

Optimal projection for parametric importance sampling in high dimension

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Abstract

In this paper we propose a dimension-reduction strategy in order to improve the performance of importance sampling in high dimension. The idea is to estimate variance terms in a small number of suitably chosen directions. We first prove that the optimal directions, i.e., the ones that minimize the Kullback–Leibler divergence with the optimal auxiliary density, are the eigenvectors associated to extreme (small or large) eigenvalues of the optimal covariance matrix. We then perform extensive numerical experiments that show that as dimension increases, these directions give estimations which are very close to optimal. Moreover, we show that the estimation remains accurate even when a simple empirical estimator of the covariance matrix is used to estimate these directions. These theoretical and numerical results open the way for different generalizations, in particular the incorporation of such ideas in adaptive importance sampling schemes.

Keywords: Importance sampling, High dimension, Gaussian covariance matrix, Kullback-Leibler divergence, Projection

Contents

```
import numpy as np
import scipy as sp
import pickle
import scipy.stats
import matplotlib.pyplot as plt
from CEIS_vMFNM import *
from IPython.display import display, Math, Latex
from IPython.display import Markdown
from tabulate import tabulate
np.random.seed(10)
np.random.rand(1,1)
 array([[0.77132064]])
# Table 2. Numerical comparison on test case 1
def Somme(x):
   n=np.shape(x)[1]
   return(np.sum(x,axis=1)-3*np.sqrt(n))
E=sp.stats.norm.cdf(-3)
n=100
d=n
# dimension
phi=Somme
def mypi(X):
   n=np.shape(X)[1]
   f0=sp.stats.multivariate_normal.pdf(X,mean=np.zeros(n),cov=np.eye(n))
   return((phi(X)>0)*f0)
N=2000
M=500
B=10 # number of runs
Eopt=np.zeros(B)
EIS=np.zeros(B)
```

```
Eprj=np.zeros(B)
Eprm=np.zeros(B)
Eprjst=np.zeros(B)
Eprmst=np.zeros(B)
Evmfn=np.zeros(B)
SI=[]
SIP=[]
SIPst=[]
SIM=[]
SIMst=[]
print(np.random.rand(1,1))
# Mstar
alpha=np.exp(-3**2/2)/(E*np.sqrt(2*np.pi))
Mstar=alpha*np.ones(d)/np.sqrt(d)
# Sigmastar
vstar=3*alpha-alpha**2+1
Sigstar = (vstar - 1)*np.ones((d,d))/d+np.eye(d)
Eigst=np.linalg.eigh(Sigstar)
logeigst=np.sort(np.log(Eigst[0])-Eigst[0])
deltast=np.zeros(len(logeigst)-1)
for i in range(len(logeigst)-1):
    deltast[i]=abs(logeigst[i]-logeigst[i+1])
## choice of the number of dimension
k_st=np.argmax(deltast)+1
indist=[]
for i in range(k_st):
    \verb|indist.append(np.where(np.log(Eigst[0])-Eigst[0] == logeigst[i])[0][0]|
P1st=np.array(Eigst[1][:,indist[0]],ndmin=2).T
for i in range(1,k_st):
    P1st=np.concatenate((P1st,np.array(Eigst[1][:,indist[i]],ndmin=2).T)\
                         ,axis=1) # matrix of influential directions
#np.random.seed(0)
for i in range(B):
########################## Estimation of the matrices
```

```
## g*-sample of size M
VA=sp.stats.multivariate_normal(np.zeros(n),np.eye(n))
X0=VA.rvs(size=M*1000)
ind=(phi(X0)>0)
X1=X0[ind,:]
X=X1[:M,:]
R=np.sqrt(np.sum(X**2,axis=1))
Xu=(X.T/R).T
## estimated gaussian mean and covariance
mm=np.mean(X,axis=0)
Xc = (X-mm).T
sigma =Xc @ Xc.T/np.shape(Xc)[1]
SI.append(sigma)
## von Mises Fisher parameters
normu=np.sqrt(np.mean(Xu,axis=0).dot(np.mean(Xu,axis=0).T))
mu=np.mean(Xu,axis=0)/normu
mu=np.array(mu,ndmin=2)
chi=min(normu, 0.95)
kappa=(chi*n-chi**3)/(1-chi**2)
## Nakagami parameters
omega=np.mean(R**2)
tau4=np.mean(R**4)
pp=omega**2/(tau4-omega**2)
Eig=np.linalg.eigh(sigma)
logeig=np.sort(np.log(Eig[0])-Eig[0])
delta=np.zeros(len(logeig)-1)
for j in range(len(logeig)-1):
    delta[j]=abs(logeig[j]-logeig[j+1])
k=np.argmax(delta)+1
indi=[]
for 1 in range(k):
    indi.append(np.where(np.log(Eig[0])-Eig[0]==logeig[1])[0][0])
P1=np.array(Eig[1][:,indi[0]],ndmin=2).T
```

```
for 1 in range(1,k):
       P1=np.concatenate((P1,np.array(Eig[1][:,indi[1]],ndmin=2).T)\
                         ,axis=1)
   diagsi=np.diag(Eig[0][indi])
   sig_opt_d=P1.dot((diagsi-np.eye(k))).dot(P1.T)+np.eye(n)
   SIP.append(sig_opt_d)
   ###
   diagsist=P1st.T.dot(sigma).dot(P1st)
   sig_opt=P1st.dot(diagsist-np.eye(k_st)).dot(P1st.T)+np.eye(n)
   SIPst.append(sig_opt)
   Norm_mm=np.linalg.norm(mm)
   normalised_mm=np.array(mm,ndmin=2).T/Norm_mm
   vhat=normalised_mm.T.dot(sigma).dot(normalised_mm)
   sig_mean_d=(vhat-1)*normalised_mm.dot(normalised_mm.T)+np.eye(n)
   SIM.append(sig_mean_d)
   ### MODIF : Sigma_opt=Sigma_mean donc inutile de recalculer
#
     Norm Mstar=np.linalg.norm(Mstar)
#
     normalised_Mstar=np.array(Mstar,ndmin=2).T/Norm_Mstar
#
     vhatst=normalised_Mstar.T.dot(sigma).dot(normalised_Mstar)
     sig_mean=(vhatst-1)*normalised_Mstar.dot(normalised_Mstar.T)+np.eye(n)
     SIMst.append(sig_mean)
Xop=sp.stats.multivariate_normal.rvs(mean=mm, cov=Sigstar,size=N)
   wop=mypi(Xop)/sp.stats.multivariate_normal.pdf(Xop,mean=mm, cov=Sigstar)
   Eopt[i]=np.mean(wop)
   ###
   Xis=sp.stats.multivariate_normal.rvs(mean=mm, cov=sigma,size=N)
   wis=mypi(Xis)/sp.stats.multivariate_normal.pdf(Xis,mean=mm, cov=sigma)
   EIS[i]=np.mean(wis)
   ###
   Xpr=sp.stats.multivariate_normal.rvs(mean=mm, cov=sig_opt_d,size=N)
```

```
wpr=mypi(Xpr)/sp.stats.multivariate_normal.pdf(Xpr,mean=mm,\
                                                    cov=sig_opt_d)
    Eprj[i]=np.mean(wpr)
   ###
    Xpm=sp.stats.multivariate_normal.rvs(mean=mm, cov=sig_mean_d,size=N)
    wpm=mypi(Xpm)/sp.stats.multivariate_normal.pdf(Xpm,mean=mm,\
                                                    cov=sig_mean_d)
    Eprm[i]=np.mean(wpm)
    Xprst=sp.stats.multivariate_normal.rvs(mean=mm, cov=sig_opt,size=N)
    wprst=mypi(Xprst)/sp.stats.multivariate_normal.pdf(Xprst,mean=mm, \
                                                        cov=sig_opt)
    Eprjst[i]=np.mean(wprst)
   ###
    Xvmfn = vMFNM_sample(mu, kappa, omega, pp, 1, N)
    Rvn=np.sqrt(np.sum(Xvmfn**2,axis=1))
    Xvnu=Xvmfn.T/Rvn
    h_log=vMF_logpdf(Xvnu,mu.T,kappa)+nakagami_logpdf(Rvn,pp,omega)
    A = np.log(n) + np.log(np.pi ** (n / 2)) - sp.special.gammaln(n / 2 + 1)
    f u = -A
    f_{chi} = (np.log(2) * (1 - n / 2) + np.log(Rvn) * (n - 1) - 0.5 * 
             Rvn ** 2 - sp.special.gammaln(n / 2))
    f_{\log} = f_{u} + f_{chi}
    W_{\log} = f_{\log} - h_{\log}
    wvmfn=(phi(Xvmfn)>0)*np.exp(W_log)
    Evmfn[i]=np.mean(wvmfn)
### KL divergences
dkli=np.zeros(B)
dklp=np.zeros(B)
dklm=np.zeros(B)
dklpst=np.zeros(B)
#dklmst=np.zeros(B)
dklpca=np.zeros(B)
for i in range(B):
    dkli[i]=np.log(np.linalg.det(SI[i]))+sum(np.diag(Sigstar.dot\
```

```
(np.linalg.inv(SI[i]))))
    dklp[i]=np.log(np.linalg.det(SIP[i]))+sum(np.diag(Sigstar.dot\
                                             (np.linalg.inv(SIP[i]))))
    dklm[i]=np.log(np.linalg.det(SIM[i]))+sum(np.diag(Sigstar.dot\
                                             (np.linalg.inv(SIM[i])))
    dklpst[i]=np.log(np.linalg.det(SIPst[i]))+sum(np.diag(Sigstar.dot\
                                             (np.linalg.inv(SIPst[i]))))
Tabresult=np.zeros((3,7)) # table of results
Tabresult[0,0]=np.log(np.linalg.det(Sigstar))+n
Tabresult[0,1]=np.mean(dkli)
Tabresult[0,2]=np.mean(dklpst)
Tabresult[0,3]=np.mean(dklpst)
Tabresult[0,4]=np.mean(dklp)
Tabresult[0,5]=np.mean(dklm)
Tabresult[0,6]=None
Tabresult[1,0]=np.mean(Eopt-E)/E*100
Tabresult[1,1]=np.mean(EIS-E)/E*100
Tabresult[1,2]=np.mean(Eprjst-E)/E*100
Tabresult[1,3]=np.mean(Eprjst-E)/E*100
Tabresult[1,4]=np.mean(Eprj-E)/E*100
Tabresult[1,5] = np.mean(Eprm-E)/E*100
Tabresult[1,6]=np.mean(Evmfn-E)/E*100
Tabresult[2,0]=np.sqrt(np.mean((Eopt-E)**2))/E*100
Tabresult[2,1]=np.sqrt(np.mean((EIS-E)**2))/E*100
Tabresult[2,2]=np.sqrt(np.mean((Eprjst-E)**2))/E*100
Tabresult[2,3]=np.sqrt(np.mean((Eprjst-E)**2))/E*100
Tabresult[2,4]=np.sqrt(np.mean((Eprj-E)**2))/E*100
Tabresult[2,5]=np.sqrt(np.mean((Eprm-E)**2))/E*100
Tabresult[2,6]=np.sqrt(np.mean((Evmfn-E)**2))/E*100
print(Tabresult)
Tabresult=np.round(Tabresult,1)
table=[["D'", Tabresult[0,0], Tabresult[0,1], Tabresult[0,2], Tabresult[0,3],
        Tabresult[0,4],Tabresult[0,5],"/"],
      ["Relative error (\%)", Tabresult[1,0], Tabresult[1,1],
       Tabresult[1,2], Tabresult[1,3], Tabresult[1,4], Tabresult[1,5], Tabresult[1,6]],
```

Table 1: Numerical comparison of the estimation of $E \approx 1.35 \cdot 10^{-3}$ considering the Gaussian model with the six covariance matrices defined in **?@sec-def_cov** and the vFMN model, when $\phi = \mathbb{I}_{\varphi \geq 0}$ with φ the linear function given by **?@eq-sum**. As explained in the text, $\widehat{\Sigma}_{\text{mean}}$ and $\widehat{\Sigma}_{\text{opt}}$ are actually equal in this case.

	Σ^*	$\widehat{\Sigma}^{\star}$	$\widehat{\Sigma}_{opt}$	$\widehat{\Sigma}_{mean}$	$\widehat{\Sigma}_{opt}^{+d}$	$\widehat{\Sigma}_{mean}^{+d}$	vMFN
D'	97.3	112	97.4	97.4	97.7	97.6	/
Relative error (%)	0.3	157.6	-1.5	-1.5	1.1	-0.6	- 0.4
Coefficient of variation (%)	6.9	548	2.3	2.3	4.8	5	3.6

```
np.random.rand(1,1)
array([[0.89818344]])
print(Tabresult)
```

- [[9.730e+01 1.120e+02 9.740e+01 9.740e+01 9.770e+01 9.760e+01 nan]
- [3.000e-01 1.576e+02 -1.500e+00 -1.500e+00 1.100e+00 -6.000e-01 -4.000e-01]
- [6.900e+00 5.480e+02 2.300e+00 2.300e+00 4.800e+00 5.000e+00 3.600e+00]]