1. **Problem.** Suppose we have a mechanical system with n degrees of freedom. Let $q_1(t), q_2(t), \ldots, q_n(t)$ be its generalized coordinates. Now consider a time-dependent coordinate transformation

$$Q_i = Q_i(t, q_1, q_2, \dots, q_n)$$
 $i = 1, 2, \dots, n.$

Show that if $q_i(t)$ solves a system of Euler-Lagrange equations involving a Lagrangian $L(t, q_i, \dot{q}_i)$, then $Q_i(t)$ solves the Euler-Lagrange equations involving $L(t, Q_i, \dot{Q}_i)$ provided the time-dependent coordinate transformation fulfills some minimal standard of good behavior. Specify this "minimal standard of good behavior."

Solution. Suppose that

$$\frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} = 0; \tag{1}$$

that is, $q_i(t)$ solve a system of Euler-Lagrange equations. We want to show that

$$\frac{\partial L}{\partial Q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{Q}_i} = 0. \tag{2}$$

Beginning with the first term of (??), we can use the chain rule for $L(t, Q_i, \dot{Q}_i)$ to write

$$\frac{\partial L}{\partial q_i} = \frac{\partial L}{\partial Q_j} \frac{\partial Q_j}{\partial q_i} + \frac{\partial L}{\partial \dot{Q}_j} \frac{\partial \dot{Q}_j}{\partial q_i},\tag{3}$$

provided there exists an inverse transformation

$$q_i = q_i(t, Q_1, Q_2, \dots, Q_n)$$
 $i = 1, 2, \dots, n$ (4)

that allows us to write $L(t, q_i, \dot{q}_i)$ in terms of t, Q_i , and \dot{Q}_i . This is only possible if there is a one-to-one correspondence between $q_i(t)$ and $Q_i(t)$, which is the "minimal standard of good behavior" for the transformation. We will assume the transformation is so well behaved.

Again using the chain rule for $Q_j = Q_j(t, q_1, q_2, \dots, q_n)$, note that

$$\dot{Q}_j = \frac{\partial Q_j}{\partial t} + \frac{\partial Q_j}{\partial q_i} \dot{q}_i \tag{5}$$

so (??) becomes

$$\frac{\partial L}{\partial q_i} = \frac{\partial L}{\partial Q_j} \frac{\partial Q_j}{\partial q_i} + \frac{\partial L}{\partial \dot{Q}_j} \left(\frac{\partial^2 Q_j}{\partial q_i \, \partial t} + \frac{\partial^2 Q_j}{\partial q_i \, \partial q_k} \dot{q}_k \right). \tag{6}$$

Applying the chain rule now to the second term of (??), we have

$$\frac{\partial L}{\partial \dot{q}_i} = \frac{\partial L}{\partial \dot{Q}_j} \frac{\partial \dot{Q}_j}{\partial \dot{q}_i} = \frac{\partial L}{\partial \dot{Q}_j} \frac{\partial Q_j}{\partial q_i} \tag{7}$$

where the right-hand side comes from (??). Then, using the product rule to take the time derivative,

$$\frac{d}{dt}\frac{\partial L}{\partial \dot{q}_i} = \left(\frac{d}{dt}\frac{\partial L}{\partial \dot{Q}_j}\right)\frac{\partial Q_j}{\partial q_i} + \frac{\partial L}{\partial \dot{Q}_j}\left(\frac{d}{dt}\frac{\partial Q_j}{\partial q_i}\right). \tag{8}$$

For the second term of (??), the chain rule for $Q_j = Q_j(t, q_1, q_2, \dots, q_n)$ gives

$$\frac{d}{dt}\frac{\partial Q_j}{\partial q_i} = \frac{\partial^2 Q_j}{\partial t \,\partial q_i} + \frac{\partial^2 Q_j}{\partial q_i \,\partial q_k}\dot{q}_k. \tag{9}$$

Substituting (??) into (??), we have

$$\frac{d}{dt}\frac{\partial L}{\partial \dot{q}_i} = \left(\frac{d}{dt}\frac{\partial L}{\partial \dot{Q}_j}\right)\frac{\partial Q_j}{\partial q_i} + \frac{\partial L}{\partial \dot{Q}_j}\left(\frac{\partial^2 Q_j}{\partial t \partial q_i} + \frac{\partial^2 Q_j}{\partial q_k \partial q_i}\dot{q}_k\right) \tag{10}$$

where the terms on the far right appeared in (??). Making this substitution,

$$\frac{d}{dt}\frac{\partial L}{\partial \dot{q}_i} = \left(\frac{d}{dt}\frac{\partial L}{\partial \dot{Q}_j}\right)\frac{\partial Q_j}{\partial q_i} + \frac{\partial L}{\partial q_i} - \frac{\partial L}{\partial Q_j}\frac{\partial Q_j}{\partial q_i},\tag{11}$$

and rearranging,

$$\frac{\partial Q_j}{\partial q_i} \left(\frac{\partial L}{\partial Q_j} - \frac{d}{dt} \frac{\partial L}{\partial \dot{Q}_j} \right) = \frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i},\tag{12}$$

Finally, substituting the original assumption (??), we find

$$\frac{\partial L}{\partial Q_j} - \frac{d}{dt} \frac{\partial L}{\partial \dot{Q}_j} = 0 \tag{13}$$

which is what we sought to prove.

2. Problem. Look at the Lagrangian

$$L = e^{\sigma t} \left(\frac{m\dot{q}^2}{2} - \frac{kq^2}{2} \right)$$

for one-dimensional motion.

- (a) Write down the associated Euler-Lagrange ODE.
- (b) Now perform a point transformation

$$Q = e^{\sigma t/2} q$$

where the new position coordinate Q is a function of t and q. What is the equation of motion for Q(t)? Are there conserved quantities?

Solution.

(a) Beginning from the general expression for the Euler-Lagrange equations,

$$\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = -e^{\sigma t} kq - \frac{d}{dt} \left(e^{\sigma t} m \dot{q} \right) = -m e^{\sigma t} \left(\ddot{q} + \sigma \dot{q} + \frac{k}{m} q \right)$$
(14)

so the ODE is

$$0 = \ddot{q} + \sigma \dot{q} + \frac{k}{m}q. \tag{15}$$

(b) It is possible to invert this transformation and write q = q(t, Q). Explicitly, this is

$$q = Qe^{-\sigma t/2} \tag{16}$$

so

$$\dot{q} = e^{-\sigma t/2} \left(\dot{Q} - \frac{\sigma}{2} Q \right). \tag{17}$$

Rewriting the Lagrangian such that $L = L(t, Q, \dot{Q})$ results in

$$L = e^{\sigma t} \left(\frac{m}{2} \left(e^{-\sigma t/2} \left(\dot{Q} - \frac{\sigma}{2} Q \right) \right)^2 - \frac{k}{2} \left(Q e^{-\sigma t/2} \right)^2 \right)$$
 (18)

$$=\frac{m}{2}\left(\dot{Q}-\frac{\sigma}{2}Q\right)^2-\frac{k}{2}Q^2\tag{19}$$

$$= \frac{m}{2} \left(\dot{Q}^2 - \sigma \dot{Q}Q + \left(\frac{\sigma^2}{4} - \frac{k}{m} \right) Q^2 \right). \tag{20}$$

Then the Euler-Lagrange equations are given by

$$0 = \frac{\partial L}{\partial Q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{Q}} = \frac{m}{2} \left(-\sigma \dot{Q} + 2 \left(\frac{\sigma^2}{4} - \frac{k}{m} \right) Q - \frac{d}{dt} \left(2 \dot{Q} - \sigma Q \right) \right)$$
(21)

which simplifies to

$$0 = \ddot{Q} + \left(\frac{k}{m} - \frac{\sigma^2}{4}\right)Q. \tag{22}$$

The solutions to (??) have the form

$$Q(t) = \begin{cases} A_1 \sin\left(\sqrt{\frac{k}{m} - \frac{\sigma^2}{4}}t\right) + A_2 \cos\left(\sqrt{\frac{k}{m} - \frac{\sigma^2}{4}}t\right) & \text{if } \frac{k}{m} > \frac{\sigma^2}{4}, \\ B_1 + B_2 t & \text{if } \frac{k}{m} = \frac{\sigma^2}{4}, \end{cases}$$

$$C_1 \exp\left\{-\sqrt{\frac{k}{m} - \frac{\sigma^2}{4}}t\right\} + C_2 \exp\left\{\sqrt{\frac{k}{m} - \frac{\sigma^2}{4}}t\right\} & \text{if } \frac{k}{m} < \frac{\sigma^2}{4}, \end{cases}$$

$$(23)$$

where A_i, B_i, C_i are real constants.

The Lagrangian in (??) does not explicitly depend on time (in contrast to the Lagrangian in the problem statement). Thus, the total energy H of the system is conserved. Explicitly,

$$H = \dot{Q}\frac{\partial L}{\partial \dot{Q}} - L = \frac{m}{2} \left(\dot{Q}^2 - \left(\frac{\sigma^2}{4} - \frac{k}{m} \right) Q^2 \right)$$
 (24)

is a conserved quantity.

3. **Problem.** Let $U(\alpha \mathbf{r}_1, \dots, \alpha \mathbf{r}_N)$ be a potential for N particles that satisfies the relation

$$U(\alpha \mathbf{r}_1, \dots, \alpha \mathbf{r}_N) = \alpha^k U(\mathbf{r}_1, \dots, \mathbf{r}_N).$$

The factor α can be any nonzero real number. The exponent k is an integer.

- (a) Show that the equations of motion associated with such a potential remain unchanged under a dilation of the distance scale if the time scale is also dilated by some other factor β . Find β as a function of α and k.
- (b) If k = 2, the forces correspond to a system of harmonic oscillators coupled to each other. Show that the result in part (a) implies the frequencies of such a system are independent of the oscillation amplitude.
- (c) If k = -1, we have an inverse square force law, such as that which arises in mutual gravitational attraction. Show that the result in part (a) implies Kepler's third law: the square of the orbital period of a planet is directly proportional to the cube of the semi-major axis of its orbit.

Solution. The Lagrangian $L = L(t, \mathbf{r}_i, \dot{\mathbf{r}}_i)$ for the system of N particles is

$$L = T - U = \frac{1}{2} m_i \dot{\mathbf{r}}_i \cdot \dot{\mathbf{r}}_i - U(\mathbf{r}_1, \dots, \mathbf{r}_N)$$
(25)

where m_i is the mass of the particle located at \mathbf{r}_i . The Euler-Lagrange equations for this Lagrangian are

$$\frac{\partial L}{\partial \mathbf{r}_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{\mathbf{r}}_i} = 0 \implies \frac{\partial U}{\partial \mathbf{r}_i} + m_i \ddot{\mathbf{r}}_i = 0.$$
 (26)

Define the time scale transformation

$$T = \beta t, \tag{27}$$

and define the coordinate transformation

$$\mathbf{R}_i = \mathbf{R}_i(T) = \alpha \mathbf{r}_i \tag{28}$$

for all N particles. Using these coordinates, the Lagrangian $L = L(T, \mathbf{R}_i, \dot{\mathbf{R}}_i)$ is

$$L = \frac{1}{2} m_i \dot{\mathbf{R}}_i \cdot \dot{\mathbf{R}}_i - U(\mathbf{R}_1, \dots, \mathbf{R}_N)$$
 (29)

and the Euler-Lagrange equations are

$$\frac{\partial L}{\partial \mathbf{R}_i} - \frac{d}{dT} \frac{\partial L}{\partial \dot{\mathbf{R}}_i} = 0 \implies \frac{\partial U}{\partial \mathbf{R}_i} + m_i \ddot{\mathbf{R}}_i = 0.$$
(30)

(a) The equations of motion associated to the Lagrangians (??) and (??) are identical if the Euler-Lagrange equations in (??) and (??) are identical. We will now show that this is the case.

The transformation $\mathbf{R}_i = \alpha \mathbf{r}_i$ is invertible, so $\mathbf{r}_i = \mathbf{R}_i/\alpha$. Likewise, $t = T/\beta$. By the chain rule,

$$\frac{d}{dT} = \frac{d}{dt}\frac{dt}{dT} = \frac{1}{\beta}\frac{d}{dt} \tag{31}$$

so

$$\dot{\mathbf{R}} = \alpha \frac{d\mathbf{r}_i}{dT} = \frac{\alpha}{\beta} \dot{\mathbf{r}}_i \tag{32}$$

and, likewise,

$$\ddot{\mathbf{R}} = \frac{\alpha}{\beta^2} \ddot{\mathbf{r}}_i. \tag{33}$$

From the given relationship for U, note that

$$U(\mathbf{R}_1, \dots, \mathbf{R}_N) = \alpha^k U(\mathbf{r}_1, \dots, \mathbf{r}_N)$$
(34)

and again using the chain rule,

$$\frac{\partial}{\partial \mathbf{R}_i} = \frac{\partial}{\partial \mathbf{r}_i} \frac{d\mathbf{r}_i}{d\mathbf{R}_i} = \frac{1}{\alpha} \frac{\partial}{\partial \mathbf{r}_i}.$$
 (35)

Making use of (??), (??), and (??), we can rewrite (??) in terms of the original coordinates:

$$0 = \frac{\alpha^k}{\alpha} \frac{\partial U}{\partial \mathbf{r}_i} + m_i \frac{\alpha}{\beta^2} \ddot{\mathbf{r}}_i \implies \alpha^{k-2} \beta^2 \frac{\partial U}{\partial \mathbf{r}_i} + m_i \ddot{\mathbf{r}}_i = 0$$
 (36)

which is equivalent to (??) so long as

$$\alpha^{k-2}\beta^2 = 1 \implies \beta = \alpha^{k-2/2}.$$
 (37)

(b) Fixing k=2, the solution of (a) gives us

$$\beta = \alpha^0 = 1. \tag{38}$$

This result indicates that the time scale is completely independent of the distance scale. That is, if we make the distance scale transformation $\mathbf{r}_i \mapsto \alpha \mathbf{r}_i$, the equations of motion will remain unchanged with respect to time $(t \mapsto \beta t = t)$.

For the system of harmonic oscillators, the oscillation amplitudes A_i have units of distance. The frequencies ω_i have units of inverse time. Making the transformation $\mathbf{r}_i \mapsto \alpha \mathbf{r}_i$ will change the distance scale and therefore the amplitudes, but the time scale and hence ω_i will remain unchanged. In other words, the frequencies are independent of the amplitudes. \square

(c) Fixing k = -1, the solution of (a) gives us

$$\beta = \alpha^{-3/2} = 1. {39}$$

This means making the distance scale transformation $\mathbf{r}_i \mapsto \alpha \mathbf{r}_i$ requires that we also make the time scale transformation $t \mapsto \alpha^{-3/2}t$.

The orbital period of a planet T has units of time, and its semi-major axis a has units of distance. agljshn

4. Problem. A particle in three-dimensional space is confined in a central potential

$$U(r) = -U_0 \left(\frac{r_0}{r}\right)^n.$$

Here $r = |\mathbf{r}|$ where $\mathbf{r}(t)$ is the location of the particle at time t, U_0 is a characteristic energy scale and r_0 is a characteristic length scale. Show that the particle motion is confined to a two-dimensional orbital plane. For what values of n are circular orbits stable?

Solution. We want to show that the particle motion is confined to a two-dimensional orbital plane.

The kinetic energy of the particle is

$$T = \frac{1}{2}m\dot{\mathbf{r}}^2\tag{40}$$

where m is the mass of the particle. Then the system has the Lagrangian

$$L = T - U = \frac{1}{2}m\dot{\mathbf{r}}^2 + U_0 \left(\frac{r_0}{r}\right)^n.$$
 (41)

We will change to a spherical coordinate system, defined (ρ, θ, ϕ) . Then

$$\dot{\mathbf{r}}^2 = \dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \dot{\phi}^2 \tag{42}$$

and the Lagrangian becomes

$$L(r, \dot{r}, \theta, \dot{\theta}, \phi, \dot{\phi}) = \frac{1}{2} m \left(\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \dot{\phi}^2 \right) + U_0 \frac{r_0^n}{r^n}. \tag{43}$$

Note that $\partial L/\partial \phi = 0$. Therefore, by Noether's theorem, ϕ is a cyclic coordinate and

$$\frac{\partial L}{\partial \dot{\phi}} = mr^2 \sin^2 \theta \dot{\phi} \tag{44}$$

is conserved.

HELP

Just use angular momentum argument

$$L(r, \dot{r}, \theta, \dot{\theta}) = \frac{1}{2} m \left(\dot{r}^2 + r^2 \dot{\theta}^2 \right) + U_0 \frac{r_0^n}{r^n}$$
(45)

For r, the Euler-Lagrange equations are

$$\frac{\partial L}{\partial r} - \frac{d}{dt} \frac{\partial L}{\partial \dot{r}} = mr\dot{\theta}^2 - nU_0 \frac{r_0^n}{r^{n+1}} - m\ddot{r}$$
(46)