

Quantum Electrodynamics

Jie Ren

Contents

I. Introduction	1
A. Representations of Lorentz group	1
II. Scattering	3
A. $e^+e^- \rightarrow \mu^+\mu^-$	3
B. $e^-p^+ \rightarrow e^-p^+$	4
C. $\gamma e^- \rightarrow \gamma e^-$	5
III. Perturbative Renormalization	5
A. Vacuum Polarization	7
1. Regularization and Renormalization	8
2. Physical Observable	8
B. One-loop Correction to Electron Propagator	9
1. Regularization and Renormalization	10
2. Physical Observables	10
C. One-loop Correction to Vertex	12
IV. Systematic Renormalization	13
A. Renormalization Group	13

I. INTRODUCTION

Quantum electrodynamics (QED) is the field theory for the interaction of charged Dirac field with the U(1) gauge field. The Lagrangian is obtained by minimal coupling:

$$\mathcal{L}_{\text{QED}} = \bar{\psi} (i\not{D} - m) \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}, \quad \text{where} \quad \not{D} = \gamma^\mu D_\mu = \gamma^\mu [\partial_\mu - iqA_\mu(x)] = \not{\partial} - iq\not{A}. \quad (1)$$

The Lagrangian is invariant under the gauge transformation: $\psi(x) \rightarrow e^{iq\alpha(x)}\psi(x)$, $A^\mu(x) \rightarrow A^\mu(x) + \partial^\mu\alpha(x)$, where q is the charge of the field. For electron, $q = -e$, and for quarks, $q = \frac{2}{3}e$ or $q = -\frac{1}{3}e$.

A. Representations of Lorentz group

For the (3 + 1)D spacetime, the Lorentz group can be represented as

$$\Lambda(\boldsymbol{\theta}, \boldsymbol{\beta}) = \exp(i\theta_i J_i + i\beta_i K_i), \quad \theta_i \equiv \frac{1}{2}\varepsilon_{ijk}\omega_{jk}, \quad \beta_i \equiv \omega_{i0}, \quad (2)$$

where the Lie algebra contains 3 rotational generators $J_i \equiv \frac{1}{2}\varepsilon_{ijk}M^{jk}$ and 3 boost generators $K_i \equiv M^{i0}$. In the fundamental representation, the generators are represented by

$$\begin{aligned} J_1 &= \begin{bmatrix} 0 & & & \\ & 0 & & \\ & & 0 & -i \\ & & i & 0 \end{bmatrix}, & J_2 &= \begin{bmatrix} 0 & & & \\ & 0 & i & \\ & & 0 & \\ -i & & & 0 \end{bmatrix}, & J_3 &= \begin{bmatrix} 0 & & & \\ & 0 & -i & \\ & i & 0 & \\ & & & 0 \end{bmatrix}, \\ K_1 &= \begin{bmatrix} 0 & -i & & \\ -i & 0 & & \\ & & 0 & \\ & & & 0 \end{bmatrix}, & K_2 &= \begin{bmatrix} 0 & -i & & \\ & 0 & & \\ -i & & 0 & \\ & & & 0 \end{bmatrix}, & K_3 &= \begin{bmatrix} 0 & & -i & \\ & 0 & & \\ & & 0 & \\ -i & & & 0 \end{bmatrix}. \end{aligned}$$

The Lie algebra of the Lorentz algebra can be explicitly done using the fundamental representation. The result is

$$[J_i, J_j] = i\varepsilon_{ijk}J_k, \quad [J_i, K_j] = i\varepsilon_{ijk}K_k, \quad [K_i, K_j] = -i\varepsilon_{ijk}J_k. \quad (3)$$

A special combination of those generators $N_i^{L/R} \equiv \frac{1}{2}(J_i \mp iK_i)$ will create two independent algebras:

$$[N_i^L, N_j^L] = i\varepsilon_{ijk}N_k^L, \quad [N_i^R, N_j^R] = i\varepsilon_{ijk}N_k^R, \quad [N_i^L, N_j^R] = 0. \quad (4)$$

In the following, we show such generators give all irreducible representations of the Lorentz group $SO(3, 1)$, which are the building blocks of the relativistic field theory.

We see from (4) that after the recombination, the Lorentz algebra breaks into two independent $\mathfrak{su}(2)$ algebra. Mathematically it means $\mathfrak{so}(3, 1) \simeq \mathfrak{su}(2) \oplus \mathfrak{su}(2)$. That is,

$$U_{j_L, j_R}(\Lambda) = \exp [i(\boldsymbol{\theta} + i\boldsymbol{\beta}) \cdot \mathbf{N}_{j_L}^L + i(\boldsymbol{\theta} - i\boldsymbol{\beta}) \cdot \mathbf{N}_{j_R}^R], \quad (5)$$

where we see that the representation of the Lorentz algebra can be labeled by j_L and j_R . In relativistic QFT, the Lorentz symmetry restricts the possible terms that can appear in the Lagrangian. Different free fields correspond to different representations of the Lorentz algebra, and the Lagrangian should be singlet under the Lorentz transformations.

Trivial representation. The $(j_L, j_R) = (0, 0)$ representation corresponds to the scalar field denoted as $\phi(x)$. Since the field itself is singlet, any polynomial of the field in principle can appear in the theory. When considering the free theory, we restrict our attention to the quadratic terms, therefore the allowed free theory can only be

$$\mathcal{L}_{\text{KG}} = \frac{1}{2}\partial^\mu\phi\partial_\mu\phi - \frac{m^2}{2}\phi^2 \simeq -\frac{1}{2}\phi(\partial^2 + m^2)\phi.$$

Besides the field Lagrangian \mathcal{L}_{KG} , there are more general Lorentz-invariant terms that can be added to the Lagrangian, which describe the interaction of the theory.

Vector representation. If we choose $(j_L = j_R = 1/2)$, the field is transformed as a Lorentz vector. We denote the field as $A^\mu(x)$. Some possible quadratic forms for the vector field that forms singlets are $A^\mu A_\mu$, $(\partial_\mu A^\mu)^2$, $A^\nu \partial^2 A_\nu$, and $\varepsilon_{\mu\nu\rho\lambda}\partial^\mu A^\nu \partial^\rho A^\lambda$. For the field theory describing the electromagnetic field, we require the theory to further have gauge symmetry, i.e., invariant under $A^\mu(x) \rightarrow A^\mu(x) + \partial^\mu \alpha(x)$. The gauge invariant forbids the first term and forces the second and third terms to combine as

$$(\partial_\mu A^\mu)^2 - A^\nu \partial^2 A_\nu \sim \frac{1}{2}(\partial^\mu A^\nu - \partial^\nu A^\mu)(\partial_\mu A^\nu - \partial_\nu A_\mu) \equiv \frac{1}{2}F^{\mu\nu}F_{\mu\nu}, \quad \text{where} \quad F^{\mu\nu} \equiv (\partial^\mu A^\nu - \partial^\nu A^\mu).$$

The Lagrangian describing the electromagnetic field is given by

$$\mathcal{L}_{\text{Maxwell}} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu}. \quad (6)$$

Note that the fourth term is called the *theta term*, which can be written as a boundary term: $\varepsilon_{\mu\nu\rho\lambda}\partial^\mu A^\nu \partial^\rho A^\lambda = \partial^\mu(\varepsilon_{\mu\nu\rho\lambda}A^\nu \partial^\rho A^\lambda)$.

Spinor representation. The spinor representations are those with $j_L = 1/2$ or $j_R = 1/2$. Specifically, we define the left-hand spinor $\psi_L = (\psi_L^1, \psi_L^2)^T$ and right-hand spinor $\psi_R = (\psi_R^1, \psi_R^2)^T$ whose transformations define Λ_L and Λ_R :

$$\Lambda_L(\boldsymbol{\theta}, \boldsymbol{\beta}) = \exp\left(\frac{i}{2}\boldsymbol{\theta} \cdot \boldsymbol{\sigma} - \frac{1}{2}\boldsymbol{\beta} \cdot \boldsymbol{\sigma}\right), \quad \Lambda_R(\boldsymbol{\theta}, \boldsymbol{\beta}) = \exp\left(\frac{i}{2}\boldsymbol{\theta} \cdot \boldsymbol{\sigma} + \frac{1}{2}\boldsymbol{\beta} \cdot \boldsymbol{\sigma}\right) = \Lambda_L(\boldsymbol{\theta}, -\boldsymbol{\beta}). \quad (7)$$

We can create a “scalar-like” object by the inner product of the spinors. Note that the $\psi_L^\dagger \psi_R$ and $\psi_R^\dagger \psi_L$ are Lorentz invariant objects, while the products of two single-handed spinors like $\psi_L^\dagger \psi_L$ or $\psi_R^\dagger \psi_R$ are not. However, note that $-i\sigma^2 \psi_L^*$ transforms like a right-handed spinor.¹ Thus, the left-hand and right-hand spinor can be interchanged by the action of a “time reversal” operation $-i\sigma^2 \mathcal{K}$ (where \mathcal{K} is complex conjugation). The bilinears

$$\psi_L \cdot \psi_L \equiv -i\psi_L^T \sigma^2 \psi_L, \quad \psi_R \cdot \psi_R \equiv -i\psi_R^T \sigma^2 \psi_R$$

are therefore invariants. The spinor field theory consisting of only the left-hand spinor is called the *Majorana theory*:

$$\mathcal{L}_{\text{Maj}} = i\psi_L^\dagger \bar{\sigma}^\mu \partial_\mu \psi_L - m(\psi_L \cdot \psi_L + \psi_L^\dagger \cdot \psi_L^\dagger). \quad (8)$$

The field theory containing both left-hand and right-hand spinors is called the Dirac theory:

$$\mathcal{L}_{\text{Dirac}} = \bar{\psi} (i\not{\partial} - m) \psi \equiv \bar{\psi} (i\gamma^\mu \partial_\mu - m) \psi, \quad (9)$$

where the Dirac spinor contains left- and right-handed Weyl spinors:

$$\psi = \begin{pmatrix} \psi_L \\ \psi_R \end{pmatrix}, \quad \bar{\psi} = (\psi_R^\dagger \quad \psi_L^\dagger), \quad \gamma^\mu = \begin{bmatrix} 0 & \sigma^\mu \\ \bar{\sigma}^\mu & 0 \end{bmatrix}.$$

We remark that under the Dirac spinor basis, the Lorentz algebra is generated by $M^{\mu\nu} = \frac{i}{4}[\gamma^\mu, \gamma^\nu]$, or

$$J_i = \begin{bmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{bmatrix}, \quad K_i = \frac{i}{2} \begin{bmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{bmatrix},$$

using the familiar parametrization (2). We can easily check that these matrices obey the transformation property (7).

Finally, we note that there is another invariants consisting of Dirac field and the vector gauge field:

$$\mathcal{L}_{\text{int}} = q\bar{\psi}\gamma^\mu A_\mu \psi, \quad (10)$$

which produce the QED Lagrangian $\mathcal{L}_{\text{QED}} = \mathcal{L}_{\text{Dirac}} + \mathcal{L}_{\text{Maxwell}} + \mathcal{L}_{\text{int}}$.

II. SCATTERING

A. $e^+ e^- \rightarrow \mu^+ \mu^-$

Consider the scattering process ($e_{p_1} + \bar{e}_{p_2} \rightarrow \mu_{p_3} + \bar{\mu}_{p_4}$). To the first order, the amplitude correspond to the simplest tree level diagram. Using the Feynman rule, this process gives the amplitude

$$i\mathcal{M} = (-ie)^2 \bar{u}(p_3) \gamma^\mu v(p_4) \frac{-ig_{\mu\nu}}{(p_1 + p_2)^2} \bar{v}(p_2) \gamma^\nu u(p_1). \quad (11)$$

Use the Mandelstam variables $s = p_1 + p_2$, the amplitude is

$$\mathcal{M} = \frac{e^2}{s} [\bar{u}(p_3) \gamma^\mu v(p_4)] [\bar{v}(p_2) \gamma_\mu u(p_1)]. \quad (12)$$

The complex conjugate of the amplitude is

$$\mathcal{M}^\dagger = \frac{e^2}{s} [\bar{u}(p_1) \gamma_\mu v(p_2)] [\bar{v}(p_4) \gamma^\mu u(p_3)]. \quad (13)$$

Note that we have use the relation

$$(\bar{u}\gamma^\mu v)^\dagger = v^\dagger \gamma^{\mu\dagger} \gamma^0 u = v^\dagger \gamma^0 \gamma^\mu u = \bar{v} \gamma^\mu u. \quad (14)$$

¹ Using the identity $\sigma^2 \cdot \sigma^* \cdot \sigma^2 = -\sigma$, we have $\sigma^2 \psi_L^* \rightarrow \sigma^2 \exp(-\frac{i}{2}\boldsymbol{\theta} \cdot \boldsymbol{\sigma}^* - \frac{1}{2}\boldsymbol{\beta} \cdot \boldsymbol{\sigma}^*) \sigma^2 \psi_L^* = \Lambda_L(\boldsymbol{\theta}, -\boldsymbol{\beta}) \sigma^2 \psi_L^*$.

Therefore,

$$|\mathcal{M}|^2 = \frac{e^4}{s^2} [\bar{v}(p_4)\gamma^\mu u(p_3)] [\bar{u}(p_3)\gamma^\nu v(p_4)] [\bar{v}(p_2)\gamma_\mu u(p_1)] [\bar{u}(p_1)\gamma_\nu v(p_2)]. \quad (15)$$

If in the scattering experiment the spins are not measured, the spin averaged cross section is just the sum of spin indices. The spin sum can actually simplify the expression, as it gives the orthogonal relations

$$\sum_s u(p)_s \bar{u}_s(p) = \not{p} + m, \quad \sum_s v(p)_s \bar{v}_s(p) = \not{p} - m. \quad (16)$$

The spin averaged amplitude is

$$\frac{1}{4} \sum_{\text{spins}} |\mathcal{M}|^2 = \frac{e^2}{4s^2} \text{Tr}[(\not{p}_4 - m_\mu)\gamma^\mu(\not{p}_3 + m_\mu)\gamma^\nu] \text{Tr}[(\not{p}_2 - m_e)\gamma_\mu(\not{p}_1 + m_e)\gamma_\nu]. \quad (17)$$

Now we evaluate the gamma traces, using the facts $\text{Tr}[\gamma^\mu\gamma^\nu] = 4g^{\mu\nu}$ and $\text{Tr}[\gamma^\mu\gamma^\nu\gamma^\rho\gamma^\sigma] = 4g^{\mu\nu}g^{\rho\sigma} - 4g^{\mu\rho}g^{\nu\sigma} + 4g^{\mu\sigma}g^{\nu\rho}$. The first trace is

$$\begin{aligned} \text{Tr}[(\not{p}_4 - m_\mu)\gamma^\mu(\not{p}_3 + m_\mu)\gamma^\nu] &= \text{Tr}[\gamma^\alpha\gamma^\mu\gamma^\beta\gamma^\nu] p_{4\alpha}p_{3\beta} - m_\mu^2 \text{Tr}[\gamma^\mu\gamma^\nu] \\ &= 4(p_3^\mu p_4^\nu + p_3^\nu p_4^\mu) - 4g^{\mu\nu}(p_3 \cdot p_4 + m_\mu^2) \end{aligned} \quad (18)$$

The second is of the same form as the first, we can similarly get

$$\text{Tr}[(\not{p}_2 - m_e)\gamma_\mu(\not{p}_1 + m_e)\gamma_\nu] = 4(p_{1\mu}p_{2\nu} + p_{1\nu}p_{2\mu}) - 4g_{\mu\nu}(p_1 \cdot p_2 + m_e^2). \quad (19)$$

The final result of the spin average amplitude is

$$\frac{1}{4} \sum_{\text{spins}} |\mathcal{M}|^2 = \frac{8e^4}{s^2} (p_{13}p_{24} + p_{14}p_{23} + m_\mu^2 p_{12} + m_e^2 p_{34} + 2m_e^2 m_\mu^2), \quad (20)$$

where $p_{ij} \equiv p_i \cdot p_j$. The amplitude is in a better form with the Mandelstam variables s , t and u :

$$\begin{aligned} s &= (p_1 + p_2)^2 = (p_3 + p_4)^2 = 2m_e^2 + 2p_{12} = 2m_\mu^2 + 2p_{34}, \\ t &= (p_1 - p_3)^2 = (p_2 - p_4)^2 = m_e^2 + m_\mu^2 - 2p_{13} = m_e^2 + m_\mu^2 - 2p_{24}, \\ u &= (p_1 - p_4)^2 = (p_2 - p_3)^2 = m_e^2 + m_\mu^2 - 2p_{14} = m_e^2 + m_\mu^2 - 2p_{23}. \end{aligned} \quad (21)$$

Note that $s + t + u = 2m_e^2 + 2m_\mu^2$. After some algebra, we get

$$\frac{1}{4} \sum_{\text{spins}} |\mathcal{M}|^2 = \frac{2e^4}{s^2} [t^2 + u^2 + 4s(m_e^2 + m_\mu^2) - 2(m_e^2 + m_\mu^2)^2]. \quad (22)$$

B. $e^- p^+ \rightarrow e^- p^+$

Next we consider the Rutherford scattering ($e_{p_1} + p_{p_2} \rightarrow e_{p_3} + p_{p_4}$). This corresponds to the same tree-level diagram as ($e_{p_1} + \bar{e}_{p_2} \rightarrow \mu_{p_3} + \bar{\mu}_{p_4}$) except a rotation. We immediately get the amplitude:

$$|\mathcal{M}|^2 = -\frac{e^4}{t^2} [\bar{u}(p_3)\gamma^\mu u(p_1)] [\bar{u}(p_1)\gamma^\nu u(p_3)] [\bar{u}(p_4)\gamma_\mu u(p_2)] [\bar{u}(p_2)\gamma_\nu u(p_4)]. \quad (23)$$

Note that the minus sign comes from the positive charge of proton. The spin-averaged amplitude is

$$\frac{1}{4} \sum_{\text{spins}} |\mathcal{M}|^2 = \frac{e^2}{4t^2} \text{Tr}[(\not{p}_3 + m_e)\gamma^\mu(\not{p}_1 + m_e)\gamma^\nu] \text{Tr}[(\not{p}_4 + m_p)\gamma_\mu(\not{p}_2 + m_p)\gamma_\nu]. \quad (24)$$

The trace is evaluated similarly:

$$\begin{aligned} \text{Tr}[(\not{p}_3 + m_e)\gamma^\mu(\not{p}_1 + m_e)\gamma^\nu] &= 4(p_1^\mu p_3^\nu + p_1^\nu p_3^\mu) - 4g^{\mu\nu}(p_{13} - m_e^2), \\ \text{Tr}[(\not{p}_4 + m_p)\gamma_\mu(\not{p}_2 + m_p)\gamma_\nu] &= 4(p_{2\mu}p_{4\nu} + p_{2\nu}p_{4\mu}) - 4g_{\mu\nu}(p_{24} - m_p^2). \end{aligned} \quad (25)$$

Therefore,

$$\frac{1}{4} \sum_{\text{spins}} |\mathcal{M}|^2 = \frac{8e^4}{t^2} (p_{12}p_{34} + p_{14}p_{23} - m_p^2 p_{13} - m_e^2 p_{24} + 2m_e^2 m_p^2). \quad (26)$$

The Mandelstam variables are

$$\begin{aligned} s &= m_e^2 + m_p^2 + 2p_{12} = m_e^2 + m_p^2 + 2p_{34}, \\ t &= 2m_e^2 - 2p_{13} = 2m_p^2 - 2p_{24}, \\ u &= m_e^2 + m_p^2 - 2p_{14} = m_e^2 + m_p^2 - 2p_{23}. \end{aligned} \quad (27)$$

After some algebra we get

$$\frac{1}{4} \sum_{\text{spins}} |\mathcal{M}|^2 = \frac{2e^4}{t^2} \left[u^2 + s^2 + 4t(m_e^2 + m_p^2) - 2(m_e^2 + m_p^2)^2 \right]. \quad (28)$$

C. $\gamma e^- \rightarrow \gamma e^-$

Next we consider the Compton scattering ($\gamma_k + e_p \rightarrow \gamma_{k'} + e_{p'}$). There are two corresponding tree-level diagrams: one is an s -process, the other is a u -process. The amplitude for s -process is

$$\mathcal{M}_s = -\frac{e^2}{s - m_e^2} \epsilon_\mu(k') \bar{u}(p') \gamma^\mu (\not{p} + \not{k} + m_e) \gamma^\nu u(p) \epsilon_\nu(k), \quad (29)$$

and the amplitude for u -process is

$$\mathcal{M}_u = -\frac{e^2}{u - m_e^2} \epsilon_\mu(k) \bar{u}(p') \gamma^\mu (\not{p} - \not{k}' + m_e) \gamma^\nu u(p) \epsilon_\nu(k'). \quad (30)$$

Together, the coherent amplitude is

$$\mathcal{M} = e^2 \epsilon_\mu(k') \bar{u}(p') \left[\frac{\gamma^\mu (\not{p} + \not{k} + m_e) \gamma^\nu}{s - m_e^2} + \frac{\gamma^\nu (\not{p} - \not{k}' + m_e) \gamma^\mu}{u - m_e^2} \right] u(p) \epsilon_\nu(k). \quad (31)$$

The conjugate amplitude is

$$\mathcal{M}^\dagger = e^2 \epsilon_\alpha(k) \bar{u}(p) \left[\frac{\gamma^\alpha (\not{p} + \not{k} + m_e) \gamma^\beta}{s - m_e^2} + \frac{\gamma^\beta (\not{p} - \not{k}' + m_e) \gamma^\alpha}{u - m_e^2} \right] u(p') \epsilon_\beta(k'). \quad (32)$$

If we do not measure the photon polarization, we can use the polarization averaged result. The photon polarization sum is

$$\sum_i \epsilon_\mu^i(k) \epsilon_\nu^i(k) = -g_{\mu\nu} + \frac{1}{2E^2} (p_\mu \bar{p}_\nu + \bar{p}_\mu p_\nu). \quad (33)$$

The second term does not contribute to the amplitude due to the Ward identity. We can then simply replace the polarization sum with the metric $-g_{\mu\nu}$. In this way,

$$\frac{1}{4} \sum_{s,p} |\mathcal{M}_s|^2 = \frac{e^4}{(s - m_e^2)^2} \text{Tr} [(\not{p}_2 - m_e) \gamma_\nu (\not{p}_1 + \not{p}_2 + m_e) \gamma_\mu (\not{p}_4 + m_e) \gamma^\mu (\not{p}_1 + \not{p}_2 + m_e) \gamma^\nu]. \quad (34)$$

III. PERTURBATIVE RENORMALIZATION

As with the scalar field, the partition function with source is defined as

$$Z[\bar{\eta}, \eta, J] = \exp \left\{ i \int d^d x \mathcal{L}_{\text{int}} \left[\frac{\delta}{i\delta J}, \frac{\delta}{i\delta \eta}, \frac{i\delta}{\delta \bar{\eta}} \right] \right\} Z_0[\bar{\eta}, \eta, J]. \quad (35)$$

A. Vacuum Polarization

Consider the one-loop correction to the photon propagator:

$$\begin{aligned}
 \text{Diagram: } & \text{A photon line with momentum } k-p \text{ enters a loop from the left, and a photon line with momentum } p \text{ exits to the right. The loop contains a fermion line with momentum } k \text{ and a fermion line with momentum } p. \text{ The loop is labeled with } \mu, \beta, \alpha, \gamma, \tau, \nu. \\
 & \simeq (-ie_R)^2 A_\mu \overbrace{\bar{\psi}_\alpha \gamma_\alpha^\mu \psi_\beta A_\nu \bar{\psi}_\gamma \gamma_\gamma^\nu \psi_\tau} \\
 & \equiv iA_\mu \Pi^{\mu\nu}(p) A_\nu.
 \end{aligned} \tag{43}$$

The self energy is:

$$\begin{aligned}
 i\Pi^{\mu\nu}(p) &= -e_R^2 \int \frac{d^4 k}{(2\pi)^4} \text{Tr} \left[\gamma^\mu G_F^{(0)}(k-p) \gamma^\nu G_F^{(0)}(k) \right] \\
 &= -e_R^2 \int \frac{d^4 k}{(2\pi)^4} \frac{\text{Tr} [\gamma^\mu (\not{k} - \not{p} + m_R) \gamma^\nu (\not{k} + m_R)]}{(k^2 - m_R^2)[(p-k)^2 - m_R^2]}.
 \end{aligned} \tag{44}$$

The trace of the Dirac matrices can be evaluated in Mathematic using the FeynCalc package:

```
(*Dirac trace using FeynCalc*)
res=DiracTrace[GA[\[Mu]] . (GS[k-p]+m) . GA[\[Nu]] . (GS[k]+m)] ;
DiracSimplify[res]
```

The Dirac trace is:

$$\begin{aligned}
 & \text{Tr} [\gamma^\mu (\not{k} - \not{p} + m_R) \gamma^\nu (\not{k} + m_R)] \\
 &= 4 [g^{\mu\nu} (k \cdot p - k^2 + m_R^2) + 2k^\mu k^\nu - k^\mu p^\nu - p^\mu k^\nu].
 \end{aligned} \tag{45}$$

Using the Feynman parameters, the denominator is:

$$\begin{aligned}
 \frac{1}{(k^2 - m_R^2)[(p-k)^2 - m_R^2]} &= \frac{1}{\{[k - p(1-x)]^2 - [m_R^2 + p^2 x(x-1)]\}^2} \\
 &\equiv \frac{1}{\{[k - p(1-x)]^2 - D_x\}^2}.
 \end{aligned} \tag{46}$$

Since the Ward identity requires that the p^μ term in the propagator do not contribute to any scattering process, we then shift $k \rightarrow k + p(1-x)$ and drop all p^μ linear term. The final result is simplified to:

$$\begin{aligned}
 i\Pi^{\mu\nu}(p) &= -4e_R^2 \int_0^1 dx \int \frac{d^4 k}{(2\pi)^4} \frac{2k^\mu k^\nu - g^{\mu\nu} [k^2 - x(1-x)p^2 - m_R^2]}{[k^2 - D_x]^2} \\
 &\simeq 4e_R^2 g^{\mu\nu} \int_0^1 dx \int \frac{d^4 k}{(2\pi)^4} \frac{\frac{1}{2}k^2 - x(1-x)p^2 - m_R^2}{[k^2 - D_x]^2} \\
 &\simeq -ie_R^2 g^{\mu\nu} \int_0^1 dx \frac{\Omega_d \tilde{\mu}^\varepsilon}{(2\pi)^d} \int_0^\infty dk k^{d-1} \frac{(4 - \frac{2}{d})k^2 + 4x(1-x)p^2 + 4m_R^2}{[k^2 + D_x]^2}.
 \end{aligned} \tag{47}$$

where we have made the Wick rotation, shifted the dimensionality to $(d = 4 - \varepsilon)$, and made the substitution (since the self-energy $i\Pi^{\mu\nu} \propto g^{\mu\nu}$):

$$k^\mu k^\nu \rightarrow \frac{1}{d} k^2 g^{\mu\nu}. \tag{48}$$

The remaining problem is to regularize and renormalize the divergent integral

$$I_\varepsilon(x) \equiv \frac{\Omega_{4-\varepsilon} \tilde{\mu}^\varepsilon}{(2\pi)^{4-\varepsilon}} \int_0^\infty dk k^{3-\varepsilon} \frac{\left(4 - \frac{8}{4-\varepsilon}\right) k^2 + 4x(1-x)p^2 + 4m_R^2}{[k^2 + D_x]^2}. \tag{49}$$

1. Regularization and Renormalization

In $(4 - \varepsilon)$ -dimensional Euclidean space, the integral is convergent. The ε -expansion is carried out in Mathematica using the following code:

```

omg = (2*Pi^(d/2))/(Gamma[d/2]);
cof = \[Mu]^(4-d)*omg/(2*Pi)^d;
nom = k^(d-1)*((4-8/(4-\[Epsilon]))k^2+4x*(1-x)p^2+4m^2);
int = cof*Integrate[nom/(k^2+D)^2,{k,0,Infinity}][[1]];
map = D->m^2-p^2*x*(1-x);
ans = Series[int/.{d->4-\[Epsilon]},{\[Epsilon],0,0}];
ans /. map // Simplify

```

The result is

$$I_\varepsilon(x) = \frac{p^2 x(1-x)}{2\pi^2} \left[\frac{2}{\varepsilon} + \ln \left(\frac{4\pi e^{-\gamma_E} \tilde{\mu}^2}{m_R^2 - p^2 x(1-x)} \right) \right] \quad (50)$$

So the photon self-energy is (also denote $\mu^2 = 4\pi e^{-\gamma_E} \tilde{\mu}^2$):

$$\Pi^{\mu\nu}(p) = -\frac{e_R^2 p^2 g^{\mu\nu}}{6\pi^2 \varepsilon} - \frac{e_R^2 p^2 g^{\mu\nu}}{2\pi^2} \int_0^1 dx x(1-x) \ln \left[\frac{\mu^2}{m_R^2 - p^2 x(1-x)} \right] \quad (51)$$

The counter term coefficient can be chosen as

$$\delta_A = -\frac{e_R^2}{6\pi^2 \varepsilon}. \quad (52)$$

The renormalized photon self-energy is then

$$\begin{aligned} \Pi^{\mu\nu}(p) &= -\frac{e_R^2 p^2 g^{\mu\nu}}{2\pi^2} \int_0^1 dx x(1-x) \ln \left[\frac{\mu^2}{m_R^2 - p^2 x(1-x)} \right] \\ &= \frac{e_R^2 p^2 g^{\mu\nu}}{2\pi^2} \left\{ \frac{1}{3} \ln \left(\frac{m_R}{\mu} \right) + \int_0^1 dx x(1-x) \ln \left[1 - \frac{p^2 x(1-x)}{m_R^2} \right] \right\}. \end{aligned} \quad (53)$$

2. Physical Observable

The photon self-energy has the form

$$\Pi^{\mu\nu}(p) = -e_R^2 [g^{\mu\nu} - (1 - \xi) p^\mu p^\nu] g^{\mu\nu} \Pi_2(p), \quad (54)$$

where

$$\Pi_2(p) = \frac{1}{2\pi^2} \int_0^1 dx x(1-x) \ln \left[\frac{\mu^2}{m_R^2 - p^2 x(1-x)} \right] \quad (55)$$

The one-loop correction to photon propagator is

$$\begin{aligned} iG_\gamma^{\mu\nu}(p) &= -i \frac{g^{\mu\nu}}{p^2} \left(1 + \sum_{n=1}^{\infty} (-e_R^2)^n \Pi_2^n(p) \right) \\ &= -i \frac{g^{\mu\nu}}{p^2 [1 + e_R^2 \Pi_2(p)]}. \end{aligned} \quad (56)$$

We can choose the on-shell condition that the photon has no rest mass:

$$\Pi_2(0) = 0 \quad \implies \quad \mu = m_R. \quad (57)$$

Note that the propagator is related to the Coulomb potential.² To the second order,

$$\begin{aligned} V(p) &= e_R^2 \frac{1 - e_R^2 \Pi_2(p)}{p^2} + O(e_R^6) \\ &= \frac{e_R^2}{p^2} \left\{ 1 + \frac{e_R^2}{2\pi^2} \int_0^1 dx \, x(1-x) \ln \left[1 - \frac{p^2 x(1-x)}{m_R^2} \right] + O(e_R^4) \right\}. \end{aligned} \quad (58)$$

Consider the small momentum limit, where the integral is approximated by

$$\int_0^1 dx \, x(1-x) \ln \left[1 - \frac{p^2 x(1-x)}{m_R^2} \right] \approx -\frac{p^2}{m_R^2} \int_0^1 dx \, x^2(1-x)^2 = -\frac{p^2}{30m_R^2}. \quad (59)$$

This implies

$$V(p) = \frac{e_R^2}{p^2} - \frac{e_R^4}{60\pi^2 m_R^2}. \quad (60)$$

The Fourier transformation gives

$$V(r) = -\frac{e_R^2}{4\pi r} - \frac{e_R^4}{60\pi^2 m_R^2} \delta^{(3)}(r). \quad (61)$$

For atomic orbit, since only the ($L=0$)-orbit have support at $r=0$, this extra potential will shift the spectrum. This effect is called the *Lamb shift*.

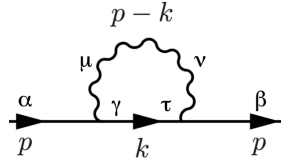
On the other hand, the large momentum limit,

$$V(p) \approx \frac{e_R^2}{p^2} \left[1 + \frac{e_R^2}{12\pi^2} \ln \frac{-p^2}{m_R^2} \right], \quad (62)$$

which predicts a *Landau pole* beyond which perturbation theory breaks down.

B. One-loop Correction to Electron Propagator

Consider the one-loop correction to the particle propagator:



$$\simeq (-ie_R)^2 \overbrace{A_\mu \bar{\psi}_\alpha \gamma_\mu^{\alpha\gamma} \psi_\gamma A_\nu \bar{\psi}_\tau \gamma_\nu^{\tau\beta} \psi_\beta} \equiv i \bar{\psi}_\alpha \Sigma^{\alpha\beta}(p) \psi_\beta. \quad (63)$$

The self energy is

$$\begin{aligned} i\Sigma_{\alpha\beta}(p) &= e_R^2 \int \frac{d^4 k}{(2\pi)^4} G_\gamma^{\mu\nu}(p-k) [\gamma_\mu G_F(k) \gamma_\nu]_{\alpha\beta} \\ &= -e_R^2 \int \frac{d^d k}{(2\pi)^d} \frac{\gamma^\mu (\not{k} + m_R) \gamma_\mu}{(p-k)^2 (k^2 - m_R^2)}. \end{aligned} \quad (64)$$

The second equality comes from the contraction:

$$(-ie)^2 \overbrace{A_\mu \bar{\psi}_\alpha \gamma_\mu^{\alpha\gamma} \psi_\gamma A_\nu \bar{\psi}_\tau \gamma_\nu^{\tau\beta} \psi_\beta} \quad (65)$$

² The Coulomb potential arises just like we derive the force (??), but the sources have additional charge e_R , and the photon is mass less, so $V(p) = \frac{e_R^2}{p^2}$ for free field.

The nominator can be simplified using the Dirac matrix identities:

$$\gamma^\mu \gamma_\mu = d, \quad \gamma^\mu \gamma^\nu \gamma_\mu = (2-d)\gamma^\nu \implies \gamma^\mu (\not{k} + m_R) \gamma_\mu = dm_R + (2-d)\not{k}. \quad (66)$$

The denominator can be simplified using the Feynman parameter:

$$\begin{aligned} \frac{1}{(p-k)^2(k^2-m_R^2)} &= \int_0^1 \frac{dx}{[(k-px)^2 - (1-x)(m_R^2 - p^2x)]^2} \\ &\rightarrow \int_0^1 \frac{dx}{(k^2 - D_x)^2} \end{aligned} \quad (67)$$

where we have shifted $k \rightarrow k + px$ (note this shift also change the numerator).

The self energy becomes (including a $\tilde{\mu}$ mass scale):

$$\begin{aligned} i\Sigma(p) &= e_R^2 \tilde{\mu}^\varepsilon \int_0^1 [(2-\varepsilon)x\not{p} - (4-\varepsilon)m_R] dx \int \frac{d^d k}{(2\pi)^d} \frac{1}{(k^2 - D_x)^2} \\ &= ie_R^2 \int_0^1 dx [(2-\varepsilon)x\not{p} - (4-\varepsilon)m_R] \frac{\tilde{\mu}^\varepsilon \Omega_d}{(2\pi)^d} \int \frac{k^{d-1} dk}{(k^2 + D_x)^2}. \end{aligned} \quad (68)$$

1. Regularization and Renormalization

The regularization procedure is carried out by the following code:

```
omg=(2*Pi^(d/2))/(Gamma[d/2]);
cof=[Mu]^(4-d)*omg/(2*Pi)^d;
int=cof*Integrate[q^(d-1)/(q^2+D)^2,{q,0,Infinity}][[1]];
map={e->Sqrt[4*Pi*Alpha],EulerGamma->Subscript[Gamma,E]};
ans=Series[int/.{d->4-[Epsilon]},{[Epsilon],0,0}];
ans/.map//Simplify
```

The result is ($\mu^2 = 4\pi e^{-\gamma_E} \tilde{\mu}^2$)

$$\begin{aligned} \Sigma(p) &= \frac{e_R^2}{16\pi^2} \int_0^1 dx [(2-\varepsilon)x\not{p} - (4-\varepsilon)m_R] \left[\frac{2}{\varepsilon} + \ln \frac{\mu^2}{(1-x)(m_R^2 - p^2x)} \right] \\ &= \frac{e_R^2}{16\pi^2} \left\{ \int_0^1 dx [2x\not{p} - 4m_R] \left[\frac{2}{\varepsilon} + \ln \frac{\mu^2}{(1-x)(m_R^2 - p^2x)} \right] - \not{p} + 2m_R \right\}. \end{aligned} \quad (69)$$

The infinity comes from

$$\frac{e_R^2}{4\pi^2\varepsilon} \int_0^1 dx (x\not{p} - 2m_R) = \frac{e_R^2}{8\pi^2\varepsilon} \not{p} - \frac{e_R^2}{2\pi^2\varepsilon} m_R. \quad (70)$$

Using the $\overline{\text{MS}}$ subtraction scheme, we choose

$$\delta_\psi = -\frac{e_R^2}{8\pi^2\varepsilon}, \quad \delta_m = -\frac{3e_R^2}{8\pi^2\varepsilon}, \quad (71)$$

and the self energy is

$$\Sigma(p) = \frac{e_R^2}{16\pi^2} \left\{ \int_0^1 dx (2x\not{p} - 4m_R) \ln \left[\frac{\mu^2}{(1-x)(m_R^2 - p^2x)} \right] - \not{p} + 2m_R \right\}. \quad (72)$$

2. Physical Observables

The Dyson series gives:

$$iG_F(p) = \frac{i}{\not{p} - m_R + \Sigma(p)} \quad (73)$$

Experimentally, for a given The on-shell subtraction requires that the m_R equals to the physical mass:

$$\Sigma(\not{p})|_{\not{p}=m_R} = 0, \quad \frac{d}{d\not{p}}\Sigma(\not{p})\Big|_{\not{p}=m_R} = 0. \quad (74)$$

To implement the on-shell condition, we have to modify the subtraction scheme to

$$\begin{aligned} \delta_\psi &= -\frac{e_R^2}{8\pi^2} \left(\frac{1}{\varepsilon} + \ln \frac{\mu}{m_R} + A \right), \\ \delta_m &= -\frac{e_R^2}{8\pi^2} \left(\frac{3}{\varepsilon} + 3 \ln \frac{\mu}{m_R} + B \right), \end{aligned} \quad (75)$$

and the self energy is

$$\begin{aligned} \Sigma(\not{p}) &= -\frac{e_R^2}{8\pi^2} \int_0^1 dx (x\not{p} - 2m_R) \ln \left[(1-x) \left(1 - \frac{p^2}{m_R^2} x \right) \right] \\ &\quad - \frac{e_R^2}{8\pi^2} \left[\left(A + \frac{1}{2} \right) \not{p} - (A + B + 1)m_R \right]. \end{aligned} \quad (76)$$

The first condition

$$\begin{aligned} \Sigma(\not{p})|_{\not{p}=m_R} &= -\frac{e_R^2}{8\pi^2} m_R \left[\int_0^1 dx (x-2) \ln(1-x)^2 - B - \frac{1}{2} \right] \\ &= -\frac{e_R^2}{8\pi^2} m_R (2 - B) = 0 \end{aligned} \quad (77)$$

gives the mass renormalization coefficient

$$\delta_m = -\frac{e_R^2}{8\pi^2} \left(\frac{3}{\varepsilon} + 3 \ln \frac{\mu}{m_R} + 2 \right). \quad (78)$$

While in the derivative of the self-energy:

$$\frac{d}{d\not{p}}\Sigma(m_R) = -\frac{e_R^2}{8\pi^2} \left\{ \int_0^1 dx \left[x \ln(1-x)^2 - \frac{2x(x-2)}{1-x} \ln(1-x) \right] + A + \frac{1}{2} \right\}, \quad (79)$$

there is a divergent integral:

$$\int_0^1 dx \frac{2x(x-2)}{1-x} \ln(1-x), \quad (80)$$

indicating an IR divergence. We can never the less get rid of it by introducing a small mass m_γ for photon (which will be set to zero). This mass term change the denominator in the loop integral:

$$\frac{1}{[(p-k)^2 - m_\gamma^2](k^2 - m_R^2)} = \int_0^1 \frac{dx}{[(k - px)^2 - D_x - x m_\gamma^2]^2}. \quad (81)$$

Most derivation remains the same, we just need to make a substitution in the finial result:

$$D_x \rightarrow D_x + x m_\gamma^2. \quad (82)$$

Especially, the introducing of the photon mass will not change the result of the mass renormalization factor we have computed.

The modified self-energy is then

$$\begin{aligned} \Sigma(\not{p}) &= -\frac{e_R^2}{8\pi^2} \int_0^1 dx (x\not{p} - 2m_R) \ln \left[(1-x) \left(1 - \frac{p^2}{m_R^2} x \right) + x \frac{m_\gamma^2}{m_R^2} \right] \\ &\quad - \frac{e_R^2}{8\pi^2} \left[\left(A + \frac{1}{2} \right) \not{p} - (A + B + 1)m_R \right]. \end{aligned} \quad (83)$$

The derivative is now

$$\frac{d}{d\not{p}}\Sigma(m_R) = -\frac{e_R^2}{8\pi^2} \left\{ \int_0^1 dx \left[x \ln(1-x)^2 + \frac{2x(2-x)(1-x)}{(1-x)^2 + x \frac{m_\gamma^2}{m_R^2}} \right] + A + \frac{1}{2} \right\}, \quad (84)$$

Note that in the $(m_\gamma \rightarrow 0)$ limit, the asymptotic behavior of the originally divergent integral is

$$\lim_{m_\gamma \rightarrow 0} \int_0^1 dx \frac{2x(2-x)(1-x)}{(1-x)^2 + x \frac{m_\gamma^2}{m_R^2}} = -1 - 2 \ln \frac{m_\gamma}{m_R}. \quad (85)$$

So the second subtraction condition is:

$$\frac{d}{d\not{p}}\Sigma(m_R) = -\frac{e_R^2}{8\pi^2} \left(A - 2 - 2 \frac{m_\gamma}{m_R} \right) = 0. \quad (86)$$

The field strength renormalization is

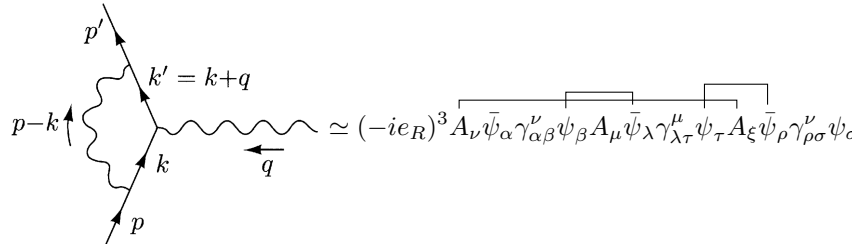
$$\delta_\psi = -\frac{e_R^2}{8\pi^2} \left(\frac{1}{\varepsilon} + \ln \frac{\mu}{m_R} + 2 + 2 \ln \frac{m_\gamma}{m_R} \right). \quad (87)$$

The final self energy is (shall take the $m_\gamma \rightarrow 0$ limit):

$$\begin{aligned} \Sigma(\not{p}) = & -\frac{e_R^2}{16\pi^2} \int_0^1 dx (2x\not{p} - 4m_R) \ln \left[(1-x) \left(1 - \frac{p^2}{m_R^2} x \right) + 2x \ln \frac{m_\gamma}{m_R} \right] \\ & - \frac{e_R^2}{16\pi^2} \left[\left(5 + 4 \ln \frac{m_\gamma}{m_R} \right) \not{p} - \left(10 + 4 \ln \frac{m_\gamma}{m_R} \right) m_R \right]. \end{aligned} \quad (88)$$

C. One-loop Correction to Vertex

Consider the one-loop correction to interaction:



$$\begin{aligned} & \simeq (-ie_R)^3 A_\nu \bar{\psi}_\alpha \gamma_{\alpha\beta}^\nu \psi_\beta A_\mu \bar{\psi}_\lambda \gamma_{\lambda\tau}^\mu \psi_\tau A_\xi \bar{\psi}_\rho \gamma_{\rho\sigma}^\xi \psi_\sigma \\ & \equiv -ie A_\mu \Gamma_{\alpha\beta}^\mu(q, p, p') \bar{\psi}_\alpha \psi_\beta. \end{aligned} \quad (89)$$

The vertex function is:

$$\begin{aligned} i\Gamma_{\alpha\beta}^\mu(q, p, p') &= -e_R^2 \int \frac{d^4 k}{(2\pi)^4} G_\gamma^{\nu\lambda}(p-k) [\gamma_\nu G_F(k') \gamma^\mu G_F(k) \gamma_\lambda]_{\alpha\beta} \\ &= e_R^2 \int \frac{d^4 k}{(2\pi)^4} \frac{[\gamma^\nu (\not{k}' + m_R) \gamma^\mu (\not{k} + m_R) \gamma_\nu]_{\alpha\beta}}{(k^2 - m_R^2)(k'^2 - m_R^2)(p-k)^2} \end{aligned} \quad (90)$$

Using the following code

```
(*numerator*)
den=Contract[GA[\[Nu]].(GS[kp]+m).GA[\[Mu]].(GS[k]+m).GA[\[Nu]]];
DiracSimplify[den]

(*Feynman parameter*)
A1=k^2-m^2;
```

```

A2=(k+q)^2-m^2;
A3=(p-k)^2;
{c,b,a}=CoefficientList[x*A1+y*A2+z*A3,{k}];
-b/2//Simplify
-c+b^2/4//Simplify

```

The numerator is

$$-2k^\mu k'^\mu - 2m_R^2 \gamma^\mu + 4m_R(k+k')^\mu. \quad (91)$$

The denominator is

$$\int \frac{dF_3}{[(k+yq-zp)^2 - D_{xyz}]^3}, \quad (92)$$

where

$$\begin{aligned} D_{xyz} &= (x+y)m^2 - z(1-z)p^2 - y(1-y)q^2 - 2yzpq \\ &= (x+y)m_R^2 - xyq^2 - yzp'^2 - xzp^2. \end{aligned}$$

Shift $k^\mu \rightarrow k^\mu + zq_1^\mu - yp^\mu$, throw away all terms with linear k^μ , and replace $k^\mu k^\nu$ with $\frac{1}{d}k^2 g^{\mu\nu}$, the result is

$$\frac{4}{d}k^2 \gamma^\mu - 2(-yq + zp)\gamma^\mu [(1-y)q + zp] + 4m_R^2 \gamma^\mu - 2m_R [(1-2y)q^\mu + 2zp^\mu].$$

Note that only the quadratic term is divergent.

$$\Gamma^\mu(p, q_1, q_2) = -i \frac{4e^2 \tilde{\mu}^\epsilon \gamma^\mu}{d} \int dF_3 \int \frac{d^d k}{(2\pi)^d} \frac{k^2}{(k^2 - D)^3} + \delta\Gamma^\mu(p, q_1, q_2).$$

where $\delta\Gamma^\mu$ stores all the finite part

$$\begin{aligned} &\delta\Gamma^\mu(p, q_1, q_2) \\ &= \int \frac{e^2 k^3 dk dF_3}{(2\pi)^2 (k^2 + D)^3} \{(-yq + zp)\gamma^\mu [(1-y)q + zp] - 2m_R^2 \gamma^\mu + m_R [(1-2y)q^\mu + 2zp^\mu]\}. \end{aligned}$$

The divergent part is

$$\frac{4e^2 \tilde{\mu}^\epsilon \Omega_d \gamma^\mu}{d(2\pi)^d} \int dF_3 \int \frac{k^{d+1} dk}{(k^2 + D)^3} = \frac{e_R^2}{16\pi^2} \gamma^\mu \int dF_3 \left(\frac{2}{\epsilon} + \ln \frac{\mu^2}{D_{xyz}} \right). \quad (93)$$

Using the $\overline{\text{MS}}$ scheme, the coefficient of the counter term is

$$\delta_e = -\frac{e_R^2}{8\pi^2 \epsilon}. \quad (94)$$

IV. SYSTEMATIC RENORMALIZATION

A. Renormalization Group

In summery, the renormalization factors are

$$\begin{aligned} Z_\psi &= 1 - \frac{e_R^2}{8\pi^2 \epsilon} + O(e_R^3), \\ Z_A &= 1 - \frac{e_R^2}{6\pi^2 \epsilon} + O(e_R^3), \\ Z_m &= 1 - \frac{e_R^2}{2\pi^2 \epsilon} + O(e_R^3), \\ Z_e &= 1 - \frac{e_R^2}{8\pi^2 \epsilon} + O(e_R^3), \end{aligned} \quad (95)$$

which means

$$\begin{aligned}
\frac{d \ln Z_\phi}{de_R} &= -\frac{e_R}{4\pi^2\epsilon} + O(e_R^2), \\
\frac{d \ln Z_A}{de_R} &= -\frac{e_R}{3\pi^2\epsilon} + O(e_R^2), \\
\frac{d \ln Z_m}{de_R} &= -\frac{e_R}{\pi^2\epsilon} + O(e_R^2), \\
\frac{d \ln Z_e}{de_R} &= -\frac{e_R}{4\pi^2\epsilon} + O(e_R^2).
\end{aligned} \tag{96}$$

The bare parameters are

$$\begin{aligned}
\psi_0 &= Z_\psi^{1/2} \psi_R, \\
A_0 &= Z_A^{1/2} A_R, \\
m_0 &= Z_m Z_\psi^{-1} m_R, \\
e_0 &= Z_e Z_\psi^{-1} Z_A^{-1/2} e_R \tilde{\mu}^{\epsilon/2}.
\end{aligned} \tag{97}$$

The RG equation for e_0 is

$$\frac{d \ln e_0}{d \ln \mu} = \left(\frac{\ln Z_e}{de_R} - \frac{\ln Z_\psi}{de_R} - \frac{1}{2} \frac{\ln Z_A}{de_R} + \frac{1}{e_R} \right) \frac{de_R}{d \ln \mu} + \frac{\epsilon}{2} = 0. \tag{98}$$

The beta function is

$$\beta(e_R) = \frac{de_R}{d \ln \mu} = -\frac{\epsilon}{2} e_R + \frac{e_R^3}{12\pi^2} + O(e_R^4). \tag{99}$$

The RG equation for m_0 is

$$\frac{d \ln m_0}{d \ln \mu} = \left(\frac{\ln Z_m}{de_R} - \frac{\ln Z_\psi}{de_R} \right) \frac{de_R}{d \ln \mu} + \frac{1}{m_R} \frac{dm_R}{d \ln \mu} = 0. \tag{100}$$

The anomalous dimension of mass is

$$\gamma_m = \frac{d \ln m_R}{d \ln \mu} = -\frac{3e_R^2}{8\pi^2} + O(e_R^3). \tag{101}$$