${\rm COM}$ S 577 - Final Exam Study Guide

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<2016-12-13 Tue 12:00>

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1 Plane Curves

Curves are differentiable functions. We denote them $\boldsymbol{\alpha}:[a,b]\to\mathbb{R}^2$ (or \mathbb{R}^3). A curve $\boldsymbol{\alpha}(t)=(x(t),y(t))$ is said to be **smooth** at $t=t_0$ if its kth derivative exists for any integer k>0.

$$\boldsymbol{\alpha}^{(k)}\left(t\right) = \left(x^{(k)}\left(t\right), y^{(k)}\left(t\right)\right)$$

A curve is **piecewise smooth** if it has a domain which is the union of a finite number of subintervals of each of which the curve is smooth. A polygon is piecewise smooth for each of its sides.

A curve is a **closed parametric curve** if $\alpha(a) = \alpha(b)$. A point of self-crossing is a point $\alpha(t_1)$ for which there exists finitely many distinct values $t_1, \ldots, t_n \in [a, b]$ for $n \geq 2$ which satisfy $\alpha(t_1) = \ldots = \alpha(t_n)$ and in the case n = 2, $[t_1, t_2] \neq [a, b]$. For example, a circle is closed, but does not have self-crossings. A figure eight is closed and self-crossing.

1.1 Velocity, Speed, and Arc Length

For a curve $\alpha(t)$, the **velocity vector** at t is $\alpha'(t)$. The **speed** at t is the length $||\alpha'(t)||$. This is clear if we see the curve as a particle moving on a path. The parametrization $\alpha(t)$ is **unit-speed** if the $||\alpha'(t)|| = 1$ for all values of t. When the speed vanishes, the curve has a **cusp**. A curve is **regular** if the velocity vectors are non-zero for all t. The **arc length** is defined by:

$$\int_{a}^{b} ||\boldsymbol{\alpha}'\left(t\right)||dt$$

1.2 Unit-Speed Reparametrization

A regular curve α has a reparametrization $\tilde{\alpha}$. Not all unit-speed reparametrizations have an explicit form, but we can use their existence for many computations on curves.

1.3 Tangent (T) and Normal (N)

If we have a curve α , then at a regular point on the curve, there exists a non-zero **tangent vector** $\alpha'(t) = (x'(t), y'(t))$. It represents the velocity of the curve at the point. The normal vector is (-y'(t), x'(t)). This is a rotation of the tangent vector counterclockwise by $\frac{\pi}{2}$. If the curve was unit-speed, then the tangent and normal are unit vectors.

The cross product is non-zero. In the case of unit-speed curves, $T \times N = 1$.

1.4 Curvature (κ)

The **curvature** of a curve is defined as:

$$\kappa = \frac{d\phi}{ds}$$

where ϕ is the angle from the x-axis to the tangent and s is the arc length. The absolute curvature is $|\kappa|$.

Since $T \cdot T = 1$ and differentiating this equation yields $T' \cdot T = 0$, we see that the change of T(s) is orthogonal to the tangential direction, so it must be in the normal direction. The curvature is defined as the turning of the tangent along the direction of the unit normal. So,

$$T' = \frac{dT}{ds} = \kappa N$$

1.5 Center of Curvature, Osculating Circle, Radius of Curvature, and Total Curvature

When the curvature is positive then the **center of curvature** lies along the direction of the normal N(s) at a distance $\frac{1}{\kappa}$ from the point $\alpha(s)$. When the curvature is negative, then the center of curvature lies in the direction of -N(s) at a distance of $-\frac{1}{\kappa}$. The **radius of curvature** is $\frac{1}{|\kappa|}$. Together, these form the **osculating circle**. This circle approximates the curve locally up to the second order. The **total curvature** over a closed interval [a,b] measures the rotation of the tangent as s changes from a to b.

$$\Phi(a,b) = \int_{a}^{b} \kappa ds$$
$$= \int_{a}^{b} \frac{d\phi}{ds} ds$$
$$= \int_{a}^{b} d\phi$$
$$= \phi(b) - \phi(a)$$

1.6 Inflection and Vertex

A point s on the curve α is a **simple inflection**, or an **inflection**, if $\kappa(s) = 0$, but $\kappa'(s) \neq 0$. Intuitively, this is a point where the curve moves from one side of the tangent to the other or, for simple closed curves, when the curve switches from convex to concave or vice versa.

A point where the curvature and first j-1 order derivatives are zero but the jth derivative is non-zero is called an **inflection point of order** j. A second order inflection point is also called a point of **simple undulation** which will maintain concavity or convexity in the neighborhood on a simple closed curve.

A simple vertex, or simply vertex, is a local minimum or maximum when $\kappa' = 0$ but $\kappa'' \neq 0$.

1.7 Arbitrary Speed Curves

If α is not necessarily unit-speed, then we obtain the tangent:

$$T = \frac{\alpha'}{||\alpha'||}$$

If we let $v = \frac{\alpha'}{||\alpha'||}$, then we have that:

$$T' = \kappa v N$$

Intuitively we are simply normalizing the vectors by the speed. We also note that:

$$\kappa = \frac{\boldsymbol{\alpha}' \times \boldsymbol{\alpha}''}{||\boldsymbol{\alpha}'||^3}$$

We also find that **total curvature** is:

$$\Phi(a,b) = \int_{a}^{b} \kappa(t) ||\alpha'(t)|| dt$$

1.8 Simple Closed Curves

We can say a **simple closed curve** is one that joins up the beginning and end.

$$\alpha(t) = \alpha(u)$$
 if and only if $t - u = ka$ for some integer k

The **area** is defined as:

$$A(a) = \int \int_{int(\alpha)} dx dy$$
$$= \frac{1}{2} \int_{0}^{a} (xy' - yx') dt$$

Then, we find that if we have a simple closed curve with length l and area A,

$$A \le \frac{1}{4\pi}l^2$$

2 Space Curves

A plane curve is determined by curvature. Space curves are not. A circle on the xy – plane and a circular helix have unit curvature everywhere, but can not be transformed to one another. Instead, we measure space curves with curvature and torsion.

If we have a unit-speed space curve $\gamma\left(s\right)$ in \mathbb{R}^{3} , we can denote the unit tangent vector as $T=\gamma'\left(s\right)$. We have the real-valued **curvature function** $kappa\left(s\right)=||T'\left(s\right)||\geq0.$

2.1 Principal Normal

The principal normal of a space curve $\gamma(s)$ is the vector:

$$N\left(s\right) = \frac{1}{\kappa\left(s\right)}T'\left(s\right)$$

Since $||T'(s)|| = \kappa$, N is a unit vector. Similarly, $T' \cdot T = 0$ so T and N are orthogonal to each other.

2.2 Binormal

The **binormal** is:

$$B = T \times N$$

Then, $\{T(s), N(s), B(s)\}$ form an **orthonormal basis** of \mathbb{R}^3 where each is orthogonal to the other two. They form the **Frenet frame**.

2.3 Torsion (τ)

We define the **torsion** as:

$$B' = -\tau N$$

We can see that:

$$\frac{d\theta}{ds}B'\left(s\right)\cdot N\left(s\right) = -\tau$$

 $-\tau$ represents the rate of binormal rotation.

2.4 Frenet Formulas

Given a unit-speed curve with positive curvature and some torsion,

$$\begin{array}{ccccc} T' & = & & \kappa N \\ N' & = & -\kappa T & & +\tau B \\ B' & = & & -\tau N \end{array}$$

Remember the matrix:

$$\begin{pmatrix}
0 & \kappa & 0 \\
-\kappa & 0 & \tau \\
0 & -\tau & 0
\end{pmatrix}$$

If we multiply by the column vector $(T, N, B)^T$, then we can get the Frenet formulas.

The Frenet frame consists of the curvature function, the torsion function, the unit tangent, the principal normal, and the binormal. In cases of arbitrary-speed curves, we define:

$$\kappa = \tilde{\kappa}(s) \tau$$
 $= \tilde{\tau}(s) T = \tilde{T}(s) N$ $= \tilde{N}(s) B = \tilde{B}(s)$

This is true because the Frenet frame is **independent** of the parametrization. For example for s being an arc-length function of the curve, $\alpha(t) = \alpha(s(t))$. Only the **derivatives are affected by parametrization** where we simply use the speed v of the curve to correct the values of T', N', and B'.

$$\begin{array}{rcl} T' & = & \kappa v N \\ N' & = & -\kappa v T & +\tau v B \\ B' & = & -\tau v N \end{array}$$

Intuitively, a unit-speed curve has speed v=1, so the Frenet formulas are as defined earlier.

2.4.1 Computing the Frenet Apparatus

Theorem 3 Let α be a regular curve in \mathbb{R}^3 , and $\alpha' \times \alpha'' \neq 0$. Then,

$$T = \frac{\alpha'}{||\alpha'||}$$

$$B = \frac{\alpha' \times \alpha''}{||\alpha' \times \alpha''||}$$

$$N = B \times T$$

$$\kappa = \frac{||\alpha' \times \alpha''||}{||\alpha'||^3}$$

$$\tau = \frac{(\alpha' \times \alpha'') \cdot \alpha'''}{||\alpha' \times \alpha''||^2}$$

2.5 Approximation

Consider the unit-speed curve $\boldsymbol{\beta} = (\beta_1, \beta_2, \beta_3)$ near the point $\boldsymbol{\beta}(0)$. For a small s, each coordinate $\beta_i(s)$ is closely approximated by the Taylor series. So,

$$\boldsymbol{\beta}(s) \approx \boldsymbol{\beta}(0) + s\boldsymbol{\beta}'(0) + \frac{s^2}{2}\boldsymbol{\beta}''(0) + \frac{s^3}{6}\boldsymbol{\beta}'''(0)$$

However, we know the first and second derivative of β at s=0 are:

$$\boldsymbol{\beta}'(0) = T_0$$
$$\boldsymbol{\beta}''(0) = \kappa_0 N_0$$

Where subscript is evaluation at s = 0. Then, the third derivative is defined as:

$$\beta'''(0) = -\kappa_0^2 T_0 + \frac{d\kappa}{ds}(0) N_0 + \kappa_0 \tau_0 B_0$$

So, the approximation becomes:

$$\beta(s) \approx \beta(0) + sT_0 + \kappa_0 \frac{s^2}{2} N_0 + \kappa_0 \tau_0 \frac{s^3}{6} B_0$$

This is the **Frenet approximation of** β **near** s = 0. We can get an approximation at some s_0 by replacing 0 with s_0 and s with $s - s_0$.

2.6 Osculating Plane

Using the **Frenet approximation**, we can get the best linear or quadratic approximation of the curve. This approximation lies in the plane through the approximated point orthogonal to the binormal. This is called the osculating plane of β near s_0 .

2.7 Spherical Image

Given a unit-speed curve $\boldsymbol{\beta}$, then the curve $\boldsymbol{\sigma}(s) = T(s) = \boldsymbol{\beta}'(s)$ is called the **spherical image** of $\boldsymbol{\beta}$.

3 Algebraic Curves

3.1 Singular Points

A **singular point** is at a point where the gradient vanishes. We can not simply solve for $\nabla f = 0$ though. We have to also find a simultaneous solution for f = 0. These are singularities of the polynomial that lie on the curve.

3.2 Local Parametrization

Algebraic curves can be parametrized locally near non-singular points. They can be parametrized by x, y or both. A **local parametrization** of an algebraic curve near a non-singular point (a, b) is a parametrization $J \to \mathbb{R}^2$ of a piece of the curve including the point (a, b).

3.3 Curvature

In general, algebraic curves cannot be parametrized in a simple way. So, we use a different method of finding the curvature using the **Hessian matrix** of a polynomial.

$$H = \begin{pmatrix} f_{xx} & f_{xy} \\ f_{xy} & f_{yy} \end{pmatrix}$$

Then, the curvature is defined as:

$$\kappa = \frac{f_y^2 f_{xx} - 2f_x f_y f_{xy} + f_x^2 f_{yy}}{\left(f_x^2 + f_y^2\right)^{3/2}}$$

3.4 Inflection Points

Algebraic curves have **inflection points** at non-singular points (a, b) if and only if

$$f_y^2 f_{xx} - 2f_x f_y f_{xy} + f_x^2 f_{yy} = 0$$

at the point, changes signs as (x, y) moves through the point. So, to find all inflection points, we solve:

$$f(x,y) = 0$$

$$f_y^2 f_{xx} - 2f_x f_y f_{xy} + f_x^2 f_{yy} = 0$$

4 Surfaces

4.1 Surface Patch

A surface is viewed as a collection of homeomorphisms $\sigma: U \to S \cup W$ called surface patches where u is a . This collection is called the atlas of the surface. Every point of the surface is in at least one surface patch.

A map is smooth if each component of the map has continuous partial derivatives of all orders. A surface patch is **regular** if it is smooth and $\sigma_u \times \sigma_v \neq 0$ at every point of U. A **smooth surface** is one whose atlas consists of regular surface patches.

4.2 Orientable Surface

We define the **tangent space** at a point p of a surface to consist of the tangent vectors of all curves on the surface that pass through p. It is uniquely defined by the **unit normal** to the surface at the point, which is perpendicular to the tangent space, also known as the **tangent plane**.

$$oldsymbol{n}_{\sigma} = rac{oldsymbol{\sigma}_{u} imes oldsymbol{\sigma}_{v}}{||oldsymbol{\sigma}_{u} imes oldsymbol{\sigma}_{v}||}$$

This is called the **standard unit normal**. A surface is **orientable** if there is a canonical choice for the unit normal at every point, obtained by taking the standard unit normal of each surface patch in the atlas.

4.3 Surface Curves

If we have a curve in a surface patch, then we can calculate the arc length as:

$$s = \int_{t_0}^{t} ||\dot{\alpha}\left(w\right)||dw$$

The **normal curvature** and **geodesic curvature** of a surface curve $\gamma(t)$ are:

$$\kappa_n = \ddot{\gamma} \cdot \hat{\boldsymbol{n}}$$

$$\kappa_g = \ddot{\gamma} \cdot (\hat{\boldsymbol{n}} \times \dot{\gamma})$$

The Gaussian curvature and mean curvature of a surface patch respectively are:

$$K = \kappa_1 \kappa_2 = \frac{LN - M^2}{EG - F^2}$$

$$H = \frac{1}{2} (\kappa_1 + \kappa_2) = \frac{1}{2} \frac{LG - 2MF + NE}{EG - F^2}$$

where κ_1 and κ_2 are the principal curvatures.

The **Gauss map** is the mapping from a surface patch to the unit sphere at from $\mathbf{p} = \boldsymbol{\sigma}(u, v)$ to the point $\hat{\boldsymbol{n}}(u, v)$.

The total Gaussian curvature of some surface is:

$$\int \int_{S} K dS$$

Which is the integral of the Gaussian curvature K over S. If the surface is defined over the domain $[a,b] \times [c,d]$, then the total Gaussian curvature is:

$$\int_{c}^{d} \int_{a}^{b} K(u, v) \sqrt{EG - F^{2}} du dv$$

4.4 First Fundamental Form

First fundamental form of a surface patch represents the principal part of the square of the increment on the surface patch when the parameters u and v are increased by du and dv respectively.

$$E = \boldsymbol{\sigma}_u \cdot \boldsymbol{\sigma}_u \quad F = \boldsymbol{\sigma}_u \cdot \boldsymbol{\sigma}_v \quad G = \boldsymbol{\sigma}_v \cdot \boldsymbol{\sigma}_v$$

So the first fundamental form is:

$$ds^2 = Edu^2 + 2Fdudv + Gdv^2$$

Thus, the **arc length** is:

$$s = \int_{t_0}^{t} \sqrt{E\dot{u}^2 + 2F\dot{u}\dot{v} + G\dot{v}^2} dt$$
$$= \int_{\alpha} = \sqrt{Edu^2 + 2Fdudv + Gdv^2}$$

The **metric tensor** of the surface is:

$$\mathcal{F}_1 = \begin{pmatrix} E & F \\ F & G \end{pmatrix}$$

The first fundamental form is invariant to parametrization.

4.5 Second Fundamental Form

Let the standard unit normal be:

$$\hat{m{n}} = rac{m{\sigma}_u imes m{\sigma}_v}{||m{\sigma}_u imes m{\sigma}_v||} = rac{m{\sigma}_u imes m{\sigma}_v}{\sqrt{EG - F^2}}$$

The second fundamental form has elements:

$$L = \boldsymbol{\sigma}_{uu} \cdot \hat{\boldsymbol{\sigma}} \quad M = \boldsymbol{\sigma}_{uv} \cdot \hat{\boldsymbol{\sigma}} \quad N = \boldsymbol{\sigma}_{vv} \cdot \hat{\boldsymbol{\sigma}}$$

And takes the form:

$$Ldu^2 + 2Mdudv + Ndv^2$$

4.6 Surface Area

The surface area is:

$$\int \int_{U} \sqrt{EG - F^2} du dv$$

4.7 Darboux Frame

If we look at a unit-speed curve γ , then we have a frame called the **Darboux frame** defined by $\dot{\gamma}$, \hat{n} , and $\hat{n} \times \dot{\gamma}$. We let these be T, U, and V respectively. Then, we have:

$$\begin{pmatrix} T' \\ U' \\ V' \end{pmatrix} = \begin{pmatrix} 0 & \kappa_g & \kappa_n \\ -\kappa_g & 0 & \tau_g \\ -\kappa_n & -\tau_g & 0 \end{pmatrix} \begin{pmatrix} T & V & U \end{pmatrix}$$

5 Geodesics

A curve is called a **geodesic** if the acceleration is either zero or parallel to its unit normal \hat{n} .

Proposition 1 Every line can be reparametrized as a geodesic. Thus, a geodesic $\gamma(t)$ on a surface S has constant speed.

Thus, unit-speed parametrizations of geodesics are also geodesics.

Proposition 2 A curve on a surface is a geodesic if and only if its geodesic curvature is zero everywhere

Theorem 3 A curve γ on a surface S is a geodesic if and only if for any part $\gamma(t) = \sigma(u(t), v(t))$ contained in a surface patch σ of S, the following two equations are satisfied:

$$\frac{d}{dt} (E\dot{u} + F\dot{v}) = \frac{1}{2} \left(E_u \dot{u}^2 + 2F_u \dot{u}\dot{v} + G_u \dot{v}^2 \right)$$
$$\frac{d}{dt} (F\dot{u} + G\dot{v}) = \frac{1}{2} \left(E_v \dot{u}^2 + 2F_v \dot{u}\dot{v} + G_v \dot{v}^2 \right)$$

where $Edu^2 + 2Fdudv + Gdv^2$ is the first fundamental form of σ .

These two equations are called the **geodesic equations**. They are non-linear and solvable analytically on rare occasions only.

5.1 Isometry

An **isometry** is a **diffeomorphism** that maps curves from one surface to curves of the same length in another surface. A **diffeomorphism** is a mapping from one surface to another that is smooth, bijective, and has an inverse map that is also smooth. The necessary conditions for f to be an isometry is any surface patch σ_1 of S_1 and its image patch of $f \circ \sigma_1$ of S_2 have the same first fundamental form.

Theorem 5 An isometry maps the geodesics of S_1 to the geodesics of S_2 .

5.2 Shortest Paths

Theorem 6 Let γ be a shortest path on a surface S connecting two points p and q. Then the part of γ contained in any surface patch σ must be a geodesic.

The converse is not true. Not all geodesics are shortest paths. Additionally, a shortest path may not exist.

5.3 Geodesic Coordinates

We can use geodesics of a surface to construct an atlas for the surface. For any given unit-speed geodesic $\gamma(v)$ that passes through a point p in S at v=0, then for any value of v, there exists a unique unit-speed geodesic $\beta_v(u)$ such that $\beta_v(0) = \gamma(v)$ and $\beta'_v(0) \perp \gamma'(v)$. We define a patch $\sigma(u,v) = \beta_v(u)$ as a **geodesic patch**, and we call u and v **geodesic coordinates**.

6 Nonlinear Optimization

Given functions that depend on one or more independent variable, we want to maximize the function. In general, finding global extremum is very difficult. The popular approaches are: i) find local extrema starting from widely varying values of the independent variables, and then pick the most extreme of these; ii) perturb a local extremum by taking a finite amplitude step away from it, and then see if your routine can get to a better point, or "always" to the same one.

6.1 Golden Section Search

The **golden section search** is similar to bisection for root finding. The basic idea is to bracket a minimum and then shrink the bracket. We can not know if we have bracketed a minimum, but if we settle for a local minimum, then we can just bracket any zero of the derivative.

$$f\left(b\right) < \min\left\{f\left(a\right), f\left(c\right)\right\}$$

Now, we choose a point x, say, halfway between a and b. Then, we if f(x) > f(b), then we bracket [x, b, c]. Otherwise, we bracket [a, x, b]. To ensure that the interval [a, c] shrinks towards a point, we alternate the bracketing tuples.

6.2 Gradient and Hessian

The **gradient** is defined as:

$$\nabla f(x) = \left(\frac{\partial f}{\partial x_1}, \dots, \frac{\partial f}{\partial x_n}\right)$$

This gives the direction in which the value of f increases the fastest. The **Hessian H** of f is defined as an $n \times n$ matrix:

$$H\left(\boldsymbol{x}\right) = \begin{pmatrix} \frac{\partial^{2} f}{\partial x_{1}^{2}} & \cdots & \frac{\partial^{2} f}{\partial x_{1} \partial x_{n}} \\ \vdots & \ddots & \vdots \\ \frac{\partial^{2} f}{\partial x_{n} \partial x_{1}} & \cdots & \frac{\partial^{2} f}{\partial x_{n}^{2}} \end{pmatrix}$$

If x^* is a **relative minimum**, then the following conditions hold:

- $\nabla f(\boldsymbol{x}^*) = 0$
- $d^T H(x^*) x$ for every $d \in \mathbb{R}^n$

The Hessian at a relative minimum is $\mathbf{symmetric}$ positive $\mathbf{semi-definite}$, so:

$$\boldsymbol{x}^T H(\boldsymbol{x}^*) \boldsymbol{x} \geq 0 \text{ for any } \boldsymbol{x}.$$

If x^* is a strict relative minimum, then the Hessian is positive definite, so:

$$\boldsymbol{x}^T H\left(\boldsymbol{x}^*\right) \boldsymbol{x} > 0 \text{ for any } \boldsymbol{x} \neq 0$$

6.3 Convex Functions

Taylor's theorem shows that every function looks like a quadratic near a strict local minimum. **Convex functions** are functions whose local minima are global minima.

Proposition 1 /Let $f \in C^2$. Then f is convex over a convex set Ω containing an interior point if and only if the Hessian matrix H is positive semi-definite in Ω .

Theorem 2 Let f be a convex function defined on a convex set Ω . Then the set Γ where f achieves its minimum value is convex. Furthermore, any relative minimum is a global minimum.

6.4 Steepest Descent

In one dimension, we apply the rules:

if f'(x) < 0 then move to the right; if f'(x) > 0 then move to the left; if f'(x) = 0 then stop.

In higher dimensions, we use $-\nabla f$ to point us toward a minimum. This method is called **steepest descent** because we are descending at the steepest rate.

6.5 Jacobian

The *Jacobian* of a function is the derivative of a vector function $f(x) = (f_1(x), ..., f_m(x))$ with respect to x which results in an $m \times n$ matrix.

6.6 Lagrange Multipliers

To solve constrained problems, we introduce the Lagrangian defined as:

$$\ell(x, y, \lambda) = f(x, y) + \lambda h(x, y)$$

Then, we have:

$$\nabla \ell = \begin{pmatrix} \frac{\partial f}{\partial x} + \lambda \frac{\partial h}{\partial x} \\ \frac{\partial f}{\partial y} + \lambda \frac{\partial h}{\partial y} \\ h \end{pmatrix}^{T} = (\nabla f + \lambda \nabla h, h)$$

 λ is known as the **Lagrangian multiplier**. We set the gradient of the Lagrangian to zero $(\nabla \ell = 0)$ to yield a system of nonlinear equations. Solving these yields the stationary points (x, y, λ) , and thus, the minimum and maximum points.

Theorem 3 (First – OrderNecessaryConditions) Let \mathbf{x}^* be a local extremum point of f subject to the constraints $\mathbf{h}(\mathbf{x}) = \mathbf{0}$. Assume further that \mathbf{x}^* is a regular point of these constraints. Then there is a $\lambda \in \mathbb{R}^m$ such that

$$\nabla f\left(\boldsymbol{x}^{*}\right) + \boldsymbol{\lambda}^{T} \nabla h\left(\boldsymbol{x}^{*}\right) = \mathbf{0}$$

Theorem 4 (Second – OrderNecessaryConditions) Suppose that \mathbf{x}^* is a local minimum of f subject to $\mathbf{h}(\mathbf{x}) = \mathbf{0}$ and that \mathbf{x}^* is a regular point of these constraints. Denote by F the Hessian of f and by H_i the Hessian of h_i for $1 \leq i \leq m$. Then there is a $\lambda \in \mathbb{R}^m$ such that

$$\nabla f(\boldsymbol{x}^*) + \boldsymbol{\lambda}^T \nabla h(\boldsymbol{x}^*) = \mathbf{0}$$

The matrix

$$L(\boldsymbol{x}^*) = F(\boldsymbol{x}^*) = \sum_{i=1}^{m} \lambda_i H_i(\boldsymbol{x}^*)$$
(1)

is positive semi-definite on the tangent space $\{y \mid \nabla h(x^*) y = 0\}$.

Theorem 5 (Second – OrderSufficientConditions) Suppose there is a point \mathbf{x}^* satisfying $\mathbf{h}(\mathbf{x}^*) = \mathbf{0}$, and a λ such that

$$\nabla f(\boldsymbol{x}^*) + \boldsymbol{\lambda}^T \nabla \boldsymbol{h}(\boldsymbol{x}^*) = \boldsymbol{0}$$

Suppose also that the matrix $L(\mathbf{x}^*)$ defined in (1) is positive definite on the tangent space. Then \mathbf{x}^* is a strict local minimum of f subject to $\mathbf{h}(\mathbf{x}) = \mathbf{0}$

7 Data Fitting

7.1 Basis Functions

The **basis functions** are the functions that make up the variable terms of k degree for an algebraic function of the form:

$$\sum_{i=0}^{n} a_i b_i \left(x \right)$$

Basis functions are not necessarily, for example, $1, x, x^2$. They are any function that spans the vector space.

7.2 Least-Squares Fitting

Using a linear combination of the chosen basis functions, we can achieve an approximation of the original function through several norms to measure the error. Most often, we use the **least-squares fitting** that is defined by the norm:

2-norm:
$$||f - F||_2 = \sqrt{\sum_{i=1}^{n} |f_i - F(x_i)|^2}$$

To solve for the least-squares fitting, we can use two methods: coefficient matrices or SVD.

The **coefficient matrix** method is sets Φ to be the matrix containing the values of the \$i\$th basis function in the \$i\$th column of n size where n is the number of points to be related. We then take $\mathbf{c} = (\Phi^T \Phi)^{-1} \Phi^T \mathbf{f}$.

The SVD method solves:

$$\Phi c = f$$

- 8 Orthogonal Polynomials
- 9 Fourier Series
- 10 Calculus of Variations

11 Probability

Probability of one or more events A given a set of possible outcomes S is defined by Laplace as:

$$Pr\left(A\right) = \frac{|A|}{|S|}$$

We also have conditional probability which is defined as the probability of event A given event B where Pr(B) > 0.

$$Pr(A \mid B) = \frac{Pr(A \cap B)}{Pr(B)}$$

We call the probability of A the **a priori** probability. The conditional probability of A given B is called the **a posteriori** probability.

11.1 Random Variables

Random variables are functions from the sample space S of an experiment to the set of real numbers. It simply assigns a real number to each possible outcome.

Random variables can be either **discrete** or **continuous**.

11.2 Cumulative Distribution Function

Given a random variable X, its **cumulative distribution function (CDF)** is defined as:

$$D\left(x\right) = Pr\left(X \le x\right)$$

If we have a **discrete** probability Pr(x), that is the probability that a discrete random variable X has value x, we have:

$$D\left(x\right) = \sum_{X \le x} Pr\left(X\right)$$

1. Properties

$$D(x) \in [0, 1]$$

$$D(-\infty) = 0$$

$$D(\infty) = 1$$

$$D(a) \le D(b) \text{ if } a < b$$

$$Pr(a < X \le b) = D(b) - D(a)$$

11.3 Probability Density Function

The **probability density function (PDF)** P(x) of a continuous random variable is defined as the derivative of the cumulative distribution function D(x).

$$P(x) = \frac{d}{dx}D(x)$$
$$D(x) = \int_{-\infty}^{x} P(\xi) d\xi$$

1. Properties

$$P(x) \ge 0$$

$$\int_{-\infty}^{\infty} P(x) dx = 1$$

$$D(a < x \le b) = \int_{a}^{b} P(x) dx$$

A uniform distribution has constant PDF. The PDF and CDF on the interval [a, b] are:

$$P(x) = \begin{cases} 0 & \text{for } x < a, \\ \frac{1}{b-a} & \text{for } a \le x \le b, \\ 0 & \text{for } x > b; \end{cases}$$
$$D(x) = \begin{cases} 0 & \text{for } x < a, \\ \frac{x-a}{b-a} & \text{for } a \le x \le b, \\ 1 & \text{for } x > b; \end{cases}$$

11.4 Generating Continuous Distributions

Built-in random functions like C++'s rand generate uniformly distributed pseudo-random numbers. If we want to generate uniform random points for a different distribution, say for a circle of radius ρ , then we can do so using CDF.

 $D\left(x\right)$ increases monotonically from zero to one. If we suppose $D\left(x\right)$ is continuous and strictly increasing, then there exists an inverse $D^{-1}\left(y\right)$ such that, for 0 < y < 1,

$$y = D(x)$$
 if and only if $x = D^{-1}(y)$

We can compute a random variable X with distribution D(x) by setting:

$$X = D^{-1}(Y)$$

where Y is a random variable with uniform distribution over [0,1]. The reasoning is:

$$Pr(X \le x) = Pr(D^{-1}(Y) \le x)$$
$$= Pr(Y \le D(x))$$
$$= D(x)$$

Example: Uniformly Distributed Random Points

Suppose we want a uniform distribution of random points in a rectangle with dimensions (l, w). We can do it by using built in pseudo-random functions such as C++'s rand to get a number in the interval [0,65535]. We can normalize this to the interval [0,1]. Now, we can plot a point (X,Y) where X and Y are uniformly distributed within the intervals [0,l] and [0,w] respectively. Simply multiply the random output by l and w after normalizing.

Now, suppose we want a uniform distribution of random points in a **circle** of radius ρ . If we simply plot a point (R, Θ) where R and Θ are uniformly distributed within the intervals $[0, \rho]$ and $[0, 2\pi]$ respectively. We will find that the distribution is more densely populated near the origin. Hence, we need to find a distribution for R. Consider the CDF:

$$D(r) = Pr(R \le r) = \frac{\pi r^2}{\pi \rho^2} = \frac{r^2}{\rho^2}$$

Clearly, $D(r) \in [0, 1]$. So, if we let s = D(r), we let $r = \rho \sqrt{s}$. Introduce a random variable S with uniform distribution over [0, 1]. So, $R = \rho \sqrt{S}$.

Thus, random points should be generated at $\rho\sqrt{S}$ (cos Θ , sin Θ), where S and Θ are random variables with uniform distributions over [0, 1] and $[0, 2\pi]$.

11.5 Mean (μ)

The **mean** or **expected value**, of a random variable X is its average value over a large number of experiments (say, N experiments) with outcomes x_i that occurs n_i times for $i \in \{1, ..., m\}$.

$$E\left(X\right) = \frac{1}{N} \sum_{i=1}^{m} x_i n_i$$

Example: Infinite Dice Rolls

Suppose we roll a die an infinite number of times. Each number appears $\frac{1}{6}$ of the time. So, expected value is:

$$E(X) = \lim_{n \to \infty} \frac{1}{n} \sum_{i=1}^{6} i \cdot \frac{n}{6}$$
$$= \frac{7}{2}$$

If the random variable X with PDF $P\left(X\right)$ is continuous, then the expected is:

$$E\left(X\right) = \int_{-\infty}^{\infty} x P\left(x\right) dx$$

 \mathbf{t}

11.6 Standard Deviation (σ)

The **variance** measure the dispersion of the values about the mean. We can compute the variance of a discrete random variable with N possible values x_1, \ldots, x_N and mean μ :

$$var(X) = \sum_{i=1}^{N} Pr(x_i) (x_i - \mu)^2$$

 $var\left(X\right) =\sigma^{2},$ and we call σ the **standard deviation**. For a continuous distribution:

$$var(X) = \int_{-\infty}^{\infty} P(X) (x - \mu)^{2} dx$$

So, we can rewrite variance as:

$$\sigma^{2} = E((X - \mu)(X - \mu))$$

$$= E(X^{2} - 2X\mu + \mu^{2})$$

$$= E(X^{2}) - 2\mu^{2} + \mu^{2}$$

$$= E(X^{2}) - \mu^{2}$$

11.7 Gaussian Distribution

A random variable X with mean μ and variance σ^2 has **Gaussian distribution** or **normal distribution** if its probability density function is given by:

$$P(x) = \frac{1}{\sigma\sqrt{2\pi}}e^{-(x-\mu)^2/\left(2\sigma^2\right)}$$

We denote the Gaussian distribution as $N\left(\mu,\sigma^2\right)$.

- 11.8 Random Variables
- 11.9 Covariant (Matrix)
- 11.10 Correlation (Matrix)