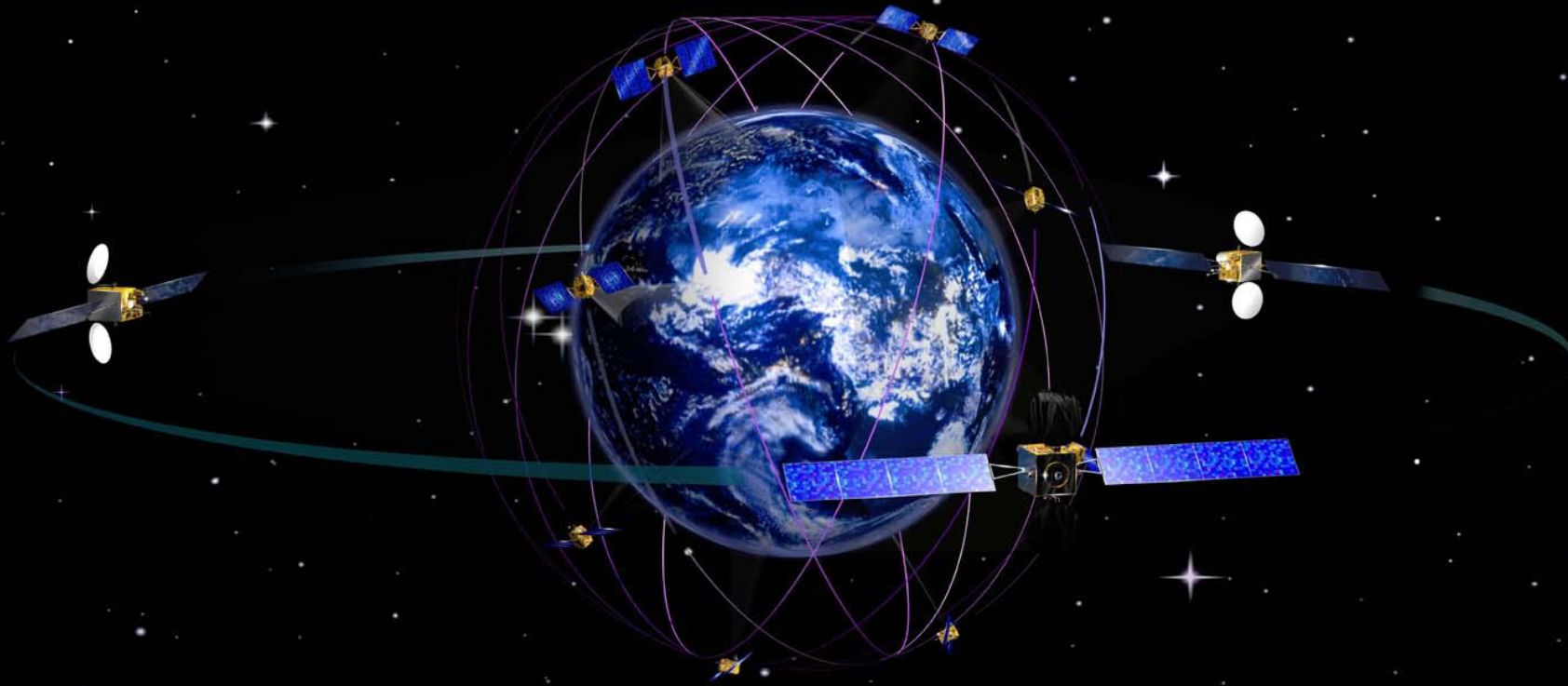
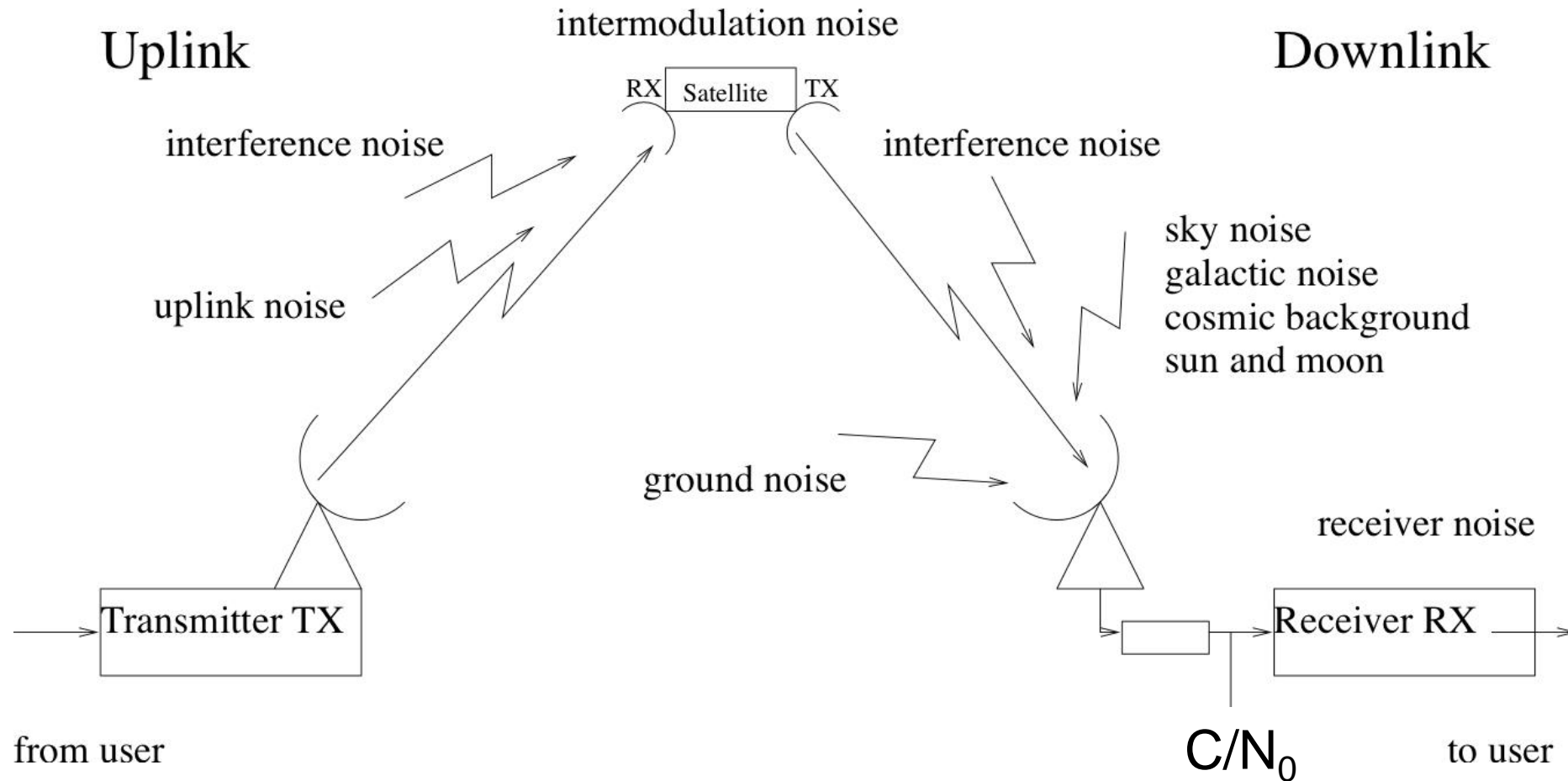


Satellite Communications

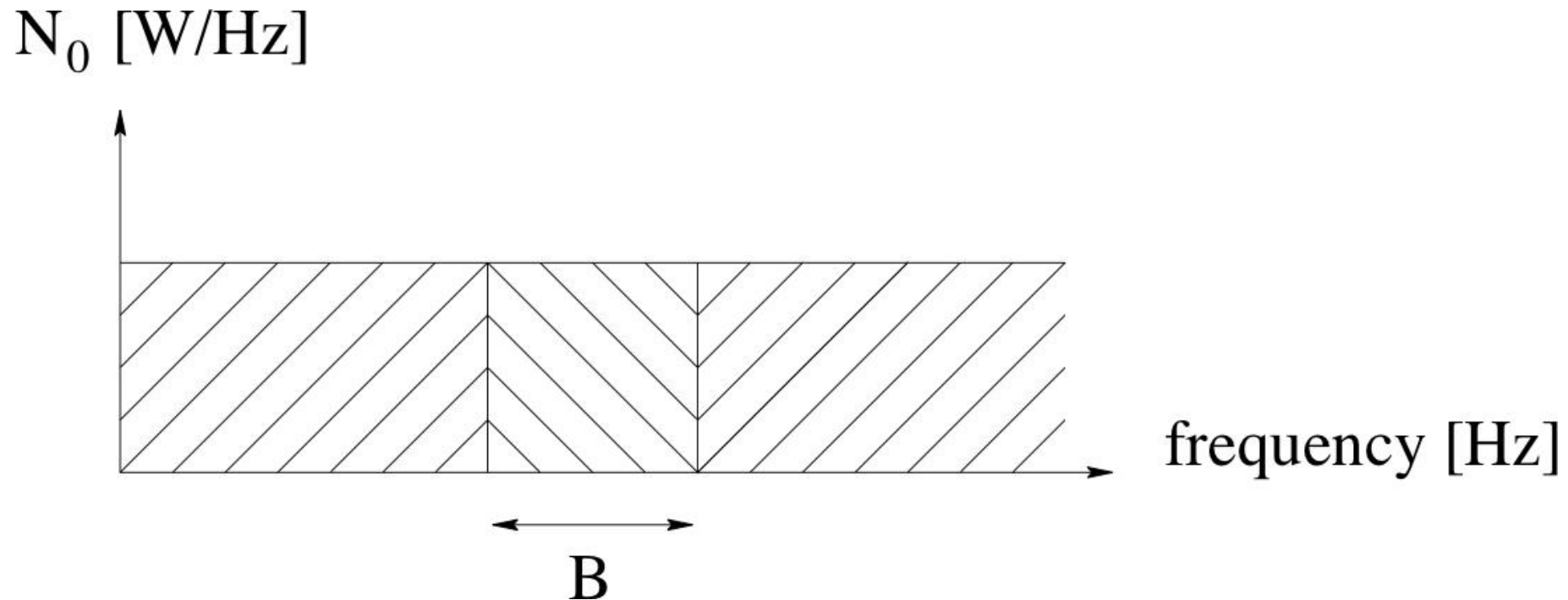
– RRY100 –



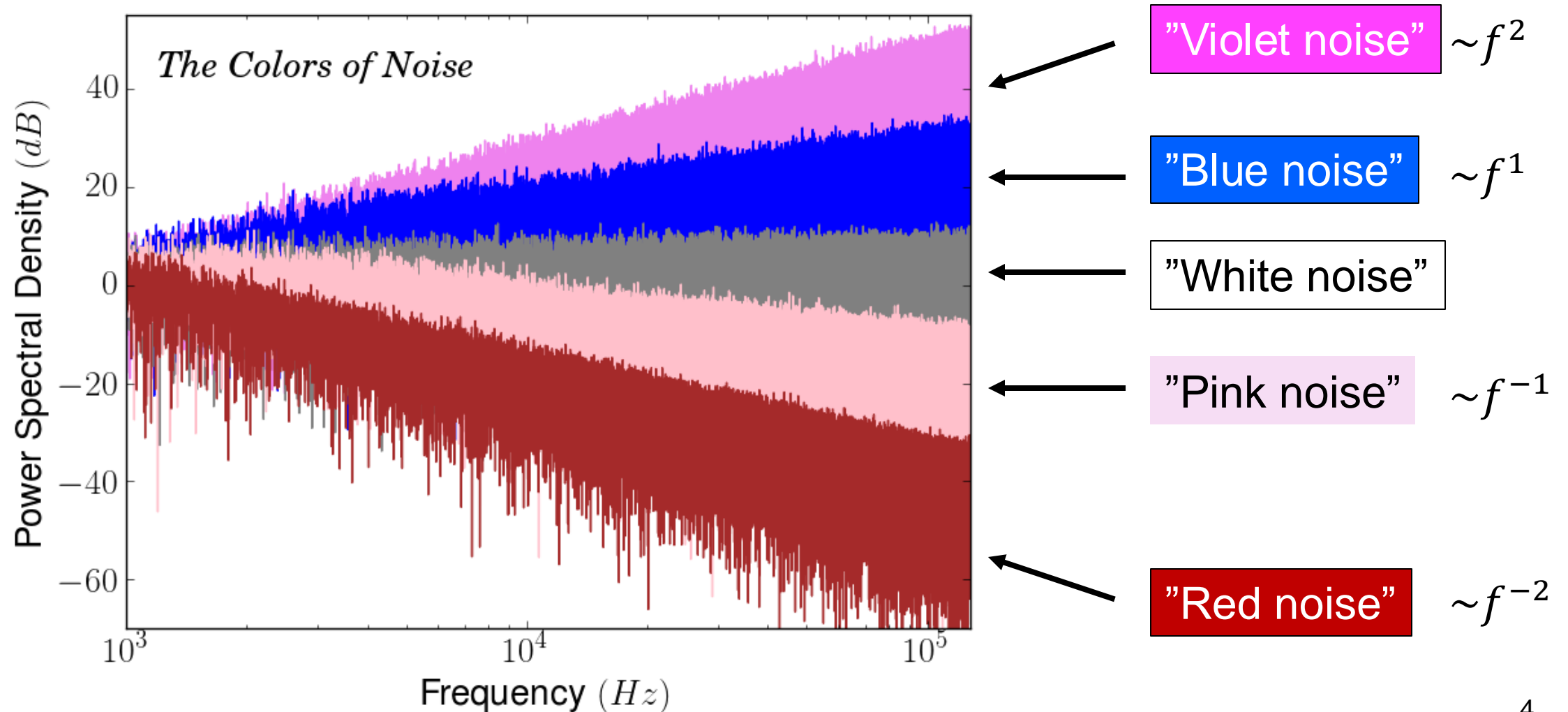
Noise in satellite communications



- Noise model:
 - We usually work with limited bandwidth only
 - An appropriate model is 'white noise'
 - Real noise over larger bandwidth usually is 'coloured', i.e. not constant over frequency

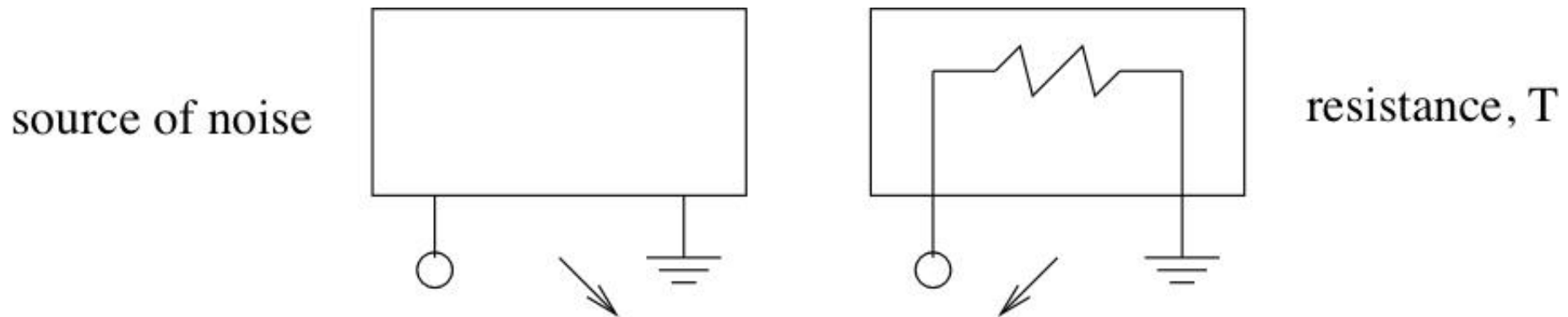


Different noise models



- Concept of 'Noise Temperature':

- All objects with a physical temperature T generate electromagnetic radiation
- We compare the noise power of the source of noise (its physical temperature may be different from T) with an object that has a physical temperature T
- If they have the same noise power N we identify the amount of noise by the 'noise temperature T' ' in Kelvin



Ludwig Boltzmann (1844-1906)

$$N_0 = k \cdot T \quad [\text{Ws}] \text{ or } [\text{W Hz}^{-1}]$$

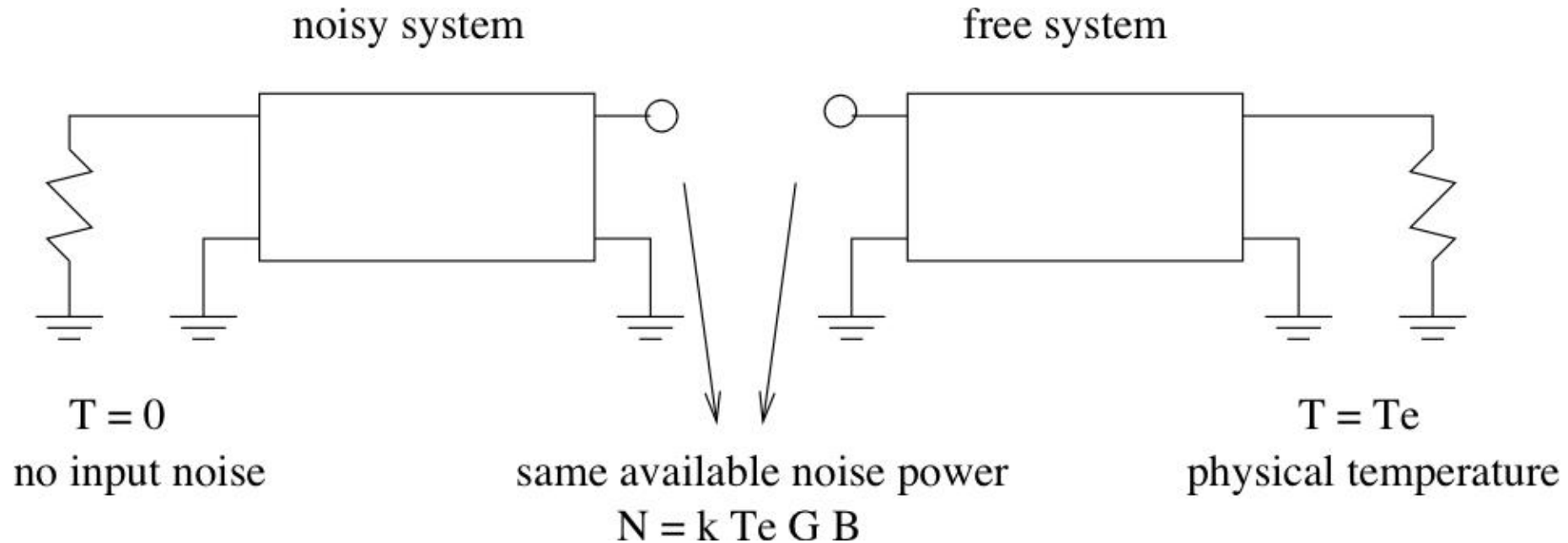
$$N = k \cdot B \cdot T \quad [\text{W}]$$

Boltzmann's constant:

$$k = 1.3806 \times 10^{-23} \quad [\text{W s K}^{-1}]$$

$$k = -228.599 \quad [\text{dB W Hz}^{-1} \text{ K}^{-1}]$$

- Concept of 'Effective Input Noise Temperature':
 - Is a measure of the noise generated by internal components of a two-port element
 - Thermodynamic temperature of resistance connected at the input to the four-port element (assumed to be noiseless) that gives the same noise power at the output of the element

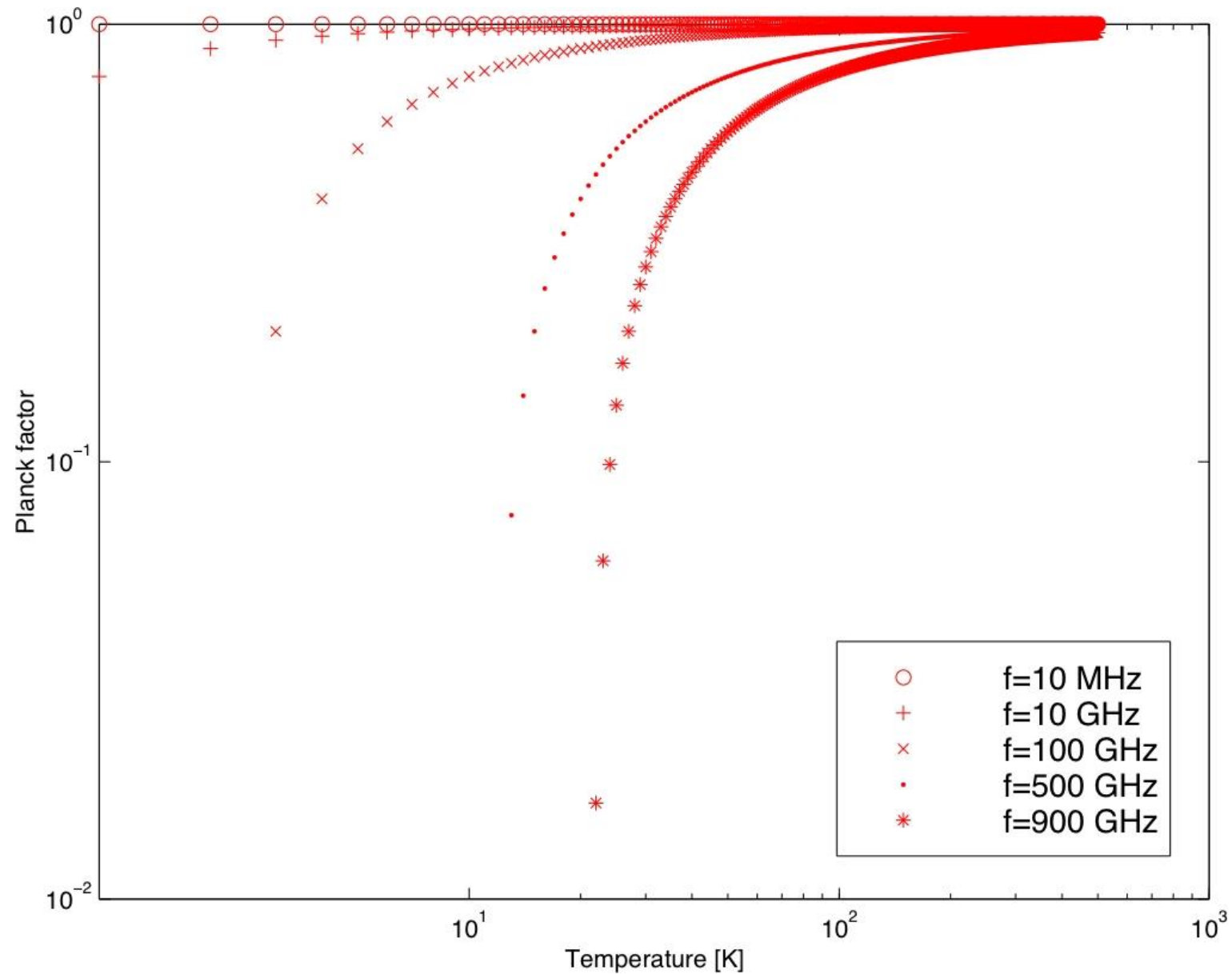


- Planck's radiation law:
 - The noise power N of an electric circuit in thermodynamical equilibrium at a temperature T within a given bandwidth B is:

$$N = k \times T \times B \times p(f)$$

- With a frequency dependent Planck's factor $p(f)$
- For satellite communications with frequencies $f < 100$ GHz the simple formula for noise is good enough

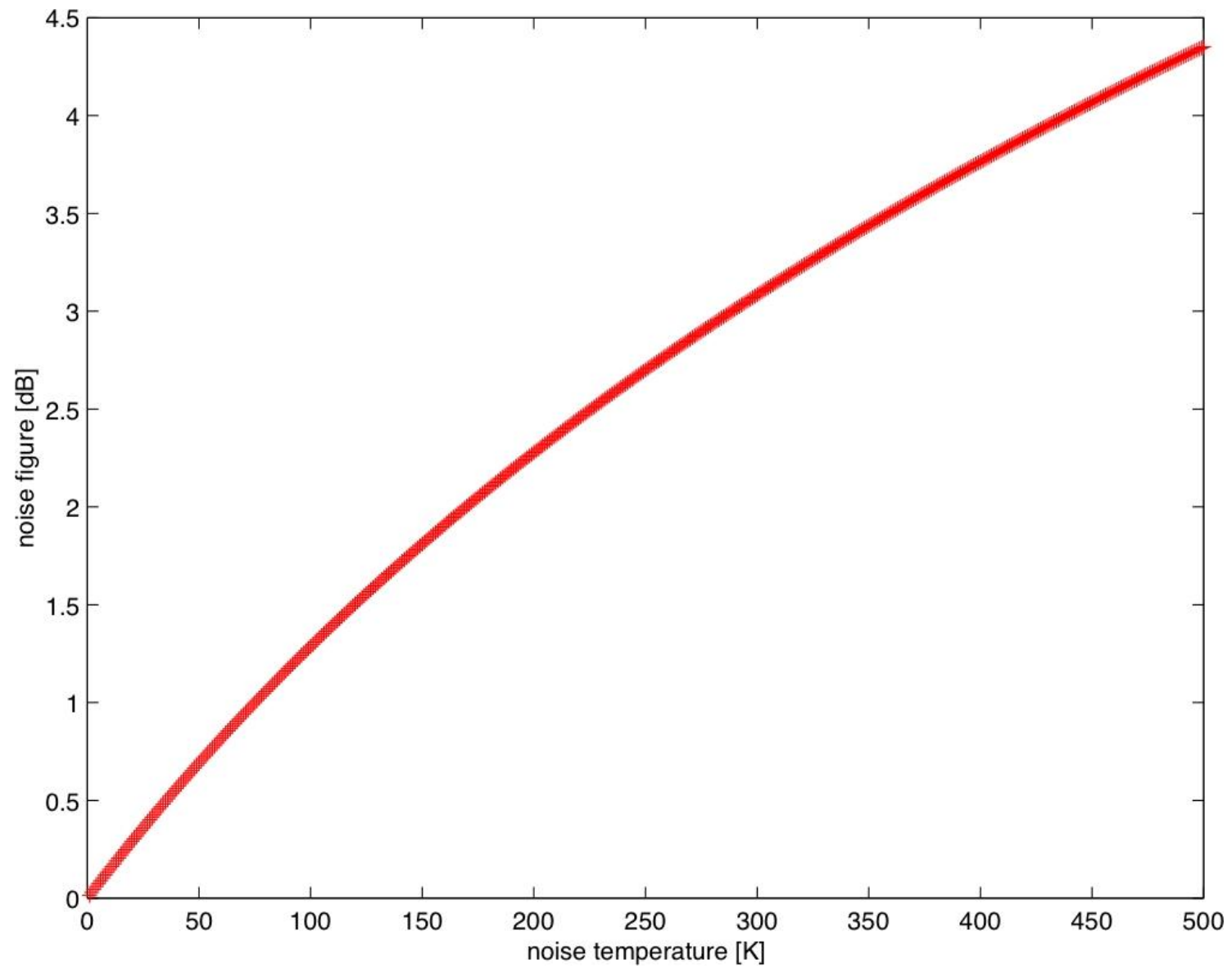
$$N = k \times T \times B$$



Planck's factor as function of temperature and frequency.

- Noise figure:
 - The ratio of total noise at the output of a 4-port element to the contribution of the input noise with reference temperature T_0
 - Reference temperature is $T_0=290$ K, (in Japan $T_0=293$ K)
 - For electronic equipment often given in dB

$$F = \frac{G \cdot k \cdot (T_e + T_0) \cdot B}{G \cdot k \cdot T_0 \cdot B} = \frac{T_e + T_0}{T_0}$$
$$= 1 + \frac{T_e}{T_0} \quad [/]$$



Noise figure F versus noise temperature T .

- System noise temperature,
 - For example at receiving earth station or satellite
 - At the receiver input (useful convention)
 - Contributions from antenna, feeder and receiving system

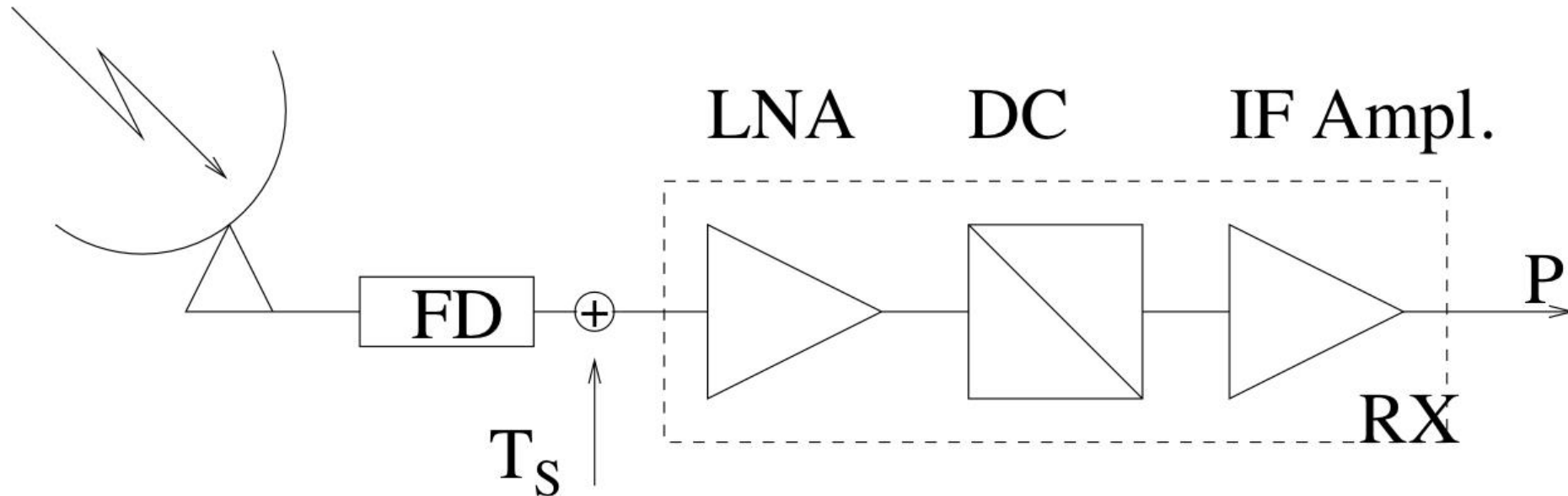


Figure: Schematic drawing of an earth station receiver.

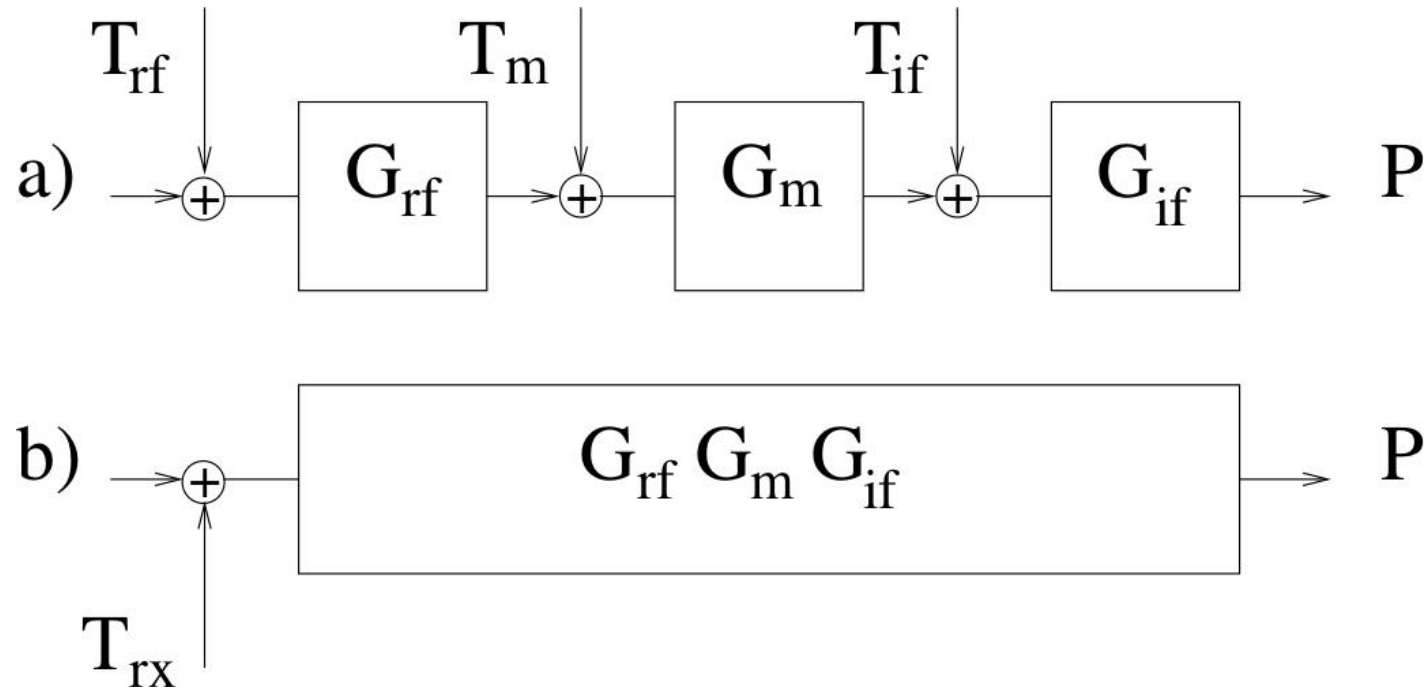
- Cascaded systems:

- Compare the power at the output of

a) the series of individual elements

and

b) one combined element



- First element should have high gain and low effective input noise temperature (==> low noise amplifier – LNA)
- The temperatures and gains of the following elements are less important

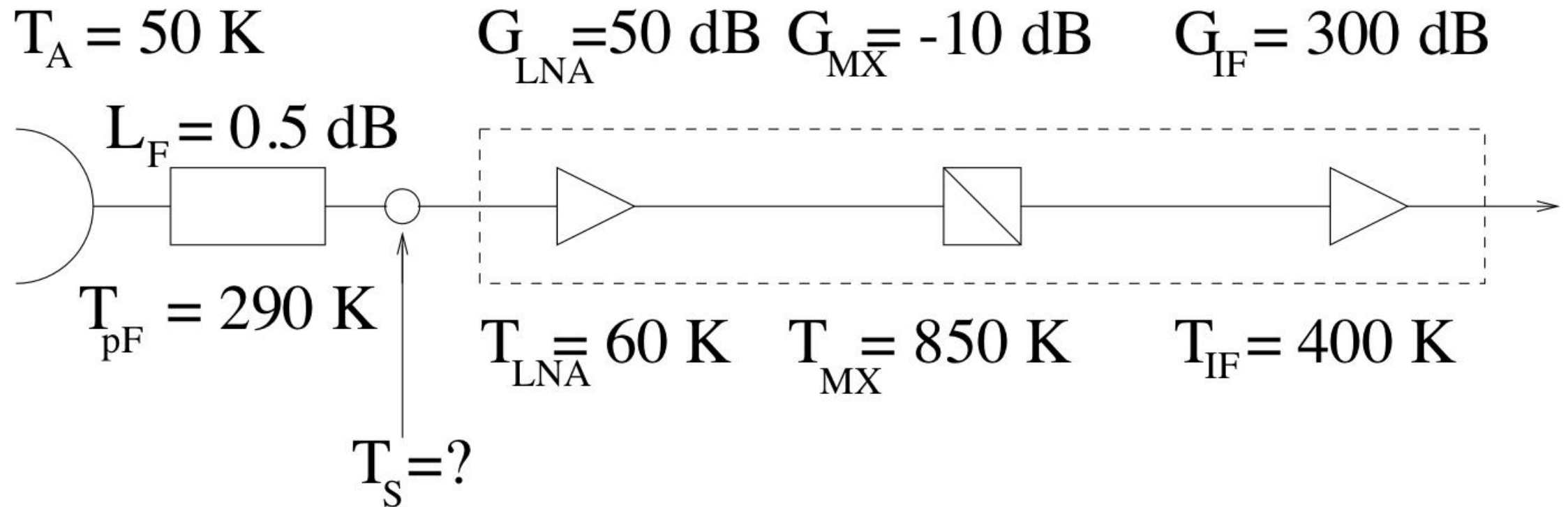
- Noise temperature of an attenuator:
 - Attenuator contains passive elements at ambient physical temperature T_p
 - Attenuation L_{att} (gain $G_{att} = 1/L_{att}$)
 - Effective input noise temperature

$$T_{e.att} = (L_{att} - 1) \times T_p$$

- An antenna picks up noise from all radiating bodies within its radiation pattern
- Brightness temperature of a radiating body in a specific direction $T_b(\theta, \phi)$
- Antenna gain in this direction $G(\theta, \phi)$
- Antenna noise temperature T_A

$$T_A = \frac{1}{4\pi} \iint T_b(q, f) \times G(q, f) \sin(q) dq df$$

- Example-1:
 - Calculate the system noise temperature for the following receiving system



- Figure of merit (G/T):
 - Relation of gain to system noise temperature (G/T)
 - Useful as a measure for the quality of a receiving station
 - Quality of the overall link (C/N₀) can be split up in contributions from (equation in dB):
 - sending side (EIRP, losses)
 - path loss (L_{FS}, L_A)
 - receiving side (Figure of merit G/T, losses)

$$\frac{C}{N_0} = \begin{aligned} & \text{sending side} \\ & (P_{tx} + G_{t.max} - L_t - L_{f.tx}) \\ & \text{path loss} \\ & (-L_{FS} - L_A) \\ & \text{receiving side} \\ & (+ G_{r.max} - L_r - L_{pol} - L_{f.rx} - k - T_s) \end{aligned}$$

Some examples:

Type	f [GHz]	Te [K]	G [dB]	cost
Cryogenic	4	15	30	high
Parametric	20	<100	30	
Cooled	4	35	30	medium
Parametric	12	85	30	
(Peltier)	20	150	30	
Ambient	4	55	30	medium
Parametric	12	150	30	
FET cryogenic	20	200	30	
FET cooled	4	40	60	low
(Peltier)	12	120	60	
	12	160	60	
	20	180	45	
FET ambient	4	70	60	low
	12	130	60	
	12	180	60	
	20	350	22	
HEMT	22	300	16	low

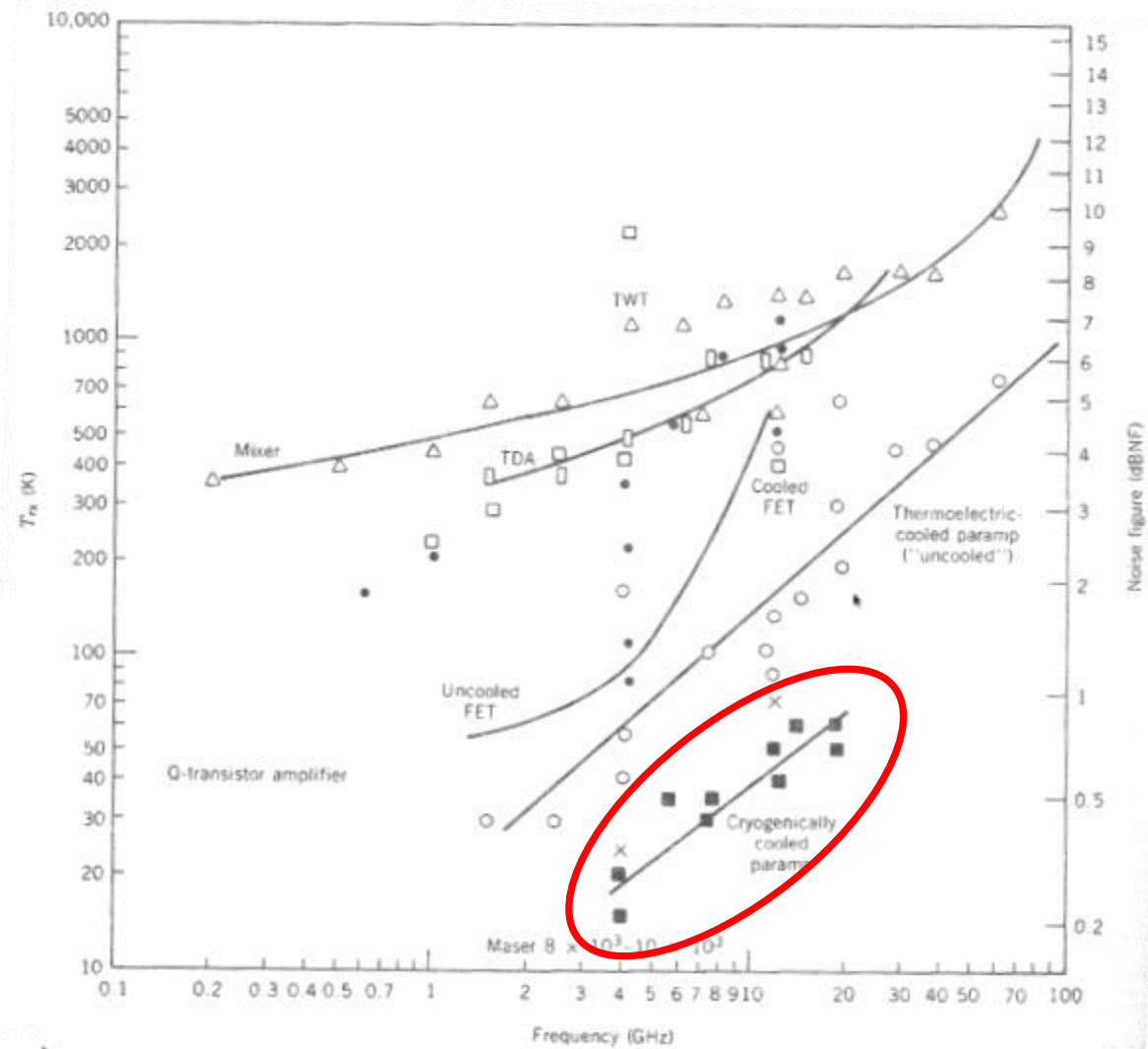
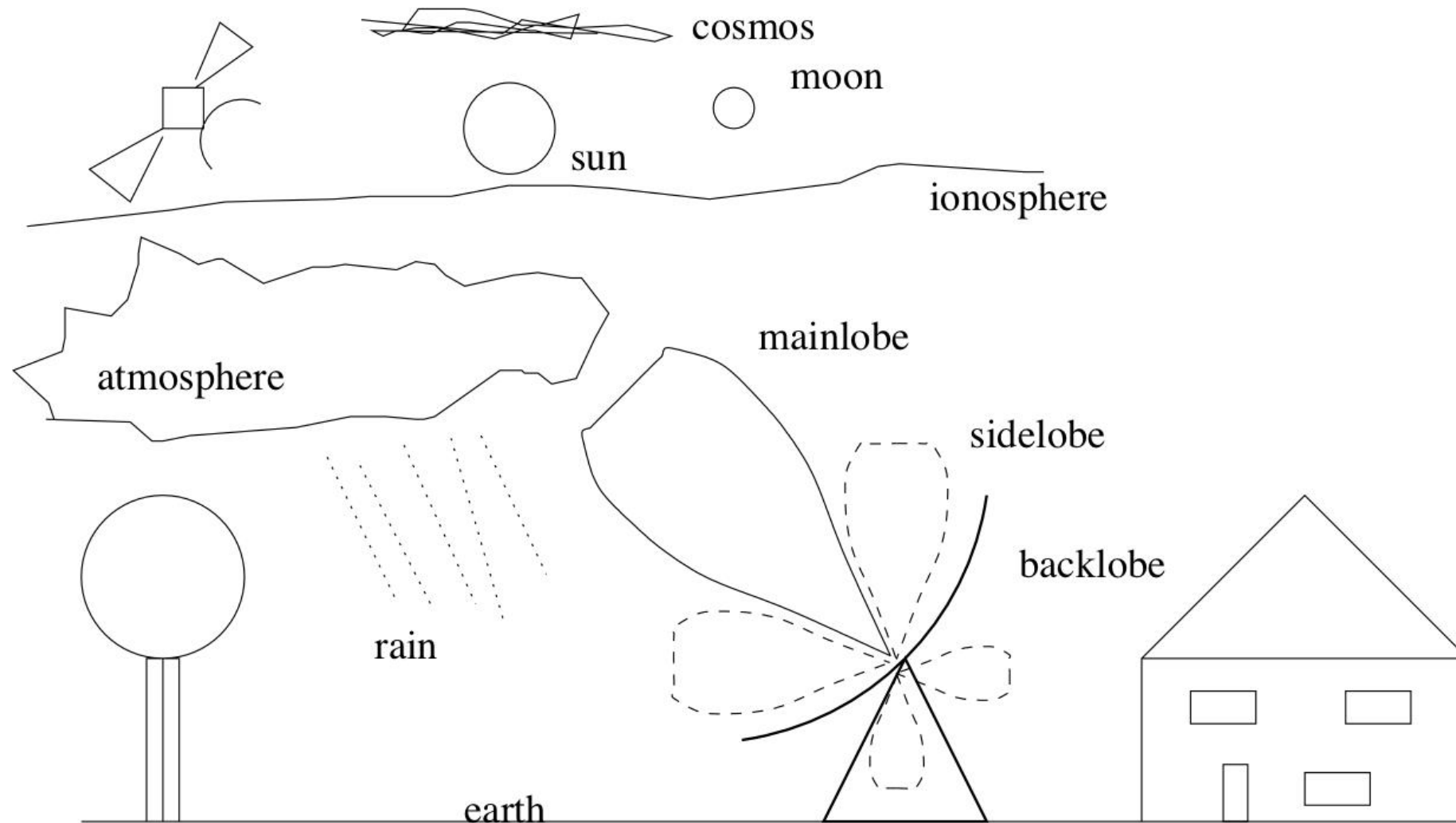
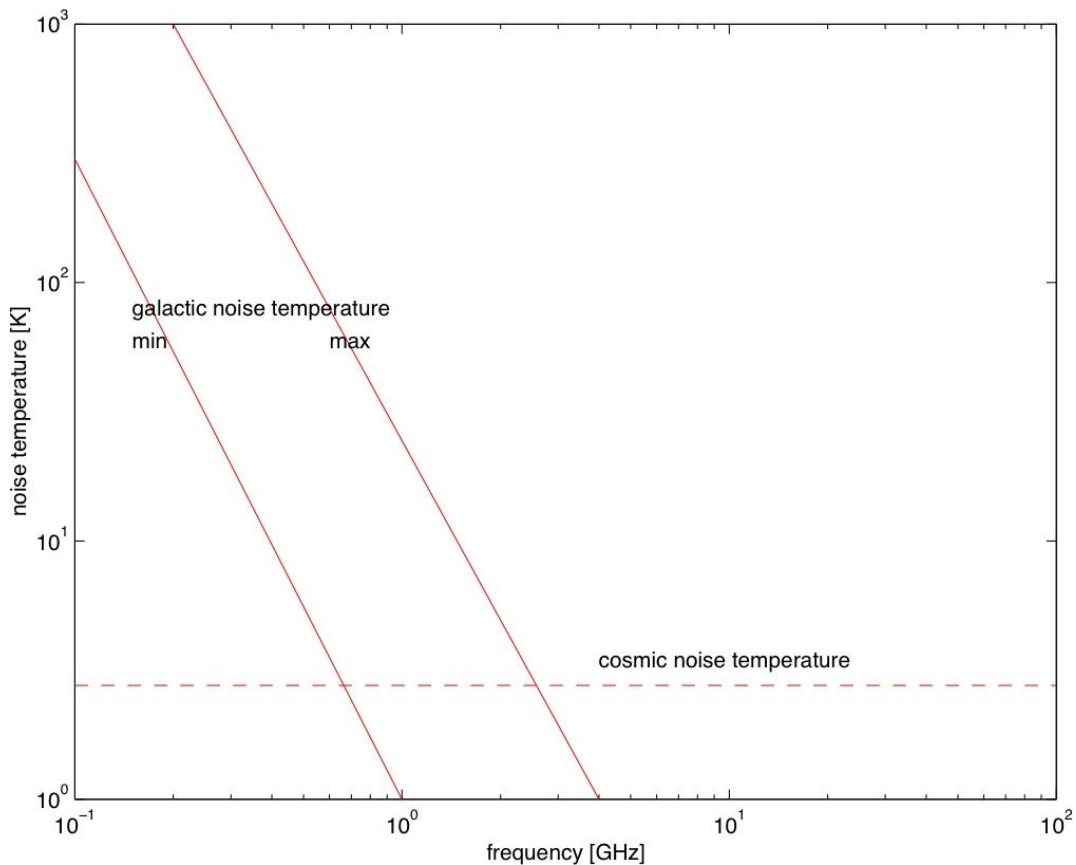


FIGURE 9.20 Noise temperature and noise figure ($T_n = 290$ K) for RF amplifiers and mixers.

- = Cryogenically cooled paramp
- = Thermoelectrically cooled paramp
- = Tunnel diode amplifier
- △ = Mixer
- = Uncooled paramp
- = Travelling wave tube amplifier
- × = High electron mobility transistor (HEMT) low noise converter (1993)

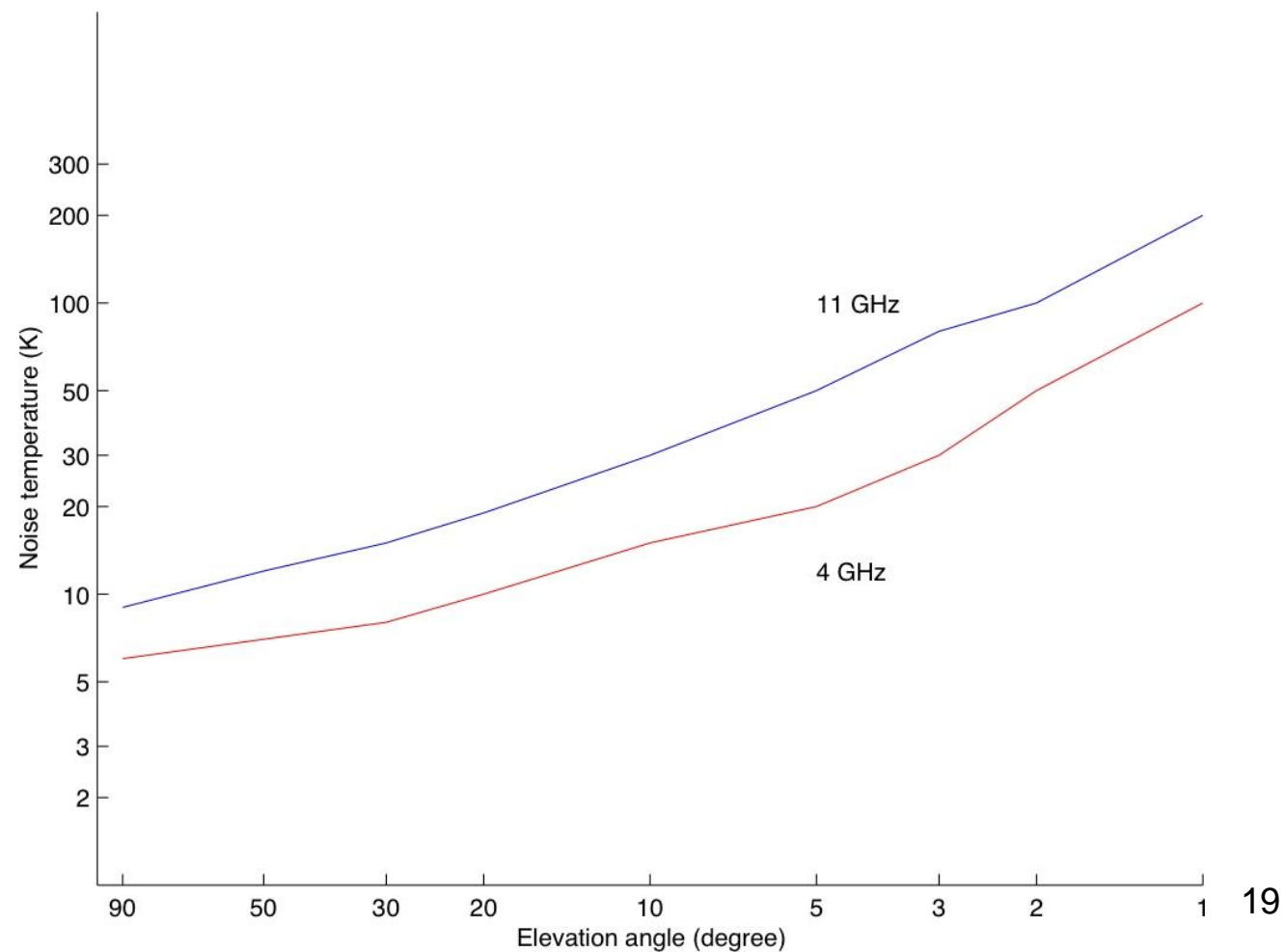
- Contributions to the antenna noise temperature of an earth station:





Example for cosmic and galactic noise temperatures

Example for atmospheric noise temperature ("clear-sky") as function of frequency and elevation



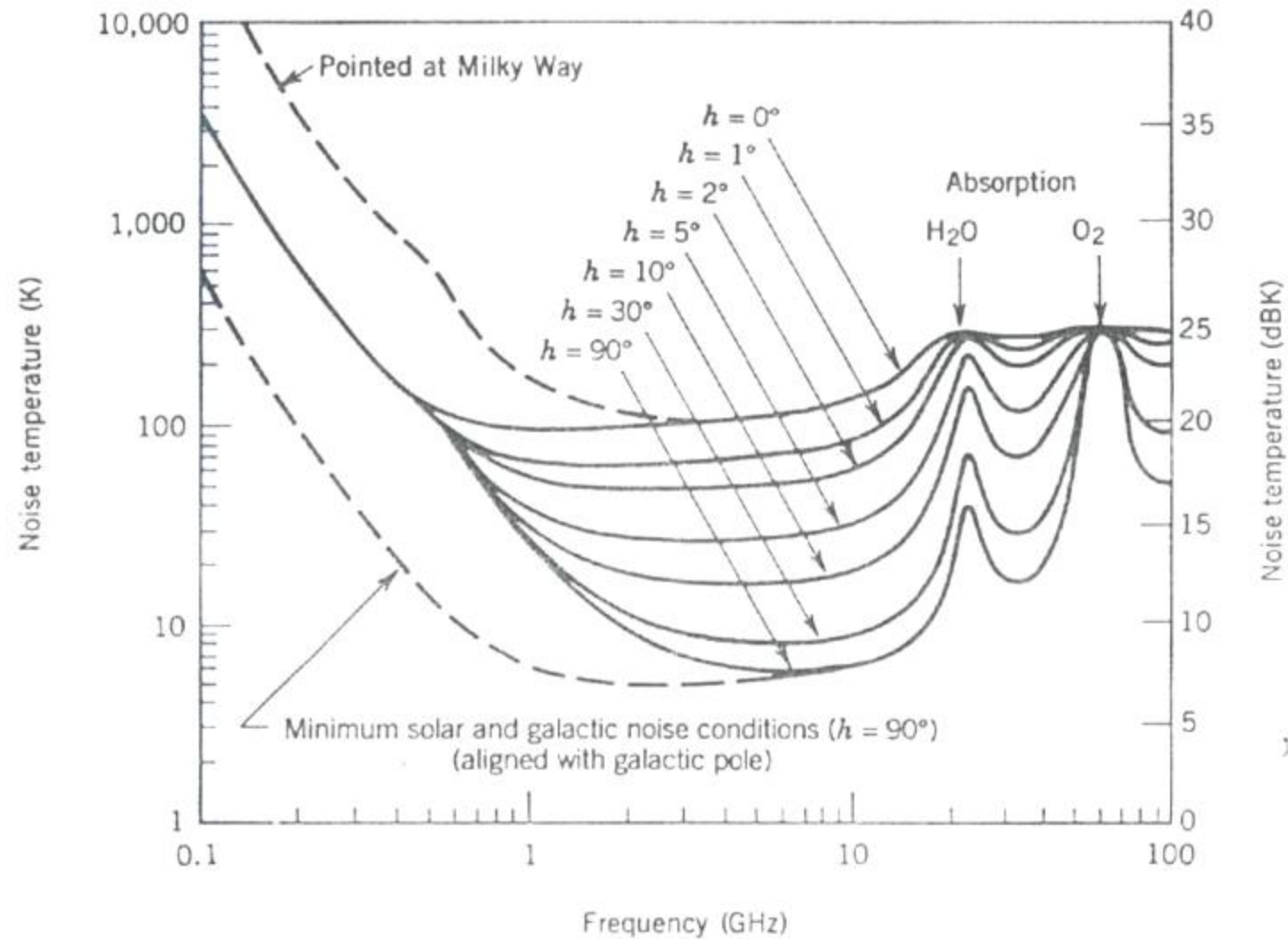
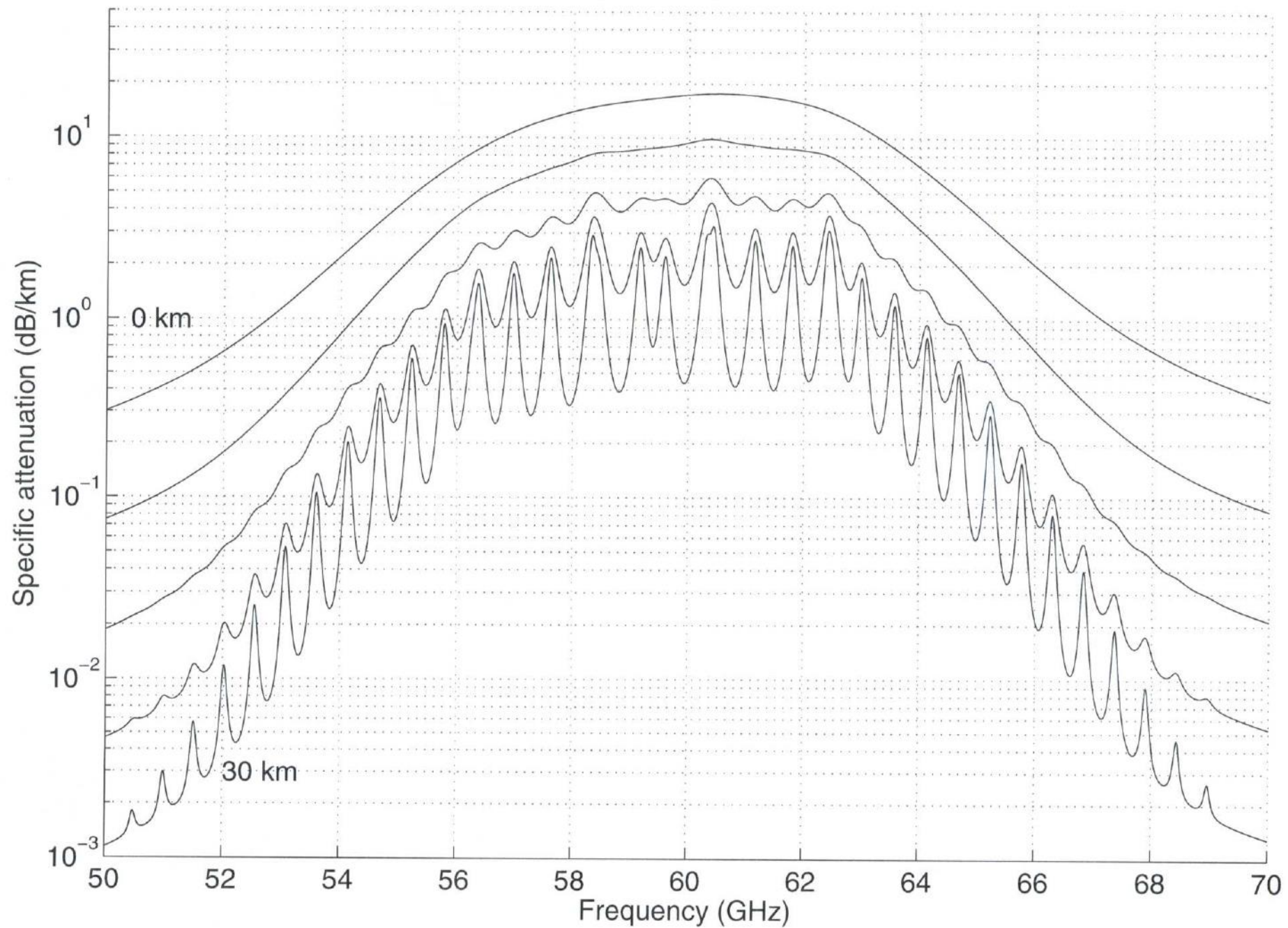


FIGURE 9.8 Earth station antenna noise temperature.

Assumptions:

- Antenna has no earth-looking sidelobes or a backlobe (zero ground noise)
- Antenna is lossless
- h is antenna elevation angle (deg)
- Sun not considered
- Cool, temperate-zone troposphere

Oxygen absorption lines



- Rain has two effects:
 - Rain acts as an attenuator with temperature $T_m = 275$ K and contributes with a noise temperature:

$$T_{Rain} = T_m \cdot (1 - 1/A_{Rain})$$

- Rain increases the sky noise (for downlink)

$$T_{skyrain} = T_{sky} / A_{Rain}$$

- Noise from the ground T_G :

- Picked up from mainlobe in low elevations
- Picked up from back- and sidelobes
- Approximate calculation:

$$T_{Ground} = \sum_i^{lobes} G_i \cdot \frac{\Omega_i}{4 \cdot \pi} \cdot T_G$$

- Where G_i and Ω_i are the mean gain and the solid angle of the lobe, respectively, T_G the ground contribution is depending on elevation

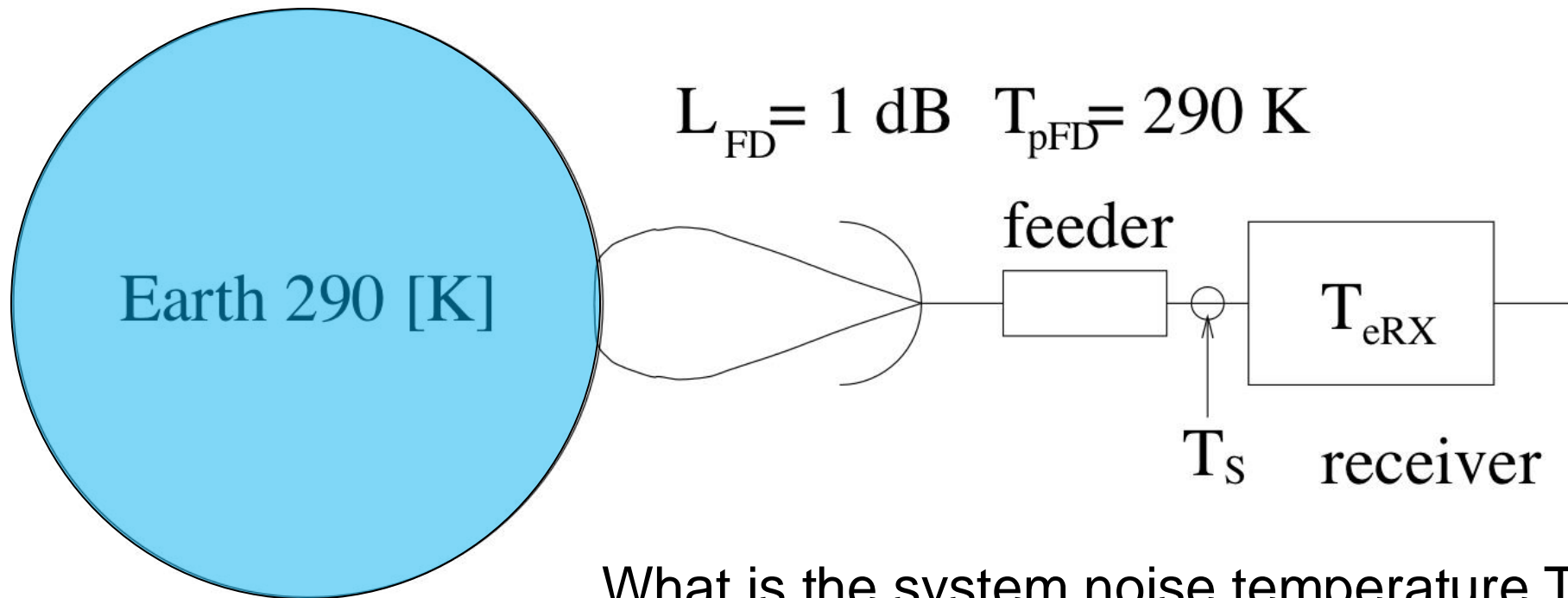
$$T_G = 290 \text{ K} \quad \text{for} \quad \varepsilon < -10^\circ$$

$$T_G = 50 \text{ K} \quad \text{for} \quad 0^\circ < \varepsilon < 10^\circ$$

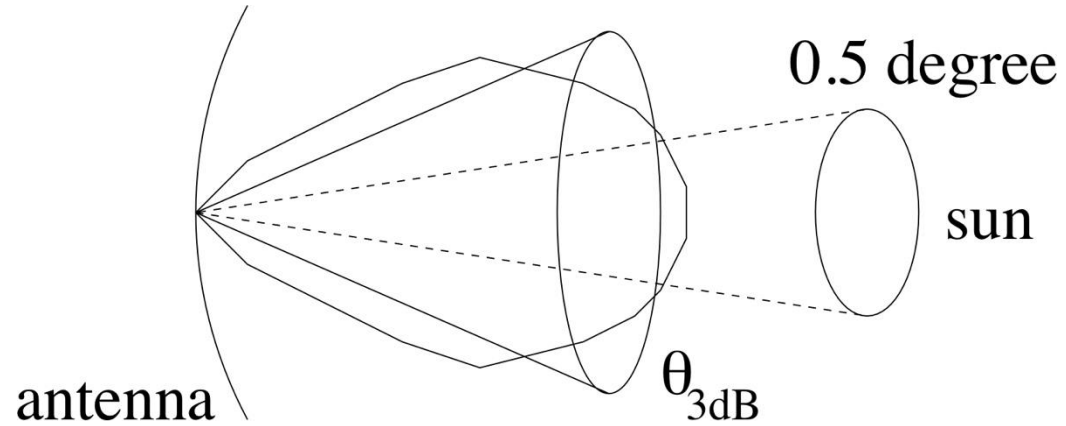
$$T_G = 150 \text{ K} \quad \text{for} \quad -10^\circ < \varepsilon < 0^\circ$$

$$T_G = 10 \text{ K} \quad \text{for} \quad 10^\circ < \varepsilon < 90^\circ$$

- Satellite antenna noise temperature:
 - Satellite captures noise from the earth and space
 - Satellites "sees" the earth approximately as "black body" of about 290 K
 - Thus there is always high antenna noise temperature onboard satellites
 - Not much can be gained by using LNAs onboard satellites



- Additional noise due to sun interference

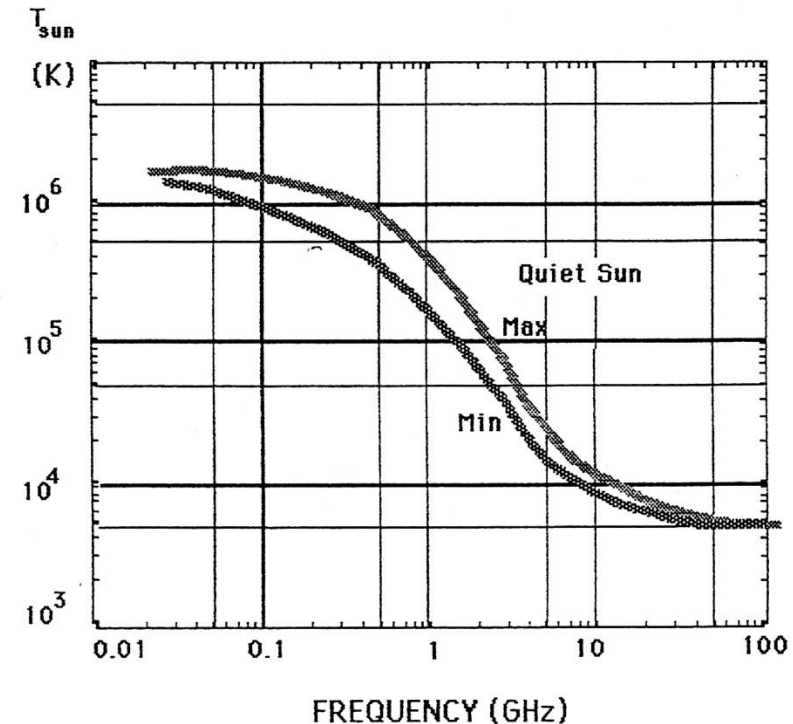


$$\Delta T_A = T_{sun} \left(\frac{0.5}{\theta_{3dB}} \right)^2 \quad \text{if } \theta_{3dB} > 0.5^\circ$$

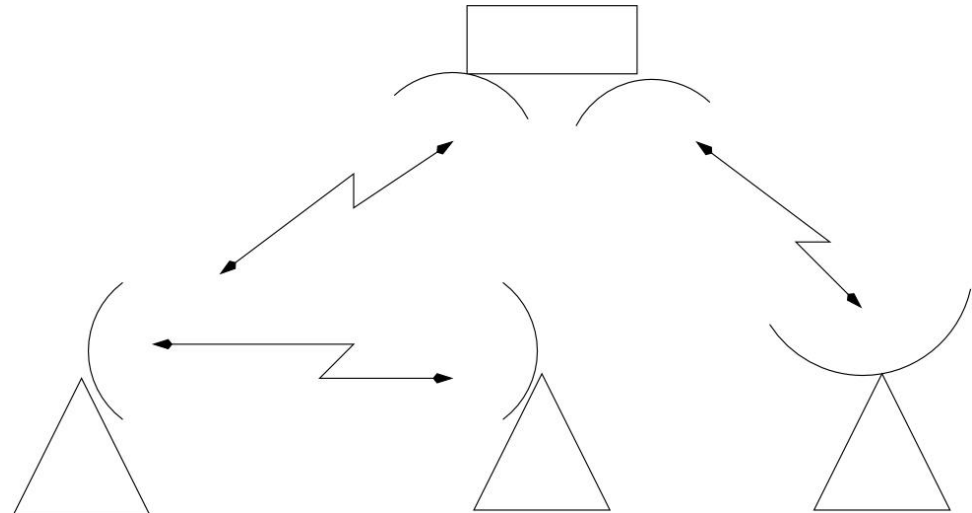
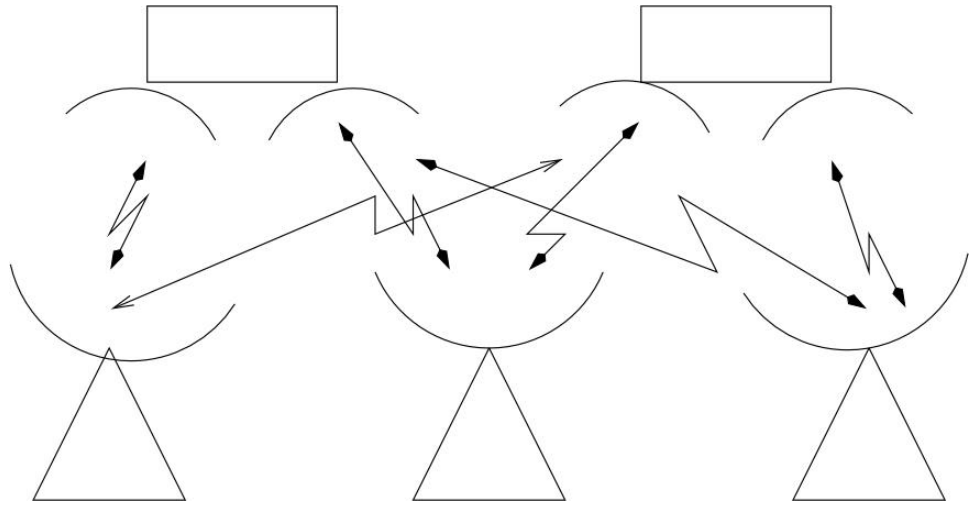
$$\Delta T_A = T_{sun} \quad \text{if } \theta_{3dB} \leq 0.5^\circ$$

The mean brightness temperature of the sun is frequency dependent, e.g.:

$$T_{Sun} = \frac{1.9610^5}{f} \left[1 + \frac{\sin 2\pi \left(\frac{\log 6(f - 0.1)}{2.3} \right)}{2.3} \right]$$

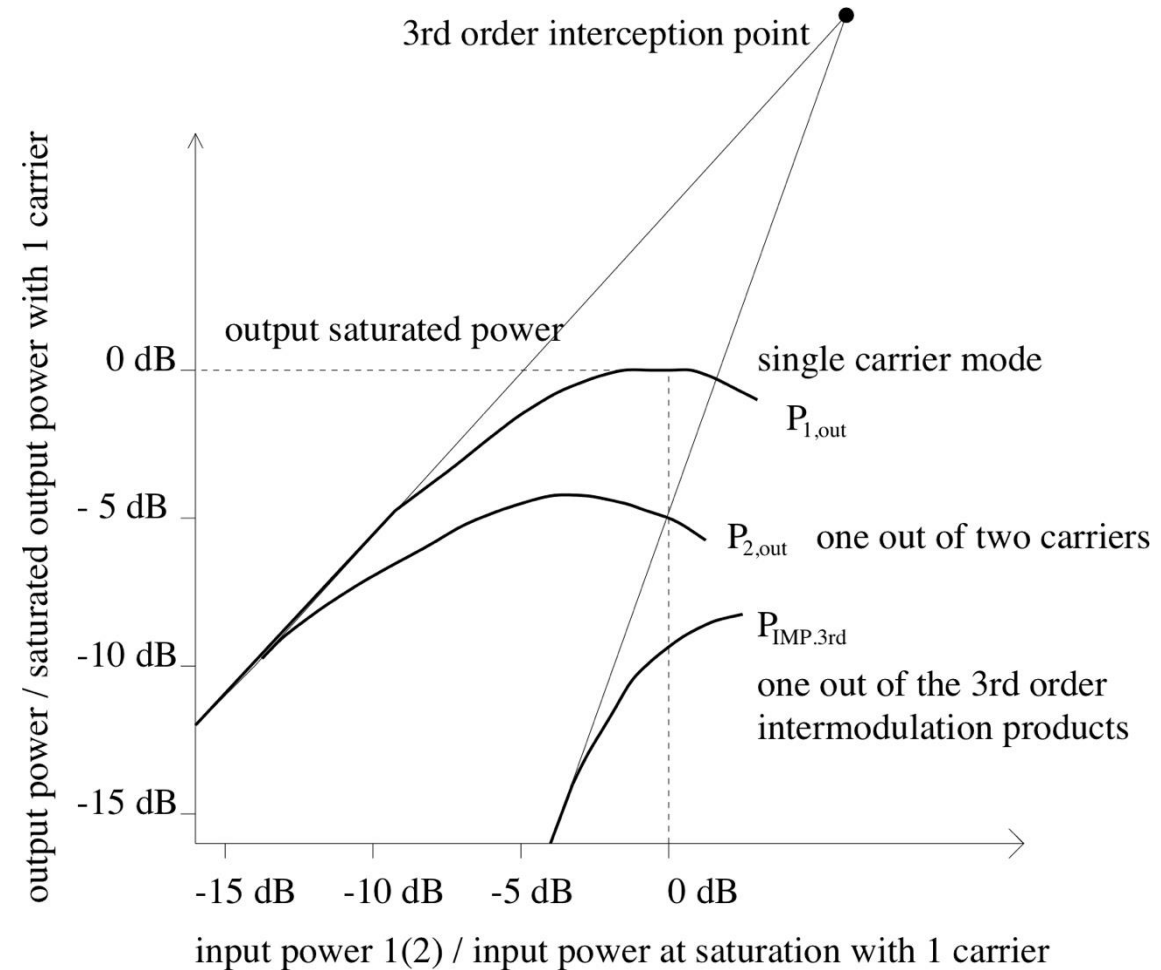


- Interference noise:
 - Between different satellite systems
 - Between satellite and terrestrial systems



- Intermodulation noise:
 - Due to non-linearity of satellite transponders
 - Transponders have a saturation point
 - Intermodulation products are generated when the transponder is operated in the saturation region

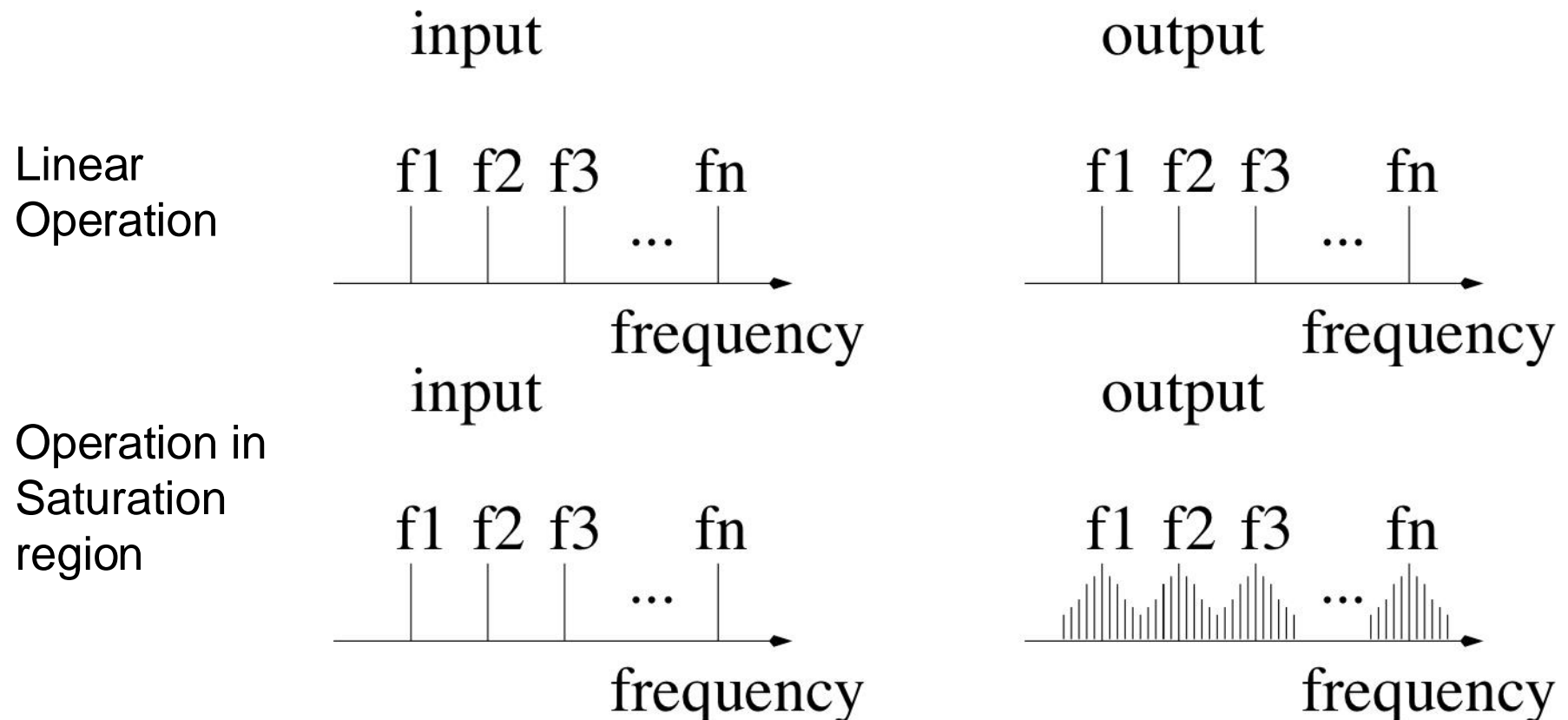
Simplified model of
a satellite channel:



- Output in linear region is the amplified carriers
- Output in the saturation region is amplified carriers plus intermodulation products f_{IM} :

$$f_{IM} = m_1 \cdot f_1 + m_2 \cdot f_2 + \dots + m_n \cdot f_n$$

- Only odd products of intermodulation noise are dangerous
- 3rd and 5th order intermodulation products are most important
- Of interest is the third order interception point



- Repeater as non-linear device:
 - Input voltage of e.g. two carriers with $f_i = \omega_i$

$$V_i = A \times \cos(W_1 \times t) + B \times \cos(W_2 \times t)$$

- Output voltage includes higher order products

$$\begin{aligned}
 V_o = & a \times \left(A \times \cos(W_1 \times t) + B \times \cos(W_2 \times t) \right) \\
 & + b \times \left(A \times \cos(W_1 \times t) + B \times \cos(W_2 \times t) \right)^2 \\
 & + c \times \left(A \times \cos(W_1 \times t) + B \times \cos(W_2 \times t) \right)^3 \\
 & + \dots + \\
 & + \dots + \frac{3}{4} c \times A^2 \times B \times \cos((2 \times W_1 - W_2) \times t) \\
 & + \dots + \frac{3}{4} c \times B^2 \times A \times \cos((2 \times W_1 - W_2) \times t)
 \end{aligned}$$

Example for a satellite payload:

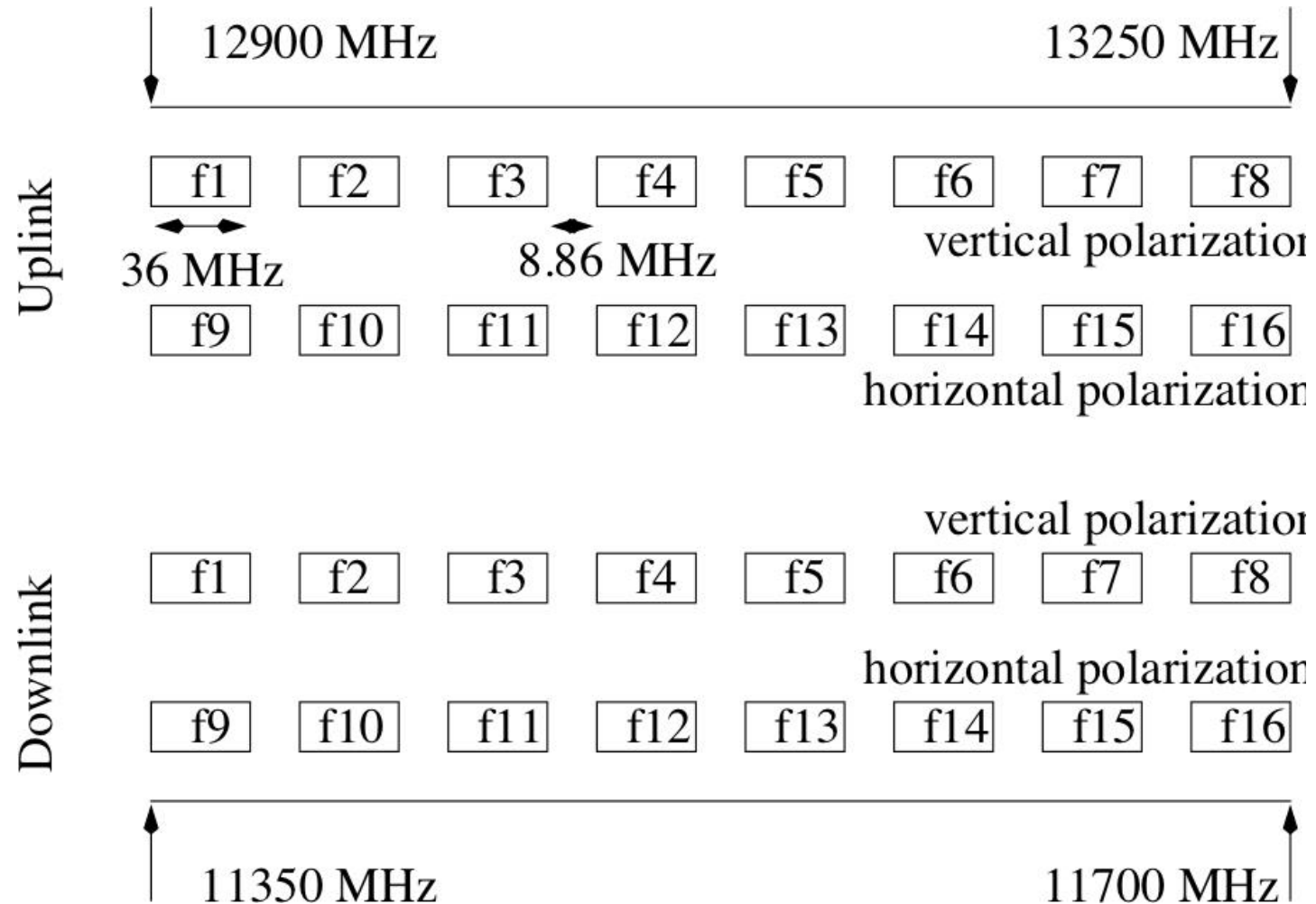
16 transponders with 36 MHz each,

uplink frequencies

12.90–13.25 GHz and

downlink frequencies

11.25–11.70 GHz



f1=12918.00 MHz	f5=13097.43 MHz
f2=12962.86 MHz	f6=13142.29 MHz
f3=13007.71 MHz	f7=13187.14 MHz
f4=13052.57 MHz	f8=13232.00 MHz

- 2nd order intermodulation products:
=> outside carriers
- 3rd order intermodulation products:
=> affect f2 and f5
- 4th order intermodulation products:
=> outside carriers
- 5th order intermodulation products:
=> affect f6 and f1

Example-2

A satellite transponder with power $P_{tx} = 10$ W is connected via a feeder (feeder loss $L_{f.tx} = 1$ dB) to a transmitting antenna with maximal gain $G_{max} = 33$ dBi and depointing loss $L_t = 3$ dB. An earth station with a reflector of diameter $D = 3.6$ m and aperture efficiency of 0.6 is connected via a feeder of loss $L_{f.rx} = 0.5$ dB with physical temperature $T_{ph.fd} = 290$ K to a receiver with effective input noise temperature of $T_{e.rx} = 60$ K. The depointing loss of the antenna is $L_t = 1$ dB and the half-power beam-width of the reflector is $\theta_{3dB} = 0.5$ degrees. The sky temperature is $T_{sky} = 8$ K and the ground noise temperature is $T_{Ground} = 20$ K. The free space loss is $L_{FS} = 206$ dB.

Questions:

1. What is the system noise temperature T_{sys} of the earth station under clear-sky conditions?
2. What is the system noise temperature of the earth station under rain conditions, assuming a rain attenuation of $A_{\text{rain}} = 6$ dB and thermodynamic temperature of $T_{\text{ph}} = 275$ K?
3. What does the the rain mean for the link performance of this downlink?

Suggested solutions

Antenna temperature in clear sky: $T_A = T_{sky} + T_{ground} = 28 \text{ K}$

$$T_{A.clear-sky} = 28 \text{ K}$$

Antenna temperature in rain: $T_A = T_{sky}/A_{rain} + T_m (1 - 1/A_{rain}) + T_{ground}$

$$T_{A.rain} = 228 \text{ K}$$

System noise temperature: $T_S = T_A/L_{f.RX} + T_{ph.frx} (1 - 1/L_{ph.frx}) + T_{e.RX}$

1) T_S in clear sky: $T_S = 117 \text{ K}$

2) T_S in rain: $T_S = 295 \text{ K}$

Suggested solutions

Satellite EIRP: $EIRP = P_{tx} + G_{t.max} - L_{ftx} - L_t$
 $EIRP = 39 \text{ dBW}$

Gain of earth station: $G_{r.max} = \eta \cdot ((\pi \cdot 70^\circ) / \theta_{3dB})^2 \text{ []}$
 $G_{r.max} = 51 \text{ dBi}$

$$G_r = G_{r.max} - L_r - L_{frx} = 49.5 \text{ dBi}$$

Suggested solutions

Figure of merit of earth station:

$$(G/T)_{es} = G_r - T_S$$

$(G/T)_{es}$ in clear sky:

$$(G/T)_{es} = 28.5 \text{ dBi K}^{-1}$$

$(G/T)_{es}$ in rain:

$$(G/T)_{es} = 24.5 \text{ dBi K}^{-1}$$

Path loss:

L_{path} in clear sky:

$$L_{path} = L_{FS} = 206 \text{ dB}$$

L_{path} in rain:

$$L_{path} = L_{FS} + A_{rain} = 212 \text{ dB}$$

Suggested solutions

Downlink performance: $(C/N_0)_{down} = EIRP - L_{path} + (G/T_S)_{es} - k$

$$(C/N_0)_{down} \text{ in clear sky: } (C/N_0)_{down} = 90.1 \text{ dBHz}$$

$$(C/N_0)_{down} \text{ in rain: } (C/N_0)_{down} = 80.1 \text{ dBHz}$$

3) Rain reduces the system performance by 10 dB.

Short summary of today's topics

- Noise models
- Noise temperature, effective input noise temperature
- Noise figure
- System noise temperature T_{sys}
- Cascades systems
- Figure of merit
- Gases and rain
- Non-linear devices, intermodulation noise