

# CHAPTER 29

## SPACE TIME GEOMETRY

*Space bends where time learns how to flow,  
Events whisper distances we cannot know.  
Light draws the rules by which causes agree,  
And geometry remembers what motion must be.*

A manifold is a space that may be curved overall, but looks flat if you zoom in close enough. A manifold is a mathematical space such that: Around every point, there exists a neighborhood that looks like ordinary Euclidean space  $\mathbb{R}^n$ . If it looks like  $\mathbb{R}^n$  we say it is an n-dimensional manifold, i.e.,  $\mathbb{R}^n = \{(x^1, x^2, \dots, x^n) \mid x^i \in \mathbb{R}\}$ .

Spacetime geometry provides the geometric framework of special and general relativity. Space and time are unified into a four-dimensional manifold equipped with a metric that encodes causal structure, intervals, and the motion of matter and fields.

This unification is not merely a change of notation, but a profound reorganization of the geometric and physical foundations of classical mechanics.

### 29.1 THE GEOMETRIC UNITY OF SPACE AND TIME

In classical physics, space and time are treated as fundamentally distinct. Space is the domain of geometry, endowed with distances and angles, while time acts as an external parameter ordering physical processes. This separation, though intuitive, is incompatible with the symmetries revealed by modern physics.

The invariance of physical laws under Lorentz transformations shows that spatial lengths and temporal durations are not absolute. Different observers assign different spatial and temporal coordinates to the same physical process, yet agree on certain invariant quantities. What is preserved is not space alone, nor time alone, but a specific combination of the two.

This insight shifts the basic objects of physical description. Rather than positions evolving in time, the fundamental entities are *events*, each specified by both spatial and temporal coordinates. Physics becomes the study of relations among events, independent of the coordinate system used to label them.

The invariant interval between events replaces Euclidean distance as the primary geometric quantity. This structure naturally classifies separations as timelike, spacelike, or null, embedding causality directly into geometry. Causal relations are no longer imposed externally, but arise from the geometric structure of spacetime itself.

Once invariance is elevated to a guiding principle, geometry becomes the natural language of physics. Tensorial objects acquire direct physical meaning, and coordinate independence becomes synonymous with physical objectivity. In this view, kinematics, causality, and measurement are all expressions of spacetime geometry.

## 29.2 SPACETIME AS A MANIFOLD

Spacetime is modeled as a four-dimensional differentiable manifold whose points represent events. Local coordinate charts label events but have no intrinsic physical significance. Physical laws must therefore be expressed in a form independent of any particular coordinate choice.

A metric tensor defined on this manifold encodes the invariant interval between events, determining lengths, angles, and causal structure. The metric serves as the central geometric object from which physical predictions follow.

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Spacetime geometry unifies kinematics and gravitation:

- ▷ Lorentz invariance arises from flat spacetime geometry
- ▷ Gravitation is interpreted as spacetime curvature
- ▷ Physical laws are expressed in covariant tensor form

## 29.3 CLOSING REMARKS

Spacetime geometry replaces the notion of absolute space and time with an intrinsic geometric structure. By describing physics in terms of invariant intervals, worldlines, and curvature, it provides the foundation for modern relativistic theories. Geometry ceases to be a passive background and becomes an essential component of physical law.

## 29.4 SPACETIME AS A MANIFOLD

Spacetime is modeled as a four-dimensional smooth manifold  $\mathcal{M}$ .

Each event is specified by coordinates

$$x^\mu = (ct, x, y, z)$$

where  $\mu = 0, 1, 2, 3$ .

Coordinate transformations between inertial frames preserve the geometric structure of spacetime.

## 29.5 MINKOWSKI METRIC

### 29.5.1 METRIC TENSOR

In special relativity, spacetime is flat and equipped with the Minkowski metric

$$\eta_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

The spacetime interval between two events is

$$ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$$

This interval is invariant under Lorentz transformations.

### 29.5.2 TIMELIKE, SPACELIKE, AND NULL INTERVALS

Intervals are classified according to the sign of  $ds^2$ .

For timelike intervals

$$ds^2 > 0$$

For spacelike intervals

$$ds^2 < 0$$

For null intervals

$$ds^2 = 0$$

This classification determines causal relationships between events.

## 29.6 FOUR-VECTORS

Physical quantities are expressed as four-vectors to ensure Lorentz invariance.

### 29.6.1 POSITION AND VELOCITY FOUR-VECTORS

The position four-vector is

$$x^\mu = (ct, x, y, z)$$

The four-velocity is defined as

$$u^\mu = \frac{dx^\mu}{d\tau}$$

where  $\tau$  is the proper time.

The normalization condition is

$$\eta_{\mu\nu} u^\mu u^\nu = c^2$$

## 29.6.2 ENERGY-MOMENTUM FOUR-VECTOR

The energy-momentum four-vector is

$$p^\mu = (E/c, \mathbf{p})$$

It satisfies the invariant relation

$$\eta_{\mu\nu} p^\mu p^\nu = m^2 c^2$$

## 29.7 PROPER TIME AND WORLDLINES

The proper time interval along a worldline is defined by

$$d\tau = \frac{1}{c} \sqrt{ds^2}$$

A particle's motion through spacetime is represented by a worldline

$$x^\mu = x^\mu(\tau)$$

Free particles follow straight worldlines in flat spacetime.

## 29.8 LORENTZ TRANSFORMATIONS AS ISOMETRIES

Lorentz transformations preserve the Minkowski metric

$$\eta_{\alpha\beta} \Lambda^\alpha_\mu \Lambda^\beta_\nu = \eta_{\mu\nu}$$

They are therefore isometries of spacetime geometry.

## 29.9 CURVED SPACETIME

In the presence of gravitation, spacetime becomes curved.

The metric is generalized to

$$ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$$

where  $g_{\mu\nu}$  varies with position.

## 29.10 GEODESICS IN SPACETIME

The motion of free particles in curved spacetime follows geodesics.

The geodesic equation is

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma^\mu_{\alpha\beta} \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0$$

where  $\Gamma^\mu_{\alpha\beta}$  are the Christoffel symbols.

## 29.11 CURVATURE AND GRAVITATION

Spacetime curvature is described by the Riemann curvature tensor

$$R^\mu_{\nu\alpha\beta}$$

Curvature determines the relative acceleration of nearby worldlines and encodes the gravitational interaction.

## 29.12 SPACETIME GEOMETRY AND PHYSICS

Spacetime geometry unifies kinematics and gravitation:

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## 29.13 CLOSING REMARKS

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