









Multi-approach gravity field models from Swarm GPS data

TN-01: Standards and Background Models

Delft University of Technology (TU Delft) Astronomical Institute of the University of Bern (AIUB) **Astronomical Institute Ondřejov (ASU) Institute of Geodesy Graz (IfG) Ohio State University (OSU)**

> **Version 1G** 2019-04-09

	Approved:	
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Version history

Version 1, 2017-10-29

• Initial release.

Version 1A, 2017-11-12

- Version numbering is now adhering to the DISC conventions;
- Added logos of Swarm and DISC to the cover page;
- Added section on the dissemination of data, specifically what concerns packaging, in Section 7.2;
- Added version history;
- Added document version to the header;
- Corrected naming convention of the deliverables, in Section 5.1;
- Added the naming convention of the version numbers, in Section 6.1.

Version 1B, 2017-11-16

- Modified the format of the Kinematic Orbits, in Section 6.2;
- Minor typographical corrections in the bibliography.

Version 1C, 2018-1-19

- Added the format description, naming convention and directory locations for the non-gravitational accelerations;
- Fixed typo at the bottom of the title page.

Version 1D, 2018-1-22

 Added the format description, naming convention and directory locations for the GPS data weights.

Version 1E, 2018-3-14

• Modified the guideline for Gravity Field Models and Normal Equations file name version numbering.

Version 1F, 2018-4-24

• Corrected units of the clock correction in the orbit format from TU Delft:

Version 1G, 2018-10-27

• Updated processing details for OSU version 02 models;

Version 1H, 2019-04-09

• Updated processing details for IfG processing;

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1 Introduction

The objective of this document is to address Task 1. According to the Statement of Work (SoW), this task includes the following activities (cf. Section 4.2.1.2 in that document):

- 1. Description of Swarm specific adaptations of the gravity field processors
- 2. Adapt gravity field processors to Swarm L1 products
- 3. Description of standards and background models
- 4. Implement GRACE and GRACE-FO standards in gravity field processor

This document pertains mostly to point 3, as presented in Section 4. Points 2, 3 and 4 have already been completed by all partners, as demonstrated by Jäggi et al. (2016) (for AIUB), Bezděk et al. (2016) (for ASU), Zehentner and Mayer-Gürr (2016) (for IfG) and Shang et al. (2017) (for OSU). Unlike what is specified in point 4, all partners adhere to European Gravity Service for Improved Emergency Management (EGSIEM) data standards, cf. Section 6.

Section 2 compares the processing of the Kinematic Orbits (KOs) for TU Delft, AIUB and IfG. Section 3 illustrates a few details of the Kinematic Baselines (KBs) for TU Delft and AIUB. Finally, Sections 5, 6 and 7 describes the file name conventions, data formats and directory structure of the data exchange server (respectively), used in the project activities.

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2 Kinematic Orbits

2.1 Delft University of Technology

Software: GPS High precision Orbit determination Software Tool (GHOST)

(Helleputte 2004; Wermuth, Montenbruck and Helleputte

2010)

Differencing Scheme: Undifferenced **Linear combination:** Undifferenced Ionosphere-free

GPS observations: Code and carrier phase

Estimator: Bayesian weighted Least-Squares (LS)

Arc length: 30 hours

Data weighting: a-priori weights equal to 1m and 1mm for code and phase

observations (resp.)

Transmitter PCV: IGS08.atx model (Schmid et al., 2007) **Receiver PCV:** Empirical, derived from 70 days of data

Data screening: minimum Signal-to-Noise Ratio (SNR) of 10, minimum of 6

GPS satellites, code and phase outlier editing threshold of 2 m and 3.5 cm, respectively, 1 meter or larger difference between estimated KO positions and with Reduced-Dynamic Precise

Science Orbit (PSO)

Earth precession model: International Astronomical Union (IAU) 1976 (Lieske et al.,

1977)

Earth nutation model: IAU 1980 (Seidelmann, 1982)

Earth orientation model: Centre for Orbit Determination in Europe (CODE) final Earth

Rotation Parameters (ERP)

2.2 Astronomical Institute of the University of Bern

Software: Bernese v5.3 (Dach et al., 2015; Jäggi, Hugentobler and Beut-

ler, 2006)

Differencing Scheme:
Linear combination:
GPS observations:
Estimator:
Arc length:
Data weighting:
Undifferenced
Ionosphere-free
Carrier phase
Batch LS
24 hours
N/A

Transmitter PCV: Official IGS08 ANTEX up to day 17/028, official IGS14 ANTEX

from day 17/029 on

Receiver PCV: Stacking of residuals from reduced-dynamic Precise Orbit

Determination (POD) of approx. 120 days, 9 iterations, 1°

binning

Data screening: 2 cm/s or larger time-differences of the geometry-free linear

combination of L1B GPS carrier phase observations

Earth precession model: International Earth Rotation Service (IERS) 2010 Conventions

(Petit and Luzum, 2010)

Earth nutation model: IERS 2010 Conventions (Petit and Luzum, 2010)

Earth orientation model: CODE final ERP

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2.3 Institute of Geodesy Graz

Software: Gravity Recovery Object Oriented Programming System (GROOPS)

Differencing Scheme: None

Linear combination: None (the ionospheric influence is co-estimated)

GPS observations: Code and carrier phase

Estimator: LS **Arc length:** 24 hours

Data weighting: Elevation and azimuth-dependent, epoch-wise Variance Com-

ponent Estimation (VCE)

Transmitter PCV: Empirical, estimated from 5.5 years of data, including data

from several Low-Earth Orbit (LEO) missions (Gravity Recovery And Climate Experiment (GRACE), Jason 2 & 3, MetOp-A & -B, Sentinel 3A, Swarm, TanDEM-X, TerraSAR-X) (Zehent-

ner, 2016)

Receiver PCV: Empirical, spherical harmonics (maximum D/O 60), derived

from 38 months of data

Data screening: Implicit in VCE

Earth precession model: IAU 2006/2000A precession-nutation model (Petit and Lu-

zum, 2010)

Earth nutation model: IAU 2006/2000A precession-nutation model (Petit and Lu-

zum, 2010)

Earth orientation model: IERS Earth Orientation Parameter (EOP) 08 C04 (Petit and

Luzum, 2010)

2.4 Common

Carrier phase ambiguities: Float

Receiver clock corrections: Co-estimated

Sampling rate: 10 or 1 seconds (depending on L1B GPS data)

Elevation cut-off angle: 0°

GPS orbits and clocks: Final orbits and 5 seconds clocks of Centre for Orbit Determ-

ination in Europe (CODE) (Dach et al. 2017)

Swarm attitude: L1B attitude data

2.5 Summary

Institut	e Software	Reference
AIUB	Bernese v5.3 (Dach et al., 2015; Jäggi, Hugentobler and Beutler, 2006)	Jäggi et al. (2016)
IfG	Gravity Recovery Object Oriented Programming	Zehentner and
110	System (GROOPS) (in-house development)	Mayer-Gürr (2016)
	GPS High precision Orbit determination Software	
TUD	Tool (GHOST) (Helleputte 2004; Wermuth,	IJssel et al. (2015)
	Montenbruck and Helleputte 2010)	

Table 1 - Overview of the Kinematic Orbits and the software packages used to estimate them

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3 Kinematic Baselines

3.1 Delft University of Technology

Software: Multiple satellites Orbit Determination using Kalman filter-

ing (MODK) (Barneveld 2012)

Linear combination: N/A (the ionospheric influence is modelled) **Estimator:** Iterative Extended Kalman Filter (EKF)

Carrier phase ambiguities: Integer, using the Least-squares Ambiguity De-correlation

Adjustment (LAMBDA) (Teunissen 1995) method

Receiver PCV: Empirical Phase Center Variations (PCVs) and Code Residual

Variations (CRVs) maps are estimated a priori for each GPS

frequency

3.2 Astronomical Institute of the University of Bern

Software: Bernese v5.3 (Dach et al., 2015; Jäggi et al., 2007)

Linear combination: Ionosphere-free

Estimator: LS

Carrier phase ambiguities: wide-lane and narrow-lane integer bootstrapping with the

Melbourne-Wübbena and the ionosphere-free linear com-

bination, respectively

Receiver PCV: Empirical

3.3 Common

Differencing Scheme: Double-differenced **GPS observations:** Code and carrier phase

Carrier phase ambiguities: Integer

3.4 Summary

Institute	e Software	Reference
AIUB	Bernese v5.3 (Dach et al., 2015; Jäggi et al., 2007)	Allende-Alba et al. (2017)
TUD	Multiple satellites Orbit Determination using Kalman filtering (MODK) (Barneveld 2012)	Mao et al. (2017)

Table 2 – Overview of the Kinematic Baselines and the software packages used to estimate them

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4 Gravity Field Models

4.1 Astronomical Institute of the University of Bern

Software: Bernese v5.3 (Dach et al. 2015)

Approach: Celestial Mechanics Approach (CMA) (Beutler et al. 2010) **Reference GFM:** AIUB GRACE-only static model, version 3 (AIUB-GRACE03S)

(Jäggi et al. 2011)

Empirical Parameters: Daily piecewise-constant, 15 minutes piecewise-constant

(constrained)

Drag Model: None **EARP and EIRP Models:** None

Non-tidal Model: Atmosphere and Ocean De-aliasing Level 1B (AOD1B) product

(Flechtner, Schmidt and Meyer 2006; Flechtner 2007; Flecht-

ner 2011)

Ocean Tidal Model: 2011 Empirical Ocean Tide model (EOT11a) (Savcenko and

Bosch 2012)

Permanent Tide System: tide-free

4.2 Astronomical Institute Ondřejov

Software: (developed in-house)

Approach: Decorrelated Acceleration Approach (DAA) (Bezděk et al. 2014;

Bezděk et al. 2016)

Reference GFM: ITG GRACE-only static model, 2010 (ITG-GRACE2010s) (Mayer-

guerr et al. 2010)

Empirical Parameters: Daily constant-piecewise

Drag Model: (US) Naval Research Laboratory Mass Spectrometer and In-

coherent Scatter Radar Atmospheric model (NRLMSISE) (Pi-

cone et al. 2002)

EARP and EIRP Models: Knocke, Ries and Tapley (1988)

Non-tidal Model: Atmosphere and Ocean De-aliasing Level 1B (AOD1B) product

(Flechtner, Schmidt and Meyer 2006; Flechtner 2007; Flecht-

ner 2011)

Ocean Tidal Model: 2004 Finite Element Solution (FES2004) global tide model

(Lyard et al. 2006)

Permanent Tide System: tide-free

4.3 Institute of Geodesy Graz

Software: GROOPS

Approach: Short-Arcs Approach (SAA) (Mayer-Gürr 2006)

Reference GFM: GOCO release 05 satellite-only gravity field model (GOCO05S)

(Mayer-Gürr 2015)

Empirical Parameters: Piecewise linear for each arc (ranging from 15 to 45 minutes)

Drag Model: Jacchia-Bowman 2008 (JB2008) (Bowman et al. 2008)

EARP and EIRP Models: Rodriguez-Solano et al. (2012)

Non-tidal Model: Atmosphere and Ocean De-aliasing Level 1B RL06 (AOD1B-

RL06) product (Dobslaw et al. 2017)

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Ocean Tidal Model: 2014 Finite Element Solution (FES2014) global tide model

(Carrere et al. 2015)

Permanent Tide System: zero tide

4.4 Ohio State University

Software: (developed in-house)

Approach: Improved Energy Balance Approach (IEBA) (Shang et al. 2015) **Reference GFM:** GRACE Intermediate Field 48 (GIF48) (Ries et al. 2011) up to

Degree and Order (D/O) 200

Empirical Parameters: 2nd order polynomial every 3 hours, 1-Cycle Per Revolution

(CPR) sinusoidal every 24 hours

Regularization: none

Drag Model: (US) Naval Research Laboratory Mass Spectrometer and In-

coherent Scatter Radar Atmospheric model (NRLMSISE) (Pi-

cone et al. 2002)

EARP and EIRP Models: Knocke, Ries and Tapley (1988)

Non-tidal Model: Atmosphere and Ocean De-aliasing Level 1B (AOD1B) product

(Flechtner, Schmidt and Meyer 2006; Flechtner 2007; Flecht-

ner 2011)

Ocean Tidal Model: 2011 Empirical Ocean Tide model (EOT11a) (Savcenko and

Bosch 2012)

Permanent Tide System: tide-free

4.5 Common

Atmospheric Tidal Model: Biancale and Bode (2006)

Solid Earth Tidal Model: IERS2010 **Pole Tidal Model:** IERS2010 **Ocean Pole Tidal Model:** IERS2010

Third body perturbations: Sun, Moon, Mercury, Venus, Mars, Jupiter and Saturn, follow-

ing the JPL Planetary and Lunar Ephemerides (Folkner et al.

2014)

 $C_{2,0}$ **coefficient:** estimated alongside other coefficients

4.6 Summary

Inst.	Approach	Reference
AIUB	Celestial Mechanics Approach (CMA) (Beutler et al. 2010)	Jäggi et al. (2016)
ASU	Decorrelated Acceleration Approach (DAA) (Bezděk et al. 2014; Bezděk et al. 2016)	Bezděk et al. (2016)
IfG	Short-Arcs Approach (SAA) (Mayer-Gürr 2006)	Zehentner and Mayer-Gürr (2016)
OSU	Improved Energy Balance Approach (IEBA) (Shang et al. 2015)	Guo et al. (2015)

Table 3 – Overview of the gravity field estimation approaches

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5 File name conventions

5.1 Documentation

The file names of the documents shall be compliant with ESA requirements for Configuration and Document Management¹. As such, the following naming syntax applies:

$$SW_TN_{p>_GS_{n>_}< dl>_< t>.pdf}$$

The file name particles are identified as:

: issuing institute (i.e. the affiliation of the corresponding Work Package (WP) manager)

DUT: Delft University of Technology²

AIUB: Astronomical Institute of the University of Bern

ASU: Astronomical Institute Ondřejov

IFG: Institute of Geodesy Graz

OSU: Ohio State University

<n>: ever increasing four digit, zero padded number designating a unique number of this document type for this institute³

<dl>: deliverable name, i.e. either TN-01, TN-02, TN-03 or TN-04

<t>: document title

Note that there is no version number in the file names. Examples:

- SW_TN_DUT_GS_0001_TN-01_Standards_and_Background_models.pdf, as a result of WP100;
- SW_TN_ASU_GS_0001_TN-02_Pre-processing_baselines_and_accelerometer_data.pdf, as a result of WP200;
- SW_TN_DUT_GS_0002_TN-03_Swarm_models_validation.pdf, as a result of WP300;
- SW_TN_IFG_GS_0001_TN-04_Swarm_models_description.pdf, as a result of WP400.

5.2 Data

The names of data files **shared internally within the project** start with the string GSWARM and are composed of a series of *particles* (identified below between the < and > characters) connected by the underscore character (_):

The file name particles are identified as:

<dt>: data type

KO: Kinematic Orbit

¹https://smart-svn.spacecenter.dk/svn/smart/SwarmESL-All/Management/Plans/SW-PL
-DTU-GS-007_ESL_CDMP.pdf

 $^{^2}$ This particle is not in agreement with the remaining naming conventions because this is the acronym attributed to TU Delft within DISC.

 $^{^3}$ The <n> particle is used to distinguish different documents of the same type and institute, since the <dl> and <t> particles are formally optional. As a result of the obvious impracticability of (the non-optional part of) this naming convention, the code<dl> and <t> particles are highly encouraged and widely used.

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KB: Kinematic Baseline

GF: Gravity Field Model

NE: Normal Equation

AC: Modelled non-gravitational accelerations

WO: GPS data weights

<s>: satellite(s)

SA, SB, SC: single satellite

SAB, SBC, SAC: two satellites (SBA, SCB and SCA are not contemplated)

SABC: all satellites (other orders of the A, B and C characters are not contemplated)

: processor

TUD: Delft University of Technology

AIUB: Astronomical Institute of the University of Bern

ASU: Astronomical Institute Ondřejov

 $\label{eq:iff} \textbf{IFG: Institute of Geodesy Graz}$

OSU: Ohio State University

COMBINED: combined solutions from AIUB, ASU, IfG and OSU (relevant only for Gravity Field Models (GFMs) and Normal Equations (NEs))

<v>: data validity

KO, KB: yyyy-mm-dd_doy

GF, NE: yyyy-mm yyyy: four digit year

mm: two digit calendar month (zero padded)

dd: two digit calendar day (zero padded)

doy: two digit Day of Year (DoY) (zero padded)

<dv>: data version

- two digits (zero padded)
- GFMs and NEs increment the version number whenever there is a change in their processing, including new versions of KOs and/or KBs (the meaning of each version should be documented in the header and/or *readme* file)
- GF_COMBINED need to specify the solutions and respective versions in the header and/or *readme* file; the version number is incremented only because of processing or data combinations changes

<sd>: source data

KO, KB, COMBINED: empty

GF, NE: a valid

<e>: file extension

KO, KB: sp3

GF: gfc

NE: snx

AC: nrtdm

OW: wgt

Examples:

- GSWARM_KO_SA_AIUB_2016-02-25_056_03.sp3
- GSWARM_KB_SAB_TUD_2016-03-25_084_01.sp3

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- GSWARM_GF_SABC_OSU_2016-02_01_TUD.gfc
- GSWARM_NE_SABC_IFG_2016-02_01_IFG.snx
- GSWARM_GF_SABC_COMBINED_2016-02_01.gfc

6 Formats

6.1 Documentation

The deliverable documents are distributed in Portable Data Format (PDF) (with extension pdf) or MicroSoft Word format (with extension docx). Templates for MEX and Word are available in the dissemination server (see Section 7), under the directory management/Templates/. A version tracking number shall be maintained authors, under the following syntax:

```
<i>[<r>][ d<dr>]
```

- <i>: ever increasing integer number (not zero-padded);
- <r>: capital letter, initially blank and progressing alphabetically;
- <dr>: capital letter reserved for draft versions of the document, initially blank and progressing alphabetically (omitted for non-drafts).

Examples:

- 1 dA: issue 1, draft A;
- 1 dB: issue 1, draft B;
 - 1: issue 1, final version;
 - 1A: issue 1, revision A, final version;
- 20 dD: issue 2, revision O, draft D.

6.2 Data

Data compression

All data files shall be compressed **individually** using the *zip* or *gzip* compression formats (usually with file extensions zip and gz, respectively). It is the responsibility of every partner to compress/uncompress the data before/after uploading/downloading it to/from the dissemination server. The file name extension resulting data compression is omitted elsewhere in this document.

Kinematic Orbits

The KO are preferably distributed in the *SP3k* format, described below, which is a modification of the Extended Standard Product 3 Orbit Format (SP3c)⁴ format. The variance-covariance information is identified by the EPx record name.

The main innovation in the *SP3k* is the increased precision to the sub-millimetre level, by adding one additional significant digit after the comma in the kinematic positions and variances (expressed in the form of STandard Deviation (STD)). The co-variances, represented by the correlation factors, remains unchanged. This means that:

1st header line: format identifier #c of the SP3c format is replaced by #k;

⁴ftp://igs.org/pub/data/format/sp3c.txt

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P record:

Columns 1 to 4: unchanged;

Columns 5 to 46: there are 7 digits after the comma (instead of 6), at the expense of a digit before the comma, as described by the format F14.7 (instead of F14.6); these fields describe the kinematic positions;

Columns 47 to 60: unchanged;

Columns 62 to 80: unchanged and optional;

Example:

```
1 1 2 2 3 3 4 4 5 5 6
5 0 5 0 5 0 5 0 5 0 5 0 5 0
PL49 -519.6320223 -1895.3792238 -6545.3236807 -0.000884
```

EP record:

Columns 1 to 2: unchanged;

Columns 3: contains the character x to distinguish from the EP record of the SP3c format:

Columns 5 to 24: these (mandatory) fields describe the STD of the positions according to the F6.1 format;

Column 25: unused;

Columns 26 to 86: unchanged relative to columns 20 to 80 in the SP3c format; these (mandatory) fields contain the cross-correlations in the form of correlation coefficients;

Example:

Although the *SP3k* format allows for multiple satellites in one data file, the team shall exchange KO data files relevant to individual Swarm satellites (as it has been traditionally done with other formats, see below).

Since numerous KO orbit files are already available at the dissemination server, the existing alternative formats shall also to be supported by all (relevant) partners, in addition to the *SP3k* format. The original file names have been replaced by the convention described in Section 5.2. The alternative formats of the KOs, **shared internally within the project**, are specific to each institute:

TU Delft:

File name extension: sigma

File header: none Data records:

Column 1-6: GPS epoch: year, month, day, hour, minute, second (fractional)

Column 7-9: position x, y and z-component [m]

Column 10: clock correction (or 't'-component, already applied to columns 1 to 6) [ms]

Column 11-20: xx, yy, zz, tt (clock-correction variance), xy, xz ,xt, yz, yt and zt-element of 4x4 epoch-wise covariance matrix $[m^2]^5$

AIUB:

 $^{^5}$ The (variance) covariance terms involving the clock correction (represented by 't') are also in units of m²; the conversion to units of time can be done by dividing by the (square of the) speed of light.

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File name extension: KIN

File header:

Line 1: description of orbit including release number and orbit generation date and time

Line 2: dummy line

Line 3: information about geodetic datum and date & time of first epoch (GPS time)

Line 4: a posteriori STD of L1/L2 GPS carrier phase obs. [m]

Line 5: description of columns

Line 6: dummy line

Data records: After the header each line represents one epoch. The columns contain the following quantities:

Column 1: internal AIUB satellite name (Swarm-A: SWMA, Swarm-B: SWMB, Swarm-C: SWMC)

Column 2: modified SVN number (Swarm-A: L47, Swarm-B: L48, Swarm-C: L49)

Column 3: GPS week

Column 4: GPS second within the given GPS week [s]

Column 5-7: position x, y, and z-componentz [m]

Column 8: quality flag (possible values are K, X, S and G, see below)

Column 9-14: xx, yy, zz, xy, xz, yz-elements of 3x3 epoch-wise cofactor matrix

Meaning of quality flags:

K: KO position could be properly determined

X: KO position could not be determined (less than 4 satellites)

S: KO position could be determined, but less than 5 satellites were available

G: KO position has been flagged during internal AIUB orbit screening

IfG:

File name extension: txt

File header:

Line 1: short description of the orbit product

Line 2: geodetic datum

Data records: The epochs are in the GPS time system.

Column 1: Modified Julian Date (MJD) in the GPS time system

Column 2-4: position x, y, and z-componentz [m]

Column 5-10: xx, yy, zz, xy, xz, yz-elements of 3x3 epoch-wise covariance matrix

Kinematic Baselines

The KB data are distributed in the *SP3k* format, with the following modification:

EB **record:** a new record containing the variance-covariance information of the KBs, following the same format as the EPx record (which describes variance-covariance information for the positions);

P **records:** describe the orbits of the **two** satellites (with 7 significant digits, under the *SP3k* format) from which it is possible to reconstruct the estimated KBs.

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Gravity Field Model

The GFMs data are distributed in the International Centre for Global Earth Models (ICGEM) format⁶.

Normal Equation

The NEs data are distributed in the Solution-Independent Exchange (SINEX) format⁷.

Modelled non-gravitational accelerations

The modelled non-gravitational accelerations are distributed **internally within the project** in the *nrtdm* format (since this format is readily available in TU Delft's Near Real-Time Density Model (NRTDM) software). It is a column-wise, plain text, self-explanatory format. The header is optional but useful, since it describes what each column contains. Maintaining the yyyy-mm-dd and hh:mm:ss.ss date and time formats is important, since otherwise it would make parsing the header mandatory. Additional columns may be appended, with their description appended after line 6 of the header.

File name extension: nrtdm

File header: (optional but highly encouraged)

Lines 1-6: Description of the contents of columns 1 to 6 (see below)

Line 7: Fortran format string **Line 8:** Column headers

Line 9: Column-wise units of the data (columns 4 to 6)

Data records: After the header each line represents one epoch. The columns contain the following quantities:

Column 1: calendar date in the yyyy-mm-dd format

Column 2: time in the hh:mm:ss.sss format

Column 3: time system, i.e. GPS, UTC, UT1 or TAI

 $\begin{tabular}{ll} \textbf{Column 4-6:} & x, y and z coordinates of the non-gravitational acceleration in the Satellite Body Reference Frame \\ \end{tabular}$

An example of this format is:

```
# Column
         1:
                     Date (yyyy-mm-dd)
# Column
                     Time (hh:mm:ss.sss)
# Column
         3:
                     Time system
# Column 4:
             E13.6 Linear acceleration in S/C X-direction (m/s/s)
             E13.6 Linear acceleration in S/C Y-direction (m/s/s)
# Column 5:
# Column 6:
              E13.6 Linear acceleration in S/C Z-direction (m/s/s)
# Format string: (a27,1x,E13.6,1X,E13.6,1X,E13.6)
# Date/time
                            Column 4
                                          Column 5
                                                        Column 6
                            m/s/s
                                          m/s/s
                                                        m/s/s
2005-01-01 00:00:00.000 GPS
                            0.424512E-07
                                           0.479988E-08
                                                        0.103282E-07
2005-01-01 00:00:10.000 GPS 0.425367E-07
                                           0.483965E-08 0.104057E-07
2005-01-01 00:00:20.000 GPS  0.428182E-07  0.513651E-08  0.104278E-07
2005-01-01 00:00:30.000 GPS  0.433768E-07  0.467810E-08  0.101609E-07
```

⁶http://icgem.gfz-potsdam.de/ICGEM-Format-2011.pdf

 $^{^7}$ www.iers.org/IERS/EN/Organization/AnalysisCoordinator/SinexFormat/sinex.html

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2005-01-01	00:00:40.000 GPS	0.440563E-07	0.492133E-08	0.993525E-08
2005-01-01	00:00:50.000 GPS	0.446449E-07	0.543075E-08	0.971838E-08
2005-01-01	00:01:00.000 GPS	0.451432E-07	0.539917E-08	0.929534E-08

GPS data weights

The GPS data weights aim to provide an optimum weighting scheme for the GPS code and phase measurements, by indicating the a-priori STD for specific GPS satellites at specific epochs. The data is formatted in column-wise, plain ASCII, **and is shared internally within the project**.

File name extension: wgt

File header:

Line 1: Description of the data and date and time it was exported

Line 2: Header separator, consisting of several – characters

Line 3: Column description

Line 4: Place-holder for the digits in the columns, using the * character

Data records: After the header each line represents one epoch. The columns contain the following quantities:

Column 1: initial epoch, in MJD

Column 2: final epoch, optionally empty indicating one single epoch is affected

Column 3: GPS satellite number

Column 4: a priori STD to assign to the corresponding data

An example of this format is:

Observation-specific weights

18-Jan-2018 15:51

Start MJD	End MJD	SAT	SIGMA
*****	*****	**	***.***
57023.001122700000		04	21.000
57023.001134300001		04	21.000
57023.001145800001		04	21.000
57023.001157400002		04	21.000
57023.001169000003		04	21.000
57023.001180599997		04	21.000

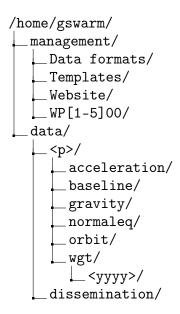
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7 Dissemination Server

The server is located at aristarchos.lr.tudelft.nl and can only be accessed using secure shell client (and related utilities). The access to this server requires dedicated credentials, usually issued in the name of the institute team leaders (TU Delft team members have individual access). Additional user credentials can be issued if needed. All users (except TU Delft team members) share the same \$HOME directory, /home/gswarm.

7.1 Directory structure

The directory structure of the distribution server is:



The contents of the directories above is:

```
Data formats: documentation on the data formats mentioned in Section 6
```

Proposal: the files used in producing the answer of the team to the Swarm DISC ITT

1.1

Templates: MTEXand Microsoft Word templates for reports

Website: github repository with the contents of https://jgte.github.io/gsw

arm/

WP [1-5] 00: reserved for the activities of the corresponding WP

: data distributed by each processor: aiub, asu, ifg, osu or tudelft

acceleration: Modelled non-gravitational acceleration data

baseline: KB data

gravity: GFM coefficients

normaleq: NE data

orbit: KO data distributed in yearly directories, with names <yyyy> (four digit

year)

dissemination: data to be disseminated to ESA

7.2 Data dissemination

The dissemination directory contains the GFMs that are uploaded to the ESA and ICGEM dissemination servers. These data are packaged according to the Swarm Level 2 (L2) Product

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Format⁸ and following all required naming conventions.

Acronyms

AA Acceleration Approach, Rummel (1979)

AIUB Astronomical Institute of the University of Bern, Switzerland,

www.aiub.unibe.ch

ASCII American Standard Code for Information Interchange

ASU Astronomical Institute (Astronomický ústav), AVCR, Ondřejov,

www.asu.cas.cz/en

AVCR Czech Academy of Sciences (Akademie věd České Republiky), Czech Republic,

www.avcr.cz/en/

CODE Centre for Orbit Determination in Europe, Dach et al. (2017)

CPR Cycle Per Revolution

CRV Code Residual Variation

D/O Degree and Order

DISC Data, Innovation and Science Cluster

DoY Day of Year

EARP Earth Albedo Radiation Pressure

EGSIEM European Gravity Service for Improved Emergency Management, EU Horizon

2020, www.egsiem.eu

EIRP Earth Infrared Radiation Pressure

EKF Extended Kalman Filter

EBA Energy Balance Approach, O'Keefe (1957) and Jekeli (1999)

EOP Earth Orientation Parameter
ERP Earth Rotation Parameters

ESA European Space Agency, www.esa.int

EU European Union

FES Finite Element Solution global tide model

GFM Gravity Field Model

GOCO Gravity Observation COmbination

GPS Global Positioning System

GRACE Gravity Recovery And Climate Experiment, Tapley, Reigher and Melbourne

(1996) and Tapley (2004)

GRACE-FO GRACE Follow On, Sheard et al. (2012), Larkin (2012) and Flechtner et al. (2014)

GROOPS Gravity Recovery Object Oriented Programming System

IAU International Astronomical Union

 $^{^8} https://earth.esa.int/web/guest/missions/esa-eo-missions/swarm/data-handbook/level-2-product-format$

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ICGEM International Centre for Global Earth Models, icgem.gfz-potsdam.de

IERS International Earth Rotation Service

IERS Conventions 2010, Petit and Luzum (2010)

Institute of Geodesy, TUG, Graz, www.ifg.tugraz.at

ITG Institut für Geodäsie und Geoinformation, Germany, www.igg.uni-bonn.de

ITT Invitation To Tenders

JPL Jet Propulsion Laboratory, USA, www.jpl.nasa.gov

JPL Planetary and Lunar Ephemerides, Folkner et al. (2014)

KB Kinematic Baseline
KO Kinematic Orbit
L1B Level 1B data
L2 Level 2 data
LEO Low-Earth Orbit
LS Least-Squares

MJD Modified Julian Date

N/A Not ApplicableNE Normal Equation

NRTDM Near Real-Time Density Model

OSU Ohio State University, www.osu.edu

PCV Phase Center Variation
PDF Portable Data Format,

en.wikipedia.org/wiki/Portable_Document_Format

POD Precise Orbit Determination

PSO Precise or Post-processed Science Orbit

RL06 Release 6

SINEX Solution-Independent Exchange Format,

www.iers.org/IERS/EN/Organization/AnalysisCoordinator/SinexFormat/sinex.htm

SoW Statement of Work, Doc. Ref. SW-SW-DTU-GS-111_ITT1-1

SP3c Extended Standard Product 3 Orbit Format,

ftp://igs.org/pub/data/format/sp3c.txt

SNR Signal-to-Noise Ratio
STD STandard Deviation

TU Delft University of Technology, Netherlands, www.tudelft.nl

TUG Graz University of Technology, Austria, www.tugraz.at

USA United States of America

VCE Variance Component Estimation

WP Work Package

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References

- Allende-Alba, Gerardo et al. (2017). **Reduced-dynamic and kinematic baseline determination for the Swarm mission**. In: *GPS Solut.* 21.3, pp. 1275–1284. DOI: 10.1007/s10291-017-0611-z (cit. on p. 8).
- Barneveld, Pieter Willem Lucas van (2012). **Orbit determination of satellite formations**. PhD thesis. Delft University of Technology. DOI: 10.4233/uuid:c5ac8599-fca2-40eb-adc6-bbfeeec38fa (cit. on p. 8).
- Beutler, Gerhard et al. (2010). **The celestial mechanics approach: theoretical foundations**. In: *J. Geod.* 84.10, pp. 605–624. DOI: 10.1007/s00190-010-0401-7 (cit. on pp. 9, 10).
- Bezděk, Aleš et al. (2014). **Gravity field models from kinematic orbits of CHAMP, GRACE and GOCE satellites**. In: *Adv. Sp. Res.* 53.3, pp. 412–429. DOI: 10.1016/j.asr.2013. 11.031 (cit. on pp. 9, 10).
- Bezděk, Aleš et al. (2016). **Time-variable gravity fields derived from GPS tracking of Swarm**. In: *Geophys. J. Int.* 205.3, pp. 1665–1669. DOI: 10.1093/gji/ggw094 (cit. on pp. 5, 9, 10).
- Biancale, R. and A. Bode (2006). **Mean annual and seasonal atmospheric tide models based on 3-hourly and 6-hourly ECMWF surface pressure data**. Tech. rep. Potsdam, Germany: Deutsches GeoForschungsZentrum GFZ. DOI: 10.2312/GFZ.b103-06011 (cit. on p. 10).
- Bowman, Bruce et al. (2008). **A New Empirical Thermospheric Density Model JB2008 Using New Solar and Geomagnetic Indices**. In: *AIAA/AAS Astrodyn. Spec. Conf. Exhib.* August. Reston, Virigina: American Institute of Aeronautics and Astronautics. DOI: 10.2514/6. 2008–6438 (cit. on p. 9).
- Carrere, L et al. (2015). **FES 2014, a new tidal model on the global ocean with enhanced accuracy in shallow seas and in the Arctic region**. In: *EGU Gen. Assem. Conf. Abstr.* Vol. 17. EGU General Assembly Conference Abstracts, p. 5481 (cit. on p. 10).
- Dach, Rolf et al. (2015). **Bernese GNSS Software Version 5.2**. Bern: Bern Open Publishing. DOI: 10.7892/boris.72297 (cit. on pp. 6–9).
- Dach, Rolf et al. (2017). **CODE final product series for the IGS**. Bern, Switzerland. DOI: 10. 7892/boris.75876.2 (cit. on pp. 7, 19).
- Dobslaw, H. et al. (2017). **A new high-resolution model of non-tidal atmosphere and ocean mass variability for de-aliasing of satellite gravity observations: AOD1B RL06**. In: *Geophys. J. Int.* 211.1, pp. 263–269. DOI: 10.1093/gji/ggx302 (cit. on p. 9).
- Flechtner, Frank (2007). **Gravity Recovery and Climate Experiment AOD1B Product Description Document for product releases 01 to 04**. Technical report GR-GFZ-AOD-0001. GeoForschungszentrum Potsdam, pp. 1–33. URL: ftp://podaac.jpl.nasa.gov/pub/grace/doc/AOD1B_20070413.pdf (cit. on pp. 9, 10).
- (2011). **GRACE AOD1B RL04 Quality Assurance**. Miscellaneous. URL: http://op.gfz-potsdam.de/grace/results/grav/g007_aod1b_rl04.html (visited on 23/07/2015) (cit. on pp. 9, 10).
- Flechtner, Frank, Roland Schmidt and Ulrich Meyer (2006). **De-aliasing of Short-term Atmospheric and Oceanic Mass Variations for GRACE**. In: *Obs. Earth Syst. from Sp.* Ed. by J. Flury et al. Springer Berlin Heidelberg, pp. 83–97. DOI: 10.1007/3-540-29522-4_7 (cit. on pp. 9, 10).
- Flechtner, Frank et al. (2014). **Status of the GRACE Follow-On Mission**. In: *Int. Assoc. Geod. Symp.* Vol. 141. Springer, Cham, pp. 117–121. DOI: 10.1007/978-3-319-10837-7_15 (cit. on p. 19).

2019-04-09 Page 22 of 23

- Folkner, William M et al. (2014). **The Planetary and Lunar Ephemerides DE430 and DE431**. In: *Interplanet. Netw. Prog. Rep* 42.196. URL: https://ipnpr.jpl.nasa.gov/progress_report/42-196/196C.pdf (cit. on pp. 10, 20).
- Guo, J. Y. et al. (2015). On the energy integral formulation of gravitational potential differences from satellite-to-satellite tracking. In: *Celest. Mech. Dyn. Astron.* 121.4, pp. 415–429. DOI: 10.1007/s10569-015-9610-y (cit. on p. 10).
- Helleputte, T. van (2004). **GPS High Precision Orbit De- termination Software Tools: User Manual**. Oberpfaffenhofen (cit. on pp. 6, 7).
- IJssel, Jose van den et al. (2015). **Precise science orbits for the Swarm satellite constellation**. In: *Adv. Sp. Res.* 56.6, pp. 1042–1055. DOI: 10.1016/j.asr.2015.06.002 (cit. on p. 7).
- Jäggi, A. et al. (2007). **Precise orbit determination for GRACE using undifferenced or doubly differenced GPS data**. In: *Adv. Sp. Res.* 39.10, pp. 1612–1619. DOI: 10.1016/j.asr.2007.03.012 (cit. on p. 8).
- Jäggi, A. et al. (2011). **AIUB-GRACE03S**. Bern, Switzerland. URL: http://icgem.gfz-potsdam.de/(cit.onp.9).
- Jäggi, A. et al. (2016). **Swarm kinematic orbits and gravity fields from 18 months of GPS data**. In: *Adv. Sp. Res.* 57.1, pp. 218–233. DOI: 10.1016/j.asr.2015.10.035 (cit. on pp. 5, 7, 10).
- Jäggi, Adrian, U. Hugentobler and G. Beutler (2006). **Pseudo-Stochastic Orbit Modeling Techniques for Low-Earth Orbiters**. In: *J. Geod.* 80.1, pp. 47–60. DOI: 10.1007/s00190-006-0029-9 (cit. on pp. 6, 7).
- Jekeli, Christopher (1999). **The determination of gravitational potential differences from satellite-to-satellite tracking**. In: *Celest. Mech. Dyn. Astron.* 75.2, pp. 85–101. DOI: 10. 1023/A:1008313405488 (cit. on p. 19).
- Knocke, P., J. Ries and B. Tapley (1988). **Earth radiation pressure effects on satellites**. In: *Astrodyn. Conf.* Reston, Virigina: American Institute of Aeronautics and Astronautics. DOI: 10.2514/6.1988-4292 (cit. on pp. 9, 10).
- Larkin, Philip (2012). **ESTO**:: News:: Laser System for GRACE Follow-On. URL: http://esto.nasa.gov/news/news_gracefollowon.html (visited on 06/07/2013) (cit. on p. 19).
- Lieske, J H et al. (1977). **Expression for the precession quantities based upon the IAU (1976) system of astronomical constants**. In: *Astron. Astrophys.* 58, pp. 1–16 (cit. on p. 6).
- Lyard, Florent et al. (2006). **Modelling the global ocean tides: modern insights from FES2004**. In: *Ocean Dyn.* 56.5-6, pp. 394–415. DOI: 10.1007/s10236-006-0086-x (cit. on p. 9).
- Mao, X. et al. (2017). **Impact of GPS antenna phase center and code residual variation maps on orbit and baseline determination of GRACE**. In: *Adv. Sp. Res.* 59.12, pp. 2987–3002. DOI: 10.1016/j.asr.2017.03.019 (cit. on p. 8).
- Mayer-guerr, Torsten et al. (2010). **ITG-Grace2010: the new GRACE gravity field release computed in Bonn**. In: *EGU Gen. Assem. Abstr.* Vol. 12.EGU2010-2446, p. 2446. URL: http://www.igg.uni-bonn.de/apmg/index.php?id=itg-grace2010 (cit. on p. 9).
- Mayer-Gürr, Torsten (2006). **Gravitationsfeldbestimmung aus der Analyse kurzer Bahnbögen am Beispiel der Satellitenmissionen CHAMP und GRACE**. PhD thesis. Rheinischen Friedrich-Wilhelms Universität Bonn. URL: http://hss.ulb.uni-bonn.de/2006/0904/0904.pdf (cit. on pp. 9, 10).
- (2015). **The Combined Satellite Gravity Field Model GOCO05s**. In: *Geophys. Res. Abstr.* 17.EGU2015-12364 (cit. on p. 9).
- O'Keefe, John A. (1957). **An application of Jacobi's integral to the motion of an earth satellite**. In: *Astron. J.* 62, p. 265. DOI: 10.1086/107530 (cit. on p. 19).

- Petit, Gérard Gerard and Brian Luzum (2010). **IERS Conventions (2010)**. Frankfurt am Main. URL: http://www.iers.org/TN36/ (cit. on pp. 6, 7, 20).
- Picone, J. M. et al. (2002). **NRLMSISE-00 empirical model of the atmosphere: Statistical comparisons and scientific issues**. In: *J. Geophys. Res. Sp. Phys.* 107.A12, SIA 15–1–SIA 15–16. DOI: 10.1029/2002JA009430 (cit. on pp. 9, 10).
- Ries, John C. et al. (2011). **Mean Background Gravity Processing Mean Gravity Field from Space and Ground**. In: *GRACE Sci. Team Meet*. (Cit. on p. 10).
- Rodriguez-Solano, C. J. et al. (2012). **Impact of Earth radiation pressure on GPS position estimates**. In: *J. Geod.* 86.5, pp. 309–317. DOI: 10.1007/s00190-011-0517-4 (cit. on p. 9).
- Rummel, R. (1979). **Determination of short-wavelength components of the gravity field from satellite-to-satellite tracking or satellite gradiometry**. In: *Manuscripta Geod.* 4.2, pp. 107–148 (cit. on p. 19).
- Savcenko, R and W Bosch (2012). **EOT11a Empirical ocean tide model from multi-mission satellite altimetry**. Tech. rep. München, Germany: Deutsches Geodätisches Forschungsinstitut. URL: https://epic.awi.de/36001/1/DGFI_Report_89.pdf (cit. on pp. 9, 10).
- Schmid, Ralf et al. (2007). **Generation of a consistent absolute phase-center correction model for GPS receiver and satellite antennas**. In: *J. Geod.* 81.12, pp. 781–798. DOI: 10. 1007/s00190-007-0148-y (cit. on p. 6).
- Seidelmann, P. K. (1982). **1980 IAU Theory of Nutation: The final report of the IAU Working Group on Nutation**. In: *Celest. Mech.* 27.1, pp. 79–106. DOI: 10.1007/BF01228952 (cit. on p. 6).
- Shang, K. et al. (2017). Low-Degree Temporal Gravity Field Solution from SWARM Constellation of Satellites Using the Energy Balance Approach. In: Fourth Swarm Sci. Meet. Geod. Mission. Work. Banff, Canada (cit. on p. 5).
- Shang, Kun et al. (2015). **GRACE time-variable gravity field recovery using an improved energy balance approach**. In: *Geophys. J. Int.* 203.3, pp. 1773–1786. DOI: 10.1093/gji/ggv392 (cit. on p. 10).
- Sheard, B. S. et al. (2012). **Intersatellite laser ranging instrument for the GRACE follow-on mission**. In: *J. Geod.* 86.12, pp. 1083–1095. DOI: 10.1007/s00190-012-0566-3 (cit. on p. 19).
- Tapley, B., C. Reigber and W Melbourne (1996). **Gravity Recovery And Climate Experiment** (**GRACE**) **mission**. Baltimore, USA (cit. on p. 19).
- Tapley, Byron D. (2004). **GRACE Measurements of Mass Variability in the Earth System**. In: *Science* (80-.). 305.5683, pp. 503–505. DOI: 10.1126/science.1099192 (cit. on p. 19).
- Teunissen, P. J. G. (1995). **The least-squares ambiguity decorrelation adjustment: a method for fast GPS integer ambiguity estimation**. In: *J. Geod.* 70.1-2, pp. 65–82. DOI: 10.1007/BF00863419 (cit. on p. 8).
- Wermuth, Martin, Oliver Montenbruck and Tom Van Helleputte (2010). **GPS high precision orbit determination software tools (GHOST)**. In: *4th Int. Conf. Astrodyn. Tools Tech.* Madrid: ESA WPP-308 (cit. on pp. 6, 7).
- Zehentner, Norbert (2016). **Kinematic orbit positioning applying the raw observation approach to observe time variable gravity**. Doctoral Dissertation. Graz University of Technology, p. 175. DOI: 10.13140/RG.2.2.33916.33927 (cit. on p. 7).
- Zehentner, Norbert and Torsten Mayer-Gürr (2016). **Precise orbit determination based on raw GPS measurements**. In: *J. Geod.* 90.3, pp. 275–286. DOI: 10.1007/s00190-015-0872-7 (cit. on pp. 5, 7, 10).