

STAT 621 HW 6 Spectral Analysis and Filtering

Yingying Xu

4/4/2018 (Due)

Verify Equation 4.5

$$\gamma_x(h) = \sum_{k=1}^q \sigma_k^2 \cos(2\pi w_k h)$$

Based on Equation 4.4:

$$x_t = \sum_{k=1}^q [U_{k1} \cos(2\pi w_k t) + U_{k2} \sin(2\pi w_k t)]$$

We have:

$$\begin{aligned} \gamma_x(h) &= \text{cov}(x_{t+h}, x_t) = \text{cov}\left(\sum_{k=1}^q [U_{k1} \cos(2\pi w_k(t+h)) + U_{k2} \sin(2\pi w_k(t+h))], \sum_{k=1}^q [U_{k1} \cos(2\pi w_k t) + U_{k2} \sin(2\pi w_k t)]\right) \\ &= \sum_{k,l=1}^q \text{cov}([U_{k1} \cos(2\pi w_k(t+h)) + U_{k2} \sin(2\pi w_k(t+h))], [U_{l1} \cos(2\pi w_l t) + U_{l2} \sin(2\pi w_l t)]) \\ &= \sum_{k=1}^q \text{cov}([U_{k1} \cos(2\pi w_k(t+h)) + U_{k2} \sin(2\pi w_k(t+h))], [U_{k1} \cos(2\pi w_k t) + U_{k2} \sin(2\pi w_k t)]) \\ &= \sum_{k=1}^q \text{cov}(U_{k1} \cos(2\pi w_k(t+h)) + U_{k1} \cos(2\pi w_k t)) + \sum_{k=1}^q \text{cov}(U_{k2} \sin(2\pi w_k(t+h)) + U_{k2} \sin(2\pi w_k t)) \\ &= \sum_{k=1}^q \sigma_k^2 \cos(2\pi w_k(t+h)) \cos(2\pi w_k t) + \sum_{k=1}^q \sigma_k^2 \sin(2\pi w_k(t+h)) \sin(2\pi w_k t) \\ &= \sum_{k=1}^q \sigma_k^2 \cos(2\pi w_k h) \end{aligned}$$

Q4.5 (Omit, get the wrong question)

$$x_t = w_t - \theta w_{t-1}$$

a) Mean and Autocovariance

$w_t \sim N(0, 1)$, mean is 0 and variance is 1. $r_w(s, t) = \text{cov}(w_s, w_t) = \sigma_w^2 = 1$, for $s = t$; $r_w(s, t) = 0$ for $s \neq t$.

Mean: $\mu_x = E[w_t - \theta w_{t-1}] = 0$;

Autocovariance: $r_x(s, t) = \text{cov}(x_s, x_t) = E[(x_s - \mu_x)(x_t - \mu_x)] = 1 + \theta^2$, for $s = t$; $r_x(s, t) = \text{cov}(x_s, x_t) = -\theta$, for $|t - s| = 1$; $r_x(s, t) = 0$, else.

Both series are stationary, because the mean is constant and the autocovariance only depends on $|t - s|$.

b) Power Spectrum

Since $\gamma_0 = 1 + \theta^2, \gamma_{\pm 1} = \theta$.

$$f(\omega) = \sum_{h=-\infty}^{\infty} \gamma_h e^{-2\pi i \omega h} = \sigma_w^2 [1 + \theta^2 + \theta(e^{2\pi i \omega} + e^{-2\pi i \omega h})] = 1 + \theta^2 + 2\theta \cos(2\pi \omega)$$

Note: we have used $\cos(\alpha) = (e^{i\alpha} + e^{-i\alpha})/2$. And in our case, $\sigma_w^2 = 1$.

4.6

$$x_t = \phi x_{t-1} - w_t$$

a) Power Spectrum

$$(1 - \phi B)x_t = w_t, \Phi(B) = 1 - \phi B$$

$$\gamma_0 = \sigma^2/(1 - \phi^2) \text{ and } \gamma_1 = \sigma^2\phi/(1 - \phi^2)$$

$$|\phi(e^{-2\pi i \omega})|^2 = 1 + \phi^2 - 2\phi e^{-2\pi i \omega} = 1 + \phi^2 - 2\phi \cos(2\pi \omega)$$

$f_w(w) = \sigma_w^2$. By property 4.4, we have:

$$f_x(\omega) = \frac{\sigma_w^2}{|\phi(e^{-2\pi i \omega})|^2} = \frac{\sigma_w^2}{1 + \phi^2 - 2\phi \cos(2\pi \omega)}$$

b) Verify the autocovariance function

From traditional methods: $\gamma_0 = \sigma^2(1 + \phi^2 + \phi^4 + \dots) = \sigma^2/(1 - \phi^2)$; $\gamma_{\pm 1} = \sigma^2(\phi + \phi^3 + \phi^5 + \dots) = \sigma^2\phi/(1 - \phi^2)$; $\gamma_{\pm 2} = \sigma^2(\phi^2 + \phi^4 + \phi^6 + \dots) = \sigma^2\phi^2/(1 - \phi^2)$; ... Therefore:

$$\gamma_x(h) = \frac{\sigma_w^2 \phi^{|h|}}{1 - \phi^2}$$

The inverse transform of $\gamma_x(h)$

$$\gamma_w(h) = \text{cov}(w_{t+h}, w_t) = \text{cov}(x_{t+h} - \phi x_{t+h-1}, x_t - \phi x_{t-1}) = \gamma_x(h) - \phi \gamma_x(h+1) - \phi \gamma_x(h-1) + \phi^2 \gamma_x(h)$$

$$= \gamma_x(h)(1 + \phi^2) - \phi(\gamma_x(h+1) + \gamma_x(h-1)) = \int_{-1/2}^{1/2} [(1 + \phi^2) - \phi(e^{2\pi i w} + e^{-2\pi i w})] e^{2\pi i w h} f_x(w) dw$$

where $(e^{2\pi i w} + e^{-2\pi i w}) = 2\cos(2\pi w)$. If $w_t \sim WN$, $g_w(w) = \sigma_w^2$:

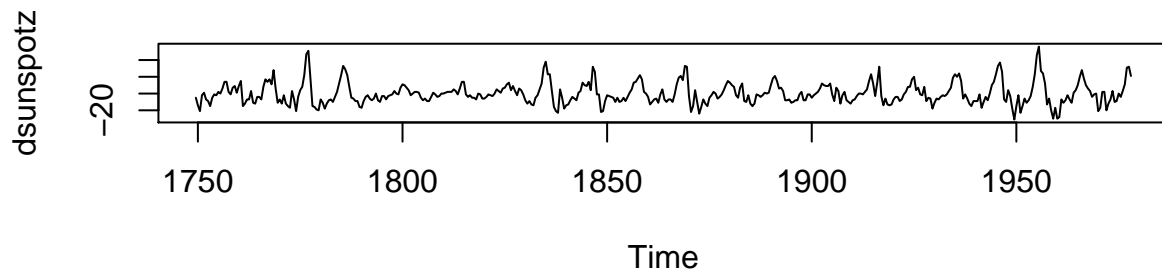
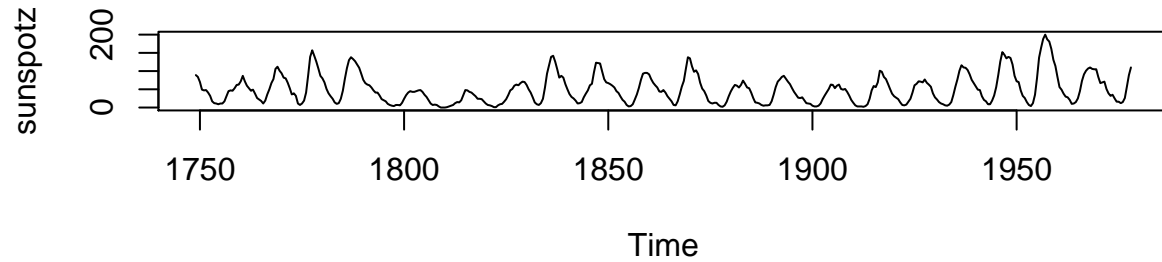
$$g_w(w) = \sigma_w^2 = [(1 + \phi^2) - 2\phi \cos(2\pi w)] f_x(w)$$

By the uniqueness of Fourier inverse transformation:

$$\gamma_x(h) = \frac{\sigma_w^2 \phi^{|h|}}{1 - \phi^2}$$

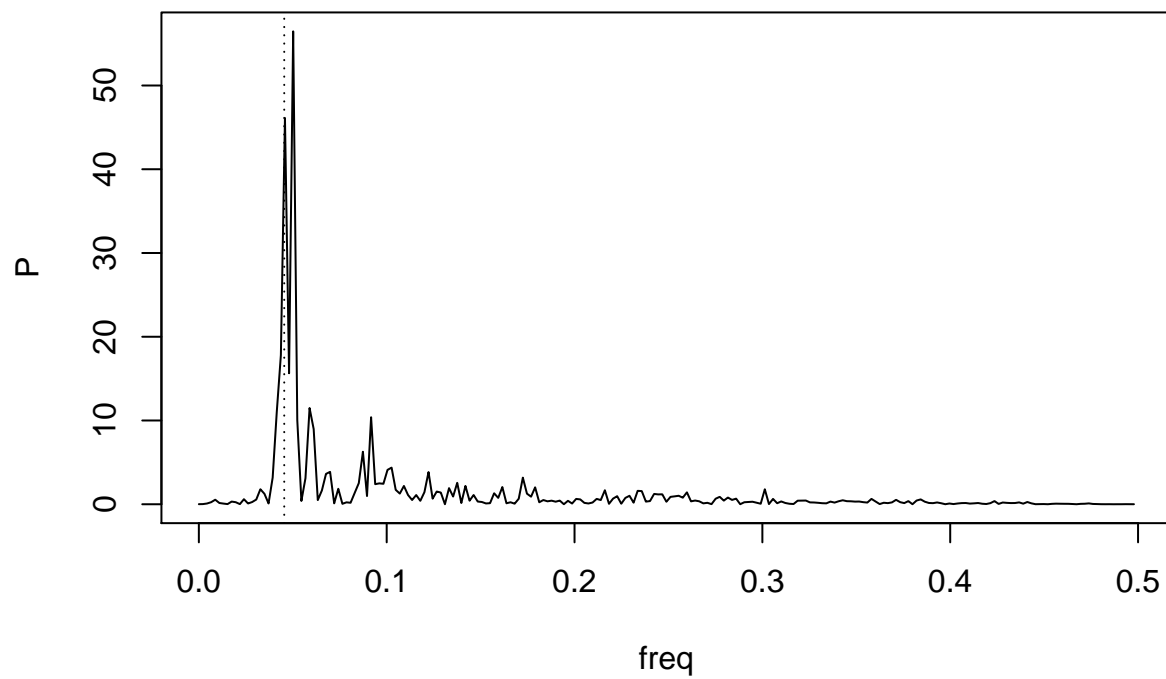
4.9 Sunspot Series

```
par(mfrow=c(2,1))
plot(sunspotz) #length=459
dsunspotz=diff(sunspotz)
plot(dsunspotz) #length=458
```



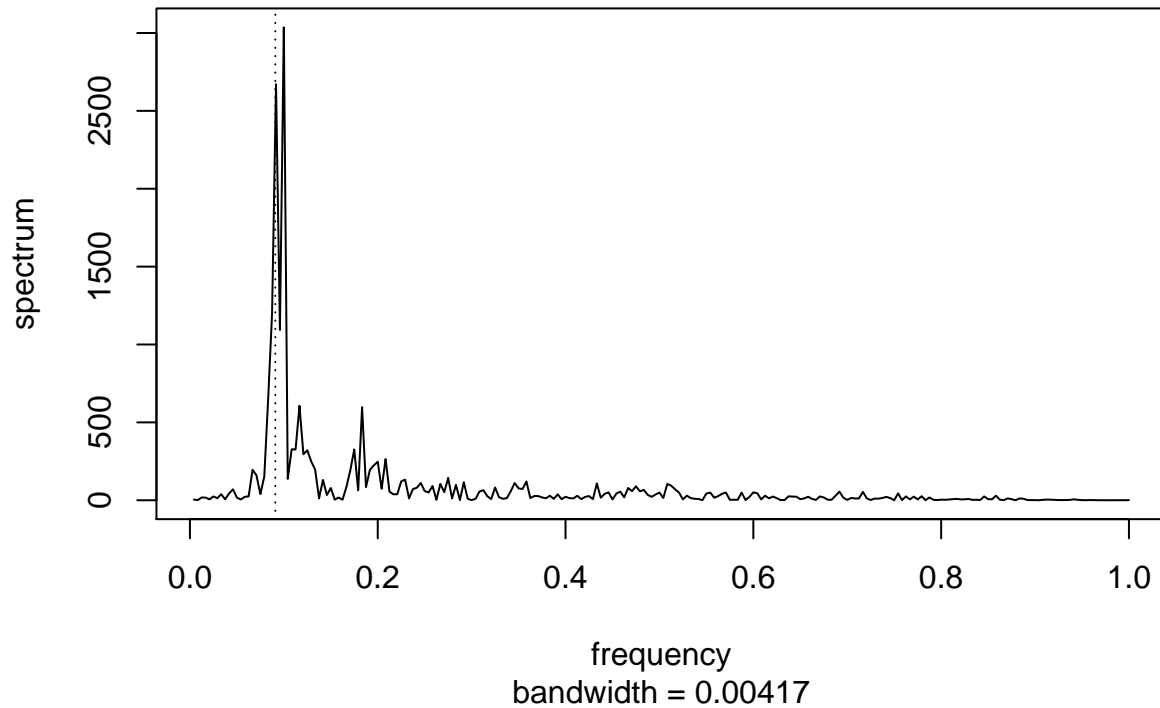
Solution 1 Use `diff(sunspotz)`

```
I=abs(fft(dsunspotz))^2/458
P=(4/458)*I[1:229]
freq=(0:228)/458
plot(freq,P,type="l")
abline(v=1/22, lty="dotted") # one cycle every 11 years
```



```
#Another code with exactly same results is:  
dsunspotz.per = mvspec(dsunspotz, log="no")  
abline(v=1/11, lty="dotted") # one cycle every 11 years
```

Series: dsunspotz Raw Periodogram



The confidence intervals for the sunspotz series at the 11 years cycle, $\omega = 1/22 \approx 21/458$.

```
invisible(dsunspotz.per$spec[21]) # 1204.158; sunspotz pgram at freq 1/22
```

```
dsunspotz.per$freq[1] #0.004166667
```

```
## [1] 0.004166667
```

```
dsunspotz.per$df #1.908333, choose df=2
```

```
## [1] 1.908333
```

```
# conf intervals - returned value:
```

```
2*dsunspotz.per$spec[21]/qchisq(.975,2) # 326.4291
```

```
## [1] 326.4291
```

```
2*dsunspotz.per$spec[21]/qchisq(.025,2) # 47561.69
```

```
## [1] 47561.69
```

An approximate 95% confidence interval (df=2) for the $\text{diff}(\text{spectrum}) f(1/22)$ is $[326.4291, 47561.69]$.

Solution 2 (not detrend)

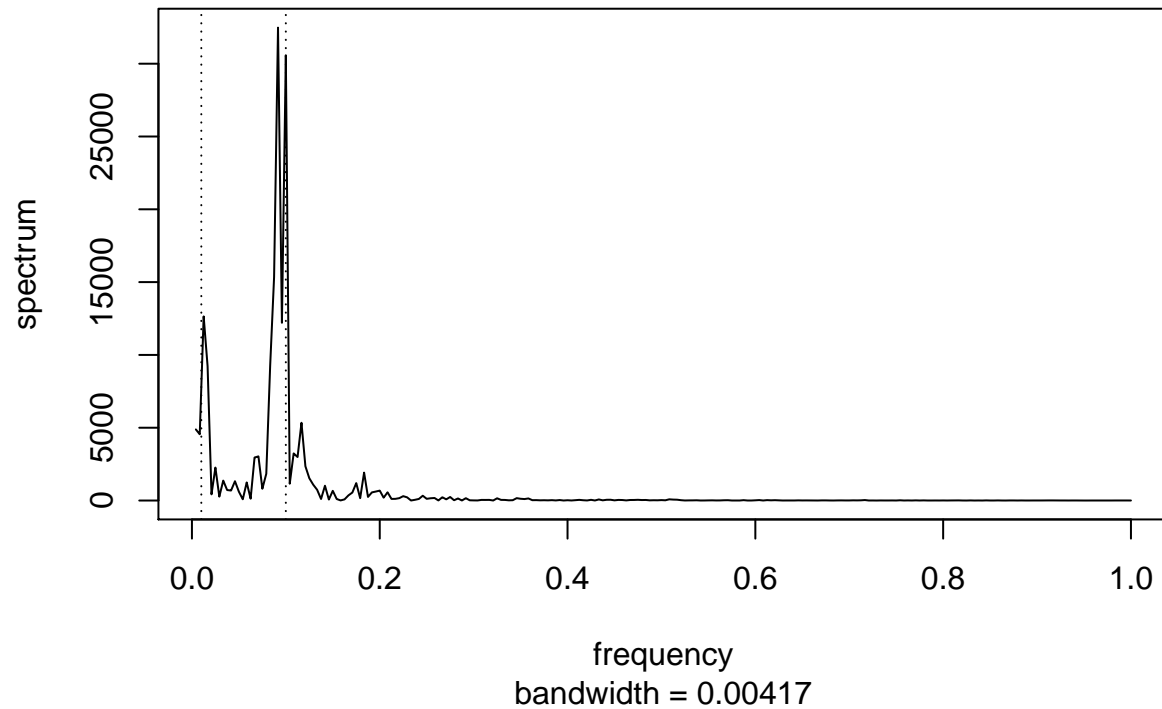
The frequency axis is labeled in multiples of $\Delta = 1/2$. The peaks at $\omega = 0.1\Delta = 1/20$, or one cycle per 10 years. And another peak at $\omega = 0.01\Delta = 1/200$, or one cycle per 100 years.

```

sunspotz.per = mvspec(sunspotz, log="no")
abline(v=0.1, lty="dotted") #w=1/10, one cycle every 10 years
abline(v=0.01, lty="dotted") #w=1/100, one cycle every 100 years

```

Series: sunspotz Raw Periodogram



The confidence intervals for the sunspotz series at the 10 years cycle, $\omega = 1/20 \approx 23/459$, and the 100 years cycle $\omega = 1/200 \approx 2.3/459$

```

sunspotz.per$spec[23] # 12205.53; sunspotz pgram at freq 1/20

```

```
## [1] 12205.53
```

```

sunspotz.per$spec[2.3] # 4551.155; sunspotz pgram at freq 1/200, same as spec[2]

```

```
## [1] 4551.155
```

```

dsunspotz.per$freq[1] #0.004166667

```

```
## [1] 0.004166667
```

```

dsunspotz.per$df #1.908333, choose df=2

```

```
## [1] 1.908333
```

```
# conf intervals - returned value:
```

```
U = qchisq(.025,2) # 0.05063
```

```
L = qchisq(.975,2) # 7.37775
```

```
2*sunspotz.per$spec[23]/L # 3308.737
```

```
## [1] 3308.737
```

```
2*sunspotz.per$spec[23]/U # 482092.8
```

```
## [1] 482092.8
```

```
2*sunspotz.per$spec[2.3]/L # 1233.75
```

```
## [1] 1233.75
```

```
2*sunspotz.per$spec[2.3]/U # 179761
```

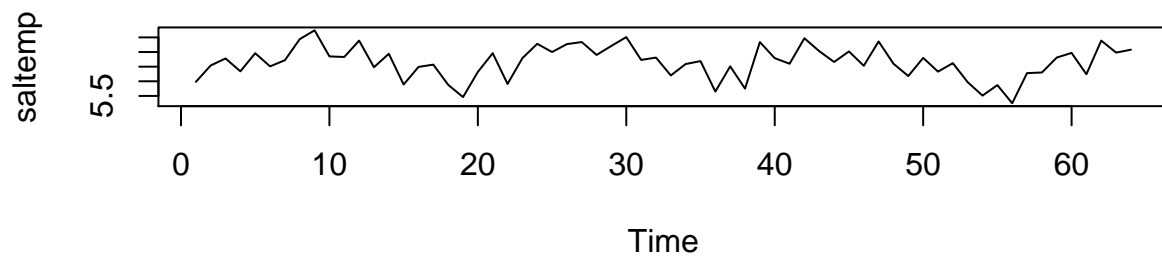
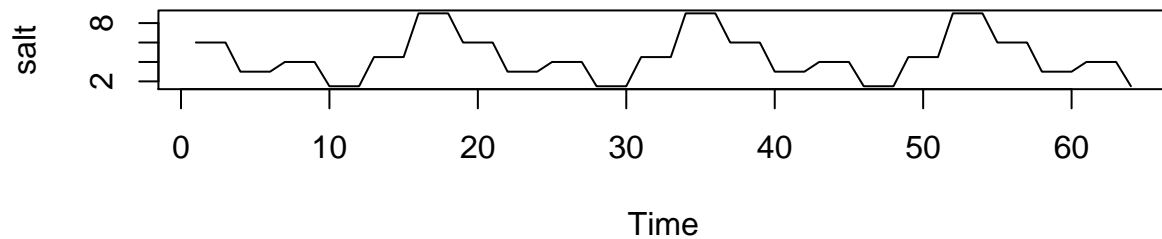
```
## [1] 179761
```

An approximate 95% confidence interval for the spectrum $f(1/20)$ is $[3308.7, 482092.8]$. An approximate 95% confidence interval for the spectrum $f(1/200)$ is $[1233.7, 179761]$.

We can also choose the circle=11 years, as shown in solution 1.

4.10 Salt and Temperature

```
par(mfrow=c(2,1))  
plot(salt)  
plot(saltemp)
```



```
length(salt) # n=64
```

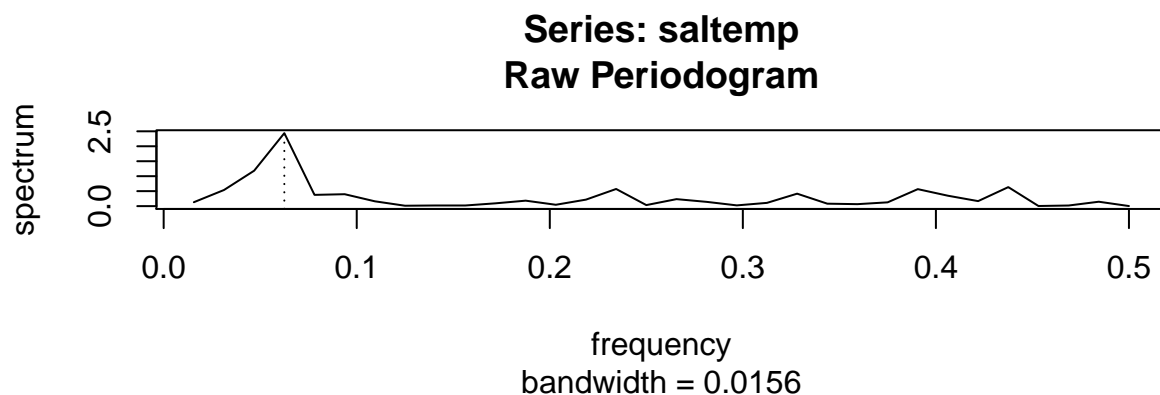
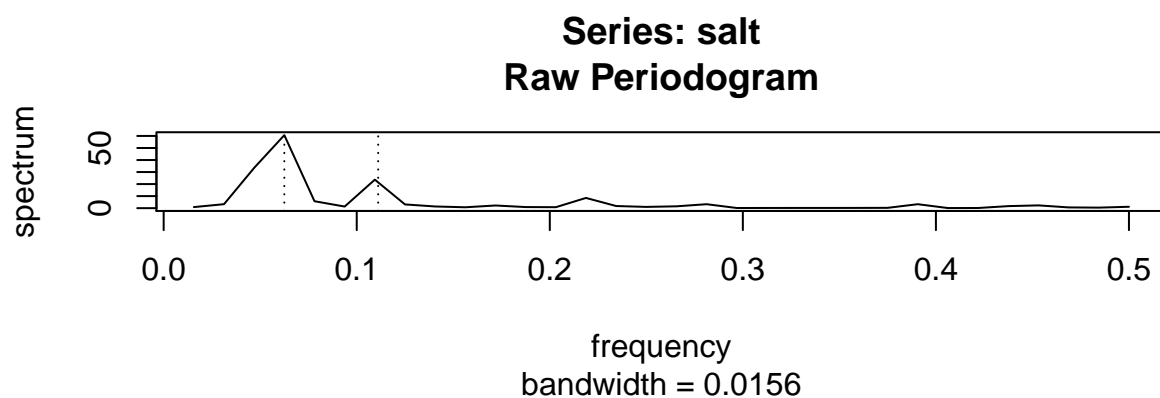
```
## [1] 64
```

```

par(mfrow=c(2,1))
salt.per = mvspec(salt, log="no")
abline(v=1/9, lty="dotted")
abline(v=1/16, lty="dotted")

saltemp.per = mvspec(saltemp, log="no")
abline(v=1/16, lty="dotted")

```



```
salt.per$freq[1] #0.015625
```

```
## [1] 0.015625
```

```
salt.per$df      #df=2
```

```
## [1] 2
```

```
saltemp.per$freq[1] #0.015625
```

```
## [1] 0.015625
```

```
saltemp.per$df    #df=2
```

```
## [1] 2
```

```
which.max(salt.per$spec) #4
```

```
## [1] 4
```

```
salt.per$freq[4] #0.0625=1/16
```



```
## [1] 0.0625
salt.per$freq[7] #0.109375=1/9

## [1] 0.109375
which.max(saltemp.per$spec) #4

## [1] 4
saltemp.per$freq[4] #0.0625=1/16

## [1] 0.0625
salt.per$spec[4] # 60.66648; salt pgram at freq 1/16 = 4/64

## [1] 60.66648
salt.per$spec[7] # 23.69029; salt pgram at freq 1/9 = 7/64

## [1] 23.69029
saltemp.per$spec[4] # 2.43787; saltemp pgram at freq 1/16 = 4/64

## [1] 2.43787
salt.per$df #df=2

## [1] 2
saltemp.per$df #df=2

## [1] 2
# conf intervals - returned value:
2*salt.per$spec[4]/qchisq(.975,2) # 16.44577

## [1] 16.44577
2*salt.per$spec[4]/qchisq(.025,2) # 2396.198

## [1] 2396.198
2*salt.per$spec[7]/qchisq(.975,2) # 6.422082

## [1] 6.422082
2*salt.per$spec[7]/qchisq(.025,2) # 935.7164

## [1] 935.7164
2*saltemp.per$spec[4]/qchisq(.975,2) # 0.6608701

## [1] 0.6608701
2*saltemp.per$spec[4]/qchisq(.025,2) # 96.29072

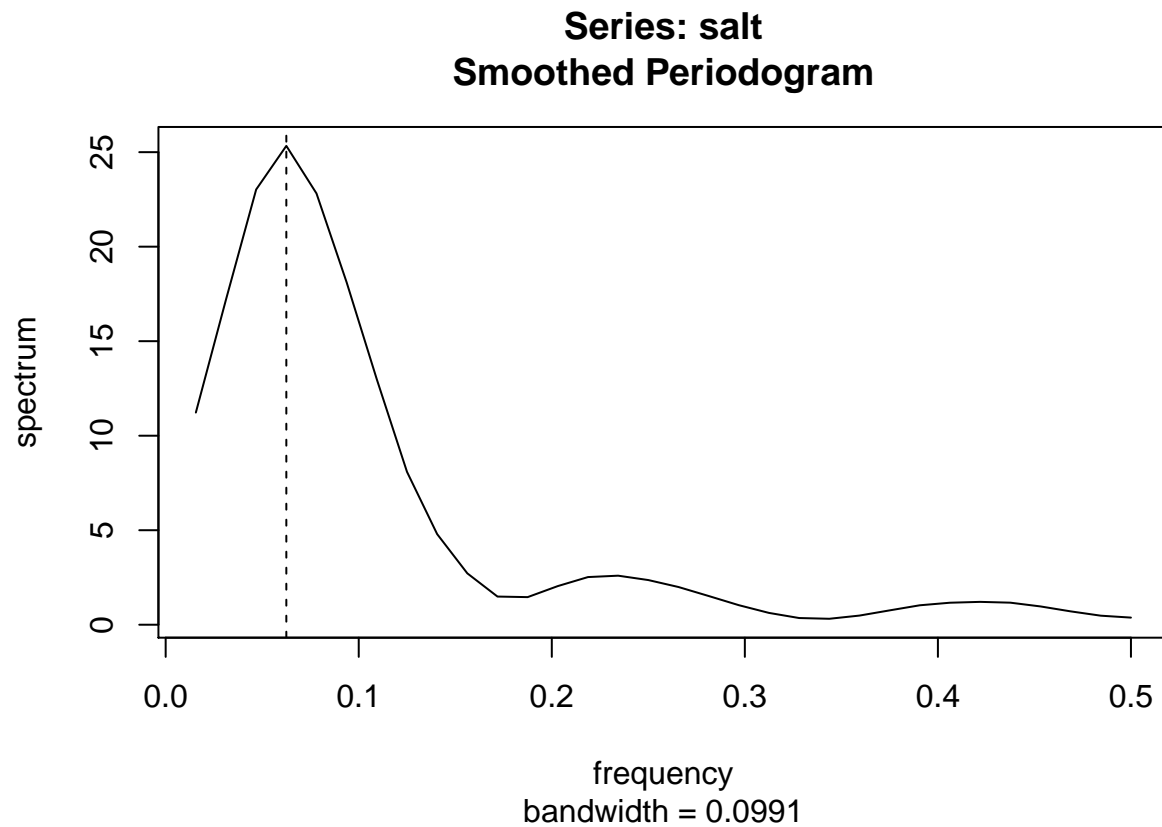
## [1] 96.29072
```

The frequency is 1/16 and 1/9 for salt, and 1/16 for saltemp.

The confidence intervals for the salt series at 1/16 freq, [16.4, 2396], and at 1/9 is [6.4, 935.7]. The confident intervals for the saltemp series at 1/16 freq is [0.66, 96.3].

4.15 Nonparametric spectral of Salt and Temperature

```
salt.smo = mvspec(salt, spans=c(5,5), taper=.1, log="no")  
abline(v = 1/16, lty=2)
```



```
salt.smo$df      # df = 11.35968
```

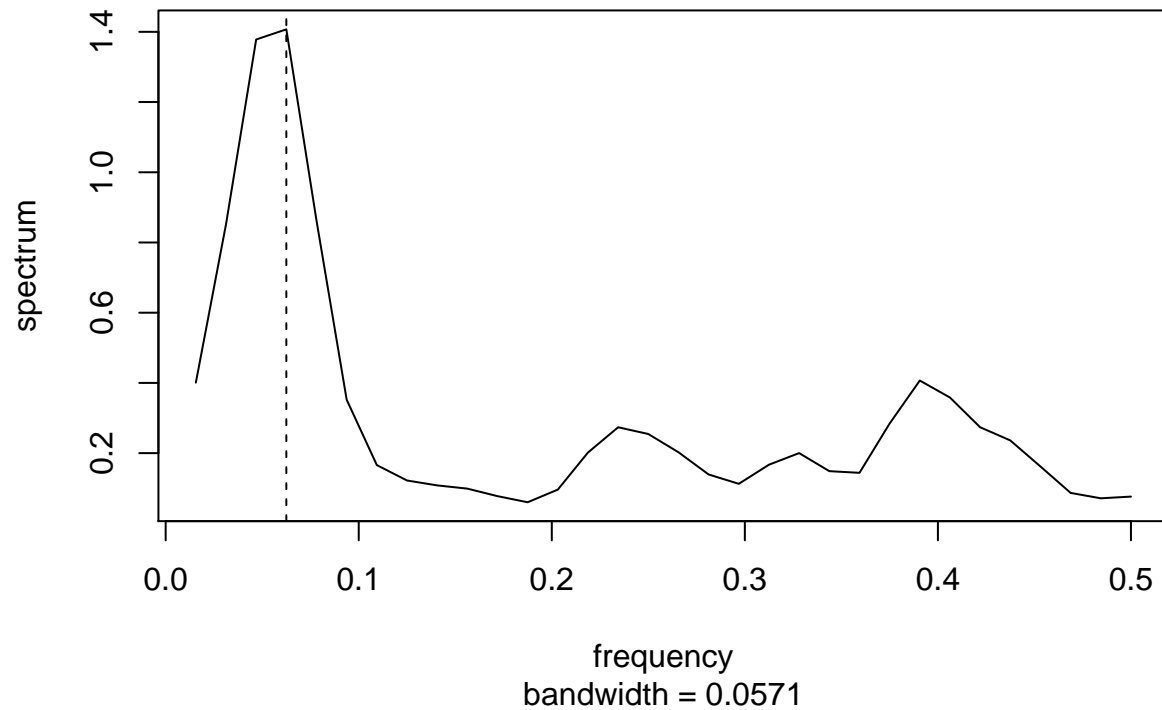
```
## [1] 11.35968
```

```
salt.smo$bandwidth # B = 0.09907121
```

```
## [1] 0.09907121
```

```
saltemp.smo = mvspec(saltemp, spans=c(3,3), taper=.1, log="no")  
abline(v = 1/16, lty=2)
```

Series: saltemp Smoothed Periodogram



```
saltemp.smo$df      # df = 6.552102
```

```
## [1] 6.552102
```

```
saltemp.smo$bandwidth # B = 0.05714286
```

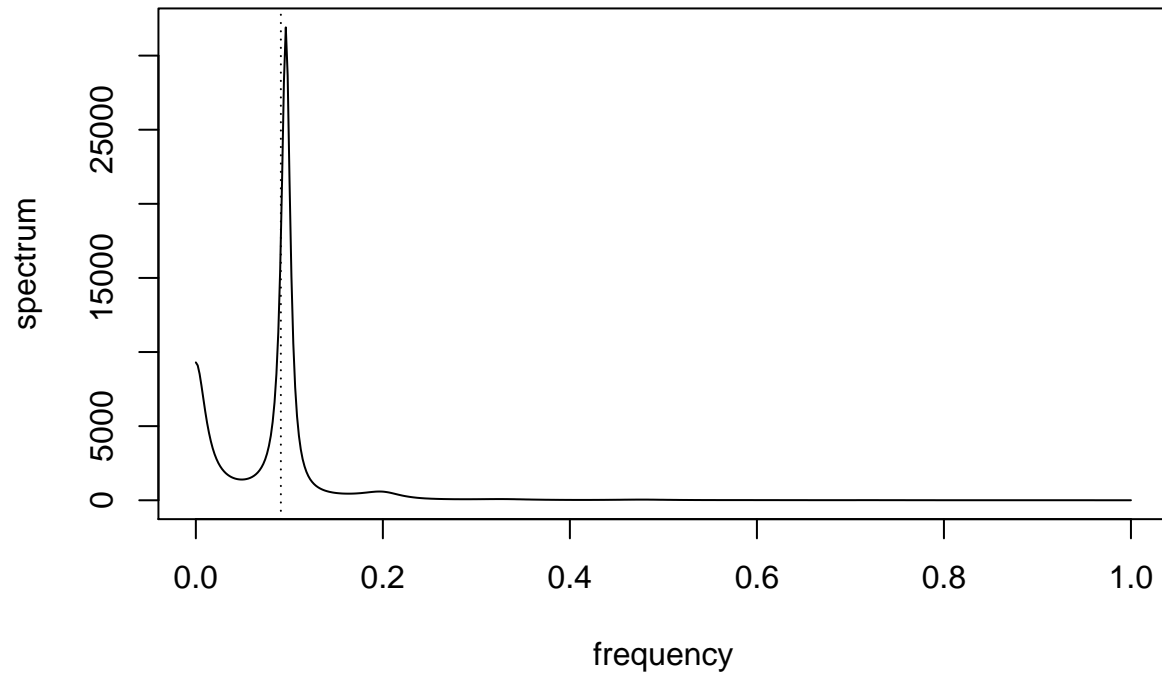
```
## [1] 0.05714286
```

Findings: When choose different spans and taper, the graph shape, degree of freedom, and bandwidth will also change accordingly. The peak will be influenced by the choice of smoothing parameter, according to the graph and theoretical proof. There is Bias-Variance tradeoff.

4.19 Sunspot with autoregressive spectral estimator

```
spaic=spec.ar(sunspotz,log="no")
abline(v=frequency(sunspotz)*1/22,lty=3)
```

Series: sunspotz AR (16) spectrum



```
(sunspotz.ar=ar(sunspotz,order.max = 30))
```

```
##
## Call:
## ar(x = sunspotz, order.max = 30)
##
## Coefficients:
##      1      2      3      4      5      6      7      8
## 1.5411 -0.7843  0.3014 -0.1899  0.0468 -0.0953  0.0925 -0.0099
##      9     10     11     12     13     14     15     16
## -0.0012 -0.1035  0.1529 -0.0960  0.0567 -0.1003  0.0326  0.0934
##
## Order selected 16  sigma^2 estimated as  73.81

n = length(sunspotz)
c() -> AIC -> AICc -> BIC
for (k in 1:30){
  fit = ar(sunspotz, order=k, aic=FALSE)
  sigma2 = fit$var.pred
  BIC[k] = log(sigma2) + (k*log(n)/n)
  AICc[k] = log(sigma2) + ((n+k)/(n-k-2))
  AIC[k] = log(sigma2) + ((n+2*k)/n)
}

dev.new()
IC = cbind(AICc, BIC+1)
ts.plot(IC, type="o", xlab="p", ylab="AIC / BIC")
```

`grid()`

AR(4) reached the minimum AIC and BIC. I will choose the AR(4) model. The conventional nonparametric spectral estimator omits the autoregression lags, and treat it as a peak.

4.30

$$y_t = \sum_r a_r x_{t-r}, z_t = \sum_s b_s x_{t-s}$$

a) Show the spectrum of the output

According to Property 4.3,

$$y_t = \sum_r a_r x_{t-r} \rightarrow f_y(w) = |A(w)|^2 f_x(w)$$

$$z_t = \sum_s b_s y_{t-s} \rightarrow f_z(w) = |B(w)|^2 f_y(w) = |B(w)|^2 [|A(w)|^2 f_x(w)] = |A(w)|^2 |B(w)|^2 f_x(w)$$

where $A(w) = \sum_r a_r e^{-2\pi i w r}$ and $B(w) = \sum_s b_s e^{-2\pi i w s}$.

b)

Apply the filter $u_t = x_t - x_{t-1}$ and $v_t = u_t - u_{t-12}$. The first difference filter will attenuate the lower frequencies and enhance the higher frequencies.

$$A(w) = \sum_{r=-\infty}^{\infty} a_r e^{-2\pi i w r} = 0 + e^{-2\pi i w (0)} - e^{-2\pi i w (1)} = 1 - e^{-2\pi i w}$$

$$|A(w)|^2 = (1 - e^{-2\pi i w})(1 - e^{2\pi i w}) = 2(1 - \cos(2\pi w))$$

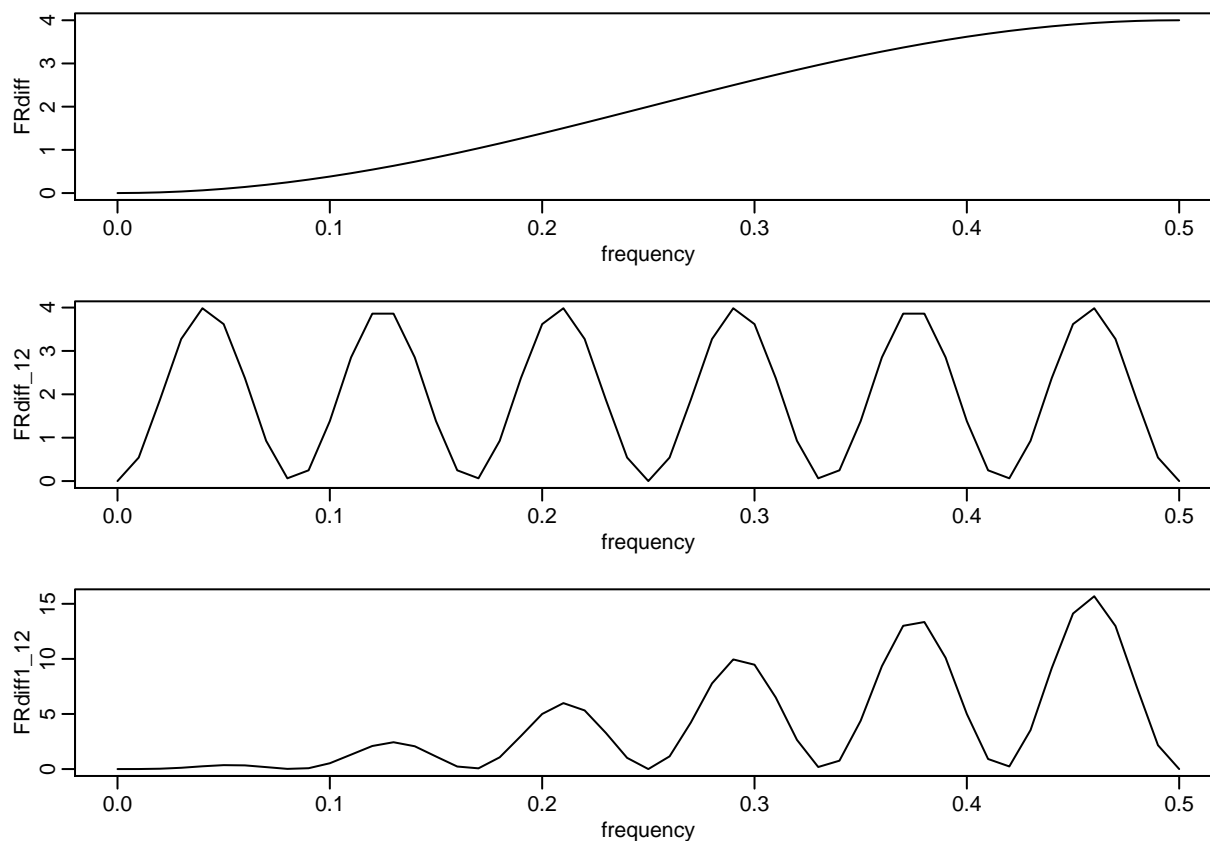
$$B(w) = \sum_{s=-\infty}^{\infty} b_s e^{-2\pi i w s} = 0 + e^{-2\pi i w (0)} - e^{-2\pi i w (12)} = 1 - e^{-24\pi i w}$$

$$|B(w)|^2 = (1 - e^{-24\pi i w})(1 - e^{24\pi i w}) = 2(1 - \cos(24\pi w))$$

$$f_z(w) = |A(w)|^2 |B(w)|^2 f_x(w) = 4(1 - \cos(2\pi w))(1 - \cos(24\pi w)) f_x(w)$$

c) Plot

```
par(mfrow=c(3,1), mar=c(3,3,1,1), mgp=c(1.6,.6,0))
w = seq(0, .5, by=.01)
FRdiff = abs(1-exp(2i*pi*w))^2
plot(w, FRdiff, type='l', xlab='frequency')
FRdiff_12=abs(1-exp(24i*pi*w))^2
plot(w, FRdiff_12, type='l', xlab='frequency')
FRdiff1_12=FRdiff*FRdiff_12
plot(w, FRdiff1_12, type='l', xlab='frequency')
```



4.31

$$y_t = ay_{t-1} + x_t, t = \pm 1, \pm 2, \dots$$

a)

$$f_y(w) = \frac{f_x(w)}{|\phi(e^{-2\pi iw})|^2} = \frac{f_x(w)}{|1 - ae^{-2\pi iw}|^2} = \frac{f_x(w)}{(1 - ae^{-2\pi iw})(1 - ae^{2\pi iw})} = \frac{f_x(w)}{1 + a^2 - a(e^{-2\pi iw} + e^{2\pi iw})} = \frac{f_x(w)}{1 + a^2 - 2a\cos(2\pi w)}$$

b)

Use function "curve" For a=0.1 and a=0.8

```
par(mfrow=c(2,1), mar=c(3,3,1,1), mgp=c(1.6,.6,0))

w = seq(0, .5, by=.01)
FRdiff_0.1 = 1/abs(1-0.1*exp(2i*pi*w))^2
plot(w,FRdiff_0.1,type='l',xlab ="frequency",ylab ="FR a=0.1")

FRdiff_0.8 = 1/abs(1-0.8*exp(2i*pi*w))^2
plot(w,FRdiff_0.8,type='l',xlab ="frequency",ylab ="FR a=0.8")
```

