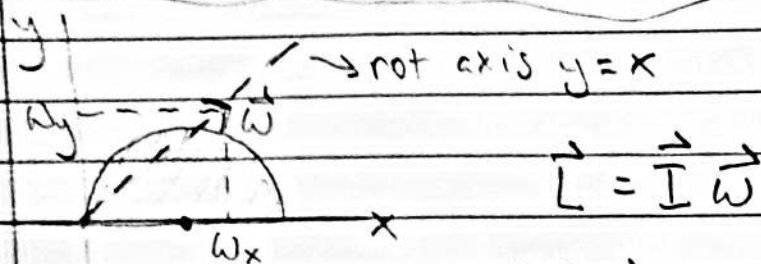


HW #32 problem 1



$$\vec{\omega} = \omega \begin{pmatrix} 1/\sqrt{2} \\ 1/\sqrt{2} \\ 0 \end{pmatrix}$$

$$\omega_x = \omega/\sqrt{2} \quad \omega_y = \omega/\sqrt{2} \\ \omega_z = 0$$

$$I = \begin{pmatrix} I_{yy} & -I_{xy} & 0 \\ -I_{xy} & I_{xx} & 0 \\ 0 & 0 & I_{yy} + I_{xx} \end{pmatrix}$$

since we have no
z coordinate

$$I_{xy} = \iint xy \sigma dx dy$$

integrate along x

dir from 0 to 2R.

For y integral, from 0
to the curve.

$$\sigma = \frac{M}{A} = \frac{M}{\frac{1}{2} \pi R^2}$$

$$I_{xy} = \int_0^{2R} \int_0^{\sqrt{x(2R-x)}} xy \sigma dy dx = \frac{2R^4 \sigma}{3}$$

$$(x-R)^2 + y^2 = R^2$$

$$I_{yy} = \int_0^{2R} \int_0^{\sqrt{x(2R-x)}} y^2 \sigma dy dx = \frac{1}{8} \pi R^4 \sigma$$

$$I_{xx} = \int_0^{2R} \int_0^{\sqrt{x(2R-x)}} x^2 \sigma dy dx = \frac{5}{8} \pi R^4 \sigma$$

Thus

$$I = \begin{bmatrix} \frac{1}{2} \pi R^4 \sigma & -\frac{2R^4 \sigma}{3} & 0 \\ -\frac{2R^4 \sigma}{3} & \frac{5}{8} \pi R^4 \sigma & 0 \\ 0 & 0 & \frac{6}{8} \pi R^4 \sigma \end{bmatrix}$$

$$\vec{L} = I \vec{\omega}$$

$$\vec{L} = \begin{bmatrix} \omega/\sqrt{2} \\ \omega/\sqrt{2} \\ 0 \end{bmatrix} = \begin{pmatrix} \frac{\omega \pi R^4 \sigma}{2\sqrt{2}} - \frac{2R^4 \omega \sigma}{3\sqrt{2}} \\ -\frac{2\omega R^4 \sigma}{3\sqrt{2}} + \frac{5\omega \pi R^4 \sigma}{\sqrt{2} 8} \\ 0 \end{pmatrix} = \begin{matrix} L_x \\ L_y \\ L_z \end{matrix}$$

② In the body frame, \vec{L} is "attached" to $\vec{\omega}$, with both precessing about \hat{e}_3 at some rate Ω . \hat{e}_3 stays constant in this frame.

In the space frame, \hat{e}_3 is no longer constant, but moves around a fixed \vec{L} along w/ $\vec{\omega}$ at some rate Ω_s . So, in both frames $\vec{\omega}$ is not constant.

③ Euler's Equations refer to axes fixed in the body frame. The axes are awkward to work with. The Lagrangian is set up to specify orientation of the body to a non-rotating frame, which is a lot easier to deal with.