The dynamics of starvation and recovery

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The eco-evolutionary dynamics of species are fundamentally linked to the energetic constraints of its constituent individuals. Of particular importance is the interplay between reproduction and the dynamics of starvation and recovery. We introduce a minimal nutritional state-structured model that incorporates two classes of consumer: nutritionally replete, reproducing consumers, and undernourished, non-reproducing consumers. We obtain strong constraints on starvation and recovery rates by deriving allometric scaling relationships and find that population dynamics are typically driven to a steady state. Moreover, we find that these rates fall within a 'refuge' in parameter space, where the probability of extinction of the consumer population is minimized. We also show that our model provides a natural framework that predicts maximum body size for mammals by determining the relative stability of an otherwise homogeneous population to a competing population with altered percent body fat, providing a principled mechanism for a within-lineage driver of Cope's rule.

8 by the energetic state of individuals, which directly influ-• ences how they invest reserves in uncertain environments. 10 Such behaviors are generally manifested as tradeoffs be-11 tween investing in somatic maintenance and growth, or $_{12}$ allocating energy towards reproduction $^{1-3}$. The timing 13 of these behaviors responds to selective pressure, as the 14 choice of the investment impacts future fitness $^{4-6}$. The 15 influence of resource limitation on an organism's ability 16 to maintain its nutritional stores may lead to repeated 17 delays or shifts in reproduction over the course of an or-18 ganism's life.

The balance between (a) somatic growth and mainte-20 nance, and (b) reproduction depends on resource avail-21 ability⁷. For example, reindeer invest less in calves born 22 after harsh winters (when the mother's energetic state is 23 depleted) than in calves born after moderate winters⁸. 24 Many bird species invest differently in broods during pe-²⁵ riods of resource scarcity compared to normal periods^{9,10}, 26 sometimes delaying or even foregoing reproduction for 27 a breeding season^{1,11,12}. Even freshwater and marine 28 zooplankton have been observed to avoid reproduction ²⁹ under nutritional stress¹³, and those that do reproduce 30 have lower survival rates². Organisms may also separate 31 maintenance and growth from reproduction over space 32 and time: many salmonids, birds, and some mammals 33 return to migratory breeding grounds to reproduce af-34 ter one or multiple seasons in resource-rich environments 35 where they accumulate reserves $^{14-16}$.

Physiology also plays an important role in regulating 37 reproductive expenditures during periods of resource lim-38 itation. Many mammals (47 species in 10 families) ex-39 hibit delayed implantation, whereby females postpone fe-40 tal development until nutritional reserves can be accumu-41 lated^{17,18}. Many other species (including humans) suf-42 fer irregular menstrual cycling and higher abortion rates 43 during periods of nutritional stress 19,20. In the extreme s1 vation, and (b) an energetically deficient (hungry) state 44 case of unicellular organisms, nutrition is unavoidably \$2 H, where the consumer does not reproduce but dies by

The behavioral ecology of all organisms is influenced 45 linked to reproduction because the nutritional state of 46 the cell regulates all aspects of the cell cycle²¹. The ex-47 istence of so many independently evolved mechanisms 48 across such a diverse suite of organisms highlights the 49 universality of the fundamental tradeoff between somatic 50 and reproductive investment.

Including individual energetic dynamics²² in a 52 population-level framework^{22,23} is challenging²⁴, and a 53 common simplifying approach is provided in the clas-54 sic Lotka-Volterra (LV) model, which assumes that 55 consumer population growth rate depends linearly on 56 resource density²⁵. Here, we introduce an alterna-57 tive approach—the Nutritional State-structured Model 58 (NSM)—that accounts for resource limitation through 59 the consequences of starvation. In contrast to the LV 60 model, the NSM incorporates two consumer states: hun-61 gry and full, with only the former susceptible to mor-62 tality and only the latter possessing sufficient energetic 63 reserves to reproduce. Additionally, we incorporate allo-64 metrically derived constraints on reproduction³, incorpo-65 rating the timescales of starvation and recovery. As we 66 shall show, our model makes several important predic-67 tions: (i) the dynamics are typically driven to a refuge 68 far from cyclic behavior and extinction risk, (ii) steady 69 state conditions of the NSM accurately predict measured 70 biomass densities for mammals as well as Damuth's en-71 ergetic equivalence principle, (iii) an allometrically con-72 strained upper-bound for mammalian body size, and (iv) 73 a selective mechanism for the evolution of larger body 74 size, known as Cope's rule.

Nutritional state-structured model (NSM). 76 We begin by defining a minimal Nutritional State-77 structured population Model (NSM), where the consumer 78 population is partitioned into two states: (a) an energet-79 ically replete (full) state F, where the consumer repro-80 duces at a constant rate λ and does not die from star-

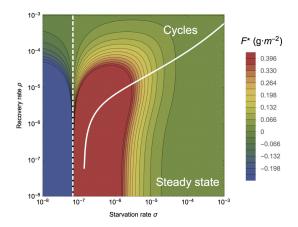


Figure 1: The transcritical (TC; dashed line) and Hopf bifurcation (solid line) as a function of the starvation rate σ and recovery rate ρ for a 100g consumer. These bifurcation conditions separate parameter space into unphysical (left of the TC), cyclic, and steady state dynamic regimes. The colors show the steady state densities for the energetically replete consumers F^* .

starvation at rate μ . The dynamics of the underlying resource R is governed by logistic growth with an intrinsic 85 growth rate α and a carrying capacity C. The rate at 86 which consumers transition between states and consume 87 resources is dependent on their number, the abundance 88 of resources, the efficiency of converting resources into 89 metabolism, and how that metabolism is partitioned be-50 tween maintenance and growth purposes. We provide a 91 physiologically and energetically mechanistic model for 92 each of these dynamics and constants (see the Supple-93 mentary Information (SI)), and show that the system 94 produces a simple non-dimensional form which we de-95 scribe below.

Consumers transition from the full state F to the hun-106 Full consumers eat resources at a constant rate β that ac-107 counts for maximal maintenance and somatic growth (see previous models incorporating starvation²⁹.

117 are governed by the rate equations

$$\dot{F} = \lambda F + \xi \rho R H - \sigma (1 - R) F,
\dot{H} = \sigma (1 - R) F - \xi \rho R H - \mu H,
\dot{R} = \alpha (1 - R) R - (\rho R + \delta) H - \beta F$$
(1)

This system of nondimensional equations follows from a set of first-principle relationships for resource consump-120 tion and growth (see the SI for a full derivation and the 121 dimensional form). Notice that the total consumer den-122 sity F + H evolves according to $\dot{F} + \dot{H} = \lambda F - \mu H$. This 123 resembles the equation of motion for the predator den-124 sity in the LV model³⁰, except that the resource density 125 does not appear in the growth term. The rate of repro-126 duction is independent of resource density because it is 127 assumed that the satiated state of the full consumer al-128 lows it to partition a constant amount of energy towards 129 reproduction, whereas a starved consumer partitions no 130 energy towards reproduction. The rate of reproduction 131 for the total consumer density is dependent on resource 132 density, which determines the size of the full and starved portions of the consumer population. Similarly, the con-134 sumer maintenance terms (δH and βF) are independent 135 of resource density because they represent a minimal en-136 ergetic requirement for consumers in the H and F state, 137 respectively. It follows that model predictions are robust 138 only when R is of the order of 1, which holds for all cases 139 that we explore.

Steady states of the NSM. From the solution 141 to the single internal fixed point (Eq. (2), see Meth-142 ods), an obvious constraint on the NSM is that the re-143 production rate λ must be less than the starvation rate σ , so that the consumer and resource densities are pos-145 itive. The condition $\sigma = \lambda$ thus represents a transcrit-146 ical (TC) bifurcation³¹ that demarcates a physical from 147 an unphysical regime where all steady-state densities be-148 come negative after intersecting the trivial fixed point 97 gry state H at a rate σ —the starvation rate—and also $(F^*, H^*, R^*) = (0, 0, 0)$. The biological implication of 98 in proportion to the absence of resources (1-R) (we 150 the constraint $\lambda < \sigma$ has a simple interpretation—the 99 assume a maximum density, the carrying capacity equal 151 rate at which a macroscopic organism loses mass due to 100 to 1). Conversely, consumers recover from state H to 152 lack of resources is generally much faster than the rate state F at rate $\xi\rho$ and in proportion to R, where ξ rep- 153 of reproduction. As we will discuss below, this inequal-102 resents a ratio between maximal resource consumption 154 ity is a natural consequence of allometric constraints of for 103 and the carrying capacity of the resource. The resources 155 organisms within empirically observed body size ranges. that are eaten by hungry consumers (at rate $\rho R + \delta$) ac- in the physical regime of $\lambda < \sigma$, the fixed point (2) 105 count for their somatic growth (ρR) and maintenance (δ) . 157 may either be a stable node or a limit cycle (Fig. 1). In 158 continuous-time systems, a limit cycle arises when a pair 159 of complex conjugate eigenvalues crosses the imaginary 108 the SI for mechanistic derivations of these rates from re- 160 axis to attain positive real parts³². This Hopf bifurcation source energetics). The NSM represents an ecologically $_{161}$ is defined by $\overline{\mathrm{Det}(\mathbf{S})} = 0$, with \mathbf{S} the Sylvester matrix, motivated fundamental extension of the idealized starv- 162 which is composed of the coefficients of the characterising random walk model of foraging, which focuses on re- 163 tic polynomial of the Jacobian matrix³³. As the system 112 source depletion, to include reproduction and resource re- 164 parameters are tuned to be within the stable regime, but plenishment^{26–28}, and is a more general formulation than 165 close to the Hopf bifurcation, the amplitude of the tran-166 sient cycles becomes large. Given that ecological systems In the mean-field approximation, in which the con- 167 are constantly being perturbed³⁴, the onset of transient 116 sumers and resources are perfectly mixed, their densities 168 cycles, even though they decay with time in the meanfield description, can increase extinction risk^{35–37}.

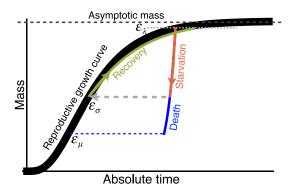


Figure 2: The growth trajectory over absolute time of an intrajectory to $m = \epsilon_{\sigma} \epsilon_{\lambda} M$. Recovery follows the green curve to the replete adult mass, where this trajectory differs from the original growth because only fat is being regrown which requires a longer time to reach $\epsilon_{\lambda}M$. Alternatively, death from starvation follows the blue trajectory to $m = \epsilon_{\mu} \epsilon_{\lambda} M$.

When the starvation rate $\sigma \gg \lambda$, a substantial ²²⁸ 172 reproducing state. Because reproduction is inhibited, 230 not independent parameters, and the biologically rele-173 there is a low steady-state consumer density and a high 231 vant portion of the phase space shown in Fig. 1 is con-180 become stable over time. $_{182} \rho$ increases, such that the range of parameter space $_{240}$ particularly in environments where resources are limit-183 giving rise to cyclic dynamics also increases with higher 241 ing. 184 recovery rates.

186 Results

185

200 (cf. ref. ³⁸).

202 termined by consumer metabolism, which can be used 260 the rate of starvation to vary, we assessed extinction risk 203 to describe a variety of organismal features³⁹. We derive 261 across a range of values for σ and ρ between ca. 10^{-7} to

204 relationships for the rates of reproduction, starvation, re-205 covery, and mortality based on first principles, and as a 206 function of an organism's body size and metabolic rate 207 (see Methods). Because we aim to explore the starvation-208 recovery dynamics as a function of an organism's body 209 mass M, we parameterize these rates in terms of the per-210 cent gain and loss of the asymptotic (maximum) body 211 mass, ϵM , where different values of ϵ define different 212 states of the consumer (Fig. 2: see Methods for deriva-213 tions of allometrically constrained rate equations). Although the rate equations (1) are general and can in 215 principle be used to explore the starvation recovery dy-216 namics for most organisms, here we focus on allomet-217 ric relationships for terrestrial-bound lower trophic level 218 endotherms (see the SI for values), specifically herbivodividual organism as a function of body mass. Initial growth 219 rous mammals, which range from a minimum of $M \approx 1 \mathrm{g}$ follows the black trajectory to an energetically replete repro- 220 (the Etruscan shrew Suncus etruscus) to a maximum of ductive adult mass $m = \epsilon_{\lambda} M$. Starvation follows the red 221 $M \approx 10^7 \mathrm{g}$ (the early Oligocene Indricotheriinae and the 222 Miocene Deinotheriinae). Investigating other classes of 223 organisms would simply involve altering the metabolic 224 exponents and scalings associate with ϵ . Moreover, we 225 emphasize that our allometric equations describe mean 226 relationships, and do not account for the (sometimes con-227 siderable) variance associated with individual species.

As the allometric derivations of the NSM rate laws 171 fraction of the consumers are driven to the hungry non- 229 reveal (see Methods), starvation and recovery rates are 174 steady-state resource density. However, if $\sigma/\lambda \to 1$ from 232 strained via covarying parameters. Given the parameters above, the population is overloaded with energetically- 233 of terrestrial endotherms, we find that the starvation rate 176 replete (reproducing) individuals, thereby promoting 234 σ and the recovery rate ρ are constrained to lie within 177 transient oscillations between the consumer and resource 235 a small region of potential values for the known range $_{178}$ densities (Fig. 1). If the starvation rate is low enough $_{236}$ of body sizes M. Indeed, starvation and recovery rates 179 that the Hopf bifurcation is crossed, these oscillations 237 across all values of M fall squarely in the steady state This threshold occurs at 238 region at some distance from the Hopf bifurcation. This 181 higher values of the starvation rate as the recovery rate 239 suggests that cyclic population dynamics should be rare,

Higher rates of starvation result in a larger flux of the 243 population to the hungry state. In this state, reproduc-244 tion is absent, thus increasing the likelihood of extinc-The allometry of extinction risk. While there are 245 tion. From the perspective of population survival, it is 188 no a priori constraints on the parameters in the NSM, 246 the rate of starvation relative to the rate of recovery that 189 we expect that each species should be restricted to a dis- 247 determines the long-term dynamics of the various species 190 tinct portion of the parameter space. We use allomet- 248 (Fig. 1). We therefore examine the competing effects of 101 ric scaling relations to constrain the covariation of rates 249 cyclic dynamics vs. changes in steady-state density on 102 in a principled and biologically meaningful manner (see 250 extinction risk, both as functions of σ and ρ . To this 193 Methods). Allometric scaling relations highlight com- 251 end, we computed the probability of extinction, where we 194 mon constraints and average trends across large ranges 252 define extinction as a population trajectory falling below 195 in body size and species diversity. Many of these rela- 253 one fifth of the allometrically constrained steady state at 196 tions can be derived from a small set of assumptions and 254 any time between $t=10^8$ and $t=10^{10}$. This procedure is 107 in the Methods we describe our framework to determine 255 repeated for 50 replicates of the continuous-time system 198 the covariation of timescales and rates across a range of 256 shown in Eq. 1 for organisms with mass ranging from 10^2 199 body sizes for each of the key parameters of our model 257 to 10⁶ grams. In each replicate the initial densities are 258 chosen to be (XF^*, XH^*, R^*) , with X a random vari-Nearly all of the rates described in the NSM are de-259 able that is uniformly distributed in [0,2]. By allowing

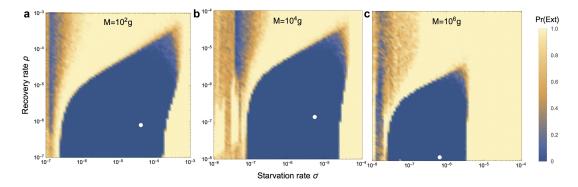


Figure 3: Probability of extinction for a consumer with (a) $M = 10^2$ g, (b) $M = 10^4$ g, and (c) $M = 10^6$ g as a function of the starvation rate σ and recovery rate ρ , where the initial density is given as (XF^*, XH^*, R^*) , where X is a random uniform variable in [0,2]. Note the change in scale for $M=10^4$ and $M=10^6$ g. Extinction is defined as the population trajectory falling below 0.2× the allometrically constrained steady state. The white points denote the allometrically constrained starvation and recovery rate.

²⁶² 10⁻³. As expected, higher rates of extinction correlate ³⁰¹ suggests that the NSM dynamics may both drive—and 263 with both high values of σ if ρ is small, and high values 302 constrain—natural animal populations. 264 of ρ if σ is small. For low values of σ and high values of ρ , 303 265 the increased extinction risk results from transient cycles 266 with larger amplitudes as the system nears the Hopf bi-267 furcation (Fig. 3). For high values of σ and low values of ρ , increased extinction risk arises because of the decrease 269 in the steady-state consumer population density (Figs. 270 1b, 3). This interplay creates an 'extinction refuge', such ₂₇₁ that for a constrained range of σ and ρ , extinction prob-272 abilities are minimized.

280 small window supports the possibility that a selective 319 and has been shown to hold in both mammalian and vas-285 that serve to minimize extinction risk. This selection 324 illustrates that the steady state values of the NSM com-287 metabolic rates, and biomass maintenance efficiencies. 326 energy equivalence. 288 We also find that as body size increases, the amount 327 Our model shows that energetic equivalence breaks 289 of low extinction risk parameter space becomes smaller 328 down at large M suggesting that this maximum is a hard 290 (Fig. 3a-c), suggesting that the population dynamics 329 limit where deviations outside of this range are energeti-291 of larger organisms are more sensitive to variability in 330 cally suboptimal. With respect to predicted steady state 202 physiological rates controlling starvation and recovery. 331 densities, the total metabolic rate of F and H becomes 293 This finding is in accordance with, and may serve as 332 infinite at a finite mass, and occurs at the same scale 204 contributing support for, observations of increased ex- 333 where the steady state resources vanish (Fig. 4). This 205 tinction risk among larger mammals⁴⁰. Moreover, larger 334 asymptotic behavior is governed by body sizes at which 206 body sizes decrease the steady state resource density, 335 ϵ_{μ} and ϵ_{λ} equal zero, causing the timescales to become 207 such that fluctuations for larger organisms will be more 336 infinite and the rates μ and λ to equal zero. A theoretical 298 likely to drive resources to extinction. To summarize, 337 upper bound on mammalian body size is given by $\epsilon_{\sigma}=0$, 209 our finding that the allometrically-determined param- 338 where mammals are entirely composed of metabolic re-300 eters fall within this low extinction probability region 330 serves, and this occurs at $M = 8.3 \times 10^8$, or 120 times the

Predicting Damuth's Law and body size lim-305 its. The NSM correctly predicts that species with 306 smaller masses have larger steady-state population densi-307 ties (Fig. 4a). Similar predictions have been made for car-308 nivore populations using alternative consumer-resource models⁴¹. Moreover, we show that the NSM provides 310 independent theoretical support for the energy equiva-311 lence hypothesis and Damuth's Law^{42–44}. The energy We find that the allometrically constrained values 312 equivalence hypothesis argues that the total energy use, of σ and ρ fall squarely within the extinction refuge $^{313}B_{\rm tot}$, of a population is constant independent of species across a range of M (Fig. 3a-c, white points). These 314 size $^{42-44}$. This hypothesis is based on observations show- $_{276}$ values are close enough to the Hopf bifurcation to avoid $_{315}$ ing that the steady state abundance, N^* , of a species 277 low steady-state densities, and far enough away to 316 is proportional to the inverse of individual metabolism, avoid large-amplitude transient cycles. The feature that 317 such that $N^* \propto M^{-3/4}/B_0^{43,44}$. This relationship imalization alloweric values of σ and ρ fall within this relatively 318 plies that $B_{\rm tot} = N^*B(M) = Q$, where Q is a constant, 281 mechanism has constrained the physiological conditions 320 cular plant communities 42-44. Figure 4a shows that both 282 that drive starvation and recovery rates within popu- 321 F^* and H^* scale as $M^{-\eta}$ over a wide range of organism 283 lations. Such a mechanism would select for organism 322 sizes and Figure 4b shows that F^*B is nearly constant 284 physiology that generates appropriate σ and ρ values 323 over this same range. This result is remarkable because it 286 could occur via the tuning of body fat percentages, 325 bined with the derived timescales naturally give rise to

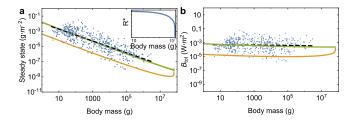


Figure 4: (a) Consumer steady states F^* (green) and H^* (orange) as a function of body mass. Inset: Resource steady state R^* as a function of consumer body mass. (b) Total energetic use B_{tot} of consumer populations at the steady state as a function of body mass. The data are from Damuth⁴² and have been converted to total population metabolism using the allometric relationships for metabolic rate (see the SI and Refs. $^{45-47}$).

340 mass of a male African elephant. In contrast, the $\mu=0$ asymptote occurs first when $f_0 M^{\gamma-1} + u_0 M^{\zeta-1} = 1$, and 342 corresponds to $(F^*, H^*, R^*) = (0, 0, 0)$. This point pre-343 dicts a more realistic upper bound on mammalian body 344 size and occurs at $M_{\rm max} = 6.54 \times 10^7$. Moreover, $M_{\rm max}$, 345 which is entirely determined by the population-level con-346 sequences of energetic constraints, is within an order of 347 magnitude of the maximum body size observed in the 348 North American mammalian fossil record⁴⁸, as well as 349 the mass predicted from an evolutionary model of body 350 size evolution⁴⁹. It should be noted that the asymptotic 351 behavior and predicted upper bound depend only on the 352 scaling of body composition and are independent of the 353 resource parameters. We also note that the prediction $_{354}$ of an asymptotic limit on mammalian size parallels work $_{355}$ on microbial life where an upper and lower bound on 356 bacterial size, and an upper bound on single cell eukary-357 otic size, is predicted from similar growth and energetic 358 scaling relationships^{3,50}.

360 port constraints are widely thought to place strict bound- 394 or against. 361 aries on biological scaling 39,51,52 and thereby lead to 395 362 specific predictions on the minimum possible body size $\frac{3}{2}$ 06 closely related species: a resident species of mass M, 363 for organisms⁵³. Above this bound, a number of ener- 397 and a competing species (denoted by ') where individ-364 getic and evolutionary mechanisms have been explored 398 uals have a different proportion of body fat such that 365 to assess the costs and benefits associated with larger 399 $M' = M(1+\chi)$. If $\chi < 0$, individuals within the com-366 body masses, particularly for mammals. One impor- 400 peting population have fewer metabolic reserves, and if 367 tant such example is the fasting endurance hypothesis, 401 $\chi > 0$, individuals have more metabolic reserves than the 368 which contends that larger body size, with consequent 402 resident species. For the allowable values of χ the ad-369 lower metabolic rates and increased ability to maintain 403 justed mass should exceed the minimal amount of body 370 more endogenous energetic reserves, may buffer organ-404 fat, $1+\chi>\epsilon_{\sigma}$, and the adjusted time to reproduce is most endogeneous energetic receives, may state organize that χ is a significant tenderal fluctuations in resource available for the positive, which given Equation 4, implies that a bility 54 . Over evolutionary time, terrestrial mammalian and $1 - \epsilon_{\lambda}^{1-\eta} (1+\chi)^{1-\eta} > 0$. Together these conditions imply lineages show a significant trend towards larger body size for that $\chi \in (-f_0 M^{\gamma-1}, 1/\epsilon_{\lambda} - 1)$ where the upper bound 374 known as Cope's rule^{48,49,55,56}, and it is thought that 408 approximately equals 0.05 and the lower bound is mass-375 within-lineage drivers generate selection towards an op- 409 dependent. The modified mass adjusts our model via 376 timal upper bound of roughly 10^7 grams⁴⁸, a value that 410 the altered rates of starvation $\sigma(M')$, recovery $\rho(M')$, 377 is likely limited by higher extinction risk for large taxa 411 and the maintenance of both starving $\delta(M')$ and full 378 over longer timescales ⁴⁹. These trends are thought to 412 consumers $\beta(M')$ (see the SI for detailed derivations of

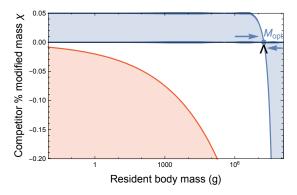


Figure 5: Competitive outcomes for a resident species with body mass M vs. a closely related competing species with modified body mass $M' = M(1 + \chi)$. The blue region denotes proportions of modified mass χ resulting in exclusion of the resident species. The red region denotes values of χ that result in a mass that is below the starvation threshold and are thus infeasible. Arrows point to the predicted optimal mass from our model $M_{\rm opt} = 1.748 \times 10^7$, which may serve as an evolutionary attractor for body mass. The black wedge points to the largest body mass known for terrestrial mammals (Deinotherium spp.) at $1.74 \times 10^7 g^{55}$.

380 availability⁵⁶; however the underpinning energetic costs 381 and benefits of larger body sizes, and how they influ-382 ence dynamics over ecological timescales, have not been 383 explored.

The NSM predicts that the steady state resource 385 density R^* decreases with increasing body size of the 386 consumer population (Fig. 4a, inset), and classic re-387 source competition theory predicts that the species sur-388 viving on the lowest resource abundance will outcompete others^{57–59}. Thus, the combined NSM steady-state 390 dynamics and allometric timescales predict that larger 391 mammals have an intrinsic competitive advantage given 392 a common resource, but these absolute limits do not offer A mechanism for Cope's rule Metabolite trans- 393 a mechanism by which larger body sizes are selected for

We directly assess competitive outcome between two ₃₇₉ be driven by a combination of climate change and niche ₄₁₃ competitor rates). Importantly, ϵ_{σ} , which determines the 414 point along the growth curve that defines the body com- 472 resource-limited environments. By incorporating allo-415 position of starved foragers, is assumed to remain un- 473 metric relations between the rates in the NSM, we found: 416 changed for the competing population.

418 to competitive exclusion, we determine which consumer 476 namic consequences of these rates may introduce ad- 419 pushes the steady-state resource density R^* to lower val- 477 ditional drivers and hard boundaries on the evolution 420 ues for a given value of χ , with the expectation that a 478 of maximum body size, and (iii) a selective mechanism 421 population capable of surviving on lower resource den- 479 for the evolution of larger body sizes known as Cope's 422 sities has a competitive advantage⁵⁷. We find that for 480 rule. We suggest that the NSM offers a means by which 423 $M \le 1.748 \times 10^7 {
m g}$, having additional body fat $(\chi > 0)$ 481 the dynamic consequences of energetic constraints can 424 results in a lower steady state resource density $(R'^* < 482 \text{ be assessed using macroscale interactions between and})$ $_{425}$ R^*), such that the competitor has an intrinsic advan- $_{483}$ among species. Future efforts will involve exploring the 426 tage over the resident species (Fig. 5). However, for 484 consequences of these dynamics in a spatially explicit 427 $M>1.748\times 10^7{
m g}$, leaner individuals ($\chi<0$) have lower 485 framework, thus incorporating elements such as move-428 resource steady state densities, switching the advantage 486 ment costs and spatial heterogeneity, which may eluci-429 from having more metabolic reserves to having less.

The observed switch in susceptibility as a function of 488 of starvation and recovery. 431 χ at $M_{\rm opt} = 1.748 \times 10^7 {\rm g}$ thus serves as an attractor, 432 such that the NSM predicts organismal mass to increase 433 if $M < M_{\rm opt}$ and decrease if $M > M_{\rm opt}$. This value is 434 close to but smaller than the asymptotic upper bound 435 for terrestrial mammal body size predicted by the NSM, 436 however it is remarkably close to independent estimates 437 of the largest land mammals, the early Oligocene 438 Indricotherium at ca. 1.5×10^7 g and the late Miocene 439 Deinotherium at ca. $1.74 \times 10^7 \text{g}^{55}$. Additionally, our 440 calculation of M_{opt} as a function of mass-dependent 441 physiological rates is similar to theoretical estimates of ⁴⁴² maximum body size⁴⁹, and provides independent theo-443 retical support for the observation of a 'maximum body 444 size attractor' for North American mammals explored by 445 Alroy⁴⁸. While the state of the environment, as well as 490 where $A = (\lambda \xi \rho + \mu \sigma)$ and $B = (\beta \mu \xi + \delta \lambda \xi - \lambda \mu)$. The stabil-446 the competitive landscape, will determine whether spe- 491 ity of this fixed point is determined by the Jacobian matrix J, 447 cific body sizes are selected for or against 56 , we propose 492 where each matrix element $J_{ij} = \partial \dot{X}_i/\partial X_j$ when evaluated 448 that the dynamics of starvation and recovery described 493 at the internal fixed point, and X is the vector (F, H, R). 449 in the NSM provide a general selective mechanism for the 494 The parameters in Eq. (1) are such that the real part of the 450 evolution of larger body size among terrestrial mammals. 495 largest eigenvalue of $\bar{\bf J}$ is negative, so that the system is sta-

452 Discussion

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453 Complementary previous work has analyzed the trade-454 offs associated with carnivory in terms of the energy ex-457 considerations of energy intake going to zero⁶⁰. Simi-458 larly, it has been shown in the past that the largest body 459 size within a particular lineage should be a function of 460 the metabolic normalization constant and a critical death 461 rate as supported by data. This approach, which con-462 sider the energetics associated with active metabolism, 463 provides a complementary prediction for maximum size 464 from energetics⁶¹, and also suggests that future connec-465 tions between the scaling of fat and muscle mass should 466 be systematically connected with B_0 .

The energetics associated with somatic maintenance, 468 growth, and reproduction are important elements that 469 influence the dynamics of all populations¹¹. The NSM 470 is a general model that incorporates the dynamics of 471 starvation and recovery that are expected to occur in

474 (i) allometrically-determined rates of starvation and re-To assess the susceptibility of the resident species 475 covery appear to minimize extinction risk, (ii) the dy-487 date additional tradeoffs associated with the dynamics

Methods

Analytical solution to the NSM Equation (1) has three fixed points: two trivial fixed points at (F^*, H^*, R^*) = (0,0,0) and (0,0,1), and one non-trivial, internal fixed point

$$F^* = (\sigma - \lambda) \frac{\alpha \lambda \mu^2 (\mu + \xi \rho)}{A(\lambda \rho B + \mu \sigma (\beta \mu + \lambda (\delta + \rho)))},$$

$$H^* = (\sigma - \lambda) \frac{\alpha \lambda^2 \mu (\mu + \xi \rho)}{A(\lambda \rho B + \mu \sigma (\beta \mu + \lambda (\delta + \rho)))},$$

$$R^* = (\sigma - \lambda) \frac{\mu}{A}.$$
(2)

496 ble with respect to small perturbations from the fixed point. 497 Because this fixed point is unique, it is the global attractor 498 for all population trajectories for any initial condition where 499 the resource and consumer densities are both nonzero.

Metabolic scaling relationships The scaling relation 500 455 penditure on hunting and prey size, where a maximum 501 between an organism's metabolic rate B and its body mass M456 carnivore size between 700 and 1,000 is predicted from 502 at reproductive maturity is known to scale as $B = B_0 M^{\eta 62}$ 503 where the scaling exponent η is typically close to 2/3 or 3/4 504 for metazoans (e.g., ref. 39), and has taxonomic shifts for uni-505 cellular species between $\eta \approx 1$ in eukaryotes and $\eta \approx 1.76$ in 506 bacteria^{3,63}

Several efforts have shown how a partitioning of B be-508 tween growth and maintenance purposes can be used to 509 derive a general equation for both the growth trajectories 510 and growth rates of organisms ranging from bacteria to 511 metazoans^{3,45–47,64}. This relation is derived from the sim-512 ple balance condition $B_0 m^{\eta} = E_m \dot{m} + B_m m$, ^{3,45–47,64} where 513 E_m is the energy needed to synthesize a unit of mass, B_m is 514 the metabolic rate to support an existing unit of mass, and m515 is the mass of the organism at any point in its development. 516 This balance has the general solution 3,65

$$\left(\frac{m(t)}{M}\right)^{1-\eta} = 1 - \left[1 - \left(\frac{m_0}{M}\right)^{1-\eta}\right] e^{-a(1-\eta)t/M^{1-\eta}},$$
 (3)

(Fig. 2; see⁴⁷ for a detailed presentation of these timescales). 548 the body fat scaling in our framework. The time that it takes to reach a particular mass ϵM is given by the timescale

$$\tau\left(\epsilon\right) = \ln\left[\frac{1 - \left(m_0/M\right)^{1-\eta}}{1 - \epsilon^{1-\eta}}\right] \frac{M^{1-\eta}}{a\left(1 - \eta\right)},\tag{4}$$

517 where we will define values of ϵ to describe a set of rates 518 within our model. The time to reproduce is given by t_{λ} 519 $\tau(\epsilon_{\lambda})$, where ϵ_{λ} is the fraction of the asymptotic mass where 520 an organism is reproductively mature and should be close to 521 one (typically $\epsilon_{\lambda} \approx 0.95^{45}$). The growth rate then is $\lambda =$ $_{522} \ln \left(\nu\right)/t_{\lambda}$ where ν is the number of offspring produced, and 523 for any constant value of ϵ_{λ} , this rate will scale as $\lambda \propto M^{\eta-1}$ 524 for $M \gg m_0^{3,45-47,64}$.

The rate of recovery $\rho = 1/t_{\rho}$ requires that an organism ac-526 crues sufficient tissue to transition from the hungry to the full 527 state. Since only certain tissues can be digested for energy (for 528 example the brain cannot be degraded to fuel metabolism), 529 we define the rates for starvation, death, and recovery by the 520 timescales required to reach, or return from, specific fractions 540 The starvation rate is then $\sigma = 1/t_{\sigma}$, which scales with 531 of the replete-state mass (see the SI, Table I, for parameterise 550 replete-state mass as $1/M^{1-\eta} \ln{(1-f_0 M^{\gamma}/M)}$. An impor-528 example the brain cannot be degraded to fuel metabolism), 532 zations). We define $m_{\sigma} = \epsilon_{\sigma} M$, where $\epsilon_{\sigma} < 1$ is the fraction 551 tant feature is that σ does not have a simple scaling depen-533 of replete-state mass where reproduction ceases. This fraction 552 dence on λ , which is important for the dynamics that we later 534 will deviate from a constant if tissue composition systemat- 553 discuss. 535 ically scales with adult mass. For example, making use of 536 the observation that body fat in mammals scales with overall 537 body size according to $M_{\rm fat} = f_0 M^{\gamma}$ and assuming that once 538 this mass is fully digested the organism starves, this would 539 imply that $\epsilon_{\sigma} = 1 - f_0 M^{\gamma}/M$. It follows that the recovery 540 timescale, t_{ρ} , is the time to go from $m = \epsilon_{\sigma} \epsilon_{\lambda} M$ to $m = \epsilon_{\lambda} M$ 541 (Fig. 2). Using Eqs. (3) and (4) this timescale is given by sim-542 ply considering an adjusted starting mass of $m'_0 = \epsilon_{\sigma} \epsilon_{\lambda} M$, in 543 which case

$$t_{\rho} = \ln \left[\frac{1 - (\epsilon_{\sigma} \epsilon_{\lambda})^{1 - \eta}}{1 - \epsilon_{\lambda}^{1 - \eta}} \right] \frac{M^{1 - \eta}}{a'(1 - \eta)} \tag{5}$$

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where, for $\eta < 1$, $M = (B_0/B_m)^{1/(1-\eta)}$ is the asymptotic 544 where $a' = B_0/E'_m$ accounts for possible deviations in the mass, $a = B_0/E_m$, and m_0 is mass at birth, itself varying 545 biosynthetic energetics during recovery (see the SI). It should allometrically (see the SI). We now use this solution to define 546 be noted that more complicated ontogenetic models explicitly the timescale for reproduction and recovery from starvation 547 handle storage 46, whereas this feature is implicitly covered by

> To determine the starvation rate, σ , we are interested in the time required for an organism to go from a mature adult that reproduces at rate λ , to a reduced-mass hungry state where reproduction is impossible. For starving individuals we assume that an organism must meet its maintenance requirements by using the digestion of existing mass as the sole energy source. This assumption implies the following simple metabolic balance $\dot{m}E'_m = -B_m m$ or $\dot{m} = -a'm/M^{1-\eta}$ where E'_m is the amount of energy stored in a unit of existing body mass, which differs from E_m , the energy required to synthesis a unit of biomass⁴⁶. Given the replete mass, M, of an organism, the above energy balance prescribes the mass trajectory of a non-consuming organism: $m(t) = Me^{-a't/M^{1-\eta}}$ The timescale for starvation is given by the time it takes m(t)to reach $\epsilon_{\sigma}M$, which gives

$$t_{\sigma} = -\frac{M^{1-\eta}}{a'} \ln\left(\epsilon_{\sigma}\right). \tag{6}$$

The time to death should follow a similar relation, but de-555 fined by a lower fraction of replete-state mass, $m_{\mu} = \epsilon_{\mu} M$ 556 where $\epsilon_{\mu} < \epsilon_{\sigma}$. Suppose, for example, that an organism dies 557 once it has digested all fat and muscle tissues, and that mus-558 cle tissue scales with body mass according to $M_{\text{musc}} = u_0 M^{\zeta}$. This gives $\epsilon_{\mu} = 1 - \left(f_0 M^{\gamma} + u_0 M^{\zeta} \right) / M$. Muscle mass has 560 been shown to be roughly proportional to body mass⁶⁶ in 561 mammals and thus ϵ_{μ} is merely ϵ_{σ} minus a constant. The 562 time to go from starvation to death is the total time to reach **563** $\epsilon_{\mu} M$ minus the time to starve, or $t_{\mu} = -M^{1-\eta} \ln (\epsilon_{\mu}) / a' - t_{\sigma}$, (5) **564** and $\mu = 1/t_{\mu}$.

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