

Physics 623 Homework 5

Jeremy Welsh-Kavan

4.3.4. Classical atom

Consider a classical electron in a circular orbit in a Coulomb potential, for which the virial theorem yields $\bar{V} = 2E$.

- a) Since the electron in a hydrogen atom is nonrelativistic, we can use the Larmor formula to describe the electromagnetic radiation of the electron as a point charge.

$$\mathcal{P} = \frac{2e^2}{3c^3} \dot{\mathbf{v}}^2 \quad (1)$$

We can use the equation of motion for the electron to rewrite (1) in terms of E .

$$\begin{aligned} m\dot{\mathbf{v}} &= -\frac{e^2}{r^2} \\ \dot{\mathbf{v}} &= -\frac{V^2}{me^2} \\ \dot{\mathbf{v}}^2 &= \frac{V^4}{m^2 e^4} \\ \overline{\dot{\mathbf{v}}^2} &\approx \frac{\overline{V^4}}{m^2 e^4} \\ \overline{\dot{\mathbf{v}}^2} &\approx \frac{16E^4}{m^2 e^4} \end{aligned} \quad (2)$$

where we have approximated $\overline{V^4}$ with \overline{V}^4 since we may assume the potential is roughly constant over one period of the electron's orbit. Therefore, the average power radiated by the electron is

$$\begin{aligned} \mathcal{P} &= \frac{2e^2}{3c^3} \frac{16E^4}{m^2 e^4} \\ \mathcal{P} &= \frac{32}{3e^2 m^2 c^3} E^4 \end{aligned} \quad (3)$$

- b) We will show that the expected value of the potential diverges to $-\infty$ at finite time. We can rewrite (3) as a differential equation in E

$$-\frac{dE}{dt} = \frac{32}{3e^2 m^2 c^3} E^4 \quad (4)$$

which we can then rewrite in terms of \bar{V} using the virial theorem.

$$\begin{aligned}\frac{1}{2} \frac{d\bar{V}}{dt} &= -\frac{1}{16} \frac{32}{3e^2 m^2 c^3} \bar{V}^4 \\ \frac{d\bar{V}}{dt} &= -\frac{4}{3e^2 m^2 c^3} \bar{V}^4 \\ \Rightarrow \bar{V}(t) &= \frac{1}{\left(\frac{4}{e^2 m^2 c^3} t + 1/\bar{V}_0^3\right)^{1/3}}\end{aligned}\tag{5}$$

We can assume that $\bar{V}_0 \approx -e^2/r_0$, where r_0 is the radius of the electron's orbit at $t = 0$. Then we have

$$\begin{aligned}\bar{V}(t) &= \frac{1}{\left(\frac{4}{e^2 m^2 c^3} t - r_0^3/e^6\right)^{1/3}} \\ \Rightarrow \lim_{t \rightarrow \alpha^-} \bar{V}(t) &= -\infty\end{aligned}\tag{6}$$

since $\frac{4}{e^2 m^2 c^3} t - r_0^3/e^6 \rightarrow 0^-$ as $t \rightarrow \alpha^-$, where $\alpha = \frac{r_0^3 m^2 c^3}{4e^4}$.

If the radius of the initial orbit is $r_0 = 10^{-8}$ cm then the time it takes for the electron to reach the nucleus is 105 ps $\approx 10^{-10}$ seconds.

4.3.5. Absence of dipole radiation

Consider a system of particles whose charge and current densities are

$$\begin{aligned}\rho(\mathbf{x}, t) &= \sum_j e_j \delta(\mathbf{x} - \mathbf{r}_j(t)) \\ \mathbf{j}(\mathbf{x}, t) &= \sum_j e_j \mathbf{v}_j(t) \delta(\mathbf{x} - \mathbf{r}_j(t))\end{aligned}\tag{7}$$

respectively. Assume that $e_j/m_j = \mu$ for all j , where m_j is the mass of the j^{th} particle and μ is fixed. Assume also that there are no external forces or torques on the system. The electric and magnetic dipole moments of this system of charges are then given by

$$\begin{aligned}
\mathbf{d}(t) &= \int d\mathbf{y} \, \mathbf{y} \rho(\mathbf{y}, t) \\
\mathbf{m}(t) &= \frac{1}{2c} \int d\mathbf{y} \, \mathbf{y} \times \mathbf{j}(\mathbf{y}, t) \\
\Rightarrow \mathbf{d}(t) &= \sum_j e_j \mathbf{r}_j(t) \\
\mathbf{m}(t) &= \frac{1}{2c} \sum_j e_j \mathbf{r}_j(t) \times \mathbf{v}_j(t)
\end{aligned} \tag{8}$$

We may rewrite (8) in terms of the masses of the particles and again in terms of the center of mass and angular momentum of the system

$$\begin{aligned}
\mathbf{d}(t) &= \mu \sum_j m_j \mathbf{r}_j(t) \\
\mathbf{m}(t) &= \frac{\mu}{2c} \sum_j m_j \mathbf{r}_j(t) \times \mathbf{v}_j(t) \\
\Rightarrow \mathbf{d}(t) &= \mu M \mathbf{r}_{\text{cm}}(t) \\
\mathbf{m}(t) &= \frac{\mu}{2c} \sum_j \mathbf{l}_j(t) = \frac{\mu}{2c} \mathbf{L}(t)
\end{aligned} \tag{9}$$

where M is the total mass of the system, $\mathbf{r}_{\text{cm}}(t)$ is the center of mass of the system, and $\mathbf{L}(t)$ is the total angular momentum of the system. Since the total force and total torque on the system are both zero, we have

$$\begin{aligned}
\ddot{\mathbf{d}}(t) &= \mu M \mathbf{a}_{\text{cm}}(t) = \mu \mathbf{F}_{\text{ext}} = \mathbf{0} \\
\dot{\mathbf{m}}(t) &= \frac{\mu}{2c} \dot{\mathbf{L}}(t) = \frac{\mu}{2c} \boldsymbol{\tau}_{\text{ext}} = \mathbf{0}
\end{aligned} \tag{10}$$

Therefore, the total power due to electric and magnetic dipole radiation is

$$\begin{aligned}
\mathcal{P} &= \frac{2}{3c^3} \left[\ddot{\mathbf{d}}^2 + \dot{\mathbf{m}}^2 \right] \\
&= 0
\end{aligned} \tag{11}$$

4.3.6. Rotating dipole

An electric dipole moment \mathbf{d} rotates uniformly with angular velocity ω in a plane. We attempt to find the radiated power per solid angle, and the total radiated power, averaged over one rotational period. Since \mathbf{d} rotates in the plane, we can let $\mathbf{d} = (d, 0, 0)$ at $t = 0$ so that, as a function of time, $\mathbf{d}(t) = (d \cos \omega t, d \sin \omega t, 0)$. The power per solid angle, Ω , radiated by the dipole is given by

$$\frac{d\mathcal{P}}{d\Omega} = \frac{1}{4\pi c^3} \left(\hat{\mathbf{x}} \times \dot{\mathbf{d}} \right)^2 \quad (12)$$

To compute the average power per solid angle radiated in one rotational period, we can simply compute

$$\overline{\frac{d\mathcal{P}}{d\Omega}} = \frac{\omega}{2\pi} \int_0^{2\pi/\omega} dt \frac{1}{4\pi c^3} \left(\hat{\mathbf{x}} \times \ddot{\mathbf{d}} \right)^2 \quad (13)$$

With $\hat{\mathbf{x}} = (\cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta)$ and $\ddot{\mathbf{d}} = -\omega^2 \mathbf{d}(t)$, the power per solid angle, averaged over one rotational period, radiated by the dipole is

$$\begin{aligned} \overline{\frac{d\mathcal{P}}{d\Omega}} &= \frac{\omega^4}{4\pi c^3} \frac{\omega}{2\pi} \int_0^{2\pi/\omega} dt \left(\hat{\mathbf{x}} \times \mathbf{d} \right)^2 \\ &= \frac{\omega^4}{4\pi c^3} \frac{\omega}{2\pi} \int_0^{2\pi/\omega} dt \frac{1}{4} (3 + \cos 2\theta - 2 \cos(2(\phi - \omega t)) \sin^2 \theta) \\ &= \frac{\omega^4 d^2}{8\pi c^3} (1 + \cos^2 \theta) \end{aligned} \quad (14)$$

To get the total power radiated, averaged over one rotational period, we just integrate (14) over Ω . This yields

$$\begin{aligned} \overline{\mathcal{P}} &= \int d\Omega \overline{\frac{d\mathcal{P}}{d\Omega}} \\ &= \frac{\omega^4 d^2}{8\pi c^3} \int_0^{2\pi} \int_0^\pi d\phi d\theta (1 + \cos^2 \theta) \sin \theta \\ &= \frac{2\omega^4 d^2}{3c^3} \end{aligned} \quad (15)$$

