

A large, bold, black number '42' is centered on a dark blue background. The number has a soft, glowing blue aura around it, giving it a three-dimensional appearance.

An Introduction to Simulation Using 42

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What is 42?

- A simulation of spacecraft attitude and orbital dynamics and control
- Intended for use from concept studies through ops
 - Rapid prototyping makes it useful for MDL studies
 - Environment models support actuator sizing, performance studies
 - High-fidelity dynamics handle multi-body, flexible-body spacecraft
 - Portability (Mac, linux, Windows) minimizes infrastructure requirements
 - Clean interface aids progression from flight software “model” to dropping in actual flight software
 - Visualization aids situational awareness from concept to operations
- Designed to be powerful, but easy to get started

Features

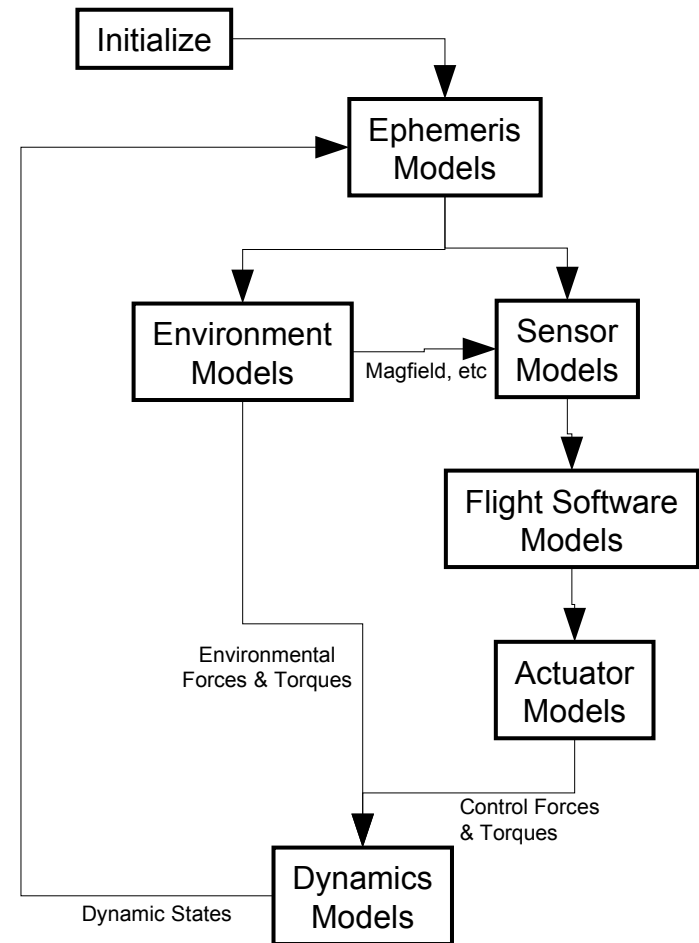
- Multiple spacecraft, anywhere in the solar system
 - Two-body, three-body orbit dynamics (with seamless transition between)
 - One sun, nine planets, 45 major moons
 - Minor bodies (comets and asteroids) added as needed
 - RQ36, Eros, Itokawa, Wirtanen, etc
- Multi-body spacecraft
 - Tree topology
 - Kane's dynamics formulation
 - Each body may be rigid or flexible
 - Flexible parameters taken (by m-file script) from Nastran output (.f06 file)
 - Joints may have any combination of rotational and translational degrees of freedom

More Features

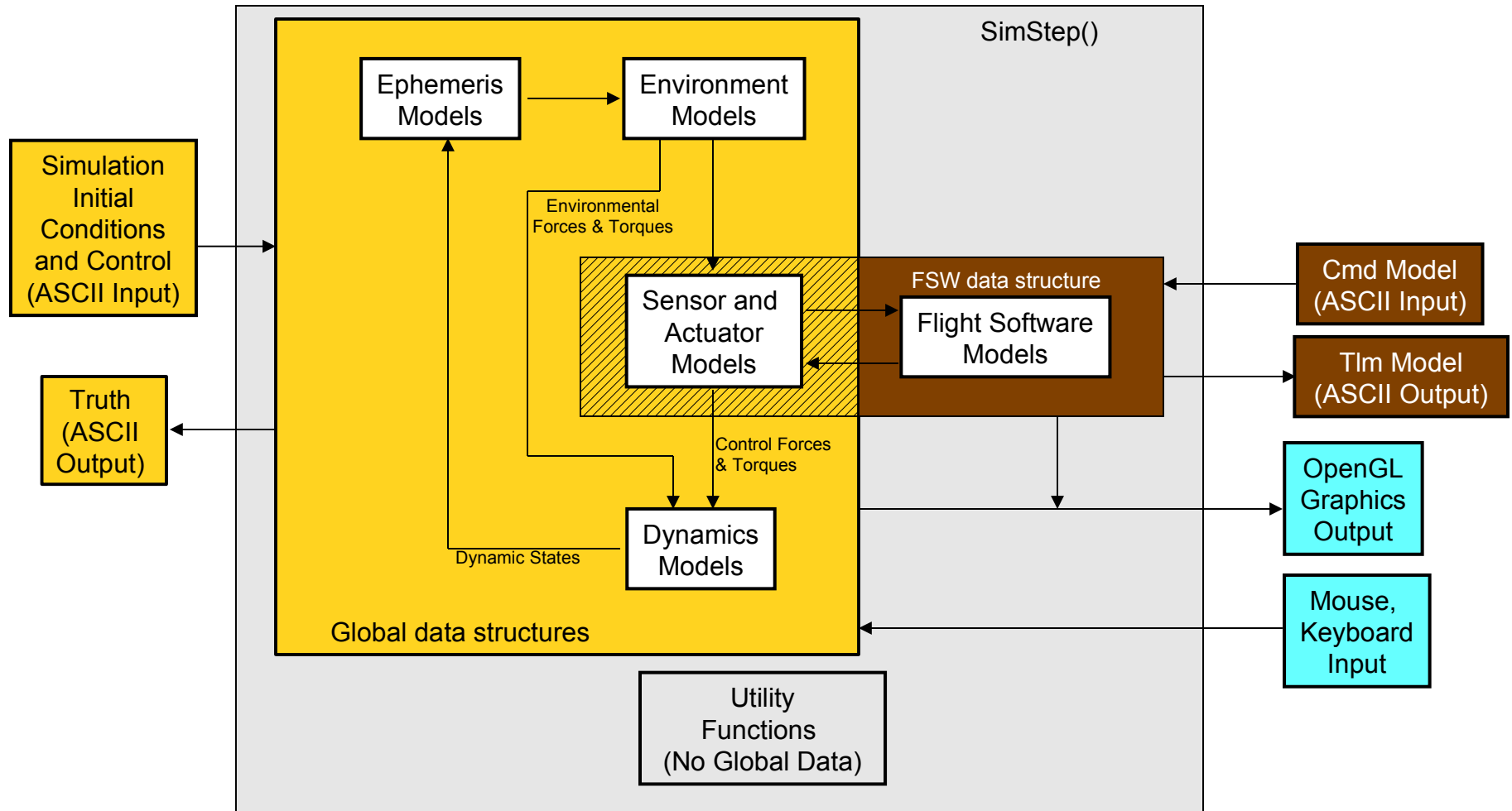
- Supports precision formation flying
 - Several S/C may be tied to a common reference orbit
 - Encke's method or Euler-Hill equations used to propagate relative orbit states
 - Precision maintained by judicious partitioning of dynamics
 - Add big things to big things, small things to small things
- Clean FSW interface facilitates FSW validation
 - Used by GLAST project for independent validation of vendor's (autocoded) GNC flight software
- Open Source, available on sourceforge and github
 - Sourceforge.net/projects/fortytwospacecraftsimulation
 - [Github/ericstoneking/42](https://github.com/ericstoneking/42)

A Basic Simulation Loop

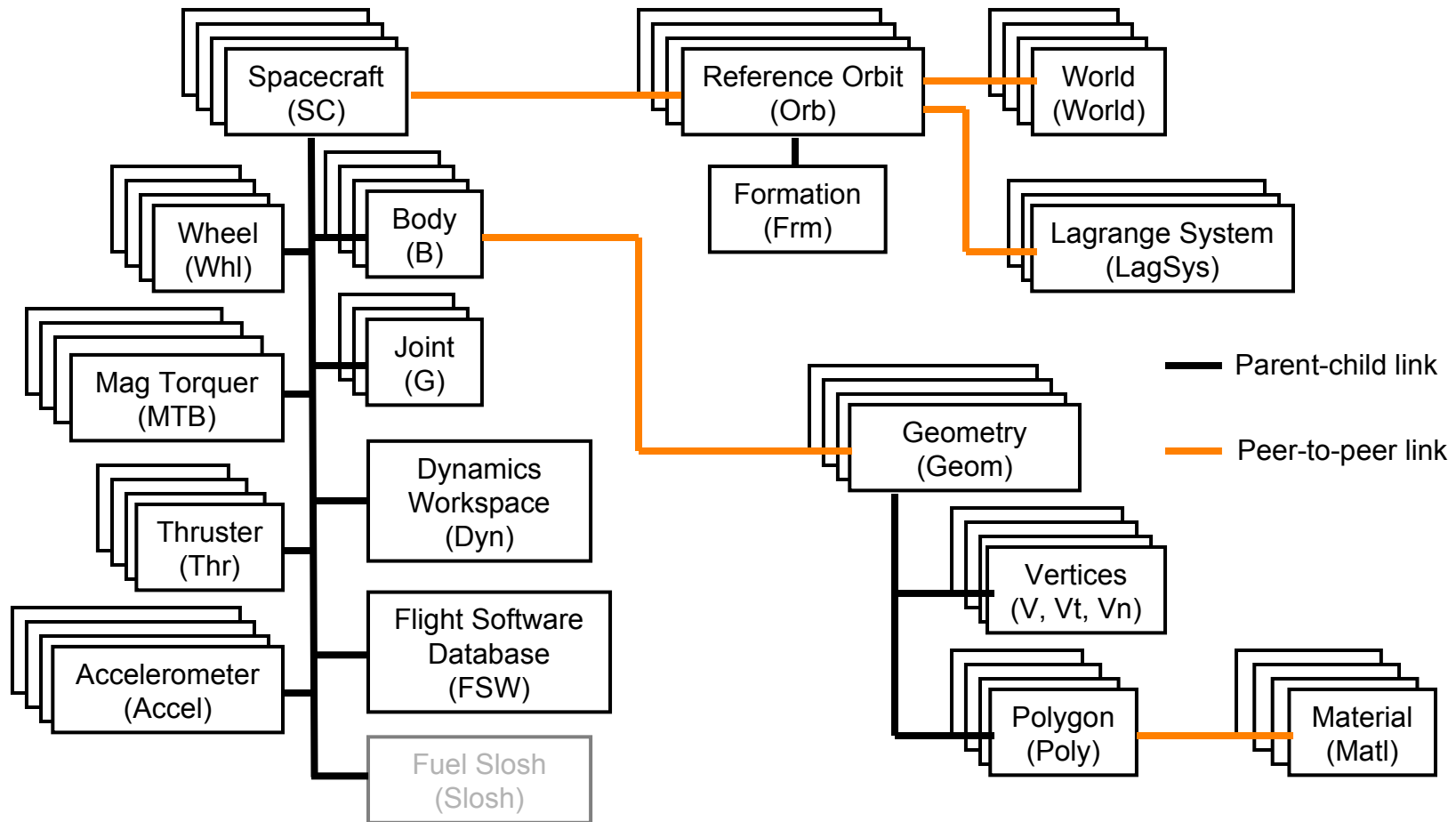
- Initialize
 - Read user inputs
 - Set up
- Ephemeris: Where is everything?
 - Sun, Earth, Moon, etc
 - Orbits
 - Spacecraft
- Environment Models: What forces and torques exerted by the environment?
- Sensor Models
 - Input truth
 - Output measurements
- Flight Software Models
 - Input Measurements
 - Process Control Laws, etc
 - Output Actuator Commands
- Actuator Models
 - Input Commands
 - Output Forces and Torques
- Dynamics: How does S/C respond to forces and torques?
 - Integrate dynamic equations of motion over a timestep
 - Advance time to next step



42's Architecture



Global Data Structure Relationships



42 Walkthrough

Models, Terminology, Input Files

Environment Models

- Planetary Ephemerides use Mean Elements or DE430
 - Mean elements from Meeus, “Astronomical Algorithms”
 - Good enough for GNC validation, not intended for mission planning
 - Use GMAT or ODTBX for that
 - DE430 from JPL
- Gravity Models have coefficients up to 18th order and degree
 - Earth: EGM96
 - Mars: GMM-2B
 - Luna: GLGM2
- Planetary Magnetic Field Models
 - IGRF up to 10th order (Earth only)
 - Tilted offset dipole field
- Earth Atmospheric Density Models
 - NRLMSISE00
 - MSIS-86 (thanks to John Downing)
 - Jacchia-Roberts Atmospheric Density Model (NASA SP-8021)
- Simple exponential Mars atmosphere density model
 - New models easily incorporated as the state of the art advances

Dynamics Models

- Full nonlinear “6DOF” (actually N-DOF) dynamics
- Attitude Dynamics
 - One or many bodies
 - Tree topology (no kinematic loops)
 - Each body may be rigid or flexible
 - Joints may combine rotational and translational DOFs
 - May be gimbaled or spherical
 - Slosh may be modeled as a pendulum (lo-fi, quick to implement and run)
 - 42 may run concurrently with Star-CCM CFD software for hi-fi slosh
 - Wheels embedded in Body[0]
 - Torques from actuators, aerodynamic drag, gravity-gradient, solar radiation pressure, joint torques
- Orbit Dynamics
 - Two- or three-body orbits
 - Encke or Euler-Hill (Clohessy-Wiltshire) for relative orbit motion (good for formation flying, prox ops)
 - Forces from actuators, aerodynamic drag, non-spherical gravity, third-body gravity, solar radiation pressure

Reference Frames are Important!

- In any dynamics problem beyond the spinning top, a systematic approach to reference frames and the relationships between them is vital
- For 42, we define several fundamental reference frames, and notational conventions to keep quaternions and direction cosines sorted out

Reference Frames (1 of 2)

- Heliocentric Ecliptic (H)
 - Planet positions expressed in this frame
- Each world has an inertial (N) and rotating (W) frame
 - For Earth, N = ECI (True of date), W = ECEF
 - N is the bedrock for orbits, S/C attitude dynamics

Reference Frames (2 of 2)

- Each reference orbit has a reference point R
 - For two-body orbit, R moves on Keplerian orbit
 - For three-body orbit, R propagates under influence of both attracting centers (as point masses)
 - S/C orbit perturbations integrated with respect to R
- Associated with each R is a LVLH frame (L) and a formation frame (F)
 - F is useful for formation-flying scenarios
 - F may be offset from R, may be fixed in N or L
- Each spacecraft has one or more Body (B) frames and one LVLH frame (L)
 - L(3) points to nadir, L(2) points to negative orbit normal
 - SC.L is distinct from Orb.L, since SC may be offset from R

Representing Attitude

- There are several ways to represent the rotation between two reference frames
 - Direction Cosines
 - Euler Angles
 - Quaternions (aka Euler Parameters)
 - and more
- They all have their strengths and weaknesses
 - Learn them all!

Strengths and Weaknesses of Attitude Representations

Representation	Strengths	Weaknesses	Best Used For
Direction Cosines	<ul style="list-style-type: none"> • Work well with vectors • Easy to concatenate rotations • Moderately intuitive (dot products) • No singularities 	<ul style="list-style-type: none"> • 9 params for 3 DOF 	<ul style="list-style-type: none"> • Transforming Vectors
Quaternions	<ul style="list-style-type: none"> • Efficient (4 params for 3 DOF) • No singularities 	<ul style="list-style-type: none"> • Not intuitive 	<ul style="list-style-type: none"> • Propagating Equations of Motion
Euler Angles	<ul style="list-style-type: none"> • Intuitive • 3 params for 3DOF 	<ul style="list-style-type: none"> • Singularities • 24 Variants 	<ul style="list-style-type: none"> • Input, Output • Gimballed Joints

Notation for Quaternions, DCMs

- The rotation from frame A to frame B may be described by the direction cosine matrix

$${}^B C^A_{ij} = \hat{b}_i \cdot \hat{a}_j$$

- Given the components of a vector in A , its components in B may be found by the multiplication

$${}^B \mathbf{v} = {}^B C^A \mathbf{v}$$

- In C, we write the DCM as CBA to preserve order of superscripts, eg

$$\text{MxV}(\text{CBA}, \mathbf{v}_a, \mathbf{v}_b)$$

- Quaternions are another way to describe rotations. We use a parallel notation:

$$\text{QxV}(\text{qba}, \mathbf{v}_a, \mathbf{v}_b)$$

- These and similar conventions promote concise, *unambiguous* code

Interfaces to Matlab

- 42 generates ASCII output files, which may be loaded into Matlab (or whatever) for post-processing and plotting
- Using Matlab's `mcc` utility, m-files may be translated into C, and compiled and linked into 42

Matlab + 42 = Monte Carlo

- 42 can be called from within Matlab using the `system` command
- Use Matlab as the MC executive
 - Generate initial conditions, parameters
 - Write to 42's input files
 - Run 42
 - Process and save data
 - Repeat
- Use 42 as the high-speed, high-fidelity component

Matlab/42 Example

```
for Irun=1:Nrun,

    % Compute initial attitude
    CRN = TRIAD(tvn(Irun,:),svn,[0 0 1],[1 0 0]);
    qrn = C2Q(CRN);

    % Write target to file
    Outdata = [TrgRA(Irun) TrgDec(Irun)];
    save -ascii ./MOMBIAS/TargetRaDec.inp Outdata

    % Write initial attitude to file
    line = sprintf('%f %f %f %f ! Quaternion\n', qrn(1),qrn(2),qrn(3),qrn(4));
    OverwriteLineInFile('./MOMBIAS/GLAST.inp',21,line);

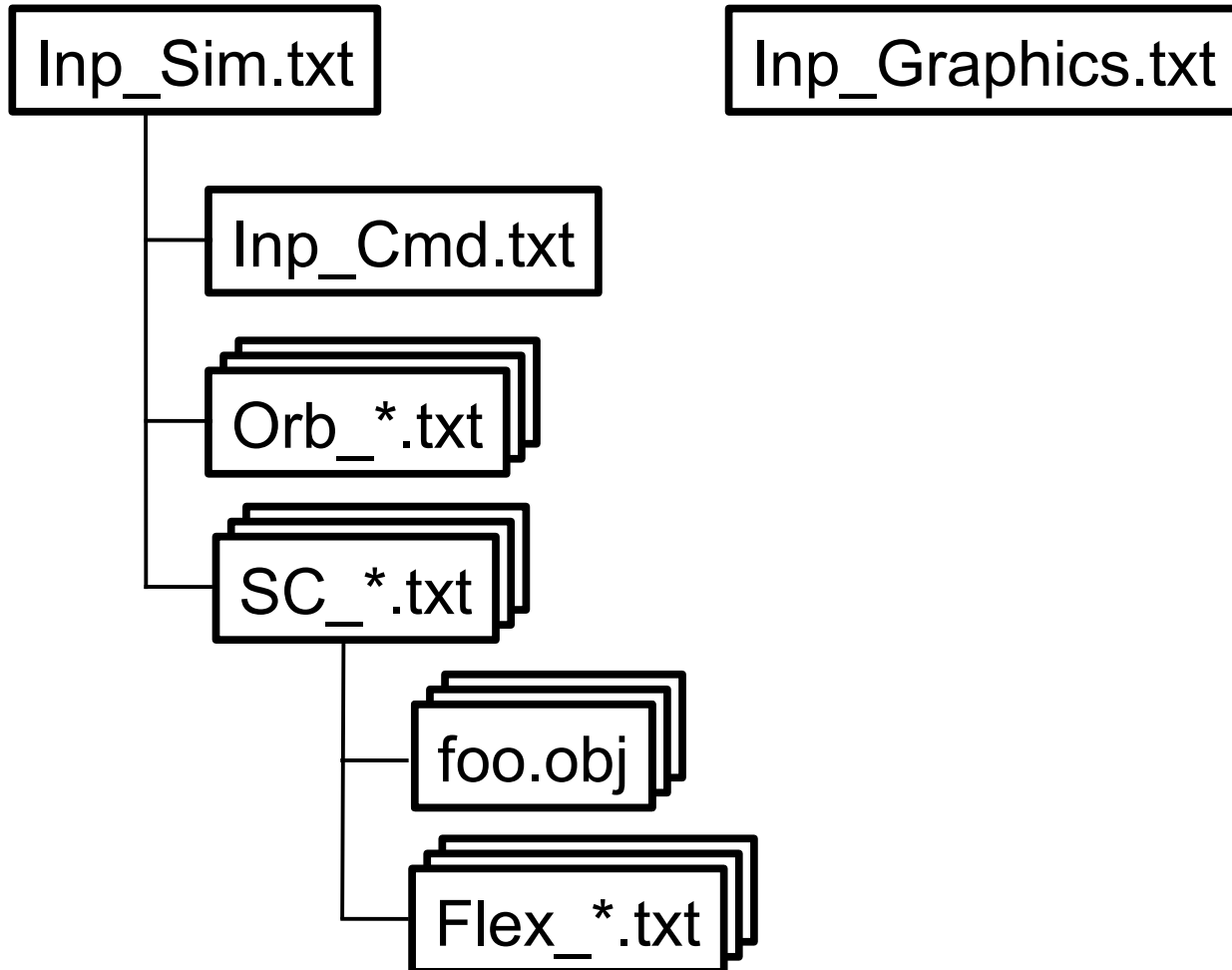
    % Run 42 for three days.
    system('./42 MOMBIAS');

    % Record pointing histogram.
    load ./MOMBIAS/AngleToGo.42
    [HistCount(Irun,:),HistAng(Irun,:)] = hist(AngleToGo,20);

end
```

Input File Overview

Input File Hierarchy



Geometry Definition

- Geometry uses Wavefront “Object” format
 - ASCII, human-readable
 - Can hand-generate models *in theory*
 - In practice, you’ll want mechanical help
- I use Wings3D solid modeling software
 - Free, multi-platform
- 42 can support myriad-polygon models, limited mainly by your patience
 - I get impatient at about 10,000 polys
- Note that aerodynamic and solar-pressure force and torque computations use these models
 - Interior polygons, self-shadowing are error sources

```
# Exported from Wings 3D 0.99.00b
mtllib Altair.mtl
o cylinder9
#12 vertices, 8 faces
v 2.11050391 1.21850000 -1.80666667
v 2.11050391 -1.21850000 -1.80666667
[...]
vn 0.00000000e+0 9.1113913e-17 1.00000000
vn 0.50000000 0.86602540 0.00000000e+0
[...]
g cylinder9_SHINY_WHITE
usemtl SHINY_WHITE
f 1//1 6//16 5//13 4//10 3//7 2//4
f 1//2 7//20 12//35 6//17
[...]
```

Excerpt from Altair.obj