The dynamics of starvation and recovery

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The eco-evolutionary dynamics of species are fundamentally linked to the energetic constraints of its constituent individuals. Of particular importance is the interplay between reproduction and the dynamics of starvation and recovery. To elucidate this interplay, we introduce a nutritional state-structured model that incorporates two classes of consumer: nutritionally replete, reproducing consumers, and undernourished, non-reproducing consumers. We obtain strong constraints on starvation and recovery rates by deriving allometric scaling relationships and find that population dynamics are typically driven to a steady state. Moreover, these rates fall within a 'refuge' in parameter space, where the probability of population extinction is minimized. We also show that our model provides a natural framework to predict maximum mammalian body size by determining the relative stability of an otherwise homogeneous population to a competing population with altered percent body fat. This framework provides a principled mechanism for a within-lineage driver of Cope's rule.

8 the energetic state of individuals, which directly impacts • how they invest reserves in uncertain environments. Such 10 behaviors are generally manifested as tradeoffs between 11 investing in somatic maintenance and growth, or allocat- $_{12}$ ing energy towards reproduction $^{1-3}$. The timing of these 13 behaviors responds to selective pressure, as the choice of the investment impacts future fitness⁴⁻⁶. The influence 15 of resource limitation on an organism's ability to main-16 tain its nutritional stores may lead to repeated delays or 17 shifts in reproduction over the course of an organism's

The balance between (a) somatic growth and mainte-20 nance, and (b) reproduction depends on resource avail-²¹ ability⁷. For example, reindeer invest less in calves born 22 after harsh winters (when the mother's energetic state is 23 depleted) than in calves born after moderate winters⁸. 24 Many bird species invest differently in broods during pe-²⁵ riods of resource scarcity compared to normal periods ^{9,10}, 26 sometimes delaying or even foregoing reproduction for 27 a breeding season^{1,11,12}. Even freshwater and marine 28 zooplankton have been observed to avoid reproduction ²⁹ under nutritional stress¹³, and those that do reproduce 30 have lower survival rates². Organisms may also separate 31 maintenance and growth from reproduction over space 32 and time: many salmonids, birds, and some mammals 33 return to migratory breeding grounds to reproduce af-34 ter one or multiple seasons in resource-rich environments 35 where they accumulate reserves $^{14-16}$.

Physiology also plays an important role in regulating 37 reproductive expenditures during periods of resource lim-38 itation. Many mammals (47 species in 10 families) ex-39 hibit delayed implantation, whereby females postpone fe-40 tal development until nutritional reserves can be accumu-41 lated ^{17,18}. Many other species (including humans) suf-42 fer irregular menstrual cycling and higher abortion rates 80 getically deficient (hungry) state H, where the consumer 43 during periods of nutritional stress 19,20 . In the extreme 81 does not reproduce but dies by starvation at rate μ . The 44 case of unicellular organisms, nutrition directly controls $_{52}$ dynamics of the underlying resource R are governed by

The behavioral ecology of all organisms is influenced by 45 growth to a reproductive state^{3,21}. The existence of so 46 many independently evolved mechanisms across such a 47 diverse suite of organisms highlights the near-universality 48 of the fundamental tradeoff between somatic and repro-49 ductive investment.

> Including individual energetic dynamics²² in a 51 population-level framework^{22,23} is challenging²⁴, and a 52 common simplifying approach is provided in the clas-53 sic Lotka-Volterra (LV) model, which assumes that 54 consumer population growth rate depends linearly on 55 resource density 25 . Here, we introduce an alterna-56 tive approach—the Nutritional State-structured Model 57 (NSM)—that accounts for resource limitation via explicit 58 starvation. In contrast to the LV model, the NSM in-59 corporates two consumer states: hungry and full, with 60 only the former susceptible to mortality and only the lat-61 ter possessing sufficient energetic reserves to reproduce. 62 Additionally, we incorporate allometrically derived con-63 straints on reproduction³, incorporating the timescales 64 of starvation and recovery. As we shall show, our model 65 makes several important predictions: (i) the dynamics 66 are typically driven to a refuge far from cyclic behavior 67 and extinction risk, (ii) steady state conditions of the 68 NSM accurately predict the measured biomass densities 69 for mammals described by Damuth's law, (iii) there is an 70 allometrically constrained upper-bound for mammalian 71 body size, and (iv) the NSM provides a selective mech-72 anism for the evolution of larger body size, known as 73 Cope's rule.

> Nutritional state-structured model (NSM). 75 We begin by defining the nutritional state-structured 76 population model, where the consumer population is par-77 titioned into two states: (a) an energetically replete (full) ₇₈ state F, where the consumer reproduces at a constant 79 rate λ and does not die from starvation, and (b) an ener-

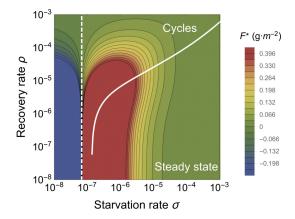


Figure 1: The transcritical (TC; dashed line) and Hopf bifurcation (solid line) as a function of the starvation rate σ and recovery rate ρ for a 100g consumer. These bifurcation conditions separate parameter space into unphysical (left of the TC), cyclic, and steady state dynamic regimes. The colors show the steady state densities for the energetically replete consumers F^* .

83 logistic growth with an intrinsic growth rate α and a car-84 rying capacity C. The rate at which consumers transi-85 tion between states and consume resources is dependent 86 on their number, the abundance of resources, the effi-87 ciency of converting resources into metabolism, and how 88 that metabolism is partitioned between maintenance and 89 growth purposes. We provide a physiologically and enof ergetically mechanistic model for each of these dynam-91 ics and constants (see the Supplementary Information 92 (SI), and show that the system produces a simple non-93 dimensional form which we describe below.

Consumers transition from the full state F to the hun-95 gry state H at a rate σ —the starvation rate—and also 96 in proportion to the absence of resources (1-R) (we 97 assume a maximum density, the carrying capacity equal 98 to 1). Conversely, consumers recover from state H to previous models incorporating starvation²⁹.

114 sumers and resources are perfectly mixed, their densities 166 reproducing state. Because reproduction is inhibited, 115 are governed by the rate equations

$$\dot{F} = \lambda F + \xi \rho R H - \sigma (1 - R) F,
\dot{H} = \sigma (1 - R) F - \xi \rho R H - \mu H,
\dot{R} = \alpha (1 - R) R - (\rho R + \delta) H - \beta F$$
(1)

This system of nondimensional equations follows from a set of first-principle relationships for resource consumption and growth (see the SI for a full derivation and the 119 dimensional form). Notice that the total consumer den-120 sity F + H evolves according to $\dot{F} + \dot{H} = \lambda F - \mu H$. This 121 resembles the equation of motion for the predator den-122 sity in the LV model³⁰, except that the resource density 123 does not appear in the growth term. The rate of repro-124 duction is independent of resource density because it is 125 assumed that the satiated state of the full consumer al-126 lows it to partition a constant amount of energy towards 127 reproduction, whereas a starved consumer partitions no 128 energy towards reproduction. The rate of reproduction 129 for the total consumer density is dependent on resource 130 density, which determines the size of the full and starved portions of the consumer population. Similarly, the con-132 sumer maintenance terms (δH and βF) are independent 133 of resource density because they represent a minimal en-134 ergetic requirement for consumers in the H and F state, 135 respectively.

Steady states of the NSM. From the solution to

137 the single internal fixed point (Eq. (2), see Methods), an 138 obvious constraint on the NSM is that the reproduction 139 rate λ must be less than the starvation rate σ , so that the 140 consumer and resource densities are positive. The condition $\sigma = \lambda$ represents a transcritical (TC) bifurcation³¹ that demarcates a physical from an unphysical (negative 143 steady-state densities) regime. The biological implication 144 of the constraint $\lambda < \sigma$ has a simple interpretation—the 145 rate at which a macroscopic organism loses mass due to $_{146}$ lack of resources is generally much faster than the rate 147 of reproduction. As we will discuss below, this inequal-148 ity is a natural consequence of allometric constraints³ for 149 organisms within empirically observed body size ranges. 99 state F at rate $\xi \rho$ and in proportion to R, where ξ rep. 150 In the physical regime of $\lambda < \sigma$, the fixed point (2) 100 resents a ratio between maximal resource consumption 151 may either be a stable node or a limit cycle (Fig. 1). In 101 and the carrying capacity of the resource. The resources 152 continuous-time systems, a limit cycle arises when a pair that are eaten by hungry consumers (at rate $\rho R + \delta$) action of complex conjugate eigenvalues crosses the imaginary count for their somatic growth (ρR) and maintenance (δ) . 154 axis to attain positive real parts³². This Hopf bifurcation Full consumers eat resources at a constant rate β that ac-105 counts for maximal maintenance and somatic growth (see 156 which is composed of the coefficients of the characteris-106 the SI for mechanistic derivations of these rates from re- 157 tic polynomial of the Jacobian matrix³³. As the system 107 source energetics). The NSM represents an ecologically 158 parameters are tuned to be within the stable regime, but notivated fundamental extension of the idealized starv- 159 close to the Hopf bifurcation, the amplitude of the tran-100 ing random walk model of foraging, which focuses on re- 160 sient cycles becomes large. Given that ecological systems 110 source depletion, to include reproduction and resource re- 161 are constantly being perturbed³⁴, the onset of transient plenishment ^{26–28}, and is a more general formulation than ¹⁶² cycles, even though they decay with time in the mean-163 field description, can increase extinction $risk^{35-37}$.

When the starvation rate $\sigma \gg \lambda$, a substantial In the mean-field approximation, in which the con- 105 fraction of the consumers are driven to the hungry non-167 there is a low steady-state consumer density and a high

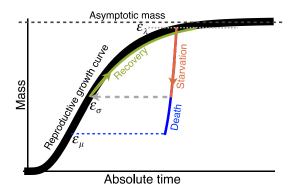


Figure 2: The growth trajectory over absolute time of an intrajectory to $m = \epsilon_{\sigma} \epsilon_{\lambda} M$. Recovery follows the green curve to the replete adult mass, where this trajectory differs from the original growth because only fat is being regrown which requires a longer time to reach $\epsilon_{\lambda}M$. Alternatively, death from starvation follows the blue trajectory to $m = \epsilon_{\mu} \epsilon_{\lambda} M$.

168 steady-state resource density. However, if $\sigma/\lambda \to 1$ from $\,^{\,\mathbf{226}}$ above, the population is overloaded with energetically- 227 reveal (see Methods), starvation and recovery rates are 170 replete (reproducing) individuals, thereby promoting 228 not independent parameters, and the biologically rele-171 transient oscillations between the consumer and resource 229 vant portion of the phase space shown in Fig. 1 is con-172 densities (Fig. 1). If the starvation rate is low enough 230 strained via covarying parameters. Given the parameters 173 that the Hopf bifurcation is crossed, these oscillations 231 of terrestrial endotherms, we find that the starvation rate 174 become stable over time. 175 higher values of the starvation rate as the recovery rate 233 a small region of potential values for the known range $_{176}$ ρ increases, such that the range of parameter space $_{234}$ of body sizes M. Indeed, starvation and recovery rates $_{177}$ giving rise to cyclic dynamics also increases with higher $_{235}$ across all values of M fall squarely in the steady state 178 recovery rates.

180 Results

The allometry of extinction risk. While there are 239 ing. 194 (cf. ref. 38).

202 recovery dynamics as a function of an organism's body 203 mass M, we parameterize these rates in terms of the per-204 cent gain and loss of the asymptotic (maximum) body 205 mass, ϵM , where different values of ϵ define different 206 states of the consumer (Fig. 2; see Methods for deriva-207 tions of allometrically constrained rate equations). Al-208 though the rate equations (1) are general and can in 209 principle be used to explore the starvation recovery dy-210 namics for most organisms, here we focus on allomet-211 ric relationships for terrestrial-bound lower trophic level 212 endotherms (see the SI for values), specifically herbivo-213 rous mammals, which range from a minimum of $M \approx 1$ g 214 (the Etruscan shrew Suncus etruscus) to a maximum of 215 $M \approx 10^7 \text{g}$ (the early Oligocene Indricotheriinae and the 216 Miocene Deinotheriinae). Investigating other classes of dividual organism as a function of body mass. Initial growth 217 organisms would simply involve altering the metabolic follows the black trajectory to an energetically replete repro- 218 exponents and scalings associate with ϵ . Moreover, we ductive adult mass $m = \epsilon_{\lambda} M$. Starvation follows the red 219 emphasize that our allometric equations describe mean 220 relationships, and do not account for the (sometimes 221 considerable) variance associated with individual species. 222 We note that including additional allometrically-scaled 223 mortality terms to both F and H does not change the 224 form of our model nor impact our quantitative findings 225 (see SI for derivation).

> As the allometric derivations of the NSM rate laws This threshold occurs at 232 σ and the recovery rate ρ are constrained to lie within 236 region at some distance from the Hopf bifurcation. This 237 suggests that cyclic population dynamics should be rare, 238 particularly in environments where resources are limit-

182 no a priori constraints on the parameters in the NSM, 240 Higher rates of starvation result in a larger flux of the 183 we expect that each species should be restricted to a dis- 241 population to the hungry state. In this state, reproduc-184 tinct portion of the parameter space. We use allomet- 242 tion is absent, thus increasing the likelihood of extinc-185 ric scaling relations to constrain the covariation of rates 243 tion. From the perspective of population survival, it is 186 in a principled and biologically meaningful manner (see 244 the rate of starvation relative to the rate of recovery that 187 Methods). Allometric scaling relations highlight com- 245 determines the long-term dynamics of the various species 188 mon constraints and average trends across large ranges 246 (Fig. 1). We therefore examine the competing effects of 189 in body size and species diversity. Many of these rela- 247 cyclic dynamics vs. changes in steady-state density on 100 tions can be derived from a small set of assumptions and 248 extinction risk, both as functions of σ and ρ . To this 191 in the Methods we describe our framework to determine 249 end, we computed the probability of extinction, where we 102 the covariation of timescales and rates across a range of 250 define extinction as a population trajectory falling below 193 body sizes for each of the key parameters of our model 251 one fifth of the allometrically constrained steady state 252 at any time between $t = 10^8$ and $t = 10^{10}$. This proce-Nearly all of the rates described in the NSM are de-253 dure was repeated for 50 replicates of the continuous-time 106 termined by consumer metabolism, which can be used 254 system shown in Eq. 1 for organisms with mass ranging 197 to describe a variety of organismal features³⁹. We derive 255 from 10² to 10⁶ grams. In each replicate the initial den-1968 relationships for the rates of reproduction, starvation, re- 256 sities were chosen to be (XF^*, XH^*, R^*) , with X a ran-199 covery, and mortality based on first principles, and as a 257 dom variable uniformly distributed in [0, 2]. By allowing 200 function of an organism's body size and metabolic rate 258 the rate of starvation to vary, we assessed extinction risk ₂₀₁ (see Methods). Because we aim to explore the starvation-₂₅₉ across a range of values for σ and ρ between ca. 10^{-8}

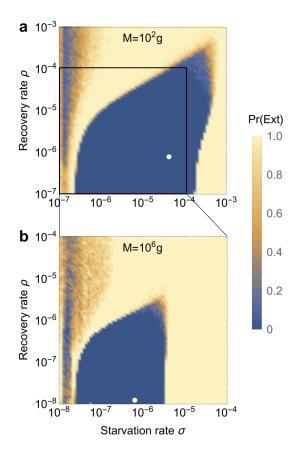


Figure 3: Probability of extinction for a consumer with (a) $M = 10^2$ g and (b) $M = 10^6$ g as a function of the starvation rate σ and recovery rate ρ , where the initial density is given as (XF^*, XH^*, R^*) , where X is a random uniform variable in [0,2]. Note the change in scale for $M=10^4$ and $M=10^4$ 10⁶g. Extinction is defined as the population trajectory falling below $0.2\times$ the allometrically constrained steady state. The white points denote the allometrically constrained starvation and recovery rate.

 $_{264}$ transient cycles as the system nears the Hopf bifurcation $_{322}$ associated with each clade. 265 (Fig. 3). For high values of σ and low values of ρ , in- $_{266}$ creased extinction risk arises because of the decrease in 267 the steady-state consumer population density (Figs. 1b, 268 3). This interplay creates an 'extinction refuge', such that for a constrained range of σ and ρ , extinction prob-270 abilities are minimized.

₂₇₂ σ and ρ , each representing different trajectories along ₃₃₀ malian body size is given by $\epsilon_{\sigma}=0$, where mammals 273 the ontogenetic curve (Fig. 2), fall squarely within the 331 are entirely composed of metabolic reserves, and this oc-274 extinction refuge across a range of M (Fig. 3a,b, white 332 curs at $M = 8.3 \times 10^8$ (g), or 120 times the mass of a 275 points). These values are close enough to the Hopf bi- 333 male African elephant. In contrast, the $\mu=0$ asymptote 276 furcation to avoid low steady-state densities, yet distant 334 occurs first when $f_0 M^{\gamma-1} + u_0 M^{\zeta-1} = 1$, and corre-277 enough to avoid large-amplitude transient cycles. That 335 sponds to $(F^*, H^*, R^*) = (0, 0, 0)$. This point predicts 278 allometric values of σ and ρ fall within this relatively 336 a more realistic upper bound on mammalian body size

279 small window supports the possibility that a selective 280 mechanism has constrained the physiological conditions 281 driving starvation and recovery rates within populations. 282 Such a mechanism would select for organism physiology 283 that generates appropriate σ and ρ values that minimize 284 extinction risk. This selection could occur via the tuning 285 of body fat percentages, metabolic rates, and/or biomass 286 maintenance efficiencies. We also find that as body size 287 increases, the amount of low extinction risk parameter 288 space becomes smaller (Fig. 3b), suggesting that the 289 population dynamics for larger organisms are more sen-290 sitive to variability in physiological rates. This finding 291 is in accordance with, and may serve as contributing 292 support for, observations of increased extinction risk 293 among larger mammals 40 .

Predicting Damuth's Law and body size limits.

296 The NSM correctly predicts that smaller species have 297 larger steady-state population densities (Fig. 4). Simi-298 lar predictions have been made for carnivore populations 299 using alternative consumer-resource models⁴¹. Moreover, 300 we show that the NSM provides independent theoretical 301 support for Damuth's Law^{42–45}. Damuth's law reveals that species abundances, N^* , follow $N^* \propto M^{-0.78}$. Fig-303 ure 4 shows that both F^* and H^* scale as $M^{-\eta}$ over a 304 wide range of organismal sizes and that $F^* + H^*$ closely 305 matches the best fit to Damuth's data. This result is 306 remarkable because it illustrates that the steady state 307 values of the NSM combined with the derived timescales 308 naturally give rise to Damuth's law. While the previous 309 metabolic studies supporting Damuth's law provided ar-310 guments for the value of the exponent⁴³, these studies 311 are only able to infer the intercept from the data (see SI 312 for a discussion of the energy equivalence hypothesis re-313 lated to these metabolic arguments). Our model predicts and not only the exponent but also the intercept by explicitly 315 including the resource dynamics and the parameters ad-316 justing growth and consumption. It should be noted that 317 density relationships of individual clades follow a more $_{260}$ to 10^{-3} . Higher rates of extinction correspond to both $_{318}$ shallow scaling relationship than predicted by Damuth's high values of σ if ρ is small, and high values of ρ if σ 319 law⁴⁵. In the context of our model, this suggests that 262 is small. For low values of σ and high values of ρ , the 320 future work may be able to anticipate these shifts by ac-263 increased extinction risk results from higher-amplitude 321 counting for differences in the physiological parameters

With respect to predicted steady state densities, the 324 total metabolic rate of F and H becomes infinite at a fi-325 nite mass, and occurs at the same scale where the steady 326 state resources vanish (Fig. 4). This asymptotic behavior 327 is governed by body sizes at which ϵ_{μ} and ϵ_{λ} equal zero, 328 causing the timescales to become infinite and the rates μ We find that the allometrically constrained values of 329 and λ to equal zero. A theoretical upper bound on mam-

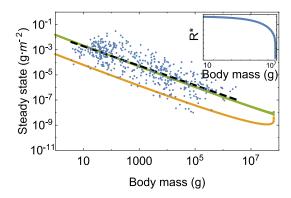


Figure 4: Consumer steady states F^* (green) and H^* (orange) as a function of body mass along with the data from Damuth⁴². Inset: Resource steady state R^* as a function of consumer body mass.

337 and occurs at $M_{\rm max}=6.54\times 10^7$ (g). Moreover, $M_{\rm max},$ 338 which is entirely determined by the population-level con-339 sequences of energetic constraints, is within an order of 340 magnitude of the maximum body size observed in the 341 North American mammalian fossil record 46 , as well as 352 els that incorporate the allometry of hunting and resting 386 explored. 353 combined with foraging time predicts a maximum carni- 387 The NSM predicts that the steady state resource 354 vore size between 7×10^5 and 1.1×10^6 (g)^{49,50}. Similarly, 388 density R^* decreases with increasing body size of the 355 the maximum body size within a particular lineage has 389 consumer population (Fig. 4, inset), and classic re-356 been shown to scale with the metabolic normalization 390 source competition theory predicts that the species sur-357 constant⁵¹. This complementary approach is based on 391 viving on the lowest resource abundance will outcom-358 the balance between growth and mortality, and suggests 392 pete others 58-60. Thus, the combined NSM steady-state 359 that future connections between the scaling of fat and 393 dynamics and allometric timescales predict that larger 366 muscle mass should systematically be connected with B_0 394 mammals have an intrinsic competitive advantage given 361 when comparing lineages.

363 port constraints are widely thought to place strict bound- 397 or against. aries on biological scaling 39,52,53 and thereby lead to 398 We directly assess competitive outcome between two

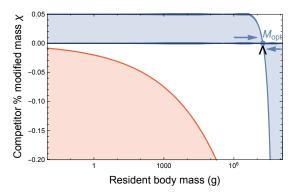


Figure 5: Competitive outcomes for a resident species with body mass M vs. a closely related competing species with modified body mass $M' = M(1 + \chi)$. The blue region denotes proportions of modified mass χ resulting in exclusion of the resident species. The red region denotes values of χ that result in a mass that is below the starvation threshold and are thus infeasible. Arrows point to the predicted optimal mass from our model $M_{\rm opt} = 1.748 \times 10^7$, which may serve as an evolutionary attractor for body mass. The black wedge points to the largest body mass known for terrestrial mammals (Deinotherium spp.) at 1.74×10^7 (g)⁵⁶.

342 the mass predicted from an evolutionary model of body 376 lineages show a significant trend towards larger body size 343 size evolution⁴⁷. It should be noted that the asymptotic 377 known as Cope's rule^{46,47,56,57}, and it is thought that behavior and predicted upper bound depend only on the 378 within-lineage drivers generate selection towards an opscaling of body composition and are independent of the 379 timal upper bound of roughly 10⁷ (g)⁴⁶, a value that 346 resource parameters. We also note that the prediction of 380 is likely limited by higher extinction risk for large taxa an asymptotic limit on mammalian size parallels work on 381 over longer timescales 47. These trends are thought to 348 microbial life where an upper and lower bound on bac- 382 be driven by a combination of climate change and niche terial size, and an upper bound on single cell eukaryotic 383 availability⁵⁷; however the underpinning energetic costs 350 size, is predicted from similar growth and energetic scal- 384 and benefits of larger body sizes, and how they influ-351 ing relationships^{3,48}. It has also been shown that mod-365 ence dynamics over ecological timescales, have not been

 $_{\tt 395}$ a common resource, but these absolute limits do not offer A mechanism for Cope's rule Metabolite trans- 396 a mechanism by which larger body sizes are selected for

365 specific predictions on the minimum possible body size 300 closely related species: a resident species of mass M, 366 for organisms⁵⁴. Above this bound, a number of ener-400 and a competing species (denoted by ') where individ-367 getic and evolutionary mechanisms have been explored 401 uals have a different proportion of body fat such that 368 to assess the costs and benefits associated with larger 402 $M' = M(1 + \chi)$. If $\chi < 0$, individuals within the com-369 body masses, particularly for mammals. One impor- 403 peting population have fewer metabolic reserves, and 370 tant such example is the fasting endurance hypothesis, 404 if $\chi > 0$, individuals have greater metabolic reserves 371 which contends that larger body size, with consequent 405 than the resident species. For the allowable values of $_{372}$ lower metabolic rates and increased ability to maintain $_{406}$ χ the adjusted mass should exceed the minimal amount 373 more endogenous energetic reserves, may buffer organ-407 of body fat, $1+\chi>\epsilon_{\sigma}$, and the adjusted time to re-374 isms against environmental fluctuations in resource avail- 408 produce must be positive, which given Eq. 4, implies 375 ability⁵⁵. Over evolutionary time, terrestrial mammalian 400 that $1 - \epsilon_{\lambda}^{1-\eta} (1+\chi)^{1-\eta} > 0$. Together these conditions 410 imply that $\chi \in (-f_0 M^{\gamma-1}, 1/\epsilon_{\lambda} - 1)$ where the upper 468 starvation and recovery that are expected to occur in 411 bound approximately equals 0.05 and the lower bound 469 resource-limited environments. We found that incorpo-412 is mass-dependent. The modified mass of the competi- 470 rating allometrically-determined rates into the NSM pre-413 tor leads to altered rates of starvation $\sigma(M')$, recovery 471 dicts: (i) that extinction risk is minimized, (ii) the pre-414 $\rho(M')$, and the maintenance of both starving $\delta(M')$ and 472 dicted steady-states quantitatively reproduce Damuth's 415 full consumers $\beta(M')$ (see the SI for detailed derivations 473 law, and (iii) the selective mechanism for the evolution of 416 of competitor rates). Importantly, ϵ_{σ} , which determines 474 larger body sizes anticipates Cope's rule. The NSM offers 417 the point along the growth curve that defines the body 475 a means by which the dynamic consequences of energetic 418 composition of starved foragers, is assumed to remain 476 constraints can be assessed using macroscale interactions 419 unchanged for the competing population.

To assess the susceptibility of the resident species 421 to competitive exclusion, we determine which consumer 422 pushes the steady-state resource density R^* to lower val-478 423 ues for a given value of χ , with the expectation that a 424 population capable of surviving on lower resource densi-425 ties has a competitive advantage⁵⁸. We find that for $M \leq$ 426 1.748×10^7 (g), having additional body fat $(\chi > 0)$ results 427 in a lower steady state resource density $(R'^* < R^*)$, such 428 that the competitor has an intrinsic advantage over the resident species (Fig. 5). However, for $M > 1.748 \times 10^7$ 430 (g), leaner individuals ($\chi < 0$) have lower resource steady 431 state densities, switching the advantage.

The observed switch in susceptibility as a function of 433 χ at $M_{\rm opt} = 1.748 \times 10^7$ (g) thus serves as an attractor, 434 such that the NSM predicts organismal mass to increase 435 if $M < M_{\rm opt}$ and decrease if $M > M_{\rm opt}$. This value is 479 where $A = (\lambda \xi \rho + \mu \sigma)$ and $B = (\beta \mu \xi + \delta \lambda \xi - \lambda \mu)$. The stabil-436 close to but smaller than the asymptotic upper bound 480 ity of this fixed point is determined by the Jacobian matrix J, 437 for terrestrial mammal body size predicted by the NSM, 481 where each matrix element $J_{ij} = \partial \dot{X}_i/\partial X_j$ when evaluated 438 however it is remarkably close to independent estimates 482 at the internal fixed point, and X is the vector (F, H, R). 439 of the largest land mammals, the early Oligocene Indri- 483 The parameters in Eq. (1) are such that the real part of the 440 cotherium at ca. 1.5×10^7 (g) and the late Miocene 484 largest eigenvalue of $\bf J$ is negative, so that the system is sta-441 Deinotherium at ca. 1.74×10^7 (g) ⁵⁶. Additionally, ⁴⁸⁵ ble with respect to small perturbations from the fixed point. 442 our calculation of M_{opt} as a function of mass-dependent 443 physiological rates is similar to theoretical estimates of 444 maximum body size⁴⁷, and provides independent theo-445 retical support for the observation of a 'maximum body 446 size attractor' explored by Alroy⁴⁶.

448 mass was predicted by Brown et al. based on re- 493 for metazoans (e.g., ref. 39), and has taxonomic shifts for uniproductive maximization and the transition between 494 cellular species between $\eta \approx 1$ in eukaryotes and $\eta \approx 1.76$ in 450 hungry and full individuals 52 . By coupling the NSM 495 bacteria 3,62 . $_{451}$ to resource dynamics as well as introducing an explicit 496 452 treatment of storage, we show that species with larger 453 body masses have an inherent competitive advantage 499 rates of organisms ranging from bacteria to metazoans 3,63-67. 454 for size classes up to $M_{\rm opt} = 1.748 \times 10^7$. While 500 This relation is derived from the simple balance condition 455 the state of the environment as well as the compet- ${}^{501}B_0m^{\eta}=E_m\dot{m}+B_mm$, ${}^{3,63-67}$ where E_m is the energy needed 456 itive landscape will determine whether specific body 502 to synthesize a unit of mass, B_m is the metabolic rate to 457 sizes are selected for or against 57 , we propose that $_{503}$ support an existing unit of mass, and m is the mass of the 458 the dynamics of starvation and recovery described in 504 organism at any point in its development. This balance has 459 the NSM provide a general selective mechanism for the 505 the general solution 3,68 460 evolution of larger body size among terrestrial mammals.

462 Discussion

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The energetics associated with somatic maintenance, 465 growth, and reproduction are important elements that 466 influence the dynamics of all populations¹¹. The NSM 467 is a general model that incorporates the dynamics of

477 between and am ong species.

Methods

Analytical solution to the NSM Equation (1) has three fixed points: two trivial fixed points at (F^*, H^*, R^*) = (0,0,0) and (0,0,1), and one non-trivial, internal fixed point

$$F^* = (\sigma - \lambda) \frac{\alpha \lambda \mu^2 (\mu + \xi \rho)}{A(\lambda \rho B + \mu \sigma (\beta \mu + \lambda (\delta + \rho)))},$$

$$H^* = (\sigma - \lambda) \frac{\alpha \lambda^2 \mu (\mu + \xi \rho)}{A(\lambda \rho B + \mu \sigma (\beta \mu + \lambda (\delta + \rho)))},$$

$$R^* = (\sigma - \lambda) \frac{\mu}{A}.$$
(2)

486 Because this fixed point is unique, it is the global attractor 487 for all population trajectories for any initial condition where 488 the resource and consumer densities are both nonzero.

Metabolic scaling relationships The scaling relation between an organism's metabolic rate B and its body mass M491 at reproductive maturity is known to scale as $B = B_0 M^{\eta 61}$. An optimal size for mammals at intermediate body 492 where the scaling exponent η is typically close to 2/3 or 3/4

Several efforts have shown how a partitioning of B between

$$\left(\frac{m(t)}{M}\right)^{1-\eta} = 1 - \left[1 - \left(\frac{m_0}{M}\right)^{1-\eta}\right] e^{-a(1-\eta)t/M^{1-\eta}}, \quad (3)$$

where, for $\eta < 1$, $M = (B_0/B_m)^{1/(1-\eta)}$ is the asymptotic mass, $a = B_0/E_m$, and m_0 is mass at birth, itself varying allometrically (see the SI). We now use this solution to define the timescale for reproduction and recovery from starvation (Fig. 2; see⁶⁴ for a detailed presentation of these timescales).

by the timescale

$$\tau(\epsilon) = \ln\left[\frac{1 - (m_0/M)^{1-\eta}}{1 - \epsilon^{1-\eta}}\right] \frac{M^{1-\eta}}{a(1-\eta)},\tag{4}$$

506 where we will define values of ϵ to describe a variety of 507 timescales, and related rates, within our model. For example, 508 the rate of reproduction is given by the timescale to go from 509 the birth mass to the adult mass. The time to reproduce is 510 given by Equation 4 as $t_{\lambda} = \tau(\epsilon_{\lambda})$, where ϵ_{λ} is the fraction of 511 the asymptotic mass where an organism is reproductively ma-512 ture and should be close to one (typically $\epsilon_{\lambda} \approx 0.95^{63}$). Our 513 reproductive rate, λ , is a specific rate, or the number of off-514 spring produced per time per individual, defined as $F = \lambda F$. 515 In isolation this functional form gives population growth fol-516 lowing $F(t) = F_0 e^{\lambda t}$ which can be related to the reproductive 517 timescale by assuming that when $t = t_{\lambda}$ it is also the case that $F = \nu F_0$, where $\nu - 1$ is the number of offspring pro-519 duced per reproductive cycle. Following this relationship the 520 growth rate is given by $\lambda = \ln(\nu)/t_{\lambda}$, which is the standard 521 relationship (e.g. ⁶⁷) and will scales as $\lambda \propto M^{\eta-1}$ for $M \gg m_0$ for any constant value of $\epsilon_{\lambda}^{3,63-66}$.

The rate of recovery $\rho = 1/t_{\rho}$ requires that an organism ac-524 crues sufficient tissue to transition from the hungry to the full 525 state. Since only certain tissues can be digested for energy (for 526 example the brain cannot be degraded to fuel metabolism), 527 we define the rates for starvation, death, and recovery by the 528 timescales required to reach, or return from, specific fractions 529 of the replete-state mass (see the SI, Table I, for parameteri-530 zations). We define $m_{\sigma} = \epsilon_{\sigma} M$, where $\epsilon_{\sigma} < 1$ is the fraction 531 of replete-state mass where reproduction ceases. This fraction 532 will deviate from a constant if tissue composition systemat-533 ically scales with adult mass. For example, making use of 534 the observation that body fat in mammals scales with overall 535 body size according to $M_{\rm fat} = f_0 M^{\gamma}$ and assuming that once 536 this mass is fully digested the organism starves, this would 537 imply that $\epsilon_{\sigma} = 1 - f_0 M^{\gamma}/M$. It follows that the recovery 538 timescale, t_{ρ} , is the time to go from $m = \epsilon_{\sigma} \epsilon_{\lambda} M$ to $m = \epsilon_{\lambda} M$ 539 (Fig. 2). Using Eqs. (3) and (4) this timescale is given by sim-540 ply considering an adjusted starting mass of $m_0' = \epsilon_{\sigma} \epsilon_{\lambda} M$, in 541 which case

$$t_{\rho} = \ln \left[\frac{1 - (\epsilon_{\sigma} \epsilon_{\lambda})^{1 - \eta}}{1 - \epsilon_{\lambda}^{1 - \eta}} \right] \frac{M^{1 - \eta}}{a'(1 - \eta)}$$
 (5)

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The time that it takes to reach a particular mass ϵM is given ϵM where $a'=B_0/E'_m$ accounts for possible deviations in the 543 biosynthetic energetics during recovery (see the SI). It should 544 be noted that more complicated ontogenetic models explicitly by handle storage⁶⁶, whereas this feature is implicitly covered by 546 the body fat scaling in our framework.

> To determine the starvation rate, σ , we are interested in the time required for an organism to go from a mature adult that reproduces at rate λ , to a reduced-mass hungry state where reproduction is impossible. For starving individuals we assume that an organism must meet its maintenance requirements by using the digestion of existing mass as the sole energy source. This assumption implies the following simple metabolic balance $\dot{m}E'_m = -B_m m$ or $\dot{m} = -a'm/M^{1-\eta}$ where E'_m is the amount of energy stored in a unit of existing body mass, which differs from E_m , the energy required to synthesis a unit of biomass⁶⁶. Given the replete mass, M, of an organism, the above energy balance prescribes the mass trajectory of a non-consuming organism: $m(t) = Me^{-a't/M^{1-\eta}}$. The timescale for starvation is given by the time it takes m(t)to reach $\epsilon_{\sigma}M$, which gives

$$t_{\sigma} = -\frac{M^{1-\eta}}{a'} \ln \left(\epsilon_{\sigma} \right). \tag{6}$$

547 The starvation rate is then $\sigma = 1/t_{\sigma}$, which scales with 548 replete-state mass as $1/M^{1-\eta} \ln (1-f_0M^{\gamma}/M)$. An important feature is that σ does not have a simple scaling depen-550 dence on λ , which is important for the dynamics that we later 551 discuss.

The time to death should follow a similar relation, but de-553 fined by a lower fraction of replete-state mass, $m_{\mu} = \epsilon_{\mu} M$ 554 where $\epsilon_{\mu} < \epsilon_{\sigma}$. Suppose, for example, that an organism dies 555 once it has digested all fat and muscle tissues, and that mus-556 cle tissue scales with body mass according to $M_{\text{musc}} = u_0 M^{\zeta}$. This gives $\epsilon_{\mu} = 1 - \left(f_0 M^{\gamma} + u_0 M^{\zeta} \right) / M$. Muscle mass has 558 been shown to be roughly proportional to body mass⁶⁹ in 559 mammals and thus ϵ_{μ} is merely ϵ_{σ} minus a constant. The 560 time to go from starvation to death is the total time to reach **561** $\epsilon_{\mu}M$ minus the time to starve, or $t_{\mu} = -M^{1-\eta} \ln \left(\epsilon_{\mu}\right)/a' - t_{\sigma}$, (5) see and $\mu = 1/t_{\mu}$.

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