

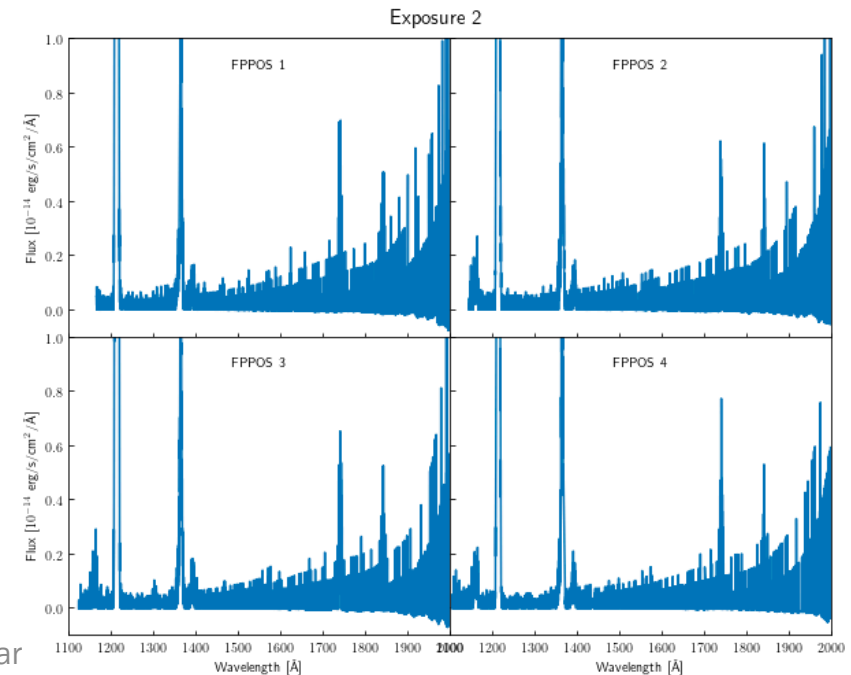
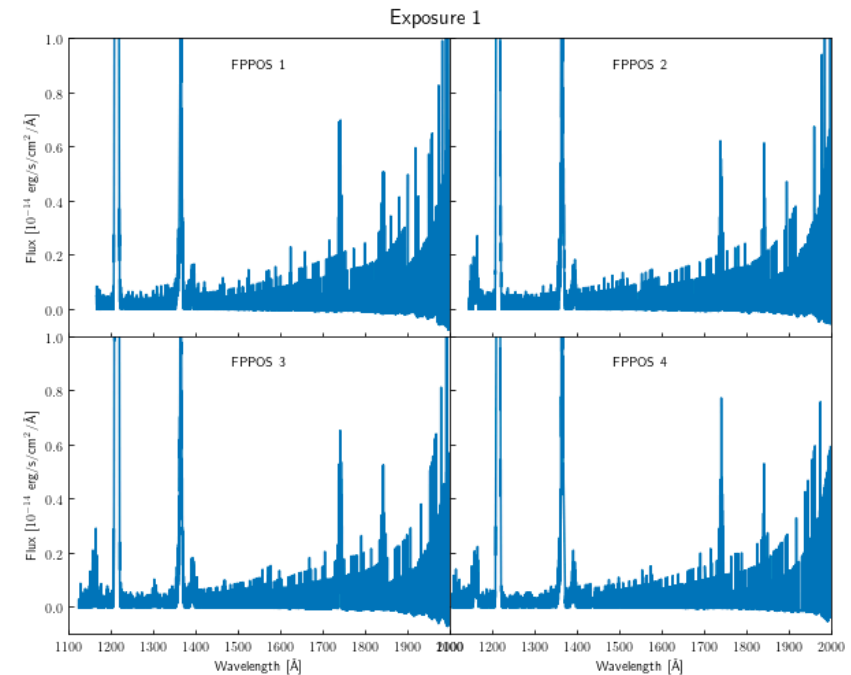
# Nuclear Spectrum

Jean J. Somalwar

# Separate Exposures

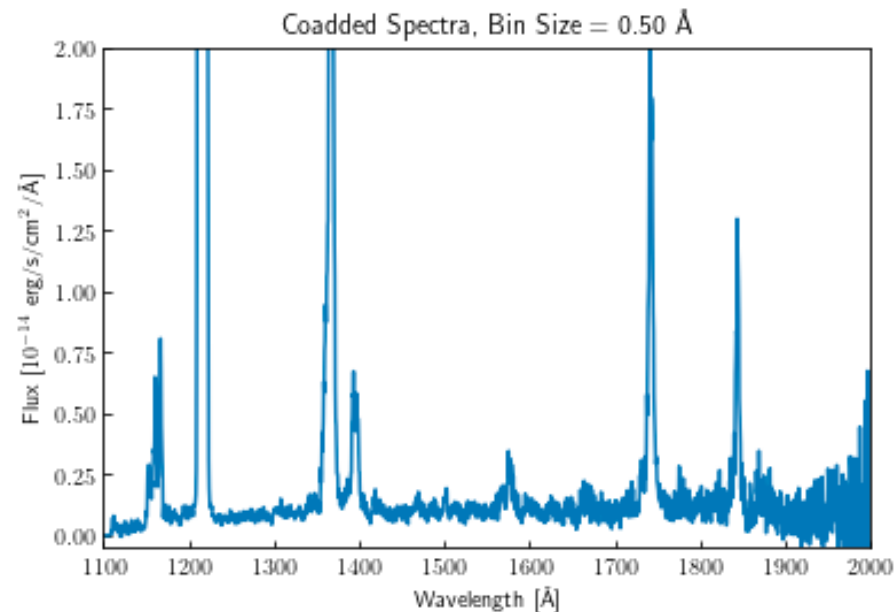
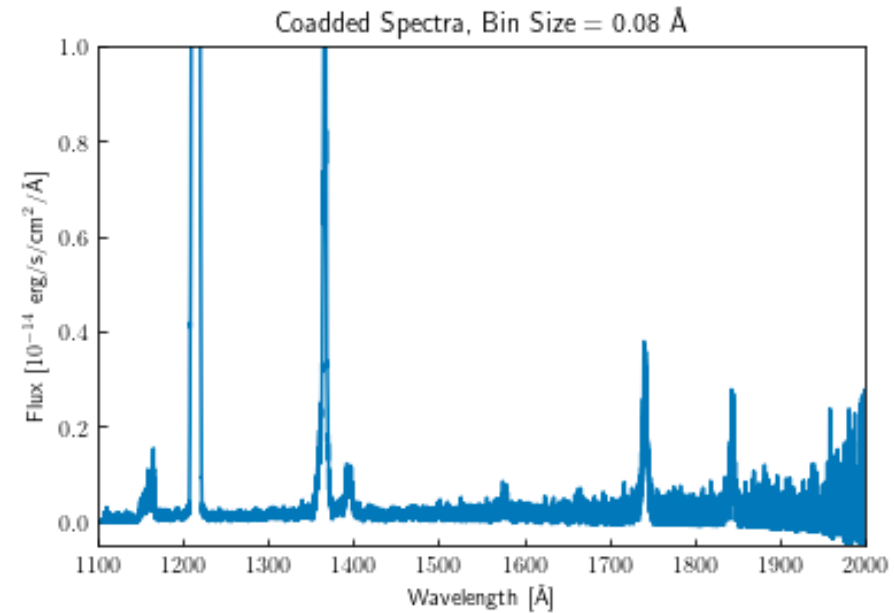
Cuts: DQ = 0 or DQ = 4

- DQ=0: No anomalies
- DQ = 4: Detector shadow



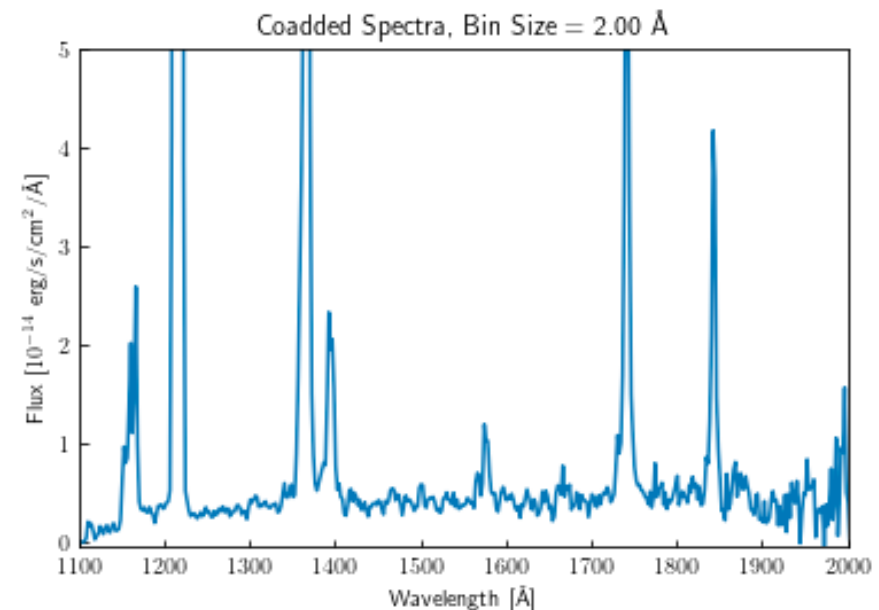
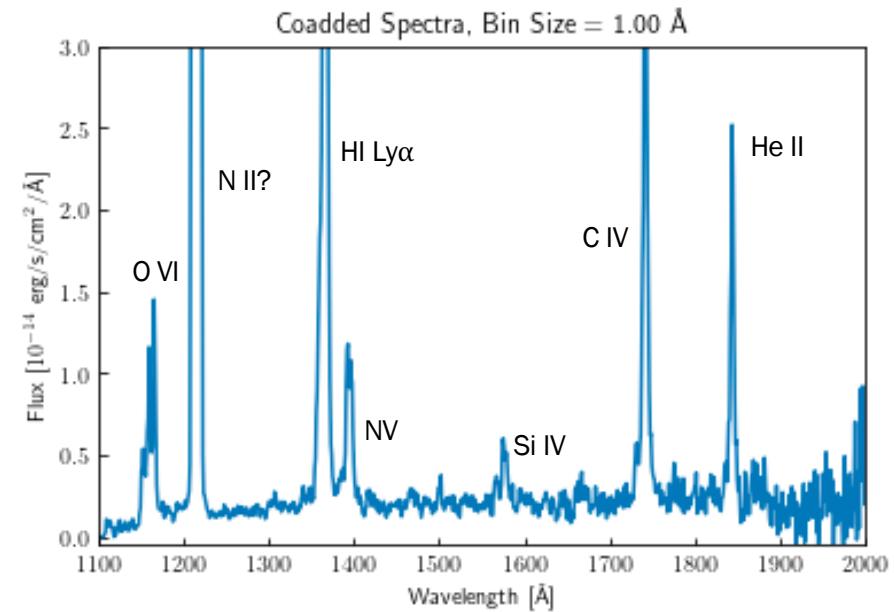
# Stacked Spectra

- Exposure weighted average of the 4 exposures
- Using wavelength bins with width given in plot titles
- Cuts: DQ=0 or DQ=4



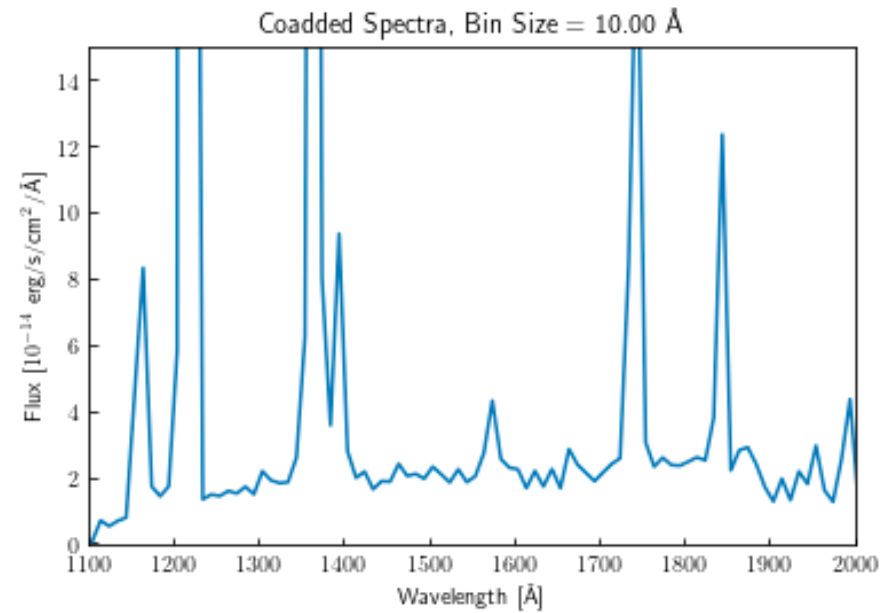
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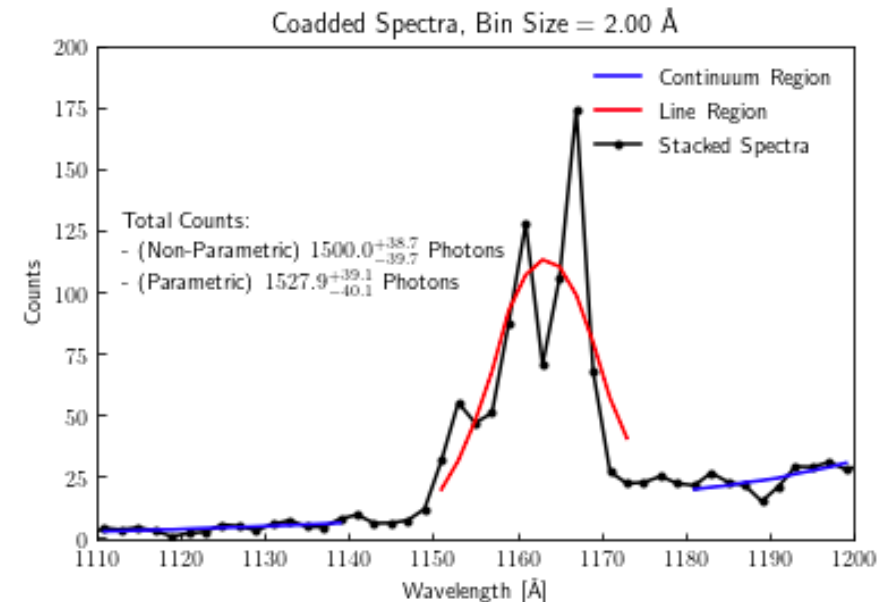
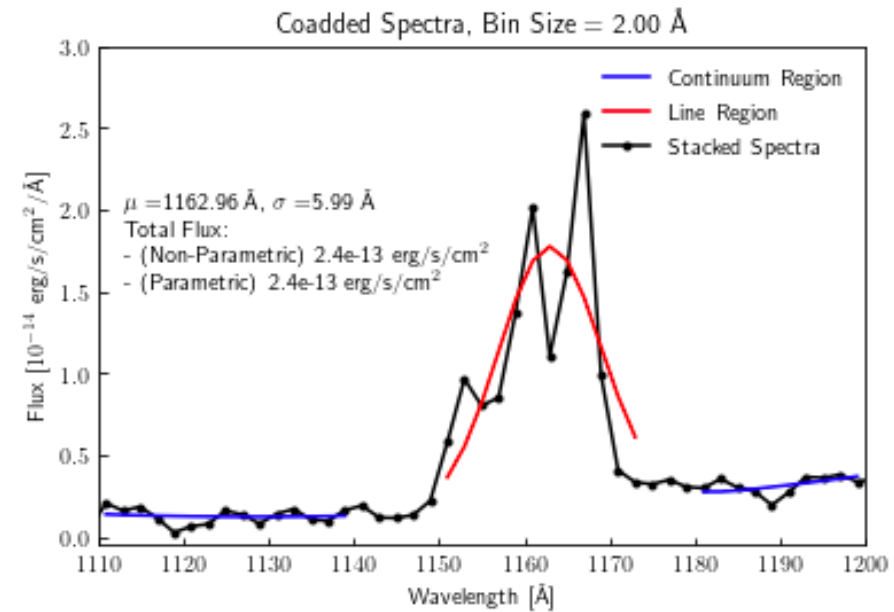


# Lines

Element	Rest-Frame Wavelength (Å)	$z = 0.1234$ Wavelength (Å)
O VI	<a href="#"><u>1031.92, 1037.61</u></a>	<a href="#"><u>1159.26, 1165.65</u></a>
N II?	<a href="#"><u>1083.99</u></a>	<a href="#"><u>1217.75</u></a>
HI Ly $\alpha$	<a href="#"><u>1215.67</u></a>	<a href="#"><u>1365.68</u></a>
N V	<a href="#"><u>1238.82, 1242.8</u></a>	<a href="#"><u>1391.69, 1396.16</u></a>
Si IV	<a href="#"><u>1393.75, 1402.77</u></a>	<a href="#"><u>1565.74, 1575.87</u></a>
C IV	<a href="#"><u>1548.19, 1550.77</u></a>	<a href="#"><u>1742.14</u></a>
He II	<a href="#"><u>1640.4</u></a>	<a href="#"><u>1842.83</u></a>
O III	1665.85	1871.42

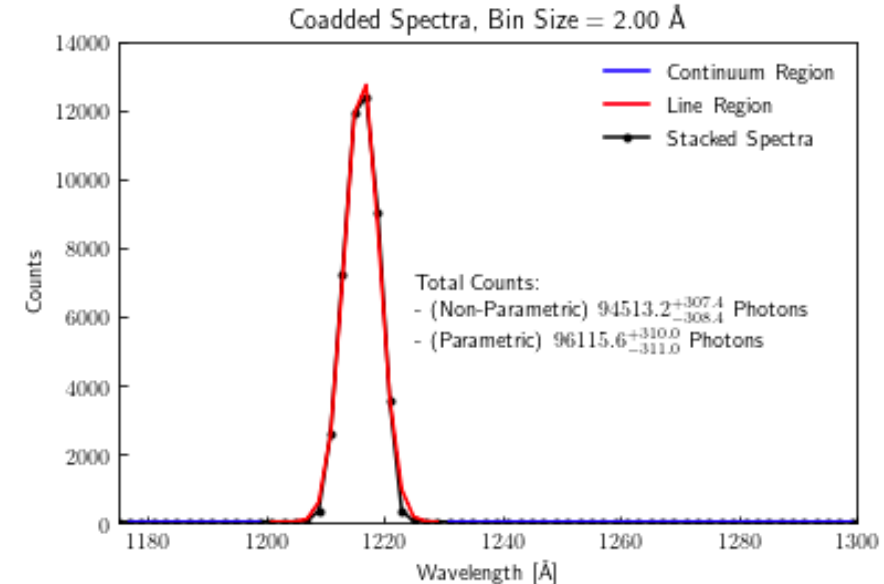
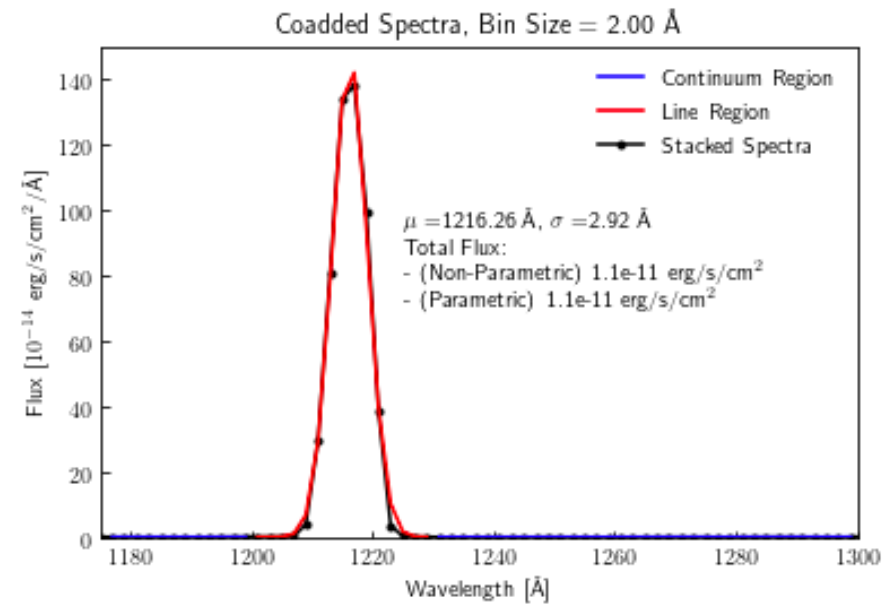
# Line Fits, O VI

- Fitting the lines using the stacked spectra with wavelength bins of width  $1 \text{ \AA}$
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region (region bounds are approximate)
- Red: gaussian fit to line region (region bounds are approximate) (I'm pretty sure some of the lines are actually two lines, but for now I am fitting them with only one gaussian)
- Errors on count values are from astropy's poisson\_conf\_interval



# Line Fits, N II?

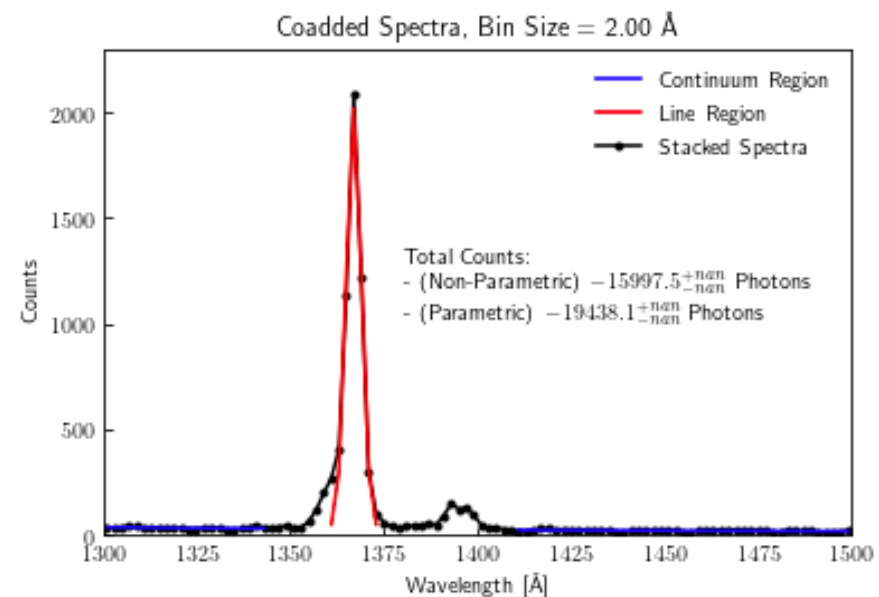
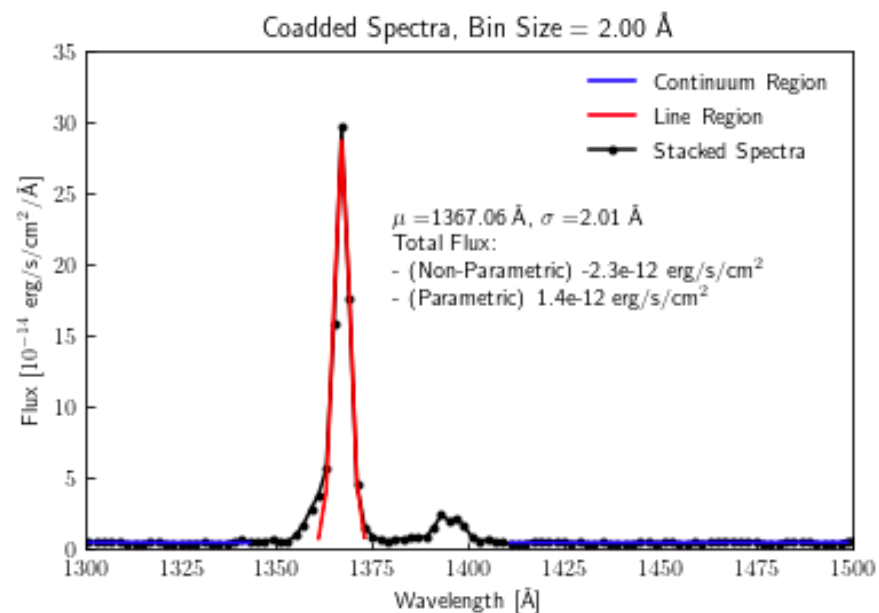
- Fitting the lines using the stacked spectra with wavelength bins of width  $1 \text{ \AA}$
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region (region bounds are approximate)
- Red: gaussian fit to line





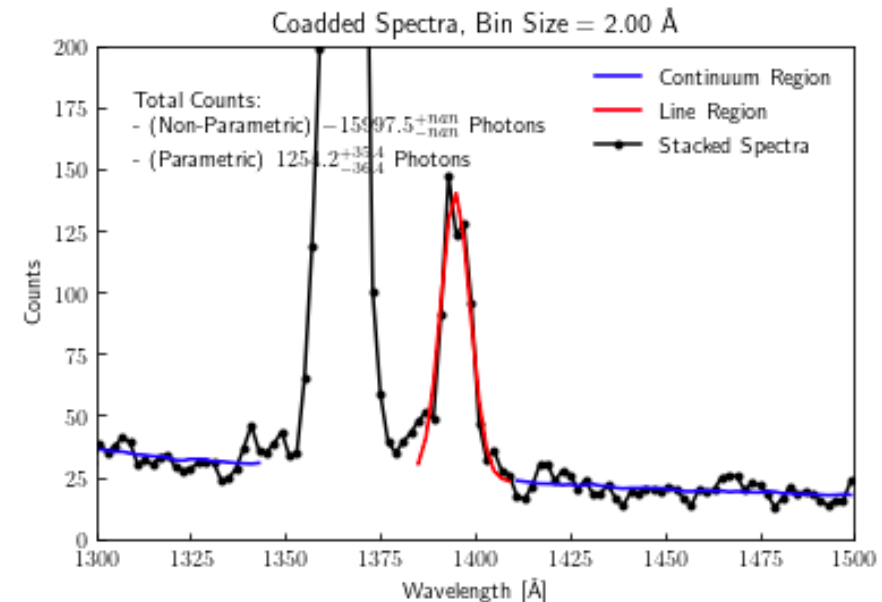
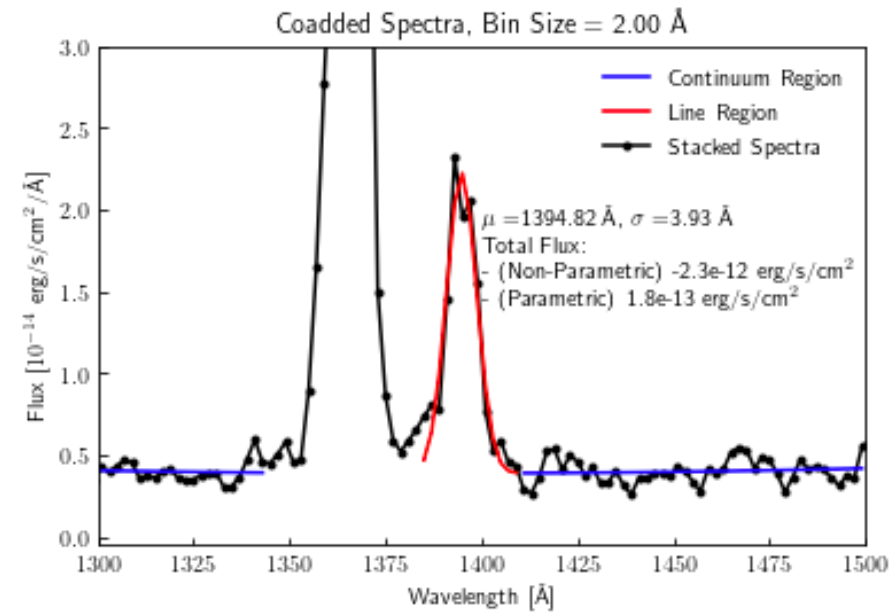
# Line Fits, H I Ly $\alpha$

- Fitting the lines using the stacked spectra with wavelength bins of width 1 Å
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region
- Red: gaussian fit to line region



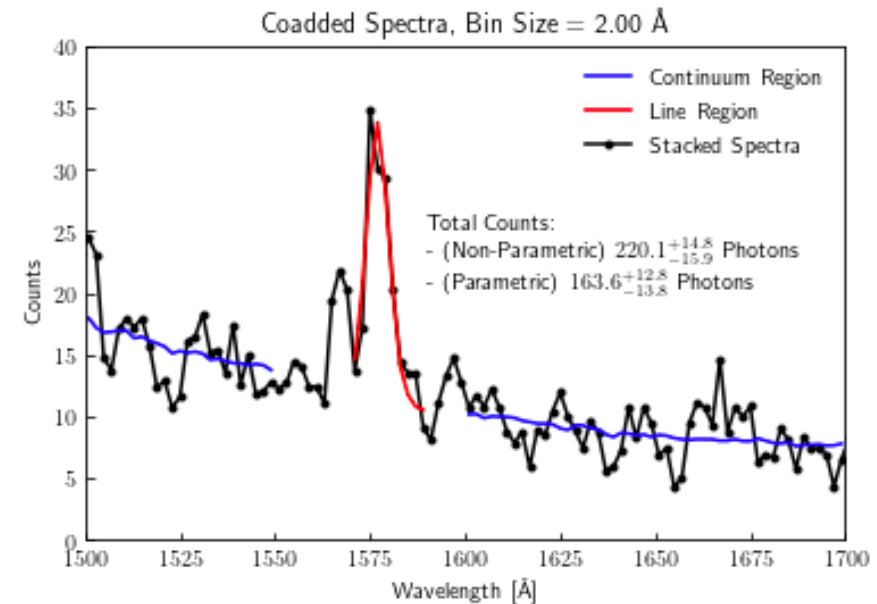
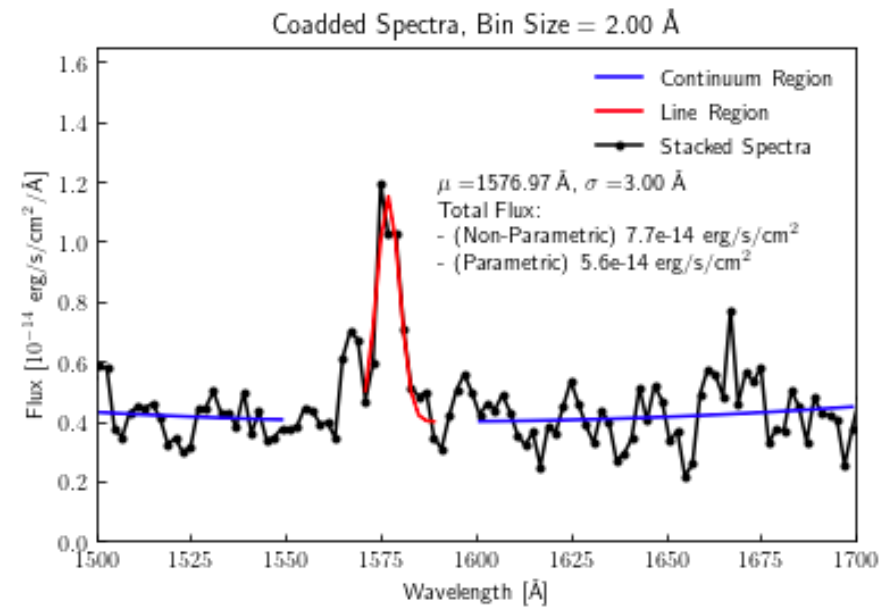
# Line Fits, N V

- Fitting the lines using the stacked spectra with wavelength bins of width  $1 \text{ \AA}$
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region
- Red: gaussian fit to line region



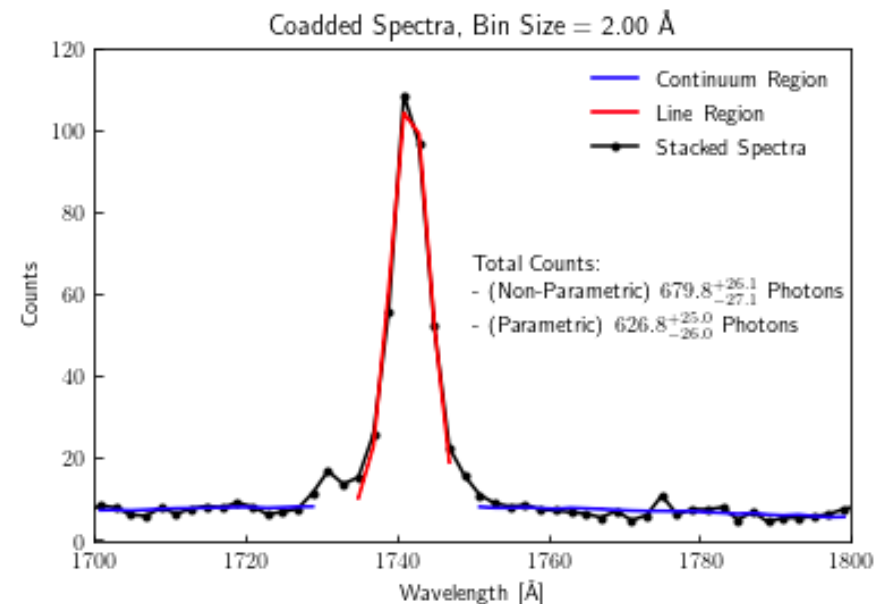
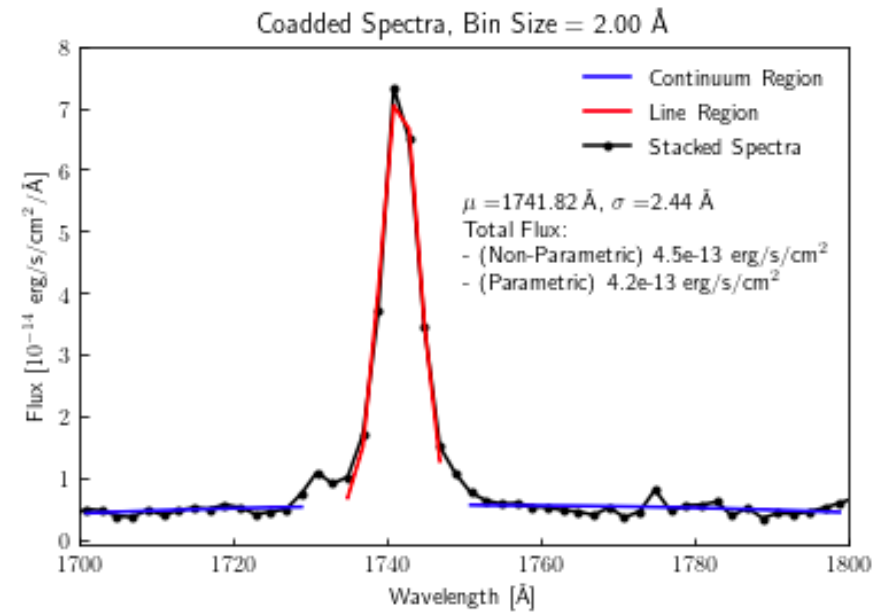
# Line Fits, Si IV

- Fitting the lines using the stacked spectra with wavelength bins of width  $1 \text{ \AA}$
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region
- Red: gaussian fit to line region



# Line Fits, C IV

- Fitting the lines using the stacked spectra with wavelength bins of width  $1 \text{ \AA}$
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region
- Red: gaussian fit to line region



# Line Fits, He II

- Fitting the lines using the stacked spectra with wavelength bins of width  $1 \text{ \AA}$
- Blue: 2<sup>nd</sup> degree polynomial fit to continuum region
- Red: gaussian fit to line region

