

Jie Ren

Contents

1	Real Klein-Gordon Field Theory			1
	1.1	Quant	tization of Free Field	1
		1.1.1	Path Integral Formalism	1
		1.1.2	Canonical Quantization	
		1.1.3	Scattering Amplitude	5
	1.2	ϕ^3 Th	eory in $(d = 6 - \epsilon)$ Space-time	5
		1.2.1	Self Energy Correction	6
		1.2.2	Vertex Correction	10
		1.2.3	Renormalization Group	11
	1.3	ϕ^4 Th	eory in $(d = 4 - \epsilon)$ Space-time	13
		1.3.1	One-loop Correction	13
		1.3.2	Renormalization Group	15
2	Quantum Electrodynamics 1			
	2.1	Free F	Field Theory	16
		2.1.1	Dirac Field	
		2.1.2	Electromagnetic Field	
		2.1.3	Perturbation Theory	
	2.2	Loop	Correction	
		2.2.1	Electron Propagator	
		2.2.2	Photon Self-energy	
		2.2.3	Vertex Correction	
		2.2.4	Renormalization Group	24
3	Non-relativistic QFT			
	3.1		Temperature Field Theory	26 27
	ŭ - <u>-</u>		Free Field Theory	

Chapter 1

Real Klein-Gordon Field Theory

In this note, we use the (+, -, -, -) metric, where the inner product of two 4-momentum and 4-coordinate is

$$k \cdot x = \omega t - \vec{k} \cdot \vec{x}. \tag{1.1}$$

In this chapter, we consider the real Klein-Gordon field. The free field Lagrangian density:

$$\mathcal{L}_0 = \frac{1}{2} \partial^{\mu} \phi \partial_{\mu} \phi - \frac{m^2}{2} \phi^2 \simeq -\frac{1}{2} \phi (\partial^2 + m^2) \phi. \tag{1.2}$$

The action for free field with source is

$$S_0[\phi, J] = \int d^d x \left(\frac{1}{2} \partial^\mu \phi \partial_\mu \phi - \frac{m^2}{2} \phi^2 + J \cdot \phi \right). \tag{1.3}$$

The space-time Fourier transformation is defined as

$$\tilde{\phi}(k) = \int d^d x e^{ik \cdot x} \phi(x),$$

$$\phi(x) = \int \frac{d^d k}{(2\pi)^d} e^{-ik \cdot x} \tilde{\phi}(k).$$
(1.4)

The Lagrangian in momentum space is

$$\tilde{\mathcal{L}}_0[\phi_k, J] = \tilde{\phi}(k)(k^2 - m^2)\tilde{\phi}(-k) + \tilde{J}(k) \cdot \tilde{\phi}(-k) + \tilde{\phi}(k) \cdot \tilde{J}(-k).$$

1.1 Quantization of Free Field

1.1.1 Path Integral Formalism

In the path integral formula,

$$Z_0[J] = \int D[\phi] \exp(iS_0[\phi, J]). \tag{1.5}$$

The partition function for free field:

$$\frac{Z_0[J]}{Z_0[0]} = \exp\left(-\frac{i}{2} \int \frac{d^d k}{(2\pi)^d} \frac{\tilde{J}(k)\tilde{J}(-k)}{k^2 - m^2 + i\epsilon}\right)
= \exp\left(-\frac{i}{2} \int d^d x_1 d^d x_2 J(x_1) \Delta_0(x_1 - x_2) J(x_2)\right).$$

where the propagator is¹

$$\Delta_0(x_1 - x_2) = \int \frac{d^d k}{(2\pi)^d} \frac{e^{-ik \cdot x}}{k^2 - m^2 + i\epsilon}.$$
 (1.6)

Remark 1. Gaussian Integral for Real Scalar Field

The real Gaussian integral formula is

$$\int d\boldsymbol{v} \exp\left(-\frac{1}{2}\boldsymbol{v}^T \cdot A \cdot \boldsymbol{v} + \boldsymbol{b}^T \cdot \boldsymbol{v}\right) = \sqrt{\frac{(2\pi)^N}{\det A}} \exp\left(\frac{1}{2}\boldsymbol{b}^T \cdot A^{-1} \cdot \boldsymbol{b}\right), \quad (1.7)$$

where $\boldsymbol{v}, \boldsymbol{b}$ are two N-dimensional vector, and A is an $N \times N$ matrix. For the field integral, we absorbed the $(2\pi)^{N/2}$ term into the measure, and express the path integral for the Gaussian field as:

$$Z[J] = \int D[\phi] \exp\left(\frac{i}{2} \int d^d x \phi \hat{A} \phi + i \int d^d x J \phi\right)$$
$$= Z[0] \exp\left[-\frac{i}{2} \int d^d x_1 d^d x_2 J(x_1) A^{-1}(x_1 - x_2) J(x_2)\right].$$

We make use of (1.7) by making the identification

$$A = \bigoplus_{|k|} \begin{pmatrix} 0 & k^2 - m^2 \\ k^2 - m^2 & 0 \end{pmatrix}, \ b = \bigoplus_{|k|} \begin{pmatrix} \tilde{J}(k) \\ \tilde{J}(-k) \end{pmatrix}.$$

This gives the propagator in the momentum space:

$$\tilde{\Delta}_0(k) = \frac{1}{k^2 - m^2}.$$

Note that $\Delta_0(x_1 - x_2)$ is related to the correlation function:

$$\langle 0|T\phi(x_1)\phi(x_2)|0\rangle = \frac{\delta}{i\delta J(x_1)} \frac{\delta}{i\delta J(x_2)} Z_0[J]$$

$$= i\Delta(x_1 - x_2). \tag{1.8}$$

For interaction theory, the partition function is

$$Z[J] = \exp\left(i \int d^d x \mathcal{L}_{int} \left[\frac{\delta}{i\delta J(x)}\right]\right) Z_0[J]. \tag{1.9}$$

¹The extra $i\epsilon$ term is use to bring the singularities infinitesimally below the real axis. This infinitesimal value can be absorbed into the mass term, by regarding the mass term m^2 as $m^2 - i\epsilon$.

The expectation values for a generic operator of the form $O(\phi)$ can be evaluated by the true partition function

$$\langle O(\phi) \rangle = \frac{1}{Z[0]} O\left[\frac{\delta}{i\delta J(x)}\right] Z[J] \bigg|_{J=0}.$$
 (1.10)

Remark 2. Connected Diagrams

The expression (1.10) can be expanded order by order using the Feynman diagram. Since the unconnected diagram can be absorbed into Z[0], we only need to calculate the connected diagram.

The procedure of perturbative expansion with only connected diagrams can be formally represented by introducing the quantity

$$Z[J] = \exp(iW[J]). \tag{1.11}$$

The perturbative expansion of W[J] contain only the connected diagrams. Eq. (1.10) can then be rephrased as

$$\langle O(\phi) \rangle = O\left[\frac{\delta}{\delta J(x)}\right] W[J] \bigg|_{I=0}.$$
 (1.12)

For example, the two-point connected correlation (propagator) is

$$\langle \mathcal{T}\phi(x_1)\phi(x_2)\rangle_c = \frac{1}{i} \left. \frac{\delta^2 W[J]}{\delta J(x_1)\delta J(x_2)} \right|_{J=0}$$

$$= -\left. \frac{\delta^2}{\delta J(x_1)\delta J(x_2)} \ln Z[J] \right|_{J=0}$$

$$= \frac{1}{Z[0]} \left. \frac{\delta^2 Z[J]}{i\delta J(x_1)i\delta J(x_2)} \right|_{J=0}.$$
(1.13)

1.1.2 Canonical Quantization

The classical equation of motion for the real Klein-Gordon field is

$$(-\partial_t^2 + \nabla^2 - m^2)\phi(\vec{x}, t) = 0.$$
 (1.14)

The solution to Eq. (1.14) is proportional to the plane wave:

$$\phi(\vec{x},t) \propto e^{-i\omega_{\mathbf{k}}t + i\vec{p}\cdot\vec{x}} + e^{i\omega_{\vec{k}}t - i\mathbf{p}\cdot\vec{x}},\tag{1.15}$$

where the energy is $\omega_{\mathbf{k}} = \mathbf{k}^2 + m^2$ and \vec{k} is the momentum as the conserved quantity. The general solution to the EOM is

$$\phi(\vec{x},t) \propto \int \frac{d^d k}{(2\pi)^d} \left(a_k e^{-i\omega_k t + i\vec{k}\cdot\vec{x}} + a_k^* e^{i\omega_k t - i\vec{k}\cdot\vec{x}} \right). \tag{1.16}$$

The canonical quantization promote the coefficient a_k/a_k^* to the particle annihilation/creation operator a_k/a_k^{\dagger} , with the commutation relation

$$[a_k, a_p^{\dagger}] = (2\pi)^d \delta^d(\vec{k} - \vec{p}).$$
 (1.17)

The single-particle state with momentum \vec{k} is created by a_k^{\dagger} operators acting on the vacuum:

 $a_k^{\dagger}|0\rangle = \frac{1}{\sqrt{2\omega_k}}|\vec{k}\rangle,$ (1.18)

where $|\vec{k}\rangle$ is a state with a single particle of momentum \vec{k} .

Remark 3. Lorentz Invariance of Single-particle State

The factor of $\sqrt{2\omega_k}$ in Eq. (1.18) is just a convention, but it will make some calculations easier. To compute the normalization of one-particle states, we start with

$$\langle 0|0\rangle = 1,\tag{1.19}$$

which leads to

$$\langle \vec{p} | \vec{k} \rangle = 2\sqrt{\omega_{p}\omega_{k}} \left\langle 0 \left| a_{p}a_{k}^{\dagger} \right| 0 \right\rangle = 2\omega_{p}(2\pi)^{d} \delta^{d}(\vec{p} - \vec{k}).$$
 (1.20)

The identity operator for one-particle states is

$$1 = \int \frac{d^d p}{(2\pi)^d} \frac{1}{2\omega_{\mathbf{p}}} |\vec{p}\rangle\langle\vec{p}|, \qquad (1.21)$$

which we can check with

$$|\vec{k}\rangle = \int \frac{d^d p}{(2\pi)^d} \frac{1}{2\omega_p} |\vec{p}\rangle \langle \vec{p}|\vec{k}\rangle = \int \frac{d^d p}{(2\pi)^d} \frac{1}{2\omega_p} 2\omega_p (2\pi)^d \delta^d (\vec{p} - \vec{k}) |\vec{p}\rangle = |\vec{k}\rangle.$$

The identity operator Eq. (1.21) is Lorentz invariant since it can be expressed as

$$1 = \int \frac{d^d p}{(2\pi)^d} \int \frac{d\omega}{2\pi} 2\pi \delta(\omega^2 - \boldsymbol{p}^2 - m^2). \tag{1.22}$$

By requiring $\langle \vec{k} | \phi | 0 \rangle = 1$, the normalization Eq. (1.18) determines the normalization for the quantized field operator:

$$\phi(\vec{x},t) = \int \frac{d^dk}{(2\pi)^d} \frac{1}{\sqrt{2\omega_k}} \left(a_k e^{-ik\cdot x} + a_k^{\dagger} e^{ik\cdot x} \right). \tag{1.23}$$

1.1.3 Scattering Amplitude

Consider the scattering process in the interaction picture,

$$\langle f|e^{-iHt}|i\rangle = \langle f|T\exp\left(-i\int dt V_{\rm int}(t)\right)|i\rangle$$

$$= \langle f|T\exp\left(i\int d^dx \mathcal{L}_{\rm int}(t)\right)|i\rangle.$$
(1.24)

The S-matrix is defined as

$$S = \mathcal{T} \exp\left(i \int d^d x \mathcal{L}_{int}(t)\right) = 1 + i\mathcal{T}.$$
 (1.25)

Because of the additional momentum conservation,

$$\mathcal{T} = (2\pi)^d \delta^d \left(\sum p\right) \mathcal{M}. \tag{1.26}$$

1.2 ϕ^3 Theory in $(d = 6 - \epsilon)$ Space-time

Now consider the interaction theory with additional Lagrangian

$$\mathcal{L}_{int}[\phi] = \frac{g}{3!}\phi^3. \tag{1.27}$$

Note that the field ϕ has the mass dimension $\left[\frac{d-2}{2}\right]$. When d=6, the coupling constant g is dimensionless. For interaction theory, the renormalized Lagrangian has the form:

$$\mathcal{L} = Z_{\phi} \frac{1}{2} \partial^{\mu} \phi \partial_{\mu} \phi - Z_{m} \frac{m^{2}}{2} \phi^{2} + Z_{g} \frac{g}{3!} \phi^{3}$$

$$= \mathcal{L}_{0} + \mathcal{L}_{int} + \mathcal{L}_{ct},$$
(1.28)

where the counter terms are:

$$\mathcal{L}_{ct}[\phi] = \frac{A}{2} \partial^{\mu} \phi \partial_{\mu} \phi - \frac{B}{2} m^2 \phi^2 + \frac{C}{3!} g \phi^3$$

$$\simeq -\frac{A}{2} \phi \partial^2 \phi - \frac{B}{2} m^2 \phi^2 + \frac{C}{3!} g \phi^3,$$
(1.29)

where

$$A = Z_{\phi} - 1, B = Z_m - 1, C = Z_g - 1.$$

The counter term for the the free field gives additional correction

$$i\tilde{\Delta}^{(\text{ct})}(k) = i\tilde{\Delta}_0(k)(Ak^2 - Bm^2)i\tilde{\Delta}_0(k)$$

$$= k \qquad k$$
(1.30)

1.2.1 Self Energy Correction

To second order, we consider the one-loop correction to the propagator with the diagram:

This correspond to

$$i\tilde{\Delta}^{(2)}(k) = i\tilde{\Delta}_0(k) \left[i\Sigma^{(2)}(k^2) \right] i\tilde{\Delta}_0(k), \tag{1.31}$$

where the self energy term to the second order $i\Pi^{(2)}(k)$ is defined as:

$$i\Sigma^{(2)}(k^2) \equiv \frac{g^2}{2} \int \frac{d^d q}{(2\pi)^d} \tilde{\Delta}_0(q) \tilde{\Delta}_0(k-q) + (Ak^2 - Bm^2).$$
 (1.32)

Remark 4. Symmetry Factor

The coefficient $g^2/2$ comes from the symmetry factor in the diagram. We can also check the coefficient explicitly, by considering the expansion to the second order (we denote $\delta/\delta J(x_i)$ as δ_i):

$$\delta_1 \delta_2 \frac{1}{2!4!} \left[\frac{ig}{3!} \int d^d y \left(\frac{\delta}{\delta J(y)} \right)^3 \right]^2 \left[-\frac{i}{2} \int d^d y_1 d^d y_2 J(y_1) \Delta(y_1 - y_2) J(y_2) \right]^4.$$

The expansion gives the coefficient

$$\left(\frac{ig}{6}\right)^2 \times \frac{1}{2! \times 4! \times 2^4}.$$

Now consider the combinatorial factor, which comes from the exchange of $\phi(x_i)$ in the propagator, the exchange of $\phi(x_i)$ in the vertex, the exchange of propagator in the diagram, and the change of vertices in the diagram:

$$(2!)^4 \times (3!)^2 \times (4 \times 3) \times 2.$$

Those two factors produce a $-g^2/2$ coefficient. Note that in the self energy expression (1.31), we put a *i* factor in front of each propagator, which absorbs the minus sign.

Once we obtain the self energy, the one-loop corrected propagator has the form:

$$i\tilde{\Delta}(k) = i\tilde{\Delta}_0(k) + i\tilde{\Delta}_0(k) \left[\sum_{n=1}^{\infty} i\Sigma(k^2) \right] i\tilde{\Delta}_0(k)$$

$$= \frac{i}{\tilde{\Delta}_0^{-1}(k) - \Sigma(k^2)}$$

$$= \frac{i}{k^2 - m^2 - \Sigma(k^2)}.$$
(1.33)

Now we are going to evaluate the divergent integral in the self energy expression, using the Feynman parameters:

$$\int \frac{d^d q}{(2\pi)^d} \frac{1}{q^2 - m^2} \frac{1}{(k - q)^2 - m^2}$$

$$= \int \frac{d^d q}{(2\pi)^d} \int_0^1 dx \frac{1}{[q^2 - m^2 + x((q - k)^2 - q^2)]^2}$$

$$= \int_0^1 dx \int \frac{d^d q}{(2\pi)^d} \frac{1}{[(q - kx)^2 - D]^2},$$

where $D = m^2 - k^2 x (1 - x)$. Then we can shift $q \to q + kx$ leaving an integral that only depends on q^2 . In this way,

$$\Sigma(k^2) = \int_0^1 I(x)dx.$$

To evaluate the self-energy, it suffices to obtain the integral

$$I(x) = \frac{g^2}{2i} \int \frac{d^d q}{(2\pi)^d} \frac{1}{[q^2 - D]^2}.$$

Remark 5. Feynman Parameters

We use Feynman's formula to combine denominators,

$$\frac{1}{A_1 \dots A_n} = \int dF_n \left(x_1 A_1 + \dots + x_n A_n \right)^{-n}, \tag{1.34}$$

where the integration measure over the Feynman parameters x_i is

$$\int dF_n = (n-1)! \int_0^1 dx_1 \dots dx_n \delta(x_1 + \dots + x_n - 1). \tag{1.35}$$

This measure is normalized so that $\int dF_n = 1$. The simplest case is

$$\frac{1}{AB} = \int_0^1 \frac{dx}{[A + (B - A)x]^2} = \int_0^1 \frac{\delta(x + y - 1)}{[xA + yB]^2} dx dy.$$
 (1.36)

Other useful identities are

$$\frac{1}{AB^n} = \int_0^1 dx dy \frac{\delta(x+y-1)ny^{n-1}}{[xA+yB]^{n+1}},$$

$$\frac{1}{ABC} = \int_0^1 dx dy dz \frac{2\delta(x+y+z-1)}{[xA+yB+zC]^3}.$$
(1.37)

By making the Wick rotation $q^0 \to iq_E^0$, the integral becomes:²

$$I(x) = \frac{g}{2} \int \frac{d^d q_E}{(2\pi)^d} \frac{1}{(q_E^2 + D)^2} = \frac{g\Omega_d}{2(2\pi)^d} \int dq \frac{q^{d-1}}{(q^2 + D)^2}.$$

$$\Omega_d = \frac{2\pi^{d/2}}{\Gamma(\frac{d}{2})},\tag{1.38}$$

 $^{^{2}}$ The *d*-dimensional solid angle is

Dimensional Regularization

We set the dimension to $d = 6 - \epsilon$, and rewrite the Lagrangian as

$$\mathcal{L} = Z_{\phi} \frac{1}{2} \partial^{\mu} \phi \partial_{\mu} \phi - Z_{m} \frac{m^{2}}{2} \phi^{2} + Z_{g} \frac{g \tilde{\mu}^{\epsilon/2}}{3!} \phi^{3}. \tag{1.40}$$

Note that the coupling constant should be changed to $g \to g\tilde{\mu}^{\epsilon/2}$ where μ is of mass dimension [1] in order to get the correct dimensionality. We then expand the expression to zeroth order of ϵ . A useful identity is:

$$\int dk \frac{k^a}{\left(k^2 + D\right)^b} = D^{\frac{a+1}{2}-b} \frac{\Gamma\left(\frac{a+1}{2}\right)\Gamma\left(b - \frac{a+1}{2}\right)}{2\Gamma(b)}.$$
(1.41)

Actually, we can compute the integral and series expansion in Mathematica all together:

```
omg=(2*Pi^(d/2))/(Gamma[d/2]);
cof=g^2*\[Mu]^(6-d)/2*omg/(2*Pi)^d;
int=cof*Integrate[q^(d-1)/(q^2+D)^2,{q,0,Infinity}][[1]];
map={g^2->\[Alpha]*(4*Pi)^3,EulerGamma->Subscript[\[Gamma],E]};
ans=Series[int/.{d->6-\[Epsilon]},{\[Epsilon],0,0}];
ans/.map//Simplify
```

The result is (where $\alpha \equiv g^2/(4\pi)^3$)

$$I(x) = \frac{\alpha D}{2} \left[\ln \left(\frac{De^{\gamma_E}}{4\pi \tilde{\mu}^2} \right) - \left(\frac{2}{\epsilon} + 1 \right) \right] + O(\epsilon).$$

Now insert $D = m^2 - k^2 x (1 - x)$. Note that

$$\int_0^1 dx D = m^2 - \frac{k^2}{6}.$$

This simplifies the result to

$$\Sigma^{(2)}(k^2) = \frac{\alpha}{2} \left(\frac{2}{\epsilon} + 1 \right) \left(\frac{k^2}{2} - m^2 \right) + \frac{\alpha}{2} \int_0^1 dx D(x) \ln \left(\frac{D(x)}{\mu^2} \right), \tag{1.42}$$

where we have replace $\tilde{\mu}$ with

$$\mu \equiv \sqrt{\frac{4\pi}{e^{\gamma_E}}}\tilde{\mu}.\tag{1.43}$$

where $\Gamma(x)$ is the gamma function, satisfing

$$\Gamma(1+x) = x\Gamma(x), \ \Gamma(\epsilon) = \frac{1}{\epsilon} - \gamma_E + O(\epsilon).$$
 (1.39)

In particular, $\Gamma(n+1) = n!$.

Renormalization

The counter terms also contribute to the perturbative correction,

$$\begin{split} \Sigma^{(2)}\left(k^2\right) = & \frac{\alpha}{2} \int_0^1 dx D \ln\left(\frac{D}{m^2}\right) + \left\{\frac{\alpha}{6} \left[\frac{1}{\varepsilon} + \ln\left(\frac{\mu}{m}\right) + \frac{1}{2}\right] + A\right\} k^2 \\ & - \left\{\alpha \left[\frac{1}{\varepsilon} + \ln\left(\frac{\mu}{m}\right) + \frac{1}{2}\right] + B\right\} m^2 + O\left(\alpha^2\right). \end{split}$$

Consider the on-shell condition for the subtraction:

$$\Sigma(m^2) = \Sigma'(m^2) = 0. {(1.44)}$$

Set $D_0 \equiv D(x)|_{k^2=m^2} = m^2(1-x+x^2)$, the self energy has the form:

$$\Sigma^{(2)}(k^2) = \frac{\alpha}{2} \int_0^1 dx D(x) \ln\left(\frac{D(x)}{D_0(x)}\right) + C_k k^2 + C_m m^2.$$
 (1.45)

The condition $\Pi(m^2) = 0$ requires

$$\Sigma^{(2)}(k^2) = \frac{\alpha}{2} \int_0^1 dx D(x) \ln\left(\frac{D(x)}{D_0(x)}\right) + C_k(k^2 - m^2).$$

The condition $\Pi'(m^2) = 0$ requires

$$\frac{d\Sigma^{(2)}(k^2)}{dk^2}\Big|_{k^2=m^2} = \frac{\alpha}{2} \int_0^1 dx \left[\frac{D(x)}{dk^2} \ln \left(\frac{D(x)}{D_0(x)} \right) + D_0(x) \right] \Big|_{q^2=m^2} + C_k$$

$$= \frac{\alpha}{2} \int_0^1 dx (x^2 - x) + C_k$$

$$= C_k - \frac{\alpha}{12} = 0.$$

In this way, we obtained the renormalized self-energy:

$$\Sigma^{(2)}(k^2) = \frac{\alpha}{2} \int_0^1 dx D(x) \ln\left(\frac{D(x)}{D_0(x)}\right) + \frac{\alpha}{12} (k^2 - m^2). \tag{1.46}$$

On the other hand, we chan choose the $\overline{\rm MS}$ subtraction scheme, i.e.,

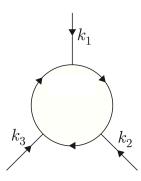
$$A = -\frac{\alpha}{6\epsilon}, \ B = -\frac{\alpha}{\epsilon}.\tag{1.47}$$

The self energy under $\overline{\rm MS}$ scheme will depend on the the mass scale μ we choose:

$$\Sigma^{(2)}\left(k^2\right) = \frac{\alpha}{2} \int_0^1 dx D \ln\left(\frac{D}{m^2}\right) + \alpha \left[\ln\left(\frac{\mu}{m}\right) + \frac{1}{2}\right] \left(\frac{k^2}{6} - m^2\right). \tag{1.48}$$

1.2.2 Vertex Correction

Now consider the simplest one-loop correction to the vertex function from the diagram:



The vertex function corresponding to such correction, together with the counter term, can be expressed as:

$$iV_3^{(3)}(k_1, k_2, k_3) = (ig)^3 i^3 \int \frac{d^a q}{(2\pi)^d} \tilde{\Delta}(q - k_1) \tilde{\Delta}(q + k_2) \tilde{\Delta}(q) + iCg, \tag{1.49}$$

Using the Feynman parameter, the integrant is

$$\tilde{\Delta}(q - k_1)\tilde{\Delta}(q + k_2)\tilde{\Delta}(q) = \int dF_3 \frac{1}{(q^2 - D)^3}$$

where we have shift the value of q, and D can be evaluate by the following code:

```
A1=(1-k1)^2-m^2;

A2=(1+k2)^2-m^2;

A3=(1)^2-m^2;

{c,b,a}=CoefficientList[x1*A1+x2*A2+(1-x1-x2)*A3,{1}];

-c+b^2/(4*a)//Expand
```

The result is

$$D = m^2 - k_1^2 x_1 (1 - x_1) - k_2^2 x_2 (1 - x_2) - 2k_1 k_2 x_1 x_2.$$

The same procedure gives:

$$V_3^{(3)}/g = \int dF_3 I(x_1, x_2, x_3) + C, \qquad (1.50)$$

where

$$I(x_1, x_2, x_3) = \frac{g^2 \Omega_d}{(2\pi)^d} \int dq \frac{q^{d-1}}{(q^2 + D)^3}.$$

The same regularization procedure in Mathematica:

The result is

$$V_3^{(3)}/g = \frac{\alpha}{\epsilon} + \frac{\alpha}{2} \int dF_3 \ln\left(\frac{4\pi\tilde{\mu}^2 e^{-\gamma_E}}{D}\right) + C + O(\epsilon)$$

$$= \frac{\alpha}{\epsilon} + \alpha \ln\left(\frac{\mu}{m}\right) - \frac{\alpha}{2} \int dF_3 \ln\left(\frac{D}{m}\right) + C.$$
(1.51)

The on-shell subtraction requires

$$V_3(0,0,0) = g, (1.52)$$

which gives

$$C = -\frac{\alpha}{\epsilon} - \alpha \ln \left(\frac{\mu}{m}\right). \tag{1.53}$$

So the vertex function to the third order is

$$V_3(k_1, k_2, k_3) = g \left\{ 1 - \frac{\alpha}{2} \int dF_3 \ln \left[\frac{D(x_1, x_2, x_3)}{m} \right] \right\}.$$
 (1.54)

The $\overline{\rm MS}$ scheme, on the other hand, sets

$$C = -\frac{\alpha}{\epsilon}.\tag{1.55}$$

1.2.3 Renormalization Group

We fist summarize the normalization factor obtained on the one-loop level (with $\overline{\rm MS}$ subtraction scheme):

$$Z_{\phi} = 1 - \frac{\alpha}{6\epsilon} + O(\alpha^{2}),$$

$$Z_{m} = 1 - \frac{\alpha}{\epsilon} + O(\alpha^{2}),$$

$$Z_{g} = 1 - \frac{\alpha}{\epsilon} + O(\alpha^{2}).$$
(1.56)

For the renormalized Lagrangian in $(6 - \epsilon)$ -dimension

$$\mathcal{L} = Z_{\phi} \frac{1}{2} \partial^{\mu} \phi \partial_{\mu} \phi - Z_{m} \frac{m^{2}}{2} \phi^{2} + Z_{g} \frac{g \tilde{\mu}^{\epsilon/2}}{3!} \phi^{3}, \tag{1.57}$$

the factors relate the original field and bare coefficients

$$\phi_0 = Z_{\phi}^{1/2} \phi, \ m_0 = Z_m^{1/2} Z_{\phi}^{-1/2} m, \ g_0 = Z_g Z_{\phi}^{-3/2} \tilde{\mu}^{\epsilon/2} g.$$
 (1.58)

The renormalization group requires that the bare parameter is independent of the mass scale μ we choose, that is:

$$\frac{d\phi_0}{d\ln\mu} = \frac{dm_0}{d\ln\mu} = \frac{dg_0}{d\ln\mu} = 0.$$
 (1.59)

Beta Function

Star with g_0 , it is more convenient to use

$$\alpha_0 \equiv \frac{g_0^2}{4\pi} = Z_g^2 Z_\phi^{-3} \tilde{\mu}^\epsilon \alpha. \tag{1.60}$$

Take logarithm on both side:

$$\ln \alpha_0 = \ln(Z_q^2 Z_\phi^{-3}) + \ln \alpha + \epsilon \ln \tilde{\mu}. \tag{1.61}$$

The RG equation is

$$\frac{d\ln\alpha_0}{d\ln\mu} = \frac{d\ln(Z_g^2 Z_\phi^{-3})}{d\alpha} \frac{d\alpha}{d\ln\mu} + \frac{1}{\alpha} \frac{d\alpha}{d\ln\mu} + \epsilon = 0.$$
 (1.62)

To the first order of α :

$$\frac{d\ln(Z_g^2 Z_\phi^{-3})}{d\alpha} = \frac{d}{d\alpha} \left(-\frac{2\alpha}{\epsilon} + \frac{\alpha}{2\epsilon} \right) = -\frac{3}{2\epsilon},\tag{1.63}$$

which leads to

$$\frac{d\alpha}{d\ln\mu}\left(1 - \frac{3\alpha}{2\epsilon} + O(\alpha^2)\right) + \epsilon\alpha = 0. \tag{1.64}$$

The beta function is defined as

$$\beta(\alpha) = \frac{d\alpha}{d \ln \mu} = \beta_1 \alpha + \beta_2 \alpha^2 + O(\alpha^3). \tag{1.65}$$

Insert such definition into the original expression, and keep track of the order of α , we get

$$(\beta_1 + \epsilon)\alpha + \left(\beta_2 - \frac{3\beta_1}{2\epsilon}\right)\alpha^2 + O(\alpha^3) = 0.$$
 (1.66)

The beta function is

$$\beta(\alpha) = -\epsilon \alpha - \frac{3}{2}\alpha^2 + O(\alpha^3). \tag{1.67}$$

Anomalous Dimension

Consider the RG equation with bare mass:

$$\frac{d \ln m_0}{d \ln \mu} = \frac{1}{2} \frac{d(\ln Z_m - \ln Z_\phi)}{d\alpha} \frac{d\alpha}{d \ln \mu} + \frac{1}{m} \frac{dm}{d \ln \mu}$$

$$= \frac{5\alpha}{12} + \frac{1}{m} \frac{dm}{d \ln \mu} + O(\alpha^2) = 0.$$
(1.68)

We get the anomalous dimension of the mass:

$$\gamma_m(\alpha) \equiv \frac{1}{m} \frac{dm}{d \ln \mu} = -\frac{5\alpha}{12} + O(\alpha^2). \tag{1.69}$$

Also, for the bare field

$$\frac{d\ln\phi_0}{d\ln\mu} = \frac{1}{2}\frac{d\ln Z_\phi}{d\ln\mu} + \frac{d\ln\phi}{d\ln\mu} = 0. \tag{1.70}$$

We can define the anomalous dimension of the field as

$$\gamma_{\phi} \equiv \frac{1}{2} \frac{d \ln Z_{\phi}}{d \ln \mu} = \frac{1}{2} \frac{d \ln Z_{\phi}}{d \alpha} \frac{d \alpha}{d \ln \mu} = \frac{\alpha}{12} + O(\alpha^2). \tag{1.71}$$

Callan-Symanzik Equation

Consider the bare propagator:

$$\tilde{\Delta}_0(k) = Z_\phi \tilde{\Delta}(k) \tag{1.72}$$

The RG condition for the bare propagator gives:

$$\frac{d\ln\tilde{\Delta}_0(k)}{d\ln\mu} = \frac{d\ln Z_\phi}{d\ln\mu} + \frac{1}{\tilde{\Delta}(k)} \left(\frac{\partial}{\partial\ln\mu} + \frac{d\alpha}{d\ln\mu} \frac{\partial}{\partial\alpha} + \frac{dm}{d\ln\mu} \frac{\partial}{\partial m} \right) \tilde{\Delta}(k) = 0.$$

The Callan-Symanzik equation is

$$\left(2\gamma_{\phi} + \frac{\partial}{\partial \ln \mu} + \beta(\alpha)\frac{\partial}{\partial \alpha} + \gamma_{m}(\alpha)m\frac{\partial}{\partial m}\right)\tilde{\Delta}(k) = 0.$$
(1.73)

1.3 ϕ^4 Theory in $(d=4-\epsilon)$ Space-time

In this section, we consider the real Klein-Gordon field with ϕ^4 interaction in $(4 - \epsilon)$ -dimension space-time:

$$\mathcal{L} = Z_{\phi} \frac{1}{2} \partial^{\mu} \phi \partial_{\mu} \phi - Z_{m} \frac{m^{2}}{2} \phi^{2} - Z_{g} \frac{g\tilde{\mu}^{\epsilon}}{4!} \phi^{4}. \tag{1.74}$$

Note that the field ϕ has mass dimension $\left[\frac{d-2}{2}\right] = [1]$, so the original coupling constant g is dimensionless.

As the ϕ^3 theory, we can rewrite the Lagrangian as:

$$\mathcal{L} = \mathcal{L}_0 + \mathcal{L}_{int} + \mathcal{L}_{ct}. \tag{1.75}$$

In the following we investigate the loop correction to the mass and the coupling constant.

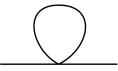
1.3.1 One-loop Correction

Self-energy

Following the same procedure, the one-loop self-energy correction is

$$i\Sigma(k^2) = -\frac{g\tilde{\mu}^{\epsilon}}{2} \int \frac{d^dq}{(2\pi)^d} \frac{1}{q^2 - m^2} + i(Ak^2 - Bm^2).$$
 (1.76)

The first term comes from the diagram



and the second term comes from the counter terms. After the Wick rotation,

$$\Sigma(k^2) = -\frac{g\tilde{\mu}^{\epsilon}}{2} \frac{\Omega_d}{(2\pi)^d} \int \frac{q^{d-1}dq}{q^2 + m^2} + (Ak^2 - Bm^2).$$
 (1.77)

The dimensional regulation is carried out using the following code:

```
 \begin{aligned} &\text{omg=}\left(2*\text{Pi}^{(d/2)}\right)/\left(\text{Gamma}\left[d/2\right]\right); \\ &\text{cof=}g*\left[\text{Mu}\right]^{(4-d)}/2*\text{omg}/\left(2*\text{Pi}\right)^{d}; \\ &\text{int=}\text{cof*}Integrate}\left[q^{(d-1)}/\left(q^{2}+\text{m}^{2}\right),\left\{q,0,Infinity\right\}\right]\left[\left[1\right]\right]; \\ &\text{map=}\left\{\text{EulerGamma-}>\text{Subscript}\left[\left[\text{Gamma}\right],E\right]\right\}; \\ &\text{ans=}\text{Series}\left[\text{int}/.\left\{d-\right\rangle-\left\{\frac{\text{Epsilon}}{2}\right\},\left\{\left(\frac{\text{Epsilon}}{2}\right\},0,0\right\}\right]; \\ &\text{ans}/.\text{map}//\text{Simplify} \end{aligned}
```

The result is

$$\Sigma(k^2) = \frac{gm^2}{32\pi^2} \left[\frac{2}{\epsilon} + 1 + \log\left(\frac{4\pi\tilde{\mu}^2 e^{-\gamma_E}}{m^2}\right) \right] + (Ak^2 - Bm^2) + O(\epsilon).$$
 (1.78)

Using the $\overline{\rm MS}$ renormalization scheme, we set

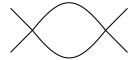
$$A = 0, \ B = \frac{g}{16\pi^2 \epsilon}. (1.79)$$

The result is

$$\Sigma(k^2) = \frac{gm^2}{16\pi^2} \log\left(\frac{\mu}{m}\right) + \frac{gm^2}{32\pi^2} + O(\epsilon). \tag{1.80}$$

Vertex Correction

Now consider the vertex correction. To the lowest order the diagram is



Together with the counter term, the vertex function is

$$iV_4^{(2)}(k_1, k_2, k_3, k_4) = \frac{g^2}{2} \left[iF(s) + iF(t) + iF(u) \right] - iCg, \tag{1.81}$$

where

$$s = (k_1 + k_2)^2, \ t = (k_1 + k_3)^2, \ u = (k_1 + k_4)^2,$$
 (1.82)

and

$$iF(k^2) = \tilde{\mu}^{\epsilon} \int \frac{d^d q}{(2\pi)^d} \tilde{\Delta}_0(q) \tilde{\Delta}_0(q+k)$$
(1.83)

$$= \frac{i\tilde{\mu}^{\epsilon}\Omega_d}{(2\pi)^d} \int_0^1 dx \int \frac{q^{d-1}dq}{\left[q^2 + m^2 + x(1-x)k^2\right]^2}.$$
 (1.84)

Then we carry out the calculation (set $D(k^2, x) = m^2 + x(1-x)k^2$)

```
omg=(2*Pi^(d/2))/(Gamma[d/2]);
cof=g^2*\[Mu]^(4-d)/2*omg/(2*Pi)^d;
int=cof*Integrate[q^(d-1)/(q^2+D)^2,{q,0,Infinity}][[1]];
map={EulerGamma->Subscript[\[Gamma],E]};
ans=Series[int/.{d->4-\[Epsilon]},{\[Epsilon],0,0}];
ans/.map//Simplify
```

The result is:

$$F(s) = \frac{1}{8\pi^{2}\epsilon} + \frac{1}{16\pi^{2}} \int_{0}^{1} dx \ln\left(\frac{4\pi\tilde{\mu}^{2}e^{-\gamma_{E}}}{D}\right)$$

$$= \frac{1}{8\pi^{2}\epsilon} + \frac{1}{8\pi^{2}} \ln\left(\frac{\mu}{m}\right) - \frac{1}{16\pi^{2}} \int_{0}^{1} dx \ln\left(\frac{D(s,x)}{m^{2}}\right).$$
(1.85)

The $\overline{\rm MS}$ scheme absorbs the $\frac{1}{8\pi^2\epsilon}$ term, i.e.,

$$C = \frac{3g}{16\pi^2}. (1.86)$$

The result is:

$$V_4(k_1, k_2, k_3, k_4) = -g + \frac{g^2}{32\pi^2} \int_0^1 dx \ln\left(\frac{\mu^6}{D(s, x)D(t, x)D(u, x)}\right). \tag{1.87}$$

To summarize, the normalization is:

$$Z_{\phi} = 1, \tag{1.88}$$

$$Z_m = 1 + \frac{g}{16\pi^2\epsilon},\tag{1.89}$$

$$Z_g = 1 + \frac{3g}{16\pi^2\epsilon}. (1.90)$$

1.3.2 Renormalization Group

Now consider the RG equation for the one-loop correction. The bare parameters are:

$$g_0 = Z_g g \tilde{\mu}^{\epsilon}, \ m_0 = Z_m^{1/2} m,$$
 (1.91)

The RG conditions are:

$$\frac{dg_0}{d\ln\mu} = \left(\frac{3}{16\pi^2\epsilon} + \frac{1}{g}\right)\frac{dg}{d\ln\mu} + \epsilon = 0, \tag{1.92}$$

$$\frac{dm_0}{d\ln\mu} = \frac{1}{32\pi^2\epsilon} \frac{dg}{d\ln\mu} + \frac{1}{m} \frac{dm}{d\ln\mu} = 0.$$
 (1.93)

Consider the series expansion of beta function:

$$\beta(g) = \frac{dg}{d \ln \mu} = \beta_1 g + \beta_2 g^2 + O(g^3). \tag{1.94}$$

The beta function is

$$\beta(g) = -\epsilon g + \frac{3g^2}{16\pi^2} + O(g^3). \tag{1.95}$$

The anomalous dimension of mass is

$$\gamma_m = \frac{1}{m} \frac{dm}{d \ln \mu} = \frac{g}{32\pi^2} + O(g^2)$$
 (1.96)

Chapter 2

Quantum Electrodynamics

The Lagrangian for quantum electrodynamics is

$$\mathcal{L}_{\text{QED}} = \bar{\psi} \left(i \gamma^{\mu} \partial_{\mu} - m \right) \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - e A_{\mu} \bar{\psi} \gamma^{\mu} \psi$$

$$= \mathcal{L}_{\text{Dirac}} + \mathcal{L}_{\text{Maxwell}} + \mathcal{L}_{\text{int}},$$
(2.1)

where

$$F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\nu} = (dA)_{\mu\nu}. \tag{2.2}$$

The Lagrangian is invariant under the gauge transformation:

$$\psi(x) \to e^{-ie\alpha(x)}\psi(x),$$

$$A^{\mu}(x) \to A^{\mu}(x) + \partial^{\mu}\alpha(x).$$
(2.3)

It is convenient to rewrite Lagrangian as

$$\mathcal{L}_{\text{QED}} = \bar{\psi} \left(i \mathcal{D} - m \right) \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}, \qquad (2.4)$$

where we have define the covariant derivative as:

$$\mathscr{D} = \gamma^{\mu} D_{\mu} = \gamma^{\mu} [\partial_{\mu} + ieA_{\mu}(x)] = \mathscr{D} + ie\mathscr{A}. \tag{2.5}$$

2.1 Free Field Theory

2.1.1 Dirac Field

The free Dirac field is

$$\mathcal{L}_{\text{Dirac}} = \bar{\psi} \left(i \gamma^{\mu} \partial_{\mu} - m \right) \psi \tag{2.6}$$

Consider the partition function with source

$$Z_{\text{Dirac}}[J] = \int D[\bar{\psi}, \psi] \exp \left[i \int d^d x \left(\mathcal{L}_{\text{Dirac}} + \bar{\eta}\psi + \bar{\psi}\eta \right) \right]$$

= $Z_{\text{Dirac}}[0] \exp \left[-i \int d^d x_1 d^d x_2 \bar{\eta}(x_1) \cdot D_F(x_1 - x_2) \cdot \eta(x_2) \right].$ (2.7)

where

$$D_F(x_1 - x_2) = \int \frac{d^d k}{(2\pi)^d} \frac{e^{-ik \cdot (x_1 - x_2)}}{\cancel{k} - m} = \int \frac{d^d k}{(2\pi)^d} e^{-ik \cdot (x_1 - x_2)} \frac{\cancel{k} + m}{k^2 - m^2}.$$
 (2.8)

Note that the propagator is

$$\langle 0|T\psi(x_1)\bar{\psi}(x_2)|0\rangle = \frac{1}{Z_{\text{Dirac}}[0]} \frac{\delta}{i\delta\bar{\eta}(x_1)} \frac{i\delta}{\delta(\eta(x_2))} Z_{\text{Dirac}}[J]$$

$$= iD_F(x_1 - x_2). \tag{2.9}$$

2.1.2 Electromagnetic Field

The Lagrangian for the classical electromagnetic field is

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + A_{\mu}J^{\mu}.$$
 (2.10)

Consider the partition function

$$\frac{Z_{\text{maxwell}}[J]}{Z_{\text{maxwell}}[0]} = \exp\left[-\frac{i}{2} \int dx_1 dx_2 J_{\mu}(x_1) \Pi^{\mu\nu}(x_1 - x_2) J_{\nu}(x_2)\right]. \tag{2.11}$$

The propagator is

$$\Pi^{\mu\nu}(x_1 - x_2) = \int \frac{d^dk}{(2\pi)^d} e^{-ik\cdot(x_1 - x_2)} \frac{-g^{\mu\nu} + (1 - \xi)k^{\mu}k^{\nu}}{k^2}.$$
 (2.12)

2.1.3 Perturbation Theory

As with the scalar field,

$$Z[\bar{\eta}, \eta, J] = \exp\left\{i \int d^d x \mathcal{L}_{\text{int}} \left[\frac{\delta}{i\delta J}, \frac{\delta}{i\delta \eta}, \frac{i\delta}{\delta \bar{\eta}}\right]\right\} Z_0[\bar{\eta}, \eta, J]. \tag{2.13}$$

We use the dimensional regularization by default. Note that ψ has the mass dimension $\left[\frac{d-1}{2}\right]$, A^{μ} had the mass dimension $\left[\frac{d}{2}-1\right]$, and e has the mass dimension $\left[2-\frac{d}{2}\right]$. When $d=4-\epsilon$, we replace e with $e\tilde{\mu}^{\epsilon/2}$, so that to make the coupling constant e dimensionless.

The renormalized Lagrangian is

$$\mathcal{L} = Z_{\psi}\bar{\psi}_{R}(i\gamma^{\mu}\partial_{\mu})\psi_{R} - Z_{m}m\bar{\psi}_{R}\psi_{R} + \frac{1}{4}Z_{A}F_{R,\mu\nu}F_{R}^{\mu\nu} - Z_{e}e_{R}A_{R}^{\mu}\bar{\psi}_{R}\gamma^{\mu}\psi_{R}$$

$$= \mathcal{L}_{0} + \mathcal{L}_{\text{int}} + \mathcal{L}_{\text{ct}}.$$
(2.14)

The we define the coefficients

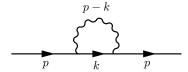
$$\delta_{\psi} = Z_{\phi} - 1, \ \delta_{m} = Z_{m} - 1, \ \delta_{Z} = Z_{A} - 1, \ \delta_{e} = Z_{e} - 1.$$
 (2.15)

The counter term also contribute to the perturbative expansion like the interactions.

¹The sign in the variational derivative comes from the anti-commutation relation of the fermionic fields.

Example 1. One-loop Correction to Electron Propagator

Consider the diagram



This contains 3 electron propagator, 1 photon propagator, and 2 vertices. The coefficient is (omit all the integration and summation for simplicity):

$$iD_F^{(2)}(p) \sim \frac{\delta^2}{\delta \bar{\eta} \delta \eta} \frac{1}{2!} \left(\frac{-ie\gamma_{\alpha\beta}^{\mu} \delta^3}{i\delta J^{\mu} \delta \eta_{\alpha} \delta \bar{\eta}_{\beta}} \right)^2 \frac{1}{3!} \left(-i\bar{\eta}_{\alpha} D_F^{\alpha\beta} \eta_{\beta} \right)^3 \left(-\frac{i}{2} J^{\mu} \Pi_{\mu\nu} J^{\nu} \right).$$

First consider the scalar coefficient. Since there is no additional symmetry, the abstract value is e^2 . There is an additional sign factor by the proper order of the fermion operators:

$$\frac{\delta^2}{\delta\bar{\eta}_f\delta\eta_i}\frac{\delta^2}{\delta\eta_1\delta\bar{\eta}_1}\frac{\delta^2}{\delta\eta_2\delta\bar{\eta}_2} = -\frac{\delta}{\delta\eta_i}\frac{\delta^2}{\delta\bar{\eta}_1\delta\eta_1}\frac{\delta^2}{\delta\bar{\eta}_2\delta\eta_2}\frac{\delta}{\delta\bar{\eta}_f}.$$

Then consider the tensor contraction,

$$\Pi_{\mu\nu}D_F^{\alpha\lambda}\gamma_{\lambda\rho}^{\mu}D_F^{\rho\tau}\gamma_{\tau\sigma}^{\nu}D_F^{\sigma\beta}.$$

The total amplitude is

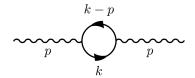
$$iD_F^{(2)}(p) = -e^2 \int \frac{d^4k}{(2\pi)^4} \Pi_{\mu\nu}(p-k) \left[D_F(p) \gamma^{\mu} D_F(k) \gamma^{\nu} D_F(p) \right]_{\alpha\beta}$$
$$= iD_F(p) i\Sigma(p^2) iD_F(p),$$

where $i\Sigma(p^2)$ is the self energy:

$$i\Sigma(p^2) = e^2 \int \frac{d^4k}{(2\pi)^4} \Pi_{\mu\nu}(p-k)\gamma^{\mu} D_F(k)\gamma^{\nu},$$
 (2.16)

Example 2. One-loop Correction to Photon Propagator

Consider the diagram



There is 2 electron propagator, 2 photon propagator, and 2 vertices. Consider the

18

perturbative expansion:

$$i\Pi^{(2)}(p) \sim \frac{\delta^2}{i\delta J i\delta J} \frac{1}{2!} \left(\frac{-e\gamma^{\mu}_{\alpha\beta}\delta^3}{\delta J^{\mu}\delta\eta_{\alpha}\delta\bar{\eta}_{\beta}} \right)^2 \frac{1}{2!} \left(-i\bar{\eta}_{\alpha}D_F^{\alpha\beta}\eta_{\beta} \right)^2 \frac{1}{2!} \left(\frac{i}{2} J^{\mu}\Pi_{\mu\nu}J^{\nu} \right)^2.$$

The diagram has no symmetry factor, but with a -1 sign, which is canceled out by the operator reordering:

$$\bar{\eta}_{\beta} D_F^{\beta \tau} \eta_{\tau} \bar{\eta}_{\sigma} D_F^{\sigma \alpha} \eta_{\alpha} = -\eta_{\alpha} \bar{\eta}_{\beta} D_F^{\beta \tau} \eta_{\tau} \bar{\eta}_{\sigma} D_F^{\sigma \alpha}. \tag{2.17}$$

The overall value is e^2 .

Then consider the tensor contraction,

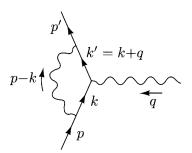
$$-i\Pi^{\mu\nu}_{(2)} \sim e^2 \Pi_{\mu\rho} \gamma^{\rho}_{\alpha\beta} D_F^{\beta\tau} \gamma^{\eta}_{\tau\sigma} D_F^{\sigma\alpha} \Pi_{\eta\nu} \sim i\Pi_{\mu\rho} i\Sigma^{\rho\sigma} i\Pi_{\sigma\nu}.$$

The photon self-energy is

$$i\Sigma^{\mu\nu}(p^2) = -e^2 \int \frac{d^4k}{(2\pi)^4} \text{Tr} \left[\gamma^{\mu} D_F(k-p) \gamma^{\nu} D_F(k) \right].$$
 (2.18)

Example 3. One-loop Correction to Vertex

Consider the diagram



There is 4 electron propagator, 2 photon propagator, and 3 vertices. Consider the perturbative expansion:

$$\frac{\delta^3}{i\delta J\delta\bar{\eta}\delta\eta}\frac{1}{2!}\left(\frac{-e\gamma^{\mu}_{\alpha\beta}\delta^3}{\delta J^{\mu}\delta\eta_{\alpha}\delta\bar{\eta}_{\beta}}\right)^3\frac{1}{2!}\left(-i\bar{\eta}_{\alpha}D_F^{\alpha\beta}\eta_{\beta}\right)^4\frac{1}{2!}\left(-\frac{i}{2}J^{\mu}\Pi_{\mu\nu}J^{\nu}\right)^2.$$

There is not symmetry factor, and an additional -i factor. The total coefficient is $-ie^3$.

Then consider the tensor contraction

$$D_F^{\alpha\gamma} \gamma_{\gamma\rho}^{\nu} D_F^{\rho\sigma} \gamma_{\sigma\tau}^{\zeta} D_F^{\tau\eta} \gamma_{\eta\varepsilon}^{\lambda} D_F^{\xi\beta} \Pi_{\nu\lambda} \Pi_{\mu\zeta}.$$

The vertex correction is:

$$iV_3(q, p, p') = [iD_F(p)][iD_F(p')][i\Pi^{\mu\nu}(q)][-ie\Gamma^{\nu}(q, p, p')]$$

where

$$i\Gamma^{\mu}(q,p,p') = -e^2 \int \frac{d^4k}{(2\pi)^4} \Pi_{\nu\lambda}(p-k) \gamma^{\nu} D_F(k') \gamma^{\mu} D_F(k) \gamma^{\lambda}.$$
 (2.19)

Example 4. Counter Terms

Consider the counter term in the diagram



The perturbative expansion is

$$iD_F^{(\mathrm{ct})} \sim \frac{\delta^2}{\delta \bar{\eta} \delta \eta} i (\delta_\psi \gamma_{\alpha\beta}^\mu k_\mu - \delta_m \mathbb{I}_{\alpha\beta}) \frac{\delta^2}{\delta \eta_\alpha \delta \bar{\eta}_\beta} \frac{1}{2!} \left(-i \bar{\eta}_\alpha D_F^{\alpha\beta} \eta_\beta \right)^2.$$

The contribution to the electron self energy is

$$\delta_{\psi} \cancel{k} - \delta_m m_R$$
.

Similarly, the digram



contribute to the photon self energy with term

$$\delta_A[-p^2g^{\mu\nu} + (1-\xi)p^{\mu}p^{\nu}],$$

and the diagram



contribute to the vertex with term

$$\delta_e \gamma^\mu$$
.

2.2 Loop Correction

In this section, we consider the QED in $(d=4-\epsilon)$ dimensional space-time.

2.2.1 Electron Propagator

Consider the one-loop correction to the electron propagator, where the self energy (2.16) is

$$i\Sigma(p^2) = e^2 \tilde{\mu}^{\epsilon} \int \frac{d^d k}{(2\pi)^d} \Pi_{\mu\nu}(p-k) \left[\gamma^{\mu} D_F(k) \gamma^{\nu} \right]_{\alpha\beta}$$
$$= -e^2 \tilde{\mu}^{\epsilon} \int \frac{d^d k}{(2\pi)^d} \frac{\gamma^{\mu} (\cancel{k} + m) \gamma_{\mu}}{(p-k)^2 (k^2 - m^2)}.$$
 (2.20)

The nominator can be simplified using the FeynCalc Package:

```
(*load FeynCalc Package*)
<< FeynCalc`

(*simplify the gamma expression*)
Contract[GA[\[Mu]].(GS[k]+m).GA[\[Mu]]]//DiracSimplify</pre>
```

The result is

$$4m-2k$$
.

The denominator can be simplify using the Feynman parameter:

$$\frac{1}{(p-k)^2(k^2-m^2)} = \int_0^1 \frac{dx}{\left[(k-b)^2-D\right]^2}$$

where b and D can be calculated by

```
A1=(k-p)^2;

A2=k^2-m^2;

{c,b,a}=CoefficientList[x*A1+(1-x)*A2, {k}];

-b/2

-c+b^2/(4*a)//Simplify
```

The result is

$$b = px$$
, $D = (1 - x)(m^2 - p^2x)$.

Shift $k \to k + px$, the self energy becomes:

$$i\Sigma(p^{2}) = 2e^{2}\tilde{\mu}^{\epsilon} \int_{0}^{1} (x\not p - 2m)dx \int \frac{d^{d}k}{(2\pi)^{d}} \frac{1}{(k^{2} - D)^{2}}$$

$$= i\frac{2e^{2}\tilde{\mu}^{\epsilon}\Omega_{d}}{(2\pi)^{d}} \int_{0}^{1} (x\not p - 2m)dx \int \frac{k^{d-1}dk}{(k^{2} + D)^{2}}.$$
(2.21)

The regularization procedure

```
omg=(2*Pi^(d/2))/(Gamma[d/2]);
cof=2*e^2*\[Mu]^(4-d)*omg/(2*Pi)^d;
int=cof*Integrate[q^(d-1)/(q^2+D)^2,{q,0,Infinity}][[1]];
map={e->Sqrt[4*Pi*\[Alpha]],EulerGamma->Subscript[\[Gamma],E]};
ans=Series[int/.{d->4-\[Epsilon]},{\[Epsilon],0,0}];
ans/.map//Simplify
```

The result is $(\mu^2 = 4\pi \tilde{\mu}^2 e^{-\gamma_E})$

$$\Sigma(p^2) = \frac{e_R^2}{8\pi^2} \int_0^1 dx (x \not p - 2m_R) \left[\frac{2}{\epsilon} + \ln\left(\frac{\mu^2}{D}\right) \right]. \tag{2.22}$$

The infinity comes from

$$\frac{e_R^2}{4\pi^2\epsilon}\int_0^1 dx (x\not\!p-2m_R) = \frac{e_R^2}{8\pi^2\epsilon}\not\!p - \frac{e_R^2}{2\pi^2\epsilon}m_R.$$

Using the $\overline{\rm MS}$ subtraction scheme, we choose

$$\delta_{\psi} = -\frac{e_R^2}{8\pi^2 \epsilon}, \ \delta_m = -\frac{e_R^2}{2\pi^2 \epsilon},$$
 (2.23)

and the self energy is

$$\Sigma(p^{2}) = \frac{e_{R}^{2}}{8\pi^{2}} \int_{0}^{1} dx (x \not p - 2m_{R}) \ln\left[\frac{\mu^{2}}{(1-x)(m_{R}^{2} - p^{2}x)}\right]$$

$$= \frac{e_{R}^{2}}{8\pi^{2}} (\not p - 4m_{R}) \ln\left(\frac{\mu}{m_{R}}\right) - \int_{0}^{1} dx \ln\left[(1-x)\left(1 - \frac{p^{2}x}{m_{R}^{2}}\right)\right].$$
(2.24)

2.2.2 Photon Self-energy

Consider the one-loop correction to the photon propagator, where the self energy (2.18) is

$$i\Sigma^{\mu\nu} = -e^2 \tilde{\mu}^{\epsilon} \int \frac{d^d k}{(2\pi)^d} \frac{\text{Tr}\left[\gamma^{\mu} D_F(k-p) \gamma^{\nu} D_F(k)\right]}{(k^2 - m^2)[(p-k)^2 - m^2]}.$$
 (2.25)

The Dirac trace and Feynman parameter is calculated by

```
(*Dirac trace*)
DiracTrace[GA[\[Mu]].(GS[k-p]+m).GA[\[Nu]].(GS[k]+m)]//DiracSimplify

(*Feynman paramater*)
A1=k^2-m^2;
A2=(k-p)^2-m^2;
{c,b,a}=CoefficientList[x*A1+(1-x)*A2,{k}];
-b/2
-c+b^2/(4*a)//Simplify
```

The nominator is

$$4\left[g^{\mu\nu}\left(k\cdot p - k^2 + m^2\right) + 2k^{\mu}k^{\nu} - k^{\mu}p^{\nu} - p^{\mu}k^{\nu}\right]$$

The denominator is:

$$\frac{1}{(k^2 - m^2)[(p-k)^2 - m^2]} = \frac{1}{\{[k - p(1-x)]^2 - [m^2 + p^2x(x-1)]\}^2}$$

Let $D = m^2 - p^2 x (1 - x)$, shift $k \to k + p (1 - x)$, and drop all p^{μ} linear term,² the result is

$$i\Sigma^{\mu\nu} = -4e^2\tilde{\mu}^{\epsilon} \int_0^1 dx \int \frac{d^dk}{(2\pi)^d} \frac{2k^{\mu}k^{\nu} - g^{\mu\nu} \left[k^2 - x(1-x)p^2 - m^2\right]}{\left[k^2 - D\right]^2}$$
(2.26)

The self-energy $i\Sigma^{\mu} \propto g^{\mu\nu}$, we can make the substitution

$$k^{\mu}k^{\nu} \to \frac{1}{d}k^2g^{\mu\nu}.$$

The Ward identity requires that the p^{μ} term in the propagator do not contribute to any scattering process.

We then need to consider the integral

$$iI(x) = 4e^{2}\tilde{\mu}^{\epsilon} \int \frac{d^{d}k}{(2\pi)^{d}} \frac{(1 - \frac{2}{d})k^{2} - x(1 - x)p^{2} - m^{2}}{[k^{2} - D]^{2}},$$

$$I(x) = -\frac{4e^{2}\tilde{\mu}^{\epsilon}\Omega_{d}}{(2\pi)^{d}} \int k^{d-1}dk \frac{(1 - \frac{2}{d})k^{2} + x(1 - x)p^{2} + m^{2}}{[k^{2} + D]^{2}}.$$

The regulation is carried out by the following code:

```
omg=(2*Pi^(d/2))/(Gamma[d/2]);
cof=-4*e^2*\[Mu]^(4-d)*omg/(2*Pi)^d;
den=q^(d-1)*((1-2/d)q^2+x*(1-x)p^2+m^2);
int=cof*Integrate[den/(q^2+D)^2,{q,0,Infinity}][[1]];
map={EulerGamma->Subscript[\[Gamma],E],D->m^2-p^2*x*(1-x)};
ans=Series[int/.{d->4-\[Epsilon]},{\[Epsilon],0,0}];
ans/.map//Simplify
```

The result is

$$\Sigma^{\mu\nu}(p^2) = -\frac{e_R^2 p^2 g^{\mu\nu}}{2\pi^2} \int_0^1 dx \ x(1-x) \left[\frac{2}{\epsilon} + \ln\left(\frac{\mu^2}{m_R^2 - p^2 x(1-x)}\right) \right]$$
 (2.27)

The divergent part is

$$-\frac{e_R^2 p^2 g^{\mu\nu}}{\pi^2 \epsilon} \int_0^1 dx \ x(1-x) = -\frac{e_R^2 p^2 g^{\mu\nu}}{6\pi^2 \epsilon}.$$

The counter term coefficient is

$$\delta_A = -\frac{e_R^2}{6\pi^2\epsilon}. (2.28)$$

The photon self-energy is then

$$\Sigma^{\mu\nu}(p^2) = -\frac{e_R^2 p^2 g^{\mu\nu}}{2\pi^2} \int_0^1 dx \ x(1-x) \ln\left[\frac{\mu^2}{m_R^2 - p^2 x(1-x)}\right]$$

$$= -\frac{e_R^2 p^2 g^{\mu\nu}}{12\pi^2} \ln\left(\frac{\mu}{m}\right) + \frac{e_R^2 p^2 g^{\mu\nu}}{2\pi^2} \int_0^1 dx \ x(1-x) \ln\left[1 - \frac{p^2}{m_R^2} x(1-x)\right]. \tag{2.29}$$

2.2.3 Vertex Correction

Consider the loop correction (2.19):

$$i\Gamma^{\mu}(p, q_1, q_2) = e^2 \tilde{\mu}^{\epsilon} \int \frac{d^4k}{(2\pi)^4} \frac{\gamma^{\nu}(\cancel{k}' + m)\gamma^{\mu}(\cancel{k} + m)\gamma_{\nu}}{(k^2 - m^2)(k'^2 - m^2)(p - k)^2}.$$
 (2.30)

Using the following code

```
(*numerator*)
den=Contract[GA[\[Nu]].(GS[kp]+m).GA[\[Mu]].(GS[k]+m).GA[\[Nu]]];
DiracSimplify[den]
(*Feynman parameter*)
```

```
A1=k^2-m^2;

A2=(k+q)^2-m^2;

A3=(p-k)^2;

{c,b,a}=CoefficientList[x*A1+y*A2+z*A3,{k}];

-b/2//Simplify

-c+b^2/4//Simplify
```

The numerator is

$$-2k\gamma^{\mu}k' - 2m^2\gamma^{\mu} + 4m(k+k')^{\mu}$$

The denominator is

$$\int \frac{dF_3}{[(k+yq-zp)^2-D]^3},$$

where

$$D = (x+y)m^2 - z(1-z)p^2 - y(1-y)q^2 - 2yzpq$$

= $(x+y)m^2 - xyq^2 - yzp'^2 - xzp^2$.

Shift $k^{\mu} \to k^{\mu} + zq_1^{\mu} - yp^{\mu}$, throw away all terms with linear k^{μ} , and replace $k^{\mu}k^{\nu}$ with $\frac{1}{d}k^2g^{\mu\nu}$, the result is

$$\frac{4}{d}k^2\gamma^{\mu} - 2(-y\not q + z\not p)\gamma^{\mu}[(1-y)\not q + z\not p] + 4m^2\gamma^{\mu} - 2m\left[(1-2y)q^{\mu} + 2zp^{\mu}\right].$$

Note that only the quadratic term is divergent.

$$\Gamma^{\mu}(p, q_1, q_2) = -i \frac{4e^2 \tilde{\mu}^{\epsilon} \gamma^{\mu}}{d} \int dF_3 \int \frac{d^d k}{(2\pi)^d} \frac{k^2}{(k^2 - D)^3} + \delta \Gamma^{\mu}(p, q_1, q_2).$$

where $\delta\Gamma^{\mu}$ stores all the finite part

$$\begin{split} &\delta \Gamma^{\mu}(p,q_{1},q_{2}) \\ &= \int \frac{e^{2}k^{3}dkdF_{3}}{(2\pi)^{2}(k^{2}+D)^{3}} \left\{ (-y\not q + z\not p)\gamma^{\mu}[(1-y)\not q + z\not p] - 2m^{2}\gamma^{\mu} + m\left[(1-2y)q^{\mu} + 2zp^{\mu}\right] \right\}. \end{split}$$

The divergent part is

$$\frac{4e^2\tilde{\mu}^{\epsilon}\Omega_d\gamma^{\mu}}{d(2\pi)^d}\int dF_3\int\frac{k^{d+1}dk}{(k^2+D)^3}=\frac{e_R^2}{16\pi^2}\gamma^{\mu}\int dF_3\left[\frac{2}{\epsilon}+\ln\left(\frac{\mu^2}{D}\right)\right].$$

Using the $\overline{\rm MS}$ scheme, the coefficient of the counter term is

$$\delta_e = -\frac{e_R^2}{8\pi^2\epsilon}. (2.31)$$

2.2.4 Renormalization Group

In summery, the renormalization factors are

$$Z_{\psi} = 1 - \frac{e_R^2}{8\pi^2 \epsilon} + O(e_R^3),$$

$$Z_A = 1 - \frac{e_R^2}{6\pi^2 \epsilon} + O(e_R^3),$$

$$Z_m = 1 - \frac{e_R^2}{2\pi^2 \epsilon} + O(e_R^3),$$

$$Z_e = 1 - \frac{e_R^2}{8\pi^2 \epsilon} + O(e_R^3),$$
(2.32)

which means

$$\frac{d \ln Z_{\phi}}{de_{R}} = -\frac{e_{R}}{4\pi^{2}\epsilon} + O(e_{R}^{2}),$$

$$\frac{d \ln Z_{A}}{de_{R}} = -\frac{e_{R}}{3\pi^{2}\epsilon} + O(e_{R}^{2}),$$

$$\frac{d \ln Z_{m}}{de_{R}} = -\frac{e_{R}}{\pi^{2}\epsilon} + O(e_{R}^{2}),$$

$$\frac{d \ln Z_{e}}{de_{R}} = -\frac{e_{R}}{4\pi^{2}\epsilon} + O(e_{R}^{2}).$$
(2.33)

The bare parameters are

$$\psi_{0} = Z_{\psi}^{1/2} \psi_{R},$$

$$A_{0} = Z_{A}^{1/2} A_{R},$$

$$m_{0} = Z_{m} Z_{\psi}^{-1} m_{R},$$

$$e_{0} = Z_{e} Z_{\psi}^{-1} Z_{A}^{-1/2} e_{R} \tilde{\mu}^{\epsilon/2}.$$
(2.34)

The RG equation for e_0 is

$$\frac{d \ln e_0}{d \ln \mu} = \left(\frac{\ln Z_e}{de_R} - \frac{\ln Z_\psi}{de_R} - \frac{1}{2} \frac{\ln Z_A}{de_R} + \frac{1}{e_R}\right) \frac{de_R}{d \ln \mu} + \frac{\epsilon}{2} = 0.$$
 (2.35)

The beta function is

$$\beta(e_R) = \frac{de_R}{d \ln \mu} = -\frac{\epsilon}{2} e_R + \frac{e_R^3}{12\pi^2} + O(e_R^4). \tag{2.36}$$

The RG equation for m_0 is

$$\frac{d \ln m_0}{d \ln \mu} = \left(\frac{\ln Z_m}{de_R} - \frac{\ln Z_\psi}{de_R}\right) \frac{de_R}{d \ln \mu} + \frac{1}{m_R} \frac{dm_R}{d \ln \mu} = 0.$$
 (2.37)

The anomalous dimension of mass is

$$\gamma_m = \frac{d \ln m_R}{d \ln \mu} = -\frac{3e_R^2}{8\pi^2} + O(e_R^3). \tag{2.38}$$

Chapter 3

Non-relativistic QFT

A general non-relativistic field has the Lagrangian¹

$$\mathcal{L} = \bar{\psi}_a(x)(i\delta_{ab}\partial_t - \hat{H}_{ab})\psi_b(x) + \mathcal{V}_{int}$$
(3.1)

where the field operator ψ can be bosonic or fermionic, which is denoted by a number $\zeta = \pm 1$, and \mathcal{V}_{int} is the interaction Lagrangian. A general interaction has the form

$$\mathcal{V}_{\text{int}} = \bar{\psi}_a(x_1)\bar{\psi}_b(x_2)V_{ab}(x_1, x_2)\psi_b(x_2)\psi_a(x_1). \tag{3.2}$$

Note that the classical equation of motion for the free field is

$$0 = \partial_{\mu} \frac{\partial \mathcal{L}}{\partial(\partial_{\mu} \bar{\psi}_{a}(x))} - \frac{\partial \mathcal{L}}{\partial \bar{\psi}_{a}(x)}$$

$$= -i\partial_{t} \phi_{a}(x) + \hat{H}_{ab} \phi_{b}(x),$$
(3.3)

which satisfies the Schrödinger equation.

We are mostly work with finite system size L^d with UV cutoff $\Lambda = \frac{2\pi}{a}$, in which case the spatial Fourier transformation is

$$\tilde{\psi}_a(k) = \int_{I^d} d^d x e^{-ik \cdot x} \psi(k), \qquad (3.4)$$

$$\psi_a(x) = \frac{1}{L^d} \sum_k e^{ik \cdot x} \tilde{\psi}(k). \tag{3.5}$$

Note that for finite size, the momentum is discretized:

$$k_i = \frac{2\pi}{L} n_i, \ n_i = -N, \cdots, N.$$
 (3.6)

By default, we take the thermodynamic limit. The summation is approximated by the integration:

$$\frac{1}{L^d} \sum_{k} \longrightarrow \int_{|k| < \Lambda} \frac{d^d k}{(2\pi)^d}.$$
 (3.7)

¹The repeated indices are automatically summed.

²We can regard a as the lattice spacing, and assume L = Na.

3.1 Finite Temperature Field Theory

The original real-time partition function is defined as³

$$Z[J] = \int D[\bar{\psi}, \psi] \exp\left\{i \int dt \int d^dx \left[\mathcal{L} + \bar{J}_a(x)\psi_a(x) + \bar{\psi}_a(x)J_a(x)\right]\right\}. \tag{3.8}$$

For finite-temperature field theory, after making the wick rotation $t \to -i\tau$, the partition function for a generic non-relativistic lattice theory is:

$$Z[J] = \int D[\bar{\psi}, \psi] e^{-S[\bar{\psi}, \psi] + \bar{J} \cdot \psi + \bar{\psi} \cdot J}, \qquad (3.9)$$

where the action is

$$S = \int_0^\beta d\tau \int d^d x \left[\bar{\psi}_a(x) (\delta_{ab}\partial_\tau + \hat{H}_{ab}) \psi_b(x) + \mathcal{V}_{\text{int}} \right]. \tag{3.10}$$

Remark 6. Temporal Fourier Transformation

The Fourier transformation on the imaginary time domain is defined as:

$$\tilde{\psi}(\omega_n) = \int_0^\beta d\tau e^{i\omega_n \tau} \psi(\tau), \qquad (3.11)$$

$$\psi(\tau) = \frac{1}{\beta} \sum_{\omega_n} e^{-i\omega_n \tau} \tilde{\psi}(\omega_n). \tag{3.12}$$

Under such convention, in the thermodynamic limit and zero-temperature limit, the spatial-temporal Fourier transformation agrees with the relativistic case (up to a Wick rotation).

3.1.1 Free Field Theory

We first consider the action of free field

$$S_0 = \int_0^\beta d\tau \int d^d x \ \bar{\psi}_a(x) (\delta_{ab}\partial_\tau + \hat{H}_{ab}) \psi_b(x). \tag{3.13}$$

The Fourier transformation

$$S_0 = \frac{1}{\beta} \sum_{\omega_n} \int_{\Lambda} \frac{d^d k}{(2\pi)^d} \tilde{\bar{\psi}}_a(k, \omega_n) \left[-i\omega_n + \tilde{H}_{ab}(k) \right] \tilde{\psi}_b(k, \omega_n). \tag{3.14}$$

The partition function with source is

$$\frac{Z_0[J]}{Z_0[0]} = \exp\left[-\frac{1}{\beta} \sum_{\omega_n} \int_{\Lambda} \frac{d^d k}{(2\pi)^d} \tilde{J}_a(k,\omega_n) \tilde{G}_{ab}(k,\omega_n) \tilde{J}_b(k,\omega_n)\right], \tag{3.15}$$

³As with the relativistic case, we introduce an auxiliary source J, which is bosonic/fermionic if the field ψ is bosonic/fermionic.

where the Green's function is

$$\tilde{G}_{ab}(k,\omega_n) = \left[\frac{1}{i\omega_n - \tilde{H}(k)}\right]_{ab}.$$
(3.16)

Remark 7. Obtaining the Partition Function

Unlike the relativistic case, the value of the value of partition function without source $Z_0[0]$ is related to the free energy. We can express it formally as

$$Z_0[0] = \left[\det(-G_{ab})^{-1} \right]^{-\zeta}.$$

To get the correct dimensionality, we set the determinant as

$$Z_0[0] \equiv \prod_{k,\omega_n} \left\{ \beta \det \left[-i\omega_n + \tilde{H}(k) \right] \right\}^{-\zeta}.$$

Thus the free energy is

$$F = -\frac{1}{\beta} \ln Z_0 = \zeta \sum_{k,\omega_n} \ln \left\{ \beta \det \left[-i\omega_n + \tilde{H}(k) \right] \right\}.$$
 (3.17)

Remark 8. Matsubara Summation

Now consider the summation on Matsubara frequency:

$$\sum_{\omega_n} f(\omega_n) = \begin{cases} \sum_n f(\frac{2n\pi}{\beta}) & \text{bosonic} \\ \sum_n f(\frac{(2n+1)\pi}{\beta}) & \text{fermionic} \end{cases}$$
 (3.18)

The frequency is capture by the singularities of the density function of the states:

$$\rho(z) = \begin{cases} \frac{1}{\exp(\beta z) - 1} & \text{bosonic} \\ \frac{1}{\exp(\beta z) + 1} & \text{fermionic} \end{cases}$$
 (3.19)

The residue on imaginary frequency $i\omega_n$ is alway $\frac{1}{\beta}$. In this way, the summation is:

$$\frac{1}{\beta} \sum_{\omega_n} f(i\omega_n) = \frac{1}{2\pi i} \oint \rho(z) f(z) = -\sum_k \text{Res } \rho(z) f(z)|_{z=z_k}.$$
 (3.20)

Example 5. Summation of Green's function

Consider the frequency summation for the correlation function:

$$\frac{1}{\beta} \sum_{\omega_p} \tilde{G}_0(k) = \frac{1}{\beta} \sum_{\omega_p} \frac{1}{i\omega_n - E_p} = -\text{Res } \frac{\rho(z)}{z - E_p} \Big|_{z = E_p} = \rho(E_p).$$

Example 6. Summation of Green's function

Consider the frequency summation for the correlation function:

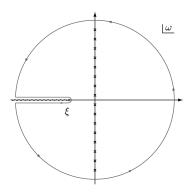
$$\sum_{\omega_n} \langle \bar{\psi}_{\vec{p},\omega_n} \psi_{\vec{p},\omega_n} \rangle = \frac{1}{\beta} \sum_{\omega_n} \frac{1}{-i\omega_n + \epsilon_{\vec{p}}} = \operatorname{Res} \left. \frac{\rho(z)}{z - \epsilon_{\vec{p}}} \right|_{z - \epsilon_{\vec{p}}} = \rho(\epsilon_{\vec{p}}).$$

Example 7. Free Energy Summation

Consider the free energy

$$F = -\frac{1}{\beta} \ln Z = -\frac{1}{\beta} \sum_{\alpha \mid r} \ln[\beta(-i\omega_n + E_{\vec{p}})] = \frac{1}{2\pi i} \oint dz \rho(z) \ln[\beta(\xi - z)].$$

To calculate the summation, we consider the line integral along the loop:



The free energy is

$$F = \frac{1}{2\pi i} \int_{-\infty}^{\infty} dx \rho(x) \ln\left(\frac{\xi - x - i\epsilon}{\xi - x + i\epsilon}\right)$$
$$= \frac{-\zeta}{2\pi i\beta} \int_{-\infty}^{\infty} dx \ln(1 - \zeta e^{-\beta z}) \left(\frac{1}{x + i\epsilon - \xi} - \frac{1}{x - i\epsilon - \xi}\right),$$

where we integrate the expression by part, noticing that

$$\frac{d}{dz}\frac{\zeta}{\beta}\ln(1-\zeta e^{-\beta z}) = \frac{1}{e^{\beta z}-\zeta} = \rho(z)$$
(3.21)

Using the identity

$$\lim_{\epsilon \to 0^+} \frac{1}{x + i\epsilon} = -i\pi\delta(x) + \mathcal{P}\frac{1}{x},$$

the above expression can be simplified to

$$F = \frac{\zeta}{\beta} \ln(1 - \zeta e^{-\beta\zeta}). \tag{3.22}$$