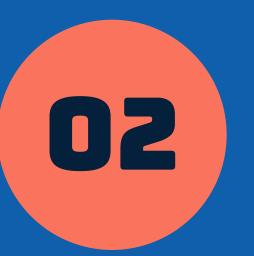


### ORBITAL MECHANICS





Varying Velocities



**Orbital Period** 

#### **KEPLER 1ST**

```
//calc position on the ellipse
//a = semi-major axis | b = semi-minor axis
//Keplers 1st law -> keeping planets on elliptical orbits around the sun
float x = a * Mathf.Cos(currentAngle);
float y = b * Mathf.Sin(currentAngle);
Vector3 desiredPosition =
new Vector3(
    particle.target.transform.position.x + x,
    particle.target.transform.position.y + y,
    particle.target.transform.position.z
);
```

# Parametric Equation of an Ellipse Sun Pos used as Offset

#### **KEPLER 2ND**

```
//calc velocity magnitude using vis-viva equation
//v = sqrt(G * M * (2/r - 1/a))
//Keplers 2nd law -> varying velocity on the ellipse
// = faster when closer to sun, slower when further YET same area convered
float velocityMagnitude = Mathf.Sqrt(gravityConstant * particle.sunMass * (2 / r - 1 / a));
```

# Higher orbital velocity nearing Major Vertex Lower orbital velocity nearing Minor Vertex Same Area!

$$v = \sqrt{G \cdot M \cdot \left(rac{2}{r} - rac{1}{a}
ight)}$$

#### **KEPLER 3RD**

```
//keplers 3rd law -> T = 2π * sqrt(a^3 / (G * M))
//the further the planet from the sn, the slower its orbitalperiod
float gravityConstant = Planet.gravConstant; // gravity constant
float a = particle.semiMajorAxis; // semi-major axis
float orbitalPeriod = 2 * Mathf.PI * Mathf.Sqrt(Mathf.Pow(a, 3) / (gravityConstant * particle.sunMass));
```

# Furthest Orbit to sun = Slowest Orbital Period Closest Orbit to sun = Fastest Orbital Period

$$T=2\pi\sqrt{rac{a^3}{G\cdot M}}$$



### GRAUITY



$$F_{g} = G \frac{m_{1}m_{2}}{r^{2}}$$

#### where

- $F_g$  is the force
- G is the gravitational constant (6.674 x  $10^{-11}$  m<sup>3</sup>.kg<sup>-1</sup>.s<sup>-2</sup>)
- ullet  $m_1$  and  $m_2$  are the masses of the objects
- r is the distance between the centers of the objects



#### GRAVITY

```
float velocityMagnitude = Mathf.Sqrt(gravityConstant * particle.sunMass * (2 / r - 1 / a));
```

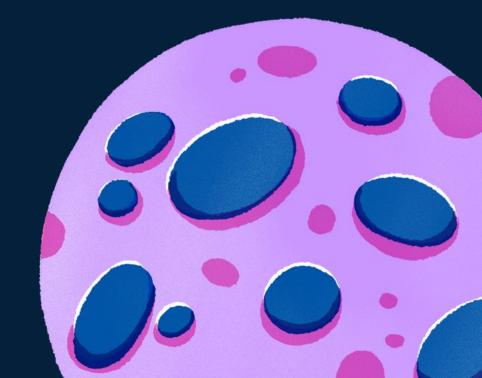
#### Velocity altered through vis-viva equation

```
Vector3 tangent = new Vector3(-Mathf.Sin(currentAngle), Mathf.Cos(currentAngle), 0).normalized;
```

#### Velocity direction altered through tangent vector



## TECHNICAL DEMO





### SCALING

Factor of: 2.8748 \* 10^-8

Sphere	Real Dist. From Sun	Scaled Dist.
Earth	1.496 * 10^11 M	4228.69
Mercury	5.79 * 10^10 M	1322.41
Neptune	4.503 * 10^12 M	128533.85





Due to floating-point precision limitations, it is recommended to bring the world coordinates of the GameObject within a smaller range.



#### **PROBLEMS**



- LineRenderer
- Gravitational Calculation

- All 3 Laws





# QUESTIONS

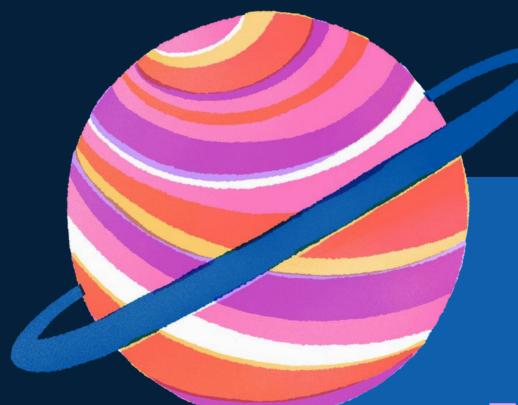


# \* THANKS:3

-JAYDEN







### Slides Carnival

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SlidesCarnival for the presentation template

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#### HAPPY DESIGNING!