

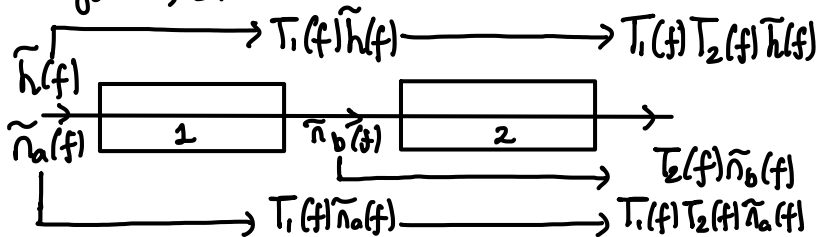
## Gravitational Waves Notes 2

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### Maggiore Vol 2 - 7.1 - Noise Spectral Density

- Output of a GW detector is a time-series f.ex. phase shift of light recombining after travelling through interferometer
- Mixture of true GW signal and noise
- GW described by tensor  $h_{ij}$ , so input  $h(t) = D_{ij} h_{ij}(t)$   
 $D_{ij}$  = constant detector tensor
- Output of detector  $\tilde{h}_{out}(f)$  in freq space is linear function of input via transfer function in linear system  $\tilde{h}_{out}(f) = T(f)\tilde{h}(f)$
- BUT output also has noise, so output in time space is  
 $s_{out}(t) = h_{out}(t) + n_{out}(t)$

- A detector can be modelled as a series of linear systems with their own transfer functions  $T_i(f)$  and their own noises  $n_i(f)$ . The overall transfer function is given by  $T(f) = \prod_{i=1}^N T_i(f)$  and the noises given by propagating the noise through the remaining transformers, shown here:



If  $n_{out}(t)$  is noise output, then define total noise input as  $\tilde{n}(f) = T^{-1}(f)\tilde{n}_{out}(f)$  and  $s(t) = h(t) + n(t)$

- Detection problem: separate  $h(t)$  from  $n(t)$
- Advantage of referring to everything as the inputs is  $n(t)$  gives measure of min of  $h(t)$ , since  $\mathcal{Q}(D_{ij}) = 1$ , whereas  $h_{out}$  depends on transfer function
- Assume noise is stationary: different Fourier components are uncorrelated, ensemble average of the components is  
 $\langle \tilde{n}^*(f)\tilde{n}(f') \rangle := \delta(f-f') \frac{1}{2} S_n(f)$
- $n(t)$  real  $\Rightarrow \tilde{n}(-f) = \tilde{n}^*(f) \Rightarrow S_n(-f) = S_n(f)$
- $n(t)$  dimensionless assumed  $\Rightarrow [S_n(f)] = \text{Hz}^{-1}$
- wlog  $\langle n(t) \rangle = 0$
- If  $f = f'$ , restrict time interval to  $-\tau/2 < t < \tau/2$  finite  
 $\delta(f=0) \rightarrow \left[ \int_{-\tau/2}^{\tau/2} dt e^{2\pi i f t} \right]_{f=0} = \tau$

$$\Rightarrow \langle |\tilde{n}(f)|^2 \rangle = \frac{1}{2} S_n(f) \tau$$

- On the interval  $[-\tau/2, \tau/2]$ , the Fourier modes have discrete frequencies  $f_n = n/\tau$ , so resolution in frequency is  $\Delta f = 1/\tau$   
 $\Rightarrow \frac{1}{2} S_n(f) = \langle |\tilde{n}(f)|^2 \rangle \Delta f$
- $\langle n^2(t) \rangle = \langle n^2(t=0) \rangle = \int_{-\infty}^{\infty} df df' \langle \tilde{n}^*(f)\tilde{n}(f') \rangle$   
 $= \frac{1}{2} \int_{-\infty}^{\infty} df S_n(f)$   
 $= \int_{-\infty}^{\infty} df S_n(f) \rightarrow S_n(f)$  is noise spectral density

### or power spectral density

Specifically, single-sided as we only integrate over physical frequencies  $f > 0$ .

- Note:  $n(t) \rightarrow \lambda n(t) \Rightarrow S_n(f) \rightarrow \lambda^2 S_n(f)$ , but strain sensitivity  $\sqrt{S_n(f)}$  increases linearly.
- Technically, one should define the auto-correlation function  $R(\tau) = \langle n(t+\tau)n(t) \rangle$  and perform the FT on this  
 $\frac{1}{2} S_n(f) \equiv \int_{-\infty}^{\infty} d\tau R(\tau) e^{2\pi i f \tau}$   
 $\Rightarrow R(\tau) = \frac{1}{2} \int_{-\infty}^{\infty} df S_n(f) e^{-2\pi i f \tau} = \langle n(t+\tau)n(t) \rangle$   
 $\Rightarrow R(0) = \langle n^2(t) \rangle = \int_{-\infty}^{\infty} df S_n(f)$  which is equivalent to before if the FT of  $n(t)$  exists.

## 7.2 Pattern Functions & Angular Sensitivity

A GW with given propagation direction  $\hat{n}$  is written  

$$h_{ij}(t, \underline{x}) = \sum_{A=+, \times} e_{ij}^A(\hat{n}) \int_{-\infty}^{\infty} df h_A(f) e^{-2\pi i f (t - \hat{n} \cdot \underline{x}/c)}$$

- For a ground based interferometer,  $2\pi f \hat{n} \cdot \underline{x} \ll 1$ , so neglect spatial dependence and we get  

$$h_{ij}(t, \underline{x}) = \sum_{A=+, \times} e_{ij}^A(\hat{n}) \int_{-\infty}^{\infty} df h_A(f) e^{-2\pi i f t} = \sum_{A=+, \times} e_{ij}^A(\hat{n}) h_A(t)$$

contracting with the tensor  $D_{ij}$  gives us

$$h(t, \underline{x}) = \sum_{A=+, \times} D_{ij} e_{ij}^A(\hat{n}) h_A(t) ; \quad \boxed{F_A(\hat{n}) = D_{ij} e_{ij}^A(\hat{n})}$$

The  $F_A(\hat{n})$  are detector pattern functions, so can write  

$$\boxed{h(t) = h_+(t) F_+(\theta, \phi) + h_\times(t) F_\times(\theta, \phi)} \quad [\text{since } \hat{n} = \hat{n}(\theta, \phi)]$$

- This is all in terms of axes  $(\hat{u}, \hat{v})$  in a plane orthogonal to  $\hat{n}$ ; if we rotate by  $\psi$  in the transverse plane:

$$\hat{u}' = \hat{u} \cos \psi - \hat{v} \sin \psi$$

$$\hat{v}' = \hat{u} \sin \psi + \hat{v} \cos \psi$$

and the amplitudes of the polarisations change to

$$h'_+ = h_+ \cos 2\psi - h_\times \sin 2\psi$$

$$h'_\times = h_+ \sin 2\psi + h_\times \cos 2\psi$$

$$(e_{ij}^+)'(\hat{n}) = \hat{u}'_i \hat{u}'_j - \hat{v}'_i \hat{v}'_j \quad (e_{ij}^\times)'(\hat{n}) = \hat{u}'_i \hat{v}'_j + \hat{v}'_i \hat{u}'_j$$

$$\hookrightarrow = e_{ij}^+(\hat{n}) \cos 2\psi - e_{ij}^\times(\hat{n}) \sin 2\psi \quad \hookrightarrow = e_{ij}^+(\hat{n}) \sin 2\psi + e_{ij}^\times(\hat{n}) \cos 2\psi$$

Together this implies

$$F'_+(\hat{n}, \psi) = F_+(\hat{n}) \cos 2\psi - F_\times(\hat{n}) \sin 2\psi$$

$$F'_\times(\hat{n}, \psi) = F_+(\hat{n}) \sin 2\psi + F_\times(\hat{n}) \cos 2\psi$$

- $\int \frac{d^2 \hat{n}}{4\pi} F_+(\hat{n}) F_\times(\hat{n}) = 0$  - useful identity, independent of  $D_{ij}$
- If we take the average over all  $\psi$  as well we get  

$$\int_0^{2\pi} \frac{d\psi}{2\pi} F_+^2(\hat{n}, \psi) = \int_0^{2\pi} \frac{d\psi}{2\pi} F_\times^2(\hat{n}, \psi)$$

$$\text{and } \langle F_+^2(\hat{n}, \psi) \rangle = \langle F_\times^2(\hat{n}, \psi) \rangle \quad \langle \dots \rangle = \int_0^{2\pi} \frac{d\psi}{2\pi} \int \frac{d^2 \hat{n}}{4\pi}$$

with  $F = \sqrt{F_x^2 + F_y^2} = \sqrt{2/5}$  angular efficiency factor

- For interferometers:  $F_y(\theta, \phi; \psi=0) = \frac{1}{2}(1 + \cos^2\theta)\cos 2\phi$   
 $F_x(\theta, \phi; \psi=0) = \cos\theta \sin 2\phi$   
 $F = \frac{2}{5}$

### 7.3 Matched Filtering

- $S(t) = h(t) + n(t)$ ; usually in the situation  $|h(t)| \ll |n(t)|$
- Suppose we know the form of  $h(t)$ ; then  
 $\frac{1}{T} \int_0^T dt S(t) h(t) = \frac{1}{T} \int_0^T dt h^2(t) + \frac{1}{T} \int_0^T dt n(t) h(t)$
- $h(t)$  and  $n(t)$  are separately oscillating, the first integral on the RHS is definitely positive and a slowly varying function of time so grows as  $T$  for large  $T \Rightarrow \frac{1}{T} \int_0^T dt h^2(t) \sim h_0^2 \mathcal{O}(1)$
- $n(t), h(t)$  uncorrelated, so  $\int_0^T dt n(t) h(t) \sim T^{1/2}$  for large  $T$  (c.f. random walk), so  
 $\frac{1}{T} \int_0^T dt n(t) h(t) \sim \left(\frac{T_0}{T}\right)^{1/2} n_0 h_0$   
 $h_0$  = characteristic amplitude of  $h$      $n_0$  = characteristic amplitude of  $n$   
 $T_0$  = characteristic time
- As  $T \rightarrow \infty$ , second term goes to 0 and we have 'filtered out' noise - not possible to set  $T = \infty$  in practice, but need only  $h_0 > (T_0/T)^{1/2} n_0$
- Look for optimal signal to noise ratio. Let:  
 $\hat{S} = \int_{-\infty}^{\infty} dt S(t) K(t)$      $K(t)$  = filter function
- Signal-to-noise ratio is  $S/N$  where  
 $S = \mathbb{E}[\hat{S} | \text{signal present}]$      $N = \text{rms}(\hat{S} | \text{signal absent})$
- $\langle n(t) \rangle = 0$  implies:  
 $S = \int_{-\infty}^{\infty} dt \langle s(t) \rangle K(t) = \int_{-\infty}^{\infty} dt h(t) K(t)$   
 $= \int_{-\infty}^{\infty} df \tilde{h}(f) \tilde{K}^*(f)$
- $N^2 = [\langle \hat{S}^2(t) \rangle - \langle \hat{S}(t) \rangle^2]_{h=0} = \langle \hat{S}^2(t) \rangle_{h=0}$   
 $= \int_{-\infty}^{\infty} dt dt' K(t) K(t') \langle n(t) n(t') \rangle$   
 $= \int_{-\infty}^{\infty} dt dt' K(t) K(t') \int_{-\infty}^{\infty} df df' e^{2\pi i f t - 2\pi i f' t'} \langle \tilde{n}^*(f) \tilde{n}(f') \rangle$
- From defn of PSD, this is  $N^2 = \int_{-\infty}^{\infty} df \frac{1}{2} S_n(f) |\tilde{K}(f)|^2$   
 $\Rightarrow \frac{S}{N} = \frac{\int_{-\infty}^{\infty} df \tilde{h}(f) \tilde{K}^*(f)}{[\int_{-\infty}^{\infty} df \frac{1}{2} S_n(f) |\tilde{K}(f)|^2]^{1/2}}$
- Now want to know which  $K$  maximises this
- Define  $(A|B) = \text{Re} \int_{-\infty}^{\infty} df \frac{\tilde{A}^*(f) \tilde{B}(f)}{(\frac{1}{2}) S_n(f)}$   
 $= 4 \text{Re} \int_{-\infty}^{\infty} df \frac{\tilde{A}^*(f) \tilde{B}(f)}{S_n(f)}$   
 $\Rightarrow \frac{S}{N} = \frac{(V|h)}{(V|V)^{1/2}}$      $V(t)$  is func. s.t.  $\tilde{V}(f) = \frac{1}{2} S_n(f) \tilde{K}(f)$
- This is a scalar product, so the solution is simple; we

want to maximise  $U$  along  $S$  choose  $\tilde{U}(f)$  proportional to  $\tilde{h}(f)$ ,  $\boxed{\tilde{K}(f) = \text{const.} \frac{\tilde{h}(f)}{S_n(f)}}$

(const. arbitrary)

- This gives optimal  $\frac{S}{N}$  as  $\boxed{\left(\frac{S}{N}\right)^2 = 4 \int_0^\infty df \frac{|\tilde{h}(f)|^2}{S_n(f)}}$