

# Pyplis\* | A Python based software package for the analysis of volcanic SO<sub>2</sub> emissions using UV SO<sub>2</sub> cameras

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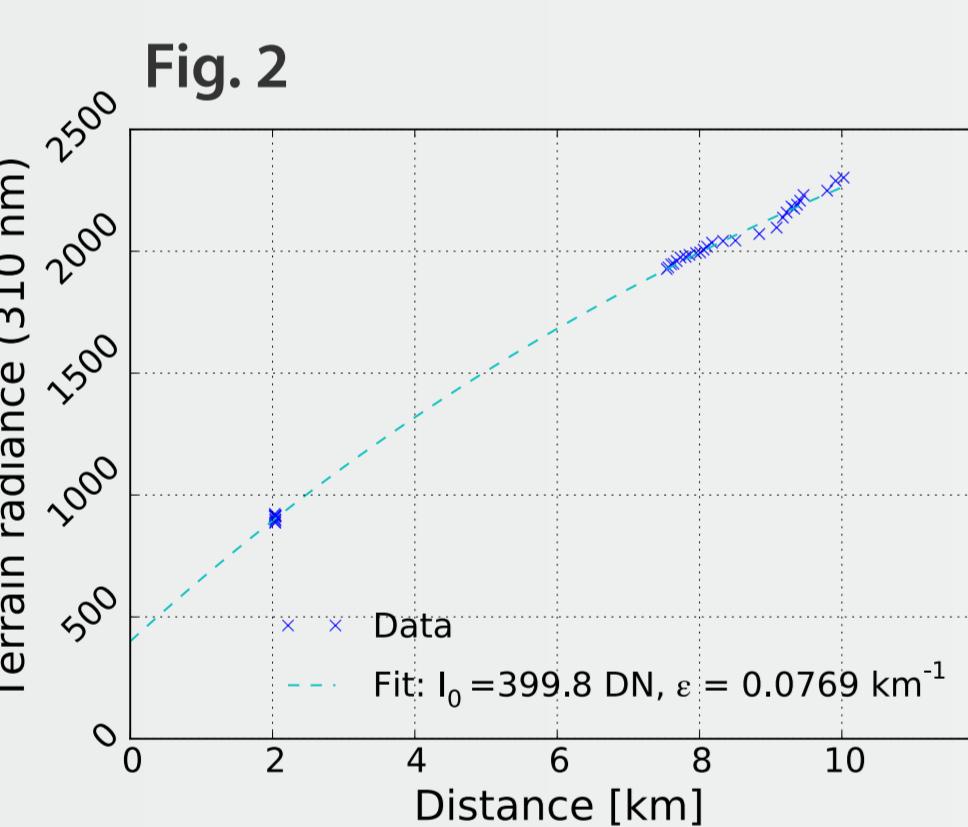
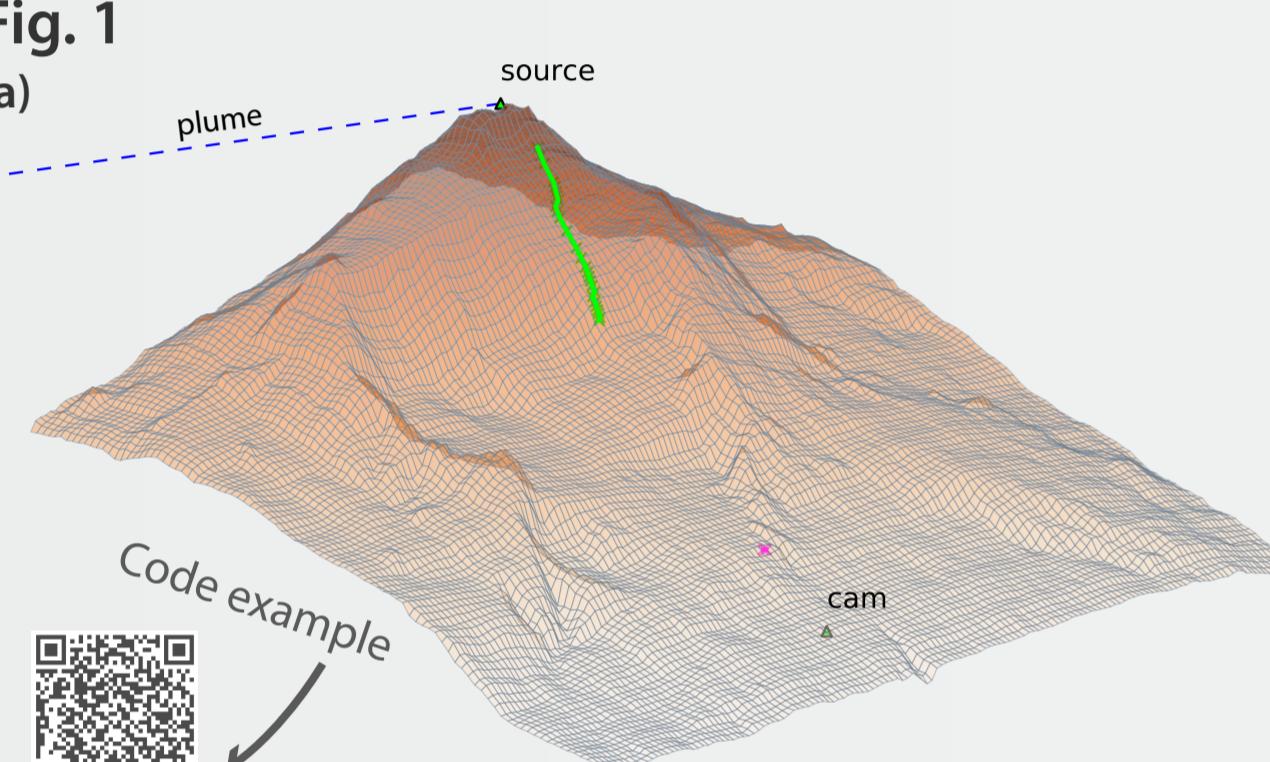


## Main Features

- △ Easy setup of various data formats and camera specifications
- △ Detailed 3D geometrical calculations
- △ Routine for image based correction of the signal dilution effect
- △ Flexible options for retrieval of plume background intensities
- △ Engines for automated cell and DOAS calibration
- △ Optical flow and cross correlation based plume velocity retrievals
- △ Simultaneous emission rate retrievals for multiple retrieval lines

\* due to a naming conflict, the software was renamed from *piscope* to *pyplis*

## 1 Measurement geometry and signal dilution correction



$$\text{Equation 1 | Atmospheric scattering model}$$

$$I_{\text{meas}}(\lambda) = I_0(\lambda)e^{-\epsilon(\lambda)d} + I_A(\lambda)(1 - e^{-\epsilon(\lambda)d})$$

Legend:  
 $I_{\text{meas}}(\lambda)$ : measured intensity of terrain feature  
 $I_0(\lambda)$ : initial intensity to terrain feature  
 $\epsilon(\lambda)$ : average ambient intensity  
 $d$ : scattering extinction coefficient

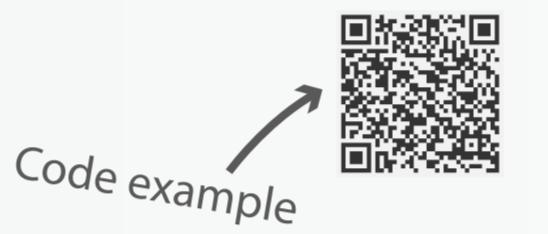
### Infobox 1 | Geometrical calculations

Plume distances (required for emission rate retrieval) are retrieved on a pixel-level based on camera coordinates and plume azimuth. More features include:

- Retrieval of camera viewing direction using distinct features in the images (e.g. volcano summit)
- Retrieval of distances to local terrain features in the images on pixel-level (see Fig. 1)

Fig. 1a: 3D terrain map of the measurement setup including 2 topographic profile lines (green and magenta, defined in the image, cf. Fig. 1b) used for the signal dilution analysis (Fig. 2)

Fig. 1b: Exemplary on-band (310 nm) image. Due to SO<sub>2</sub> absorption the Etna plume appears darker than the sky background. Measured intensities along the two indicated lines are used to estimate the scattering extinction (cf. Fig. 2).



## 2 Plume background analysis

$$\text{Equation 2 | Optical density (OD)}$$

$$\tau = \ln \left( \frac{I_0}{I} \right)$$

$I$ : measured plume intensity  
 $I_0$ : corresponding background intensity

$$\text{Equation 3 | SO}_2\text{ apparent absorbance (AA)}$$

$$\tau_{\text{AA}} = \tau_{\text{on}} - \tau_{\text{off}}$$

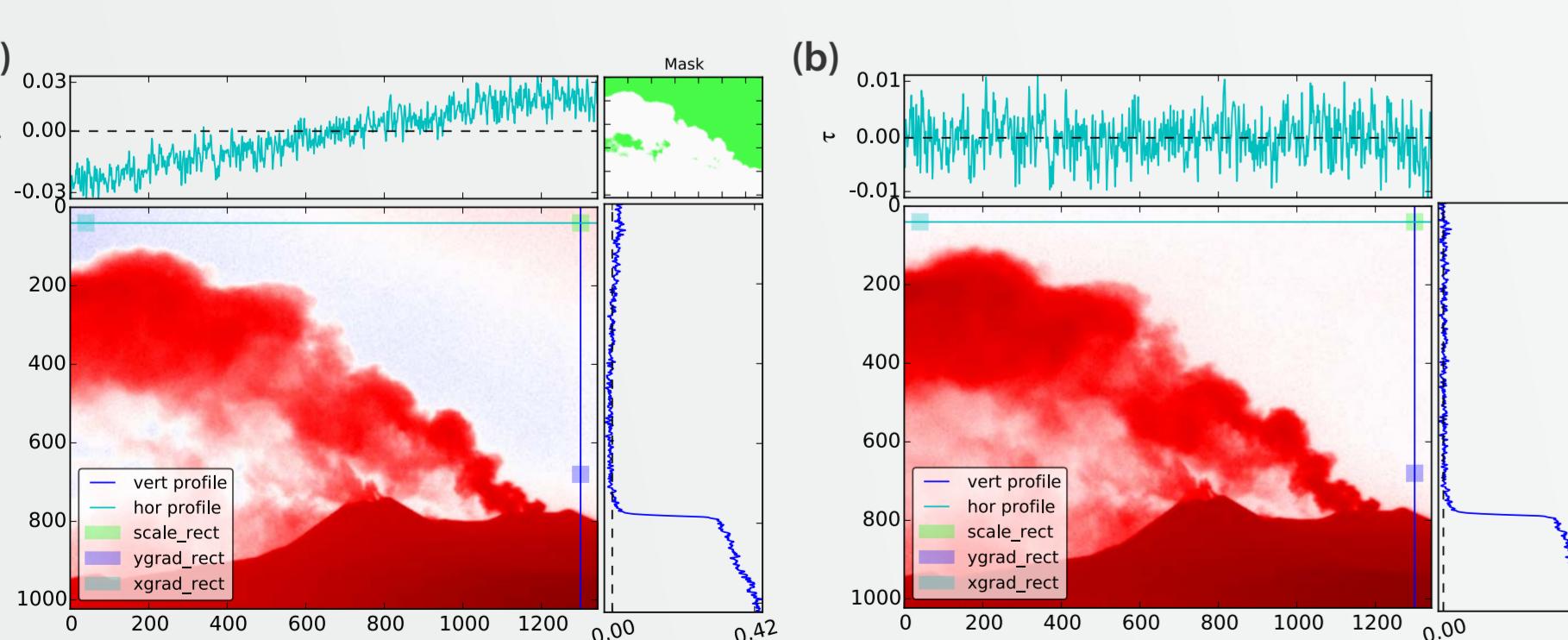


Fig. 3: Exemplary on-band OD images calculated using a masked 2D polynomial fit (a) and using an additional sky background image, (b) including 2. order horizontal and vertical sky gradient corrections along drawn profile lines.

### Infobox 3 | Calculation of optical density (OD) images

The calculation of OD images (Eq. 2) requires knowledge of the sky background intensities behind the plume. These can either be retrieved from the plume images directly (by fitting a 2D polynomial to clear sky areas, cf. Fig. 3a) or using a sky reference image which is scaled using a gas free sky area in the plume image (cf. scale\_rect in Fig. 3). The latter may include further corrections to account for variations in the sky background curvature between the two images (cf. Fig. 3b).

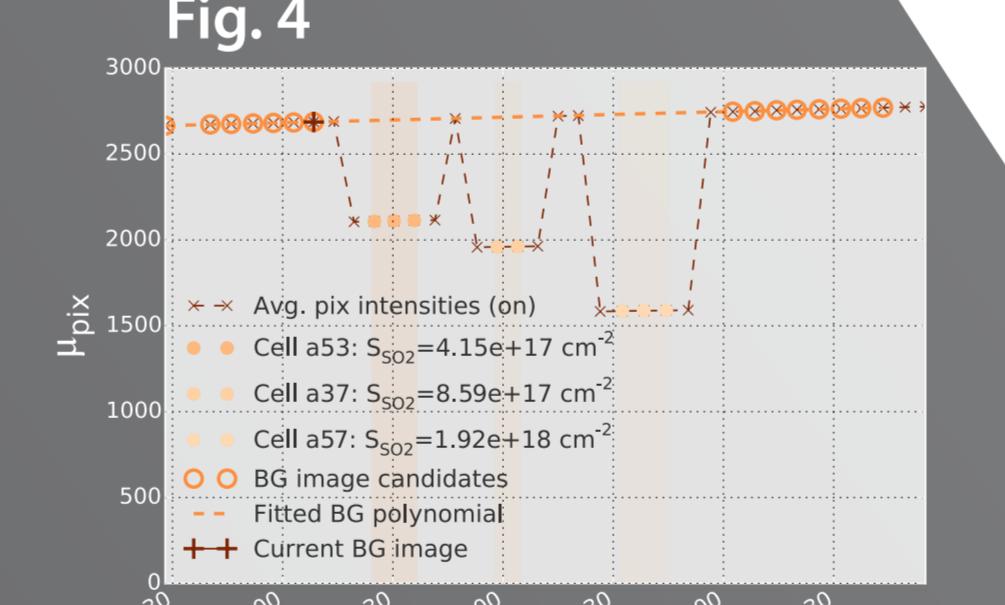
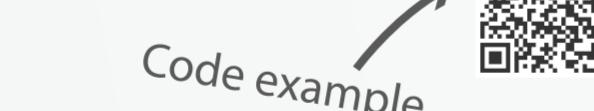


Fig. 4: Automatic cell search algorithm  
Time series of average pixel intensities from on-band images (dashed line) including automatically detected time windows of individual cells based on a gradient analysis of the signal.

Fig. 5 (left): Cell and DOAS calibration curves  
Figure showing exemplary AA cell (magenta) and DOAS calibration data (green) including fitted polynomials.

### Infobox 5 | DOAS calibration

The DOAS calibration is performed using plume SO<sub>2</sub>-CDs retrieved with a co-located DOAS spectrometer. Position and shape of the DOAS FOV within the camera images (cf. Fig. 6) can be retrieved using either of the two methods suggested by [2] and [3]. Both methods identify the pixel region showing highest correlation between a time series of DOAS-CDs and camera images. The calibration curve (Fig. 5) is then retrieved from a time series of AA values (extracted within the DOAS FOV) and the corresponding DOAS SO<sub>2</sub>-CDs. The DOAS calibration is more accurate in case aerosols are abundant in the plume [2].

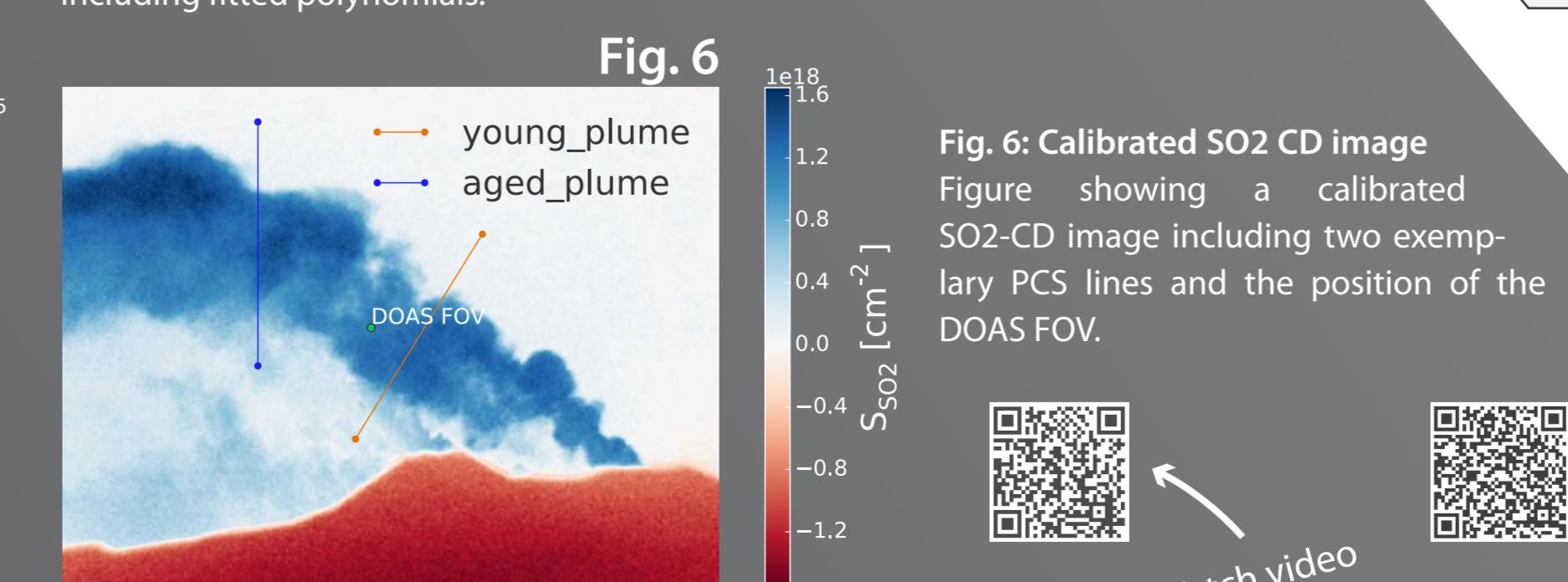
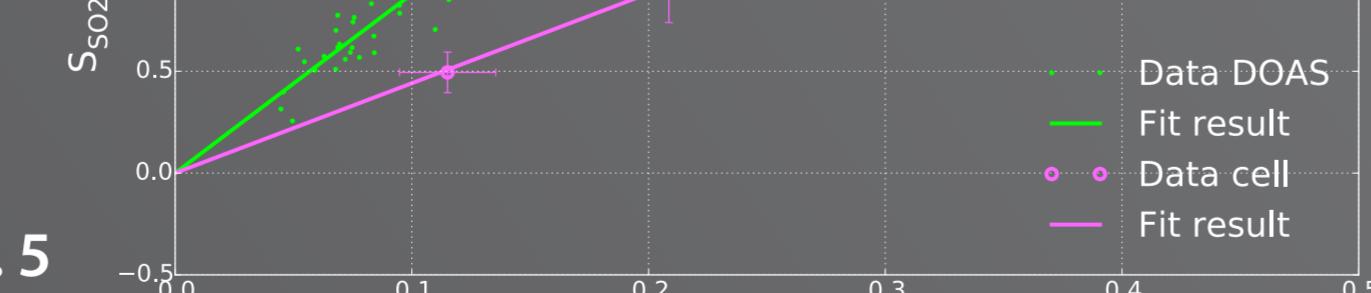
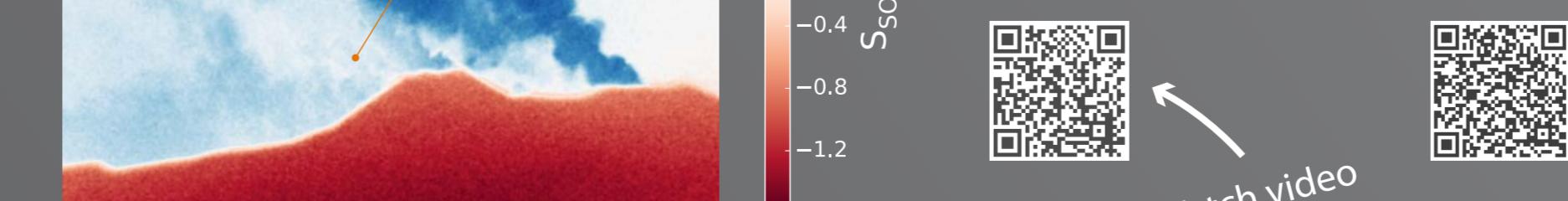


Fig. 6: Calibrated SO<sub>2</sub> CD image  
Figure showing a calibrated SO<sub>2</sub>-CD image including two exemplary PCS lines and the position of the DOAS FOV.



## Background

UV SO<sub>2</sub> cameras have become a common tool for monitoring SO<sub>2</sub> emission rates from volcanoes. The latter are inferred based on plume optical densities measured in two UV wavelength windows around 310 nm (distinct SO<sub>2</sub> absorption, ‘‘on-band’’) and 330 nm (weak SO<sub>2</sub> absorption, ‘‘off-band’’). The data analysis comprises geometrical calculations, the retrieval of plume background intensities and the camera calibration as well as the retrieval of gas velocities within the plume. Emission rates are then retrieved by integrating a projected plume cross section (PCS, e.g. a straight line crossing the plume).

Pyplis is a cross platform, open source software toolbox for the analysis of UV SO<sub>2</sub> camera data written in Python. The software comprises a comprehensive selection of algorithms based on the most commonly used analysis methods, and furthermore, including recent developments and newly developed methods.

Here, we present the main features of Pyplis using an exemplary dataset recorded at Mt. Etna, Italy, on 16 September 2015. Pyplis is freely available and can be downloaded from the website together with an extensive collection of example scripts, providing an easy start into Pyplis. The Etna example data is freely accessible on the website.

Fig. 7

Infobox 6 | Cross correlation  
The plume velocity is retrieved based on the cross correlation lag between two time series of integrated-column-amounts (ICA) of two plume intersections (located at different positions in the images, e.g. [4]).

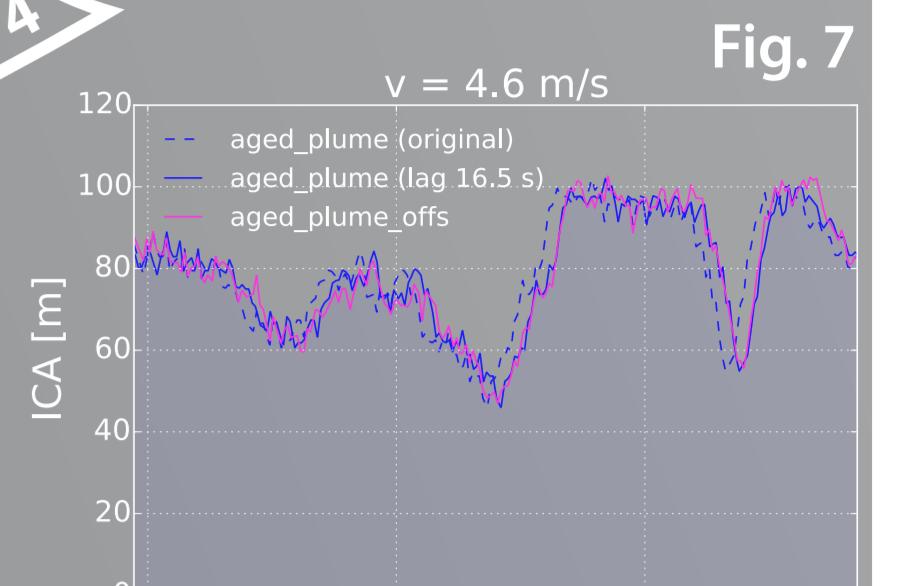


Fig. 7

Fig. 8

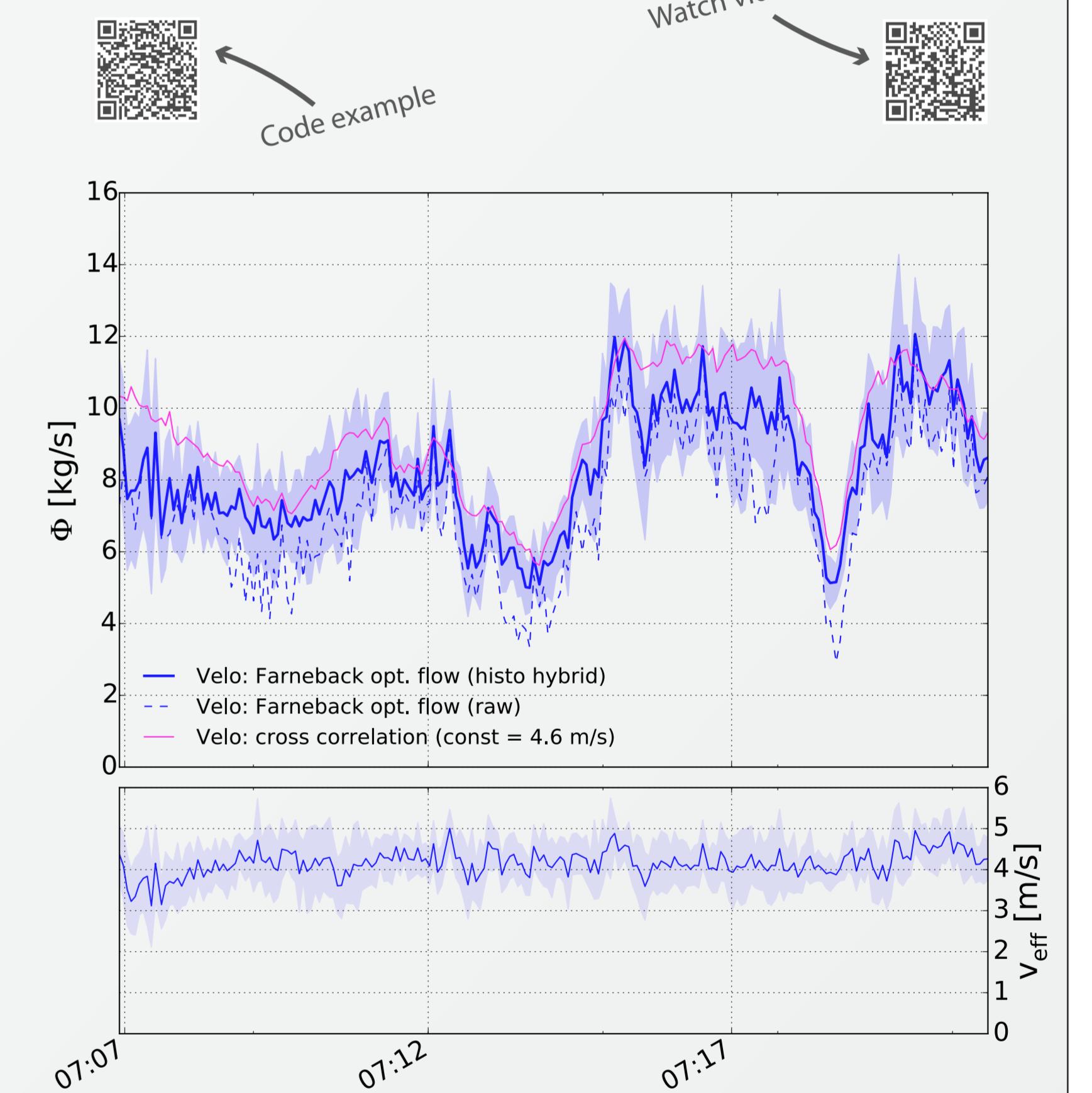
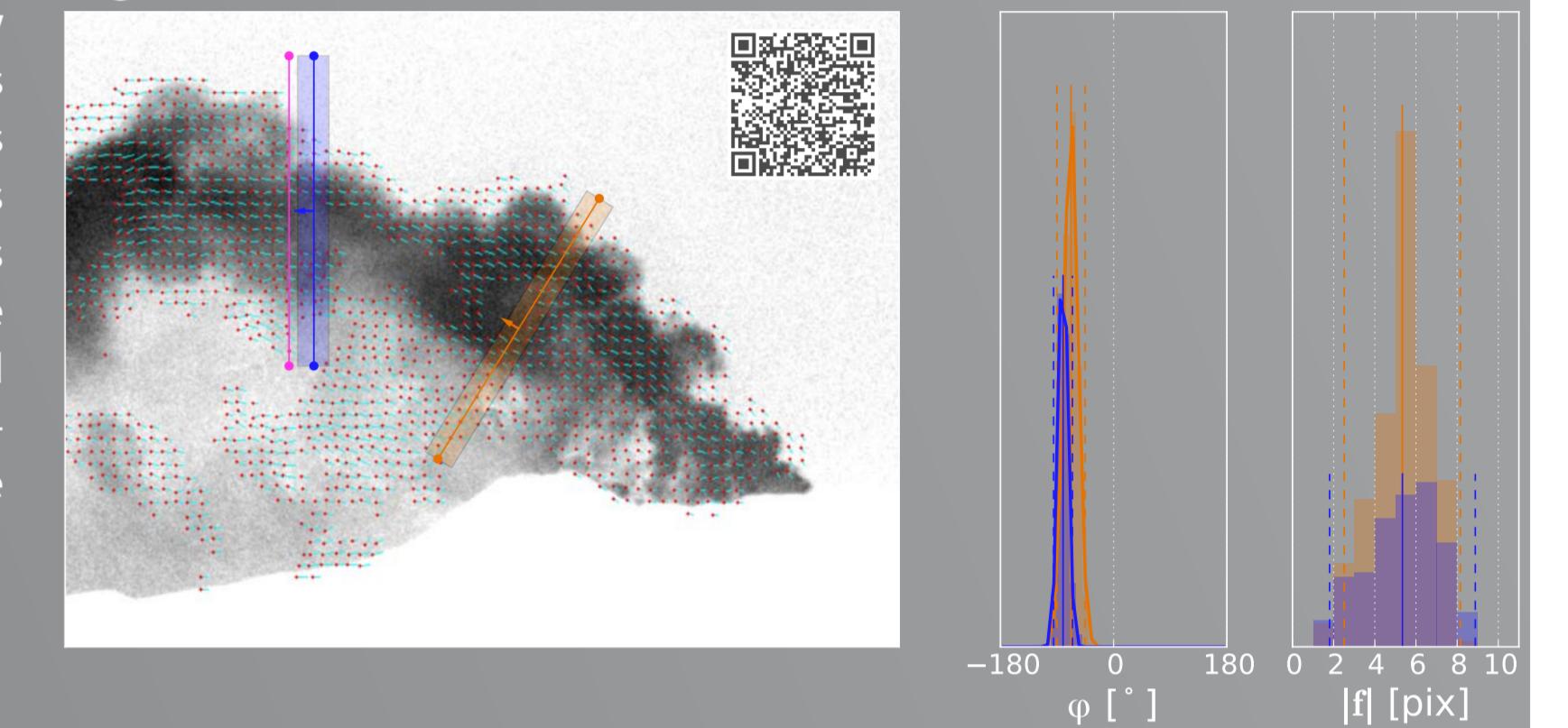


Fig. 9

Fig. 9