# Notes on Physics from Symmetry

Author: Juvid Aryaman Last compiled: May 21, 2021

This document contains my personal notes on Jakob Schwichtenberg's Physics from Symmetry (Schwichtenberg, 2015), with a sprinkling of notes from my undergraduate physics course in quantum field theory (and, to a lesser extent, general relativity).

### 1. Special relativity

#### 1.1. Definitions and postulates

In special relativity, **inertial frames of reference** are coordinate systems moving with constant velocity relative to each other. Special relativity has two basic postulates:

- 1. The principal of relativity: The laws of physics are the same in all inertial frames of reference.
- 2. The invariance of the speed of light: The velocity of light has the same value c in all inertial frames of reference.

**Theorem 1.1 (Invariant of special relativity).** Consider two events A and B in an inertial observer O's frame of reference. Let the time interval measured by O between the two events be  $(\Delta t)$ , and the three spatial intervals be  $(\Delta x)$ ,  $(\Delta y)$ ,  $(\Delta z)$ . Then, the quantity

$$(\Delta s)^2 := (\Delta ct)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta y)^2 \tag{1.1}$$

is invariant between all frames of reference. I.e.

$$(\Delta s') = (\Delta s) \tag{1.2}$$

for any inertial frame of reference O'.

Theorem 1.1 follows directly from the invariance of the speed of light (consider a pair of mirrors, for two observers with relative velocity).

**Definition 1.1 (Proper time).** Proper time,  $\tau$ , is the time measured by an observer in the special frame of reference where the object in question is at rest. In this frame of reference,

$$(\Delta s)^2 = (c\Delta \tau)^2. \tag{1.3}$$

In the infinitesimal limit

$$(\mathrm{d}s)^2 = (c\,\mathrm{d}\tau)^2. \tag{1.4}$$

Physically, Defn. 1.1 means that all observers agree on the time interval between events for an observer who travels with the object in question. However, different observers **do not** in general agree on the time interval between events generally:  $(\Delta t) \neq (\Delta t')$  – this is called **time dilation**.

#### 1.2. c is an upper speed limit

All observers agree on the value of  $(\mathrm{d}s)^2=(c\,\mathrm{d}\tau)^2$ . Furthermore, we commonly assume that there exists a minimal proper time of  $\tau=0$  for two events if  $\Delta s^2=0$ . We can therefore write that when  $\tau=0$ 

$$c^{2} = \frac{(\mathrm{d}x)^{2} + (\mathrm{d}y)^{2} + (\mathrm{d}z)^{2}}{(\mathrm{d}t)^{2}}$$
(1.5)

between two events with an infinitesimal distance. We can equate the right-hand side with a squared velocity, and hence

$$\tau = 0 \implies c^2 = v^2 \tag{1.6}$$

so

$$(\mathrm{d}s)^2 \ge 0 \implies c^2 \ge v^2 \tag{1.7}$$

for **any** pair of events (which are causally connected, although how this follows is not immediately clear to me right now).

#### 1.3. Tensor notation and Minkowski spacetime

**Definition 1.2 (Four-vector (contravariant)).** A position four-vector is defined as

$$x^{\mu} = \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix} \equiv \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}. \tag{1.8}$$

**Definition 1.3 (Minkowski metric).** The Minkowski metric is defined as

$$\eta_{\mu\nu} = \text{diag}(1, -1, -1, -1).$$
(1.9)

 $\eta$  is used to compute distances and lengths in Minkowski space.

We define  $\eta^{\mu\nu}$  through the relation

$$\eta^{\mu\nu}\eta_{\nu\sigma} = \delta^{\mu}_{\ \sigma} \tag{1.10}$$

where we have appled the **Einstein summation convention**, where a repeated Greek index implies a summation from 0 to 3 (where the zeroth index is time), and a repeated Roman index is summed from 1 to 3. Hence, for a matrix multiplication between two  $3\times3$  matricies A and B,  $(AB)_{ij}=A_{ik}B_{kj}$ , and  $(A^T)_{ij}=A_{ji}$ .

**Definition 1.4 (One-form (covariant vector)).** We define a one-form as

$$x_{\mu} = \eta_{\mu\nu} x^{\nu}.\tag{1.11}$$

Thus,

$$ds^2 = \eta_{\mu\nu} dx^{\mu} dx^{\nu}. \tag{1.12}$$

**Definition 1.5 (Scalar product).** A scalar product between four-vectors x and y is defined as

$$x \cdot y := x^{\mu} y^{\nu} \eta_{\mu\nu} = x_{\mu} y_{\nu} \eta^{\mu\nu} = x^{\mu} y_{\mu} = x_{\nu} y^{\nu}$$
 (1.13)

due to the symmetry of the metric:  $\eta_{\mu\nu} = \eta_{\nu\mu}$ .

**Ordering (spacing) of indicies** In order to be able to freely raise/lower indicies (without repeatedly writing the metric tensor), we can impose an ordering upon indicies of tensor fields – which we can represent typographically with spacing between tensor indicies. A metric  $g_{ij}$  (or  $g^{ij}$ ) has the effect of lowering (or raising) a repeated index. For example,

$$g_{iq}T^{abcd}_{efgh}^{ijkl}_{mnop} = T^{abcd}_{efghq}^{jkl}_{mnop}. {(1.14)}$$

(Proof of this, I imagine, requires background in differential geometry?)

#### 1.4. Lorentz transformations

From the invariant of SR (Theorem 1.1), we have

$$ds'^{2} = dx'_{\mu} dx'_{\nu} \eta^{\mu\nu} = dx_{\mu} dx_{\nu} \eta^{\mu\nu}$$
(1.15)

for all reference frames. We denote  $\Lambda$  as a (1,1) tensor field, which transforms a four-vector from one reference frame to another:

$$\mathrm{d}x^{\prime\mu} = \Lambda^{\mu}{}_{\nu}\,\mathrm{d}x^{\nu} \tag{1.16}$$

which leaves the  $ds^2$  invariant, i.e.  $ds'^2 = ds^2$ . It follows that

$$\eta_{\mu\nu} = \Lambda^{\sigma}{}_{\mu}\Lambda^{\delta}{}_{\nu}\eta_{\sigma\delta} 
\eta = \Lambda^{T}\eta\Lambda.$$
(1.17)

The physical meaning of Eq.(1.17) is that Lorentz transformations leave the scalar product of Minkowski spacetime invariant: i.e. changes between frames of reference that respect the two postualtes of special relativity (Section 1.1). Conservation of the scalar product is analogous to rigid rotation (O) in Euclidean space  $(a \cdot b = a' \cdot b' = a^T O^T O b \implies O^T I O = I)$ , which preserves orientation  $(\det(\Lambda) = 1)$ .

Note that  $\Lambda^{\mu}_{\ \nu} \neq \Lambda^{\mu}_{\nu}$ . Beginning with Eq.(1.17),

$$\Lambda^{\mu}{}_{\rho}\Lambda^{\nu}{}_{\sigma}\eta_{\mu\nu} = \eta_{\rho\sigma}$$

we can raise one index, and lower one index, of  $\Lambda^{\nu}_{\ \sigma}$ 

$$\begin{split} & \Lambda^{\mu}_{\phantom{\mu}\rho} \eta_{\mu\nu} \Lambda^{\nu}_{\phantom{\nu}\sigma} \eta_{\nu\mu} \eta^{\sigma\lambda} = \eta_{\rho\sigma} \eta_{\nu\mu} \eta^{\sigma\lambda} \\ & \Lambda^{\mu}_{\phantom{\mu}\rho} \Lambda_{\mu}^{\phantom{\mu}\lambda} \eta_{\mu\nu} = \eta_{\mu\nu} \delta_{\rho}^{\phantom{\rho}\lambda} \\ & \Lambda^{\mu}_{\phantom{\mu}\rho} \Lambda_{\mu}^{\phantom{\mu}\lambda} = \delta_{\rho}^{\phantom{\rho}\lambda} \end{split} \tag{1.18}$$

so we see that  $\Lambda_{\nu}^{\mu}$  is the inverse of  $\Lambda^{\mu}_{\nu}$ .

### 2. Lie group theory

#### 2.1. Invariance, symmetry, and covariance

We call a quantity **invariant** if it does not change under particular transformations. E.g. if we transform  $A, B, C, ... \to A', B', C', ...$  and we have

$$F(A', B', C', ...) = F(A, B, C, ...)$$
(2.1)

then we say F is invariant under this transformation. **Symmetry** is defined as invariance under a transformation (or class of transformations). An equation is covariant if it takes the same form when objects in it are transformed. *All physical laws must be covariant under Lorentz transformations*.

Group theory describes the properties of particular sets of transformations: the invariances under such groups allows us to mathematically describe symmetry. For example, the set of rotations about the origin of a square by  $n\pi/2$  form a **discrete group**, and leave the set of points which constitute the square invariant under the transformation. The set of rotations about the origin of a circle form a **continuous group**. We can use group theory to work with *all* kinds of symmetries: symmetries which operate on vectors, equations, ...

## 2.2. Groups

**Definition 2.1 (Group axioms).** A group  $(G, \circ)$  is a set G, together with a binary operation  $\circ$  defined on G, that satisfies the following axioms

- Closure: For all  $g_1, g_2 \in G$ ,  $g_1 \circ g_2 \in G$
- Identity element: There exists an identity element  $e \in G$  such that for all  $g \in G$ ,  $g \circ e = g = e \circ g$
- Inverse element: For each  $g \in G$ , there exists an inverse element  $g^{-1} \in G$  such that  $g \circ g^{-1} = e = g^{-1}g$ .
- Associativity: For all  $g_1, g_2, g_3 \in G$ ,  $g_1 \circ (g_2 \circ g_3) = (g_1 \circ g_2) \circ g_3$

The set of all transformations that leave a given object invariant is called a **symmetry group**. For Minkowski spacetime, the object that is left invariant is the Minkowski metric, and the corresponding symmetry group is called the **Poincaré group**. Notice that the transformations which constitute a group are defined entirely independently from the object on which the transformations act.

## **2.2.1.** Rotations in two dimensions and SO(2)

Consider the 2D rotation matrix

$$R_{\theta} = \begin{pmatrix} \cos(\theta) & -\sin(\theta) \\ \sin(\theta) & \cos(\theta) \end{pmatrix} \tag{2.2}$$

and the two reflection matrices

$$P_x = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix} \qquad P_y = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \tag{2.3}$$

These matrices satisfy the group axioms. We can uncover this group from a symmetry perspective. The above transformations leave the length of a vector unchanged, i.e.

$$a.a = a'.a'. \tag{2.4}$$

Letting the transformation be represented by a'=Oa, it follows that all members of the group must satisfy

$$O^T O = I. (2.5)$$

This condition defines the group O(2), which is the group of all **orthogonal**  $2 \times 2$  matrices. It follows that  $\det(O) = \pm 1$  – i.e. the transformations are area-preserving. The subgroup with  $\det(O) = 1$  is called SO(2), which corresponds to rigid rotations preserving the orientation of the system – "S" denoting **special**.

## 2.2.2. Rotations with unit complex numbers and U(1)

A unit complex number is a complex number z which satisfies  $|z|^2 = z^*z = 1$ . The group U(1) is the set of unit complex numbers, together with ordinary complex number multiplication. The U stands for 'unitary', which generally stands for the condition

$$U^{\dagger}U = 1, \tag{2.6}$$

where  $U^\dagger=U^{T*}$  is the **Hermitian conjugate** of U. For scalars, the Hermitian conjugate is equivalent to the complex conjugate. Note that a unit complex number can also be denoted as

$$R_{\theta} = e^{i\theta} = \cos(\theta) + i\sin(\theta) \tag{2.7}$$

which makes the interpretation of U(1) as rotations on the unit complex numbers evident. We can connect this description of rotations (U(1)) to the previous (SO(2)) by defining

$$1 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \qquad , \qquad i = \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}. \tag{2.8}$$

For an arbitrary unit complex number z = a + ib, let

$$f(z) = a \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} + b \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} = \begin{pmatrix} a & -b \\ b & a \end{pmatrix}.$$
 (2.9)

Since  $z=R_{\theta}=\cos(\theta)+i\sin(\theta)$ , we can plug in the real and imaginary components of z into Eq.(2.9) to arrive at Eq.(2.2). We then have  $z'=R_{\theta}z$ , to perform rotations. There therefore exists an **isomorphism** between SO(2) and U(1):

**Definition 2.2 (Group isomorphism).** Given two groups (G,\*),  $(H,\odot)$ , a group isomorphism is a bijective function  $f:G\to H$  such that

$$f(u * v) = f(u) \odot f(v) \ \forall \ u, v \in G.$$
 (2.10)

which is written as

$$(G,*) \cong (H,\odot). \tag{2.11}$$

f(z) in Eq.(2.9) is therefore a group isomorphism between U(1) and SO(2).

# References

Schwichtenberg, J., 2015 Physics from symmetry. Springer.