

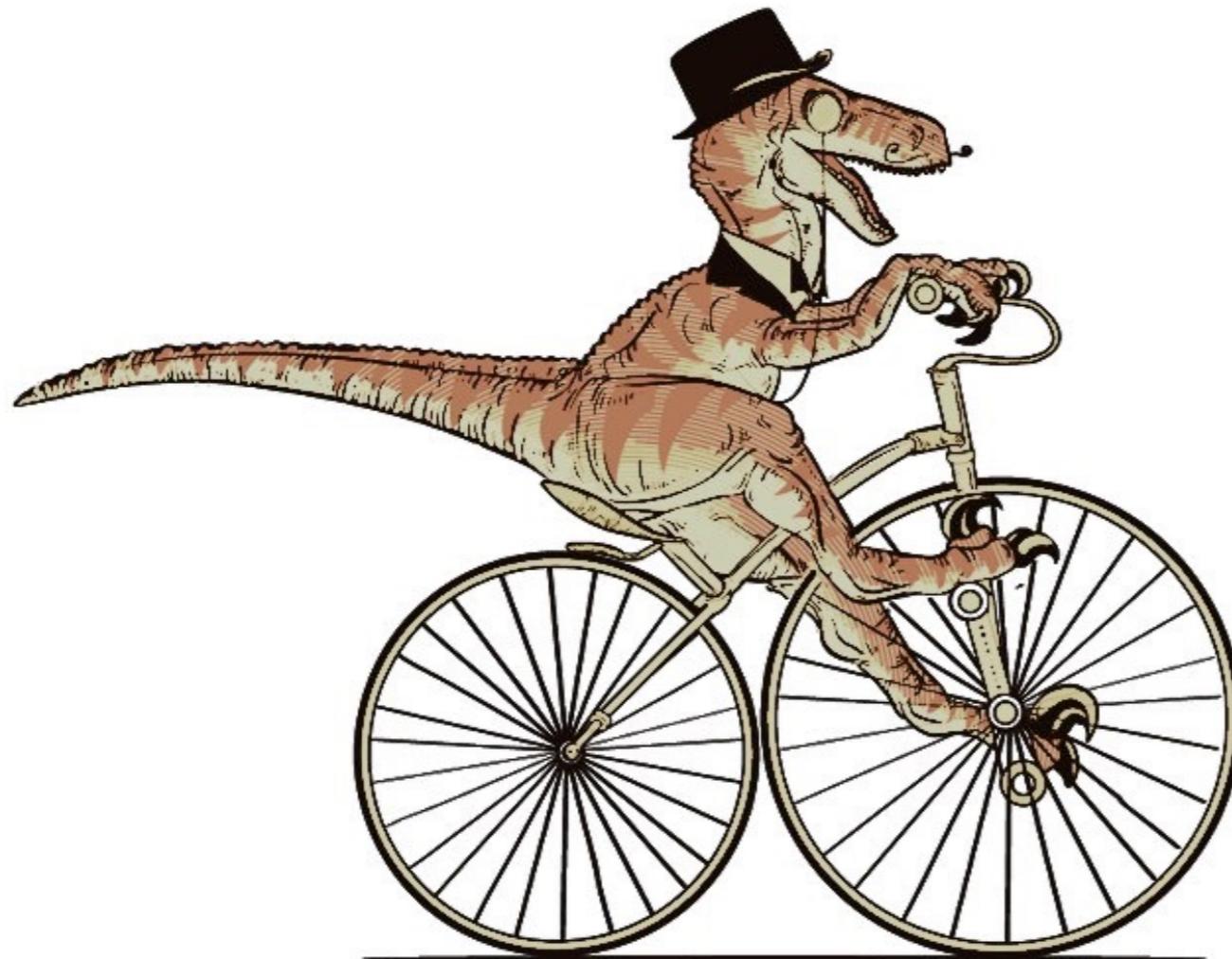
The Natural History of
DINOSAURS

Two websites:

1) The main course website

<http://jdyeakel.github.io/teaching/dinos/>

2) Catcourses (Canvas)



How to contact



BIO 065 01/ESS 065 01

Spring 2024

Home

S24-BIO 065 01/ESS 065 01

Edit

⋮

Announcements

Assignments



Discussions

Grades

People

Pages

Files

Syllabus

Outcomes



Rubrics

Quizzes



Modules



BigBlueButton

Collaborations

Chat

Attendance

New Analytics

Course Feedback

Natural History of Dinosaurs Syllabus (Spring, 2024)

[< Back to Course Page](#)

[> To Course Schedule](#)

- Lecture Time: Monday and Wednesday, 10:30 - 11:45AM
- Lecture Location: SSB 170 unless otherwise notified
- Instructor: Justin D. Yeakel
- Office Hours: W 9-10am or by appointment - SE1 288
- Course Websites

Everything that we do in this course will revolve around two principle websites. In some cases, information will be duplicated on both sites (link to readings), and I will cross-link as necessary.

The [CatCourse Website \(Canvas\)](#): Turning in assignments, grades

The [Dinosaurs Course Website](#): Basic course information, updated schedule, links to readings

- Course Schedule:

[Course Schedule](#)

- Course Description

This course fulfills a lower division requirement for the Ecology and Evolutionary Biology (EEB) emphasis track of the Biological Sciences Major, as well a lower division requirement for the Earth Systems Science Major. This course provides an introduction to the history of life, with an emphasis on the evolutionary ecology of dinosaurs during the Mesozoic Era. Prerequisite: none.

Homework assignments and exams will be submitted on Catcourses unless otherwise stated

Class time and location:

MW 10:30-11:45am

SSB 170

Important information:

Office hours: W 9-10am SEI 288

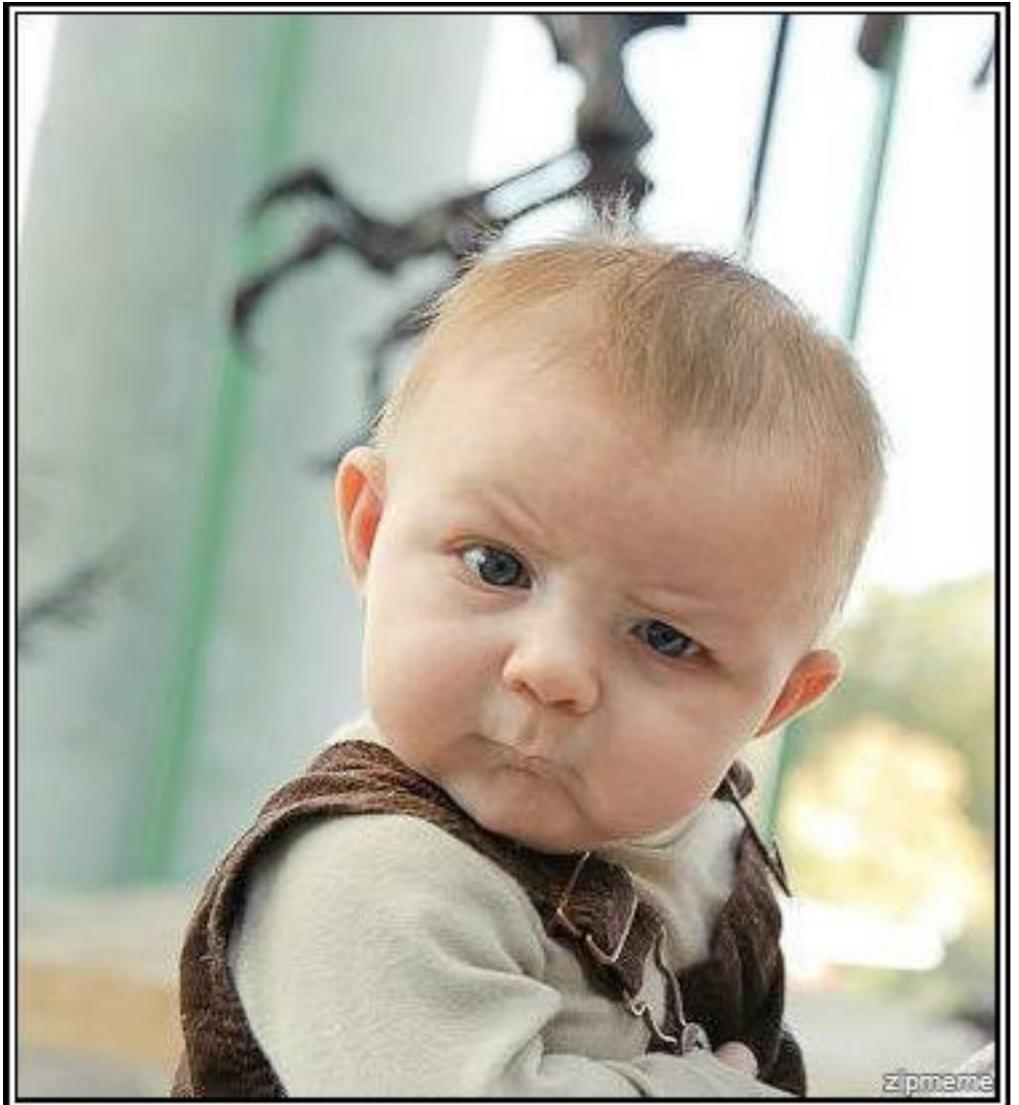
Important Dates

Class add/drop deadline: 2/5



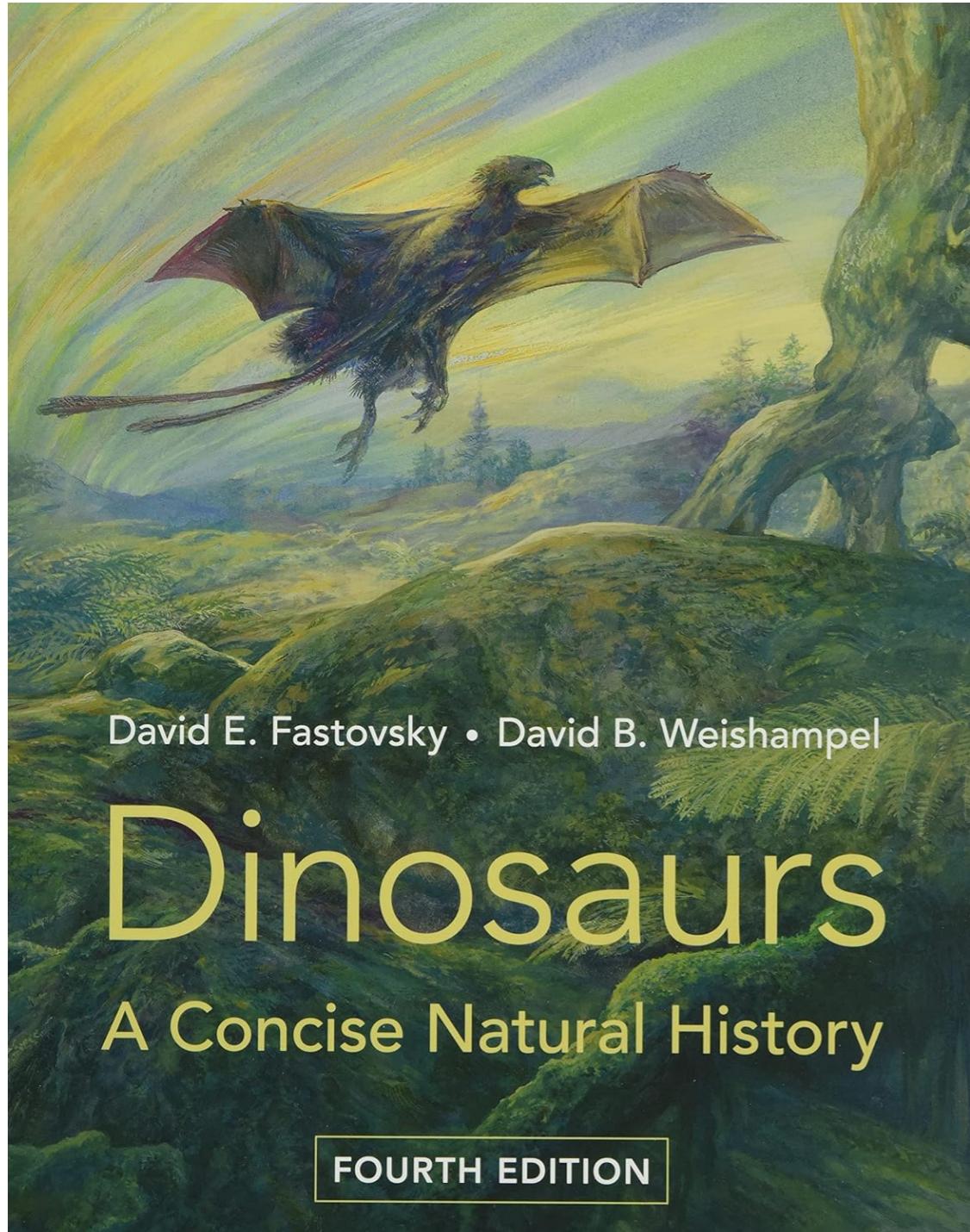
Who are you and why are you here?

Catcourses - participation credit (due tonight at midnight)



Name
Major
Year
What you want from Dinosaurs

The Textbook



Reading for this week:
Fastovsky & Weishampel
Pgs 2-43 (through Chapter 2)

- **Required Textbook:** Fastovsky & Weishampel. Dinosaurs: A Concise Natural History, 4th edition.
- Note: earlier editions (1st, 2nd, 3rd) of this textbook may be used but however the student will be responsible for differences!
- This course is part of the UC Merced Inclusive Access program. You will have immediate access to your digital course materials on the first day of class through CatCourses. The materials that are provided are at a discounted fee, you will see the material charge on your student account in the Mybill section.
- Please do not PRE-Pay for any content or access. If you are prompted to pay for access, try using another web browser such as Chrome or Firefox.
- If you wish to opt-out of Inclusive Access, you must do so by 02/06/24.

Lecture Schedule

Week	Month/Day	Lecture No.	Lecture Topic	Description	Required Readings
1	1/17	W-L1	Introduction	Paleontology and fossils	Fast. Ch. 1,2
2	1/22	M-L2	Evolution 1	Natural selection	Fast. Chp. 3
	1/24	W-L3	Evolution 2	Cladistics	
3	1/29	M-L4	Early life	Cambrian through Devonian	Fast. Chp. 4
	1/31	W-L5	Early life	Devonian through Permian	
4	2/5	M-L6	Dinosaurs	Basal dinosaurs	Fast. Chp. 5
	2/7	W-L7	Thyreophorans 1	Stegosaurs	Fast. Chp. 10
5	2/12	M-L8	Thyreophorans 2	Ankylos	Fast. Chp. 10
	2/14	W	<i>Exam I</i>	In class	

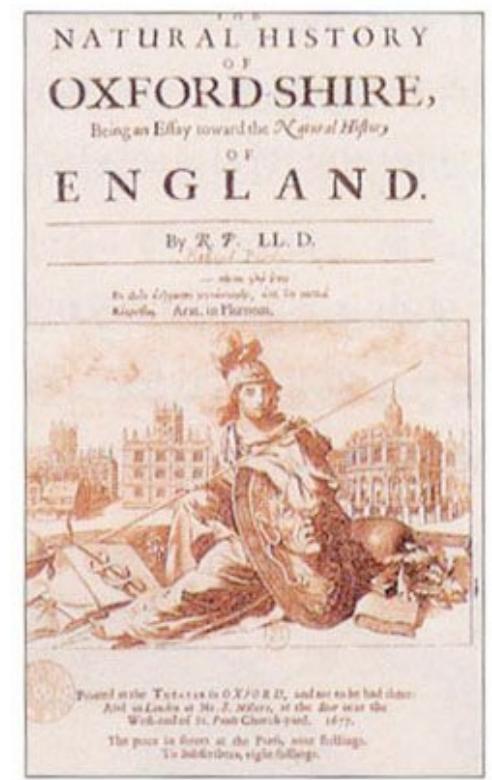
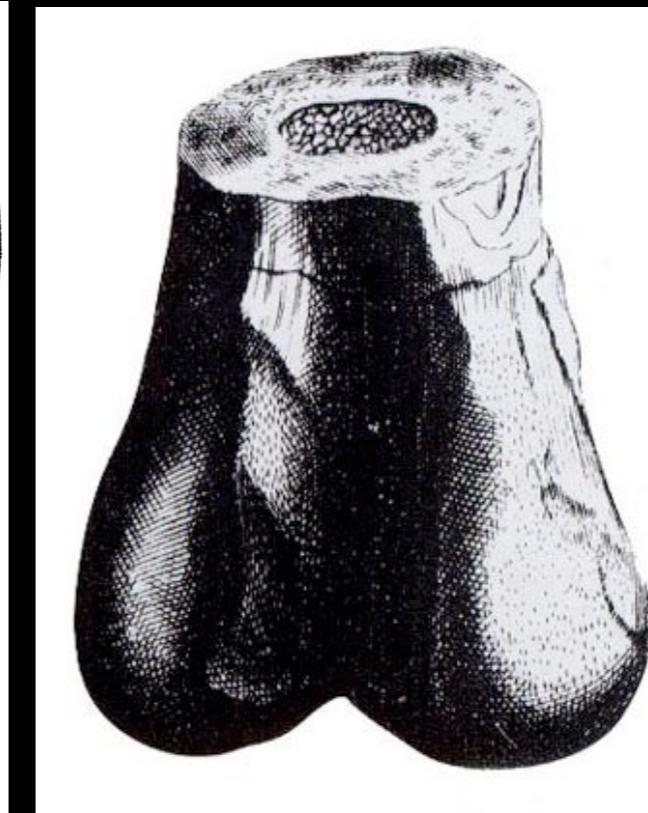
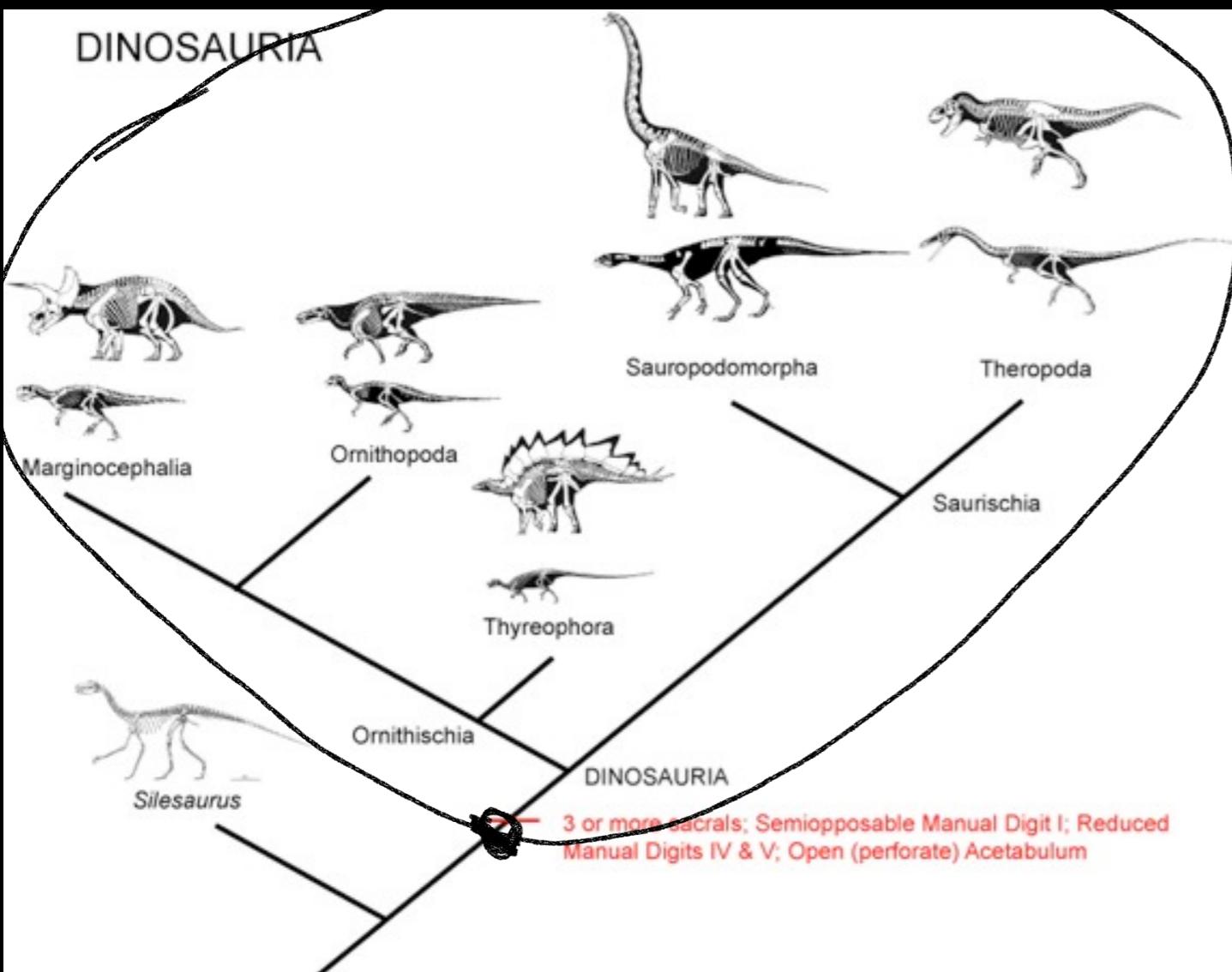
What is a dinosaur???

Scales: Space

Scales: Time

Where do Dinos fit in?

Fossilization and Taphonomy



Scrotum humanum
The distal femur of *Megalosaurus*

Dinosauria:
A monophyletic clade



Go to www.menti.com

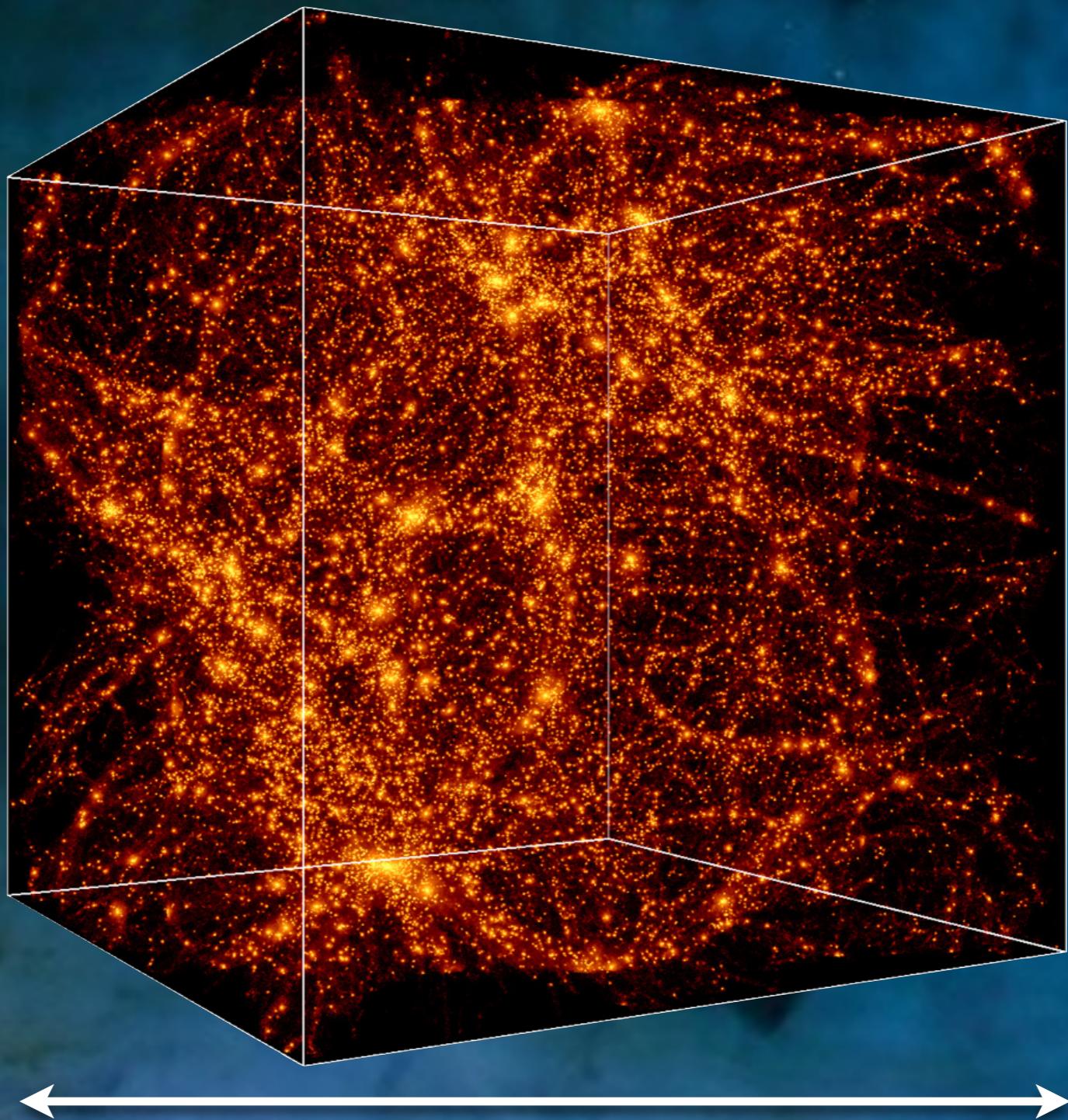
Orders of Magnitude

$$\begin{aligned}10^{10} &= 10000000000 \\&= 10 * 10 * 10 * 10 * 10 * 10 * 10 * 10 * 10\end{aligned}$$

$$\begin{aligned}10^{-10} &= 0.0000000001 \\&= .10 * .10 * .10 * .10 * .10 * .10 * .10 * .10 * .10\end{aligned}$$

Spatial Scales

Orders of Magnitude



$| 10^{26}$

Orders of Magnitude



10^{15}



Orders of Magnitude

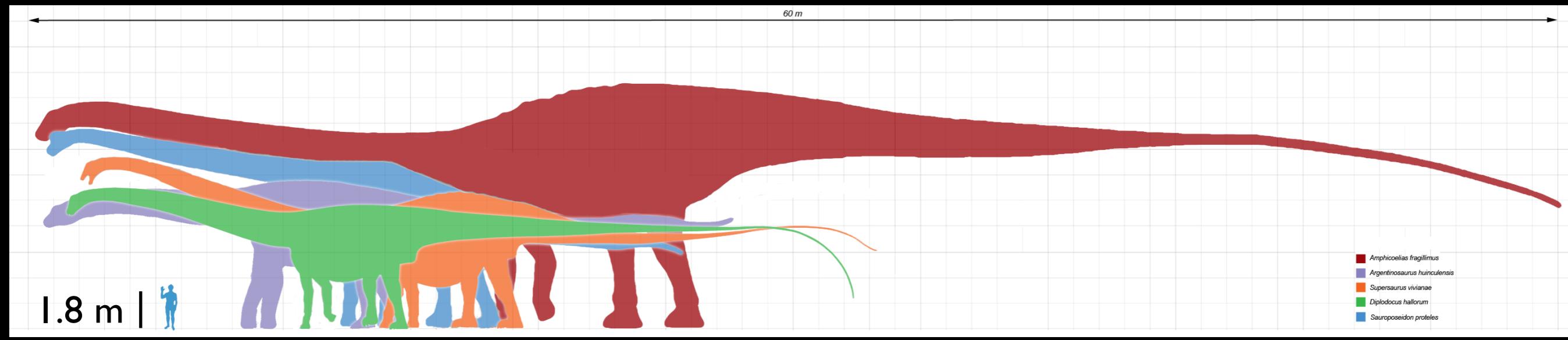


10^{10}

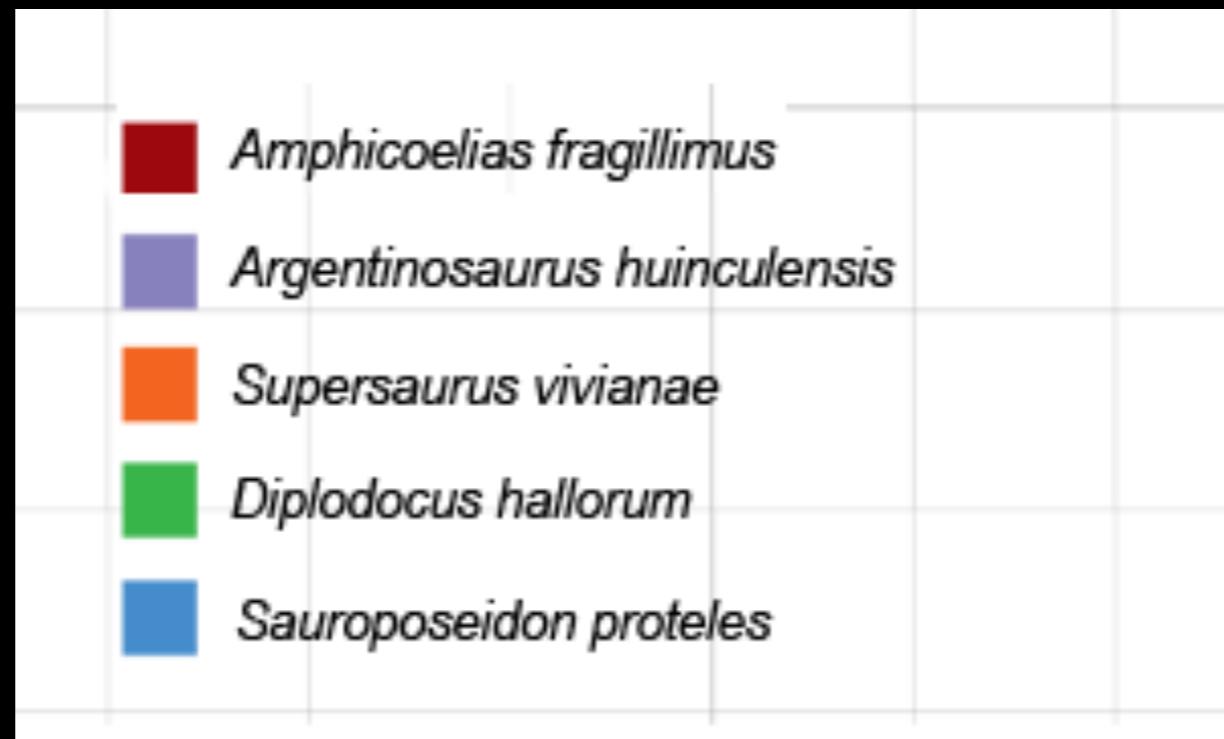
Orders of Magnitude



| 10⁶

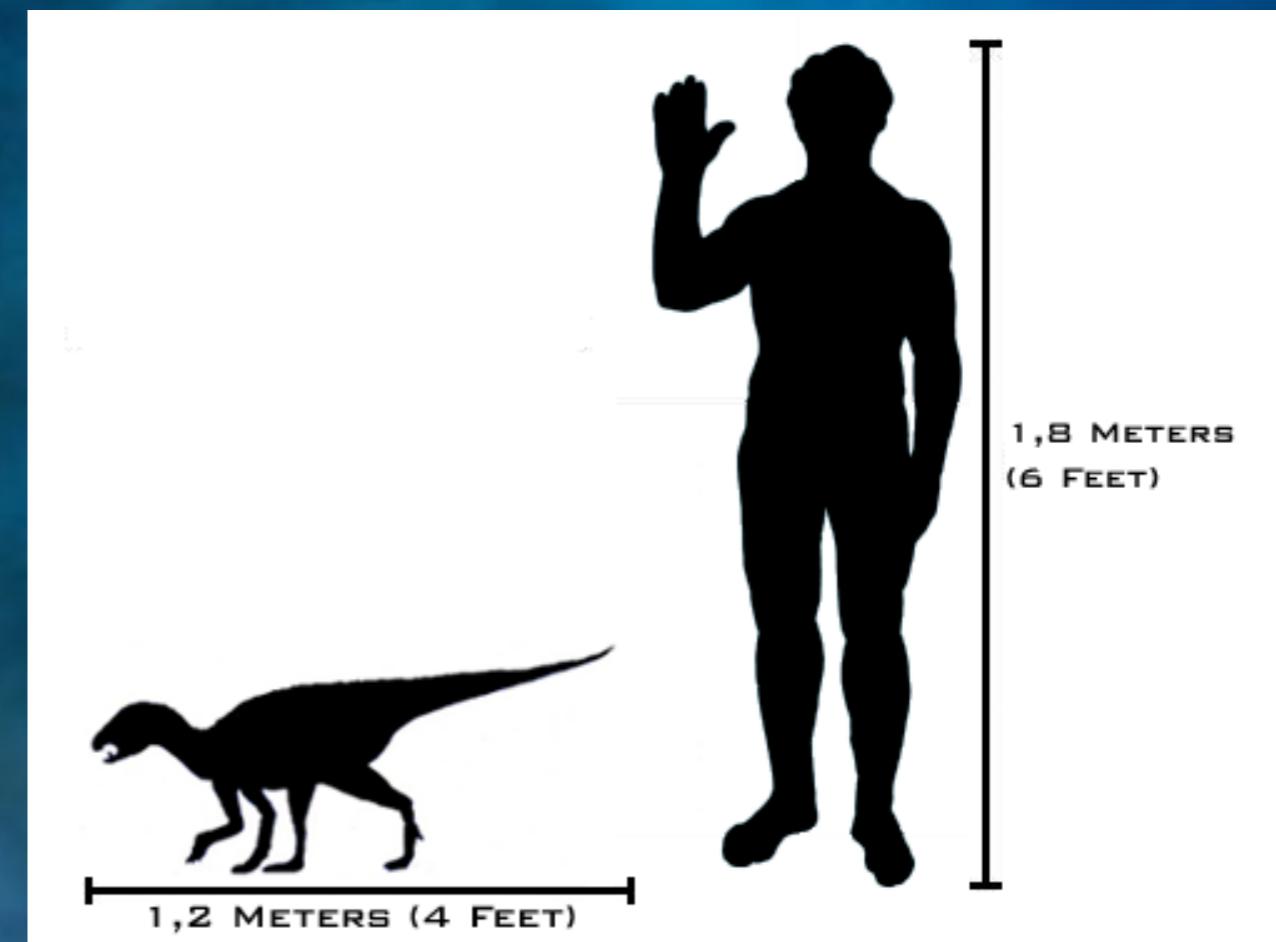


60 m



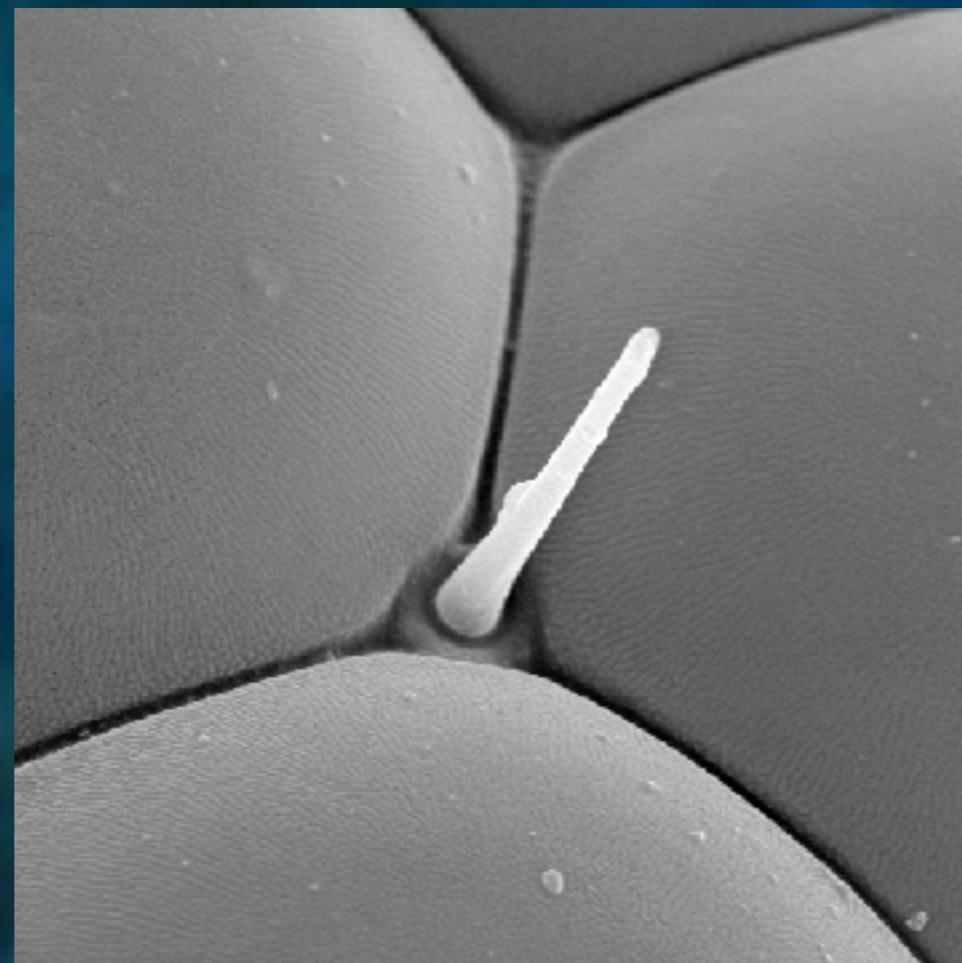


Orders of Magnitude



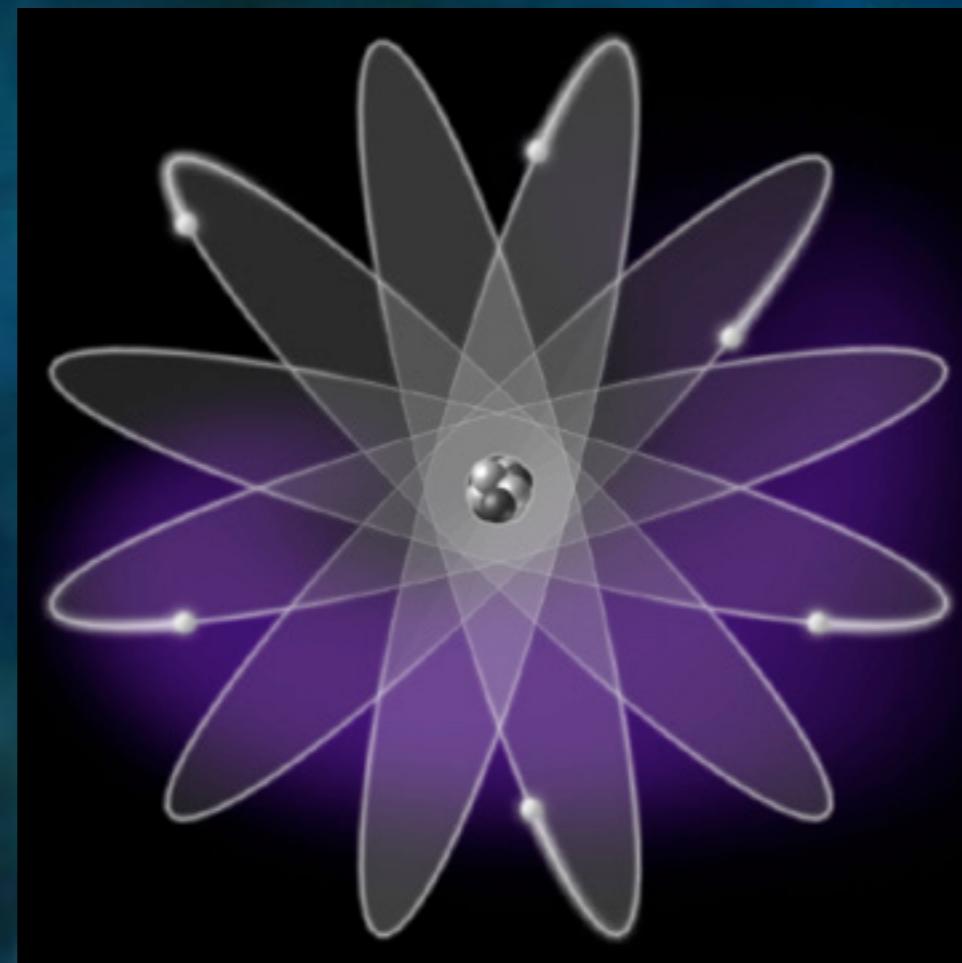
| 10⁰

Orders of Magnitude



10^{-5}

Orders of Magnitude



10^{-10}



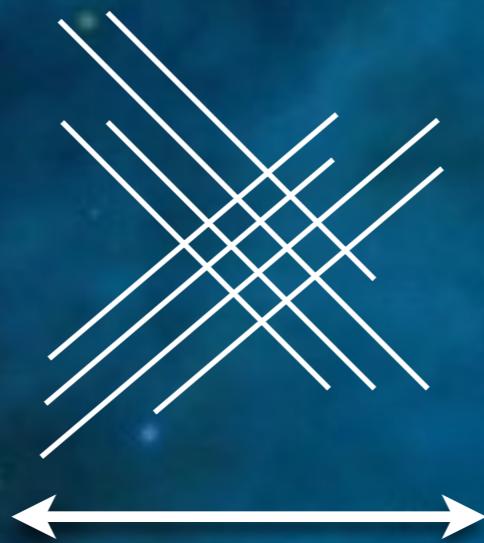
Orders of Magnitude

10^{-15}



Orders of Magnitude

10^{-34}



Planck's Constant
 $h = 6.63 \times 10^{-34}$

Temporal Scales

13.3-13.9 Ga

0

THE BIG BANG

INFLATION

COSMIC MICROWAVE
BACKGROUND
400,000 YEARS AFTER
BIG BANG

GALAXY EVOLUTION
CONTINUES...

FIRST STARS
400,000,000 YEARS
AFTER BIG BANG

THE DARK AGES

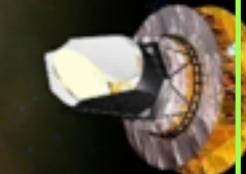
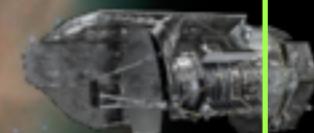
FIRST GALAXIES
1000,000,000 YEARS
AFTER BIG BANG

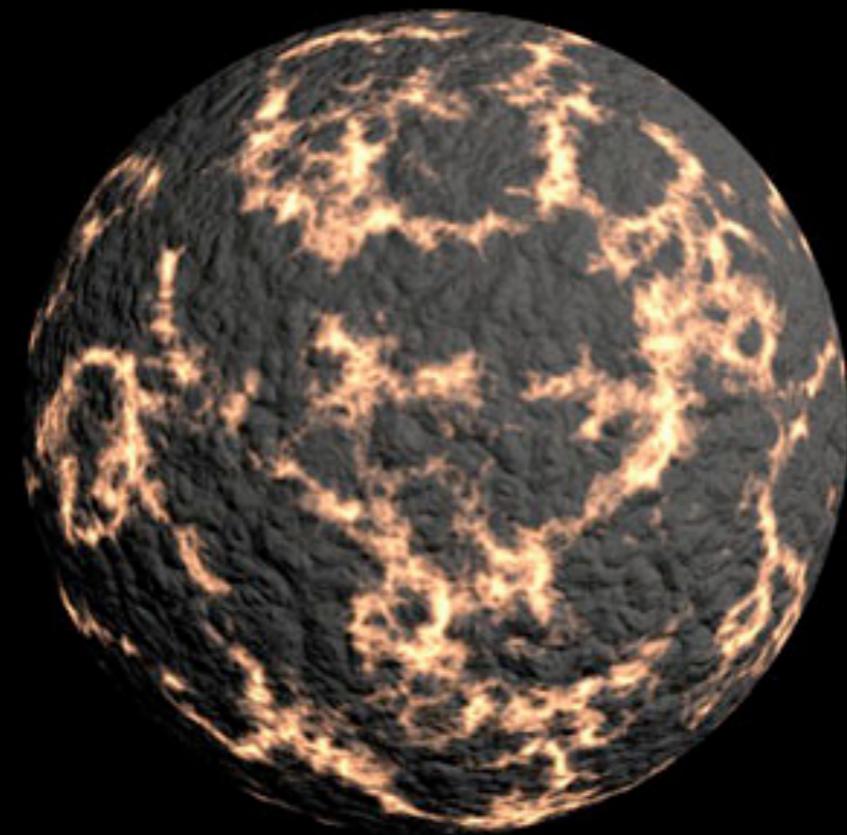
FORMATION OF
THE SOLAR SYSTEM
8,700,000,000 YEARS
AFTER BIG BANG

DARK ENERGY?

life

NOW
13,700,000,000 YEARS
AFTER BIG BANG





4.56 gyr



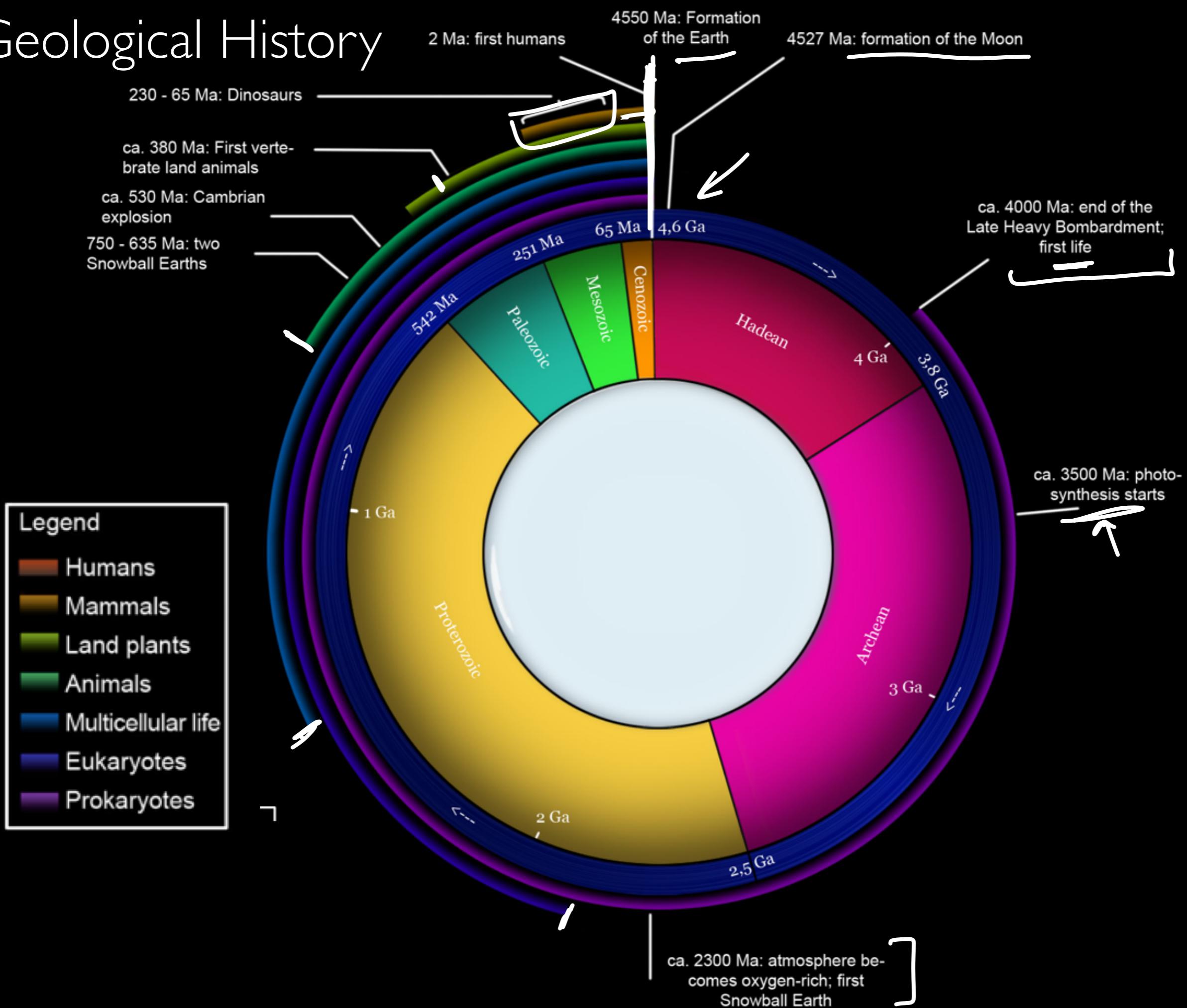
4.527 gyr



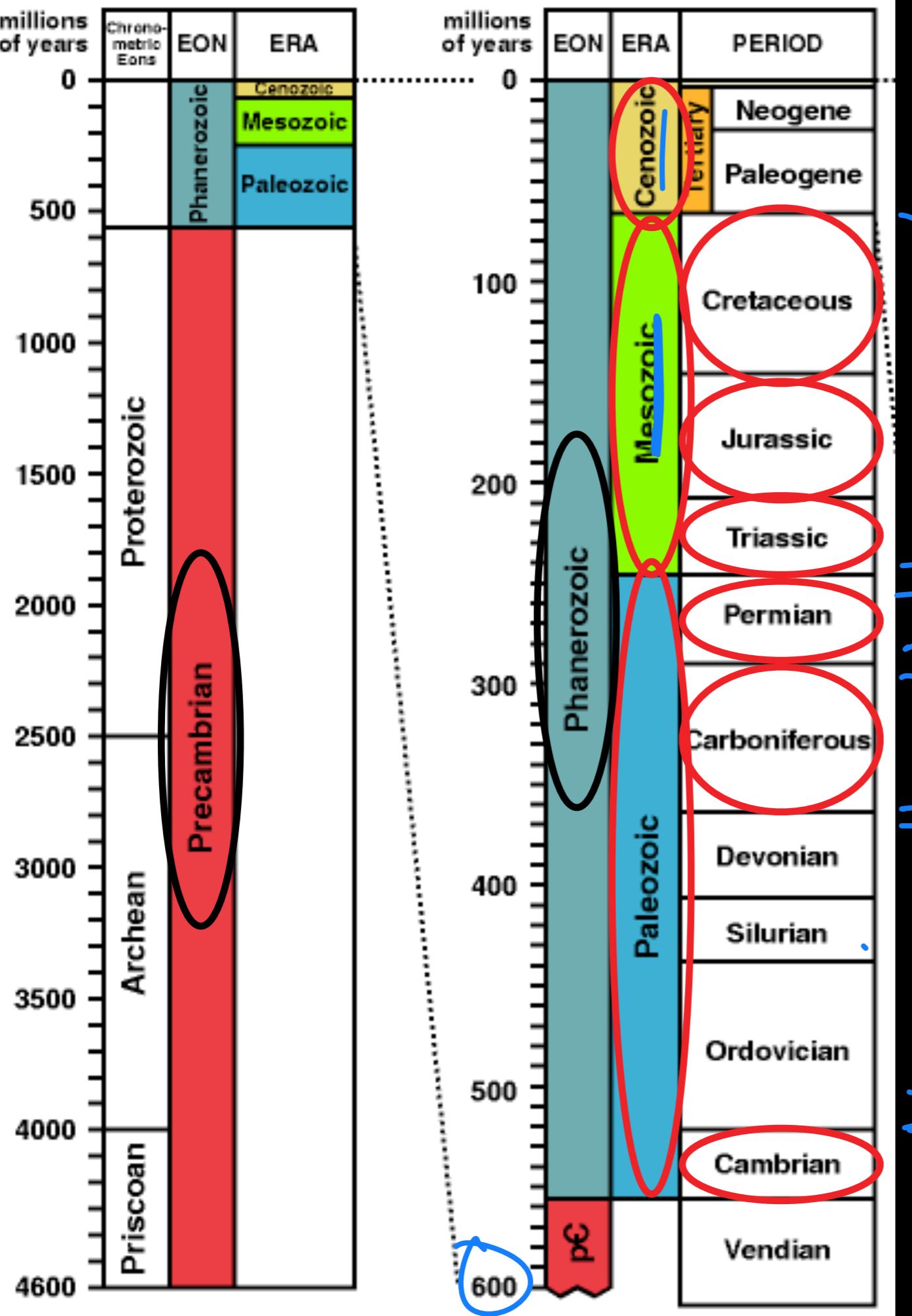
© Mark A. Garlick
space-art.co.uk

$> 4.527 \text{ gyr}$

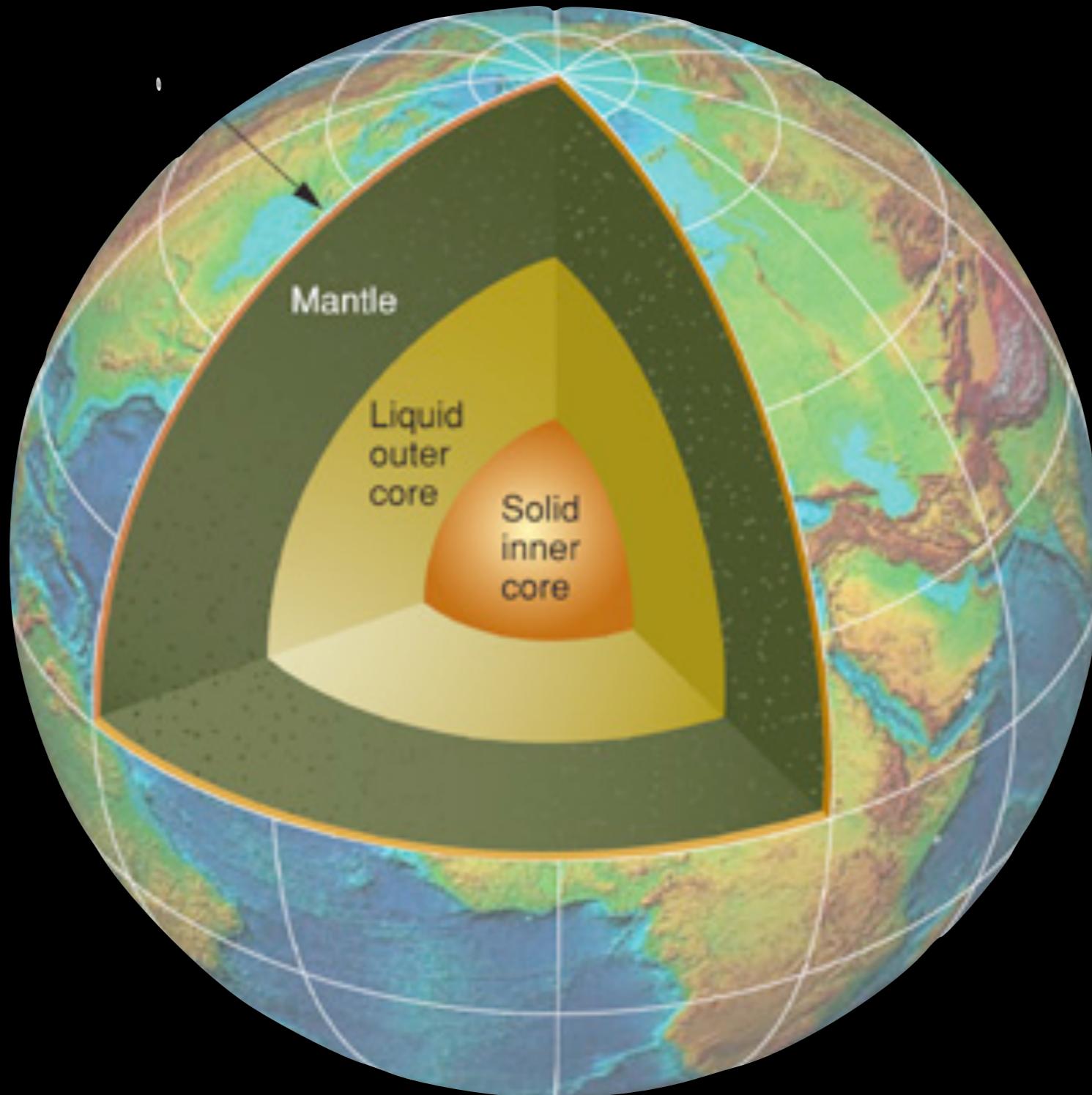
Geological History



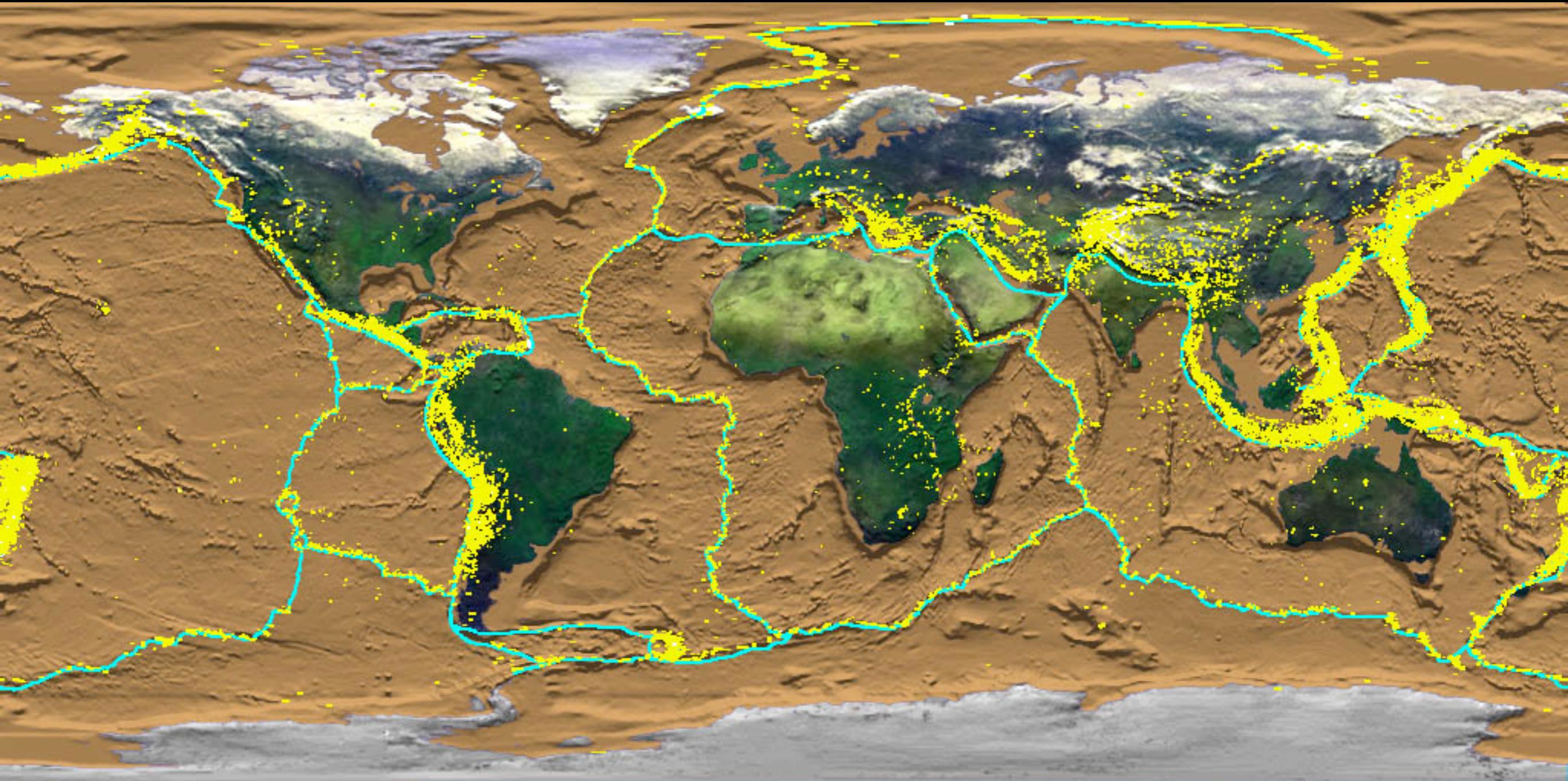
Be Familiar Know



- Precambrian Eon
- Phanerozoic Eon
- Paleozoic
 - Cambrian (542:488 Ma)
 - Carboniferous (360:299 Ma)
 - Permian (299:251 Ma)
- Mesozoic
 - Triassic (251:200 Ma)
 - Jurassic (200:146 Ma)
 - Cretaceous (146:65.5)
- Cenozoic (65.5:now)

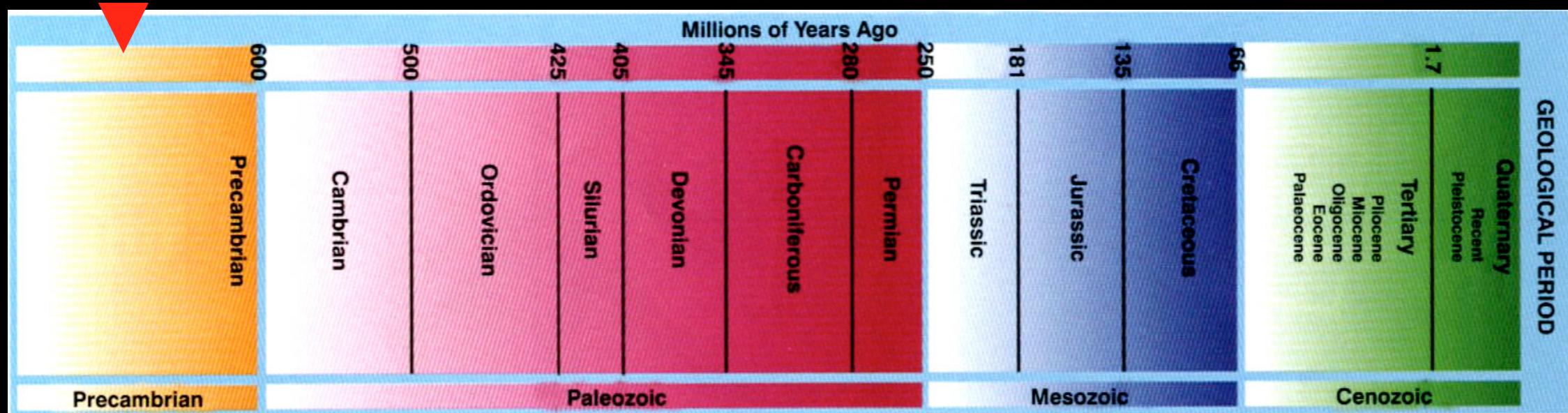
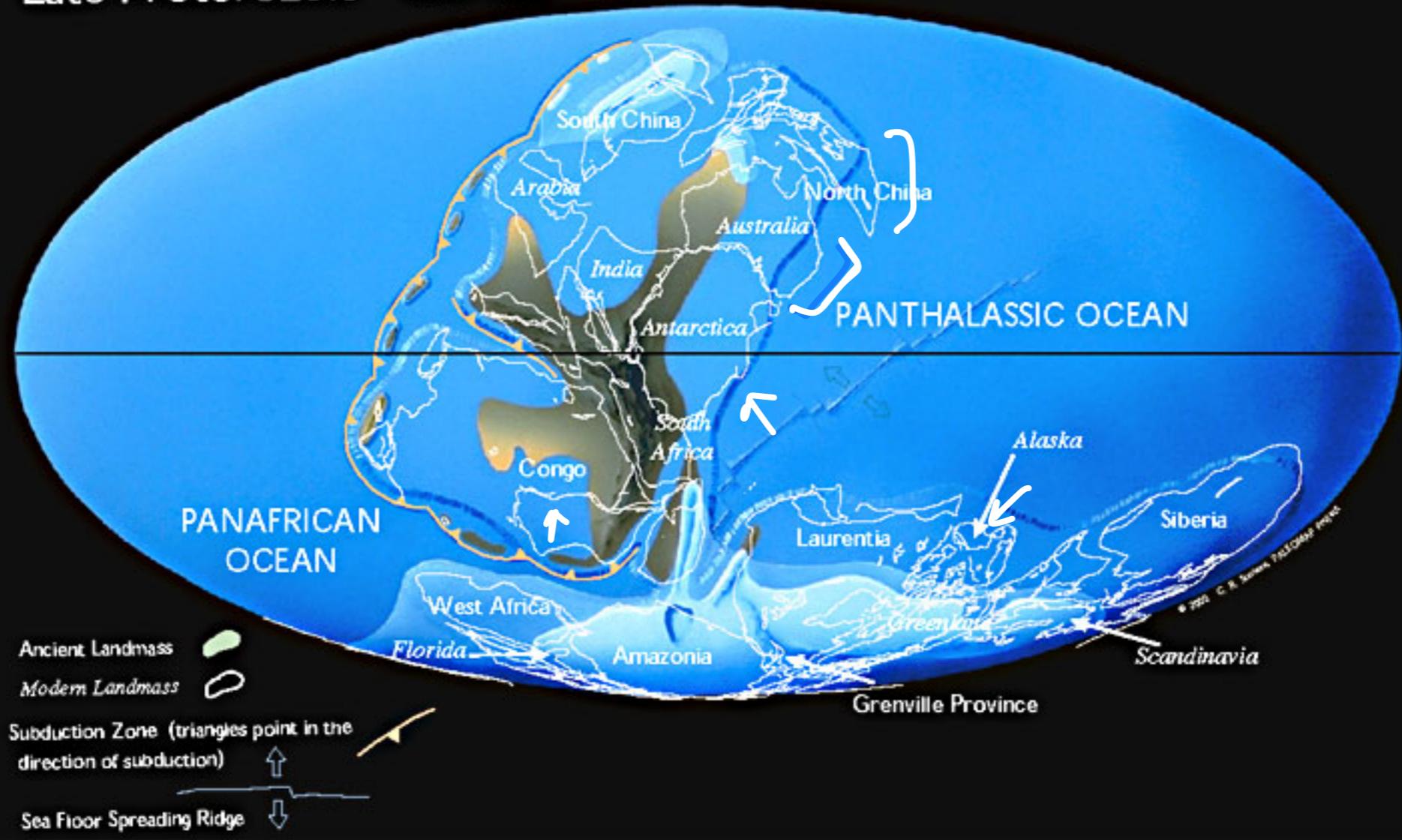


Continental Plates

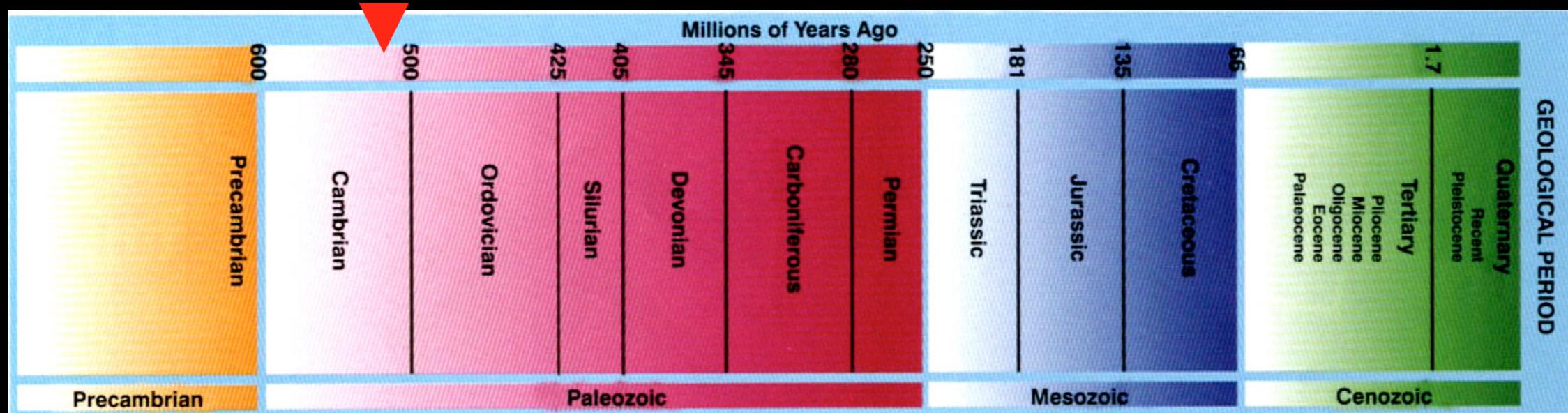
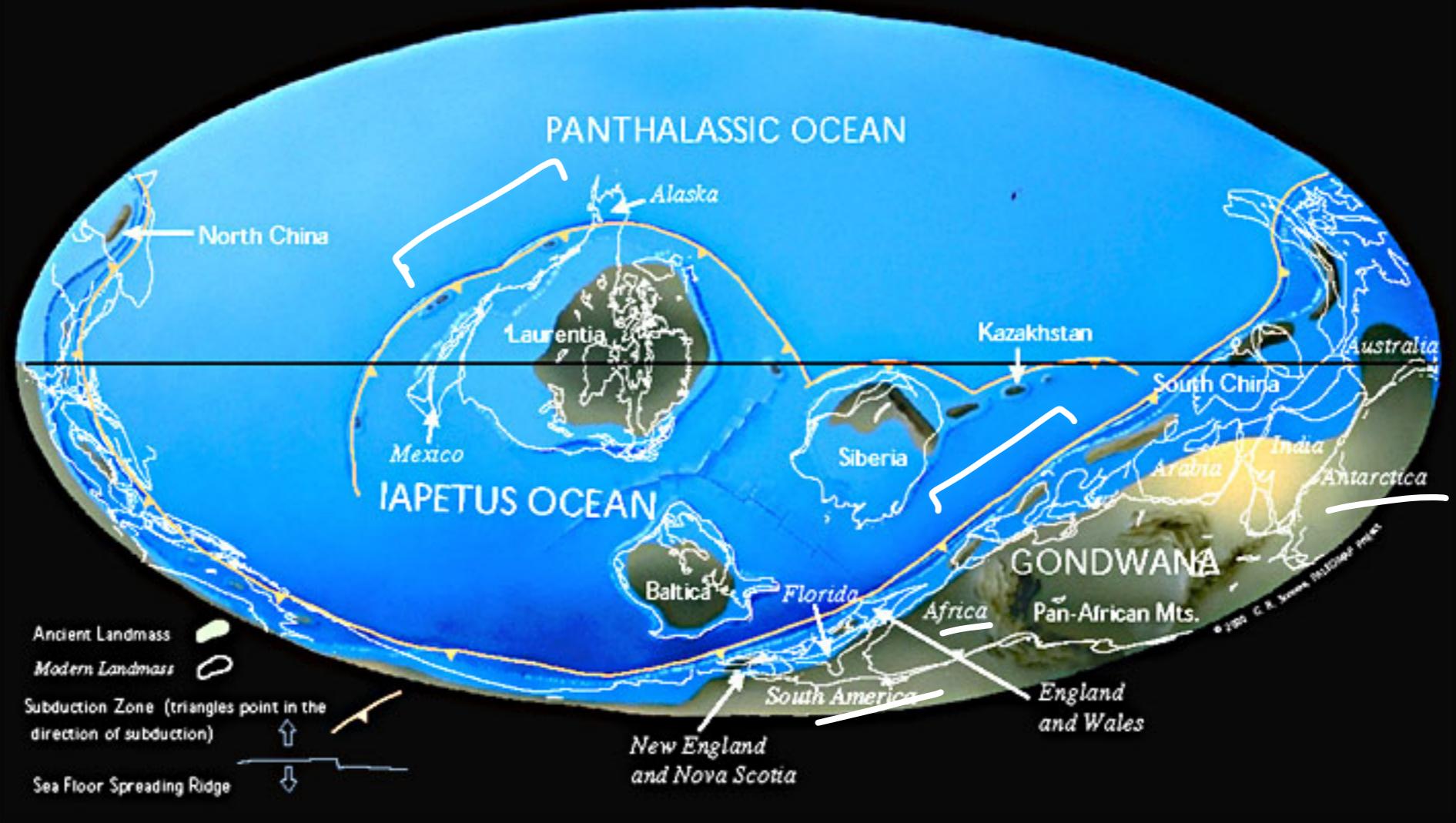


- = earthquakes

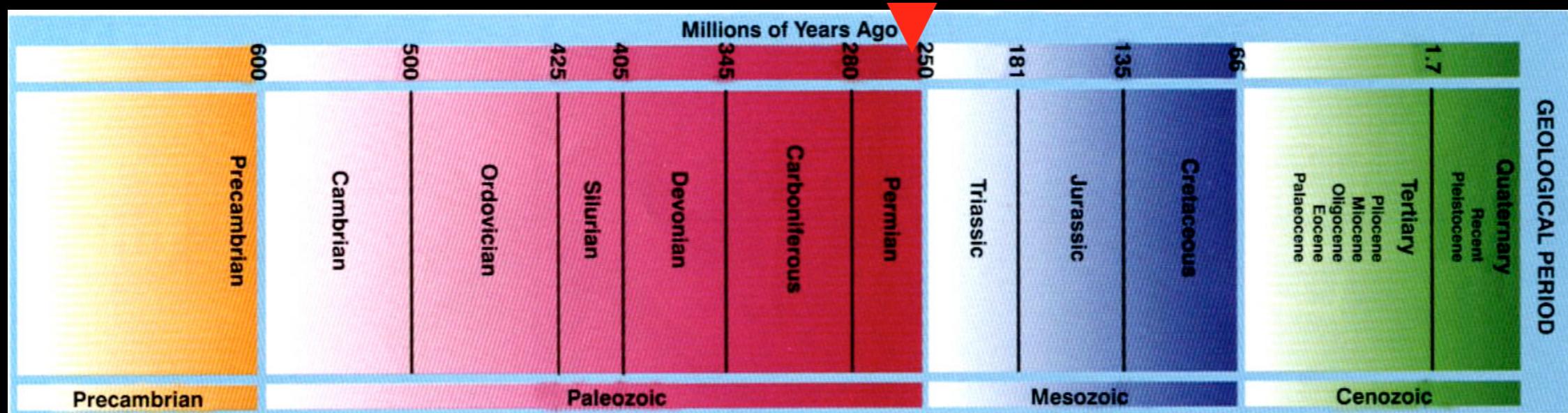
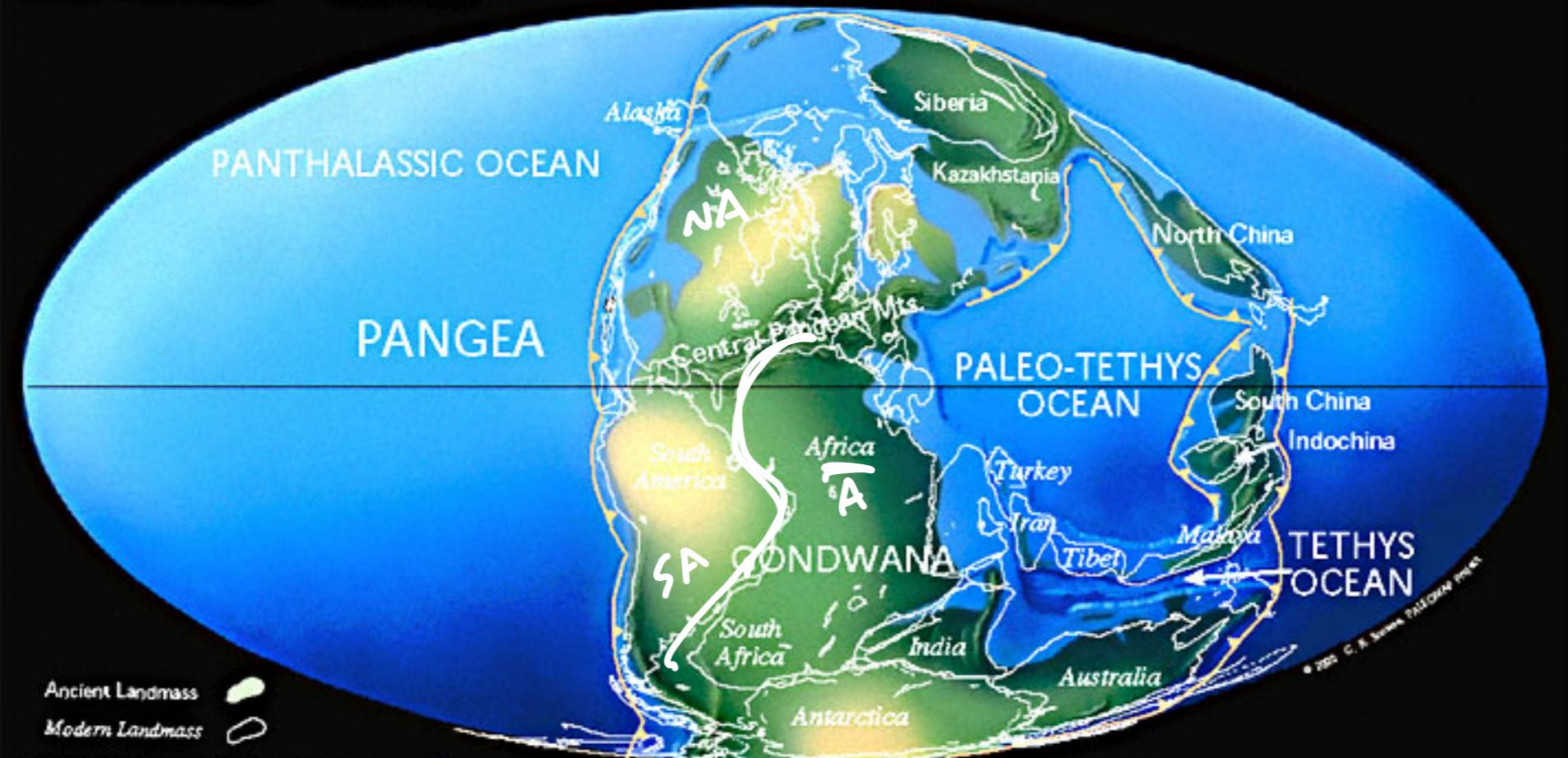
Late Proterozoic 650 Ma



Late Cambrian 514 Ma

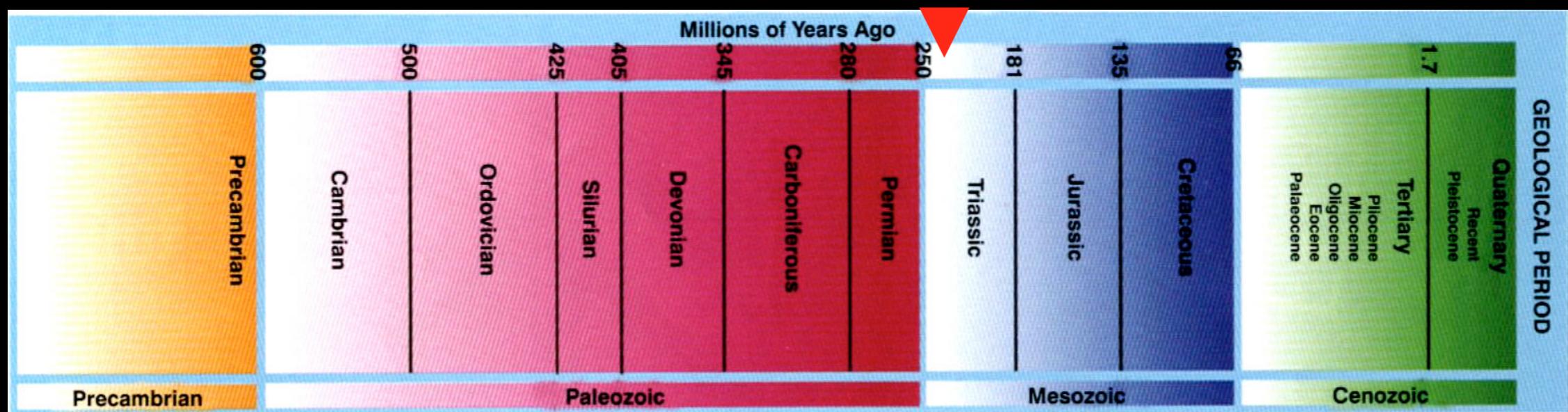
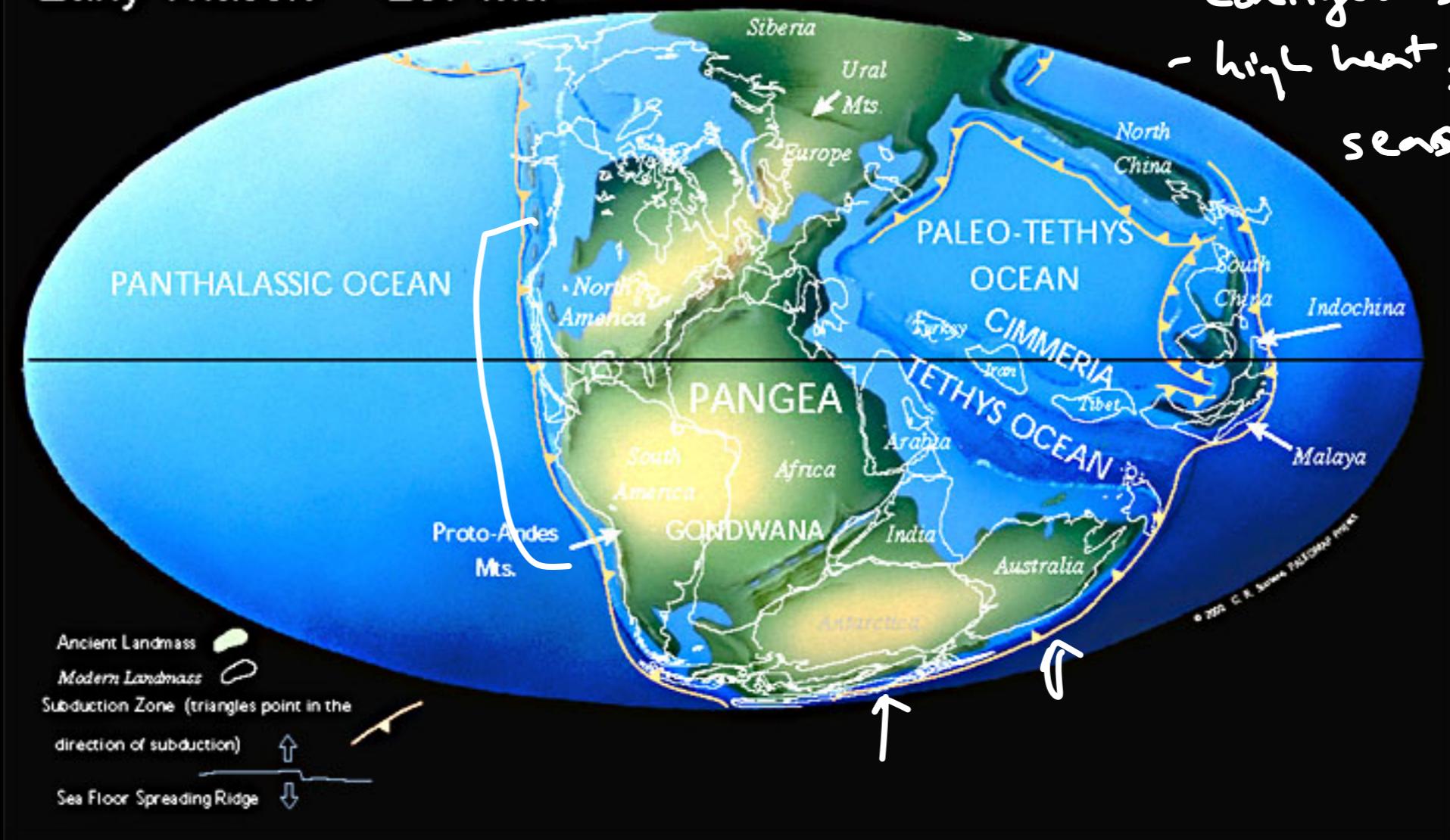


Late Permian 255 Ma

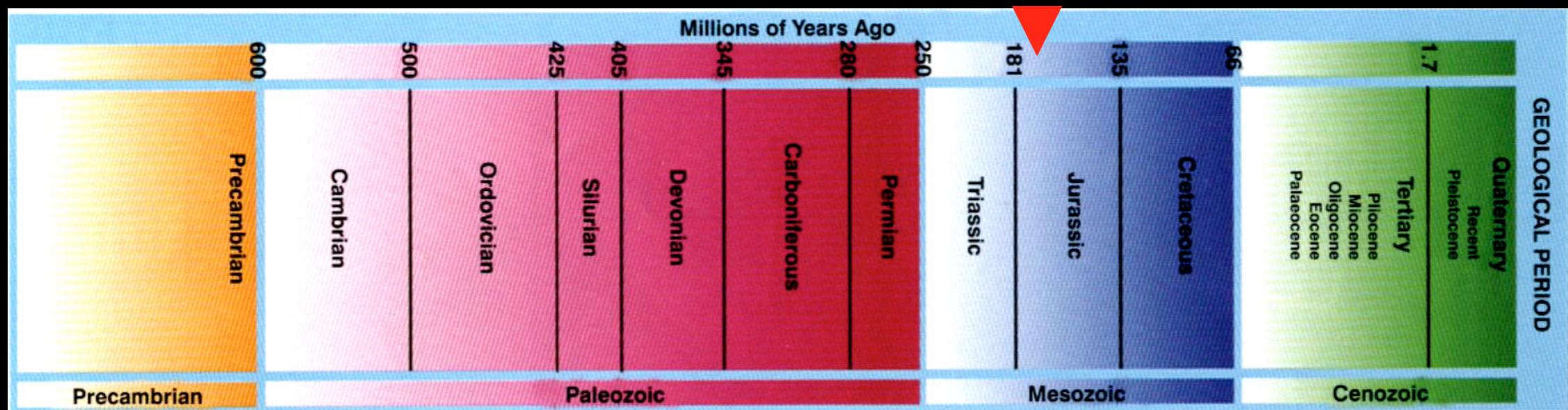
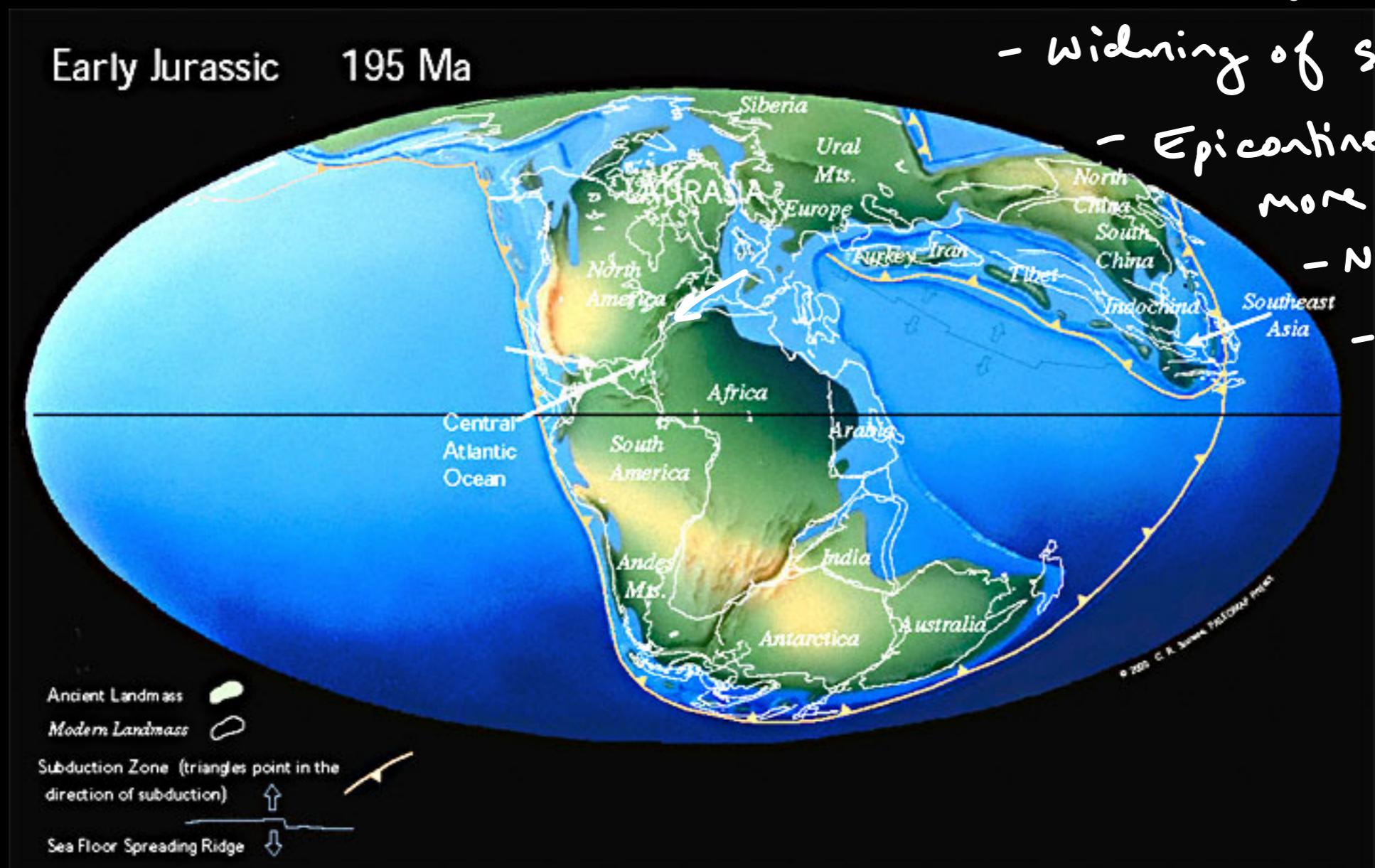


Early Triassic 237 Ma

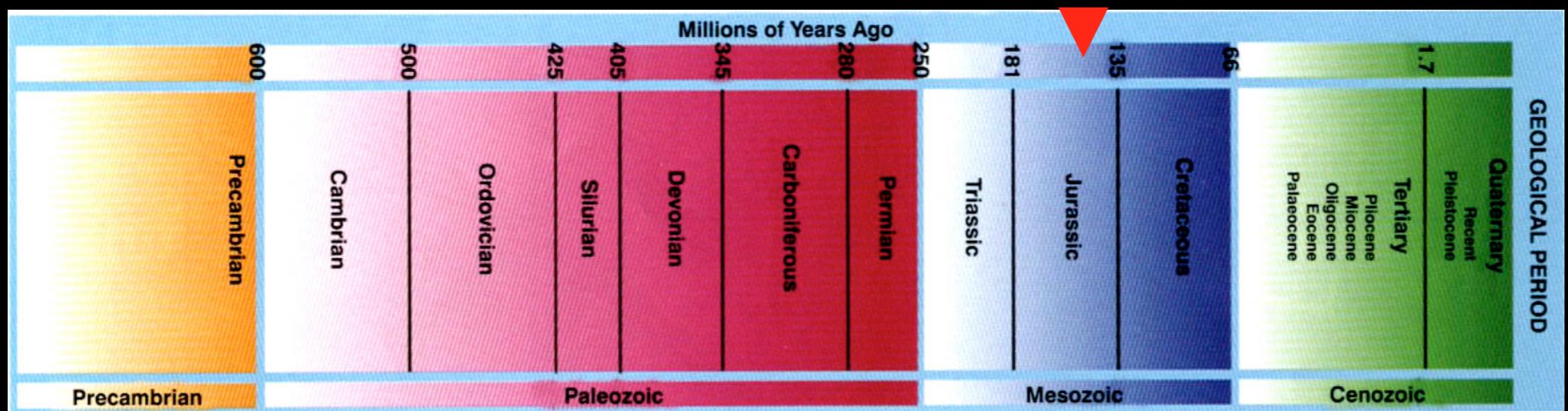
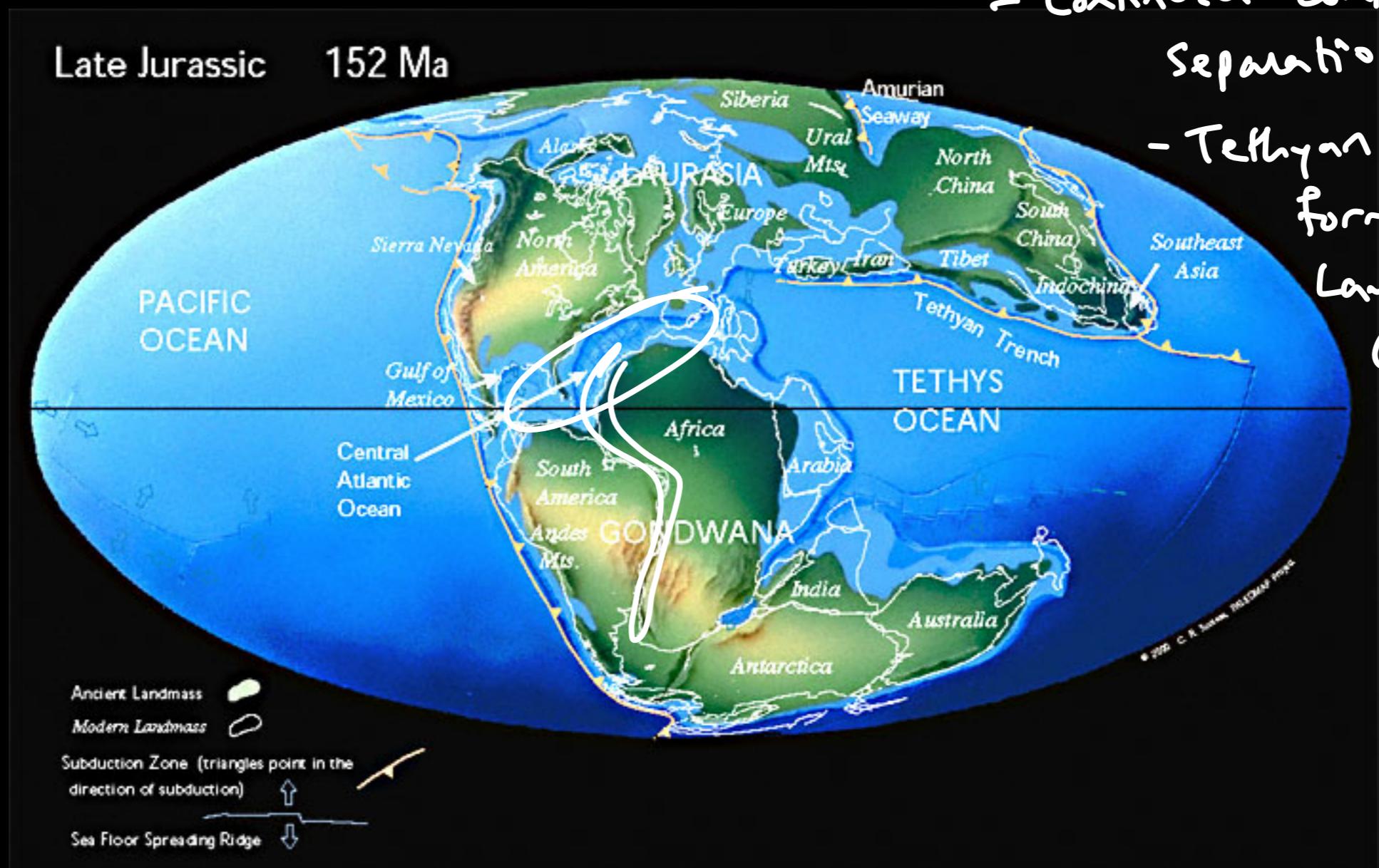
- Pangea
- Contiguous
- high heat, dry, seasonal

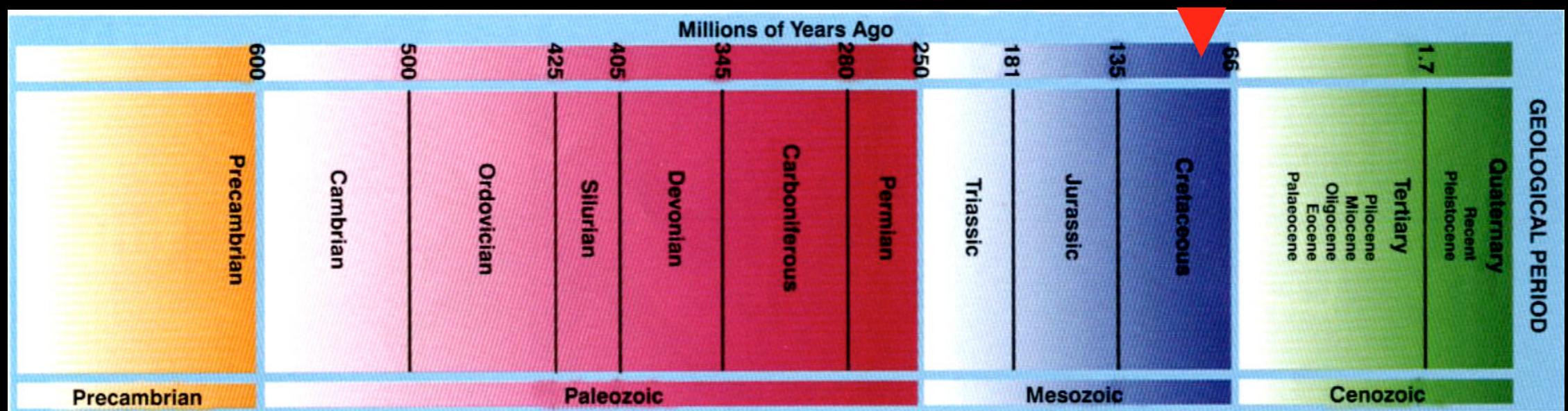
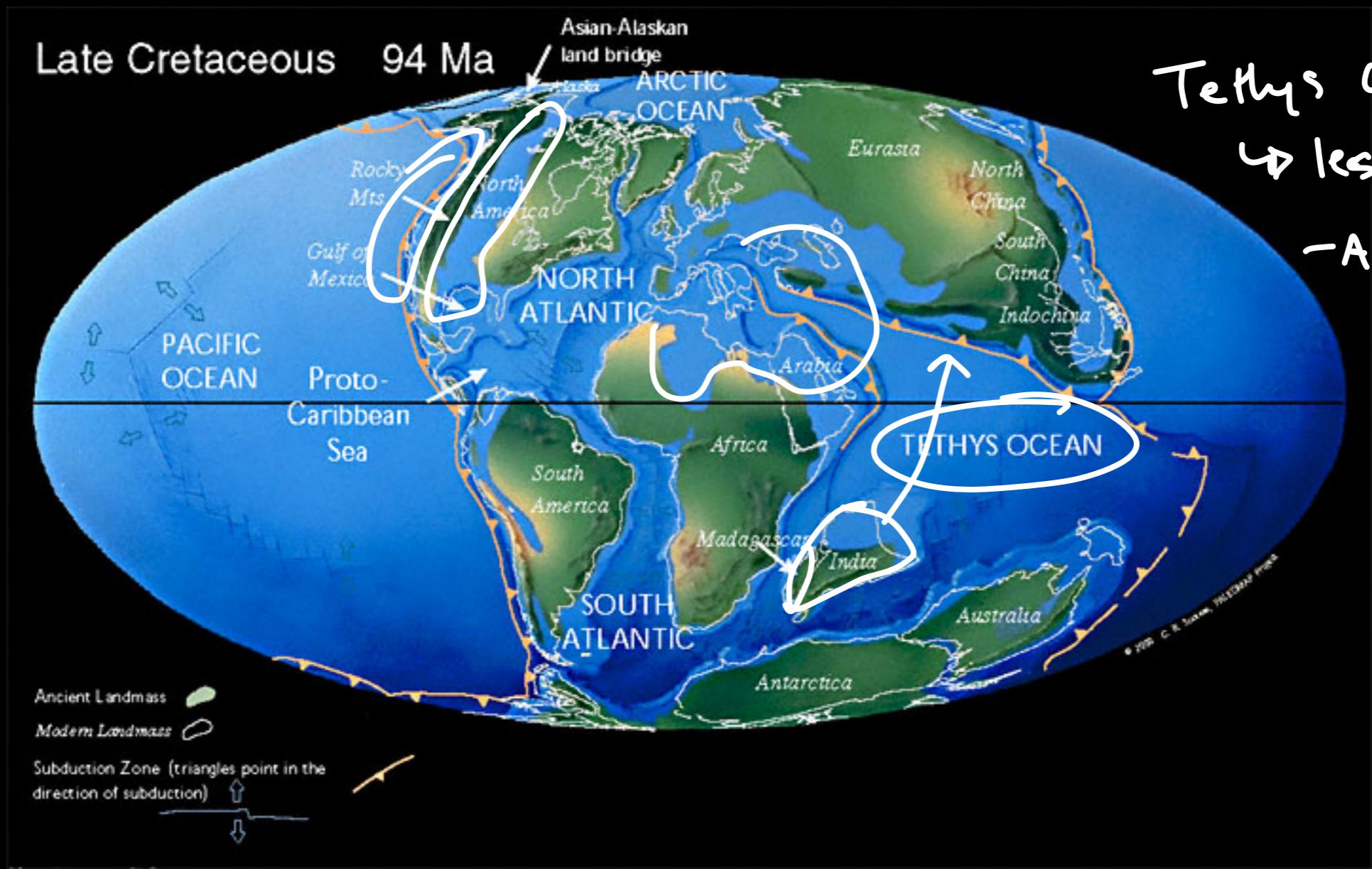


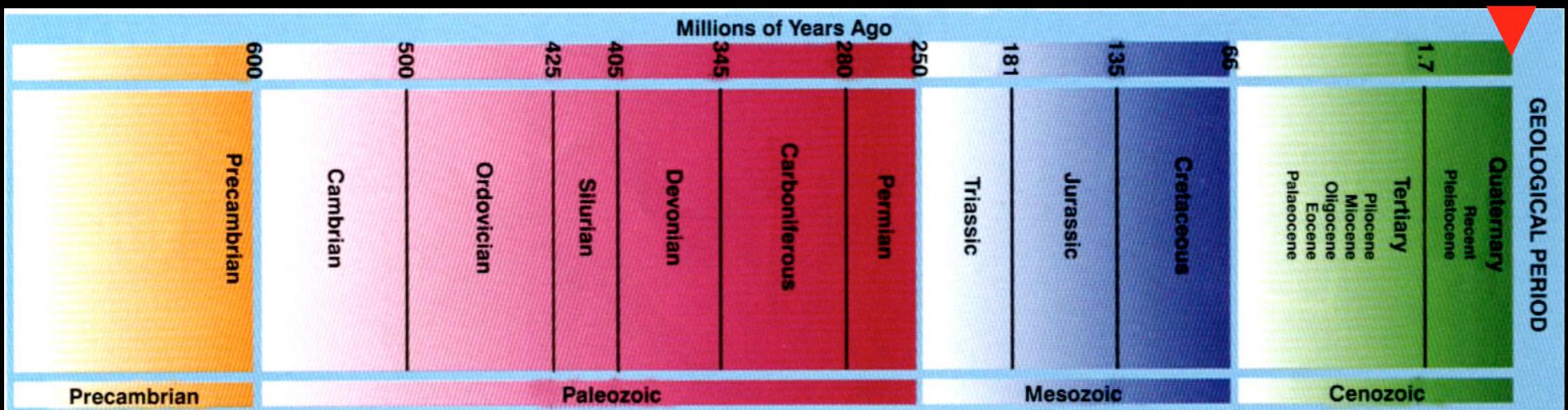
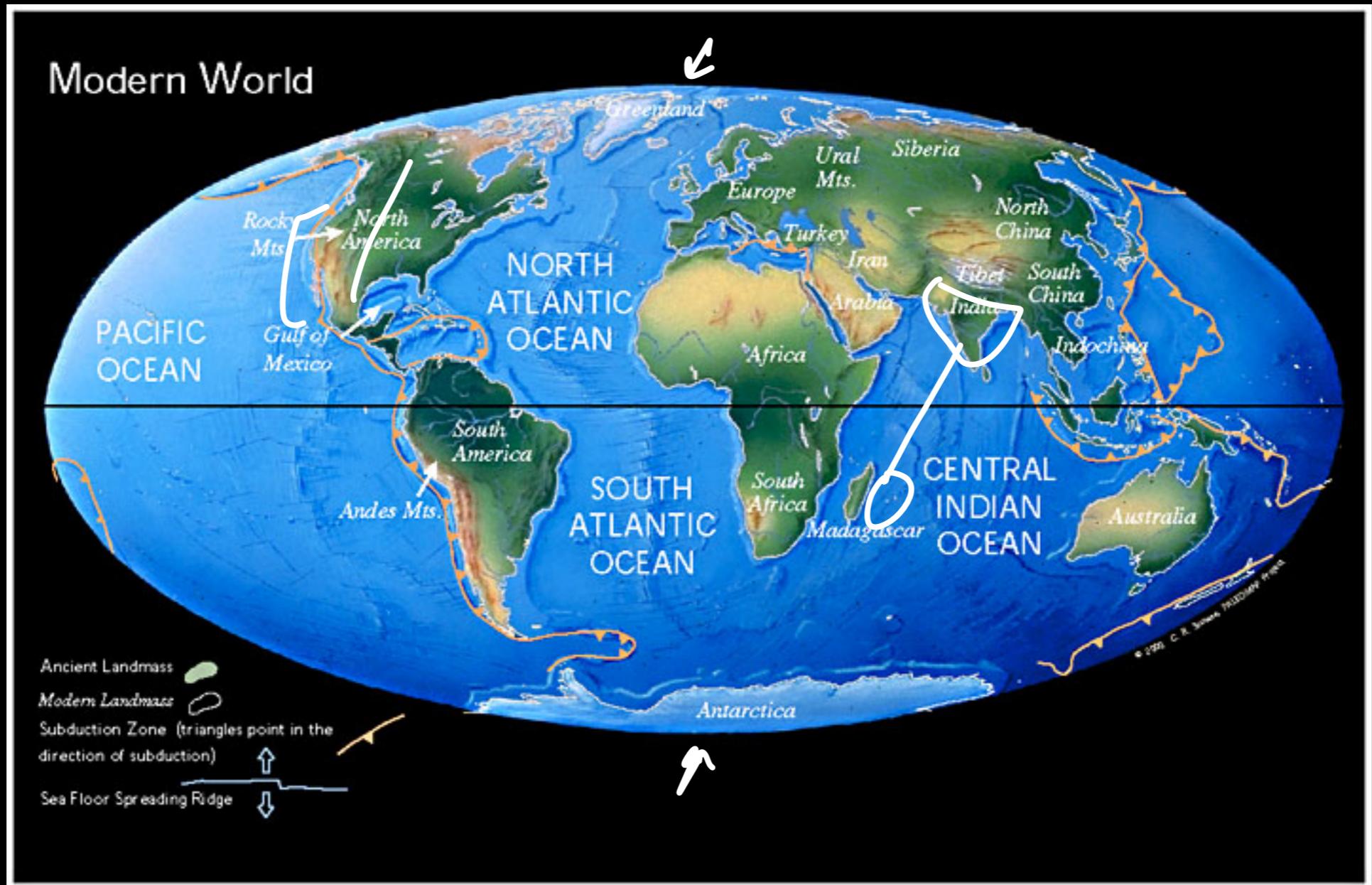
- Break-up of Pangaea
- Widening of seaways
- Epicontinental seas more common
- No ice caps
- Sea levels higher



- Continued continental separation
 - Tethyan Seaway forms between Laurasia & Gondwana





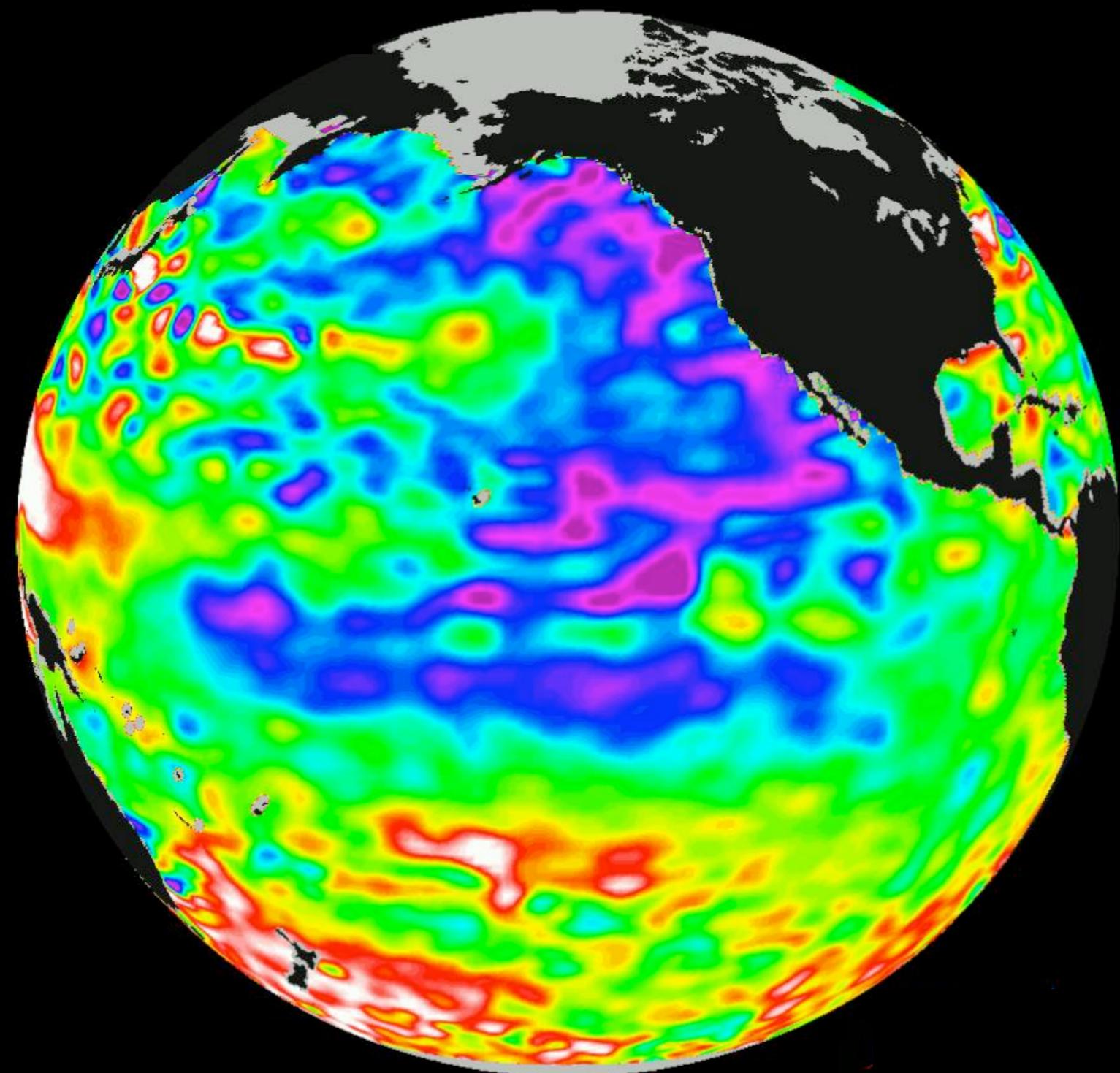


Consequences

- 1) Pangea : More rapid warming/cooling - more intense than today
 - * ~~Modern~~ Modern continents tempered by oceans between them. Pangea didn't have this
- 2) Heat is distributed more equally through fluids than solids ~ oceans are slower to warm/cool than continents

Effects of large seas increase as Pangea breaks up.
[lowers the magnitude/rapidity of temperature fluctuations]

Why do we care?





“Being a paleontologist is like being a coroner except that all the witnesses are dead and all the evidence has been left out in the rain for 65 million years.”

Mike Brett-Surman, 1994



Herrerasaurus



There is a lot that we don't know about dinosaurs...

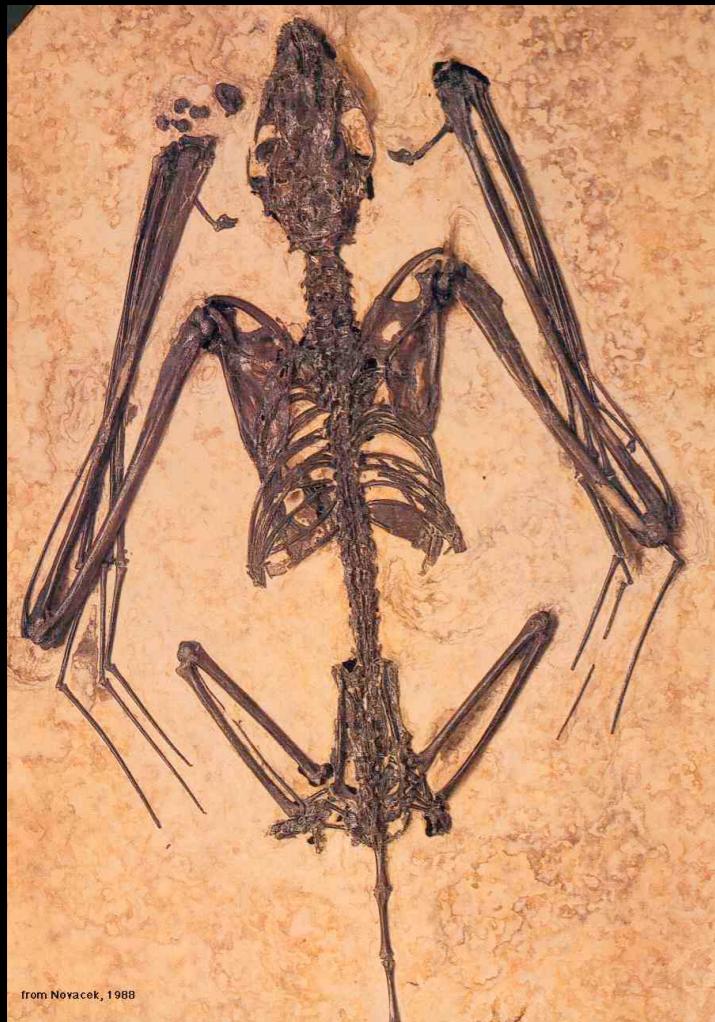
Science is dynamic



Fossils and Preservation:

Fossils

From greek *Fossilis* = 'Dug Up'



Many meanings... types of fossils

Trace Fossils

- ↪ Poo (Coprolites)
- ↪ Gastroliths
- ↪ Trackways



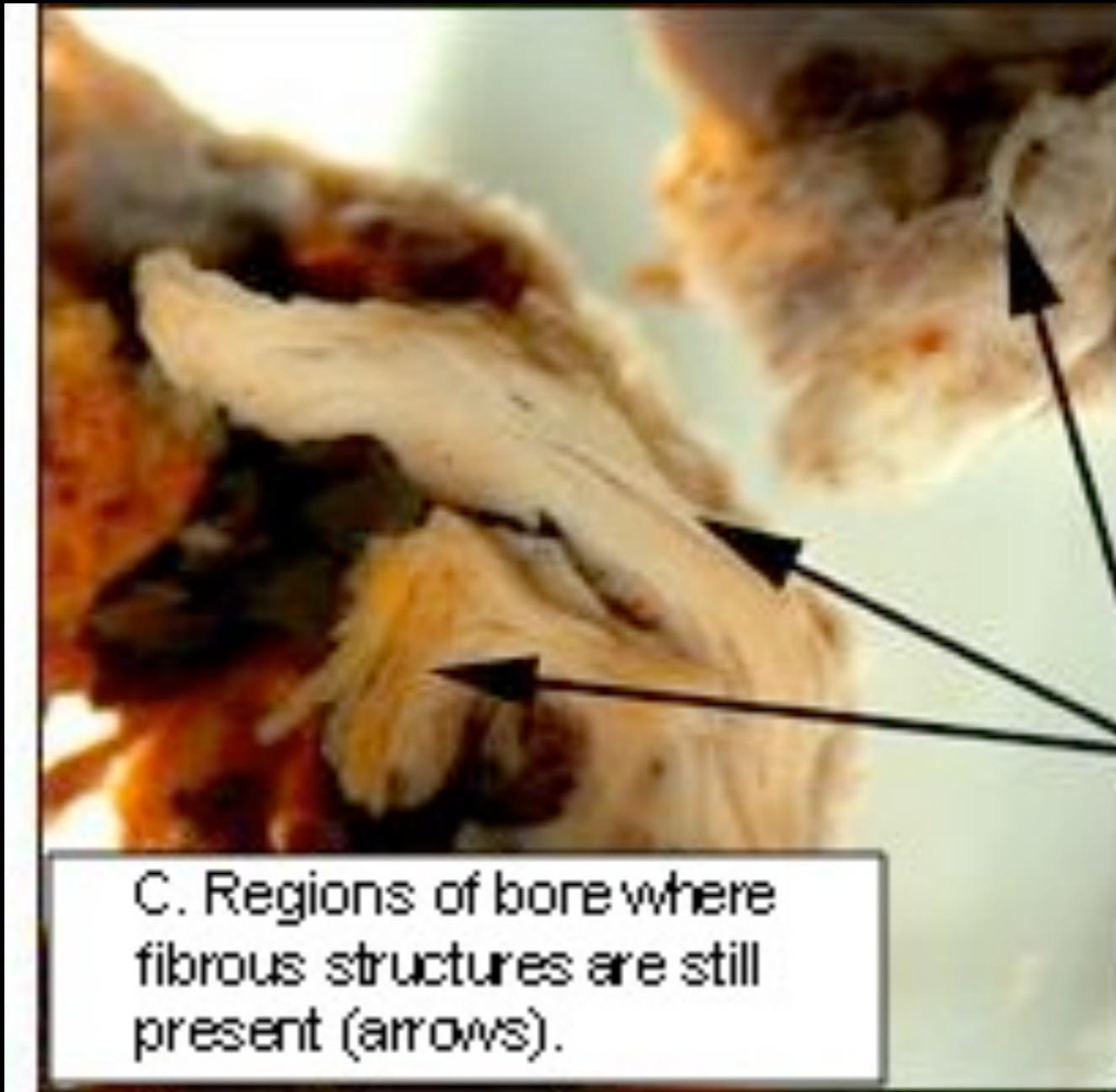
Soft Parts:
Impressions
Amber
Protein???



Hadrosaur skin



T. rex collagen?



68 Ma

Hard Parts: Living vs. Fossilized

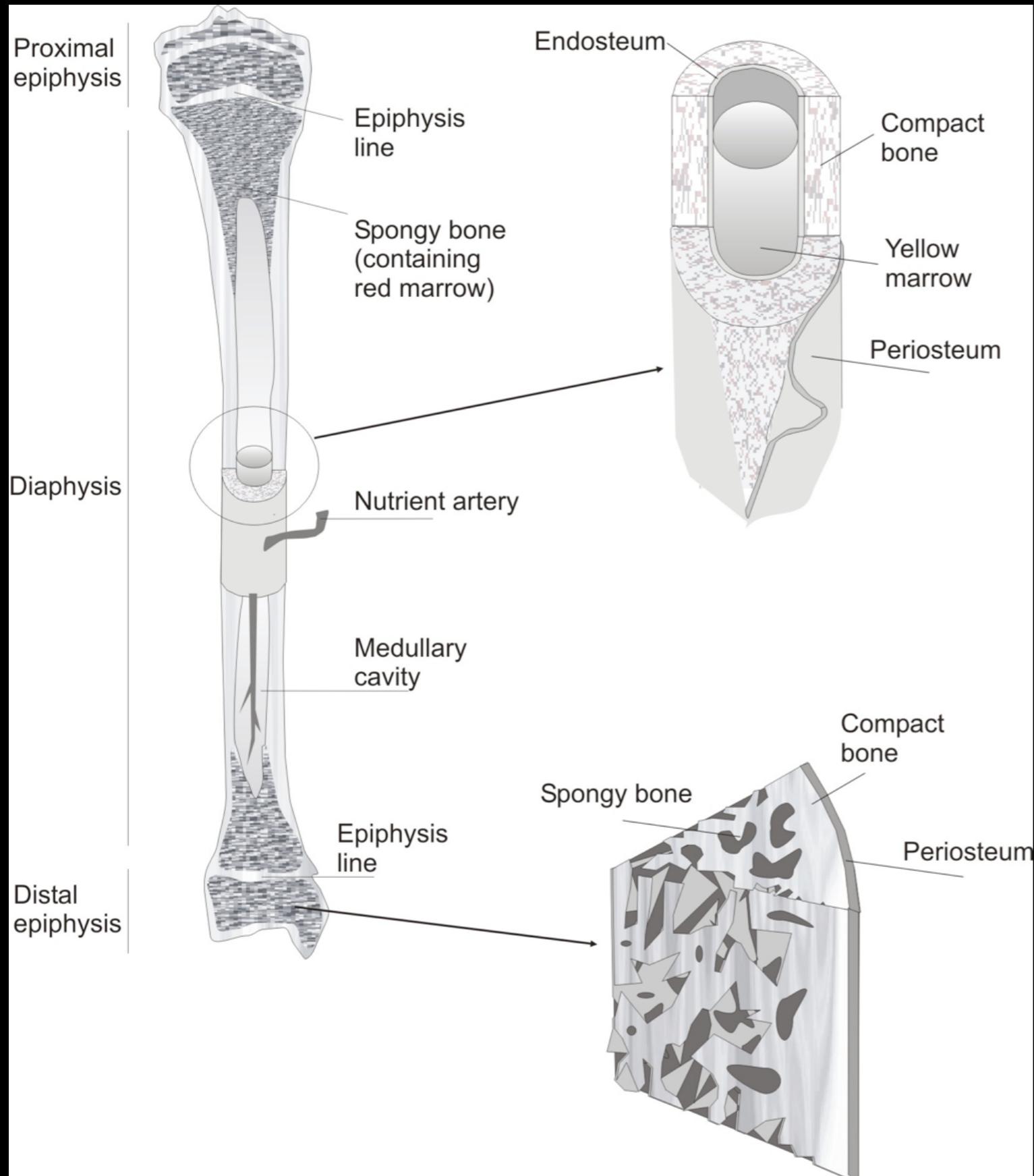


Tooth



Egg

Bone



Living Bone

Bone Matrix

- Organic
 - **Collagen**

protein

- Nonorganic
 - Hydroxyapatite

mineral

Structure of Bone

Diagram illustrating the structure of skeletal long bones comprising solid outer cortical (compact) bone and inner trabecular (spongy) bone in which the bone marrow is housed. Redrawn and adapted from Baron, 1996. Copyright BTR©

Fossil Bone

$\text{Ca}_{10}(\text{PO}_4)_6(\text{F}, \text{OH}, \text{Cl})_2$



Apatite Mineral Francolite
No longer biological- it's a rock.

Unaltered remains

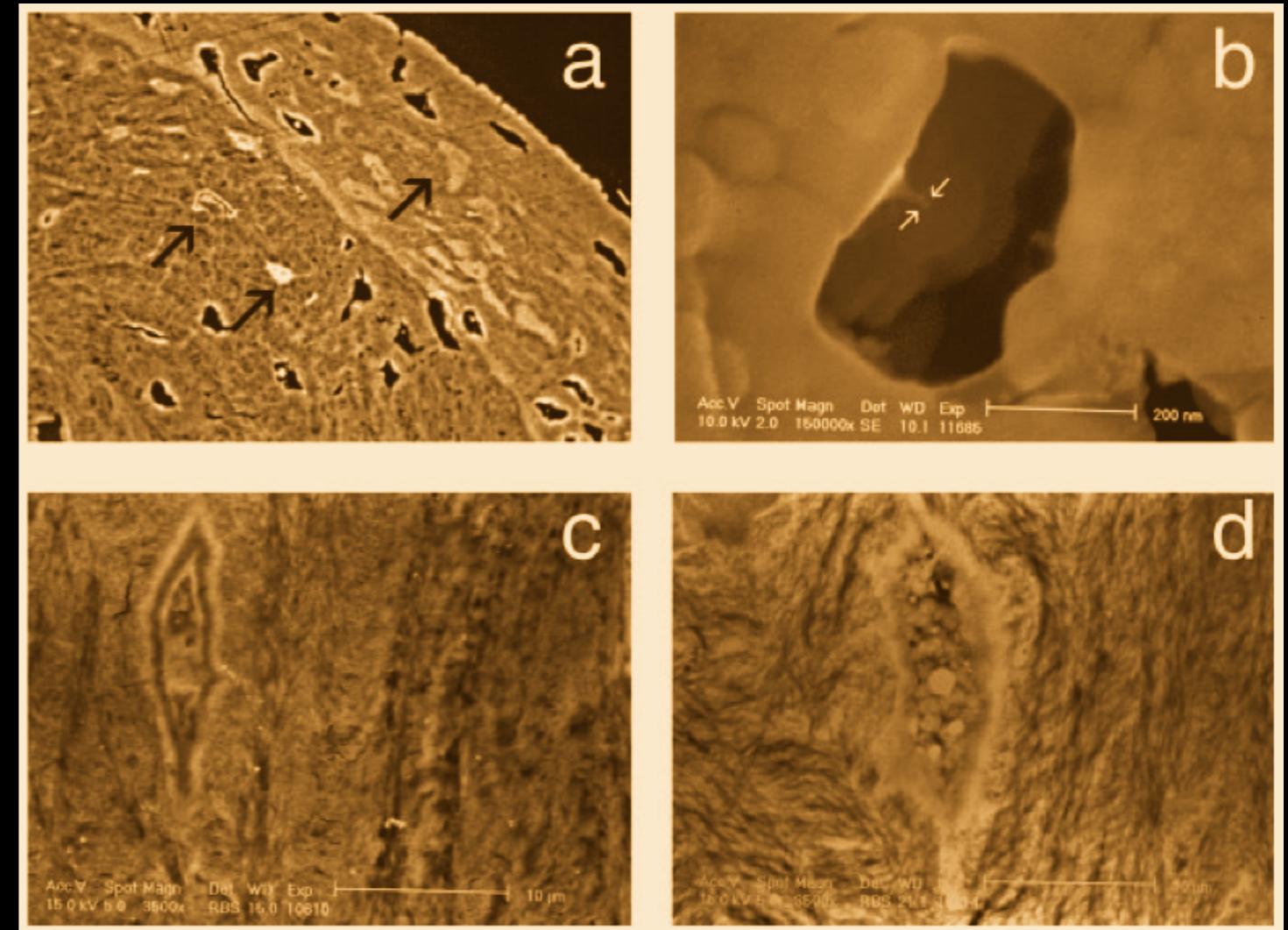


Bogs



- Acidic peat bogs, tar pits
- Pickling

Fossilization: A fine scale process



Even osteocytes (bone cells) are preserved

Routes to Fossilization



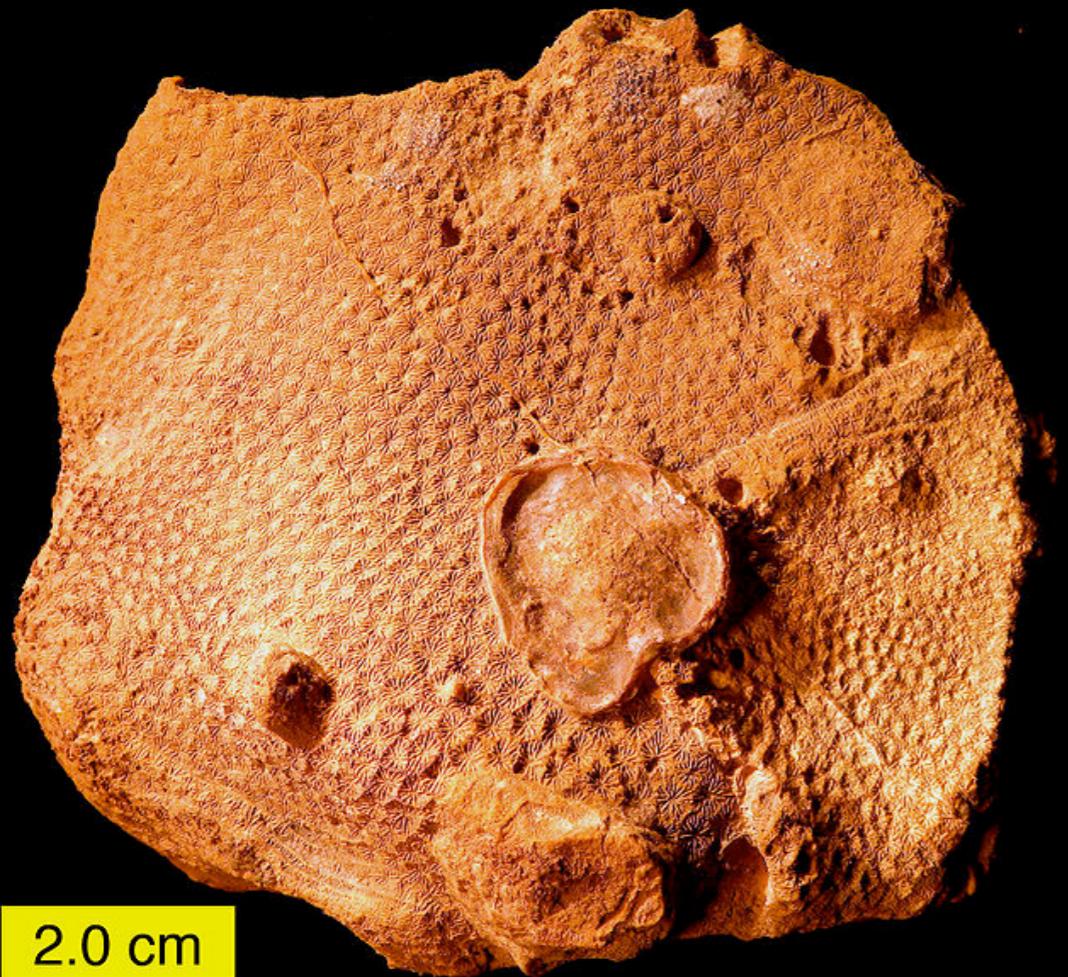
Permineralization

Trilobite

- **Open spaces in organic material is filled up with minerals**
 - Crystals form within cell walls
 - This type of preservation conserves cell structure



Routes to Fossilization



2.0 cm

Fossil Coral: Jurassic
Mineralization and Recrystallization

- Shell, bone, tissue **replaced** with another mineral into a crystal

Routes to Fossilization



Structure is typically compressed

Pressure, heat force out gasses, liquids

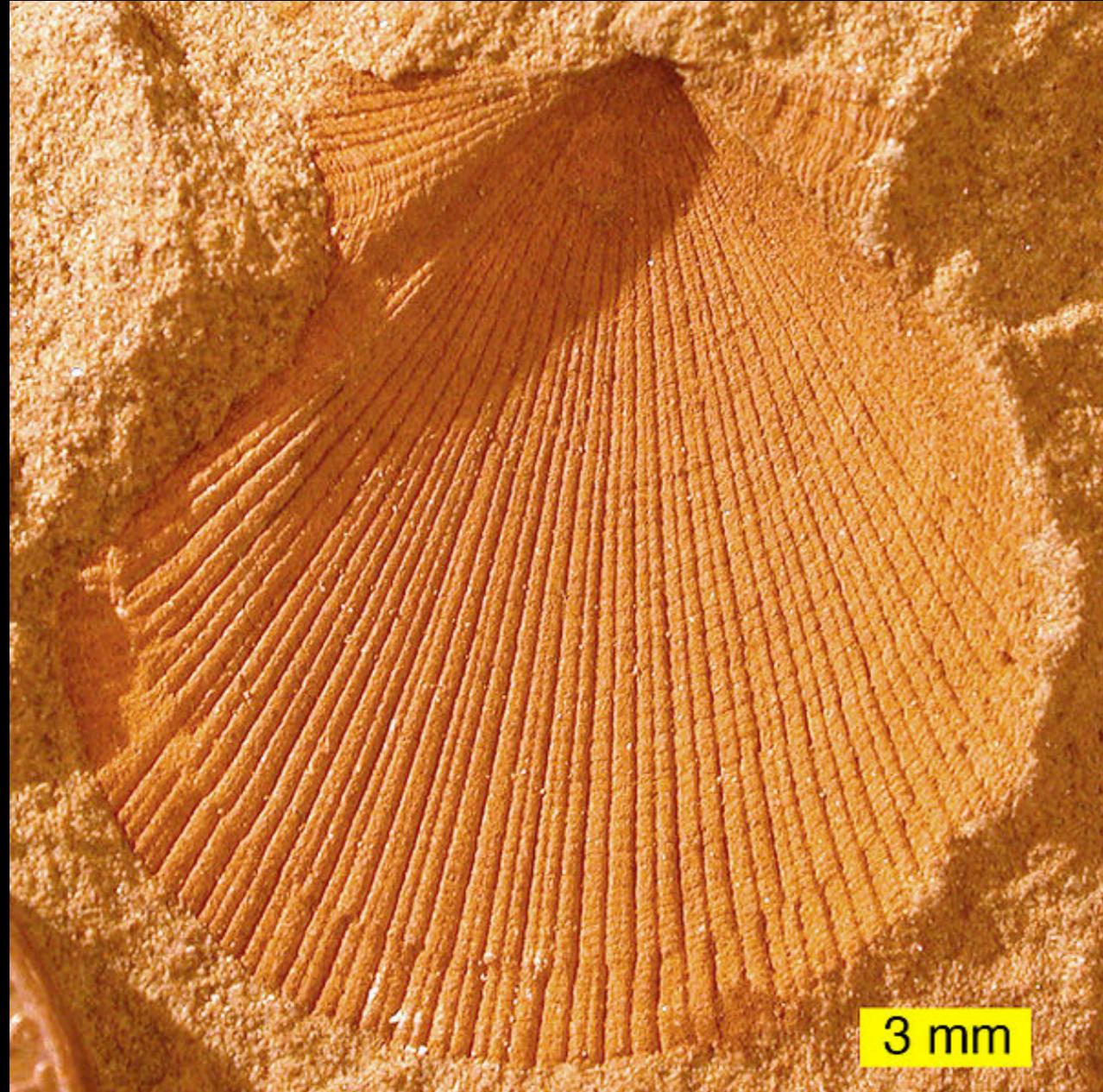
Leaves behind a carbon film

!! Soft parts !!

Carbonization



Routes to Fossilization

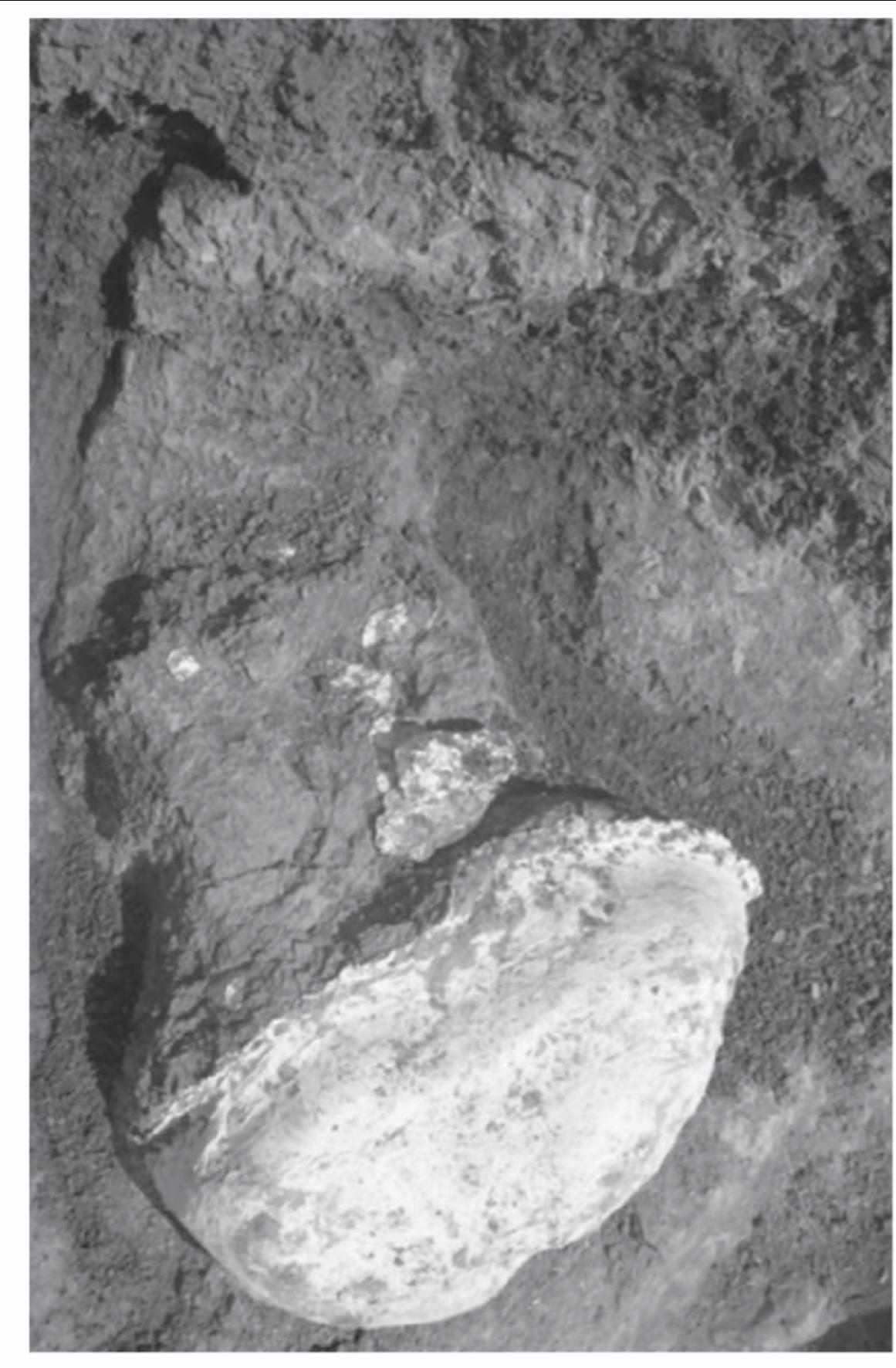


Molds, casts

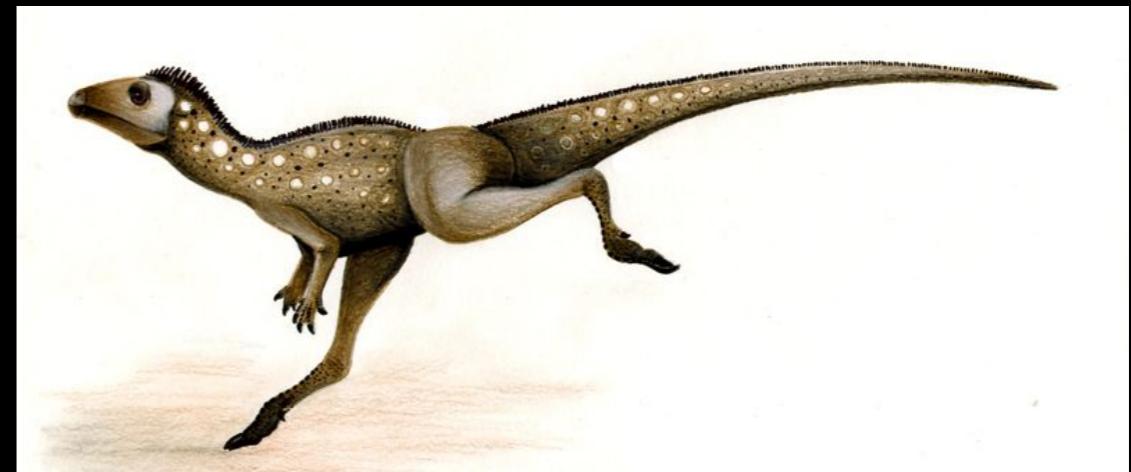
- Molds
- Casts
- Little or no original material

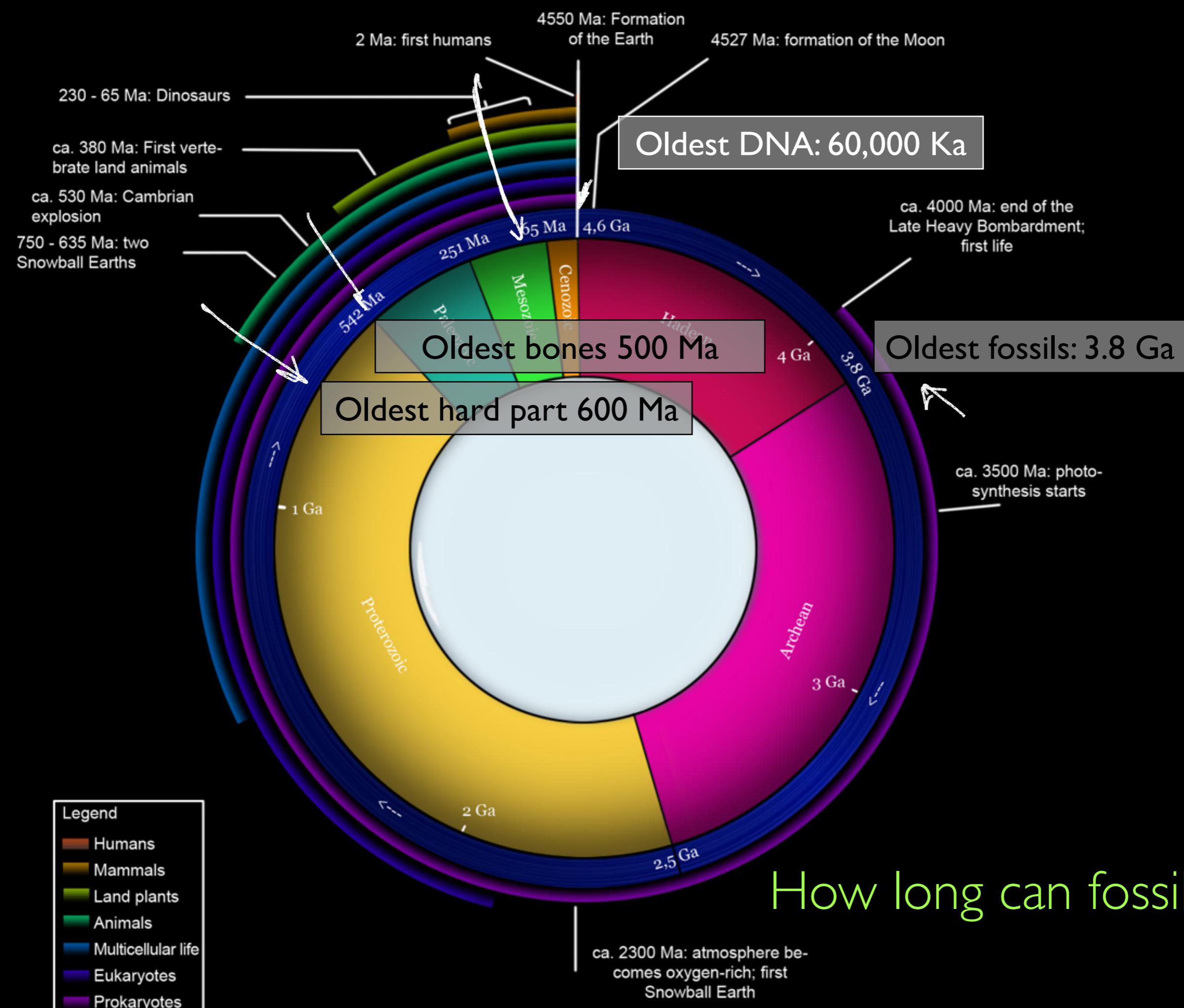


T. rex brain cast



Oryctodromeus





Taphonomy: study of the transition from the biosphere to the lithosphere



Quick burial

(a)

Replacement and/or permineralization

Nearly complete specimen exposed



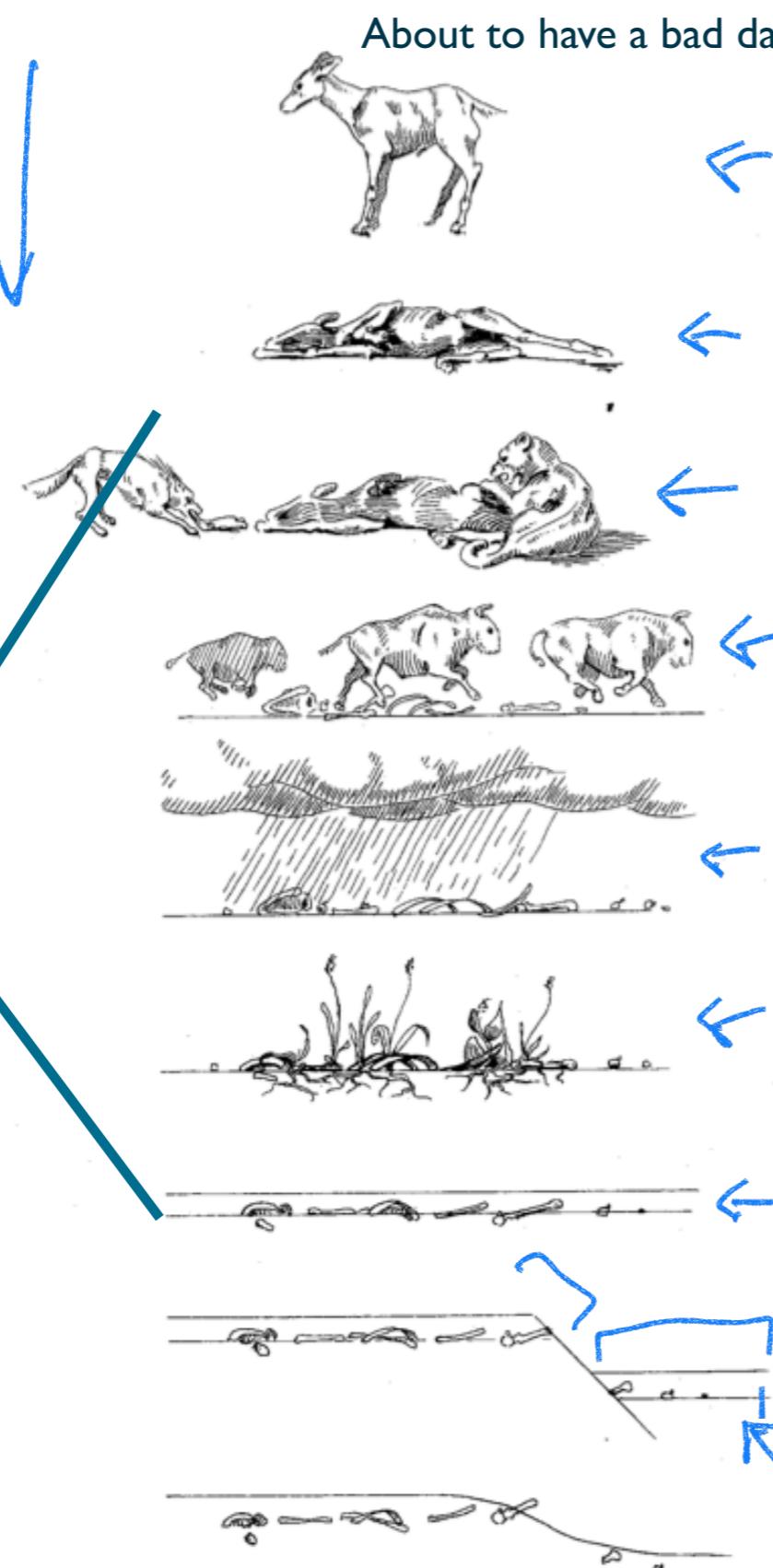
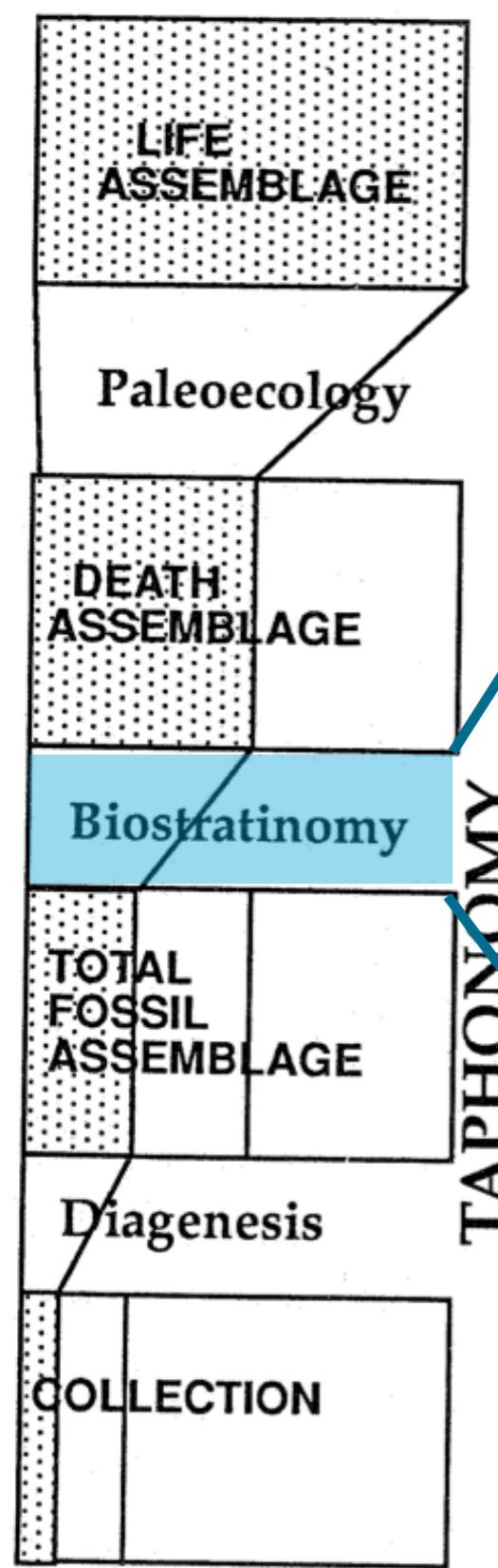
(b)



Dismemberment before burial –
scavenging and other natural
processes

Isolated bones buried and mineralized

Isolated bones exposed



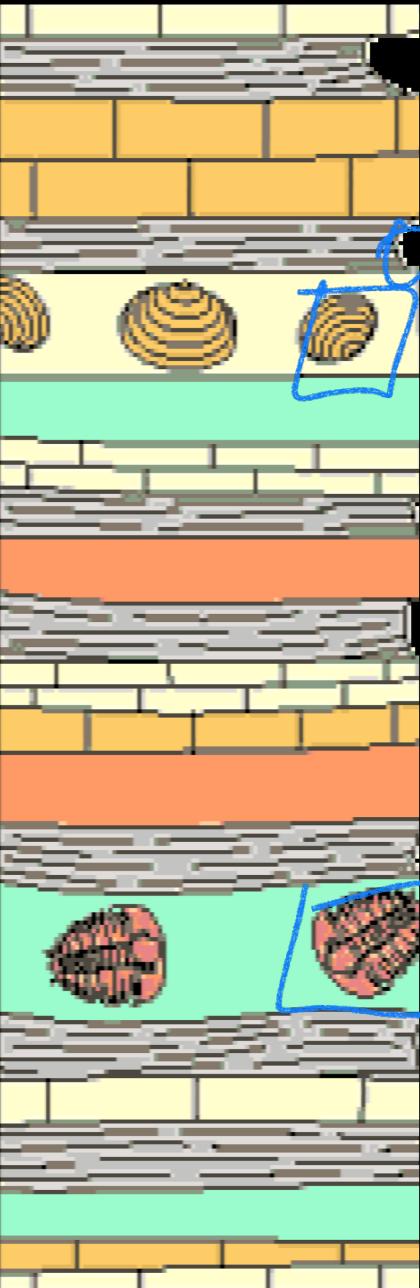
Key: Rapid Burial!

Placing fossils in TIME



Relative vs. Absolute Dating

YOUNGER



Principle of superposition ✘

OLDER

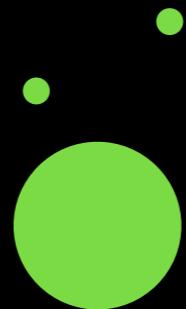
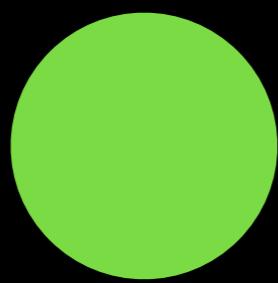
Absolute dates
Volcanic Ash

Relative ages

Issues:

- Last appearance
- Lazarus Taxa

Radiometric Dating via radioactive (UNSTABLE) isotopes



- Material must be chemically inert
- Must have radioactive isotopes
 - must be 'right type' of radioactive isotopes

If we know:

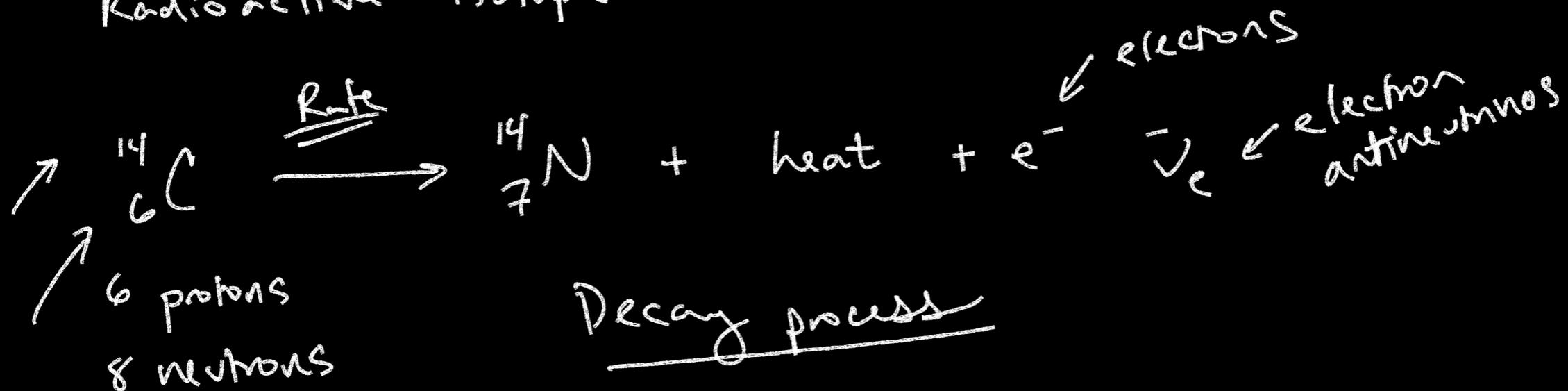
- * -Original amt of parent isotope
- * -How much of the parent isotope is left
- * -Rate of decay of that isotope

Then we can estimate:

Amount of elapsed time

Absolute dating!

Radioactive Isotope:



Isotopes ^{have} ~~at~~ the same # of ~~pro~~tons, different numbers of neutrons

Carbon: half life (50% is broken down)

5730 years

$^{40}\text{K} - ^{40}\text{Ar}$: 1280 My

$^{235}\text{U} - ^{207}\text{Pb}$: 707 My

Biostratigraphy

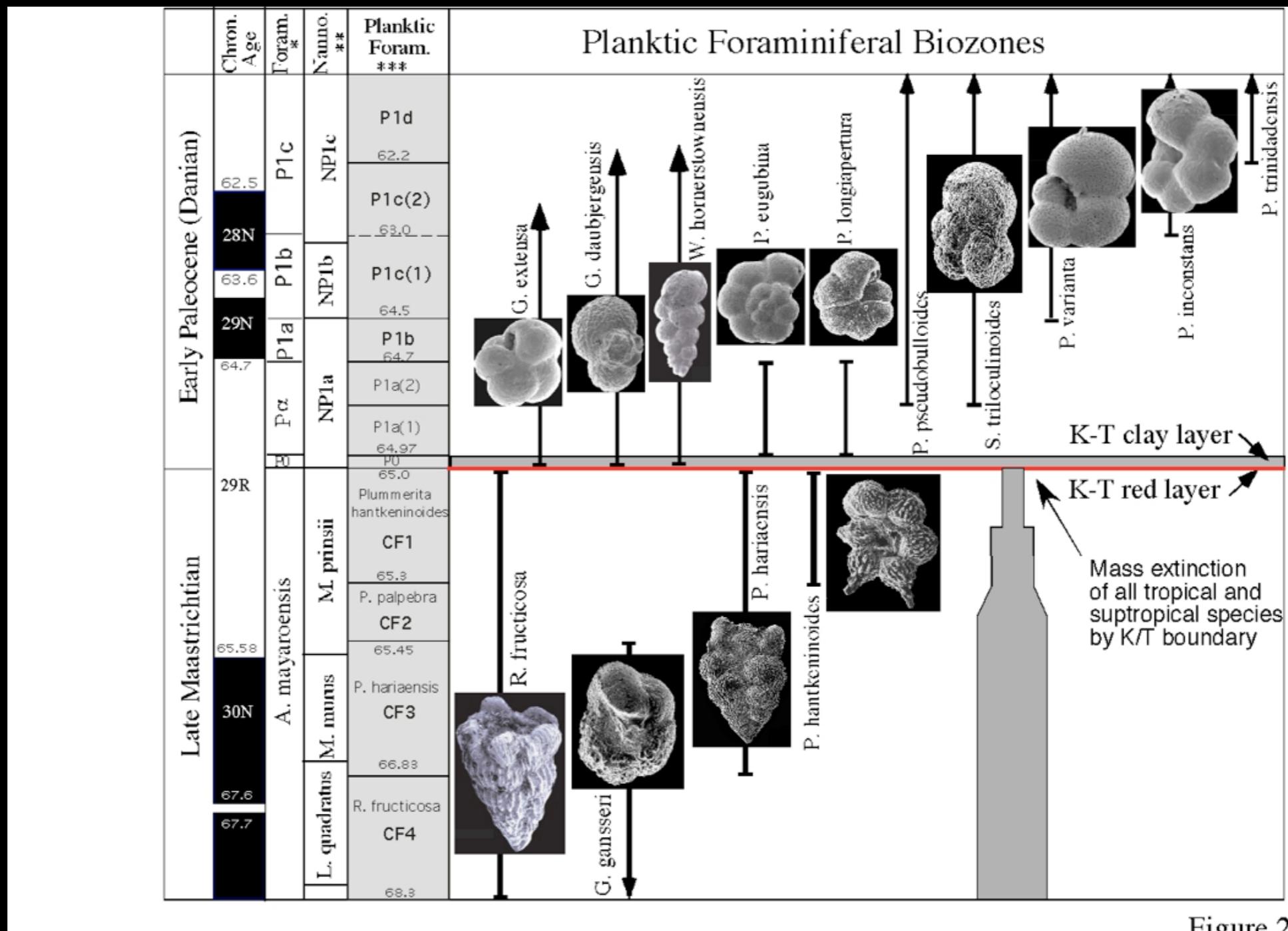


Figure 2

Relative Dating