Lectures on Statistical Mechanics and Thermodynamics

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Chapter 1

Foundations of Statistical Mechanics

The object of Statistical Mechanics is to study systems with a large number of degrees of freedom by focusing solely on the macroscopic properties of the system.

Why, you may ask, are we only focusing on these macroscopic properties? Surely, given enough information we could model any system by following its individual members! This is not necessarily the case because the sheer volume of information necessary to do such a calculation is in fact an impediment to performing such a calculation. Let's illustrate this using a simple example of a system consisting of 20 particles in the ground state. Suppose, through divine intervention or otherwise, we have obtained the complete wave equation for each particle in the system. Before we begin to calculate any properties of this system we must first store this information.

As shown in figure 2, every particle wavefunction will be reasonably approximated by 10 points in each dimension. Taking the 3-dimensional case, we find that the amount of memory we would need would be $(10)^{3N} * 8 \simeq 10^{61}$ bytes. We include the factor of 8 since each real number will be represented by 8 bytes. Assuming we store all of this data on terabyte drives that have a storage capacity of $\frac{10^{12}bytes}{0.4kgs}$, we find that the total mass of all of the drives required to store the information describing this 20 particle system to be on the order of 10^{49} kilograms, which is much much more than the mass of the Milky Way galaxy! Clearly, it is impossible to store this much information, much less manipulate it. Therefore, we employ statistics to deal with the system at a macroscopic level.

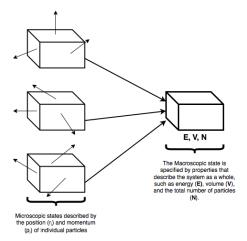


Figure 1.1: Multiple microscopic states make up a single macroscopic state of a system.

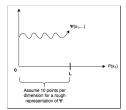


Figure 1.2: Wavefunction of our system in dimension x_1

1.1 Hamiltonian Dynamics

Consider a hypothetical system consisting of some number of particles which can be described by their individual positions, q_i and their momentums, p_i . We can indicate the paths these particles travel on a **Phase Space** diagram.

This diagram displays the evolution of the systems constituent particles as their positions and momentums change. The evolution of this system is governed by the Hamilton Equations: $\frac{dq_1}{dt} = \frac{\delta \mathcal{H}}{\delta p_i}$ and $\frac{dp_i}{dt} = -\frac{\delta \mathcal{H}}{\delta q_i}$ The evolution of p and q are clearly deterministic, therefore we can express these values as time-dependant functions. $q_i(t), p_i(t) \to q_i(t=0) = q_i^0$ and $p_i(t=0) = p_i^0$

There are a couple of tricks we can employ to idealize the measurement process and simplify manipulations of these functions.

1.2 Idealization of the Measurement Process

Let's start by defining some value \overline{f} , which is the result of measuring function $f(p_i, q_i)$.

$$\overline{f} = \lim_{T \to \infty} \frac{1}{T} \int_{\delta}^{\delta + T} dt f(q(t), p(t)) = \widetilde{f}(q_i^0, p_i^0)$$

T= duration of the measurement, $\delta=$ time the measurement starts, $\widetilde{f}=$ some function of the initial conditions.

This expression is an idealization of an arbitrary measurement. In reality, the time T is much larger than the microscopic scales in which the functions of q(t) and p(t) vary. For example, suppose the function f is the pressure on the wall of a gas containing vessel. The quantity f(q(t), p(t)) fluctuates like the diagram below.

However, except in specialized cases, the measuring apparatus is not accurate enough the capture every tiny fluctuation of the force on the wall of the vessel. Instead, it measures an average of f(q(t), p(t)) over a finite period of time:

Taking the $\lim_{T\to\infty}$ is a mathematical idealization of this averaging out of function fluctuations. Returning to our analysis of the hypothetical measurement \overline{f} , we see that the function $\widetilde{f}(q_i^0,p_i^0)$ is not really a function of the initial condition. Rather, it is a function of the trajectory of the system evolution. This may not be immediately apparent, but can be illustrated by choosing another initial point on the trajectory of our function $\widetilde{f}=(q_i^0,p_i^0)$. As $T\to\infty$, we see that our new initial points $q_i^{0'}$ and $p_i^{0'}$ yield the same value of \widetilde{f} as q_i^0 and p_i^0 . i.e., $\widetilde{f}(q_i^0,p_i^0)=\widetilde{f}(q_i^{0'},p_i^{0'})$ This is because the difference between the two trajectories is neglible compared to the infinite extended trajectories as $T\to\infty$.

Now, it could be possible that two different trajectories give different values for \overline{f} .

However, it turns out that, for systems we are interested in, that it does not happen. These systems have the property that any individual trajectory will pass through almost **every** point in phase space. This is called the **Ergodic Property**.

A couple of important things to keep in mind. The assumption that the trajectory passes through every point in the phase space is contingent on the the assumption that these points have the same value of energy (and any other conserved quantities) as the initial point. After all, these values are conserved by the Hamiltonian Flow. The phrase "almost every point" also has a specific meaning, in that it is "almost every point" in the measure theory sense.

It is really difficult to prove that a realistic system has the ergodic property. Only a few systems have actually been proven rigorously to be ergodic (for example, the Sinai Stadium). Moving forward, we will operate on the assumption that the systems we are interested are, in fact, ergodic.

To expand on the concept of ergodicity, it is very easy to find systems that are not ergodic. Any conserved value restricts the trajectories to lie on a submanifold of the phase space. If the only conserved quantities are the ones resulting from the standard symmetries (i.e. energy, momeentum, etc.) we can question the validity of ergodicity within the restricted subspace with constant values of these quantities. There are systems with so many other conserved quantities that the subspace is one dimensional (integrable systems), thereby invalidating the ergodicity of the system. The Integrable/Ergodic classification is also non-binary, there are intermediate types of systems that exist between the two extremes, such as mixing systems.

1.3 Introduction to Thermodynamics

At its core level, many thermodynamic problems can be phrased as a comparison between two states of a system. The pre-evolutionary equilibrium state and the post-evolutionary equilibrium state.

Since every microstate of the system is equally probably, the final macroscopic state will be the one which corresponds to the largest number of microstates:

$$\Gamma(E'_1, V'_1, N'_1, E'_2, V'_2, N'_2) = \Gamma_1(E'_1, V'_1, N'_1)\Gamma_2(E'_2, V'_2, N'_2)$$

Where $\Gamma(E_1', V_1', N_1', E_2', V_2', N_2')$ is the number of states of the composite system, $\Gamma_1(E_1', V_1', N_1')$ is the number of microstates of subsystem 1, and $\Gamma_2(E_2', V_2', N_2')$ is the number of microstates of subsystem 2.

Let's look at the case where the wall between the two subsystems allows the exchange of energy but not of volume or particle number. In this case we find:

$$E_1 + E_2 = E_1' + E_2'$$

$$V_1 = V_1', V_2 = V_2'$$

$$N_1 = N_1', N_2 = N_2'$$

$$\Gamma(E_1', E_2') = \Gamma_1(E_1')\Gamma_2(E_2') \quad (1.1)$$

This implies that

$$\Gamma_{Max,E_{total}fixed} \Leftrightarrow \frac{1}{\Gamma_1} \frac{d\Gamma_1(E_1')}{dE_1'} - \frac{1}{\Gamma_2(E_2')} \frac{d\Gamma_2(E_2')}{dE_2'} = 0$$

So, knowing the total number of initial microstates of the two subsystems allows us to predict the final state of the overall system. This is summed up in a very clean and useful fashion in the definition of **entropy**:

$$S = k_B \ln(\Gamma)$$

Notice that entropy is extensive as long as the subsystems are large enough that the boundary effects are negligible.

How can Entropy grow?

Within the constraints of the problem, there is an apparent paradox in staying that the entropy is maximized. That's because, using Hamiltonian equations, we can show S is a constant.

Define $\rho(q, p, t = 0)$ to be an initial distribution of identical systems (this is called an **ensemble**.) As the Hamiltonian Flow evolves in time every element of the ensemble is carried with it and the distribution evolves to $\rho(q, p, t) = \rho(q(t), p(t))$.

Because the number of systems in the ensemble is fixed ρ obeys the continuity equation:

$$\frac{d\rho}{dt} + \nabla * (\rho \nabla) = 0$$

or

$$0 = \frac{\delta \rho}{\delta t} + \sum_{i=1}^{3N} \left[\frac{\delta(\rho \dot{q}_i)}{\delta q_i} + \frac{\delta \rho \dot{p}_i}{\delta p_i} \right]$$

$$= \frac{\delta \rho}{\delta t} +$$

$$(1.2)$$

This result is known as the **Liouville Theorem**. It implies that the density of elements on the ensemble does not change as it evolves along the Hamiltonian Flow (of course, it does change at any fixed position in phase space). As the number of systems in the ensemble is fixed, the Liouville Theorem implies that the volume of the phase space occupied by the ensemble is fixed.

This means that the entropy, being the log of the volume, cannot change either. So how can the entropy grow if the microscopic equations of motion show that it does not? A resolution to this paradox is to notice that, for systems in which statistical mechanics holds, the Hamiltonian Flow takes nice looking, civilized ensembles and turns them into convoluted shapes:

The average of smooth observables (i.e. the value of f(q,p)) does not distinguish between the average performed with the true $\rho(t)$ or the "coarse grained" $\tilde{\rho}(t)$, as shown below.

$$\overline{f} = \int dq dp f(q, p) \rho(q, p, t) \simeq \int dq dp f(q, p) \widetilde{\rho}(q, p, t)$$

For all practical purposes, the entropy grows so long as we only look at observables that are very smooth. Macroscopic observables don't care about the precise poition of the particles and tend to be smooth in phase space. Systems whose Hamiltonian Flow have this property are called "Mixing." It is easy to show that ergodic systems are mixing, but keep in mind that mixing systems are not necessarily ergodic.

Chapter 2

Thermodynamics

From the general property of entropy we can derive a massive amount of information about a system. In fact, if the entropy as a function of extensive variables is known, everything else about the thermodynamic behavior of the system can be found! The definition of entropy is reiterated below:

$$S = k_B \ln(\Gamma)$$

We include the Boltzmann Constant, k_B , for historical reasons. However, if we judiciously choose our units (by measuring temperature in units of energy, as we should) we can set $k_B = 1$. A good thing to remember in this regard is that $10,000K \simeq 1eV$.

Since entropy is so important, let's review a few properties satisfied by the entropy:

- S as a function of the <u>extensive</u> variables contains all the information about a system that you need.
- S is itself extensive: $S(\lambda * E, \lambda * V, \lambda * N) = \lambda * S(E, V, N)$
- $\frac{\delta S}{\delta E} \geq 0$ We will eventually see that this implies $T \geq 0$
- S(E, V, N) is a concave function. We will eventually see that this implies stability:

$$S(E+\Delta*E,V+\Delta*V,N+\Delta*N)+S(E-\Delta*E,V-\Delta*V,N-\Delta*N)\leq 2S(E,V,N)$$

- S=0 at $\frac{1}{\frac{\delta S}{\delta E}}=0(T=0)$ This property depends on how we normalize the number of microstates, Γ . Quantum mechanics determines the right normalization that implies S(T=0)-0
- S is maximized in equilibrium within the constraints set by the physical setup. This is a direct consequence of the fact that all microstates are equally probable.

Inverting the entropy function S(E, V, N) allows us to obtain a function for the energy of the system in terms of entropy: E(S, V, N). If S, V, N are changed slowly (i.e. "adiabatically") so that the system stays in equilibrium at all intermediate states we find:

$$dE = \frac{\delta E}{\delta S}|_{V,N} dE + \frac{\delta S}{\delta V}|_{E,N} dV + \frac{\delta S}{\delta N}|_{E,V} dN$$

where

$$\frac{\delta S}{\delta E}|_{V,N} = \frac{1}{T}, \frac{\delta S}{\delta V}|_{E,N} = \frac{P}{T}, \frac{\delta S}{\delta N}|_{E,V} = \frac{\mu}{T}$$

To derive the first relation, we used $\frac{\delta E}{\delta S}|_{V,N}=\frac{1}{den}$ It's important to note that S is defined for equilibrium macroscopic states only. So dS or dE refer to changes slow enough such that the system continuously remains in equilibrium during its transition from its initial state to its final state. i.e., at any given time during the transition the system is in an equilibrium state. For instance, if the volume is changed suddenly and then the system is left to relax to equilibrium the change in entropy is not given by $\frac{p}{T}dV$ (it will, in fact, be larger).

From these relations we can derive the Equations of State:

$T = \frac{\delta E}{\delta S} _{V,N} = T(S, V, N)$	$rac{1}{T} = rac{\delta S}{\delta E} _{V,N} = rac{1}{T(E,V,N)}$
$-P = \frac{\delta E}{\delta V} _{S,N} = P(S, V, N)$	$-\frac{P}{T} = \frac{\delta S}{\delta V} _{E,N} = -\frac{P(E,V,N)}{T(E,V,N)}$
$\mu = \frac{\delta E}{\delta N} _{S,V} = \mu(S, V, N)$	$\frac{\mu}{T} = \frac{\delta S}{\delta N} _{E,V} = \frac{\mu(E,V,N)}{T(E,V,N)}$

These equations of state are particularly important because they are frequently accessible to experiments. Notice that the same quantities as functions of other variables are not equations of state and may not contain the same information.

As mentioned earlier, all thermodynamical problems can be reduced to a transition between two states of a system:

Depending on the properties of the wall between the two subsystems, we find that:

- Wall allows Energy exchange $\to T_1' = T_2'$
- Wall allows Volume exchange $\rightarrow V_1' = V_2'$
- Wall allows Particle exchange $\rightarrow \mu'_1 = \mu'_2$

2.1 Non-Relativistic, Non-Interacting, Monoatomic Gas

Let us compute the entropy, S, for one example:

$$\mathcal{H}(\mathbf{n},\mathbf{n}_i) = \sum_{i=1}^N \frac{\mathbf{n}_i^2}{2m}, (0 < \mathbf{n}_i < L)$$

2.1. NON-RELATIVISTIC, NON-INTERACTING, MONOATOMIC GAS 11

$$\begin{split} \Gamma(E,V,N) &= \int \prod_{i=1}^{N} d \times_{i} |_{i} \delta(E-sumseri=1N\frac{|_{i}^{2}}{2m}) \\ &= [\int_{0}^{L} dx]^{3N} \int_{-\infty}^{\infty} dp_{ix} dp_{iy} dp_{iz}...dp_{nx} dp_{ny} dp_{nz} * 2m\delta(2mE - \sum_{i=1}^{3N} p_{i}^{2})_{S} = k_{B} \ln \Gamma = k_{B}[N \ln V + \frac{3N-1}{2} \ln E - \ln \Gamma(1+\frac{N}{2}) + \frac{N}{2} \ln E - \frac{N}$$

The above expression violates extensivity (2) and the Neust Postulate (5). The violation of the Neust Postualte is not surprising: the behavior of a gas at T=0 is dictated by quantum mechanics, not classical mechanics (the atoms wavelength is much larger than the separation between them). What is surprising is that quantum mechanics connects the calculation above even at large T. What is missing from the classical calculation above is that the gas is composed of identical particles (bosons or fermions, it doesn't matter at high T). So the correct way of counting states is:

$$S = k_B \ln \Gamma = k_B \int \prod_{i=1}^{N} \frac{dq_i dp_i}{N!h^N} \delta(E - \sum_{i=1}^{N} \frac{\frac{1^2}{2m}}{2m})$$

The denominator factor of N! avoids double counting of microstates with the same particle position and the factor of h^N becomes important when considering the classical limit of quantum gases.

As the number of particles in the system becomes very large we find that the expression becomes

$$S = k_B * N \ln \frac{VE^{\frac{3}{2}}}{N^{\frac{5}{2}}}$$

From this entropy we can find the equations of state:

$\frac{1}{T} = \frac{\delta S}{\delta E} _{V,N} = \frac{3}{2} \frac{k_B * N}{E}$	$E = \frac{3}{2}k_B * N * T$
$-\frac{P}{T} = \frac{\delta S}{\delta V} _{E,N} = -\frac{k_B * N}{V}$	$PV = k_B * NT$
$\frac{\mu}{T} = \frac{\delta S}{\delta N} _{E,V} =$	
$k_B \ln \frac{VE^{\frac{3}{2}}}{N^{\frac{5}{2}}} - \frac{5}{2}k_B$	

Note that we found the Boyle-Mariott law without even discussing atoms colliding against walls or other kinetic theory considerations. There's a lot behind the assumption of ergodicity.

This leads us to....

2.2 Euler Relation

$$S(\lambda * E, \lambda * V, \lambda * N) = \lambda * S(E, V, N)$$

Taking the derivative of this expression in terms of λ yields

$$E()\frac{\delta S}{\delta E} = \frac{1}{T}) + V(\frac{\delta S}{\delta V} = -\frac{P}{T}) + N(\frac{\delta S}{\delta N} = \frac{\mu}{T}) = S$$

Which can be re-arranged to yield the following equations:

$$E = TS - PV + \mu * N$$

$$S = \frac{E}{T} - \frac{PV}{T} + \frac{\mu * N}{T}$$

2.2.1 Recovering the Entropy or Energy from an Equation of State

If three equations of state are available:

$$T(S, V, N) = \frac{\delta E}{\delta S}|_{V,N}$$
$$-P(S, V, N) = \frac{\delta E}{\delta V}|_{S,N}$$
$$\mu(S, V, N) = \frac{\delta E}{\delta N}|_{S,V}$$

$$\rightarrow E(S, V, N) = T(S, V, N) = T(S, V, N)S - P(S, V, N)V + \mu(S, V, N)N$$

If only two equations of state are available: $E(S, V, N) = NE(\frac{S}{N}, \frac{V}{N}, 1)$ The equality above is due to extensivity.

$$T = \frac{\delta E}{\delta S}|_{V,N} = T(S, V, N) = T(\frac{S}{N}, \frac{V}{N}, 1)$$
$$-P = \frac{\delta E}{\delta V}|_{S,N} = P(S, V, N) = P(\frac{S}{N}, \frac{V}{N}, 1)$$

Note that there is no N!. T and P are intensive.



$$E(\frac{S}{V}, \frac{V}{N}, 1) = \int_{0}^{\frac{S}{N}} d(\frac{S}{N}) T(\frac{S}{N}, \frac{V}{N}, 1) + E(0, \frac{V}{N}, 1) = \int_{0}^{\frac{S}{N}} d(\frac{S}{N}) T(\frac{S}{N}, \frac{V}{N}, 1) - \int_{0}^{\frac{V}{N}} d(\frac{V}{N}) P(0, \frac{V}{N}, 1) + E(0, \frac{V}{N}, 1) = \int_{0}^{\frac{S}{N}} d(\frac{S}{N}) T(\frac{S}{N}, \frac{V}{N}, 1) + E(0, \frac{V}{N}, \frac{V}{N}, 1) = \int_{0}^{\frac{S}{N}} d(\frac{S}{N}) T(\frac{S}{N}, \frac{V}{N}, 1) + E(0, \frac{V}{N}, \frac{V}{N}, 1) = \int_{0}^{\frac{S}{N}} d(\frac{S}{N}) T(\frac{S}{N}, \frac{V}{N}, \frac{V}$$

In the above expressions, $T(\frac{S}{N}, \frac{V}{N}, 1) = \frac{\delta E}{\delta S}|_{V,N}$ and $E(0,0,1) = E(\frac{S}{N}, \frac{V}{N}, 1)$. The latter equality can be found up to one constant independant of S,V, or N.

Now, if only one equation of state is available, you cannot recover the entropy or energy. There is simply not enough information in a single equation of state to extract these system properties.

2.2.2 EXAMPLE: Non-relativistic, Monoatomic, Non-interacting Gas

$$\begin{split} S &= k_B * N \ln \frac{VE^{\frac{3}{2}}}{N^{\frac{5}{2}}}, E = (\frac{N^{\frac{5}{2}}}{V}e^{\frac{S}{k_B*N}})^{\frac{2}{3}} = \frac{N^{\frac{5}{2}}}{V^{\frac{2}{3}}} * e^{\frac{2S}{3k_B*N}} \\ & T = \frac{\delta E}{\delta S}|_{V,N} = \frac{2}{3k_B*N}\frac{N^{\frac{5}{3}}}{V^{\frac{2}{3}}}e^{\frac{2S}{3k_B*N}} \\ & -P = \frac{\delta E}{\delta V}|_{S,N} = -\frac{2}{3}\frac{N^{\frac{5}{3}}}{V^{\frac{5}{3}}}e^{\frac{2S}{3k_B*N}} \\ & E(\frac{S}{N}, \frac{V}{N}, 1) = \int_0^{\frac{S}{N}} d(\frac{S}{N})\frac{2}{3k_B}(\frac{N}{V})^{\frac{2}{3}}e^{\frac{2S}{3k_B*N}} - \int_0^{\frac{V}{N}} d(\frac{V}{N})\frac{2}{3}(\frac{N}{V})^{\frac{5}{3}} + E(0, 0, 1) \\ & E(\frac{S}{N}, \frac{V}{N}, 1) = \frac{2}{3k_B}(\frac{N}{V})^{\frac{2}{3}}\frac{3k_B*N}{2}[e^{\frac{2S}{3k_B*N}} - 1] - \frac{2}{3}(\frac{-3}{2})(\frac{N}{V})^{\frac{2}{3}} \\ & E(\frac{S}{N}, \frac{V}{N}, 1) = \frac{N^{\frac{5}{3}}}{V^{\frac{1}{3}}}e^{\frac{2S}{3k_B*N}} \end{split}$$

2.3 Minimum Energy Principle

Before formally stating the minimum energy principle, let's begin with an analogy. We know that a circle is the figure with the largest area for a given perimeter. Of course, it is also the the figure with the smallest perimeter for a given area. Even if a given circle was built by someone with a fixed length string trying to enclose the largest area it will still be true that the final circle minimizes the perimeter for the given area.

The same thing is true for the maximum entropy principle. After some constraints are changed the final state will be the one with the largest entropy for a given **total** energy (the constraint that the total energy is unchanged is always there). But the final equilibrium state can also be characterized as the state with the smallest energy for a fixed entropy:

Convince yourself that S as a function of total E and subsystem energy E_1^1 has the general shape depicted above. The minimum energy principle follows from the picture above.

Now let's see how the minimum energy principle works in our archetypal example :

$$E = E_1(S_1^1, V_1, N_1) + E_2(S_2', V_2, N_2)$$

Where $S_2' = S - S_1'$.

Therefore, we find that the minimum energy, E, for fixed entropy, S, is:

$$\frac{\delta E}{\delta S_1'} = 0 \to \frac{\delta E_1}{\delta S_1'} - \frac{\delta E_2}{\delta S_2'} = 0 \to T_1 = T_2$$

Which confirms what we expected: In its final state our example subsystem pair reaches an equilibrium temperature.

2.4 Efficiency of Engines

In our thermodynamic terminology, an 'engine' is a machine that extracts work from the the energy difference between two systems in contact. To clarify, the engine takes energy from the 'hot' source, uses some of that energy to perform work, and ejects the remaining energy into the 'cold' sink. In many cases, the 'cold' sink is the atmosphere: one example where this is true is a car engine, which throws off heat as it performs the work of moving the car.

In designing our thermodynamic engine, we would like it to perform this energy transfer over and over again in a closed cycle. I.e. we want the system to end up in the same macroscopic state as it started. For our purposes right now we will assume that all changes in the system caused by the energy transfers occur slowly (adiabatically) so, that at every point in the cycle the machine is in equilibrium and its macroscopic state is specified by E,V,N or S,V,N. It turns out, that for thermodynamic calculations it is better to use the variables T,S,N instead. In our engine we want to keep N fixed by having impermeable walls separating the machine from the rest of the Universe. This machine cycle can be represented by a closed loop in the TxS space:

Let us start by considering one specific cycle, the **Carnot Cycle**:

For a fixed N: dE = TdS - PdV so that

$$\Delta Q_1 = \int_{S_1}^{S_2} T dS = T_h(S_2 - S_1)$$
$$\Delta Q_2 = \int_{S_2}^{S_1} T dS = -T_c(S_2 - S_1)$$

Since after a full cycle the energy of the system goes back to what it was in the beginning we find:

$$\Delta Q_1 + \Delta Q_2 = \Delta W =$$

the area inside the cycles in the TxS plane.

In the expression above, $\Delta Q_1 + \Delta Q_2$ is the net heat absorbed by the system and ΔW is the work done by the system.

For practical reasons, we would like to get as much work as possible from a given amount of heat absorbed. The efficiency of the engine is defined as"

$$\epsilon = \frac{\Delta W}{\Delta Q_1} = \frac{(T_h - T_c)(S_2 - S_1)}{T_h(S_2 - S_1)} = 1 - \frac{T_c}{T_h} < 1$$

Now, one very important thing to note about this expression is that the total efficiency is *less than* not *less than or equal to* one because some of the energy is wasted and dispersed into the cold sink. This begs the question: is there a way that we could change the cycle to obtain a larger efficiency?

Going back to our TxS plane, we would want to maximize the area inside of the cycle (ΔW) and minimize the area below the top curve (ΔQ_1) . It's clear, for instance, that the cycle on the left can have its efficiency improved by changing it as shown below. We end up then with the Carnot cycle, which is the best that can be done. Thus, we have the universal result:

$$\epsilon \le 1 - \frac{T_c}{T_h}$$

Notice that we only used the fact that an ideal gas was used inside the engine to describe the different phases of the cycle in words. The main result, $E \leq 1 - \frac{T_c}{T_b}$ does not depend on that and is valid universally.

Finally, let us revisit the assumption of adiabaticity, namely, that all phases of the cycle were infinitely slow. In each phase of the cycle the machine foes from the one initial state A to a final state B exchanging a certain amount of heat ΔQ (with a source at temperature T) and performing some work ΔW (positive or negative). I'll assume the work is done on a reversible work source, namely, a system which can exchange work but not heat (a paper weight on top of a piston is an example). Similarly, heat is exchanged with a reversible heat source.

For given initial and final states A and B (and consequently, given changes in energy $E_B - E_A$ and entropy) the maximum amount of work will be delivered if ΔQ is a minimum as $E_B - E_A = -(\Delta W + \Delta Q)$. The process with the smallest ΔQ will correspond to the smallest possible increase of the entropy of the universe (system+heat source, the work source has no entropy). How much is the minimum amount of heat to be transferred to the heat source? As the total entropy of the universe must increase (otherwise the system would stay at A) we have

$$S_B - S_A + T_r \Delta Q \ge 0$$

Where T_r is the temperature of the heat source. That means that the minimum amount of heat "wasted" on the heat source happens when the total entropy of the universe is unchanged as is given by

$$\Delta Q = \frac{S_A - S_B}{T_r}$$

If more heat than $\frac{S_A - S_B}{T_r}$ is delivered to the heat source the universe's entropy will increase and less work will be delivered to the work source. A similar analysis applies when the system is absorbing heat or work. The main point is that **reversible processes**, those where the entropy of the universe does not change, lead to the maximum efficiency. Thus, our bound on the efficiency of an engine obtained assuming adiabaticity, is not violated by non-adiabatic processes. Notice that the adiabaticity condition means that the engine works really slowly. Practical Machines, which need to do their job fast, are not adiabatic and much less efficient than the theoretical limit would expect (roughly a factor of 3 less efficient).