

Exploring the cosmic infinite horizons of modern

Astronomy

Today.

By J. Pinkney
Ohio Northern University

Goals for this course

- 1) Obtain knowledge about astronomy.**
- 2) Obtain understanding of some basic physics concepts.**
- 3) Improve science “skills”: use math for solving problems, communicating science, observing, critical thinking.**
- 4) Learn about science and how it differs from pseudosciences and other belief systems.**
- 5) Awareness of science in current events - “astro news”.**
- 6) Expand your personal “theory of everything” - your cosmology.**

Week 1 of Stellar and Galactic Astronomy

**View the film “*Powers of 10 A Film About the Relative Sizes of Things*
1977**

by Charles and Ray Eames

Narrated by Phillip Morrison

Pre-Questions

- 1) Note the largest scale achieved.**
- 2) Note the smallest scale achieved.**
- 3) In which powers of 10 do we find great “voids” where nothing new enters the view?**

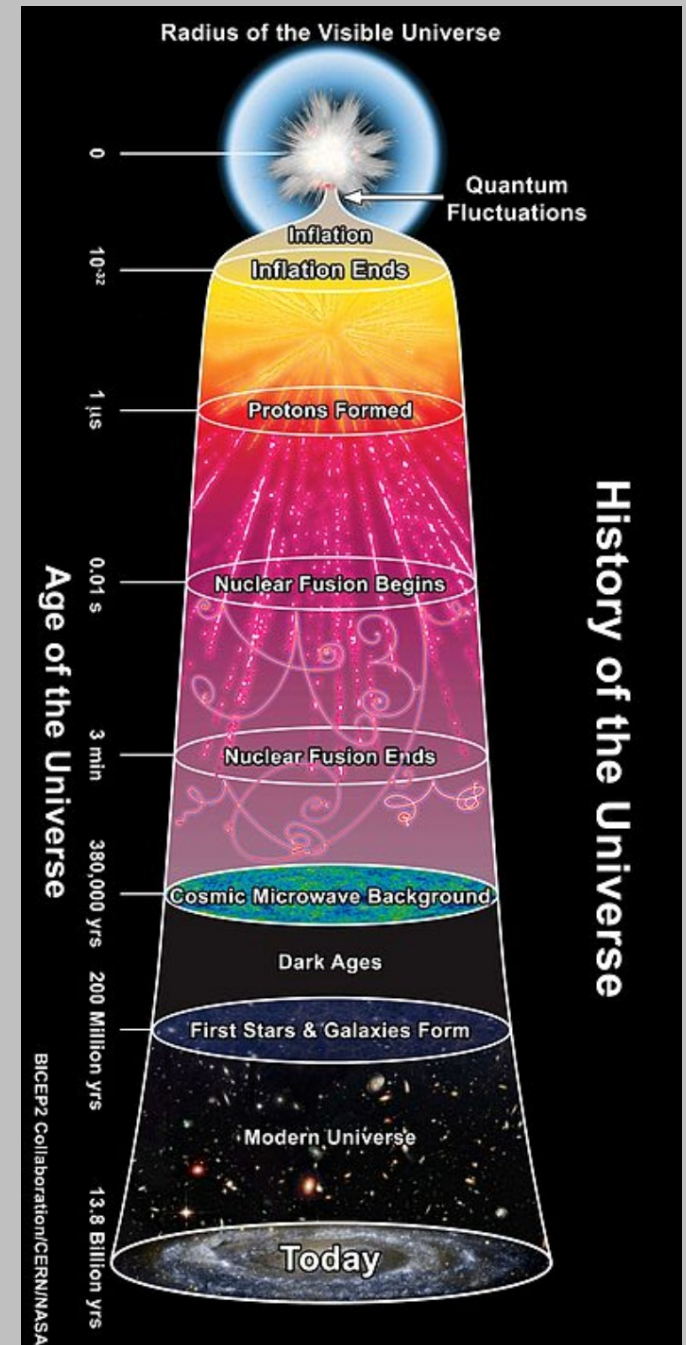
Also consider ...

- 4) In every 10 seconds, the view expands by a factor of 10. Could this entire journey be experienced while flying in a rocket ship?**

Why do we need powers of 10?

Many quantities have a vast range of values in Astronomy ...

- 1) *Distances.* 10^{-16} to 10^{24} meters (actually $>10^{26}$) for the scale of an atomic nucleus compared to the scale of largest structures in the universe.
- 2) *Time.* 10^{-46} second to 10^{16} seconds (10^9 yrs) for the Planck time to the age of the universe.
- 3) *Masses:* 10^{-31} kg (electrons) to 10^{45} kg (clusters of galaxies)
- 4) *Energies:* and 10^{-19} Joules for H-alpha photon to 10^{39} Joules for Gamma-Ray Bursts.
- 5) *Speeds:* continental drift (cm/yr) to the speed of light 300,000,000 m/s.



And that's why we use “powers of 10” -- to make them more manageable!

Understanding Powers of 10, orders of magnitude, and Scientific Notation

Scientific Notation: a way of writing a number in which the decimal point is placed to the right of the most significant digit, and this is multiplied by 10^P where P =an integer (the exponent, or “power of 10”)

Exponential Notation Format: Coefficient X Base^{exponent}
(where Base=10)

Example: $58400 = 5.84 \times 10^4$

Example: $0.01093 = 1.093 \times 10^{-2}$

Example: The average Earth-Sun distance is 93,000,000 miles or 9.3×10^7 miles.

Power of 10: one can approximate a number by giving only the exponent of that number expressed in scientific notation, rounded up or down depending on the coefficient.

Example: $5.84 \times 10^4 = 10^{4.7664} \sim 10^5 = 10 \times 10 \times 10 \times 10 \times 10 = 100,000$.

Example of usage: The distance to the Sun from the Earth is about 10^8 miles. Thus, the Earth-Sun distance is 8 powers of 10 greater than a mile.

Order of magnitude: the “order of magnitude” of a number is the same thing as a number's “power of ten”, it is just used differently in sentences.

Example: “The Earth-Sun distance is 8 orders of magnitude larger than a mile.”)

Example: “If you thought the US population was 3 million, you were off by 2 orders of magnitude.”

Understanding Powers of 10, orders of magnitude, and Scientific Notation

Rounding to the nearest power of 10.

Previous example: $5.84 \times 10^4 = 10^{4.7664} \sim 10^5$. But what if we had ...

Example: $4.84 \times 10^4 = 10^{4.6848} \sim 10^5$.

Example: $3.84 \times 10^4 = 10^{4.5843} \sim 10^5$. Perhaps if the exponent dropped below 4.5 ...

Example: $2.84 \times 10^4 = 10^{4.4533} \sim 10^4$. Finally, we don't round up!

For which coefficient will the exponent be exactly 4.5? Answer: $3.162278 (= \sqrt{10})$

Example: $3.1623 \times 10^4 \sim 10^5$.

Example: $3.1622 \times 10^4 \sim 10^4$.

Try these:

Example: $9.99 \times 10^2 \sim 10^3$

Example: $9.9 \times 10^{-2} \sim 10^{-1}$.

Example: $5.1 \times 10^{-4} \sim 10^{-3}$.

Example: $3.10 \times 10^6 \sim 10^6$.

Example: $3.20 \times 10^9 \sim 10^{10}$.

Example: 401,000 $\sim 10^6$.

Example: 301,000 $\sim 10^5$.

Example: 73,162,055,319 $\sim 10^{11}$.

Why do we need these new ways of expressing numbers?

1) To compress long numbers.

Example: mass of the Sun in kilograms:

$$1 M_{\odot} = 2.0 \times 10^{30} \text{ kg (sci notation)}$$

$$1 M_{\odot} = 10^{30} \text{ kg (nearest power of 10)}$$

(Now try writing this number as a 1 with 30 zeros!)

2) To simplify multiplication and division.

Example: if the Earth's mass is 6×10^{24} kg, how many Earth masses go into the Sun?

$$M_{\odot}/M_{\oplus} = 10^{30} \div 10^{25} = 10^{30-25} = 10^5 \quad (\text{actual } 3.35 \times 10^5)$$

Example: if there are 86400 seconds per day, and 365 days in a year, roughly how many seconds are in a year?

$$10^5 * 10^2 = 10^{5+2} = 10^7 \quad (\text{actual } 3.156 \times 10^7)$$

Simplified multiplication and division allows easy rough estimates called “order of magnitude calculations” or “back of the envelope calculations”.

Summary:

With powers of 10 math, division becomes subtraction and multiplication becomes addition.

GROUP EXERCISE

Small to
large -
guess

Rank these lengths from smallest to largest. First guess using left column, then use data provided by the instructor to fill in right column.

Sizes
smallest
to largest

1. <u> H </u>	A. Black Hole, 1 solar mass, (Schwarzschild radius)	<u> </u>	1. <u> H </u>
2. <u> </u>	B. Cosmic Microwave Background (distance)	<u> </u>	2. <u> </u>
3. <u> </u>	C. Distance between stars in Sun's neighborhood	<u> </u>	3. <u> </u>
4. <u> </u>	D. Earth (diam)	<u> </u>	4. <u> </u>
5. <u> </u>	E. Galaxies, Dwarf Ellipticals (diam)	<u> </u>	5. <u> </u>
6. <u> </u>	F. Galaxies, Giant Ellipticals (diam)	<u> </u>	6. <u> </u>
7. <u> </u>	G. Galaxies, Milky Way (diam)	<u> </u>	7. <u> </u>
8. <u> </u>	H. Human (diam)	<u> </u>	8. <u> </u>
9. <u> </u>	I. Jupiter (diam)	<u> </u>	9. <u> </u>
10. <u> </u>	J. Local Group (diam)	<u> </u>	10. <u> </u>
11. <u> </u>	K. Moon (diam)	<u> </u>	11. <u> </u>
12. <u> </u>	L. Moon (distance)	<u> </u>	12. <u> </u>
13. <u> </u>	M. Neptune (dist from Sun. size of sol sys)	<u> </u>	13. <u> </u>
14. <u> </u>	N. neutron star (diam)	<u> </u>	14. <u> </u>
15. <u> </u>	O. Rich clusters of galaxies (diam)	<u> </u>	15. <u> </u>
16. <u> </u>	P. Star Clusters, globular (diam)	<u> </u>	16. <u> </u>
17. <u> </u>	Q. Sun (diam)	<u> </u>	17. <u> </u>
18. <u> </u>	R. Sun (distance)	<u> </u>	18. <u> </u>
19. <u> </u>	S. Superclusters of galaxies (length)	<u> </u>	19. <u> </u>
20. <u> </u>	T. White dwarf	<u> </u>	20. <u> </u>

GROUP EXERCISE**NAME:** _____**Rank these lengths from smallest to largest. (Use lengths in meters.)**

	Sizes	smallest to largest
A. Black Hole, 1 solar mass, (Schwarzschild radius)	$3 \times 10^3 \text{ m}$	1. <u>H</u>
B. Cosmic Microwave Background (distance)	10^{26} m	2. _____
C. Distance between stars in Sun's neighborhood	10^{16} m	3. _____
D. Earth (diam)	10^7 m	4. _____
E. Galaxies, Dwarf Ellipticals (diam)	10^{19} m	5. _____
F. Galaxies, Giant Ellipticals (diam)	10^{22} m	6. _____
G. Galaxies, Milky Way (diam)	10^{21} m	7. _____
H. Human (diam)	10^0 m	8. _____
I. Jupiter (diam)	10^8 m	9. _____
J. Local Group (diam)	10^{22} m	10. _____
K. Moon (diam)	10^6 m	11. _____
L. Moon (distance)	10^8 m	12. _____
M. Neptune (dist from Sun. size of sol sys)	10^{13} m	13. _____
N. neutron star (diam)	10^4 m	14. _____
O. Rich clusters of galaxies (diam)	10^{23} m	15. _____
P. Star Clusters, globular (diam)	10^{17} m	16. _____
Q. Sun (diam)	10^9 m	17. _____
R. Sun (distance)	10^{11} m	18. _____
S. Superclusters of galaxies (length)	10^{24} m	19. _____
T. White dwarf	10^7 m	20. _____

Other ways to make large numbers manageable

1) Use prefixes

small: deci, centi, milli, micro, nano, pico, femto

10 to the: -1 -2 -3 -6 -9 -12 -15

large: deca, hecto, kilo, mega, giga, tera, peta, exa

10 to the: 1 2 3 6 9 12 15 18

Example) What is a convenient unit for 10^{-6} seconds? Ans: a microsecond ($1 \mu\text{s}$).

2) Invent new units

In astronomy we have ... (**red** ones are new units)

a) The “solar mass”, $1 M_{\odot} = 2 \times 10^{30} \text{ kg}$

b) The “**astronomical unit**”, $1 \text{ AU} = 1.5 \times 10^8 \text{ km}$, 93,000,000 miles.

The average distance between the Earth and Sun.

c) The **Light year**, $1 \text{ LY} = 9.5 \times 10^{12} \text{ km}$

The distance light travels through space in a year.

Good for distances between stars.

d) The **parsec**, $1 \text{ pc} = 3.1 \times 10^{13} \text{ km}$.

The distance one must be from the Solar system so that the Earth-Sun separation appears to be 1 arcsecond.

Good for distances between stars.

e) The **kiloparsec**, $1 \text{ kpc} = 1000 \text{ pc}$

Good for distances inside a galaxy

f) The **megaparsec**, $1 \text{ Mpc} = 1000,000 \text{ pc}$

Good for distances between galaxies, clusters, superclusters.

Other ways to make large numbers manageable

1) Use prefixes

NUMBER PREFIXES

<i>Multiple</i>	<i>Prefix</i>	<i>Symbol</i>	<i>Common Name</i>
$(10^{100})^{100}$	anton	A	antonplex
$(10^{10})^{10}$	-	-	googolplex
10^{100}	-	-	googol
10^{24}	yotta	Y	heptillion
10^{21}	zetta	Z	hexillion
10^{18}	exa	E	quintillion
10^{15}	peta	P	quadrillion
10^{12}	tera	T	trillion
10^9	giga	G	billion
10^6	mega	M	million
10^3	kilo	k	thousand
10^2	hecto	h	hundred
10^1	deca	da	ten
10^{-1}	deci	d	tenth
10^{-2}	centi	c	hundreth
10^{-3}	milli	m	thousandth
10^{-6}	micro	(Greek mu)	millionth
10^{-9}	nano	n	billionth
10^{-12}	pico	p	trillionth
10^{-15}	femto	f	quadrillionth
10^{-18}	atto	a	quintillionth
10^{-21}	zepto	z	hexillionth
10^{-24}	yocto	y	heptillionth

Side notes on astronomical distances / sizes

1) The magic number, 110.

110 (roughly) comes up many times in distance ratios.

$$110 = \text{DiamSun}/\text{DiamE} = \text{distSun}/\text{DiamSun} = \text{distMoon}/\text{DiamMoon}$$

2) The AU and Light Year.

There are 63,000 AU in 1 LY.

Coincidentally, there are 63,000 inches in a mile!

3) The distance to the Moon is 240,000 miles. A good car typically lasts about 240,000 miles. So you *might* be able to drive to the Moon if there were a direct route from Earth!

4) The ratio 400.

$$400 = \text{DistSun}/\text{distMoon} = \text{DiamSun}/\text{diamMoon}$$

Because of this coincidence the Sun and Moon subtend about the same angle in the sky ($\frac{1}{2}$ degree) and we can observe both total and annular solar eclipses.

The Naked – Eye Universe

Astronomical things that we can see with the naked eye

1. Sun
2. Moon
3. 5 planets (+Uranus, just visible)
4. 6500 stars (down to +6.0 mag)
5. 3 galaxies (M31,LMC,SMC. Some can see M33)
6. Comets
7. Supernovae, novae
8. Meteors (in our atmosphere)
9. Aurora (in our atmosphere)

The Naked – Eye Universe

The Top Ten Brightest Objects in the Sky

1. **Sun**
2. **Moon**
3. **Venus**
4. **Mars**
5. **Jupiter**
6. **Mercury**
7. **Sirius**
8. **Saturn**
9. **Canopus** (in Carina, Southern Hem)
10. **Arcturus** (Bootes)

Rigil Kent (Alpha Cen), **Vega**, and **Capella** are almost a tie for 11th!

The Naked – Eye Universe

Constellations and Asterisms

Constellation: a designated region in the sky containing one or more historical star patterns.

Examples) Orion, Taurus, Ursa Major (see next slides)

Asterism: a recognizable pattern of stars.

Ex) Orion, the hunter

Ex) Taurus, the bull; the Pleiades; the Hyades

Ex) Ursa Major (the great bear); the Big Dipper; La Cassarole

Ex) The Summer Triangle

Ex) The Coathanger (Brocchi's cluster)

*** 88 total constellations**

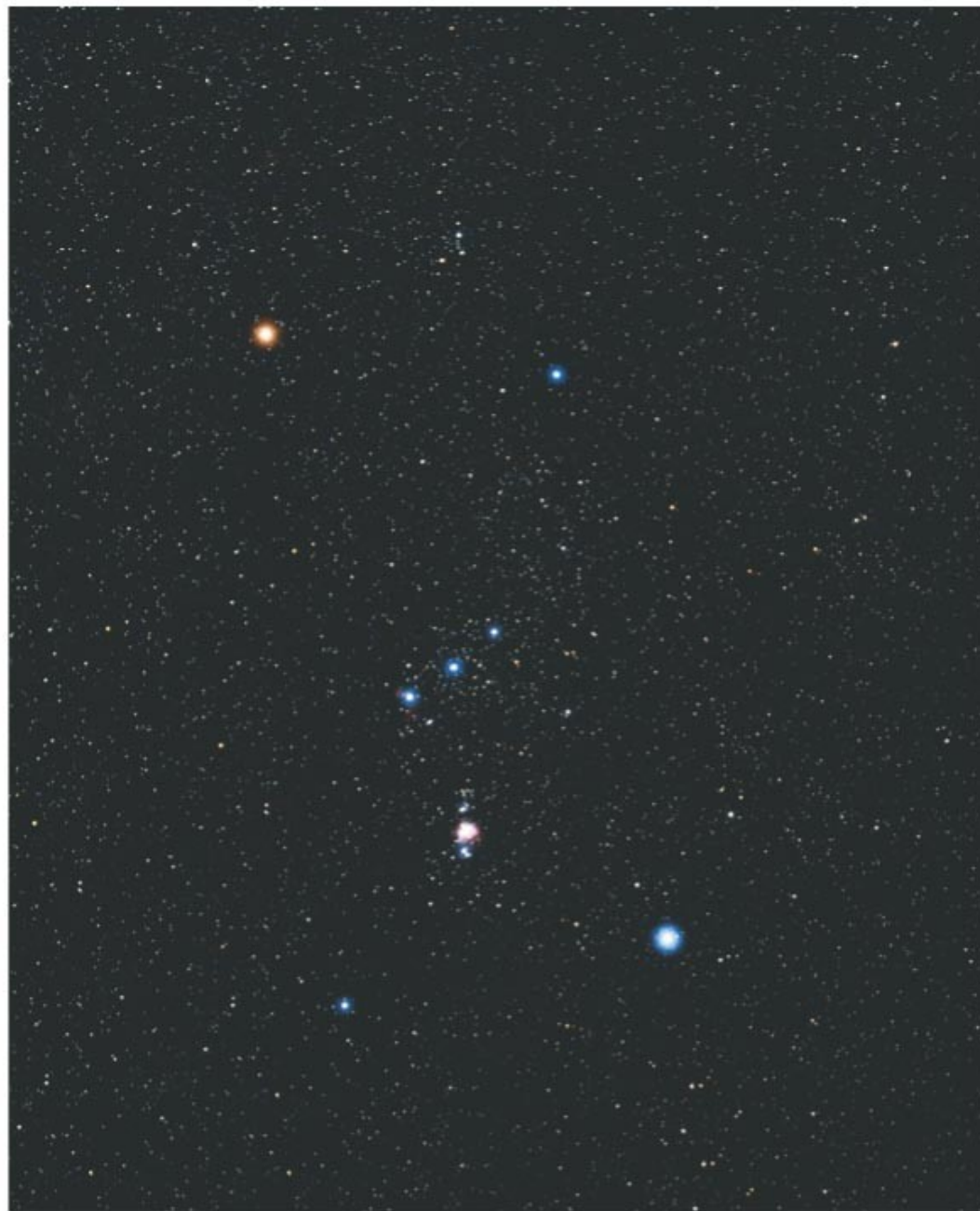
*** More than 88 asterisms**

*** Northern constellations named after Greek Mythological characters**

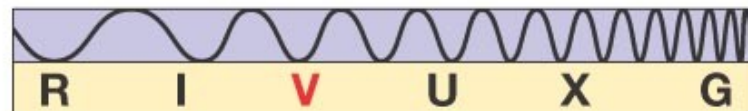
**Example:
Orion.**

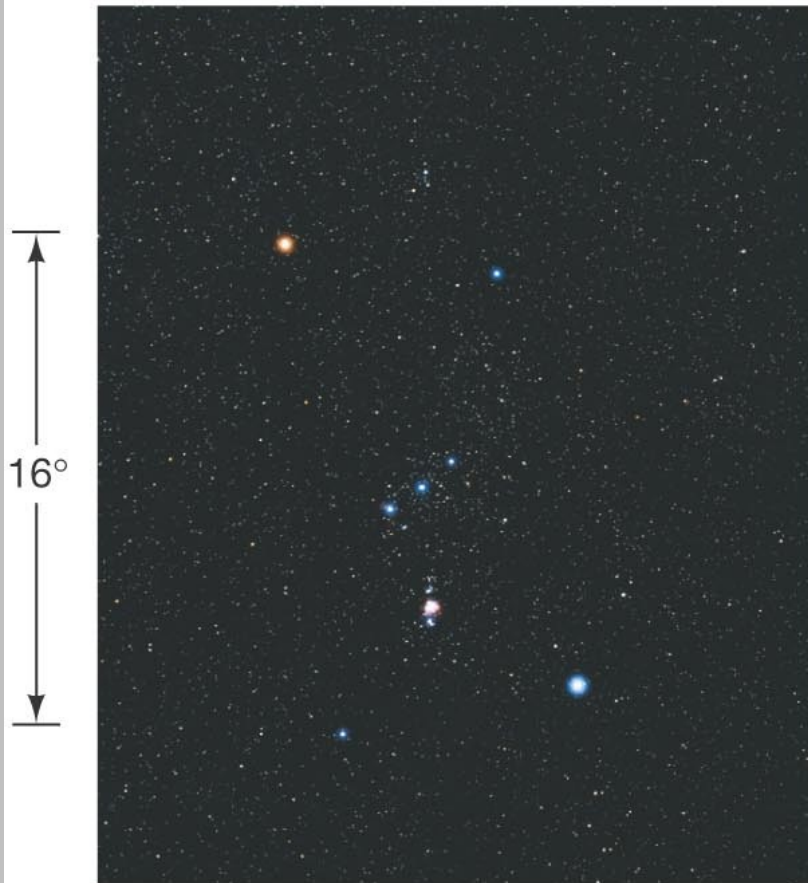
**An easily
recognized
constellation!**

↑
16°
↓

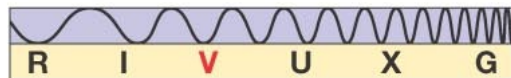


(a)

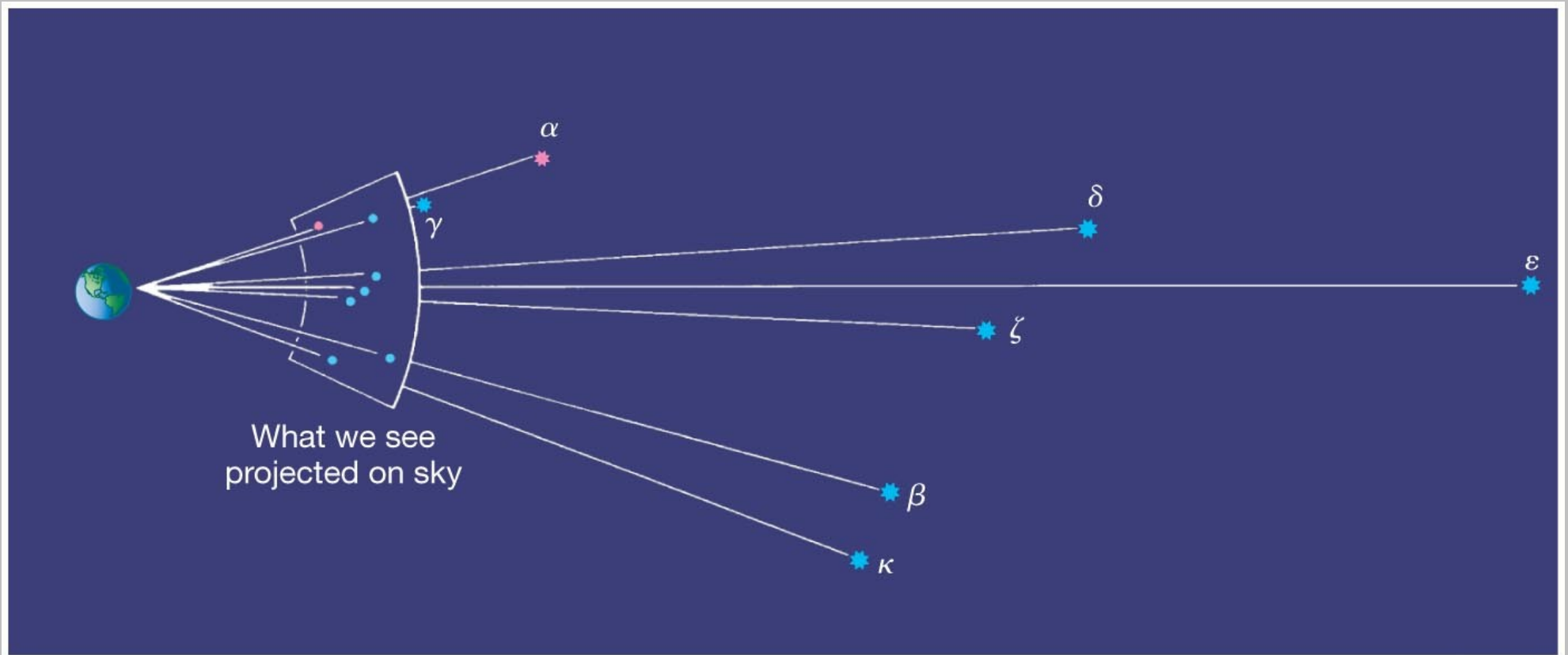


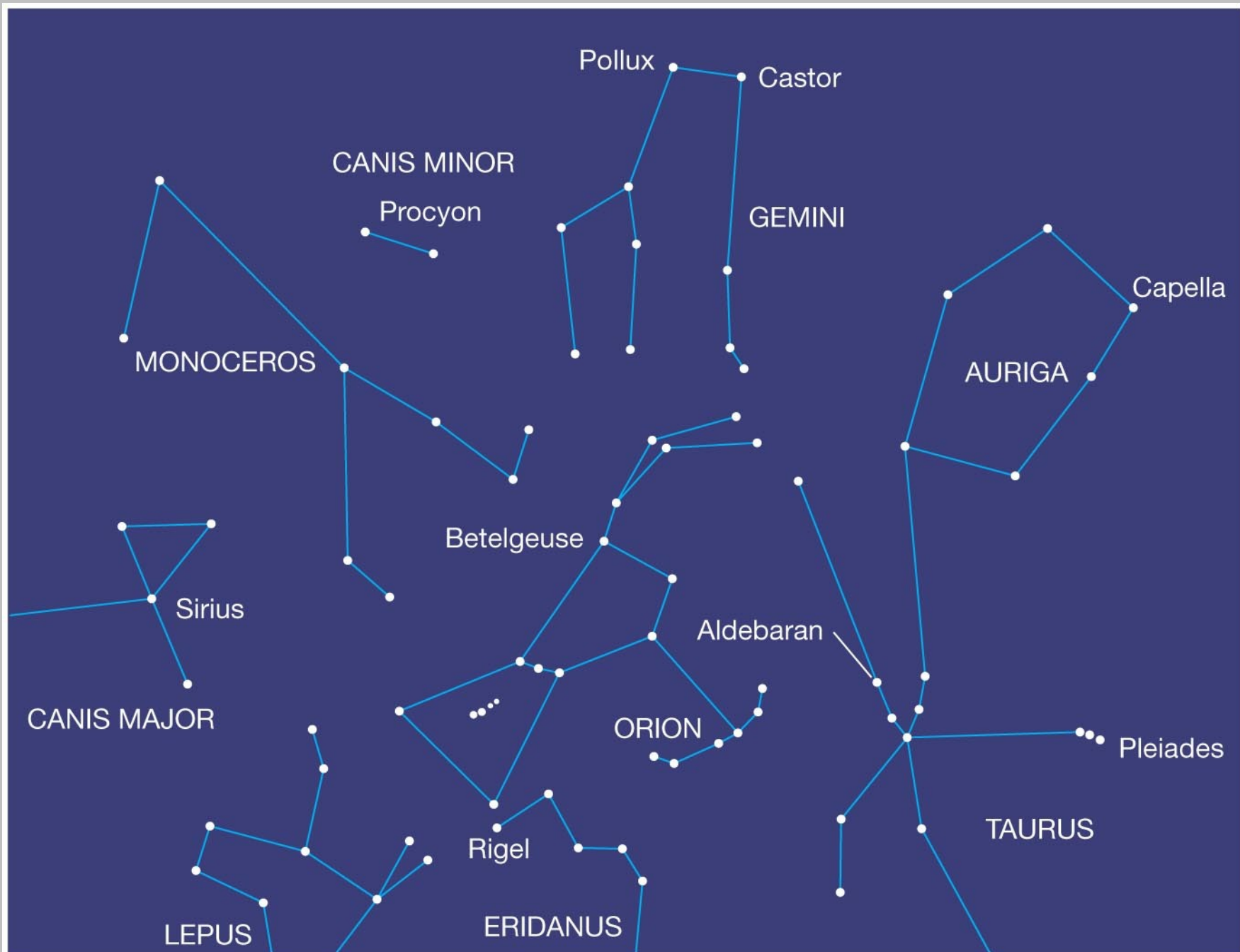


(a)



(b)





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Try a planetarium program like “Stellarium” or “Celestia” to see the sky in motion.