

# Notes on Physics from Symmetry

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This document contains my personal notes on Jakob Schwichtenberg's Physics from Symmetry ([Schwichtenberg, 2015](#)), with a sprinkling of notes from my undergraduate physics course in quantum field theory (and, to a lesser extent, general relativity).

## 1. Special relativity

### 1.1. Definitions and postulates

In special relativity, **inertial frames of reference** are coordinate systems moving with constant velocity relative to each other. Special relativity has two basic postulates:

1. **The principal of relativity:** The laws of physics are the same in all inertial frames of reference.
2. **The invariance of the speed of light:** The velocity of light has the same value  $c$  in all inertial frames of reference.

**Theorem 1.1 (Invariant of special relativity).** *Consider two events  $A$  and  $B$  in an inertial observer  $O$ 's frame of reference. Let the time interval measured by  $O$  between the two events be  $(\Delta t)$ , and the three spatial intervals be  $(\Delta x)$ ,  $(\Delta y)$ ,  $(\Delta z)$ . Then, the quantity*

$$(\Delta s)^2 := (\Delta ct)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta z)^2 \quad (1.1)$$

*is invariant between all frames of reference. I.e.*

$$(\Delta s')^2 = (\Delta s)^2 \quad (1.2)$$

*for any inertial frame of reference  $O'$ .*

Theorem 1.1 follows directly from the invariance of the speed of light (consider a pair of mirrors, for two observers with relative velocity).

**Definition 1.1 (Proper time).** *Proper time,  $\tau$ , is the time measured by an observer in the special frame of reference where the object in question is at rest. In this frame of reference,*

$$(\Delta s)^2 = (c\Delta\tau)^2. \quad (1.3)$$

*In the infinitesimal limit*

$$(ds)^2 = (cd\tau)^2. \quad (1.4)$$

Physically, Defn. 1.1 means that all observers agree on the time interval between events for an observer who travels with the object in question. However, different observers **do not** in general agree on the time interval between events generally:  $(\Delta t) \neq (\Delta t')$  – this is called **time dilation**.

### 1.2. $c$ is an upper speed limit

All observers agree on the value of  $(ds)^2 = (cd\tau)^2$ . Furthermore, we commonly assume that there exists a minimal proper time of  $\tau = 0$  for two events if  $\Delta s^2 = 0$ . We can therefore write that when  $\tau = 0$

$$c^2 = \frac{(dx)^2 + (dy)^2 + (dz)^2}{(dt)^2} \quad (1.5)$$

between two events with an infinitesimal distance. We can equate the right-hand side with a squared velocity, and hence

$$\tau = 0 \implies c^2 = v^2 \quad (1.6)$$

so

$$(ds)^2 \geq 0 \implies c^2 \geq v^2 \quad (1.7)$$

for **any** pair of events (which are causally connected, although how this follows is not immediately clear to me right now).

### 1.3. Tensor notation and Minkowski spacetime

**Definition 1.2 (Four-vector (contravariant)).** A position four-vector is defined as

$$x^\mu = \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix} \equiv \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}. \quad (1.8)$$

**Definition 1.3 (Minkowski metric).** The Minkowski metric is defined as

$$\eta_{\mu\nu} = \text{diag}(1, -1, -1, -1). \quad (1.9)$$

$\eta$  is used to compute distances and lengths in Minkowski space.

We define  $\eta^{\mu\nu}$  through the relation

$$\eta^{\mu\nu}\eta_{\nu\sigma} = \delta^\mu_\sigma \quad (1.10)$$

where we have applied the **Einstein summation convention**, where a repeated Greek index implies a summation from 0 to 3 (where the zeroth index is time), and a repeated Roman index is summed from 1 to 3. Hence, for a matrix multiplication between two  $3 \times 3$  matrices  $A$  and  $B$ ,  $(AB)_{ij} = A_{ik}B_{kj}$ , and  $(A^T)_{ij} = A_{ji}$ .

**Definition 1.4 (One-form (covariant vector)).** We define a one-form as

$$x_\mu = \eta_{\mu\nu}x^\nu. \quad (1.11)$$

Thus,

$$ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu. \quad (1.12)$$

**Definition 1.5 (Scalar product).** A scalar product between four-vectors  $x$  and  $y$  is defined as

$$x \cdot y := x^\mu y^\nu \eta_{\mu\nu} = x_\mu y_\nu \eta^{\mu\nu} = x^\mu y_\mu = x_\nu y^\nu \quad (1.13)$$

due to the symmetry of the metric:  $\eta_{\mu\nu} = \eta_{\nu\mu}$ .

**Ordering (spacing) of indicies** In order to be able to freely raise/lower indicies (without repeatedly writing the metric tensor), we can impose an ordering upon indicies of tensor fields – which we can represent typographically with spacing between tensor indicies. A metric  $g_{ij}$  (or  $g^{ij}$ ) has the effect of lowering (or raising) a repeated index. For example,

$$g_{iq} T^{abcd}{}_{efgh}{}^{ijkl}{}_{mnop} = T^{abcd}{}_{efghq}{}^{ijkl}{}_{mnop}. \quad (1.14)$$

(Proof of this, I imagine, requires background in differential geometry?)

## 1.4. Lorentz transformations

From the invariant of SR (Theorem 1.1), we have

$$ds'^2 = dx'_\mu dx'_\nu \eta^{\mu\nu} = dx_\mu dx_\nu \eta^{\mu\nu} \quad (1.15)$$

for all reference frames. We denote  $\Lambda$  as a  $(1,1)$  tensor field, which transforms a four-vector from one reference frame to another:

$$dx'^\mu = \Lambda^\mu{}_\nu dx^\nu \quad (1.16)$$

which leaves the  $ds^2$  invariant, i.e.  $ds'^2 = ds^2$ . It follows that

$$\begin{aligned} \eta_{\mu\nu} &= \Lambda^\sigma{}_\mu \Lambda^\delta{}_\nu \eta_{\sigma\delta} \\ \eta &= \Lambda^T \eta \Lambda. \end{aligned} \quad (1.17)$$

The physical meaning of Eq.(1.17) is that Lorentz transformations leave the scalar product of Minkowski spacetime invariant: i.e. changes between frames of reference that respect the two postulates of special relativity (Section 1.1). Conservation of the scalar product is analogous to rigid rotation ( $O$ ) in Euclidean space ( $a \cdot b = a' \cdot b' = a^T O^T O b \implies O^T O = I$ ), which preserves orientation ( $\det(O) = 1$ ).

Note that  $\Lambda^\mu{}_\nu \neq \Lambda_\nu{}^\mu$ . Beginning with Eq.(1.17),

$$\Lambda^\mu{}_\rho \Lambda^\nu{}_\sigma \eta_{\mu\nu} = \eta_{\rho\sigma}$$

we can raise one index, and lower one index, of  $\Lambda^\nu{}_\sigma$

$$\begin{aligned} \Lambda^\mu{}_\rho \eta_{\mu\nu} \Lambda^\nu{}_\sigma \eta^{\sigma\lambda} &= \eta_{\rho\sigma} \eta^{\sigma\lambda} \\ \Lambda^\mu{}_\rho \Lambda_\mu{}^\lambda \eta_{\mu\nu} &= \eta_{\mu\nu} \delta_\rho{}^\lambda \\ \Lambda^\mu{}_\rho \Lambda_\mu{}^\lambda &= \delta_\rho{}^\lambda \end{aligned} \quad (1.18)$$

so we see that  $\Lambda_\nu{}^\mu$  is the inverse of  $\Lambda^\mu{}_\nu$ .

## 2. Lie group theory

### 2.1. Invariance, symmetry, and covariance

We call a quantity **invariant** if it does not change under particular transformations. E.g. if we transform  $A, B, C, \dots \rightarrow A', B', C', \dots$  and we have

$$F(A', B', C', \dots) = F(A, B, C, \dots) \quad (2.1)$$

then we say  $F$  is invariant under this transformation. **Symmetry** is defined as invariance under a transformation (or class of transformations). An equation is covariant if it takes the same form when objects in it are transformed. *All physical laws must be covariant under Lorentz transformations.*

Group theory describes the properties of particular sets of transformations: the invariances under such groups allows us to mathematically describe symmetry. For example, the set of rotations about the origin of a square by  $n\pi/2$  form a **discrete group**, and leave the set of points which constitute the square invariant under the transformation. The set of rotations about the origin of a circle form a **continuous group**. We can use group theory to work with *all* kinds of symmetries: symmetries which operate on vectors, equations, ...

## 2.2. Groups

**Definition 2.1 (Group axioms).** A group  $(G, \circ)$  is a set  $G$ , together with a binary operation  $\circ$  defined on  $G$ , that satisfies the following axioms

- *Closure:* For all  $g_1, g_2 \in G$ ,  $g_1 \circ g_2 \in G$
- *Identity element:* There exists an identity element  $e \in G$  such that for all  $g \in G$ ,  $g \circ e = g = e \circ g$
- *Inverse element:* For each  $g \in G$ , there exists an inverse element  $g^{-1} \in G$  such that  $g \circ g^{-1} = e = g^{-1} \circ g$ .
- *Associativity:* For all  $g_1, g_2, g_3 \in G$ ,  $g_1 \circ (g_2 \circ g_3) = (g_1 \circ g_2) \circ g_3$

The set of all transformations that leave a given object invariant is called a **symmetry group**. For Minkowski spacetime, the object that is left invariant is the Minkowski metric, and the corresponding symmetry group is called the **Poincaré group**. Notice that the transformations which constitute a group are defined entirely independently from the object on which the transformations act.

### 2.2.1. Rotations in two dimensions and $SO(2)$

Consider the 2D rotation matrix

$$R_\theta = \begin{pmatrix} \cos(\theta) & -\sin(\theta) \\ \sin(\theta) & \cos(\theta) \end{pmatrix} \quad (2.2)$$

and the two reflection matrices

$$P_x = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} \quad P_y = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \quad (2.3)$$

These matrices satisfy the group axioms. We can uncover this group from a symmetry perspective. The above transformations leave the length of a vector unchanged, i.e.

$$a \cdot a = a' \cdot a'. \quad (2.4)$$

Letting the transformation be represented by  $a' = Oa$ , it follows that all members of the group must satisfy

$$O^T O = I. \quad (2.5)$$

This condition defines the group  $O(2)$ , which is the group of all **orthogonal**  $2 \times 2$  matrices. It follows that  $\det(O) = \pm 1$  – i.e. the transformations are area-preserving. The subgroup with  $\det(O) = 1$  is called  $SO(2)$ , which corresponds to rigid rotations preserving the orientation of the system – “S” denoting **special**.

### 2.2.2. Rotations with unit complex numbers and $U(1)$

A unit complex number is a complex number  $z$  which satisfies  $|z|^2 = z^* z = 1$ . The group  $U(1)$  is the set of unit complex numbers, together with ordinary complex number multiplication. The  $U$  stands for ‘**unitary**’, which generally stands for the condition

$$U^\dagger U = 1, \quad (2.6)$$

where  $U^\dagger = (U^T)^*$  is the **Hermitian conjugate** of  $U$ . For scalars, the Hermitian conjugate is equivalent to the complex conjugate. Note that a unit complex number can also be denoted as

$$R_\theta = e^{i\theta} = \cos(\theta) + i \sin(\theta) \quad (2.7)$$

which makes the interpretation of  $U(1)$  as rotations on the unit complex numbers evident.

We can connect this description of rotations ( $U(1)$ ) to the previous ( $SO(2)$ ) by defining

$$1 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \quad i = \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}. \quad (2.8)$$

For an arbitrary unit complex number  $z = a + ib$ , let

$$f(z) = a \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} + b \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} = \begin{pmatrix} a & -b \\ b & a \end{pmatrix}. \quad (2.9)$$

Since  $z = R_\theta = \cos(\theta) + i \sin(\theta)$ , we can plug in the real and imaginary components of  $z$  into Eq.(2.9) to arrive at Eq.(2.2). We then have  $z' = R_\theta z$ , to perform rotations. There therefore exists an **isomorphism** between  $SO(2)$  and  $U(1)$ :

**Definition 2.2 (Group isomorphism).** *Given two groups  $(G, *)$ ,  $(H, \odot)$ , a group isomorphism is a bijective function  $f : G \rightarrow H$  such that*

$$f(u * v) = f(u) \odot f(v) \quad \forall u, v \in G. \quad (2.10)$$

which is written as

$$(G, *) \cong (H, \odot). \quad (2.11)$$

$f(z)$  in Eq.(2.9) is therefore a group isomorphism between  $U(1)$  and  $SO(2)$ . This realization has an analogue in three dimensions, which will reveal something fundamental about nature.

### 2.2.3. Rotations in three dimensions

Rotations in three dimensions can be described by the following “basis rotations”

$$\begin{aligned} R_x &= \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos(\theta) & -\sin(\theta) \\ 0 & \sin(\theta) & \cos(\theta) \end{pmatrix} & R_y &= \begin{pmatrix} \cos(\theta) & 0 & \sin(\theta) \\ 0 & 1 & 0 \\ -\sin(\theta) & 0 & \cos(\theta) \end{pmatrix} \\ R_z &= \begin{pmatrix} \cos(\theta) & -\sin(\theta) & 0 \\ \sin(\theta) & \cos(\theta) & 0 \\ 0 & 0 & 1 \end{pmatrix}. \end{aligned} \quad (2.12)$$

So, to rotate a vector  $v$  around the  $z$ -axis by  $\theta$ , we would compute  $R_z(\theta)v$ .

### 2.2.4. Quaternions

To get a second description of rotations in three dimensions, we must generalize complex numbers in higher dimensions. Astonishingly, it turns out that there are no 3-dimensional complex numbers. Instead, we can find 4-dimensional complex numbers called quaternions, which will turn out to be able to describe rotations in 3-dimensions. The fact that quaternions are 4-dimensional will reveal something deep about the universe. We could have anticipated this result, because we will be using unit quaternions, which have 3 degrees of freedom.

To construct quaternions, we introduce three complex units satisfying the relations

$$\mathbf{i}^2 = \mathbf{j}^2 = \mathbf{k}^2 = -1 \quad (2.13)$$

$$\mathbf{ijk} = -1 \quad (2.14)$$

$$q = a\mathbf{1} + b\mathbf{i} + c\mathbf{j} + d\mathbf{k}. \quad (2.15)$$

×	1	i	j	k
1	1	i	j	k
i	i	-1	k	-j
j	j	-k	-1	i
k	k	j	-i	-1

**Figure 1.** Quaternion multiplication table, read as row  $\times$  column = value. E.g.  $\mathbf{j}\mathbf{i} = -\mathbf{k}$ . In general, the basic quaternions anti-commute.

All other relations can be computed from the above. For example, the relation  $\mathbf{ij} = \mathbf{k}$  can be derived by multiplying both sides of Eq.(2.15) by  $\mathbf{k}$ . Notice that it follows that the **basic quaternions anticommute** with each other, see Fig. 1.

The set of unit quaternions satisfy

$$q^\dagger q = 1 \tag{2.16}$$

$$\implies a^2 + b^2 + c^2 + d^2 = 1. \tag{2.17}$$

## References

Schwichtenberg, J., 2015 *Physics from symmetry*. Springer.