

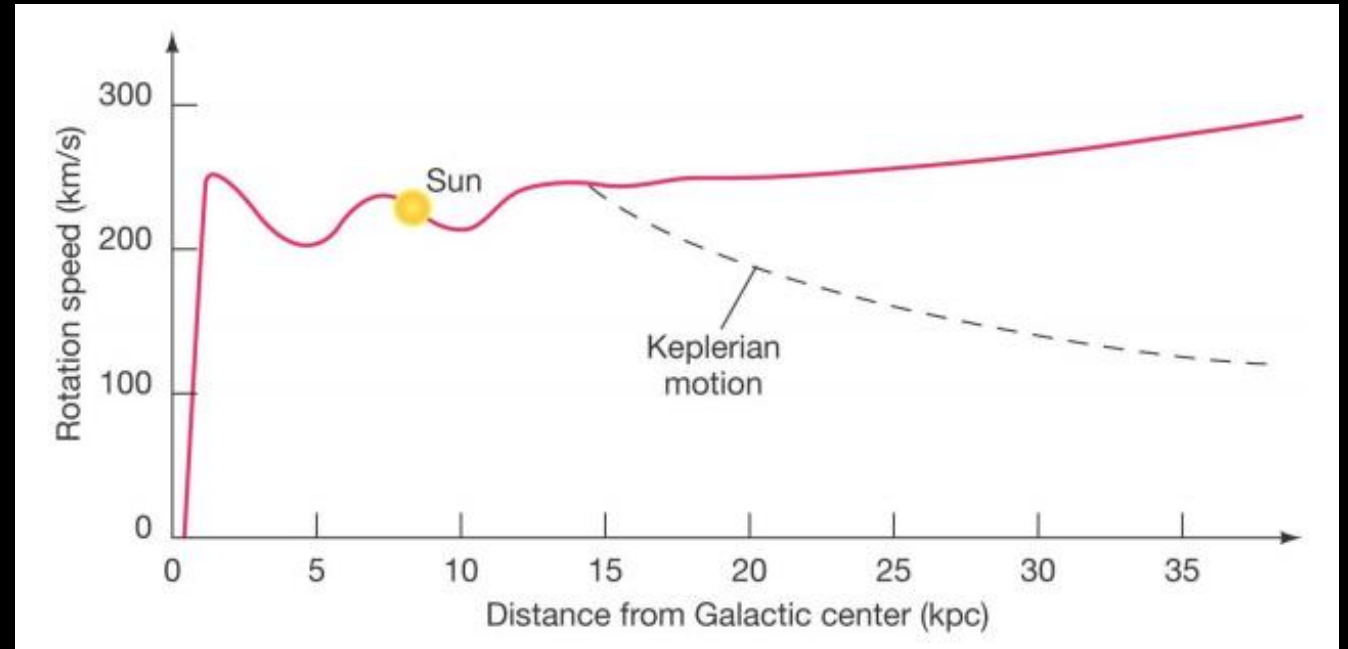
# Dark Matter

And the possible candidates for dark matter

Ashley Ernst

# Mass of the Milky Way Galaxy

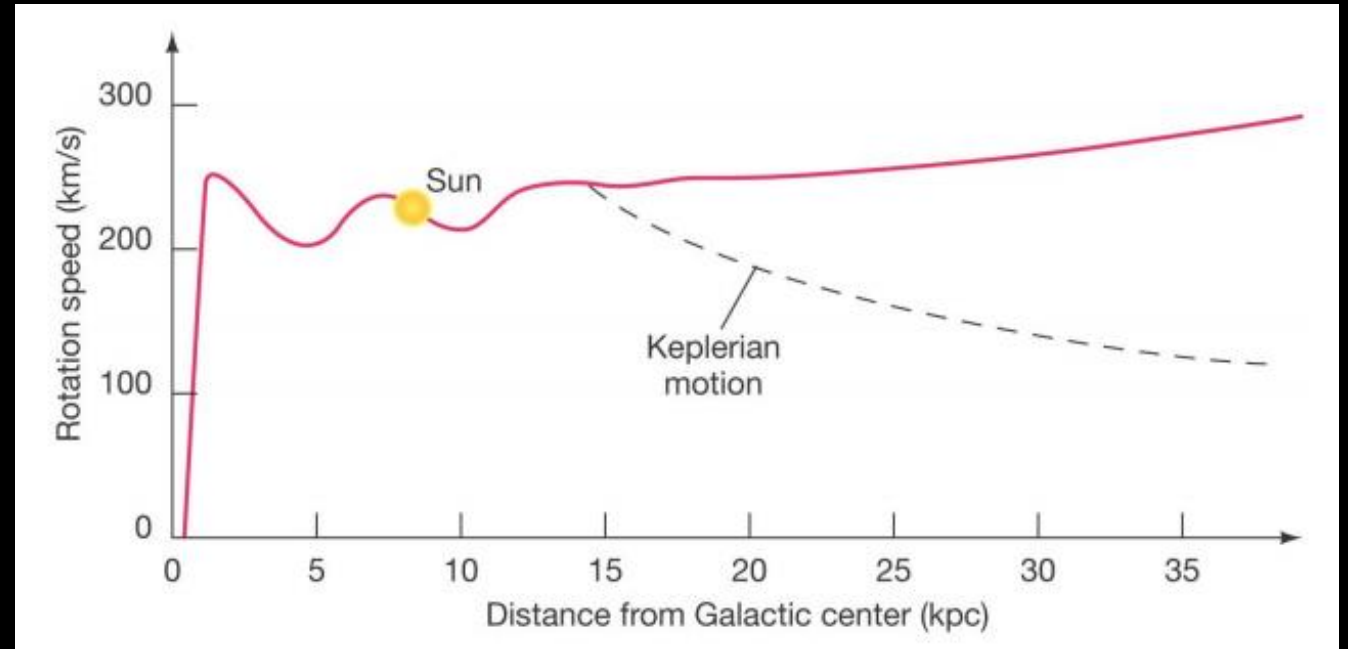
- Using Kepler's third law,
  - Total mass=(orbital size)<sup>3</sup>/(period)<sup>2</sup>and radio observations of gas in the Galactic disk, the rotation curve on the right can be plotted
- Radius of 15 kpc is the edge of the visible part of the Galaxy
- Rotation speed should fall off for distances larger than 15 kpc



Galaxy rotation curve showing the expected rotation speed (dashed line) and the observed rotation speed (red line) in the Milky Way Galaxy at various distances

# Mass of the Milky Way Galaxy

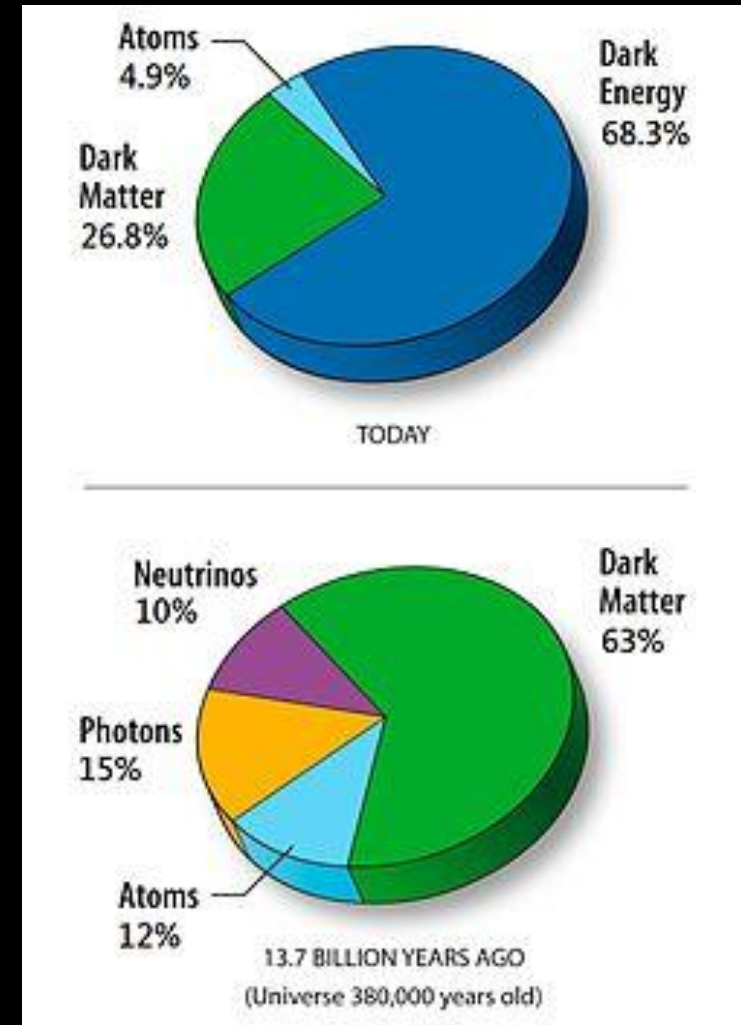
- Instead, the velocity rises at greater distances
- For 40 kpc, at least twice as much mass is outside of the visible part of the galaxy as inside
- An invisible dark halo must be surrounding the luminous portion of the galaxy
- Observed mass  $\ll$  expected mass
- Dark matter accounts for the difference



Galaxy rotation curve showing the expected rotation speed (dashed line) and the observed rotation speed (red line) in the Milky Way Galaxy at various distances

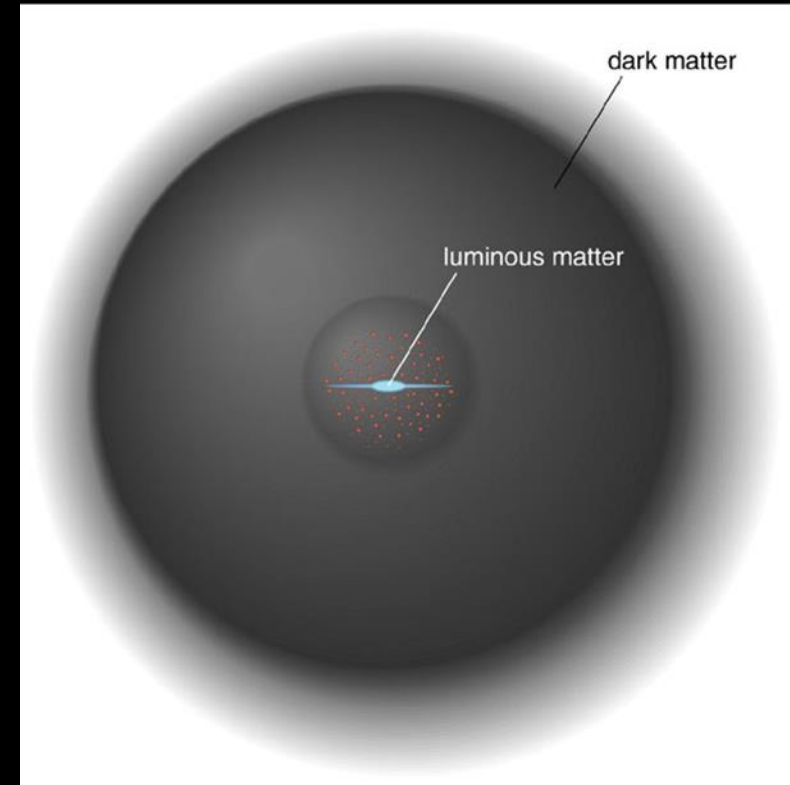
# What is Dark Matter?

- Composition of the universe
  - ~68% dark energy, ~27% dark matter, ~5% normal matter
- Not detectable at any wavelengths, from radio to gamma rays
- Inferred from gravitational pull on visible matter and from gravitational lensing of background radiation



# Classifications of Dark Matter

- Baryonic dark matter
- Non-baryonic dark matter
- Cold dark matter
- Hot dark matter



# Baryonic vs Non-Baryonic Dark Matter

- Baryonic dark matter
  - Astronomical bodies made of ordinary matter that emit little or no radiation
  - Massive Compact Halo Objects (MaCHOs)
  - Theory of Big Bang Nucleosynthesis shows that only a small fraction of dark matter
- Non-Baryonic Dark Matter
  - Thought to be the most prevalent form of dark matter
  - Weakly Interacting Massive Particles (WIMPs)

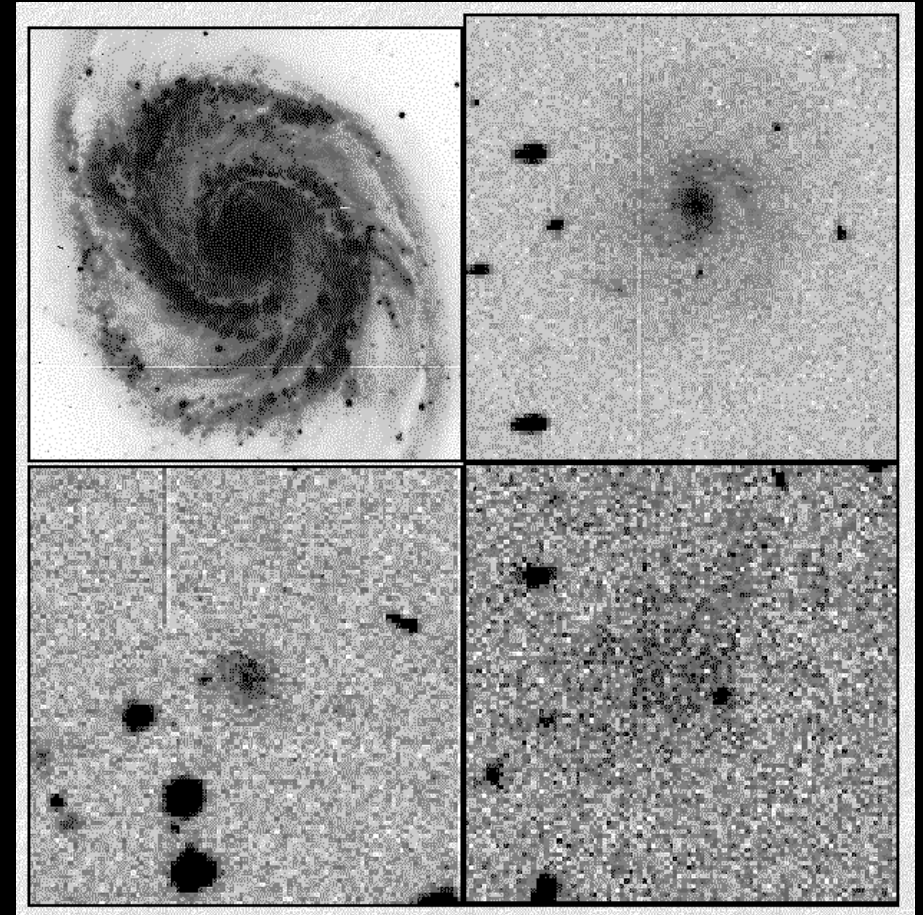
# Hot vs Cold Dark Matter

Hot and cold don't refer to temperature but refer to the velocity of the particle in the early universe

- Hot dark matter
  - Non baryonic particles moving with relativistic velocities (less than  $10^2 \text{ eV}/c^2$ )
  - Most common example are neutrinos
  - Not favored because of the process of the formation of the universe
- Cold dark matter
  - Slow moving exotic particles
  - WIMPs and axions
  - Have yet to be detected
  - Favored theory for dark matter

# MaCHOs

- Massive Compact Halo Objects
- Baryonic dark matter
- Hard to detect ordinary matter that emits little to no electromagnetic radiation
- Can be detected through gravitational microlensing
- Examples: Brown dwarfs, black holes, faint galaxies



High surface brightness galaxy compared to three low surface brightness galaxies



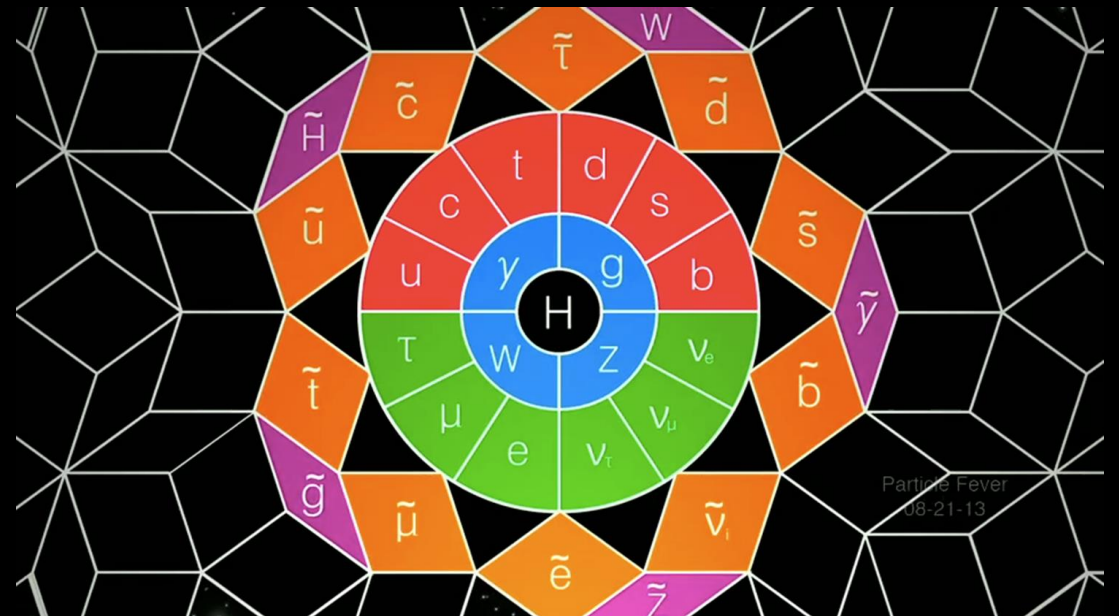
# WIMPs

- Weakly Interacting Massive Particles ( $10^{10}$ - $10^{12}$  eV/c<sup>2</sup>)
- Non-baryonic dark matter
- Interact only through gravity and the weak force
- Generally supersymmetric particles
- Main focus in particle astrophysics research but have yet to be detected
- Current experiments include SuperCDMS and LUX

	Fermions			Bosons	
Quarks	<i>u</i> up	<i>c</i> charm	<i>t</i> top	$\gamma$ photon	Force carriers
	<i>d</i> down	<i>s</i> strange	<i>b</i> bottom	<i>Z</i> Z boson	
Leptons	$\nu_e$ electron neutrino	$\nu_\mu$ muon neutrino	$\nu_\tau$ tau neutrino	<i>W</i> W boson	
	<i>e</i> electron	$\mu$ muon	$\tau$ tau	<i>g</i> gluon	

Typical standard model (left) and an example of a supersymmetric standard model (below)

Supersymmetric particles include neutralinos, gravitinos, photinos, higgsino, etc.



Particle Fever  
08-21-13

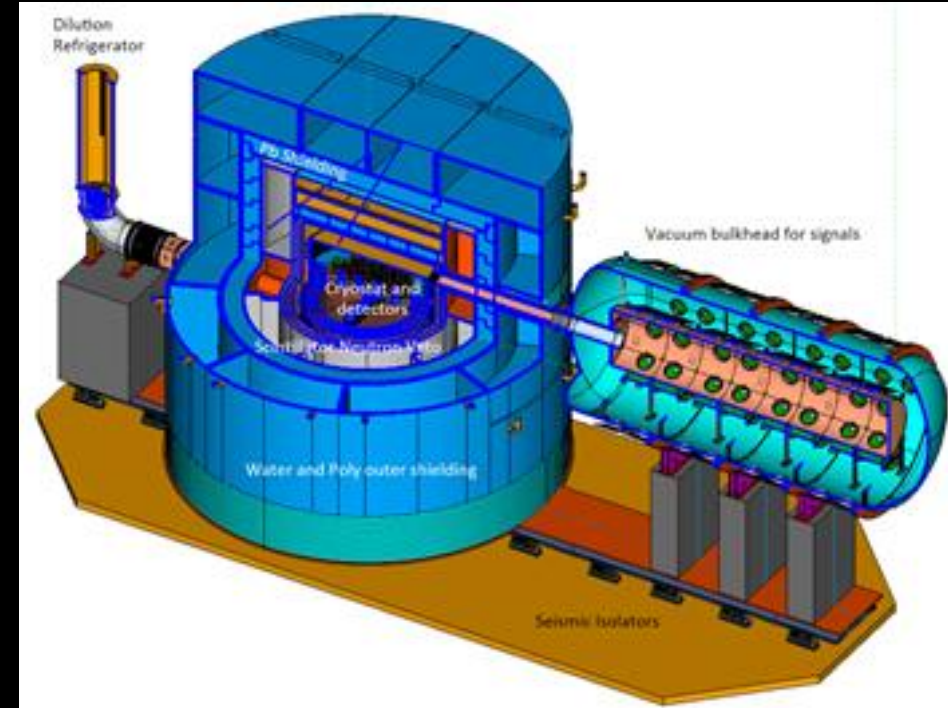
# Axions

- Non-baryonic cold dark matter
- Low mass boson ( $10^{-12}$  to  $10^3$  eV/c<sup>2</sup>)
- Believed to be the most numerous particle in existence if they exist
- Predicted by quantum chromodynamics
- Axion Dark Matter eXperiment (ADMX)



# Current Experiments

- SuperCDMS
  - Super Cryogenic Dark Matter Search
  - Experiment in search of WIMPs
  - Cryogenic germanium detectors deep underground at SNOLAB in Canada



# Current Experiments

- LUX
  - Large Underground Xenon dark matter experiment
  - In search of WIMPs
  - Liquid xenon chamber 1 mile underground in South Dakota



# References

- Astronomy Today Textbook
- B.O.B.
- <http://science.nasa.gov/astrophysics/focus-areas/what-is-dark-energy/>
- <http://cdms.berkeley.edu/experiment.html>
- [http://www2.astro.psu.edu/~caryl/a480/lecture24\\_10.pdf](http://www2.astro.psu.edu/~caryl/a480/lecture24_10.pdf)
- <http://arxiv.org/pdf/hep-ph/0002126.pdf>
- <http://atropos.as.arizona.edu/aiz/teaching/nats102/lecture23.html>

# Questions

- Which seems more reasonable, WIMPs or MaCHOs?
- Could dark matter be the product of 'massless' particles that have mass much smaller than scientists can measure?
  - My research advisor at Argonne suggested looking into the possibility that photons having negligible mass could account for dark matter