## Preliminary Examination

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prelim exam

#### Dirac equation and charge conjugation

$$\mathcal{L} = rac{i}{2} ig[ \, \overline{\psi} \gamma^{\mu} \, (\partial_{\mu} \psi) 
onumber \ - ig( \partial_{\mu} \overline{\psi} ig) \, \gamma^{\mu} \psi ig] - ig( m{m} \overline{\psi} \psi ig)$$

$$e^{-} \iff e^{+}$$

$$\hat{\psi} \to \hat{\psi}^{C} = C\gamma^{0}\hat{\psi}^{*}$$

$$C = \begin{pmatrix} i\sigma_{2} & 0\\ 0 & -i\sigma_{2} \end{pmatrix}$$

$$(i\partial - m)\psi = 0$$
$$\psi = \begin{pmatrix} \varphi \\ \chi \end{pmatrix}$$

The Quantum Theory of the Electron,

By P. A. M. Dmac, St. John's College, Cambridge.

(Communicated by R. H. Fowler, F.R.S.—Received January 2, 1928.)

The new quantum mechanics, when applied to the problem of the structure of the atom with point-charge electrons, does not give results in genement with experiment. The discrepancies consist of "duplexity" phenomena, he observed number of stationary states for an electron in an atom being twice the number given by the theory. To sense the difficulty, condentia and Otherheek have introduced the idea of an electron with a spin angular momentum of half a quantum and a magnetic moment of one Bolt rangenton. This model for the electron has been fitted into the new mechanics by Pauli,\* and Darwin,† wowking with an equivalent theory, has shown that it gives results in agreement.

#### Majorana fermions and neutrinos

- Majorana fermions can be seen as solutions to the Dirac equation with an extra constraint
- $\psi = \psi^{C}$
- implies these fermions are chargeless, and their own anti-particles

The left and right handed components of the spinor are no longer independent

$$\Psi_{majorana} = \psi_L + \psi_R$$

mass terms:

$$\begin{split} & \left( \overline{\hat{\psi^C}}_L \hat{\psi}_L + \overline{\hat{\psi}}_L \hat{\psi}_L^C \right)_{\textit{Majorana}} \\ & \left( \overline{\hat{\psi}}_L \hat{\psi}_R + \overline{\hat{\psi}}_R \hat{\psi}_L \right)_{\textit{Dirac}} \end{split}$$

# why is it still unknown if neutrinos are Dirac or Majorana fermions?

- Both types of fermions are compatible with observed features of neutrinos, including neutrino oscillations. majorana neutrinos and their anti-matter partners are only distinguishable by their chirality.
- In the ultra-relativistic limit, helicity and chirality (lepton number). we can never boost into a frame where the neutrino has right handed helicity. (does the goldhaber experiment measure helicity or chirality? chirality is the sign of the phase of the wavefunction for the dirac fermion, it describes which path
- it takes in the complex plane.

  expand the fermionic propagator
  in the limit are and the

The left and right handed components of the spinor are no longer independent

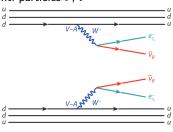
$$\Psi_{majorana} = \psi_L + \psi_R$$

mass terms:

$$egin{aligned} \left( \overline{\hat{\psi}^{\mathcal{C}}}_L \hat{\psi}_L + \overline{\hat{\psi}}_L \hat{\psi}_L^{\mathcal{C}} 
ight)_{Majorana} \ \left( \overline{\hat{\psi}}_L \hat{\psi}_R + \overline{\hat{\psi}}_R \hat{\psi}_L 
ight)_{Dirac} \end{aligned}$$

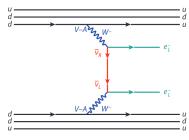
#### SM and BSM double-beta decay

 $2\nu\beta\beta$ : the neutrinos are Dirac fermions, and have distinct anti-particles  $\nu$ ,  $\overline{\nu}$ 



extremely rare process, observed in direct detection experiments (CUORE-0, GERDA, XENON, EXO, NEMO)  $T_{1/2}^{2\nu\beta\beta}\sim 10^{19}-10^{24}yr$ 

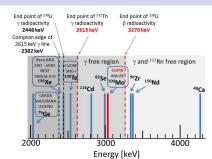
 $\mathbf{0}
uetaeta$ : the neutrinos are Majorana fermions and are absorbed as virtual particles

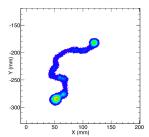


Never before observed in any experiment:  $T_{1/2}^{0
uetaeta}>10^{26}yr$  (Lepton number violation)

#### experimental signature and detection

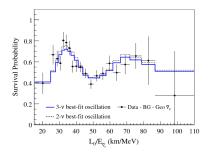
- mono-energetic peak with 2 electrons leaving the reaction.
- tomography of electron tracks can help reduced backgrounds, except for the  $2\nu\beta\beta$  irreducible background
- experimental goal is to measure  $T_{1/2}^{0
  uetaeta}$  to constrain the effective neutrino mass  $m_{etaeta}$ .
- this measurement complements measurements of Σ and m<sub>β</sub> from cosmology and kinematic experiments, respectively.





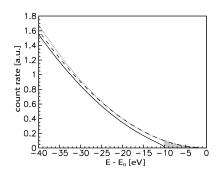
#### neutrino mass and oscillations

- Homestake Mine and the solar neutrino anomaly  ${}^{37}Cl + \nu_e \rightarrow {}^{37}Ar + e^-$  observes factor of 3 discrepancy in  $\nu_e$  flux.
- Sudbury Neutrino Observatory (SNO) measures total neutrino flux:  $\nu_{\mu}, \nu_{\tau}$  account for  $\nu_{e}$  deficit.
- KAMLAND, a reactor neutrino experiment shows survival probability of electron neutrinos  $P = 1 \sin^2(2\theta)\sin^2(\Delta m^2L/4E)$



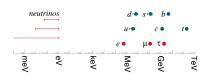
#### neutrino mass and the $\beta$ -spectrum

- Tritium ( $^3H$ ) is a viable isotope for directly probing the neutrino mass low Q value (decays in region  $E_0 E$  scale with  $(E_0 E)^3/Q^3$  and short-life to reduce line broadening.....?
- Measure deviation from the mass-less spectrum's linearity at large energies.
- This method is most-direct and model independent probe of the neutrino's mass. Constructing a sensitive experiment is very challenging (KATRIN may be last of its kind)



#### mass scale

- Physicists believed neutrinos to be massless for decades, due to Standard Model predictions (Higgs mechanism for leptons) and experimental data
- Neutrinos are orders of magnitude lighter than any other fundamental particle. This in addition to them being chargeless, makes it hard to measure their presence in experiments.
- The common model for describing BSM neutrino mass states is the light-neutrino mixing model (3 mass eigenstates)



#### neutrino mixing

A given flavor eigenstate is a linear combination of mass eigenstates (PMNS-matrix)

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = U \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

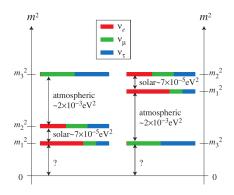
parameters: angles  $\theta_{12}, \theta_{13}, \theta_{23}$  CP-violating phases  $\delta_{CP}, \alpha_1, \alpha_2$ , and neutrino masses  $m_1, m_2, m_3$ 

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{CP}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{CP}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{CP}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{CP}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{CP}} & c_{23}c_{13} \end{pmatrix}$$

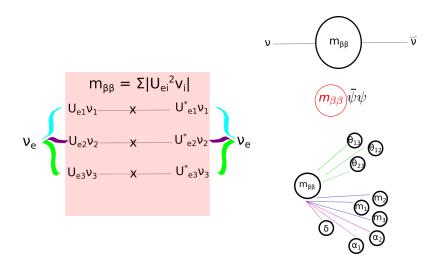
$$\times \begin{pmatrix} e^{i\alpha_{1}/2} & 0 & 0 \\ 0 & e^{i\alpha_{2}/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

#### mass hierarchy

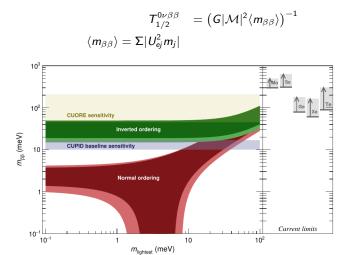
Neutrino oscillation experiments are sensitive to  $\Delta m_{ij}^2 = m_i^2 - m_j^2$  This, coupled with the neutrinos' very small mass scale, leads to ambiguity in their ordering. **normal ordering**  $\implies m_3^2 > m_2^2 > m_1^2$  **inverted ordering**  $\implies m_2^2 > m_1^2 > m_2^2$ 



#### effective Majorana neutrino mass

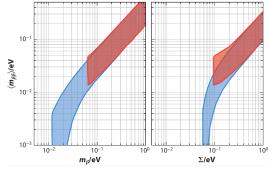


#### zero neutrino half-life



#### complementary limits

- $\begin{tabular}{ll} $\tt Limits on the effective $0\nu\beta\beta$ \\ halflife are also constrained by kinematic and cosmological searches \\ \end{tabular}$
- $m_{\beta} = \sqrt{U_{ei}^2 m_i^2} \text{ and }$   $\Sigma = m_1 + m_2 + m_3$
- $m_{eta} < 1.1 \mathrm{eV}$  $\Sigma < 0.3 \mathrm{eV}$



inverted normal

#### designing a direct-detection experiment

Isotope	$Q_{\beta\beta}$ (keV)	I.A. (%)	$G^{0\nu}$	$H^{0\nu}$
<sup>48</sup> Ca	4272	0.187	24.81	826.2
<sup>76</sup> Ge	2039	7.8	2.36	49.6
<sup>82</sup> Se	2995	8.73	10.16	198.1
$^{96}Zr$	3350	2.8	20.58	342.7
$^{100}$ Mo	3034	9.63	15.92	254.5
<sup>110</sup> Pd	2018	11.72	4.82	70.0
116 Cd	2814	7.49	16.70	230.1
<sup>124</sup> Sn	2287	5.79	9.04	116.5
<sup>128</sup> Te	866	31.69	0.59	7.4
<sup>130</sup> Te	2527	33.8	14.22	174.8
<sup>136</sup> Xe	2458	8.9	14.58	171.4
<sup>148</sup> Nd	1929	5.76	10.10	109.1
150 Nd	3371	5.64	63.03	671.7
<sup>154</sup> Sm	1215	22.7	3.02	31.3
<sup>160</sup> Gd	1730	21.86	9.56	95.5
<sup>198</sup> Pt	1047	7.2	7.56	61.0

Rate 
$$(\mathcal{O}_{majorana}) \propto rac{m_{
u}^2}{Q_{etaeta}}$$

$$T_{1/2}^{0
uetaeta} = \left(G|\mathcal{M}|^2\langle m_{etaeta}
ight)^{-1}$$
  $F_{0
u} = au_{1/2}^{bkg.fluct.} = \ln(2)N_{etaeta}\epsilonrac{t}{n_{bkg}}$   $F_{0
u} = ln(2) imes rac{N_A}{W} \left[a\epsilon\sqrt{rac{Mt}{b\Delta}}
ight]$ 

#### backgrounds

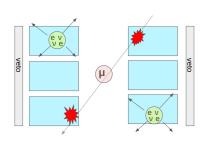
- Primordial radioisotopes:  $^{238}U$  and  $^{232}Th$  generate chain of  $\beta, \gamma$  near  $Q_{\beta\beta}$ .
- =  $^{222}$  Rn can exist in gaseous state produces  $\beta + \gamma$  impinging on  $Q_{\beta\beta}$
- <sup>208</sup> TI emits γ at 2615 keV, affecting a number of candidate source isotopes.
- the intrinsically irreducible background  $2\nu\beta\beta$ :  $\Gamma_{2\nu\beta\beta}$  is negligible in some isotopes ( $^{130}$  Te), but not others ( $^{100}$  Mo)
- muons enter the backgrounds in a number of ways: prompt signal + long-lived isotope activation + neutron production

**goal**:  $B.I. = 0.1 counts / (FWHM \cdot tonne \cdot year)$ 

#### $2\nu\beta\beta$ events and muons

- The rate of  $2\nu\beta\beta$  events is not negligible<sup>1</sup> in the CUPID array
- Minimizing the distance cut helps avoid mis-labelling random  $2\nu\beta\beta$  coincidences as multiplicity 2.
- Assuming a simple muon veto geometry, increasing the distance cut rejects more muon events.

$$T_{1/2}^{2
uetaeta}=7.1*10^{18} yr$$
  $ightarrow$  single crystal rate  $\sim 3mHz$ 

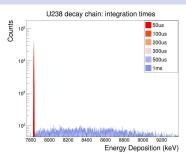


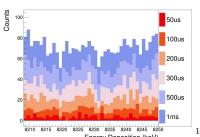
<sup>&</sup>lt;sup>1</sup>arxiv.org/abs/1912.07272.

## Integration time in the CUPID array

$$\begin{array}{ccc} ^{238}U \xrightarrow{\alpha} ^{234}Th \rightarrow ... \\ ^{214}Bi \xrightarrow{\beta} ^{214}Po \xrightarrow{\alpha} ^{210}Pb \ ... \end{array}$$

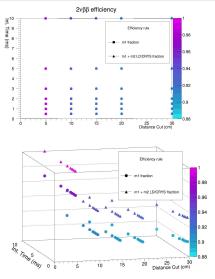
- 100,000 Uranium-238 events (full chain)
- $T_{1/2}$  <sup>214</sup>Po  $\sim 160 \mu s$
- lacksquare  $T_{1/2}$   $^{214}$ Bi  $\sim$  20 minutes





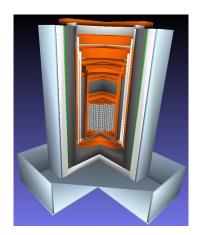
#### $2\nu\beta\beta$ tagging simulation

- 40 million events generated in crystal volume.
- expect  $\sim$  90% efficiency for  $2\nu\beta\beta$  tagging, when using approximate CUORE parameters
- larger distance cut causes more random coincidences and more variation with integration time
- bremsstrahlung, escape, random coincidences are primary contributors



#### implementation of muon background

- incorporating scintillating muon veto geometry into monte-carlo simulations
- muon flux at LNGS is  $3 \cdot 10^{-8}$  muons  $/(s \cdot cm^2)$
- muon suppression versus distance cut optimization.



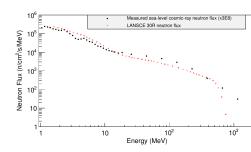
preliminary muon veto geometry

## spallation in <sup>100</sup>Mo

$$\sigma = \sigma_0 f(A) f(E) e^{-P\Delta A}$$

$$\times \exp(-R|Z| - SA + TA^2 + UA^3|^{\nu} \Omega \eta \xi)$$

- semi-empirical cross section formula
- at LANL, a beam bombarding protons on a tungsten target mimicks the cosmic ray spectrum.
- 3 · 10<sup>8</sup> scaling in flux ⇒ 2 days of beam time equates to 1 million years of sea-level exposure.



## far reaching consequences of observing $0\nu\beta\beta$

- baryogenesis through leptogenesis
- charge parity violating process in the standard model
- lepton number conservation (accidental symmetry of the standard model)

$$m_{
u} = U egin{pmatrix} m_1 & 0 & 0 \ 0 & m_2 & 0 \ 0 & 0 & m_3 \end{pmatrix} U^T$$

$$\langle m_{\beta\beta} \rangle = |U_{ij}^2 m_j|$$

#### $0\nu\beta\beta$ half-life and sensitivity

 direct detection experiments typically characterize the half-life in terms of the effective neutrino mass m<sub>ββ</sub>

$$T_{1/2}^{0
uetaeta} = \left(G|\mathcal{M}|^2\langle m_{etaeta}
ight)^{-1} \ \simeq 10^{27-28}\left(rac{0.01~{
m eV}}{\langle m_{etaeta}
angle}
ight)^2$$
 years

$$F_{0
u} = au_{1/2}^{\mathit{bkg.fluct.}} = \ln(2) \mathit{N}_{etaeta} \epsilon rac{t}{\mathit{n}_{B}}$$

#### $0\nu\beta\beta$ half-life and sensitivity

to probe

$$egin{aligned} T_{1/2}^{0
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angle
ight)^{-1} \ &\simeq 10^{27-28}\left(rac{0.01~{
m eV}}{\langle m_{etaeta}
angle}
ight)^2 {
m years} \ &T_{1/2}^{0
uetaeta} \sim ~\epsilon\sqrt{rac{M_{iso}\,t}{b\Delta E}} \end{aligned}$$