Preliminary Examination

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prelim exam

Dirac equation and charge conjugation

$$\mathcal{L} = \frac{i}{2} \left[\ \overline{\psi} \gamma^{\mu} \left(\partial_{\mu} \psi \right) - \left(\partial_{\mu} \overline{\psi} \right) \gamma^{\mu} \psi \right] - \boxed{m \overline{\psi} \psi}$$

$$e^{-} \iff e^{+}$$

$$\hat{\psi} \to \hat{\psi}^{C} = C\gamma^{0}\hat{\psi}^{*}$$

$$C = \begin{pmatrix} i\sigma_{2} & 0\\ 0 & -i\sigma_{2} \end{pmatrix}$$

$$(i\partial - m)\psi = 0$$
$$\psi = \begin{pmatrix} \varphi \\ \chi \end{pmatrix}$$

The Quantum Theory of the Electron,

By P. A. M. Dmac, St. John's College, Cambridge.

(Communicated by R. H. Fowler, F.R.S.—Received January 2, 1928.)

The new quantum mechanics, when applied to the problem of the structure of the atom with point-charge electrons, does not give results in agreement with experiment. The discrepancies consist of "duplexity" phenomens, the observed number of stationary states for an electron in an atom being twice the number given by the theory. To meet the difficulty, condemit and Otherheek have introduced the idea of an electron with a spin angular momentum of half a quantum and a magnetic moment of one Bolt magneton. This model for the electron has been fitted into the new mechanics by Pauli,* and Darwin,† working with an equivalent theory, has shown that it gives results in agreement.

Majorana fermions and neutrinos

Majorana fermions can be seen as solutions to the Dirac equation with an extra constraint: $\psi = \psi^{C}$

$$\begin{split} \hat{\psi}(x_{\mu}) &= \int \frac{d^{3}\overline{k}}{(2\pi)^{3}2\omega_{k}} \sum_{s} \\ &\left[\hat{\boldsymbol{b}}(\overline{k},s)u(\overline{k},s)e^{-ik^{\mu}x_{\mu}} + \hat{\boldsymbol{b}}^{\dagger}(\overline{k},s)v(\overline{k},s)e^{+ik^{\mu}x_{\mu}} \right] \\ &Q &= \int d^{3}k \ \left(b(k)b^{\dagger}(k) - b(k)b^{\dagger}(k) \right) \end{split}$$

⇒ Majorana fermions are chargeless, and their own anti-particles

The left and right handed components of the spinor are no longer independent

$$\hat{\psi} = \begin{pmatrix} \hat{\varphi} \\ -i\sigma_2(\varphi^{\dagger})^T \end{pmatrix} = \begin{pmatrix} \hat{\psi}_L \\ \hat{\psi}_R \end{pmatrix}$$

mass terms:

$$\begin{split} & \left(\overline{\hat{\psi}^{\mathsf{C}}}_{\mathsf{L}} \hat{\psi}_{\mathsf{L}} + \overline{\hat{\psi}}_{\mathsf{L}} \hat{\psi}_{\mathsf{L}}^{\,\mathsf{C}} \right)_{\mathsf{Majorana}} \\ & \left(\overline{\hat{\psi}}_{\mathsf{L}} \hat{\psi}_{\mathsf{R}} + \overline{\hat{\psi}}_{\mathsf{R}} \hat{\psi}_{\mathsf{L}} \right)_{\mathsf{Dirac}} \end{split}$$

why is Dirac/Majorana nature still unknown?

- Both types of fermions are compatible with observed features of neutrinos, including neutrino oscillations. majorana neutrinos and their anti-matter partners are only distinguishable by their chirality.
- In the ultra-relativistic limit, helicity and chirality are equivalent e.g. we can never boost into a frame where a neutrino has right handed helicity.
- V-A interaction has left-handed projection operator, so we only observe β-decay that is also consistent with Dirac neutrinos that respect lepton number conservation (an accidental symmetry).

$$\mathcal{L}_{cc} = -\frac{g}{\sqrt{2}} \left[\overline{e} \gamma^{\mu} \left(\frac{1 - \gamma_5}{2} \right) \nu W_{\mu}^{-} \right.$$

$$\left. + \overline{\nu} \gamma^{\mu} \left(\frac{1 - \gamma_5}{2} \right) e W_{\mu}^{+} \right]$$

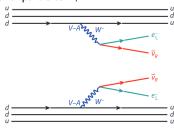
$$\hat{\nu} = \int \frac{d^3 \overline{k}}{(2\pi)^3 2\omega_k} \sum_s$$

$$\left[\hat{b}(\overline{k}, s) u(\overline{k}, s) e^{-ikx} \right.$$

$$\left. + \hat{b}^{\dagger}(\overline{k}, s) v(\overline{k}, s) e^{+ikx} \right]$$

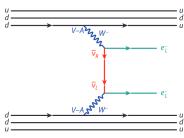
SM and BSM double-beta decay

 $2\nu\beta\beta$: the neutrinos are Dirac fermions, and have distinct anti-particles ν , $\overline{\nu}$



extremely rare process, observed in direct detection experiments $T_{1/2}^{2\nu\beta\beta}\sim 10^{19}-10^{24}yr$

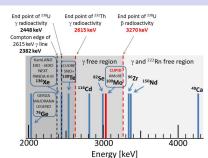
0
uetaeta: the neutrinos are Majorana fermions and are absorbed as virtual particles

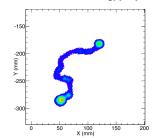


Never before observed in any experiment: $T_{1/2}^{0
uetaeta}>10^{26}yr$ (Lepton number violation)

experimental signature and detection

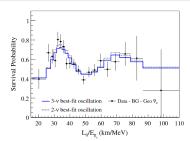
- mono-energetic peak with 2 electrons leaving the reaction.
- tomography of electron tracks can help reduced backgrounds, except for the $2\nu\beta\beta$ irreducible background
- experimental goal is to measure $T_{1/2}^{0
 uetaeta}$ to constrain the effective neutrino mass m_{etaeta} .
- this measurement complements measurements of Σ and m_β from cosmology and kinematic experiments, respectively.

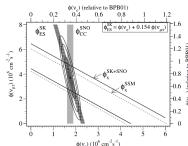




neutrino mass and oscillations

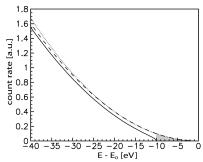
- Homestake Mine and the solar neutrino anomaly ${}^{37}Cl + \nu_e \rightarrow {}^{37}Ar + e^-$ observes factor of 3 discrepancy in ν_e flux.
- Sudbury Neutrino Observatory (SNO) measures total neutrino flux: ν_{μ}, ν_{τ} account for ν_{e} deficit.
- KAMLAND, a reactor neutrino experiment shows survival probability of electron neutrinos $P = 1 \sin^2(2\theta)\sin^2(\Delta m^2L/4E)$





neutrino mass and the β -spectrum

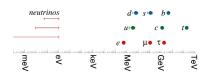
- Tritium (3H) is a viable isotope for directly probing the neutrino mass low Q value (decays in region $E_0 E$ scale with $(E_0 E)^3/Q^3$ and short-life to reduce line broadening.....?
- Measure deviation from the mass-less spectrum's linearity at large energies.
- This method is most-direct and model independent probe of the neutrino's mass. Constructing a sensitive experiment is very challenging (KATRIN may be last of its kind)





mass scale

- Physicists believed neutrinos to be massless for decades, due to Standard Model predictions (Higgs mechanism for leptons) and experimental data
- Neutrinos are orders of magnitude lighter than any other fundamental particle. This in addition to them being chargeless, makes it hard to measure their presence in experiments.
- The common model for describing BSM neutrino mass states is the light-neutrino mixing model (3 mass eigenstates)



neutrino mixing

A given flavor eigenstate is a linear combination of mass eigenstates (PMNS-matrix)

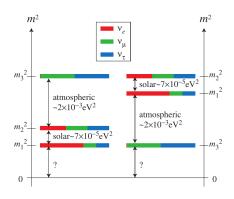
$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = U \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

parameters: angles $\theta_{12}, \theta_{13}, \theta_{23}$ CP-violating phases $\delta_{CP}, \alpha_1, \alpha_2$, and neutrino masses m_1, m_2, m_3

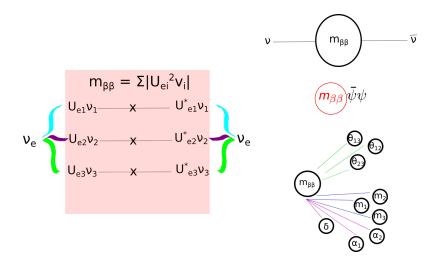
$$\begin{split} U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{CP}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{CP}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{CP}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{CP}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{CP}} & c_{23}c_{13} \end{pmatrix} \\ \times \begin{pmatrix} e^{i\alpha_{1}/2} & 0 & 0 \\ 0 & e^{i\alpha_{2}/2} & 0 \\ 0 & 0 & 1 \end{pmatrix} \end{split}$$

mass hierarchy

Neutrino oscillation experiments are sensitive to $\Delta m_{ij}^2 = m_i^2 - m_j^2$ This, coupled with the neutrinos' very small mass scale, leads to ambiguity in their ordering. **normal ordering** $\implies m_3^2 > m_2^2 > m_1^2$ **inverted ordering** $\implies m_2^2 > m_1^2 > m_2^2$

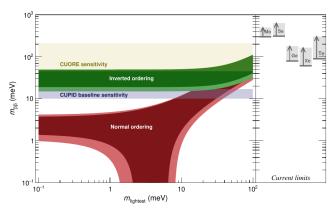


effective Majorana neutrino mass



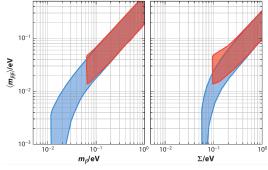
zero neutrino half-life

$$T_{1/2}^{0
uetaeta} = \left(G|\mathcal{M}|^2\langle m_{etaeta}
angle
ight)^{-1} \ \langle m_{etaeta}
angle = \Sigma|U_{ej}^2m_j|$$



complementary limits

- $\begin{tabular}{ll} \blacksquare Limits on the effective <math>0 \nu \beta \beta $$ halflife are also constrained by kinematic and cosmological searches $$$
- $m{m}_eta = \sqrt{U_{ei}^2 m_i^2}$ and $\Sigma = m_1 + m_2 + m_3$
- $m_{eta} < 1.1 \mathrm{eV}$ $\Sigma < 0.3 \mathrm{eV}$



inverted normal

designing a direct-detection experiment

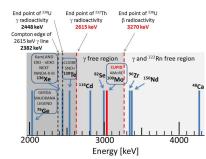
Isotope	$Q_{\beta\beta}$ (keV)	I.A. (%)	$G^{0\nu}$	$H^{0\nu}$
⁴⁸ Ca	4272	0.187	24.81	826.2
⁷⁶ Ge	2039	7.8	2.36	49.6
⁸² Se	2995	8.73	10.16	198.1
96Zr	3350	2.8	20.58	342.7
¹⁰⁰ Mo	3034	9.63	15.92	254.5
¹¹⁰ Pd	2018	11.72	4.82	70.0
116 Cd	2814	7.49	16.70	230.1
¹²⁴ Sn	2287	5.79	9.04	116.5
¹²⁸ Te	866	31.69	0.59	7.4
¹³⁰ Te	2527	33.8	14.22	174.8
¹³⁶ Xe	2458	8.9	14.58	171.4
148 Nd	1929	5.76	10.10	109.1
150 Nd	3371	5.64	63.03	671.7
154 Sm	1215	22.7	3.02	31.3
¹⁶⁰ Gd	1730	21.86	9.56	95.5
¹⁹⁸ Pt	1047	7.2	7.56	61.0

$$T_{1/2}^{0
uetaeta}=\left(G|\mathcal{M}|^2\langle m_{etaeta}
angle
ight)^{-1}$$

$$F_{0
u} = \ln(2)N_{etaeta}\epsilonrac{t}{n_{bkg}}$$
 $F_{0
u} = \ln(2) imesrac{N_A}{W}\left[a\epsilon\sqrt{rac{Mt}{b\Delta}}
ight]$

backgrounds

- Primordial radioisotopes: ^{238}U and ^{232}Th generate chain of β, γ near $Q_{\beta\beta}$.
- $lacksquare ^{222}Rn$ can exist in gaseous state produces $eta+\gamma$ impinging on Q_{etaeta}
- ²⁰⁸ TI emits γ at 2615 keV, affecting a number of candidate source isotopes.
- the intrinsically irreducible background $2\nu\beta\beta$: $\Gamma_{2\nu\beta\beta}$ is negligible in some isotopes (130 Te), but not others (100 Mo)
- muons enter the backgrounds in a number of ways: prompt signal + long-lived isotope activation + neutron production



Next Generation $0\nu\beta\beta$ experiments goal: probe $0\nu\beta\beta$ to a half life of 10^{28} years

requirement: $B.I. = 0.1 counts / (FWHM \cdot tonne \cdot year)$

CUPID conservative goal: B.I. =

 $0.5 \ counts \ / \ (FWHM \cdot tonne \cdot year)$

CUORE experiment

- Array of 988 TeO₂ bolometers instrumented with NTD sensors.
- CUORE cryostat maintains about 200kg of ¹³⁰ Te around 10mK
- current limit on $0\nu\beta\beta$ in ¹³⁰ Te: $T_{1/2}^{0\nu} > 3.2 \cdot 10^{25} yr$



Cuore

cryostat is a dilution refrigerator that cools 750kg of mass to 10mK

CUORE backgrounds

- irreducible $2\nu\beta\beta$ rate is negligible relative to
- CUORE cryostat maintains about 200kg of ¹³⁰ Te around 10mK.
- current background index is $10^{-2} counts/(keV \cdot kg \cdot yr)$



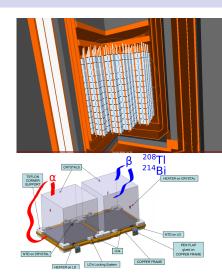
Cuore

cryostat is a dilution refrigerator that cools 750kg of mass to 10mK

CUPID experiment

- Proposed array of 1596 Li₂Mo₄ to be deployed in the CUORE cryostat
- lacksquare aims to address dominant background of lpha-particles in CUORE.
- capitalize on processes (CR6) and technology (cryostat and calibration) developed for CUORE
- current research is to address new backgrounds introduced with new source material to improve BI

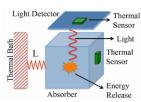


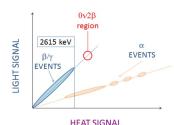


Lithium molybdate

- Li₂MoO₄ crystals allow for discrimination of α backgrounds from $\beta\beta$ events (Q=3034keV) via thermal + scintillation signals.
- relatively high isotopic abundance of ¹⁰⁰Mo (10%)
- enrichment above 95% already demonstrated in CUPID-Mo¹.

Scintillating Bolometer



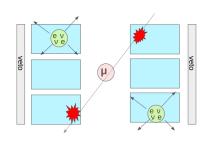


¹arxiv.org/abs/1909.02994.

$2\nu\beta\beta$ events and muons

- The rate of $2\nu\beta\beta$ events is not negligible² in the CUPID array
- Minimizing the distance cut helps avoid mis-labelling random $2\nu\beta\beta$ coincidences as multiplicity 2.
- Assuming a simple muon veto geometry, increasing the distance cut rejects more muon events.

$$T_{1/2}^{2
uetaeta}(^{100}\mbox{Mo}) = 7.1*10^{18}\mbox{yr}$$
 $ightarrow$ single crystal rate $\sim 3\mbox{mHz}$ $T_{1/2}^{2
uetaeta} \sim 10^{19} - 10^{24}\mbox{yr}$

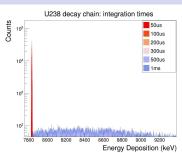


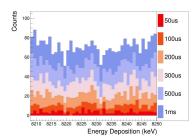
²arxiv.org/abs/1912.07272.

Integration time in the CUPID array

$$\begin{array}{ccc} ^{238}U \stackrel{\alpha}{\rightarrow} & ^{234}Th \rightarrow ... \\ ^{214}Bi \stackrel{\beta}{\rightarrow} & ^{214}Po \stackrel{\alpha}{\rightarrow} & ^{210}Pb \ ... \end{array}$$

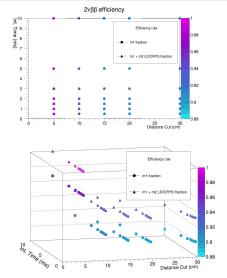
- 100,000 Uranium-238 events (full chain)
- lacktriangleright $T_{1/2}$ 214 Po $\sim 160 \mu s$
- ${f T}_{1/2}$ $^{214}{
 m Bi}\sim 20$ minutes





$2\nu\beta\beta$ tagging simulation

- 40 million events generated in crystal volume.
- expect \sim 90% efficiency for $2\nu\beta\beta$ tagging, when using approximate CUORE parameters
- larger distance cut causes more random coincidences and more variation with integration time
- bremsstrahlung, escape, random coincidences are primary contributors

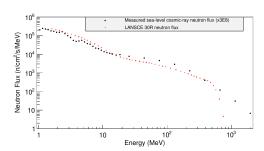


spallation in ¹⁰⁰Mo

$$\sigma = \sigma_0 f(A) f(E) e^{-P\Delta A}$$

$$\times \exp(-R|Z| - SA + TA^2 + UA^3|^{\nu} \Omega \eta \xi)$$

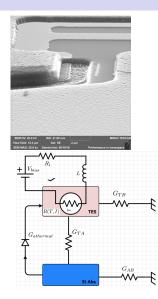
- semi-empirical cross section formula
- at LANL, a beam bombarding protons on a tungsten target mimics the cosmic ray spectrum.
- 3 · 10⁸ scaling in flux ⇒ 2 days of beam time equates to 1 million years of sea-level exposure.
- currently implementing spectrum for parasitic neutron simulation and are proposing to perform a measurement at LANL neutron beam.



product	Q(MeV)	$T_{1/2}$	A _{initial} (Bq)	A_{1yr} (Bq)
⁵⁶ Co	4.5	77 d	$5.1 \cdot 10^{-8}$	$2.1 \cdot 10^{-9}$
88 Y	3.6	106 d	$2.2 \cdot 10^{-5}$	$2 \cdot 10^{-6}$
²⁶ AI	4	$7 \cdot 10^5 yr$	$2.1 \cdot 10^{-8}$	$2.1 \cdot 10^{-8}$
⁴⁸ V	4	16 d	$4.1 \cdot 10^{-8}$	$5.4 \cdot 10^{-15}$
⁹⁶ Nb	3.2	1 d	$4.7 \cdot 10^{-3}$	≈ 0

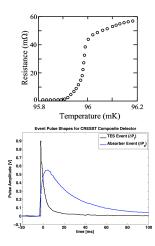
transition edge sensors

- Motivation is two-fold: good energy resolution improves efficiency in the region of interest + higher bandwidth operation enables $2\nu\beta\beta$ suppression
- superconducting films biased in transition region have exceptional sensitivity $\frac{\alpha_{\it TES}}{\omega_{\it semi}} \sim 100 \text{ and bandwidth}$ $\omega \sim G/C$
- voltage biased, electrothermal-feedback TES solve bias and thermal-runaway, but also requires the implementation of SQUID's as readout devices.



challenges applying to $0\nu\beta\beta$

- Require good coverage for light detector application
- sensor-absorber bandwidth mismatch is limiting factor in energy resolution
- low-T_c devices are also necessary because of operating temperature of the detector itself.
- TES dynamic range is limited by the width of the transition $\mathcal{O}(mK)$. $Q_{\beta\beta} \sim MeV$



Robeson 1 low-temperature laboratory



