

A note on Schultz

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1 These notes

Feynman's 1955 paper 'Slow electrons in polar crystals'[1] introduces his path-integral solution to the polaron problem. This solves the Fröhlich Hamiltonian, which considers effective-mass electrons interacting via polar vibrations induced in the lattice.

Schultz[2] received his thesis, "Electron-Lattice Interactions in Polar Crystals" in 1956. A two-year postdoctoral position in Birmingham working with Rudolf Peierls led to his 1959 Physical Review, "Slow Electrons in Polar Crystals: Self-Energy, Mass, and Mobility"[3]. These notes are chiefly concerned with the early part of this paper, where the Feynman theory is 'fleshed out' with more experimental observables.

Implementing numerical methods is far easier today than it would have been for Schultz (or rather, for Hannah Landsman acknowledged 'for most of the numerical work') with the 'Whirlwind Digital Computer'[4], a very early vacuum tube digital computer built at MIT. These notes were developed having read the original literature, in order to develop a modern set of codes to implement the methods, and so enable calculation of polaron states for arbitrary materials. The codes are developed in the Julia language[5]. This language is high level and mathematically expressive, as such these notes are an adjunct

to reading the program source codes directly. These are available online at <https://github.com/jarvist/PolaronMobility.jl>.

2 Schultz's polaron radii

Schultz1959[3] defines the Feynman polaron radius as the standard deviation of the Gaussian Wavefunction. This Gaussian is fully specified by the Simple Harmonic Oscillator parameters for the polaron system, which in turn are provided by Feynman's variational solution (Feynman1955[1]) to the Fröhlich polaron problem.

The reduced mass is defined in terms of the internal polaron parameters v and w ,

$$\mu = m(v^2 - w^2)/v^2 \quad (1)$$

Here m is the band effective-mass of the electron (hole).

The (Gaussian) wavefunction is fully specified (plotted above),

$$\psi(r) = (\mu v/\pi)^{\frac{3}{4}} \exp(-\frac{\mu v r^2}{2}). \quad (2)$$

And from the standard deviation of this wavefunction, a radius is defined:

$$r_f \equiv (\langle \rho^2 \rangle)^{\frac{1}{2}} = (3/2\mu v)^{\frac{1}{2}} \quad (2.4, \text{Schultz}) \quad (3)$$

Here ρ is the density of the wavefunction, μ is the reduced (effective) mass of the electron and interacting phonon-cloud, while v and w are internal polaron parameters characterising the harmonic motion of the polaron. The units of v and w are $\hbar\omega$.

Schultz: 'Using the weak- α coupling expansions given by Feynman for w and v , we find'

$$r_f(\alpha \rightarrow 0) \approx (3/0.44\alpha)^{\frac{1}{2}}(2m\omega)^{-\frac{1}{2}} \quad (4)$$

To understand where this magic number (0.44) comes from, one must return to Feynman1955 and repeat the derivation used to get the small- α expansion for energy. This is the part around Eqn 35 to 36 in Feynman 1955.

Finding the term $2\mu v$ is our aim.

Following Feynman, assuming α is small, $v = (1 + \epsilon)w$. You can then go back to the energy integral Feynman(31), substituting your v expansion in. You can expand the square-root on on the denominator, by pulling out terms of w to simplify, before taking a term of $\tau^{\frac{1}{2}}$ out to place the root in the form of $(1 + x)^n$. You can then use a binomial expansion for this square root, and keep only zeroth and first order in ϵ terms. The resulting integral in the total energy has two parts, the zeroth-order contribution integrating to 1, and the linear in ϵ term (P) apparently analytic and provided by Feynman (35),

$$2w^{-1}[(1 + w)^{\frac{1}{2}} - 1] = P. \quad (5)$$

This makes the total energy,

$$E = \frac{3}{4v}(v - w)^2 - A, \quad (\text{Feynman33}) \quad (6)$$

which with our linear-in- ϵ form of the integral A ,

$$E = \frac{3}{4v}(v - w)^2 - \alpha \frac{v}{w}[1 - P] \quad (7)$$

Again, substituting $v = (1 + \epsilon)w$,

$$E = \frac{3}{4} \frac{((1 + \epsilon)w - w)^2}{(1 + \epsilon)w} - \alpha \frac{(1 + \epsilon)w}{w}[1 - P], \quad (8)$$

and discarding high order (in ϵ) terms,

$$E = \frac{3}{4}w\epsilon^2 - \alpha - \alpha\epsilon[1 - P], \quad (9)$$

as given by Feynman.

If we reorder terms,

$$E = \frac{3}{4}w\epsilon^2 - \alpha\epsilon[1 - P] - (\alpha), \quad (10)$$

it becomes clear the total energy has a quadratic ($E = a\epsilon^2 + b\epsilon + c$) form in ϵ .

We will want a variational (minimum energy) solution. This found at the minimum, $x = \frac{-b}{2a}$ in the standard quadratic formula, and so,

$$\epsilon = \frac{2}{3}\alpha \frac{(1 - P)}{w}. \quad (11)$$

This is as given by Feynman, valid for small- α only, due to the various occasions where ϵ is taken as small.

We can now return to the key $2\mu v$ term.

Expanding this identity with our definition of μ ,

$$2\mu v = 2m_e \frac{v^2 - w^2}{v}. \quad (12)$$

Substituting in $v = (1 + \epsilon)w$, following through the algebra, and as epsilon is small approximating $\frac{1}{1 + \epsilon} = \frac{1}{1}$, we eventually have

$$2\mu v = 2m_e[2w\epsilon]. \quad (13)$$

Substituting in our form for ϵ , by simple algebraic rearrangement we get

$$v = (1 + \frac{2\alpha(1-P)}{3w})w = w + \frac{2}{3}\alpha(1-P) \quad (14)$$

Feynman states that ‘the variational solution is least for $w = 3$ ’, and goes on to state that the energy correction is anyway insensitive to the choice of w .

If $w = 3$, $P = \frac{2}{3}$, and we have that,

$$v = 3 + \frac{2}{9}\alpha. \quad (15)$$

Thus, $2\mu v = 2m_e \frac{4}{9}\alpha = 2m_e 0.4\alpha$.

Schultz1959 Eqn. 2.5a approximates this as

$$2\mu v = 2m_e 0.44. \quad (16)$$

It is not obvious whether this arises as a printing error, or was an intentional (1% error) approximation.

References

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