

Although we are calling these states 'particle', they aren't localised - they are momentum eigenstates.

We can create a localised state via a Fourier transform

$$|x\rangle = \int \frac{d^3p}{(2\pi)^3} e^{-ip \cdot x} |p\rangle$$

More generally we create a wavepacket and insert $\psi(p)$

$$|\psi\rangle = \int \frac{d^3p}{(2\pi)^3} e^{-ip \cdot x} \psi(p) |p\rangle \quad \text{which is partially localised in space and momentum}$$

$$\text{(e.g. } |\psi(p)\rangle \propto e^{-p^2/2m}$$

Neither $|x\rangle$ nor $|\psi\rangle$ are e'states of H like non-relativistic QM

Relativistic normalisation

We define vacuum s.t. $\langle 0|0\rangle = 1$ and 1 particle state $|p\rangle = a_p^\dagger |0\rangle$ satisfies $\langle p|q\rangle = (2\pi)^3 \delta^3(p - q)$

Is this Lorentz invariant? It's not obvious because we've only written 3-vectors. What could go wrong with Lorentz invariance?

Suppose we have a Lorentz transformation $p^\mu \rightarrow \Lambda^\mu_\nu p^\nu = p'^\mu$

We would want the 2 states to be related by a unitary transformation on $|p\rangle \rightarrow U(L) |p'\rangle = |p\rangle$ but we haven't been careful about normalisation, so there's no reason why we wouldn't get

$$|p\rangle \rightarrow f(p, p') |p'\rangle$$

some unknown f

The trick is to look at an object we know to be Lorentz invariant. e.g. the identity operator.

$$1 = \int \frac{d^3p}{(2\pi)^3} |p\rangle \langle p|$$

individually not L.I.

Claim: $\int \frac{d^3p}{2E_p}$ is L.I.

Proof: $\int \frac{d^4p}{(2\pi)^4}$ is L.I. relativistic dispersion relation $p_0^2 = E_p^2 = \overset{\text{3-momentum}}{p^2} + m^2$
is L.I.

The sol for p_0 has two branches, $\pm E_p$, but the choice of branch is L.I.
(can't change the sign by L. boosts)

So the following comb is L.I.

$$\int d^4p \delta(p_0^2 - p^2 - m^2) = \int \frac{d^3p}{2p_0} \Big|_{p_0 = E_p} \\ \text{take } p_0 > 0 \text{ branch}$$

Claim 2:

$2E_p \delta^3(p - \underline{q})$ is L.I.

proof: $\int \frac{d^3p}{E_p} 2E_p \delta^3(p - \underline{q}) = 1$
L.I.

From this, we learn that the correctly normalized states are
4-vector

$$|p\rangle = \sqrt{2E_p} |\tilde{p}\rangle = \sqrt{2E_p} a_p^\dagger |0\rangle$$

These new states satisfy $\langle p | \underline{q} \rangle = (2\pi)^3 2E_p \delta^3(p - \underline{q})$ and we can re-write the identity as $1 = \int \frac{d^3p}{(2\pi)^3 2E_p} |p\rangle \langle p|$
relativistic states

We can define relativistically normalized creation/annihilation operators

$$a^\dagger(p) = \sqrt{2E_p} \tilde{a}_p^\dagger$$

$$\phi = \int \frac{d^3p}{(2\pi)^3} \frac{1}{2E_p} a_p^\dagger e^{ip \cdot x} + \dots$$

① Scalar field

$$\mathcal{L} = \partial_\mu \psi^\dagger \partial^\mu \psi - \mu^2 \psi^\dagger \psi$$

real
↓
 $\psi^\dagger = \psi^*$

E-L $\partial_\mu \partial^\mu \psi + \mu^2 \psi = 0$

$\partial_\mu \partial^\mu \psi^* + \mu^2 \psi^* = 0$

check:

Expand $\psi = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} (b_p e^{ip \cdot x} + c_p^\dagger e^{-ip \cdot x})$

$$\psi^\dagger = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} (b_p^\dagger e^{-ip \cdot x} + c_p e^{ip \cdot x})$$

and $\pi = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} = \dot{\psi}^\dagger = \int \frac{d^3p}{(2\pi)^3} \frac{\sqrt{E_p}}{\sqrt{2E_p}} (b_p^\dagger e^{-ip \cdot x} - c_p e^{ip \cdot x})$

$$\pi^\dagger = \int \frac{d^3p}{(2\pi)^3} (-i) \frac{\sqrt{E_p}}{2} (b_p e^{ip \cdot x} - c_p^\dagger e^{-ip \cdot x})$$

Recall that the theory has a conserved charge classically. (11)

$$Q = i \int d^3x \dot{\psi}^* \psi - \psi^* \dot{\psi} = i \int d^3x (\pi \psi - \pi^* \psi^*)$$

~~after normal ordering~~

Commutator relations are $\left\{ \begin{aligned} [\psi(x), \pi(y)] &= i \delta^3(x-y) \\ [\psi(x), \pi^*(y)] &= 0 \end{aligned} \right\} \Leftrightarrow \left\{ \begin{aligned} [b_p, b_q^\dagger] &= (2\pi)^3 \delta^3(p-q) \\ [c_p, c_q^\dagger] &= (2\pi)^3 \delta^3(p-q) \end{aligned} \right.$

• Claim, after normal ordering

$$Q = \int \frac{d^3p}{(2\pi)^3} (c_p^\dagger c_p - b_p^\dagger b_p) = \underbrace{\int \frac{d^3p}{(2\pi)^3} N_c}_{\# \text{ c particles}} - \underbrace{\int \frac{d^3p}{(2\pi)^3} N_b}_{\# \text{ b particles}}$$

They are interpreted as part^c & anti-part^c (both spin 0, mass μ)

For a real scalar field ϕ , part^c = anti-part^c

$[Q, H] = 0 \Rightarrow Q$ is conserved (we are in a free theory, no big deal as N_c and N_b are separately conserved. But in an interacting theory, N_c and N_b are not conserved but Q is)

The Heisenberg Picture

Although our theory is L.I., it's not so obvious.

$\phi(x)$ depends on x not t , and the states evolve in t by the S. equation.

● $i \frac{d|p\rangle}{dt} = H|p\rangle = E_p|p\rangle \Rightarrow |p(t)\rangle = e^{iE_p t} |p(0)\rangle$

Things look better in the H. picture, with t -dep assigned to the op.s.

$$O_H(t) = e^{iHt} O_S e^{-iHt} \quad \frac{dO_H(t)}{dt} = i[H, O_H]$$

$O_H = O_S$ at $t=0$. In this language, op. satisfy equal + comm relations.

$$[\phi(x, t), \phi(y, t)] = 0 = [\pi(x, t), \pi(y, t)]$$

$$[\phi(x, t), \pi(y, t)] = i \delta^3(x-y)$$

We can check H. eq for ϕ , i.e. $\frac{d\phi}{dt} = i[H, \phi]$ means that the H of ϕ satisfies the K.G. e.g. $\partial_\mu \partial^\mu \phi + m^2 \phi = 0$

We write the Fourier transform of $\phi(x)$ by using

$$e^{iHt} a_p e^{-iHt} = e^{-iE_p t} a_p \quad [H, a_p] = -E_p a_p$$

$$e^{iHt} a_p^\dagger e^{-iHt} = e^{iE_p t} a_p^\dagger$$

gives $\phi(x, t) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{E_p}} \left(a_p e^{-iE_p t + i\vec{p}\cdot\vec{x}} + a_p^\dagger e^{iE_p t + i\vec{p}\cdot\vec{x}} \right)$

← absorbs the energy factor

Causality

ϕ, π satisfy equal time comm. relations

$$[\phi(x, t), \pi(y, t)] = i\delta^3(x-y) \quad \text{What if we chose } \pi(y, t')?$$

In particular, causality requires that all space-like operators commute.

$$\text{i.e. } [O(x), O(y)] = 0 \quad \forall (x-y)^2 < 0$$

↑
4-vector

This ensures that measurements at x can't affect measurements at y .

Do we have this in our set-up.

Define $\Delta(x-y) = [\phi(x), \phi(y)]$

RHS ops LHS $c-p^n$ ← what do we know about this?

$$\Delta(x-y) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{E_p}} \left(e^{-ip(x-y)} - e^{ip(x-y)} \right)$$

L.I. L.I. We know it is L.I.

• It doesn't vanish for timelike separation

$$[\phi(x, 0), \phi(x, t)] \sim e^{-imt} - e^{imt}$$

• It vanishes for space-like separation

Note that $\Delta(x-y) = 0$ at equal times. But L.I. \Rightarrow can only depend on $(x-y)^2$ so it must vanish $\forall (x-y)^2 < 0$.

M.B.

$$[\phi(x, t), \phi(y, t)] = \frac{i}{2} \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{p^2 + m^2}} \left(e^{ip \cdot (x-y)} - e^{-ip \cdot (x-y)} \right) \quad \oplus$$

$= 0$

↑ change sign of p integral to flip the sign here

- Also holds in interacting theory. But comm. is only a c fⁿ in free theory.