

Data format for visualization/analysis of collisional  
 $N$ -body simulation Ver 0.0 — Aug 4 2011  
Ver 0.1 — Aug 17 2011

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# 1 Introduction

- need for common data format
- need to be able to reproduce complete trajectory of particles of interests
- machine independence
- performance
- parallel processing

## 2 Changes

### 2.1 Aug 17, 2011

Data format for binary components (or any hierarchical object) defined.

## 3 Basic idea

We need to store the data sufficient to reconstruct the orbits of individual particles. Conceptually, what we need is a stream of phase-space information of particles, such as:

```
particle_id, time, mass, x, y, z, vx, vy, vz, ...  
particle_id, time, mass, x, y, z, vx, vy, vz, ...  
particle_id, time, mass, x, y, z, vx, vy, vz, ...
```

However, the data format must be flexible enough to be able to include more information such as

- radius, and other info related to stellar evolution
- merger history
- whatever else you can think of

One way to construct such a flexible data format is to use self-describing data format, such as XML or YAML. For simplicity, we adopt YAML here.

### 3.1 Some basics of YAML

The following is a simple example of YAML data which can be directly converted to/from an Ruby object:

```
--- !!Particle
id: 0
x:
  - 0.1
  - 0.2
  - 0.3
v:
  - -1
  - -2
  - -3
m: 1.0
```

If this text is in a file “test.yaml”, one can read in it by

```
class Particle
  def taguri
    return 'x-private:Particle'
  end
end
require 'yaml'
YAML.add_private_type('Particle') do |type, val|
  YAML.object_maker(Particle, val)
end
obj= YAML.load_file('test.yaml')
```

The content of the object is now:

```
p obj
-> #<Particle:0xb7d151c0 @v=[-1, -2, -3], @id=0,
    @x=[0.1, 0.2, 0.3], @m=1.0>
```

As one can see, in this way we created an object of class Particle, with index, position, velocity, and mass given in the text. In the above example, the line

```
--- !!Particle
```

indicates that the following text should generate a ruby object with name “Particle”.

```
id: 0
```

defines a member variable with name “id”, and value 0 for this particular instance

```
x:
  - 0.1
  - 0.2
  - 0.3
```

means the member variable “x” is an array with three elements. The first “-” means this line is a data for an array. By default, numbers without “.” are regarded as integers, and with “.” floating point. Note that indentation has meaning here and “-” must be indented the same level or deeper than “x” and should be aligned. I guess I do not need to explain “v” and “m”

## 4 Particle Stream Data Format

### 4.1 normal particles

With the minimal description of YAML in the previous section, now we can define the generic data format (not yet for binaries, though): The data format is the stream of YAML representation of particle object, like

```
--- !!Particle
id: 0
t: 0
x:
  - 0.1
  - 0.2
  - 0.3
v:
  - -1
  - -2
  - -3
m: 1.0
--- !!Particle
id: 1
t: 0
x:
  - 0.2
  - 0.3
```

```

- 0.4
v:
- 0
- 0
- 0
m: 1.0
....

```

We need to have some naming convention, so that different programs can understand at least the basic common part of data. So we have

name	description
id	index (can be arbitrary text)
parent_id	index of parent
m	mass
t	time
dt	timestep
x	position, array with three elements
v	velocity, array with three elements
p	potential
a	acceleration, array with three elements
j	jerk, array with three elements
s	snap, array with three elements
c	crackle, array with three elements
pop	pop, array with three elements
X_something	User-defined extension

We call this format with basic naming convention PSDF (Particle Stream Data Format). Here, “X-something” implies that one can use any name starting with “X-” to record whatever thing one likes. If different users use the same name, like “X\_predicted\_position” for different things, that might cause confusion. So it might be better to have your own tag, like “NBODY6” or “Starlab” as the second-level tag, so that you have “X\_Starlab\_predicted\_position”.

We call a YAML document which corresponds to a single particle object as “world-point”, since it is a point on the worldline associated with a particle. It is translated to an object of Particle class, to call it particle is a bit confusing.

## 4.2 hierarchical subsystems

In many dynamics code, it is necessary to apply some special treatment for binary stars or other compact subsystems, and it is desirable to be able to retain such data structure in the I/O format. We express such data structure by “upward” pointers.

A binary, with “name” “0+1”, composed of two particles “0” and “1” appears in the data file just as a normal particle:

```
--- !!Particle
id: 0+1
t: 0
x:
  - 0.1
  - 0.2
  - 0.3
v:
  - -1
  - -2
  - -3
m: 2.0
```

It looks just as a normal particle. If one wants to say more, it is possible to include some more info, using user-defined extension (X\_something).

Components need to have the index of their parent, and we assume that the coordinates of components (and time derivatives) are relative to that of their parent.

So the worldpoint data of two components look like:

```
--- !!Particle
id: 0
parent_id: 0+1
t: 0
x:
  - 0.1
  - 0.0
  - 0.0
v:
  - -1
  - 0
  - 0
m: 1.0
```

and

```
--- !!Particle
id: 1
parent_id: 0+1
t: 0
```

```

x:
  - -0.1
  - 0.0
  - 0.0
v:
  - 1
  - 0
  - 0
m: 1.0

```

Note that there is no need to explicitly express formation and destruction of the “0+1” particle. It can just appear in the file when it is constructed, and from then on particles 0 and 1 will have “parent\_id” in the output data and that means their coordinates are relative. When the composite particle disappear, the parent\_id tag just disappear from the output fields of particles 0 and 1.

## 5 Some remarks

### 5.1 File format

PSDF does not define the on-memory data description. It does not define how the individual YAML documents are stored.

There are variety of ways to store worldpoints in file(s). In one extreme, all worldpoints are stored in a single big file, in the order defined by the time sequence. In the opposite extreme, one worldpoint is stored in one file. We could think of anything in between, such as single file for one particle or single file for some period in time.

In order to do some parallel I/O, using multiple files may be more natural. We do not yet specify any particular form of parallel access, though. We might provide some examples, though.

### 5.2 Interpolation etc

In the example given in the main text, a binary is given by:

```

--- !!Particle
id: 0+1
t: 0
x:
  - 0.1
  - 0.2

```



```

- 0.3
v:
- -1
- -2
- -3
m: 2.0

```

It looks just as a normal particle. If one wants to say more, it is possible to include some more info, using user-defined extension (X\_something).

Components need to have the index of their parent, and we assume that the coordinates of components (and time derivatives) are relative to that of their parent.

So the worldpoint data of two components look like:

```

--- !!Particle
id: 0
parent_id: 0+1
t: 0
x:
- 0.1
- 0.0
- 0.0
v:
- -1
- 0
- 0
m: 1.0

```

and

```

--- !!Particle
id: 1
parent_id: 0+1
t: 0
x:
- -0.1
- 0.0
- 0.0
v:
- 1
- 0
- 0
m: 1.0\begin{verbatim}

```

```

--- !!Particle
id: 0+1
t: 0
x:
  - 0.1
  - 0.2
  - 0.3
v:
  - -1
  - -2
  - -3
m: 2.0

```

It looks just as a normal particle. If one wants to say more, it is possible to include some more info, using user-defined extension (X\_something).

Components need to have the index of their parent, and we assume that the coordinates of components (and time derivatives) are relative to that of their parent.

So the worldpoint data of two components look like:

```

--- !!Particle
id: 0
parent_id: 0+1
t: 0
x:
  - 0.1
  - 0.0
  - 0.0
v:
  - -1
  - 0
  - 0
m: 1.0

```

and

```

--- !!Particle
id: 1
parent_id: 0+1
t: 0
x:
  - -0.1
  - 0.0

```

```
- 0.0
v:
- 1
- 0
- 0
m: 1.0
```

When you do the time integration, for a while only the internal motion is updated, resulting in sequence of data for particles 0 and 1. Particle 0+1 will be updated only some time later. If one wants the absolute position of particle 0 at any given time between 0 and the next output time of 0+1, one need to estimate the position of 0+1 in some way. If it has velocity, one can do linear prediction using one wp, or third-order Hermite interpolation using both the time 0 and the next time. Interpolation of even higher orders can be done if the wp has acceleration or higher order terms.

## 6 To do

- Add more examples

## A Repository

```
git@github.com:jmakino/Particle-Stream-Data-Format.git
```

you can download it by:

```
mkdir foo
cd foo
git clone git://github.com/jmakino/Particle-Stream-Data-Format.git
```

To update:

```
git pull
```

## B Sample codes

Also in the github.

## B.1 Base library

psdf.rb:

```
require 'yaml'
class Particle
  def taguri
    return 'x-private:Particle'
  end
end
YAML.add_private_type('Particle') do |type, val|
  YAML.object_maker(Particle, val)
end
```

This is current minimal “library”, which defines particle class and “Particle” tag.

## B.2 Simple program to generate a psdf file

writetest.rb:

```
require "psdf.rb"
class Particle
  attr_accessor :id, :x
  def initialize
    @id=0
    @t=0
    @x=[0,1,2]
  end
end
(0..10).each{|id|
  obj=Particle.new
  obj.id =id;
  obj.x[0]=id*0.1
  print YAML.dump(obj)
}
```

This one generate 10 worldpoints.

## B.3 Simple program to read a psdf file

readtest.rb:

```

require "psdf.rb"
a = []
while s = gets("---- ")
  print s
  print "\n\nend of one gets\n\n"
  print s, "\n"
  obj = YAML.load(s)
  a.push obj if obj
end
p a

```

This one reads in 10 particles and store everything in a single array.

## B.4 An ACS Body method to write a psdf worldpoint

acs\_psdf.rb:

```

require "psdf.rb"
class Particle
  attr_accessor :id, :x, :v, :a, :p, :j, :m, :t, :dt
end

class Body
  def to_psdf
    obj=Particle.new
    obj.id = @body_id
    obj.t = @time
    obj.x = Array[*@pos]
    obj.v = Array[*@vel]
    obj.a = Array[*@acc]
    obj.j = Array[*@jerk]
    obj
  end
  def psdf_output
    print YAML.dump(self.to_psdf)
  end
end

```

In the ACS Body class, position, velocity and other vector data are of class **Vector**, not **Array**. So you need to convert them to **Array**.

## B.5 An OpenGL animation program

The file `testplot.rb` gives a simple example. It is intended to be a skeleton code, with minimal UI and other stuff. It takes an input file with name “testin”, in which a sequence of psdf world points is stored (can come from shared-timestep or individual-timestep code).

This program requires ruby-opengl library. To install this (if not already installed) try:

```
sudo gem install opengl
```

On my machine with CentOS 5.5, it failed with:

```
sudo gem install ruby-opengl
Building native extensions. This could take a while...
ERROR: Error installing ruby-opengl:
        ERROR: Failed to build gem native extension.
```

```
/usr/bin/ruby -rubygems /usr/lib/ruby/gems/1.8/gems/rake-0.8.7/bin/rake RUBYARCHDIR=/
/usr/bin/ruby mkrf_conf.rb
(in /usr/lib/ruby/gems/1.8/gems/ruby-opengl-0.60.1)
rake
gcc -fPIC -g -O2 -Wall -DRUBY_VERSION=187 -I/usr/include -I/usr/lib/ruby/1.8/x86
(in /usr/lib/ruby/gems/1.8/gems/ruby-opengl-0.60.1/ext/gl)
gl-enums.c:6 include :
../common/common.h:45:21: error: GL/glut.h:
                                rake aborted!
Command failed with status (1): [gcc -fPIC -g -O2 -Wall -DRUBY_VERSION=...]
/usr/lib/ruby/gems/1.8/gems/ruby-opengl-0.60.1/ext/gl/Rakefile:29
(See full trace by running task with --trace)
rake aborted!
Command failed with status (1): [rake...]
```

(See full trace by running task with --trace)

Gem files will remain installed in `/usr/lib/ruby/gems/1.8/gems/ruby-opengl-0.60.1` for  
Results logged to `/usr/lib/ruby/gems/1.8/gems/ruby-opengl-0.60.1/gem_make.out`

In this case, you need to install GLUT library, by

```
sudo yum install freeglut freeglut-devel
```