

# WS2: Finite Differencing and Interpolation

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## 1 Finite Difference Approximation and Convergence

### 1.1 Forward Differencing vs. Central Differencing

I computed the first derivative of  $f(x) = x^3 - 5x^2 + x$  on the interval  $[-2,6]$  using finite differencing and central differencing. Figure 1 shows the error on the computed derivative using forward differencing, where the error is the differencing between the computed value of the derivative and the analytic value. Two step sizes were used to compute the derivative,  $dx=0.1$  and  $dx=0.2$ . Since central differencing is first order, the errors for  $dx=0.1$  are one half the errors for  $dx=0.2$ .

Figure 2 shows the same quantities as Figure 1, but for central differencing. Central differencing is second order, so the errors for  $dx=0.1$  are one quarter of the errors for  $dx=0.2$ .

### 1.2 Finite Difference Estimate of the Second Derivative

To get a central-differencing estimate for the second derivative of  $f$ , we start with the Taylor expansions for  $f(x_0 + h)$  and  $f(x_0 - h)$ :

$$f(x_0 + h) = f(x_0) + hf'(x_0) + \frac{h^2}{2}f''(x_0) + \frac{h^3}{6}f'''(x_0) + O(h^4) \quad (1)$$

$$f(x_0 - h) = f(x_0) - hf'(x_0) - \frac{h^2}{2}f''(x_0) - \frac{h^3}{6}f'''(x_0) + O(h^4) \quad (2)$$

By adding these two formulas and then solving for  $f''(x_0)$ , we get:

$$f''(x_0) = \frac{f(x_0 + h) - 2f(x_0) + f(x_0 - h)}{h^2} + O(h^2) \quad (3)$$

## 2 Interpolation: Cepheid Lightcurve

Figure 3 shows the measured lightcurve of a Cepheid variable star with a period of one day (measurements denoted by asterisks). The same figure also shows

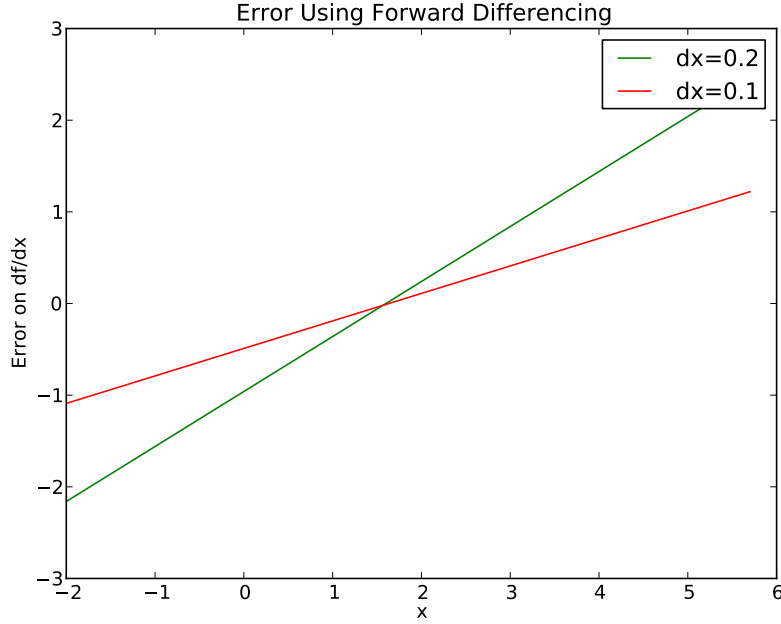


Figure 1: Error on the first derivative of  $f(x) = x^3 - 5x^2 + x$  computed using forward differencing. The error on forward differencing for step size  $dx$  is  $O(dx)$ .

3 methods of interpolation: Lagrange interpolation (an  $n=8$  polynomial fit using all 9 data points), piecewise linear interpolation, and piecewise quadratic interpolation. The piecewise quadratic interpolation uses 3 data points to interpolate a magnitude  $p(t)$  at time  $t$ . I chose to use the two nearest data points at time less than or equal to  $t$ , plus the nearest data point at time greater than  $t$ . The  $n=8$  polynomial has very large errors far from the center of the data. The piecewise linear and quadratic interpolation both have sharp changes in derivative where the interpolated segments meet at the data points.

Figure 4 shows 2 additional methods of interpolation: piecewise cubic Hermite interpolation, and spline interpolation. The Hermite interpolation improves upon the previous methods since it accounts for the derivative at the data points. However, the spline provides the smoothest fit.

For all the methods except the  $n=8$  polynomial, I took advantage of the periodicity of the data by extending the data set by one period (repeating the same values) on either end of the range of the interest from 0 to 1 days, in order to avoid dealing with edge effects. If I do this for the polynomial, and increase the degree to be one less than the number of data points (including those repeated), it looks much improved but still has some alarming wiggles.

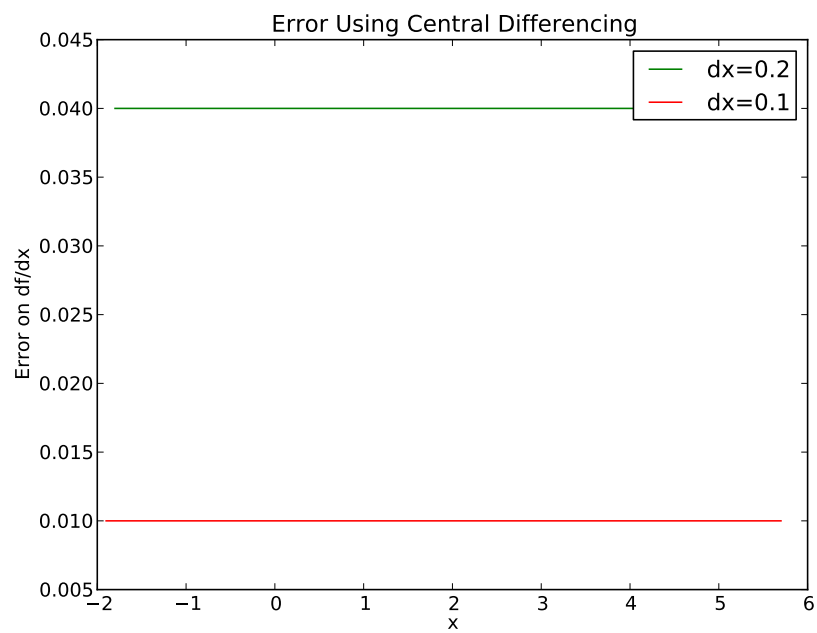


Figure 2: Error on the first derivative of  $f(x) = x^3 - 5x^2 + x$  computed using central differencing. The error on central differencing for step size  $dx$  is  $O(dx^2)$ .

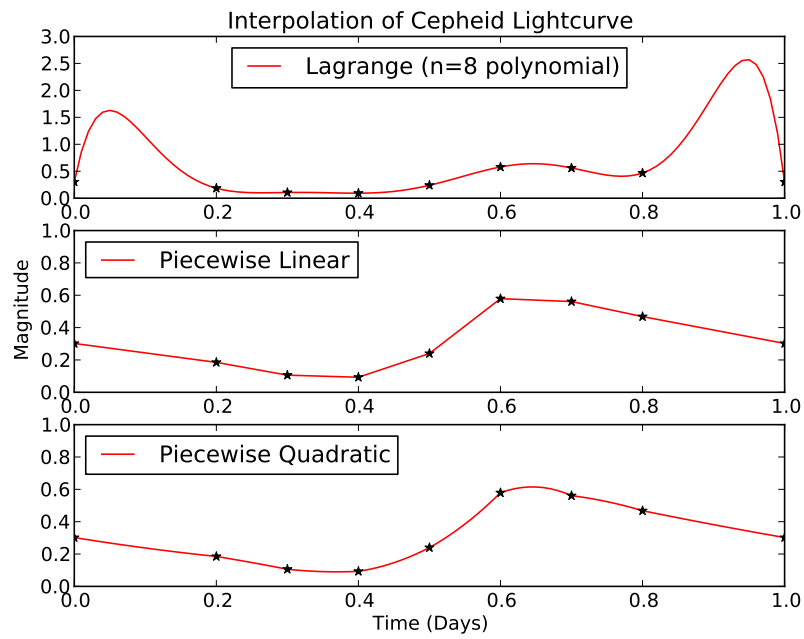


Figure 3: Cepheid lightcurve with a period of one day. 3 methods of interpolation shown: Lagrange ( $n=8$  polynomial) interpolation, piecewise linear interpolation, and piecewise quadratic interpolation.

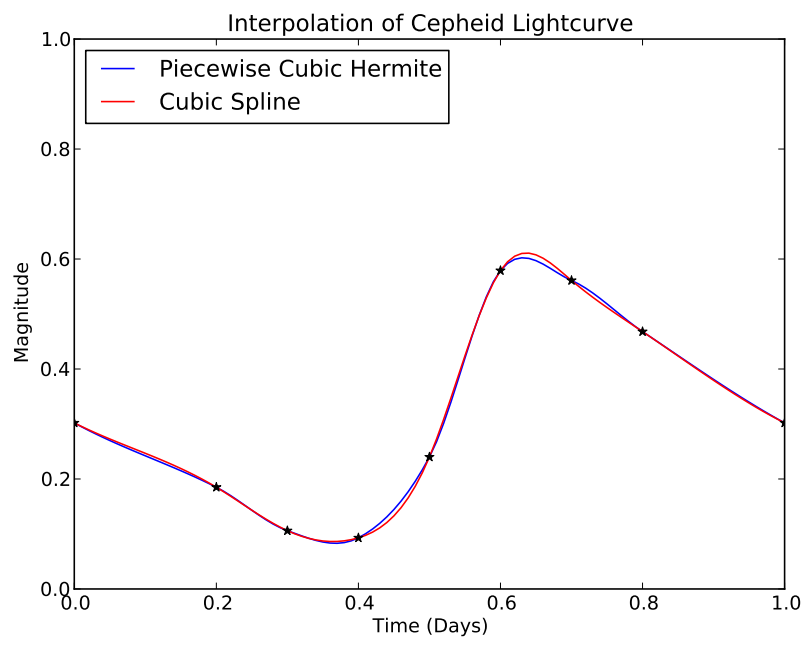


Figure 4: Cepheid lightcurve with a period of one day. 2 methods of interpolation shown: Piecewise cubic Hermite interpolation, and cubic spline interpolation.