

# Dynamics of a System

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ASEN 5010

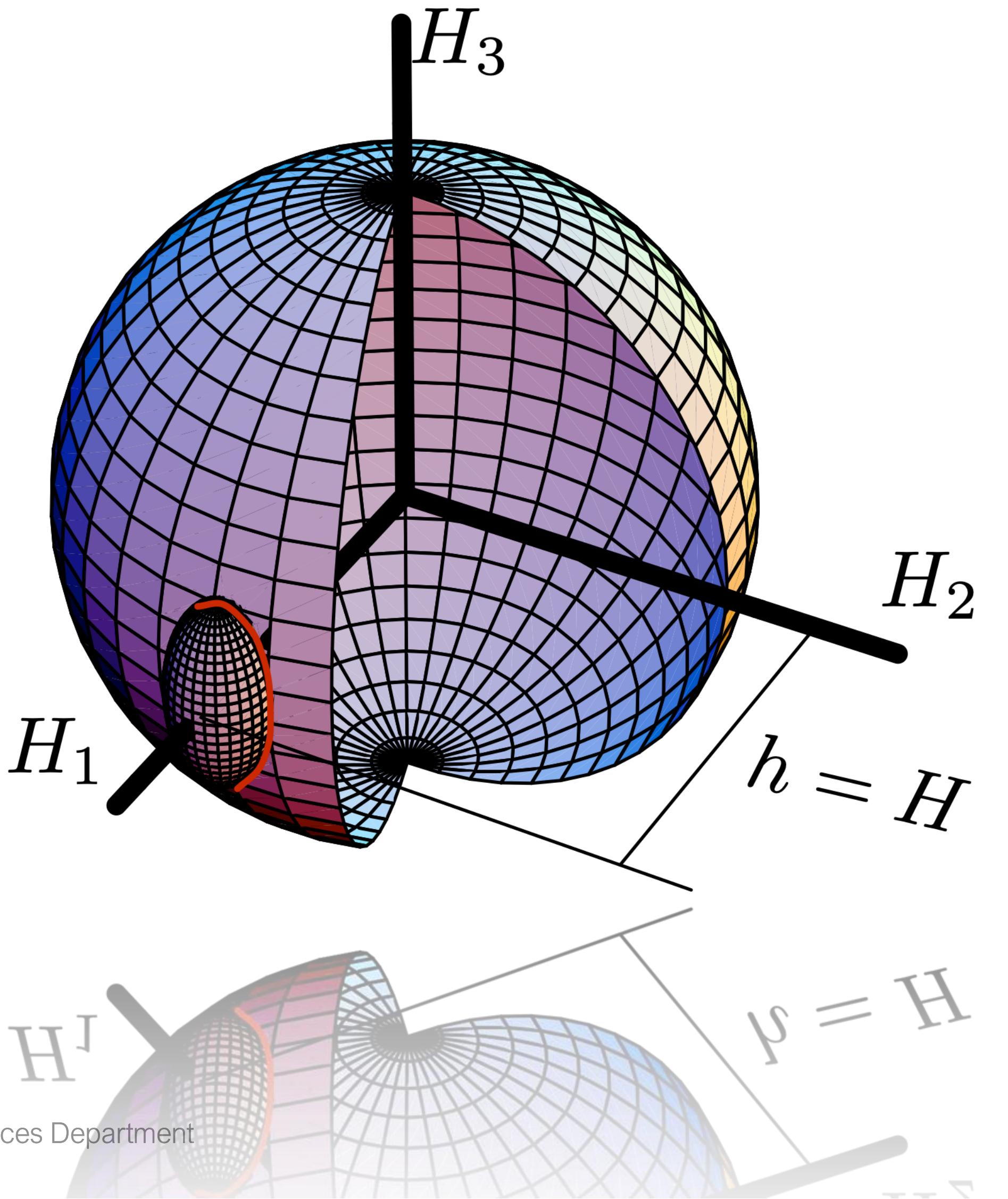
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# Outline

- Continuous System
- Rigid Body
  - Inertia tensor
  - Energy
  - Equations of motion
  - Spin Stability
- Dual Spin Spacecraft
- Gravity Gradient Torque
  - zero GG Torque attitudes
  - Equilibrium stabilities



# Continuous System

What does jello look like in space?



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# Equations of Motion

Newton's Law:

$$d\mathbf{F} = \ddot{\mathbf{R}} dm$$

Force Vector:

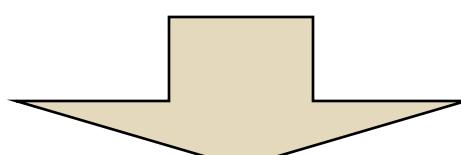
$$d\mathbf{F} = d\mathbf{F}_E + d\mathbf{F}_I$$

Total Force acting  
on System:

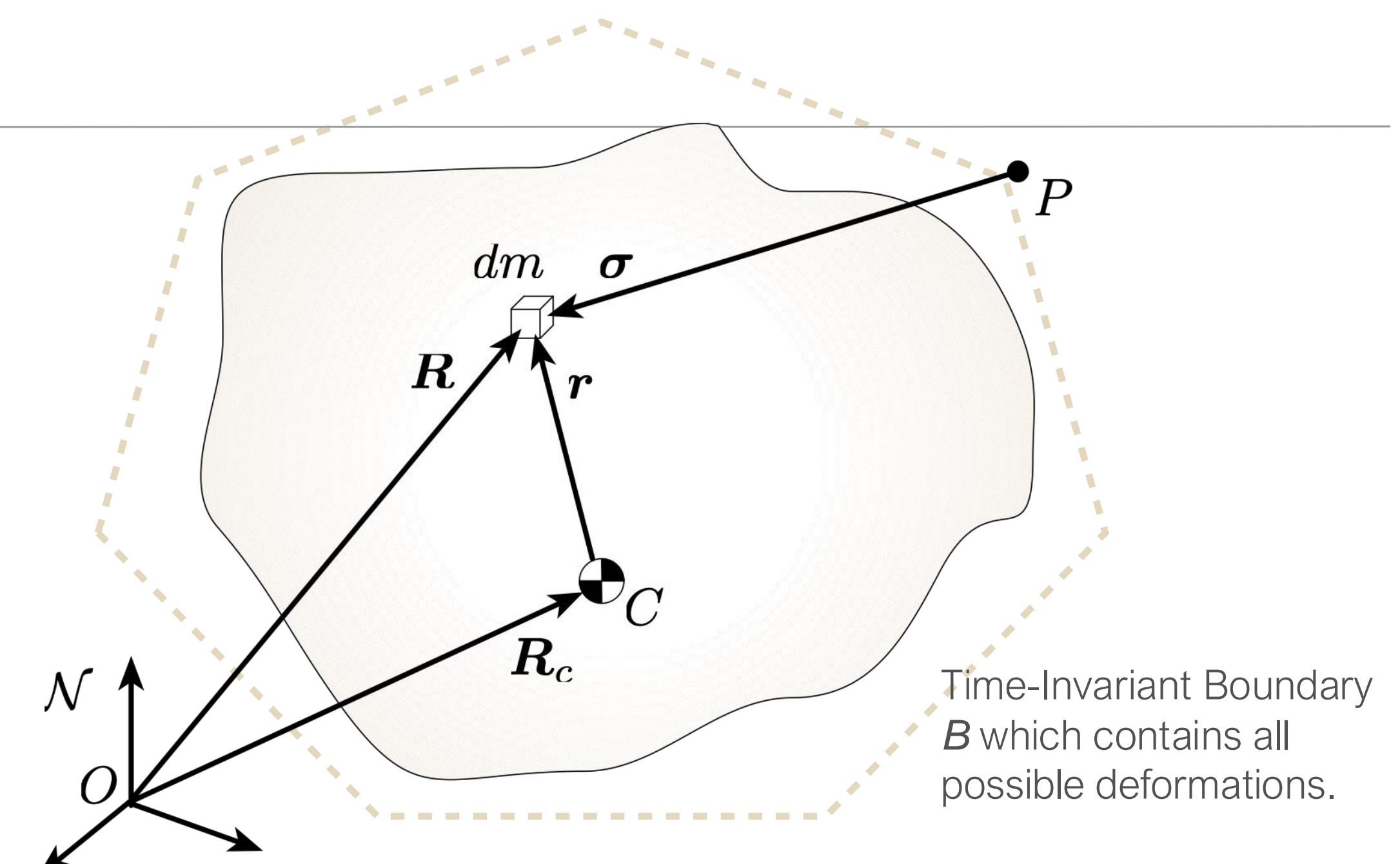
$$\mathbf{F} = \int_B d\mathbf{F} = \int_B d\mathbf{F}_E$$

Center of Mass:

$$M\mathbf{R}_c = \int_B \mathbf{R} dm = \int_{\mathcal{B}} (\mathbf{R}_c + \mathbf{r}) dm \rightarrow \int_{\mathcal{B}} \mathbf{r} dm = \mathbf{0}$$



$$M\ddot{\mathbf{R}}_c = \int_B \ddot{\mathbf{R}} dm = \int_B d\mathbf{F}$$



$$M\ddot{\mathbf{R}}_c = \mathbf{F}$$

Super Particle Theorem



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# Kinetic Energy

Definition:

$$T = \frac{1}{2} \int_B \dot{\mathbf{R}} \cdot \dot{\mathbf{R}} dm$$

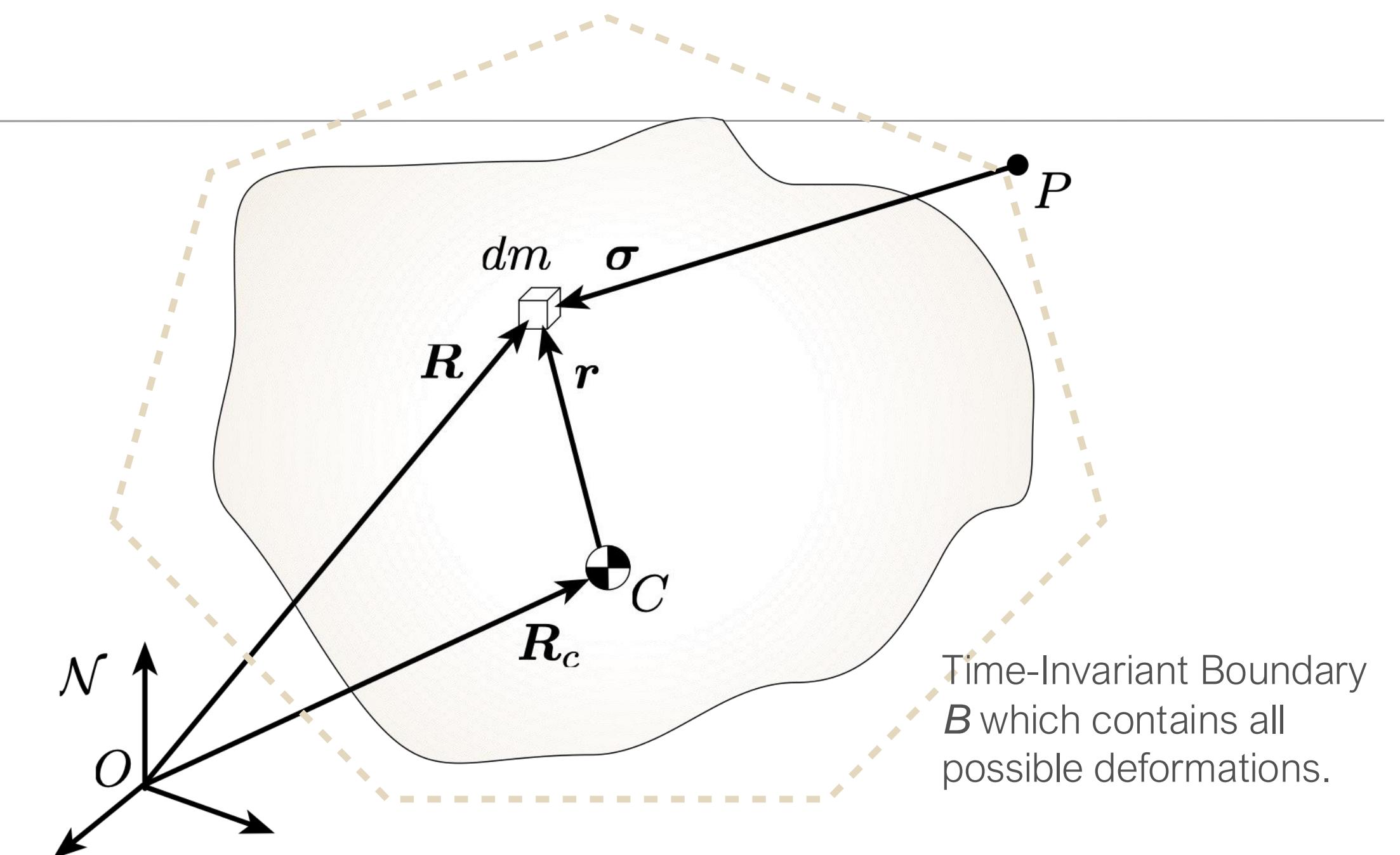
$$\dot{\mathbf{R}} = \dot{\mathbf{R}}_c + \dot{\mathbf{r}}$$

$$T = \frac{1}{2} \left( \int_B dm \right) \dot{\mathbf{R}}_c \cdot \dot{\mathbf{R}}_c + \cancel{\dot{\mathbf{R}}_c \cdot \cancel{\dot{\mathbf{r}} dm}}$$
$$+ \frac{1}{2} \int_B \dot{\mathbf{r}} \cdot \dot{\mathbf{r}} dm$$

$$T = \frac{1}{2} M \dot{\mathbf{R}}_c \cdot \dot{\mathbf{R}}_c + \frac{1}{2} \int_B \dot{\mathbf{r}} \cdot \dot{\mathbf{r}} dm$$

Energy of CM

Energy about CM



# Work/Energy Principle

Differentiate Energy:

$$\frac{dT}{dt} = M \ddot{\mathbf{R}}_c \cdot \dot{\mathbf{R}}_c + \int_B \dot{\mathbf{r}} \cdot \ddot{\mathbf{r}} dm$$

$$M \ddot{\mathbf{R}}_c = \mathbf{F}$$
$$\ddot{\mathbf{r}} = \ddot{\mathbf{R}} - \ddot{\mathbf{R}}_c$$

$$\frac{dT}{dt} = \mathbf{F} \cdot \dot{\mathbf{R}}_c + \int_B (\ddot{\mathbf{R}} dm \cdot \dot{\mathbf{r}}) - \ddot{\mathbf{R}}_c \cdot \cancel{\int_B \dot{\mathbf{r}} dm}$$

C.M.

$$\frac{dT}{dt} = \mathbf{F} \cdot \dot{\mathbf{R}}_c + \int_B d\mathbf{F} \cdot \dot{\mathbf{r}}$$

$$T(t_2) - T(t_1) = \int_{\mathbf{R}(t_1)}^{\mathbf{R}(t_2)} \mathbf{F} \cdot d\dot{\mathbf{R}}_c dt + \int_{t_1}^{t_2} \int_{\mathbf{r}(tB)}^{\mathbf{r}(t_2)} d\mathbf{F} \cdot d\mathbf{r} dt$$

Work energy/principle for system of particles

# Linear Momentum

Definition:

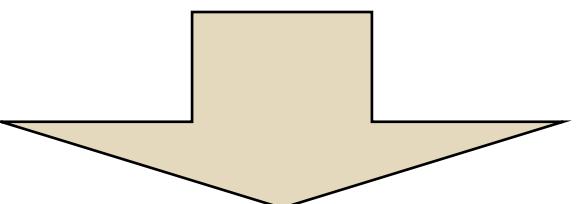
$$dp = \dot{R}dm$$

$$p = \int_{\mathcal{B}} dp = \int_{\mathcal{B}} \dot{R}dm = \int_{\mathcal{B}} (\dot{R}_c + \dot{r})dm = \left( \int_{\mathcal{B}} dm \right) \dot{R}_c + \cancel{\int_{\mathcal{B}} \dot{r}dm}$$

$$p = M \dot{R}_c$$

Linear Momentum  
Rate:

$$\dot{p} = \int_{\mathcal{B}} \ddot{R}dm = \int_{\mathcal{B}} dF = F$$



$$F = \frac{\mathcal{N}_d}{dt} (p)$$



# Angular Momentum

Ang. Momentum about  $P$ :  $\mathbf{H}_P = \int_B \boldsymbol{\sigma} \times \dot{\boldsymbol{\sigma}} dm$

$$\boldsymbol{\sigma} = \mathbf{R} - \mathbf{R}_P$$

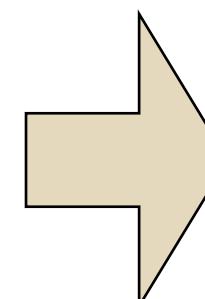
Inertial Time Derivative:  $\dot{\mathbf{H}}_P = \cancel{\int_B \dot{\boldsymbol{\sigma}} \times \dot{\boldsymbol{\sigma}} dm} + \int_B \boldsymbol{\sigma} \times \ddot{\boldsymbol{\sigma}} dm$

$$\dot{\mathbf{H}}_P = \boxed{\int_B \boldsymbol{\sigma} \times \ddot{\mathbf{R}} dm} - \boxed{\left( \int_B \boldsymbol{\sigma} dm \right)} \times \ddot{\mathbf{R}}_P$$

$$\int_B \boldsymbol{\sigma} dm = \int_B \mathbf{R} dm - \left( \int_B dm \right) \mathbf{R}_P = \boxed{M(\mathbf{R}_c - \mathbf{R}_P)}$$

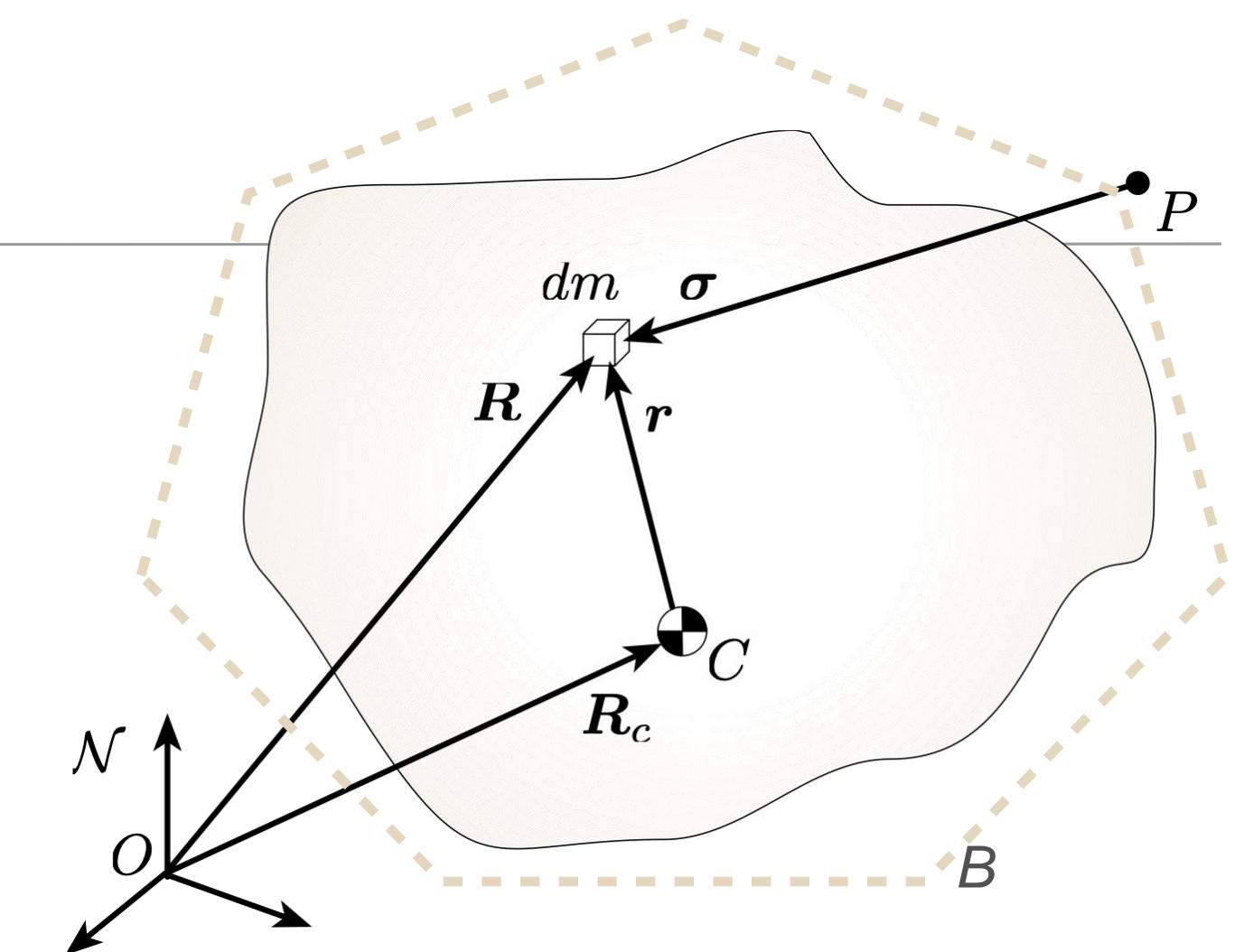
Torque about  $P$ :  $\mathbf{L}_P = \boxed{\int_B \boldsymbol{\sigma} \times \ddot{\mathbf{R}} dm} = \int_B \boldsymbol{\sigma} \times d\mathbf{F}$

$$\dot{\mathbf{H}}_P = \mathbf{L}_P + M \ddot{\mathbf{R}}_P \times (\mathbf{R}_c - \mathbf{R}_P)$$



$$\boxed{\dot{\mathbf{H}}_P = \mathbf{L}_P}$$

If  $P$  is CM or Inertial



# Rigid Body Dynamics

The 101 of spacecraft dynamics...



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# General Angular Momentum

Definition:

$$\mathbf{H}_O = \int_B \mathbf{R} \times \dot{\mathbf{R}} dm$$

or

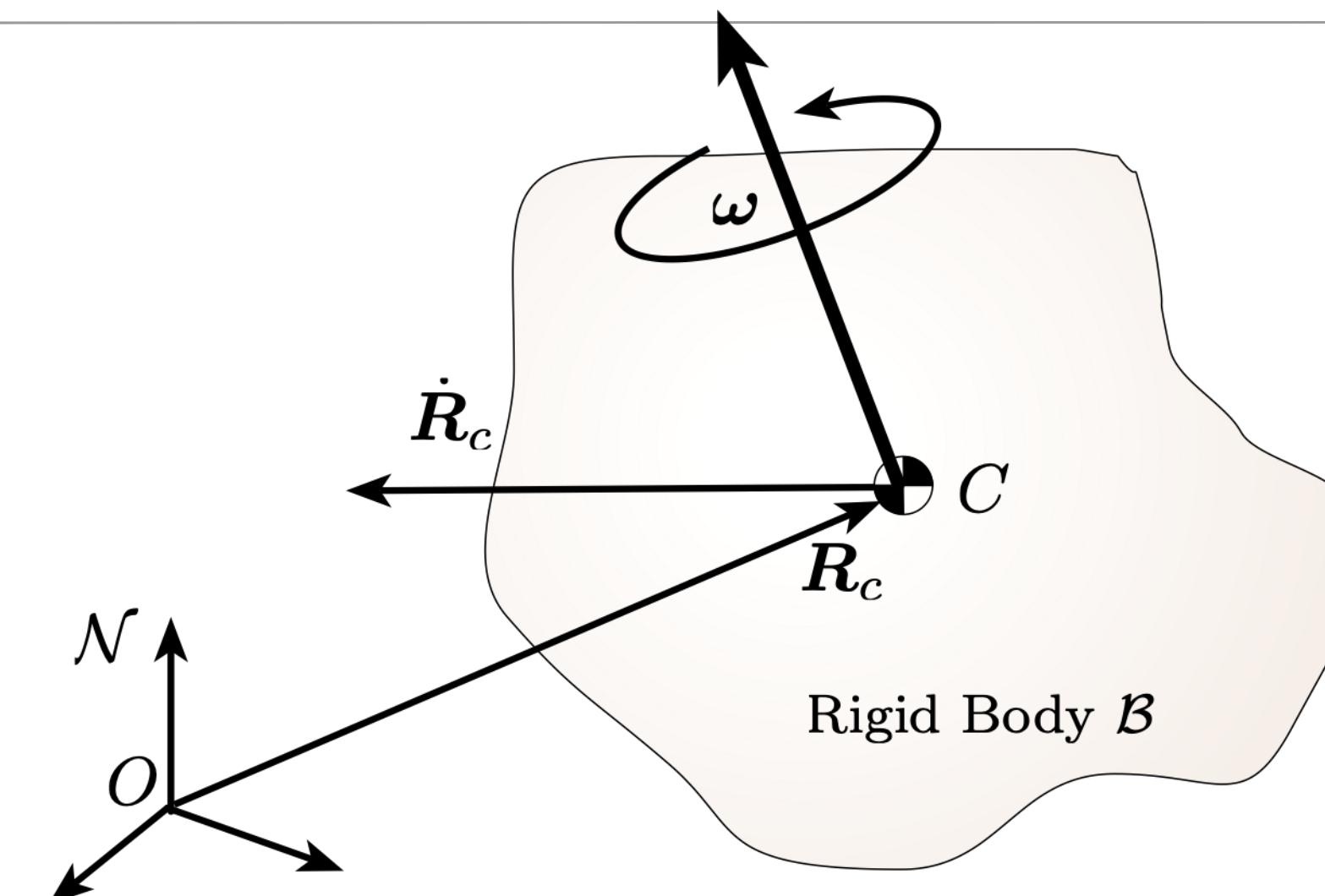
$$\mathbf{H}_O = \mathbf{R}_c \times M \dot{\mathbf{R}}_c + \int_B \mathbf{r} \times \dot{\mathbf{r}} dm$$

Momentum about CM:

$$\mathbf{H}_c = \int_B \mathbf{r} \times \dot{\mathbf{r}} dm$$

$$\dot{\mathbf{r}} = \frac{d\mathbf{r}}{dt} = \cancel{\frac{d\mathbf{r}}{dt}} + \omega \times \mathbf{r} = \omega \times \mathbf{r}$$

$$\mathbf{H}_c = \int_B \mathbf{r} \times (\omega \times \mathbf{r}) dm = \left( \int_B -[\tilde{\mathbf{r}}][\tilde{\mathbf{r}}] dm \right) \omega$$



# Inertia Tensor Properties

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Definition:

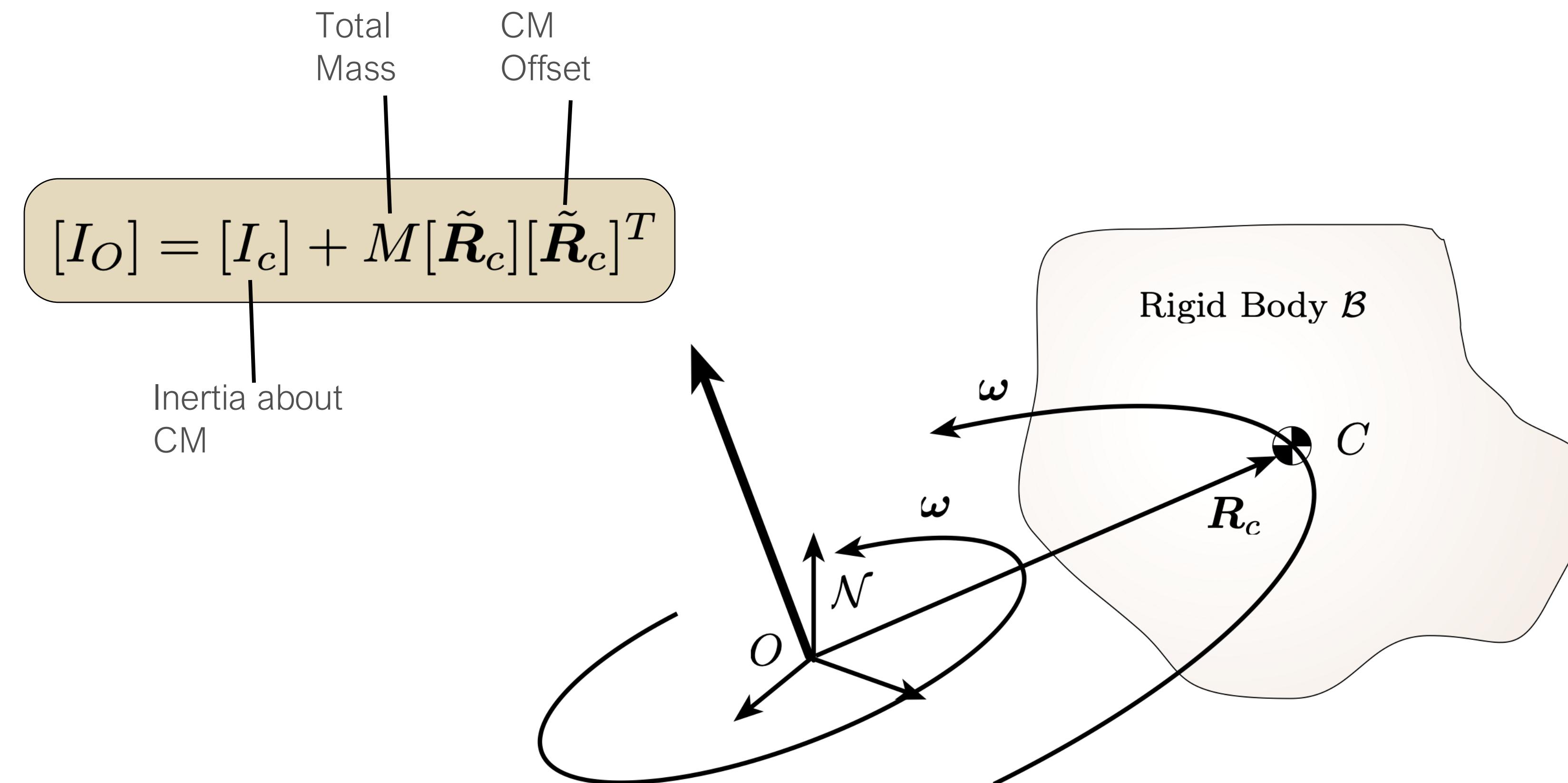
$$\mathcal{B}[I_c] = \int_B -[\tilde{\mathbf{r}}][\tilde{\mathbf{r}}]dm = \int_B \begin{bmatrix} r_2^2 + r_3^2 & -r_1r_2 & -r_1r_3 \\ -r_1r_2 & r_1^2 + r_3^2 & -r_2r_3 \\ -r_1r_3 & -r_2r_3 & r_1^2 + r_2^2 \end{bmatrix} dm$$

Angular Momentum Expression:

$$\mathbf{H}_c = \begin{pmatrix} \mathcal{B}(H_{c_1}) \\ \mathcal{B}(H_{c_2}) \\ \mathcal{B}(H_{c_3}) \end{pmatrix} = \int_B \begin{bmatrix} r_2^2 + r_3^2 & -r_1r_2 & -r_1r_3 \\ -r_1r_2 & r_1^2 + r_3^2 & -r_2r_3 \\ -r_1r_3 & -r_2r_3 & r_1^2 + r_2^2 \end{bmatrix} \begin{pmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix} dm = [I_c]\boldsymbol{\omega}$$



# Parallel Axis Theorem



# Coordinate Transformation

$$\mathcal{F}[I] = [FB]^{\mathcal{B}}[I][FB]^T$$

$B$  - Body Frame

$F$  - 2<sup>nd</sup> Coordinate Frame

$$\begin{bmatrix} C_{11} & C_{12} & C_{13} \\ C_{21} & C_{22} & C_{23} \\ C_{31} & C_{32} & C_{33} \end{bmatrix}^{\mathcal{B}} \begin{bmatrix} I_{11} & I_{12} & I_{13} \\ I_{12} & I_{22} & I_{23} \\ I_{13} & I_{23} & I_{33} \end{bmatrix} = \begin{bmatrix} I_1 & 0 & 0 \\ 0 & I_2 & 0 \\ 0 & 0 & I_3 \end{bmatrix} \begin{bmatrix} C_{11} & C_{12} & C_{13} \\ C_{21} & C_{22} & C_{23} \\ C_{31} & C_{32} & C_{33} \end{bmatrix}$$

Rotation matrix  $[C]$   
contains the  
eigenvectors of  $[I]$

$$[C] = [V]^T = \begin{bmatrix} \mathbf{v}_1^T \\ \mathbf{v}_2^T \\ \mathbf{v}_3^T \end{bmatrix}$$

Principal inertia  
matrix whose  
diagonal entries  
are the  
eigenvalues of  $[I]$



# Kinetic Energy

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Total Energy:

$$T = \frac{1}{2} M \dot{\mathbf{R}}_c \cdot \dot{\mathbf{R}}_c + \frac{1}{2} \int_B \dot{\mathbf{r}} \cdot \dot{\mathbf{r}} dm = T_{\text{trans}} + T_{\text{rot}}$$

Rotational Energy:

$$T_{\text{rot}} = \frac{1}{2} \int_B \dot{\mathbf{r}} \cdot \dot{\mathbf{r}} dm = \frac{1}{2} \int_B (\boldsymbol{\omega} \times \mathbf{r}) \cdot (\boldsymbol{\omega} \times \mathbf{r}) dm$$

$$T_{\text{rot}} = \frac{1}{2} \boldsymbol{\omega} \cdot \int_B \mathbf{r} \times (\boldsymbol{\omega} \times \mathbf{r}) dm = \frac{1}{2} \boldsymbol{\omega} \cdot \mathbf{H}_c = \frac{1}{2} \boldsymbol{\omega}^T [I] \boldsymbol{\omega}$$

Energy Rate:

$$\dot{T} = \mathbf{F} \cdot \dot{\mathbf{R}}_c + \mathbf{L}_c \cdot \boldsymbol{\omega}$$

Work/Energy Principle:

$$W = T(t_2) - T(t_1) = \int_{t_1}^{t_2} \mathbf{F} \cdot \dot{\mathbf{R}}_c dt + \int_{t_1}^{t_2} \mathbf{L}_c \cdot \boldsymbol{\omega} dt$$



# Equations of Motion

Euler's equation:

$$\dot{\mathbf{H}}_c = \frac{\mathcal{B}_d}{dt} (\mathbf{H}_c) + \boldsymbol{\omega} \times \mathbf{H}_c = \mathbf{L}_c$$

$$\frac{\mathcal{B}_d}{dt} (\mathbf{H}_c) = \frac{\mathcal{B}_d}{dt} ([I]) \boldsymbol{\omega} + [I] \frac{\mathcal{B}_d}{dt} (\boldsymbol{\omega}) = [I] \dot{\boldsymbol{\omega}}$$

Euler's rotational  
equations of motion:

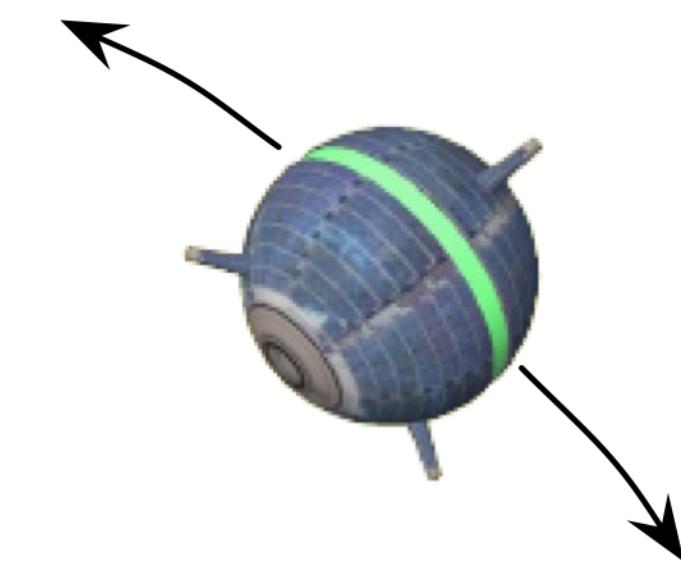
$$[I] \dot{\boldsymbol{\omega}} = -[\tilde{\boldsymbol{\omega}}][I] \boldsymbol{\omega} + \mathbf{L}_c$$

Principal axis version of  
rotational EOM:

$$I_{11} \dot{\omega}_1 = -(I_{33} - I_{22}) \omega_2 \omega_3 + L_1$$

$$I_{22} \dot{\omega}_2 = -(I_{11} - I_{33}) \omega_3 \omega_1 + L_2$$

$$I_{33} \dot{\omega}_3 = -(I_{22} - I_{11}) \omega_1 \omega_2 + L_3$$



Discuss how to integrate full EOM

# Example: Slender Rod Falling

Rod Inertia about CM:  $I_c = \frac{m}{12}L^2$

Momentum about CM:  $\mathbf{H}_c = I_c \dot{\theta} \hat{\mathbf{e}}_3$

Torque:

$$\mathbf{L}_c = \left( -\frac{L}{2} \hat{\mathbf{e}}_L \right) \times N \hat{\mathbf{n}}_2 = \frac{L}{2} N \sin \theta \hat{\mathbf{e}}_3$$

Euler's Eqn:

$$\dot{\mathbf{H}}_c = \mathbf{L}_c$$

$$\frac{m}{12} L^2 \ddot{\theta} - \frac{L}{2} N \sin \theta = 0$$

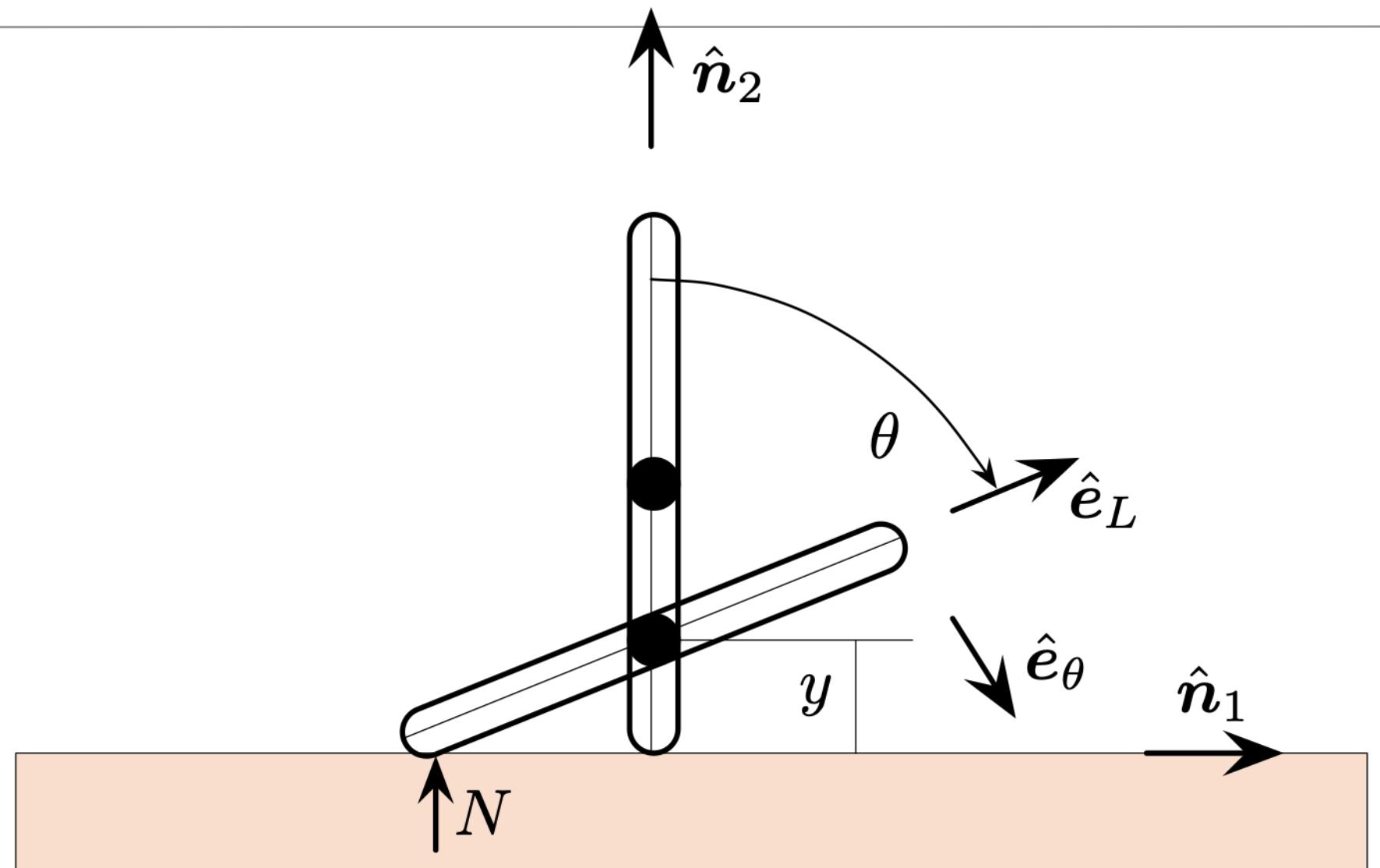
Newton's Eqn:

$$m \ddot{y} \hat{\mathbf{n}}_2 = (N - mg) \hat{\mathbf{n}}_2$$

$$y = \frac{L}{2} \cos \theta \quad \ddot{y} = -\frac{L}{2} \ddot{\theta} \sin \theta - \frac{L}{2} \dot{\theta}^2 \cos \theta$$

EOM:

$$\boxed{\frac{m}{12} L^2 \ddot{\theta} (1 + 3 \sin^2 \theta) + \frac{m}{4} L^2 \dot{\theta}^2 \sin \theta \cos \theta - \frac{m}{2} L g \sin \theta = 0}$$



$$N = mg - m \frac{L}{2} \ddot{\theta} \sin \theta - m \frac{L}{2} \dot{\theta}^2 \cos \theta$$



# Example: Slender Rod Falling

Energy functions:

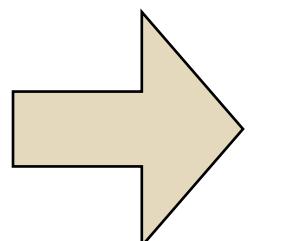
$$V(\theta) = mgy = mg\frac{L}{2} \cos \theta$$

$$T(\theta, \dot{\theta}) = \frac{m}{2}\dot{y}^2 + \frac{I_c}{2}\dot{\theta}^2 = \frac{mL^2}{24}(1 + 3\sin^2 \theta)\dot{\theta}^2$$

Initial energy level:

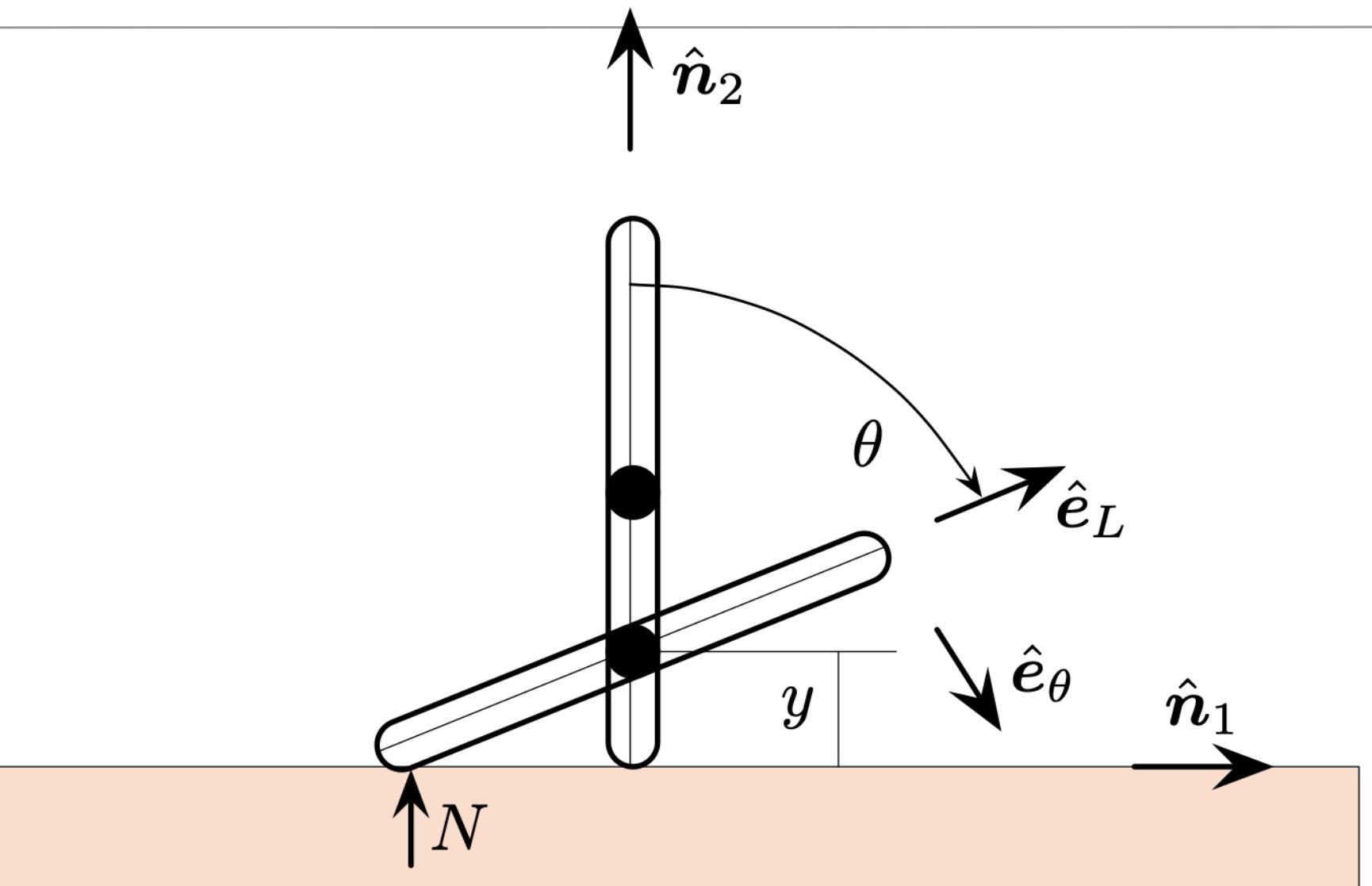
$$E(t_0) = T_0 + V_0 = mg\frac{L}{2}$$

$$T(\theta, \dot{\theta}) + V(\theta) = E(t_0)$$



$$\dot{\theta}^2 = \frac{12g(1 - \cos \theta)}{L(1 + 3\sin^2 \theta)}$$

Energy Conservation avoids having to solve ODE.



# Momentum/Energy Surfaces

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- Let's study a special class of rigid body motion where no external torque is acting on the body.
- In this case the kinetic energy and the angular momentum of the rigid body are conserved!
- In inertial frame vector components, we find

$$\begin{aligned}\mathbf{H} &= {}^N\mathbf{H} = [BN]^T {}^B\mathbf{H} \\ \mathbf{H} &= {}^N\mathbf{H} = \begin{bmatrix} c\theta_2c\theta_1 & c\theta_2s\theta_1 & -s\theta_2 \\ s\theta_3s\theta_2c\theta_1 - c\theta_3s\theta_1 & s\theta_3s\theta_2s\theta_1 + c\theta_3c\theta_1 & s\theta_3c\theta_2 \\ c\theta_3s\theta_2c\theta_1 + s\theta_3s\theta_1 & c\theta_3s\theta_2s\theta_1 - s\theta_3c\theta_1 & c\theta_3c\theta_2 \end{bmatrix} \begin{pmatrix} I_1\omega_1 \\ I_2\omega_2 \\ I_3\omega_3 \end{pmatrix} \\ \dot{\mathbf{H}} &= 0 = \mathbf{f}(\theta_1, \theta_2, \theta_3, \omega_1, \omega_2, \omega_3, \dot{\omega}_1, \dot{\omega}_2, \dot{\omega}_3)\end{aligned}$$

- Notice that the constant angular momentum condition can become very complicated and difficult to study!
- Instead of writing  $\mathbf{H}$  in inertial frame components, we chose to write it in the body frame where the inertia matrix is a constant for a rigid body.
- Assume the angular momentum vector  $\mathbf{H}$  is written in body frame components, and that principal axes were chosen for the body frame  $B$ .

$$\begin{aligned}\boldsymbol{\omega} &= \omega_1 \hat{\mathbf{b}}_1 + \omega_2 \hat{\mathbf{b}}_2 + \omega_3 \hat{\mathbf{b}}_3 \\ \mathbf{H} = {}^B\mathbf{H} &= H_1 \hat{\mathbf{b}}_1 + H_2 \hat{\mathbf{b}}_2 + H_3 \hat{\mathbf{b}}_3 \quad [I] = \begin{bmatrix} I_1 & 0 & 0 \\ 0 & I_2 & 0 \\ 0 & 0 & I_3 \end{bmatrix} \\ &= I_1 \omega_1 \hat{\mathbf{b}}_1 + I_2 \omega_2 \hat{\mathbf{b}}_2 + I_3 \omega_3 \hat{\mathbf{b}}_3\end{aligned}$$

- This allows us to write  $\mathbf{H}$  as:

$$\mathbf{H} = {}^{\mathcal{B}}\mathbf{H} = \begin{pmatrix} H_1 \\ H_2 \\ H_3 \end{pmatrix} = \begin{pmatrix} {}^{\mathcal{B}}(I_1\omega_1) \\ {}^{\mathcal{B}}(I_2\omega_2) \\ {}^{\mathcal{B}}(I_3\omega_3) \end{pmatrix}$$

- Because  $\mathbf{H}$  is constant, all possible rigid body angular velocities must lie on the surface of the following momentum ellipsoid:

$$H^2 = \mathbf{H}^T \mathbf{H} = I_1^2\omega_1^2 + I_2^2\omega_2^2 + I_3^2\omega_3^2 = \text{constant}$$

- Similarly, since the kinetic energy is conserved, all possible rigid body angular velocities must also lie on the surface of the following energy ellipsoid:

$$T = \frac{1}{2}I_1\omega_1^2 + \frac{1}{2}I_2\omega_2^2 + \frac{1}{2}I_3\omega_3^2$$

- The final admissible angular velocities will be on the intersection of these two ellipsoids.

- Using the momenta coordinates

$$H_1 = I_1\omega_1 \quad H_2 = I_2\omega_2 \quad H_3 = I_3\omega_3$$

- we can write the momentum magnitude constraint as

$$H^2 = H_1^2 + H_2^2 + H_3^2 \quad \rightarrow \text{Sphere}$$

- and the kinetic energy constraint as

$$1 = \frac{H_1^2}{2I_1T} + \frac{H_2^2}{2I_2T} + \frac{H_3^2}{2I_3T}$$

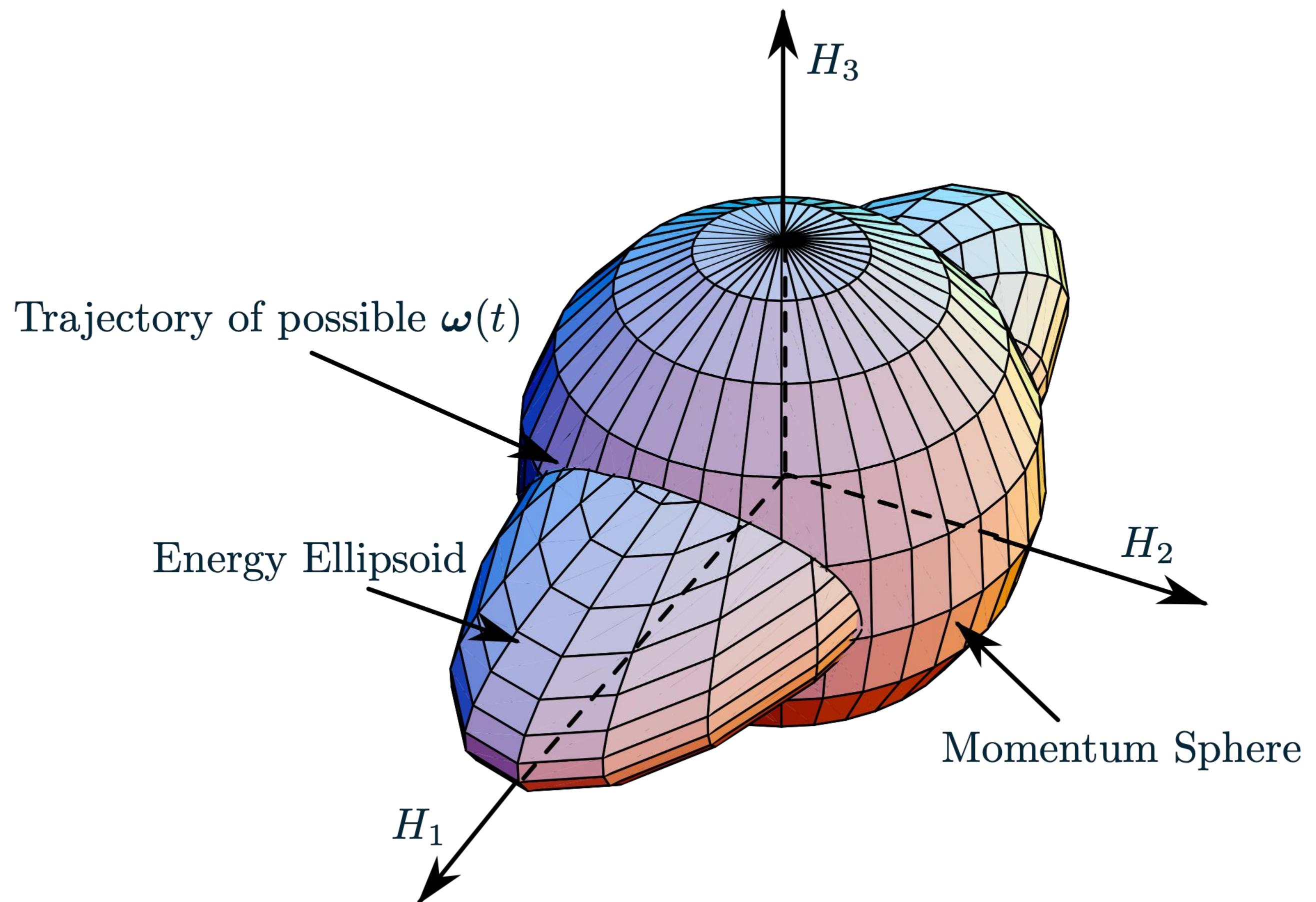
Compare to ellipsoid equation:

$$1 = \frac{x^2}{a^2} + \frac{y^2}{b^2} + \frac{z^2}{c^2}$$

$$\sqrt{2I_iT} \quad \rightarrow \text{semi-axes of ellipsoid}$$

- Clearly, for a given  $H$ , only a certain range of kinetic energies is possible.
- Let's assume the common notation:

$$I_1 \geq I_2 \geq I_3$$



- Let's look at the Minimum Energy Case:

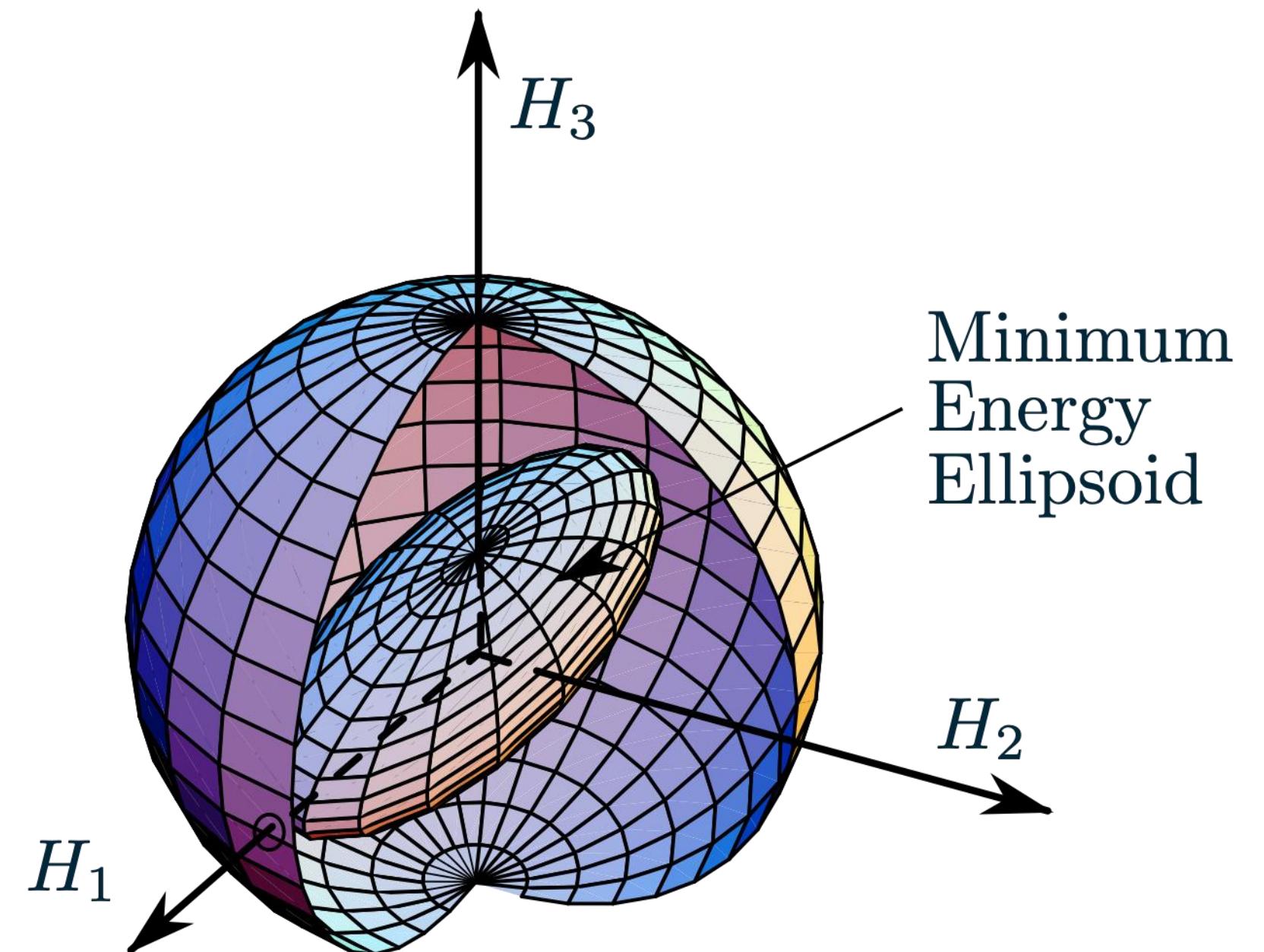
- The surfaces will only intersect at:

$$\begin{aligned} {}^B\mathbf{H} &= \pm H \hat{\mathbf{b}}_1 \\ H_1 &= H \quad H_2 = H_3 = 0 \end{aligned}$$

The kinetic energy is:

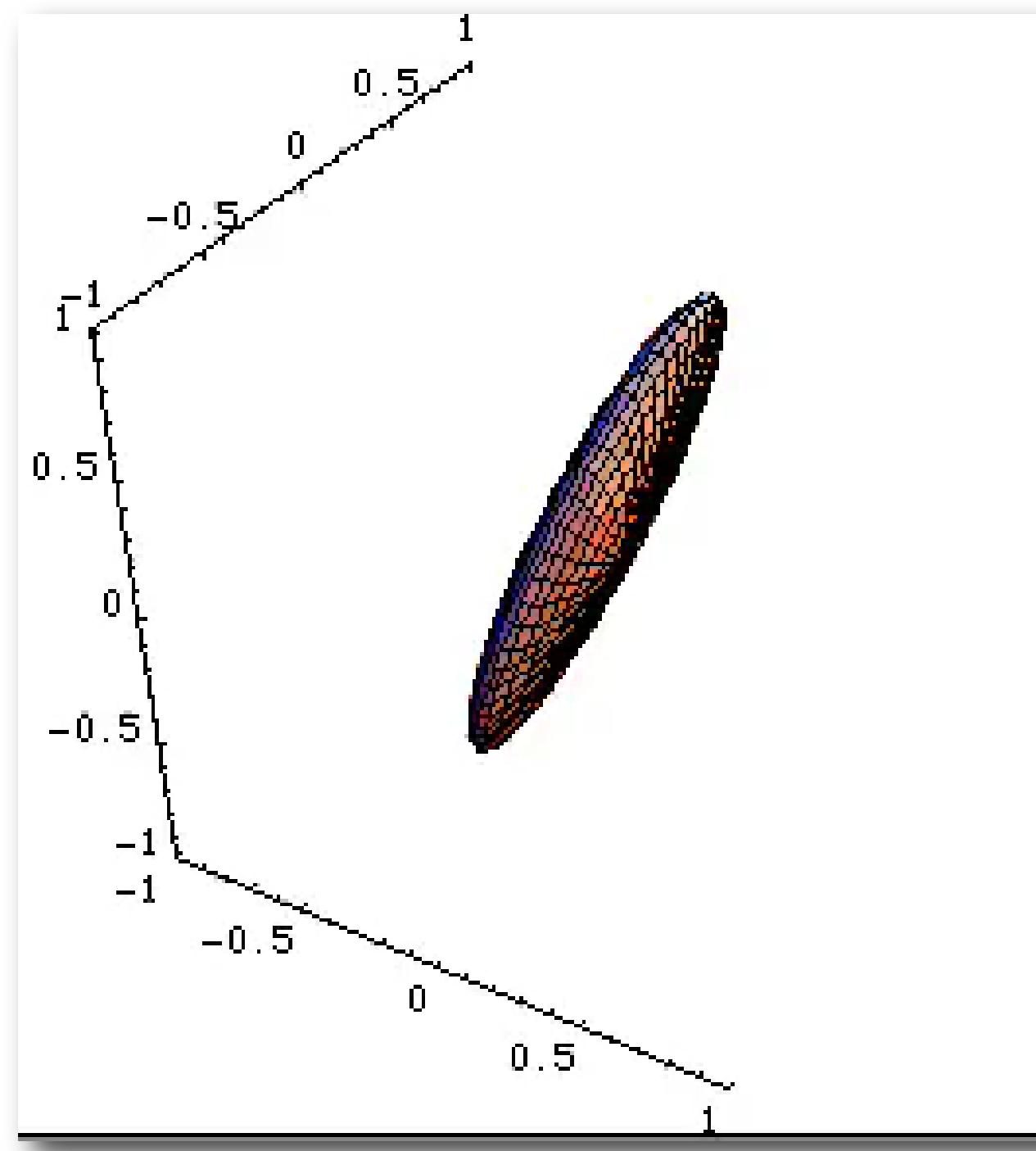
$$1 = \frac{H_1^2}{2I_1 T} + \frac{H_2^2}{2I_2 T} + \frac{H_3^2}{2I_3 T}$$

$$T_{\min} = \frac{H^2}{2I_1}$$



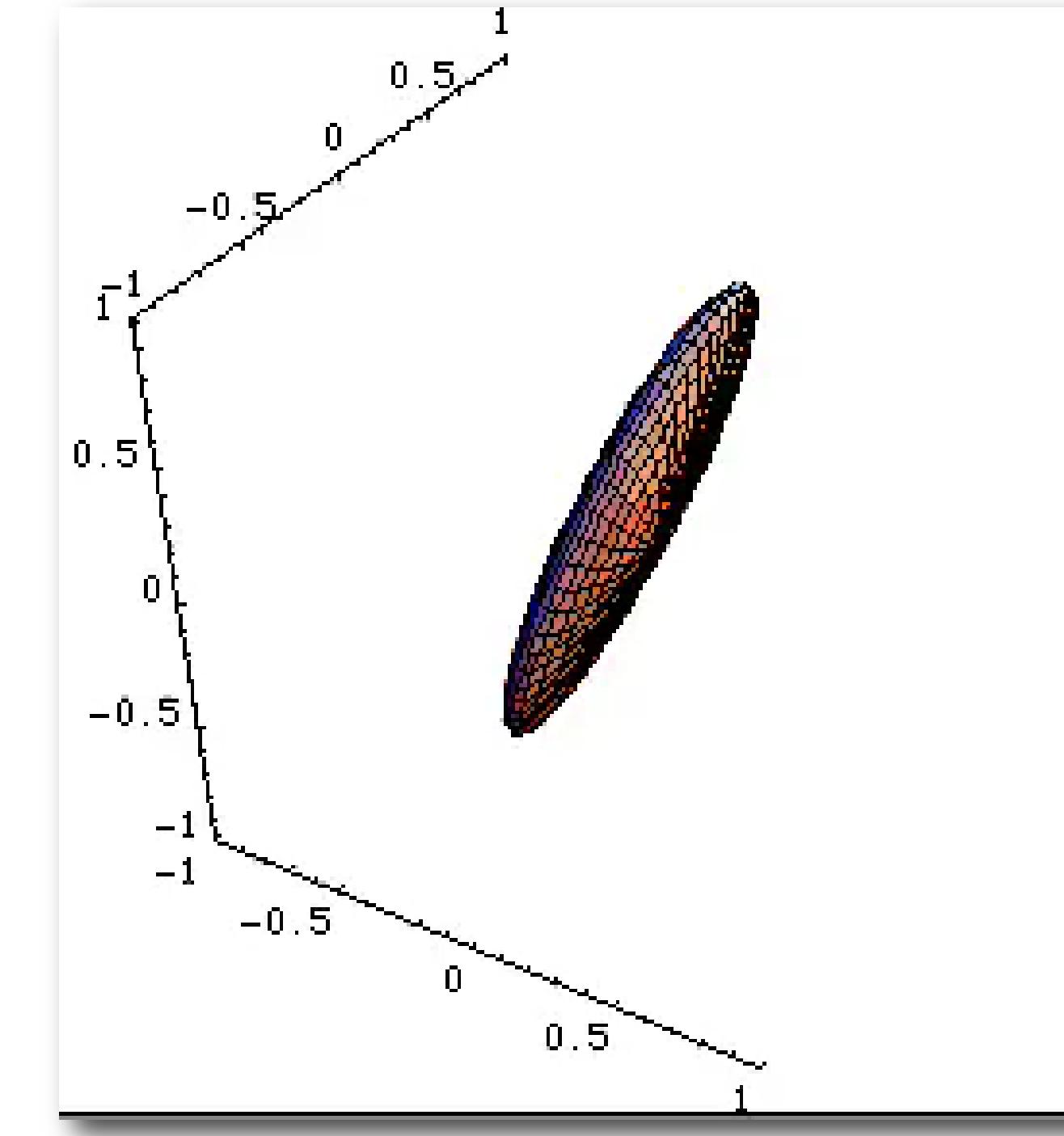
## General Rotation of Rigid Body with

$$I_1 \geq I_2 \geq I_3$$



Pure Spin

$$\boldsymbol{\omega}_0 = (10^\circ, 0^\circ, 0^\circ) / \text{s}$$



Slightly Off Spin

$$\boldsymbol{\omega}_0 = (10^\circ, 0.5^\circ, 0.5^\circ) / \text{s}$$

- Let's look at the Intermediate Energy Case:

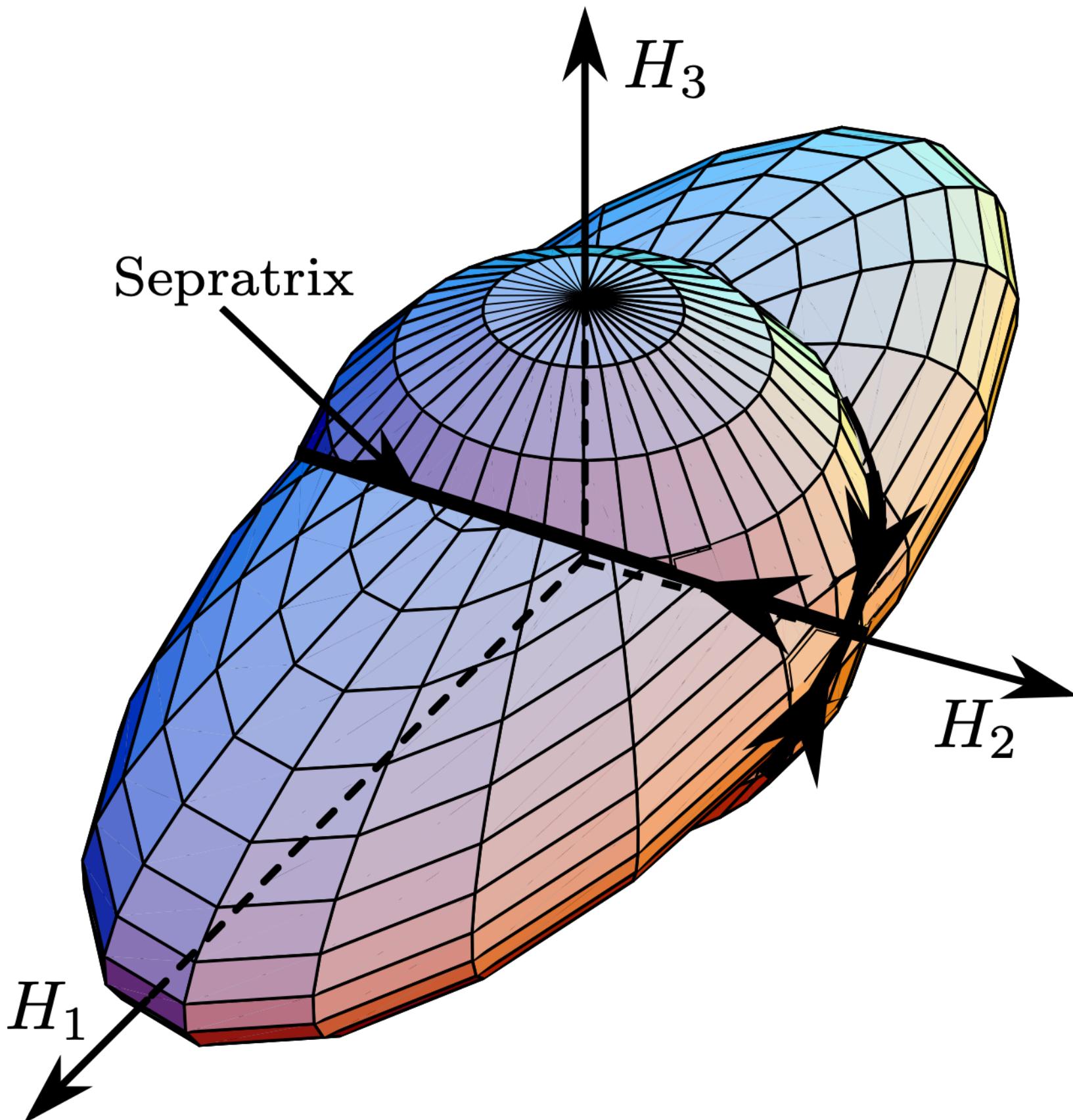
- The surfaces will only intersect at:

$${}^{\mathcal{B}}\boldsymbol{H} = \pm H \hat{\boldsymbol{b}}_2$$

The kinetic energy is:

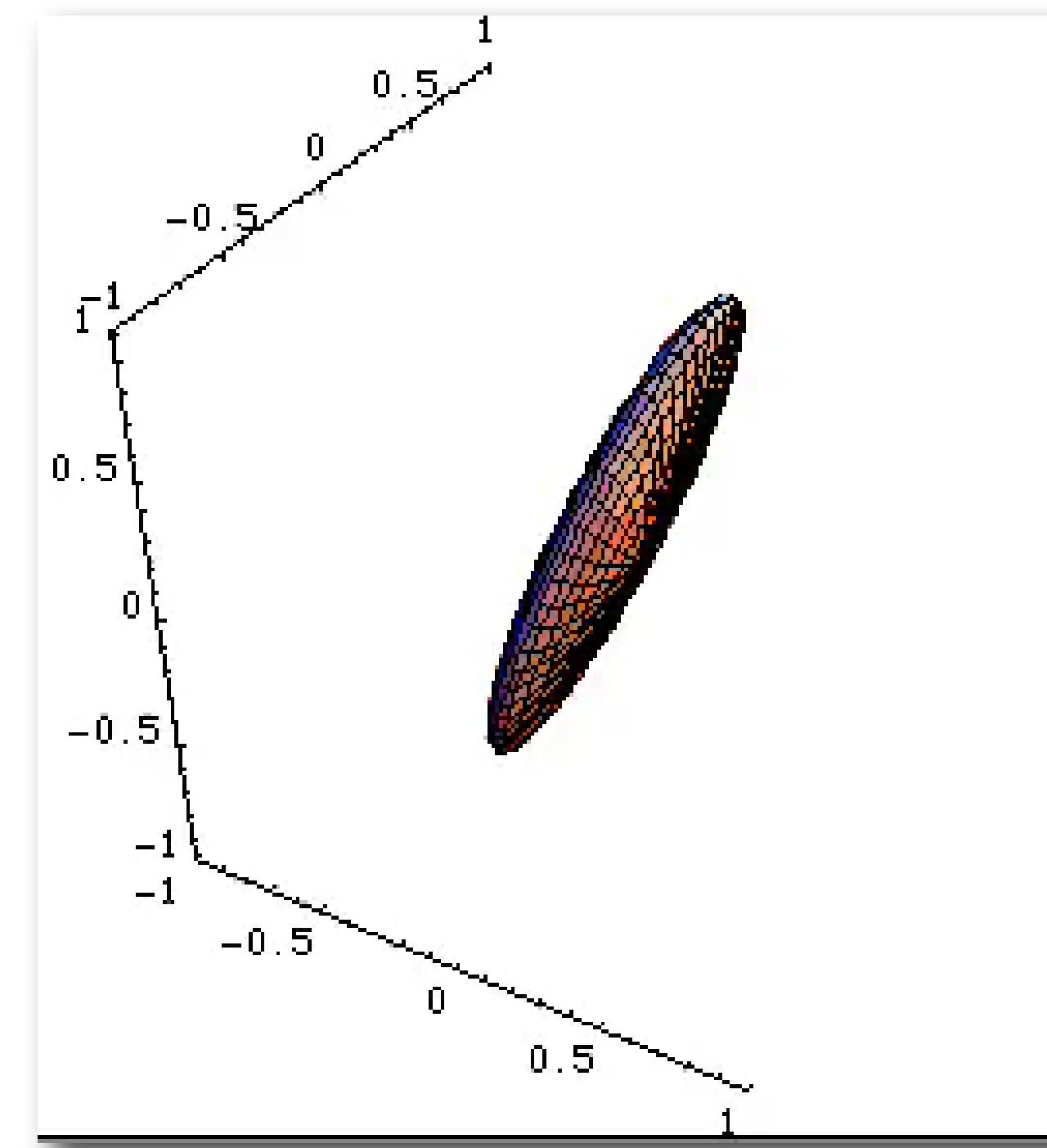
$$1 = \frac{H_1^2}{2I_1 T} + \frac{H_2^2}{2I_2 T} + \frac{H_3^2}{2I_3 T}$$

$$T_{\text{int}} = \frac{H^2}{2I_2}$$



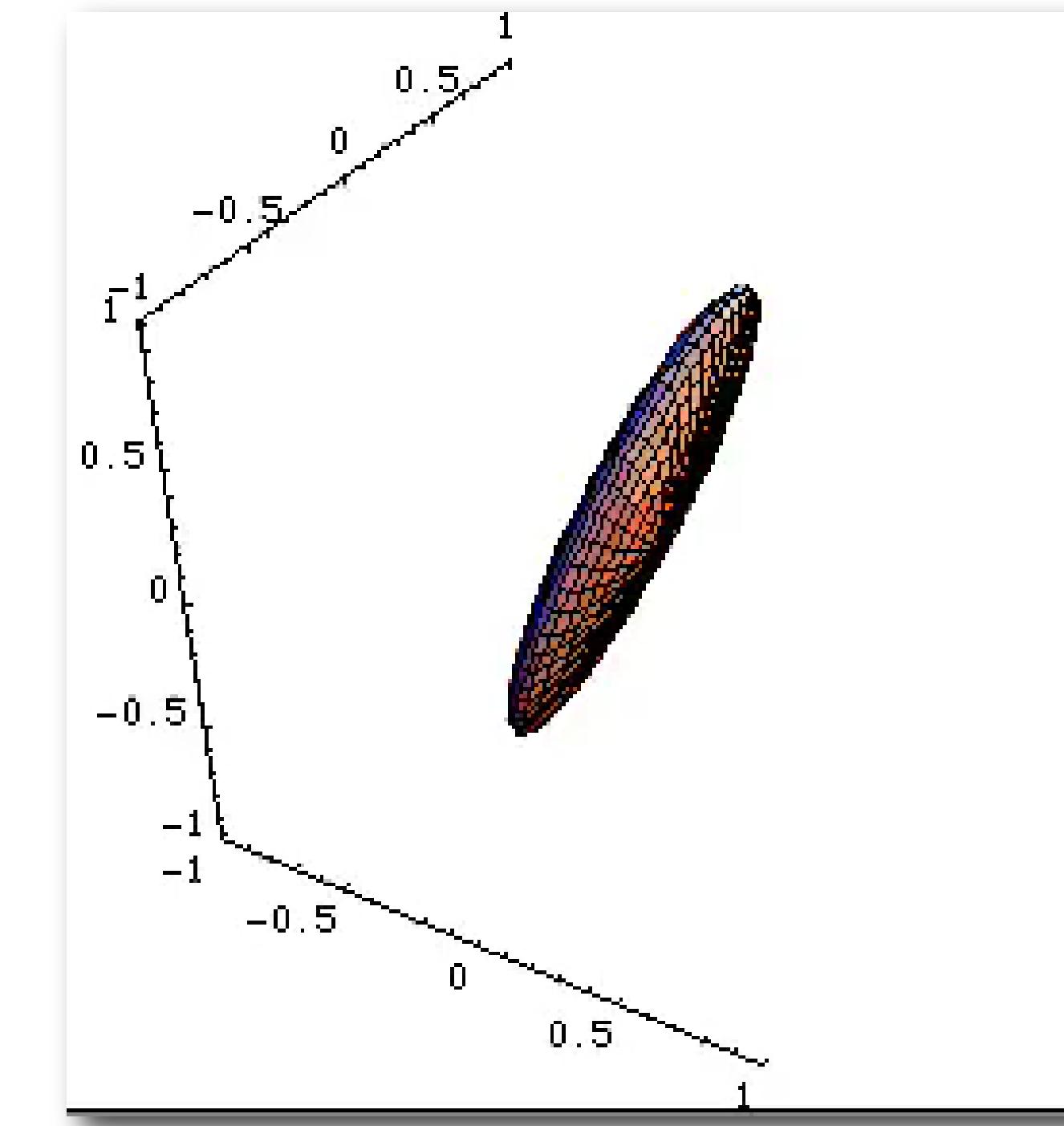
General Rotation of Rigid Body with

$$I_1 \geq I_2 \geq I_3$$



Pure Spin

$$\omega_0 = (0^\circ, 10^\circ, 0^\circ) / \text{s}$$



Slightly Off Spin

$$\omega_0 = (0.5^\circ, 10^\circ, 0.5^\circ) / \text{s}$$

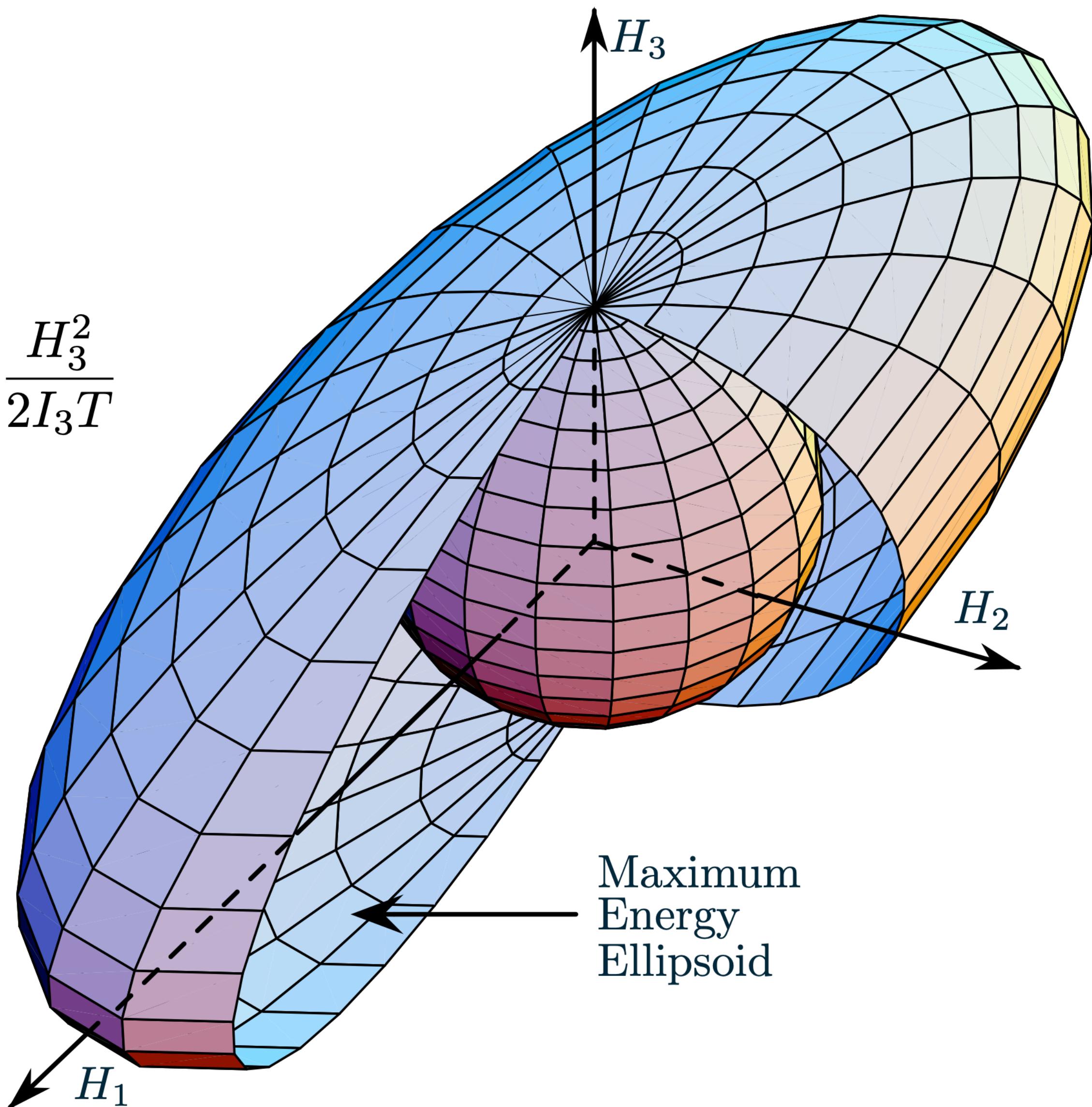
- Maximum Energy Case:

$${}^B\boldsymbol{H} = \pm H \hat{\boldsymbol{b}}_3$$

The kinetic energy is:

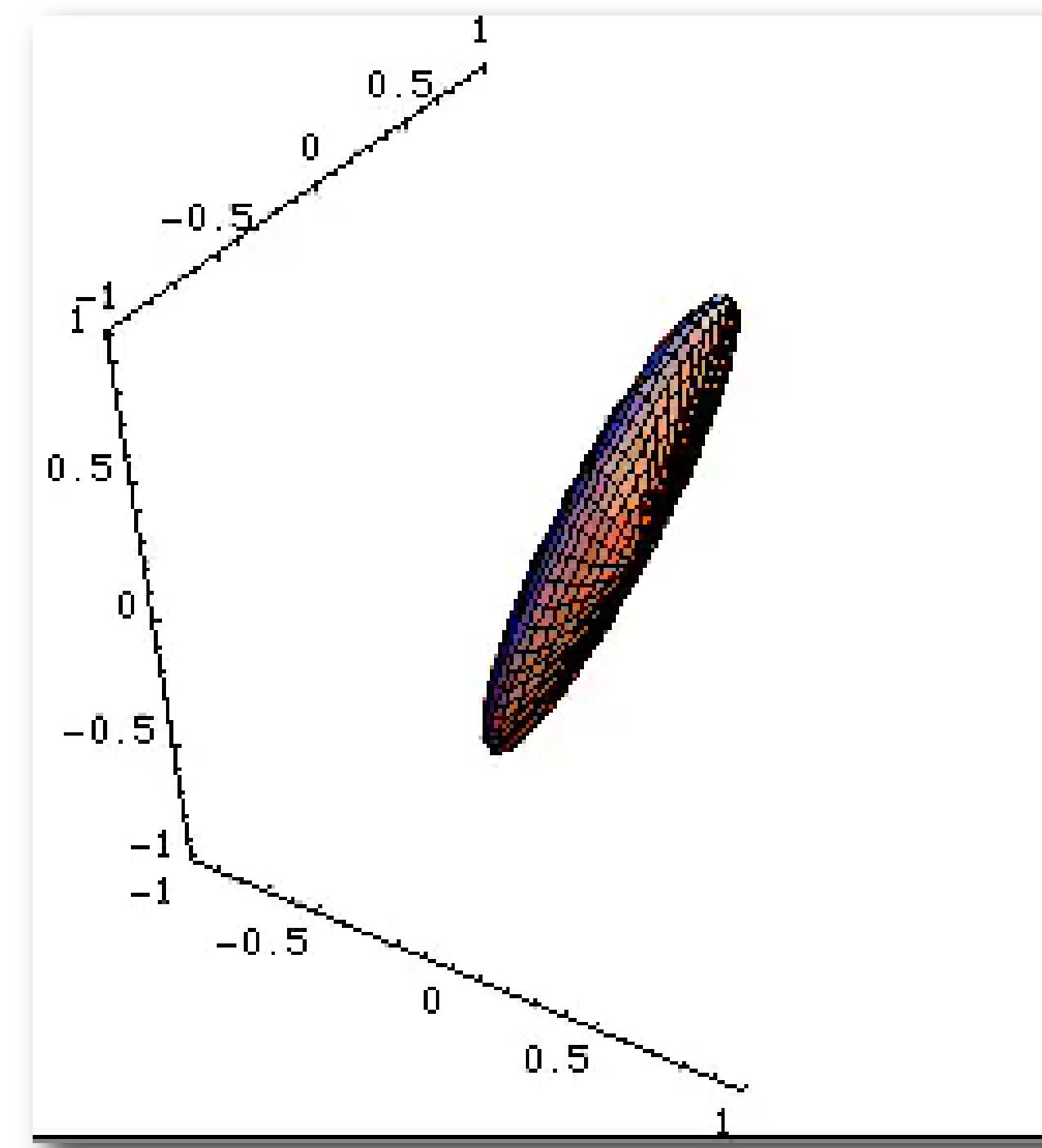
$$1 = \frac{H_1^2}{2I_1T} + \frac{H_2^2}{2I_2T} + \frac{H_3^2}{2I_3T}$$

$$T_{\max} = \frac{H^2}{2I_3}$$



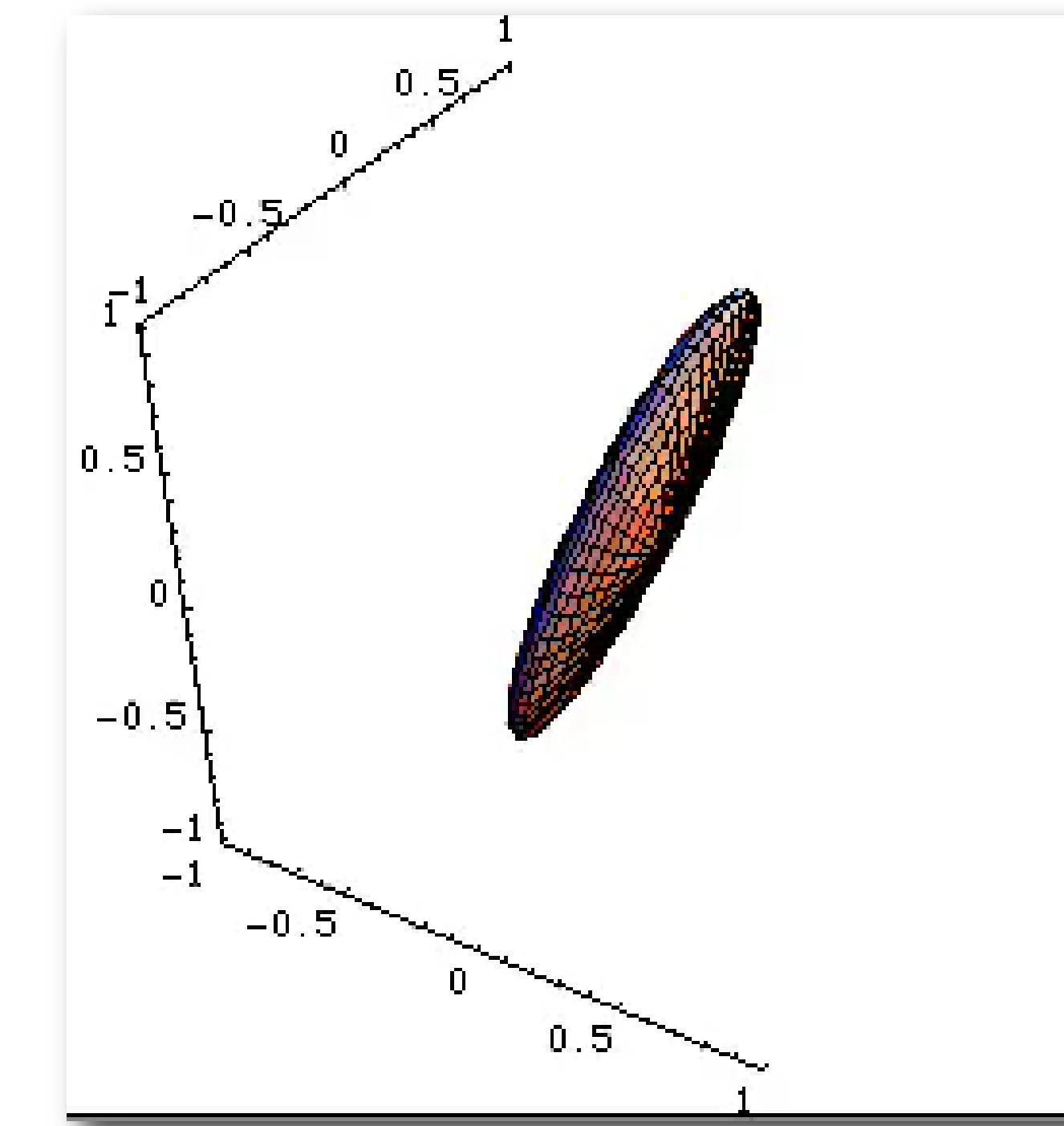
General Rotation of Rigid Body with

$$I_1 \geq I_2 \geq I_3$$



Pure Spin

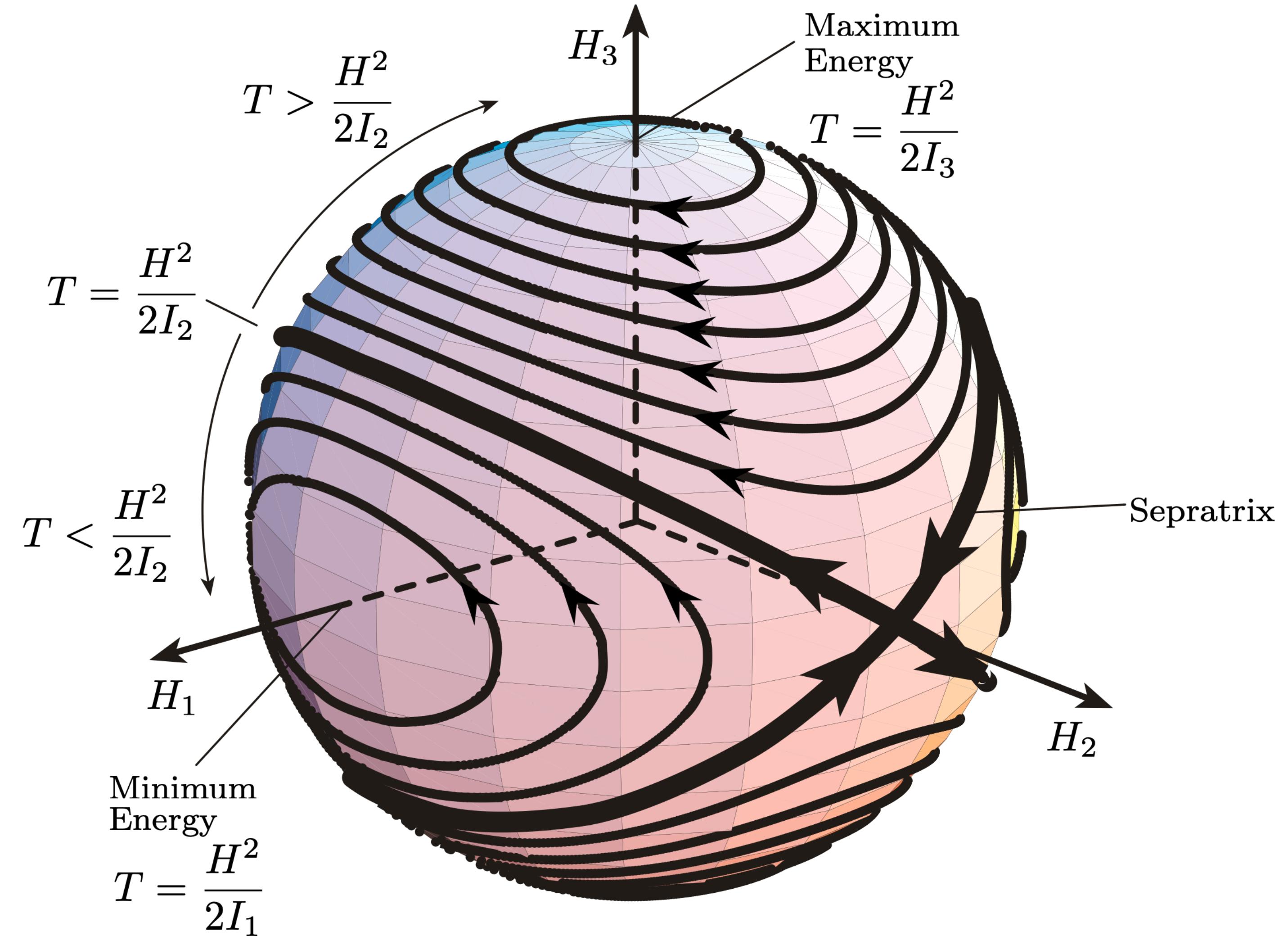
$$\omega_0 = (0^\circ, 0^\circ, 10^\circ) / \text{s}$$



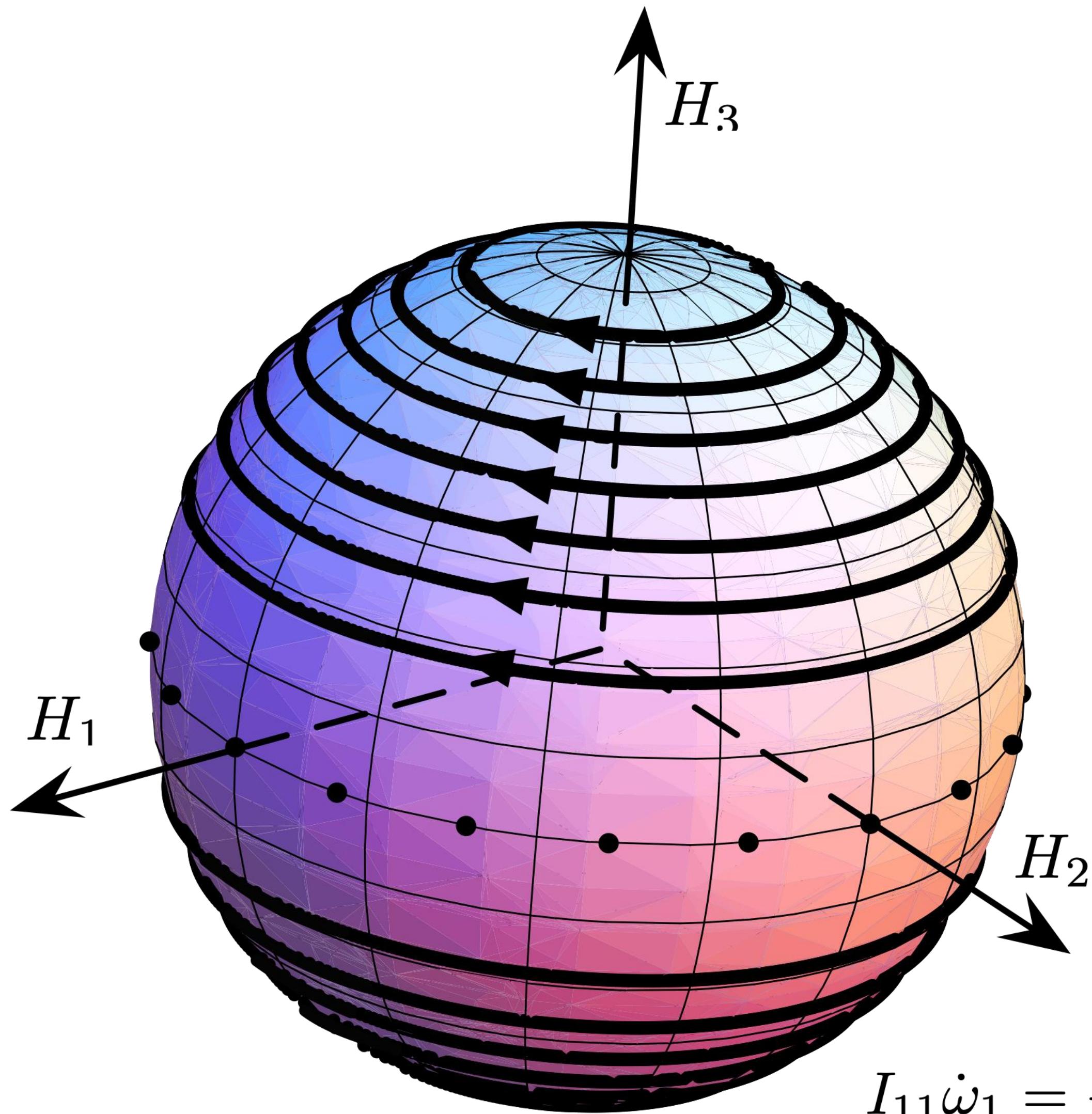
Slightly Off Spin

$$\omega_0 = (0.5^\circ, 0.5^\circ, 10^\circ) / \text{s}$$

$$I_1 \geq I_2 \geq I_3$$



Family of energy ellipsoid and momentum sphere intersections.

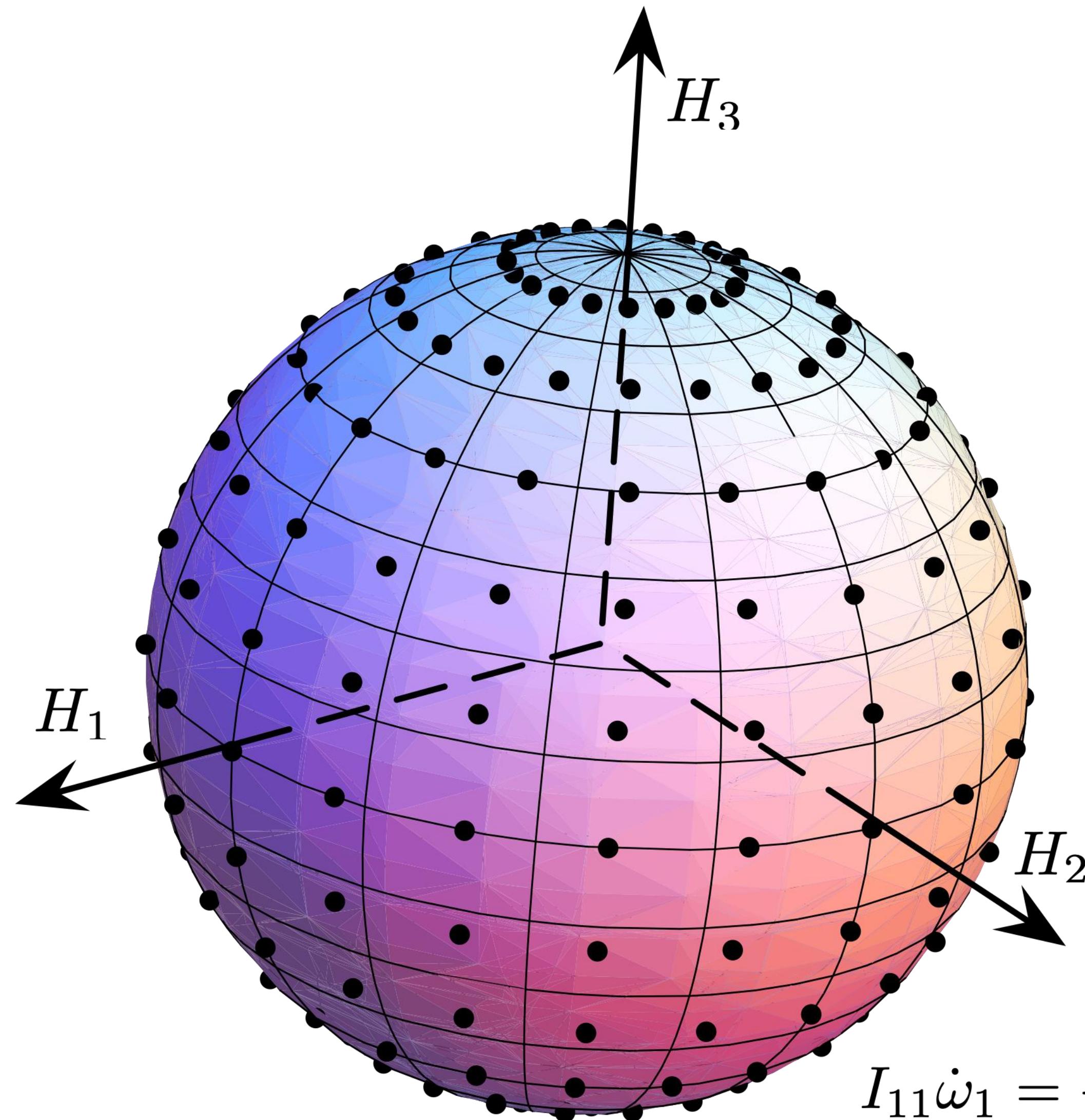


What type of spacecraft body would yield this?

$$I_{11}\dot{\omega}_1 = -(I_{33} - I_{22})\omega_2\omega_3$$

$$I_{22}\dot{\omega}_2 = -(I_{11} - I_{33})\omega_3\omega_1$$

$$I_{33}\dot{\omega}_3 = -(I_{22} - I_{11})\omega_1\omega_2$$



What type of spacecraft body would yield this?

$$I_{11}\dot{\omega}_1 = -(I_{33} - I_{22})\omega_2\omega_3$$

$$I_{22}\dot{\omega}_2 = -(I_{11} - I_{33})\omega_3\omega_1$$

$$I_{33}\dot{\omega}_3 = -(I_{22} - I_{11})\omega_1\omega_2$$

# Analytical Torque-Free Motion

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- Let us assume that there are no external torques acting on the rigid body, and the equations of motion are given by:

$$I_{11}\dot{\omega}_1 = -(I_{33} - I_{22})\omega_2\omega_3$$

$$I_{22}\dot{\omega}_2 = -(I_{11} - I_{33})\omega_3\omega_1$$

$$I_{33}\dot{\omega}_3 = -(I_{22} - I_{11})\omega_1\omega_2$$

- We are looking for analytical solutions to the angular motion.
- Assume that the body coordinate frame is a principal frame, and the inertia matrix is diagonal.



# Axi-Symmetric Case

- Let the external torque be zero. Consider the special principal inertia case where

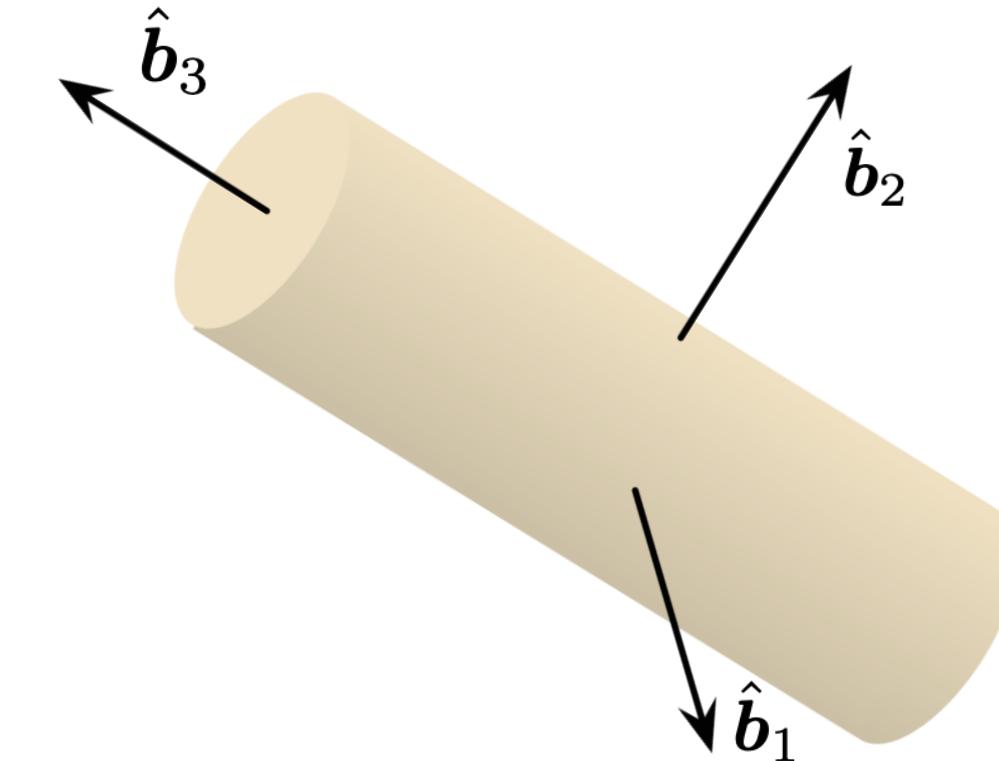
$$I_T = I_{11} = I_{22}$$

- Here the EOM are given by

$$I_T \dot{\omega}_1 = -(I_{33} - I_T) \omega_2 \omega_3$$

$$I_T \dot{\omega}_2 = (I_{33} - I_T) \omega_3 \omega_1$$

$$I_{33} \dot{\omega}_3 = 0$$

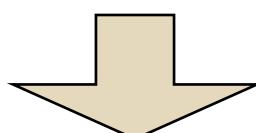


- From this equation it is clear that the third angular velocity component will be constant.

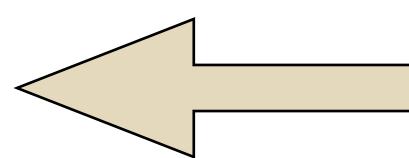
$$\omega_3(t) = \omega_3(t_0) = \text{constant}$$

- Let's examine the remaining two differential equations more carefully. Taking the derivative of the first one we find

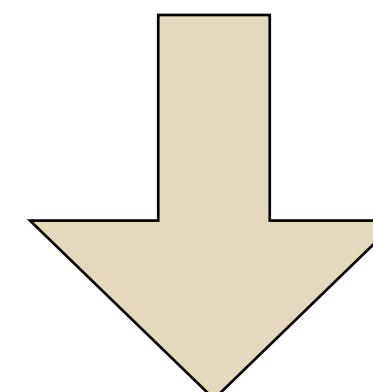
$$I_t \dot{\omega}_1 = -(I_{33} - I_T) \omega_2 \omega_3$$



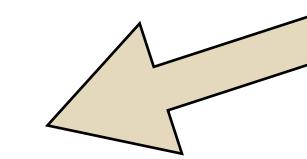
$$I_T \ddot{\omega}_1 = -(I_{33} - I_T) \dot{\omega}_2 \omega_3$$



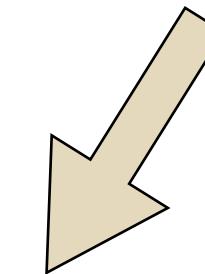
$$\dot{\omega}_2 = \frac{1}{I_T} ((I_{33} - I_T) \omega_3 \omega_1)$$



$$\ddot{\omega}_1 + \left(\frac{I_{33}}{I_T} - 1\right)^2 \omega_3^2 \omega_1 = 0$$



Mathematically equivalent to  
simple Spring-Mass Systems!



Similarly, we can find:

$$\ddot{\omega}_2 + \left(\frac{I_{33}}{I_T} - 1\right)^2 \omega_3^2 \omega_2 = 0$$

- The analytical solution to a spring-mass dynamical system is the simple oscillator equation

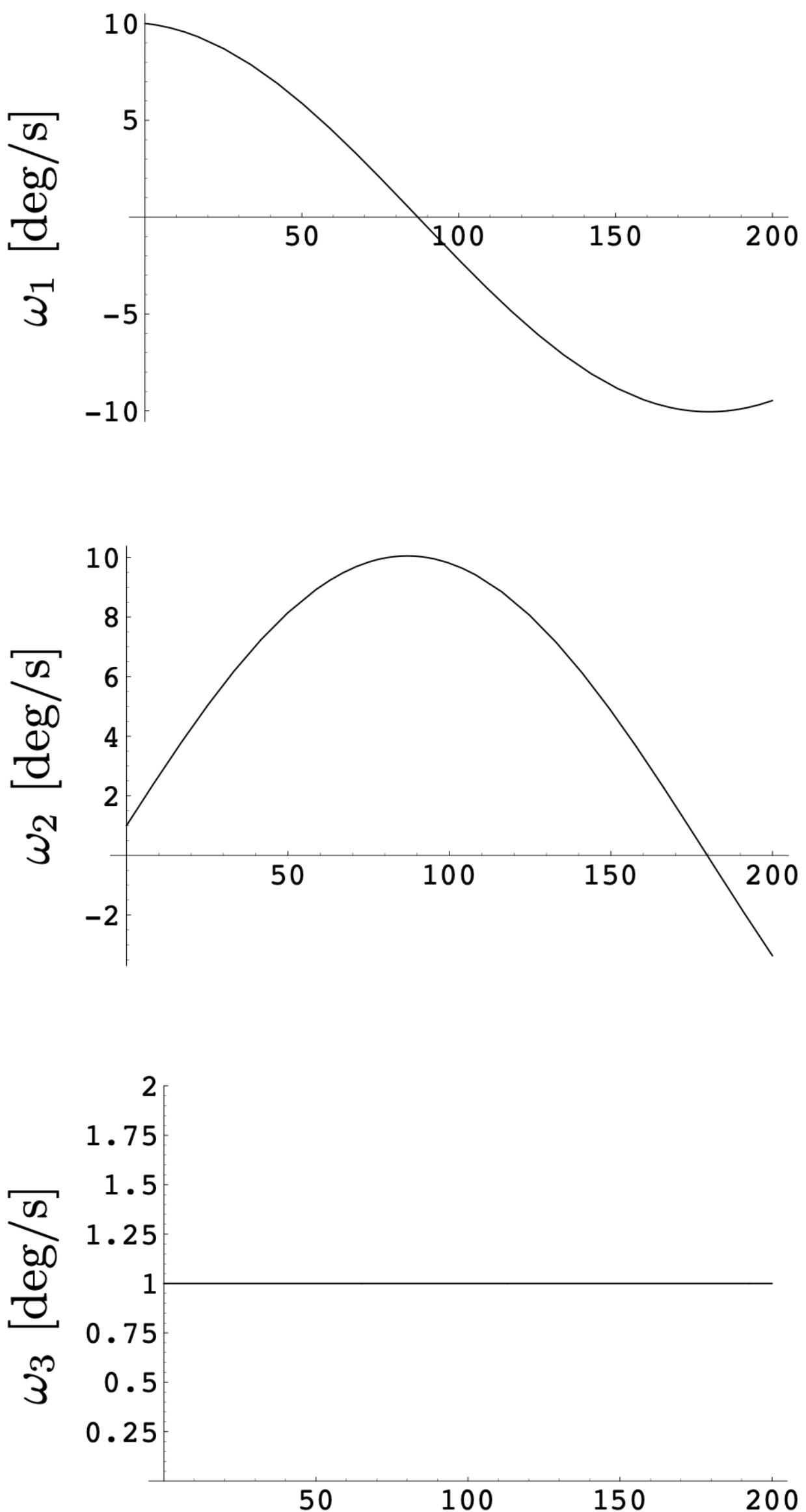
$$\begin{aligned}\omega_1(t) &= A_1 \cos \omega_p t + B_1 \sin \omega_p t \\ \omega_2(t) &= A_2 \cos \omega_p t + B_2 \sin \omega_p t\end{aligned}$$

- Using the initial conditions, we find the analytical solution of the body angular velocity components for the axi-symmetric spacecraft case:

$$\omega_p = \left( \frac{I_{33}}{I_T} - 1 \right) \omega_3$$

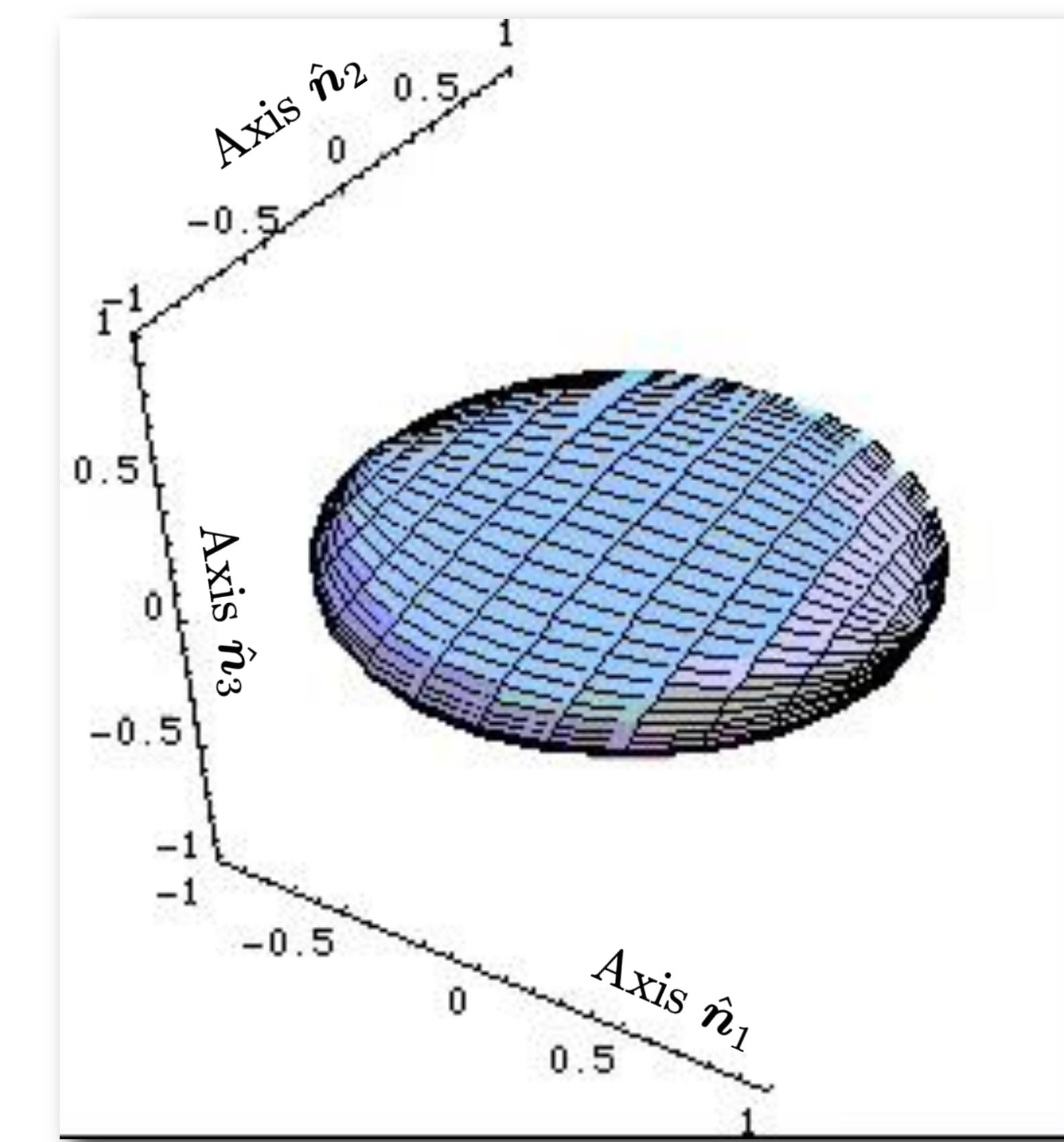
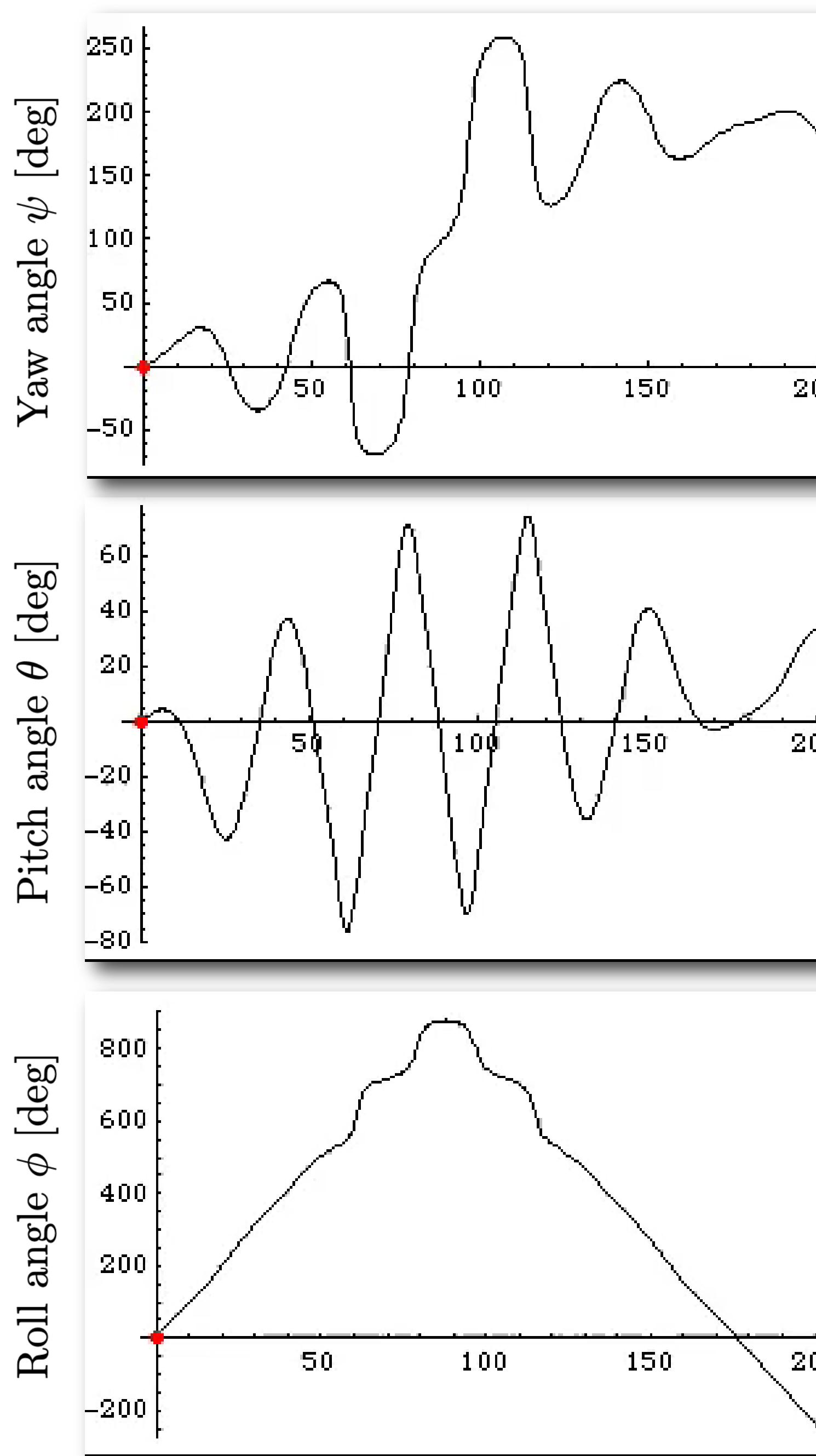
where

$$\begin{aligned}\omega_1(t) &= \omega_{10} \cos \omega_p t - \omega_{20} \sin \omega_p t \\ \omega_2(t) &= \omega_{20} \cos \omega_p t + \omega_{10} \sin \omega_p t \\ \omega_3(t) &= \omega_{30}\end{aligned}$$



The first and second body angular velocity components are sinusoidal in nature.

As predicted, the third body angular velocity component remains constant here.

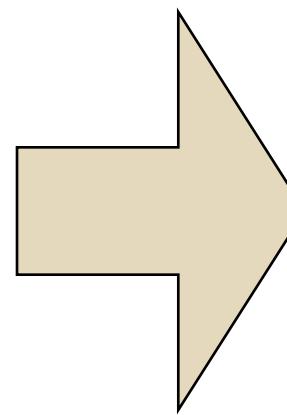


# General Inertia Case\*

$$H^2 = I_1^2\omega_1^2 + I_2^2\omega_2^2 + I_3^2\omega_3^2$$

$$2T = I_1\omega_1^2 + I_2\omega_2^2 + I_3\omega_3^2$$

Momentum magnitude and kinetic energy conservation yield two integrals of the torque-free motion.



$$\omega_2^2 = \left( \frac{2I_3T - H^2}{I_2(I_3 - I_2)} \right) - \frac{I_1(I_3 - I_1)}{I_2(I_3 - I_2)}\omega_1^2$$

$$\omega_3^2 = \left( \frac{2I_2T - H^2}{I_3(I_2 - I_3)} \right) - \frac{I_1(I_2 - I_1)}{I_3(I_2 - I_3)}\omega_1^2$$

We can use these two equations to solve for two of the angular rates!

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Analogously, we can solve for the two angular velocities in terms of other angular rates.

$$\omega_1^2 = \left( \frac{2I_3T - H^2}{I_1(I_3 - I_1)} \right) - \frac{I_2(I_3 - I_2)}{I_1(I_3 - I_1)}\omega_2^2$$

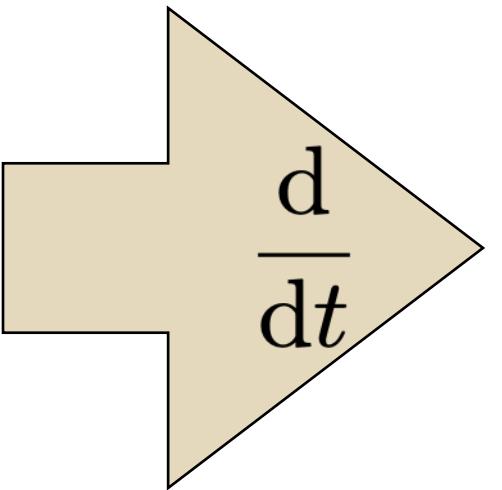
$$\omega_3^2 = \left( \frac{2I_1T - H^2}{I_3(I_1 - I_3)} \right) - \frac{I_2(I_1 - I_2)}{I_3(I_1 - I_3)}\omega_2^2$$

$$\omega_1^2 = \left( \frac{2I_2T - H^2}{I_1(I_2 - I_1)} \right) - \frac{I_3(I_2 - I_3)}{I_1(I_2 - I_1)}\omega_3^2$$

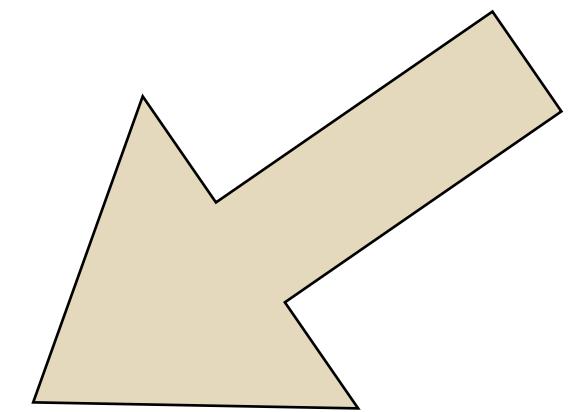
$$\omega_2^2 = \left( \frac{2I_1T - H^2}{I_2(I_1 - I_2)} \right) - \frac{I_3(I_1 - I_3)}{I_2(I_1 - I_2)}\omega_3^2$$

\* Junkins, J. L., Jacobson, I. D., and Blanton, J. N., "A Nonlinear Oscillator Analog of Rigid Body Dynamics," *Celestial Mechanics*, Vol. 7, pp. 398 – 407, 1973.

$$\begin{aligned} I_1 \dot{\omega}_1 &= -(I_3 - I_2) \omega_2 \omega_3 \\ I_2 \dot{\omega}_2 &= -(I_1 - I_3) \omega_3 \omega_1 \\ I_3 \dot{\omega}_3 &= -(I_2 - I_1) \omega_1 \omega_2 \end{aligned}$$



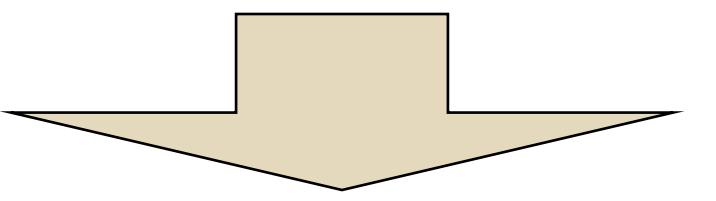
$$\begin{aligned} \ddot{\omega}_1 &= \frac{I_2 - I_3}{I_1} (\dot{\omega}_2 \omega_3 + \omega_2 \dot{\omega}_3) \\ \ddot{\omega}_2 &= \frac{I_3 - I_1}{I_2} (\dot{\omega}_3 \omega_1 + \omega_3 \dot{\omega}_1) \\ \ddot{\omega}_3 &= \frac{I_1 - I_2}{I_3} (\dot{\omega}_1 \omega_2 + \omega_1 \dot{\omega}_2) \end{aligned}$$



$$\begin{aligned} \ddot{\omega}_1 &= \frac{I_2 - I_3}{I_1} \left( \frac{I_1 - I_2}{I_3} \omega_1 \omega_2^2 + \frac{I_3 - I_1}{I_2} \omega_1 \omega_3^2 \right) \\ \ddot{\omega}_2 &= \frac{I_3 - I_1}{I_2} \left( \frac{I_1 - I_2}{I_3} \omega_2 \omega_1^2 + \frac{I_2 - I_3}{I_1} \omega_2 \omega_3^2 \right) \\ \ddot{\omega}_3 &= \frac{I_1 - I_2}{I_3} \left( \frac{I_3 - I_1}{I_2} \omega_3 \omega_1^2 + \frac{I_2 - I_3}{I_1} \omega_3 \omega_2^2 \right) \end{aligned}$$

$$\ddot{\omega}_1 = \frac{I_2 - I_3}{I_1} \left( \frac{I_1 - I_2}{I_3} \omega_1 \boxed{\omega_2^2} + \frac{I_3 - I_1}{I_2} \omega_1 \boxed{\omega_3^2} \right)$$

$$\ddot{\omega}_2 = \frac{I_3 - I_1}{I_2} \left( \frac{I_1 - I_2}{I_3} \omega_2 \boxed{\omega_1^2} + \frac{I_2 - I_3}{I_1} \omega_2 \boxed{\omega_3^2} \right)$$

$$\ddot{\omega}_3 = \frac{I_1 - I_2}{I_3} \left( \frac{I_3 - I_1}{I_2} \omega_3 \boxed{\omega_1^2} + \frac{I_2 - I_3}{I_1} \omega_3 \boxed{\omega_2^2} \right)$$


$$\omega_2^2 = \left( \frac{2I_3T - H^2}{I_2(I_3 - I_2)} \right) - \frac{I_1(I_3 - I_1)}{I_2(I_3 - I_2)} \omega_1^2$$

$$\omega_3^2 = \left( \frac{2I_2T - H^2}{I_3(I_2 - I_3)} \right) - \frac{I_1(I_2 - I_1)}{I_3(I_2 - I_3)} \omega_1^2$$

$$\omega_1^2 = \left( \frac{2I_3T - H^2}{I_1(I_3 - I_1)} \right) - \frac{I_2(I_3 - I_2)}{I_1(I_3 - I_1)} \omega_2^2$$

$$\omega_3^2 = \left( \frac{2I_1T - H^2}{I_3(I_1 - I_3)} \right) - \frac{I_2(I_1 - I_2)}{I_3(I_1 - I_3)} \omega_2^2$$

$$\omega_1^2 = \left( \frac{2I_2T - H^2}{I_1(I_2 - I_1)} \right) - \frac{I_3(I_2 - I_3)}{I_1(I_2 - I_1)} \omega_3^2$$

$$\omega_2^2 = \left( \frac{2I_1T - H^2}{I_2(I_1 - I_2)} \right) - \frac{I_3(I_1 - I_3)}{I_2(I_1 - I_2)} \omega_3^2$$

$$\ddot{\omega}_i + A_i \omega_i + B_i \omega_i^3 = 0 \quad \text{for } i = 1, 2, 3$$

homogenous, undamped Duffing equation

Duffing equations are often found studying nonlinear mechanical oscillations, where the cubic “stiffness” term arises to approximately account for nonlinear departure from Hooke’s law. For the torque-free motion, this equation is the *exact differential equation!*



$$\ddot{\omega}_i + A_i \omega_i + B_i \omega_i^3 = 0 \quad \text{for } i = 1, 2, 3$$

- These equations form three *uncoupled nonlinear oscillators*.
- Notice that while the oscillators are *uncoupled*, they are not *independent*! The six spring constants are all uniquely determined from initially evaluated inertia, energy and momentum constants.

$i$	$A_i$	$B_i$
1	$\frac{(I_1 - I_2)(2I_3T - H^2) + (I_1 - I_3)(2I_2T - H^2)}{I_1 I_2 I_3}$	$\frac{2(I_1 - I_2)(I_1 - I_3)}{I_2 I_3}$
2	$\frac{(I_2 - I_3)(2I_1T - H^2) + (I_2 - I_1)(2I_3T - H^2)}{I_1 I_2 I_3}$	$\frac{2(I_2 - I_1)(I_2 - I_3)}{I_1 I_3}$
3	$\frac{(I_3 - I_1)(2I_2T - H^2) + (I_3 - I_2)(2I_1T - H^2)}{I_1 I_2 I_3}$	$\frac{2(I_3 - I_1)(I_3 - I_2)}{I_1 I_2}$

- The oscillator differential equations have three immediate integrals of the form

$$\dot{\omega}_i^2 + A_i \omega_i^2 + \frac{B_i}{2} \omega_i^4 = K_i \quad \text{for } i = 1, 2, 3$$

- Here  $K_1$ ,  $K_2$  and  $K_3$  are the three oscillator “energy-type” integral constants of the motion.

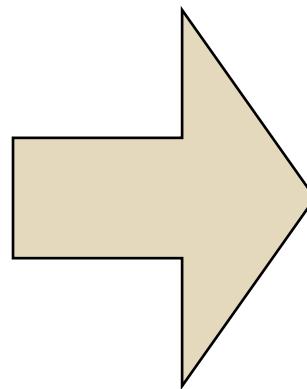
$$K_1 = \frac{(2I_2T - H^2)(H^2 - 2I_3T)}{I_1^2 I_2 I_3}$$

$$K_2 = \frac{(2I_3T - H^2)(H^2 - 2I_1T)}{I_1 I_2^2 I_3}$$

$$K_3 = \frac{(2I_1T - H^2)(H^2 - 2I_2T)}{I_1 I_2 I_3^2}$$

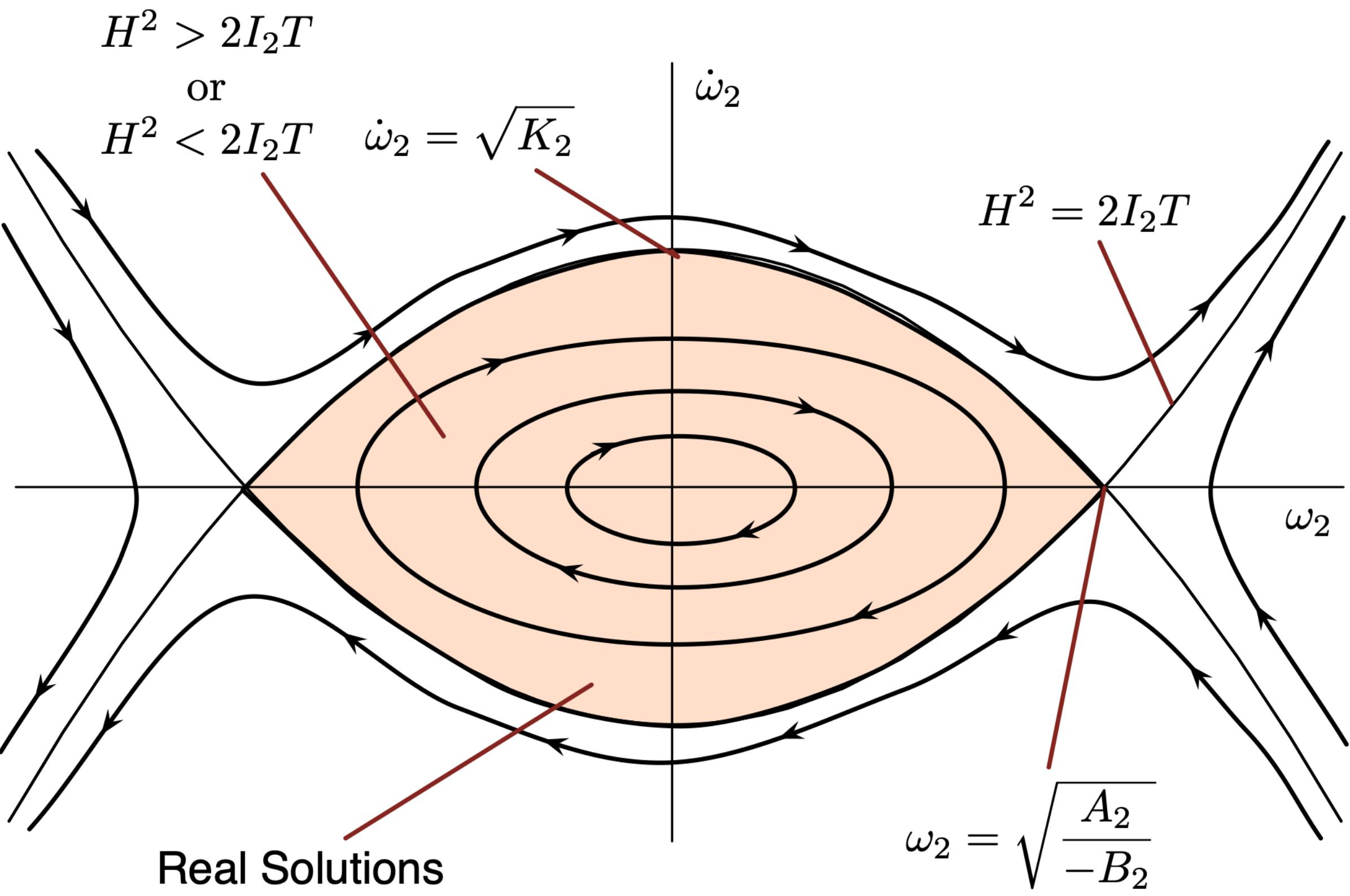
Assume:

$$I_1 \geq I_2 \geq I_3$$



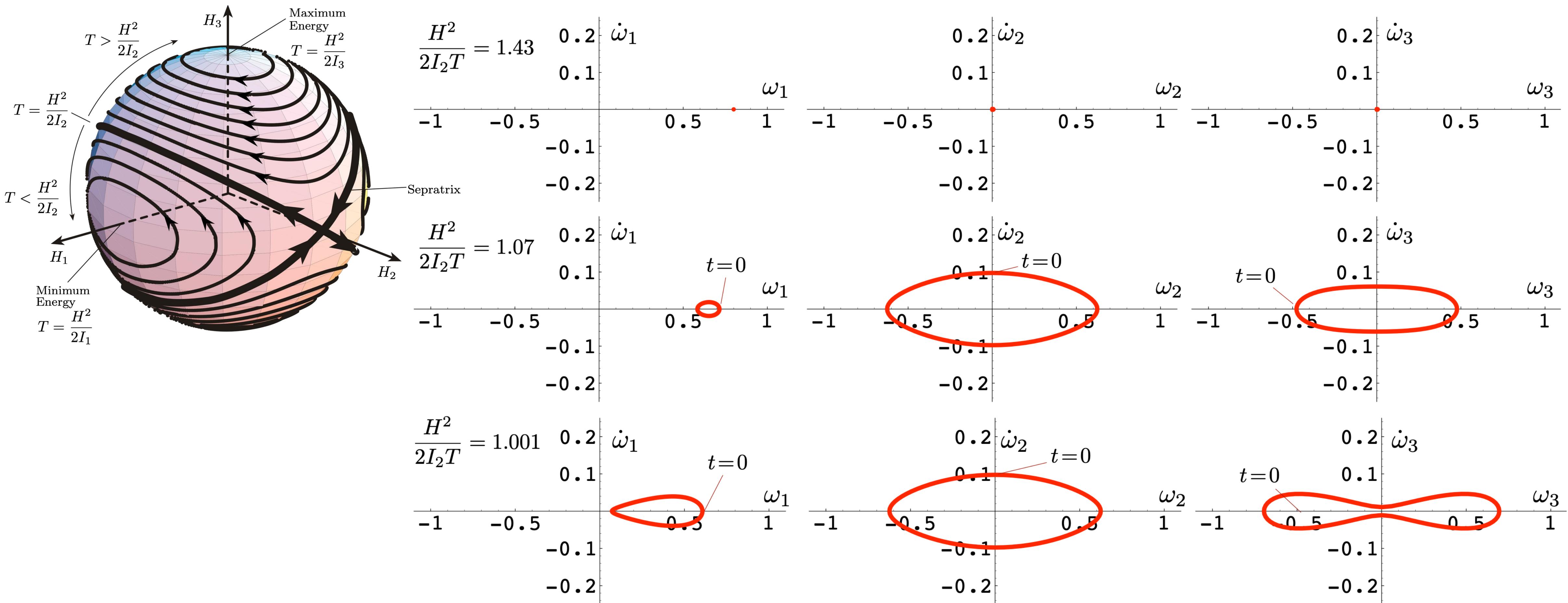
$i$	$A_i$	$B_i$
1	not defined	$>0$
2	$>0$	$<0$
3	not defined	$>0$

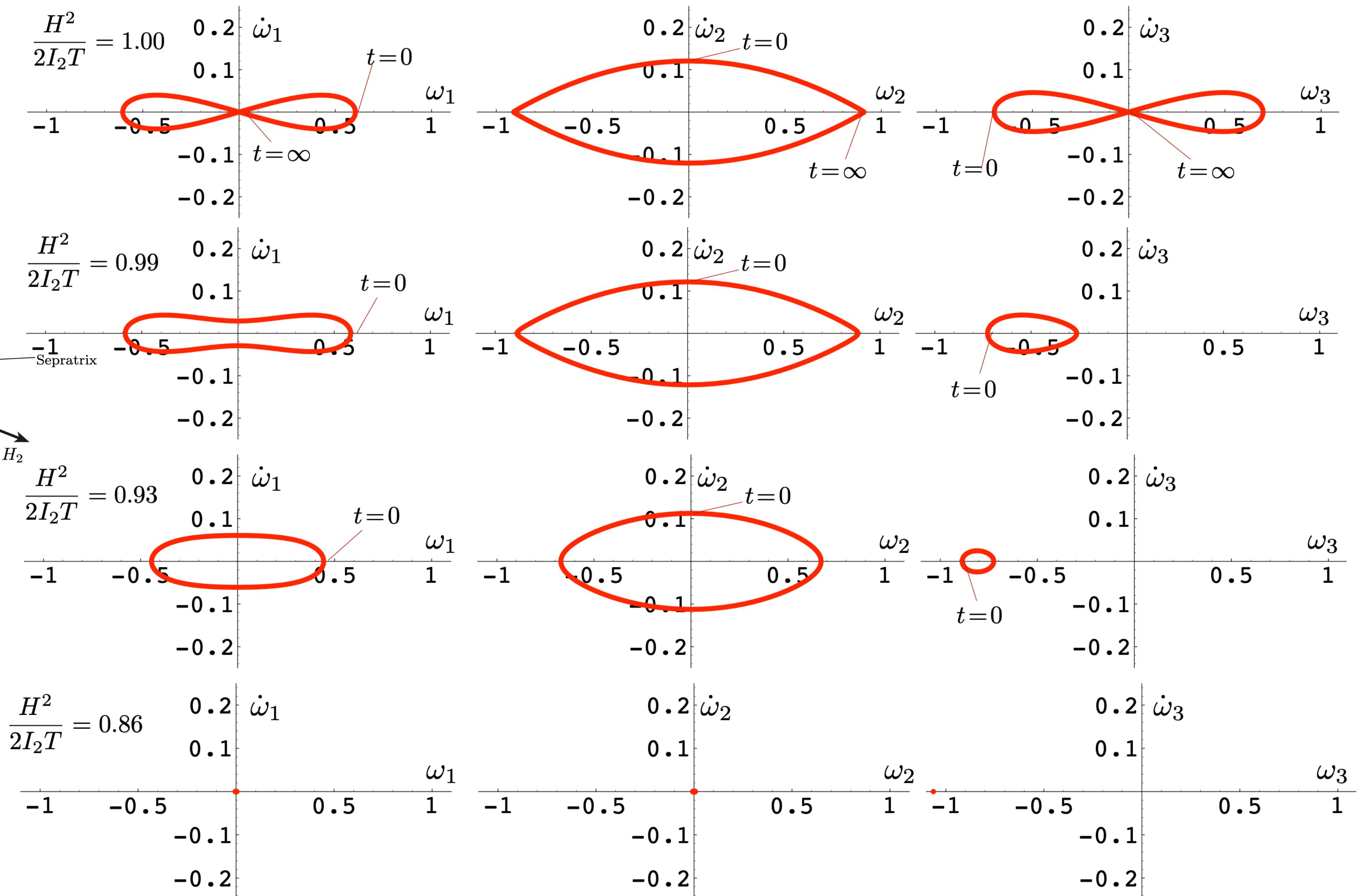
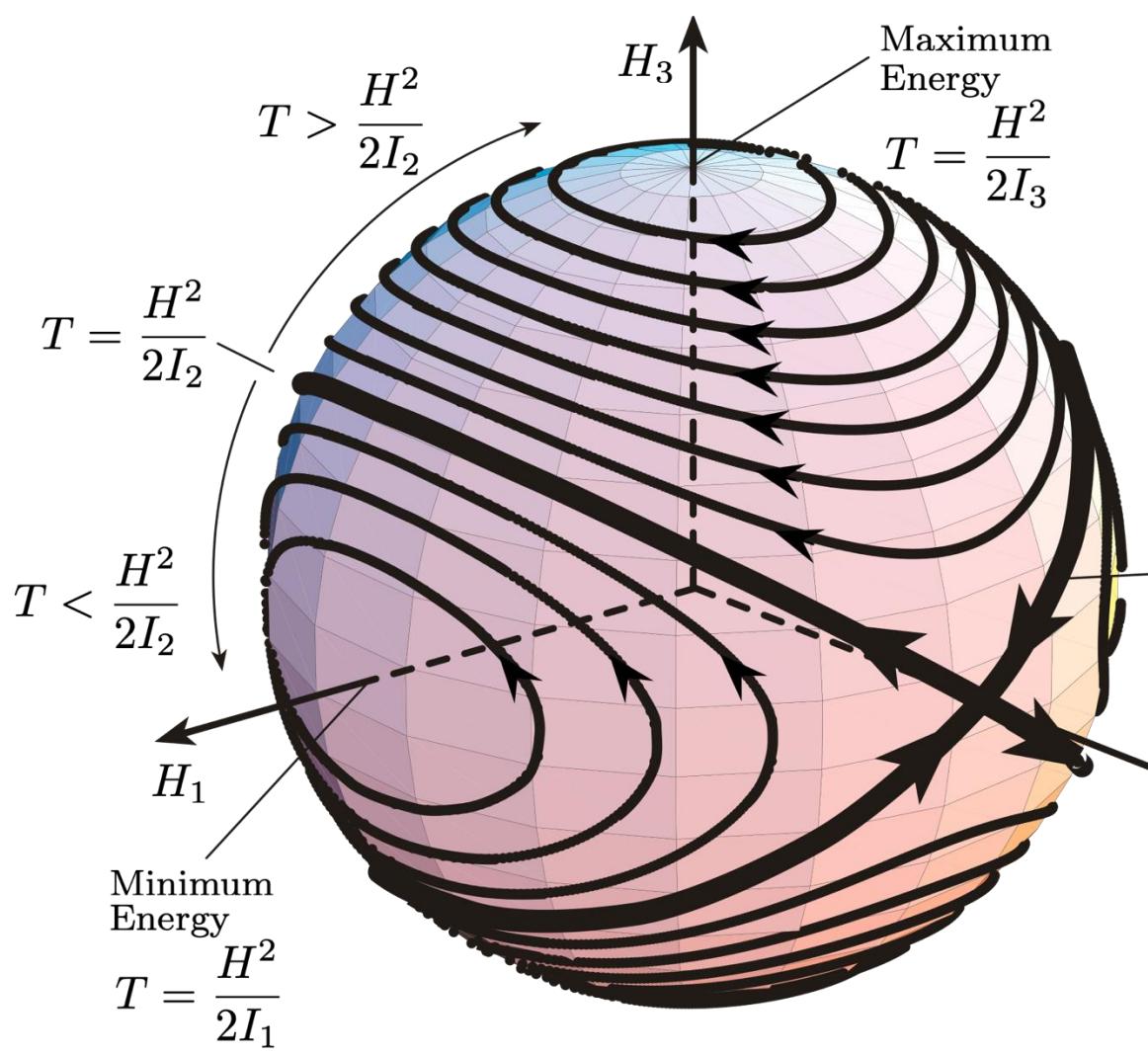
- The linear “spring constants”  $A_1$  and  $A_3$  can produce de-stabilizing spring forces (negative spring effect).
- The positive cubic “spring constants”  $B_1$  and  $B_3$  always produce restoring forces and are therefore hard springs. Because cubic springs will override linear springs for sufficiently large displacements, all trajectories of the 1<sup>st</sup> and 3<sup>rd</sup> phase planes must be closed.
- The cubic spring constant  $B_2$  produces a de-stabilizing force (soft spring), and will eventually override the stabilizing linear spring force.



- Only solutions with  $K_2 \geq 0$  are physically possible
- The limiting trajectory occurs if
  - $I_1 \rightarrow I_3$
  - $H^2 \rightarrow 2 I_2 T$  (pure spin about intermediate inertia axis)
  - $I_1 I_2 I_3 \rightarrow \infty$

Let's sweep through cases from a minimum energy case to a maximum energy case. The momentum is held constant here.





# General Free Rotation

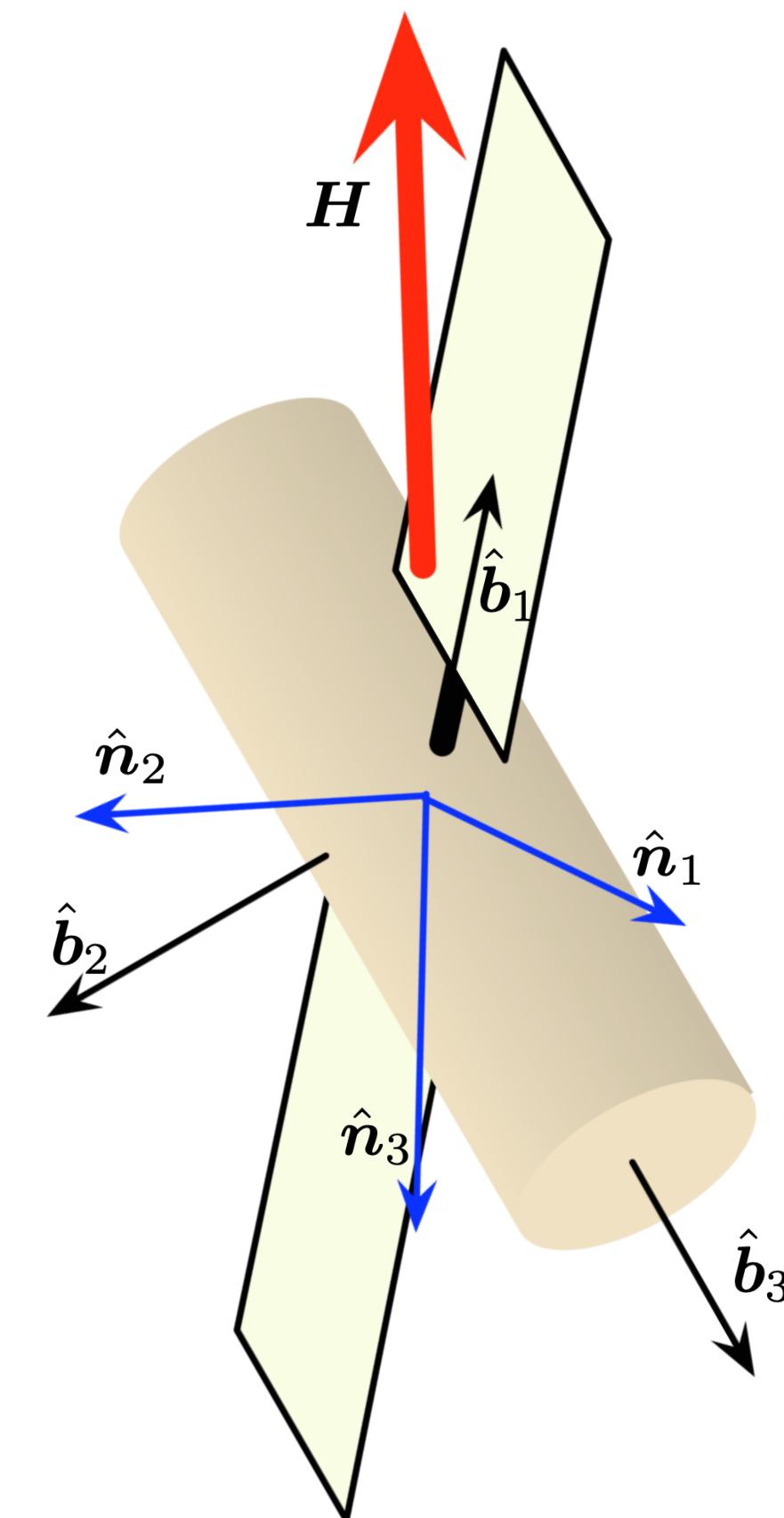
- We would like to study the general free rotation of a rigid body using the 3-2-1 Euler angles.
- Because the inertial angular momentum vector  $H$  is constant as seen by the inertial frame, we can always align our inertial frame such that

$$H = {}^N\mathbf{H} = -H\hat{\mathbf{n}}_3 = \begin{pmatrix} 0 \\ 0 \\ -H \end{pmatrix}$$

•

- Using the rotation matrix  $[BN]$ , we find

$${}^B\mathbf{H} = [BN] {}^N\mathbf{H}$$



- Recall the mapping between the rotation matrix  $[BN]$  and the 3-2-1 Euler angles:

$$[BN] = \begin{bmatrix} c\theta_2c\theta_1 & c\theta_2s\theta_1 & -s\theta_2 \\ s\theta_3s\theta_2c\theta_1 - c\theta_3s\theta_1 & s\theta_3s\theta_2s\theta_1 + c\theta_3c\theta_1 & s\theta_3c\theta_2 \\ c\theta_3s\theta_2c\theta_1 + s\theta_3s\theta_1 & c\theta_3s\theta_2s\theta_1 - s\theta_3c\theta_1 & c\theta_3c\theta_2 \end{bmatrix}$$

This leads to

$$\mathcal{B}H = \begin{pmatrix} H_1 \\ H_2 \\ H_3 \end{pmatrix} = [BN]^N H = \begin{pmatrix} H \sin \theta \\ -H \sin \phi \cos \theta \\ -H \cos \phi \cos \theta \end{pmatrix} = \begin{pmatrix} I_1 \omega_1 \\ I_2 \omega_2 \\ I_3 \omega_3 \end{pmatrix}$$

Which can be solved for the rigid body angular velocity.

$$\begin{pmatrix} \frac{H}{I_1} \sin \theta \\ -\frac{H}{I_2} \sin \phi \cos \theta \\ -\frac{H}{I_3} \cos \phi \cos \theta \end{pmatrix} = \begin{pmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix}$$

- Recall the 3-2-1 Euler angle differential kinematic equation:

$$\boldsymbol{\omega} = \begin{pmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix} = \begin{bmatrix} -\sin \theta & 0 & 1 \\ \sin \phi \cos \theta & \cos \phi & 0 \\ \cos \phi \cos \theta & -\sin \phi & 0 \end{bmatrix} \begin{pmatrix} \dot{\psi} \\ \dot{\theta} \\ \dot{\phi} \end{pmatrix}$$

Solving these equations for the Euler angle rates, we obtain:

$$\begin{aligned} \dot{\psi} &= -H \left( \frac{\sin^2 \phi}{I_2} + \frac{\cos^2 \phi}{I_3} \right) \quad \rightarrow \text{cannot be positive} \\ \dot{\theta} &= \frac{H}{2} \left( \frac{1}{I_3} - \frac{1}{I_2} \right) \sin 2\phi \cos \theta \\ \dot{\phi} &= H \left( \frac{1}{I_1} - \frac{\sin^2 \phi}{I_2} - \frac{\cos^2 \phi}{I_3} \right) \sin \theta \end{aligned}$$

These are the spinning top equations of motion.

# Axi-Symmetric Coning Motion

- Assume the spacecraft is axi-symmetric with  $I_2 = I_3$ , and align the inertial frame such that

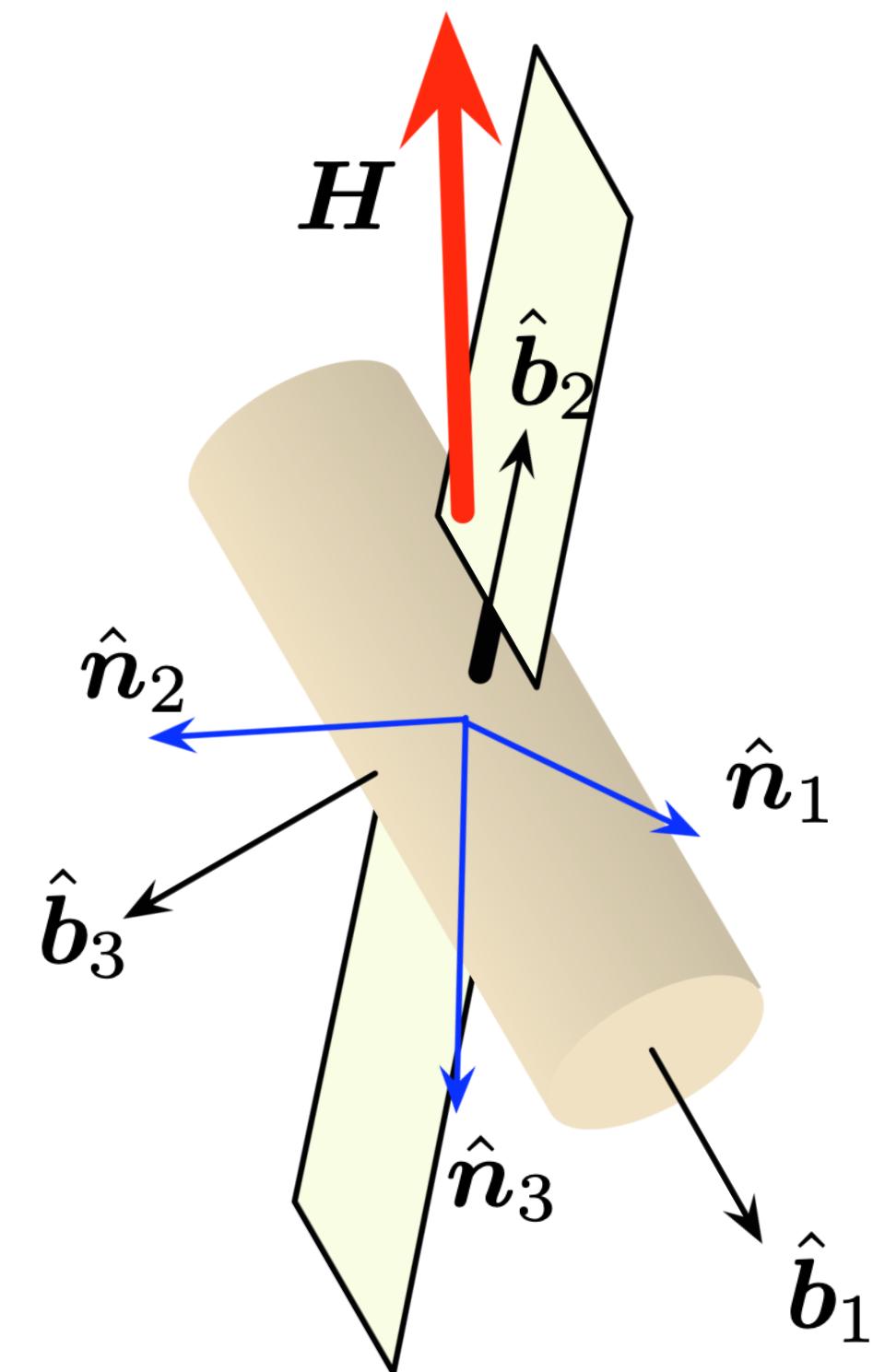
$$\mathbf{H} = {}^N\mathbf{H} = -H\hat{\mathbf{n}}_3 = \begin{pmatrix} {}^N \\ 0 \\ 0 \\ -H \end{pmatrix}$$

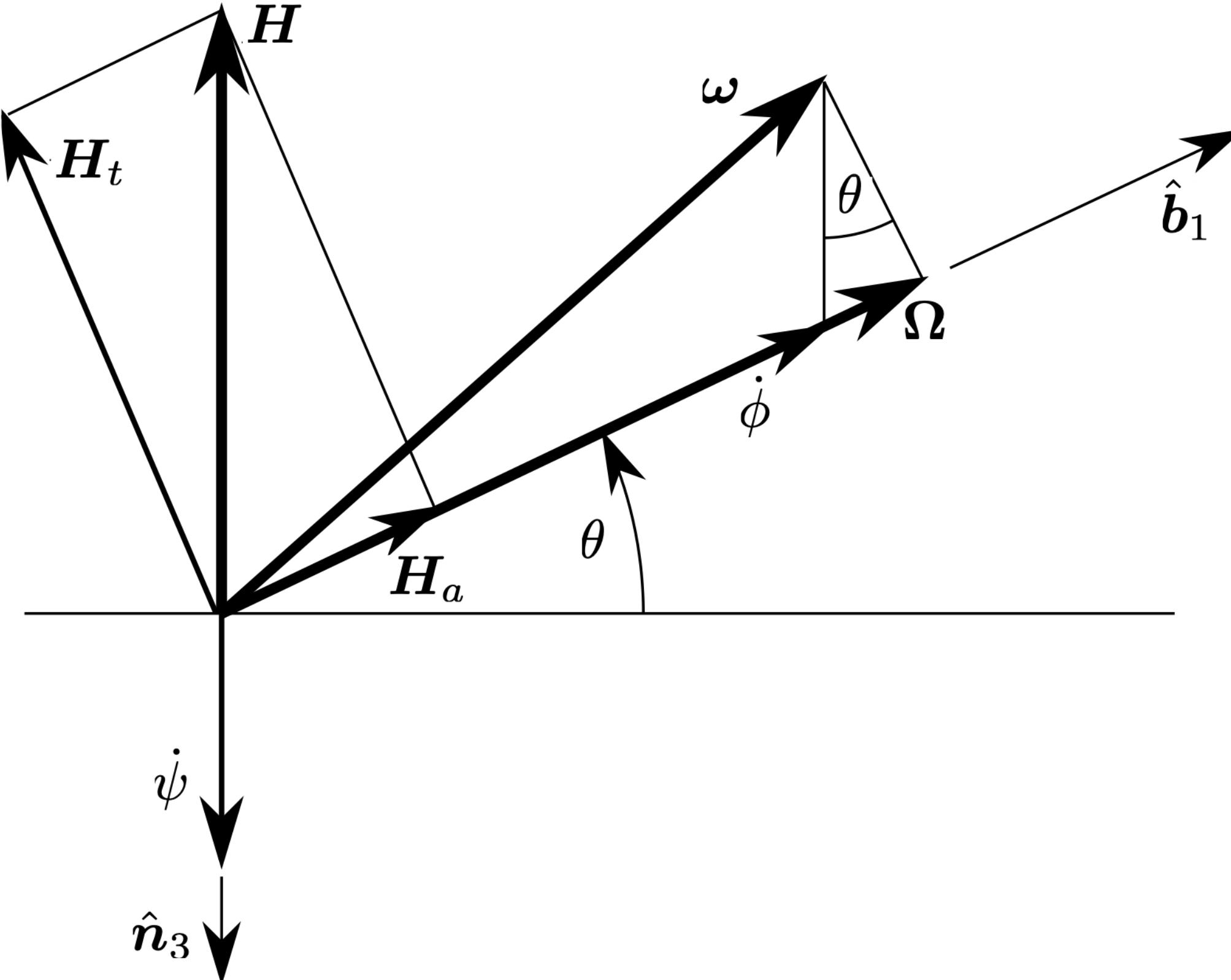
- The 3-2-1 Euler angle differential equation are then given by:

$$\dot{\psi} = -\frac{H}{I_2}$$

$$\dot{\theta} = 0$$

$$\dot{\phi} = H \left( \frac{I_2 - I_1}{I_1 I_2} \right) \sin \theta$$



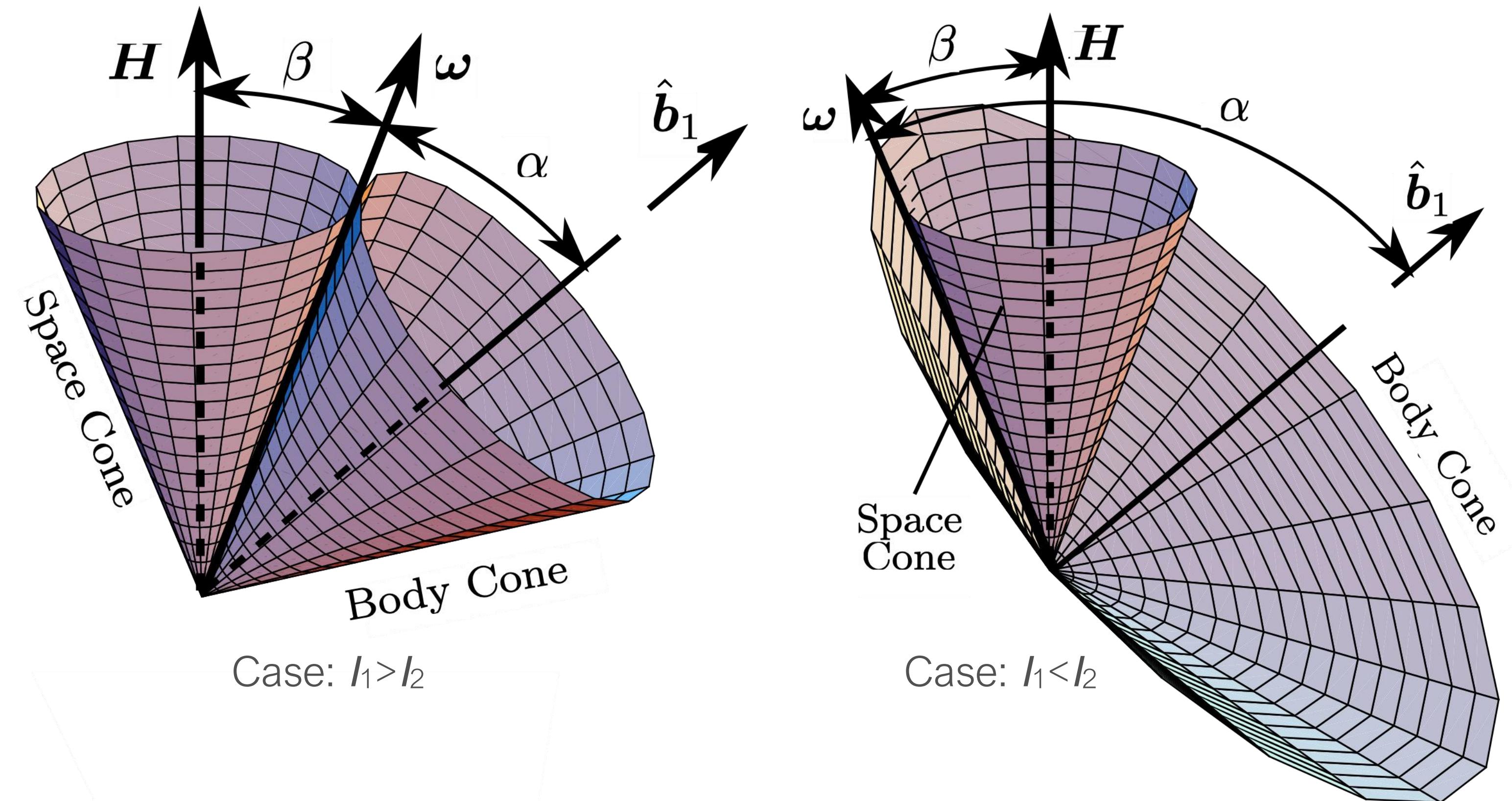


Let  $\Omega = \omega_1 \longrightarrow \Omega = \frac{H}{I_1} \sin \theta$

Note that for  $0 \leq \theta \leq \pi/2$   
we find that  $\Omega > 0$

The EOM can be written as

$$\dot{\psi} = -\frac{I_1}{I_2} \frac{\Omega}{\sin \theta} \quad \dot{\phi} = \frac{I_2 - I_1}{I_2} \Omega$$



Since the pitch angle  $\theta$  is shown to remain constant during this torque-free rotation, the resulting motion can be visualized by two cones rolling on each other. The **space cone** is fixed in space and its cone axis is always aligned with the angular momentum vector  $\mathbf{H}$ . The cone angle  $\beta$  is defined as the angle between the vectors  $\mathbf{H}$  and  $\boldsymbol{\omega}$ . The **body cone** axis is aligned with the first body axis and has the cone angle  $\alpha$  which is the angle between  $\boldsymbol{\omega}$  and first body axis.