The Infrared Telescope Facility (IRTF) spectral library: Spectral diagnostics for cool stars

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ABSTRACT

Context. The near-infrared (NIR) wavelength range offers some unique spectral features, and it is less prone to the extinction than the optical one. Recently, the first flux calibrated NIR library of cool stars from the NASA Infrared Telescope Facility (IRTF) have become available, and it has not been fully exploited yet.

Aims. We want to develop spectroscopic diagnostics for stellar physical parameters based on features in the wavelength range $1-5 \mu m$. In this work we test the technique in the *I* and *K* bands. The study of the *Y*, *J*, *H*, and *L* bands will be presented in the following paper. *Methods*. An objective method for semi-empirical definition of spectral features sensitive to various physical parameters is applied to the spectra. It is based on sensitivity map – i.e., derivative of the flux in the spectra with respect to the stellar parameters at a fixed wavelength. New optimized indices are defined and their equivalent widths (EWs) are measured.

Results. The method is applied in the *I*- and *K*-band windows of the IRTF stellar spectra to verify the new technique by comparing the results with the known behavior of well-studied spectral features. A number of sensitive features to the effective temperature and surface gravity are re-identified or newly identified clearly showing the reliability of the sensitivity map analysis.

Conclusions. The sensitivity map allows to identify the best bandpass limits for the line and nearby continuum. It reliably predicts the trends of spectral features with respect to a given physical parameter but not their absolute strengths. Line blends are easy to recognize when blended features have different behavior with respect to some physical stellar parameter. The use of sensitivity map is therefore complementary to the use of indices. We give the EWs of the new indices measured for the IRTF star sample. This new and homogeneous set of EWs will be useful for stellar population synthesis models and can be used to get element-by-element abundances for unresolved stellar population studies in galaxies.

Key words. infrared: stars - line: identification - stars: abundances - supergiants - stars: late-type - stars: fundamental parameters

1. Introduction

The interpretation of the spectral absorption features in the integrated light of galaxies requires the understanding of the behavior of these features in the spectra of the stars as function of their effective temperature $T_{\rm eff}$, surface gravity $\log{(g)}$, and metal abundance (usually parametrized with the Iron abundance [Fe/H] and the ratio [α /Fe] of alpha-elements to Iron). This explains why some major observational and theoretical efforts are invested into the development of extensive stellar libraries for population synthesis e.g., (Burstein et al. 1984; Worthey et al. 1994; Trager et al. 1998; Tripicco & Bell 1995; Korn et al. 2005). The knowledge derived from the analysis of the stellar spectra is promptly transferred to the galaxy integrated light because $T_{\rm eff}$ and $\log{(g)}$ are related to the age of the dominant stellar population, and [Fe/H] is related to the overall galactic chemical abundance.

Usually, first the sensitivity of the spectral absorption features to the physical parameters of stars is empirically investigated, and then stellar photosphere models are adopted to explain the behavior of the features within the theory of line formation. The former step is typically limited to looking for correlations between of the line strengths and $T_{\rm eff}$, $\log(g)$, or [Fe/H]. A notable exception is the work by Worthey et al. (1994, see their

Tables 2 and 3) who determined the derivatives of these relations and adopted them as *sensitivity indices* for individual lines in order to compare the relative sensitivity of various features to the physical parameters of the stellar population.

Prompted by this idea, we attempt here to develop it further to its logical conclusion by defining the sensitivity map, that is a continuous derivative across the stellar spectra of stars for different $T_{\rm eff}$, $\log{(g)}$, or [Fe/H] spanning the full range of parameters for a given wavelength.

This new method allowed us to characterize the behavior of the individual spectral features and it is an useful and objective tool to define at the best the spectral indices more sensible to physical stellar parameters. Moreover, it accounts for the variable sensitivity of individual lines. It was developed in order to explore near-infrared (NIR) spectral ranges, which have never been used for stellar population analysis before (i.e., $1-1.8\,\mu\text{m}$). First we applied it to some well-understood features, such as the calcium triplet (CaT hereafter) at $0.85\,\mu\text{m}$ and CO band at $2.29\,\mu\text{m}$, to demonstrate its feasibility. In this paper we also present the sensitivity map for a number of other NIR indices with respect to spectral type (SpT hereafter) and surface gravity. In a forthcoming paper we will extend such an analysis to other spectral ranges to define new indices, optimized for stellar

population analysis. The new set of NIR indices will be useful for both stellar and extragalactic astrophysics.

The structure of the paper is as follows. Section 2 gives a short description of the adopted Infrared Telescope Facility (IRTF) spectral library. Section 3 summarizes the physical parameters of the stars in the library. The method used to create the sensitivity map is described in Sect. 4. The absorption features and line indices in the *I* band and their use as spectral diagnostics are discussed in Sects. 5 and 6, respectively. Sections 7 and 8 are devoted to *K*-band indices. Section 9 discusses and summarizes the results.

2. The IRTF spectral library

The IRTF spectral library (Cushing et al. 2005; Rayner et al. 2009), which is the first NIR library of flux-calibrated stellar spectra and extends the optical library described in Gorgas et al. (1993), contains 210 cool stars. The spectra were taken with the cross-disperser medium-resolution infrared spectrograph SpeX (Rayner et al. 2003) mounted at the 3.0-m NASA Infrared Telescope Facility (IRTF) on Mauna Kea, Hawaii. The observations were carried out with two different set-ups with a resolving power $R \approx 2000$ at $0.8-2.4 \,\mu\text{m}$ and $R \approx 2500$ at $2.4-5 \,\mu\text{m}$, respectively. Several spectral orders were simultaneously recorded during a single exposure with significant wavelength overlap between the adjacent orders making it easier to preserve the continuum shape. It results to be reliable to within a few percent, as verified by generating a set of synthetic Two-Micron All Sky Survey (2MASS) colors from the spectra (Rayner et al. 2009). Therefore, the IRTF library allowed us to measure the strong and broad molecular absorption bands that are common in the NIR.

Most of the stars were observed in the $0.8-2.5\,\mu\mathrm{m}$ range, and for a small fraction of them the wavelength coverage extends up to $5\,\mu\mathrm{m}$. The wide wavelength range permits to connect some well-studied regions as the I and K bands, with the relatively unexplored J, H, and L bands. The sample contains the F, G, K, M, and L stars, spanning luminosity classes from I to V. It also includes some AGB, carbon, and S stars. The SpTs were derived from optical spectra in the framework of the MK classification system (Rayner et al. 2009).

The sample is limited to bright stars (Fig. 1), guaranteeing high signal-to-noise ratio (S/N) even in the thermal infrared region ($\sim 2.3-5\,\mu m$). It is $S/N \geq 100$ across most of the wavelength range, with the exception of regions with poor atmospheric transmission and with $\lambda > 4\,\mu m$. The brightness limit has a downside. Indeed, most of the sample stars are located nearby in the Milky Ways disk, and they have near-solar chemical composition. The abundance distribution of the sample stars with known [Fe/H] (Cayrel de Strobel et al. 1997) is shown in Fig. 1(a) of Rayner et al. (2009), and it is representative for the stars in the solar neighborhood (Nordström et al. 2004).

3. Physical parameters of the sample stars

The sample stars span a wide range in SpTs and luminosity classes. The photospheric data for the sample stars are taken from literature. This is an inhomogeneous compilation, and in an effort to homogenize the data as much as possible (at least in terms of abundance estimate methods) we used spectroscopic determinations when possible. We adopted the metallicities based on narrow-band photometry when these were the only ones available in the literature. The *K*-band absolute luminosities were calculated by convolving the spectra with a filter

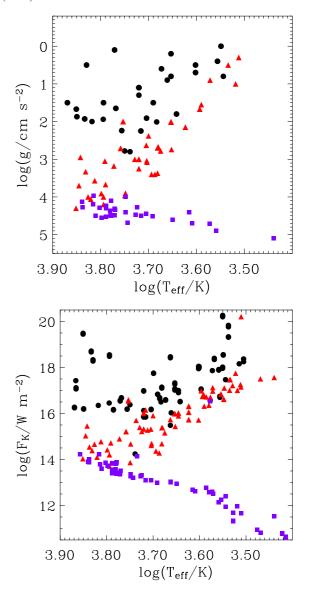


Fig. 1. Surface gravity $\log(g)$ (top panel) and K-band flux $\log(F_K)$ (bottom panel) as a function of effective temperature $\log(T_{\rm eff})$ for the sample stars.

transmission curve. If $\log{(T_{\rm eff})}$ was not available from literature, we estimated it from the SpT- $T_{\rm eff}$ relation by Carroll & Ostlie (1996). Part of the sample lacks gravity determination. The parallaxes were obtained from the HIPPARCOS catalog (Perryman et al. 1997). All the collected parameters of the IRTF stars are listed in Table 3.

Figure 1 shows how the sample stars populate the $\log(g) - \log(T_{\rm eff})$ and $\log(L_K) - \log(T_{\rm eff})$ planes. Their metallicity distribution is shown in Fig. 2. The average [Fe/H] is about -0.1 dex with a dispersion of 0.2 dex, which is typical of the solar neighborhood.

4. Method

This section presents the empirical method we used to identify the most sensitive line-strength indices to a particular stellar physical parameter, and the procedure to define them. The method is free from any a priori assumption and it can be applied to any star or line-strength index. But here it was constrained to the range of SpTs, luminosity classes, and spectral coverage

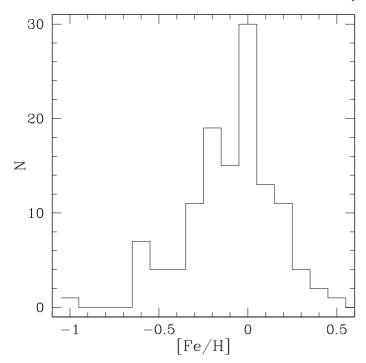


Fig. 2. Metallicity distribution of the sample stars with spectroscopic measurements of [Fe/H].

spanned by the IRTF library. Our goal is to find proxies for $T_{\rm eff}$, $\log{(g)}$ and [Fe/H], except for that the temperature is replaced by SpT, as it is more readily available. We numbered the SpT assigning value of 0 to F0 stars, 1 to G0, 2 to K0, etc., with decimals for the subtypes (e.g., F5 = 0.5; G8 = 1.8; M0.5 = 3.05). In the first application described in this work we considered the spectral ranges that include the CaT and CO bands that have already been widely used for stellar population analysis (e.g., Jones et al. 1984; Armandroff & Da Costa 1991; Mayya 1997; Ivanov et al. 2000; Vazdekis et al. 2003; Mármol-Queraltó et al. 2008). They provided a test-bed for verifying the method.

The first step was to homogenize the spectra re-binning them with a step of $3.872 \times 10^{-4} \mu \text{m}$ pixel⁻¹. The supergiants, giants, and dwarfs were analyzed separately. After experimenting with absolute flux calibrated spectra and continuum-normalized spectra we opted the latter because the normalization lowered the scatter in the final results. The spectra were normalized to unity at $\lambda = 0.88 \, \mu \text{m}$ and $2.20 \, \mu \text{m}$ for the analysis of the *I* and *K* bands, respectively. Both these wavelength regions were chosen because they were free of relevant absorption features.

Next, for each wavelength bin we derived a second-order polynomial least-square fit of the normalized intensity as a function of SpT. The fit was performed under the IDL¹ environment using a specially developed script. An independent fit is obtained for each wavelength bin. The choice of the second-order is discussed in Sect. 9. We defined as *model spectrum* the combined fitting functions for all wavelengths. Such a fit was introduced to provide continuity across the entire parameter space, otherwise, the sensitivity could be evaluated only for discrete values of the physical parameter (the SpT in this case) for which stellar spectra were available.

Finally, we calculated the derivative of the model spectrum with respect to the SpT by building the so-called *sensitivity map*. A spectral index is sensitive to the SpT if the derivative with

respect to the SpT of the model spectrum at the index central wavelength is different from the derivative of the surrounding region. The wavelength regions where such a strong difference is observed are characterized by sharp features in the sensitivity map and can be easily identified. Cuts of the fit and sensitivity map for a fixed SpT in the *I* band are shown for sample supergiants in Figs. 3 and 6, respectively.

To study the features sensitive to gravity we divided the sample into four SpT bins (F, G, K, and M; the L stars in the sample have no gravity measurements and therefore were excluded). In each bin we sorted the flux-normalized spectra in a sequence of increasing $\log(g)$. We then followed the same steps as for the SpT, except for that the fit and derivative were performed along the $\log(g)$ axis. Figures 9 and 10 show the results of the analysis for fixed surface gravity in the I band for the F and G stars and the K and M stars, respectively.

5. Spectral indices in the I band

5.1. Main I-band spectral features

Here we shortly summarize the main *I*-band $(0.80-0.90 \,\mu\text{m})$ spectral features in later type stars. The spectra of F stars are dominated by the neutral hydrogen (H I) absorption lines of the Paschen series. Their strength decreases with increasing wavelength. The Paschen series becomes weaker progressing from supergiants through giants to dwarfs, and from F to late-type G stars. The absorption lines of neutral metals are stronger in G stars than in F stars, and reach a maximum depth in the spectra of K and M stars. The lines of ionized metals (the strongest feature is CaT at $0.86 \,\mu\text{m}$) weaken towards later SpTs. The molecular absorption increases in later types, affecting significantly the slope of the local continuum. In M stars the titanium oxide (TiO) bands are significant and blend with CaT, which is weaker than in earlier types. Progressing from M to L stars, the metal oxides (TiO and VO) are replaced by metal hydrides (CrH at 0.8611 μ m and FeH at $0.8692 \mu m$) as the main molecular species (Rayner et al. 2009).

The CaT was used in different studies over a wide range of atmospheric parameters and was applied to both individual stars and integrated stellar populations in different environments, the calibration is both empirical (i.e., Cenarro et al. 2002) and theoretical (i.e., Du et al. 2012). Several definitions for the CaT index exist: the classical approach consists in establishing a central bandpass covering the spectral feature, and one or more adjacent bandpasses to trace the reference level of the local continuum. Cenarro et al. (2001b, Cen01 hereafter) analyzed previous CaT index definitions and defined a new one, which was specifically designed to avoid contamination from molecular bands and to cover the line wings completely. The latter is an important issue, because the main contributors to the strength of the CaT lines are their wings, whereas the core is not very sensitive to the atmospheric and stellar parameters (Erdelyi-Mendes & Barbuy 1991). The Ca II lines are heavily affected by metallicity and gravity: their strengths increase as metallicity increases and gravity decreases. For a detailed discussion see Sect. 2 of Cen01. To summarize, the CaT in cool stars (~F7-M0) follows a complex behavior with varying temperature, metallicity, and gravity.

5.2. Sensitivity map for the spectral type

The *I*-band model spectrum and sensitivity map for the different SpTs (i.e., for different $T_{\rm eff}$) are shown in Figs. 3–5 and in Figs. 6–8 for supergiant, giant, and dwarf stars, respectively. The different plots are shown with a constant vertical offset for

¹ The interactive Data Language is distributed by ITT Visual Information Solutions.

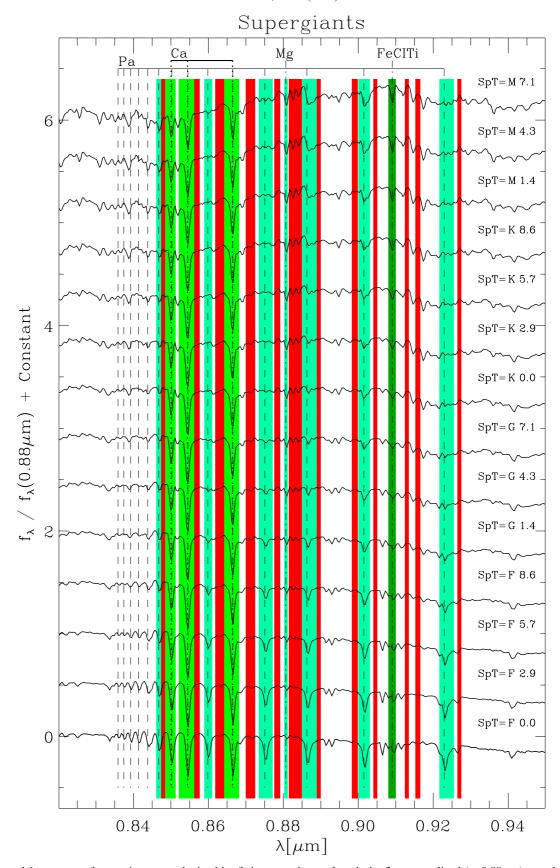


Fig. 3. *I*-band *model spectrum* of supergiant stars obtained by fitting at each wavelength the flux-normalized (at $0.88\,\mu\text{m}$) sample spectra along SpTs. The model spectrum for different SpTs is offset for displaying purposes and the SpT is given. The green and red regions mark the bandpasses of the newly defined indices and their adjacent continuum, respectively (see Table 1). In particular the light green, green and dark green regions mark the Pa, Ca, Mg and FeClTe features, respectively. Some relevant absorption features are marked.

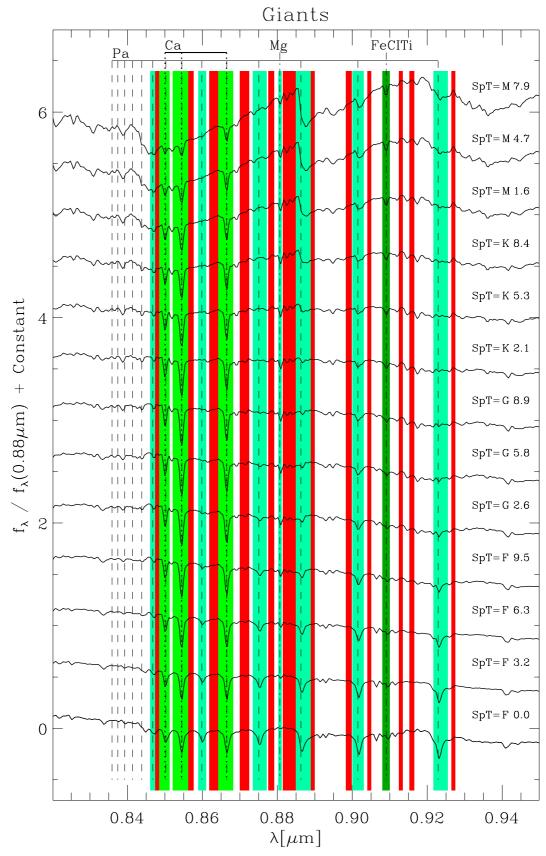


Fig. 4. As in Fig. 3 but for giant stars.

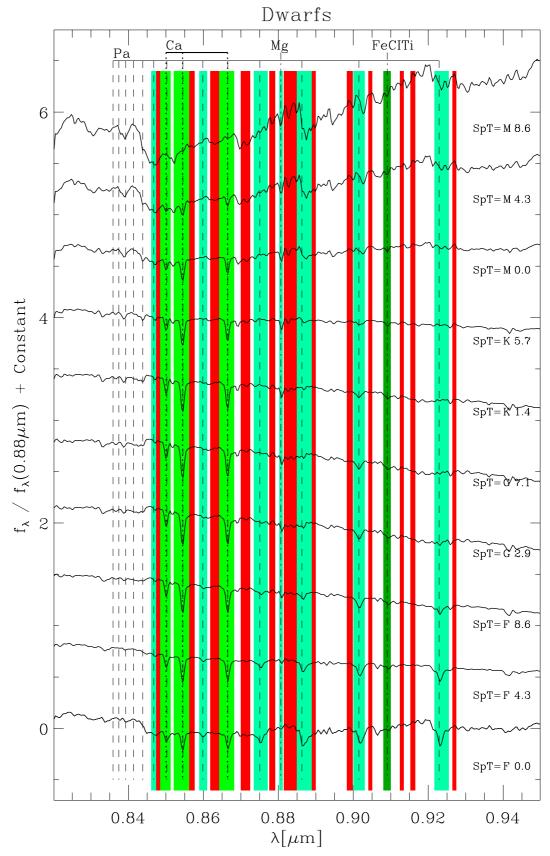


Fig. 5. As in Fig. 3 but for dwarf stars.

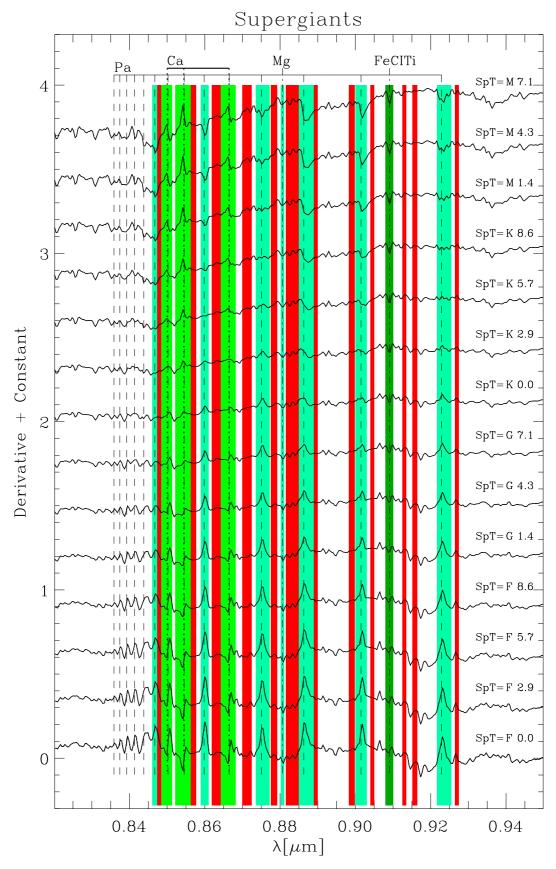


Fig. 6. *I*-band sensitivity map for SpT of supergiant stars. The sensitivity map for different SpTs is offset for displaying purposes and the SpT is given. Symbols are as in Fig. 3.

display purposes and the position of the CaT band, Paschen lines, the Mg and the FeClTi band are marked.

The analysis of the supergiant stars can be used to summarize the features present in the *I* band (Fig. 6). The H I lines from the Paschen series correspond to positive peaks in the sensitivity map over about the F0-K0 SpT range. Their equivalent width (EW hereafter) decreases with SpT, because the flux within the absorption line is growing more rapidly than the flux of the surrounding continuum. From about K0.0 to about K8.6 the sensitivity of the Paschen lines to the SpT is negligible, and from about K8.6 to about M7.1 the sensitivity map shows negative peaks at the Paschen lines because these features decrease toward later types over this range of SpTs. The CaT generally shows just the opposite trend, but with asymmetric sensitivity map because the features are contaminated by H I, so there is a superposition of a negative and a positive peak. Higherresolution spectroscopy with sufficient S/N to allow line decomposition is required when CaT is used as a diagnostic tool. The Magnesium sensitivity map show negative peaks for stars of all spectral types and luminosity classes.

The samples of giant and dwarf stars show similar behavior (Figs. 7 and 8) but the H I lines are less-sensitive temperature indicators than in supergiant stars. As a result, the variation of the CaT comes across clearer, with symmetric, less contaminated peaks. Some cases of contamination are still present, e.g., Pa4 at $0.886\,\mu\mathrm{m}$ is affected by Ni and Fe absorption lines. This index should only be used for hotter stars that show no TiO in their atmospheres. Finally, the variations of molecular features are particularly strong in giant stars, and affect most of the continuum in *I* band, confirming that they are good temperature indicators. The FeClTi band shows no variation in dwarf stars.

5.3. Sensitivity map for the surface gravity

The I-band sensitivity map of the surface gravity is displayed in Figs. 9 and 10 for the F and G stars and the K and M stars, respectively. For the purpose of this analysis the sample was divided according to SpT and the L stars were excluded because of insufficient gravity coverage. As expected, the CaT band shows the strongest variation across all the SpTs, with positive peaks at the core of the lines, consistent with CaT becoming weaker moving from supergiant and giant towards dwarf stars. The sensitivity to surface gravity (i.e., the strongest peaks) decreases for stars with lowest log(g). The Paschen lines follow a similar trend but they are useful only in F and, to some degree, in G stars, disappearing in K and M stars, as discussed in the previous section. The Mg line and FeClTi band show no noticeable variation. The change in the overall shape of the sensitivity map in K and M stars for low log(g) stars is probably due to the broad molecular features.

5.4. Definition of new I-band indices

The sensitivity map helped us to identify $11\ I$ -band features that can be used to measure the SpT and $\log(g)$. We defined spectral indices in order to quantitatively analyze the relation between the EW of these features and stellar physical parameters. The indices consist of a central bandpass covering the feature of interest, and two other adjacent bandpasses, at the red and blue sides, tracing the local continuum. The central bandpass was selected to include the peak of the sensitivity map and the continuum bandpasses were placed on spectral regions where the sensitivity map is (nearly) constant. For a perfectly uniform stellar library the

Table 1. Definition of the bandpasses of the *I*-band indices.

Index	Element	Central bandpass	Continuum bandpasses
		(μm)	(μm)
Pa1	H I $(n = 3)$	0.8461-0.8474	0.8474-0.8484, 0.8563-0.8577
Ca1	Ca II	0.8484-0.8513	$0.8474 - 0.8484, \ 0.8563 - 0.8577$
Ca2	Са п	0.8522 - 0.8562	$0.8474 - 0.8484, \ 0.8563 - 0.8577$
Pa2	H I $(n = 3)$	0.8577-0.8619	$0.8563 - 0.8577, \ 0.8619 - 0.8642$
Ca3	Ca II	0.8642 - 0.8682	$0.8619 – 0.8642, \ 0.8700 – 0.8725$
Pa3	H I $(n = 3)$	0.8730 - 0.8772	$0.8700 – 0.8725, \ 0.8776 – 0.8792$
Mg	Mg I	0.8802 - 0.8811	$0.8776 – 0.8792, \ 0.8815 – 0.8850$
Pa4	H I $(n = 3)$	0.8850 - 0.8890	$0.8815 - 0.8850, \ 0.8890 - 0.8900$
Pa5	H I $(n = 3)$	0.9000-0.9030	$0.8983 - 0.8998, \ 0.9040 - 0.9050$
FeClTi	Fe I,Cl I,Ti I	0.9080-0.9100	$0.9040 - 0.9050, \ 0.9125 - 0.9135$
Pa6	H I $(n = 3)$	0.9217-0.9255	$0.9152 – 0.9165, \ 0.9265 – 0.9275$

central bandpass can be defined to guarantee optimal extraction, but this was not possible in the NIR region so we were conservative and defined slightly wider bandpasses to assure that the spectral features were fully encompassed. If possible, the combined width of the continuum bandpasses was equal or larger than the width of the central bandpass to avoid S/N being degraded. For the Ca II feature the central bandpass overlaps with the Cen01 definition, whereas the continuum bandpasses are different. Six indices were defined to characterize the Paschen series: three have narrower bandpasses than in Cen01 and three have new bandpasses. The Mg feature has the central bandpass overlapping with the Rayner et al. (2009) definition, whereas the continuum bandpasses are different. Finally, we defined a new index centered at $0.9090 \,\mu\mathrm{m}$ to measure the combined contribution of Fe I, Cl I, and Ti I. The bandpasses of the *I*-band indices are listed in Table 1.

The EW is defined as

$$EW = \int_{\lambda 1}^{\lambda 2} (1 - F_{\text{line}}/F_{\text{cont}}) \, d\lambda, \tag{1}$$

where F_{line} is the flux density of the observed spectrum $F(\lambda)$ inside the line bandpass between λ_1 and λ_2 , F_{cont} is the value of the local continuum at the central wavelength of the line bandpass as obtained by linearly interpolating between the two continuum bandpasses, and $\Delta \lambda = \lambda_2 - \lambda_1$ is the width of the line bandpass. The measurements were performed under the IDL environment using a specially developed script. The EW is measured directly on the observed spectrum. To derive the EW errors we first computed the rms of the residual of the spectrum and the continuum (where the continuum is assumed to be the straight line interpolating the two continuum bandpasses) in the two continuum regions. We then estimated the EW errors by means of Monte Carlo simulations on the observed spectrum taking as noise the root mean square (rms) above derived. We also measured the CaT, PaT and CaT* indices as defined by Cen01. The results are listed in Tables 4 and 5.

6. Spectral diagnostics in the I band

6.1. Spectral diagnostics for the spectral type

The new *I*-band indices are plotted as a function of the SpT and $T_{\rm eff}$ in Fig. 11. The plots confirm that the CaT lines are not sensitive to the SpT for stars from F to early-M type, (although they have some sensitivity to the luminosity class, i.e., $\log(g)$, see Sect. 6.2) and show a marked decrease for SpT later than M, vanishing beyond the M5 type. The negative values are due to

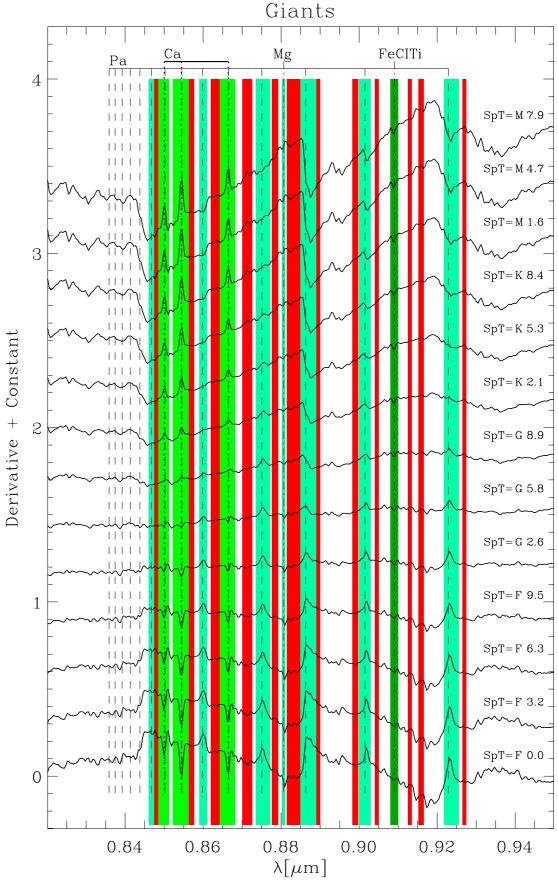


Fig. 7. As in Fig. 6 but for giant stars.

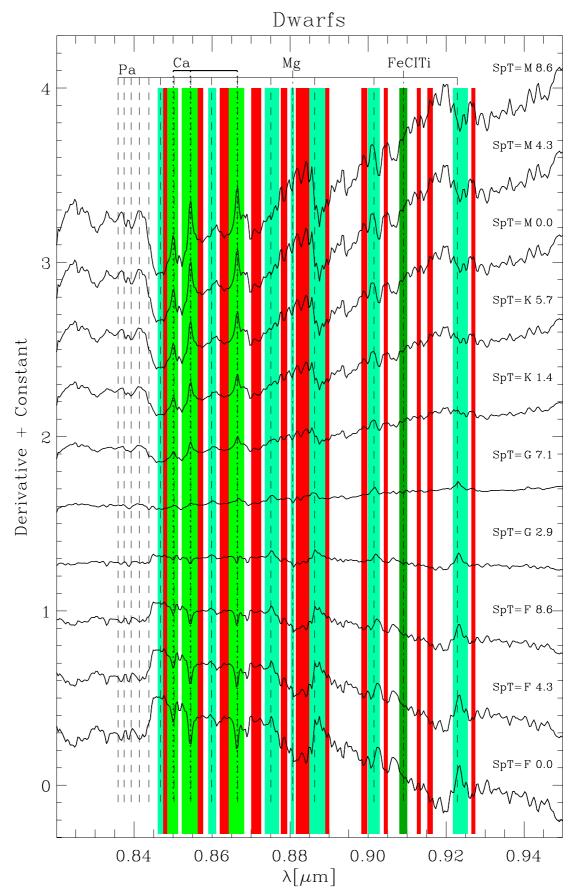


Fig. 8. As in Fig. 6 but for dwarf stars.

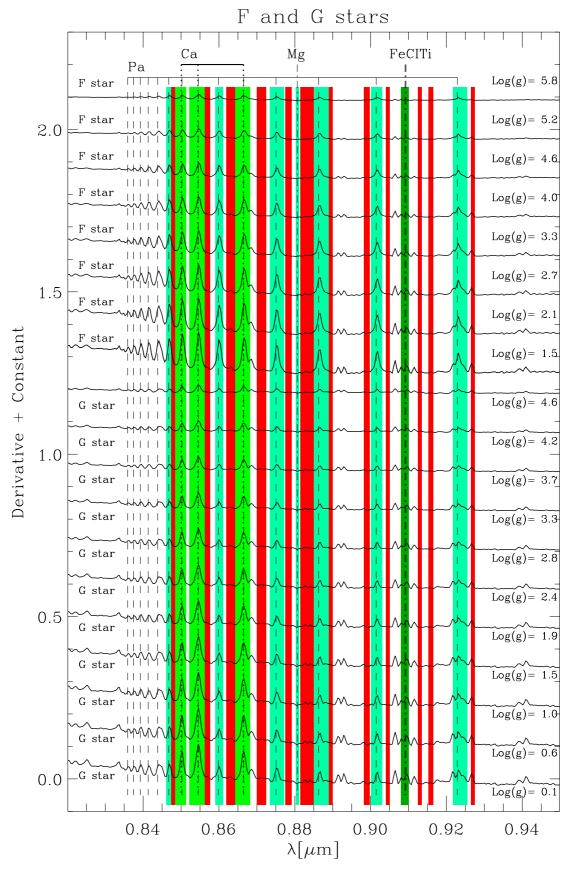


Fig. 9. *I*-band sensitivity map for surface gravity of F (top) and G-type stars (bottom). The sensitivity map for different gravity values is offset for displaying purposes and the central values of the corresponding log(g) bins are given. Symbols are as in Fig. 3.

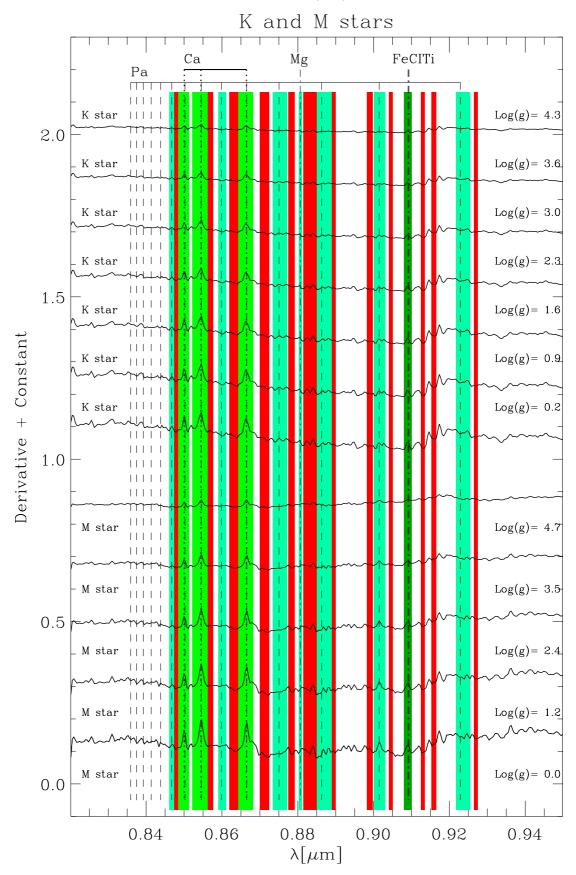


Fig. 10. As in Fig. 9 but for K (*top*) and M-type stars (*bottom*).

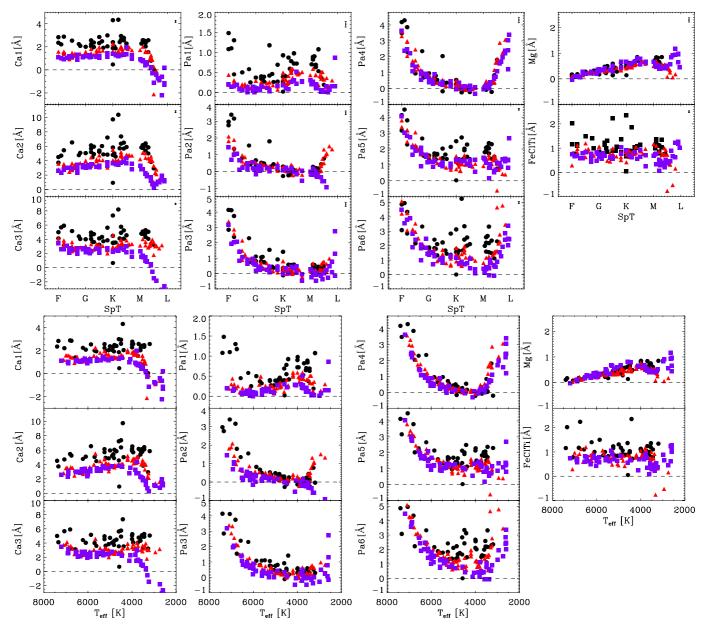


Fig. 11. Equivalent width of the *I*-band indices as function of spectral type (*top panels*) and effective temperature (*bottom panels*). The different symbols correspond to the supergiant (circles), giant (triangles), and dwarf stars (squares), respectively. For each index on the right side of the plot the median of the errors is shown (*top panels*).

the broad TiO bands affecting the continuum, particularly in the dwarf and, to some extent, in giant stars.

Pa1 shows negligible variation with SpT, whereas Pa2 and the other H I lines decrease, most notably from F to early-G stars. The sharp rise beyond the M type is due to molecular contamination, which is strongest in Pa4. The scatter is the largest for the reddest lines, probably due to the increasing of the sky background and worsening of the atmospheric transmission. Mg shows a constant increasing with SpT. Finally, the FeClTi band appears insensitive to the SpT for all luminosity classes.

We conclude that the sensitivity map method correctly recovers the different behavior of Mg, CaT and Paschen lines and it allows to determine the SpT of stars. However, we notice that the CaT in supergiant stars is characterized by a larger scatter than in dwarf and giant stars. This is not due to a metallicity effect (see Sect. 6.2).

6.2. Spectral diagnostics for the surface gravity and metallicity

The Paschen indices show no trend with surface gravity. The sensitivity map indicates a mild decrease of the Paschen lines with increasing $\log(g)$ for F and G stars, but the large EW scatter for the hotter sample stars prevented us from measuring any gradient. K and M stars are characterized by a smaller EW scatter and we found constant Paschen values as expected. The Mg and FeClTi feature do not show any correlation with surface gravity as expected from the sensitivity map.

Weak trends are observed for Ca1 and FeClTi but the narrow metallicity range of the stellar library prevented us from drawing firm conclusions whether these indices can be used to derive [Fe/H]. The CaT index is indeed a well understood metallicity indicator (Terlevich et al. 1989; Tolstoy et al. 2009), although it degenerates at higher metallicities (Vazdekis et al. 2003).

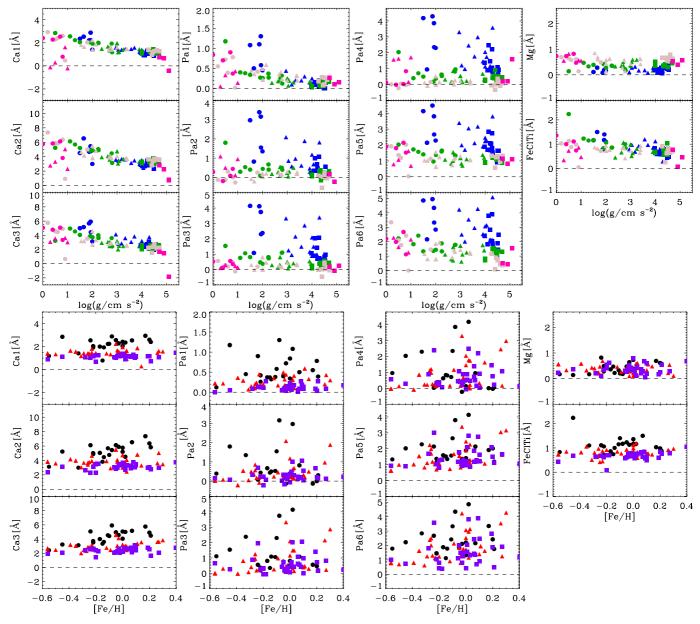


Fig. 12. Equivalent width of the *I*-band indices as function of surface gravity (*top panels*) and metallicity (*bottom panels*). Symbols for supergiant, giant, and dwarf stars and errors are as in Fig. 11. Different colors in top panels correspond to F (blue), G (green), K (grey), and M stars (pink), respectively.

7. Spectral indices in the K band

7.1. Main K-band spectral features

The same analysis described in the previous sections for the I band was carried out for the K band $(1.92-2.40\,\mu\text{m})$, which also contains some well-studied features. The most prominent are the series of Ca I lines at $1.95\,\mu\text{m}$, Na I doublet at $2.21\,\mu\text{m}$, Ca I doublet at $2.26\,\mu\text{m}$, and the series of first-overtone bandheads of ^{12}CO extending redward of $2.29\,\mu\text{m}$ and of ^{13}CO extending redward of $2.34\,\mu\text{m}$ (Cushing et al. 2005). The F stars show H I absorption lines of the Bracket series. They are the Br δ at $1.94\,\mu\text{m}$, which falls in a wavelength region of moderate telluric absorption and Br γ at $2.16\,\mu\text{m}$. The metal features and the CO are known to increase with SpT, whereas the H I lines decrease with SpT (e.g., Ivanov et al. 2000). Finally, broad waterabsorption bands appear on both sides of the K band in late-M

type stars and they smoothly decrease in strength from supergiants through giants to dwarfs (e.g., Lançon et al. 2007).

7.2. Sensitivity map for the spectral type

The K-band model spectrum and sensitivity map as a function of SpT are shown in Figs. 13–15 and in Figs. 16–18 for supergiant, giant, and dwarf stars, respectively. Each luminosity class was considered separately. The analysis of the supergiant stars can be used to summarize the features present in the K band as done for the I band. The bluest part of the K band contains a complex of Fe I and Ca I lines at about $1.95-1.99~\mu m$. The sensitivity map shows dips with variable strength, corresponding to an increase of the line strengths from F0 to late-F stars, a plateau for the F–G types, and a further increase for SpTs from K5 to M7.

The sensitivity map of the Si I features shows a dip for the F stars and it is flat – within the scatter – for later SpTs, indicating that the line looses its sensitivity to $T_{\rm eff}$ for redder stars. The

Bry has a peak which decreases with SpT from F0.0 to K8.6, because the flux in the line rises towards later types faster than the neighboring continuum, making the line weaker. Br δ also corresponds to a peak in the sensitivity map and the line almost completely disappears after K2-5 type. The blend of Mg I at 2.106 μ m and Al I at 2.110 μ m exhibits a shallow dip in the sensitivity map, i.e. it increases mildly with the SpT. The EWs of the Na I doublet at 2.21 μ m and Ca I doublet at 2.26 μ m follow the pattern of the previously discussed Ca lines. The two Fe I at 2.23 and 2.24 μ m are not sensitive to the SpT in the range between F and K and they increase only for M stars. The Mg I at 2.28 μ m shows no marked variations for different SpTs. The 12 CO at 2.29 μ m shows a dip in the sensitivity map and it strongly increases with the SpT.

Summarizing, the spectral features in the supergiant, giant, and dwarf stars show similar behavior – perhaps, with slightly different strengths of the gradients. A difference may be noted in the overall shape of the derivatives due to the broad water vapor absorption at the blue and the red edges of the *K* band in the late-K and M-type dwarf stars.

7.3. Sensitivity map for the surface gravity

The K-band sensitivity map as a function of surface gravity is plotted in Figs. 19 and 20 for the F and G stars and the K and M stars, respectively. The sample is divided according the SpT and the L stars are excluded, as done in the I band. The peaks at the Br δ and Br γ lines are the most prominent features for the F and G stars. They indicate that flux at the core of the lines increases faster than the continuum flux with increasing gravity, i.e., the lines become weaker for more compact stars. Mg, and to lesser extent Na, follow opposite trends, as indicated by the small dips at these lines. The H I lines nearly disappear in the K and M stars and the peaks corresponding to the CO bandheads become the strongest features in the sensitivity map, being characterized by a strong decrease towards lower surface gravity. The Ca and Fe lines in the range $1.95-1.99 \mu m$ show small dips which indicate a mild increase with increasing log(g). On the contrary, the prominent Na and Ca doublets appear insensitive to the stellar gravity.

7.4. Definition of new K-band indices

The sensitivity map allowed us to identify 16 K-band features that could be used as indicators of the SpT and/or $\log(q)$. Eight of them, falling in the spectral range $\lambda = 2.10-2.30 \,\mu\text{m}$, correspond to absorption lines already studied. We followed the previous definition of the line-strength indices by Ivanov et al. (2004) for Mg I at $2.11 \,\mu m$ and Br γ , by Silva et al. (2008) for Fe I at $2.23 \,\mu\text{m}$, Fe I at $2.24 \,\mu\text{m}$, and Mg I at $2.28 \,\mu\text{m}$, and by Cesetti et al. (2009) for the Ca I and Na I doublets and the ¹²CO absorption band at $2.29 \,\mu\text{m}$. For some of them in literature are present multiple index definitions. In this case we preferred to use the index definition more suitable for extragalactic studies. This is the case of Ca I, Na I, and 12 CO. The features for $\lambda > 2.29 \,\mu m$ are not considered in the analysis due to the difficulty of defining a reliable continuum on the red side of the features. Eight indices, falling in the wavelength range $\lambda = 1.92-2.01 \,\mu\text{m}$ are defined by adopting the common continuum passband. The estimated continuum in the wavelength range of the absorption features is shown with a straight line in Figs. 13–15. All the above definitions are listed in Table 2. The indices were measured for all the sample stars and their EWs are listed in Tables 6 and 7.

Table 2. Definition of the bandpasses of the *K*-band indices.

Index	Element	Central bandpass	Continuum bandpasses	Ref.
		(μm)	(μm)	
Fe1	Fe I	1.9297-1.9327	1.9220-1.9260, 2.0030-2.0100	(1)
${ m Br}\delta$	H I $(n = 4)$	1.9425-1.9470	1.9220-1.9260, 2.0030-2.0100	(1)
Ca1	Ca I	1.9500-1.9526	1.9220-1.9260, 2.0030-2.0100	(1)
Fe23	Fe I	1.9583-1.9656	1.9220-1.9260, 2.0030-2.0100	(1)
Si	Si I	1.9708-1.9748	1.9220-1.9260, 2.0030-2.0100	(1)
Ca2	Ca I	1.9769-1.9795	1.9220-1.9260, 2.0030-2.0100	(1)
Ca3	Ca I	1.9847-1.9881	1.9220-1.9260, 2.0030-2.0100	(1)
Ca4	Ca I	1.9917-1.9943	1.9220-1.9260, 2.0030-2.0100	(1)
Mg1	Mg I	2.1040-2.1110	2.1000-2.1040, 2.1110-2.1150	(2)
Brγ	H I $(n = 4)$	2.1639-2.1686	2.0907-2.0951, 2.2873-2.2900	(2)
Na_d	Na I	2.2000-2.2140	2.1934-2.1996, 2.2150-2.2190	(3)
FeA	Fe I	2.2250-2.2299	2.2133-2.2176, 2.2437-2.2479	(4)
FeB	Fe I	2.2368-2.2414	2.2133-2.2176, 2.2437-2.2479	(4)
Ca_d	Ca I	2.2594-2.2700	2.2516-2.2590, 2.2716-2.2888	3 (3)
Mg2	Mg I	2.2795-2.2845	2.2700-2.2720, 2.2850-2.2874	1 (4)
¹² CO	¹² CO(2,0)	2.2910-2.3070	2.2516-2.2590, 2.2716-2.2888	

References. (1) This paper, (2) Ivanov et al. (2004), (3) Cesetti et al. (2009), and (4) Silva et al. (2008).

8. Spectral diagnostics in the K band

8.1. Spectral diagnostics for the spectral type

The new K-band indices are plotted as a function of spectral type in Fig. 21. All indices within the $1.92-2.01\,\mu\mathrm{m}$ range suffer from large scatter – not surprising, given their weakness and the poor atmospheric transmission in this spectral region – but a general increase of most Ca and Fe lines towards later type stars is observed. The Ca2 and Ca3 indices have a turnover at early-to mid-M stars, but their scatter is significant. The Br δ index decreases in hotter stars stars until hitting a turnover at early- to mid-K stars.

The Mg lines increase with SpT from F to K, but for latertype stars the trend flattens or even reverts, and dwarf stars show abnormal scatter. The Na, Ca, and Fe indices follow similar pattern but the strength of the lines makes it more pronounced, especially the turnover for K8-M0 supergiants. The Bry index decreases nearly monotonically towards later SpTs. The 12 CO index is insensitive to SpT for F and early-G stars and it increases for later SpTs. Moreover, it is also able to distinguish between luminosity classes: the index is progressively weaker ranging from supergiant, to giant, and to dwarf stars. The same results were obtained in the analysis of the spectral diagnostics for the $T_{\rm eff}$.

With the exception of a few features, the behavior of indices is similar for stars of the three luminosity classes.

8.2. Spectral diagnostics for surface gravity and metallicity

We also studied how the indices vary as function of the surface gravity and metallicity (Fig. 22). Most features show no significant changes within the errors with two exceptions. The molecular CO band is sensitive to $\log(g)$, as expected. The metal lines of K and M stars peak at the extremely high gravity. This effect did not appear for the *I*-band indices and it is probably due to a combination of two factors: in late-type dwarfs the $T_{\rm eff}$ and $\log(g)$ are correlated (i.e., the start are ordered along the main sequence) and the features do show temperature dependence, as demonstrated in Fig. 21. The large scatter of Na and Ca indices for low gravity K stars is puzzling.

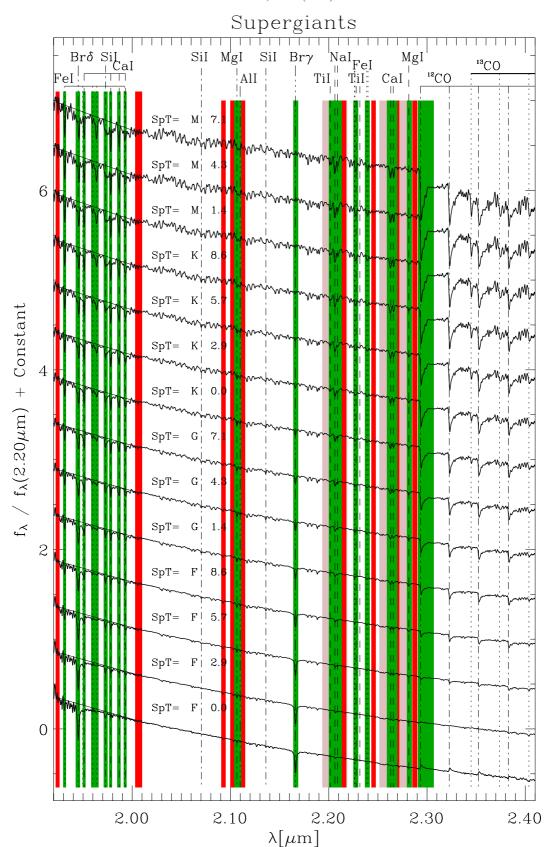


Fig. 13. *I*-band *model spectrum* of supergiant stars obtained by fitting at each wavelength the flux-normalized (at $2.20 \,\mu\text{m}$) sample spectra along SpTs. The model spectrum for different SpTs is offset for displaying purposes and the SpT is given. The green regions mark the bandpasses of the newly defined indices and, the red and grey regions, mark their adjacent continuum as defined in this paper and in literature respectively (see Table 2). Some relevant absorption features are marked.

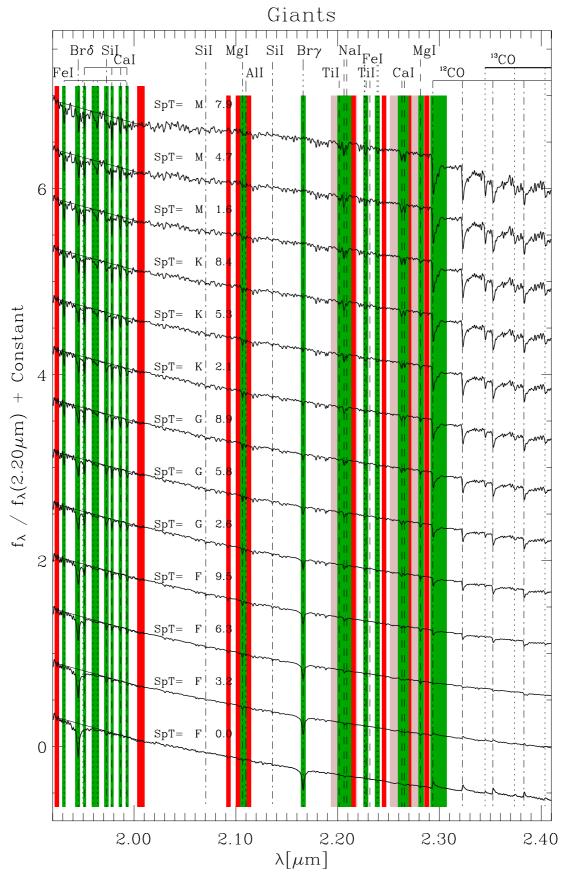


Fig. 14. As in Fig. 13 but for giant stars.

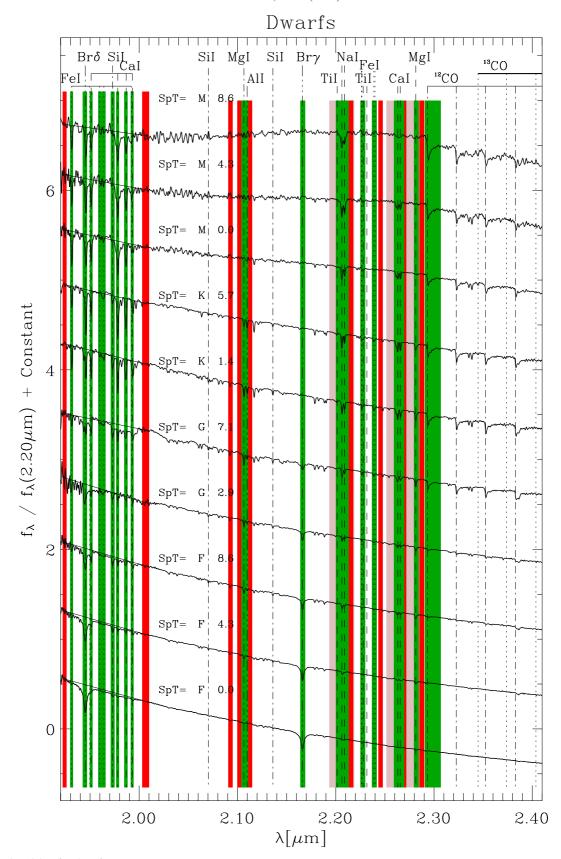


Fig. 15. As in Fig. 13 but for dwarf stars.

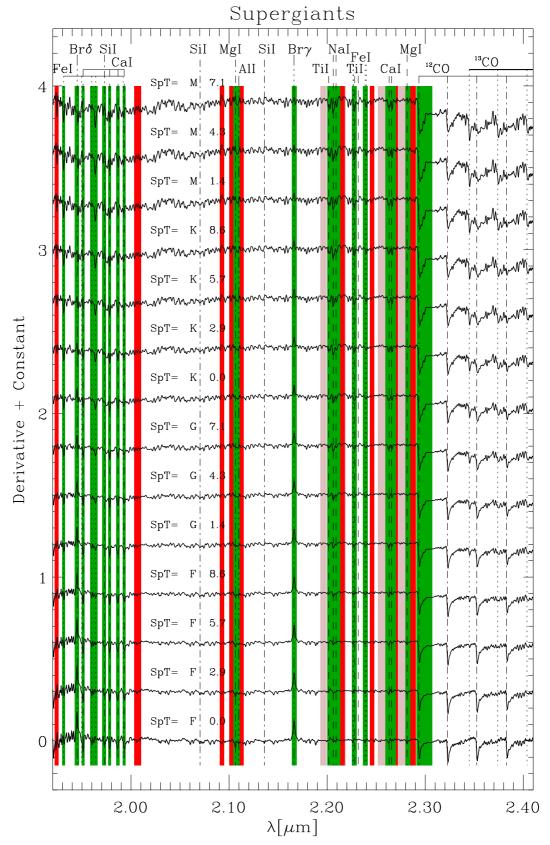


Fig. 16. *K*-band sensitivity map for SpT of supergiant stars. The sensitivity map for different SpTs is offset for displaying purposes and the SpT is given. Symbols are as in Fig. 13.

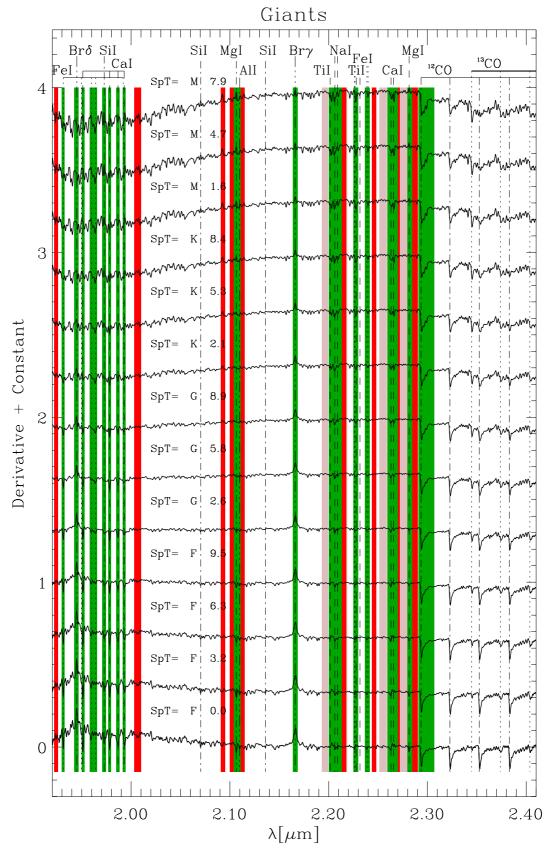


Fig. 17. As in Fig. 16 but for giant stars.

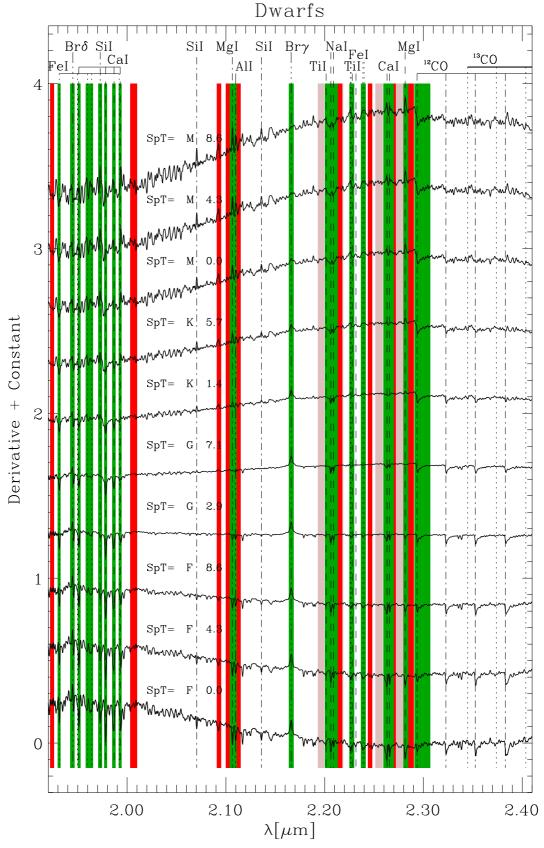


Fig. 18. As in Fig. 16 but for dwarf stars.

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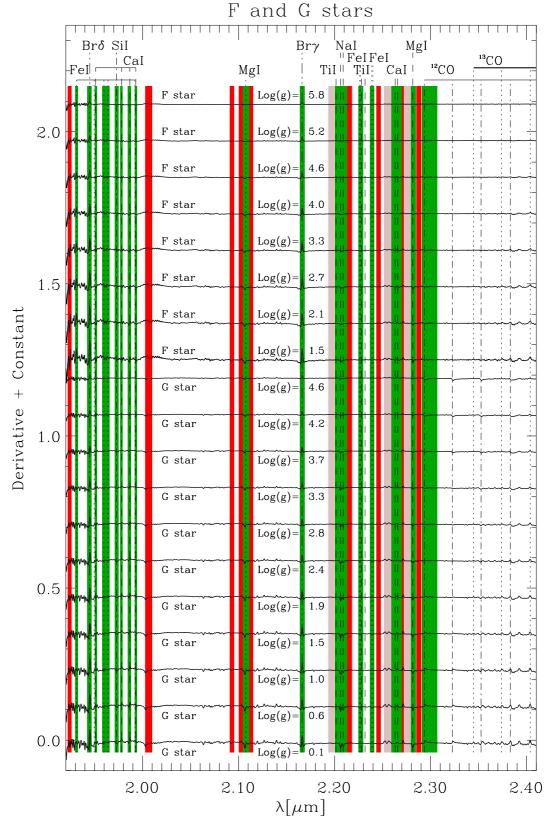


Fig. 19. *K*-band sensitivity map for surface gravity of F (top) and G-type stars (bottom). The sensitivity map for different gravity values is offset for displaying purposes and the central values of the corresponding log(g) bins are given. Symbols are as in Fig. 16.

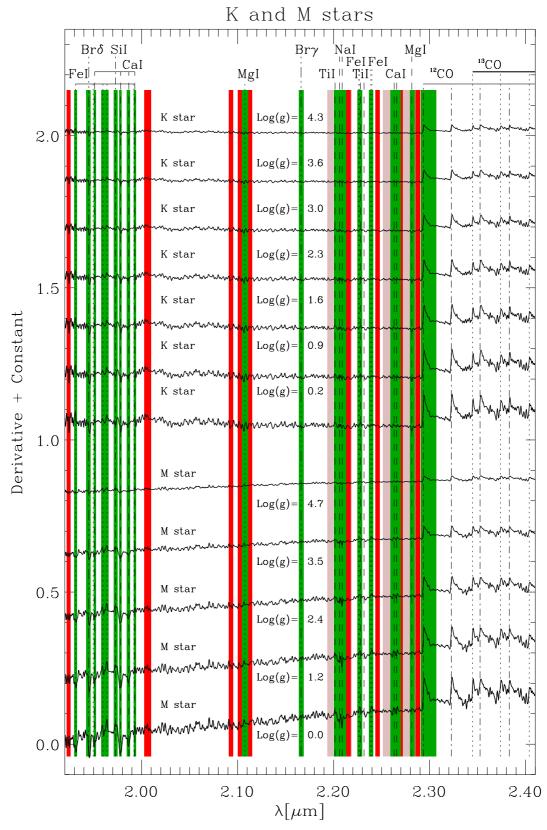


Fig. 20. As in Fig. 19 but for K (top) and M-type stars (bottom).

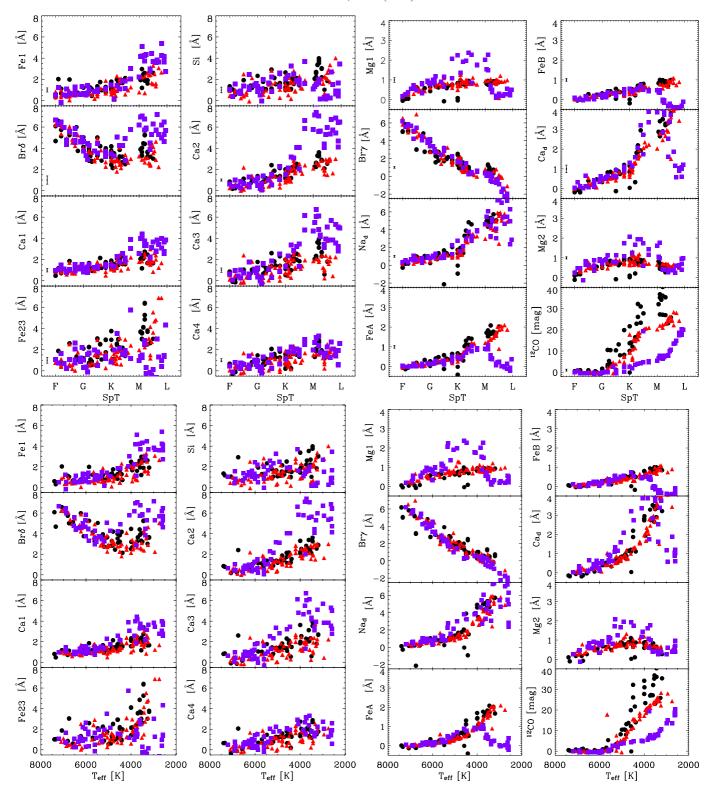


Fig. 21. As in Fig. 11 but for *K*-band indices.

Finally, no trend is observed with [Fe/H] for any of the investigated lines, as expected for the narrow metallicity range of the sample.

9. Discussion and conclusions

The Cenarro et al. (2001a, 2002) library, which is limited to the *I*-band spectral region, is characterized by a larger number of

stars (706 stars with $T_{\rm eff} = 3300-25\,000\,\rm K$) compared to and our sample (198 stars with $T_{\rm eff} = 2600-7200\,\rm K$) and by a higher spectral resolution. This gives us the opportunity to check if our results are affected by poor stellar parameter sampling, spectral resolution, and differences in data reduction.

Spectra direct comparison. We compared directly spectra from the IRTF and Cenarro et al. (2001a, 2002) libraries, rebinning the latter to the resolution of the IRTF spectra. Two

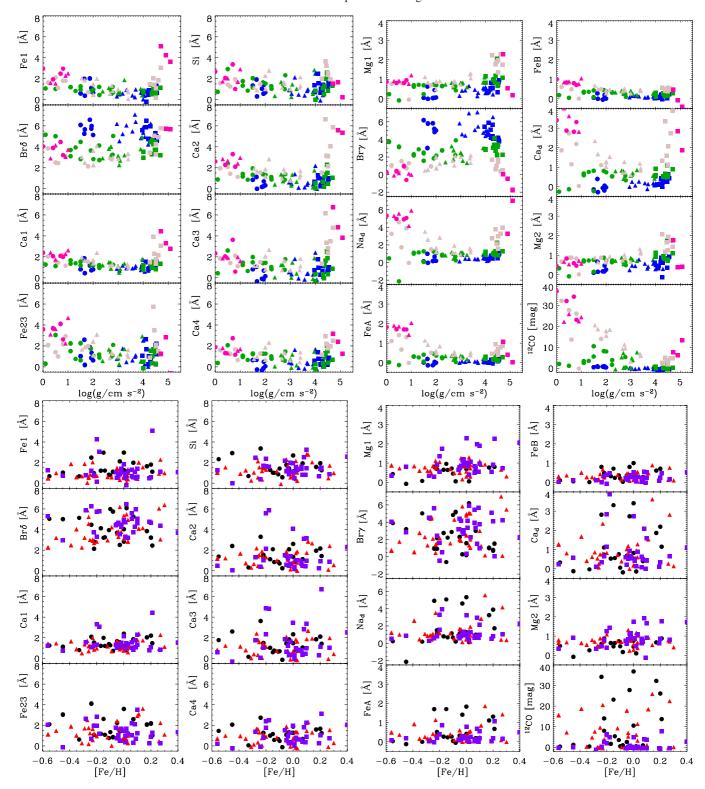


Fig. 22. As in Fig. 12 but for K-band indices.

extreme cases were considered: the M supergiant Betelgeuse (HD 39801), one of the brightest stars in both samples, and a flaring K3 dwarf HD 219134. In Fig. 23 we show for each star the two spectra superimposed and their difference. In the case of HD 219134 we notice a slight difference in the continuum slope. We therefore compared the two spectra also after the continuum normalization (Fig. 23 bottom panel). We see that no telluric corrections residuals are present and the strength of the individual spectral features are the same. We therefore are confident

that the IRTF and Cenarro et al. (2001a, 2002) libraries are consistent.

Calcium triplet comparison. Indices as the CaT were extensively studied both with empiric that theoretic methods. A comprehensive study was done by Cen01 that extensively studied the behavior of this spectral feature in respect to $T_{\rm eff}$, $\log{(g)}$, and [Fe/H]. We compare our sample CaT behavior with Cen01 adopting their definition of the CaT* index that removes the Paschen line contamination. The CaT* is defined

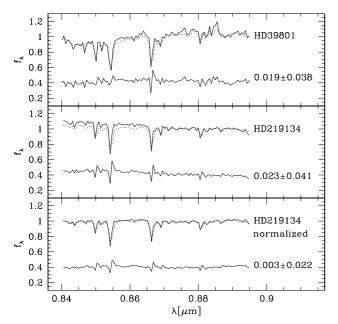


Fig. 23. Direct comparison of the spectra of Betelgeuse (HD 39801, *top panel*) and HD 219134 (*middle panel*) from the libraries of Rayner et al. (2009, solid lines) and Cenarro et al. (2001a, 2002, dotted lines). In the *bottom panel* the continuum normalized spectra of HD 219134 are shown. In each panel the difference of the two corresponding spectra are plotted (shifted by +0.4 for displaying purposes) and the rms value is given.

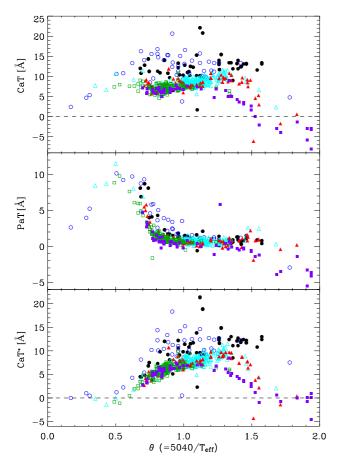


Fig. 24. EW of the CaT (top panel), PaT (middle panel), and CaT* index (bottom panel) as a function of $\theta = 5040/T_{\rm eff}$ for the stars of the IRTF sample (filled symbols) and the sample by Cenarro et al. (2002, open symbols) sample. The different symbols correspond to the supergiant (stars), giant (triangles), and dwarf stars (squares), respectively.

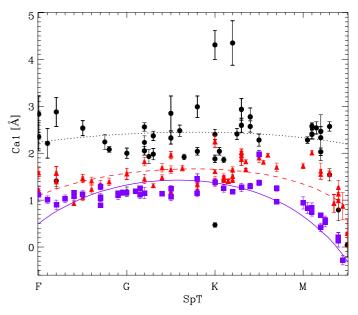


Fig. 25. Equivalent width of the Ca1 index as a function of spectral type for the observed and model spectrum of the sample stars. The different symbols correspond to the measurements in the observed spectra of the supergiant (circles), giant (triangles), and dwarf stars (squares), respectively. The different lines are for measurements in the model spectrum of of the supergiant (dotted line), giant (dashed line), and dwarf stars (solid line), respectively.

as: $CaT^* = CaT - 0.93 \times PaT$, where the CaT = Ca1 + Ca2 + Ca3 and PaT = Pa1 + Pa2 + Pa3 indices measure the strength of the raw CaT and three Paschen lines. We limited the comparison to stars within the same metallicity range as in Cenarro et al. (2002). Different luminosity classes were analyzed separately. The comparison of the samples shows excellent agreement and similar scatter (Fig. 24).

Sensitivity map comparison. We tested if the library limited number of stars and systematic errors in the stellar parameters can affect the sensitivity maps. This has been done by deriving the sensitivity maps with the Cenarro et al. (2001a, 2002) library. The maps are remarkably similar.

We demonstrated that the study of derivatives yields qualitatively consistent results with respect to the direct investigation of the behaviour of line-strength indices. However, it is not possible to predict the strength of a gradient and the intrinsic scatter of the correlations. The latter is due to the smoothing applied to the data using a second-degree polynomial as fitting function in the derivation of the model spectrum. In fact, the introduction of the fitting function was necessary to fill gaps in the parametric space due to irregular sampling and to smooth the data before perform its derivative.

To address this problem we measured the EWs on both the model and observed spectra. Let us consider, as an example, the case of Ca1 dependence on SpT. The EWs of the observed spectra are characterized by a significant scatter which is different for the three luminosity classes, whereas the EWs measured on the model spectrum change smoothly with SpT (Fig. 25). This means that, although for each wavelength an independent fit is applied along the SpT direction, the different fits maintain a degree of coherence along the wavelength direction such that a smooth trend of the EW is observed when it is measured on the model spectrum. On the other hand the EWs of the model spectrum do follow the trend measured on real data with different degrees of goodness. In this respect, the choice of the degree

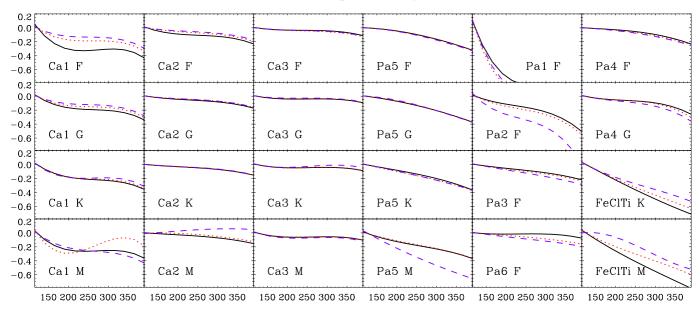


Fig. 26. *I*-band indices sensitivity to the velocity dispersion broadening. The indices relative variations for different luminosity classes and spectral types are shown: supergiants (solid line), giants (dashed line) and dwarfs (dotted line).

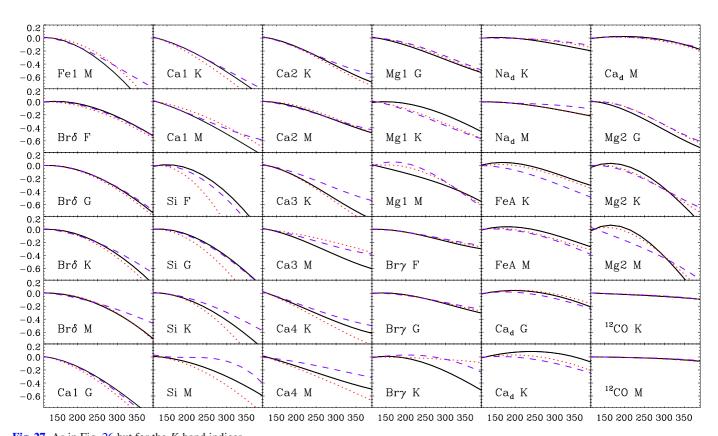


Fig. 27. As in Fig. 26 but for the *K*-band indices.

of the fitting function plays a relevant role. The introduction of a higher-order polynomial generally allows a better description of the data but the noise on the derivative increases severely limiting its effectiveness in defining indices. Our conclusion is that a second-order polynomial leads to a sensitivity map that qualitatively predict the variation of a spectral feature with respect to a given physical parameter. Therefore the sensitivity map is useful as an objective tool to define the best bandpasses for the spectral feature and their adjacent continuum, whereas

the measurements on the real spectra yield more reliable correlations between the EWs and the stellar physical parameters. Last but not least, the index definitions derived from the sensitivity map depend on the spectral resolution of the library.

Sensitivity of indices to velocity dispersion broadening. A straightforward application of the indices system is the study of the unresolved stellar population in galaxies. It is therefore important to investigate how our measurements are affected by

the unavoidable velocity broadening due to the internal galaxy kinematics.

We broadened all the spectra by convolving them with a Gaussian of σ varying from 115 to 400 km s⁻¹, in steps of 25 km s⁻¹. The indices were measured for all the broadened spectra and a third-order polynomial fit was done to the relative changes of each index value as a function of velocity dispersion. Figures 26 and 27 illustrate the $\Delta(Index)/Index(\sigma_0)$ values for the different luminosity classes and spectral types. It has to be noted that some indices have a low EW value, typically <1 Å for some of the spectral types considered. The EW of these indices shows a large variation even for small values of σ and goes to zero for $\sigma > 200-300 \text{ km s}^{-1}$. The faint indices are not shown in Fig. 26 and 27. The analysis shows that there is a very strong correlation between the line strength and the effect of the broadening, being the strongest lines less affected. This is true in both bands. The K-band indices are generally more affected by broadening due to the intrinsic faintness of the lines.

This work extended the previously defined sensitivity indices (Worthey et al. 1994) to a broader concept of sensitivity map and tested if they allows us to identify the spectral features that can be used as a proxy for $T_{\rm eff}$, $\log{(g)}$, and [Fe/H]. This empirical method is applied to I- and K-band absorption features for the star spectra of the IRTF library (Cushing et al. 2005; Rayner et al. 2009). The main results are:

- sensitivity map allows to fine tune the best definition for the line-strength indices (i.e., the bandpass limits for the line and nearby continuum);
- sensitivity map reliably predicts the variation of a spectral feature with respect to a given physical parameter but not its absolute strength;
- spectral line blends are obvious on sensitivity map when the blended features are characterized by a different behavior with respect to some physical stellar parameters;
- the EWs of new indices were measured for the IRTF star sample, and they will be useful for stellar population synthesis models and can be used to get element-by-element abundances for unresolved stellar population studies in galaxies:
- a systematic search for reliable T_{eff}, log (g), and [Fe/H] for the IRTF sample stars was carried out and the available physical parameters are reported.

We develop a fast and efficient method to identify those features that are sensible to different physical stellar parameters. The method is free from any assumption and it can be applied to any star or line-strength index. In a forthcoming paper we will extend such an analysis to Y, J, H, and L bands to define new indices. A straightforward next step is the extensive use of the sensitivity map in the upcoming X-Shooter Spectral Library (XLS, Chen et al. 2011).

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Tables 3–7 that follow are an important part of the project, but they are placed here because they are large.

Table 3. Physical parameters of the stars of the IRTF library.

(1) (2) (3) (4) (5) (6) (7) (8) (9) (10) (10) (10007927) (1) 20 04.9 +58 13 5.3 8 F0Ia 5.01 7341 -	Star IDs	RA	Dec	Sp. Type	M_V	$T_{ m eff}$	$\log(g)$	[Fe/H]	Parallax	Ref.
HD007927 01 20 04.9 +58 13 53.8 F0la 5.01 7341 -		(J2000)	(J2000)			K				
HD 135153 1514 373	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
HD006130	HD 007927	01 20 04.9	+58 13 53.8	F0Ia	5.01	7341		_	1.40	
HD089025 10 16 41 4 23 25 02.3 FOIII 3.44 6950 2.95 -0.03 12.56 2 HD013174 02.09 25.3 +25.56 23.6 F2III 4.98 7000 3.7 0.3 10.19 4 HD027397 04 19 57.7 +14 02 06.7 FOIV 5.56 7100 4.3 - 22.31 5 HD108519 12 746.3 +27 25 21.9 FOVn - 7200* - 6.00 - 6.00 - 6.00 - 6.00 - 7200* - 6.00 - 6.00 - 7200* - 7200*	HD 135153		$-31\ 31\ 08.8$	F1II	4.92	7070	1.87	_		2
HD013174	HD 006130	01 03 37.0	+61 04 29.3	FOII	5.91	7400		0.02	2.16	3
HD 1027397 04 19 57.7 +14 02 06.7 FOIV 5.56 7100 4.3 - 22.31 5 HD 108519 12 27 46.3 +27 25 21.9 FOVN - 7200* - - - - 6.90 - HD 173638 18 46 43.3 -10 07 30.1 F2Ib-II 5.73 7090 1.67 - - - - - 2 HD 213135 22 29 46.0 - - 27 06 26.2 FIV 4.19 6918 - - - - - -	HD 089025	10 16 41.4	+23 25 02.3	FOIII	3.44	6950	2.95	-0.03	12.56	
HD 108519	HD 013174	02 09 25.3	+25 56 23.6		4.98	7000	3.7	0.3	10.19	
HD 173638	HD 027397	04 19 57.7	+14 02 06.7	F0IV	5.56	7100	4.3	_	22.31	5
HD 13135 22 29 46.0 -27 06 26.2 FIV 4.19 6918 - -0.24 23.76 6 6 BD+38-2803 16 35 57.3 +37 58 02.1 F2-F5Ib - 7350* - - -0.15 - -0.15 - HD 182835 19 26 31.1 +00 20 18.8 F2Iab 4.68 6810 1.93 -0.08 0.28 2 HD 040535 05 59 01.1 -09 22 56.0 F2III-IV - 6870* - - - 0.32 - - HD 164136 17 58 30.1 +30 11 21.4 F2II 4.11 6575 2.00 -0.33 4.10 7 HD 113139 13 00 43.7 +56 21 58.8 F2V 4.93 6890 4.13 0.02 40.06 2 HD 020615 04 07 42.0 +15 09 46.0 F3V 6.05 6880 4.28 0.09 21.27 2 4 4 4 4 4 4 4 4 4	HD 108519	12 27 46.3	+27 25 21.9		_	7200*	_	_	6.90	
BD+38-2803	HD 173638	18 46 43.3	-100730.1	F2Ib-II	5.73	7090	1.67	-0.10	-0.07	2
HD 182835 19 26 3 1.1 -00 20 18.8 F21ab 4.68 6810 1.93 -0.08 0.28 2 HD 1040535 05 59 01.1 -09 22 56.0 F2III-IV - 6870* - 10.32 - HD 164136 17 58 30.1 +30 11 21.4 F2II 4.41 6575 2.00 -0.33 4.10 7 HD 113139 13 00 43.7 +56 21 58.8 F2V 4.93 6890 4.13 0.02 40.06 2 HD 206015 04 07 42.0 +15 09 46.0 F3V 6.05 6880 4.28 0.09 21.27 2 HD 201770 03 32 26.2 +46 03 24.7 F4III 5.31 6615 4.06 0.03 27.46 8 HD 1016232 02 36 57.7 +24 38 53.0 F4V 7.09 6462 4.50 0.27 23.36 9 HD 878822 10 08 15.9 +31 36 14.5 F4V 6.23 6545 4.20 0.19 15.80 10 HD 1075555 08 52 21.8 +44 53 51.4 F5 8.11 6490 3.57 0.02 3.61 2 HD 213306 22 29 10.3 +58 24 54.7 F5Ib 4.07 5864 1.65 0.07 3.32 11 HD 17918 02 53 11.7 +16 29 00.4 F2III 5.07 6780 3.33 0.26 20.51 2 HD 218804 23 10 27.2 +43 32 39.2 F5IV 6.00 6222 4.00 -0.27 35.39 10 HD 1075254 04 21 31.6 +21 02 23.5 F5V 6.77 6519 0.06 3.97 19.55 12 HD 10160365 17 38 57.8 +13 19 45.3 F6II 6.14 6180 3.05 -0.26 10.68 2 HD 124850 14 16 00.9 -06 000 0.19 F7IV 4.20 6000 4.10 -0.28 61.54 13 HD 201078 21 06 30.2 +31 11 04.7 F7.5Ib-IIV 5.82 6230 1.94 -0.10 1.72 2 HD 124850 14 16 00.9 -06 000 0.19 F7IV 4.10 6338 4.29 -0.05 68.63 8 HD 102870 11 50 41.7 +01 45 52.9 F9V 3.61 6146 4.30 0.20 91.74 15 HD 19963 23 52 28 823 41.4.7 F8IV 4.40 5920 3.18 -0.26 18.83 2 HD 114710 13 11 52.4 +27 52 41.4 F9.5V 4.26 6146 4.52 0.06 109.23 17 HD 105908 18 701.6 +32 54 04.5 60V+kiV 5.56 5730 2.71 -0.35 8.09 2 HD 1160918 19 36 52.4 +111 623.5 60Ib 5.98 5700 2.24 -0.17 2.66 6.76 2 HD 105908 18 070.5 5.30 33 34.7	HD 213135	22 29 46.0	$-27\ 06\ 26.2$	F1V	4.19	6918	_	-0.24	23.76	6
HD 040535 05 59 01.1 -09 22 56.0 F2III-IV - 6870* - - 10.32 -	BD+38-2803	16 35 57.3	+37 58 02.1	F2-F5Ib	_	7350*		_	-0.15	
HD 164136		19 26 31.1		F2Iab	4.68	6810	1.93	-0.08	0.28	
HD 113139	HD 040535	05 59 01.1	$-09\ 22\ 56.0$		_	6870*		_	10.32	
HD 026015 04 07 42.0 +15 09 46.0 F3V 6.05 6880 4.28 0.09 21.27 2 HD 021770 03 32 26.2 +46 03 24.7 F4III 5.31 6615 4.06 0.03 27.46 8 HD 016232 02 36 57.7 +24 38 53.0 F4V 7.09 6462 4.50 0.27 23.36 9 HD 087822 10 08 15.9 +31 36 14.5 F4V 6.23 6545 4.20 0.19 15.80 10 HD 075555 08 52 21.8 +44 53 51.4 F5 8.11 6490 3.57 0.02 3.61 2 HD 213306 22 29 10.3 +58 24 54.7 F51b 4.07 5864 1.65 0.07 3.32 11 HD 017918 02 53 11.7 +16 29 00.4 F2III 6.30 6700 4.0 0.0 9.27 4 HD 186155 19 40 50.2 +45 31 29.7 F51I-III 5.07 6780 3.33 0.26 20.51 2 HD 218804 23 10 27.2 +43 32 39.2 F5IV 6.00 6222 4.00 -0.27 35.39 10 HD 027524 04 21 31.6 +21 02 23.5 F5V 6.77 6519 0.06 3.97 19.55 12 HD 160365 17 38 57.8 +13 19 45.3 F6III 6.14 6180 3.05 -0.26 10.68 2 HD 1143 015 30.49 +29 34 43.7 F6IV 3.41 6288 3.91 0.00 50.87 8 HD 215648 22 46 41.6 +12 10 22.4 F7V 4.20 6000 4.10 -0.28 61.54 13 HD 201078 21 06 30.2 +31 11 0.47 F7.51b-IIV 5.82 6230 1.94 -0.10 1.72 2 HD 124850 14 16 0.09 -06 00 01.9 F7IV 4.10 6338 4.29 -0.05 68.63 8 HD 102870 11 50 41.7 +01 45 52.9 F9V 3.61 6146 4.30 0.20 91.74 15 HD 102870 15 041.7 +01 45 52.9 F9V 3.61 6146 4.30 0.20 91.74 15 HD 102667 23 25 22.8 +23 24 14.7 F8IV 4.40 5920 3.18 -0.26 18.83 2 HD 116233 20 149.6 +14 58 58.7 G01a 4.2 5900 0.1 1.23 16 HD 107383 04 19 54.9 +16 31 21.3 F9V 6.86 6280 4.55 0.10 23.27 2 HD 114710 13 11 52.4 +27 52 41.4 F9.5V 4.26 6100 4.50 0.07 19.46 15 HD 109358 12 33 44.5 +41 12 16.9 G0V 4.26 6000 4.50 0.07 119.46 15 HD 109358 12 33 44.5 +41 12 16.9 G0V 4.26 6000 4.50 0.07 119.46 15 HD 109358 12 33 44.5 +41 12 16.9 G0V 4.26 6000 4.50	HD 164136	17 58 30.1	+30 11 21.4	F2II	4.41	6575	2.00	-0.33	4.10	
HD 021770 03 32 26.2	HD 113139	13 00 43.7	+56 21 58.8	F2V	4.93	6890	4.13	0.02	40.06	
HD 016232	HD 026015	04 07 42.0	+15 09 46.0	F3V	6.05	6880	4.28	0.09	21.27	
HD 087822	HD 021770	03 32 26.2	+46 03 24.7	F4III	5.31	6615	4.06	0.03	27.46	8
HD 075555 08 52 21.8	HD 016232	02 36 57.7	+24 38 53.0	F4V	7.09	6462	4.50	0.27	23.36	9
HD 213306 22 29 10.3 +58 24 54.7 F51b 4.07 5864 1.65 0.07 3.32 11 HD 017918 02 53 11.7 +16 29 00.4 F2III 6.30 6700 4.0 0.0 9.27 4 HD 186155 19 40 50.2 +45 31 29.7 F5II-III 5.07 6780 3.33 0.26 20.51 2 HD 218804 23 10 27.2 +43 32 39.2 F5IV 6.00 6222 4.00 -0.27 35.39 10 HD 027524 04 21 31.6 +21 02 23.5 F5V 6.77 6519 0.06 3.97 19.55 12 HD 160365 17 38 57.8 +13 19 45.3 F6III 6.14 6180 3.05 -0.26 10.68 2 HD 011443 01 53 04.9 +29 34 43.7 F6IV 3.41 6288 3.91 0.00 50.87 8 HD 215648 22 46 41.6 +12 10 22.4 F7V 4.20 6000 4.10 -0.28 61.54 13 HD 201078 21 06 30.2 +31 11 04.7 F7.51b-IIV 5.82 6230 1.94 -0.10 1.72 2 HD 124850 14 16 00.9 -06 00 01.9 F7IV 4.10 6222 4.20 -0.09 46.74 14 HD 126660 14 25 11.8 +51 51 02.6 F7V 4.10 6338 4.29 -0.05 68.63 8 HD 102870 11 50 41.7 +01 45 52.9 F9V 3.61 6146 4.30 0.20 91.74 15 HD 190323 20 03 49.6 +14 58 58.7 G0Ia 4.2 5900 0.1 - -1.23 16 HD 051956 06 59 31.7 +00 55 00.3 F8Ib - 6100* - - - 1.37 - HD 219623 23 16 42.3 +53 12 48.5 F7V 5.60 6145 4.23 0.01 49.31 8 HD 11844 12 51 54.4 +19 10 05.2 F8IV 7.80 6160 3.67 0.02 -12.61 2 HD 219623 23 16 42.3 +53 12 48.5 F7V 5.60 6145 4.23 0.01 49.31 8 HD 027383 04 19 54.9 +16 31 21.3 F9V 6.86 6280 4.55 0.10 23.27 2 HD 114710 13 11 52.4 +27 52 41.4 F9.5V 4.26 6146 4.52 0.06 109.23 17 HD 006903 01 09 49.2 +19 39 30.2 G0III 5.56 5730 2.71 -0.35 8.09 2 HD 176051 18 57 01.6 +32 54 04.5 G0V + kIV 5.25 6030 4.37 -0.04 66.76 2 HD 109358 12 33 44.5 +41 21 26.9 G0V 4.26 6000 4.50 0.07 119.46 15	HD 087822	10 08 15.9		F4V	6.23	6545	4.20	0.19	15.80	10
HD 017918 02 53 11.7 +16 29 00.4 F2III 6.30 6700 4.0 0.0 9.27 4 HD 186155 19 40 50.2 +45 31 29.7 F5II-III 5.07 6780 3.33 0.26 20.51 2 4 4 2 2 2 2 4 2 3 2 2 2 4 3 2 3 2 2 5 5 5 5 6.00 6222 4.00 -0.27 35.39 10 4 4 0 10 2 2 2 2 4 0 0 -0.27 3 5 3 10 4 10 2 2 2 2 2 4 0 0 -0.27 3 5 3 10 4 10 2 2 2 2 4 0 0 -0.27 3 5 3 10 4 2 4 4	HD 075555	08 52 21.8	+44 53 51.4	F5	8.11	6490	3.57	0.02	3.61	2
HD 186155	HD 213306	22 29 10.3	+58 24 54.7	F5Ib	4.07	5864	1.65	0.07		11
HD 218804 23 10 27.2 +43 32 39.2 F5IV 6.00 6222 4.00 -0.27 35.39 10 HD 027524 04 21 31.6 +21 02 23.5 F5V 6.77 6519 0.06 3.97 19.55 12 HD 160365 17 38 57.8 +13 19 45.3 F6III 6.14 6180 3.05 -0.26 10.68 2 HD 011443 01 53 04.9 +29 34 43.7 F6IV 3.41 6288 3.91 0.00 50.87 8 HD 215648 22 46 41.6 +12 10 22.4 F7V 4.20 6000 4.10 -0.28 61.54 13 HD 201078 21 06 30.2 +31 11 04.7 F7.5lb-IIV 5.82 6230 1.94 -0.10 1.72 2 HD 124850 14 16 00.9 -06 00 01.9 F7IV 4.10 6222 4.20 -0.09 46.74 14 HD 126660 14 25 11.8 +51 51 02.6 F7V 4.10 6338 4.29 -0.05 68.63 8 HD 102870 11 50 41.7 +01 45 52.9 F9V 3.61 6146 4.30 0.20 91.74 15 HD 190323 20 03 49.6 +14 58 58.7 G0Ia 4.2 5900 0.11.23 16 HD 051956 06 59 31.7 +00 55 00.3 F8Ib - 6100* 1.37 - HD 220657 23 25 22.8 +23 24 14.7 F8IV 4.40 5920 3.18 -0.26 18.83 2 HD 11844 12 51 54.4 +19 10 05.2 F8IV 7.80 6160 3.67 0.02 -12.61 2 HD 219623 23 16 42.3 +53 12 48.5 F7V 5.60 6145 4.23 0.01 49.31 8 HD 027383 04 19 54.9 +16 31 21.3 F9V 6.86 6280 4.55 0.10 23.27 2 HD 114710 13 11 52.4 +27 52 41.4 F9.5V 4.26 6146 4.52 0.06 109.23 17 HD 006903 01 09 49.2 +19 39 30.2 G0III 5.56 5730 2.71 -0.35 8.09 2 HD 176051 18 57 01.6 +32 54 04.5 G0Ib 5.98 5700 2.24 -0.17 2.86 2 HD 185018 19 36 52.4 +11 16 23.5 G0Ib 5.98 5700 2.24 -0.17 2.86 2 HD 19358 12 33 44.5 +41 126.9 G0V 4.26 6000 4.50 0.07 119.46 15 HD 02018 03 23 39.0 +04 52 55.5 G5III 6.40 5250 3.0 0.0 0.0 2.92 21 HD 170018 03 23 39.0 +04 52 55.5 G5III 6.40 5250 3.0 0.0 0.0 2.92 21 HD 101807 01 41 47.1 +42 36 48.1 G1.5V 4.90 5898 4.31 -0.02 79.09 18	HD 017918	02 53 11.7	+16 29 00.4	F2III	6.30	6700	4.0	0.0	9.27	4
HD 027524	HD 186155	19 40 50.2	+45 31 29.7	F5II-III	5.07	6780	3.33	0.26	20.51	2
HD 160365	HD 218804	23 10 27.2		F5IV	6.00	6222	4.00	-0.27	35.39	10
HD 011443	HD 027524	04 21 31.6	+21 02 23.5	F5V	6.77	6519		3.97	19.55	
HD 215648 22 46 41.6 +12 10 22.4 F7V 4.20 6000 4.10 -0.28 61.54 13 HD 201078 21 06 30.2 +31 11 04.7 F7.51b-IIv 5.82 6230 1.94 -0.10 1.72 2 HD 124850 14 16 00.9 -06 00 01.9 F7IV 4.10 6222 4.20 -0.09 46.74 14 HD 126660 14 25 11.8 +51 51 02.6 F7V 4.10 6338 4.29 -0.05 68.63 8 HD 102870 11 50 41.7 +01 45 52.9 F9V 3.61 6146 4.30 0.20 91.74 15 HD 190323 20 03 49.6 +14 58 58.7 G0Ia 4.2 5900 0.1 - -1.23 16 HD 051956 06 59 31.7 +00 55 00.3 F8Ib - 6100* - - 1.37 - HD 220657 23 25 22.8 +23 24 14.7 F8IV 4.40 5920 3.18 -0.26 18.83 2 HD 111844 12 51 54.4 +19 10 05.2 F8IV 7.80 6160 3.67 0.02 -12.61 2 HD 219623 23 16 42.3 +53 12 48.5 F7V 5.60 6145 4.23 0.01 49.31 8 HD 027383 04 19 54.9 +16 31 21.3 F9V 6.86 6280 4.55 0.10 23.27 2 HD 114710 13 11 52.4 +27 52 41.4 F9.5V 4.26 6146 4.52 0.06 109.23 17 HD 006903 01 09 49.2 +19 39 30.2 G0III 5.56 5730 2.71 -0.35 8.09 2 HD 165908 18 07 01.5 +30 33 43.7 F7V 5.06 6020 4.48 -0.56 63.88 18 HD 185018 19 36 52.4 +11 16 23.5 G0Ib 5.98 5700 2.24 -0.17 2.86 2 HD 109358 12 33 44.5 +41 21 26.9 G0V 4.26 6000 4.50 0.07 119.46 15 HD 020619 03 19 01.9 -02 50 35.5 G1.5V 7.10 5600 4.00 -0.45 40.72 20 HD 074395 08 43 40.4 -07 14 01.4 G2Iab: 4.64 5250 3.0 0.0 2.92 21 HD 216219 22 50 52.1 +18 00 07.5 G0IIp 7.44 5478 2.80 -0.55 10.74 22 HD 095128 10 59 28.0 +40 25 48.9 G1V 5.10 5882 4.34 0.01 71.04 18 HD 010307 01 41 47.1 +42 36 48.1 G1.5V 4.90 5898 4.31 -0.02 79.09 18	HD 160365		+13 19 45.3	F6III	6.14	6180		-0.26	10.68	
HD 201078										
HD 124850	HD 215648	22 46 41.6	+12 10 22.4						61.54	
HD 126660	HD 201078		+31 11 04.7					-0.10		
HD 102870	HD 124850	14 16 00.9	$-06\ 00\ 01.9$		4.10			-0.09	46.74	
HD 190323	HD 126660									
HD 051956	HD 102870							0.20		
HD 220657			+14 58 58.7		4.2		0.1	_		16
HD 111844			+00 55 00.3		_			_		
HD 219623										2
HD 027383	HD 111844									
HD 114710										
HD 006903										
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HD 074395 08 43 40.4 -07 14 01.4 G2Iab: 4.64 5250 1.30 -0.11 5.40 18 HD 021018 03 23 39.0 +04 52 55.5 G5III 6.40 5250 3.0 0.0 2.92 21 HD 216219 22 50 52.1 +18 00 07.5 G0IIp 7.44 5478 2.80 -0.55 10.74 22 HD 095128 10 59 28.0 +40 25 48.9 G1V 5.10 5882 4.34 0.01 71.04 18 HD 010307 01 41 47.1 +42 36 48.1 G1.5V 4.90 5898 4.31 -0.02 79.09 18										
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	-									18

Notes. The temperatures with asterisks are obtained with the luminosity-temperature relation of Carroll & Ostlie (1996).

References. The Ref. column give references for the literature values of $T_{\rm eff}$, $\log(g)$ and [Fe/H]: (1) Kovtyukh (2007); (2) Gray et al. (2001); (3) Venn (1995); (4) Jasniewicz et al. (2006); (5) Soubiran et al. (2010); (6) Holmberg et al. (2008); (7) Luck & Wepfer (1995); (8) Balachandran (1990); (9) Boesgaard & Friel (1990); (10) Boesgaard & Tripicco (1986); (11) Cenarro et al. (2007); (12) Arellano Ferro (2010); (13) Boesgaard & Lavery (1986); (14) Edvardsson et al. (1984); (15) Gehren (1981); (16) Andrievsky et al. (2002); (17) Cornide & Rego (1984); (18) Edvardsson et al. (1993); (19) Luck & Bond (1980); (20) Barbuy & Erdelyi-Mendes (1989); (21) Lèbre et al. (2009); (22) Krishnaswamy & Sneden (1985); (23) Mallik (1998); (24) Cayrel de Strobel et al. (1981); (25) Lyubimkov et al. (2010); (26) Santos et al. (2004); (27) Hekker & Meléndez (2007); (28) Kipper (2008); (29) Takeda et al. (2008); (30) Castro et al. (1999); (31) McWilliam (1990); (32) Gratton & Sneden (1991); (33) Lambert & Ries (1981); (34) Spite & Spite (1980); (35) Fuhrmann (1998); (36) Goss et al. (1982); (37) Luck & Heiter (2007); (38) Gonzalez et al. (1999); (39) Cottrell & Sneden (1986); (40) Randich et al. (1999); (41) Perrin (1983); (42) Soubiran et al. (2008); (43) Malkan et al. (2002); (44) Luck (1982); (45) Levesque et al. (2005); (46) Tomkin & Lambert (1999); (47) Strassmeier & Schordan (2000); (48) Carr et al. (2000); (49) Zboril & Byrne (1998); (50) Smith & Lambert (1986); (51) Mirtorabi et al. (2003); (52) Lançon et al. (2007); (53) Kučinskas et al. (2005); (54) Smith & Lambert (1990).

Table 3. continued.

Star IDs	RA	Dec	Sp. Type	M_V	$T_{ m eff}$	$\log(g)$	[Fe/H]	Parallax	Ref.
	(J2000)	(J2000)		mag	K			mas	
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
HD 039949	05 57 05.5	+27 18 59.9	G2Ib	7.25	5250	1.10	-0.16	-0.83	19
HD 042454	06 12 05.5	+29 29 31.7	G2Ib	7.32	5250	1.10	-0.05	3.20	19
HD 003421	00 37 21.2	+35 23 58.2	G2.5IIa	5.43	5620	2.78	-0.14	3.19	2
HD 219477	23 15 46.3	+28 14 52.4	G2II-III	-	5450*	-	- 0.02	3.41	-
HD 126868	14 28 12.1	-02 13 40.6	G2IV	4.84	5600	3.90	-0.02	24.15	23
HD 076151	08 54 17.9	-05 26 04.0	G2V	6.00	5600 5000*	4.40	-0.02	58.50	24
HD 192713 HD 176123	20 15 30.2 18 59 26.8	+23 30 32.0 -18 33 59.1	G3Ib-II G3II	- 6.39	5000* 5200	2.25	_	0.75 3.93	25
HD 088639	10 13 49.7	+27 08 08.9	G3III	0.39	5300*	_	_	8.39	_
HD 010697	01 44 55.8	+20 04 59.3	G5IV	6.29	5641	4.05	0.14	30.71	26
HD 006474	01 06 59.7	+63 46 23.3	G4Ia	7.60	6222	1.50	0.25	-0.20	11
HD 108477	12 27 49.4	-16 37 54.6	G4III	_	5200*	_	-	2.55	_
HD 094481	10 54 17.8	-13 45 28.9	KOIII	5.65	5355	3.00	_	7.97	27
HD 179821	19 13 58.6	+00 07 31.9	G5Ia	8.12	6750	0.50	-0.45	0.18	28
HD 214850	22 40 52.7	+14 32 56.9	G4V	5.73	5420	_	-0.22	30.49	6
HD 190113	20 02 02.8	+35 38 28.0	G5Ib	_	4850*	_	_	1.28	_
HD 193896	20 23 00.8	-09 39 16.9	G5IIIa	_	5150*	_	_	7.49	_
HD 018474	02 59 49.8	+47 13 14.5	G4pIII	5.24	5013	2.38	-0.23	5.85	29
HD 165185	18 06 23.7	-36 01 11.2	G5V	5.94	5895	4.49	-0.06	57.58	30
HD 115617	13 18 24.3	$-18\ 18\ 40.3$	G5V	4.74	5600	4.00	-0.03	117.30	14
HD 161664	17 47 45.6	-22 28 40.0	G6Ib	_	4800*	_	_	1.80	_
HD 202314	21 14 10.3	+29 54 03.4	G2Ib	6.18	4900	1.50	0.00	1.34	16
HD 058367	07 25 38.9	+09 16 33.9	G6.5IIb	4.99	4820	2.01	-0.22	3.30	31
HD 027277	04 20 53.5	+50 16 18.2	G6III	_	5100*	_	_	_	_
HD 016139	02 36 18.0	+27 28 20.3	G7.5IIIa	_	4950*	_	_	_	_
HD 333385 HD 025877	20 02 27.4	+30 04 25.5	G7Ia		4650*	- 1.91	_	2.04	25
HD 182694	04 09 27.6 19 23 56.5	+59 54 29.0 +43 23 17.4	G8IIa G6.3IIIa	6.29 5.85	5060 5067	2.63	-0.04	3.04 8.06	23 29
HD 114946	13 14 10.9	-19 55 51.4	G8III/IV	5.33	5056	3.04	-0.04	25.89	32
HD 020618	03 19 55.8	+27 04 16.1	G8IV	5.91	5049	3.08	-0.21	15.88	29
HD 208606	21 55 20.6	+61 32 30.5	G8Ib	J.91 —	4600*	-	-0.22	0.36	
HD 104979	12 05 12.5	+08 43 58.7	G8IIIa	4.13	5250	3.25	-0.29	19.08	33
HD 135722	15 15 30.2	+33 18 53.4	G8III	3.47	4800	2.70	-0.50	27.94	2
HD 122563	14 02 31.8	+09 41 09.9	KIIvw	6.20	4582	0.90	-2.50	3.76	34
HD 101501	11 41 03.0	+34 12 05.9	G8V	5.32	5538	4.69	0.03	104.81	17
HD 075732	08 52 35.8	+28 19 50.9	G8V	5.95	5336	4.47	0.40	79.80	35
HD 170820	18 32 13.1	$-19\ 07\ 26.3$	K0III	7.36	5663	2.00	-0.05	2.34	3
HD 222093	23 37 39.6	-13 03 36.9	KOIII	5.66	4730	2.78	-0.25	11.50	31
HD 164349	18 00 03.4	+16 45 03.3	K0.5IIb	4.67	4383	1.80	-0.22	4.97	36
HD 009852	01 37 51.2	+61 51 41.7	K0.5III	_	4750*	_	_	_	_
HD 165782	18 08 26.5	$-18\ 33\ 07.9$	K0Ia	_	4420*	_	_	2.11	_
HD 044391	06 22 47.9	+27 59 12.0	K0Ib	7.68	4710	0.60	0.21	1.39	18
HD 179870	19 13 53.6	+09 01 59.6	KOII	_ 5.55	4705	-	- 0.12	2.62	-
HD 100006	11 30 29.0	+18 24 35.2	KOIII	5.55	4785	2.67	-0.12	10.03	37
HD 145675	16 10 24.3	+43 49 03.5	K0V	6.67	5300	4.27	0.50	55.11	38
HD 124897 HD 063302	14 15 39.7 07 47 38.5	+19 10 56.6	K1.5III	-0.04 6.33	4500 4500	2.01 0.20	-0.56 0.17	88.85 1.43	33 18
HD 003302 HD 091810	10 37 20.5	-15 59 26.4 +56 25 52.8	K3Iab/Ib K1-IIIb		4600*			5.86	10
HD 036134	05 29 23.7	-03 26 47.0	K1-III0 K1-III	_	4600	_	_	6.98	_
HD 025975	03 29 23.7	+37 43 38.9	K1-III K1III	6.09	4941	3.40	-0.20	22.66	39
HD 142091	15 51 13.9	+35 39 26.5	K1IIVa	4.82	4800	3.37	-0.04	32.13	31
HD 165438	18 06 15.2	-04 45 04.5	K1IV	5.74	4862	3.40	0.02	28.31	40
HD 010476	01 42 29.8	+20 16 06.6	K1V	5.20	5196	4.50	-0.20	133.91	41
HD 023082	03 44 05.8	+44 53 04.9	K2.5II	_	_	_	_	1.16	_
HD 002901	00 32 47.5	+54 07 11.8	K2III	_	4420*	_	_	3.56	_
HD 132935	15 02 04.2	$-08\ 20\ 40.9$	K2III	_	4420*	_	_	4.16	_
HD 137759	15 24 55.8	+58 57 57.8	K2III	3.31	4490	2.74	0.03	31.92	31
HD 212466	22 23 07.0	+55 57 47.6	K2O-Ia	_	_	_	_	1.19	_
HD 003765	00 40 49.3	+40 11 13.8	K2V	6.15	5067	4.45	0.10	57.90	42
HD 114960	13 13 57.6	+01 27 23.2	K3.5IIIb		4000*	_	_	6.65	_
HD 187238	19 48 11.8	+22 45 46.3	K3Ia0-Ia	7.05	4500	0.80	0.20	2.16	18
HD 099998	11 30 18.9	-03 00 12.6	K3.5III	4.77	3920	1.67	-0.39	5.40	31
HD 035620	05 27 38.9	+34 28 33.2	K3IIICN+	5.07	4200	2.15	0.11	8.14	31

Table 3. continued.

Star IDs	RA	Dec	Sp. Type	M_V	$T_{ m eff}$	$\log(g)$	[Fe/H]	Parallax	Ref.
	(J2000)	(J2000)	1 31	mag	K	C (0)		mas	
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
HD 178208	19 05 09.8	+49 55 23.4	K3III	_	4200*	-	-	5.59	
HD 221246	23 30 07.4	+49 07 59.3	K3III	_	4200*	_	_	3.81	_
HD 016068	02 36 52.8	+55 54 55.4	K2	6.29	6427	_	-0.31	_	6
HD 219134	23 13 17.0	+57 10 06.1	K3Vvar	6.46	4913	4.51	0.08	153.24	42
HD 185622	19 39 25.3	+16 34 16.0	K4Ib	6.38	3990	_	_	1.42	43
HD 201065	21 05 35.8	+46 57 47.7	K4Ib-II	-	3950*	_	_	1.20	_
HD 207991	21 51 55.4	+48 26 13.6	K5Ib	6.88	3777	_	_	2.99	1
HD 045977	06 30 07.3	-11 48 32.1	K4V	-	4590*	_	-	35.00	_
HD 120477	13 49 28.6	+15 47 52.4	K5.5III	4.05	3890	1.55	-0.23	13.29	31
HD 216946	22 56 26.0	+49 44 00.7	K5Iab	5.00	4000	0.50	-0.03	1.74	44
HD 181596	19 18 30.1	+50 13 39.4	K5III	_ 7.64	3950*	- 4.61	-	1.20	-
HD 036003	05 28 26.1	-03 29 58.4	K5V	7.64	4465	4.61	0.0	77.03	11
HD 003346	00 36 46.4	+44 29 18.9	K6IIIa	-	3900*	_	_	4.96	_ 45
HD 181475	19 20 48.3	-04 30 09.0	M1II	6.96	3700	_	_	0.65	45
HD 194193 HD 201092	20 22 45.3 21 06 55.3	+41 01 33.6 +38 44 31.4	K7III K7V	6.03	3850* 4120	- 4.40	-0.63	3.81 285.50	- 46
HD 201092 HD 237903	10 30 25.3	+55 59 56.8	K7V K7V	8.76	4070	4.40	-0.03 -0.18	91.00	11
HD 237903 HD 236697	01 19 53.6	+58 18 30.7	M0.5Ib	6.70 —	3600*	4.70 -	-0.18	1.93	-
HD 209290	22 02 10.3	+01 24 00.8	M0.5V	_	3800*	_	_	96.98	_
HD 209290 HD 213893	22 34 35.9	+01 24 00.8	MOIIIb	_	3800*	_	_	3.81	_
HD 019305	03 06 26.7	+00 55 42.6	M0V	_	3850*	_	_	67.69	_
IRAS14086-0703	14 11 17.6	-07 44 50.0	M10+III	_	3240*	_	_	07.09	_
HD 035601	05 27 10.2	+29 55 15.7	M1.5Ia0-Ia	7.32	4000	0.70	-0.24	1.06	18
BD+60-265	01 33 33.1	+61 33 31.0	M1.5Ib	-	3500*	-	-0.24	-	10
HD 036395	05 31 27.4	-03 40 38.0	M1V	- 7.92	3742	4.71	0.21	175.72	42
HD 014404	02 21 42.4	+57 51 46.1	M1-Iab-Ib	-	3550*	-	0.21	0.06	- -
HD 339034	19 50 11.9	+24 55 24.2	M1Ia	_	3550*	_	_	-	_
HD 204724	21 29 56.9	+23 38 19.8	M1III	4.53	3773	_	_	7.37	47
HD 039801	05 55 10.3	+07 24 25.4	M1Iab	0.58	3540	0.00	0.00	7.63	48
HD 042581	06 10 34.6	-21 51 52.7	M1V	_	3720*	_	-	173.19	_
HD 219734	23 17 44.6	+49 00 55.1	M2III	4.85	3730	0.90	0.27	4.98	11
Gl381	10 12 04.7	-02 41 05.0	M2.5V	_	3500*	_	_	81.23	_
Gl581	15 19 26.8	$-07\ 43\ 20.2$	M2.5V	_	3500*	_	_	159.52	_
HD 206936	21 43 30.5	+58 46 48.1	M2-Ia	_	3450*	_	_	0.62	_
HD 010465	01 43 11.1	+48 31 00.3	M2Ib	_	3450*	_	_	0.24	_
HD 023475	03 49 31.3	+65 31 33.5	M2II	_	3620*	_	_	3.38	_
HD 120052	13 47 25.4	-175135.4	M2III	5.44	3729	_	_	5.13	47
G1806	20 45 04.1	+44 29 56.6	M2V	_	3580*	_	_	80.01	_
HD 095735	11 03 20.2	+35 58 11.5	M2V	7.49	3620	4.90	-0.20	392.50	49
HD 014488	02 22 24.3	+57 06 34.4	M3.5Iab	_	3090*	_	_	_	_
HD 028487	04 29 38.9	+05 09 51.3	M3.5III	_	3580	_	_	4.28	_
G1273	07 27 24.5	+05 13 33.0	M3.5V	_	3420*	_	_	_	_
CD-31-49	07 41 02.6	$-31\ 40\ 59.1$	M3Iab-Ia	_	3550*	_	_	-0.53	_
HD 040239	05 59 56.1	+45 56 12.2	M3IIb	_	_	_	_	3.88	_
HD 039045	05 51 25.7	+32 07 28.9	M3III	_	3530*	_	_	5.62	_
HD 014469	02 22 06.9	+56 36 15.0	M3-M4Iab	_	3200*	_	_	_	_
RWCyg	20 28 50.6	+39 58 54.4	M3toM4Ia	_	3200*	_	_	1.28	_
Gl388	10 19 36.3	+19 52 12.0	M3V	_	3470*	_	_	_	_
HD 204585	21 28 59.8	+22 10 45.9	M4.5IIIa	_	3380*	_	_	5.41	_
Gl268AB	07 10 01.8	+38 31 46.0	M4.5V	_	3300*	_	_	157.24	_
HD 019058	03 05 10.6	+38 50 24.9	M4II	3.42	3500	0.80	-0.15	10.03	50
HD 214665	22 38 37.9	+56 47 44.2	M4+III	_	3430*	_	_	7.59	_
HD 027598	04 20 41.3	-16 49 47.9	M4-III	- 5.40	3430*	_	_	1.76	_
HD 004408	00 46 32.9	+15 28 31.8	M4IIIa	5.42	3522	_	_	5.55	51
Gl213	05 42 09.3	+12 29 21.6	M4V	_	3370*	_	_	172.78	_
Gl299	08 11 57.6	+08 46 22.0	M4V	- 5.00	3370*	1.0	_	143.00	- 52
HD 094705	10 56 01.5	+06 11 07.3	M5.5III	5.98	3300	1.0	_	10.03	52
HD 014386	02 19 20.8	-02 58 39.5	M5e-M9eIII	_	3330*	_	_	7.79	_
HD 156014	17 14 38.8	+14 23 25.2	M5Ib-II	- 4.20	2800*	- 0.50	0.14	8.53	_ 11
HD 175865	18 55 20.1 01 03 19.7	+43 56 45.9 +62 21 55 7	M5III M5V	4.20	3420 3240*	0.50	0.14	9.33 95.50	11
Gl51	01 03 19.7	+62 21 55.7	IVIJ V	_	3240"	_	_	93.30	_

Table 3. continued.

Star IDs	RA	Dec	Sp. Type	M_V	$T_{ m eff}$	$\log(g)$	[Fe/H]	Parallax	Ref.
	(J2000)	(J2000)		mag	K			mas	
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
Gl866	22 38 33.7	-15 17 57.3	M6	12.18	2747	5.09	_	300.00	11
GJ1111	08 29 49.3	+26 46 33.7	M6.5V	_	2990*	_	_	275.80	_
HD 069243	08 16 33.8	+11 43 34.4	M6e-M9eIII	- 5.00	3240*	-	_	-0.27	- 11
HD 018191	02 55 48.5	+18 19 53.9	M6III	5.80	3250	0.30	_	8.08	11
HD 196610	20 37 54.7	+18 16 06.8	M6III	6.23	3243	_	_	9.16	53
G1406	10 56 28.9	+07 00 52.8	M6V	_	3050*	_	_	_	_
BRI2339-0447	23 42 02.7	-04 31 05.0	M7-8III	- 0.24	3200*	_	_	- 5.60	-
HD 108849	12 30 21.0	+04 24 59.1	M7III:	8.24	2944	_	_	5.68	53
HD 207076	21 46 31.8	-02 12 45.9 +60 49 39.0	M8IIIv	6.78	2750	_	_	7.39	11
MY-Cep	22 54 31.7		M7-M7Iab	_	2600*	-	_	15450	_
Gl644C	16 55 35.3	-08 23 40.1	M7V	_	2940*	-	_	154.50	_
IRAS14303-1042	14 32 59.9	-10 56 03.6	M8	_	_	_	_	_	_
IRAS14436-0703	14 46 18.4	-07 15 49.8	M8	_	_	_	_	_	_
IRAS21284-0747	21 31 06.5	-07 34 20.5	M8	_	-	_	_	_	_
IRAS01037+1219	01 06 26.0	+12 35 53.0	M8III	_	3200*	_	_	164.20	_
Gl752B	19 16 57.6	+05 09 02.2	M8V	_	2640*	_	_	164.30	_
LP412-31	03 20 59.7	+18 54 23.3	M8V	_	2640*	_	_	68.90	_
BRIB0021-0214	00 24 24.6	-01 58 20.1	M9.5V	_	2600*	_	_	86.60	_
BRIB1219-1336	15 08 25.8	+09 36 18.2	M9III	_	3000*	_	_	_	_
IRAS15060+0947	12 21 52.5	-13 53 10.0	M9III	_	3000*	_	_	_	_
PJ1048-3956	10 48 14.6	-39 56 06.0	M9V	_	2600*	_	_	_	_
LHS2065	03 39 35.2	-35 25 44.0	M9V	_	2600*	_	_	117.30	_
LHS2924	08 53 36.2	-03 29 32.1	M9V	_	2600*	_	_	103.80	_
LP944-20	14 28 43.2	+33 10 39.1	M9V	_	2600*	_	_	201.20	_
HD 31996	04 59 36.3	-14 48 22.5	C7	_	_	_	_	3.99	_
HD 70138	08 19 43.1	-18 15 52.8	C	_	_	_	_	1.19	_
HD 57160	07 20 59.0	+24 59 58.0	C	_	_	_	_	1.83	_
HD 92055	10 37 33.3	-13 23 04.3	C	_	_	_	_	6.18	_
HD 76221	08 55 22.9	+17 13 52.5	C	_	_	_	_	0.90	_
HD 44984	06 25 28.2	+14 43 19.1	C	_	_	_	_	2.51	_
HD 48664	06 44 40.7	+03 18 58.6	C	_	_	_	_	_	_
HD 76846	08 59 48.9	+33 46 26.4	C	_	_	_	_	-0.68	_
HD 142143	15 50 46.6	+48 28 58.8	M6.5S	_	_	_	_	3.22	_
BD+44-2267	13 21 18.7	+43 59 14.0	S2.5Z	-	_	_	_	_	_
HD 064332	07 53 05.3	$-11\ 37\ 29.3$	S4.5	7.64	3500	0.50	-0.34	3.01	54
HD 62164	07 42 17.5	-105247.2	S5-S6	_	_	_	_	_	_
HD 44544	06 22 23.8	+03 25 27.8	SC5.5	_	_	_	_	0.30	_
2MASSJ0746+2000AB	07 46 42.6	+20 00 32.0	L0.5	_	_	_	_	81.90	_
2MASSJ0208+2542	02 08 18.3	+25 42 53.0	L1	_	_	_	_	10.67	_
2MASSJ1439+1929	14 39 28.4	+19 29 15.0	L1	_	_	_	_	69.60	_
Kelu-1AB	13 05 40.2	$-25\ 41\ 06.0$	L2	_	_	_	_	53.60	_
2MASSJ1146+2230AB	11 46 34.5	+22 30 53.0	L3	_	_	_	_	36.80	_
2MASSJ1506+1321	15 06 54.4	+13 21 06.0	L3	_	_	_	_	70.92	_
2MASSJ0036+1821	00 36 16.2	+18 21 10.0	L3.5	_	_	_	_	114.20	_
2MASSJ2224-0158	22 24 43.8	$-01\ 58\ 52.0$	L4.5	_	_	_	_	88.10	_
2MASSJ1507-1627	15 07 47.7	-162739.0	L5	_	_	_	_	_	_
SDSSJ0539-0059	05 39 52.0	-00 59 02.0	L5	_	_	_	_	_	_
2MASSJ1515+4847	15 15 00.8	+48 47 42.0	L6	_	_	_	_	95.24	_
2MASSJ0825+2115	08 25 19.7	+21 15 52.0	L7.5	_	_	_	_	93.80	_
DENISJ0255-4700	02 55 03.6	$-47\ 00\ 51.0$	L8	_	_	_	_	201.40	_
SDSSJ1254-0122	12 54 53.9	$-01\ 22\ 47.0$	T2	_	_	_	_	84.90	_
2MASSJ0559-1404	05 59 19.1	-14 04 49.0	T4.5					97.70	

Table 4. EW of the *I*-band indices. Part 1.

Star IDs	Ca1	Ca2	Ca3	Pa1	Pa2	Pa3	Mg
				rgiants			
HD 007927	2.83 ± 0.20	3.72 ± 0.16	4.12 ± 0.09	1.48 ± 0.12	2.76 ± 0.21	2.85 ± 0.15	0.11 ± 0.06
HD 006130	2.35 ± 0.15	4.50 ± 0.12	5.00 ± 0.06	1.08 ± 0.09	2.97 ± 0.16	4.14 ± 0.11	0.17 ± 0.03
HD 135153	2.21 ± 0.16	4.69 ± 0.12	5.67 ± 0.05	1.10 ± 0.10	3.39 ± 0.16	4.12 ± 0.09	0.08 ± 0.07
HD 164136 HD 182835	1.41 ± 0.06 2.88 ± 0.16	2.99 ± 0.05 5.44 ± 0.13	3.13 ± 0.04 5.92 ± 0.07	0.45 ± 0.04 1.30 ± 0.10	1.35 ± 0.07 3.16 ± 0.20	2.38 ± 0.08 3.74 ± 0.11	0.17 ± 0.04 0.18 ± 0.05
HD 102033 HD 213306	2.58 ± 0.10 2.54 ± 0.08	6.53 ± 0.06	5.92 ± 0.07 5.16 ± 0.07	0.50 ± 0.10 0.50 ± 0.05	0.82 ± 0.17	1.08 ± 0.11	0.18 ± 0.03 0.28 ± 0.06
HD 201078	2.24 ± 0.08	4.76 ± 0.06	4.42 ± 0.05	0.50 ± 0.05 0.58 ± 0.05	1.55 ± 0.10	2.30 ± 0.12	0.20 ± 0.06 0.20 ± 0.06
HD 51956	2.08 ± 0.03	5.08 ± 0.02	4.04 ± 0.06	0.25 ± 0.02	0.43 ± 0.07	0.77 ± 0.14	0.32 ± 0.06
HD 185018	2.00 ± 0.05	5.05 ± 0.04	4.00 ± 0.07	0.37 ± 0.03	0.67 ± 0.15	0.81 ± 0.13	0.36 ± 0.06
HD 216219	1.18 ± 0.05	3.13 ± 0.04	2.42 ± 0.03	0.12 ± 0.03	0.39 ± 0.09	1.11 ± 0.07	0.17 ± 0.04
HD 074395	2.23 ± 0.07	5.69 ± 0.06	4.54 ± 0.07	0.38 ± 0.04	0.53 ± 0.14	1.06 ± 0.15	0.47 ± 0.08
HD 042454	2.56 ± 0.04	6.12 ± 0.04	4.95 ± 0.09	0.40 ± 0.03	0.29 ± 0.15	0.76 ± 0.19	0.33 ± 0.06
HD 202314	2.05 ± 0.05	5.20 ± 0.04	4.05 ± 0.08	0.34 ± 0.03	0.40 ± 0.13	0.81 ± 0.18	0.39 ± 0.08
HD 003421	1.93 ± 0.07	4.46 ± 0.06	3.60 ± 0.06	0.32 ± 0.04	0.50 ± 0.15	1.09 ± 0.11	0.22 ± 0.06
HD 192713 HD 176123	2.37 ± 0.05	5.94 ± 0.04 4.91 ± 0.05	4.73 ± 0.09	0.39 ± 0.03	0.48 ± 0.14 0.40 ± 0.10	0.75 ± 0.19	0.35 ± 0.07
HD 170123 HD 179821	1.98 ± 0.06 2.85 ± 0.19	4.91 ± 0.03 5.23 ± 0.15	3.62 ± 0.07 3.23 ± 0.13	0.27 ± 0.04 1.18 ± 0.12	0.40 ± 0.10 1.80 ± 0.36	0.73 ± 0.16 1.54 ± 0.26	0.50 ± 0.09 0.59 ± 0.09
HD 190113	2.33 ± 0.19 2.33 ± 0.06	5.23 ± 0.13 5.41 ± 0.05	4.12 ± 0.11	0.42 ± 0.04	0.11 ± 0.20	0.27 ± 0.20	0.39 ± 0.09 0.14 ± 0.05
HD 161664	2.48 ± 0.06	6.03 ± 0.05	4.33 ± 0.12	0.46 ± 0.04	0.33 ± 0.24	0.27 ± 0.20 0.44 ± 0.21	0.43 ± 0.10
HD 058367	1.92 ± 0.03	4.87 ± 0.02	3.72 ± 0.08	0.26 ± 0.02	0.44 ± 0.11	0.37 ± 0.18	0.28 ± 0.07
HD 025877	2.04 ± 0.04	4.84 ± 0.03	3.85 ± 0.07	0.27 ± 0.02	0.08 ± 0.10	0.47 ± 0.13	0.42 ± 0.05
HD 208606	2.99 ± 0.12	7.21 ± 0.09	5.63 ± 0.14	0.69 ± 0.07	0.30 ± 0.32	0.92 ± 0.24	0.62 ± 0.12
HD 122563	0.47 ± 0.03	0.93 ± 0.02	0.66 ± 0.04	0.02 ± 0.02	-0.27 ± 0.10	-0.12 ± 0.05	0.44 ± 0.17
HD 165782	4.31 ± 0.16	9.73 ± 0.13	7.35 ± 0.10	0.93 ± 0.10	0.08 ± 0.28	1.41 ± 0.20	0.14 ± 0.04
HD 044391	2.40 ± 0.05	5.83 ± 0.04	4.48 ± 0.12	0.39 ± 0.03	0.30 ± 0.20	0.20 ± 0.22	0.58 ± 0.10
HD 179870	1.88 ± 0.04	4.79 ± 0.03	3.49 ± 0.10	0.27 ± 0.02	0.07 ± 0.16	0.36 ± 0.19	0.49 ± 0.05
HD 164349	2.04 ± 0.04	5.01 ± 0.03	3.52 ± 0.11	0.42 ± 0.02	0.25 ± 0.16	0.19 ± 0.21	0.49 ± 0.09
HD 091810 HD 212466	1.86 ± 0.03 4.35 ± 0.26	4.56 ± 0.02 10.37 ± 0.21	3.54 ± 0.11 8.20 ± 0.17	0.40 ± 0.02 0.87 ± 0.16	0.18 ± 0.12 -0.23 \pm 0.46	0.22 ± 0.22 0.57 ± 0.34	0.62 ± 0.08 0.62 ± 0.24
HD 023082	4.33 ± 0.26 2.41 ± 0.06	5.72 ± 0.05	6.20 ± 0.17 4.12 ± 0.12	0.87 ± 0.10 0.59 ± 0.04	-0.23 ± 0.40 -0.18 ± 0.18	0.37 ± 0.34 0.04 ± 0.25	0.62 ± 0.24 0.63 ± 0.13
HD 063302	2.41 ± 0.00 2.94 ± 0.12	7.36 ± 0.10	5.76 ± 0.12	0.59 ± 0.04 0.55 ± 0.07	-0.13 ± 0.16 -0.13 ± 0.26	0.56 ± 0.25	0.63 ± 0.13 0.61 ± 0.14
HD 187238	2.60 ± 0.08	6.36 ± 0.06	4.91 ± 0.12	0.78 ± 0.07	0.00 ± 0.29	0.47 ± 0.22	0.69 ± 0.15
HD 185622	2.78 ± 0.09	6.50 ± 0.08	5.25 ± 0.15	0.98 ± 0.06	0.09 ± 0.27	0.58 ± 0.27	0.85 ± 0.19
HD 201065	2.57 ± 0.06	5.68 ± 0.05	4.09 ± 0.13	0.72 ± 0.04	0.04 ± 0.19	0.37 ± 0.23	0.71 ± 0.13
HD 207991	1.98 ± 0.04	4.37 ± 0.03	3.01 ± 0.10	0.51 ± 0.02	-0.16 ± 0.18	0.48 ± 0.17	0.64 ± 0.14
HD 216946	2.28 ± 0.06	5.89 ± 0.05	4.56 ± 0.11	0.72 ± 0.04	-0.13 ± 0.19	0.29 ± 0.19	0.57 ± 0.09
HD 236697	2.28 ± 0.04	5.83 ± 0.03	4.79 ± 0.12	0.70 ± 0.02	-0.20 ± 0.23	0.13 ± 0.21	0.73 ± 0.17
HD 181475	2.58 ± 0.04	6.15 ± 0.03	4.77 ± 0.14	0.82 ± 0.02	-0.16 ± 0.32	0.34 ± 0.23	0.84 ± 0.16
HD 014404	2.40 ± 0.04	6.11 ± 0.04	5.10 ± 0.12	0.89 ± 0.03	-0.14 ± 0.28	0.59 ± 0.22	0.75 ± 0.20
HD 039801 HD 035601	2.41 ± 0.05 2.54 ± 0.06	5.76 ± 0.04 6.11 ± 0.05	5.06 ± 0.11 5.07 ± 0.15	0.84 ± 0.03 0.90 ± 0.04	0.28 ± 0.21 0.08 ± 0.32	0.52 ± 0.17 0.55 ± 0.25	0.73 ± 0.16 0.82 ± 0.17
HD 033601 HD 206936	2.34 ± 0.06 2.33 ± 0.25	6.39 ± 0.20	5.07 ± 0.13 5.28 ± 0.16	0.90 ± 0.04 0.69 ± 0.15	-0.22 ± 0.32	0.33 ± 0.23 0.28 ± 0.38	0.82 ± 0.17 0.55 ± 0.25
HD 010465	2.33 ± 0.23 2.47 ± 0.05	6.09 ± 0.20 6.09 ± 0.04	4.89 ± 0.10	0.69 ± 0.13 0.63 ± 0.03	0.12 ± 0.23	0.28 ± 0.38 0.29 ± 0.20	0.53 ± 0.25 0.61 ± 0.15
HD 023475	2.03 ± 0.04	5.37 ± 0.03	4.32 ± 0.09	0.65 ± 0.03	0.12 ± 0.23 0.21 ± 0.17	0.18 ± 0.18	0.61 ± 0.15 0.61 ± 0.15
HD 040239	1.54 ± 0.06	5.00 ± 0.05	4.06 ± 0.08	0.53 ± 0.04	0.19 ± 0.16	0.45 ± 0.14	0.85 ± 0.21
RWCyg	2.57 ± 0.05	5.84 ± 0.04	5.06 ± 0.16	1.08 ± 0.03	-0.04 ± 0.34	0.28 ± 0.25	0.52 ± 0.12
HD 019058	0.79 ± 0.11	3.82 ± 0.08	3.34 ± 0.07	0.37 ± 0.07	0.44 ± 0.15	0.35 ± 0.15	0.47 ± 0.12
HD 156014	0.04 ± 0.12	2.93 ± 0.09	3.23 ± 0.05	0.19 ± 0.08	0.77 ± 0.23	0.50 ± 0.11	0.41 ± 0.13
			C:	ante			
HD 089025	1.58 ± 0.06	3.58 ± 0.05	4.13 ± 0.03	ants 0.44 ± 0.04	2.09 ± 0.08	3.33 ± 0.06	0.10 ± 0.05
HD 089023 HD 027397	1.38 ± 0.06 1.20 ± 0.04	3.38 ± 0.03 2.79 ± 0.04	4.13 ± 0.03 3.75 ± 0.04	0.44 ± 0.04 0.19 ± 0.03	1.82 ± 0.08	3.39 ± 0.00 3.39 ± 0.09	0.10 ± 0.03 0.03 ± 0.09
HD 027397	1.58 ± 0.04	3.48 ± 0.05	3.84 ± 0.04	0.19 ± 0.03 0.30 ± 0.04	1.82 ± 0.09 1.89 ± 0.11	2.87 ± 0.09	0.03 ± 0.09 0.11 ± 0.03
HD 40535	1.57 ± 0.07	3.24 ± 0.06	3.49 ± 0.04	0.30 ± 0.01 0.30 ± 0.04	1.42 ± 0.12	2.57 ± 0.07 2.57 ± 0.08	0.11 ± 0.05 0.12 ± 0.06
HD 017918	1.39 ± 0.07	3.53 ± 0.05	3.20 ± 0.05	0.25 ± 0.04	1.18 ± 0.12	2.09 ± 0.09	0.11 ± 0.05
HD 021770	0.93 ± 0.02	2.45 ± 0.02	2.44 ± 0.02	0.10 ± 0.02	0.68 ± 0.04	1.49 ± 0.06	0.14 ± 0.03
HD 075555	1.47 ± 0.04	3.13 ± 0.03	3.07 ± 0.04	0.20 ± 0.02	0.61 ± 0.05	1.49 ± 0.09	0.22 ± 0.06
HD 186155	1.46 ± 0.04	3.63 ± 0.03	3.60 ± 0.05	0.20 ± 0.03	0.97 ± 0.09	2.15 ± 0.14	0.22 ± 0.04
HD 218804	1.06 ± 0.03	2.68 ± 0.03	2.32 ± 0.04	0.13 ± 0.02	0.35 ± 0.07	0.89 ± 0.07	0.20 ± 0.03
HD 160365	1.41 ± 0.04	3.42 ± 0.03	2.81 ± 0.05	0.21 ± 0.02	0.86 ± 0.07	1.25 ± 0.10	0.24 ± 0.03
HD 011443	1.23 ± 0.04	3.00 ± 0.03	2.47 ± 0.04	0.15 ± 0.02	0.46 ± 0.07	1.16 ± 0.09	0.21 ± 0.03
HD 124850	1.33 ± 0.04	2.92 ± 0.04	2.46 ± 0.04	0.17 ± 0.03	0.46 ± 0.10	0.98 ± 0.08	0.26 ± 0.04
HD 220657 HD 006903	1.39 ± 0.04 1.56 ± 0.06	3.66 ± 0.03 3.98 ± 0.05	2.85 ± 0.05 3.05 ± 0.06	0.10 ± 0.02 0.29 ± 0.04	0.57 ± 0.08 0.54 ± 0.12	0.95 ± 0.11 0.81 ± 0.11	0.22 ± 0.04 0.32 ± 0.04
HD 219477	1.82 ± 0.00	4.46 ± 0.03	3.03 ± 0.00 3.31 ± 0.06	0.29 ± 0.04 0.26 ± 0.03	0.54 ± 0.12 0.51 ± 0.10	0.81 ± 0.11 0.73 ± 0.12	0.32 ± 0.04 0.36 ± 0.05
HD 126868	1.45 ± 0.05	3.31 ± 0.04	2.52 ± 0.05	0.16 ± 0.03	0.28 ± 0.12	0.54 ± 0.12	0.28 ± 0.04

Table 4. continued.

Star IDs	Ca1	Ca2	Ca3	Pa1	Pa2	Pa3	Mg
HD 088639	1.31 ± 0.02	3.55 ± 0.02	2.55 ± 0.06	0.14 ± 0.01	0.40 ± 0.09	0.31 ± 0.13	0.32 ± 0.05
HD 108477	1.69 ± 0.02	4.18 ± 0.04	3.31 ± 0.07	0.14 ± 0.01 0.24 ± 0.03	0.35 ± 0.09	0.68 ± 0.16	0.32 ± 0.03 0.35 ± 0.03
HD 018474	1.48 ± 0.03	3.93 ± 0.02	2.93 ± 0.06	0.14 ± 0.02	0.18 ± 0.11	0.29 ± 0.12	0.33 ± 0.04
HD 021018	1.96 ± 0.04	4.74 ± 0.03	3.18 ± 0.07	0.25 ± 0.02	0.23 ± 0.10	0.50 ± 0.15	0.38 ± 0.04
HD 010697	1.18 ± 0.04	3.38 ± 0.03	2.48 ± 0.07	0.14 ± 0.02	0.33 ± 0.12	0.39 ± 0.14	0.32 ± 0.05
HD 193896	1.65 ± 0.04	4.10 ± 0.03	3.11 ± 0.07	0.26 ± 0.02	0.24 ± 0.10	0.52 ± 0.14	0.38 ± 0.04
HD 182694	1.68 ± 0.03	4.08 ± 0.02	3.06 ± 0.07	0.18 ± 0.02	0.27 ± 0.10	0.17 ± 0.15	0.43 ± 0.05
HD 114946	1.18 ± 0.02	3.05 ± 0.02	2.01 ± 0.06	0.07 ± 0.01	0.05 ± 0.06	0.08 ± 0.14	0.33 ± 0.05
HD 020618	1.21 ± 0.04	3.40 ± 0.03	2.52 ± 0.06	0.19 ± 0.02	0.30 ± 0.12	0.24 ± 0.13	0.43 ± 0.04
HD 104979	1.43 ± 0.02	3.69 ± 0.02	2.48 ± 0.06	0.15 ± 0.01	0.50 ± 0.07	0.41 ± 0.12	0.33 ± 0.05
HD 135722	1.34 ± 0.03	3.59 ± 0.03	2.48 ± 0.05	0.22 ± 0.02	0.08 ± 0.08	0.31 ± 0.13	0.40 ± 0.05
HD 094481	1.61 ± 0.02	4.07 ± 0.02	2.93 ± 0.06	0.20 ± 0.01	0.30 ± 0.07	0.44 ± 0.12	0.47 ± 0.05
HD 170820	2.25 ± 0.06	5.50 ± 0.05 3.79 ± 0.02	4.58 ± 0.09	0.49 ± 0.04	0.19 ± 0.14	0.80 ± 0.22	0.44 ± 0.11
HD 222093 HD 100006	1.41 ± 0.02 1.40 ± 0.05	3.79 ± 0.02 3.91 ± 0.04	2.85 ± 0.07 2.77 ± 0.09	0.24 ± 0.02 0.35 ± 0.03	0.17 ± 0.11 0.24 ± 0.13	0.19 ± 0.14 0.26 ± 0.17	0.41 ± 0.05 0.46 ± 0.06
HD 036134	1.40 ± 0.03 1.49 ± 0.03	4.07 ± 0.04	2.77 ± 0.09 2.77 ± 0.08	0.33 ± 0.03 0.11 ± 0.02	0.24 ± 0.13 0.05 ± 0.10	0.20 ± 0.17 0.18 ± 0.16	0.40 ± 0.00 0.47 ± 0.06
HD 025975	1.45 ± 0.03	3.69 ± 0.02	2.55 ± 0.07	0.20 ± 0.02	0.27 ± 0.07	0.29 ± 0.16	0.45 ± 0.06
HD 142091	1.43 ± 0.03	3.83 ± 0.02	2.59 ± 0.09	0.32 ± 0.02	0.18 ± 0.11	0.33 ± 0.20	0.51 ± 0.08
HD 165438	1.47 ± 0.03	3.71 ± 0.02	2.54 ± 0.08	0.31 ± 0.02	0.18 ± 0.09	0.34 ± 0.17	0.58 ± 0.07
HD 124897	1.41 ± 0.02	3.82 ± 0.02	2.67 ± 0.06	0.22 ± 0.01	0.02 ± 0.07	-0.01 ± 0.14	0.56 ± 0.06
HD 002901	1.42 ± 0.02	3.75 ± 0.02	2.60 ± 0.06	0.24 ± 0.01	0.05 ± 0.11	0.29 ± 0.11	0.61 ± 0.08
HD 132935	1.51 ± 0.04	4.07 ± 0.04	3.25 ± 0.06	0.27 ± 0.03	0.12 ± 0.10	0.14 ± 0.12	0.53 ± 0.06
HD 137759	1.65 ± 0.03	4.07 ± 0.02	2.97 ± 0.10	0.41 ± 0.02	0.13 ± 0.09	0.15 ± 0.19	0.45 ± 0.06
HD 035620	1.87 ± 0.06	4.81 ± 0.05	3.34 ± 0.11	0.59 ± 0.04	0.20 ± 0.14	0.46 ± 0.24	0.67 ± 0.11
HD 178208	2.02 ± 0.03	4.65 ± 0.03	3.28 ± 0.11	0.53 ± 0.02	0.24 ± 0.13	0.27 ± 0.22	0.54 ± 0.09
HD 221246	1.95 ± 0.04	5.17 ± 0.04	3.91 ± 0.12	0.49 ± 0.03	0.40 ± 0.19	0.31 ± 0.21	0.56 ± 0.11
HD 114960 HD 099998	1.82 ± 0.02 1.65 ± 0.02	4.54 ± 0.02 4.17 ± 0.01	3.41 ± 0.12	0.59 ± 0.01 0.36 ± 0.01	0.31 ± 0.08	0.15 ± 0.23	0.58 ± 0.10 0.77 ± 0.09
HD 181596	1.03 ± 0.02 1.90 ± 0.05	4.82 ± 0.04	3.19 ± 0.08 3.84 ± 0.07	0.30 ± 0.01 0.48 ± 0.03	-0.21 ± 0.11 0.11 ± 0.14	-0.02 ± 0.17 0.27 ± 0.15	0.77 ± 0.09 0.56 ± 0.10
HD 120477	1.80 ± 0.03 1.81 ± 0.02	4.65 ± 0.01	3.19 ± 0.07	0.45 ± 0.03 0.45 ± 0.01	-0.08 ± 0.13	0.27 ± 0.13 0.30 ± 0.14	0.69 ± 0.10
HD 003346	1.96 ± 0.01	4.80 ± 0.01	3.30 ± 0.08	0.48 ± 0.01	0.17 ± 0.16	0.22 ± 0.14	0.66 ± 0.11
HD 194193	1.70 ± 0.04	4.71 ± 0.04	3.34 ± 0.09	0.51 ± 0.03	-0.19 ± 0.13	-0.20 ± 0.17	0.59 ± 0.09
HD 213893	1.73 ± 0.03	4.21 ± 0.03	3.12 ± 0.06	0.43 ± 0.02	-0.15 ± 0.10	0.17 ± 0.15	0.50 ± 0.09
HD 204724	2.01 ± 0.03	5.10 ± 0.02	4.03 ± 0.09	0.59 ± 0.02	0.23 ± 0.14	0.65 ± 0.18	0.69 ± 0.11
HD 219734	1.62 ± 0.08	5.04 ± 0.07	3.63 ± 0.09	0.43 ± 0.05	0.11 ± 0.18	0.23 ± 0.17	0.54 ± 0.11
HD 120052	1.55 ± 0.05	4.40 ± 0.04	3.14 ± 0.07	0.44 ± 0.03	0.03 ± 0.19	0.18 ± 0.13	0.57 ± 0.09
HD 039045	1.56 ± 0.03	4.32 ± 0.02	2.85 ± 0.08	0.18 ± 0.02	-0.13 ± 0.20	0.01 ± 0.13	0.57 ± 0.10
HD 028487	0.93 ± 0.09	4.00 ± 0.07	3.06 ± 0.06	0.41 ± 0.06	0.03 ± 0.14	0.27 ± 0.15	0.43 ± 0.10
HD 214665 HD 027598	1.13 ± 0.06 0.99 ± 0.08	4.15 ± 0.05 3.95 ± 0.06	3.62 ± 0.07 3.24 ± 0.08	0.51 ± 0.04 0.34 ± 0.05	0.33 ± 0.10	0.60 ± 0.18 0.22 ± 0.17	0.53 ± 0.13 0.44 ± 0.09
HD 027398 HD 004408	0.99 ± 0.08 1.27 ± 0.07	4.60 ± 0.06	3.24 ± 0.08 3.64 ± 0.08	0.34 ± 0.05 0.42 ± 0.05	0.05 ± 0.17 0.45 ± 0.12	0.22 ± 0.17 0.23 ± 0.15	0.44 ± 0.09 0.56 ± 0.11
HD 204585	0.87 ± 0.07	3.98 ± 0.10	3.95 ± 0.06	0.42 ± 0.03 0.51 ± 0.09	0.76 ± 0.12	0.23 ± 0.13 0.72 ± 0.18	0.50 ± 0.11 0.53 ± 0.15
HD 014386	-2.11 ± 0.30	1.04 ± 0.20	1.79 ± 0.19	0.32 ± 0.22	1.32 ± 0.61	0.78 ± 0.13	0.08 ± 0.12
HD 175865	0.29 ± 0.13	3.30 ± 0.10	3.56 ± 0.06	0.33 ± 0.09	0.50 ± 0.19	0.30 ± 0.14	0.48 ± 0.14
HD 094705	-0.21 ± 0.11	2.31 ± 0.08	3.02 ± 0.06	0.14 ± 0.07	0.95 ± 0.24	0.15 ± 0.12	0.38 ± 0.12
HD 018191	-0.12 ± 0.14	2.77 ± 0.10	3.07 ± 0.06	0.05 ± 0.10	1.09 ± 0.20	0.10 ± 0.14	0.34 ± 0.12
HD 196610	-0.12 ± 0.16	2.13 ± 0.11	3.02 ± 0.07	0.36 ± 0.11	1.04 ± 0.29	0.62 ± 0.14	0.25 ± 0.10
HD 108849	-0.91 ± 0.15	1.40 ± 0.11	2.72 ± 0.11	0.14 ± 0.11	1.51 ± 0.33	0.98 ± 0.18	0.06 ± 0.12
HD 207076	-0.61 ± 0.19	1.59 ± 0.13	3.16 ± 0.11	0.30 ± 0.13	1.33 ± 0.38	0.82 ± 0.17	0.18 ± 0.11
			D				
HD 108519	1.11 ± 0.07	2.60 ± 0.05	3.46 ± 0.02	0.20 \pm 0.04	1.44 ± 0.09	3.18 ± 0.08	-0.03 ± 0.07
HD 108519 HD 213135	1.11 ± 0.07 1.01 ± 0.04	2.80 ± 0.03 2.32 ± 0.03	3.46 ± 0.02 2.58 ± 0.03	0.20 ± 0.04 0.14 ± 0.02	0.69 ± 0.06	3.18 ± 0.08 1.95 ± 0.07	-0.03 ± 0.07 0.11 ± 0.05
HD 113139	0.90 ± 0.04	2.72 ± 0.03 2.72 ± 0.04	2.61 ± 0.03	0.14 ± 0.02 0.16 ± 0.03	1.00 ± 0.08	2.03 ± 0.07 2.03 ± 0.07	0.11 ± 0.05 0.11 ± 0.05
HD 026015	1.03 ± 0.04	3.02 ± 0.03	3.09 ± 0.04	0.10 ± 0.03 0.19 ± 0.02	1.07 ± 0.08	2.03 ± 0.07 2.07 ± 0.10	0.06 ± 0.07
HD 016232	1.08 ± 0.04	2.99 ± 0.03	2.10 ± 0.05	0.08 ± 0.02	0.12 ± 0.07	0.83 ± 0.12	0.28 ± 0.03
HD 087822	1.16 ± 0.02	2.99 ± 0.02	2.74 ± 0.02	0.09 ± 0.01	0.67 ± 0.05	1.41 ± 0.06	0.17 ± 0.04
HD 027524	1.13 ± 0.03	2.89 ± 0.03	2.56 ± 0.04	0.05 ± 0.02	0.22 ± 0.08	1.07 ± 0.10	0.20 ± 0.03
HD 215648	1.06 ± 0.04	2.65 ± 0.03	2.39 ± 0.03	0.09 ± 0.02	0.33 ± 0.06	0.89 ± 0.07	0.21 ± 0.03
HD 126660	1.04 ± 0.03	2.86 ± 0.03	2.17 ± 0.05	0.06 ± 0.02	0.33 ± 0.08	0.73 ± 0.11	0.26 ± 0.03
HD 219623	1.29 ± 0.03	3.42 ± 0.03	2.70 ± 0.05	0.13 ± 0.02	0.48 ± 0.08	0.71 ± 0.10	0.27 ± 0.04
HD 165908	0.89 ± 0.02	2.35 ± 0.01	1.94 ± 0.04	-0.00 ± 0.01	0.14 ± 0.03	0.65 ± 0.09	0.23 ± 0.03
HD 102870	1.15 ± 0.02	3.07 ± 0.02	2.58 ± 0.05	0.04 ± 0.01	0.21 ± 0.04	0.71 ± 0.14	0.25 ± 0.04
HD 027383 HD 114710	1.10 ± 0.04 1.16 ± 0.03	3.33 ± 0.04 3.28 ± 0.02	2.77 ± 0.05 2.76 ± 0.04	0.14 ± 0.03 0.01 ± 0.02	0.55 ± 0.13 0.12 ± 0.05	0.98 ± 0.08 0.62 ± 0.12	0.19 ± 0.04 0.36 ± 0.03
HD 176051	1.16 ± 0.03 1.16 ± 0.03	3.28 ± 0.02 2.99 ± 0.02	2.76 ± 0.04 2.25 ± 0.06	0.01 ± 0.02 0.10 ± 0.02	-0.12 ± 0.05 -0.16 ± 0.05	0.02 ± 0.12 0.19 ± 0.16	0.38 ± 0.06 0.38 ± 0.06
HD 109358	1.16 ± 0.03 1.16 ± 0.03	3.05 ± 0.02	2.15 ± 0.05	0.10 ± 0.02 0.09 ± 0.02	0.17 ± 0.09	0.19 ± 0.10 0.57 ± 0.11	0.38 ± 0.00 0.29 ± 0.04

Table 4. continued.

Star IDs	Cal	Ca2	Ca3	Pa1	Pa2	Pa3	Mg
HD 095128	1.19 ± 0.02	3.34 ± 0.02	2.61 ± 0.05	0.06 ± 0.01	0.36 ± 0.07	0.44 ± 0.10	0.34 ± 0.04
HD 020619	1.11 ± 0.04	3.12 ± 0.03	2.04 ± 0.04	0.05 ± 0.02	0.06 ± 0.09	0.25 ± 0.10	0.36 ± 0.04
HD 010307	1.25 ± 0.02	3.31 ± 0.02	2.09 ± 0.05	0.09 ± 0.01	0.32 ± 0.08	0.39 ± 0.13	0.35 ± 0.05
HD 076151	1.15 ± 0.05	3.48 ± 0.04	2.29 ± 0.05	0.15 ± 0.03	0.13 ± 0.11	0.31 ± 0.12	0.42 ± 0.05
HD 214850	1.14 ± 0.02	3.15 ± 0.02	2.66 ± 0.04	0.12 ± 0.01	0.22 ± 0.07	0.26 ± 0.10	0.35 ± 0.04
HD 165185	1.07 ± 0.04	3.08 ± 0.03	2.09 ± 0.05	0.05 ± 0.02	0.01 ± 0.06	0.43 ± 0.11	0.52 ± 0.05
HD 115617	1.24 ± 0.03	3.32 ± 0.03	2.60 ± 0.04	0.03 ± 0.02	0.14 ± 0.05	0.10 ± 0.12	0.30 ± 0.04
HD 101501	1.15 ± 0.04	3.32 ± 0.03	2.62 ± 0.06	0.17 ± 0.02	0.31 ± 0.14	0.43 ± 0.12	0.46 ± 0.05
HD 075732	1.46 ± 0.05	3.75 ± 0.04	2.74 ± 0.07	0.17 ± 0.03	0.23 ± 0.09	0.24 ± 0.14	0.67 ± 0.06
HD 145675	1.38 ± 0.03	3.67 ± 0.03	2.25 ± 0.10	0.18 ± 0.02	0.09 ± 0.08	-0.01 ± 0.22	0.66 ± 0.07
HD 010476	1.25 ± 0.04	3.48 ± 0.03	2.45 ± 0.06	0.11 ± 0.02	0.22 ± 0.09	0.34 ± 0.16	0.55 ± 0.06
HD 003765	1.18 ± 0.02	3.44 ± 0.02	2.40 ± 0.07	0.08 ± 0.01	0.04 ± 0.08	0.01 ± 0.15	0.60 ± 0.08
HD 219134	1.27 ± 0.04	3.72 ± 0.04	2.75 ± 0.07	0.27 ± 0.03	0.19 ± 0.11	0.43 ± 0.18	0.67 ± 0.09
HD 045977	1.30 ± 0.03	3.82 ± 0.02	2.50 ± 0.08	0.30 ± 0.02	-0.02 ± 0.12	0.33 ± 0.21	0.81 ± 0.13
HD 207991	1.98 ± 0.04	4.37 ± 0.03	3.00 ± 0.10	0.51 ± 0.02	-0.16 ± 0.18	0.49 ± 0.17	0.57 ± 0.09
HD 036003	1.37 ± 0.03	3.65 ± 0.02	2.18 ± 0.06	0.22 ± 0.02	-0.06 ± 0.07	0.06 ± 0.14	0.79 ± 0.10
HD 201092	0.97 ± 0.02	2.80 ± 0.02	1.52 ± 0.04	0.28 ± 0.01	-0.54 ± 0.06	-0.48 ± 0.12	0.66 ± 0.11
HD 237903	1.25 ± 0.03	3.57 ± 0.02	2.25 ± 0.06	0.27 ± 0.02	0.06 ± 0.13	-0.05 ± 0.14	0.62 ± 0.11
HD 019305	0.94 ± 0.06	3.29 ± 0.05	2.34 ± 0.05	0.30 ± 0.04	-0.02 ± 0.09	0.27 ± 0.15	0.67 ± 0.11
HD 209290	0.83 ± 0.06	2.78 ± 0.05	1.53 ± 0.06	0.21 ± 0.04	-0.20 ± 0.10	-0.23 ± 0.13	0.52 ± 0.10
HD 036395	0.74 ± 0.04	2.79 ± 0.03	1.68 ± 0.05	0.27 ± 0.02	0.16 ± 0.08	0.18 ± 0.13	0.54 ± 0.11
HD 042581	0.84 ± 0.05	2.80 ± 0.04	1.50 ± 0.05	0.20 ± 0.03	-0.15 ± 0.08	0.01 ± 0.11	0.47 ± 0.10
G1806	0.42 ± 0.04	2.21 ± 0.03	0.97 ± 0.05	0.21 ± 0.03	-0.02 ± 0.07	-0.48 ± 0.08	0.51 ± 0.10
HD 095735	0.68 ± 0.02	2.27 ± 0.02	1.46 ± 0.03	0.11 ± 0.01	-0.22 ± 0.03	-0.07 ± 0.07	0.40 ± 0.07
Gl381	0.57 ± 0.04	2.06 ± 0.03	0.92 ± 0.05	0.11 ± 0.03	-0.09 ± 0.08	-0.03 ± 0.11	0.41 ± 0.10
Gl581	0.53 ± 0.05	1.81 ± 0.04	0.82 ± 0.04	0.12 ± 0.03	-0.30 ± 0.07	0.01 ± 0.08	0.40 ± 0.10
Gl213	0.13 ± 0.06	1.18 ± 0.05	0.34 ± 0.06	0.02 ± 0.04	-0.35 ± 0.09	-0.15 ± 0.10	0.41 ± 0.09
G1299	0.21 ± 0.04	1.71 ± 0.03	0.45 ± 0.05	0.05 ± 0.03	-0.34 ± 0.04	-0.09 ± 0.12	0.43 ± 0.10
Gl268AB	-0.28 ± 0.05	0.73 ± 0.04	-0.56 ± 0.06	0.04 ± 0.03	-0.57 ± 0.07	-0.37 ± 0.22	0.45 ± 0.11
Gl51	-1.12 ± 0.09	0.26 ± 0.07	-1.49 ± 0.11	0.02 ± 0.06	-0.57 ± 0.15	0.02 ± 0.31	0.52 ± 0.16
G1866	-0.41 ± 0.09	0.76 ± 0.07	-1.82 ± 0.11	0.15 ± 0.06	-0.93 ± 0.15	0.24 ± 0.36	0.57 ± 0.21
GJ1111	-0.73 ± 0.10	1.11 ± 0.07	-3.73 ± 0.16	0.01 ± 0.07	-2.12 ± 0.22	-0.28 ± 0.53	0.84 ± 0.25
G1644C	-0.85 ± 0.21	1.01 ± 0.15	-4.40 ± 0.19	0.07 ± 0.15	-2.26 ± 0.30	0.49 ± 0.67	0.92 ± 0.32
Gl752B	-0.60 ± 0.23	1.32 ± 0.16	-3.09 ± 0.18	0.15 ± 0.17	-2.05 ± 0.45	0.78 ± 0.63	0.94 ± 0.31
LP412-31	-2.21 ± 0.14	1.92 ± 0.09	-5.90 ± 0.20	-0.47 ± 0.10	-2.75 ± 0.41	0.07 ± 0.83	1.17 ± 0.35
LHS2065	0.19 ± 0.09	1.30 ± 0.06	-3.09 ± 0.25	0.86 ± 0.06	-2.06 ± 0.51	-0.15 ± 0.74	0.97 ± 0.34
LHS2924	-1.30 ± 0.56	-1.90 ± 0.42	-2.63 ± 0.27	0.15 ± 0.36	-3.36 ± 0.96	2.75 ± 0.74	0.95 ± 0.48
LP944-20	-1.14 ± 0.37	1.13 ± 0.24	-2.93 ± 0.42	-0.84 ± 0.28	-1.76 ± 1.14	1.30 ± 0.73	0.65 ± 0.34
BRIB0021	-9.51 ± 1.36	-3.60 ± 0.94	-8.81 ± 0.48	-1.35 ± 0.96	-4.30 ± 1.62	-3.91 ± 1.29	0.45 ± 0.55

Table 5. EW of the *I*-band indices. Part 2.

Star IDs	Pa3	Pa4	Pa5	Pa6	CaT	PaT	CaT*
HD 007927	3.48 ± 0.06	3.16 ± 0.04	Supergia 2.02 ± 0.02	ants 3.09 ± 0.35	10.8 ± 0.9	6.9 ± 0.4	4.4 ± 0.8
HD 006130	4.18 ± 0.12	4.13 ± 0.04	1.16 ± 0.02	4.86 ± 0.17	12.0 ± 0.6	8.1 ± 0.3	4.5 ± 0.5
HD 135153	4.29 ± 0.12	4.48 ± 0.09	0.78 ± 0.04	4.96 ± 0.16	12.8 ± 0.6	8.7 ± 0.3	4.8 ± 0.5
HD 164136	2.30 ± 0.08	2.28 ± 0.04	1.10 ± 0.03	2.84 ± 0.07	7.4 ± 0.3	4.1 ± 0.1	3.7 ± 0.3
HD 182835	3.86 ± 0.08	3.80 ± 0.06	1.17 ± 0.02	4.35 ± 0.26	14.3 ± 0.6	8.1 ± 0.3	6.8 ± 0.6
HD 213306	0.95 ± 0.10	2.14 ± 0.07	1.50 ± 0.02	2.16 ± 0.06	14.0 ± 0.4	2.3 ± 0.2	12.0 ± 0.4
HD 201078	2.35 ± 0.11	2.66 ± 0.05	1.40 ± 0.02	3.33 ± 0.09	11.3 ± 0.4	4.3 ± 0.2	7.3 ± 0.3
HD 51956	0.70 ± 0.10	1.58 ± 0.09	1.00 ± 0.02	1.89 ± 0.05	10.9 ± 0.6	1.1 ± 0.3	9.8 ± 0.5
HD 185018 HD 216219	0.63 ± 0.09 0.96 ± 0.08	1.59 ± 0.04 1.33 ± 0.02	0.86 ± 0.02 0.83 ± 0.01	1.88 ± 0.04 1.78 ± 0.03	10.8 ± 0.4 6.5 ± 0.3	1.7 ± 0.2 1.5 ± 0.1	9.2 ± 0.4 5.1 ± 0.3
HD 074395	0.70 ± 0.08 0.71 ± 0.12	1.53 ± 0.02 1.57 ± 0.05	1.15 ± 0.02	2.42 ± 0.02	12.2 ± 0.5	1.8 ± 0.1 1.8 ± 0.2	10.6 ± 0.4
HD 042454	0.40 ± 0.12	1.72 ± 0.09	1.23 ± 0.02	1.83 ± 0.04	13.3 ± 0.6	1.2 ± 0.3	12.2 ± 0.6
HD 202314	0.50 ± 0.14	1.41 ± 0.04	1.17 ± 0.03	1.33 ± 0.06	11.0 ± 0.6	1.4 ± 0.2	9.7 ± 0.5
HD 003421	0.81 ± 0.11	1.35 ± 0.02	1.00 ± 0.02	1.78 ± 0.04	9.7 ± 0.5	1.7 ± 0.2	8.1 ± 0.4
HD 192713	0.43 ± 0.15	1.69 ± 0.05	1.31 ± 0.02	2.11 ± 0.05	12.8 ± 0.6	1.5 ± 0.3	11.4 ± 0.5
HD 176123	0.54 ± 0.10	1.65 ± 0.06	0.86 ± 0.02	1.33 ± 0.03	10.4 ± 0.4	1.3 ± 0.2	9.1 ± 0.4
HD 179821	2.04 ± 0.08	2.03 ± 0.15	2.24 ± 0.06	2.23 ± 0.43	11.2 ± 1.3	4.0 ± 0.6	7.5 ± 1.2
HD 190113	-0.17 ± 0.15	1.43 ± 0.05	0.89 ± 0.03	2.23 ± 0.06	11.2 ± 0.9	0.4 ± 0.4	10.8 ± 0.8
HD 161664 HD 058367	0.34 ± 0.17 0.37 ± 0.11	1.65 ± 0.10 1.37 ± 0.04	0.89 ± 0.02 1.06 ± 0.05	1.67 ± 0.07 1.44 ± 0.08	12.5 ± 0.7 10.2 ± 0.5	1.1 ± 0.3 0.9 ± 0.2	11.5 ± 0.6 9.4 ± 0.5
HD 038307 HD 025877	0.37 ± 0.11 0.18 ± 0.09	0.96 ± 0.03	0.84 ± 0.02	1.44 ± 0.08 1.37 ± 0.05	10.2 ± 0.5 10.3 ± 0.6	0.9 ± 0.2 0.5 ± 0.3	9.4 ± 0.5 9.8 ± 0.6
HD 208606	0.18 ± 0.09 0.27 ± 0.20	0.90 ± 0.03 2.13 ± 0.03	0.84 ± 0.02 1.37 ± 0.02	2.84 ± 0.03	10.3 ± 0.0 15.3 ± 1.0	0.5 ± 0.5 1.6 ± 0.4	9.8 ± 0.0 13.8 ± 0.8
HD 122563	0.27 ± 0.20 0.03 ± 0.04	0.03 ± 0.04	0.05 ± 0.02	0.00 ± 0.05	1.7 ± 0.4	-0.6 ± 0.2	2.3 ± 0.4
HD 165782	0.54 ± 0.30	2.49 ± 0.14	2.36 ± 0.05	3.14 ± 0.16	20.8 ± 0.9	2.1 ± 0.4	18.9 ± 0.8
HD 044391	0.06 ± 0.16	1.19 ± 0.07	0.88 ± 0.02	1.76 ± 0.04	12.3 ± 0.8	0.6 ± 0.4	11.7 ± 0.7
HD 179870	0.26 ± 0.09	1.27 ± 0.04	0.67 ± 0.02	1.55 ± 0.04	9.8 ± 0.7	0.5 ± 0.3	9.3 ± 0.6
HD 164349	0.16 ± 0.15	1.13 ± 0.05	0.72 ± 0.03	1.46 ± 0.06	10.2 ± 0.7	0.7 ± 0.3	9.6 ± 0.6
HD 091810	0.20 ± 0.14	1.36 ± 0.05	0.89 ± 0.04	1.65 ± 0.03	9.7 ± 0.7	0.6 ± 0.3	9.1 ± 0.6
HD 212466	-0.24 ± 0.40	2.72 ± 0.16	1.86 ± 0.06	5.27 ± 0.12	22.1 ± 1.4	0.8 ± 0.6	21.4 ± 1.2
HD 023082 HD 063302	-0.10 ± 0.20 0.01 ± 0.27	1.36 ± 0.04 1.92 ± 0.08	0.88 ± 0.06 1.03 ± 0.08	1.68 ± 0.05 3.35 ± 0.06	11.7 ± 0.9 15.5 ± 1.1	0.1 ± 0.4 0.6 ± 0.5	11.6 ± 0.8 14.9 ± 1.0
HD 187238	-0.03 ± 0.23	2.02 ± 0.08	0.98 ± 0.05	2.43 ± 0.05	13.4 ± 0.9	1.0 ± 0.3	12.4 ± 0.8
HD 185622	0.01 ± 0.32	2.28 ± 0.14	1.07 ± 0.05	2.57 ± 0.08	14.0 ± 1.1	1.4 ± 0.5	12.7 ± 1.0
HD 201065	-0.05 ± 0.22	1.51 ± 0.08	1.02 ± 0.06	2.01 ± 0.04	11.9 ± 0.9	0.9 ± 0.4	11.1 ± 0.8
HD 207991	0.17 ± 0.14	1.28 ± 0.07	0.93 ± 0.06	1.10 ± 0.04	9.0 ± 0.7	0.6 ± 0.3	8.5 ± 0.6
HD 216946	-0.22 ± 0.23	1.36 ± 0.02	1.13 ± 0.02	2.11 ± 0.05	12.4 ± 0.8	0.6 ± 0.3	11.8 ± 0.7
HD 236697	-0.01 ± 0.29	1.80 ± 0.07	1.00 ± 0.05	1.99 ± 0.06	12.5 ± 1.0	0.3 ± 0.4	12.2 ± 0.8
HD 181475	-0.08 ± 0.27	1.84 ± 0.11	1.24 ± 0.04	1.59 ± 0.04	13.0 ± 0.9	0.8 ± 0.4	12.3 ± 0.8
HD 014404 HD 039801	0.03 ± 0.33 0.13 ± 0.27	2.08 ± 0.10 1.89 ± 0.08	1.20 ± 0.04 1.36 ± 0.04	2.59 ± 0.07 2.21 ± 0.03	13.2 ± 0.9 13.1 ± 0.8	1.1 ± 0.4 1.5 ± 0.3	12.2 ± 0.8 11.7 ± 0.7
HD 035601	0.13 ± 0.27 0.03 ± 0.29	2.14 ± 0.07	1.30 ± 0.04 1.11 ± 0.04	2.68 ± 0.07	13.1 ± 0.8 13.2 ± 1.0	1.3 ± 0.3 1.3 ± 0.4	12.1 ± 0.7
HD 206936	0.03 ± 0.29 0.03 ± 0.40	2.37 ± 0.07	0.79 ± 0.16	3.33 ± 0.09	13.4 ± 1.4	0.4 ± 0.6	13.0 ± 1.2
HD 010465	0.06 ± 0.25	1.72 ± 0.05	1.10 ± 0.05	2.18 ± 0.09	13.1 ± 0.8	0.8 ± 0.4	12.4 ± 0.7
HD 023475	0.03 ± 0.24	2.04 ± 0.12	0.86 ± 0.06	1.55 ± 0.06	11.6 ± 0.7	0.9 ± 0.3	10.8 ± 0.6
HD 040239	0.33 ± 0.21	1.38 ± 0.07	1.04 ± 0.07	1.48 ± 0.05	10.7 ± 0.7	1.2 ± 0.3	9.6 ± 0.6
RWCyg	-0.21 ± 0.36	2.30 ± 0.16	1.34 ± 0.10	2.36 ± 0.11	13.1 ± 1.0	1.1 ± 0.4	12.0 ± 0.9
HD 019058	0.84 ± 0.29	1.52 ± 0.09	0.74 ± 0.06	1.12 ± 0.04	8.5 ± 0.9	1.4 ± 0.4	7.2 ± 0.7
HD 156014	1.81 ± 0.55	0.77 ± 0.10	0.41 ± 0.06	2.30 ± 0.07	6.0 ± 1.2	1.3 ± 0.5	4.8 ± 0.9
			Giant	:s			
HD 089025	3.30 ± 0.10	3.39 ± 0.06	0.79 ± 0.04	4.23 ± 0.11	9.2 ± 0.2	5.8 ± 0.1	3.8 ± 0.2
HD 027397	3.58 ± 0.14	3.80 ± 0.11	0.29 ± 0.04	5.09 ± 0.10	7.7 ± 0.3	5.5 ± 0.1	2.7 ± 0.2
HD 013174	2.98 ± 0.08	3.16 ± 0.06	0.97 ± 0.02	4.28 ± 0.09	8.9 ± 0.3	5.0 ± 0.1	4.2 ± 0.3
HD 40535	2.94 ± 0.10	3.20 ± 0.06	0.82 ± 0.04	4.26 ± 0.06	8.2 ± 0.3	4.2 ± 0.1	4.3 ± 0.3
HD 017918	2.05 ± 0.06	2.46 ± 0.04	1.12 ± 0.04	3.07 ± 0.05	8.0 ± 0.4	3.4 ± 0.1	4.9 ± 0.3
HD 021770	1.74 ± 0.05	2.27 ± 0.02	0.59 ± 0.01	3.14 ± 0.03	5.7 ± 0.2	2.2 ± 0.1	3.7 ± 0.2
HD 075555	1.77 ± 0.08	2.12 ± 0.05	1.20 ± 0.04	2.96 ± 0.06	7.5 ± 0.4	2.1 ± 0.2	5.5 ± 0.3
HD 186155 HD 218804	2.46 ± 0.08 1.14 ± 0.05	3.05 ± 0.04 2.24 ± 0.02	0.92 ± 0.03 0.63 ± 0.01	3.54 ± 0.11 2.39 ± 0.01	8.5 ± 0.5 5.9 ± 0.2	3.1 ± 0.2 1.3 ± 0.1	5.6 ± 0.4 4.7 ± 0.2
HD 160365	1.14 ± 0.05 1.21 ± 0.05	1.99 ± 0.02	0.03 ± 0.01 0.97 ± 0.01	2.59 ± 0.01 2.50 ± 0.03	3.9 ± 0.2 7.5 ± 0.3	1.3 ± 0.1 2.2 ± 0.1	4.7 ± 0.2 5.5 ± 0.3
HD 011443	1.46 ± 0.05	1.91 ± 0.03 1.91 ± 0.02	0.86 ± 0.02	2.50 ± 0.03 2.53 ± 0.04	6.5 ± 0.3	1.6 ± 0.1	5.0 ± 0.3 5.0 ± 0.3
HD 124850	1.17 ± 0.07	1.67 ± 0.02	0.77 ± 0.03	2.14 ± 0.01	6.6 ± 0.3	1.5 ± 0.1	5.2 ± 0.3
HD 220657	0.98 ± 0.06	1.70 ± 0.05	0.69 ± 0.02	2.01 ± 0.02	7.8 ± 0.3	1.6 ± 0.1	6.3 ± 0.3
HD 006903	0.86 ± 0.06	1.54 ± 0.06	1.06 ± 0.01	1.63 ± 0.02	8.4 ± 0.4	1.5 ± 0.2	7.0 ± 0.3
HD 219477	0.52 ± 0.09	1.46 ± 0.04	0.76 ± 0.02	1.51 ± 0.04	9.4 ± 0.4	1.3 ± 0.2	8.2 ± 0.4
HD 126868	0.68 ± 0.07	1.38 ± 0.02	0.68 ± 0.01	1.76 ± 0.03	7.1 ± 0.4	0.8 ± 0.2	6.3 ± 0.4

Table 5. continued.

Star IDs	Pa3	Pa4	Pa5	Pa6	CaT	PaT	CaT*
HD 088639	0.36 ± 0.08	1.11 ± 0.02	0.48 ± 0.01	0.93 ± 0.02	7.2 ± 0.4	0.7 ± 0.2	6.5 ± 0.4
HD 108477	0.28 ± 0.07	1.43 ± 0.04	0.77 ± 0.03	1.16 ± 0.02	9.0 ± 0.5	1.1 ± 0.2	7.9 ± 0.4
HD 018474	0.32 ± 0.06	1.08 ± 0.04	0.45 ± 0.01	1.11 ± 0.02	8.1 ± 0.4	0.4 ± 0.2	7.7 ± 0.4
HD 021018	0.54 ± 0.07	1.18 ± 0.04	0.99 ± 0.01	1.71 ± 0.04	9.5 ± 0.6	0.7 ± 0.3	8.9 ± 0.6
HD 010697	0.33 ± 0.06	1.20 ± 0.06	0.77 ± 0.04	1.40 ± 0.03	6.8 ± 0.5	0.7 ± 0.2	6.2 ± 0.4
HD 193896	0.35 ± 0.09	1.21 ± 0.07	0.70 ± 0.02	1.41 ± 0.03	8.7 ± 0.5	0.9 ± 0.2	7.9 ± 0.4
HD 182694	0.17 ± 0.09	0.83 ± 0.05	0.66 ± 0.01	0.79 ± 0.02	8.5 ± 0.6	0.4 ± 0.3	8.1 ± 0.5
HD 114946 HD 020618	0.24 ± 0.06	0.83 ± 0.03	0.51 ± 0.03	0.62 ± 0.03	6.0 ± 0.5 7.0 ± 0.4	-0.0 ± 0.2 0.7 ± 0.2	6.0 ± 0.5
HD 020018 HD 104979	0.15 ± 0.08 0.23 ± 0.09	0.70 ± 0.04 1.11 ± 0.03	0.63 ± 0.02 0.81 ± 0.02	1.11 ± 0.03 1.25 ± 0.04	7.0 ± 0.4 7.4 ± 0.4	0.7 ± 0.2 0.9 ± 0.2	6.4 ± 0.3 6.5 ± 0.4
HD 135722	0.23 ± 0.09 0.14 ± 0.08	1.00 ± 0.05	0.81 ± 0.02 0.85 ± 0.02	0.70 ± 0.03	7.4 ± 0.4 7.2 ± 0.4	0.5 ± 0.2 0.5 ± 0.2	6.8 ± 0.4
HD 094481	0.23 ± 0.08	1.30 ± 0.03	0.73 ± 0.01	1.43 ± 0.02	8.4 ± 0.4	0.7 ± 0.2	7.7 ± 0.4
HD 170820	0.52 ± 0.19	1.55 ± 0.03	0.93 ± 0.04	1.85 ± 0.04	11.9 ± 0.8	1.3 ± 0.3	10.7 ± 0.7
HD 222093	0.27 ± 0.09	1.19 ± 0.03	0.42 ± 0.02	1.30 ± 0.03	7.8 ± 0.4	0.5 ± 0.2	7.4 ± 0.4
HD 100006	0.17 ± 0.10	0.73 ± 0.06	0.85 ± 0.02	0.91 ± 0.03	7.9 ± 0.5	0.7 ± 0.2	7.2 ± 0.5
HD 036134	-0.00 ± 0.10	0.89 ± 0.03	0.87 ± 0.04	0.72 ± 0.05	8.0 ± 0.6	0.1 ± 0.2	7.9 ± 0.5
HD 025975	0.24 ± 0.10	1.08 ± 0.02	0.92 ± 0.02	1.23 ± 0.04	7.6 ± 0.5	0.6 ± 0.2	7.0 ± 0.4
HD 142091	0.17 ± 0.13	1.04 ± 0.05	0.64 ± 0.02	1.24 ± 0.03	7.5 ± 0.6	0.6 ± 0.3	7.0 ± 0.5
HD 165438 HD 124897	0.08 ± 0.12 -0.01 ± 0.10	1.06 ± 0.06	0.76 ± 0.01	1.16 ± 0.03	7.6 ± 0.6	0.7 ± 0.2 0.0 ± 0.2	6.9 ± 0.5 7.6 ± 0.4
HD 124897 HD 002901	0.16 ± 0.10	0.62 ± 0.04 1.05 ± 0.02	0.74 ± 0.01 0.67 ± 0.02	0.81 ± 0.03 0.86 ± 0.04	7.6 ± 0.5 7.6 ± 0.5	0.0 ± 0.2 0.5 ± 0.2	7.0 ± 0.4 7.1 ± 0.4
HD 132935	0.10 ± 0.10 0.04 ± 0.10	1.05 ± 0.02 1.05 ± 0.03	0.67 ± 0.02 0.69 ± 0.04	0.60 ± 0.04 0.60 ± 0.05	8.6 ± 0.3	0.3 ± 0.2 0.4 ± 0.2	8.2 ± 0.4
HD 137759	-0.00 ± 0.10	1.03 ± 0.03 1.11 ± 0.05	0.83 ± 0.03	1.42 ± 0.02	8.4 ± 0.7	0.5 ± 0.2	8.0 ± 0.6
HD 035620	0.15 ± 0.17	1.24 ± 0.07	0.77 ± 0.08	1.95 ± 0.03	9.6 ± 0.8	1.0 ± 0.4	8.7 ± 0.7
HD 178208	0.04 ± 0.18	1.43 ± 0.09	0.88 ± 0.05	1.53 ± 0.03	9.6 ± 0.7	0.8 ± 0.3	8.8 ± 0.6
HD 221246	0.10 ± 0.15	1.60 ± 0.02	0.82 ± 0.02	1.71 ± 0.03	10.8 ± 0.7	1.0 ± 0.3	9.8 ± 0.6
HD 114960	-0.05 ± 0.15	1.08 ± 0.08	1.06 ± 0.05	1.18 ± 0.03	9.5 ± 0.8	0.8 ± 0.3	8.8 ± 0.7
HD 099998	-0.08 ± 0.16	1.02 ± 0.06	0.82 ± 0.03	0.80 ± 0.02	8.6 ± 0.7	-0.2 ± 0.3	8.8 ± 0.7
HD 181596	0.02 ± 0.16	1.22 ± 0.04	0.88 ± 0.04	1.24 ± 0.04	10.4 ± 0.6	0.7 ± 0.2	9.7 ± 0.5
HD 120477	0.12 ± 0.16	1.03 ± 0.05	0.87 ± 0.03	1.15 ± 0.06	9.4 ± 0.7	0.4 ± 0.3	9.0 ± 0.6
HD 003346	-0.02 ± 0.19	0.89 ± 0.08	0.76 ± 0.01	1.12 ± 0.05	9.8 ± 0.5	0.7 ± 0.2	9.2 ± 0.5
HD 194193 HD 213893	0.04 ± 0.16 0.16 ± 0.14	1.07 ± 0.06 0.93 ± 0.03	0.97 ± 0.05 1.17 ± 0.03	0.96 ± 0.10 0.81 ± 0.03	9.6 ± 0.6 8.7 ± 0.6	-0.0 ± 0.2 0.2 ± 0.3	9.7 ± 0.5 8.5 ± 0.5
HD 204724	0.10 ± 0.14 0.47 ± 0.18	0.93 ± 0.03 1.43 ± 0.10	1.17 ± 0.03 1.10 ± 0.08	1.17 ± 0.06	10.8 ± 0.8	0.2 ± 0.3 1.2 ± 0.3	9.7 ± 0.7
HD 219734	0.17 ± 0.18	1.62 ± 0.07	0.86 ± 0.04	1.25 ± 0.06	10.3 ± 0.5 10.3 ± 0.7	0.7 ± 0.3	9.6 ± 0.6
HD 120052	0.29 ± 0.15	1.41 ± 0.07	0.85 ± 0.04	0.58 ± 0.03	9.0 ± 0.4	0.7 ± 0.2	8.4 ± 0.4
HD 039045	0.38 ± 0.17	1.06 ± 0.04	0.72 ± 0.04	0.90 ± 0.03	8.6 ± 0.5	-0.1 ± 0.2	8.7 ± 0.5
HD 028487	0.54 ± 0.20	1.34 ± 0.05	0.73 ± 0.04	1.37 ± 0.04	8.2 ± 0.6	0.8 ± 0.3	7.4 ± 0.5
HD 214665	1.17 ± 0.25	1.49 ± 0.21	0.75 ± 0.06	1.49 ± 0.08	9.2 ± 0.8	1.6 ± 0.3	7.7 ± 0.7
HD 027598	0.42 ± 0.22	1.59 ± 0.10	0.51 ± 0.07	0.56 ± 0.06	8.5 ± 0.7	0.8 ± 0.3	7.8 ± 0.6
HD 004408	0.36 ± 0.23	1.29 ± 0.06	1.00 ± 0.04	1.57 ± 0.04	9.7 ± 0.7	1.1 ± 0.3	8.7 ± 0.6
HD 204585	1.63 ± 0.41	1.42 ± 0.10	0.92 ± 0.03	1.61 ± 0.07	9.1 ± 1.0	2.2 ± 0.4	7.1 ± 0.8
HD 014386	1.87 ± 1.18	-0.61 ± 0.16	-0.76 ± 0.11	4.66 ± 0.16	-6.1 ± 7.7 7.5 ± 1.0	-1.9 ± 4.2	-4.3 ± 4.4 6.4 ± 0.8
HD 175865 HD 094705	1.63 ± 0.43 1.71 ± 0.59	1.07 ± 0.10 0.69 ± 0.10	0.73 ± 0.06 0.47 ± 0.06	1.63 ± 0.07 2.15 ± 0.07	4.6 ± 1.3	1.3 ± 0.4 0.8 ± 0.5	3.8 ± 1.0
HD 018191	1.71 ± 0.59 1.57 ± 0.58	0.09 ± 0.10 0.79 ± 0.14	0.47 ± 0.06 0.34 ± 0.06	2.13 ± 0.07 2.28 ± 0.06	5.3 ± 1.2	0.8 ± 0.5 1.0 ± 0.5	4.4 ± 0.9
HD 196610	2.88 ± 0.76	0.97 ± 0.11	0.31 ± 0.06 0.32 ± 0.06	2.78 ± 0.08	3.0 ± 2.2	0.9 ± 1.0	2.2 ± 1.6
HD 108849	2.25 ± 0.97	0.43 ± 0.12	-0.52 ± 0.08	4.79 ± 0.08	-1.7 ± 4.7	-0.4 ± 2.4	-1.4 ± 2.9
HD 207076	3.07 ± 0.92	0.40 ± 0.11	0.16 ± 0.04	3.46 ± 0.07	0.6 ± 3.7	0.2 ± 1.8	0.4 ± 2.4
			Dwarfs				
HD 108519	3.63 ± 0.13	4.05 ± 0.10	0.74 ± 0.05	4.48 ± 0.10	7.1 ± 0.3	4.7 ± 0.1	2.7 ± 0.2
HD 213135	2.39 ± 0.09	3.02 ± 0.05	0.67 ± 0.03	3.56 ± 0.05	5.8 ± 0.2	2.7 ± 0.1	3.3 ± 0.2
HD 113139 HD 026015	2.48 ± 0.08	2.83 ± 0.04	0.69 ± 0.02	3.81 ± 0.05	6.1 ± 0.2	3.1 ± 0.1	3.2 ± 0.2
HD 026015 HD 016232	2.22 ± 0.11 1.08 ± 0.05	2.75 ± 0.07 1.65 ± 0.02	0.58 ± 0.04 0.79 ± 0.02	3.90 ± 0.04 1.91 ± 0.02	7.0 ± 0.3 6.0 ± 0.4	3.3 ± 0.1 0.8 ± 0.2	4.0 ± 0.3 5.2 ± 0.4
HD 010232 HD 087822	1.08 ± 0.03 1.46 ± 0.06	2.11 ± 0.05	0.79 ± 0.02 0.59 ± 0.02	2.83 ± 0.05	6.0 ± 0.4 6.7 ± 0.3	0.8 ± 0.2 2.0 ± 0.1	3.2 ± 0.4 4.8 ± 0.2
HD 027524	1.38 ± 0.05	1.89 ± 0.03	0.83 ± 0.02 0.83 ± 0.02	2.43 ± 0.03 2.43 ± 0.03	6.4 ± 0.4	1.1 ± 0.2	5.4 ± 0.4
HD 215648	0.90 ± 0.06	1.69 ± 0.02	0.64 ± 0.02	1.82 ± 0.02	6.1 ± 0.2	1.3 ± 0.1	4.9 ± 0.2
HD 126660	1.06 ± 0.04	1.78 ± 0.04	0.61 ± 0.02	2.02 ± 0.03	5.8 ± 0.4	0.9 ± 0.2	5.0 ± 0.4
HD 219623	0.89 ± 0.05	2.09 ± 0.03	0.76 ± 0.02	2.05 ± 0.03	7.3 ± 0.3	1.2 ± 0.1	6.1 ± 0.3
HD 165908	0.56 ± 0.04	0.93 ± 0.03	0.44 ± 0.01	1.00 ± 0.01	5.0 ± 0.4	0.6 ± 0.2	4.5 ± 0.4
HD 102870	0.76 ± 0.06	1.51 ± 0.03	0.84 ± 0.03	1.86 ± 0.02	6.4 ± 0.6	0.7 ± 0.2	5.8 ± 0.5
HD 027383	0.93 ± 0.07	1.63 ± 0.03	0.49 ± 0.03	2.47 ± 0.03	7.0 ± 0.3	1.6 ± 0.1	5.5 ± 0.3
HD 114710	0.73 ± 0.05	1.17 ± 0.03	0.78 ± 0.02	1.44 ± 0.03	7.0 ± 0.4	0.6 ± 0.2	6.5 ± 0.3
HD 176051 HD 109358	0.40 ± 0.09	1.19 ± 0.03 1.19 ± 0.02	0.74 ± 0.01	1.49 ± 0.03	6.0 ± 0.7	-0.2 ± 0.3 0.7 ± 0.1	6.2 ± 0.6 5.5 ± 0.3
סככנטו חוו	0.52 ± 0.07	1.17 エ U.U2	0.64 ± 0.01	1.37 ± 0.01	6.2 ± 0.3	U./ ± U.1	J.J ± U.J

Table 5. continued.

Star IDs	Pa3	Pa4	Pa5	Pa6	CaT	PaT	CaT*
HD 095128	0.55 ± 0.06	1.20 ± 0.02	0.57 ± 0.01	1.56 ± 0.02	6.9 ± 0.4	0.7 ± 0.2	6.3 ± 0.4
HD 020619	0.29 ± 0.07	0.89 ± 0.02	0.68 ± 0.01	0.72 ± 0.02	6.0 ± 0.4	0.1 ± 0.2	5.9 ± 0.4
HD 010307	0.42 ± 0.07	1.19 ± 0.04	0.79 ± 0.01	1.50 ± 0.02	6.5 ± 0.4	0.6 ± 0.2	5.9 ± 0.4
HD 076151	0.89 ± 0.08	1.27 ± 0.02	0.69 ± 0.02	1.26 ± 0.02	7.1 ± 0.4	0.6 ± 0.2	6.5 ± 0.4
HD 214850	0.31 ± 0.06	1.02 ± 0.02	0.61 ± 0.01	1.03 ± 0.02	6.7 ± 0.4	0.4 ± 0.2	6.3 ± 0.3
HD 165185	0.43 ± 0.07	1.18 ± 0.03	0.57 ± 0.01	1.13 ± 0.02	6.0 ± 0.4	0.3 ± 0.2	5.8 ± 0.4
HD 115617	0.23 ± 0.07	0.92 ± 0.03	0.65 ± 0.01	0.81 ± 0.02	6.9 ± 0.4	0.1 ± 0.2	6.8 ± 0.3
HD 101501	0.47 ± 0.08	0.87 ± 0.02	0.76 ± 0.00	1.32 ± 0.03	6.8 ± 0.4	0.7 ± 0.2	6.2 ± 0.4
HD 075732	0.14 ± 0.08	1.06 ± 0.03	1.04 ± 0.03	1.23 ± 0.04	7.7 ± 0.6	0.3 ± 0.3	7.4 ± 0.6
HD 145675	-0.03 ± 0.10	1.32 ± 0.03	0.91 ± 0.01	1.17 ± 0.03	7.0 ± 0.7	-0.1 ± 0.3	7.1 ± 0.7
HD 010476	0.38 ± 0.10	1.28 ± 0.03	0.70 ± 0.02	0.80 ± 0.04	7.0 ± 0.4	0.5 ± 0.2	6.5 ± 0.4
HD 003765	0.03 ± 0.12	1.23 ± 0.03	0.63 ± 0.01	0.89 ± 0.05	6.7 ± 0.6	-0.2 ± 0.3	6.9 ± 0.6
HD 219134	0.15 ± 0.15	1.27 ± 0.02	0.84 ± 0.02	0.70 ± 0.03	7.6 ± 0.5	0.8 ± 0.2	6.9 ± 0.4
HD 045977	0.10 ± 0.20	1.06 ± 0.05	0.56 ± 0.02	0.80 ± 0.03	7.4 ± 0.6	0.4 ± 0.3	7.0 ± 0.6
HD 207991	0.17 ± 0.14	1.28 ± 0.07	0.93 ± 0.06	1.10 ± 0.04	9.0 ± 0.7	0.6 ± 0.3	8.5 ± 0.6
HD 036003	-0.07 ± 0.14	0.95 ± 0.05	0.65 ± 0.01	0.43 ± 0.02	6.9 ± 0.5	-0.0 ± 0.2	7.0 ± 0.5
HD 201092	-0.32 ± 0.18	0.56 ± 0.01	0.25 ± 0.03	0.09 ± 0.03	4.8 ± 0.7	-1.1 ± 0.3	5.8 ± 0.7
HD 237903	0.08 ± 0.17	1.32 ± 0.03	0.70 ± 0.02	0.37 ± 0.02	6.9 ± 0.4	0.1 ± 0.2	6.8 ± 0.4
HD 019305	-0.05 ± 0.18	1.44 ± 0.05	0.69 ± 0.02	0.41 ± 0.02	6.5 ± 0.5	0.4 ± 0.2	6.1 ± 0.4
HD 209290	0.08 ± 0.15	1.35 ± 0.04	0.75 ± 0.04	0.12 ± 0.04	5.0 ± 0.5	-0.4 ± 0.2	5.4 ± 0.5
HD 036395	0.22 ± 0.18	1.13 ± 0.03	0.76 ± 0.05	0.49 ± 0.04	5.2 ± 0.5	0.5 ± 0.2	4.8 ± 0.5
HD 042581	0.16 ± 0.17	1.14 ± 0.02	0.34 ± 0.03	0.37 ± 0.03	5.0 ± 0.5	-0.1 ± 0.2	5.2 ± 0.4
G1806	0.02 ± 0.19	0.17 ± 0.08	0.53 ± 0.07	-0.06 ± 0.09	3.7 ± 0.6	-0.5 ± 0.3	4.1 ± 0.6
HD 095735	0.21 ± 0.12	0.82 ± 0.03	0.07 ± 0.01	0.43 ± 0.02	4.3 ± 0.3	-0.3 ± 0.1	4.6 ± 0.3
Gl381	0.40 ± 0.17	0.81 ± 0.06	0.29 ± 0.02	1.04 ± 0.06	3.5 ± 0.5	-0.2 ± 0.2	3.7 ± 0.5
Gl581	0.31 ± 0.19	1.05 ± 0.07	0.46 ± 0.02	0.78 ± 0.04	3.2 ± 0.6	-0.4 ± 0.2	3.5 ± 0.5
Gl213	0.75 ± 0.24	0.61 ± 0.04	0.39 ± 0.03	-0.07 ± 0.05	1.7 ± 0.8	-0.7 ± 0.3	2.4 ± 0.7
G1299	0.62 ± 0.24	0.55 ± 0.02	0.22 ± 0.01	0.64 ± 0.07	2.6 ± 0.9	-0.6 ± 0.3	3.1 ± 0.7
Gl268AB	0.74 ± 0.28	0.71 ± 0.05	0.53 ± 0.03	0.53 ± 0.07	-0.0 ± 1.4	-1.4 ± 0.6	1.2 ± 1.2
Gl51	1.10 ± 0.37	1.15 ± 0.06	0.37 ± 0.03	0.90 ± 0.06	-2.1 ± 2.2	-1.2 ± 0.8	-1.0 ± 1.8
G1866	1.59 ± 0.42	1.09 ± 0.12	0.46 ± 0.03	1.53 ± 0.12	-1.3 ± 2.8	-1.4 ± 1.1	0.0 ± 2.3
GJ1111	1.69 ± 0.56	0.87 ± 0.23	0.89 ± 0.07	1.28 ± 0.10	-3.0 ± 4.6	-3.9 ± 1.8	0.6 ± 3.7
Gl644C	1.97 ± 0.64	1.34 ± 0.24	0.66 ± 0.09	1.94 ± 0.19	-4.0 ± 5.6	-3.5 ± 2.2	-0.7 ± 4.5
G1752B	2.22 ± 0.71	1.38 ± 0.27	0.62 ± 0.11	2.86 ± 0.23	-3.0 ± 6.6	-3.3 ± 2.7	0.1 ± 5.1
LP412-31	2.38 ± 0.79	1.22 ± 0.18	1.23 ± 0.07	2.41 ± 0.18	-5.8 ± 7.8	-5.5 ± 3.1	-0.7 ± 6.1
LHS2065	2.42 ± 0.69	1.28 ± 0.25	1.27 ± 0.11	3.37 ± 0.18	-3.0 ± 7.1	-4.1 ± 2.8	0.9 ± 5.7
LHS2924	3.05 ± 1.01	0.81 ± 0.36	1.19 ± 0.33	2.02 ± 0.21	-8.1 ± 9.3	-3.8 ± 3.9	-4.6 ± 6.8
LP944-20	1.21 ± 0.75	1.31 ± 0.32	1.18 ± 0.27	2.47 ± 0.29	-3.2 ± 8.5	-3.6 ± 3.6	0.1 ± 6.4
BRIB0021	3.39 ± 0.53	2.69 ± 0.45	1.01 ± 0.31	2.46 ± 0.20	-19.5 ± 8.3	-10.8 ± 3.0	-9.5 ± 7.0

Table 6. EW of the *K*-band indices. Part 1.

Star IDs	Fe1	$\mathrm{Br}\delta$	Ca1	Fe23	Si	Ca2	Ca3	Ca4
				Supergiant				
HD 007927	0.30 ± 0.10	4.69 ± 0.20	0.48 ± 0.09	1.05 ± 0.14	1.17 ± 0.14	0.78 ± 0.05	0.87 ± 0.10	0.67 ± 0.07
HD 006130	0.60 ± 0.15	6.09 ± 0.30	0.81 ± 0.13	1.03 ± 0.21	1.35 ± 0.20	0.85 ± 0.08	0.82 ± 0.16	0.65 ± 0.11
HD 135153	2.03 ± 0.52	6.58 ± 1.04	1.18 ± 0.47	1.89 ± 0.76	1.05 ± 0.74	0.18 ± 0.28	-0.09 ± 0.58	-0.33 ± 0.42
HD 164136	0.67 ± 0.13	5.13 ± 0.25	0.80 ± 0.11	0.62 ± 0.18	0.69 ± 0.17	0.32 ± 0.06	0.15 ± 0.13	0.06 ± 0.10
HD 182835	1.12 ± 0.17	5.75 ± 0.33	0.72 ± 0.15	0.87 ± 0.24	0.59 ± 0.23	0.25 ± 0.09	-0.13 ± 0.18	-0.12 ± 0.13
HD 213306	1.99 ± 0.37	5.25 ± 0.74	1.86 ± 0.33	2.60 ± 0.53	1.93 ± 0.52	1.13 ± 0.20	1.05 ± 0.41	0.68 ± 0.29
HD 201078	0.77 ± 0.21	5.99 ± 0.42	1.27 ± 0.19	1.34 ± 0.30	1.42 ± 0.29	0.70 ± 0.11	0.69 ± 0.23	0.48 ± 0.16
HD 51956	0.73 ± 0.08	3.75 ± 0.17	1.11 ± 0.07	1.06 ± 0.12	1.20 ± 0.12	1.06 ± 0.04	1.05 ± 0.09	0.82 ± 0.06
HD 185018	1.22 ± 0.17	4.13 ± 0.33	1.21 ± 0.15	1.18 ± 0.24	1.29 ± 0.23	1.12 ± 0.08	1.03 ± 0.18	1.00 ± 0.13
HD 216219 HD 074395	0.68 ± 0.25	5.04 ± 0.49	1.25 ± 0.22	2.10 ± 0.35	2.33 ± 0.34	1.38 ± 0.12	1.77 ± 0.26	1.46 ± 0.18
HD 042454	0.83 ± 0.09 1.20 ± 0.43	3.94 ± 0.19 3.70 ± 0.86	1.39 ± 0.08 1.67 ± 0.39	0.76 ± 0.13 2.26 ± 0.62	0.86 ± 0.13 1.78 ± 0.61	0.86 ± 0.05 1.42 ± 0.23	0.65 ± 0.10 1.45 ± 0.47	0.66 ± 0.07 1.08 ± 0.34
HD 042434 HD 202314	1.20 ± 0.43 1.27 ± 0.23	2.83 ± 0.45	1.07 ± 0.39 1.18 ± 0.20	0.63 ± 0.32	1.78 ± 0.01 1.14 ± 0.31	1.42 ± 0.23 1.33 ± 0.12	1.43 ± 0.47 1.48 ± 0.24	1.08 ± 0.34 1.28 ± 0.17
HD 003421	1.27 ± 0.23 1.22 ± 0.18	3.25 ± 0.45	1.18 ± 0.20 1.21 ± 0.16	1.08 ± 0.32	1.00 ± 0.24	0.85 ± 0.12 0.85 ± 0.09	0.76 ± 0.19	0.58 ± 0.17 0.58 ± 0.14
HD 192713	1.22 ± 0.18 1.07 ± 0.12	2.80 ± 0.23	1.21 ± 0.10 1.16 ± 0.10	0.56 ± 0.25 0.56 ± 0.16	1.00 ± 0.24 1.07 ± 0.16	1.17 ± 0.06	0.70 ± 0.19 1.19 ± 0.12	0.38 ± 0.14 1.14 ± 0.09
HD 176123	0.44 ± 0.12	2.74 ± 0.24	0.95 ± 0.10	0.90 ± 0.10 0.90 ± 0.17	1.04 ± 0.16 1.04 ± 0.16	0.89 ± 0.06	0.83 ± 0.13	0.78 ± 0.09
HD 179821	1.03 ± 0.26	5.02 ± 0.51	1.13 ± 0.23	3.06 ± 0.17	2.93 ± 0.35	2.41 ± 0.13	2.61 ± 0.27	2.07 ± 0.09
HD 190113	1.35 ± 0.20 1.35 ± 0.11	3.29 ± 0.31	1.49 ± 0.29	1.35 ± 0.15	1.14 ± 0.15	1.40 ± 0.05	1.38 ± 0.11	1.26 ± 0.08
HD 161664	1.75 ± 0.32	4.21 ± 0.63	2.15 ± 0.28	2.53 ± 0.15 2.53 ± 0.45	2.15 ± 0.44	1.88 ± 0.16	1.83 ± 0.34	1.59 ± 0.05
HD 058367	0.97 ± 0.11	2.83 ± 0.21	1.21 ± 0.09	1.07 ± 0.15	1.54 ± 0.14	1.32 ± 0.05	1.43 ± 0.11	1.26 ± 0.08
HD 025877	0.79 ± 0.21	3.65 ± 0.40	1.23 ± 0.18	1.75 ± 0.29	1.65 ± 0.28	1.38 ± 0.10	1.41 ± 0.21	1.26 ± 0.05 1.26 ± 0.15
HD 208606	1.28 ± 0.33	4.00 ± 0.66	1.76 ± 0.29	2.95 ± 0.47	2.52 ± 0.46	2.14 ± 0.17	2.32 ± 0.36	1.82 ± 0.26
HD 122563	1.79 ± 0.29	3.11 ± 0.59	1.16 ± 0.26	2.93 ± 0.42	1.65 ± 0.41	1.14 ± 0.15	1.11 ± 0.32	0.65 ± 0.23
HD 165782	1.28 ± 0.27	4.19 ± 0.54	1.57 ± 0.24	0.98 ± 0.39	1.02 ± 0.38	1.21 ± 0.14	0.56 ± 0.29	0.22 ± 0.21
HD 044391	1.12 ± 0.10	2.45 ± 0.21	1.47 ± 0.09	1.01 ± 0.15	1.26 ± 0.14	1.43 ± 0.05	1.42 ± 0.11	1.25 ± 0.08
HD 179870	1.04 ± 0.43	3.76 ± 0.85	1.90 ± 0.38	3.76 ± 0.62	2.65 ± 0.61	2.01 ± 0.23	2.00 ± 0.48	1.57 ± 0.34
HD 164349	0.71 ± 0.30	2.14 ± 0.60	1.53 ± 0.27	1.99 ± 0.43	1.50 ± 0.42	1.55 ± 0.16	1.49 ± 0.32	1.27 ± 0.23
HD 091810	0.82 ± 0.11	3.53 ± 0.23	1.59 ± 0.10	1.15 ± 0.16	1.61 ± 0.16	1.73 ± 0.06	1.36 ± 0.12	1.38 ± 0.09
HD 212466	1.99 ± 0.37	5.06 ± 0.74	2.14 ± 0.33	2.56 ± 0.54	1.11 ± 0.53	1.20 ± 0.20	1.01 ± 0.41	0.25 ± 0.30
HD 023082	1.77 ± 0.28	2.89 ± 0.56	1.50 ± 0.25	3.00 ± 0.40	2.10 ± 0.39	2.22 ± 0.15	2.28 ± 0.30	1.62 ± 0.22
HD 063302	1.64 ± 0.25	3.86 ± 0.49	2.04 ± 0.22	2.10 ± 0.35	1.59 ± 0.35	1.96 ± 0.13	1.86 ± 0.27	1.60 ± 0.19
HD 187238	1.81 ± 0.19	3.23 ± 0.38	2.21 ± 0.17	2.21 ± 0.27	1.95 ± 0.27	2.17 ± 0.10	2.12 ± 0.21	1.68 ± 0.15
HD 185622	1.90 ± 0.21	3.21 ± 0.42	2.34 ± 0.18	2.17 ± 0.30	1.85 ± 0.29	2.24 ± 0.11	2.00 ± 0.23	1.53 ± 0.16
HD 201065	1.66 ± 0.20	2.68 ± 0.39	1.58 ± 0.17	2.18 ± 0.28	2.12 ± 0.27	2.52 ± 0.10	2.55 ± 0.21	2.04 ± 0.15
HD 207991	1.14 ± 0.12	2.85 ± 0.23	1.89 ± 0.10	1.92 ± 0.16	1.56 ± 0.16	2.02 ± 0.06	1.35 ± 0.12	1.52 ± 0.09
HD 216946	2.12 ± 0.19	2.51 ± 0.39	1.57 ± 0.17	1.41 ± 0.28	1.44 ± 0.27	2.09 ± 0.10	1.84 ± 0.21	1.33 ± 0.15
HD 236697	1.94 ± 0.39	3.91 ± 0.79	2.06 ± 0.35	3.71 ± 0.57	2.08 ± 0.57	2.28 ± 0.21	2.05 ± 0.44	1.32 ± 0.32
HD 181475	2.27 ± 0.41	3.65 ± 0.82	2.27 ± 0.36	3.58 ± 0.58	2.61 ± 0.57	2.63 ± 0.21	2.22 ± 0.44	2.09 ± 0.32
HD 014404	1.20 ± 0.20	3.32 ± 0.39	1.18 ± 0.17	3.03 ± 0.28	2.22 ± 0.27	2.26 ± 0.10	2.32 ± 0.21	1.86 ± 0.15
HD 039801	2.97 ± 0.19	4.04 ± 0.39	2.39 ± 0.17	3.60 ± 0.28	2.70 ± 0.27	2.61 ± 0.10	2.34 ± 0.21	1.88 ± 0.15
HD 035601 HD 206936	2.49 ± 0.35 2.59 ± 0.69	4.47 ± 0.71 5.25 ± 1.40	2.10 ± 0.31 3.43 ± 0.62	4.13 ± 0.50 6.41 ± 1.01	3.38 ± 0.50 3.99 ± 1.01	3.31 ± 0.19 3.26 ± 0.38	3.63 ± 0.38	2.74 ± 0.27 2.70 ± 0.57
HD 200936 HD 010465	2.39 ± 0.09 1.89 ± 0.32	5.25 ± 1.40 5.27 ± 0.64	3.43 ± 0.62 2.74 ± 0.28	6.41 ± 1.01 5.27 ± 0.45	3.99 ± 1.01 3.76 ± 0.45	3.20 ± 0.38 3.61 ± 0.17	3.55 ± 0.78 3.91 ± 0.34	2.70 ± 0.57 2.88 ± 0.25
HD 010403	1.89 ± 0.32 2.59 ± 0.62	3.27 ± 0.04 4.42 ± 1.26	3.29 ± 0.56	4.87 ± 0.43	3.70 ± 0.43 3.54 ± 0.90	3.39 ± 0.17	3.91 ± 0.34 3.13 ± 0.70	2.32 ± 0.23 2.32 ± 0.51
HD 023473 HD 040239	2.59 ± 0.62 1.60 ± 0.11	4.42 ± 1.26 3.23 ± 0.22	3.29 ± 0.36 2.47 ± 0.10	4.87 ± 0.91 2.49 ± 0.16	3.34 ± 0.90 2.20 ± 0.15	3.39 ± 0.34 2.82 ± 0.06	3.13 ± 0.70 2.30 ± 0.12	2.32 ± 0.31 1.94 ± 0.08
RWCyg	1.00 ± 0.11 1.92 ± 0.25	3.23 ± 0.22 3.66 ± 0.50	2.47 ± 0.10 1.68 ± 0.22	3.80 ± 0.16	2.20 ± 0.13 3.06 ± 0.35	2.82 ± 0.06 2.91 ± 0.13	2.30 ± 0.12 2.71 ± 0.27	1.79 ± 0.08 1.79 ± 0.20
HD 019058	1.92 ± 0.23 2.96 ± 0.23	3.00 ± 0.30 3.20 ± 0.46	1.08 ± 0.22 1.99 ± 0.20	2.62 ± 0.32	1.02 ± 0.32	2.91 ± 0.13 2.09 ± 0.12	0.59 ± 0.24	1.79 ± 0.20 1.16 ± 0.17
HD 156014	2.62 ± 0.23	3.02 ± 0.48	1.64 ± 0.21	3.02 ± 0.32 3.02 ± 0.34	1.85 ± 0.34	2.54 ± 0.12	1.80 ± 0.26	1.50 ± 0.17 1.50 ± 0.19
112 13001 1	2.02 ± 0.23	3.02 ± 0.10	1.01 ± 0.21	3.02 ± 0.3 i	1.05 ± 0.5 1	2.5 (± 0.12	1.00 ± 0.20	1.50 ± 0.17
				Giants				
HD 089025	0.32 ± 0.05	6.20 ± 0.10	0.88 ± 0.04	0.77 ± 0.07	0.92 ± 0.07	0.61 ± 0.03	0.54 ± 0.05	0.43 ± 0.04
HD 027397	0.84 ± 0.25	6.22 ± 0.49	0.95 ± 0.22	0.79 ± 0.35	0.44 ± 0.34	0.25 ± 0.13	0.12 ± 0.26	0.05 ± 0.19
HD 013174	0.98 ± 0.17	6.39 ± 0.33	1.13 ± 0.15	1.11 ± 0.24	0.99 ± 0.23	0.56 ± 0.09	0.43 ± 0.18	0.29 ± 0.13
HD 40535	0.33 ± 0.05	6.03 ± 0.11	0.87 ± 0.05	0.39 ± 0.08	0.75 ± 0.07	0.62 ± 0.03	0.54 ± 0.06	0.48 ± 0.04
HD 017918	0.51 ± 0.05	5.36 ± 0.10	0.80 ± 0.05	0.53 ± 0.07	0.72 ± 0.07	0.51 ± 0.03	0.46 ± 0.05	0.34 ± 0.04
HD 021770	0.38 ± 0.16	4.95 ± 0.32	0.60 ± 0.14	0.27 ± 0.23	0.23 ± 0.22	-0.00 ± 0.08	-0.17 ± 0.17	-0.22 ± 0.12
HD 075555	0.41 ± 0.19	5.48 ± 0.38	0.91 ± 0.17	0.63 ± 0.27	0.40 ± 0.27	0.14 ± 0.10	-0.16 ± 0.21	-0.16 ± 0.15
HD 186155	0.93 ± 0.20	6.05 ± 0.39	0.92 ± 0.17	0.47 ± 0.28	0.49 ± 0.27	0.11 ± 0.10	-0.05 ± 0.21	-0.09 ± 0.15
HD 218804	0.04 ± 0.22	5.88 ± 0.43	1.71 ± 0.19	2.52 ± 0.31	2.50 ± 0.30	2.19 ± 0.11	2.41 ± 0.23	1.68 ± 0.17
HD 160365	0.52 ± 0.06	4.55 ± 0.12	0.70 ± 0.05	0.19 ± 0.08	0.56 ± 0.08	0.47 ± 0.03	0.34 ± 0.06	0.35 ± 0.04
HD 011443	0.69 ± 0.19	4.84 ± 0.38	0.75 ± 0.17	0.56 ± 0.27	0.77 ± 0.26	0.43 ± 0.10	0.32 ± 0.20	0.26 ± 0.14
HD 124850	0.57 ± 0.18	4.23 ± 0.36	0.69 ± 0.16	0.03 ± 0.26	-0.07 ± 0.25	-0.34 ± 0.09	-0.77 ± 0.20	-0.71 ± 0.14
HD 220657	0.59 ± 0.16	5.23 ± 0.31	1.18 ± 0.14	1.75 ± 0.22	2.13 ± 0.22	1.40 ± 0.08	1.81 ± 0.16	1.63 ± 0.12
HD 006903	0.59 ± 0.15	4.03 ± 0.30	0.89 ± 0.13	1.02 ± 0.21	1.24 ± 0.21	0.73 ± 0.08	0.79 ± 0.16	0.71 ± 0.11
HD 219477	0.61 ± 0.11	3.64 ± 0.22	1.10 ± 0.10	0.65 ± 0.16	0.85 ± 0.15	0.38 ± 0.06	0.28 ± 0.12	0.28 ± 0.09
HD 126868	1.04 ± 0.27	3.78 ± 0.53	1.00 ± 0.24	0.51 ± 0.38	0.56 ± 0.37	0.14 ± 0.14	-0.10 ± 0.29	-0.06 ± 0.21

Table 6. continued.

Star IDs	Fe1	$\mathrm{Br}\delta$	Ca1	Fe23	Si	Ca2	Ca3	Ca4
HD 088639	0.56 ± 0.07	3.58 ± 0.14	1.12 ± 0.06	0.72 ± 0.10	0.97 ± 0.10	0.93 ± 0.04	0.73 ± 0.07	0.82 ± 0.05
HD 108477	0.17 ± 0.07	2.77 ± 0.14	0.92 ± 0.06	1.49 ± 0.10	1.52 ± 0.09	1.21 ± 0.03	1.37 ± 0.07	1.13 ± 0.05
HD 018474	0.80 ± 0.20	2.92 ± 0.39	1.20 ± 0.17	0.57 ± 0.28	0.36 ± 0.27	0.41 ± 0.10	0.05 ± 0.21	0.05 ± 0.15
HD 021018	0.64 ± 0.14	3.12 ± 0.28	1.12 ± 0.12	0.84 ± 0.20	0.92 ± 0.19	0.85 ± 0.07	0.68 ± 0.15	0.73 ± 0.11
HD 010697	0.43 ± 0.36	4.92 ± 0.72	1.79 ± 0.32	2.92 ± 0.52	2.83 ± 0.51	1.75 ± 0.19	1.88 ± 0.39	1.46 ± 0.28
HD 193896 HD 182694	0.42 ± 0.09	3.60 ± 0.18 3.82 ± 0.48	1.51 ± 0.08 1.49 ± 0.21	1.44 ± 0.13	1.62 ± 0.13 1.69 ± 0.33	1.32 ± 0.05 1.37 ± 0.12	1.43 ± 0.10 1.23 ± 0.25	1.28 ± 0.07 1.12 ± 0.18
HD 114946	1.13 ± 0.24 0.68 ± 0.13	3.82 ± 0.48 2.26 ± 0.26	0.50 ± 0.12	1.67 ± 0.34 0.27 ± 0.19	0.26 ± 0.33	0.31 ± 0.12	0.02 ± 0.23 0.14	-0.08 ± 0.10
HD 020618	0.69 ± 0.13	2.70 ± 0.26 2.70 ± 0.36	0.85 ± 0.12	1.21 ± 0.25	1.46 ± 0.24	1.35 ± 0.07	1.46 ± 0.19	1.18 ± 0.13
HD 104979	1.88 ± 0.47	3.21 ± 0.93	1.36 ± 0.42	1.71 ± 0.67	1.04 ± 0.66	0.67 ± 0.24	0.41 ± 0.51	0.15 ± 0.37
HD 135722	0.99 ± 0.22	3.20 ± 0.44	1.46 ± 0.19	1.48 ± 0.31	1.54 ± 0.30	1.22 ± 0.11	1.23 ± 0.23	1.00 ± 0.17
HD 094481	1.00 ± 0.12	2.86 ± 0.24	1.08 ± 0.11	1.01 ± 0.17	1.04 ± 0.17	0.79 ± 0.06	0.72 ± 0.13	0.63 ± 0.09
HD 170820	1.48 ± 0.26	3.64 ± 0.52	1.72 ± 0.23	2.44 ± 0.37	2.30 ± 0.36	1.87 ± 0.13	2.08 ± 0.28	1.87 ± 0.20
HD 222093	0.85 ± 0.09	3.00 ± 0.18	1.24 ± 0.08	1.24 ± 0.13	1.55 ± 0.12	1.55 ± 0.04	1.59 ± 0.09	1.32 ± 0.07
HD 100006 HD 036134	0.60 ± 0.10	2.96 ± 0.20	1.16 ± 0.09	0.84 ± 0.14	1.17 ± 0.14	1.01 ± 0.05 0.89 ± 0.08	0.74 ± 0.11	0.82 ± 0.08 0.41 ± 0.12
HD 030134 HD 025975	1.15 ± 0.16 0.72 ± 0.10	2.76 ± 0.31 2.83 ± 0.20	1.08 ± 0.14 0.98 ± 0.09	0.73 ± 0.22 0.95 ± 0.14	0.52 ± 0.22 1.31 ± 0.14	0.89 ± 0.08 1.31 ± 0.05	0.46 ± 0.17 1.12 ± 0.10	0.41 ± 0.12 0.90 ± 0.07
HD 142091	0.72 ± 0.10 0.98 ± 0.26	3.30 ± 0.20	1.31 ± 0.23	1.19 ± 0.37	1.15 ± 0.14 1.15 ± 0.36	1.08 ± 0.03	0.73 ± 0.27	0.46 ± 0.20
HD 165438	0.09 ± 0.26	3.35 ± 0.51	1.15 ± 0.23	1.62 ± 0.36	2.16 ± 0.35	1.71 ± 0.13	1.84 ± 0.27	1.40 ± 0.19
HD 124897	0.47 ± 0.07	2.27 ± 0.13	1.42 ± 0.06	1.13 ± 0.09	1.03 ± 0.09	1.17 ± 0.03	0.82 ± 0.07	0.71 ± 0.05
HD 002901	2.31 ± 0.16	1.81 ± 0.33	1.40 ± 0.15	0.66 ± 0.24	0.16 ± 0.23	0.87 ± 0.08	0.29 ± 0.18	0.17 ± 0.13
HD 132935	0.93 ± 0.14	2.99 ± 0.27	1.94 ± 0.12	1.82 ± 0.19	1.84 ± 0.19	2.09 ± 0.07	2.05 ± 0.14	1.73 ± 0.10
HD 137759	1.35 ± 0.24	2.82 ± 0.48	1.22 ± 0.21	1.39 ± 0.34	1.67 ± 0.33	1.87 ± 0.12	1.63 ± 0.25	1.33 ± 0.18
HD 035620	2.01 ± 0.21	4.09 ± 0.42	2.40 ± 0.18	3.10 ± 0.30	2.79 ± 0.29	2.84 ± 0.11	2.68 ± 0.22	2.33 ± 0.16
HD 178208 HD 221246	1.74 ± 0.18 1.22 ± 0.09	2.76 ± 0.36 2.46 ± 0.17	1.88 ± 0.16 1.89 ± 0.07	1.97 ± 0.25 1.25 ± 0.12	2.43 ± 0.25 1.70 ± 0.12	2.50 ± 0.09 2.13 ± 0.04	2.46 ± 0.19 1.99 ± 0.09	1.96 ± 0.14 1.65 ± 0.06
HD 114960	2.08 ± 0.17	2.40 ± 0.17 2.14 ± 0.34	0.99 ± 0.07 0.99 ± 0.15	0.61 ± 0.12	0.86 ± 0.12	1.85 ± 0.04	1.25 ± 0.09 1.25 ± 0.18	0.86 ± 0.13
HD 099998	1.87 ± 0.13	2.49 ± 0.27	1.66 ± 0.12	1.91 ± 0.19	0.92 ± 0.19	1.68 ± 0.07	1.16 ± 0.14	0.97 ± 0.10
HD 181596	0.59 ± 0.17	3.39 ± 0.33	1.71 ± 0.15	2.96 ± 0.24	2.29 ± 0.23	2.53 ± 0.09	2.47 ± 0.18	1.80 ± 0.13
HD 120477	2.04 ± 0.20	3.03 ± 0.40	1.98 ± 0.18	1.66 ± 0.29	1.18 ± 0.28	1.89 ± 0.10	1.14 ± 0.22	0.77 ± 0.16
HD 003346	0.98 ± 0.16	2.65 ± 0.32	1.16 ± 0.14	1.73 ± 0.22	1.14 ± 0.22	1.61 ± 0.08	1.38 ± 0.16	1.02 ± 0.12
HD 194193	1.26 ± 0.17	3.00 ± 0.33	2.20 ± 0.15	2.40 ± 0.24	1.92 ± 0.23	2.40 ± 0.09	1.94 ± 0.18	1.68 ± 0.13
HD 213893	0.79 ± 0.08	3.25 ± 0.17	1.79 ± 0.07	1.97 ± 0.12	1.42 ± 0.11	2.10 ± 0.04	1.86 ± 0.09	1.49 ± 0.06
HD 204724 HD 219734	2.21 ± 0.22 2.29 ± 0.21	2.93 ± 0.45 3.05 ± 0.43	1.85 ± 0.20 2.28 ± 0.19	1.53 ± 0.32 2.24 ± 0.30	1.74 ± 0.31 1.71 ± 0.29	2.64 ± 0.12 2.41 ± 0.11	2.07 ± 0.24 1.98 ± 0.23	1.58 ± 0.17 1.55 ± 0.16
HD 120052	0.71 ± 0.21	2.67 ± 0.43	2.28 ± 0.19 2.58 ± 0.27	2.24 ± 0.30 2.70 ± 0.44	1.71 ± 0.29 1.92 ± 0.43	2.41 ± 0.11 2.51 ± 0.16	1.79 ± 0.23 1.79 ± 0.33	1.67 ± 0.10 1.67 ± 0.24
HD 039045	2.58 ± 0.25	3.11 ± 0.50	1.73 ± 0.22	2.58 ± 0.36	2.03 ± 0.35	2.84 ± 0.13	2.37 ± 0.27	1.76 ± 0.19
HD 028487	2.39 ± 0.17	2.60 ± 0.34	1.99 ± 0.15	3.05 ± 0.24	1.60 ± 0.24	2.13 ± 0.09	1.23 ± 0.18	1.22 ± 0.13
HD 028487	2.39 ± 0.17	2.59 ± 0.34	1.98 ± 0.15	3.04 ± 0.24	1.60 ± 0.24	2.13 ± 0.09	1.22 ± 0.18	1.22 ± 0.13
HD 214665	3.07 ± 0.56	4.32 ± 1.14	2.99 ± 0.51	4.16 ± 0.82	2.26 ± 0.81	1.98 ± 0.30	1.06 ± 0.63	0.87 ± 0.46
HD 027598	3.13 ± 0.26	2.37 ± 0.53	1.52 ± 0.24	1.50 ± 0.38	0.68 ± 0.37	1.89 ± 0.14	0.90 ± 0.28	0.65 ± 0.20
HD 004408	2.43 ± 0.21	2.48 ± 0.41	1.76 ± 0.18	1.18 ± 0.30	0.54 ± 0.29	1.55 ± 0.11	0.24 ± 0.22	0.23 ± 0.16
HD 204585 HD 014386	1.65 ± 0.21 3.30 ± 0.67	4.48 ± 0.43 3.55 ± 1.33	1.84 ± 0.19 0.85 ± 0.61	2.75 ± 0.30 5.53 ± 0.94	2.17 ± 0.30 2.63 ± 0.93	2.94 ± 0.11 1.49 ± 0.35	2.37 ± 0.23 3.78 ± 0.70	1.66 ± 0.16 2.51 ± 0.51
HD 175865	1.55 ± 0.07	4.15 ± 0.54	0.85 ± 0.01 1.75 ± 0.24	3.62 ± 0.34 3.62 ± 0.39	2.03 ± 0.93 2.04 ± 0.38	2.67 ± 0.33	1.93 ± 0.70	1.31 ± 0.31 1.31 ± 0.21
HD 094705	2.51 ± 0.19	3.73 ± 0.39	2.63 ± 0.17	4.68 ± 0.28	2.91 ± 0.27	2.96 ± 0.10	1.60 ± 0.30	1.70 ± 0.21 1.70 ± 0.15
HD 018191	1.61 ± 0.15	2.91 ± 0.29	2.41 ± 0.13	3.10 ± 0.21	1.75 ± 0.20	2.84 ± 0.07	1.87 ± 0.15	1.81 ± 0.11
HD 196610	1.35 ± 0.34	4.63 ± 0.67	2.01 ± 0.30	4.81 ± 0.47	3.36 ± 0.47	3.01 ± 0.17	2.25 ± 0.36	1.94 ± 0.26
HD 108849	3.07 ± 0.49	2.24 ± 0.98	1.23 ± 0.44	6.89 ± 0.69	2.79 ± 0.68	1.65 ± 0.25	0.45 ± 0.52	1.51 ± 0.37
HD 207076	2.66 ± 0.40	5.04 ± 0.81	1.68 ± 0.36	6.89 ± 0.58	4.01 ± 0.58	3.00 ± 0.22	2.23 ± 0.45	2.10 ± 0.32
				Dwarfs				
HD 108519	0.51 ± 0.06	6.68 ± 0.12	0.99 ± 0.05	1.08 ± 0.09	1.03 ± 0.08	0.65 ± 0.03	0.63 ± 0.07	0.47 ± 0.05
HD 213135	1.23 ± 0.14	6.25 ± 0.12	1.07 ± 0.03	0.72 ± 0.19	0.39 ± 0.08	0.40 ± 0.03	0.05 ± 0.07 0.06 ± 0.15	-0.10 ± 0.10
HD 113139	-0.19 ± 0.16	6.48 ± 0.32	1.48 ± 0.14	1.66 ± 0.23	1.55 ± 0.22	1.02 ± 0.08	0.98 ± 0.17	0.80 ± 0.12
HD 026015	0.50 ± 0.07	5.97 ± 0.13	1.09 ± 0.06	0.70 ± 0.09	0.82 ± 0.09	0.46 ± 0.03	0.43 ± 0.07	0.36 ± 0.05
HD 016232	1.11 ± 0.17	4.37 ± 0.32	0.85 ± 0.14	0.25 ± 0.23	0.89 ± 0.22	0.86 ± 0.08	1.03 ± 0.17	0.74 ± 0.12
HD 087822	0.47 ± 0.09	4.78 ± 0.17	0.80 ± 0.08	0.24 ± 0.12	0.65 ± 0.12	0.26 ± 0.04	0.28 ± 0.09	0.24 ± 0.07
HD 027524	0.67 ± 0.09	5.57 ± 0.18	1.04 ± 0.08	1.06 ± 0.13	1.38 ± 0.13	0.69 ± 0.05	0.73 ± 0.10	0.59 ± 0.07
HD 215648 HD 126660	0.77 ± 0.30 0.51 ± 0.08	5.85 ± 0.60 4.52 ± 0.17	1.54 ± 0.27 1.01 ± 0.07	2.63 ± 0.43 0.76 ± 0.12	2.46 ± 0.42 1.18 ± 0.12	1.29 ± 0.15 0.60 ± 0.04	1.77 ± 0.32 0.64 ± 0.09	1.40 ± 0.23 0.54 ± 0.06
HD 120000 HD 219623	0.31 ± 0.08 0.92 ± 0.23	4.32 ± 0.17 4.71 ± 0.46	0.93 ± 0.07	0.76 ± 0.12 1.01 ± 0.33	1.18 ± 0.12 1.04 ± 0.32	0.00 ± 0.04 0.36 ± 0.12	0.04 ± 0.09 0.31 ± 0.25	0.34 ± 0.06 0.22 ± 0.18
HD 165908	1.27 ± 0.35	5.28 ± 0.69	1.23 ± 0.31	2.05 ± 0.49	1.04 ± 0.32 1.27 ± 0.48	0.50 ± 0.12 0.50 ± 0.18	0.57 ± 0.23 0.57 ± 0.38	0.22 ± 0.13 0.21 ± 0.27
HD 102870	0.58 ± 0.10	4.66 ± 0.19	1.10 ± 0.09	0.88 ± 0.14	1.41 ± 0.13	0.94 ± 0.05	1.12 ± 0.10	1.02 ± 0.07
HD 027383	0.64 ± 0.17	4.60 ± 0.34	1.32 ± 0.15	1.22 ± 0.25	1.47 ± 0.24	0.99 ± 0.09	0.99 ± 0.18	0.88 ± 0.13
HD 114710	1.31 ± 0.21	5.03 ± 0.41	1.54 ± 0.18	2.20 ± 0.29	1.61 ± 0.29	0.90 ± 0.11	0.79 ± 0.22	0.62 ± 0.16
HD 176051	0.67 ± 0.08	4.65 ± 0.16	1.16 ± 0.07	1.03 ± 0.11	1.34 ± 0.11	1.21 ± 0.04	1.21 ± 0.08	0.92 ± 0.06

Table 6. continued.

Star IDs	Fe1	$\mathrm{Br}\delta$	Cal	Fe23	Si	Ca2	Ca3	Ca4
HD 109358	0.60 ± 0.08	3.74 ± 0.15	0.83 ± 0.07	0.32 ± 0.11	0.84 ± 0.10	0.47 ± 0.04	0.34 ± 0.08	0.39 ± 0.06
HD 095128	0.51 ± 0.12	3.31 ± 0.23	0.94 ± 0.10	0.34 ± 0.17	0.50 ± 0.16	0.19 ± 0.06	-0.04 ± 0.13	0.02 ± 0.09
HD 020619	0.73 ± 0.14	2.95 ± 0.28	0.76 ± 0.12	-0.12 ± 0.20	-0.02 ± 0.19	0.06 ± 0.07	-0.31 ± 0.15	-0.21 ± 0.11
HD 010307	1.41 ± 0.34	4.46 ± 0.67	1.31 ± 0.30	1.52 ± 0.49	1.10 ± 0.47	0.66 ± 0.18	0.35 ± 0.37	0.36 ± 0.26
HD 076151	1.04 ± 0.22	4.26 ± 0.44	1.52 ± 0.20	1.67 ± 0.31	2.00 ± 0.31	1.39 ± 0.11	1.63 ± 0.23	1.29 ± 0.17
HD 214850	0.84 ± 0.18	3.73 ± 0.36	1.27 ± 0.16	1.97 ± 0.26	1.67 ± 0.25	1.01 ± 0.09	1.05 ± 0.19	0.80 ± 0.14
HD 165185	1.04 ± 0.29	4.43 ± 0.56	1.34 ± 0.25	1.77 ± 0.40	2.20 ± 0.39	1.52 ± 0.14	1.87 ± 0.30	1.50 ± 0.21
HD 115617	0.86 ± 0.15	3.53 ± 0.31	1.18 ± 0.14	0.51 ± 0.22	0.85 ± 0.21	0.51 ± 0.08	0.32 ± 0.16	0.24 ± 0.12
HD 101501	1.16 ± 0.20	3.20 ± 0.39	1.36 ± 0.17	1.16 ± 0.28	1.31 ± 0.27	1.04 ± 0.10	0.85 ± 0.21	0.66 ± 0.15
HD 075732	1.08 ± 0.14	3.71 ± 0.28	1.53 ± 0.12	1.34 ± 0.20	2.58 ± 0.19	2.29 ± 0.07	2.52 ± 0.15	2.04 ± 0.10
HD 145675	1.45 ± 0.29	3.44 ± 0.58	1.51 ± 0.26	0.85 ± 0.42	2.02 ± 0.40	1.52 ± 0.15	1.34 ± 0.31	1.14 ± 0.22
HD 010476	0.16 ± 0.18	3.28 ± 0.35	1.33 ± 0.15	1.95 ± 0.25	2.36 ± 0.24	2.15 ± 0.09	2.35 ± 0.18	1.77 ± 0.13
HD 003765	1.41 ± 0.40	5.04 ± 0.80	2.28 ± 0.35	3.51 ± 0.57	3.25 ± 0.56	3.16 ± 0.21	3.37 ± 0.43	2.23 ± 0.31
HD 219134	1.10 ± 0.32	4.26 ± 0.63	1.76 ± 0.28	2.13 ± 0.45	2.50 ± 0.44	3.10 ± 0.16	2.84 ± 0.33	1.80 ± 0.24
HD 045977	1.86 ± 0.09	4.24 ± 0.18	2.58 ± 0.08	1.56 ± 0.12	1.89 ± 0.12	4.05 ± 0.04	3.21 ± 0.09	2.08 ± 0.07
HD 207991	1.13 ± 0.12	2.85 ± 0.23	1.88 ± 0.10	1.90 ± 0.16	1.55 ± 0.16	2.02 ± 0.06	1.34 ± 0.12	1.51 ± 0.09
HD 036003	1.85 ± 0.08	4.90 ± 0.15	2.15 ± 0.07	1.47 ± 0.11	1.95 ± 0.10	4.06 ± 0.04	3.47 ± 0.08	1.95 ± 0.06
HD 201092	2.03 ± 0.89	8.21 ± 1.77	3.90 ± 0.79	5.73 ± 1.30	3.69 ± 1.29	6.61 ± 0.48	6.15 ± 0.98	2.94 ± 0.73
HD 237903	3.04 ± 0.06	5.79 ± 0.12	3.06 ± 0.05	1.18 ± 0.08	1.62 ± 0.08	5.86 ± 0.03	4.80 ± 0.06	2.50 ± 0.04
HD 019305	2.22 ± 0.11	6.19 ± 0.23	2.59 ± 0.10	0.99 ± 0.16	1.09 ± 0.16	5.77 ± 0.06	4.68 ± 0.12	1.97 ± 0.09
HD 209290	4.17 ± 0.09	6.95 ± 0.18	3.97 ± 0.08	1.36 ± 0.13	1.33 ± 0.12	7.12 ± 0.04	5.48 ± 0.09	2.66 ± 0.07
HD 036395	5.09 ± 0.10	8.28 ± 0.21	4.42 ± 0.09	1.26 ± 0.15	1.50 ± 0.15	8.15 ± 0.05	6.72 ± 0.11	3.14 ± 0.08
HD 042581	4.19 ± 0.07	7.04 ± 0.14	4.18 ± 0.06	1.50 ± 0.10	1.57 ± 0.09	7.39 ± 0.03	6.04 ± 0.07	3.05 ± 0.05
G1806	3.72 ± 0.26	6.75 ± 0.53	3.76 ± 0.23	2.49 ± 0.38	2.01 ± 0.37	7.03 ± 0.14	6.13 ± 0.27	3.29 ± 0.20
HD 095735	4.25 ± 0.40	5.74 ± 0.81	3.31 ± 0.36	2.85 ± 0.58	1.65 ± 0.57	5.56 ± 0.21	4.85 ± 0.43	2.38 ± 0.31
Gl381	4.33 ± 0.24	5.80 ± 0.49	3.32 ± 0.21	0.32 ± 0.35	0.14 ± 0.34	5.63 ± 0.12	4.34 ± 0.25	1.79 ± 0.18
G1581	4.35 ± 0.18	5.89 ± 0.36	3.42 ± 0.16	-0.24 ± 0.26	0.13 ± 0.25	5.63 ± 0.09	4.18 ± 0.19	1.75 ± 0.13
Gl213	2.97 ± 0.22	4.49 ± 0.44	2.24 ± 0.19	-0.80 ± 0.32	0.30 ± 0.30	4.42 ± 0.11	3.43 ± 0.23	1.54 ± 0.16
G1299	2.77 ± 0.29	3.86 ± 0.59	2.13 ± 0.26	-0.17 ± 0.42	0.27 ± 0.40	4.10 ± 0.15	3.12 ± 0.30	1.34 ± 0.22
Gl268AB	4.46 ± 0.16	5.29 ± 0.32	3.15 ± 0.14	-0.38 ± 0.23	0.15 ± 0.22	5.32 ± 0.08	3.94 ± 0.17	1.68 ± 0.12
Gl51	3.68 ± 0.33	6.27 ± 0.67	3.66 ± 0.29	0.13 ± 0.48	0.45 ± 0.47	5.96 ± 0.17	5.09 ± 0.35	2.26 ± 0.25
G1866	3.60 ± 0.23	5.70 ± 0.46	2.79 ± 0.20	-0.62 ± 0.33	0.24 ± 0.32	5.36 ± 0.12	3.84 ± 0.24	1.27 ± 0.18
GJ1111	3.84 ± 0.38	7.17 ± 0.77	3.83 ± 0.33	3.00 ± 0.55	2.68 ± 0.54	6.73 ± 0.20	5.67 ± 0.40	2.73 ± 0.29
G1644C	3.66 ± 0.46	5.31 ± 0.93	3.44 ± 0.41	1.29 ± 0.67	1.52 ± 0.64	5.72 ± 0.24	3.86 ± 0.48	2.06 ± 0.35
G1752B	5.40 ± 0.42	4.95 ± 0.85	3.09 ± 0.37	0.40 ± 0.61	0.63 ± 0.59	5.44 ± 0.22	3.54 ± 0.44	1.40 ± 0.32
LP412-31	4.23 ± 0.35	6.17 ± 0.71	4.00 ± 0.31	0.97 ± 0.51	0.90 ± 0.49	6.85 ± 0.18	5.31 ± 0.36	1.83 ± 0.26
LHS2065	3.62 ± 0.45	6.49 ± 0.91	3.65 ± 0.40	1.44 ± 0.66	1.68 ± 0.64	6.44 ± 0.24	5.05 ± 0.48	1.91 ± 0.35
LHS2924	3.71 ± 0.53	5.28 ± 1.06	2.57 ± 0.47	-0.99 ± 0.77	0.44 ± 0.74	4.70 ± 0.27	3.25 ± 0.56	0.57 ± 0.41
LP944-20	4.14 ± 0.37	4.63 ± 0.75	2.76 ± 0.33	1.36 ± 0.54	0.92 ± 0.52	4.14 ± 0.19	2.79 ± 0.40	0.61 ± 0.29
BRIB0021	2.73 ± 0.54	5.68 ± 1.10	3.83 ± 0.48	4.32 ± 0.79	3.44 ± 0.78	6.46 ± 0.29	5.00 ± 0.59	2.60 ± 0.43

Table 7. EW of the *K*-band indices. Part 2.

$ D007977 - 0.27 \pm 0.04 - 0.18 \pm 0.06 - 0.26 \pm 0.15 - 0.05 \pm 0.06 \pm 0.07 \pm 0.01 - 0.06 \pm 0.03 - 0.04 \pm 0.03 - 0.11 \pm 0.08 \\ D101313 - 0.17 \pm 0.01 - 0.06 \pm 0.01 - 0.06 \pm 0.01 - 0.06 \pm 0.03 - 0.01 \pm 0.04 - 0.06 \pm 0.03 - 0.01 \pm 0.08 - 0.01 \pm 0.08 \\ D101313 - 0.07 - 0.01 \pm 0.00 \\ D101313 - 0.07 - 0.01 \pm 0.00 - 0.01 \pm 0.01 - 0.01 \pm 0.00 - 0.01 \pm 0.00 - 0.01 \\ D101313 - 0.07 - 0.01 \pm 0.00 - 0.01 - 0.01 - 0.01 + 0.00 - 0.01 \pm 0.00 \\ D101313 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 + 0.00 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D101301 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 - 0.01 \\ D1013$	Star IDs	Na _d	Ca _d	¹² CO	Mg1	Brγ	FeA	FeB	Mg2
	IID 007027	0.27 0.04	0.10 0.06	0.26 0.15	Supergiants	5.05 0.01	0.06 0.03	0.04.002	0.11 0.00
$ \begin{aligned} & \text{ID 15153} & 0.17 \pm 0.10 & -0.06 \pm 0.10 & 0.80 \pm 0.24 & 0.04 \pm 0.12 & 0.88 \pm 0.05 & -0.01 \pm 0.07 & -0.04 \pm 0.08 & 0.17 \pm 0.08 \\ & \text{ID 161350} & 0.19 \pm 0.06 & -0.08 \pm 0.01 & 0.09 \pm 0.025 & 0.05 \pm 0.01 & 5.01 \pm 0.05 & -0.00 \pm 0.04 & 0.01 \pm 0.08 & 0.17 \pm 0.08 \\ & \text{ID 21306} & 0.99 \pm 0.08 & 0.04 \pm 0.01 & 0.95 \pm 0.25 & 0.05 \pm 0.01 & 5.21 \pm 0.05 & -0.00 \pm 0.05 & 0.03 \pm 0.08 & 0.17 \pm 0.01 \\ & \text{ID 21078} & 0.09 \pm 0.08 & 0.04 \pm 0.01 & 0.03 \pm 0.03 & 0.03 \pm 0.03 & 0.03 \pm 0.03 & 0.03 \pm 0.03 & 0.07 \pm 0.09 \\ & \text{ID 21078} & 0.09 \pm 0.08 & 0.04 \pm 0.01 & 0.03 \pm 0.03 & 0.03 \pm 0.03 & 0.04 \pm 0.07 & 0.04 \pm 0.07 & 0.03 \pm 0.08 & 0.07 \pm 0.09 \\ & \text{ID 10108} & 0.08 \pm 0.07 & 0.03 \pm 0.01 & 0.03 \pm 0.03 & 0.04 \pm 0.07 & 0.04 \pm 0.07 & 0.02 \pm 0.08 & 0.07 \pm 0.09 \\ & \text{ID 10108} & 0.08 \pm 0.07 & 0.03 \pm 0.01 & 1.68 \pm 0.03 & 0.05 \pm 0.07 & 2.74 \pm 0.09 & 0.02 \pm 0.08 & 0.04 \pm 0.07 & 0.05 \pm 0.01 \\ & \text{ID 10108} & 0.08 \pm 0.07 & 0.03 \pm 0.01 & 1.68 \pm 0.03 & 0.05 \pm 0.07 & 2.74 \pm 0.09 & 0.02 \pm 0.00 & 0.03 \pm 0.00 & 0$									
HD 14136 $0.09 \pm 0.06 - 0.03 \pm 0.09 $ 0.41 $\pm 0.21 $ 0.07 $\pm 0.01 $ 5.02 $\pm 0.05 - 0.05 \pm 0.05 + 0.05 + 0.05 + 0.05 $ 0.07 $\pm 0.05 + 0.05 $ 0.07 $\pm 0.05 + 0.05 $ 0.07 $\pm 0.05 $ 0.07 $\pm 0.05 $ 0.09 $\pm 0.08 $ 0.09 $\pm 0.08 $ 0.06 $\pm 0.16 $ 0.32 $\pm 0.05 $ 0.05 $\pm 0.05 $ 0.06 $\pm 0.05 $ 0.07 $\pm 0.05 $ 0.07 $\pm 0.05 $ 0.09 $\pm 0.08 $ 0.06 $\pm 0.06 $ 0.07 $\pm 0.05 $ 0.09 $\pm 0.08 $ 0.07 $\pm 0.05 $ 0.09 $\pm 0.05 $ 0.07 $\pm 0.05 $ 0.09 $\pm 0.05 $ 0.07 $\pm 0.05 $ 0.09 $\pm 0.05 $ 0.07 $\pm 0.05 $ 0.05 $\pm 0.05 $ 0									
HD 18235 0.21 ± 0.06 -0.18 ± 0.10 0.95 ± 0.25 0.05 ± 0.10 5.21 ± 0.05 -0.05 ± 0.05 0.03 ± 0.05 0.77 ± 0.15 HD 21300 0.99 ± 0.08 0.40 ± 0.18 ± 0.10 0.25 ± 0.11 0.06 ± 0.16 0.53 ± 0.13 3.03 ± 0.10 0.19 ± 0.07 0.28 ± 0.08 0.67 ± 0.09 HD 201078 0.67 ± 0.11 0.06 ± 0.16 0.53 ± 0.16 1.68 ± 0.09 0.45 ± 0.09 0.27 ± 0.07 0.20 ± 0.08 0.67 ± 0.09 0.10 ± 0.10 0.10 ± 0.10 0.10 ± 0.10 0.10									
HD 210306 0.99 ± 0.08 0.40 ± 0.18 0.91 ± 0.44 0.63 ± 0.13 3.03 ± 0.10 0.19 ± 0.07 0.28 ± 0.08 0.67 ± 0.01 HD 210179 0.07 ± 0.07 0.07 ± 0.07 0.07 ± 0.07 HD 21056 1.04 ± 0.05 0.47 ± 0.16 0.01 ± 0.04 0.58 ± 0.09 0.27 ± 0.01 0.34 ± 0.06 0.25 ± 0.07 0.55 ± 0.14 HD 185018 0.88 ± 0.07 0.53 ± 0.16 1.06 ± 0.08 0.35 ± 0.04 0.05 0.47 ± 0.12 0.27 ± 0.07 0.32 ± 0.08 0.64 ± 0.06 HD 185018 0.889 ± 0.07 0.33 ± 0.14 0.32 ± 0.05 0.33 ± 0.14 0.33 ± 0.14 0.32 ± 0.05 0.33 ± 0.14 0.33 ± 0.14 0.32 ± 0.05 0.34 ± 0.05 0.32 ± 0.08 0.44 ± 0.06 0.44 ± 0.06 HD 1920214 1.27 ± 0.09 0.82 ± 0.07 0.35 ± 0.04 0.08 ± 0.08 $0.$									
$ \begin{aligned} & \text{HD 19596} & & 1.04 \pm 0.05 & 0.47 \pm 0.16 & -0.01 \pm 0.41 & 0.58 \pm 0.09 & 2.78 \pm 0.10 & 0.34 \pm 0.06 & 0.25 \pm 0.07 & 0.25 \pm 0.08 & 1.04 \pm 0.05 & 0.35 \pm 0.16 & 1.08 \pm 0.39 & 0.65 \pm 0.07 & 2.74 \pm 0.09 & 0.27 \pm 0.07 & 0.32 \pm 0.08 & 0.64 \pm 0.06 & 1.00 \pm 0.00 & 0.00 \pm 0.00 & 0.00 \pm 0.00 &$		0.99 ± 0.08	0.40 ± 0.18						
HD185018 0.80 ± 0.07 0.53 ± 0.16 1.68 ± 0.39 0.65 ± 0.07 2.76 ± 0.09 0.27 ± 0.07 0.32 ± 0.08 0.64 ± 0.06 HD216219 0.24 ± 0.05 0.33 ± 0.14 0.37 ± 0.35 0.45 ± 0.10 3.89 ± 0.06 0.03 ± 0.04 0.10 ± 0.05 0.49 ± 0.03 HD042351 1.15 ± 0.08 0.76 ± 0.18 3.52 ± 0.44 0.72 ± 0.10 3.89 ± 0.06 0.03 ± 0.04 0.10 ± 0.05 0.49 ± 0.03 HD042354 1.11 ± 0.07 0.68 ± 0.18 2.57 ± 0.45 0.70 ± 0.09 2.32 ± 0.10 0.32 ± 0.07 0.36 ± 0.07 0.36 ± 0.07 0.73 ± 0.09 HD042354 1.11 ± 0.07 0.68 ± 0.18 2.57 ± 0.45 0.70 ± 0.09 2.32 ± 0.10 0.32 ± 0.07 0.36 ± 0.07 0.36 ± 0.07 0.68 ± 0.18 0.57 ± 0.09 0.09 ± 0.09 0.23 ± 0.01 0.44 ± 0.07 0.34 ± 0.07 0.36 ± 0.07 0.68 ± 0.18 0.58 ± 0.06 0.65 ± 0.15 5.35 ± 0.36 0.64 ± 0.08 2.21 ± 0.07 0.32 ± 0.08 0.06 ± 0.08 0.08 ± 0.18 0.08 ± 0.19 1.03 ± 0.04 0.08 ± 0.08 0.08 ± 0.19 0.08 ± 0.08 0.08 ± 0.08 0.08 ± 0.08 0.08 ± 0.09 0.08 ± 0.09 0.08 ± 0.08 0.08 ± 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09 0.09									
$\begin{array}{c c c c c c c c c c c c c c c c c c c $									
$ \begin{aligned} & \text{HD} 162(219) & 0.24 \pm 0.05 & 0.33 \pm 0.14 & 0.37 \pm 0.35 & 0.44 \pm 0.10 & 3.89 \pm 0.06 & 0.03 \pm 0.04 & 0.10 \pm 0.05 & 0.49 \pm 0.03 \\ & \text{HD} 042(454) & 1.11 \pm 0.07 & 0.68 \pm 0.18 & 2.57 \pm 0.45 & 0.70 \pm 0.09 & 2.32 \pm 0.10 & 0.32 \pm 0.07 & 0.36 \pm 0.07 & 0.78 \pm 0.09 \\ & \text{HD} 042(454) & 1.11 \pm 0.07 & 0.68 \pm 0.18 & 2.57 \pm 0.45 & 0.70 \pm 0.09 & 2.32 \pm 0.10 & 0.32 \pm 0.07 & 0.36 \pm 0.07 & 0.68 \pm 0.14 \\ & \text{HD} 020(314) & 1.27 \pm 0.09 & 0.82 \pm 0.17 & 1.05 \pm 0.04 & 0.06 \pm 0.10 \\ & \text{HD} 030(321) & 0.68 \pm 0.06 & 0.56 \pm 0.15 & 5.35 \pm 0.36 & 0.64 \pm 0.08 & 2.21 \pm 0.07 & 0.23 \pm 0.05 & 0.23 \pm 0.06 & 0.64 \pm 0.08 \\ & \text{HD} 19(2713) & 1.36 \pm 0.10 & 0.85 \pm 0.19 & 1.04 \pm 0.03 & 0.06 \pm 0.01 & 0.08 \pm 0.01 \\ & \text{HD} 1792(1) & -2.14 \pm 0.10 & -0.12 \pm 0.07 & 1.17 \pm 0.17 & -0.08 \pm 0.06 & 0.02 \pm 0.01 & 0.03 \pm 0.03 & 0.03 \pm 0.01 & 0.03 \pm 0.01 \\ & \text{HD} 10131 & 1.67 \pm 0.10 & 0.96 \pm 0.19 & 10.33 \pm 0.45 & 0.76 \pm 0.14 & 1.78 \pm 0.09 & 0.58 \pm 0.06 & 0.55 \pm 0.07 & 0.73 \pm 0.06 \\ & \text{HD} 038867 & 1.11 \pm 0.07 & 0.74 \pm 0.18 & 8.02 \pm 0.43 & 0.70 \pm 0.10 & 2.16 \pm 0.08 & 0.05 \pm 0.05 & 0.03 \pm 0.05 & 0.05 \pm 0.05 & 0.03 \pm 0.05 & 0.05 & 0.05 \pm 0.05 & 0.05 \pm 0.05 & 0.05 \pm 0.05 & 0.05 \pm 0.05 & 0.0$									
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$									
$\begin{aligned} & \text{HD} 0242454 & 1.11 \pm 0.07 & 0.68 \pm 0.18 & 2.57 \pm 0.45 & 0.70 \pm 0.09 & 2.32 \pm 0.10 & 0.32 \pm 0.07 & 0.36 \pm 0.07 & 0.68 \pm 0.14 \\ & \text{HD} 0203421 & 0.68 \pm 0.06 & 0.56 \pm 0.15 & 5.35 \pm 0.36 & 0.64 \pm 0.08 & 2.21 \pm 0.07 & 0.23 \pm 0.05 & 0.23 \pm 0.06 & 0.64 \pm 0.08 \\ & \text{HD} 192713 & 1.36 \pm 0.10 & 0.85 \pm 0.19 & 10.43 \pm 0.45 & 0.62 \pm 0.12 & 1.79 \pm 0.09 & 0.23 \pm 0.05 & 0.33 \pm 0.05 & 0.73 \pm 0.06 \\ & \text{HD} 179612 & 0.89 \pm 0.07 & 0.53 \pm 0.15 & 3.88 \pm 0.38 & 0.66 \pm 0.09 & 2.75 \pm 0.09 & 0.23 \pm 0.05 & 0.33 \pm 0.05 & 0.73 \pm 0.06 \\ & \text{HD} 179612 & 1.55 \pm 0.10 & 0.96 \pm 0.19 & 10.33 \pm 0.45 & 0.76 \pm 0.14 & 1.78 \pm 0.09 & 0.58 \pm 0.06 & 0.55 \pm 0.07 & 0.77 \pm 0.06 \\ & \text{HD} 161664 & 1.35 \pm 0.09 & 0.87 \pm 0.16 & 13.13 \pm 0.37 & 0.99 \pm 0.13 & 2.02 \pm 0.03 & 0.40 \pm 0.06 & 0.55 \pm 0.07 & 0.77 \pm 0.06 \\ & \text{HD} 1058367 & 1.11 \pm 0.07 & 0.74 \pm 0.18 & 8.02 \pm 0.43 & 0.70 \pm 0.10 & 2.16 \pm 0.08 & 0.36 \pm 0.05 & 0.38 \pm 0.06 & 0.72 \pm 0.06 \\ & \text{HD} 025837 & 1.24 \pm 0.09 & 0.81 \pm 0.18 & 8.12 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.36 \pm 0.05 & 0.38 \pm 0.06 & 0.72 \pm 0.06 \\ & \text{HD} 1025860 & 2.05 \pm 0.15 & 1.20 \pm 0.23 & 19.89 \pm 0.51 & 0.76 \pm 0.23 & 1.55 \pm 0.10 & 0.64 \pm 0.10 & 0.52 \pm 0.11 & 0.92 \pm 0.09 \\ & \text{HD} 1025680 & -0.04 \pm 0.03 & 0.04 \pm 0.02 & -0.12 \pm 0.04 & 0.08 \pm 0.22 & 2.53 \pm 0.08 & -0.41 \pm 0.08 & -0.18 \pm 0.09 & 0.10 \pm 0.03 & 0.02 \pm 0.01 \\ & \text{HD} 105782 & -0.93 \pm 0.13 & -0.92 \pm 0.19 & 1.56 \pm 0.40 & 0.08 \pm 0.22 & 2.53 \pm 0.08 & -0.41 \pm 0.08 & -0.18 \pm 0.09 & 0.10 \pm 0.33 & 0.02 \pm 0.01 \\ & \text{HD} 1079870 & 1.35 \pm 0.10 & 0.84 \pm 0.20 & 10.07 \pm 0.47 & 0.84 \pm 0.12 & 1.34 \pm 0.08 & 0.47 \pm 0.08 & 0.04 \pm 0.00 & 0.34 \pm 0.02 & 0.02 \pm 0.01 & 0.04 \pm 0.00 &$									
$\begin{aligned} & \text{HD} 102314 & 1.77 \pm 0.09 & 0.82 \pm 0.17 & 10.53 \pm 0.41 & 0.69 \pm 0.10 & 1.58 \pm 0.06 & 0.41 \pm 0.07 & 0.43 \pm 0.07 & 0.76 \pm 0.05 \\ & \text{HD} 192713 & 1.36 \pm 0.10 & 0.85 \pm 0.19 & 10.43 \pm 0.45 & 0.62 \pm 0.12 & 1.79 \pm 0.09 & 0.48 \pm 0.08 & 0.43 \pm 0.09 & 0.85 \pm 0.07 \\ & \text{HD} 179821 & -2.14 \pm 0.10 & -0.12 \pm 0.07 & 1.17 \pm 0.17 & -0.08 \pm 0.06 \\ & \text{HD} 197821 & -2.14 \pm 0.10 & -0.12 \pm 0.07 & 1.17 \pm 0.17 & -0.08 \pm 0.06 \\ & \text{HD} 197821 & 1.67 \pm 0.10 & 0.96 \pm 0.19 & 10.33 \pm 0.45 & 0.62 \pm 0.00 & 2.75 \pm 0.09 & 0.32 \pm 0.05 & 0.05 \pm 0.05 & -0.08 \pm 0.13 \\ & \text{HD} 10161664 & 1.55 \pm 0.09 & 0.87 \pm 0.16 & 13.13 \pm 0.37 & 0.59 \pm 0.13 & 20.2 \pm 0.08 & 0.40 \pm 0.06 & 0.41 \pm 0.07 & 0.63 \pm 0.05 \\ & \text{HD} 1025877 & 1.12 \pm 0.09 & 0.81 \pm 0.18 & 8.18 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.36 \pm 0.05 & 0.38 \pm 0.06 & 0.72 \pm 0.06 \\ & \text{HD} 1025877 & 1.24 \pm 0.09 & 0.81 \pm 0.18 & 8.18 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.46 \pm 0.06 & 0.45 \pm 0.07 & 0.87 \pm 0.07 \\ & \text{HD} 1025860 & 2.05 \pm 0.15 & 1.20 \pm 0.23 & 1.98 \pm 0.51 & 1.07 \pm 0.22 & 1.11 \pm 0.07 & 0.08 \pm 0.12 \\ & \text{HD} 1025873 & -0.04 \pm 0.03 & 0.04 \pm 0.02 & -0.12 \pm 0.04 & -0.04 \pm 0.02 & 1.55 \pm 0.10 & 0.44 \pm 0.10 & 0.45 \pm 0.10 & 0.22 \pm 0.11 & 0.92 \pm 0.09 \\ & \text{HD} 122563 & -0.04 \pm 0.03 & 0.04 \pm 0.02 & -0.12 \pm 0.04 & -0.04 \pm 0.02 & 1.29 \pm 0.01 & -0.04 \pm 0.02 & 0.01 \pm 0.03 & 0.02 \pm 0.01 \\ & \text{HD} 1045782 & -0.93 \pm 0.13 & 1.09 \pm 0.23 & 1.03 \pm 0.94 & 0.08 \pm 0.22 & 2.55 \pm 0.08 & 0.45 \pm 0.05 & 0.25 \pm 0.11 & 0.92 \pm 0.09 \\ & \text{HD} 1165782 & -0.03 \pm 0.13 & 1.15 \pm 0.20 & 1.369 \pm 0.46 & 0.80 \pm 0.14 & 1.54 \pm 0.09 & 0.67 \pm 0.07 & 0.94 \pm 0.06 \\ & \text{HD} 1043491 & 1.73 \pm 0.11 & 1.15 \pm 0.20 & 13.69 \pm 0.46 & 0.80 \pm 0.14 & 1.54 \pm 0.09 & 0.67 \pm 0.07 & 0.94 \pm 0.06 \\ & \text{HD} 104391 & 1.97 \pm 0.14 & 1.10 \pm 0.20 & 1.00 \pm 0.04 \pm 0.00 & 0.94 \pm 0.06 & 0.94 \pm 0.06 \\ & \text{HD} 104391 & 1.97 \pm 0.14 & 1.10 \pm 0.20 & 1.01 \pm 0.04 & 0.09 \pm 0.06 & 0.94 \pm 0.0$									
$\begin{array}{c} \text{HD192713} & 1.36 \pm 0.10 & 0.85 \pm 0.07 & 10.43 \pm 0.45 & 0.62 \pm 0.02 & 2.75 \pm 0.09 & 0.48 \pm 0.08 & 0.43 \pm 0.09 & 0.85 \pm 0.07 \\ \text{HD179221} & -2.14 \pm 0.10 & -0.12 \pm 0.07 & 1.17 \pm 0.17 & -0.08 \pm 0.06 \\ \text{HD179231} & 1.67 \pm 0.10 & 0.96 \pm 0.19 & 10.33 \pm 0.45 & 0.76 \pm 0.014 & 1.78 \pm 0.09 \\ \text{HD161664} & 1.35 \pm 0.09 & 0.87 \pm 0.16 & 13.13 \pm 0.37 & 0.59 \pm 0.13 & 20.2 \pm 0.08 & 0.40 \pm 0.06 & 0.41 \pm 0.07 & 0.63 \pm 0.05 \\ \text{HD052877} & 1.14 \pm 0.09 & 0.81 \pm 0.18 & 81.8 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.36 \pm 0.05 & 0.38 \pm 0.06 & 0.72 \pm 0.06 \\ \text{HD025877} & 1.24 \pm 0.09 & 0.81 \pm 0.18 & 81.8 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.36 \pm 0.05 & 0.25 \pm 0.017 & 0.72 \pm 0.06 \\ \text{HD025877} & 1.24 \pm 0.09 & 0.81 \pm 0.18 & 81.8 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.46 \pm 0.06 & 0.45 \pm 0.07 & 0.87 \pm 0.07 \\ \text{HD122563} & -0.04 \pm 0.03 & 0.04 \pm 0.02 & -0.12 \pm 0.04 & -0.04 \pm 0.02 & 1.55 \pm 0.10 & 0.04 \pm 0.10 & 0.25 \pm 0.01 & 0.92 \pm 0.09 \\ \text{HD122563} & -0.04 \pm 0.03 & 0.04 \pm 0.02 & -0.12 \pm 0.04 & -0.04 \pm 0.02 & 1.29 \pm 0.01 & -0.01 \pm 0.02 & 0.01 \pm 0.03 & 0.02 \pm 0.01 \\ \text{HD167872} & -0.93 \pm 0.13 & 1.59 \pm 0.10 & 0.08 \pm 0.22 & 2.53 \pm 0.08 & 0.46 \pm 0.06 & 0.45 \pm 0.07 & 0.85 \pm 0.06 \\ \text{HD164391} & 1.73 \pm 0.11 & 1.15 \pm 0.20 & 13.69 \pm 0.46 & 0.80 \pm 0.14 & 1.54 \pm 0.09 & 0.67 \pm 0.07 & 0.94 \pm 0.08 & 8.80 \pm 0.05 \\ \text{HD164349} & 1.59 \pm 0.12 & 1.11 \pm 0.20 & 14.01 \pm 0.47 & 0.69 \pm 0.16 & 0.94 \pm 0.07 & 0.51 \pm 0.05 & 0.39 \pm 0.06 & 0.87 \pm 0.07 \\ \text{HD163782} & 3.32 \pm 0.19 & 1.81 \pm 0.22 & 2.29 \pm 0.48 & 0.99 \pm 0.17 & 1.99 \pm 0.08 & 0.88 \pm 0.05 & 0.35 \pm 0.06 & 1.11 \pm 0.05 \\ \text{HD163393} & 3.32 \pm 0.19 & 1.81 \pm 0.22 & 2.29 \pm 0.48 & 0.09 \pm 0.17 & 1.99 \pm 0.08 & 0.88 \pm 0.05 & 0.35 \pm 0.06 & 0.35 \pm 0.06 & 1.11 \pm 0.05 \\ \text{HD163782} & 3.39 \pm 0.12 & 1.11 \pm 0.20 & 1.64 \pm 0.49 & 0.79 \pm 0.27 & 1.55 \pm 0.10 & 0.31 \pm 0.11 & 0.35 \pm 0.09 & 0.35 \pm 0.06 & 0.35 $									
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$\begin{array}{c} \mathrm{HD} 025877 & 1.24 \pm 0.09 & 0.81 \pm 0.18 & 8.18 \pm 0.44 & 0.73 \pm 0.09 & 1.86 \pm 0.08 & 0.46 \pm 0.06 & 0.45 \pm 0.07 & 0.87 \pm 0.07 \\ \mathrm{HD} 122563 & -0.04 \pm 0.03 & 0.04 \pm 0.02 & -0.12 \pm 0.04 & -0.04 \pm 0.02 & 1.29 \pm 0.01 & -0.01 \pm 0.02 & 0.01 \pm 0.03 & 0.02 \pm 0.01 \\ \mathrm{HD} 165782 & -0.93 \pm 0.13 & -0.92 \pm 0.19 & 21.68 \pm 0.40 & 0.08 \pm 0.22 & 2.53 \pm 0.08 & -0.41 \pm 0.08 & -0.18 \pm 0.09 & 0.02 \pm 0.01 \\ \mathrm{HD} 165782 & -0.93 \pm 0.13 & -0.92 \pm 0.19 & 21.68 \pm 0.40 & 0.08 \pm 0.22 & 2.53 \pm 0.08 & -0.41 \pm 0.08 & -0.18 \pm 0.09 & 0.07 \pm 0.07 & 0.49 \pm 0.08 \\ \mathrm{HD} 179870 & 1.35 \pm 0.10 & 0.84 \pm 0.20 & 10.07 \pm 0.47 & 0.08 \pm 0.11 & 1.54 \pm 0.09 & 0.67 \pm 0.07 & 0.49 \pm 0.06 & 0.88 \pm 0.05 \\ \mathrm{HD} 191810 & 1.97 \pm 0.14 & 1.09 \pm 0.24 & 12.10 \pm 0.56 & 0.99 \pm 0.17 & 1.99 \pm 0.08 & 0.58 \pm 0.06 & 0.33 \pm 0.06 & 1.11 \pm 0.02 \\ \mathrm{HD} 212466 & 1.38 \pm 0.16 & 0.33 \pm 0.22 & 16.48 \pm 0.49 & 0.79 \pm 0.27 & 1.55 \pm 0.10 & 0.31 \pm 0.11 & 0.23 \pm 0.12 & 0.21 \pm 0.33 \\ \mathrm{HD} 023082 & 3.32 \pm 0.19 & 1.81 \pm 0.22 & 22.90 \pm 0.48 & 0.76 \pm 0.29 & 1.04 \pm 0.10 & 1.03 \pm 0.09 & 0.52 \pm 0.10 & 0.88 \pm 0.05 \\ \mathrm{HD} 187238 & 3.90 \pm 0.24 & 2.19 \pm 0.24 & 26.10 \pm 0.52 & 0.89 \pm 0.35 & 0.89 \pm 0.11 & 1.31 \pm 0.13 & 0.07 \pm 0.14 & 0.08 \pm 0.14 \\ \mathrm{HD} 210665 & 3.48 \pm 0.19 & 1.93 \pm 0.23 & 23.56 \pm 0.50 & 0.82 \pm 0.28 & 1.04 \pm 0.11 & 1.01 \pm 0.02 \pm 0.11 & 0.61 \pm 0.11 & 0.67 \pm 0.01 \\ \mathrm{HD} 2206697 & 4.96 \pm 0.024 & 2.88 \pm 0.25 & 31.02 \pm 0.38 & 0.95 \pm 0.45 & 0.78 \pm 0.15 & 1.44 \pm 0.16 & 0.78 \pm 0.17 & 0.79 \pm 0.09 \\ \mathrm{HD} 201665 & 3.48 \pm 0.19 & 1.79 \pm 0.23 & 23.36 \pm 0.50 & 0.82 \pm 0.28 & 1.04 \pm 0.15 & 1.03 \pm 0.07 & 0.51 \pm 0.07 & 0.09 \pm 0.06 \\ \mathrm{HD} 2036697 & 4.96 \pm 0.24 & 2.88 \pm 0.25 & 32.12 \pm 0.50 & 0.82 \pm 0.28 & 1.04 \pm 0.11 & 1.13 \pm 0.13 & 0.71 \pm 0.14 & 0.88 \pm 0.07 \\ \mathrm{HD} 103406 & 5.31 \pm 0.05 & 1.75 \pm 0.15 & 1.96 \pm 0.35 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.41 & 0.13 & 0.07 \pm 0.07 \pm 0.09 \\ \mathrm{HD} 206936 & 4.96 \pm 0.24 & 2.88 \pm 0.25 & 32.12 \pm 0.50 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.40 \pm 0.10 & 0.61 \pm 0.11 & 0.88 \pm 0.05 \\ \mathrm{HD} 023475 & 5.31 \pm 0.24 & 3.51 \pm 0.21 & 3.53 \pm 0.05 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.04 \pm 0$									
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$\begin{array}{c} \mathrm{HD} 1079870 & 1.35 \pm 0.10 & 0.84 \pm 0.20 & 10.07 \pm 0.47 & 0.84 \pm 0.12 & 1.34 \pm 0.08 & 0.47 \pm 0.05 & 0.49 \pm 0.06 & 0.94 \pm 0.06 \\ \mathrm{HD} 106449 & 1.500 \pm 0.12 & 1.11 \pm 0.20 & 14.01 \pm 0.47 & 0.69 \pm 0.16 & 0.94 \pm 0.07 & 0.51 \pm 0.05 & 0.49 \pm 0.06 & 0.87 \pm 0.07 \\ \mathrm{HD} 091810 & 1.97 \pm 0.14 & 1.09 \pm 0.24 & 12.10 \pm 0.56 & 0.99 \pm 0.17 & 1.99 \pm 0.08 & 0.58 \pm 0.06 & 0.53 \pm 0.06 & 1.11 \pm 0.05 \\ \mathrm{HD} 212466 & 1.88 \pm 0.16 & 0.33 \pm 0.22 & 16.48 \pm 0.49 & 0.79 \pm 0.27 & 1.55 \pm 0.10 & 0.31 \pm 0.11 & 0.23 \pm 0.12 & 0.21 \pm 0.33 \\ \mathrm{HD} 023082 & 3.32 \pm 0.19 & 1.81 \pm 0.22 & 22.90 \pm 0.48 & 0.76 \pm 0.29 & 1.04 \pm 0.10 & 0.31 \pm 0.11 & 0.23 \pm 0.12 & 0.21 \pm 0.33 \\ \mathrm{HD} 023032 & 3.25 \pm 0.25 & 1.87 \pm 0.23 & 32.22 \pm 0.47 & 0.77 \pm 0.38 & 1.07 \pm 0.14 & 1.11 \pm 0.12 & 0.61 \pm 0.13 & 0.67 \pm 0.05 \\ \mathrm{HD} 187238 & 3.99 \pm 0.24 & 2.19 \pm 0.23 & 32.22 \pm 0.47 & 0.77 \pm 0.38 & 1.07 \pm 0.14 & 1.11 \pm 0.12 & 0.61 \pm 0.13 & 0.67 \pm 0.05 \\ \mathrm{HD} 185622 & 4.83 \pm 0.31 & 2.95 \pm 0.28 & 31.02 \pm 0.58 & 0.95 \pm 0.45 & 0.78 \pm 0.15 & 1.44 \pm 0.16 & 0.78 \pm 0.17 & 0.79 \pm 0.09 \\ \mathrm{HD} 201693 & 3.48 \pm 0.19 & 1.93 \pm 0.23 & 32.66 \pm 0.50 & 0.82 \pm 0.28 & 1.04 \pm 0.15 & 1.44 \pm 0.16 & 0.78 \pm 0.17 & 0.79 \pm 0.09 \\ \mathrm{HD} 216946 & 4.60 \pm 0.24 & 2.88 \pm 0.25 & 32.12 \pm 0.50 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.41 \pm 0.11 & 0.61 \pm 0.17 & 0.66 \pm 0.03 \\ \mathrm{HD} 216946 & 4.60 \pm 0.24 & 2.88 \pm 0.25 & 32.12 \pm 0.50 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.41 \pm 0.11 & 0.86 \pm 0.12 & 0.88 \pm 0.07 \\ \mathrm{HD} 101494 & 4.64 \pm 0.24 & 2.65 \pm 0.24 & 35.23 \pm 0.47 & 0.76 \pm 0.37 & 3.33 \pm 0.11 & 1.53 \pm 0.11 & 0.78 \pm 0.12 & 0.71 \pm 0.09 \\ \mathrm{HD} 0139801 & 5.33 \pm 0.26 & 3.41 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.33 \pm 0.11 & 1.53 \pm 0.11 & 0.98 \pm 0.06 & 0.00 \\ \mathrm{HD} 010465 & 5.31 \pm 0.24 & 3.52 \pm 0.25 & 34.19 \pm 0.52 & 0.81 \pm 0.45 & 1.01 \pm 0.15 & 0.80 \pm 0.14 & 0.95 \pm 0.39 & 0.01 \\ \mathrm{HD} 010465 & 5.31 \pm 0.24 & 3.52 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.33 \pm 0.15 & 1.83 \pm 0.13 & 0.09 \pm 0.14 & 0.55 \pm 0.34 & 0.00 \\ \mathrm{HD} 019958 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 2.31 \pm 0.46 & 0.92 \pm 0.05 & 0.36 \pm 0.24 & 1.72 \pm 0.16 & 0.91 \pm 0.16 & 0.95 \pm 0$									
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$\begin{array}{c} \text{HD} 091810 & 1.97 \pm 0.14 & 1.09 \pm 0.24 \\ \text{HD} 212466 & 1.38 \pm 0.16 & 0.33 \pm 0.22 & 16.48 \pm 0.49 & 0.79 \pm 0.27 & 1.55 \pm 0.10 & 0.31 \pm 0.11 \\ \text{O} 232 \pm 0.12 & 0.21 \pm 0.33 \\ \text{HD} 202082 & 3.32 \pm 0.19 & 1.81 \pm 0.22 & 22.99 \pm 0.48 & 0.76 \pm 0.29 & 1.04 \pm 0.10 & 1.03 \pm 0.09 & 0.52 \pm 0.10 & 0.88 \pm 0.07 \\ \text{HD} 063302 & 3.25 \pm 0.25 & 1.87 \pm 0.23 & 32.22 \pm 0.47 & 0.77 \pm 0.38 & 1.07 \pm 0.14 & 1.11 \pm 0.12 & 0.61 \pm 0.13 & 0.67 \pm 0.05 \\ \text{HD} 1063303 & 3.90 \pm 0.24 & 2.19 \pm 0.24 & 26.10 \pm 0.52 & 0.89 \pm 0.35 & 0.80 \pm 0.11 & 1.31 \pm 0.13 & 0.71 \pm 0.14 & 0.83 \pm 0.07 \\ \text{HD} 185622 & 48.3 \pm 0.31 & 2.95 \pm 0.28 & 31.02 \pm 0.58 & 0.95 \pm 0.45 & 0.78 \pm 0.15 & 1.44 \pm 0.16 & 0.78 \pm 0.17 & 0.79 \pm 0.09 \\ \text{HD} 201065 & 3.48 \pm 0.19 & 1.93 \pm 0.23 & 23.36 \pm 0.50 & 0.82 \pm 0.28 & 1.04 \pm 0.10 & 1.10 \pm 0.10 & 0.61 \pm 0.11 & 0.89 \pm 0.06 \\ \text{HD} 207991 & 2.73 \pm 0.10 & 1.75 \pm 0.15 & 19.36 \pm 0.34 & 0.86 \pm 0.12 & 0.48 \pm 0.05 & 1.03 \pm 0.07 & 0.51 \pm 0.07 & 0.60 \pm 0.03 \\ \text{HD} 216946 & 4.60 \pm 0.24 & 2.78 \pm 0.22 & 27.79 \pm 0.50 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.41 \pm 0.11 & 0.80 \pm 0.12 & 0.87 \pm 0.13 \\ \text{HD} 181475 & 4.34 \pm 0.23 & 2.74 \pm 0.22 & 27.28 \pm 0.48 & 0.95 \pm 0.32 & 0.59 \pm 0.11 & 1.53 \pm 0.11 & 0.78 \pm 0.12 & 0.87 \pm 0.13 \\ \text{HD} 1914404 & 4.64 \pm 0.24 & 2.65 \pm 0.24 & 3.52 \pm 0.47 & 0.76 \pm 0.37 & 1.33 \pm 0.11 & 1.72 \pm 0.14 & 0.79 \pm 0.15 & 0.54 \pm 0.07 \\ \text{HD} 035601 & 4.91 \pm 0.28 & 2.82 \pm 0.26 & 34.11 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.38 & 0.12 & 1.17 \pm 0.11 & 0.78 \pm 0.12 & 0.74 \pm 0.09 \\ \text{HD} 020375 & 5.13 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.09 \pm 0.14 & 0.54 \pm 0.06 \\ \text{HD} 020375 & 5.13 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.09 \pm 0.14 & 0.54 \pm 0.06 \\ \text{HD} 0203475 & 5.13 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \text{HD} 023377 & 0.73 \pm 0.37 & 3.23 \pm 0.21 & 36.09 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.75 \pm 0.06 \\ \text{HD} 023397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.26 & 0.33 $									
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$\begin{array}{c} \mathrm{HD} 023082 3.32 \pm 0.19 1.81 \pm 0.22 22.90 \pm 0.48 0.76 \pm 0.29 1.04 \pm 0.10 1.03 \pm 0.09 0.52 \pm 0.10 0.88 \pm 0.07 \\ \mathrm{HD} 063302 3.25 \pm 0.25 1.87 \pm 0.23 32.22 \pm 0.47 0.77 \pm 0.38 1.07 \pm 0.14 1.11 \pm 0.12 0.61 \pm 0.13 0.67 \pm 0.05 \\ \mathrm{HD} 187238 3.90 \pm 0.24 2.19 \pm 0.24 26.10 \pm 0.52 0.89 \pm 0.35 0.80 \pm 0.11 1.31 \pm 0.13 0.71 \pm 0.14 0.83 \pm 0.07 \\ \mathrm{HD} 185622 4.83 \pm 0.31 2.95 \pm 0.28 31.02 \pm 0.58 0.95 \pm 0.45 0.78 \pm 0.15 1.44 \pm 0.16 0.78 \pm 0.17 0.79 \pm 0.09 \\ \mathrm{HD} 201065 3.48 \pm 0.19 1.93 \pm 0.23 23.36 \pm 0.50 0.88 \pm 0.28 1.04 \pm 0.10 1.10 \pm 0.10 0.61 \pm 0.11 0.79 \pm 0.09 \\ \mathrm{HD} 207991 2.73 \pm 0.10 1.75 \pm 0.15 19.36 \pm 0.34 0.86 \pm 0.12 0.48 \pm 0.05 1.03 \pm 0.07 0.51 \pm 0.07 0.60 \pm 0.03 \\ \mathrm{HD} 216946 4.60 \pm 0.24 2.73 \pm 0.23 27.79 \pm 0.50 0.85 \pm 0.34 0.30 \pm 0.13 1.04 \pm 0.10 0.61 \pm 0.12 0.87 \pm 0.07 \\ \mathrm{HD} 181475 4.34 \pm 0.23 2.74 \pm 0.22 27.28 \pm 0.48 0.95 \pm 0.32 0.59 \pm 0.11 1.53 \pm 0.11 0.84 \pm 0.12 0.87 \pm 0.07 \\ \mathrm{HD} 181475 4.34 \pm 0.23 2.74 \pm 0.22 27.28 \pm 0.48 0.95 \pm 0.32 0.59 \pm 0.11 1.53 \pm 0.11 0.78 \pm 0.12 0.71 \pm 0.09 \\ \mathrm{HD} 014040 4.64 \pm 0.24 2.65 \pm 0.24 32.32 \pm 0.47 0.76 \pm 0.37 1.33 \pm 0.11 1.72 \pm 0.14 0.79 \pm 0.15 0.45 \pm 0.07 \\ \mathrm{HD} 035601 4.91 \pm 0.28 2.82 \pm 0.26 34.19 \pm 0.52 0.81 \pm 0.45 1.01 \pm 0.15 1.70 \pm 0.15 0.80 \pm 0.16 \\ \mathrm{HD} 035601 4.91 \pm 0.28 2.82 \pm 0.26 34.19 \pm 0.52 0.81 \pm 0.45 1.01 \pm 0.15 1.70 \pm 0.15 0.80 \pm 0.16 \\ \mathrm{HD} 034075 5.13 \pm 0.24 3.51 \pm 0.21 36.09 \pm 0.41 0.84 \pm 0.39 0.52 \pm 0.12 2.08 \pm 0.13 0.99 \pm 0.14 0.51 \pm 0.06 \\ \mathrm{HD} 010465 5.31 \pm 0.24 3.51 \pm 0.21 36.09 \pm 0.41 0.84 \pm 0.39 0.52 \pm 0.12 2.08 \pm 0.13 0.99 \pm 0.14 0.51 \pm 0.06 \\ \mathrm{HD} 010405 5.73 \pm 0.37 3.22 \pm 0.28 2.02 26.97 \pm 0.42 0.91 \pm 0.26 0.34 \pm 0.10 1.77 \pm 0.13 0.94 \pm 0.14 0.51 \pm 0.06 \\ \mathrm{HD} 023475 5.13 \pm 0.24 3.51 \pm 0.21 36.09 \pm 0.14 0.84 \pm 0.39 0.52 \pm 0.12 2.08 \pm 0.13 0.99 \pm 0.14 0.51 \pm 0.06 \\ \mathrm{HD} 013174$									
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$\begin{array}{llllllllllllllllllllllllllllllllllll$	HD 187238	3.90 ± 0.24		26.10 ± 0.52	0.89 ± 0.35			0.71 ± 0.14	0.83 ± 0.07
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$\begin{array}{c} \mathrm{HD}216946 & 4.60 \pm 0.24 & 2.73 \pm 0.23 & 27.79 \pm 0.50 & 0.85 \pm 0.34 & 0.30 \pm 0.13 & 1.41 \pm 0.11 & 0.80 \pm 0.12 & 0.87 \pm 0.07 \\ \mathrm{HD}236697 & 4.96 \pm 0.24 & 2.88 \pm 0.25 & 32.12 \pm 0.50 & 0.81 \pm 0.34 & 0.42 \pm 0.12 & 1.71 \pm 0.11 & 0.84 \pm 0.12 & 0.83 \pm 0.13 \\ \mathrm{HD}181475 & 4.34 \pm 0.23 & 2.74 \pm 0.22 & 27.28 \pm 0.48 & 0.95 \pm 0.32 & 0.59 \pm 0.11 & 1.53 \pm 0.11 & 0.78 \pm 0.12 & 0.71 \pm 0.09 \\ \mathrm{HD}014404 & 4.64 \pm 0.24 & 2.65 \pm 0.24 & 35.23 \pm 0.47 & 0.76 \pm 0.37 & 1.33 \pm 0.11 & 1.72 \pm 0.14 & 0.79 \pm 0.15 & 0.54 \pm 0.07 \\ \mathrm{HD}039801 & 5.33 \pm 0.26 & 3.41 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.38 & 0.22 \pm 0.15 & 1.83 \pm 0.13 & 1.00 \pm 0.14 & 0.45 \pm 0.06 \\ \mathrm{HD}035601 & 4.91 \pm 0.28 & 2.82 \pm 0.26 & 34.19 \pm 0.52 & 0.81 \pm 0.45 & 1.01 \pm 0.15 & 1.70 \pm 0.15 & 0.80 \pm 0.16 & 0.59 \pm 0.09 \\ \mathrm{HD}206936 & 4.66 \pm 0.41 & 2.90 \pm 0.28 & 40.26 \pm 0.54 & 0.82 \pm 0.65 & 0.36 \pm 0.24 & 1.72 \pm 0.16 & 0.91 \pm 0.18 & 0.65 \pm 0.06 \\ \mathrm{HD}013455 & 5.31 \pm 0.24 & 3.52 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \mathrm{HD}023475 & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \mathrm{HD}040239 & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \mathrm{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \mathrm{HD}019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 2.31 \pm 4 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \mathrm{HD}027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.30 \pm 0.05 \\ \mathrm{HD}013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.06 \\ \mathrm{HD}017918 & 0.44 \pm 0.04 & 0.09 \pm 0.90 & 0.07 \pm 0.23 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.04 & 0.02 \pm 0.04 & 0.33 \pm 0.09 \\ \mathrm{HD}017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.04 & 0.03 \pm 0.05 \\ \mathrm{HD}0217$									
$\begin{array}{c} \text{HD} 236697 & 4.96 \pm 0.24 & 2.88 \pm 0.25 & 32.12 \pm 0.50 & 0.81 \pm 0.34 & 0.42 \pm 0.12 & 1.71 \pm 0.11 & 0.84 \pm 0.12 & 0.71 \pm 0.09 \\ \text{HD} 0181475 & 4.34 \pm 0.23 & 2.74 \pm 0.22 & 27.28 \pm 0.48 & 0.95 \pm 0.32 & 0.59 \pm 0.11 & 1.53 \pm 0.11 & 0.78 \pm 0.12 & 0.71 \pm 0.09 \\ \text{HD} 014404 & 4.64 \pm 0.24 & 2.65 \pm 0.24 & 35.23 \pm 0.47 & 0.76 \pm 0.37 & 1.33 \pm 0.11 & 1.72 \pm 0.14 & 0.79 \pm 0.15 & 0.54 \pm 0.07 \\ \text{HD} 039801 & 5.33 \pm 0.26 & 3.41 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.38 & 0.22 \pm 0.15 & 1.83 \pm 0.13 & 1.00 \pm 0.14 & 0.45 \pm 0.06 \\ \text{HD} 035601 & 4.91 \pm 0.28 & 2.82 \pm 0.26 & 34.19 \pm 0.52 & 0.81 \pm 0.45 & 1.01 \pm 0.15 & 1.70 \pm 0.15 & 0.80 \pm 0.16 & 0.59 \pm 0.09 \\ \text{HD} 206936 & 4.66 \pm 0.41 & 2.90 \pm 0.28 & 40.26 \pm 0.54 & 0.82 \pm 0.65 & 0.36 \pm 0.24 & 1.72 \pm 0.16 & 0.91 \pm 0.18 & 0.65 \pm 0.06 \\ \text{HD} 010465 & 5.31 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \text{HD} 023475 & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \text{HD} 040239 & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \text{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD} 019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD} 0273797 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.30 \pm 0.05 \\ \text{HD} 013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.06 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.07 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.04 & 0.02 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.04 & 0.02 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.05 & 0.$									
$\begin{array}{c} \text{HD} 181475 & 4.34 \pm 0.23 & 2.74 \pm 0.22 & 27.28 \pm 0.48 & 0.95 \pm 0.32 & 0.59 \pm 0.11 & 1.53 \pm 0.11 & 0.78 \pm 0.12 & 0.71 \pm 0.09 \\ \text{HD} 014404 & 4.64 \pm 0.24 & 2.65 \pm 0.24 & 35.23 \pm 0.47 & 0.76 \pm 0.37 & 1.33 \pm 0.11 & 1.72 \pm 0.14 & 0.79 \pm 0.15 & 0.54 \pm 0.07 \\ \text{HD} 035901 & 5.33 \pm 0.26 & 3.41 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.38 & 0.22 \pm 0.15 & 1.83 \pm 0.13 & 1.00 \pm 0.14 & 0.45 \pm 0.06 \\ \text{HD} 035601 & 4.91 \pm 0.28 & 2.82 \pm 0.26 & 34.19 \pm 0.52 & 0.81 \pm 0.45 & 1.01 \pm 0.15 & 1.70 \pm 0.15 & 0.80 \pm 0.16 & 0.59 \pm 0.09 \\ \text{HD} 206936 & 4.66 \pm 0.41 & 2.90 \pm 0.28 & 40.26 \pm 0.54 & 0.82 \pm 0.65 & 0.36 \pm 0.24 & 1.72 \pm 0.16 & 0.91 \pm 0.18 & 0.65 \pm 0.06 \\ \text{HD} 010465 & 5.31 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \text{HD} 023475 & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \text{HD} 040239 & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \text{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD} 019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD} 027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.30 \pm 0.05 \\ \text{HD} 013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.03 \\ \text{HD} 0157918 & 0.44 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.10 \pm 0.03 & 0.42 \pm 0.03 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 5.06 \pm 0.06 & 0.05 & 0.04 \pm 0.05 & 0.38 \pm 0.04 \\ $									
$\begin{array}{c} \text{HD}014404 & 4.64 \pm 0.24 & 2.65 \pm 0.24 & 35.23 \pm 0.47 & 0.76 \pm 0.37 & 1.33 \pm 0.11 & 1.72 \pm 0.14 & 0.79 \pm 0.15 & 0.54 \pm 0.07 \\ \text{HD}039801 & 5.33 \pm 0.26 & 3.41 \pm 0.22 & 36.81 \pm 0.44 & 0.87 \pm 0.38 & 0.22 \pm 0.15 & 1.83 \pm 0.13 & 1.00 \pm 0.14 & 0.45 \pm 0.06 \\ \text{HD}035601 & 4.91 \pm 0.28 & 2.82 \pm 0.26 & 34.19 \pm 0.52 & 0.81 \pm 0.45 & 1.01 \pm 0.15 & 1.70 \pm 0.15 & 0.80 \pm 0.16 & 0.59 \pm 0.09 \\ \text{HD}206936 & 4.66 \pm 0.41 & 2.90 \pm 0.28 & 40.26 \pm 0.54 & 0.82 \pm 0.65 & 0.36 \pm 0.24 & 1.72 \pm 0.16 & 0.91 \pm 0.18 & 0.65 \pm 0.06 \\ \text{HD}010465 & 5.31 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \text{HD}023475 & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \text{HD}040239 & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \text{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD}019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD}027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.09 \pm 0.20 & 0.16 \pm 0.11 & 2.03 \pm 0.12 & 1.00 \pm 0.03 & 0.08 \pm 0.03 \\ \text{HD}027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.33 \pm 0.05 \\ \text{HD}013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.04 \pm 0.03 & 0.04 \pm 0.03 \\ \text{HD}047991 & 0.44 \pm 0.04 & 0.07 \pm 0.09 & 0.07 \pm 0.23 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD}0179718 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.32 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.04 & 0.33 \pm 0.09 \\ \text{HD}017970 & 0.19 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.06 & 0.57 \pm 0.01 \\ \text{HD}0188804 & 0.94 \pm 0.05 & 0.54 \pm 0.14 & 0.13 \pm 0.35 & 0.60 \pm 0.09 & 0.49 \pm 0.05 & 0.03 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.05 \\ \text{HD}014433 & 0.99 \pm 0.05 & 0.5$									
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$\begin{array}{c} \text{HD 206936} & 4.66 \pm 0.41 & 2.90 \pm 0.28 & 40.26 \pm 0.54 & 0.82 \pm 0.65 & 0.36 \pm 0.24 & 1.72 \pm 0.16 & 0.91 \pm 0.18 & 0.65 \pm 0.06 \\ \text{HD 010465} & 5.31 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \text{HD 023475} & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \text{HD 040239} & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \text{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD 019058} & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD 156014} & 5.74 \pm 0.24 & 3.83 \pm 0.20 & 25.87 \pm 0.44 & 0.88 \pm 0.34 & -0.14 \pm 0.11 & 2.03 \pm 0.12 & 1.00 \pm 0.14 & 0.51 \pm 0.14 \\ \text{HD 089025} & 0.14 \pm 0.03 & 0.00 \pm 0.08 & 0.09 \pm 0.20 & 0.16 \pm 0.11 & 5.83 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.06 \\ \text{HD 013174} & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.10 \pm 0.03 & 0.42 \pm 0.03 \\ \text{HD 40535} & 0.34 \pm 0.04 & 0.09 \pm 0.09 & 0.07 \pm 0.23 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD 017918} & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 4.94 \pm 0.05 & 0.03 \pm 0.04 & 0.14 \pm 0.04 & 0.29 \pm 0.03 \\ \text{HD 027770} & 0.19 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 5.01 \pm 0.06 & 0.03 \pm 0.04 & 0.14 \pm 0.04 & 0.29 \pm 0.03 \\ \text{HD 186155} & 0.43 \pm 0.07 & 0.27 \pm 0.17 & -0.11 \pm 0.43 & 0.66 \pm 0.11 & 6.98 \pm 0.09 & 0.09 \pm 0.06 & 0.13 \pm 0.04 & 0.55 \pm 0.05 \\ \text{HD 160365} & 0.88 \pm 0.04 & 0.26 \pm 0.11 & -0.13 \pm 0.29 & 0.42 \pm 0.07 & 4.59 \pm 0.05 & 0.07 \pm 0.04 & 0.15 \pm 0.04 & 0.57 \pm 0.05 \\ \text{HD 1248800} & 0.57 \pm 0.04 & 0.04 & 0.07 \pm 0.12 & -0.29 \pm 0.31 & 0.44 \pm 0.09 & 4.44 \pm 0.06 & 0.05 \pm 0.04 & 0.15 \pm 0.04 & 0.45 \pm 0.04 \\ \text{HD 220657} & 0.44 \pm 0.04 & 0.07 \pm 0.12 & -0.29 \pm 0.31 & 0.44 \pm 0.09 & 4.44 \pm 0.06 & 0.05 \pm 0.04 & 0.15 \pm 0.04 & 0.45$		5.33 ± 0.26	3.41 ± 0.22			0.22 ± 0.15	1.83 ± 0.13	1.00 ± 0.14	0.45 ± 0.06
$\begin{array}{c} \text{HD}010465 & 5.31 \pm 0.24 & 3.51 \pm 0.21 & 36.09 \pm 0.41 & 0.84 \pm 0.39 & 0.52 \pm 0.12 & 2.08 \pm 0.13 & 0.99 \pm 0.14 & 0.51 \pm 0.08 \\ \text{HD}023475 & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \text{HD}040239 & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \text{RW}\text{Cyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD}019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD}0156014 & 5.74 \pm 0.24 & 3.83 \pm 0.20 & 25.87 \pm 0.44 & 0.88 \pm 0.34 & -0.14 \pm 0.11 & 2.03 \pm 0.12 & 1.00 \pm 0.14 & 0.51 \pm 0.14 \\ \hline \\ \text{HD}089025 & 0.14 \pm 0.03 & 0.00 \pm 0.08 & 0.09 \pm 0.20 & 0.16 \pm 0.11 & 5.83 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.06 \\ \text{HD}027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.30 \pm 0.05 \\ \text{HD}013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.10 \pm 0.03 & 0.42 \pm 0.03 \\ \text{HD}040535 & 0.34 \pm 0.04 & 0.09 \pm 0.09 & 0.07 \pm 0.23 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD}017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.14 \pm 0.04 & 0.29 \pm 0.03 \\ \text{HD}021770 & 0.19 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 4.94 \pm 0.05 & 0.03 \pm 0.05 & 0.04 \pm 0.05 & 0.38 \pm 0.04 \\ \text{HD}075555 & 0.41 \pm 0.09 & 0.19 \pm 0.15 & 0.29 \pm 0.37 & 0.47 \pm 0.13 & 5.08 \pm 0.09 & 0.09 \pm 0.06 & 0.13 \pm 0.06 & 0.75 \pm 0.11 \\ \text{HD}186155 & 0.43 \pm 0.07 & 0.27 \pm 0.17 & -0.11 \pm 0.43 & 0.66 \pm 0.11 & 6.98 \pm 0.09 & 0.09 \pm 0.06 & 0.13 \pm 0.06 & 0.78 \pm 0.07 \\ \text{HD}188040 & 0.94 \pm 0.05 & 0.54 \pm 0.14 & 0.13 \pm 0.35 & 0.60 \pm 0.09 & 4.40 \pm 0.07 & 0.18 \pm 0.03 & 0.13 \pm 0.04 & 0.50 \pm 0.05 \\ \text{HD}14433 & 0.49 \pm 0.06 & 0.19 \pm 0.15 & 0.07 \pm 0.39 & 0.45 \pm 0.04 & 0.44 \pm 0.06 & 0.05 \pm 0.04 & 0.15 \pm 0.04 & 0.45 \pm$	HD 035601	4.91 ± 0.28	2.82 ± 0.26	34.19 ± 0.52	0.81 ± 0.45	1.01 ± 0.15	1.70 ± 0.15	0.80 ± 0.16	0.59 ± 0.09
$\begin{array}{c} \text{HD}023475 & 5.13 \pm 0.24 & 3.52 \pm 0.21 & 27.13 \pm 0.46 & 0.92 \pm 0.36 & 0.40 \pm 0.13 & 1.66 \pm 0.11 & 0.88 \pm 0.12 & 0.57 \pm 0.14 \\ \text{HD}040239 & 4.96 \pm 0.18 & 3.29 \pm 0.20 & 26.97 \pm 0.42 & 0.91 \pm 0.26 & 0.34 \pm 0.10 & 1.77 \pm 0.13 & 0.94 \pm 0.14 & 0.58 \pm 0.08 \\ \text{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD}019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD}156014 & 5.74 \pm 0.24 & 3.83 \pm 0.20 & 25.87 \pm 0.44 & 0.88 \pm 0.34 & -0.14 \pm 0.11 & 2.03 \pm 0.12 & 1.00 \pm 0.14 & 0.51 \pm 0.14 \\ \text{HD}089025 & 0.14 \pm 0.03 & 0.00 \pm 0.08 & 0.09 \pm 0.20 & 0.16 \pm 0.11 & 5.83 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.06 \\ \text{HD}027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.30 \pm 0.05 \\ \text{HD}013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.10 \pm 0.03 & 0.42 \pm 0.03 \\ \text{HD}04535 & 0.34 \pm 0.04 & 0.09 \pm 0.09 & 0.07 \pm 0.23 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD}017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 4.94 \pm 0.05 & 0.03 \pm 0.04 & 0.14 \pm 0.04 & 0.29 \pm 0.03 \\ \text{HD}021770 & 0.19 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.09 & 4.94 \pm 0.05 & 0.03 \pm 0.04 & 0.14 \pm 0.04 & 0.29 \pm 0.03 \\ \text{HD}0186155 & 0.43 \pm 0.07 & 0.27 \pm 0.17 & -0.11 \pm 0.43 & 0.66 \pm 0.11 & 6.98 \pm 0.09 & 0.09 \pm 0.06 & 0.13 \pm 0.06 & 0.78 \pm 0.07 \\ \text{HD}018645 & 0.88 \pm 0.04 & 0.26 \pm 0.11 & -0.13 \pm 0.29 & 0.42 \pm 0.07 & 4.59 \pm 0.05 & 0.07 \pm 0.04 & 0.15 \pm 0.04 & 0.45 \pm 0.04 \\ \text{HD}017443 & 0.49 \pm 0.05 & 0.54 \pm 0.14 & 0.13 \pm 0.35 & 0.60 \pm 0.09 & 4.40 \pm 0.07 & 0.18 \pm 0.03 & 0.13 \pm 0.04 & 0.50 \pm 0.05 \\ \text{HD}01443 & 0.49 \pm 0.06 & 0.19 \pm 0.13 & -0.36 \pm 0.32 & 0.56 \pm 0.11 & 4.77 \pm 0.06 & 0.06 \pm 0.04 & 0.15 \pm 0.04 & 0.45 \pm 0.04 \\ \text{HD}0220657 & 0.44 \pm 0.04 & 0.07 \pm 0.12 & -0.29 \pm 0.31 & 0.44 \pm 0.09 & 4.44 \pm 0.06 & 0.05 \pm 0.04 & 0.15 \pm 0.04 & 0.45 \pm 0.04 $									
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$\begin{array}{c} \text{RWCyg} & 5.73 \pm 0.37 & 3.22 \pm 0.28 & 35.36 \pm 0.56 & 0.93 \pm 0.69 & 0.75 \pm 0.24 & 1.68 \pm 0.23 & 0.84 \pm 0.26 & 0.27 \pm 0.10 \\ \text{HD} 019058 & 5.07 \pm 0.20 & 3.31 \pm 0.15 & 23.14 \pm 0.34 & 0.96 \pm 0.20 & -0.61 \pm 0.08 & 1.70 \pm 0.11 & 0.71 \pm 0.12 & 0.46 \pm 0.10 \\ \text{HD} 156014 & 5.74 \pm 0.24 & 3.83 \pm 0.20 & 25.87 \pm 0.44 & 0.88 \pm 0.34 & -0.14 \pm 0.11 & 2.03 \pm 0.12 & 1.00 \pm 0.14 & 0.51 \pm 0.14 \\ \hline \\ \text{HD} 089025 & 0.14 \pm 0.03 & 0.00 \pm 0.08 & 0.09 \pm 0.20 & 0.16 \pm 0.11 & 5.83 \pm 0.05 & 0.00 \pm 0.03 & 0.08 \pm 0.03 & 0.24 \pm 0.06 \\ \text{HD} 027397 & 0.45 \pm 0.05 & 0.04 \pm 0.08 & 0.19 \pm 0.20 & 0.23 \pm 0.06 & 6.55 \pm 0.05 & -0.04 \pm 0.04 & 0.02 \pm 0.04 & 0.30 \pm 0.05 \\ \text{HD} 013174 & 0.38 \pm 0.04 & 0.14 \pm 0.10 & 0.12 \pm 0.24 & 0.31 \pm 0.08 & 5.47 \pm 0.05 & 0.00 \pm 0.03 & 0.10 \pm 0.03 & 0.42 \pm 0.03 \\ \text{HD} 40535 & 0.34 \pm 0.04 & 0.09 \pm 0.09 & 0.07 \pm 0.23 & 0.25 \pm 0.09 & 5.66 \pm 0.06 & 0.07 \pm 0.04 & 0.10 \pm 0.04 & 0.33 \pm 0.09 \\ \text{HD} 017918 & 0.44 \pm 0.04 & 0.17 \pm 0.09 & 0.28 \pm 0.22 & 0.30 \pm 0.09 & 4.94 \pm 0.05 & 0.03 \pm 0.04 & 0.14 \pm 0.04 & 0.29 \pm 0.03 \\ \text{HD} 021770 & 0.19 \pm 0.04 & 0.14 \pm 0.09 & -0.44 \pm 0.23 & 0.30 \pm 0.07 & 5.01 \pm 0.06 & 0.03 \pm 0.05 & 0.04 \pm 0.05 & 0.38 \pm 0.04 \\ \text{HD} 075555 & 0.41 \pm 0.09 & 0.19 \pm 0.15 & 0.29 \pm 0.37 & 0.47 \pm 0.13 & 5.08 \pm 0.08 & 0.06 \pm 0.05 & 0.07 \pm 0.06 & 0.57 \pm 0.11 \\ \text{HD} 186155 & 0.43 \pm 0.07 & 0.27 \pm 0.17 & -0.11 \pm 0.43 & 0.66 \pm 0.11 & 6.98 \pm 0.09 & 0.09 \pm 0.06 & 0.13 \pm 0.06 & 0.78 \pm 0.07 \\ \text{HD} 160365 & 0.88 \pm 0.04 & 0.26 \pm 0.11 & -0.13 \pm 0.29 & 0.42 \pm 0.07 & 4.59 \pm 0.05 & 0.07 \pm 0.04 & 0.16 \pm 0.04 & 0.47 \pm 0.06 \\ \text{HD} 011443 & 0.49 \pm 0.06 & 0.19 \pm 0.15 & 0.07 \pm 0.39 & 0.48 \pm 0.09 & 4.40 \pm 0.07 & 0.09 \pm 0.06 & 0.15 \pm 0.07 & 0.70 \pm 0.05 \\ \text{HD} 124850 & 0.57 \pm 0.07 & 0.39 \pm 0.15 & 0.07 \pm 0.39 & 0.48 \pm 0.09 & 4.44 \pm 0.06 & 0.05 \pm 0.04 & 0.15 \pm 0.07 & 0.70 \pm 0.05 \\ \text{HD} 220657 & 0.44 \pm 0.04 & 0.07 \pm 0.12 & -0.29 \pm 0.31 & 0.44 \pm 0.09 & 4.44 \pm 0.06 & 0.05 \pm 0.04 & 0.15 \pm 0.04 & 0.45 \pm 0.04 \\ \text{HD} 240657 & 0.44 \pm 0.04 & 0.07 \pm 0.12 & -0.29 \pm 0.31 & 0.44 \pm 0.09 & 4.44 \pm 0.06 & 0.05 \pm 0.$									
$\begin{array}{c} \text{HD}019058 & 5.07 \pm 0.20 \\ \text{HD}156014 & 5.74 \pm 0.24 \\ \end{array} \begin{array}{c} 3.31 \pm 0.15 \\ 3.83 \pm 0.20 \\ \end{array} \begin{array}{c} 23.14 \pm 0.34 \\ 25.87 \pm 0.44 \\ \end{array} \begin{array}{c} 0.96 \pm 0.20 \\ 0.88 \pm 0.34 \\ \end{array} \begin{array}{c} -0.61 \pm 0.08 \\ -0.14 \pm 0.11 \\ \end{array} \begin{array}{c} 1.70 \pm 0.11 \\ 2.03 \pm 0.12 \\ \end{array} \begin{array}{c} 0.71 \pm 0.12 \\ 1.00 \pm 0.14 \\ \end{array} \begin{array}{c} 0.46 \pm 0.10 \\ 0.51 \pm 0.14 \\ \end{array} \end{array}$									
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Table 7. continued.

Sign Discription Naj									
$ \begin{aligned} & \text{HD} 2 4977 & 0.75 \pm 0.06 & 0.37 \pm 0.15 & 1.97 \pm 0.37 & 0.68 \pm 0.09 & 2.78 \pm 0.08 & 0.22 \pm 0.06 & 0.24 \pm 0.06 & 0.47 \pm 0.07 & 0.07 \pm 0.01 & 0.09 \pm 0.01 & 0.07 \pm 0.01 & 0.07 \pm 0.01 & 0.07 \pm 0.01 & 0.01 & 0.01 & 0.07 \pm 0.01 & 0.07 \pm 0.01 & 0.01 & 0.01 & 0.07 \pm 0.01 & 0.07$	Star IDs	Nad	Caa	¹² CO	Mø1	Brγ	FeA	FeB	Mø2
$ \begin{aligned} & \text{HD108608} & 193 + 0.017 & 0.57 + 0.17 & 0.79 + 0.44 & 0.84 + 0.07 & 3.02 + 0.107 & 0.25 + 0.05 & 0.66 + 0.04 \\ & \text{HD108477} & 1.09 + 0.017 & 0.55 + 0.15 & -0.57 + 0.35 & 0.06 + 0.04 \\ & \text{HD108477} & 1.09 + 0.017 & 0.55 + 0.18 & -0.05 & 0.08 + 0.017 & 0.08 + 0.10 \\ & \text{HD101847} & 1.09 + 0.017 & 0.55 + 0.18 & 0.31 + 0.16 & 0.82 + 0.08 & 2.04 + 0.08 & 0.34 + 0.05 & 0.33 + 0.05 & 0.72 + 0.05 \\ & \text{HD101087} & 1.01 + 0.018 & 0.57 + 0.18 & 0.31 + 0.16 & 0.82 + 0.08 & 2.04 + 0.09 & 0.34 + 0.05 & 0.33 + 0.05 & 0.35 + 0.03 & 0.07 \\ & \text{HD101087} & 1.01 + 0.016 & 0.65 + 0.02 & 0.01 + 0.02 & 0.01 + 0.011 & 3.00 + 0.01 & 0.25 + 0.00 & 0.25 + 0.07 & 0.25 + 0.07 \\ & \text{HD102094} & 1.17 + 0.07 & 0.83 + 0.019 & 5.25 + 0.45 & 0.82 + 0.08 & 2.03 + 0.05 & 0.05 + 0.02 & 0.07 + 0.06 \\ & \text{HD1020618} & 1.13 + 0.04 & 0.07 + 0.017 & 3.89 + 0.42 & 0.83 + 0.05 & 1.07 + 0.08 & 0.25 + 0.05 & 0.27 + 0.06 & 0.08 + 0.01 \\ & \text{HD103479} & 0.97 + 0.09 & 0.65 + 0.017 & 7.05 + 0.02 & 0.00 + 0.018 & 0.02 + 0.02 & 0.07 + 0.06 \\ & \text{HD1105722} & 1.02 + 0.04 & 0.81 + 0.16 & 7.25 + 0.02 & 0.07 + 0.06 & 1.32 + 0.07 & 0.18 + 0.06 & 0.20 + 0.07 & 0.07 + 0.06 \\ & \text{HD1105720} & 1.66 + 0.16 & 1.03 + 0.18 & 1.73 + 1 + 0.40 & 0.39 + 0.02 & 0.00 & 0.25 + 0.06 & 0.20 + 0.07 & 0.05 + 0.04 \\ & \text{HD10200948} & 1.19 + 0.06 & 0.01 + 7 & 2.02 & 9.60 + 0.04 & 0.30 + 0.02 & 0.07 & 0.08 + 0.04 \\ & \text{HD110006} & 1.13 + 0.109 & 0.72 + 0.20 & 9.60 + 0.04 & 0.30 + 0.02 & 0.00 & 0.20 + 0.06 & 0.05 + 0.05 \\ & \text{HD110006} & 1.13 + 0.109 & 0.72 + 0.20 & 9.60 + 0.04 & 0.30 + 0.02 & 0.00 & 0.02 + 0.00 & 0.05 + 0.05 \\ & \text{HD110006} & 1.13 + 0.109 & 0.72 + 0.20 & 9.60 + 0.04 & 0.30 + 0.02 & 0.00 & 0.02 + 0.00 & 0.05 + 0.05 \\ & \text{HD110006} & 1.13 + 0.109 & 0.72 + 0.20 & 9.60 + 0.04 & 0.30 + 0.00 & 0.20 + 0.00 & 0.20 + 0.00 & 0.05 + 0.00 & 0.05 + 0.00 \\ & \text{HD110006} & 1.13 + 0.109 & 0.72 + 0.02 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0.00 & 0.00 + 0$									
HD 186839 0.83 \pm 0.04 0.51 \pm 0.15 2.75 \pm 0.38 0.79 \pm 0.06 2.73 \pm 0.07 0.28 \pm 0.05 0.33 \pm 0.05 0.65 \pm 0.05 HD 181474 0.82 \pm 0.07 0.85 \pm 0.07 0.45 \pm 0.05 HD 181474 0.82 \pm 0.07 0.85 \pm 0.08 \pm 0.15 0.46 \pm 0.26 \pm 0.08 0.26 \pm 0.05 0.31 \pm 0.06 0.75 \pm 0.05 HD 181474 0.82 \pm 0.07 0.85 \pm 0.08 \pm 0.15 0.16 \pm 0.16 0.16 0.16 0.17 \pm 0.17 0									
HD 186839 0.83 \pm 0.04 0.51 \pm 0.15 2.75 \pm 0.38 0.79 \pm 0.06 2.73 \pm 0.07 0.28 \pm 0.05 0.33 \pm 0.05 0.66 \pm 0.75 \pm 0.05 HD 18474 0.82 \pm 0.07 0.85 \pm 0.07 0.47 \pm 0.03 0.74 \pm 0.00 0.26 \pm 0.05 0.31 \pm 0.05 0.75 \pm 0.05 HD 18474 0.82 \pm 0.07 0.85 \pm 0.08 0.15 \pm 0.07 0.85 \pm 0.08 HD 010697 1.01 \pm 0.01 0.42 \pm 0.27 \pm 0.18 0.31 \pm 0.04 0.82 \pm 0.07 0.35 \pm 0.05 0.31 \pm 0.05 0.75 \pm 0.05 HD 010697 1.01 \pm 0.01 0.42 \pm 0.17 \pm 0.17 1.07 \pm 0.07 \pm 0.27 \pm 0.16 0.31 \pm 0.06 0.25 \pm 0.07 0.25	HD 126868	0.93 ± 0.07	0.57 ± 0.17	0.79 ± 0.44	0.84 ± 0.07	3.02 ± 0.10	0.19 ± 0.07	0.27 ± 0.07	0.76 ± 0.05
$ \begin{aligned} & \text{HD 108477} & 1.09 \pm 0.07 & 0.54 \pm 0.17 & 4.64 \pm 0.42 & 0.68 \pm 0.10 & 2.06 \pm 0.08 & 0.34 \pm 0.05 & 0.31 \pm 0.06 & 0.75 \pm 0.08 \\ & \text{HD 021048} & 1.14 \pm 0.08 & 0.57 \pm 0.18 & 0.31 \pm 0.46 & 0.82 \pm 0.08 & 2.49 \pm 0.09 & 0.34 \pm 0.05 & 0.33 \pm 0.06 & 0.72 \pm 0.05 \\ & \text{HD 1021048} & 1.14 \pm 0.08 & 0.57 \pm 0.18 & 0.31 \pm 0.46 & 0.82 \pm 0.08 & 2.49 \pm 0.09 & 0.34 \pm 0.05 & 0.35 \pm 0.08 \\ & \text{HD 1021049} & 1.01 \pm 0.04 & 0.47 \pm 0.07 \pm 0.05 \pm 0.08 & 0.07 \pm 0.05 & 0.08 \pm 0.08 \\ & \text{HD 113896} & 1.01 \pm 0.06 & 0.77 \pm 0.07 \pm 0.08 \pm 0.08 & 0.25 \pm 0.08 & 0.32 \pm 0.01 & 0.32 \pm 0.08 & 0.33 \pm 0.01 & 0.32 \pm 0.03 & 0.32 \pm 0.08 & 0.33 \pm 0.01 & 0.32 \pm 0.03 & 0.32 \pm 0.08 & 0.33 \pm 0.01 & 0.32 \pm 0.03 & 0.03 \pm 0.03 & 0$	HD 088630								
$ \begin{aligned} & \text{HD108474} & 0.82 \pm 0.07 & 0.89 \pm 0.15 & -0.47 \pm 0.39 & 0.74 \pm 0.06 & 1.41 \pm 0.08 & 0.34 \pm 0.05 & 0.33 \pm 0.05 & 0.72 \pm 0.06 \\ & \text{HD1010097} & 1.01 \pm 0.04 & 0.42 \pm 0.24 & 1.10 \pm 0.59 & 1.10 \pm 0.11 & 3.00 \pm 0.10 & 0.25 \pm 0.07 & 0.25 \pm 0.07 & 0.25 \pm 0.08 \\ & \text{HD1010989} & 1.04 \pm 0.06 & 0.71 \pm 0.07 & 0.82 \pm 0.08 & 2.04 \pm 0.07 & 0.25 \pm 0.07 & 0.$									
$ \begin{aligned} & \text{HD} 1020188 & 1.14 \pm 0.08 & 0.57 \pm 0.18 & 0.31 \pm 0.46 & 0.82 \pm 0.08 & 2.49 \pm 0.09 & 0.36 \pm 0.06 & 0.44 \pm 0.07 & 0.85 \pm 0.08 \\ & \text{HD} 103996 & 1.04 \pm 0.06 & 0.71 \pm 0.17 & 3.87 \pm 0.42 & 0.75 \pm 0.07 & 1.92 \pm 0.08 & 0.32 \pm 0.05 & 0.37 \pm 0.06 & 0.71 \pm 0.05 \\ & \text{HD} 138996 & 1.04 \pm 0.06 & 0.71 \pm 0.17 & 3.87 \pm 0.42 & 0.75 \pm 0.07 & 1.92 \pm 0.08 & 0.32 \pm 0.05 & 0.37 \pm 0.06 & 0.71 \pm 0.05 \\ & \text{HD} 114946 & 1.09 \pm 0.06 & 0.62 \pm 0.18 & 4.81 \pm 0.24 & 0.78 \pm 0.08 & 2.29 \pm 0.08 & 0.26 \pm 0.05 & 0.27 \pm 0.06 & 0.88 \pm 0.05 \\ & \text{HD} 114946 & 1.09 \pm 0.06 & 0.62 \pm 0.18 & 4.81 \pm 0.24 & 0.78 \pm 0.08 & 2.29 \pm 0.08 & 0.26 \pm 0.05 & 0.27 \pm 0.06 & 0.88 \pm 0.05 \\ & \text{HD} 104979 & 0.87 \pm 0.09 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.66 \pm 0.06 & 1.32 \pm 0.07 & 0.18 \pm 0.06 & 0.20 \pm 0.07 & 0.76 \pm 0.04 \\ & \text{HD} 104979 & 0.87 \pm 0.00 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.66 \pm 0.06 & 1.32 \pm 0.07 & 0.04 & 0.25 \pm 0.06 & 0.67 \pm 0.04 \\ & \text{HD} 104079 & 0.67 \pm 0.06 & 0.61 \pm 0.07 & 1.27 \pm 0.09 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD} 104020 & 0.67 \pm 0.07 & 0.07 \pm 0.02 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD} 104020 & 0.67 \pm 0.07 & 0.07 \pm 0.02 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD} 104000 & 1.15 \pm 0.09 & 0.72 \pm 0.02 & 0.92 \pm 0.47 & 0.69 \pm 0.08 & 1.56 \pm 0.08 & 0.03 \pm 0.06 & 0.39 \pm 0.07 & 0.24 \pm 0.06 \\ & \text{HD} 104000 & 1.15 \pm 0.09 & 0.07 \pm 0.02 & 0.07 \pm 0.06 & 1.26 \pm 0.09 & 1.85 \pm 0.07 & 0.00 \pm 0.00 & 0.39 \pm 0.07 & 0.27 \pm 0.06 \\ & \text{HD} 104001 & 1.16 \pm 0.00 & 0.33 \pm 0.05 & 0.95 \pm 0.02 & 0.07 \pm 0.00 & 0.04 \pm 0.00 & 0.39 \pm 0.07 & 0.27 \pm 0.06 \\ & \text{HD} 104001 & 1.12 \pm 0.00 & 0.33 \pm 0.05 & 0.95 \pm 0.05 & 0.09 \pm 0.02 & 0.00 \pm 0.00 & 0.0$		1.09 ± 0.07		4.64 ± 0.42		2.06 ± 0.08			
$ \begin{aligned} & \text{HD} 1020188 & 1.14 \pm 0.08 & 0.57 \pm 0.18 & 0.31 \pm 0.46 & 0.82 \pm 0.08 & 2.49 \pm 0.09 & 0.36 \pm 0.06 & 0.44 \pm 0.07 & 0.85 \pm 0.08 \\ & \text{HD} 103996 & 1.04 \pm 0.06 & 0.71 \pm 0.17 & 3.87 \pm 0.42 & 0.75 \pm 0.07 & 1.92 \pm 0.08 & 0.32 \pm 0.05 & 0.37 \pm 0.06 & 0.71 \pm 0.05 \\ & \text{HD} 138996 & 1.04 \pm 0.06 & 0.71 \pm 0.17 & 3.87 \pm 0.42 & 0.75 \pm 0.07 & 1.92 \pm 0.08 & 0.32 \pm 0.05 & 0.37 \pm 0.06 & 0.71 \pm 0.05 \\ & \text{HD} 114946 & 1.09 \pm 0.06 & 0.62 \pm 0.18 & 4.81 \pm 0.24 & 0.78 \pm 0.08 & 2.29 \pm 0.08 & 0.26 \pm 0.05 & 0.27 \pm 0.06 & 0.88 \pm 0.05 \\ & \text{HD} 114946 & 1.09 \pm 0.06 & 0.62 \pm 0.18 & 4.81 \pm 0.24 & 0.78 \pm 0.08 & 2.29 \pm 0.08 & 0.26 \pm 0.05 & 0.27 \pm 0.06 & 0.88 \pm 0.05 \\ & \text{HD} 104979 & 0.87 \pm 0.09 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.66 \pm 0.06 & 1.32 \pm 0.07 & 0.18 \pm 0.06 & 0.20 \pm 0.07 & 0.76 \pm 0.04 \\ & \text{HD} 104979 & 0.87 \pm 0.00 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.66 \pm 0.06 & 1.32 \pm 0.07 & 0.04 & 0.25 \pm 0.06 & 0.67 \pm 0.04 \\ & \text{HD} 104079 & 0.67 \pm 0.06 & 0.61 \pm 0.07 & 1.27 \pm 0.09 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD} 104020 & 0.67 \pm 0.07 & 0.07 \pm 0.02 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD} 104020 & 0.67 \pm 0.07 & 0.07 \pm 0.02 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD} 104000 & 1.15 \pm 0.09 & 0.72 \pm 0.02 & 0.92 \pm 0.47 & 0.69 \pm 0.08 & 1.56 \pm 0.08 & 0.03 \pm 0.06 & 0.39 \pm 0.07 & 0.24 \pm 0.06 \\ & \text{HD} 104000 & 1.15 \pm 0.09 & 0.07 \pm 0.02 & 0.07 \pm 0.06 & 1.26 \pm 0.09 & 1.85 \pm 0.07 & 0.00 \pm 0.00 & 0.39 \pm 0.07 & 0.27 \pm 0.06 \\ & \text{HD} 104001 & 1.16 \pm 0.00 & 0.33 \pm 0.05 & 0.95 \pm 0.02 & 0.07 \pm 0.00 & 0.04 \pm 0.00 & 0.39 \pm 0.07 & 0.27 \pm 0.06 \\ & \text{HD} 104001 & 1.12 \pm 0.00 & 0.33 \pm 0.05 & 0.95 \pm 0.05 & 0.09 \pm 0.02 & 0.00 \pm 0.00 & 0.0$	HD 018474	0.82 ± 0.07	0.80 ± 0.15	-0.47 ± 0.39	0.74 ± 0.06	1.41 ± 0.08	0.34 ± 0.05	0.33 ± 0.05	0.72 ± 0.06
$\begin{array}{c} \text{HD1010997} & 1.01 \pm 0.04 & 0.42 \pm 0.24 & 1.10 \pm 0.59 & 1.10 \pm 0.11 & 3.00 \pm 0.10 & 0.25 \pm 0.07 & 0.93 \pm 0.07 \\ \text{HD1182094} & 1.17 \pm 0.07 & 0.83 \pm 0.19 & 5.25 \pm 0.46 & 0.82 \pm 0.08 & 3.23 \pm 0.08 & 0.35 \pm 0.04 & 0.42 \pm 0.05 & 0.88 \pm 0.05 \\ \text{HD118404} & 1.09 \pm 0.00 & 0.02 \pm 0.01 & 8.18 \pm 0.04 & 0.82 \pm 0.08 & 3.23 \pm 0.08 & 0.35 \pm 0.04 & 0.42 \pm 0.05 & 0.88 \pm 0.05 \\ \text{HD1000618} & 1.13 \pm 0.04 & 0.70 \pm 0.01 & 3.89 \pm 0.42 & 0.83 \pm 0.05 & 1.57 \pm 0.08 & 0.32 \pm 0.04 & 0.25 \pm 0.05 & 0.76 \pm 0.04 \\ \text{HD100737} & 0.87 \pm 0.00 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.66 \pm 0.06 & 1.20 \pm 0.06 & 0.20 \pm 0.06 \\ \text{HD157222} & 1.02 \pm 0.04 & 0.81 \pm 0.16 & 7.25 \pm 0.39 & 0.73 \pm 0.06 & 1.75 \pm 0.09 & 0.73 \pm 0.06 & 0.20 \pm 0.06 & 0.05 \pm 0.05 & 0.05 \\ \text{HD1003481} & 1.00 \pm 0.05 & 0.06 \pm 0.17 & 3.26 \pm 0.42 & 0.73 \pm 0.06 & 0.17 \pm 0.06 & 0.04 \pm 0.06 & 0.04 \pm 0.06 \\ \text{HD150209} & 1.66 \pm 0.15 & 1.03 \pm 0.18 & 17.81 \pm 0.40 & 0.50 \pm 0.02 & 2.04 \pm 0.08 & 0.26 \pm 0.06 & 0.29 \pm 0.06 & 0.07 \pm 0.06 & 0.07 \pm 0.05 & 0.07 \pm 0.06 \\ \text{HD100006} & 1.15 \pm 0.09 & 0.72 \pm 0.29 & 0.56 \pm 0.47 & 0.85 \pm 0.07 & 2.10 \pm 0.06 & 0.04 \pm 0.06 & 0.04 \pm 0.06 & 0.08 \pm 0.05 & 0.07 \pm 0.06 & 0.07 \pm 0.05 & 0.08 \pm 0.05 & 0.07 \pm 0.06 & 0.07 \pm 0.06 & 0.08 \pm 0.05 & 0.08 \pm 0.05 & 0.07 \pm 0.06 & 0.08 \pm 0.05 & 0.08 \pm 0.05$									
$\begin{array}{c} \mathrm{HD198996} 1.04 \pm 0.06 0.71 \pm 0.07 3.87 \pm 0.04 0.75 \pm 0.07 1.92 \pm 0.08 0.32 \pm 0.05 0.37 \pm 0.06 0.71 \pm 0.05 \\ \mathrm{HD114946} 1.09 \pm 0.06 0.62 \pm 0.18 4.81 \pm 0.04 0.78 \pm 0.04 2.29 \pm 0.08 0.26 \pm 0.05 0.27 \pm 0.06 0.88 \pm 0.05 \\ \mathrm{HD114946} 1.09 \pm 0.06 0.62 \pm 0.18 4.81 \pm 0.04 0.78 \pm 0.04 2.29 \pm 0.08 0.26 \pm 0.05 0.27 \pm 0.06 0.80 \pm 0.04 \\ \mathrm{HD104979} 0.87 \pm 0.09 0.86 \pm 0.17 7.65 \pm 0.42 0.60 \pm 0.06 1.32 \pm 0.07 0.01 \pm 0.06 0.20 \pm 0.07 0.71 \pm 0.06 \\ \mathrm{HD104481} 0.90 \pm 0.06 0.66 \pm 0.01 3.25 \pm 0.39 0.73 \pm 0.06 1.52 \pm 0.07 0.37 \pm 0.04 0.26 \pm 0.06 0.68 \pm 0.05 \\ \mathrm{HD094481} 0.90 \pm 0.06 0.66 \pm 0.17 3.26 \pm 0.42 0.73 \pm 0.07 0.27 \pm 0.09 0.26 \pm 0.06 0.29 \pm 0.06 0.67 \pm 0.05 \\ \mathrm{HD1022093} 1.10 \pm 0.06 0.75 \pm 0.02 9.60 \pm 0.04 0.05 \pm 0.02 2.04 \pm 0.08 0.60 \pm 0.08 0.04 \pm 0.09 \\ \mathrm{HD122093} 1.10 \pm 0.06 0.75 \pm 0.02 9.60 \pm 0.04 0.71 \pm 0.06 0.08 \pm 0.05 \\ \mathrm{HD1022093} 1.10 \pm 0.06 0.75 \pm 0.02 9.60 \pm 0.04 0.71 \pm 0.08 0.06 \pm 0.08 0.06 0.06 0.08 \pm 0.05 \\ \mathrm{HD1022093} 1.18 \pm 0.06 0.83 \pm 0.23 8.95 \pm 0.55 1.66 0.09 2.18 \pm 0.08 0.06 0.06 0.08 0.04 \pm 0.09 \\ \mathrm{HD142091} 1.62 \pm 0.10 1.19 \pm 0.27 7.04 \pm 0.06 1.26 \pm 0.09 1.85 \pm 0.09 0.39 \pm 0.07 0.47 \pm 0.08 \\ \mathrm{HD124091} 1.62 \pm 0.10 1.19 \pm 0.27 7.04 \pm 0.06 1.26 \pm 0.09 1.85 \pm 0.09 0.39 \pm 0.07 0.47 \pm 0.08 0.05 \pm 0.05 \\ \mathrm{HD124097} 1.23 \pm 0.11 1.28 \pm 0.18 1.55 \pm 0.44 0.88 \pm 0.08 0.72 \pm 0.09 0.49 \pm 0.07 0.47 \pm 0.08 0.05 \pm 0.05 \\ \mathrm{HD124097} 1.23 \pm 0.11 1.28 \pm 0.18 1.55 \pm 0.44 0.88 \pm 0.08 0.72 \pm 0.09 0.08 \pm 0.05 0.05 \pm 0.05 0.05 \pm 0.05 \\ \mathrm{HD124097} 1.23 \pm 0.11 1.28 \pm 0.18 1.55 \pm 0.44 0.88 \pm 0.08 0.72 \pm 0.07 0.47 \pm 0.08 0.05 \pm 0.05 0.05 \pm 0.05 $									
HB1128994 1.17 ± 0.07 0.83 ± 0.19 5.25 ± 0.46 0.82 ± 0.08 3.23 ± 0.08 0.35 ± 0.04 0.42 ± 0.05 0.80 ± 0.04 HB1020618 1.13 ± 0.04 0.70 ± 0.17 3.89 ± 0.42 0.83 ± 0.05 1.57 ± 0.08 0.32 ± 0.04 0.25 ± 0.05 0.76 ± 0.04 HB102079 0.87 ± 0.09 0.86 ± 0.17 7.55 ± 0.39 0.73 ± 0.06 0.32 ± 0.04 0.25 ± 0.07 0.71 ± 0.06 HB155722 1.02 ± 0.04 0.81 ± 0.16 7.25 ± 0.39 0.73 ± 0.06 1.75 ± 0.07 0.71 ± 0.06 HB155722 1.02 ± 0.04 0.81 ± 0.16 7.25 ± 0.39 0.73 ± 0.06 1.75 ± 0.07 0.71 ± 0.06 0.04 0.05 ± 0.05 0.06 0.07 ± 0.05 0.06 0.07 ± 0.05 0.05 0.07 ± 0.09 0.06 0.06 0.07 ± 0.05 0.07 0.07 0.07 ± 0.09 0.06 0.07 ± 0.05 0.07 0.07 0.07 ± 0.09 0.06 0.07 ± 0.05 0.07 0.07 ± 0.09 0.06 0.07 ± 0.05 0.07 0.07 ± 0.07 0.07 ± 0.07 0.07 ± 0.07 0.07	HD 010697	1.01 ± 0.04	0.42 ± 0.24	1.10 ± 0.59		3.00 ± 0.10	0.25 ± 0.07	0.26 ± 0.07	0.93 ± 0.07
HB1128994 1.17 ± 0.07 0.83 ± 0.19 5.25 ± 0.46 0.82 ± 0.08 3.23 ± 0.08 0.35 ± 0.04 0.42 ± 0.05 0.80 ± 0.04 HB1020618 1.13 ± 0.04 0.70 ± 0.17 3.89 ± 0.42 0.83 ± 0.05 1.57 ± 0.08 0.32 ± 0.04 0.25 ± 0.05 0.76 ± 0.04 HB102079 0.87 ± 0.09 0.86 ± 0.17 7.55 ± 0.39 0.73 ± 0.06 0.32 ± 0.04 0.25 ± 0.07 0.71 ± 0.06 HB155722 1.02 ± 0.04 0.81 ± 0.16 7.25 ± 0.39 0.73 ± 0.06 1.75 ± 0.07 0.71 ± 0.06 HB155722 1.02 ± 0.04 0.81 ± 0.16 7.25 ± 0.39 0.73 ± 0.06 1.75 ± 0.07 0.71 ± 0.06 0.04 0.05 ± 0.05 0.06 0.07 ± 0.05 0.06 0.07 ± 0.05 0.05 0.07 ± 0.09 0.06 0.06 0.07 ± 0.05 0.07 0.07 0.07 ± 0.09 0.06 0.07 ± 0.05 0.07 0.07 0.07 ± 0.09 0.06 0.07 ± 0.05 0.07 0.07 ± 0.09 0.06 0.07 ± 0.05 0.07 0.07 ± 0.07 0.07 ± 0.07 0.07 ± 0.07 0.07	HD 193896	1.04 ± 0.06	0.71 ± 0.17	3.87 ± 0.42	0.75 ± 0.07	1.92 ± 0.08	0.32 ± 0.05	0.37 ± 0.06	0.71 ± 0.05
$\begin{aligned} & \text{HD114946} & 1.09 \pm 0.06 & 0.62 \pm 0.18 & 4.81 \pm 0.44 & 0.78 \pm 0.04 & 2.29 \pm 0.08 & 0.26 \pm 0.05 & 0.77 \pm 0.06 & 0.80 \pm 0.04 \\ & \text{HD104979} & 0.87 \pm 0.09 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.60 \pm 0.06 & 1.32 \pm 0.07 & 0.04 & 0.25 \pm 0.05 & 0.76 \pm 0.04 \\ & \text{HD104979} & 0.87 \pm 0.09 & 0.86 \pm 0.17 & 7.65 \pm 0.42 & 0.60 \pm 0.06 & 1.32 \pm 0.07 & 0.01 \pm 0.06 & 0.06 \pm 0.06 & 0.06 \pm 0.07 & 0.07 \pm 0.06 \\ & \text{HD104481} & 0.90 \pm 0.06 & 0.66 \pm 0.01 & 3.26 \pm 0.42 & 0.73 \pm 0.06 & 1.52 \pm 0.07 & 0.03 \pm 0.06 & 0.25 \pm 0.06 & 0.68 \pm 0.05 \\ & \text{HD104481} & 0.90 \pm 0.06 & 0.66 \pm 0.01 & 3.26 \pm 0.42 & 0.73 \pm 0.07 & 2.01 \pm 0.08 & 0.06 \pm 0.08 & 0.04 \pm 0.09 \\ & \text{HD1020973} & 1.10 \pm 0.06 & 0.75 \pm 0.20 & 9.06 \pm 0.47 & 0.85 \pm 0.07 & 2.10 \pm 0.07 & 0.42 \pm 0.06 & 0.44 \pm 0.09 & 0.64 \pm 0.04 \\ & \text{HD1202093} & 1.10 \pm 0.06 & 0.75 \pm 0.20 & 9.06 \pm 0.47 & 0.85 \pm 0.07 & 2.10 \pm 0.07 & 0.42 \pm 0.06 & 0.43 \pm 0.06 & 0.88 \pm 0.05 \\ & \text{HD10205973} & 1.18 \pm 0.06 & 0.83 \pm 0.23 & 5.29 \pm 0.55 & 1.06 \pm 0.09 & 1.85 \pm 0.09 & 0.38 \pm 0.06 & 0.47 \pm 0.07 & 0.81 \pm 0.05 \\ & \text{HD1164381} & 1.43 \pm 0.09 & 1.00 \pm 0.17 & 10.79 \pm 0.41 & 0.71 \pm 0.08 & 1.74 \pm 0.07 & 0.50 \pm 0.05 & 0.05 & 0.04 & 1.74 \pm 0.05 \\ & \text{HD1164381} & 1.64 \pm 0.08 & 0.85 \pm 0.25 & 6.09 \pm 0.03 & 1.30 \pm 0.09 & 1.33 \pm 0.09 & 0.33 \pm 0.07 & 0.47 \pm 0.05 & 1.32 \pm 0.09 \\ & \text{HD1164381} & 1.64 \pm 0.08 & 0.95 \pm 0.26 & 6.09 \pm 0.03 & 1.30 \pm 0.09 & 1.33 \pm 0.09 & 0.33 \pm 0.07 & 0.47 \pm 0.05 & 1.32 \pm 0.09 \\ & \text{HD1164381} & 1.64 \pm 0.08 & 0.95 \pm 0.26 & 6.09 \pm 0.06 & 1.26 \pm 0.09 & 1.85 \pm 0.09 & 0.39 \pm 0.07 & 0.47 \pm 0.07 & 1.17 \pm 0.08 \\ & \text{HD1020971} & 1.32 \pm 0.18 & 1.5.39 \pm 0.01 & 0.88 \pm 0.05 & 0.72 \pm 0.07 & 0.49 \pm 0.07 & 0.30 \pm 0.07 & 0.65 \pm 0.03 \\ & \text{HD1137789} & 1.23 \pm 0.18 & 1.5.49 \pm 0.18 & 1.5.39 \pm 0.01 & 0.87 \pm 0.09 & 0.30 \pm 0.07 & 0.47 \pm 0.07 & 0.55 \pm 0.02 \\ & \text{HD1178208} & 2.99 \pm 0.14 & 1.58 \pm 0.27 & 1.579 \pm 0.01 & 1.00 \pm 0.08 & 0.05 \pm 0.00 & 0.03 \pm 0.07 & 0.07 \pm 0.04 \\ & \text{HD119409} & 3.27 \pm 0.18 & 1.58 \pm 0.27 & 1.579 \pm 0.01 & 1.00 \pm 0.08 & 0.07 \pm 0.00 & 0.07 \pm 0.00 & 0.07 \pm 0.00 \\ & \text{HD118700} & 3.27 \pm 0.18 & 1.09 \pm $									
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$									
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 114946	1.09 ± 0.06	0.62 ± 0.18	4.81 ± 0.44	0.78 ± 0.04	2.29 ± 0.08	0.26 ± 0.05	0.27 ± 0.06	0.80 ± 0.04
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 020618	1.13 ± 0.04	0.70 ± 0.17	3.89 ± 0.42	0.83 ± 0.05		0.32 ± 0.04	0.25 ± 0.05	0.76 ± 0.04
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 135722	1.02 ± 0.04	0.81 ± 0.16	7.25 ± 0.39	0.73 ± 0.06	1.75 ± 0.07	0.37 ± 0.04	0.36 ± 0.05	0.68 ± 0.05
$\begin{array}{c} \mathrm{HD} 170820 \\ \mathrm{HD} 170820 \\ \mathrm{I} 1.09 \pm 0.00 \\ \mathrm{I} 1.00 \pm 0.00 \\ \mathrm{I} 1.0$	HD 004481	0.00 ± 0.06							0.67 ± 0.05
$\begin{array}{c} \mathrm{HB} 1222993 1.10 \pm 0.06 0.75 \pm 0.20 9.60 \pm 0.47 0.85 \pm 0.07 2.10 \pm 0.07 0.42 \pm 0.06 0.43 \pm 0.06 0.88 \pm 0.05 \\ \mathrm{HD} 1036134 1.43 \pm 0.09 1.00 \pm 0.17 10.79 \pm 0.41 0.71 \pm 0.08 1.74 \pm 0.07 0.50 \pm 0.05 0.47 \pm 0.06 0.81 \pm 0.05 \\ \mathrm{HD} 1025975 1.18 \pm 0.06 0.83 \pm 0.23 5.92 \pm 0.55 1.06 \pm 0.09 2.18 \pm 0.09 0.39 \pm 0.06 0.47 \pm 0.05 0.81 \pm 0.05 \\ \mathrm{HD} 1124991 1.62 \pm 0.10 1.19 \pm 0.27 7.04 \pm 0.66 1.26 \pm 0.09 1.85 \pm 0.09 0.39 \pm 0.07 0.47 \pm 0.08 1.32 \pm 0.09 \\ \mathrm{HD} 1124897 1.23 \pm 0.13 1.28 \pm 0.18 1.55 \pm 0.41 0.88 \pm 0.06 0.72 \pm 0.07 0.49 \pm 0.07 0.30 \pm 0.07 0.47 \pm 0.08 1.32 \pm 0.09 \\ \mathrm{HD} 1037935 1.71 \pm 0.06 1.33 \pm 0.13 1.48 \pm 0.99 0.85 \pm 0.05 0.63 \pm 0.06 0.62 \pm 0.05 0.47 \pm 0.07 0.30 \pm 0.07 0.47 \pm 0.08 \\ \mathrm{HD} 1337935 1.71 \pm 0.06 1.33 \pm 0.17 15.39 \pm 0.40 0.87 \pm 0.09 0.53 \pm 0.05 0.75 \pm 0.04 0.55 \pm 0.04 0.74 \pm 0.08 \\ \mathrm{HD} 103793 1.92 \pm 0.11 1.28 \pm 0.24 1.138 \pm 0.57 0.91 \pm 0.13 1.04 \pm 0.07 0.55 \pm 0.05 0.75 \pm 0.04 0.55 \pm 0.04 0.74 \pm 0.04 \\ \mathrm{HD} 103793 1.92 \pm 0.11 1.53 \pm 0.23 1.847 \pm 0.52 0.91 \pm 0.28 1.24 \pm 0.11 0.85 \pm 0.08 0.48 \pm 0.09 0.39 \pm 0.07 0.49 \pm 0.09 0.74 \pm 0.08 \\ \mathrm{HD} 103793 1.92 \pm 0.11 1.53 \pm 0.23 1.847 \pm 0.52 0.91 \pm 0.28 1.24 \pm 0.11 0.85 \pm 0.08 0.48 \pm 0.09 0.73 \pm 0.04 0.55 \pm 0.04 0.74 \pm 0.04 \\ \mathrm{HD} 114960 3.27 \pm 0.18 1.62 \pm 0.22 17.19 \pm 0.50 0.08 \pm 0.17 0.05 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.08 0.09 0.08 0.09 $									
$\begin{array}{c} \mathrm{HD100006} 1.15\pm0.09 0.72\pm0.20 9.26\pm0.47 0.69\pm0.08 1.56\pm0.08 0.39\pm0.06 0.39\pm0.07 0.74\pm0.06 0.81\pm0.05 \\ \mathrm{HD1025975} 1.18\pm0.06 0.83\pm0.33 5.92\pm0.55 1.06\pm0.09 2.18\pm0.09 0.38\pm0.06 0.40\pm0.07 1.09\pm0.06 0.91\pm0.05 \\ \mathrm{HD1025975} 1.18\pm0.06 0.83\pm0.33 5.92\pm0.55 1.06\pm0.09 2.18\pm0.09 0.38\pm0.06 0.40\pm0.07 1.09\pm0.06 0.91\pm0.05 0.09\pm0.01 1.09\pm0.07 0.09\pm0.01 $	HD 170820	1.66 ± 0.16	1.03 ± 0.18	17.81 ± 0.40		2.04 ± 0.08	0.60 ± 0.08		0.64 ± 0.04
$\begin{array}{c} \mathrm{HD100006} 1.15\pm0.09 0.72\pm0.20 9.26\pm0.47 0.69\pm0.08 1.56\pm0.08 0.39\pm0.06 0.39\pm0.07 0.74\pm0.06 0.81\pm0.05 \\ \mathrm{HD1025975} 1.18\pm0.06 0.83\pm0.33 5.92\pm0.55 1.06\pm0.09 2.18\pm0.09 0.38\pm0.06 0.40\pm0.07 1.09\pm0.06 0.91\pm0.05 \\ \mathrm{HD1025975} 1.18\pm0.06 0.83\pm0.33 5.92\pm0.55 1.06\pm0.09 2.18\pm0.09 0.38\pm0.06 0.40\pm0.07 1.09\pm0.06 0.91\pm0.05 0.09\pm0.01 1.09\pm0.07 0.09\pm0.01 $	HD 222093	1.10 ± 0.06	0.75 ± 0.20	9.60 ± 0.47	0.85 ± 0.07	2.10 ± 0.07	0.42 ± 0.06	0.43 ± 0.06	0.88 ± 0.05
$\begin{array}{c} \mathrm{HD026915} \\ \mathrm{HD025915} \\ \mathrm{L184} \\ \mathrm{L006} \\ \mathrm{OR} \\$									
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$\begin{array}{c} \mathrm{HD} 1142091 & 1.62 \pm 0.10 & 1.19 \pm 0.27 & 7.04 \pm 0.66 & 1.26 \pm 0.09 & 1.83 \pm 0.09 & 0.32 \pm 0.07 & 0.47 \pm 0.07 & 1.17 \pm 0.08 \\ \mathrm{HD} 11624897 & 1.23 \pm 0.13 & 1.28 \pm 0.18 & 15.55 \pm 0.41 & 0.88 \pm 0.06 & 0.72 \pm 0.07 & 0.49 \pm 0.07 & 0.30 \pm 0.06 & 0.03 \pm 0.06 \\ \mathrm{HD} 1123935 & 1.71 \pm 0.06 & 1.33 \pm 0.17 & 15.39 \pm 0.40 & 0.87 \pm 0.09 & 0.53 \pm 0.05 & 0.30 \pm 0.06 & 0.73 \pm 0.08 \\ \mathrm{HD} 1132935 & 1.71 \pm 0.06 & 1.33 \pm 0.17 & 15.39 \pm 0.40 & 0.87 \pm 0.09 & 0.53 \pm 0.05 & 0.73 \pm 0.03 \pm 0.06 & 0.73 \pm 0.03 \\ \mathrm{HD} 1132935 & 1.71 \pm 0.06 & 1.33 \pm 0.17 & 15.39 \pm 0.40 & 0.87 \pm 0.09 & 0.53 \pm 0.05 & 0.73 \pm 0.04 & 0.55 \pm 0.04 & 0.74 \pm 0.07 \\ \mathrm{HD} 1137759 & 1.92 \pm 0.11 & 1.28 \pm 0.24 & 11.38 \pm 0.57 & 0.91 \pm 0.13 & 1.04 \pm 0.07 & 0.55 \pm 0.05 & 0.55 \pm 0.04 & 0.74 \pm 0.05 \\ \mathrm{HD} 1132026 & 2.24 \pm 0.19 & 1.58 \pm 0.22 & 11.59 \pm 0.05 & 0.91 \pm 0.28 & 1.24 \pm 0.11 & 0.85 \pm 0.08 & 0.48 \pm 0.09 & 0.93 \pm 0.07 \\ \mathrm{HD} 113208 & 2.90 \pm 0.14 & 1.58 \pm 0.27 & 15.59 \pm 0.61 & 1.00 \pm 0.18 & 0.85 \pm 0.09 & 0.80 \pm 0.08 & 0.62 \pm 0.09 & 0.93 \pm 0.07 \\ \mathrm{HD} 114960 & 3.27 \pm 0.18 & 1.96 \pm 0.26 & 17.19 \pm 0.59 & 0.98 \pm 0.23 & 0.72 \pm 0.08 & 0.97 \pm 0.08 & 0.61 \pm 0.06 & 0.64 \pm 0.04 \\ \mathrm{HD} 1181596 & 3.38 \pm 0.12 & 2.06 \pm 0.19 & 20.11 \pm 0.43 & 0.80 \pm 0.18 & 0.26 \pm 0.06 & 1.22 \pm 0.08 & 0.64 \pm 0.09 & 0.74 \pm 0.06 \\ \mathrm{HD} 102477 & 3.52 \pm 0.17 & 2.35 \pm 0.22 & 20.57 \pm 0.48 & 1.12 \pm 0.23 & 0.46 \pm 0.08 & 1.14 \pm 0.07 & 0.68 \pm 0.018 & 0.93 \pm 0.07 & 0.68 \pm 0.07 & 0.07 \pm 0.00 \\ \mathrm{HD} 102493 & 3.27 \pm 0.14 & 1.23 \pm 0.18 & 20.77 \pm 0.40 & 0.83 \pm 0.17 & 0.15 \pm 0.06 & 1.17 \pm 0.08 & 0.62 \pm 0.09 & 0.70 \pm 0.07 \\ \mathrm{HD} 102493 & 3.28 \pm 0.09 & 2.03 \pm 0.18 & 20.28 \pm 0.40 & 0.95 \pm 0.12 & 0.82 \pm 0.05 & 1.20 \pm 0.08 & 0.53 \pm 0.09 & 0.66 \pm 0.05 \\ \mathrm{HD} 202734 & 4.17 \pm 0.16 & 2.82 \pm 0.19 & 22.48 \pm 0.42 & 0.77 \pm 0.22 & 0.12 \pm 0.07 & 1.47 \pm 0.09 & 0.65 \pm 0.05 & 0.71 \pm 0.04 \\ \mathrm{HD} 1020952 & 3.18 \pm 0.13 & 2.46 \pm 0.17 & 23.60 \pm 0.33 & 1.04 \pm 0.11 & 1.07 \pm 0.12 & 0.05 \pm 0.01 & 0.75 \pm 0.01 & 0.75$	HD 036134	1.43 ± 0.09	1.00 ± 0.17	10.79 ± 0.41	0.71 ± 0.08	1.74 ± 0.07	0.50 ± 0.05	0.47 ± 0.05	0.81 ± 0.05
$\begin{array}{c} \mathrm{HD} 1142091 & 1.62 \pm 0.10 & 1.19 \pm 0.27 & 7.04 \pm 0.66 & 1.26 \pm 0.09 & 1.83 \pm 0.09 & 0.32 \pm 0.07 & 0.47 \pm 0.07 & 1.17 \pm 0.08 \\ \mathrm{HD} 11624897 & 1.23 \pm 0.13 & 1.28 \pm 0.18 & 15.55 \pm 0.41 & 0.88 \pm 0.06 & 0.72 \pm 0.07 & 0.49 \pm 0.07 & 0.30 \pm 0.06 & 0.03 \pm 0.06 \\ \mathrm{HD} 1123935 & 1.71 \pm 0.06 & 1.33 \pm 0.17 & 15.39 \pm 0.40 & 0.87 \pm 0.09 & 0.53 \pm 0.05 & 0.30 \pm 0.06 & 0.73 \pm 0.08 \\ \mathrm{HD} 1132935 & 1.71 \pm 0.06 & 1.33 \pm 0.17 & 15.39 \pm 0.40 & 0.87 \pm 0.09 & 0.53 \pm 0.05 & 0.73 \pm 0.03 \pm 0.06 & 0.73 \pm 0.03 \\ \mathrm{HD} 1132935 & 1.71 \pm 0.06 & 1.33 \pm 0.17 & 15.39 \pm 0.40 & 0.87 \pm 0.09 & 0.53 \pm 0.05 & 0.73 \pm 0.04 & 0.55 \pm 0.04 & 0.74 \pm 0.07 \\ \mathrm{HD} 1137759 & 1.92 \pm 0.11 & 1.28 \pm 0.24 & 11.38 \pm 0.57 & 0.91 \pm 0.13 & 1.04 \pm 0.07 & 0.55 \pm 0.05 & 0.55 \pm 0.04 & 0.74 \pm 0.05 \\ \mathrm{HD} 1132026 & 2.24 \pm 0.19 & 1.58 \pm 0.22 & 11.59 \pm 0.05 & 0.91 \pm 0.28 & 1.24 \pm 0.11 & 0.85 \pm 0.08 & 0.48 \pm 0.09 & 0.93 \pm 0.07 \\ \mathrm{HD} 113208 & 2.90 \pm 0.14 & 1.58 \pm 0.27 & 15.59 \pm 0.61 & 1.00 \pm 0.18 & 0.85 \pm 0.09 & 0.80 \pm 0.08 & 0.62 \pm 0.09 & 0.93 \pm 0.07 \\ \mathrm{HD} 114960 & 3.27 \pm 0.18 & 1.96 \pm 0.26 & 17.19 \pm 0.59 & 0.98 \pm 0.23 & 0.72 \pm 0.08 & 0.97 \pm 0.08 & 0.61 \pm 0.06 & 0.64 \pm 0.04 \\ \mathrm{HD} 1181596 & 3.38 \pm 0.12 & 2.06 \pm 0.19 & 20.11 \pm 0.43 & 0.80 \pm 0.18 & 0.26 \pm 0.06 & 1.22 \pm 0.08 & 0.64 \pm 0.09 & 0.74 \pm 0.06 \\ \mathrm{HD} 102477 & 3.52 \pm 0.17 & 2.35 \pm 0.22 & 20.57 \pm 0.48 & 1.12 \pm 0.23 & 0.46 \pm 0.08 & 1.14 \pm 0.07 & 0.68 \pm 0.018 & 0.93 \pm 0.07 & 0.68 \pm 0.07 & 0.07 \pm 0.00 \\ \mathrm{HD} 102493 & 3.27 \pm 0.14 & 1.23 \pm 0.18 & 20.77 \pm 0.40 & 0.83 \pm 0.17 & 0.15 \pm 0.06 & 1.17 \pm 0.08 & 0.62 \pm 0.09 & 0.70 \pm 0.07 \\ \mathrm{HD} 102493 & 3.28 \pm 0.09 & 2.03 \pm 0.18 & 20.28 \pm 0.40 & 0.95 \pm 0.12 & 0.82 \pm 0.05 & 1.20 \pm 0.08 & 0.53 \pm 0.09 & 0.66 \pm 0.05 \\ \mathrm{HD} 202734 & 4.17 \pm 0.16 & 2.82 \pm 0.19 & 22.48 \pm 0.42 & 0.77 \pm 0.22 & 0.12 \pm 0.07 & 1.47 \pm 0.09 & 0.65 \pm 0.05 & 0.71 \pm 0.04 \\ \mathrm{HD} 1020952 & 3.18 \pm 0.13 & 2.46 \pm 0.17 & 23.60 \pm 0.33 & 1.04 \pm 0.11 & 1.07 \pm 0.12 & 0.05 \pm 0.01 & 0.75 \pm 0.01 & 0.75$	HD 025975	1.18 ± 0.06			1.06 ± 0.09		0.38 ± 0.06		1.09 ± 0.06
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 165438	1.46 ± 0.08	0.95 ± 0.26	6.09 ± 0.63	1.30 ± 0.09	1.33 ± 0.09		0.47 ± 0.07	1.17 ± 0.05
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 132935	1.71 ± 0.06	1.33 ± 0.17	15.39 ± 0.40	0.87 ± 0.09	0.53 ± 0.05	0.73 ± 0.04	0.55 ± 0.04	0.74 ± 0.04
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 137759	1.92 ± 0.11	1.28 ± 0.24		0.91 ± 0.13		0.55 ± 0.05		1.03 ± 0.10
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 178208	2.90 ± 0.14	1.58 ± 0.27	15.59 ± 0.61	1.00 ± 0.18	0.85 ± 0.09	0.80 ± 0.08	0.62 ± 0.09	1.11 ± 0.07
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 221246	2.71 ± 0.15	1.62 ± 0.22	17.19 ± 0.50	0.68 ± 0.17	0.92 ± 0.07	0.89 ± 0.07	0.64 ± 0.08	0.93 ± 0.04
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 099998		1.68 ± 0.17	18.69 ± 0.38	0.84 ± 0.13	0.58 ± 0.05	0.93 ± 0.06	0.51 ± 0.06	
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 181596	3.38 ± 0.12	2.06 ± 0.19	20.11 ± 0.43	0.80 ± 0.18	0.26 ± 0.06	1.22 ± 0.08	0.64 ± 0.09	0.74 ± 0.06
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$\begin{array}{c} \mathrm{HD} 213893 & 2.98 \pm 0.09 & 2.03 \pm 0.18 & 20.28 \pm 0.40 & 0.95 \pm 0.12 & 0.82 \pm 0.05 & 1.20 \pm 0.08 & 0.53 \pm 0.09 & 0.66 \pm 0.05 \\ \mathrm{HD} 204724 & 4.54 \pm 0.20 & 2.79 \pm 0.25 & 22.04 \pm 0.55 & 0.97 \pm 0.23 & 0.26 \pm 0.11 & 1.27 \pm 0.12 & 0.79 \pm 0.13 & 1.06 \pm 0.12 \\ \mathrm{HD} 219734 & 4.17 \pm 0.16 & 2.82 \pm 0.19 & 22.48 \pm 0.42 & 0.77 \pm 0.22 & 0.12 \pm 0.07 & 1.43 \pm 0.09 & 0.76 \pm 0.10 & 0.72 \pm 0.05 \\ \mathrm{HD} 120052 & 3.18 \pm 0.13 & 2.46 \pm 0.17 & 21.06 \pm 0.38 & 1.04 \pm 0.11 & -0.14 \pm 0.05 & 1.36 \pm 0.10 & 0.57 \pm 0.11 & 0.69 \pm 0.02 \\ \mathrm{HD} 039045 & 4.00 \pm 0.14 & 2.67 \pm 0.18 & 22.97 \pm 0.39 & 0.85 \pm 0.17 & 0.57 \pm 0.07 & 1.47 \pm 0.09 & 0.73 \pm 0.10 & 0.69 \pm 0.06 \\ \mathrm{HD} 028487 & 4.18 \pm 0.17 & 2.96 \pm 0.17 & 23.60 \pm 0.37 & 0.91 \pm 0.19 & -0.35 \pm 0.07 & 1.59 \pm 0.08 & 0.72 \pm 0.09 & 0.68 \pm 0.07 \\ \mathrm{HD} 0214665 & 4.93 \pm 0.23 & 3.46 \pm 0.19 & 25.49 \pm 0.42 & 0.87 \pm 0.33 & -0.49 \pm 0.12 & 1.64 \pm 0.08 & 0.83 \pm 0.09 & 0.66 \pm 0.01 \\ \mathrm{HD} 004408 & 5.12 \pm 0.22 & 3.35 \pm 0.19 & 24.18 \pm 0.42 & 1.02 \pm 0.29 & 0.79 \pm 0.10 & 1.74 \pm 0.09 & 0.99 \pm 0.10 & 0.68 \pm 0.09 \\ \mathrm{HD} 024585 & 5.91 \pm 0.26 & 4.03 \pm 0.23 & 26.11 \pm 0.50 & 0.99 \pm 0.34 & -0.39 \pm 0.12 & 2.02 \pm 0.13 & 0.93 \pm 0.14 & 0.58 \pm 0.16 \\ \mathrm{HD} 014386 & 2.43 \pm 0.22 & 1.51 \pm 0.42 & 24.30 \pm 0.36 & 0.99 \pm 0.34 & -5.03 \pm 0.24 & 0.90 \pm 0.19 & -0.19 \pm 0.22 & 0.39 \pm 0.16 \\ \mathrm{HD} 0175865 & 5.57 \pm 0.23 & 3.62 \pm 0.22 & 25.94 \pm 0.49 & 0.93 \pm 0.27 & 0.17 \pm 0.11 & 1.86 \pm 0.13 & 0.89 \pm 0.15 & 0.54 \pm 0.16 \\ \mathrm{HD} 0194705 & 5.92 \pm 0.25 & 4.05 \pm 0.21 & 24.43 \pm 0.45 & 1.21 \pm 0.29 & -0.01 \pm 0.12 & 0.09 \pm 0.13 & 0.93 \pm 0.14 & 0.58 \pm 0.16 \\ \mathrm{HD} 019610 & 5.03 \pm 0.17 & 3.95 \pm 0.18 & 28.36 \pm 0.40 & 0.91 \pm 0.26 & -0.15 \pm 0.10 & 2.04 \pm 0.11 & 1.05 \pm 0.13 & 0.57 \pm 0.12 \\ \mathrm{HD} 113135 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.02 & 0.08 \pm 0.14 & 0.67 \pm 0.12 \\ \mathrm{HD} 106321 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.04 & 0.26 \pm 0.02 \\ \mathrm{HD} 2131359 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.35 \pm 0.08 & 0.99 \pm 0.26 & -0.05 \pm 0.10 & 0.05 \pm 0.13 & 0.05 $				20.77 ± 0.40					0.70 ± 0.07
$\begin{array}{c} \mathrm{HD} 213893 & 2.98 \pm 0.09 & 2.03 \pm 0.18 & 20.28 \pm 0.40 & 0.95 \pm 0.12 & 0.82 \pm 0.05 & 1.20 \pm 0.08 & 0.53 \pm 0.09 & 0.66 \pm 0.05 \\ \mathrm{HD} 204724 & 4.54 \pm 0.20 & 2.79 \pm 0.25 & 22.04 \pm 0.55 & 0.97 \pm 0.23 & 0.26 \pm 0.11 & 1.27 \pm 0.12 & 0.79 \pm 0.13 & 1.06 \pm 0.12 \\ \mathrm{HD} 219734 & 4.17 \pm 0.16 & 2.82 \pm 0.19 & 22.48 \pm 0.42 & 0.77 \pm 0.22 & 0.12 \pm 0.07 & 1.43 \pm 0.09 & 0.76 \pm 0.10 & 0.72 \pm 0.05 \\ \mathrm{HD} 120052 & 3.18 \pm 0.13 & 2.46 \pm 0.17 & 21.06 \pm 0.38 & 1.04 \pm 0.11 & -0.14 \pm 0.05 & 1.36 \pm 0.10 & 0.57 \pm 0.11 & 0.69 \pm 0.02 \\ \mathrm{HD} 039045 & 4.00 \pm 0.14 & 2.67 \pm 0.18 & 22.97 \pm 0.39 & 0.85 \pm 0.17 & 0.57 \pm 0.07 & 1.47 \pm 0.09 & 0.73 \pm 0.10 & 0.69 \pm 0.06 \\ \mathrm{HD} 028487 & 4.18 \pm 0.17 & 2.96 \pm 0.17 & 23.60 \pm 0.37 & 0.91 \pm 0.19 & -0.35 \pm 0.07 & 1.59 \pm 0.08 & 0.72 \pm 0.09 & 0.68 \pm 0.07 \\ \mathrm{HD} 0214665 & 4.93 \pm 0.23 & 3.46 \pm 0.19 & 25.49 \pm 0.42 & 0.87 \pm 0.33 & -0.49 \pm 0.12 & 1.64 \pm 0.08 & 0.83 \pm 0.09 & 0.66 \pm 0.01 \\ \mathrm{HD} 004408 & 5.12 \pm 0.22 & 3.35 \pm 0.19 & 24.18 \pm 0.42 & 1.02 \pm 0.29 & 0.79 \pm 0.10 & 1.74 \pm 0.09 & 0.99 \pm 0.10 & 0.68 \pm 0.09 \\ \mathrm{HD} 024585 & 5.91 \pm 0.26 & 4.03 \pm 0.23 & 26.11 \pm 0.50 & 0.99 \pm 0.34 & -0.39 \pm 0.12 & 2.02 \pm 0.13 & 0.93 \pm 0.14 & 0.58 \pm 0.16 \\ \mathrm{HD} 014386 & 2.43 \pm 0.22 & 1.51 \pm 0.42 & 24.30 \pm 0.36 & 0.99 \pm 0.34 & -5.03 \pm 0.24 & 0.90 \pm 0.19 & -0.19 \pm 0.22 & 0.39 \pm 0.16 \\ \mathrm{HD} 0175865 & 5.57 \pm 0.23 & 3.62 \pm 0.22 & 25.94 \pm 0.49 & 0.93 \pm 0.27 & 0.17 \pm 0.11 & 1.86 \pm 0.13 & 0.89 \pm 0.15 & 0.54 \pm 0.16 \\ \mathrm{HD} 0194705 & 5.92 \pm 0.25 & 4.05 \pm 0.21 & 24.43 \pm 0.45 & 1.21 \pm 0.29 & -0.01 \pm 0.12 & 0.09 \pm 0.13 & 0.93 \pm 0.14 & 0.58 \pm 0.16 \\ \mathrm{HD} 019610 & 5.03 \pm 0.17 & 3.95 \pm 0.18 & 28.36 \pm 0.40 & 0.91 \pm 0.26 & -0.15 \pm 0.10 & 2.04 \pm 0.11 & 1.05 \pm 0.13 & 0.57 \pm 0.12 \\ \mathrm{HD} 113135 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.02 & 0.08 \pm 0.14 & 0.67 \pm 0.12 \\ \mathrm{HD} 106321 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.04 & 0.26 \pm 0.02 \\ \mathrm{HD} 2131359 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.35 \pm 0.08 & 0.99 \pm 0.26 & -0.05 \pm 0.10 & 0.05 \pm 0.13 & 0.05 $	HD 194193	3.27 ± 0.14	2.23 ± 0.18	21.03 ± 0.41	0.77 ± 0.15	-0.03 ± 0.06	0.94 ± 0.08	0.59 ± 0.08	0.71 ± 0.04
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 219734	4.17 ± 0.16	2.82 ± 0.19	22.48 ± 0.42	0.77 ± 0.22	0.12 ± 0.07	1.43 ± 0.09	0.76 ± 0.10	0.72 ± 0.05
$\begin{array}{c} \text{HD}039045 & 4,00\pm0.14 & 2.67\pm0.18 & 22.97\pm0.39 & 0.85\pm0.17 & 0.57\pm0.07 & 1.47\pm0.09 & 0.73\pm0.10 & 0.69\pm0.06 \\ \text{HD}028487 & 4.18\pm0.17 & 2.96\pm0.17 & 23.60\pm0.37 & 0.91\pm0.19 & -0.35\pm0.07 & 1.59\pm0.08 & 0.72\pm0.09 & 0.68\pm0.07 \\ \text{HD}214665 & 4.93\pm0.23 & 3.46\pm0.19 & 25.49\pm0.42 & 0.87\pm0.33 & -0.49\pm0.01 & 1.56\pm0.10 & 0.66\pm0.11 & 0.57\pm0.04 \\ \text{HD}027598 & 3.88\pm0.18 & 2.89\pm0.16 & 24.34\pm0.36 & 0.80\pm0.18 & -0.49\pm0.06 & 1.56\pm0.10 & 0.66\pm0.11 & 0.57\pm0.04 \\ \text{HD}004408 & 5.12\pm0.22 & 3.35\pm0.19 & 24.18\pm0.42 & 1.02\pm0.29 & 0.79\pm0.10 & 1.74\pm0.09 & 0.99\pm0.10 & 0.68\pm0.09 \\ \text{HD}204585 & 5.91\pm0.26 & 4.03\pm0.23 & 26.11\pm0.50 & 0.99\pm0.34 & -0.39\pm0.12 & 2.02\pm0.13 & 0.93\pm0.14 & 0.58\pm0.16 \\ \text{HD}014386 & 2.43\pm0.22 & 1.51\pm0.42 & 24.30\pm0.89 & 0.09\pm0.48 & -5.03\pm0.24 & 0.90\pm0.19 & -0.19\pm0.22 & 0.39\pm0.19 \\ \text{HD}175865 & 5.57\pm0.23 & 3.62\pm0.22 & 25.94\pm0.49 & 0.93\pm0.27 & 0.17\pm0.11 & 1.86\pm0.13 & 0.89\pm0.15 & 0.54\pm0.16 \\ \text{HD}094705 & 5.92\pm0.25 & 4.25\pm0.21 & 24.43\pm0.45 & 1.21\pm0.29 & -0.01\pm0.12 & 2.04\pm0.11 & 1.05\pm0.13 & 0.57\pm0.12 \\ \text{HD}191819 & 5.31\pm0.25 & 4.02\pm0.20 & 22.24\pm0.45 & 0.91\pm0.21 & -0.09\pm0.07 & 1.90\pm0.13 & 0.85\pm0.14 & 0.67\pm0.10 \\ \text{HD}108849 & 5.80\pm0.24 & 4.07\pm0.23 & 28.05\pm0.49 & 0.93\pm0.41 & -1.48\pm0.18 & 2.10\pm0.09 & 0.74\pm0.10 & 0.61\pm0.19 \\ \text{HD}207076 & 4.91\pm0.17 & 3.44\pm0.16 & 24.34\pm0.36 & 0.99\pm0.26 & -0.15\pm0.10 & 2.04\pm0.16 & 1.09\pm0.17 & 0.38\pm0.14 \\ \text{HD}113139 & 0.35\pm0.03 & 0.13\pm0.11 & -0.48\pm0.28 & 0.39\pm0.09 & 5.88\pm0.06 & -0.01\pm0.04 & 0.01\pm0.04 & 0.47\pm0.04 \\ \text{HD}101339 & 0.35\pm0.05 & 0.04\pm0.06 & -0.00\pm0.16 & 0.18\pm0.09 & 5.89\pm0.06 & -0.01\pm0.04 & 0.01\pm0.04 & 0.47\pm0.04 \\ \text{HD}102632 & 1.05\pm0.04 & 0.06\pm0.14 & -0.48\pm0.28 & 0.39\pm0.99 & 5.88\pm0.06 & -0.00\pm0.10 & -0.04 & 0.06\pm0.03 & -0.13\pm0.09 \\ \text{HD}026023 & 1.05\pm0.04 & 0.06\pm0.14 & -0.48\pm0.35 & 0.50\pm0.19 & 4.21\pm0.08 & 0.15\pm0.05 & 0.10\pm0.04 & 0.02\pm0.05 & -0.05 \\ \text{HD}215648 & 0.28\pm0.08 & 0.10\pm0.14 & -0.48\pm0.35 & 0.50\pm0.09 & 5.44\pm0.07 & 0.02\pm0.04 & 0.08\pm0.05 & 0.64\pm0.07 \\ \text{HD}1026600 & 0.61\pm0.04 & 0.22\pm0.16 & -0.43\pm0.40 & 0.79\pm0.09 & 4.37\pm0.07 & 0.02\pm0.04 & 0.08\pm0.05 & 0.61$									
$\begin{array}{c} \text{HD}028487 \\ \text{HD}214665 \\ 4.93 \pm 0.23 \\ 3.46 \pm 0.17 \\ 2.96 \pm 0.17 \\ 2.5.49 \pm 0.42 \\ 0.87 \pm 0.33 \\ 0.91 \pm 0.19 \\ 0.87 \pm 0.33 \\ -0.49 \pm 0.07 \\ 1.59 \pm 0.08 \\ 0.72 \pm 0.09 \\ 1.59 \pm 0.08 \\ 0.72 \pm 0.09 \\ 0.66 \pm 0.11 \\ 0.66 \pm 0.11 \\ 0.66 \pm 0.11 \\ 0.66 \pm 0.11 \\ 0.57 \pm 0.04 \\ \text{HD}004408 \\ 5.12 \pm 0.22 \\ 3.35 \pm 0.19 \\ 2.4.18 \pm 0.42 \\ 1.02 \pm 0.29 \\ 0.79 \pm 0.10 \\ 1.02 \pm 0.29 \\ 0.79 \pm 0.10 \\ 1.74 \pm 0.09 \\ 0.99 \pm 0.10 \\ 0.99 \pm 0.10 \\ 0.99 \pm 0.10 \\ 0.99 \pm 0.10 \\ 1.74 \pm 0.09 \\ 0.99 \pm 0.10 \\ 0.17 \pm 0.09 \\ 0.17 \pm 0.11 \\ 0.86 \pm 0.13 \\ 0.89 \pm 0.15 \\ 0.85 \pm 0.11 \\ 0.85 \pm 0.16 \\ 0.15 \pm 0.12 \\ 0.99 \pm 0.12 \\ 0.99 \pm 0.10 \\ 0.17 \pm 0.11 \\ 0.86 \pm 0.13 \\ 0.89 \pm 0.15 \\ 0.54 \pm 0.16 \\ 0.15 \pm 0.10 \\ 0.85 \pm 0.14 \\ 0.55 \pm 0.14 \\ 0.57 \pm 0.12 \\ 0.10 \pm 0.13 \\ 0.58 \pm 0.14 \\ 0.57 \pm 0.12 \\ 0.10 \pm 0.13 \\ 0.58 \pm 0.14 \\ 0.57 \pm 0.12 \\ 0.10 \pm 0.13 \\ 0.58 \pm 0.14 \\ 0.15 \pm 0.10 \\ 0.15$									
$\begin{array}{c} \text{HD} 214665 \\ \text{HD} 027598 \\ 3.88 \pm 0.18 \\ 2.43 \pm 0.23 \\ 3.46 \pm 0.19 \\ 2.43 \pm 0.36 \\ 0.80 \pm 0.18 \\ 2.43 \pm 0.36 \\ 0.80 \pm 0.18 \\ -0.49 \pm 0.06 \\ 1.56 \pm 0.10 \\ 0.06 \pm 0.11 \\ 1.56 \pm 0.10 \\ 0.66 \pm 0.11 \\ 0.66 \pm 0.11 \\ 0.66 \pm 0.11 \\ 0.57 \pm 0.04 \\ 0.68 \pm 0.09 \\ 0.68 \pm 0.09 \\ 0.10 \pm 0.26 \\ 0.99 \pm 0.10 \\ 0.83 \pm 0.12 \\ 0.99 \pm 0.10 \\ 0.99$					0.85 ± 0.17				
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$\begin{array}{c} \text{HD}027598 & 3.88 \pm 0.18 & 2.89 \pm 0.16 & 24.34 \pm 0.36 & 0.80 \pm 0.18 & -0.49 \pm 0.06 & 1.56 \pm 0.10 & 0.66 \pm 0.11 & 0.57 \pm 0.04 \\ \text{HD}004408 & 5.12 \pm 0.22 & 3.35 \pm 0.19 & 24.18 \pm 0.42 & 1.02 \pm 0.29 & 0.79 \pm 0.10 & 1.74 \pm 0.09 & 0.99 \pm 0.10 & 0.68 \pm 0.09 \\ \text{HD}024585 & 5.91 \pm 0.26 & 4.03 \pm 0.23 & 26.11 \pm 0.50 & 0.99 \pm 0.34 & -0.39 \pm 0.12 & 2.02 \pm 0.13 & 0.93 \pm 0.14 & 0.58 \pm 0.16 \\ \text{HD}014386 & 2.43 \pm 0.22 & 1.51 \pm 0.42 & 24.30 \pm 0.89 & 0.09 \pm 0.48 & -5.03 \pm 0.24 & 0.90 \pm 0.19 & -0.19 \pm 0.22 & 0.39 \pm 0.19 \\ \text{HD}175865 & 5.57 \pm 0.23 & 3.62 \pm 0.22 & 25.94 \pm 0.49 & 0.93 \pm 0.27 & 0.17 \pm 0.11 & 1.86 \pm 0.13 & 0.89 \pm 0.15 & 0.54 \pm 0.16 \\ \text{HD}014705 & 5.92 \pm 0.25 & 4.25 \pm 0.21 & 24.43 \pm 0.45 & 1.21 \pm 0.29 & -0.01 \pm 0.12 & 2.04 \pm 0.11 & 1.05 \pm 0.13 & 0.57 \pm 0.12 \\ \text{HD}018191 & 5.31 \pm 0.25 & 4.02 \pm 0.20 & 22.24 \pm 0.45 & 0.91 \pm 0.21 & -0.09 \pm 0.07 & 1.90 \pm 0.13 & 0.85 \pm 0.14 & 0.67 \pm 0.10 \\ \text{HD}108849 & 5.80 \pm 0.24 & 4.07 \pm 0.23 & 28.05 \pm 0.49 & 0.93 \pm 0.41 & -1.48 \pm 0.18 & 2.10 \pm 0.09 & 0.74 \pm 0.10 & 0.61 \pm 0.19 \\ \text{HD}108519 & 0.32 \pm 0.05 & 0.04 \pm 0.06 & -0.00 \pm 0.16 & 0.18 \pm 0.09 & 5.89 \pm 0.06 & -0.01 \pm 0.04 & 0.01 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}113139 & 0.35 \pm 0.03 & 0.13 \pm 0.11 & -0.48 \pm 0.28 & 0.39 \pm 0.09 & 5.86 \pm 0.06 & -0.01 \pm 0.04 & 0.01 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}016232 & 1.05 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.03 & -0.44 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.03 & -0.13 \pm 0.02 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.15 & -0.04 \pm 0.35 & 0.50 \pm 0.09 & 5.67 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.64 \pm 0.07 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.15 & -0.04 \pm 0.35 & 0.50 \pm 0.09 & 5.67 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.15 & -0.04 \pm 0.35 & 0.50 \pm 0.09 & 5.67 \pm 0.07 & 0.02 \pm 0.04 & 0.03 $	HD 214665		3.46 ± 0.19						
$\begin{array}{c} \text{HD}004408 & 5.12 \pm 0.22 & 3.35 \pm 0.19 & 24.18 \pm 0.42 & 1.02 \pm 0.29 & 0.79 \pm 0.10 & 1.74 \pm 0.09 & 0.99 \pm 0.10 & 0.68 \pm 0.09 \\ \text{HD}204585 & 5.91 \pm 0.26 & 4.03 \pm 0.23 & 26.11 \pm 0.50 & 0.99 \pm 0.34 & -0.39 \pm 0.12 & 2.02 \pm 0.13 & 0.93 \pm 0.14 & 0.58 \pm 0.16 \\ \text{HD}014386 & 2.43 \pm 0.22 & 1.51 \pm 0.42 & 24.30 \pm 0.89 & 0.09 \pm 0.48 & -5.03 \pm 0.24 & 0.90 \pm 0.19 & -0.19 \pm 0.22 & 0.39 \pm 0.19 \\ \text{HD}175865 & 5.57 \pm 0.23 & 3.62 \pm 0.22 & 25.94 \pm 0.49 & 0.93 \pm 0.27 & 0.17 \pm 0.11 & 1.86 \pm 0.13 & 0.89 \pm 0.15 & 0.54 \pm 0.16 \\ \text{HD}094705 & 5.92 \pm 0.25 & 4.25 \pm 0.21 & 24.43 \pm 0.45 & 1.21 \pm 0.29 & -0.01 \pm 0.12 & 2.04 \pm 0.11 & 1.05 \pm 0.13 & 0.57 \pm 0.12 \\ \text{HD}018191 & 5.31 \pm 0.25 & 4.02 \pm 0.20 & 22.24 \pm 0.45 & 0.91 \pm 0.21 & -0.09 \pm 0.07 & 1.90 \pm 0.13 & 0.85 \pm 0.14 & 0.67 \pm 0.10 \\ \text{HD}196610 & 5.03 \pm 0.17 & 3.95 \pm 0.18 & 28.36 \pm 0.40 & 0.91 \pm 0.26 & -0.15 \pm 0.10 & 2.04 \pm 0.16 & 1.09 \pm 0.17 & 0.38 \pm 0.14 \\ \text{HD}108849 & 5.80 \pm 0.24 & 4.07 \pm 0.23 & 28.05 \pm 0.49 & 0.93 \pm 0.41 & -1.48 \pm 0.18 & 2.10 \pm 0.09 & 0.74 \pm 0.10 & 0.61 \pm 0.19 \\ \text{HD}207076 & 4.91 \pm 0.17 & 3.44 \pm 0.16 & 24.34 \pm 0.36 & 0.99 \pm 0.26 & -1.09 \pm 0.11 & 1.87 \pm 0.08 & 0.90 \pm 0.08 & 0.50 \pm 0.11 \\ \text{Dwarfs} \\ \text{HD}113139 & 0.32 \pm 0.05 & 0.04 \pm 0.06 & -0.00 \pm 0.16 & 0.18 \pm 0.09 & 5.88 \pm 0.06 & -0.01 \pm 0.04 & 0.04 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}213135 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & -0.48 \pm 0.28 & 0.39 \pm 0.09 & 5.88 \pm 0.06 & -0.01 \pm 0.04 & 0.01 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}016232 & 1.05 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.03 & -0.13 \pm 0.02 \\ \text{HD}016232 & 1.05 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 4.21 \pm 0.08 & 0.15 \pm 0.05 & 0.14 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.64 \pm 0.07 \\ \text{HD}027634 & 0.28 \pm 0.08 & 0.10 \pm 0.15 & -1.15 \pm 0.37 & 0.33 \pm 0.07 & 4.52 \pm 0.08 & 0.02 \pm 0.06 & 0.10 \pm 0.07 & 0.34 \pm 0.04 \\ \text{HD}026600 & 0.61 \pm 0.04 & 0.22 \pm 0.16 & -0.43 \pm 0.40 & 0.79 \pm 0.09 & 4.37 \pm 0.07 & 0.02 \pm 0.$									
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 004408	5.12 ± 0.22	3.35 ± 0.19	24.18 ± 0.42	1.02 ± 0.29	0.79 ± 0.10	1.74 ± 0.09	0.99 ± 0.10	0.68 ± 0.09
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 204585								
$\begin{array}{c} \text{HD 175865} & 5.57 \pm 0.23 \\ \text{HD 094705} & 5.92 \pm 0.25 \\ \text{4.25} \pm 0.21 \\ \text{24.43} \pm 0.45 \\ \text{1.21} \pm 0.29 \\ \text{-0.01} \pm 0.12 \\ \text{0.93} \pm 0.77 \\ \text{-0.01} \pm 0.12 \\ \text{0.01} \pm 0.12 \\ \text{-0.01} \pm 0.13 \\ \text{0.89} \pm 0.15 \\ \text{0.15} \\ \text{0.54} \pm 0.16 \\ \text{0.57} \pm 0.12 \\ \text{HD 01819} \\ \text{1.31} \pm 0.25 \\ \text{4.02} \pm 0.25 \\ \text{4.02} \pm 0.20 \\ \text{22.24} \pm 0.45 \\ \text{0.91} \pm 0.21 \\ \text{-0.91} \pm 0.21 \\ \text{-0.09} \pm 0.07 \\ \text{-0.15} \pm 0.10 \\ \text{-0.15} \pm 0.13 \\ \text{0.85} \pm 0.14 \\ \text{0.67} \pm 0.10 \\ \text{0.88} \pm 0.14 \\ \text{0.18} \pm 0.13 \\ \text{0.85} \pm 0.14 \\ \text{0.67} \pm 0.10 \\ \text{0.85} \pm 0.14 \\ \text{0.67} \pm 0.10 \\ \text{0.88} \pm 0.14 \\ \text{0.93} \pm 0.14 \\ \text{0.99} \pm 0.26 \\ \text{-1.09} \pm 0.11 \\ \text{1.87} \pm 0.08 \\ \text{0.90} \pm 0.08 \\ \text{0.90} \pm 0.08 \\ \text{0.50} \pm 0.11 \\ \text{0.61} \pm 0.19 \\ \text{0.99} \pm 0.26 \\ \text{0.11} \\ \text{0.99} \pm 0.26 \\ \text{0.11} \\ \text{0.99} \pm 0.11 \\ \text{0.87} \pm 0.08 \\ \text{0.99} \pm 0.08 \\ 0.99$									
$\begin{array}{c} \text{HD}094705 & 5.92 \pm 0.25 & 4.25 \pm 0.21 & 24.43 \pm 0.45 & 1.21 \pm 0.29 & -0.01 \pm 0.12 & 2.04 \pm 0.11 & 1.05 \pm 0.13 & 0.57 \pm 0.12 \\ \text{HD}018191 & 5.31 \pm 0.25 & 4.02 \pm 0.20 & 22.24 \pm 0.45 & 0.91 \pm 0.21 & -0.09 \pm 0.07 & 1.90 \pm 0.13 & 0.85 \pm 0.14 & 0.67 \pm 0.10 \\ \text{HD}196610 & 5.03 \pm 0.17 & 3.95 \pm 0.18 & 28.36 \pm 0.40 & 0.91 \pm 0.26 & -0.15 \pm 0.10 & 2.04 \pm 0.16 & 1.09 \pm 0.17 & 0.38 \pm 0.14 \\ \text{HD}108849 & 5.80 \pm 0.24 & 4.07 \pm 0.23 & 28.05 \pm 0.49 & 0.93 \pm 0.41 & -1.48 \pm 0.18 & 2.10 \pm 0.09 & 0.74 \pm 0.10 & 0.61 \pm 0.19 \\ \text{HD}207076 & 4.91 \pm 0.17 & 3.44 \pm 0.16 & 24.34 \pm 0.36 & 0.99 \pm 0.26 & -1.09 \pm 0.11 & 1.87 \pm 0.08 & 0.90 \pm 0.08 & 0.50 \pm 0.11 \\ \text{Dwarfs} \\ \text{HD}108519 & 0.32 \pm 0.05 & 0.04 \pm 0.06 & -0.00 \pm 0.16 & 0.18 \pm 0.09 & 6.20 \pm 0.03 & 0.04 \pm 0.03 & 0.10 \pm 0.04 & 0.26 \pm 0.02 \\ \text{HD}213135 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.04 & 0.01 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}113139 & 0.35 \pm 0.03 & 0.13 \pm 0.11 & -0.48 \pm 0.28 & 0.39 \pm 0.09 & 5.88 \pm 0.06 & 0.00 \pm 0.02 & 0.06 \pm 0.03 & 0.47 \pm 0.04 \\ \text{HD}026015 & 0.65 \pm 0.05 & 0.01 \pm 0.04 & -0.41 \pm 0.10 & 0.55 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.03 & -0.13 \pm 0.02 \\ \text{HD}016232 & 1.05 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 4.21 \pm 0.08 & 0.15 \pm 0.05 & 0.17 \pm 0.05 & 0.82 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.64 \pm 0.07 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}0275648 & 0.28 \pm 0.08 & 0.10 \pm 0.15 & -1.15 \pm 0.37 & 0.33 \pm 0.07 & 4.52 \pm 0.08 & 0.02 \pm 0.04 &$									
$\begin{array}{c} \text{HD}094705 & 5.92 \pm 0.25 & 4.25 \pm 0.21 & 24.43 \pm 0.45 & 1.21 \pm 0.29 & -0.01 \pm 0.12 & 2.04 \pm 0.11 & 1.05 \pm 0.13 & 0.57 \pm 0.12 \\ \text{HD}018191 & 5.31 \pm 0.25 & 4.02 \pm 0.20 & 22.24 \pm 0.45 & 0.91 \pm 0.21 & -0.09 \pm 0.07 & 1.90 \pm 0.13 & 0.85 \pm 0.14 & 0.67 \pm 0.10 \\ \text{HD}196610 & 5.03 \pm 0.17 & 3.95 \pm 0.18 & 28.36 \pm 0.40 & 0.91 \pm 0.26 & -0.15 \pm 0.10 & 2.04 \pm 0.16 & 1.09 \pm 0.17 & 0.38 \pm 0.14 \\ \text{HD}108849 & 5.80 \pm 0.24 & 4.07 \pm 0.23 & 28.05 \pm 0.49 & 0.93 \pm 0.41 & -1.48 \pm 0.18 & 2.10 \pm 0.09 & 0.74 \pm 0.10 & 0.61 \pm 0.19 \\ \text{HD}207076 & 4.91 \pm 0.17 & 3.44 \pm 0.16 & 24.34 \pm 0.36 & 0.99 \pm 0.26 & -1.09 \pm 0.11 & 1.87 \pm 0.08 & 0.90 \pm 0.08 & 0.50 \pm 0.11 \\ \text{Dwarfs} \\ \text{HD}108519 & 0.32 \pm 0.05 & 0.04 \pm 0.06 & -0.00 \pm 0.16 & 0.18 \pm 0.09 & 6.20 \pm 0.03 & 0.04 \pm 0.03 & 0.10 \pm 0.04 & 0.26 \pm 0.02 \\ \text{HD}213135 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.04 & 0.01 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}113139 & 0.35 \pm 0.03 & 0.13 \pm 0.11 & -0.48 \pm 0.28 & 0.39 \pm 0.09 & 5.88 \pm 0.06 & 0.00 \pm 0.02 & 0.06 \pm 0.03 & 0.47 \pm 0.04 \\ \text{HD}026015 & 0.65 \pm 0.05 & 0.01 \pm 0.04 & -0.41 \pm 0.10 & 0.55 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.03 & -0.13 \pm 0.02 \\ \text{HD}016232 & 1.05 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 4.21 \pm 0.08 & 0.15 \pm 0.05 & 0.17 \pm 0.05 & 0.82 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.64 \pm 0.07 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}0275648 & 0.28 \pm 0.08 & 0.10 \pm 0.15 & -1.15 \pm 0.37 & 0.33 \pm 0.07 & 4.52 \pm 0.08 & 0.02 \pm 0.04 &$	HD 175865	5.57 ± 0.23	3.62 ± 0.22	25.94 ± 0.49	0.93 ± 0.27	0.17 ± 0.11	1.86 ± 0.13	0.89 ± 0.15	0.54 ± 0.16
$\begin{array}{c} \text{HD}018191 & 5.31 \pm 0.25 & 4.02 \pm 0.20 & 22.24 \pm 0.45 & 0.91 \pm 0.21 & -0.09 \pm 0.07 & 1.90 \pm 0.13 & 0.85 \pm 0.14 & 0.67 \pm 0.10 \\ \text{HD}196610 & 5.03 \pm 0.17 & 3.95 \pm 0.18 & 28.36 \pm 0.40 & 0.91 \pm 0.26 & -0.15 \pm 0.10 & 2.04 \pm 0.16 & 1.09 \pm 0.17 & 0.38 \pm 0.14 \\ \text{HD}108849 & 5.80 \pm 0.24 & 4.07 \pm 0.23 & 28.05 \pm 0.49 & 0.93 \pm 0.41 & -1.48 \pm 0.18 & 2.10 \pm 0.09 & 0.74 \pm 0.10 & 0.61 \pm 0.19 \\ \text{HD}207076 & 4.91 \pm 0.17 & 3.44 \pm 0.16 & 24.34 \pm 0.36 & 0.99 \pm 0.26 & -1.09 \pm 0.11 & 1.87 \pm 0.08 & 0.90 \pm 0.08 & 0.50 \pm 0.11 \\ \hline \\ \text{Dwarfs} \\ \text{HD}108519 & 0.32 \pm 0.05 & 0.04 \pm 0.06 & -0.00 \pm 0.16 & 0.18 \pm 0.09 & 6.20 \pm 0.03 & 0.04 \pm 0.03 & 0.10 \pm 0.04 & 0.26 \pm 0.02 \\ \text{HD}213135 & 0.58 \pm 0.05 & 0.33 \pm 0.11 & 0.37 \pm 0.27 & 0.28 \pm 0.07 & 5.59 \pm 0.06 & -0.01 \pm 0.04 & 0.01 \pm 0.04 & 0.47 \pm 0.04 \\ \text{HD}113139 & 0.35 \pm 0.03 & 0.13 \pm 0.11 & -0.48 \pm 0.28 & 0.39 \pm 0.09 & 5.88 \pm 0.06 & 0.00 \pm 0.02 & 0.06 \pm 0.03 & 0.47 \pm 0.04 \\ \text{HD}026015 & 0.65 \pm 0.05 & 0.01 \pm 0.04 & -0.41 \pm 0.10 & 0.55 \pm 0.09 & 5.67 \pm 0.07 & 0.01 \pm 0.03 & -0.02 \pm 0.03 & -0.13 \pm 0.02 \\ \text{HD}016232 & 1.05 \pm 0.04 & 0.65 \pm 0.18 & -0.36 \pm 0.46 & 0.72 \pm 0.09 & 4.21 \pm 0.08 & 0.15 \pm 0.05 & 0.17 \pm 0.05 & 0.82 \pm 0.05 \\ \text{HD}087822 & 0.54 \pm 0.04 & 0.12 \pm 0.15 & -0.44 \pm 0.38 & 0.55 \pm 0.10 & 5.12 \pm 0.07 & 0.02 \pm 0.04 & 0.08 \pm 0.05 & 0.64 \pm 0.07 \\ \text{HD}027524 & 0.62 \pm 0.04 & 0.08 \pm 0.14 & -0.48 \pm 0.35 & 0.50 \pm 0.09 & 5.44 \pm 0.07 & 0.02 \pm 0.04 & 0.03 \pm 0.05 & 0.62 \pm 0.05 \\ \text{HD}216660 & 0.61 \pm 0.04 & 0.22 \pm 0.16 & -0.43 \pm 0.40 & 0.79 \pm 0.09 & 5.44 \pm 0.07 & 0.08 \pm 0.05 & 0.16 \pm 0.05 & 0.61 \pm 0.06 \\ \text{HD}219623 & 0.77 \pm 0.05 & 0.43 \pm 0.20 & -0.54 \pm 0.49 & 0.83 \pm 0.08 & 4.17 \pm 0.10 & 0.14 \pm 0.05 & 0.24 \pm 0.05 & 0.85 \pm 0.02 \\ \text{HD}102870 & 0.82 \pm 0.04 & 0.34 \pm 0.20 & -0.63 \pm 0.50 & 0.91 \pm 0.10 & 4.09 \pm 0.09 & 0.18 \pm 0.05 & 0.24 \pm 0.05 & 0.86 \pm 0.03 \\ \text{HD}102870 & 0.82 \pm 0.04 & 0.34 \pm 0.20 & -0.63 \pm 0.50 & 0.91 \pm 0.10 & 4.09 \pm 0.09 & 0.18 \pm 0.05 & 0.24 \pm 0.05 & 0.86 \pm 0.03 \\ \text{HD}102870 & 0.82 \pm 0.04 & 0.34 \pm 0.20 & -0.63 \pm 0.50 & 0.91 \pm 0.10 & 4.09 \pm 0.09 & 0.18 \pm 0.05$	HD 094705	5.92 ± 0.25	4.25 ± 0.21	24.43 ± 0.45	1.21 ± 0.29	-0.01 ± 0.12	2.04 ± 0.11	1.05 ± 0.13	0.57 ± 0.12
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$									
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$									
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 196610	5.03 ± 0.17	3.95 ± 0.18	28.36 ± 0.40	0.91 ± 0.26	-0.15 ± 0.10	2.04 ± 0.16	1.09 ± 0.17	0.38 ± 0.14
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	HD 108849	5.80 ± 0.24	4.07 ± 0.23	28.05 ± 0.49	0.93 ± 0.41	-1.48 ± 0.18	2.10 ± 0.09	0.74 ± 0.10	0.61 ± 0.19
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$									
$\begin{array}{llllllllllllllllllllllllllllllllllll$	11D 207070	4.91 ± 0.17	3.44 ± 0.10	24.34 ± 0.30	0.99 ± 0.20	-1.09 ± 0.11	1.67 ± 0.06	0.90 ± 0.08	0.50 ± 0.11
$\begin{array}{llllllllllllllllllllllllllllllllllll$									
$\begin{array}{llllllllllllllllllllllllllllllllllll$					Dwarfs				
$\begin{array}{llllllllllllllllllllllllllllllllllll$	HD 108519	0.32 ± 0.05	0.04 ± 0.06	-0.00 + 0.16	0.18 ± 0.09	6.20 ± 0.03	0.04 ± 0.03	0.10 ± 0.04	0.26 ± 0.02
$\begin{array}{llllllllllllllllllllllllllllllllllll$									
$\begin{array}{cccccccccccccccccccccccccccccccccccc$									
$\begin{array}{llllllllllllllllllllllllllllllllllll$	HD 113139	0.35 ± 0.03	0.13 ± 0.11	-0.48 ± 0.28	0.39 ± 0.09	5.88 ± 0.06	0.00 ± 0.02	0.06 ± 0.03	0.47 ± 0.04
$\begin{array}{llllllllllllllllllllllllllllllllllll$	HD 026015	0.65 ± 0.05	0.01 ± 0.04	-0.41 ± 0.10	0.55 ± 0.09	5.67 ± 0.07	0.01 ± 0.03	-0.02 ± 0.03	-0.13 ± 0.02
$\begin{array}{llllllllllllllllllllllllllllllllllll$									
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$\begin{array}{llllllllllllllllllllllllllllllllllll$		0.54 ± 0.04	0.12 ± 0.15	-0.44 ± 0.38	0.55 ± 0.10	5.12 ± 0.07	0.02 ± 0.04	0.08 ± 0.05	0.64 ± 0.07
$\begin{array}{llllllllllllllllllllllllllllllllllll$	HD 027524	0.62 ± 0.04	0.08 ± 0.14	-0.48 ± 0.35	0.50 ± 0.09	5.44 ± 0.07	0.02 ± 0.04	0.03 ± 0.05	0.62 ± 0.05
$\begin{array}{llllllllllllllllllllllllllllllllllll$									
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$ \begin{array}{llllllllllllllllllllllllllllllllllll$									
$ \begin{array}{llllllllllllllllllllllllllllllllllll$	HD 219623	0.77 ± 0.05	0.43 ± 0.20	-0.54 ± 0.49	0.83 ± 0.08	4.17 ± 0.10	0.14 ± 0.05	0.21 ± 0.05	0.87 ± 0.05
$ \text{HD } 102870 0.82 \pm 0.04 0.34 \pm 0.20 -0.63 \pm 0.50 0.91 \pm 0.10 4.09 \pm 0.09 \qquad 0.18 \pm 0.05 0.24 \pm 0.05 0.86 \pm 0.03 $	HD 165908	0.51 ± 0.06	0.25 ± 0.10	-0.32 ± 0.26	0.34 ± 0.07	4.06 ± 0.06	0.07 ± 0.03	0.03 ± 0.03	0.35 ± 0.02
$ \text{HD} 027383 1.01 \pm 0.04 0.44 \pm 0.21 -0.49 \pm 0.54 0.96 \pm 0.07 3.77 \pm 0.10 \qquad 0.18 \pm 0.06 0.24 \pm 0.06 0.92 \pm 0.06 $									
	HD 027383	1.01 ± 0.04	0.44 ± 0.21	-0.49 ± 0.54	0.96 ± 0.07	3.77 ± 0.10	0.18 ± 0.06	0.24 ± 0.06	0.92 ± 0.06

Table 7. continued.

Star IDs	Na _d	Ca _d	¹² CO	Mg1	Brγ	FeA	FeB	Mg2
HD 114710	0.87 ± 0.06	0.63 ± 0.19	-0.51 ± 0.48	0.84 ± 0.10	3.96 ± 0.11	0.13 ± 0.05	0.24 ± 0.06	0.85 ± 0.05
HD 176051	0.80 ± 0.06	0.51 ± 0.19	-0.07 ± 0.48	0.86 ± 0.08	4.14 ± 0.10	0.14 ± 0.06	0.13 ± 0.07	0.83 ± 0.05
HD 109358	0.81 ± 0.03	0.50 ± 0.17	-0.16 ± 0.44	0.77 ± 0.07	3.37 ± 0.09	0.11 ± 0.04	0.16 ± 0.05	0.72 ± 0.03
HD 095128	1.09 ± 0.06	0.51 ± 0.21	0.01 ± 0.52	1.01 ± 0.07	3.38 ± 0.09	0.15 ± 0.04	0.25 ± 0.05	0.88 ± 0.05
HD 020619	0.88 ± 0.06	0.58 ± 0.20	0.03 ± 0.51	0.82 ± 0.05	2.84 ± 0.10	0.19 ± 0.04	0.12 ± 0.04	0.92 ± 0.06
HD 010307	0.99 ± 0.06	0.54 ± 0.20	0.19 ± 0.51	0.97 ± 0.07	3.06 ± 0.10	0.18 ± 0.04	0.28 ± 0.05	0.94 ± 0.05
HD 076151	1.31 ± 0.05	0.66 ± 0.25	-0.10 ± 0.63	1.19 ± 0.08	2.69 ± 0.11	0.29 ± 0.06	0.37 ± 0.07	1.14 ± 0.06
HD 214850	0.82 ± 0.05	0.64 ± 0.20	2.62 ± 0.50	0.91 ± 0.06	2.44 ± 0.08	0.21 ± 0.05	0.32 ± 0.06	0.80 ± 0.04
HD 165185	0.82 ± 0.05	0.61 ± 0.19	-0.58 ± 0.47	0.66 ± 0.07	3.16 ± 0.11	0.12 ± 0.06	0.20 ± 0.06	0.77 ± 0.04
HD 115617	1.22 ± 0.06	0.93 ± 0.24	0.81 ± 0.62	1.16 ± 0.06	2.88 ± 0.11	0.30 ± 0.05	0.32 ± 0.06	1.12 ± 0.05
HD 101501	1.26 ± 0.06	0.90 ± 0.24	0.33 ± 0.60	1.08 ± 0.06	2.25 ± 0.12	0.27 ± 0.08	0.28 ± 0.09	1.09 ± 0.04
HD 075732	2.25 ± 0.08	1.11 ± 0.38	2.62 ± 0.95	2.06 ± 0.11	2.24 ± 0.14	0.50 ± 0.08	0.52 ± 0.09	1.73 ± 0.07
HD 145675	2.78 ± 0.11	1.32 ± 0.41	2.58 ± 1.05	2.24 ± 0.09	1.70 ± 0.14	0.47 ± 0.09	0.56 ± 0.10	2.07 ± 0.09
HD 010476	1.56 ± 0.04	1.14 ± 0.28	1.87 ± 0.70	1.36 ± 0.07	1.61 ± 0.10	0.40 ± 0.08	0.39 ± 0.09	1.26 ± 0.07
HD 003765	2.13 ± 0.08	1.59 ± 0.34	3.58 ± 0.85	1.91 ± 0.07	0.98 ± 0.12	0.56 ± 0.08	0.51 ± 0.09	1.67 ± 0.07
HD 219134	2.98 ± 0.07	2.08 ± 0.39	3.95 ± 0.97	1.98 ± 0.05	1.48 ± 0.14	0.69 ± 0.09	0.56 ± 0.10	1.94 ± 0.08
HD 045977	3.79 ± 0.08	3.20 ± 0.41	4.22 ± 1.03	2.37 ± 0.05	0.27 ± 0.12	0.88 ± 0.10	0.74 ± 0.11	1.93 ± 0.04
HD 207991	2.73 ± 0.10	1.75 ± 0.15	19.36 ± 0.34	0.86 ± 0.12	0.45 ± 0.05	1.03 ± 0.07	0.51 ± 0.07	0.60 ± 0.03
HD 036003	3.26 ± 0.06	2.74 ± 0.37	4.80 ± 0.93	2.28 ± 0.03	0.36 ± 0.10	0.78 ± 0.09	0.52 ± 0.10	1.75 ± 0.07
HD 201092	3.80 ± 0.10	3.36 ± 0.26	5.13 ± 0.66	1.77 ± 0.11	0.60 ± 0.06	0.94 ± 0.08	0.42 ± 0.09	1.17 ± 0.08
HD 237903	3.95 ± 0.06	3.88 ± 0.27	4.80 ± 0.68	1.74 ± 0.05	0.51 ± 0.07	0.86 ± 0.07	0.48 ± 0.07	1.44 ± 0.04
HD 019305	4.63 ± 0.08	3.87 ± 0.31	5.47 ± 0.77	2.01 ± 0.05	0.35 ± 0.07	0.91 ± 0.08	0.49 ± 0.09	1.61 ± 0.05
HD 209290	5.46 ± 0.05	4.31 ± 0.20	6.57 ± 0.49	1.55 ± 0.08	0.05 ± 0.04	0.89 ± 0.04	0.39 ± 0.05	1.06 ± 0.03
HD 036395	8.39 ± 0.08	5.70 ± 0.31	7.68 ± 0.76	2.27 ± 0.10	0.06 ± 0.05	1.07 ± 0.07	0.47 ± 0.08	1.76 ± 0.05
HD 042581	6.05 ± 0.05	4.71 ± 0.19	6.91 ± 0.46	1.50 ± 0.09	0.34 ± 0.04	0.89 ± 0.05	0.35 ± 0.05	0.99 ± 0.03
G1806	4.41 ± 0.11	4.05 ± 0.10	6.58 ± 0.25	0.70 ± 0.18	0.29 ± 0.09	0.58 ± 0.06	-0.00 ± 0.07	0.43 ± 0.04
HD 095735	3.28 ± 0.09	2.85 ± 0.09	6.34 ± 0.21	0.55 ± 0.22	-0.47 ± 0.12	0.40 ± 0.05	-0.08 ± 0.06	0.39 ± 0.02
Gl381	4.96 ± 0.08	3.74 ± 0.09	7.17 ± 0.21	0.50 ± 0.22	-0.07 ± 0.11	0.43 ± 0.05	-0.03 ± 0.05	0.42 ± 0.05
Gl581	5.19 ± 0.10	3.86 ± 0.09	7.87 ± 0.21	0.38 ± 0.27	-0.63 ± 0.12	0.37 ± 0.05	-0.16 ± 0.06	0.41 ± 0.04
Gl213	4.09 ± 0.12	2.53 ± 0.11	8.14 ± 0.27	0.20 ± 0.34	-0.60 ± 0.15	0.15 ± 0.08	-0.30 ± 0.08	0.27 ± 0.05
G1299	3.06 ± 0.12	1.61 ± 0.14	7.81 ± 0.32	0.10 ± 0.36	-0.72 ± 0.15	0.08 ± 0.09	-0.34 ± 0.10	0.29 ± 0.08
Gl268AB	7.31 ± 0.10	3.38 ± 0.11	9.99 ± 0.26	0.21 ± 0.30	-0.34 ± 0.14	0.23 ± 0.07	-0.20 ± 0.07	0.41 ± 0.04
Gl51	7.40 ± 0.14	3.74 ± 0.15	10.60 ± 0.34	0.31 ± 0.40	-0.90 ± 0.17	0.25 ± 0.10	-0.34 ± 0.11	0.37 ± 0.08
G1866	7.05 ± 0.14	1.88 ± 0.16	13.49 ± 0.36	0.19 ± 0.40	-1.74 ± 0.18	0.03 ± 0.11	-0.40 ± 0.12	0.40 ± 0.07
GJ1111	5.42 ± 0.21	1.17 ± 0.20	13.67 ± 0.46	0.19 ± 0.45	-1.42 ± 0.19	0.01 ± 0.14	-0.38 ± 0.16	0.26 ± 0.09
G1644C	4.69 ± 0.24	0.58 ± 0.25	14.95 ± 0.56	0.28 ± 0.54	-2.24 ± 0.21	-0.04 ± 0.16	-0.42 ± 0.18	0.59 ± 0.12
G1752B	5.47 ± 0.26	0.90 ± 0.27	17.37 ± 0.60	0.27 ± 0.54	-2.30 ± 0.24	-0.10 ± 0.18	-0.49 ± 0.19	0.63 ± 0.12
LP412-31	7.25 ± 0.24	1.04 ± 0.23	18.69 ± 0.50	0.56 ± 0.41	-1.97 ± 0.20	0.15 ± 0.16	-0.37 ± 0.17	0.54 ± 0.06
LHS2065	6.30 ± 0.24	1.02 ± 0.27	19.73 ± 0.58	0.22 ± 0.46	-2.92 ± 0.19	-0.06 ± 0.17	-0.31 ± 0.19	0.70 ± 0.04
LHS2924	4.95 ± 0.31	1.08 ± 0.37	20.33 ± 0.82	0.36 ± 0.55	-3.23 ± 0.26	-0.15 ± 0.25	-0.34 ± 0.27	0.96 ± 0.09
LP944-20	2.39 ± 0.32	0.61 ± 0.29	17.74 ± 0.64	0.44 ± 0.47	-3.21 ± 0.23	-0.21 ± 0.20	-0.66 ± 0.22	0.40 ± 0.10
BRIB0021	2.85 ± 0.35	1.20 ± 0.41	20.02 ± 0.90	0.51 ± 0.58	-4.00 ± 0.23	0.38 ± 0.23	-0.10 ± 0.26	0.97 ± 0.15