

ASEN 6060

ADVANCED ASTRODYNAMICS

Orbital Stability

Objectives:

- Define common approach to assessing orbital stability
- Present an overview of insight gained by assessing orbital stability across a family of periodic orbits
- Define common bifurcations

Periodic Orbit Stability

- Orbital stability describes the behavior of trajectories in the local neighborhood of the periodic orbit
- Lyapunov definition of stability of a reference solution $\bar{x}_R(t)$:

Linearizing About a Periodic Orbit

Linearizing relative to a trajectory produces variational equations

$$\delta \dot{\bar{x}}(t) = \mathbf{A}(t)\delta \bar{x}(t)$$

With the following solution:

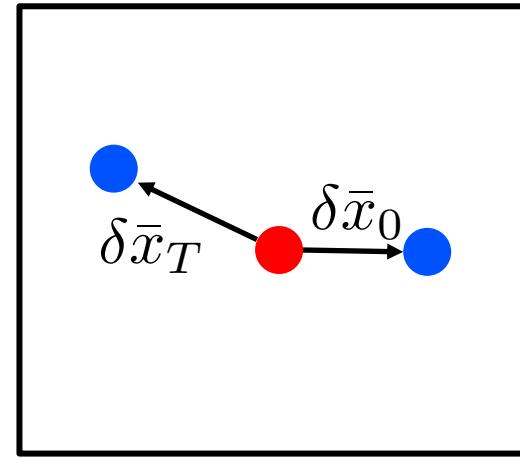
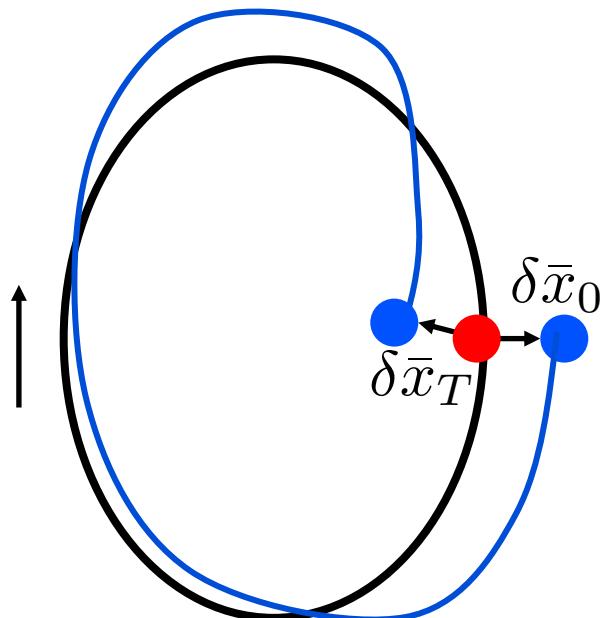
$$\delta \bar{x}(t) = \Phi(t, t_0)\delta \bar{x}(t_0)$$

where $\dot{\Phi}(t, t_0) = \mathbf{A}(t)\Phi(t, t_0)$ $\Phi(t_0, t_0) = \mathbf{I}_{6 \times 6}$

For a periodic reference path, $A(t)$ is time-varying and periodic

Monodromy Matrix

Monodromy matrix \mathbf{M} (6x6 matrix):



Eigendecomposition of Monodromy Matrix

If the solution to the linear system is $\delta\bar{x}_T = \mathbf{M}\delta\bar{x}_0$, study types of motion near periodic orbit by decomposing \mathbf{M} into eigenvalues and eigenvectors:

where

And recall the definition of an eigenvector:

$$\delta\bar{x}_0 = \bar{v}_i$$



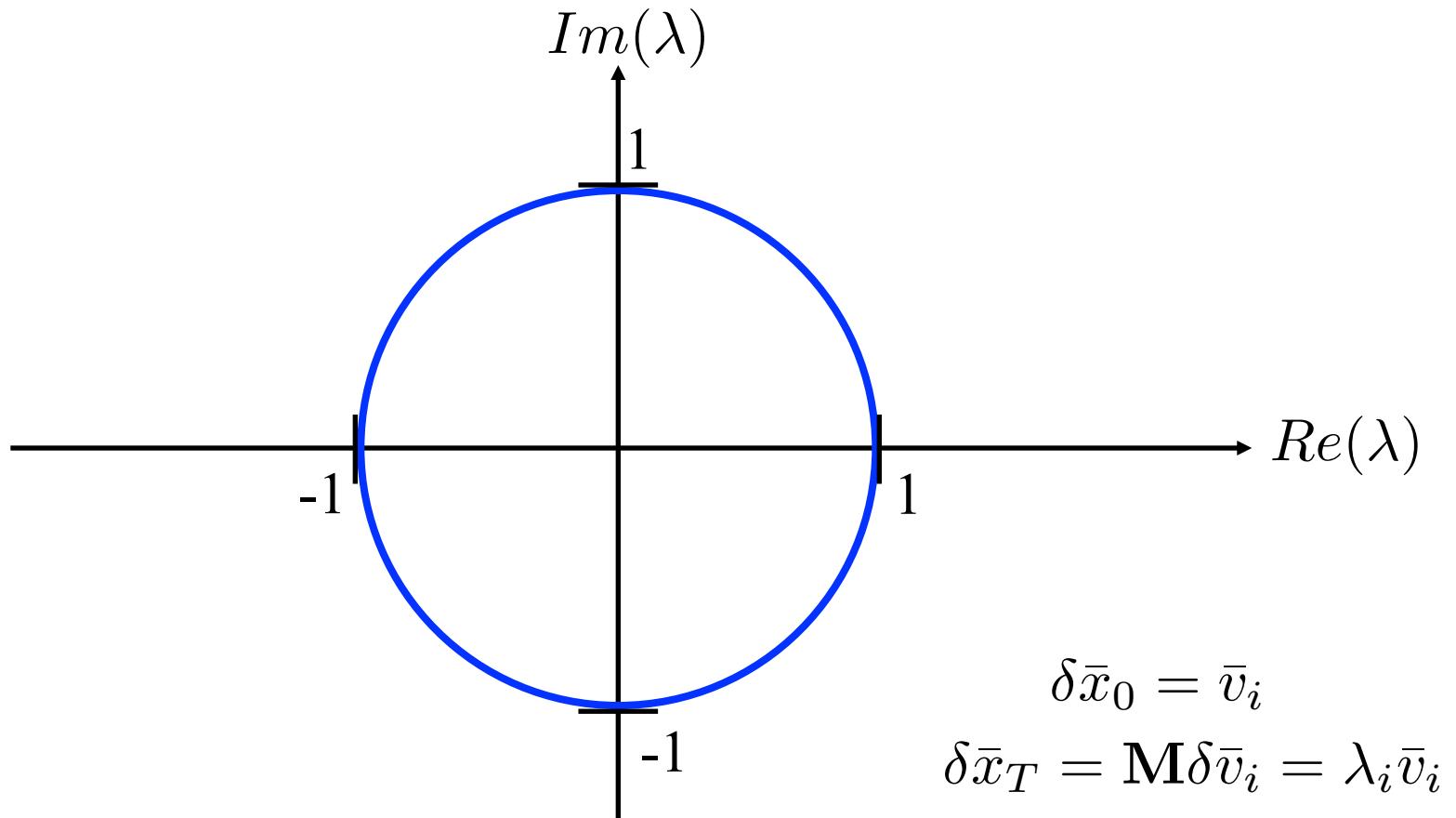
Eigendecomposition of Monodromy Matrix

Decomposing \mathbf{M} into eigenvalues and eigenvectors:

$$\mathbf{M} = \Phi(T, 0) = \mathbf{V}\Lambda\mathbf{V}^{-1}$$

Eigenvalues of Monodromy Matrix

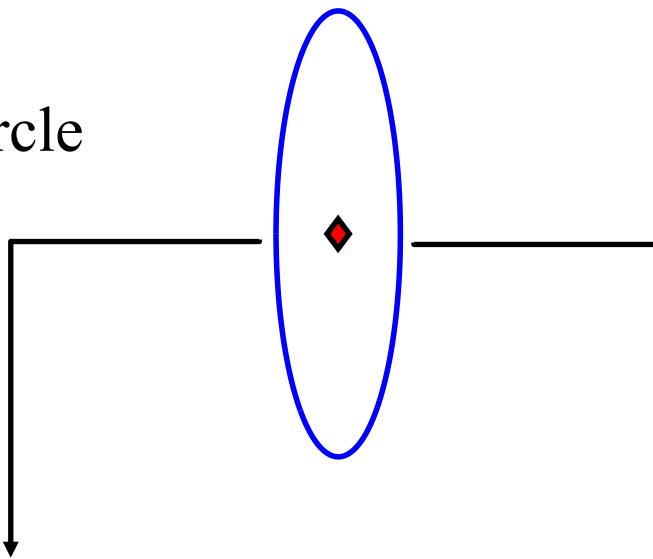
$$\lambda = a \pm bi$$



Eigenvalues of Monodromy Matrix

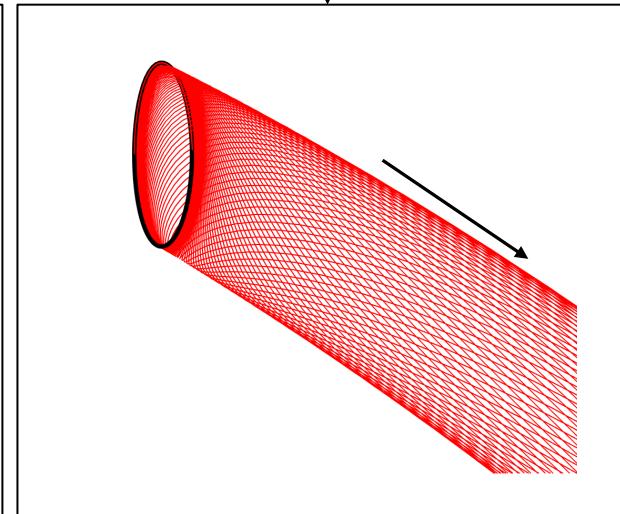
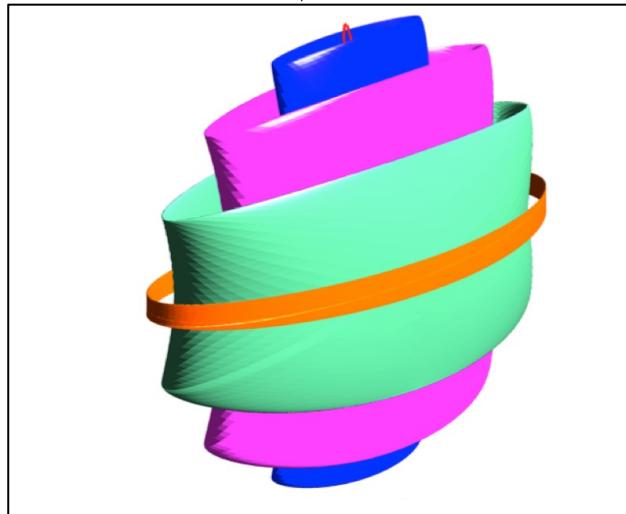
One pair of complex eigenvalues on unit circle

$$|\lambda| = 1$$



One reciprocal pair of real eigenvalues with
 $|\lambda| < 1$
 $|\lambda| > 1$

**Quasi-
Periodic
Trajectory**
tracing out
the surface
of a torus



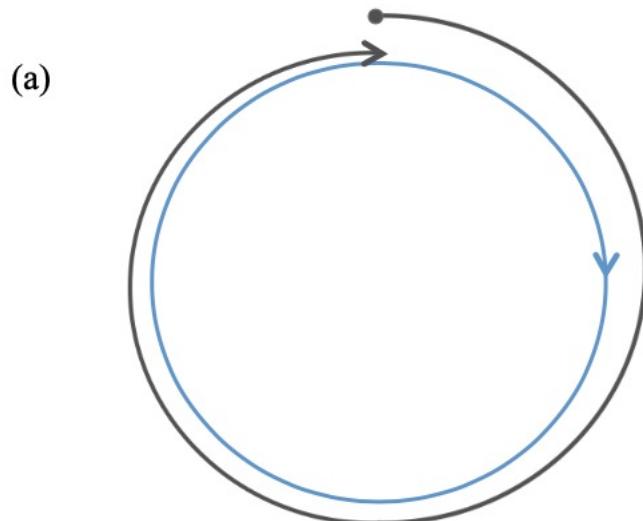
**Stable/
Unstable
Manifolds**

Eigenvalues of Monodromy Matrix

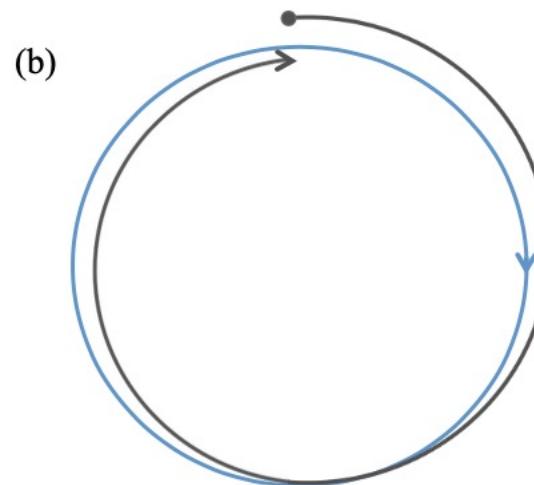
Sign of real eigenvalues indicates relative direction of motion and twisting about the orbit when exciting that mode

Example: $|a| < 1$

$$\lambda = a > 0$$



$$\lambda = a < 0$$



$$\delta \bar{x}_0 = \bar{v}_i$$

$$\delta \bar{x}_T = \mathbf{M} \delta \bar{v}_i = \lambda_i \bar{v}_i$$

Image credit: Bosanac 2016

Stability Index

Tracking the eigenvalues of the monodromy matrix for members of a periodic orbit family supplies insight into the stability of those members, and the identification of any structural changes in the local neighborhood of those orbits

Challenging to visualize the nontrivial eigenvalues in complex plane

Often use stability indices:

2 stability indices for periodic orbit (don't calculate for +1,+1 pair)

There are also other (less intuitive) ways to represent stability information succinctly!

Stability Index

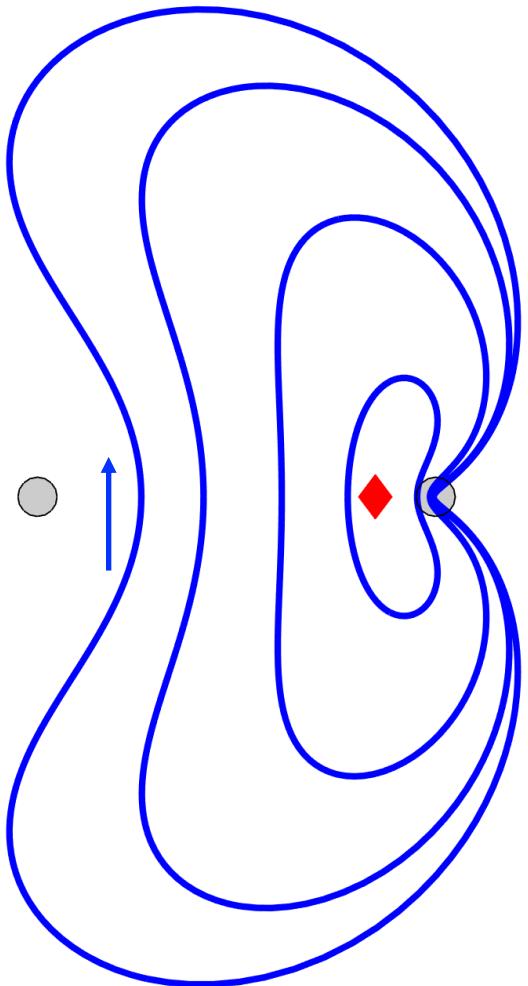
Stability indices:

$$s_1 = \lambda_1 + \lambda_2$$

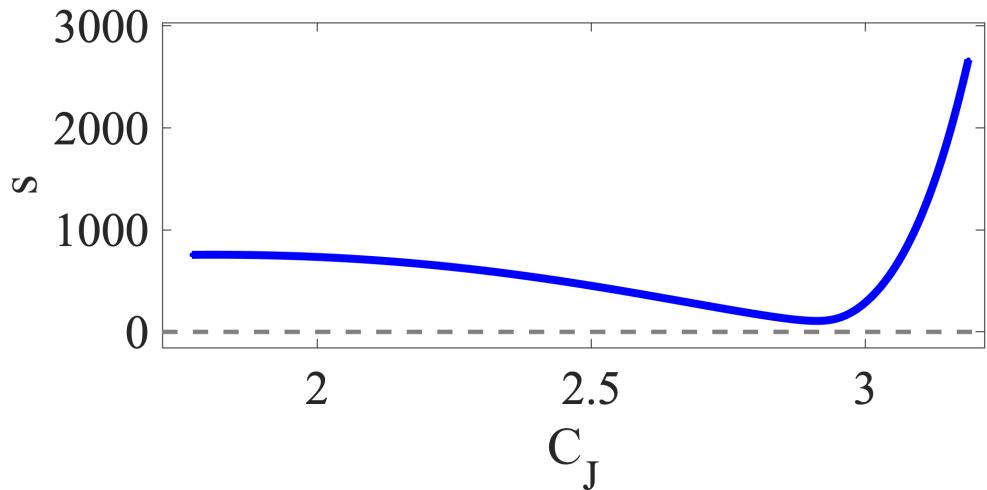
$$s_2 = \lambda_3 + \lambda_4$$

Observations:

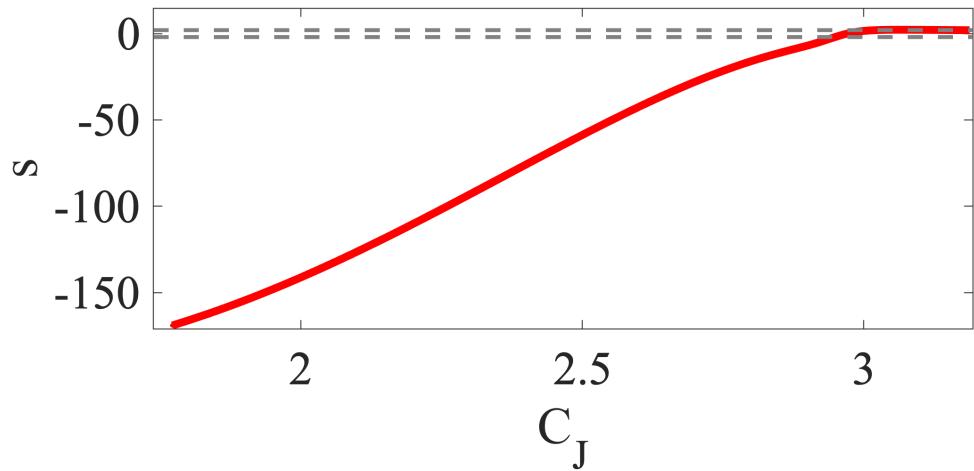
Example: L_1 Lyapunov Orbits



Stability index for in-plane modes



Stability index for out-of-plane modes



Local Bifurcations

Recall: Periodic orbits exist in a family that may be described by a continuously varying natural parameter

A local bifurcation occurs if a change in the natural parameter results in a change in the qualitative behavior of the trajectories in the vicinity of a periodic orbit

A variety of bifurcations exist and may produce a change in the stability of periodic orbits along the family, formation of a new family/ies of periodic orbits, or termination of the current family

- Selected examples discussed in these notes.

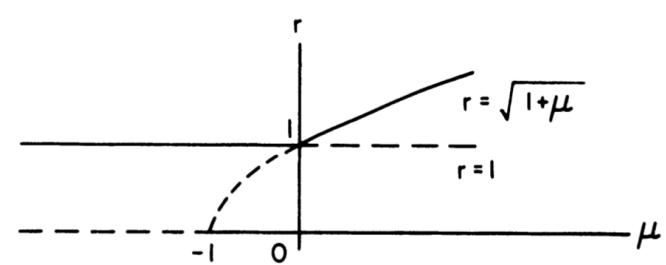
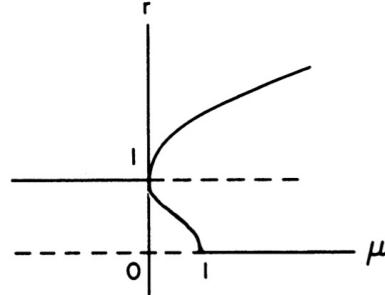
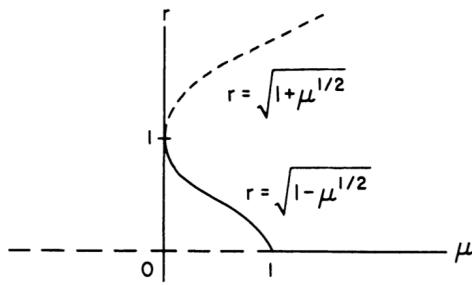
(Interesting reading: Ch. 4 of Perko, 2000, “Differential Equations and Dynamical Systems” 3rd edition, See link on Canvas page to free download via CU libraries)

Local Bifurcations

Nontrivial eigenvalues pass through critical values $+1, +1, s = 2$

Three well-known types:

- *Saddle-node/fold bifurcation*: stability change, no new/intersecting families
- *Pitchfork bifurcation*: as stability of periodic orbits change along family, two additional families of similar period are formed with same stability as members of original family before bifurcation
- *Transcritical bifurcation*: family of stable orbits intersects family of unstable orbits and the two families exchange stability characteristics at the intersection



Period-Multiplying Bifurcation

A family of period- mq orbits emerges from a family of period- q orbit (with orbit period T)

Example: period-doubling bifurcation

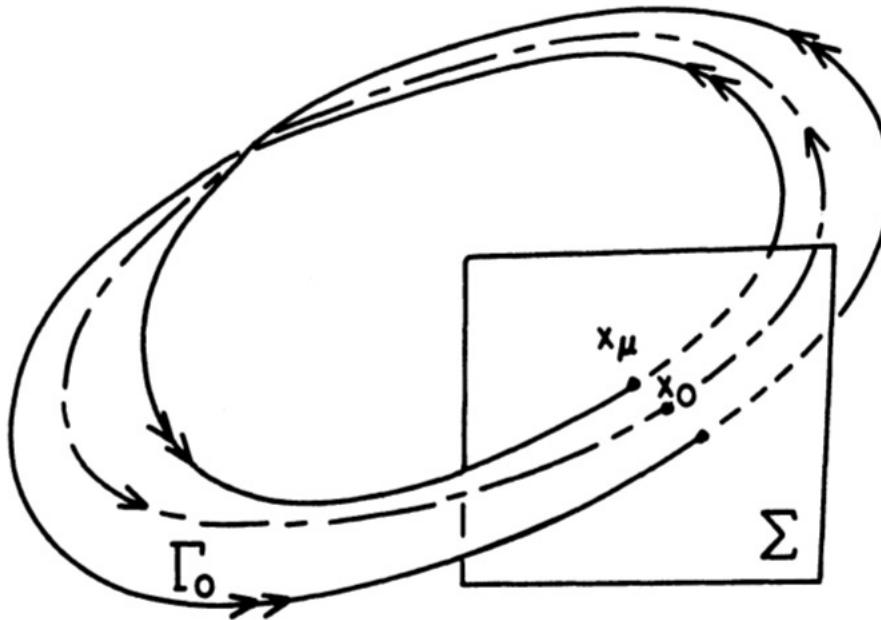


Image credit: Perko, 2000, “Differential Equations and Dynamical Systems” 3rd edition

Period-Multiplying Bifurcation

Express monodromy matrix of period- mq orbit, $\Phi_{mq}(mT, 0)$, in terms of monodromy matrix of period- q orbit, $\Phi_q(T, 0)$, at the intersection of the two families

$$\Phi_{mq}(mT, 0) = \Phi_q(mT, (m-1)T) \Phi_q((m-1)T, (m-2)T) \dots \Phi_q(2T, T) \Phi_q(T, 0)$$

$$\Phi_{mq}(mT, 0) = [\Phi_q(T, 0)]^m$$

The eigenvalues of the period- mq orbit are related to the eigenvalues of the period- q orbit

$$\lambda_{mq} = \lambda_q^m$$

Period-Multiplying Bifurcation

At the intersection of the two families when the period- mq orbits form, there is one nontrivial pair with:

- $\lambda_{mq} = 1, 1$
- λ_q are the roots of unity
- Detect the bifurcation in the original period- q family when the eigenvalues equal:

and when the stability index equals:

Period-Doubling Bifurcation

Limiting case is the period-doubling bifurcation when:

$$\lambda_i, \lambda_j = -1, -1$$

and is accompanied by a change in stability

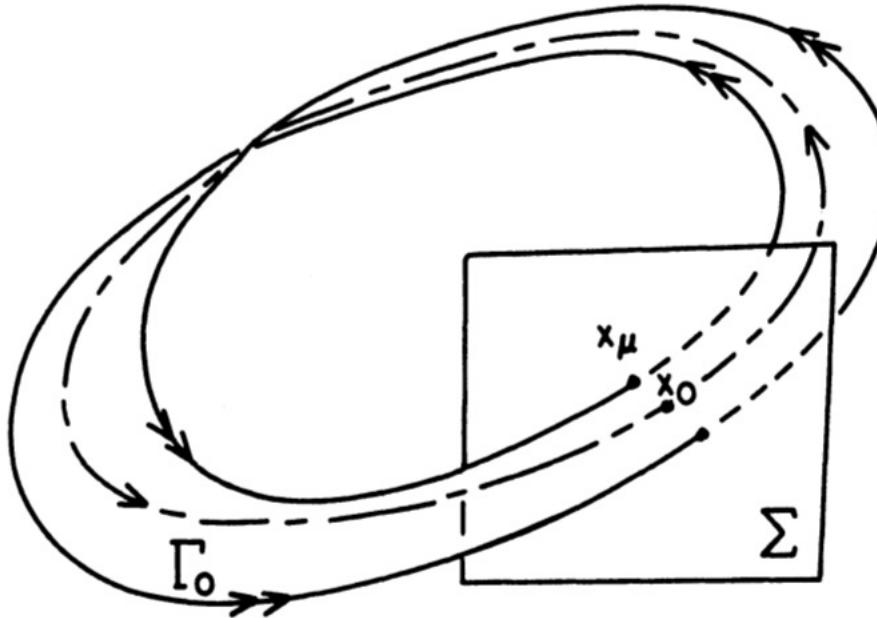
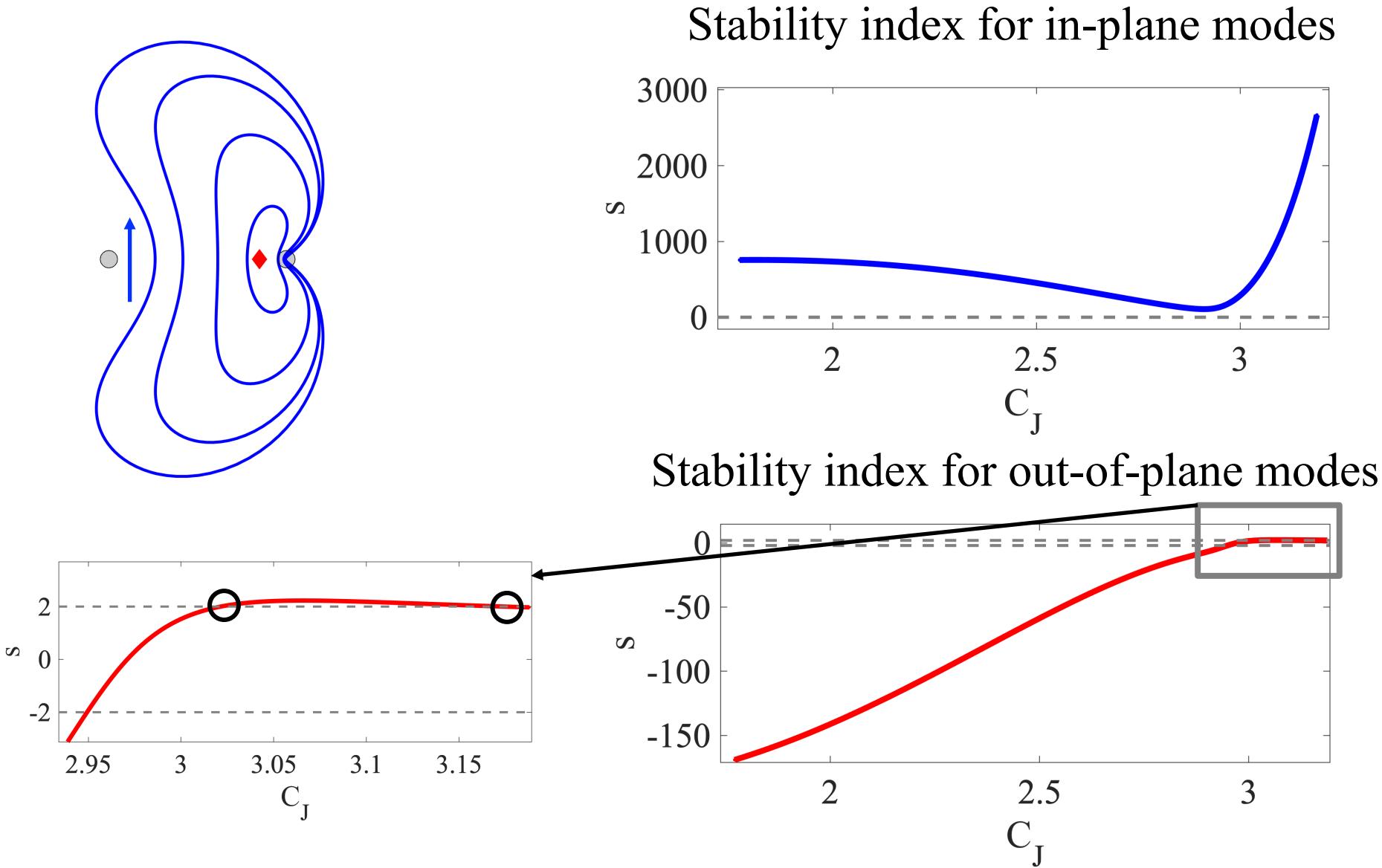


Image credit: Perko, 2000, “Differential Equations and Dynamical Systems” 3rd edition

Example: L_1 Lyapunov Orbits



Periodic Orbit Families near L_1

Bifurcations along L_1 Lyapunov orbit family correspond to intersections with L_1 halo and axial orbit families

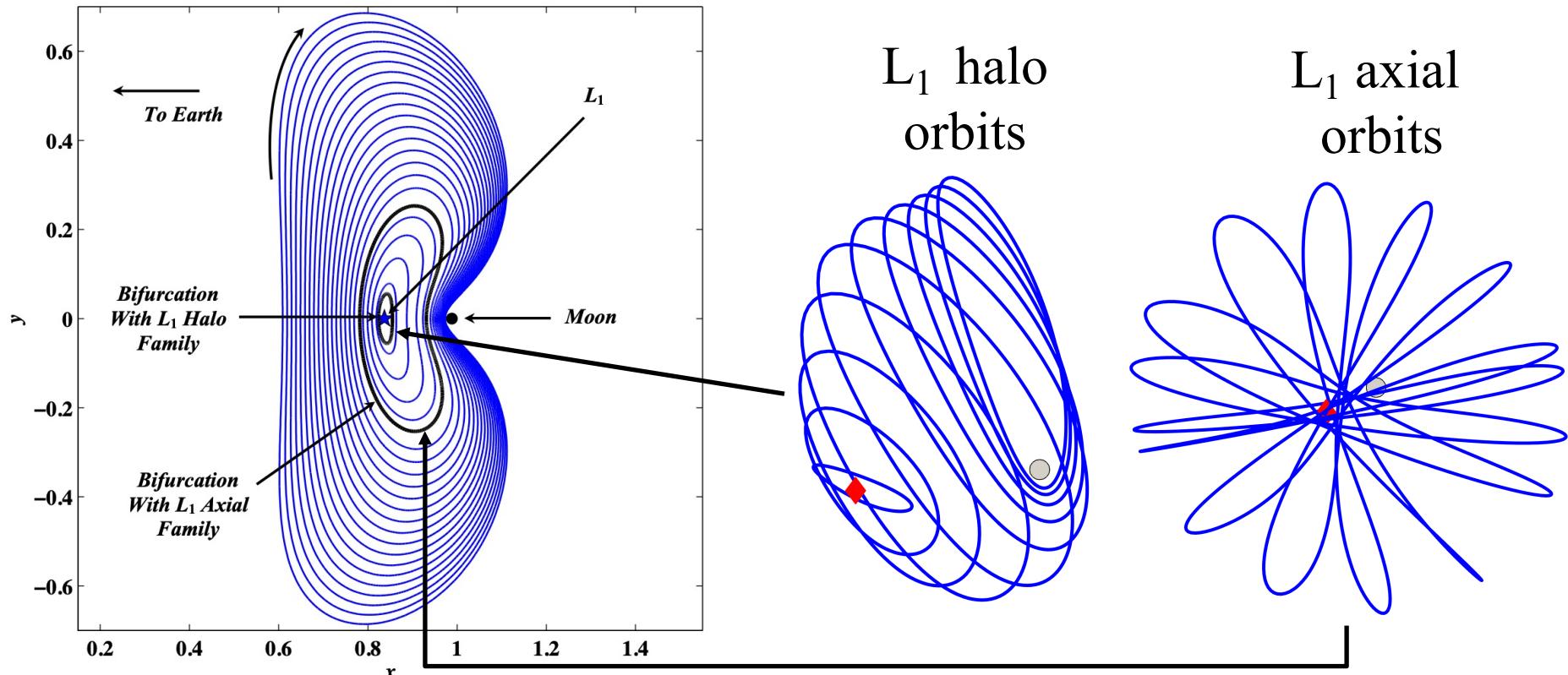
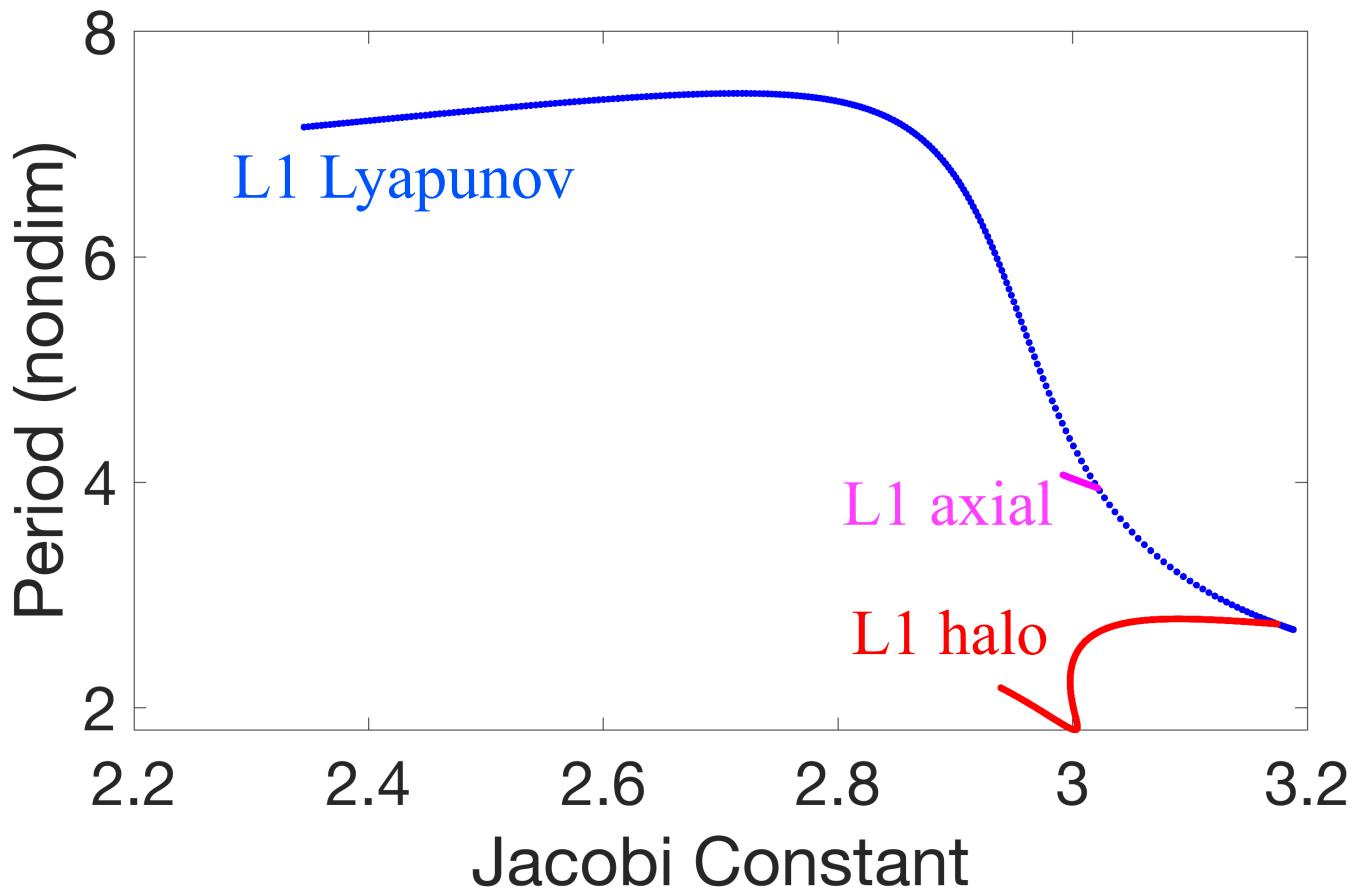


Figure 3.5 The L_1 Lyapunov Family

Left image credit: Grebow, D. 2006, "GENERATING PERIODIC ORBITS IN THE CIRCULAR RESTRICTED THREE- BODY PROBLEM WITH APPLICATIONS TO LUNAR SOUTH POLE COVERAGE", MS Thesis

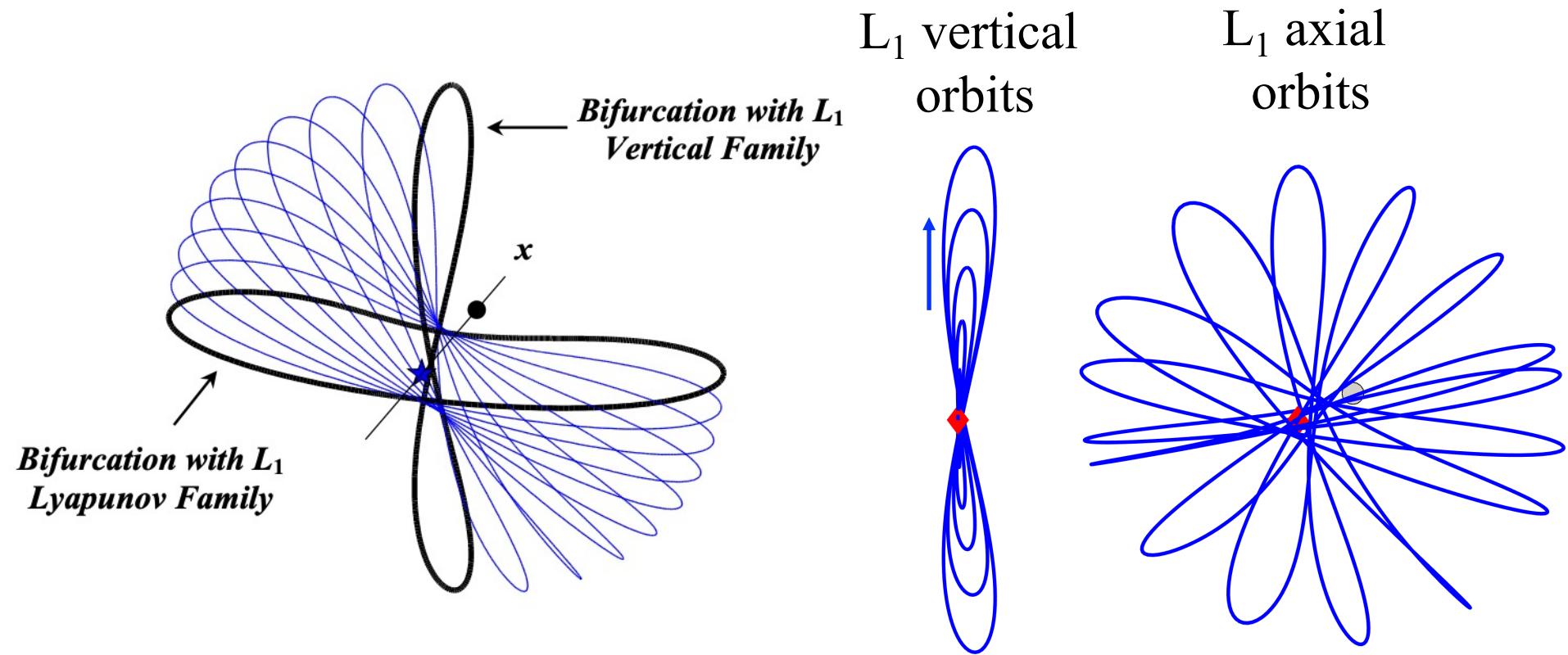
Periodic Orbit Families near L_1

Bifurcations along L_1 Lyapunov orbit family correspond to intersections with L_1 halo and axial orbit families



Periodic Orbit Families near L_1

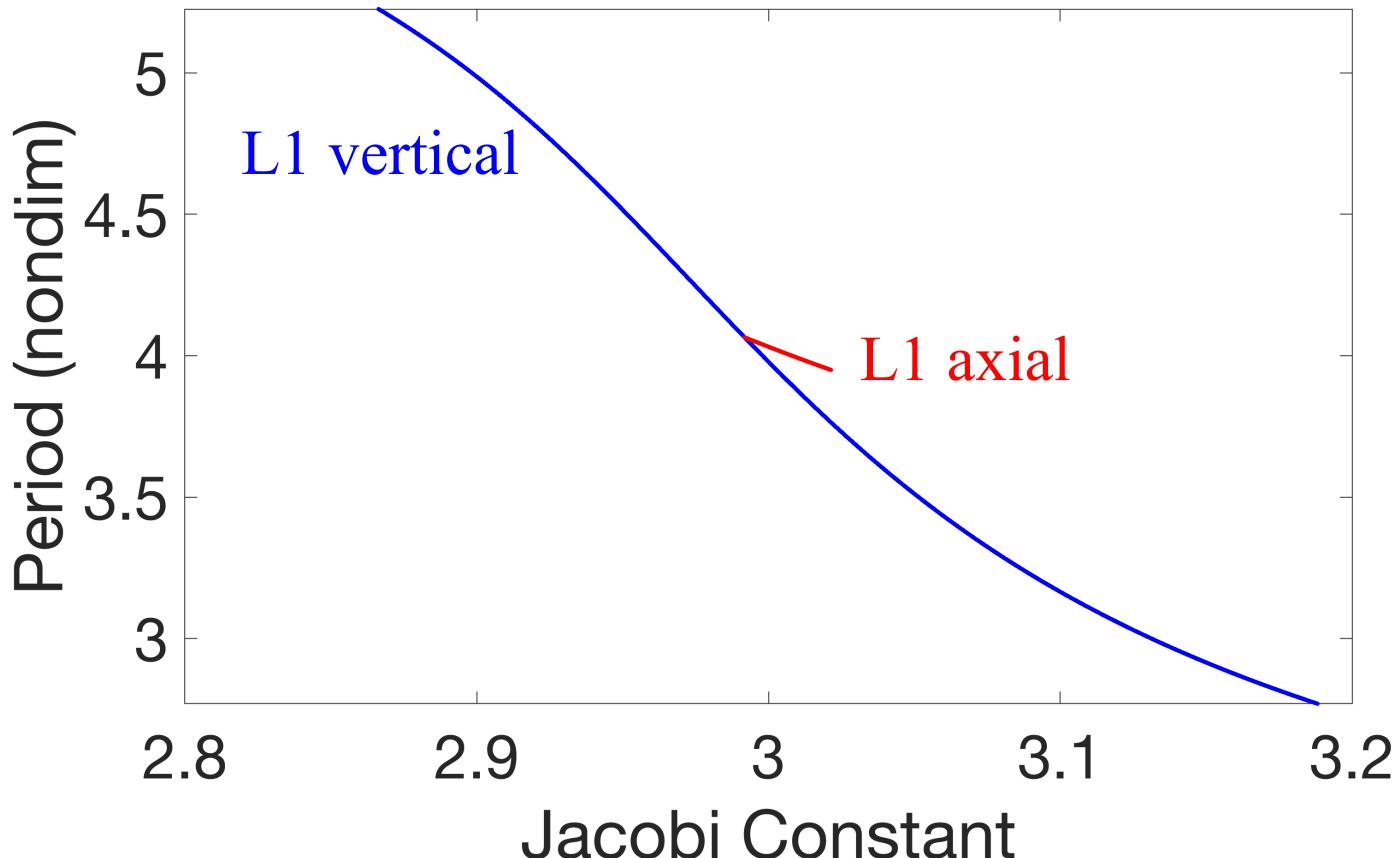
Bifurcations along L_1 Lyapunov orbit family correspond to intersections with L_1 halo and axial orbit families



Left image credit: Grebow, D. 2006, "GENERATING PERIODIC ORBITS IN THE CIRCULAR RESTRICTED THREE-BODY PROBLEM WITH APPLICATIONS TO LUNAR SOUTH POLE COVERAGE", MS Thesis

Periodic Orbit Families near L_1

Bifurcation along L_1 vertical orbit family correspond to intersection with L_1 axial orbit family



EM Libration Point Orbit Families

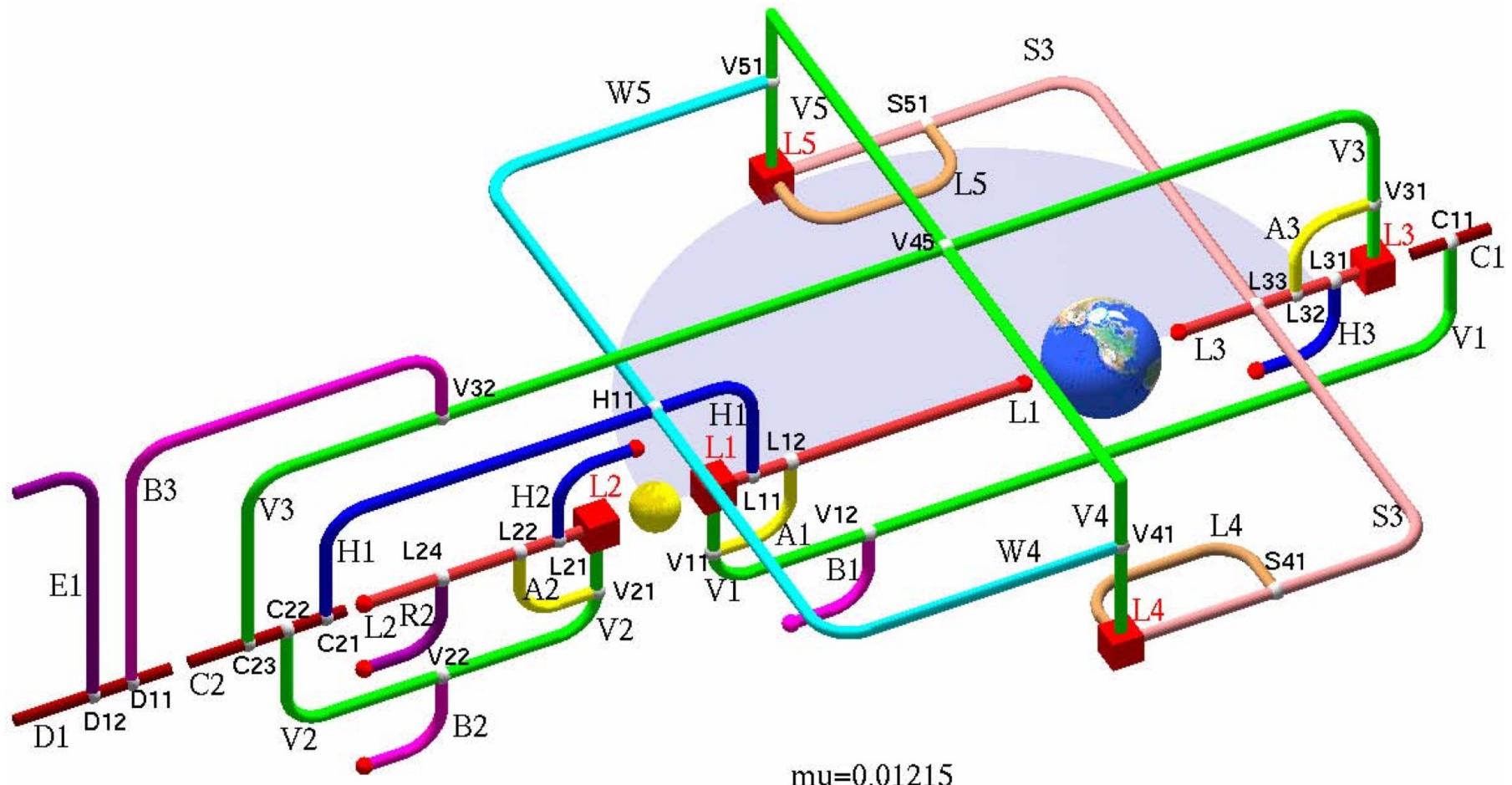


Image credit: Doedel et. al., 2007, “Elemental Periodic Orbits Associated with the Libration Points in the Circular Restricted Three-Body Problem” (See Canvas for link to paper)

Impact of Mass Ratio

Existence of some families depends on mass ratio

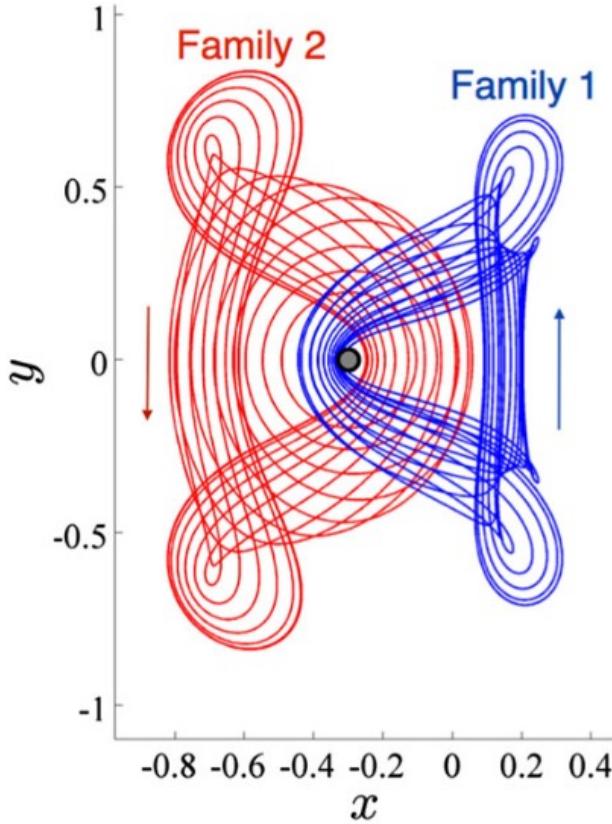


Fig. 11 Stability representation for ‘family 1’, comprised of prograde orbits about P_1 . Orbital stability is indicated via color: stable (blue), positive unstable (red), and negative unstable (purple). **a** In-plane stability. **b** Out-of-plane stability

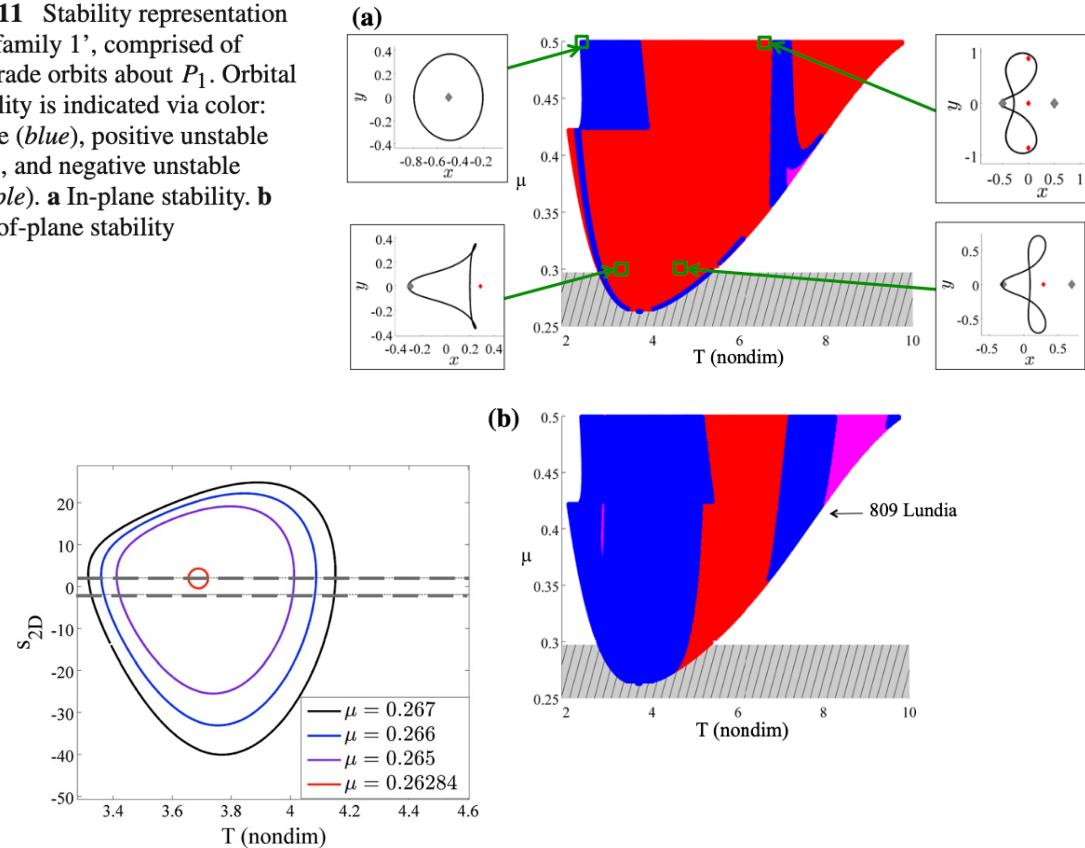


Image credits: Bosanac, N; Howell, K.C; Fischbach, E, 'Stability of Orbits Near Large Mass Ratio Binary Systems,' May 2015, Volume 122, pp. 27-52, Celestial Mechanics and Dynamical Astronomy.

Impact of Mass Ratio

Families can significantly change as mass ratio changes

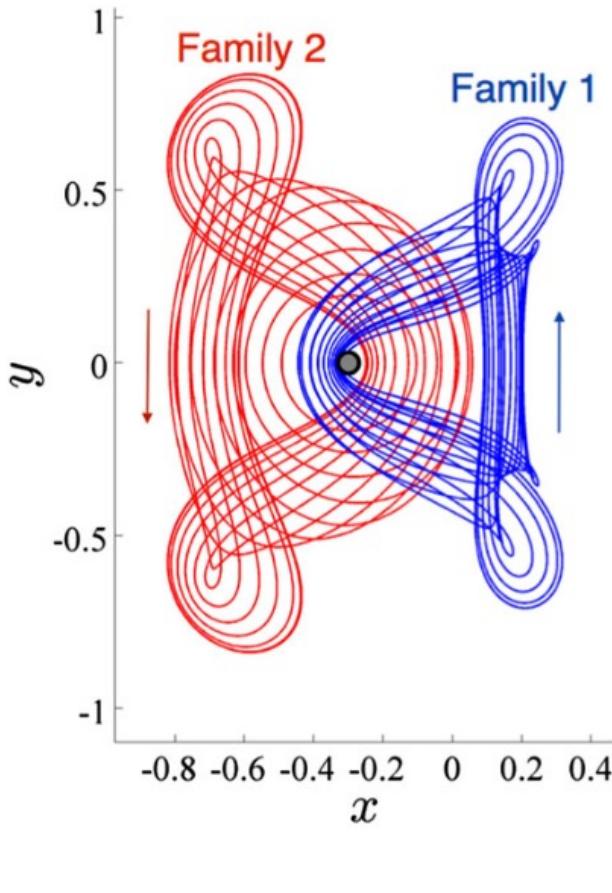


Fig. 15 Stability representation for ‘family 2’, consisting of prograde orbits about P_1 . Orbital stability is indicated via color: stable (blue), positive unstable (red), and negative unstable (purple). **a** In-plane stability. **b** Out-of-plane stability

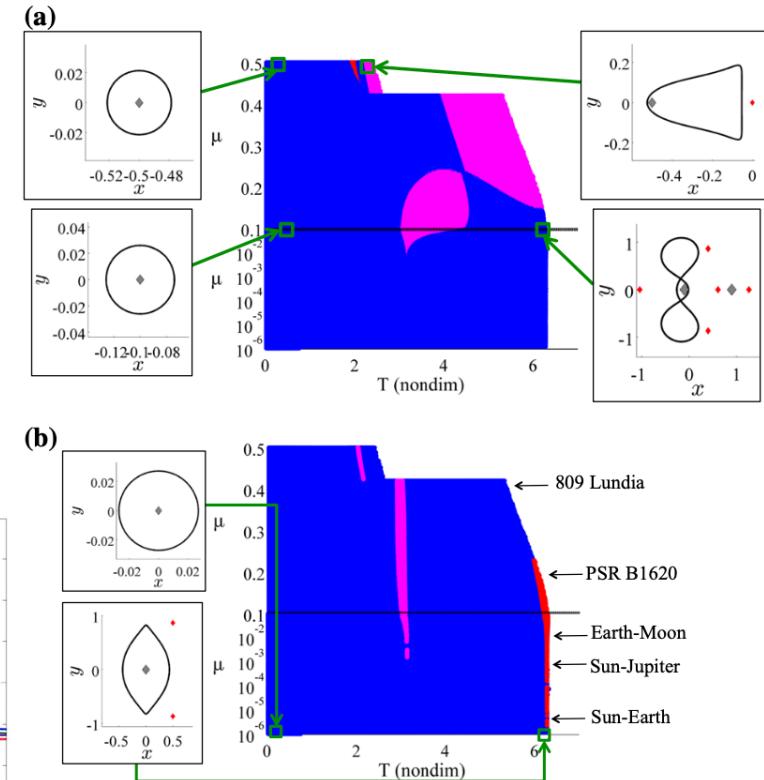
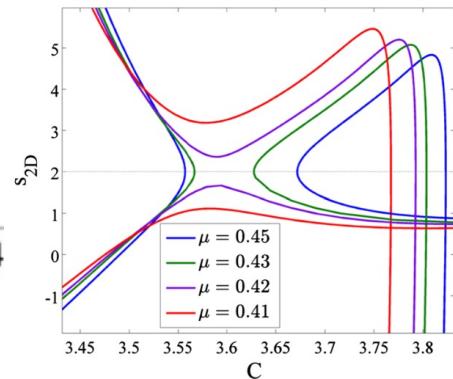


Image credits: Bosanac, N; Howell, K.C; Fischbach, E, 'Stability of Orbits Near Large Mass Ratio Binary Systems,' May 2015, Volume 122, pp. 27-52, Celestial Mechanics and Dynamical Astronomy.