Notes on Physics from Symmetry

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This document contains my personal notes on Jakob Schwichtenberg's Physics from Symmetry (Schwichtenberg, 2015).

1. Special relativity

1.1. Definitions and postulates

In special relativity, **inertial frames of reference** are coordinate systems moving with constant velocity relative to each other. Special relativity has two basic postulates:

- 1. **The principal of relativity**: The laws of physics are the same in all inertial frames of reference.
- 2. The invariance of the speed of light: The velocity of light has the same value c in all inertial frames of reference.

Theorem 1.1 (Invariant of special relativity). Consider two events A and B in an inertial observer O's frame of reference. Let the time interval measured by O between the two events be (Δt) , and the three spatial intervals be (Δx) , (Δy) , (Δz) . Then, the quantity

$$(\Delta s)^2 := (\Delta ct)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta y)^2 \tag{1.1}$$

is invariant between all frames of reference. I.e.

$$(\Delta s') = (\Delta s) \tag{1.2}$$

for any inertial frame of reference O'.

Theorem 1.1 follows directly from the invariance of the speed of light (consider a pair of mirrors, for two observers with relative velocity).

Definition 1.1 (Proper time). Proper time, τ , is the time measured by an observer in the special frame of reference where the object in question is at rest. In this frame of reference,

$$(\Delta s)^2 = (c\Delta \tau)^2. \tag{1.3}$$

In the infinitesimal limit

$$(ds)^2 = (c d\tau)^2.$$
 (1.4)

Physically, Defn. 1.1 means that all observers agree on the time interval between events for an observer who travels with the object in question. However, different observers **do not** in general agree on the time interval between events generally: $(\Delta t) \neq (\Delta t')$ – this is called **time dilation**.

1.2. c is an upper speed limit

All observers agree on the value of $(ds)^2 = (c d\tau)^2$. Furthermore, we commonly assume that there exists a minimal proper time of $\tau = 0$ for two events if $\Delta s^2 = 0$. We can therefore write that when $\tau = 0$

$$c^{2} = \frac{(\mathrm{d}x)^{2} + (\mathrm{d}y)^{2} + (\mathrm{d}z)^{2}}{(\mathrm{d}t)^{2}}$$
(1.5)

between two events with an infinitesimal distance. We can equate the right-hand side with a squared velocity, and hence

$$\tau = 0 \implies c^2 = v^2 \tag{1.6}$$

SO

$$(\mathrm{d}s)^2 \ge 0 \implies c^2 \ge v^2 \tag{1.7}$$

for **any** pair of events (which are causally connected, although how this follows is not immediately clear to me right now).

1.3. Minkowski Notation

Definition 1.2 (Four-vector (contravariant)). A position four-vector is defined as

$$x^{\mu} = \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix} \equiv \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}. \tag{1.8}$$

Definition 1.3 (Minkowski metric). The Minkowski metric is defined as

$$\eta_{\mu\nu} = \text{diag}(1, -1, -1, -1).$$
(1.9)

We define $\eta^{\mu\nu}$ through the relation

$$\eta^{\mu\nu}\eta_{\nu\sigma} = \delta^{\mu}_{\sigma} \tag{1.10}$$

where we have appled the **Einstein summation convention**, where a repeated Greek index implies a summation from 0 to 3 (where the zeroth index is time), and a repeated Roman index is summed from 1 to 3. Hence, for a matrix multiplication between two 3×3 matricies A and B, $(AB)_{ij}=A_{ik}B_{kj}$, and $(A^T)_{ij}=A_{ji}$.

Definition 1.4 (One-form (covariant vector)). We define a one-form as

$$x_{\mu} = \eta_{\mu\nu} x^{\nu}.\tag{1.11}$$

Thus,

$$ds^2 = \eta_{\mu\nu} dx^{\mu} dx^{\nu}. \tag{1.12}$$

Definition 1.5 (Scalar product). A scalar product between four-vectors x and y is defined as

$$x \cdot y := x^{\mu} y^{\nu} \eta_{\mu\nu} = x_{\mu} y_{\nu} \eta^{\mu\nu} = x^{\mu} y_{\mu} = x_{\nu} y^{\nu}$$
 (1.13)

due to the symmetry of the metric: $\eta_{\mu\nu} = \eta_{\nu\mu}$.

In general, a (q,r) tensor field can have q upper indicies and r lower indicies, defined through the following transformation law:

Definition 1.6 ((q,r) tensor). A (q,r) tensor is defined through the transformation law

$$T_{j'_1j'_2...j'_r}^{i'_1i'_2...i'_q}(x') = T_{j_1j_2...j_r}^{i_1i_2...i_q}(x) (M_{i_1}^{i'_1}...M_{i_q}^{i'_q}) (M_{j_1}^{j'_1}...M_{j_r}^{j'_r}). \tag{1.14}$$

1.4. Lorentz transformations

From the invariant of SR (Theorem 1.1), we have

$$ds'^{2} = dx'_{\mu} dx'_{\nu} \eta^{\mu\nu} = dx_{\mu} dx_{\nu} \eta^{\mu\nu}$$
(1.15)

for all reference frames. We denote Λ as a transformation to move from one reference frame to another

$$\mathrm{d}x^{\prime\mu} = \Lambda^{\mu}_{\ \nu} \, \mathrm{d}x^{\nu} \tag{1.16}$$

which leaves the $\mathrm{d}s^2$ invariant, i.e. $\mathrm{d}s'^2=ds^2.$ It follows that

$$\eta_{\mu\nu} = \Lambda^{\sigma}_{\mu} \Lambda^{\delta}_{\nu} \eta_{\sigma\delta}
\eta = \Lambda^{T} \eta \Lambda$$
(1.17)

References

Schwichtenberg, J., 2015 Physics from symmetry. Springer.