

ASTR 405
Planetary Systems
Astrometry

Fall 2025
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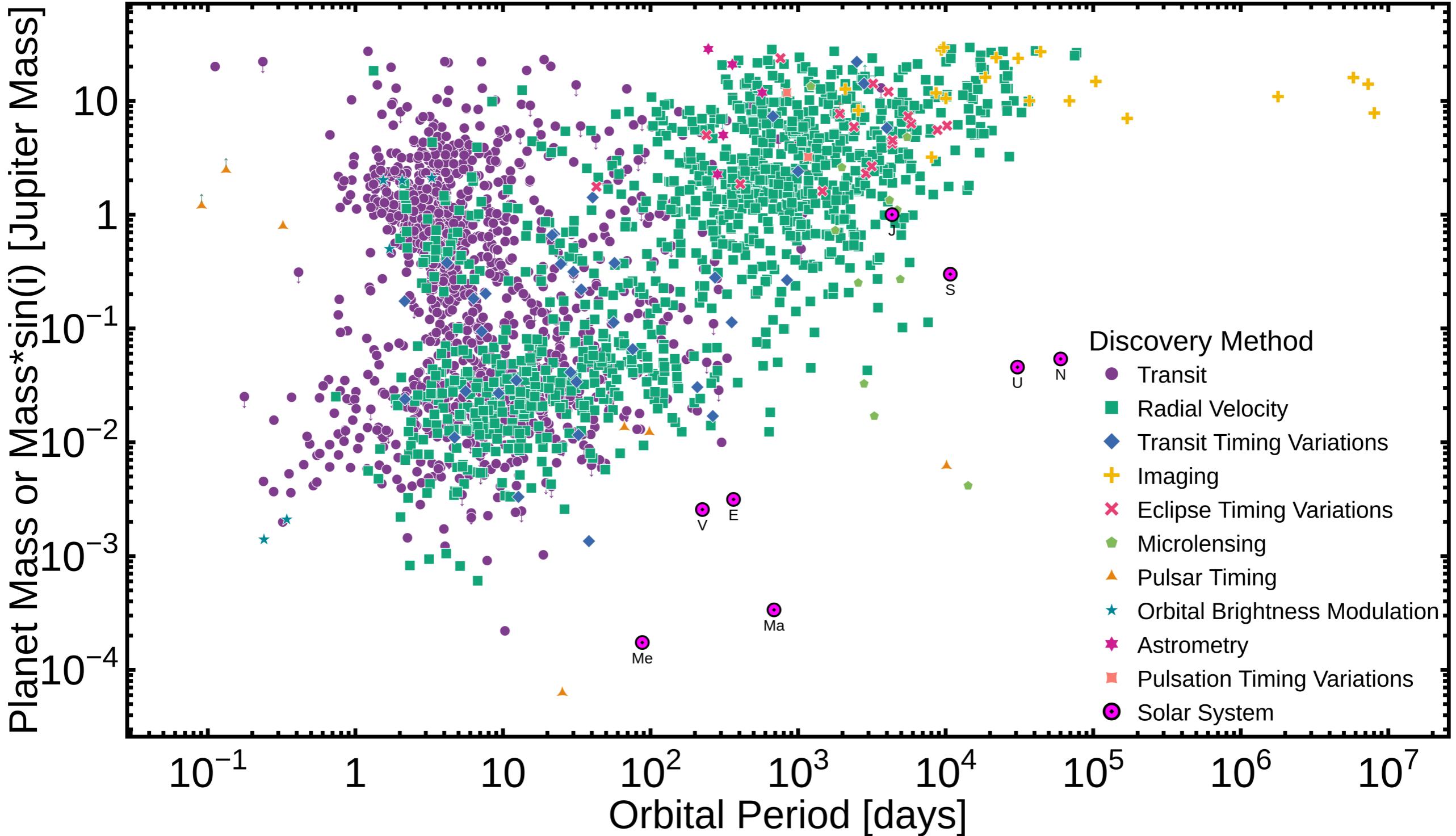
Module I: Exoplanet Detection Methods

- Radial Velocity: detecting exoplanets by measuring Doppler shifts from a star's **radial** reflex motion along our line of sight
- Astrometry: detecting exoplanets by measuring tiny changes in a star's sky position from its **tangential** reflex motion.
- Transit
- Microlensing
- Direct Imaging

Exoplanet Mass–Period Distribution

Planet Mass or Mass $\cdot\sin(i)$ vs Orbital Period

exoplanetarchive.ipac.caltech.edu, 2025-08-14



Exoplanet Detection via Astrometry

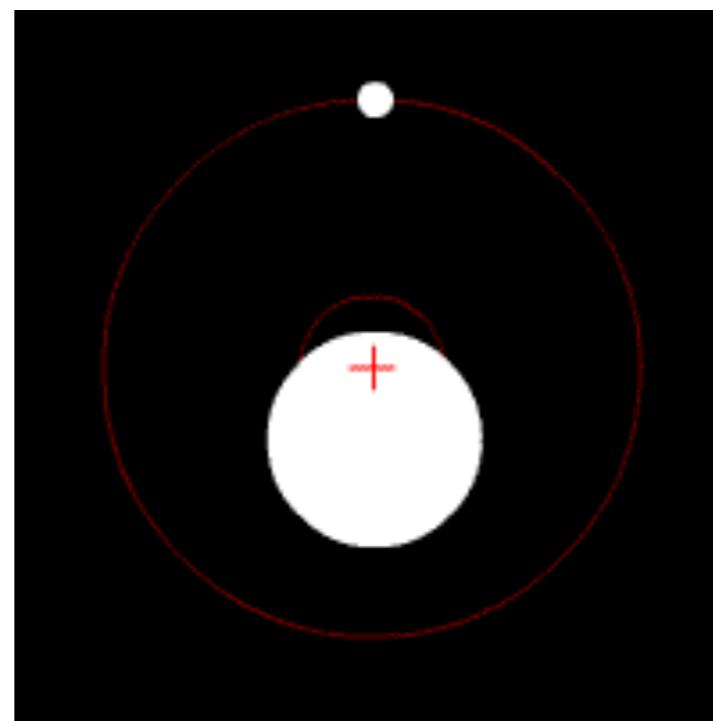


Diagram showing how an exoplanet orbiting a star could produce changes in position and velocity of the star as they orbit their common center of mass (red cross). *Wikipedia*

Proper Motion

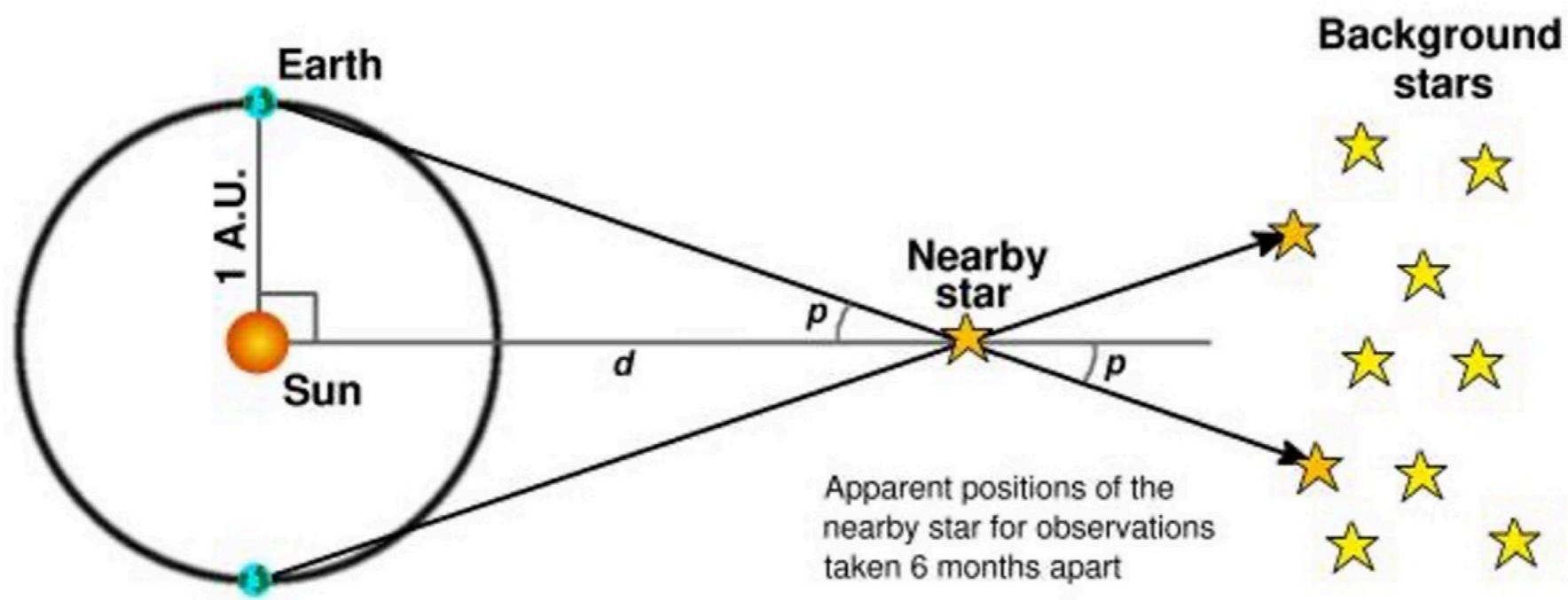
- Proper motion: $\mu \equiv \frac{d\theta}{dt} = \frac{v_\theta}{d}$
- The apparent shift in angular location (θ) of a star as it moves across the sky due to the star's tangential velocity (v_θ)
- Tangential velocity (v_θ) is orthogonal to radial velocity (v_r)



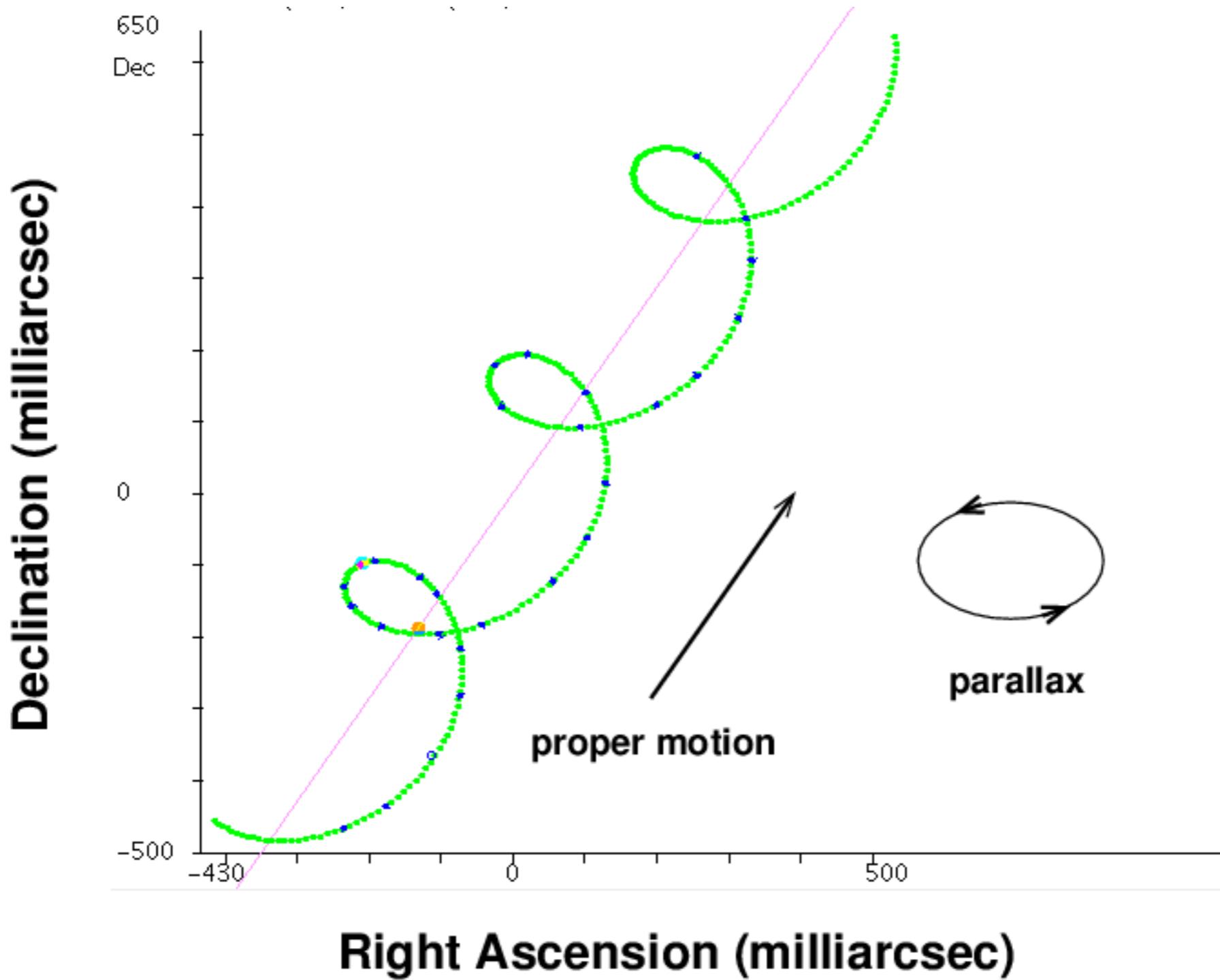
Barnard's star has a high proper motion of 10.3 arcsec/yr.

Parallax

- Parallax: $p \text{ (arcsec)} = \frac{1}{d \text{ (pc)}}$
- Apparent shift in position of a star on sky (relative to distant background stars) due to the motion of Earth around the Sun



Proper Motion + Parallax

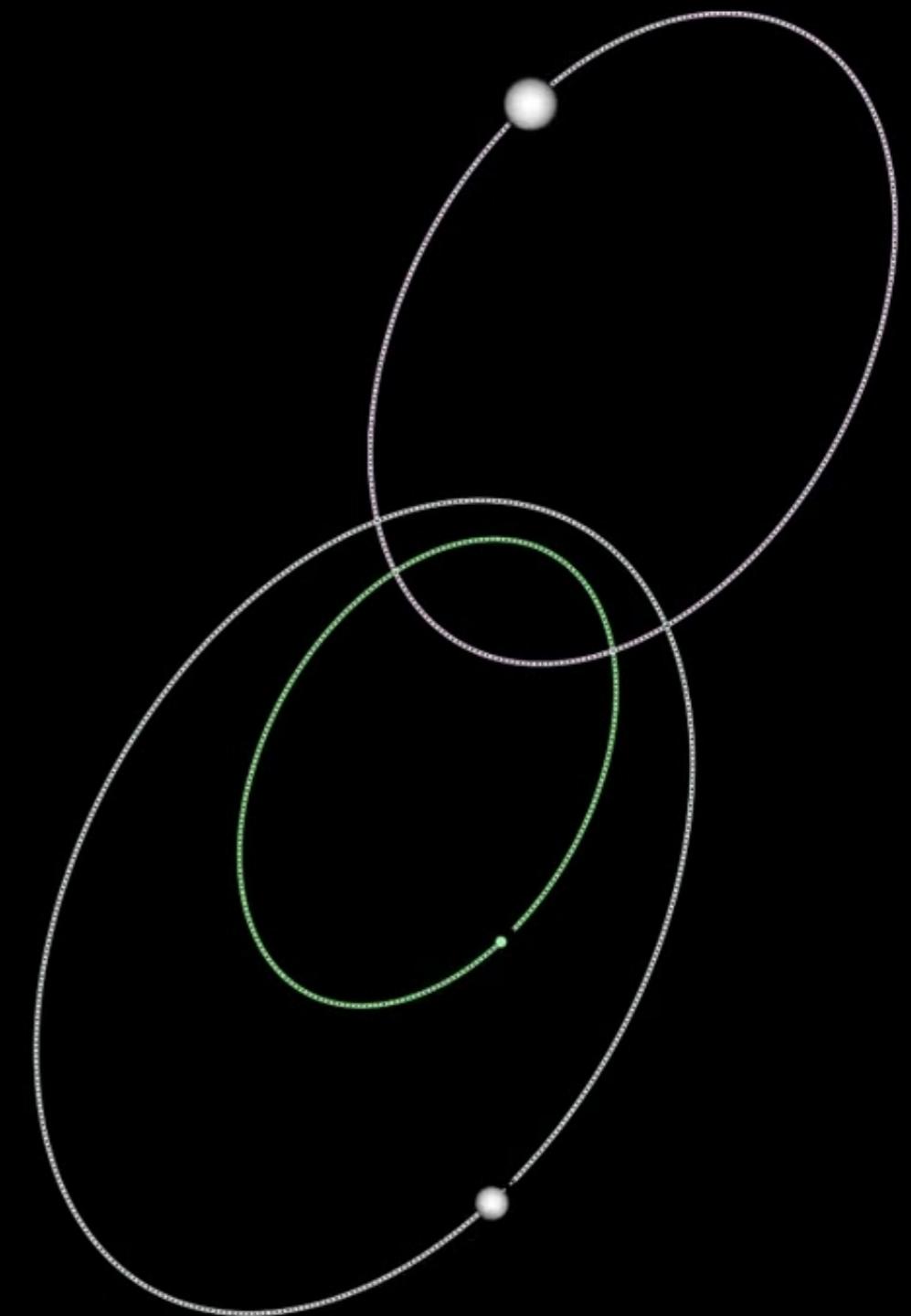
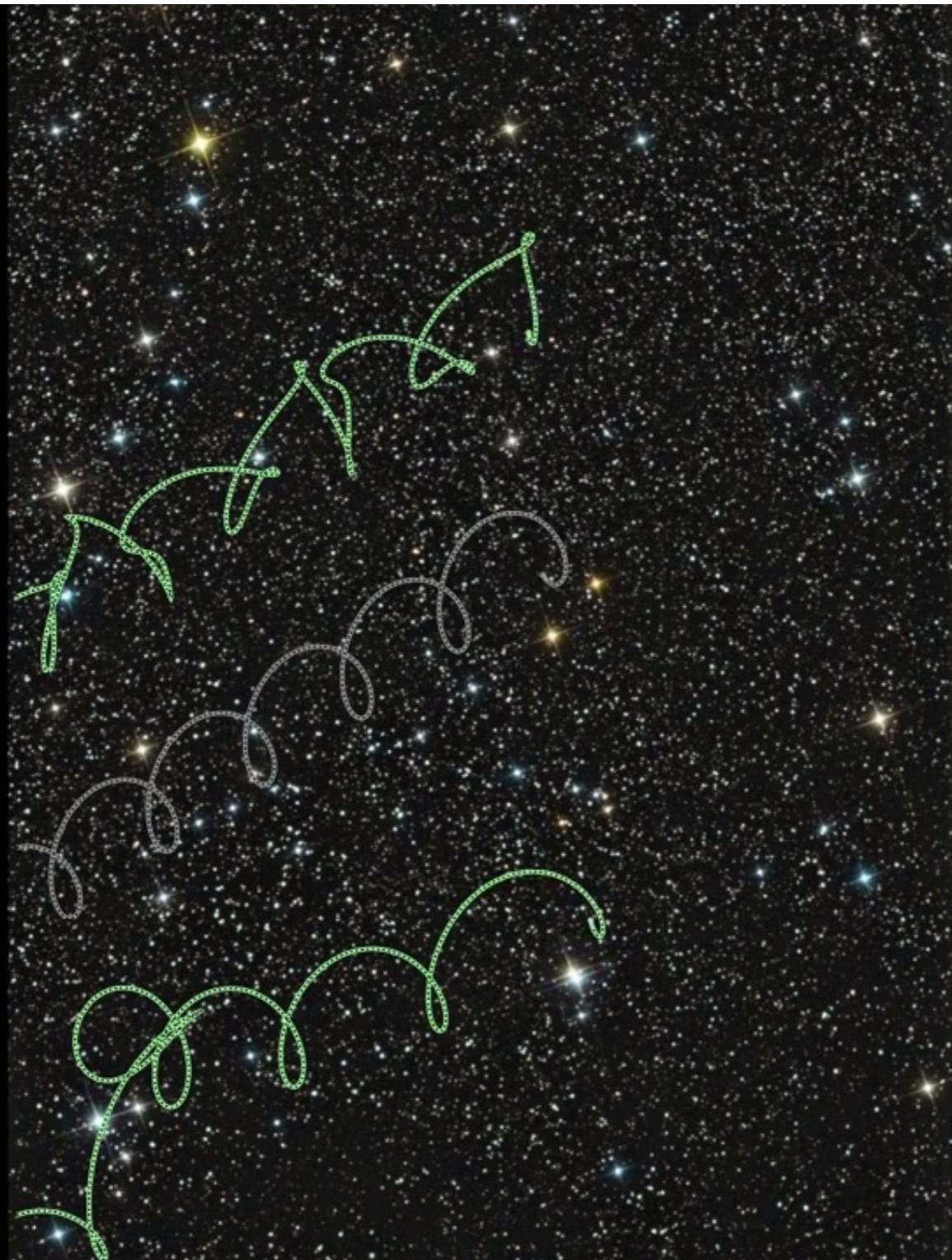


$t = 5.56$ (yr)



Gaia
DPAC

ULB
belspo



<https://www.youtube.com/watch?v=p4ICN8Ch2JA>

Astrometric Wobble of a Star Induced by Its Companion Planet

- Center of Mass $M_\star r_\star = M_p r_p$
- Semi-major axis $a = r_p + r_\star$
- Thus, $r_\star = a \frac{M_p}{M_\star + M_p} \approx a \frac{M_p}{M_\star}$
- Angular astrometric shift $\alpha = r_\star/d$,

$$\alpha = \frac{a}{d} \frac{M_p}{M_\star + M_p} \approx \frac{a}{d} \frac{M_p}{M_\star} \approx \left(\frac{M_p}{M_\star} \right) \left(\frac{a}{1 \text{ au}} \right) \left(\frac{d}{1 \text{ pc}} \right)^{-1} \text{ arcsec}$$

In-Class Activity

The wobbling of Barnard's star

Astrometric Missions

Ground-based

- Challenging due to atmospheric turbulence, limited seeing → Solutions: Adaptive Optics, Interferometry

Space-based

- Hipparcos (1989-1993)
 - Achieved ~1 mas accuracy for $V = 10$
- Gaia (2013-2025)
 - Targeting at ~10 μ as accuracy for $G = 10$