

# The Distance Ladder I.

## The Milky Way Galaxy



(Ch. 23)

# Units of Chapter 23

## 1. Our Milky Way Galaxy

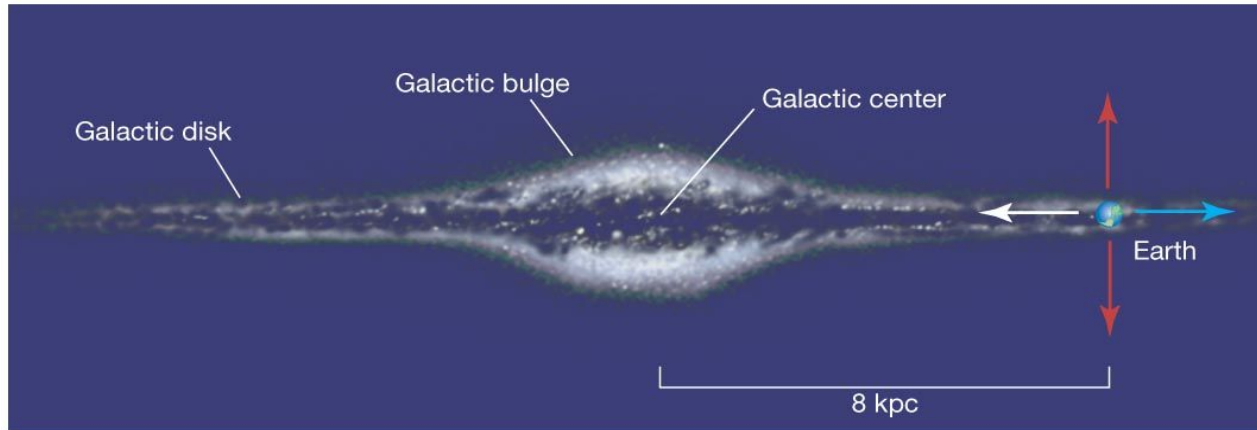
- a) Dimensions and structure
- b) Spiral Arms
- c) Mass and Dark Matter
- d) Nucleus

## 2. Distances within the Milky Way

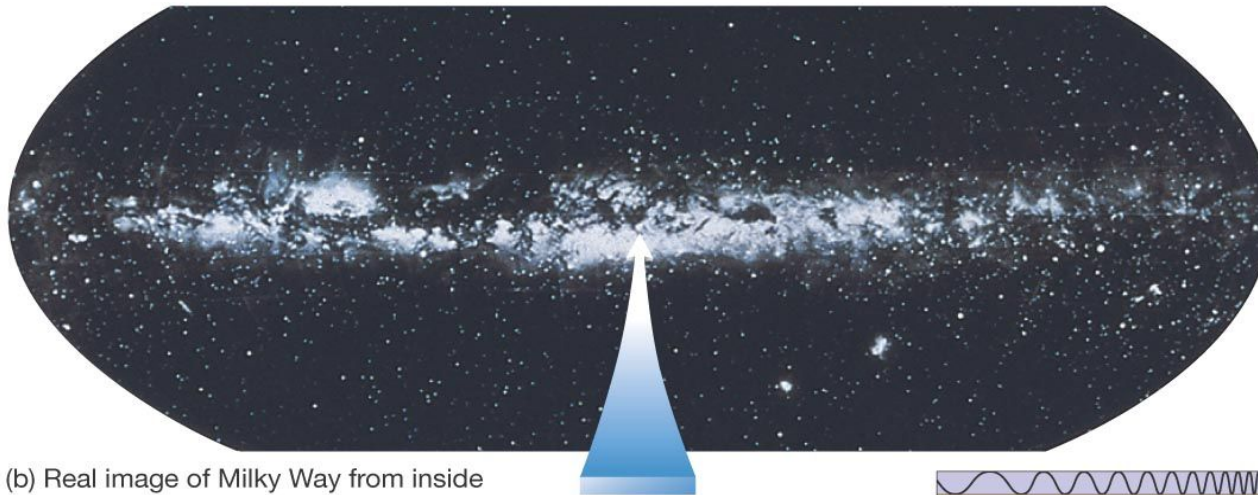
- a) Stellar and spectroscopic parallax
- b) “Standard Candles” or “Beasts of a kind” concept
- c) Herschel's star counts
- d) “Intrinsic” Variable Stars
- e) Other Distance Indicators

# 23.1 Our Parent Galaxy

**Milky Way** is what our galaxy appears as in the night sky.



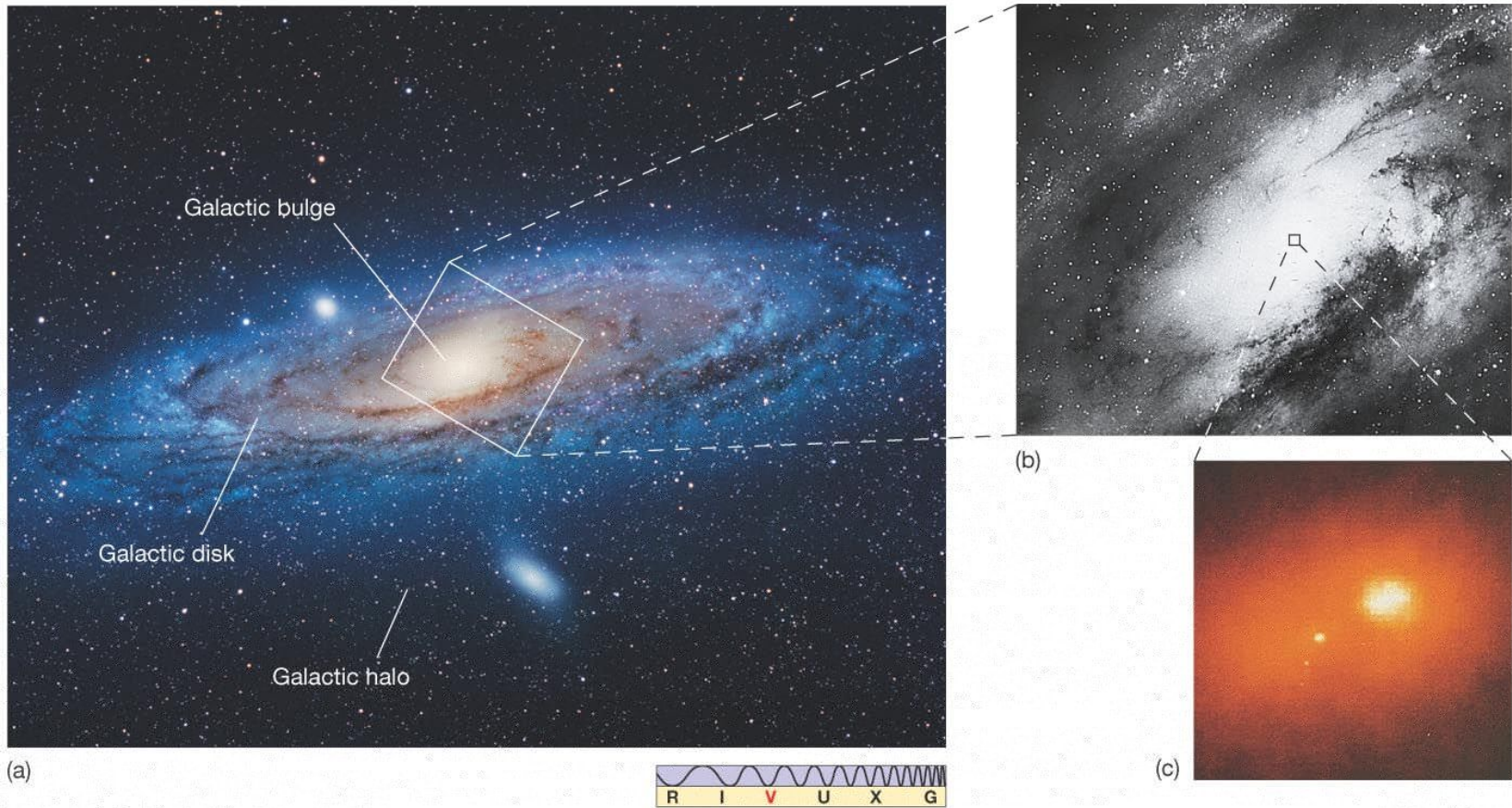
(a) Artist's view of Milky Way from afar



(b) Real image of Milky Way from inside

# 23.1 Our Parent Galaxy

Our galaxy is a **spiral** galaxy. The Andromeda Galaxy, our closest spiral neighbor, probably resembles the Milky Way fairly closely.















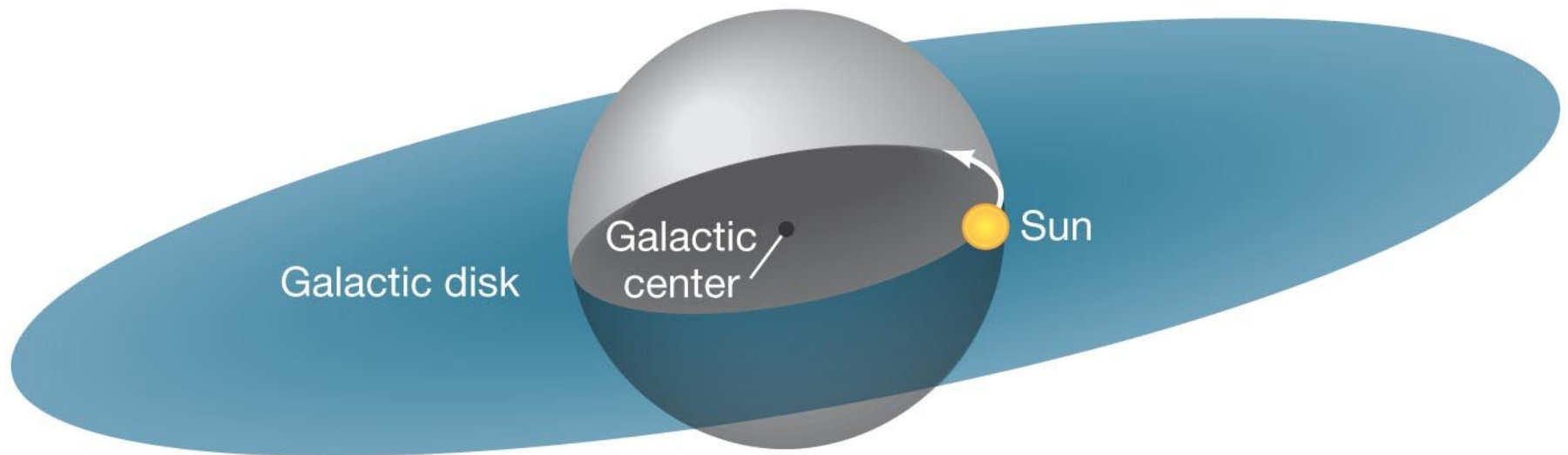






## 23.6 The Mass of the Milky Way Galaxy

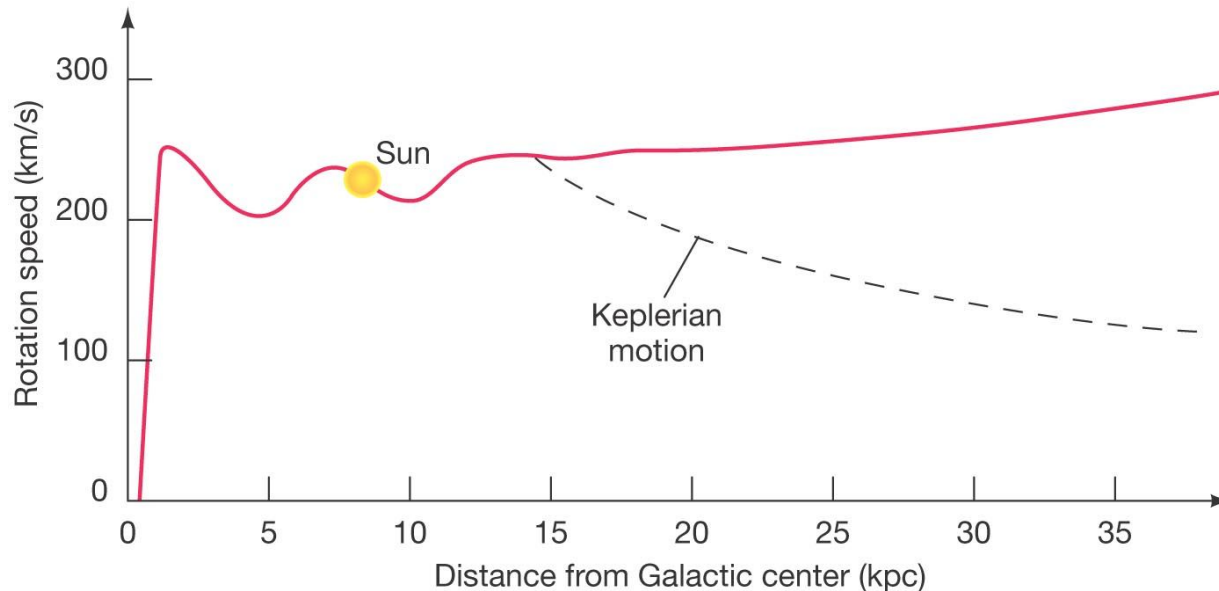
The **orbital speed** of an object depends only on the **amount of mass** between it and the **galactic center**:



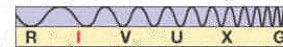
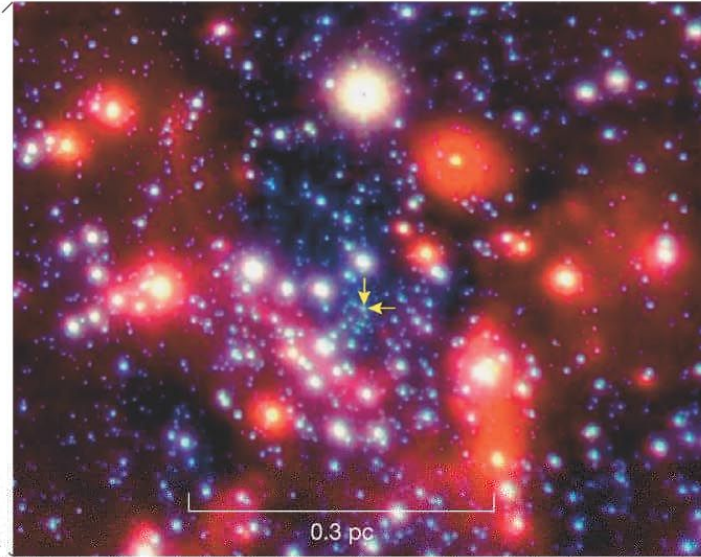
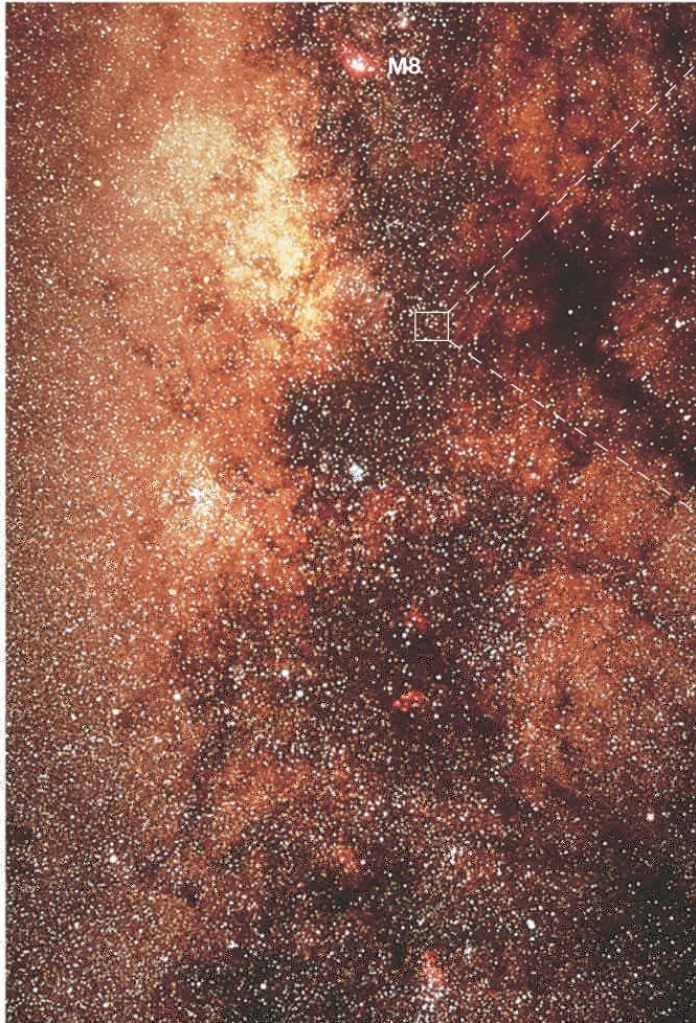
# 23.6 The Mass of the Milky Way Galaxy

Once all the galaxy is within an orbit, the velocity should **diminish** with distance, as the dashed curve shows.

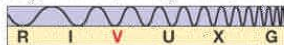
It doesn't; more than **twice** the mass of the galaxy would have to be outside the visible part to reproduce the observed curve.



# 23.7 The Galactic Center



**Two views toward the galactic center.**











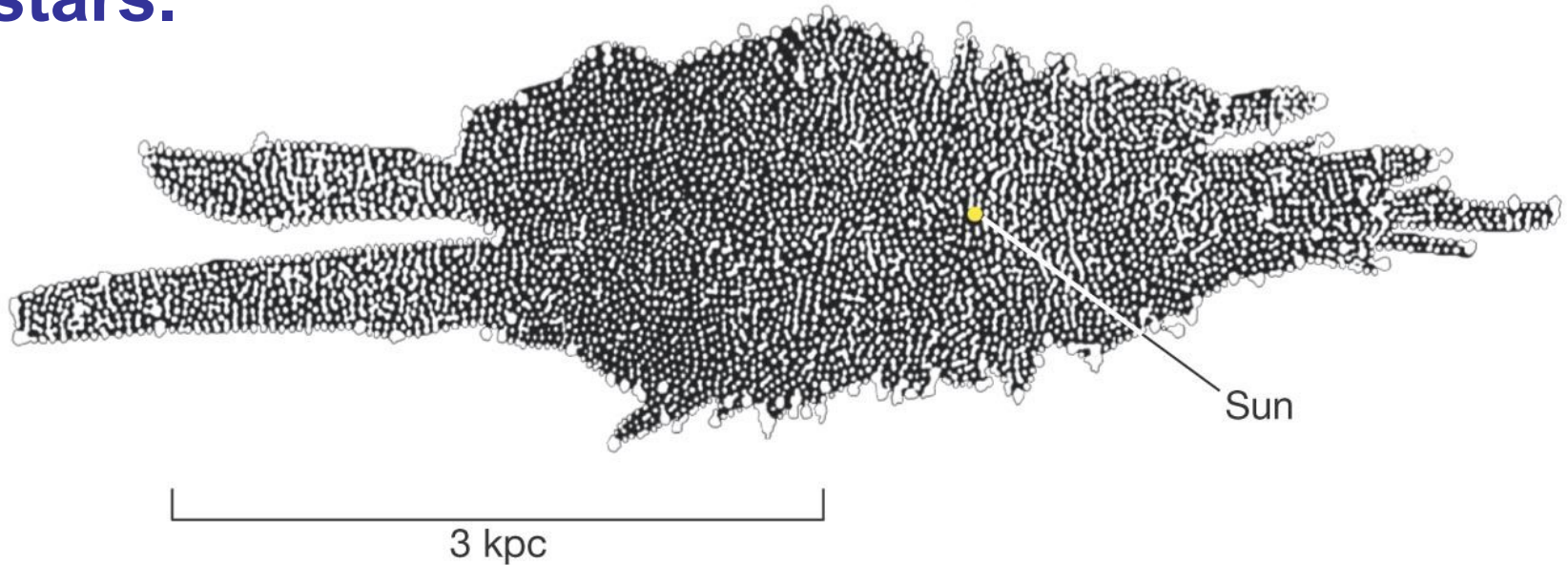






## 23.2 Measuring the Milky Way

One of the first attempts to **measure** the Milky Way was done by W. Herschel using visible stars.

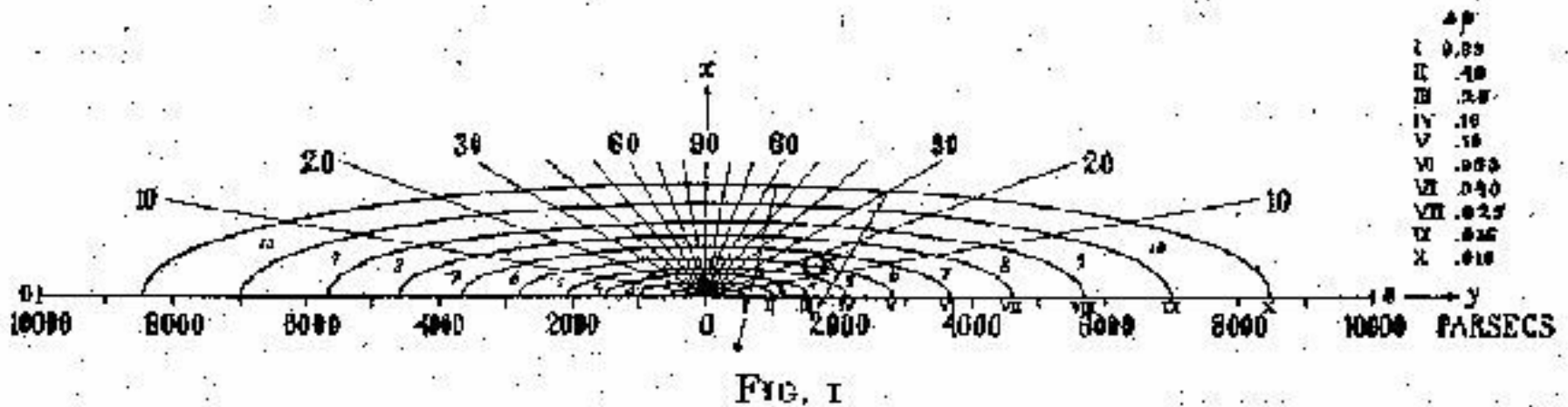


### Problems:

1. patchy dust blocked view! (extinction)
2. all stars do not have the same luminosity!!

# 23.2 Measuring the Milky Way

A model based on star brightnesses, types, & motions was done by Jacobus Kapteyn (1850-1922).



Still lacked corrections for extinction.

Sun 2000 LY from center.

Kapteyn Universe (circa 1920)

6,500  
light years



solar system



30,000 light years



## 23.2 Variable stars & distances

Extrinsic variables: eclipsing binaries

Cataclysmic variables: novae, supernovae.

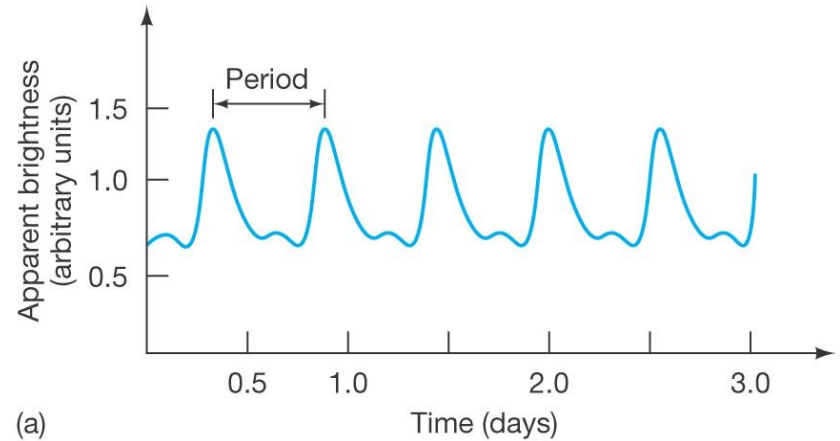
**“Intrinsic variables”** - pulsating regularly:  
**RR Lyrae stars and Cepheids.** - very good  
for distances!

Long period, semi-regular variables (like Mira) –  
not good for distances

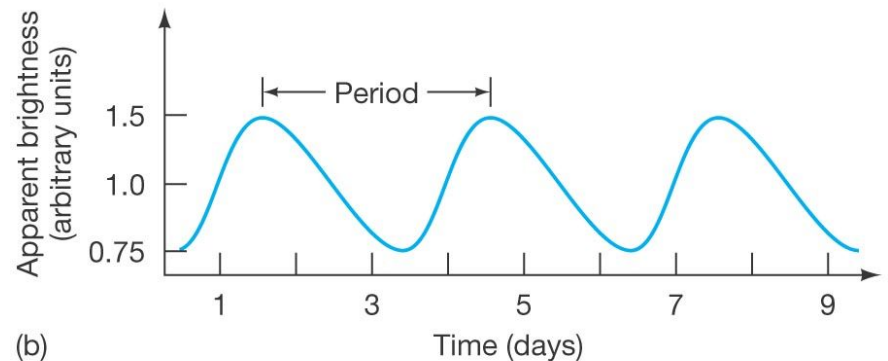


## 23.2 Variable stars & distances

The upper plot is an **RR Lyrae** star. All such stars have essentially the same luminosity curve with periods from **0.5 to 1 day**.



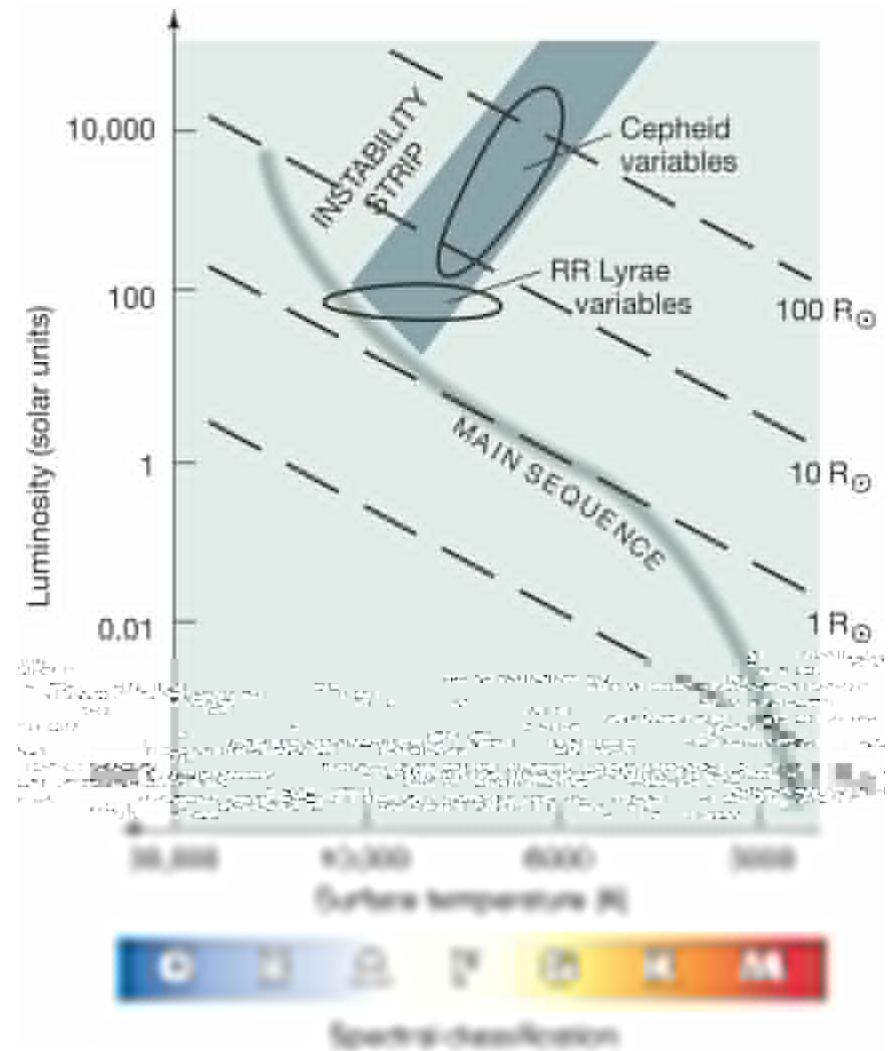
The lower plot is a **Cepheid** variable; Cepheid periods range from about **1 to 100 days**.



## 23.2 Variable stars & distances

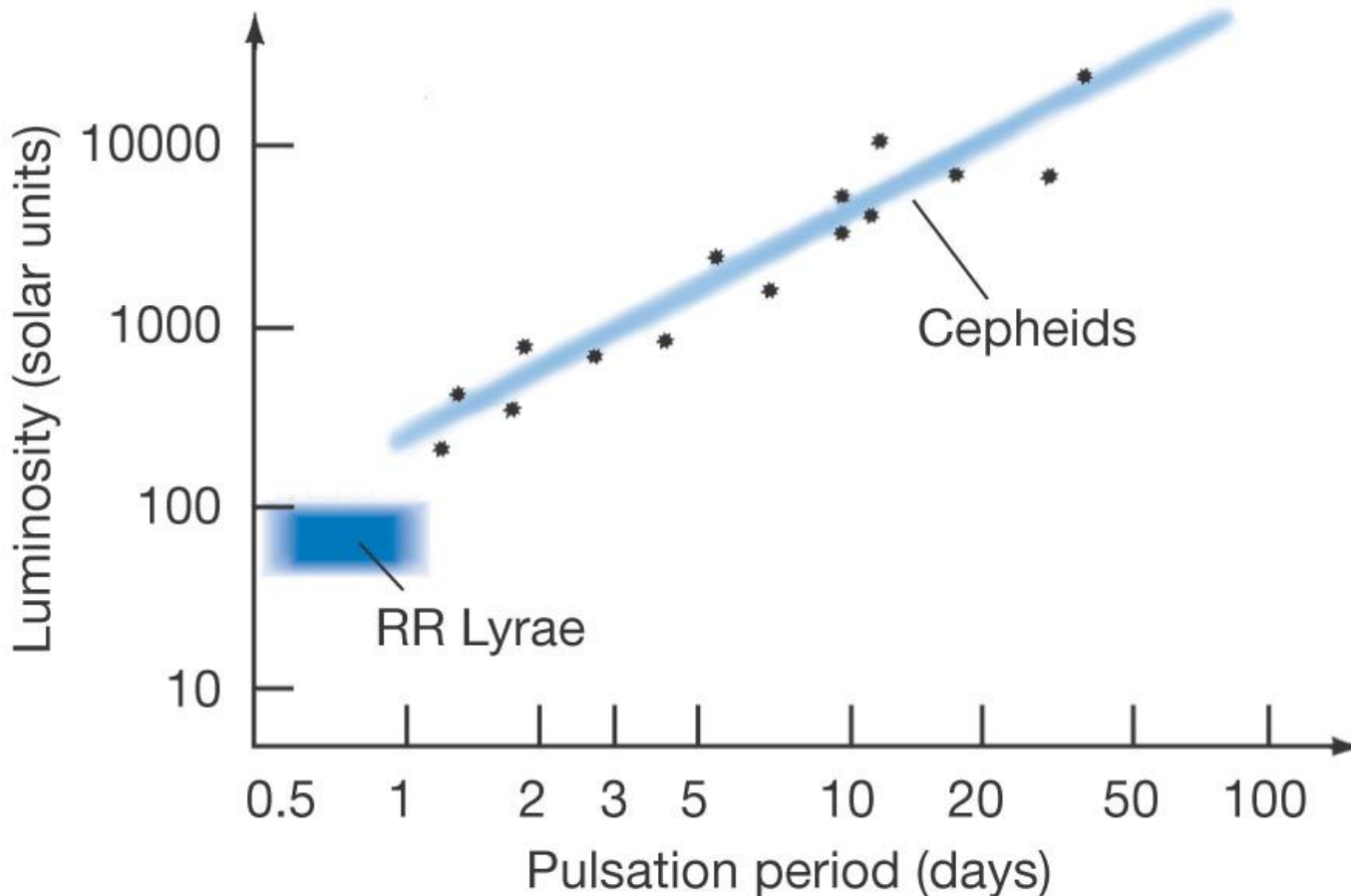
The variability of these stars comes from a dynamic balance between gravity and pressure—they have large oscillations around stability.

- Helium opacity-



## 23.2 Variable stars & distances

The usefulness of these stars comes from **their period-luminosity relation:**



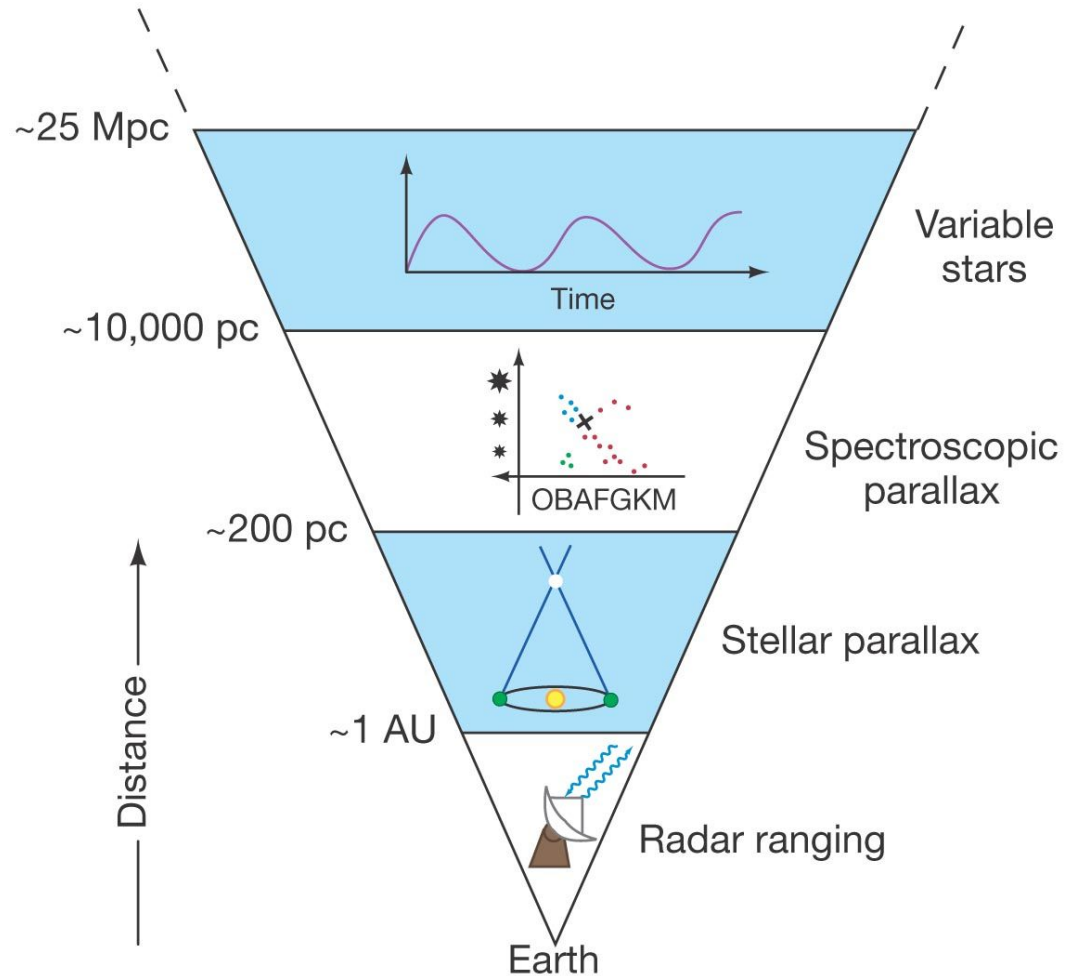
## 23.2 Variable stars & distances

**This allows us to measure the distances to these stars:**

- **RR Lyrae stars all have about the same luminosity; knowing their apparent magnitude allows us to calculate the distance.**
- **Cepheids have a luminosity that is strongly correlated with the period of their oscillations; once the period is measured, the luminosity is known and we can proceed as above.**

# 23.2 Measuring the Milky Way

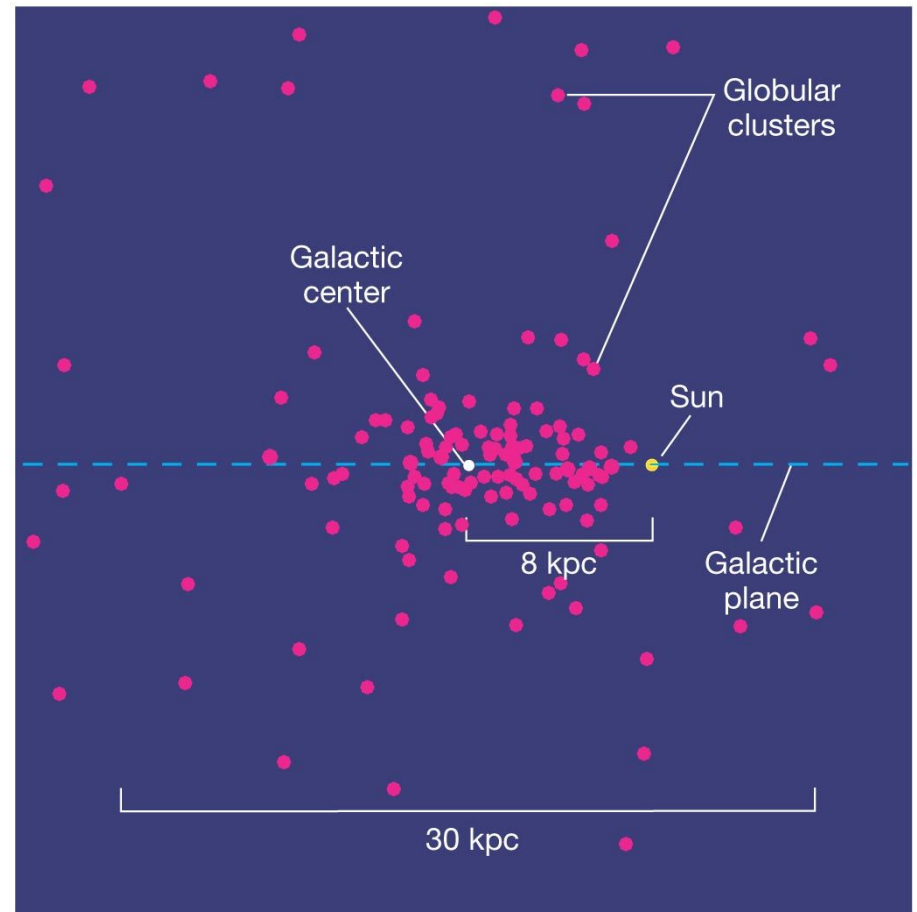
**We have now  
expanded our  
cosmic distance  
ladder one more  
step:**



## 23.2 Measuring the Milky Way

Many RR Lyrae stars are found in **globular clusters**. Harlow Shapley used these to estimate the size of the M.W.

This yields a much more **accurate** picture of the extent of our galaxy and our place within it.



## 23.2 Distance indicators

**So Far:**

***Novae, supernovae, cepheids, RR Lyrae***

**Other Standard Candles:**

***Brightest blue stars, Brightest red stars***

***Tip of the red giant branch (TRGB)***

***Planetary nebula luminosity function (PNLF)***

***globular cluster luminosity function (GCLF)***

**Standard Yardsticks:**

***Open Clusters, Globular Clusters, HII regions,***

***Size of galaxies of specific types***

**Other techniques:**

***Eclipsing Binaries, spectro. parallax, stel. Parallax,***

***Globular cluster spatial distribution***

**Trig parallax using radio interferometry! (2009)**

# Summary

- A good “standard candle” is luminous
- Variable stars can be used for distance measurement through the period–luminosity relationship.
- The extent of the Galaxy can be determined using globular clusters.
- Modern mapping of the MW is done with radio interferometry of gas clouds.



# Summary of Chapter 23 (cont.)

- **Spiral arms may be density waves.**
- **The galactic rotation curve shows large amounts of undetectable mass at large radii called dark matter.**
- **Activity near galactic center suggests presence of a 2 to 3 million solar-mass black hole**