

ASTR 405

Planetary Systems

Solar System

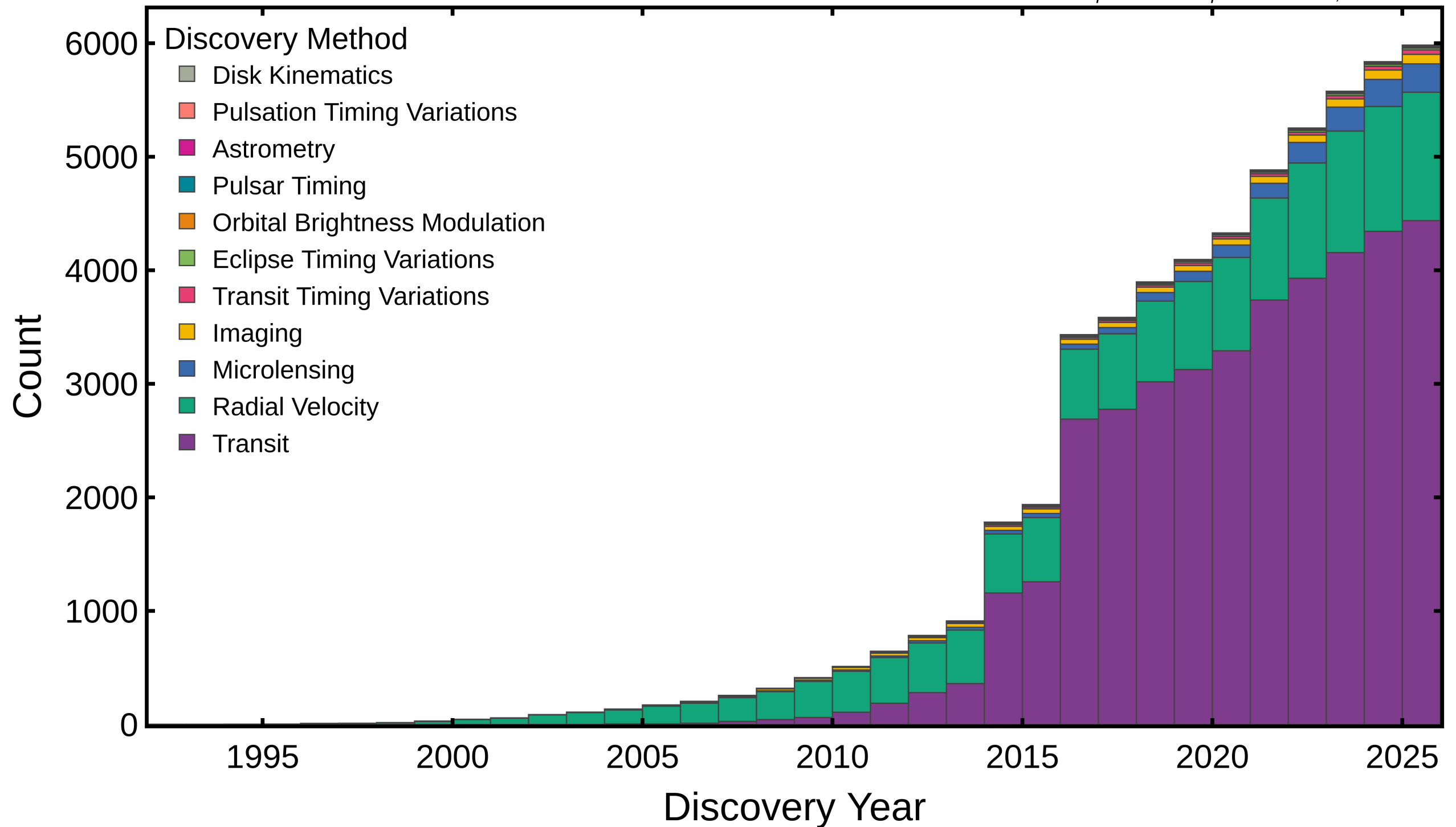
Fall 2025

Prof. Jiayin Dong

Cumulative Exoplanet Discoveries

Cumulative Counts vs Discovery Year

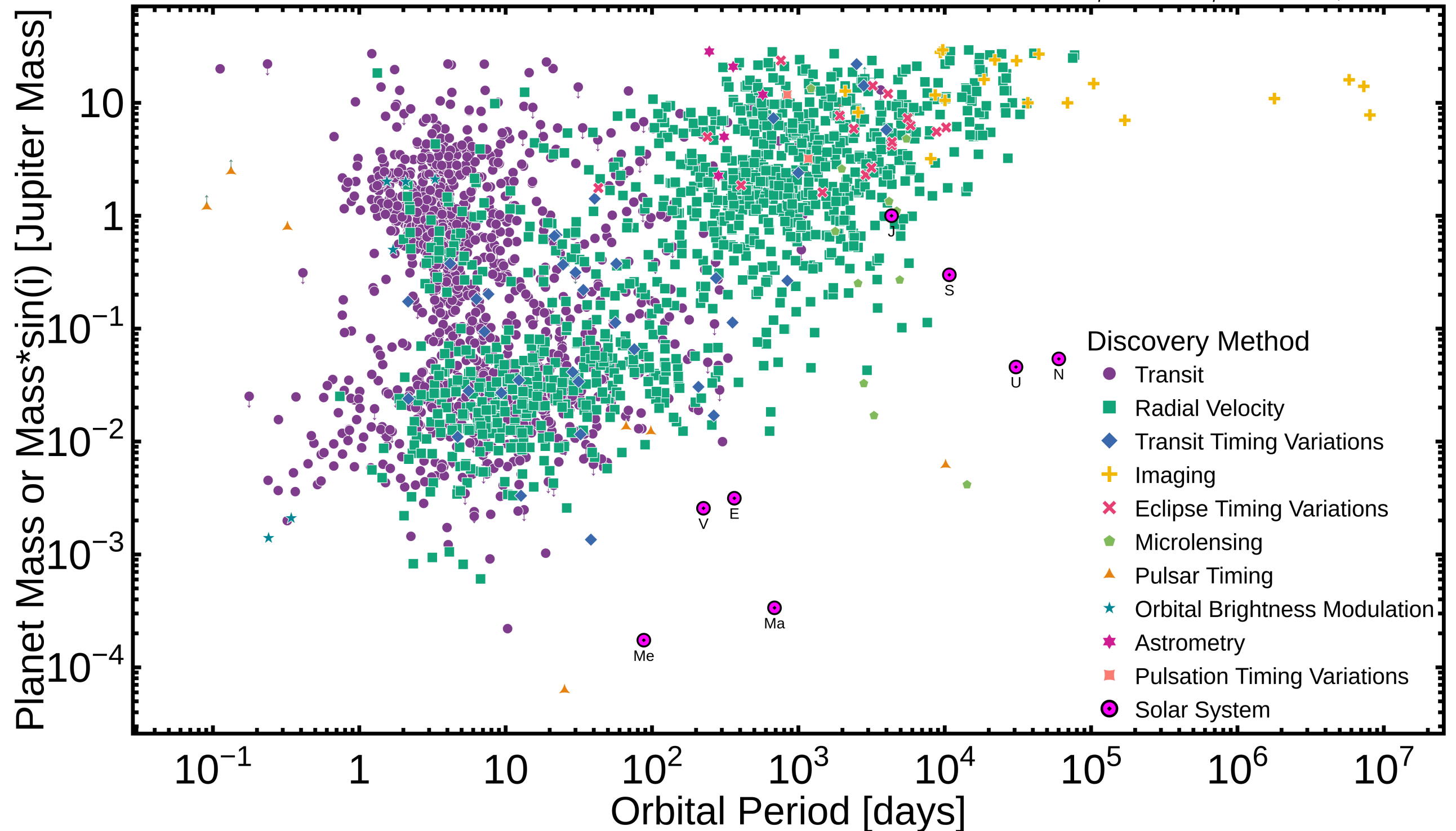
exoplanetarchive.ipac.caltech.edu, 2025-08-14



Exoplanet Mass–Period Distribution

Planet Mass or Mass* $\sin(i)$ vs Orbital Period

exoplanetarchive.ipac.caltech.edu, 2025-08-14



Scales

Jupiter in Solar Units

$$R_{\text{Jup}} \approx 0.1 R_{\odot}$$

$$M_{\text{Jup}} \approx 10^{-3} M_{\odot}$$

$$a_{\text{Jup}} \approx 5 \text{ au}$$

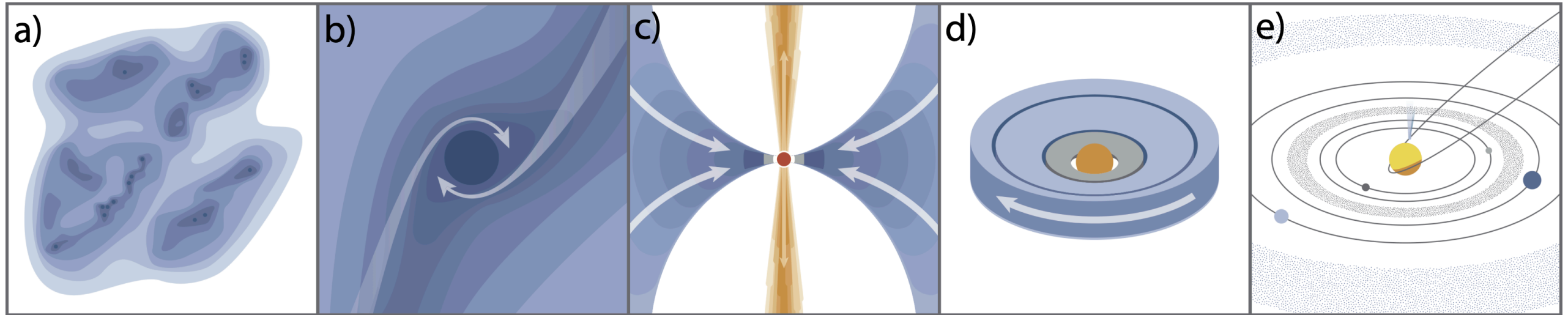
Earth in Jupiter Units

$$R_{\oplus} \approx 0.1 R_J$$

$$M_{\oplus} \approx 3 \times 10^{-3} M_J$$

$$a_{\oplus} = 1 \text{ au}$$

Formation of the Solar System



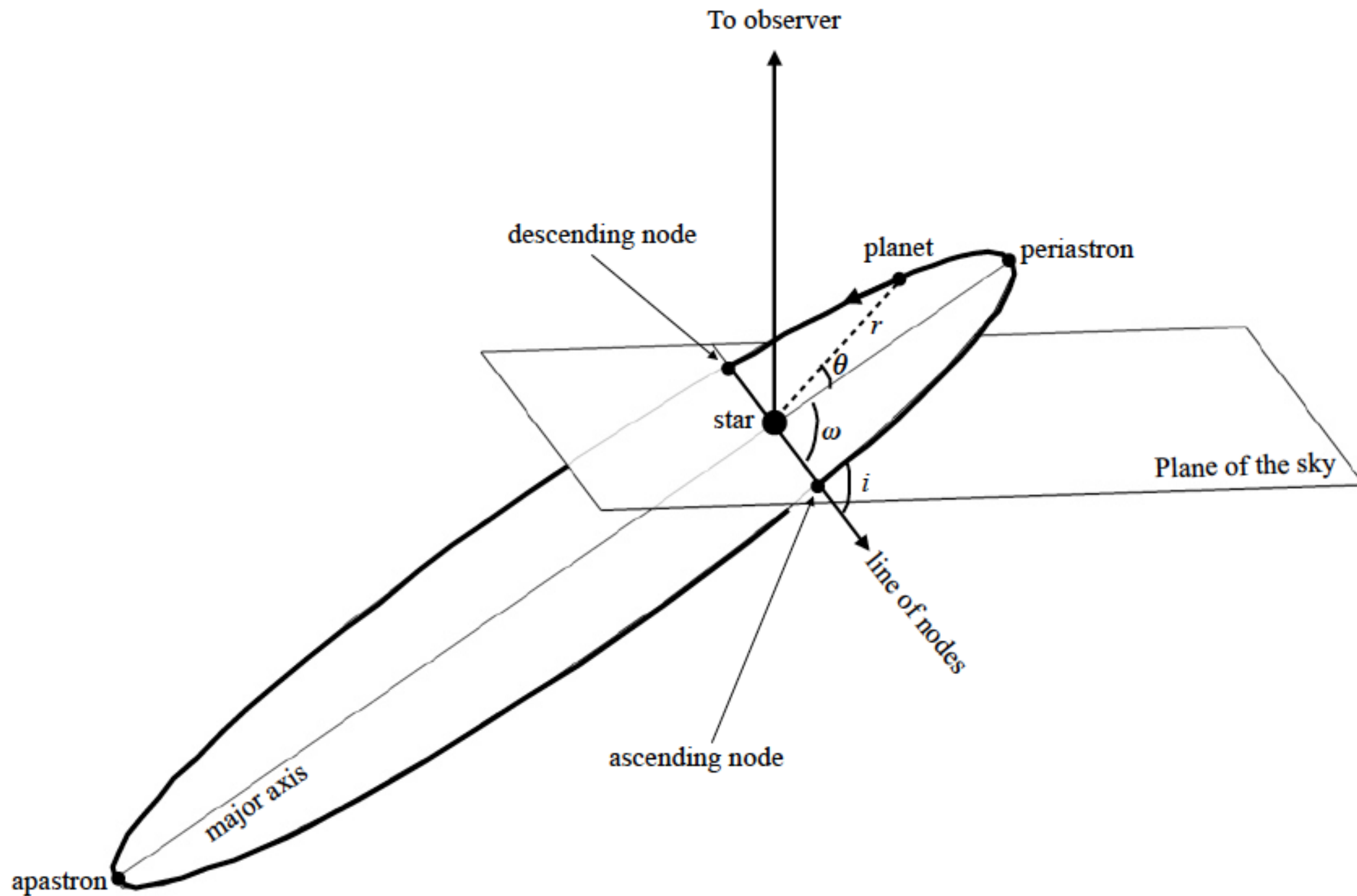
Oberg & Bergin 2021, <https://arxiv.org/pdf/2010.03529>

- a) Stars form in dense cores in interstellar molecular clouds.
- b) Star formation begins when such a dense core begins to collapse due to self-gravity.
- c) As the collapse proceeds the center heats up forming a protostar. Accretion of remnant cloud material continues, funneled through a disk, which is formed as a consequence of cloud angular momentum. This stage is also characterized by outflows of material.
- d) Following dispersal of the cloud remnant the now pre-main sequence star is left with a circumstellar disk, which is the formation sites of planets. The disk gas is dispersed through disk winds within $\sim 2\text{-}5$ Myrs, putting a halt to Gas Giant formation.
- e) Rocky and icy planets can continue to grow for another 100 Myrs at which time a mature planetary system exists.

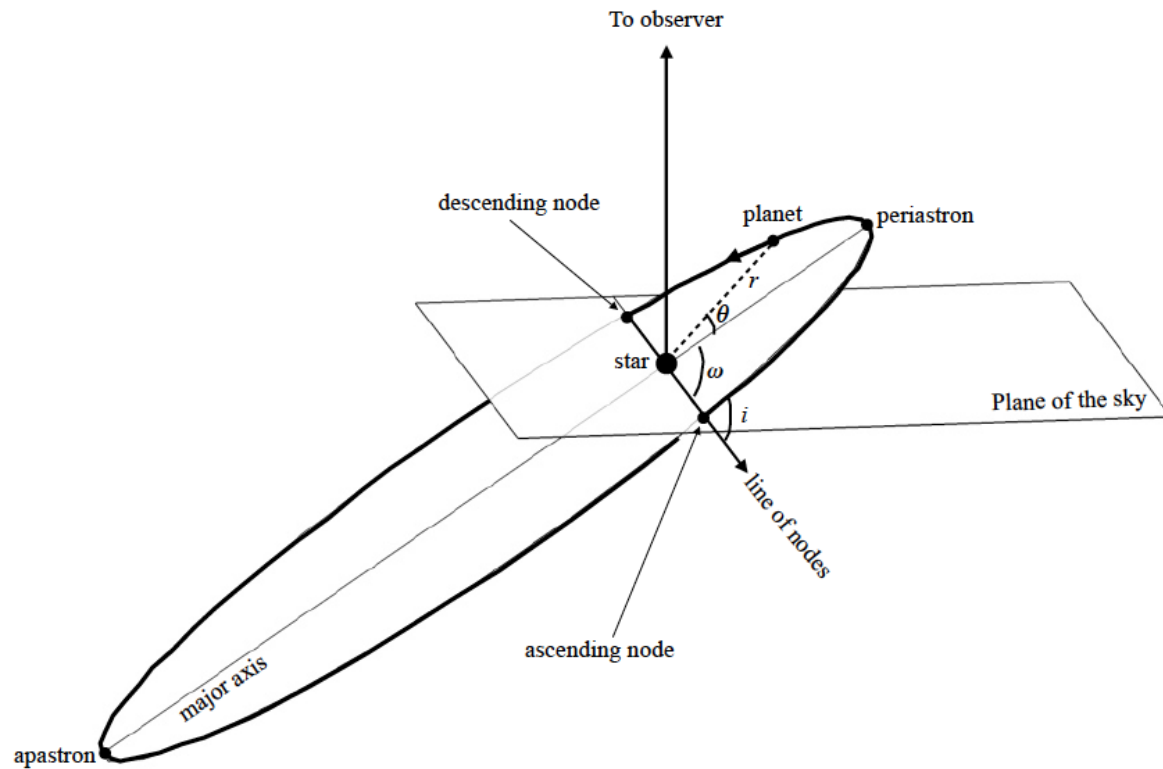
In-Class Activity

Angular momentum in the Solar System

Geometry of a 3D Orbit



Geometry of a 3D Orbit



- Plane of the sky: The plane oriented perpendicular to the observer's line of sight, which passes directly through the star being orbited.
- Plane of the orbit: The two-dimensional plane in which the planet and star orbit.
- Inclination angle i : The angle between the plane of the sky and the plane of the orbit. It is 90 degrees for edge-on orbits and 0 degrees for face-on orbits.
- Periastron: The location of closest approach between the star and planet, similar to perihelion in Sun-centered orbits.
- Apastron: The location of greatest separation between the star and planet, similar to aphelion in Sun-centered orbits.
- Ascending and descending nodes: The two locations where the planet's orbit crosses the plane of the sky. The ascending node is where the planet moves from below to above the plane of the sky; the descending node is where it moves from above to below.
- Line of nodes: The line connecting the ascending and descending nodes.
- Argument of periastron ω : The angle between the line of nodes and periastron. This defines how the orbit's ellipse is oriented relative to the observer.
- True anomaly (θ as a function of time): The angle between periastron and the planet's current location in its orbit. This angle changes over time.

Kepler's Three Laws

- **First Law:** All planets move in elliptical orbits with the Sun

at one focus. $r = \frac{a(1 - e^2)}{1 + e \cos \theta(t)}$

- **Second Law:** Equal areas swept in equal time (planets

move faster when closer). $\frac{dA}{dt} = \frac{r^2}{2} \frac{d\theta}{dt} = \text{constant}$

- **Third Law:** Square of period \propto Cube of semi-major axis

$$T^2 \propto a^3$$