

Optimal Trajectories ASEN 6020

Definition, Computation & Analysis of the
optimal design, guidance & control of space
trajectories and the properties of these trajectories.

Optimal → Depends on the situation & goals

Design & Control → Where do you want to go? What technology are
you using?

Guidance → Are you on track? What's the best way to correct
errors?

Space Trajectory → Motion in a vacuum, interface w/ "top" of atmosphere

Properties \rightarrow Is a trajectory stable or unstable?
Is a trajectory truly optimal?

Analysis \rightarrow Math. Properties

Computation \rightarrow Realization of the theory.

2 basic approaches

Parametric Optimization

- Set up a sequence of 2 body arcs & optimize total ΔV
- Specify a control law & choose \vec{a}_i

$$\vec{a} = \vec{a}_0 + \vec{a}_1 t + \vec{a}_2 t^2 + \dots$$

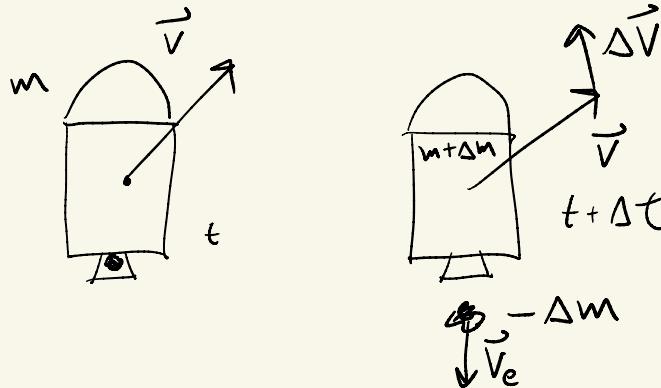
Functional Optimization

- How to continuously choose your control law $\vec{a}(t)$ to satisfy optimality conditions ...
 - $\vec{a}(t)$ "infinite dimensional"
 - Ex: Min. fuel traj. from LEO + Luna orbit

Basics

Rocket Equation

$$g = 0$$



Cons. of Momentum

$$m\vec{v} = (m + \Delta m)(\vec{v} + \Delta \vec{v}) - \Delta m(\vec{v}_e + \vec{v})$$

$$\boxed{\Delta m \vec{v}_e = m \Delta \vec{v}}$$

turn into
a rate

$$\left(\frac{dm}{dt} \vec{v}_e \right) = m \frac{d\vec{v}}{dt} = \underline{\text{Thrust}}$$

$$\frac{d\vec{v}}{dt} = \frac{1}{m} \frac{dm}{dt} \vec{v}_e \quad \frac{dm}{dt} < 0, \vec{v}_e \text{ opposes acceleration}$$

Assume $\vec{v}_e \approx \text{constant}$ \Rightarrow

$$d\vec{v} = \frac{1}{m} dm \vec{v}_e \Rightarrow \boxed{\Delta \vec{v} = \vec{v}_e \ln\left(\frac{m_0}{m_f}\right)} = -\vec{v}_e \ln\left(\frac{m_0}{m_f}\right)$$

$$|\vec{V}_e| = c \sim I_{sp}$$

$$|\Delta \vec{V}| = \Delta V = c \ln\left(\frac{m_0}{m_f}\right) \Rightarrow \text{Direct link between } \Delta V + \text{Payload}$$

Series of ΔV 's

$$\sum_{i=0}^{N-1} \Delta V_i = c \sum_{i=0}^N \ln\left(\frac{m_i}{m_{i-1}}\right) = c \ln\left(\frac{m_0}{m_{N-1}}\right)$$

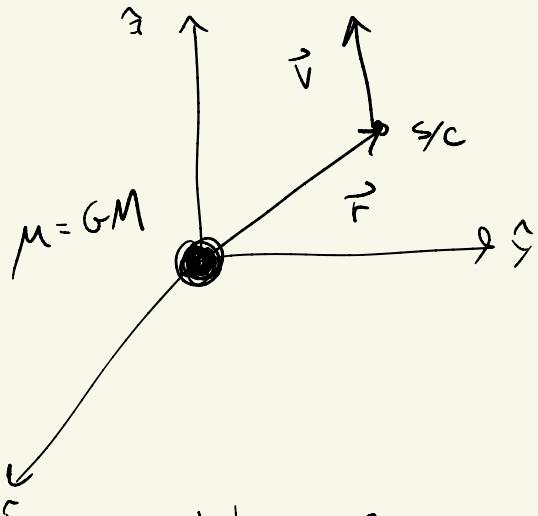
General results

$$\boxed{\frac{d\vec{v}}{dt} = \frac{1}{m} (\dot{m} \vec{V}_e - \vec{c})}$$

\Rightarrow mass can become an additional "state"
 $(\vec{r}, \vec{v}) \Rightarrow (\vec{r}, \vec{v}, m)$

$$\dot{m} = f(m, \vec{r}, t)$$

2-Body Orbit Mechanics



$$\begin{cases} \vec{v} = \frac{d\vec{r}}{dt} = \dot{\vec{r}} \\ m\ddot{\vec{v}} = -\frac{\mu m}{r^3} \vec{r} + \vec{T} \end{cases}$$

$$\boxed{\ddot{\vec{v}} = -\frac{\mu}{r^3} \vec{r} + \frac{1}{m} \vec{a}(t)}$$

= "acceleration control"

When $\dot{m} = 0 \Rightarrow \ddot{\vec{v}} = -\frac{\mu}{r^3} \vec{r} \Rightarrow \text{Integrate (completely solved)}$

Impulsive Moves: When $\dot{m} \neq 0$ over a short time interval, & if we can assume \vec{r} is relatively stationary, then ...

$$\frac{d\vec{v}}{dt} = -\frac{M}{r^3} \vec{r} + \frac{1}{m} \vec{c} \frac{dm}{dt}$$

\Rightarrow

$$d\vec{v} = -\frac{M}{r^3} \vec{r} dt + \frac{1}{m} \vec{c} dm$$

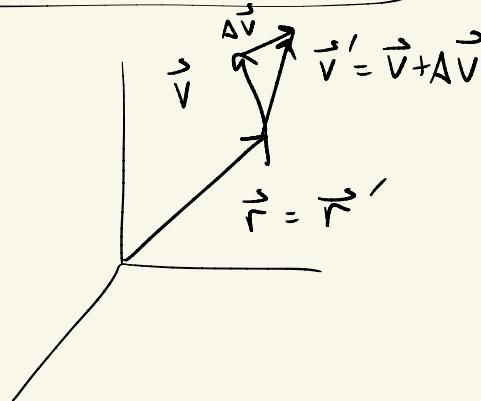
$$\int_0^{At} d\vec{v} = \int_0^{At} -\frac{M}{r^3} \vec{r} dt + \int_0^{At} \frac{\vec{c}}{m} dm$$

$$\Rightarrow |\vec{g}| \ll |\vec{T}|$$

$$\Delta \vec{v} = \frac{1}{m} \vec{c} \ln\left(\frac{m_0}{m_f}\right)$$

\Rightarrow "Impulsive Change" in velocity

Instantaneous change in velocity.



Kepler's Problem

Integrals of Motion:

$$\left\{ \begin{array}{l} E = \frac{1}{2} \vec{v} \cdot \vec{v} - \frac{M}{r} \quad \text{Energy} \\ \vec{H} = \vec{r} \times \vec{v} \quad \text{Ang. Mmnt} \\ |\vec{H}| = |\vec{r} \times \vec{v}| = rV \cos \gamma \quad \gamma = \text{Flight path angle} \\ \vec{e} = \frac{1}{M} \vec{v} \times \hat{\vec{H}} - \hat{\vec{r}} \end{array} \right.$$

6 orbit elements

$$a = -\frac{M}{2E} ; \quad e^2 = 1 + \frac{2EH^2}{M^2} ; \quad e = |\vec{e}| \quad \tan 2\nu = \frac{\vec{r} \cdot (\hat{\vec{z}} \times \hat{\vec{e}})}{\vec{r} \cdot \vec{e}}$$

$$\cos i = \hat{\vec{z}} \cdot \hat{\vec{H}} \quad \tan \Omega = \frac{\hat{x} \cdot \hat{\vec{H}}}{\hat{y} \cdot \hat{\vec{H}}}$$

$$\tan \omega = \dots$$

a = semi-major axis

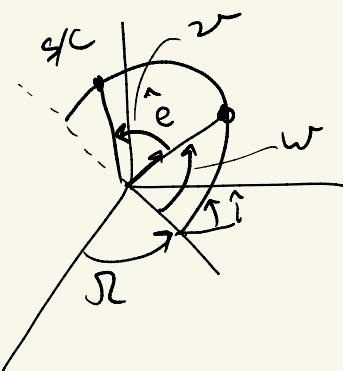
e = eccentricity

i = inclination

ν = true anomaly

Ω = long. of asc. node

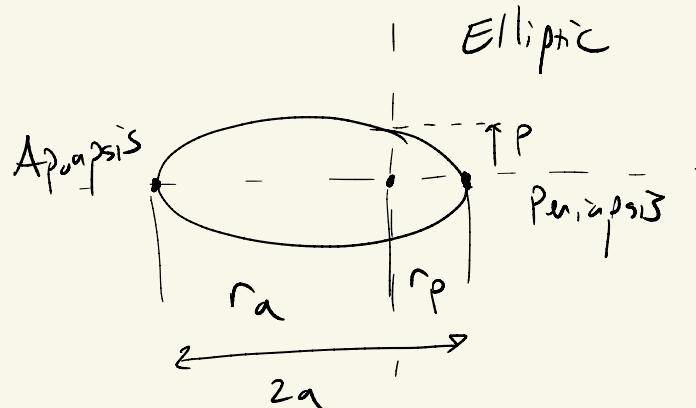
ω = arg. of periapsis



$$r = \frac{p}{1 + e \cos \alpha}$$

$$V = \sqrt{\frac{M}{p}} \sqrt{1 + 2e \cos \alpha + e^2}$$

$E=0 \Rightarrow$ parabolic, $a \rightarrow \pm\infty$, $e=1$



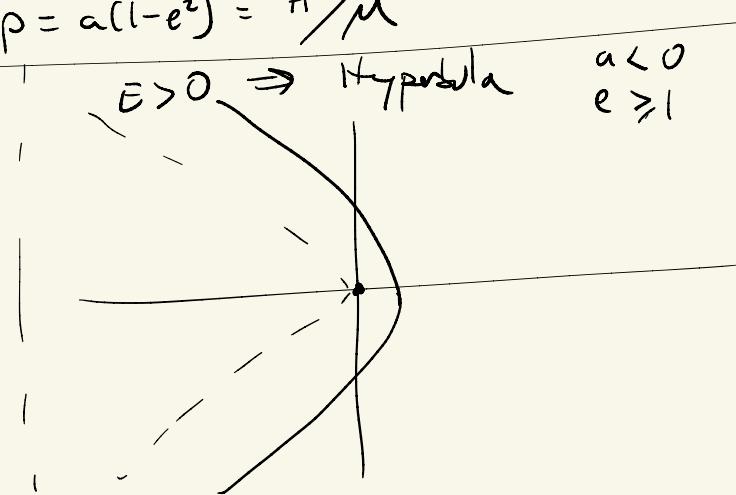
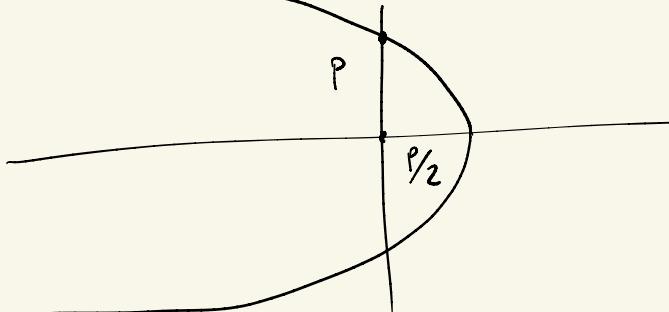
$E < 0$
 $a > 0$
 $1 \geq e \geq 0$

$$\tan \delta = \frac{e \sin \alpha}{1 + e \cos \alpha}$$

$$p = a(1 - e^2) = \frac{h^2}{\mu}$$

$E > 0 \Rightarrow$ hyperbola

$a < 0$
 $e \geq 1$



Important "points" on three orbits

Periapsis $\Rightarrow r_{\min}, v = 0 ; r_p = a(1-e) = \frac{P}{1+e}$

Always exists (r_p, g)

$$v_p = \sqrt{\frac{M}{P}} (1+e) = \sqrt{\frac{M(1+e)}{a(1-e)}} +$$

Apoapsis $\Rightarrow r_{\max}, v = \pi$ $\left(\begin{array}{l} = a \cos\left(-\frac{1}{e}\right) \\ r_{\max} \rightarrow \infty \end{array} \right)$

Does not

always exist

$$r_a = \frac{P}{1-e} = a(1+e)$$

$$v_a = \sqrt{\frac{M}{P}} (1-e) = \sqrt{\frac{M(1-e)}{a(1+e)}} +$$

Note:

$$a = \frac{1}{2}(r_a + r_p)$$

$$e = \frac{r_a - r_p}{r_a + r_p}$$

$$v_p = \sqrt{\frac{2M}{(r_a + r_p)}} \frac{r_a}{r_p} +$$

$$v_a = \sqrt{\frac{2M}{(r_a + r_p)}} \frac{r_p}{r_a} +$$

Some special orbits:

For a given AM , what is the minimum energy orbit? Circular

$$E = -\frac{M}{2a} ; H = \sqrt{\mu a(1-e^2)} \Rightarrow H^2 = \mu a(1-e^2)$$

$$\left[E = -\frac{M}{2} \cdot \frac{\mu(1-e^2)}{H^2} \right] = \left[\frac{-\mu^2}{2H^2} + \frac{\mu^2 e^2}{2H^2} \right] \Rightarrow \text{drive } e \rightarrow 0 \Rightarrow \text{Circular orbit}$$

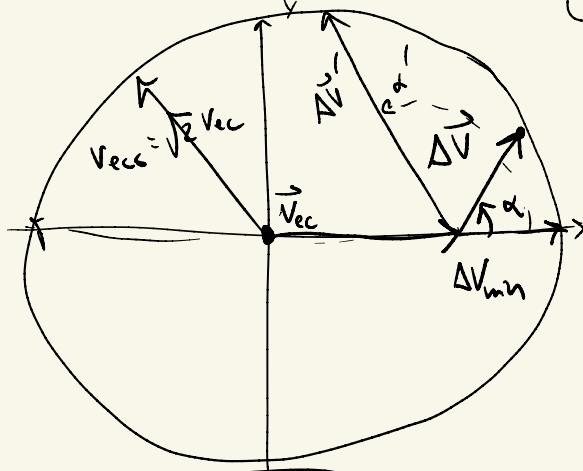
What is the minimum increment in velocity (ΔV) to make a S/C escape from a circular orbit? at radius "r"

$$E_{circ} = \frac{1}{2} \frac{M}{r} - \frac{M}{r} = -\frac{M}{2r}$$

$$E' = \left[\frac{1}{2} (\vec{V}_{ec} + \Delta \vec{v}) \cdot (\vec{V}_{ec} + \Delta \vec{v}) - \frac{M}{r} \right] = 0 \Rightarrow \vec{V}_{ec} + \Delta \vec{v} = \vec{V}_{esc}$$

$$(\vec{V}_{ec} + \Delta \vec{V}) \cdot (\vec{V}_{ec} + \Delta \vec{V}) = \frac{2M}{r} = V_{esc}^2 \quad \checkmark$$

$$\Delta V^2 + 2\vec{V}_{ec} \cdot \Delta \vec{V} + \frac{M}{r} - \frac{2M}{r} = \boxed{\Delta V^2 + 2V_{ec} \Delta V \cos \alpha - \frac{M}{r} = 0}$$



$$\Delta V_{min} = (\sqrt{2} - 1) V_{ec}$$

$$\Delta V = V_{ec} \left[\sqrt{1 + \cos^2 \alpha} - \cos \alpha \right]$$

$$\frac{\partial \Delta V}{\partial \alpha} = V_{ec} \sin \alpha \left[1 - \frac{\cos \alpha}{\sqrt{1 + \cos^2 \alpha}} \right] = 0 \Rightarrow \alpha = 0, \pi$$

$$\frac{\partial^2 \Delta V}{\partial \alpha^2} \approx V_{ec} \cos \alpha \left[1 - \frac{\cos \alpha}{\sqrt{1 + \cos^2 \alpha}} \right] + V_{ec} \sin \alpha \left[\dots \right]$$

$$\alpha = \pi \quad -V_{ec} \left[1 + \frac{1}{\sqrt{2}} \right] < 0$$

$$\alpha = 0 \quad V_{ec} \left[1 - \frac{1}{\sqrt{2}} \right] > 0 \Rightarrow \underline{\text{minimum}}$$

Other questions

Where is the best location to change energy?
 " " " " " " " " A. P. ?

$$\ddot{\vec{r}} = -\frac{\mu}{r^3} \vec{r} + \vec{a}$$

$\vec{E} \Rightarrow \dot{\vec{E}} = \vec{V} \cdot \vec{a} \Rightarrow$ where $|\vec{V}|$ is max & $\vec{a} \parallel \vec{V}$
 perigapsis

$\vec{H} \Rightarrow \dot{\vec{H}} = \vec{r} \times \vec{a} \Rightarrow$ where $|\vec{r}|$ is max & $\vec{a} \perp \vec{V}$

$$\vec{e} \Rightarrow \dot{\vec{e}} = \frac{1}{\mu} [\vec{a} \times (\vec{r} \times \vec{v}) + \vec{V} \times (\vec{r} \times \vec{a})]$$

apogapsis

$$e^2 = 1 + \frac{2EH^2}{\mu^2} \Rightarrow e \Delta e = \frac{4EH}{\mu^2} \Delta H + \frac{2H^2 \Delta E}{\mu^2}$$