

If the energy is bounded below, there must be a ground state $|0\rangle$

satisfies $a|0\rangle = 0$. Excited states arise from repeated application of a^\dagger .

$$|n\rangle = (a^\dagger)^n |0\rangle \quad \text{with} \quad H|n\rangle = (n + \frac{1}{2})\omega |n\rangle \quad (\text{ignored normalisation } \langle n|n\rangle \neq 1).$$

this algebraic approach tells us the spectrum but not the explicit form of the wave function. In the Schrödinger rep, $\hat{p} = -i\frac{\partial}{\partial q}$.

$$a|0\rangle = 0 \Rightarrow \left(\frac{1}{\sqrt{2m}} \frac{\partial}{\partial q} + \sqrt{\frac{m\omega}{2}} q \right) \Psi_0(q) = 0 \Rightarrow \left(\frac{\partial}{\partial q} + \omega q \right) \Psi_0(q) = 0$$

$$\Psi_0 \propto e^{-\omega q^2/2}$$

Free Field theory

let's apply SHO to free fields. Write $\phi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_p}} \left(a_p e^{ip \cdot x} + a_p^\dagger e^{-ip \cdot x} \right)$

$$\pi(x) = \int \frac{d^3p}{(2\pi)^3} (-i) \frac{\omega_p}{2} \left(a_p e^{ip \cdot x} - a_p^\dagger e^{-ip \cdot x} \right)$$

$$\left[\begin{array}{l} \text{Can rewrite } \phi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{e^{ip \cdot x}}{\sqrt{2\omega_p}} (a_p + a_{-p}^\dagger) \\ \text{now looks like } q = \frac{1}{\sqrt{2m}} (a + a^\dagger) \end{array} \right] \quad \leftarrow \text{relabelled } p \rightarrow -p$$

Claim We can use comm relations to show

$$[\phi(x), \phi(y)] = 0$$

$$[\pi(x), \pi(y)] = 0$$

$$[\phi(x), \pi(y)] = i\delta^3(x-y)$$

$$\left. \begin{array}{l} [\phi(x), \phi(y)] = 0 \\ [\pi(x), \pi(y)] = 0 \\ [\phi(x), \pi(y)] = i\delta^3(x-y) \end{array} \right\} \Leftrightarrow \left\{ \begin{array}{l} [a_p, a_q] = [a_p^\dagger, a_q^\dagger] = 0 \\ [a_p, a_q^\dagger] = (2\pi)^3 \delta^3(p-q) \end{array} \right.$$

check one way: assume this side, prove LHS

$$\begin{aligned} [\phi(x), \pi(y)] &= \int \frac{d^3p}{(2\pi)^3} \frac{d^3q}{(2\pi)^3} \left(\frac{-i}{2} \right) \sqrt{\frac{\omega_q}{\omega_p}} \left(-[a_p, a_q^\dagger] e^{ip \cdot x - iq \cdot y} + [a_p^\dagger, a_q] e^{-ip \cdot x + iq \cdot y} \right) \\ &= \int \frac{d^3p}{(2\pi)^3} \left(\frac{-i}{2} \right) \left[e^{-ip \cdot (x-y)} - e^{ip \cdot (x-y)} \right] = i\delta^3(x-y) \end{aligned}$$

Hamiltonian in terms of a_p and a_p^\dagger .

$$H = \frac{1}{2} \int d^3x \left(\pi^2 + (\nabla\phi)^2 + m^2 \phi^2 \right)$$


$$= \frac{1}{2} \int \frac{d^3x d^3p d^3q}{(2\pi)^6} \left[\frac{-\sqrt{\omega_p \omega_q}}{2} (a_p e^{ip \cdot x} - a_p^\dagger e^{-ip \cdot x}) (a_q e^{iq \cdot x} - a_q^\dagger e^{-iq \cdot x}) + \right.$$

$$\begin{aligned}
& + \frac{1}{2\sqrt{\omega_p \omega_q}} \left(i p a_p e^{i p \cdot x} - i q a_q^\dagger e^{-i q \cdot x} \right) \left(i q a_q e^{i q \cdot x} - i p a_p^\dagger e^{-i p \cdot x} \right) \\
& + \frac{m^2}{2\sqrt{\omega_p \omega_q}} \left(a_p e^{i p \cdot x} + a_p^\dagger e^{-i p \cdot x} \right) \left(a_q e^{i q \cdot x} + a_q^\dagger e^{-i q \cdot x} \right) \Big] \\
& = \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \left[\left(-\omega_p + \frac{p^2}{\omega_p} + \frac{m^2}{\omega_p} \right) (a_p a_{-p} + a_p^\dagger a_{-p}^\dagger) + \left(\omega_p + \frac{p^2}{\omega_p} + \frac{m^2}{\omega_p} \right) (a_p a_p^\dagger + a_p^\dagger a_p) \right] \\
& = \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\omega_p} (\omega_p^2 + p^2 + m^2) (a_p a_p^\dagger + a_p^\dagger a_p) \quad \text{as } \omega_p^2 = p^2 + m^2 \\
& \Rightarrow H = \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \omega_p \left(a_p^\dagger a_p + \frac{1}{2} [a_p, a_p^\dagger] \right) \rightarrow (2\pi)^3 \delta^3(0)
\end{aligned}$$

which is simply the Hamiltonian for an ∞ of uncoupled SHOs each with frequency $\omega_p = \sqrt{p^2 + m^2}$.
The vacuum

Following the SHO, we define the vacuum $|0\rangle$ s.t. $a_p |0\rangle = 0$. So the energy of the ground state comes from the 2nd term in H . Since $[a_p, a_q^\dagger] = (2\pi)^3 \delta^3(p - q)$.

We have $H|0\rangle = \frac{1}{2} \int \frac{d^3 p}{(2\pi)^3} \omega_p (2\pi)^3 \delta^3(0) |0\rangle = \infty |0\rangle$. QFT are full of ∞ s!

But they tell us something important. Often that we are asking a stupid question. Let's explore that ∞ . In fact $\int 2 \infty$ is there. The first is due to space being big ("infrared divergence"). Put universe in a box  and impose periodic boundary conditions on the field. $(2\pi)^3 \delta^3(0) = \lim_{L \rightarrow \infty} \int_{-L/2}^{L/2} e^{i p \cdot x} \Big|_{p=0} = V$

the volume of ∞ universe. So $\delta^3(0)$ divergence was because we are computing the total energy, not the energy density \mathcal{E}_0 .

$$\mathcal{E}_0 = \frac{E}{V} = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2} \omega_p \quad \text{which is still } \infty \quad (\omega_p^2 = p^2 + m^2).$$

It's the sum of ground state energies for each SHO but $\mathcal{E}_0 \rightarrow \infty$ from $|p| \rightarrow \infty$. This is a high frequency divergence ("ultraviolet divergence") at short distance.

this arose because we assumed that our theory is valid to arbitrarily short distances / high momenta. So the \int could be cut off at high momentum in some way (e.g. lattice).

In non-gravitational physics, we only care about energy differences. So we can simply redefine H by subtracting the $\infty \Rightarrow H = \int \frac{d^3 p}{(2\pi)^3} \omega_p a_p^\dagger a_p$ s.t. $H|0\rangle = 0$.

The difference between this H and the previous one is an ordinary ambiguity in going from the classical to the quantum theory.

e.g. if we defined $H = \frac{1}{2} (\omega q - ip) (\omega q + ip)$ (classically $H = \frac{p^2}{2} + \frac{\omega^2 q^2}{2}$).

Classically the same as our original choice. Upon quantisation $H = \omega a^\dagger a$

We write a normal ordered string of ops

$$\phi_1(x_1) \dots \phi_n(x_n) \text{ as } :\phi_1(x_1) \dots \phi_n(x_n):$$

as the usual product but with all annihilation ops on the R.H.S., so we can then write

$$:H: = \int \frac{d^3 p}{(2\pi)^3} a_p^\dagger a_p$$