

# Gravitational Wave Exercises @ Jena

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Exercises for the Gravitational Waves course. The exercise sheets can be found on the [webpage](#).

## 1 Quadrupole approximation

The procedure to find the quadrupole approximation for the GW emitted by a Newtonian system is in the form:

1. determine the density  $\rho(\vec{x}, t)$ ;
2. calculate the trace-free inertia tensor  $Q^{ij}(t)$ ;
3. calculate the gravitational wave strain  $h_{ij}$ .

### Two point particles

In order to model the point-like nature of the particles we can write an expression for the density as a sum of two delta-functions, whose locations rotate around an axis — let us fix it to be the  $z$  axis for simplicity, and also let us suppose that we are on the  $z = 0$  plane:

$$\rho(\vec{x}, t) = m\delta(\vec{x} - \vec{x}_1(t)) + m\delta(\vec{x} - \vec{x}_2(t)), \quad (1.1)$$

where

$$\vec{x}_1(t) = r \begin{bmatrix} \cos(\omega t + \phi) \\ -\sin(\omega t + \phi) \\ 0 \end{bmatrix} \quad \text{and} \quad \vec{x}_2(t) = -\vec{x}_1(t). \quad (1.2)$$

For simplicity we will also assume that the radius  $r$  is constant — this cannot be precisely the case, since there will be some amount of power lost because of the GW emission, but it is a reasonable approximation.

So, the trace-free inertia tensor will be given by

$$Q^{ij}(t) = \int \rho(\vec{x}, t) \left( x^i x^j - \frac{1}{3} \delta^{ij} r^2 \right) d^3x \quad (1.3)$$

$$= m \sum_{k=1,2} \left( x_k^i x_k^j - \frac{1}{3} \delta^{ij} r^2 \right) \quad (1.4)$$

$$= 2mr^2 \begin{bmatrix} \cos^2 - 1/3 & -\cos \sin \\ -\cos \sin & \sin^2 - 1/3 \end{bmatrix}, \quad (1.5)$$

where we write only the upper-left 2x2 submatrix in  $Q^{ij}$ , since the other entries are constant, and we omit the argument of the sines and cosines (which is always  $\omega t + \phi$ ).

The second derivative of this tensor will be given by

$$\ddot{Q}^{ij}(t) = 4mr^2\omega^2 \begin{bmatrix} \sin^2 - \cos^2 & 2\sin \cos \\ 2\sin \cos & \cos^2 - \sin^2 \end{bmatrix}, \quad (1.6)$$

Now, in order to compute the gravitational wave strain we need the projection tensor  $\Lambda_{ij,kl}$ . If the propagation direction we are interested in is  $\vec{k}$ , then the tensor

$$P_{ij} = \delta_{ij} - \frac{k_i k_j}{|\vec{k}|^2} \quad (1.7)$$

will project a vector onto the subspace orthogonal to  $\vec{k}$ ; and the tensor

$$\Lambda_{ij,kl} = P_{ik}P_{jl} - \frac{1}{2}P_{ij}P_{kl} \quad (1.8)$$

$$= \delta_{ik}\delta_{jl} - n_i n_k \delta_{jl} - \delta_{ik} n_j n_l + \frac{1}{2}n_i n_k n_j n_l - \frac{1}{2}\delta_{ij}\delta_{kl} + \frac{1}{2}\delta_{ij}n_k n_l + \frac{1}{2}n_i n_j \delta_{kl}, \quad (1.9)$$

will project a rank-2 tensor onto the corresponding subspace.