

CHROMOSPHERIC Ca II EMISSION IN NEARBY F, G, K, AND M STARS¹

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Received 2003 November 11; accepted 2004 February 17

ABSTRACT

We present chromospheric Ca II H and K activity measurements, rotation periods, and ages for ~1200 F, G, K, and M type main-sequence stars from ~18,000 archival spectra taken at Keck and Lick Observatories as a part of the California and Carnegie Planet Search Project. We have calibrated our chromospheric *S*-values against the Mount Wilson chromospheric activity data. From these measurements we have calculated median activity levels and derived R'_{HK} , stellar ages, and rotation periods from general parameterizations for 1228 stars, ~1000 of which have no previously published *S*-values. We also present precise time series of activity measurements for these stars.

Subject headings: stars: activity — stars: chromospheres — stars: rotation

On-line material: machine-readable tables

1. INTRODUCTION

The California and Carnegie Planet Search Program has included observations of ~2000 late-type main-sequence stars at high spectral resolution as the core of its ongoing survey of bright, nearby stars to find extrasolar planets through precision radial velocity measurements (e.g., Cumming, Marcy, & Butler 1999; Butler et al. 2003). One source of error in the measured velocities is that due to “photospheric jitter”: flows and inhomogeneities on the stellar surface can produce variations in the measured radial velocity of a star and may even mimic the signature of planetary companions (Henry, Donahue, & Baliunas 2002; Queloz et al. 2001; Santos et al. 2003). In addition to providing a precision radial velocity, each of our radial velocity observations provides a measurement of the strength of the stellar chromospheric Ca II H and K emission cores. These measurements are an indicator of stellar magnetic activity and can provide an estimate of the photospheric jitter and rotation period of a star, both critical values for understanding and interpreting the noise present in radial velocity measurements (Noyes et al. 1984; Saar & Fischer 2000; Santos et al. 2000).

The largest campaign to measure and monitor Ca II H and K emission has been the Mount Wilson program begun by O. C. Wilson (1968) and continued and improved since then by Vaughan, Preston, & Wilson (1978) and others (Baliunas et al. 1998). From 1966 to 1977 this program used the “HKP-1” photometer, which employed a photoelectric scanner at the coudé focus of the 100 inch telescope. Since 1977 the “HKP-2” photometer has been used, which is a new, specially designed photomultiplier mounted at the Cassegrain focus of the 60 inch telescope (e.g., Baliunas et al. 1995).

Duncan et al. (1991) published data from this program in the form of “season averages” of H and K line strengths from 65,263 observations of 1296 stars (of all luminosity classes) in the Northern Hemisphere, and later as detailed analyses of 171,300 observations of 111 stars characterizing the varieties and evolution of stellar activity in dwarf stars. This program defined the Mount Wilson “*S*-value,” which has become a standard metric of chromospheric activity.

Henry et al. (1996) published data from a survey of stars in the southern hemisphere, providing *S*-values from 961 observations of 815 stars. Other surveys include the Vienna-KPNO survey (Strassmeier et al. 2000), whose motivation was to find Doppler-imaging candidates by using Ca II H and K as a tracer of rotation period in 1058 late-type stars, and that of our Anglo-Australian Planet Search (Tinney et al. 2002), which reported *S*-values for 59 planet search stars not observed by previous surveys.

2. OBSERVATIONS

Observations for the California and Carnegie Planet Search Program have used the HIRES spectrometer at Keck Observatory for 6 years, measuring precision velocities of ~700 stars as part of a campaign to find and characterize extrasolar planetary systems (e.g., Butler et al. 1996). HIRES is an echelle spectrometer, which yields high-resolution (67,000) spectra from 3850 to 6200 Å. Typical exposures in the Ca II H and K region yield a signal-to-noise ratio of 60 in the continuum, although this number can be smaller for very red stars since our requisite signal-to-noise ratio in the iodine region of the spectrum dictates exposure time.

The detector on HIRES is a Tektronix 2048EB2 engineering-grade CCD optimized for the optical. The quantum efficiency degrades significantly blueward of the H and K lines but is still 60% at 0.38 μm. Observations at Keck always employ an image rotator to keep the position angle parallactic, thereby minimizing the effects of atmospheric dispersion (Vogt et al. 1994).

The Planet Search program has also included observations made at Lick Observatory since 1987 with the Hamilton spectrograph fed by the Shane 3 m telescope and the 0.6 m Coudé Auxiliary Telescope (CAT) (Vogt 1987). The Hamilton spectrograph is also an echelle spectrometer with high resolution

¹ Based on observations obtained at Lick Observatory, which is operated by the University of California, and on observations obtained at the W. M. Keck Observatory, which is operated jointly by the University of California and the California Institute of Technology. The Keck Observatory was made possible by the generous financial support of the W. M. Keck Foundation.

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(60,000). In 2001 the CCD readout window was expanded to include the Ca II H and K region, where we typically achieve a signal-to-noise ratio between 10 and 60 in the continuum. This large range of S/N is partly due to the fact that the program does not employ an image rotator at Lick in order to keep the overall system efficiency high. As a result, significant blue flux can be lost for those observations at large hour angles and air masses (Filippenko 1982).

Two detectors have been used on the Hamilton spectrograph since the readout window was expanded. The first CCD, referred to as “Dewar 6” for the Dewar it sits in, is an overthinned 2048×2048 chip with $15 \mu\text{m}$ pixels. The second, “Dewar 8,” is a Lawrence Berkeley Laboratory high-resistivity CCD with the same dimensions as Dewar 6.

3. THE STELLAR SAMPLE

The Planet Search at Lick and Keck observatories has included over 1000 stars over its duration, with many stars being added and a few dropped along the way as resources and circumstances dictate. Cumming et al. (1999) analyzed the frequency of planets around many (76) of the best-observed Lick program stars, and Nidever et al. (2002) published absolute radial velocities for most program stars at Lick and Keck, and many other stars (889) observed in the course of the program. These same spectra were analyzed by LTE atmosphere modeling to yield T_{eff} , $\log g$, $v \sin i$, and chemical abundances by J. A. Valenti & D. A. Fischer (in preparation) and D. A. Fischer & J. A. Valenti (in preparation). The sample of this paper includes every star observed at Keck and Lick for which an accurate S can be obtained. It is composed of 1231 stars, 1199 of which have at least one measured S from Keck and 132 of which have at least one from Lick.

The Planet Search initially included single, late type, dwarf stars accessible from Lick Observatory. As the resources of the Planet Search have grown, fainter F, G, and K dwarfs, M dwarfs, subdwarfs, and some subgiant stars have been added, and as the effect of activity on velocity precision has been uncovered, some more active stars have been dropped. This paper’s sample constitutes the stars that were still being monitored when activity measurements began (some now dropped), as well as stars added to the search since then and a few incidental targets.

4. DATA REDUCTION

4.1. S -Values

The S -value is defined by the operation of the Mount Wilson spectrometers (Duncan et al. 1991), which measure a quotient of the flux in two triangular bandpasses centered on the H and K emission cores and two continuum regions on either side. Duncan et al. (1991) refer to these channels as the H, K, R , and V channels (where the R and V channels are the continuum channels on the red and blue sides, respectively, of the H and K channels.) The HKP-1 spectrometer measured two 25 \AA -wide R and V channels separated by about 250 \AA about the rest position of the H and K lines, and the H and K channels, which had a triangular instrumental profile with FWHM close to 1 \AA . The HKP-2 spectrometer consisted of two 20 \AA -wide R and V channels centered on 4001.07 and 3901.07 \AA in the star’s frame, and triangular bandpass H and K channels with a FWHM of 1.09 \AA . Because the two Mount Wilson spectrometers had different bandpasses defined for the four channels, Duncan et al. derived a transformation from the HKP-1 measurements (referred to as “ F -values”) to the HKP-2 S -values.

The S -values were constructed as

$$S = \alpha \frac{H + K}{R + V}, \quad (1)$$

where H, K, R , and V refer to the flux in the corresponding bandpasses and α was calculated to be 2.4 to make the mean S correspond to the mean F determined from the HKP-1 observations.

The differences in the continuum regions resulted in a transformation being necessary from S to F :

$$F = 0.033 + 0.9978S - 0.2019S^2. \quad (2)$$

We follow a similar prescription to extract measurements of activity from our spectra and transform those values into S -values on the Mount Wilson scale.

4.2. Reduction of Spectra and Calibration of S

4.2.1. The Planet Search Reduction Pipeline

For all Planet Search spectra, extraction from raw CCD images is performed in an automated pipeline. We measured S -values from these archival, reduced spectra.

We apply a scattered light subtraction to the HIRES echellograms before extraction. HIRES echellograms have many pixels of interorder real estate, from which we can make good measurements of scattered light. We fit B -splines to the signal in these interorder regions and interpolate linearly between them to estimate the scattered light in each order.

For both Lick and Keck data, a cosmic-ray removal algorithm removes the strongest cosmic rays from each two-dimensional echellogram before extraction. This is performed by modeling the profile of each order in the spatial direction by averaging over a suitably large region in the wavelength direction. Having determined the spatial profile in a region of the echellogram, cosmic rays are identified as extreme excursions from the mean profile. This technique is nearly identical to the technique used in optimal extraction (Horne 1986, §§ IID and IIE).

One difficulty in extracting S -values from Planet Search program spectra is that the spectra are not flux calibrated. Further, to properly account for photon statistics, the blaze function and throughput of the spectrometers are not removed.

4.2.2. Extraction of S from Keck Observations

We extracted S -values from our Keck spectra following the prescription of Duncan et al. (1991) as closely as possible. To remove the blaze function, which the standard extraction retains, as noted above, we smoothed and normalized a representative flat-field spectrum and divided it out of all spectra in the region of interest. This replaced the blaze function, which is a strong function of position along each spectral order, with the much more slowly varying continuum of the quartz lamp used in our flat-field images.

We simulated the measurement of the Mount Wilson spectrometers by summing the counts within four wavelength bins. We defined two triangular, 1.09 \AA FWHM bandpasses centered on the H and K lines, and two fixed continuum channels 20 \AA wide, just as in the case of the HKP-2 spectrometer. One difference we employed was to fix the continuum regions in the observer’s frame rather than shift them into the star’s frame. This prevented stars with particularly large

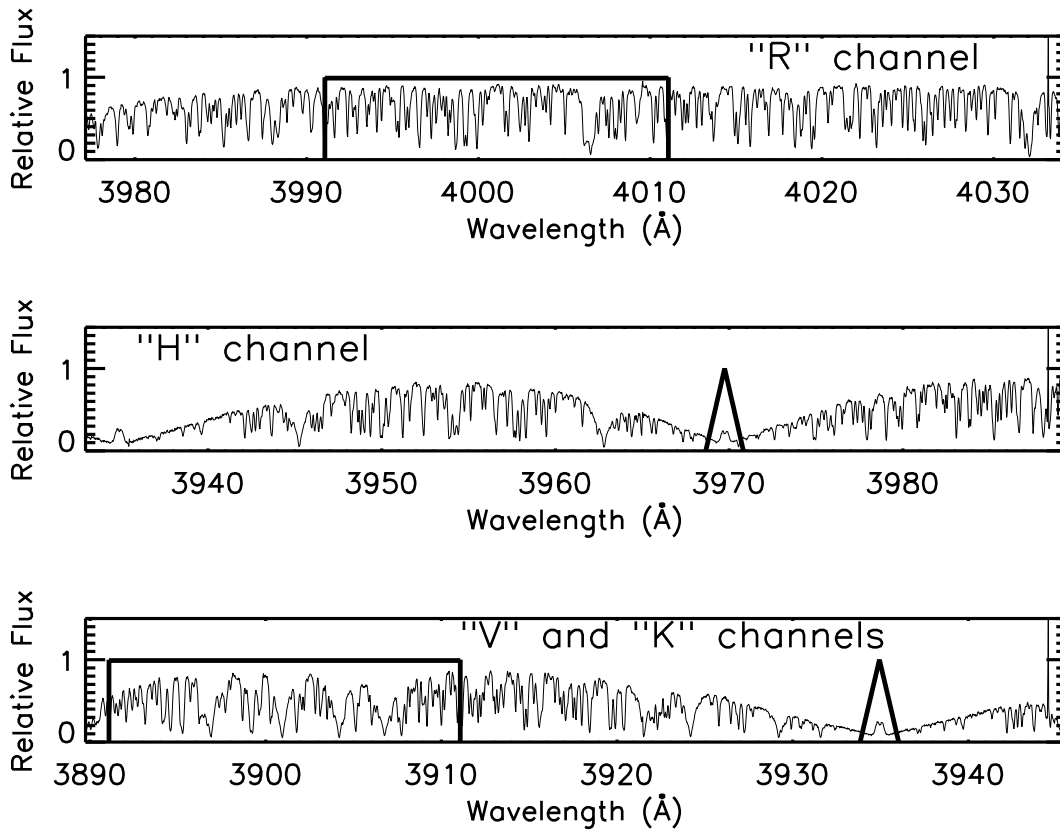


FIG. 1.—*R*, *H*, *K*, and *V* channels in a representative Keck spectrum. The ordinate is relative photon flux in arbitrary units. Wavelength is in the rest frame of the star. The *H* and *K* channels are always centered on the line cores; the *R* and *V* channels are fixed in the observer's frame.

Doppler velocities from shifting the channel into the next order, and it allowed us to correct for the effects of the imperfect flux calibration, as discussed below. The effect of choosing this frame, rather than the stellar frame, was extremely small, causing changes in S of less than 1%.

Figure 1 shows the positions of the *R*, *H*, *K*, and *V* channels in one of our Keck spectra. Because we do not flux calibrate our spectra, the relative fluxes of neighboring orders is not correct and the small effect of dividing out a quartz-lamp flat-field remains. Because our continuum channels are fixed in the observer's frame, these spurious effects are essentially a constant function of the spectrometer, not of the object being observed. As a result, the relative fluxes calculated by integrating the fluxes within each of the four bandpasses is different from the proper value by a constant multiplicative factor.

Rather than model and calculate these factors, we elected to fit for them and solve for any additional calibrations required to match Mount Wilson S -values without invoking additional degrees of freedom. We thus constructed S -values from our Keck spectra as

$$S = \frac{aH + bK}{cR + dV}, \quad (3)$$

where a , b , c , and d are relative weights to be determined. We found that there are 114 suitable stars in our sample that have S -values published in Duncan et al. and that have been observed more than once by each of our projects. We used unweighted averages of the Mount Wilson seasonal data to construct a mean S -value and performed a nonlinear least-

squares fit in log space for our relative weights. We found a good solution over all spectral types:

$$S = \frac{1.68H + 0.585K}{0.497R + 1.72V}. \quad (4)$$

The addition of constant or $B-V$ terms did not improve χ^2_ν of our fit.

Not all of our extracted spectra were of sufficient quality for our purposes. The automated extraction pipeline occasionally failed to properly trace an order on the echellogram or it improperly extracted the background scattered light. We found that an excellent diagnostic of the quality of a spectra was to simply examine the ratio of the counts in the *H* and *K* channels. We modeled the dependence of the *H*/*K* ratio as a function of derived S with a spline, and rejected all points for which the *H*/*K* ratio differed from this model curve by a factor of 1.35 or more.

A second rejection was based on the empirical observation that a few very low- S /N spectra passed the ratio test but were clearly useless. We rejected all stars with fewer than 200 counts per pixel in the *V* channel. These cuts culled 759 of 15,274 spectra, bringing to total remaining Keck spectra to 14,515.

The results of the data rejection and calibration are shown in Figure 2, which shows Median S from Keck versus mean S from Duncan et al. for the 199 stars our programs have in common. The 13% scatter in Figure 2 is primarily due to the chromospheric variability intrinsic to the stars. The Mount Wilson measurements used for the Keck calibration were made between 1966 and 1983, and many of these stars simply

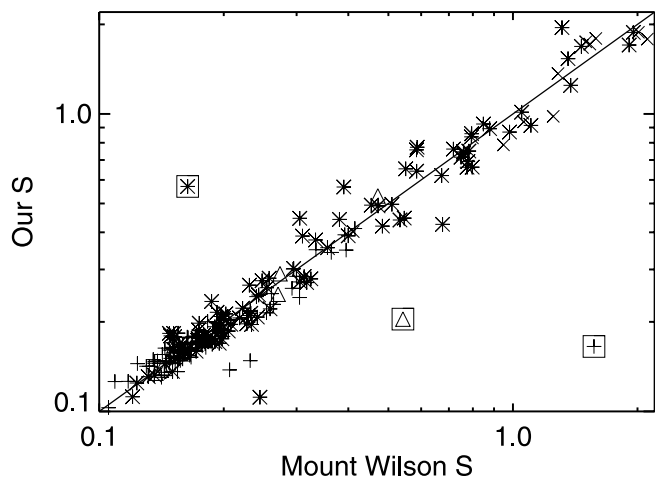


FIG. 2.—Median Keck S from this work vs. mean Mount Wilson S from Duncan et al. The scatter of 13% (excluding the boxed outliers) is due in large part to the intrinsic chromospheric variability of these stars. Triangles represent F stars, pluses G stars, asterisks K stars, and crosses M stars. We discuss the boxed outliers in § 5.4.

have different levels of chromospheric activity today than their means over that period.

4.2.3. Extraction of S from Lick Observations

We calibrated Dewars 6 and 8 from Lick independently because different extraction programs were used for the two Dewars. Data reduction for our Lick spectra proved more difficult than it was for the Keck spectra. One complication is the lack of use of an image rotator at Lick: differential refraction can impose a spurious slope on the flux spectrum which is impossible to calculate a priori and difficult to calibrate. Also, since most of the Lick spectra are from the CAT, the typical signal-to-noise ratio is considerably lower than for the Keck spectra.

Dewar 8 spectra posed additional difficulties as well. The automated extraction pipeline and observational plan for our Lick spectra were not optimized for the blue orders, where no Doppler information is gathered. This causes difficulty because the orders on the Hamilton spectrograph are closely spaced, making the extraction algorithm sensitive to any error in the assumed position of the orders which might arise from the low signal. The Dewar 8 reduction algorithm seems to have suffered along these lines, causing the continuum normalization to be uncertain by 10%. Further, the high-resistivity CCD in Dewar 8 yields a high incidence of cosmic-ray-like “worms,” which are difficult to remove and can contribute a significant fraction of the flux in the H and K and continuum channels. As a result, the Dewar 8 S -values are less precise than those of the other instruments.

To mitigate these problems we employed a simpler extraction algorithm for our Lick spectra. We simply defined a continuum region, which we denote the “C” channel, just redward of the H line and constructed a simple ratio of the H channel to the C channel (see Fig. 3). We thus defined our raw, uncalibrated Lick “ S -value” as

$$L = \frac{H}{C}, \quad (5)$$

where we use the symbol L to denote the fact that this is not a Mount Wilson S -value, but a ratio we have constructed related

to it. The proximity of the continuum region and its employment of fewer pixels reduced the severity of the problems outlined above.

As a check, we examined the dependence of the extracted L on the signal-to-noise ratio in the H and K region for a well-observed star (τ Ceti). The L -values for exposures with low signal were clearly discrepant from nominal values, which is probably a result of poor extraction or poor accounting of the flux zero point. Such problems are not unexpected, since the extraction and background subtraction algorithm was designed for and tested on the typically high-signal iodine region. To correct for this we rejected all data in the low signal regime, which we defined as spectra with a signal-to-noise ratio of less than 50 in the continuum. The effect of this strict rejection scheme is severe. Of 3014 spectra, only 1400 survived this cut.

Because there are very few stars in our Lick sample that have never been observed at Keck, we performed a secondary calibration against our Keck data. We found that to match the Keck S -values required quadratic transformations. For Dewar 6

$$S = 0.507L + 0.189L^2, \quad (6)$$

and for Dewar 8

$$S = 0.607L + 0.239L^2. \quad (7)$$

Figures 4 and 5 show how the final Lick S -values compare with the Keck measurements. The scatter in Figure 4 is considerably lower than that of Figure 2 because the measurements were contemporaneous.

4.3. Differential S -Values

It is possible to make measurements of the flux change in an emission line of a particular star that are much more precise than the absolute strength indicated by S , as measured above. Since changes in S can be important diagnostics of rotation and activity, we have also measured sensitive differential S -values, S_{diff} , for all of our Keck observations, inspired by the technique of Shkolnik, Walker, & Bohlender (2003). We made these measurements from the same reduced data as the S -values measured above, but with an independent technique, as described below.

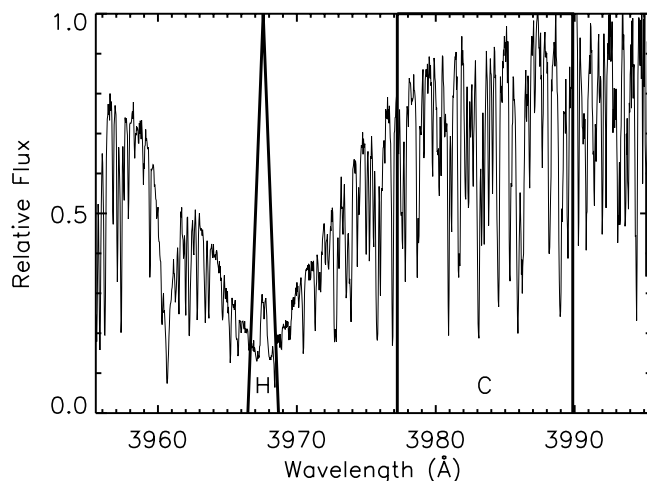


FIG. 3.—The H and C channels for a representative Lick spectrum. The ordinate units are arbitrary, the abscissa is in Angstroms.

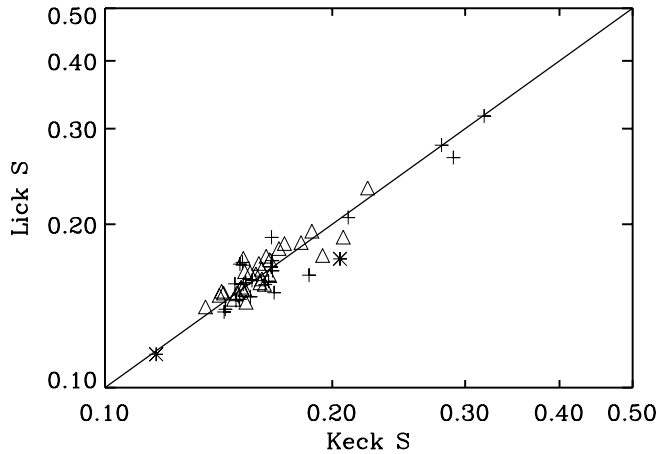


FIG. 4.— S measurements from Dewar 6 at Lick vs. Keck S . The rms scatter is 6%, some of which is due to the intrinsic variability of these stars. Triangles represent F stars, crosses G stars, and asterisks K stars.

For each star, we chose the observation from Keck with the highest signal-to-noise ratio as a template and reference spectrum. We scaled and shifted all other Keck spectra for that star such that we could directly compare the H line of every observation with the reference spectrum. When necessary, we added a small constant offset or slope to each spectrum to match the reference spectrum as closely as possible. We then defined E_i , the summed (scaled) counts in a 1 Å rectangular bandpass in observation i .

To transform E to the same scale as the S -values measured in § 4.2 we compared the fractional changes with time in E , E_i/E , to that in S , S_i/S , for those stars with more than nine Keck spectra and more than 2% variation in E , as shown in Figure 6. We fitted a single line to these points for all such stars using a robust line-fitting routine which iteratively rejects outliers using the more precise S_{diff} -values as the independent variable. The best-fit line which passes through (1,1) has a slope of 1.17. Therefore, we adopted the relation

$$S_{\text{diff},i} = \left(\frac{E_i}{\bar{E}} \right)^{1.2} \bar{S}, \quad (8)$$

where \bar{E} is the mean value of E_i for the star, and \bar{S} is the grand S for the star (see § 5.3).

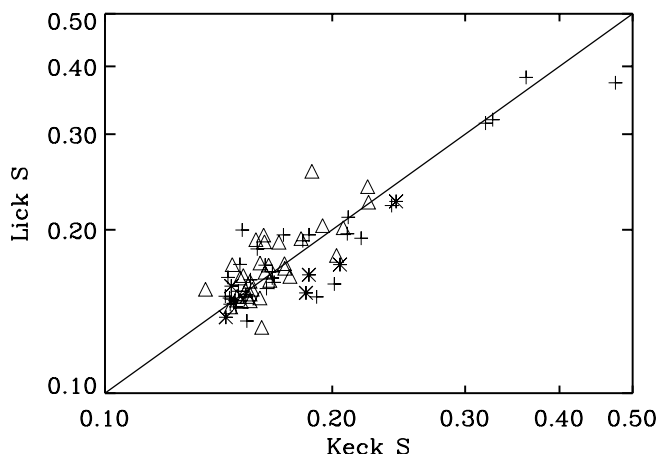


FIG. 5.— S measurements from Dewar 8 at Lick vs. Keck S . The rms scatter is 12%, some of which due to the intrinsic variability of these stars. Triangles represent F stars, crosses G stars, asterisks K stars.

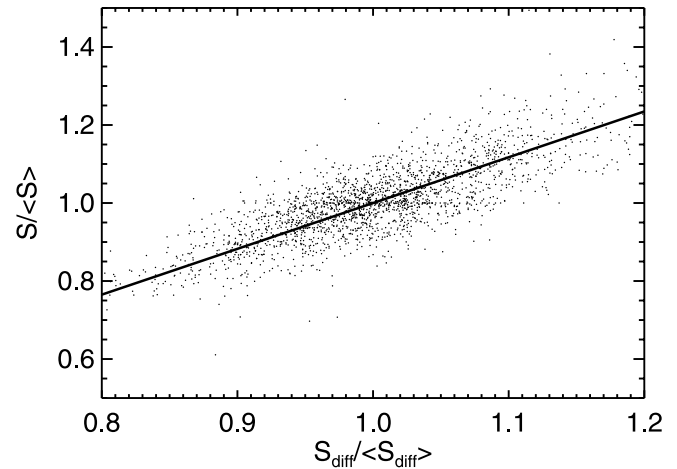


FIG. 6.—This figure demonstrates the calibration of the differential S -values (S_{diff}) to the scale of the less precise, absolute measurements of S . Each point shows the S and S_{diff} measurements from a single Keck spectrum, scaled by the median S and mean S_{diff} , respectively, for all such observations of a given star. 2203 observations of the 124 most chromospherically variable stars are represented here. The best-fit line is shown as a solid line and has a slope of 1.2. The 6% scatter of this distribution about the solid line is consistent with the estimate of the errors in S in § 5.1.1.

5. DATA FROM H AND K MEASUREMENTS

5.1. Uncertainties

Estimating uncertainties in our final values of R'_{HK} or other chromospheric-based quantities is difficult because they vary intrinsically with time on all timescales including those for rotation and stellar activity cycles. Activity cycles with periods longer than the duration of our observations will not be well measured; our final R'_{HK} -values for our sample therefore represent median activity levels during the time we observed them, and not true averages for the star.

Measurement errors in these R'_{HK} -values stem from the modest signal-to-noise ratio in the spectra and the quality of the calibration to the Mount Wilson S -value. We discuss below uncertainties from random errors from finite signal-to-noise ratios and short-term stellar variability. Calibration errors are negligible as a result of the large number of stars we used in the calibration.

The 13% scatter in Figure 2 is partly due to stellar variability, since our data are not contemporaneous with the Mount Wilson data and many stars are in different parts of their activity cycles. Thus, our quoted values of R'_{HK} carry uncertainties of no more than 13%; that is, they lie within 13% of the long-term average for the typical star. Measurements for stars observed more frequently and for the full duration of the Planet Search program will have correspondingly lower uncertainties.

5.1.1. Random Errors

To estimate the random errors in our S -values we used τ Ceti (=HD 10700, HR 509, $V = 3.5$) as a test case. τ Ceti serves as an excellent diagnostic star because we have a large number of observations of it at Keck and with both Dewars at Lick. We observe τ Ceti often because of the extraordinary velocity precision we can achieve for this star with short exposure times. This makes it an excellent source with which to search for any systematic errors in our precision velocities.

τ Ceti is also very well observed by the Mount Wilson project. Baliunas et al. (1995) note that despite its late spectral

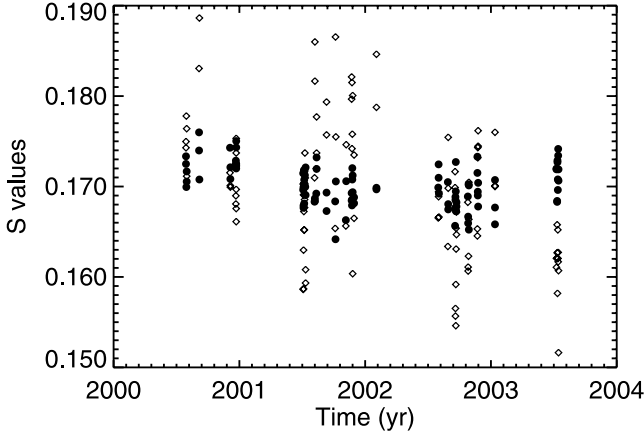


FIG. 7.— τ Ceti S -values (open diamonds) and differential S -values, S_{diff} (filled circles) from Keck. Activity variations apparent in the S -values are revealed to be due to uncertainties by the more precise differential values.

type and color, it exhibits only 1% variation in its S -values, suggesting that it may be in a “Maunder Minimum.” They also note that the rotation period (33 days) implied by its S -value and the observed $v \sin i$ (1 km s⁻¹) suggest that the star may be viewed nearly pole-on.

The standard deviation of all τ Ceti S -values is 6% for the Keck observations, 5.5% for Dewar 6, and 10.5% for Dewar 8. The differential S -values from Keck (§ 4.3) for this star (Fig. 7) have a standard deviation of 1.3%, which is consistent with the 1% variations reported by the Mount Wilson project. These estimates are consistent with the 6.5% scatter about the fit in Figure 6.

The disparity in the scatter in S and S_{diff} quantifies the errors introduced by difficulties in the raw reduction and extraction of S -values from the echellograms, as described in § 4.2. These include the difficulty of properly correcting for scattered light, properly removing the blaze function of the spectrometer, and properly extracting orders with modest signal-to-noise ratio. Much of the systematic component of these errors is removed by our calibration procedure, but a component that varies with time clearly remains. While the errors induced by these difficulties are not necessarily characterized by Gaussian noise, Figure 7 shows that they are averaged out over many observations.

Errors in S for other stars will be similar to those for τ Ceti because we use an exposure meter to ensure a uniform signal-to-noise ratio across our sample. This exposure meter is sensitive to light in the iodine region, so blue stars will have smaller Poisson errors in the Ca II H and K region than red stars.

We can calculate uncertainties in S_{diff} another way, as well. Occasionally during the course of the planet search, we take two or more consecutive exposures of a star. Under the assumption that the chromospheric activity of the star does not change over the course of several minutes, we can look at these sets of exposures and use the variation in the S_{diff} -values measured in these sets as an estimate of the precision of S_{diff} . Based on these sets we estimate a typical uncertainty in S_{diff} of 1.2%. This value, which is significantly higher than the Poisson noise, represents a reasonable estimate of the precision of the S_{diff} -values. This value is also consistent with the smallest variations in S_{diff} seen among stars in our sample.

There are also a small systematic (and therefore correlated) errors in the differential S -values, on the order of 1%. For

instance, our values of S_{diff} have a small and complex dependence on the focus of the spectrometer. This can be seen in Figure 7, where there appears to be a slight decrease in the S_{diff} -values for τ Ceti after 2001 caused by an improvement in the focus of HIRES at that time.

5.2. R'_{HK} , Ages, and Rotation Periods

The S index includes both chromospheric and photospheric contributions. To remove the photospheric component and determine the fraction of a star’s luminosity that is in the Ca II H and K lines, we follow the prescription of Noyes et al. (1984) to generate the $\log R'_{\text{HK}}$ -values which appear in Table 2. The transformation of S indices to R'_{HK} is a function of $B-V$ and is only calibrated for $0.44 < B-V < 0.9$.

The transformation used by Noyes et al. is that of Middelkoop (1982):

$$R_{\text{HK}} = 1.340 \times 10^{-4} C_{\text{cf}} S, \quad (9)$$

where

$$C_{\text{cf}}(B-V) = 1.13(B-V)^3 - 3.91(B-V)^2 + 2.84(B-V) - 0.47 \quad (10)$$

transforms the flux in the R and V channels to total continuum flux and S is the Mount Wilson S -value of the star. This number must be corrected for the photospheric contribution to the flux in the Ca II H and K line cores. Noyes et al. use the expression in Hartmann et al. (1984)

$$\log R_{\text{phot}} = -4.898 + 1.918(B-V)^2 - 2.893(B-V)^3 \quad (11)$$

to make the correction

$$R'_{\text{HK}} = R_{\text{HK}} - R_{\text{phot}}. \quad (12)$$

From these R'_{HK} -values one can derive rotation periods from the empirical fits of Noyes et al.:

$$\log(P_{\text{rot}}/\tau) = 0.324 - 0.400 \log R_5 - 0.283(\log R_5)^2 - 1.325(\log R_5)^3, \quad (13)$$

where R_5 is defined as $R'_{\text{HK}} \times 10^5$ and τ is the convective turnover time:

$$\log \tau = \begin{cases} 1.362 - 0.166x + 0.025x^2 - 5.323x^3, & x > 0, \\ 1.362 - 0.14x, & x < 0, \end{cases} \quad (14)$$

where $x = 1 - (B-V)$ and the ratio of mixing length to scale height is 1.9. Finally, we can calculate ages (Donahue 1993, cited in Henry et al. 1996):

$$\log t = 10.725 - 1.334R_5 + 0.4085R_5^2 - 0.0522R_5^3, \quad (15)$$

where t is the stellar age in years. The age calibration is certainly invalid in the T Tauri regime; therefore, in Table 2 for stars so active that this relation yields unreasonably low ages of $\log(\text{Age/yr}) < 7$ we simply quote “<7.”

Noyes et al. (1984) report that the rms in the calibration for equation (14) is 0.08 dex. Henry et al. (1996) notes that the age relation yields ages such that in 15 of 22 binaries where it has been tested the ages differ by less than 0.5 Gyr. On the

TABLE 1
 S_{diff} FOR PLANET SEARCH PROGRAM STARS

HD	HIP	Other	Date JD -2,400,000	S_{diff}	Notes
		Sol	50667.07115	0.168	Vesta
		Sol	50667.07810	0.165	Vesta
		Sol	50667.08616	0.167	Vesta
		Sol	50667.09427	0.165	Vesta
		Sol	50667.10113	0.166	Vesta
225261.....	400	GJ 3003	50366.92005	0.176	
225261.....	400	GJ 3003	50667.02786	0.172	
225261.....	400	GJ 3003	50689.03329	0.177	
225261.....	400	GJ 3003	50716.03115	0.171	
225261.....	400	GJ 3003	51010.02209	0.170	
225261.....	400	GJ 3003	51071.00666	0.175	
225261.....	400	GJ 3003	51072.00845	0.170	
225261.....	400	GJ 3003	51072.91727	0.177	
225261.....	400	GJ 3003	51173.80237	0.173	
225261.....	400	GJ 3003	51368.01875	0.175	

NOTE.—Table 1 is available in its entirety in the electronic edition of the *Astrophysical Journal Supplement*. A portion is shown here for guidance regarding its form and content.

other hand, during the solar cycle the Sun's age as calculated by the relation varies from 2.2 to 8.0 Gyr. Finally, this relation is probably only accurate for stars on the main sequence.

For all three of these values, R'_{HK} , P_{rot} , and age, the *mean* S -value of a star is the dispositive quantity and that many program stars have been observed only once or twice. Only for stars with many years of observation can we claim good knowledge of a mean value of S .

5.3. Tables of Measurements

We present two tables of S -values here. Table 1 contains S_{diff} -values from all of our Keck observations. The first three columns of these tables identify the star observed by HD number, *Hipparcos* number, and an “other” designation such as Gliese, HR, or SAO number. For some stars we have added binary component letters “A” and “B” to HD numbers of the brighter and fainter components, respectively, for uniqueness even when these qualifiers do not appear in the HD catalog. The fourth column specifies the Julian Date of the observation, and the fifth column lists the differential S -value (on the absolute Mount Wilson scale, as described in § 4.3) for that observation. The final column contains alternate names for some stars and coordinates for stars for which only one catalog name is given.

Table 2 contains our median, final S -value for each star in our sample, which we refer to as the “grand S .” To remove the effects of highly uneven sampling of stars which vary in activity, we took the median S -values in 30 day bins and used the median value of those medians. For simplicity, and to reduce the chance of breaking up observations taken within days of each other, these bins are not adjacent but are rather defined algorithmically such that some observation always lies at the beginning of a 30 day interval. For instance if observations occur on days 1, 25, 62, 63, 90, 91 and 99, then the bins would be from days 1 through 30, days 62 through 91, and days 99 through 128.

To calculate grand S -values we use only Keck and Dewar 6 spectra, if possible. For stars with only (less precise) Dewar 8 data, those are used and so noted. We combined all Dewar 6 and Keck observations for each of our stars and calculated 30 day medians. The median of all of these 30 day S -values we call the “grand S ” value of the star, and the standard deviation of the

differential S -values is quoted as a fractional uncertainty, $\sigma_{S_{\text{diff}}}/S$.

Table 2 contains 15 columns. The first three identify each star with the conventions described above for Table 1. The fourth and fifth columns list $B-V$ and M_V as reported by the *Hipparcos* catalog (Perryman et al. 1997) or, if unavailable, by SIMBAD. The sixth and seventh columns list the Julian dates of the first and last observations used in calculating the grand S -value. The eighth column lists the total number of observations, and the ninth column lists the number of monthly bins used to calculate the grand S -value. The tenth column contains the grand S -value as described above. The eleventh column contains the fractional standard deviation of the differential S -values for that star. Since the precision of S_{diff} is 1%, entries in this column near 1% represent stars without significant detected activity variations.

The next three columns list the quantities derived from these measurements and target notes, $\log R'_{\text{HK}}$, $\log (\text{Age}/\text{yr})$, and rotation period in days, as described in § 5.2. The final column contains target notes. The note “d8” refers to an entry based solely on Dewar 8 data, for which the uncertainty of each measurement contributing to the grand S is around 10%. All other entries are based on Keck and Dewar 6 data which have per-measurement uncertainties of around 6%. Again, alternate names for stars are noted, and J2000 coordinates are given for stars with only one catalog name.

5.4. Target Notes

5.4.1. The Sun

Our sample contains five observations of the asteroid Vesta obtained in 1997, about 1 year after solar minimum. We include these at the top of Table 2 under the name “Sol.” The S -value of 0.167 is consistent with solar minimum (Baliunas et al. 1995).

5.4.2. HD 531 A and B

The binary system of HD 531 consists of two stars of similar colors and magnitudes separated by $5''$. Since there seems to be confusion in SIMBAD regarding the properties of and nomenclature for these stars, we have deemed the eastern object HD 531B and the western one HD 531A, and we have not listed colors for these objects.

TABLE 2
MEASURED AND DERIVED QUANTITIES FOR ALL STARS

HD	HIP	Other	$B-V$	M_V	Begin (JD -2400000)	End (JD -2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
		Sol	0.656	4.83	50667	50667	5	1	0.166	0.70	-4.96	25.	9.69	Vesta; see § 5.4
224983.....	184		0.888	5.85	52833	52833	1	1	0.177	...	-5.01	46.	9.76	
225261.....	400	GJ 3003	0.755	5.78	50367	52836	32	18	0.174	1.36	-4.96	36.	9.68	
	428	GJ 2F	1.472	9.65	51757	52829	16	10	1.788	9.44	... ^a	... ^a	... ^a	
225213.....	439	GJ 1	1.462	10.36	52601	52832	2	2	0.451	16.21	... ^a	... ^a	... ^a	
38 A	473	GJ 4A	1.440	8.66	51757	52829	12	9	1.647	4.28	... ^a	... ^a	... ^a	
38 B		GJ 4B	1.410	...	51757	52829	19	10	1.820	6.12	... ^a	... ^a	... ^a	
105.....	490	SAO 214961	0.595	4.49	52516	52833	7	3	0.337	2.92	-4.41	5.	8.63	
283.....	616	GJ 9003	0.798	6.14	50367	52807	17	14	0.179	2.14	-4.96	39.	9.68	
377.....	682	SAO 109027	0.626	4.59	51343	52836	27	14	0.377	1.46	-4.36	4.	8.38	
400.....	699	HR 17	0.504	3.61	51014	51014	1	1	0.166	...	-4.87	8.	9.55	
531 B	795		51343	52829	18	10	0.462	3.83	Eastern of pair; see § 5.4
531 A	795		51343	52829	18	9	0.435	2.93	Western of pair; see § 5.4
691.....	919		0.755	5.29	52515	52830	8	3	0.446	3.97	-4.38	8.	8.45	V344 And
984.....	1134	SAO 128650	0.521	4.00	51014	51014	1	1	0.307	...	-4.42	3.	8.69	
	1368	GJ 14	1.370	8.11	52097	52829	9	7	1.979	6.65	... ^a	... ^a	... ^a	
1388.....	1444	GJ 9008	0.599	4.42	50366	52601	45	20	0.155	1.33	-5.00	19.	9.75	
1326 A	1475	GJ 15A	1.560	10.33	50462	52829	24	17	0.606	12.04	... ^a	... ^a	... ^a	
1326 B		GJ 15B	1.800	...	51052	52829	6	5	1.191	14.91	... ^a	... ^a	... ^a	
1461.....	1499	HR 72	0.674	4.62	50367	52829	18	17	0.156	1.75	-5.03	29.	9.80	
	1734	GJ 1009	1.485	9.86	52832	52832	1	1	1.801 ^a	... ^a	... ^a	
1815.....	1768	GJ 17.2	0.888	6.55	50367	50367	1	1	0.454	...	-4.51	17.	8.96	
1854.....		SAO 192490	0.610	...	50366	50984	12	6	0.172	1.69	-4.91	19.	9.60	
1835.....	1803	HR 88	0.659	4.84	50367	50367	1	1	0.342	...	-4.44	8.	8.78	BE Cet
1832.....	1813	SAO 73940	0.639	4.53	50806	52829	16	14	0.170	2.59	-4.93	22.	9.64	
2025.....	1936	GJ 18.0	0.940	6.64	50367	52807	14	13	0.348	5.28	... ^a	... ^a	... ^a	
2475.....	2237	HR 108	0.594	3.90	50367	50463	4	3	0.237	1.99	-4.64	10.	9.18	
2589.....	2422		0.883	3.20	52834	52836	3	1	0.127	1.64	-5.23	56.	10.07	
3074 A	2663	SAO 192609	0.617	3.62	50366	52829	15	13	0.154	3.02	-5.02	22.	9.78	
3079.....	2712	GJ 9016	0.549	4.01	51014	51174	5	2	0.152	0.81	-4.99	13.	9.73	
3141.....	2742		0.870	5.71	52833	52833	1	1	0.187	...	-4.97	44.	9.69	
3443.....	2941	HR 159	0.715	4.61	50367	50807	6	3	0.175	3.01	-4.94	31.	9.65	
		GJ 26	1.540	...	52102	52834	19	6	0.719	12.52	... ^a	... ^a	... ^a	LHS 119
3651.....	3093	HR 166	0.850	5.65	50367	52834	34	18	0.169	2.56	-5.02	45.	9.77	
3674.....	3119	SAO 92011	0.538	3.95	51014	51014	1	1	0.170	...	-4.88	11.	9.55	
3770.....	3169	SAO 128888	0.571	3.58	51014	51014	1	1	0.156	...	-4.98	16.	9.72	
3795.....	3185	HR 173	0.718	3.86	50366	52601	31	19	0.151	1.14	-5.07	35.	9.86	
3821.....	3203	GJ 9020A	0.620	4.96	51174	51174	1	1	0.303	...	-4.49	8.	8.92	
3765.....	3206	GJ 28	0.937	6.17	50462	52834	24	17	0.210	7.78	... ^a	... ^a	... ^a	
3861.....	3236	SAO 109369	0.528	3.90	51014	52160	2	2	0.183	...	-4.80	9.	9.43	
4208.....	3479	GJ 9024	0.664	5.21	50367	52833	41	20	0.170	1.72	-4.95	26.	9.66	
4203.....	3502	SAO 74235	0.771	4.24	51757	52806	19	10	0.136	1.84	-5.18	45.	10.00	
4256.....	3535	GJ 31.4	0.983	6.32	50367	52830	29	16	0.261	9.54	... ^a	... ^a	... ^a	
4271.....	3540	SAO 128932	0.537	3.94	51014	51014	1	1	0.163	...	-4.91	11.	9.61	
4307.....	3559	HR 203	0.603	3.63	50366	52830	38	16	0.143	2.03	-5.10	21.	9.89	18 Cet
4398.....	3607	HR 210	0.978	0.44	50367	50367	1	1	0.109 ^a	... ^a	... ^a	
4614 A	3821	HR 219	0.587	4.59	51171	51174	27	1	0.165	1.09	-4.93	17.	9.64	η Cas

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{Sdiff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
4747.....	3850	GJ 36	0.769	5.78	50367	52573	24	17	0.248	4.10	−4.72	26.	9.31	
4635.....	3876		0.900	6.09	52834	52834	1	1	0.338	...	−4.67	28.	9.23	
4915.....	3979	SAO 128986	0.663	5.26	51174	51174	1	1	0.186	...	−4.86	23.	9.53	
4903.....	4005	SAO 54203	0.559	3.65	51174	52160	2	2	0.166	...	−4.91	13.	9.61	
4967.....	4022	GJ 40	1.290	8.04	50366	50366	2	1	1.026	1.41	... ^a	... ^a	... ^a	
5035.....	4103		0.707	5.42	52833	52833	1	1	0.173	...	−4.94	30.	9.66	
5065.....	4127	SAO 54224	0.595	3.64	51014	52162	2	2	0.141	...	−5.11	20.	9.91	
5133.....	4148	GJ 42.0	0.936	6.41	51174	51174	1	1	0.407 ^a	... ^a	... ^a	
5372.....	4393	SAO 21850	0.667	4.47	50807	52835	20	15	0.189	2.38	−4.85	23.	9.51	
5470.....	4423	SAO 92167	0.631	4.20	52133	52602	14	6	0.154	1.75	−5.03	24.	9.80	
		GJ 47	51793	52834	7	4	1.025	3.63	LHS 1176
6101.....	4849	GJ 3071	1.008	6.50	51012	51070	3	2	0.500	4.31	... ^a	... ^a	... ^a	
	4856	GJ 48	1.501	10.49	50667	52834	14	11	0.769 ^a	... ^a	... ^a	
	4872	GJ 49	1.463	9.55	50667	52834	17	13	2.495	7.33	... ^a	... ^a	... ^a	
	5004	SAO 192901	0.764	6.31	51582	52833	9	9	0.177	2.20	−4.95	36.	9.67	
6558.....	5189	SAO 129121	0.606	3.86	52133	52833	10	6	0.155	2.80	−5.01	20.	9.76	
6611.....	5276	SAO 36943	0.541	3.74	51014	51014	1	1	0.157	...	−4.96	12.	9.68	
6660.....	5286	GJ 53.1A	1.122	6.83	50367	50367	1	1	0.773 ^a	... ^a	... ^a	
6734.....	5315		0.847	3.11	50367	52833	46	20	0.131	1.49	−5.20	53.	10.04	29 Cet
6715.....	5335		0.710	5.06	52833	52833	1	1	0.158	...	−5.03	33.	9.79	
	5476	SAO 22049	0.536	3.87	52096	52603	19	6	0.155	1.16	−4.97	12.	9.69	
6872 A	5480	SAO 22050	0.470	3.69	52096	52834	12	7	0.163	1.86	−4.86	6.	9.53	
6963.....	5521	SAO 36995	0.730	5.52	52515	52834	7	3	0.290	0.79	−4.60	18.	9.12	
7047.....	5534	SAO 109718	0.568	4.24	51014	51014	1	1	0.162	...	−4.94	15.	9.65	
7228.....	5682	SAO 54518	0.551	3.50	51014	52162	2	2	0.146	...	−5.04	14.	9.81	
7438.....		GJ 54.2B	0.780	...	50367	50367	1	1	0.275	...	−4.67	24.	9.22	
7483.....	5881	SAO 54557	0.671	5.01	51793	52308	12	6	0.202	7.87	−4.80	22.	9.43	
7661.....	5938	SAO 147702	0.753	5.43	52515	52602	5	2	0.412	2.62	−4.42	9.	8.67	
7590.....	5944	SAO 37069	0.594	4.72	51174	51174	1	1	0.278	...	−4.53	7.	9.00	
7727.....	5985		0.563	4.16	51014	51229	4	2	0.162	0.16	−4.94	14.	9.65	40 Cet
7895.....	6130	GJ 56.3A	0.780	5.79	50367	50367	1	1	0.256	...	−4.71	26.	9.29	
8038.....	6197	SAO 166958	0.701	4.79	52095	52601	14	7	0.260	5.54	−4.64	18.	9.19	
	6276	SAO 147747	0.791	5.71	52515	52836	6	3	0.529	1.96	−4.32	6.	8.01	
	6339	SAO 54624	0.916	6.17	52833	52833	1	1	0.487 ^a	... ^a	... ^a	
7924.....	6379	GJ 56.5	0.826	6.04	52188	52576	7	5	0.222	6.21	−4.83	35.	9.48	
8326.....	6390	GJ 3091A	0.970	6.29	50367	50367	1	1	0.367 ^a	... ^a	... ^a	
8262.....	6405	SAO 92398	0.627	4.92	51014	52159	2	2	0.170	...	−4.93	21.	9.63	
8331.....	6442	SAO 92406	0.681	3.86	51014	51014	1	1	0.141	...	−5.14	33.	9.96	
8389.....	6456	GJ 57.1A	0.900	5.45	50367	52833	26	20	0.184	11.36	−5.00	46.	9.74	
8328.....	6498	SAO 37167	0.691	3.80	52097	52807	10	7	0.131	2.74	−5.23	37.	10.07	
8467.....	6575	SAO 54666	1.100	5.71	52515	52835	6	3	0.279	3.23	... ^a	... ^a	... ^a	
8553.....	6613		0.912	5.89	52833	52833	1	1	0.172 ^a	... ^a	... ^a	
8574.....	6643	SAO 74702	0.577	3.90	51014	52834	11	8	0.144	0.81	−5.07	18.	9.86	
8648.....	6653	SAO 109878	0.675	4.40	50806	52833	31	17	0.153	0.97	−5.05	30.	9.82	
8673.....	6702	HR 410	0.500	3.43	51014	52159	2	2	0.200	...	−4.71	6.	9.29	
8765.....	6712	SAO 129296	0.707	3.76	52134	52574	11	5	0.185	2.28	−4.89	29.	9.57	
8941.....	6869	SAO 92453	0.526	3.10	51014	51014	1	1	0.195	...	−4.75	8.	9.34	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
8907.....	6878	SAO 37248	0.505	3.99	52515	52835	6	3	0.268	1.59	−4.50	3.	8.94	
8997.....	6917	GJ 58.2	0.966	5.92	50367	50367	1	1	0.567 ^a	... ^a	... ^a	
9070.....	6978	SAO 54737	0.710	4.78	52097	52836	10	7	0.195	4.36	−4.85	27.	9.51	
9280.....	7080	SAO 147853	0.760	4.22	51171	51174	6	1	0.145	0.87	−5.12	42.	9.92	
9224.....	7090	SAO 74767	0.622	4.15	51014	52162	2	2	0.168	...	−4.93	20.	9.64	
9312.....	7143	SAO 92492	0.929	2.95	52101	52101	1	1	0.267 ^a	... ^a	... ^a	
9331.....	7221	SAO 22381	0.710	4.87	51883	52602	11	6	0.151	2.41	−5.07	34.	9.85	
9540.....	7235	GJ 59	0.766	5.52	50366	52833	11	6	0.300	10.78	−4.60	20.	9.13	
9472.....	7244	SAO 74789	0.666	5.02	52515	52836	7	3	0.308	3.02	−4.51	10.	8.97	
9562.....	7276	HR 448	0.639	3.39	50367	52833	21	17	0.135	1.51	−5.19	28.	10.01	
9407.....	7339	GJ 59.1	0.686	4.91	50807	52835	13	12	0.160	1.35	−5.01	30.	9.76	
9770.....	7372		0.909	5.24	51043	51070	3	1	0.508	13.41	... ^a	... ^a	... ^a	BB Scl
9518.....		SAO 11842	0.480	...	52097	52834	11	7	0.132	2.12	−5.12	8.	9.92	
		SAO 11844	0.490	...	52128	52834	12	6	0.146	3.07	−5.00	8.	9.74	BD +60 271
9826.....	7513	HR 458	0.536	3.45	52112	52282	48	6	0.146	...	−5.04	12.	9.80	ν And
10002.....	7539	GJ 62	0.830	5.51	50463	52832	18	14	0.161	1.66	−5.04	45.	9.81	
9939.....	7564	SAO 74830	0.900	3.87	52101	52101	1	1	0.201	...	−4.94	44.	9.66	
9986.....	7585	SAO 92543	0.648	4.72	50806	52601	19	14	0.175	3.12	−4.91	23.	9.60	
10126.....	7733	SAO 74857	0.733	5.08	51174	51174	1	1	0.168	...	−4.98	34.	9.71	
10086.....	7734	GJ 65.1	0.690	4.95	51014	52161	2	2	0.275	...	−4.60	15.	9.12	
10145.....	7902	GJ 9059	0.691	4.88	50462	52834	23	15	0.160	1.66	−5.01	30.	9.76	
10476.....	7981	HR 493	0.836	5.87	50276	52835	268	25	0.186	3.42	−4.95	41.	9.66	
10486.....	8044	HR 495	1.020	2.61	52101	52101	1	1	0.201 ^a	... ^a	... ^a	
	8051	GJ 70	1.525	10.67	52134	52833	5	5	1.265	20.74	... ^a	... ^a	... ^a	
10436.....	8070	GJ 69	1.202	7.78	51369	52834	13	9	0.869	5.34	... ^a	... ^a	... ^a	
10700.....	8102	HR 509	0.727	5.68	51755	52836	394	15	0.168	1.35	−4.98	33.	9.71	τ Cet
10697.....	8159	HR 508	0.720	3.71	50367	52835	57	25	0.149	2.53	−5.08	36.	9.87	109 Psc
10853.....	8275	GJ 74	1.044	7.10	50367	50367	1	1	0.762 ^a	... ^a	... ^a	
11020.....	8346	GJ 76	0.795	6.04	50687	52573	14	11	0.169	2.05	−5.00	40.	9.74	
10780.....	8362	HR 511	0.804	5.64	51174	51174	1	1	0.271	...	−4.69	27.	9.27	
11226.....	8548	SAO 129526	0.569	3.90	51014	51174	5	2	0.150	0.51	−5.02	16.	9.79	
11507.....	8768	GJ 79	1.424	8.67	50463	51582	2	2	1.846	6.91	... ^a	... ^a	... ^a	
11505.....	8798	SAO 129551	0.635	4.55	51014	51014	1	1	0.155	...	−5.02	24.	9.78	
11373.....	8867		1.010	6.74	52834	52834	1	1	0.576 ^a	... ^a	... ^a	
11833.....	9035	SAO 110238	0.590	3.25	52129	52189	5	2	0.163	...	−4.94	17.	9.66	
11850.....	9073	SAO 75038	0.711	5.24	52515	52835	7	3	0.339	3.36	−4.49	11.	8.91	
11964 A	9094	GJ 81.1A	0.817	3.76	50366	52836	33	21	0.138	1.66	−5.16	49.	9.98	
12051.....	9269	GJ 82.1	0.773	5.19	50419	52836	22	16	0.157	1.76	−5.05	40.	9.82	
12235.....	9353	HR 582	0.610	3.43	50367	51014	8	5	0.152	2.34	−5.03	21.	9.79	ZI 106
12328.....	9406	SAO 129630	0.960	3.14	52101	52101	1	1	0.133 ^a	... ^a	... ^a	
		GJ 3126	1.540	...	52133	52574	3	3	1.108	5.40	... ^a	... ^a	... ^a	G 244−47
12414.....	9473	SAO 110286	0.510	3.72	50840	50840	1	1	0.157	...	−4.93	9.	9.64	
12661.....	9683	SAO 75125	0.710	4.58	50840	52836	52	16	0.150	1.44	−5.08	35.	9.87	
12786.....	9716	GJ 3133	0.832	5.75	50367	50367	1	1	0.541	...	−4.36	8.	8.32	
	9724	GJ 84	1.514	10.32	50840	52489	9	8	1.230	11.55	... ^a	... ^a	... ^a	
12846.....	9829	SAO 75144	0.662	5.06	51174	52652	20	8	0.163	1.07	−4.98	26.	9.71	
13043.....	9911	GJ 9073A	0.624	4.04	50367	52830	57	21	0.150	1.56	−5.05	23.	9.83	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
13357 A	10175	SAO 92824	52308	52652	6	5	0.218	1.68	BD +13 343B
13357 B	10175		52308	52652	6	5	0.282	4.19	
13382.....	10218	SAO 75191	0.980	4.66	52516	52835	7	3	0.324	3.37	... ^a	... ^a	... ^a	
	10279	GJ 87	1.431	9.96	51410	52574	10	8	0.544	13.08	... ^a	... ^a	... ^a	
13612 B	10303	GJ 87.1B	0.716	2.09	50367	52601	22	17	0.162	2.56	−5.00	33.	9.75	
13507.....	10321	SAO 37865	0.672	5.10	50840	51229	8	3	0.319	1.74	−4.49	10.	8.93	
13531.....	10339	SAO 37868	0.700	5.31	50840	50840	1	1	0.390	...	−4.40	7.	8.59	
	10449	SAO 129772	0.582	5.12	51411	52574	18	9	0.163	1.73	−4.94	16.	9.65	
13945.....	10492	SAO 167658	0.737	4.88	52129	52308	7	4	0.225	4.18	−4.76	26.	9.37	
13825.....	10505	SAO 75231	0.690	4.69	50840	50840	1	1	0.169	...	−4.96	29.	9.68	
13836.....	10510		0.705	5.27	52836	52836	1	1	0.290	...	−4.58	15.	9.09	
13579.....	10531	GJ 90	0.920	5.27	51174	51174	1	1	0.377 ^a	... ^a	... ^a	
14001.....	10542	GJ 91.2	1.033	6.20	50419	50419	1	1	0.528 ^a	... ^a	... ^a	8 Tri
13997.....	10599	SAO 92865	0.790	5.33	52133	52832	16	6	0.160	2.54	−5.04	41.	9.81	
13931.....	10626	SAO 37918	0.642	4.32	50807	52836	17	14	0.154	1.00	−5.03	25.	9.80	
13974.....	10644	HR 660	0.607	4.66	52516	52538	4	1	0.217	0.41	−4.71	13.	9.29	
14082 B	10679	SAO 75264	0.622	5.09	50840	50840	1	1	0.370	...	−4.37	4.	8.43	
14082 A	10680	SAO 75265	0.518	4.01	50840	50840	1	1	0.309	...	−4.41	3.	8.67	
14412.....	10798	HR 683	0.724	5.81	50366	52574	23	18	0.195	5.12	−4.85	29.	9.52	
14651.....	11028		0.720	5.27	52836	52836	1	1	0.154	...	−5.05	35.	9.82	
	11048	GJ 96	1.466	9.02	51582	52603	17	6	1.791	8.93	... ^a	... ^a	... ^a	κ For 13 Tri
14802.....	11072	HR 695	0.608	3.48	50462	51171	9	6	0.149	4.92	−5.05	21.	9.83	
15335.....	11548	HR 720	0.591	3.45	50367	52603	17	13	0.142	1.61	−5.10	20.	9.89	
15468.....	11565	GJ 100	1.120	7.40	50462	50462	1	1	0.883 ^a	... ^a	... ^a	
15814.....	11843	HR 741	0.572	3.71	50367	50715	3	2	0.181	6.76	−4.83	13.	9.48	29 Ari
16141.....	12048		0.670	4.05	50366	52835	70	23	0.145	1.44	−5.11	31.	9.91	79 Cet
16270.....	12110	GJ 1048	1.069	6.70	50366	50366	1	1	0.876 ^a	... ^a	... ^a	BX Cet
16160.....	12114	HR 753	0.918	6.50	50367	50367	1	1	0.266 ^a	... ^a	... ^a	
		GJ 105B	1.620	...	52830	52830	1	1	0.727 ^a	... ^a	... ^a	
16287.....	12158	GJ 9087	0.944	6.17	51012	52576	10	8	0.659	4.78	... ^a	... ^a	... ^a	
16275.....	12198	SAO 93028	0.665	4.30	52134	52603	16	5	0.163	2.10	−4.98	27.	9.72	
16397.....	12306	GJ 9089	0.583	4.59	50462	52575	12	10	0.161	1.15	−4.96	16.	9.68	
16548.....	12350	SAO 130036	0.702	3.36	52101	52101	1	1	0.137	...	−5.17	37.	10.00	
16623.....	12364	GJ 105.3	0.601	4.67	50689	52575	16	12	0.156	1.94	−5.00	19.	9.74	
16909.....	12709	GJ 106	1.074	6.89	50367	50367	1	1	0.890 ^a	... ^a	... ^a	
17037.....	12764	SAO 130084	0.534	3.85	50839	52239	17	11	0.157	1.55	−4.95	11.	9.67	
16895.....	12777	HR 799	0.514	3.85	52161	52219	2	2	0.152	...	−4.97	10.	9.70	
16895.....	12777	GJ 107	...	9.36	52238	52652	4	4	1.604	10.54	
	12781	GJ 109	1.530	11.18	51410	52576	11	8	0.886	9.31	... ^a	... ^a	... ^a	VY Ari
17152.....	12797	SAO 168014	0.772	5.17	52129	52237	5	3	0.158	1.75	−5.04	40.	9.81	
17190.....	12926	GJ 112.0	0.840	5.84	50367	52835	24	15	0.210	5.96	−4.87	38.	9.55	
17230.....	12929	GJ 112.1	1.269	7.58	50367	51174	2	2	0.915	5.46	... ^a	... ^a	... ^a	
17332 A	13027	SAO 93105	0.680	4.33	51411	51900	10	4	0.387	2.87	−4.39	6.	8.53	
17382.....	13081	GJ 113.0	0.820	5.81	50367	50367	1	1	0.439	...	−4.45	12.	8.79	
17433.....	13118		0.956	3.72	52101	52101	1	1	1.951 ^a	... ^a	... ^a	
17660.....	13258	GJ 114	1.194	7.12	51227	52575	14	10	1.014	5.64	... ^a	... ^a	... ^a	
18144.....	13601	SAO 93185	0.749	5.33	51174	51174	1	1	0.203	...	−4.84	31.	9.49	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD -2400000)	End (JD -2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
18143.....	13642	GJ 118.2A	0.953	5.72	50367	52835	25	16	0.188	10.34	... ^a	... ^a	... ^a	
18262.....	13679	HR 870	0.478	2.73	52652	52652	1	1	0.140	...	-5.04	7.	9.80	
18445.....	13769	GJ 120.1C	0.960	5.79	51043	52652	18	12	0.415	6.20	... ^a	... ^a	... ^a	
18632.....	13976	SAO 110894	0.926	6.12	51012	52601	15	10	0.616	5.09	... ^a	... ^a	... ^a	
18907.....	14086	HR 914	0.794	3.47	51171	51173	5	1	0.143	0.78	-5.13	45.	9.94	ϵ For
18803.....	14150		0.696	4.99	50367	52834	34	18	0.185	5.10	-4.88	27.	9.56	51 Ari
18940.....	14230	SAO 75711	0.624	4.45	52516	52516	1	1	0.258	...	-4.60	11.	9.12	
19034.....	14241	GJ 121.2	0.677	5.34	50366	52835	28	16	0.180	1.83	-4.90	26.	9.59	
19019.....	14258	SAO 110937	0.520	4.27	52516	52602	5	2	0.211	2.72	-4.68	7.	9.24	
19467.....	14501	GJ 3200	0.645	4.48	50366	52835	28	18	0.156	1.30	-5.02	25.	9.78	
19308.....	14532	SAO 56178	0.672	4.21	50839	52575	12	9	0.156	2.63	-5.03	29.	9.79	
19632.....	14623	SAO 168331	0.678	4.89	52516	52516	1	1	0.319	...	-4.50	10.	8.94	
19373.....	14632	HR 937	0.595	3.94	52188	52575	9	5	0.153	1.12	-5.02	19.	9.77	ι Per
19668.....	14684	SAO 130311	0.790	5.43	52516	52602	5	2	0.476	1.31	-4.38	8.	8.45	
19994.....	14954	HR 962	0.575	3.32	52202	52517	12	4	0.173	0.69	-4.88	14.	9.55	94 Cet
19902.....	14976	SAO 56256	0.732	5.03	51074	51074	1	1	0.286	...	-4.60	18.	9.13	
20165.....	15099	GJ 9112	0.861	6.09	50366	52603	17	13	0.221	6.35	-4.86	38.	9.52	
20367.....	15323	SAO 56323	0.574	4.23	50839	50840	2	1	0.282	0.53	-4.50	6.	8.95	
20619.....	15442	GJ 135	0.655	5.09	50366	52601	21	13	0.193	4.55	-4.83	21.	9.48	
20727.....	15572	GJ 138.1A	0.690	4.93	50367	50367	1	1	0.164	...	-4.99	30.	9.73	
20675.....	15669	HR 1001	0.468	2.64	51899	52652	52	10	0.178	1.62	-4.78	5.	9.40	
21019.....	15776	HR 1024	0.702	3.36	50366	52601	19	13	0.147	1.11	-5.10	34.	9.90	
	15904		0.571	6.27	51411	52575	13	9	0.154	2.23	-4.99	16.	9.73	BD +11 468
21197.....	15919	GJ 141.0	1.153	6.96	50366	52601	17	10	0.743	9.81	... ^a	... ^a	... ^a	
21313.....	16107	SAO 56465	0.623	3.92	51793	52601	16	6	0.145	1.39	-5.09	24.	9.89	
21531.....	16134	GJ 142	1.337	7.89	50463	50463	1	1	1.606 ^a	... ^a	... ^a	
21774.....	16405	SAO 75971	0.680	4.59	52133	52576	9	5	0.150	1.60	-5.07	31.	9.86	
	16404		0.667	6.14	51171	52653	12	9	0.127	2.43	-5.27	35.	10.12	BD +66 268
21847.....	16517	SAO 56530	0.503	3.84	50839	52575	15	10	0.226	2.12	-4.62	5.	9.15	
22049.....	16537	HR 1084	0.881	6.18	52142	52543	13	4	0.447	...	-4.51	17.	8.95	ϵ Eri
21962.....	16538	SAO 93484	0.503	3.33	50839	50840	2	1	0.189	0.05	-4.76	7.	9.36	
22072.....	16641	HR 1085	0.891	3.00	50420	52833	16	11	0.132	1.45	-5.20	55.	10.03	
22282.....	16727	SAO 130589	0.789	5.02	52134	52652	15	7	0.169	2.28	-4.99	40.	9.73	
22484.....	16852	HR 1101	0.575	3.60	52146	52229	4	2	0.139	...	-5.12	18.	9.92	10 Tau
22713.....	17027	HR 1111	0.921	3.24	52652	52652	1	1	0.142 ^a	... ^a	... ^a	21 Eri
22879.....	17147	GJ 147.1	0.554	4.75	50366	52603	16	14	0.163	1.19	-4.92	13.	9.62	
22918.....	17183	GJ 3244	0.954	3.70	52652	52652	1	1	0.145 ^a	... ^a	... ^a	
23249.....	17378	HR 1136	0.915	3.74	50463	52602	20	14	0.136	1.59	... ^a	... ^a	... ^a	δ Eri
23356.....	17420	SAO 149134	0.927	6.36	51043	52536	12	9	0.345	6.25	... ^a	... ^a	... ^a	
23588.....	17544	GJ 154.1	1.012	6.52	50420	50420	1	1	0.581 ^a	... ^a	... ^a	
23439.....	17666	GJ 1064A	0.750	6.23	50463	52603	17	14	0.178	1.06	-4.94	34.	9.65	
23596.....	17747	SAO 39110	0.634	3.67	51901	51901	1	1	0.150	...	-5.06	25.	9.84	
		TYC 65- 1471-1	1.100	...	52516	52575	2	2	0.735	0.89	... ^a	... ^a	... ^a	03:48:58.70 +01:10:53.9
24040.....	17960	SAO 93630	0.653	4.16	50839	52576	14	10	0.147	1.85	-5.09	28.	9.88	
24213.....	18106	SAO 76286	0.583	3.79	50839	52652	14	10	0.148	1.92	-5.05	18.	9.82	
24365.....	18208	GJ 3254	0.840	2.66	50463	52603	14	11	0.136	1.33	-5.17	51.	9.99	
24496.....	18267	GJ 3255	0.719	5.23	50839	52576	15	11	0.192	5.06	-4.87	29.	9.53	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD -2400000)	End (JD -2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
	18280	GJ 156	1.366	8.10	50462	52576	18	12	1.158	8.50	... ^a	... ^a	... ^a	
24341.....	18309	SAO 24291	0.683	3.80	50463	52652	15	12	0.153	1.56	-5.05	31.	9.83	
24238.....	18324	GJ 155.2	0.831	6.20	50463	52652	20	14	0.176	0.91	-4.98	42.	9.72	
24727.....	18388	SAO 130824	0.500	3.62	50839	52576	13	10	0.147	1.87	-5.00	9.	9.74	
24892.....	18432	GJ 3258	0.748	3.94	50366	52601	29	18	0.148	1.48	-5.09	39.	9.89	
24916.....	18512	GJ 157A	1.115	7.08	50366	52576	13	10	0.947	3.02	... ^a	... ^a	... ^a	
25069.....	18606	HR 1232	1.001	2.65	50366	52308	19	13	0.122	1.38	... ^a	... ^a	... ^a	
24451.....	18774	GJ 156.2	1.132	7.18	51227	52575	12	7	0.925	7.56	... ^a	... ^a	... ^a	
25535.....	18824	SAO 194709	0.628	3.14	50366	50806	3	2	0.150	1.53	-5.05	24.	9.83	
25457.....	18859	HR 1249	0.500	3.96	52516	52602	5	2	0.317	1.24	-4.39	2.0	8.54	
25329.....	18915	GJ 158	0.863	7.18	50463	52601	13	11	0.193	5.37	-4.94	42.	9.66	
25682.....	19024	GJ 9142	0.769	5.01	52133	52574	10	5	0.163	1.11	-5.01	39.	9.77	
25790.....	19070	SAO 130932	0.752	3.00	52652	52652	1	1	0.179	...	-4.93	34.	9.64	
281540.....	19143		0.964	7.21	51412	52602	11	7	0.273	6.17	... ^a	... ^a	... ^a	
25825.....	19148	SAO 93760	0.593	4.50	52133	52602	13	5	0.297	2.72	-4.48	6.	8.90	
	19165	GJ 160.2	1.206	7.84	52601	52601	1	1	0.596 ^a	... ^a	... ^a	
26151.....	19232	SAO 169142	0.832	5.18	51171	52601	11	9	0.186	7.24	-4.95	41.	9.66	
25918.....	19301	SAO 39329	0.725	5.32	51584	51584	1	1	0.170	...	-4.97	33.	9.69	
25665.....	19422	GJ 161	0.952	6.37	51174	51174	1	1	0.284 ^a	... ^a	... ^a	
26161.....	19428	SAO 57026	0.550	3.88	50839	52537	12	8	0.147	1.34	-5.03	14.	9.80	
26767.....	19786	SAO 93830	0.640	4.78	51411	51975	9	5	0.314	3.68	-4.48	8.	8.90	
26794.....	19788	GJ 165.2	0.943	6.17	50420	52602	17	13	0.208	6.36	... ^a	... ^a	... ^a	
26965.....	19849	HR 1325	0.820	5.92	52236	52602	7	4	0.196	2.80	-4.90	38.	9.60	DY Eri
26990.....	19911	SAO 111707	0.661	4.79	52516	52602	6	2	0.247	0.68	-4.65	16.	9.20	
27466.....	20218	SAO 131119	0.650	5.05	52516	52602	5	2	0.250	3.43	-4.64	14.	9.18	
28187.....	20638	SAO 195001	0.629	4.69	50366	52601	15	12	0.162	1.58	-4.97	22.	9.71	
28005.....	20800	SAO 39540	0.711	4.35	50839	52652	21	12	0.151	1.78	-5.07	35.	9.85	
28388.....	20802	SAO 169470	0.752	2.76	50366	51171	8	6	0.156	1.56	-5.05	38.	9.82	
28237.....	20826		0.560	4.12	52516	52603	5	2	0.290	1.53	-4.48	5.	8.88	
28344.....	20899		0.609	4.45	51411	51584	7	3	0.314	4.38	-4.46	6.	8.84	V920 Tau
28343.....	20917	GJ 169	1.363	8.00	51227	52602	12	7	1.534	11.31	... ^a	... ^a	... ^a	
28447.....	21010	SAO 76634	0.722	3.52	52652	52652	1	1	0.142	...	-5.14	38.	9.95	
28676.....	21158	SAO 76645	0.641	4.12	51174	51174	1	1	0.155	...	-5.03	25.	9.79	
28946.....	21272	GJ 9158	0.779	5.78	51174	51174	1	1	0.235	...	-4.76	29.	9.37	
28495.....	21276		0.759	5.55	52537	52603	6	2	0.478	1.28	-4.34	6.	8.20	
237287.....	21433	GJ 171	0.902	5.99	50463	50463	1	1	0.159 ^a	... ^a	... ^a	
29150.....	21436	SAO 76668	0.685	4.94	51174	51174	1	1	0.180	...	-4.90	27.	9.59	
232979.....	21553	GJ 172	1.423	8.58	51544	52603	9	6	1.705	4.38	... ^a	... ^a	... ^a	
	21556	GJ 173	1.505	10.12	51581	52516	8	5	0.677	5.38	... ^a	... ^a	... ^a	
29461.....	21654	SAO 94049	0.655	4.55	51411	52601	15	8	0.284	1.73	-4.55	12.	9.05	
29528.....	21703	SAO 94057	0.830	5.16	51793	52601	13	6	0.155	2.01	-5.07	46.	9.86	
29587.....	21832	SAO 39690	0.633	5.03	50463	51073	5	4	0.186	2.69	-4.85	20.	9.51	
29836.....	21923	SAO 94078	0.677	3.95	51901	51901	1	1	0.142	...	-5.13	32.	9.94	
285968.....	21932	GJ 176	1.523	10.08	50840	52575	18	12	1.367	8.40	... ^a	... ^a	... ^a	
29883.....	21988	GJ 176.2	0.907	6.25	51174	51174	1	1	0.215 ^a	... ^a	... ^a	
30495.....	22263	HR 1532	0.632	4.87	51975	51975	1	1	0.259	...	-4.60	12.	9.12	58 Eri
30508.....	22319	SAO 112091	0.846	3.12	52652	52652	1	1	0.151	...	-5.10	48.	9.89	

TABLE 2—Continued

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
30562.....	22336	HR 1536	0.631	3.65	51584	51584	3	1	0.152	0.60	−5.04	24.	9.81	
30339.....	22429	SAO 24853	0.611	3.88	51793	52603	14	6	0.151	1.93	−5.04	21.	9.81	
30652.....	22449	HR 1543	0.484	3.67	52280	52280	2	1	0.214	...	−4.65	4.	9.19	ZI 311
30708.....	22576	SAO 57460	0.717	4.20	50839	52652	12	9	0.153	1.54	−5.06	35.	9.83	
30649.....	22596	GJ 9168	0.586	4.56	50463	52537	18	12	0.162	1.18	−4.95	17.	9.67	
30825.....	22633	SAO 57468	0.875	2.91	52652	52652	1	1	0.134	...	−5.19	54.	10.02	
30973.....	22715	GJ 2035	1.019	6.63	50420	50420	2	1	0.642	0.52	... ^a	... ^a	... ^a	
	22762	GJ 180	1.547	10.44	52220	52603	3	3	0.687	25.48	... ^a	... ^a	... ^a	
31253.....	22826	SAO 94193	0.583	3.48	50839	52576	22	14	0.141	1.66	−5.11	19.	9.91	
31560.....	22907	GJ 2037	1.072	6.86	50366	52603	10	9	0.537	11.21	... ^a	... ^a	... ^a	
31412.....	22919	GJ 9169	0.561	4.24	50839	52536	16	11	0.173	2.21	−4.87	13.	9.54	
31966.....	23286	GJ 182.1	0.673	4.02	50839	52654	17	12	0.141	1.56	−5.14	32.	9.95	
32450.....	23452	GJ 185A	1.430	8.66	50840	51073	3	2	1.400	3.06	... ^a	... ^a	... ^a	
32387.....	23550	GJ 3324	0.810	2.95	50366	51072	5	4	0.140	1.07	−5.15	47.	9.96	
32850.....	23786	GJ 3330	0.804	5.84	50366	52536	4	4	0.317	3.26	−4.60	21.	9.13	
32923.....	23835	HR 1656	0.657	3.91	51174	51174	1	1	0.155	...	−5.03	27.	9.79	104 Tau
33021.....	23852	HR 1662	0.625	3.89	50366	52603	19	15	0.145	1.33	−5.10	24.	9.89	13 Ori
32963.....	23884	SAO 76970	0.664	4.87	50806	52602	11	10	0.156	1.12	−5.02	28.	9.78	
	23932	GJ 190	1.520	10.43	51552	51900	4	3	0.965	6.27	... ^a	... ^a	... ^a	
33555.....	24130	HR 1685	0.984	2.82	52652	52652	1	1	0.123 ^a	... ^a	... ^a	
33793.....	24186		1.543	10.89	51171	52713	6	6	0.325	4.98	... ^a	... ^a	... ^a	VZ Pic
33636.....	24205	SAO 112506	0.588	4.71	50839	52603	25	13	0.180	2.32	−4.85	15.	9.51	
33632.....	24332	SAO 57754	0.542	4.41	50840	51230	6	2	0.171	3.96	−4.87	11.	9.54	
34101.....	24419	GJ 193	0.719	4.94	50366	52308	8	7	0.169	2.59	−4.97	32.	9.70	
34445.....	24681	SAO 112601	0.661	4.04	50839	52652	21	13	0.154	2.04	−5.04	28.	9.80	
34721.....	24786	HR 1747	0.572	3.98	50366	52601	18	14	0.149	1.45	−5.03	16.	9.79	
34411.....	24813	HR 1729	0.630	4.18	51898	52713	9	6	0.151	1.25	−5.05	24.	9.83	λ Aur
34579 A.....	24820	HR 1741	1.020	0.51	50366	50366	1	1	0.111 ^a	... ^a	... ^a	
34745.....	24864	SAO 112633	0.531	4.13	50839	52713	15	11	0.155	1.10	−4.96	11.	9.69	
34865.....	24874	GJ 3346	1.003	6.79	50420	50420	1	1	0.668 ^a	... ^a	... ^a	
34575.....	25094	SAO 25144	0.750	4.75	51544	52601	11	7	0.147	1.58	−5.10	40.	9.90	
35627.....	25388	SAO 132095	0.572	3.67	50839	52652	12	9	0.145	0.90	−5.07	17.	9.85	
35850.....	25486		0.553	4.16	52536	52603	5	2	0.445	1.89	−4.22	1.1	< 7	
35974.....	25490	SAO 195865	0.600	3.34	50419	52601	12	11	0.147	1.74	−5.06	20.	9.84	
35681.....	25580	SAO 58065	0.510	3.76	50840	51230	5	2	0.183	1.05	−4.79	7.	9.41	
36003.....	25623	GJ 204	1.113	7.08	50367	52713	17	13	0.367	12.69	... ^a	... ^a	... ^a	
35956.....	25662	GJ 3347A	0.582	4.40	50366	52188	15	12	0.162	2.10	−4.95	16.	9.66	
36308.....	25873	SAO 94610	0.814	5.64	51544	52713	9	6	0.442	3.77	−4.44	12.	8.76	
36395.....	25878		1.474	9.19	50420	52713	18	14	2.173	4.99	... ^a	... ^a	... ^a	WR 69
37213.....	26273	SAO 170581	0.707	3.60	50420	52713	23	14	0.150	1.18	−5.08	34.	9.86	
245409.....	26335	GJ 208	1.415	8.50	50420	52575	19	12	3.297	4.63	... ^a	... ^a	... ^a	
37124.....	26381	GJ 209	0.667	5.07	50420	52602	38	17	0.179	2.48	−4.90	25.	9.59	
37484.....	26453		0.404	3.39	52538	52713	6	3	0.252	0.89	... ^a	... ^a	... ^a	
37008.....	26505	GJ 3358	0.834	6.18	51174	51174	1	1	0.181	...	−4.96	42.	9.69	
37216.....	26653	SAO 25310	0.764	5.63	52536	52603	6	3	0.358	1.34	−4.50	14.	8.94	
37588.....	26689	SAO 94743	0.524	3.61	50840	50840	1	1	0.154	...	−4.97	11.	9.70	
37962.....	26737	SAO 196074	0.648	5.01	50367	52603	14	11	0.208	1.87	−4.77	19.	9.38	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
233153.....	26801	GJ 212	1.473	9.30	51581	52603	8	6	2.543	11.57	... ^a	... ^a	... ^a	γ Lep B
38207.....		SAO 170732	0.360	...	52538	52575	3	2	0.239	1.43	... ^a	... ^a	... ^a	
38392.....		HR 1982	0.940	...	52236	52603	6	3	0.533	3.98	... ^a	... ^a	... ^a	
38230.....	27207	GJ 217	0.833	5.77	50420	52713	20	14	0.173	2.96	−4.99	43.	9.74	
37006.....	27225	SAO 5630	0.722	5.45	52536	52713	6	4	0.393	2.69	−4.41	8.	8.66	
38529.....	27253	HR 1988	0.773	2.81	50419	52603	49	19	0.174	6.92	−4.96	37.	9.69	γ Lep B
38949.....	27417	SAO 170859	0.566	4.65	52538	52713	6	3	0.263	2.05	−4.55	7.	9.03	
38858.....	27435	HR 2007	0.639	5.01	50419	52538	18	13	0.168	0.84	−4.95	23.	9.66	
39156.....	27641	SAO 94924	0.828	3.14	52652	52652	1	1	0.135	...	−5.18	50.	10.01	
39715.....	27918	GJ 223	1.014	6.71	50420	51174	3	2	0.442	2.78	... ^a	... ^a	... ^a	
39881.....	28066	HR 2067	0.650	4.36	50366	52601	22	14	0.154	1.35	−5.04	26.	9.80	V1386 Ori
40397.....	28267	GJ 3377C	0.720	5.16	50839	52652	13	10	0.166	1.34	−4.98	33.	9.72	
40650.....	28634	SAO 40802	0.567	4.22	50840	50840	1	1	0.160	...	−4.95	15.	9.67	
41700.....	28764		0.517	4.22	52574	52602	4	1	0.291	0.93	−4.45	3.	8.81	
40979.....	28767	SAO 40830	0.573	4.13	50840	52713	25	8	0.234	2.52	−4.63	9.	9.17	
40647.....	28902		0.783	5.77	52537	52713	6	3	0.384	3.90	−4.48	13.	8.90	V1386 Ori
41593.....	28954		0.814	5.81	50839	50840	2	1	0.490	0.33	−4.39	9.	8.51	
	29277	GJ 226	1.514	10.62	52008	52220	2	2	0.876	17.56	... ^a	... ^a	... ^a	
42581.....	29295	GJ 229	1.487	9.34	50547	52603	17	13	1.675	8.08	... ^a	... ^a	... ^a	
42618.....	29432	GJ 3387	0.642	5.03	50366	52654	22	14	0.168	2.51	−4.94	23.	9.66	
43162.....	29568	HR 2225	0.713	5.26	50366	50366	1	1	0.400	...	−4.40	8.	8.57	71 Ori; d8: higher rand. err
43042.....	29650	HR 2220	0.430	3.58	52535	52607	12	2	0.171 ^a	... ^a	... ^a	
42250.....	29761	GJ 9205	0.776	5.38	50462	52713	18	15	0.163	1.79	−5.02	39.	9.77	
43745.....	29843	HR 2254	0.576	3.04	50366	52603	19	13	0.137	1.24	−5.14	19.	9.96	
43587.....	29860	HR 2251	0.610	4.27	50366	52004	15	11	0.156	1.02	−5.00	20.	9.75	
43523.....	30023	SAO 41016	0.561	4.31	50840	50840	1	1	0.214	...	−4.69	9.	9.26	ZI 520
43947.....	30067	SAO 95538	0.562	4.41	50839	52652	13	10	0.159	1.22	−4.95	14.	9.67	
44420.....	30243	SAO 133153	0.686	4.56	51544	52603	11	7	0.157	2.11	−5.03	30.	9.79	
45184.....	30503	HR 2318	0.626	4.65	50366	52574	19	14	0.167	2.66	−4.95	21.	9.66	
45067.....	30545	HR 2313	0.564	3.28	50419	52603	21	14	0.140	1.45	−5.10	16.	9.90	
44985.....	30552	SAO 95656	0.593	4.42	50839	52652	14	10	0.159	2.06	−4.97	18.	9.70	ZI 520
45588.....	30711	HR 2349	0.545	3.67	50366	52601	19	12	0.151	1.86	−5.00	13.	9.74	
45350.....	30860	SAO 59126	0.740	4.44	51544	52712	19	10	0.147	1.32	−5.10	39.	9.90	
45391.....	30862	SAO 59131	0.613	5.10	51174	51174	1	1	0.176	...	−4.89	18.	9.57	
46090.....	31083	SAO 113993	0.710	4.90	50839	52574	13	9	0.177	3.67	−4.93	30.	9.63	
46375.....	31246	SAO 114040	0.860	5.29	51070	52713	69	14	0.186	1.97	−4.96	43.	9.69	56 Aur
47391.....	31623	SAO 197008	0.703	4.92	50419	50548	4	3	0.166	5.26	−4.98	31.	9.72	
260655.....	31635	GJ 239	1.480	9.66	50840	52574	12	9	0.919	9.65	... ^a	... ^a	... ^a	
47157.....	31655	SAO 114147	0.735	4.70	51314	52713	19	11	0.154	1.87	−5.06	37.	9.83	
47127.....	31660	SAO 95907	0.725	4.65	50840	50840	1	1	0.160	...	−5.02	35.	9.78	
47752.....	32010	GJ 241	1.021	6.86	50462	51174	2	2	0.497	6.02	... ^a	... ^a	... ^a	56 Aur
48938.....	32322	HR 2493	0.550	4.31	50419	52603	17	13	0.156	1.64	−4.96	13.	9.69	
48682.....	32480	HR 2483	0.575	4.15	52349	52349	1	1	0.177	...	−4.85	14.	9.52	
49197.....	32702		0.546	4.05	52576	52653	3	2	0.310	0.71	−4.43	4.	8.72	
49736.....	32874	SAO 78795	0.600	4.14	50840	50840	1	1	0.154	...	−5.01	19.	9.76	
49674.....	32916	SAO 41390	0.729	5.05	51883	52713	37	9	0.211	4.61	−4.80	27.	9.43	56 Aur
50499.....	32970	SAO 197294	0.614	3.84	50419	52712	25	17	0.153	1.67	−5.02	21.	9.79	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
50281.....	32984	GJ 250A	1.071	6.88	50419	52653	22	13	0.694	7.51	... ^a	... ^a	... ^a	
50806.....	33094	HR 2576	0.708	3.99	50366	52712	22	15	0.147	1.28	−5.10	35.	9.89	
50639.....	33109	SAO 133840	0.566	4.04	50839	52574	13	10	0.156	1.11	−4.98	15.	9.71	
50635 B.....		GJ 9220B	0.720	5.64	50419	50419	1	1	0.315	...	−4.54	14.	9.02	
50554.....	33212	SAO 78855	0.582	4.38	50840	52653	20	7	0.161	1.49	−4.95	16.	9.67	
265866.....	33226	GJ 251	1.580	11.18	50784	52576	14	10	0.760	7.92	... ^a	... ^a	... ^a	
50692.....	33277	HR 2569	0.573	4.55	50463	52576	18	13	0.162	1.81	−4.94	15.	9.65	37 Gem
51219.....	33382	SAO 114659	0.690	4.88	50806	52574	16	11	0.164	1.83	−4.99	30.	9.73	
51419.....	33537	SAO 78921	0.620	5.02	51174	51174	1	1	0.180	...	−4.87	19.	9.54	
52265.....	33719	HR 2622	0.572	4.05	50839	52573	26	12	0.150	1.87	−5.02	16.	9.78	
52698.....	33817	GJ 259	0.882	5.89	51171	51171	1	1	0.378	...	−4.59	23.	9.11	
52456.....	33848	SAO 114791	0.863	5.90	51069	52574	13	9	0.290	8.67	−4.71	30.	9.29	
51866.....	33852	GJ 257.1	0.986	6.43	50462	52576	17	13	0.335	8.61	... ^a	... ^a	... ^a	
52919.....	33955		1.078	7.02	50419	52574	3	3	0.532	8.03	... ^a	... ^a	... ^a	
52711.....	34017	HR 2643	0.595	4.53	50546	52576	23	12	0.162	1.65	−4.96	18.	9.68	
53665.....	34239	SAO 134179	0.517	3.22	50839	52653	15	11	0.144	1.41	−5.04	11.	9.81	
54563.....	34608	HR 2692	0.880	3.11	50419	51071	5	5	0.134	1.77	−5.19	54.	10.02	
55575.....	35136	HR 2721	0.576	4.41	51174	51174	1	1	0.161	...	−4.95	16.	9.66	
56274.....	35139	GJ 3436	0.607	5.14	50419	52573	16	12	0.178	2.38	−4.87	18.	9.55	
56303.....	35209	SAO 115185	0.609	4.27	50839	52574	13	10	0.150	1.43	−5.04	21.	9.82	
56124.....	35265	SAO 59945	0.631	4.74	50840	52334	2	2	0.175	...	−4.90	21.	9.59	
	35519	SAO 115271	0.876	5.71	51171	52219	9	7	0.175	3.12	−5.01	46.	9.76	
	36208	GJ 273	1.573	11.97	50806	52574	13	10	0.844	15.14	... ^a	... ^a	... ^a	
58781.....	36249	SAO 96918	0.734	4.89	50839	52712	16	13	0.158	1.59	−5.03	36.	9.80	
58830.....	36322	SAO 60106	0.950	1.55	50419	50548	4	3	0.204	1.04	... ^a	... ^a	... ^a	See § 5.4
	36357	GJ 273.1	0.923	6.51	51174	51174	2	1	0.564	0.28	... ^a	... ^a	... ^a	
59747.....	36704	SAO 60168	0.863	6.21	51174	51174	1	1	0.565	...	−4.37	9.	8.43	
60491.....	36827	SAO 134849	0.900	6.19	51171	52574	11	7	0.546	2.87	−4.43	13.	8.74	
	36834	GJ 277.1	51581	52576	5	5	0.891	6.66	
61606.....	37349	GJ 282A	0.891	6.42	50419	52778	11	9	0.581	4.48	−4.39	10.	8.55	
63077.....	37853	HR 3018	0.589	4.45	50366	50548	3	3	0.159	3.53	−4.97	17.	9.70	171 Pup
61994.....	38018		0.712	4.79	52574	52601	3	1	0.241	0.91	−4.70	21.	9.27	
63754.....	38216	HR 3048	0.579	2.97	50419	52778	16	13	0.132	2.02	−5.19	20.	10.02	
63433.....	38228	SAO 79729	0.682	5.21	50840	50840	1	1	0.387	...	−4.39	7.	8.54	
64090.....	38541	GJ 1104	0.621	6.01	50462	52576	21	11	0.143	0.86	−5.11	24.	9.91	
64606.....	38625	GJ 292.2	0.739	6.01	50419	50419	1	1	0.176	...	−4.94	34.	9.66	
64324.....	38647		0.659	5.04	52574	52601	3	1	0.241	3.34	−4.67	16.	9.23	
64468.....	38657	GJ 292.1	0.950	6.26	50463	52004	14	10	0.187	1.17	... ^a	... ^a	... ^a	
62613.....	38784		0.719	5.39	52220	52576	4	2	0.199	1.24	−4.84	28.	9.49	
65277.....	38931	GJ 293.1	1.065	6.84	50462	52778	18	13	0.312	6.37	... ^a	... ^a	... ^a	
65486.....	38939		1.043	7.12	50462	52575	3	3	0.757	4.73	... ^a	... ^a	... ^a	
65430.....	39064	GJ 3471	0.833	5.86	50462	52576	28	14	0.169	1.38	−5.01	44.	9.77	
65583.....	39157	GJ 295	0.716	5.84	50419	52574	17	12	0.172	0.95	−4.95	32.	9.68	
66428.....	39417	SAO 135426	0.715	4.55	51883	52778	17	7	0.149	1.99	−5.08	35.	9.87	
67458.....	39710	GJ 296.1	0.600	4.76	50462	52712	20	12	0.162	1.35	−4.96	18.	9.68	
67228.....	39780	HR 3176	0.642	3.46	51314	52334	2	2	0.143	...	−5.12	27.	9.92	10 Cnc
66171.....	39822	SAO 6424	0.621	4.81	50462	52713	12	11	0.168	2.34	−4.94	20.	9.65	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{Sdiff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
66751.....	40015		0.567	4.21	52573	52603	4	2	0.165	1.44	−4.92	14.	9.62	ψ Cnc
67767.....	40023	HR 3191	0.825	2.61	50419	52712	26	14	0.187	11.61	−4.94	40.	9.65	
68017.....	40118	GJ 9256	0.679	5.10	50462	52712	35	17	0.176	1.32	−4.92	26.	9.62	
68168.....	40133	SAO 97640	0.667	4.69	50806	52778	12	9	0.155	0.92	−5.03	28.	9.80	
68978.....	40283	SAO 198958	0.618	4.54	50462	52778	19	14	0.168	2.89	−4.93	20.	9.64	
69076.....	40419		0.706	5.61	52574	52601	3	1	0.190	0.80	−4.87	28.	9.54	χ Cnc
	40501	GJ 2066	1.510	10.29	50806	52575	11	8	0.933	9.66	... ^a	... ^a	... ^a	
68988.....	40687	SAO 14494	0.652	4.35	51552	52713	24	11	0.154	2.46	−5.04	26.	9.80	
69830.....	40693	HR 3259	0.754	5.45	51975	51975	1	1	0.176	...	−4.95	35.	9.67	
69809.....	40761	SAO 97727	0.674	4.34	51551	52778	12	9	0.147	2.20	−5.09	31.	9.88	
69897.....	40843	HR 3262	0.487	3.84	52283	52283	1	1	0.164	...	−4.87	7.	9.55	See § 5.4
70516.....	41184		0.652	4.87	52573	52603	4	2	0.396	1.94	−4.35	5.	8.30	
70843.....	41226	SAO 97797	0.539	3.72	50840	50840	1	1	0.156	...	−4.96	12.	9.68	
71334.....	41317	GJ 306.1	0.643	4.86	50462	52575	28	11	0.161	1.36	−4.99	24.	9.73	
71479.....	41479	SAO 135935	0.646	4.06	50839	52778	13	9	0.149	2.32	−5.07	26.	9.85	
71148.....	41484	HR 3309	0.624	4.63	50840	50840	1	1	0.165	...	−4.95	21.	9.67	See § 5.4
71881.....	41844	SAO 26892	0.630	4.39	50806	52713	16	11	0.159	1.19	−5.00	23.	9.74	
72673.....	41926	HR 3384	0.780	5.95	50419	52778	20	16	0.179	3.46	−4.95	37.	9.66	
72528.....	41968	SAO 136028	0.525	3.82	51975	51975	1	1	0.146	...	−5.02	11.	9.78	
72659.....	42030	SAO 136045	0.612	3.91	50839	52806	17	10	0.154	2.05	−5.02	21.	9.78	
72760.....	42074	GJ 3507	0.791	5.63	50840	50840	1	1	0.472	...	−4.38	9.	8.49	55 Cnc
72954.....	42075	HR 3397	0.752	2.42	50419	50463	2	2	0.136	2.25	−5.17	43.	10.00	
72780.....	42112	SAO 97927	0.513	3.87	50840	52334	2	2	0.161	...	−4.91	9.	9.61	
73350.....	42333	GJ 9273	0.655	4.87	51174	51174	1	1	0.315	...	−4.49	9.	8.92	
73344.....	42403	SAO 80310	0.547	4.17	50863	51230	6	2	0.207	2.52	−4.71	9.	9.29	
73512.....	42418	SAO 116990	0.897	5.87	51171	51227	3	2	0.219	1.67	−4.89	41.	9.58	STF 1321B
72905.....	42438		0.618	4.86	52573	52601	3	1	0.349	1.83	−4.40	5.	8.62	
73668.....	42488	SAO 117000	0.610	4.50	50840	51230	14	3	0.177	1.83	−4.88	18.	9.56	
	42491	SAO 117001	0.810	5.57	50840	52806	5	3	0.289	1.93	−4.66	25.	9.22	
73667.....	42499	GJ 315.0	0.832	6.27	50462	52805	23	15	0.181	1.33	−4.97	42.	9.69	
74014.....	42634	SAO 136179	0.760	4.96	51174	51174	1	1	0.154	...	−5.06	40.	9.84	STF 1321B
74156.....	42723	SAO 117040	0.585	3.56	52008	52778	9	6	0.144	1.97	−5.08	19.	9.86	
75302.....	43297	SAO 117163	0.689	5.08	50840	50840	1	1	0.256	...	−4.64	17.	9.19	
75393.....	43299		0.536	4.18	52573	52713	4	2	0.307	2.45	−4.43	3.	8.73	
75332.....	43410	HR 3499	0.549	3.93	50863	51230	11	3	0.290	2.34	−4.47	4.	8.87	
75732 A.....	43587	HR 3522	0.869	5.47	52065	52805	21	8	0.165	3.11	−5.04	47.	9.81	STF 1321B
75782.....	43634	SAO 61111	0.609	2.80	51975	51975	1	1	0.134	...	−5.19	24.	10.02	
76218.....	43852		0.771	5.60	52573	52807	5	3	0.388	2.50	−4.46	12.	8.85	
76752.....	44089	SAO 80548	0.680	4.46	50863	50863	1	1	0.164	...	−4.98	28.	9.72	
76909.....	44137	SAO 117323	0.756	4.45	51544	52805	17	10	0.143	0.98	−5.13	42.	9.93	
77407.....	44458	SAO 61224	0.609	4.65	50863	50863	1	1	0.385	...	−4.34	3.	8.17	STF 1321B
79210.....	45343	GJ 338A	1.410	8.68	50608	52712	15	13	1.785	8.38	... ^a	... ^a	... ^a	
79211.....	120005		1.420	8.71	50608	52805	18	14	1.879	4.39	... ^a	... ^a	... ^a	
79555.....	45383	GJ 339	1.035	6.58	50462	51171	2	2	0.692	3.91	... ^a	... ^a	... ^a	
80133.....	45621	SAO 136740	0.869	5.20	51174	51174	1	1	0.488	...	−4.45	13.	8.80	
80367.....	45737	GJ 340.2	0.860	5.81	51171	52713	11	9	0.172	0.90	−5.01	45.	9.77	STF 1321B
80632.....	45839	GJ 340.3	1.163	7.19	50462	50462	1	1	0.639 ^a	... ^a	... ^a	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{diff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
80715.....	45963		0.987	5.76	50462	50462	1	1	1.084 ^a	... ^a	... ^a	BF Lyn
80606.....	45982		0.765	5.23	52008	52806	22	10	0.149	1.41	−5.09	41.	9.88	ZI 748
81133.....	45995	SAO 200301	0.557	4.37	50462	50862	7	4	0.186	2.36	−4.80	11.	9.43	
		G 161−29	1.400	...	51173	52601	7	5	0.204 ^a	... ^a	... ^a	LTT 3472
81809.....	46404	HR 3750	0.642	2.91	50419	50419	1	1	0.169	...	−4.94	23.	9.66	
82106.....	46580	GJ 349	1.002	6.68	50863	51230	4	3	0.662	2.39	... ^a	... ^a	... ^a	
233641.....	46639	SAO 27275	0.548	4.31	51883	52712	12	5	0.137	0.82	−5.12	15.	9.92	
82558.....	46816		0.933	6.50	51171	51581	5	4	1.568	1.85	... ^a	... ^a	... ^a	LQ Hya
82943.....	47007	SAO 155312	0.623	4.35	51975	52652	17	8	0.172	6.32	−4.92	20.	9.61	
	47103	GJ 357	1.572	11.14	50840	52713	2	2	0.425 ^a	... ^a	... ^a	
83443.....	47202	SAO 221348	0.811	5.04	51898	52712	37	8	0.216	2.62	−4.84	35.	9.49	
	47513	GJ 361	1.490	10.10	51581	52806	6	6	1.474	9.25	... ^a	... ^a	... ^a	
	47650	GJ 362	1.488	10.88	52713	52713	1	1	3.208	6.04	... ^a	... ^a	... ^a	
84035.....	47690	GJ 365	1.133	6.88	50462	52713	18	14	0.559	6.33	... ^a	... ^a	... ^a	
84737.....	48113	HR 3881	0.619	3.75	52334	52334	1	1	0.130	...	−5.23	27.	10.07	15 LMi
85488.....	48411	GJ 371	1.230	7.25	50463	50463	1	1	1.025 ^a	... ^a	... ^a	
85301.....	48423		0.718	5.20	52601	52807	5	3	0.361	2.99	−4.46	10.	8.83	
85725.....	48468	HR 3916	0.622	2.50	50419	51171	8	6	0.130	1.36	−5.24	27.	10.08	
	48714	GJ 373	1.438	8.90	51552	52805	12	8	2.060	4.30	... ^a	... ^a	... ^a	
86264.....	48780	SAO 155612	0.510	3.11	51975	51975	1	1	0.193	...	−4.74	7.	9.34	
86680.....	49060	SAO 81137	0.607	3.00	51884	52236	10	3	0.156	4.47	−5.00	20.	9.75	
86728.....	49081	HR 3951	0.676	4.50	52334	52334	1	1	0.152	...	−5.06	30.	9.84	20 LMi
86972.....	49127	GJ 3581	1.016	6.56	50463	50463	1	1	0.744 ^a	... ^a	... ^a	
87359.....	49350	SAO 137336	0.689	4.98	50806	52805	14	11	0.185	4.27	−4.88	27.	9.56	
87424.....	49366	SAO 155709	0.891	6.32	51171	52683	11	7	0.522	1.96	−4.44	13.	8.78	
87836.....	49680	SAO 61885	0.708	4.30	51314	52829	17	10	0.157	4.32	−5.03	33.	9.80	
87883.....	49699	SAO 61890	0.965	6.28	51174	51174	1	1	0.250 ^a	... ^a	... ^a	
88072.....	49756		0.647	4.70	50806	52806	18	13	0.163	1.24	−4.98	24.	9.71	
88218.....	49769	HR 3992	0.615	3.70	50419	52601	40	17	0.148	1.58	−5.06	22.	9.85	
88371.....	49942	SAO 81224	0.637	4.44	50463	52713	14	13	0.160	1.00	−4.99	24.	9.74	
	49986	GJ 382	1.487	9.82	50608	52805	19	13	2.031	6.16	... ^a	... ^a	... ^a	
88725.....	50139	GJ 9322	0.609	4.96	50463	52804	20	15	0.169	1.32	−4.92	19.	9.62	
88986.....	50316	HR 4027	0.635	3.93	50420	52713	17	14	0.148	1.90	−5.08	25.	9.86	24 LMi
89307.....	50473	SAO 99049	0.594	4.57	50863	52334	2	2	0.162	...	−4.95	18.	9.67	
89391.....	50478	GJ 3594	0.940	2.53	50462	52683	21	16	0.122	2.66	... ^a	... ^a	... ^a	
89221.....	50485	SAO 43276	0.925	3.72	52102	52102	1	1	0.153 ^a	... ^a	... ^a	
89269.....	50505	GJ 3593	0.653	5.09	50419	52713	21	15	0.169	1.29	−4.94	24.	9.66	
89319.....	50546	HR 4046	1.022	2.82	52102	52334	2	2	0.121 ^a	... ^a	... ^a	
89744.....	50786	HR 4067	0.531	2.78	52215	52346	12	3	0.158	...	−4.94	11.	9.65	
90156.....	50921	GJ 3597	0.659	5.20	50419	52804	16	14	0.169	1.13	−4.95	25.	9.66	
90125.....	50939	HR 4085	0.991	1.32	50419	52804	20	17	0.112	2.15	... ^a	... ^a	... ^a	
	51007	GJ 390	1.459	9.67	51581	52804	8	6	1.734	11.24	... ^a	... ^a	... ^a	
90711.....	51257	GJ 3603	0.810	5.36	50462	52804	22	16	0.169	8.29	−5.00	42.	9.75	
90722.....	51258	SAO 156003	0.724	4.35	51551	52804	14	10	0.142	2.10	−5.13	38.	9.94	
	51317	GJ 393	1.507	10.34	50608	52806	17	12	1.004	7.15	... ^a	... ^a	... ^a	
90905.....	51386		0.562	4.38	52573	52807	5	2	0.311	2.51	−4.43	4.	8.75	
		SAO 27668	1.340	...	52805	52805	1	1	1.296 ^a	... ^a	... ^a	GJ 394

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
90839.....	51459	HR 4112	0.541	4.28	52739	52739	3	1	0.173	...	−4.86	11.	9.53	36 UMa A; d8: higher rand. err
90875.....	51468	GJ 3606	1.150	7.00	51200	52805	12	10	0.593	7.41	
	51525	GJ 397	1.330	7.87	51200	52805	14	9	1.502	10.11	
91204.....	51579	SAO 99171	0.651	4.25	51552	52601	13	6	0.147	1.83	−5.09	27.	9.88	
91545.....	51781	SAO 81423	1.063	2.85	52102	52102	1	1	0.130	
91638.....	51784	SAO 137653	0.506	3.93	50863	51230	6	2	0.159	1.60	−4.92	9.	9.62	
91816.....	51884		0.855	5.37	51172	51228	3	2	0.454	1.36	−4.47	14.	8.86	LR Hya
92222 B.....		SAO 62181	0.790	...	52063	52713	9	5	0.370	2.78	−4.51	15.	8.96	
92222 A.....		SAO 62182	0.700	...	52063	52713	12	8	0.362	4.65	−4.44	9.	8.78	
92588.....	52316	HR 4182	0.880	3.57	52102	52102	1	1	0.146	...	−5.13	51.	9.94	33 Sex
92788.....	52409	SAO 137743	0.694	4.76	50863	52683	26	11	0.153	2.90	−5.05	32.	9.83	
92945.....	52462	GJ 3615	0.873	6.05	50462	51680	13	6	0.641	2.72	−4.32	7.	8.00	
92855.....	52498		0.565	4.51	52601	52807	5	4	0.308	4.59	−4.44	4.	8.78	
93745.....	52888	SAO 201782	0.596	3.76	50462	52806	23	16	0.148	3.06	−5.06	20.	9.83	
	52940		0.551	3.38	51171	52308	16	10	0.151	2.23	−5.00	14.	9.75	BD +13 2311B
	52942	SAO 99310	0.506	3.18	51171	52806	18	13	0.142	1.82	−5.04	10.	9.81	
94280.....	53196	SAO 137858	0.560	3.97	51975	51975	1	1	0.155	...	−4.98	14.	9.71	
94340.....	53217	GJ 9336	0.645	3.94	50462	50548	4	2	0.269	1.36	−4.58	12.	9.10	
94765.....	53486	GJ 3633	0.920	6.15	51174	51174	1	1	0.543	
95091.....	53618	SAO 201951	0.659	3.74	50462	50462	1	1	0.152	...	−5.05	28.	9.83	
		BD −10 3166	0.903	...	51172	52806	30	13	0.216	7.92	TYC 5503- 946−1
95128.....	53721	HR 4277	0.624	4.29	51171	51201	29	1	0.154	1.18	−5.02	23.	9.78	47 UMa
95188.....	53747		0.760	5.70	52602	52807	6	4	0.431	1.05	−4.40	9.	8.59	
	53767	GJ 408	1.525	10.92	50608	52829	16	12	0.890	10.39	
95650.....	53985		1.437	9.24	51581	52834	17	8	3.990	8.01	DS Leo
95735.....	54035	GJ 411	1.502	10.46	50603	52713	34	16	0.402	9.75	
	54211	GJ 412A	1.491	10.30	50606	52805	16	12	0.367	8.20	
96418.....	54347	SAO 81674	0.513	3.14	50863	50863	1	1	0.145	...	−5.02	10.	9.78	
96574.....	54383	SAO 99459	0.551	3.81	50863	50863	1	1	0.149	...	−5.02	14.	9.78	
96700.....	54400	HR 4328	0.606	4.42	50419	52806	22	15	0.160	1.94	−4.97	19.	9.70	
	54532	GJ 413.1	1.529	10.30	51552	52713	9	7	0.906	6.00	
97037.....	54582	SAO 138023	0.613	4.24	50863	51229	4	2	0.154	1.10	−5.02	21.	9.78	
97004.....	54614	SAO 43637	0.767	4.89	51174	51174	1	1	0.152	...	−5.07	41.	9.86	
97101 B.....		GJ 414B	1.570	...	52065	52829	4	4	1.491	5.24	
97101 A.....	54646	GJ 414A	1.255	7.93	50463	52805	19	15	1.192	7.70	
97233.....	54677	GJ 416	1.203	7.28	50463	50463	1	1	0.596	
97343.....	54704	GJ 3648	0.760	5.37	50419	52805	25	17	0.166	1.66	−5.00	37.	9.74	
97503.....	54810	GJ 418	1.179	7.43	50462	50462	1	1	1.056	
97561.....	54844		0.750	3.38	52102	52102	1	1	0.146	...	−5.11	40.	9.91	ZI 876
97658.....	54906	GJ 3651	0.845	6.12	50463	52805	21	17	0.195	4.03	−4.92	41.	9.63	
97584.....	54952		1.043	6.85	52308	52681	3	3	0.757	4.87	
98281.....	55210	GJ 423.1	0.732	5.58	50462	52806	17	14	0.176	3.57	−4.94	33.	9.65	
98388.....	55262	SAO 99552	0.507	3.73	50863	50863	1	1	0.152	...	−4.96	9.	9.69	
	55360	GJ 424	1.412	9.52	51552	52681	9	8	0.778	7.52	
98553.....	55363		0.594	4.89	52652	52807	5	3	0.154	1.44	−5.01	19.	9.76	
98697.....	55455	SAO 118806	0.523	3.52	50863	50863	1	1	0.148	...	−5.01	11.	9.77	
98618.....	55459	SAO 27996	0.642	4.71	50838	52681	13	11	0.159	1.35	−5.00	24.	9.75	

TABLE 2—Continued

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
98744.....	55508	SAO 62514	0.538	2.72	52065	52778	10	5	0.128	2.14	−5.21	15.	10.04	
98745.....	55509	SAO 62515	0.536	0.28	52065	52335	10	3	0.145	3.44	−5.04	13.	9.81	
99109.....	55664	SAO 138182	0.874	5.19	51172	52835	31	14	0.161	6.27	−5.06	48.	9.84	
99491.....	55846	HR 4414	0.778	5.25	50419	52804	31	20	0.208	7.37	−4.84	33.	9.49	83 Leo
99492.....	55848	GJ 429.0B	1.002	6.30	50462	52807	28	20	0.254	9.41	... ^a	... ^a	... ^a	
99610.....	55900	SAO 156690	0.722	4.70	51975	51975	1	1	0.174	...	−4.95	32.	9.66	
99995.....	56145	SAO 43786	0.912	3.40	52102	52102	1	1	0.140 ^a	... ^a	... ^a	
100180.....	56242	HR 4437	0.570	4.46	50419	52806	29	16	0.162	1.68	−4.94	15.	9.65	88 Leo
100167.....	56257	SAO 43797	0.617	4.63	50863	50863	1	1	0.206	...	−4.76	16.	9.36	
100238.....	56260	SAO 138251	1.050	2.92	52102	52102	1	1	0.119 ^a	... ^a	... ^a	
100623.....	56452	HR 4458	0.811	6.06	50419	52805	16	14	0.198	7.09	−4.89	37.	9.57	
	56528	GJ 433	1.489	10.03	51200	52804	8	7	0.830	6.45	... ^a	... ^a	... ^a	
101177.....	56809	HR 4486	0.566	4.45	51314	51314	1	1	0.163	...	−4.93	14.	9.64	
101206.....	56829	GJ 3678	0.980	6.74	50463	51551	9	7	0.686	5.15	... ^a	... ^a	... ^a	
101259.....	56830	HR 4489	0.821	2.34	50462	52804	24	15	0.131	1.31	−5.21	51.	10.04	
101472.....	56960		0.549	4.46	52652	52807	5	3	0.253	2.57	−4.56	6.	9.06	
101501.....	56997	HR 4496	0.723	5.41	52739	52739	3	1	0.309	...	−4.55	15.	9.05	d8: higher rand. err
101563.....	57001	HR 4498	0.651	3.35	50462	50956	6	5	0.152	1.52	−5.05	27.	9.83	
	57087	GJ 436	1.493	10.62	51552	52834	17	9	0.726	4.12	... ^a	... ^a	... ^a	
101959.....	57217	SAO 180141	0.552	4.42	52652	52807	5	3	0.158	1.27	−4.95	13.	9.67	
102071.....	57271		0.805	5.56	52652	52807	5	3	0.240	1.51	−4.77	31.	9.38	
102158.....	57349	GJ 27	0.622	4.47	50463	52806	19	15	0.156	1.62	−5.01	22.	9.76	
	57450		0.582	5.58	51173	52681	14	10	0.155	1.55	−4.99	17.	9.73	BD +51 1696
102357.....	57488	SAO 81988	0.521	4.00	50863	50863	1	1	0.284	...	−4.47	3.	8.86	
	57544	GJ 445	1.572	12.14	50863	51581	2	2	0.609 ^a	... ^a	... ^a	
102540.....	57574	GJ 3687	0.756	2.89	50462	50956	6	5	0.151	2.03	−5.08	40.	9.87	
102634.....	57629	HR 4533	0.518	3.48	50863	50863	1	1	0.154	...	−4.96	10.	9.69	
102870.....	57757	HR 4540	0.518	3.40	52288	52335	3	2	0.157	...	−4.94	10.	9.65	β Vir
102902.....	57759	SAO 202865	0.701	2.64	50462	51680	11	10	0.150	8.49	−5.07	33.	9.86	
	57802	GJ 450	1.477	10.10	50840	52712	16	9	1.619	5.67	... ^a	... ^a	... ^a	
103095.....	57939	HR 4550	0.754	6.61	51200	51314	5	2	0.200	1.38	−4.85	31.	9.51	CF UMa
103432.....	58067	GJ 452.3A	0.710	5.36	50462	52804	17	13	0.203	4.24	−4.82	26.	9.46	
103829.....	58318	SAO 28203	0.668	4.49	51900	52681	12	7	0.151	1.85	−5.06	29.	9.84	
103932.....	58345	GJ 453	1.128	6.95	50462	52804	21	14	0.486	17.52	... ^a	... ^a	... ^a	
104067.....	58451	GJ 1153	0.974	6.33	50462	52804	18	14	0.437	9.27	... ^a	... ^a	... ^a	
104304.....	58576	HR 4587	0.760	4.99	51975	51975	1	1	0.183	...	−4.92	35.	9.62	
104526.....	58698	GJ 454.3	1.173	−1.01	50463	50610	4	3	0.145	1.10	... ^a	... ^a	... ^a	
104556.....	58708	SAO 44005	0.860	2.92	50463	52835	22	18	0.126	1.27	−5.23	55.	10.07	
104800.....	58843	SAO 119191	0.585	5.23	51553	52713	10	8	0.163	1.27	−4.94	16.	9.65	
104860.....	58876		0.596	4.52	52778	52807	4	1	0.349	2.20	−4.39	4.	8.52	
104985.....	58952		1.029	0.74	52835	52835	1	1	0.076 ^a	... ^a	... ^a	
105113 B.....	59021		50806	51884	13	8	0.180	3.14	CD−32 8503B
105113 A.....	59021	SAO 203123	0.623	2.94	50462	52804	24	16	0.145	2.18	−5.09	24.	9.89	
105405.....	59175	SAO 62852	0.515	4.02	50863	50863	1	1	0.153	...	−4.96	10.	9.69	
105590.....	59272	GJ 9390A	0.666	4.62	50462	50956	6	5	0.173	1.88	−4.93	25.	9.63	
105618.....	59278	SAO 99967	0.710	4.51	51883	52804	17	7	0.145	1.93	−5.11	36.	9.91	
105631.....	59280	GJ 3706	0.794	5.53	50463	52806	17	14	0.289	5.40	−4.65	24.	9.20	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
106116.....	59532	GJ 458.1	0.701	4.78	50462	52804	18	15	0.156	2.20	−5.03	32.	9.80	
106156.....	59572	GJ 3715	0.792	5.47	50463	52804	17	15	0.239	6.00	−4.76	30.	9.37	
106252.....	59610	SAO 99998	0.635	4.54	50863	50863	1	1	0.163	...	−4.97	23.	9.70	
106423.....	59690	SAO 119282	0.616	3.90	51975	51975	1	1	0.148	...	−5.06	22.	9.84	
107146.....	60074	SAO 100038	0.604	4.76	51341	52804	8	5	0.383	3.04	−4.34	3.	8.17	
107148.....	60081	SAO 138714	0.707	4.46	51553	52806	24	12	0.157	2.61	−5.03	33.	9.80	
107213.....	60098	HR 4688	0.523	2.90	51584	52334	2	2	0.138	...	−5.10	12.	9.89	9 Com
107434.....	60239	SAO 203374	0.536	3.98	50984	50984	2	1	0.296	...	−4.45	4.	8.81	
107705.....	60353	HR 4708	0.567	4.09	50863	51229	4	2	0.167	3.39	−4.90	14.	9.60	17 Vir
	60559	GJ 465	1.574	11.57	52654	52713	2	2	0.308 ^a	... ^a	... ^a	
108510.....	60816	SAO 138799	0.584	4.43	50863	50863	1	1	0.197	...	−4.77	13.	9.38	
		G 60−6	0.920	...	51172	52804	13	11	0.207 ^a	... ^a	... ^a	LP 555−23
108874.....	61028	SAO 82344	0.738	4.58	51341	52834	34	15	0.150	1.81	−5.08	38.	9.87	
109358.....	61317	HR 4785	0.588	4.63	51551	52806	11	9	0.167	1.98	−4.92	16.	9.62	
109333.....	61329	GJ 3735	1.103	7.06	50463	50463	1	1	0.759 ^a	... ^a	... ^a	
110315.....	61901	GJ 481	1.109	7.13	50463	52806	22	17	0.347	7.86	... ^a	... ^a	... ^a	
110501.....	61995	SAO 63146	1.074	2.86	52102	52102	1	1	0.125 ^a	... ^a	... ^a	
110537.....	62039	SAO 138928	0.675	4.49	50806	52835	55	19	0.154	1.53	−5.05	29.	9.82	
111031.....	62345	GJ 3746	0.695	4.45	50463	52807	21	17	0.151	1.46	−5.07	33.	9.85	
111066.....	62349	SAO 82490	0.540	3.64	50863	50863	1	1	0.151	...	−5.00	12.	9.74	
	62452	GJ 486	1.563	11.82	52335	52804	3	3	0.652 ^a	... ^a	... ^a	
111261.....	62472	GJ 1164A	1.149	7.41	50463	50463	1	1	0.596 ^a	... ^a	... ^a	
111312.....	62505	GJ 1165	0.946	6.30	50463	52007	10	9	0.522	4.09	... ^a	... ^a	... ^a	
111395.....	62523	HR 4864	0.703	5.12	50463	51975	2	2	0.290	13.83	−4.58	15.	9.09	
111398.....	62536	SAO 100279	0.660	4.31	50863	50863	1	1	0.152	...	−5.05	28.	9.83	
111484 A.....	62596		0.568	4.45	50984	52829	19	14	0.211	3.48	−4.71	10.	9.28	
111484 B.....	62596		0.568	4.55	50984	52829	19	14	0.185	2.69	−4.81	12.	9.45	BD +04 2658B
111515.....	62607	GJ 3752	0.686	5.52	50463	52829	21	16	0.170	1.09	−4.95	28.	9.67	
111631.....	62687	GJ 488	1.409	8.33	51582	52806	12	7	1.793	6.06	... ^a	... ^a	... ^a	
112060.....	62904	SAO 100321	0.805	3.20	52102	52102	1	1	0.156	...	−5.06	44.	9.84	
112257.....	63048	SAO 82565	0.665	4.69	50806	52806	16	13	0.167	1.84	−4.96	26.	9.68	
113194.....	63618		1.216	7.12	52308	52804	6	3	0.701	7.96	... ^a	... ^a	... ^a	
114174.....	64150	GJ 9429	0.667	4.68	50463	52712	54	22	0.164	1.89	−4.98	27.	9.71	
114335.....	64264	SAO 204187	0.559	3.99	50984	50984	1	1	0.155	...	−4.98	14.	9.71	
114710.....	64394	HR 4983	0.572	4.42	52335	52450	7	2	0.198	...	−4.76	12.	9.36	β Com
114783.....	64457	GJ 3769	0.930	6.01	50984	52835	48	20	0.215	12.25	... ^a	... ^a	... ^a	
114729.....	64459	SAO 204237	0.591	3.96	50463	52778	44	21	0.147	1.55	−5.05	19.	9.83	
114946.....	64577	HR 4995	0.862	2.38	50463	52804	20	18	0.131	1.48	−5.21	54.	10.04	55 Vir
115383.....	64792	HR 5011	0.585	3.92	52268	52268	2	1	0.337	...	−4.40	4.	8.60	59 Vir; d8: higher rand. err
115404.....	64797		0.926	6.24	52308	52805	6	3	0.446	3.54	... ^a	... ^a	... ^a	
115589.....	64905		0.864	5.45	51200	52834	22	14	0.155	2.47	−5.08	48.	9.87	
115617.....	64924	HR 5019	0.709	5.09	52389	52738	7	3	0.156	...	−5.04	33.	9.80	61 Vir; d8: higher rand. err
115781.....	64956		1.123	0.91	50546	50546	1	1	0.906 ^a	... ^a	... ^a	BL CVn
116442.....	65352	GJ 3781A	0.780	6.04	50463	52806	26	18	0.180	1.66	−4.94	37.	9.66	
116443.....	65355	GJ 3782B	0.869	6.22	50463	52829	23	20	0.194	5.27	−4.94	42.	9.66	
116858.....	65574	GJ 511A	0.920	6.36	50546	50546	1	1	0.315 ^a	... ^a	... ^a	
117126.....	65708	SAO 139353	0.651	4.13	50863	50863	1	1	0.154	...	−5.04	26.	9.80	

TABLE 2—Continued

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
117176.....	65721	HR 5072	0.714	3.68	52334	52349	6	1	0.165	...	−4.99	32.	9.73	70 Vir
117207.....	65808	SAO 204517	0.724	4.67	50463	52805	33	22	0.152	1.77	−5.06	36.	9.84	
	65859	GJ 514	1.493	9.64	50546	52806	21	17	1.272	6.77	... ^a	... ^a	... ^a	
117635.....	65982	GJ 3788	0.781	4.87	50463	50955	5	4	0.181	2.27	−4.94	37.	9.65	
117936.....	66147	GJ 2102	1.026	6.65	50984	52806	16	14	0.565	6.79	... ^a	... ^a	... ^a	
	66459	GJ 519	1.391	8.87	51581	52807	10	9	1.468	5.63	... ^a	... ^a	... ^a	
118914.....	66621	SAO 63667	0.661	4.55	51756	52806	13	8	0.150	3.40	−5.07	28.	9.85	
119850.....	67155	GJ 526	1.435	9.79	50546	52807	18	15	0.787	9.88	... ^a	... ^a	... ^a	
120066.....	67246	HR 5183	0.630	3.90	50463	52807	23	17	0.139	1.82	−5.15	26.	9.97	
120136.....	67275	HR 5185	0.508	3.53	52334	52381	5	2	0.202	...	−4.70	6.	9.28	
120476.....	67422	GJ 528	1.110	6.37	51227	52806	14	11	0.736	12.59	... ^a	... ^a	... ^a	
120476 B	67422	GJ 528B	1.200	...	51227	52806	17	12	1.114	8.63	... ^a	... ^a	... ^a	
120467.....	67487	GJ 529	1.257	7.40	50546	52807	17	14	0.684	9.33	... ^a	... ^a	... ^a	
120690.....	67620	HR 5209	0.703	4.93	50463	50609	3	2	0.223	3.30	−4.75	23.	9.35	
121320.....	67904		0.687	5.29	52654	52807	5	3	0.195	0.46	−4.84	25.	9.49	
121560.....	68030	HR 5243	0.518	4.24	50863	50863	1	1	0.160	...	−4.92	10.	9.62	
122064.....	68184	HR 5256	1.040	6.47	51581	52834	19	12	0.235	5.32	... ^a	... ^a	... ^a	
122120.....	68337	GJ 535	1.160	7.16	50546	52806	16	15	0.642	5.15	... ^a	... ^a	... ^a	
122303.....	68469	GJ 536	1.461	9.67	51411	52806	10	9	1.064	17.34	... ^a	... ^a	... ^a	
122652.....	68593	SAO 63905	0.563	4.31	50863	52806	15	13	0.191	2.29	−4.78	11.	9.40	
122676.....	68634	GJ 3824	0.740	5.03	50863	50863	1	1	0.175	...	−4.95	34.	9.66	
123760.....	69160	SAO 100878	0.656	4.02	52095	52489	8	4	0.198	2.62	−4.81	21.	9.44	
124257 B			1.100	...	52063	52806	5	5	0.141	1.60	... ^a	... ^a	... ^a	BD +50 2054B
124257 A		SAO 29032	0.800	...	52063	52806	5	5	0.141	1.84	−5.14	46.	9.95	
124115.....	69340	HR 5307	0.479	3.11	51975	51975	1	1	0.189	...	−4.74	5.	9.34	
124106.....	69357	GJ 3827	0.865	6.11	50548	52804	18	14	0.340	5.33	−4.63	25.	9.17	
124292.....	69414	GJ 3830	0.733	5.31	50863	50863	1	1	0.169	...	−4.97	34.	9.71	
124694.....	69518	SAO 44946	0.530	4.26	50863	50863	1	1	0.266	...	−4.52	5.	8.98	
124642.....	69526	GJ 3833	1.064	6.84	51174	51174	1	1	0.877 ^a	... ^a	... ^a	
124755.....	69569	HR 5335	1.071	2.98	52101	52101	1	1	0.121 ^a	... ^a	... ^a	
125040.....	69751	HR 5346	0.504	3.71	50863	50863	1	1	0.255	...	−4.53	4.	9.00	
125184.....	69881	HR 5353	0.723	3.89	50277	52835	34	22	0.151	3.02	−5.07	36.	9.85	
125455.....	70016	GJ 544A	0.867	5.99	50276	52829	20	18	0.198	6.64	−4.93	42.	9.63	
	70218	GJ 546	1.275	7.77	50284	51174	3	2	1.371	19.68	... ^a	... ^a	... ^a	
126053.....	70319	HR 5384	0.639	5.02	50276	52806	47	20	0.168	1.42	−4.94	23.	9.66	
125968.....	70330	SAO 182489	0.658	3.95	52095	52128	6	2	0.187	0.59	−4.86	23.	9.52	
126532.....	70500		0.853	5.93	52834	52834	1	1	0.214	...	−4.87	39.	9.54	
126583.....	70557		0.750	5.39	52833	52833	1	1	0.183	...	−4.92	34.	9.62	
126614.....	70623	SAO 139932	0.810	4.64	51200	52777	28	16	0.142	2.75	−5.14	47.	9.95	
126961.....	70782	SAO 120481	0.549	3.98	50863	50863	1	1	0.162	...	−4.92	12.	9.63	
128642.....	70857		0.774	5.42	52834	52834	1	1	0.187	...	−4.91	35.	9.61	
127334.....	70873	HR 5423	0.702	4.50	50863	52452	4	3	0.152	...	−5.06	33.	9.84	
128165.....	71181	GJ 556	0.997	6.60	51174	51174	1	1	0.446 ^a	... ^a	... ^a	
	71253	GJ 555	1.633	12.38	52334	52805	4	4	0.884	4.69	... ^a	... ^a	... ^a	
128167.....	71284	HR 5447	0.364	3.52	52739	52796	4	2	0.220 ^a	... ^a	... ^a	d8: higher rand. err
128311.....	71395	GJ 3860	0.973	6.38	50984	52835	49	21	0.700	4.27	... ^a	... ^a	... ^a	
128428.....	71462	SAO 140035	0.755	4.20	50546	51707	51	13	0.147	1.78	−5.10	41.	9.90	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{diff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
129333.....	71631		0.626	4.95	52654	52834	6	3	0.530	4.20	−4.18	1.2	< 7	
129010.....	71774	SAO 182760	0.591	3.33	50546	52806	18	16	0.129	6.08	−5.24	23.	10.08	
129191.....	71803	SAO 140072	0.682	4.52	51583	52807	11	8	0.155	2.22	−5.04	30.	9.80	
	71898	GJ 9492	1.615	10.91	50840	52806	8	7	0.716	15.38	... ^a	... ^a	... ^a	
129814.....	72043	SAO 101175	0.636	4.41	50863	52712	17	11	0.165	2.02	−4.96	23.	9.69	
130004.....	72146		0.931	6.42	52334	52778	4	3	0.267	4.67	... ^a	... ^a	... ^a	
130087.....	72190	SAO 101198	0.612	3.78	51583	52806	11	8	0.148	1.48	−5.06	22.	9.84	
129926 B		GJ 561.1B	0.500	...	50277	50277	2	1	0.298	5.08	−4.43	2.	8.72	
130307.....	72312	GJ 3867	0.893	6.29	50276	52829	15	13	0.410	5.54	−4.56	21.	9.06	
130322.....	72339	SAO 140142	0.781	5.67	51756	52806	11	9	0.230	6.47	−4.78	30.	9.39	
130948.....	72567	HR 5534	0.576	4.59	52452	52452	2	1	0.285	...	−4.50	6.	8.94	
130871.....	72577	GJ 563.3	0.957	6.64	50276	52806	19	15	0.280	1.46	... ^a	... ^a	... ^a	
131023.....	72634	GJ 3868A	0.760	5.23	50547	50547	1	1	0.341	...	−4.53	15.	9.00	
131156 B.....	72659	GJ 566B	1.160	7.84	50276	50276	1	1	1.280 ^a	... ^a	... ^a	
131156 A	72659	HR 5544	0.720	5.41	52376	52376	2	1	0.437	...	−4.35	6.	8.30	ξ Boo; d8: higher rand. err
130992.....	72688	GJ 565	1.036	6.66	50546	52807	15	13	0.389	6.74	... ^a	... ^a	... ^a	
131117.....	72772	HR 5542	0.605	3.29	50463	52829	18	16	0.134	1.23	−5.18	23.	10.01	
131509.....	72830	GJ 3872	0.896	3.69	50547	52834	23	20	0.136	1.73	−5.18	54.	10.01	
	72944		1.500	10.15	52063	52804	9	6	6.110	8.46	... ^a	... ^a	... ^a	CE Boo
132142.....	73005	GJ 569.1	0.785	5.88	50546	52834	25	20	0.168	1.17	−5.00	40.	9.74	
131977.....	73184	HR 5568	1.024	6.86	50463	50608	12	3	0.504	2.22	... ^a	... ^a	... ^a	
132307.....	73245		0.780	5.55	52833	52833	1	1	0.179	...	−4.95	37.	9.67	
132173.....	73269		0.554	4.21	52516	52829	5	3	0.278	3.09	−4.50	5.	8.94	
132375.....	73309	HR 5583	0.509	3.37	50863	50863	1	1	0.153	...	−4.96	9.	9.68	
132425.....	73314		0.834	5.50	52833	52833	1	1	0.292	...	−4.68	27.	9.24	
132505.....	73321	SAO 101312	0.651	3.82	52095	52807	10	6	0.144	1.64	−5.11	28.	9.91	
132756.....	73449	SAO 120789	0.691	4.33	50863	50863	1	1	0.168	...	−4.96	29.	9.69	
133161.....	73593	SAO 101345	0.599	4.27	50863	50863	1	1	0.153	...	−5.02	19.	9.78	
133460.....	73700	SAO 83637	0.559	3.16	51975	51975	1	1	0.161	...	−4.94	14.	9.65	
133295.....	73754		0.573	4.53	52516	52829	5	4	0.297	3.36	−4.47	5.	8.87	
134319.....	73869		0.677	5.17	52515	52834	6	4	0.412	4.14	−4.35	5.	8.28	
134044.....	73941	HR 5630	0.537	4.00	50863	52452	7	3	0.159	...	−4.94	11.	9.65	
134083.....	73996	HR 5634	0.429	3.46	52797	52797	2	1	0.217 ^a	... ^a	... ^a	d8: higher rand. err
134353.....	74118		0.852	5.41	52833	52833	1	1	0.252	...	−4.78	34.	9.39	
134440.....	74234	GJ 579.2B	0.850	7.08	51229	52713	12	9	0.221	1.06	−4.85	37.	9.51	
134439.....	74235	GJ 579.2A	0.770	6.74	50548	52829	41	15	0.227	0.93	−4.78	29.	9.40	
		SAO 101438	0.743	...	50276	52834	58	26	0.157	1.61	−5.04	37.	9.81	GJ 3897B
135101.....	74432	HR 5659	0.680	4.42	50276	52834	31	21	0.145	1.48	−5.11	32.	9.90	
134987.....	74500	HR 5657	0.691	4.42	50277	52829	53	20	0.147	1.89	−5.09	33.	9.89	23 Lib
135599.....	74702	SAO 120922	0.830	5.96	50863	51174	5	2	0.392	2.88	−4.52	17.	8.98	
136118.....	74948	SAO 140452	0.553	3.34	50863	50863	1	1	0.156	...	−4.97	13.	9.70	
136274.....	74954		0.737	5.36	52834	52834	1	1	0.152	...	−5.07	37.	9.85	
	74995	GJ 581	1.600	11.57	51410	52829	13	9	0.528	5.41	... ^a	... ^a	... ^a	
136580.....	75039	SAO 45491	0.510	3.87	50863	52450	3	3	0.149	...	−4.99	10.	9.73	
136442.....	75101	HR 5706	1.062	3.50	50984	52834	37	19	0.129	1.63	... ^a	... ^a	... ^a	
136654.....	75104	SAO 64647	0.532	3.68	50863	52451	2	2	0.156	...	−4.96	11.	9.68	
136544.....	75158	SAO 140480	0.475	3.46	51975	51975	1	1	0.137	...	−5.06	7.	9.84	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{diff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
136713.....	75253	GJ 1191	0.970	6.28	50277	52834	36	22	0.323	14.47	... ^a	... ^a	... ^a	
136834.....	75266	GJ 1192	0.992	6.25	50276	52807	16	16	0.243	11.94	... ^a	... ^a	... ^a	
136923.....	75277	SAO 101515	0.804	5.64	50863	50863	1	1	0.239	...	−4.77	31.	9.38	
136925.....	75281	SAO 101514	0.656	4.58	50863	52834	19	12	0.157	1.44	−5.01	26.	9.77	
137510.....	75535	HR 5740	0.618	3.16	51975	52339	3	2	0.155	...	−5.01	22.	9.77	
137303.....	75542	GJ 3905	1.045	6.88	50548	50548	1	1	0.414 ^a	... ^a	... ^a	
138004.....	75676		0.676	4.94	52515	52807	5	3	0.173	0.66	−4.93	27.	9.64	
137763.....	75718	GJ 586A	0.788	5.40	50277	50277	1	1	0.173	...	−4.97	39.	9.71	
137778.....	75722	GJ 586B	0.868	5.98	50277	52835	15	14	0.570	4.38	−4.37	9.	8.43	See § 5.4
137812.....	75762	SAO 159285	0.821	3.73	52101	52101	1	1	0.143	...	−5.13	48.	9.94	
139813.....	75829		0.803	5.62	52515	52807	5	3	0.469	1.30	−4.40	10.	8.58	
138573.....	76114	SAO 101603	0.656	4.77	50863	50863	1	1	0.160	...	−5.00	26.	9.75	
138549.....	76200	SAO 206764	0.717	5.36	50547	52834	20	14	0.189	5.57	−4.88	29.	9.55	
138776.....	76228	SAO 140612	0.745	4.62	51311	52835	14	11	0.143	1.47	−5.12	40.	9.93	
139477.....	76315	SAO 16775	1.063	6.99	51227	52834	17	12	0.892	2.76	... ^a	... ^a	... ^a	
139323.....	76375	GJ 591	0.946	5.91	50546	52835	26	21	0.237	13.30	... ^a	... ^a	... ^a	
139341.....	76382	GJ 593A	0.906	5.09	50546	50610	3	2	0.202	2.32	... ^a	... ^a	... ^a	
139324.....	76398	SAO 64802	0.633	3.88	51975	51975	1	1	0.152	...	−5.05	24.	9.82	
139457.....	76543	SAO 101658	0.531	3.69	51341	52835	13	10	0.145	0.76	−5.03	12.	9.80	
140538.....	77052	HR 5853	0.684	5.03	52308	52835	7	4	0.197	1.45	−4.83	24.	9.47	ψ Ser
141004.....	77257	HR 5868	0.604	4.07	50602	52452	24	6	0.160	1.02	−4.97	19.	9.70	ZI 1157
141399.....	77301		0.770	4.41	52834	52835	2	1	0.137	0.40	−5.17	45.	9.99	
141103.....	77335	SAO 140761	0.512	3.52	50863	50863	1	1	0.146	...	−5.02	10.	9.78	
141272.....	77408	GJ 3917	0.801	5.79	50276	50276	1	1	0.476	...	−4.39	9.	8.53	
141937.....	77740		0.628	4.63	52516	52835	7	3	0.169	1.26	−4.94	21.	9.65	
142373.....	77760	HR 5914	0.563	3.60	50602	50610	11	1	0.139	1.10	−5.11	16.	9.91	χ Her
142626.....	77790		0.835	5.91	52834	52834	1	1	0.259	...	−4.75	31.	9.35	
142267.....	77801	HR 5911	0.598	4.86	50276	51975	4	3	0.181	1.84	−4.85	16.	9.51	
142229.....	77810	SAO 121238	0.627	5.04	51341	52516	12	8	0.325	2.36	−4.45	7.	8.81	
142860.....	78072	HR 5933	0.478	3.62	52451	52451	1	1	0.172	...	−4.82	6.	9.46	
143291.....	78241	GJ 9533	0.757	5.94	50276	52806	24	17	0.179	2.48	−4.94	35.	9.65	
143006.....			0.730	...	52516	52829	4	3	0.839	5.52	−4.03	0.5	< 7	V1149 Sco
143313.....	78259		0.995	3.47	50276	50276	1	1	1.693 ^a	... ^a	... ^a	MS Ser
143761.....	78459	HR 5968	0.612	4.18	50602	50610	10	1	0.145	1.32	−5.08	22.	9.87	ρ CrB
144287.....	78709		0.771	5.44	52834	52835	2	1	0.162	1.29	−5.02	39.	9.78	
144087.....	78738	GJ 9541A	0.750	5.18	50277	50277	1	1	0.267	...	−4.66	22.	9.22	
144088.....	78739	GJ 9541B	0.850	5.47	50277	50277	1	1	0.344	...	−4.60	23.	9.13	
144579.....	78775	GJ 611A	0.734	5.87	50546	52806	30	15	0.170	1.02	−4.97	34.	9.70	
144179.....	78842	GJ 9543A	0.818	4.73	50548	50984	12	4	0.194	1.37	−4.91	39.	9.61	
144253.....	78843	GJ 610.0	1.043	6.05	50547	51756	12	10	0.247	7.37	... ^a	... ^a	... ^a	
144585.....	78955	HR 5996	0.660	4.02	50547	52829	19	15	0.146	3.11	−5.10	29.	9.89	
145148.....	79137	HR 6014	0.988	3.51	52101	52101	1	1	0.135 ^a	... ^a	... ^a	
145435.....	79152	SAO 45913	0.530	3.86	51013	52451	2	2	0.151	...	−4.99	11.	9.73	
145229.....	79165	SAO 101968	0.604	4.86	51013	51013	1	1	0.294	...	−4.50	7.	8.93	
144988.....	79214	SAO 207405	0.596	2.85	50548	52829	16	14	0.139	1.63	−5.13	21.	9.94	
145675.....	79248		0.877	5.32	50546	52834	46	27	0.161	4.63	−5.06	48.	9.84	14 Her
	79308	SAO 45950	0.714	5.36	52833	52833	1	1	0.161	...	−5.01	33.	9.76	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
145934.....		SAO 102017	1.050	...	50606	52834	43	17	0.113	1.71	... ^a	... ^a	... ^a	
145958 B	79492	GJ 615.1B	0.800	4.87	50276	52712	35	20	0.182	2.43	−4.94	39.	9.66	
145958 A	79492	GJ 615.1A	0.764	4.75	50276	52712	33	18	0.179	4.26	−4.94	36.	9.65	
145809.....	79524	GJ 3944	0.617	3.72	50277	52829	19	15	0.149	2.10	−5.06	22.	9.84	
146362.....		HR 6064	0.640	...	50546	52834	21	17	0.229	4.38	−4.69	16.	9.26	
147231.....	79619	GJ 9551	0.722	4.78	50546	52538	33	16	0.177	2.31	−4.93	31.	9.64	
146233.....	79672	HR 6060	0.652	4.76	50284	52805	139	18	0.169	1.95	−4.95	24.	9.66	ZI 1223
147379 A	79755	GJ 617A	1.409	8.47	51680	52805	14	8	1.593	5.42	... ^a	... ^a	... ^a	
147379 B	79762		1.510	10.54	51707	52806	9	8	1.220	6.04	... ^a	... ^a	... ^a	EW Dra
147044.....	79862	SAO 65213	0.631	4.71	51013	51013	1	1	0.168	...	−4.94	22.	9.66	
146775.....	79946	GJ 9553	0.616	4.61	50548	52829	32	18	0.166	2.02	−4.94	20.	9.66	
147512.....	80218	SAO 141126	0.718	5.13	51013	51013	1	1	0.170	...	−4.96	32.	9.69	
147750.....	80262		0.724	5.43	52833	52833	1	1	0.178	...	−4.93	32.	9.63	
	80295		1.009	6.98	51311	52807	13	11	0.431	6.52	... ^a	... ^a	... ^a	BD −11 4126
147776.....	80366	GJ 621	0.950	6.73	50547	51314	2	2	0.392	6.61	... ^a	... ^a	... ^a	
	80459	GJ 625	1.591	11.04	50862	52806	14	12	0.612	18.39	... ^a	... ^a	... ^a	
148467.....	80644	GJ 626	1.253	7.59	50276	51314	2	2	0.716	10.83	... ^a	... ^a	... ^a	
148427.....	80687	SAO 159932	0.950	3.03	52101	52101	1	1	0.124 ^a	... ^a	... ^a	
	80824	GJ 628	1.604	11.95	51410	52805	20	14	0.892	11.75	... ^a	... ^a	... ^a	
150706.....	80902		0.607	4.84	52515	52807	5	3	0.249	3.61	−4.61	11.	9.14	
149200.....	81062	SAO 141237	0.533	3.64	51013	51013	1	1	0.141	...	−5.08	13.	9.86	
149652.....	81279	SAO 121722	0.523	3.68	51013	51013	1	1	0.143	...	−5.05	11.	9.82	
149661.....	81300	HR 6171	0.827	5.82	50602	52834	32	13	0.355	7.40	−4.57	20.	9.07	V2133 Oph
149724.....	81347	SAO 141271	0.783	4.12	52095	52807	10	6	0.139	1.54	−5.16	45.	9.98	
149806.....	81375	SAO 121731	0.828	5.57	51013	51013	1	1	0.222	...	−4.83	35.	9.48	
149933.....	81421		0.763	5.17	52833	52833	1	1	0.146	...	−5.11	42.	9.91	
150554.....	81662		0.591	4.41	52515	52834	6	4	0.186	0.92	−4.82	15.	9.46	
150433.....	81681	GJ 634.1	0.631	4.86	50276	52806	43	18	0.162	1.53	−4.98	22.	9.71	
150437.....	81767	SAO 184553	0.683	4.13	52095	52829	11	6	0.152	1.56	−5.06	31.	9.84	
151044.....	81800	SAO 30055	0.541	4.14	51975	51975	1	1	0.151	...	−5.00	13.	9.74	
151541.....	81813	GJ 637.1	0.769	5.63	50546	52834	37	23	0.168	1.62	−4.99	38.	9.73	
150933.....	81880	SAO 84562	0.573	3.99	51013	52451	2	2	0.153	...	−5.00	16.	9.75	
150698.....	81910	SAO 184581	0.674	3.36	50548	52829	17	13	0.135	1.48	−5.19	34.	10.02	
151090.....	81991		0.890	3.17	52101	52101	1	1	0.138	...	−5.17	53.	9.99	41 Her
151288.....	82003	GJ 638	1.310	8.15	50602	52807	22	10	1.248	7.32	... ^a	... ^a	... ^a	
151877.....	82267	GJ 639	0.821	5.85	50546	52806	16	15	0.208	7.42	−4.87	37.	9.53	
151995.....	82389	GJ 640	1.020	6.71	50276	52807	18	16	0.272	8.27	... ^a	... ^a	... ^a	
152446.....	82568	SAO 102445	0.529	3.42	51013	51013	1	1	0.157	...	−4.95	11.	9.67	
152391.....	82588		0.749	5.51	50602	51703	25	5	0.392	5.93	−4.44	10.	8.76	V2292 Oph
152311.....	82621	HR 6269	0.685	3.63	50548	50714	4	3	0.135	2.80	−5.19	35.	10.02	
152792.....	82636	SAO 46350	0.631	3.43	50547	52806	17	16	0.144	1.08	−5.11	25.	9.91	
152555.....	82688		0.591	4.43	52516	52834	7	3	0.334	4.30	−4.41	4.	8.65	
152781.....	82861	HR 6284	0.952	3.30	52101	52101	1	1	0.134 ^a	... ^a	... ^a	
153525.....	83006	GJ 649.1C	1.004	6.71	50547	52835	13	7	0.613	6.27	... ^a	... ^a	... ^a	
153557.....	83020	GJ 649.1A	0.980	6.49	50547	52834	16	7	0.636	5.72	... ^a	... ^a	... ^a	
	83043	GJ 649	1.522	9.58	51410	52833	25	13	1.588	5.65	... ^a	... ^a	... ^a	
153458.....	83181	SAO 141474	0.652	4.78	50863	52829	16	13	0.257	13.19	−4.62	14.	9.15	

TABLE 2—Continued

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
153627.....	83204	SAO 102522	0.594	4.23	51013	51013	1	1	0.167	...	−4.93	17.	9.63	
154345.....	83389	GJ 651	0.728	5.48	50547	52834	18	15	0.183	5.31	−4.91	31.	9.60	
154160.....	83435	HR 6339	0.770	3.72	52101	52101	1	1	0.137	...	−5.17	45.	9.99	
154088.....	83541	GJ 652	0.814	5.30	50548	52834	26	19	0.166	4.01	−5.02	42.	9.77	
154363.....	83591	GJ 653	1.100	7.54	50276	52835	18	14	0.518	10.85	... ^a	... ^a	... ^a	
154417.....	83601	HR 6349	0.578	4.45	51013	51013	1	1	0.248	...	−4.59	8.	9.12	V2213 Oph
154653.....	83689	GJ 654.2	0.960	0.30	50276	50608	2	2	0.098	0.57	... ^a	... ^a	... ^a	
155060.....	83827	SAO 65809	0.562	4.42	51013	52449	2	2	0.157	...	−4.97	14.	9.70	
154962.....	83906	HR 6372	0.697	3.61	51013	51013	1	1	0.133	...	−5.21	37.	10.04	
155456.....	84028		0.869	5.44	52833	52833	1	1	0.168	...	−5.03	46.	9.80	
155423.....	84082	SAO 122157	0.569	3.56	51013	51013	1	1	0.144	...	−5.07	17.	9.85	
	84099	GJ 3992	...	11.14	52390	52833	5	4	0.817	8.02	
156279.....	84171		0.801	5.33	52833	52833	1	1	0.155	...	−5.06	43.	9.84	
155712.....	84195		0.941	6.39	52833	52833	1	1	0.205 ^a	... ^a	... ^a	
156146.....	84417	SAO 102690	0.734	4.25	51371	52190	18	8	0.207	5.06	−4.82	28.	9.46	
156026.....	84478		1.144	7.45	50984	52805	18	14	0.858	2.90	... ^a	... ^a	... ^a	V2215 Oph
156668.....	84607		1.015	6.52	52833	52833	1	1	0.194 ^a	... ^a	... ^a	
156985.....	84616		1.019	6.59	52190	52806	11	6	0.305	7.72	... ^a	... ^a	... ^a	
156365.....	84636	GJ 665.1	0.647	3.22	50277	52829	46	19	0.137	1.64	−5.17	29.	10.00	
		GJ 667C	1.570	10.97	52007	52804	10	7	0.633	4.80	... ^a	... ^a	... ^a	LHS 443
	84790	GJ 671	1.560	10.91	52390	52833	3	3	0.601 ^a	... ^a	... ^a	
156826.....	84801	HR 6439	0.850	2.67	50276	52778	24	15	0.135	1.63	−5.18	52.	10.01	
157214.....	84862	HR 6458	0.619	4.59	50602	50608	9	1	0.151	0.99	−5.04	22.	9.81	
157466.....	85007	SAO 85045	0.526	4.51	51013	52449	2	2	0.165	...	−4.89	10.	9.58	
157172.....	85017	SAO 160504	0.783	5.22	52095	52829	10	6	0.177	3.97	−4.96	38.	9.68	
157347.....	85042	HR 6465	0.680	4.83	50955	52829	17	13	0.155	1.51	−5.04	30.	9.80	
157338.....	85158	SAO 208769	0.588	4.34	50548	52807	13	12	0.153	1.13	−5.01	18.	9.76	
158633.....	85235	HR 6518	0.759	5.90	51975	51975	1	1	0.181	...	−4.93	35.	9.63	
158222.....	85244	SAO 30377	0.667	4.76	50956	51369	6	4	0.177	1.44	−4.91	25.	9.61	
157881.....	85295	GJ 673	1.359	8.10	50276	52805	18	10	1.686	7.26	... ^a	... ^a	... ^a	
158332.....	85436		0.820	5.31	52833	52833	1	1	0.150	...	−5.10	46.	9.89	
159062.....	85653		0.737	5.47	52833	52833	1	1	0.159	...	−5.03	36.	9.79	
	85665	GJ 678.1A	1.461	9.33	50955	52804	16	11	1.367	12.71	... ^a	... ^a	... ^a	
159063.....	85799	SAO 102891	0.534	3.47	51013	52804	24	11	0.146	2.03	−5.03	12.	9.79	
159222.....	85810	HR 6538	0.639	4.65	50547	52835	19	17	0.177	3.62	−4.90	22.	9.58	
	86162	GJ 687	1.505	10.87	50604	52806	25	17	0.771	11.58	... ^a	... ^a	... ^a	
159909.....	86193	SAO 102962	0.693	4.46	51013	51013	1	1	0.180	...	−4.91	28.	9.60	
	86287	GJ 686	1.530	10.08	50605	52804	17	12	0.691	10.16	... ^a	... ^a	... ^a	
160346.....	86400	GJ 688	0.959	6.38	50276	50276	1	1	0.284 ^a	... ^a	... ^a	
160693.....	86431	SAO 66228	0.576	4.70	50547	52835	29	18	0.160	1.33	−4.95	16.	9.67	
161897.....	86540		0.720	5.29	52515	52835	6	3	0.199	3.14	−4.84	28.	9.49	
161198.....	86722	GJ 692.1	0.752	5.65	50276	50667	4	3	0.171	1.40	−4.97	36.	9.70	
	86776	GJ 694	1.528	10.60	50956	52805	18	13	1.028	9.79	... ^a	... ^a	... ^a	
	86961	GJ 2130	1.463	11.53	52007	52488	7	5	2.045	9.18	... ^a	... ^a	... ^a	
161797.....	86974	HR 6623	0.750	3.80	50602	51411	48	4	0.145	1.35	−5.11	40.	9.92	86 Her
161555.....	86985	SAO 141859	0.671	3.63	51013	51013	1	1	0.144	...	−5.11	31.	9.91	
	87062		0.605	5.67	51410	52804	11	9	0.165	6.86	−4.94	19.	9.66	BD −08 4501

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD -2400000)	End (JD -2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
161848.....	87089	GJ 9605	0.822	6.00	50276	52829	18	16	0.183	2.17	-4.95	40.	9.67	
162826.....	87382	HR 6669	0.541	3.93	51013	52452	3	3	0.144	...	-5.05	13.	9.83	
	87579	GJ 697	0.940	6.52	50276	50276	1	1	0.602	
163102.....	87678	SAO 141954	0.516	3.76	51013	51013	1	1	0.146	...	-5.02	10.	9.77	
163153.....	87710	SAO 141956	0.759	3.66	51013	52834	40	21	0.133	1.69	-5.20	45.	10.03	
163489.....		GJ 4035	1.110	...	50276	52833	51	20	0.103	1.92	
163675.....	87834	SAO 103234	1.002	3.46	52101	52101	1	1	0.135	
163840.....	87895	HR 6697	0.642	4.04	50276	50276	1	1	0.161	...	-4.99	24.	9.73	
	87937	GJ 699	1.570	13.21	50602	52835	67	18	0.761	17.62	
164595.....	88194	SAO 85632	0.635	4.76	50956	52835	15	13	0.159	1.22	-5.00	23.	9.74	
164507.....	88217	HR 6722	0.747	3.01	50548	52835	53	18	0.143	1.80	-5.13	41.	9.94	
164651.....	88324		0.746	5.09	52833	52833	1	1	0.155	...	-5.05	38.	9.82	
164922.....	88348	GJ 700.2	0.799	5.31	50276	52834	47	24	0.158	1.45	-5.05	43.	9.82	
165045.....	88481		0.796	5.42	52833	52833	1	1	0.344	...	-4.55	18.	9.05	
165173.....	88511		0.762	5.39	52833	52833	1	1	0.162	...	-5.02	38.	9.78	
165567.....	88533	HR 6764	0.509	3.02	51013	52452	2	2	0.186	...	-4.77	7.	9.39	
165222.....	88574	GJ 701	1.508	9.91	50602	52835	28	18	0.990	4.69	
165341.....	88601		1.150	7.49	50284	50284	1	1	0.945	70 Oph B
165360.....	88656	SAO 142081	0.526	3.38	51013	51013	1	1	0.229	...	-4.62	6.	9.15	
165438.....	88684	HR 6756	0.968	3.02	52101	52101	1	1	0.124	
165908.....	88745	GJ 704A	0.500	3.95	50547	50547	1	1	0.145	...	-5.02	9.	9.77	
165908.....	88745	GJ 704B	1.100	7.31	50603	50603	1	1	0.146	
	88778	SAO 103400	1.276	1.22	50548	50666	4	3	0.113	1.16	
166435.....	88945	SAO 85784	0.633	4.83	51013	51014	2	1	0.450	0.39	-4.27	3.	7.27	
166620.....	88972	HR 6806	0.876	6.15	50602	52452	12	4	0.186	1.89	-4.97	44.	9.70	
	89215		0.755	6.52	51410	52804	12	9	0.175	1.52	-4.96	35.	9.68	BD +05 3640
167215.....	89270	SAO 85832	0.578	3.60	52063	52538	11	6	0.139	1.50	-5.12	19.	9.93	
167216.....	89275	SAO 85834	0.529	3.70	52063	52835	14	8	0.143	1.05	-5.05	12.	9.83	
167389.....	89282	SAO 47313	0.649	4.76	50956	52835	16	12	0.214	3.73	-4.75	18.	9.35	
168009.....	89474	HR 6847	0.641	4.52	51013	52450	3	3	0.147	...	-5.08	26.	9.87	
167665.....	89620	HR 6836	0.536	4.00	50284	52575	20	16	0.152	2.92	-4.99	12.	9.73	
168603.....	89771		0.771	5.45	52833	52833	1	1	0.373	...	-4.49	13.	8.90	
168443.....	89844	GJ 4052	0.724	4.03	50277	52835	102	31	0.143	1.67	-5.12	38.	9.93	
168723.....	89962	HR 6869	0.941	1.84	52148	52148	2	1	0.118	
168746.....	90004	SAO 161386	0.713	4.78	51756	52804	13	9	0.155	1.16	-5.05	34.	9.82	
169822.....	90355	SAO 123474	0.699	5.67	51373	52489	26	13	0.169	1.23	-4.96	30.	9.69	
169889.....	90365		0.764	5.42	52833	52833	1	1	0.172	...	-4.97	37.	9.70	
169830.....	90485	HR 6907	0.517	3.10	51756	52835	11	9	0.140	1.87	-5.07	11.	9.86	
170174.....		SAO 123515	1.080	...	50548	52835	33	17	0.096	2.25	
170778.....	90586	SAO 47529	0.619	4.66	51013	51013	1	1	0.308	...	-4.48	7.	8.89	
170469.....	90593	SAO 103765	0.677	4.15	51706	52804	22	10	0.147	1.93	-5.09	31.	9.88	
170493.....	90656	GJ 715	1.074	6.67	50276	52829	23	18	0.431	13.61	
170829.....	90729	HR 6950	0.795	3.72	52101	52101	1	1	0.136	...	-5.17	47.	10.00	
170657.....	90790	GJ 716	0.861	6.21	50277	52804	28	17	0.320	5.30	-4.65	26.	9.21	
171067.....	90864	GJ 1229	0.692	5.20	50276	52836	25	19	0.176	4.66	-4.92	28.	9.62	
171314.....	90959		1.181	7.05	50276	50276	1	1	0.425	V774 Her
		GJ 4063	1.420	...	51441	52833	4	4	1.281	8.44	LP 229-17

TABLE 2—Continued

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
171665.....	91287	GJ 9630	0.687	5.03	50284	52836	19	18	0.178	6.45	−4.91	27.	9.61	
171918.....	91332	SAO 142443	0.679	4.18	51706	52804	9	8	0.151	1.53	−5.06	30.	9.84	
172310.....	91381	GJ 9631	0.704	5.70	50547	52835	18	15	0.184	6.97	−4.89	28.	9.57	
172051.....	91438	HR 6998	0.673	5.28	50284	52835	33	18	0.179	2.89	−4.90	25.	9.59	
172649.....	91507	SAO 67221	0.525	4.12	51013	51013	1	1	0.287	...	−4.46	3.	8.85	
	91699	GJ 4070	1.579	11.00	52099	52833	5	5	0.523	
172513.....	91700	SAO 210475	0.748	5.34	50548	52834	40	22	0.208	3.74	−4.82	30.	9.47	
173739.....	91768	GJ 725A	1.504	11.18	50602	52776	21	14	0.440	12.40	
173740.....	91772	GJ 725B	1.561	11.97	50602	52804	20	13	0.654	19.83	
173701.....	91949	GJ 725.1	0.843	5.36	50548	52804	16	14	0.214	12.06	−4.87	38.	9.53	
173667.....	92043	HR 7061	0.483	2.79	52123	52123	1	1	0.166	...	−4.86	6.	9.53	
173818.....	92200	GJ 726	1.295	8.06	51313	52835	11	10	1.221	10.30	
174080.....	92283	GJ 727	1.070	6.78	51314	51314	1	1	0.837	
229590.....	92311	GJ 728	1.284	8.01	51442	51442	1	1	0.941	
174457.....	92418	SAO 104160	0.621	3.84	51373	52516	14	8	0.150	1.13	−5.05	23.	9.82	
174912.....	92532	SAO 67481	0.594	4.76	51013	52450	2	2	0.164	...	−4.94	17.	9.65	
175317.....	92882	HR 7126	0.445	3.03	51707	51707	1	1	0.184	...	−4.74	3.	9.33	
175541.....	92895	GJ 736	0.869	2.49	50284	52834	30	20	0.126	3.06	−5.23	56.	10.07	
175518.....	92918	SAO 142809	0.747	4.83	51013	51013	1	1	0.160	...	−5.02	37.	9.78	
175726.....	92984	SAO 124077	0.583	4.55	51013	51013	1	1	0.331	...	−4.41	4.	8.65	
176051.....	93017	GJ 738B	1.000	6.46	50548	50548	1	1	0.154	
176029.....	93101	GJ 740	1.444	8.99	51442	51442	1	1	1.747	
176377.....	93185	GJ 9639	0.606	4.95	50547	52835	87	15	0.178	1.71	−4.87	17.	9.54	
230409.....	93341	SAO 104353	0.703	6.07	51410	52804	13	9	0.175	1.59	−4.93	30.	9.64	
176733.....	93377		0.705	4.88	52833	52833	1	1	0.155	...	−5.04	33.	9.81	
177153.....	93427	SAO 48012	0.569	4.12	51013	52123	3	2	0.154	...	−4.99	15.	9.73	
176982.....	93518	GJ 740.1A	0.738	3.47	50284	52835	19	16	0.142	1.64	−5.14	40.	9.95	
177830.....	93746	GJ 743.2	1.093	3.32	50276	52833	51	22	0.125	1.60	
		GJ 745B	1.570	11.02	50983	52833	10	9	0.282	
349726 A.....		GJ 745A	1.580	11.01	50983	52833	9	7	0.310	G 185−4
178911 B.....	94075	SAO 67875	0.750	4.62	51342	52805	15	9	0.168	3.86	−4.98	36.	9.72	
179957.....	94336	HR 7293	0.650	4.77	50548	52836	17	15	0.152	1.00	−5.05	27.	9.83	
179958.....	94336	HR 7294	0.650	4.59	50548	52836	15	14	0.148	1.34	−5.08	27.	9.86	
180161.....	94346		0.804	5.53	52833	52833	1	1	0.372	...	−4.52	16.	8.98	
230999.....	94615		0.687	4.69	51793	52833	10	6	0.153	2.46	−5.05	31.	9.82	
179949.....	94645	HR 7291	0.548	4.09	51793	52835	14	9	0.188	1.66	−4.79	10.	9.41	
180684.....	94751	SAO 104678	0.588	3.18	51373	52516	15	9	0.141	1.32	−5.11	20.	9.91	
180617.....	94761		1.464	10.28	52062	52836	23	10	0.981	7.46	V1428 Aql
182189.....	94802		0.729	4.11	52833	52833	1	1	0.135	...	−5.19	41.	10.02	
181144.....	94905	SAO 104707	0.541	4.20	51013	51013	1	1	0.249	...	−4.57	6.	9.07	
180702.....	94926	SAO 211161	0.579	3.35	50366	50366	1	1	0.150	...	−5.02	17.	9.78	
181655.....	94981	HR 7345	0.676	4.28	50806	51707	35	10	0.171	1.32	−4.94	27.	9.66	
181234.....	95015	SAO 143270	0.841	5.15	51342	52835	19	11	0.153	2.15	−5.08	47.	9.87	
181720.....	95262	SAO 211218	0.599	4.10	50366	50366	1	1	0.147	...	−5.07	20.	9.85	
182488.....	95319	HR 7368	0.804	5.42	51013	51013	1	1	0.183	...	−4.94	39.	9.65	
182619.....	95428		0.718	5.19	52833	52833	1	1	0.155	...	−5.05	35.	9.82	
182572.....	95447	HR 7373	0.761	4.27	50367	52834	49	19	0.148	2.65	−5.10	41.	9.89	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
183255.....	95575	GJ 1237	0.929	6.01	50548	50716	4	4	0.213	5.30	... ^a	... ^a	... ^a	
183263.....	95740	SAO 124664	0.678	4.25	52095	52833	20	11	0.145	1.87	−5.11	32.	9.91	
183341.....	95772	SAO 104879	0.623	4.20	51013	51013	1	1	0.157	...	−5.00	22.	9.75	
183650.....	95821	GJ 761.1	0.718	4.14	50548	52833	27	18	0.145	2.00	−5.11	37.	9.92	
183756.....	95926	SAO 104914	0.915	2.95	52101	52101	1	1	0.128 ^a	... ^a	... ^a	
183658.....	95962	GJ 9659	0.640	4.60	50955	52835	17	14	0.160	1.41	−4.99	24.	9.73	
183870.....	96085	GJ 1240	0.922	6.25	50277	52835	26	15	0.546	2.44	... ^a	... ^a	... ^a	
185144.....	96100	HR 7462	0.786	5.87	50602	52834	96	10	0.206	8.61	−4.85	34.	9.51	
184385.....	96183	GJ 762.2	0.745	5.37	51013	51013	1	1	0.314	...	−4.56	16.	9.06	
184962.....	96434	SAO 105038	0.531	3.54	51013	51013	1	1	0.250	...	−4.56	5.	9.06	
185395.....	96441	HR 7469	0.395	3.14	52112	52517	92	8	0.187 ^a	... ^a	... ^a	θ Cyg
184860.....	96471	GJ 764.1A	1.011	5.97	50284	52489	22	16	0.237	8.37	... ^a	... ^a	... ^a	
185501.....	96576		0.719	4.97	52833	52833	1	1	0.229	...	−4.74	24.	9.33	
185295.....		SAO 124877	0.840	...	50366	52833	20	16	0.138	1.47	−5.16	50.	9.98	
185720.....	96813	SAO 143674	0.525	3.57	51013	51013	1	1	0.157	...	−4.94	10.	9.66	
186408.....	96895	HR 7503	0.643	4.32	50602	52489	13	3	0.145	1.78	−5.10	27.	9.90	16 Cyg
186427.....	96901	HR 7504	0.661	4.60	52489	52489	1	1	0.148	...	−5.08	29.	9.87	
186704.....	97255	SAO 125036	0.612	4.62	51342	51412	5	2	0.380	3.04	−4.35	4.	8.26	
187123.....	97336	SAO 68845	0.661	4.43	50806	52834	60	17	0.155	1.41	−5.03	27.	9.80	
187237.....	97420	SAO 87733	0.660	4.79	51013	52805	20	12	0.210	3.00	−4.77	20.	9.38	
187691.....	97675	HR 7560	0.563	3.68	52106	52160	3	2	0.146	...	−5.05	16.	9.82	o Aql
187923.....	97767	HR 7569	0.642	3.95	50277	52835	26	18	0.146	1.59	−5.09	26.	9.88	
188015.....	97769	SAO 87842	0.727	4.63	51755	52836	23	10	0.155	4.65	−5.05	36.	9.82	
187897.....	97779	SAO 125154	0.647	4.52	51013	52573	39	14	0.243	5.46	−4.65	15.	9.21	
187760.....	97805	SAO 188654	1.155	7.21	51442	51442	1	1	1.135 ^a	... ^a	... ^a	
188088.....	97944	HR 7578	1.017	5.46	50277	50277	2	1	0.643	0.45	... ^a	... ^a	... ^a	V4200 Sgr
188510.....	98020	SAO 105417	0.599	5.85	51342	52834	16	11	0.148	1.57	−5.05	20.	9.83	
188512.....	98036	HR 7602	0.855	3.03	50602	50609	9	1	0.130	1.00	−5.21	54.	10.04	
188376.....	98066	HR 7597	0.748	2.82	50284	50715	6	5	0.178	4.71	−4.94	34.	9.65	ω Sgr
189087.....	98192	GJ 773.2	0.797	5.88	50276	50367	2	2	0.302	4.97	−4.63	22.	9.16	
188807.....	98204	GJ 773	1.318	7.91	51442	51442	1	1	0.941 ^a	... ^a	... ^a	
189067.....	98206	SAO 87974	0.641	4.04	51013	51013	1	1	0.150	...	−5.06	26.	9.84	
189733.....	98505		0.932	6.25	52833	52833	1	1	0.525 ^a	... ^a	... ^a	
189561.....	98575	HR 7643	0.981	1.10	50284	50367	2	2	0.154	1.32	... ^a	... ^a	... ^a	
189625.....	98589	SAO 163190	0.654	4.67	50806	52835	20	16	0.200	3.29	−4.80	20.	9.43	
190067.....	98677	GJ 775.1	0.714	5.72	50277	52834	25	19	0.189	5.90	−4.88	29.	9.55	
190007.....	98698		1.128	6.87	50984	52835	25	15	0.662	6.36	... ^a	... ^a	... ^a	V1654 Aql
190228.....	98714	SAO 88118	0.793	3.33	52238	52835	8	5	0.135	1.46	−5.18	47.	10.01	
190360.....	98767	HR 7670	0.749	4.72	50366	52836	67	21	0.148	1.54	−5.09	40.	9.89	
190404.....	98792	GJ 778	0.815	6.32	50276	52833	25	20	0.176	1.35	−4.98	41.	9.71	
190406.....	98819	HR 7672	0.600	4.56	50602	52576	43	14	0.199	4.39	−4.77	15.	9.39	
190470.....	98828	GJ 779.1	0.924	6.15	50276	50366	2	2	0.342	11.40	... ^a	... ^a	... ^a	
190412.....	98878		0.705	5.08	52832	52832	1	1	0.170	...	−4.96	31.	9.68	
190771.....	98921	HR 7683	0.654	4.80	51013	51013	1	1	0.350	...	−4.43	7.	8.72	
191022.....	98978	SAO 49176	0.661	3.95	51013	51013	1	1	0.153	...	−5.05	28.	9.82	
191391.....	99385	GJ 782	1.296	7.93	51442	51442	1	1	1.298 ^a	... ^a	... ^a	
193202.....	99427	GJ 786	1.320	7.75	51342	52805	12	5	0.892	3.21	... ^a	... ^a	... ^a	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{diff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
191785.....	99452	GJ 783.2A	0.830	5.78	50277	52833	31	21	0.165	1.60	−5.03	44.	9.79	
192020.....		GJ 4138	0.860	...	50548	52834	21	15	0.447	5.27	−4.48	15.	8.90	
192263.....	99711	SAO 144192	0.938	6.30	50984	52834	28	16	0.488	5.53	... ^a	... ^a	... ^a	
192343.....	99727	SAO 125595	0.680	3.92	50366	52834	15	11	0.145	1.83	−5.11	32.	9.91	
192344.....	99729	SAO 125597	0.704	3.72	51755	52834	13	8	0.136	1.79	−5.18	37.	10.01	
193216.....	99965	SAO 32429	0.747	5.70	52515	52540	2	1	0.184	0.56	−4.91	33.	9.61	
193017.....	100072	SAO 144266	0.567	4.45	51013	51013	1	1	0.209	...	−4.71	10.	9.29	
193795.....	100363	SAO 88563	0.683	4.23	51793	52836	14	10	0.150	1.14	−5.07	31.	9.85	
194035.....	100500	SAO 106041	0.726	3.71	51013	51013	1	1	0.147	...	−5.10	37.	9.90	
193901.....	100568	SAO 189226	0.554	5.45	51410	52835	15	10	0.155	1.37	−4.97	14.	9.71	
194766.....	100895	SAO 144449	0.520	4.15	51013	51013	1	1	0.159	...	−4.93	10.	9.63	
194765.....	100896	SAO 144450	0.519	3.29	51707	52159	2	2	0.157	...	−4.94	10.	9.65	
195034.....	100963		0.642	4.84	52515	52836	6	4	0.166	2.32	−4.96	23.	9.68	
195019 A.....	100970	SAO 106138	0.662	4.01	51013	52835	37	18	0.147	1.87	−5.09	29.	9.88	
195019 B.....	100970		1.150	7.34	51439	52835	13	7	0.675	12.21	... ^a	... ^a	... ^a	4"1 from component A
195104.....	101059	SAO 125857	0.512	4.25	51013	51013	1	1	0.166	...	−4.88	9.	9.56	
	101180	GJ 793	1.542	11.04	50984	52805	9	8	1.600	7.76	... ^a	... ^a	... ^a	
195564.....	101345	HR 7845	0.689	3.74	50366	52778	28	17	0.142	1.59	−5.13	34.	9.94	
195987.....	101382		0.796	5.35	52832	52832	1	1	0.175	...	−4.97	39.	9.70	
196201.....	101597	SAO 106276	0.759	5.58	51368	52540	16	7	0.181	5.33	−4.93	35.	9.63	
196850.....	101875	GJ 794.3	0.610	4.61	50548	52835	24	15	0.161	1.97	−4.97	20.	9.70	
196885.....	101966	HR 7907	0.559	3.80	51013	52450	2	2	0.151	...	−5.01	15.	9.76	
196761.....	101997	HR 7898	0.719	5.53	50277	52835	29	18	0.179	3.00	−4.92	31.	9.63	
197076.....	102040	HR 7914	0.611	4.82	50366	52836	21	16	0.170	3.25	−4.92	19.	9.62	
197214.....	102264	GJ 4157	0.671	5.20	50277	50715	5	4	0.175	1.46	−4.92	26.	9.62	
	102401	GJ 806	1.491	10.29	50602	52805	12	11	0.948	10.55	... ^a	... ^a	... ^a	
197711.....	102486	SAO 189666	0.914	5.21	50984	51051	3	2	0.187	0.94	... ^a	... ^a	... ^a	
198089.....	102610	SAO 106494	0.587	4.47	51013	51013	1	1	0.158	...	−4.98	17.	9.71	
198387.....	102642	HR 7972	0.883	3.18	52101	52101	1	1	0.134	...	−5.19	54.	10.02	
198425.....	102766		0.939	6.38	52832	52832	1	1	0.448 ^a	... ^a	... ^a	
199019.....	102791		0.767	5.51	52515	52807	6	4	0.457	1.48	−4.37	8.	8.44	
198550.....	102851	GJ 808.2	1.068	6.74	50276	50276	1	1	0.938 ^a	... ^a	... ^a	
199476.....	102970	GJ 1255D	0.685	5.45	50604	52836	19	11	0.168	1.43	−4.96	28.	9.69	
	103039		1.653	12.71	52805	52835	2	1	1.294	6.96	... ^a	... ^a	... ^a	LP 816−60
198802.....	103077	HR 7994	0.661	3.11	50366	52833	28	16	0.146	3.58	−5.10	29.	9.89	
199305.....	103096	GJ 809	1.483	9.31	50602	52835	19	14	1.553	7.27	... ^a	... ^a	... ^a	
	103256	GJ 1259	1.054	7.04	50366	50366	2	1	0.555	0.26	... ^a	... ^a	... ^a	
	103269		0.590	6.05	51343	52833	20	10	0.156	1.84	−4.99	18.	9.74	BD +41 3931
199598.....	103455	SAO 89320	0.584	4.32	51013	52160	4	3	0.207	...	−4.73	12.	9.32	
199960.....	103682	HR 8041	0.635	4.10	50366	52836	21	16	0.147	2.13	−5.08	25.	9.87	11 Aqr
199918.....	103735	SAO 212635	0.620	4.41	50366	50957	6	5	0.172	2.64	−4.91	20.	9.61	
200560.....	103859	GJ 816.1A	0.970	6.26	50548	50548	1	1	0.621 ^a	... ^a	... ^a	
200565.....	103983	SAO 126508	0.650	4.37	51341	52806	31	12	0.195	2.98	−4.82	21.	9.46	
200538.....	104071	SAO 212692	0.606	4.01	50366	52836	18	15	0.146	2.22	−5.07	21.	9.85	
200746.....	104075	SAO 126530	0.654	4.74	51373	52515	14	7	0.326	2.34	−4.47	8.	8.86	
200779.....	104092	GJ 818	1.119	7.41	50277	50366	4	2	0.750	10.45	... ^a	... ^a	... ^a	
201651.....	104225		0.766	5.61	52832	52832	1	1	0.163	...	−5.01	38.	9.77	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	σ_{diff}/S (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
200968.....	104239	GJ 819A	0.901	5.89	50984	52833	15	9	0.420	5.14	... ^a	... ^a	... ^a	
201219.....	104318		0.692	5.23	52515	52836	8	3	0.267	6.38	−4.62	16.	9.16	
201203.....	104367	SAO 190052	0.511	3.09	52095	52836	14	7	0.179	1.19	−4.81	8.	9.44	
	104432	GJ 821	1.490	10.45	51704	52829	8	6	0.362 ^a	... ^a	... ^a	
201924.....	104587		0.780	5.40	52834	52836	2	1	0.151	2.04	−5.08	42.	9.87	
201891.....	104659	SAO 106908	0.525	4.63	51013	52162	2	2	0.170	...	−4.86	10.	9.53	
202108.....	104733	SAO 71071	0.666	5.17	51013	51013	1	1	0.210	...	−4.77	20.	9.38	
201989.....	104809		0.689	5.01	52516	52833	6	3	0.313	5.43	−4.52	12.	8.98	
202573.....	105000	GJ 822.2	0.887	0.69	50276	51075	21	9	0.103	1.76	−5.39	71.	10.25	
202575.....	105038	GJ 824	1.020	6.84	50367	52835	20	13	0.746	2.94	... ^a	... ^a	... ^a	
202751.....	105152	GJ 825.3	0.990	6.73	50366	52832	20	16	0.246	8.31	... ^a	... ^a	... ^a	
203030.....	105232		0.750	5.39	52515	52830	7	3	0.450	3.10	−4.37	7.	8.39	
203040.....	105341	GJ 826.1	1.340	8.02	51442	51442	1	1	1.642 ^a	... ^a	... ^a	
202917.....	105388		0.690	5.34	52539	52833	4	2	0.537	5.64	−4.22	2.	< 7	
204277.....	105918		0.529	4.07	52515	52807	6	3	0.250	0.75	−4.56	5.	9.06	
	106106	GJ 829A	1.620	11.19	50957	50957	1	1	0.850 ^a	... ^a	... ^a	
204814.....	106122		0.759	5.56	52833	52833	1	1	0.156	...	−5.05	39.	9.82	
204587.....	106147	GJ 830	1.261	7.83	50366	52807	23	15	0.755	9.41	... ^a	... ^a	... ^a	
205905.....	106913		0.623	4.71	52516	52807	6	3	0.247	3.54	−4.63	12.	9.17	
	106924		0.551	6.27	51410	52830	16	8	0.130	3.09	−5.20	16.	10.03	BD +59 2407
206332.....	107040	SAO 89899	0.600	3.91	51013	52160	2	2	0.159	...	−4.98	19.	9.71	
207897.....	107038		0.868	6.07	52832	52832	1	1	0.251	...	−4.79	35.	9.42	
206374.....	107070		0.686	5.30	52515	52834	7	4	0.256	2.22	−4.64	17.	9.19	
206387.....	107107	SAO 126991	0.720	4.57	51341	52829	18	10	0.429	2.95	−4.36	7.	8.38	
207485.....	107457		0.727	5.10	52832	52832	1	1	0.387	...	−4.43	9.	8.72	
207491.....		GJ 838.1A	1.040	6.71	50366	50366	2	1	0.568	0.29	... ^a	... ^a	... ^a	
207740.....	107821	SAO 90050	0.737	4.50	51342	51793	13	4	0.217	1.56	−4.79	27.	9.41	
207839.....	107822		0.777	5.15	52833	52833	1	1	0.196	...	−4.88	34.	9.56	
207804.....	107840	SAO 71762	1.062	0.45	50604	50689	3	2	0.197	5.00	... ^a	... ^a	... ^a	
207966.....	107920		0.798	5.50	52833	52833	1	1	0.194	...	−4.90	37.	9.59	
207874.....	107941	SAO 127122	0.880	5.71	50984	52807	15	10	0.175	4.98	−5.01	46.	9.77	
207992.....	107958		0.720	5.23	52832	52832	1	1	0.157	...	−5.03	35.	9.80	
208038.....	108028		0.937	6.28	52832	52832	1	1	0.533 ^a	... ^a	... ^a	
208313.....	108156	GJ 840	0.911	6.19	51374	51374	1	1	0.279 ^a	... ^a	... ^a	
208527.....	108296	HR 8372	1.698	−1.34	50276	50419	7	3	0.288	2.92	... ^a	... ^a	... ^a	
208776.....	108473	SAO 127201	0.577	3.82	51013	51013	1	1	0.147	...	−5.05	17.	9.83	
208801.....	108506	HR 8382	0.971	3.46	50366	52836	26	16	0.125	1.33	... ^a	... ^a	... ^a	
208880.....	108525		0.755	5.70	52832	52832	1	1	0.206	...	−4.83	31.	9.48	
209393.....	108774		0.693	5.33	52515	52834	7	3	0.346	2.04	−4.46	10.	8.84	
209290.....	108782	GJ 846	1.453	9.10	51410	52806	16	7	1.717	7.81	... ^a	... ^a	... ^a	
209253.....	108809		0.504	4.24	52515	52807	7	3	0.278	0.79	−4.47	3.	8.87	
209458.....	108859		0.594	4.29	51341	52836	56	14	0.154	1.40	−5.00	19.	9.75	V376 Peg
209599.....	108947		0.816	5.85	52832	52832	1	1	0.197	...	−4.90	38.	9.59	
209779.....	109110	SAO 145866	0.674	4.82	51374	51374	1	1	0.349	...	−4.44	8.	8.78	
209875.....	109144	SAO 127300	0.537	3.72	51013	52829	23	12	0.147	2.23	−5.02	12.	9.78	
210144.....	109162		0.788	5.33	52832	52832	1	1	0.174	...	−4.97	39.	9.70	
211681.....	109169	SAO 3704	0.735	3.85	52096	52602	10	6	0.147	2.03	−5.10	38.	9.89	

TABLE 2—Continued

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
210312.....	109355	SAO 107694	0.670	4.89	52095	52806	22	8	0.156	2.09	−5.03	28.	9.79	
210277.....	109378	GJ 848.4	0.773	4.90	50277	52829	66	26	0.155	1.98	−5.06	41.	9.83	
	109388	GJ 849	1.531	10.65	51410	52834	15	7	1.042	4.08	... ^a	... ^a	... ^a	
210302.....	109422	HR 8447	0.489	3.58	51342	52807	12	9	0.147	1.33	−4.99	8.	9.73	τ PsA
210392.....	109428	SAO 145912	0.702	3.95	51793	52829	16	9	0.154	2.25	−5.05	33.	9.83	
210460.....	109439	HR 8455	0.688	2.46	50276	52834	51	21	0.215	12.69	−4.77	22.	9.38	
210667.....	109527	GJ 850	0.812	5.48	51013	51440	17	5	0.389	3.47	−4.50	15.	8.95	
	109555	GJ 851	1.465	9.96	52095	52834	7	5	1.838	10.47	... ^a	... ^a	... ^a	
210752.....	109646	SAO 145939	0.535	4.56	51013	51013	1	1	0.167	...	−4.89	11.	9.57	
211038.....	109822	GJ 851.3	0.890	3.64	50277	52832	62	22	0.135	1.58	−5.18	54.	10.01	
211080.....	109836	SAO 145965	0.756	3.27	52095	52807	10	6	0.136	1.14	−5.18	44.	10.01	
		SAO 51891	0.960	...	52515	52829	7	3	1.011	1.64	... ^a	... ^a	... ^a	V383 Lac
212291.....	110508		0.682	5.36	52515	52836	14	3	0.198	2.78	−4.82	24.	9.46	
212733.....	110716		0.902	6.00	52832	52832	1	1	0.165 ^a	... ^a	... ^a	
212801.....	110853	SAO 213830	0.810	3.19	50367	52832	33	18	0.189	7.95	−4.92	38.	9.62	
239960 A	110893	GJ 860A	1.613	11.58	50606	52601	19	13	0.707	10.39	... ^a	... ^a	... ^a	
239960 B	110893		1.800	13.31	51439	51439	1	1	0.659 ^a	... ^a	... ^a	DO Cep
213042.....	110996	GJ 862	1.080	6.71	52189	52807	8	6	0.374	9.14	... ^a	... ^a	... ^a	
213472.....	111136		0.700	4.13	52095	52807	12	7	0.145	1.63	−5.11	35.	9.91	ZI 2096
213519.....	111148	SAO 52097	0.649	4.50	50806	52833	19	12	0.156	1.18	−5.02	26.	9.78	
213575.....	111274	SAO 146142	0.668	4.17	51013	51013	1	1	0.147	...	−5.09	30.	9.88	
213628.....	111349	GJ 9787	0.721	5.27	50367	52807	19	14	0.175	3.39	−4.94	32.	9.65	
214557.....	111748	SAO 52209	0.582	3.54	51013	52162	6	3	0.143	0.94	−5.09	18.	9.88	
214683.....	111888		0.938	6.70	52832	52832	1	1	0.590 ^a	... ^a	... ^a	
214749.....	111960	GJ 868.0	1.143	7.17	50277	52832	23	14	1.156	4.80	... ^a	... ^a	... ^a	
215152.....	112190	GJ 4291	0.966	6.45	51010	52807	16	11	0.293	11.11	... ^a	... ^a	... ^a	
215500.....	112245		0.719	5.50	52832	52832	1	1	0.165	...	−4.99	33.	9.73	
215578.....		SAO 108160	0.970	...	50367	52536	24	16	0.112	1.76	... ^a	... ^a	... ^a	
215704.....	112426		0.803	5.50	52833	52833	1	1	0.196	...	−4.90	37.	9.58	
215648.....	112447	HR 8665	0.502	3.15	52121	52229	16	3	0.139	...	−5.07	10.	9.86	ξ Peg
216520.....	112527	SAO 3796	0.867	6.03	52189	52833	11	7	0.181	2.49	−4.98	44.	9.72	
216191.....	112768		0.859	5.75	52832	52832	1	1	0.205	...	−4.90	40.	9.59	
216182.....	112813		1.192	0.18	50366	50786	5	4	0.112	1.96	... ^a	... ^a	... ^a	72 Aqr; see § 5.4
216275.....	112812		0.590	4.74	52515	52807	7	3	0.165	1.87	−4.93	17.	9.64	
216320.....	112829		0.817	5.45	52833	52833	1	1	0.259	...	−4.73	29.	9.32	
216259.....	112870	GJ 9798	0.849	6.68	50276	52807	30	17	0.188	1.65	−4.95	42.	9.67	
	113020	GJ 876	1.597	11.79	50602	52835	115	24	1.020	10.44	... ^a	... ^a	... ^a	
216625.....	113086	SAO 108240	0.530	3.82	51013	52162	6	4	0.160	0.72	−4.93	11.	9.63	
216770.....	113238		0.821	5.22	52832	52832	1	1	0.193	...	−4.92	39.	9.62	
216803.....	113283		1.094	7.07	52515	52807	4	3	0.913	4.21	... ^a	... ^a	... ^a	
216899.....	113296	GJ 880	1.507	9.49	50666	52807	20	14	1.882	6.74	... ^a	... ^a	... ^a	
217014.....	113357	HR 8729	0.666	4.52	52121	52190	17	3	0.148	...	−5.08	29.	9.87	
217004.....	113386	SAO 191529	0.679	3.41	50367	52807	23	15	0.140	1.57	−5.15	33.	9.96	
217107.....	113421	HR 8734	0.744	4.70	51014	52832	53	22	0.150	1.36	−5.08	39.	9.87	
217165.....	113438	SAO 127869	0.617	4.47	51342	52829	33	16	0.165	1.28	−4.95	20.	9.67	
217357.....	113576	GJ 884	1.379	8.33	50366	52830	37	15	1.569	6.14	... ^a	... ^a	... ^a	
217618.....	113695	SAO 72937	0.724	4.22	51374	51374	1	1	0.148	...	−5.09	37.	9.88	

TABLE 2—*Continued*

HD	HIP	Other	$B-V$	M_V	Begin (JD −2400000)	End (JD −2400000)	Number of Obs.	Number of Month Bins	Grand S	$\sigma_{S_{\text{diff}}}/S$ (%)	$\log R'_{\text{HK}}$	P_{rot} (days)	$\log (\text{Age/yr})$	Note
217813.....	113829		0.620	4.72	51374	52159	2	2	0.316	...	−4.47	7.	8.85	MT Peg
217877.....	113896	HR 8772	0.581	4.24	51014	52829	14	7	0.152	2.00	−5.01	17.	9.76	
218168.....	113905	SAO 10607	0.632	4.68	52096	52834	14	8	0.302	1.59	−4.50	8.	8.94	
218209.....	113989	GJ 9808	0.646	5.12	50667	52806	25	14	0.180	2.26	−4.89	22.	9.57	
218101.....	113994	HR 8784	0.886	3.40	52101	52101	1	1	0.164	...	−5.06	48.	9.83	
218133.....	114028	SAO 108390	0.597	4.21	51014	51014	1	1	0.153	...	−5.02	19.	9.78	
217987.....	114046	GJ 887	1.483	9.76	50984	52830	29	12	1.070	11.30	... ^a	... ^a	... ^a	
218235.....	114081	HR 8788	0.482	2.98	52158	52160	3	1	0.206	...	−4.67	5.	9.23	
218261.....	114096	HR 8792	0.544	4.18	52123	52159	2	2	0.213	...	−4.68	8.	9.25	
218566.....	114322	GJ 4313	1.012	6.21	50367	52829	23	15	0.245	17.18	... ^a	... ^a	... ^a	
218687.....	114378	SAO 108437	0.607	4.48	51014	51014	1	1	0.343	...	−4.41	5.	8.62	
218739.....	114385	SAO 52754	0.658	4.77	51014	51014	1	1	0.301	...	−4.52	10.	8.99	
	114411	GJ 891	1.566	10.29	52829	52829	1	1	1.168	13.73	... ^a	... ^a	... ^a	
218730.....	114424	SAO 146541	0.604	4.57	51014	51014	1	1	0.177	...	−4.88	17.	9.55	
218868.....	114456	SAO 52768	0.750	5.13	51014	51014	1	1	0.231	...	−4.75	27.	9.35	
219172.....	114670	SAO 108468	0.561	4.05	51014	52162	2	2	0.178	...	−4.85	12.	9.50	
219175 B	114703	SAO 146578	0.651	5.96	50367	50367	1	1	0.179	...	−4.89	23.	9.57	
219420.....	114834	SAO 128069	0.537	3.52	51014	51014	1	1	0.156	...	−4.96	12.	9.68	
219538.....	114886	GJ 4320	0.871	6.15	50462	52829	20	16	0.233	9.67	−4.84	38.	9.49	
219542 B	114914		0.654	3.91	52446	52830	15	6	0.204	4.76	−4.78	20.	9.40	
219542 A	114914	SAO 146605	0.640	4.50	52446	52805	11	5	0.158	1.92	−5.00	24.	9.75	
219834 A	115126	HR 8866	0.787	3.62	50366	50366	1	1	0.159	...	−5.04	41.	9.81	94 Aqr A
219834 B	115125	GJ 894.2B	0.880	6.06	50366	52829	34	19	0.230	12.35	−4.85	39.	9.51	
219953.....	115194	GJ 9822	0.811	6.36	50276	52829	27	19	0.175	1.23	−4.98	41.	9.71	
220077.....	115279		0.561	4.08	51369	52488	24	9	0.171	1.34	−4.88	13.	9.55	HEI 88
220096.....	115312	HR 8883	0.817	0.63	50366	50366	1	1	0.492	...	−4.39	9.	8.52	
220182.....	115331	GJ 894.4	0.801	5.66	50667	50667	1	1	0.494	...	−4.37	8.	8.41	
	115332	GJ 4333	1.524	11.55	52488	52829	3	3	0.945	9.28	... ^a	... ^a	... ^a	
220221.....	115341		1.056	6.51	52834	52834	1	1	0.627 ^a	... ^a	... ^a	
220339.....	115445	GJ 894.5	0.881	6.35	50367	52830	20	15	0.					

NOTES.—Values of $B-V$ and M_V are from *Hipparcos* (Perryman et al. 1997), where available. “Begin” and “End” are the dates of the first and last observations used to compute the grand S . “Number of Month Bins” denotes total number of time bins used to determine the grand S as defined in § 5.1.1. “Number of Obs.” denotes the total number of observations of the target included in grand S . “ $\sigma_{S_{\text{diff}}}/S$ ” is the fractional standard deviation of the differential S -values for a star (for those stars with Keck data) and measures the star’s intrinsic variability. The overall uncertainty in the grand S varies from star to star, but we conservatively estimate it to be around 13%, in general (see § 5.1). “log (Age/yr)” values less than 7 are entered as “< 7.” Coordinates in the “Note” column are J2000 *Hipparcos* coordinates. Other entries in the “Note” column are explained in §5.3. Table 2 is also available in machine-readable form in the electronic edition of the *Astrophysical Journal Supplement*.

^a The chromospheric activity relation used to derive this quantity is only calibrated for stars with $0.44 < B-V < 0.9$.

5.4.3. *HR 3309*

HR 3309 (=HD 71148), which appears as a boxed cross in Figure 2, has a grand S -value of 0.17, which is highly discrepant from the mean Duncan et al. value of 1.57. The higher value is inconsistent with the *Hipparcos* $B-V$ value of 0.67 for this star. Soderblom (1985), in his analysis of the Mount Wilson S -values, quotes a mean S -value of 0.169, consistent with our value, and the Mount Wilson project archives confirm that the Duncan et al. value is a transcription error (S. Baliunas 2003, private communication).

5.4.4. *HD 137778*

HD 137778 (=GJ 586B), which appears as a boxed asterisk in Figure 2, has a grand S -value of 0.57, significantly higher than the mean Duncan et al. value of 0.16. Both values are plausible with its *Hipparcos* $B-V$ color of 0.87, but the higher value would imply a very young age. Strassmeier et al. (2000) report an “ R_{HK} ” of 7×10^{-5} , which apparently corresponds to the R'_{HK} -values calculated here. Strassmeier’s R_{HK} would imply $S = 0.875$ for this star, which is extremely high for a star of this color, and would imply extraordinary activity.

This star is in the *ROSAT* All-Sky Survey bright source catalog with a flux of $\sim 5 \times 10^{13}$ ergs cm $^{-2}$ s $^{-1}$. With a parallax of 48 mas, this corresponds to $L_X/L_{\text{Bol}} \sim 10^{-5}$, consistent with a very active, young star.

These discrepancies are puzzling, although it appears that most measurements imply significant activity. Perhaps this star has simply become significantly more active since the Mount Wilson data were taken. Observations by the Mount Wilson project between 1995 May and 2001 March give a mean S -value of 0.64, much closer to the value reported here (S. Baliunas 2003, private communication). The Strassmeier value may represent an extraordinary event or may be calibrated differently than values derived from S -values.

5.4.5. *HD 58830*

Based upon its parallax, HD 58830 (=GJ 9233), which appears as a boxed triangle in Figure 2, appears not to be a main-sequence star at all, but a giant. Our S -value of 0.20 is significantly different from the Duncan et al. value of 0.54. Observations by the Mount Wilson project between 1997 November and 1998 November give a mean S -value of 0.21, consistent with our observations (S. Baliunas 2003, private communication).

6. CONCLUSIONS

We have measured chromospheric activity as S -values from over 15,000 archival spectra taken over the course of the California and Carnegie Planet Search Program. These spectra were taken with the HIRES spectrograph at Keck Observatory and the two detectors at the Hamilton spectrograph at Lick observatory and contain both precision velocity information and the Ca II H and K lines from which S -values were derived.

Extraction of activity measurements from the Keck spectra was successful, with over 95% of all Keck spectra used

yielding useful S -values. The Lick spectra were more problematic: only $\sim 50\%$ of the spectra proved useful owing to low signal-to-noise ratios and poor extraction in this very blue region by the automated extraction pipeline. Nonetheless, the 1400 good Lick measurements and the 14514 good Keck measurements combine to give a record of the chromospheric activity for over 1000 late-type main-sequence stars. Analysis of the measured activity level of τ Ceti demonstrates a typical per observation random error of 6% at Keck and 6.5% and 10.5% for the two detectors at Lick.

We combined the Keck data with the good Lick data to create median S -values in 30 day bins. We have generated “grand S ” by taking the median of these monthly values. These grand S -values represent median activity levels for our program stars for the periods we observed them. For stars with $0.44 < B-V < 0.9$, there are well-calibrated relationships between mean activity level, age, and rotation period, allowing the determination of those quantities for our stellar sample. We present our grand S -values and derived ages and rotation periods in Table 2. We also have measured differential S -values for each star that are more precise. We present these data electronically as Table 1.

For each star, these measurements of activity, S and R'_{HK} represent median activity levels over the duration of the observations, which may be significantly shorter than a stellar activity cycle. This represents a source of error when deriving stellar properties such as age from these measurements.

The differential S -values, S_{diff} , measured here are much more precise than S but no more accurate: we scaled them to the median of the S -values, so while they are excellent measurements of temporal variations in activity for a given star, they contain no additional information regarding differences in the overall activity among stars.

The authors are indebted to Debra Fischer for her work with the Lick Observatory Planet Search, particularly in regards to gathering Ca II H and K data, and for donating her time and expertise regarding the Hamilton spectrograph and its data products.

The authors wish to recognize and acknowledge the very significant cultural role and reverence that the summit of Mauna Kea has always had within the indigenous Hawaiian community. We are most fortunate to have the opportunity to conduct observations from this mountain.

The authors also thank the many observers who helped gather the data herein, and the referee, Sallie Baliunas, for her insightful and constructive report.

This research has made use of the SIMBAD database, operated at CDS, Strasbourg, France.

This research has made use of NASA’s Astrophysics Data System Bibliographic Services.

This research was made possible by support from NASA, the NSF, and Sun Microsystems.

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