

# Week 14 Worksheet Solutions

## Black Holes

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The Kruskal coordinates  $V, U$  are defined by

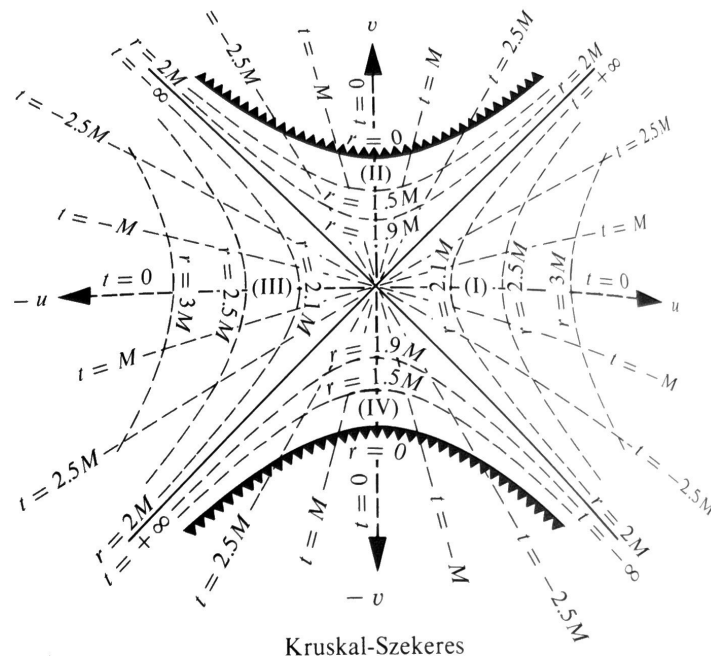
$$\left(\frac{r}{2M} - 1\right) e^{r/2M} = U^2 - V^2$$

$$\frac{t}{2M} = \ln\left(\frac{V+U}{U-V}\right) = 2 \tanh^{-1}(V/U).$$

The Schwarzschild metric in Kruskal coordinates is

$$ds^2 = \frac{32M^3 e^{-r/2M}}{r} (-dV^2 + dU^2) + r^2(d\theta^2 + \sin^2\theta d\varphi^2).$$

The Kruskal diagram of a black hole is



**Exercise.**

- Identify the worldlines of photons traveling radially in a Kruskal diagram.
- Show that the worldline of a photon traveling nonradially makes an angle of less than 45 degrees with the vertical axis of the Kruskal diagram.
- Use part (b) to show that particles with finite mass always move at an angle less than 45 degrees with the vertical axis.
- If someone falls past the radius  $r = 2M$ , he or she will always hit the singularity at  $r = 0$ .
- Once someone has fallen past  $r = 2M$ , he or she can't send messages to friends located at  $r > 2M$  but can still receive messages.
- Show that once someone falling in reaches the gravitational radius  $r = 2M$ , then *no matter what he or she does subsequently—no matter in what direction, how long, and how hard he or she blasts his or her rocket engines*—he or she will be killed by the singularity at  $r = 0$  in a proper time of

$$\tau < 1.54 \cdot 10^{-5} \frac{M}{M_{\odot}} \text{ seconds,}$$

where  $M_{\odot} = 2 \cdot 10^{30}$  kg is the mass of the Sun (and  $G = 6.7 \cdot 10^{-11} \text{ m}^3/\text{kg s}^2$ ).

*Hint:* Note that

$$\left(\frac{dr}{d\tau}\right)^2 = e^2 - \left(1 - \frac{2M}{r}\right) \left(1 + \frac{\ell^2}{r^2}\right).$$

- Radial null geodesics are those with  $d\varphi = 0$  and  $d\theta = 0$ , so

$$0 = -dV^2 + dU^2.$$

Hence,

$$\frac{dV}{dU} = \pm 1$$

for such geodesics.

- Write the line element for photons

$$0 = \frac{32M^3}{r} e^{-r/2M} [-(p^V)^2 + (p^U)^2] + r^2 [(p^\theta)^2 + (p^\varphi)^2].$$

Since the second term in brackets is  $> 0$  for nonradial travel, so  $p^V > |p^U|$ .

- The null geodesics define the light cone, so if they have less than a 45 degree angle with the vertical axis, then necessarily so will timelike geodesics which correspond to massive particles.

- d) Follows immediately from the diagram and part (c) together with causality. Namely, we have identified that

$$\left| \frac{dV}{dU} \right| \geq 1$$

for light rays; however, this does not tell us about the *direction* that light rays. If we identify that light rays travel “up” in the Kruskal diagram, then we will have identified the future/past light cones. We can do this by analyzing the formula

$$\frac{t}{2M} = \ln \left( \frac{V + U}{U - V} \right).$$

Consider two points  $(U_1, V_1)$  and  $(U_2, V_2)$ . Then

$$t_2 - t_1 = 2M \ln \left( \frac{(U_2 + V_2)(U_1 - V_1)}{(U_2 - V_2)(U_1 + V_1)} \right).$$

Now, let  $(U_1, V_1) = (0, 1)$  and  $(U_2, V_2) = (1, 2)$ . Then we find that

$$t_2 - t_1 > 0.$$

Similarly, if  $(U_1, V_1) = (0, 1)$  and  $(U_2, V_2) = (-1, 2)$ , we again find that

$$t_2 - t_1 > 0.$$

It follows that on slope  $\pm 1$  lines light travels up (other such lines will be displaced by constants from these). Similarly, one checks that on slope  $> 1$  lines light travels up with a  $\Delta t$  larger than in the slope 1 case, implying that light always travels up. This implies the same is true for massive particles, since we have found the future light cones to be pointing up.

- e) Follows immediately from the diagram and part (b).  
 f) This problem is effectively on the homework, so I won't post solutions to it.