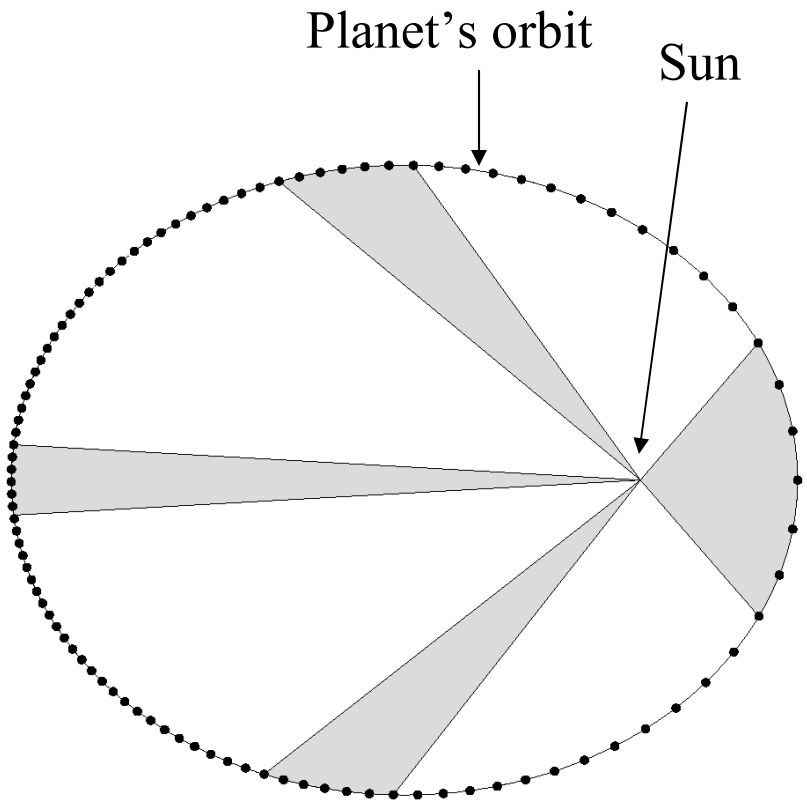
Kepler's Laws of Motion

- 1609 in Astronomica Nova (The New Astronomy)

- First Law – A planet orbits the Sun in an ellipse, with the Sun at one focus of the ellipse.
- Second Law – A line connecting a planet to the Sun sweeps out equal areas in equal time intervals
  - Several areas associated with the time interval of "six" are shown
    - They all have equal areas



Kepler's Third Law of Motion

From *Harmonica Mundi* (1619) (Harmony of the Worlds)

$$P^2 = a^3$$

P = orbital period

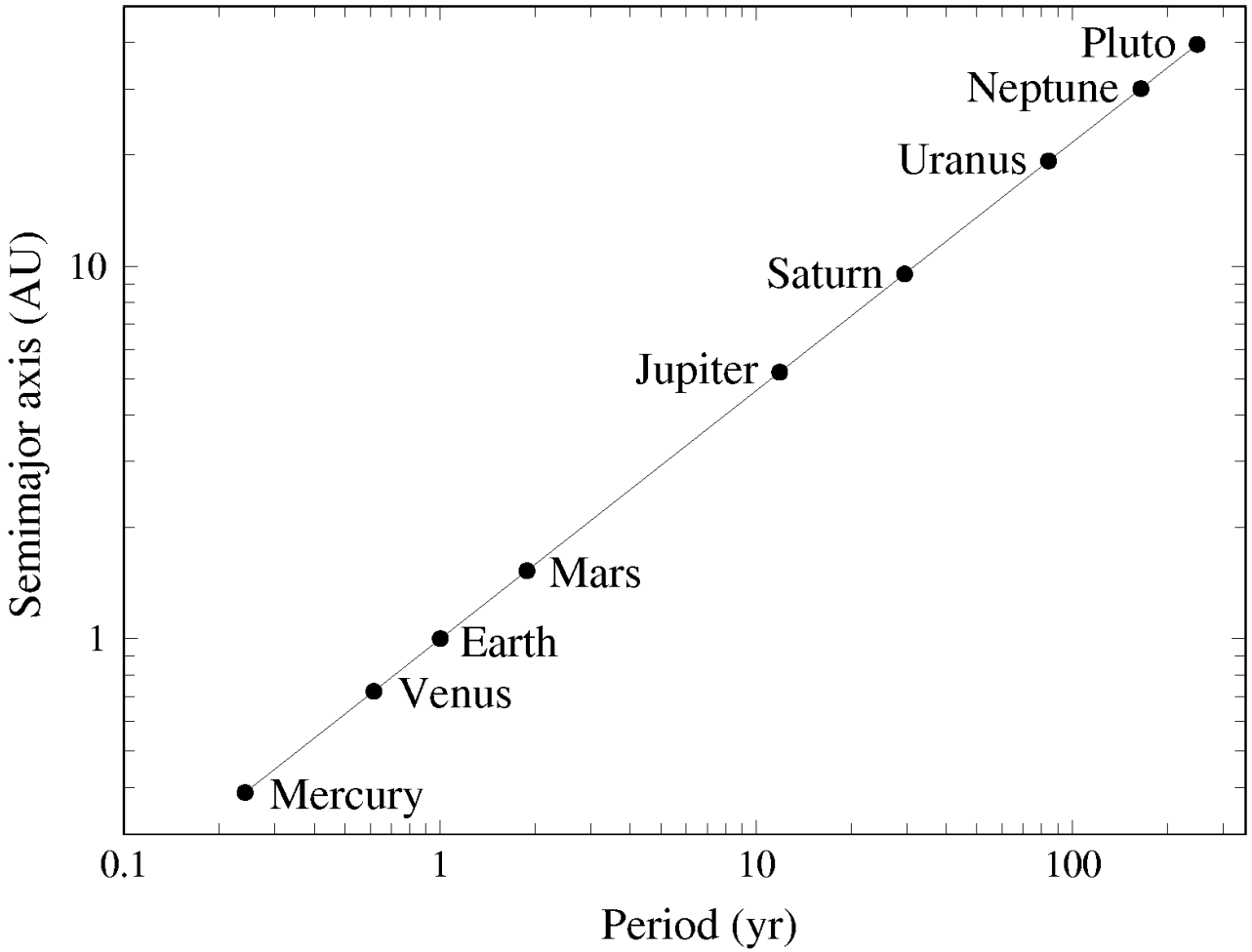
a = semimajor axis

"Power law" slope is 2/3:

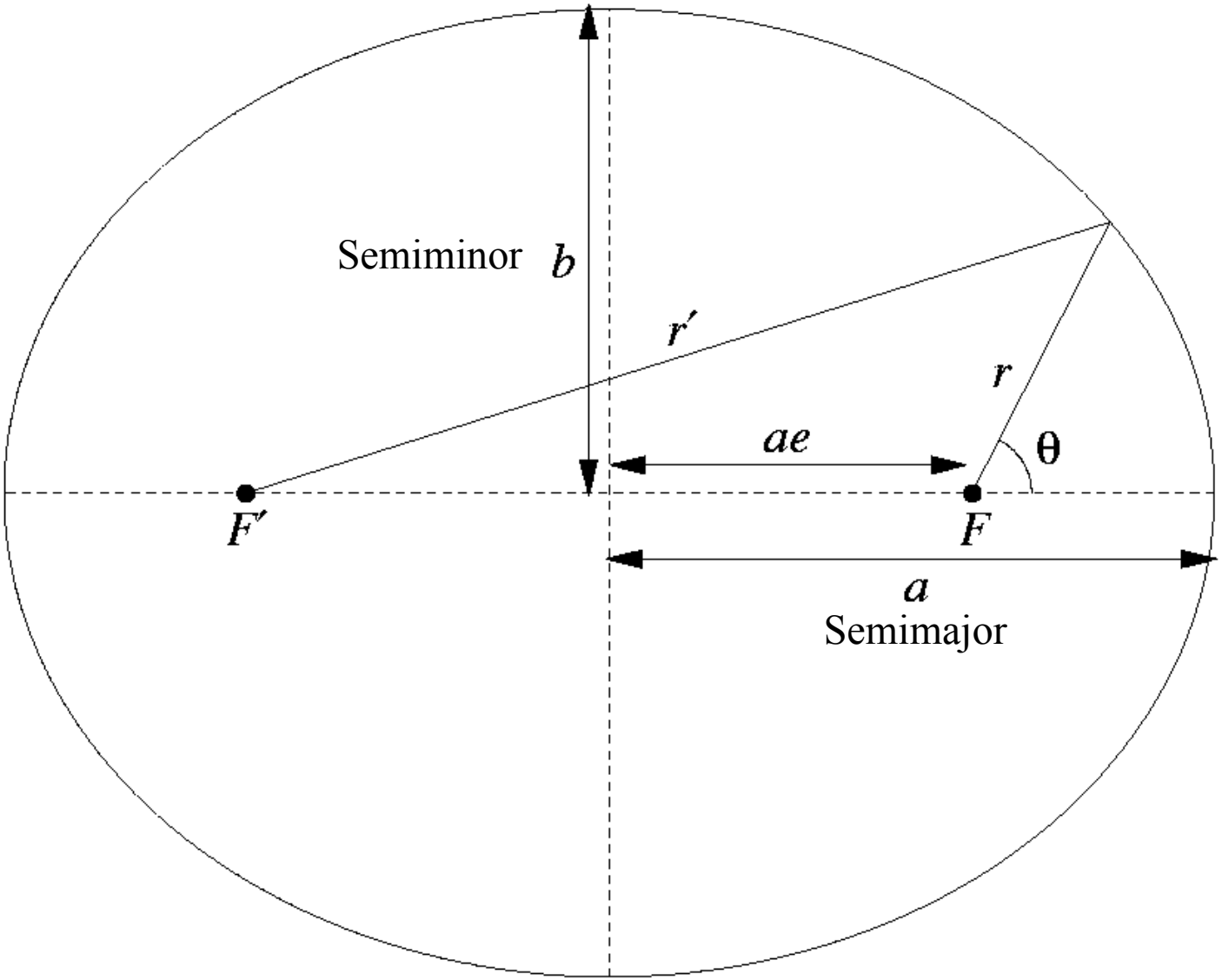
$$\log\left(P^2\right) = \log\left(a^3\right)$$

$$2\log\left(P\right) = 3\log\left(a\right)$$

$$\log\left(a\right) = \frac{2}{3}\log\left(P\right)$$



Ellipses



Ellipses – actual orbits

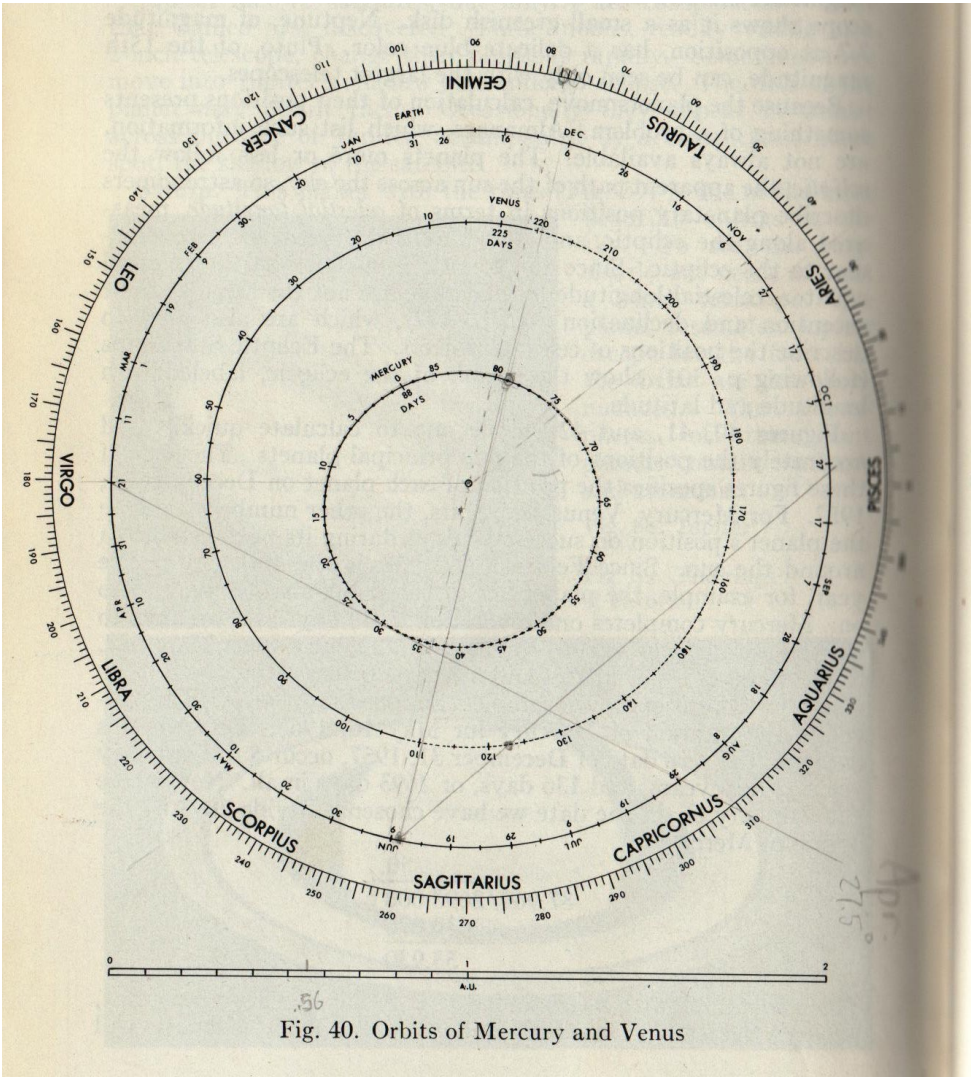


Fig. 40. Orbits of Mercury and Venus

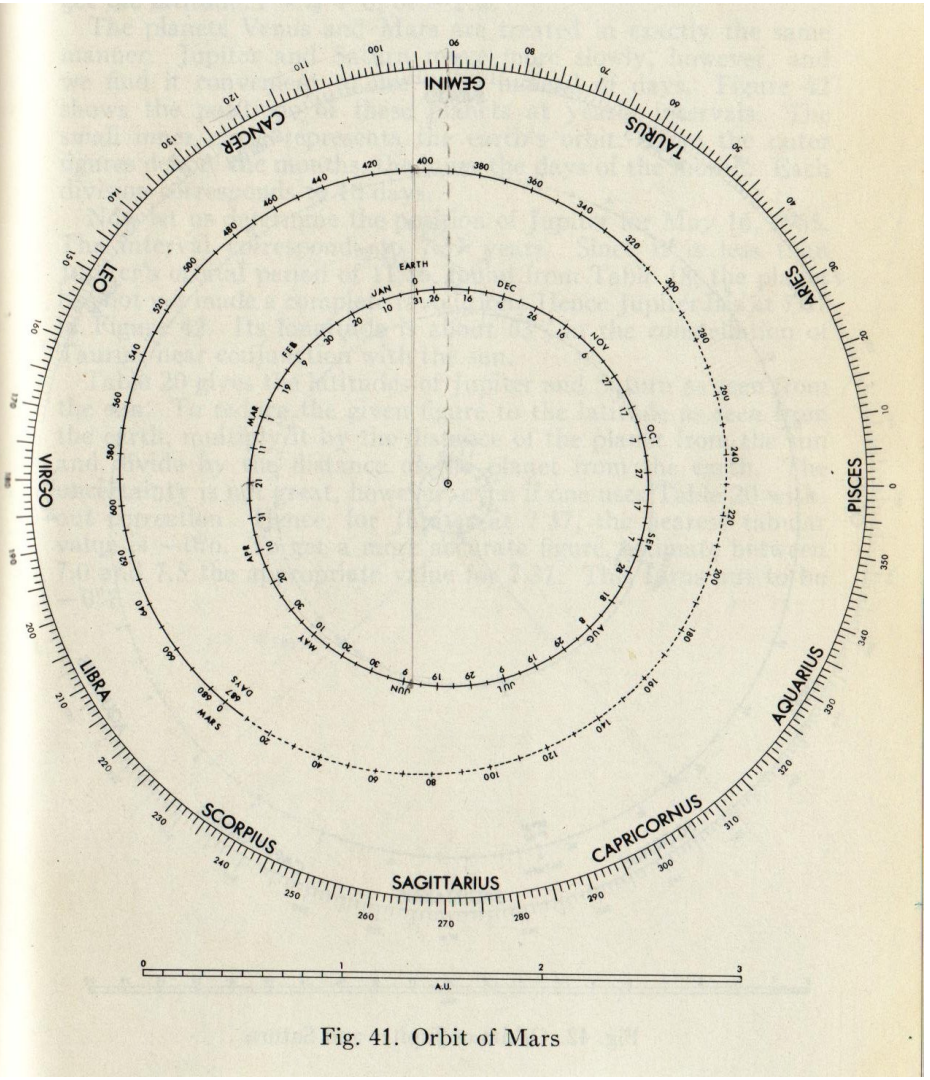
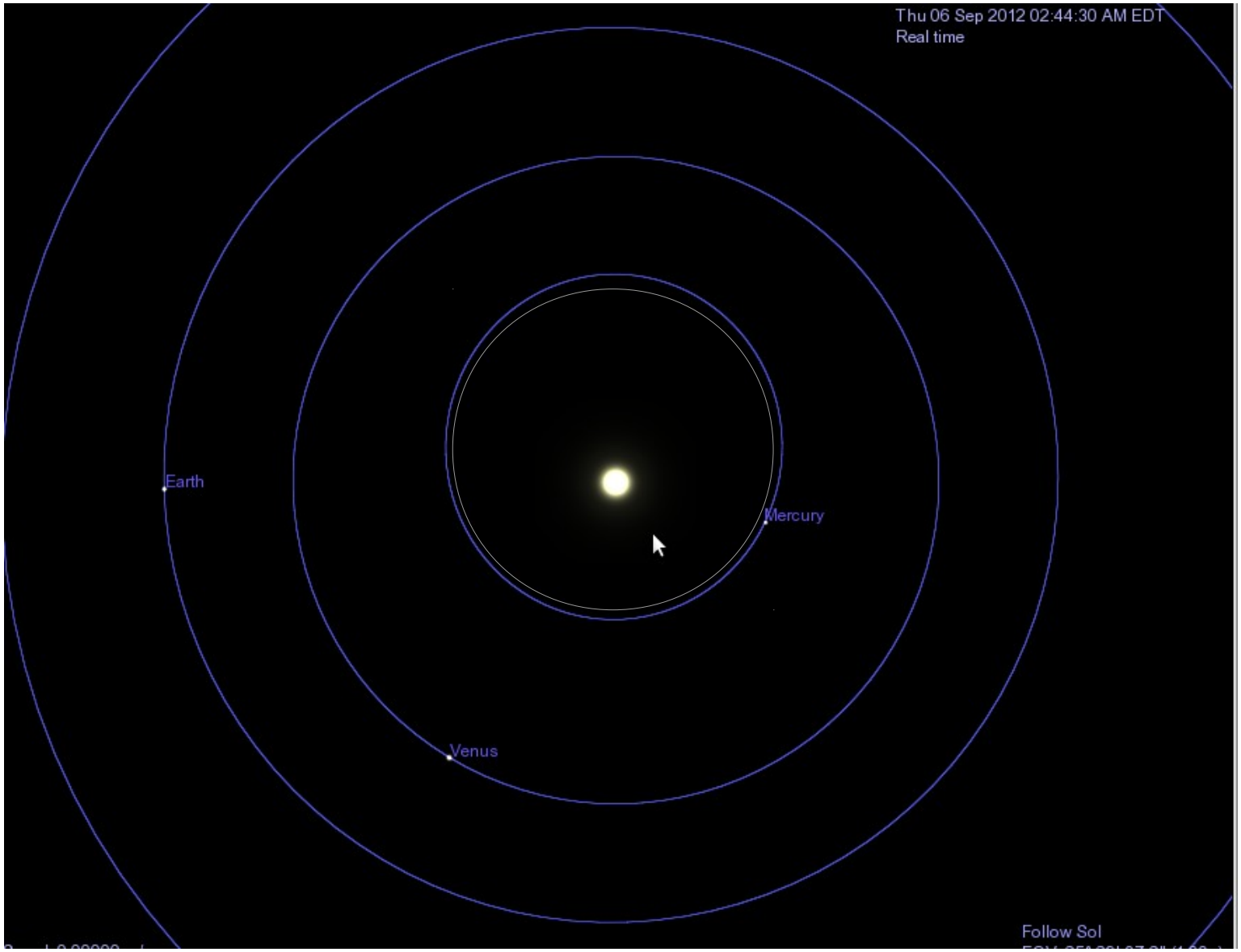


Fig. 41. Orbit of Mars

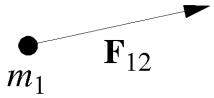
Ellipses – actual orbits

September 2012



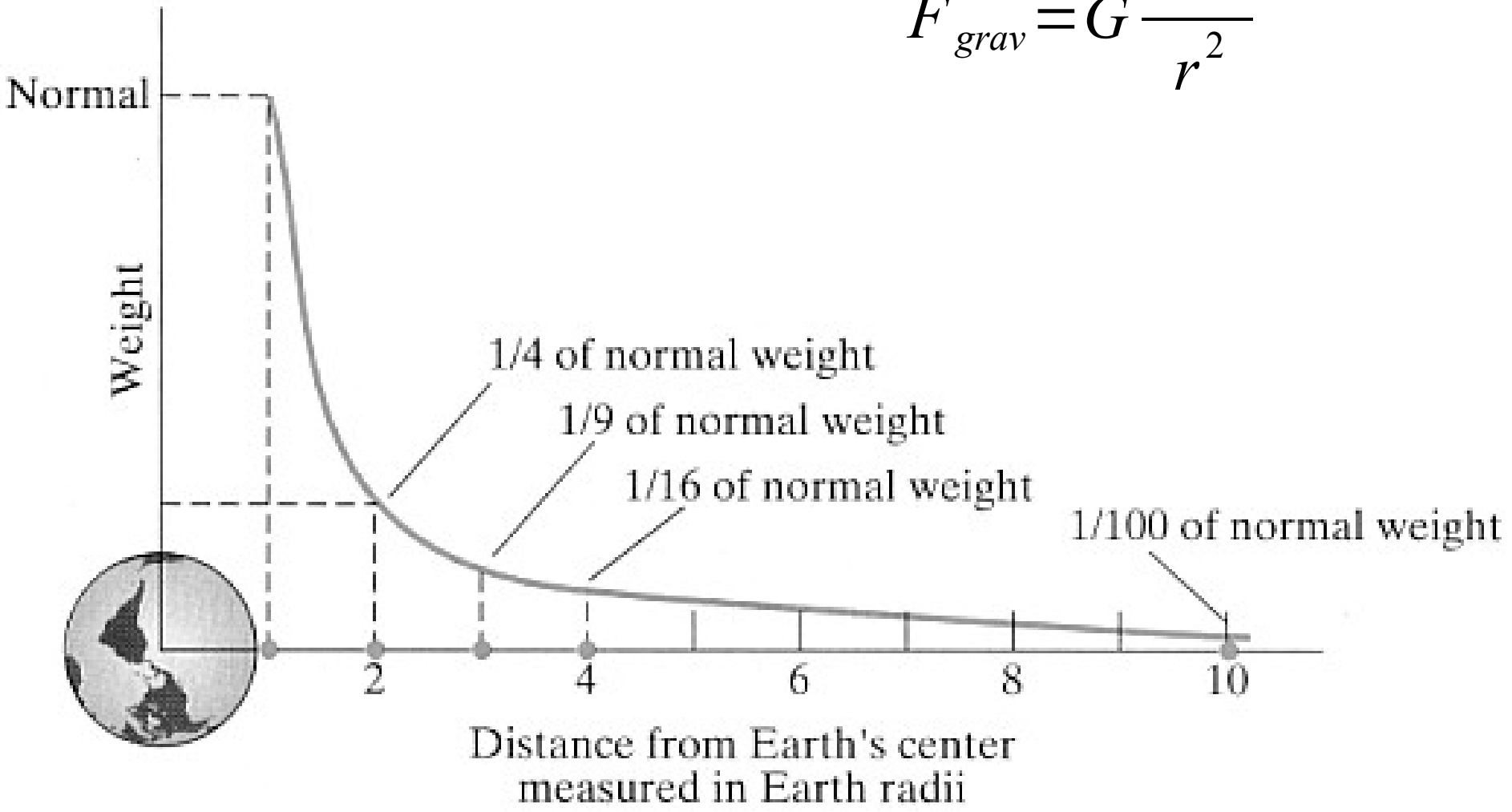
Newton's Laws of Motion

- Brachistochrone problem…
- 1st Law – Law of inertia
  - An object at rest remains at rest and an object in uniform motion remains in uniform motion unless acted upon by an unbalanced force.
  - An ***inertial reference frame*** is needed for 1st law to be valid
  - A non-inertial reference frame is being accelerated (e.g. In car going around a curve you feel a fictitious force)
- 2nd Law – $\mathbf{a} = \mathbf{F}_{net}/m$ or $\mathbf{F}_{net} = m\mathbf{a}$
  - The net force (sum of all forces) acting on an object is proportional to the object's mass and its resultant acceleration.
  - Inertial mass, m, does not appear to be different from gravitational mass
- 3rd Law
  - For every action there is an equal but opposite reaction



Universal Law of Gravitation

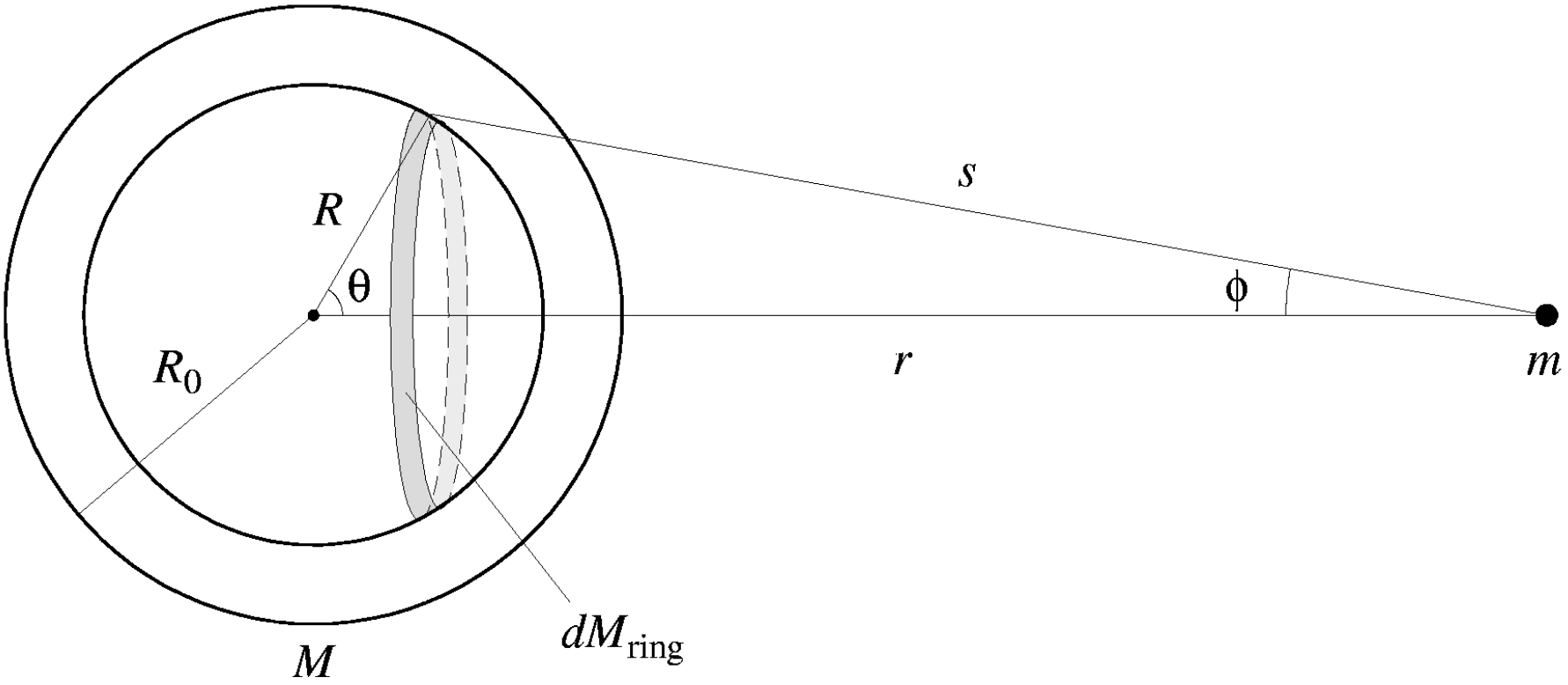
$$F_{grav} = G\frac{Mm}{r^2}$$



Shell theorems for gravity:

) The Force on *m* due to a uniform shell of mass is the same as the force due to a point mass at the center of the shell with the same total mass as the shell.

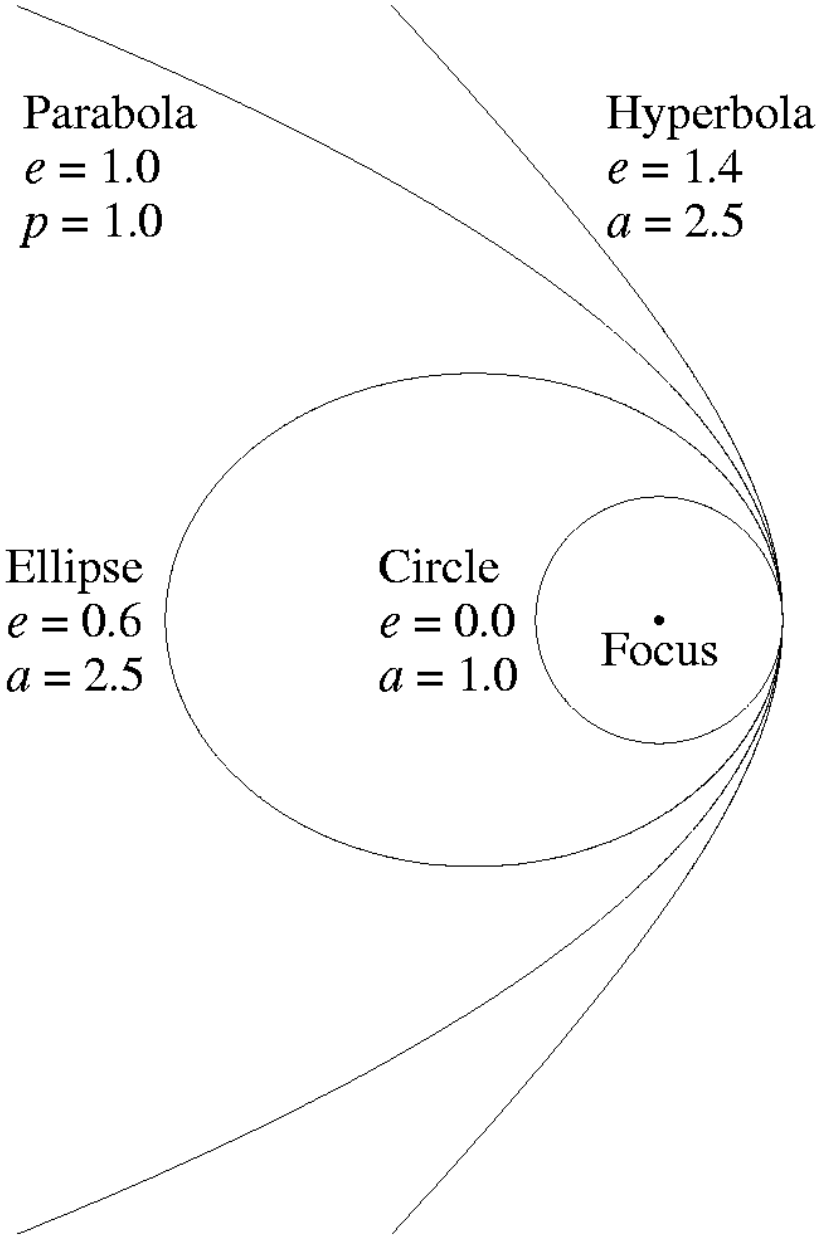
) The force of gravity inside of a uniform shell is zero.



Conic Sections

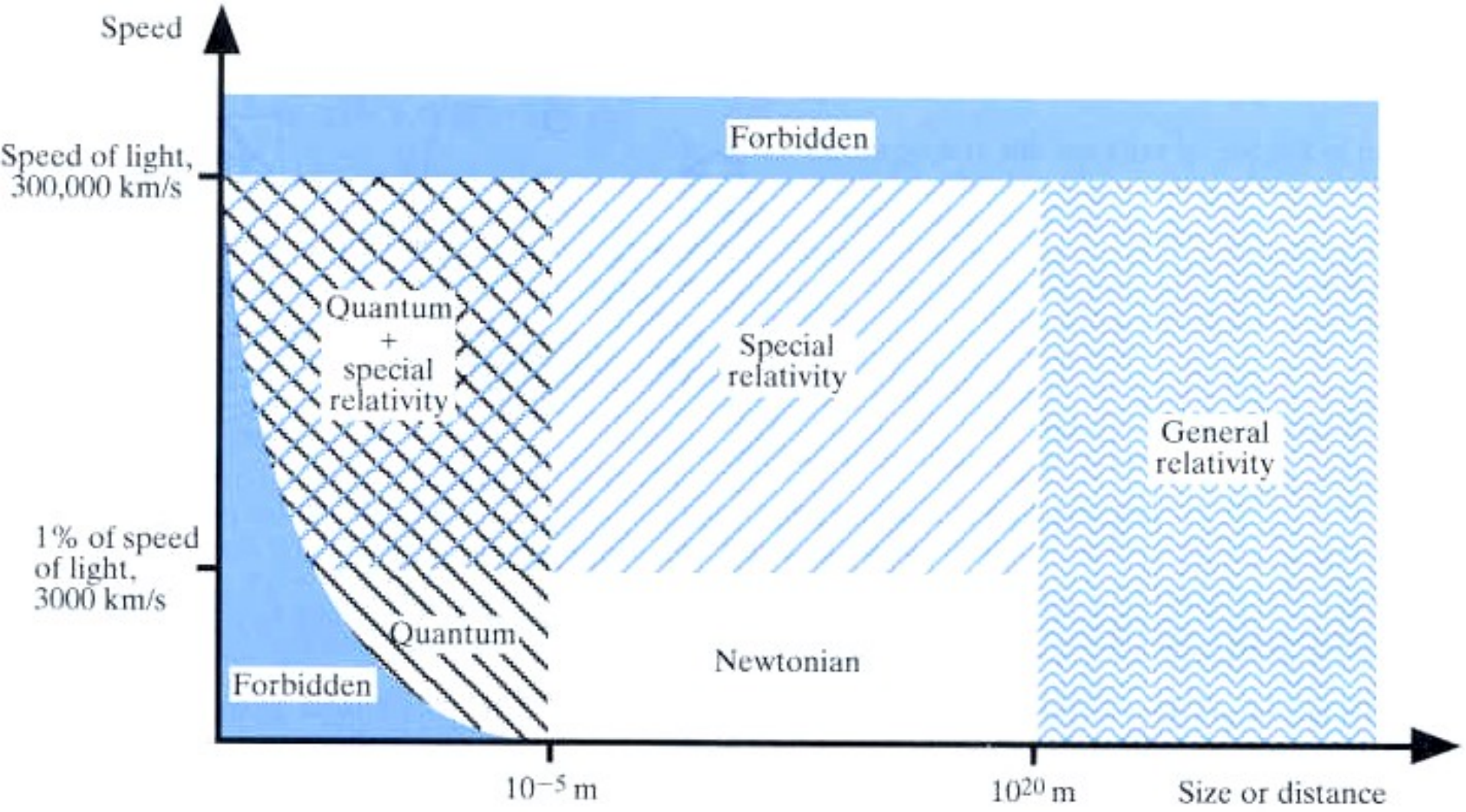
- All are possible in celestial mechanics
- Orbits are Ellipses

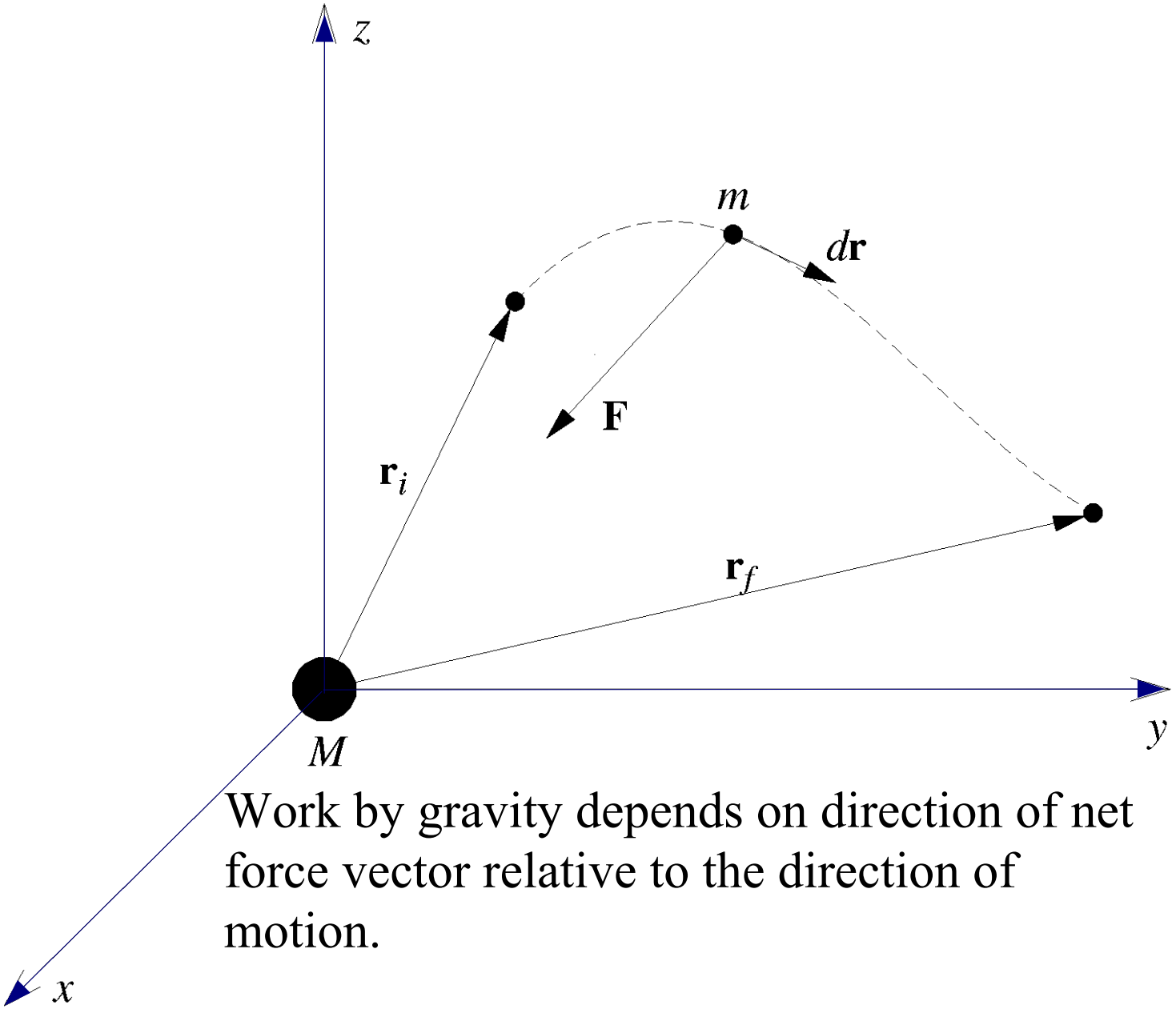
r = constant	e = 0	Circle
$r = \frac{a\left(1-e^2\right)}{1+e\cos\theta}$	$0 \le e < 1$	ellipse
$r = \frac{2p}{1+\cos\theta}$	e = 1	parabola
$r = \frac{a\left(e^2-1\right)}{1+e\cos\theta}$	e > 1	hyperbola



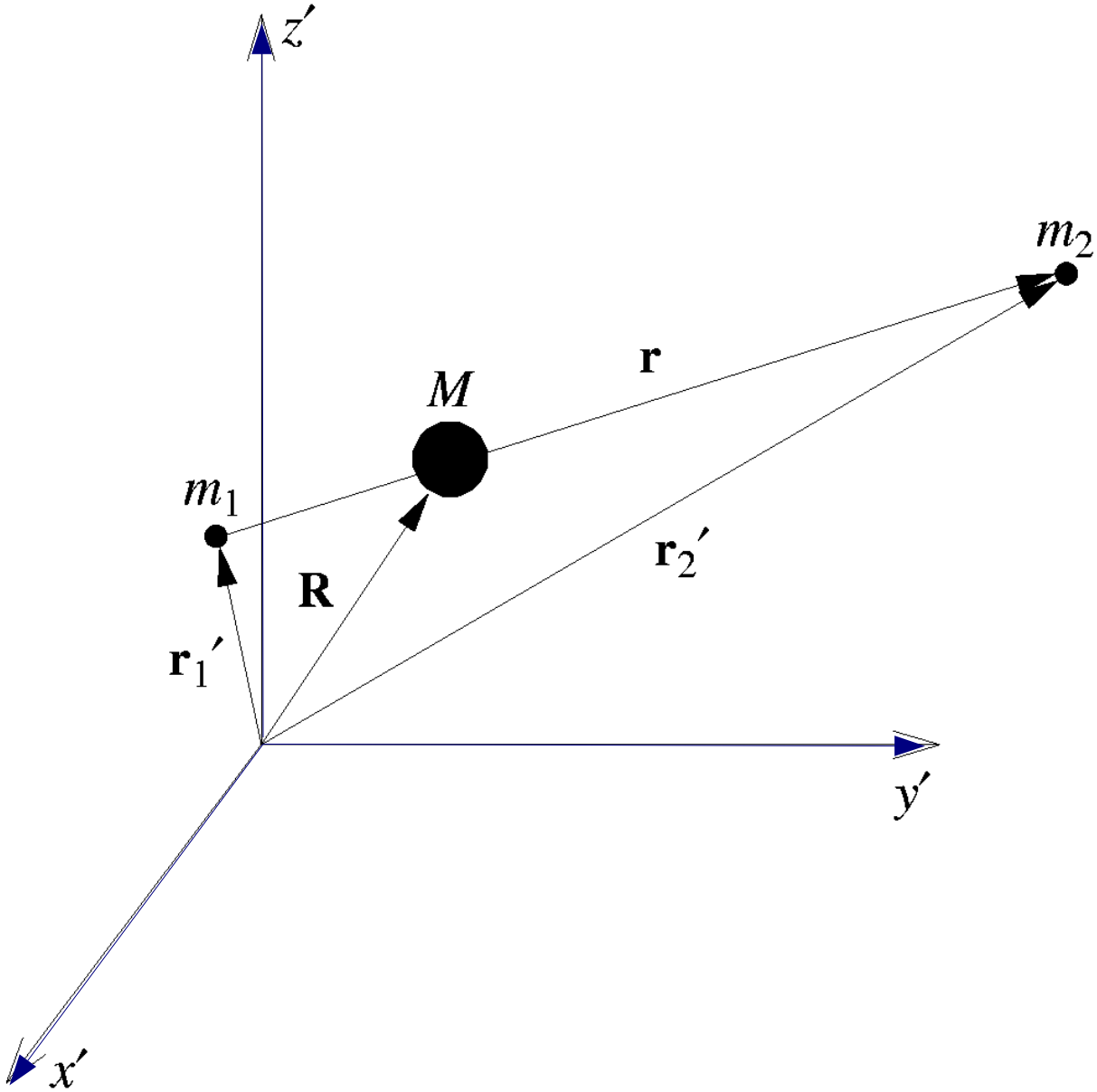
Beyond Newton

- Gravity passed every test until ~1890s
- Newton's gravity and motion is incorrect when …





Work by gravity depends on direction of net force vector relative to the direction of motion.


z'
m_2
$\mathbf{r}$
M
m_1
$\mathbf{r}_2'$
$\mathbf{R}$
$\mathbf{r}_1'$
y'
x'