

# Worksheet 7 - Solutions

## 1. The Meaning of "Never."

Let us consider to have  $N_{\text{monkeys}} = 10^{10}$ .

We need to count how likely it is to have  $10^5$  characters in one specific order, having 44 possibilities for each character.

$$\left(\frac{1}{44}\right)^{100000} = \frac{1}{44^{100000}}$$

but if we use that

$$\log_{10} 44 \approx 1,64345$$

↳  $44 \approx 10^{1,64345}$

So, the probability is

$$\left(\frac{1}{44}\right)^{100000} = 10^{-164345}$$

b) Now, knowing the writing speed, we have that since of the beginning of the universe;

$$(\text{age of universe}) \times \frac{10 \text{ keys}}{\text{second}} = 10^{18} \cancel{s} \left( \frac{10 \text{ keys}}{\cancel{s}} \right)$$

$= 10^{19}$  keys.

So, the 10 monkeys typed out

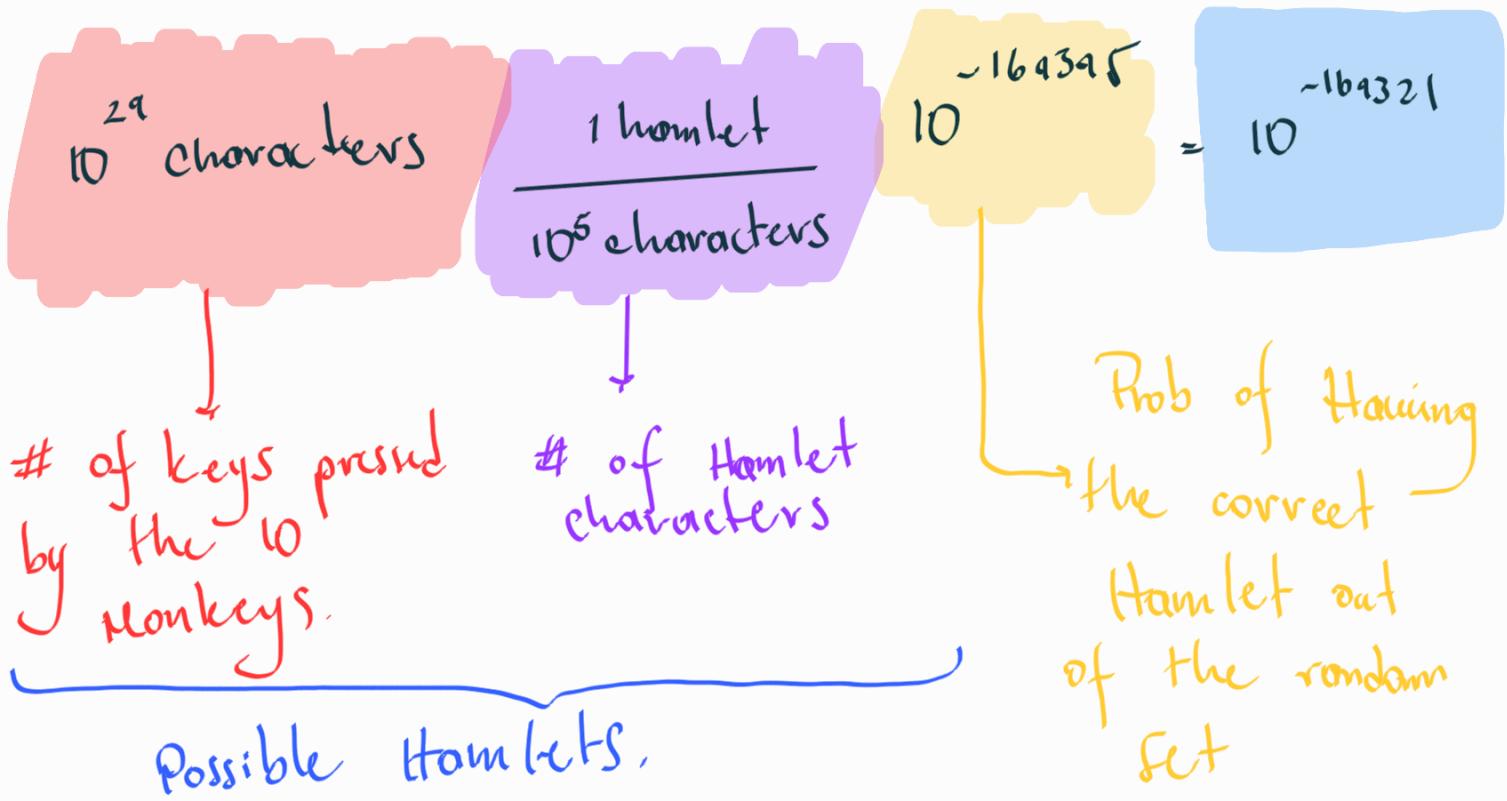
$$10^{19} \text{ keys} \times 10^0 \text{ monkeys} = 10^{29} \text{ keys.}$$

so

$$10^{29} \text{ keys typed out} \times \frac{1 \text{ hamlet}}{10^5 \text{ characters}}$$

$$= 10^{24} \text{ possible "Hamlets".}$$

As we saw before, the "Chance" of getting a Hamlet is  $10^{-164345}$ , so





### 3 Ideal Gas in the Canonical ensemble

let us start by considering a macrostate  $M = (T, V, N)$

\*The canonical ensemble is characterised by its constant temperature.

Now the energy of the system changes, but its average and fluctuations are well defined and known.

$$P(\mu) = P(\{\vec{p}_i, \vec{q}_i\}) = \frac{1}{Z} \exp \left[ -\beta \sum_{i=1}^N \frac{p_i^2}{2m} \right] \times \begin{cases} 1 & \text{for } \{\vec{q}_i\} \in \text{Box} \\ 0 & \text{otherwise} \end{cases}$$

↓  
Canonical partition function.

as we do not have a potential term:

$$Z(T, V, N) = \frac{1}{N!} \int \frac{d\vec{q}_1 d\vec{q}_2 \dots d\vec{q}_N d\vec{p}_1 \dots d\vec{p}_N}{h^{3N}} e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}}$$

↓  
undistinguishable particles      ↓  
phase space minimum volume

$$= \frac{1}{N! h^{3N}} \underbrace{\left[ \prod_{i=1}^N \int_{-\infty}^{\infty} d\vec{p}_i e^{-\beta \frac{p_i^2}{2m}} \right]}_{\text{Gaussian integral.}} \times \prod_{i=1}^N \int_{\text{Box}} d\vec{q}_i$$

↓  
 $V^N$

Gaussian integral.

$$= (2\pi m k_B T)^{3N/2}$$

$$= \frac{V^N}{N!} \left( \frac{2\pi m k_B T}{h^2} \right)^{3N/2} = \frac{1}{N!} \left( \frac{V}{\lambda(T)^3} \right)^N \quad \text{with} \quad \lambda = \frac{h}{\sqrt{2\pi m k_B T}}$$

Once we have the partition function,  $Z$ , we can compute the thermodynamics.

The free energy

$$F = -k_B T \ln(Z) = -k_B T \ln \left[ \frac{1}{N!} \left( \frac{V}{\lambda^3} \right)^N \right]$$

$$= -k_B T \left[ N \ln(V) - \underbrace{N \ln(N) + N}_{\text{Stirling formula.}} - \frac{3N}{2} \ln(\lambda^2) \right]$$

$$= -k_B T N \left[ \ln \left( \frac{V}{N} \right) - 1 - \frac{3}{2} \ln \left( \frac{\hbar^2}{2\pi m k_B T} \right) \right]$$

$$-1 = -\ln(e)$$

$$= \ln(1/e)$$

$$= -k_B T N \left[ \ln \left( \frac{Ve}{N} \right) + \frac{3}{2} \ln \left( \frac{2\pi m k_B T}{\hbar^2} \right) \right]$$

let us use this expression to compute more thermodynamics.

We have that  $F$  is an equation of state;

$$dF = -SdT - PdV + \mu dN$$

$$\hookrightarrow -S = \frac{\partial F}{\partial T} \Big|_{V,N}; -P = \frac{\partial F}{\partial V} \Big|_{T,N}; \mu = \frac{\partial F}{\partial N} \Big|_{T,V}$$

$$-S = \frac{\partial F}{\partial T} \Big|_{V,N} = -Nk_B \left[ \ln \left( \frac{V_e}{N} \right) + \frac{3}{2} \ln \left( \frac{2\pi m k_B T}{h^2} \right) \right]$$

$-Nk_B T \left( \frac{3}{2} \frac{1}{T} \right)$

$\underbrace{\frac{3}{2} Nk_B T = E \equiv \text{internal energy}}$

$= \frac{F - E}{T}$

To get the equation of state, we might use the equation for pressure.

$$-P = \frac{\partial F}{\partial V} \Big|_{T,N} = -\frac{Nk_B T}{V}$$

$$PV = Nk_B T$$

lastly, the chemical potential.

$$\mu = \frac{\partial F}{\partial N} \Big|_{T,V} = -k_B T \left[ \ln \left( \frac{V_e}{N} \right) + \frac{3}{2} \left( \frac{2\pi m k_B T}{h^2} \right) \right] + k_B T N \left( \frac{f_1}{N} \right)$$

$$= \frac{E}{N} + k_B T = \frac{E - TS + PV}{N}$$

4 The ideal gas in the isobaric ensemble.

We have a similar situation as before, but now one has to add the correct Lagrange multiplier.

$$P(\{\vec{p}_i, \vec{q}_i\}, V) = \frac{\exp \left[ -\beta \sum_{i=1}^N \frac{p_i^2}{2m} - \beta PV \right]}{\mathcal{Z}} \times \begin{cases} 1 & \text{for } \{\vec{q}_i\} \in \text{Box}(V) \\ 0 & \text{otherwise} \end{cases}$$

Now, the partition function is a bit different.

$$\mathcal{Z}(N, T, P) = \frac{1}{N! h^{3N}} \int_0^\infty dV \int \frac{d\vec{q}_1 d\vec{q}_2 \dots d\vec{q}_N d\vec{p}_1 \dots d\vec{p}_N}{h^{3N}} e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}} e^{-\beta PV}$$

↓  
Undistinguishable particles

↓  
phase space  
minimum volume

$$= \frac{1}{N! h^{3N}} \underbrace{\left[ \prod_{i=1}^N \int_{-\infty}^{\infty} d\vec{p}_i e^{-\beta \frac{p_i^2}{2m}} \right] \times \int_0^\infty dV e^{-\beta PV} \prod_{i=1}^N \int_{\text{Box}(V)} d\vec{q}_i}_{\text{Gaussian integral.}}$$

$$= \frac{1}{\lambda(T)^{3N}} \left[ \frac{1}{N!} \int_0^\infty dV V^N e^{-\beta PV} \right]$$

$$U = \beta P V \quad \frac{U}{\beta P} = V$$

$$dU = \beta P dV$$

$$\frac{du}{\beta P} = dV$$

$$= \frac{1}{\lambda(\tau)^{3N}} \left[ \frac{1}{N!} \int_0^\infty \frac{du}{\beta P} \left( \frac{u}{\beta P} \right)^N e^{-u} \right]$$

$$= \frac{1}{\lambda(\tau)^{3N}} \left[ \frac{1}{N!} \left( \frac{1}{\beta P} \right)^{N+1} \int_0^\infty du u^N e^{-u} \right]$$

$\Gamma(N+1) = N!$

$$\mathcal{Z}(N, T, P) = \frac{1}{\lambda(\tau)^{3N} (\beta P)^{N+1}}$$

Then, the free energy

$$G = -k_B T \ln \mathcal{Z} = k_B T \ln \left( \lambda(\tau)^{3N} (\beta P)^{N+1} \right)$$

$$= k_B T \left[ 3N \ln(\lambda) + (N+1) \ln(\beta P) \right]$$

$$= k_B T \left[ 3N \ln(\lambda) + (N+1) (\ln(\beta) + \ln(P)) \right]$$

$$\begin{aligned}
&= k_B T N \left[ 3 \ln(\lambda) + \ln(P) + \ln(\beta) \right] \\
&\quad + k_B T \ln(\beta) + k_B T \ln(P) \\
&= k_B T N \left[ \frac{3}{2} \ln \left( \frac{h^2 \beta}{2\pi m} \right) + \ln(P) + \ln(\beta) \right] \\
&\quad + k_B T \ln(\beta) + k_B T \ln(P) \\
&= k_B T N \left[ \ln(P) - \frac{5}{2} \ln(k_B T) + \frac{3}{2} \ln \left( \frac{h^2}{2\pi m} \right) \right] \\
&\quad + k_B T \ln(P\beta)
\end{aligned}$$

in the same way we did in the Canonical case.

$$dG = -SdT + VdP + \mu dN.$$

The Volume

$$\begin{aligned}
V &= \left. \frac{\partial G}{\partial P} \right|_{T, N} = \frac{N k_B T}{P} + \frac{k_B T}{P} \\
&= \frac{(N+1) k_B T}{P} \quad \text{as } N \gg 1 \\
&\approx \frac{N k_B T}{P} \rightarrow PV = N k_B T
\end{aligned}$$

The enthalpy  $H = \langle E + PV \rangle$

$$H = -\frac{\partial \ln Z}{\partial \beta} = \frac{-\partial}{\partial \beta} (\rho G) \\ = \frac{-\partial}{\partial \beta} \left( \frac{G}{k_B T} \right)$$

$$= \frac{5}{2} N k_B T - \frac{k_B T}{P}$$

The same argument as before

$$N k_B T \gg \frac{1}{P}$$

$$H \approx \frac{5}{2} k_B T N$$

so

$$C_P = \frac{dH}{dT} = \frac{5}{2} N k_B$$

## 5 Ideal Gas in the Grand Canonical ensemble.

a) the Grand Partition function

$$\Xi(T, \mu, V) = \sum_{N=0}^{\infty} e^{\beta \mu N} \frac{1}{N!} \underbrace{\int \prod_i^N \left( \frac{d^3 \vec{q}_i d^3 \vec{p}_i}{h^3} \right) \exp \left[ -\beta \sum_i \frac{p_i^2}{2m} \right]}_{\text{canonical Distribution}}$$

$$= \sum_{N=0}^{\infty} e^{\beta \mu N} \frac{1}{N!} \left( \frac{V}{\lambda^3} \right)^N \quad \text{with} \quad \lambda = \frac{h}{\sqrt{2\pi m k_B T}}$$

$$= \exp \left( e^{\beta \mu} \frac{V}{\lambda^3} \right) \quad \rightarrow \text{Using } e^x = \sum_{n=0}^{\infty} \frac{x^n}{n!}$$

So, the grand potential

$$G(T, \mu, V) = -k_B T \ln(\Xi) = -k_B T e^{\beta \mu} \frac{V}{\lambda^3}$$

and using  $G = E - TS - \mu N = -PV$ .

$$P = -\frac{\partial G}{\partial V} = -\left. \frac{\partial G}{\partial V} \right|_{\mu, T} = \frac{k_B T e^{\beta \mu}}{\lambda^3}$$

and the number of particles;

$$N = -\left. \frac{\partial G}{\partial \mu} \right|_{T, V} = \frac{e^{\beta \mu} V}{\lambda^3}$$

from these equations one can get the equation of state

$$P = k_B T N / V$$

from the equation of  $N$

$$\frac{\lambda^3 N}{V} = e^{\beta \mu} \rightarrow \mu = k_B T \ln \left( \frac{\lambda^3 N}{V} \right) = k_B T \ln \left( \frac{P \lambda^3}{k_B T} \right)$$

# Grand partition Function of a two level system.

$$\begin{aligned} \Xi &= \sum_N e^{\mu N} \sum_i e^{-\beta E_i} \\ &= e^{\mu \times 0} \cdot e^{-\beta E_0} + e^{\mu - \beta E_1} e^{-\beta E_1} + e^{\mu - \beta E_2} e^{-\beta E_2} \\ &\quad \text{Not occupied } \downarrow \quad \text{occupied } \downarrow \quad \text{occupied } \downarrow \\ &\quad E_0 = 0 \quad \text{Not excited } \downarrow \quad \text{excited } \downarrow \\ &\quad \text{state } E_1 = 0 \quad \epsilon_2 = \epsilon \\ &= 1 + \lambda \times 1 + \lambda \times e^{-\beta \epsilon} \end{aligned}$$

$$= 1 + \lambda + \lambda e^{-\beta \epsilon}$$

the number of particles

$$\langle N \rangle = \frac{0 \times 1 + (1) \times (\lambda + \lambda e^{-\beta \epsilon})}{\Xi} = \frac{\lambda + \lambda e^{-\beta \epsilon}}{\Xi}$$

the number of particles with energy  $\epsilon$

$$\langle N(\epsilon) \rangle = \frac{0 \times 1 + 0 \times \lambda + 1 \times \lambda e^{-\beta \epsilon}}{\Xi} = \frac{\lambda e^{-\beta \epsilon}}{\Xi}$$

the average of energy

$$\langle \epsilon \rangle = \frac{0 \times 1 + 0 \times \lambda + \epsilon \times \lambda e^{-\beta \epsilon}}{\Xi} = \epsilon \langle N(\epsilon) \rangle = \frac{\epsilon \lambda e^{-\beta \epsilon}}{\Xi}$$

thus,

$$\Xi = \sum_N e^{\mu N} \sum_i e^{-\beta E_i} = e^0 e^{-\beta \epsilon} + e^{\mu 1} e^{-\beta \epsilon} + e^{\mu 1} e^{-\beta \epsilon_2} + e^{\mu 2} e^{-\beta \epsilon_3}$$

$$= 1 + \lambda + \lambda e^{-\beta \epsilon} + \lambda^2 e^{-2\beta \epsilon}$$

└ The only new term including two particles.

