

UNIVERSIDAD DE LOS ANDES

# Three body problem in the spherical geometry

by

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*“Write a funny quote here.”*

If the quote is taken from someone, their name goes here

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# *Abstract*

Faculty of Sciences  
Department of Physics

Bachelor in Physics

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The Thesis Abstract is written here (and usually kept to just this page). The page is kept centered vertically so can expand into the blank space above the title too...

# *Acknowledgements*

The acknowledgements and the people to thank go here, don't forget to include your project advisor...

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# Abbreviations

**LAH** List Abbreviations **Here**

# Physical Constants

Speed of Light  $c = 2.997\,924\,58 \times 10^8 \text{ ms}^{-\text{s}}$  (exact)

# Symbols

$a$	distance	m
$P$	power	W ( $\text{Js}^{-1}$ )
$\omega$	angular frequency	$\text{rads}^{-1}$

*For/Dedicated to/To my...*

# Chapter 1

## Introduction

### 1.1 Motivation

### 1.2 State of the art

## Chapter 2

# The three body problem in the plane

In this chapter a classic approach of a somehow general case of the three body problem in 2 dimensions is going to be presented. This will give some necessary intuition to develop the analogous problem in the spherical geometry. To begin with, the problem is going to be described in great detail; then its integrability is going to be proven; and finally, a formalism to describe the movement of the particles is going to be presented.

### 2.1 The definition of the problem

The three body problem presented here is that of three particles of electrical charge  $e$  and mass  $m$  confined to a plane, under the influence of a strong magnetic field perpendicular to it, and forces whose potentials satisfy translational and rotational symmetries in the plane.

Given this information, the Hamiltonian associated with this system has the form:

$$H = \sum_{i=1}^3 \frac{1}{2m} \left\| \vec{p}_i - e\vec{A}(\vec{q}_i) \right\|^2 + V(\vec{q}_1, \vec{q}_2, \vec{q}_3) + \frac{\omega_c^2}{2m} \sum_{i=1}^3 \|\vec{q}_i\|^2 \quad (2.1)$$

Where  $\vec{q}_i = x_i\hat{i} + y_i\hat{j}$ ,  $\vec{p}_i = p_{x_i}\hat{i} + p_{y_i}\hat{j}$  and  $\vec{A}(\vec{q})$  is the magnetic vector potential, which satisfies  $\nabla \times \vec{A} = B\hat{k}$ .

Besides, the potential  $V(\vec{q}_1, \vec{q}_2, \vec{q}_3)$  satisfies the symmetries:

$$V(R\vec{q}_1 + \vec{a}, R\vec{q}_2 + \vec{a}, R\vec{q}_3 + \vec{a}) = V(\vec{q}_1, \vec{q}_2, \vec{q}_3) \quad (2.2)$$

For any rotation  $R$  and translation  $\vec{a}$  in the plane.

## 2.2 The canonical transformation of the guiding centres

For the proof of integrability for this system, and for further analysis of the trajectories of the particles, let us perform the well known transformation of the guiding centres.

This transformation is defined by the following two equations:

$$\vec{\pi}_i = \vec{p}_i - e\vec{A}(\vec{q}_i) \quad (2.3)$$

$$\vec{R}_i = \vec{q}_i - \frac{\hat{k} \times \vec{\pi}_i}{eB} \quad (2.4)$$

The equation (2.3) passes from the canonical momentum  $\vec{p}_i$  to the linear momentum  $\vec{\pi}_i$ , which is much more intuitive and understandable; while the equation (2.4) transforms the general position  $\vec{q}_i$  to the position of the instantaneous guiding centre  $\vec{R}_i$ .

In a system without the interaction potentials, the electrically charged particles are known to perform the circular motion of the cyclotron with radii that depends on the initial linear momenta. In this case, the guiding centres would be constant in time as would be the linear momenta. However, with the introduction of an interacting potential, the momentum of each particle may vary making the guiding centre change too, which is why the instantaneous interpretation of the guiding centres is necessary.

Now, let us calculate the Poisson brackets for this new set of coordinates in a specific particle.

$$\begin{aligned}
\{\pi_1, \pi_2\} &= \frac{\partial \pi_1}{\partial q_\alpha} \frac{\partial \pi_2}{\partial p_\alpha} - \frac{\partial \pi_2}{\partial q_\alpha} \frac{\partial \pi_1}{\partial p_\alpha} \\
&= -e\delta_{\alpha 2} \frac{\partial A_1}{\partial q_\alpha} + e\delta_{\alpha 1} \frac{\partial A_2}{\partial q_\alpha} \\
&= -e \frac{\partial A_1}{\partial q_2} + e \frac{\partial A_2}{\partial q_1} \\
&= e(\nabla \times \vec{A})_3 = eB
\end{aligned}$$

$$\begin{aligned}
\{R_1, R_2\} &= \{q_1, q_2\} + \left\{q_1, -\frac{\pi_1}{eB}\right\} + \left\{\frac{\pi_2}{eB}, q_2\right\} + \left\{\frac{\pi_2}{eB}, -\frac{\pi_1}{eB}\right\} \\
&= \frac{1}{eB} \left( \cancel{\{p_1, q_1\}}^{-1} - e \cancel{\{A_1, q_1\}}^0 + \cancel{\{p_2, q_2\}}^{-1} - e \cancel{\{A_2, q_2\}}^0 \right) + \frac{eB}{(eB)^2} \\
&= \frac{-2}{eB} + \frac{1}{eB} = -(eB)^{-1}
\end{aligned}$$

$$\begin{aligned}
\{R_1, \pi_2\} &= \{q_1, \pi_2\} + \left\{\frac{\pi_2}{eB}, \pi_2\right\} \\
&= \cancel{\{q_1, p_2\}}^0 - e \cancel{\{q_1, A_2\}}^0 \\
&= \{R_2, \pi_1\} = 0
\end{aligned}$$

This Poisson brackets can be generalised to the transformation for the three particles. Taking  $i, j = \{1, 2, 3\}$  and  $\alpha, \beta = \{1, 2\}$ :

$$\{\pi_{i,\alpha}, \pi_{j,\beta}\} = (eB) \delta_{ij} \epsilon_{\alpha\beta} \quad (2.5)$$

$$\{R_{i,\alpha}, R_{j,\beta}\} = -(eB)^{-1} \delta_{ij} \epsilon_{\alpha\beta} \quad (2.6)$$

$$\{R_{i,\alpha}, \pi_{j,\beta}\} = 0 \quad (2.7)$$

Equations (2.5)-(2.6) allow us to identify the proposed transformation as canonical. However, this is not the usual canonical transformation where the position coordinates and the momentum coordinates are canonical conjugates. In this special case, one component of the momentum is canonical conjugate with the other momentum coordinate, as equally happens for the position coordinates.



Now, with a huge magnetic field, if the potential of the interaction forces does not vary abruptly in space, we can use the approximation  $\vec{R}_i \approx \vec{q}_i$  to average the potentials over the guiding centres, that is, we can replace  $\vec{q}_i$  for  $\vec{R}_i$  in  $V(\vec{q}_1, \vec{q}_2, \vec{q}_3)$ .

We can support the last approximation as follows: In the cyclotron problem, the radius of the circular motion described is proportional to the linear momentum and inversely proportional to the magnetic field. Then, in the presence of a big  $B$ , the radius of the cyclotron would shrink to a very small size. Regarding the case we are working with, the radii of the instantaneous cyclotron motion would be proportional to  $\|(\hat{k} \times \vec{\pi}_i)(eB)^{-1}\|$  and its frequency to  $\sqrt{B}$ . As the potential  $V$  does not vary abruptly in the radii scale, the averaging of this motion over the guiding centres means that this potential does not sense that circular motion. Moreover, given the big frequency of the cyclotrons and the scale of variance of the potential, the scale of time of the local circular motions is far smaller than that of the motion of the guiding centres. Therefore, we can ignore the instantaneous quality of the circular motion, and take it as constant in a scale of time small enough for the motion of the guiding centres. In this sense we say that the coordinates for the guiding centres decouple from that of the linear momenta of the particles.

Before replacing the new set of coordinates in the Hamiltonian, it is necessary to do a scale transformation to obtain the proper Poisson brackets for the formal definition of canonical transformation, that is:

$$\begin{aligned}\vec{\pi}_i &\rightarrow (eB)^{-1/2} \vec{\pi}_i \\ \vec{R}_i &\rightarrow \sqrt{eB} \vec{R}_i\end{aligned}$$

With this consideration, the Hamiltonian of the system in the new set of rescaled coordinates is given by:

$$H = \sum_{i=1}^3 \frac{eB}{2m} \|\vec{\pi}_i\|^2 + V\left((eB)^{-1/2} \vec{R}_1, (eB)^{-1/2} \vec{R}_2, (eB)^{-1/2} \vec{R}_3\right) + \frac{\omega_c^2}{2meB} \sum_{i=1}^3 \|\vec{R}_i\|^2 \quad (2.8)$$

This Hamiltonian, given equation (2.7) can be decomposed in a Hamiltonian that describes the movement of the guiding centres, and other that describes the movement of the linear momenta. In one hand, the Hamiltonian for the linear momenta is easily identified with the harmonic oscillator, whereas the one that characterises the movement of

the guiding centres needs a deeper analysis.

## 2.3 Integrability of the system

As the Hamiltonian describing the trajectories of the linear momenta of the particles is that of an harmonic oscillator, this part of the problem is integrable and its solutions are widely known. The guiding centre Hamiltonian, in turn, needs to be analysed more deeply. For this purpose, let us take the following convention:

$$H_{gc} = \frac{\omega_c^{*2}}{2m} \sum_{i=1}^3 \|\bar{x}^2\| + \|\bar{y}^2\| + V^*(\bar{x}, \bar{y}) \quad (2.9)$$

Where  $\bar{x} = (x_1, x_2, x_3)$  and  $\bar{y} = (y_1, y_2, y_3)$ , being  $x_i, y_i$  the rescaled coordinates of the guiding centres of the particles. For simplicity, the potential  $V$  and the constant  $\omega_c$  have been rescaled to take into account the scale transform of the coordinates and maintain the original form of the Hamiltonian:

$$\omega^* = \frac{\omega}{\sqrt{eB}}$$

$$V^*(\bar{x}, \bar{y}) = V\left(\frac{\bar{x}}{\sqrt{eB}}, \frac{\bar{y}}{\sqrt{eB}}\right)$$

Clearly, the new potential  $V^*$  still has the symmetries expressed in the equation (2.2). Furthermore, in the new order for the scaled guiding centres coordinates, the Poisson brackets take the form:

$$\{y_i, x_j\} = \delta_{ij} \quad (2.10)$$

Now that the guiding centres Hamiltonian has been expressed in terms of the proper canonical set of coordinates, the fastest way to prove the integrability of the system is via the Liouville-Arnol'd theorem [?, Sect. 49]. For this theorem, it is only necessary to find 2 more independent integrals in involution (besides the Hamiltonian).

To get this 2 integrals, let us exploit the symmetries of the guiding centres Hamiltonian. We then define the generators of translations and rotation in the plane, which are symmetries of the potential:

$$\begin{aligned} T_x &= \sum_{i=1}^3 x_i \\ T_y &= \sum_{i=1}^3 y_i \\ R_z &= \frac{1}{2} \sum_{i=1}^3 (x_i^2 + y_i^2) \end{aligned} \tag{2.11}$$

It is easily verifiable that these are indeed the symmetries generators. To see that, take the first order infinitesimal transformations of translation and rotation:

$$\begin{aligned} x_i &\rightarrow x_i + \epsilon \\ y_i &\rightarrow y_i + \epsilon \\ (x_i, y_i) &\rightarrow (x_i + \epsilon y_i, y_i - \epsilon x_i) \end{aligned}$$

Now note that for the infinitesimal translations, the potential of the primed coordinates is related to the potential of the normal coordinates by a directional derivative, which can be identified with the Poisson bracket of the potential  $V^*$  and each generator:

$$\begin{aligned} 0 &= V^*(\bar{x} + \epsilon, \bar{y}) - V^*(\bar{x}, \bar{y}) = \epsilon \sum_{i=1}^3 \frac{\partial V^*(\bar{x}, \bar{y})}{\partial x_i} = \epsilon \{V^*, T_x\} = 0 \\ 0 &= V^*(\bar{x}, \bar{y} + \epsilon) - V^*(\bar{x}, \bar{y}) = \epsilon \sum_{i=1}^3 \frac{\partial V^*(\bar{x}, \bar{y})}{\partial y_i} = \epsilon \{T_y, V^*\} = 0 \end{aligned}$$

For the infinitesimal rotation, the relation is analogous:

$$0 = V^*(\bar{x} + \epsilon \bar{y}, \bar{y} - \epsilon \bar{x}) - V^*(\bar{x}, \bar{y}) = \frac{\partial V^*(\bar{x}, \bar{y})}{\partial x_i} (\epsilon y_i) - \frac{\partial V^*(\bar{x}, \bar{y})}{\partial y_i} (\epsilon x_i) = \epsilon \{V^*, R_z\}$$

Therefore, we conclude that the generators of translations and rotations in the plane commute with the potential  $V^*$  due to its symmetries. Besides, the generator of rotations is multiple of the harmonic-like part of the guiding centres Hamiltonian which validates that  $R_z$  is other integral in involution. The generators of translations are not integrals in involution, for they do not commute with the harmonic potential, however, we can calculate a quantity in terms of these generators, which already commute with the potential  $V^*$ , to make it commute with the remaining part of  $H_{gc}$ :

$$L = T_x^2 + T_y^2 \quad (2.12)$$

This new quantity  $L$  clearly commutes with the potential  $V^*$  because the Poisson bracket is a linear differential operator in one component and it obeys the Leibniz rule. Moreover, it also commutes with the rotation generator  $R_z$ :

$$\begin{aligned} \{T_x^2 + T_y^2, R_z\} &= \sum_{i,j,k} \{x_i x_j + y_i y_j, x_k^2 + y_k^2\} \\ &= \sum_{i,j,k} \{x_i x_j, y_k^2\} + \{y_i y_j, x_k^2\} = \sum_{i,j,k} y_k \{x_i x_j, y_k\} + x_k \{y_i y_j, x_k\} \\ &= \sum_{i,j,k} y_k x_i \delta_{jk} + y_k x_j \delta_{ik} - x_k y_i \delta_{jk} - x_k y_j \delta_{ik} \\ &= \sum_{i,j} y_j x_i + y_i x_j - x_j y_i - x_i y_j = 0 \end{aligned}$$

As we found  $L$  as the last integral in involution, we conclude, by the Liouville-Arnol'd theorem, that the subsystem of guiding centres is integrable by quadratures.

## 2.4 Analysis of the motion

## Chapter 3

# The problem of a charged particle in the magnetic field of a monopole

With the objective to obtain some intuition about the N-body problem restricted to a spherical geometry, we study in this chapter the symmetries and trajectories of a charged particle under the influence of the magnetic field of a monopole. To achieve this we present the deduction, via the Lagrangian formalism, of the so called Poincaré cone [?] that characterises the trajectory of a particle in this situation. We then extrapolate the important symmetries used in the Lagrangian formalism to the Hamiltonian formalism to retrieve some important aspects of the classical counterparts of the known Haldane formalism [?].

### 3.1 The Poincaré cone

Let  $L$  be the Lagrangian of a charged particle of charge  $-e$  and mass  $m$  under the influence of a magnetic monopole of magnitude  $g$ . Then  $L$  takes the form:

$$L(\vec{x}, \dot{\vec{x}}) = \frac{m}{2} \|\dot{\vec{x}}\|^2 - e\vec{A}_{\hat{u}}(\vec{x}) \cdot \dot{\vec{x}} \quad (3.1)$$

Where  $\vec{A}(\vec{x})$  is the vector potential of the magnetic monopole with singularity along the direction defined by the unit vector  $\hat{u}$ . This potential takes the form [?]:

$$\vec{A}_{\hat{u}}(\vec{x}) = \frac{g}{r} \frac{\hat{u} \times \hat{r}}{1 + \hat{u} \cdot \hat{r}} \quad (3.2)$$

The first thing to note in this Lagrangian is its time independence, which yields to the conservation of the Jacobi integral:

$$\frac{\partial L}{\partial \dot{\vec{x}}} \cdot \dot{\vec{x}} - L = m \left\| \dot{\vec{x}} \right\|^2 = m \left\| \dot{\vec{x}}_0 \right\|^2$$

With  $\dot{\vec{x}}_0$  the initial velocity of the particle. Here the Jacobi integral clearly represents the kinetic energy of the particle, which is constant in time because of the linear dependence of the term associated with the magnetic interaction that basically means that the magnetic fields do no work.

For the following symmetry of the system, let us take the action integral:

$$S = \int_{t_1}^{t_2} L(\vec{x}, \dot{\vec{x}}) dt = \int_{t_1}^{t_2} \left( \frac{m}{2} \left\| \dot{\vec{x}} \right\|^2 - e \vec{A}_{\hat{u}}(\vec{x}) \cdot \dot{\vec{x}} \right) dt \quad (3.3)$$

We note that this action is invariant under the following scale transformation:

$$\begin{aligned} \vec{x}' &= \exp^s \vec{x} \\ t' &= \exp^{2s} t \end{aligned}$$

Which, by the Noether's theorem must imply a conserved quantity. To calculate it we prefer the method stated in [?, Thm (dontforget)], however, to apply it, the Lagrangian must be reparametrised to include the time as a generalised coordinate. To achieve this, note that:

$$\dot{\vec{x}} = \frac{d\vec{x}}{dt} = \frac{\frac{d\vec{x}}{d\tau}}{\frac{dt}{d\tau}} := \frac{\dot{\vec{x}}}{\dot{t}}$$

Then the action can be written in terms of the new parameter  $\tau$ :

$$S = \int_{t_1}^{t_2} L(\vec{x}, \dot{\vec{x}}) dt = \int_{\tau_1}^{\tau_2} \dot{t} L(\vec{x}, \dot{\vec{x}} \dot{t}^{-1}) d\tau = \int_{\tau_1}^{\tau_2} L'(\vec{x}, \dot{\vec{x}}, \dot{t}) d\tau$$

This gives us the new parametrised Lagrangian  $L'$ :

$$L'(\vec{x}, \overset{\circ}{\vec{x}}, \overset{\circ}{t}) = \overset{\circ}{t} L(\vec{x}, \overset{\circ}{\vec{x}} \overset{\circ}{t}^{-1})$$

**Appendix A**

**An Appendix**