

whence we have for the heat capacity of an electron gas, when  $\tau \ll \tau_F$ ,

$$C_{el} = \frac{1}{3}\pi^2 \mathfrak{D}(\varepsilon_F) \tau. \quad (34)$$

In conventional units,

$$C_{el} = \frac{1}{3}\pi^2 \mathfrak{D}(\varepsilon_F) k_B^2 T. \quad (35)$$

We found that the density of orbitals at the Fermi energy is

$$\mathfrak{D}(\varepsilon_F) = 3N/2\varepsilon_F = 3N/2\tau_F \quad (36)$$

for a free electron gas, with  $\tau_F \equiv \varepsilon_F$ . Do not be deceived by the notation  $\tau_F$ ; it is *not* the temperature of the Fermi gas, but only a convenient reference point. For  $\tau \ll \tau_F$  the gas is degenerate; for  $\tau \gg \tau_F$  the gas is in the classical regime. Thus (34) becomes

$$C_{el} = \frac{1}{2}\pi^2 N \tau / \tau_F. \quad (37)$$

In conventional units there is an extra factor  $k_B$ , so that


$$C_{el} = \frac{1}{2}\pi^2 N k_B T / T_F, \quad (38)$$

where  $k_B T_F \equiv \varepsilon_F$ . Again,  $T_F$  is not an actual temperature, but only a reference point.

We can give a physical explanation of the form of the result (37). When the specimen is heated from absolute zero, chiefly those electrons in states within an energy range  $\tau$  of the Fermi level are excited thermally, because the FD distribution function is affected over a region of the order of  $\tau$  in width, illustrated by Figures 7.3 and 7.5. Thus the number of excited electrons is of the order of  $N\tau/\varepsilon_F$ , and each of these has its energy increased approximately by  $\tau$ . The total electronic thermal energy is therefore of the order of  $U_{el} \approx N\tau^2/\varepsilon_F$ . Thus the electronic contribution to the heat capacity is given by

$$C_{el} = dU_{el}/d\tau \approx N\tau/\varepsilon_F \approx N\tau/\tau_F, \quad (39)$$

which is directly proportional to  $\tau$ , in agreement with the exact result (34) and with the experimental results.

$\mathfrak{D}(\varepsilon_F) \sim \frac{N}{2\varepsilon_F}$   
  
 $2 \times \frac{2}{3}$