

Data archiving and data dissemination for the next generation of high-resolution solar telescopes

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Scientist at the Experimental Solar Physics group at KIS
Head of the KIS Science Data Centre

- Space-based (solar) telescopes provide science-ready data to the scientific community

This translates into the use and re-use of data \Rightarrow
high scientific output

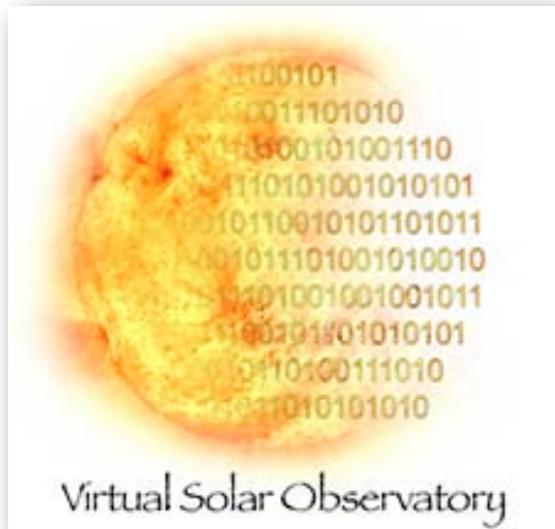
(backed-up by a high number of publications)

- This is still not the case for ground-based solar observatories..
for which the PI-mode is still in use

Space-based vs. ground-based solar observatories

- There is a need to make data discovery easy & to distribute data, especially, for ground-based observations

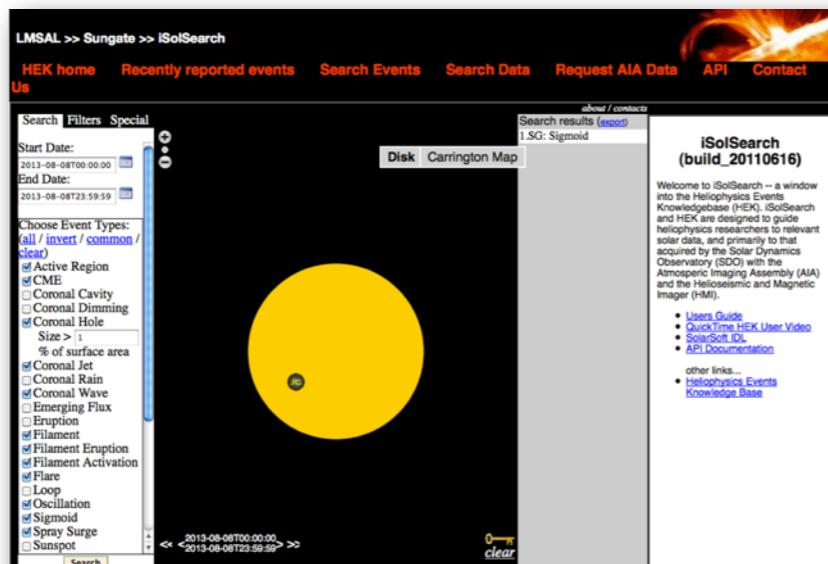
Data archives, e.g., VSO



Browsing tools, e.g., (j)Helioviewer



Event searching, e.g., HEK



Next generation 4m-class ground-based solar telescopes

Next generation large solar telescopes

- Next generation ground-based solar observatories
(US) DKIST & (European) EST 4m-class telescopes

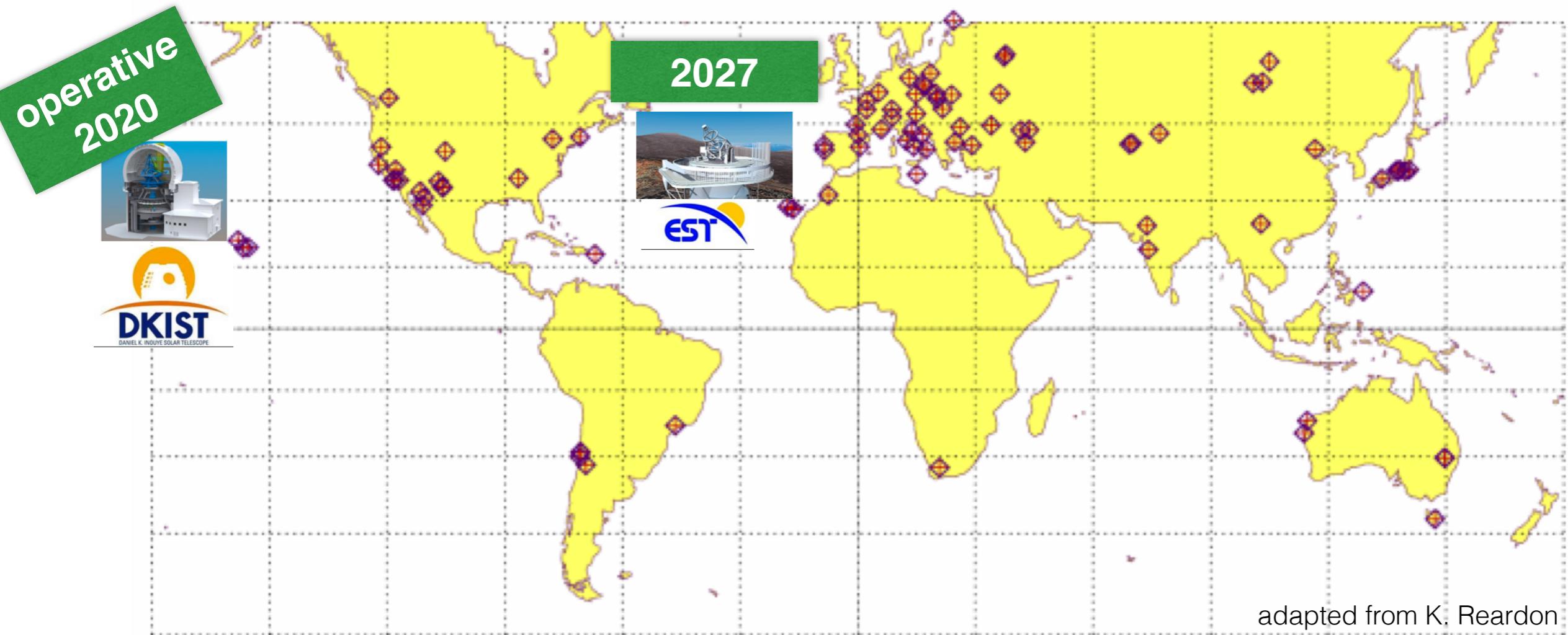
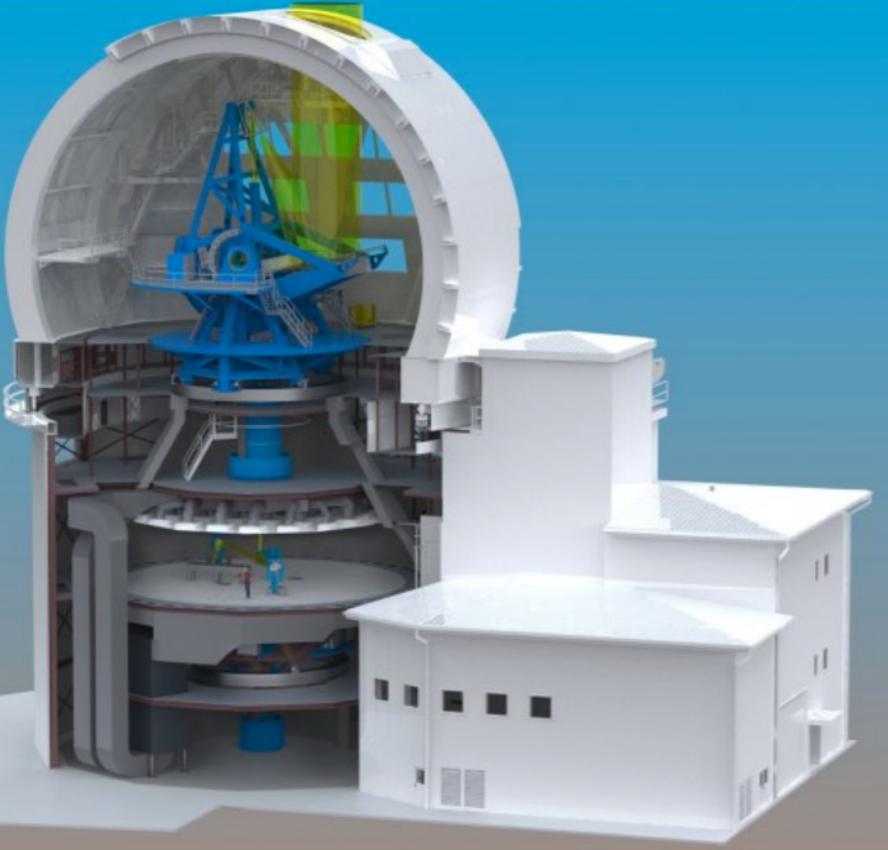


Fig. 1: The distribution of the observatories around the globe.

Daniel K. Inouye Solar Telescope DKIST



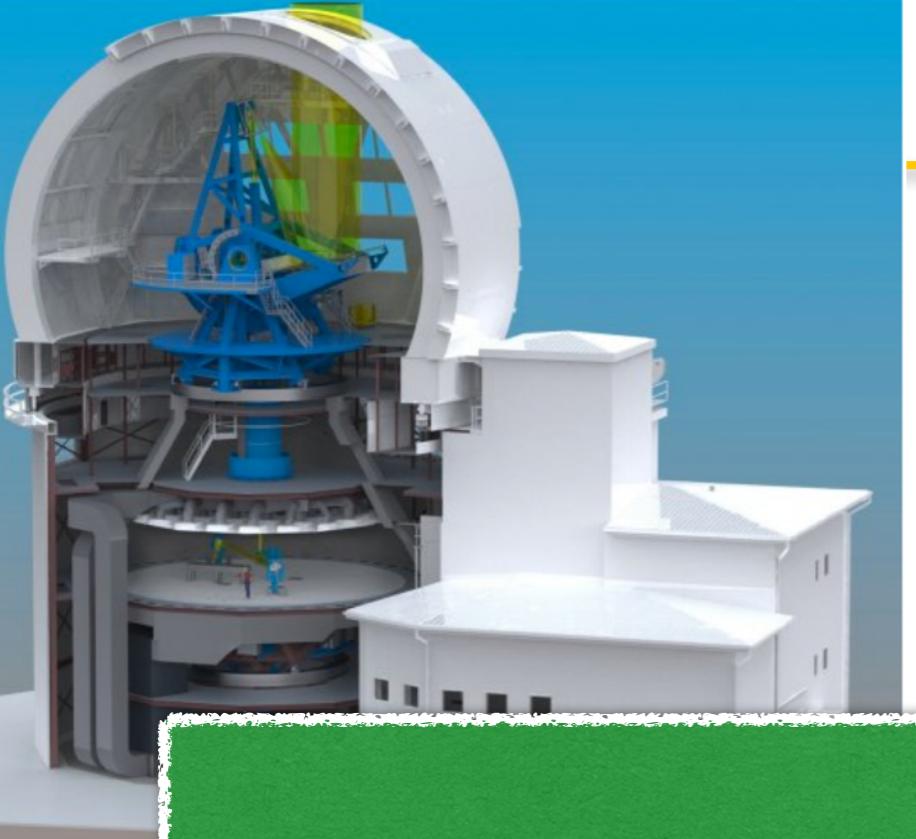
Haleakala Obs., Maui

Rendering of proposed ATST facility at the primary Mees site on Haleakala, Maui, Hawaii by Tom Kekona, K. C. Environmental, Inc. Original aerial photo by Frank Rizzo.



DKIST is a 4m solar telescope facility under construction by the (US) National Solar Observatory atop the Haleakala volcano in Maui

In operation on mid 2020, it will be the world's largest solar telescope



Haleakala

Rendering of proposed ATST facility at the p...



KIS (Germany) is contributing to the DKIST project with the *Visible Tunable Filter* (VTF) post-focus instrument, a 2D spectro-polarimeter

→ European contribution

DKIST is a 4m solar telescope facility under construction by the (US) National Solar Observatory atop the Haleakala volcano in Maui

In operation on mid 2020, it will be the world's largest solar telescope

- ☀ The upcoming ground-based large solar telescopes + new post-focus instrumentation (detectors) ➔ vast increase of the data volume

Example: expected DKIST data stream

DKIST Instrument	Detectors	Hourly Data Volume
VBI	1 x 4096 x 4096 detector, 30 fps	6.5 TB/hour
ViSP	4 x 4096 x 4096 detector, 12.5/6.25 fps	2.1 TB/hour
VTF	3 x 4096 x 4096 detector, 10 fps	8.6 TB/hour
DL-AUDOR	1 x 4096 x 4096 detector, 10.5/5.25 fps	4.3 TB/hour
Cryo		

~ 160 x increased data volume!

Adapted from

The experience with VTF/DKIST data handling
will be essential to get ready for EST

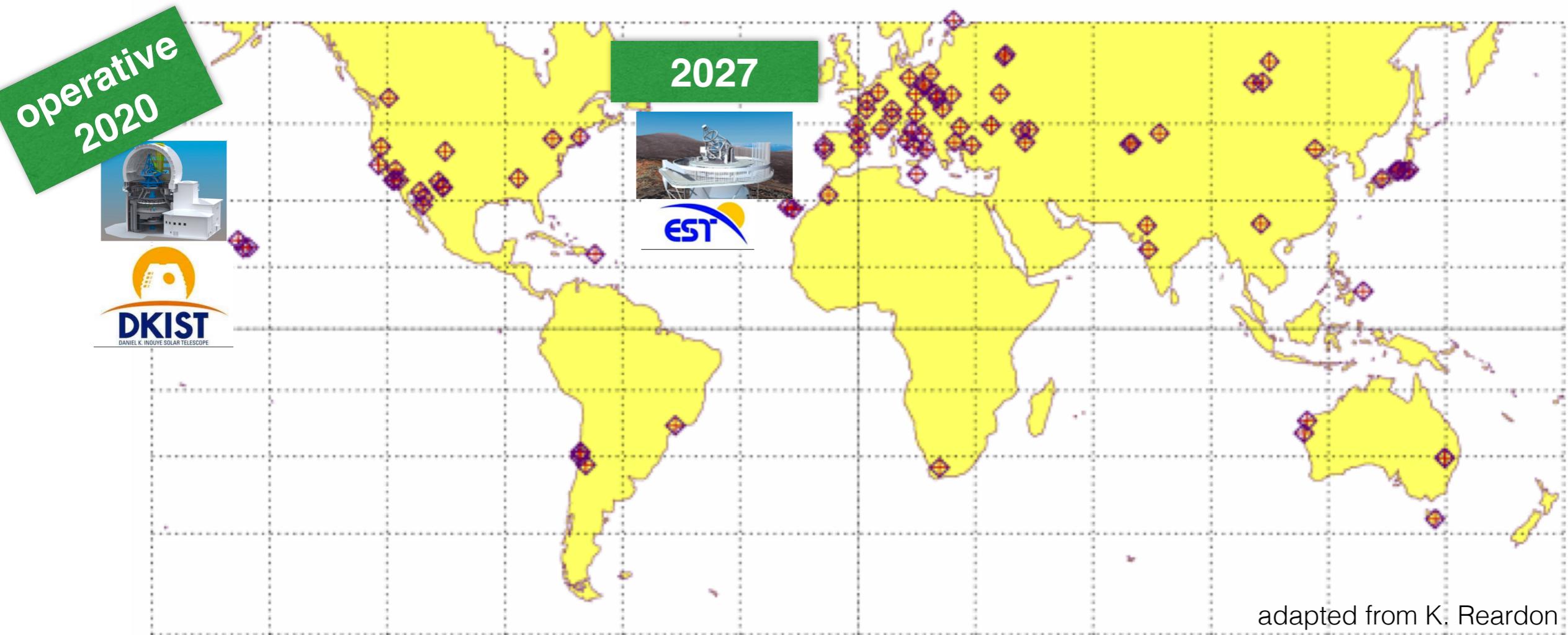


Fig. 1: The distribution of the observatories around the globe.

European Solar Telescope

Institute		Location	
IGAM	Institutsbereich Geophysik, Astrophysik und Meteorologie		Graz
HVO	Hvar Observatory		Hvar
AIASCR	Astronomical Institute AS CR		Ondrejov
THEMIS	THEMIS S.L., [n]		
KIS	Kiepenheuer-In		
UniDeb	Heliophysical C		
INAF	Istituto Naziona		
UU	Utrecht Univers		
ITA	Institute of The		
IA UW	Astronomical In		
AISAS	Astronomical In		
IAC	Instituto de Astr		
SU	The Institute for		
IRSOL	Istituto Ricerche		
UCL-MSSL	University Coll		

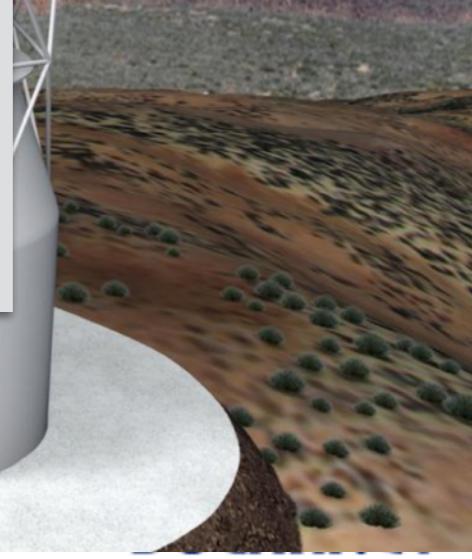
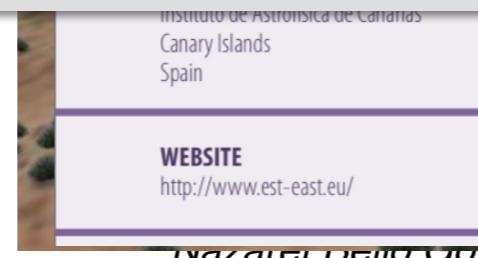


The *European Solar Telescope* (EST) is a 4m solar telescope for high-resolution solar observations to be located in the Canary Islands

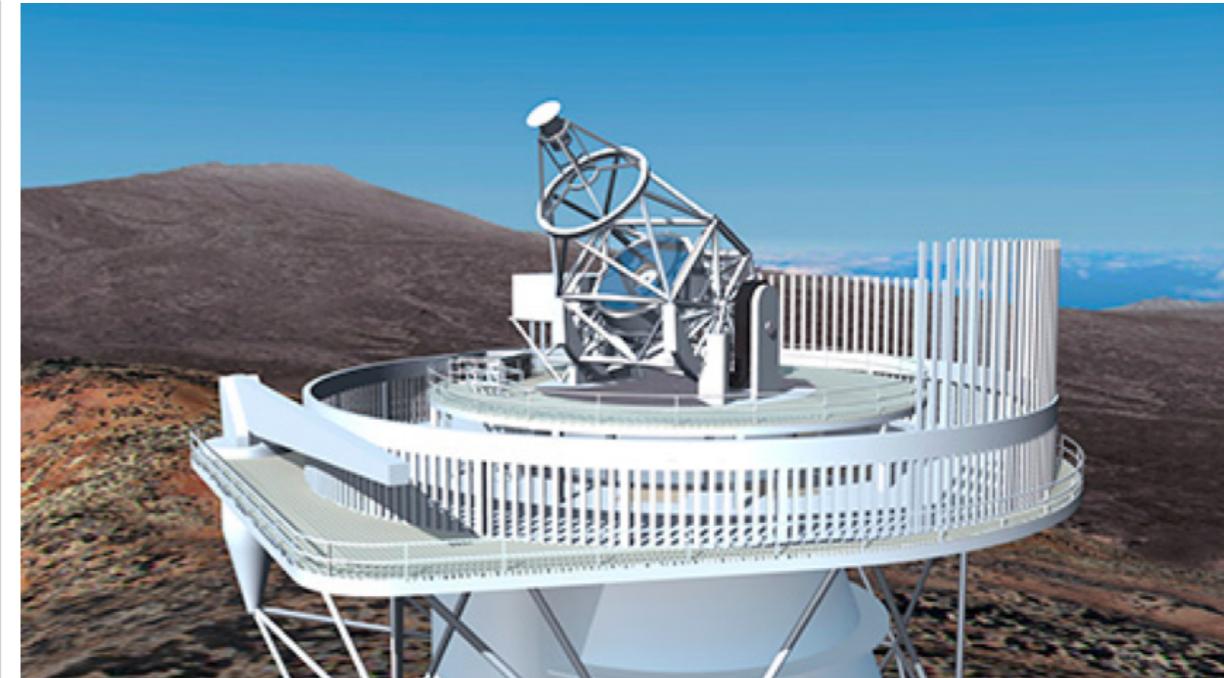
The EST project engages 15 European institutions lead by the IAC (Spain) and KIS (Germany)

The (3 years) conceptual design study conducted by research institutions and industrial companies was finalised in May 2011

It was co-financed by the European Commission under the EU's Seventh Framework Programme for Research (FP7)



Institute		Location	
IGAM	Institutsbereich Geophysik, Astrophysik und Meteorologie		Graz
HVO	Hvar Observatory		Hvar
AIASCR	Astronomical Institute AS CR		Ondrejov
THEMIS	THEMIS S.L., [note 1] INSU-CNRS, CNR		Paris
KIS	Kiepenheuer-Institut für Sonnenphysik		Freiburg
UniDeb	Heliophysical Observatory Debrecen		Debrecen
INAF	Istituto Nazionale di Astrofisica		Rome
UU	Utrecht University, Sterrekundig Instituut		Utrecht
ITA	Institute of Theoretical Astrophysics		Oslo
IA UW	Astronomical Institute of the Wroclaw University		Wroclaw
AISAS	Astronomical Institute of the Slovak, Academy of Sciences		Tatranská Lomnica
IAC	Instituto de Astrofísica de Canarias		La Laguna
SU	The Institute for Solar Physics		Stockholm
IRSOL	Istituto Ricerche Solari		Locarno
UCL-MSSL	University College London - MSSL		London



TIMELINE

- ESFRI Roadmap entry: 2016
- Preparation phase: 2011-2019
- Construction phase: 2019-2025
- Operation start: 2026

ESTIMATED COSTS

- Capital value: Not Available
- Preparation: 10 M€
- Construction: 200 M€
- Operation: 9 M€/year

HEADQUARTERS

Instituto de Astrofísica de Canarias
Canary Islands
Spain

WEBSITE

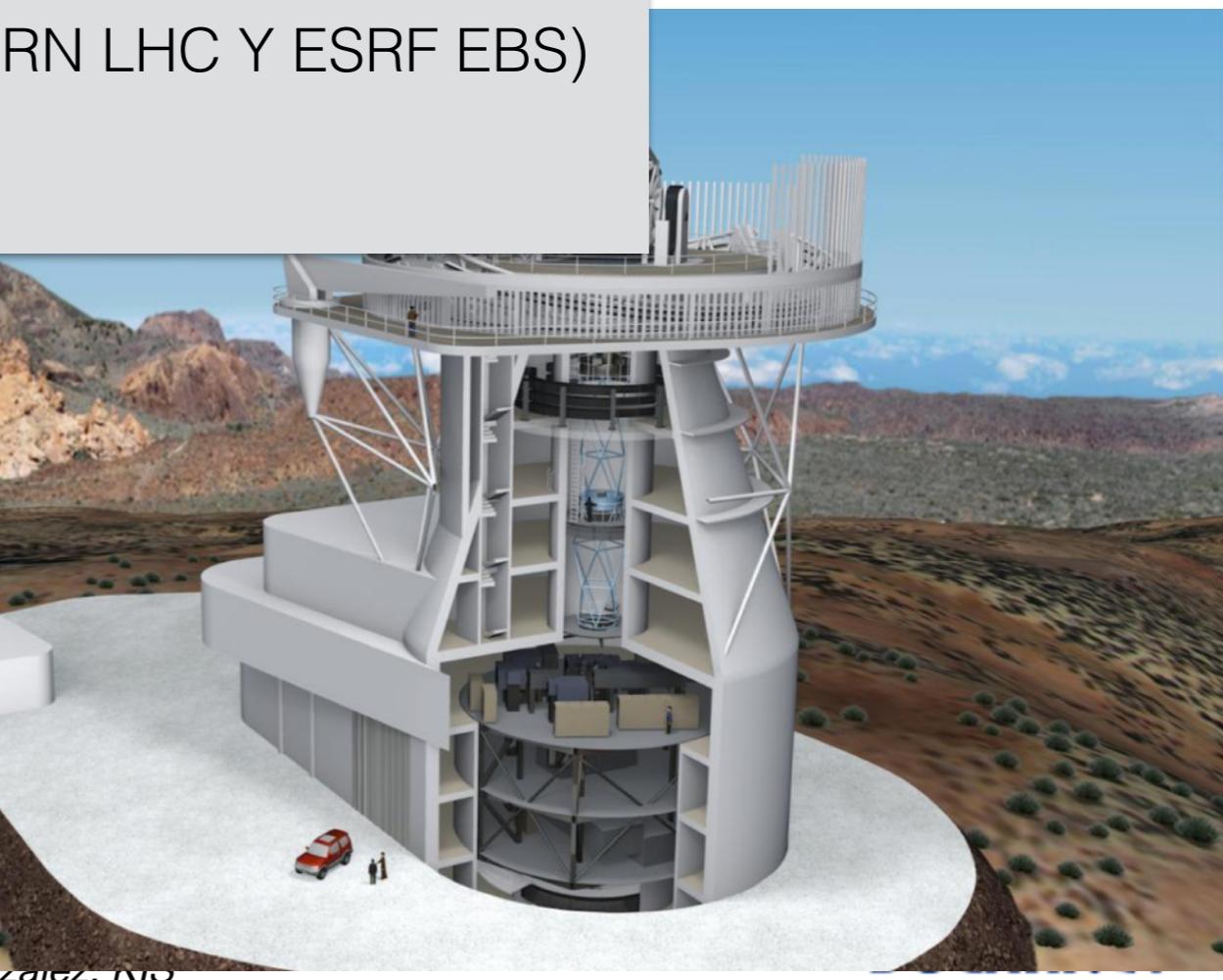
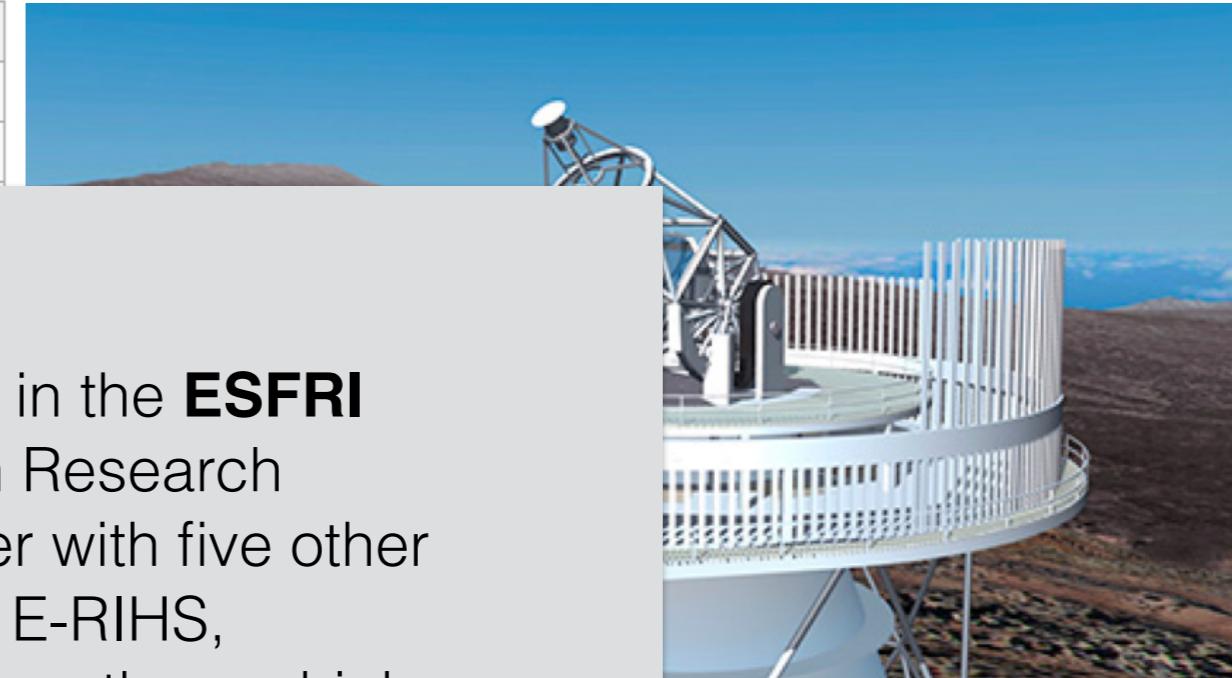
<http://www.est-east.eu/>

Nazaret Benito González, KIS

Canaries, Spain

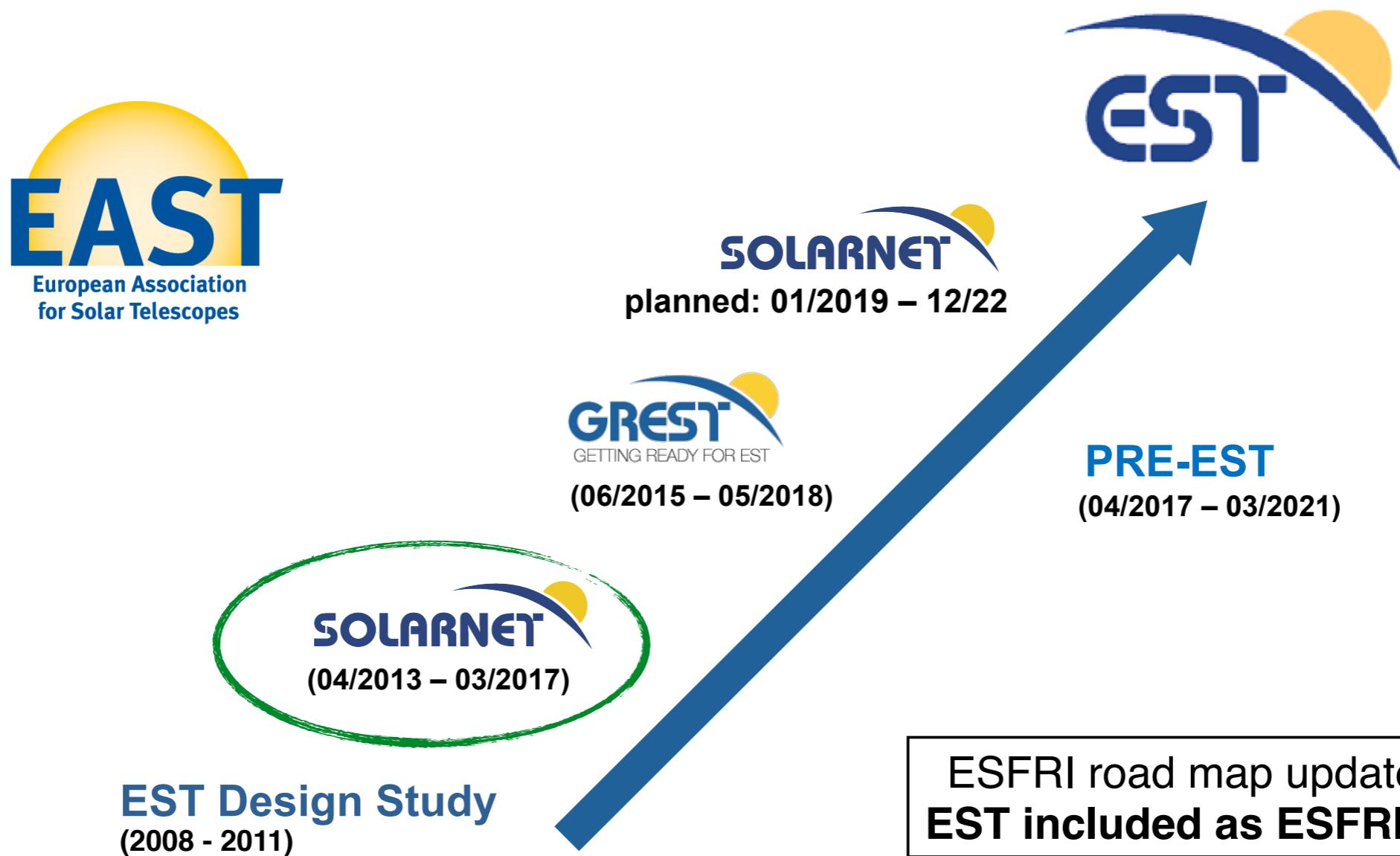
Institute		Location	
IGAM	Institutsbereich Geophysik, Astrophysik und Meteorologie		Graz
HVO	Hvar Observatory		Hvar
AIASCR	Astronomical Institute of the Academy of Sciences of the Czech Republic		
THEMIS	THEMIS		
KIS	Kiepenheuer-Institut für Sonnenphysik		
UniDeb	University of Debrecen		
INAF	Istituto Nazionale di Astrofisica		
UU	Utrecht University		
ITA	Italian Institute of Telematics		
IA UWr	Astronomical Institute of the University of Warsaw		
AISAS	Astronomical Institute of the Slovak Academy of Sciences		
IAC	Instituto de Astrofísica de Canarias		
SU	The Institute of Space and Astronautical Science		
IRSO	Istituto di Ricerca per le Scienze dello Spazio e della Terra		
UCL-MSSL	University College London - Mullard Space Science Laboratory		

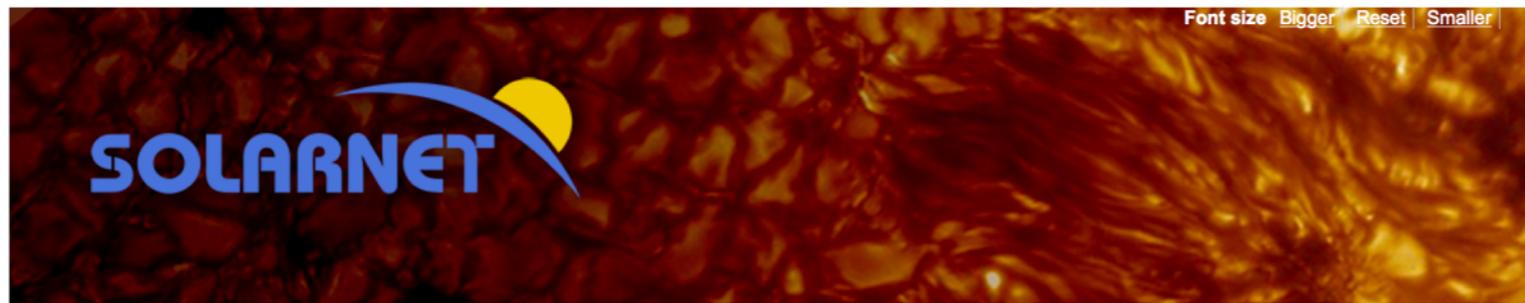
In March 2016, EST was included in the **ESFRI** (Forum for a European Strategy in Research Infrastructures) route map, together with five other projects (ACTRIS, DANUBIUS-RI, E-RIHS, EMPHASIS Y KM3NeT 2.0) and two others which are considered to be emblematic (CERN LHC Y ESRF EBS)



Canaries, Spain

The EST project is driving other projects..





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HIGHLIGHTS

Login Form

User Name

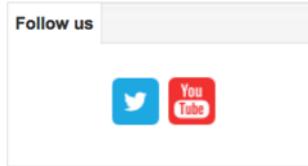
Password

Remember Me

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Latest News

- ▶ April 20, 2016. 4th SOLARNET Workshop "Solar Eruptive Events: Observations and Modelling" (London, April 20-23, 2016).
- ▶ April 13, 2016. 4th SOLARNET School Started Today in London.
- ▶ March 10, 2016. The EST in ESFRI Roadmap 2016.
- ▶ Feb. 3, 2016. SOLARNET Announcement of Opportunity. Mobility Programme of Young Researchers. DEADLINE: March 15th, 2016.

Agenda

June 2016						
M	T	W	T	F	S	S
30	31	1	2	3	4	5
6	7	8	9	10	11	12
13	14	15	16	17	18	19
20	21	22	23	24	25	26
27	28	29	30	1	2	3

Coming Events

June 20, 2016 IRIS-6 Workshop: The Solar Chromosphere
June 26, 2016 SPIE Astronomical Telescopes + Instrumentation
July 04, 2016 EWASS 2016
July 30, 2016 41st COSPAR Scientific Assembly
August 29, 2016 Partially Ionised Plasmas in Astrophysics (PIPA)
September 05, 2016 Hinode 10
September 12, 2016 Solar Polarization 8

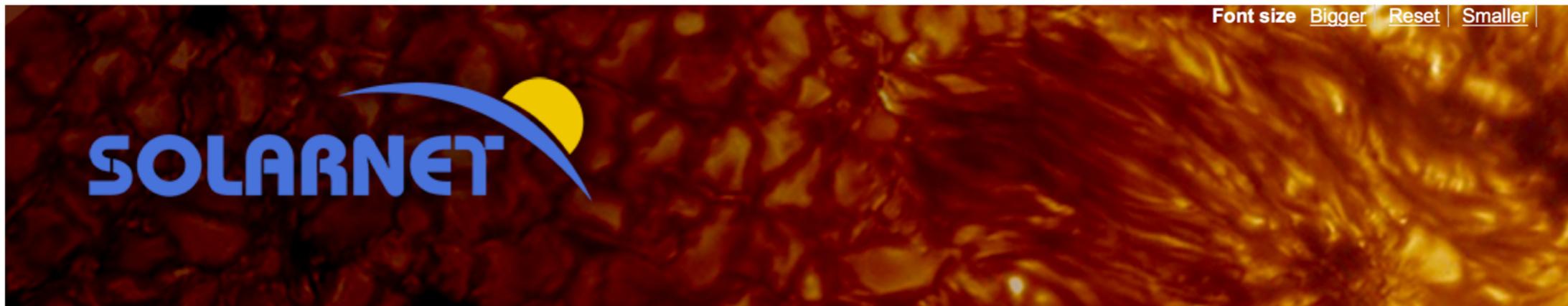
[View Full Calendar](#)

SOLARNET integrated
the major research
infrastructures in
high-res solar physics

Funded by the FP7
since 2013



This project is supported by the European Commission's FP7 Capacities Programme for the period April 2013 – March 2017 under the Grant Agreement number 312495.



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Joint Research Activities



The following joint research activities will be carried out to improve the service provided by owners/operators of research infrastructures for research on solar physics:

WP50. Tools for Innovative Data Handling: Pipelines, Databases & SVO

Develop data-reduction pipelines for the most important European ground-based high resolution solar instruments. Enhancement of observational procedures for increased productivity and easier coobserving and combination of data. The pipelines will produce data and meta-data fulfilling the requirements of a Solar Virtual Observatory (SVO). A SVO archive prototype will be implemented.

WP60. Advanced Instrumentation Development

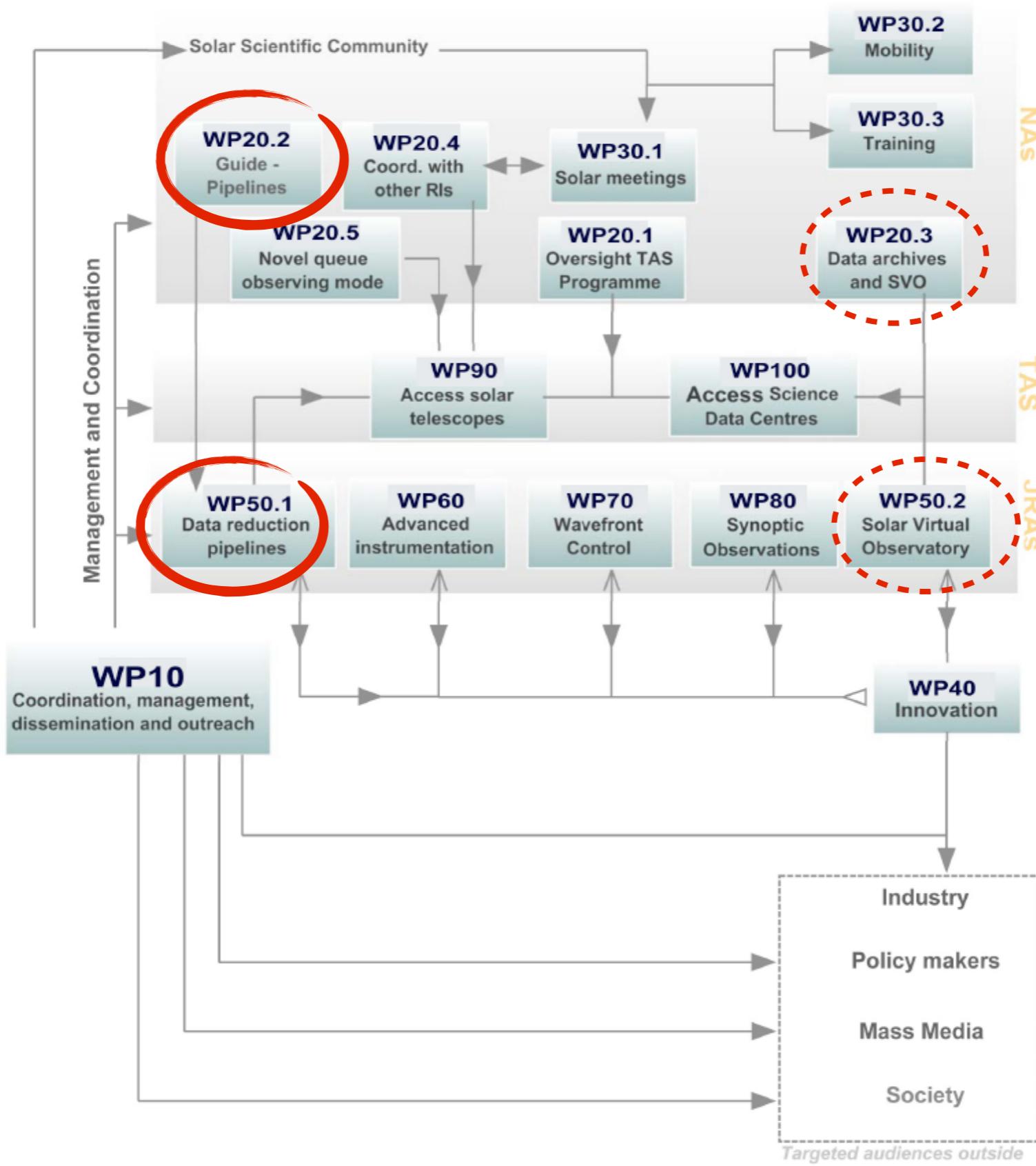
Development of instrumentation to improve the existing solar telescopes and with possible application to the future large aperture solar telescopes. The instrumentation developments included in this WP are the following: large diameter FPIs (100 to 300mm), image slicer for 2D spectroscopy, microlens-fed spectrograph and Fast Imaging Polarimeter.

HIGHLIGHTS

Login Form

User Name

WP60.2 Image Slicers for 2D spectroscopy



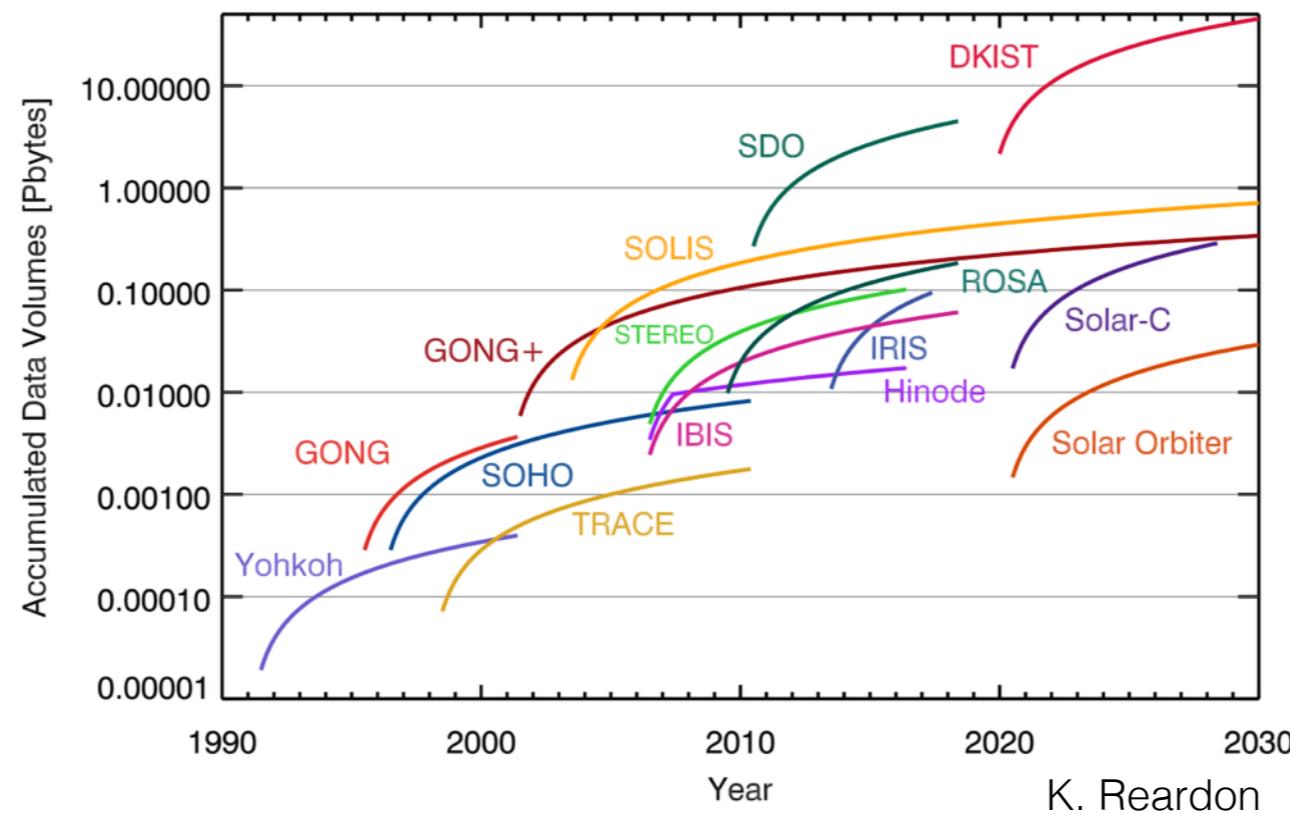
Challenges in data search & data discovery from solar observations

Traditionally, solar observation archives and VOs have been used primarily to locate data from data sets that researchers have already known existed, namely from space-based solar observatories

http://sdc.uio.no/open/solarnet-20.3/WP20.3%20Deliverable%20D20.4_v1.2.pdf

However, the number of data sets available has grown, and will continue to grow as an increasing amount of data flow from ground-based observations are made and will be available

http://sdc.uio.no/open/solarnet-20.3/WP20.3%20Deliverable%20D20.4_v1.2.pdf



K. Reardon

The use of multi-instrument analysis of solar phenomena has grown over the last decade, but the ability of solar VOs to locate multi-instrument observations has **not** grown with it

http://sdc.uio.no/open/solarnet-20.3/WP20.3%20Deliverable%20D20.4_v1.2.pdf

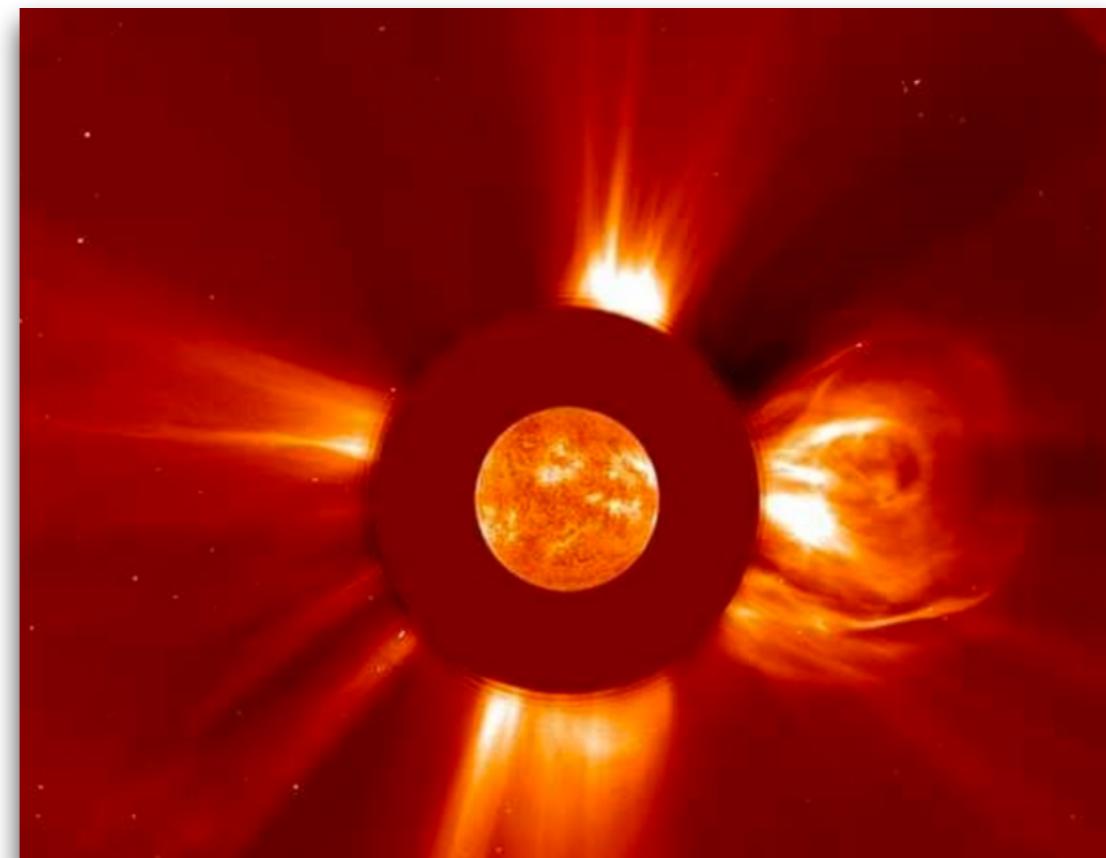


An ideal Solar Virtual Observatory (SVO) should be able to find sets of successful observations matching a hypothetical ideal observation proposal:

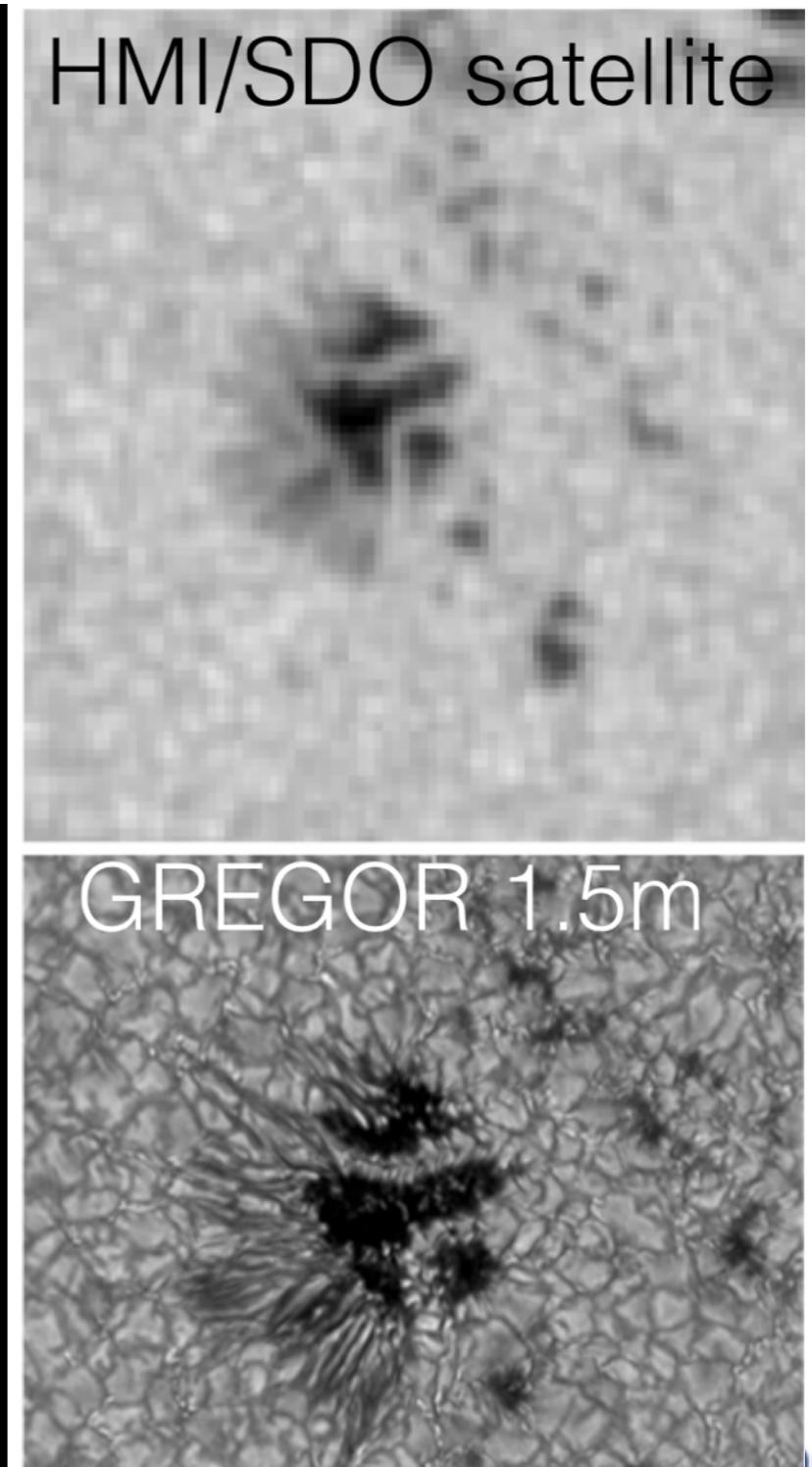
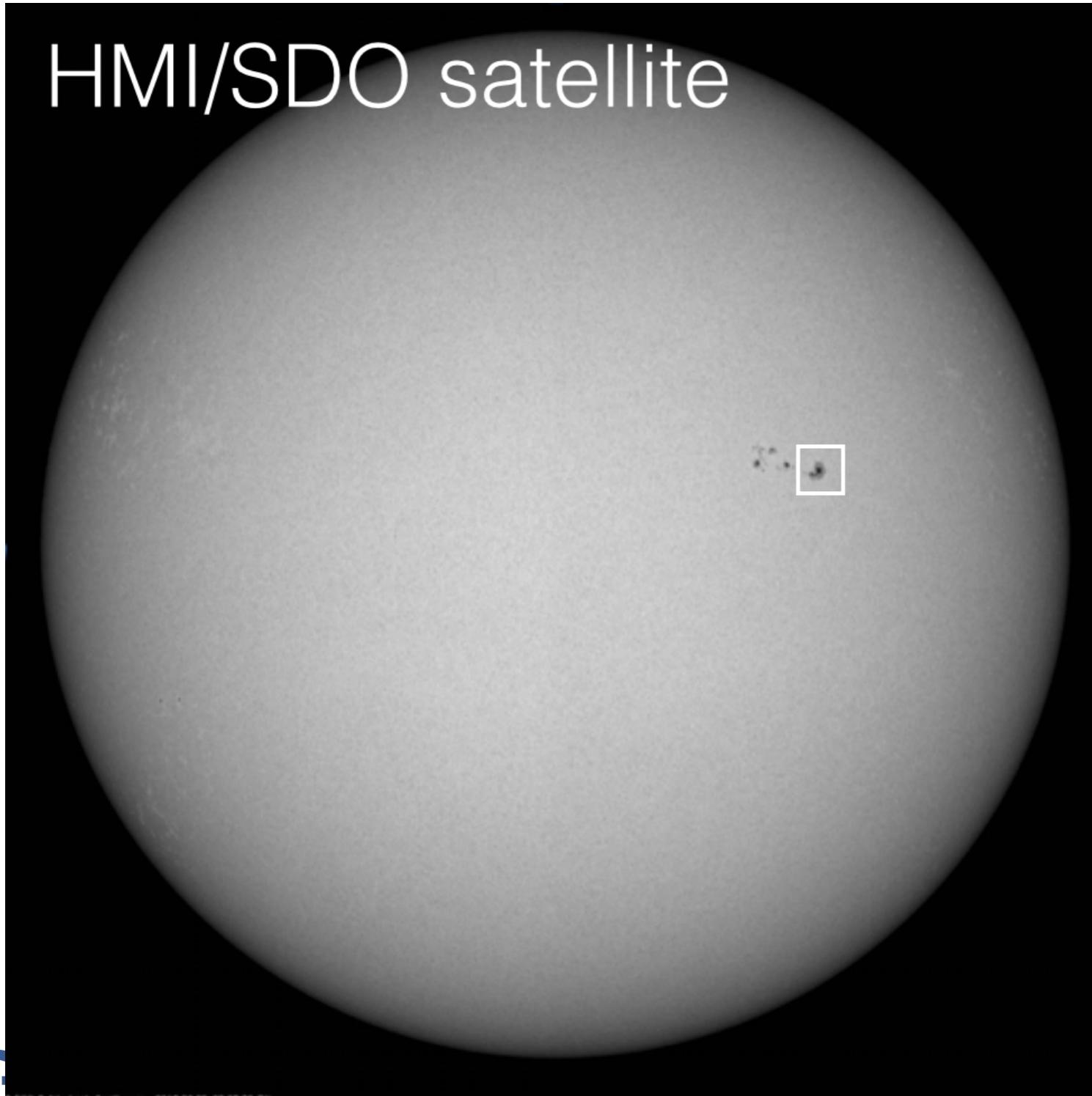
joint observations of specific targets/events from multiple instruments

Such a scenario may even involve observations that do not overlap in time, e.g., solar disc observations of events vs. *in situ* observations of particles/shocks/interactions at a later time

http://sdc.uio.no/open/solarnet-20.3/WP20.3%20Deliverable%20D20.4_v1.2.pdf

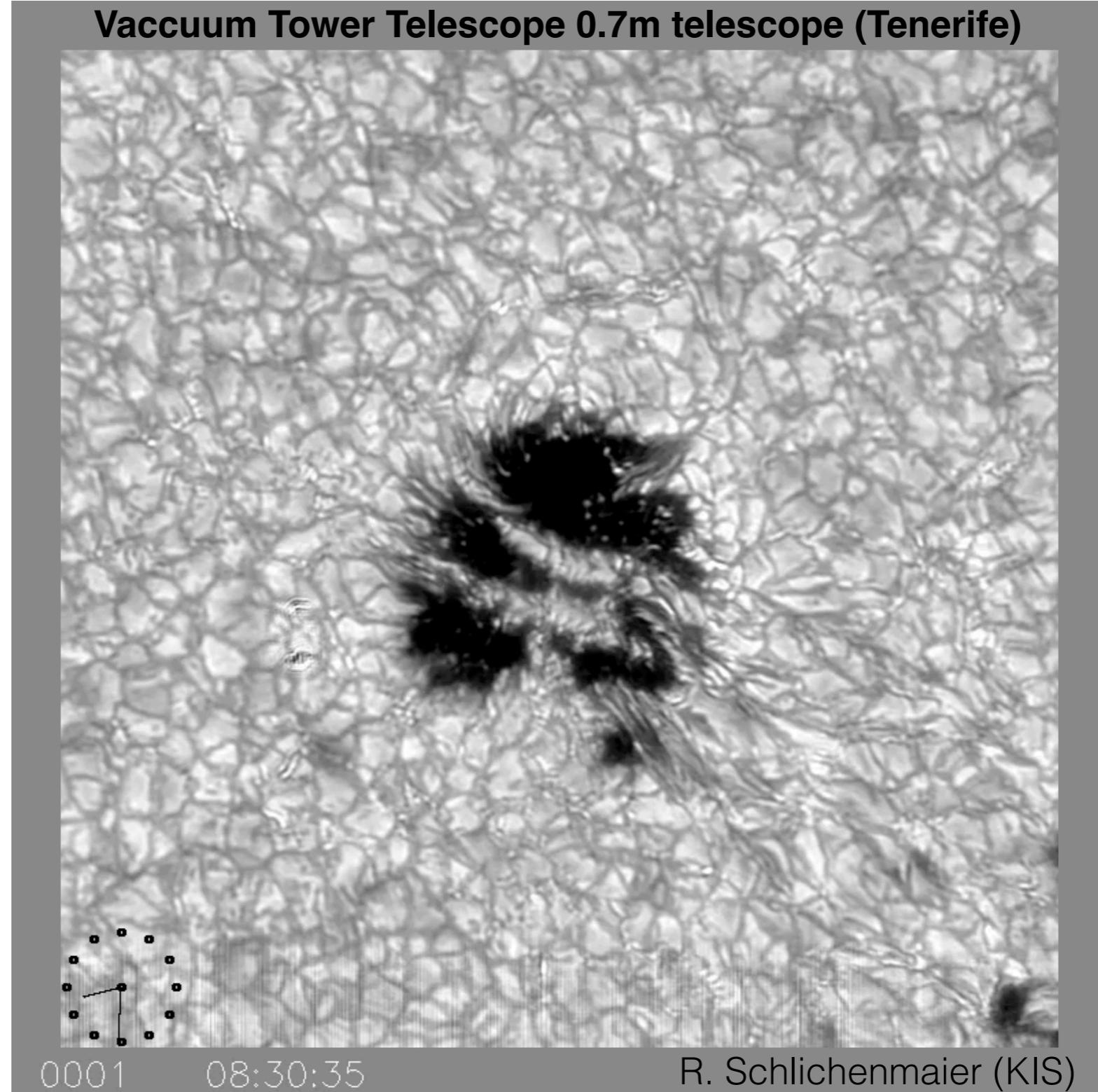
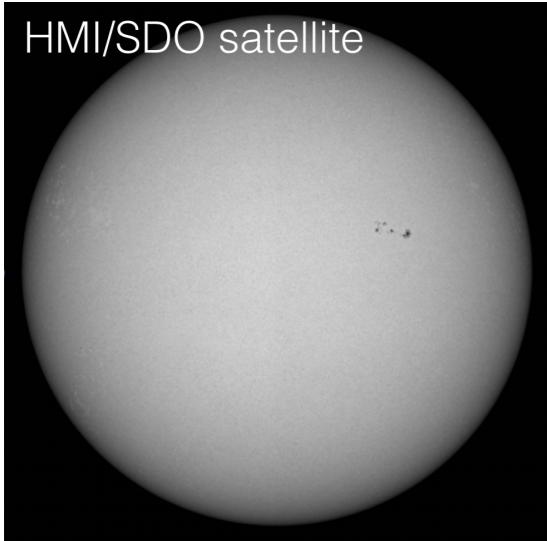


Characteristics of ground-based solar observations:

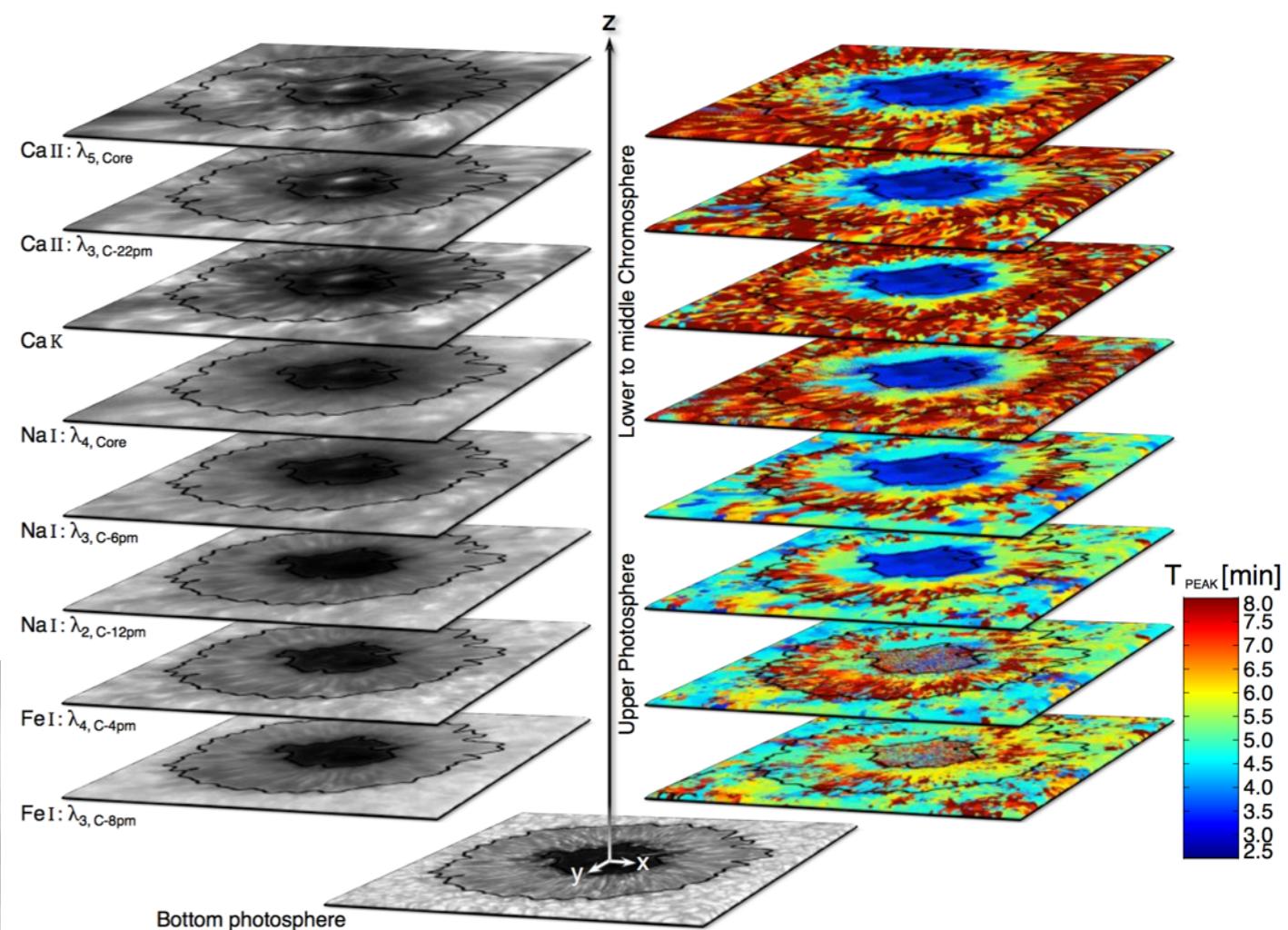
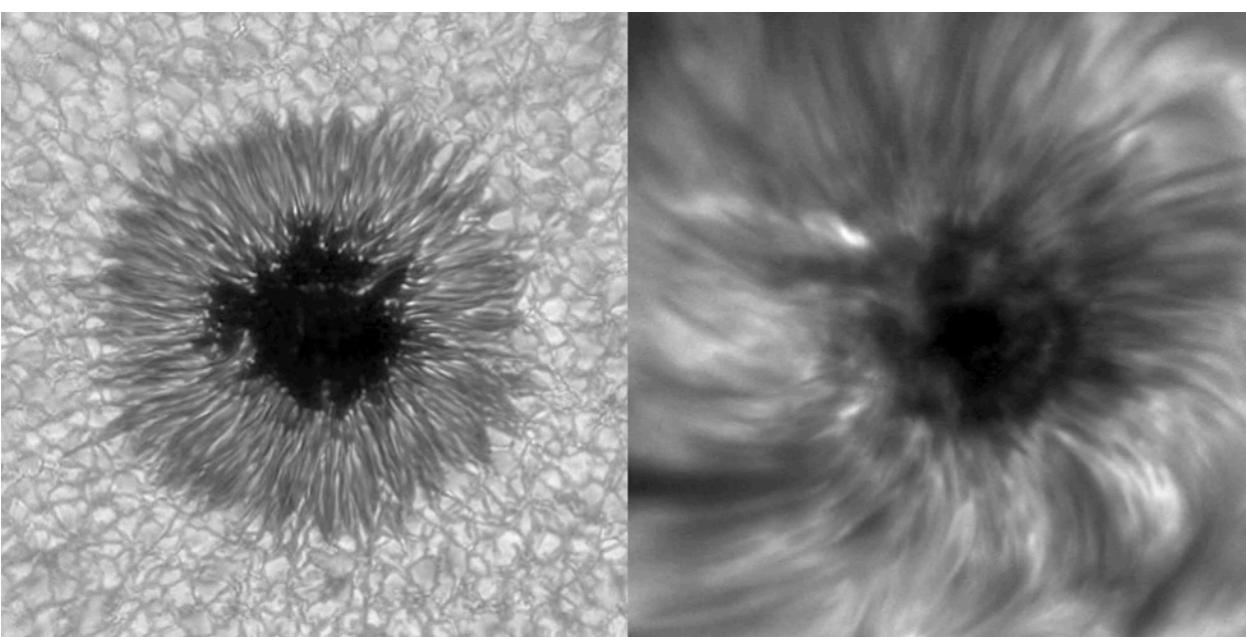
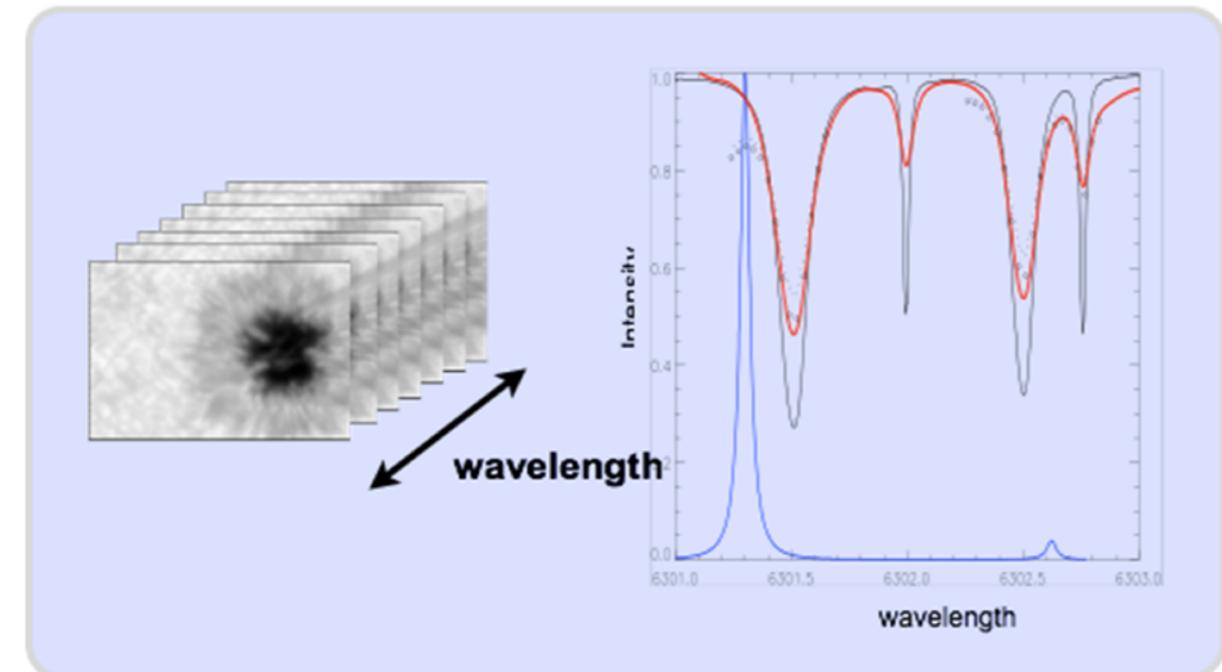


Schlichenmaier et al. (2016)

Characteristics of ground-based solar observations:

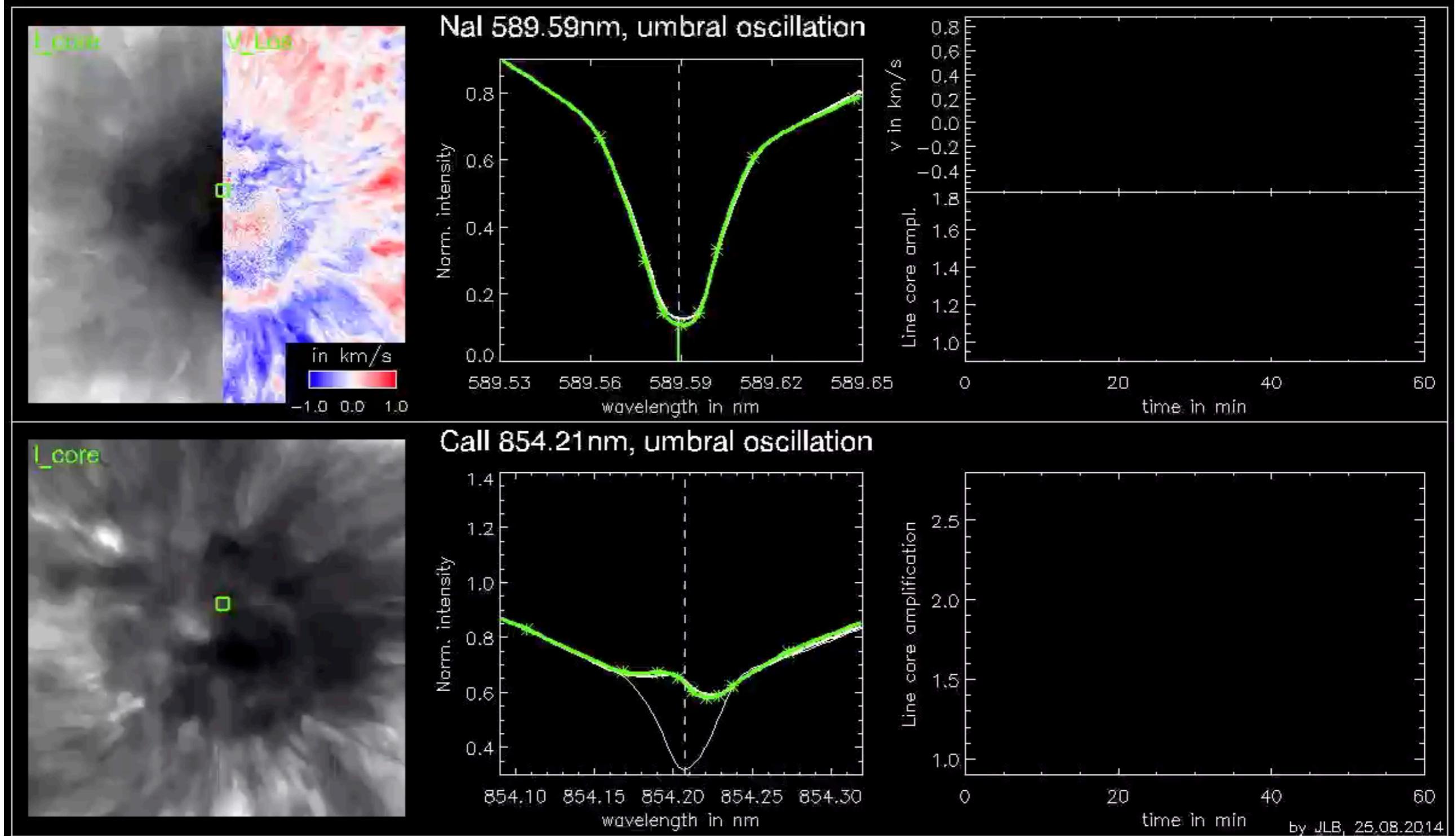


Characteristics of ground-based solar observations:



IBIS & ROSA data (J. Löhner-Bötcher, KIS)

Characteristics of ground-based solar observations:



Characteristics of ground-based solar observations:

- Limited FoV (non-full disc) — Target dependent:
quiet Sun, sunspots, pores, plages, faculae, prominences,...
- *Seeing* conditions, cadence variations, # dropped frames, polarimetric accuracy, etc. as **quality/success parameters**
- Versatile observing modes: non-standardised observing runs — novel science (multi wavelength,...) — difficulty in unifying data pipelines
- Upgrade of instrumentation — changes in data characteristics for a given (upgraded) instrument

The ideal SVO should address:

1. Efficient presentation of search results
2. Visualisation: quick-look and movies, using external existing websites
3. Type of observations, targets and events must be identified
4. Instrument specific criteria: ideally, the archive should extract generic parameters matching specific criteria

http://sdc.uio.no/open/solarnet-20.3/WP20.3%20Deliverable%20D20.4_v1.2.pdf

In order to fulfil the “vision” of an ideal SVO, it is necessary to ensure that the ground-based data to be served contains the necessary metadata

http://sdc.uio.no/open/solarnet-20.3/WP20.3%20Deliverable%20D20.4_v1.2.pdf

This is our current challenge!

Ground-based solar data archive — An example: GREGOR data

Data archive for the GREGOR Infrared Spectrograph

2017

March: 28 29
April: 02 03 04
May: 05
June: 12 13 14 16 17 18 19 20
September: 01 02 03 07 08 09 11 12 13 22 28 29
October: 02 03 30 31
November: 01

2016

May: 09 13 14 15 18 19 20 21 22 23 26 27 29 30
June: 02 04 05 06 07 08 09 10 11 12 13 14 15 16 19 20 21 22 28
July: 02 03 19 27
August: 08 10 13 14 15 16 17 19 22 23 24 25 26 29 30
September: 02 21 22 23 24 26 28 29
November: 25 28

2015

April: 15 16 17 18 19 21 23 26 27 29 30
May: 01 02 07 08 09 10 11 18 19 21 22 23 24 25 28 29 30 31
June: 01 02 03
August: 04 06 19
September: 08 09 10 12 13 14 15 16 17 19 20

2014

April: 26 27 28 29 30
May: 01 02 03 05 07 08 09 10 11 12
June: 17 18 19 20 21 22 23 24 25 26 27 28 29
July: 01 02 03 05 08 09
September: 02 03 04 05 08 10 11 13 17 18 20 22 23

Data archive for the GREGOR Infrared Spectrograph

2017

March: 28 29

April: 02 03 04

May: 05

June: 12 13 14 16 17

September: 01 02 03 07 08

October: 02 03 30 31

November: 01

[Back to main page](#) [Go to archive folder](#)

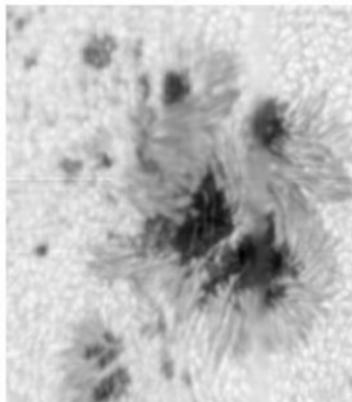
HMI context data: The arrow in the box indicates the 'slit direction', the arrow outside the box the scanning direction.

Blue (red) color of the box indicates that the GRIS scan is flipped in the scanning direction with respect to HMI (or not).

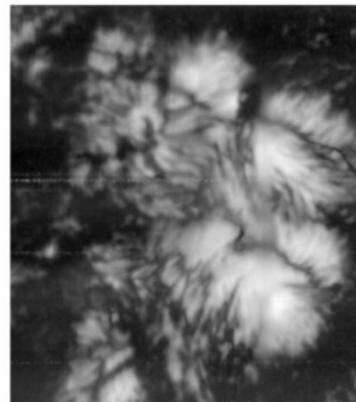
Please note that the coordinates ('x/y-pos') given in the GRIS preview images are those from the fits headers, so they are not necessarily correct.

08sep14.001

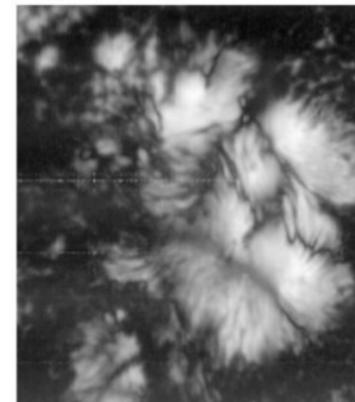
TOT I



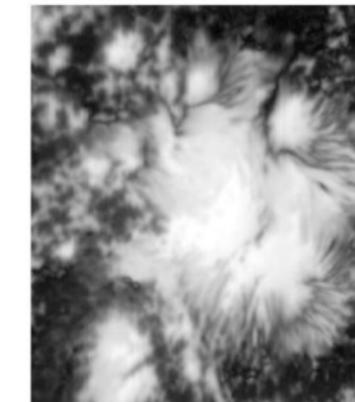
TOT Q



TOT U



TOT V



08sep14.001

1564 nm

07:56:21-08:32:07 UT

30.0 ms / 20 accum.

of steps: 400

x/y-pos: -205'' / -444''

May: 09 13 14 15 16

June: 02 04 05 06 07

July: 02 03 19 27

August: 08 10 13 14 15

September: 02 21 22 23 24

November: 25 28

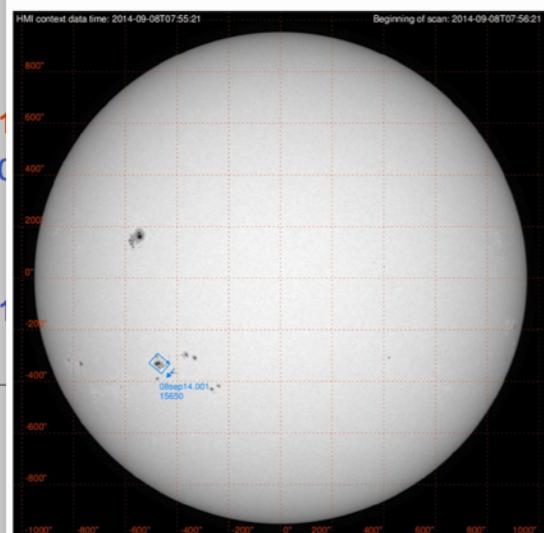
April: 15 16 17 18 19

May: 01 02 07 08 09

June: 01 02 03

August: 04 06 19

September: 08 09 10 12 13



April: 26 27 28 29 30

May: 01 02 03 05 07 08 09 10 11 12

June: 17 18 19 20 21 22 23 24 25 26 27 28 29

July: 01 02 03 05 08 09

September: 02 03 04 05 08 10 11 13 17 18 20 22 23

M. Franz, KIS

The SOLARNET VO — A prototype

<http://solarnet.oma.be/>

SOLARNET Virtual Observatory Prototype



This web server is a prototype for the [SOLARNET](#) Virtual Observatory, and is hosted currently at the Royal Observatory of Belgium

For explanations on how to search and download the data, please see our detailed [User Manual](#)

Access data via a web application

The purpose of the web application is to give a very simple access to search and download solar data. For more complex search, you are invited to use the Python or IDL clients.

The web application presents the following features:

- Cross dataset search by date of observation, wavelength, tags and telescope
- Specific dataset search that is dependant on the dataset
- Search by solar events date and type
- Co-observation searches i.e. the date of observation overlap
- Quick-look with thumbnail (if available) and FITS header
- Data selection download by FTP and ZIP (if not too large)

N.B. : This is the version 2 of the application. The version 1 corresponding to the deliverable of March 2106 is not online anymore, but the code can be found at <https://github.com/bmampaey/SDA/tree/1.1>

Access data via IDL

To search and download solar data from IDL, you will need IDL version 8.0 or higher and to download the following library on your computer [SOLARNET.pro](#)

You can then compile it and use it as in the examples in the [README](#)

Access data via Python

To search and download data from python, install the SOLARNET python library. If you have pip install, it is as simple as doing

```
pip install solarnet
```

You can then import it and use it as in the examples in the [Readme](#)

Access data via RESTful API

All metadata and data locations are available through a RESTful API. The documentation is accessible at <http://solarnet.oma.be/SDA/api/doc> If you develop tools using the API, please let us know.

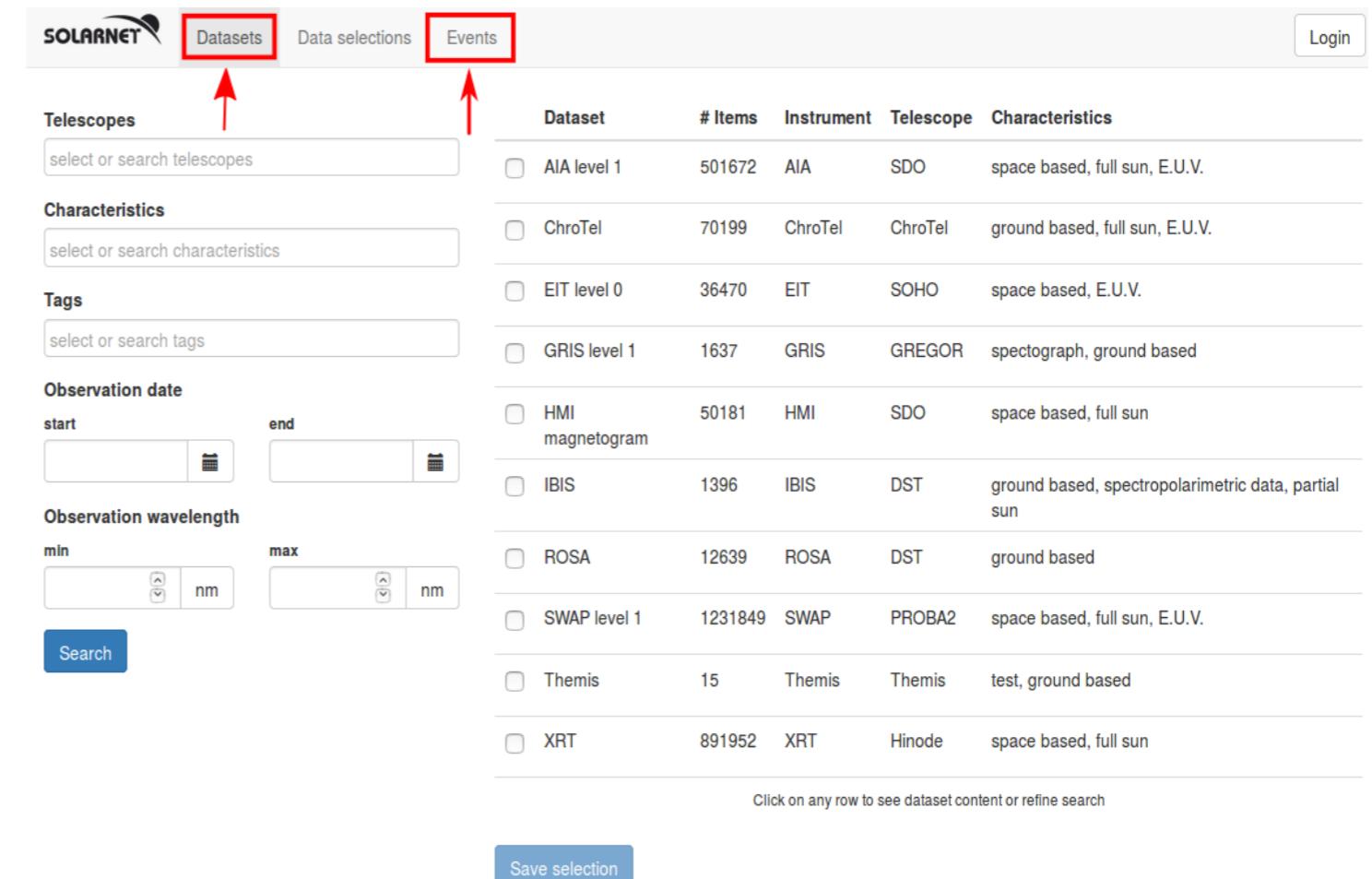
 

This project is supported by the European Commission's FP7 Capacities Programme for the period April 2013 – March 2017 under the Grant Agreement number 312495.

SOLARNET WP lead by the Royal Observatory of Belgium (Brussels)

The SOLARNET Solar Virtual Observatory

- In SOLARNET-1 a SVO prototype was developed.
- The goal of the SVO is to increase awareness of available datasets
- It's possible to search on datasets, events and then cross search your search results with other Datasets
- Some data is viewable as quick-look.
- URL: solarnet.oma.be



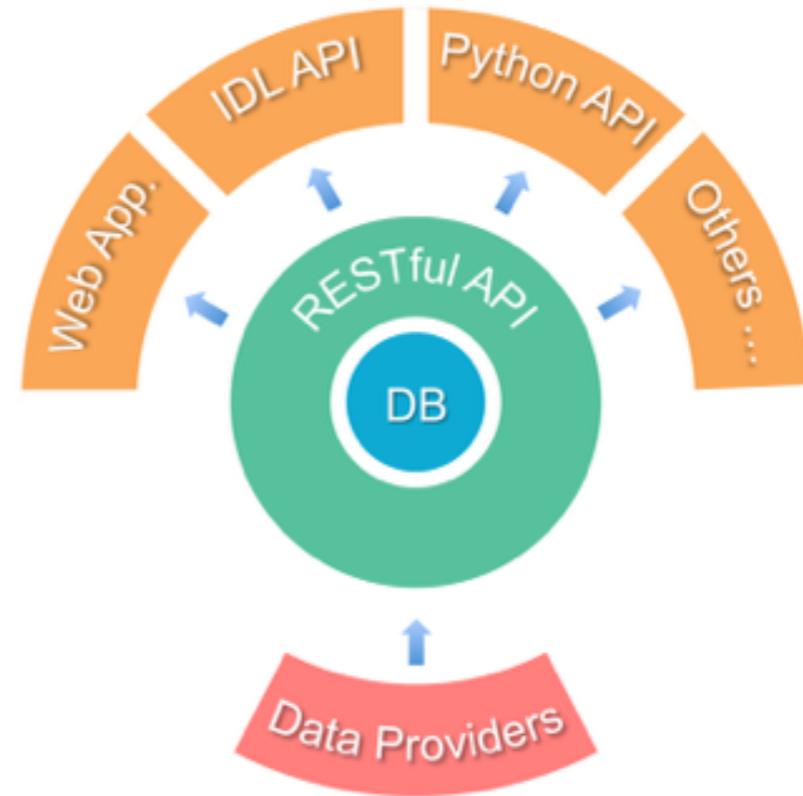
Dataset	# Items	Instrument	Telescope	Characteristics
AIA level 1	501672	AIA	SDO	space based, full sun, E.U.V.
ChroTel	70199	ChroTel	ChroTel	ground based, full sun, E.U.V.
EIT level 0	36470	EIT	SOHO	space based, E.U.V.
GRIS level 1	1637	GRIS	GREGOR	spectograph, ground based
HMI magnetogram	50181	HMI	SDO	space based, full sun
IBIS	1396	IBIS	DST	ground based, spectropolarimetric data, partial sun
ROSA	12639	ROSA	DST	ground based
SWAP level 1	1231849	SWAP	PROBA2	space based, full sun, E.U.V.
Themis	15	Themis	Themis	test, ground based
XRT	891952	XRT	Hinode	space based, full sun

Click on any row to see dataset content or refine search

Save selection

The SOLARNET Solar Virtual Observatory

- At the core of the SVO is a database containing meta-data from all datasets.
- The meta-data can be searched with the web App and IDL and Python API
- Through the RESTful API other developers can interface with the meta-database.
- The data can be downloaded from the providers server.

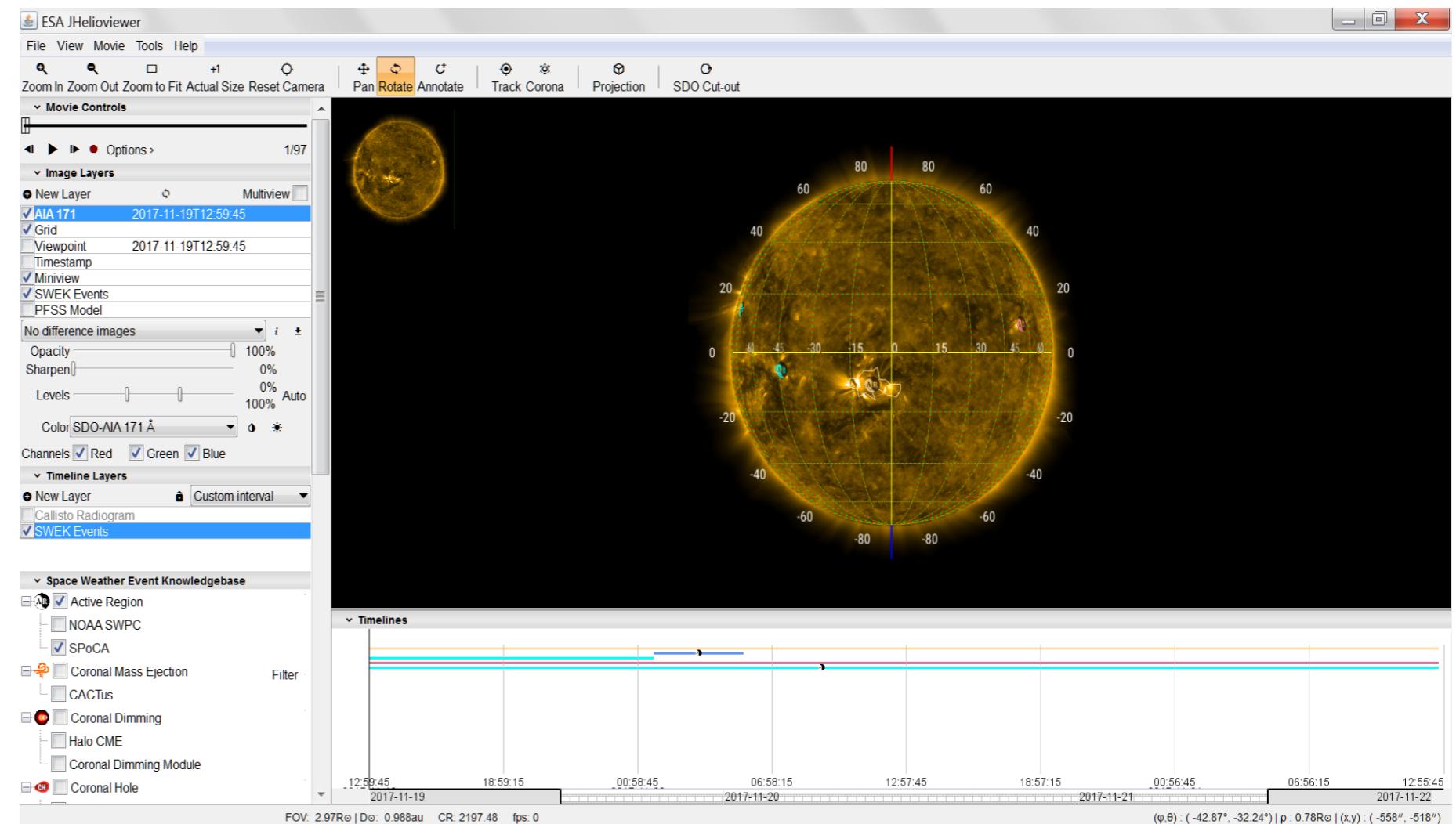


SOLARNET WP lead by the Royal Observatory of Belgium (Brussels)

Visualisation tools — Another way of data discovery

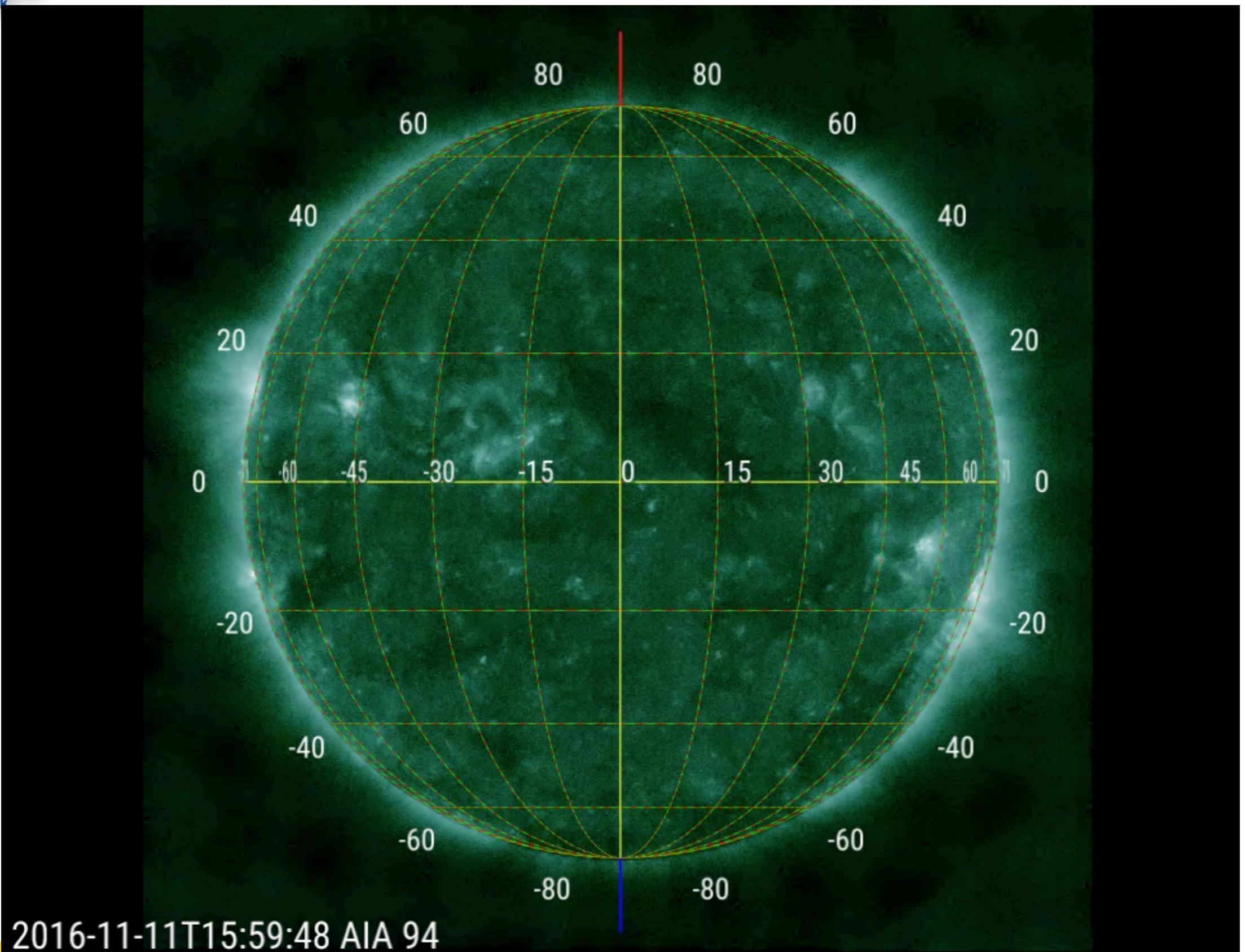
jHelioviewer a quicklook viewer for Solar data

- Access to jpeg2000 quicklook data from different dataset.
- Several viewing options like running, base difference and multiview mode
- Several Image projection
- Time line synchronized with image time series
- Showing feature events from Heliophysics Events knowledge base (HEK)
- URL:
www.jhelioviewer.org



11/20/17

JHelioviewer — Discovering data



Summary

The ground-based solar physics community is

1. Experimenting a change of paradigm on data handling, data archiving and data dissemination driven by the steady development of the observing capabilities and upcoming of large observing facilities (EST)
2. Aware of the need/the challenge on:
 - Standardising the observation procedure
 - Developing efficient data-pipelines
 - Storing the necessary metadata (time, WCS coordinates, event, wavelength..) to build a comprehensive database
3. Aware there is no need of 're-inventing the wheel' — Joint efforts with
 - The space-based solar community on data catalogue
 - The astronomical community — ASTERICS (OBELICS & DADI)