# The dynamics of starvation and recovery

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6 The eco-evolutionary dynamics of species are fundamentally 64 vest differently in broads during periods of resource scarcity  $\tau$  linked to the energetic constraints of its constituent individuals.  $_{65}$  compared to normal periods [9, 10], sometimes delaying or even 8 Of particular importance are the tradeoffs between reproduction 9 and the dynamics of starvation and recovery in resource-limited 10 environments. To elucidate the consequences of this tradeoff, 11 we introduce a minimal nutritional state-structured model that in-12 corporates two classes of consumer: nutritionally replete con-13 sumers that reproduce, and undernourished, non-reproducing 14 consumers that are susceptible to mortality. As a function of 15 the transition rates between these two states that are determined 72 migratory breeding grounds to reproduce after one or multiple 16 by the abundance of resources, the consumer populations can ei- 73 seasons in resource-rich environments where they accumulate 17 ther undergo cyclic dynamics or reach a steady state. We obtain 74 nutritional reserves [14, 15, 16]. 18 strong constraints on starvation and recovery rates by deriving 19 allometric scaling relationships between body size and a variety 20 of traits and find that population dynamics subject to these constraints are typically driven to a steady state. Moreover, we find 22 that these rates fall within a 'refuge' in parameter space, where 23 the probability of extinction of the consumer population is min-24 imized. Thus we identify a potential mechanism that may both  $_{25}$  drive and constrain the dynamics of animal populations. Our  $_{81}$  other species (including humans) suffer irregular menstrual cy-26 model provides a natural framework that predicts maximum body 82 cling and higher abortion rates during periods of nutritional 27 size for mammals by determining the relative stability of an oth- 83 stress [19, 20]. In the extreme case of unicellular organisms, erwise homogeneous population to a mutant population with al- $_{29}$  tered percentage of body fat. For body masses  $< 1.748 \times 10^7$  g, 30 individuals with increased energetic reserves can invade resident  $_{31}$  populations, and vice versa for body mass  $> 1.748 \times 10^7 \mathrm{g}$ , thus 32 providing a principled mechanism for a within-lineage driver of 33 Cope's rule.

34 foraging | starvation | reproduction

35 Significance Statement Energetic investment in somatic mainte-36 nance and growth vs. reproduction directly impacts the dynamics of 37 populations among species. Here, we construct a Nutritional State-38 structured Model (NSM) to assess the population-level effects of star-39 vation and recovery of a consumer population in a resource-limited en-40 vironment, and use allometric scaling relationships for mammals to es-41 tablish all timescales and rates. Our model: i. reveals that mammalian 42 energetic rates minimize the probability of stochastic extinction, ii. es-43 tablishes dynamic bounds on mammalian body size while providing 44 independent theoretical support for the energy equivalence hypothesis, 45 and iii. provides a mechanistic driver for the evolutionary trend towards 46 larger body size known as Cope's rule.

## 47 Introduction

49 ergetic state of individuals, which directly influences how they 107 the idea that reproduction is strongly constrained allometrically 50 invest reserves in uncertain environments. Such behaviors are 108 [3], and is not generally linearly related to resource density. As 51 generally manifested as tradeoffs between investing in somatic 52 maintenance and growth, or allocating energy towards repro-53 duction [1, 2, 3]. The timing of these behaviors responds to 54 selective pressure, as the choice of the investment impacts fu-55 ture fitness [4, 5, 6]. The influence of resource limitation on an 56 organism's ability to maintain its nutritional stores may lead to 57 repeated delays or shifts in reproduction over the course of an 58 organism's life.

The balance between (a) somatic growth and maintenance, 60 and (b) reproduction depends on resource availability [7]. For 61 example, reindeer invest less in calves born after harsh win-62 ters (when the mother's energetic state is depleted) than in 63 calves born after moderate winters [8]. Many bird species in-

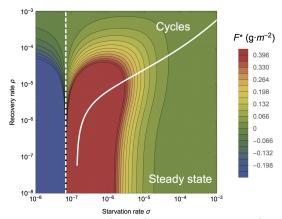
66 foregoing reproduction for a breeding season [1, 11, 12]. Even 67 freshwater and marine zooplankton have been observed to avoid 68 reproduction under nutritional stress [13], and those that do 69 reproduce have lower survival rates [2]. Organisms may also 70 separate maintenance and growth from reproduction over space 71 and time: many salmonids, birds, and some mammals return to

Physiology also plays an important role in regulating repro-76 ductive expenditures during periods of resource limitation. The 77 data collected thus far has shown that diverse mammals (47 78 species in 10 families) exhibit delayed implantation, whereby 79 females postpone fetal development (blastocyst implantation) so until nutritional reserves can be accumulated [17, 18]. Many 84 nutrition is unavoidably linked to reproduction because the nu-85 tritional state of the cell regulates all aspects of the cell cycle 86 [21]. The existence of so many independently evolved mecha-87 nisms across such a diverse suite of organisms highlights the im-88 portance and universality of the fundamental tradeoff between 89 somatic and reproductive investment. However the general dy-90 namic implications of these constraints are unknown.

Though straightforward conceptually, incorporating the en-92 ergetic dynamics of individuals [22] into a population-level 93 framework [22, 23] presents numerous mathematical obsta-94 cles [24]. An alternative approach involves modeling the 95 macroscale relations that guide somatic versus reproductive 96 investment in a consumer-resource system. For example, 97 macroscale Lotka-Volterra models assume that the growth rate 98 of the consumer population depends on resource density, thus 99 implicitly incorporating the requirement of resource availability 100 for reproduction [25].

In this work, we adopt an alternative approach in which we 102 explicitly account for resource limitation and the subsequent 103 effects of starvation. Namely, only individuals with sufficient 104 energetic reserves can reproduce. Such a constraint leads to 105 reproductive time lags due to some members of the population 48 The behavioral ecology of all organisms is influenced by the en- 106 going hungry and then recovering. Additionally, we incorporate

#### **Reserved for Publication Footnotes**



function of the star vation rate  $\sigma$  and recovery rate  $\rho$  for a 100g consumer. These bifurcation conditions separate parameter space into unphysical, cyclic, and steady state dynamic regimes. The colors show the steady state densities for the energetically replete consumers  $F^*$ , (warmer colors denote higher densities).

109 we shall show, these constraints influence the ensuing popula-110 tion dynamics in dramatic ways.

#### 112 Nutritional state-structured model (NSM)

113 We begin by defining a minimal Nutritional State-structured population Model (NSM), where the consumer population is  $^{160}$  where  $A = (\lambda \xi \rho + \mu \sigma)$  and  $B = (\beta \mu \xi + \delta \lambda \xi - \lambda \mu)$ . The standard population is  $^{160}$  where  $A = (\lambda \xi \rho + \mu \sigma)$  and  $A = (\beta \mu \xi + \delta \lambda \xi - \lambda \mu)$ . partitioned into two states: (a) an energetically replete (full) 161 bility of this fixed point is determined by the Jacobian matrix state F, where the consumer reproduces at a constant rate  $\lambda$  162  $\mathbf{J}$ , where each matrix element  $J_{ij} = \partial \dot{X}_i/\partial X_j$  when evaluated and does not die from starvation, and (b) an energetically defi-  $^{163}$  at the internal fixed point, and  $\mathbf{X}$  is the vector (F, H, R). The 118 cient (hungry) state H, where the consumer does not reproduce 164 parameters in Eq. [1] are such that the real part of the largest 119 but dies by starvation at rate  $\mu$ . The underlying resource R 165 eigenvalue of J is negative, so that the system is stable with  $_{120}$  evolves by logistic growth with an intrinsic growth rate  $\alpha$  and  $_{166}$  respect to small perturbations from the fixed point. Because  $_{121}$  a carrying capacity C. The rate at which consumers transition  $_{167}$  this fixed point is unique, it is the global attractor for all pop-122 between states and consume resources is dependent on their 168 ulation trajectories for any initial condition where the resource 123 overall abundance, the abundance of resources, the efficiency of 169 and consumer densities are both nonzero. 124 converting resources into metabolism, and how that metabolism 170 From Eq. [2], an obvious constraint on the NSM is that 125 is partitioned between maintenance and growth purposes. In 171 the reproduction rate  $\lambda$  must be less than the starvation rate 126 the supplementary information (SI) we provide a fully mecha- 172  $\sigma$ , so that the consumer and resource densities are positive. 127 nistic model for each of these dynamics and constants, and show 173 The condition  $\sigma = \lambda$  thus represents a transcritical (TC) bi-128 that the system produces a simple non-dimensional form which 174 furcation [31] that demarcates a physical from an unphysical 129 we describe below.

Consumers transition from the full state F to the hungry 131 state H at a rate  $\sigma$ —the starvation rate—and also in propor-132 tion to the absence of resources (1-R). Conversely, consumers 133 recover from state H to state F at rate  $\xi \rho$  and in proportion to 134 R, where  $\xi$  represents a ratio between maximal resource con-135 sumption and the carrying capacity of the resource. Resources are eaten by the hungry consumers at rate  $\rho R + \delta$ , that accounts 137 for their somatic growth  $(\rho R)$  and maintenance  $(\delta)$ . Full con-138 sumers eat resources at a constant rate  $\beta$  that accounts for 139 maximal maintenance and somatic growth (see SI for mecha-140 nistic derivations of these rates from resource energetics). The 141 NSM represents an ecologically motivated fundamental exten-142 sion of the idealized starving random walk model of foraging, 143 which focuses on resource depletion, to include reproduction 144 and resource replenishment [26, 27, 28], and is a more general 145 formulation than previous models incorporating starvation [29].

In the mean-field approximation, in which the consumers 147 and resources are perfectly mixed, their densities evolve accord-148 ing to the rate equations

$$\dot{F} = \lambda F + \xi \rho R H - \sigma (1 - R) F,$$

$$\dot{H} = \sigma (1 - R) F - \xi \rho R H - \mu H,$$

$$\dot{R} = \alpha (1 - R) R - (\rho R + \delta) H - \beta F$$
[1]

This system of nondimensional equations follows from a 150 set of first-principle relationships for resource consumption and 151 growth (see SI for a full derivation and the dimensional form). 152 Notice that the total consumer density F + H evolves accord-153 ing to  $\dot{F} + \dot{H} = \lambda F - \mu H$ . This resembles the equation of 154 motion for the predator density in the classic Lotka-Volterra 155 model [30], except that the resource density does not appear in 156 the growth term. As discussed above, the attributes of repro-157 duction and mortality have been explicitly apportioned to the 158 full and hungry consumers, respectively, so that the growth in Fig. 1. The transcritical (dashed) and Hopf bifurcation (solid) as a 159 the total density is decoupled from the resource density.

> Equation [1] has three fixed points: two trivial fixed points at  $(F^*, H^*, R^*) = (0, 0, 0)$  and (0, 0, 1), and one non-trivial, internal fixed point at

$$F^* = (\sigma - \lambda) \frac{\alpha \lambda \mu^2 (\mu + \xi \rho)}{A(\lambda \rho B + \mu \sigma (\beta \mu + \lambda (\delta + \rho)))},$$

$$H^* = (\sigma - \lambda) \frac{\alpha \lambda^2 \mu (\mu + \xi \rho)}{A(\lambda \rho B + \mu \sigma (\beta \mu + \lambda (\delta + \rho)))},$$

$$R^* = (\sigma - \lambda) \frac{\mu}{\Delta}.$$
[2]

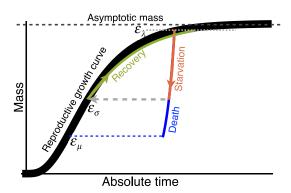


Fig. 2. The growth trajectory over absolute time of an individual organism as a function of body mass. Initial growth follows the black trajectory to an energetically replete reproductive adult mass  $m = \epsilon_{\lambda} M$ . Starvation follows the red trajectory to  $m = \epsilon_{\sigma} \epsilon_{\lambda} M$ . Recovery follows the green curve to the replete adult mass, where this trajectory differs from the original growth because only fat is being regrown which requires a longer time to reach  $\epsilon_{\lambda}M$ . Alternatively, death from starvation follows the blue trajectory to  $m = \epsilon_{\mu} \epsilon_{\lambda} M$ .

175 regime where all steady-state densities become negative after in- 242 This balance has the general solution [46, 3] 176 tersecting the trivial fixed point  $(F^*, H^*, R^*) = (0, 0, 0)$ . The 177 biological implication of the constraint  $\lambda < \sigma$  has a simple 178 interpretation—the rate at which a macroscopic organism loses mass due to lack of resources is generally much faster than the rate of reproduction. As we will discuss below, this inequality is rate of reproduction. As we will discuss below, this inequality is rate of reproduction. As we will discuss below, this inequality is rate of reproduction. As we will discuss below, this inequality is rate of reproduction. As we will discuss below, this inequality is rate of reproduction. As we will discuss below, this inequality is represented as a superior of allowatric constraints [3] for organisms. within empirically observed body size ranges.

In the physical regime of  $\lambda < \sigma$ , the fixed point [2] may 184 either be a stable node or a limit cycle (Fig. 1). In continuous-185 time systems, a limit cycle arises when a pair of complex con-186 jugate eigenvalues crosses the imaginary axis to attain positive real parts [32]. This Hopf bifurcation is defined by  $Det(\mathbf{S}) = 0$ , 189 ficients of the characteristic polynomial of the Jacobian ma-  $^{249}$  where we will define values of  $\epsilon$  to describe a set of rates within 188 with S the Sylvester matrix, which is composed of the coefdescription, can increase the extinction risk [35, 36, 37].

When the starvation rate  $\sigma \gg \lambda$ , a substantial fraction of the consumers are driven to the hungry non-reproducing state. Because reproduction is inhibited, there is a low steadystate consumer density and a high steady-state resource density. However, if  $\sigma / \lambda = 1$  from above the starvation rate  $\sigma \gg \lambda$ , a substantial fraction 256 [42, 43, 44, 45, 3].

The rate of recovery  $\rho = 1/t_{\rho}$  requires that an organism 258 accrues sufficient tissue to transition from the hungry to the 259 full state. Since only certain tissues can be digested for energy 260 (for example the brain connect be decreased to find the property of the prope 208 increases with higher recovery rates.

## 210 Role of allometry

214 to constrain the covariation of rates in a principled and biologi274 ing mass of  $m_0' = \epsilon_\sigma \epsilon_\lambda M$ , in which case 215 cally meaningful manner. Allometric scaling relations highlight 216 common constraints and average trends across large ranges in 217 body size and species diversity. Many of these relations can be 218 derived from a small set of assumptions and below we describe  $_{275}$  where  $a'=B_0/E'_m$  accounts for possible deviations in the  $_{219}$  our framework to determine the covariation of timescales and  $_{276}$  biosynthetic energetics during recovery (see SI). It should be 220 rates across the range of mammals for each of the key parame221 ters of our model (cf. ref. [38]). We are thereby able to define
222 the regime of dynamics occupied by the entire class of mam223 body fat scaling in our framework. 223 mals, along with the key differences between the largest and 280 224 smallest mammals.

238 ance condition  $B_0 m^{\eta} = E_m \dot{m} + B_m m$ , [42, 43, 44, 45, 3] where 239  $E_m$  is the energy needed to synthesize a unit of mass,  $B_m$  is the metabolic rate to support an existing unit of mass, and  $m_{241}$  is the mass of the organism at any point in its development.

The starvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the starvation rate is then  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is the mass as  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate is  $\sigma = 1/t_{\sigma}$ , which scales with repletestarvation rate

$$\left(\frac{m(t)}{M}\right)^{1-\eta} = 1 - \left[1 - \left(\frac{m_0}{M}\right)^{1-\eta}\right] e^{-a(1-\eta)t/M^{1-\eta}},$$
[3]

180 rate of reproduction. 12 is a natural consequence of allometric constraints [3] for organisms  $\frac{244}{245}$  cally (see SI). We now use this solution to define the timescale 246 for reproduction and recovery from starvation (Fig. 2; see [43] 247 for a detailed presentation of these timescales). The time that 248 it takes to reach a particular mass  $\epsilon M$  is given by the timescale

$$\tau\left(\epsilon\right) = \ln\left[\frac{1 - \left(m_0/M\right)^{1-\eta}}{1 - \epsilon^{1-\eta}}\right] \frac{M^{1-\eta}}{a\left(1 - \eta\right)},\tag{4}$$

190 trix [33]. As the system parameters are tuned to be within the 250 our model. For the time to reproduce,  $t_{\lambda} = \tau(\epsilon_{\lambda})$ , where  $\epsilon_{\lambda}$ 191 stable regime, but close to the Hopf bifurcation, the amplitude  $^{251}$  is the fraction of the asymptotic mass where an organism is 192 of the transient cycles becomes large. Given that ecological 252 reproductively mature and should be close to one (typically 193 systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34], the onset of transient walk systems are constantly being perturbed [34]. 194 sient cycles, even though they decay with time in the mean-field  $^{254}$  where v is the number of offspring produced, and for any con-255 stant value of  $\epsilon_{\lambda}$ , this rate will scale as  $\lambda \propto M^{\eta-1}$  for  $M \gg m_0$ 

200 sity. However, if  $\sigma/\lambda \to 1$  from above, the population is  $^{260}$  (for example the brain cannot be degraded to fuel metabolism), 201 overloaded with energetically-replete (reproducing) individuals, 261 we define the rates for starvation, death, and recovery by the 202 thereby promoting transient oscillations between the consumer 262 timescales required to reach, or return from, specific fractions and recovery densities (Fig. 4). If the replete state were formulated to the inetabolism), 203 and resource densities (Fig. 1). If the starvation rate is low 263 of the replete-state mass (see SI, Table I for parameterizations). 204 enough that the Hopf bifurcation is crossed, these oscillations  $^{264}$  We define  $m_{\sigma} = \epsilon_{\sigma} M$ , where  $\epsilon_{\sigma} < 1$  is the fraction of replete-205 become stable over time. This threshold occurs at higher values 265 state mass where reproduction ceases. This fraction will deviate 206 of the starvation rate as the recovery rate  $\rho$  increases, such that 266 from a constant if tissue composition systematically scales with 207 the range of parameter space giving rise to cyclic dynamics also 267 adult mass. For example, making use of the observation that 268 body fat in mammals scales with overall body size according to  $_{269}~M_{\rm fat}=f_0M^{\gamma}$  and assuming that once this mass is fully digested 270 the organism starves, this would imply that  $\epsilon_{\sigma} = 1 - f_0 M^{\gamma}/M$ . While there are no a priori constraints on the parameters in  $^{271}$  It follows that the recovery timescale,  $t_{\rho}$ , is the time to go from  $^{212}$  the NSM, most organisms correspond to restricted portions of  $^{272}m = \epsilon_{\sigma}\epsilon_{\lambda}M$  to  $m = \epsilon_{\lambda}M$  (Fig. 2). Using Eqs. [3] and [4] 213 the parameter space. Here we use allometric scaling relations 273 this timescale is given by simply considering an adjusted start-

$$t_{\rho} = \ln \left[ \frac{1 - (\epsilon_{\sigma} \epsilon_{\lambda})^{1 - \eta}}{1 - \epsilon_{\lambda}^{1 - \eta}} \right] \frac{M^{1 - \eta}}{a' (1 - \eta)}$$
 [5]

To determine the starvation rate,  $\sigma$ , we are interested in the Nearly all of the rates described in the NSM are determined  $^{281}$  time required for an organism to go from a mature adult that  $^{282}$  reproduces at rate  $\lambda$ , to a reduced-mass hungry state where reby consumer metabolism, which can be used to describe a vari- 283 production is impossible. For starving individuals we assume ety of organismal features [39]. The scaling relation between an <sub>284</sub> that an organism must meet its maintenance requirements by 228 organism's metabolic rate B and its body mass M at reproduc229 tive maturity is known to scale as  $B = B_0 M^{\eta}$  [40], where the
230 scaling exponent  $\eta$  is typically close to 2/3 or 3/4 for metazoans
231 (e.g., ref. [39]), and has taxonomic shifts for unicellular species
238 that an organism must need to mathematical the following mass as the sole energy source.
248 that an organism must need to mathematic requirements as using the digestion of existing mass as the sole energy source.
250 This assumption implies the following simple metabolic balance
251  $mE'_m = -B_m m$  or  $m = -a'm/M^{1-\eta}$  where  $E'_m$  is the amount constant of the production of existing body mass, which differs the following simple metabolic balance are the mathematic requirements of the production of existing mass as the sole energy source.
252  $mE'_m = -B_m m$  or  $m = -a'm/M^{1-\eta}$  where  $E'_m$  is the amount constant of the production of existing mass as the sole energy source.
253  $mE'_m = -B_m m$  or  $m = -a'm/M^{1-\eta}$  where  $E'_m$  is the amount constant of the production of existing mass as the sole energy source.
254  $mE'_m = -B_m m$  or  $m = -a'm/M^{1-\eta}$  where  $E'_m = -B_m m$  or  $m = -a'm/M^{1-\eta}$  wher 231 (e.g., fel. [95]), and has taxonomic since the matrix for time that a species 288 of energy stored in a unit of existing body mass, which differs 232 between  $\eta \approx 1$  in eukaryotes and  $\eta \approx 1.76$  in bacteria [41, 3]. 289 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from 290 [45]. Given the replete mass, M, of an organism, the above energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from  $E_m$ , the energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from  $E_m$ , the energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from  $E_m$ , the energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from  $E_m$ , the energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of existing body mass, which differs 280 from  $E_m$ , the energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy stored in a unit of existing body mass, which differs 280 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from  $E_m$ , the energy required to synthesis a unit of biomass 280 from  $E_m$ , the energy required to synthesis a unit of existing body mass, which differs 280 from  $E_m$  f

$$t_{\sigma} = -\frac{M^{1-\eta}}{a'} \ln \left( \epsilon_{\sigma} \right).$$
 [6]

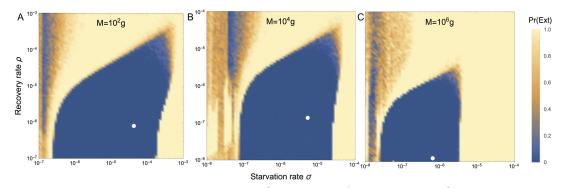


Fig. 3. Probability of extinction for a consumer with (A)  $M=10^2 \mathrm{g}$ , (B)  $M=10^4 \mathrm{g}$ , and (C)  $M=10^6 \mathrm{g}$  as a function of the starvation rate  $\sigma$  and recovery rate  $\rho$ , where the initial density is given as  $(XF^*,XH^*,R^*)$ , where X is a random uniform variable in [0,2]. Note the change in scale for  $M=10^4$  and  $M=10^6$ g. Extinction is defined as the population trajectory falling below  $0.2\times$  the allometrically constrained steady state. The white points denote the allometrically constrained starvation and recovery rate.

297 is important for the dynamics that we later discuss.

299 fined by a lower fraction of replete-state mass,  $m_{\mu} = \epsilon_{\mu} M$  where 347 replicate the initial densities are chosen to be  $(XF^*, XH^*, R^*)$ ,  $\epsilon_{\mu} < \epsilon_{\sigma}$ . Suppose, for example, that an organism dies once it 348 with X a random variable that is uniformly distributed in [0, 2]. 301 has digested all fat and muscle tissues, and that muscle tissue 349 By allowing the rate of starvation to vary, we assessed extinc\_  $_{302}$  scales with body mass according to  $M_{\rm musc}=u_0M^{\zeta}$ . This gives  $_{350}$  tion risk across a range of values for  $\sigma$  and  $\rho$  between ca.  $10^{-7}$  $_{303}$   $\epsilon_{\mu}=1-(f_0M^{\gamma}+u_0M^{\zeta})/M$ . Muscle mass has been shown  $_{351}$  to  $10^{-3}$ . As expected, higher rates of extinction correlate with to be roughly proportional to body mass [47] in mammals and 352 both high values of  $\sigma$  if  $\rho$  is small, and high values of  $\rho$  if  $\sigma$  305 thus  $\epsilon_{\mu}$  is merely  $\epsilon_{\sigma}$  minus a constant. The time to go from 353 is small. For low values of  $\sigma$  and high values of  $\rho$ , the in-

309 on parameterizations for terrestrial-bound endotherms, specif- 357 risk arises because of the decrease in the steady-state consumer 310 ically mammals, which range from a minimum of  $M \approx 1$ g 358 population density (Figs. 1B, 3). This interplay creates an 'ex-<sub>311</sub> (the Etruscan shrew Suncus etruscus) to a maximum of  $M \approx _{359}$  tinction refuge', such that for a constrained range of  $\sigma$  and  $\rho$ , 312 10<sup>7</sup>g (the early Oligocene Indricotheriinae and the Miocene 360 extinction probabilities are minimized. <sup>313</sup> Deinotheriinae). Investigating other classes of organisms would <sup>361</sup>  $_{314}$  simply involve altering the metabolic exponents and scalings  $_{362}$   $\rho$  fall squarely within the extinction refuge across a range of 315 associate with  $\epsilon$ . Moreover, we emphasize that our allometric 363 M (Fig. 3A-C, white points). These values are close enough 316 equations describe mean relationships, and do not account for 364 to the Hopf bifurcation to avoid low steady-state densities, and 317 the (sometimes considerable) variance associated with individ- 365 far enough away to avoid large-amplitude transient cycles. The 318 ual species.

## 320 Extinction risk

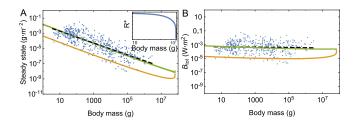
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321 As the allometric derivations of the NSM rate laws reveal, star-369 drive starvation and recovery rates within populations. Such a 322 vation and recovery rates are not independent parameters, and 370 mechanism would select for organism physiology that generates  $_{323}$  the biologically relevant portion of the phase space shown in  $_{371}$  appropriate  $\sigma$  and  $\rho$  values that serve to minimize extinction 324 Fig. 1 is constrained via covarying parameters. Given the pa-372 risk. This selection could occur via the tuning of body fat per-325 rameters of terrestrial endotherms, we find that the starvation 373 centages, metabolic rates, and biomass maintenance efficiencies. 326 rate  $\sigma$  and the recovery rate  $\rho$  are constrained to lie within a 327 small region of potential values for the known range of body  $_{328}$  sizes M. Indeed, starvation and recovery rates across all values  $_{329}$  of M fall squarely in the steady state region at some distance 330 from the Hopf bifurcation. This suggests that cyclic population 331 dynamics should be rare, particularly in environments where 332 resources are limiting.

Higher rates of starvation result in a larger flux of the popu-334 lation to the hungry state. In this state, reproduction is absent, 335 thus increasing the likelihood of extinction. From the perspec-336 tive of population survival, it is the rate of starvation relative 337 to the rate of recovery that determines the long-term dynamics 338 of the various species (Fig. 1). We therefore examine the com-339 peting effects of cyclic dynamics vs. changes in steady-state 340 density on extinction risk, both as functions of  $\sigma$  and  $\rho$ . To 341 this end, we computed the probability of extinction, where we 342 define extinction as a population trajectory falling below one 343 fifth of the allometrically constrained steady state at any time

<sub>296</sub> is that  $\sigma$  does not have a simple scaling dependence on  $\lambda$ , which <sub>344</sub> between  $t=10^8$  and  $t=10^{10}$ . This procedure is repeated for 345 50 replicates of the continuous-time system shown in Eq. 1 for The time to death should follow a similar relation, but de-  $_{346}$  organisms with mass ranging from  $10^2$  to  $10^6$  grams. In each so that  $\epsilon_{\mu}$  is interest  $\epsilon_{\sigma}$  minute to each term to reach  $\epsilon_{\mu}M$  minus the sign reaches extinction risk results from transient cycles with larger sign time to starve, or  $t_{\mu} = -M^{1-\eta} \ln{(\epsilon_{\mu})}/a' - t_{\sigma}$ , and  $\mu = 1/t_{\mu}$ . sign amplitudes as the system nears the Hopf bifurcation (Fig. 3). Although the rate equations [1] are general, here we focus sign For high values of  $\sigma$  and low values of  $\rho$ , increased extinction

> We find that the allometrically constrained values of  $\sigma$  and 366 feature that allometric values of  $\sigma$  and  $\rho$  fall within this rel-367 atively small window supports the possibility that a selective 368 mechanism has constrained the physiological conditions that



**Fig. 4.** (A) Consumer steady states  $F^*$  (green) and  $H^*$  (orange) as a function of body mass. Inset: Resource steady state  $R^*$  as a function of consumer body mass. (B) Total energetic use  $B_{\rm tot}$  of consumer populations at the steady state as a function of body mass. The data are from Damuth [56] and have been converted to total population metabolism using the allometric relationships for metabolic rate (please see SI and Refs. [42, 45, 43]).

We also find that as body size increases, the amount of low 441 sequences of energetic constraints, is close to the maximum 375 extinction risk parameter space becomes smaller (Fig. 3A-C), 442 body size observed in the North American mammalian fossil 376 suggesting that the population dynamics of larger organisms are 443 record [52], as well as the mass predicted from an evolutionary 377 more sensitive to smaller changes in physiological rates control-444 model of body size evolution [53]. It should be noted that the 378 ling starvation and recovery. To summarize, our finding that 445 asymptotic behavior and predicted upper bound depend only 379 the allometrically-determined parameters fall within this low 446 on the scaling of body composition and are independent of the 380 extinction probability region suggests that the NSM dynamics 447 resource parameters. We also note that the prediction of an may both drive—and constrain—natural animal populations. 448 asymptotic limit on mammalian size parallels work on micro-381

### 383 Dynamic and energetic barriers to body size

384 Metabolite transport constraints are widely thought to place 451 similar growth and energetic scaling relationships [3, 60]. 385 strict boundaries on biological scaling [48, 49, 39] and thereby 452 386 lead to specific predictions on the minimum possible body size 453 standing of the energetic dynamics that give rise to both ob-387 for organisms [50]. Above this bound, a number of energetic and 454 served limitations on mammalian body size, as well as the ob-388 evolutionary mechanisms have been explored to assess the costs 455 served trend towards larger body size over evolutionary time.  $_{389}$  and benefits associated with larger body masses, particularly  $_{456}$  The NSM predicts that the steady state resource density  $R^*$ 390 for mammals. One important such example is the fasting en-457 decreases with increasing body size of the consumer popula-393 tain more endogenous energetic reserves, may buffer organisms 460 dance will outcompete others [61, 62, 63]. Thus, the combined 394 against environmental fluctuations in resource availability [51]. 461 NSM steady-state dynamics and allometric timescales predict 396 significant trend towards larger body size (known as Cope's 463 given a common resource, but does not offer a within-lineage 397 rule) [52, 53, 54, 55], and it is thought that within-lineage 464 mechanism by which larger body sizes are selected for. 398 drivers generate selection towards an optimal upper bound of 399 roughly 10<sup>7</sup> grams [52], a value that is likely limited by higher 400 extinction risk for large taxa over longer timescales [53]. These 401 trends are thought to be driven by a combination of climate 402 change and niche availability [55]; however the underpinning 403 energetic costs and benefits of larger body sizes, and how they 404 influence dynamics over ecological timescales, have not been ex-405 plored. We argue that the NSM provides a suitable framework to explore these issues.

The NSM correctly predicts that species with smaller  $_{408}$  masses have larger steady-state population densities (Fig. 4A). Similar predictions have been made for carnivore populations 410 using alternative consumer-resource models [57]. Moreover, we show that the NSM provides independent theoretical support for the energy equivalence hypothesis and Damuth's Law [56, 58, 59]. The energy equivalence hypothesis argues that the 414 total energy use,  $B_{\text{tot}}$ , of a population is constant independent 415 of species size [56, 58, 59]. This hypothesis is based on ob-416 servations showing that the steady state abundance,  $N^*$ , of a 417 species is proportional to the inverse of individual metabolism, 418 such that  $N^* \propto M^{-3/4}/B_0$  [58, 59]. This relationship implies 419 that  $B_{\rm tot} = N^*B(M) = Q$ , where Q is a constant, and has been 420 shown to hold in both mammalian and vascular plant communi- $_{421}$  ties [56, 58, 59]. Figure 4A shows that both  $F^*$  and  $H^*$  scale as  $_{422}~M^{-\eta}$  over a wide range of organism sizes and Figure 4B shows  $_{423}$  that  $F^*B$  is nearly constant over this same range. This result is  $_{465}$ 423 that I be heady state values of 466 nism, we begin by noting that a theoretical upper bound on  $^{423}$  the NSM combined with the derived timescales naturally give  $^{467}$  mammalian body size is given by  $\epsilon_{\sigma}=0$ , where mammals 426 rise to energy equivalence.

428 maximum observed body sizes for mammals, suggesting that 470 phant. Next we examine to what extent a more realistic upper 428 maximum observed body sizes for maininals, suggesting that  $_{470}$  phant. Next we examine to what extent a more realistic appear this maximum is a hard limit where deviations outside of this  $_{471}$  bound to body mass may serve as an evolutionary attractor,  $_{430}$  range are energetically suboptimal. In the framework of our  $_{472}$  thus providing a suitable within-lineage mechanism for Cope's rule at a finite mass, and occur at the same scale where the  $_{472}$  mogeneous population to invasion by a mutated subset of the  $_{472}$  mogeneous population (denoted by ') where individuals have a modified  $_{472}$  proportion of body fat  $_{473}$  proportion of body fat  $_{474}$  proportion of body fat  $_{475}$  population (denoted by ') where individuals have a modified  $_{475}$  proportion of body fat  $_{476}$  proportion 435 zero, causing the timescales to become infinite and the rates  $_{477}$  of  $\chi$  the adjusted mass should exceed the amount of body fat, 435  $\mu$  and  $\lambda$  to equal zero. The  $\mu=0$  asymptote occurs first when 478  $1+\chi>\epsilon_{\sigma}$ , and the adjusted time to reproduce must be positive, 437  $f_0M^{\gamma-1}+u_0M^{\zeta-1}=1$ , and corresponds to  $(F^*,H^*,R^*)=479$  which given Equation 4, implies that  $1-\epsilon_{\lambda}^{1-\eta}(1+\chi)^{1-\eta}>0$ . 438 (0,0,0). This point predicts a strong upper bound on mam-439 malian body size and occurs at  $M_{\rm max}=6.54\times10^7$ . Moreover, 480 Together these conditions imply that  $\chi\in(-f_0M^{\gamma-1},1/\epsilon_{\lambda}-1)$  481 where the upper bound approximately equals 0.05. The modi-

449 bial life where an upper and lower bound on bacterial size, and 450 an upper bound on single cell eukaryotic size, is predicted from

We contend that the NSM provides a mechanistic underdurance hypothesis, which contends that larger body size, with 458 tion (Fig. 4A, inset), and classic resource competition theory consequent lower metabolic rates and increased ability to main- 459 predicts that the species surviving on the lowest resource abun-Over evolutionary time, terrestrial mammalian lineages show a 462 that larger mammals have an intrinsic competitive advantage

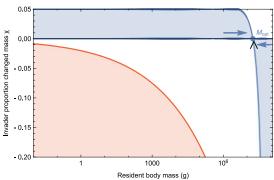


Fig. 5. Invasion feasibility for organisms with a proportional change in mass  $\chi$  against a population with a resident body mass M. The blue region denotes proportions of modified mass  $\chi$  resulting in successful invasion. The red region denotes values of  $\chi$  that result in a mass that is below the starvation threshold and are thus infeasible. Arrows point to the predicted optimal mass from our model  $M_{\rm opt} = 1.748 \times 10^7$ which may serve as an evolutionary attractor for body mass. The black wedge points to the largest body mass known for terrestrial mammals (Deinotherium spp.) at  $1.74 \times 10^7$ g [54].

To examine whether the NSM could provide such a mechato energy equivalence.

Our model shows that the equivalence breaks down at the  $^{468}_{469}$  at  $M=8.3\times10^8$ , or 120 times the mass of a male African ele $_{483}$   $\sigma(M')$ , recovery  $\rho(M')$ , and the maintenance of both starv-  $_{511}$  and provides independent theoretical support for the observa- $_{484}$  ing  $\delta(M')$  and full consumers  $\beta(M')$ . Importantly,  $\epsilon_{\sigma}$ , which  $_{512}$  tion of a 'maximum body size attractor' for North American 485 determines the point along the growth curve that defines the 513 mammals outlined by Alroy [52]. While the state of the envi-486 body composition of starved foragers, is assumed to remain un-514 ronment, as well as the competitive landscape, will determine 487 changed for the invader population (see SI for detailed deriva-515 whether specific body sizes are selected for or against [55], we tions of invader rates).

489 490 invasion, we determine which consumer has a lower steady- 518 evolution of larger body size among terrestrial mammals. 491 state resource density for a given value of  $\chi$ , again with the 519 492 expectation that populations able to survive on lower resource 520 growth, and reproduction are important elements that influence 493 densities have a competitive advantage [61]. We find that for 521 the dynamics of all populations [11]. The NSM is a minimal and  $_{494}~M \le 1.748 \times 10^7 \mathrm{g}$ , having additional body fat  $(\chi > 0)$  results in  $_{522}$  general model that incorporates the dynamics of starvation and 495 a lower steady state resource density  $(R'^* < R^*)$ , such that the 523 recovery that are expected to occur in resource-limited envi-496 invader has an intrinsic competitive advantage over the resident 524 ronments. By incorporating allometric relations between the 497 population (Fig. 5). However, for  $M > 1.748 \times 10^7$  g, leaner 525 rates in the NSM, we found: (i) different organismal masses 498 individuals ( $\chi < 0$ ) have lower resource steady state densities, 526 have distinct population dynamic regimes, (ii) allometrically-499 switching the advantage for higher values of M.

 $_{501}$   $M_{\rm opt} = 1.748 \times 10^7 {\rm g}$  thus serves as an attractor, such that  $_{529}$  may introduce additional drivers and hard boundaries on the  $_{502}$  the NSM predicts organismal mass to increase if  $M < M_{
m opt}$   $_{530}$  evolution of maximum body size. We suggest that the NSM  $_{503}$  and decrease if  $M>M_{\mathrm{opt}}$ . This value is close to but smaller  $_{531}$  offers a means by which the dynamic consequences of energetic 504 than the asymptotic upper bound for terrestrial mammal body 532 constraints can be assessed using macroscale interactions be-505 size predicted by the NSM, however it is remarkably close to 533 tween and among species. Future efforts will involve exploring 506 independent estimates of the largest land mammals, the early 534 the consequences of these dynamics in a spatially explicit frame-507 Oligocene Indricotherium at ca.  $1.5 \times 10^7$  g and the late Miocene 535 work, thus incorporating elements such as movement costs and  $_{508}$  Deinotherium at ca.  $1.74 \times 10^7$  g [54]. Additionally, our calcula- $_{536}$  spatial heterogeneity, which may elucidate additional tradeoffs  $M_{\text{opt}}$  as a function of mass-dependent physiological rates  $M_{\text{opt}}$  associated with the dynamics of starvation and recovery.

482 fied mass adjusts our model via the altered rates of starvation 510 is similar to theoretical estimates of maximum body size [53], 516 propose that the dynamics of starvation and recovery described To assess the susceptibility of the resident population to 517 in the NSM provide a general within-lineage mechanism for the

The energetics associated with somatic maintenance, 527 determined rates of starvation and recovery appear to minimize The observed switch in susceptibility as a function of  $\chi$  at 528 extinction risk, and (iii) the dynamic consequences of these rates

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