## Monte Carlo Markov Chain

Likelihoods, Priors, Model comparison



#### Binomial data - Beta prior

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$$\pi(p; a, b) = \frac{\Gamma(a+b)}{\Gamma(a)\Gamma(b)} p^{a-1} (1-p)^{b-1}$$

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#### Normal data - Normal prior

Let's consider normally distributed Data, with a normal prior on the mean (we keep the variance fixed here):

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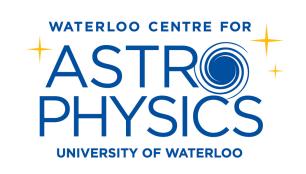
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- If the data distribution is binomial, then the conjugate family of distributions is Beta.
- If the data distribution is normal with known variance, the conjugate family of distributions is normal.

## The Bayesian guy

# "A subjective Bayesian is a person who really buys the Bayesian philosophy."

Someone on Internet about Bayesian Statistics.



#### Loosing priors

Priors are somewhat important in Bayesian statistics as they contribute to the overall uncertainty of the quantities we are interested in.

It is not uncommon that people don't really bother about priors in their Bayesian analyses. When the signal in our data is large, the likelihood outweighs the prior and it becomes subdominant...

So PAY ATTENTION TO IT!



#### Loosening priors

However, we may have enough S/N in our data to allow for less "physics motivated" priors.

In such cases, we could choose priors for mathematical convenience rather to accurately express uncertainties.

We would use priors with a very spread distribution, that represent "extreme" uncertainty. Something we might define "vague" or "diffuse', despite it is not a rigorous mathematical definition...

Those very loose priors are called improper priors.



#### Normal example

Let's consider the normal likelihood with normal prior case:

$$\mathcal{L}(d) = N(\mu, \sigma) \propto e^{-\frac{(\mu - d)^2}{2\sigma^2}} \qquad \pi(\mu) \propto e^{-\frac{(\mu - \mu_0)^2}{2\sigma_0^2}}$$

In the limit of the prior variance  $\sigma_0^2 \to \infty$  the prior tends towards a constant unitary value.

In this case, the posterior is therefore just the likelihood:

$$P(\theta \mid d) = \mathcal{L}(d)\pi(\theta) = \mathcal{L}(d)$$



#### Frequentist and Bayesian agreement

Interestingly, the Bayesian with such a prior agrees with the Frequentist:

The maximum likelihood estimatore (MLE) is  $\hat{\mu}_n = \bar{x}_n$  and we know the exact sampling distribution of the MLE is:

$$\hat{\mu}_n \sim \mathcal{N}(\mu, \frac{\sigma^2}{n})$$

Where we are interested in determining  $\mu$ , whereas the variance is fixed.



#### Frequentist and Bayesian agreement

From a frequentist point of view, the confidence interval would computed from:

$$\hat{\mu} - \mu \sim \mathcal{N}(0, \frac{\sigma^2}{n})$$

The Bayesian sees  $\hat{\mu}$  as fixed by the data and  $\mu$  as a random quantity, whereas the Frequentist tells  $\mu$  is fixed by the model, and  $\hat{\mu}$  is random... but both agree on the distribution of  $\hat{\mu} - \mu$  with this improper prior definition!



Be very very mindful about this..

We this kind of priors improper because they don't really make sense from a mathematical point of view.

Let's look at the Bayes' rule  $P(\theta | d) = \mathcal{L}(d)\pi(\theta)$ .

Here the prior is a probability distribution, whereas we just defined a prior whose integral diverges (can't be normalized, and therefore does not live in the mathematical space of probability distributions).

However the posterior results in something that can be normalized!



Be very very mindful about this...

We are using the Bayes rule's form but not the content; some people say we are using the **formal Bayes rule**.

There is no guarantee that an improper prior leads to a posterior that does integrate. If it is not the case, our result makes no sense.



# Improper priors Summary

Improper priors are very questionable.

- Subjective Bayesians think they are nonsense. They do not correctly describe the uncertainty of anyone.
- Everyone has to be careful using them, because they don't always yield proper posteriors. Everyone agrees improper posteriors are nonsense.
- Because the joint distribution of data and parameters is also improper, paradoxes arise. These can be puzzling.

However they are widely used and need to be understood.



# "Objective" Bayesian Inference Flat prior

The most obvious "default" prior is flat (constant), which seems to express no preference for any parameter value over any other.

If the parameter space is unbounded, then the prior is improper.

The real issue with flat priors is that they are flat for only one parameterization.



## "Objective" Bayesian Inference

#### Reparameterization

Recall the change-of-variable formulas; in many dimensions:

$$f_Y(y) = f_X(h(y)) \cdot |\det(\nabla h(y))|$$

Where x = h(y) and there is the Jacobian term  $\nabla h(y)$ .

Let's use a flat prior, and change variable into  $\varphi = \theta^2$ :

$$h(\theta) = \theta^2 \quad h'(\theta) = 1/2\sqrt{\theta}$$

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# "Objective" Bayesian Inference Jeffreys prior

Another choice revolves around the Fisher information  $I(\theta)$ . The Jeffreys prior is therefore defined as:

$$\pi(\theta) \propto \sqrt{I(\theta)}$$

This is objective because any reparameterization yields the Jeffreys prior for that parameter. Can you see why?

If the parameter space is unbounded, then the Jeffreys prior is usually improper.



# "Objective" Bayesian Inference Jeffreys prior

In a multivariate case, for a parameter vector  $\vec{ heta}$  and Fisher matrix  $\mathbf{I}(\vec{ heta})$ :

$$\pi(\vec{\theta}) \propto \sqrt{\det(\mathbf{I}(\vec{\theta}))}$$

With the same property of the univariate Jeffreys prior: any reparameterization yields the Jeffreys prior for that parameter.



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Point estimates are not really interesting from a Bayesian point of view because in principle we have access to the full parameter posterior distribution, which gives a whole lot of information —much more than a single compressed number.



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We usually compute Mean, Median and Mode (usually referred as best-fit in cosmology) of our parameter posterior distribution.



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Long story short it is a maximum likelihood where instead of maximizing  $\mathcal{L}(d \mid \theta)$  we maximize  $\mathcal{L}(d \mid \theta)\pi(\theta)$ , the posterior.



#### **Credible intervals**

Guess where the Bayesian gets the information on uncertainty from...

From a Bayesian point of view, the interval estimate is inferred from the posterior distribution.

The Bayesian does not quite like the term "confidence interval", for that is associated with a frequentist notion and refers to the confidence one has that repeating an experiment many times, one gets the parameter within a certain interval (remember that the parameter is a number in the frequentist approach).

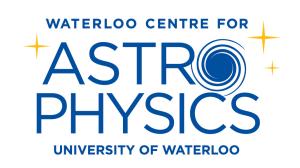
In the Bayesian approach "credible interval" sounds much more appropriate, because it describes the probability of a parameter (that is a PDF) to be within some interval.

#### Let's include an additional term

When giving the Bayes rule, we always implicitly assumed a model to our data, and dropped the relative notation. However, we should now make it explicit that this is something that affects the analysis overall:

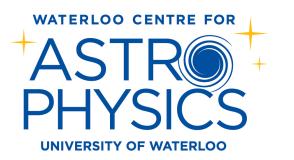
$$P(\theta \mid d, M) = \frac{P(d \mid \theta, M)P(\theta \mid M)}{\int P(d \mid \theta, M)P(\theta \mid M)d\theta}$$

So, our parameters actually depend (of course) on the model choice. The question is, what if we had different models, and wanted to test data to see which model better describes them?



#### **Bayes factor**

Let  $\mathscr{M}$  be a finite or countable set of models. For each model  $m \in \mathscr{M}$  we have a prior probability  $\pi(m)$ . Each model has a parameter space  $\Theta_m$  with its prior  $\pi(\theta \mid m)$ , with  $\theta \in \Theta_m$ .



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The data likelihood then is, for each model:

$$\mathcal{L}(d) = P(d \mid m, \theta)$$



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This gives the unnormalized posterior of the model, the normalized one is:

$$P(m \mid d) = \frac{P(d \mid m)}{\sum_{i \in \mathcal{M}} P(d \mid m_i)}$$



#### **Bayes factor**

Let's define the quantity:

$$b(d \mid m) = \int_{\Omega_m} P(d \mid \theta, m) P(\theta \mid m) d\theta$$

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Bear in mind that the normalization is the real problem in Bayesian model selection. We need to explore a large parameter space and we need to compute its value in the whole phase space.