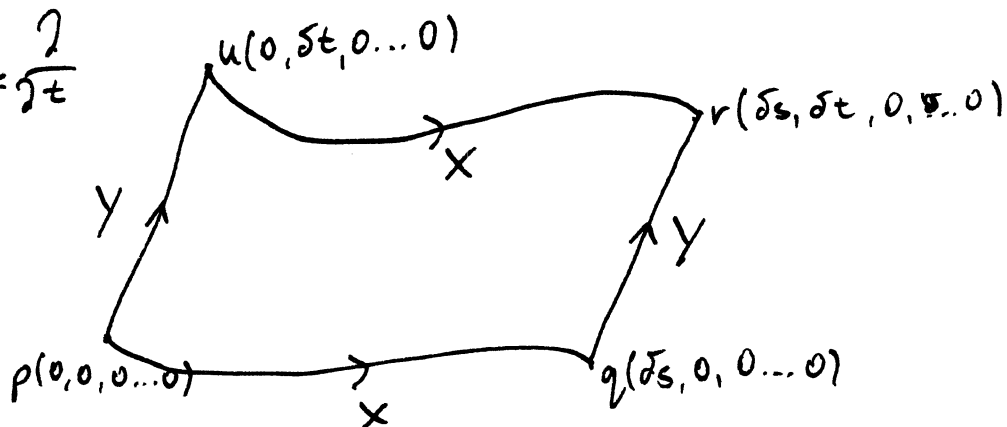


Relation between Riemann tensor and parallel transport

Let X, Y be the commuting vector fields; $[X, Y] = 0$ on M

Frobenius then $\Rightarrow \exists$ coordinate system (s, t, \dots) on M such that

$$X = \frac{\partial}{\partial s}, \quad Y = \frac{\partial}{\partial t}$$



Let $Z_p \in T_p M$, parallel transport it twice

- along pqr to $Z_r \in T_r M$

- along pqr to $Z'_r \in T_r M$

Compute the difference $\Delta Z_r = Z'_r - Z_r$ for a torsion-free connection.

[Taylor series - see notes]

$$(\Delta Z_r)^a = (Z'_r)^a - (Z_r)^a = (R^a_{\nu\rho\sigma} Z^\nu X^\rho Y^\sigma) \big|_r \delta s \delta t + O(\delta s^2, \delta t^2)$$

Rewrite as

$$R^a_{bcd} Z^b X^c Y^d \big|_r = \lim_{\delta s, \delta t \rightarrow 0} \frac{\Delta Z_r^a}{\delta s \delta t} \quad \left. \vphantom{\lim} \right\} \begin{array}{l} \text{allowed as LHS \& RHS} \\ \text{are tensors} \end{array}$$

Symmetries of the Riemann tensor

* From definition $R^a_{b(cd)} = 0$

* For torsion free ∇ , $R^a_{[bcd]} = 0$

Proof Use normal coordinates at p [enough, as this is an algebraic property of a tensor]

$$\Gamma^M_{\nu\rho}(p) = 0 \rightarrow R^M_{\nu\rho\sigma} \big|_p = (\partial_\rho \Gamma^M_{\nu\sigma} - \partial_\sigma \Gamma^M_{\nu\rho}) \big|_p$$

$$[\rho\nu\sigma] \rightarrow 0 \quad (\text{symmetries } \Gamma^M_{\nu\rho}) \quad \square$$

* Bianchi identity: If ∇ is torsion free, then $R^a{}_{b[cd;e]} = 0$.

Proof (again) normal coordinates

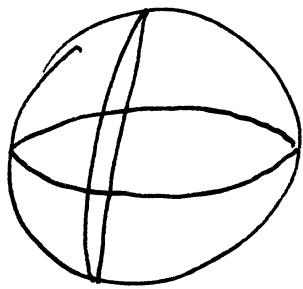
$$\left. \begin{aligned} \nabla_\tau R^\tau_{\nu\rho\sigma}|_p &= 2\tau R^\tau_{\nu\rho\sigma}|_p \\ \nabla_\tau R^\tau_{\nu\rho\sigma}|_p &= 2\tau R^\tau_{\nu\rho\sigma}|_p \end{aligned} \right\} \begin{aligned} \text{Formally} \\ R &= \partial\Gamma - \Gamma\Gamma \xrightarrow{\text{normal}} \partial\Gamma \\ \partial R &= \partial\partial\Gamma + 2\Gamma\partial\Gamma \xrightarrow{\text{normal}} \partial\partial\Gamma \end{aligned}$$

$$\text{so } R^{\Gamma}_{\nu\rho\sigma;\tau}|_p = (\partial_\tau \partial_\rho \Gamma^\mu_{\nu\sigma} - \partial_\tau \partial_\sigma \Gamma^\mu_{\nu\rho})|_p$$

$$[\rho, \sigma_\tau] \rightarrow 0 \quad \square$$

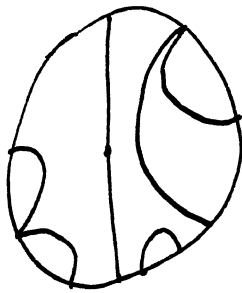
2.5 Geodesic deviation equations

In \mathbb{R}^n (flat) initially parallel lines remain parallel. For curved spaces no notion of parallel, so instead ask: given two geodesics initially close, do they remain close?

 s^2

great circles intersect

Gauss curvature $K > 0$


$$K \angle \theta$$

Poincare d'ic

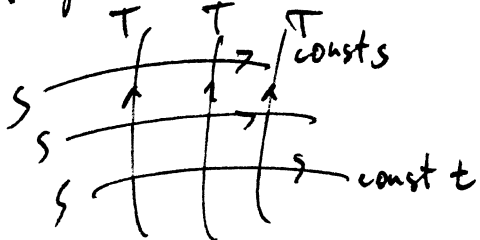
$$g = \frac{dx^2 + dy^2}{(1 - x^2 - y^2)^2}$$

$$x^2 + y^2 < 1$$

Def (M, ∇) mfd with torsion free connection. A 1-param family of geodesics in a map $\gamma: I \times I' \rightarrow M$, where

- I, I' open intervals in \mathbb{R}
- for fixed $s \in I$, $\gamma(s, t)$ is a geodesic with affine parameter t .
- $\text{Map } (s, t) \rightarrow \gamma(s, t)$ is a smooth 1:1 map, with a smooth inverse,

Family of gradients \equiv smooth surface $\Sigma \in M$



T = tangent to quadrics

So derivation vector Y^M coords on M , on Σ $X^M(s, t)$

$$\left. \begin{aligned} S^M &= \frac{\partial X^M}{\partial s} & ; & \quad S = S^M \frac{\partial}{\partial X^M} = \frac{\partial}{\partial s} \\ T^M &= \frac{\partial X^M}{\partial t} & ; & \quad T = T^M \frac{\partial}{\partial X^M} = \frac{\partial}{\partial t} \end{aligned} \right\} \begin{array}{l} \text{extend } (s, t) \text{ to a coordinate system} \\ \text{on } M, \text{ defined in a neighborhood} \\ \text{of } \Sigma \subset M. \end{array}$$

$$[T, S] = 0 \quad ; \quad \begin{aligned} \delta S \cdot \nabla_T S &= \text{relative velocity} \\ \nabla_T \nabla_T S &= - \text{acceleration} \end{aligned}$$

Prop $\nabla_T \nabla_T S = R(T, S)T$ for which $\nabla_T S - \nabla_S T = [T, S] = 0$

Proof $\nabla_T \nabla_T S = \nabla_T \nabla_S T = \nabla_S \underbrace{\nabla_T T}_0 + R(T, S)T$ } definition of R
 \uparrow Geodesic deviation eq. $\circ T$ tangent to geodesics. \square

In abstract indices

$$T^c \nabla_c (T^b \nabla_b S^a) = R^a{}_{bcd} T^b T^c S^d$$

Try to "measure" LHS $\rightarrow R^a{}_{(bc)d}$.

Claim: this determines the full curvature by [exercise]

$$R^a{}_{bcd} = \frac{2}{3} (R^a{}_{(bc)d} - R^a{}_{(bd)c})$$

Riemann = tidal forces