# ECE 6390: Homework 2

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September 7, 2025

## 1 Problem Setup.

From the previous homework, for continuous time,

$$\ddot{r} = r \, \dot{\theta}^2 - \frac{\mu}{r^2}, \qquad \ddot{\theta} = -\frac{2 \, \dot{r} \, \dot{\theta}}{r}.$$

Here,  $\mu$  is the standard gravitational parameter, defined as the product of the universal gravitational constant (G) and the mass of the primary body  $(M_p)$ , such that  $\mu = GM_p$ .

$$\frac{\Delta r_n}{\Delta t} = \dot{r}_n, \qquad \frac{\Delta \theta_n}{\Delta t} = \dot{\theta}_n, \qquad r_{\text{mid},n} = r_n + \frac{1}{2} \Delta r_n.$$

Then the step  $n \rightarrow n+1$  is

$$\begin{split} r_{n+1} &= r_n + \Delta r_n, \\ \theta_{n+1} &= \theta_n + \Delta \theta_n, \\ \Delta r_{n+1} &= \Delta r_n + \left( r_{\text{mid},n} \, \Delta \theta_n^2 - \frac{\mu}{r_n^2} \, \Delta t^2 \right), \\ \Delta \theta_{n+1} &= \Delta \theta_n - \frac{2 \, \Delta r_n \, \Delta \theta_n}{r_{\text{mid},n}} \, . \end{split}$$

#### Solar Pressure Model

**Assumptions.** (1) Effective reflectivity  $\alpha_r = 0.5$ . (2) Solar pressure acts in the equatorial plane. (3) Sun direction rotates uniformly:  $\alpha(t) = \omega_{\rm yr} t + \phi_0$ , with  $\omega_{\rm yr} = \frac{2\pi}{365.25636 \times 86400}$ .

Area density and magnitude.

$$\begin{split} \rho_A &= \frac{m_s}{A_s}, \qquad F_{\rm solar} = 9.08~\mu\text{N/m}^2 \times \alpha_r A_s, \\ a_{\rm solar} &= \frac{F_{\rm solar}}{m_s} = \frac{9.08 \times 10^{-6}~\alpha_r}{\rho_A}~[\text{m/s}^2] = \frac{9.08 \times 10^{-9}~\alpha_r}{\rho_A}~[\text{km/s}^2]. \end{split}$$

Components in polar  $(r, \theta)$ . Let  $\psi(t) = \alpha(t) - \theta(t)$ . Then radial component and tangential component is,

$$a_r = a_{\text{solar}} \cos \psi, \qquad a_\theta = a_{\text{solar}} \sin \psi.$$

Modified update equations. The solar pressure adds acceleration components  $a_r$  and  $a_\theta/r$  to the continuous-time equations,

$$\ddot{r} = r \,\dot{\theta}^2 - \frac{\mu}{r^2} + a_r, \qquad \ddot{\theta} = -\frac{2 \,\dot{r} \,\dot{\theta}}{r} + \frac{a_\theta}{r}.$$

The corresponding discrete update for the step  $n \rightarrow n+1$  becomes

$$\Delta r_{n+1} = \Delta r_n + \left( r_{\text{mid},n} \, \Delta \theta_n^2 - \frac{\mu}{r_n^2} \, \Delta t^2 + a_{\text{solar}} \cos \psi_n \, \Delta t^2 \right),$$

$$\Delta \theta_{n+1} = \Delta \theta_n - \frac{2 \, \Delta r_n \, \Delta \theta_n}{r_{\text{mid},n}} + \frac{a_{\text{solar}} \sin \psi_n}{\mathbf{r}_{\text{mid},n}} \, \Delta t^2,$$

where  $t_n = n\Delta t$ ,  $\alpha_n = \phi_0 + \omega_{yr}t_n$ ,  $\psi_n = \alpha_n - \theta_n$ .

## 2 Results

### 2.1 Minimum area density vs. target drift

Table 1 compares the minimum area density  $\rho_A$  required to produce 1°, 5°, and 15° of azimuth drift over one sidereal year for two time steps. The 1° and 5° cases agree within about 1% and 4%, respectively, while the 15° case shows strong step—size sensitivity. We therefore report the  $dt=0.5\,\mathrm{s}$  values as our final results (without extrapolation). The signed percentage indicates  $(\rho_A^{1\mathrm{s}}-\rho_A^{0.5\mathrm{s}})/\rho_A^{0.5\mathrm{s}}\times 100\%$ .

Target drift [deg]	$ ho_A~( ext{dt}= ext{0.5 s})~ ext{[kg/m}^2]$	$ ho_A \; ( ext{dt} = 1 \;  ext{s}) \; [ ext{kg/m}^2]$	Change vs. 0.5 s [%]
1	14.544	14.681	+0.94
5	2.784	2.675	-3.92
15	0.792	0.231	-70.83

Table 1: Minimum area density  $\rho_A$  versus time step. We use the  $dt = 0.5 \,\mathrm{s}$  column as the reported values. Negative percentages indicate that the  $dt = 1 \,\mathrm{s}$  run produced a smaller  $\rho_A$  than the  $dt = 0.5 \,\mathrm{s}$  run.

#### 2.2 Minimizing 1-year azimuth deviation at $\rho_A = 0.5$

**Method.** With  $\rho_A = 0.5 \text{ kg/m}^2$  fixed, we chose a small tangential impulse  $\Delta v_t$  to drive the 1-year mean azimuth drift (slope) to  $\approx 0 \text{ deg/day}$  using a 4-term regression fit [1, t (day),  $\sin(2\pi t/1y)$ ,  $\cos(2\pi t/1y)$ ]. Keeping  $\Delta v_t$  fixed, we then selected the initial longitude offset  $\Delta \lambda_0$  to minimize the annual (1/year) harmonic amplitude. Dynamics were integrated in planar ECI using RK4 with a rotating solar-pressure acceleration  $a_{\text{solar}} = 9.08 \times 10^{-9} \alpha_r / \rho_A \text{ km/s}^2$ . Look angles were computed from local ENU(East North Up) via Az =  $\tan 2(E, N)$  and El =  $\tan 2(U, \sqrt{E^2 + N^2})$ .

$\Delta v_t  [\mathrm{m/s}]$	$\Delta \lambda_0  [\mathrm{deg}]$	Slope [deg/day]	Annual amp [deg]	Max  Az-lin  [deg]	RMS [deg]
+0.010	-0.691	0.001350	0.030	9.156	4.579

Table 2: Summary of the minimized-azimuth design at  $\rho_A = 0.5 \text{ kg/m}^2$ . The slope corresponds to  $\approx 0.49^{\circ}$  per year (viable). The annual component is strongly suppressed. The remaining  $\pm$  few-degree swing is dominated by the daily geosynchronous east—west oscillation.

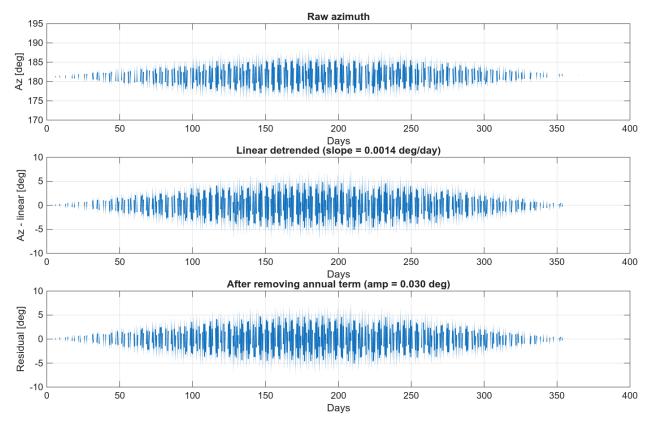


Figure 1: One-year azimuth as seen from Atlanta for the minimized-deviation design ( $\rho_A = 0.5 \text{ kg/m}^2$ ). Top: raw azimuth. Middle: azimuth after removing the linear trend (slope = 0.00135 deg/day). Bottom: residual after additionally removing the annual component (amplitude = 0.030°).

Takeaway. With  $\Delta v_t = +0.010$  m/s and  $\Delta \lambda_0 = -0.691^{\circ}$ , the 1-year mean drift is  $< 1^{\circ}$  and the annual wobble is  $\approx 0.03^{\circ}$ . The residual variation is primarily the daily geosynchronous oscillation.

## 3 Code.

#### Minimum area density

```
function minimum_area_density()
2
   % Finds the minimum area density rho_A that yields a target azimuth drift
   % (1 , 5 , 15 ) over one sidereal year as seen from Atlanta (Van Leer).
3
   %% Constants
5
   mu = 398600.4418; %
                         =GM [km^3/s^2]
   r_earth = 6378.137; % [km]
   w_earth = 2*pi/86164.0905; % Earth sidereal rotation [rad/s]
   sidereal_year = 365.25636*86400; % Sidereal year [s]
   w_yr = 2*pi/sidereal_year; % Sun-direction rotation [rad/s]
11
   alpha_r = 0.5; % Effective reflectivity (Given in problem)
   % Station (Van Leer, Atlanta) latitute, longitude [rad]
   station_lat = deg2rad(33.7758);
   station_lon = deg2rad(-84.39738);
16
   % GEO initial conditions (circular, equatorial)
17
   r0 = (mu / w_earth^2)^(1/3); % [km]
   vtheta0 = r0 * w_earth;
   theta0 = 0.0;
```

```
lon0 = station_lon;
22 \text{ vr0} = 0.0;
23
24 % Integrator
25 % Test for 1, 0.5
26 dt = 0.5; % [s]
27
28 %% Solve for rho_A at three drift targets
29 targets = [1, 5, 15]; % degrees
rho_out = zeros(size(targets));
for k = 1:numel(targets)
       rho_out(k) = rho_for_target_deg(targets(k));
32
зз end
g5 fprintf('\nMinimum area density over 1 sidereal year\n');
36 for k = 1:numel(targets)
       fprintf('Target drift %2 d : rho_A = %.3f kg/m^2\n', targets(k), rho_out(k));
37
зв end
39
   %%
40
41
       function rho = rho_for_target_deg(target_deg)
           % Find rho_A such that | Az | target_deg after one sidereal year.
42
           % Drift decreases as rho_A increases.
43
44
           % Initial bracket [rho_lo, rho_hi] where drift_lo > target > drift_hi
45
46
           rho 10 = 0.5:
47
           rho_hi = 200; % broad bracket [kg/m^2]
48
49
           drift_lo = drift_for_rho(rho_lo);
50
51
           drift_hi = drift_for_rho(rho_hi);
52
           it = 0;
           while ~(drift_lo>target_deg && drift_hi<target_deg) && it < 50</pre>
54
               if drift_lo <= target_deg, rho_lo = rho_lo/2; drift_lo = drift_for_rho(rho_lo); end</pre>
55
               if drift_hi >= target_deg, rho_hi = rho_hi*2; drift_hi = drift_for_rho(rho_hi); end
56
               it = it + 1;
57
           end
           % Bisection
59
60
           for i=1:50
               mid = 0.5*(rho_lo+rho_hi);
61
               d = drift_for_rho(mid);
62
63
               if d > target_deg
                   rho_lo = mid;
64
65
                   rho_hi = mid;
66
67
68
           end
           rho = 0.5*(rho_lo+rho_hi);
69
70
       end
72
       function drift_deg = drift_for_rho(rho_A)
           [r_hist, theta_hist, tt] = propagate(r0, theta0, vr0, vtheta0, ...
               mu,alpha_r,rho_A,w_yr,dt,sidereal_year);
74
75
           lsh = wrapToPi( (theta_hist - w_earth*tt) + lon0 ); % ECI->ECEF subpoint longitude
           azimuth_0 = az_from_station(station_lat, station_lon,0,lsh(1),r_earth);
76
           azimuth_f = az_from_station(station_lat, station_lon,0,lsh(end),r_earth);
78
           drift_deg = abs(rad2deg(angdiff(azimuth_0,azimuth_f)));
79
80 end
81
82 %% Propagator with solar-pressure
sa function [r_hist,theta_hist,t_hist] = propagate(r0,theta0,vr0,vtheta0,...
       mu,alpha_r,rho_A,w_yr,dt,simTime)
84
85
so n = ceil(simTime/dt);
87  r_hist = zeros(n+1,1);
88 t_hist = (0:n)'*dt;
```

```
r_{hist}(1) = r0;
90 theta_hist = zeros(n+1,1);
91 theta_hist(1) = theta0;
93 % Delta variables store
                               over one step
94 delta_r = vr0*dt;
95 delta_theta = (vtheta0/r0)*dt;
97 % Solar acceleration magnitude [km/s^2]
98 a_solar = (9.08e-6*alpha_r/rho_A)/1000; % m/s^2 to km/s^2
100 r = r0; theta = theta0;
101 for k = 1:n
        t = (k-1)*dt;
103
        % Position update
104
        r_next = r + delta_r;
105
        theta_next = theta + delta_theta;
106
107
        % Mid-step radius
108
109
        r_mid = r + 0.5*delta_r;
        theta_mid = theta + 0.5*delta_theta;
110
112
        % Sun direction and components
        alpha = w_yr*(t + 0.5*dt); % Sun angle (equatorial)
113
114
        psi = alpha - theta_mid; % Relative angle
        a_r = a_solar*cos(psi);
115
        a_theta = a_solar*sin(psi);
116
117
        % Modified updates
118
        dr_next = delta_r + (r_mid*(delta_theta^2) - (mu/(r_mid^2))*dt^2 + a_r*dt^2);
119
        dth_next = delta_theta - (2*delta_r*delta_theta)/r_mid + (a_theta/r_mid)*dt^2;
120
121
        % Roll
122
        r = r_next;
123
124
        theta = theta_next;
        delta_r = dr_next;
125
126
        delta_theta = dth_next;
127
128
        r_hist(k+1) = r;
        theta_hist(k+1) = theta;
129
130 end
131 end
132
   %% Azimuth from station (station_lat, station_lon) to sub-satellite point (target_lat, target_lon)
133
   function azimuth = az_from_station(station_lat, station_lon, target_lat, target_lon, earth_radius)
134
135
136
        % Spherical law of cosines for central angle between station and target
        cos_gamma = sin(target_lat)*sin(station_lat) + ...
137
                    cos(target_lat)*cos(station_lat)*cos(target_lon - station_lon);
138
        central_angle = acos( min(1, max(-1, cos_gamma)) );
139
140
        % Magnitude of the azimuth offset (an acute angle in [0, /2])
141
        sin_azimuth_abs = abs( sin(abs(station_lon - target_lon)) * cos(target_lat) / ...
142
                                max(1e-12, sin(central_angle)) );
143
        azimuth_offset = asin(min(1, max(0, sin_azimuth_abs))); % in [0, /2]
144
145
146
        % Quadrant resolution (for GEO with target_lat = 0 and station_lat > 0, target is to the south)
        if target_lon > station_lon % South & East
147
            azimuth = pi - azimuth_offset;
148
149
        else % South & West
            azimuth = pi + azimuth_offset;
150
        end
151
152
153
        azimuth = mod(azimuth, 2*pi);
154 end
155
156 %%
```

## Minimum azimuth look angles

```
function result = minimum_azimuth_lookangles()
2 % Choose v_t so that the 1-year mean azimuth drift (slope)
3 % Choose 0 to minimize the 1-year (solar-pressure) harmonic amplitude.
4 % Run a high-fidelity simulation and plot 1-year look angles/metrics.
5 %
6 % Output:
_{7} % Console: v_t , 0 , slope, annual amplitude, detrended/annual-removed Max/RMS
8 % Figure: (1) Raw azimuth, (2) Linear detrended, (3) After removing annual term
9 % result struct: dv_t_mps, dlambda0_deg, metrics
10
11 %% Constants
mu = 398600.4418: \% = GM [km^3/s^2]
13 r_earth = 6378.137; % Earth radius [km]
u_earth = 2*pi/86164.0905; % Earth sidereal rotation [rad/s]
sidereal_year = 365.25636*86400; % Sidereal year [s]
sidereal_days = 365.25636; % Sidereal year [day]
w_yr = 2*pi/sidereal_year; % Sun-direction rotation [rad/s]
alpha_r = 0.5; % Effective reflectivity (given)
rho_A = 0.5; % Area density [kg/m^2] (fixed)
20 r0 = 42164.17; % GEO radius [km]
22 % Station (Van Leer, Atlanta)
                                    latitude, longitude [rad]
23  station_lat = deg2rad(33.7758);
24 station_lon = deg2rad(-84.39738);
26 %% Find v_t such that mean drift slope
27 dt_A = 120; % [s] 2 min (tighter slope estimation)
slope_fn = @(dv_t) slope_deg_per_day( ...
       dv_t, 0.0, dt_A, mu, r_earth, w_earth, sidereal_year, w_yr, ...
31
       alpha_r, rho_A, r0, station_lat, station_lon, sidereal_days);
33 % Grid to bracket a sign change for fzero
34 cand = linspace(-1, 1, 81); % v_t in [-1, +1] m/s
svals = arrayfun(slope_fn, cand);
s6 ix = find(sign(svals(1:end-1)).*sign(svals(2:end)) <= 0, 1, 'first');</pre>
37
38 if ~isempty(ix)
      bracket = [cand(ix), cand(ix+1)];
39
40
      cand2 = linspace(-5, 5, 81);
41
       svals2 = arrayfun(slope_fn, cand2);
42
       ix2 = find(sign(svals2(1:end-1)).*sign(svals2(2:end)) \le 0, 1, 'first');
43
      if ~isempty(ix2)
44
          bracket = [cand2(ix2), cand2(ix2+1)];
45
46
       else
47
          bracket = [-5, 5];
       end
48
49 end
51 dv_t_star = fzero(slope_fn, bracket); % [m/s]
53 %% Minimize 1-year amplitude by setting 0
54 dt_B = dt_A;
amp_fn = @(d_lambda0) annual_amp_deg( ...
```

```
dv_t_star, d_lambda0, dt_B, mu, r_earth, w_earth, sidereal_year, w_yr, ...
56
57
        alpha_r, rho_A, r0, station_lat, station_lon, sidereal_days);
58
59 % Search range 15 deg
60 d_lambda0_star = fminbnd(amp_fn, deg2rad(-15), deg2rad(15));
61
62 %% Simulation & plots
63 dt_F = 60; % [s]
65 [t_hist, az_deg_raw, az_deg_detr_lin, resid_noannual_deg, metrics] = ...
        sim_and_metrics( ...
66
            \label{eq:dv_t_star} dv_t_star, \ d_lambda0\_star, \ dt_F, \ mu, \ r\_earth, \ w\_earth, \ \dots
67
            sidereal_year, w_yr, alpha_r, rho_A, r0, station_lat, station_lon, sidereal_days);
68
fprintf('\nMin-azimuth design (rho_A = %.3f kg/m^2)\n', rho_A);
71 fprintf(' v_t (slope0) = %+7.3f m/s\n', dv_t_star);
72 fprintf(' 0 (min annual amp) = %+7.3f deg\n', rad2deg(d_lambda0_star));
fprintf('Mean drift slope = %10.6f deg/day\n', metrics.slope_deg_per_day);
74 fprintf('Annual amp (Az) = %7.3f deg (peak-to-peak
                                                             %7.3f deg)\n', ...
            metrics.annual_amp_deg, 2*metrics.annual_amp_deg);
75
    fprintf('Max | Az detrended| = %7.3f deg; RMS = %7.3f deg \ ...
            metrics.max_abs_detrended_deg, metrics.rms_detrended_deg);
78 fprintf('After removing annual: Max = %7.3f deg; RMS = %7.3f deg\n', ...
            metrics.max_abs_after_annual_deg, metrics.rms_after_annual_deg);
80
81 % Plots
82 days = t_hist/86400;
sa figure('Name','Az over sidereal year','Color','w');
84 subplot(3,1,1);
plot(days, az_deg_raw, 'LineWidth', 1); grid on;
s6 xlabel('Days'); ylabel('Az [deg]');
87 title('Raw azimuth');
89 subplot(3,1,2);
plot(days, az_deg_detr_lin, 'LineWidth', 1); grid on; stabel('Days'); ylabel('Az - linear [deg]');
p2 title(sprintf('Linear detrended (slope = %.4f deg/day)', metrics.slope_deg_per_day));
94 subplot(3.1.3):
plot(days, resid_noannual_deg, 'LineWidth', 1); grid on;
96 xlabel('Days'); ylabel('Residual [deg]');
pr title(sprintf('After removing annual term (amp = %.3f deg)', metrics.annual_amp_deg));
pp result = struct('dv_t_mps', dv_t_star, ...
                     'dlambda0_deg', rad2deg(d_lambda0_star), ...
100
                     'metrics', metrics);
101
102 end
103
104 %%
    function slope = slope_deg_per_day(dv_t_mps, d_lambda0, dt, ...
           mu, r_earth, w_earth, T, w_yr, alpha_r, rho_A, r0, station_lat, station_lon, sidereal_days)
106
107 % Estimate mean azimuth drift slope (deg/day) via 4-term regression:
_{1} _{08} % [1, t(day), sin(2 t /1y), cos(2 t /1y)] to separate the annual component.
109
110 [t_hist, az_deg] = forward_az( ...
        dv_t_mps, d_lambda0, dt, mu, r_earth, w_earth, T, w_yr, ...
        alpha_r, rho_A, r0, station_lat, station_lon);
112
113
114 tt_day = t_hist/86400;
1/15 y_deg = unwrap(deg2rad(az_deg))*180/pi; % unwrap in rad, then convert to deg
116 X = [ones(size(tt_day)), tt_day, ...
              sin(2*pi*tt_day/ sidereal_days), cos(2*pi*tt_day/ sidereal_days)];
118 b = X \setminus y_deg;
slope = b(2); % deg/day
120 end
121
function A_deg = annual_amp_deg(dv_t_mps, d_lambda0, dt, ...
mu, r_earth, w_earth, T, w_yr, alpha_r, rho_A, r0, station_lat, station_lon, sidereal_days)
```

```
124 % Return annual amplitude (deg) from the sin/cos coefficients of the 4-term fit.
125
126 [t_hist, az_deg] = forward_az( ...
        dv_t_mps, d_lambda0, dt, mu, r_earth, w_earth, T, w_yr, ...
        alpha_r, rho_A, r0, station_lat, station_lon);
128
129
130 tt_day = t_hist/86400;
131 y_deg = unwrap(deg2rad(az_deg))*180/pi;
X = [ones(size(tt_day)), tt_day, ...
              sin(2*pi*tt_day/ sidereal_days), cos(2*pi*tt_day/ sidereal_days)];
133
b = X \setminus y_{deg}
A_{deg} = hypot(b(3), b(4));
136 end
137
138 function [t_hist, az_deg, az_deg_detr_lin, resid_noannual_deg, m] = ...
        sim_and_metrics(dv_t_mps, d_lambda0, dt, ...
            \verb|mu|, \verb|r_earth|, \verb|w_earth|, \verb|T|, \verb|w_yr|, \verb|alpha_r|, \verb|rho_A|, \verb|r0|, \verb|station_lat|, \verb|station_lon|, \verb|sidereal_days|)
140
141 % Final simulation and metric computation.
142
143 [t_hist, az_deg] = forward_az( ...
        dv_t_mps, d_lambda0, dt, mu, r_earth, w_earth, T, w_yr, ...
145
        alpha_r, rho_A, r0, station_lat, station_lon);
146
147 tt_day = t_hist/86400;
1/48 y_deg = unwrap(deg2rad(az_deg))*180/pi;
149
150 % Remove linear trend
P = polyfit(tt_day, y_deg, 1);
152 trend = polyval(P, tt_day);
az_deg_detr_lin = y_deg - trend;
154
155 % Remove annual term (sin/cos at 1/year)
156 X = [sin(2*pi*tt_day/ sidereal_days), cos(2*pi*tt_day/ sidereal_days)];
ab = X\az_deg_detr_lin;
158 annual = X*ab;
resid = az_deg_detr_lin - annual;
160
161 m = struct();
162 m.slope_deg_per_day = P(1);
163 m.annual_amp_deg = hypot(ab(1), ab(2));
164 m.max_abs_detrended_deg = max(abs(az_deg_detr_lin));
165 m.rms_detrended_deg = sqrt(mean(az_deg_detr_lin.^2));
166 m.max_abs_after_annual_deg = max(abs(resid));
167 m.rms_after_annual_deg = sqrt(mean(resid.^2));
resid_noannual_deg = resid;
170 end
171
172 function [t_hist, az_deg] = forward_az( ...
            dv_t_mps, d_lambda0, dt, mu, r_earth, w_earth, T, w_yr, ...
173
            alpha_r, rho_A, r0, station_lat, station_lon)
174
175 % Earth-centered initial (ECI)
176 % Earth-centered - Earth-fixed (ECEF)
177 % Planar ECI -> ECEF -> East North Up -> azimuth time-series over T seconds.
178 % East North Up (ENU)
179
_{
m 180} % Propagate in Cartesian with RK4 (planar, solar pressure rotating with sun)
181 [s_hist, t_hist] = propagate_rk4( ...
182
        init_state(dv_t_mps, d_lambda0, r0, mu, station_lon), ...
183
        dt, T, mu, alpha_r, rho_A, w_yr);
184
185 % ECI -> ECEF
x = s_hist(:,1); y = s_hist(:,2);
theta_e = w_earth * t_hist;
188 c = cos(theta_e); s = sin(theta_e);
189 	ext{ xe} = c.*x + s.*y;
190 ye = -s.*x + c.*y;
ze = zeros(size(xe)); % planar z = 0
```

```
192
193 % Station ECEF
194 [xs, ys, zs] = station_ecef(r_earth, station_lat, station_lon);
196 % Line-of-sight in ECEF
197 dx = xe - xs; dy = ye - ys; dz = ze - zs;
198
199 % ECEF -> East North Up
200 [E, N, U] = ecef_to_enu(dx, dy, dz, station_lat, station_lon); %#ok<ASGLU>
201
202 % Look angles
203 azimuth = atan2(E, N);
                                      % [rad], 0..2
204 azimuth = mod(azimuth, 2*pi);
206 az_deg = rad2deg(azimuth);
207 end
208
gog function s0 = init_state(dv_t_mps, d_lambda0, r0, mu, station_lon)
210 % Initial planar ECI state from v_t and 0
211
212  v_circ = sqrt(mu/r0); % [km/s]
213 v_t = v_circ + dv_t_mps/1000; % [km/s]
theta0 = station_lon + d_lambda0; % start longitude offset
215
216 \times 0 = r0 \times cos(theta0);
_{217} y0 = r0*sin(theta0);
218  vx0 = -v_t*sin(theta0);
v_{19} = v_{t} \cos(theta0);
220
221 s0 = [x0; y0; vx0; vy0];
222 end
223
224 function [S, t_hist] = propagate_rk4(s0, dt, T, mu, alpha_r, rho_A, w_yr)
225 % 4th-order Runge Kutta propagation (planar 2D + rotating solar-pressure accel).
226
_{227} n = floor(T/dt);
228 t_hist = (0:n)'*dt;
S = zeros(n+1, 4);
231 S(1,:) = s0.';
232 a_solar = 9.08e-9 * alpha_r / rho_A; % [km/s^2] (from 9.08e-6 m/s^2)
233
_{234} for k = 1:n
       tk = t_hist(k);
235
       s = S(k,:).';
236
237
        k1 = f(tk,
238
                             s):
        k2 = f(tk + 0.5*dt, s + 0.5*dt*k1);
239
        k3 = f(tk + 0.5*dt, s + 0.5*dt*k2);
240
241
        k4 = f(tk + dt,
                             s + dt*k3);
242
        S(k+1,:) = (s + (dt/6)*(k1 + 2*k2 + 2*k3 + k4)).';
243
244 end
245
246
        function ds = f(ti, s)
           x=s(1); y=s(2); vx=s(3); vy=s(4);
247
248
            r = hypot(x,y);
249
            aG = -mu/r^3 * [x; y]; % gravity
250
            alpha = w_yr * ti; % sun direction angle
251
            aS = a_solar * [cos(alpha); sin(alpha)]; % solar-pressure accel
            ds = [vx; vy; aG(1)+aS(1); aG(2)+aS(2)];
252
253
        end
254 end
255
256 function [xs, ys, zs] = station_ecef(r_earth, station_lat, station_lon)
257 % Spherical Earth station to ECEF.
258 xs = r_earth*cos(station_lat)*cos(station_lon);
ys = r_earth*cos(station_lat)*sin(station_lon);
```

```
zs = r_earth*sin(station_lat);
end

261
262
263 function [E, N, U] = ecef_to_enu(dx, dy, dz, station_lat, station_lon)
264 % ECEF delta -> local East North Up.
265 sl = sin(station_lon);
266 cl = cos(station_lon);
267 sp = sin(station_lat);
268 cp = cos(station_lat);
268 cp = cos(station_lat);
269 E = -sl.*dx + cl.*dy;
270 N = -sp.*cl.*dx - sp.*sl.*dy + cp.*dz;
271 U = cp.*cl.*dx + cp.*sl.*dy + sp.*dz;
272 end
```