

# 't Hooft-Polyakov Monopoles

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## Abstract

In this work we showed the fundamental steps on Dirac and 't Hooft Poljakov monopolar theories, deducing charge quantization for electromagnetism and generalizing the approach per non abelian gauge theories under Higgs mechanism.

## 1 Introduction

Electric charge quantization has been proved experimentally by R.A. Millikan in 1909: from that date, all the experiments that followed confirmed this hypothesis. In fact, Maxwell's equations in their classical formulation do not formally admit the existence of magnetic monopole and since today we have no proof of its empirical existence. Nonetheless, several physicists (Pierre Curie for the first time) reflected about the reason of asymmetry between electric and magnetic charge in nature. In 1931 Dirac showed that from an appropriate duality operation in Maxwell equations it follows magnetic charge quantization, which implies formal possibility for the existence of magnetic monopole in a dual electromagnetic theory. A different approach has been introduced in QFT by 't Hooft and Poljakov in 1975, deducing charge quantization in a general non abelian gauge theory. This approach is different from Dirac one, since it obtains quantization from the lagrangian action of the theory considered and some reasonable physical assumptions, in both static and extended dynamic case. Nonetheless, it has been shown that this theory can explain Dirac results for QED, but can go further and explain quantization, for instance, for electroweak mechanism and other non abelian gauge theory.

In this paper we will recall the fundamental steps of monopolar theories, starting from Dirac quantization for the electromagnetic case and then...

## 2 Dirac Monopoles

In 1931 Dirac showed that postulating the existence of a magnetic charge, we can provide a theoretical explanation for the quantisation of the electric charge [1]. In this section we will recover the basic ideas exposed in Dirac's paper and show how they were generalised to a particle with both charges, called *dyon*, by Schwinger in 1966 [2] and Zwanziger in 1968 [3].

### 2.1 Magnetic Monopole and Electromagnetic Duality

The electromagnetic tensor  $F^{\mu\nu}$ , where both electric and magnetic fields are encoded:

$$E^i = F^{0i} \quad B^i = \epsilon^{ijk} F^{jk}, \quad (2.1)$$

is an antisymmetric tensor and can be written as a 2-form (see appendix A):

$$F = \frac{1}{2} F_{\mu\nu} dx^\mu \wedge dx^\nu \quad (2.2)$$

The tensor  $F^{\mu\nu}$  obeys Maxwell's equations:

$$\partial_\mu F^{\mu\nu} = j^\nu \quad (2.3a)$$

$$\partial_{[\mu} F_{\nu\rho]} = 0. \quad (2.3b)$$

The duality operation (see appendix A) acts on the electromagnetic tensor as follows:

$$\tilde{F}^{\mu\nu} = *F^{\mu\nu} = -\frac{1}{2} \epsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}, \quad (2.4)$$

and allows us to write Maxwell's equations in this way:

$$dF = 0 \quad d * F = *j, \quad (2.5)$$

where we introduced the current form  $j = j_\mu dx^\mu$ . Now, it is clear to see that, if  $j = 0$ , equations 2.5 are invariant with respect to the duality transformation  $F \rightarrow *F$ , since  $*^2 = -1$  for 2-forms. Such symmetry is broken when  $j \neq 0$ : however, we can preserve invariance under duality when  $j \neq 0$  if we introduce a “magnetic current”  $j_m$  such that  $dF = *j_m$ , and complement the duality transformation as such:

$$\begin{cases} j \rightarrow j_m \\ j_m \rightarrow -j \end{cases} \quad (2.6)$$

The electromagnetic duality transformation corresponds to the formal substitution of  $(\vec{E}, \vec{B}) \rightarrow (\vec{B}, -\vec{E})$ :

$$\tilde{E}^i = \tilde{F}^{i0} = \frac{1}{2} \epsilon^{i0jk} F_{jk} = -\frac{1}{2} \epsilon^{ijk} F^{ik} = B^i \quad (2.7a)$$

$$\tilde{B}^i = -\frac{1}{2} \epsilon^{ijk} \tilde{F}_{jk} = -\frac{1}{2} \epsilon^{ijk} \epsilon^{jkl0} F_{l0} = -\frac{1}{2} \epsilon^{ijk} \epsilon^{ljk} E^l = -E^i \quad (2.7b)$$

If we consider the configuration where an electrically charged point-like particle lies in the origin, the electric and magnetic field are:

$$E^i = \frac{e}{2r^2} x^i, \quad (2.8a)$$

$$B^i = 0, \quad (2.8b)$$

and satisfy:

$$\partial_i E^i = 2\pi e \delta^3(\vec{r}), \quad (2.9a)$$

$$e = \frac{1}{2\pi} \int_{S^2} E_i dS^i. \quad (2.9b)$$

Now, we might apply a duality transformation to this configuration and obtain a magnetic point-like monopole, which has charge  $m$  and is placed in the origin. Now the fields are:

$$B^i = \frac{m}{2r^2} x^i \quad (2.10a)$$

$$E^i = 0, \quad (2.10b)$$

and correspond to this electromagnetic tensor:

$$F_{ij} = \frac{m}{2r^3} \varepsilon_{ijk} x^k \quad F_{i0} = 0. \quad (2.11)$$

We now compute the 2-form associated to the tensor in equation 2.11:

$$F = F_{ij} dx^i dx^j = \frac{m}{4r^3} \varepsilon_{ijk} x_k dx^i dx^j = \frac{m}{2} \sin \theta d\theta \wedge d\phi, \quad (2.12)$$

and deduce from it the value of the magnetic charge in the origin:

$$\int_{\mathbb{R}^3} dF = \int_{S_\infty^2} F = \frac{m}{2} \int_{S_\infty^2} \sin \theta d\theta \wedge d\phi = 2\pi m. \quad (2.13)$$

In the first passage of equation 2.13, we applied Stokes' theorem to an arbitrarily large sphere, because the integrand function has no radial dependence.

## 2.2 Dirac Quantization for the Monopole

We cannot find for the magnetic monopole fields, represented by the 2-form  $F$ , a global vector potential  $A$  such that  $F = dA$ . If it existed, we would have the following paradox:

$$m = \int_{\mathbb{R}^3} \frac{dF}{2\pi} = \int_{S_\infty^2} \frac{F}{2\pi} = \int_{S_\infty^2} \frac{dA}{2\pi} = \int_{\partial S_\infty^2} \frac{A}{2\pi} = 0 \neq m, \quad (2.14)$$

Even though the form  $F$  is closed ( $dF = 0$ ), it is not exact. This happens because the manifold we are considering, is not contractible and therefore Poincaré's lemma cannot be applied.

We can, however, find potentials which are defined *locally*: a possible choice of  $A$  is, in cylindric coordinates,

$$A = \frac{m}{2} (c - \cos \theta) d\phi, \quad (2.15)$$

where  $c = \pm 1$ . Differentiating this equation we have

$$dA = \frac{m}{2} (c - \cos \theta) d^2\phi - \frac{m}{2} d\cos \theta \wedge d\phi = \frac{m}{2} \sin \theta d\theta \wedge d\phi, \quad (2.16)$$

where the first term vanishes because  $d^2 = 0$ . We will show that for both choices of  $c$  the potential is singular somewhere:

- for  $c = 1$  the potential  $A$  is singular along the  $z < 0$  axis;
- for  $c = -1$ , instead,  $A$  is singular along the  $z > 0$  axis.

The  $z \leq 0$  ray is called a *Dirac string*. In both cases, we see that removing the string makes the topology trivial: the manifold becomes contractible, and the local potential is defined on the whole of either open set  $\mathbb{R}^3 \setminus \{z \leq 0\}$ . We call the two choices of  $A$  respectively  $A^+$  and  $A^-$ . We say that  $U_+ = \mathbb{R}^3 \setminus \{z < 0\}$  is the chart where we defined  $A^+$  and  $U_- = \mathbb{R}^3 \setminus \{z > 0\}$  is the one where we defined  $A^-$ .

Let us now show that the potentials indeed diverge near their Dirac string. We wish to integrate in a similar domain to the one of equation (2.14), however now we must exclude the Dirac string: the resulting domain is a 2D surface, which we denote by  $S_2 - D_\epsilon$ ; its boundary is a small circle  $C_\epsilon$  around the Dirac string.

$$2\pi m = \int_{S_2} F = \lim_{\epsilon \rightarrow 0} \int_{S_2 - D_\epsilon} F = \lim_{\epsilon \rightarrow 0} \int_{S_2 - D_\epsilon} dA = \lim_{\epsilon \rightarrow 0} \int_{C_\epsilon} A, \quad (2.17)$$

which implies the divergence of  $A$ , since its integral along an arbitrarily small circle is requested to be fixed and finite.

We now want to calculate the transition function between the two potentials in the region where their domains overlap  $U_+ \cap U_-$ :

$$A^+ - A^- = m d\phi = d(m\phi) \equiv d\lambda(x). \quad (2.18)$$

This transformation can be seen as a  $U(1)$  gauge transformation: let  $h$  be the map  $h : U_+ \cap U_- \rightarrow G = U(1)$ , such that  $x \rightarrow \exp\left(\frac{ie\lambda(x)}{2\pi}\right)$ , where  $e$  is the electric charge. The corresponding transformation of the potential is:

$$A^+ = A^- + d\lambda = h^{-1} A^- h - \frac{i}{e} h dh. \quad (2.19)$$

In order to have the function  $h$  well defined, we must have  $h(m\phi) = h(m(\phi + 2\pi))$ . This condition is satisfied if we assume that:

$$me \in 2\pi\mathbb{Z}. \quad (2.20)$$

Expression (2.20) is the so-called *Dirac Quantization Condition* and guarantees that:

$$h(m(\phi + 2\pi)) = \exp(2\pi me) h(m\phi) = h(m\phi). \quad (2.21)$$

We now want to show that the integer  $em$  has a clear geometrical meaning: it is the winding number of the map  $h$ , i.e. the number of times  $h$  wraps around the circle  $S^1$  as  $x$  goes around the center once. To prove this, we can split the integral in (2.17) into two contributions from both the potentials and integrate them over two hemispheres, whose borders can both be made to coincide with

the equator  $S^1$ , but will have opposite orientations, whence the minus sign before  $A_-$ :

$$m = \int_{S^2} \frac{F}{2\pi} = \int_{S^2_+} \frac{dA^+}{2\pi} + \int_{S^2_-} \frac{dA^-}{2\pi} = \int_{S^1} \frac{A^+ - A^-}{2\pi} = \int_{S^1} \frac{d\lambda}{2\pi} = \frac{\Delta\lambda}{2\pi}, \quad (2.22)$$

Multiplying  $\Delta\lambda$  by the charge  $e$ , we can deduce how many radians the argument of the exponent has elapsed. Therefore, the winding number is:

$$S_1 = \frac{e\Delta\lambda}{2\pi} = me. \quad (2.23)$$

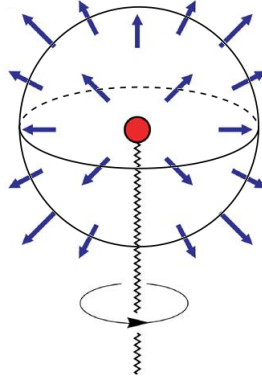


Figure 1: A schematization of Dirac string

### 2.3 Zwanziger-Schwinger Quantization for the Dyon

Let  $L$  be the now orbital angular momentum of a dyon with mass  $m$  and charge  $e$ . This particle experiments a variation of  $L^i = \epsilon^{ijk} r_j m \dot{r}_k$ , such that

$$\frac{dL^i}{dt} = \epsilon^{ijk} r_j m \ddot{r}_k = \epsilon^{ijk} r_j \epsilon^{klm} e \dot{r}_l B_m = \epsilon^{ijk} r_j \epsilon^{klm} e \dot{r}_l \frac{g}{4\pi r^3} B_m = \frac{d}{dt} \left( \frac{egr^i}{4\pi r} \right). \quad (2.24)$$

This implies that a particle into the monopolar field conservs another quantity that is the whole momentum

$$J^i = L^i - \frac{egr^i}{4\pi r}. \quad (2.25)$$

We define now a dyon as an hypothetic particle which has both electric and magnetic charge, respectively  $e$  and  $g$ . Suppose now to have another dyon

with charges  $e'$  and  $g'$ , electric and magnetic. In this case we can hold a similar reasoning to obtain the equivalent condition on charges involved: since

$$\vec{J}_{em} = \int_{\mathbb{R}^3} d^3r \vec{r} \times (\vec{E} \times \vec{B}), \quad (2.26)$$

which is the angular momentum of fields generated by the particles involved, we can use the relation that holds quantization

$$|\vec{J}_{em}| = \frac{n}{4\pi} \quad (2.27)$$

where  $n \in \mathbb{Z}$ . From this it naturally follows:

$$\frac{q_1 g_2 - q_2 g_1}{2\pi\hbar} \in \mathbb{Z}. \quad (2.28)$$

### 3 't Hooft-Polyakov Monopoles

In this section, it will be shown that free magnetic monopoles in the physical vacuum are possible as regular solutions of the field equations of the Georgi-Glashow model, following the notorious works published by Gerard 't Hooft [8] and Aleksander Polyakov [7] in 1974.

#### 3.1 An Intuitive Example

A simple and very clarifying example, taken from [8], will be discussed in order to give an intuitive idea of how it is possible to have monopoles in quantum field theory.

We should consider a two-dimensional spherical surface  $S^2$  in the usual 3D space, with some magnetic flux  $\Phi$  entering only in a spot. If the spot is surrounded by a contour  $C_0$  along which the fields are null, the potential around  $C_0$  should be equal to:

$$A_i = \partial_i \Lambda \quad (3.1)$$

where  $\Lambda$  is a gauge transformation, which acts on the wavefunction in this way:  $\psi \rightarrow e^{i\Lambda} \psi$ . While  $\Lambda$  is multivalued,  $\Phi$ , being a physical field, is required to be a single-valued function. Since,

$$\Phi = \int_{C_0} \partial_i \Lambda dl^i, \quad (3.2)$$

$\Phi$  should be equal to an integer multiplied by  $2\pi$ , a complete gauge rotation along the contour. In an Abelian gauge theory, there must be another spot from which the flux comes out, because the  $2k\pi$  rotation cannot be continuously changed into a constant lowering the contour  $C_0$  over the sphere.

Instead, in a non Abelian gauge theory, with a compact covering group, such as  $O(3)$ , a  $2k\pi$  rotation with even  $k$  can be shifted into a constant without singularity. This implies that if we have a non Abelian gauge theory with a compact covering group and such compact covering group has the electromagnetic group  $U(1)$  as a subgroup, the existence of the magnetic monopole would be allowed without supposing the existence of any singularity anywhere on the sphere. A theory that satisfies those requirements is the Georgi-Glashow model for the electroweak interaction.

### 3.2 The bosonic part of the Georgi-Glashow Model

The Georgi-Glashow model was proposed as a theory for the electroweak interaction and, in its bosonic part, is based on the gauge group  $SU(2)$  which two-to-one homomorphic to  $SO(3)$ .

The lagrangian density  $\mathcal{L}$ , whose associated action is invariant under gauge transformation belonging to  $SO(3)$ , contains the Higgs field  $\vec{\phi} = (\phi^1, \phi^2, \phi^3)$ , which is a vector in the adjoint representation of  $SO(3)$  and three gauge potentials  $\vec{W}^\mu$ , with values in the Lie algebra of  $SO(3)$ . We use superscript arrows to denote objects which take values in the adjoint representation of  $SO(3)$ .  $\mathcal{L}$  has the following form:

$$\mathcal{L} = \frac{1}{2} D_\mu \phi^a D^\mu \phi^a - \frac{1}{4} G_{\mu\nu}^a G^{a\mu\nu} - \underbrace{\frac{1}{4} \lambda (\phi^2 - a^2)}_{V(\phi)}, \quad (3.3)$$

where  $\vec{G}_{\mu\nu}$  is the gauge field-strength, defined as follows:

$$\vec{G}_{\mu\nu} = \partial_\mu \vec{W}^\nu - \partial_\nu \vec{W}^\mu - e \vec{W}^\nu \times \vec{W}^\mu, \quad (3.4)$$

and the operator  $D_\mu$  is the gauge covariant derivative:

$$D_\mu \vec{\phi} = \partial_\mu \vec{\phi} - e \vec{W}_\mu \times \vec{\phi}. \quad (3.5)$$

Let us look at the physical meaning of each term in the lagrangian density:

- The term  $\frac{1}{2} D_\mu \phi^a D^\mu \phi^a$  is the free field term;
- The term  $-\frac{1}{4} G_{\mu\nu}^a G^{a\mu\nu}$  describes the dynamics of the potential  $\vec{W}_\mu$ ;
- The term  $V(\phi)$  describes the self interaction of the Higgs field and contains a non negative constant  $\lambda$ .

**Spectrum of the Model** Now we are interested in the spectrum of the model. It can be obtained through perturbation theory, considering a Higgs potential with little fluctuation around the minimum  $\vec{a}$ :

$$\vec{\phi} = \vec{a} + \vec{\chi}. \quad (3.6)$$



Four bosons figure in our model: the photon  $A_\mu = \frac{1}{a} \vec{a} \cdot \vec{W}_\mu$ , the Higgs Boson, associated to the field  $\vec{\phi}$  and the bosons  $\vec{W}_\mu^\pm$ . Their properties are shown in table 1.

Field	Definition	Mass	Charge
$A_\mu$	Photon	0	0
$\phi$	Higgs Boson	$a\sqrt{2\lambda}\hbar$	0
$W_\mu^\pm$	Two Massive Bosons	$ae\hbar$	$\pm e\hbar$

Table 1: Bosonic Spectrum of the Georgi-Glashow model.

For the sake of simplicity, we choose as a minimum the vector  $\vec{a} = (0, 0, a)$ . Then, working in the so-called unitary gauge, the fluctuation vector  $\vec{\chi}$  can be written in such a way that  $\vec{\chi} = (0, 0, \chi)$ . At this point, we should expand the Higgs field around  $\vec{a}$  in order to read the masses of the bosons from the quadratic term coefficients:

$$V(\phi) \simeq \frac{1}{2} \left( \sqrt{2\lambda}a \right)^2 \chi^2 = \left( \frac{M_\chi}{\hbar} \right)^2 \chi^2 \quad (3.7a)$$

$$\frac{1}{2} D_\mu \vec{\phi} \cdot D^\mu \vec{\phi} \simeq \frac{1}{2} D_\mu \vec{a} \cdot D^\mu \vec{a} + \frac{1}{2} (ea)^2 = \frac{1}{2} D_\mu \vec{a} \cdot D^\mu \vec{a} + \frac{1}{2} \left( \frac{M_W}{\hbar} \right)^2 \quad (3.7b)$$

To deduce the charges of the bosons, instead, we should see how the covariant derivative, which describes the minimal coupling for the photon  $A_\mu$ :

$$\nabla_\mu = \partial_\mu + i \frac{Q}{\hbar} A_\mu \quad (3.8)$$

is embedded in the SO(3) covariant derivative. Then, the value given to  $Q$  for each boson, represents the charge of the bosons.

**Vacuum Configurations** Another aspect of the Georgi-Glashow model, which is relevant for understanding the 't Hooft-Polyakov monopoles, is the study of the vacuum configurations.

Using Noether's theorem, the stress energy tensor can be calculated. To be more specific, we are interested in its component  $T^{00}$ , which describes the energy density:

$$T^{00} = \frac{1}{2} \vec{E}^i \cdot \vec{E}^i + \frac{1}{2} \vec{B}^i \cdot \vec{B}^i + \frac{1}{2} D_0 \vec{\phi} \cdot D_0 \vec{\phi} + \frac{1}{2} D_i \vec{\phi} \cdot D_i \vec{\phi} + V(\phi), \quad (3.9)$$

where  $\vec{E}^i = -\vec{G}^{0i}$  and  $\vec{B}^i$  is defined through this relation:  $\vec{G}_{ij} = \epsilon_{ijk} \vec{B}_k$ .

Imposing that the energy density  $T_{00}$ , contained in expression 3.9, is null, we get the conditions for the vacuum configurations:

$$\vec{G}_{\mu\nu} = 0 \quad D^\mu \vec{\phi} = 0 \quad V(\phi) = 0 \quad (3.10)$$

The last two of those conditions give the so-called *Higgs Vacuum* configurations. In the Higgs vacuum, the Higgs field is such that  $\vec{\phi}^2 = \vec{a}^2$ . It follows that such vacuum configurations are not invariant under the whole  $\text{SO}(3)$  but under the subgroup  $\text{SO}(2) \simeq \text{U}(1)$ , which is the gauge group of electromagnetism. Hence, the Georgi-Glashow model undergoes a spontaneous symmetry breaking process.

### 3.3 The 't Hooft-Polyakov Ansatz

**Hypotheses** We are now interested at solutions of the Euler-Lagrange equations of the Georgi-Glashow model, characterised by the following properties:

1. Finite Energy;
2. Stability;
3. Spherical Symmetry;
4. Static Nature.

In order to satisfy the first condition, we require that the following integral has a finite value:

$$E = \int_{\mathbb{R}^3} T_{00} \, d^3x. \quad (3.11)$$

This is equivalent to asking that at spatial infinity the Higgs Field  $\vec{\phi}$  approaches the Higgs Vacuum, which is represented by the set  $\mathcal{M}_0 = \{\vec{\phi}(\vec{r}) \mid \vec{\phi}^2 = \vec{a}^2\}$ . Hence, denoting as  $\Sigma_\infty$  the spherical surface in  $\mathbb{R}^3$  with an infinite radius, we require the existence of a continuous function,

$$\phi_\infty: \Sigma_\infty \rightarrow \mathcal{M}_0, \quad (3.12)$$

such that:

$$\lim_{|\vec{r}| \rightarrow \infty} \vec{\phi}(|\vec{r}| \hat{r}) = \vec{\phi}_\infty(\hat{r}). \quad (3.13)$$

Those functions belong to the second homotopy group  $\Pi_2(\mathbb{R}^3)$  and can be classified according to their *winding number*, which is a topological invariant and specifies the number of times the map wraps around  $\mathcal{M}_0$ .

The second condition can be read as the request for non-dissipative solutions, that will never evolve in time to a configuration where  $\vec{\phi}_\infty$  is constant. In mathematical terms, we are asking for the *winding number* of the map  $\vec{\phi}_\infty$  to be different from 0.

The third and fourth conditions are imposed in order to simplify the problem: we will discuss a more general result, obtained without them, in section 3.4. By static nature of the solutions, we mean that that the fields have to be time-independent and that  $\vec{W}^{\mu=0} = 0$  at any time. The latter is more than a

simple gauge-fixing, because we require the time component of the gauge field to be null *at any time*, which equates to saying that the gauge-fixing transformation, which brings us to  $\vec{W}^{\mu=0} = 0$ , has to be time-independent. Taking the expression 3.9 for the energy density and imposing the static nature condition, we get that:

$$E = \int_{\mathbb{R}^3} T_{00} = - \int_{\mathbb{R}^3} \mathcal{L} = -L \quad (3.14)$$

**Serach for the solutions** The spherical symmetry condition, instead, allows us to claim that there might exist two real functions of the parameter  $\xi \equiv aer$ , where  $a$  is the module of the Higgs minimum  $\vec{a}$  and  $e$  is the electric charge:  $H(\xi)$  and  $K(\xi)$ . We rescale the radial variable with  $\xi$  for convenience in our future calculations. Hence, the ansatz for the solution to the Euler-Lagrange equations of the Georgi-Glashow model, might have this form:

$$\vec{\phi}(\vec{r}) = \frac{\vec{r}}{er^2} H(\xi) \quad (3.15a)$$

$$W_a^{\mu=i} = -\epsilon_{aij} \frac{r^j}{er^2} (1 - K(\xi)) \quad (3.15b)$$

$$W_a^{\mu=0} = 0. \quad (3.15c)$$

Plugging these fields into the Euler-Lagrange equations of the model, we obtain two differential equations describing the dynamics of  $H(\xi)$  and  $K(\xi)$ :

$$\xi^2 \frac{d^2 K}{d\xi^2} = KH^2 + K(K^2 - 1) \quad (3.16a)$$

$$\xi^2 \frac{d^2 H}{d\xi^2} = 2K^2 H + \frac{\lambda}{e^2} H(H^2 - \xi^2), \quad (3.16b)$$

In order to find the boundary conditions for the differential equations above, we should write the energy of the system as a function of  $H(\xi)$  and  $K(\xi)$ :

$$E = \frac{4\pi a}{e} \int_0^\infty \frac{d\xi}{\xi^2} \left( \xi^2 \frac{dH}{d\xi} + \frac{1}{2} \left( \xi \frac{dH}{d\xi} - H \right)^2 + \frac{1}{2} (K^2 - 1)^2 + K^2 H^2 + \frac{\lambda}{4e^2} (H^2 - \xi^2)^2 \right) \quad (3.17)$$

and require that such an integral has a finite value. It follows that:

$$\xi \rightarrow \infty: \quad \frac{H}{\xi} \rightarrow 1 \quad K \rightarrow 0 \quad (3.18a)$$

$$\xi \rightarrow 0: \quad H \sim O(\xi) \quad K \sim 1 + O(\xi) \quad (3.18b)$$

We should note that the condition that  $H(\xi)$  approaches spatial infinity linearly in  $\xi$  implies that asymptotically the Higgs field has the following form:

$$\vec{\phi}_\infty(\vec{r}) = \lim_{|\vec{r}| \rightarrow \infty} \frac{\vec{r}}{er^2} H(aer) = a\hat{r}. \quad (3.19)$$

This means that  $\vec{\phi}_\infty$  is homotopic to the identity and has a winding number equal to 1. Therefore, those boundary conditions also ensure the non-dissipative nature of the 't Hooft-Polyakov solution.

Another aspect that deserves to be discussed is the existence of a solution for the system of equations 3.16a and 3.16b, given the boundary conditions 3.18a and 3.18b. It has been proved that a solution actually exists [9].

**Asymptotic form for the 't Hooft-Polyakov Solutions** We wish now to study the behaviour of the solutions of equation 3.16a and 3.16b in the limits for  $r \rightarrow 0$  and  $r \rightarrow \infty$  and then finally show that they describe an object of finite size that at big distances behaves analogously to the Dirac Monopole.

In the  $r \rightarrow 0$  limit, we observe that the solutions for the fields, given the boundary conditions 3.18a and 3.18b, show no singularity:

$$\begin{cases} H \sim \xi^2 \\ K \sim 1 + C\xi^2 \end{cases} \implies \begin{cases} \vec{\phi} \sim \vec{r} \\ W_a^i \sim \epsilon_{aij} x^j \end{cases} \quad (3.20)$$

This implies that also  $\vec{G}_{\mu\nu}$  is smooth and regular in the origin. Such evidence is a relevant difference, between the Dirac Monopole where the fields show a singularity in the origin, and the 't Hooft-Polyakov Monopole where no singularity in the origin occurs.

In the limit  $r \rightarrow \infty$ , equations 3.16a and 3.16b, provided the boundary conditions 3.18a and 3.18b, have such form:

$$\begin{cases} \frac{d^2 K}{d\xi^2} = K \\ \frac{d^2 h}{d\xi^2} = \frac{2\lambda}{e^2} h \end{cases} \implies \begin{cases} K \sim e^{-\xi} = e^{-\frac{M_W r}{h}} \\ h \sim e^{-\xi \sqrt{\frac{2\lambda}{e^2}}} = e^{-\frac{M_H r}{h}} \end{cases} \quad (3.21)$$

where  $h := H - \xi$ . Hence, the object we are looking at has finite dimensions, comparable with the Compton wave lengths  $\frac{\hbar}{M_H}$  and  $\frac{\hbar}{M_W}$ .

In order to prove that such an object at spatial infinity behaves like a monopole, we should look at the electromagnetic field, which, in the Georgi-Glashow model is described by the following expression:

$$F_{\mu\nu} = \frac{1}{a} \vec{\phi} \cdot \vec{G}_{\mu\nu}, \quad (3.22)$$

and in the limit  $r \rightarrow \infty$  gives:

$$F_{0i} = 0 \quad F_{ij} = \epsilon_{ijk} \frac{r^k}{er^3}. \quad (3.23)$$

Hence, the asymptotic magnetic and electric fields are:

$$\vec{E} = 0 \quad \vec{B} = -\frac{1}{e} \frac{\vec{r}}{r^3}; \quad (3.24)$$

they are compatible with a magnetic charge of:

$$m = \frac{-4\pi}{e}, \quad (3.25)$$

which is admitted by the Dirac quantization condition.

Discutere  
un possibile  
commento  
sul fatto che  
tale carica  
sia la min-  
ima possi-  
bile

### 3.4 Extension to non-static and time-dependent solutions

In this section we will try to extend the 't Hooft-Polyakov ansatz to generic finite-energy and non-dissipative configurations. Now, unlike section 3.3, we do not require the solution to be static and spherically symmetric.

In the static and spherically symmetric case, we proved that the Higgs field  $\vec{\phi}$  approaches the Higgs vacuum at spatial infinity up to terms of the order of  $O(\exp(-r/R))$ , where  $R$  is the dimension of the monopole. Therefore, it seems reasonable to suppose that in a generic finite-energy configuration the solutions to the Georgi-Glashow model satisfy the Higgs vacuum condition everywhere except for a finite number of localised and compact regions, whose dimension is of the order of  $R$ .

We now consider a closed surface  $\Sigma$  that lies in the Higgs vacuum region and try to estimate the amount of magnetic charge in it. First, we should notice that in the Higgs vacuum, holding that  $\vec{\phi} \times \vec{W}_\mu = -\frac{1}{e} \partial_\mu \vec{\phi}$ , the field  $\vec{W}_\mu$  has the component orthogonal to  $\vec{\phi}$  that can be deduced from  $\vec{\phi}$ , whereas the one parallel to  $\vec{\phi}$  is unknown and will be called  $A_\mu$ . Hence, the field  $\vec{W}_\mu$  has this form:

$$\vec{W}_\mu = \frac{1}{a^2 e} \vec{\phi} \times \partial_\mu \vec{\phi} + \frac{1}{a} \vec{\phi} A_\mu. \quad (3.26)$$

Such an expression allows us to calculate the field-strength  $\vec{G}_{\mu\nu}$  and the electromagnetic tensor  $F_{\mu\nu} = \frac{1}{a} \vec{\phi} \cdot \vec{G}_{\mu\nu}$  in the Higgs vacuum, as a function of  $\vec{\phi}$  and  $A_\mu$ :

$$F_{\mu\nu} = \frac{1}{a^3 e} \vec{\phi} \cdot (\partial_\mu \vec{\phi} \times \partial_\nu \vec{\phi}) + \partial_\mu A_\nu - \partial_\nu A_\mu \quad (3.27)$$

Since  $\Sigma$  lies in the Higgs vacuum, the form of the magnetic field on  $\Sigma$  can be deduced from expression 3.27:

$$B_i = \epsilon_{ijk} F_{jk} = \epsilon_{ijk} \left[ \frac{1}{a^3 e} \vec{\phi} \cdot (\partial_j \vec{\phi} \times \partial_k \vec{\phi}) + \partial_j A_k - \partial_k A_j \right] \quad (3.28)$$

Therefore, the total magnetic charge  $g_\Sigma$  contained in  $\Sigma$  can be calculated as:

$$\begin{aligned} g_\Sigma &= \int_\Sigma \vec{B} \cdot d\vec{S} \\ &= -\frac{1}{2ea^2} \int_\Sigma \epsilon_{ijk} \vec{\phi} \cdot (\partial_j \vec{\phi} \times \partial_k \vec{\phi}) dS_i \end{aligned} \quad (3.29)$$

It can be proved that the quantity:

$$N_\Sigma := \frac{1}{8\pi a^3} \int_\Sigma \epsilon_{ijk} \vec{\phi} \cdot (\partial_j \vec{\phi} \times \partial_k \vec{\phi}) dS_i \quad (3.30)$$

is an integer and corresponds to the winding number of the map  $\vec{\phi}: \Sigma \rightarrow \mathcal{M}_0 = \{\vec{\phi} : \vec{\phi}^2 = a^2\}$ . Hence we have just derived a quantisation condition for the magnetic charge analogue to the Dirac's one apart from a  $\frac{1}{2}$  coefficient:

$$g_\Sigma = -\frac{4\pi}{e} N_\Sigma. \quad (3.31)$$

Furthermore we have that the magnetic charge in  $\Sigma$  depends only on a topological property of the map  $\vec{\phi}$  on  $\Sigma$  and is invariant under continuous deformation of  $\vec{\phi}$  preserving the Higgs vacuum, such as:

- Time Evolution of  $\vec{\phi}$ ;
- Gauge Transformations of  $\vec{\phi}$ ;
- Changes of  $\Sigma$  in the Higgs vacuum.

In order to obtain the Dirac Quantization condition, we should consider that the lagrangian from which we started (expression 3.3) is invariant under  $SO(3)$  even if the model is based on the gauge group  $SU(2)$ . This was possible thanks to the fact that  $SO(3) \simeq SU(2)/\mathbb{Z}_2$ . However, it implies that the minimum charge allowed using  $SO(3)$  as a gauge field instead of  $SU(2)$  is  $e_{min} = \frac{e}{2}$ . It then follows that the 't Hooft-Polyakov monopole is characterised by the same quantisation condition as the Dirac one:

$$g = -\frac{4\pi}{e} N_\Sigma = -\frac{2\pi}{e_{min}} N_\Sigma \quad (3.32)$$

### 3.5 Relation with the Dirac Monopole

In conclusion, the Dirac Monopole and the 't Hooft-Polyakov Monopole share the same quantization condition:

$$g = -\frac{2\pi}{e} N \quad N \in \mathbb{Z} \quad (3.33)$$

However, there are many relevant differences between the two, which are summarized below:

- **Postulate vs Necessity**

The Dirac Monopole was introduced in the framework given by electromagnetism and its existence was assumed by Dirac as a "postulate". Claiming that nature should respect duality symmetry, he showed that assuming the monopole existence, we could derive the well observed and otherwise theoretically unexplained, quantization of the electric charge. The 't Hooft-Polyakov Monopole, instead, is intrinsic to the Georgi-Glashow model and can be deduced for any non-Abelian gauge theory with a compact covering group. It arises from a spontaneous symmetry breaking and fulfills the same quantization condition of the Dirac Monopole.

- **Singularities vs Regularities**

The existence of the Dirac Monopole causes the existence of an origin-located singularity for the field-strength of electromagnetism  $F_{\mu\nu}$ . The 't Hooft-Polyakov ansatz for the solutions of the Georgi-Glashow model does not require the existence of any singularity for the field strength  $\vec{G}_{\mu\nu}$ .

- **Different derivations for the quantization condition**

The quantisation of the Dirac Monopole, is obtained defining two different potentials up to a gauge transformation, and imposing the particle wavefunction to be single-valued. The quantisation of the 't Hooft-Polyakov monopole charge is due to a topological property: the winding number of the map  $\vec{\phi}_\infty$ , which goes from a compact closed surface containing the monopoles to a 2-dimensional sphere. For such a reason it is often addressed as a *topological charge*.

### 3.6 Generalisation of the Ansatz to Dyons

The 't Hooft-Polyakov Ansatz can be extended to dyons relaxing the hypotheses imposed at the beginning of section 3.3, as it was done in the 1975 work of B.Julia and A.Zee. [6].

In fact we may take the Lagrangian of the Georgi-Glashow model, which can be found in expression 3.3, and find solutions satisfying the following hypotheses:

1. Finite Energy;
2. Stability;
3. Spherical Symmetry;
4. Time Independence.

It should be noted that we have only changed condition 4, relaxing our request on the static nature of the solutions: we want them not to depend on time, but  $\vec{W}^{\mu=0}$  can be different from zero. Hence we may think that, given three radial real functions  $J(\xi)$ ,  $H(\xi)$  and  $K(\xi)$ , the ansatz for the solutions to the Euler-Lagrange equations of the system has such form:

$$\vec{\phi}(\vec{r}) = \frac{\vec{r}}{er^2} H(\xi) \quad (3.34)$$

$$W_a^{\mu=0} = \frac{\vec{r}}{er^2} J(\xi) \quad (3.35)$$

$$W_a^{\mu=i} = -\epsilon_{aij} \frac{r^j}{er^2} (1 - K(\xi)) \quad (3.36)$$

Plugging them into the Euler-Lagrange equations we obtain a system of three differential equations describing the dynamics of the functions  $J(\xi)$ ,  $H(\xi)$  and  $K(\xi)$ :

$$\xi^2 \frac{d^2 J}{d\xi^2} = 2JK^2 \quad (3.37a)$$

$$\xi^2 \frac{d^2 H}{d\xi^2} = 2K^2 H + \frac{\lambda}{e^2} H(H^2 - \xi^2), \quad (3.37b)$$

$$\xi^2 \frac{d^2 K}{d\xi^2} = K(K^2 - J^2 + H^2 - 1) \quad (3.37c)$$

It should be noticed that the equations found by 't Hooft and Polyakov, 3.16a and 3.16b, can be obtained from those above computing the limit  $J \rightarrow 0$ . The boundary conditions on  $H(\xi)$  and  $K(\xi)$  are analogous to expressions 3.18a and 3.18b, whereas the boundary conditions on  $J(\xi)$  are the following ones:

$$\xi \rightarrow \infty: J \rightarrow 0 \quad (3.38a)$$

$$\xi \rightarrow 0: J \sim 1 + O(\xi) \quad (3.38b)$$

With a procedure similar to the one used in section 3.3, we can calculate the electromagnetic tensor:

$$F_{\mu\nu} = \frac{1}{a} \vec{\phi} \cdot \vec{G}_{\mu\nu}, \quad (3.39)$$

and notice that at spatial infinity the magnetic field behaves as that produced by the 't Hooft-Polyakov Monopole:

$$F_{ij} = \epsilon_{ija} \frac{r_a}{r} \left( -\frac{1}{er^2} \right). \quad (3.40)$$

However, what is new from the 't Hooft-Polyakov Monopoles, is that a non-zero electric field appears:

$$E^i = F^{0i} = -\frac{r^i}{r} \frac{d}{dr} \left[ \frac{J(r)}{r} \right], \quad (3.41)$$



and gives a total charge, which can be written as:

$$\begin{aligned}
Q &= \int dS_i F_{0i} = \\
&= \int d^3x \partial_i F_{0i} \\
&= -\frac{8\pi}{e} \int_0^\infty dr \frac{IK^2}{r}.
\end{aligned} \tag{3.42}$$

In conclusion, we get an object, which has a quantised magnetic charge but has an electric charge with no quantisation condition.

### 3.7 The Monopole Mass

We wish to estimate the mass (or, equivalently, energy) of the 't Hooft-Polyakov monopole: in general, the mass of a dyon is given by the integral

$$E = \int_{\mathbb{R}^3} \frac{1}{2} \left( E_i^a E_i^a + B_i^a B_i^a + D_\mu \phi^a D_\mu \phi^a \right) + V(\phi). \tag{3.43}$$

where  $G_{0i}^a = E_i^a$  and  $G_{ij}^a = B_i^a$ .

Now, we will derive a lower bound for this mass. Then, we will explicitly show that this bound can be reached by a specific solution: the BPS monopole.

#### 3.7.1 The Bogomol'nyi Bound

First of all we drop the (always positive) terms  $D_0 \phi^a D_0 \phi^a$  and  $V(\phi)$ . Then, we apply the famous “squaring trick”: we add and subtract some terms depending on a generic angle  $\theta$ :

$$\begin{aligned}
E \geq \frac{1}{2} \int_{\mathbb{R}^3} & E_i^a E_i^a + B_i^a B_i^a + D_i \phi^a D_i \phi^a + 2E_i^a D_i \phi^a \sin \theta - 2E_i^a D_i \phi^a \sin \theta \\
& + 2B_i^a D_i \phi^a \cos \theta - 2B_i^a D_i \phi^a \cos \theta,
\end{aligned} \tag{3.44}$$

and then we make them into squares, which we can discard while still maintaining the bound:

$$\begin{aligned}
E \geq \frac{1}{2} \int_{\mathbb{R}^3} & \left( E_i^a E_i^a - 2E_i^a D_i \phi^a \sin \theta + D_i \phi^a D_i \phi^a \sin^2 \theta \right) + \\
& + \left( B_i^a B_i^a - 2B_i^a D_i \phi^a \cos \theta + D_i \phi^a D_i \phi^a \cos^2 \theta \right) + \\
& + 2E_i^a D_i \phi^a \sin \theta + 2B_i^a D_i \phi^a \cos \theta.
\end{aligned} \tag{3.45}$$

The terms in parentheses are positive, therefore this is just bounded by:

$$E \geq \int_{\mathbb{R}^3} E_i^a D_i \phi^a \sin \theta + B_i^a D_i \phi^a \cos \theta \tag{3.46}$$

for *any*  $\theta$ . Now, we use the following facts: by the Bianchi identities  $D_\mu {}^* G^{\mu\nu,a} = 0$  we have  $D_i B_i^a = 0$ , therefore  $B_i^a D_i \phi^a D_i (B_i^a \phi^a)$ . Also, the equations of motion read  $D_\nu G^{\mu\nu,a} = e \varepsilon_{abc} \phi^c D^\mu \phi^b$ . Also,  $D_i G_{0i,a} = D_i E_i^a$ . Therefore:

$$D_i \phi^a E_i^a = D_i (\phi^a E_i^a) - \phi^a D_i E_i^a = D_i (\phi^a E_i^a) - \phi^a e \varepsilon_{abc} \phi^c D^\mu \phi^b = D_i (\phi^a E_i^a). \quad (3.47)$$

With these two facts we can frame all of the integrand as a divergence:

$$E \geq \int_{\mathbb{R}^3} D_i (E_i^a \phi^a \sin \theta + B_i^a \phi^a \cos \theta) d^3x \quad (3.48)$$

and apply Stokes' theorem:

$$E \geq \int_{\Sigma_\infty} \phi^a (E_i^a \sin \theta + B_i^a \cos \theta) dS_i. \quad (3.49)$$

Now, by the boundary conditions on  $\phi$  it will be radially outward and with absolute value  $a$  at infinity: therefore its product with the electric and magnetic fields (which are also radial at infinity) reduces to

$$E \geq a(q \sin \theta + g \cos \theta), \quad (3.50)$$

where  $q$  and  $g$  are the electric and magnetic charges. The sharpest bound occurs when the  $(q, g)$  and  $(\sin \theta, \cos \theta)$  2D vectors are aligned, and corresponds to their euclidean scalar product: the final bound we get is thus found by the saturation of the Cauchy-Schwarz inequality:  $E \geq a \sqrt{q^2 + g^2}$ .

In the 't Hooft-Polyakov case, this simplifies to  $E \geq a|g|$ .

### 3.7.2 Saturation of the Bogomol'nyi Bound

A state which attains the lower bound is called a Bogomol'nyi-Prasad-Sommerfield — for short BPS — monopole.

To derive the bound we discarded in the expression of the energy integrals of always-positive quantities, so to reach the bound they must be assumed to vanish everywhere. In addition, we assume our solution to be static. This yields the conditions:

$$V(\phi) = 0 \quad (3.51a)$$

$$D_0 \phi^a = 0 \quad (3.51b)$$

$$E_i^a = 0 \quad (3.51c)$$

$$B_i^a = \pm D_i \phi^a. \quad (3.51d)$$

Now, we make an interesting observation: if  $\lambda \neq 0$  in condition (3.51a) we must have that  $\phi^a \phi^a = a^2$  everywhere: therefore  $D_i (\phi^a \phi^a) = 2\phi^a D_i \phi^a = 0$ , and an analogous equation holds with  $D_i \rightarrow \partial_i$ . However, we also have equation (3.51d): using it we can also derive  $\phi^a D_i \phi^a \propto \phi^a B_i^a = 0$ .

If we insert the 't Hooft-Polyakov *ansatz* in the Bogomol'nyi condition (3.51d) we get the following set of differential equations:

$$\xi \frac{dK}{d\xi} = -KH \quad (3.52a)$$

$$\xi \frac{dH}{d\xi} = -K^2 + H + 1; \quad (3.52b)$$

these can be solved analytically, when coupled to the regular monopole equations of motion.

Add computations to show that the energy density of the monopole is finite at the origin.

## A Differential forms and Poincaré's lemma

In this appendix we present a brief overview of differential forms, following [10].

Differential forms are a useful tool when working with completely antisymmetric tensors: in a space of dimension  $d$  one can define tensors of any rank  $n$ , but if we restrict ourselves to the antisymmetric ones, which satisfy:

$$F_{\mu_1\mu_2\ldots\mu_n} = F_{[\mu_1\mu_2\ldots\mu_n]} \quad (A.1)$$

we can see that they only have  $\binom{d}{n}$  independent components: they can be nontrivial only if  $n \leq d$ .

The differential form associated with the tensor  $F_{\mu_1\ldots\mu_n}$  is defined as:

$$F = \frac{1}{n!} F_{\mu_1\ldots\mu_n} dx^{\mu_1} \wedge \cdots \wedge dx^{\mu_n}, \quad (A.2)$$

where  $dx^{\mu_1} \wedge \cdots \wedge dx^{\mu_n}$  is the canonical basis of the exterior algebra  $\wedge^n(V)$ , where  $V$  is the vector space over which the tensors are constructed. This basis is antisymmetric:  $dx^{\mu_1} \wedge \cdots \wedge dx^{\mu_n} = dx^{[\mu_1} \wedge \cdots \wedge dx^{\mu_n]}$ .

The wedge product is an antisymmetrized tensor product: it maps a  $p$ -form  $A$  and a  $q$ -form  $B$  to a  $(p+q)$ -form  $A \wedge B$ , defined by:

$$A \wedge B = \frac{1}{p!q!} A_{[\mu_1\ldots\mu_p} B_{\nu_1\ldots\nu_q]} dx^{\mu_1} \wedge \cdots \wedge dx^{\mu_p} \wedge dx^{\nu_1} \wedge \cdots \wedge dx^{\nu_q}. \quad (A.3)$$

It is *graded* antisymmetric:  $A \wedge B = (-1)^{pq} B \wedge A$ .

The number of independent components tells us that  $n$ -forms and  $d-n$ -forms have the same number of independent components: it is natural to find a connection between them. This is realized by *Hodge duality*: an  $n$ -form  $F$  is mapped to the  $(d-n)$ -form  $*F$ , defined by:

$$*F = \frac{1}{n!(d-n)!} \varepsilon_{\mu_1\ldots\mu_d} F^{\mu_1\ldots\mu_n} dx^{\mu_{n+1}} \wedge \cdots \wedge dx^{\mu_d}. \quad (A.4)$$

The exterior differential maps an  $n$ -form  $F$  to an  $(n + 1)$ -form  $dF$  as such: consider the derivative operator  $\partial \equiv dx^\mu \partial_\mu$ . Then,  $dF \equiv \partial \wedge F$ .

The functions usually considered are sufficiently regular, therefore by Schwarz's theorem  $d^2 = 0$ .

An  $n$ -form  $F$  can be called *closed* if  $dF = 0$  and *exact* if there exists an  $(n - 1)$ -form  $A$  such that  $F = dA$ . By  $d^2 = 0$  we have that exact implies closed, while the inverse is not generally true.

A *contractible* subset of a topological space is one which can be continuously deformed to a point. *Poincaré's lemma* states that, if a form is defined on a contractible set and it is closed, then it is also exact.

The non-contractibility of spaces is caused by topological nontriviality: the quotient of the closed  $n$ -forms modulo the exact  $n$ -forms characterizes this and is called the  $n$ -th de Rham cohomology group.

## B Homotopy

If we have a topological space  $X$ , the set of all maps  $f$  from  $[0, 1]^n$  such that  $f(\vec{0}) = f(\vec{1}) = x_0 \in X$  considered modulo homotopy and with the operation of composition is called the  $n$ -th homotopy group  $\Pi_n(X; x_0)$ . The dependence on  $x_0$  is actually trivial, it only matters to which connected component of  $X$  the element  $x_0$  belongs.

Composition is to be considered as such:

$$(f + g)(\vec{x}) = \begin{cases} f(2x_1, x_2, \dots, x_n) & x_1 \in [0, 1/2] \\ g(2x_1 - 1, x_2, \dots, x_n) & x_1 \in [1/2, 1]. \end{cases} \quad (\text{B.1})$$

Some notable cases are:  $\Pi_n(S^m) = 0$  for  $n < m$  and  $\Pi_n(S^n) = (\mathbb{Z}, +)$ .

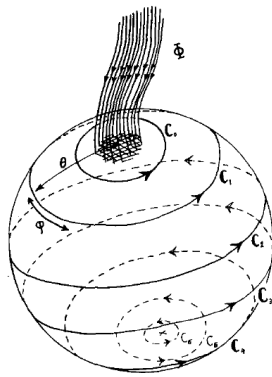


Figure 2: 't Hooft schematization of the spherical surface and the spot

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