1}{2} \left(\hat{a}_0^\dagger \hat{a}_0 + \hat{a}_1^\dagger \hat{a}_0 + \hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_1 - \left(\hat{a}_0^\dagger \hat{a}_0 - \hat{a}_1^\dagger \hat{a}_0 - \hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_1 \right) \right) = \\ &= \hat{a}_0^\dagger \hat{a}_1 + \hat{a}_1^\dagger \hat{a}_0 = \left(\hat{a}_0^\dagger e^{i\varphi} + \hat{a}_0 e^{-i\varphi} \right) |\alpha_H|,\end{aligned}$$

where it was used that the local oscillator is in a coherent state with large amplitude

$$\alpha_H = |\alpha_H| e^{i\varphi},$$

and \hat{a}_1 is replaced by the classical field $|\alpha_H| e^{i\varphi}$ ($|\alpha_H| \gg 1$).

Let us find the expectation value of the operator \hat{n}_{23} in the case of the signal in a squeezed vacuum state $|z, 0\rangle$, where for simplicity let $z = r$ be real ($r = |z|$). Also for simplicity, set $\varphi = 0$. In this case

$$\hat{n}_{23} = \left(\hat{a}_0 + \hat{a}_0^\dagger \right) |\alpha_H| = 2\hat{X}_1 |\alpha_H|.$$

This means we consider squeezing of the quadrature component \hat{X}_1 . The expectation value of \hat{n}_{23} writes as

$$\begin{aligned}\langle \hat{n}_{23} \rangle &= 2 |\alpha_H| \langle \hat{X}_1 \rangle = |\alpha_H| \langle r, 0 | \hat{a}_0 + \hat{a}_0^\dagger | r, 0 \rangle = \\ &= |\alpha_H| \langle 0 | \hat{S}^\dagger(r) \left(\hat{a}_0 + \hat{a}_0^\dagger \right) \hat{S}(r) | 0 \rangle = \\ &= |\alpha_H| \left(\langle 0 | \hat{a}_0 \cosh r - \hat{a}_0^\dagger \sinh r | 0 \rangle + \langle 0 | \hat{a}_0^\dagger \cosh r - \hat{a}_0 \sinh r | 0 \rangle \right) = 0,\end{aligned}$$

since $\hat{a}_0 |0\rangle = 0$, $\langle 0 | \hat{a}_0^\dagger = 0$. Thus, the mean value $\langle \hat{n}_{23} \rangle = 0$. However, the mean square is not zero; it reflects noise limiting measurement accuracy.

Calculate the uncertainty $(\Delta X_{1,2})^2$, which determines the noise level for the considered balanced scheme:

$$\begin{aligned} (\Delta X_{1,2})^2 &= \langle r, 0 | \frac{1}{4} \left(\hat{a}_0 \pm \hat{a}_0^\dagger \right)^2 | r, 0 \rangle - \\ &\quad - \langle r, 0 | \frac{1}{2} \left(\hat{a}_0 \pm \hat{a}_0^\dagger \right) | r, 0 \rangle^2. \end{aligned} \quad (9.28)$$

Since we showed that the last term in (9.28) equals zero, we obtain:

$$\begin{aligned} (\Delta X_{1,2})^2 &= \langle 0 | \hat{S}^\dagger(r, 0) \frac{1}{4} \left(\hat{a}_0 \pm \hat{a}_0^\dagger \right)^2 \hat{S}(r, 0) | 0 \rangle = \\ &= \langle 0 | \hat{S}^\dagger(r, 0) \frac{1}{4} \left(\hat{a}_0 \pm \hat{a}_0^\dagger \right) \hat{S}^\dagger(r, 0) \hat{S}(r, 0) \left(\hat{a}_0 \pm \hat{a}_0^\dagger \right) \hat{S}(r, 0) | 0 \rangle. \end{aligned} \quad (9.29)$$

To find $(\Delta X_{1,2})^2$, calculate the averages:

$$\begin{aligned} \langle \hat{a}_0 \rangle &= \langle 0 | \hat{S}^\dagger(r, 0) \hat{a}_0 \hat{S}(r, 0) | 0 \rangle = \langle 0 | \hat{a}_0 \cosh r - \hat{a}_0^\dagger \sinh r | 0 \rangle = 0, \\ \langle \hat{a}_0^\dagger \rangle &= \langle 0 | \hat{a}_0^\dagger \cosh r - \hat{a}_0 \sinh r | 0 \rangle = 0, \end{aligned} \quad (9.30)$$

since $\hat{a}_0 |0\rangle = 0$, $\langle 0 | \hat{a}_0^\dagger = 0$.

Similarly, one can show

$$\begin{aligned} \langle \hat{a}_0^\dagger \hat{a}_0 \rangle &= \sinh^2 r, \\ \langle \hat{a}_0 \hat{a}_0^\dagger \rangle &= 1 + \langle \hat{a}_0^\dagger \hat{a}_0 \rangle = 1 + \sinh^2 r = \cosh^2 r, \\ \langle \hat{a}_0 \hat{a}_0 \rangle &= \langle \hat{a}_0^\dagger \hat{a}_0^\dagger \rangle^* = -\cosh r \sinh r. \end{aligned} \quad (9.31)$$

Transforming (9.29) using (9.30) and (9.31), we get the following expression:

$$(\Delta X_{1,2})^2 = \cosh 2r \mp \sinh 2r = e^{\mp 2r},$$

from which

$$(\Delta n_{2,3})^2 = n_H e^{-2r}$$

for the X_1 component. The variance is

$$\Delta n_{2,3} = \sqrt{n_H} e^{-r},$$

where $n_H = |\alpha_H|^2$ is the photon number in the local oscillator field. Using the balanced scheme, it is possible to measure the degree of squeezing in the same way as with the simplest homodyne scheme. The advantage here is that the local oscillator noise is compensated, and measurement accuracy increases.

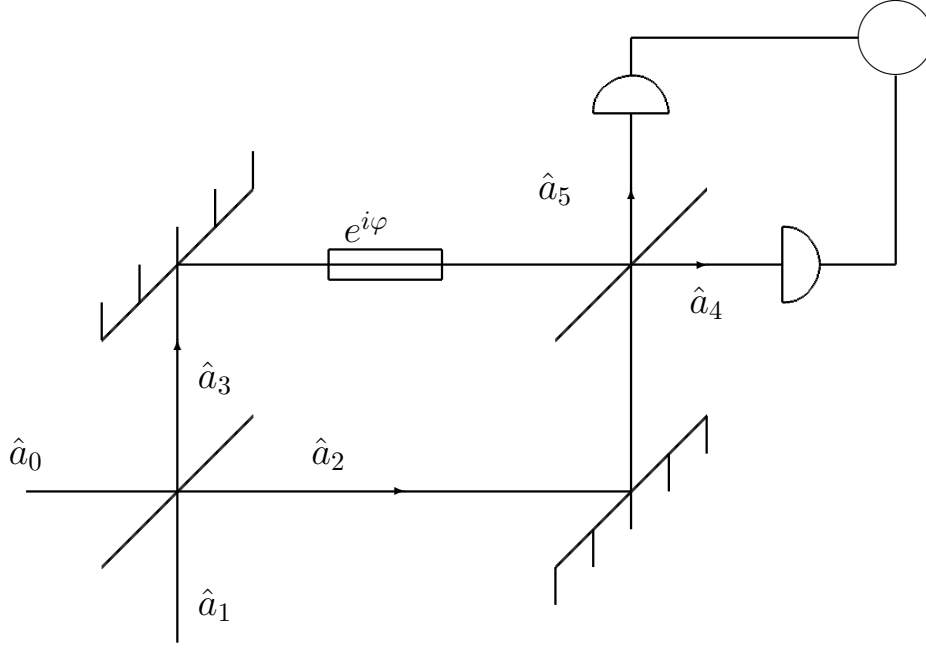


Figure 9.12: Mach-Zehnder Interferometer

9.6 Interference Measurements Using Squeezed Light

Consider the Mach-Zehnder interferometer scheme shown in [Figure 9.12](#). This scheme has been previously examined during the study of interference measurement errors. The difference is that at the zero input, squeezed vacuum radiation is fed instead of vacuum radiation. We still assume that the input and output mirrors are partially transparent ($t = r = \frac{1}{\sqrt{2}}$), and the angular mirrors are opaque ($r = 1, t = 0$). The scattering matrix of the mirrors is

$$\hat{S} = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & i \\ i & 1 \end{pmatrix}.$$

The interferometer is considered as a sensor of some physical quantity affecting the optical length of one arm. On the scheme, the phase shifter included in the upper arm can respond to the action of this quantity.

The interferometer equations are given by [\(6.12\)](#):

$$\begin{aligned} \hat{a}_2 &= \frac{1}{\sqrt{2}} (\hat{a}_0 + i\hat{a}_1), \quad \hat{a}_3 = \frac{1}{\sqrt{2}} (i\hat{a}_0 + \hat{a}_1), \\ \hat{a}_4 &= \frac{1}{\sqrt{2}} (i\hat{a}_2 + e^{i\varphi}\hat{a}_3) = \frac{1}{2} [i(1 + e^{i\varphi})\hat{a}_0 - (1 - e^{i\varphi})\hat{a}_1], \\ \hat{a}_5 &= \frac{1}{\sqrt{2}} (\hat{a}_2 + ie^{i\varphi}\hat{a}_3) = \frac{1}{2} [(1 - e^{i\varphi})\hat{a}_0 + i(1 + e^{i\varphi})\hat{a}_1]. \end{aligned} \quad (9.32)$$

Let the input signal field (input 0, operator \hat{a}_0) be in a squeezed vacuum state.

The second input field (input 1, operator \hat{a}_1) is in a coherent state with a large amplitude. So large that the input field is two-mode

$$|\psi\rangle_{\text{in}} = |\alpha\rangle |re^{i\theta}, 0\rangle,$$

where $|\alpha\rangle$ is the coherent state, $|re^{i\theta}, 0\rangle$ is the squeezed vacuum state. Channels 4 and 5 are each detected by their own photodetector. The signals from the photodetectors are subtracted and recorded.

Thus, the output signal is determined by the average value of the operator

$$\begin{aligned} \hat{n}_{54} &= \hat{a}_5^\dagger \hat{a}_5 - \hat{a}_4^\dagger \hat{a}_4 = \\ &= \left(\hat{a}_1^\dagger \hat{a}_1 - \hat{a}_0^\dagger \hat{a}_0 \right) \cos \varphi - \left(\hat{a}_0^\dagger \hat{a}_1 - \hat{a}_1^\dagger \hat{a}_0 \right) \sin \varphi \end{aligned} \quad (9.33)$$

Here the equality (9.32) was used.

Let us find the average value of the operator $\langle \psi | \hat{n}_{54} | \psi \rangle$. We take into account that each operator acts only on its own mode. We have:

$$\begin{aligned} \langle \psi | \hat{n}_{54} | \psi \rangle &= \left(\langle \alpha | \hat{a}_1^\dagger \hat{a}_1 | \alpha \rangle - \langle re^{i\theta}, 0 | \hat{a}_0^\dagger \hat{a}_0 | re^{i\theta}, 0 \rangle \right) \cos \varphi - \\ &- \left(\langle re^{i\theta}, 0 | \hat{a}_0^\dagger | re^{i\theta}, 0 \rangle \alpha + \alpha \langle re^{i\theta}, 0 | \hat{a}_0 | re^{i\theta}, 0 \rangle \right) \sin \varphi. \end{aligned} \quad (9.34)$$

The first term in the first bracket equals the average number of pump photons $|\alpha|^2$, and the second term, as we know, equals the number of photons in the squeezed state, which is sh^2r . The second bracket on the right, as we found when examining the balanced detector, is zero. Thus we get

$$\langle \psi | \hat{n}_{54} | \psi \rangle = \left(|\alpha|^2 - sh^2r \right) \cos \varphi.$$

If we set $\varphi = \frac{\pi}{2}$, then the average value of the output signal is zero:

$$\langle \psi | \hat{n}_{54} | \psi \rangle = 0.$$

The average value of the second term in (9.34) is zero, but its mean square is not zero. This means it is a source of noise which limits measurement accuracy.

Let us calculate the mean square of \hat{n}_{54} :

$$(\Delta n_{54})^2 = \langle \psi | \hat{n}_{54}^2 | \psi \rangle - \langle \psi | \hat{n}_{54} | \psi \rangle^2.$$

The mean in our case is zero, so we have

$$\begin{aligned} (\Delta n_{54})^2 &= \langle \psi | \hat{n}_{54}^2 | \psi \rangle = \\ &= |\alpha|^2 \langle re^{i\theta}, 0 | \left(\hat{a}_0^\dagger + \hat{a}_0 \right)^2 | re^{i\theta}, 0 \rangle = \\ &= |\alpha|^2 \langle re^{i\theta}, 0 | \left(\hat{a}_0^\dagger \right)^2 + (\hat{a}_0)^2 + \hat{a}_0^\dagger \hat{a}_0 + \hat{a}_0 \hat{a}_0^\dagger | re^{i\theta}, 0 \rangle. \end{aligned}$$

Previously, we had (9.30):

$$\begin{aligned}\langle \hat{a}_0 \rangle &= \langle 0 | \hat{S}^\dagger(r, 0) \hat{a}_0 \hat{S}(r, 0) | 0 \rangle = \langle 0 | \hat{a}_0 \cosh r - \hat{a}_0^\dagger \sinh r | 0 \rangle = 0, \\ \langle \hat{a}_0^\dagger \rangle &= \langle 0 | \hat{a}_0^\dagger \cosh r - \hat{a}_0 \sinh r | 0 \rangle = 0,\end{aligned}$$

from which we get

$$(\Delta n_{54})^2 = \langle \psi | \hat{n}_{54}^2 | \psi \rangle = |\alpha|^2 (\cosh 2r - \sinh 2r) = |\alpha|^2 e^{-2r}.$$

The root mean square noise value will be

$$|\Delta n_{54}| = |\alpha| e^{-r}.$$

We take this value as the threshold below which we will not be able to detect or measure the signal. Hence, the signal must be above the threshold.

If the initial value $\varphi = \frac{\pi}{2}$ changes by an amount $\Delta\varphi$, a signal equal to

$$\Delta n_{54} = |\alpha|^2 \Delta\varphi$$

will appear. It is necessary that

$$|\alpha|^2 \Delta\varphi > |\alpha| e^{-r},$$

from which we obtain

$$\Delta\varphi > \frac{1}{|\alpha|} e^{-r}, \quad (9.35)$$

where $|\alpha| = \sqrt{\bar{n}}$, and \bar{n} is the average number of photons in the homodyne field.

Expression (9.35) gives a smaller value than the one obtained earlier when considering the accuracy of interference measurements. Using the squeezed state can increase measurement accuracy.

9.7 Squeezed Light in Nonlinear Phenomena

The efficiency of generating higher harmonics can be increased if squeezed light is used as the EMF instead of a coherent state. In particular, the $2N$ harmonic is generated $(2N - 1)!!$ times more effectively [42, pp. 72-76], where

$$n!! = \prod_{k=1}^{\frac{n}{2}} (2k)$$

denotes the double factorial.

TBD

9.8 Nonclassicality of the Squeezed State

In conclusion, we show that the squeezed state is nonclassical. The condition for a squeezed state is

$$(\Delta X_1)^2 < \frac{1}{4},$$

where

$$(\Delta X_1)^2 = \langle \hat{X}_1^2 \rangle - \langle \hat{X}_1 \rangle^2.$$

The operator

$$\hat{X}_1 = \frac{\hat{a} + \hat{a}^\dagger}{2},$$

then

$$\begin{aligned} \hat{X}_1^2 &= \frac{\hat{a}^2 + (\hat{a}^\dagger)^2 + \hat{a}^\dagger \hat{a} + \hat{a} \hat{a}^\dagger}{4} = \\ &= \frac{\hat{a}^2 + (\hat{a}^\dagger)^2 + 2\hat{a}^\dagger \hat{a} + 1}{4}. \end{aligned}$$

At the same time,

$$(\hat{N} \hat{X}_1^2) = \frac{\hat{a}^2 + (\hat{a}^\dagger)^2 + 2\hat{a}^\dagger \hat{a}}{4},$$

where $\hat{N} \hat{X}_1^2$ is the operator subjected to the action of the operator \hat{N} , whose action consists in setting the normal ordering of the operators \hat{a}^\dagger and \hat{a} , regardless of the commutation relation.

From this we get

$$(\Delta N X_1)^2 = (\Delta X_1)^2 - \frac{1}{4},$$

or

$$(\Delta N X_1)^2 < 0,$$

since

$$(\Delta X_1)^2 < \frac{1}{4}.$$

Suppose the squeezed state is described by a density operator, which is written in the coherent state representation as

$$\hat{\rho} = \int P(\alpha) |\alpha\rangle \langle \alpha| d^2\alpha,$$

then for the expectation value

$$(\Delta NX_1)^2 = \langle \hat{N} \hat{X}_1^2 \rangle - \langle \hat{N} \hat{X}_1 \rangle^2 \quad (9.36)$$

we can write

$$\begin{aligned} (\Delta NX_1)^2 &= \int P(\alpha) \langle \alpha | \hat{N} \hat{X}_1^2 | \alpha \rangle d^2\alpha - \int P(\alpha) \langle X_1 \rangle^2 d^2\alpha = \\ &= \int P(\alpha) \left(\frac{\alpha^2 + \alpha^{*2} + 2\alpha\alpha^*}{4} - \langle X_1 \rangle^2 \right) d^2\alpha = \\ &= \int P(\alpha) (\Delta X_1)^2 d^2\alpha \end{aligned} \quad (9.37)$$

where $(\Delta X_1)^2$ is obtained from (9.36) if all \hat{a}^\dagger are replaced by α^* , and \hat{a} by α . It can be shown that

$$(\Delta X_1)^2 = (\Delta\alpha^* + \Delta\alpha)^2$$

which is clearly non-negative. By the squeezing condition, the integral must be less than zero:

$$(\Delta NX_1)^2 = \int P(\alpha) (\Delta X_1(\alpha^*\alpha))^2 d^2\alpha < 0.$$

This is obviously possible if at least on part of the plane (α) , $P(\alpha) < 0$. This, as we know, is the condition for nonclassicality.

9.9 Second-Order Coherence for the Squeezed Vacuum

Let us calculate the second-order coherence for the squeezed vacuum $|z, 0\rangle$. From (8.1) we have

$$G^{(2)} = \frac{\langle z, 0 | \hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{a} | z, 0 \rangle - \langle n \rangle^2}{\langle n \rangle^2}.$$

The expression $\langle z, 0 | \hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{a} | z, 0 \rangle$ can be rewritten as

$$\begin{aligned} \langle z, 0 | \hat{a}^\dagger \hat{a} \hat{a}^\dagger \hat{a} | z, 0 \rangle &= \\ &= \langle 0 | \hat{a}^\dagger \hat{S} \hat{S}^\dagger \hat{a} \hat{S} \hat{S}^\dagger \hat{a}^\dagger \hat{S} \hat{S}^\dagger \hat{a} \hat{S} | 0 \rangle = \langle \phi | \phi \rangle, \end{aligned}$$

where

$$|\phi\rangle = \hat{S}^\dagger \hat{a}^\dagger \hat{S} \hat{S}^\dagger \hat{a} \hat{S} | 0 \rangle.$$

Using (9.8) one obtains

$$\begin{aligned} |\phi\rangle &= (\hat{a}^\dagger ch r - \hat{a} e^{-i\theta} sh r) (\hat{a} ch r - \hat{a}^\dagger e^{i\theta} sh r) |0\rangle = \\ &= e^{i\theta} sh r (\hat{a}^\dagger ch r - \hat{a} e^{-i\theta} sh r) \hat{a}^\dagger |0\rangle = sh^2 r |0\rangle + \sqrt{2} e^{i\theta} sh r ch r |2\rangle. \end{aligned}$$

Thus,

$$\langle\phi|\phi\rangle = sh^4 r + 2sh^2 r ch^2 r.$$

Therefore, taking into account (9.15),

$$\begin{aligned} G^{(2)} &= \frac{sh^4 r + 2sh^2 r ch^2 r - sh^2 r}{sh^4 r} = 1 + 2\frac{ch^2 r}{sh^2 r} - \frac{1}{sh^2 r} = \\ &= 1 + \frac{ch^2 r}{sh^2 r} + \frac{ch^2 r - 1}{sh^2 r} = 1 + \frac{ch^2 r}{sh^2 r} + \frac{sh^2 r}{sh^2 r} = 2 + \frac{ch^2 r}{sh^2 r} = \\ &= 2 + \frac{1 + sh^2 r}{sh^2 r} = 3 + \frac{1}{\langle n \rangle} \end{aligned}$$

9.10 Exercises

1. Derive the relations for the uncertainties (9.2) and (9.3).
2. Derive the relations for the uncertainties (9.4) and (9.5).
3. Prove (9.7) and (9.8).
4. Prove the uncertainty relations (9.13).
5. Prove the expressions (9.16) for the displacement operators $\hat{D}(\alpha)$.
6. Prove that (9.23) and (9.24) are the solution of equation (9.21) considering the initial conditions (9.22).

Chapter 10

Entangled States

Since the inception of quantum mechanics, the question of the completeness of this theory has arisen. The accuracy of quantum mechanics' predictions is very high, but it only predicts the probabilities of various events. In particular, there is no possibility to measure both the coordinate and momentum of a particle with arbitrary precision. It seems that the probabilities inherent in the quantum mechanical description reflect its incompleteness, and possibly, there exists another theory that would have the accuracy of quantum mechanics without using a probabilistic approach.

It turned out that the probabilities underlying quantum mechanics have profound physical meaning and there are no theories where these can be discarded or where, for example, one could measure coordinates and momentum with arbitrary precision. A special role is played by what are known as entangled states, which describe systems consisting of several particles, with the behavior of such a composite system described by a general wave function.

Since entangled states are purely quantum, i.e., they have no classical analogs, they offer the possibility to observe phenomena that seem utterly impossible from a classical perspective, such as quantum teleportation. Moreover, in recent times, practical applications of entangled states have emerged, such as quantum dense coding (see [11.2.3](#)) and quantum cryptography (see [12.3](#)).

There are several methods for obtaining entangled photons [[34](#)], related to quantum optics, among which polarization entangled states are noteworthy. This is because there exists a large number of methods for controlling polarization characteristics as well as methods for measuring the polarization properties of light.

10.1 EPR Paradox for Stokes Parameters and Entangled States

Consider the single-photon state (7.3). To explain the physical meaning of the parameters α and β , one needs to refer to the experimental setup scheme shown in Figure 7.1. Suppose that the source S emits photons in the state (7.3). Then the average current of photodetector D_x will be proportional to $|\alpha|^2$, and that of photodetector D_y - $|\beta|^2$, i.e., $P_x = |\alpha|^2$ and $P_y = |\beta|^2$ describe the probabilities of detecting a photon in a state polarized along x - $|x\rangle$ or y - $|y\rangle$. The following question arises - "what is behind these probabilities?". Two possible answers exist. The first one assumes that we actually have incomplete knowledge about the source S , i.e., if we knew everything about this source, we could predict which photon, polarized along x or along y , is created at an arbitrary moment in time and, correspondingly, the readings of photodetectors D_x and D_y . From this point of view, quantum mechanics represents some intermediate theory, which can later be replaced by a more precise one, describing quantum properties of objects with absolute accuracy.

The second answer variant assumes that it is impossible to determine all parameters describing some quantum system with arbitrary precision because the measurement result is not predetermined in advance (as in our case by properties of source S), but rather defined at the moment of measurement as a result of the interaction between the quantum system and the macroscopic measuring apparatus.

The first attempt to answer this question was made by A. Einstein, B. Podolsky, and N. Rosen in 1935 in a paper often called EPR [8]. In EPR, a physical theory is considered complete if every element of physical reality has a counterpart in the physical theory. The following was adopted as a definition of an element of physical reality: "if, without disturbing a system, we can predict with certainty (probability equal to unity) the value of some physical quantity, then there exists an element of reality corresponding to this quantity" [40]. Non-commuting operators are of particular interest here. In our case, these could be operators \hat{S}_1 and \hat{S}_2 . Since these operators do not commute (7.5), the values of the corresponding observables cannot be measured with arbitrary precision. From this, two assumptions can be made:

1. The quantum mechanical description of reality is incomplete
2. The quantities defined by operators \hat{S}_1 and \hat{S}_2 cannot be simultaneously real

As a test system, following EPR, we consider a complex system consisting of

several particles. At the same time, some overall wave function is used to describe the entire system.

To describe the polarization properties of a complex system consisting of two photons, one must use a wave function of the following form:

$$\begin{aligned} |\psi\rangle &= \sum_{i,j=x,y} c_{ij} |i\rangle_1 |j\rangle_2 = \\ &= c_{xx} |x\rangle_1 |x\rangle_2 + c_{xy} |x\rangle_1 |y\rangle_2 + c_{yx} |y\rangle_1 |x\rangle_2 + c_{yy} |y\rangle_1 |y\rangle_2. \end{aligned} \quad (10.1)$$

If two or more coefficients c_{ij} in (10.1) are nonzero, then the wave function is non-factorizable, i.e.,

$$|\psi\rangle \neq |\psi\rangle_1 |\psi\rangle_2.$$

Thus, no single wave function can be assigned to each separate photon. Such states are called entangled.

Remark 10.1.1. Consider the following pure entangled state

$$|\psi_{12}\rangle = \frac{|x\rangle_1 |y\rangle_2 - |y\rangle_1 |x\rangle_2}{\sqrt{2}},$$

whose density matrix has the form

$$\begin{aligned} \hat{\rho}_{12} &= |\psi_{12}\rangle \langle \psi_{12}| = \\ &= \frac{1}{2} (|x\rangle_1 |y\rangle_2 - |y\rangle_1 |x\rangle_2) (\langle x|_1 \langle y|_2 - \langle y|_1 \langle x|_2) = \\ &= \frac{|x\rangle_1 |y\rangle_2 \langle y|_2 \langle x|_1 + |y\rangle_1 |x\rangle_2 \langle x|_2 \langle y|_1 -}{2} \\ &\quad - \frac{|x\rangle_1 |y\rangle_2 \langle x|_2 \langle y|_1 + |y\rangle_1 |x\rangle_2 \langle y|_2 \langle x|_1}{2}. \end{aligned}$$

At the same time, the density matrix of the second photon has the form (see [section 14.5](#))

$$\begin{aligned} \hat{\rho}_2 &= \text{Tr}_1 \hat{\rho}_{12} = \langle x|_1 \hat{\rho}_{12} |x\rangle_1 + \langle y|_1 \hat{\rho}_{12} |y\rangle_1 = \\ &= \frac{|y\rangle_2 \langle y|_2 + |x\rangle_2 \langle x|_2}{2}. \end{aligned}$$

Thus, the state of the second photon is mixed.

This astonishing fact greatly distinguishes quantum systems from classical ones: if we have a pure state of two particles, then the corresponding states of individual particles are not necessarily pure. In particular, if the state of the Universe is a pure quantum state [12], then the states of individual parts of this Universe (stars, planets, etc.) are not necessarily pure and can be mixed, i.e., classical.

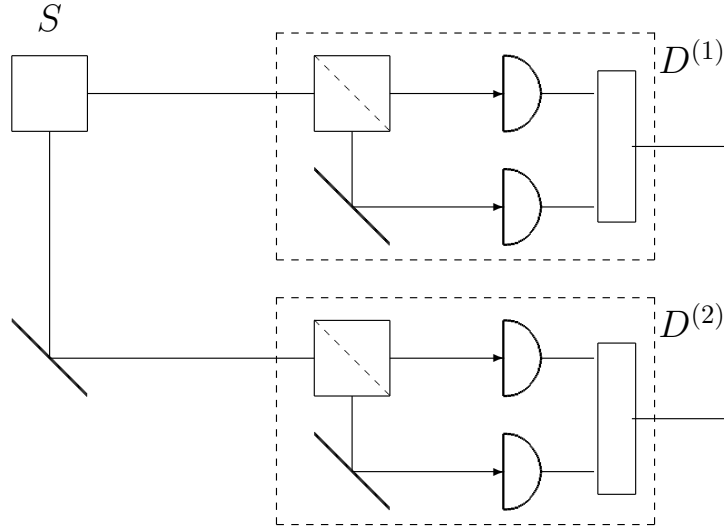


Figure 10.1: Scheme for measuring Stokes parameters for an entangled two-photon state. Source of entangled photons S . Detectors of entangled photons $D^{(1,2)}$

As an example, we study the following wave function:

$$|\psi\rangle = \frac{|+\rangle_1 |-\rangle_2 - |-\rangle_1 |+\rangle_2}{\sqrt{2}}, \quad (10.2)$$

i.e., all we know about the photons is that they have different circular polarizations. With 50% probability the first photon has left circular polarization, and in this case the second photon has right circular polarization. Conversely, with 50% probability the first photon has right circular polarization, and the second photon has left circular polarization.

To measure the polarization properties of the state (10.2), the scheme shown in Figure 10.1 can be used. In this scheme, entangled photons from the source S are fed to two detectors of Stokes parameters $D^{(1,2)}$. Detector $D^{(1)}$ measures the average values of the Stokes operators for the first photon - $\hat{S}_k^{(1)}$, and $D^{(2)}$ measures the average values of the Stokes operators $\hat{S}_k^{(2)}$ for the second photon. The setup of each detector is identical to the scheme shown in Figure 7.2.

Since the eigenvalues of the Stokes operators (7.4) are $s = \pm 1$, at the detector outputs we will obtain either $+1$ or -1 . Suppose we measure the Stokes parameter $\hat{S}_1^{(1)}$ for the first photon. Suppose the measurement outcome is $+1$. At this point, the wave function (10.2) undergoes reduction. The reduction can be described by the projection operator (see subsection 14.1.6) onto the state $|x\rangle_1$:

$$\hat{P}_{|x\rangle_1} = |x\rangle_1 \langle x|_1,$$

and the wave function (10.2) transforms as follows:

$$\begin{aligned}\hat{P}_{|x\rangle_1} |\psi\rangle &= |x\rangle_1 \langle x|_1 |\psi\rangle = \\ &= |x\rangle_1 \langle x|_1 \frac{(|x\rangle_1 + i|y\rangle_1)|-\rangle_2 - (|x\rangle_1 - i|y\rangle_1)|+\rangle_2}{2} = \\ &= |x\rangle_1 \frac{|-\rangle_2 - |+\rangle_2}{2} = -\frac{1}{2\sqrt{2}} |x\rangle_1 (2i) |y\rangle_2 = -\frac{i}{\sqrt{2}} |x\rangle_1 |y\rangle_2.\end{aligned}$$

Therefore, after measuring $\hat{S}_1^{(1)}$, the wave function of the entire system will be written as

$$|\psi\rangle_{red} = |x\rangle_1 |y\rangle_2.$$

This means that the second photon will be in the state $|y\rangle_2$, i.e., the reading of detector $D^{(2)}$ will be -1 .

Similarly, if the reading of the first detector was -1 , the wave function after measurement will be equal to

$$\begin{aligned}P_{|y\rangle_1} |\psi\rangle &= |y\rangle_1 \langle y|_1 |\psi\rangle = \\ &= |y\rangle_1 \langle y|_1 \frac{(|x\rangle_1 + i|y\rangle_1)|-\rangle_2 - (|x\rangle_1 - i|y\rangle_1)|+\rangle_2}{2} = \\ &= |y\rangle_1 i \frac{|-\rangle_2 + |+\rangle_2}{2} = \frac{i}{2\sqrt{2}} |y\rangle_1 (2) |x\rangle_2 = \frac{i}{\sqrt{2}} |y\rangle_1 |x\rangle_2,\end{aligned}$$

i.e.,

$$|\psi\rangle_{red} = |x\rangle_2 |y\rangle_1,$$

the state of the second photon will be described by the state $|x\rangle_2$ and the reading of detector $D^{(2)}$ will be $+1$. Thus, for any measurement of $\hat{S}_1^{(1)}$, we can reliably (with probability equal to 1) predict the measurement result of $\hat{S}_1^{(2)}$, without directly affecting the second photon, i.e., the quantity $\hat{S}_1^{(2)}$ should be considered an element of physical reality.

At the same time, the Stokes parameters for the second particle, $\hat{S}_1^{(2)}$ and $\hat{S}_2^{(2)}$ have different eigenvectors: $|x\rangle_2, |y\rangle_2$ (7.8, 7.9) and $\frac{1}{\sqrt{2}}(|x\rangle_2 \pm |y\rangle_2)$ (7.12, 7.13), respectively. Thus, the first and second Stokes parameters for the second particle cannot be measured simultaneously; indeed, otherwise, the measured value would correspond to a certain state vector that would be an eigenvector of both operators $\hat{S}_1^{(2)}$ and $\hat{S}_2^{(2)}$ (see also subsection 14.3.3).

Now, if we assume that the results of experiments are predetermined in advance, i.e., do not depend on which quantity (S_1 or S_2) was measured, and since measurements on the first photon are made in arbitrary order, it follows that for the second photon $\hat{S}_1^{(2)}$ and $\hat{S}_2^{(2)}$ must simultaneously be elements of physical reality. This entails the thesis about the incompleteness of quantum mechanics.

We have a different situation when we accept the assumption that the result of some measurement is determined at the moment the experiment is performed. In this case, we can no longer claim that $\hat{S}_1^{(2)}$ and $\hat{S}_2^{(2)}$ should be **simultaneously** considered as elements of physical reality, so there are no contradictions with the basic principles of quantum mechanics.

10.2 Bell Inequality for Stokes Parameters

The answer to the question about the completeness of quantum mechanics can be given by the following experiment [4, 34, 9].

Suppose that the system under consideration consists of two photons with the joint wave function of the form

$$|\psi\rangle = \frac{|x\rangle_1 |y\rangle_2 - |y\rangle_1 |x\rangle_2}{\sqrt{2}}. \quad (10.3)$$

Let detector $D^{(1)}$ (Figure 10.1) measure the quantity

$$\hat{A} = \hat{S}_1^{(1)}, (\xi = 0)$$

or

$$\hat{A}' = \hat{S}_2^{(1)}, (\xi = \frac{\pi}{4}).$$

As noted earlier, the eigenvalues of operators \hat{A} and \hat{A}' will be ± 1 , i.e., the readings of detector $D^{(1)}$ can only be two numbers: ± 1 .

Receiver $D^{(2)}$ will measure the following quantities

$$\hat{B} = \frac{1}{\sqrt{2}} \left(\hat{S}_1^{(2)} + \hat{S}_2^{(2)} \right), (\xi = \frac{\pi}{8})$$

or

$$\hat{B}' = \frac{1}{\sqrt{2}} \left(\hat{S}_1^{(2)} - \hat{S}_2^{(2)} \right), (\xi = -\frac{\pi}{8}).$$

To find the possible readings of $D^{(2)}$, it is necessary to find the eigenvalues of operators \hat{B} and \hat{B}' . Using the matrix representations of operators \hat{S}_1 (7.7) and \hat{S}_2 (7.11), we obtain for operator \hat{B} the following matrix representation:

$$\hat{B} = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \end{pmatrix}. \quad (10.4)$$

From (10.4) we can obtain the characteristic equation:

$$(1 - \sqrt{2}b) (-1 - \sqrt{2}b) - 1 = 0, \quad (10.5)$$

from which the eigenvalues are: $b = \pm 1$. In the case of operator \hat{B}' , we can obtain from (7.7) and (7.11) the following matrix representation:

$$\hat{B}' = \begin{pmatrix} \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ -\frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \end{pmatrix}. \quad (10.6)$$

and the corresponding characteristic equation

$$(1 - \sqrt{2}b)(-1 - \sqrt{2}b) - 1 = 0. \quad (10.7)$$

As is easy to see, equation (10.7) has the same solutions as (10.5), thus the readings of $D^{(2)}$, just like for $D^{(1)}$, will be only two values ± 1 .

Four series of experiments are conducted, each with N trials, in which pairs of operators (\hat{A}, \hat{B}) , (\hat{A}', \hat{B}) , (\hat{A}, \hat{B}') , and (\hat{A}', \hat{B}') are measured. As a result, the following sets of numbers (a_i, b_i) , (a'_i, b_i) , (a_i, b'_i) , and (a'_i, b'_i) are obtained. As was just shown, each of the obtained numbers can be either $+1$ or -1 .

Next, from these pairs, the following value is calculated

$$f_i = \frac{1}{2} (a_i b_i + a'_i b_i + a_i b'_i - a'_i b'_i)$$

for which the average is computed as

$$\langle F \rangle_N = \frac{1}{N} \sum_i f_i. \quad (10.8)$$

As $N \rightarrow \infty$ one can assume

$$\langle F \rangle_N \rightarrow \langle F \rangle.$$

Remark 10.2.1 (About the Bell Experiment). *As noted above, in the experiment we perform measurements of the following pairs of physical quantities (\hat{A}, \hat{B}) , (\hat{A}', \hat{B}) , (\hat{A}, \hat{B}') , and (\hat{A}', \hat{B}') , and for computing (10.8) it is necessary to assume that all these measurements are made in one go, i.e., simultaneously. In the classical case, unlike the quantum case, this is allowed. In addition, from the classical point of view, \hat{A} and \hat{A}' are two independent observables, as are \hat{B} and \hat{B}' . At the same time, from the quantum perspective, they are no longer independent, for example, for \hat{A}, \hat{A}' we have (7.5):*

$$[\hat{A}, \hat{A}'] = 2i\hat{S}_3.$$

Furthermore, from the classical point of view, the measurements of quantities \hat{A}, \hat{A}' on one hand and \hat{B}, \hat{B}' on the other are also independent because they are

performed at different points in space. Thus, in the classical case the four pairs (\hat{A}, \hat{B}) , (\hat{A}', \hat{B}) , (\hat{A}, \hat{B}') , and (\hat{A}', \hat{B}') are independent, and it does not matter whether they are measured in the same experiment or in several.

It should be noted that although \hat{A} and \hat{A}' are dependent from the quantum point of view, we cannot measure them in a single experiment simultaneously (see [subsection 14.3.3](#)); at the same time, measurements at different points in space are connected through the joint wave function, which leads to correlation between the quantities \hat{A}, \hat{A}' on the one hand and \hat{B}, \hat{B}' on the other, and consequently to differing results of the averaging for the quantity (10.8).

The quantum mechanical approach gives the following expression for the average in the considered state (10.3)

$$\begin{aligned} \langle F \rangle &\sim \langle F \rangle_{quant} = \frac{1}{2} \langle \psi | \hat{A}\hat{B} + \hat{A}'\hat{B} + \hat{A}\hat{B}' - \hat{A}'\hat{B}' | \psi \rangle = \\ &= \frac{1}{2} \langle \psi | \hat{A} (\hat{B} + \hat{B}') + \hat{A}' (\hat{B} - \hat{B}') | \psi \rangle = \\ &= \frac{1}{\sqrt{2}} \langle \psi | \hat{S}_1^{(1)} \hat{S}_1^{(2)} + \hat{S}_2^{(1)} \hat{S}_2^{(2)} | \psi \rangle = \\ &= \frac{1}{\sqrt{2}} (-1 - 1) = -\sqrt{2}. \end{aligned} \quad (10.9)$$

If we accept the thesis of the incompleteness of quantum mechanics, then we assume that the polarization of photons, and hence the Stokes parameters, are defined immediately after the two photons leave the source S . Thus, there exist a priori values a , a' , b , and b' , and the properties of the source can be described using classical probability theory. Thus, there exist 16 elementary probabilities $p(a, b, a', b')$ such that the average value can be written as

$$\langle F \rangle \sim \langle F \rangle_{class} = \sum_{a, b, a', b' = \pm 1} p(a, b, a', b') f(a, b, a', b'), \quad (10.10)$$

where

$$\begin{aligned} f(a, b, a', b') &= \frac{1}{2} (ab + a'b + ab' - a'b') = \\ &= \frac{1}{2} (a(b + b') + a'(b - b')). \end{aligned} \quad (10.11)$$

Function (10.11) can only take two values: $f = \pm 1$. Indeed, there are two possibilities: $b = b'$ or $b = -b'$. In the first case,

$$f = \frac{1}{2} (2ab) = \pm 1.$$

In the second case,

$$f = \frac{1}{2} (2a'b) = \pm 1.$$

Thus, function f can take values within the following interval $f_{min} = -1 \leq f \leq f_{max} = +1$ (unlike the measurable quantity $f_i = 0, \pm 1, \pm 2$). It is clear that in the classical case

$$|\langle F \rangle_{class}| \leq 1. \quad (10.12)$$

Comparing expressions (10.9) and (10.12), which are called Bell inequalities, one can see that quantum correlations have a greater quantitative value. This quantitative difference can be tested in an experiment. The first experiment was conducted in 1972 [9] and its results confirm the completeness of the quantum description. The experiment verifying the completeness of quantum mechanics can have not only theoretical importance but also practical applications, for instance, in quantum cryptography, which we will consider below (see 12.3).

It should be noted that expression (10.9) can be obtained using formula (10.10), provided it is accepted that some of the probabilities $p(a, b, a', b') < 0$. Thus we can speak about the non-classicality of the entangled state.

10.3 Bell Basis States

Any two-photon entangled state of the form (10.1):

$$\begin{aligned} |\psi\rangle = & c_{xx} |x\rangle_1 |x\rangle_2 + c_{xy} |x\rangle_1 |y\rangle_2 + \\ & + c_{yx} |y\rangle_1 |x\rangle_2 + c_{yy} |y\rangle_1 |y\rangle_2 \end{aligned}$$

can be decomposed into the following Bell basis states:

$$\begin{aligned} |\psi^+\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 + |y\rangle_1 |x\rangle_2), \\ |\psi^-\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 - |y\rangle_1 |x\rangle_2), \\ |\phi^+\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |x\rangle_2 + |y\rangle_1 |y\rangle_2), \\ |\phi^-\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |x\rangle_2 - |y\rangle_1 |y\rangle_2). \end{aligned} \quad (10.13)$$

To prove the decomposability, it is sufficient to show that the wavefunctions (10.13) are orthogonal. For this, we write each of these functions in matrix form

in the basis formed by $|x\rangle_1 |x\rangle_2$, $|x\rangle_1 |y\rangle_2$, $|y\rangle_1 |x\rangle_2$, and $|y\rangle_1 |y\rangle_2$, resulting in

$$\begin{aligned}
 |\psi^\dagger\rangle_{12} &= \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ 1 \\ 1 \\ 0 \end{pmatrix}, \\
 |\psi^-\rangle_{12} &= \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ 1 \\ -1 \\ 0 \end{pmatrix}, \\
 |\phi^\dagger\rangle_{12} &= \frac{1}{\sqrt{2}} \begin{pmatrix} 1 \\ 0 \\ 0 \\ 1 \end{pmatrix}, \\
 |\phi^-\rangle_{12} &= \frac{1}{\sqrt{2}} \begin{pmatrix} 1 \\ 0 \\ 0 \\ -1 \end{pmatrix}.
 \end{aligned} \tag{10.14}$$

When checking the orthogonality of $|\psi^\dagger\rangle_{12}$ and $|\psi^-\rangle_{12}$, we have

$$\begin{aligned}
 \langle \psi^\dagger |_{12} \psi^- \rangle_{12} &= |\psi^\dagger\rangle_{12} = \frac{1}{\sqrt{2}} \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 1 & 0 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \\ -1 \\ 0 \end{pmatrix} = \\
 &= \frac{1}{2} (1 - 1) = 0.
 \end{aligned}$$

Similarly, the orthogonality of the other states (10.13) can be shown.

Moreover, using (10.14), it can be shown that the states (10.13) are normalized; for example, for $|\psi^\dagger\rangle_{12}$ we get

$$\begin{aligned}
 \langle \psi^\dagger |_{12} \psi^\dagger \rangle_{12} &= |\psi^\dagger\rangle_{12} = \frac{1}{\sqrt{2}} \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 1 & 0 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \\ 1 \\ 0 \end{pmatrix} = \\
 &= \frac{1}{2} (1 + 1) = 1.
 \end{aligned}$$

optical axis of the crystal

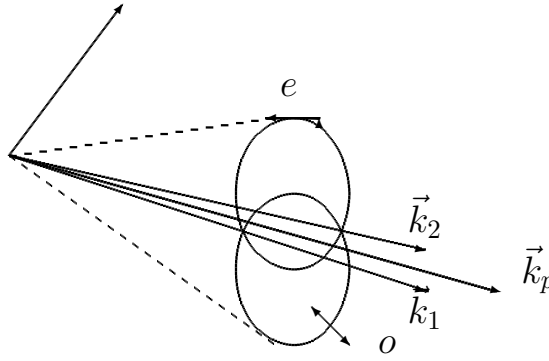


Figure 10.2: Spontaneous parametric scattering with type II phase matching. \vec{k}_p - pump wave vector, e - cone along which extraordinarily polarized photons propagate, o - cone along which ordinarily polarized photons propagate, $\vec{k}_{1,2}$ - directions in which entangled photons propagate

10.4 Generation of Bell States

To obtain entangled states, the process of parametric light scattering can be used, which we previously considered for describing the process of obtaining squeezed states of the electromagnetic field (see 9.4). At that time, we considered phase synchronism of the first kind, where the polarizations of the two photons are the same. In second-kind phase synchronism, the polarizations of the photons are different. In this case, the photons propagate along two cones, as shown in Figure 10.2.

Along one cone, the radiation is polarized as ordinary waves, and along the second as extraordinary. Thus, at the points of intersection of the cones (see Figure 10.3), the state of the light is recorded as follows:

$$|\psi\rangle = \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 + e^{i\alpha} |y\rangle_1 |x\rangle_2),$$

where the phase difference α arises due to the difference in refractive indices for ordinary and extraordinary photons. By using an additional birefringent phase plate, any phase difference α can be obtained, for example, 0 or π . As a result, we will obtain the following Bell states:

$$\begin{aligned} |\psi^+\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 + |y\rangle_1 |x\rangle_2), \\ |\psi^-\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 - |y\rangle_1 |x\rangle_2). \end{aligned} \quad (10.15)$$

Now, if a half-wave phase plate that changes the polarization to the orthogonal

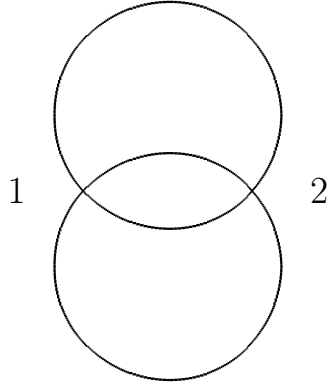


Figure 10.3: Schematic representation obtained when photographing radiation emitted by a nonlinear crystal. Numbers 1 and 2 indicate the directions in which entangled photons propagate.

one is placed in front of one of the beams, for example, the first:

$$|x\rangle_2 \rightarrow |y\rangle_2, |y\rangle_2 \rightarrow |x\rangle_2,$$

then from (10.15) we get

$$\begin{aligned} |\phi^\dagger\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |x\rangle_2 + |y\rangle_1 |y\rangle_2), \\ |\phi^-\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |x\rangle_2 - |y\rangle_1 |y\rangle_2). \end{aligned} \quad (10.16)$$

Thus, combining (10.15) and (10.16), we obtain the complete basis (10.13).

10.5 Registration of Bell States

Of the four basis states (10.13), only $|\psi^-\rangle_{12}$ is antisymmetric, i.e., it changes its sign upon particle (photon) permutation. This allows us to distinguish this state from all others. For this purpose, the scheme shown in Figure 10.4 can be used, where two photons will be detected in different photodetectors only for the antisymmetric state; for all other states, both photons are detected either in $D^{(1)}$ or $D^{(2)}$.

To prove this, let us consider how a beamsplitter acts on a single-photon state. A photon can impinge on the beamsplitter from half-plane (1) (Figure 10.5); in this case, with probability 0.5 the photon ends up in half-plane (1) and with probability 0.5 in (2). If the photon comes from half-plane (2), it again can end up in either half-plane with equal probability. Thus, to describe the beamsplitter we will use the scattering matrix expression in the form (6.4), where $r = t = \frac{1}{\sqrt{2}}$,

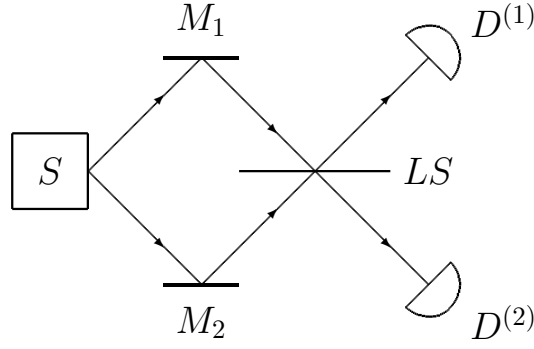


Figure 10.4: Scheme for registering Bell states. S - source of entangled photons, M_1 and M_2 - two blind mirrors, LS - beamsplitter, $D^{(1)}$ and $D^{(2)}$ - two photodetectors

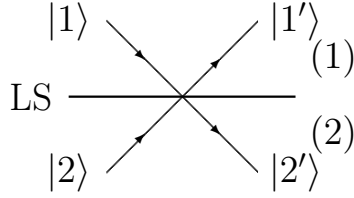


Figure 10.5: Passage of a photon through a beamsplitter

i.e., the action of the beamsplitter can be expressed by the following operator (Hadamard[34]):

$$\begin{aligned}\hat{H} |1\rangle &= \frac{1}{\sqrt{2}} (|1'\rangle + |2'\rangle), \\ \hat{H} |2\rangle &= \frac{1}{\sqrt{2}} (|1'\rangle - |2'\rangle).\end{aligned}\tag{10.17}$$

In our case, two photons fall on the beamsplitter from different sides, so only two wavefunctions are possible, describing the spatial part of our initial state:

$$\begin{aligned}|S\rangle_{12} &= \frac{1}{\sqrt{2}} (|1\rangle_1 |2\rangle_2 + |2\rangle_1 |1\rangle_2), \\ |A\rangle_{12} &= \frac{1}{\sqrt{2}} (|1\rangle_1 |2\rangle_2 - |2\rangle_1 |1\rangle_2).\end{aligned}\tag{10.18}$$

As seen from (10.18), the state $|S\rangle_{12}$ is symmetric, i.e., it remains unchanged under photon permutation, while the state $|A\rangle_{12}$ is antisymmetric (changes sign upon photon permutation).

Let us consider how the states (10.18) transform upon passing through the beamsplitter. Since the operator \hat{H} acts independently on each photon, for $|S\rangle_{12}$

we have

$$\hat{H} |S\rangle_{12} = |S'\rangle_{12} = \frac{1}{\sqrt{2}} \left(\hat{H} |1\rangle_1 \hat{H} |2\rangle_2 + \hat{H} |2\rangle_1 \hat{H} |1\rangle_2 \right),$$

from which, using (10.17), we obtain

$$\begin{aligned} \hat{H} |S\rangle_{12} &= \frac{1}{\sqrt{2}} \left(\frac{1}{2} (|1'\rangle_1 + |2'\rangle_1) (|1'\rangle_2 - |2'\rangle_2) + \right. \\ &\quad \left. + \frac{1}{2} (|1'\rangle_1 - |2'\rangle_1) (|1'\rangle_2 + |2'\rangle_2) \right) = \\ &= \frac{1}{2\sqrt{2}} (|1'\rangle_1 |1'\rangle_2 + |2'\rangle_1 |1'\rangle_2 - |1'\rangle_1 |2'\rangle_2 - |2'\rangle_1 |2'\rangle_2 + \\ &\quad + |1'\rangle_1 |1'\rangle_2 - |2'\rangle_1 |1'\rangle_2 + |1'\rangle_1 |2'\rangle_2 - |2'\rangle_1 |2'\rangle_2) = \\ &= \frac{1}{2\sqrt{2}} (|1'\rangle_1 |1'\rangle_2 - |2'\rangle_1 |2'\rangle_2 + |1'\rangle_1 |1'\rangle_2 - |2'\rangle_1 |2'\rangle_2) = \\ &\quad \frac{1}{\sqrt{2}} (|1'\rangle_1 |1'\rangle_2 - |2'\rangle_1 |2'\rangle_2). \end{aligned}$$

Thus, for the state $|S'\rangle_{12}$ both photons appear either in half-plane (1) or in half-plane (2).

For the antisymmetric state, we have

$$\begin{aligned} \hat{H} |A\rangle_{12} &= |A'\rangle_{12} = \frac{1}{\sqrt{2}} \left(\hat{H} |1\rangle_1 \hat{H} |2\rangle_2 - \hat{H} |2\rangle_1 \hat{H} |1\rangle_2 \right) = \\ &= \frac{1}{\sqrt{2}} \left(\frac{1}{2} (|1'\rangle_1 + |2'\rangle_1) (|1'\rangle_2 - |2'\rangle_2) - \right. \\ &\quad \left. - \frac{1}{2} (|1'\rangle_1 - |2'\rangle_1) (|1'\rangle_2 + |2'\rangle_2) \right) = \\ &= \frac{1}{2\sqrt{2}} (|1'\rangle_1 |1'\rangle_2 + |2'\rangle_1 |1'\rangle_2 - |1'\rangle_1 |2'\rangle_2 - |2'\rangle_1 |2'\rangle_2 - \\ &\quad - |1'\rangle_1 |1'\rangle_2 + |2'\rangle_1 |1'\rangle_2 - |1'\rangle_1 |2'\rangle_2 + |2'\rangle_1 |2'\rangle_2) = \\ &= \frac{1}{2\sqrt{2}} (|2'\rangle_1 |1'\rangle_2 - |1'\rangle_1 |2'\rangle_2 + |2'\rangle_1 |1'\rangle_2 - |1'\rangle_1 |2'\rangle_2) = \\ &\quad -\frac{1}{\sqrt{2}} (|1'\rangle_1 |2'\rangle_2 - |2'\rangle_1 |1'\rangle_2). \end{aligned}$$

Thus, in the case of the antisymmetric state, after passing through the beamsplitter, the photons will be in different half-planes.

Since photons are bosons, the total wavefunction describing both polarization and spatial properties must be symmetric [58] (see section 14.8). Thus, from the

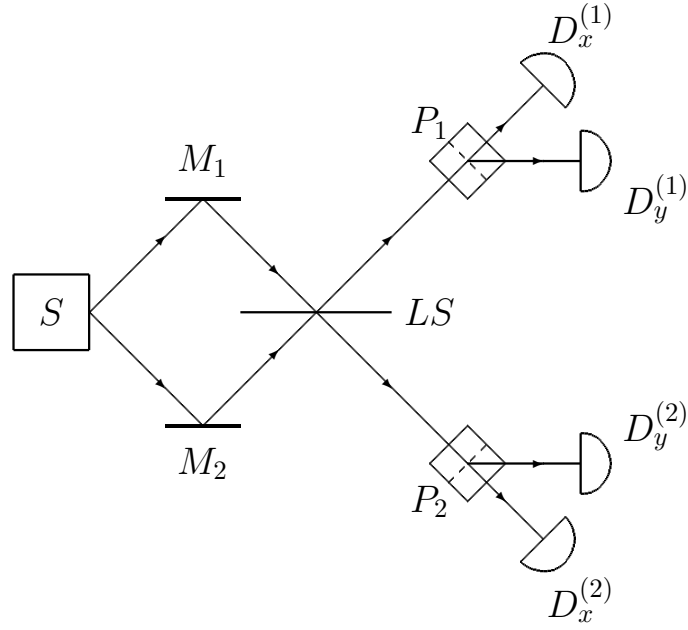


Figure 10.6: Diagram of Bell state registration. S - source of entangled photons, M_1 and M_2 - two blind mirrors, LS - beam splitter, P_1 and P_2 - Nicol prisms, $D_x^{(1)}$ and $D_x^{(2)}$ - photodetectors for x polarization, $D_y^{(1)}$ and $D_y^{(2)}$ - photodetectors for y polarization

combination of (10.13) and (10.18), only the following combinations are possible:

$$\begin{aligned} &|\psi^+\rangle_{12} |S\rangle_{12}, \\ &|\psi^-\rangle_{12} |A\rangle_{12}, \\ &|\phi^+\rangle_{12} |S\rangle_{12}, \\ &|\phi^-\rangle_{12} |S\rangle_{12}. \end{aligned}$$

Consequently, for Bell states $|\psi^+\rangle_{12}$, $|\phi^+\rangle_{12}$, and $|\phi^-\rangle_{12}$, both photons after passing through the beamsplitter will hit the same photodetector, while for $|\psi^-\rangle_{12}$ they will go to different ones, which allows us to uniquely distinguish the state $|\psi^-\rangle_{12}$ from all others.

The scheme in Figure 10.4 can be modified so that it can also register the state $|\psi^+\rangle_{12}$. To do this, note that among all symmetric Bell states only in this one do the photons have different polarizations. Thus, this state can be singled out by measuring the photon polarizations.

Figure 10.6 shows a scheme that allows registration of $|\psi^-\rangle_{12}$ and $|\psi^+\rangle_{12}$. For the state $|\psi^+\rangle_{12}$, photodetectors $D_x^{(1)}$ and $D_y^{(1)}$ or $D_x^{(2)}$ and $D_y^{(2)}$ will fire simultaneously. For $|\psi^-\rangle_{12}$, it is $D_x^{(1)}$ and $D_y^{(2)}$ or $D_x^{(2)}$ and $D_y^{(1)}$. For $|\phi^+\rangle_{12}$ and $|\phi^-\rangle_{12}$, both photons will be detected simultaneously in one of the 4 photodetectors. Recently, works have appeared [16] where registration of all four Bell states has been performed.

10.6 No-Cloning Theorem and Quantum Teleportation.

In quantum mechanics, there exists the so-called no-cloning theorem [29]. The meaning of this theorem can be explained using our example of a photon with two mutually orthogonal polarizations, whose wave function is given by relation (7.3). The no-cloning theorem states that it is impossible to create a device where a particle in the state (7.3) is input and two particles in the same state are output.

Indeed, suppose we have a device that performs state cloning. The action of this device on a photon polarized along x is described by the cloning operator \hat{D} as follows:

$$\hat{D} |D_I\rangle |x\rangle = |D_{F_x}\rangle |x\rangle |x\rangle ,$$

where $|D_I\rangle$ is the wave function describing the initial state of the cloning device, and $|D_{F_x}\rangle$ is the wave function describing the state of the device after cloning the photon with vertical x -polarization. For a photon polarized along y , we have

$$\hat{D} |D_I\rangle |y\rangle = |D_{F_y}\rangle |y\rangle |y\rangle ,$$

where $|D_{F_y}\rangle$ is the wave function describing the state of the device after cloning this photon.

If we apply the cloning operator \hat{D} to a photon in an arbitrary state (7.3), we get

$$\hat{D} |D_I\rangle |\psi\rangle = \hat{D} |D_I\rangle (\alpha |x\rangle + \beta |y\rangle) = \alpha |D_{F_x}\rangle |x\rangle |x\rangle + \beta |D_{F_y}\rangle |y\rangle |y\rangle ,$$

from which it is clearly impossible to obtain the expected result

$$\hat{D} |D_I\rangle |\psi\rangle = |D_{F_{xy}}\rangle (\alpha |x\rangle + \beta |y\rangle) (\alpha |x\rangle + \beta |y\rangle) .$$

But if the state of a photon cannot be cloned, it turns out that it can be transmitted from one point in space to another (naturally, with destruction of the original state), as demonstrated by quantum teleportation experiments.

The scheme of the quantum teleportation protocol is shown in Figure 10.7. In this figure, Alice wants to send to Bob the state of photon 1, which is described by the wave function (7.3).

$$|\psi\rangle_1 = \alpha |x\rangle_1 + \beta |y\rangle_1 ,$$

There is also an entangled photon source S that emits pairs of photons 2 and 3 in the state described by the following wave function:

$$|\psi\rangle_{23} = |\psi^-\rangle_{23} = \frac{1}{\sqrt{2}} (|x\rangle_2 |y\rangle_3 - |y\rangle_2 |x\rangle_3) .$$

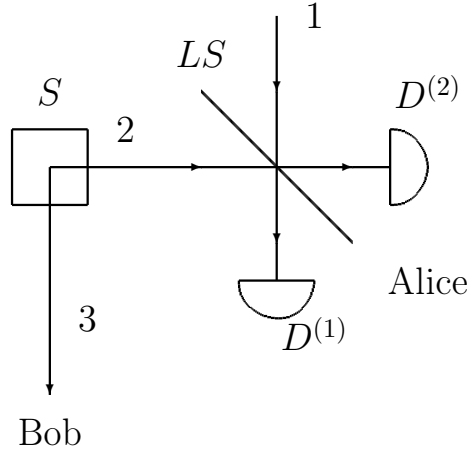


Figure 10.7: Scheme of the quantum teleportation protocol. S is the source of entangled photon pairs, LS - beamsplitter, $D^{(1)}$ and $D^{(2)}$ - two photodetectors, 1 - teleported photon, 2 and 3 - pair of entangled photons.

Alice mixes photons 1 and 2 on a beamsplitter LS and then registers a certain Bell state using detectors $D^{(1,2)}$. The easiest registration is of the state $|\psi^-\rangle_{12}$, in which both photodetectors must trigger simultaneously.

The total state of the three particles is written as

$$|\psi\rangle_{123} = |\psi\rangle_1 |\psi\rangle_{23},$$

which can be expanded into Bell states of photons 1 and 2:

$$\begin{aligned} |\psi\rangle_{123} = & c_{|\psi^\dagger\rangle_{12}} |\psi^\dagger\rangle_{12} + c_{|\psi^-\rangle_{12}} |\psi^-\rangle_{12} + \\ & + c_{|\phi^\dagger\rangle_{12}} |\phi^\dagger\rangle_{12} + c_{|\phi^-\rangle_{12}} |\phi^-\rangle_{12}. \end{aligned} \quad (10.19)$$

In the expansion (10.19), we are interested in the coefficient $c_{|\psi^-\rangle_{12}}$, since it describes the state of the third photon when Alice registers the Bell state $|\psi^-\rangle_{12}$. For the desired coefficient, we obtain:

$$\begin{aligned} c_{|\psi^-\rangle_{12}} &= \langle \psi^- |_{12} \psi \rangle_{123} = \\ &= \frac{1}{\sqrt{2}} (\langle x|_1 \langle y|_2 - \langle y|_1 \langle x|_2) (\alpha |x\rangle_1 + \beta |y\rangle_1) \frac{1}{\sqrt{2}} (|x\rangle_2 |y\rangle_3 - |y\rangle_2 |x\rangle_3) = \\ &= \frac{1}{2} (\alpha \langle y|_2 - \beta \langle x|_2) (|x\rangle_2 |y\rangle_3 - |y\rangle_2 |x\rangle_3) = -\frac{1}{2} (\alpha |x\rangle_3 + \beta |y\rangle_3). \end{aligned}$$

This means that whenever Alice registers the pair of photons 1 and 2 in the state $|\psi^-\rangle_{12}$, i.e., when both detectors $D^{(1)}$ and $D^{(2)}$ simultaneously trigger, photon 3 on Bob's side ends up in the state identical to the original photon 1 state, that is, teleportation of photon 1 to Bob occurs.

10.7 Quantum Pseudo-Telepathic Games

The nonclassicality of entangled states can be vividly demonstrated in the so-called quantum pseudo-telepathic games [7, 3]

10.7.1 Game Description

Consider the Mermin-Peres game in which two players, Alice and Bob, play against a casino. Alice and Bob fill a 3×3 square with numbers ± 1 . The casino tells Alice the number of the row that Alice must fill, and Bob the number of the column. Alice's product of all digits must be equal to $+1$, while Bob's must be -1 . The players win if the number at the intersection of the chosen row and column matches. They lose otherwise.

Alice and Bob are isolated from each other until they receive from the casino the row and column numbers, respectively. Thus, they can only agree on a strategy before this moment.

Example 10.7.1 (Mermin-Peres Game). *Suppose Alice (A) received row number 1, and Bob (B) column number 3. In this case, a winning combination may look like*

$$A = \begin{pmatrix} +1 & -1 & -1 \\ * & * & * \\ * & * & * \end{pmatrix},$$

$$B = \begin{pmatrix} * & * & -1 \\ * & * & +1 \\ * & * & +1 \end{pmatrix}.$$

This combination is winning because both Alice and Bob have the same number -1 at the intersection of the selected row and column.

In the opposite case (row 1, column 1)

$$A = \begin{pmatrix} +1 & -1 & -1 \\ * & * & * \\ * & * & * \end{pmatrix},$$

$$B = \begin{pmatrix} -1 & * & * \\ +1 & * & * \\ +1 & * & * \end{pmatrix},$$

the players lose, since Alice obtains $+1$ at the intersection, while Bob has -1 .

10.7.2 Classical Strategy

In the classical strategy, Alice and Bob cannot agree on all elements of the matrix because Alice's and Bob's conditions conflict: according to Alice, the product of all elements of the matrix equals $(+1)^3 = 1$, while according to Bob it equals $(-1)^3 = -1$. Thus, sometimes Alice and Bob win, but there will always be games in which they lose, since the probability of winning $p < 1$.

10.7.3 Quantum Strategy

As we have found, the classical strategy does not exist, but on the other hand, if Alice and Bob can create in advance some entangled state and agree on the measurements they will perform on this state ¹, then there is a way to ensure a winning probability $p = 1$.

The entangled state used by Alice and Bob:

$$|\psi\rangle = |\psi\rangle_{A_1, B_1} \otimes |\psi\rangle_{A_2, B_2}$$

The matrix used for measurements

$$X = \begin{pmatrix} \hat{S}_0 \otimes \hat{S}_3 & \hat{S}_3 \otimes \hat{S}_0 & \hat{S}_3 \otimes \hat{S}_3 \\ \hat{S}_1 \otimes \hat{S}_0 & \hat{S}_0 \otimes \hat{S}_1 & \hat{S}_1 \otimes \hat{S}_1 \\ -\hat{S}_1 \otimes \hat{S}_3 & -\hat{S}_3 \otimes \hat{S}_1 & \hat{S}_2 \otimes \hat{S}_2 \end{pmatrix} \quad (10.20)$$

Since

$$\hat{S}_1^2 = \hat{S}_2^2 = \hat{S}_3^2 = \hat{S}_0^2 = \hat{I}$$

for each element \hat{x}_{ij} of the matrix (10.20) we have $\hat{x}_{ij}^2 = \hat{I}$, and therefore upon measurement we get eigenvalues ± 1 .

Moreover, Alice and Bob can perform measurements since the operators in the same row or the same column commute, i.e., can be measured simultaneously. The values Alice obtains satisfy the requirements imposed on the row, and Bob's on the column.

Finally, the measurements performed by Alice and Bob at the matching position also coincide. This follows from the relation

$$(\hat{x}_{ij}^A \otimes \hat{x}_{ij}^B) |\psi\rangle = |\psi\rangle,$$

where \hat{x}_{ij}^A is Alice's measurement, and \hat{x}_{ij}^B is Bob's measurement. Thus the product of the results obtained by Alice and Bob is equal to 1, or, since the results can only be ± 1 , Alice's and Bob's results coincide.

TBD

¹naturally, the measurements are performed by Alice and Bob separately, each in their own laboratory

10.8 Exercises

1. Prove the orthonormality of the Bell basis states (10.13).
2. Which Bell state should Alice register in the scheme shown in Figure 10.7, to confirm the fact of teleportation in the case when the source of entangled photon pairs S produces photons in the state

$$|\psi\rangle_{23} = |\psi^\dagger\rangle_{23} = \frac{1}{\sqrt{2}} (|x\rangle_2 |y\rangle_3 + |y\rangle_2 |x\rangle_3).$$

3. Derive expressions for the coefficients $c_{|\psi^\dagger\rangle_{12}}$, $c_{|\phi^\dagger\rangle_{12}}$ and $c_{|\phi^-\rangle_{12}}$ in decomposition (10.19)

Part 4

Quantum Information Theory

Chapter 11

Introduction to Quantum Information Theory

Classical information theory was developed by Claude Shannon [23] applied to communication theory problems and is currently finding increasing application. In this section, we will consider the basic concepts and results of information theory, including its application to quantum phenomena.

11.1 Information and Entropy

The very concept of information is closely related to the measure of uncertainty of a system, i.e., what in statistical physics is called entropy. The more disordered a system is, the less information we know about it. Acquiring information reduces the measure of uncertainty. Accordingly, the more disordered the original system was, the more information we acquire about it as a result of the outcome of some experiment. Thus, as a measure of information, one can take the complexity (uncertainty) of the system. Following Shannon, we will call this measure entropy, while we will not distinguish between "amount of information measure" and "entropy," i.e., we will consider these concepts synonyms.

11.1.1 Entropy in Classical Information Theory

We begin the definition of entropy with an experiment with equally probable outcomes. Suppose we have an event \mathcal{A} , which has a equally likely outcomes. As entropy (a measure of uncertainty), it is reasonable to use some positive function that depends on the number of outcomes a and increases with increasing a :

$$H_{\mathcal{A}} = f(\mathcal{A}) = f(a).$$

In this case, it is reasonable to assume that $f(1) = 0$, since the result of an experiment with a single possible value is fully determined.

Now consider another experiment \mathcal{B} with a number of outcomes b . Obviously, for it the relation

$$H_{\mathcal{B}} = f(b).$$

holds. If we consider the event \mathcal{AB} , which consists of simultaneous occurrence of \mathcal{A} and \mathcal{B} , then the number of equally likely outcomes of such an event is $a \cdot b$, so we can write

$$H_{\mathcal{AB}} = f(ab).$$

On the other hand, we would like to express $H_{\mathcal{AB}}$ through $H_{\mathcal{A}}$ and $H_{\mathcal{B}}$. To do this, recall the property of entropy increase with increasing system complexity, i.e., the following inequality holds:

$$H_{\mathcal{AB}} \geq H_{\mathcal{A}}, H_{\mathcal{B}}. \quad (11.1)$$

If we use only the four simplest arithmetic operations to find the desired relation, then only addition and multiplication satisfy property (11.1). Moreover, if we consider the situation where the number of outcomes of event \mathcal{B} equals 1, then for the desired relation only the operation of addition can be used:

$$H_{\mathcal{AB}} = H_{\mathcal{A}} + H_{\mathcal{B}},$$

from which

$$f(ab) = f(a) + f(b).$$

Now, let's combine what we know about the desired function f

$$\begin{aligned} f(1) &= 0, \\ f(a) &> f(b), \text{ if } a > b, \\ f(ab) &= f(a) + f(b). \end{aligned} \quad (11.2)$$

There is only one continuous function that satisfies the requirements (11.2) (proof see [32]):

$$f(a) = K \log a. \quad (11.3)$$

The coefficient K in (11.3) determines the logarithm base and the unit of information measurement. For example, if we use as the unit of entropy the experiment with two equally probable outcomes (such as heads or tails in coin tossing), then (11.3) uses the logarithm base 2:

$$f(a) = \log_2 a, \quad (11.4)$$

and the corresponding unit of information is called a **bit**¹. This unit choice for measuring information is explained by the fact that often information is encoded using objects which have only two stable states (e.g., signal power exceeding a certain threshold or not). Along with the logarithm base 2, the natural logarithm is often used in (11.3):

$$f(a) = \ln a, \quad (11.5)$$

where the corresponding unit of information is called a **nat**. In what follows, we will use entropy definition in the form (11.3) with the logarithm base taken as 2.

Expression (11.3) was obtained under the assumption of equally probable outcomes. As is known, the probability of event \mathcal{A} is defined as

$$p_{\mathcal{A}} = \frac{1}{a},$$

from which entropy $H_{\mathcal{A}}$ can be written as

$$H_{\mathcal{A}} = -\log p_{\mathcal{A}}. \quad (11.6)$$

Now consider the case when event \mathcal{A} has unequal probabilities of outcomes. In this case, each outcome ξ of event \mathcal{A} should be associated with the probability of that outcome, p_{ξ} . To define the measure of information acquired upon realization of outcome ξ , we will still use expression (11.6):

$$H_{\xi} = -\log p_{\xi}. \quad (11.7)$$

The quantity H_{ξ} in (11.7) characterizes only one particular outcome ξ of the event and, accordingly, cannot be used as a measure of uncertainty of the whole event \mathcal{A} . As a measure, it is reasonable to use the averaged value of entropy:

$$H_{class} = H_{\mathcal{A}} = \sum_{\xi} -p_{\xi} \log p_{\xi}. \quad (11.8)$$

In expression (11.8) it is taken that

$$0 \cdot \log 0 = 0. \quad (11.9)$$

Despite the seeming convention in introducing (11.8), this expression indeed describes the amount of information; in particular, living organisms process information at a speed proportional to (11.8). In [32] experiments on human reaction speed to external stimuli are described. As the stimulus, one of several light bulbs illuminated, each with different frequency. Before the experiment, examinees were

¹from English **b**inary digit - binary digit

given time to train, after which it was found that reaction speed is determined precisely by expression (11.8).

Entropy (11.8) has the following property:

$$H_{min} \leq H_{\mathcal{A}} \leq H_{max}, \quad (11.10)$$

where $H_{min} = 0$ and is achieved when the probability of one of the outcomes of event \mathcal{A} equals 1 and the rest are 0 (minimal system uncertainty). $H_{max} = \log a$ is achieved when all outcomes are equally probable (maximum system uncertainty).

11.1.2 Entropy in Quantum Systems

The analog of the probability distribution $\{p_\xi\}$ in the quantum case is the density matrix $\hat{\rho}$. The averaging here uses the trace operation in some basis. That is, (11.8) in quantum systems corresponds to:

$$H_{quant} = -\text{Tr}(\hat{\rho} \log \hat{\rho}). \quad (11.11)$$

When analyzing relation (11.11), the following questions arise:

- How to compute the entropy (11.11)
- In which cases the entropy will be maximal and minimal, i.e., what is the quantum analog of the classical expression 11.10.

To answer the first question, recall that the trace operation does not depend on the representation (see subsection 14.1.7), so it is enough to consider the density matrix in the basis where it is diagonal. If we denote the diagonal elements of the density matrix by ρ_{nn} , then using (14.27) we can rewrite (11.11) as

$$H_{quant} = -\text{Tr}(\hat{\rho} \log \hat{\rho}) = -\sum_n \rho_{nn} \log \rho_{nn}. \quad (11.12)$$

Let's start analyzing expression (11.12) by considering pure states. If we take an arbitrary pure state $|\psi\rangle$, written in some basis $\{|n\rangle\}$:

$$|\psi\rangle = \sum_n c_n |n\rangle,$$

then one can always find a basis $\{|m\rangle\}$ in which $|\psi\rangle$ is one of the basis states, for example the first:

$$|\psi\rangle = \sum_m c_m |m\rangle = |m\rangle|_{m=1},$$

thus, the diagonal density matrix will have only one non-zero element $\rho_{11} = 1$, and the entropy will be

$$H_{pure} = 0, \quad (11.13)$$

where we used expression (11.9). This expression tells us that the pure state is completely determined.

If we consider a mixed state, the density matrix will be diagonal, with diagonal elements equal to the probabilities of the system being in the corresponding states:

$$\rho_{nn} = p_n.$$

If now we use (11.12), we get that the expression for entropy of the mixed state coincides with the classical entropy expression (11.8):

$$H_{misc} = - \sum_n p_n \log p_n = H_{class}, \quad (11.14)$$

which, as mentioned earlier, is maximal for uniform probability distribution and minimal for degenerate one (one outcome probability 1, the others 0).

Along with expressions (11.12) and (11.14), in which the unit of information is nat, it is convenient to use a logarithm base two - then the unit is called a qubit (**q**uantum **bit**).

11.2 Information Transmission. Communication Channel.

Such general characteristics of information as entropy do not take into account the concrete content of information, which may also be important, and therefore may not always be applicable. Nevertheless, there are situations when the specific value is unimportant, and what matters is the amount of information. As an example of such systems we can consider information transmission systems.

11.2.1 Classical Communication Channel

A typical information transmission system is shown in [Figure 11.1](#). As seen in this figure, before information is transmitted, it must be encoded, i.e., represented in states that can be transmitted through the channel. At the channel output, information must be decoded, i.e., transformed back into the form used at the input.

Encoding information with classical objects that have only two stable states (e.g., signal power exceeding or not a threshold) is the most convenient. Obviously, classical states are chosen so that they can be easily detected.

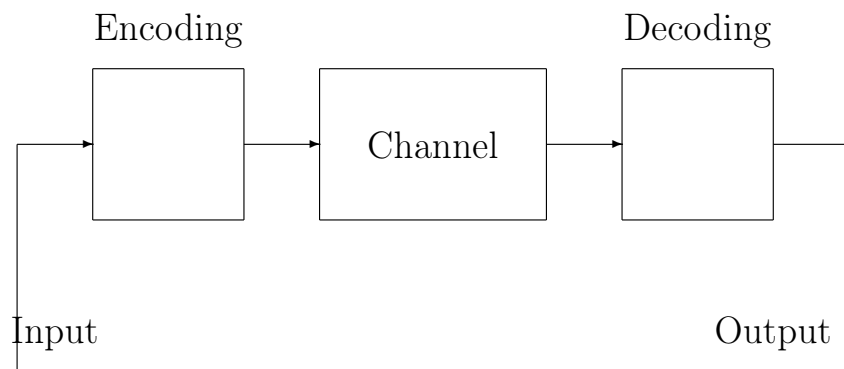


Figure 11.1: Information Transfer. Communication Channel

11.2.2 Quantum Communication Channel

Using quantum objects for information transmission significantly expands the capabilities of the communication channel because a quantum system has a much larger choice of states that can be used for transmission.

At the same time, the quantum channel itself has some special properties. For example, consider transmitting information encoded using pure quantum states. As mentioned earlier (see 10.6), an arbitrary quantum state cannot be copied, whence it follows that in such a system, unlike the classical communication channel, not only substitution of information is impossible, but even unauthorized reading. This is particularly important for information security.

11.2.3 Quantum Dense Coding

As an example of a quantum communication channel, consider the quantum dense coding protocol proposed by Bennett and Wiesner [5], [18]. This protocol uses the polarization-entangled two-photon state, considered earlier by us (see chapter 10).

Each of the photons individually has two orthogonal states $|x\rangle_{1,2}$ and $|y\rangle_{1,2}$. Thus, from the classical point of view using these two photons, one can encode four bits of information using the following states: $|x\rangle_1 |x\rangle_2$, $|x\rangle_1 |y\rangle_2$, $|y\rangle_1 |x\rangle_2$, and $|y\rangle_1 |y\rangle_2$.

In the quantum dense coding protocol, one can also encode 4 bits of information, but while operating on *only one* of the two photons. The scheme of this protocol is shown in Figure 11.2. Source S produces a pair of entangled photons in some Bell state, for example $|\Psi^\dagger\rangle_{12}$. The first photon of this pair is sent to Alice, and the second to Bob. By changing the properties of his (second) photon,

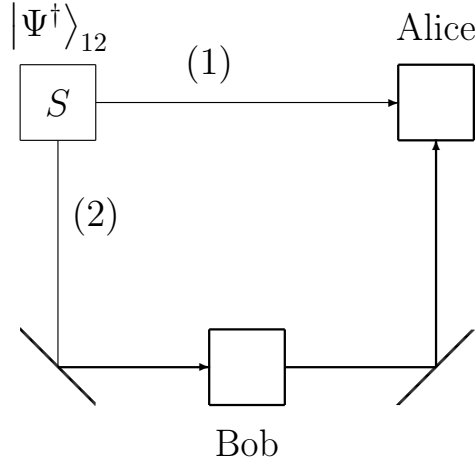


Figure 11.2: Quantum Dense Coding Protocol

Bob transforms the initial state into one of the 4 Bell states (10.13):

$$\begin{aligned}
 |\psi^+\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 + |y\rangle_1 |x\rangle_2), \\
 |\psi^-\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |y\rangle_2 - |y\rangle_1 |x\rangle_2), \\
 |\phi^+\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |x\rangle_2 + |y\rangle_1 |y\rangle_2), \\
 |\phi^-\rangle_{12} &= \frac{1}{\sqrt{2}} (|x\rangle_1 |x\rangle_2 - |y\rangle_1 |y\rangle_2).
 \end{aligned} \tag{11.15}$$

In the case of the state $|\Psi^\dagger\rangle_{12}$ no additional action is required since it is the initial:

$$|y\rangle_2 \rightarrow |y\rangle_2, |x\rangle_2 \rightarrow |x\rangle_2. \tag{11.16}$$

To obtain the state $|\psi^-\rangle_{12}$ Bob must produce a polarization-dependent phase shift:

$$|y\rangle_2 \rightarrow |y\rangle_2, |x\rangle_2 \rightarrow e^{i\pi} |x\rangle_2. \tag{11.17}$$

In the case of the state $|\phi^+\rangle_{12}$ Bob performs polarization flip to the orthogonal

$$|y\rangle_2 \rightarrow |x\rangle_2, |x\rangle_2 \rightarrow |y\rangle_2. \tag{11.18}$$

Finally, for the state $|\phi^-\rangle_{12}$ Bob performs both polarization flip and polarization-dependent phase shift:

$$|y\rangle_2 \rightarrow |x\rangle_2, |x\rangle_2 \rightarrow e^{i\pi} |y\rangle_2. \tag{11.19}$$

After Bob performs one of the transformations (11.16) - 11.19, he sends his particle to Alice. Alice, in turn, measures the received Bell state (11.15), for example using the detector described in 10.5, and determines which of the four actions Bob took. Thus, Bob can transmit 4 bits of information to Alice.

11.3 Information Encoding

11.3.1 Shannon's Coding Theorem

Suppose we have a message that may consist of a certain limited set of symbols (alphabet). Suppose our alphabet has k symbols:

$$\{a_1, \dots, a_k\} \quad (11.20)$$

Each symbol appears with some known probability $p_k = p(a_k)$. Now consider a message consisting of $n \gg 1$ symbols. We are interested in the question: is it possible to compress this message, i.e., can we encode this message in such a way that the result contains fewer than n symbols while still allowing the original message to be retrieved without loss of information?

An answer was given by Shannon, stating that the maximally compressed message will consist of nH bits, where $H = -\sum_k p_k \log p_k$ is the amount of information characterizing our alphabet. This result is known as Shannon's theorem for a noiseless channel.

As a convenient example, consider a binary alphabet, where 0 occurs with probability p , and 1 with probability $1 - p$. Thus, if we have a message of length n bits, it is possible to compress it to nH bits. It is clear that

$$n \geq nH > 0. \quad (11.21)$$

The equality in (11.21) holds only with equiprobable distribution, $p = \frac{1}{2}$. In this case,

$$nH = n \left(-\frac{1}{2} \log \frac{1}{2} - \frac{1}{2} \log \frac{1}{2} \right) = n.$$

11.3.2 Quantum Coding Theorem

What is the quantum analog of Shannon's theorem for a noiseless channel? First, consider the quantum analog of the alphabet (11.20). A quantum message is encoded by certain states

$$\{|a_1\rangle, \dots, |a_k\rangle\} \quad (11.22)$$

Each state appears with some probability $p_k = p(|a_k\rangle)$. Thus, we can write a density matrix for each symbol of our message:

$$\hat{\rho} = \sum_k p_k |a_k\rangle \langle a_k|. \quad (11.23)$$

A message consisting of n symbols has the following density matrix:

$$\hat{\rho}^n = \hat{\rho} \otimes \cdots \otimes \hat{\rho}.$$

Can this message be compressed, and if so, to what extent? The answer is given by the quantum coding theorem for a noiseless channel (Schumacher's theorem). The limit is achieved using

$$nH = -nSp(\hat{\rho} \log \hat{\rho}),$$

As an example, consider encoding with the polarization states of a photon. Suppose our alphabet consists of two states:

1. $|x\rangle$ - photon polarized along the x axis
2. $|y\rangle$ - photon polarized along the y axis

Assume state $|x\rangle$ occurs with probability $p_{|x\rangle} = p$, and state $|y\rangle$ with probability $p_{|y\rangle} = 1 - p_{|x\rangle} = 1 - p$. Thus the density matrix (11.23) will have the form

$$\hat{\rho} = p|x\rangle \langle x| + (1 - p)|y\rangle \langle y|. \quad (11.24)$$

In this case, it is convenient to use a qubit as the unit of information, and accordingly, if the original message is coded by n photons, it can be compressed to $nH = -nSp(\hat{\rho} \log \hat{\rho})$ qubits, i.e., it is possible to use fewer than n photons to transmit this message. The main difference in the quantum case from the classical case is that compression of information is possible even when $p = \frac{1}{2}$.

11.4 Exercises

1. Prove the quantum expression for entropy (11.12).

Chapter 12

Cryptography

From the moment the importance of information was realized, means of its protection began to appear.

New encryption methods were invented, such as the Caesar cipher, in which each letter of the alphabet was replaced with another one (for example, the one three positions further in the alphabet). Alongside new encryption methods, ways to crack these ciphers appeared, such as using statistical properties of the language in which the original message was written to crack the Caesar cipher.

Very often, the security of a cipher was ensured by keeping the encryption algorithm secret, as in the case of the Caesar cipher discussed above. In modern classical cryptography, algorithms are usually published and available for anyone to study. Secrecy is maintained by mixing the message with a secret key according to a certain open algorithm.

Suppose we need to transmit a message from Alice to Bob over a secure communication channel. The message must be presented in some digital form. The protocol describing this transmission involves several stages. Initially, Alice and Bob must obtain a common random sequence of numbers, which will be called a key. This process is called key distribution.

In the next stage, Alice must use a certain algorithm E to obtain an encrypted message C from the original message P and key K . This process can be described by the following equation:

$$E_K(P) = C. \quad (12.1)$$

In the third stage, the obtained encrypted message must be transmitted to Bob.

In the final stage, Bob, using a known algorithm D and the key K obtained in the first stage, must recover the original message P from the received encrypted message C . This process can be described by the following equation:

$$D_K(C) = P. \quad (12.2)$$

When analyzing this protocol, the following questions arise. How to implement secure key distribution? Secondly, is there an absolutely secure algorithm? And finally, is it possible to safely transmit an encrypted message such that it cannot be intercepted or altered?

Classical cryptography provides a clear answer only to the second question. An absolutely secure algorithm exists—it is called the one-time pad. A detailed description of this algorithm is presented below.

12.1 One-Time Pad

The one-time pad scheme was proposed in 1917 by Major J. Mauborgne and G. Vernam. The classic one-time pad consists of a set of random keys, each equal in size to the message being sent and used only once.

Let's assume we want to encrypt a message in some language (e.g., in English). The number of symbols (letters) used in the alphabet is denoted as X . For the English language (without punctuation and case sensitivity) $X = 26$. Then we assign a number c to each symbol of the language such that $0 \leq c < X$. For example, in English, we can write

$$\begin{aligned} A &\rightarrow 0 \\ B &\rightarrow 1 \\ &\dots \\ Z &\rightarrow 25 \end{aligned}$$

The encryption procedure (12.1) is described by the following expression

$$E_{K_i}(P_i) = P_i + K_i \mod X = C_i, \quad (12.3)$$

where i is the index of the encrypted symbol.

The decryption procedure (12.2) is described by the following expression

$$D_{K_i}(C_i) = C_i - K_i \mod X = P_i, \quad (12.4)$$

where i is the index of the encrypted symbol.

This procedure easily generalizes to the case of binary data, where instead of modular addition, the XOR operation ($a \oplus b$) is used for both encryption and decryption:

Claude Shannon showed [24] that if the key is truly random, has the same length as the original message, and is not reused, the proposed one-time pad scheme is perfectly secure.

According to Shannon, perfect security can be defined as follows.

| a | b | $a \oplus b$ |
|-----|-----|--------------|
| 0 | 0 | 0 |
| 0 | 1 | 1 |
| 1 | 0 | 1 |
| 1 | 1 | 0 |

Table 12.1: XOR $a \oplus b$

Definition 12.1.1. A cipher (E, D) is perfectly secure if for any two messages of the same length m_0 and m_1 , some ciphertext c , and key $k \leftarrow_R K$, the probabilities that the original text is m_0 or m_1 are equal:

$$P(E(m_0, k) = c) = P(E(m_1, k) = c)$$

Rephrasing this definition, one can say that from the initial statistics of the ciphertext, no information about the original message can be obtained.

Theorem 12.1.1 (Cryptographic Strength of the One-Time Pad). *The one-time pad scheme is perfectly secure.*

Proof. Let $|K|$ be the number of all possible keys of length l . Where l is also the length of the original messages: $|m_{0,1}| = l$. Since the key used to encrypt the message is uniquely defined:

$$k_{0,1} = c \oplus m_{0,1},$$

we obtain for the probabilities

$$P(E(m_0, k) = c) = P(E(m_1, k) = c) = \frac{1}{|K|}.$$

□

12.2 Problems of Classical Cryptography

If there exists a perfectly secure cryptographic system (one-time pad), what is wrong with classical cryptography? The problem lies in obtaining keys that meet the requirements of the one-time pad (the key length is equal to the message length, the key consists of random data, and it is not reused) and transferring these keys to Bob and Alice.

Problems arise both at the stage of key generation,¹ and at the stage of transferring these keys.

¹obtaining large sequences of random numbers is a non-trivial mathematical task

In classical cryptography, so-called public key algorithms are used to transfer keys. There are several key exchange protocols based on public key cryptographic systems. They are all based on the existence of two keys: the first, called the public key, is used only for encryption, and the second, the private key, is used for decryption. To obtain the private key from the public one, a complex mathematical operation is required. For example, the security of one of the most popular public key systems—RSA (see 19.6)—is based on the difficulty of factoring² large numbers.

A key distribution protocol scheme based on public key cryptography can be described as follows. In the first stage, Alice creates public and private keys and sends the first to Bob. Bob, for his part, creates the key that Alice and Bob want to have (which needs to be distributed). This key is encrypted (for example, via RSA) using Alice's public key and sent to her. Alice, receiving this encrypted key, can decrypt it with her private key.

If an intruder (Eve) wants to learn the transmitted key, she must solve a complex mathematical problem of factoring large numbers. It is believed, but not proven, that the complexity of factoring grows exponentially with the number of digits in the number [34].³ Thus, as the number of digits increases, the task quickly becomes unsolvable.

This scheme has several problems. The first is that the complexity of factoring is not proven. Moreover, there are algorithms for quantum computers—Shor's algorithm (see 13.3)—that solve the factoring problem for number N in time $O(\log N)$, i.e., on the order of the number of digits in N . Thus, at the moment a quantum computer is built, all systems based on RSA will become obsolete.

12.3 Quantum Cryptography

The randomness inherent to quantum objects suggests using them for key distribution. There are many different secure key distribution schemes based on the use of quantum objects. We explore a scheme based on testing Bell inequalities (see 10.2). The considered scheme is depicted in Figure 12.1.

The source of entangled photons S creates pairs of photons, one of which is sent to Alice and the other to Bob, who perform measurements of the Stokes parameters of their photons.

Alice randomly measures either $\hat{A} = \hat{S}_1^{(1)}$ or $\hat{A}' = \hat{S}_2^{(1)}$. Bob randomly per-

²decomposition into prime factors

³The fastest known algorithm solves the factoring problem for number N in time approximately $O\left(\exp\left(\log^{\frac{1}{3}} N (\log \log N)^{\frac{2}{3}}\right)\right)$.

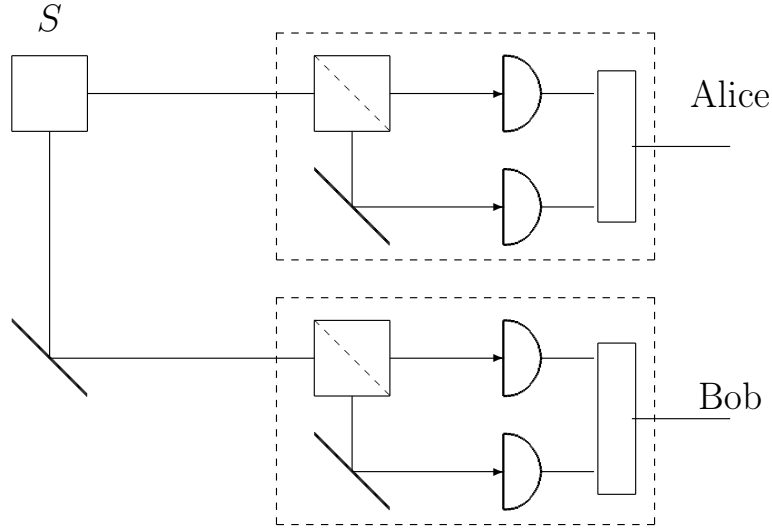


Figure 12.1: Key distribution scheme based on Bell inequalities verification

forms measurements of the following quantities

$$\begin{aligned}\hat{B} &= \frac{1}{\sqrt{2}} \left(\hat{S}_1^{(2)} + \hat{S}_2^{(2)} \right), \\ \hat{B}' &= \frac{1}{\sqrt{2}} \left(\hat{S}_1^{(2)} - \hat{S}_2^{(2)} \right) \\ \hat{C} &= \hat{S}_1^{(2)}, \\ \hat{C}' &= \hat{S}_2^{(2)}.\end{aligned}$$

Thus, as a result of the experiment, we will obtain 8 pairs of values that can be combined into three groups.

The first group contains correlating combinations of operators \hat{A} , \hat{A}' , \hat{C} , and \hat{A}' : (a, c) and (a', c') . For these combinations, we can say that if one of the numbers a or a' is equal to ± 1 , then the other (c or c') is equal to ∓ 1 . Therefore, these numbers can be used to obtain a random sequence of numbers, which will be subsequently used as a key.

The second group contains 4 pairs of values that will be used to verify Bell inequalities: (a, b) , (a', b) , (a, b') , and (a', b') .

The last group contains pairs of values (a, c') and (a', c) . This group of values is subsequently discarded.

At the initial stage, Bob and Alice randomly (independently of each other) perform measurements. After completing the measurements, they tell each other (via a regular open communication channel) which quantities they measured in each specific trial, without disclosing the results of the measurements themselves. Subsequently, the trials where either of them could not register a photon, and the results relating to the third group of measurements, are discarded. The results of

the second group of measurements are openly published, and the average value $\langle F \rangle$ (10.8) is calculated from them.

If the obtained value in absolute terms is close to the predictions of quantum mechanics (10.9)

$$\langle F \rangle_{quant} = -\sqrt{2},$$

then the result of the first group of measurements can be interpreted as a key.

If an eavesdropper Eve wants to try to learn the results of the first group of measurements, which constitute the distributed key, one way to do so is to substitute herself as the source of entangled photon pairs and send Alice and Bob a pair of photons with certain polarization properties, so their experimental results will be predetermined. However, in this case $\langle F \rangle$, according to (10.12), will lie in the interval:

$$-1 \leq \langle F \rangle_{class} \leq 1.$$

Thus, Alice and Bob can determine the fact of Eve's interference and consider the obtained key compromised if the check establishes the fact of Eve's presence.

Chapter 13

Quantum Computations

Algorithms play a significant role in computational techniques. An algorithm represents a sequence of steps necessary to obtain an answer to a certain problem. Each problem is characterized by some number, which determines its size. The complexity of an algorithm is evaluated as the number of elementary operations necessary to solve the given problem. Obviously, in most cases (but not always), this number grows with the size of the problem.

Example 13.0.1. *Array Element Search* The task is to find an element of an array that satisfies certain conditions. The size of the problem is the number of elements in the array N .

In the general case (unstructured data array), the search is conducted through a simple enumeration. This search requires a number of operations (comparisons) that grows linearly with the size of the array $O(N)$.

In the case of structured data, the number of operations required for the search can be reduced. For example, in the case of a sorted array, the complexity of the task grows as $O(\log N)$.

Moreover, the existence of an algorithm does not guarantee its practical implementability. In particular, algorithms requiring an exponential number of steps based on the size of the original problem are considered practically unfeasible, despite the theoretical existence of a solution.

One example is the problem of factoring a natural number, i.e., the problem of decomposing it into prime factors (see example [13.0.2](#)).

Example 13.0.2. *Factorization of Natural Numbers* The task is to find the decomposition of a number into prime factors. The size of the problem is the digit length of the original number. For example, in the case of a digit length $r = 4$: $1 \leq N = 15 \leq 2^r = 2^4 = 16$). The result can be found easily and quickly: $15 = 3 \cdot 5$.

With an increase in the number of digits r , the number of operations necessary for factorization in classical algorithms grows as $O(2^r)$, which for the case $r = 1000 - 2000$ means the practical impossibility of factorizing such numbers.

Quantum objects have properties different from classical objects; accordingly, algorithms based on quantum objects may, in some cases, have characteristics inaccessible to classical algorithms. For example, Grover's quantum algorithm [11] solves the search problem in an unstructured data array (see example 13.0.1) using $O(\sqrt{N})$ operations. Shor's algorithm [25] allows solving the problem of number factorization (see example 13.0.2) using a linear number of operations $O(r)$.

13.1 Basic Principles of Quantum Computing

13.1.1 Information Representation. Classical and Quantum States

The main difference between quantum and classical computers lies in how they store information.

In the classical case, information is stored in certain memory cells. The state of each memory cell is described by a single number that can take the value 0 or 1. If m memory cells are combined, the overall state of the classical system (which it can assume at a specific moment in time) is described by m numbers.

In the quantum case, a memory cell is represented by a qubit, which requires two complex numbers α_0 and α_1 to describe it ¹:

$$|\psi\rangle_1 = \alpha_0 |0\rangle + \alpha_1 |1\rangle.$$

To describe a composite system consisting of m qubits, 2^m complex numbers are needed. Alternatively, one could say that a quantum state contains all possible classical states as a superposition. For example, consider a system consisting of 3 qubits:

$$\begin{aligned} |\psi\rangle_3 = & \alpha_0 |000\rangle + \alpha_1 |001\rangle + \alpha_2 |010\rangle + \alpha_3 |011\rangle + \\ & + \alpha_4 |100\rangle + \alpha_5 |101\rangle + \alpha_6 |110\rangle + \alpha_7 |111\rangle. \end{aligned} \quad (13.1)$$

As can be seen, any classical state of a 3-bit system is represented as one of the terms in the superposition (13.1). For example, the number $5_{10} = 101_2$ is included in (13.1) with the coefficient α_5 .

¹More accurately, three real numbers are required because $\alpha_{0,1}$ are subject to the constraint $|\alpha_0|^2 + |\alpha_1|^2 = 1$, from which, considering $\alpha_{0,1} = r_{0,1}e^{i\theta_{0,1}}$, we obtain $r_0^2 = 1 - r_1^2$

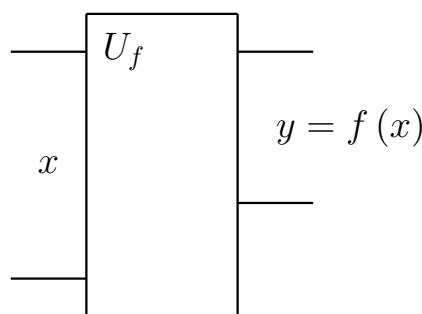


Figure 13.1: Classical computations. The input is a number x consisting of n bits, and the output is the result $y = f(x)$ described by m bits

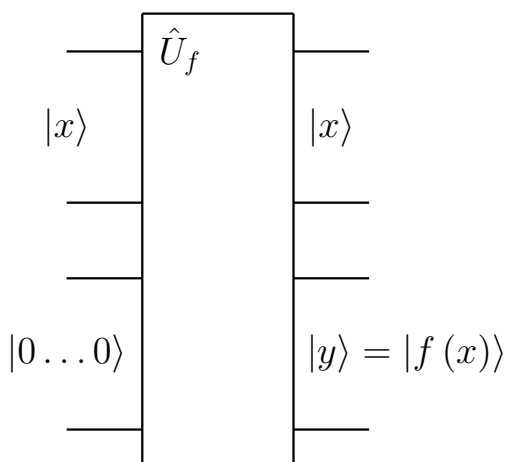


Figure 13.2: Quantum reversible computations. The input is a number $|x\rangle$ consisting of n qubits and a seed of zero states (m qubits), and the output is the result $|y\rangle = |f(x)\rangle$ described by m qubits and the original state $|x\rangle$

13.1.2 Reversible Computations

In the classical case, computation involves transforming the initial n bits into a result described by m bits (see [Figure 13.1](#)). This transformation is defined by some function $f(x)$. A typical example is addition modulo 2 (see [Table 12.1](#)), where the input is 2 bits ($n = 2$), and the output is 1 bit ($m = 1$).

Such a scheme will not work in the quantum case primarily because the change of pure quantum states over time must occur via a unitary evolution operator ([14.35](#)), i.e., it must be reversible, which is impossible for our classical example ². Therefore, quantum computations use a different scheme (see [Figure 13.2](#)) where reversible computations are possible.

At the input, together with the initial data x described by n qubits, m qubits in the state $|0\rangle$ are provided, so that the total number of inputs and outputs corresponds to each other. Consequently, the relationship between input and

²it is impossible to obtain two bits of initial information from one bit (the result)

output can be described in the form ³

$$\underbrace{|x\rangle}_n \underbrace{|f(x)\rangle}_m = \hat{U}_f \underbrace{|x\rangle}_n \underbrace{|0\dots 0\rangle}_m. \quad (13.2)$$

13.2 Quantum Logic Elements

How can an element be constructed that performs the transformation \hat{U}_f (13.2)? There is a set of elements from which an element performing the necessary transformation \hat{U}_f can be constructed with a given accuracy. Such sets are called universal.

13.2.1 Universal Set of Quantum Gates

Identity Transformation

$$\hat{\sigma}_0 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

Negation

$$\hat{\sigma}_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

Phase Shift

$$\hat{\sigma}_2 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

³More accurately written in a general form as $\underbrace{|x\rangle}_n \underbrace{|f(x)\rangle}_m \underbrace{|r\rangle}_k = \hat{U}_f \underbrace{|x\rangle}_n \underbrace{|0\dots 0\rangle}_{m+k}$, where $|r\rangle$ is a remainder of size k qubits that is not used in the calculations and serves the purpose of ensuring the unitarity of the operator \hat{U}_f

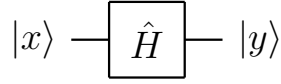
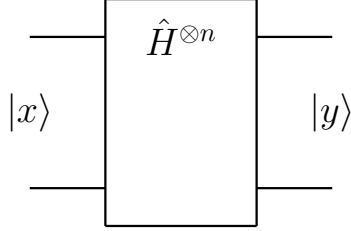


Figure 13.3: Hadamard Transformation on One Qubit

Figure 13.4: Hadamard transform $\hat{H}^{\otimes n}$ on multiple qubits

Hadamard Transformation

One of the basic quantum logic elements is the Hadamard transformation (see [Figure 13.3](#)), which is defined by the following relations

$$\begin{aligned}\hat{H} |0\rangle &= |+\rangle = \frac{|0\rangle + |1\rangle}{\sqrt{2}}, \\ \hat{H} |1\rangle &= |-\rangle = \frac{|0\rangle - |1\rangle}{\sqrt{2}},\end{aligned}$$

In matrix form this transformation can be written as

$$\hat{H} = \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & 1 \\ 1 & -1 \end{pmatrix}, \quad (13.3)$$

where the basis vectors are chosen as

$$|0\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$

and

$$|1\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}.$$

From (13.3), we can obtain the following property of the operator \hat{H} :

$$\hat{H}\hat{H} = \frac{1}{2} \begin{pmatrix} 1 & 1 \\ 1 & -1 \end{pmatrix} \begin{pmatrix} 1 & 1 \\ 1 & -1 \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}. \quad (13.4)$$

This transformation is used to obtain a superposition of states containing all possible values of the argument of the computed function (see [Figure 13.4](#)).

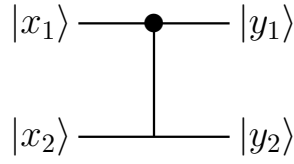


Figure 13.5: Controlled CNOT Element

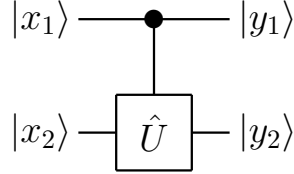


Figure 13.6: Control element

CNOT

The transformation matrix is as follows

$$CNOT = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \end{pmatrix}$$

This gate (see [Figure 13.5](#)) is used for two qubits and inverts the state of the second qubit only if the first qubit is equal to one.

Thus, if our initial state of two qubits was

$$|\psi_i\rangle = a|00\rangle + b|01\rangle + c|10\rangle + d|11\rangle$$

it transforms to

$$CNOT|\psi_i\rangle = |\psi_f\rangle = a|00\rangle + b|01\rangle + c|11\rangle + d|10\rangle$$

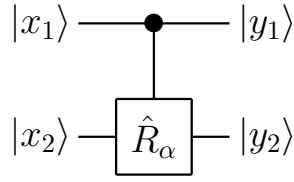
Universal Set

Definition 13.2.1 (Universal Set of Quantum Gates). A set of quantum gates is called universal if any unitary transformation can be approximated to a given accuracy by a finite sequence of gates from this set.

Theorem 13.2.1 (Kitaev). *The set $\hat{\sigma}_0, \hat{\sigma}_1, \hat{\sigma}_2, \hat{H}, CNOT$ is universal.*

Proof. TBD

□

Figure 13.7: Controlled phase shift \hat{R}_α

13.2.2 Control Elements

13.3 Shor's Algorithm

One of the most popular RSA encryption algorithms (see [section 19.6](#)) is based on the assumption of the complexity of factoring (decomposing into prime factors) large numbers. Accordingly, algorithms that enable decomposition into prime factors are of particular interest. Below is a description of such an algorithm proposed by Shor [\[25\]](#).

13.3.1 Factoring Numbers and Finding the Period of Functions

The task of factoring a number N is closely related to finding the period of functions. Consider the following function, which is called the exponentiation function modulo

$$f(x, a) = a^x \mod N. \quad (13.5)$$

Function (13.5) depends on the analyzed number N and two arguments x and a . The argument a is chosen under the following conditions

$$\begin{aligned} 0 < a < N, \\ \text{GCD}(N, a) &= 1. \end{aligned} \quad (13.6)$$

A typical graph of the function (13.5) is shown in [Figure 13.8](#).

The conditions for choosing coefficient a (13.6) are such that a and N do not have common divisors. If such divisors exist, they are the desired solution for the factoring problem and can be easily found using the Euclidean algorithm (see [section 19.1](#)).

Function (13.5) is periodic, i.e., there exists a number r such that $f(x + r, a) = f(x, a)$. The smallest possible number r is called the period of the function (13.5).

To prove periodicity, note that $f(x, a)$ cannot be zero. Indeed, if the condition $f(x, a) = 0$ is fulfilled, then

$$\exists x \in \{0, 1, \dots\} : a^x = k \cdot N,$$

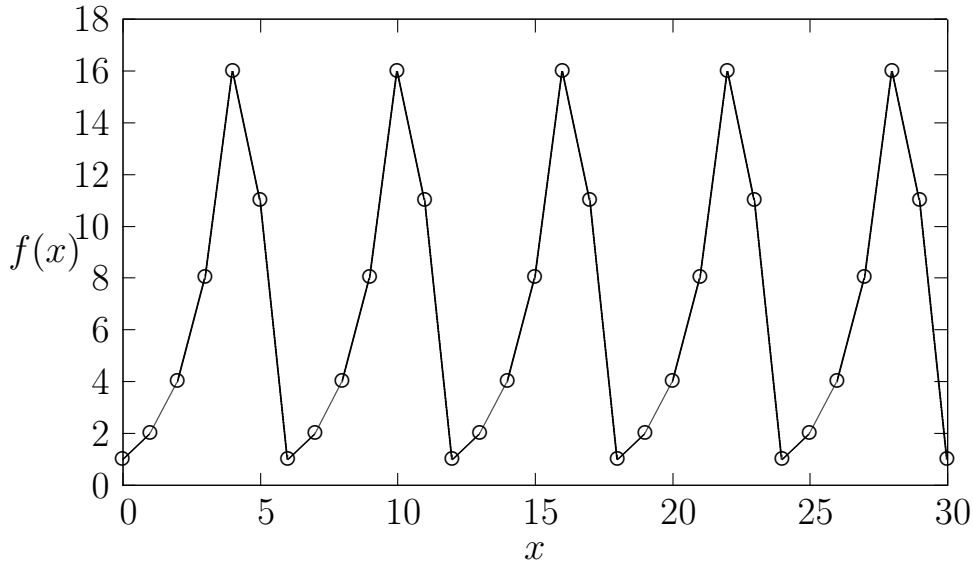


Figure 13.8: Plot of the function $f(x, a) = a^x \bmod N$ for $a = 2$, $N = 21$. The period of the function is $r = 6$.

where k is an integer, which is not possible due to the coprimality of a and N (13.6)⁴.

Thus, the range of values of function (13.5) is limited to the set

$$f(x, a) \in \{1, \dots, N-1\},$$

from which

$$\exists k, j : k > j, k, j \in \{0, 1, \dots, N\}, f(k, a) = f(j, a),$$

which proves the periodicity of the function (13.5).

Let $k = j + r$, then

$$a^k \bmod N = a^{j+r} \bmod N = a^j a^r \bmod N = a^j \bmod N,$$

since a and N are coprime, we can write

$$a^r \equiv 1 \bmod N. \quad (13.7)$$

The period of the function (13.5) can be either even or odd. In Shor's algorithm, we are interested in the first case: period - even number. Otherwise, a new number a is chosen and the period is found again. Thus, considering $r = 2 \cdot l$ we can rewrite (13.7) as

$$a^{2 \cdot l} \equiv 1 \bmod N,$$

⁴It is assumed here, obviously, that $N > 1$

and since r is the smallest number satisfying the periodicity condition,

$$a^l \not\equiv 1 \pmod{N}.$$

If, in addition, a number a is selected such that

$$a^l \not\equiv -1 \pmod{N},$$

then we have

$$(a^l - 1)(a^l + 1) = k \cdot N, \quad (13.8)$$

where k is some integer. From (13.8) we find that $a^l \pm 1$ have nontrivial common divisors (other than 1) with N .

Example 13.3.1. Finding divisors of the number $N = 21$ *As an example, consider the task of finding the divisors of the number $N = 21$. Choosing $a = 2$, we get the period of function (13.5) $r = 6$ (see Figure 13.8). Obviously,*

$$2^3 \equiv 8 \pmod{21} \not\equiv -1 \pmod{21}.$$

Thus, finding the corresponding greatest common divisors solves the task

$$\begin{aligned} \text{GCD}(2^3 - 1, 21) &= \text{GCD}(7, 21) = 7, \\ \text{GCD}(2^3 + 1, 21) &= \text{GCD}(9, 21) = 3, \\ 21 &= 7 \cdot 3. \end{aligned}$$

Thus, the task of factoring the number N can be reduced to the task of finding the period of a certain function using the following algorithm:

Algorithm 1 Shor's Algorithm

$a \leftarrow 0$

repeat

 Choose a new number a such that $0 < a < N$

if $\text{GCD}(a, N) \neq 1$ **then**

return a

end if

 Find the period r of the function $f(x, a) = a^x \pmod{N}$

until $(r \not\equiv 0 \pmod{2})$ **or** $(a^{\frac{r}{2}} \equiv -1 \pmod{N})$

return $\text{GCD}(a^{\frac{r}{2}} \pm 1, N)$

13.4 Quantum Fourier Transform

For the analysis of periodic sequences (functions), discrete Fourier transform can be used (see [section 18.2](#)), which is defined by the following relation ([18.4](#)):

$$\tilde{X}_k = \sum_{m=0}^{M-1} x_m e^{-\frac{2\pi}{M} k \cdot m}, \quad (13.9)$$

where the original sequence of numbers $\{x_m\}$ has M members.

13.4.1 Quantum Fourier Transform Scheme

Quantum Fourier transform ⁵ deals with states of the form

$$|x\rangle = \sum_{k=0}^{M-1} x_k |k\rangle, \quad (13.10)$$

where there is a sequence of amplitudes $\{x_k\}$ which defines the original sequence for the Fourier transform ([13.9](#)). The basis vector $|k\rangle$ denotes the member number of this sequence.

Obviously, the members of the sequence ([13.10](#)) must satisfy the normalization condition

$$\sum_k |x_k|^2 = 1.$$

Let's assume that a certain operator \hat{F}^M (quantum Fourier transform operator) transforms the basis vector $|k\rangle$ according to the rule given by ([18.4](#)):

$$\hat{F}^M |k\rangle = \frac{1}{\sqrt{M}} \sum_{j=0}^{M-1} e^{-i\omega k j} |j\rangle_{inv} \quad (13.11)$$

The systems of basis vectors $\{|k\rangle\}$ and $\{|k\rangle_{inv}\}$ represent the same set of vectors but are numbered differently.

⁵The work [[15](#)] was used for the analysis of the quantum Fourier transform scheme

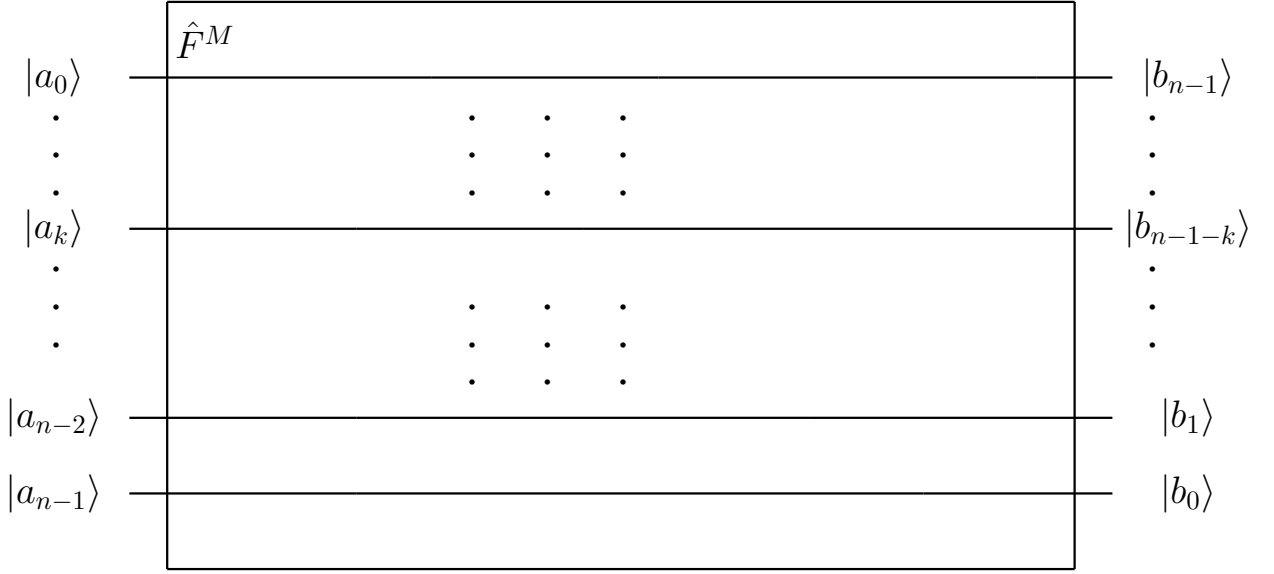


Figure 13.9: Quantum Fourier transform scheme based on the fast Fourier transform algorithm. Input data and output data

From (13.10) and (13.11) we get

$$\begin{aligned}
 \hat{F}^M |x\rangle &= \sum_{j=0}^{M-1} x_j \hat{F}^M |j\rangle = \\
 &= \frac{1}{\sqrt{M}} \sum_{k=0}^{M-1} \sum_{j=0}^{M-1} e^{-i\omega k j} x_k |j\rangle_{inv} = \\
 &= \sum_{j=0}^{M-1} \left\{ \frac{1}{\sqrt{M}} \left(\sum_{k=0}^{M-1} e^{-i\omega k j} x_k \right) \right\} |j\rangle_{inv} = \\
 &= \sum_{j=0}^{M-1} \tilde{X}_j |j\rangle_{inv} = |\tilde{X}\rangle_{inv},
 \end{aligned}$$

where

$$\tilde{X}_j = \tilde{X}_j^M = \frac{1}{\sqrt{M}} \sum_{k=0}^{M-1} e^{-i\omega k j} x_k. \quad (13.12)$$

The expression (13.12) repeats the classical analogue (18.4), i.e., it can be written as

$$|x\rangle \longleftrightarrow |\tilde{X}\rangle_{inv}.$$

Suppose now that a state of the form (13.10) is fed into our system, representing a superposition of M basis states $\{|k\rangle\}$ (see Figure 13.9). Assume the

number of basis states is a power of two, i.e., the basis state can be represented as a tensor product of $n = \log_2 M$ qubits:

$$|k\rangle = |a_0^{(k)}\rangle \otimes |a_1^{(k)}\rangle \otimes \cdots \otimes |a_{n-1}^{(k)}\rangle,$$

where

$$k = a_0^{(k)} + 2^1 a_1^{(k)} + \cdots + 2^{n-1} a_{n-1}^{(k)},$$

$$a_i^{(k)} \in \{0, 1\}.$$

At the output (see [Figure 13.9](#)), we have a superposition of M basis states $\{|j\rangle_{inv}\}$, where for the state $|j\rangle_{inv}$ we obtain

$$|j\rangle_{inv} = |b_{n-1}^{(j)}\rangle \otimes |b_{n-2}^{(j)}\rangle \otimes \cdots \otimes |b_0^{(j)}\rangle,$$

where

$$j = b_0^{(j)} + 2^1 b_1^{(j)} + \cdots + 2^{n-1} b_{n-1}^{(j)},$$

$$b_i^{(j)} \in \{0, 1\}.$$

From formula (18.10), it can be noted that if we have an input signal x consisting of $n = \log_2 M$ bits, then the bit $a_0^{(k)}$ can be used to select even (first term of the sum (18.10)) or odd (second term of the sum (18.10)).

Indeed, excluding $a_0^{(k)}$, the state (13.10) can be represented as a sum of even and odd components:

$$\begin{aligned} |x\rangle &= \sum_{k=0}^{M-1} x_k |k\rangle = \sum_{k=0}^{M-1} x_k |a_0^{(k)}\rangle \otimes |a_1^{(k)}\rangle \otimes \cdots \otimes |a_{n-1}^{(k)}\rangle = \\ &= \sum_{m=0}^{\frac{M}{2}-1} x_{k=2m} |0\rangle \otimes |a_1^{(k)}\rangle \otimes \cdots \otimes |a_{n-1}^{(k)}\rangle + \\ &+ \sum_{m=0}^{\frac{M}{2}-1} x_{k=2m+1} |1\rangle \otimes |a_1^{(k)}\rangle \otimes \cdots \otimes |a_{n-1}^{(k)}\rangle = \\ &= \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |0\rangle \otimes |m\rangle + \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |1\rangle \otimes |m\rangle = \\ &= \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |2m\rangle + \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |2m+1\rangle, \end{aligned}$$

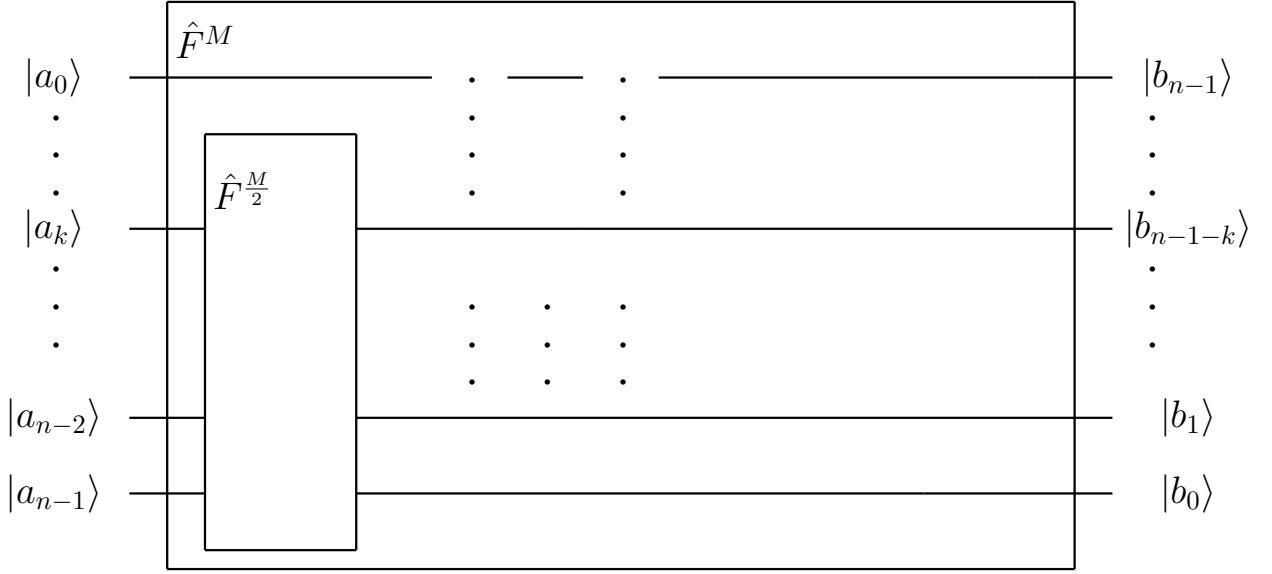


Figure 13.10: Quantum Fourier transform scheme based on the fast Fourier transform algorithm. Step 1: $|x\rangle \rightarrow \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |2m\rangle + \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |2m+1\rangle$

where

$$m = a_1^{(k)} + 2^1 a_2^{(k)} + \dots + 2^{n-2} a_{n-1}^{(k)}.$$

Applying the Fourier transform only to the higher bits $\hat{F}^{\frac{M}{2}}$, i.e., excluding $a_0^{(k)}$, we get (see [Figure 13.10](#)):

$$\begin{aligned}
 |x\rangle &\rightarrow \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |2m\rangle + \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |2m+1\rangle = \\
 &= \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |0\rangle \otimes |m\rangle + \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |1\rangle \otimes |m\rangle = \\
 &= \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |0\rangle \otimes \hat{F}^{\frac{M}{2}} |m\rangle + \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |1\rangle \otimes \hat{F}^{\frac{M}{2}} |m\rangle. \tag{13.13}
 \end{aligned}$$

Considering the expression (13.11), we get

$$\hat{F}^{\frac{M}{2}} |m\rangle = \sqrt{\frac{2}{M}} \sum_{j=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} |j\rangle_{inv}.$$

Thus, for (13.13), we have

$$\begin{aligned}
|x\rangle &\rightarrow \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |0\rangle \otimes \hat{F}^{\frac{M}{2}} |m\rangle + \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |1\rangle \otimes \hat{F}^{\frac{M}{2}} |m\rangle = \\
&= \sqrt{\frac{2}{M}} \sum_{j=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |0\rangle \otimes |j\rangle_{inv} + \\
&+ \sqrt{\frac{2}{M}} \sum_{j=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |1\rangle \otimes |j\rangle_{inv} = \\
&= \sum_{j=0}^{\frac{M}{2}-1} \left(\sqrt{\frac{2}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} x_{2m} \right) |j\rangle_{inv} + \\
&+ \sum_{j=0}^{\frac{M}{2}-1} \left(\sqrt{\frac{2}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} x_{2m+1} \right) \left| \frac{M}{2} + j \right\rangle_{inv} = \\
&= \sum_{j=0}^{\frac{M}{2}-1} \tilde{A}_j |j\rangle_{inv} + \sum_{j=0}^{\frac{M}{2}-1} \tilde{B}_j \left| \frac{M}{2} + j \right\rangle_{inv},
\end{aligned}$$

where

$$\begin{aligned}
\tilde{A}_j &= \sqrt{\frac{2}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} x_{2m} \\
\tilde{B}_j &= \sqrt{\frac{2}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} x_{2m+1}
\end{aligned} \tag{13.14}$$

If we now add a phase shift for the odd elements, i.e., for those with $a_0^k = 1$, we get the scheme shown in Figure 13.11:

$$\begin{aligned}
|x\rangle &\rightarrow \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |2m\rangle + \hat{R} \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} |2m+1\rangle = \\
&= \sum_{j=0}^{\frac{M}{2}-1} \tilde{A}_j |j\rangle_{inv} + \sum_{j=0}^{\frac{M}{2}-1} \tilde{B}_j \hat{R} \left| \frac{M}{2} + j \right\rangle_{inv}, \\
&= \sum_{j=0}^{\frac{M}{2}-1} \tilde{A}_j |j\rangle_{inv} + \sum_{j=0}^{\frac{M}{2}-1} \tilde{C}_j \left| \frac{M}{2} + j \right\rangle_{inv}.
\end{aligned} \tag{13.15}$$



Figure 13.11: Diagram of quantum Fourier transform based on the fast Fourier transform algorithm. Step 2: $|x\rangle \rightarrow \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |2m\rangle + \hat{R} \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1}$

Using the expression

$$\hat{R}_l |b_l^{(j)}\rangle = \exp\left(-2\pi i \frac{b_l^{(j)}}{2^{n-l}}\right) |b_l^{(j)}\rangle$$

we find that the operator \hat{R} acts on the state $|\frac{M}{2} + j\rangle_{inv}$ as follows:

$$\begin{aligned} \hat{R} \left| \frac{M}{2} + j \right\rangle_{inv} &= \hat{R} |1\rangle \otimes |j\rangle_{inv} = \\ &= |1\rangle \otimes \hat{R}_0 |b_0^{(j)}\rangle \otimes \cdots \otimes \hat{R}_{n-2} |b_{n-2}^{(j)}\rangle = \\ &= \prod_{l=0}^{n-2} \exp\left(-2\pi i \frac{2^l b_l^{(j)}}{2^n}\right) |1\rangle \otimes |j\rangle_{inv} = \\ &= \exp\left(-2\pi i \frac{j}{M}\right) \left| \frac{M}{2} + j \right\rangle_{inv} \end{aligned} \quad (13.16)$$

Derivation (13.16) took into account that $j = b_0^{(j)} + 2^1 b_1^{(j)} + \cdots + 2^{n-2} b_{n-2}^{(j)}$.

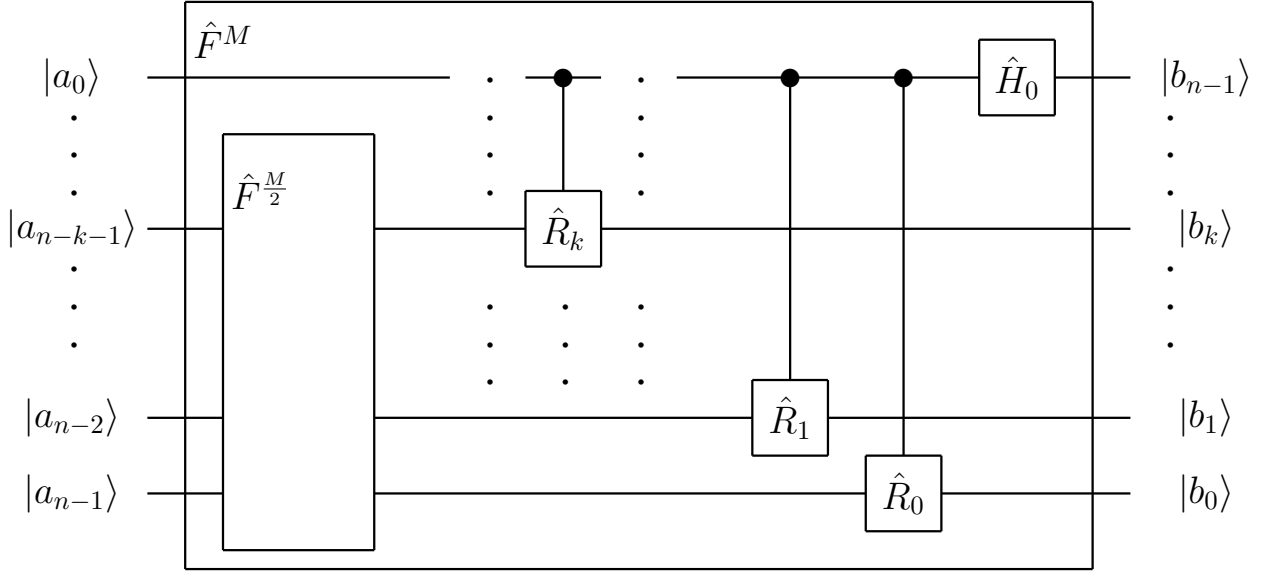


Figure 13.12: Diagram of quantum Fourier transform based on the fast Fourier transform algorithm

Thus, for \tilde{C}_j in (13.15) we have

$$\begin{aligned}
 \tilde{C}_j &= \sqrt{\frac{2}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-2\pi i \frac{j}{M}} e^{-i \frac{4\pi}{M} m j} x_{2m+1} = \\
 &= \sqrt{\frac{2}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i \frac{2\pi}{M} (2m+1) j} x_{2m+1}
 \end{aligned} \tag{13.17}$$

If we now apply the Hadamard transform to the qubit $|a_0\rangle$, we get the scheme shown in Figure 13.12. The initial state is transformed according to the following

law:

$$\begin{aligned}
|x\rangle &\rightarrow \hat{H}_0 \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m} |2m\rangle + \hat{H}_0 \hat{R} \hat{F}^{\frac{M}{2}} \sum_{m=0}^{\frac{M}{2}-1} x_{2m+1} = \\
&= \sum_{j=0}^{\frac{M}{2}-1} \tilde{A}_j \hat{H} |0\rangle \otimes |j\rangle_{inv} + \sum_{j=0}^{\frac{M}{2}-1} \tilde{C}_j \hat{H} |1\rangle \otimes |j\rangle_{inv} = \\
&= \frac{1}{\sqrt{2}} \sum_{j=0}^{\frac{M}{2}-1} \tilde{A}_j (|0\rangle + |1\rangle) \otimes |j\rangle_{inv} + \frac{1}{\sqrt{2}} \sum_{j=0}^{\frac{M}{2}-1} \tilde{C}_j (|0\rangle - |1\rangle) \otimes |j\rangle_{inv} = \\
&= \sum_{j=0}^{\frac{M}{2}-1} \frac{\tilde{A}_j + \tilde{C}_j}{\sqrt{2}} |0\rangle \otimes |j\rangle_{inv} + \sum_{j=0}^{\frac{M}{2}-1} \frac{\tilde{A}_j - \tilde{C}_j}{\sqrt{2}} |1\rangle \otimes |j\rangle_{inv} = \\
&= \sum_{j=0}^{\frac{M}{2}-1} \frac{\tilde{A}_j + \tilde{C}_j}{\sqrt{2}} |j\rangle_{inv} + \sum_{j=0}^{\frac{M}{2}-1} \frac{\tilde{A}_j - \tilde{C}_j}{\sqrt{2}} \left| \frac{M}{2} + j \right\rangle_{inv}. \quad (13.18)
\end{aligned}$$

For the members (13.18), taking into account (13.14) and (13.17), we have:

$$\begin{aligned}
\frac{\tilde{A}_j + \tilde{C}_j}{\sqrt{2}} &= \sqrt{\frac{1}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} x_{2m} + \sqrt{\frac{1}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{2\pi}{M}(2m+1)j} x_{2m+1} = \\
&= \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}mj} x_m \quad (13.19)
\end{aligned}$$

and

$$\begin{aligned}
\frac{\tilde{A}_j - \tilde{C}_j}{\sqrt{2}} &= \sqrt{\frac{1}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{4\pi}{M}mj} x_{2m} - \sqrt{\frac{1}{M}} \sum_{m=0}^{\frac{M}{2}-1} e^{-i\frac{2\pi}{M}(2m+1)j} x_{2m+1} = \\
&= \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}mj} x_m \frac{1 + e^{-i\pi m}}{2} - \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}mj} x_m \frac{1 - e^{-i\pi m}}{2} = \\
&= \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}mj} e^{-i\pi m} x_m = \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}mj} e^{-i\frac{2\pi}{M}m\frac{M}{2}} x_m = \\
&= \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}m(\frac{M}{2}+j)} x_m \quad (13.20)
\end{aligned}$$

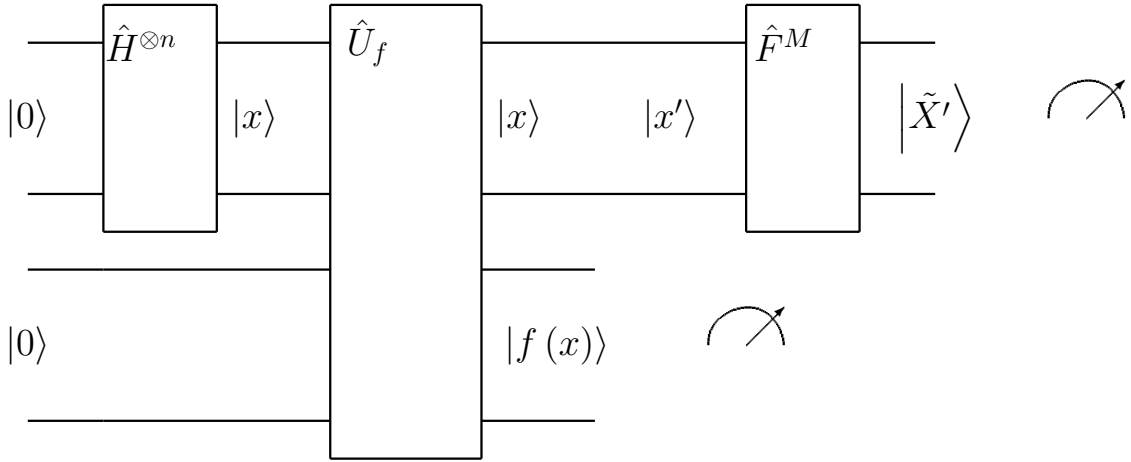


Figure 13.13: Determination of the period of functions using the quantum Fourier transform

Combining (13.18), (13.19), and (13.20), we finally obtain

$$\begin{aligned}
 |x\rangle &\rightarrow \sum_{j=0}^{\frac{M}{2}-1} \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}mj} x_m |j\rangle_{inv} + \\
 &+ \sum_{j=0}^{\frac{M}{2}-1} \sqrt{\frac{1}{M}} \sum_{m=0}^{M-1} e^{-i\frac{2\pi}{M}m(\frac{M}{2}+j)} x_m \left| \frac{M}{2} + j \right\rangle_{inv} = \\
 &= \sum_{j=0}^{M-1} \tilde{X}_j^M |j\rangle_{inv}
 \end{aligned}$$

13.4.2 Finding the Period of Functions Using Quantum Fourier Transform

To determine the period of the function (13.5), the scheme shown in Figure 13.13 is used.

The first element is the Hadamard transform for n qubits, which prepares the initial state as follows:

$$|in\rangle = \frac{1}{\sqrt{2^n}} \sum_{x=0}^{2^n-1} |x\rangle \otimes |0\rangle.$$

After the element computing the function \hat{U}_f , we have for the state

$$\hat{U}_f |in\rangle = \frac{1}{\sqrt{2^n}} \sum_{x=0}^{2^n-1} |x\rangle \otimes |f(x)\rangle.$$

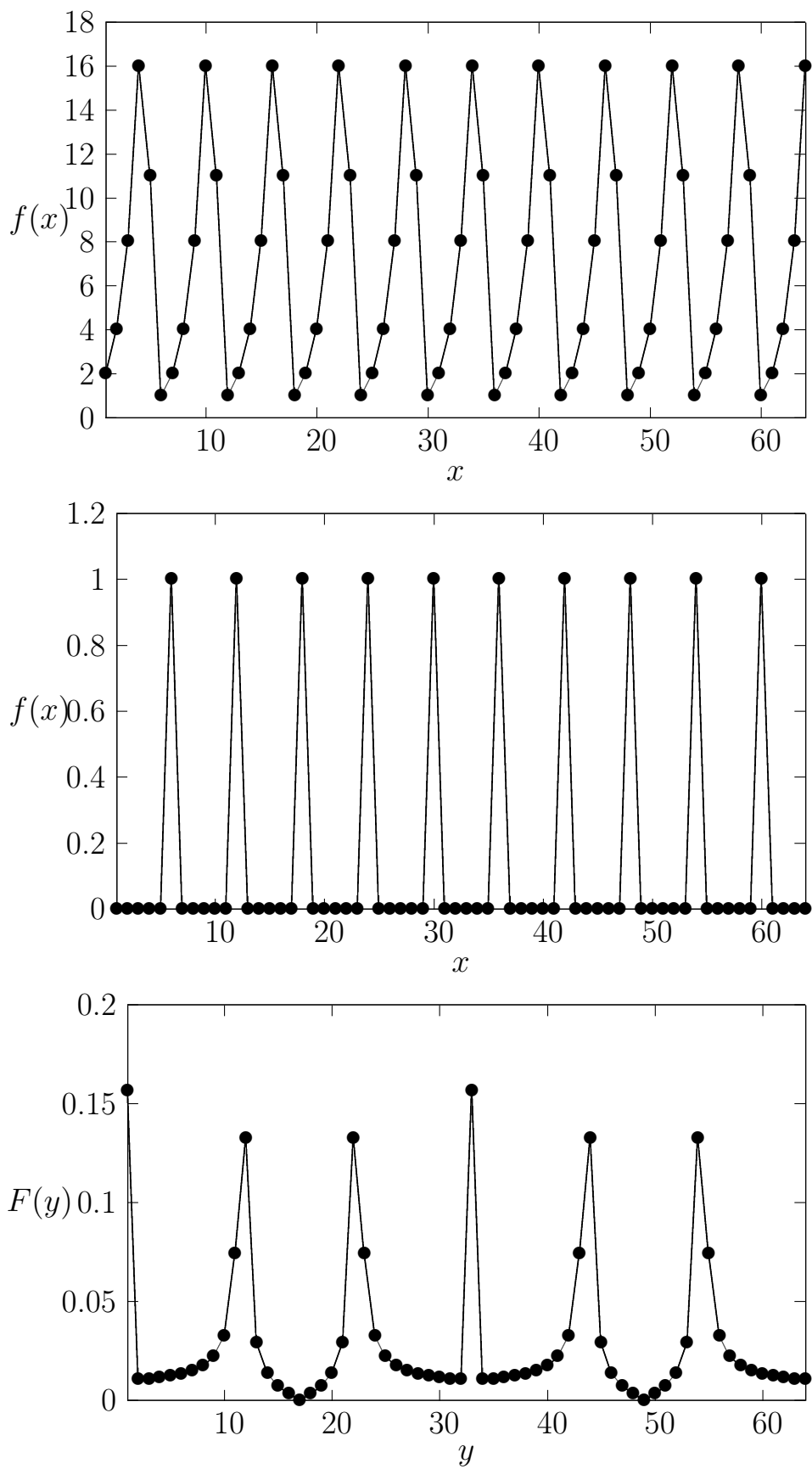


Figure 13.14: Shor's Algorithm. Finding the period of the function $f(x, a) = a^x \bmod N$ with $a = 2$, $N = 21$ (top graph). The value of the function 1 repeats with period $r = 6$ (middle graph). The local maxima of the Fourier transform of the middle graph occur with period $\frac{M}{r} \approx 10.67$ (bottom graph). See note [13.4.1](#)

After measuring the function value, only those elements for which the function value equals the measured value will remain in the list of coordinates. As a result, the input to the Fourier transformation measurement element is a state of the form

$$|in'\rangle = \sum_{x'} |x'\rangle,$$

where all non-zero elements have the same amplitude and follow with a period equal to the period of the investigated function. The initial value will be shifted depending on the experiment (the shift will vary in different experiments). By lemma 18.2.1, the Fourier image will be the same for different function measurements.

Further, by lemma 18.2.2, (on periodicity) (see also comment 18.2.1), it follows that the most probable readings (probability maxima) follow with a period related to the original period of the function. Thus, as a result of several experiments, the desired function period can be found with the required probability level (see Figure 13.14).

Example 13.4.1. Finding the period of the function $f(x) = 2^x \bmod 21$ As an example, consider the task of finding the period of the function $f(x, a) = a^x \bmod N$ with $a = 2$, $N = 21$; see Figure 13.14.

The number of samples M must be a power of two. In our example, we choose $M = 2^6 = 64$ as the number of samples. Thus, 6 qubits are needed for our example.

The initial state after the Hadamard transform is as follows:

$$|in\rangle = \frac{1}{8} \sum_{x=0}^{63} |x\rangle \otimes |0\rangle,$$

where $|x\rangle$ represents the tensor product of 6 qubits that encode the binary representation of the function argument. For example, for $x = 5_{10} = 000101_2$ we have

$$|x\rangle = |0\rangle \otimes |0\rangle \otimes |0\rangle \otimes |1\rangle \otimes |0\rangle \otimes |1\rangle$$

After computing the function, we have a state of the form (see the top plot in Figure 13.14)

$$\begin{aligned} \hat{U}_f |in\rangle &= \frac{1}{8} \sum_{x=0}^{63} |x\rangle \otimes |f(x)\rangle = \\ &= \frac{1}{8} (|0\rangle \otimes |2\rangle + |1\rangle \otimes |4\rangle + |2\rangle \otimes |8\rangle + \dots + \\ &\quad + |62\rangle \otimes |8\rangle + |63\rangle \otimes |16\rangle). \end{aligned} \tag{13.21}$$

If the result of the function measurement was 1, then from the entire sum (13.21), the terms for which the function value equals 1 will remain (see the middle plot in Figure 13.14):

$$|in'\rangle = \frac{1}{\sqrt{10}} (|5\rangle \otimes |1\rangle + |11\rangle \otimes |1\rangle + |17\rangle \otimes |1\rangle + \cdots + |60\rangle \otimes |1\rangle). \quad (13.22)$$

Expression (13.22) contains 10 terms of equal amplitude, so the normalization factor is $\frac{1}{\sqrt{10}}$.

The Fourier transform for the sequence (13.22) is shown in the bottom plot Figure 13.14. The most probable values of the Fourier transform measurement result will correspond to local maxima that repeat with a period of $\frac{M}{r} \approx 10.67$ from which the period of the desired function $r = 6$ can be found.

13.5 Quantum Fourier Transform and Discrete Logarithms

The discrete logarithm (see section 19.7) is the basis for many modern cryptographic algorithms (see section 19.9, section 19.8). Meanwhile, Shor's method for factoring integers can also be applied to compute discrete logarithms, making it possible to break the corresponding cryptographic algorithms.

Let's set the task as follows: there is an expression

$$b = a^x \mod p,$$

where the numbers a , b , and p are given, and the number x is unknown, which needs to be determined. Analogous to the application of the quantum Fourier transform for factoring numbers (see subsection 13.3.1), we must construct some periodic function whose period will allow us to determine the unknown number x . We choose the function to be analyzed as

$$f(x_1, x_2) = b^{x_1} a^{x_2} = a^{x \cdot x_1} a^{x_2} \mod p \quad (13.23)$$

As an example, we will consider the quantum analog of solving the problem from example 19.7.1:

Example 13.5.1. ($ind_3 14 \mod 17$) In our example $p = 17$, $b = 14$, and $a = 3$. The function (13.23) is

$$f(x_1, x_2) = b^{x_1} a^{x_2} = 14^{x_1} 3^{x_2}.$$

and is depicted in Figure 13.15.

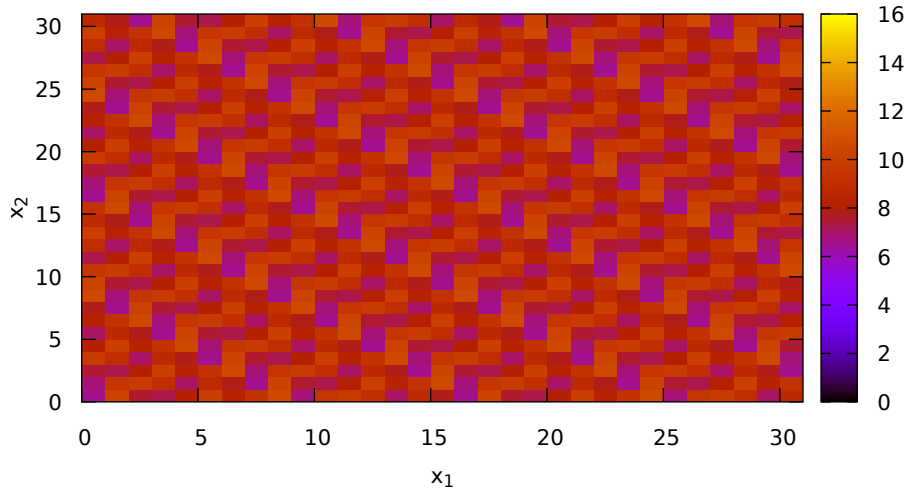


Figure 13.15: Investigated function $f(x_1, x_2) = 14^{x_1}3^{x_2}$

Both $b = 14$ and $a = 3$ are generators of $(\mathbb{Z}/17\mathbb{Z})^\times$. Moreover, $3 \equiv 14^9 \pmod{17}$. The periods of the function shown, as seen from Figure 13.15, will be the following numbers

$$\begin{aligned} t_1 &\equiv 1 \pmod{16}, \\ t_2 &\equiv 9 \pmod{16} \end{aligned} \tag{13.24}$$

In analogy with the factorization problem, you measure this function. Suppose the measurement yields a number $c \in (\mathbb{Z}/p\mathbb{Z})^\times$. Given that a is a generating element (see definition 16.1.4) of the multiplicative group $(\mathbb{Z}/p\mathbb{Z})^\times$ (see definition 16.1.5) $\exists x_0 : c = a^{x_0}$. Then, considering Fermat's Little Theorem (Theorem 19.4.1) $a^{p-1} \equiv 1 \pmod{p}$, it follows that

$$x_0 \equiv xx_1 + x_2 \pmod{q},$$

where $q = p - 1$. From this expression, it follows that

$$x_2 \equiv x_0 - xx_1 \pmod{q}.$$

This means if the function is periodic in the first argument:

$$f(x_1 + t_1, x_2) = f(x_1, x_2),$$

then it will also be periodic in the second argument

$$f(x_1, x_2 + t_2) = f(x_1, x_2),$$

where

$$t_2 \equiv xt_1 \pmod{q}. \tag{13.25}$$

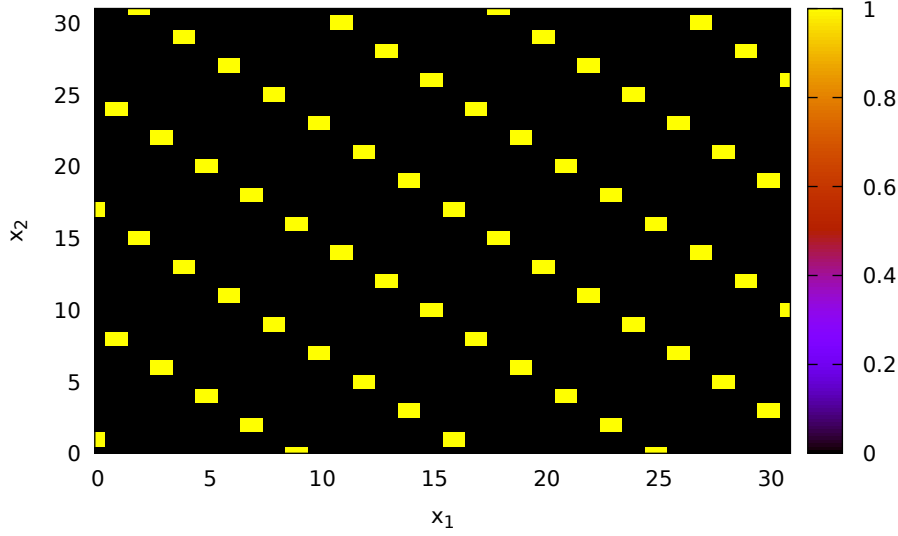


Figure 13.16: The function under study $f(x_1, x_2) = 14^{x_1}3^{x_2}$. Only those pairs x_1, x_2 for which $f(x_1, x_2) = 3, x_0 = 1$ are marked.

13.5.1 Two-Dimensional Fourier Transform and Period of the Function $f(x_1, x_2)$

Our function of interest will be:

$$f'(x_1, x_2) = \begin{cases} 1, & xx_1 + x_2 \equiv x_0 \pmod{q} \\ 0, & xx_1 + x_2 \not\equiv x_0 \pmod{q} \end{cases} \quad (13.26)$$

Example 13.5.2. ($\text{ind}_3 14 \pmod{17}$) Continuing example 13.5.1, suppose the measurement of function f yields $f = 3$. That is, $f = a^{x_0} = 3^{x_0} \equiv 3 \pmod{17}$. The only values for x_1, x_2 that remain are those corresponding to the observed function value (see Figure 13.16), i.e., $xx_1 + x_2 = x_0 \equiv 1 \pmod{16}$. With fixed values of x, x_1 , and the number of counts $M = q = 16$.

For the Fourier transform \tilde{f}' , we have

$$\tilde{f}'(j_1, j_2) = \frac{1}{M} \sum_{x_1=0}^{M-1} \sum_{x_2=0}^{M-1} f'(x_1, x_2) e^{-i\omega(x_1 j_1 + x_2 j_2)}, \quad (13.27)$$

where $\omega = \frac{2\pi}{M}$, M is the number of samples.

First, consider the case when $M = q$. There are two options for x_2 :

1. $x_2 = x_0 - xx_1$, if $x_0 \geq xx_1$
2. $x_2 = x_0 + q - xx_1$, if $x_0 < xx_1$

Thus,

$$\begin{aligned} e^{-i\omega x_2 j_2} &= e^{-i\omega(x_0 - xx_1 + q)j_2} = \\ &= e^{-i\omega(x_0 - xx_1)j_2 - i\omega q j_2} = e^{-i\omega(x_0 - xx_1)j_2}, \end{aligned}$$

i.e., both options coincide and can be reduced to the first: $x_2 = x_0 - xx_1$.

Continuing from (13.27), we have

$$\begin{aligned} \tilde{f}'(j_1, j_2) &= \frac{1}{M} \sum_{x_1=0}^{M-1} \sum_{x_2=0}^{M-1} f'(x_1, x_2) e^{-i\omega(x_1 j_1 + x_2 j_2)} = \\ &= \frac{1}{M} \sum_{x_1=0}^{M-1} e^{-i\omega(x_1 j_1 + (x_0 - xx_1)j_2)} = \\ &= e^{-i\omega x_0 j_2} \frac{1}{M} \sum_{x_1=0}^{M-1} e^{-i\omega x_1(j_1 - x j_2)} = \frac{1}{M} e^{-i\omega x_0 j_2} \sum_{x_1=0}^{M-1} e^{-i\omega x_1(j_1 - x j_2)}. \end{aligned} \quad (13.28)$$

In expression (13.28), $\tilde{f}'(j_1, j_2) = e^{-i\omega x_0 j_2} \neq 0$, if

$$j_1 \equiv x j_2 \pmod{M}. \quad (13.29)$$

Otherwise, using the geometric progression formula:

$$\begin{aligned} \tilde{f}'(j_1 \neq x j_2, j_2) &= e^{-i\omega x_0 j_2} \frac{1}{M} \sum_{x_1=0}^{M-1} e^{-i\omega x_1(j_1 - x j_2)} = \\ &= \frac{e^{-i\omega x_0 j_2}}{M} \frac{e^{-i\omega M(j_1 - x j_2)} - 1}{e^{-i\omega(j_1 - x j_2)} - 1} = \\ &= \frac{e^{-i\omega x_0 j_2}}{M} \frac{e^{-i\frac{2\pi}{M} M(j_1 - x j_2)} - 1}{e^{-i\omega(j_1 - x j_2)} - 1} = 0. \end{aligned}$$

Thus, to determine the period, we need to find the coordinates (j_1, j_2) of a maximum of the Fourier transform and use the expression

$$x \equiv j_1 j_2^{-1} \pmod{M}, \quad (13.30)$$

which follows from (13.29).

Remark 13.5.1 (On Zero Divisors in $\mathbb{Z}/M\mathbb{Z}$). *If there exists such a number y that*

$$j_2 y \equiv 0 \pmod{M},$$

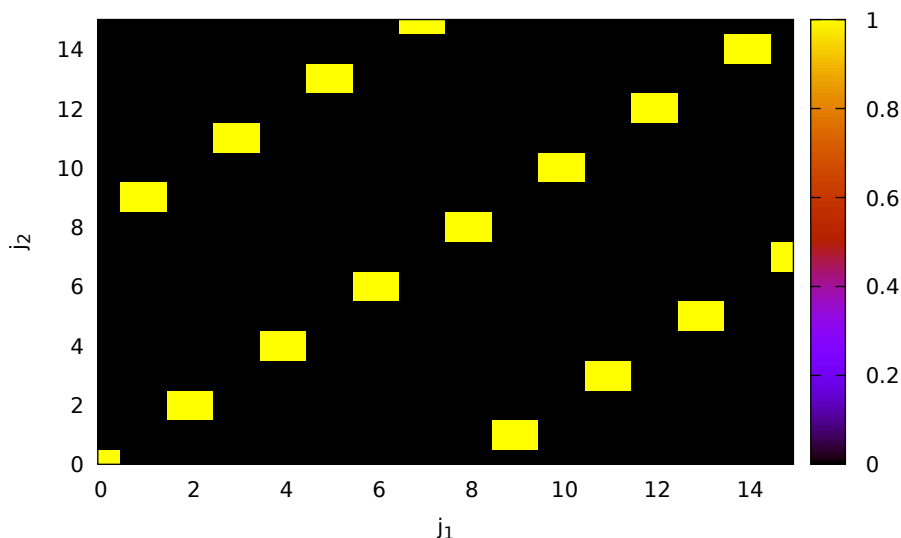


Figure 13.17: Fourier image of the function from Figure 13.16. The number of samples $M = 16$. Coordinates of the maximum $j_1 = 9$, $j_2 = 1$. The solution of the equation $3^x \equiv 14 \pmod{17}$ is $x \equiv 9 \cdot 1^{-1} \equiv 9 \pmod{16}$. Considering that the period is less than the number of samples, we can conclude that $x = 9$

then j_2 is called a zero divisor. It is obvious that

$$\gcd(j_2, M) \neq 1,$$

thus, from subsection 19.2.2, it follows that j_2^{-1} does not exist. Therefore, for such j_2 , expression (13.30) is undefined. In that case, different coordinates (j_1, j_2) must be used.

Example 13.5.3. ($\text{ind}_3 14 \pmod{17}$)

The Fourier transform of the function from Figure 13.16 is shown in Figure 13.17, from which it is clear that the counts most likely to be registered follow at intervals of $T_{j_1} = 9$ along the j_1 coordinate and at intervals of $T_{j_2} = 1$ along the j_2 coordinate. Considering the number of counts $M = 16$, the coordinates of the Fourier transform maximum can be obtained as $j_1 = 9$ and $j_2 = 1$. The solution to the equation $3^x \equiv 14 \pmod{17}$ is, according to (13.30), $x = 9 \cdot 1^{-1} = 9$, which corresponds to the result of example 19.7.1.

A similar result can be obtained if the coordinates $j_1 = 11, j_2 = 3$ are considered. Given $3 \cdot 11 = 33 \equiv 1 \pmod{16}$, it follows that $j_2^{-1} \equiv 11 \pmod{16}$, i.e., $x \equiv 11 \cdot 11 \equiv 121 \equiv 9 \pmod{16}$, which again corresponds to the result of example 19.7.1.

It is worth noting that points lying on the diagonal, for example, $j_1 = 6, j_2 = 6$, will not yield a correct result because $\gcd(6, 16) = 2 \neq 1$.

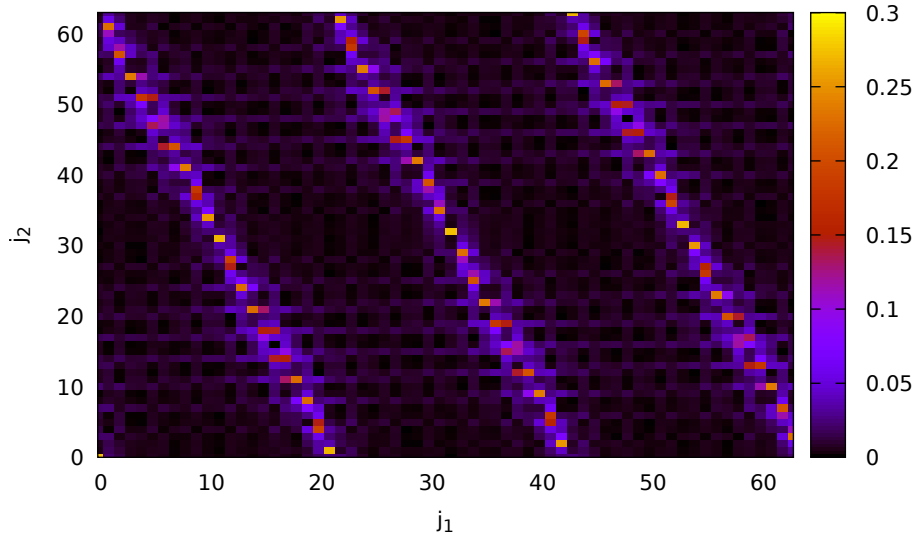


Figure 13.18: Fourier transform of the samples of the function $f'(x_1, x_2)$. Number of samples $M = 64$. The three lower maxima have coordinates $\approx (20, 1), (41, 2.2), (62, 3)$, giving the following estimates for x : $x \approx 20, 18.6, 20.6$, which is close to the real value $x = 19$.

It is worth noting that the obtained result (13.30) is in direct correspondence with lemma 18.2.2 for the one-dimensional Fourier transform. Furthermore, an analogue of comment 18.2.1 holds, stating that when the number of Fourier transform counts is not equal to q : $M \neq q$, but $M \approx q$, we can still approximately consider expression (13.30) to be valid [21].

Example 13.5.4. ($\text{ind}_2 14 \bmod 59$)

As an example, consider $p = 59$ where the number of counts $M = 64 \approx q = p - 1 = 58$. The generator of the group \mathbb{F}_{59} (see subsection 19.2.3) is $g = 2$, i.e., $\mathbb{F}_{59} = \langle 2 \rangle$. This means that the equation $2^x \equiv b \bmod 59$ will have a solution for any b , in particular, $x = 19$ is a solution to the equation

$$2^x \equiv 14 \bmod 59.$$

The function under study is

$$f(x_1, x_2) = 14^{x_1} 2^{x_2} \bmod 59,$$

Suppose $x_0 = 50$, i.e., the registered value of the function is $f(x_1, x_2) = 2^{x_0} = 2^{50} \equiv 3 \bmod 59$.

As mentioned earlier, the number of Fourier transform counts is $M = 64$. Note that $q = p - 1 = 58 \not\equiv 0 \bmod 64$.

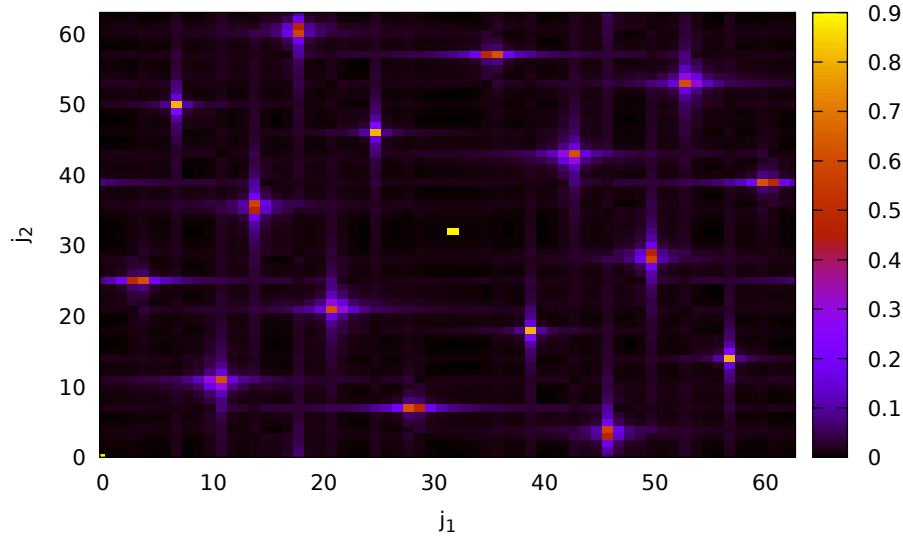


Figure 13.19: Fourier image of the samples of the function $f'(x_1, x_2)$ Number of samples $M = 64$. Coordinates of the maximum $j_1 \approx 46$, $j_2 \approx 3.5$. The solution to the equation $3^x \equiv 14 \pmod{19}$ is $x = 13 \approx \frac{46}{3.5} \approx 13.14$

The Fourier transform of the function

$$f'(x_1, x_2) = \begin{cases} 1, & \text{if } 14^{x_1} 2^{x_2} \equiv 3 \pmod{59} \\ 0, & \text{if } 14^{x_1} 2^{x_2} \not\equiv 3 \pmod{59} \end{cases}$$

is shown in [Figure 13.18](#). The three lower maxima have coordinates

$$(j_1, j_2) \approx (20, 1), (41, 2.2), (62, 3),$$

which give the following estimates for x : $x \approx 20, 18.6, 20.6$, close to the actual value $x = 19$.

Example 13.5.5. ($\text{ind}_3 14 \pmod{19}$)

As an example, consider the task of determining x such that

$$3^x \equiv 14 \pmod{19}.$$

The function under study is

$$f(x_1, x_2) = 14^{x_1} 3^{x_2} \pmod{19},$$

Suppose $x_0 = 1$, i.e., the registered value of the function is $f(x_1, x_2) = 3$.

The number of Fourier transform counts is chosen as $M = 64$. Note that $18 \not\equiv 0 \pmod{64}$.

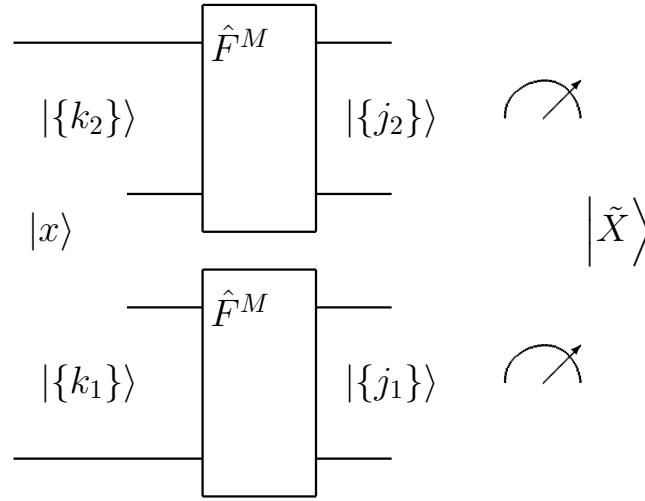


Figure 13.20: Scheme of two-dimensional quantum Fourier transform. The input signal is $|x\rangle = \sum_{k_1, k_2=0}^{M-1} x(k_1, k_2) |k\rangle_1 \otimes |k\rangle_2$. The output signal $|\tilde{X}\rangle = \sum_{j_1, j_2=0}^{M-1} \tilde{X}(j_1, j_2) |j_1\rangle \otimes |j_2\rangle$ is the two-dimensional Fourier transform of the original $|x\rangle$.

The Fourier transform of the function

$$f'(x_1, x_2) = \begin{cases} 1, & \text{if } 14^{x_1} 3^{x_2} \equiv 3 \pmod{19} \\ 0, & \text{if } 14^{x_1} 3^{x_2} \not\equiv 3 \pmod{19} \end{cases}$$

is shown in [Figure 13.19](#). The lowest maximum has coordinates $j_1 = 46, j_2 = 3.5$, from which we get an estimate

$$x \approx \frac{46}{3.5} \approx 13.14.$$

It is noteworthy that the solution to the desired equation $x = 13$ is close to the approximate solution found.

13.5.2 Two-Dimensional Quantum Fourier Transform

To determine the periods of a function with two arguments, we can use a two-dimensional Fourier transform, which can be constructed from blocks implementing the one-dimensional Fourier transform, as depicted in [Figure 13.20](#). For analyzing this circuit, consider the trivial case where the input is (see also [\(13.10\)](#))

$$\begin{aligned} |x\rangle &= |x\rangle_1 \otimes |x\rangle_2, \\ |x\rangle_{1,2} &= \sum_{k_{1,2}=0}^{M-1} x_{k_{1,2}}^{(1,2)} |k_{1,2}\rangle. \end{aligned}$$

Given that the output is

$$|\tilde{X}\rangle = |\tilde{X}_1\rangle \otimes |\tilde{X}_2\rangle,$$

where

$$|\tilde{X}_{1,2}\rangle = \sum_{j_{1,2}=0}^{M-1} \tilde{X}_{j_{1,2}}^{(1,2)} |j_{1,2}\rangle$$

and in accordance with (13.12)

$$\tilde{X}_{j_{1,2}}^{(1,2)} = \frac{1}{\sqrt{M}} \sum_{k_{1,2}=0}^{M-1} e^{-i\omega_{1,2}k_{1,2}j_{1,2}} x_{k_{1,2}}^{(1,2)}.$$

we obtain

$$\begin{aligned} |\tilde{X}\rangle &= |\tilde{X}_1\rangle \otimes |\tilde{X}_2\rangle = \\ &= \sum_{j_1=0}^{M-1} \sum_{j_2=0}^{M-1} \tilde{X}_{j_1}^{(1)} \tilde{X}_{j_2}^{(2)} |j_1\rangle \otimes |j_2\rangle = \\ &= \sum_{j_1=0}^{M-1} \sum_{j_2=0}^{M-1} \tilde{X}_{j_1,j_2} |j_1\rangle \otimes |j_2\rangle, \end{aligned}$$

where

$$\begin{aligned} \tilde{X}_{j_1,j_2} &= \frac{1}{(\sqrt{M})^2} \sum_{k_1=0}^{M-1} \sum_{k_2=0}^{M-1} e^{-i\omega(k_1j_1+k_2j_2)} x_{k_1}^{(1)} x_{k_2}^{(2)} = \\ &= \frac{1}{(\sqrt{M})^2} \sum_{k_1=0}^{M-1} \sum_{k_2=0}^{M-1} e^{-i\omega(k_1j_1+k_2j_2)} x_{k_1,k_2} \end{aligned}$$

which, in accordance with definition 18.3.1, is a two-dimensional Fourier transform of the initial two-dimensional signal

$$|x\rangle = \sum_{k_1=0}^{M-1} \sum_{k_2=0}^{M-1} x_{k_1}^{(1)} x_{k_2}^{(2)} |k_1\rangle \otimes |k_2\rangle = \sum_{k_1=0}^{M-1} \sum_{k_2=0}^{M-1} x_{k_1,k_2} |k_1\rangle \otimes |k_2\rangle.$$

Therefore, using the scheme shown in Figure 13.21, the coordinates of the maxima of the two-dimensional Fourier transform j_1, j_2 can be determined and subsequently used in (13.30) to find the desired x .

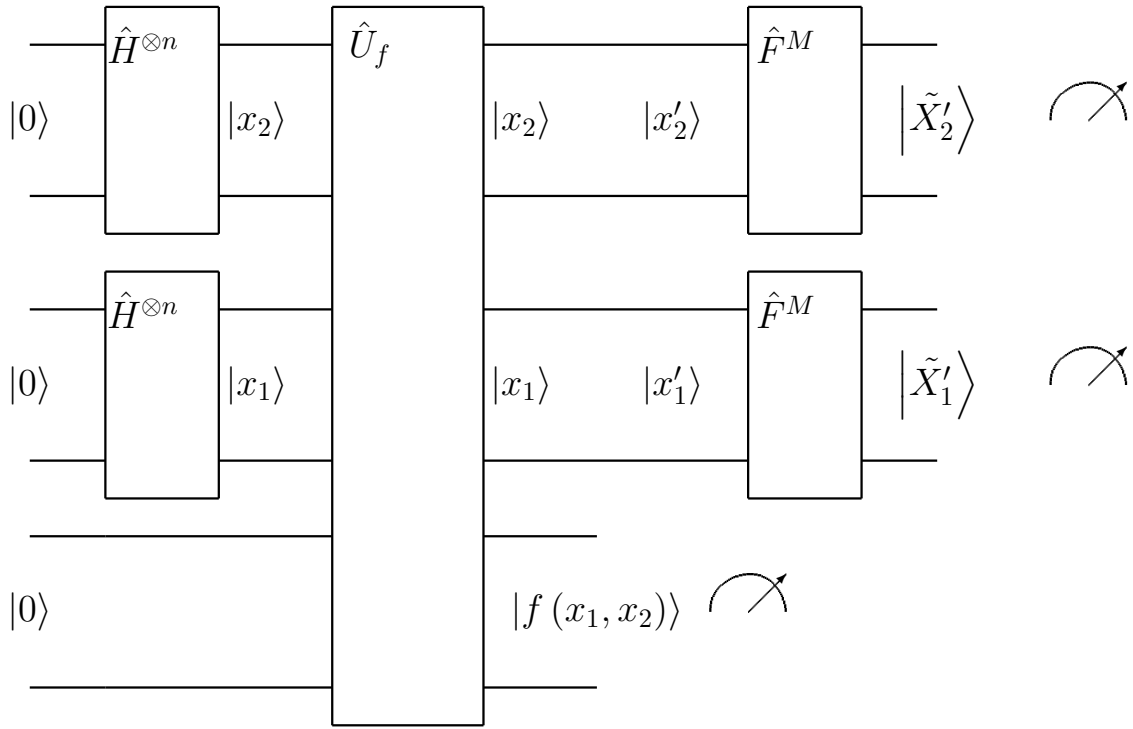


Figure 13.21: Determining the period of functions of two arguments using the quantum Fourier transform

13.6 Shor's Algorithm and the Discrete Logarithm on Elliptic Curves

Consider the elliptic curve

$$E(\mathbb{F}) = \{(x, y) \in \mathbb{F}_p \times \mathbb{F}_p, y^2 = x^3 + ax + b, \}$$

with a given base point $g \in E(\mathbb{F}_p)$ such that:

$$ng = 0.$$

The task is to solve the following problem: given $q \in E(\mathbb{F}_p)$, find x such that

$$xg = q \pmod{n} \quad (13.31)$$

Consider the following auxiliary function

$$f(x_1, x_2) = x_1q + x_2g = (xx_1 + x_2)g, \quad (13.32)$$

where $q, g \in E(\mathbb{F}_p)$ are taken from the conditions of our problem (13.31). This function is analogous to (13.23) used in solving the discrete logarithm problem. Next, a measurement of this function is taken. The result of this measurement is

some point $c \in E(\mathbb{F}_p)$. Furthermore, from (13.32) it follows that $c \in \langle g \rangle$, i.e. $\exists x_0$ such that $c = x_0 g$.

Thus, by analogy with (13.26), we compose the following function

$$f'(x_1, x_2) = \begin{cases} 1, & xx_1 + x_2 \equiv x_0 \pmod{n} \\ 0, & xx_1 + x_2 \not\equiv x_0 \pmod{n} \end{cases} \quad (13.33)$$

The coordinates (j_1, j_2) of the maximum of the Fourier transform \tilde{f}' give, according to the formula (13.30), some value of the desired number x . In our case, almost always $n \neq M$, so we can only use an approximate estimate

$$x \approx \frac{j_1}{j_2}.$$

Example 13.6.1 (Discrete Logarithm on an Elliptic Curve). *Consider the problem from example 19.11.1 The curve and base point (see example 19.10.2) are given by*

$$(p, a, b, g, n, h) = (97, -7, 10, (96, 93), 41, 2)$$

Assume we know Alice's public key

$$A = (37, 35)$$

and we want to find $x \in \{0, 1, \dots, 40\}$ such that $xg = A$, as follows from example 19.11.1 the answer is $x = d_a = 5$. Our studied function will be

$$f(x_1, x_2) = x_1 A + x_2 g = x_1(37, 35) + x_2(96, 93)$$

As a result of measurement, we choose $c = g$, i.e., $x_0 = 1$. The graph of the function $f'(x_1, x_2)$ corresponding to this measurement is shown in Figure 13.22. It is worth noting that the function $f'(x_1, x_2)$ is periodic and if you take any two closely spaced points, for example, $(8, 2)$ and $(16, 3)$, you can notice that the period in the x_1 coordinate is $T_1 = 8$, and in the x_2 coordinate - $T_2 = 1$. Solving the equation $x = T_2 T_1^{-1} \pmod{n}$ we get $x = 8^{-1} \equiv 5 \pmod{41}$, which corresponds to the desired solution.

The Fourier transform of the function $f'(x_1, x_2)$ is shown in Figure 13.23. From this, the desired $x = 5$ can be found.

13.7 Grover's Algorithm

Consider the following problem. Suppose we have a large dataset consisting of N elements in which it is necessary to find an element satisfying certain conditions

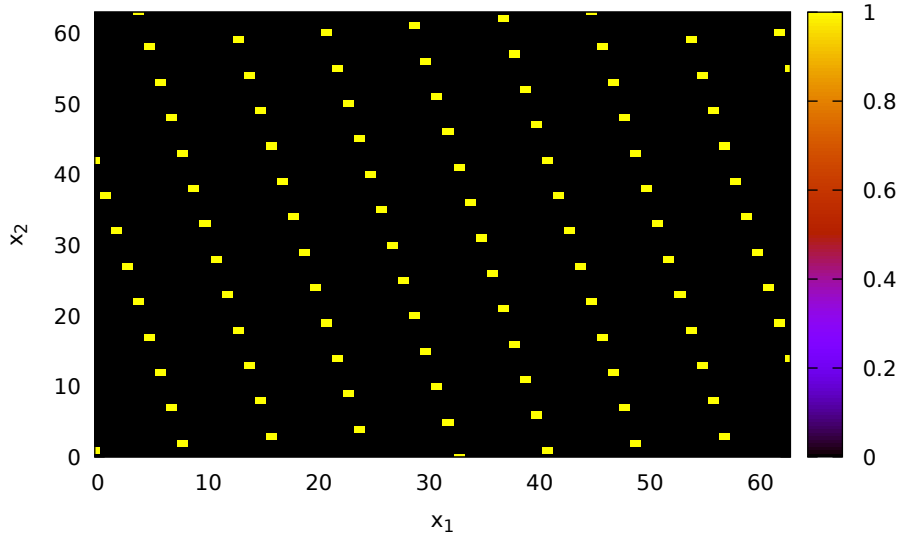


Figure 13.22: Graph of the function $f'(x_1, x_2)$ at $x_0 = 1$. Thus, points x_1, x_2 corresponding to the relation $x_1 A + x_2 g = g$ are depicted: $x_1(37, 35) + x_2(96, 93) = (96, 93)$, for example, $7(37, 35) + 7(96, 93) = (96, 93)$, $8(37, 35) + 2(96, 93) = (96, 93)$, or $16(37, 35) + 3(96, 93) = (96, 93)$. It is worth noting that the chosen pairs of points correspond to the condition (13.33), indeed we have $x \cdot 7 + 7 \equiv x \cdot 8 + 2 \pmod{41}$. That is, if we subtract one from the other, we get an equation of the form $x(8 - 7) = -(2 - 7) = 5 \pmod{41}$. Or $x \equiv 5 \pmod{41}$

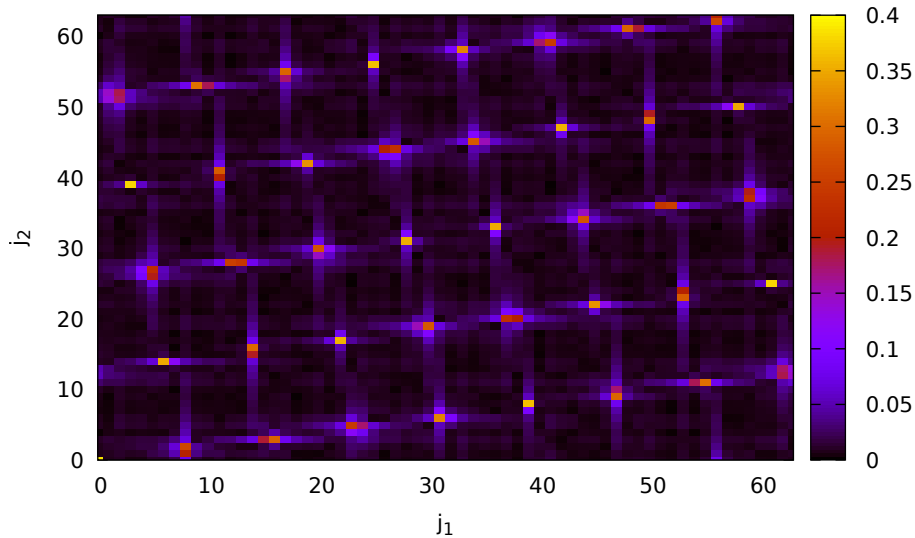


Figure 13.23: Fourier transform of the samples of function $f'(x_1, x_2)$. Number of samples $M = 64$. The three lower left maxima have coordinates $\approx (8, 2), (15, 3), (24, 5)$, which give the following estimates for x : $x \approx 4, 5, 4.8$, which is close to the actual value $x = 5$

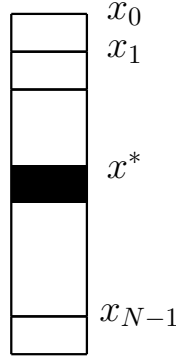


Figure 13.24: Search in an unstructured volume of data (looking for "a needle in a haystack")

(see Figure 13.24). If the data is sorted, the desired element can be found using "divide and conquer" type algorithms in time $O(\log N)$ (see section 20.2). In some cases, the original dataset cannot be prepared for fast search, and in this case classical search is performed in time $O(N)$.

One example is symmetric encryption algorithms where the task is to determine the key from the known encrypted text and its corresponding original text. In this case, preprocessing the data is impossible, and the "brute force" solution is simply to try all possible values.

Grover's algorithm [11] solves the problem of unstructured search in time $O(\sqrt{N})$.

13.7.1 Description of the Algorithm

Suppose we have a quantum circuit that computes the value of the function $f(x)$ which can take only two values: 0 and 1. The value 1 is true only for the desired element:

$$\begin{aligned} f(x)|_{x=x^*} &= 1, \\ f(x)|_{x \neq x^*} &= 0. \end{aligned} \quad (13.34)$$

In Figure 13.25 a scheme for calculating the desired function is shown. At the output, we have a state of the form

$$|out\rangle = \frac{1}{\sqrt{N}} \left(\sum_{x \neq x^*} |x\rangle \otimes |0\rangle + |x^*\rangle \otimes |1\rangle \right), \quad (13.35)$$

where N is the total number of elements in the sequence in which the search is conducted.

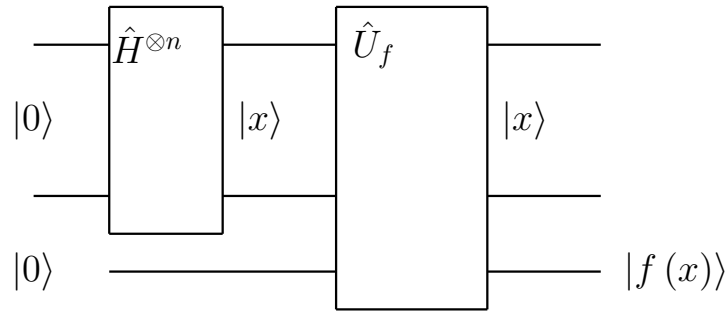


Figure 13.25: Computation of the function $f(x)$. At the output of the circuit, we have a superposition of states of the form $\frac{1}{\sqrt{N}} \left(\sum_{x \neq x^*} |x\rangle \otimes |0\rangle + |x^*\rangle \otimes |1\rangle \right)$

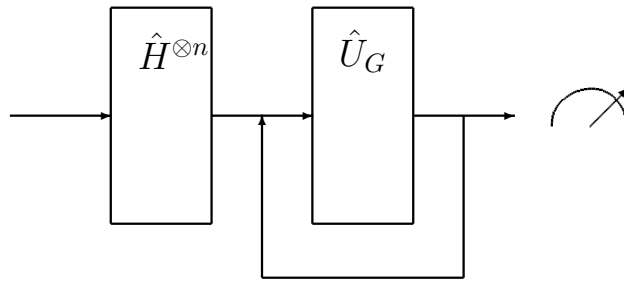


Figure 13.26: Grover's Algorithm

If you look at expression (13.35), you can note that although the scheme calculates the function at the desired point, it does not allow selecting the desired element, because all elements of the resulting sequence are equally probable, i.e., each element can be selected (as a result of measurement) with equal probability: $\frac{1}{N}$.

Grover proposed an algorithm that increases the probability of detecting the desired element in the resulting superposition (13.35).

The scheme implementing Grover's algorithm represents a block described by the operator \hat{U}_G , which is repeated a certain number of times (see Figure 13.26). With each iteration step, the probability of detecting the desired element in-

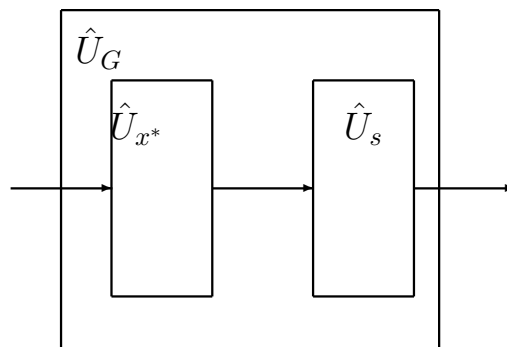


Figure 13.27: Grover's Algorithm. Basic Element



Figure 13.28: Grover's Algorithm. Phase Inversion. Described by the following equation $\hat{U}_{x^*} (\sum_x \alpha_x |x\rangle) = \sum_x \alpha_x (-1)^{f(x)} |x\rangle$

creases.

The basic element \hat{U}_G is a sequential action of two operators (see [Figure 13.27](#)):

$$\hat{U}_G = \hat{U}_s \hat{U}_{x^*},$$

where \hat{U}_{x^*} is the phase inversion operator, \hat{U}_s is the inversion about the mean operator.

The action of the operator \hat{U}_{x^*} is described by the following relation (see [Figure 13.28](#)):

$$\hat{U}_{x^*} \left(\sum_x \alpha_x |x\rangle \right) = \sum_x \alpha_x (-1)^{f(x)} |x\rangle. \quad (13.36)$$

The operator \hat{U}_{x^*} can be rewritten as

$$\hat{U}_{x^*} = \hat{I} - 2 |x^*\rangle \langle x^*|.$$

Indeed,

$$\begin{aligned} & \left(\hat{I} - 2 |x^*\rangle \langle x^*| \right) \left(\sum_x \alpha_x |x\rangle \right) = \\ &= \sum_x \alpha_x |x\rangle - 2 \alpha_{x^*} |x^*\rangle = \sum_{x \neq x^*} \alpha_x |x\rangle - \alpha_{x^*} |x^*\rangle = \\ &= \sum_x \alpha_x (-1)^{f(x)} |x\rangle, \end{aligned}$$

which coincides with (13.36).



Figure 13.29: Grover's Algorithm. Inversion about the average. Described by the following relation $\hat{U}_s (\sum_x \alpha_x |x\rangle) = \sum_x (2\mathcal{M} - \alpha_x) |x\rangle$

The action of the operator \hat{U}_s is described by the following relation (see [Figure 13.29](#)):

$$\hat{U}_s \left(\sum_x \alpha_x |x\rangle \right) = \sum_x (2\mathcal{M} - \alpha_x) |x\rangle, \quad (13.37)$$

where $\mathcal{M} = \sum_x \frac{\alpha_x}{N}$.

The operator \hat{U}_s can be rewritten as follows

$$\hat{U}_s = 2|s\rangle\langle s| - \hat{I},$$

where $|s\rangle = \frac{1}{\sqrt{N}} \sum_x |x\rangle$ is the initial state in Grover's algorithm. Indeed,

$$\begin{aligned} & \left(2|s\rangle\langle s| - \hat{I} \right) \left(\sum_x \alpha_x |x\rangle \right) = \\ &= 2 \sum_x \alpha_x \langle s|x\rangle |s\rangle - \sum_x \alpha_x |x\rangle = \\ &= \frac{2}{N} \sum_x \alpha_x \sum_x |x\rangle - \sum_x \alpha_x |x\rangle = \\ &= \sum_x (2\mathcal{M} - \alpha_x) |x\rangle, \end{aligned}$$

which coincides with (13.37).

13.7.2 Analysis of Grover's Algorithm

The schematic form of Grover's algorithm is given in Alg. [2](#).

Algorithm 2 Grover's Algorithm

```

 $|\psi\rangle_0 \Leftarrow \frac{1}{\sqrt{N}} \sum_x |x\rangle$ 
 $t \Leftarrow 1$ 
repeat
   $|\psi\rangle_t \Leftarrow \hat{U}_s \hat{U}_{x^*} |\psi\rangle_{t-1}$ 
   $t \Leftarrow t + 1$ 
until  $(t < \frac{\pi}{4} \sqrt{N})$ 
return result of measuring the state  $|\psi\rangle_t$ 

```

We are interested in two questions: what is the algorithmic complexity of Grover's algorithm and are there any algorithms that can perform the search task in an unstructured set of data more efficiently than Grover's algorithm.

The criterion for the efficiency of an algorithm is as follows: a good algorithm should find the desired value with a minimal number of function calls (13.34).

Let's consider the very first iteration. The initial state $|\psi\rangle_0$ has the following form

$$|\psi\rangle_0 = \sum_x \alpha_x |x\rangle = |s\rangle = \frac{1}{\sqrt{N}} \sum_x |x\rangle = \frac{1}{\sqrt{N}} \sum_{x \neq x^*} |x\rangle + \frac{1}{\sqrt{N}} |x^*\rangle.$$

Thus, the coefficient before the desired element is $\alpha_{x^*}^* = \frac{1}{\sqrt{N}}$.

After applying the phase inversion operator U_{x^*} from (13.36), we get

$$\hat{U}_{x^*} |\psi\rangle_0 = \frac{1}{\sqrt{N}} \sum_{x \neq x^*} |x\rangle - \frac{1}{\sqrt{N}} |x^*\rangle = \sum_x \beta_x |x\rangle,$$

where $\beta_{x^*} = -\frac{1}{\sqrt{N}}$ and $\beta_{x \neq x^*} = \frac{1}{\sqrt{N}}$.

After applying the inversion about the mean operator \hat{U}_s from (13.37), we get

$$\begin{aligned} \hat{U}_G |\psi\rangle_0 &= \hat{U}_s \hat{U}_{x^*} |\psi\rangle_0 = \hat{U}_s \sum_x \beta_x |x\rangle = \\ &= \sum_x (2M - \beta_x) |x\rangle \approx \sum_{x \neq x^*} \left(2\frac{1}{\sqrt{N}} - \frac{1}{\sqrt{N}} \right) |x\rangle + \\ &+ \left(2\frac{1}{\sqrt{N}} + \frac{1}{\sqrt{N}} \right) |x^*\rangle = \frac{1}{\sqrt{N}} \sum_{x \neq x^*} |x\rangle + \frac{3}{\sqrt{N}} |x^*\rangle. \end{aligned} \quad (13.38)$$

In deriving (13.38), it was assumed

$$\mathcal{M} = \frac{\sum_x \alpha_x}{N} \approx \frac{N}{N\sqrt{N}} = \frac{1}{\sqrt{N}}.$$

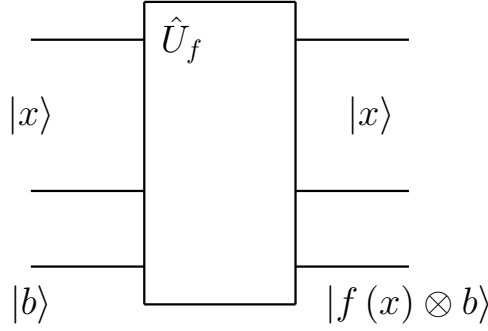


Figure 13.30: Grover's Algorithm. Phase Inversion Implementation (operator \hat{U}_{x^*}). Assuming $b = |-\rangle = \frac{|0\rangle - |1\rangle}{\sqrt{2}}$, for the depicted circuit we obtain $\hat{U}_f(\sum_x \alpha_x |x\rangle) \otimes |-\rangle = \sum_x \alpha_x (-1)^{f(x)} |x\rangle \otimes |-\rangle$

Thus, after the first iteration of Grover's algorithm, the amplitude α_{x^*} increased by $\frac{2}{\sqrt{N}}$. If this result is approximated for an arbitrary iteration, it can be determined that a 50% probability of finding $|x^*\rangle$ will be achieved after the following number of iterations:

$$\frac{1}{\sqrt{2}} / \frac{2}{\sqrt{N}} = \frac{\sqrt{N}}{2\sqrt{2}} = O(\sqrt{N}).$$

More accurate calculations [19] give for the number of iterations $\frac{\pi}{4}\sqrt{N}$.

One could ask about the optimality of Grover's algorithm: Is there a quantum algorithm that performs a search in an unstructured set of data faster than $O(\sqrt{N})$ calls to the function (13.34)? In the article [6], it is shown that such an algorithm does not exist.

13.7.3 Implementation of Basic Elements of Grover's Algorithm

Phase Inversion

How can phase inversion be implemented: what does the quantum logic element look like that performs the transformation (13.36), i.e., how can $f(x)$ be "sent" into the phase?

Consider the circuit shown in Figure 13.30. The proposed circuit performs the following transformation:

$$|x\rangle \otimes |b\rangle \rightarrow |x\rangle \otimes |b \oplus f(x)\rangle,$$

where the following notation is introduced: $a \oplus b = a + b \pmod{2}$.

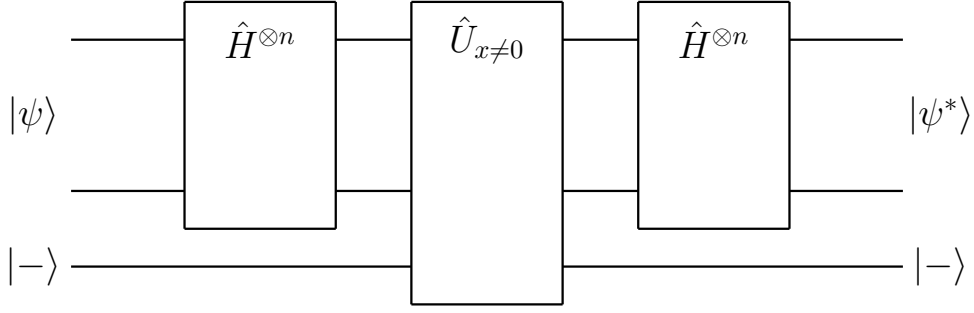


Figure 13.31: Grover's Algorithm. Implementation of inversion about the mean (operator \hat{U}_s): $|\psi\rangle \otimes |-\rangle \rightarrow |\psi^*\rangle \otimes |-\rangle$, where $|\psi\rangle = \sum_x \alpha_x |x\rangle$, $|\psi^*\rangle = \sum_x (2\mathcal{M} - \alpha_x) |x\rangle$. The proposed scheme performs the following transformation: $\hat{H}^{\otimes n} \hat{U}_{x \neq 0} \hat{H}^{\otimes n} \sum_x \alpha_x |x\rangle \otimes |-\rangle = \sum_x (2\mathcal{M} - \alpha_x) |x\rangle \otimes |-\rangle$

For the case $|b\rangle = |-\rangle = \frac{|0\rangle - |1\rangle}{\sqrt{2}}$ we have

$$\begin{aligned}
 |x\rangle \otimes |-\rangle &\rightarrow |x\rangle \otimes \left(\frac{|0 \oplus 0\rangle - |1 \otimes 0\rangle}{\sqrt{2}} \right) = \\
 &= |x\rangle \otimes \left(\frac{|0\rangle - |1\rangle}{\sqrt{2}} \right) = |x\rangle \otimes |-\rangle, x \neq x^*, \\
 |x\rangle \otimes |-\rangle &\rightarrow |x\rangle \otimes \left(\frac{|0 \oplus 1\rangle - |1 \oplus 1\rangle}{\sqrt{2}} \right) = \\
 &= |x\rangle \otimes \left(\frac{|1\rangle - |0\rangle}{\sqrt{2}} \right) = -|x\rangle \otimes |-\rangle, x = x^*,
 \end{aligned}$$

thus we have the following transformation

$$|x\rangle \otimes |-\rangle \rightarrow (-1)^{f(x)} |x\rangle \otimes |-\rangle. \quad (13.39)$$

Inversion about the Mean

Consider the circuit shown in [Figure 13.31](#). Element $\hat{U}_{x \neq 0}$ performs a transformation similar to (13.39) whereby the function $f(x=0) = 0$, and $f(x \neq 0) = 1$, thus

$$\begin{aligned}
 \hat{U}_{x \neq 0} |x\rangle \otimes |-\rangle &= |x\rangle \otimes |-\rangle, x = 0, \\
 \hat{U}_{x \neq 0} |x\rangle \otimes |-\rangle &= -|x\rangle \otimes |-\rangle, x \neq 0,
 \end{aligned}$$

i.e., the transformation matrix looks as follows

$$\begin{aligned}\hat{U}_{x \neq 0} &= \begin{pmatrix} 1 \otimes |-\rangle & 0 & 0 & \cdots & 0 \\ 0 & -1 \otimes |-\rangle & 0 & \cdots & 0 \\ 0 & 0 & -1 \otimes |-\rangle & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & -1 \otimes |-\rangle \end{pmatrix} = \\ &= \left\{ \begin{pmatrix} 2 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & 0 \end{pmatrix} - \begin{pmatrix} 1 & 0 & 0 & \cdots & 0 \\ 0 & 1 & 0 & \cdots & 0 \\ 0 & 0 & 1 & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & 1 \end{pmatrix} \right\} \otimes |-\rangle.\end{aligned}$$

Combining this result with two Hadamard transforms, and using (13.4), we get:

$$\begin{aligned}\hat{H}^{\otimes n} &\left\{ \begin{pmatrix} 2 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & 0 \end{pmatrix} - \begin{pmatrix} 1 & 0 & 0 & \cdots & 0 \\ 0 & 1 & 0 & \cdots & 0 \\ 0 & 0 & 1 & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & 1 \end{pmatrix} \right\} \otimes |-\rangle \hat{H}^{\otimes n} = \\ &= \left\{ \hat{H}^{\otimes n} \begin{pmatrix} 2 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & 0 \end{pmatrix} \hat{H}^{\otimes n} - \hat{H}^{\otimes n} \hat{I} \hat{H}^{\otimes n} \right\} \otimes |-\rangle = \\ &= \left\{ \begin{pmatrix} \frac{2}{N} & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} \end{pmatrix} - \hat{I} \right\} \otimes |-\rangle = \\ &= \left\{ \begin{pmatrix} \frac{2}{N} - 1 & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} \\ \frac{2}{N} & \frac{2}{N} - 1 & \frac{2}{N} & \cdots & \frac{2}{N} \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} - 1 & \cdots & \frac{2}{N} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} - 1 \end{pmatrix} \right\} \otimes |-\rangle. \quad (13.40)\end{aligned}$$

If we apply the operator $\hat{H}^{\otimes n} \hat{U}_{x \neq 0} \hat{H}^{\otimes n}$, then using the result (13.40) we get:

$$\begin{aligned}
 & \hat{H}^{\otimes n} \hat{U}_{x \neq 0} \hat{H}^{\otimes n} \sum_x \alpha_x |x\rangle = \\
 & = \begin{pmatrix} \frac{2}{N} - 1 & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} \\ \frac{2}{N} & \frac{2}{N} - 1 & \frac{2}{N} & \cdots & \frac{2}{N} \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} - 1 & \cdots & \frac{2}{N} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \frac{2}{N} & \frac{2}{N} & \frac{2}{N} & \cdots & \frac{2}{N} - 1 \end{pmatrix} \begin{pmatrix} \alpha_0 \\ \alpha_1 \\ \alpha_2 \\ \vdots \\ \alpha_{N-1} \end{pmatrix} = \\
 & = \begin{pmatrix} \frac{2}{N} \sum_x \alpha_x - \alpha_0 \\ \frac{2}{N} \sum_x \alpha_x - \alpha_1 \\ \frac{2}{N} \sum_x \alpha_x - \alpha_2 \\ \vdots \\ \frac{2}{N} \sum_x \alpha_x - \alpha_{N-1} \end{pmatrix} = \sum_x (2\mathcal{M} - \alpha_x) |x\rangle.
 \end{aligned}$$

Thus, the circuit proposed in Figure 13.31 indeed performs inversion about the mean.

13.8 Adiabatic Quantum Computers

Adiabatic computers are based on the so-called adiabatic theorem [Theorem 14.7.1](#).
TBD

13.9 Algorithm Complexity Theory and Quantum Computing

13.9.1 Algorithm Complexity Classes P , NP , BQP

TBD

Part 5

Appendices

Chapter 14

Basic Principles of Quantum Mechanics

The necessity of this section is primarily due to the specifics of teaching quantum mechanics in technical universities. The main emphasis is placed on practical application, and in the process, the theoretical and philosophical foundations of quantum mechanics are often overlooked.

For example, practical techniques for processing the results of many similar measurements are described in detail, i.e., the rules for calculating the average values of physical quantities in a given quantum state are described. However, answers to questions arising from the analysis of single experiments, such as the relationship between the instrument's readings and the change in the wave function during measurement, what decoherence is, and many others, are omitted. The relevance of these issues has increased recently due to the necessity to analyze the results of single experiments, leading to the possibility of designing new instruments that rely on these properties of pure quantum states. In particular, the Nobel Prize in Physics in 2012 was awarded to Serge Haroche and David J. Wineland for "ground-breaking experimental methods that enable measuring and manipulation of individual quantum systems." ¹

14.1 Dirac Formulation of Quantum Mechanics

In the course of lectures on quantum optics, we will everywhere use the Dirac formalism [55]. In the usual formulation of quantum mechanics, we deal with wave functions, for example $\psi(q, t)$ - the wave function in the coordinate representation. The same state of the system can be described by wave functions in different

¹"ground-breaking experimental methods that enable measuring and manipulation of individual quantum systems."

representations related by linear transformations. For instance, the wave function in the momentum representation is related to the wave function in the coordinate representation by the equality

$$\phi(p, t) = \frac{1}{2\pi\hbar} \int_{-\infty}^{+\infty} \psi(q, t) e^{-i\frac{pq}{\hbar}} dq \quad (14.1)$$

The main point here is that the same state can be described by wave functions expressed through different variables. From this, it follows that one can introduce a more general entity that characterizes the state of the system independently of the representation. For such an entity, Dirac introduced the concept of the wave vector, or state vector, denoted as:

$$|\dots\rangle \quad (14.2)$$

and called a ket-vector.

14.1.1 Ket-vector

$|\dots\rangle$ is a general notation for a ket-vector; $|a\rangle$, $|x\rangle$, $|\psi\rangle$, etc. denote ket-vectors describing some particular states, whose symbols are written inside the brackets.

14.1.2 Bra-vectors

Each ket-vector has a conjugate bra-vector. The bra-vector is denoted:

$$\langle \dots |, \quad \langle a |, \quad \langle \psi |. \quad (14.3)$$

The names bra- and ket-vectors come from the first and second halves of the English word *bra-cket*.

Thus, to bra-vectors $\langle a|$, $\langle x|$, $\langle \psi|$ correspond conjugate ket-vectors $|a\rangle$, $|x\rangle$, $|\psi\rangle$ and vice versa. For state vectors, the same fundamental relations hold as for wave functions:

$$|u\rangle = |a\rangle + |b\rangle, \quad \langle u| = \langle a| + \langle b|, \quad |v\rangle = l|a\rangle, \quad \langle v| = l\langle a|. \quad (14.4)$$

Bra- and ket-vectors are connected by the operation of Hermitian conjugation:

$$|u\rangle = (\langle u|)^\dagger, \quad \langle u| = (|u\rangle)^\dagger. \quad (14.5)$$

In known cases, this reduces to the following relations:

$$(\psi(q))^\dagger = \psi^*(q)$$

for the wave function in the coordinate representation;

$$\begin{pmatrix} a_1 \\ a_2 \\ \vdots \\ a_n \end{pmatrix}^\dagger = (a_1^*, a_2^*, \dots, a_n^*)$$

in the matrix representation.

Using bra- and ket-vectors, we can define the scalar product

$$\langle v|u\rangle = \langle u|v\rangle^*. \quad (14.6)$$

In specific cases, this means:

$$\langle \psi| \phi \rangle = \int \psi^* \phi dq$$

in the coordinate representation;

$$\langle a|b\rangle = (a_1^*, a_2^*, \dots, a_n^*) \begin{pmatrix} b_1 \\ b_2 \\ \vdots \\ b_n \end{pmatrix} = a_1^* b_1 + a_2^* b_2 + \dots + a_n^* b_n$$

in the matrix representation.

From relation (14.6) it follows that the norm of a vector is real. Additionally, we assume that the norm of the vector is positive or zero: $\langle a|a\rangle \geq 0$.

14.1.3 Operators

In quantum mechanics, linear operators are used. Operators connect one state vector with another:

$$|q\rangle = \hat{L} |p\rangle \quad (14.7)$$

The conjugate equality has the form

$$\langle q| = \langle p| \hat{L}^\dagger \quad (14.8)$$

where \hat{L}^\dagger is the operator conjugate to \hat{L} .

Let us give some relations valid for linear operators:

$$\begin{aligned}
 \hat{L}^{++} &= \hat{L}, \quad \left(l \hat{L} |a\rangle \right)^\dagger = l^* \langle a| \hat{L}^\dagger, \\
 \left(\left(\hat{L}_1 + \hat{L}_2 \right) |a\rangle \right)^\dagger &= \langle a| \left(\hat{L}_1^\dagger + \hat{L}_2^\dagger \right), \\
 \left(\left(\hat{L}_1 \hat{L}_2 \right) |a\rangle \right)^\dagger &= \langle a| \left(\hat{L}_2^\dagger \hat{L}_1^\dagger \right), \\
 \left(\left(\hat{L}_1 \hat{L}_2 \hat{L}_3 \right) |a\rangle \right)^\dagger &= \langle a| \left(\hat{L}_3^\dagger \hat{L}_2^\dagger \hat{L}_1^\dagger \right), \text{ and so on}
 \end{aligned} \tag{14.9}$$

Note that the algebra of operators coincides with the algebra of square matrices. The matrix elements of operators are denoted as follows:

$$\langle a| \hat{L} |b\rangle = L_{ab} \tag{14.10}$$

For matrix elements, the following equalities hold:

$$\langle a| \hat{L} |b\rangle^* = \langle b| \hat{L}^\dagger |a\rangle, \quad \langle a| \hat{L}_1 \hat{L}_2 |b\rangle^* = \langle b| \hat{L}_2^\dagger \hat{L}_1^\dagger |a\rangle \tag{14.11}$$

14.1.4 Eigenvalues and eigenvectors of operators

Eigenvalues and eigenvectors of operators are defined by the equality

$$\hat{L} |l_n\rangle = l_n |l_n\rangle, \tag{14.12}$$

where l_n is an eigenvalue; $|l_n\rangle$ an eigenvector.

For bra-vectors we have analogous equalities:

$$\langle d_n| \hat{D} = d_n \langle d_n|. \tag{14.13}$$

If operators correspond to observables, they must be self-adjoint:

$$\hat{L} = \hat{L}^\dagger. \tag{14.14}$$

The eigenvalues of a self-adjoint (Hermitian) operator are real. Indeed, from

$$\hat{L} |l\rangle = l |l\rangle$$

it follows that

$$\langle l| \hat{L} |l\rangle = l \langle l|l\rangle.$$

On the other hand, recalling (14.9): $\langle l| \hat{L}^\dagger = l^* \langle l|$, from (14.14) we have

$$\langle l| \hat{L} |l\rangle = l^* \langle l|l\rangle.$$

Thus $l \langle l|l \rangle = l^* \langle l|l \rangle$, i.e. $l = l^*$.

Eigenvectors of a self-adjoint operator are orthogonal. Indeed, consider two eigenvectors $|l_1\rangle$ and $|l_2\rangle$:

$$\hat{L} |l_1\rangle = l_1 |l_1\rangle, \quad \hat{L} |l_2\rangle = l_2 |l_2\rangle$$

From the second relation we get

$$\langle l_1 | \hat{L} |l_2\rangle = l_2 \langle l_1 |l_2\rangle$$

Taking into account the reality of eigenvalues and relation (14.14) for the vector $|l_1\rangle$ we obtain:

$$\langle l_1 | \hat{L} = l_1 \langle l_1 |.$$

Hence

$$\langle l_1 | \hat{L} |l_2\rangle = l_1 \langle l_1 |l_2\rangle.$$

Thus

$$(l_1 - l_2) \langle l_1 |l_2\rangle = 0, \quad \text{i.e. } \langle l_1 |l_2\rangle = 0, \text{ since } l_1 \neq l_2.$$

14.1.5 Observables. Expansion in eigenvectors. Completeness of set of eigenvectors

Operators corresponding to physical observables are self-adjoint operators. This ensures the reality of the values of the observable physical quantity. We have a set of eigenstates of some Hermitian operator $|l_n\rangle$, $\hat{L} |l_n\rangle = l_n |l_n\rangle$. If the set of eigenstates is complete, according to the principles of quantum mechanics any state can be represented as a superposition of states $|l_n\rangle$:

$$|\psi\rangle = \sum_{(n)} c_n |l_n\rangle. \quad (14.15)$$

From here for the expansion coefficients we have: $c_n = \langle l_n | \psi \rangle$, and thus, the equality holds

$$|\psi\rangle = \sum_{(n)} \langle l_n | \psi \rangle |l_n\rangle = \sum_{(n)} |l_n\rangle \langle l_n | \psi \rangle. \quad (14.16)$$

From equality 14.16 follows an important relation:

$$\sum_{(n)} |l_n\rangle \langle l_n| = \hat{I}. \quad (14.17)$$

where \hat{I} is the identity operator. This equality is the condition of completeness of the set of eigenvectors (the condition of decomposability).



Figure 14.1: Projection Operator. The action of the operator can be interpreted as the projection of the vector $|\psi\rangle$ onto the axis $|l_n\rangle$

14.1.6 Projection operator

Consider the operator $\hat{P}_n = |l_n\rangle \langle l_n|$. The result of action of this operator on the state $|\psi\rangle$ will be

$$\hat{P}_n |\psi\rangle = \sum_{(k)} |l_n\rangle \langle l_n| c_k |l_k\rangle = c_n |l_n\rangle. \quad (14.18)$$

The operator $\hat{P}_n = |l_n\rangle \langle l_n|$ is called the projection operator.

One can write the following properties of this operator

$$\sum_{(n)} \hat{P}_n = \hat{I}. \quad (14.19)$$

$$\hat{P}_n^2 = \hat{P}_n. \quad (14.20)$$

The action of the projection operator has a simple geometric interpretation (see [Figure 14.1](#)):

$$\hat{P}_n |\psi\rangle = \cos \theta |l_n\rangle,$$

where $\cos \theta = \langle \psi | l_n \rangle = c_n$.

14.1.7 Trace of an operator

In an orthonormal basis $\{|l_n\rangle\}$, the quantity

$$Sp \hat{L} = \sum_n \langle l_n | \hat{L} | l_n \rangle \quad (14.21)$$

is called the trace of the operator \hat{L} . Under certain conditions [\[39\]](#), the series [14.21](#) converges absolutely and does not depend on the choice of basis.

If we use the matrix representation

$$L_{kn} = \langle l_k | \hat{L} | l_n \rangle ,$$

then the trace of the operator is the sum of the diagonal elements of the matrix representation

$$Sp \hat{L} = \sum_n L_{nn}$$

One can write the following properties of the trace:

$$\begin{aligned} Sp(l\hat{L} + m\hat{M}) &= lSp\hat{L} + mSp\hat{M}, \\ Sp(\hat{L}\hat{M}) &= Sp(\hat{M}\hat{L}). \end{aligned} \quad (14.22)$$

14.1.8 Mean values of operators

The mean value of operator \hat{L} in the state $|\psi\rangle$ is given by the equality

$$\langle \hat{L} \rangle_\psi = \langle \psi | \hat{L} | \psi \rangle \quad (14.23)$$

under the condition

$$\langle \psi | \psi \rangle = 1.$$

Indeed, if one assumes that $|\psi\rangle$ is expanded in the series of eigenfunctions of the operator \hat{L} as follows:

$$|\psi\rangle = \sum_n c_n |l_n\rangle ,$$

then $\hat{L}|\psi\rangle$ can be written as

$$\hat{L}|\psi\rangle = \sum_n l_n c_n |l_n\rangle ,$$

where l_n is the eigenvalue corresponding to the eigenstate $|l_n\rangle$. If now the last two expressions are substituted into (14.23) one obtains:

$$\langle \psi | \hat{L} | \psi \rangle = \sum_{n,m} l_n c_n c_m^* \langle l_m | l_n \rangle = \sum_n l_n c_n c_n^* = \sum_n l_n |c_n|^2 ,$$

which (under the condition $\langle \psi | \psi \rangle = 1$) proves that expression (14.23) indeed represents the expression for the mean value of the operator \hat{L} in the state $|\psi\rangle$. ²

²For this, it is enough to recall that $|c_n|^2$ gives the probability to find the system in the state $|l_n\rangle$, i.e., to obtain a reading of the measuring device equal to l_n

If one takes some orthonormal basis $\{|n\rangle\}$, forming a complete set, i.e., satisfying the condition (14.17): $\sum_n |n\rangle \langle n| = \hat{I}$, then expression (14.23) can be rewritten as follows:

$$\begin{aligned} \langle \hat{L} \rangle_\psi &= \langle \psi | \hat{L} | \psi \rangle = \langle \psi | \hat{I} \hat{L} | \psi \rangle = \\ &= \sum_n \langle \psi | n \rangle \langle n | \hat{L} | \psi \rangle = \sum_n \langle n | \hat{L} | \psi \rangle \langle \psi | n \rangle = Sp(\hat{L} \hat{\rho}), \end{aligned}$$

where $\hat{\rho} = |\psi\rangle \langle \psi| = \hat{P}_\psi$ is the projection operator onto the state $|\psi\rangle$. Taking into account (14.22), one can write

$$\langle \hat{L} \rangle_\psi = Sp(\hat{\rho} \hat{L}). \quad (14.24)$$

14.1.9 Representation of operators through outer products of eigenvectors

Using the completeness condition (14.16) twice, we get

$$\hat{A} = \hat{I} \hat{A} \hat{I} = \sum_{(l)} \sum_{(l')} |l\rangle \langle l| \hat{A} |l'\rangle \langle l'| = \sum_{(l)} \sum_{(l')} |l\rangle \langle l'| A_{ll'}, \quad (14.25)$$

where $A_{ll'} = \langle l | \hat{A} | l' \rangle$ is the matrix element of the operator \hat{A} in the $|l\rangle$ representation.

The operator expressed through its own eigenvectors can be represented by the expansion ³

$$\hat{L} = \sum_{(l)} l |l\rangle \langle l|. \quad (14.26)$$

The generalization of this equality for an operator function has the form

$$F(\hat{L}) = \sum_{(l)} F(l) |l\rangle \langle l|. \quad (14.27)$$

14.1.10 Wave functions in coordinate and momentum representations

Transition from the state vector to the wave function is carried out by scalar multiplying this state vector by the eigenvector of the corresponding observable. For example, for the wave function in the coordinate representation

$$\phi(q) = \langle q | \psi \rangle. \quad (14.28)$$

³Under the condition of normalization of eigenvectors: $\langle l | l \rangle = 1$

where $\langle q|$ is the eigenvector of the coordinate operator. In the momentum representation, we get

$$\phi(p) = \langle p| \psi \rangle. \quad (14.29)$$

where $\langle p|$ is the eigenvector of the momentum operator.

14.2 Dynamics of wave function change

The wave function $|\phi\rangle$ can change through two mechanisms:

- Collapse of the wave function during measurement
- Schrödinger equation in the intervals between two consecutive measurements

14.2.1 Schrödinger Equation

The change of the state of a pure quantum system between two consecutive measurements is described by the following equation (Schrödinger)

$$i\hbar \frac{\partial |\phi\rangle}{\partial t} = \hat{\mathcal{H}} |\phi\rangle. \quad (14.30)$$

Equation (14.30) is reversible and, accordingly, not applicable to describing the change of the wave function at the moment of measurement.

It is worth noting the connection of the Schrödinger equation with Stone's theorem (Theorem 17.2.1) TBD.

Schrödinger equation in the interaction picture

Suppose that the Hamiltonian can be separated into two parts:

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_0 + \hat{\mathcal{V}}.$$

We introduce the following transformation of the wave function:

$$|\phi\rangle_I = \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} |\phi\rangle$$

and examine the value of the following expression:

$$\begin{aligned}
i\hbar \frac{\partial |\phi\rangle_I}{\partial t} &= i\hbar \frac{i\hat{\mathcal{H}}_0}{\hbar} \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} |\phi\rangle + \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} i\hbar \frac{\partial |\phi\rangle}{\partial t} = \\
&= -\hat{\mathcal{H}}_0 \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} |\phi\rangle + \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} (\hat{\mathcal{H}}_0 + \hat{\mathcal{V}}) |\phi\rangle = \\
&= -\hat{\mathcal{H}}_0 \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} |\phi\rangle + \hat{\mathcal{H}}_0 \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} |\phi\rangle + \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} \hat{\mathcal{V}} |\phi\rangle = \\
&= \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} \hat{\mathcal{V}} |\phi\rangle = \\
&= \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} \hat{\mathcal{V}} \exp\left\{\left(-\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} |\phi\rangle = \\
&= \hat{\mathcal{V}}_I |\phi\rangle_I,
\end{aligned}$$

where

$$\hat{\mathcal{V}}_I = \exp\left\{\left(\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} \hat{\mathcal{V}} \exp\left\{\left(-\frac{i\hat{\mathcal{H}}_0 t}{\hbar}\right)\right\} \quad (14.31)$$

is the interaction Hamiltonian in the interaction picture.

Thus, we obtain the Schrödinger equation in the interaction picture:

$$i\hbar \frac{\partial |\phi\rangle_I}{\partial t} = \hat{\mathcal{V}}_I |\phi\rangle_I. \quad (14.32)$$

Equation of motion for the density matrix

From relation (14.30) we have

$$\begin{aligned}
i\hbar \frac{\partial |\phi\rangle}{\partial t} &= \hat{\mathcal{H}} |\phi\rangle, \\
-i\hbar \frac{\partial \langle\phi|}{\partial t} &= \hat{\mathcal{H}} \langle\phi|,
\end{aligned}$$

thus, for the density matrix $\hat{\rho} = |\phi\rangle \langle\phi|$ we get

$$\begin{aligned}
i\hbar \frac{\partial \hat{\rho}}{\partial t} &= i\hbar \frac{\partial |\phi\rangle \langle\phi|}{\partial t} = i\hbar \left(\frac{\partial |\phi\rangle}{\partial t} \langle\phi| + |\phi\rangle \frac{\partial \langle\phi|}{\partial t} \right) = \\
&= \hat{\mathcal{H}} |\phi\rangle \langle\phi| - |\phi\rangle \langle\phi| \hat{\mathcal{H}} = [\hat{\mathcal{H}}, \hat{\rho}]
\end{aligned} \quad (14.33)$$

Equation (14.33) is often called the quantum Liouville equation and the von Neumann equation.

Evolution operator. Heisenberg and Schrödinger pictures

The change of the wave function according to the law (14.30) can also be described using a certain operator (evolution) $\hat{U}(t, t_0)$:

$$|\phi(t)\rangle = \hat{U}(t, t_0) |\phi(t_0)\rangle. \quad (14.34)$$

Equation (14.30) can be rewritten as

$$|\phi(t)\rangle = \exp\left(-\frac{i}{\hbar}\hat{\mathcal{H}}(t - t_0)\right) |\phi(t_0)\rangle,$$

from which for the evolution operator we have

$$\hat{U}(t, t_0) = \exp\left(-\frac{i}{\hbar}\hat{\mathcal{H}}(t - t_0)\right) \quad (14.35)$$

The evolution operator is unitary. Indeed:

$$\begin{aligned} & \hat{U}(t, t_0) \hat{U}^\dagger(t, t_0) = \\ & = \exp\left(-\frac{i}{\hbar}\hat{\mathcal{H}}(t - t_0)\right) \exp\left(+\frac{i}{\hbar}\hat{\mathcal{H}}(t - t_0)\right) = \hat{I} \end{aligned}$$

Alongside the Schrödinger picture where operators do not depend on time and wave functions evolve, there exists the Heisenberg picture where operators evolve with time.

Obviously, the expectation values of operators should not depend on the picture:

$$\begin{aligned} \langle \phi_H(t_0) | \hat{A}_H(t) | \phi_H(t_0) \rangle &= \langle \phi_S(t) | \hat{A}_S | \phi_S(t) \rangle = \\ &= \langle \phi_H(t_0) | \hat{U}^\dagger(t, t_0) \hat{A}_S \hat{U}(t, t_0) | \phi_H(t_0) \rangle, \end{aligned}$$

from which, given $\hat{A}_H(t_0) = \hat{A}_S(t_0)$, we obtain the evolution law for operators in the Heisenberg picture:

$$\hat{A}_H(t) = \hat{U}^\dagger(t, t_0) \hat{A}_H(t_0) \hat{U}(t, t_0) \quad (14.36)$$

At the same time, the equation for the operator \hat{A}_H looks like:

$$\begin{aligned} \frac{\partial \hat{A}_H}{\partial t} &= \frac{i}{\hbar} \hat{\mathcal{H}} \hat{U}^\dagger(t, t_0) \hat{A}_H(t_0) \hat{U}(t, t_0) - \\ &- \frac{i}{\hbar} \hat{U}^\dagger(t, t_0) \hat{A}_H(t_0) \hat{U}(t, t_0) \hat{\mathcal{H}} = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{A}_H] \end{aligned} \quad (14.37)$$

14.2.2 Differences between pure and mixed states. Decoherence

Definition 14.2.1 (Pure state). If the state of a system is described by a density matrix $\hat{\rho}$ which can be written as

$$\hat{\rho} = |\psi\rangle \langle\psi| \quad (14.38)$$

then this state is called pure.

Definition 14.2.2 (Mixed state). If the state of a system is described by a density matrix $\hat{\rho}$ which **cannot** be written in the form (14.38), i.e.

$$\hat{\rho} \neq |\psi\rangle \langle\psi|$$

then this state is called mixed.

Of particular interest is the difference between pure and mixed states, especially how the transition from pure to mixed states occurs.

Consider a two-level state (see Figure 14.2). In a pure state it is described by the following wave function:

$$|\phi\rangle = c_a |a\rangle + c_b |b\rangle,$$

the corresponding density matrix is

$$\begin{aligned} \hat{\rho} &= |\phi\rangle \langle\phi| = \\ &= |c_a|^2 |a\rangle \langle a| + |c_b|^2 |b\rangle \langle b| + \\ &\quad + c_a c_b^* |a\rangle \langle b| + c_b c_a^* |b\rangle \langle a|, \end{aligned} \quad (14.39)$$

or in matrix form

$$\hat{\rho} = \begin{pmatrix} |c_a|^2 & c_a c_b^* \\ c_b c_a^* & |c_b|^2 \end{pmatrix}.$$

The density matrix for a mixed state has only diagonal elements:

$$\begin{aligned} \hat{\rho} &= \begin{pmatrix} |c_a|^2 & 0 \\ 0 & |c_b|^2 \end{pmatrix} = \\ &= |c_a|^2 |a\rangle \langle a| + |c_b|^2 |b\rangle \langle b|. \end{aligned} \quad (14.40)$$

The transition from (14.39) to (14.40) is called decoherence. In describing the decoherence process we will follow [53].



Figure 14.2: Model of a two-level atom used to describe decoherence.

The difference of mixed states from pure states manifests in the influence of the environment \mathcal{E} . In the case of pure states the considered system and its environment are independent, i.e.

$$|\phi\rangle_{pure} = |\phi\rangle_{at} \otimes |\mathcal{E}\rangle. \quad (14.41)$$

In the case of mixed states the atom and its environment form what is called an entangled state where the states $|a\rangle$ and $|b\rangle$ correspond to distinguishable environment states $|\mathcal{E}_a\rangle$ and $|\mathcal{E}_b\rangle$.

$$|\phi\rangle_{mix} = c_a |a\rangle |\mathcal{E}_a\rangle + c_b |b\rangle |\mathcal{E}_b\rangle. \quad (14.42)$$

The density matrix corresponding to (14.42) is

$$\begin{aligned} \hat{\rho}_{mix} &= |\phi\rangle_{mix} \langle\phi|_{mix} = \\ &= |c_a|^2 |a\rangle \langle a| \otimes |\mathcal{E}_a\rangle \langle \mathcal{E}_a| + |c_b|^2 |b\rangle \langle b| \otimes |\mathcal{E}_b\rangle \langle \mathcal{E}_b| + \\ &\quad + c_a c_b^* |a\rangle \langle b| \otimes |\mathcal{E}_a\rangle \langle \mathcal{E}_b| + c_b c_a^* |b\rangle \langle a| \otimes |\mathcal{E}_b\rangle \langle \mathcal{E}_a|. \end{aligned} \quad (14.43)$$

If we now apply averaging over the environment variables to expression (14.43), we obtain

$$\begin{aligned} \langle \hat{\rho}_{mix} \rangle_{\mathcal{E}} &= Sp_{\mathcal{E}} (\hat{\rho}) = \\ &= \langle \mathcal{E}_a | \hat{\rho}_{mix} | \mathcal{E}_a \rangle + \langle \mathcal{E}_b | \hat{\rho}_{mix} | \mathcal{E}_b \rangle = \\ &= |c_a|^2 |a\rangle \langle a| + |c_b|^2 |b\rangle \langle b|. \end{aligned} \quad (14.44)$$

Expression (14.44) is obtained under the assumption of an orthonormal basis $\{|\mathcal{E}_a\rangle, |\mathcal{E}_b\rangle\}$:

$$\begin{aligned} \langle \mathcal{E}_a | \mathcal{E}_a \rangle &= \langle \mathcal{E}_b | \mathcal{E}_b \rangle = 1, \\ \langle \mathcal{E}_a | \mathcal{E}_b \rangle &= \langle \mathcal{E}_b | \mathcal{E}_a \rangle = 0. \end{aligned} \quad (14.45)$$

Conditions (14.45) are key to understanding why the considered basis of the atomic system is preferred and why, for instance, in the case of mixed states

one does not consider other bases such as the basis obtained by the Hadamard transform relative to the original :

$$\begin{aligned} |\mathcal{A}\rangle &= \frac{|a\rangle + |b\rangle}{\sqrt{2}}, \\ |\mathcal{B}\rangle &= \frac{|a\rangle - |b\rangle}{\sqrt{2}}. \end{aligned} \quad (14.46)$$

The environment states corresponding to the basis (14.46) are not orthogonal, hence it is impossible to use (14.46) as basis vectors for mixed states.

The decoherence process, i.e., the transition from (14.41) to (14.42), can be described using the Schrödinger equation and therefore theoretically reversible. The only requirement is the orthogonality of distinguishable environment states: $\langle \mathcal{E}_a | \mathcal{E}_b \rangle = 0$. This requirement is always fulfilled for macroscopic systems, where the state depends on a very large number of variables. Moreover, in the case of macroscopic systems there exist many possible final states $|\mathcal{E}_{a,b}\rangle$, which makes the reverse process practically unrealizable because it is necessary to control a large number of possible variables describing the environment state. In this sense, the decoherence process has the same nature as the second law of thermodynamics (increasing entropy), which describes irreversible processes.⁴

The decoherence process is very fast; in particular, [31] provides the following estimate: for a system of mass 1 g at a separation $\Delta x = 1$ cm and temperature $T = 300$ K, with relaxation time equal to the lifetime of the Universe $\tau_R = 10^{17}$ s, the decoherence process takes 10^{-23} s

14.2.3 Collapse of the wave function. Measurement in quantum mechanics

The process of selection (measurement result) is one of the most complex in quantum mechanics. Unlike deterministic change of the wave function described by the Schrödinger equation (14.30), the measurement process has a random character and different equations must be used to describe it.

Let us first consider pure states and suppose that a measurement is performed of a physical observable described by the operator \hat{L} . The eigenvalues and eigenfunctions of this operator are $\{l_k\}$ and $\{|l_k\rangle\}$ respectively. At the moment of measurement, the instrument readout can take values corresponding to the eigenvalues of the measured operator (see Figure 14.3). Suppose that the readout is l_n , in which case the wave function must be $|l_n\rangle$, i.e. the following change of the

⁴One should be somewhat careful here since the second law of thermodynamics applies to closed systems, while the decoherence processes happen in open systems.



Figure 14.3: Measurement process. The instrument reading corresponds to one of the eigenvalues of the operator \hat{L} : $\{l_k\}$

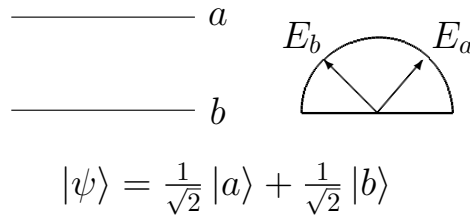


Figure 14.4: The process of measuring the energy of a two-level atom in a pure state $|\psi\rangle = \frac{1}{\sqrt{2}} |a\rangle + \frac{1}{\sqrt{2}} |b\rangle$. The instrument records the energy value E_a or E_b .

wave function occurred:

$$|\phi\rangle \rightarrow |l_n\rangle,$$

which can be described by the action of the projection operator $\hat{P}_n = |l_n\rangle \langle l_n|$ (14.18):

$$\hat{P}_n |\phi\rangle = c_n |l_n\rangle.$$

Example 14.2.1 (Measurement of energy of a two-level atom). Consider a two-level atom in a pure state (see Figure 14.4) $|\psi\rangle = \frac{1}{\sqrt{2}} |a\rangle + \frac{1}{\sqrt{2}} |b\rangle$.

Our instrument measures the energy of this atom and the Hamiltonian operator has 2 eigenfunctions $|a, b\rangle$, corresponding to eigenvalues E_a, E_b . Thus, possible readings of the instrument belong to the set $\{E_a, E_b\}$.

When the instrument pointer shows E_a , the following collapse occurs (see Figure 14.5)

$$|\psi\rangle \rightarrow |a\rangle.$$

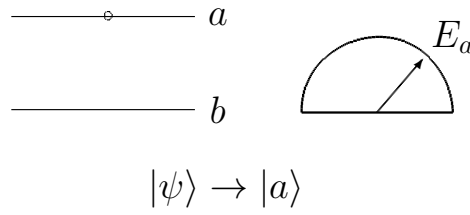


Figure 14.5: The process of measuring the energy of a two-level atom in a pure state $|\psi\rangle = \frac{1}{\sqrt{2}}|a\rangle + \frac{1}{\sqrt{2}}|b\rangle$. The device registers the energy value E_a . During the measurement, the following reduction occurs $|\psi\rangle \rightarrow |a\rangle$

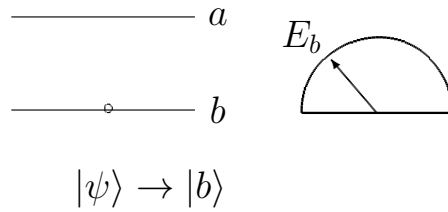


Figure 14.6: The process of measuring the energy of a two-level atom in a pure state $|\psi\rangle = \frac{1}{\sqrt{2}}|a\rangle + \frac{1}{\sqrt{2}}|b\rangle$. The instrument registers the energy value E_b . During the measurement, the following reduction occurs $|\psi\rangle \rightarrow |b\rangle$

Similarly, when the reading is E_b , the following collapse occurs (see [Figure 14.6](#))

$$|\psi\rangle \rightarrow |b\rangle.$$

There is no way to predict the result that will be obtained from a single measurement. However, one can specify the probability of each possible result.

Indeed, in the case of a mixed state

$$\hat{\rho} = \sum_n |c_n|^2 |l_n\rangle \langle l_n|$$

the coefficients $P_n = |c_n|^2$ give the probabilities to find the system in state $|l_n\rangle$.

For a pure state

$$|\phi\rangle = \sum_n c_n |l_n\rangle$$

we also have that the probability to find the system in state $|l_n\rangle$ is given by the number $P_n = |c_n|^2$.

The main difference between pure and mixed states from the point of view of measurement is that in the former (pure states) the wave function and thus the state itself changes during measurement. In particular, if a certain final state $|l_i\rangle$ was obtained during measurement, one cannot say that it was the same before measurement. Mixed states behave like classical objects, i.e. if a state $|l_i\rangle$ was

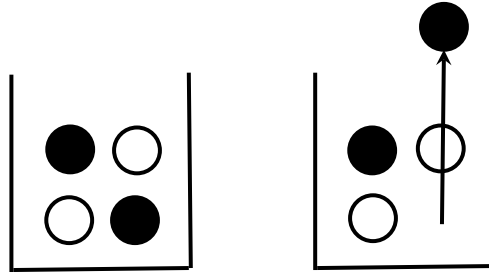


Figure 14.7: Example of a mixed state. The color of the sphere does not change as a result of "measurement"

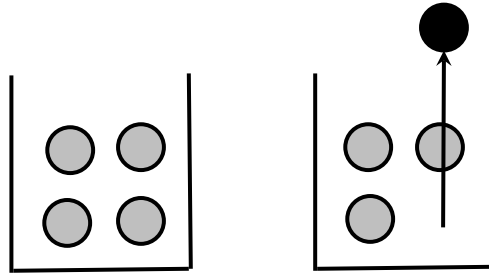


Figure 14.8: Example of a pure state. The color of the sphere changes as a result of a "measurement"

obtained during measurement, it can be claimed that it was the same before measurement and the measurement itself represents choosing one state from many possible ones.

Example 14.2.2. Choosing from an urn with balls of two colors *Suppose we have an urn with 4 balls. With probability $\frac{1}{2}$ either a white or a black ball will be drawn. Suppose that as a result of the experiment a black ball was obtained. If the system under consideration is quantum and in a mixed state (see [Figure 14.7](#)), then the state of the drawn ball (color) did not change as a result of the experiment.*

If the system under consideration is pure (see [Figure 14.8](#)), then the state of each ball is described by a superposition of two colors - black and white. Thus, as a result of the experiment this superposition collapses and the ball acquires a definite color (black in our case), i.e., one can say that the color of the ball changes.

14.3 Measurability of Physical Quantities

14.3.1 Heisenberg Uncertainty Principle

Suppose two operators \hat{A} and \hat{B} do not commute with each other, i.e.

$$[\hat{A}\hat{B}] = \hat{A}\hat{B} - \hat{B}\hat{A} = i\hat{C} \neq 0,$$

where \hat{C} is some Hermitian operator.

Suppose the system is in the state $|\psi\rangle$. Then the average values of the operators are expressed by the following relations:

$$\begin{aligned}\langle\psi|\hat{A}|\psi\rangle &= \langle\hat{A}\rangle, \\ \langle\psi|\hat{B}|\psi\rangle &= \langle\hat{B}\rangle.\end{aligned}$$

Define the uncertainty of measuring the quantities A and B as their variances as follows:

$$\begin{aligned}\Delta A &= \sqrt{\langle\psi|\hat{\mathcal{A}}^2|\psi\rangle}, \\ \Delta B &= \sqrt{\langle\psi|\hat{\mathcal{B}}^2|\psi\rangle},\end{aligned}$$

where

$$\begin{aligned}\hat{\mathcal{A}} &= \hat{A} - \langle\hat{A}\rangle, \\ \hat{\mathcal{B}} &= \hat{B} - \langle\hat{B}\rangle.\end{aligned}$$

Introduce the operator \hat{D} as follows:

$$\hat{D} = \hat{\mathcal{A}} + i\lambda\hat{\mathcal{B}}.$$

Consider the operator $\hat{D}^\dagger\hat{D}$, which is Hermitian. Its expectation in the state $|\psi\rangle$ is:

$$\langle\psi|\hat{D}^\dagger\hat{D}|\psi\rangle = \langle\phi|\phi\rangle \geq 0,$$

where $|\phi\rangle = \hat{D}|\psi\rangle$. On the other hand,

$$\begin{aligned}\langle\psi|\hat{D}^\dagger\hat{D}|\psi\rangle &= \langle\psi|(\hat{\mathcal{A}} - i\lambda\hat{\mathcal{B}})(\hat{\mathcal{A}} + i\lambda\hat{\mathcal{B}})|\psi\rangle = \\ &= \langle\psi|\hat{\mathcal{A}}^2|\psi\rangle + \lambda^2\langle\psi|\hat{\mathcal{B}}^2|\psi\rangle + i\lambda\langle\psi|[\hat{\mathcal{A}},\hat{\mathcal{B}}]|\psi\rangle = \\ &= (\Delta A)^2 + \lambda^2(\Delta B)^2 + i\lambda\langle\psi|[\hat{A},\hat{B}]|\psi\rangle = \\ &= \lambda^2(\Delta B)^2 - \lambda\langle C\rangle + (\Delta A)^2 \geq 0.\end{aligned}$$

Consider the polynomial

$$f(\lambda) = \lambda^2(\Delta B)^2 - \lambda\langle C\rangle + (\Delta A)^2.$$

We have $f(\pm\infty) > 0$, so $f(\lambda) \geq 0$ if this polynomial has at most one real root, i.e.

$$\langle C \rangle^2 - 4(\Delta A)^2(\Delta B)^2 \leq 0$$

or

$$\Delta A \Delta B \geq \frac{|\langle C \rangle|}{2}, \quad (14.47)$$

which is the Heisenberg inequality.

14.3.2 Energy-Time Uncertainty Relation

Time in quantum mechanics does not have a corresponding operator, and therefore to estimate the time one should use some separate observable \hat{O} . Using (14.47), the following relation can be obtained:

$$\Delta E \Delta O \geq \frac{|\langle C \rangle|}{2},$$

where, taking into account (14.37),

$$\frac{d\hat{O}}{dt} = \frac{i}{\hbar} [\hat{\mathcal{H}}, \hat{O}]$$

and therefore

$$C = \frac{1}{i} [\hat{\mathcal{H}}, \hat{O}] = -\hbar \frac{d\hat{O}}{dt}.$$

Thus,

$$\Delta E \Delta O \geq \frac{|\langle C \rangle|}{2} = \frac{\hbar}{2} \left| \left\langle \frac{d\hat{O}}{dt} \right\rangle \right|.$$

Denoting $\Delta t = \frac{\Delta O}{\left| \left\langle \frac{d\hat{O}}{dt} \right\rangle \right|}$ ⁵ finally we get

$$\Delta E \Delta t \geq \frac{\hbar}{2}. \quad (14.48)$$

It is worth noting that states with definite energy do not contradict relation (14.48) because although $\Delta E = 0$, for any observable we have $\left\langle \frac{d\hat{O}}{dt} \right\rangle = 0$ and consequently, $\Delta t = \infty$.

⁵We can assume that at small times the observable \hat{O} changes linearly, i.e. $\left\langle \frac{d\hat{O}}{dt} \right\rangle = \frac{\Delta O}{\Delta t}$. Thus, Δt is the time of significant change of the observable O .

14.3.3 Simultaneous Measurability of Physical Quantities

The question of simultaneous measurability of two physical quantities is very meaningful, especially when studying quantum mechanical paradoxes such as the EPR paradox (see [section 10.1](#)).

Suppose we have two observables \hat{A}, \hat{B} , each corresponding to a set of eigenfunctions $|a_i\rangle, |b_i\rangle$ and eigenvalues $\{a_i\}, \{b_i\}$, i.e.

$$\begin{aligned}\hat{A}|a_i\rangle &= a_i|a_i\rangle, \\ \hat{B}|b_i\rangle &= b_i|b_i\rangle.\end{aligned}$$

If at some moment we can measure the values of these quantities simultaneously, then the obtained values can be described by numbers a and b , where obviously $a \in \{a_i\}, b \in \{b_i\}$ and

$$\begin{aligned}\hat{A}|a\rangle &= a|a\rangle, \\ \hat{B}|b\rangle &= b|b\rangle.\end{aligned}$$

Since the measurement of the quantities was performed simultaneously, due to the collapse of the wavefunction (see [subsection 14.2.3](#)), we have

$$|a\rangle = |b\rangle,$$

that is, the operators \hat{A} and \hat{B} have common eigenfunctions. It is not necessary that all eigenvectors of these operators coincide, it is sufficient that a single vector matches.

Note that if the operators commute, then their sets of eigenvectors coincide, which is consistent with the Heisenberg inequality, which in this case takes the form

$$\Delta a \Delta b \geq 0.$$

Indeed, if

$$\hat{A}|a\rangle = a|a\rangle,$$

then, considering $\hat{A}\hat{B} = \hat{B}\hat{A}$,

$$\hat{A}\hat{B}|a\rangle = \hat{B}\hat{A}|a\rangle = a\hat{B}|a\rangle.$$

Denoting $|b\rangle = \hat{B}|a\rangle$, we have

$$\hat{A}|b\rangle = a|b\rangle,$$

i.e., $|b\rangle = |a\rangle$ and this vector is an eigenvector of both operators \hat{A} and \hat{B} . Thus, if the operators commute, they have a common basis (set of eigenvectors). The converse is also true, see [\[36\]](#).

14.4 Quantization Procedure

As we have already found out, physical quantities in quantum mechanics correspond to self-adjoint operators. If we know the wave function (for example, by means of the Schrödinger equation), then knowing the expression for the operator of the physical quantity of interest, we can perform practical calculations, such as expectation values.

The procedure of deriving the relations that operators obey is called quantization.

Below is information about the connection between classical mechanics and the quantization procedure, as well as an example of quantizing the angular momentum operator.

14.4.1 Generalized Coordinates and Momenta

To fully describe a mechanical system, it is necessary to specify its coordinates and momenta. The coordinate is not always a classical coordinate, for example, positions in Cartesian coordinate system (x, y, z) . Sometimes it is convenient to use special coordinates to describe the system under consideration, such as an angle in polar coordinates. In this case, the question arises of what should be considered the momentum corresponding to such a coordinate. To answer this question, one should consider the Lagrangian, which is defined as the difference between kinetic and potential energies:

$$\mathcal{L} = T - U.$$

If we have a set of generalized coordinates q_n , then the Lagrangian \mathcal{L} can be presented as a function of these variables and their time derivatives:

$$\mathcal{L} = \mathcal{L}(q_1, q_2, \dots, \dot{q}_1, \dot{q}_2, \dots).$$

We can define the generalized momentum p_i corresponding to the generalized coordinate q_i as

$$p_i = \frac{\partial \mathcal{L}}{\partial \dot{q}_i}.$$

In quantum mechanics, the equations of motion in Hamiltonian form play a special role:

$$\begin{aligned} \dot{q}_i &= \frac{\partial \mathcal{H}}{\partial p_i}, \\ \dot{p}_i &= -\frac{\partial \mathcal{H}}{\partial q_i}, \end{aligned} \tag{14.49}$$

where \mathcal{H} is the Hamiltonian (total energy) of the system as a function of generalized coordinates and momenta.

Example 14.4.1 (Rectilinear Motion of a Material Point). *Suppose we have a material point moving along the x -axis under the action of a force $F = -\frac{dU}{dx}$, where U is the potential energy. In this case:*

$$\mathcal{L} = T - U = \frac{m\dot{x}^2}{2} - U(x).$$

The generalized momentum p corresponding to coordinate x is

$$p = \frac{\partial \mathcal{L}}{\partial \dot{x}} = m\dot{x},$$

which coincides with the classical definition of momentum.

The Hamiltonian is

$$\mathcal{H}(x, p) = T + U(x) = \frac{p^2}{2m} + U(x).$$

Equations of motion:

$$\begin{aligned}\dot{x} &= \frac{\partial \mathcal{H}}{\partial p} = \frac{p}{m}, \\ \dot{p} &= -\frac{\partial \mathcal{H}}{\partial x} = -\frac{dU}{dx} = F,\end{aligned}$$

where the first equation matches the definition of momentum, and the second represents the well-known Newton's law

$$F = m\ddot{x}.$$

14.4.2 Quantization of Angular Momentum

Classical Mechanics

The system under consideration consists of a material particle moving on a circle. The generalized coordinates describing the particle are the angle θ and the radius of the circle r , which is considered constant: $r = \text{const}|_t$ (see [Figure 14.9](#)).

The Hamiltonian of the system is given by

$$\begin{aligned}\mathcal{H} = T + U &= \frac{mv^2}{2} + U(r) = \\ &= \frac{mr^2\dot{\theta}^2}{2} + U(r),\end{aligned}$$

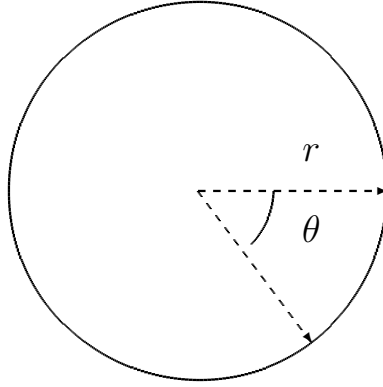


Figure 14.9: Quantization of angular momentum. The position of the particle is determined by two generalized coordinates: r and angle θ .

where $T = \frac{mv^2}{2}$ is the kinetic energy of the particle, and U is the potential energy, which, due to the symmetry of the problem, does not depend on the angle θ and depends only on the distance r .

The Lagrangian of the system is

$$\mathcal{L} = T - U = \frac{mr^2\dot{\theta}^2}{2} - U(r).$$

The angular momentum (generalized momentum corresponding to the generalized coordinate θ) is defined as

$$l = \frac{d\mathcal{L}}{d\dot{\theta}} = mr^2\dot{\theta} = I\dot{\theta}, \quad (14.50)$$

where I denotes $I = mr^2$ — the moment of inertia.

Equations of motion for the coordinate r :

$$\frac{\partial \mathcal{L}}{\partial r} = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{r}},$$

from which

$$\begin{aligned}\frac{\partial \mathcal{L}}{\partial r} &= 0, \\ \frac{\partial T}{\partial r} - \frac{\partial U}{\partial r} &= 0, \\ \frac{\partial U}{\partial r} &= mr\dot{\theta}^2,\end{aligned}$$

or equivalently

$$U = \frac{mr^2\dot{\theta}^2}{2} = \frac{l^2}{2I}.$$

Thus, the Hamiltonian of the system under consideration is

$$\begin{aligned}\mathcal{H} &= \frac{mr^2\dot{\theta}^2}{2} + \frac{l^2}{2I} = \\ &= \frac{l^2}{2I} + \frac{l^2}{2I} = \frac{l^2}{I}\end{aligned}\tag{14.51}$$

Quantization

Let $|\psi\rangle$ be an eigenfunction of the angular momentum operator \hat{L} corresponding to the eigenvalue l :

$$\hat{L}|\psi\rangle = l|\psi\rangle.\tag{14.52}$$

This same wave function must satisfy the Schrödinger equation

$$i\hbar\frac{\partial|\psi\rangle}{\partial t} = \hat{\mathcal{H}}|\psi\rangle.$$

From (14.51), we have

$$\hat{\mathcal{H}} = \frac{\hat{L}^2}{I},$$

and taking into account (14.52),

$$i\hbar\frac{\partial|\psi\rangle}{\partial t} = \frac{1}{I}\hat{L}\hat{L}|\psi\rangle = \frac{l^2}{I}|\psi\rangle,$$

or equivalently

$$\frac{\partial|\psi\rangle}{\partial t} = \frac{-il^2}{\hbar I}|\psi\rangle.$$

Thus,

$$|\psi(t)\rangle = C \cdot \exp\left\{\frac{-il^2}{\hbar I}t\right\}.$$

The wave function must satisfy the periodicity condition

$$|\psi(t)\rangle = |\psi(t+T)\rangle,$$

where T is the period of oscillations. It can be found from (14.50):

$$\begin{aligned}\dot{\theta} &= \frac{l}{I}, \\ \theta &= \theta_0 + \frac{l \cdot t}{I}, \\ 2\pi &= \frac{l \cdot T}{I}, \\ T &= \frac{2\pi I}{l}.\end{aligned}$$

Thus,

$$2i\pi n = \frac{-il^2}{\hbar I} T = \frac{2\pi I}{l} \frac{-il^2}{\hbar I},$$

from which

$$l = -\hbar n.$$

14.5 Composite Systems

Systems consisting of multiple parts behave fundamentally differently for the case of classical and quantum systems.

As an example, consider a system of two particles. The position of the first one is described by 3 coordinates, which represent a point in some linear space L_1 : $(x_1, y_1, z_1) \in L_1$. For the second, we assume we have 2 coordinates in space L_2 that fully determine the location: $(x_2, y_2) \in L_2$. Obviously, to fully describe the position of the two particles we need 5 coordinates: $(x_1, y_1, z_1, x_2, y_2) \in L_1 \times L_2$. Thus, composite systems, in the classical case, are described by points in space $L^{classical}$, which is the Cartesian product of the originals: $L = L_1 \times L_2$. A distinguishing feature of the Cartesian product is that the dimensions of the corresponding spaces add:

$$\dim L^{classical} = \dim L_1 + \dim L_2.$$

In the quantum case, a composite system is described by vectors (points) in space that is the tensor product of the originals:

$$L^{quantum} = L_1 \otimes L_2.$$

In this space, there are 6 basis vectors, and accordingly, to describe the state of the system, 6 numbers are required: $(x_1 \cdot x_2, y_1 \cdot x_2, z_1 \cdot x_2, x_1 \cdot y_2, y_1 \cdot y_2, z_1 \cdot y_2) \in L^{quantum}$. Accordingly, the dimensions multiply:

$$\dim L^{quantum} = \dim L_1 \cdot \dim L_2.$$

Thus, if we have two independent quantum systems with wave functions $|\psi_1\rangle$ and $|\psi_2\rangle$, then the composite system will have the following wave function

$$|\psi_{12}\rangle = |\psi_1\rangle \otimes |\psi_2\rangle,$$

and it is obvious that the following equalities hold

$$\begin{aligned} |\psi_1\rangle &= Sp_2 |\psi_{12}\rangle, \\ |\psi_2\rangle &= Sp_1 |\psi_{12}\rangle \end{aligned} \tag{14.53}$$

i.e., to obtain the state of subsystem 1, one must take the trace over the states of system 2 from the overall wave function.

Composite systems have an interesting property where, despite the parts of this system being mixed states, the entire system as a whole is pure. As an example, consider the so-called entangled states :

$$|\psi\rangle = \frac{|0_1\rangle |1_2\rangle - |1_1\rangle |0_2\rangle}{\sqrt{2}}$$

which is pure. At the same time, for the first particle, we have

$$\begin{aligned} \rho_1 &= Sp_2 |\psi\rangle \langle \psi| = \\ &= \langle 0_2 | \psi \rangle \langle \psi | 0_2 \rangle + \langle 1_2 | \psi \rangle \langle \psi | 1_2 \rangle = \\ &= \frac{|1_1\rangle \langle 1_1| + |0_1\rangle \langle 0_1|}{2} \end{aligned}$$

i.e., the state of the first particle is mixed.

14.6 Coordinate representation of a coherent state

Let's start with the vacuum state, which satisfies the equation

$$\hat{a} |0\rangle = 0. \tag{14.54}$$

In the coordinate representation $\hat{q} = q$, $\hat{p} = -i\frac{\partial}{\partial q}$, then

$$\begin{aligned} \hat{a} &= \frac{1}{\sqrt{2\hbar\omega}} \left(\omega q + \hbar \frac{\partial}{\partial q} \right), \\ \hat{a}^\dagger &= \frac{1}{\sqrt{2\hbar\omega}} \left(\omega q - \hbar \frac{\partial}{\partial q} \right). \end{aligned} \tag{14.55}$$

Substituting (14.55) into (14.54), we obtain the equation for the wave function of the vacuum state of the mode in the coherent representation:

$$\left(\omega q + \hbar \frac{\partial}{\partial q}\right) \Phi_0(q) = 0, \quad (14.56)$$

where $\Phi_0(q) = \langle q|0\rangle$. The normalized solution of this equation is

$$\Phi_0(q) = \left(\frac{\omega}{\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{\omega q^2}{2\hbar}}.$$

The uncertainty of this solution $\Delta q \Delta p = \frac{\hbar}{2}$ is minimal. The probability density to detect a certain coordinate q (in our case a certain electric field strength) is:

$$\Phi_0(q) \Phi_0^*(q) = \left(\frac{\omega}{\pi\hbar}\right)^{\frac{1}{2}} e^{-\frac{\omega q^2}{\hbar}}. \quad (14.57)$$

For the coherent state, the following equation holds

$$\hat{a} |\alpha\rangle = \alpha |\alpha\rangle,$$

which in the coordinate representation has the form

$$\left(\omega q + \hbar \frac{\partial}{\partial q}\right) \Phi_\alpha(q) = \alpha \Phi_\alpha(q), \quad (14.58)$$

where

$$\Phi_\alpha(q) = \langle q | \alpha \rangle.$$

Comparing (14.56) and (14.58), it is easy to see that the solution of equation (14.58) differs from (14.57) by a shift of $\alpha \sqrt{\frac{2\hbar}{\omega}}$. Indeed, equation (14.58) can be rewritten as

$$\left(\omega \left(q - \frac{\alpha}{\omega}\right) + \hbar \frac{\partial}{\partial q}\right) \Phi_\alpha(q) = 0.$$

The solution of this equation is:

$$\Phi_\alpha(q) = \left(\frac{\omega}{\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{\omega}{2\hbar} \left(q - \sqrt{\frac{2\hbar}{\omega}} \alpha\right)^2}. \quad (14.59)$$

The coherent state is not a stationary state. It must satisfy the non-stationary Schrödinger equation. The parameter $\alpha = \alpha(t)$ depends on time. Substituting (14.59) into the Schrödinger equation, this dependence can be determined, but even without that it is clear that $\alpha(t)$ must oscillate harmonically at the mode frequency. We assume

$$\alpha(t) = |\alpha| e^{-i\omega t}.$$

Then the probability density takes the form

$$\Phi_{\alpha}(q) \Phi_{\alpha}^{*}(q) = N e^{-\frac{\omega}{2\hbar} \left[\left(q - \sqrt{\frac{2\hbar}{\omega}} \alpha \right)^2 + \left(q - \sqrt{\frac{2\hbar}{\omega}} \alpha^{*} \right)^2 \right]}.$$

Let's transform the exponent's argument

$$\begin{aligned} & \left[\left(q - \sqrt{\frac{2\hbar}{\omega}} \alpha \right)^2 + \left(q - \sqrt{\frac{2\hbar}{\omega}} \alpha^{*} \right)^2 \right] = \\ &= q^2 + \frac{2\hbar}{\omega} \alpha^2 - 2q\alpha \sqrt{\frac{2\hbar}{\omega}} + q^2 + \frac{2\hbar}{\omega} \alpha^{*2} - 2q\alpha^{*} \sqrt{\frac{2\hbar}{\omega}} = \\ &= 2 \left[q^2 + \frac{2\hbar}{\omega} \frac{\alpha^2 + \alpha^{*2}}{2} - 2q \sqrt{\frac{2\hbar}{\omega}} \frac{\alpha + \alpha^{*}}{2} + \frac{2\hbar}{\omega} (\alpha\alpha^{*}) - \frac{2\hbar}{\omega} (\alpha\alpha^{*}) \right] = \\ &= 2 \left[q^2 + \frac{4\hbar}{\omega} \left(\frac{\alpha + \alpha^{*}}{2} \right)^2 - 2q \sqrt{\frac{2\hbar}{\omega}} \frac{\alpha + \alpha^{*}}{2} - \frac{2\hbar}{\omega} (\alpha\alpha^{*}) \right] = \\ &= 2 \left[q - \sqrt{\frac{2\hbar}{\omega}} \frac{\alpha + \alpha^{*}}{2} \right]^2 + const = \\ &= 2 \left[q - \sqrt{\frac{2\hbar}{\omega}} |\alpha| \cos \omega t \right]^2 + const. \end{aligned}$$

The constant will go into the normalization factor, so we obtain:

$$\Phi_{\alpha}(q, t) \Phi_{\alpha}^{*}(q, t) = N e^{-\frac{\omega}{\hbar} \left[q - \sqrt{\frac{2\hbar}{\omega}} |\alpha| \cos \omega t \right]^2}. \quad (14.60)$$

From expression (14.60) it is seen that the probability density for q (in our case for the electric field strength) at a fixed time is the same as for the vacuum state, but shifted by $\sqrt{\frac{2\hbar}{\omega}} \alpha(t)$. Over time, the center of the distribution moves harmonically, as depicted in Figure 14.10

Since the width of the distribution does not depend on the amplitude, the relative uncertainty decreases with increasing amplitude. Therefore, the state will tend to become classical.

14.7 Adiabatic Theorem

Theorem 14.7.1 (Adiabatic Theorem). *A physical system remains in its instantaneous eigenstate if the perturbation acts slowly enough and if this state is separated from the rest of the Hamiltonian spectrum by an energy gap.*

Proof. TBD □

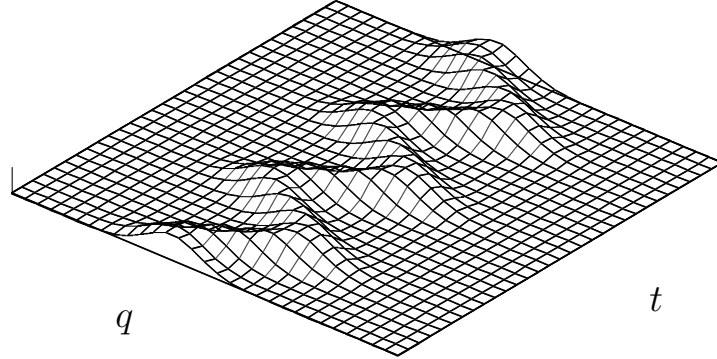


Figure 14.10: Evolution of the distribution function of the coherent state $|\Phi_\alpha(q, t)|^2$ over time.

14.8 Bosons and Fermions

Consider an ensemble of identical particles. The state of a system consisting of n identical particles can be written in the following form

$$|\psi\rangle = |x_1, x_2, \dots, x_k, \dots, x_j, \dots, x_n\rangle. \quad (14.61)$$

For each of the particles in expression (14.61) we are interested in some physical characteristic (for example, coordinate) which is denoted by x .

Obviously, from the classical point of view, there should be no difference if we swap two particles. At the same time, the wave function, which is defined up to a phase, may change. Thus, denoting by $\hat{S}_{k,j}$ the operator that swaps two particles with indices k and j , we get

$$\begin{aligned} \hat{S}_{k,j} |\psi\rangle &= |x_1, x_2, \dots, x_j, \dots, x_k, \dots, x_n\rangle = \\ &= e^{i\phi} |x_1, x_2, \dots, x_k, \dots, x_j, \dots, x_n\rangle = e^{i\phi} |\psi\rangle \end{aligned}$$

If the operator $\hat{S}_{k,j}$ is applied twice, it is obvious that the system will be brought completely back to the original state, i.e.,

$$\hat{S}_{k,j}^2 |\psi\rangle = e^{2i\phi} |\psi\rangle = |\psi\rangle$$

from which it follows that ϕ can take two values 0 and π . Thus, we can divide all particles into two classes. For the first of them the wave function does not change when swapping two particles ($\psi = 0$), and for the second it changes sign ($\psi = \pi$).

Example 14.8.1. Consider a system consisting of two particles. Suppose the first particle can be in one of two states $|\psi_1^{(1)}\rangle$ or $|\psi_2^{(1)}\rangle$. Similarly, the second particle can be in one of two states $|\psi_1^{(2)}\rangle$ or $|\psi_2^{(2)}\rangle$.

Assume the total wave function $|\psi^{(12)}\rangle$ of the two particles can be represented as a superposition of the single-particle wave functions, i.e.

$$|\psi^{(12)}\rangle = |\psi_{1,2}^{(1)}\rangle \pm |\psi_{1,2}^{(2)}\rangle$$

The case of a symmetric (under particle exchange) wave function will have the form

$$|\psi^{(12)}\rangle = |\psi_{1,2}^{(1)}\rangle + |\psi_{1,2}^{(2)}\rangle$$

and both particles can be in the same state, i.e., have the same wave function

$$|\psi\rangle = |\psi_1^{(1)}\rangle = |\psi_1^{(2)}\rangle. \quad (14.62)$$

In this case

$$|\psi^{(12)}\rangle = 2|\psi\rangle,$$

i.e., the particles can simultaneously be in the same state.

For the antisymmetric wave function case we have

$$|\psi^{(12)}\rangle = |\psi_{1,2}^{(1)}\rangle - |\psi_{1,2}^{(2)}\rangle$$

and if both particles are in the same state (14.62), then for $\psi^{(12)}$ we get

$$|\psi^{(12)}\rangle = |\psi\rangle - |\psi\rangle = 0,$$

i.e., in this case particles cannot be in the same state simultaneously.

Bearing in mind Example 14.8.1, we can consider the general case of an ensemble of identical particles, each of which is in the same state or equivalently has the same wave function. In this case, under particle exchange, the actual wave function of the ensemble does not change:

$$|\psi^{(12)}\rangle = |\psi^{(21)}\rangle.$$

That is, in the antisymmetric case

$$|\psi^{(12)}\rangle = -|\psi^{(21)}\rangle = -|\psi^{(12)}\rangle = 0.$$

Thus, we can say that particles whose ensemble states are symmetric can be simultaneously in the same state. Such particles are called bosons.

For the antisymmetric case, we assume that such particles cannot simultaneously be in the same state. Such particles are called fermions.

Consider a state with a definite number of particles n $|\psi\rangle = |n\rangle$. In the case of bosons we will have

$$|\psi_b\rangle = |n\rangle_b. \quad (14.63)$$

The creation operators \hat{a}_b^\dagger and annihilation operators \hat{a}_b act on the state (14.63) as follows:

$$\begin{aligned} \hat{a}_b^\dagger |n\rangle_b &= \sqrt{n+1} |n+1\rangle_b, \\ \hat{a}_b |n\rangle_b &= \sqrt{n} |n-1\rangle_b \end{aligned}$$

For fermions there are two possible states: $|0\rangle_f$ and $|1\rangle_f$, with the corresponding creation operators \hat{a}_f^\dagger and annihilation operators \hat{a}_f acting as follows:

$$\begin{aligned} \hat{a}_f^\dagger |0\rangle_f &= |1\rangle_f, \\ \hat{a}_f^\dagger |1\rangle_f &= 0, \\ \hat{a}_f |0\rangle_f &= 0, \\ \hat{a}_f |1\rangle_f &= |0\rangle_f. \end{aligned}$$

14.9 Fermi's Golden Rule

Fermi's golden rule allows one to use time-dependent perturbation theory to calculate the transition probability between two states of a quantum system.

Theorem 14.9.1 (Fermi's Golden Rule). *Suppose that a quantum system is initially in the state $|i\rangle$ and we want to estimate the probability of transition to the state $|f\rangle$. The Hamiltonian of the considered system consists of two parts: a stationary Hamiltonian \hat{H}_0 and a weak perturbation \hat{V} :*

$$\hat{H} = \hat{H}_0 + \hat{V}$$

In this case, the desired probability is given by

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} \left| \langle f | \hat{V} | i \rangle \right|^2 \rho,$$

where ρ is the density of final states.

Proof. Suppose there exists a set of states $\{|n\rangle\}$ forming a complete set of eigenvectors of the Hamiltonian \hat{H}_0 , i.e.,

$$\hat{H}_0 |n\rangle = E_n |n\rangle,$$

while the considered initial state $|i\rangle$ and final state $|f\rangle$ belong to this vector system, i.e.,

$$|i\rangle, |f\rangle \in \{|n\rangle\}.$$

At some moment in time, the system is in the state $|\psi\rangle$, which satisfies the Schrödinger equation:

$$i\hbar \frac{\partial |\psi\rangle}{\partial t} = \hat{H} |\psi\rangle. \quad (14.64)$$

Our task is to calculate the probability that upon observation the system will be found in the state $|f\rangle$, given that the initial condition for (14.64) is

$$|\psi\rangle|_{t=0} = |i\rangle. \quad (14.65)$$

Due to the completeness of the system $\{|n\rangle\}$, the state $|\psi\rangle$ can be represented in the form

$$|\psi\rangle = \sum_n a_n(t) |n\rangle e^{\frac{-iE_n t}{\hbar}} \quad (14.66)$$

Substituting (14.66) into the Schrödinger equation (14.64) yields

$$i\hbar \sum_n \frac{\partial a_n(t)}{\partial t} |n\rangle e^{\frac{-iE_n t}{\hbar}} + \sum_n a_n(t) |n\rangle i\hbar \frac{\partial e^{\frac{-iE_n t}{\hbar}}}{\partial t} = \quad (14.67)$$

$$\begin{aligned} &= i\hbar \sum_n e^{\frac{-iE_n t}{\hbar}} |n\rangle \left(\frac{\partial a_n(t)}{\partial t} - \frac{iE_n}{\hbar} a_n(t) \right) = \\ &= i\hbar \sum_n e^{\frac{-iE_n t}{\hbar}} \frac{\partial a_n(t)}{\partial t} |n\rangle + \sum_n E_n a_n(t) |n\rangle e^{\frac{-iE_n t}{\hbar}} = \\ &= i\hbar \sum_n e^{\frac{-iE_n t}{\hbar}} \frac{\partial a_n(t)}{\partial t} |n\rangle + \sum_n E_n a_n(t) |n\rangle e^{\frac{-iE_n t}{\hbar}}. \end{aligned} \quad (14.68)$$

On the other hand, expression (14.68) must equal

$$\hat{H} |\psi\rangle = \sum_n E_n a_n(t) |n\rangle e^{\frac{-iE_n t}{\hbar}} + \sum_n a_n(t) e^{\frac{-iE_n t}{\hbar}} \hat{V} |n\rangle,$$

i.e.,

$$i\hbar \sum_n e^{\frac{-iE_n t}{\hbar}} \frac{\partial a_n(t)}{\partial t} |n\rangle = \sum_n a_n(t) e^{\frac{-iE_n t}{\hbar}} \hat{V} |n\rangle. \quad (14.69)$$

We will solve expression (14.69) by perturbation theory, i.e.,

$$a_n(t) = a_n^{(0)}(t) + a_n^{(1)}(t) + \dots,$$

and suppose that

$$a_n^{(0)}(t) = \text{const}|_t = a_n^{(0)}(0),$$

from which, given the initial condition (14.65),

$$a_n^{(0)}(t) = \delta_{ni}.$$

Thus, for the first order of perturbation theory from (14.69) we get

$$i\hbar \sum_n e^{\frac{-iE_n t}{\hbar}} \frac{\partial a_n^{(1)}(t)}{\partial t} |n\rangle = \sum_n \delta_{ni} e^{\frac{-iE_n t}{\hbar}} \hat{V} |n\rangle = e^{\frac{-iE_i t}{\hbar}} \hat{V} |i\rangle. \quad (14.70)$$

Therefore, from (14.70) we have

$$\frac{\partial a_n^{(1)}(t)}{\partial t} = -\frac{i}{\hbar} e^{\frac{-i(E_i - E_n)t}{\hbar}} \langle n | \hat{V} | i \rangle,$$

denoting $E_i - E_n = \hbar\omega_{in}$, where ω_{in} is the transition frequency between the states $|i\rangle$ and $|n\rangle$, after integrating we get

$$a_n^{(1)}(t) = -\frac{i}{\hbar} \int_0^t e^{-i\omega_{in}t'} \langle n | \hat{V} | i \rangle dt',$$

i.e.,

$$\begin{aligned} a_n^{(1)}(t) &= \frac{1}{\hbar\omega_{in}} \langle n | \hat{V} | i \rangle (e^{-i\omega_{in}t} - 1) = \\ &= \frac{1}{\hbar\omega_{in}} \langle n | \hat{V} | i \rangle e^{\frac{-i\omega_{in}t}{2}} \left(e^{\frac{-i\omega_{in}t}{2}} - e^{\frac{i\omega_{in}t}{2}} \right) = \\ &= -\frac{2i}{\hbar\omega_{in}} \langle n | \hat{V} | i \rangle e^{\frac{-i\omega_{in}t}{2}} \sin\left(\frac{\omega_{in}t}{2}\right) \end{aligned} \quad (14.71)$$

We are interested in the rate $W_{i \rightarrow f}$ of transition to the state $|f\rangle$, which is given by the expression

$$W_{i \rightarrow f} = \frac{|a_f^{(1)}(t)|^2}{t}.$$

Or

$$W_{i \rightarrow f} = \frac{t}{\hbar^2} \left| \langle f | \hat{V} | i \rangle \right|^2 \frac{\sin^2\left(\frac{\omega_{if}t}{2}\right)}{\left(\frac{\omega_{if}t}{2}\right)^2}. \quad (14.72)$$

In the case where the final state $|f\rangle$ represents some continuous spectrum near the energy E_f , the expression (14.72) must be summed over all final states for which $\omega_{nf} \approx 0$. We introduce the concept of the density of states $\rho(E) = \frac{dE}{dn}$, so that the number of states dn corresponding to an infinitesimal energy interval dE can be represented as $dn = \rho dE$. Therefore, the transition rate can be rearranged as:

$$W_{i \rightarrow f} = \sum_{n: E_n \approx E_f} \rho(E_n) W_{i \rightarrow f}(E_n),$$

i.e., assuming for the considered energy interval $\rho(E_n) \approx \rho(E_f)$,

$$\left| \langle n | \hat{V} | i \rangle \right|^2 \approx \left| \langle f | \hat{V} | i \rangle \right|^2$$

and replacing summation by integration, we get

$$\begin{aligned} W_{i \rightarrow f} &\approx \int_{E_n \approx E_f} \rho(E_n) dE_n \frac{t}{\hbar^2} \left| \langle n | \hat{V} | i \rangle \right|^2 \frac{\sin^2\left(\frac{\omega_{in}t}{2}\right)}{\left(\frac{\omega_{in}t}{2}\right)^2} = \\ &= \int_{\omega_{in} \approx \omega_{if}} \rho(E_n) dE_n \frac{t}{\hbar^2} \left| \langle n | \hat{V} | i \rangle \right|^2 \frac{\sin^2\left(\frac{\omega_{in}t}{2}\right)}{\left(\frac{\omega_{in}t}{2}\right)^2} \approx \\ &\approx \frac{2}{\hbar} \left| \langle f | \hat{V} | i \rangle \right|^2 \rho \int_{-\infty}^{\infty} d\left(\frac{\omega t}{2}\right) \frac{\sin^2\left(\frac{\omega t}{2}\right)}{\left(\frac{\omega t}{2}\right)^2} = \frac{2\pi}{\hbar} \left| \langle f | \hat{V} | i \rangle \right|^2 \rho, \end{aligned} \quad (14.73)$$

which coincides with the expression to be proved. □

Chapter 15

Probability Theory

When talking about quantum mechanics, it is very often considered an applied part of classical probability theory. Such an assumption is only partially correct, as the probabilities underlying effects such as entangled states and non-classical light have a fundamentally non-classical nature, i.e., they are not described by classical Kolmogorov probability theory. The material in this section mainly relies on the monographs [35] and [37], which compare classical and quantum probability theory.

15.1 Set algebra and basic concepts of probability theory

Definition 15.1.1 (σ -algebra). Suppose that Ω is some set, then a family \mathcal{F} of subsets of the set Ω is called a σ -algebra if the following conditions are satisfied:

- \mathcal{F} contains Ω : $\Omega \in \mathcal{F}$
- if $\Upsilon \in \mathcal{F}$ then $\Omega \setminus \Upsilon \in \mathcal{F}$
- the union or intersection of a countable subfamily of \mathcal{F} belongs to \mathcal{F}

Definition 15.1.2 (Measure). If \mathcal{F} is some σ -algebra then the following function

$$\mu : \mathcal{F} \rightarrow [0, \infty]$$

is called a measure if

- $\mu(\emptyset) = 0$
- $\forall A, B \in \mathcal{F}, A \cap B = \emptyset$ we have $\mu(A \cup B) = \mu(A) + \mu(B)$

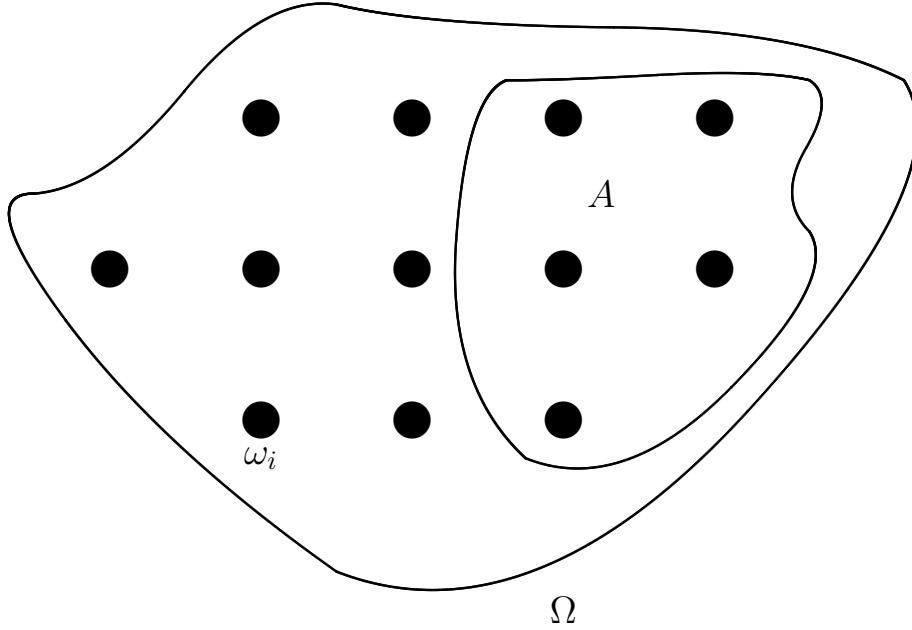


Figure 15.1: The event space Ω is a set consisting of 12 elements (elementary events) $\omega_i \in \Omega$. A subset $A \subset \Omega$ (event) contains 5 elements (elementary events) of the set Ω

15.2 Classical Kolmogorov Probability Theory

15.2.1 Basic Concepts and Axioms of Probability Theory

Classical probability theory deals with set theory and is based on several simple axioms. This axiomatics was proposed in the 1930s by Kolmogorov A. N. [33].

First of all, a few definitions of the objects with which we will be dealing.

Definition 15.2.1 (Event Space). \mathcal{F} is some σ -algebra defined on some set Ω and is called an event space.

Remark 15.2.1 (Event Space). *By virtue of the properties of a σ -algebra, we have the following facts: if $\Omega \subset \Omega$, $\emptyset \subset \Omega$ and for any $A, B \subset \Omega$ it holds that $A \cup B \subset \Omega$, $A \cap B \subset \Omega$ and $A \setminus B \subset \Omega$. An element of the set $\omega \in \Omega$ is called an elementary event. A subset $A \subset \Omega$ or $A \in \mathcal{F}$ is called an event.*

Definition 15.2.2 (Probability P). A finite measure defined on \mathcal{F} over the set Ω is called a probability if $P(\Omega) = 1$.

Example 15.2.1 (Event Space). In [Figure 15.1](#), the set Ω is depicted, which we call the event space. Elements of the set $\omega_i \in \Omega$ (black dots in [Figure 15.1](#)) are elementary events. A subset $A \subset \Omega$ is an event.

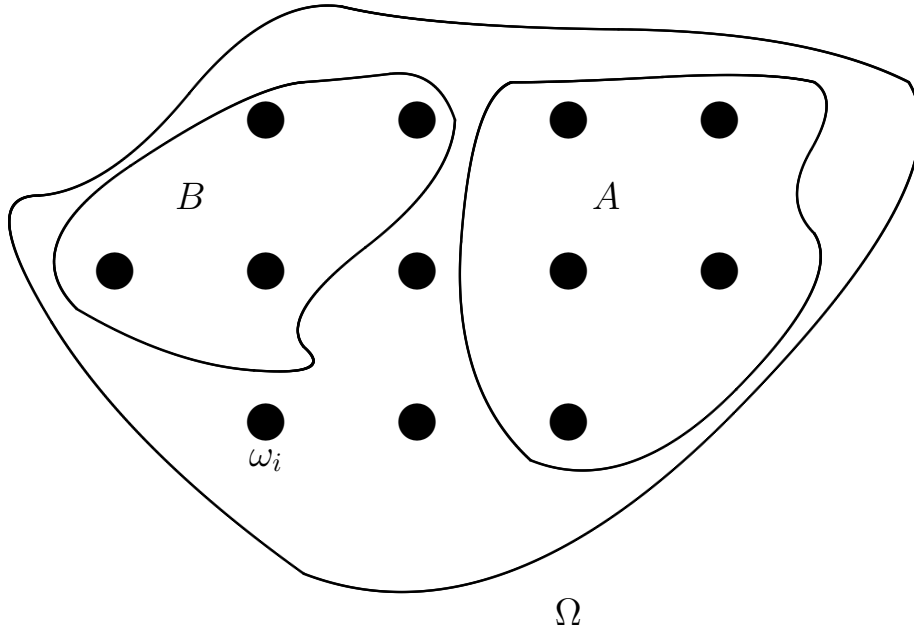


Figure 15.2: Event space Ω , elementary event $\omega \in \Omega$, events $A \subset \Omega$ and $B \subset \Omega$. $P(\omega) = \frac{1}{12}$, $P(\Omega) = \frac{12}{12} = 1$, $P(A) = \frac{5}{12}$, $P(B) = \frac{4}{12}$, $P(A \cup B) = P(A) + P(B) = \frac{5}{12} + \frac{4}{12} = \frac{3}{4}$

Now, strictly speaking, the axioms.

Axiom 15.2.1 (Non-negativity). *The probability of an event $A \subset \Omega$ is a non-negative real number, i.e. $P(A) \geq 0$*

Axiom 15.2.2 (Normalization). *The probability of the event space Ω is 1, i.e. $P(\Omega) = 1$*

Axiom 15.2.3 (Additivity). *If $A_i \cap A_j = \emptyset$, then $P(A_i \cup A_j) = P(A_i) + P(A_j)$*

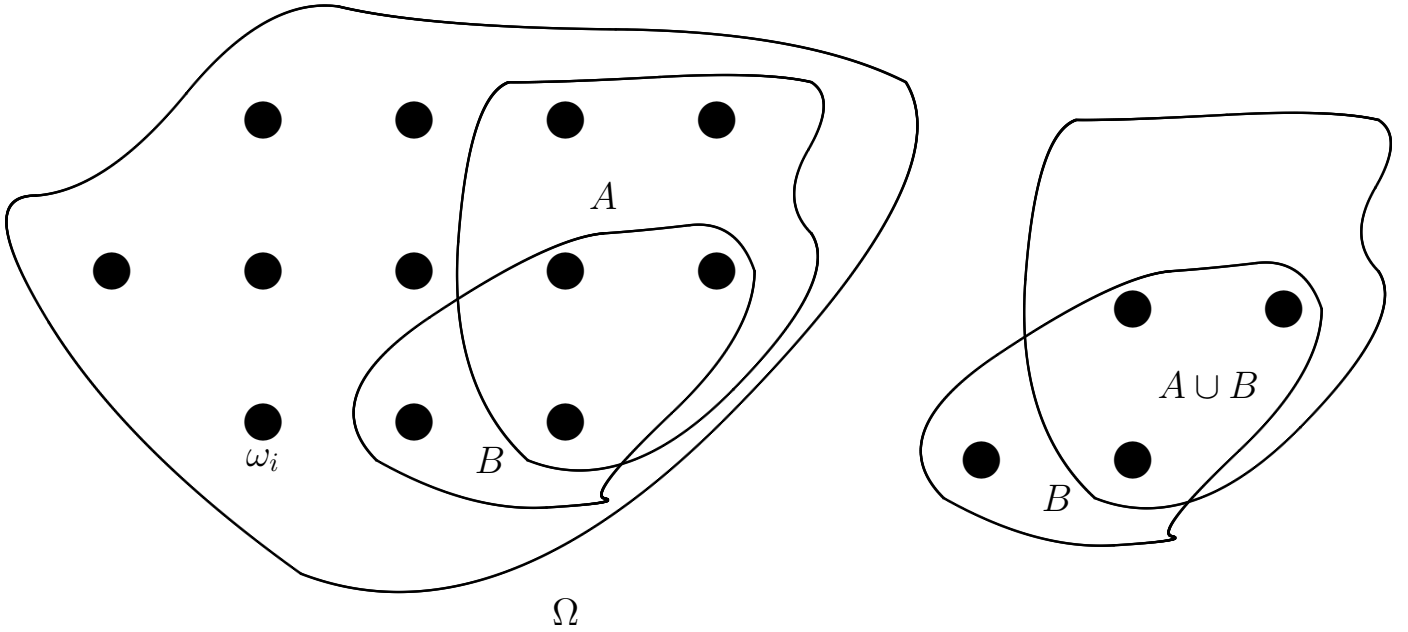
Example 15.2.2 (Axioms of Kolmogorov's Probability Theory). *Assign to each elementary event $\omega_i \in \Omega$ in Figure 15.2 a non-negative number $P(\omega_i) = \frac{1}{12}$. The probability of event A is $P(A) = \frac{5}{12}$ and for event B we have $P(B) = \frac{4}{12}$. Thus, since $A \cap B = \emptyset$, using axiom 15.2.3, we get:*

$$P(A \cup B) = P(A) + P(B) = \frac{5}{12} + \frac{4}{12} = \frac{3}{4}.$$

For the event Ω , by axiom 15.2.2 we have $P(\Omega) = 1$. On the other hand, $\Omega = \cup_i \omega_i$, i.e.

$$P(\Omega) = \sum_i P(\omega_i) = 12 \cdot \frac{1}{12} = 1.$$

All known facts of probability theory are derived from these three axioms.
TBD

Figure 15.3: Conditional probability $P(A|B)$

15.2.2 Random Variable. Distribution Functions

Definition 15.2.3 (Random Variable). If Ω is an event space (see definition 15.2.1), then the mapping $X : \Omega \rightarrow \mathbb{R}$ is called a random variable.

It is worth noting that if a random variable is denoted with a capital letter X , then a specific realization is denoted with a lowercase letter x , i.e., if $\omega \in \Omega$, then $x = X(\omega)$. In particular, for discrete random variables, each realization $x = X(\omega)$ can be associated with some probability $P(x) = P(\omega)$. Otherwise, distribution functions are used.

Definition 15.2.4 (Distribution Function of a Random Variable). The distribution function F_X of a random variable X is a function $F_X : \mathbb{R} \rightarrow [0, 1]$, which is given by the formula

$$F_X(x) = P(X \leq x),$$

i.e., the distribution function at argument x is equal to the probability of those events $\omega \in \Omega$ for which $X(\omega) \leq x$.

Definition 15.2.5 (Density of Distribution of a Random Variable). In the case of continuous random variables, it makes sense to speak of the density of distribution $f_X(x)$, which is defined by the following expression

$$f_X(x) = \frac{dF_X(x)}{dx}$$

Definition 15.2.6 (Expectation of a Random Variable). For a discrete random variable, i.e., if the set of elementary events is countable, the expectation $E(X)$ of a random variable X is defined by the following relation

$$E(X) = \sum_{\omega} p_{\omega} \cdot X(\omega) = \sum_x p_x \cdot x,$$

where the summation is over all possible elementary outcomes ω , $x = X(\omega)$, $p_{\omega} = p_{X(\omega)} = p_x$. For continuous random variables, this relation transforms into

$$E(X) = \int_{-\infty}^{\infty} x \cdot f_X(x) dx.$$

TBD

15.2.3 Poisson Distribution

The Poisson distribution plays a very important role in quantum optics and therefore deserves a more detailed consideration.

Definition 15.2.7 (Poisson Distribution).

TBD

15.2.4 Random Processes

To fully describe a random process $x(t)$, it is necessary to know all the joint probability density functions

$$p_n(\{x\}_n, \{t\}_n) = p_n(x_1, x_2, \dots, x_n, t_1, t_2, \dots, t_n).$$

For example, using $p_1 = p_1(x, t)$, one can calculate the mean value of the random variable

$$\langle x(t) \rangle = \int x p_1(x, t) dx, \quad (15.1)$$

and using $p_2 = p_2(x_1, x_2, t_1, t_2)$ — the so-called two-time autocorrelation function

$$r(t_1, t_2) = \langle x^*(t_1) x(t_2) \rangle = \int \int x_1 x_2 p_2(x_2, t_2, x_1, t_1) dx_1 dx_2. \quad (15.2)$$

In the case where the mean of the random variable (15.1) does not depend on t :

$$\langle x(t) \rangle = \text{const}_t, \quad (15.3)$$

and the autocorrelation function (15.2) depends only on the difference $t_1 - t_2 = \tau$:

$$r(t_1, t_2) = r(t_1 - t_2) = r(\tau), \quad (15.4)$$

then the random process is called wide-sense stationary. Obviously, for such processes, the following identity is valid:

$$r(-\tau) = r^*(\tau), \quad (15.5)$$

indeed, from (15.4) we have

$$r(-\tau) = r(t_2 - t_1) = r(t_2, t_1) = r^*(t_1, t_2) = r^*(\tau).$$

For wide-sense stationary random processes, the following theorem holds:

Theorem 15.2.1 (Khinchin-Wiener). *The autocorrelation function of wide-sense stationary random processes and their spectral density are linked by the direct and inverse Fourier transforms.*

Proof. The proof will be informal and serves as an attempt to explain the meaning of this theorem. Consider the following integral

$$\tilde{x}(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} x(t) e^{i\omega t} dt, \quad (15.6)$$

which can be interpreted as the Fourier transform (see section 18.1) of the random process $x(t)$. For stationary random processes, it is obvious that the integral (15.6) generally does not exist. Indeed, if this integral existed, the Parseval relation would hold:

$$\int_{-\infty}^{\infty} |\tilde{x}(\omega)|^2 d\omega = \int_{-\infty}^{\infty} |x(t)|^2 dt, \quad (15.7)$$

which obviously diverges in the case of a stationary random process. At the same time, by virtue of (15.7), $S(\omega) = |\tilde{x}(\omega)|^2$ represents the distribution of energy along the frequency axis for a particular realization of the process, and the averaged value $\langle S(\omega) \rangle$ represents a certain characteristic of the random process that describes the energy distribution over frequencies.

Although the integral (15.7) diverges for the considered random processes, as shown by Khinchin and Wiener, the integral

$$\langle \tilde{x}^*(\omega) \tilde{x}(\omega') \rangle = \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \langle x^*(t) x(t') \rangle e^{i(\omega' t' - \omega t)} dt dt', \quad (15.8)$$

exists and can be interpreted as the power spectral density $S(\omega, \omega')$ of the considered random process. Thus, taking into account stationarity,

$$\langle x^*(t) x(t') \rangle = r(t' - t) = r(\tau),$$

we obtain

$$\begin{aligned}
 S(\omega, \omega') &= \langle \tilde{x}^*(\omega) \tilde{x}(\omega') \rangle = \\
 &= \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \langle x^*(t) x(t') \rangle e^{i(\omega't' - \omega t)} dt dt' = \\
 &= \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} r(\tau) e^{i(\omega' - \omega)t} e^{i\omega'(t' - t)} dt d\tau = \\
 &= \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} e^{i(\omega' - \omega)t} dt \int_{-\infty}^{\infty} r(\tau) e^{i\omega'\tau} d\tau = \\
 &= \tilde{r}(\omega') \delta(\omega' - \omega), \tag{15.9}
 \end{aligned}$$

where the integral representation of the delta function (18.3) has been used.

Thus, from (15.9) it follows that the Fourier transform of the autocorrelation function $\tilde{r}(\omega)$ can be regarded as the spectral density of the random process (15.8). \square

15.3 Quantum Probability Theory

When speaking about quantum probability theory, we will rely on the classical (Kolmogorov) theory, while trying to choose a formulation that is applicable in both cases. As a model, we will consider a system with n possible outcomes, i.e. one where the classical probability distribution is given by the following distribution:

$$p = \{p_1, p_2, \dots, p_n\} \tag{15.10}$$

A quantum state is fully described in the general case (both pure and mixed states) by the density matrix

$$\hat{\rho} = \begin{bmatrix} \rho_{11} & \rho_{12} & \cdots & \rho_{1n} \\ \rho_{21} & \rho_{22} & \cdots & \rho_{2n} \\ \vdots & \vdots & \ddots & \vdots \\ \rho_{n1} & \rho_{n2} & \cdots & \rho_{nn} \end{bmatrix}.$$

In the classical case, this object corresponds to the probability distribution (15.10), which can also be represented as a certain matrix

$$\hat{p} = \begin{bmatrix} p_1 & 0 & \cdots & 0 \\ 0 & p_2 & \cdots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \cdots & p_n \end{bmatrix}.$$

It is worth noting that for both cases the equality holds

$$Sp(\hat{\rho}) = Sp(\hat{p}) = 1,$$

which is a rewritten version of axiom 15.2.2.

An observable \hat{x} corresponds to a certain matrix

$$\hat{x} = \begin{bmatrix} x_{11} & x_{12} & \cdots & x_{1n} \\ x_{21} & x_{22} & \cdots & x_{2n} \\ \vdots & \vdots & \ddots & \vdots \\ x_{n1} & x_{n2} & \cdots & x_{nn} \end{bmatrix}, \quad (15.11)$$

which in the classical case is diagonal

$$\hat{x} = \begin{bmatrix} x_1 & 0 & \cdots & 0 \\ 0 & x_2 & \cdots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \cdots & x_n \end{bmatrix}. \quad (15.12)$$

It is also worth noting that the calculation of the mean then takes the familiar classical form

$$\bar{x} = Sp(\hat{p}\hat{x}) = \sum_{i=1}^n p_i x_i.$$

Remark 15.3.1. *It is worth noting that for any observable \hat{X} , due to self-adjointness $\hat{X} = \hat{X}^\dagger$, one can choose a basis in which the corresponding matrix will be diagonal, with eigenvalues on the diagonal. Moreover, if two observables \hat{X} and \hat{Y} commute:*

$$\hat{X}\hat{Y} = \hat{Y}\hat{X},$$

then they can be diagonalized in a common basis.

The difference between quantum probability theory and classical theory can be shown in several ways [35, 37], one of which is based on the noncommutativity of the quantum description. Indeed, it can be noted that in the general case quantum observables, represented by matrices of the form (15.11), do not commute with each other. At the same time, classical observables, written as diagonal matrices (15.12) always commute with each other. From this, many facts follow TBD

Chapter 16

Abstract Algebra

Here you can find definitions and basic properties of groups, rings, and fields.

16.1 Introduction to Group Theory

Definition 16.1.1. A group (\mathcal{G}, \circ) is called a set of elements $g \in \mathcal{G}$ for which a certain binary operation \circ (often called multiplication or addition) is defined:

$$\begin{aligned} \forall g_1, g_2 \in \mathcal{G}, \\ g_1 \circ g_2 \in \mathcal{G}. \end{aligned} \tag{16.1}$$

The operation defined by (16.1) satisfies the property of associativity:

$$g_1 \circ (g_2 \circ g_3) = (g_1 \circ g_2) \circ g_3.$$

The considered set must contain an element $e_{\mathcal{G}}$ that has the following property, valid for any element of the set g :

$$g \circ e_{\mathcal{G}} = e_{\mathcal{G}} \circ g = g.$$

For each element of the group g , there must exist an inverse element $g^{-1} \in \mathcal{G}$, possessing the following property:

$$g \circ g^{-1} = g^{-1} \circ g = e_{\mathcal{G}}$$

Definition 16.1.2 (Monoid). If for some set of elements \mathcal{G} the last property of groups (existence of inverse element) is not fulfilled, then this set is called a monoid or a semigroup.

Definition 16.1.3 (Abelian Group). A group (\mathcal{A}, \circ) is called abelian, or commutative, if $\forall a_1, a_2 \in \mathcal{A}$: $a_1 \circ a_2 = a_2 \circ a_1$.

Example 16.1.1. Group $(\mathbb{Z}, +)$ *The set of integers $\mathbb{Z} = \{0, \pm 1, \pm 2, \dots\}$ forms a group under the operation of addition.*

Definition 16.1.4 (Cyclic Group). A cyclic group G is a group generated by a single element $g : G = \langle g \rangle$, i.e. all its elements are powers of g . The element g is called the generating element or the generator of the group G .

Definition 16.1.5 (Multiplicative Group of the Ring of Residues). Consider the set of integers coprime with n and less than n , denoted by $(\mathbb{Z}/n\mathbb{Z})^\times$. As the operation of multiplication of two elements $a, b \in (\mathbb{Z}/n\mathbb{Z})^\times$ we take

$$a \circ b = a \cdot b \pmod{n}.$$

The identity element $e_{(\mathbb{Z}/n\mathbb{Z})^\times}$ is 1. Moreover, it can be shown that for each $a \in (\mathbb{Z}/n\mathbb{Z})^\times$ there exists $a^{-1} \in (\mathbb{Z}/n\mathbb{Z})^\times$ such that $a \circ a^{-1} = 1$. Thus, $(\mathbb{Z}/n\mathbb{Z})^\times$ is a group.

Theorem 16.1.1 (On the order of $(\mathbb{Z}/n\mathbb{Z})^\times$). *The order of the group $(\mathbb{Z}/n\mathbb{Z})^\times$ is given by the formula*

$$|(\mathbb{Z}/n\mathbb{Z})^\times| = \phi(n),$$

where $\phi(n)$ is the Euler's totient function (19.3.1). Moreover, if $n = p$ is a prime number, then

$$|(\mathbb{Z}/p\mathbb{Z})^\times| = \phi(p),$$

and if $a \in (\mathbb{Z}/p\mathbb{Z})^\times, a \neq 1$, then a is a generator of the considered group, i.e.,

$$a^{p-1} = 1.$$

Proof. TBD □

Theorem 16.1.2 (Lagrange). *For any finite group \mathcal{G} , the order (number of elements) of any subgroup \mathcal{H} divides the order of \mathcal{G} :*

$$|\mathcal{G}| = h |\mathcal{H}|,$$

where the integer h is called the index of the subgroup.

16.2 Fields

Definition 16.2.1 (Field (algebra)). Let there be an Abelian group (see definition 16.1.3) $(\mathcal{F}, +)$. The unit element of this group $e_{\mathcal{F}}$ is 0. Let also $(\mathcal{F} \setminus \{0\}, \cdot)$ be some

other group (also Abelian) with the unit element 1. Moreover, the operations $+$, \cdot satisfy the distributive property, i.e. $\forall a, b, c \in \mathcal{F}$:

$$\begin{aligned}c \cdot (a + b) &= c \cdot a + c \cdot b, \\(a + b) \cdot c &= a \cdot c + b \cdot c.\end{aligned}$$

In this case, $(\mathcal{F}, +, \cdot)$ is called a field.

Example 16.2.1 (Field \mathbb{Q}). *Note that \mathbb{Z} is not a field because not every integer has an inverse with respect to the multiplication operation. However, the following set will be a field: $\mathbb{Q} = \{a/b \mid a \in \mathbb{Z}, b \in \mathbb{Z} \setminus \{0\}\}$. The inverse with respect to $a/b \in (\mathbb{Q} \setminus \{0\}, \cdot)$ will be b/a .*

Example 16.2.2 (Field \mathbb{R}). *Real numbers form a field.*

Example 16.2.3 (Field \mathbb{C}). *Complex numbers form a field.*

Chapter 17

Linear Algebra

17.1 Linear Space

Definition 17.1.1 (Linear Space). TBD

17.1.1 Direct Sum of Vector Spaces and Tensor Product

Suppose we have two vector spaces \mathbb{A} and \mathbb{B} with bases $\{a_i\} \subset \mathbb{A}$ and $\{b_j\} \subset \mathbb{B}$ respectively.

Definition 17.1.2 (Tensor Product of Vector Spaces). A vector space $\mathbb{L} = \mathbb{A} \otimes \mathbb{B}$ with basis $l_{ij} = a_i \otimes b_j$ is called the tensor product of the vector spaces \mathbb{A} and \mathbb{B} .

Definition 17.1.3 (Tensor (External) Product of Linear Operators). Consider two linear operators $\hat{A} : \mathbb{A} \rightarrow \mathbb{A}$ and $\hat{B} : \mathbb{B} \rightarrow \mathbb{B}$ then the operator $\hat{L} = \hat{A} \otimes \hat{B} : \mathbb{L} \rightarrow \mathbb{L}$ acting on the basis by the rule

$$\hat{L}l_{ij} = \hat{L}a_i \otimes b_j = \hat{A}a_i \otimes \hat{B}b_j$$

is called the tensor product of the operators.

17.1.2 Theorem on Operator Decomposition

Theorem 17.1.1 (On Operator Identity).

$$e^{\hat{A}}\hat{B}e^{-\hat{A}} = \hat{B} + [\hat{A}, \hat{B}] + \frac{1}{2!} [\hat{A}, [\hat{A}, \hat{B}]] + \dots \quad (17.1)$$

Proof. To prove this, we introduce the following auxiliary function:

$$f(\alpha) = e^{\alpha\hat{A}}\hat{B}e^{-\alpha\hat{A}}, \quad (17.2)$$

which coincides with the desired expression at $\alpha = 1$. Expand (17.2) in a Taylor series:

$$f(\alpha) = f(0) + \frac{\alpha}{1!}f'(0) + \frac{\alpha^2}{2!}f''(0) + \cdots + \frac{\alpha^k}{k!}f^{(k)}(0) + \dots \quad (17.3)$$

The following recurrence relation holds:

$$\begin{aligned} f^{(k)}(\alpha) &= [\hat{A}, f^{(k-1)}(\alpha)] = \\ &= \underbrace{[\hat{A}[\dots[\hat{A}, f^{(0)}(\alpha)]\dots]]}_{k \text{ times}}, \end{aligned} \quad (17.4)$$

where

$$f^{(0)}(\alpha) = f(\alpha).$$

Indeed, by induction we have for the first derivative:

$$\begin{aligned} f^{(1)}(\alpha) &= \hat{A}e^{\alpha\hat{A}}\hat{B}e^{-\alpha\hat{A}} - e^{\alpha\hat{A}}\hat{B}\hat{A}e^{-\alpha\hat{A}} = \\ &= \hat{A}e^{\alpha\hat{A}}\hat{B}e^{-\alpha\hat{A}} - e^{\alpha\hat{A}}\hat{B}e^{-\alpha\hat{A}}\hat{A} = \\ &= [\hat{A}, e^{\alpha\hat{A}}\hat{B}e^{-\alpha\hat{A}}] = [\hat{A}, f(\alpha)] = [\hat{A}, f^{(0)}(\alpha)], \end{aligned} \quad (17.5)$$

i.e. (17.4) is valid for $k = 1$. In deriving (17.5), we considered the fact that

$$\hat{A}e^{-\alpha\hat{A}} = e^{-\alpha\hat{A}}\hat{A}.$$

For the $k + 1$ derivative we obtain:

$$\begin{aligned} f^{(k+1)}(\alpha) &= \frac{\partial f^{(k)}(\alpha)}{\partial \alpha} = \\ &= \frac{\partial}{\partial \alpha} \left(\underbrace{[\hat{A}[\dots[\hat{A}, f^{(0)}(\alpha)]\dots]]}_k \right) = \underbrace{[\hat{A}[\dots[\hat{A}, \frac{\partial}{\partial \alpha} f^{(0)}(\alpha)]\dots]]}_k = \\ &= \underbrace{[\hat{A}[\dots[\hat{A}, f^{(1)}(\alpha)]\dots]]}_k = \underbrace{[\hat{A}[\dots[\hat{A}, [\hat{A}, f^{(0)}(\alpha)]]\dots]]}_k = \\ &= \underbrace{[\hat{A}[\dots[\hat{A}, f^{(0)}(\alpha)]\dots]]}_{k+1}, \end{aligned}$$

which along with (17.5) proves the validity of (17.4).

Considering that

$$f^{(0)}(\alpha) = \hat{B},$$

one can obtain from (17.3) the desired identity (17.1) at $\alpha = 1$. □

17.2 Hilbert Space

Hilbert space is a generalization of Euclidean space allowing infinite dimensions.

Definition 17.2.1 (Hilbert Space). TBD

17.2.1 Stone's Theorem

Theorem 17.2.1 (Stone). *TBD*

Chapter 18

Fourier Transform

18.1 Fourier Transform

In this course, we will use the following form of the Fourier transform ¹

$$\tilde{x}(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} x(t) e^{i\omega t} dt, \quad (18.1)$$

while the inverse Fourier transform is as follows

$$x(t) = \int_{-\infty}^{\infty} \tilde{x}(\omega) e^{-i\omega t} d\omega, \quad (18.2)$$

The following table contains important transform formulas

| Function $x(t)$ | Transform $\tilde{x}(\omega)$ |
|-----------------|-------------------------------|
| 1 | $\delta(\omega)$ |
| $\delta(t)$ | $\frac{1}{2\pi}$ |

Table 18.1: Fourier Transform

From [Table 18.1](#), an important formula follows (integral representation of the delta function) [\[1\]](#)

$$\delta(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{i\omega t} dt \quad (18.3)$$

¹Most commonly, the following definition of the Fourier transform is used

$$\tilde{x}(\omega) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} x(t) e^{-i\omega t} dt.$$

However, according to [\[2\]](#), there are alternative definitions differing in the sign before the frequency and the coefficient before the integral.

18.2 Discrete Fourier Transform

The Fourier transform plays an important role in digital signal processing, particularly for the analysis of periodic sequences.

18.2.1 Definition

Definition 18.2.1. Suppose there are M samples x_0, x_1, \dots, x_{M-1} then the discrete Fourier transform is given by the relation

$$\tilde{X}_k = \frac{1}{\sqrt{M}} \sum_{m=0}^{M-1} x_m e^{-\frac{2\pi i}{M} k \cdot m}, \quad (18.4)$$

which is also written as

$$\{x_m\} \longleftrightarrow \{\tilde{X}_k\}.$$

The inverse Fourier transform can be obtained by the analogous relation

$$x_k = \frac{1}{\sqrt{M}} \sum_{m=0}^{M-1} \tilde{X}_m e^{\frac{2\pi i}{M} k \cdot m},$$

Figure [Figure 18.1](#) shows the graph of some periodic function and its Fourier transform.

Expression (18.4) can also be rewritten in the matrix form

$$\vec{\tilde{X}} = \hat{F} \vec{x},$$

where

$$\vec{x} = \begin{pmatrix} x_0 \\ x_1 \\ \vdots \\ x_{M-1} \end{pmatrix}, \quad \vec{\tilde{X}} = \begin{pmatrix} \tilde{X}_0 \\ \tilde{X}_1 \\ \vdots \\ \tilde{X}_{M-1} \end{pmatrix},$$

and the matrix \hat{F} has the form

$$\hat{F} = \frac{1}{\sqrt{M}} \begin{pmatrix} 1 & 1 & 1 & \dots & 1 \\ 1 & e^{-i\omega} & e^{-2i\omega} & \dots & e^{-(M-1)i\omega} \\ 1 & e^{-2i\omega} & e^{-4i\omega} & \dots & e^{-2(M-1)i\omega} \\ 1 & e^{-3i\omega} & e^{-6i\omega} & \dots & e^{-3(M-1)i\omega} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 1 & e^{-(M-1)i\omega} & e^{-2(M-1)i\omega} & \dots & e^{-(M-1)(M-1)i\omega} \end{pmatrix}, \quad (18.5)$$

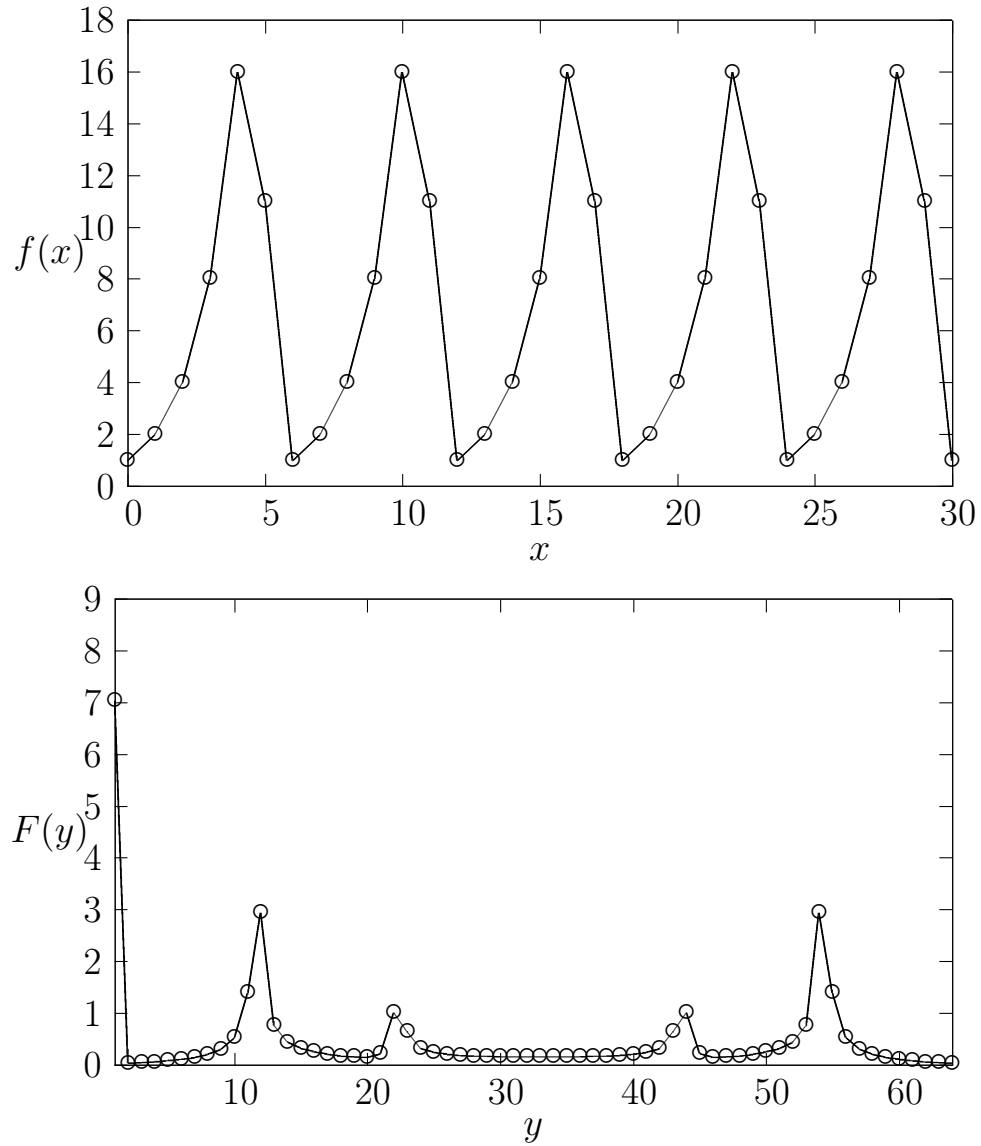


Figure 18.1: Periodic function $f(x, a) = a^x \bmod N$ for $a = 2$, $N = 21$ (upper graph) and its discrete Fourier transform (lower graph). The period of the original function $r = 6$. Number of samples $M = 64$, showing local maxima with period $T \approx \frac{M}{r} \approx 10.67$

where

$$\omega = \frac{2\pi}{M}.$$

For a matrix element of the matrix (18.5), one can write

$$F_{nm} = \frac{1}{\sqrt{M}} e^{-i\omega nm}, \quad (18.6)$$

where $n, m \in \{0, 1, \dots, M-1\}$.

18.2.2 Properties of the Discrete Fourier Transform

The following properties of the Fourier transform should be noted, which play a fundamental role in Shor's algorithm:

Lemma 18.2.1. (Shift) *If $\{x_n\} \longleftrightarrow \{\tilde{X}_k\}$ then $\{x_{(n-m) \bmod M}\} \longleftrightarrow \{e^{-i\omega mk} \tilde{X}_k\}$*

Proof. For the sequence $\{x_{(n-m) \bmod M}\}$ one can write

$$\{x_{(n-m) \bmod M}\} = \begin{pmatrix} x_{-m \bmod M} \\ x_{-m+1 \bmod M} \\ \vdots \\ x_{-1 \bmod M} \\ x_0 \\ x_1 \\ \vdots \\ x_{M-m-1} \end{pmatrix} = \begin{pmatrix} x_{M-m} \\ x_{M-m+1} \\ \vdots \\ x_{M-1} \\ x_0 \\ x_1 \\ \vdots \\ x_{M-m-1} \end{pmatrix},$$

and

$$\begin{aligned}
& \hat{F} \{x_{(n-m) \bmod M}\} = \\
& = \frac{1}{\sqrt{M}} \begin{pmatrix} 1 & 1 & 1 & \dots & 1 \\ 1 & e^{-i\omega} & e^{-2i\omega} & \dots & e^{-(M-1)i\omega} \\ 1 & e^{-2i\omega} & e^{-4i\omega} & \dots & e^{-2(M-1)i\omega} \\ 1 & e^{-3i\omega} & e^{-6i\omega} & \dots & e^{-3(M-1)i\omega} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 1 & e^{-(M-1)i\omega} & e^{-2(M-1)i\omega} & \dots & e^{-(M-1)(M-1)i\omega} \end{pmatrix} \begin{pmatrix} x_{M-m} \\ x_{M-m+1} \\ \vdots \\ x_0 \\ \vdots \\ x_{M-m-1} \end{pmatrix} = \\
& = \begin{pmatrix} \frac{x_{M-m}}{\sqrt{M}} + \frac{x_{M-m+1}}{\sqrt{M}} + \dots + \frac{x_0}{\sqrt{M}} + \dots + \frac{x_{M-m-1}}{\sqrt{M}} \\ \frac{x_{M-m}}{\sqrt{M}} + \frac{e^{-i\omega} x_{M-m+1}}{\sqrt{M}} + \dots + \frac{e^{-i\omega m} x_0}{\sqrt{M}} + \dots + \frac{e^{-i\omega(M-1)} x_{M-m-1}}{\sqrt{M}} \\ \frac{x_{M-m}}{\sqrt{M}} + \frac{e^{-2i\omega} x_{M-m+1}}{\sqrt{M}} + \dots + \frac{e^{-2i\omega m} x_0}{\sqrt{M}} + \dots + \frac{e^{-2i\omega(M-1)} x_{M-m-1}}{\sqrt{M}} \\ \vdots \\ \frac{x_{M-m}}{\sqrt{M}} + \frac{e^{-mi\omega} x_{M-m+1}}{\sqrt{M}} + \dots + \frac{e^{-mi\omega m} x_0}{\sqrt{M}} + \dots + \frac{e^{-mi\omega(M-1)} x_{M-m-1}}{\sqrt{M}} \\ \vdots \end{pmatrix} \quad (18.7)
\end{aligned}$$

Taking into account the relation

$$e^{-i\omega k M} = 1, k \in \{0, 1, \dots\},$$

the expression (18.7) can be rewritten in the following form

$$\begin{aligned}
& \hat{F} \{x_{(n-m) \bmod M}\} = \\
& = \frac{1}{\sqrt{M}} \begin{pmatrix} x_{M-m} + \dots + x_{M-m-1} \\ e^{-i\omega m} e^{-i\omega(M-m)} x_{M-m} + \dots + e^{-i\omega 2m} e^{-i\omega(M-m-1)} \\ e^{-i\omega 2m} e^{-i\omega(M-m)} x_{M-m} + \dots + e^{-i\omega 2m} e^{-i\omega(M-1)} \\ \vdots \end{pmatrix} = \\
& = \begin{pmatrix} \tilde{X}_0 \\ e^{-i\omega m} \tilde{X}_1 \\ e^{-i\omega 2m} \tilde{X}_2 \\ \vdots \end{pmatrix}.
\end{aligned}$$

□

Lemma 18.2.2. (Periodicity) *If the sequence $\{x_n\}$ has period r : $x_n = x_{n+r}$, and the number of samples M is divisible by r , then the nonzero terms of the Fourier transform appear with period $\frac{M}{r}$.*

Proof. Indeed, if $M \bmod r = 0$ and $kr \bmod M \neq 0$, then

$$1 - e^{-i\omega kr} \neq 0,$$

whence

$$\begin{aligned} \tilde{X}_k &= \frac{1}{\sqrt{M}} \sum_{n=0}^{M-1} e^{-i\omega kn} x_n = \\ &= \frac{1}{\sqrt{M}} \left(\sum_{n=0}^{r-1} e^{-i\omega kn} x_n + \sum_{n=0}^{r-1} e^{-i\omega k(n+r)} x_{n+r} + \right. \\ &\quad \left. + \sum_{n=0}^{r-1} e^{-i\omega k(n+2r)} x_{n+2r} + \dots \right) = \\ &= \frac{1}{\sqrt{M}} \left(\sum_{n=0}^{r-1} e^{-i\omega kn} x_n + \sum_{n=0}^{r-1} e^{-i\omega k(n+r)} x_n + \sum_{n=0}^{r-1} e^{-i\omega k(n+2r)} x_n + \dots \right) = \\ &= \frac{1}{\sqrt{M}} \left(\sum_{n=0}^{r-1} x_n e^{-i\omega kn} \sum_{l=0}^{\frac{M}{r}-1} e^{-i\omega klr} = \sum_{n=0}^{r-1} x_n e^{-i\omega kn} \frac{1 - e^{-i\omega k \frac{M}{r} r}}{1 - e^{-i\omega kr}} \right) = \\ &= \frac{1}{\sqrt{M}} \frac{1 - e^{-i\omega kM}}{1 - e^{-i\omega kr}} \sum_{n=0}^{r-1} x_n e^{-i\omega kn} = 0. \quad (18.8) \end{aligned}$$

If $M \bmod r = 0$ and $kr \bmod M = 0$, then

$$e^{-i\omega kr} = e^{-i\frac{2\pi}{M}kr} = 1,$$

whence

$$\begin{aligned} \tilde{X}_k &= \frac{1}{\sqrt{M}} \sum_{n=0}^{M-1} e^{-i\omega kn} x_n = \\ &= \frac{1}{\sqrt{M}} \left(\sum_{n=0}^{r-1} e^{-i\omega kn} x_n + \sum_{n=0}^{r-1} e^{-i\omega kn} x_{n+r} + \sum_{n=0}^{r-1} e^{-i\omega kn} x_{n+2r} + \dots \right) = \\ &= \frac{1}{\sqrt{M}} \left(\sum_{n=0}^{r-1} e^{-i\omega kn} x_n + \sum_{n=0}^{r-1} e^{-i\omega kn} x_n + \sum_{n=0}^{r-1} e^{-i\omega kn} x_n + \dots \right) = \\ &= \frac{1}{\sqrt{M}} \frac{M}{r} \sum_{n=0}^{r-1} e^{-i\omega kn} x_n \neq 0. \quad (18.9) \end{aligned}$$

Thus, from expressions (18.8) and (18.9) it follows that the nonzero terms appear with a period $T = \frac{M}{r}$. \square

Remark 18.2.1. (Periodicity of maxima) *It is worth noting that expression (18.8) (in the case when the period is not a divisor of the number of samples: $M \bmod r \neq 0$) will be approximately equal to 0 for those values of k that are far from multiples of $\frac{M}{r}$, i.e. local maxima of the Fourier transform will repeat with period $\frac{M}{r}$.*

18.2.3 Fast Fourier Transform

Calculations using formula (18.4) have complexity of order $O(M^2)$, where M is the number of elements (samples) ².

There exists an algorithm for fast calculation using formula (18.4) that has complexity $O(M \log M)$.

Using the "divide and conquer" paradigm (see section 20.2) one can pay attention to the form of the expression (18.6) and notice that the expression (18.4) can be rewritten as

$$\tilde{X}_k = \sum_{m=0}^{M-1} F_{km}^M x_m,$$

where the notation F_{km}^M indicates that the matrix (18.6) of size $M \times M$ is used. If M is even, then

$$\tilde{X}_k = \sum_{m=0}^{M-1} F_{k,m}^M x_m = \sum_{m=0}^{\frac{M}{2}-1} F_{k,2m}^M x_{2m} + \sum_{m=0}^{\frac{M}{2}-1} F_{k,2m+1}^M x_{2m+1},$$

where

$$\begin{aligned} F_{k,2m}^M &= e^{-i\omega k 2m} = e^{-ikm \frac{2\pi}{M}} = F_{k,m}^{\frac{M}{2}}, \\ F_{k,2m+1}^M &= e^{-i\omega k (2m+1)} = e^{-i\omega k} e^{-ikm \frac{2\pi}{M}} = e^{-2\pi i \frac{k}{M}} F_{k,m}^{\frac{M}{2}}, \end{aligned}$$

i.e.

$$\tilde{X}_k = \sum_{m=0}^{\frac{M}{2}-1} F_{k,m}^{\frac{M}{2}} x_{2m} + \exp\left\{\left(-2\pi i \frac{k}{M}\right)\right\} \sum_{m=0}^{\frac{M}{2}-1} F_{k,m}^{\frac{M}{2}} x_{2m+1}. \quad (18.10)$$

The complexity of computations using formula (18.10) is determined by the following relation

$$T(M) = 2T\left(\frac{M}{2}\right) + O(M). \quad (18.11)$$

²Indeed, it is necessary to obtain M elements, for the computation of each of which requires M multiplication operations

The truth of (18.11) can be verified by noting that the calculations with complexity $T(M)$ in (18.10) split into two subproblems each with complexity $T\left(\frac{M}{2}\right)$.

Using the Master Theorem for recurrence relations (case 2) (Theorem 20.1.1) we obtain

$$T(M) = O(M \log M).$$

18.3 Two-dimensional Fourier Transform

18.3.1 Definition

Definition 18.3.1 (Two-dimensional Fourier Transform). Suppose we have a two-dimensional signal $x(k_1, k_2)$, where $k_1, k_2 \in \{0, \dots, M-1\}$. The two-dimensional Fourier transform is called a two-dimensional function

$$\tilde{X}(j_1, j_2), j_1, j_2 \in \{0, \dots, M-1\}$$

such that

$$\tilde{X}(j_1, j_2) = \frac{1}{M} \sum_{k_1=0}^{M-1} \sum_{k_2=0}^{M-1} x(k_1, k_2) e^{-i\omega(k_1 j_1 + k_2 j_2)},$$

where

$$\omega = \frac{2\pi}{M}.$$

Chapter 19

Discrete Mathematics

19.1 Greatest Common Divisor. Euclidean Algorithm

Definition 19.1.1. The greatest common divisor of numbers a and b ($\text{GCD}(a, b)$) is called the largest of their common divisors.

Theorem 19.1.1 (Theorem on common divisors). *Suppose the following inequalities hold $a > b > 0$, and the number r is the remainder of the division of a by b . Thus, we can write*

$$a = x \cdot b + r, \quad (19.1)$$

where $x \geq 1$, $b > r \geq 0$. If $r = 0$, then b is the largest number that divides both a and b without remainder. In the case when $r > 0$, then

$$\text{GCD}(a, b) = \text{GCD}(b, r). \quad (19.2)$$

Proof. To prove (19.2), we show that any divisor of the pair of numbers (a, b) is a divisor of the pair of numbers (b, r) . Let d be a common divisor of numbers a and b , i.e. $a = d \cdot x_1$, $b = d \cdot x_2$. Thus, from (19.1) it follows

$$r = a - x \cdot b = d \cdot (x_1 - x \cdot x_2),$$

i.e. d is a divisor of the number r .

Now we prove that any common divisor of numbers b and r is also a common divisor of numbers a and b : indeed, let d be a common divisor of numbers b and r , i.e. $b = y_1 \cdot d$ and $r = y_2 \cdot d$, thus (19.1) can be rewritten as

$$a = x \cdot y_1 \cdot d + y_2 \cdot d = d \cdot (x \cdot y_1 + y_2),$$

i.e. d is a divisor of the number a .

Thus, the pairs of numbers (a, b) and (b, r) have common divisors, including the greatest divisor for one pair is also such for the other. \square

Relation (19.2) leads to the following algorithm for computing the greatest common divisor

Algorithm 3 Euclidean Algorithm

```

 $a > b$ 
if  $b = 0$  then
    return  $a$ 
end if
 $a \Leftarrow 0$ 
 $r \Leftarrow b$ 
 $b \Leftarrow a$ 
repeat
     $a \Leftarrow b$ 
     $b \Leftarrow r$ 
     $r \Leftarrow$  remainder of division of  $a$  by  $b$ 
until  $(r \neq 0)$ 
return  $b$ 

```

To estimate the complexity of algorithm 3, we write it in the following form

$$\begin{aligned}
 f_k &= x_k \cdot f_{k+1} + f_{k+2}, \\
 f_0 &= a, \quad f_1 = b, \\
 x_k &\geq 1, \quad f_k > f_{k+1} > f_{k+2},
 \end{aligned}$$

thus $f_k > 2 \cdot f_{k+2}$, or else $f_0 > 2 \cdot f_2 > \dots > 2^n f_{2n}$ i.e. the algorithm stops at $n = \log_2(f_0) = \log_2(a)$ ¹. The number of steps of the algorithm is obviously equal to $2n$ or $2 \cdot \log_2(a)$. Thus the algorithmic complexity of the Euclidean algorithm can be written as $O(\log(a))$.

Example 19.1.1. (GCD (2345, 1456))

$$\begin{aligned}
 2345 &= 1456 + 889, \\
 1456 &= 889 + 567, \\
 889 &= 567 + 322, \\
 567 &= 322 + 245, \\
 322 &= 245 + 77, \\
 245 &= 3 \cdot 77 + 14, \\
 77 &= 5 \cdot 14 + 7, \\
 14 &= 2 \cdot 7.
 \end{aligned}$$

¹The following derivations are made assuming that $\log_2(a)$ is an integer

thus $\text{GCD}(2345, 1456) = 7$. The number of algorithm steps is $8 < 2 \cdot \log_2 2345 \approx 2 \cdot 11.2 = 22.4$.

19.1.1 Bezout's Identity

Theorem 19.1.2 (Bezout). *If numbers a and b are coprime, then the equation*

$$ax + by = 1$$

has integer solutions.

Proof. We use the Euclidean algorithm 3 to find $\text{GCD}(a, b)$. Assuming $a > b$, then

$$\begin{aligned} r_1 &= a - bq_0, \\ r_2 &= b - r_1q_1 = b - (a - bq_0)q_1 = b(1 + q_1q_0) - aq_1, \\ r_3 &= r_1 - r_2q_2 = a - bq_0 - (b(1 + q_1q_0) - aq_1)q_2 = \\ &= a(1 + q_1q_2) - b(q_0 + q_2 + q_0q_1q_2), \\ &\quad \dots \\ \text{GCD}(a, b) &= r_n = r_{n-2} - r_{n-1}q_{n-1} = \dots = ax + by, \end{aligned} \tag{19.3}$$

which proves our statement. \square

Remark 19.1.1 (On the complexity of computing Bezout's identity). *Calculations by (19.3) are equivalent to the steps in the algorithm 3. The number of these steps is $O(\log_2(a))$, thus the algorithm described by equations (19.3) is sufficiently efficient and has complexity $O(\log_2(a))$.*

Example 19.1.2 (Bezout's Identity). *Suppose $a = 25, b = 14$. The formulas of the Euclidean algorithm are*

$$\begin{aligned} 11 &= 25 - 14 \cdot 1, \\ 3 &= 14 - 11 \cdot 1, \\ 2 &= 11 - 3 \cdot 3, \\ 1 &= 3 - 2 \cdot 1. \end{aligned}$$

From these formulas we get

$$\begin{aligned} 11 &= 25 - 14 \cdot 1, \\ 3 &= 14 - 25 + 14 = 2 \cdot 14 - 25, \\ 2 &= 11 - 3 \cdot 3 = 25 - 14 - 3 \cdot (2 \cdot 14 - 25) = 4 \cdot 25 - 7 \cdot 14, \\ 1 &= 3 - 2 \cdot 1 = 9 \cdot 14 - 5 \cdot 25. \end{aligned}$$

Thus,

$$25 \cdot (-5) + 14 \cdot 9 = 1.$$

19.2 Modular Arithmetic

Definition 19.2.1. The notation

$$a \equiv b \pmod{c} \quad (19.4)$$

means that a and b have the same remainders when divided by c or that a and b are congruent modulo the natural number c . Here, the number c is called the modulus of the congruence.

Definition 19.4 can also be interpreted as saying that the difference $a - b$ is divisible by c .

Example 19.2.1. Modular Congruence: $30 \equiv 8 \pmod{11}$, because $30 = 2 \cdot 11 + 8$.

Definition 19.2.2 (Negative Element). If $a < n$, then $n - a$ will be called the negative element relative to a and denoted by $-a \pmod{n}$.

Example 19.2.2. Negative Element:

$$-5 \equiv 6 \pmod{11},$$

since $5 < 11$ and $6 = 11 - 5$.

19.2.1 Arithmetic Operations

Lemma 19.2.1 (Addition modulo). If $a_1 \equiv a_2 \pmod{n}$, $b_1 \equiv b_2 \pmod{n}$, then

$$a_1 + b_1 \equiv a_2 + b_2 \pmod{n}$$

Proof. We can write $a_1 = k_1n + r_a$, $a_2 = k_2n + r_a$, $b_1 = l_1n + r_b$, $b_2 = l_2n + r_b$ from which

$$a_1 + b_1 = (k_1 + l_1)n + r_a + r_b \equiv r_a + r_b \pmod{n}$$

and

$$a_2 + b_2 = (k_2 + l_2)n + r_a + r_b \equiv r_a + r_b \pmod{n}$$

hence

$$a_1 + b_1 \equiv a_2 + b_2 \equiv r_a + r_b \pmod{n}$$

□

Lemma 19.2.2 (Multiplication modulo). If $a_1 \equiv a_2 \pmod{n}$, $b_1 \equiv b_2 \pmod{n}$, then

$$a_1 \cdot b_1 \equiv a_2 \cdot b_2 \pmod{n}$$

Proof. If $a_1 \equiv a_2 \pmod n$, $b_1 \equiv b_2 \pmod n$, then

$$a_1 + b_1 \equiv a_2 + b_2 \pmod n$$

We can write $a_1 = k_1n + r_a$, $a_2 = k_2n + r_a$, $b_1 = l_1n + r_b$, $b_2 = l_2n + r_b$ from which

$$a_1 \cdot b_1 = k_1l_1n + l_1nr_a + k_1nr_b + r_ar_b \equiv r_ar_b \pmod n$$

and

$$a_2 \cdot b_2 = k_2l_2n + l_2nr_a + k_2nr_b + r_ar_b \equiv r_ar_b \pmod n$$

hence

$$a_1 \cdot b_1 \equiv a_2 \cdot b_2 \equiv r_ar_b \pmod n$$

□

19.2.2 Solving Equations

In cryptography, one often deals with equations of the form

$$ax \equiv b \pmod n, \tag{19.5}$$

where a, b, n are known integers, and x is an unknown parameter to be determined.

Obviously, if we find an integer a^{-1} such that

$$aa^{-1} \equiv 1 \pmod n,$$

then

$$x \equiv ba^{-1} \pmod n.$$

If $\gcd(a, n) = 1$, then according to Bézout's identity (see Bézout's theorem ([Theorem 19.1.2](#))) there exist integers x, y such that $ax + ny = 1$, i.e.

$$x \equiv a^{-1} \pmod n.$$

Moreover, according to remark [19.1.1](#), a^{-1} and the solution of equation [\(19.5\)](#) can be found quite efficiently.

19.2.3 The Field \mathbb{F}_p

As we have seen, in modular arithmetic one can add, subtract, multiply, and even divide if the modulo is a prime number. Moreover, addition and multiplication are commutative and satisfy the distributive property. Thus, the remainders modulo a prime form a field (see definition [16.2.1](#)) which is called the Galois field and denoted by \mathbb{F}_p .

19.3 Euler's Function

19.3.1 Definition

Definition 19.3.1 (Euler's Function). Euler's function $\phi(n)$ shows how many numbers $k \in \{1, \dots, n-1\}$ are coprime with n , i.e. $\gcd(k, n) = 1$.

Example 19.3.1 (Euler's Function). *If we take the number $n = 15$, there are 8 numbers coprime with 15: 1, 2, 4, 7, 8, 11, 13, 14. The remaining 7 numbers are not coprime with n because they have a greatest common divisor other than 1, for example $\gcd(6, 15) = 3$. Thus, $\phi(15) = 8$.*

19.3.2 Properties

Property 19.3.1 (Euler's Function of a Prime Number). *If p is a prime number, then $\phi(p) = p - 1$*

Proof. This follows from definition 19.3.1. □

Property 19.3.2 (Euler's Function of a Product (Generalized Multiplicativity)). *If $\gcd(n, m) = 1$, then $\phi(n \cdot m) = \phi(n) \phi(m)$*

Proof. TBD □

Remark 19.3.1 (On the Complexity of Calculating Euler's Function). *Calculating Euler's function for large numbers is a very complex task. It is often computed using property 19.3.1 in combination with 19.3.2. Applying these properties to an arbitrary number requires its factorization, so the complexity of calculating Euler's function is comparable to the complexity of the factorization problem.*

19.4 Fermat's Little Theorem

Theorem 19.4.1 (Fermat's Little Theorem). *If p is a prime number, and a is not divisible by p , then*

$$a^{p-1} \equiv 1 \pmod{p} \tag{19.6}$$

Proof. Consider the following relation

$$a \cdot k_i \pmod{p},$$

where $k_i \in \{1, \dots, p-1\}$.

Obviously,

$$a \cdot k_i \equiv k_j \pmod{p}. \quad (19.7)$$

Indeed,

$$a \cdot k_i \pmod{p} \subset \{1, \dots, p-1\},$$

since any remainder from division by p takes values $0, 1, \dots, p-1$. The zero remainder is impossible since a and p are coprime.

Moreover, each of the remainders $a \cdot k_i \pmod{p}$ occurs only once. Indeed, suppose $a \cdot k_i \pmod{p} = a \cdot k_j \pmod{p}$ or $a \cdot (k_i - k_j) \equiv 0 \pmod{p}$, i.e., a is divisible by p , which contradicts the condition of coprimeness.

Multiplying all expressions from Equation 19.7, we get

$$a \cdot 2a \cdot 3a \cdot \dots \cdot a(p-1) \equiv 1 \cdot 2 \cdot 3 \cdot \dots \cdot (p-1) \pmod{p}.$$

or

$$a^{p-1} (p-1)! \equiv (p-1)! \pmod{p},$$

from which we obtain the required equality due to the coprimeness of p and $(p-1)!$:

$$a^{p-1} \equiv 1 \pmod{p}$$

□

19.4.1 Pseudoprime Numbers

The generalization of Fermat's Little Theorem is not true, i.e., if a and p are coprime numbers that satisfy the relation Equation 19.6, then p may not be prime. For example,

$$2^{341-1} \equiv 1 \pmod{341},$$

given that $341 = 11 \cdot 31$.

Numbers p satisfying the relation Equation 19.6, but not being prime, are called pseudoprime numbers to the base a . For example, 341 is the first pseudoprime number to the base 2.

19.5 Chinese Remainder Theorem

Theorem 19.5.1. *If there are pairwise coprime integers n_1, n_2, \dots, n_k , then for any set of integers a_1, a_2, \dots, a_k , $\exists x$ such that*

$$\begin{aligned} x &\equiv a_1 \pmod{n_1}, \\ x &\equiv a_2 \pmod{n_2}, \\ &\vdots \\ x &\equiv a_k \pmod{n_k}, \end{aligned} \quad (19.8)$$

Moreover, for any x_1, x_2 satisfying this relation, the equality holds

$$x_1 \equiv x_2 \pmod{N},$$

where $N = n_1 \cdot n_2 \cdot \dots \cdot n_k$.

19.6 RSA Algorithm

The RSA algorithm (an acronym for the surnames Rivest, Shamir, and Adleman) is a public-key encryption algorithm ², based on the difficulty of factoring a number into prime factors.

19.6.1 Key Generation

Consists of several steps

- Two prime numbers p and q are chosen.
- The product of the chosen prime numbers is calculated $n = p \cdot q$.
- Euler's function (19.3.1) is calculated (see properties 19.3.1 and 19.3.2)
 $\phi(n) = (p - 1)(q - 1)$
- An integer e is chosen such that $1 < e < \phi(n)$ and e and $\phi(n)$ are coprime, i.e. $\gcd(e, \phi(n)) = 1$.
- Calculate $d \equiv e^{-1} \pmod{\phi(n)}$, i.e. $d \cdot e \equiv 1 \pmod{\phi(n)}$.

The public key consists of two numbers: the modulus n and the public exponent e . These two numbers are used to encrypt the original message.

The private key also consists of two numbers: the modulus n and the private exponent d .

The original numbers p and q are kept secret, since they make computing $\phi(n)$ trivial.

It should be noted that to obtain the private key from the public key, one must compute $\phi(n)$ given n . This is a difficult problem (if p and q are unknown), as noted in remark 19.3.1.

Example 19.6.1. (RSA. Key Generation) Choose two prime numbers $p = 3$ and $q = 7$. The product of these numbers is $n = 21$. Euler's Function (19.3.1)
 $\phi(n) = (p - 1)(q - 1) = 2 \cdot 6 = 12$.

²A public-key (asymmetric) encryption algorithm is one in which two different keys are used: one for encryption and the other for decryption

Choose a number e (public exponent) such that $1 < e < 12$ and $\gcd(e, 12) =$

1. Obviously, $e = 5$ satisfies the stated conditions.

Calculate the private exponent $d \equiv 5^{-1} \pmod{12}$, i.e., $d = 5$. Indeed, $5 \cdot 5 = 25 = 2 \cdot 12 + 1$, i.e., $5 \cdot 5 \equiv 1 \pmod{12}$.

Thus, we obtain

- Public key ($n = 21, e = 5$)
- Private key ($n = 21, d = 5$)

19.6.2 Encryption

Suppose we need to encrypt some message M . Initially, it is converted into an integer(s) m such that $0 < m < \phi(n)$. Then the ciphertext c is computed as:

$$c \equiv m^e \pmod{n} \quad (19.9)$$

Example 19.6.2. (RSA. Encryption) Suppose we have a public key ($n = 21, e = 5$) (see example 19.6.1) and we want to encrypt the following message $m = 1101_2 = 11_{10}$. The ciphertext is computed using formula (19.9) $c \equiv 11^5 \pmod{21} = 2$.

19.6.3 Decryption

m can be recovered from c using the following formula:

$$m \equiv c^d \pmod{n}. \quad (19.10)$$

Having m , the original message M can be restored.

Example 19.6.3. (RSA. Decryption) Suppose we have a private key ($n = 21, d = 5$) (see ex. 19.6.1) and ciphertext $c = 2$ from example 19.6.2.

The original text is calculated using the formula (19.10) $m \equiv 2^5 \pmod{21} = 11 = 1101_2$.

19.6.4 Proof

We want to prove that

$$(m^e)^d \equiv m \pmod{p \cdot q}$$

for any positive number m when p and q are prime numbers, and e and d satisfy the relation

$$d \cdot e \equiv 1 \pmod{\phi(p \cdot q)},$$

which can be rewritten as

$$d \cdot e - 1 = h(p-1)(q-1).$$

Thus,

$$m^{e \cdot d} = m \cdot m^{h(p-1)(q-1)}.$$

Next, two cases are possible: when m is divisible by p , and when m and p are coprime.

In the first case,

$$m^{e \cdot d} \equiv m \equiv 0 \pmod{p}.$$

In the second case, we use Fermat's little theorem ([Theorem 19.4.1](#)) :

$$m \cdot m^{h(p-1)(q-1)} = m (m^{p-1})^{h(q-1)} \equiv m \cdot 1^{h(q-1)} \equiv m \pmod{p}.$$

Similarly, we have either

$$m^{e \cdot d} \equiv m \equiv 0 \pmod{q}$$

or, by Fermat's little theorem,

$$m \cdot m^{h(p-1)(q-1)} = m (m^{q-1})^{h(p-1)} \equiv m \cdot 1^{h(p-1)} \equiv m \pmod{q}.$$

Thus, we have the following two types of congruences: the trivial equalities

$$\begin{aligned} x_1 &= m \equiv m \pmod{p}, \\ x_1 &= m \equiv m \pmod{q}, \end{aligned}$$

and the just obtained congruences

$$\begin{aligned} x_2 &= m^{ed} \equiv m \pmod{p}, \\ x_2 &= m^{ed} \equiv m \pmod{q}, \end{aligned}$$

from which, by the Chinese remainder theorem ([Theorem 19.5.1](#)) , we get

$$x_1 \equiv x_2 \pmod{p \cdot q},$$

i.e.

$$m^{e \cdot d} \equiv m \pmod{n}.$$

19.7 Discrete Logarithm

Definition 19.7.1 (Discrete Logarithm). Consider an Abelian multiplicative group G and the equation

$$g^x = a \quad (19.11)$$

The solution to this equation, i.e., a whole non-negative number x satisfying the equality (19.11) is called the discrete logarithm.

The equation (19.11) generally does not have a solution for all values of a . However, if g is a generating element of G , i.e., $G = \langle g \rangle$, then a solution always exists and it is unique. In further discussion of the discrete logarithm, we will assume that g is chosen such that $G = \langle g \rangle$.

In applied cryptography, one often deals with a special kind of discrete logarithm in the ring of residues modulo p :

Definition 19.7.2 (Discrete Logarithm in the Ring of Residues Modulo p). The discrete logarithm $\text{ind}_g(a) \bmod p$ ³ is called the minimum number x that satisfies the following equation (if such a number exists):

$$g^x \equiv a \bmod p \quad (19.12)$$

Example 19.7.1. ($\text{ind}_3 14 \bmod 17$) We solve the problem by enumeration [28]:

$$\begin{aligned} 3^1 &\equiv 3 \bmod 17, & 3^2 &\equiv 9 \bmod 17, & 3^3 &\equiv 10 \bmod 17, & 3^4 &\equiv 13 \bmod 17, \\ 3^5 &\equiv 5 \bmod 17, & 3^6 &\equiv 15 \bmod 17, & 3^7 &\equiv 11 \bmod 17, & 3^8 &\equiv 16 \bmod 17, \\ 3^9 &\equiv 14 \bmod 17, & 3^{10} &\equiv 8 \bmod 17, & 3^{11} &\equiv 7 \bmod 17, & 3^{12} &\equiv 4 \bmod 17, \\ 3^{13} &\equiv 12 \bmod 17, & 3^{14} &\equiv 2 \bmod 17, & 3^{15} &\equiv 6 \bmod 17, & 3^{16} &\equiv 1 \bmod 17, \end{aligned}$$

thus, it can be seen that $\text{ind}_3 14 \bmod 17 = 9$, since $3^9 \equiv 14 \bmod 17$.

The problem of finding the discrete logarithm is a difficult problem. The fastest known algorithm [10] solves it in time of the order $O\left(c \cdot \exp\left(\log p^{\frac{1}{3}} \log \log p^{\frac{2}{3}}\right)\right)$, where c is some constant, which is why algorithms using discrete logarithms are widely applied in cryptography.

19.8 Diffie-Hellman Protocol (DH)

Suppose there are two parties, Alice and Bob. They are aware of two numbers g and p , which are not secret.

³From the word **index** - an alternative name for the discrete logarithm

Alice chooses a random number a and sends the following value to Bob

$$A \equiv g^a \pmod{p}. \quad (19.13)$$

Bob calculates the following number (using a secret random variable b)

$$B \equiv g^b \pmod{p}. \quad (19.14)$$

Alice, using only the number a known to her, calculates the key

$$K \equiv B^a \pmod{p} \equiv g^{ab} \pmod{p}. \quad (19.15)$$

Bob can obtain the same key value using his secret number b :

$$K \equiv A^b \pmod{p} \equiv g^{ab} \pmod{p}. \quad (19.16)$$

Thus, Alice and Bob obtain the same key, which can later be used for message transfer using symmetric encryption algorithms (for example, AES).

Example 19.8.1. (Diffie-Hellman) *Initial data (public information): $g = 2$, $p = 23$. Alice chooses a random number $a = 6$ and calculates the number (19.13) $A = 18$ and sends it to Bob. Bob chooses a random number $b = 9$ and, using the formula (19.14), gets $B = 6$ and sends this number to Alice.*

Alice calculates the key $K = 12$ using the formula (19.15). Bob can obtain the same key value $K = 12$ using (19.16)

The eavesdropper Eve knows the numbers g , p , A , and B . To obtain the key K , Eve needs to get one of the secret numbers a or b :

$$a \equiv \text{ind}_g(A) \pmod{p},$$

from which, using (19.15), the desired value K is obtained.

19.9 Elgamal Scheme

One variation of the Diffie-Hellman protocol is the Elgamal scheme. It's important to distinguish between the Elgamal encryption algorithm and the Elgamal digital signature algorithm. The Elgamal digital signature is the basis for the digital signature standards of the USA (DSA) and Russia (GOST R 34.10-94).

Below we will consider the algorithm in encryption mode.

19.9.1 Key Generation

- A prime number p is generated.
- An integer g is chosen.
- A random integer $x : 1 < x < p$ is chosen.
- Compute $y = g^x \mod p$

The public key is the triple p, g, y . The private key is the number x .

Example 19.9.1 (Key Generation (Elgamal)). *Choose $p = 21, g = 10, x = 3$. $y = 10^3 \mod 21 = 13$.*

19.9.2 Encryption

The message M to be encrypted must satisfy the criterion $0 < M < p$.

- A session key is chosen - a random number $k : 1 < k < p - 1$.
- Compute $a = g^k \mod p$
- Compute $b = y^k M \mod p$

The pair of numbers (a, b) is considered the ciphertext.

Example 19.9.2 (Encryption (Elgamal)). *Suppose we want to encrypt $M = 6$. Choose $k = 7$. $a = 10^7 \mod 21 = 10$, $b = 13^7 \cdot 6 \mod 21 = 15$.*

19.9.3 Decryption

Knowing the private key x , we can recover the original message using

$$M = b \cdot (a^x)^{-1} \mod p, \quad (19.17)$$

due to the fact that

$$(a^x)^{-1} = g^{-kx} \mod p$$

we have

$$b \cdot (a^x)^{-1} = y^k M g^{-kx} = g^{kx} M g^{-kx} = M \mod p.$$

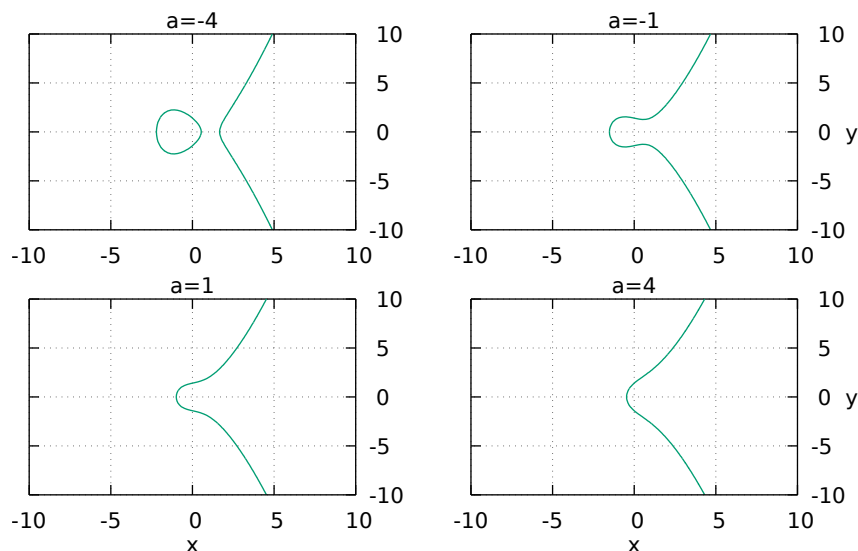


Figure 19.1: Elliptic curve $y^2 = x^3 + ax + 2$ over the field of real numbers \mathbb{R} for various values of the parameter a

Example 19.9.3 (Decryption (Elgamal)). *The encrypted message from ex. 19.9.3 $C = (a = 10, b = 15)$. Using (19.17) we have*

$$M = 15 \cdot 13 \mod 21 = 6$$

where it was used

$$(10^3)^{-1} \equiv 13 \mod 21,$$

because $13 \cdot 10^3 = 1 \mod 21$. Thus we recovered the message $M = 6$ encrypted in ex. 19.9.3.

19.10 Elliptic Cryptography

19.10.1 Elliptic Curves over the Field \mathbb{R}

In elliptic cryptography, certain sets of objects that form a group are considered (see section 16.1). As such a set, we will consider points belonging to a certain curve (see Figure 19.1):

$$E : y^2 = x^3 + ax + b,$$

where the coefficients a, b must satisfy the following relation

$$4a^3 + 27b^2 \neq 0,$$

in this case the cubic equation $x^3 + ax + b = 0$ will have 3 distinct real roots [26].

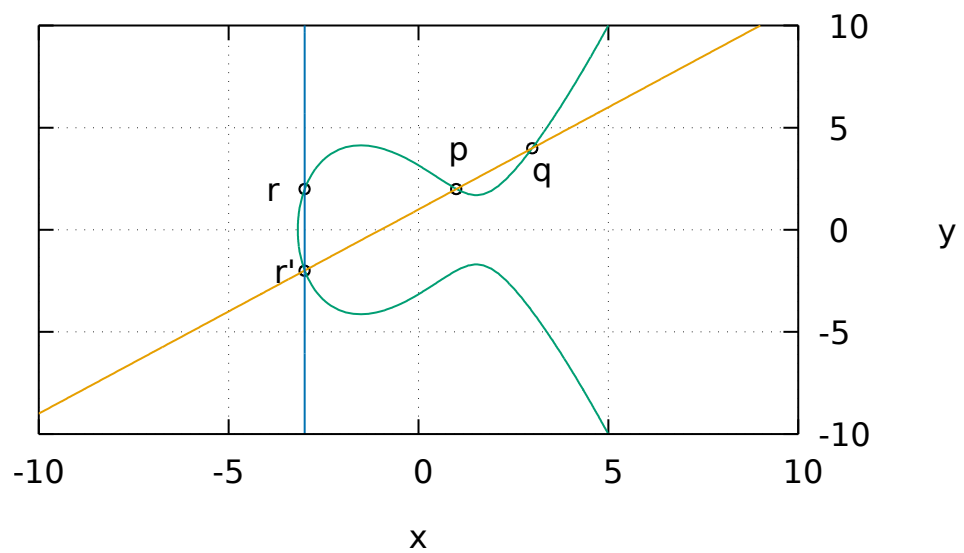


Figure 19.2: Elliptic curve $y^2 = x^3 - 7x + 10$ over the field of real numbers \mathbb{R} . Addition of two points $p(1, 2)$ and $q(3, 4)$. The line passing through these points intersects the curve at a third point $r'(-3, -2)$. The point $r(-3, 2)$, symmetric with respect to the curve, is the sum of the original two: $p + q = r$

According to definition 16.1.1, for points on the curve there is a certain binary operation which assigns a third point r with coordinates (r_x, r_y) to two points p, q with coordinates (p_x, p_y) and (q_x, q_y) , respectively. We will call this operation addition:

$$p + q = r.$$

There is a simple geometric interpretation of the addition operation (see Figure 19.2). Suppose there are 2 points on the curve that we want to add: p and q with coordinates $(x_p, y_p), (x_q, y_q)$ respectively. If $x_p \neq x_q$, then through these points one can draw a straight line with slope

$$m = \frac{y_p - y_q}{x_p - x_q}.$$

and intersects the curve at a point r' . If this point has coordinates $(x_{r'}, y_{r'})$, then since it lies on the line with slope m :

$$m = \frac{y_{r'} - y_p}{x_{r'} - x_p},$$

therefore

$$y_{r'} = y_p + m(x_{r'} - x_p).$$

This point must belong to the curve, i.e.

$$y_{r'}^2 = (y_p + m(x_{r'} - x_p))^2 = x_{r'}^3 + ax_{r'} + b$$

which can be rewritten as

$$x_{r'}^3 - m^2 x_{r'}^2 + \dots = 0.$$

The equation $x^3 - m^2 x^2 + \dots = 0$ has 3 roots: $x_p, x_q, x_{r'}$, i.e., it can also be rewritten as

$$(x - x_{r'})(x - x_p)(x - x_q) = x^3 - (x_{r'} + x_p + x_q)x^2 + \dots = 0.$$

Thus,

$$x_{r'} + x_p + x_q = m^2.$$

Therefore

$$\begin{aligned} x_{r'} &= m^2 - x_p - x_q, \\ y_{r'} &= y_p + m(x_{r'} - x_p), \end{aligned}$$

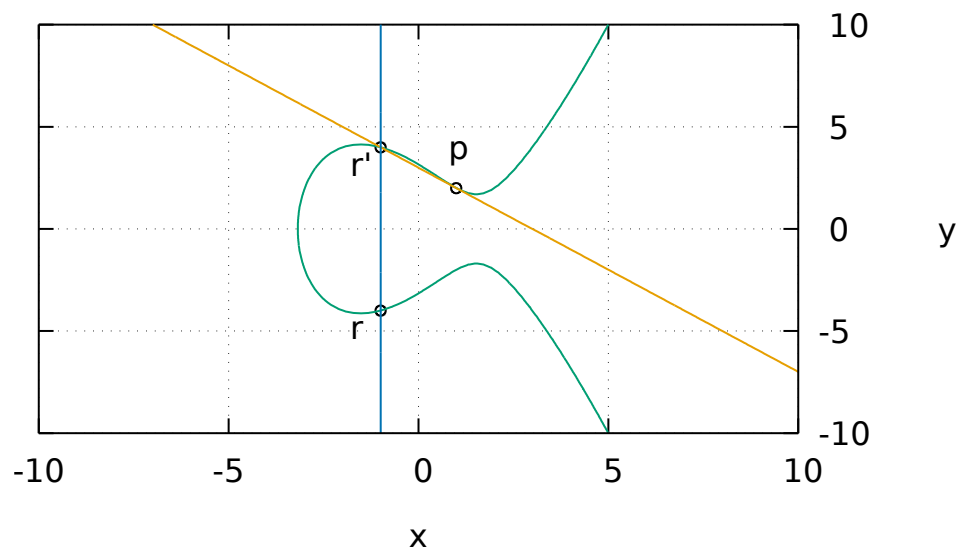


Figure 19.3: Elliptic curve $y^2 = x^3 - 7x + 10$ over the field of real numbers \mathbb{R} . Addition of two points with identical coordinates $p(1, 2)$: $2p = r(-1, -4)$

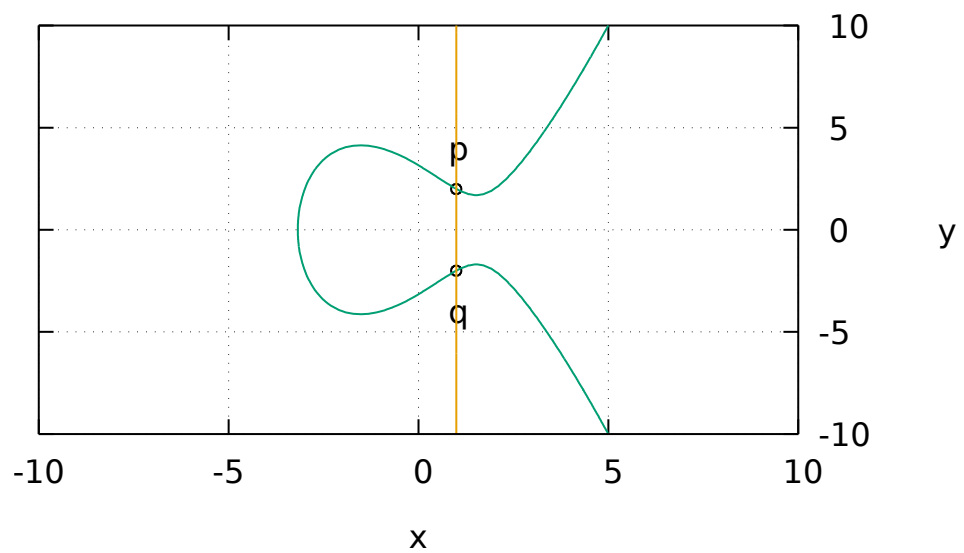


Figure 19.4: Elliptic curve $y^2 = x^3 - 7x + 10$ over the field of real numbers \mathbb{R} . Addition of two points $p(1, 2)$ and $q(1, -2)$. The line passing through these points does not intersect the curve. The result of the addition is the zero element: $p + q = 0$, i.e. $q = -p$

Reflecting this point relative to the X-axis we get the final point r which we will call the sum of the original points (see [Figure 19.2](#)). The coordinates of this point x_r, y_r can be obtained by the following formulas

$$\begin{aligned}x_r &= m^2 - x_p - x_q, \\y_r &= -y_p + m(x_p - x_r).\end{aligned}\tag{19.18}$$

In the case $x_p = x_q$ there are two possibilities:

1. $y_p = y_q$ (see [Figure 19.3](#)). When the points on the curve approach each other, the straight line drawn through them tends to the tangent line. The coefficient m can be found by the following formula: $m = \frac{dy}{dx}$. Taking into account $2ydy = 3x^2dx + adx$ we have $m = \frac{dy}{dx} = \frac{3x^2+a}{2y}$. Further calculation is carried out using formulas (19.18).
2. $y_p \neq y_q$ (see [Figure 19.4](#)). In this case, due to the symmetry of the curve with respect to the X-axis, only one possibility exists: $y_p = y_q$ and the line passing through these two points does not intersect the curve at a third point. For this case, an additional point 0 is introduced, where the line passing through the two points goes. Thus, in this case we have $p + q = 0, q = -p$.

Thus, we can define an elliptic curve over the field of real numbers \mathbb{R} as the following set of points

$$E(\mathbb{R}) = \{(x, y) \in \mathbb{R} \times \mathbb{R} : y^2 = x^3 + ax + b, \} \cup \{0\}.\tag{19.19}$$

where $a, b \in \mathbb{R}$.

For these points, a binary operation is defined, which we called addition. There is a zero element on the set, and for each element an inverse element can be defined. It can be proven that the introduced operation is associative: $(a + b) + c = a + (b + c)$ [26]. Thus, the set $E(\mathbb{R})$ forms a group with respect to addition. Considering the obvious equality $p + q = q + p$, this group will be commutative, i.e., Abelian (see definition 16.1.3).

Remark 19.10.1 (On the Addition Operation). *The introduced addition operation for points on an elliptic curve initially looks unnatural; why not use more obvious operations of addition of points on the plane, for example, vector addition rules. In this case, if $p, q \in E(\mathbb{R})$, then it is quite possible that $p + q \notin E(\mathbb{R})$ and, consequently, the main property of the group — closure — would be violated. Meanwhile, for the binary operation (19.18) defined by us, all group properties hold, and accordingly, all properties following from this, such as Lagrange's theorem (see theorem 16.1.2), which can be applied to the introduced objects.*

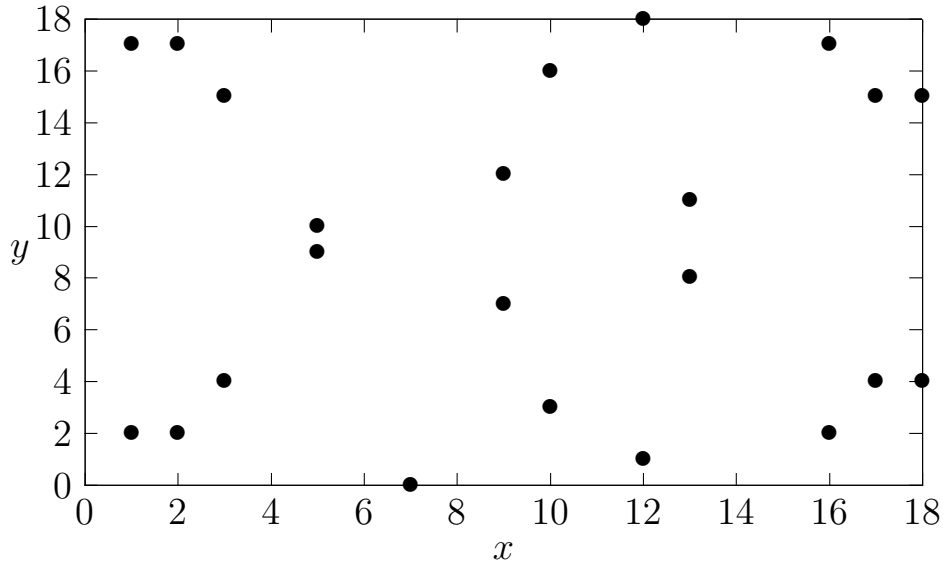


Figure 19.5: Elliptic curve $y^2 = x^3 - 7x + 10$ over the field \mathbb{F}_{19}

19.10.2 Elliptic Curves over the Field \mathbb{F}_p

The set (19.19) together with the addition operation (19.18) can be defined over an arbitrary field (see definition 16.2.1), i.e., not only over \mathbb{R} . From the cryptographic point of view, special interest lies in the field \mathbb{F}_p (see subsection 19.2.3). It is possible to define the set of elements of an elliptic curve over the field \mathbb{F}_p similarly to expression (19.2.3):

$$E(\mathbb{F}_p) = \{(x, y) \in \mathbb{F}_p \times \mathbb{F}_p : y^2 \equiv x^3 + ax + b, \} \cup \{0\}. \quad (19.20)$$

where $a, b \in \mathbb{F}_p$.

Definition 19.10.1 (Order of an Elliptic Curve). The number of points of an elliptic curve E is called its order and is denoted as $|E|$

Figure 19.5 shows such a set for the field \mathbb{F}_{19} , i.e., $p = 19$. The curve equation is $y^2 \equiv x^3 - 7x + 10 \pmod{19}$.

For each point a with coordinates x_a, y_a , the inverse element $-a$ is defined with coordinates $x_{-a} = x_a, y_{-a} \equiv -y_a \pmod{p}$.

For points on this curve, the following addition law is defined $a + b = c, b \neq -a$

$$\begin{aligned} x_c &\equiv m^2 - x_a - x_b \pmod{p}, \\ y_c &\equiv -y_a + m(x_a - x_c) \pmod{p}, \end{aligned} \quad (19.21)$$

where $(x_{a,b,c}, y_{a,b,c})$ are the coordinates of points a, b and c respectively. For the coefficient m , the following relations are used:

$$\begin{aligned} m &= (y_a - y_b)(x_a - x_b)^{-1} \pmod{p}, \text{ if } x_a \neq x_b \\ m &= (3x^2 + a)(2y)^{-1} \pmod{p}, \text{ if } x_a = x_b. \end{aligned}$$

Obviously, if $b = -a$, then $a + b = a + (-a) = 0$.

19.10.3 Scalar Multiplication and Discrete Logarithm

Assume n is some natural number $n \in \mathbb{N}$ and $a \in E(\mathbb{F}_p)$. Define scalar multiplication as follows:

$$n \cdot a = a + a + \cdots + a = \sum_{k=1}^n a$$

A naive implementation requires $O(n)$ addition operations, but using the divide and conquer paradigm (see [section 20.2](#)) it can be reduced to $O(\log n)$ addition operations.

Example 19.10.1 (Scalar Multiplication). *Consider the elliptic curve*

$$E(\mathbb{F}_{19}) = \{(x, y) \in \mathbb{F}_{19} \times \mathbb{F}_{19} : y^2 \equiv x^3 - 7x + 10\} \cup \{0\},$$

shown in [Figure 19.5](#). Choose the point $p = (13, 8)$, then

$$\begin{aligned} 0 \cdot p &= 0, \\ 1 \cdot p &= p = (13, 8), \\ 2 \cdot p &= p + p = (16, 7), \\ 3 \cdot p &= 2 \cdot p + p = (18, 5), \\ 4 \cdot p &= 3 \cdot p + p = (12, 1), \\ 5 \cdot p &= 4 \cdot p + p = (5, 10), \\ 6 \cdot p &= 5 \cdot p + p = (7, 0), \\ 7 \cdot p &= 6 \cdot p + p = (5, 9), \\ 8 \cdot p &= 7 \cdot p + p = (12, 18), \\ 9 \cdot p &= 8 \cdot p + p = (18, 4), \\ 10 \cdot p &= 9 \cdot p + p = (16, 2), \\ 11 \cdot p &= 10 \cdot p + p = (13, 11), \\ 12 \cdot p &= 11 \cdot p + p = 0 \end{aligned}$$

As seen from [example 19.10.1](#), every element of an elliptic curve is a generator of some cyclic subgroup. At the same time, the entire group of points on the elliptic curve is not necessarily cyclic (TBD). On the other hand, for formulating the discrete logarithm problem, we need exactly a cyclic group. Thus, for a given

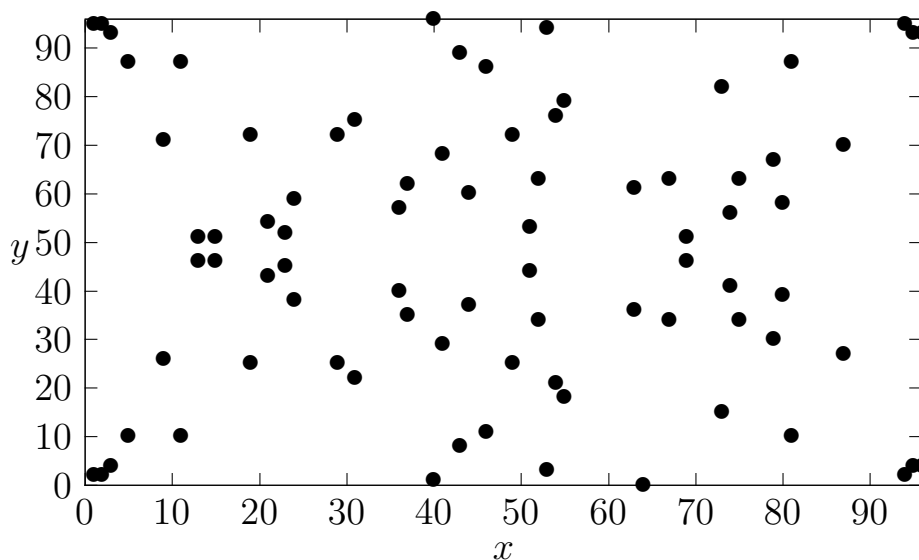


Figure 19.6: Elliptic curve $y^2 = x^3 - 7x + 10$ over the field \mathbb{F}_{97}

elliptic curve, first its order (see def. 19.10.1) is calculated, for which there exists an efficient Schoof's algorithm [22]. Then a prime divisor of the found order is located and a point is sought which is a generator of a subgroup of the chosen order. To do this, the following fact is used. For any point $g \in E$, the relation holds

$$Ng = 0,$$

where $N = |E|$ is the order (number of points) of the elliptic curve. Assume p is a prime divisor of N :

$$N = hp$$

then

$$Ng = p(hg) = 0.$$

Thus, if $hg \neq 0$, then the point $g' = hg$ will be a generator of a cyclic subgroup of order p .

Remark 19.10.2. *It makes sense to choose the order of the subgroup to be a prime number, since if the order of the group is prime then by Lagrange's theorem (see theorem 16.1.2) it has only trivial subgroups — the group itself and the subgroup consisting of the identity element. Thus, due to $hg \neq 0$, only the group itself remains, which is cyclic with order p .*

Example 19.10.2 (Choosing a Base Point). *Consider the elliptic curve*

$$E = E(\mathbb{F}_{97}) = \{(x, y) \in \mathbb{F}_{97} \times \mathbb{F}_{97} : y^2 \equiv x^3 - 7x + 10\} \cup \{0\},$$

shown in [Figure 19.6](#). The order of this curve is:

$$N = |E| = 82.$$

The number 82 has 2 divisors: 41 and 2. Thus, there exists a cyclic subgroup of order 41, i.e., $h = 2$.

Take the point $g = (1, 2) \in E$, its order: $|\langle g \rangle| = 82$, i.e., this point is not suitable. Calculate

$$g' = hg = 2g = (96, 93) \neq 0.$$

At the same time, $|\langle g' \rangle| = 41$, i.e., we have found the required base point.

If there are two points on the curve $a, b \in E(\mathbb{F}_p)$, the question of the existence of such $x \in \mathbb{N}$ makes sense:

$$x \cdot a = b$$

this problem is called the discrete logarithm problem on elliptic curves.

19.11 ECDH Algorithm

The ECDH algorithm is a modification of the Diffie-Hellman algorithm (see [section 19.8](#)) for elliptic curves. The Diffie-Hellman protocol is a key exchange protocol. In our case, the following elliptic curve parameters are published: (p, a, b, g, n, h) , where p, a, b define the curve

$$E(\mathbb{F}_p) = \{(x, y) : y^2 \equiv x^3 + ax + b \pmod{p}\} \cup \{0\},$$

g is the base point of order n : $|\langle g \rangle| = n$, h is the cofactor of the group $\langle g \rangle$, i.e., the order of the curve (see definition [19.10.1](#)) $|E| = nh$.

Alice chooses a private key $d_a \in \{1, \dots, n-1\}$ and forms a public key $A = d_a g$. Bob also forms a private key $d_b \in \{1, \dots, n-1\}$ and a public key $B = d_b g$. Alice and Bob exchange these keys. Then each of them calculates the actual key according to the rule $K = d_a B = d_b A$.

Example 19.11.1 (ECDH Algorithm). *Let's take the curve and base point from example [19.10.2](#). Thus,*

$$(p, a, b, g, n, h) = (97, -7, 10, (96, 93), 41, 2)$$

Alice chooses $d_a = 5$, i.e., $A = (37, 35)$. Bob chooses $d_b = 15$, thus $B = (15, 51)$. Alice's key $K = d_a B = (46, 11)$ and Bob's key $K = d_b A = (46, 11)$ match.

Chapter 20

Theory of Algorithms

20.1 Main Theorem on Recurrence Relations

Theorem 20.1.1 (Main Theorem on Recurrence Relations). *If there is the following recurrence relation for the complexity of some algorithm*

$$T(n) = aT\left(\frac{n}{b}\right) + f(n),$$

then it is possible to determine the asymptotic behavior of the function $T(n)$ in the following cases

1. *If $f(n) = O(n^{\log_b a - \epsilon})$, for some $\epsilon > 0$, then $T(n) = \Theta(n^{\log_b a})$*
2. *If $f(n) = \Theta(n^{\log_b a} \log^k n)$, then $T(n) = \Theta(n^{\log_b a} \log^{k+1} n)$*
3. *If $f(n) = \Omega(n^{\log_b a + \epsilon})$, for some $\epsilon > 0$ and $af\left(\frac{n}{b}\right) \leq cf(n)$ for some constant $c < 1$ and large n , then $T(n) = \Theta(f(n))$*

20.2 Divide and Conquer

"Divide and conquer" is an important paradigm in solving algorithmic tasks, which involves dividing the original problem into simpler ones.

20.3 Boolean Logic. Boolean Formulas in CNF

Definition 20.3.1 (Conjunction (logical "AND")). For the logical operation "AND" ($a \& b$), the following truth table [20.1](#) holds.

Definition 20.3.2 (Disjunction (logical "OR")). For the logical operation "OR" ($a \parallel b$), the following truth table [20.2](#) holds.

| a | b | $a \& b$ |
|-----|-----|----------|
| 0 | 0 | 0 |
| 0 | 1 | 0 |
| 1 | 0 | 0 |
| 1 | 1 | 1 |

Table 20.1: Conjunction $a \& b$

| a | b | $a \parallel b$ |
|-----|-----|-----------------|
| 0 | 0 | 0 |
| 0 | 1 | 1 |
| 1 | 0 | 1 |
| 1 | 1 | 1 |

Table 20.2: Disjunction $a \parallel b$

Definition 20.3.3 (Boolean Formula). A Boolean formula is a set of Boolean literals combined by logical operations.

Definition 20.3.4 (CNF - Conjunctive Normal Form (SAT)). In Boolean logic, a conjunctive normal form is a Boolean formula that is a conjunction of disjunctions of literals.

Example 20.3.1. SAT Problem *The SAT problem is the problem of determining the satisfiability of CNF. Suppose the following Boolean formula is given in CNF form:*

$$(x_1 \parallel x_2 \parallel \bar{x}_3) \& (x_1 \parallel \bar{x}_2 \parallel x_3)$$

It is satisfiable; it is sufficient to set $x_1 = 1$, and the values of $x_{2,3}$ can be arbitrary. Meanwhile, the following formula is unsatisfiable:

$$(x_1 \parallel x_2 \parallel \bar{x}_3) \& (\bar{x}_1 \parallel \bar{x}_2 \parallel x_3).$$

Theorem 20.3.1 (Reducibility to CNF). *Any Boolean formula can be reduced to CNF.*

Proof. For the proof, you may use: the law of double negation, De Morgan's laws, distributivity. □

Theorem 20.3.2 (Reducibility to 3-CNF (3-SAT)). *Any Boolean formula in conjunctive normal form can be reduced to 3-CNF.*

Proof. TBD □

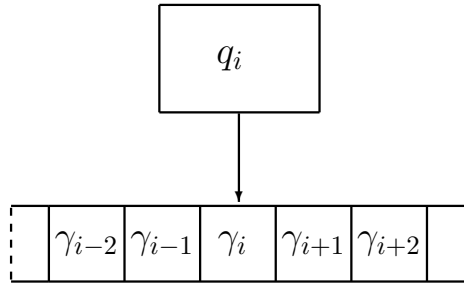


Figure 20.1: Turing Machine. The current state of the machine is described by two numbers (q_i, γ_i) , where $q_i \in Q$ is the state of the control unit, which points to the tape element $\gamma_i \in \Gamma$.

20.4 Turing Machine

A Turing Machine TM is a mathematical object consisting of a control device (head) and an infinite tape divided into cells, each of which can hold one symbol.

The control device TM can be in one of the states $q_i \in Q$, where Q is a set of different states. The state in which TM was at the initial time $q_0 \in Q$ is called the initial state.

Among the elements of the set Q , a subset $F \subset Q$ of final states is selected. If TM is in a state $f \in F \subset Q$, it is said that the Turing Machine has terminated its operation.

The symbols recorded on the tape form a certain set (alphabet) $\Gamma = \{\gamma_i\}$.

The input alphabet Σ of the Turing Machine is a subset of the elements Γ with which the control device TM operates, i.e., the symbols it can recognize on the tape. Those symbols that the control device cannot recognize will be denoted by the symbol β (blank).

In the process of operation, TM performs the following actions:

- reads the element pointed to by the control device (see [Figure 20.1](#))
- changes (or leaves unchanged) this symbol to an element $\sigma \in \Sigma \subset \Gamma$
- changes the position of the control device by one cell to the right (R), to the left (L) or does not change it (S)

The change of state TM is described by a transition table Δ . An element of the transition table is recorded in the following form: $\delta_j(q_{old}, \gamma_{old})$ where $q_{old} \in Q$ is the current state, $\gamma_{old} \in \Gamma$ is the selected element on the tape. The value of the transition can be undefined or recorded in the form $\{q_{new}, \gamma_{new}, d\}$ (see [Figure 20.2](#)), where $q_{new} \in Q$ is the new state of the Turing Machine, $\gamma_{new} \in \Gamma$ is the new value of the element on the tape pointed to by the control device, i.e., γ_{old} is replaced by γ_{new} . $d \in D = \{L, R, S\}$ is the direction of movement of

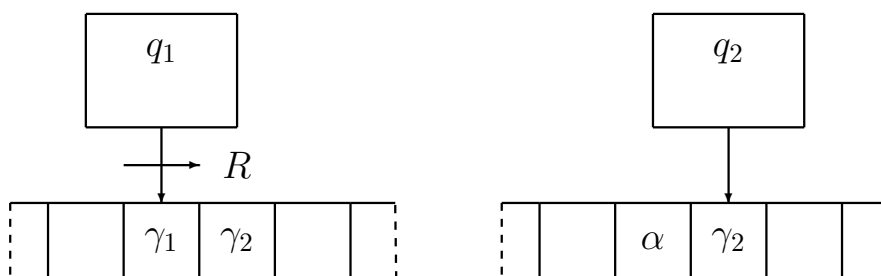


Figure 20.2: Transition of a Turing machine from state (q_1, γ_1) to state (q_2, γ_2) . The corresponding element of the transition table $\delta : \{q_1, \gamma_1\} \rightarrow \{q_2, \alpha, R\}$. That is, the tape element initially pointed to by the control device (γ_1) changes its value to α , and the control device changes its position by one cell to the right (R).

the control device (L left, R right, S stay in the same place). In the case of an undefined transition, it is said that the Turing Machine has halted in state q_{old} .

Definition 20.4.1 (Language of the Turing Machine $L(M)$). The original symbols on the tape can be interpreted as a certain string x that is fed to the input of the Turing Machine M . All those strings that take the machine M to a certain final state $f \in F$ are called the language $L(M)$ of machine M

Definition 20.4.2 (Deterministic Turing Machine). If in the transition table the mapping $(q_{old}, \gamma_{old}) \rightarrow \{q_{new}, \gamma_{new}, d\}$ is a bijection ¹, then the corresponding Turing Machine is called a deterministic Turing Machine

Definition 20.4.3 (Non-deterministic Turing Machine). If in the transition table the mapping $(q_{old}, \gamma_{old}) \rightarrow \{q_{new}, \gamma_{new}, d\}$ is not a bijection, i.e., the state (q_{old}, γ_{old}) may correspond to several elements of the transition table $\{q_{new}, \gamma_{new}, d\}$, then the corresponding Turing Machine is called a non-deterministic Turing Machine

20.5 Computability Theory

A Turing machine M (see Figure 20.3) can be interpreted as a certain function m whose argument is a set of symbols representing the initial state of the tape x . As a result of processing x , the machine M can finish its work, i.e., end up in some final state $f \in F \subset Q$. In this case, it is said that M accepts x . Then the corresponding function value is $m(x) = 1$.

At the same time, the computation of the function $m(x)$ can be interpreted as a certain problem.

¹Bijection is a one-to-one mapping, i.e., in our case, each state (q_{old}, γ_{old}) corresponds to only one element of the transition table $\{q_{new}, \gamma_{new}, d\}$

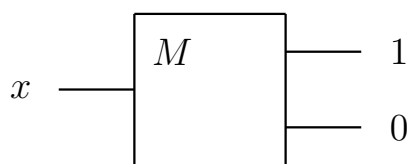


Figure 20.3: Turing machine M as a function $m(x)$. The input consists of a certain sequence of initial data x (initial symbols on the tape). On output, M can be in some final state (1 - accept x), halt (0 - reject x) - i.e., be in a state with no further transitions. The machine can also enter an infinite loop and never complete its work.

Definition 20.5.1 (Problem). We call a problem with respect to a Turing machine a yes/no answer to a question about some set of possible objects.

Example 20.5.1 (Hamiltonian Graph). *The Hamiltonian graph problem - for a given graph, whether there exists a path (chain) containing each vertex of this graph exactly once.*

In the case of halting (halt) of M , i.e., when M reaches a state from which no further transition is defined, it is said that M rejects x . In this case, the corresponding function value is $m(x) = 0$.

Also, the machine M on input x may never terminate its operation. In this case, the value of the function $m(x)$ is undefined.

Definition 20.5.2 (Recursive Language). The set of strings L ² is called a recursive language if there exists a Turing machine M such that $\forall x \in L$ the corresponding function $m(x) = 1$, and $\forall x \notin L$ the corresponding function $m(x) = 0$.

Thus, M corresponding to a recursive language $L(M)$ halts on any input. The set of recursive languages is denoted R .

Definition 20.5.3 (Recursively Enumerable Language). The set of strings L is called a recursively enumerable language if there exists a Turing machine M such that $\forall x \in L$, M halts in some final state, i.e., the corresponding function $m(x) = 1$.

Thus, M corresponding to a recursively enumerable language $L(M)$ halts on every input belonging to the language, but may never halt on input not belonging to the language $L(M)$. The set of recursively enumerable languages is denoted RE . Obviously $R \subset RE$.

A Turing machine defines a class of functions for which the notion of computability can be defined.

²the set of initial tape states of some Turing machine

Definition 20.5.4 (Computable Function).³ A function m is called computable, and the corresponding problem - decidable if there exists some Turing machine M for it, and the language corresponding to this machine is recursive, $L(M) \in R$. Otherwise, the corresponding problem is called undecidable.

Theorem 20.5.1 (On the cardinality of the set of Turing Machines). *The set of Turing machines is countable.*

Proof. A set is called countable [54] if there exists a natural number assigned to each element. A Turing machine is uniquely determined by the following group of elements:

- Set of states Q
- Start state q_0
- Set of final states $F \subset Q$
- Alphabet Σ
- Transition table Δ

Since each of the listed elements of a Turing machine takes on a finite number of values, a unique (for the corresponding machine) number can be assigned to each element, and therefore to the entire Turing machine,⁴ which proves the countability of the set of all Turing machines. \square

After introducing the notion of a computable function, the question arises about the computability of an arbitrary function. Suppose we have some input data (function argument) x . Does there exist a Turing machine M that accepts this data, i.e., M halts in some final state? The following theorem answers this question.

Theorem 20.5.2 (On Undecidability). *There exists a language L_d that is not accepted by any Turing machine, i.e., the corresponding problem L_d is fundamentally undecidable (not computable).*

Language L_d is often called the diagonalization language because if one constructs a matrix [13] whose row indices denote the numbers of input data of Turing machines, and the column numbers correspond to the Turing machine number, then the value of this matrix a_{ij} is equal to 1 if M_j accepts x_i and 0 otherwise. The inverted⁵ numbers on the diagonal define the language L_d .

³In fact, the notion of computability cannot be defined strictly mathematically and is given here in the formulation of the Church-Turing thesis.

⁴For example, one can encode each machine by a unique sequence of numbers 0 and 1. The resulting sequence is finite due to the finiteness of the set of elements comprising the Turing machine. Any finite binary sequence corresponds to a unique natural number for which the considered sequence is the binary representation.

⁵0 replaced by 1 and 1 by 0

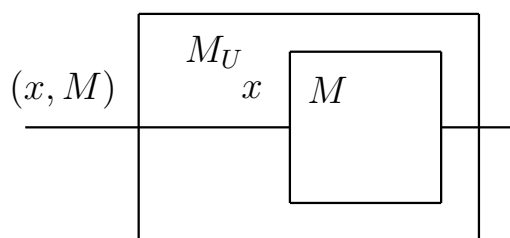


Figure 20.4: Universal Turing machine M_U . M_U receives as input a pair of strings (x, M) , where M is the code of some Turing machine on which the initial data x must be executed

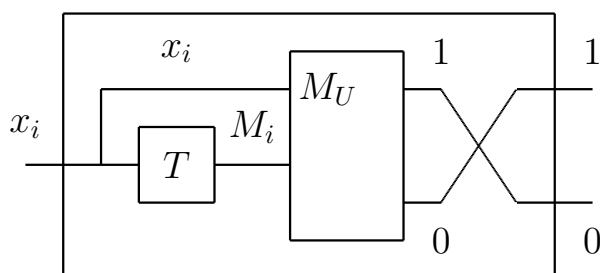


Figure 20.5: Proof that $L_U \notin R$. If $L_U \in R$, then the universal Turing machine M_U can be constructed for the Turing machine M_d , which accepts all strings of the language L_d , which is impossible due to the theorem 20.5.2. T is a special device that creates a code for the Turing machine M_i based on the string x_i , which does not accept x_i .

Proof. By theorem [On the cardinality of the set of Turing Machines](#) (Theorem 20.5.1), one can talk about the i -th Turing machine M_i , where i is some natural number. Suppose x_i is a string not recognized by M_i , i.e., $x_i \notin L(M_i)$. The strings x_i form some language L_d . By the definition of L_d , no Turing machine exists that accepts all strings of this language. \square

Definition 20.5.5 (Universal Turing Machine). A universal Turing machine M_u is a machine that receives on input a string x and a code for some Turing machine M and computes the result of applying x to M .

The set of pairs of strings (x, M) form some language L_U of the machine M_U .

Theorem 20.5.3 (On the Language of the Universal Turing Machine). *The language of the universal Turing machine is recursively enumerable but not recursive, i.e., $L_U \in RE$ and $L_U \notin R$. Thus, the problem of recognizing the language of the universal Turing machine is undecidable.*

Proof. The statement $L_U \in RE$ follows directly from the definition of the universal Turing machine: if M accepts x , then M_U accepts the pair of strings (x, M) .

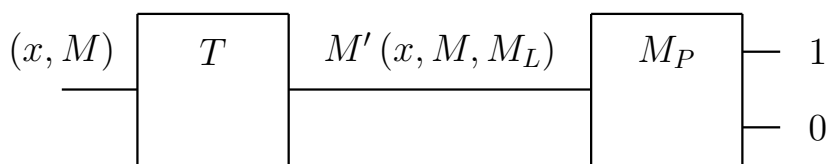


Figure 20.6: Rice's Theorem. Turing machine M_P recognizing a nontrivial property of language P is used to construct a universal Turing machine M_U which can either accept or reject pairs of strings (x, M) . An adapter T is used to transform the pair of strings (x, M) into the code for a Turing machine M' which satisfies the property P if and only if M accepts x .

We prove $L_U \notin R$ by contradiction. Suppose $L_U \in R$. Using such a machine M_U , construct a Turing machine M_d that will accept all strings of the language L_d . For each string $x_i \in L_d$ there is a corresponding Turing machine M_i which does not accept this string, i.e., $x_i \notin L(M_i)$. The pair (x_i, M_i) can be input to the universal Turing machine as shown in Figure 20.5. If $L_U \in R$, then machine M_U can decide whether M_i accepts input x_i or not. If M_i does not accept x_i , then machine M_d should accept this data. Thus, we have constructed a Turing machine M_d that accepts all strings $x_i \in L_d$ and rejects all strings $x_i \notin L_d$, which is impossible due to theorem [On Undecidability](#) (Theorem 20.5.2). \square

Every Turing machine M has some code which can be interpreted as input data for another Turing machine M_P . The machine can either accept or reject this code. Due to the one-to-one correspondence between a Turing machine M and its language $L(M)$, one can talk about a Turing machine that accepts or rejects languages L . If machine M_P accepts some set of languages, it is said that these languages possess some property P .

Definition 20.5.6 (Nontrivial Property of a Turing Machine Language). A property of a language is called nontrivial if among all languages one can distinguish both those which satisfy the property and those which do not.

Theorem 20.5.4 (Rice's Theorem). *The problem of determining any nontrivial property of the language $L(M)$ of a Turing machine M is undecidable.*

Proof. Suppose we have a machine M_P that can determine some nontrivial language property P , i.e., it can classify the languages of all Turing machines into two sets (each of which is nonempty). The first set contains all languages satisfying P (the machine M_P accepts languages from this set), and the second set contains all languages that do not satisfy property P .

Let us try to construct on the basis of M_P a universal Turing machine M_U that accepts (or rejects) pairs (x, M) (see Figure 20.6).

Suppose we have a machine M_L that accepts strings l from some language L ⁶.

As the first step, construct a machine M' that accepts strings l . At the same time, the pair (x, M) is fixed and embedded in the code of M' . The key property of M' is the following: M' accepts strings $l \in L$ if and only if M accepts x . Thus, two cases are possible:

- $L(M') = L(M_L)$ if M' accepts l as a result of accepting the pair (x, M)
- $L(M') = \emptyset$ if M' rejects any input l , i.e., in the case where M never halts on input x .

Next, recall the property P and suppose that the language L (accepted by machine M') has property P , that is, $L \in P$. Also assume that $\emptyset \notin P$, ⁷. Thus, if the code for M' is given as input to machine M_P , then M_P accepts the input (code M') if $L(M') \in P$ and rejects it if $L(M') = \emptyset$.

On the other hand, the pair (x, M) is embedded in the original code of M' , so we can assume that M_P receives exactly the pair (x, M) as input, accepting it if M accepts x ($L(M') \in P$) and rejecting this pair if M does not accept x ($L(M') = \emptyset \notin P$). Thus, using M_P we have constructed a universal Turing machine and shown that its language is recursive, which is impossible by theorem [On the Language of the Universal Turing Machine](#) (Theorem 20.5.3). \square

20.6 Complexity Classes P , NP

When solving a particular problem, two questions arise primarily. Does the given problem have a solution in principle, i.e. can an algorithm be constructed that solves the given problem? Once a positive answer to the first question is found, i.e. an algorithm solving the problem is constructed, the next issue is the practical feasibility of the proposed solution. Here, the theory of algorithm complexity helps us address the feasibility of various algorithms.

In the following table [38] some large numbers useful for analyzing algorithm complexity are given

Example 20.6.1. Prime factorization of integers by trial division *Suppose we have the problem of factorizing an integer x (see example 13.0.2). In binary representation, x consists of N digits, i.e. $2^N \leq x < 2^{N+1}$. In the simplest approach (trial division), we can try all integers from 2 up to $2^{\frac{N}{2}}$, and for each such integer $2 < y \leq 2^{\frac{N}{2}}$ find the remainder of division of x by y . If x is divisible*

⁶There is no relation between language L and property P at this step

⁷If $\emptyset \in P$, then one can consider the property $\bar{P} = \{L : L \notin P\}$. In this case, $\emptyset \notin \bar{P}$.

| Physical analogy | Number |
|--|--|
| Probability of being struck by lightning (in one day) | 1 in 9 billion (2^{33}) |
| Probability of winning the grand prize in the US state lottery | 1 in 4,000,000 (2^{22}) |
| Probability of winning the grand prize in the US state lottery and being struck by lightning on the same day | 1 in 2^{61} |
| Probability of drowning (in the USA within a year) | 1 in 59,000 (2^{16}) |
| Probability of dying in a car accident (in the USA in a year) | 1 in 6,100 (2^{13}) |
| Probability of dying in a car accident (in the USA during a lifetime) | 1 in 88 (2^7) |
| Time until next glaciation | 14,000 (2^{14}) years |
| Time until the Sun becomes a supernova | 10^9 (2^{30}) years |
| Age of the planet | 10^9 (2^{30}) years |
| Age of the Universe | 10^{10} (2^{34}) years |
| Number of atoms in the planet | 10^{51} (2^{170}) |
| Number of atoms in the Sun | 10^{57} (2^{190}) |
| Number of atoms in the galaxy | 10^{67} (2^{223}) |
| Number of atoms in the Universe | 10^{77} (2^{265}) |
| Volume of the Universe | 10^{84} (2^{280}) cm^3 |
| If the Universe is finite: | |
| Total lifetime of the universe | 10^{11} (2^{37}) years 10^{18} (2^{61}) seconds |
| If the Universe is infinite: | |
| Time until cooling of low-mass stars | 10^{14} (2^{47}) years |
| Time until planets detach from stars | 10^{15} (2^{50}) years |
| Time until stars detach from galaxies | 10^{19} (2^{64}) years |
| Time until orbital decay by gravitational radiation | 10^{20} (2^{67}) years |
| Time until black holes evaporate by Hawking processes | 10^{64} (2^{213}) years |
| Time until matter becomes liquid at zero temperature | 10^{65} (2^{216}) years |

Table 20.3: Large numbers [38]

by y , the problem is solved and y is declared a divisor of x . If none of the numbers y divide x , then x is declared prime. Thus, the original problem is divided into $2^{\frac{N}{2}}$ subproblems involving computing remainders.

Assume x has 20 digits, i.e. $N = 20$. In that case, $2^{10} = 1024$ division operations need to be performed to find a solution to the original problem, which can be easily done on simple computers.

However, if the number of digits increases 30-fold to $N = 600$, then to solve the given problem, one must perform 2^{300} division operations. If you look at Table 20.3, you can see that this number is much larger than the number of atoms in the universe, which is on the order of 2^{265} . That is, even if each atom in the universe represents a very simple computational device that computes the remainder of division of two integers, using all atoms in the universe would be insufficient to solve the problem.

Thus, despite the existence of a simple algorithm, the problem of prime factorization, using this algorithm, becomes practically infeasible for numbers of length 600 bits, even though the number 600 does not seem very large compared to 20 (where the task is practically feasible).

20.6.1 Hierarchy of Complexity Classes

Below are formal definitions for problems that are practically feasible, P . The class of problems that are practically infeasible (until it is proven that $P = NP$) is also defined.

Definition 20.6.1 (Problem size). The size of the problem N is defined as the number of symbols from the alphabet of the Turing machine contained on the tape in the initial state, that is, in other words, the size of the problem is considered to be the length of the string input to the Turing machine.

For example, in example 20.6.1, the problem size is the length of number x (whose divisors are being sought) in binary representation.

Definition 20.6.2 (Class P). A problem (algorithm) belongs to class P (Polynomial) if it can be solved in $O(N^k)$ steps on a deterministic Turing machine. Here N is the size of the original problem, k is any integer independent of N .

The class P defines the class of problems that are practically feasible. For example, the problem of sorting unordered data (a list) can be solved in $O(N \log N) = O(N^2)$ steps, where N is the number of elements in the list to be sorted (the problem size).

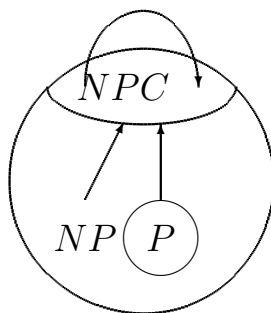


Figure 20.7: The relationship of complexity classes P , NP , and NPC . The presented picture corresponds to the assumption $P \neq NP$. In this case, $P \subset NP$, $NPC \subset NP$. Any problem from P and NP can be reduced to NPC in polynomial time.

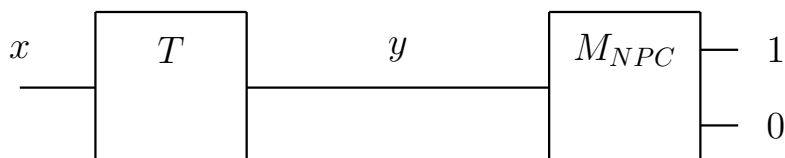


Figure 20.8: Class NPC . The input for any NP problem x can be transformed into an input y for some problem in the class NPC . The transformation $x \rightarrow y$ can be done in $O(N^k)$ steps, where $N = |x|$ is the length of x , k is some integer independent of N .

Definition 20.6.3 (Class NP). A problem (algorithm) belongs to class NP (Nondeterministic Polynomial) if it can be solved in $O(N^k)$ steps on a nondeterministic Turing machine. Here N is the size of the original problem, k is any integer independent of N .

One example of an NP problem is integer factorization. A nondeterministic Turing machine can guess a solution (in one step) and verify it (multiply the two guessed numbers and compare the product to the original number) in polynomial time. At the same time, the question of whether the factorization problem belongs to class P remains open.

The problem of equality of classes P and NP is one of the most important unsolved problems in mathematics and probably the most important unsolved problem in algorithm theory. To show how an NP problem can be reduced to a P problem, it is necessary to introduce the class of NP -complete problems, which are, in some sense, the hardest NP problems.

Definition 20.6.4 (Class NPC (NP -complete)). A problem M belongs to class NPC if any NP problem can be reduced to M (see [Figure 20.8](#)).

Thus, if a polynomial algorithm is found for at least one NPC problem, then polynomial algorithms will be found for all NP problems (see [Figure 20.7](#)).

20.6.2 Cook-Levin Theorem

Once we introduced the complexity class NPC , the question arises about the problems included in it. Below it will be shown that the SAT problem (on the satisfiability of a boolean formula in conjunctive normal form) is NP complete.

Theorem 20.6.1 (Cook-Levin). *The problem of the satisfiability of a boolean formula in CNF is NP complete, i.e., belongs to the class NPC .*

Proof. The proof consists of two parts. First, we must prove that the SAT problem (see example [20.3.1](#)) belongs to the class NP . Second, we must show that any problem from the class NP can be reduced to the SAT problem in polynomial time.

Membership in the class NP is proven in an obvious way, - since there is the possibility of verifying the truth of a solution in $O(N)$ steps, where N is the number of boolean operations used in the given boolean formula.

TBD

□

Part 6

Answers to Questions

Chapter 21

Quantum Electrodynamics

21.1 Quantization of the Electromagnetic Field

21.1.1 5: Number of modes

How many electromagnetic field modes with wavelength $\lambda \geq 500\text{nm}$ are in the quantization cube with side $L = 1\text{mm}$?

The wavelength range $\lambda \geq \lambda_0 = 500\text{nm}$ corresponds to frequency range $\omega \leq \omega_0 = \frac{2\pi c}{\lambda_0}$.

Using formula 1.47

$$dN = 2 \left(\frac{L}{2\pi c} \right)^3 \omega^2 d\omega d\Omega$$

hence

$$\begin{aligned} N &= \int_0^{\omega_0} 2 \left(\frac{L}{2\pi c} \right)^3 \omega^2 d\omega \int_{4\pi} d\Omega = \\ &= 8\pi \int_0^{\omega_0} \left(\frac{L}{2\pi c} \right)^3 d\omega = 8\pi \left(\frac{L}{2\pi c} \right)^3 \frac{\omega_0^3}{3} = \\ &= \frac{8\pi}{3} \left(\frac{L}{\lambda_0} \right)^3 \approx 67 \cdot 10^9 \end{aligned}$$

21.2 Interaction of Light with an Atom

21.2.1 13: Determination of lithium atom transition frequency

Let us denote by $\Omega_R^{(1,2)}$ the effective Rabi frequencies for measurements 1 and 2, respectively. The frequency detunings for the two experiments are $\delta_{1,2}$. Thus,

we have

$$\begin{aligned}\Omega_R^{(1)} &= \sqrt{\omega_R^2 + \delta_1^2}, \\ \Omega_R^{(2)} &= \sqrt{\omega_R^2 + \delta_2^2}\end{aligned}\tag{21.1}$$

On the other hand, from the graphs [Figure 2.18](#) it is seen that the maximum probability to detect lithium atoms in state $|2\rangle$ for frequency $\omega_2 = 2\pi \cdot 228.4\text{MHz}$ will be $P_2 = 0.49$, while from [\(2.27\)](#) we have

$$P_2 = \left(\frac{\omega_R}{\Omega_R^{(2)}} \right)^2 = \frac{1}{1 + \frac{\delta_2^2}{\omega_R^2}},$$

from which

$$\omega_R = 0.98 \cdot \delta_2.\tag{21.2}$$

On the other hand, one can conclude that

$$\frac{\Omega_R^{(1)} \cdot 6\mu\text{s}}{2} = \pi$$

and

$$\frac{\Omega_R^{(2)} \cdot 12.5\mu\text{s}}{2} = \pi,$$

thus

$$\Omega_R^{(1)} = 2\pi \cdot 0.167\text{MHz},$$

and

$$\Omega_R^{(2)} = 2\pi \cdot 0.08\text{MHz},$$

From [\(21.1\)](#) and [\(21.2\)](#) one can find

$$\delta_2 = 2\pi \cdot 0.057\text{MHz},$$

i.e. two candidates

$$f_0 = 228.4 \pm 0.057\text{MHz} = 228.343\text{MHz}, 228.457\text{MHz}\tag{21.3}$$

For δ_1 we have

$$\delta_1 = \sqrt{\left(\Omega_R^{(1)}\right)^2 - \omega_R^2} = 2\pi \sqrt{0.167^2 - (0.98 \cdot 0.057)^2}\text{MHz} = 2\pi \cdot 0.157\text{MHz}$$

thus we get the next two candidates

$$f_0 = 228.5 \pm 0.157\text{MHz} = 228.343\text{MHz}, 228.657\text{MHz}\tag{21.4}$$

From (21.3) and (21.4) we finally have

$$f_0 = 228.343\text{MHz}.$$

It is worth noting that the obtained value differs from the real $f_0 = 228.205\text{MHz}$. This deviation is caused by the influence of the Zeeman effect from the magnetic field, which is usually present in such experiments.

Chapter 22

Quantum Optics

22.1 Quantum Theory of the Laser

22.1.1 Natural Linewidth

From (3.24) we have

$$D = \frac{A}{2\bar{n}_{st}} = \frac{10^6}{2 \cdot 10^6} = 0.5\text{Hz}.$$

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